



# 1 Evolution of Maud Rise Polynya during the last 250 years – a 2 multiproxy ice core reconstruction from coastal Dronning Maud 3 Land, Antarctica

4 Rahul Dey<sup>1,2</sup>, Chavarukonam M. Laluraj<sup>1</sup>, Kenichi Matsuoka<sup>3</sup>, Ashish Paiguinkar<sup>1</sup>, Bhikaji L  
5 Redkar<sup>1</sup> and Meloth Thamban<sup>1</sup>

6 <sup>1</sup>National Centre for Polar and Ocean Research (NCPOR), Ministry of Earth Sciences, Vasco da Gama, Goa  
7 403804, India

8 <sup>2</sup>Physics of Ice, Climate and Earth, Niels Bohr Institute, University of Copenhagen, Copenhagen N 2200,  
9 Denmark

10 <sup>3</sup>Norwegian Polar Institute, Framsentret, Postboks 6606, Langnes, 9296 Tromsø, Norway

11 *Correspondence to:* Rahul Dey (rdey1801@gmail.com)

## 12 Abstract

13 Open ocean polynyas drive deep ocean convection, influencing regional carbon and heat  
14 budgets, which in turn influence the ocean circulation and overall climate of Antarctica. The  
15 Maud Rise Polynya (MRP), also known as the Weddell Polynya, is one such polynya that forms  
16 in the Southern Ocean during early spring or winter months. The extensively studied MRP  
17 opening, which occurred during 2016-2017 and 1974–1976, triggered intense convection,  
18 ventilating heat from the deep ocean and modifying water mass properties. However, polynya  
19 evolution beyond the satellite era remains poorly understood. Here, we develop a polynya index  
20 using multiple proxy records from an ice core in coastal Dronning Maud Land, East Antarctica.  
21 Our approach—integrating records of snow accumulation,  $\delta^{18}\text{O}$ , deuterium excess, Na flux,  
22 and Na/SO<sub>4</sub> ratios—enhances polynya reconstruction, thereby overcoming the limitations of  
23 single-proxy methods. The index replicates the 1974–1976 polynya and extends the record to  
24 1774, revealing three major events comparable to the 1974–1976 great polynya event, in the  
25 past 250 years, totalling 47 polynya years. We identified distinct clusters of polynya activity,  
26 possibly corresponding to a specific combination of atmospheric circulation patterns and  
27 oceanographic preconditioning for MRP development. This study offers a long-term  
28 perspective on MRP variability, providing insights into its drivers and climate-related impacts.



## 29 1. Introduction

30 Open ocean polynyas represent one of the most dynamic and influential features in the  
31 Southern Ocean system, serving as critical regulators of heat and gas exchange between the  
32 ocean and atmosphere. These persistent areas of open water within the sea ice cover act as  
33 windows through which complex oceanographic and atmospheric processes interact,  
34 influencing global ocean circulation, carbon cycling, and climate patterns (Bennetts et al.,  
35 2024; The SO-CHIC consortium, 2023; Zheng et al., 2021). The significance of polynyas  
36 extends beyond their immediate vicinity, influencing deep-water formation, marine ecosystem  
37 dynamics, and global climate processes. In the Southern Ocean, polynyas manifest in two  
38 distinct forms: (1) coastal polynyas, which form along the Antarctic margin through  
39 mechanical forcing by katabatic winds, and (2) open-ocean polynyas, which develop hundreds  
40 of kilometres offshore through complex thermodynamic processes. While coastal polynyas  
41 occur regularly and have been extensively studied (Arrigo and van Dijken, 2003; Årthun et al.,  
42 2013; Jacobs et al., 1979; Visbeck et al., 1996; Xu et al., 2023), open-ocean polynyas represent  
43 rare but profoundly influential events that can significantly impact global ocean circulation  
44 patterns. The Maud Rise Polynya (MRP), also known as Weddell Polynya, occurring over the  
45 Maud Rise seamount in the eastern Weddell Sea (Fig. 1), is the most enigmatic open-ocean  
46 polynya in the Southern Ocean (Holland, 2001). First documented through satellite  
47 observations in the 1970s, the MRP reached its largest extent during the winter and early spring  
48 months from 1974 to 1976, exceeding 300,000 km<sup>2</sup> at its peak (Carsey, 1980; Zwally et al.,  
49 1985). This event drove intense ocean convection, reaching depths of over 3000 meters, and  
50 fundamentally altered water mass properties throughout the region, influencing global ocean  
51 circulation patterns (Gordon and Comiso, 1988).

52 Following this dramatic episode, the Maud Rise region has exhibited intermittent  
53 polynya activity, typically manifesting as halos of reduced sea ice concentration (Lindsay et  
54 al., 2004; McHedlishvili et al., 2022). However, the phenomenon garnered renewed attention  
55 when a large polynya exceeding 40,000 km<sup>2</sup> reappeared during the two winters of 2016 and  
56 2017 (Jena et al., 2019). Early polynya studies focused primarily on documenting polynya  
57 occurrence and extent through satellite imagery (Carsey, 1980; Comiso and Gordon, 1987;  
58 Zwally et al., 1985), while more recent works have attempted to unravel the complex  
59 mechanisms controlling polynya formation and maintenance (de Lavergne et al., 2014; Wilson  
60 et al., 2019). Satellite observations have revealed that open-ocean polynya formation often



61 follows a characteristic pattern, beginning with the appearance of small-scale openings that can  
62 rapidly expand under favourable conditions, such as wind-driven upwelling (de Lavergne et  
63 al., 2014; Jena et al., 2019), baroclinic instabilities (Akitomo, 2006; Campbell et al., 2019),  
64 and deep warm water intrusions (Gülk et al., 2024; Heuzé et al., 2021). However, the precise  
65 triggers initiating this process remain incompletely understood, because the relative importance  
66 of these processes likely varies temporally and may be interconnected (Campbell et al., 2019;  
67 McPhee, 2003; Narayanan et al., 2024), and also because large polynyas openings are rare and  
68 satellite-based observations range only for the past few decades, limiting the observational  
69 opportunity. Similarly, model-based assessments are limited in their ability to understand  
70 polynya formation in the weakly stratified Southern Ocean (Heuzé et al., 2013; Sallée et al.,  
71 2013). This modelling challenge hampers our ability to assess both the role of polynyas in  
72 natural climate variability and their potential responses to future climate change.

73 To obtain longer polynya records, Goosse et al. (2021) examined surface mass balance  
74 (SMB) records reconstructed from six ice cores and two automatic weather stations. In these  
75 records, they found that increased snowfall during the 2016-17 MRP opening assumed that this  
76 had also occurred in past MRP opening periods. Their study constructed a polynya index using  
77 two data assimilation approaches and one statistical method to reconstruct historical climate  
78 states based on SMB reconstruction from ice cores. Their study found multiple polynya  
79 openings since 1250 CE, but observed that large polynya openings are rare. However, they rely  
80 only on snow accumulation records, which can be influenced by many factors other than  
81 polynya occurrences. The polynya signal in their ice-core SMB records is also weak compared  
82 to natural atmospheric variability, making it difficult to distinguish polynya-induced anomalies  
83 from SMB caused by other climatic fluctuations. The short instrumental record from automatic  
84 weather stations prevents a robust calibration of polynya reconstructions, increasing the risk of  
85 false positives (unrelated high snowfall events misinterpreted as polynya activity) and false  
86 negatives (missed polynya events). Additionally, the ice-core and instrumental records they  
87 used were collected from a large region, including the Weddell Sea sector and the ice sheet  
88 inland, where the surface elevation exceeds 2000 m a.s.l., which can hardly reflect changes in  
89 the vapour sources in the MRP regions directly (Fig. 1).

90 To overcome the limitations of the previous study, we develop a novel, multi-proxy  
91 approach to reconstruct past MRP activity using a high-resolution ice core, IND-36/9, from



92 Djupranen Ice Rise in coastal Dronning Maud Land. While polynya opening results in  
 93 increased heat and moisture exchange, the polynya years, characterised by a decrease in sea ice  
 94 cover, also result in increased sea salt production. These signatures can be carried to the coastal  
 95 region of Antarctica and deposited in the low-elevation ice shelves and ice rises. The back-  
 96 trajectory analysis (Section x) clearly indicates the direct atmospheric link between MRP and  
 97 the IND-36/9 core site (Fig. 1). Our ice core site is situated within ~1000 km of the typical  
 98 MRP location, close enough to potentially capture atmospheric signatures transported from  
 99 polynya formation near the Maud Rise. By combining records of snow accumulation, water  
 100 isotopes, and major ion concentrations, we aim to provide a more robust reconstruction of  
 101 polynya activity. This approach leverages the strengths of selecting an idea core site and  
 102 employing multiple proxies to overcome the limitations of single-proxy reconstructions,  
 103 thereby capturing the complex signatures of polynya events that extend beyond the last few  
 104 decades, when satellite and instrumental records are available.

