

1 **Optimizing the precision of infrared measurements using the Eppley**  
2 **Laboratory, Inc. model PIR pyrgeometer**

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13

14 **Abstract.** The Eppley Precision Infrared Radiometer (PIR) is widely used for broadband (3.5-50  
15  $\mu\text{m}$ ), thermal infrared wavelength measurements of the downwelling and upwelling radiation  
16 from the atmosphere and surface, respectively. The field of view of the instrument is  $2\pi$   
17 steradians with a receiver that has an approximate cosine response. In this paper we examine four  
18 equations suggested by the literature that have been used to transfer irradiance calibrations from  
19 our standard PIRs that are calibrated at the World Radiation Center to field units used for  
20 network operations. We first discuss various equations used to convert the resistance  
21 measurements of the thermistors to temperatures of the body and dome that are used in the  
22 derivation of incoming irradiance. We then use the four related, but distinct, equations for the  
23 transfer of the calibration from standard PIRs to field instruments. A clear choice for the  
24 preferred equation to use for calibration and transfer of calibration to field PIRs emerges from  
25 this study.

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28 **1. Introduction**

29

30 The Eppley Precision Infrared Radiometer (PIR) is a pyrgeometer developed to measure  
31 broadband (3.5-50  $\mu\text{m}$ ) thermal infrared (hereafter, IR) radiation emitted by the atmosphere and  
32 surface. Detecting even small changes in atmospheric and surface IR in the environment are  
33 important to understand as they may indicate temperature changes in the environment caused by  
34 changes in the atmospheric composition or cloud cover.

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36 The PIR originally came equipped with a battery-powered circuit to compensate for the radiation  
37 emitted by the body so the net signal from the instrument was a measure of actual incoming IR  
38 radiation. Users of this instrument that are interested in the optimum precision in their IR  
39 measurements do not use the battery-powered circuit, but, instead, use temperatures from two  
40 thermistors connected to the body and dome of the instrument along with the thermopile output  
41 to calculate the incoming IR irradiance signal.

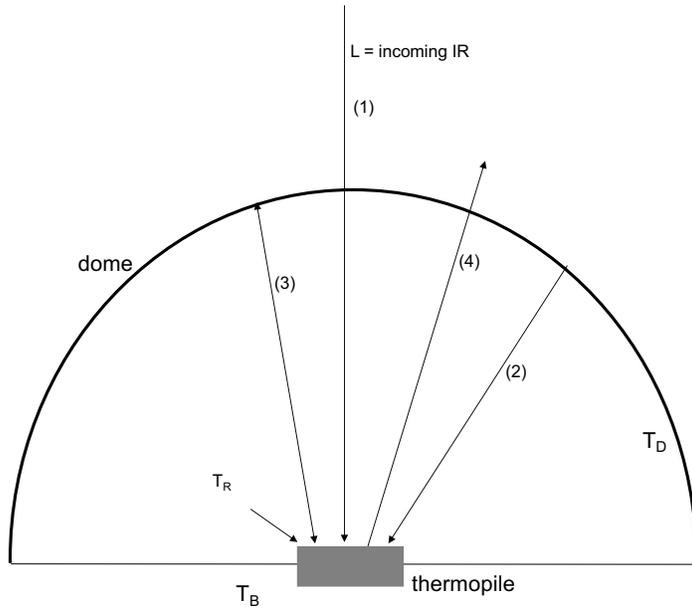
42

43 Fig. 1 illustrates the most significant incoming and outgoing IR irradiances at the thermopile  
44 surface. To derive  $L$ , the incoming IR radiation from the hemisphere outside the instrument,

45 radiative equilibrium of the instrument must be defined. To do that, the sum of the incoming  
 46 radiation transmitted by the dome (labeled 1), radiation emitted by the dome to the receiver  
 47 (labeled 2), and the radiation emitted by the thermopile surface and reflected by the dome  
 48 (labeled 3) are set equal to the radiation emitted by the thermopile surface (labeled 4, which is  
 49 the largest of the IR signals). Note that the dome transmits IR radiation between 3.5 and 50  $\mu\text{m}$ ,  
 50 however, the transmission is not perfect, nor uniform. Considering these components Albrecht  
 51 and Cox (1977) formulated Eq. (1) for the externally received radiation as  
 52

$$L = U_{thermopile}(c_1 + c_2 T_B^3) + \epsilon_o \sigma T_B^4 - k \sigma (T_D^4 - T_B^4), \quad (1)$$

53 where  $U_{thermopile}$  is the voltage measured across the thermopile,  $T_B$  and  $T_D$  are the body and dome  
 54 temperatures in K,  $\sigma$  is the Stefan-Boltzmann constant,  $\epsilon_o$  is the emissivity of the detector, and  
 55  $c_1$ ,  $c_2$ , and  $k$  are constants to be determined in calibration.  
 56



57 **Figure 1.** Schematic for the most significant incoming and outgoing IR radiation components on  
 58 the thermopile surface (numbered 1 – 4) that are considered in calculating the incident  
 59 atmospheric IR irradiance. The top of the dark rectangle is the receiving surface surrounded by  
 60 the dome that transmits IR in the range 3-50  $\mu\text{m}$ . YSI 44031 thermistors are used to measure  $T_B$   
 61 of the body (thermistor buried in the brass body of the PIR) and  $T_D$  of the dome in K.  $T_R$  is the  
 62 estimated receiver temperature in K.  
 63  
 64

65 In practice Albrecht and Cox (1977) dropped the  $c_2 T_B^3$  term as negligible relative to the  $c_1$  term  
 66 and set the emissivity of the body of the instrument  $\varepsilon_o$  to 1 yielding this commonly expressed  
 67 form of their equation  
 68

$$L = \frac{U_{thermopile}}{C} + \sigma T_B^4 - k\sigma(T_D^4 - T_B^4), \quad (2)$$

69  
 70 where  $c_1$  has been replaced by  $1/C$ .  
 71  
 72

73 Philipona et al. (1995), however, used Eq. (1) in its entirety, but to compare symbolically to Eq.  
 74 (2) it is written  
 75

$$L = \frac{U_{thermopile}}{C} (1 + k_1 \sigma T_B^3) + k_2 \sigma T_B^4 - k_3 \sigma (T_D^4 - T_B^4), \quad (3)$$

76 where the  $T_B^3$  term in Eq. (1) is retained, the emissivity of the body is  $k_2$ , and  $k_3$  is the same as  $k$   
 77 in Eq. (2). All constants,  $C$ ,  $k_1$ ,  $k_2$ , and  $k_3$ , are determined in calibration.  
 78  
 79

