



1 Chloric Acid-Driven Nucleation Enhanced by
2 Dimethylamine and Sulfuric acid in the Arctic:
3 Mechanistic Study

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5 Shengming Wang¹, Huidi Zhang², Xiangli Shi², Qingzhu Zhang^{1,*},
6 Wenxing Wang¹, Qiao Wang¹

7
8 ¹Environment Research Institute, Shandong University,

9 Qingdao 266237, P. R. China

10 ²College of Geography and Environment, Shandong Normal
11 University, Jinan 250014, PR China

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22 *Corresponding authors. E-mail: zqz@sdu.edu.cn.



Abstract

Chlorine radicals are strong oxidizing agents in the atmosphere, and the process of chlorine oxidation results in the formation of chloric acid (HClO_3 , CA). Recent studies have shown that CA is prevalent in the Arctic boundary layer. However, the contribution of chlorine-containing species to oceanic new particle formation (NPF) has not been fully revealed. It is expected that CA is involved in the oceanic nucleation process. In this study, the enhancement of CA-based NPF by dimethylamine (DMA) and sulfuric acid (SA) was comparatively investigated at the molecular level using density-functional theory (DFT) and atmospheric cluster dynamics simulation (ACDC). The results show that DMA can form clusters with CA through hydrogen bonding, halogen bonding and proton transfer, which reduces the energy barrier for CA-based cluster formation and significantly improves the thermodynamic stability of CA clusters. The cluster formation rate of CA-DMA cluster system is higher than that of the CA-SA cluster system. The CA-DMA cluster system in the Arctic atmosphere contributes to NPF. These findings may help to reveal some of the missing sources of the Arctic NPF. The present study contributes to a deeper understanding of the influence of oceanic chlorine-containing constituents on the oceanic NPF.



44 **1 Introduction**

45

46 Marine aerosols as the main natural aerosol system have a major global impact
47 by regulating the radiative balance and climate of clouds(O' dowdg). New particle
48 formation (NPF) contributes to more than half of the global cloud condensation nuclei,
49 which in turn contributes to cloud formation (Gordon et al., 2017; Takegawa et al.,
50 2020; Williamson et al., 2019; Zhang et al., 2012). Compared with clouds over land,
51 ocean clouds cover a wider area and significantly increase the albedo of the ocean, so
52 ocean clouds contribute more to the climate system (Merikanto et al., 2009; Wood,
53 2012; Zheng et al., 2021). Sulfuric acid (SA, H₂SO₄), methane sulfonic acid (MSA,
54 CH₃HSO₃), and iodic acid (IA, HIO₃) are generally considered to contribute to the
55 formation of oceanic particles (Hodshire et al., 2019; Yin et al., 2021; Facchini et al.,
56 2008; Perraud et al., 2015; Arquero et al., 2017; Hopkins et al., 2008). However, there
57 is still a significant difference between the particle formation rates observed in the
58 field and those predicted by simulation (Kirkby et al., 2011; Kirkby et al., 2016;
59 Zhang et al., 2004; Ehn et al., 2014; Dawson et al., 2012). Therefore, it is necessary to
60 consider whether other gaseous precursors are involved in NPF to narrow the gap
61 between experiments and simulations.

62 Compared to the main atmospheric oxidants, hydroxyl radicals, chlorine radicals
63 act as strong oxidants in the polar troposphere at relatively high concentration levels
64 (Stone et al., 2012). The active chlorine cycle in the Arctic boundary layer during the
65 spring after polar sunrise depletes O₃ in the region (Custard et al., 2016; Foster et al.,



2001; Thompson et al., 2015). Chloric acid (CA) has no photoactivity and CA is ubiquitous in the spring in the Arctic, with concentrations estimated to range from 1×10^5 to 7×10^6 molecules cm^{-3} (Tham et al., 2023). CA was not found in the particle phase of the aerosol, thus it is difficult to determine whether CA is involved in the NPF phase.

Many studies have shown that atmospheric bases such as methylamine (MA), dimethylamine (DMA), trimethylamine (TMA) and ammonia can effectively enhance SA-based NPF (Yao et al., 2018; Almeida et al., 2013). Although amines emit 10–20 times less than ammonia in the ocean, amines can effectively form clusters with substances such as SA in the ocean (Myriokefalitakis et al., 2010; Semeniuk and Dastoor, 2018; Almeida et al., 2013). Of these amines, DMA has been found as a component of marine secondary aerosols (Facchini et al., 2008). Widely dispersed DMA has an atmospheric concentration of 0.4 – 10 pptv over the ocean and plays a key role in marine NPF (Van Pinxteren et al., 2019). DMA has been identified as the strongest enhancing atmospheric amine for SA and IA-driven NPF (Olenius et al., 2017; Ning et al., 2022). Thus, DMA may have a higher enhancing potential (EP) than NH_3 for CA-based NPF.