## 105 **2. Study Area and Methods**

### 106 **2.1. Study area**

107 The DML coast is characterised by distinct topographic features like ice rises having  
 108 associated local ice flow, climate regime and SMB variability (Drews et al., 2015; Goel et al.,  
 109 2017; Lenaerts et al., 2014; Matsuoka et al., 2015; Pratap et al., 2022; Rignot et al., 2019). As  
 110 part of the Indo-Norwegian project MADICE, an ice core was drilled on the summit of the  
 111 Djupranen Ice Rise (70.18° S, 9.18° E; elevation 321 m a.s.l.), at the western margin of the  
 112 Nivlisen Ice Shelf in coastal Dronning Maud Land (DML), East Antarctica (Fig. 1). Our site  
 113 survey found that this ice rise summit has been at the current position at least in the past few  
 114 millennia (Pratap et al., 2022). This location also offers several advantages for polynya  
 115 reconstruction: its high SMB allows annual layer counting (Dey et al., 2023) and its coastal  
 116 proximity may provide sensitivity to maritime signals (Ejaz et al., 2021; Wauthy et al., 2024),  
 117 while its elevation causes insignificant surface melt and local noise while maintaining regional  
 118 signal strengths (Dey, 2023). Its position is also within primary atmospheric transport pathways  
 119 from the Maud Rise region (Fig. 1).



## 120      **2.2. Ice core: from the field to the laboratory**

121            The drill site of the Djupranen ice rise was located at its summit, based on analysis of  
122            satellite altimetry data and satellite image analysis, followed by an ice-penetrating radar survey  
123            (Pratap et al., 2022). The radar survey also revealed Raymond arches, indicative of a relatively  
124            stable summit position in the past (Goel et al., 2020). The ice core was drilled using an  
125            electromechanical ice core drilling system (Model D2, GeoTec, Japan). Over nine days, a 122  
126            m ice core (IND 36/B9; hereafter IND36/9) was retrieved. The drilled ice core sections were  
127            sealed in high-density polyethylene core bags, packed in expanded polypropylene boxes, and  
128            stored in a reefer container at  $-20^{\circ}\text{C}$  until their transport to the National Centre for Polar and  
129            Ocean Research (NCPOR), Goa, India, where samples were stored in the in-house Ice Core  
130            Laboratory maintained at  $-20^{\circ}\text{C}$ . The ice cores were processed in the  $-15^{\circ}\text{C}$  core processing  
131            facility at NCPOR. The cores were initially cut into 3.5 cm thick, 10 cm wide slabs for line  
132            scanning and then subsequently sub-sampled at 5 cm resolution for stable isotope and chemical  
133            analysis. The samples for chemical analysis were cut into cuboids; the three dimensions of the  
134            samples were measured using a calliper and weighed using a weighing balance. The density  
135            for the samples was calculated as mass divided by the volume of each sample. The error in  
136            measuring the sample dimensions was  $\pm 0.5$  mm, while the weighing balance's uncertainty was  
137             $\pm 0.1$  g for measurements up to 100 g. As a result, the density measurements have a propagated  
138            uncertainty of 5%.

139            Major inorganic anions ( $\text{Cl}^{-}$ ,  $\text{SO}_4^{2-}$ ,  $\text{MSA}^{-}$  and  $\text{NO}_3^{-}$ ) and cations ( $\text{Na}^{+}$ ,  $\text{NH}_4^{+}$ ,  $\text{K}^{+}$ ,  $\text{Mg}_2^{+}$ ,  
140            and  $\text{Ca}_2^{+}$ ) were measured in the samples using an ICS 5000+ ion chromatograph (Thermo  
141            Dionex) equipped with a conductivity detector. Anions were separated on an AS11 (2 mm)  
142            column with potassium hydroxide as eluent and an AG11 (2 mm) guard column with AERS  
143            500, 2 mm suppressor. Cations were separated on a CS17 (0.4 mm) capillary column with  
144            methane sulphonic acid as the eluent and the CG17 (0.4 mm) capillary guard column with a  
145            CCES 300 capillary suppressor. A 10 mg/L stock solution of  $\text{Na}^{+}$ ,  $\text{NH}_4^{+}$ ,  $\text{K}^{+}$ ,  $\text{Ca}^{2+}$ , and  $\text{Mg}_2^{+}$   
146            was mixed and then diluted with MilliQ ultrapure water to prepare standards for cation  
147            exchange chromatography. The anion standards were prepared from 10 ppm stock solutions of  
148             $\text{MSA}^{-}$ ,  $\text{Cl}^{-}$ ,  $\text{SO}_4^{2-}$ , and  $\text{NO}_3^{-}$ . The dilutions were conducted volumetrically and were freshly  
149            prepared within a few days of each run, ranging from 5 ppb to 1 ppm. Eight standards were  
150            selected from this range for calibration. Before analysis, the ice core samples were melted in a  
151            Class 100 clean room. The analytical precision for all ions was better than 10%.



152           The ice core samples were analysed for oxygen and hydrogen isotopic ratios at NCPOR  
153 using a Triple Isotope Water Analyser (TIWA-45EP from Los Gatos Research, USA), which  
154 works on the principle of off-axis integrated cavity output spectroscopy (OA-ICOS). The  
155 melted ice core samples were introduced into the TIWA-45EP without sample conversion  
156 through a PAL HTC-xt auto-injector (CTC Analytics) equipped with a heated ( $\approx$ approximately  
157 85 °C) injector block (LGR) (Berman et al., 2013). Using a Hamilton 1.2  $\mu$ L, zero-dead-volume  
158 syringe, samples were injected into the injector block and evaporated for direct isotope  
159 analysis. Measurements were completed at a speed of  $\sim$ 90 s per individual injection. To  
160 eliminate sample-to-sample memory, a total of nine injections were made, with the first three  
161 being discarded for analysis. The last six injections were averaged to produce a single, high-  
162 throughput (HT) sample measurement. One commercially available working standard from  
163 LGR1C and two in-house laboratory standards (CDML1 and HL1) with known isotopic  
164 composition, spanning the entire range of our sample measurements ( $-46.19\text{‰}$  to  $-19.49\text{‰}$   
165 for  $\delta^{18}\text{O}$  and  $-362.85\text{‰}$  to  $-154.0\text{‰}$  for  $\delta\text{D}$ ) were analysed routinely as reference waters after  
166 every five ice core samples to check the instrument performance. Laboratory standards are  
167 calibrated on the VSMOW/SLAP scale. The external precision obtained using our laboratory  
168 standards (CDML1 and HL1) for  $\delta^{18}\text{O}$  was  $\pm 0.046\text{‰}$  and  $\pm 0.068\text{‰}$ , respectively, and for  $\delta\text{D}$   
169 was  $\pm 0.32\text{‰}$  and  $\pm 0.23\text{‰}$  ( $1\sigma$  standard deviation) for 30 samples. Replicate analyses  
170 performed based on ten samples yield repeatability of  $\pm 0.76\text{‰}$  for  $\delta\text{D}$  and  $\pm 0.09\text{‰}$  for  $\delta^{18}\text{O}$ .  
171 All the raw instrumental OA-ICOS data were processed in the LGR post-analysis software.  
172 Any measured injection with water number density outside the manufacturer's suggested range  
173 of  $2\text{--}4.5 \times 10^{16} \text{ H}_2\text{O molecules/cm}^3$  was discarded. Injections with incomplete evaporation were  
174 detected by examining the standard deviation of the measured water number density ( $\sigma_{\text{nmeas}}$ ) as  
175 reported by the instrument (Berman et al., 2013). Processed raw data directly gives  $\delta^{18}\text{O}$  and  
176  $\delta\text{D}$ , which are further used to calculate deuterium excess [ $\text{d-excess} = \delta\text{D}_{\text{ice}} - 8 \times \delta^{18}\text{O}_{\text{ice}}$ ].

