

Impacts of the Three-dimensional Radiative Effects on Cloud Droplet Number Concentration Retrieval and Albedo Susceptibility Estimates

Deleted: Aerosol Cloud Interaction Analysis

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Abstract

Cloud droplet number concentration (N_d) in warm liquid clouds plays a crucial role in understanding cloud microphysical processes and the influence of aerosol–cloud interactions (ACI) on Earth’s climate. N_d from satellite-retrieved cloud properties such as the cloud optical thickness (τ) and cloud droplet effective radius (r_e) can be biased due to the three-dimensional (3D) radiative transfer (RT) effects. Using Large-Eddy Simulation (LES) cloud fields and RT simulations, this study investigates how biases in cloud property retrievals caused by 3D-RT effects impact the derived N_d and subsequent [albedo-susceptibility estimates](#). Our sensitivity studies confirm that the bi-spectral retrievals using the $3.7\ \mu\text{m}$ channel—whose r_e retrieval is closest to cloud top—shows better agreement with N_d from our LES models, compared to results based on the 1.6 and $2.1\ \mu\text{m}$ retrievals. At native LES resolution, N_d across all absorbing channels is strongly impacted by the 3D-effects, with the magnitude depending on the solar zenith angles (SZAs); on average, for high/low sun conditions N_d under [1D-RT overestimates /underestimate](#) its [3D-RT](#) counterpart, which indicates dominant darkening/brightening effects. At coarser satellite-like resolutions, average statistics between 1D and 3D retrievals agree better, indicating compensation between 3D and plane-parallel [averaging assumptions](#). Furthermore, [albedo susceptibility results show that 3D RT greatly modifies the local \$\alpha\$ – \$N_d\$ relationship at native LES resolution, especially under more oblique solar geometry, but the differences between 1D- and 3D-based susceptibility estimates are substantially reduced at coarser resolution and low to moderate LWP/ \$\tau\$ regimes. These results indicate that although 3D RT can strongly affect susceptibility at the LES resolution, its impact is reduced at satellite-like spatial resolution for low to moderate LWP/ \$\tau\$ regimes.](#)

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Deleted: the impact of 3D-effects on ACI analyses produced similar results across all spectral band pairings, with minimal disagreement between 1D and 3D at coarse spatial resolution. Together, these results indicate that 3D retrieval artifacts in bi-spectral N_d retrievals do not seem to drive uncertainties associated with radiative impact applications, resulting in reliable ACI and flux-related analyses.

1. Introduction

Aerosols influence Earth's radiative budget both directly by scattering and absorbing radiation and indirectly by modifying cloud properties. An increase in aerosol particles that serve as cloud condensation nuclei (CCN) in warm clouds typically leads to a decrease in cloud droplet effective radius (r_e), and an increase in the cloud droplet number concentration (N_d). This results in a more reflective cloud for a given liquid water path (LWP), known as the first aerosol indirect effect or Twomey effect (Twomey, 1974, 1977), leading to the radiative forcing from aerosol–cloud interactions (ACI), (RF_{aci}) (Bellouin et al., 2020; Forster et al., 2021). In addition to the Twomey effect, increased aerosol concentrations can cause additional cloud responses, such as changes in LWP and suppression of precipitation (Albrecht, 1989).

In most remote sensing–based studies of ACI, aerosol effects on clouds are estimated through changes in N_d . Therefore, continuous and accurate measurements of N_d at regional and global scales are essential to improve understanding of aerosol impacts on cloud microphysical and optical properties, as well as to evaluate RF_{aci} . Satellite-based remote sensing is commonly employed for such observations because alternative methods for measuring N_d are often limited or unavailable.

Operational [passive](#) satellite instruments such as the Moderate-resolution Imaging Spectroradiometer (MODIS) do not directly retrieve N_d . Instead, N_d is derived from other retrieved parameters—namely, cloud optical thickness (τ) and r_e —under assumptions about cloud adiabatic growth and the constancy of N_d throughout the cloud depth (Boers et al., 2006; Grosvenor et al., 2018; Quaas et al., 2006). Such estimates of N_d derived from passive imager observations rely on the τ and r_e retrieved via the so-called bi-spectral retrieval techniques (e.g., Nakajima and King, 1990; [Twomey and Seton, 1980](#)). The bi-spectral retrieval method

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79 simultaneously retrieves τ and r_e using cloud reflectance measurements from two spectral bands.
80 Typically, one band is selected from a non-absorbing visible or near-infrared (VNIR) spectral
81 region (e.g., wavelength centered near 0.86 μm), which is primarily sensitive to τ , while the other
82 is chosen from a moderately absorbing shortwave infrared (SWIR) (e.g., wavelength centered near
83 2.1 μm) or mid-wave infrared (MWIR) (e.g., wavelength centered near 3.7 μm) spectral region,
84 which is more sensitive to r_e .

85 While bi-spectral retrievals from passive instruments like MODIS have significantly
86 advanced our understanding of clouds, N_d derived from these retrievals may contain large errors
87 and uncertainties. These arise from assumptions made both in the retrieval process and in the
88 computation of N_d , which affect their applicability for aerosol–cloud interaction and other process
89 studies (Gryspeerdt et al., 2016; McCoy et al., 2017; Quaas et al., 2020). [All](#) operational bi-spectral
90 retrievals rely on one-dimensional (1D) radiative transfer (RT) theory, which assumes that the
91 atmosphere within each pixel is horizontally homogeneous (plane-parallel assumption) and that
92 each pixel is independent of its neighbors (independent pixel assumption), primarily for
93 computational efficiency. However, real clouds have complex 3D structures, requiring RT
94 simulations that account for both vertical and horizontal radiation transport, known as “3D RT.”
95 This is challenging because detailed 3D cloud structures are generally unknown a priori, and even
96 when available, 3D RT simulations are computationally expensive and not operationally feasible.
97 The deviation of the 3D RT in real cloud from the 1D RT theory as the backbone of operational
98 cloud retrieval algorithm is often referred to as the 3D radiative effects. These effects can introduce
99 substantial biases in cloud property retrievals and radiative quantities that are based on 1D RT
100 (Marshak et al., 2006; Várnai and Marshak, 2002; Zhang et al., 2012, 2016; [Zinner et al., 2010](#);
101 [Cornet and Davies 2008](#)). This paper focuses on investigating how errors due to 3D radiative

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103 effects associated with retrieved cloud properties (τ and r_e) impact derived N_d statistics, and how
104 these errors affect [albedo susceptibility estimates](#) and interpretations of the first aerosol indirect
105 effect. Evaluating 3D effect errors on derived N_d from LES model fields can improve constraints
106 on N_d , thereby provide better understanding of ACI and RF_{aci} .

107 From the perspective of results, 3D RT can cause brightening or darkening phenomena in
108 observed cloud reflectance compared to 1D RT. Brightening occurs when cloud reflectance in 3D
109 RT is higher than in 1D RT, while darkening occurs when clouds appear darker under 3D RT
110 relative to 1D RT. These effects can significantly impact the retrievals of τ and r_e (Kato and
111 Marshak, 2009; Marshak et al., 2006; Várnai and Marshak, 2002). For example, brightening effects
112 typically lead to overestimated τ and underestimated r_e retrievals, while the darkening effects
113 typically result in underestimated τ and overestimated r_e when compared to 1D RT retrieved results
114 (Marshak et al., 2006). [Several factors can contribute to 3D radiative effects in broken cloud fields,](#)
115 [including variability in cloud-top height, cloud horizontal and vertical heterogeneity, solar and](#)
116 [viewing geometry, and light scattering from optically thick to optically thin regions. Previous](#)
117 [studies have shown that solar and observation geometry strongly modulate how these 3D effects](#)
118 [appear in bi-spectral cloud-property retrievals, influencing both the sign and magnitude of retrieval](#)
119 [errors in \$\tau\$ and \$r_e\$ \(Kato and Marshak, 2009; Marshak et al., 2006; Zhang et al., 2012\), with these](#)
120 [errors being generally more pronounced in broken cloud fields.](#)

121 Over the past decade, several studies related to aerosol–cloud interactions have utilized N_d
122 estimates or the N_d –LWP relationship obtained from bi-spectral retrievals to study the impact of
123 aerosols on clouds, evaluate RF_{aci} , and quantify related uncertainties. For example, Grosvenor et
124 al. (2018) estimated that N_d from passive remote sensing instruments at the pixel scale have a

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Deleted: Although several factors—such as, high variability in cloud top height, solar and view geometry, and photons leaking from optically thick to optically thin cloudy regions—potentially influences the 3D radiative effects, previous studies (e.g., Kato and Marshak, 2009; Marshak et al., 2006; Zhang et al., 2012) have shown that solar and observation geometry play a significant role in influencing both the sign and magnitude of 3D effects errors in bi-spectral retrievals of cloud properties (τ and r_e), with these errors being more pronounced in broken cloud fields.

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137 relative uncertainty of 78 % (dominated by errors in r_e retrievals) for optically thick single layer
 138 stratiform clouds. Quaas et al. (2020) reviewed the challenges and uncertainties in constraining the
 139 Twomey effect from satellite observations, suggesting that past studies have likely underestimated
 140 the sensitivity of N_d to aerosol perturbations, leading to an underestimation of the Twomey effect's
 141 radiative forcing. Arola et al. (2022) used satellite and simulated data to show that natural spatial
 142 variability and errors in bi-spectral retrievals of τ and r_e propagate into N_d and LWP estimates and
 143 can cause positive LWP adjustments to be misinterpreted as negative, often leading to an
 144 underestimation of the cooling effect of ACI. Recent studies by Loveridge and Di Girolamo (2024)
 145 used a combination of synthetic cloud fields, RT simulations, and satellite observations to show
 146 that commonly used sampling strategies for N_d derived from bi-spectral retrievals, do not eliminate
 147 systematic biases caused by cloud heterogeneity, potentially leading to an overestimation of
 148 aerosol indirect effects in climate studies.

149 While previous studies have quantified uncertainties in N_d derivations, investigated cloud
 150 adjustments under varying environmental conditions, constrained the Twomey effect using
 151 satellite observations, and tested filtering strategies on retrieved N_d , few have specifically
 152 examined how errors arising from 3D radiative effects propagate into in N_d calculations and
 153 influence the estimation of the Twomey effect from satellite data (e.g., Arola et al. 2022).
 154 Understanding this linkage is crucial, as 3D radiative effects can introduce biases in cloud property
 155 retrievals that may significantly affect [albedo susceptibility estimates and](#) our interpretation of
 156 ACI.

157 The goal of this work is to build on previous studies by investigating how 3D radiative
 158 effects, which influence retrievals of τ and r_e , propagate into N_d derivations [calculated from the](#)

Deleted: Gryspeerdt et al. (2019) utilized the relationship between N_d and LWP to constrain ACI. Their findings showed a highly nonlinear N_d –LWP relationship over global oceans, suggesting that LWP increases at lower N_d regimes and decreases at higher N_d regimes.

164 [retrievals](#). We further aim to examine how these effects impact [albedo susceptibility estimates](#)
 165 derived from such retrievals. Specifically, using our model LES fields, we seek to answer the
 166 following questions: *How do the brightening and darkening effects that impact the retrieved τ_e*
 167 *and τ affect calculations of N_d ? Does the derived N_d change significantly when different*
 168 *wavelength pairs are utilized in the bi-spectral retrievals used for the N_d calculations? And how*
 169 *do the 3D effects impact these changes? Finally, how do 3D effects impact [albedo susceptibility](#)*
 170 *[estimates](#) derived from bi-spectral retrievals?*

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171 The paper's remaining structure is arranged as follows: Sect. 2 briefly describes the data and
 172 theory for the study. Results and discussions on how the 3D radiative effects influence calculations
 173 of N_d and [albedo susceptibility estimates](#) are presented in Sect. 3. The summary and conclusion
 174 are given in Sect. 4.

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175 2. Data and Theory

176 2.1 Cloud Field Dataset

177 Since clouds in the real atmosphere are always influenced to some extent by 3D radiative
 178 effects, relying solely on observational data to understand these effects and their impacts on
 179 derived N_d is challenging. To address this challenge, many studies (e.g., Ademakinwa et al., 2024;
 180 Miller et al., 2018; Rajapakshe and Zhang, 2020; Zhang et al., 2012) have utilized synthetic cloud
 181 fields and RT simulations to mimic the satellite observation-retrieval process and study the 3D
 182 radiative effects on retrieved τ and τ_e . A major advantage of this approach is that the LES cloud
 183 field provides the model “truth” which is difficult to obtain in real world observations. Also, the
 184 flexibility of this approach allows investigation of various cloud mechanisms at different levels of
 185 complexity. For example, we can analyze the 3D RT impacts by comparing retrievals from 3D RT
 186

193 with those from 1D RT. Disadvantages of this approach are that the usefulness of such a study
194 depends on the realism of the synthetic microphysical and water content fields; which, like the
195 satellite retrievals themselves are difficult to validate, as well as the limited variability in scene
196 properties that can be run and processed compared to global satellite sampling.

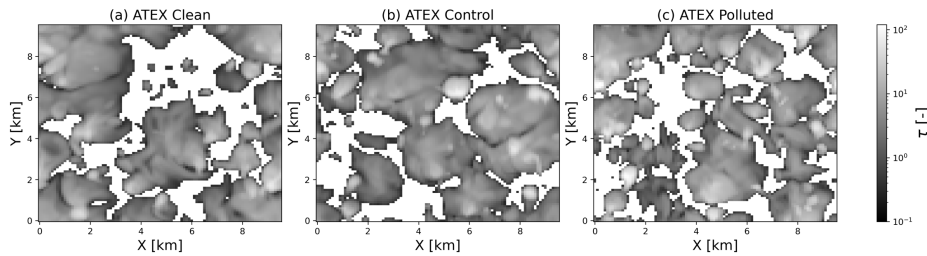
197 This study utilizes a similar state-of-the-art satellite retrieval simulator as in Zhang et al.
198 (2012) and Ademakinwa et al. (2024) that applies 1D and 3D RT simulations to synthetic cloud
199 fields. In this work, our cloud fields are Large Eddy Simulation (LES) models (Distributed
200 Hydrodynamic-Aerosol-Radiation Modeling Application (DHARMA)) with bin microphysics
201 (Ackerman et al., 2004; Miller et al., 2016; Zhang et al., 2012). They consist of different initial
202 aerosol loadings derived from an idealized case study by Stevens et al. (2001) conducted during
203 the Atlantic Trade Wind Experiment (ATEX). Three LES cases of different initial aerosol loadings
204 are examined throughout this study. The first case has a CCN loading of 40 cm^{-3} (hereafter
205 referred to as “ATEX clean”), the second case has a CCN loading of 75 cm^{-3} (hereafter referred
206 to as “ATEX control”), while the third case has a CCN loading of 600 cm^{-3} (hereafter referred to
207 as “ATEX polluted”). The LES provides a model “truth” for the 3D cloud microphysical
208 properties, which are used as a baseline for comparisons with numerically simulated retrievals.
209 The droplet sizes in the LES consist of 25 bins ($0.874 \mu\text{m} \leq \text{radius} \leq 281.76 \mu\text{m}$) used to represent
210 the distribution of droplet sizes (Ackerman et al., 1995) and Mie scattering properties are bulk
211 averaged over a highly resolved flat sub-bin to obtain the optical properties for each size bin. This
212 serves as input for RT simulations based on the size distributions of the LES cloud fields. The LES
213 cases have a domain size of $9.6 \times 9.6 \times 3 \text{ km}$ ($x \times y \times z$), with a horizontal spatial resolution of
214 $\Delta x = \Delta y = 100 \text{ m}$ and a constant vertical grid spacing of $\Delta z = 40 \text{ m}$. These grid resolutions were
215 selected such that individual cumulus clouds are appropriately resolved; however, it is expected to

Deleted: These LES are representative of a trade wind cumulus regime, characterized by scattered cumulus clouds rising into a thin, broken stratocumulus layer.

