A deep-atmosphere extension in a nonhydrostatic dynamical core: formulation and idealized tests

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Abstract: ~~-~~We have developed a deep-atmosphere extension within the GMCORE (Grid-Model dynamical CORE) using a dry terrain-following mass-based coordinate. The new dycore retains the full three-dimensional Coriolis force, accounts for the height dependence of gravity, and introduces radial metric corrections into the discrete operators, while preserving the main structure of the original solver. Numerical consistency is further examined through convergence and balanced-state stability tests. Two idealized test cases are used for evaluation: a baroclinic wave test and a tropical cyclone test. In the standard Earth-radius baroclinic wave experiment, the deep- and shallow-atmosphere solutions are similar, whereas in the reduced-radius X20 ~~configuration~~experiment the GMCORE deep-atmosphere dycore reproduces the clear separation reported in earlier studies. In the tropical-cyclone experiment, the deep-atmosphere ~~configuration~~formulation has little effect on storm intensity and warm~~-~~-core structure, but leads to a systematic south-westward trajectory shift relative to the shallow-atmosphere simulation. The trajectory difference is linked to an organised inner asymmetry. This includes a dipole-like pattern connected to ~~$2\Omega cos\phi \frac{\partial w}{\partial y}$~~ $2\Omega \cos\varphi \frac{\partial w}{\partial y}$ and a low-level easterly anomaly. These results ~~demonstrate that~~ support the numerical consistency of the GMCORE ~~can be extended to a~~ deep-atmosphere ~~dycore~~implementation and ~~provide a useful work for idealized tropical cyclone trajectory sensitivity and deep-atmosphere dynamics~~suggest that nontraditional Coriolis effects can influence low-latitude vortex motion, although a complete dynamical attribution requires further diagnosis.

1 Introduction

Modern global atmospheric models are extending their upper boundaries into the mesosphere and lower thermosphere, motivating the use of deep-atmosphere dynamics. High-top models, for example, the WACCM (Whole Atmosphere Community Climate Model, Version 6) extends to approximately 140 km (Gettelman et al., 2019), while its thermosphere-ionosphere extension version (WACCM-X) extends to nearly 500 km (Liu et al., 2018). The HIAMCM models (Becker and Vadas, 2020), with a nonhydrostatic correction, achieves the explicit simulation of gravity waves reaching the thermosphere. In addition, global nonhydrostatic dycores such as NICAM (Satoh et al., 2008), the Met Office model ENDGame (Wood et

al., 2014), MPAS-A (Skamarock et al., 2021), ICON (Zängl et al., 2015; Borchert et al., 2019), NUMA (Kelly et al., 2026) and CliMA (Yatunin et al., 2026) have introduced deep-atmosphere modifications to support the full equations and upper atmosphere simulations.

As models extend into higher layers, the shallow-atmosphere approximation becomes less accurate because the atmosphere vertical scale becomes not negligible compared to earth radius (White and Bromley, 1995; Akmaev, 2011). In practical modelling, the shallow-atmosphere approximation is commonly accompanied with traditional approximation, by neglecting horizontal component of the planetary rotation and curvature terms in order to preserve angular momentum (Phillips 1966; Vallis, 2017). While these omissions are usually small in the lower atmosphere, their importance increases with height and may become more significant in simulations involving large-scale flow and wave propagation in the upper atmosphere (Baldwin and Dunkerton, 2001; Fritts and Alexander, 2003; Kim et al., 2003).

The deep-atmosphere equations ~~can be viewed as an extension of~~are generally considered to be a more fundamental set than the traditional shallow-atmosphere equations, which relax a set of approximations involving spherical geometry, gravity and planetary rotation and change normal mode structure on vertical (Thuburn~~, 2002a~~ et al., 2002; White et al., 2005). Among these differences, when the model top is not very high, the nontraditional Coriolis terms (NCTs)—which are neglected in the shallow-atmosphere equations—often play a significant role in the atmospheric motion (Colin de Verdière and Schopp, 1994; Marshall et al., 1997; Gerkema et al., 2008). Standard shallow-atmosphere equations project the planetary rotation vector onto the local vertical and neglect the horizontal component, simplifying the full rotation to the traditional Coriolis term ~~$2\Omega\sin\phi$~~ $2\Omega \sin\varphi$. However, a deep-atmosphere treatment retains the full rotation vector, including the components proportional to ~~$2\Omega\cos\phi$~~$2\Omega \cos\varphi$, which couple horizontal and vertical motions.

A growing body of work has demonstrated that such NCTs can be numerically significant, especially in the tropics where the horizontal component of planetary rotation becomes comparable to other dynamical terms and affects wave dispersion and balance relationships (Gerkema et al~~.~~, 2008; Ong and Roundy, 2019; Igel and Biello, 2020). Recent studies have further shown that retaining the full Coriolis acceleration can influence the stability characteristics of atmospheric wave modes (Chew et al., 2023). By eliminating the Coriolis simplifications, the deep-atmosphere equations more truly represent rotational influences across latitudes, ensuring a consistent description of vortical and wave motions in a stratified, rotating atmosphere.

In this study, we develop a deep-atmosphere version based on GMCORE (Grid-Model dynamical CORE). GMCORE is a fully compressible nonhydrostatic dycore formulated on a latitude–longitude C grid with a dry terrain-following mass-based coordinate (Li et al., 2020; Li and Dong, 2024), and has been applied to Mars climate simulation ~~researches~~studies (Dong et al., 2024; Liu et al., ~~2025~~2026). Our goal is to extend the original shallow-atmosphere dycore so that it retains the full three-dimensional Coriolis force, allows gravity to vary with height, and accounts for the radial dependence of metric terms in the discrete operators, while preserving the basic structure of the existing solver. An important feature of the present formulations is that the deep-atmosphere extension is represented within the dry terrain-following mass-based coordinate while most nonhydrostatic models use height-based vertical coordinate (Wood and Staniforth, 2003).

To evaluate the new dycore, we use two idealized test cases. The first is the deep baroclinic-wave test case (Ullrich et al, 2014), which is used here primarily as a validity check case. As emphasized by Skamarock et al. (2021), the interpretation of the X20 experiment is not straightforward because part of the reported deep-shallow differences may arise from differences in the analytic initial states rather than directly from the temporal integration of the two equation sets (Staniforth and White, 2011). We therefore use this case mainly to assess whether the GMCORE deep-atmosphere dycore reproduces the canonical numerical behaviour reported in the literature. The second test is an idealized tropical-cyclone test case (Reed and Jablonowski, 2011), which is used to examine the performance of the dycore in a strongly rotating, vertically deep, low-latitude vortex. Compared with the baroclinic-wave test, this case is also useful for discussing how the retention of the full Coriolis vector may influence storm trajectory and asymmetric structure.

The paper is organized as follows. Section 2 describes the governing equations of the deep-atmosphere GMCORE and its numerical implementation in the dry terrain-following mass-based coordinate. Section 3 presents results from the idealized baroclinic-wave and tropical-cyclone tests. Section 4 summarizes the main findings and discusses their implications and limitations.

2. Model description

2.1. The deep-atmosphere governing equations

Following ~~Wood and Staniforth (2003) and other~~the vector-invariant formulation based on the Lamb transformation (Gassmann and Herzog, 2008), together with the deep-atmosphere formulations ~~(~~of Wood and Staniforth (2003), Borchert et al~~.,~~. (2019~~;~~), and Skamarock et al~~.,~~. (2021), GMCORE solves the fully compressible, nonhydrostatic deep-atmosphere equations in a dry terrain-following mass-based coordinate in vector-invariant form. Compared to the shallow-atmosphere equations, the deep extension introduces three important changes: the constant planetary radius a is replaced by the local radial distance $r = a + z$, the constant gravity ~~G~~g_0 becomes altitude-dependent $g(r) = g_0 \left(\frac{a}{r}\right)^2$, and the full Coriolis and spherical curvature terms are retained in the momentum equations. The resulting GMCORE governing equations are written as follows:

~~$$\frac{\partial \mathbf{V}}{\partial t} = -\frac{\xi}{\zeta_a}\mathbf{k}\times\mathbf{V} - \nabla_\eta^d K - \dot{\eta}\frac{\partial \mathbf{V}}{\partial \eta} - \frac{\rho_d}{\rho}\left(\frac{1}{\rho_d}\nabla_\eta^d p + \frac{\partial p}{\partial \pi}\nabla_\eta^d \phi\right) - 2\Omega\cos\varphi\, w\mathbf{i} - \frac{w}{r}\mathbf{V} + F_V \quad (1)$$~~

~~$$\frac{\partial w}{\partial t} = -\mathbf{V}\cdot\nabla_\eta^d w - \dot{\eta}\frac{\partial w}{\partial \eta} - g(r) + \frac{\rho_d}{\rho}G\left(\frac{r}{a}\right)^2\frac{\partial p}{\partial \pi} + 2\Omega\cos\varphi\, u + \frac{u^2+v^2}{r} + F_w \quad (2)$$~~

~~$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}\right) = -\nabla_\eta^d\cdot\left(\frac{\partial \pi}{\partial \eta}\mathbf{V}\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\right) \quad (3)$$~~

~~$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}\theta_m\right) = -\nabla_\eta^d\cdot\left(\frac{\partial \pi}{\partial \eta}\mathbf{V}\theta_m\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\theta_m\right) + F_{\theta_m} \quad (4)$$~~

~~$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}q_i\right) = -\nabla_\eta^d\cdot\left(\frac{\partial \pi}{\partial \eta}\mathbf{V}q_i\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}q_i\right) + F_{q_i},\ i = 1,\ldots,N \quad (6)$$~~

~~$\frac{\partial \phi}{\partial t} = -\mathbf{V} \cdot \nabla_{\eta}^{d}\phi - \dot{\eta}\frac{\partial \phi}{\partial \eta} + wg(r) + F_{\phi}$~~ (7)

~~where~~ ~~$\mathbf{V} = (u\mathbf{i}, v\mathbf{j})$~~

$$\frac{\partial \mathbf{V}}{\partial t} = -\xi_a \mathbf{k} \times \mathbf{V} - \nabla_{\eta}^{d} K - \dot{\eta}\frac{\partial \mathbf{V}}{\partial \eta} - L\left(\frac{1}{\rho_d}\nabla_{\eta}^{d} p + \left(\frac{r}{a}\right)^2 \frac{\partial p}{\partial \pi}\nabla_{\eta}^{d}\phi\right) - 2\Omega\cos\varphi\, w\mathbf{i} - \frac{w}{r}\mathbf{V} + S_V, \quad (1)$$

$$\frac{\partial w}{\partial t} = -\mathbf{V} \cdot \nabla_{\eta}^{d} w - \dot{\eta}\frac{\partial w}{\partial \eta} - g(r) + \frac{\rho_d}{\rho} g_0 \left(\frac{r}{a}\right)^2 \frac{\partial p}{\partial \pi} + 2\Omega\cos\varphi\, u + \frac{u^2 + v^2}{r} + S_w\,, \quad (2)$$

$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}\right) = -\nabla_{\eta}^{d} \cdot \left(\frac{\partial \pi}{\partial \eta}\mathbf{V}\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\right), \quad (3)$$

$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}\theta_m\right) = -\nabla_{\eta}^{d} \cdot \left(\frac{\partial \pi}{\partial \eta}\mathbf{V}\theta_m\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\theta_m\right) + S_{\theta_m}\,, \quad (4)$$

$$\frac{\partial}{\partial t}\left(\frac{\partial \pi}{\partial \eta}q_i\right) = -\nabla_{\eta}^{d} \cdot \left(\frac{\partial \pi}{\partial \eta}\mathbf{V}q_i\right) - \frac{\partial}{\partial \eta}\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}q_i\right) + S_{q_i},\ i = 1, \ldots, N\,, \quad (5)$$

$$\frac{\partial \phi}{\partial t} = -\mathbf{V} \cdot \nabla_{\eta}^{d}\phi - \dot{\eta}\frac{\partial \phi}{\partial \eta} + wg(r) + S_{\phi}\,. \quad (6)$$

Here $\mathbf{V} = (u, v)$ is the wind vector, ~~$\dot{\eta}$~~$\dot{\eta}$ is the vertical coordinate velocity, ~~∇_{η}^{d}~~∇_{η}^{d} is the deep-atmosphere horizontal differential operator, $K = (u^2 + v^2)$~~/~~/2 is the horizontal kinetic energy, ~~π~~π is the hydrostatic dry air mass, Ω is the planetary angular rotation ~~rate of planetary~~velocity, $L = \frac{\rho_d}{\rho} = \frac{1}{1+q_m}$ is the moist contribution where the q_m represent total moist dry mixing ratio, ρ_d and ρ denote the dry air and total density, ϕ is the geopotential and ~~F~~S represents the unresolved physical tendencies. The deep-atmosphere absolute vorticity ξ_a is correspondingly written as

~~$\xi_a = f + \frac{1}{r^2\cos\varphi}\left[\frac{\partial (rv)}{\partial \lambda} - \frac{\partial (ru\cos\varphi)}{\partial \varphi}\right]$~~ (8)

$$\xi_a = f + \frac{1}{r^2\cos\varphi}\left[\frac{\partial (rv)}{\partial \lambda} - \frac{\partial (ru\cos\varphi)}{\partial \varphi}\right]. \quad (7)$$

The layer-averaged mass continuity and thermodynamic equations are then expressed in conservative form as

~~$\frac{\partial}{\partial t}(\delta\pi) = -\nabla_{\eta}^{d} \cdot (\delta\pi\mathbf{V}) - \delta\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\right)$~~ (9)

~~$\frac{\partial}{\partial t}(\delta\pi\theta_m) = -\nabla_{\eta}^{d} \cdot (\delta\pi\mathbf{V}\theta_m) - \delta\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\theta_m\right) + F_{\theta_m}$~~ (10)

$$\frac{\partial}{\partial t}(\delta\pi) = -\nabla_{\eta}^{d} \cdot (\delta\pi\mathbf{V}) - \delta\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\right), \quad (8)$$

$$\frac{\partial}{\partial t}(\delta\pi\theta_m) = -\nabla_{\eta}^{d} \cdot (\delta\pi\mathbf{V}\theta_m) - \delta\left(\frac{\partial \pi}{\partial \eta}\dot{\eta}\theta_m\right) + S_{\theta_m}\,, \quad (9)$$

where θ_m is the modified potential temperature including the thermodynamic effects of moisture species (Lauritzen et al., 2018~~).~~) and $\delta(\cdot) = (\cdot)_{k-\frac{1}{2}} - (\cdot)_{k+\frac{1}{2}}$ denotes the vertical increment. In this form it is already evident that the deep-atmosphere

extension changes not only the explicit dynamical terms, such as $g(r)$ and the curvature contributions, but also the geometric structure of the transport operators that must be discretized.