### 177           **2.3. Chronology development**

178           The chronology of the IND36/9 core is based on a multiproxy approach involving  
179 annual layer determination from the stratigraphy of  $\delta^{18}\text{O}$ , major ions, and pixel intensity data  
180 from the line scanner, following Dey et al. (2023). The core has experienced limited surface  
181 melt, with annual melt proportion varying between 0 and 4.4%, with a median melt proportion  
182 of 0.25% (Dey et al., 2023). Since diffusion in the firn attenuates high-frequency water-



183 isotope information in ice cores (The Firn Symposium team, 2024), even in high accumulation  
184 sites of coastal Antarctica (Mahalinganathan et al., 2022), we diffusion-corrected our water  
185 isotope records following Jones et al. (2023) as detailed in the supplementary materials (Fig.  
186 S1). The shift in seasonal peaks in our corrected record is less than 5 cm, which falls within  
187 the sampling interval of 5 cm and is therefore insignificant in affecting the accuracy of the  
188 chronology.

189 A five-point smoothing was applied to the pixel intensity data to simplify annual layer  
190 counting and reduce noise in the record. We also used age tie points of known volcanic  
191 eruptions identified from non-sea-salt sulphate (nssSO<sub>4</sub>) peaks, as well as the tritium bomb  
192 peak of 1962. Volcanic indicators (nssSO<sub>4</sub>) have been used to identify specific, dated volcanic  
193 eruptions, allowing us to reduce the uncertainties resulting from the relative dating procedure.  
194 However, unambiguous eruption identifications are challenging in ice cores from coastal  
195 regions, where the nssSO<sub>4</sub> background signals are commonly highly variable due to the  
196 proximity of the ocean and ocean-related MSA products (Philippe et al., 2016).

197 A preliminary chronology was obtained from the annual counts, which was then  
198 refined using the nssSO<sub>4</sub> peaks (Fig. 2) from the well-established volcanic eruptions of  
199 Pinatubo (1991), El Chichon (1982), Agung (1963), Cerro Azul (1932), Santa Maria (1902),  
200 Krakatoa (1882), Cosiguina/Babuyan (1834), Tambora (1815) and unknown volcanic  
201 eruption (1809). Similar to Dey et al. (2023), we refined the chronology between the volcanic  
202 and Tritium tie points using the StratiCounter algorithm (Winstrup et al., 2012). The manual  
203 counts provide a basic framework from which StratiCounter develops and refines its statistical  
204 characterisation of annual layers, enabling adaptation to varying layer properties with depth  
205 (Winstrup et al., 2012). To minimise reliance on initial manual inputs, we conducted multiple  
206 iterations using refined layer templates derived from previous algorithm outputs. The  
207 chronology of the IND36/9 ice core provides a robust temporal framework that extends back  
208 to 1774 CE at a depth of 122m (Fig. 3).

#### 209 **2.4. Satellite-based polynya metrics**

210 We followed Heuzé et al. (2021) to define the "polynya-prone" region of the Weddell  
211 Sea (6°W to 12°E, 68°S to 60°S, approximately 600 km x 900 km), focusing on the winter and  
212 early spring months (June 1 to October 31). This region was selected based on historical



213 observations of polynya formation and its significance in Antarctic bottom water production  
214 (Campbell et al., 2019; Heuzé et al., 2021). Polynya activity was quantified using two  
215 independent metrics derived from satellite data. These metrics complement each other to  
216 provide a comprehensive assessment of polynya dynamics.

217         The first metric, Polynya Days, represents the annual count of days (between June 1  
218 and October 31) with sea ice concentration below 60% in our study zone. This threshold was  
219 determined through sensitivity analysis of satellite imagery and validation against in situ  
220 observations from previous studies (Heuzé et al., 2021). Sea ice concentration data were  
221 derived from passive microwave measurements using the NASA Team algorithm (Comiso and  
222 Nishio, 2008), which has demonstrated robust performance in discriminating between open  
223 water and sea ice in polar regions. The second metric, Cumulative Polynya Area, measures the  
224 total extent of all polynya occurrences within a year (June 1 to October 31). This metric was  
225 calculated by summing the daily polynya areas, defined as contiguous regions with sea ice  
226 concentration below the 60% threshold. The combination of these metrics allows us to  
227 characterise both the temporal persistence and spatial extent of polynyas, providing insights  
228 into their formation mechanisms and potential impact on regional oceanographic processes.

## 229         **2.5. Air mass trajectory modelling**

230         The HYbrid Single-Particle Lagrangian Integrated Trajectory (HySPLIT) model,  
231 developed by the NOAA Air Resources Laboratory (ARL), provides a means of generating  
232 back trajectories to identify the source of moisture uptake for precipitation (Markle et al.,  
233 2012). We used HYSPLIT version 5.3 for back trajectory computations, initialized with  
234 conditions from the 2.5 by 2.5-degree resolution NCEP/National Center for Atmospheric  
235 Research (NCAR) reanalysis data for the period 1948 to 2016. Forward trajectories were run  
236 for 240 hours (10 days), initialised every hour, starting over the polynya-prone region of the  
237 Weddell Sea. Groups of forward trajectories were computed with initial altitudes of 100, 200,  
238 300, 400, and 500 meters above ground level, and in all cases, the vertical velocity from the  
239 meteorological data was used as the input for vertical motion.

240         We calculated trajectory density from multiple trajectories using HYSPLIT (Miller et  
241 al., 2002), because individual trajectories are highly sensitive to meteorological uncertainties,  
242 turbulence, and small initial condition variations, often leading to misleading or unreliable





243 airflow pathways and potential position errors of up to 20% of the distance travelled. It also  
 244 aids in discriminating between local circulation patterns and regional-scale flow features  
 245 (Dorling and Davies, 1995) and has been previously applied to interpret polar ice core  
 246 paleoclimate records (Dixon et al., 2012; Ejaz et al., 2021; Neff and Bertler, 2015). We  
 247 calculated air mass transport densities for each year 1979 – 2016 from the HySPLIT output.  
 248 The trajectory endpoints in each equal-area (1 by 1 degree) oceanic pixel were summed and  
 249 divided by the total number of air mass trajectories.

## 250 **3. Results**

### 251 **3.1. Identification of polynya years in satellite data**

252 Satellite imageries are crucial for polynya studies as they provide continuous, high-  
 253 resolution observations of sea ice dynamics, allowing researchers to monitor polynya  
 254 formation, extent, and variability over time. Satellite observations spanning four decades reveal  
 255 distinct periods of polynya activity in the Weddell Sea region since 1979 (Fig. 4). The most  
 256 notable and well-documented event occurred during 1974-1976, when the polynya reached an  
 257 exceptional size of 300,000 km<sup>2</sup> (Carsey, 1980). This event, often referred to as the Great  
 258 Weddell Polynya, represented a significant perturbation to the regional ocean-atmosphere  
 259 system and has served as a benchmark for subsequent polynya observations. Following this  
 260 major event, several smaller polynyas and polynya-like features ("halos") have been observed  
 261 during the late 1980s to early 1990s and early 2000s (Heuzé et al., 2021). The two independent  
 262 metrics used for polynya detection—Polynya Days and Cumulative Area—reveal distinct but  
 263 complementary temporal patterns. This dual-metric approach enables a more comprehensive  
 264 understanding of polynya dynamics than either metric alone could provide. While the  
 265 magnitudes of these metrics do not exhibit direct correlation ( $r = 0.42$ ,  $p < 0.01$ ), both metrics  
 266 successfully identify known major polynya/halo events between 1979 and 2016, and provide  
 267 unique insights into their temporal and spatial characteristics. The Polynya Days metric  
 268 effectively captures persistent small-scale features, while the Cumulative Area metric better  
 269 represents brief but large openings. While the record of polynya openings from satellite  
 270 imagery provides crucial insights, their influence on coastal Antarctic ice cores could depend  
 271 majorly on atmospheric transport.



