



Investigating KDP signatures inside and below the dendritic growth layer with W-band Doppler Radar and in situ snowfall camera

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Abstract. Polarimetric radars provide variables like the specific differential phase (K_{DP}) to detect fingerprints of dendritic growth in the dendritic growth layer (DGL) and secondary ice production, both critical for precipitation formation. A key challenge in interpreting radar observations is the lack of in situ validation of particle properties within the radar measurement volume. While high K_{DP} in snow is usually associated with high particle number concentrations, only few studies attributed K_{DP} to certain hydrometeor types and sizes. To address this, we combined surface in situ observations from the Video In Situ Snowfall Sensor (VISSS) with remote sensing data from a polarimetric W-band radar and an X-band radar, along with modeling approaches. Data was collected during the CORSIPP project, part of the ARM SAIL campaign (winter 2022/2023, Colorado Rocky Mountains). We found that at W-band, high K_{DP} magnitudes can result from a broad range of particle number concentrations, between 1 and 1001^{-1} . Blowing snow and increased ice collisional fragmentation in a turbulent layer enhanced observed K_{DP} values. T-matrix simulations indicated that high K_{DP} values were primarily produced by particles smaller than 0.8 mm in the DGL and 1.2 mm near the surface. Discrete dipole approximation simulations based on VISSS data suggested that dendritic aggregates larger than 2.5 mm contributed 10-20 % to the measured W-band K_{DP} near the surface. These findings highlight the complexity of interpreting W-band K_{DP} in snowfall and emphasize the need for combined in situ observations and radar forward simulations to better understand snowfall microphysical processes.

1 Introduction

More than 60 % of global precipitation is generated through the ice phase (Heymsfield et al., 2020), therefore, research into the microphysics of the ice and mixed phase has intensified in recent years. The dendritic growth layer (DGL), a temperature region roughly between -10 and -20 °C, plays a major role in the formation of precipitation through the ice phase. In the DGL, the difference between the saturation vapor pressure over ice and water is maximized (Wegener-Bergeron-Findeisen process (Wegener, 1911; Bergeron, 1935)), which leads to a rapid growth of ice into very anisotropic plate-like shapes (Takahashi,



2014) while the number concentration is increased through secondary ice production (SIP) mechanisms (e.g. Von Terzi et al., 2022). A distinct secondary particle mode linked to the formation of new anisotropic particles may be observed at the top of the DGL as a bimodality in Doppler spectra from vertical pointing measurements (e.g. Moisseev et al., 2015). The origin of this secondary mode is still debated, theories about its formation were summarized in Von Terzi et al. (2022). Potential sources include sedimentation of ice particles, primary nucleation due to upward air motion, and some form of SIP mechanism.

Secondary ice production (SIP) has been identified as a key mechanism that enhances ice particle concentrations at temperatures above the homogeneous freezing temperature (Korolev and Leisner, 2020). SIP involves mechanisms that increase the number of ice particles without directly relying on primary ice nucleation through ice nucleating particles (INPs) (e.g., Hoose and Möhler, 2012; Kanji et al., 2017) or homogeneous freezing at very low temperatures (e.g., Sanz et al., 2013). As mountainous regions around the globe act as natural water reservoirs, storing water in the form of snow and ice, a special focus is put on microphysical processes leading to snow formation in mountain regions. Wind shear and associated turbulence were found to impact precipitation production and ice microphysics in multiple studies in recent years (e.g. Grazioli et al., 2015; Ramelli et al., 2021; Dedekind et al., 2023). Grazioli et al. (2015) proposed that splinters from SIP processes can be recirculated in the layer of shear and grow into small oblate ice particles. The area above a shear layer is supplied with supercooled liquid water (SLW) and ice crystals produced inside the turbulent updrafts. This can favor the growth of anisotropic ice crystals in this region and is suspected to cause enhanced Z_{DR} signatures (Hogan et al., 2002; Grazioli et al., 2015). Ramelli et al. (2021) investigated a turbulent layer in an inner-Alpine valley and found evidence of riming, ice collisional fragmentation and aggregation inside this layer. During another study in the Alps, Dedekind et al. (2023) recently found that SIP is best predicted by wind shear. Chellini and Kneifel (2024) found larger ice aggregates and higher concentrations of ice particles in turbulent layers of Arctic clouds, likely due to increased fragmentation. All SIP processes cause an increase in ice number concentration, and depending on the temperature range, the newly formed ice splinters can serve as embryos for the growth of new anisotropic ice particles (e.g. Grazioli et al., 2015; Luke et al., 2021).

Polarimetric radars can help to detect the fingerprints of dendritic growth in the DGL and SIP processes by transmitting and receiving polarized electromagnetic waves. The specific differential phase (K_{DP}) is a polarimetric radar variable utilized across various frequency bands that is intrinsically dependent on the size and number concentration of hydrometeors. Because K_{DP} is very sensitive to high concentrations of anisotropic particles, it is a useful tool for detecting signatures of SIP in slanted radar measurements. High-frequency radars operating at smaller wavelengths, such as Ka- or W-band, are particularly effective for detecting small ice particles produced by SIP. These shorter wavelengths yield larger K_{DP} magnitudes and reduced noise compared to centimeter-wavelength radars (Bringi et al., 2001). Moreover, K_{DP} signals are easier to interpret with respect to other polarimetric variables as they are less affected by wavelength-dependent scattering effects, such as non-Rayleigh effects (Lu et al., 2015). K_{DP} can be complemented with other polarimetric variables, such as integrated differential reflectivity (Z_{DR}), which is independent of number concentration but also sensitive to particle density and aspect ratio (Kumjian, 2013). Due to the presence of high number concentrations of anisotropic particles, areas of enhanced K_{DP} and Z_{DR} are often found as distinct layers in the DGL. These maxima are often not collocated, a discrepancy that remains under debate (e.g., Andrić et al., 2013; Moisseev et al., 2015; Oue et al., 2018; Von Terzi et al., 2022, among others).



To explain the behavior of K_{DP} and Z_{DR} in the DGL, the different types of particles present in the upper DGL and also how they influence Z_{DR} and K_{DP} need to be considered. Schrom and Kumjian (2016) distinguished particles into two classes. The first class are isometric (I-type) particles, containing a broad variety of aggregates and ice crystals with irregular or nearly spherical shape which can be generated throughout the whole cloud depth. The second class are dendritic (D-type) particles, containing all pristine dendrites, plates and needles, generally having higher density and more anisotropic shape than I-type particles. D-type particles grow only in defined temperature regimes: from -20°C to -10°C for dendrites and hexagonal plates and between -3°C and -8°C for needles (Schrom and Kumjian, 2016; Griffin et al., 2018). Griffin et al. (2018) followed this particle classification of Schrom and Kumjian (2016) and developed a theory for the origin of Z_{DR} and K_{DP} signatures inside the DGL. Griffin et al. (2018) state that I-type particles can produce moderate Z_{DR} and quite significant K_{DP} if their concentration is sufficiently high, while D-type particles lead to extremely large Z_{DR} and tangible K_{DP} , if their concentration is high enough. They suggested that high Z_{DR} signatures are produced by particle populations dominated by D-type particles because of their very low (or high) aspect ratio and their high density, whereas high K_{DP} is linked to the overwhelming presence of I-type particles. Griffin et al. (2018) also made clear that I and D-type particles will mostly coexist, and that polarimetric radar variables depend on the relative contributions of the I-type and D-type ice particles in the mixture. We will revisit their hypothesis later on in our study. SIP in the DGL, although not mentioned by Griffin et al. (2018), adds an additional possibility of how a population of I-type particles with low Z_{DR} and K_{DP} signatures transitions into a mix of I and D-types with tangible Z_{DR} and K_{DP} to the pathways suggested by Griffin et al. (2018).

The primary challenge in interpreting radar variables lies in the lack of in situ validation of the properties of the particle population within the radar measurement volume. Various approaches have sought to address this, including the use of radar data in conjunction with airborne in situ measurements, as e.g. in the HALO-AC3 (Wendisch et al., 2024) and IMPACTS (Dunnavan et al., 2022) campaigns. Alternatively, surface in situ precipitation measurements have been used to validate polarimetric radar observations, as demonstrated in studies employing scanning C-band radars and a NASA Particle Video Imager (Moisseev et al., 2015), or collocated polarimetric radars and disdrometers during field campaigns such as GLACE (Grazioli et al., 2015) and TRIPEX (Dias Neto et al., 2019).

The unique dataset presented in our study features continuous measurements from the slanted polarimetric W-band radar (LIMRAD94, type RPG-FMCW-94-DP (Küchler et al., 2017)), situated in close proximity to the Video In Situ Snowfall Sensor (VISSS; (Maahn et al., 2024)). It enables us to directly link polarimetric radar variables to hydrometeor population properties observed by VISSS at the ground and investigate the origin of K_{DP} signatures at W-band. The data was collected within the Characterization of orography-influenced riming and secondary ice production and their effects on precipitation rates using radar polarimetry and Doppler spectra (CORSIPP) project, which is part of the DFG Priority Programm SPP-PROM (Trömel et al., 2021). CORSIPP was embedded in the Atmospheric Radiation Measurement (ARM) Surface Atmosphere Integrated Field Laboratory (SAIL) campaign (Feldman et al., 2023) during the very snow-rich winter season 2022/2023 in the Colorado Rocky Mountains. The data of LIMRAD94 is complemented by other collocated radars like the dual polarization Colorado State University X-band radar (CSU X-band) (Heflin et al., 2024), and the single polarization Ka-band ARM Zenith Radar

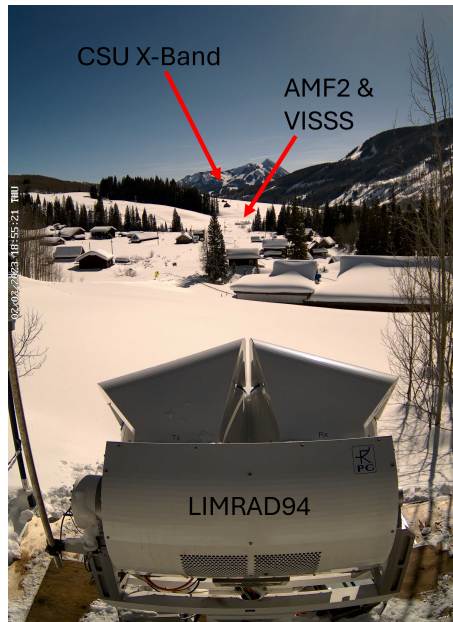


Figure 1. Field of view of LIMRAD94 during CORSIPP pointed at 151° azimuth. The red arrows mark the locations of AMF2, VISSS and CSU X-band.

90 (KAZR, (Widener et al., 2012)). The measurement site was strongly influenced by orographic turbulence, hence we also aim to elucidate the influence of turbulence on polarimetric variables and investigate possible SIP fingerprints.

In Sect. 2, we describe the used data with a special focus on polarimetric radar variables in 2.1.1. The modeling approaches are described in Sect. 2.4. The fall streak tracking method and eddy dissipation rate retrieval are described in Sect. 2.5 and Sect. 2.6, respectively. In Sect. 3.1, we provide statistics on VISSS-derived number concentration and the ratio of the third to the second measured PSD moments (D_{32}) along with collocated LIMRAD94 K_{DP} measurements. In a case study of four fall streaks, we analyze K_{DP} , integrated and spectral Z_{DR} , reflectivity (Z_e), spectral width (SW), and correlation coefficient (ρ_{hv}) to identify the particles and microphysical processes contributing to the observed K_{DP} magnitudes (Sect. 3.2). Combining T-matrix modeling and CSU X-band data, we try to constrain the particle size responsible for the observed K_{DP} magnitudes (Sect. 3.3). We also aim to quantify the contribution of aggregates larger than 2.5 mm to the observed W-band K_{DP} in Sect.

100 3.4. The main findings are summarized in Sect. 4.

2 Data and Methods

In the following sections, the instruments used during this study and their variables are described. We also provide an overview over the used modeling approaches as well as the retrievals for the fall streaks and the eddy dissipation rate.



2.1 Radars

During CORSIPP, LIMRAD94 was mounted on a novel cold-temperature scanning unit, also manufactured by RPG. LIMRAD94 was deployed at an altitude of 2905 m at the Rocky Mountain Biological Laboratory (RMBL) about 525 m away from the second ARM Mobile Facility (AMF2) in Gothic (CO) near the RMBL Ore House between November 15, 2022 and June 5, 2023. LIMRAD94 observations were collected at constant elevation of 40° and an azimuth of 151° (towards the VISSS, AMF2 and the CSU X-band radar). The temporal resolution of LIMRAD94 data is about 3 s. Further technical details on the deployment of LIMRAD94 and VISSS during CORSIPP can be found in Kalesse-Los et al. (2023). The technical details of the radar chirp program employed during CORSIPP are listed in Tab. A1. For SAIL, the ARM KAZR (Widener et al., 2012) and a scanning dual polarization X-band radar (CSU X-band) from Colorado State University (CSU) were deployed in a distance of 525 m (at 151° azimuth and 2889 m ASL) and 8018 m (at 149° azimuth and 3147 m ASL), respectively from LIMRAD94 (also see Appendix, Table B1 and Fig. B1). From the CSU X-band radar data, RHI scans at 329° azimuth (towards the AMF2 and LIMRAD94) and plane position indicator (PPI) scans at 4° elevation were used in this study. In the appendix, Sect. B we also briefly discuss the volume matching of the CSU X-band radar and LIMRAD94 data.