For the numerical discretization, the operator modifications follow naturally from the replacement of the constant radius by the local radial distance on the generalized η-surfaces. The horizontal gradient and pseudo-horizontal divergence therefore become

$$\nabla_\eta^d \alpha = \frac{1}{a\cos\varphi}\frac{a}{r}\frac{\partial \alpha}{\partial \lambda}\mathbf{i} + \frac{1}{a}\frac{a}{r}\frac{\partial \alpha}{\partial \varphi}\mathbf{j}, \tag{10}$$

~~$$\nabla_{\tilde{\eta}}^d \alpha = \frac{1}{a\cos\varphi}\frac{a}{r}\frac{\partial \alpha}{\partial \lambda}\mathbf{i} + \frac{1}{a}\frac{a}{r}\frac{\partial \alpha}{\partial \varphi}\mathbf{j}$$~~ ~~(11)~~

~~$$\nabla_{\tilde{\eta}}^d \cdot \mathbf{B} \equiv \frac{1}{a\cos\phi}\left(\frac{\partial}{\partial \lambda}\left(\frac{a}{r}B_\lambda\right) + \frac{\partial}{\partial \phi}\left(\frac{a}{r}\cos\phi\, B_\phi\right)\right)$$~~ ~~(12)~~

$$\nabla_\eta^d \cdot \mathbf{B} \equiv \frac{1}{a\cos\varphi}\left(\frac{\partial}{\partial \lambda}\left(\frac{a}{r}B_\lambda\right) + \frac{\partial}{\partial \phi}\left(\frac{a}{r}\cos\varphi\, B_\phi\right)\right), \tag{11}$$

where α and $\mathbf{B}$ are arbitrary scalar and vector. Here the factor ~~a/r~~ (a/r) is part of the local horizontal metric on the η-surface. It cannot, in general, be treated as a purely layer-dependent multiplicative factor, because the terrain-following η-surface may still carry horizontal variations of r. Consequently, the deep metric does not generally commute with the horizontal derivatives, and the metric factor must remain attached to the face fluxes or, equivalently, inside the deep divergence operator used by the discrete scheme.

It is also necessary to distinguish the pseudo-horizontal divergence appearing in the transport operators from the true three-dimensional divergence D_3 of the velocity field. Following Wood and Staniforth (2003), the exact relationship is

~~$$D_3 = \nabla_{\tilde{r}}^d \cdot \boldsymbol{V} + \frac{\partial w}{\partial r} + \frac{2w}{r}$$~~ ~~(13)~~

$$D_3 = \nabla_r^d \cdot \mathbf{V} + \frac{\partial w}{\partial r} + \frac{2w}{r}. \tag{12}$$

In the continuous equations, the factor 1~~$/r$~~ may be understood either as part of the modified horizontal mass flux or as part of the deep horizontal divergence operator (Wood and Staniforth, 2003). These two understandings are equivalent only when the metric factor remains inside the horizontal derivative and is evaluated at the same face location as the flux. For the GMCORE mass-based discretization, this means that the local radius must be diagnosed and interpolated consistently to the staggered positions required by the horizontal mass flux, pressure-gradient force and advection process.

The vertical transport operator requires a different geometric correction. Because the upper and lower interfaces of a control volume lie on spherical shells of different radii, the vertical mass and tracer fluxes must be weighted by the shell-area factor ~~$(r/a)^2$~~ $(r/a)^2$ before calculating the vertical flux divergence. This shell-area weighting belongs specifically to the vertical transport, whereas the horizontal transport ~~is controlled by the~~uses a local edge modification ~~and~~, i.e., the radius-dependent scaling of each horizontal grid-edge length by the local factor (r/a), together with the deep ~~version~~ atmosphere horizontal

divergence operator on the staggered ~~grids.~~grid. It is crucial for both conservation and numerical stability to consistently use appropriate and consistent discrete deep-atmosphere geometric extensions in calculations involving pressure gradients, horizontal divergence, vertical flux divergence, and advection.

2.2. Mass-based hybrid vertical coordinate

~~We use the dry mass terrain-following vertical coordinate (Kasahara, 1974; Simmons and Burridge, 1981), denoted as η.~~ As demonstrated by Wood and Staniforth (2003), a mass-based vertical coordinate requires adjustment for the deep-atmosphere. They denoted s as the generalized vertical coordinate. In contrast to Laprise (1992), the deep-atmosphere mass coordinate chooses $\rho r^2 \frac{\partial r}{\partial s}$ as a constant instead of $\rho a^2 \frac{\partial r}{\partial s}$. This makes the definition of hydrostatic pressure coordinate turn to a more physical description of the atmosphere mass above a cell.

We use the dry mass π to substitute the generalized coordinate s and introduce terrain-following coordinate η (Kasahara, 1974; Simmons and Burridge, 1981):

$$\pi(x, y, \eta, t) = \eta_A \pi_0 + \eta_B \pi_s(x, y, t), \tag{13}$$

where π_0 is the constant default pressure (usually, 1000 hPa) and π_s represents the surface mass. The coefficients η_A and η_B prescribe a set of hybrid coordinate defined at the model interfaces. The uppermost interface defines the model top, whereas the lowest interface follows the lower boundary. The specific vertical resolution and model-top setting used in each experiment are given in Sect. 3.

An obvious difference in this deep-atmosphere mass-coordinate dycore is the diagnostic calculation of the radial distance r. In the η-coordinate, to determine the radial distance r without relying on complex iterative solvers, GMCORE directly integrates the exact hydrostatic differential equation. The deep-atmosphere hydrostatic relation is given by:

~~$$\frac{\partial \pi}{\partial \eta} = -\rho g_0 \left(\frac{r}{a}\right)^2 \frac{\partial r}{\partial \eta} \tag{14}$$~~

By rearranging the terms to isolate the geometric dependence, we obtain:

~~$$r^2 \frac{\partial r}{\partial \eta} = -\frac{a^2}{\rho g_0} \frac{\partial \pi}{\partial \eta} \tag{15}$$~~

$$r^2 \frac{\partial r}{\partial \eta} = -\frac{a^2}{\rho g_0} \frac{\partial \pi}{\partial \eta}. \tag{15}$$

Recognizing the exact differential form $r^2 \partial r dr = \frac{1}{3} \partial(r^3) d(r^3)$, this relation can be analytically integrated from the planetary surface (where ~~$r = r_s =$~~ r_s and $\pi = \pi_s$) upwards to any given model level k:

~~$$r_k^3 = r_s^3 - \frac{3a^2}{g_0} \int_{\pi_k}^{\pi_s} \frac{1}{\rho} d\pi \tag{16}$$~~

In the numerical implementation of the dynamical core, this integral is evaluated layer-by-layer from the surface upward. Let ~~$\Delta\pi$~~$\Delta\pi$ represents the hydrostatic pressure depth of a layer, which is directly linked to the layer mass ~~Δm~~Δm. The radial distance $r_{k+\frac{1}{2}}$ at a specific half model level is diagnosed through a discrete cumulative summation of the density-weighted layers:

~~$$r_{k+\frac{1}{2}} = \left(r_s^3 + \frac{3a^2}{g_0} \sum_{l=k}^{k_{sfc}} \frac{\Delta\pi_k}{\bar{\rho}_l} \right)^{1/3} \quad (17)$$~~

$$r_{k+\frac{1}{2}} = \left(r_s^3 + \frac{3a^2}{g_0} \sum_{l=k}^{k_{sfc}} \frac{\Delta\pi_l}{\bar{\rho}_l} \right)^{1/3}, \quad (17)$$

where~~Δ~~ $\Delta\pi_k = \pi_{k+\frac{1}{2}} - \pi_{k-\frac{1}{2}}$ is the dry-air weight between two-layer interfaces, the ~~$\bar{\rho}_l$~~$\bar{\rho}_l$ is the density at the layer center (evaluated at full level).

~~Furthermore, the choice of a dry mass terrain-following coordinate offers advantages for coupling the dycore with physical parameterizations. Because the vertical coordinate inherently dictates the atmosphere mass above each model level, the calculation of optical depth for shortwave radiation becomes straightforward and direct. This bypasses the complex density-height conversions required in purely geometric height-based models, significantly enhancing the thermodynamic consistency of short wave heating (Borchert et al., 2019).~~

The present deep-atmosphere geometry uses a spherical reference planet. Orography can enter through the lower-boundary surface geopotential and the terrain-following η coordinate, but the idealized experiments reported here use zero orography. Deviations of the geoid from a perfect sphere, such as oblateness, are not represented in the present version.

2.3. Spatial discretization and grid staggering

~~In practical numerical implementation, the discrete spatial discretization must strictly matches the grid geometry expansion. In shallow-atmosphere dycore, the horizontal grid cell area is assumed constant throughout all vertical levels. On the contrast, in the deep-atmosphere dycore, the discrete finite-volume cell area $A(r)$ and grid edge lengths $\Delta l(r)$ physically change as the altitude increases. Following the illustration and discretization shown in Borchert et al. (2019) and Skamarock et al. (2021), GMCORE explicitly incorporates this radial scaling into its spatial operators. Specifically, the cell area and edge lengths at a given radial distance r are scaled by geometric factors of $(r/a)^2$ and (r/a), respectively, relative to their reference values at the planetary surface a.This guarantees the conservation of total atmosphere mass and systematically helps prevent the spurious generation of global angular momentum.~~

~~GMCORE utilizes a regular latitude-longitude grid organized on an Arakawa C-grid staggering. This staggering optimally represents inertia-gravity wave dispersion characteristics and suppresses computational mode. For advection process, GMCORE employs a modified flux-form semi-Lagrangian (FFSL) advection scheme SWIFT with a third-order polynomial~~

~~reconstruction and flux limiters, to enable large advection time step size and maintain the monotonicity of tracers (Lin and Rood, 1996; Bendall and Kent, 2025).~~

The regular latitude–longitude grid remains widely used in several established atmospheric models (Williamson, 2007; Staniforth and Thuburn, 2012) and is still under continuous development (Lima, 2024). Its orthogonal geometry aligns the coordinate directions with the zonal and meridional flow components, facilitating the formulation of fluxes and differential operators. GMCORE retains this grid structure as a practical choice that is consistent with its existing numerical framework. However, the convergence of meridians toward the poles leads to increasingly small zonal grid spacing and restrictive stability constraints, requiring the polar filtering treatment described in Sect. 2.5.

GMCORE uses an Arakawa C-grid staggering in the horizontal and a Lorenz staggering in the vertical, as illustrated in Fig. 1a. The full levels are indexed by $k = 1, \ldots, n_{lev}$ from the model top downward. Dry mass and thermodynamic variables are defined at the full levels. The Arakawa C-grid is well suited for representing geostrophic adjustment and inertia–gravity waves, while the Lorenz vertical staggering has a relatively simple structure and facilitates conservative discretization (Arakawa and Lamb, 1977; Adcroft et al, 1997; Simmons and Burridge, 1981). The horizontal velocity components u and v are also defined at the full levels, while retaining their horizontal C-grid staggering. The vertical velocity w and vertical coordinate velocity $\dot{\eta}$ are defined at the half levels.