## 272      **3.2. Air mass transport patterns**

273            We use forward air trajectory analysis to examine the transport pathways and determine  
 274 whether polynya-derived signals can reach and be recorded in the ice cores. This approach  
 275 allows us to trace the path of air masses from the polynya-prone area to potential ice core sites,  
 276 providing insight into which locations are most likely to capture polynya signals. Forward  
 277 trajectory analysis reveals consistent and well-defined transport pathways from the Maud Rise  
 278 region to coastal Dronning Maud Land (Fig. 5). Frequency analysis indicates that our study  
 279 site receives approximately 2% of all trajectories originating from the MRP region, a  
 280 statistically significant proportion ( $p < 0.001$ ) that indicates reliable capture of polynya-related  
 281 atmospheric signals. This percentage remains relatively stable across different seasons and  
 282 years, suggesting a robust atmospheric connection between the source and deposition regions.  
 283 The trajectory density analysis indicates that our ice core location falls within a primary  
 284 atmospheric transport corridor originating from the MRP region. This corridor exhibits  
 285 enhanced stability during the winter months when polynya formation typically occurs.  
 286 However, trajectory calculations are susceptible to significant spatial error of 15 – 30% of  
 287 distance travelled (Draxler, 2008). Therefore, a combination of forward and backward  
 288 trajectories would provide a better overview of the travel pathways. Our back-trajectory  
 289 analysis (Fig. 6) further confirms the findings from the forward trajectory patterns, showing  
 290 trajectories originating from the polynya-prone region influencing our ice core site. Trajectory  
 291 analysis of the transport pathways, therefore, confirms that winds can carry polynya-derived  
 292 signals from the open water to our study site and preserve them in the ice core records,  
 293 regardless of polynya presence.

## 294      **3.3. Development of the polynya index**

### 295      *3.3.1. Ice core record and proxy selection*

296            The ice core proxy dataset from 1774 to 2016 exhibits distinct variability across all  
 297 measured properties. We choose five major proxies for developing our polynya index: sea-salt  
 298 sodium (ssNa), deuterium excess (d-excess),  $\delta^{18}\text{O}$ , snow accumulation, and the Na/SO<sub>4</sub> ratio  
 299 (Fig. 7a). Sodium concentration ranges from a low of 19.10 to 218.64 ppb, with a mean value  
 300 of 89.87 ppb. Some periods show sustained higher values, such as the late 1700s and early  
 301 1800s, while others, including the late 19th and 20th centuries, display more variable levels.



302 D-excess varies between -1.83 to 10.75‰, with a mean of 4.31‰, showing notable peaks in  
303 the late 19th century and sharp declines in the early 20th century.  $\delta^{18}\text{O}$  values fluctuate between  
304 -21.64 to -15.65‰, with a mean of -17.85‰, displaying alternating periods of enrichment and  
305 depletion, including a gradual decline over recent decades. Annual snow accumulation ranges  
306 from 0.12 to 0.83 m w.e., with a mean of 0.39 m w.e., exhibiting distinct fluctuations, including  
307 lower values in the early 20th century and a more variable pattern in recent decades. The  
308 Na/SO<sub>4</sub> ratio ranges from 0.02 to 1.31, with a mean of 0.42, exhibiting periodic fluctuations  
309 throughout the dataset.

310 Our polynya index incorporates the five selected parameters, each chosen based on  
311 established physical mechanisms that link polynya formation to specific signatures in ice core  
312 records. The opening of a polynya results in increased heat and moisture exchange between the  
313 warm, open water and the cold air. Therefore, during polynya years (decrease in the sea ice  
314 cover), there is also an increase in local precipitation in the Weddell Sea region (Moore et al.,  
315 2002) and resultant snow accumulation further inland (Goosse et al., 2021). The water isotope  
316 ratios of precipitation are often related to temperature at the precipitation site (Ejaz et al., 2022;  
317 Naik et al., 2010) and the distance of transport (Goursaud et al., 2017; Klein et al., 2019), and  
318 the pattern of annual mean water isotopes of precipitation associated with polynya formation  
319 would relatively be similar to that for increased temperature from the heat exchange due to  
320 polynya opening and shorter transport distance, especially over a coastal site. Extensive open  
321 water polynyas can also act as a large factory for sea ice production, comparable to the  
322 production of the largest coastal polynyas (Zhou et al., 2023). Previous studies have  
323 demonstrated that the surface of fresh sea ice, including frost flowers, is the major source of  
324 sea salt to the Antarctic, and the production of frost flowers is controlled by the amount of new  
325 sea ice production (Wolff et al., 2003). Therefore, the formation of large open-ocean polynyas  
326 would lead to an increased supply of sea salt sodium to the ice core site.

327 Additionally, Sulphate (SO<sub>4</sub>) fractionation during freezing of seawater (ice) is another  
328 well-known process to characterise sea ice coverage (Richardson, 1976). It is driven by the  
329 crystallisation of Na<sub>2</sub>SO<sub>4</sub>·10H<sub>2</sub>O (mirabilite) occurring at temperatures below -8.2°C. Since the  
330 salt crystals are associated with the ice lattice, the remaining brine becomes increasingly  
331 depleted in SO<sub>4</sub> (and to a lesser extent in Na) as it progresses below this critical temperature.  
332 This process, therefore, increases the Na/SO<sub>4</sub> ratio of the sea salt source relative to bulk



333 seawater by preferentially removing sulphate ions (Wagenbach et al., 1998). Since this process  
334 only occurs at temperatures below  $-8.2^{\circ}\text{C}$ , temperatures not encountered at the summer ocean  
335 surface, but frequently encountered at the winter sea ice surface (Levine et al., 2014). The  
336 anomalously high  $\text{Na}/\text{SO}_4$  ratio can be used as an indicator of sea salt sourced from the  
337 production of fresh sea ice at the polynyas.

### 338 *3.3.2. Development of the integrated polynya index*

339 Our ice core record shows an anomalous increase in all proxy records during the  
340 prominent polynyas of 1964 and 1974–1976 (Fig. 7a). However, in the case of the transient  
341 polynyas and halos in more recent years, the relative changes are weaker but still identifiable.  
342 This could be due in part to the inherent complexity of air-ocean interactions and transportation  
343 processes, as well as the uncertainty associated with the measurements and chronology of the  
344 ice core. Therefore, to develop a more robust proxy, rather than using the proxies individually,  
345 we derive a multi-proxy index to quantify the likelihood of polynya occurrence more  
346 accurately.

347 The polynya index is defined using anomalous changes in snow accumulation,  $\delta^{18}\text{O}$ ,  
348 deuterium excess,  $\text{Na}_{\text{flux}}$ , and  $\text{Na}/\text{SO}_4$  ratio as possible indicators of polynya activity. We used  
349 Monte Carlo simulations to account for uncertainties associated with these individual proxies  
350 and ultimately to constrain better the influence of these uncertainties on the multi-proxy  
351 polynya index. This is because individual measurement uncertainties are independent of  
352 chronological uncertainties. We run ten thousand simulations to minimise statistical bias and  
353 define background data as the sum of a 30-year running median (RM) and two times of median  
354 absolute deviations (MAD). A polynya year (a year when a polynya activity is recorded) is  
355 defined as a year in which the polynya index exceeds the likelihood threshold (0.6, 0.8, or 1).  
356 The polynya index is, therefore, an indicator of polynya occurrence from ice core observations  
357 and not a measure of the absolute extent or duration of the polynya events.

### 358 **3.4 Past Maud Rise Polynya activity**

359 Our ~250-year reconstruction reveals complex and significant temporal variability in  
360 polynya occurrence, shedding new light on the long-term behaviour of the Maud Rise Polynya  
361 (MRP). We identified forty-seven distinct polynya years from 1774 to 2016 (Fig. 7b); however,  
362 the period from 1920 to 1950 was the longest interval without significant polynya activity in



our record. The 1830s were the most active decade, with seven polynya years recorded between 1830 and 1840. This high frequency of events suggests that atmospheric and oceanic conditions during this period were highly conducive to polynya formation. Similar clusters occurred in the 1880s, with five polynya years, and in the 1970s, with four polynya years, including the well-documented 1974-1976 sequence (Fig. 7b). These clusters suggest potential links to large-scale climate oscillations or recurring oceanographic conditions that favour polynya development. The reconstruction also identifies two past events (1833 – 1884 and 1911) that produced index values comparable to or exceeding those of the well-documented 1974-1976 polynya. The high index values for these events likely indicate that they were large-scale and persistent openings; however, exact quantification of the polynya extent is beyond the scope of this study.