219 exhibit vertical artifacts near the inversion due to unresolved entrainment processes at these
 220 resolutions (Stevens et al., 2002). Additional information about the model setup for these LES
 221 cases can be found in Fridlind and Ackerman (2011) and Zhang et al. (2012). For each LES case,
 222 snapshots of cloud microphysical properties are taken every half hour after the first 4 hours of each
 223 simulation, resulting in 9 cloud scenes per LES case and a total of 24 cloud fields.

224 The LES cloud fields comprise of spatially inhomogeneous microphysical properties, and
 225 these cloud properties change as the LES cloud field evolves at each time step. Figure 1 provides
 226 a map of the τ derived from the LES droplet size distributions for the ATEX clean, control and
 227 polluted cases at 4.0 hours of simulation time. The map shows that the scenes are typically
 228 characterized by broken clouds with cloud fraction (CF; defined as a fraction of columns with $\tau >$
 229 0.3) greater than 70 % in each LES case (CF of 73.7, 77.2 and 73.3 % for the ATEX clean, control
 230 and polluted cases respectively).

231



232

233 **Figure 1.** Map of the Large Eddy Simulation (LES) cloud optical thickness (τ) for the (a) ATEX clean, (b) ATEX
 234 control, and (c) ATEX polluted case at the 4 h simulation time.

235 2.2. Radiative Transfer Setup

236 The spherical harmonics discrete ordinate method (SHDOM) RT model developed by
 237 Evans (1998) was utilized to model reflectance for both 3D and 1D RT at the native LES resolution

238 of 100 m. The RT simulations were performed for 6 solar zenith angles (SZAs) spanning from 10°
239 to 60° in increments of 10°. A fixed nadir viewing zenith angle (VZA), and a constant relative
240 azimuthal angle ($\Delta\Phi=30^\circ$) were used throughout the study. The simulations for cloud retrievals
241 are performed based on MODIS spectral response function for MODIS Band 2 (0.86 μm), Band 6
242 (1.6 μm), Band 7 (2.1 μm), and Band 20 (3.7 μm). We hereafter refer to these bands by their
243 respective nominal wavelengths. It is important to note that although the 3.7 μm channel normally
244 comprise of both solar and thermal portion and recent study by Loveridge and Di Girolamo (2024)
245 suggest that contribution of the thermal emission to the error budget in the 3.7 μm retrieved r_e can
246 sometimes be significant—partly due to their analysis utilizing partly cloudy pixels. This work
247 assumes that the thermal emission in the 3.7 μm band can be removed with no error and thus have
248 used only the solar portion in our RT calculations. For all RT simulations in this study, the surface
249 was assumed to be Lambertian with an albedo of 5 % (an assumption approximating ocean surfaces
250 under diffuse illumination), and double periodic horizontal boundary conditions were applied.

251 Observational studies usually estimate cloud susceptibility from derived N_d and broadband
252 flux observations from Clouds and the Earth's Radiant Energy System (CERES) (e.g Painemal and
253 Minnis 2012; Platnick and Oreopoulos 2008). The CERES TOA flux is obtained from directional
254 radiance measurements using scene-dependent anisotropy factors (often described via angular
255 distribution models, ADMs), and therefore the CERES radiance to flux procedure explicitly
256 accounts for the angular redistribution of radiation. In this study, rather than approximating fluxes
257 from a single viewing direction or adopting an ADM assumption at LES scales, we compute the
258 shortwave flux and albedo directly from radiative transfer by hemispherically integrating the
259 upwelling radiance (I) at the top of the RT domain over a discrete angular grid. The integration is
260 performed using radiance simulations at 37 viewing zenith angles (θ_v) spanning 0–90° in 2.5°

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increments and 16 viewing azimuth angles (ϕ_v) spanning 0–360° in 22.5° increments. The upwelling flux ($F^\uparrow$) at solar zenith angle θ_o is defined as:

$$F^\uparrow(\theta_o) = \int_0^{2\pi} \int_0^{\pi/2} I(\theta_v, \phi; \theta_o) \cos \theta_v \sin \theta_v d\theta_v d\phi, \quad (1)$$

where ϕ is the azimuth integration variable, $\phi = \phi_v - \phi_o$ is the relative azimuth angle, with ϕ_o and ϕ_v denoting the viewing and solar azimuth angles, respectively.

To ensure accurate flux and albedo estimates from the discrete angular sampling, the hemispheric integration was evaluated using finite-volume angular quadrature, in which each viewing-zenith bin was weighted by the exact integral of $\cos \theta_v \sin \theta_v$ over its bin edges and each azimuth bin was weighted by its corresponding $\Delta\phi$. The resulting hemispherically integrated upwelling flux was then normalized by the incident solar flux at the top of the domain to obtain the broadband albedo (α), defined as:

$$\alpha(\theta_o) = \frac{F^\uparrow(\theta_o)}{F_o \mu_o} \quad (2)$$

where F_o is the incident solar flux at domain top and $\mu_o = \cos \theta_o$. The broadband shortwave (0.3 – 5 μm) radiances used in the hemispheric integration are calculated for 13 out of the 14 Rapid Radiative Transfer Model (RRTM) spectral bands in the shortwave. Notably, we have excluded RRTM band 28 (0.2 - 0.26 μm) from these calculations because our Mie property computations failed for large droplet size bins (large size parameter). This exclusion is unlikely to impact the

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316 results significantly because this band contributes very little to the solar shortwave radiative energy
317 budget. [Unless otherwise stated, all albedo and susceptibility results presented in this study are](#)
318 [based on these flux-derived broadband quantities computed consistently for both 1D and 3D](#)
319 [forward RT simulations.](#)

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321 2.3. Bi-Spectral Retrieval Method

322 The bi-spectral retrieval method ([Nakajima and King 1990](#)) described in Sect. 1 was
323 applied to the simulated reflectance (see Sect. 2.3 in Ademakinwa et al., 2024). This method relies
324 exclusively on homogeneous 1D RT assumptions to interpret the observed cloud reflectance. Its
325 implementation is done using a precomputed lookup table (LUT), which includes computed 1D
326 spectral reflectance for different τ and r_e combinations and solar-view geometries. The LUT then
327 is used to find the τ and r_e whose computed reflectance best matches the observed (here, modeled
328 LES) cloud reflectance. Notably, for small τ , retrieval uncertainty increases because the isolines
329 of the LUT are less orthogonal and more tightly packed. This LUT non-orthogonality has
330 significant consequences for observations with high spatial inhomogeneity below the pixel level
331 resolution (Zhang et al., 2012, 2016) leading to the plane parallel homogeneous approximation
332 (PPHA) bias. The consequences of both unresolved inhomogeneity and 3D radiative effects
333 changes with spatial resolution and are examined in Sect. 3.3.

334 The VNIR reflectances used for the bi-spectral retrievals were calculated for 0.86 μm ,
335 while the SWIR and MWIR reflectances were calculate for the 1.6, 2.1 and 3.7 μm channel. We
336 use a simplified cloud masking method that makes use of the 0.86 μm reflectance (i.e., $R(0.86$
337 $\mu\text{m}) > 0.07$) to identify cloudy pixels after radiative transfer simulations. The LUTs for this study

338 have 73 effective radii spanning from 4 to 40 μm and 105 log-spaced τ values spanning from 0.1
339 to 150. A constant effective variance (v_e) value of 0.1 in the gamma representation of the droplet
340 size distribution is applied, consistent with operational MODIS retrievals. For consistency, the
341 same surface albedo of 5 % used in all our LES RT reflectance simulations is applied in the LUT
342 reflectance simulations.

343 We implement retrievals using reflectance simulations at the native LES resolution (100
344 m) and spatially average to a coarser MODIS-like resolution. For the MODIS-like resolution
345 retrievals, we averaged reflectance from the LES resolution of 100 m to 800 m resolution and
346 utilized this area-averaged reflectance as input into the LUT for retrievals. Note that we have used
347 800 m resolution instead of the MODIS nadir 1 km because it divides evenly into our LES domain.
348 Also, sensitivity analyses (not presented) suggest minimal differences between τ and r_e retrievals
349 at 800 m and 1 km resolutions.

350 When averaging from high resolution to coarse resolution, we distinguish between overcast
351 and partially cloudy pixels, depending on whether or not clear-sky pixels are included in the
352 average reflectance. Firstly, if one or more pixels used for the area average is a clear-sky pixel, we
353 classify the resulting pixel as a “partially cloudy pixel”. Secondly, if all pixels utilized for the area
354 average are cloudy, we classify the averaged resulting pixel as an “overcast cloudy pixel”. With
355 this, we define a category “all cloudy pixels” to consist of both partially cloudy and overcast cloudy
356 pixels.

357 Cloud property retrievals that utilize area-averaged reflectances are impacted by
358 unresolved (i.e., sub-pixel) spatial cloud inhomogeneity that can bias retrieval results, especially
359 if the average is done over a highly inhomogeneous τ region. For example, Zhang et al. (2012)

demonstrated using RT simulations how the nonlinearity in the LUT space (non-orthogonality of the τ and r_e LUT grid) can lead to plane-parallel albedo bias (Cahalan et al., 1994; the retrieved τ from the average reflectance of inhomogeneous cloud pixels tends to be smaller than the average of the sub-pixel τ) and plane-parallel r_e bias (r_e retrieved from area-averaged reflectances over inhomogeneous τ region is overestimated compared the original r_e).

2.4. N_d and LWP Estimates

2.4.1. N_d and LWP from LES

Cloud vertical structures in realistic cloud fields, such as the LES cases considered in this study, have microphysical properties (e.g., N_d and LWC) that vary vertically within each profile. Therefore, describing a single representative microphysical property value as a reference in the LES, requires accounting for the vertical distribution of the profile and on the application of interest. For example, most ACI and in situ studies typically take the reference N_d as an average value of N_d across the cloud vertical extent from cloud base to cloud top, or cloud-base N_d (e.g., Gryspeerd et al., 2022; Painemal and Zuidema, 2011). In contrast, we adopt a different definition of N_d due to our application of interest: this study evaluates 3D effects biases in passive, radiance-based bi-spectral retrievals and examines how these biases propagate into N_d and albedo susceptibility estimates. Because passive retrievals and reflected SW radiative flux are most sensitive to the optically active portions of the cloud, a simple vertical mean does not necessarily provide the most radiatively relevant LES reference. We therefore define a radiation-relevant reference droplet number concentration in the LES as an extinction-weighted vertical mean, denoted N_{d_LES} . This is achieved by weighting N_d over the extinction coefficient (β_{ext}) at each layer (z) as given by:

Deleted: for in situ studies is typically taken

Deleted: , from the cloud base height (CBH) to the cloud top height (CTH)

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$$N_{d_LES} = \frac{1}{\tau} \int_{CTH}^{CBH} N_d(z) \beta_{ext}(z) dz, \quad (3)$$

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391 where $\tau = \int_{CTH}^{CBH} (\beta_{ext}) dz$. With this definition, layers with higher β_{ext} contribute more to the
 392 weighted value because they play a larger role in how the cloud interacts with scattered radiation,
 393 while reducing the influence of optically tenuous layers near cloud boundaries. We emphasize that
 394 N_{d_LES} is not intended to replace the vertically averaged or cloud-base N_d definitions commonly
 395 used in in situ and field-validation studies. Rather, it is used here as a radiatively weighted LES
 396 reference that is more directly aligned with the passive retrievals and broadband albedo analyzed
 397 in this study. Therefore, agreement between retrieval-derived N_d and N_{d_LES} should be interpreted
 398 as agreement with a radiatively effective cloud-column quantity, not necessarily as agreement with
 399 the vertically averaged or cloud-base N_d . This choice limits the theoretical interpretation of the
 400 retrieved N_d as a purely microphysical quantity, but it provides a more appropriate reference for
 401 evaluating how retrieval errors propagate into albedo susceptibility estimates. Also, the column
 402 LWP from the LES (LWP_{LES}) is obtained by integrating vertically the cloud liquid water content
 403 (LWC) from CBH to CTH in each column, expressed as:

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$$LWP_{LES} = \int_{CTH}^{CBH} LWC(z) dz. \quad (4)$$

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404 2.4.2. Satellite-Based N_d and LWP Retrievals

405 Following previous studies, we calculate N_d from bi-spectral retrievals of τ and r_e
 406 following a pseudo-adiabatic model which assumes that (i) the LWC increases linearly (as a fixed

fraction of its adiabatic value) with the cloud geometrical height and that (ii) N_d is constant vertically. Here, N_d is [calculated](#) from τ and r_e (denoted as N_{d_cal}) according to Grosvenor et al. (2018) [which consolidates on prior effort by several previous studies \(e.g., Brenguier et al. 2000; Quaas et al. 2006; Boers et al. 2006\). Thus, we follow Grosvenor et al. \(2018\) simplified equation to calculate \$N_d\$ from satellite-based retrievals given by:](#)

$$N_{d_cal} = \frac{\sqrt{5}}{2k\pi} \left(\frac{f_{ad} c_w \tau}{Q_{ext} \rho_w r_e^5} \right)^{1/2} \quad (5)$$

where f_{ad} is the pseudo-adiabatic factor (constrained between 0 and 1), c_w is the effective condensation rate ($g\ m^{-4}$), Q_{ext} is the bulk extinction efficiency derived from Mie theory (approximated as 2 in the geometric scattering limit), and ρ_w is the density of liquid water (taken as $1\ g\ cm^{-3}$) and the parameter k is defined as:

$$k = \left(\frac{r_v}{r_e} \right)^3 \quad (6)$$

where r_v is the droplet mean volume radius. Several values of k have been applied in N_d related studies and in operational remote-sensing that derive N_d from τ and r_e retrieved by the bi-spectral method. For example, the MODIS N_d calculation assume a constant value of $k = 0.72$ (Grosvenor et al., 2018). Brenguier et al. (2011) utilized in situ measurements collected during five distinct field experiments to show that k values vary between 0.7 and 0.9, with uncertainties ranging from 10 % to 14 % across different cloud types and atmospheric conditions. All these suggests different values of k depending on the cloud regime. Studies based on in situ measurements have shown that k can vary significantly (Martin et al., 1994) and may also change under particularly extreme

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Deleted: Lloveridge and Di Girolamo (2024) applied a k value obtained from domain-average of their stochastically generated cloud fields in deriving N_d .

aerosol conditions (Noone et al., 2000). Thus, if k is not properly represented, it can introduce significant uncertainty into the calculation of N_d . To avoid these unconstrained constants driving our understanding of retrieval behavior, throughout this work we have calculated both k , and the quasi-adiabatic lapse rate ($f_{ad}c_w$) in each cloudy column directly from the LES cloud field.

