

Empirical classification of dry-wet snow status in Antarctica using multi-frequency passive microwave observations

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Reviewer comment 2

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This work proposes a novel synthesis of binary melt determinations from several different passive microwave sources, as well as accounting for varying diurnal observation timings, in order to assert features about the melt state of Antarctic firn. The authors propose a classification system to relate observations from 3 different satellites, and 6 different melt detection algorithms, into a set of categories for the spatial and diurnal variation of liquid water in a firn column. These classes go beyond what has previously been derived from passive microwave observation synthesis, and this work provides a valuable advance in passive microwave observation analysis of melt. This work provides a comparison with ERA5 reanalysis skin temperatures to give confidence that the diurnality and spatial distribution of surface melt in this work at least generally reflects real world patterns, which is a reasonable approach for passive microwave melt analysis. Aside from several minor questions below, I believe that this work is suitable for publication in this journal.

We would like to thank the reviewer for the time devoted to this review. Please, find our answers in blue to the comments hereinafter.

My two biggest questions have to do with the new melt detection techniques presented in this work.

1) A new method for melt detection at 37 GHz is proposed here, and given its novelty I would like to see a little additional discussion about this method. In particular, I would like to see an explanation for the rationale of using a running mean of recent dry brightness temperature values and a 1-sigma increase to detect melt. Additionally, I am curious to see a short discussion of how well the method performs, or at least how well it agrees with other melt detection results.

Thank you for this comment. We propose to better describe the new melt detection method used for 37 GHz and split Section 3.1 into subsections dedicated to each indicator to improve clarity: 3.1.1 19 GHz-based dry-wet snow indicator, 3.1.2 1.4 GHz-based dry-wet snow indicator, 3.1.3 37 GHz-based dry-wet snow indicator, 3.1.4 Full-partial melting pixel indicator. We propose adding the following explanation in L157:

“Nonetheless, the Torinesi et al. (2003) method previously used for melt detection at 19 GHz and 1.4 GHz is inadequate.

Firstly, 37 GHz is strongly affected by snow metamorphism in the first centimeters and some large and rapid brightness temper-

ature variations may be related to change in snow grain size and density rather than melt (Brucker et al., 2011; Champollion et al., 2019). Even at Dome C where no melt occurs some rapid variations of 5-10 K can be observed during winter at 37 GHz (Brucker et al., 2011). Moreover, the large brightness temperature seasonal cycle at 37 GHz makes it difficult to use the Torinesi et al. (2003) method, which is based on the hypothesis of moderate variations of dry brightness temperature. Secondly, the seasonal melt, refreezing cycles and precipitations change the ice properties at the surface and can generate strong variations in brightness temperature at 37 GHz. A constant threshold over the April year N to March year N+1 period, as used at lower frequencies (Torinesi et al., 2003), is unadapted.

As we need to distinguish the brightness temperature variations related to rapid changes in the ice surface properties, such as grain size or density, from those related to the liquid water presence, we propose to use a running mean instead of a fixed annual threshold as in Torinesi et al. (2003). A new threshold definition was adopted: $T_{37} = M_{37} + \sigma_{37}$ where M_{37} is the 5-day moving mean timeseries of the brightness temperatures when the 19 GHz indicator is dry and σ_{37} is its standard deviation between 1 April year N to 31 March year N+1. The M_{37} timeseries is then linearly interpolated to fill the gaps when the 19 GHz indicator is wet. T_{37} is computed from the brightness temperature in vertical polarisation acquired at ascending passes and subsequently applied to both ascending and descending passes.