## 4. Discussion

### 4.1. Performance of the new polynya index

We compared our polynya index with those from Goosse et al. (2021) and found that all major polynyas identified in their study over the common time period are also identified in this study. This provides a degree of validation for both approaches, suggesting that they are capturing similar large-scale polynya events. However, our index shows enhanced sensitivity to polynya formations, identifying the polynya occurrence of 1964, even though with a low likelihood. This low likelihood of the 1964 polynya is possibly due to the lower persistence as compared to the well-documented events during 1974 – 1976 (Meier et al., 2013). Furthermore, our index shows better performance in identifying transient polynyas and halos, which were missed in the reconstruction by Goosse et al. (2021). These shorter-lived or less intense polynya events, while not as dramatic as major openings, play a crucial role in regional oceanography and climate dynamics. Their detection provides a more comprehensive picture of polynya activity over time.

Visual inspection shows that some years show synchronised peaks across the indices from the two studies, with leads/lags associated with the timescale of the ice cores (Fig. 8). However, some peaks are present in our polynya index which are absent/muted in Goosse et al. (2021). This discrepancy between the two records is possibly due to the selection of ice core sites used for calculating the polynya index. Our analysis of air mass transport frequency from



the MRP regions shows a sink in the coastal regions of Dronning Maud Land, with a significant proportion of the trajectories reaching our study area. In contrast, four out of the six ice core sites used in Goosse et al. (2021) fall outside the primary sink of these trajectories (Fig. 5). This difference in site selection could be a major reason for the observed discrepancy between the two polynya indices. Ice cores from sites that rarely receive air masses from the polynya region may not be reliably used to reconstruct polynya events or may do so with reduced sensitivity. Whereas our ice core site is well located to capture the occurrence of the MRP opening, as it frequently receives air masses originating from the polynya region. However, they still manage to detect the large, multi-year polynyas as they lead to a widespread positive anomaly in precipitation over the coastal and continental region. Another possible reason for this observed difference in sensitivity could be the robustness of our multiproxy approach in identifying polynyas, as it is less susceptible to noise or artefacts in any single proxy. It is crucial to note that the observed differences in peak occurrences do not necessarily invalidate either dataset but rather highlight the complexity in reconstructing past polynya activity. These discrepancies underscore the challenges inherent in paleoclimate reconstruction and the potential complementarity of different methodologies in building a comprehensive understanding of past polynya dynamics.

#### 4.2. Historical polynya activity and its link to maritime climate

Our polynya reconstruction aligns with and extends satellite era observations of MRP formation mechanisms. Recent studies have shed light on the complex interplay of oceanic and atmospheric factors controlling MRP formation and persistence (Campbell et al., 2019; Jena et al., 2019). However, extending these insights to longer timescales remains challenging. Our multi-century reconstruction offers a unique opportunity to explore the long-term drivers of MRP variability. Autonomous profiling float observations during 2016 and 2017 revealed that the MRP were initiated and modulated by the passage of severe storms, and the intense heat loss drove deep overturning within them (Campbell et al., 2019). Wind-driven upwelling of record strength weakened haline stratification in the upper ocean, thus favouring destabilisation. A recent study, however, highlighted the role of salt transport through northward Ekman transport as an additional mechanism of the polynya formation and these processes were driven by intensified eastward surface stresses during 2015 – 2018 (Narayanan et al., 2024).



424           To investigate these factors during the past polynya openings, we used ocean reanalysis  
425 outputs from the Simple Ocean Data Assimilation (SODA) version 2.2.4, which spans the  
426 period from 1871 to 2010 (Carton and Giese, 2008; Giese and Ray, 2011). It is forced using  
427 data assimilation of in-situ temperature and salinity profiles and in-situ and satellite SSTs. This  
428 long-term dataset provides monthly fields of ocean temperature, salinity, currents, and sea  
429 surface height at a horizontal resolution of  $\sim 0.5^\circ$ , making it particularly suited for studying  
430 large-scale oceanic features and their variability over decadal time scales. Our analysis reveals  
431 that periods with high likelihood of polynya occurrences are consistently characterised by  
432 increased eastward zonal wind stress due to westerly wind, enhanced northward water flow,  
433 and elevated sea surface salinity (Fig. 9). This consistency lends credence to current theories  
434 of MRP formation mechanisms and suggests that similar processes operate on multi-decadal  
435 to centennial timescales. However, we also observed multiple periods of increased meridional  
436 velocity, zonal wind stress, and sea surface salinity that did not invariably lead to a high  
437 likelihood of polynya occurrence. This non-linearity underscores the complexity of polynya  
438 formation processes and suggests the existence of critical thresholds in wind stress, ocean  
439 circulation, or stratification that must be exceeded for polynya formation to occur. The  
440 temporal sequencing and duration of favourable conditions also play crucial roles in polynya  
441 initiation and maintenance. It could also mean that there are still one or more unexplored  
442 processes that control the opening of major polynyas in the Weddell Sea. Our multiproxy  
443 approach, combining snow accumulation, water isotopes, and major ion concentrations,  
444 provides a more robust reconstruction of polynya activity compared to single-proxy studies.  
445 This approach allows us to capture diverse signatures of polynya events and mitigate the  
446 limitations of individual proxies.

447           Although beyond the scope of this study, it is crucial to understand the maritime climate  
448 variability during the polynya and non-polynya years. Our reconstruction shows alternation  
449 between polynya-active and quiescent periods, which may provide valuable context for  
450 assessing how future MRP behaviour may evolve under anthropogenic climate change. The  
451 clustering of MRPs in our record possibly indicates that once favourable preconditioning  
452 develops, the region remains susceptible to repeated polynya openings over several years or  
453 decades. As global temperatures rise, projected changes in Southern Ocean wind patterns, sea-  
454 ice extent, and surface freshening may shift the thresholds required for deep convection,  
455 thereby altering both the frequency and persistence of MRPs. Understanding how atmospheric



456 and oceanic conditions differ between the polynya and non-polynya periods identified in this  
457 study will be critical for refining projections. Integrating these insights with high-resolution  
458 coupled ocean–atmosphere models could help identify the physical thresholds and feedbacks  
459 governing MRP formation. Such integration is crucial for predicting how rare but climatically  
460 significant features such as the Maud Rise Polynya may influence Southern Ocean overturning,  
461 carbon exchange, and global climate in a warming world.

## 462 **5. Conclusions and Outlook**

463 We provide new insights into the temporal variability of Maud Rise Polynya, one of  
464 the largest and well-known open ocean polynya, by integrating multiple proxies using an ice  
465 core from coastal Dronning Maud Land, representing the past nearly 250 years. We identified  
466 multiple polynya openings during the 1830s, 1880s, and 1970s. These clusters of polynya  
467 openings possibly correspond to a specific combination of atmospheric circulation patterns and  
468 oceanographic preconditioning, suggesting a more deterministic framework for MRP  
469 development. Over the 250-year record, we identify large events during 1833 – 34 and 1911,  
470 which were comparable in magnitude to the documented 1974–1976 polynya, indicating that  
471 such openings are not isolated anomalies but represent recurring features of the Southern Ocean  
472 system. These events are characterised by proxy signatures consistent with modern polynya  
473 conditions, supporting the persistence of underlying formation mechanisms throughout the  
474 observational period. Conversely, the interval from 1920 to 1950, marked by a notable absence  
475 of polynya activity, coincides with documented changes in Southern Ocean circulation and  
476 atmospheric forcing. This prolonged quiescent phase provides valuable baseline information  
477 for assessing natural variability in MRP activity on multi-decadal timescales.

478 Our multiproxy approach significantly enhances the sensitivity of the reconstruction,  
479 particularly in detecting smaller-magnitude and shorter-duration events that may have been  
480 overlooked in previous studies. As a result, the reconstructed history reveals a more dynamic  
481 and variable pattern of MRP activity than previously recognised, with implications for  
482 understanding its role in ocean ventilation, carbon cycling, and regional climate variability.  
483 Our study reveals that synthesising paleoclimate records, modern observations, and modelling  
484 approaches can provide a robust foundation for advancing the understanding of the MRP as a  
485 critical and recurrent component of the Southern Ocean climate system.





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499 Maud Land coast, East Antarctica’ (MADICE).