Several studies that derive N_d from bi-spectral retrieved cloud properties commonly utilize filtering techniques based on several criteria to minimize uncertainties in N_d estimates, which arise from errors in retrieving r_e and τ from the bi-spectral method, especially in highly variable cloud fields (Grosvenor et al., 2018; Gryspeerdt et al., 2022; Loveridge and Di Girolamo, 2024; Quaas et al., 2006). However, this filtering methodology can unintentionally exclude certain types of clouds, limiting the analysis to specific cloud regimes. To avoid this bias, and to effectively quantify all possible sources of uncertainty, we incorporate all available N_d data without applying any pre-filtering method beyond the cloud mask. Furthermore, prior filtering schemes (e.g., Gryspeerdt et al., 2022; Zhu et al., 2018) implemented for coarser-resolution satellite observations may not be appropriate for the native LES resolution (100 m). For analysis involving LWP, we obtain the derived LWP (LWP_{cal}) from the retrieved τ and r_e by the adiabatic relationship given as (Wood and Hartmann 2006):

$$LWP_{cal} = \frac{5}{9} \rho_w r_e \tau \quad (7)$$

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3. Results and discussion

3.1. Comparison of LES N_d and N_d Derived from Cloud Properties Retrieved from 1D RT Simulated Reflectance

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We first perform a RT closure comparison between the reference N_d obtained directly from the LES ($N_{d,LES}$) and the N_d derived from Eq. (5) using τ and r_e retrieved from homogeneous

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plane parallel RT-based simulated reflectance (i.e., 1D RT) from the 0.86 μm channel paired with an absorbing wavelength channel λ_{abs} , ($N_{d_cal}^{1D}(\lambda_{abs})$). This closure comparison allows us to (i) check the methodology of our retrieval algorithm and (ii) investigate N_d errors associated with adiabatic cloud model assumptions. Figure 2 shows these comparisons as a regression plot of $N_{d_cal}^{1D}(\lambda_{abs})$ vs. N_{d_LES} for the ATEX clean, control and polluted cases at 4.0 hours of simulation time, when λ_{abs} is 1.6, 2.1 and 3.7 μm . The comparison shows a general good agreement between $N_{d_cal}^{1D}(\lambda_{abs})$ and N_{d_LES} for the ATEX clean, ATEX control and to some extent the ATEX polluted case with correlation coefficient (R) of 0.847, 0.852 and 0.748, respectively, when λ_{abs} is 1.6 μm ; 0.885, 0.867 and 0.748 when λ_{abs} is 2.1 μm ; and 0.859, 0.856 and 0.790 when λ_{abs} is 3.7 μm . The disparities observed in this comparison. [The disparities observed between \$N_{d_cal}^{1D}\(\lambda_{abs}\)\$ and \$N_{d_LES}\$ reflect the combined influence of vertical microphysical inhomogeneity, extinction weighting, and the assumptions embedded in the adiabatic \$N_d\$ retrieval equation \(Equation 5\). Therefore, agreement with \$N_{d_LES}\$ indicates consistency with the optically dominant portions of the cloud column sampled by the passive retrieval.](#)

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Deleted: indicate that deviations from adiabatic conditions are common in the LES simulations, which can bias the derived N_d . Furthermore, the ATEX polluted case appears to deviate the most from the adiabatic assumption. ¶

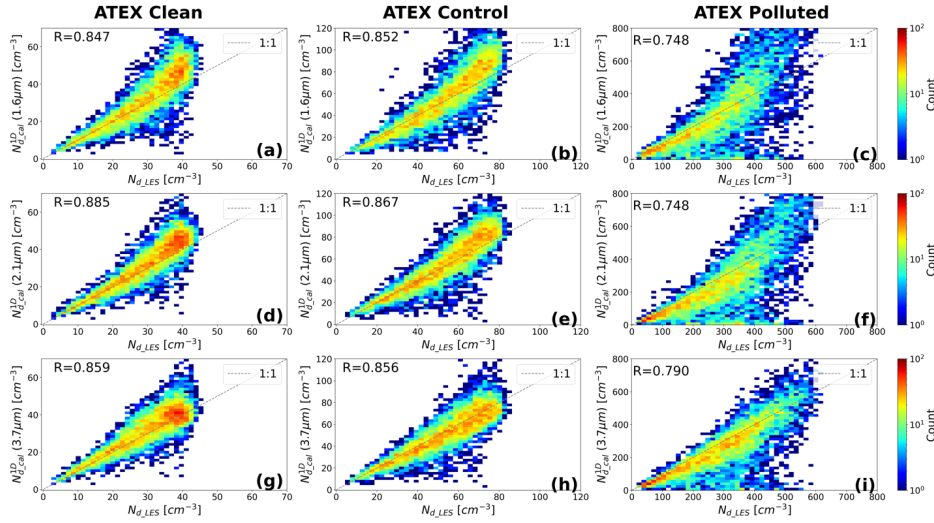


Figure 2. Comparison of N_d derived from bi-spectral retrievals based on 1D RT reflectance ($N_{d_cal}^{1D}(\lambda_{abs})$) versus N_d obtained from the LES (N_{d_LES}) for (a) ATEX clean, (b) ATEX control and (c) ATEX polluted for retrievals at absorbing channel, (λ_{abs}) of $1.6 \mu m$, for (d) ATEX clean, (e) ATEX control and (f) ATEX polluted for retrievals at λ_{abs} of $2.1 \mu m$ and for (g) ATEX clean, (h) ATEX control and (i) ATEX polluted for retrievals at λ_{abs} of $3.7 \mu m$ at 4 h of simulation time. Note: 1D RT and bi-spectral retrievals utilized for this comparison were carried out at SZA 50° . Dotted lines indicate 1:1 relationship between the derived and LES N_d .

Establishing the underlying physical processes influencing the N_d estimates obtained from the LES and retrievals are challenging when relying solely on direct one-to-one comparison. Instead, we evaluate the probability density functions (PDFs) of N_{d_LES} against those obtained from the derived N_d . For the PDF comparison, data across all time steps within each LES case are combined to ensure robust statistical analysis. Apart from evaluating the PDFs of N_{d_LES} and $N_{d_cal}^{1D}(\lambda_{abs})$, we also investigate errors associated with the adiabatic cloud model assumptions in Eq. (5). This is done by introducing constraints on the retrieved r_e and τ used to compute N_d , which increasingly force conformity with the level of adiabaticity in the LES itself. First, we utilize the vertically weighted (VW) r_e (denoted as $r_e(VW)$), computed using the two-way-transmittance-

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Deleted: Although some deviations occur in the 1D RT closure comparison, they still provide greater confidence in our derived N_d and in our definition of the reference N_d (for reference, Appendix A1 shows comparisons of N_d from 1D RT and N_d from the LES when defined as a vertical column average without extinction weighting).

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513 [weighted effective radius](#) relationship [derived from the optically weighted droplet size distribution](#)
514 defined by Miller et al. (2016) (check their Eq. (14)), and optical thickness (τ_{tot}) from the LES
515 into Eq. (5) to derive N_d (denoted as $N_{d_cal}^{r_e^{(VW)}, \tau_{tot}}$). Second, we use droplet effective radius of the
516 layer where the LWC is maximum as a representative of the adiabatic cloud top r_e (denoted as r_e^*)
517 and τ_{tot} as inputs in Eq. (5) to derive N_d (denoted as $N_{d_cal}^{r_e^*, \tau_{tot}}$). The idea of testing with r_e^* is
518 because it is representative of the droplet radius profile that follows the adiabatic assumption of
519 the retrieval — absent significant entrainment modification at cloud top or model resolution
520 artifacts. PDFs of N_d obtained from the LES, those inferred from the retrievals [obtained from](#) 1D
521 RT [simulated reflectance](#) (at SZA 50°) when λ_{abs} is 1.6, 2.1 and 3.7 μm , and those derived from
522 retrievals [when the optical thickness and droplet effective radius are constrained as](#) earlier
523 discussed, for the ATEX clean, ATEX control, and ATEX polluted cases are presented in Figure
524 3. In the ATEX clean and control cases (Figure 3a and b), the N_{d_LES} distributions peak around
525 their respective aerosol loading values (i.e., peak around 40 cm^{-3} and 75 cm^{-3} for the ATEX
526 clean and control cases respectively). This indicates that most of the available CCN are activated
527 into cloud droplets. On the other hand, the lower values of the N_{d_LES} observed in the distribution
528 are attributed to increased collision-coalescence as well as N_d removal processes driven by
529 entrainment and precipitation.

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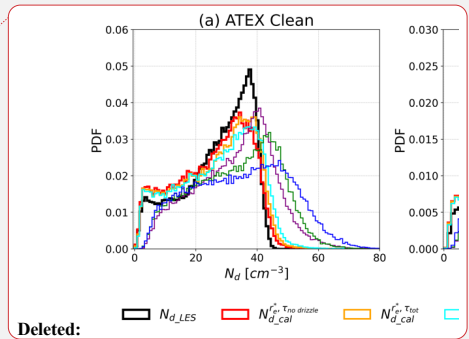
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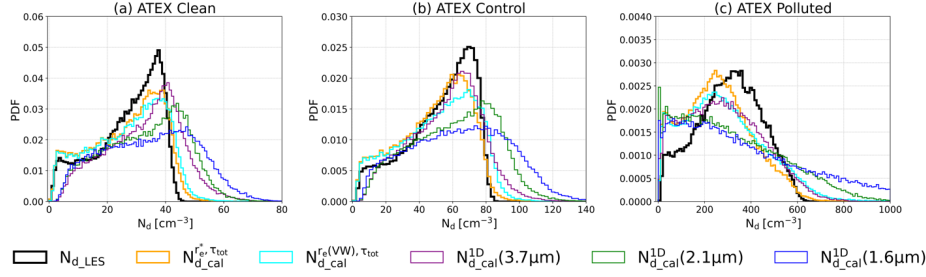


Figure 3. Probability Density Functions (PDFs) of N_d obtained from the LES (N_{d_LES}) and N_d calculated from the adiabatic equation using 1D RT derived cloud properties (for absorbing channels $1.6\ \mu m$, $2.1\ \mu m$ and $3.7\ \mu m$), and under different levels of controls across all time steps for the ATEX clean, control and polluted cases in (a), (b) and (c) respectively. ▼

Interestingly, the PDF distribution of N_{d_LES} for the ATEX polluted case (black line in Figure 3c) is somewhat different from those of the clean and control cases. Its PDF has a peak around $300\ \text{to}\ 350\ cm^{-3}$, which is significantly less than the initial aerosol loading value (CCN $600\ cm^{-3}$). This is probably because not all the CCNs are activated in the ATEX polluted case. When we consider the PDF of the $N_{d_cal}^{1D}(\lambda_{abs})$ (for the three absorbing channels considered in this study) for the ATEX clean and control cases, similar cloud processes to the N_{d_LES} are observed, but with PDFs having a broader range and higher N_d peak values, and notably long tails towards larger values, indicating overestimation. The N_d overestimation statistics is more pronounced in $N_{d_cal}^{1D}(1.6\ \mu m)$, reduces for $N_{d_cal}^{1D}(2.1\ \mu m)$ and is smallest for $N_{d_cal}^{1D}(3.7\ \mu m)$. This suggests that $N_{d_cal}^{1D}(3.7\ \mu m)$ agrees the most with N_{d_LES} compared to results from the other two absorbing channels. When we consider the $N_{d_cal}^{1D}(\lambda_{abs})$ PDFs from the ATEX polluted case (blue, green and purple lines in Figure 3c), the peaks for all three absorbing channels are at smaller values than that of the N_{d_LES} , although all have long tails towards larger values indicating significant overestimation remains. For the ATEX polluted case, the $N_{d_cal}^{1D}(3.7\ \mu m)$ PDF still has the smallest overestimation compared to the $N_{d_cal}^{1D}(1.6\ \mu m)$ and $N_{d_cal}^{1D}(2.1\ \mu m)$ results.

Deleted: Note the $N_{d_cal}^{r_e, \tau_{tot}} \text{ no drizzle}$ PDF is being covered by the $N_{d_ret}^{r_e, \tau_{tot}}$ PDF in the ATEX control and polluted cases because drizzle is significantly minimal and not driving any vertical profile changes. ¶

570 In general, the disparity in $N_{d_cal}^{1D}(\lambda_{abs})$ obtained for all λ_{abs} (1.6, 2.1 and 3.7 μm) is due
 571 primarily to differences in the retrieved r_e —in vertically inhomogeneous clouds, satellite-based
 572 r_e retrievals via the bi-spectral method can vary for different channels due to spectrally-varying
 573 liquid water absorption (Meyer et al., 2025) leading to different penetration depths within a cloud
 574 (Platnick, 2000). The 3.7 μm being the most absorptive and 1.6 μm being the least absorptive of
 575 the three channels. For ideal clouds, the 3.7 μm channel is most absorptive and yield r_e retrievals
 576 closer to the cloud top compared to the less absorbing 2.1 μm channel, and the 1.6 μm channel is
 577 the least absorptive and yields retrievals deepest into the cloud. Since the N_d equation (Eq. (5)) is
 578 most sensitive to r_e (r_e has the largest exponent in Eq. (5); $N_{d_cal} \propto r_e^{-5/2}$), small changes in r_e
 579 will greatly impact the derived N_d . Thus, the reasonable agreement observed between the PDF of
 580 $N_{d_cal}^{1D}(3.7\mu m)$ and N_{d_LES} is not surprising since Eq. (5) assumes r_e is at cloud top. This is
 581 corroborated by previous studies (e.g., Zhang and Platnick, 2011) that indicate that in adiabatic
 582 clouds, the 3.7 μm r_e retrieval is expected to be larger than retrievals from shorter wavelengths
 583 (and 2.1 μm r_e retrieval > 1.6 μm r_e retrieval), although this relationship can be affected by other
 584 factors including retrieval biases.

585 In general, the $N_{d_cal}^{1D}(\lambda_{abs})$ results in Figure 3 from the three absorbing channels indicate
 586 that, even when the RT makes the same plane-parallel assumptions as the retrievals, there are
 587 notable variabilities in the derived $N_{d_cal}^{1D}(\lambda_{abs})$ when compared to N_{d_LES} . These variabilities
 588 could be due to various reasons, which include inconsistencies between the vertical distribution of
 589 cloud microphysics in the LES truth and the assumptions that form the basis of the N_d equation
 590 utilized for the retrievals. Recall, the N_d derived from Eq. (5) assumes that the cloud is adiabatic
 591 i.e., LWC increases monotonically with height from cloud base to cloud top and N_d is constant

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vertically. But not all columns in the LES are adiabatic. Various processes such as entrainment, coalescence, etc. which are prevalent in natural clouds (and evident in this study LES cloud fields) introduce sub- or super-adiabatic behaviors in cloud vertical profiles, which impacts derived N_d .

When we consider the PDF of calculated N_d when τ and r_e are constrained to various degrees using values derived directly from the LES cloud field, the $N_{d_cal}^{r_e(VW),\tau_{tot}}$ PDF shows substantial deviations from the PDF of N_{d_LES} . A reason for this is that $r_e(VW)$ represents a weighted vertical average which is highly sensitive to the cloud microphysics and vertical structure of r_e , especially if the optical extinction in the cloud entrainment region is large enough to impact the $r_e(VW)$. The agreement between the PDF's of N_d derived under the τ and r_e constraints (i.e., $N_{d_cal}^{r_e^*,\tau_{tot}}$ and $N_{d_cal}^{r_e(VW),\tau_{tot}}$) and N_{d_LES} reduces as the adiabatic realism of the input properties decreases; $N_{d_cal}^{r_e^*,\tau_{tot}}$ showing better agreement with N_{d_LES} than $N_{d_cal}^{r_e(VW),\tau_{tot}}$ and N_{d_LES} comparison because $N_{d_cal}^{r_e^*,\tau_{tot}}$ utilizes r_e^* which is most representative of an adiabatic cloud top r_e value which is a key requirement for formulation of Equation 5. A sensitivity test was also performed where N_d is calculated using r_e^* and cloud optical depth computed from the LES that excludes cloud optical depth contributions from drizzle drop sizes (r) by isolating $r > 30 \mu m$ from the droplet size distributions. The resulting N_{d_cal} PDFs were mostly unchanged from the PDFs of $N_{d_cal}^{r_e^*,\tau_{tot}}$ (not shown). This was expected because drizzle-size drops typically occur in small numbers and contribute weakly to N_d , although their radiative impact could become important in cases with substantial drizzle populations.

