

# Simulation of a contrail formation and early life cycle for a realistic airliner geometry

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**Abstract.** Contrails—ice clouds that form in aircraft wakes—are thought to have a radiative impact up to twice that of CO<sub>2</sub> emissions accumulated by aviation, although this estimate remains debated due to significant uncertainties. These uncertainties underline the need for further research into the entire life cycle of contrails, from the formation of initial ice crystals to their potential evolution into persistent cirrus clouds. The challenge lies in the wide range of spatial and temporal scales involved in contrail development.

This work presents a novel numerical methodology for simulating contrails from the onset of ice crystal formation to the dissipation of wingtip vortices, that is to say a few minutes after the formation of the first ice crystals. Unlike conventional methods that rely on analytical initialization of a far-field simulation using a pair of Lamb-Oseen vortices, our approach couples Reynolds-averaged Navier–Stokes (RANS) with Large Eddy Simulation (LES) and synthetic turbulence techniques. This allows for a more accurate capture of near-field effects and a detailed consideration of how aircraft geometry influences the aerodynamic wake and subsequent contrail evolution.

Applied to a realistic aircraft geometry under standard atmospheric conditions, our methodology revealed that horizontal tailplane vortices can trigger short-wavelength instabilities in wingtip vortices, significantly modifying the structure of the secondary wake. Comparisons with conventional methods show that the contrails generated through this work methodology are wider and more opaque, at least in the first few minutes of the contrail lifecycle. A modified version of the conventional methods, incorporating a quadripolar wake—comprising two main wingtip vortices and two secondary vortices from the horizontal tail plane—is proposed. This new approach more closely matches simulations obtained through this work methodology and that account for near-field effects.

## 1 Introduction

Contrails, short for condensation trails, are the long ice clouds visible in the wake of airplanes on clear days. Those anthropogenic clouds result from the condensation of the water vapor located in the gas mixture between exhaust gases emitted from aircraft engines and cold ambient air. In ice-supersaturated regions of the troposphere, contrails can evolve into cirrus clouds and persist for several hours over hundreds of kilometers before dissipating, thereby increasing regional cloudiness in the meantime. Like any cloud, cirrus clouds induced by contrails interact with both solar and terrestrial infrared radiation,

25 thus impacting Earth's radiative balance. The effect of such artificial clouds on Earth's radiative balance has been quantified in a recent compilation of studies Lee et al. (2021). This work enabled the estimation of contrails' Effective Radiative Forcing (ERF) and other forcing agents such as  $\text{CO}_2$ ,  $\text{NO}_x$ , aerosols, and water vapor. The ERF of cirrus clouds induced by contrails has been estimated at  $57.4 \text{ mW} \cdot \text{m}^{-2}$  which is almost twice as much as the ERF of  $\text{CO}_2$ , estimated at  $34.3 \text{ mW} \cdot \text{m}^{-2}$ .

30 However, the estimated value of contrails' ERF is currently debated due to the high uncertainty surrounding this value. This uncertainty primarily reflects the difficulty in accurately estimating the optical properties of contrails, which depend on their microphysical properties (ice crystal concentration, crystal size distribution, etc.), as well as their vertical and horizontal extent. Moreover, the microphysical properties of a contrail several hours old are strongly influenced by these same microphysical properties in the initial moments (see Lewellen and Lewellen (2001); Unterstrasser and Görsch (2014)). During these initial moments, that is, within the first minutes of its lifecycle, the evolution of contrails results from the interaction between wake 35 aerodynamics, plume chemistry, and microphysical processes. This interaction and the underlying physics must be correctly understood to accurately predict contrails' radiative impact.

The formation and evolution of contrails can be divided into four successive regimes (see Gerz et al. (1998); Paoli et al. (2013); Misaka et al. (2015); Jacquin and Garnier (1996)), each governed by specific physical phenomena. The *jet regime*, lasting only a few seconds (see Jacquin and Garnier (1996)), is dominated by strong turbulent diffusion in the engine jet while 40 the wing's vortex sheet rolls into a pair of counter-rotating wingtip vortices. As the jet plume cools down, ice crystals begin to form, making the contrail visible at about one to two wingspans behind the aircraft (see Khou et al. (2015); Ramsay et al. (2024)). In the case of soot-rich emissions, that is to say for a soot emission index greater than  $5 \times 10^{14} \text{ kg}^{-1}$  as stated in Yu et al. (2024), ice crystals primarily form through the condensation of water vapor onto soot particles. Conversely, if soot 45 concentration is low in the plume, ice crystals predominantly form on volatile particles that have formed earlier in the plume, such as sulfuric acid droplets (see Kärcher (1998)). Interaction between the jet and vortices becomes significant around ten wingspans for a typical airliner and can be quantified by looking at the  $R_3$  ratio defined in Jacquin and Garnier (1996).

The *vortex regime* begins when the plume's evolution is dominated by the action of wake vortices, action characterized by the plume roll-up around the wingtip vortices. As time passes, the vortex pair descends due to mutual induction, carrying a significant fraction of the ice crystals downward. This descent results in the heating of those crystals, potentially leading 50 to their sublimation. However, the vortex pair interacts with atmospheric stratification, generating baroclinic vorticity and a secondary wake that redistributes some ice crystals back to flight altitude and further as observed in Unterstrasser (2014) and where ambient conditions are more favorable for crystals growth. After several tens of seconds, Crow instability, which was theorized in Crow (1970); Crow and Bate Jr (1976), deforms the vortex pair, marking the transition to the dissipation regime. This regime, lasting a few minutes, ends with the destruction of the vortices and the formation of a chain of vortex 55 rings periodically distributed along the flight direction and that can often be observed in the sky on clear days. This marks the beginning of the *diffusion regime*. It is important to note that under strong atmospheric stratifications, short-wave vortex instabilities may dominate and destroy the vortices before Crow instability develops, preventing the formation of vortex rings. During the *diffusion regime*, contrail evolution is governed by a certain number of physical processes which are affected by atmospheric conditions such as atmospheric turbulence, wind shear, and stratification. If the air remains supersaturated with

60 respect to ice, the contrail may transform into cirrus clouds, potentially persisting for hours. This happens in the so-called ice super-saturated regions (ISSR) that can frequently be found in the tropopause. A significant part of the *diffusion regime* was previously simulated by Paoli et al. (2017) through a 3D temporal LES approach, by Lewellen (2014) with a Quasi-3D approach and by Unterstrasser (2014) through a 2D approach.

65 Simulating a contrail over its entire lifecycle with a satisfactory level of fidelity presents a significant challenge. This is primarily due to the wide range of physical phenomena and scales involved in the four regimes making up the life of a contrail. Directly simulating all the scales of a contrail while accounting for its full spatiotemporal development is currently prohibitive in terms of computational cost. Nevertheless, understanding the formation and the evolution of a contrail is crucial to developing mitigation techniques, such as the use of sustainable aviation fuels (SAF) or hydrogen fuel that leads to a significant reduction of nucleation sites (see Voigt et al. (2021)). A commonly adopted solution in the literature consists of performing 70 so-called temporal LES for which periodic boundary conditions are prescribed perpendicularly to the flight direction. The temporal development of a contrail in the fixed ground reference frame can then be computed. Such an approach is valid at a distance from the airplane when jet/vortex interaction becomes significant. The problem that arises then is how to initialize such temporal simulations. In most studies in the literature, a proposed solution is to initialize the wake with analytical solutions modeling a pair of vortices and jet plumes. For instance, the vortices are often defined using the Lamb-Oseen solution while 75 turbulence inside the jet is randomly generated with a normal distribution as in Holzäpfel et al. (2001); Paoli et al. (2013); Unterstrasser (2014); Unterstrasser and Görsch (2014); Lewellen and Lewellen (2001). Such an approach simplifies the aircraft wake by idealizing it. Consequently, some aspects of the near-field aerodynamics are neglected, and their potential effect on the far-field is not taken into account. The hypothesis that the effects of the near-field on the far-field can be neglected must be examined. This is all the more important given the increase in geometrical complexity of future aircraft concepts, such as 80 blended wing bodies and rhomboidal wings.

Recently, a methodology has been developed in Bouhafid et al. (2024) in order to simulate the evolution of an aircraft wake from vortex formation, up to their dissipation several minutes later. This methodology was successively used to bridge the gap between the near-field aerodynamics of the jet regime and the far-field aerodynamics of the vortex and dissipation regimes without using analytical solution initialization for the temporal LES of the wake. This allowed for the full effects of the aircraft 85 geometry on the wake aerodynamics to be taken into account. The same study also compared the wake dynamic when their methodology is used compared to the usual analytical initializations. This comparison has shown significant impact of the horizontal tailplane vortices on the wake evolution and its interaction with the plume.

Therefore, the objective of this work is to apply the methodology developed in Bouhafid et al. (2024) to simulate a contrail from the formation of the first ice crystals to the end of the dissipation regime and to compare the results with the usual 90 analytical initialization in order to see how the differences observed in wake dynamic impact the properties of the contrail in its early life, that is to say before contrail-to-cirrus transition. In addition, the influence of horizontal tailplane vortices on contrails is investigated.

The outline of this paper is as follows. Section 2 describes the methodology and the corresponding numerical framework for simulating a contrail from the jet regime to the dissipation regime. Section 3 gives details about the different simulation

95 setups and the corresponding numerical parameters. Section 4 presents the results of the methodology applied to the NASA  
Common Research Model (CRM) geometry in cruise flight conditions, geometry representative of Boeing B777 airliner. An  
examination is conducted regarding the influence of the far-field initialization strategy. Finally, Section 5 concludes this work  
and highlights its different perspectives.

## 2 Methodology and numerical framework

### 100 2.1 Governing equations

Classical compressible Reynolds-averaged Navier-Stokes (RANS) and compressible Large Eddy Simulation (LES) equations  
are solved for a multi-species gas mixture using ONERA in-house finite volumes solver CHARME (see Refloch et al. (2011)).

105 Closing the RANS equations requires a turbulence model to compute the correlation terms. The turbulence model used in  
this work is the DRSM (Differential Reynolds Stress Modeling) ATAAC model (see Schwamborn and Strelets (2012)) where  
a transport equation is solved for each of the six Reynolds stresses, with a modelization of the pressure-strain correlation  
term based on Wilcox's work in Wilcox et al. (1998). Two additional equations are solved: one for the specific dissipation  $\omega$   
and another to model damping for near-wall turbulence. Unlike most two-equations turbulence models, DRSM models do not  
make the Boussinesq hypothesis, which makes them more accurate for solving strong vortical flows compared to eddy-viscosity  
models as observed by Churchfield and Blaisdell (2009).

110 For the LES, subgrid terms are modeled using the Smagorinsky model. In the context of contrail simulation, heat production  
is due to latent heat release caused by water condensation or sublimation. In practice, this term can be neglected in comparison  
to aerodynamic phenomena for recently formed contrails, as stated by Lewellen and Lewellen (2001) and Unterstrasser (2014).  
However, it is important to note that latent heat release cannot be neglected for older contrails in the diffusion regime (see Paoli  
et al. (2017)), which are beyond the scope of this work.

115 For both RANS and LES equations, convective fluxes are computed using the HLLC scheme with second-order accuracy. A  
second-order scheme is used for the computation of the diffusive fluxes. Time integration is performed using Backward Euler  
method for RANS calculations and Implicit Second-Order Runge Kutta for the LES. The GMRES method is used to solve the  
linear system obtained after the finite volumes discretization of the RANS/LES equations.