Figure 1 shows an example of the new threshold definition for one grid cell from September 2015 to June 2016, with which 30 wet snow days were identified. For comparison, the Torinesi et al. (2003) threshold applied to 37 GHz timeseries detects 20 melt days, and the main differences are observed from mid-January to deb-February. During this period, 37 GHz brightness temperature has strong variations of more than 40 K in one day, which could be attributed to liquid water. Moreover, the ERA daily maximum temperature around 270 K also suggest the possibility of wet snow. In general, over the 2012-2023 period, the 37 GHz dry-wet indicator computed from Torinesi et al. (2003) and this study are in agreement in 99.1% of the cases when at least one wet snow day is detected by one of the two methods. The Torinesi et al. (2003) indicator detected wet (dry) snow whereas the indicator of this study detected dry (wet) in 0.5% (0.4%) of cases.

Finally, note that this new threshold still exhibits a strong sensitivity to brightness temperature variations, leading to occasional unexpected melt detection during winter (e.g. 296 pixels in July-August on average, i.e. 0.13 % of the total wet days detected with 37 GHz). These false alarms underscore that melt detection at this frequency remains difficult and subject to uncertainties."

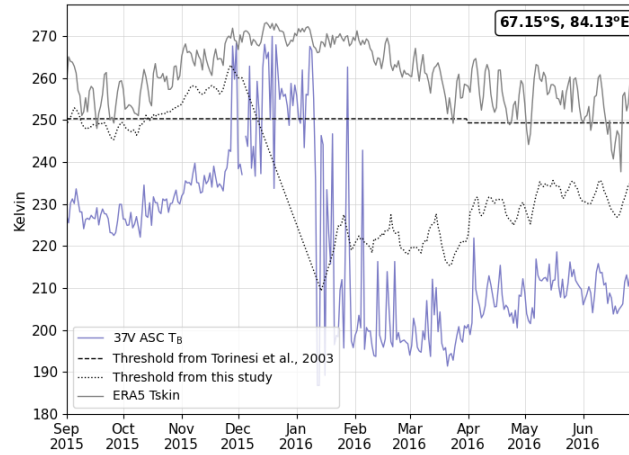


Figure 1. Brightness temperature at 37 GHz (purple) at 67.15°S, 84.13°E on West ice shelf in 2015/16 and the thresholds from Torinesi et al. (2003) (dashed) and from this study (dotted). ERA5 daily maximum skin temperature (grey).

2) A method for identifying if 80% of a 19 GHz pixel is melting is introduced. This method uses the difference between dry and wet snow brightness temperatures to produce a threshold value. I am curious, why did you use a single value for 19 GHz T_{dry} for all surfaces? I would expect the dry snow 19 GHz value to vary spatially as it is a function of temperature profile, grain size, and ice lenses in a snowpack.

We agree that the dry snow brightness temperature at 19 GHz depends on snow grain size, ice lenses and this generates spatial variations. Our hypothesis was that when snow becomes wet, brightness temperature tends to be independent of the grain size because the liquid water presence masked out the sensitivity to the other parameters and to the layers under the wet horizon.

However, we noted that using a T_{dry} spatially and periodically varying, computed as in Torinesi et al. (2003), provides $T_{80\%}$ ranges between 248 K and 270 K with a median of 261 K. Figure 2 shows the difference of the percentage of occurrences for each signature. In total, about 4% of the wet occurrences are affected by a signature change. The most impacted signature is the signature 32 which is composed by the $T_{80\%}$ at one and all the others indicator at zero. This signature is defined as invalid and excluded from the classification because the $T_{80\%}$ indicator and the 19 GHz in ascending pass indicator are in opposition. Using an adaptive $T_{80\%}$ strongly reduces the occurrence of this absurd signature (from 1.9% to 0.3%), enabling 1.6% of occurrences that would be excluded by a constant $T_{80\%}$ to be included in the classification.

We propose to use an adaptive $T_{80\%}$ to keep the consistency with the other indicators, which all use adaptive thresholds. Note that, no sharp change has been observed in the analysis performed in the article by using an adaptive $T_{80\%}$. We will add the following description in the text and update figures and numbers in the article: " M_{dry} is the mean brightness temperature of the dry snow computed between 1 April year N to 31 March year $N+1$, as described in Section 3.1.1 from Torinesi et al. (2003)."