## 500 **Author contributions**

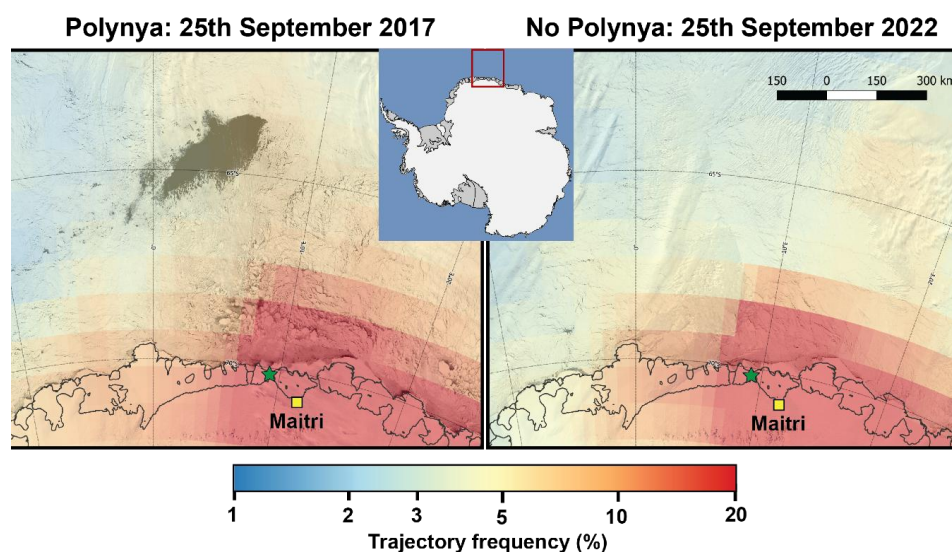
501 RD and MT defined the study objectives. RD led the ice core processing and analysis with  
502 support from AP, BLR and CML. RD led the data analysis and interpretation with inputs from  
503 KM and CML. RD prepared the manuscript with feedback from all co-authors. MT and KM  
504 were the project leaders

## 505 **Data Availability**

506 All ice core data and polynya indices are being prepared for submission to the NCPOR Polar  
507 Data Centre (<https://data.ncpor.res.in>). For final archival, open data formats will be used. A  
508 DOI will be generated and included in the final version of the paper. Data access can be granted  
509 to reviewers promptly upon request.

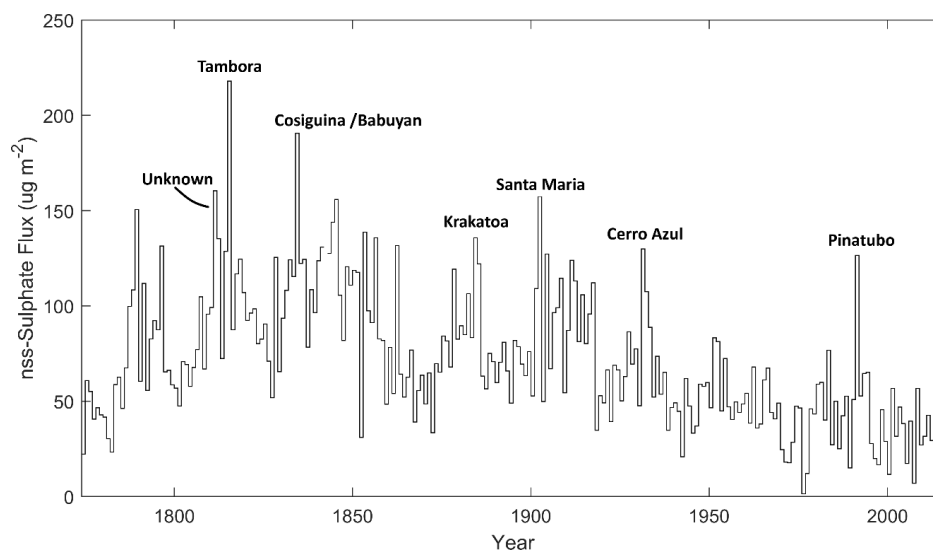
## 510 **Declaration of competing interest**

511 The authors declare that they have no known competing financial interests or personal  
512 relationships that could have appeared to influence the work reported in this paper.



513

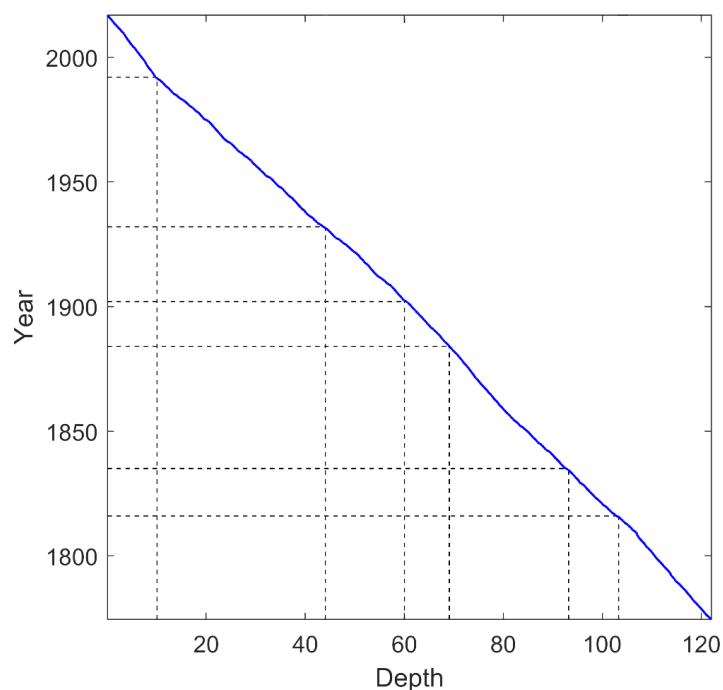
514 **Figure 1 | Study area and wind-trajectory patterns. (a)** Maud Rise Polynya during its largest  
 515 recent opening on 25 September 2017, detected with Moderate Resolution Imaging  
 516 Spectroradiometer (MODIS) on NASA's Terra satellite (from NASA Worldview,  
 517 <https://worldview.earthdata.nasa.gov/>, last access: 15 September 2024). **(b)** MODIS Satellite  
 518 imagery taken on the same day in a non-polynya year, 25 September 2022. Images are overlaid  
 519 with the endpoint frequency of back-trajectories ending at the Djupranen ice core site (green  
 520 star) from June to October of the years 2017 and 2022, respectively (see methods). The Indian  
 521 Maitri Station is marked by a yellow square. The colourmap is on a logarithmic scale. The inset  
 522 shows the location of the main map. Grounding line and calving front are marked (Matsuoka  
 523 et al., 2015).



524

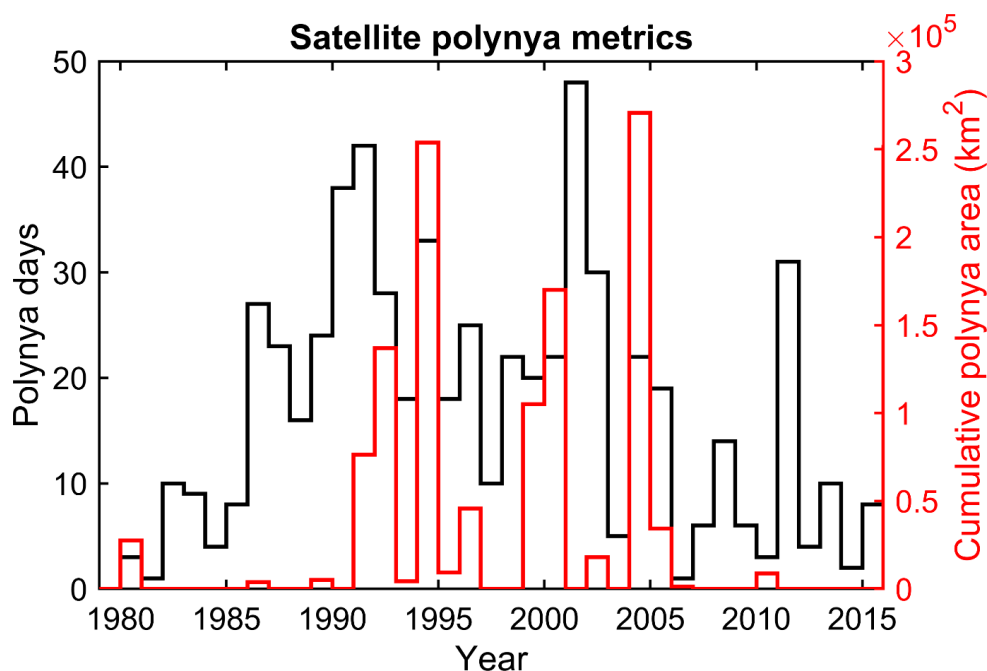
525 **Figure 2 | Tie points for ice core chronology.** Volcanic events are identified from the non-  
 526 sea-salt sulphate flux records. Only major volcanic events used as age tie points for  
 527 reconstructing the chronology are marked.