### 2.2 Microphysical model

120 The microphysical model used for the formation of contrail ice crystals is a simplified version of the bulk model developed  
by Jean-Charles Khou et al. (see Khou (2016); Khou et al. (2015); Montreuil et al. (2018)). More precisely, the exhaust plume  
is modeled through a multi-species gas mixture approach and is composed of seven species: dry air, water vapor  $H_2O$  (gas  
phase  $H_2O$ ), ice  $H_2O_s$  (ice phase  $H_2O$ ), sulfur trioxide  $SO_3$ , sulfuric acid  $H_2SO_4$ , sulfur trioxide in the adsorbed state  $SO_{3,ads}$   
and sulfuric acid in the adsorbed state  $H_2SO_{4,ads}$ . A transport equation is solved for the mass fraction  $Y_i$  of each species.  
125 The corresponding transport equations can be found in Montreuil et al. (2018). Moreover, ice crystal formation is modeled

through water vapor condensation done exclusively on soot particles previously activated by sulfur trioxide ( $\text{SO}_3$ ) and sulfuric acid ( $\text{H}_2\text{SO}_4$ ) molecules. In reality, water vapor first turns into liquid water before freezing to ice. However, freezing is not modeled, implying that any liquid water formed on the surface of a soot turns instantly into ice. The approach used by this model is purely Eulerian and as a result, the quantities associated with soot particles, ice crystals, and sulfur species are treated 130 as fields whose values depend on spatial coordinates and time. In the complete version of Khou's model (see Khou (2016)), the formation chemistry of  $\text{SO}_3$  and  $\text{H}_2\text{SO}_4$  molecules is considered through a reaction scheme involving 23 species and 60 reactions. This reaction scheme is not included in this work to reduce computational costs. Instead, it is assumed that  $\text{SO}_3$  and  $\text{H}_2\text{SO}_4$  molecules form very quickly after the engine exhaust gases are emitted, allowing their mass fractions to be directly 135 prescribed in the engine outlet boundary conditions. This assumption is supported by contrail simulations for similar cases using the complete model performed at ONERA.

Mathematically, the microphysical model first solves a transport equation for the soot density number  $N_s$ :

$$\frac{\partial}{\partial t} (\rho N_s) + \frac{\partial}{\partial x_i} (\rho v_i N_s) = \frac{\partial}{\partial x_i} \left( \rho D_{\text{diff}} \frac{\partial N_s}{\partial x_i} \right) \quad (1)$$

Thus, the equation for  $N_s$  is the equation of a passive scalar that does not influence flow properties (velocity, pressure, and temperature).  $D_{\text{diff}}$  is the diffusion coefficient of soot particles for which the corresponding Schmidt number  $\text{Sc}$  with respect to 140 air kinematic viscosity is set to 1. The turbulent Schmidt number  $\text{Sc}_t$  for the corresponding RANS and LES equation is set to 0.9. The corresponding LES equation is closed with Smagorinsky model and gradient hypothesis.

Soot activation occurs solely through the adsorption of  $\text{SO}_3$  and  $\text{H}_2\text{SO}_4$  molecules. Two new species are introduced:  $\text{SO}_{3,\text{ads}}$  and  $\text{H}_2\text{SO}_{4,\text{ads}}$ , with their respective mass fractions denoted as  $Y_{\text{SO}_{3,\text{ads}}}$  and  $Y_{\text{H}_2\text{SO}_{4,\text{ads}}}$ . These species represent  $\text{SO}_3$  and  $\text{H}_2\text{SO}_4$  molecules in the "adsorbed" state, as opposed to their "free" state. The transition from "free" to "adsorbed" states occurs through 145 the reactions  $\text{SO}_3 \rightarrow \text{SO}_{3,\text{ads}}$  and  $\text{H}_2\text{SO}_4 \rightarrow \text{H}_2\text{SO}_{4,\text{ads}}$

We define the activated surface fraction  $\theta_{\text{act}}$  of a soot particle. This quantity is equal to the ratio of the hydrophilic surface area to the total surface area of a given soot particle. A mass conservation equation is solved for the adsorbed species  $\text{SO}_{3,\text{ads}}$  and  $\text{H}_2\text{SO}_{4,\text{ads}}$ . The corresponding source term  $\dot{\omega}_{i,\text{ads}}$  is given by:

$$\dot{\omega}_{i,\text{ads}} = \frac{\pi D_s^2 \alpha_{\text{ads}} \rho Y_i N_s v_{th,i} (1 - \theta_{\text{act}})}{4 M_i} \quad (2)$$

150 where  $i = \{\text{H}_2\text{SO}_4, \text{SO}_3\}$ ,  $M_i$  the molar mass of species  $i$ ,  $v_{th,i} = \sqrt{\frac{8k_b T}{\pi m_{\text{mol},i}}}$  the thermal velocity of species  $i$ ,  $m_{\text{mol},i}$  the molecular mass of species  $i$ ,  $D_s$  is the diameter of a spherical dry soot particle, and  $\alpha_{\text{ads}}$  is the sticking probability, equal to 1 if  $T \leq 420$  K and 0 otherwise.  $D_s$  is set to 54 nm in all the calculations performed in this work. This value is in the order of magnitude of experimental measurements of soot particles diameter of a turbofan engine (see Petzold and Schröder (1998)). Note that desorption is not accounted for, meaning that adsorbed molecules cannot leave the soot surface. To ensure mass 155 conservation, the source term  $-\dot{\omega}_{i,\text{ads}}$  is added to  $\text{H}_2\text{SO}_4$  and  $\text{SO}_3$  mass conservation equations.

The fraction  $\theta_{\text{act}}$  is defined as:

$$\theta_{\text{act}} = \sum_{i,\text{ads}} \frac{\text{Na} \rho Y_{i,\text{ads}}}{\pi D_s^2 N_s \sigma_0 M_i} \quad (3)$$

where  $\text{Na}$  is Avogadro constant and  $\sigma_0 = \frac{4}{\pi (D_{\text{H}_2\text{SO}_4})^2} \approx 5 \times 10^{18} \text{ m}^{-2}$  is the surface density of active sites, based on the molecular diameter of  $\text{H}_2\text{SO}_4$ .

160 Contrail ice is modeled by introducing a new species,  $\text{H}_2\text{O}_s$ , representing water in the solid phase formed through heterogeneous nucleation of water vapor on activated soot. More precisely, activated soot particles serve as condensation nuclei. Water vapor diffuses onto these particles and forms an ice cap if the ambient air is supersaturated with respect to liquid water. As stated above, the model assumes that freezing happens instantly once liquid water forms on the soot particle. Currently, the microphysical model makes the simplifying assumption that all the ice spreads evenly over the soot particle surface, creating spherical ice crystals. These newly formed ice crystals then act as condensation nuclei and continue to grow as long as the relative humidity with respect to ice exceeds 100 %. The transformation of water vapor to ice and vice-versa is modeled through the reaction  $\text{H}_2\text{O} \leftrightarrow \text{H}_2\text{O}_s$ . The source term  $\dot{\omega}_{\text{ice}}$ , representing the mass production rate of ice crystals due to water vapor deposition around soot particles, is given by:

$$\dot{\omega}_{\text{ice}} = D_{\text{con}} \frac{4\pi r_p \theta_{\text{act}} N_s D_{\text{H}_2\text{O}} G_\alpha(r_p) \left( p_v - p_{\text{sol}}^{\text{sat},S}(r_p, T) \right)}{R_{\text{H}_2\text{O}} T} \quad (4)$$

170

$$D_{\text{con}} = \begin{cases} 1 & \text{if } p_v > p_{\text{liq}}^{\text{sat},S}(r_p, T) \text{ or } r_p > r_s \\ 0 & \text{else} \end{cases} \quad (5)$$

$$p^{\text{sat},S}(r_p, T) = \exp \left( \frac{2\sigma M_{\text{H}_2\text{O}}}{R \rho_{\text{H}_2\text{O}} r_p T} \right) p^{\text{sat}}(T) \quad (6)$$

175  $p_v$  is the partial pressure of water vapor,  $T$  is the static temperature,  $D_{\text{H}_2\text{O}}$  is the diffusion coefficient of water vapor,  $\sigma$  is the surface tension coefficient of liquid water,  $R_{\text{H}_2\text{O}}$  is the specific gas constant for water vapor, and  $\rho_{\text{H}_2\text{O}}$  is the density of water in the liquid phase.  $r_p$  is the mean radius of ice crystals, locally defined as:

$$r_p^3 = \frac{3}{4\pi} \frac{\rho Y_{\text{H}_2\text{O}_s}}{\rho_{\text{sol}} N_s} + r_s^3 \quad (7)$$

where  $\rho$  is the density of the gas mixture,  $\rho_{\text{sol}}$  is the density of ice—not to be confused with  $\rho_{\text{H}_2\text{O}_s} = \rho Y_{\text{H}_2\text{O}_s}$  the density of the  $\text{H}_2\text{O}_s$  species in the gas mixture which corresponds to the ice water content quantity—, and  $r_s = \frac{1}{2} D_s$  is the radius of a dry

180 soot particle. The saturation vapor pressures with respect to liquid water and ice are given by the semi-empirical formulas from Murphy and Koop (2005) and the density  $\rho_{\text{sol}}$  is given by Tabazadeh et al. (2000)

$G_\alpha$  is a dimensionless function used to account for cases where the continuum assumption is no longer valid in the immediate vicinity of an ice crystal:

$$G_\alpha(r_p) = \left( \frac{1}{1 + \text{Kn}} + \frac{4\text{Kn}}{3} \right)^{-1} \quad (8)$$

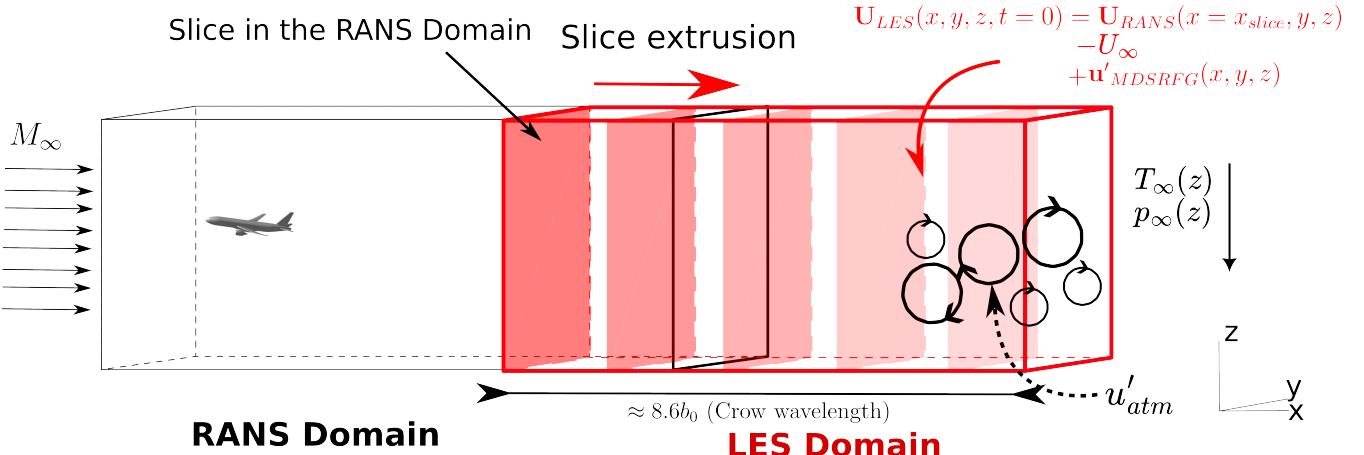
185 where  $\text{Kn} = \frac{\lambda_{\text{air}}}{r_p}$  is the Knudsen number, and  $\lambda_{\text{air}}$  is the mean free path of an air molecule in the plume.

Several remarks can be made regarding the expression for  $\dot{\omega}_{\text{ice}}$ . First, the introduction of the multiplier  $D_{\text{con}}$  prevents water vapor condensation on dry soot when the air is supersaturated with respect to ice but not with respect to liquid water. However, if ice has already formed around the soot, the nucleation process can continue even if the air is no longer saturated with respect to liquid water. Additionally, the sign of the term  $p_v - p_{\text{sol}}^{\text{sat},S}$  indicates whether the air is supersaturated with respect to ice or 190 not. In the first case, the source term is positive, and the crystals grow. In the second case, the source term is negative, and the ice crystals sublimate, returning to water vapor. The conservation of the total mass of water is thus ensured by adding the term  $-\dot{\omega}_{\text{ice}}$  to the source term of the  $\text{H}_2\text{O}$  species.