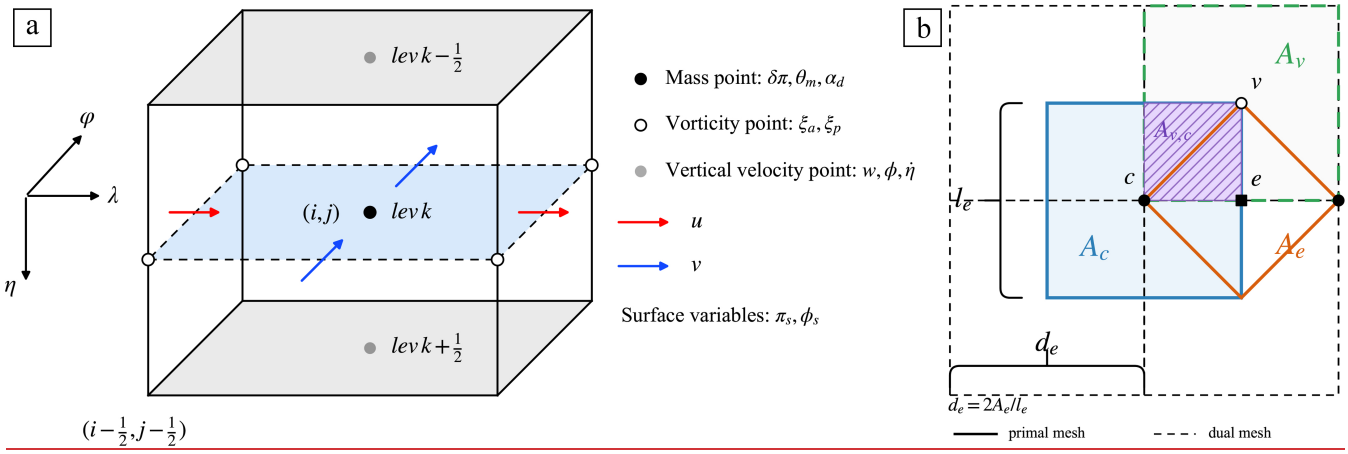


Figure 1: Schematic of the staggered-grid arrangement and the associated horizontal control volumes. The horizontal discretization uses an Arakawa C-grid and the vertical discretization uses Lorenz staggering. The black dot, open circles, and gray dots denote the mass point, vorticity points, and vertical-velocity points, respectively. Red and blue arrows denote the u and v wind components. The light-blue plane denotes the full level, and the gray planes denote the half levels. **(b)** Horizontal C-grid control-volume geometry around the edge and vertex locations. The primal mesh and dual mesh are shown by solid and dashed lines, respectively. The blue square denotes the cell-centered area A_c, the green dashed square denotes the vertex-associated area A_v, and the orange diamond denotes the edge-associated area A_e. The hatched region indicates the overlap area $A_{v,c}$ between the vertex-associated and cell-centered control volumes. The characteristic edge length l_e and dual distance d_e are also indicated, with $d_e = 2A_e/l_e$.

Following Ringler et al. (2010), the subscripts c, e, and v denote a scalar-cell center, a velocity edge, and a vorticity vertex (the center of the corresponding dual control volume), respectively. In fig.1b, A_c and A_v are the primal- and dual-control-volume areas, l_e is the primal edge length, and d_e is the distance between the two neighboring cell centers sharing edge e. The sets $E(c)$, $C(e)$, $E(v)$, $C(v)$, and $E_\perp(e)$ denote the edges around cell c, the two cells sharing edge e, the edges around vertex v, the cells surrounding vertex v, and the neighboring transverse edges of a target edge, respectively. The orientation factors $n_{e,c} = \pm 1$ and $t_{e,v} = \pm 1$ give the outward flux sign for a cell and the positive circulation sign around a vertex.

Let a be the reference planetary radius and r_c, r_e, and r_v the local radial distances at cell, edge, and vertex locations. A superscript (a) denotes the corresponding horizontal geometry on the reference sphere. The deep-atmosphere metric is therefore introduced directly through the local lengths and areas:

$$l_e = \frac{r_e}{a}\, l_e^{(a)}, d_e = \frac{r_e}{a}\, d_e^{(a)}, A_c = \left(\frac{r_c}{a}\right)^2 A_c^{(a)}, A_v = \left(\frac{r_v}{a}\right)^2 A_v^{(a)} . \tag{18}$$

The shallow-atmosphere form is recovered when $r_c = r_e = r_v = a$. Vertical transport uses the analogous shell-area factor $(r/a)^2$ at layer interfaces.

For the basic finite-volume operators, F_e below denotes a generic edge-normal flux per unit edge length (distinct from the subgrid tendencies above). Its divergence at cell c is

$$[\nabla \cdot \mathbf{F}]_c = \frac{1}{A_c} \sum_{e \in E(c)} n_{e,c}\, F_e\, l_e . \tag{19}$$

For a scalar q_c defined at cell centers, the edge-normal gradient is

$$[\nabla q]_e = \frac{1}{d_e} \sum_{c \in C(e)} (-n_{e,c})\, q_c = \frac{q_{c_2} - q_{c_1}}{d_e} , \tag{20}$$

where c_1 and c_2 are the two cells in $C(e)$, ordered along the oriented edge normal from c_1 to c_2. With V_e denoting the edge-normal wind component (u_e or v_e according to the edge orientation), the relative vorticity at vertex v is obtained from the circulation around the dual control volume:

$$\xi_v = \left[\hat{\mathbf{k}} \cdot (\nabla \times \mathbf{V})\right]_v = \frac{1}{A_v} \sum_{e \in E(v)} t_{e,v}\, V_e\, d_e , \tag{21}$$

where $\hat{\mathbf{k}}$ is the local vertical unit vector. At the exact poles, GMCORE evaluates the same Stokes circulation over the complete polar cap (Li et al., 2020).

At a full model level, u_e and v_e denote the zonal and meridional wind components on their respective C-grid edges, w_e is the vertical velocity horizontally collocated with that edge, φ_e is the latitude. The semi-discrete horizontal momentum equations are:

$$\frac{\partial u_e}{\partial t} = R_e^{(u)} - \left[\nabla_\eta K\right]_e^{(u)} - \dot{\eta}\frac{\partial u}{\partial \eta} - L\,\mathrm{PGF}_{u_e} - \left(2\Omega\cos\varphi_e + \frac{u_e}{r_e}\right) w_e + S , \tag{22}$$

$$\frac{\partial v_e}{\partial t} = R_e^{(v)} - \left[\nabla_\eta K\right]_e^{(v)} - \dot{\eta}\frac{\partial v}{\partial \eta} - L\,\mathrm{PGF}_{v_e} - \frac{v_e w_e}{r_e} + S . \tag{23}$$

Here $R_e^{(u)}$ and $R_e^{(v)}$ are the PV-flux rotational tendencies; K is the horizontal kinetic energy and ∇_η denotes the horizontal

gradient on an η surface; PGF_{u_e} and PGF_{v_e} are the pressure-gradient forces; and S represents subgrid tendencies. The nontraditional Coriolis and curvature terms are kept explicitly in Eq. (22) and Eq. (23).

The first term is Coriolis rotation tendency, let $\delta\pi_c$ be the layer dry-air mass at cell c. The mass is interpolated to vertex v with the primal-dual overlap area weight $A_{v,c}$. The vertex layer mass and layer-mass potential vorticity are

$$\delta\pi_v = \frac{1}{A_v}\sum_{c\in C(v)} A_{v,c}\,\delta\pi_c,\ \ \xi_{p,v} = \frac{\xi_v + f_v}{\delta\pi_v},\ \ f_v = 2\Omega\sin\varphi_v\ . \tag{24}$$

Thus $\delta\pi_v$ is the layer mass collocated with the vorticity, f_v is the traditional Coriolis parameter, and $\xi_{p,v}$ is the layer-mass potential vorticity.

At edge e, $\delta\pi_e$ is the mean layer mass of the two cells in $C(e)$, V_e is the edge-normal wind, and F_e is the normal dry-mass flux. The transverse flux $F_e^{\perp}$ at the same target edge is reconstructed as

$$F_e = \delta\pi_e V_e,\ \ \delta\pi_e = \frac{1}{2}\sum_{c\in C(e)}\delta\pi_c,\ \ F_e^{\perp} = \frac{1}{d_e}\sum_{e'\in E_{\perp}(e)}\omega_{e,e'}\,l_{e'}\,F_{e'}\ , \tag{25}$$

where e' runs over the four transverse neighbors in $E_{\perp}(e)$, and $\omega_{e,e'}$ is the corresponding geometric interpolation weight and we take $\omega_{e,e'} = \frac{1}{4}$. This is the Ringler-type normal-to-transverse flux mapping used to collocate the mass flux with the rotational tendency.

GMCORE then reconstructs PV from the vertices to the target edge with its default third-order upwind stencil. Let v_{-1} and v_{+1} be the two nearest vertices on the two sides of edge e in the transverse direction and v_{-2}, v_{+2} the next outward vertices. The reconstructed edge PV is

$$\xi_{p,e}^{(3)} = \frac{7}{12}\left(\xi_{p,v_{+1}} + \xi_{p,v_{-1}}\right) - \frac{1}{12}\left(\xi_{p,v_{+2}} + \xi_{p,v_{-2}}\right) + \frac{\beta_{\mathrm{pv}}}{12}\,\mathrm{sgn}(\mathcal{F}_e^{\perp})\left[\xi_{p,v_{+2}} - \xi_{p,v_{-2}} - 3\left(\xi_{p,v_{+1}} - \xi_{p,v_{-1}}\right)\right], \tag{26}$$

where β_{pv} controls the upwind contribution and $\mathrm{sgn}(\mathcal{F}_e^{\perp})$ selects the upstream side from the transverse mass flux. The PV-weighted rotational tendency is then

$$R_e = \sigma_e\frac{1}{d_e}\sum_{e'\in E_{\perp}(e)}\omega_{e,e'}\,l_{e'}\,\mathcal{F}_{e'}\,\frac{\xi_{p,e}^{(3)} + \xi_{p,e'}^{(3)}}{2}\ , \tag{27}$$

where $\sigma_e = \pm 1$ accounts for the orientation of the zonal or meridional momentum equation, and $\xi_{p,e}^{(3)}$ and $\xi_{p,e'}^{(3)}$ are the third-order reconstructed PV values at the target and neighboring transverse edges. According to the edge orientation, R_e in Eq. (27) supplies $R_e^{(u)}$ or $R_e^{(v)}$ in Eq. (22) and Eq. (23).

The kinetic energy is diagnosed at cell centers. Define the edge-associated area A_e, the edge-based estimate $K_c^{(e)}$, the optional vertex-based estimate $K_c^{(v)}$, and the prescribed blending coefficient α_K:

$$A_e = \frac{1}{2}l_e d_e,\qquad K_c^{(e)} = \frac{1}{A_c}\sum_{e\in E(c)}\frac{A_e}{4}V_e^2,\qquad K_c = \alpha_K K_c^{(e)} + (1-\alpha_K)K_c^{(v)}\ . \tag{28}$$

Its contribution at edge e is obtained with the same normal-gradient operator

$$[\nabla K]_e = \frac{1}{d_e}\sum_{c\in C(e)}\left(-n_{e,c}\right)K_c = \frac{K_{c_2}-K_{c_1}}{d_e}\,, \tag{29}$$

where K_{c_1} and K_{c_2} are the kinetic energies in the two cells adjacent to e. Because d_e already contains the local-radius scaling in Eq. (18), the deep-atmosphere metric enters this gradient directly.

Vertical momentum advection follows Simmons and Burridge (1981). At full level k, $M_{e,k+\frac{1}{2}}$ and $M_{e,k-\frac{1}{2}}$ are the dry-mass fluxes through the upper and lower interfaces, $\delta\pi_{e,k}$ is the layer dry-air mass at edge e, and $V_{e,k}$ denotes either $u_{e,k}$ or $v_{e,k}$. The operator is

$$\left(\dot{\eta}\frac{\partial V_e}{\partial \eta}\right)_k = \frac{M_{e,k+\frac{1}{2}}\left(V_{e,k+1}-V_{e,k}\right)+M_{e,k-\frac{1}{2}}\left(V_{e,k}-V_{e,k-1}\right)}{\delta\pi_{e,k}}\,. \tag{30}$$

The interface mass fluxes $M_{e,k+\frac{1}{2}}$ and $M_{e,k-\frac{1}{2}}$ are diagnosed from the same layer continuity equation, in the deep-atmosphere configuration each interface flux carries its own shell-area factor.