Sulfuric acid (H_2SO_4 , SA) has been detected in both gas and particulate phases in polluted coastal areas of China (Zhu et al., 2019; Yu et al., 2019). It is noteworthy that the concentration of SA is two orders of magnitude higher (up to 10^8 molecules cm^{-3}) in the coastal polluted areas due to urban air pollution compared to the clean marine atmosphere (Zhu et al., 2019). Considering its strong nucleation ability (Sipilä et al.,



2010; Faloona, 2009), it is likely that SA molecules nucleate in marine regions along with CA.

Using density functional theory (DFT) and the Atmospheric Clusters Dynamic Code (ACDC), the involvement of DMA and SA in the initial phase of CA-based NPF has been investigated. We obtained the minimum free energy structures of the $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters. The temperatures used in this study are within the temperature range of the atmospheric boundary layer in the ordinary range (Miřijovský and Langhammer, 2015). The concentration of CA was estimated to be in the range of $1.0 \times 10^6 - 1.0 \times 10^8$ molecules cm^{-3} based on measured data. Further study of CA-DMA clusters under Arctic atmospheric conditions and the corresponding thermodynamic data used as input to the ACDC reveals the growth pathways and formation rates of the clusters.

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101 **2 Computational methods**

102 **Configurational Sampling**

We employed a multi-step global minimum sampling scheme to search for the global minimum of $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters. In this study, the initial structures of 1000-10000 $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters were randomly generated using the ABCluster software to determine their global minima (clusters with the lowest Gibbs free energies). In the multistep sampling scheme, the geometry optimization is performed at the PM7, $\omega\text{B97X-D/6-31+G(d,p)}$ and $\omega\text{B97X-D/6-31++G(d,p)}$ levels of theory, and the single-point energy calculations are



performed at the DLPNO-CCSD(T)/aug-cc-pVTZ level of theory based on the ω B97X-D/6-31++G(d,p) theory level is performed on the geometry. The GAUSSIAN 09 program package(Frisch et al., 2016) was used to perform the PM7 and ω B97X-D calculations. DLPNO-CCSD(T) calculations were performed in the ORCA 4.0.0 program(Neese, 2012). For convergence problems and failures such as ending with a false frequency in the optimization of $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and $(\text{CA})_{1-4}(\text{SA})_{1-4}$ cluster geometries, the initial structures will be modified and re-optimized until the optimization is successful. The free energy of formation (ΔG) of individual clusters is calculated at different temperatures 238, 258, and 278 K. The ΔG of individual clusters is calculated at different temperatures. The structures of $(\text{SA})_{1-4}$ and $(\text{DMA})_{1-4}$ clusters were obtained from previous studies and are recalculated here(Xie et al., 2017).

Atmospheric Cluster Dynamics Code (ACDC) Simulation

Time-evolving cluster formation rates, steady-state concentrations, and growth paths for $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters were calculated using ACDC(Mcgrath et al., 2012). There is good agreement between the conclusions of the ACDC simulations and the experimental results obtained using the birth and death equations (Almeida et al., 2013; Lu et al., 2020; Kürten et al., 2018). In this study, ACDC simulations were performed to model the formation process of CA-DMA and CA-SA neutral clusters without considering the effects of charge and water. The $(\text{CA})_5(\text{DMA})_5$ clusters are set as boundary clusters (see Supporting Information (SI) for details). The concentration ranges of [CA], [SA] and [DMA] were set to $10^6 - 10^8$



132 cm^{-3} , 10^6 – 10^8 cm^{-3} and 0.1–100 ppt, respectively.

133

134 **3 Results and discussion**

135 3.1 Cluster Structures and Cluster Formation Free Energy

136 To evaluate the thermodynamic stability of the formed CA-DMA clusters, the
137 formation free energies (ΔG , kcal mol⁻¹) at 278 K are calculated for the
138 (CA)₁₋₄(DMA)₁₋₄ clusters at the
139 DLPNO-CCSD(T)/aug-cc-pVTZ// ω B97X-D/6-31++G(d,p) level of theory. As shown
140 in Figure. 1, hydrogen bonds play an important role in the formation of CA-DMA
141 clusters. Proton transfer reactions are not observed in the pure (CA)₂ and (CA)₃
142 clusters, whereas spontaneous proton transfer reactions are observed in all
143 (CA)₁₋₄(DMA)₁₋₄ clusters. For most (CA)₁₋₄(DMA)₁₋₄ clusters, protons are
144 transferred from CA to DMA.