2.1.1 Polarimetric radar data

LIMRAD94 is capable of measuring spectrally resolved polarimetric variables like differential phase shift and differential reflectivity, among others. Polarimetric calibration of LIMRAD94 was performed at the beginning of the field campaign during a time of vertical pointing measurements as described in Myagkov et al. (2016a).

The differential phase shift (Φ_{DP} , °) consists of a backscatter and a propagational part as given by:

$$\Phi_{DP}(r) = \delta(r) + 2 \int_0^r K_{DP}(s) ds \quad (1)$$

where δ (°) is the backscatter differential phase that occurs due to non-spherical hydrometeors large enough relative to the radar wavelength such that the scattering is in the non-Rayleigh regime (Trömel et al., 2013). The specific differential phase shift (K_{DP} , ° km⁻¹) is the range derivative of the propagational component of Φ_{DP} . Hence, to accurately calculate K_{DP} , δ needs to be estimated and subtracted from Φ_{DP} first. The estimation of δ in rain was thoroughly discussed in e.g., Trömel et al. (2013) and Myagkov et al. (2020). However, for most frozen hydrometeors, δ was found to be negligible in the S, C, and X-bands (Balakrishnan and Zrnic, 1990; Ryzhkov et al., 2011; Trömel et al., 2013). In spectrally resolved Φ_{DP} data, δ can be detected by a sudden increase or decrease of Φ_{DP} in parts of the Doppler spectrum. Consequently, also integrated Φ_{DP} often reveals characteristic "bumps" if δ is present (Trömel et al., 2013). During the case study presented, we checked integrated and spectral Φ_{DP} for the above mentioned signs of δ that may impair the interpretation of K_{DP} signatures. However, we found δ during our case study to be negligible. For LIMRAD94 and CSU X-band, the K_{DP} was calculated from Φ_{DP} using a convolution with low-noise Lanczos differentiators with window length of 23 gates. This method provides results that are analogue to a moving window linear regression and is implemented in the python package wradlib (Heistermann et al., 2013).



135 K_{DP} depends on the shape, orientation, concentration, density and the size of ice particles. High concentrations of dense particles with more anisotropic shape and uniform orientation cause high K_{DP} values. In rain, K_{DP} is proportional to the 3rd power of the particle size ($K_{DP} \sim D^3$). In low density snow, K_{DP} is proportional to the 1st power of the particle size ($K_{DP} \sim D^1$) due to density of snow being inversely depended on the diameter ($\rho \sim D^{-1}$). As the degree of riming increases, the proportionality of K_{DP} shifts towards a higher power of the particle size, reflecting changes in the snow density. However, 140 the influence of riming on K_{DP} is contradictory: while riming increases the density of particles and hence K_{DP} , riming also makes particles more spherical reducing K_{DP} (Ryzhkov and Zrnica, 2019). K_{DP} is also influenced by turbulence through its dependency on the width of the canting angle distribution (σ) via the orientation factor (F_0), which is a function of σ (Ryzhkov et al., 2002; Ryzhkov and Zrnica, 2019). The canting angle is defined as the angle between the projection of a symmetry axis of a spheroidal particle and the vertical axis onto the polarization plane (Ryzhkov et al., 2002). If the canting angles of particles 145 in a population differ a lot, the width of the canting angle distribution increases. The stronger the turbulence, the larger σ , the smaller F_0 and hence the smaller K_{DP} .

The differential reflectivity (Z_{DR} , dB) is commonly defined as:

$$Z_{DR} = 10 \cdot \log_{10} \left(\frac{Z_{eH}}{Z_{eV}} \right) \quad (2)$$

150 where (Z_{eH} , $\text{mm}^6 \text{m}^{-3}$) is the radar reflectivity measured at horizontal polarization and (Z_{eV} , $\text{mm}^6 \text{m}^{-3}$) is the radar reflectivity measured at vertical polarization. At W-band, the largest contribution to positive Z_{DR} values (in the ice phase) is usually made by the most dense anisotropic ice, small enough to be in the Rayleigh scattering regime, usually located on the slow edge of the Doppler spectrum. Contrary to K_{DP} , Z_{DR} is immune to the total concentration factor, so the combined analysis K_{DP} and Z_{DR} can be helpful to decouple effects of shape and concentration. Larger, strongly rimed, or aggregated hydrometeors, 155 typically located on the fast side of the Doppler spectrum, are more spherical and exhibit low Z_{DR} . Since Z_{DR} is reflectivity-weighted, its magnitude is often reduced by the influence of these larger, spherical particles (e.g. Oue et al., 2018). To use Z_{DR} without suffering from this masking effect, we also calculate the maximum spectral Z_{DR} value ($sZ_{DR_{max}}$) for each Doppler spectrum. As demonstrated in Von Terzi et al. (2022), this parameter is efficient for detection of small non-spherical particles masked by presence of large, more spherical particles like aggregates. Similar to K_{DP} , Z_{DR} is depend on σ (Hubbert and Bringi, 2003; Ryzhkov and Zrnica, 2019) and hence also decreased by turbulence. $sZ_{DR_{max}}$ is furthermore decreased by 160 turbulence through the broadening of the Doppler spectrum. Imagine that only one type of particle in a radar volume, moving with a certain terminal fall velocity, produces Z_{DR} larger than zero. $sZ_{DR_{max}}$ is high because the positive Z_{DR} is measured at one certain Doppler velocity bin. If the fall velocity is now determined by turbulent air motion rather than the particles inherent terminal fall velocity, the particles producing the positive Z_{DR} are "distributed" over a larger range of Doppler velocities and 165 the Z_{DR} at one certain Doppler velocity bin decreases. It is therefore essential to consider turbulence before interpreting any K_{DP} or Z_{DR} signatures.



LIMRAD94 was measuring at 40° elevation, which will cause Z_{DR} and also K_{DP} to be smaller than at lower elevation angles (e.g., Myagkov et al., 2016b). This should be taken into account when comparing the bulk values with other studies using Z_{DR} and K_{DP} but at different elevation angles.

170 2.2 VISSS

The first generation VISSS was installed next to the AMF2 facility, in the line of sight of LIMRAD94 with a horizontal distance to LIMRAD94 of approximately 550 m (Fig. 1). VISSS1 has a pixel resolution of $58.832 \mu\text{m px}^{-1}$, a frame rate of 140 Hz, and an observation volume of $(wxdxh) = 75.2 \times 75.2 \times 60.1 \text{ mm}^3$ (Maahn et al., 2024). This yields an observational volume of 0.000339 m^3 . VISSS level2match variables (see Fig. 2 in Maahn et al. (2024)) used in this study include particle
 175 size distributions (PSD), D_{32} , complexity and total number concentration (N_{tot}). D_{32} is the ratio of the third to the second measured PSD moment. Assuming that particle mass is proportional to the particle maximum dimension squared, D_{32} is a proxy for the mass-weighted mean diameter of the particle population (Maahn et al., 2015). Complexity is derived from the ratio of the particle perimeter to the perimeter of a circle with the same area (Maahn et al., 2024). The total number concentration reaches values of over 1001^{-1} in this case study, which may seem unusual at first when comparing to other
 180 literature. VISSS is however more sensitive to small particles than laser disdrometers as shown by Maahn et al. (2024). This is likely the reason for the observed particle concentrations. All diameters used refer to the maximum diameter (D_{max}), unless otherwise specified.

2.3 Additional products

During our study, we used additional data acquired at the AMF2 such as 2 m temperature, pressure, relative humidity, and
 185 wind speed as well as ARM 3-channel microwave radiometer liquid water path (LWP) and integrated water vapor (IWV); a detailed list can be found in the code and data availability section. Vertical temperature profiles on site are provided by the ARM interpolated sonde and gridded sonde VAP (Fairless et al., 2021). The sounding data, collected on site through two radiosonde launches per day (11 and 23 UTC) is transformed into continuous daily files with 1-minute time resolution and combined with ARM 3-channel microwave radiometer temperature data.

190 2.4 Modeling of radar observables

To interpret radar observations and constrain quantitative estimates of the microphysical properties of ice particles, scattering simulations are employed.

2.4.1 T-matrix

The first approach approximates the shape of a particle as a spheroid and simulates its backscattering and forward scatter-
 195 ing properties using the T-matrix scattering model (Mishchenko, 2000), hereafter referred to as the T-matrix approach. For a given wavelength, this method requires only three parameters: the equivolumetric diameter, the axis ratio, and the refractive



index. The refractive index, which depends on the fraction of ice and air (Oguchi, 1983), can be adjusted to model ice particles with varying ice densities (Myagkov et al., 2016b). For simplicity, we assume the horizontal alignment of ice particles, albeit acknowledge that a broader distribution of canting angles generally leads to less pronounced polarimetric signatures. The T-matrix approach is computationally efficient; however, it is most applicable when the particles are smaller than the radar wavelength. Large ice particles, which tend to have highly irregular shapes, are not well-represented by the spheroidal approximation, particularly for short radar wavelengths. This limitation makes the T-matrix approach less suitable for modeling the scattering of rimed particles and aggregates at millimeter wavelengths (Leinonen et al., 2012; Kneifel et al., 2015).

2.4.2 DDA

For simulating the scattering of large ice particles at millimeter wavelengths, the Discrete Dipole Approximation (DDA) is often employed. This method allows for an accurate approximation of the scattering properties of individual particles with arbitrary shapes. However, DDA is computationally intensive and requires detailed specifications of the particle geometry which is generally unknown. Here, we randomly selected 105 unrimed dendritic aggregates from the snowScatt database (Ori et al., 2021), spanning maximum dimensions of 100 μm to 4 cm. As the falling behavior of snowflakes is determined by the complex interactions between aerodynamical and gravitational forces and is further complicated by e.g., turbulence, we assume that the snowflakes fall with their maximum extension oriented horizontally. The DDA for the dendritic aggregates was calculated using *adda* (Yurkin and Hoekstra, 2011) at 94 GHz and 40° elevation, and averaged over 16 azimuth orientations.

2.5 Fall streak tracking

Hydrometeor populations are influenced by horizontal and vertical winds as they precipitate to the surface and are hence exposed to advection. To estimate their microphysical evolution and to link ground observations with radar measurements aloft, these hydrometeor populations need to be observed along slanted fall streaks in vertical radar measurements. The analysis of the fall streaks is mostly achieved by reconstructing the advection using the present wind measurements (Marshall, 1953; Kalesse et al., 2016; Pfizenmaier et al., 2017). As no sufficiently high resolution wind data is available for our case study, we decided to analyze fall streaks with the simple approach of following maximum Z_e values backwards from surface to cloud top. Specifically, the Z_e at the start point is compared to the median Z_e of the next three higher range gates. The time difference between two neighboring range gates is not allowed to exceed $\frac{2\Delta h}{v_t}$ where Δh is the height difference in m between the range gates and v_t is the mean Doppler velocity (MDV) measured by KAZR. The maximum Z_e is chosen as the next point along the fall streak, and the process is repeated. This iterative process determines the suspected course of the fall streak in time and height. We acknowledge that horizontal homogeneity has to be assumed, especially for slanted measurements. Because the main wind direction in our case study is westerly, but the radar is pointing south-east it is obvious that we will never observe the exact same particles along their way to the ground.



2.6 Turbulence eddy dissipation rate retrieval (*EDR*)

Turbulence plays a critical role in understanding the evolution of hydrometeor populations and the behavior of polarimetric radar variables. The turbulence eddy dissipation rate (*EDR*) is used as a proxy for turbulence. *EDR* is the rate at which
230 turbulence kinetic energy dissipates in the atmosphere (Foken and Mauder, 2024). High *EDR* values indicate a more turbulent environment. We calculated *EDR* from the KAZR MDV for five minute time intervals, using the approach described in (Vogl et al., 2024). To mask turbulence, we applied an empirical threshold of $0.001 \text{ m}^2\text{s}^{-3}$ similar to the one found by Vogl et al. (2022). Note that the radar beams of KAZR and LIMRAD94 during CORSIPP were not aligned, so the measurement volume from which *EDR* is retrieved has a horizontal displacement from the LIMRAD94 measurement volume. This distance
235 increases with increasing height.

3 Results and Discussion

First, we show a statistical comparisons of VISSS N_{tot} and LIMRAD94 K_{DP} (Sect. 3.1). A case study with four K_{DP} fall streaks is shown in (Sect. 3.2). For this case study, we present a combination of T-matrix K_{DP} modeling and collocated CSU X-band observations in (Sect. 3.3) and K_{DP} simulations using DDA based on VISSS data (Sect. 3.4).

240 3.1 K_{DP} vs VISSS statistics

One advantage of the CORSIPP dataset is the close proximity of in situ observations by VISSS and detailed polarimetric W-band measurements of LIMRAD94. To understand how K_{DP} is related to properties of the PSD, we compare K_{DP} measurements of LIMRAD94 to VISSS surface measurements of N_{tot} and D_{32} for the period December 1, 2022 to February 28, 2023. LIMRAD94 K_{DP} was spatially averaged (median) between 100 and 500 m above ground (52 range gates) to reduce
245 effects of the measurement error, for the horizontal displacement of the two measurement volumes and for inhomogeneities in the observed precipitation. LIMRAD94 K_{DP} was also temporally averaged (median) to match the VISSS time resolution of one minute. To compensate for the vertical displacement of the measurement volumes of VISSS and LIMRAD94 and the resulting time lag, LIMRAD94 data was shifted backwards in time as precipitation would be first visible in LIMRAD94 measurements and later in the VISSS. Given that the center of the averaged height is 300 m and assuming a terminal fall velocity
250 of 1 m s^{-1} , we chose five minutes as value for the time shift. We also tested time shifts of 3 minutes and 10 minutes; none of them significantly changed the overall data distribution (not shown). Negative K_{DP} values and range gates with $\text{SNR} < 10 \text{ dB}$ were excluded. LIMRAD94 data was only used if all range gates between 100 and 500m contained data.