Finally, the horizontal pressure-gradient force is written in perturbation form to reduce cancellation on the terrain-following coordinate. Here p_c, ϕ_c, $\alpha_{d,c}=1/\rho_{d,c}$, and π_c retain their definitions from Sect.2.1; the subscript ref denotes a time-independent hydrostatically balanced reference state and a superscript prime denotes the perturbation from that state:

$$p_c = p_{\mathrm{ref},c}+p'_c,\phi_c=\phi_{\mathrm{ref},c}+\phi'_c,\alpha_{d,c}=\alpha_{d,\mathrm{ref},c}+\alpha'_{d,c},\pi_c=\pi_{\mathrm{ref},c}+\pi'_c\,. \tag{31}$$

For any cell-centered quantity q_c, the overbar $\overline{q_e}=\left(q_{c_1}+q_{c_2}\right)/2$ denotes centered interpolation to edge e, while the bracketed gradients use Eq. (20). The symbol δ denotes the vertical increment across a model layer. The pressure-gradient force at an edge is

$$\mathrm{PGF}_e = \bar{\alpha}_{d,e}[\nabla p']_e+\bar{\alpha}'_{d,e}[\nabla p_{\mathrm{ref}}]_e+[\nabla\phi']_e+\frac{\delta p'_e-\delta\pi'_e}{\delta\pi_e}[\nabla\phi]_e\,. \tag{32}$$

The corresponding term in Eq. (22) and Eq. (23) is $-L\,\mathrm{PGF}_e$.

The model employs a centered difference scheme for layer-mass transport, while a third-order upwind scheme is adopted for potential temperature, geopotential, and vertical velocity. Tracer advection is handled by the SWIFT scheme (Bendall and Kent, 2025), which is derived from the FFSL formulation (Lin and Rood, 1996) with specific modifications to preserve monotonicity in two-dimensional transport and relies on PPM for the spatial reconstruction of tracers.

A challenge for latitude-longitude grids is the severe convergence of meridians at the poles, which typically induces numerical noise and structural deformation in cross-polar flows. To circumvent this singularity, GMCORE redefines the relative vorticity around the poles via Stokes' theorem (Li et al., 2020). This specific physical constraint governs the evolution of potential enstrophy, effectively suppressing artificial computational noise sources precisely at the polar singularities.

2.4. Time integration: semi-implicit scheme

~~To circumvent~~After the ~~Courant–Friedrichs–Lewy (CFL) stability limits imposed by~~spatial discretization described in Sect. 2.3, the governing equations can be written in the following prototype form:

$$\frac{\partial \boldsymbol{X}}{\partial t} = -\mathcal{L}(\boldsymbol{X}), \tag{33}$$

where $\boldsymbol{X}$ represents any prognostic variable and $\mathcal{L}(\boldsymbol{X})$ denotes the corresponding spatially discretized tendency, including the terms described in Sect. 2.3. GMCORE advances Eq. (33) using a modified three-stage Runge–Kutta scheme (Wicker and Skamarock, 2002). Starting from the model state $\boldsymbol{X}^n$ at time level n, the three stages are

$$\boldsymbol{X}^* = \boldsymbol{X}^n - \frac{\Delta t}{3}\mathcal{L}(\boldsymbol{X}^n)\,, \tag{34a}$$

$$\boldsymbol{X}^{**} = \boldsymbol{X}^n - \frac{\Delta t}{2}\mathcal{L}(\boldsymbol{X}^*)\,, \tag{34b}$$

$$\boldsymbol{X}^{n+1} = \boldsymbol{X}^n - \Delta t\mathcal{L}(\boldsymbol{X}^{**})\,, \tag{34c}$$

where Δt is the dynamical time step, $\boldsymbol{X}^*$ and $\boldsymbol{X}^{**}$ denote the intermediate states of the first and second Runge–Kutta stages, respectively, and $\boldsymbol{X}^{n+1}$ is the updated state. At each stage, the spatial tendency is recomputed using the most recently obtained intermediate state, whereas the stage update is based on $\boldsymbol{X}^n$. Thus, the three evaluations of the spatial operator are successively performed at $\boldsymbol{X}^n$, $\boldsymbol{X}^*$, and $\boldsymbol{X}^{**}$.

Each Runge–Kutta stage can be regarded as a forward-in-time update over the corresponding stage interval. Within such an update, most dycore tendencies are evaluated explicitly, whereas the vertically propagating fast acoustic ~~waves, GMCORE employs a semi-implicit (SI) time integration scheme.~~ modes require a different treatment. When using the nonhydrostatic solver, the vertical momentum and geopotential equations (in Eq. (2) and Eq. (6)) form a vertically coupled subsystem: w is coupled to the vertical pressure-gradient force, while ϕ evolves directly with w, and the pressure response is linked to them through the equation of state.

Following the vertical implicit treatment used in nonhydrostatic models GRIST (Zhang et al., 2019) and ICON (e.g., Zängl et al., 2015), the prognostic variables are computed directly as full-field quantities without relying on a stationary, global background reference profile.

In this semi-implicit formulation, horizontal advection, vertical advection~~,~~ and the full Coriolis force ~~and subgrid physical tendencies~~ are evaluated explicitly. Conversely, the terms responsible for high-frequency vertical acoustic wave propagation—specifically the vertical pressure gradient force in the vertical momentum equation, and the associated $w \cdot g(r)$ term in the geopotential equation—are treated implicitly. By introducing an off-centering weight ~~α~~$\beta \in [0.5,1.0]$, the implicit equations for the vertical velocity w and the geopotential ϕ can be conceptually simplified as:

~~$$w^{n+1} = \Delta t \cdot \mathcal{R}_w^{\pi} + \alpha \Delta t L g \left(\frac{\delta p}{\delta \pi}\right)^{n+1} \tag{18}$$~~

~~$$\phi^{n+1} = \Delta t \cdot \mathcal{R}_{\phi}^{\pi} + \alpha \Delta t g w^{n+1} \tag{19}$$~~

$$w^{\tau} = \Delta\tau \cdot R_w^* + \beta\Delta\tau L g_0 \left(\frac{r}{a}\right)^2 \frac{\delta(p^{\tau})}{\delta(\pi^{\tau})}, \tag{35}$$

$$\phi^{\tau} = \Delta\tau \cdot R_{\phi}^* + \beta\Delta\tau g_0 \left(\frac{a}{r}\right)^2 w^{\tau}, \tag{36}$$

where ~~$\mathcal{R}_w^n R_w^*$~~ and ~~$\mathcal{R}_\phi^n R_\phi^*$~~ represent the total explicit residual tendencies evaluated at time ~~step n, $L = \frac{p_d}{p} = \frac{1}{1+q_m}$~~stage $*$ and $\Delta\tau$ is the ~~moist contribution. The~~sub-timestep. At full levels, the temporal increment of pressure ~~$\delta p p^{\tau} - p^*$~~ is diagnostically linked to ~~$\delta\phi$~~vertical increment of $\delta\phi$ and the thermodynamic properties through the linearized equation of state:

~~$$\frac{p^{n+1} - p^n}{p^n} \approx \gamma \left(\frac{(\delta\pi\theta_m)^{n+1}}{(\delta\pi\theta_m)^n} - \frac{\delta\phi^{n+1}}{\delta\phi^n} \right) \tag{20}$$~~

$$\frac{p^{\tau} - p^*}{p^*} \approx \gamma \left(\frac{(\delta\pi\theta_m)^{\tau}}{(\delta\pi\theta_m)^*} - \frac{\delta(\phi^{\tau})}{\delta(\phi^*)} \right), \tag{37}$$

where $\gamma = c_p/c_v$ is the ratio of the specific heats at constant pressure and constant volume, the Eq. (37) can be discretized on the vertical:

$$\delta(p^{\tau}) = \delta(p^*) + \gamma\delta\left(p^* \frac{(\delta\pi\theta_m)^{\tau}}{(\delta\pi\theta_m)^*}\right) - \gamma\delta\left(p^n \frac{\delta(\phi^{\tau})}{\delta(\phi^*)}\right). \tag{38}$$

By substituting the implicit ~~increments of~~ geopotential ~~($\delta\phi$) and pressure (δp~~relation in Eq. (36) into ~~the discretized~~ Eq. (38), and then substituting the resulting pressure increment into the vertical momentum equation, Eq. (35), we eliminate the dependencies on other prognostic variables. This procedure reduces the coupled equations into a one-dimensional vertical equation for the vertical velocity w within each independent grid column. When discretized on the vertically staggered levels, it directly forms a standard tri-diagonal linear system:

~~$$A_{k-\frac{1}{2}} w_{k-\frac{1}{2}} + B_{k+\frac{1}{2}} w_{k+\frac{1}{2}} + C_{k+\frac{3}{2}} w_{k+\frac{3}{2}} = \widetilde{R_{w,k+\frac{1}{2}}} \tag{21}$$~~

$$A_{k-\frac{1}{2}} w^{\tau}_{k-\frac{1}{2}} + B_{k+\frac{1}{2}} w^{\tau}_{k+\frac{1}{2}} + C_{k+\frac{3}{2}} w^{\tau}_{k+\frac{3}{2}} = R^*_{w,k+\frac{1}{2}}, \tag{39}$$

where ~~$\widetilde{R_{w,k+\frac{1}{2}}} R^*_{w,k+\frac{1}{2}}$~~ includes all the explicitly resolved forcing terms. The matrix coefficients $A_{k-\frac{1}{2}}, B_{k+\frac{1}{2}}$, and $C_{k+\frac{3}{2}}$ are dynamically updated at each time step based on the local atmospheric state.

This one-dimensional tri-diagonal system is solved independently for every vertical column using the highly efficient Thomas algorithm, demanding no horizontal communication. Once the vertical velocity w is obtained, the increments for the remaining implicitly treated variables (such as $\delta\phi$ and ~~δp~~δp) are sequentially back-substituted. This semi-implicit vertical nonhydrostatic solver successfully decouples the fast vertical acoustic wave speeds from the integration stability criteria. This approach allows GMCORE to avoid solving global elliptic equations during nonhydrostatic calculations and to increase the time step.

2.5. Polar filtering and damping

Due to the convergence of the latitude and longitude grids at the poles, the zonal grid spacing shrinks significantly in the polar regions. As a result, the local zonal CFL number increases sharply in the polar region and can impose a severe stability constraint on the explicit part of the dynamical time integration.

In GMCORE, we apply an adaptive zonal gaussian convolution filter for dynamical tendencies along each latitude circle. For an arbitrary tendency field ~~$\mathcal{M}M$~~ in $u, v,$ ~~θ_m~~θ_m, ~~π_s~~π_s and advection tendency of w, ϕ in nonhydrostatic module, the filtered field ~~$\mathcal{M}M^{filt}$~~ can be written

~~$$\mathcal{M}_{i,j,k}^{filt} = \sum_{m=-H_j}^{H_j} w_{j,m} \mathcal{M}_{i+m,j,k}, \quad (22)$$~~

$$M_{i,j,k}^{filt} = \sum_{m=-H_j}^{H_j} G_{j,m} M_{i+m,j,k}, \quad (40)$$

where ~~$w_{j,m}$~~$G_{j,m}$ is the gaussian weight at offset m on the j-th latitude circle, and H_j is the corresponding half-width of the stencil. Because the zonal direction is periodic, the index $i + m$ is wrapped cyclically around the latitude circle.

The kernel width is not prescribed as a constant~~,~~ but diagnosed adaptively from the local zonal stability constraint. In the implementation, a dimensionless width parameter is first defined as

~~$$W_j = \alpha_f \frac{c_f \Delta t}{\Delta x_j} \quad (23)$$~~

$$W_j = \alpha_f \frac{c_f \Delta t}{\Delta x_j}, \quad (41)$$

where Δx_j is the zonal grid spacing on latitude ϕ_j, Δt is the dynamical time step, and c_f is a characteristic filtering wave speed now set to 300 m s^{-1}. The coefficient α_f is the parameter controlling the overall filter strength. The factor $\frac{c_f \Delta t}{\Delta x_j}$ has the form of a zonal CFL number, so that the filter automatically becomes stronger where the grid is more restrictive. This latitude-dependent feature provides a smooth transition of the kernel width with latitude and helps avoid abrupt changes in filtering strength.

The diagnosed continuous width W_j is converted to the discrete stencil size by taking the smallest odd integer not smaller than W_j, with a minimum stencil size of three grid points: $N_j = max\left[3, oddceil\left(W_j\right)\right]$. The continuous kernel is represented on the discrete stencil through a gaussian profile. For a stencil of width N_j ~~$== 2H_j + 1$~~, the unnormalized weights are defined by

~~$$\tilde{w}_{j,m} = \frac{1}{s_j \sqrt{2\pi}} \exp\left(-\frac{m^2}{2s_j^2}\right), m = -H_j, \ldots, H_j \quad (24)$$~~

$$\tilde{G}_{j,m} = \frac{1}{s_j \sqrt{2\pi}} \exp\left(-\frac{m^2}{2s_j^2}\right), m = -H_j, \ldots, H_j, \quad (42)$$

with $s_j = \frac{w_j}{\sigma_f}$, where σ_f is a prescribed shape parameter controlling the relation between the diagnosed width W_j and the standard deviation of the discrete gaussian kernel. The final gaussian filter weights are then normalized as

~~$$w_{j,m} = \tilde{w}_{j,m} \Big/ \sum_{m=-H_j}^{H_j} \tilde{w}_{j,m}, \quad \sum_{m=-H_j}^{H_j} w_{j,m} = 1 \tag{25}$$~~

$$G_{j,m} = \tilde{G}_{j,m} \Big/ \sum_{m=-H_j}^{H_j} \tilde{G}_{j,m}, \quad \sum_{m=-H_j}^{H_j} G_{j,m} = 1 . \tag{43}$$

The normalization procedure guarantees that the filter preserves constant fields. In this way, the filter automatically strengthens near the poles while remaining weak in lower latitudes. Because the method uses only finite-width local convolution, it avoids the global communication required by FFT-based polar filtering and is therefore well suited to parallel latitude–longitude grid implementations. It should be noted that the wider communication halo is required only for the variables to which the filter is applied. In the parallel implementation, processes covering higher latitudes therefore use wider halos for those variables. In practice with a horizontal resolution of 1 °, this half width halo usually reaches 90 points at the two poles of north and south. Nevertheless, the additional communication cost remains manageable for the resolutions considered here.