### 2.3 Bridging the gap between the near-field and the far-field

The methodology used to simulate the evolution of a contrail from the formation of the first ice crystals in the near-field to the 195 wingtip vortices destruction in the far-field is the methodology described and validated in the work of Bouhafid et al. (2024). This methodology was used to simulate the CRM geometry aerodynamic wake from the jet regime to the dissipation regime. A microphysical model, described above, is now added to take into account contrail formation and evolution. A diagram of the methodology is shown in Fig.1. The main ideas of the methodology are as follows. A RANS simulation of the jet regime of a realistic airliner geometry is carried out in the airplane reference frame. Anisotropic mesh adaptation techniques are used to 200 ensure a sufficient resolution of the flow and conservation of wake circulation up to twenty wingspans  $b$  behind the airplane. Next, a slice perpendicular to the flight direction is extracted from the RANS simulation. This slice is located far enough behind the airliner that jet/vortex interaction becomes important, and axial gradients become negligible in comparison to tangential gradients. The slice and its corresponding flow field are then extruded following the flight direction with an extrusion distance equal to one Crow instability wavelength  $\lambda_{\text{Crow}} = 8.6 \times \frac{\pi}{4}b$ . This extrusion step generates the LES domain for which periodic 205 boundary conditions are prescribed for the boundaries perpendicular to the flight direction. Moreover, the LES computation is now performed in the ground reference frame instead of the aircraft reference frame. Physically speaking, the temporal LES simulation is equivalent to a certain extent to observing the evolution of a contrail portion located inside a region of the sky after an airliner went through it.

In order to transfer the information from RANS turbulence-related fields (Reynolds tensor, dissipation rate) to the LES 210 domain, synthetic turbulence fluctuations noted  $u'_{\text{MDSRFG}}$  are generated using the MDSRFG method developed in Castro and Paz (2013) and added inside the plume, which is defined for a given position  $x$  by the isosurface  $U_x > \frac{1}{2}(U_j + U_\infty)$  with



**Figure 1.** Diagram of the initialization methodology of a far-field calculation using a near-field calculation. ( $M_\infty$ ,  $U_\infty$ ,  $T_\infty$  and  $p_\infty$  are respectively flight Mach number, flight speed, ambient static temperature and ambient static pressure).

$U_j$  the maximum axial velocity in the plume and  $U_\infty$  the flight speed. Another synthetic turbulence fluctuation field  $u'_{\text{atm}}$  is generated in the entirety of the LES domain using the MDSRFG method. This fluctuation field aims to model atmospheric turbulence responsible for the vortices destruction. The mathematical formulation of the synthetic turbulence is fully described in Bouhafid et al. (2024). Finally, stratification is taken into account by including the effect of gravity in the LES and allowing both ambient static temperature and ambient static pressure to vary with altitude  $z$  for a given Brunt-Väisälä frequency  $N_b$ .

### 3 Computational domain and numerical parameters

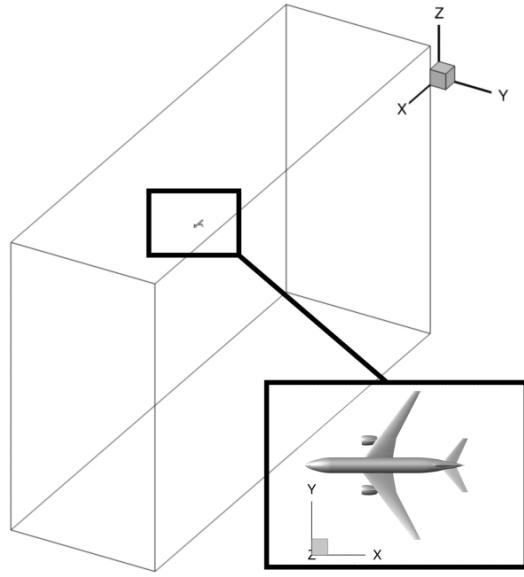
The methodology described in the previous section is now applied to the NASA Common Research Model geometry to which jet engines were added. First, a RANS calculation of the near-field is carried out and the results are discussed in Section 4.1. Afterward, the RANS calculation is extended to the far-field in Section 4.2 using a temporal LES according to the methodology developed in this work. The RANS and LES computational domains are described below.

#### 3.1 RANS domain

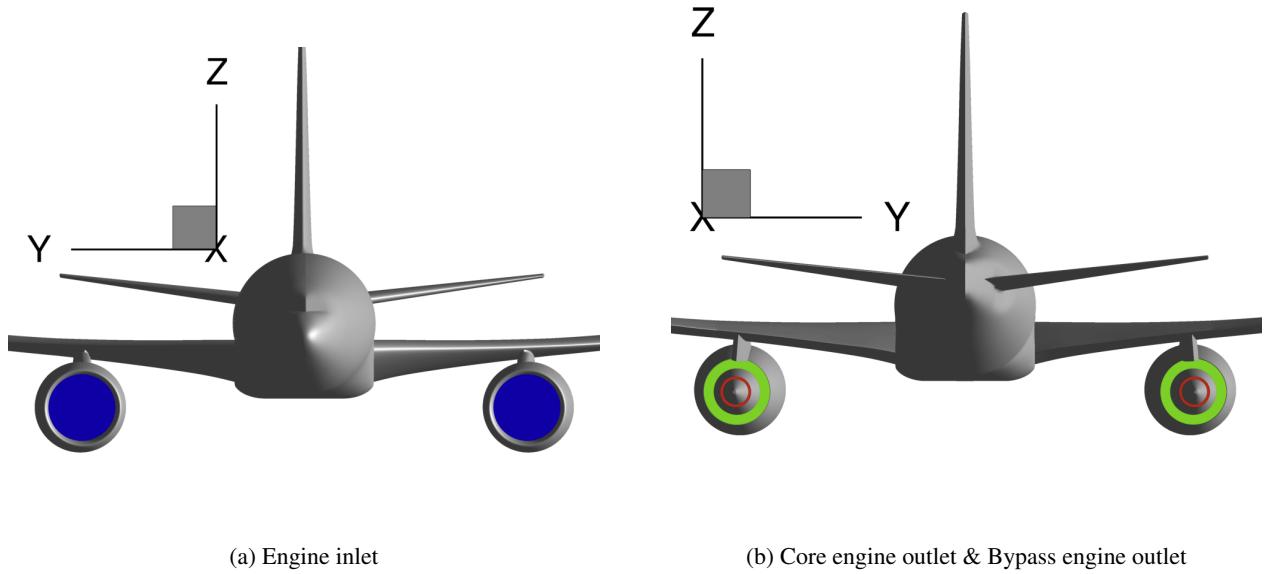
The RANS domain is the same as the one in Bouhafid et al. (2024) and consists of the B777-like CRM geometry inside a box. In terms of design parameters for the CRM geometry, wingspan  $b$  is 58.8 m, wing area  $S$  is 383 m<sup>2</sup>, and mean chord  $c_m$  is 6.5 m. The box dimensions are  $L_x \times L_y \times L_z$  with  $L_x = 30b$ ,  $L_y = 20b$  and  $L_z = 20b$  as illustrated in Fig.2.

The distance behind the aircraft in the RANS domain is equal to  $20b$ . Only half of the aircraft is taken into account in the simulation as the CRM geometry is cut in half by the symmetry plane  $y = 0$ . This is a common hypothesis justified by the symmetry of the mean flow with respect to the plane  $y = 0$ .

The contrail is studied for cruise flight conditions with a flight Mach number  $M_\infty = 0.85$  and an angle of attack  $\alpha = 3^\circ$  (see Tab.1). The ambient static pressure  $p_\infty$  and ambient static temperature  $T_\infty$  are typical for an altitude of 11 km inside

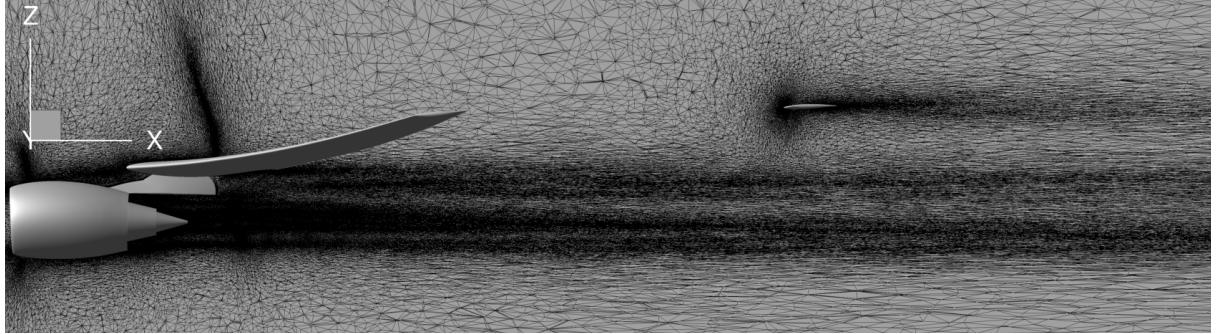


**Figure 2.** RANS computational domain with the CRM geometry. Only half of the geometry is taken into account in the computational domain.



**Figure 3.** CRM geometry with jet engines

the troposphere. The effect of stratification can be neglected in the jet regime and is thus not taken into account in the RANS calculation. Ambient water vapor mass fraction was chosen to impose  $RH_{ice} = 109\%$  and  $RH_{liq} = 68\%$  (see Tab.2), with



**Figure 4.** Mesh slice  $y = 0.17b$

$RH_{\text{ice}}$  and  $RH_{\text{liq}}$  being respectively the ambient relative humidity with respect to ice and the ambient relative humidity with respect to liquid water.

235 Turbofan engines were added to the CRM geometry by ONERA, as seen in Fig.3. Their position relative to the wing-  
tip is  $b_j/b = 34\%$ . The engine inlet and engine outlet are modeled with inlet/outlet boundary conditions (outlet boundary  
conditions for the engine inlet and inlet boundary conditions for the engine outlet). The thermodynamic values used to define  
these boundary conditions were chosen to mimic an Ultra-high bypass ratio (UHBR) turbofan in terms of bypass ratio while  
ensuring mass conservation between the inlet and the outlet. The relevant microphysical parameters for the engine outlet are  
240 given in Tab.2 and were chosen to mimic the emissions of a turbofan burning kerosene *Jet A1*. The soot number density at the  
engine outlet and the dry soot particle radius are of the order of magnitude of previous studies (see Montreuil et al. (2018);  
Khou et al. (2015); Ramsay et al. (2024)) and is in the same order of magnitude of in-situ measurements performed by Petzold  
et al. (1999). The corresponding soot emission index is  $\text{EI}_{\text{soot}} = \frac{N_s A_{\text{core}} U_{\text{core}}}{Y_{\text{H}_2\text{O}} \dot{m}_{\text{core}}} \text{EI}_{\text{H}_2\text{O}}$  with  $A_{\text{core}}$ ,  $U_{\text{core}}$  and  $\dot{m}_{\text{core}}$  being respectively  
245 the surface area of the core flow, the core flow exit velocity and the core flow mass flow rate. The water vapor emissions are  
based on the study conducted by Garnier et al. (1997), which characterizes the emissions of a Boeing 767 equipped with Rolls-  
Royce RB211 engines. It is important to note that these emissions are unlikely to be representative of a modern UHBR engine,  
given that the RB211 has a significantly lower bypass ratio of approximately five. Consequently, it would be interesting for  
future work to use more realistic engine emissions. Using those water vapor emissions yields the value  $\text{EI}_{\text{soot}} \approx 10^{15} \text{ kg}^{-1}$ .

250 The mesh of the RANS domain was generated using an anisotropic mesh adaptation technique based on the total kinetic  
energy (mean flow kinetic energy + turbulent kinetic energy) field. This was done thanks to the software Feflo.a from INRIA  
based on Loseille and Alauzet's work (see Loseille and Alauzet (2011a, b)). The final mesh contains around 110 million cells  
with around 40 million prisms and 70 million tetrahedra. It is the same mesh that was used in Bouhafid et al. (2024) for the  
255 study of the near-field aerodynamics of the CRM geometry in cruise flight conditions. Keeping the same mesh after activating  
the microphysical processes is equivalent to assuming that these processes are not energetic enough to significantly influence  
the aerodynamic properties of both the vortex wake and the jet. To illustrate the impact of anisotropic mesh adaptation, a mesh  
cut around the engine is shown in Fig.4. More details about the anisotropic mesh adaptation process and its application to the  
CRM geometry can be found in Bouhafid et al. (2024).