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3.2. Impact of 3D Radiative Effects on Retrieval-Derived N_d at the Native LES Resolution

To investigate the 3D radiative effects impacts on the retrieval-derived N_d , we compare $N_{d_cal}^{1D}(\lambda_{abs})$ with N_d derived from Eq. (5) using τ and r_e retrieved from 3D RT-based simulated reflectance at λ_{abs} ($N_{d_cal}^{3D}(\lambda_{abs})$). We carry out this comparison at different SZAs, ranging from high to low sun positions. Figure 4 shows the PDF of $N_{d_cal}^{1D}(\lambda_{abs})$ and $N_{d_cal}^{3D}(\lambda_{abs})$ derived at the LES resolution (100 m) for a high sun (SZA 10°), moderate sun (SZA 30°) and a more oblique sun position (SZA 50°), as well as the reference $N_d(N_{d_LES})$ across all time steps (combined) in each of the ATEX clean, control and polluted LES cases. Here, N_{d_LES} PDF provides the LES-based reference distribution, so shifts or broadening in the $N_{d_cal}^{1D}(\lambda_{abs})$ and $N_{d_cal}^{3D}(\lambda_{abs})$ PDFs relative to N_{d_LES} indicate how retrieval assumptions and 3D radiative effects alter the inferred droplet number concentration.

When the sun is high, the PDFs of $N_{d_cal}^{3D}(\lambda_{abs})$ are generally similar to their $N_{d_cal}^{1D}(\lambda_{abs})$ counterparts (where $\lambda_{abs} = 1.6, 2.1$, and $3.7 \mu\text{m}$), especially for the $\lambda_{abs} = 2.1 \mu\text{m}$ retrieval (Figures 4a, d, and g). A reason for this similarity is the minimal 3D effects (lack of extreme darkening and brightening in the simulated 3D RT reflectance) when the sun is high, leading to the reflectance, utilized for the cloud property retrieval under 3D RT at high sun, comparable to its 1D RT counterpart. This in turn yields comparable derived N_d . Results for SZA 30° , show that the PDF of $N_{d_cal}^{3D}(\lambda_{abs})$ broadens slightly at both ends (Figure 4b, e and h) compared to the high sun case (Figures 4a, d, and g). This slight broadening of the $N_{d_cal}^{3D}(\lambda_{abs})$ PDFs occurs because 3D-induced darkening and brightening effects occurring simultaneously. The darkening effects

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When the sun is high, the PDFs of $N_{d_cal}^{1D}(2.1\mu\text{m})$ is almost similar to those of $N_{d_cal}^{3D}(2.1\mu\text{m})$ (black line in Figures 4a, d and g). Some level of similarity is also observed for the $N_{d_cal}^{1D}(\lambda_{abs})$ and $N_{d_cal}^{3D}(\lambda_{abs})$ PDFs obtained when $\lambda_{abs} = 1.6$ and $3.7 \mu\text{m}$, as they peak around similar N_d regions (although the peaks of the $N_{d_cal}^{3D}(3.7\mu\text{m})$ are smaller compared to $N_{d_cal}^{1D}(3.7\mu\text{m})$). They also exhibit similar physical cloud processes as the N_{d_LES} described in Sect. 3.1. ...

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makes typically lead to retrieval of smaller τ and larger r_e , which together produce smaller $N_{d_cal}^{3D}(\lambda_{abs})$ values compared to $N_{d_cal}^{1D}(\lambda_{abs})$.

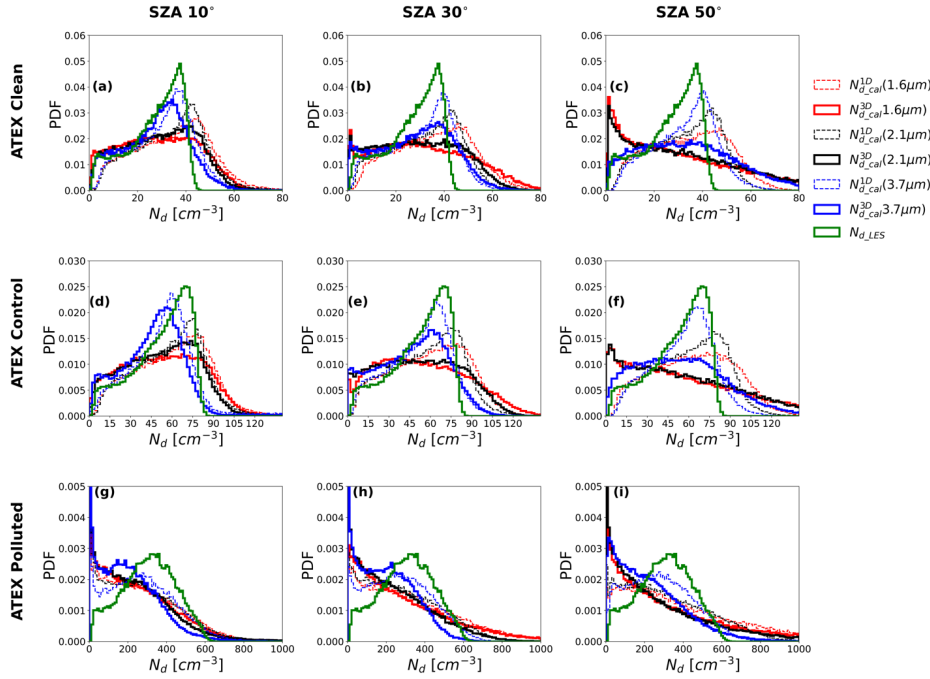
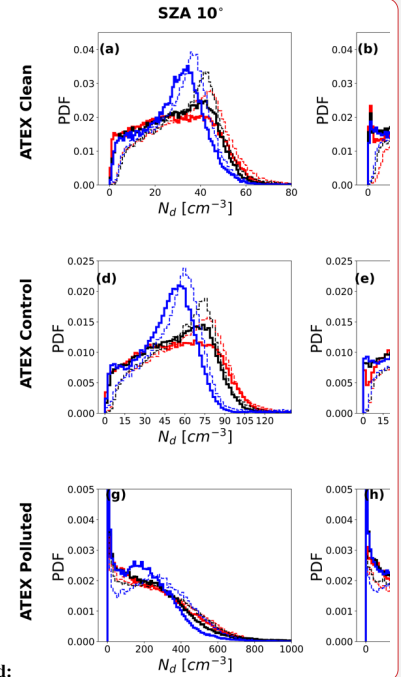


Figure 4. Probability Density Functions (PDFs) of $N_{d_cal}^{1D}(\lambda_{abs})$ and $N_{d_cal}^{3D}(\lambda_{abs})$ at native LES 100 m resolution at SZA 10, 30 and 50 degrees and the reference LES $N_d(N_{d_LES})$ for the ATEX clean, control and polluted case at all time steps. Where λ_{abs} is the absorbing channel paired with $0.86 \mu m$ in the bi-spectral retrieval.

Also, the PDF of $N_{d_cal}^{3D}(\lambda_{abs})$ extends to larger N_d values, compared to $N_{d_cal}^{1D}(\lambda_{abs})$. These larger $N_{d_cal}^{3D}(\lambda_{abs})$ values correspond to where τ is overestimated and r_e is underestimated (a

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681 result of the brightening effects). Overall, though, it appears that the shadowing effect is
 682 predominant in $N_{d_cal}^{3D}(\lambda_{abs})$ for both high and moderate sun positions (discussed further in the
 683 domain-averaged plot in Figure 5). When the sun is low, the PDF of $N_{d_cal}^{3D}(\lambda_{abs})$ become very
 684 broad, with larger peaks at smaller N_d , longer tails at large N_d , and a flattening of the peaks at
 685 moderate N_d that are observed in the high and moderate sun cases. This observed pattern is due to
 686 both brightening and darkening effects occurring simultaneously in the cloud property retrievals
 687 and contribute to the overestimate of $N_{d_cal}^{1D}(\lambda_{abs})$ (darkening effect) and underestimate of
 688 $N_{d_cal}^{1D}(\lambda_{abs})$ (brightening effect) compared to $N_{d_cal}^{3D}(\lambda_{abs})$. To examine the overall 3D radiative
 689 effect impact on the derived N_d at the LES domain scale, we calculate the in-cloud domain average
 690 of the retrieved N_d for each LES case at the different SZA's listed in Sect. 2 (SZA= [10°: 60°] in
 691 steps of 10°). These domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ and $N_{d_cal}^{1D}(\lambda_{abs})$ comparisons will provide
 692 insight on which 3D radiative effect (brightening or darkening) error is dominant (on the domain
 693 scale) in the cloud property retrievals (r_e and τ) that are utilized for the N_d calculations. Figure 5
 694 shows the domain average of $N_{d_cal}^{3D}(\lambda_{abs})$ and $N_{d_cal}^{1D}(\lambda_{abs})$ at the six SZA's for λ_{abs} of 1.6, 2.1,
 695 and 3.7 μm as well as domain average of the N_d reference for the ATEX clean, control and polluted
 696 cases averaged over all time steps.

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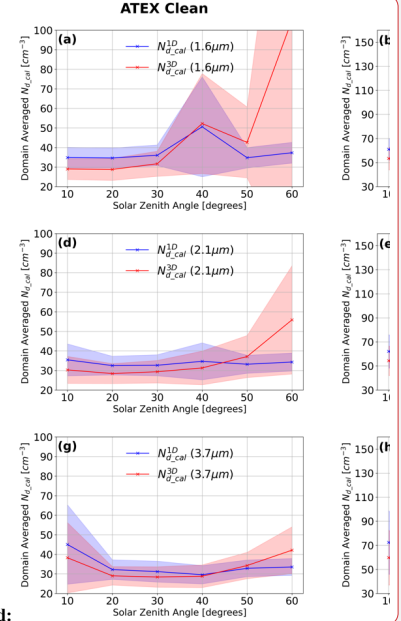
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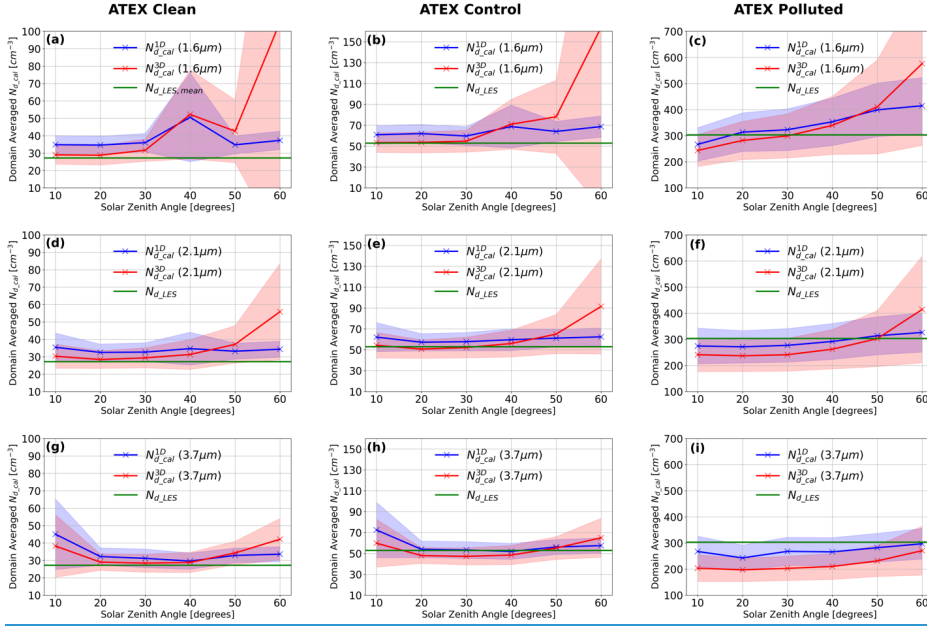


Figure 5. Domain average of $N_{d,cal}^{1D}(\lambda_{abs})$ and $N_{d,cal}^{3D}(\lambda_{abs})$ as a function of solar zenith angle as well as domain average of the $N_{d,LES}$ reference for the ATEX clean, ATEX controlled and ATEX polluted case averaged across all time steps for $\lambda_{abs} = 1.6, 2.1$, and $3.7 \mu\text{m}$. The shaded region is the associated standard error computed from the population variance. Note that the y scale is different for each LES case due to the range of N_d associated with each LES case.

In Figure 5, we generally observe that the domain-averaged $N_{d,cal}^{3D}(\lambda_{abs})$ is underestimated compare to the domain-averaged $N_{d,cal}^{1D}(\lambda_{abs})$ under high to moderate sun positions. However, this reverses towards more oblique SZAs, with the domain-averaged $N_{d,cal}^{3D}(\lambda_{abs})$ becoming larger than $N_{d,cal}^{1D}(\lambda_{abs})$ when the SZA increases towards larger values. It is important to note that domain-averaged $N_{d,LES}$ computed directly from the LES microphysical fields and independent of SZA, absorbing channel, and RT retrieval assumptions is therefore used as the reference for assessing whether the retrieval-derived $N_{d,cal}^{3D}(\lambda_{abs})$ and $N_{d,cal}^{1D}(\lambda_{abs})$ are biased high or low at the domain scale. The separation between the red/blue lines and the green reference

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line shows that the magnitude and sign of the retrieval-derived N_d bias depend on SZA, absorbing channel, and LES case.

The physical impact of the 3D effects changes with the sun position: when the sun is high, absorption dominates horizontal transport differences between bands, but when the sun is low, horizontal transport is dominated by enhanced forward scattering across cloud boundaries. Thus our N_d result indicates that when the sun is high, horizontal transport is limited by absorption leading to darkening effects in the SWIR that dominate in the $N_{d_cal}^{3D}(\lambda_{abs})$ domain-averaged statistics, whereas brightening effects in the VNIR becomes dominant when the sun is oblique. The SZA at which the domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ becomes larger than its 1D counterpart varies with the different scene configuration and absorbing channel utilized for the bi-spectral retrieval.

Interestingly, due to the strong absorption in the $3.7\ \mu m$ band, in the ATEX polluted case (Figure 6i), the $N_{d_cal}^{1D}(3.7\ \mu m)$ is always larger than the $N_{d_cal}^{3D}(3.7\ \mu m)$ for all SZAs considered in this study.

3.3. N_d Derived at at Coarse Spatial Resolution

Because the relative amount of horizontal energy transport between pixels is reduced as the pixel size becomes larger, the behavior of 3D radiative effects is spatial resolution dependent. So far, this study has focused on retrievals at the native LES resolution (100 m). At these high spatial resolutions, retrievals are more sensitive to 3D radiative effects, including cloud vertical inhomogeneity, and microphysical assumptions. However, at the coarser spatial resolutions common in global satellite imager retrievals (e.g., MODIS, Visible Infrared Imaging Radiometer

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Suite (VIIRS)), cloud-property retrievals are also affected by sub-pixel horizontal heterogeneity. This can introduce plane-parallel homogeneous approximation (PPHA) biases because the retrieval assumes each pixel is horizontally homogeneous, whereas the actual coarse pixel may contain unresolved variability in cloud optical thickness, cloud fraction, and cloud structure. Thus, at coarse resolution, the retrieved N_d can be influenced by both 3D radiative effects between neighboring cloudy regions and PPHA biases associated with unresolved sub-pixel heterogeneity. Therefore, extending the analysis of derived N_d to coarser satellite-like resolutions is necessary to determine whether the 3D-RT-induced retrieval errors observed at the native LES resolution persist after spatial aggregation and under resolutions more comparable to operational passive imager observations.

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We examine the behavior of N_d derived from a MODIS or VIIRS-like resolution (800 m is convenient for our LES models) using retrievals from area-averaged 3D RT simulated reflectances, and then compare those N_d PDFs with derived N_d obtained from 3D RT simulations at the native LES resolution (100 m). The PDF comparison of $N_{d_cat}^{3D}(2.1\mu m)$ obtained at both LES and coarse 800 m resolution when all cloudy pixels (i.e., partially cloudy and overcast cloudy pixels; refer to Sect. 2.3 for definition) are considered, is shown in Figure 6 and when only overcast cloudy pixels are considered is presented in Figure 7. In Figure 6 and 7, the green N_{d_LES} PDFs provides the LES reference distribution for each case and allows the 100 m and 800 m retrieval-derived N_d PDFs to be interpreted relative to them. From Figure 6, we clearly see that when all cloudy pixels are considered, the derived N_d PDF at the coarse 800 m resolution are distinct from those at the native LES resolution: the PDF of $N_{d_cat}^{3D}(2.1\mu m)$ at the LES resolution is mostly skewed towards smaller N_d values, for the three solar geometries shown, with smaller in-cloud mean values for the coarse resolution compared to the corresponding LES resolution. The reason for this is due to the

contribution of partially cloudy pixels in the coarse resolution retrievals, further enhancing the impact of the plane-parallel r_e and τ bias contribution and result in smaller N_d values.