80 Payne and Anderson (1999) used the functional form of Eq. (2), but substituted  $T_R$  for the  $T_B$ ,  
 81 where  $T_R$  is the empirically calculated approximate temperature of the receiving surface rather  
 82 than the measured body temperature as illustrated in Fig. 1. Thus,  
 83  
 84

$$L = \frac{U_{thermopile}}{C} + \sigma T_R^4 - k\sigma(T_D^4 - T_R^4). \quad (4)$$

85 Payne and Anderson (1999) estimated  $T_R$  using Eq. 5  
 86  
 87

$$T_R = T_B + 0.694 \cdot U_{thermopile} \quad (5)$$

88 where  $U_{thermopile}$  is in millivolts, and the emissivity  $\varepsilon_o$  is set to unity.  
 89  
 90

91 Reda et al. (2002) used a form similar to Eq. (4)  
 92

$$L = k_0 + \frac{U_{thermopile}}{C} + k_2 \sigma T_R^4 - k_3 \sigma (T_D^4 - T_R^4), \quad (6)$$

93 where the instrument body emissivity  $k_2$  is derived during calibration and a constant term  $k_0$  is  
 94 introduced.  $T_R$  is nearly the same as Eq. (5) with 0.704 replacing the constant 0.694. In this paper  
 95 we drop the constant term  $k_0$  since there is no physical justification for including it.  
 96  
 97

98 Since there are four versions of the original Albrecht and Cox (1977) Eq. (1), this paper attempts  
 99 to determine which version is best, in the precision sense, to use for the PIR calibration transfer.  
 100 The organization of this paper is as follows. Because accurate internal thermistor temperatures  
 101 are critical to pyrogeometer IR measurements, we first examine various versions of the standard

102 cubic equation used to convert the YSI 44031 thermistor resistance to the temperatures of the  
103 PIR body and dome. We then calibrate six test PIRs by transferring the calibrations of our three  
104 standard PIRs that were in turn calibrated at the World Radiation Center (WRC) in Davos,  
105 Switzerland, using the World Infrared Standard Group (WISG). Comparisons are then made  
106 between the mean irradiance of the three standard PIRs and the computed irradiance from the  
107 test PIRs using the four different forms of the original Albrecht and Cox (1977) formula, i.e., Eq.  
108 (1) to calibrate each. Boxplots are used to demonstrate the level of agreement between the  
109 standard PIRs and test PIRs for the various formulations. Lastly, a clear conclusion with regards  
110 to the preferred technique to use for calibrations and field measurements is suggested.  
111

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## 113 2. PIR Temperature Measurements

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115 The body and dome temperatures in the Eppley PIR pyrgeometer are measured using the YSI  
116 44031 thermistor. The YSI company provided a table of resistance versus temperature at one K  
117 resolution. To get a finer resolution, a mathematical fit to the tabulated data is required. Steinhart  
118 and Hart (1968) found that a cubic fit of inverse temperature to the log of measured resistance  
119 matched many of the manufacturer's thermistor data points over a wide temperature range. Their  
120 equation is  
121

$$\frac{1}{T} = a + b \cdot \ln(R) + c \cdot (\ln R)^2 + d \cdot (\ln R)^3, \quad (7)$$

122

123 where T is the temperature in K and R is the measured resistance in ohms or kilohms. Note that  
124 in the standard Steinhart-Hart equation, the 'c' is set to zero. Coefficients *a*, *b*, and *d* differ  
125 depending on whether ohms or kilohms are used, and on the temperature range over which the  
126 fit is made.  
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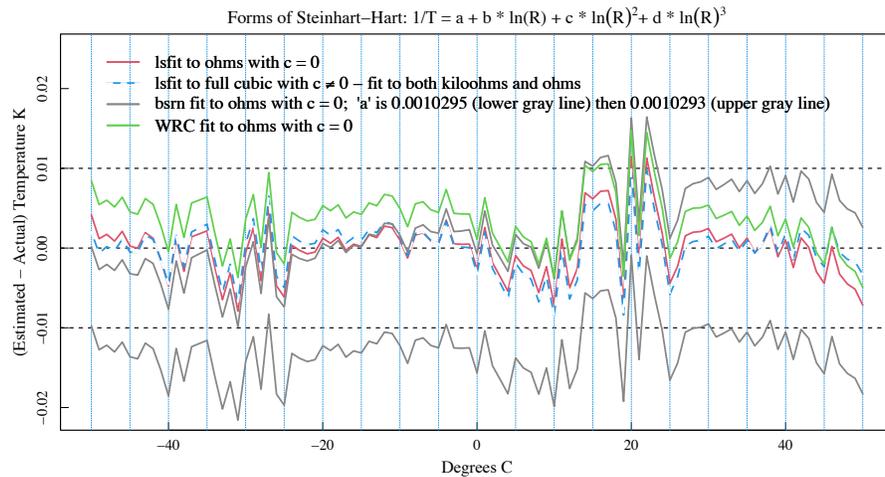
129 Fig. 2 is a plot of six independently-derived fits to the same manufacturer's YSI 44031  
130 thermistor data, and the table below lists the coefficients of Eq. 7 for those fits. The y-axis is the  
131 temperature estimate based on the fit minus the tabulated thermistor data to which the fit is  
132 made. The least-squares fit to Eq. (7) (no quadratic term) is indicated by the red line if ohms are  
133 used. If a full cubic relationship, including a quadratic term, is used to fit the tabulated data in  
134 kilohms, then similar, but not identical, agreement to the red line is obtained (blue-white line).  
135 Interestingly, fitting the full cubic relationship to either ohms or kilohms yields identical results  
136 (again, note the blue-white line). This is not the case if the quadratic term is not included in the  
137 fitting to ohms versus kilohms as discussed in the appendix. Two gray fits in Fig. 2 are from the  
138 Baseline Surface Radiation Network (BSRN) - Operation Manual Version 2.1 (McArthur, 2005),  
139 and use YSI 44031 resistance in ohms. That published equation yields the bottom gray curve that  
140 is displaced from the main grouping in Figure 2. If their coefficient *a* is modified slightly from  
141 the published 0.0010295 to 0.0010293, as shown in the legend, then an improved fit (upper gray  
142 curve) is obtained that agrees well with the others. The green curve, used by the World Radiation  
143 Center (WRC) in Davos, Switzerland, was fit over a -30 to +40 °C range, but does well over the  
144 entire range of -50 to 50 °C considered here. Differences among the various fits in Fig. 2 are  
145 small. All but the bottom gray fit in Fig. 2 cause less than 0.1 Wm<sup>-2</sup> of uncertainty in the  
irradiance estimate.