145 In order to compare the thermodynamic stability of the formed CA-DMA and
146 CA-SA clusters, we calculated the formation Gibbs free energy values (ΔG , kcal
147 mol⁻¹) for the (CA)₁₋₄(DMA)₁₋₄ clusters at 278 K at the
148 DLPNO-CCSD(T)/aug-cc-pVTZ// ω B97X-D/6-31++G(d,p) level of theory. As shown
149 in Figure. 1, hydrogen bonding plays an important role in the formation of CA-DMA
150 clusters. Proton transferring wasn't observed in pure (CA)₂ and (CA)₃ clusters,
151 whereas spontaneous proton transfer reactions were present in all (CA)₁₋₄(DMA)₁₋₄
152 clusters. For most (CA)₁₋₄(DMA)₁₋₄ clusters, protons are transferred from CA to
153 DMA, forming Cl-O...H-N hydrogen bonding, accompanied by the production of



154 ClO_3^- negative ions and DMA^+ ions. In CA-DMA clusters containing two or more
155 chlorine atoms, including $(\text{CA})_2(\text{DMA})_1$, $(\text{CA})_3(\text{DMA})_1$, $(\text{CA})_3(\text{DMA})_1$,
156 $(\text{CA})_3(\text{DMA})_1$, $(\text{CA})_3(\text{DMA})_4$, $(\text{CA})_4(\text{DMA})_1$, $(\text{CA})_4(\text{DMA})_2$, $(\text{CA})_4(\text{DMA})_3$, and
157 $(\text{CA})_4(\text{DMA})_4$ clusters, O-O...O-Cl halogen bonds and Cl-O...H-N hydrogen bonds
158 together stabilize these clusters described above. Halogen bonds are not present in
159 clusters containing single CA atom and in $(\text{CA})_2(\text{DMA})_2$, $(\text{CA})_2(\text{DMA})_3$,
160 $(\text{CA})_2(\text{DMA})_4$, as well as $(\text{CA})_3(\text{DMA})_3$ clusters. For the CA-SA clusters, the ClO_3^-
161 negative ions generated in the CA-DMA cluster system were not found in the CA-SA
162 cluster system because proton transfer do not occur. In contrast to the O-Cl...O-Cl
163 halogen bond found in the CA-DMA cluster system, the CA-SA cluster adds
164 S-O...Cl-O halogen bond (in the Figure S2).

165 The formation Gibbs free energy values of (A) $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and (B)
166 $(\text{CA})_{1-4}(\text{SA})_{1-4}$ at 278 K and 1 atm are shown in Figure. 2. The Gibbs free energy
167 values for the formation of $(\text{CA})_2$, $(\text{CA})_3$, and $(\text{CA})_4$ clusters in the pure acid system
168 are 1.31, 2.52, and 5.37 kcal mol⁻¹, respectively, implying that at 278 K, pure CA
169 clusters are thermodynamically difficult to form. The ΔG values of $(\text{CA})_1(\text{SA})_4$ and
170 pure SA clusters are lower than the corresponding ΔG values of the corresponding
171 CA-DMA system. The difference between the ΔG values of $(\text{CA})_4(\text{DMA})_4$ cluster and
172 $(\text{CA})_4(\text{SA})_4$ cluster is the largest, up to 61.62 kcal mol⁻¹. $(\text{CA})_1(\text{SA})_1$ and
173 $(\text{CA})_1(\text{DMA})_1$ are both very important in their respective cluster systems, and their
174 ΔG values are -2.62 and 10.08 kcal mol⁻¹, respectively⁻¹. The ΔG values of $(\text{CA})_{1-4}$
175 clusters are 10.08 – 28.16 kcal mol⁻¹ higher than those of the corresponding



(CA)₁₋₄(DMA)₁ clusters, suggesting that pure CA clusters may grow by collision with DMA. As the size of CA-DMA clusters increases, the clusters gradually form a cage-symmetric structure. The ΔG values of the majority of clusters in the CA-DMA system are 3.55 – 61.62 kcal mol⁻¹ lower than the corresponding ΔG values of the CA-SA system. This indicates that the CA-DMA cluster system is more thermally stable compared to the CA-SA cluster system.