In addition to the polar gaussian filter, GMCORE applies optional horizontal divergence damping to the momentum field (Lauritzen et al., 2011). This damping is to diffuse the divergent part of flow noise that may otherwise accumulate near steep gradients, in the upper levels, or in the polar region. The damping is formulated in terms of the horizontal divergence, let

~~$$\delta_{\hbar} = \nabla_{\hbar} \cdot \boldsymbol{v}_{\hbar} \tag{26}$$~~

$$\delta_h = \nabla_h \cdot \mathbf{V} , \tag{44}$$

where ~~$\boldsymbol{v}_{\hbar}\mathbf{V} = (u, v)$~~ is the wind vector. δ_h is the divergence. The second-order divergence damping is written as

~~$$\left.\frac{\partial \boldsymbol{v}_{\hbar}}{\partial t}\right|_{\delta,2} = \nu_{\delta,2} \nabla_{\hbar} \delta_{\hbar} \tag{27}$$~~

$$\left.\frac{\partial \mathbf{V}}{\partial t}\right|_{\delta,2} = \nu_{\delta,2} \nabla_h \delta_h , \tag{45}$$

and a more scale-selective fourth-order option is

~~$$\left.\frac{\partial \boldsymbol{v}_{\hbar}}{\partial t}\right|_{\delta,4} = -\nu_{\delta,4} \nabla_{\hbar} \nabla_{\hbar}^2 \delta_{\hbar} \tag{28}$$~~

$$\left.\frac{\partial \mathbf{V}}{\partial t}\right|_{\delta,4} = -\nu_{\delta,4} \nabla_h \nabla_h^2 \delta_h . \tag{46}$$

The damping coefficients $\nu_{\delta,2}$ and $\nu_{\delta,4}$ are scaled with the local grid area and time step and can be smoothly enhanced near the poles or upper levels.

~~$$\nu_{\delta,2} \sim C_2 \frac{A}{\Delta t}, \quad \nu_{\delta,4} \sim C_4 \frac{A^2}{\Delta t} \tag{29}$$~~

$$\nu_{\delta,2} \sim C_2 \frac{A}{\Delta t}, \qquad \nu_{\delta,4} \sim C_4 \frac{A^2}{\Delta t}. \tag{47}$$

GMCORE also provides an optional upper-level Rayleigh damping of vertical velocity. This damping is not activated in the two idealized cases presented here; with their comparatively low model tops, both cases remain numerically stable without a sponge layer. For future high-top applications, the same damping option can be used as an upper-boundary treatment to attenuate upward-propagating disturbances and reduce artificial reflection, rather than as a correction to the deep-atmosphere equations themselves.

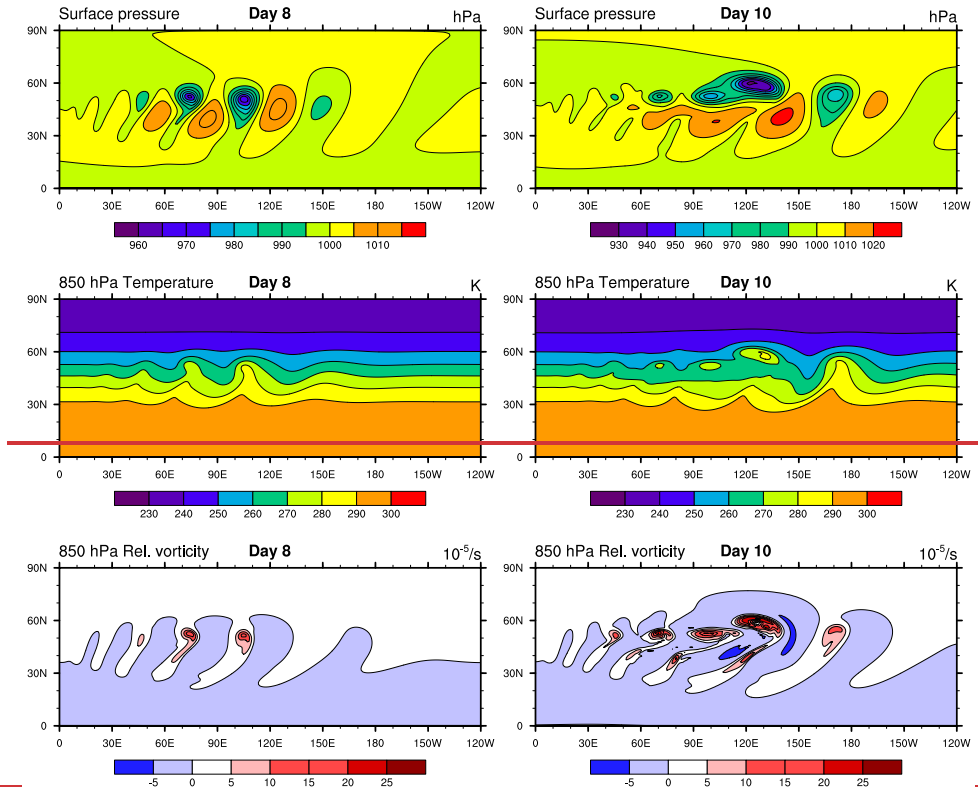
3. Test cases and simulation results

To test the validity and performance of deep-atmosphere version of GMCORE, particularly the deep-atmosphere modifications described in Section 2, two standard idealized test cases are selected.

3.1. Deep-atmosphere baroclinic wave

The idealized baroclinic wave test case evaluates the model's ability to simulate mid-latitude baroclinic processes. Following the model parameters proposed by Ullrich et al. (2014), we employ this test to demonstrate the characteristics of the deep-atmosphere configurations. The simulations use the 30-level vertical grid adopted from CAM, with the model top located at 2.255 hPa, corresponding to an approximate geometric height of 30 km. The horizontal resolution is 1°, corresponding to a 360×180 regular latitude–longitude grid. The domain is decomposed along the latitude direction for parallel computation. The simulations were performed using 2 compute nodes with 120 CPU cores in total, and the wall-clock time for the integration reported here was approximately 5 min.

The initial conditions feature a mid-latitude jet in hydrostatic and geostrophic balance (Staniforth and White, ~~2012), which is subsequently perturbed to trigger baroclinic instability. Simulations are conducted using both the deep-atmosphere and the shallow-atmosphere solvers~~2011), which is subsequently perturbed to trigger baroclinic instability. Before the baroclinic simulation, the barotropic steady geostrophic flow (see Williamson et al., 1992, their test Case2) is employed, for which an analytical solution is available to verify the convergence of GMCORE. Another reason for adopting this barotropic test case is to avoid the non-convergence issues that may arise from initialization in a three-dimensional baroclinic model. We then use the unperturbed standard-Earth baroclinic-wave basic state to examine the stability of GMCORE.



~~**Figure 1:** Evolution of the surface pressure, 850 hPa temperature and relative vorticity for the deep baroclinic wave test at days 8 and 10.~~

~~Figure 1~~

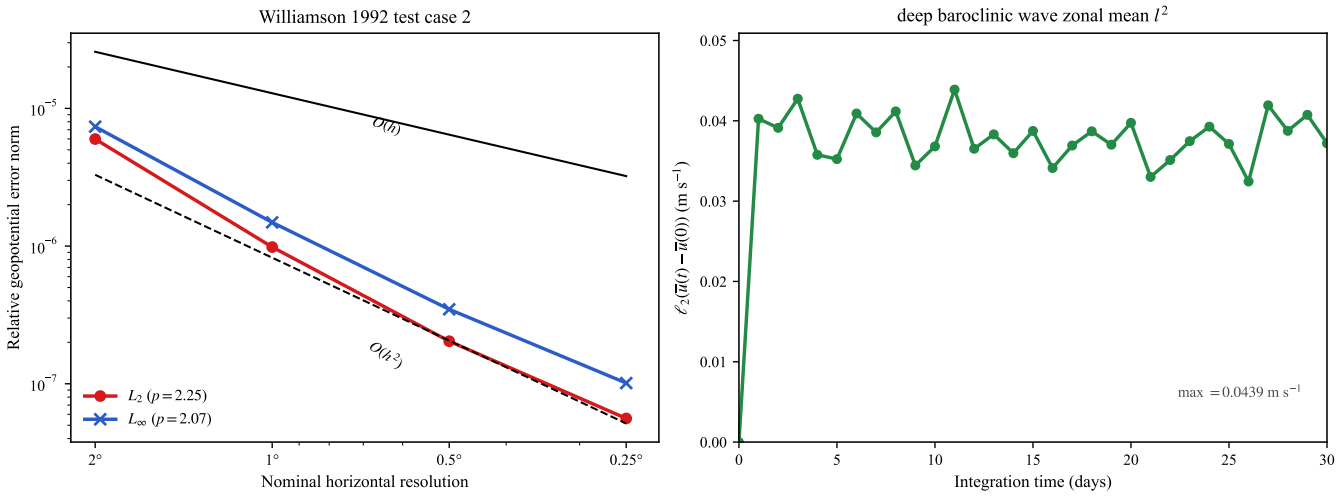


Figure 2: Convergence and stability diagnostics of GMCORE. The left panel presents the relative (L_2) and (L_∞) geopotential error norms for Williamson 2007 test case 2 at resolutions from2° to 0.25°, with reference $O(h)$ and $O(h^2)$ slopes. The right panel shows the (l^2) norm

of the zonal-mean wind deviation from its initial state in the unperturbed deep baroclinic-wave basic-state test, demonstrating a bounded drift over 30 days.

Steady geostrophic flow is integrated for 5 days at nominal horizontal resolutions of 2°, 1°, 0.5°, and 0.25°, with time steps of 600, 300, 150, and 75 s, respectively. The fitted convergence orders of the relative geopotential errors are 2.25 for L^2 and 2.07 for L^∞, indicating near-second-order convergence (in Fig. 2' left panel). For the three-dimensional model, stability is tested using the unperturbed standard-Earth baroclinic-wave basic state. Following Jablonowski and Williamson (2006), we monitor the weighted norms $l_2\left(\overline{u}(t) - \overline{u}(0)\right)$. On the 1° grid, the norm remains below 0.044 m/s throughout the 30-day integration, indicating stable preservation of the balanced basic state (in Fig. 2' right panel).

Simulations with perturbation are conducted using both the deep-atmosphere and the shallow-atmosphere solvers. Figure 3 shows the evolution of surface pressure, 850 hPa temperature, and 850 hPa relative vorticity in the deep configuration at days 8 and 10. The developing wave exhibits the familiar life cycle of baroclinic instability: a midlatitude disturbance grows along the jet, amplifies downstream, and evolves into a mature cyclone–anticyclone structure. Overall, the simulated fields are consistent with the standard behaviour reported in previous model intercomparisons of this test case (Ullrich et al., 2014; Wood et al., 2014; Borchert et al, 2019; Skamarock et al., 2021; Kelly et al., 2026; Yatunin et al., 2026), and indicate that the GMCORE deep-atmosphere dycore is able to maintain the balanced basic state and reproduce the expected evolution of the wave under Earth-like conditions.