146

	a =	b =	c =	d =
Red (ohms)	0.001029607	0.0002390769	0.0	1.567609E-07
White (ohms)	0.001020630	0.0002416721	-2.47485E-07	1.64547E-07
Blue (kiloohms)	0.002732470	0.0002618082	3.162474E-06	1.645474E-07
Gray (ohms)	0.0010295 Better 0.0010293	0.0002391	0.0	1.568E-07
Magenta (ohms)	0.00102972	0.00023906	0.0	1.5677E-07

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The larger uncertainty in thermistor temperature measurements is the fundamental accuracy of the thermistors used in the PIR, which are specified to be replaceable to 0.2 K. At 300 K, a difference of 0.2 K is a little over  $1 \text{ Wm}^{-2}$ . In this paper the full cubic (blue-white curve) is used to compute PIR temperatures in the analyzed data.



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**Figure 2.** Six independent fits using forms of Eq. (7) to the YSI 44031 tabulated data after subtraction of the tabulated data over the range -50 to 50 °C in 1 °C increments. Similar agreement among all fits ensues if the small change to the published BSRN constant  $a$  is made. The full cubic fits overlap whether ohms or kiloohms are used (with different coefficients, of course).

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### 3. Four Methods of PIR Calibration Transfer

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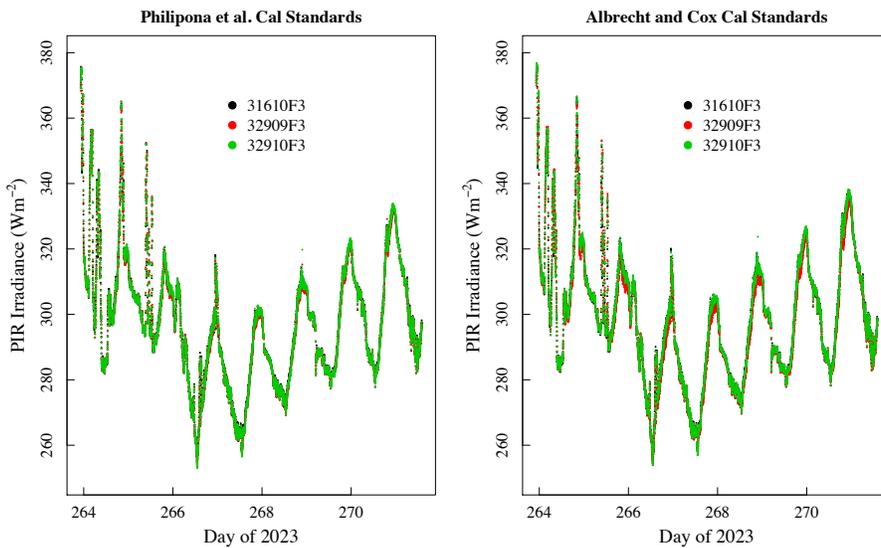
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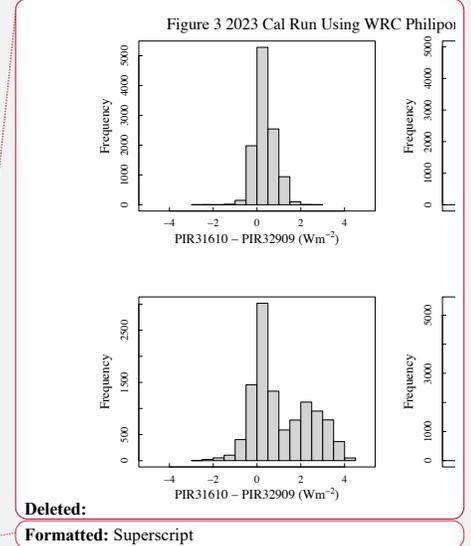
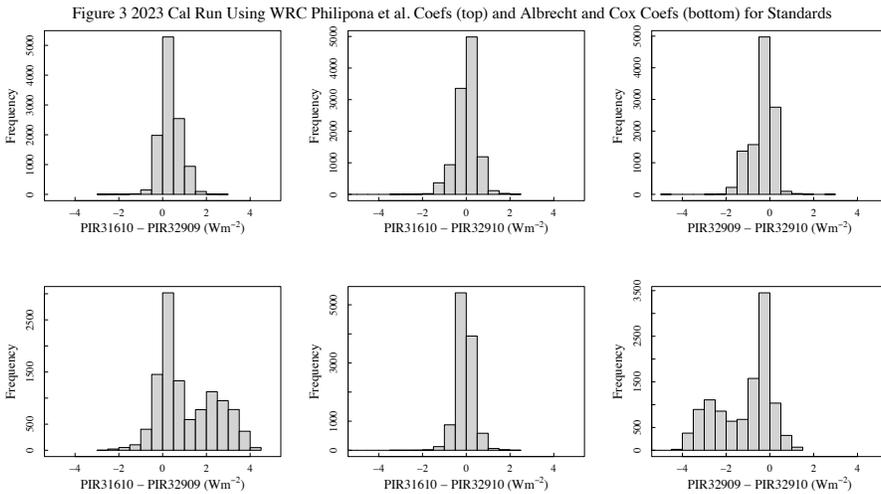
In this section, we examine the performance of Eq. (2), (3), (4), and (6) in transferring calibrations from our three “standard” PIRs to field PIRs. Our standard PIRs were calibrated against the world reference at the WRC (<https://www.pmodwrc.ch/en/world-radiation-center-2/irs/wisg/>) in 2018, 2022, and 2024 at the WRC, which is part of the PMOD in Davos, Switzerland. Each standard was returned with two sets of coefficients, one set for Eq. (2) and one set for Eq. (3). The WRC calibration of our standard PIRs uses their blackbody over a range of

168 pyrgeometer and cavity temperatures to determine the  $k_i$ 's in Eq. (3). The C determined by this  
169 regression is not used, but is set using clear and stable skies outdoors.

170  
171 To transfer the standard PIRs' calibration to field radiometers, our WRC-calibrated standard  
172 PIRs and test PIRs are arranged side-by-side for a week or more on an outdoor horizontal  
173 observing platform, with no significant obstructions surrounding the platform. Note that all  
174 calibration coefficients of Eq. (3) for our test PIRs are obtained by regression analysis, i.e., we  
175 used no blackbody for calibration transfer. For this paper two calibration periods and three PIRs  
176 from each are analyzed. However, all figures for this paper only display results from the 2023  
177 calibration performed at Table Mountain near Boulder, Colorado USA. Diurnal variability  
178 shown in Fig. 3(a) indicates the type of conditions used for calibration. On the left are the three  
179 standard PIR's outputs with the WRC's Philipona et al. (1995) coefficients applied. On the right  
180 are the same standard PIR's with the WRC's Albrecht and Cox (1977) coefficients applied.  
181 Agreement among the three on the left is very good because the last-plotted PIR readings (green)  
182 overplot the first two (black and red). Agreement on the right is nearly as good but with some  
183 underestimation by PIR 32909F3 (red dots). Histograms in Fig. 3(b) indicates the degree of  
184 agreement among our standard PIRs more clearly with closer agreement among the results using  
185 the Philipona et al. (1995) coefficients.  
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188 **Figure 3(a).** Calculated IR irradiance from our three standard PIRs (serial numbers in the  
189 legend) using Philipona et al. coefficients provided by WRC are overplotted on the left and the  
190 Albrecht and Cox coefficients are used on the right. This demonstrates the agreement among the  
191 standard PIRs using the two methods. The number of days used is typical for our calibration  
192 runs.  
193