<b>Atmospheric &amp; engine conditions</b>					
Flight conditions	$M_\infty$ 0.85	$\alpha$ 3°	$p_\infty$ 264.37 hPa	$T_\infty$ 223.15 K	
Core engine outlet	<b>Stagnation temperature</b> 626.41 K		<b>Stagnation pressure</b> 530.01 hPa		
Bypass engine outlet	297.23 K		699.21 hPa		
Engine inlet	$p$ 370.12 hPa				

**Table 1.** Atmospheric & engine conditions

<b>Microphysical model parameters</b>			
Atmospheric conditions at flight altitude	$RH_{liq}$ 68%	$RH_{ice}$ 109%	$N_s$ 0
Core engine outlet	$Y_{air}$ $\approx 0.98$	$Y_{H_2O}$ $\approx 0.02$	$Y_{SO_3}$ $3 \cdot 10^{-10}$
	$Y_{H_2SO_4}$ $1.4 \cdot 10^{-8}$	$N_s$ $3.4 \cdot 10^{12} \text{ m}^{-3}$	$r_s$ 27 nm
Bypass engine outlet	$Y_{air}$ $\approx 0.9999$	$Y_{H_2O}$ $\approx 1 \cdot 10^{-4}$	

**Table 2.** Atmospheric & engine emissions at cruise flight conditions for the microphysical model

### 3.2 LES domain

The LES domain is obtained by extruding along the flight direction  $x$  the slice  $x/b = 18$  extracted from the RANS calculation as described in Sec.2.3. The extrusion distance is equal to one Crow instability wavelength  $\lambda_{\text{Crow}} = 8.6b_0 = 396$  m  $\approx 400$  m. The LES domain is  $20b \approx 1.2$  km long in the  $y$  and  $z$  direction in order to avoid spurious boundary effects. The LES mesh contains approximately 160 million hexahedral cells, with 100 cells in the flight direction. Such resolution in the axial direction is enough to correctly solve Crow instability but might not be enough if we aim to correctly resolve all short wavelength instabilities that could appear, as stated in Bouhafid et al. (2024). The mesh resolution in the  $y$  and  $z$  direction is maximum in the zone where the vortices come down because of their mutual induction. The length of the cells was chosen in order to provide 10 points per vortex core radius  $r_c$ , i.e.  $dy = dz \approx 0.1r_c$ .

Periodic boundary conditions are imposed for the boundaries perpendicular to the flight direction  $x$  while far-field boundary conditions are chosen for the other boundaries where the fluid velocity is equal to 0 and the pressure and temperature fields are imposed to ambient values according to stratification. Temporal integration is performed using Implicit Runge-Kutta 2nd-order method with a time step chosen to verify  $CFL < 0.5$ ,  $CFL$  being the Courant number. Each LES calculation simulates about 200 s of physical time.

The synthetic turbulence in the jet region is generated using a Karman-pao spectrum. The corresponding turbulence dissipation rate  $\epsilon_{\text{jet}}$  and integral length scale are deduced from the RANS calculation and are respectively equal to  $50 \text{ m}^2/\text{s}^{-3}$  and  $r_j = 6$  m, which is the jet radius at  $x/b = 18$ . The turbulence spectrum is discretized from the integral length scale to the smallest cell length.

Atmospheric conditions in the temporal LES are determined by atmospheric turbulence dissipation rate  $\epsilon_{\text{atm}}$  and Brunt-Väisälä frequency  $N_b$  characterizing stratification. Atmospheric turbulence is generated by adding synthetic turbulence fluctuations to the LES initial velocity field. Horizontal and vertical fluctuations are respectively modeled using the horizontal spectrum  $E_h = C_k \epsilon_{\text{atm}}^{2/3} k^{-5/3}$  and vertical spectra  $E_v = \frac{C_k}{r^2} \epsilon_{\text{atm}}^{2/3} k^{-5/3}$ .  $C_k$  is the Kolmogorov constant set at 1.5 (see Pope (2001)) and  $r$  is the anisotropy factor defined by  $r = \frac{u'_h}{u'_v}$  with  $u'_h$  and  $u'_v$  being respectively the root mean square values of the horizontal and vertical fluctuations of atmospheric turbulence. The two spectra are discretized from  $k = k_{\text{Crow}} = \frac{2\pi}{\lambda_{\text{Crow}}}$  to  $k = \sqrt{8}k_{\text{Crow}}$ .  $r$  is set to 5/4 (see Holzapfel et al. (2001)) and  $\epsilon_{\text{atm}}$  to  $6 \times 10^{-5} \text{ m}^2/\text{s}^{-3}$ , a typical value in the upper troposphere as measured in Cho and Lindborg (2001). The correspond turbulent time scale is  $\tau_b = \left( \frac{\lambda_{\text{Crow}}^2}{\epsilon_{\text{atm}}} \right)^{1/3}$ . For the CRM geometry and the given value of  $\epsilon_{\text{atm}}$ ,  $\tau_b \approx 23$  min. Since the dissipation regime lasts only a few minutes, it is not necessary to force atmospheric turbulence to avoid turbulence decay.

Concerning stratification, two Brunt-Väisälä frequencies are considered in this work:  $N_b = 0.012 \text{ s}^{-1}$  and  $N_b = 0.03 \text{ s}^{-1}$ . The first case corresponds to a standard value for medium stratification while the second case corresponds to a strong stratification scenario. These values are consistent with previous observations made in Schumann et al. (2017). Knowing the value of  $N_b$ , the vertical profile for ambient pressure and ambient temperature can be deduced by the equations:

$$\frac{N_b^2}{g} = \frac{1}{\Theta_\infty} \frac{d\Theta_\infty}{dz} \quad (9)$$

$$\frac{dp_\infty}{dz} = -\rho_\infty(z)g \quad (10)$$

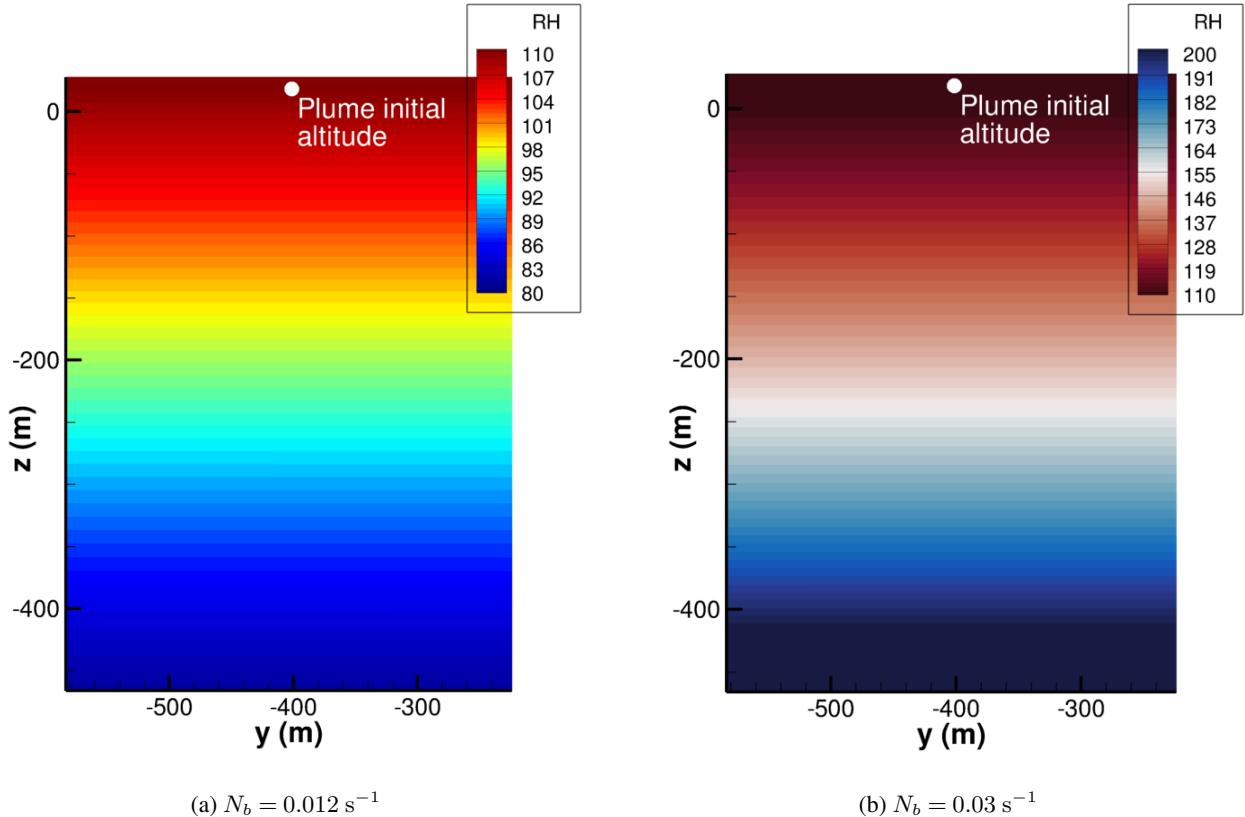
where  $\Theta_\infty$  is ambient potential temperature,  $g$  is gravitational acceleration, and  $\rho_\infty$  ambient air density. The solution equations can be found in Paoli et al. (2013). Solving those equations for the case  $N_b = 0.03 \text{ s}^{-1}$  with the chosen ambient pressure and ambient temperature at flight altitude gives rise to a temperature inversion phenomenon. That is to say, the temperature decreases when altitude decreases instead of increasing as observed for  $N_b = 0.012 \text{ s}^{-1}$ . This gives rise to an increase in relative humidity when altitude decreases as seen in Fig.5 where the ice relative humidity contour is plotted. Temperature inversion layers in the tropopause were already observed in the past, and their characteristics were examined in previous works (see Birner (2006); Grise et al. (2010)). Note that it is assumed in this work that the ambient water vapor molar fraction is constant in the whole domain. Its value is equal to the ambient water vapor molar fraction at cruise flight set for the precursor near-field calculation. Note that with this hypothesis, very high values of RH are reached at the bottom of domain in the strong stratification scenario. However, no ice crystals can descend to such low altitudes for this scenario because of the intense baroclinic vorticity that develop in the secondary wake.

## 4 Results and discussion

The results of the contrail simulations from the near-field to the far-field are discussed in this section for the two stratification scenarios. In order to evaluate the impact of the temporal LES initialization strategy, another contrail simulations of the far field is carried out using the classical analytic initialization strategy commonly used in literature, where the vortices are idealized by a pair of two Lamb-Oseen vortices as in Unterstrasser and Görsch (2014); Unterstrasser (2014); Holzäpfel et al. (2001); Paoli et al. (2013). A comparison is made between the RANS initialization strategy and the analytic initialization strategy. Explanations are given to understand the observed differences. In light of the obtained results, a new simulation is performed using an analytic initialization strategy with four Lamb-Oseen vortices, two for the main wing-tip vortices and two for the horizontal tailplane vortices. Finally, a LES simulation with two Lamb-Oseen vortices but with the plume obtained from the RANS simulation is carried out to investigate the effects of plume initialization. Note that all the analysis made below concern mainly the contrail evolution and its associated microphysics fields since the analysis of the aerodynamic fields was already done in Bouhafid et al. (2024).

### 4.1 Evolution of the contrail in the near-field

The contrail in the near-field can be visualized in Fig.6 where the cells satisfying the condition  $\theta > 0.01$ ,  $N_s > 10^7 \text{ m}^{-3}$  and  $r_p > 1.5r_s$  are blanked. For those conditions and the given atmospheric conditions, the contrail appears at around one wingspan behind the engines as seen in Fig.6b. Moreover, Fig.6a allows us to observe the jet deviation caused by the wing-tip vortex, a deviation that is followed by the beginning of the jet roll-up around the vortex.