528



529

530 **Figure 3 | Age-depth scale for the Djupranen ice core.** Tie points used for chronology (Fig.  
 531 2) are shown with dashed lines. Major time markers are shown using the black dashed lines.  
 532 Minimal deviation from the intersection points of the vertical and horizontal lines indicates the  
 533 robustness of the chronology.

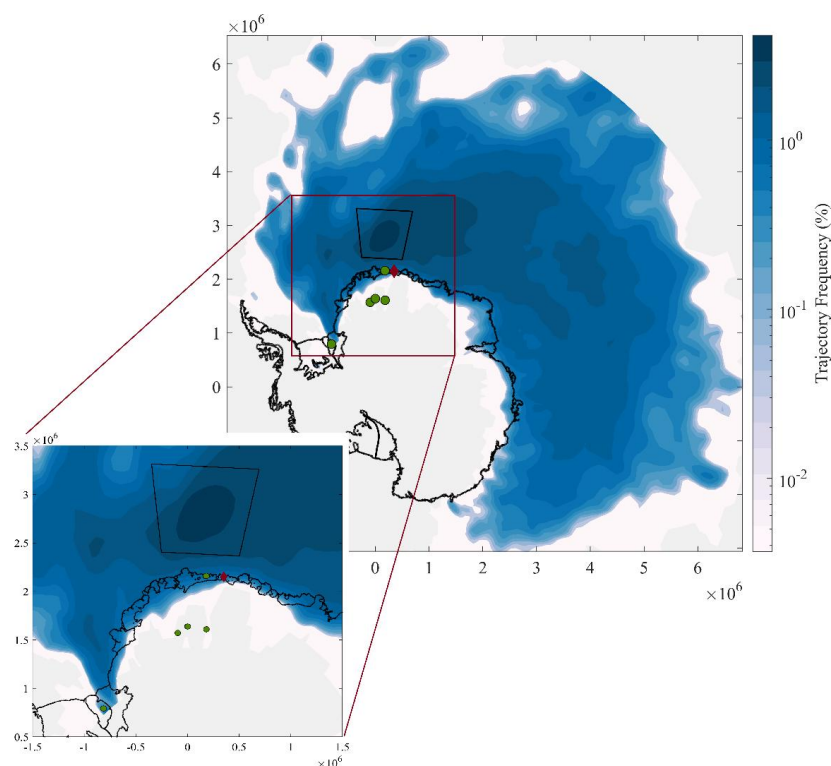


534

535 **Figure 4 | Satellite polynya metrics.** Two different annual metrics of satellite-derived polynya  
 536 occurrence. Since the polynya is highly dynamic temporally and spatially, the metrics do not  
 537 show a one-to-one resemblance.

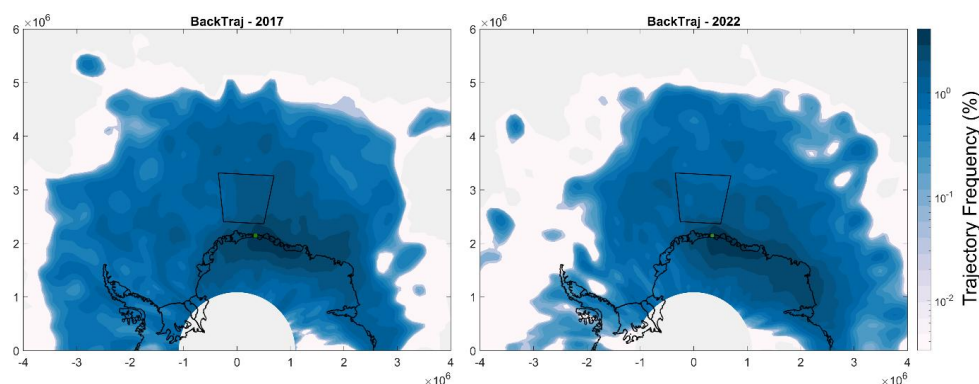


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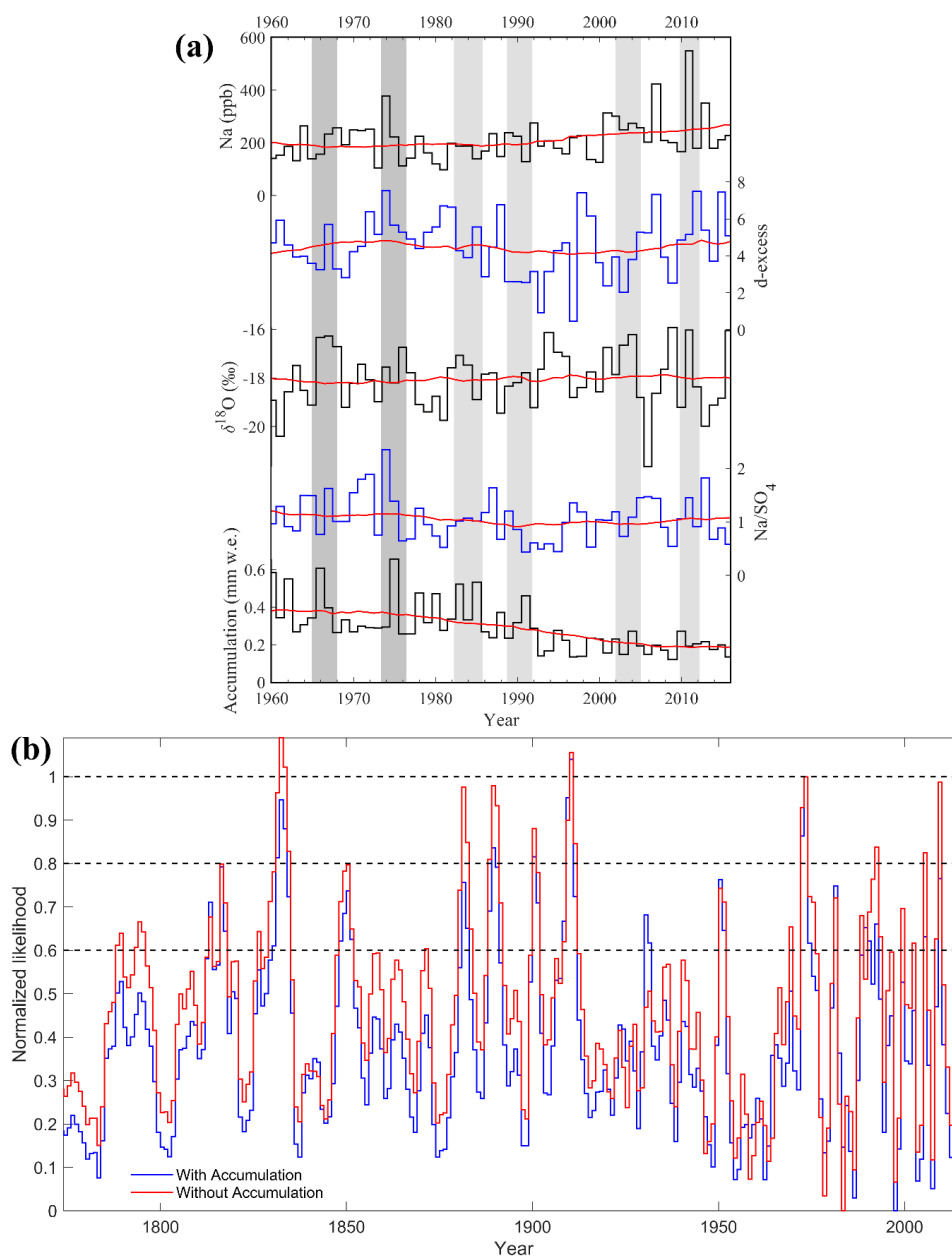
539

540 **Figure 5 | Airmass transport pathways originating from MRP:** Endpoint frequency of  
 541 forward trajectories originating over the polynya-prone region (Heuzé et al., 2021) of the  
 542 Weddell Sea, showing the prevalent transport pathway during the winter to early spring months  
 543 for the period 1948 – 2016. A significant proportion of the trajectories end over our ice core  
 544 site (red diamond). Location of ice cores used by Goosse et al., (2021) are shown with green  
 545 circles. The gray region shows area with no trajectory endpoints.