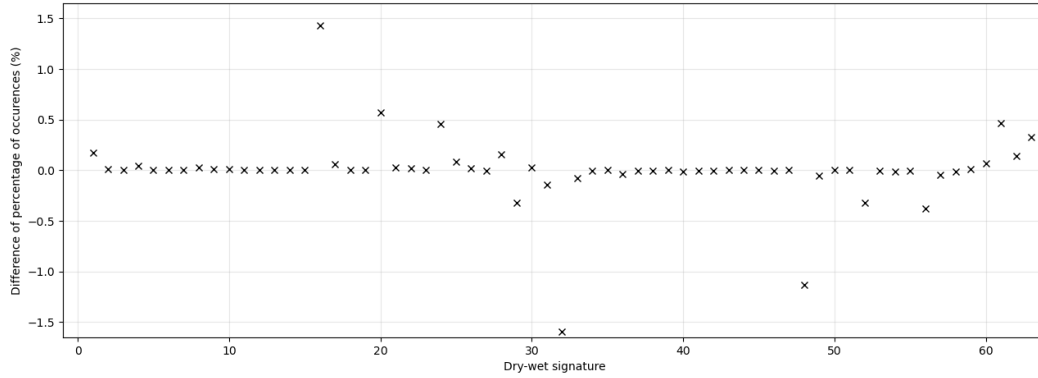


Figure 2. Difference of the percentage of occurrences for each signature.

A few specific comments by line:

Lines 141-142: You apply a bound of 20 to 35 K for the min and maximum threshold when applying the Torinesi method on
 75 19V GHz data. How often are these bounds used? If they are required often, then how sensitive is the melt detection to those bounds?

For the 19 GHz indicator, the upper (and lower) bounds are used for 0.3% (and 78%) of pixels that experienced wet snow
 for at least one day during the 12-year period. The upper bound was used very rarely, only in some marginal ice shelf areas.
 To assess the effect of the lower bound, we computed statistics using pixels where wet snow is detected at least 1 day over the
 80 2012-2023 period without applied a lower bound. Figure 3 presents the annual mean percentage of pixels for which the lower
 bound has been activated and the associated cumulative melting surface (CMS). CMS is defined as the number of wet pixels
 for an entire period over which the threshold is computed (i.e. April year N to March year N+1) and an entire region multiplied
 by the surface of a pixel (12.5 x 12.5 km). Using a lower bound of 15 K reduces by 23 % the use of this bound, and induces a
 increase in CMS of only 6%, which is relatively small. We also observed that, using a 20 K lower bound reduces by 6% the wet
 85 snow detection during winter (July-August-September period) and of 18% the wet snow detection at surface elevation higher
 than 1700 m.

Note also that theoretical analysis over the Antarctic Plateau suggested that variations of brightness temperature in vertical
 polarisation at 19 GHz lower than 20 K are probably related to snow surface metamorphism, given the 19 GHz sensitivity to
 variations in grain size (Brucker et al., 2011).

90 All of this supports the use of a lower bound of 20 K. This choice results in a conservative dry-wet snow detection that tends
 to reduce false alarm relative to undetected events.

We propose adding in the section: " *The upper (and lower) bounds are used for 0.3 % (and 78 %) of pixels that experienced
 wet snow for at least one day during the 2012-2023 period. The upper bound was used very rarely, only in some marginal ice
 shelf areas. Sensitivity analysis showed that using a lower bound of 20 K instead of 15 K reduces the detection of wet snow by*