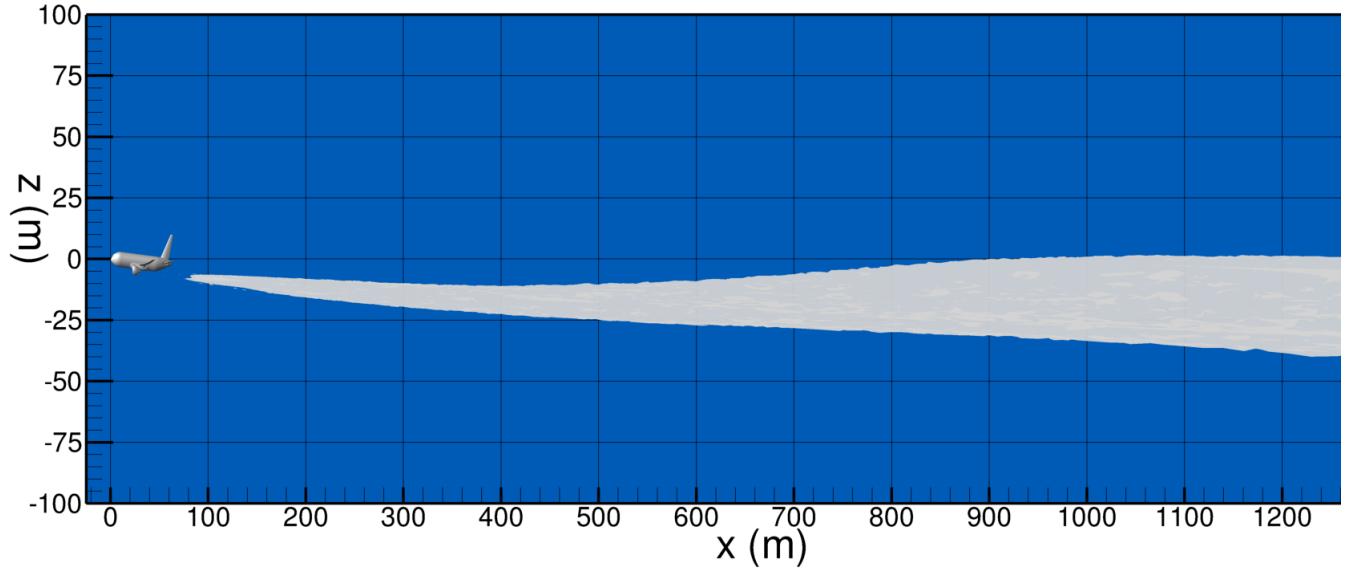
**Figure 5.** Ice relative humidity profile in % for the LES simulations for the two considered stratification scenarios

In order to quantify the evolution of the contrail in the near-field, the relevant microphysical quantities are averaged in a cross-section of the contrail for a given position  $x$ . The contrail is identified by the cells satisfying the condition  $r_p > r_s$  (strict inequality) and the average is weighted by the local number of ice crystals in a given cell. The averaged microphysics quantities are plotted in Fig.7. Fig.7a displays the evolution of the average ice crystal number  $N_p$  (not to be mistaken for soot number density  $N_s$ ) as a function of  $x/b$ . It can be observed that the dilution of the plume reduces ice crystal number by two orders of magnitude by the end of the jet regime. The dilution is faster for the first few wingspans because of the jet strong turbulent diffusion. As turbulence inside the plume decreases when the distance behind the aircraft increases, dilution becomes slower, and the influence of the vortex on plume dilution becomes more important than turbulent diffusion.

The evolution of the average ice water content  $IWC = \rho Y_{ice}$  plotted in Fig.7b can be divided into three zones. First, the IWC evolution shows a sharp growth for the first few wingspans, with an increase equal to four orders of magnitude compared to  $x/b = 0.5$ . This strong ice production results from the condensation of the excess water vapor present in important quantities in the plume in the early stage of the jet regime. This can be seen too in Fig.7e where relative humidity strongly increases for the first three wingspans. Second, for  $x/b$  between 5 and 15, the evolution of IWC is no longer monotonous as two maxima



(a) 3D view



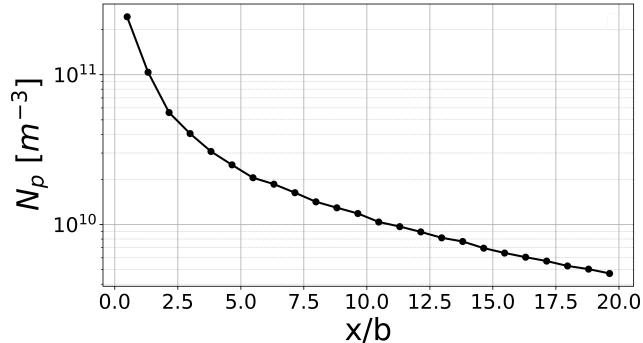
(b)  $z - x$  view

**Figure 6.** Mesh cells satisfying  $\theta > 0.01$ ,  $N_s > 10^7 \text{ m}^{-3}$  and  $r_p > 1.5r_s$

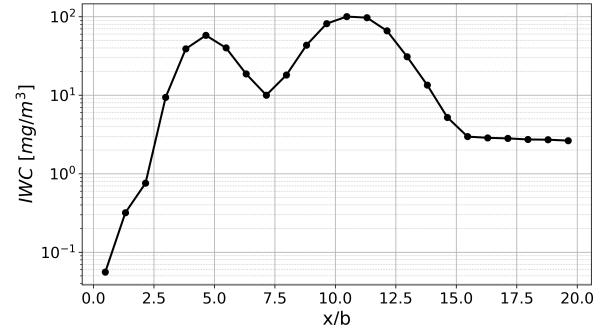
appear for  $x/b = 4.5$  and  $x/b = 11$ . This can be explained by the fact that the evolution of IWC results from the competition 335 between two phenomena: ice water dilution and ice water production. Indeed, when plume dilution is greater than ice water production, IWC will decrease and vice-versa when plume dilution is smaller. For the first three wingspans, the effect of ice water production is superior to the effect of dilution because of the important quantity of excess water vapor in the plume. However, looking at Fig.7e, relative humidity reaches a maximum at  $x/b = 3$ , that is to say, shortly before IWC does. This

results in a decrease in ice water production. Finally, for  $x/b > 15$ , the value of IWC reaches a plateau that can be interpreted  
340 as an equilibrium between ice water dilution and ice water production.

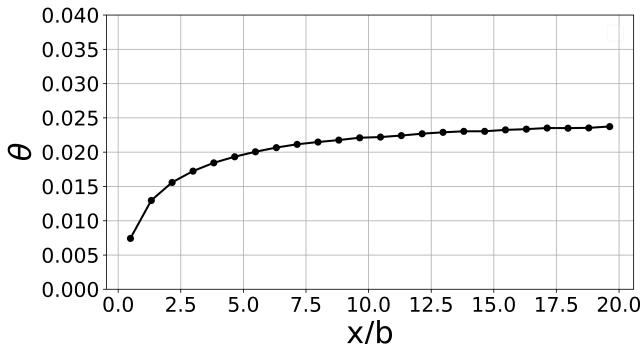
Fig.7c plots the evolution of the averaged activated surface fraction  $\theta_{act}$ . It appears that soot activation occurs very quickly  
after the exhaust gases are ejected from the engine. An asymptotic value  $\theta_{act} \approx 2.5\%$  is reached at the end of the calculation



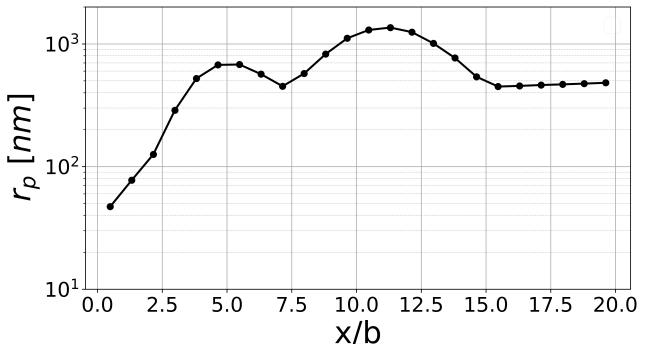
(a) Ice crystal number



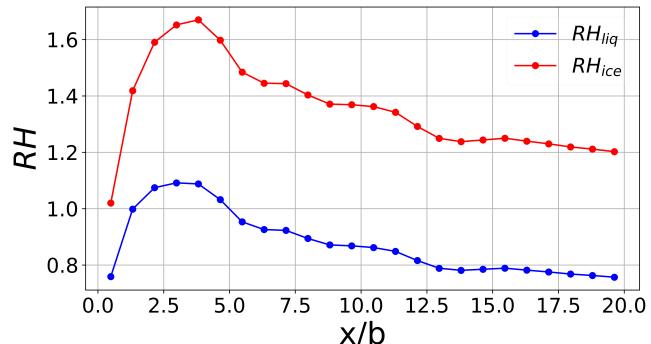
(b) Ice water content



(c) Activated surface fraction

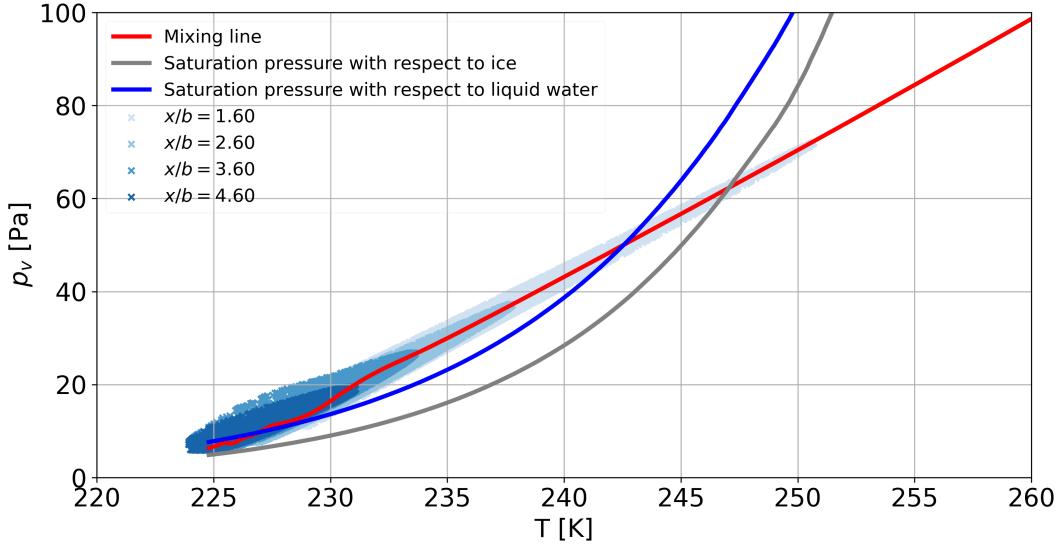


(d) Mean crystal radius



(e) Relative humidity with respect to ice water and liquid water

**Figure 7.** Microphysical quantities averaged in a cross-section of the contrail as a function of the distance behind the aircraft  $x$



**Figure 8.** Mixing line of the plume in the near-field plotted in a  $(p_v, T)$  diagram. The plume is obtained by averaging fluid variables on thirty streamlines passing through the engine core flow.

domain. The existence of an asymptotic value was already observed in Khou (2016) and Kärcher 1D model developed in Kärcher (1998). This was explained by the effect of dilution on sulfur species and the absence of desorption in the used model,

345 meaning that the activated surface fraction of a soot particle cannot decrease.