194  
 195 **Figure 3(b).** Histograms of the differences ( $Wm^{-2}$ ) among our standard PIRs applying either the  
 196 Philipona et al. (1995) WRC-assigned calibration coefficients (top) or the Albrecht and Cox  
 197 (1997) WRC-assigned coefficients (bottom). Clearly, there is closer agreement in top row.  
 198

199 Before comparing results from Eq. (2), (3), (4), and (6), we first compare results from only  
 200 Albrecht and Cox (Eq. 2) and Philipona et al. (Eq. 3) for which the WRC provided both sets of  
 201 coefficients. In this test, the mean IR irradiance of the three standard PIRs is compared to  
 202 computed IR from a test PIR that was calibrated using the mean of these standard PIRs. The  
 203 least-squares fitting technique to determine the calibration coefficients for the test instrument  
 204 uses a robust function in the R language (MASS::rlm) that de-weights outliers to reduce the  
 205 effects of noisy, for example, rain-contaminated, and other outlier data. As we shall discuss in  
 206 the final section, this appears to be comparable to the strict criteria used at the WRC for  
 207 calibration transfer. However, it should be noted that days with known rain events are removed  
 208 from all test data sets before a calibration transfer is attempted.  
 209

210 In Figure 4, boxplots are used to compare the performance of the two calibration methods  
 211 applied to the standard PIRs at the WRC, i.e., Albrecht-and-Cox (1977) and Philipona et al.  
 212 (1995). The “box” in these plots contain 50% of the data, and the lines extending from the top  
 213 and bottom of the box, or “whiskers,” include about 95% for normally distributed data. In the  
 214 top-left panel of Fig. 4 the three standard PIRs use the WRC-provided Albrecht and Cox (1977)  
 215 coefficients, and the average of the three standard PIRs is compared to coincident test PIR (SN  
 216 23215F3) data, also calculated using Albrecht and Cox (1977). The boxplot summarizes those  
 217 differences over the entire calibration period for PIR 23215F3. The boxplot on the top right  
 218 summarizes differences following the same procedure, but using Philipona et al. (1995)  
 219 coefficients for the standard PIRs (WRC-provided) and for the test PIR. Comparing the top  
 220 panels of Fig. 4, the one on the right, where Philipona coefficients are used exclusively, has a

222 smaller box, shorter whiskers, and a median nearer to zero compared to the panel on the left  
223 where Albrecht and Cox was used exclusively.

224

225 The bottom panels of Fig. 4 show the same comparison for a different test PIR (SN 38805F3).

226 The same comments apply, with the Philipona et al. (1995) calibrated data (bottom right) giving

227 smaller spread in the box and whiskers, and the median nearer zero, while there is more spread in

228 the bottom-left panel where Albrecht and Cox (1977) is used. Differences in the lower panels of

229 Fig. 4 are generally greater than those in the top panels. The calibration data for these two test

230 instruments were collected concurrently, which suggests that the disparity arises from inherent

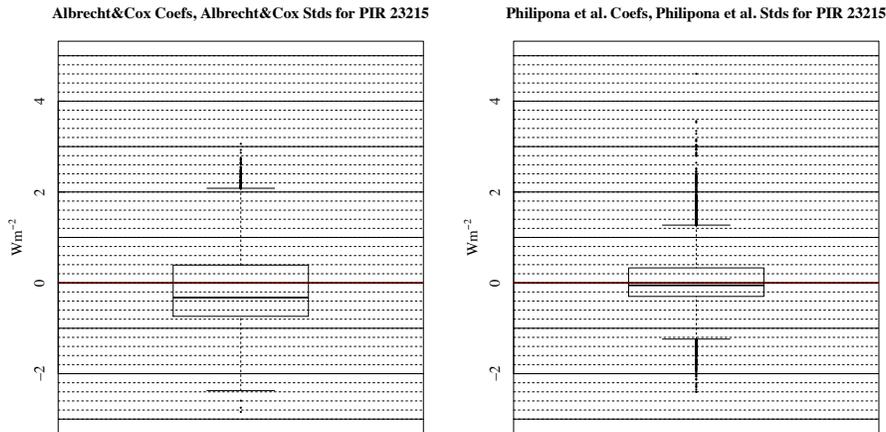
231 characteristics of the instruments themselves. We studied a total of six instruments from two

232 distinct calibration periods in this way and found that in every case using the Philipona et al.

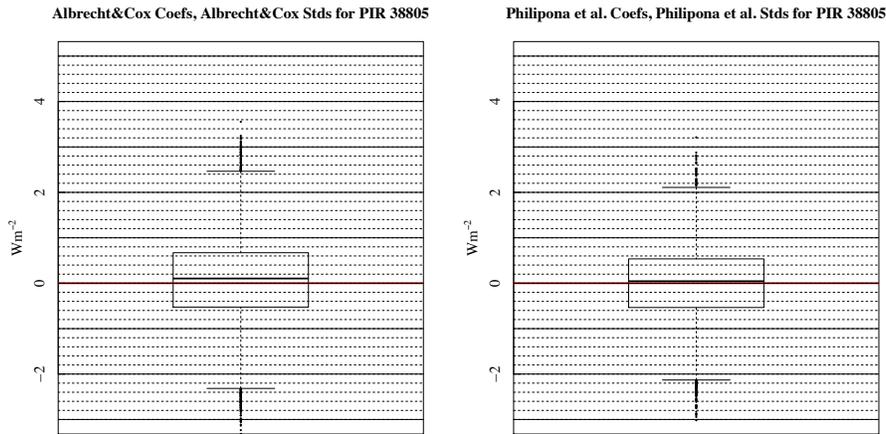
233 (1995) form (Eq. 3) gave better results than the formulation of Albrecht and Cox (1977) (Eq.

234 (2)).