546

547 **Figure 6 | Air mass transport pathways ending at the ice core site:** Endpoint frequency of  
 548 back trajectories originating from the ice core site during a well-known polynya year (2017;  
 549 left) and a non-polynya year (2022; right) is shown. The endpoint frequency pattern during  
 550 both years is very similar, indicating a consistent transport pathway during most of the years  
 551 and suggesting that our ice core records activity over the polynya-prone region of the Weddell  
 552 Sea (black trapezium) in all years. The gray region shows areas with no trajectory endpoints.

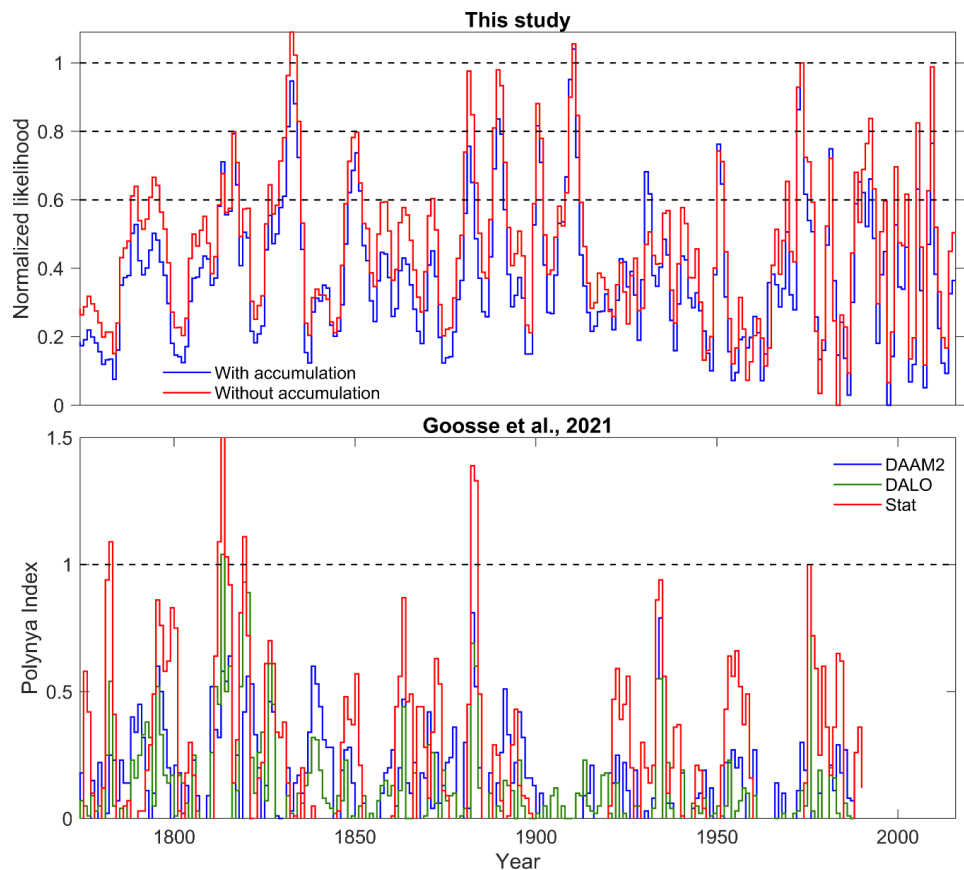


**Figure 7 | (a) Ice core proxies and polynya identification.** Annual variability of selected ice core proxy variability from 1960 – 2016. Known events from the satellite data are marked with dark gray patches for polynya years and light gray patches for halo years (Heuzé et al., 2021). **(b) Ice core derived polynya indices:** Ice core derived polynya indices in the past 250 years



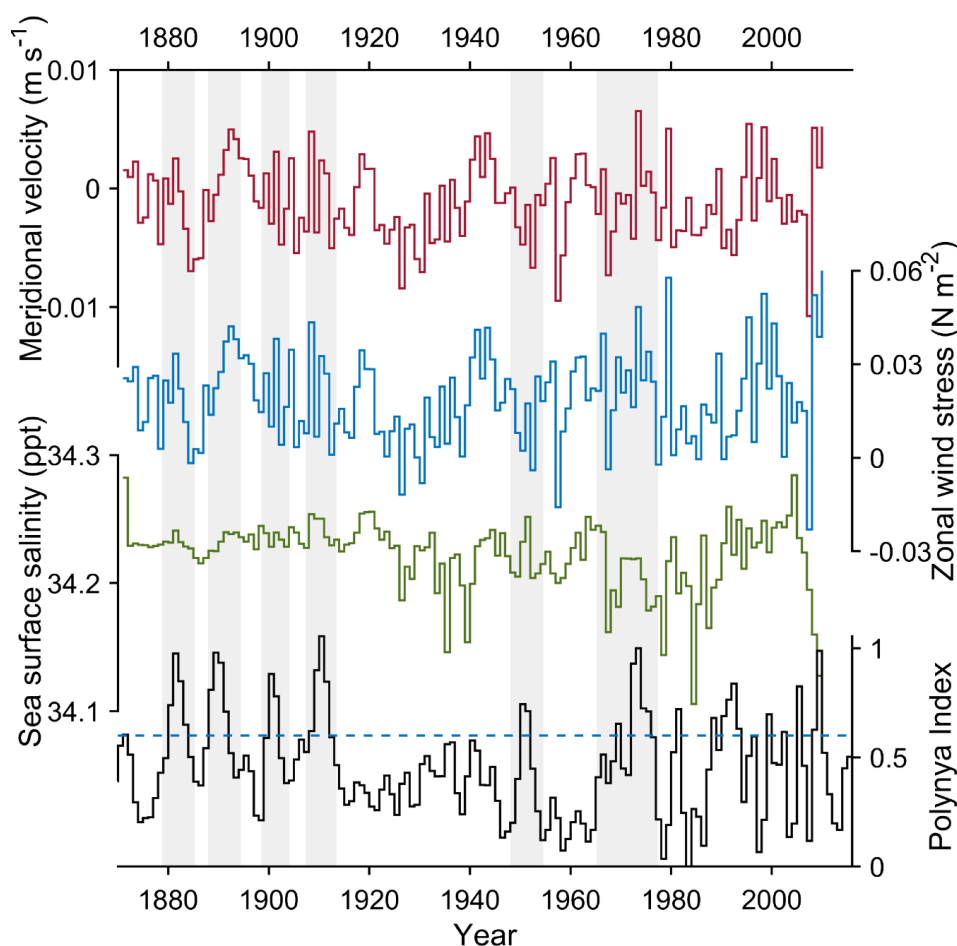


558 using the anomalies in the annual record of five ice-core proxies: Na, deuterium excess,  $\delta^{18}\text{O}$ ,  
559 snow accumulation and Na/SO<sub>4</sub> ratio. In order to test the dependence of the polynya index on  
560 snow accumulation, we calculate two indices, one using snow accumulation (blue curve) and  
561 another without (red curve). The two indices behave similarly over the entire time period;  
562 however, the index without snow accumulation shows a higher range of variability. Three  
563 different thresholds, 0.6, 0.8, and 1, are set for detecting polynyas.



564

565 **Figure 8 | Comparison of the polynya indices.** Polynya indices reconstructed in this study  
566 (upper panel) are shown with the polynya indices from Goosse et al. (2021) (lower panel). The  
567 polynya indices from Goosse et al. (2021) are based on six surface mass balance records using  
568 data assimilation with two control simulations performed with the SPEAR (Seamless system  
569 for Prediction and EArth system Research) global climate model, SPEAR\_AM2 (DAAM2;  
570 Blue) and SPEAR\_LO (DALO; green), and a simple average of the standardized time series  
571 (Stat, red).



572

573 **Figure 9 | Factors influencing polynya formations (a) Mean meridional velocity, (b) zonal**  
 574 **wind stress, and (c) sea surface salinity record over the polynya-prone region of the Weddell**  
 575 **Sea. (d) Polynya index reconstructed in this study. Grey bars are the major polynya events prior**  
 576 **to 1976.**



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