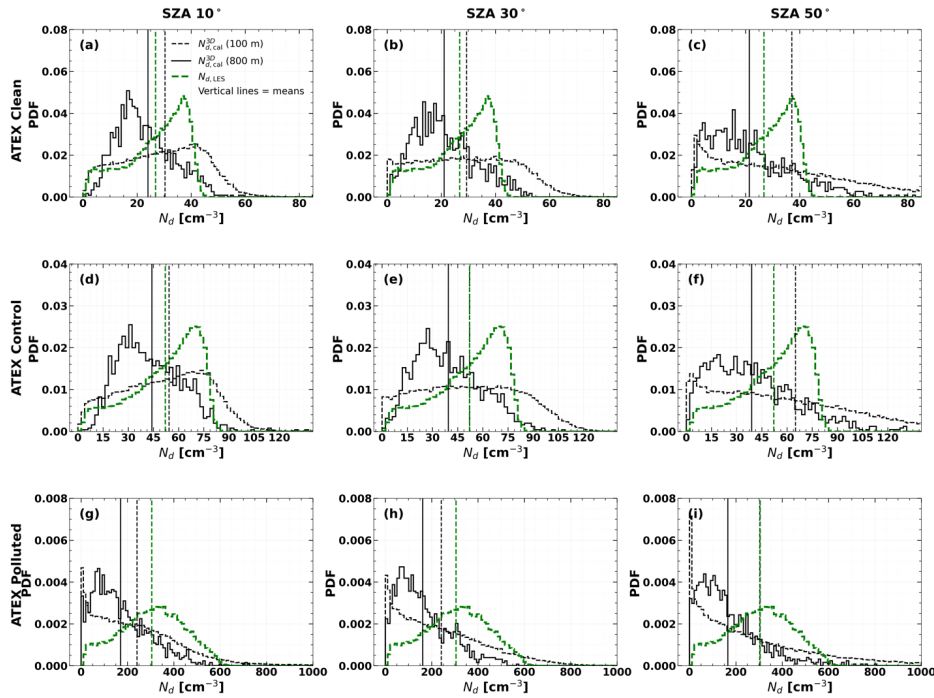
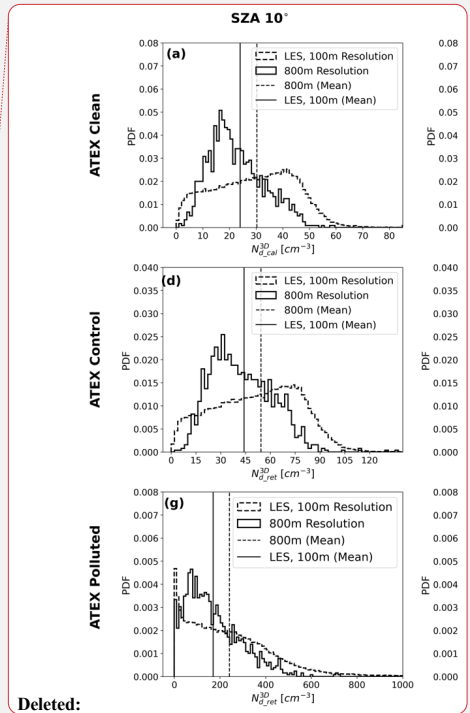


Figure 6. Probability Density Functions (PDFs) comparison of $N_{d-cal}^{3D}(2.1\mu m)$ for LES (100 m) and coarse satellite-like (800 m) resolution for all cloudy pixels, at SZA 10°, 30° and 50° for the ATEX clean, control and polluted case across all time steps. Vertical lines represent mean values. The green line represents the LES reference ($N_{d,LES}$).



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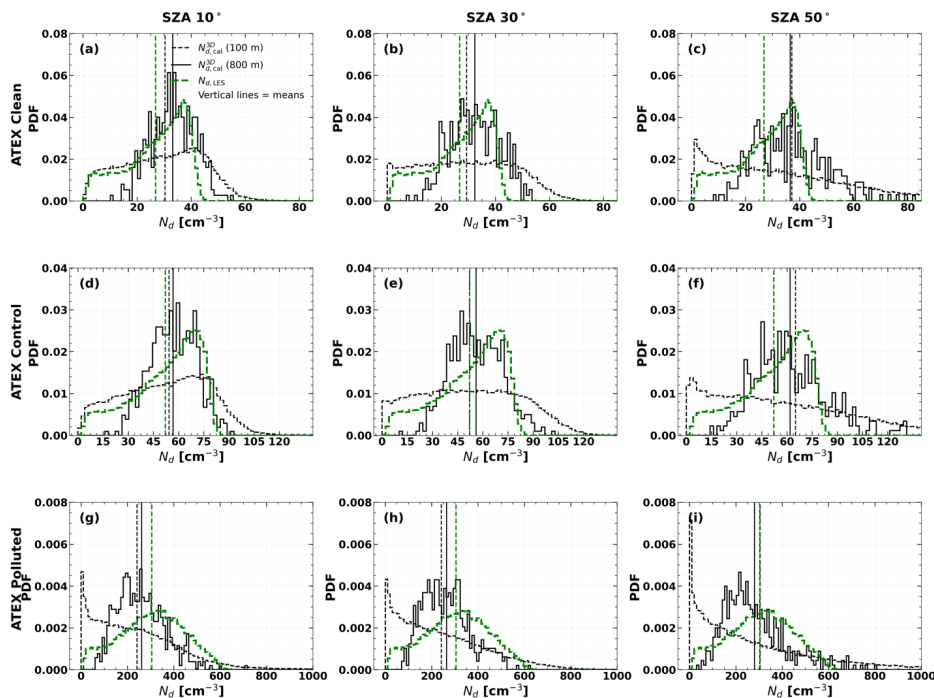
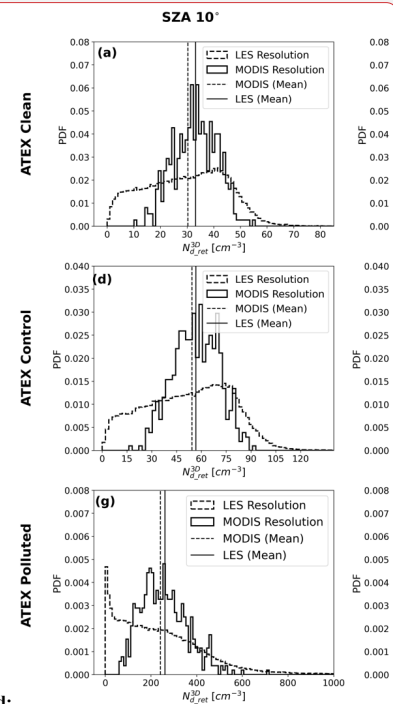


Figure 7. Probability Density Functions (PDFs) comparison of $N_{d_cal}^{3D}(2.1\mu m)$ for LES (100 m) and coarse satellite-like (800 m) resolution for overcast cloudy only pixels, at SZA 10° , 30° and 50° for the ATEX clean, control and polluted case across all time steps. Vertical lines represent mean values. [The green line represents the LES reference \(\$N_{d_LES}\$ \)](#)

When only overcast cloudy pixels are considered (Figure 7), the contribution of the partially cloudy pixels is reduced and the smaller N_d values observed in Figure 6 are mostly absent. As seen in Figure 7, the PDFs of $N_{d_cal}^{3D}(2.1\mu m)$ at the coarse (800 m) resolution, are mostly centered around moderate N_d values, and perhaps surprisingly, happen to have mean values that closely match corresponding values at the LES (100 m) resolution compared to results which utilize all cloudy pixels in the retrievals. These results can provide some additional confidence in using N_d calculations at coarse resolution for our LES cloud fields, as the effects of cloud heterogeneity



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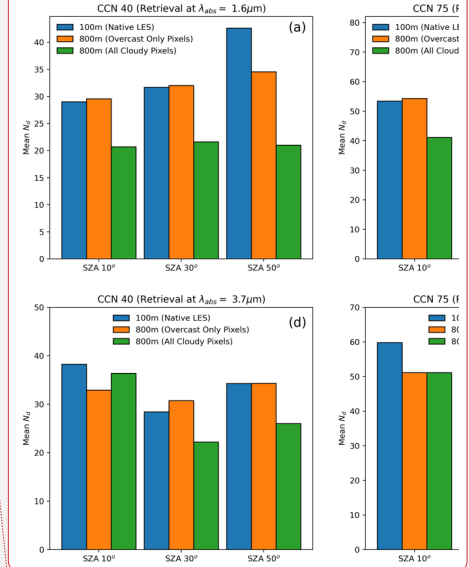
and opposing brightening and darkening 3D effects appear to cancel out (to a large extent) and still provide comparable mean values (although this depends on how well the satellite can identify partially cloudy pixels).

We also examined the sensitivity of $N_{d_cal}^{3D}(\lambda_{abs})$ to spatial aggregation by comparing native LES-resolution (100 m) results with coarse satellite-like resolution (800 m) results for $\lambda_{abs} = 1.6$ and $3.7 \mu\text{m}$ (not shown). In general, the coarse-resolution results based on overcast 800 m pixels are more consistent with the native-resolution mean $N_{d_cal}^{3D}$ than results based on all cloudy 800 m pixels. This behavior is found for both absorbing channels, although for $\lambda_{abs}=3.7 \mu\text{m}$ at $\text{SZA}=10^\circ$, the all-cloudy and overcast-only estimates are both comparable to the native-resolution result. These tests indicate that the treatment of partially cloudy coarse pixels can influence aggregated N_d , but the main resolution-dependent behavior remains robust across the two absorbing channels considered.

3.4. Implications for Estimating Twomey Effect from Retrievals

Having investigated the impacts of 3D effects on the retrievals of τ and r_e , and the subsequently derived N_d , in this section we examine the consequent implications for estimating Twomey effect from these retrievals. In many previous studies, Twomey effect is usually estimated from observations using the so-called absolute cloud albedo susceptibility (S_α) that connects the change of cloud albedo (α) with the change of N_d at a given LWP (Ackerman et al., 2000; Platnick and Twomey, 1994),

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Deleted: Figure 8. Comparison of mean $N_{d_cal}^{3D}(\lambda_{abs})$ obtained at LES resolution of 100 m and coarse 800 m resolution for overcast cloudy only and all cloudy pixels, at SZA 10° , 30° and 50° for the ATEX clean, control and polluted case across all time steps for $\lambda_{abs} = 1.6$ and $3.7 \mu\text{m}$.

Mean values of $N_{d_cal}^{3D}(\lambda_{abs})$ obtained at the native LES (100 m) and coarse satellite-like (800 m) resolution when λ_{abs} is 1.6 and $3.7 \mu\text{m}$ are presented in Figure 8. The results generally show that for $\lambda_{abs} = 1.6 \mu\text{m}$, the mean $N_{d_cal}^{3D}$ at the LES native resolution, is more comparable to results from the coarse resolution, when considering only overcast cloudy pixels than to those obtained when all cloudy pixels are included (Figures 8 a, b and c). Results when $\lambda_{abs} = 3.7 \mu\text{m}$, also show that the mean $N_{d_cal}^{3D}$ at the LES native resolution, is more comparable to results from the coarse resolution, when considering only overcast cloudy pixels than to those obtained when all cloudy pixels are considered, except for the SZA 10° case, where results which utilize all cloudy pixels are slightly better than the overcast pixels only, but regardless, both are still comparable to the LES resolution result (Figures 8d, e and f).

$$S_{\alpha} \equiv \frac{\partial \alpha}{\partial N_{d,cal}} \Big|_{LWP} \approx \frac{\partial \alpha}{\partial \tau} \cdot \frac{\partial \tau}{\partial N_{d,cal}} \Big|_{LWP}. \quad (8)$$

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While S_{α} is often derived empirically through numerical regression of satellite retrievals, many researchers try to develop theoretical expectations for these relationships. This involves using the chain rule of differentiation (Eq. (8)), where the first term—how cloud albedo depends on cloud properties (e.g., $\frac{\partial \alpha}{\partial \tau}$), is usually estimated from simplified radiative transfer models such as the two-stream approximation. The second term—how τ changes with respect to N_d at constant LWP ($\frac{\partial \tau}{\partial N_{d,cal}} \Big|_{LWP}$) is often derived from idealized cloud models assuming constant LWC or sub-adiabatic conditions. Of course, the change of N_d can also potentially lead to the change of LWP known as cloud adjustment, but it is beyond the scope of this study. Here, we first examine the impacts of 3D effect on the estimation of $\frac{\partial \tau}{\partial N_{d,cal}} \Big|_{LWP}$, which we define as the retrieval sensitivity $s(\tau)$. Because a constant N_d is assumed, regardless of the cloud vertical profile assumptions, the relationship $\tau \propto N_d^{1/3} LWP^{5/6}$ holds (Ackerman et al., 2000; Platnick and Twomey, 1994), and implies that,

$$s(\tau) \equiv \frac{\partial \ln \tau}{\partial \ln N_{d,cal}} \Big|_{LWP} \approx \frac{1}{3}. \quad (9)$$

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Constraining fixed LWP regimes is required to evaluate $s(\tau)$ and S_{α} . Thus, we group LWP into eight bins ranges (B1 to B8) as given in Table 1. [The bin edges were derived from percentile boundaries of valid retrieved LWP distribution computed at 10% intervals; however, the first two percentile ranges characterized by small LWP were excluded, leaving the eight retained bins now](#)

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labeled B1 to B8. We will point it out explicitly in the analysis whether the LWP bin categorization is based on the 1D or 3D retrievals or from the LES LWP (LWP_{LES}) cloud field.

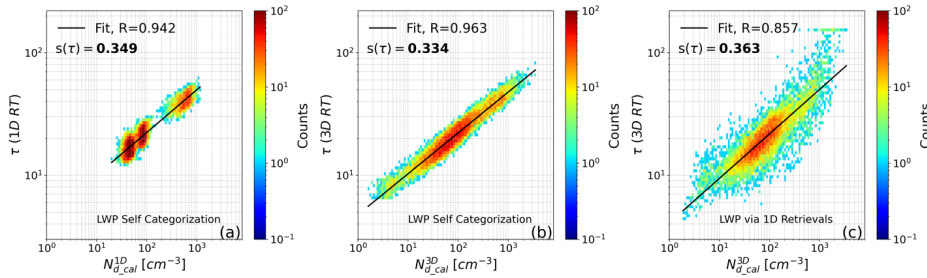
Bin Category	B1	B2	B3	B4	B5	B6	B7	B8
LWP range [gm^{-2}]	43.66	52.72	62.50	74.10	88.55	109.20	142.61	212.19
	–	–	–	–	–	–	–	–
	52.72	62.50	74.10	88.55	109.20	142.61	212.19	Max(LWP)

Table 1. Liquid Water Path (LWP) bins used in this study. B1 to B8 are the bin number representation. Max(LWP) = 3333.33, is maximum LWP value obtained from the retrievals (computed by utilizing Max allowable r_e ($40 \mu m$) and Max allowable τ (150) in Eq. (7)). The bin edges were derived from percentile boundaries of the valid retrieved LWP distribution computed at 10% intervals. The first two percentile ranges were excluded, leaving the eight retained bins labeled B1–B8.