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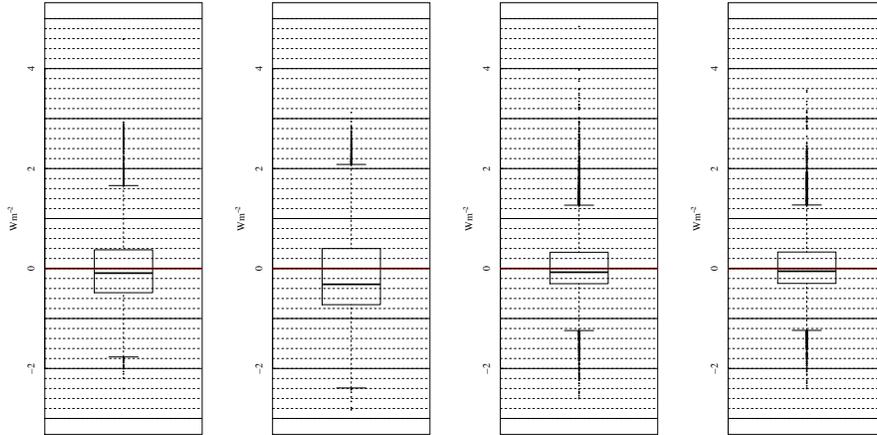


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 239 **Figure 4.** (top) Boxplots of the differences between applying Albrecht and Cox (1977)  
 240 calibrations and applying Philipona et al. (1995) calibrations for PIR 23215. Note the differences  
 241 in box widths, whisker lengths, and median values. (bottom) Boxplots for a different PIR  
 242 (38805) that was calibrated at the same time as the one in Fig. 4 (top).  
 243  
 244  
 245  
 246 Next, we compare results from all four equations (2), (3), (4), and (6) for the same two PIRs as in  
 247 Fig. 4. Since Fig. 4 suggests that the Philipona et al. (1995) Eq. (3) produces better results than  
 248 Albrecht and Cox (1977) Eq. (2), we use Philipona et al. (1995) coefficients provided by WRC  
 249 to compute IR irradiance for the standard PIRs and average these as “truth” for all of the  
 250 comparisons. For both test PIRs in Fig. 5 (top and bottom) the last boxplot on the right  
 251 (Philipona et al., 1995) gives the best results followed by the adjacent boxplot (Reda).  
 252  
 253

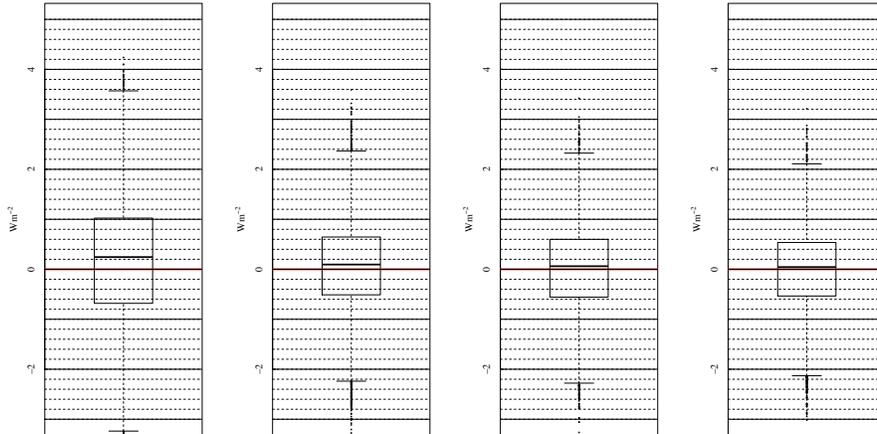
Left to Right, Albrecht Coefs, Payne Coefs, Reda Coefs, Philipona Coefs Using Philipona Stds for PIR 23215 2023 Cal



Albrecht = Albrecht and Cox (1977); Payne = Payne and Anderson (1999); Reda = Reda et al. (2002); Philipona = Philipona et al. (1995)

254

Left to Right, Albrecht Coefs, Payne Coefs, Reda Coefs, Philipona Coefs Using Philipona Stds for PIR 38805 2023 Cal



Albrecht = Albrecht and Cox (1977); Payne = Payne and Anderson (1999); Reda = Reda et al. (2002); Philipona = Philipona et al. (1995)

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256

257 **Figure 5.** Boxplots of differences using the WRC's Philipona et al. (1995) coefficients for the  
 258 standard PIRs and calibrated PIRs to this standard using the four equations to calculate incoming  
 259 IR. Top is for PIR 23215 and bottom is PIR 38805 as in Fig. 4. Compare box widths, whisker  
 260 lengths, and medians.

261

262 In these comparison plots the standard used was calibrated with Philipona et al. (1995)  
 263 coefficients rather than Albrecht and Cox (1977) as in Fig. 4. Similar results were obtained for

264 the other four PIRs with the best results always obtained with the Philipona et al. (1995)  
265 formulation. In only one case out of the six the Payne and Anderson (1999) formula performed  
266 slightly better than the Reda et al. (2002) formula (not shown).  
267

268 T-tests were performed to assess differences when using non-Philipona et al. (1995) calibration  
269 coefficients for all six calibrated instruments. If one assumes that there are no significant  
270 differences in the calculation of IR irradiances using the Philipona et al. (1995) formula versus  
271 each of the other three methods discussed, this assumption is rejected with 95% confidence in 15  
272 of the 18 cases studied (six calibrated PIRs and three formulae). The three cases where the null  
273 hypothesis cannot be rejected with 95% confidence are for three of the six PIRs using the Reda  
274 et al. (2002) formula.  
275

276 Reda et al. (2002) and Payne and Anderson (1999) did not use the measured body temperature  
277  $T_B$  in their formulae, but estimated the receiver temperature  $T_R$  using a form of Eq. (5) for their  
278 particular PIR configuration. As a test we replaced  $T_B$  with  $T_R$  in the Philipona et al. (1995) Eq.  
279 (3). The extremely small changes in the rightmost boxplots of Fig. 5 were imperceptible. We  
280 would, therefore, suggest keeping Eq. 3 in its original form for calibration transfer.  
281