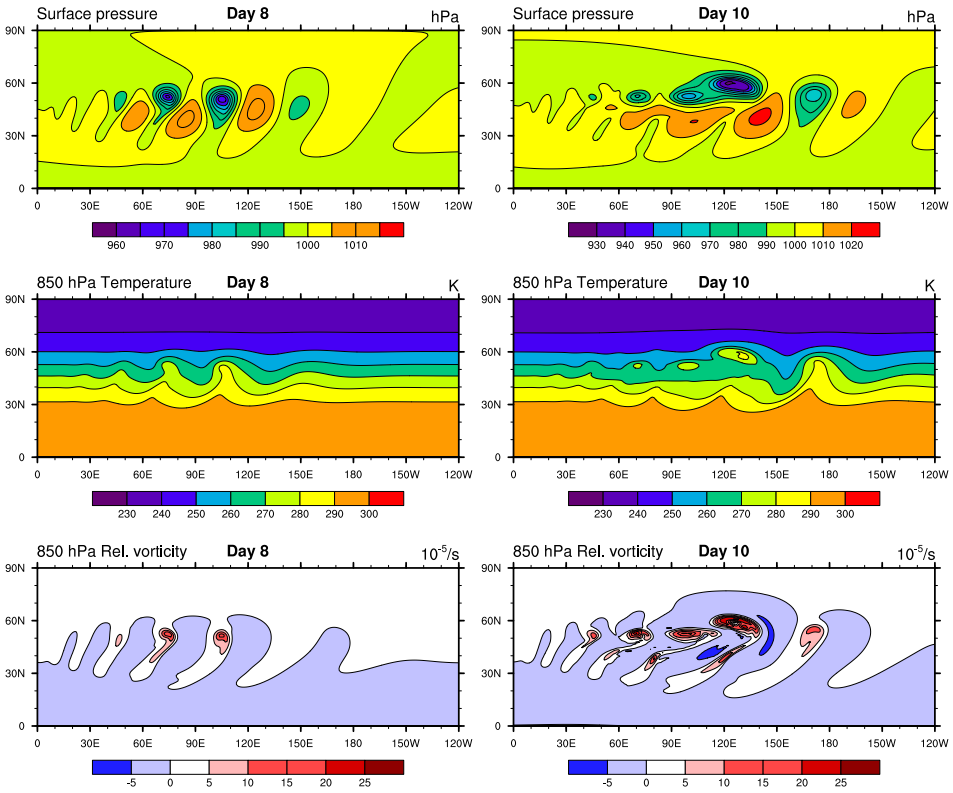


Figure 3: Evolution of the surface pressure, 850 hPa temperature and relative vorticity for the deep baroclinic wave test at days 8 and 10. These can be compared with Ullrich et at.'s (2014) Figs. 4 and 5, and with Skamarock et al.'s (2021) Fig. 2.

With a timestep of 300 s, the gravity-wave CFL number is approximately 0.27 near the equator, 0.39 at midlatitudes, and 0.64 at 60°. It exceeds unity poleward of about 70° because of the reduced zonal grid spacing, while the gaussian filter is adaptively strengthened according to Eq (41). Vertical acoustic terms are treated semi-implicitly to relax their CFL constraint.

In a standard Earth-radius ~~configuration~~experiment (a ≈ 6371 km), the troposphere is extremely thin relative to the planet's radius. Consequently, the deep-atmosphere geometrical scaling ~~$(r/a)^2$~~ $(r/a)^2$ and the variable gravity $g(r)$ introduce very small dynamical differences compared to the shallow-atmosphere.

To amplify the dynamical differences between the deep- and shallow-atmosphere equations, we follow Ullrich et al. (2014) and employ the reduced-radius, or "small-Earth", baroclinic-wave experiment, in which the planetary radius is decreased by a factor of 20 while the planetary angular rotation ~~rate~~velocity is increased by the same factor. Under the standard Earth-radius ~~configuration~~experiment, the two ~~configurations~~experiments produce nearly indistinguishable baroclinic-wave evolutions, whereas in the X20 ~~configuration~~experiment the shallow-atmosphere assumption is no longer strictly valid and the two solutions separate much more clearly.

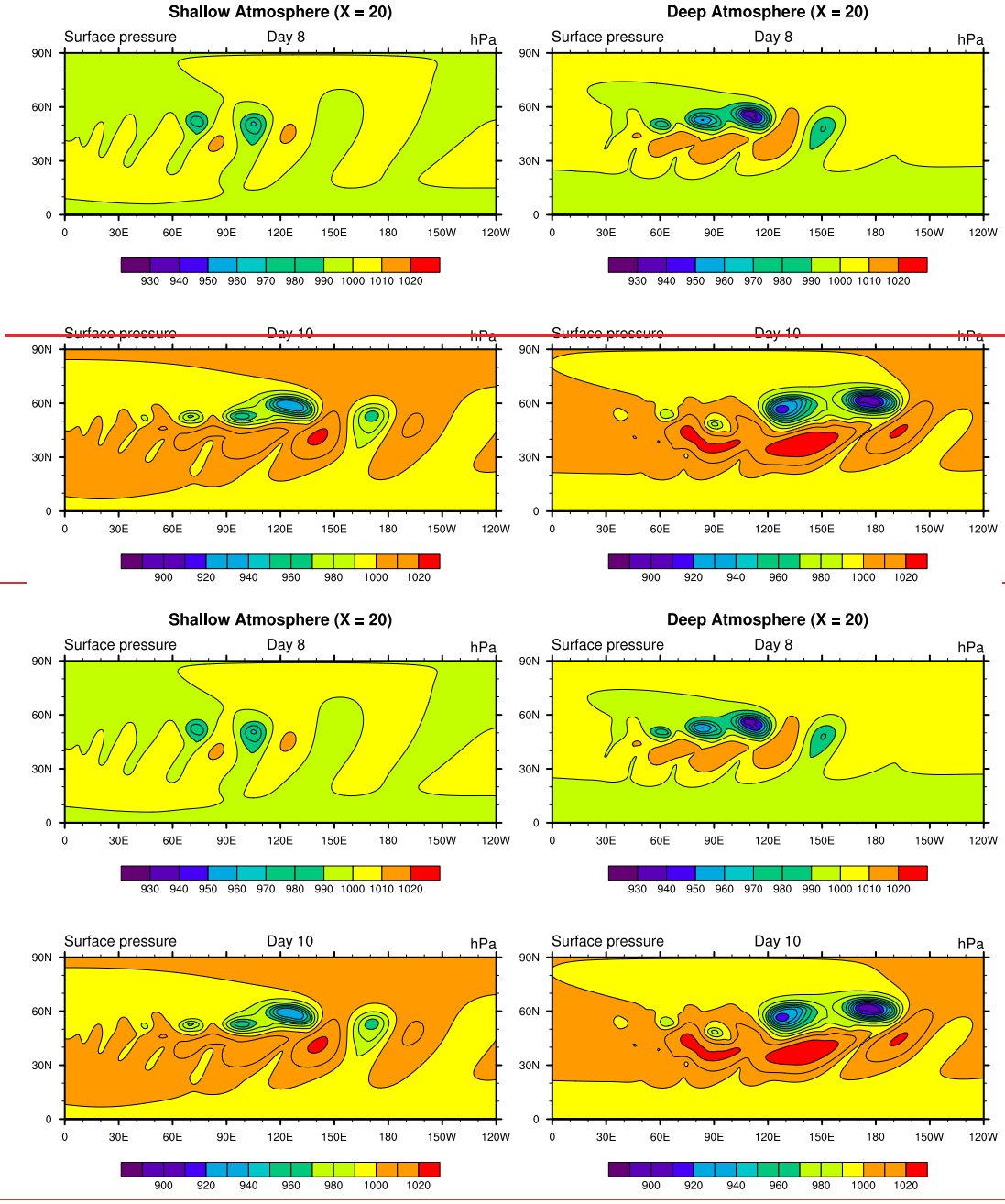


Figure ~~24~~: Surface pressure fields in the X20 reduced-radius baroclinic-wave experiment for the shallow-atmosphere and deep-atmosphere configurations at days 8 and 10. Following Ullrich et al. (2014), the planetary radius is reduced by a factor of 20 and the planetary angular rotation ~~rate~~velocity is increased by the same factor in order to exaggerate the differences between the two equation sets. These can be compared with Fig. 9 in Ullrich et al. (2014), Fig. 6 in Borchert et al.'s (2019) and Fig.6 in Skamarock et al.'s (2021).

In our X20 simulations, the shallow-atmosphere solution retains a comparatively weaker and smoother baroclinic-wave structure, broadly consistent with the reference behaviour reported by Ullrich et al. (2014). In contrast, the deep-atmosphere solution exhibits a more rapid amplification of the wave and a substantially stronger cyclonic development. By day 8, the deep-atmosphere field already shows a more compact and intense low-pressure center together with a more amplified downstream wave; by day 10, the contrast becomes even more pronounced, with the deep-atmosphere solution characterized by a deeper and stronger mature cyclone, while the shallow-atmosphere counterpart remains less developed. This behaviour is consistent with Ullrich et al. (2014), who showed that in the X20 experiment the deep-atmosphere integration intensifies much more rapidly and drives the minimum surface pressure below 900 hPa by day 10, whereas the shallow-atmosphere result remains much closer to the standard baroclinic-wave evolution.

To sum up, ~~in~~ the ~~reduced-radius~~ numerical consistency tests and the baroclinic-wave ~~experiment, GMCORE reproduces the pronounced separation between~~experiments serve complementary purposes. The steady geostrophic-flow test shows near-second-order convergence, while the unperturbed standard-Earth basic state remains close to its balanced state throughout the 30-day integration. For the perturbed baroclinic wave, the deep- and shallow-atmosphere solutions ~~reported in earlier studies. This performance is~~ are nearly indistinguishable under the standard-Earth-radius setup but separate clearly in the X20 reduced-radius experiment, consistent with ~~the purpose of the small-Earth configuration proposed by Ullrich et al. (2014), namely to exaggerate the differences between the two equation sets once the shallow~~published deep-atmosphere ~~assumption is no longer strictly valid. At the same time, we note that the interpretation of this test case remains non-trivial. As discussed by Skamarock et al. (2021), a substantial~~benchmarks. Because part of the ~~previously reported deep–shallow differences in the X20 case~~X20 contrast may ~~arise~~originate from differences in the analytic initial states, ~~rather than directly from the temporal integration of the deep versus shallow equations themselves. For this reason,~~ we regard ~~the X20~~this experiment primarily as a benchmark-reproduction test~~. Its main value in the present study is to demonstrate that the~~ rather than as a quantitative attribution of the deep-atmosphere terms. Together, these results provide complementary evidence for the numerical consistency and correctness of the GMCORE deep-atmosphere ~~extension of GMCORE is able to recover the canonical numerical behaviour documented in the literature, thereby providing confidence in the correctness of the model equation and~~ implementation.

3.2. Tropical cyclone simulation

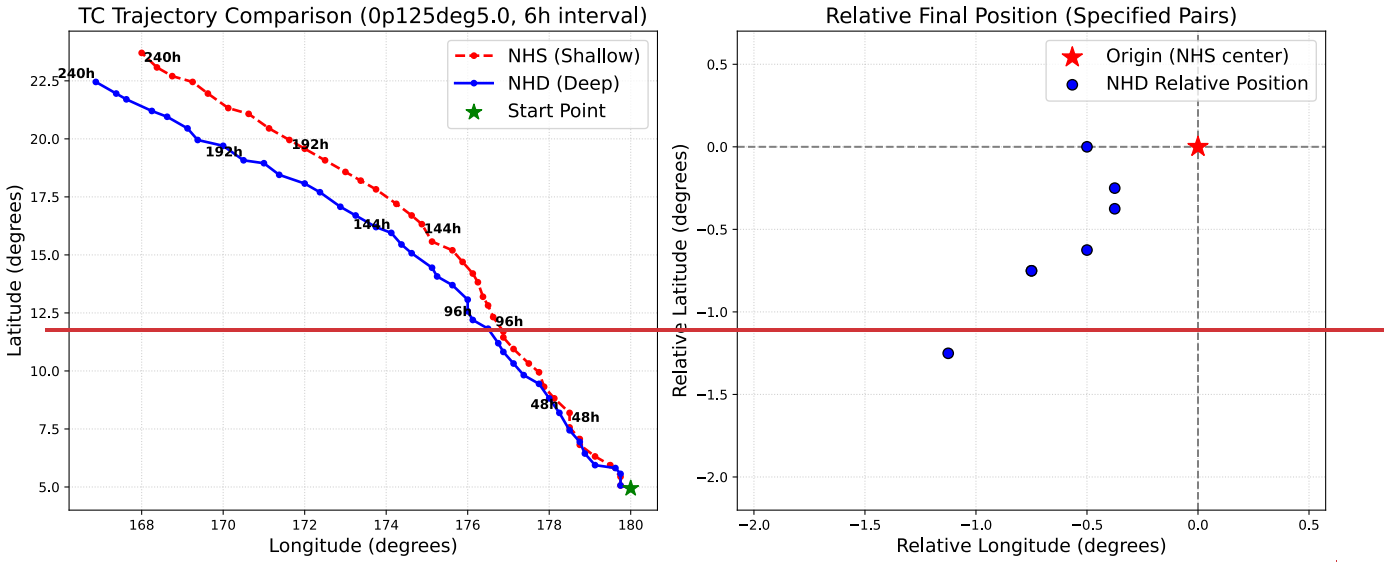
Next, we try to use an idealized tropical cyclone (TC) test case, which was proposed by Reed and Jablonowski (2011). In this case, an initial warm-core vortex that satisfies gradient wind balance is inserted in an aquaplanet without topography. This ~~configuration~~ eliminates the land surfaces and the real environmental flow effects, making it an ideal test-bed for isolating and evaluating the deep-atmosphere modification. More precisely, due to the low model top (about ~~20km~~30 km) of this case, geometric and variable-gravity effects are expected to be much smaller than the retained nontraditional Coriolis contribution. This case is therefore used mainly to examine the impact of the NCTs-related deep-atmosphere terms.

We follow the standard ~~parameters~~setup of Reed and Jablonowski (2011) and run this case at a horizontal resolution of 0.25° and 0.~~1~~25° for 10 days with the same vertical grid as the baroclinic-wave. The 0.25° simulations were performed using 480

CPU cores with a walltime around 120 minutes, whereas the 0.125° simulations used 600 CPU cores with a walltime around 900 minutes. To isolate the dynamical impact of the deep-atmosphere terms, the experiment is executed using both the nonhydrostatic deep-atmosphere (NHD) and nonhydrostatic shallow-atmosphere (NHS) configurations.