Regarding the evolution of the averaged mean ice crystal radius  $r_p$  plotted in Fig.7d, we can make the same observations and analysis as the ones made for IWC. Indeed, since  $r_p \propto \sqrt[3]{\text{IWC}/N_p}$  (see Eq.7), the evolution of  $r_p$  will closely follow the evolution of IWC when IWC evolves more rapidly than  $N_p$ . At the end of the domain,  $r_p$  reaches an asymptotic value equal to 500 nm. This value is of the same order of magnitude compared to values obtained in previous similar studies (see

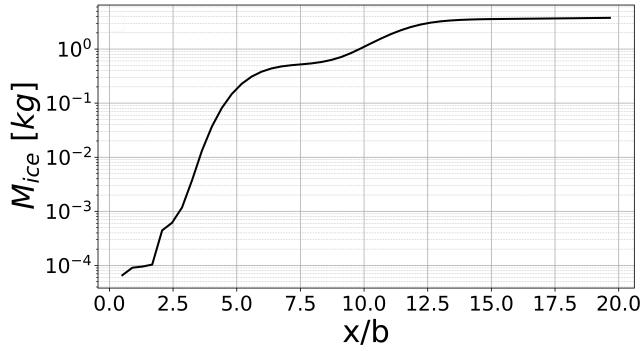
350 Khou et al. (2015); Ramsay et al. (2024); Paoli et al. (2013)). Note that the exact value will strongly depend on the ambient relative humidity, soot number density at the engine exit, and soot activation in the plume. Indeed, the work done by Ramsay et al. (2024) used a very similar microphysical model and assumed that every soot particles is activated, which explains to some

extent why the ice crystals obtained in their simulations are larger compared to this work's. This may appear counterintuitive, as a higher number of available nucleation sites generally leads to increased competition for water vapor, resulting in smaller ice 355 crystals. However, the source term for ice formation used in this work microphysical model (see Eq. 4) is directly proportional to the activated surface fraction  $\theta_{\text{act}}$  with  $\theta_{\text{act}} \leq 1$ . This amounts to assuming that the ice formed on the activated fraction of the soot is redistributed over the soot without increasing the activated surface fraction. In reality, initial ice formation creates a cap on the activated surface, increasing the hydrophilic surface fraction. Since the soot radius is negligible compared to the

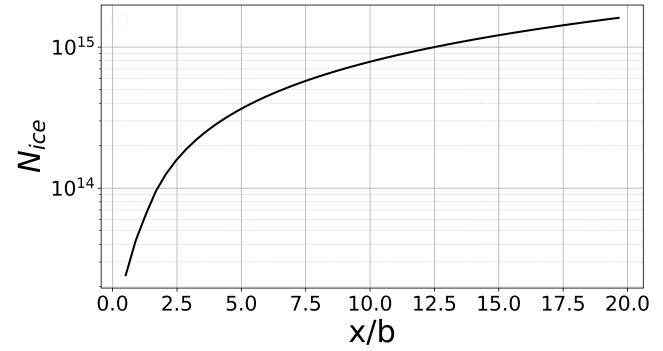
final ice crystal size, the effective hydrophilic surface fraction should reach unity. This behavior is not captured by the current 360 model, leading to smaller ice crystals than in Ramsay et al., where  $\theta_{\text{act}} = 1$  is assumed from the start.

The mixing line of the plume is plotted in Fig.8 in  $(p_v, T)$  diagram, where  $p_v$  is the water vapor pressure and  $T$  is the temperature. This line is obtained by averaging fluid variables over twenty streamlines originating from the engine core outlet. As the plume ages, both temperature and water vapor pressure decrease due to dilution with the cold ambient air. When the mixing line goes above the saturation pressure curve for liquid water, ice crystals begin to form. These crystals will grow as 365 long as the mixing line remains above the saturation pressure curve for ice, which occurs at approximately  $T = 243$  K for the flight conditions considered.

To examine the distribution of  $(p_v, T)$  within the plume, the  $(p_v, T)$  values in every cell within the plume region — defined as those where  $N_s > 0.1N_{s,\max}$ , with  $N_{s,\max}$  being the maximum  $N_s$  value at a given position  $x$  — are plotted for various 370 positions  $x$  downstream of the aircraft. The obtained scatterplot provides a clear visualization of how temperature and water vapor pressure vary within the plume for a given plume age. Notably, the older the plume, the less scattered the points become, reflecting the increasing uniformity caused by plume dilution. Finally, we observe that the mixing line differs from a straight line for  $T < 230$  K as  $p_v$  decreases more abruptly. This is most likely due to the effect of water vapor condensation which decreases  $p_v$  more rapidly than the plume isobaric mixing.



(a) Total ice mass



(b) Total ice crystal number

**Figure 9.** Evolution of the total ice mass and total ice crystal number

The evolution of the total ice mass  $M_{\text{ice}}$  and the total ice crystal number  $N_{\text{ice}}$  are plotted in Fig.9 as a function of the 375 distance  $x/b$  behind the aircraft. More precisely,  $M_{\text{ice}}(x)$  is defined as  $\int_0^x IWCdV$  and  $N_{\text{ice}}$  as  $\int_0^x N_p dV$ . The integrals are calculated using the rectangle method. Concerning  $M_{\text{ice}}(x)$ , we observe a sharp increase around  $x/b = 2$ , which corresponds to the moment when the averaged relative humidity with respect to liquid water is greater than 100% in the plume. The total mass of ice grows before reaching a plateau around  $x/b = 13$ . When it comes to the total ice crystal number, the evolution of  $N_{\text{ice}}$  seems to be logarithmic with respect to  $x$  in  $y - \log$  scale. This corresponds to a linear evolution in  $y$ —linear scale. We 380 notice that the total ice crystal number increases by almost two orders of magnitude between  $x/b = 0.5$  and  $x/b = 20$ .

## 4.2 Evolution of the contrail in the far-field

The results of four contrail far-field LES simulations are analyzed and discussed in this subsection. The first two simulations are initialized using the near-field RANS simulation. The other three simulations are initialized using an analytic initialization strategy which proceeds as follows. A vortex pair is initialized with two Lamb-Oseen vortices located at the same position as 385 the RANS vortex pair. The velocity inside the vortices is purely tangential and is defined by:

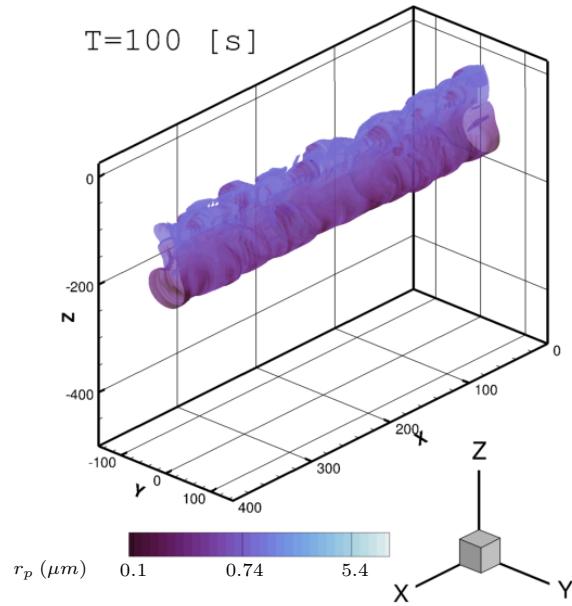
$$V_{\theta, \text{LO}}(r) = \frac{\Gamma}{2\pi r} \left( 1 - e^{-\beta \frac{r^2}{r_c^2}} \right) \quad (11)$$

$$\frac{dp_{\text{LO}}}{dr} = \rho \frac{V_{\theta, \text{LO}}}{r} \quad (12)$$

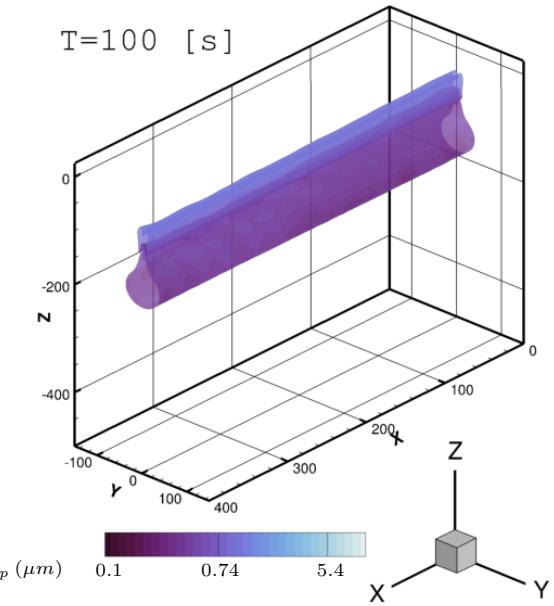
with  $V_{\theta, \text{LO}}$  is the tangential velocity,  $r$  the radial coordinate defined from the center of the vortex,  $\Gamma$  the vortex circulation,  $r_c$  the 390 vortex core radius,  $p_{\text{LO}}$  the static pressure inside the vortex and  $\beta \approx 1.256$ .  $r_c$  and  $\Gamma$  are deduced from the RANS calculation at  $x/b = 18$ . The plume is modeled by adding two circular jets with purely axial flow and located at the position of the jet engines in the CRM geometry. The radius of the jet is obtained from the RANS calculation as it was done in Bouhafid et al. (2024). All the flow field variables are uniform inside the jet, with a hyperbolic tangent decrease outside the jets. The value of the flow 395 field variables in the jets is deduced from the RANS calculation. Moreover, the parameters of the hyperbolic tangent decrease are calibrated for the mass fraction fields in order for the total mass of each species to be identical to the total mass obtained from the RANS initialization. The analytic initialization strategy with two Lamb-Oseen vortices will be noted 2LO from now on. Finally, time is scaled by the vortex pair characteristic time  $t_b = \frac{2\pi b_0^2}{\Gamma} = 25.7$  s, which is the theoretical time for the vortex 400 pair to descend a distance equal to  $b_0$  as a result of their mutual induction.

Fig.10 shows at  $t/t_b = 3.89$  and  $t/t_b = 7.78$  the soot number density isosurface  $N_s = 10^7 \text{ m}^{-3}$  colored by mean crystal 405 radius  $r_p$  for the medium stratification scenario  $N_b = 0.012 \text{ s}^{-1}$ . The important differences regarding the aerodynamic field observed and explained in Bouhafid et al. (2024) between the 2LO initialization and RANS initialization significantly impact the contrail shape and properties. While there are some similarities between the two cases, like the formation of a vortex ring following Crow instability, the contrail in the 2LO case appears much smoother. Conversely, the turbulent nature of the secondary wake in the RANS initialization case originating from short-wave vortex instabilities gives the contrail a more 410 chaotic appearance, while increasing the size of the contrail at flight altitude.

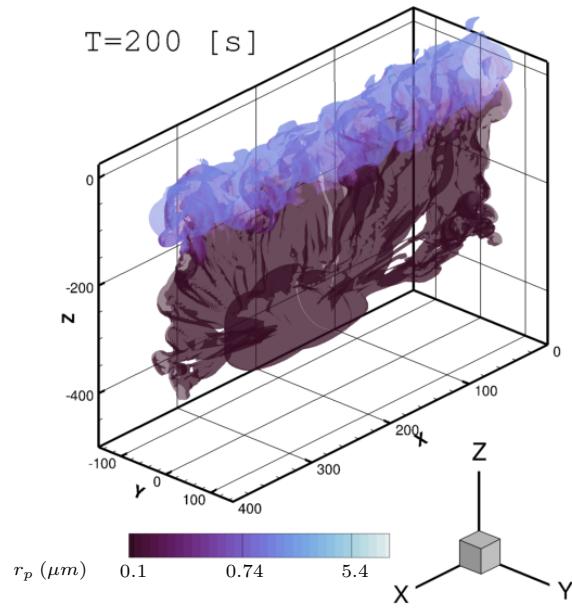
The differences between the 2LO initialization case and RANS case are even greater for the strong stratification scenario  $N_b = 0.03 \text{ s}^{-1}$  for which the same soot number density isosurface is plotted in Fig.11 at  $t/t_b = 3.89$  and  $t/t_b = 7.78$ . As 415 expected, more ice crystals are found in the secondary wake for the two initialization strategies compared to the medium stratification scenario. Since atmospheric conditions are favorable for contrail persistence in the considered strong stratification scenario because of thermal inversion, the resulting contrail will contain a higher ice water content and larger crystals. However, we observe that the contrail goes significantly lower in the 2LO case. This results in a wider contrail for the RANS case.