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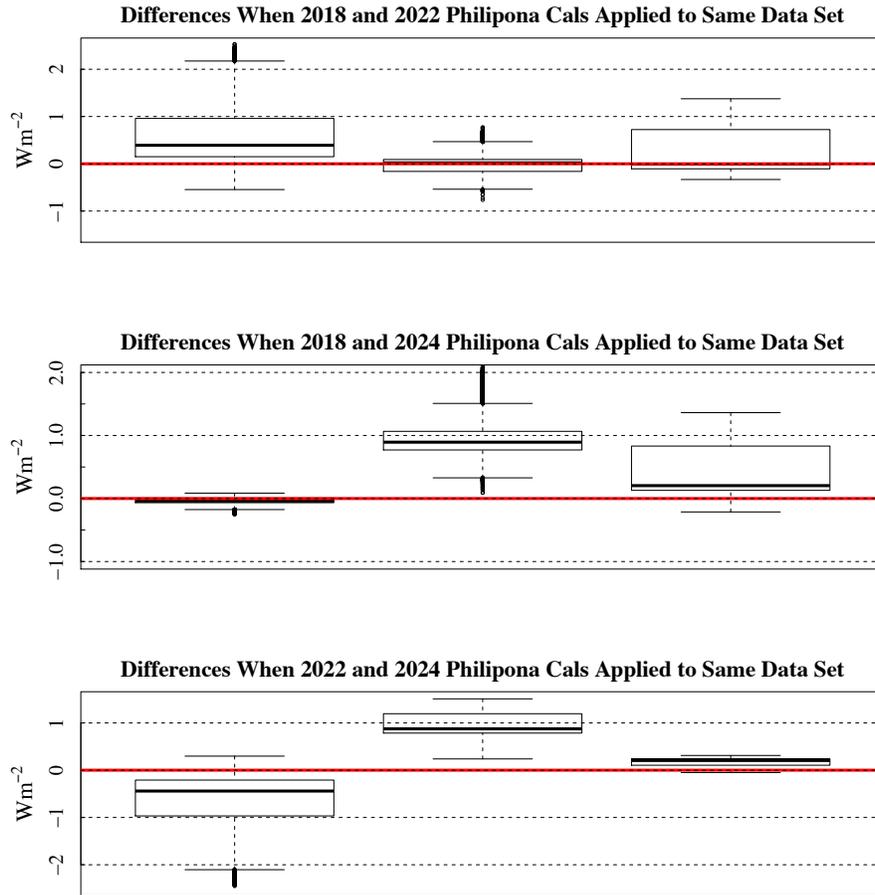
#### 283 4. Precision of the PIR Standards

284

285 The calibration of our same three standard PIRs at the WRC leads to slightly different calibration  
286 results. Here, the consistency and repeatability of those calibration events is assessed. The PIRs  
287 that we use for standards were sent to WRC in 2018, 2022, and 2024. For each of those events,  
288 coefficients for the Albrecht and Cox (1977) and Philipona et al. (1995) forms of the PIR  
289 processing equation for calculating incoming IR were provided by the WRC.  
290

291 Our calibration seasons typically run from late Spring to early Fall. Therefore, our three PIR  
292 standard PIRs experience roughly six months of exposure to the weather each year. In Fig. 6  
293 differences from applying three sets of WRC Philipona calibration coefficients (from 2018,  
294 2022, and 2024) to the same dataset (that used for Fig. 3) are summarized. For example,  
295 calibrations from 2018 and 2022 were applied to the same dataset and differences in irradiance  
296 for each minute were tallied and summarized in boxplots. Differences for all permutations are  
297 mostly within  $1 \text{ Wm}^{-2}$  and suggest that errors from applying one of the WRC calibrations from  
298 any of the three calibration years to any year would be less than the uncertainty of the WRC  
299 calibrations themselves ( $\sim 4 \text{ Wm}^{-2}$ ; <https://www.pmodwrc.ch/en/?s=wisg>). This suggests that the  
300 Eppley PIR is very stable and should be suitable for monitoring long-term changes in the thermal  
301 IR.  
302

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Three Standards (31610, 32909, 32910) w/ WRC Cals from Three Different Years

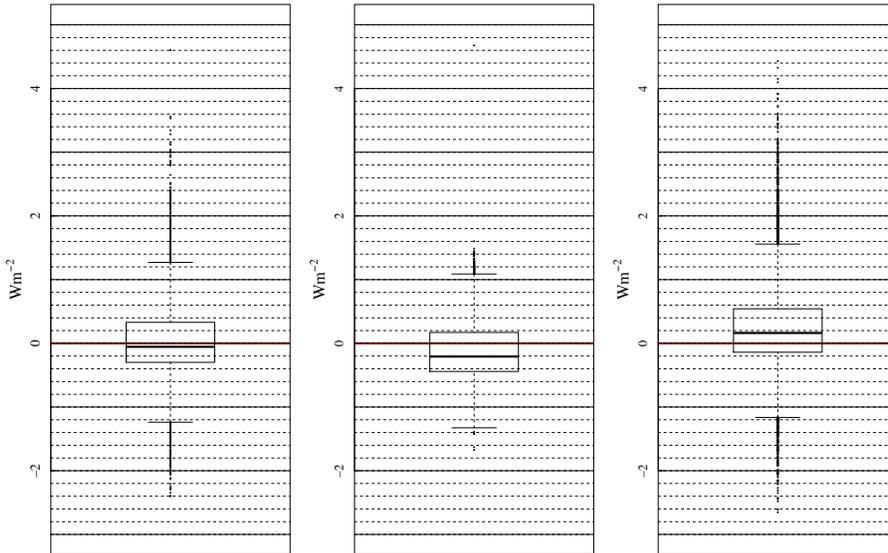
304  
 305 **Figure 6.** Comparisons of three sets of Philipona et al. (1995) calibration coefficients provided  
 306 by the WRC in 2018, 2022, and 2024 applied to the same data set as in Fig. 3 for the three PIRs  
 307 used as standard PIRs with serial numbers in the subtitle. The medians are all within  $1 \text{ Wm}^{-2}$  and  
 308 most are within  $0.5 \text{ Wm}^{-2}$ .

309  
 310 In section 3 the average output of the three standard PIRs is used to derive new calibration  
 311 coefficients for each test PIR. Using those new calibrations, the test instrument measurements  
 312 are compared to the standard PIRs' average over the entire calibration period. For the left panels  
 313 in Fig. 7 we use Philipona et al. (1995) coefficients for the standard PIRs to calibrate the three  
 314 test PIRs (serial numbers shown at the top of each subplot). We apply those new calibrations and

315 subtract the results from the standard PIRs' average for each minute and summarize the  
316 distribution of differences in boxplots. Therefore, the leftmost panels of Fig. 7 replicates the  
317 rightmost panels of Figs. 4 and 5. This is not an independent test of the reliability of the  
318 calibration because the same dataset is used for calibration and verification.