Overall, a clear connection between higher K_{DP} and N_{tot} is evident. Particle populations with D_{32} below 2 mm (blueish dots in Fig. 2) reach higher N_{tot} up to 100 l^{-1} more frequently. Within these particle populations, higher K_{DP} values of up to
255 2° km^{-1} are mostly linked to N_{tot} greater than 10 l^{-1} . In contrast, particle populations with D_{32} of 4 mm or higher (red dots in Fig. 2) mostly occur when N_{tot} is below 10 l^{-1} , but can also result in K_{DP} values larger than 2° km^{-1} . Particle populations with low N_{tot} and high D_{32} mainly form through aggregation, which decreases the overall N_{tot} by reducing the amount of

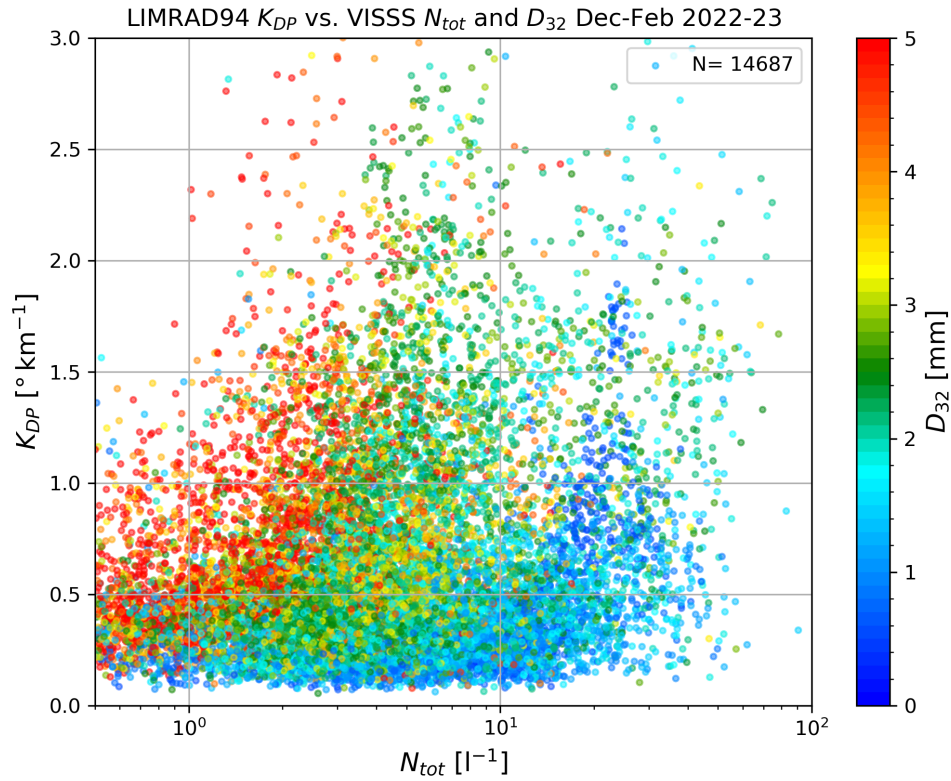


Figure 2. Scatterplot of LIMRAD94 K_{DP} vs. VISSS N_{tot} for DJF 2022/23. The x-scale is logarithmic. LIMRAD94 K_{DP} was spatially averaged between 100 and 500 m above ground and temporally averaged to fit the VISSS time resolution of one minute. Colors show D_{32} of the respective PSD. A total of 14773 data points (14687 minutes) were used for the plot.

smaller particles while increasing the mean particle diameter. There are three possible explanations for the K_{DP} produced by these particle populations:

1. The larger aggregates are the main contributors to the observed W-band K_{DP} .
2. The small ice particles are the main contributors to the observed W-band K_{DP} while their number concentration is anti-correlated with the presence of large aggregates.
3. Large aggregates and small ice particles contribute to the observed K_{DP} , their relative contributions depends on the particle size distribution and the properties of the particles.

The possible contribution of larger aggregates to increased W-band K_{DP} was recently discussed in Von Terzi et al. (2022). Since K_{DP} does not suffer from the non-Rayleigh effects as Z_e and Z_{DR} , it can keep increasing with particle size as long as

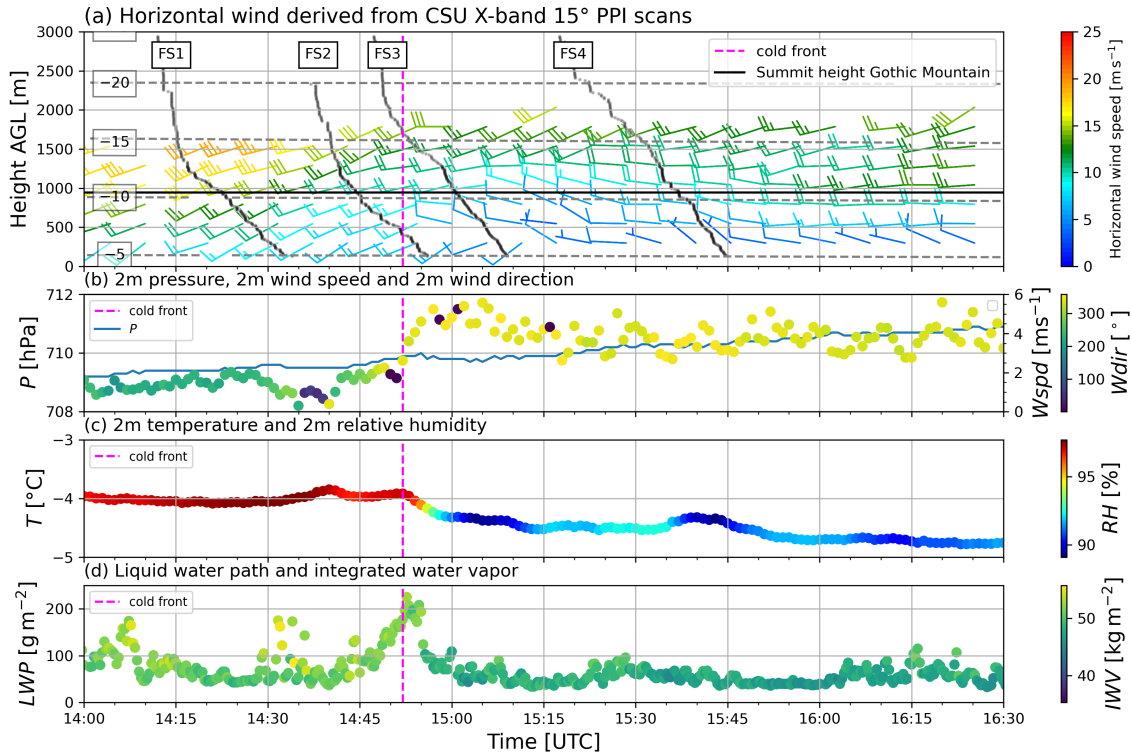


Figure 3. Meteorological observations during the case study on December 6, 2022. The vertical dashed magenta line in all plots indicates the passing of a cold front. a): Horizontal wind speed and direction. Horizontal grey dashed lines represent air temperature in °C from the interpolated radiosonde product. The analyzed fall streaks are displayed as black lines with their respective labels. The black solid horizontal line denotes the height of Gothic Mountain. b) AMF2-based meteorological variables at 2 m: pressure, wind speed and direction as colorcode. c): AMF2-based 2 m temperature and relative humidity. d): AMF2-based MWR LWP and IWV.

the shape is somewhat preserved (Lu et al., 2015; Von Terzi et al., 2022). Especially early-stage aggregates can be reasonably assumed to be still highly non-spherical and contribute substantially to the ensemble K_{DP} signal (Moisseev et al., 2015).

3.2 Snowfall case study: December 6, 2022

270 To investigate the formation of K_{DP} signatures and the influence of different microphysical processes and particle sizes on K_{DP} , we chose the case study on December 6, 2022. Over a span of 1.5 h, we identified and analyzed four fall streaks (FS1–4) characterized by enhanced K_{DP} . Despite exhibiting similar K_{DP} magnitudes, VISSS revealed varying particle population properties within these fall streaks.



3.2.1 Synoptic discussion

December 6, 2022 featured a complex synoptic situation with a stationary front over the Colorado Rocky Mountains and a surface low developing along this front over eastern central Colorado, slowly moving southward during the day. The measurement site was influenced by the development of this surface low. An overview of surface meteorological variables and the vertical wind profile during the case study is given in Fig. 3. Just before 15 UTC we observed a sudden change in 2 m wind direction from south-west to north winds, an increase in 2 m wind speed from under 2 m s^{-1} to over 5 m s^{-1} (Fig. 3b) and a decrease in 2 m temperature and humidity (Fig. 3c). After 15 UTC, we saw a steady increase in surface pressure (Fig. 3b). The evolution of the surface meteorological variables in Gothic points to a cold front passage just before 15 UTC (see magenta lines in Fig. 3). The wind throughout the lower troposphere up to 3000 m above LIMRAD94 was from the south-west before the cold front with a slight veer and an increase in speed toward higher altitudes, indicating the advection of warm air (Fig. 3a). With the passage of the cold front, the wind turned to westerly and later north-westerly directions up to 1500 m above LIMRAD94. Aloft, the winds backed to south-westerly direction again, indicating cold air advection. LWP and IWV decreased with the passage of the cold front (Fig. 3d), indicating that a dryer air mass being advected. The nature of the precipitation changed during the cold front passage; from low N_{tot} and high D_{32} precipitation dominated by aggregates before the cold front to a high N_{tot} and low D_{32} precipitation with very mixed particle types (see Fig. 4e). Therefore, we especially focus on this time period between 14:30 and 16:30 UTC in our case study.

3.2.2 The turbulence layer

Radar variables observed by LIMRAD94 as well as VISSS N_{tot} and D_{32} throughout the case study are shown in Fig. 4. Due to the orography surrounding the measurement site, we observed a distinct layer of enhanced turbulence, mostly at and below 1000 m above ground level (AGL). This is visualized by the EDR -masked data (white where EDR exceeds $0.001 \text{ m}^2 \text{ s}^{-3}$) in Fig. 4c).

The main wind direction above 1000 m AGL was south-west before the cold front and west to north-west after the cold front passage (Fig. 3a). The inner valley flow was weak south-westerly before the cold front and moderate northerly after the cold front (Fig. 3b). Hence, the wind shear was enhanced along the interface between the blocked valley and the cross-barrier flow aloft, which would lead to shear-induced turbulence as explained in Ramelli et al. (2021). The main blocking feature west of Gothic is Gothic Mountain (3850 m). The height of the summit of Gothic Mountain above the location of LIMRAD94 (roughly 1000 m AGL) corresponds well to the top of the observed turbulence layer.

Fig. 5 shows the MDV of CSU X-band from a PPI (a) and a RHI (b) scan at around 15:48 UTC and 15:46 UTC, respectively. During that time, EDR indicates enhanced turbulence between 500 and 1200 m AGL (Fig. 4c). The MDV measured during the PPI scan can be approximated as the radial component of the horizontal wind. Negative values (blue) in Fig. 5 correspond to motion towards the radar, positive values (yellow-red) to motion away from the radar. The PPI is partly blocked by mountains, however the westerly wind is apparent by blue colors west of CSU X-band and orange-red colors to the east. Where MDV is close to zero, the wind is perpendicular to the radar beam. Downstream of Gothic Mountain, there is a small corridor where