3.2.1 Tropical cyclone trajectories comparison

During the dynamical evolution of a tropical cyclone, the full three-dimensional Coriolis force—specifically the NCTs—is conventionally neglected in traditional shallow-atmosphere models. However, in the deep tropics where ~~$\cos\varphi$~~ $\cos\varphi$ approaches its maximum, and within the TC eyewall region characterized by exceptionally strong vertical updrafts (with w frequently exceeding 10 m s^{-1}), the horizontal and vertical momentum forcings generated by the NCTs become non-negligible in magnitude. Although previous idealized cloud-resolving studies (e.g., Igel and Biello, 2020) have highlighted the potential impact of NCTs on equatorial convection, the macroscopic dynamical effects of this mechanism within a global deep-atmosphere dycore during a complete TC life cycle remain largely unverified.



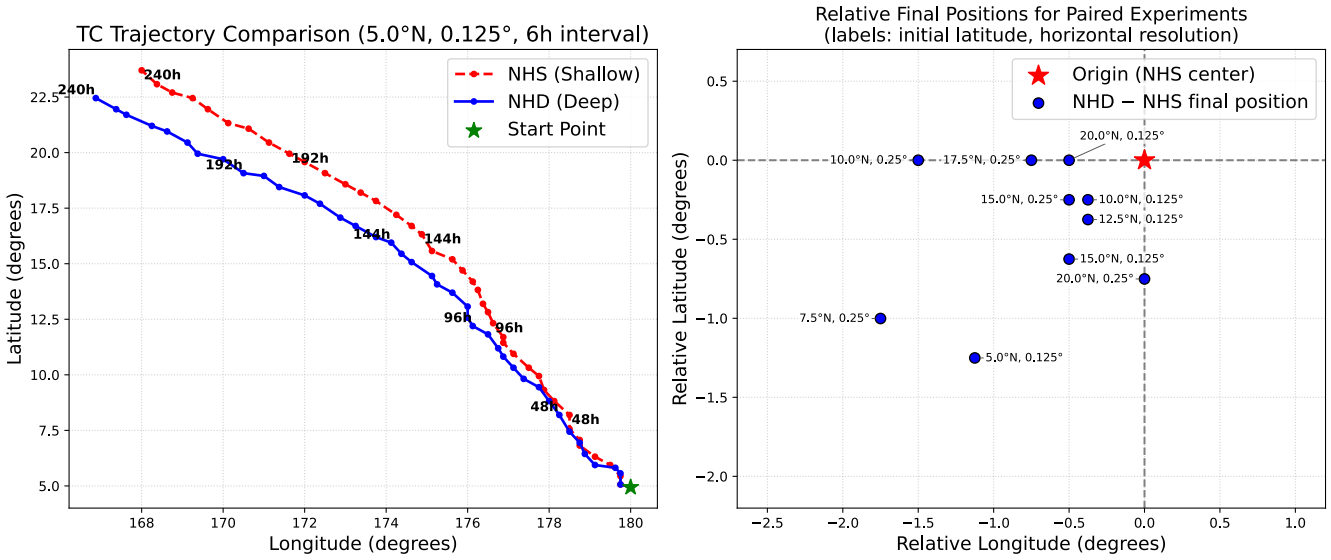


Figure ~~3~~5: Simulated tropical cyclone trajectories for the nonhydrostatic shallow-atmosphere (NHS, red line) and nonhydrostatic deep-atmosphere (NHD, blue line) configurations at a horizontal resolution of 0.125°. ~~Over~~The left panel shows the ~~integration period,~~10-day trajectories for the ~~perturbation~~ experiment initialized at 5° N. In the right panel, each blue dot shows the NHD–NHS difference in final storm-center ~~propagates north-westward and~~ position at 240 h for one independent paired experiment, after translating the corresponding NHS center to (0, 0). Labels indicate the initial perturbation latitude ~~is set at 5°N, showing on left panel. For right panel, all NHS TC center at final timestep set to (0,0), the blue points~~and horizontal resolution; the dots do not represent ~~the relative positions~~a time series. Storm centers are diagnosed directly from the grid-point minimum of ~~all 0.125° and 0.25° experiments with different perturbation latitudes~~surface pressure.

As shown in Fig. ~~3~~5a, both NHS and NHD simulations produce a north-westward-moving tropical cyclone, but the NHD trajectory exhibits a systematic south-westward displacement relative to the NHS trajectory as the vortex develops. This result is qualitatively similar to that reported by Liang and Chan (2005), who found that inclusion of the ~~$2\Omega \cos\varphi\, w$~~ $2\Omega \cos\varphi\, w$ term had little effect on tropical-cyclone intensity but led to a westward to south-westward trajectory deviation. We also test different perturbations at different latitudes and get the same conclusion (Fig. ~~3b~~5b).

In our results, both simulations remain similar in terms of storm intensity and axisymmetric structure. As already indicated in the ~~present supplementary~~Supplementary Material, the two simulations show no essential difference in the minimum central pressure, maximum low-level wind speed, or the warm-core structure, suggesting that the leading-order thermodynamic processes controlling storm intensification are not strongly altered by the deep-atmosphere extension.

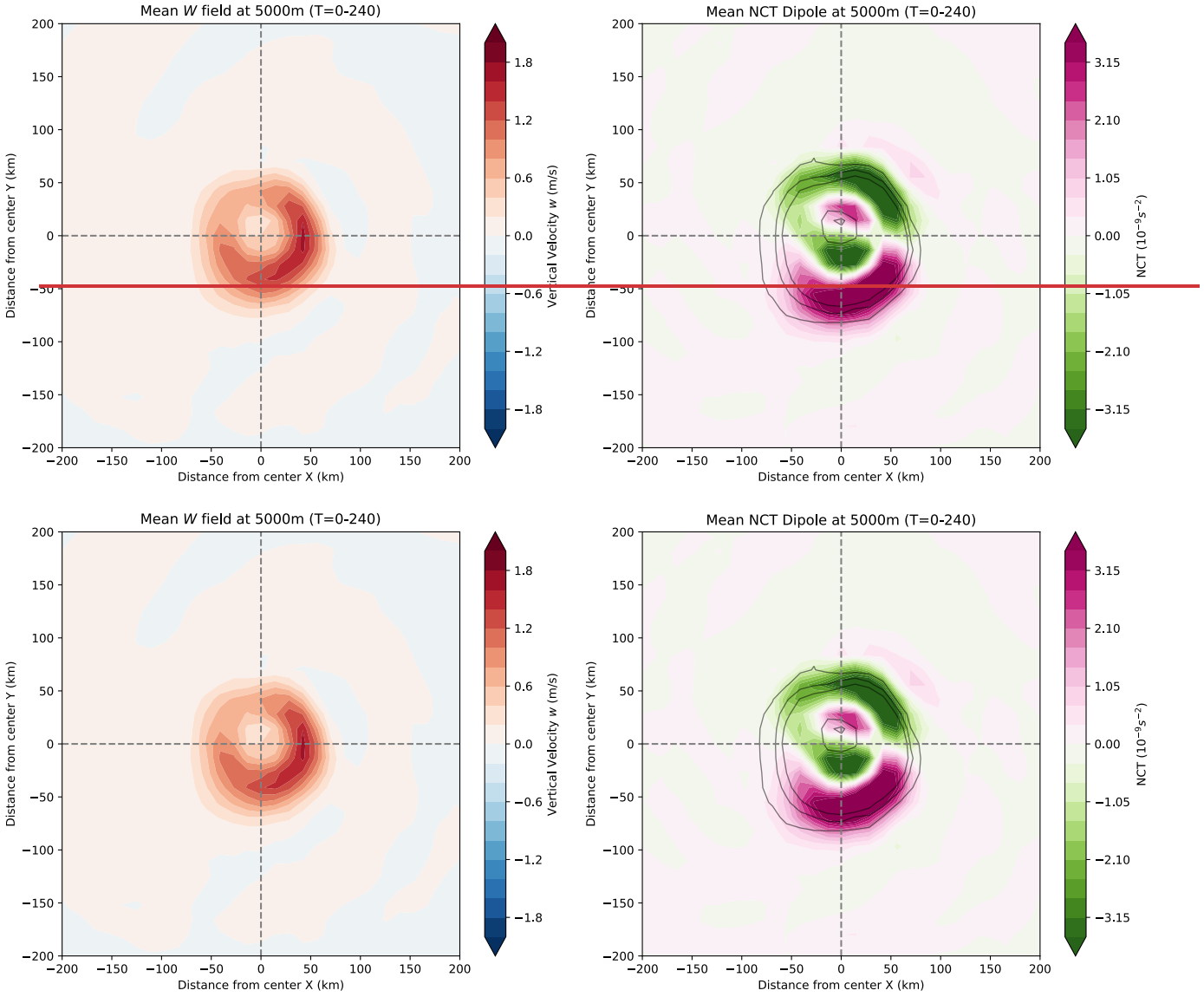


Figure ~~4~~6**:** Horizontal cross-sections of time-averaged (T = 0–240h) atmospheric variables at an altitude of 5000 m. **(Left)** Mean vertical velocity (w) field (m/s), illustrating a core of intense upward motion centered at the origin~~, surrounded by weaker spiral rainband structures.~~. **(Right)** Mean NCTs dipole (10^{-9} s^{-2}), showing a distinct positive (~~green~~pink) and negative (~~pink)~~green) coupled structure near the vortex core. Both panels are centered on the system's axis, with distances indicated in kilometers from the center (X, Y). The center is first identified within a neighbourhood of the storm position at the preceding output time. A two-dimensional quadratic surface is then fitted to the surrounding 5×5 grid points, and the local minimum of the fitted surface is used as the sub-grid storm center.

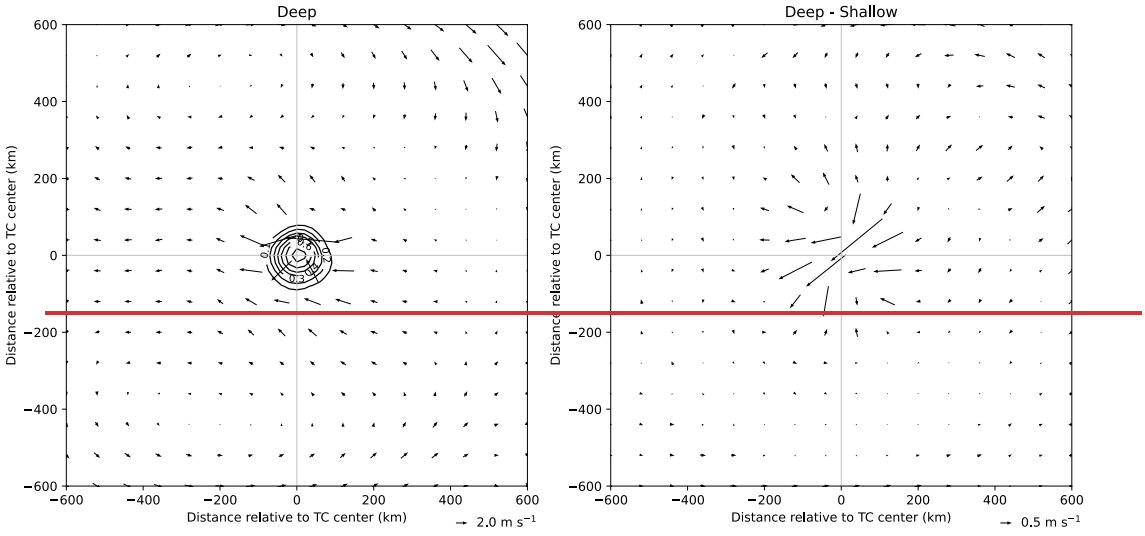
To interpret this trajectory difference, Figure ~~4~~6 shows the time-mean horizontal structure at 5 km altitude. In Fig. ~~4a~~6a, the mean vertical velocity field exhibits a compact core of strong ascent near the storm center~~, surrounded by weaker spiral~~

~~rainband structures~~. This indicates that the mature vortex maintains a well-defined eyewall-like updraft region. Closely collocated with this convective core, in Fig. ~~4b~~6b, we show the quantity

~~$$\Gamma = 2\Omega \cos\phi \frac{\partial w}{\partial y} \quad (30)$$~~

$$\Gamma = 2\Omega \cos\varphi \frac{\partial w}{\partial y}\,, \quad (48)$$

which represents the vorticity tendency associated with the NCTs in the x-momentum equation. Here, the dipole-like structure refers to the pair of opposite-signed Γ anomalies on the two sides of the storm center. In other words, the meridional gradient of the eyewall vertical motion produces a positive tendency on one side of the core and a negative tendency on the other, giving rise to an organized inner-core asymmetry rather than an axisymmetric distribution.