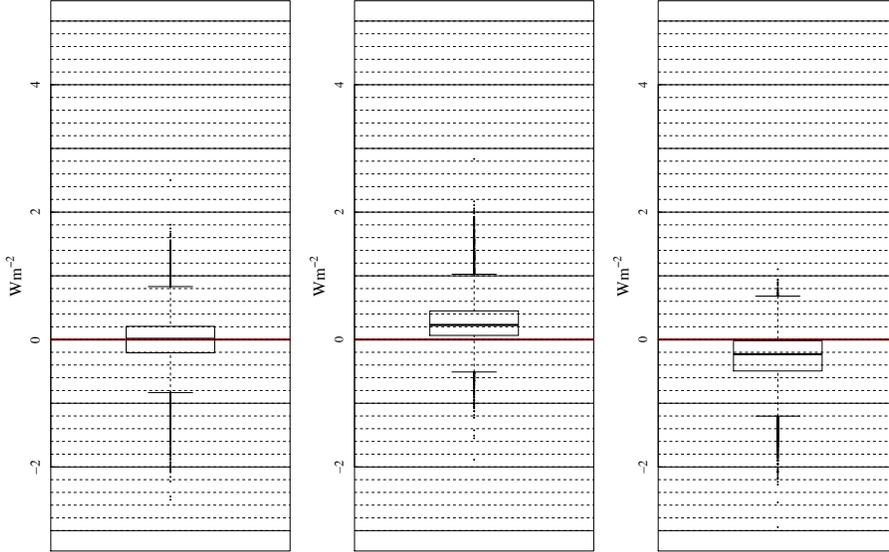
319  
320 To test new calibrations with an independent dataset, the time series in Figure 3 is divided in  
321 half. The middle panels of Fig. 7 use the first half of the data in Fig. 3 to derive a calibration and  
322 the second half of the data to validate the new calibration against the standard PIRs' average.  
323 Then, we reverse this process using the second half of the Fig. 3 data for calibrating and the first  
324 half to validate. If we examine the time series in Fig. 3, it is apparent that the first half of the data  
325 stream is noisier than the second half. Using the first half of the data to calibrate and applying to  
326 the second half and vice versa is likely responsible for the offsets in the medians, but the offsets  
327 are less than one  $Wm^{-2}$ . Note that when the less noisy data of the second half are used to validate  
328 (middle boxplots) the differences have a smaller spread. When the noisier first half data  
329 (rightmost boxplots) are used to validate, the differences have a larger spread. Examining the  
330 top, middle, and bottom plots, there are differences inherent in the instruments themselves since  
331 boxplots are not replicated from PIR to PIR. Attribution to the instruments themselves is  
332 warranted because the standard PIRs and test data used for Fig. 7 were collected simultaneously.  
333

Left to Right: Full Cal Period; 1st Half Used for Cal, Applied to 2nd Half; and Vice Versa for PIR 23215



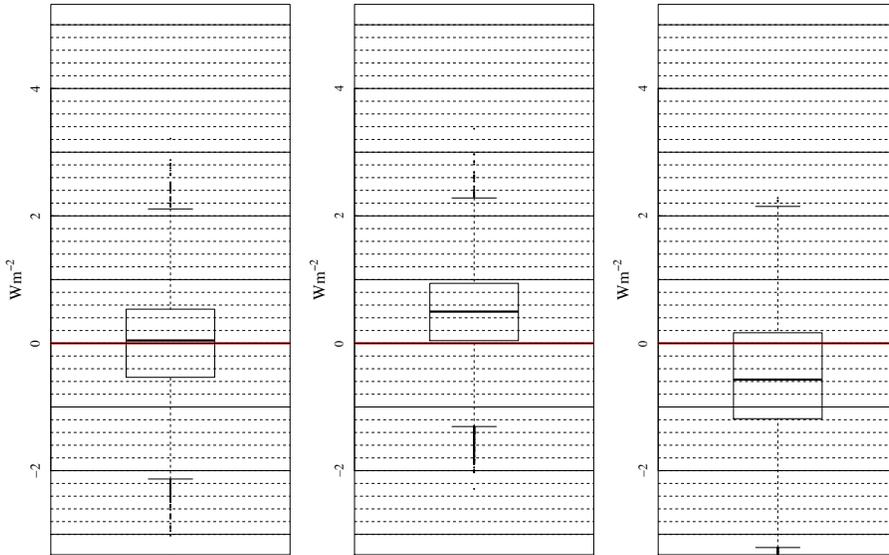
334

Left to Right: Full Cal Period; 1st Half Used for Cal, Applied to 2nd Half; and Vice Versa for PIR 28139



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Left to Right: Full Cal Period; 1st Half Used for Cal, Applied to 2nd Half; and Vice Versa for PIR 38805



336

337 **Figure 7.** The leftmost panel uses the entire period in Fig. 3 to calibrate the named PIR and then  
338 compares the calibrated PIR data to the standard PIRs' average. The middle panel uses the first  
339 half of the period to calibrate and then assesses the application of those calibrations to an  
340 independent data set in the second half. The rightmost panel reverses this using the second half  
341 of the period for calibration and the first half for assessment.

## 342 5. Summary and Conclusions

343  
344 In this paper we investigate four formulations for converting raw voltage and body and dome  
345 temperature measurements of an Eppley pyrgeometer, model PIR, to thermal IR irradiance.  
346 These methods are described in Albrecht and Cox (1977), Philipona et al. (1995), Reda et al.  
347 (2002) and Payne and Anderson (1999). All are slight variations of the original formulation of  
348 Albrecht and Cox (1977). Because the temperature measurements are critical to the infrared  
349 calculations, we also investigated various fits that have been applied to the Steinhart-Hart (1968)  
350 equation that converts thermistor-measured resistance to temperature.

351  
352 Regarding the computation of thermistor temperatures, we found that fitting the manufacturer-  
353 supplied table of resistance and temperature (1°C interval) to the range -50° to 50°C provides the  
354 least variability as opposed to fits to shorter temperature ranges. However, differences of the fit  
355 to the provided data are < 0.01°C, regardless of the range used. Based on this result, we conclude  
356 that differences in thermistor temperature calculations from fits based on various temperature  
357 ranges do not have a significant impact on PIR measurements.

358  
359 The three standard PIRs that we use to transfer calibrations from the world standard to field PIRs  
360 are calibrated frequently against the World Infrared Standard Group (WISG) at the World  
361 Radiation Center in Davos, Switzerland. They are returned with calibration coefficients for the  
362 Albrecht and Cox (1977) and Philipona (1995) methods, although the Albrecht and Cox  
363 coefficients provided are for the shortened form of their equation (Eq. 2). Comparing the  
364 application of the two methods to the standard PIRs revealed that the Philipona (1995) method is  
365 more precise and less noisy than the Albrecht and Cox formulation; the differences are quantified  
366 in Fig. 3(b) histograms. Comparisons were also made among three distinct WRC calibration  
367 results for the standard PIRs in 2018, 2022, and 2024. They showed that the three standard PIRs  
368 are stable, with the calibration coefficients changing minimally between WRC calibrations, and  
369 differences in irradiance calculations among applications of the separate biennial calibrations are  
370 within one  $Wm^{-2}$  of each other.

371  
372 Application of the four methods for converting PIR raw measurements to irradiance was  
373 analyzed using six test instruments. The major conclusion is that use of the Philipona et al.  
374 (1995) form, i.e., Eq. (3), consistently does the best in transferring the mean calibration of the  
375 standard PIRs to field-deployed PIRs. Note that Reda et al. (2002) and Payne and Anderson  
376 (1999) coefficients are not available for the standard PIRs calibrated at the WRC, which may  
377 have led to some of the differences in Fig. 5. Of the six calibration comparisons, like those in  
378 Fig. 5, Reda et al. (2002) calibration results were close to Philipona et al. (1995) results on three  
379 of the six PIRs according to t-tests performed at the 95% level. However, this agreement was  
380 found to be insignificant for the t-tests on the other three PIRs.