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Describing retrieval sensitivity requires numerical regression of $\partial \ln \tau$ with $\partial \ln N_d$. As will be demonstrated in this section, 3D effects can influence the estimation of $s(\tau)$ through two main mechanisms:

- The errors in τ and $N_{d,cal}$ retrievals caused by the 3D effects as shown in Sect. 3.2 and 3.3 can in turn lead to errors in the $\partial \ln \tau - \partial \ln N_{d,cal}$ regression. We refer to this as the “**retrieval error effect**”.
- To obtain the theoretical $s(\tau) \approx \frac{1}{3}$, the $\partial \ln \tau - \partial \ln N_{d,cal}$ regression must be performed at constant LWP. Since LWP is also derived from retrievals in reality, it is subject to errors caused by 3D effects. These errors can cause mis-categorization of LWP bins in the regression analysis. We refer to this as the “**LWP categorization error**”.



894

895 **Figure 8.** Regression plots of retrieved cloud optical thickness (τ) versus derived cloud droplet number concentration
 896 (N_d) for Liquid Water Path (LWP) bin 7 (B7; range $142.61 - 121.19 \text{ g m}^{-2}$), based on bi-spectral retrievals using the
 897 $0.86 \mu\text{m}$ and $2.1 \mu\text{m}$ channels at a large-eddy simulation (LES) resolution of 100 m and a solar zenith angle (SZA) of
 898 50° . Panel (a) shows results under 1D radiative transfer (RT) using LWP bin categorization based on 1D retrievals
 899 (i.e., LWP Self-Categorization); panel (b) shows 3D RT results using LWP binning from 3D retrievals (i.e., LWP
 900 Self-Categorization); and panel (c) presents 3D RT results using LWP bin categorization based on 1D retrievals.
 901 Values in bold print indicate sensitivity $s(\tau)$ computed according to Eq. (9) and R is the correlation coefficient.

902

903 We first demonstrate the impact of retrieval and LWP categorization errors by calculating
 904 $s(\tau)$ from regressions of a single LWP bin (B7; $142.61 - 212.19 \text{ g m}^{-2}$). This analysis combines
 905 all three ATEX cases (clean $\text{CCN} = 40 \text{ cm}^{-3}$, control $\text{CCN} = 75 \text{ cm}^{-3}$, and polluted $\text{CCN} = 600$
 906 cm^{-3}) utilizing the $0.86 \mu\text{m}$ and $2.1 \mu\text{m}$ channels at SZA 50° . The regressions shown in **Figure 8**
 907 depict τ vs. $N_{d,cal}$ and corresponding $s(\tau)$ obtained from regression. For both 1D and 3D retrievals
 908 based on LWP self-categorization (i.e., where LWP is determined consistently from cloud
 909 properties retrieved from the respective RT reflectance fields), the regressions in **Figure 8a** and **b**
 910 clearly demonstrate a consistent power law scaling with different slopes; broadened variability in
 911 3D retrievals, caused by 3D effects, which subsequently leads to a more consistent $s(\tau)$ value
 912 ($s(\tau) = 0.334$), even better than 1D RT result ($s(\tau) = 0.349$). However, in **Figure 8c**, we observe
 913 that clustering 3D retrievals into LWP bins using a reference dataset that is insensitive to the 3D
 914 radiative effects (in this case the reference LWP is calculated from 1D retrievals) leads to
 915 noticeably different distributions, where regression becomes less well correlated (correlation

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921 coefficient (R)= 0.857, compared to $R > 0.9$ for both LWP Self-categorization cases) and alters
 922 the calculated $s(\tau)$ value ($s(\tau) = 0.363$). This mismatch between retrievals and LWP classification
 923 highlights the importance of using a self-consistent LWP binning approach when calculating
 924 sensitivity. Maintaining consistency in LWP categorization is also important for accurately
 925 analyzing susceptibility estimates: thus, the use of consistent LWP categorization will be explicitly
 926 stated wherever it is applied in the remainder of this study.

927 Having demonstrated the impact of 3D effects on $s(\tau)$ using a single representative LWP
 928 bin, we next use our simulations to investigate how these errors influence the cloud susceptibility.
 929 To evaluate this, it is useful to examine how albedo responds to variations in the relative droplet
 930 number concentration. Thus, we follow the approach of Painemal and Minnis (2012) and define
 931 relative albedo susceptibility ($S_{\alpha,rel}$), which accounts for the dependence of droplet number
 932 concentration on spatial variability, as:

$$934 \quad S_{\alpha,rel} = N_{d,cal} \frac{\partial \alpha}{\partial N_{d,cal}} \Big|_{LWP} = \frac{\partial \alpha}{\partial \ln N_{d,cal}} \Big|_{LWP}. \quad (10)$$

936 Analyzing $S_{\alpha,rel}$, which is based on fractional (logarithmic) changes in number concentration, will
 937 help to minimize the impact of absolute error biases in $N_{d,cal}$ retrievals.

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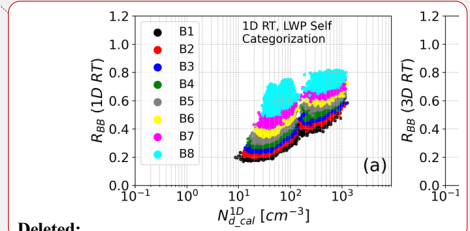
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 simulating radiative fluxes and albedos in our RT solver,
 here (and in the remainder of this text), we would
 approximate,

$$S_{alb,rel} \propto \frac{\partial R_{BB}}{\partial \ln N_{d,cal}} \Big|_{LWP}.$$

(9)



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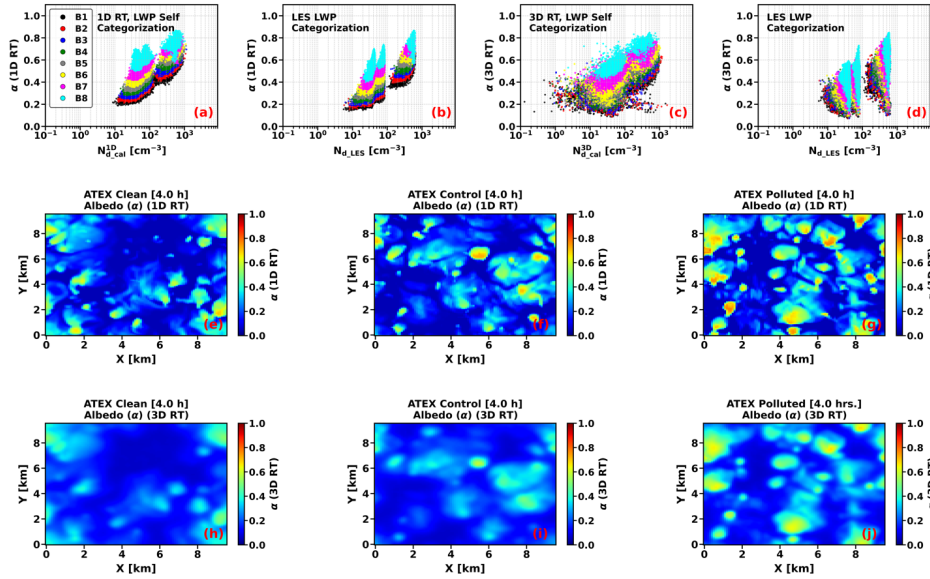


Figure 9. Scatter plots for similar liquid water path (LWP) bins showing albedo (α) vs. N_d (logarithmic scale). α under 1D RT vs. derived N_d obtained from retrievals based on 1D RT reflectance when LWP bin is categorized by 1D retrievals (LWP Self Categorization) in (a), α under 1D RT vs. $N_{d,LES}$ when LWP bin is categorized by the LES LWP in (b), α under 3D RT vs. derived N_d obtained from retrievals based on 3D RT reflectance when LWP bin is categorized by 3D retrievals (LWP Self Categorization) in (c), α under 3D RT vs. $N_{d,LES}$ when LWP bin is categorized by the LES LWP in (d). Map of α for the ATEX clean, control and polluted LES scenes under 1D RT at 4 h of simulation time in (e), (f) and (g) respectively and under 3D RT in (h), (i) and (j) respectively. All RT simulations are carried out at SZA 20° and bi-spectral retrievals utilize the 0.86 and $2.1\mu m$ channel at LES resolution of 100 m.

Figures 9 and 10 (panels a-d) present the α - N_d relationships at the native LES resolution of 100 m across all LES cases and time steps, stratified by the eight LWP bins defined in Table 1. Figure 9 corresponds to SZA = 20° , while Figure 10 shows the same set of diagnostics for a more oblique illumination condition at SZA = 60° . In both figures, panels (a) and (c) show α versus $N_{d,cal}$, where both $N_{d,cal}$ and LWP are inferred from bi-spectral retrievals applied to 1D- and 3D-RT simulated reflectances, respectively, using the 0.86 and $2.1\mu m$ wavelength pair. Panels (b) and (d) in both figures show α versus the LES reference droplet concentration ($N_{d,LES}$), with LWP binning based directly on the LWP_{LES} , and with α obtained from 1D and 3D RT, respectively. The

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corresponding albedo maps for the ATEX clean, control, and polluted cases at 4.0 h are shown in panels (e)–(g) for 1D RT and panels (h)–(j) for 3D RT. Together, Figures 9a–d and 10a–d illustrate a key feature of the first aerosol indirect effect (Twomey, 1977): under fixed LWP conditions, increases in droplet number concentration are generally associated with increases in cloud optical thickness and albedo.

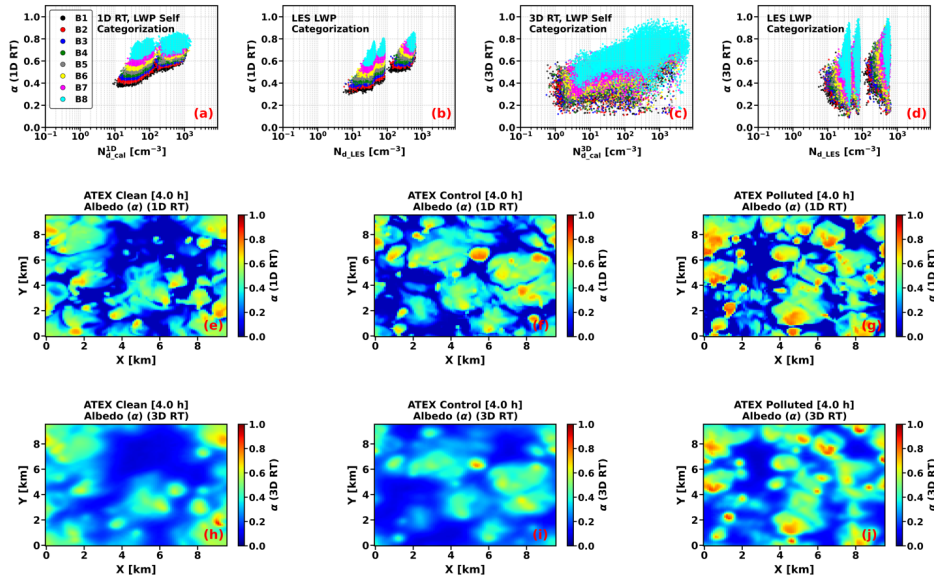


Figure 10. Scatter plots for similar liquid water path (LWP) bins showing albedo (α) vs. N_d (logarithmic scale). α under 1D RT vs. derived N_d obtained from retrievals based on 1D RT reflectance when LWP bin is categorized by 1D retrievals (LWP Self Categorization) in (a), α under 1D RT vs. $N_{d,LES}$ when LWP bin is categorized by the LES LWP in (b), α under 3D RT vs. derived N_d obtained from retrievals based on 3D RT reflectance when LWP bin is categorized by 3D retrievals (LWP Self Categorization) in (c), α under 3D RT vs. $N_{d,LES}$ when LWP bin is categorized by the LES LWP in (d). Map of α for the ATEX clean, control and polluted LES scenes under 1D RT at 4 h of simulation time in (e), (f) and (g) respectively and under 3D RT in (h), (i) and (j) respectively. All RT simulations are carried out at SZA 60° and bi-spectral retrievals utilize the 0.86 and $2.1\mu\text{m}$ channel at LES resolution of 100 m.

For the retrieval-based comparisons, the α – $N_{d,cal}$ relationships are well correlated across all LWP bins, with correlation coefficients greater than 0.77 in Figures 9a and 9c, as well as in Figures 10a

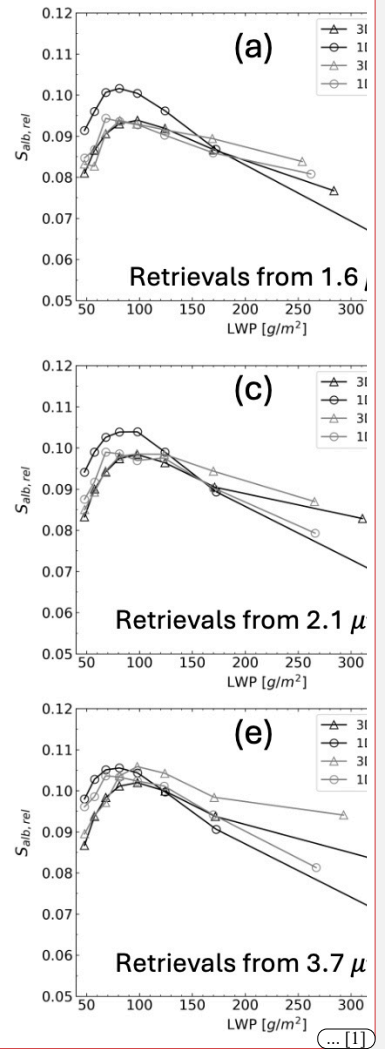
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and 10c. However, the rate at which α increases with N_{d_cal} differs among the cases and depends on whether the RT used to compute α is treated as 1D or 3D and also the RT method used to generate the reflectances from which N_{d_cal} is obtained. For 1D RT based results, the α (1D RT) vs. $N_{d_cal}^{1D}$ relationships in Figure 9a and 10a are very similar, showing increasing α with N_{d_cal} and mainly distinct across each LWP bins. When we consider the corresponding 3D RT cases in Figures 9c and 10c, the α (3D RT) vs. $N_{d_cal}^{3D}$ relationship still shows an overall increase in α (3D RT) with increasing $N_{d_cal}^{3D}$. However, the relationship becomes more dispersed under the more oblique illumination condition. Specifically, Figure 10c, corresponding to $SZA = 60^\circ$, shows a larger spread in α and stronger overlap among LWP bins than Figure 9c, corresponding to $SZA = 20^\circ$. This indicates that 3D radiative effects become more pronounced under the more oblique SZA 60° case, as seen by the larger spread and stronger overlap among LWP bins in Figure 10c ($SZA = 60^\circ$) compared to Figure 9c ($SZA = 20^\circ$). This increased variability reflects the combined influence of 3D radiative effects on the retrieved τ and r_e , which subsequently affect $N_{d_cal}^{3D}$ and derived LWP, together with the fact that α from 3D RT is less local than α from 1D RT. This nonlocality is illustrated by the albedo maps in Figures 9e–j and 10e–j. For the ATEX clean, control, and polluted cases at 4.0 h, the 1D RT albedo fields retain sharper small-scale cloud structure, whereas the corresponding 3D RT albedo fields are smoother and more spatially blurred. This blurring is consistent with radiative smoothing caused by horizontal photon transport and multiple scattering, such that the top-of-domain albedo assigned to a given pixel contains contributions from surrounding cloud pixels. Consequently, the α – N_d relationship under 3D RT becomes less tightly controlled by the local cloud column properties alone, especially at larger SZA .