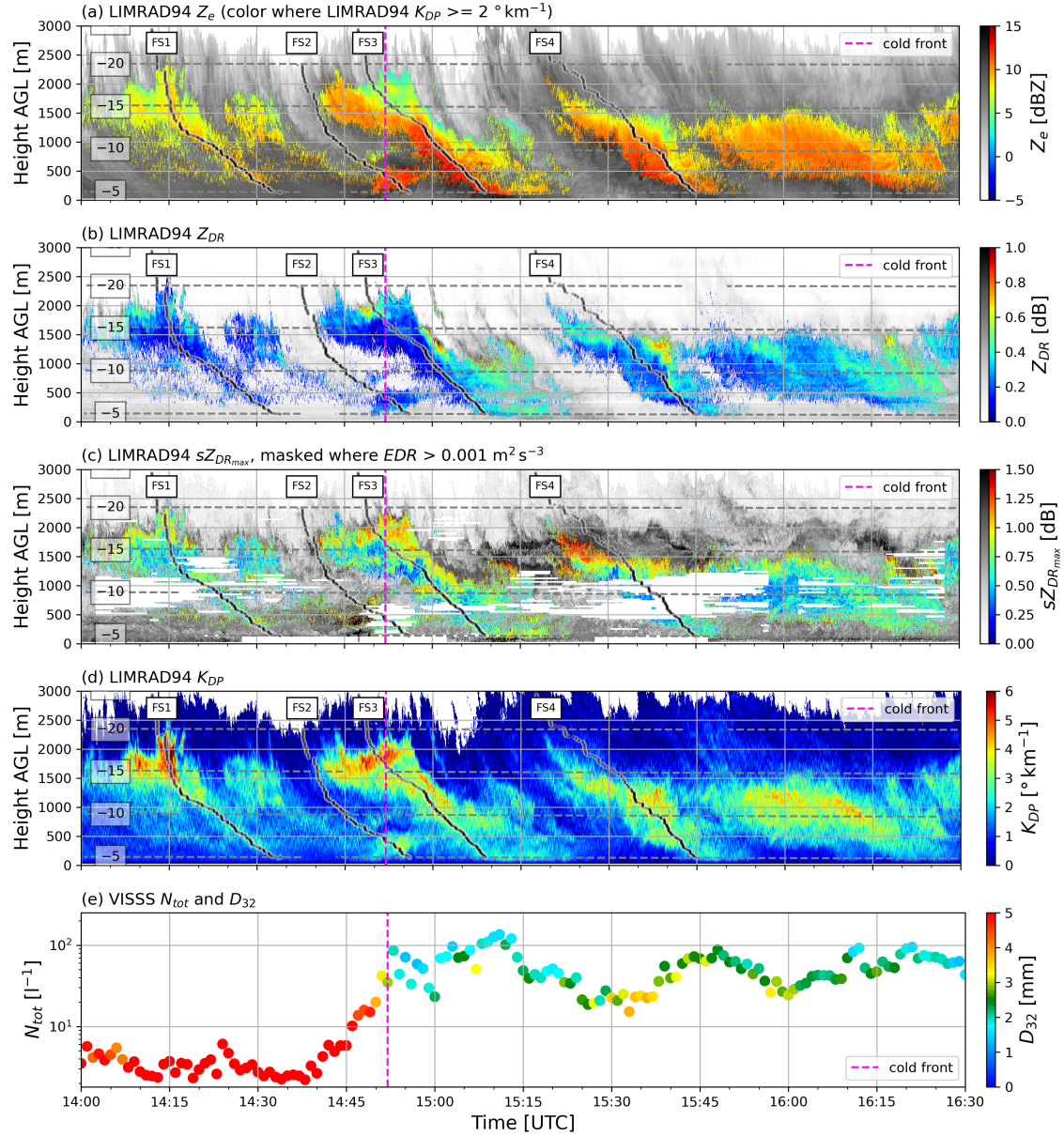


Figure 4. LIMRAD94 time-height plots during the case study on December 6, 2022. Analyzed fall streaks shown as black lines with labels. In every panel except d), data is only colored where LIMRAD94 $K_{DP} > 2^\circ \text{ km}^{-1}$. The remainder of the Z_e and $sZ_{DR_{max}} / Z_{DR}$ observations where K_{DP} was below this threshold are displayed in black and white. Grey lines are isotherms. a) Z_e . b) Z_{DR} . c) $sZ_{DR_{max}}$ with areas masked where EDR exceeds the turbulence threshold of $0.001 \text{ m}^2 \text{ s}^{-3}$. d) K_{DP} . e) Minutely VISSS N_{tot} on the y-axis scattered against time on the x-axis. Colors are the respective D_{32} values.

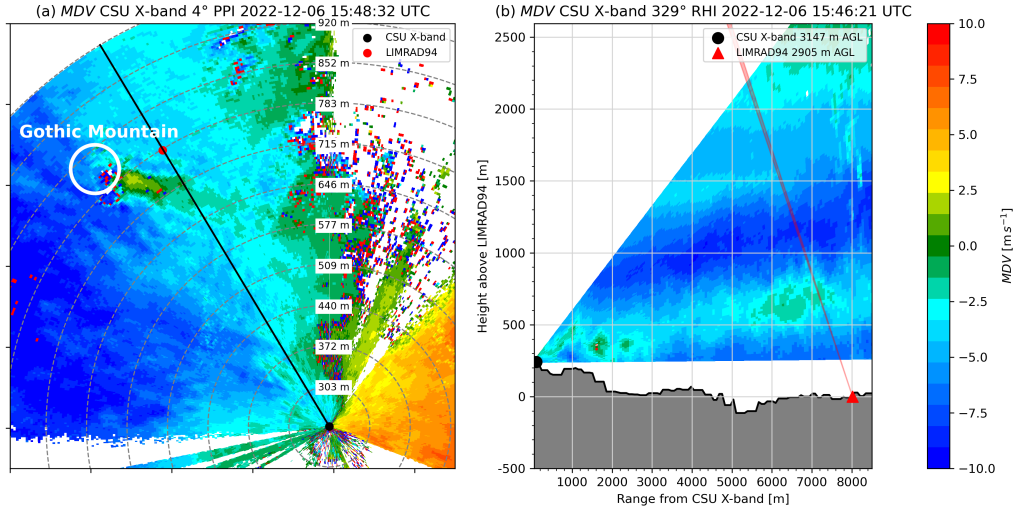


Figure 5. a) Mean Doppler velocity (approx. radial component of horizontal wind) of CSU X-band PPI scan at 4° elevation angle with location of Gothic Mountain (white circle) and direction of RHI scan at azimuth 329° (black line) at 15:48:32 UTC on Dec 6, 2022. Grey dashed circles are placed at every 1000 m range from CSU X-band. Labels are the respective heights above LIMRAD94 of the CSU X-band radar beam. b) Mean Doppler velocity of CSU X-band RHI scan (azimuth 329°) at 15:46:21 UTC on Dec 6, 2022 with LIMRAD94 beam direction (red). Grey shading shows the elevation profile.

the MDV values turn positive in sharp contrast to the surrounding negative values. This implies that downstream of Gothic Mountain, the wind turns to more southerly directions. The effect diminishes quickly further downstream. In the RHI scan (Fig. 5b), the southerly wind component is visible by the area of more green colors at 6000-7000 m range from CSU X-band and between 400-1000 m above LIMRAD94. This area of lee-induced flow disturbance (termed ALIFD) could be partly caused by a counterflow, indicative of a blocked low-level flow, as, for example, described in Ramelli et al. (2021). Along the edges of the ALIFD, wind shear is enhanced which could lead to the observed turbulence. The beam of LIMRAD94 is shown in Fig. 5b) as a red cone, and it is apparent that the measurements of LIMRAD94 are impacted by the turbulence in and around the ALIFD.

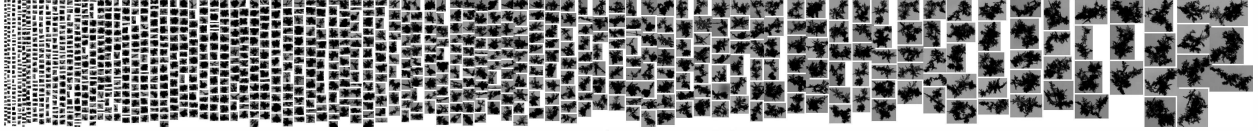
3.2.3 Evolution of polarimetric variables along fall streaks (FS)

In the following sections, the development of W-band polarimetric variables along the four fall streaks of enhanced K_{DP} before, during, and after the cold front are analyzed in detail. VISSS images of hydrometeors during the fall streaks and the respective PSD are shown in Fig. 6 and Fig. 7, respectively. The polarimetric radar observations during the four selected fallstreaks (FS1-4) are shown in Fig. 8. All subplots in Fig. 8 have the same structure, showing profiles of K_{DP} , Z_e , $sZ_{DR_{max}}$, Z_{DR} , spectral width (SW) and correlation coefficient (ρ_{hv}) as a function of the temperature. Fall streaks were derived as explained in Chapter 2.5. A grey line was added for every height / temperature where EDR exceeded $0.001 \text{ m}^2 \text{ s}^{-3}$. SW of

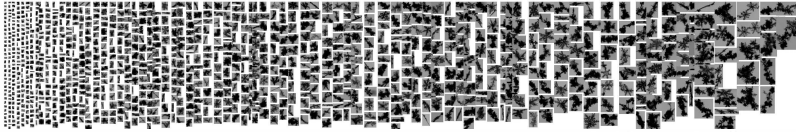


SAIL: 2022-12-06 VISSS Particles during Fallstreaks, 10 mm = 170.2 px

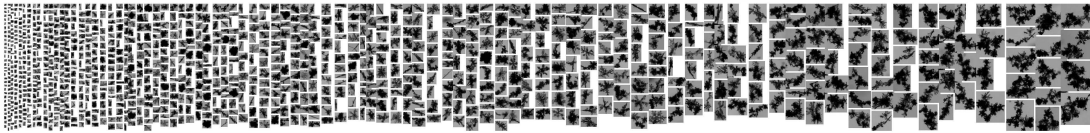
(a) FS1: 14:33-14:35 UTC



(b) FS2: 14:55-14:57 UTC



(c) FS3: 15:08-15:10 UTC



(d) FS4: 15:44-15:46 UTC

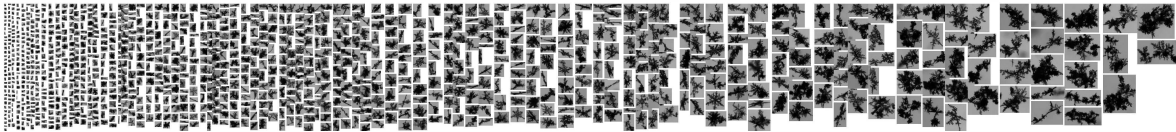


Figure 6. Particles observed on the ground by VISSS during the four fall streaks. For each fall streak, 3 minutes of data are displayed. Particles were divided into size bins and a fixed number of particles is displayed for each bin (if the bin contains a sufficient number of particles). The number of displayed particles is decreasing with increasing particle size.

LIMRAD94 was added to the plots as an additional proxy for turbulence (as for example done in Majewski et al., 2023). For each of the analyzed fall streaks, we will try to assess which particle types and microphysical growth processes lead to the observed K_{DP} signatures and how they are influenced by turbulence.

325 Before the cold front (FS1)

The first K_{DP} fall streak (FS1) we analyze appears around 14:15 UTC, originating in the DGL at around 2300 m AGL and temperatures between -20 and -15 °C (Fig. 4d). It reaches the ground well before the cold front at around 14:30 UTC. Before the cold front passage, the dominant particle type is aggregates, smaller heavily rimed or graupel-like particles were also observed by the VISSS (see Fig. 6a). Before 14:45 UTC, N_{tot} is always below 101^{-1} and D_{32} is around 4-5 mm (Fig. 4e),
 330 indicating that the particle population at the surface contains larger particles (see also Fig. 7).

The polarimetric variables along FS1 are shown in Fig. 8a). At temperatures colder than -20 °C, K_{DP} is close to zero and $sZ_{DR_{max}}$ rather low with values around 0.25 dB. Z_{DR} is slightly negative with -0.2 dB. The negative Z_{DR} observed above the DGL is most likely caused by differential attenuation within the DGL. Differential attenuation at W-band by ice is

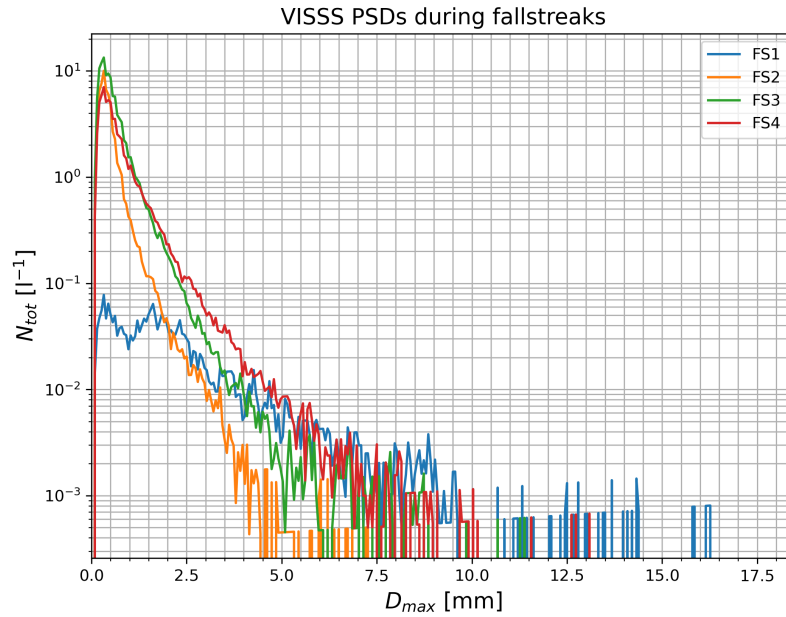


Figure 7. VISSS PSDs (one minute averages) observed during the times when the four selected fall streaks reached the surface.

either related to large particle number concentrations or high mass densities or a combination of both (Liao et al., 2008). ρ_{hv} in this upper part of FS1 is high with values above 0.995. This combination of Z_{DR} , K_{DP} and ρ_{hv} suggests that the particle population is rather uniform and consists either of nearly spherical particles, particles with a low effective refractive index or a combination of the two (e.g. Andrić et al., 2013).