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3.1 Evaporation Rates and Cluster Stability

The total evaporation rate ($\sum \gamma$, s⁻¹) of (CA)₁₋₄(DMA)₁₋₄ clusters and (CA)₁₋₄(SA)₁₋₄ clusters formed at T = 278 K is shown in Figure. 3. The smaller value of $\sum \gamma$ means that the stability of CA-DMA clusters is higher and the clusters shrink further. The smaller value of $\sum \gamma$ implies the higher stability of CA-DMA clusters and further contraction of the clusters. The clusters with the same number of CA molecules and the number of DMA molecules include (CA)₁(DMA)₁, (CA)₂(DMA)₂, (CA)₃(DMA)₃, and (CA)₄(DMA)₄ clusters, which have the values of $\sum \gamma$ of 10², 2 × 10⁻⁴, 5 × 10⁻¹, and 3 × 10¹ s⁻¹, respectively. (CA)₂(DMA)₂ cluster has the lowest $\sum \gamma$ value, implying that (CA)₂(DMA)₂ cluster is the most stable cluster in the “4 × 4” box system of CA-DMA clusters. For clusters with different numbers of CA and DMA molecules, the $\sum \gamma$ value of (CA)₂(DMA)₁ cluster is significantly lower than that of other clusters, which indicates that (CA)₂(DMA)₁ cluster has a high probability of competing for the growth path of nucleation. In this study, we compare the evaporation rates of clusters from the CA-DMA system with those from the CA-SA



198 system at 278K. In contrast to the other clusters, the evaporation rate of $(\text{CA})_1(\text{SA})_4$
199 clusters is lower than that of the corresponding $(\text{CA})_1(\text{DMA})_4$ clusters. The
200 evaporation rates of most $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ clusters are much smaller than those of
201 the corresponding $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters, indicating that the CA-DMA cluster
202 system is kinetically more stable than the CA-SA cluster system.

203

204 3.2 Cluster Formation Rates and Steady-State Cluster Concentrations

205 Cluster formation rate (J) and steady-state CA dimer concentration ($\sum[(\text{CA})_2]$)
206 are important indicators for assessing the enhancement potential of DMA for
207 CA-based nucleation. Figure 4 shows the variation of $\sum[(\text{CA})_2]$ and J values at 278 K
208 with CA concentration ($[\text{CA}] = 10^6 - 10^8 \text{ cm}^{-3}$) and DMA concentration ($[\text{DMA}] = 1,$
209 10 and 100 ppt). The J value of the CA-DMA system showed a positive correlation
210 with CA and DMA concentrations as CA and DMA concentrations increased. The
211 dependence of the cluster formation rate on the DMA concentration does not decrease
212 with increasing CA concentration, which means that the dependence of the system on
213 DMA does not saturate when the CA concentration is high. The $\sum[(\text{CA})_2]$ and J of the
214 CA-DMA system with the full range of acid-base concentrations considered (CA: 10^6
215 $- 10^8 \text{ molecules cm}^{-3}$, DMA: $1, 10$ and 100 ppt) were significantly higher than the
216 CA-SA system (Figure. S4). However, at high concentrations of $[\text{CA}] = 1 \times 10^7$
217 molecules cm^{-3} and $[\text{DMA}] = 10 \text{ ppt}$, the J value of the CA-DMA cluster system only
218 reaches $6.70 \times 10^{-7} \text{ cm}^{-3} \text{ s}^{-1}$. The contribution of the CA-DMA cluster system to the
219 NPF is not significant under the atmospheric conditions of 278 K. In addition, the



220 CA-PA cluster system has a lower J value (Figure. S15). This may be due to the weak
221 bond energies of Cl-O...Cl-O halogen bonds in the process of CA-PA nucleation.

222 To further systematically explore the effect of temperature on the J of the
223 CA-DMA cluster system, Figure. 5 shows the simulated J at other temperatures (238
224 and 258 K), $[CA] = 10^6 - 10^8$ molecules cm^{-3} , $[DMA] = 0.1$ ppt (red line), 1 ppt
225 (green line), 10 ppt (blue line), and 100 ppt (black line). A comparison of the
226 simulations at 258 K (Figure. 5a) and 238 K (Figure. 5b) reveals that the decrease in
227 temperature further increases the J value of the CA-DMA cluster system to a higher
228 level. However, at a low temperature (258 K), the J values of the CA-DMA cluster
229 system do not reach higher levels at high concentrations of $[CA] = 1 \times 10^7$ molecules
230 cm^{-3} and $[DMA] = 10$ ppt. The J values of the CA-DMA cluster system were further
231 investigated under cold Arctic conditions (238 K). It was found that the J value of
232 CA-DMA cluster system at 238 K atmospheric condition was significantly higher
233 than that of CA-DMA cluster system at other higher temperature conditions, which
234 was mainly due to the fact that the low temperature attenuates the evaporation of
235 CA-DMA clusters.