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When we consider the plots of α (1D RT) versus $N_{d,LES}$ in Figures 9b and 10b, they show a pattern that is broadly similar to the 1D RT based retrieval results in Figures 9a and 10a, respectively. In both SZA cases, α increases with increasing N_d , and the LWP bins remain relatively well separated. This similarity is expected because both panels (a) and (b) use α simulated under 1D RT, where the radiative response is largely controlled by the cloud and surface properties in the cloud column. Therefore, whether $N_{d,*}$ is taken from the retrievals in panels (a) or from the LES reference in panels (b), the α - N_d relationship still retains a clear Twomey-like behavior. Panels (d) in Figures 9 and 10 provide the clearest view of how 3D RT affects the α - N_d relationship when retrieval-related effects are removed, because they use α from 3D RT, the LES reference $N_{d,LES}$, and LWP_{LES} binning. Compared with panels (a) and (b), the relationship in panels (d) is less organized, with larger spread in α and stronger overlap among LWP bins. This indicates that even when $N_{d,LES}$ and LWP are used, α from 3D RT is still influenced by surrounding cloud structure through horizontal photon transport and radiative smoothing, rather than by radiative properties of the local cloud column alone. The effect is more pronounced in Figure 10d than in Figure 9d, showing that the nonlocal influence of 3D RT increases under the more oblique illumination condition at SZA = 60°. Thus, panels (d) show that 3D RT can weaken the local Twomey-like α - N_d relationship, with direct implications for susceptibility estimates under 3D RT. Figure 11 shows $S_{\alpha,rel}$ from diagnostics in Figure 9 (at SZA 20°) and Figure 10 (at SZA 60°) computed within the same LWP bins B1 to B8 and plotted as a function of the mean LWP and mean τ of each bin. The LES-truth reference is computed from the regression of α (3D RT) versus $N_{d,LES}$

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relationship using the LWP_{LES} binning, so it provides a reference to compare our retrieval-based susceptibility estimates.

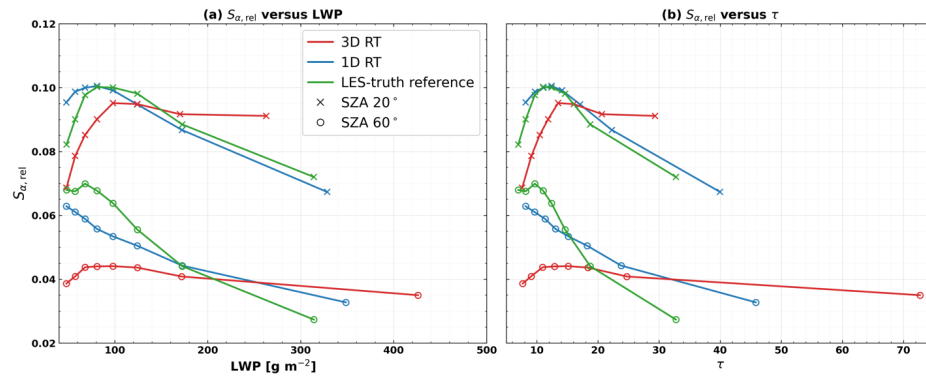


Figure 11. Mean relative cloud albedo susceptibility ($S_{\alpha,rel}$) across all LES cases and time steps at solar zenith angles (SZAs) 20 and 60 degrees, shown as a function of liquid water path (LWP) in panel (a), and as a function of cloud optical thickness (τ) in panels (b). Where, N_d and corresponding LWP or τ are retrieved by bi-spectral method using the reflectance pairs of $0.86 \mu\text{m}$ combined with $2.1 \mu\text{m}$ at the LES-native resolution (100 m). Reference plots in green are obtained from 3D RT albedo and LES microphysics.

For $\text{SZA} = 20^\circ$, the 1D RT result follows the LES-truth reference closely at low to moderate LWP/ τ values, where $S_{\alpha,rel}$ increases and then begins to decrease. The 3D RT result is lower than the reference at the smallest LWP/ τ , but becomes comparable at intermediate values and remains larger than the reference at higher LWP/ τ , suggesting that 3D radiative effects modify the LWP/ τ -dependence of the susceptibility. For the more oblique case at $\text{SZA} = 60^\circ$, all susceptibility values are generally smaller than those at $\text{SZA} = 20^\circ$, and the 3D RT retrieval produces a flatter response with increasing LWP compared with the LES-truth reference and the 1D RT estimate. This indicates that, under more oblique illumination, 3D radiative smoothing and retrieval effects weaken the local α - N_d sensitivity within the LWP bins. The α - $N_{d,cal}$ relationship obtained when

Deleted: The R_{BB} vs. $N_{d,cal}$ relationship under 3D RT for all LWP bins considered (B1 to B8; different colors in Figure 10b) spans a wider range of values, compared to its 1D counterpart (Figure 10a). The wide range of $N_{d,cal}$ values in the 3D is due to the brightening and darkening 3D effects associated with the cloud property retrievals at oblique SZA ($\text{SZA} = 50^\circ$) which results in overestimation and underestimation of $N_{d,cal}$. As expected, categorizing the 3D RT results using LWP derived from the LES model—which is unaffected by 3D radiative effects—reveals a greater spread in each bin, suggesting LWP mismatch (Figure 10c) compared to 3D results with LWP self-categorization (Figure 10b). This underscores the importance of applying a consistent LWP categorization strategy when evaluating susceptibility estimate at fixed LWP. Subsequently, our relative susceptibility calculations will be all based on LWP self-categorization.

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1117 $N_{d,cal}$ is derived from cloud properties retrieved by pairing the $0.86\ \mu\text{m}$ with either the 1.6 or 3.7
 1118 μm band exhibits a similar pattern to that observed with the 0.86 and $2.1\ \mu\text{m}$ combination (not
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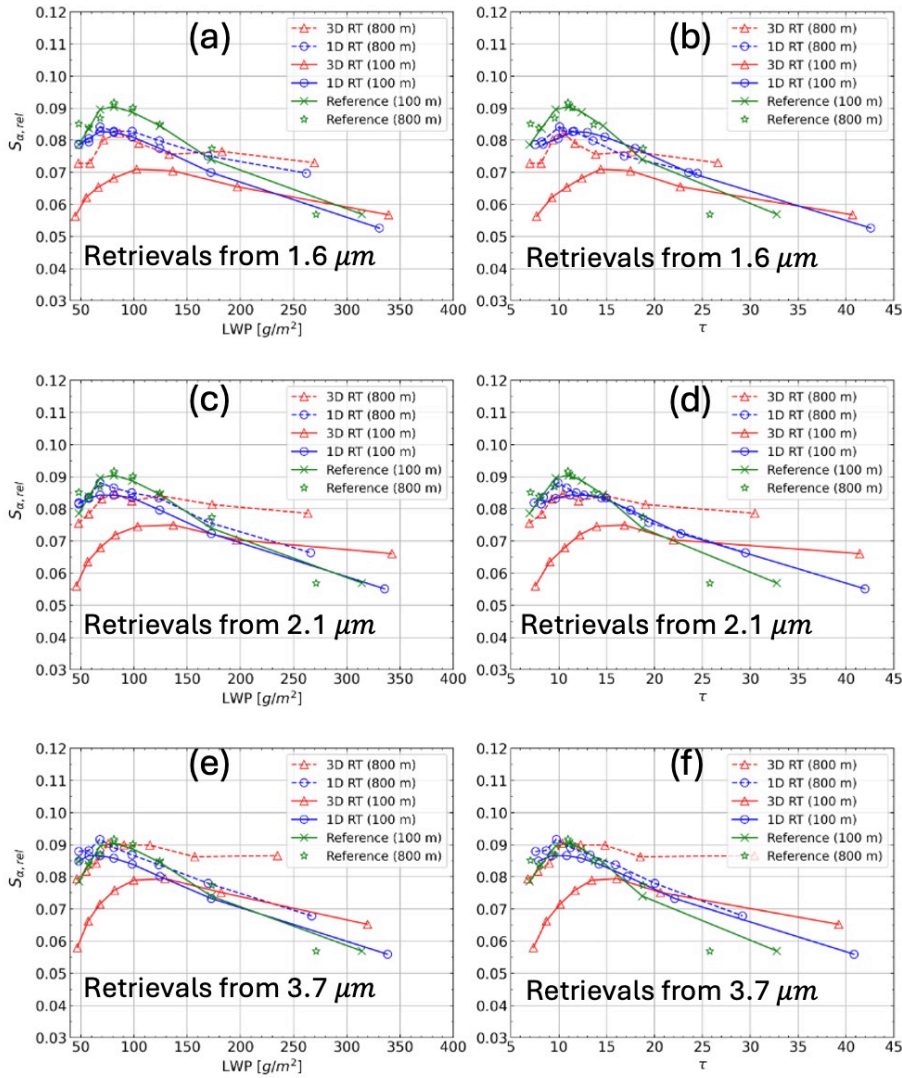


Figure 12. Mean relative cloud albedo susceptibility ($S_{\alpha,rel}$) across all LES cases, time steps and six solar zenith angles (SZAs), shown as a function of liquid water path (LWP) in panels (a), (c), and (e), and as a function of cloud optical thickness (τ) in panels (b), (d) and (f). In each case, N_d and corresponding LWP or τ are retrieved by bi-spectral method using the reflectance pairs of $0.86 \mu m$ combined with $1.6 \mu m$ in (a) and (b), $2.1 \mu m$ in (c) and (d), and $3.7 \mu m$ in (e) and (f). Triangles represent $S_{\alpha,rel}$ derived from 3D RT results, while circles denote values from 1D RT. broken lines indicate result of 800 m coarse resolution, and solid lines represent LES-native resolution (100 m). Reference plots in green are obtained from 3D RT albedo and LES microphysics.

After examining the low- and high-SZA cases, we next evaluate whether the same susceptibility behavior persists across the full set of illumination conditions used in this study i.e., across all SZA's (SZA= $[10^\circ : 60^\circ]$ in 10° intervals). To achieve this, we calculate the mean relative cloud albedo susceptibility, $S_{\alpha,rel}$, averaged across all LES cases, time steps, and all six SZAs. Figure 12 presents these mean $S_{\alpha,rel}$ as a function of mean LWP within each bin in panels (a), (c) and (e) and as a function of mean τ in panels (b), (d) and (f) for retrievals using the $0.86 \mu m$ band paired with 1.6 , 2.1 , and $3.7 \mu m$, respectively. The LES-truth reference curves are also included at both 100 m and 800 m scales so that the retrieval-based susceptibility estimates can be compared against a reference defined at the same spatial resolution. From Figure 12, we observe that, overall, $S_{\alpha,rel}$ generally increases from the smallest LWP/ τ bins to low-to-moderate LWP/ τ values, and then decreases toward larger LWP/ τ regimes. This behavior is consistent with the reduction in albedo sensitivity as clouds become optically thicker and albedo begins to saturate. The LES-truth reference (green line) shows the same general pattern, indicating that the retrieval-based susceptibility estimates broadly capture the expected dependence of albedo sensitivity on clouds.

At the native LES resolution of 100 m, the 1D RT retrievals generally follow the LES-truth reference more closely than the 3D RT retrievals, especially at low-to-moderate LWP/ τ . The 3D RT retrievals exhibit smaller $S_{\alpha,rel}$ in the lower LWP/ τ regim and show a more gradual decrease at larger LWP/ τ . This difference reflects the combined influence of 3D radiative smoothing on α

and 3D effects induced biases in the retrieved τ and r_e , which propagate into N_d and subsequent susceptibility estimate. At the coarser 800 m resolution, the differences between the 1D and 3D RT susceptibility estimates are reduced at the low LWP/ τ regimes although notable differences occur at the larger LWP/ τ values. The 800 m retrieval curves are also more comparable to the 800 m LES-truth reference than would be expected from comparison with the 100 m reference alone, confirming that the susceptibility benchmark is resolution dependent. This indicates that spatial aggregation partly reduces fine-scale 3D variability and promotes compensation between 3D radiative effects, plane-parallel retrieval biases, and subpixel averaging.

The same broad behavior of $S_{\alpha,rel}$ is observed for all three absorbing-channel combinations at the 100 m resolution, with $S_{\alpha,rel}$ generally increasing at low LWP/ τ before gradually decreasing towards larger LWP/ τ values. At the coarse 800 m resolution, the $S_{\alpha,rel}$ vs. LWP/ τ relationship is similar to those at 100 m, but the 1D vs 3D differences is band dependent. The 1.6 and 2.1 μm retrieval results show somewhat better agreement between the 1D and 3D curves at low to moderate LWP/ τ values. They all have significant differences between 1D and 3D at larger LWP/ τ values. This, indicates that 3D RT effects can greatly modify $S_{\alpha,rel}$ at native LES resolution, but their impact is substantially reduced at coarse satellite-like resolution in low LWP/ τ regimes.

4. Summary and Conclusion

Errors associated with 3D radiative effects and cloud inhomogeneity can influence bi-spectral retrievals of cloud properties (τ and r_e), and subsequently impact estimates of N_d derived from these properties. Because N_d is so important in the representation of clouds in models and the interpretation of aerosol–cloud effects, it is important to disentangle and explain these sources

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To examine how 3D radiative effects impact our relative susceptibility calculations, we combine data across all LES cases (ATEX clean, control and polluted) and time steps for each corresponding R_{BB} and $N_{d,cal}$ at the different SZA's (SZA = $[10^\circ : 60^\circ]$ in 10° intervals) in our study. Hereafter, the mean $S_{alb,rel}$ across all SZAs is calculated and the results analyzed. Figure 11 shows the mean $S_{alb,rel}$ computed under 3D and 1D RT as a function of mean LWP in each bin (Figure 11a, c and e) and mean $S_{alb,rel}$ as a function of mean τ in each LWP bin (Figure 11b, d and f) where cloud retrievals (r_e , τ) utilized to derive LWP and $N_{d,cal}$ are retrieved from pairing the 0.86 μm channel with λ_{abs} of 1.6, 2.1 and 3.7 μm channel. At the native LES resolution, the plot shows that $S_{alb,rel}$ calculated from both 1D and 3D RT (for cloud retrievals from all three absorbing bands) exhibits a similar pattern, characterized by an initial increase with LWP until it reaches a maximum (around $\tau \sim 13$, LWP $\sim 75 - 100 \text{ gm}^{-2}$), followed by a slight decrease, and then a more pronounced decrease at larger LWP/ τ values.

When we compare $S_{alb,rel}$ results derived from 1D and 3D RT simulations. In the 1D case, R_{BB} is from 1D RT simulations, and $N_{d,cal}$ is computed from τ and r_e retrieved using 1D reflectance. In the 3D case, R_{BB} is from 3D RT simulations, and $N_{d,cal}$ is computed from τ and r_e retrieved using 3D reflectance. This comparison shows that for the three absorbing retrieval bands, the 3D results tend to underestimate its 1D counterpart at low to moderate LWP values, whereas the opposite is observed at higher LWP. This difference mostly disappears at coarse spatial resolution (gray symbols) as competition between plane parallel biases and 3D radiative effects apparently compensate one another. Because flux and these bi-spectral microphysical retrievals are radiatively traceable, the 1D and 3D results at coarse spatial resolution largely agree with one another in terms of radiative susceptibility. This indicates that at coarse spatial resolution, 3D effects and plane parallel biases in retrievals are compensating one another for relative radiative measures like $S_{alb,rel}$. Thus, concerns about pixel-level 3D effects retrieval biases at satellite resolution are unlikely to impact estimates of cloud radiative effects significantly. Retrieval artifacts can apparently bias our understanding of microphysical processes in clouds while not impacting our understanding of cloud radiative effects.