As these particles fall into the DGL, they experience depositional growth through enhanced supersaturation over ice. $sZ_{DR_{max}}$ and Z_{DR} increase because depositional growth in the DGL promotes anisotropic shape enhancement (e.g. Andrić et al., 2013; Takahashi, 2014). The sharp decrease in ρ_{hv} together with the difference in $sZ_{DR_{max}}$ and Z_{DR} could be linked to a secondary particle mode consisting of freshly formed, anisotropic particles is often observed in the DGL (described in Sect. 1). $sZ_{DR_{max}}$ starts to increase at around -20°C to values of around 0.75 dB (center panel in Fig. 8a). The 99th percentile of $sZ_{DR_{max}}$ during the 2.5 hours of our case study is 1.32 dB, so 0.75 dB is still comparably small. The comparably low $sZ_{DR_{max}}$ in the DGL during FS1 indicates either more spherical, less dense or smaller anisotropic particles. Collocated with the maximum $sZ_{DR_{max}}$ and Z_{DR} , we see a sharp increase in K_{DP} (Fig. 8a) to over 6°km^{-1} . Following the hypothesis of Griffin et al. (2018) described in the Introduction, the peaks of low Z_{DR} and high K_{DP} in FS1 between -20 and -15°C could be caused by a high concentration of I-type particles. If I-type particles enter the DGL, they will experience rapid depositional growth because of the enhanced supersaturation over ice. Larger particles will grow slower and mostly maintain their shape, while smaller I-types will grow faster and could serve as embryos of rapid growing D-type particles (e.g. Sheridan et al., 2009; Griffin et al., 2018). The existing water vapor would be distributed over a large number of particles, which, as we speculate,

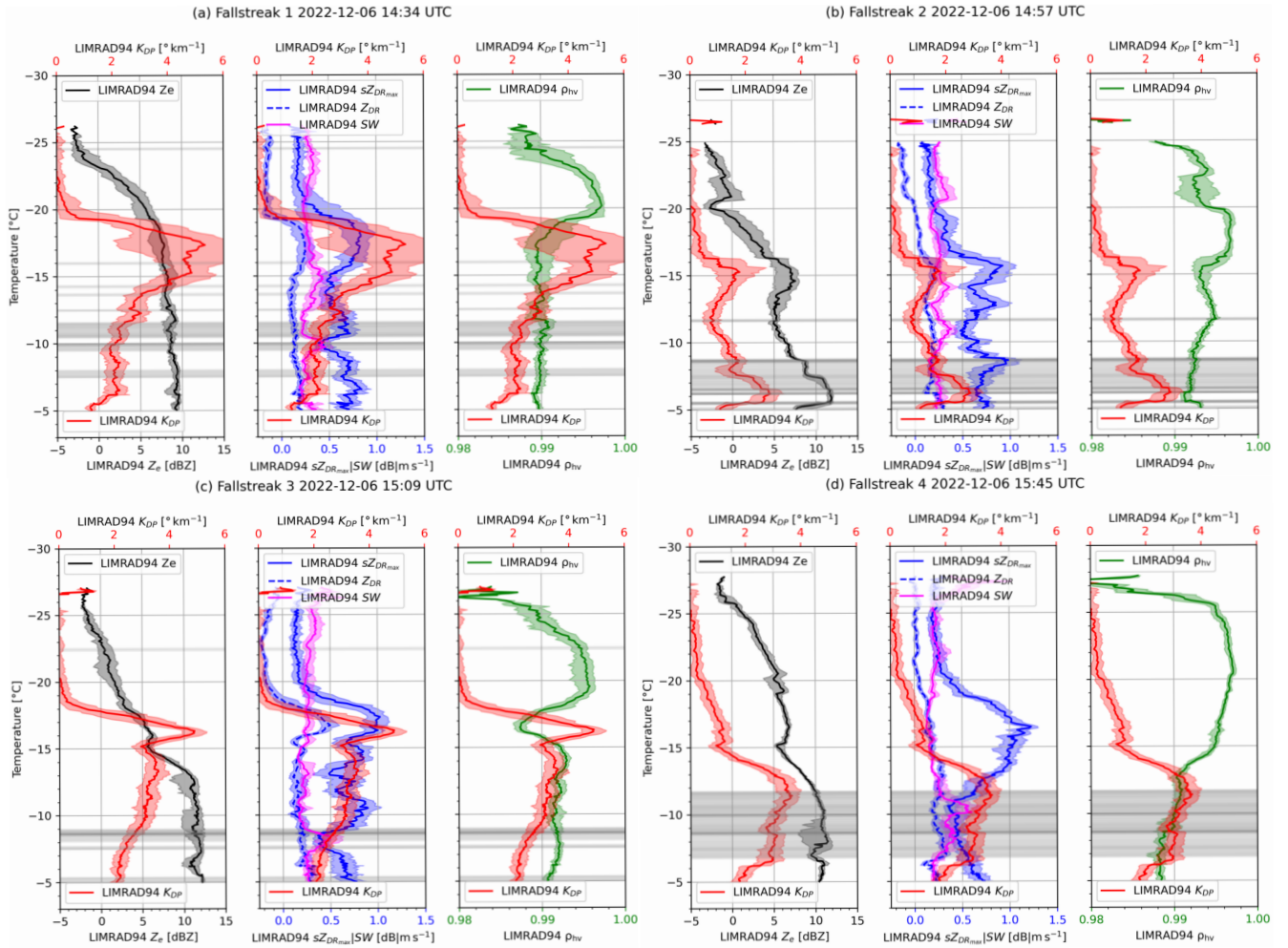


Figure 8. Profiles of polarimetric variables along the FS1 (a), FS2 (b), FS3 (c) and FS4 (d). Profiles of K_{DP} (red), Z_e (black), $sZ_{DR_{max}}$ (blue), Spectral Width (magenta) and ρ_{hv} (green) are plotted against the temperature. All polarimetric variables were averaged (median) for ± 2 min around the fall streak track (solid lines). Colored shaded areas mark 25/75th percentile values of the respective variable. Grey shaded areas mean that the EDR exceeded the turbulence threshold of $0.001 \text{ m}^2 \text{ s}^{-3}$.

would result in overall lower supersaturation and cause the newly formed D-type particles of larger quantities to grow into overall lower diameters and lower density. So while Z_{DR} could be masked by more spherical I-type particles, $sZ_{DR_{max}}$ in FS1 could be low because of D-type particles with an overall smaller diameter. At the same time, these particles are more numerous, which enhances K_{DP} and causes a collocation of the maximum $sZ_{DR_{max}}$ and K_{DP} . This theory would imply that the closer the maxima in K_{DP} and $sZ_{DR_{max}}$ are in height, the higher the overall number concentration of particles in the upper DGL. Another potential cause of the observed radar signatures could be additional SIP at the upper part of the DGL as e.g. suggested in Andrić et al. (2013) and Von Terzi et al. (2022). A large number of I-type particles with a wider size distribution would also

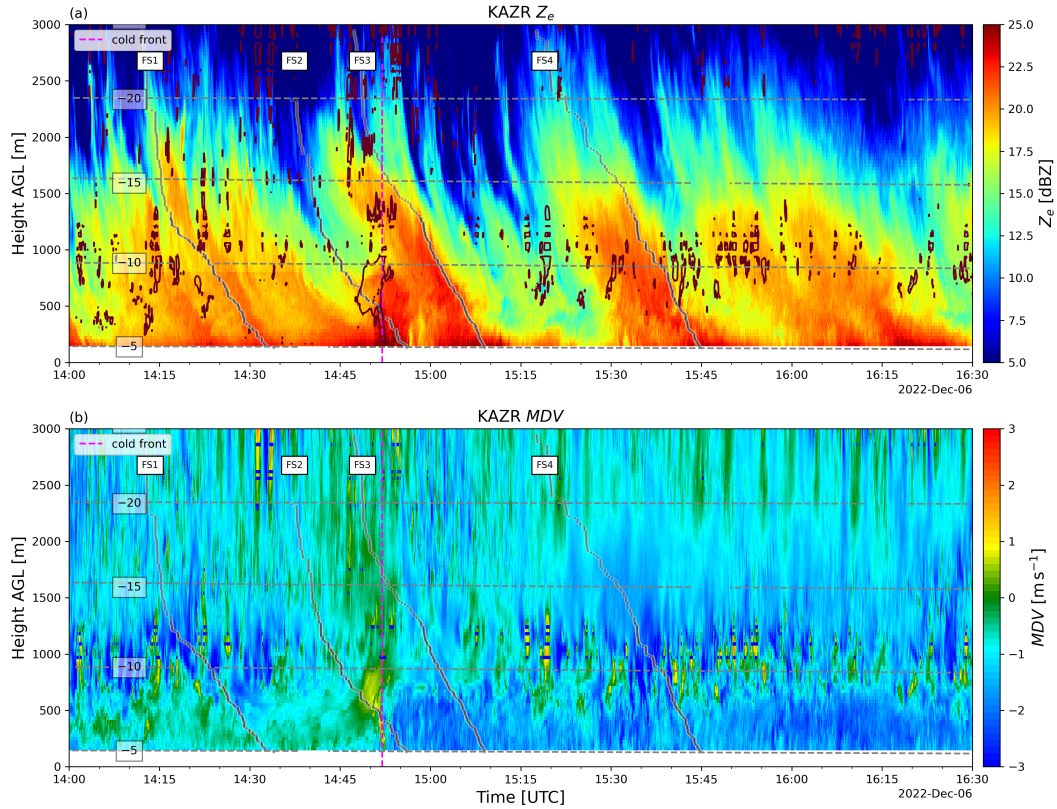


Figure 9. Time-height plots of KAZR MDV (panel a and c) and KAZR Z_e (panel b and d) of the December 6, 2022 case study. The red contour in panel b and d marks the 0 $m s^{-1}$ isotach. Dashed grey lines in panel b and d mark the isotherms (in °C) with respective labels. The analyzed fall streaks retrieved from LIMRAD94 measurements are displayed as black lines with their respective labels. Note that these fall streak paths do not exactly match the KAZR Z_e because of the horizontal displacement of the measurement volumes.

enhance the likelihood of ice collisional fragmentation, providing even more embryos for the growth of D-type particles. A second, smaller peak of K_{DP} at around $-15^{\circ}C$ in FS1 could be caused by additional dendritic growth at this temperature. A similar structure of K_{DP} in the DGL was for example shown in simulations by Andrić et al. (2013), caused by dendritic growth at this temperature. At temperatures warmer than $-15^{\circ}C$, we see an overall decrease in K_{DP} towards the ground with a collocated moderate increase of Z_e , a signal commonly attributed to the aggregation of anisotropic particles into less dense, more spherical shapes and the reduction of N_{tot} (e.g. Andrić et al., 2013; Trömel et al., 2019; Von Terzi et al., 2022; Hu et al., 2024).

The summit height of Gothic Mountain is roughly located around the $-10^{\circ}C$ isotherm. Turbulence is found at this height (Fig. 8a). Just above the turbulent layer, we see an increase in $sZ_{DR_{max}}$ together with a small peak in K_{DP} . A similar signature of enhanced Z_{DR} above a turbulent layer was found by (Hogan et al., 2002; Grazioli et al., 2015). They explained that SLW and ice splinters are transported upward by updrafts and grow into anisotropic particles. In our case, the splinters would grow into



D-type particles given temperatures between -10 and -15°C which might explain the enhanced $sZ_{DR_{max}}$ and K_{DP} . Below
 370 the turbulent layer, we find that $sZ_{DR_{max}}$ and also K_{DP} increase slightly again while Z_{DR} remains almost constant. LWP
 up to 200 g m^{-2} (see Fig. 3d) during FS1 enables riming, and graupel is detected by VISSS (see Fig. 6 FS1). The enhanced
 $sZ_{DR_{max}}$ and K_{DP} could be attributed to graupel-snowflake collisions, which are known to produce anisotropic ice fragments
 with sizes up to 2 mm (Grzegorzczak et al., 2023). In the first stage of riming, only the density of particles increases without
 changes in density of shape (Erfani and Mitchell, 2017), which might also cause the observed increase in $sZ_{DR_{max}}$ and K_{DP}
 375 below the turbulent layer. The PSD observed by VISSS during the time FS1 reaches the surface shows a comparably low
 number of small particles (Fig. 7) together with the presence of larger aggregates (see Fig. 6 FS1). It is likely that aggregation
 rapidly depletes the number of small particles towards the ground and therefore decreases N_{tot} .

During the cold front (FS2)

With the passage of the cold front at the surface, we see a distinct increase in N_{tot} in the VISSS measurements from less than
 380 101^{-1} to around 1001^{-1} (Fig. 4e). At the same time, D_{32} (colors in Fig. 4e) decreases to around 2 mm, indicating smaller
 particles are dominant in the measurement volume. Compared to FS1 before the cold front, the VISSS PSD in FS2 (Fig. 7),
 shows an overall increase/decrease in particles smaller/larger than 2 mm, respectively. The sudden increase in the number of
 small particles could be caused by blowing snow, which is supported by multiple observations. As the hours before the cold
 front featured conditions with weak 2 m winds and large aggregates, loose snow on trees or the ground was likely lifted along
 385 the convergence of the surface cold front and behind it by the suddenly increased northerly wind. KAZR data during the case
 study is shown in Fig. 9. The lift associated with the cold front can be seen as positive KAZR MDV (i.e., upward motion) in
 Fig. 9b. The strongest upward motion was observed at around 14:52 UTC in the lowest 500 m along the surface convergence.
 The low-level uplift provided by the cold front supports the theory that blowing snow was lifted and kept elevated in the vicinity
 of the surface convergence. In the presented case, blowing snow could also be detectable at higher altitudes because it is lifted
 390 off the higher valley slopes and advected over the lower valley floor. Figure . 9a shows the KAZR Z_e together with the 0 m s^{-1}
 isotach (red line). Between 14:50 and 14:55 UTC and roughly between 250 and 900 m AGL, there is an area of enhanced Z_e
 with no connection to a FS, supporting the idea of Z_e -enhancement caused by blowing snow.

Between 14:50 and 14:55 UTC, we also see a small area of enhanced LIMRAD94 K_{DP} in the lowest 500 m (Fig. 3b).
 Figure . 8b shows a possible fall streak track during the time of the cold front passage. FS2 displays a sudden change in the
 395 polarimetric variables and hence likely the particle mode when entering the pocket of cool and dryer air between -8 and -5°C
 (Fig. 8b). The boundary between the two air masses is marked by an increase in turbulence (Fig. 8b). Below the top of the
 turbulent layer, $sZ_{DR_{max}}$ and K_{DP} decrease sharply which could however be due to a turbulence masking effect. Within the
 turbulent layer, K_{DP} then increases again downwards to values of up to 3° km^{-1} at the lower edge of the turbulent layer.
 Assuming that K_{DP} is still masked by turbulence, the true K_{DP} values without turbulence would be much higher. Z_e also
 400 increases by about 5 dBZ within the turbulence layer, its peak value is collocated with the K_{DP} peak. We conclude that this
 K_{DP} signature is likely caused by blowing snow.