236

237 3.3 Cluster growth pathway

238 Figure 6 shows the growth paths of CA-DMA and CA-SA clusters at 278 K, $[CA]$
239 $= 10^6 \text{ cm}^{-3}$, $[DMA] = 1$ ppt, and $[SA] = 10^6 \text{ cm}^{-3}$. The first step in the growth of
240 CA-DMA clusters is the collision of a CA molecule and a DMA molecule to form a
241 $(CA)_1(DMA)_1$ cluster. There are two growth paths for $(CA)_1(DMA)_1$ clusters: a CA



242 molecule collides and combines with a $(\text{CA})_1(\text{DMA})_1$ cluster to form a $(\text{CA})_2(\text{DMA})_1$
243 cluster, and a DMA molecule is subsequently added to form a $(\text{CA})_2(\text{DMA})_2$ cluster; a
244 $(\text{CA})_1(\text{DMA})_1$ cluster combines with another $(\text{CA})_1(\text{DMA})_1$ cluster to form a
245 $(\text{CA})_2(\text{DMA})_2$ cluster. This growth pattern is mainly due to the high stability of
246 $(\text{CA})_2(\text{DMA})_1$ clusters. After the $(\text{CA})_2(\text{DMA})_2$ clusters, the growth route of the
247 CA-DMA cluster system extends along the direct binding to the $(\text{CA})_1(\text{DMA})_1$
248 clusters. For $(\text{CA})_3(\text{DMA})_3$ clusters, one route is the addition of acid and base, and the
249 other route is the direct binding to $(\text{CA})_1(\text{DMA})_1$ clusters to produce $(\text{CA})_4(\text{DMA})_4$
250 clusters.

251 The growth pathway of the CA-SA cluster system differs considerably from that
252 of the CA-DMA cluster system. The cluster initially formed in the CA-SA cluster
253 system is the $(\text{CA})_2$ cluster not the $(\text{CA})_1(\text{SA/DMA})_1$ cluster in the CA-DMA system.
254 After the generation of $(\text{CA})_2$ cluster, the system can generate $(\text{CA})_3(\text{SA})_1$ clusters by
255 successive addition of acid molecules. Subsequently, $(\text{CA})_3(\text{SA})_1$ clusters combine
256 with $(\text{CA})_1(\text{SA})_1$ clusters to generate $(\text{CA})_3(\text{SA})_2$ clusters. The final $(\text{CA})_4(\text{SA})_4$
257 cluster of the system is generated by collision of a $(\text{CA})_3(\text{SA})_2$ cluster with an $(\text{SA})_2$
258 cluster.

259

260 3.4 Atmospheric implications and conclusions

261 In this paper, combination method of quantum chemistry and ACDC were used
262 to elucidate the molecular structure mechanism of DMA and SA enhancing role of
263 CA nucleation by comparing the CA-DMA and CA-SA nucleation systems. Proton



264 transfer was observed in all $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ clusters while no proton transfer
265 occurred within $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters. The ClO_3^- groups generated by the
266 deprotonation of CA are involved in the formation of at least one hydrogen bond.
267 Hydrogen and halogen bonds together stabilize the CA-DMA and CA-SA nucleation
268 systems. The vast majority of CA-SA cluster systems have higher ΔG values than the
269 corresponding CA-DMA cluster systems. The cluster formation rates of the pure
270 CA-PA and CA-SA nucleation systems are relatively low, and the contribution of
271 DMA to CA nucleation is stronger than that of SA. The CA-DMA nucleation system
272 contributes to the NPF at low Arctic temperatures. Clusters with the same number of
273 CA and DMA molecules ($(\text{CA})_1(\text{DMA})_1$, $(\text{CA})_2(\text{DMA})_2$, $(\text{CA})_3(\text{DMA})_3$, and
274 $(\text{CA})_4(\text{DMA})_4$ clusters) play a key role in the growth path of CA-DMA clusters. This
275 study is important for a deeper understanding of CA Arctic atmospheric nucleation.

276 The study clarifies the role of CA in the marine NPF and reveals the mechanism
277 of DMA acting as a key enhancer to CA-based NPF through intermolecular
278 interactions. The results suggest that CA-DMA synergistic nucleation in the Arctic
279 atmosphere is an under-recognized source of NPF, which provides a theoretical basis
280 for improving regional climate models (e.g., cloud condensation nucleation prediction)
281 and assessing the Arctic aerosol-climate feedback effect. The current simulations do
282 not take into account charge effects, the involvement of water molecules, and the
283 influence of complex atmospheric matrices (e.g., organic matter). In the future, it is
284 necessary to validate the simulation results with field observations and extend it to
285 multi-component (e.g., IA/SA/DMA mixing) nucleation systems in order to quantify



286 the contribution of chlorine-containing substances to the global NPF in a more
287 comprehensive way.