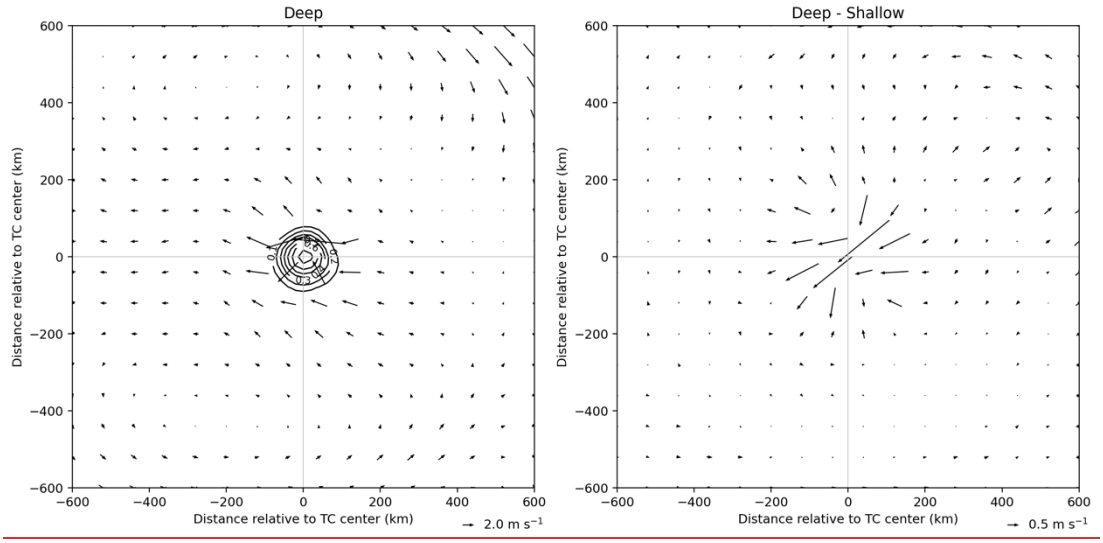


Figure ~~5~~7: Horizontal wind velocity vectors and vertical velocity contours at an altitude of 1500 m. **(Left)** ~~"Deep" case~~NHS wind field, showing a cyclonic circulation with a localized maximum in wind speed (contours, m/s) near the Tropical Cyclone (TC) center. Reference vector indicates a magnitude of 2.0 m~~-~~/s^{-1}. **(Right)** Difference in wind velocity between the ~~"Deep"~~NHD and ~~"Shallow"~~NHS cases (~~Deep – Shallow~~NHD – NHS). This panel highlights the variation in environmental flow and the intensity of the low-level circulation between the two experiments, with a reference vector of 0.5 m~~-~~/s^{-1}. All coordinates are relative to the TC center (km).

This dipole is dynamically consistent with the near-core easterly anomaly shown in Fig. ~~5~~7, which in turn is consistent with the south-westward displacement of the NHD trajectory. Figure ~~5~~7 shows the horizontal winds at 1.5 km. The left panel gives the circulation in the NHD run, and the right panel shows the wind difference between the deep and shallow simulations. The main feature in the difference field is an easterly anomaly near the storm center. This anomaly is important because it is consistent with the south-westward displacement of the NHD trajectory in Fig. ~~3~~5. In the Liang and Chan (2005) experiments, the additional Coriolis term associated with vertical motion produced a generally easterly flow near the TC center, and this easterly flow was identified as the immediate cause of the westward or south-westward displacement. Our result is not identical to their diagnosis, but the low-level easterly anomaly in Fig. ~~5~~7 points in the same direction. It suggests that the south-westward shift in the NHD trajectory is tied to a coherent change in the inner-core flow rather than to random differences accumulated during the integration.

At the same time, the present analysis remains more limited than that in Liang and Chan (2005). We have not yet constructed a full asymmetric wind composite like their ~~Fig~~Figure. 7, so the phase relationship among eyewall ascent, vorticity tendency, and asymmetric gyres is not shown explicitly here. For this reason, the mechanism discussed above should be taken as an interpretation of the present results rather than as a complete dynamical proof. The result from this test is simpler: the ~~GMCORE~~ deep-atmosphere ~~dycore~~GMCORE gives a TC evolution that is similar to the shallow-atmosphere run in intensity,

but different in trajectory, and this trajectory difference is accompanied by an organized inner-core asymmetry and a low-level easterly anomaly in the deep-minus-shallow wind field.

~~In summary, by strictly retaining the complete $2\Omega\cos\varphi$ dynamical terms, the GMCORE deep-atmosphere dynamical core successfully captures the asymmetric outflow mechanism excited by intense convective vertical motions, reasonably reproducing the south-westward deflection of the TC trajectory. This further verifies the physical fidelity and necessity of the nonhydrostatic deep-atmosphere dycore for high-resolution simulations of extreme weather events.~~

~~Existing studies have documented tropical-cyclone trajectory errors characterized by premature recurvature and north/eastward displacement, often linked directly to biases in the environmental steering flow. Meanwhile, deep-atmosphere and NCTs studies suggest that the omitted $2\Omega cos\phi$ terms can systematically modify tropical convective asymmetries and cyclone motion. Therefore, neglect of these terms may represent a dynamically plausible, but not yet conclusively established, contributor to such trajectory biases.~~

The systematic trajectory response found here can also be viewed in the broader context of tropical-cyclone asymmetric vortex dynamics. In the present simulations, retaining the NCTs introduces an organized inner-core forcing while leaving the axisymmetric intensity comparatively weakly affected. For an approximately axisymmetric eyewall updraft $w_0(r)$, the Cartesian meridional derivative can be expressed in storm-centered polar coordinates as

$$\frac{\partial}{\partial y} = \sin\theta\frac{\partial}{\partial r} + \frac{\cos\theta}{r}\frac{\partial}{\partial \theta}, \tag{49}$$

where θ is the azimuthal angle. Since w_0 is axisymmetric, $\frac{\partial w_0}{\partial \theta} = 0$, and therefore

$$\frac{\partial w_0}{\partial y} = \frac{dw_0}{dr}\sin\theta\ . \tag{50}$$

Hence, combining Eq. (48) and Eq. (50), we get $\Gamma = 2\Omega\cos\varphi\frac{dw_0}{dr}\sin\theta$, which projects directly onto an azimuthal-wavenumber-one vorticity tendency. This connection is potentially important because wavenumber-one disturbances have long been associated with displacement and tilt modes of tropical-cyclone-like vortices, whose evolution can involve vortex-Rossby-wave adjustment, precession, alignment, and trochoidal vortex-center motion (Polvani, 1991; Reasor and Montgomery, 2001; Nolan et al., 2001; Schecter et al., 2002). More recent analyses of strongly tilted vortices further show that dipolar vertical-motion asymmetries can modify vortex-centerline evolution and steering depending on their structure relative to the vortex tilt (Päschke et al., 2012; Dörffel et al., 2021). The NCT-related dipole and accompanying near-core easterly anomaly identified here therefore suggest that the retained $2\Omega\cos\varphi$ terms preferentially modify the low-wavenumber asymmetric circulation that is dynamically relevant to vortex motion. Determining how this forcing is partitioned among vortex-Rossby-wave adjustment, trochoidal displacement, and three-dimensional tilt or centerline precession provides an interesting direction for further investigation.

In summary, the NHD and NHS simulations remain similar in storm intensity and warm-core structure but exhibit a systematic difference in trajectory, with the NHD vortex displaced south-westward across the paired experiments. The NCT-related

vorticity-tendency dipole projects naturally onto a wavenumber-one asymmetry and is accompanied by a near-core low-level easterly anomaly, providing a dynamically consistent interpretation of the trajectory response. However, the present diagnostics do not establish a closed causal chain between the NCT forcing, the asymmetric circulation, and vortex translation. More complete analyses of the angular-momentum, potential-vorticity, and asymmetric-flow budgets will be required to quantify this mechanism.

4. Conclusions

In this study, we implemented ~~an extension for the~~ a deep-atmosphere extension within ~~a~~the nonhydrostatic dynamical core ~~(~~GMCORE~~)~~ and ~~validated~~evaluated the dycore~~'s performance~~ through numerical simulations of ~~ideal~~ baroclinic waves and tropical cyclones. The new version of GMCORE extends the original shallow-atmosphere equations to the deep-atmosphere ~~ones~~equations, retaining the full three-dimensional Coriolis force and variable gravity ~~terms~~, and incorporating the corresponding spherical–geometry ~~related~~ terms. A key feature of this extension is the ~~adoption~~use of a generalized dry terrain-following mass-based coordinate ~~system~~ and the ~~use of the~~ Wood and Staniforth (2003) ~~diagnostic formula~~formulation for diagnosing the distance from the Earth's center. This ~~ensures that the structural form of~~allows the deep-atmosphere ~~hydrostatic relations remains consistent with that of the shallow-atmosphere formulation, while retaining the advantages~~ formulation to retain the basic structure and numerical framework of the original solver. In this way, GMCORE has ~~evolved~~been extended from its original nonhydrostatic framework ~~into a dycore that supports~~to support fully compressible deep-atmosphere nonhydrostatic dynamics.

The extension was evaluated using ~~two~~ standard numerical tests and the idealized ~~test cases. In the deep~~ baroclinic~~-~~-wave experiment. The steady geostrophic-flow test shows near-second-order horizontal convergence, while the unperturbed standard-Earth baroclinic-wave basic state remains close to its balanced initial state throughout the integration. In the perturbed baroclinic-wave test, the deep- and shallow-atmosphere solutions remain close under the standard Earth-radius ~~configuration~~experiment, as expected when the atmospheric depth is ~~still~~ small relative to the planetary radius. In the reduced-radius X20 ~~configuration~~experiment, however, the two solutions separate clearly, and the deep-atmosphere version of GMCORE reproduces the deep–shallow differences reported in previous studies. We therefore regard ~~this test~~the X20 experiment primarily as a correctness-oriented benchmark~~, to check~~ for the new equation set and its numerical implementation. The tropical cyclone case highlights a different aspect of the deep-atmosphere extension. In this low-latitude, strongly rotating~~,~~ and vertically active regime, the deep~~-atmosphere effect~~–shallow difference is ~~therefore~~primarily associated with the retained ~~NCTs, the~~nontraditional Coriolis terms proportional to ~~$2\Omega\cos\phi$~~ $2\Omega\cos\varphi$, which couple vertical and horizontal motion in the tropical vortex core. ~~It~~The deep-atmosphere simulation produces a systematic south-westward trajectory shift relative to the shallow-atmosphere ~~configuration.~~simulation, while storm intensity and warm-core structure remain similar. This ~~difference~~displacement is accompanied by an organized inner-core asymmetry ~~and~~, including an NCT-related wavenumber-one tendency, and by a low-level easterly anomaly in the deep-minus-shallow wind field. ~~The result suggests~~These results

suggest that the retained NCTs can influence the evolution and motion of weather-scale vortices even when ~~the model top is not especially high and~~ the geometric part of the deep-atmosphere correction remains comparatively weak.

Several limitations should be kept in mind. The X20 baroclinic-wave case is useful for verification, but is not ideal for attributing deep–shallow differences purely to the ~~time~~temporal integration of the two equation sets, because part of the contrast may originate from differences in the analytic initial states. The tropical-cyclone experiment is likewise idealized, and the mechanism linking the ~~track shift~~trajectory displacement to the inner-core asymmetry still requires more systematic diagnosis, including analyses of ~~angular momentum, potential vorticity,~~vortex-Rossby-wave adjustment, vortex tilt and ~~asymmetric circulation.~~precession, and trochoidal center motion, and which of these processes contributes most directly to the simulated trajectory differences. Future work will therefore focus on extending ~~the~~these diagnostics, raising the model top together with appropriate physical parameterizations, and ~~applying the deep-atmosphere GMCORE to problems in tropical dynamies~~further evaluating gravity-wave and upper-atmosphere ~~modelling~~dynamics.

Overall, this study demonstrates that GMCORE can be extended from a shallow-atmosphere dycore to a fully compressible deep-atmosphere dycore while retaining the structural advantages of its original mass-based coordinate system and numerical design. ~~In future work, GMCORE~~Future applications will ~~continue to~~further investigate ~~tropical dynamics and~~ the ~~top~~role of ~~the model, while incorporating real physical processes to study the upper~~ deep-atmosphere dynamics in tropical and upper-atmosphere simulations.

Code and data availability

The deep-atmosphere extension code and scripts are on https://doi.org/10.5281/zenodo.19705647.

Author contributions

The deep-atmosphere extension ~~is a collaboration work between~~has been developed collaboratively by the authors. LD ~~finishes~~has completed the ~~based~~basic nonhydrostatic ~~work~~framework and ~~gives~~conceived the idea ~~of~~for this ~~work~~study. QC ~~finishes~~has implemented the deep-atmosphere ~~programming~~extension and ~~test. BW gives~~conducted the numerical experiments. BW has provided suggestions ~~to some modeling questiones.The~~on model development. All authors ~~write this paper~~have contributed to the writing and revision of the manuscript.

Competing interests

The authors declare that they have no conflict of interest.

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