At the top of the turbulent layer, we also see an increase in $sZ_{DR_{max}}$ together with a small peak in K_{DP} , similar to FS1. The explanation could again be small ice splinters being transported upward by updrafts (Hogan et al., 2002; Grazioli et al., 2015). Ice splinters could originate either from SIP processes inside the turbulent layer or from lifted blowing snow. At temperatures
 405 between -5 and -8 °C the splinters can grow into columns or needles (Bailey and Hallett, 2009). Indeed, some needles were observed by the VISSS for FS2 (Fig. 6b).

After the cold front (FS3, FS4)

FS3 forms between 14:45 and 15:00 UTC at around 2000 m and reaches the ground around 15:10 (FS3). FS4 forms between 15:25 UTC and 15:35 UTC at 1500-2000 m and reaches the surface just after 15:45 UTC (FS4). The surface relative humidity
 410 (RH) decreases noticeably before each fall streak reaches the ground (Fig 3c) and increases when the FS reaches the ground. At the same time, surface N_{tot} increases (Fig 4e). This could point to drier air that is advected behind the cold front, which is moistened again by the sublimation of precipitating snow.

FS3 originates in the upper dendritic growth layer (DGL) at around 2000 m and temperatures between -15 and -20 °C. The maximum K_{DP} values in the DGL reach 5 ° km $^{-1}$ (Fig. 8c), which is slightly lower than in FS1. The peaks of $sZ_{DR_{max}}$ and
 415 Z_{DR} have values of around 1 dB and 0.5 dB, respectively which is higher than in FS1. Comparing Z_e at -20 °C in FS1 and FS3, we find Z_e to be about 6 dBZ in FS1 and 2 dBZ in FS3. Assuming that the size of I-type particles is similar in both FS, this would point to a lower concentration of I-type particles sedimenting in the DGL in FS3 and thus less I-types serving as embryos for D-type growth. This in turn would lead to larger D-type particles (since the existing water vapor is distributed over fewer particles) with larger $sZ_{DR_{max}}$ (and higher Z_{DR} due to less masking) and still in a sufficiently high number concentration to
 420 produce high K_{DP} .

In FS3, the increase in $sZ_{DR_{max}}$ and K_{DP} coincides with a sharp decrease in ρ_{hv} (Fig. 8c) which quickly recovers near -15 °C, while K_{DP} experiences a sudden drop of approximately 2 ° km $^{-1}$. This signature could indicate rapid aggregation or riming, processes typically associated with an increase in Z_e . However, during the K_{DP} decrease, Z_e also shows a slight decline. This behavior may reflect the limitations of interpreting a three-dimensional system with two-dimensional data. Be-
 425 tween -15 °C and -12 °C, Z_e increases strongly, indicating a possible increase in particle diameter which might point to aggregation. In the same T-range, K_{DP} is nearly constant with high values of around 3 ° km $^{-1}$. This points to the formation of new anisotropic particles in parallel to the aggregation process, which continues to enhance the number concentration and hence keeps K_{DP} constant. Nevertheless, it is not ruled out that aggregates contribute to the magnitude of the K_{DP} in this part of FS3.

The whole area of enhanced K_{DP} around the black line marking FS3 between 14:40 and 15 UTC reveals another interesting feature: inside the area of high K_{DP} values, a horizontal gradient in Z_e , Z_{DR} and $sZ_{DR_{max}}$ is visible (Fig. 4a-d). Z_{DR} and $sZ_{DR_{max}}$ are especially low on the left side of black line marking FS3 while K_{DP} exceeds 4 ° km $^{-1}$. Following the hypothesis of Griffin et al. (2018), this might point to I-type particles seeding in the DGL in parts of the fall streak and dominating the particle population. Z_{DR} might additionally be masked by more spherical, low density aggregates. The presence of these larger
 435 aggregates could also explain the higher Z_e on the left side of FS3. On the right side of FS3, higher $sZ_{DR_{max}}$ and Z_{DR} might



be caused by the presence of dense, anisotropic D-type particles. The lower Z_e and less masked Z_{DR} might be related to the absence of large aggregates.

Below the turbulent layer in FS3, K_{DP} decreases, but still has values of $2\text{--}2.5^\circ \text{ km}^{-1}$ until the fall streak reaches the surface, while Z_e increases slightly towards the surface (Fig. 8c). Where K_{DP} decreases between -10 and -7°C , we also see
 440 an increase in ρ_{hv} which might point to enhanced aggregation due to the turbulent layer as for example observed by Chellini and Kneifel (2024). The observation of graupel, dendrites, and many dendrite fragments by VISSS (see Fig. 6c) points to frequent graupel-snowflake collisions. These collisions were found to produce up to 500 fragments per collision Grzegorzczuk et al. (2023). The fragment size distributions of such collisions revealed a mode at around 0.4 mm in their study. In the PSD measured by VISSS (Fig. 6 FS3), particles with a D_{max} of less than 0.5 mm were the most frequently observed. A portion
 445 of these particles might be fragments of graupel-snowflake collisions, which could also explain the very high N_{tot} of around 1101^{-1} at the surface during FS3.

In FS4 (see Fig. 8d), we find a very prominent peak in $sZ_{DR_{max}}$ of 1.25 dB between -17 and -13°C . Contrary to the previous fall streaks, this peak is firstly occurring much further above the K_{DP} peak and secondly has higher values than in FS1 and FS3. No collocated Z_{DR} signature exists, which can only mean that the Z_{DR} signature is completely masked by larger spherical
 450 particles. This fits to the theory of (Griffin et al., 2018): only very few and large D-type particles are present and produce high $sZ_{DR_{max}}$ while K_{DP} is negligible due to their low number concentration. At T warmer than -15°C , $sZ_{DR_{max}}$ decreases while Z_e and K_{DP} increase. This signature might point to the early stage formation of D-type aggregates. If there was a layer of SLW at the top of the turbulent layer as mentioned by Grazioli et al. (2015), it is also possible that the dendrites from above become rimed, forming even graupel. The resulting velocity gradient in the PSD would favour collisions and hence collisional
 455 breakup. Resulting splinters would quickly grow into anisotropic ice crystals, enhance the number concentration and therefore K_{DP} . However, our data does not allow a definitive conclusion about the origin of this polarimetric signature.

Between around 1200 m and 500 m AGL corresponding to -12 and -7°C , FS4 is influenced by the turbulent layer (Fig. 4c and Fig. 8d) which is also evident in Fig. 9b) as strongly fluctuating MDV layer. In general, in FS4 the turbulent layer is more pronounced than during the previous fall streaks leading to a decrease in $sZ_{DR_{max}}$ and Z_{DR} as well as an increase in SW .
 460 K_{DP} is rather homogeneous throughout FS4 inside the turbulent layer, horizontally and vertically. A decrease near the ground is only seen in the lowest 200 m . It seems likely that the constant K_{DP} is caused by increased ice collisional fragmentation that keep the number concentration of anisotropic particles high throughout the turbulent layer. VISSS images at the time of FS4 shows a wide mix of particles, many dendrite branches, aggregates in various stages of riming, and some graupel (see Fig. 6d). This confirms that ice collisional fragmentation and riming are likely major processes inside the turbulent layer. This
 465 is consistent with the findings of (e.g., Ramelli et al., 2021; Chellini and Kneifel, 2024). Similar to FS3, the ice collisional fragmentation in the observed turbulent layer might be the reason for the high N_{tot} of 1001^{-1} during FS4.

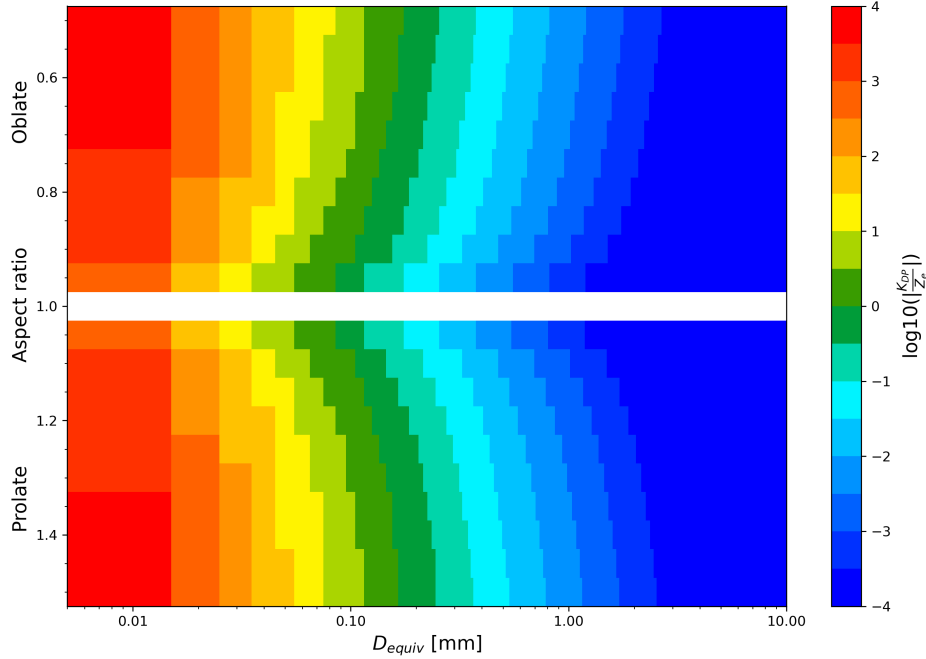


Figure 10. T-matrix simulation results: Logarithm of the ratio between X-band K_{DP} and Z_e plotted for different aspect ratios (y-axis) and equivolumetric diameters (logarithmic x-axis)

3.3 Estimating particle size responsible for W-band K_{DP} signatures using T-matrix simulations and CSU X-band Z_e and K_{DP} observations

In the previous section, we discussed the fingerprints of microphysical processes influencing polarimetric radar variables and hypothesized about the particles responsible for the observed K_{DP} . Here, we use T-matrix simulations and CSU X-band radar observations to constrain the sizes of particles contributing to K_{DP} signatures in the observed fall streaks. Instead of using LIMRAD94 data, the ratio of K_{DP} to Z_e at X-band is analyzed, as its Z_e is not affected by non-Rayleigh scattering for particle diameters smaller than 10 mm.

T-matrix simulations were performed to calculate the ratio of K_{DP} to Z_e at X-band for artificial spheroidal particles with varying aspect ratios. These calculations used Z_e in linear units ($\text{mm}^6 \text{m}^{-3}$). Since both K_{DP} and Z_e are proportional to the concentration of particles, the ratio of these radar variables is therefore independent of the concentration. The ratio is thus mostly sensitive to size and shape of ice particles. The ratio decreases with an increase in particle diameter, because in snow, Z_e and K_{DP} are roughly proportional to the 4th and 1st power of the particle's size, respectively. Figure 10 shows the results of the T-matrix simulations, the logarithm of the ratio between K_{DP} and Z_e at X-band is plotted for different aspect ratios. Even though the ratio depends on the axis ratio of particles, it is suitable for roughly estimating size of particles producing K_{DP} signatures.

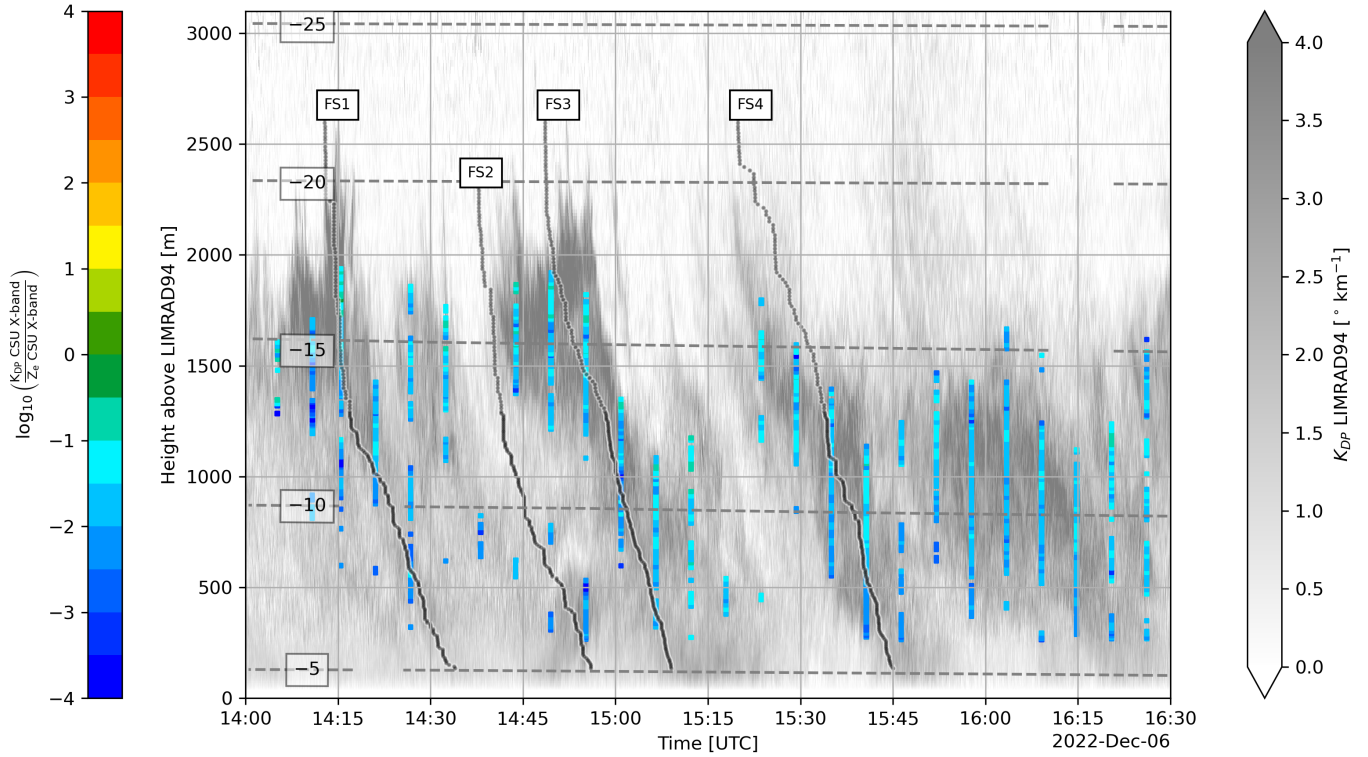


Figure 11. Time-height plot of LIMRAD94 K_{DP} for the Dec 6, 2022 case study. CSU X-band data is superimposed on LIMRAD94 K_{DP} when the latter is $\geq 2^\circ \text{ km}^{-1}$. Colored squares: Logarithm of the ratio between X-band radar K_{DP} and Z_e for RHI scans. The analysed fall streaks as well as isotherms are indicated in grey.