1215 of retrieval bias. Therefore, this study focuses on investigating the impact of the 3D radiative
 1216 effects on N_d calculated from τ and r_e retrieved by the bi-spectral method (N_{d_cal}) and further
 1217 probes how an understanding of [albedo susceptibility which quantifies](#) the first aerosol indirect
 1218 effect is impacted when interpreted from such retrievals. We address this by a satellite observation-
 1219 retrieval simulator framework, which consists of synthetic cloud fields, a radiative transfer solver
 1220 and retrieval algorithm. The cloudy scenes examined here spanned numerous temporal snapshots
 1221 from three LES cases, each of which was initialized based on observations made during the 1969
 1222 ATEX field campaign (NE Atlantic trade wind region $\sim 12^\circ\text{N}$, 35°W). These LES cases are each
 1223 characterized by different initial aerosol loadings: a clean ($\text{CCN} = 40 \text{ cm}^{-3}$), control ($\text{CCN} = 75$
 1224 cm^{-3}), and polluted ($\text{CCN} = 600 \text{ cm}^{-3}$) case. From these LES cloud fields, simulated reflectances
 1225 (based on 1D and 3D RT) for MODIS Band 2 ($0.86 \mu\text{m}$), Band 6 ($1.6 \mu\text{m}$), Band 7 ($2.1 \mu\text{m}$), and
 1226 Band 20 ($3.7 \mu\text{m}$) were carried out with the SHDOM RT model (Evans, 1998). Thereafter, bi-
 1227 spectral retrievals were implemented on the simulated reflectance by pairing the $0.86 \mu\text{m}$ VNIR
 1228 band with each of the absorbing wavelength bands ($\lambda_{abs} = 1.6, 2.1$ and $3.7 \mu\text{m}$). Subsequently,
 1229 N_{d_cal} were obtained from the bi-spectral microphysical retrievals at six solar geometries ($\text{SZA} =$
 1230 $[10^\circ : 60^\circ]$ in steps of 10°) and the results were analyzed.

1231 We utilize N_d obtained from the LES as a reference and compare with N_{d_cal} derived from
 1232 1D RT simulations, and with N_{d_cal} inferred from the adiabatic assumption (Eq. (5)) when τ and
 1233 r_e retrievals are constrained toward scenes that more closely approximate that assumption. Results
 1234 demonstrates that, N_{d_cal} from τ and r_e retrieved from homogeneous plane parallel RT-based
 1235 simulated reflectance can still be biased when compared to the reference N_d . This bias arises from
 1236 the complexity of how cloud microphysics is vertically distributed within realistic clouds, which
 1237 may sometimes deviate from the adiabatic assumptions in which the N_{d_cal} equation is based. Also,

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1239 the choice of the absorbing channel in the bi-spectral retrieval is a factor that contributes to this
 1240 derived bias. Due to different liquid water absorption and therefore vertical sensitivities at different
 1241 channels, the $3.7 \mu\text{m}$ band which has the strongest absorption, retrieves r_e closer to the cloud top,
 1242 and agrees the most with the reference LES N_d , compared to results derived from pairing the 0.86
 1243 μm band with the 1.6 and $2.1 \mu\text{m}$ band.

1244 To investigate how 3D radiative effects impact N_d estimates, we calculate N_{d_cal} using
 1245 cloud properties derived from 1D RT-based reflectances and compare directly with N_{d_cal}
 1246 calculated using cloud properties derived from 3D RT-based reflectances. These comparisons are
 1247 performed for multiple bi-spectral band pairs (pairing VNIR= $0.86 \mu\text{m}$ with either of λ_{abs} = 1.6 ,
 1248 2.1 , and 3.7). Results from this comparison carried out at different SZAs show that, for high sun
 1249 (SZA= 10°), the statistical distributions of $N_{d_cal}(\lambda_{abs})$ from 3D and 1D RT agree well for all three
 1250 LES cases. Domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ and $N_{d_cal}^{1D}(\lambda_{abs})$ was also compared to determine the
 1251 overall impact of the 3D (brightening and darkening) effects on the domain scale. For the high sun
 1252 case, the domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ is underestimated compared to $N_{d_cal}^{1D}(\lambda_{abs})$. This
 1253 indicates that the darkening effect which dominates cloud property retrievals when the sun is high
 1254 also dominates the domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ result (i.e., overestimated r_e and underestimated
 1255 τ dominates the cloud property retrievals and yields lower $N_{d_cal}^{3D}(\lambda_{abs})$ values compared to
 1256 $N_{d_cal}^{1D}(\lambda_{abs})$ results). A similar pattern is also observed for the domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$
 1257 derived from not too oblique SZAs (e.g., SZA 30°).

1258 When the sun is low at SZA 50° , 3D radiative effects become stronger and produce larger
 1259 variability in the retrieval-derived N_d . This occurs because both brightening and darkening effects
 1260 occur in significant amount at the low sun angle. Brightening effects tend to produce larger τ and
 1261 smaller r_e , leading to larger $N_{d_cal}^{3D}(\lambda_{abs})$ compared to corresponding $N_{d_cal}^{1D}(\lambda_{abs})$ values, while

Deleted: When the sun is low at SZA 50° , the PDF distribution of $N_{d_cal}^{3D}(\lambda_{abs})$ are more stretched and broad when compared to $N_{d_cal}^{1D}(\lambda_{abs})$ results. A reason for this is that the cloud properties from which $N_{d_cal}^{3D}(\lambda_{abs})$ is calculated, exhibits both underestimation and overestimation when compared to their 1D RT counterpart. This is expected since brightening and darkening 3D effects both occur in the 3D RT-simulated reflectance observation under low sun, which manifest in the cloud property retrieval as overestimation of τ and underestimation of r_e in brightened pixels and underestimation of τ and overestimation of r_e in darkened pixels. When we examine the domain-averaged results at low sun angles (typically SZA $> 45^\circ$), the brightening effect appears to be dominant with domain-averaged $N_{d_cal}^{3D}(\lambda_{abs})$ being overestimated compared to domain-averaged $N_{d_cal}^{1D}(\lambda_{abs})$ as the SZA becomes large. Although, the exact SZA at which this overestimation occurs varies for the different λ_{abs} considered in this study.

darkening effects tend to produce smaller τ and larger r_e , leading to smaller $N_{d_cal}^{3D}(\lambda_{abs})$ compared to $N_{d_cal}^{1D}(\lambda_{abs})$ values. At the domain scale, the relative magnitude of these effects changes with SZA and absorbing channel used for retrievals. For larger SZAs, the brightening effect becomes more dominant in several cases, causing the domain-averaged $N_{d_cal}^{1D}(\lambda_{abs})$ to exceed its 1D RT counterpart, although the SZA at which this transition occurs depends on the absorbing channel and cloud scene.

The statistical distributions of N_{d_cal} at satellite-like (MODIS, VIIRS) pixel footprints were also examined. Compared to retrievals at the native LES resolution of 100 m, ~~the~~ 800 m ~~results~~ exhibits some subtle differences depending on how one defines the cloud mask for the coarse resolution observations. We divide the retrieval population into two groups: all cloud retrievals (including partly cloudy) or overcast cloudy pixels (i.e., removing partly cloudy pixels). When all valid cloudy pixels are included, including partially cloudy pixels, the N_{d_cal} distributions tend to shift toward smaller values, indicating that partially cloudy coarse pixels can influence the retrieved N_d . However, when only confident overcast cloudy pixels are retained, the mean N_{d_cal} values at the native LES and coarse satellite-like resolutions become more comparable. This behavior is observed consistently for the 1.6, 2.1, and 3.7 μm absorbing-channel retrievals, which suggests that it is not just restricted to a specific band. Therefore, satellite-like coarse-resolution retrievals can provide comparable mean N_d estimates to the native-resolution results when partly cloudy pixels are excluded, and only confidently cloudy pixels are utilized.

We show that the sensitivity $s(\tau)$, defined as the change in τ with respect to N_d at constant LWP ($\frac{\partial \tau}{\partial N_{d_cal}}|_{LWP}$), is influenced by both retrieval errors due to 3D radiative effects and LWP categorization errors. The results reveal that 3D effects retrieval error cancel out in $s(\tau)$ calculated

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Deleted: When all valid cloudy pixels are utilized for the averaging, the PDF of $N_{d_cal}^{3D}(2.1\mu\text{m})$ skews toward smaller N_{d_cal} values, indicating that the contribution of partially cloudy pixels, significantly impacts N_{d_cal} calculations. However, when only confident cloudy pixels are used, the PDF shifts away from these small N_{d_cal} values, and the mean N_{d_cal} at the LES and coarse satellite-like resolutions are comparable. Similar observations are also seen in the 1.6 and 3.75 μm derived N_d , where the mean N_{d_cal} at the LES and coarse satellite-like resolutions are comparable, suggesting that retrievals from larger footprints can still be reliable if only confident cloudy pixels are considered.

under 3D RT using LWP self-categorization (i.e., LWP derived from r_e and τ retrieved from 3D reflectance) to give $s(\tau)$ values that aligns closely to theoretical expectation (1/3), compared to corresponding $s(\tau)$ values obtained when non-self-consistent LWP categorization is applied.

At the native LES resolution, the average relative susceptibility ($S_{\alpha,rel}$) computed from retrievals which utilize 3D RT reflectance is generally smaller than the corresponding 1D RT estimates at low LWP/ τ for all three absorbing-channel retrievals. However, this behavior reverses toward larger LWP/ τ , where the $S_{\alpha,rel}$ from 3D becomes comparable to, or larger than, the 1D RT result. This indicates that 3D RT modifies not only the magnitude of $S_{\alpha,rel}$, but also its dependence on cloud optical state. At the coarser 800 m resolution, the differences between the 1D- and 3D-based susceptibility estimates are substantially reduced over the low-to-moderate LWP/ τ regimes, although the degree of agreement remains somewhat dependent on the absorbing channel. This closer agreement at coarse resolution low- to moderate LWP regimes suggest compensation between plane-parallel retrieval biases and 3D radiative effects, yielding more comparable susceptibility estimates than those obtained at the native LES resolution.

In conclusion, we demonstrate the impact of 3D effects on bi-spectral retrieval of τ and r_e and their subsequent impact on the ability to infer N_d from such retrievals. Retrieval errors are known to have far-reaching implications, biasing aerosol–cloud interaction studies that attempt to constrain the impact of the first aerosol indirect effect (Quaas et al., 2020). Our study shows that 3D effects have significant impact on relative susceptibility at the native LES resolution than at coarser satellite-like footprints, and low to moderate LWP and optical depth regimes. At the coarse 800 m resolution, differences between 1D- and 3D-based relative susceptibility estimates are substantially reduced across the low-to-moderate LWP/ τ regimes, suggesting partial compensation between plane-parallel retrieval biases, and 3D radiative effects. Improvements in satellite

Deleted: At the LES resolution, the average relative susceptibility ($S_{alb,rel}$) computed under 3D RT, for retrievals computed from all three absorbing bands, underestimates the corresponding 1D RT results at low LWP, and vice versa at larger LWP values. We further show that, at the coarse 800 m resolution, $S_{alb,rel}$, computed under 1D RT closely aligns with corresponding 3D RT results, when evaluated as a function of average LWP and τ in each LWP bin across the six SZAs considered in this study. This is due to cancellation of plane parallel bias and 3D effects errors to give comparable results.

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1358 retrievals of cloud properties, from which N_d is inferred, remain important—especially the
 1359 development of retrieval methods that are less sensitive to 3D radiative effects. Recently, there has
 1360 been increased development of polarimetric retrievals of cloud properties, offering the potential to
 1361 constrain biases associated with r_e retrieval to the extent that the uppermost cloud layer
 1362 microphysics captures the desired physics. However, the method of retrieving τ still utilizes the
 1363 spectral technique, so any N_d retrieval derived from a polarimeter may not benefit from some of
 1364 the compensating error impacts observed in this study for [low to moderate LWP and \$\tau\$ regimes at](#)
 1365 [the](#) coarse-resolution observations. Other retrieval techniques such as cloud property retrievals
 1366 using a combination of active and passive instruments, [can also provide a better means to retrieve](#)
 1367 N_d and should be greatly encouraged.

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 1370 **DATA AVAILABILITY** The SHDOM radiative transfer code used in this study is freely
 1371 available online from <https://coloradolinex.com/shdom/> (SHDOM for Atmospheric Radiative
 1372 Transfer is described in detail in a journal article available at [https://doi.org/10.1175/1520-](https://doi.org/10.1175/1520-0469(1998)055<0429:TSHDOM>2.0.CO;2)
 1373 [0469\(1998\)055<0429:TSHDOM>2.0.CO;2](https://doi.org/10.1175/1520-0469(1998)055<0429:TSHDOM>2.0.CO;2); Evans, 1998). The post-processed LES fields and
 1374 radiative transfer simulation results for this study are available at
 1375 <https://doi.org/10.5281/zenodo.16945607> (Ademakinwa, 2025).

1376
 1377 **AUTHOR CONTRIBUTIONS** Conceptualization, ZZ; methodology, ASA, ZZ and DM;
 1378 software, ASA, DM; validation, ASA, DM and ZZ; formal analysis, ASA and DM; investigation,
 1379 ASA and ZZ; data curation, ASA, ZZ; writing (original draft preparation), ASA; writing (review
 1380 and editing), ASA, DM, JW, KGM, , SP, SP, ZHT, and ZZ; visualization, ASA; supervision, ZZ;
 1381 project administration, ZZ; funding acquisition, ZZ, JW, KGM, and SP. All authors have read and
 1382 agreed to the published version of the paper.

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Deleted: Appendix A: N_d from the LES When Obtained as Column Average

For reference here we display our comparison of $N_{d,cal}^{1D}(\lambda_{abs})$ vs. N_d from the LES cloud field, however this time we obtain the definition for the reference N_d from the in-cloud vertical column average N_d without weighting by β_{ext} as used in Eq. (1) (denoted as $N_{d,LES}$). Figure A1 shows these comparisons as a regression plot of $N_{d,cal}^{1D}(\lambda_{abs})$ vs. $N_{d,LES}$ for the ATEX clean, control and polluted cases at 4.0 hours of simulation time, when λ_{abs} is 1.6, 2.1 and 3.7 μm . The comparison shows $N_{d,cal}^{1D}(\lambda_{abs})$ mostly being overestimated compared to $N_{d,LES}$ for the three LES cases. This is expected since $N_{d,LES}$ represents the central tendency and smaller values of N_d in the cloud column, particularly in regions close to the cloud top could directly impact the average value.

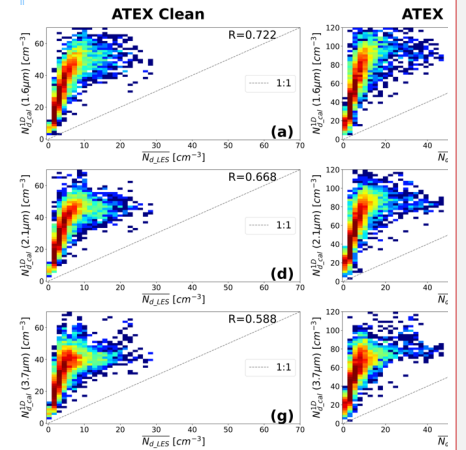


Figure A1. Comparison of N_d derived from 1D RT bi-spectral retrievals versus LES N_d obtained by vertical column average for (a) ATEX clean, (b) ATEX control and (c) ATEX polluted for retrievals at absorbing channel 1.6 μm , for (d) ATEX clean, (e) ATEX control and (f) ATEX polluted for retrievals at absorbing channel 2.1 μm and for (g) ATEX clean, (h) ATEX control and (i) ATEX polluted for retrievals at absorbing channel 3.7 μm at 4 h of simulation time. Note: 1D RT and bi-spectral retrievals utilized for this comparison were carried out at SZA 50°. Dotted lines indicate 1:1 relationship between the retrieved and LES N_d .

1418 **ACKNOWLEDGMENTS** This study has been supported by the National Aeronautics and Space
1419 Administration ACCESS project (grant no. 80NSSC21M0027). The hardware used in the
1420 computational studies is part of the UMBC High Performance Computing Facility (HPCF). The
1421 facility is supported by the U.S. National Science Foundation through the MRI program (grants
1422 CNS-0821258 and CNS-1228778) and the SCREMS program (grant DMS-0821311), with
1423 additional substantial support from the University of Maryland, Baltimore County (UMBC).

1424 **Competing interests:** None of the authors have any competing interests.

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