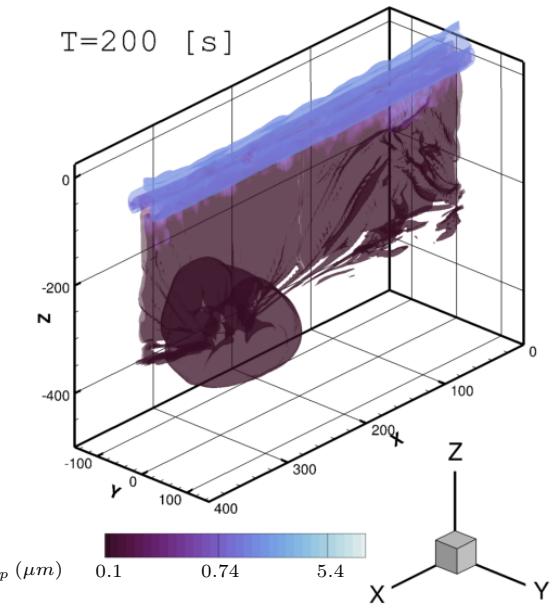
(a)  $t/t_b = 3.89$ , RANS initialization



(b)  $t/t_b = 3.89$ , 2LO initialization



(c)  $t/t_b = 7.78$ , RANS initialization

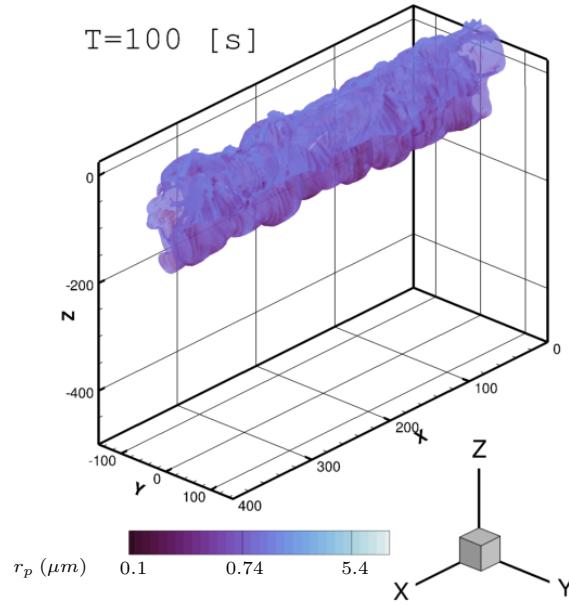


(d)  $t/t_b = 7.78$ , 2LO initialization

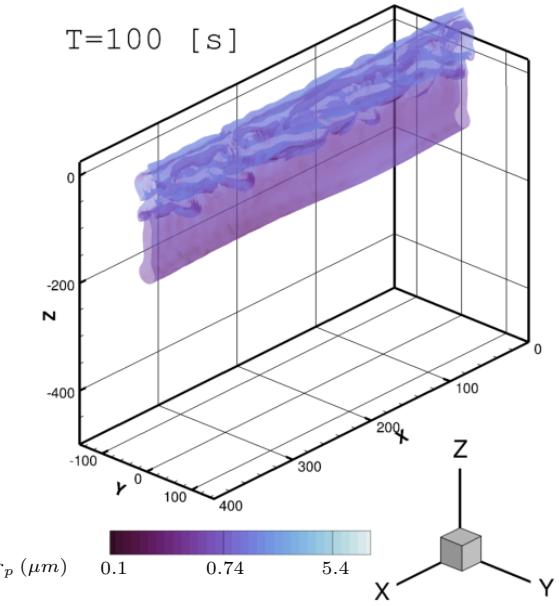
**Figure 10.** Soot number density isosurface  $N_s = 10^7 \text{ m}^{-3}$  colored by mean crystal radius  $r_p$  in logarithmic scale.  $N_b = 0.012 \text{ s}^{-1}$  (medium stratification). Comparison between RANS initialization strategy and 2LO initialization strategy. Distances are in meters. ( $t_b = 25.7 \text{ s}$ ).

A possible explanation for this phenomenon could be that the vortices are more unstable in the RANS case and are destroyed

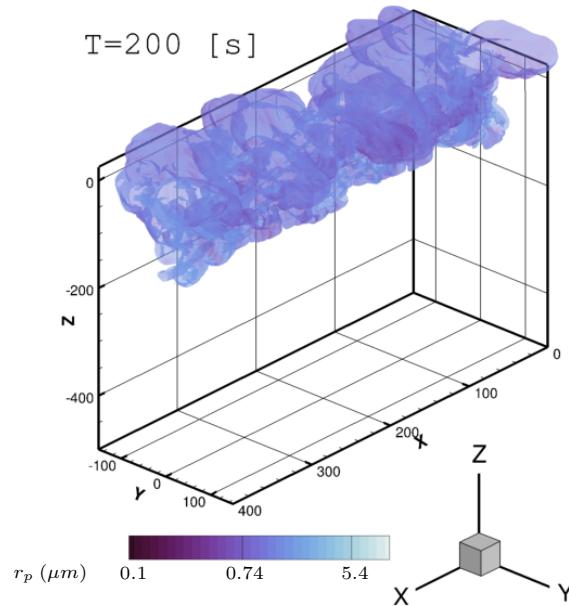
sooner, thus preventing the vortices from going downwards and carrying ice crystals to lower altitudes. The unstable nature of



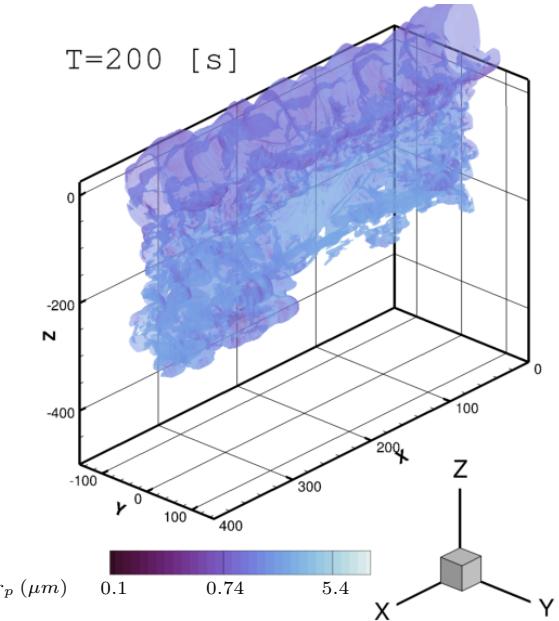
(a)  $t/t_b = 3.89$ , RANS initialization



(b)  $t/t_b = 3.89$ , 2LO initialization

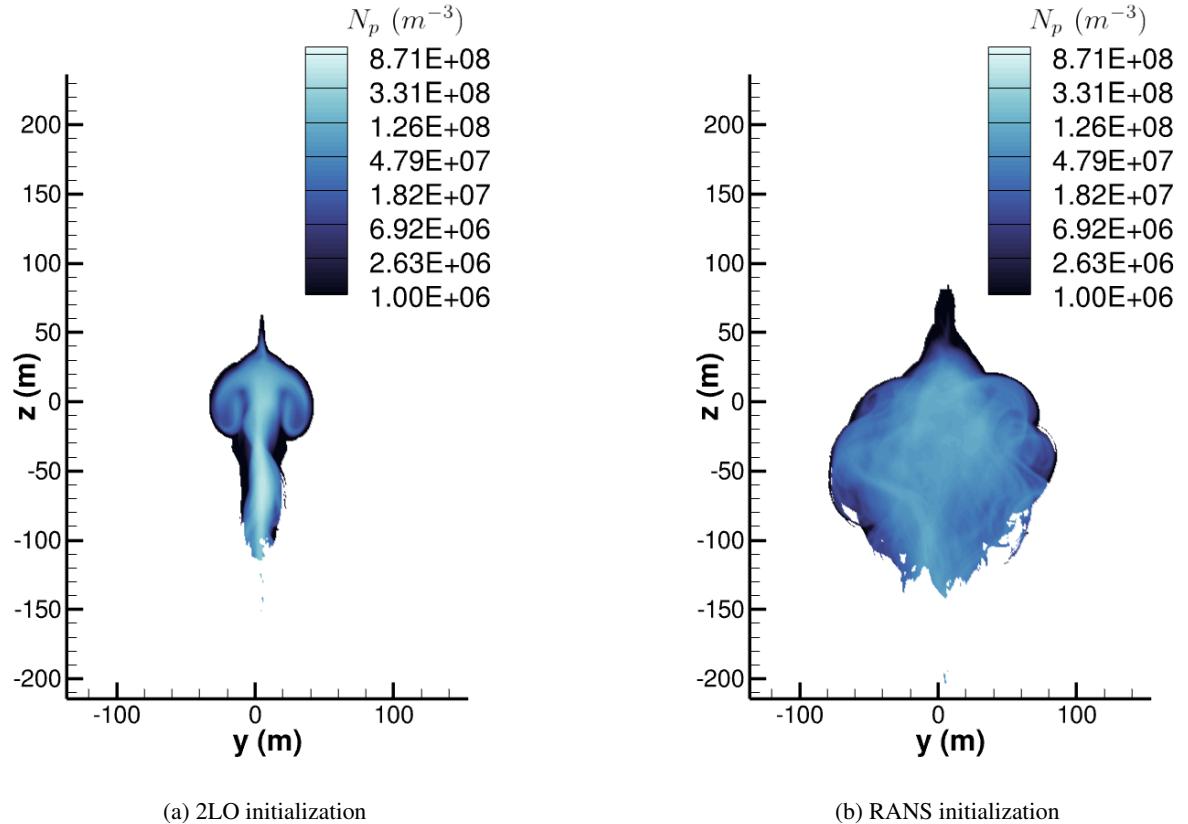


(c)  $t/t_b = 7.78$ , RANS initialization



(d)  $t/t_b = 7.78$ , 2LO initialization

**Figure 11.** Soot number density isosurface  $N_s = 10^7 \text{ m}^{-3}$  colored by mean crystal radius  $r_p$  in logarithmic scale.  $N_b = 0.03 \text{ s}^{-1}$  (strong stratification). Comparison between RANS initialization strategy and 2LO initialization strategy. Distances are in meters. ( $t_b = 25.7 \text{ s}$ ).

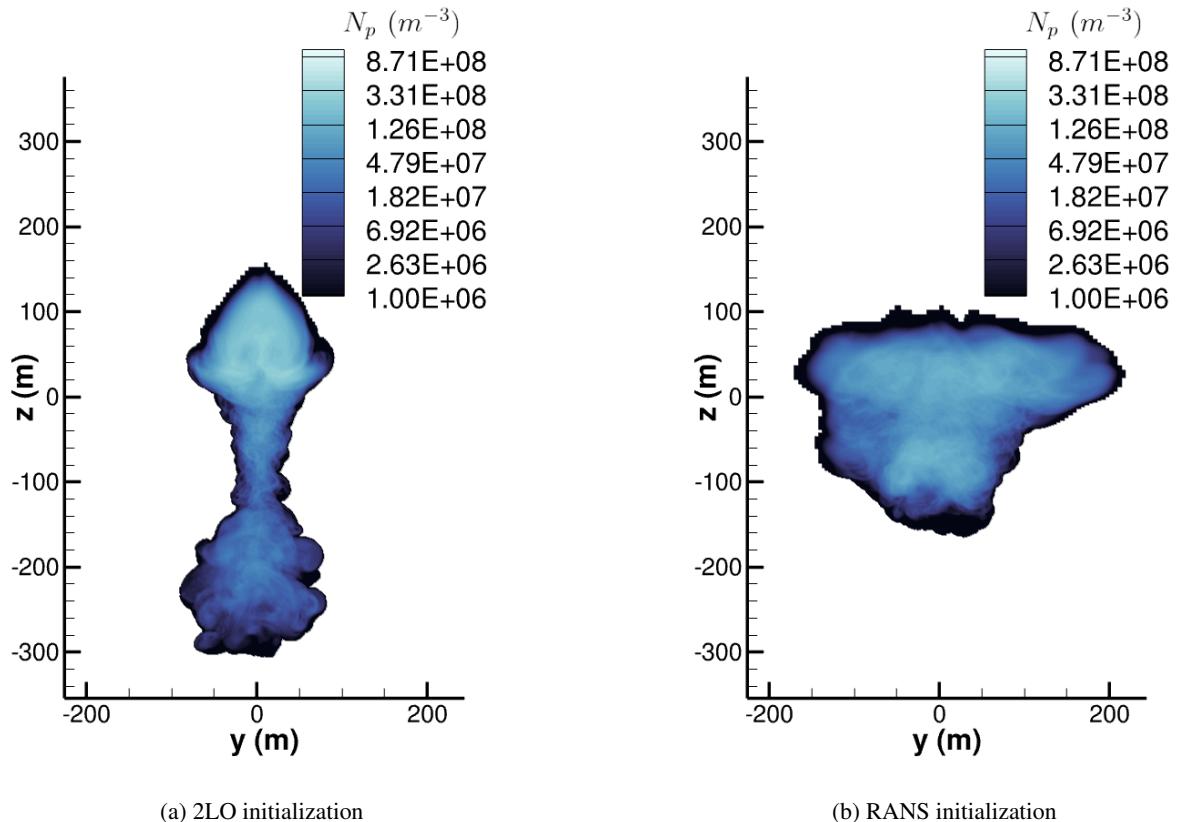


**Figure 12.** Spatially averaged ice crystal number field  $N_p$  in  $\text{m}^{-3}$  at  $t/t_b = 8$  ( $t_b = 25.7$  s) and for  $N_b = 0.012 \text{ s}^{-1}$ . A logarithmic scale is used for the color levels.