To assess the observed ratios of K_{DP} and Z_e , we analyzed CSU X-band data within the K_{DP} fall streaks (see Fig. 11). The CSU X-band K_{DP} was divided by $\cos^2$ of the elevation angle to correct for the different elevation angles during the RHI scans (Ryzhkov and Zrnich, 2019). The measured ratio of K_{DP} to Z_e typically ranges from -2.5 to -1 during fall streaks. In the upper DGL, values tend to be less spread, ranging from -2 to -1 while they decrease down to -2.5 towards the surface which might be linked to the increasing particle size through aggregation and/or depositional growth. T-matrix calculations (Fig. 10) show that K_{DP} to Z_e ratio values of $-2/-2.5$ cannot be achieved when K_{DP} signatures are produced solely by ice particles larger than approximately $0.8/1.2$ mm, respectively. Considering that during the fall streaks, VISSS showed mean mass-weighted diameters ranging from 1.5 to 5 mm, we conclude that the contribution of large ice crystals to K_{DP} is small and a major part of the K_{DP} signatures is produced by ice particles smaller than 0.8 mm in diameter in the DGL and smaller than approximately 1.2 mm closer to the surface.

The ratio of K_{DP} and Z_e of course depends on the absolute calibration of CSU X-band, and for this case study we roughly estimated a calibration offset of 4 dB by comparing Z_e of CSU X-band with Z_e from LIMRAD94 (not shown). However,



considering the several orders of magnitude range of the K_{DP} to Z_e ratios, this 4 dB offset would not significantly affect this analysis. In addition, we emphasize that the ratio alone is insufficient to estimate the exact size of particles when different populations of ice particles coexist within the same volume. To illustrate this, consider two populations of particles, such as large aggregates and small ice particles. The large aggregates primarily determine Z_e due to their size, while the high concentration of small particles may dominate the K_{DP} values. Thus, Z_e and K_{DP} effectively characterize different particles within the volume. Nevertheless, the ratio can still be used to constrain the maximum possible size of particles dominating the K_{DP} values.

3.4 DDA simulations of W-band K_{DP} based on VISSS data

Based on the results obtained from the analysis of CSU X-band data, we expect that a major part of the K_{DP} signatures at W-band are produced by small ice particles. To further investigate the possible contribution of large particles, a more detailed analysis is required. The T-matrix method is not applicable to W-band measurements because large particles may produce resonance effects. In such cases, the spheroidal approximation does not always adequately represent the scattering properties (e.g. Schrom and Kumjian, 2018). To test the possible contribution of larger aggregates to our observed LIMRAD94 K_{DP} signatures, we calculated K_{DP} and Z_e based on VISSS in situ observations using DDA simulations. The observations of LIMRAD94 K_{DP} and Z_e are temporally averaged to 1 min and vertically averaged between 30 m (lowest range gate) and 400 m AGL. The ratio between observed and simulated K_{DP} and Z_e values at W-band are displayed in Fig.12 (a/b) (blue/black) solid line, respectively. For DDA simulations of K_{DP} , only particles ≥ 2.5 mm were used which we assume to be aggregates. We determined this size threshold based on images and the spectral complexity measurements from VISSS (not shown). The particles for the simulations were selected from the DDA database using a nearest neighbor regression, considering the nearest 5 particles and weighting them with their inverse distance to calculate the scattering properties expected for the in situ particle based on their maximum dimension. All of these particles were assumed to be dendrite aggregates, which fits the VISSS observations of particles larger than 2.5 mm.

Overall, the temporal evolution of the simulated K_{DP} and Z_e appears to match the observations quite well, except for the time during FS2 which is strongly influenced by blowing snow. During FS1, we see that the K_{DP} ratio is about 20 %, which is the value that Von Terzi et al. (2022) found as contribution of larger aggregates to W-band K_{DP} in their study. The simulated Z_e is about 80 to 100 % of the observed Z_e . This is about the range we would expect, given that the PSD during FS1 contains a comparably low number of particles < 2.5 mm. This also underlines that even when the larger aggregates determine Z_e , the smaller ice crystals still produce the majority of K_{DP} without contributing much to Z_e . This is consistent with our findings from the previous section 3.3.

During the cold front in the vicinity of FS2, there is only a very low number of larger aggregates and K_{DP} is mainly produced by small blowing snow particles as we explained in Sect. 3.2.3. Meanwhile, there are only few aggregates ≥ 2.5 mm (see blue line in Fig. 12c), hence their contribution to Z_e and K_{DP} is very small. During FS3, the number concentration of aggregates ≥ 2.5 mm is similar to FS1 while the number of particles < 2.5 mm is about two orders of magnitude higher than during FS1. This might explain why the simulated Z_e of particles ≥ 2.5 mm only explains 20-30 % of the observed Z_e . The

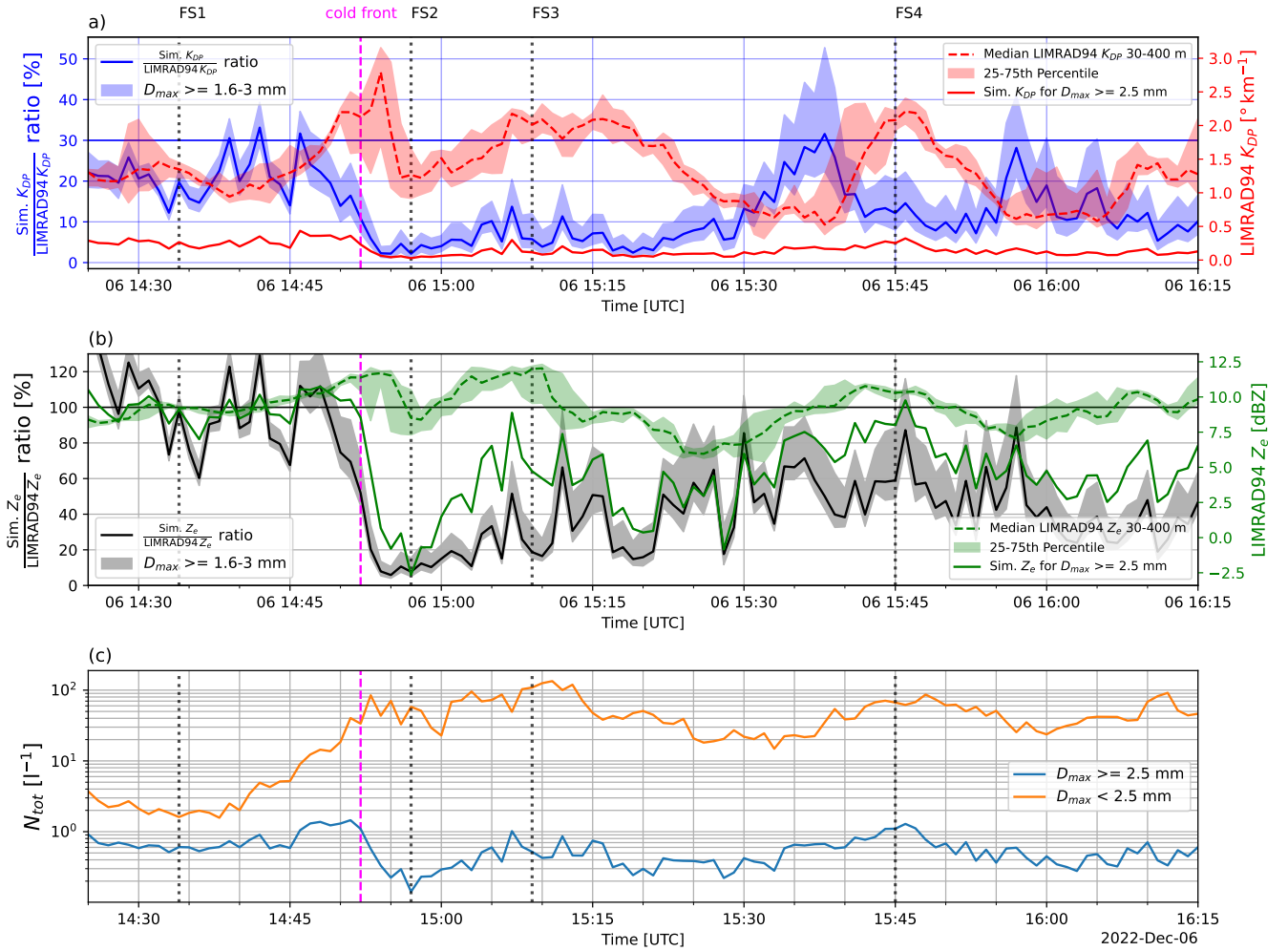


Figure 12. Panel a): Median LIMRAD94 K_{DP} between 30 and 400 m AGL (red dashed line) with the 25-75th percentile (red shading) and ratio of DDA-based forward simulated K_{DP} (Sim. K_{DP}) to observed LIMRAD94 K_{DP} (blue line). For the blue line, only particles with diameter ≥ 2.5 mm were used, the blue shading shows the range of the simulations when just particles ≥ 3 mm (lower edge) or ≥ 1.6 mm (upper edge) are included. Panel b): Median LIMRAD94 Z_e between 30 and 400 m AGL (green dashed line) with the 25-75th percentile (green shading) and ratio between DDA-based forward simulated Z_e (Sim. Z_e) and observed LIMRAD94 Z_e (black line). For the black line, only particles with diameter ≥ 2.5 mm were used, the grey shading shows the range of the simulations when just particles ≥ 3 mm (lower edge) or ≥ 1.6 mm (upper edge) are included. In panel c), VISSS particle number concentrations for particles with $D \geq 2.5$ mm (blue) and $D < 2.5$ mm (orange) are shown. Magenta line marks the passage of a cold front while the black dotted lines mark the analyzed fall streaks.

observed K_{DP} is higher than during FS1 and the simulations suggest that larger aggregates contribute less than 10% to the observed K_{DP} . Right before and after FS3 we see a higher number of larger aggregates which may contribute up to 15% of the observed K_{DP} .



During FS4, the number of particles ≥ 2.5 mm is highest compared to the other fall streaks, while the concentration of the smaller particles is slightly lower than during FS3. The contribution of bigger particles to K_{DP} is higher (around 15 %), while simulated Z_e of these particles contributes about 80 to 90 % which seems reasonable.

In general, we conclude that a contribution to K_{DP} of 10-15 % can be attributed to aggregates with a maximum diameter ≥ 2.5 mm. When the PSD contains a comparably large number of aggregates, we even see contributions of up to 20 % to K_{DP} without overestimating Z_e in the simulations. In this analysis we completely omit the effects of riming. If riming enhances the density of aggregates, K_{DP} might be initially increased by an increase in density, provided that the aggregates maintain their anisotropic shape. The opposing effect of riming on K_{DP} was briefly discussed in 2.1.1.

4 Summary and conclusions

We presented results obtained from in situ snowfall camera (VISS) measurements and collocated polarimetric Doppler W-band (LIMRAD94) observations collected within the CORSIPP project embedded within the framework of the Atmospheric Radiation Measurement (ARM) Surface Atmosphere Integrated Field Laboratory SAIL campaign (Feldman et al., 2023). LIMRAD94 K_{DP} and additional polarimetric variables along four high- K_{DP} fall streaks during a cold-front passage related snowfall case study on December 6, 2022 were analyzed and linked to VISS observations. To evaluate which particles contribute to the observed K_{DP} signatures, we further tried to constrain the particle sizes that contribute the most to the observed LIMRAD94 K_{DP} using T-matrix simulations and collocated CSU X-band measurements. Lastly, we determined the amount of LIMRAD94 K_{DP} likely produced by aggregates ≥ 2.5 mm using DDA simulations and VISS in situ observations.