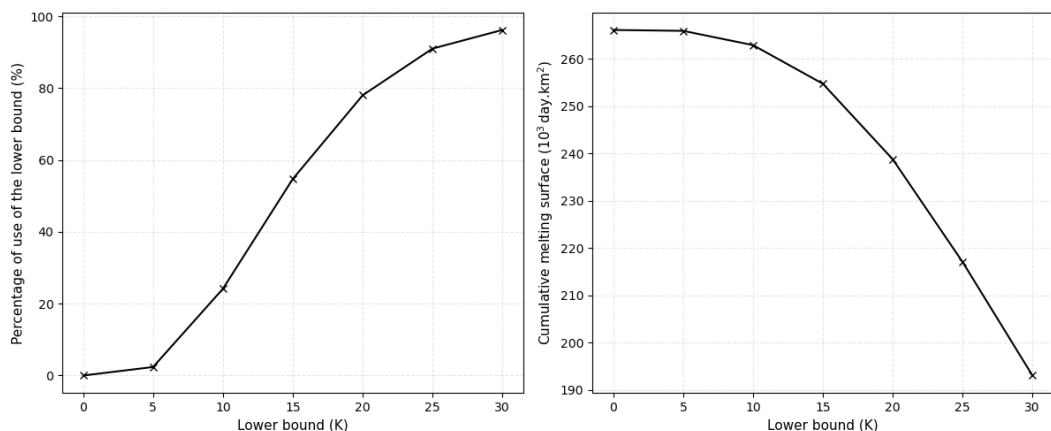


Figure 3. Annual mean of (left) the percentage of pixels for which the lower bound has been activated and (right) the associate cumulative melting surface.

95 6 % in winter (the July-September period) and by 18 % at surface elevation higher than 1700 m. This choice of a lower bound of 20 K results in a conservative dry-wet snow detection that tends to reduce false alarm relative to undetected events. "

Line 150: as with my previous comment, how often are these bounds used?

For the 1.4GHz indicator, the upper (and lower) bounds are used for 1.1% (and 96.5%) of pixels that experienced wet snow for at least one day during the 12-year period. The upper bound was used very rarely. To assess the effect of the lower bound, we computed statistics using pixels where wet snow is detected at least 1 day over the 2012-2023 period without applied a lower bound. Figure 4 presents the annual mean percentage of pixels for which the lower bound has been activated and the associated cumulative melting surface (CMS). CMS is defined as the number of wet pixels for an entire period over which the threshold is computed (i.e. April year N to March year N+1) and an entire region multiplied by the surface of a pixel (12.5 x 12.5 km). Using a lower bound of 5 K reduces by 58 % the use of this bound, and induces a increase in CMS of 45 %. We observed that this large variation is related to the fact that, using a 10 K lower bound reduces by 85 % the wet snow detection during winter(July-August-September period) and by 89 % the wet snow detection at surface elevation higher than 1700 m.

Note also that theoretical analysis over the Antarctic Plateau suggested that variations of brightness temperature in horizontal polarisation at 1.4GHz lower than 5 K are probably related to snow surface metamorphism (Brucker et al., 2014; Leduc-Leballeur et al., 2017).

110 All of this supports the use of a lower bound of 10 K. This choice results in a conservative dry-wet snow detection that tends to reduce false alarm relative to undetected events.

We propose adding in the section: " The upper (and lower) bounds are used for 1.1 % (and 96.5 %) of pixels that experienced wet snow for at least one day during the 2012-2023 period. The upper bound was used very rarely. Sensitivity analysis showed that using a lower bound of 10 K instead of 5 K reduces the detection of wet snow by 85 % in winter (the July-September