381  
382

383 Given the differences in Figs. 6 and 7, it is probable that there is greater uncertainty caused by  
 384 the particular atmospheric conditions under which calibrations are carried out. With the  
 385 assumption that the PIR is very stable, the variations among the instruments in Fig. 6 could be  
 386 subtle differences in atmospheric conditions during the three calibration sessions at the WRC in  
 387 2018, 2022, and 2024. This is reinforced by the differences in Fig. 7, where independent stable  
 388 (i.e., clear), and unstable (e.g., intermittent clouds) periods were used to calibrate test  
 389 instruments, with differing results. Note that PIR measurements for any arbitrary weather  
 390 condition are often going to have larger uncertainties than discussed here.  
 391

392 The WISG, which is used for calibration at the WRC, is the current standard for broadband IR  
 393 measurements. It has an uncertainty of  $2.6 \text{ Wm}^{-2}$ . Recent studies, which are summarized in  
 394 Gröbner et al. (2024), further suggest that the current WISG may be low by as much as  $4 \text{ Wm}^{-2}$  if  
 395 the water vapor column exceeds 1 cm, but the difference is smaller if the atmosphere is dryer  
 396 approaching no difference for vanishing water vapor (see Fig. 2 in Gröbner et al., 2024).  
 397 Nevertheless, a new standard for broadband IR radiation is not expected to be established until  
 398 the next WMO congress in 2027 at the very earliest (Laurent Vuilleumier, private  
 399 communication).

400

#### 401 **Appendix A1**

402

403 Fitting the manufacturer-supplied temperature (at 1 °C intervals) and resistance data, separately  
 404 in ohms and in kilohms, led to an unexpected outcome. First, if a full cubic (i.e., non-zero  
 405 coefficient for the squared term) least-squares fit of the Steinhart-Hart equation with YSI 44031  
 406 data in kilohms is compared to a least-squares fit using ohms, identical fits are obtained (blue  
 407 and white dashed line in Fig. 2). If the quadratic term is set to zero and the fits are made to ohms  
 408 and then kilohms, we see a significant difference as shown in Fig. A1. This difference is due to  
 409 numerical reasons, which are explained in the following.  
 410

411

412 First, it must be noted that the lack of significant digits when using kilohms is not an issue  
 413 because for the fits here, kilohms are computed simply by dividing the resistance value in ohms  
 414 by 1000, keeping significant digits in the decimal places.  
 415

416

417 The requirement of a quadratic term for expressing the Steinhart-Hart equation in kilohms can  
 418 be demonstrated by substituting for  $R$  in Eq. (7)  $1000R_k$ , where  $R_k$  is in units of kilohms as  
 419 shown in Eq. (A1).

420

$$\frac{1}{T} = a + b \cdot \ln(1000R_k) + d \cdot \ln(1000R_k)^3 \quad (\text{A1})$$

421

422 Applying logarithm rules to Eq. (A1) results in Eq. (A2).

423

$$\frac{1}{T} = a + b(\ln(1000) + \ln(R_k)) + d(\ln(1000) + \ln(R_k))^3 \quad (\text{A2})$$

424

425 Expanding and regrouping terms in Eq. (A2) then gives Eq. (A3) through Eq. (A7).  
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$$\frac{1}{T} = a_k + b_k \cdot \ln(R_k) + c_k \cdot \ln(R_k)^2 + d_k \cdot \ln(R_k)^3 \quad (A3)$$

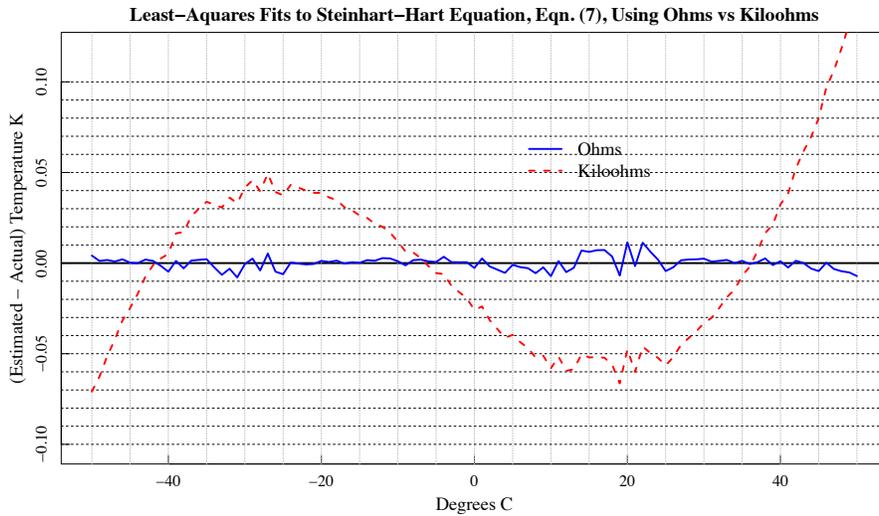
$$a_k = a + b \cdot \ln(1000) + d \cdot \ln(1000)^3 \quad (A4)$$

$$b_k = b + 3d \cdot \ln(1000)^2 \quad (A5)$$

$$c_k = 3d \cdot \ln(1000) \quad (A6)$$

$$d_k = d \quad (A7)$$

Thus, when data are in kiloohms an equation of the form of Eq. (A3) (i.e. full cubic) is required to match the results of Eq. (7) when data are in units of ohms. Thus, changing units of  $R$  in Eq. (7) results in a full cubic equation. This implies that a full cubic equation can be more robust than Eq. (7) when fitting data where units other than ohms are used for  $R$ . It also demonstrates that it is possible to change units for  $R$  in Eq. (7) analytically using the substitution process shown above rather than refitting if desired.



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**Figure A1.** Steinhart-Hart equation (i.e., no quadratic term) fit to ohms (blue solid line) versus kiloohms (red dashed line).

454 *Code availability.* Codes used to generate the results in this paper were original functions written in the  
455 programming language R and are available by contacting [joseph.michalsky@noaa.gov](mailto:joseph.michalsky@noaa.gov).

456 *Data availability.* Data can be made available by contacting [joseph.michalsky@noaa.gov](mailto:joseph.michalsky@noaa.gov).

457 *Author contributions.* JJM did most of the analyses, drafted the paper, and produced the figures. JAA  
458 provided the World Radiation Center calibrations and much useful discussion of the results. EH provided  
459 the experimental data from the calibration table used for these analyses. BRS did the analysis for and  
460 wrote the Appendix. All authors read and offered corrections to parts of the manuscript.

461

462 *Competing interests.* The contact author has declared that none of the authors has any competing interests.

463

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468

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