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289 **Data availability**

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291 All data supported the paper are available from the article, supplementary information,
292 and the corresponding author upon reasonable request.

293

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295

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301 **Additional information**

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303 **Competing interests:** The authors declare no competing financial interests.

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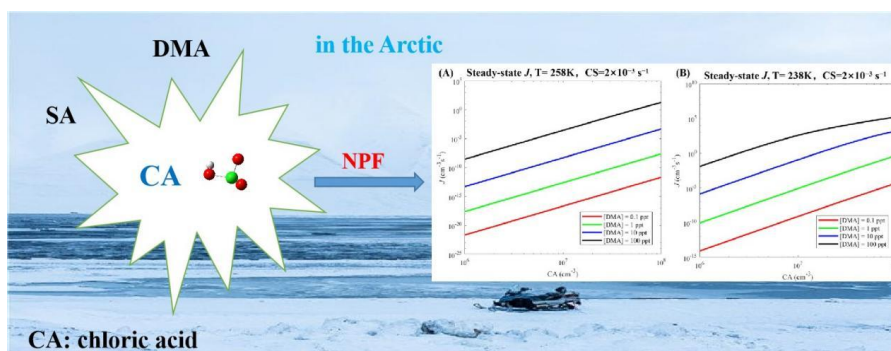
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Graphical Abstract





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Figure Captions

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468 **Figure 1.** Identified lowest free energy structures of the $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ clusters at
469 the DLPNO-CCSD(T)/aug-cc-pVTZ// $\omega\text{B97X-D/6-31++G(d,p)}$ level of theory. The
470 red, blue, gray, green and white balls represent oxygen, nitrogen, carbon, chlorine and
471 hydrogen atoms, respectively. The dashed white and black lines indicate hydrogen
472 and halogen bonds, respectively.

473 **Figure 2.** The formation free energy (ΔG) (in kcal mol^{-1}) of (A) $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and
474 (B) $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters at the
475 DLPNO-CCSD(T)/aug-cc-pVTZ// $\omega\text{B97X-D/6-31++G(d,p)}$ level of theory. The
476 calculations are performed at 278 K and 1 atm.

477 **Figure 3.** Evaporation rates for (A) $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and (B) $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters

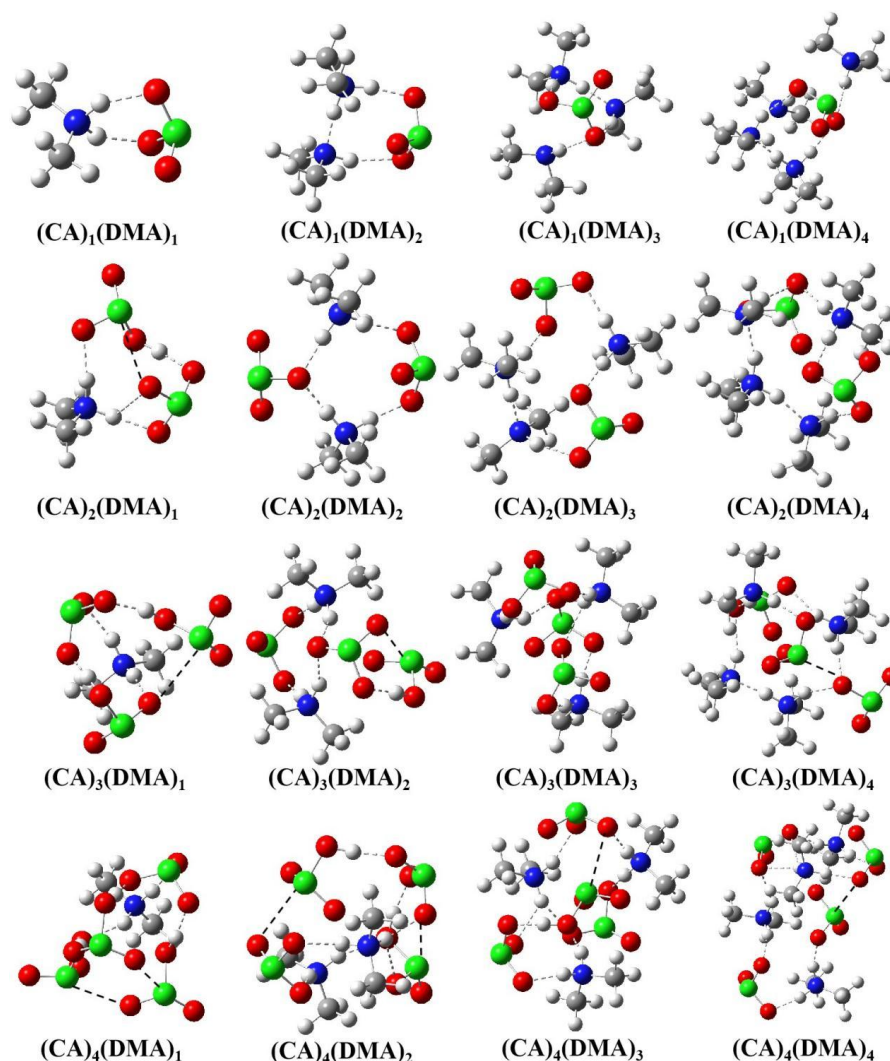


478 clusters at 278 K and 1 am.

479 **Figure 4.** Simulated steady-state CA dimer concentration $\Sigma[(CA)_2]$ (cm^{-3}) (A) and
480 the cluster formation rates J ($\text{cm}^{-3} \text{s}^{-1}$) out of the simulation systems (B) as a function
481 of $[CA]$ at 278 K.

482 **Figure 5.** The simulated cluster formation rate J ($\text{cm}^{-3} \text{s}^{-1}$) of the CA-DMA system at
483 different temperatures (A) 258, and (B) 238 K; $[CA] = 10^6 - 10^8 \text{ molec. cm}^{-3}$; $[DMA]$
484 $= 0.1, 1, 10, \text{ and } 100 \text{ ppt}$; and $CS = 2 \times 10^{-3} \text{ s}^{-1}$.

485 **Figure 6.** (A) Main clustering pathways of $(CA)_{1-4}(DMA)_{1-4}$ clusters at 278 K, $[CA]$
486 $= 10^6 \text{ cm}^{-3}$, and $[DMA] = 1 \text{ ppt}$. (B) Main clustering pathways of $(CA)_{1-4}(SA)_{1-4}$
487 clusters at 278 K, $[CA] = 10^6 \text{ cm}^{-3}$, and $[SA] = 10^6 \text{ cm}^{-3}$.



488

489 **Figure 1.** Identified lowest free energy structures of the $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ clusters at
 490 the DLPNO-CCSD(T)/aug-cc-pVTZ// $\omega\text{B97X-D/6-31++G(d,p)}$ level of theory. The
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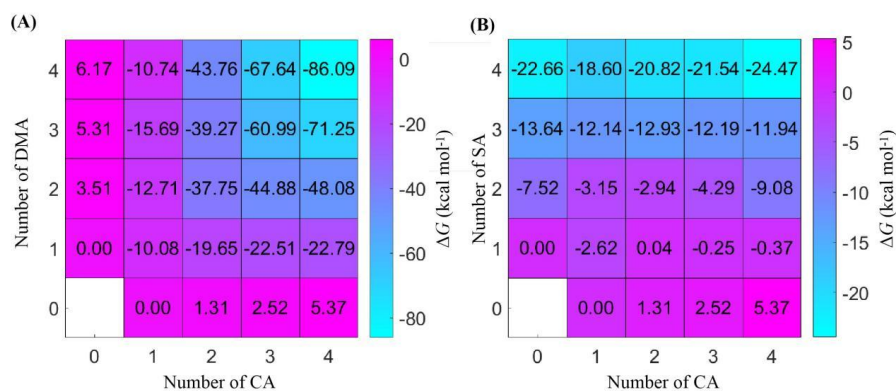


Figure 2. The formation free energy (ΔG) (in kcal mol^{-1}) of (A) $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ and (B) $(\text{CA})_{1-4}(\text{SA})_{1-4}$ clusters at the DLPNO-CCSD(T)/aug-cc-pVTZ// $\omega\text{B97X-D/6-31++G(d,p)}$ level of theory. The calculations are performed at 278 K and 1 atm.