To conclude our study, we would like to highlight the following key findings:

1. By combining in situ and remote sensing measurements, we found that at W-band, high K_{DP} magnitudes are produced by a broad range of number concentrations (N_{tot}), ranging between 1 and 100 l^{-1} (see Sect. 3.1 and Fig. 2). Particle populations with N_{tot} below 5 l^{-1} often featured D_{32} above 5 mm, indicating that the PSD contains more large particles. We thus argue that when interpreting W-band K_{DP} signatures, it should be kept in mind that similar W-band K_{DP} values can be produced by strongly variable particle number concentrations.
2. Based on T-matrix simulations using a spheroidal approach and CSU X-band observations, we found that the main contributors to W-band K_{DP} signatures in the analyzed the fall streaks was likely by ice particles with a diameter below 0.8 mm in the DGL and below 1.2 mm closer to the surface (see Sect. 3.3 and Fig. 11). DDA simulations of K_{DP} based on VISS PSDs during our analyzed the fall streaks show, that 10-15 %, in extreme cases up to 20 % of the K_{DP} observed close to the surface during the fall streaks can be attributed to aggregates ≥ 2.5 mm (see Sect. 3.4 and Fig. 12). This agrees with the results of Von Terzi et al. (2022). Revisiting the three possible explanations for high K_{DP} in particle populations with low N_{tot} and high D_{32} (Sect. 3.1), we conclude that small ice particles are the main contributors to the observed W-band K_{DP} and up to one-fifth of the observed K_{DP} can be attributed to large dendrite aggregates. When



interpreting W-band K_{DP} signatures, it should be remembered that the contribution of aggregates to W-band K_{DP} is usually not negligible.

3. We identified blowing snow as another contributing source to increased K_{DP} values during a cold front passage. Specifically, lifted blowing snow likely caused W-band K_{DP} of up to 3° km^{-1} (see Sect. 3.2.3).
 4. Turbulence, more specifically spectral broadening and the increase in the width of the canting angle distribution, significantly reduces the values of Z_{DR} and K_{DP} significantly. The magnitude of $sZ_{DR_{max}}$ seems to be more affected by turbulence than K_{DP} . Therefore, we conclude that it is essential to consider turbulence before interpreting any Z_{DR} or K_{DP} signatures.
 5. Despite the dampening effect of turbulence mentioned above, we observed an increase in K_{DP} inside a strong turbulent layer. The turbulent layer was caused by wind shear along the interface between the blocked valley flow and the cross-barrier flow aloft (see Sect. 3.2.2). VISSS observations measured below the turbulent layers revealed number concentrations of over 100 l^{-1} . Due to high number of dendrite branches in VISSS observations, we suggest that this is caused by increased ice-collisional fragmentation that keeps the number concentration of anisotropic particles high throughout the turbulent layer and hence causes enhanced K_{DP} . This SIP mechanism of enhanced ice-collisional fragmentation in the presence of turbulence was also found in the studies of Ramelli et al. (2021) and Dedekind et al. (2023). Few graupel particles coexisting with unrimed ice observed by VISSS likely further increased the collision efficiency (see Sect. 3.2.3). Furthermore, we found an increase in $sZ_{DR_{max}}$ and K_{DP} at the top of the turbulent layers during FS1, FS2 and FS3, as observed by (Hogan et al., 2002; Grazioli et al., 2015).
 6. Z_{DR} , $sZ_{DR_{max}}$ and K_{DP} maxima of different magnitudes were well collocated in FS1 and FS3 and strongly displaced in FS4. This is likely linked to different concentrations of I-type and D-type particles as described in Griffin et al. (2018). When Z_{DR} signatures were masked by bigger and more spherical particles, $sZ_{DR_{max}}$ proved to be a very useful additional variable. We therefore emphasize the advantage of using both, integrated and spectrally resolved Z_{DR} .
- During FS1, FS3 and FS4, we also hypothesized that early stage aggregates in the lower part of the DGL cloud have contributed to the observed W-band K_{DP} (as, for example, already discussed in Moisseev et al. (2015)). This speculation seems reasonable, as we did show that dendrite aggregates can contribute up to 20 % to W-band K_{DP} . However, in the observed case studies, the lower part of the DGL was about 1000 m AGL. Thus, early stage aggregates were not observed at the ground by the VISSS. Early stage aggregates in the DGL might have a different shape and other polarimetric properties. A topic of further investigation of the CORSIPP dataset could thus be the analysis of polarimetric signatures of early stage aggregates for cases where the lower edge of the DGL was nearly at ground level.

We do not in particular focus on the effects of riming in our modeling studies. Riming however may even increase the contributions of aggregates to K_{DP} . In the early stages of riming, the density may be enhanced without an initial change in the particle shape (Erfani and Mitchell, 2017). K_{DP} might therefore initially increase by an increase in density, provided that the aggregates maintain their anisotropic shape. This influence of riming on aggregate contribution to K_{DP} may therefore be a



topic of further investigation. Naturally, there are caveats when linking surface in situ observations to radar measurements. The observational volume of LIMRAD94, even at close range, is several orders of magnitude larger than the one of VISSS. This can be partly compensated by temporal averaging and consideration of time shifts between the measurements, but generally requires us to assume some homogeneity in the precipitations.

600 In summary, this study highlights the complexity of interpreting W-band K_{DP} in snow. We show that polarimetric W-band radar observations combined with surface in situ observations of snow PSD by the VISSS and radar forward simulations can give valuable insights into cloud and snowfall microphysical processes. Given the increasing use of polarimetric W-band radar observations and that even space-borne mission concepts have been proposed (Illingworth et al., 2018; Battaglia et al., 2022), the sensitivities of the polarimetric variables to microphysical properties should be further explored.



Table A1. Chirp table of LIMRAD94 used during CORSIPP.

Attributes	Chirp 1	Chirp 2	Chirp 3	Chirp 4
Integration time (s)	0.654	0.950	1.198	1.198
Range interval (m)	40-493	493-2000	2000-6000	6000-9000
Range vertical resolution (m)	11.9	11.9	23.8	23.8
Nyquist velocity (m s^{-1})	11.3	7.7	6.1	6.1
Doppler velocity resolution (m s^{-1})	0.044	0.061	0.048	0.048
Doppler velocity bins	512	256	256	256

605 Appendix A: Chirp tables of the polarimetric W-band radar

The chirp table used for the presented statistics and case studies is presented in Tab. A1.

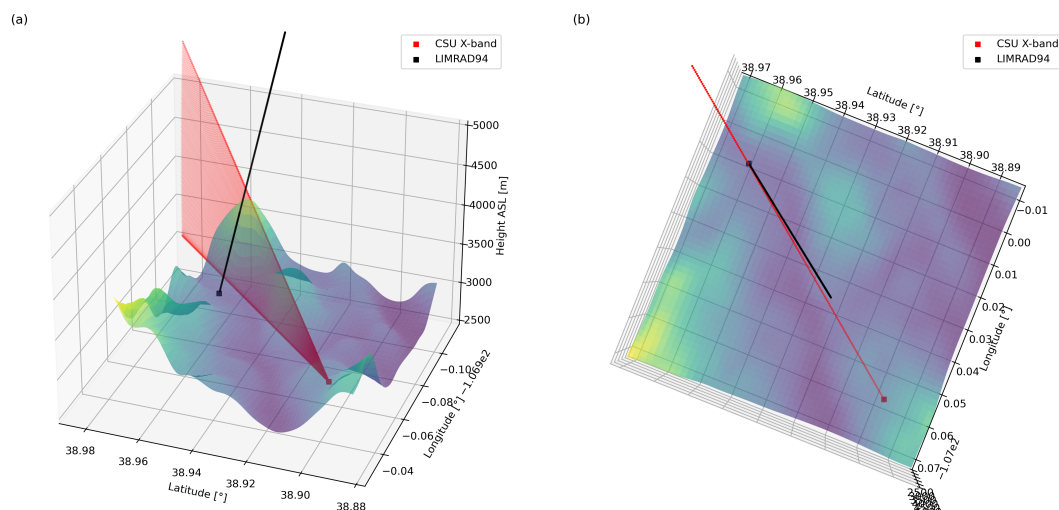


Figure B1. Radar beam overlaps of LIMRAD94 and CSU X-band RHI scan (azimuth 329°) from sideview (panel a) and top view (panel b). Red lines represent the CSU X-band radar beam and black lines the LIMRAD94 radar beam. The size of the actual radar volume is not accounted for in the plot.

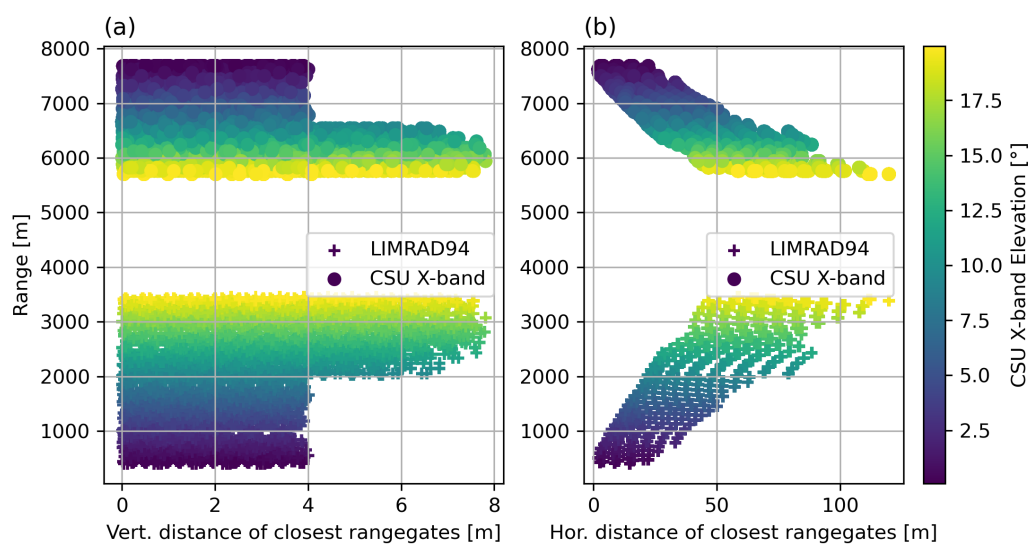


Figure B2. Vertical (panel a) and horizontal distance (panel b) of the closest radar beams from LIMRAD94 and CSU X-band during the RHI scans of CSU X-band plotted against the respective radar range. The colors indicate the respective elevation angle of CSU X-band.



Table B1. Locations and distances of instruments.

Instrument	Latitude/Longitude	Altitude AMSL [m]	Hor. dist. from LIMRAD94 [m]	Azimuth from LIMRAD94 [°]
VISSS	38.955912, -106.987834	2886	550	151
LIMRAD94	38.96031952, -106.99072266	2905	-	-
KAZR	38.956158, -106.987854	2886	525	151
CSU X-band	38.89838791, -106.94324493	3147	8010	149

Appendix B: Evaluation of antenna pointing and measurement comparability

The exact north alignment of LIMRAD94 was ensured by using a solar scan routine that is integrated into the LIMRAD94 software. The sun is a source of strong microwave radiation which can be detected by the passive 89 GHz channel of LIMRAD94. That way, the azimuth value at which the sun is visible in the measurements of LIMRAD94 is determined. The true azimuth of the sun is known from the GPS module onboard LIMRAD94. The offset from the true north is then simply the difference between the two azimuth values. During the first months of CORSIPP, LIMRAD94 measurements were performed at a constant elevation (CEL) of 40° and an azimuth of 151°. The field of view during the CEL measurements is shown in Fig. 1. The azimuth is chosen so that KAZR and VISSS are in the line of sight of LIMRAD94. The azimuth angle from LIMRAD94 towards CSU X-band is roughly 149° so there is a slight directional mismatch of both radar beams during the CEL measurements of LIMRAD94. The range height indicator (RHI) scans of CSU X-band towards an azimuth of 329° overlap best with the beam of LIMRAD94 (Fig. B1). However, the beams of both radars overlap enough to compare the measurements. In order to make the RHI scans of CSU X-band and the CEL measurements of LIMRAD94 comparable, the closest range gates were selected for each time step. With increasing altitude, hence the elevation angle of CSU X-band, the horizontal distance of the closest range gates increases. The maximum horizontal distance is 120 m, the vertical distance does not exceed 8 m (Fig. B2). Coordinates and relative locations of all instruments can be found in Tab. B1.



Author contributions. AK collected and processed the LIMRAD94 data from CORSIPP, analyzed and plotted the data and wrote the manuscript with contributions from all coauthors. LT and DO provided the DDA simulations, LT wrote the section 2.4.2. AM provided the T-matrix simulations and wrote the Sect. 2.4.1. VE created Fig. 6 and wrote the caption. TV processed the *EDR* data. MM processed the VISSS data. HKL and MM acquired funding and guided the research project. AK, AM, MM, VE, AR, PB, LT, DO, HKL participated in joint meetings to discuss the interpretation of polarimetric variables observed during the presented case study. All authors reviewed and edited the manuscript. The AI tool Writefull was used to correct spelling and typesetting.

Competing interests. The authors declare that they have no competing interests.

Code and data availability. SAIL data were obtained from the Atmospheric Radiation Measurement (ARM) user facility, a U.S. Department of Energy (DOE) Office of Science user facility managed by the Biological and Environmental Research Program.: LIMRAD94 (<https://doi.org/10.5439/2229846>, last access: December 5, 2024), VISSS (<https://doi.org/10.5439/2278627>, last access: December 5, 2024), the meteorological in situ 415 data of AMF2 (<https://doi.org/10.5439/1786358>, last access: December 5, 2024), the microwave radiometer retrieval products (<https://doi.org/10.5439/1027369>, last access: 5 Dec 2024), the ARM KAZR (<https://doi.org/10.5439/1498936>, last access: 5 Dec 2024), the CSU X-band (<https://doi.org/10.5439/1888379>, last access: 5 Dec 2024).

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