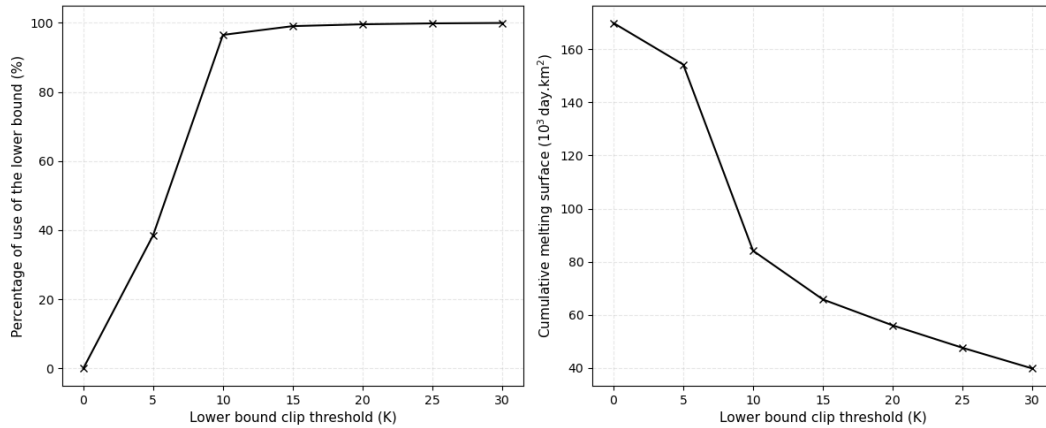


Figure 4. Annual mean of (left) the percentage of pixels for which the lower bound has been activated and (right) the associate cumulative melting surface.

115 *period) and by 89 % at surface elevation higher than 1700 m. The 10 K lower bound enables a significant reduction in false alarms. "*

Lines 150-153: Just to be clear, does “filtered out” mean that these pixels are given no result for the year, or listed as no melt all year.

We will clarify the text as: *"Moreover, pixels with a standard deviation of brightness temperature in vertical polarisation lower than 2.8 K from 1 April year N to 31 March year N+1 are marked as dry for this period (Leduc-Leballeur et al., 2020)."*

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Lines 158-159: Is the M37 value calculated from the 5 most recent days that the snow was dry at 19 GHz? The text was unclear what the algorithm does if there are more than 5 days of 19 GHz melting.

This information is indeed missing in the method description. We propose to add:

"The M_{37} timeseries is then linearly interpolated to fill the gaps when the 19 GHz indicator is wet."

125 Table 1: I am interested to see the occurrence rate of each of these classes or signatures listed somewhere, which could go in this table or Figure 1A in the appendix. Alternatively, in Figure 6 a second row could be added plotting the relative prevalence of each melt class by day of year.

For the 64 dry-wet signature, the relative occurrence rates are presented in Figure 2. For the snowpack classes, we agree to follow your suggestion, and we propose adding an inset plot in the Figure 6 of the article as in Figure 5.

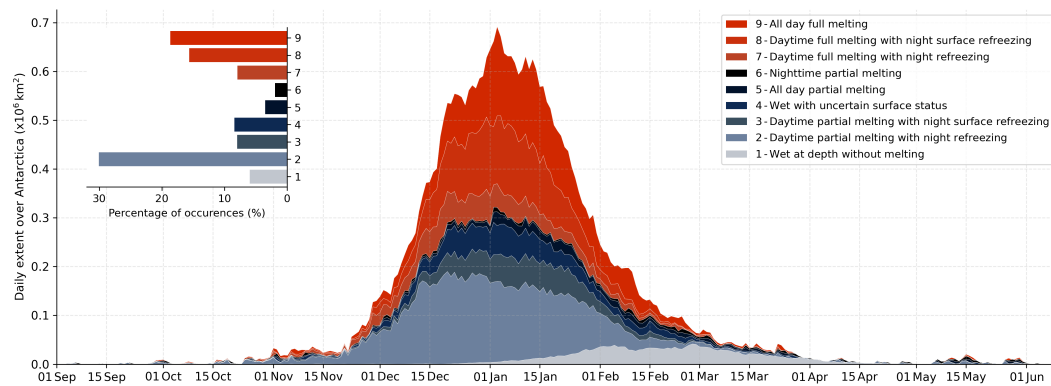


Figure 5. Daily extent of each snowpack status over Antarctica from September to June on average over 2012-2023. The percentage of occurrences of each class over this period is shown in the upper left inset plot.

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