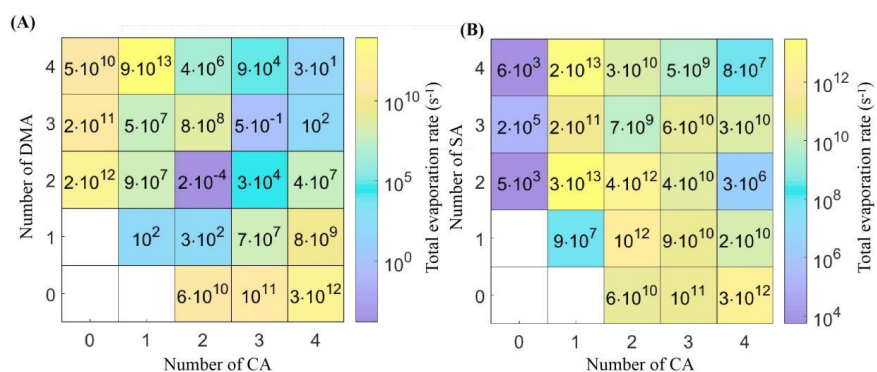


Figure 3. Evaporation rates for (A) (CA)₁₋₄(DMA)₁₋₄ and (B) (CA)₁₋₄(SA)₁₋₄ clusters at 278 K and 1 am.

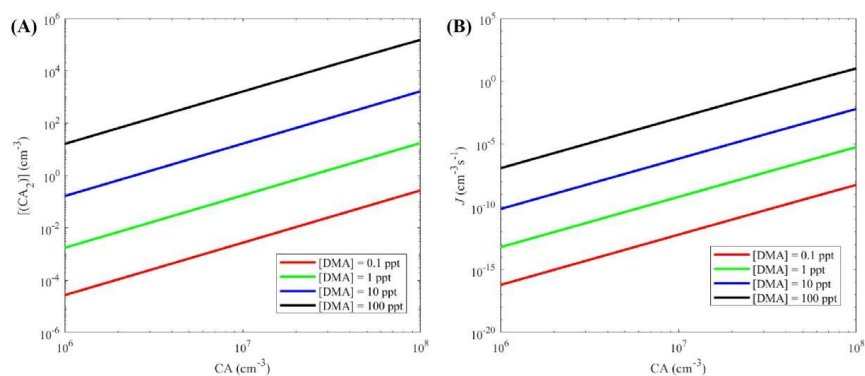


Figure 4. Simulated steady-state CA dimer concentration $\sum[(\text{CA})_2] (\text{cm}^{-3})$ (A) and the cluster formation rates $J (\text{cm}^{-3} \text{s}^{-1})$ out of the simulation systems (B) as a function of $[\text{CA}]$ at 278 K.

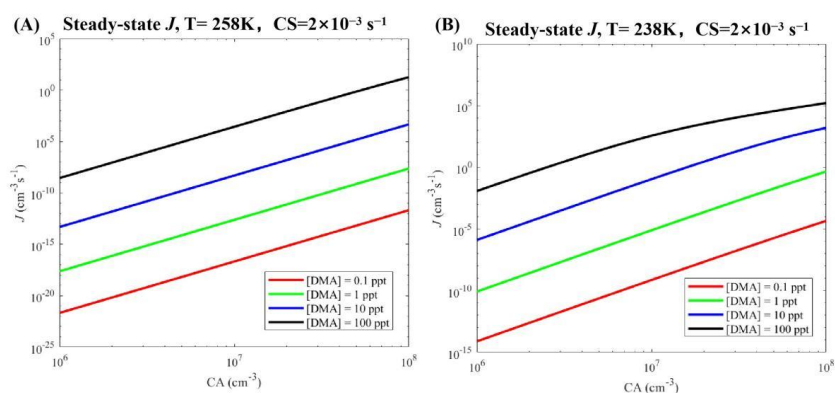
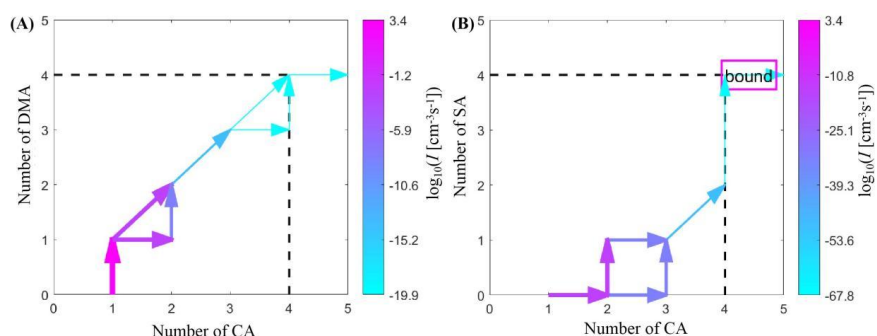


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545 **Figure 6.** (A) Main clustering pathways of $(\text{CA})_{1-4}(\text{DMA})_{1-4}$ clusters at 278 K, $[\text{CA}]$

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547 clusters at 278 K, $[\text{CA}] = 10^6 \text{ cm}^{-3}$, and $[\text{SA}] = 10^6 \text{ cm}^{-3}$.

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