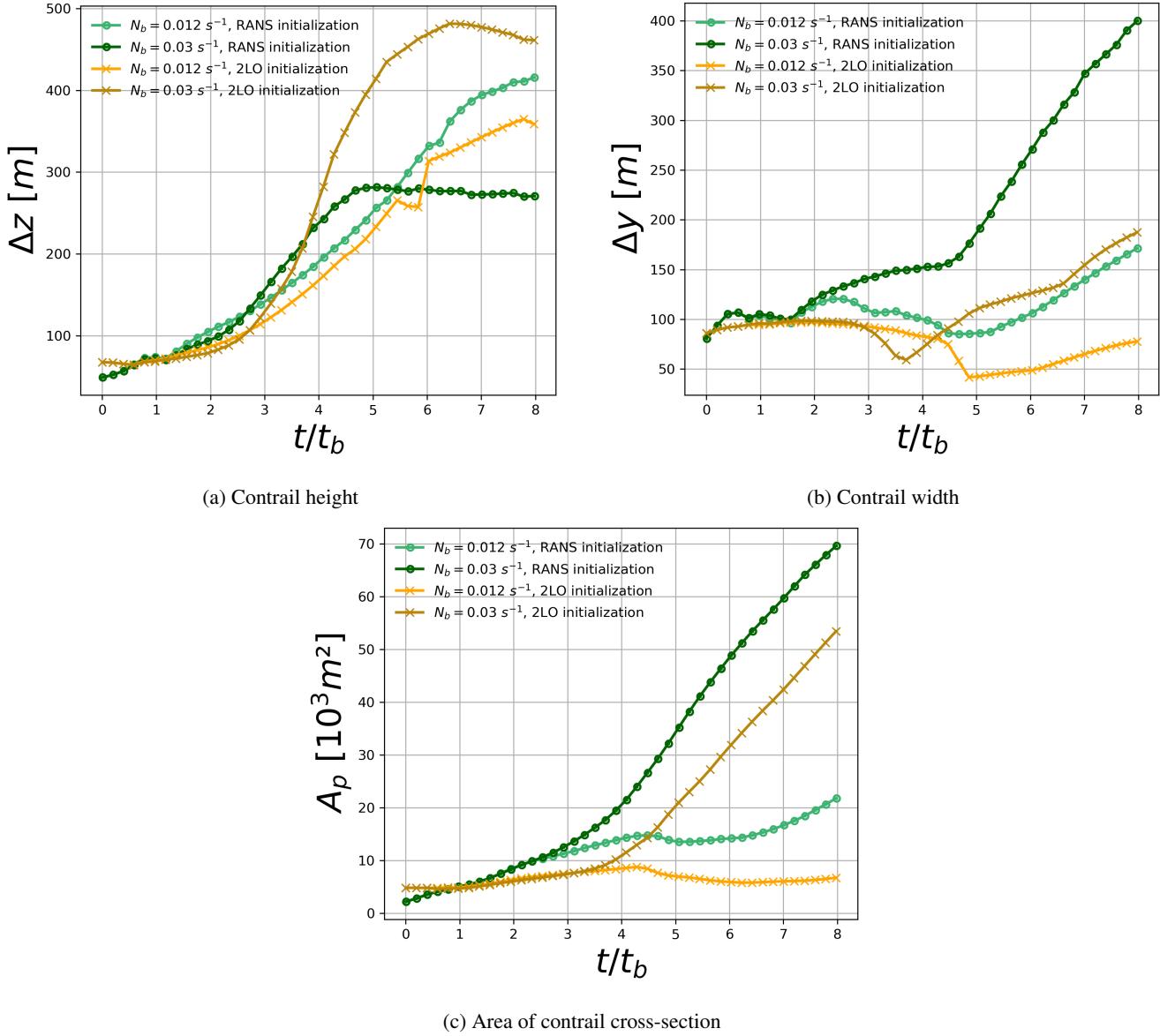
the wake can be amplified either by the development of Windall instability in the RANS initialization—manifesting under certain conditions but absent in the 2LO initialization computation—or by the presence of horizontal tailplane vortices, resulting in a four-vortex system. This will be further examined in Section 4.3.

The differences observed can be quantified by spatially averaging the contrail along flight direction  $x$ , with the cells in the contrail being identified by  $r_p > r_s$ . The strict inequality ensures that only ice crystals are considered and not dry soot particles. The ice crystal number field obtained in the averaged contrail cross section is plotted in Fig.12 and Fig.13 for  $t/t_b = 8$  ( $t = 200$  s). More precisely, the ice crystal number corresponds to the soot number density but only for soot particles with an ice layer on the surface. Thus, ice crystal number is noted  $N_p$  in order to distinguish it from the soot number density  $N_s$ . Fig.14 displays the relevant contrail length scales for the RANS initialization and 2LO initialization cases. Contrail height  $\Delta z$ , contrail width  $\Delta y$ , and contrail cross-sectional  $A_p$  area are plotted versus scaled time  $t/t_b$ . Contrail height  $\Delta z$  is computed by taking the difference between the maximum value of  $z$  and the minimum value of  $z$  in the contrail. Contrail width is computed the same way but with the  $y$  coordinate. Contrail cross-sectional area is calculated by summing the area of every cell inside the contrail.

Regarding contrail height, the differences between the RANS initialization and 2LO initialization remain small for the medium stratification scenario. However, for the strong stratification scenario, the contrail is 1.6 times taller for the 2LO initialization case compared to the RANS initialization case. The differences between the four different cases are larger for the contrail width. For both stratification scenarios, the contrail is two times wider for the RANS initialization compared to the equivalent 2LO initialization. For contrail cross-sectional area, two observations can be made. First, for the RANS initialization, the influence of stratification becomes visible at  $t/t_b = 2.5$  while it happens later for the 2LO initialization, at around  $t/t_b = 3.5$ . Second, the contrail is bigger and grows more rapidly for the strong stratification scenario. This is simply due to the higher ambient relative humidity  $RH$  for  $N_b = 0.03 \text{ s}^{-1}$  resulting from thermal inversion. For  $N_b = 0.012 \text{ s}^{-1}$ ,  $RH$  eventually decreases below 100% as the vortices descend, resulting in ice water sublimation and a decrease of the contrail cross-sectional area at  $t/t_b = 4.3$ . Nonetheless, the contrail begins to grow again at later times because of the ice crystals in the secondary wake ascending back to flight altitude where air is supersaturated with respect to ice water. It is interesting to note that for the strong stratification scenario, contrail height stops increasing around  $t/t_b = 4.5$  for RANS initialization and  $t/t_b = 6.3$  for 2LO initialization while contrail width keeps increasing. This can be explained by the impact of the vortex

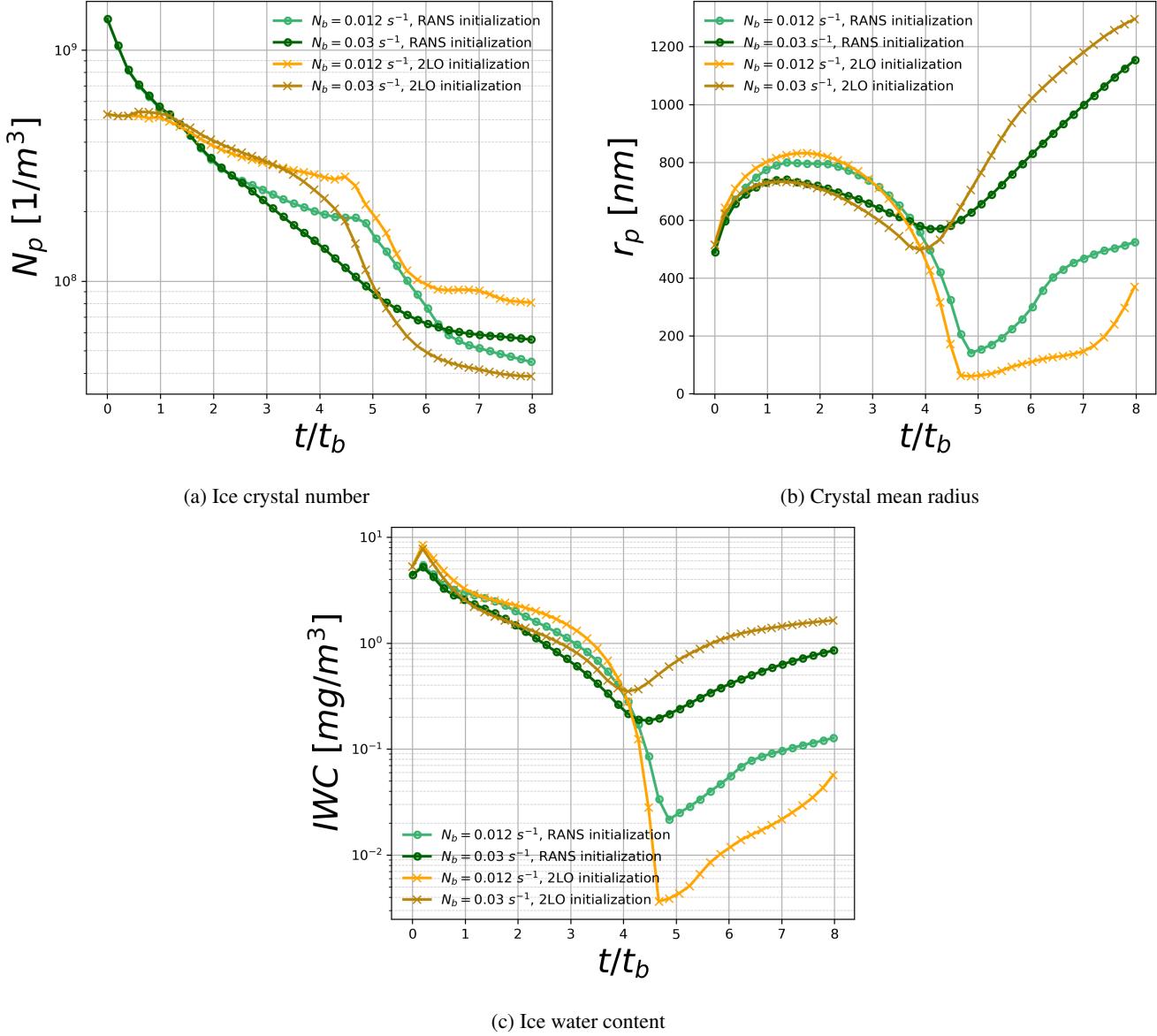


**Figure 13.** Spatially averaged ice crystal number field  $N_p$  in  $\text{m}^{-3}$  at  $t/t_b = 8$  ( $t_b = 25.7 \text{ s}$ ) and for  $N_b = 0.03 \text{ s}^{-1}$ . A logarithmic scale is used for the color levels.



**Figure 14.** Contrail length scales for the different stratification scenarios. Comparison between the 2LO initialization strategy and the RANS initialization strategy. ( $t_b = 25.7$  s).

break-up which occurs sooner for  $N_b = 0.03 \text{ s}^{-1}$  compared to  $N_b = 0.012 \text{ s}^{-1}$ . The ice crystals are no longer brought downward and the contrail will predominantly spread in the horizontal direction. We expect this phenomenon to occur at later times for  $N_b = 0.012 \text{ s}^{-1}$  as we can see that the inflection point for  $\Delta z(t)$  is attained at  $t/t_b = 6.5$  for the RANS initialization for  $N_b = 0.012 \text{ s}^{-1}$ .



**Figure 15.** Average contrail microphysics quantities for the different stratification scenarios. Comparison between the RANS initialization strategy and the 2LO initialization strategy. ( $t_b = 25.7$  s).

Microphysical quantities are now averaged inside the contrail cross-section at every time  $t$ . Their evolution as a function of  $t/t_b$  is plotted in Fig.15. Fig.15a shows ice crystal number evolution, which decreases about one order of magnitude between the beginning of the vortex regime and the end of the dissipation regime for all cases. The smallest decrease is obtained for the 2LO initialization case in the medium stratification scenario. This is due to the smooth and weakly turbulent nature of the secondary wake compared to the RANS initialization, resulting in less mixing of the plume with the ambient air.

Fig.15b displays the evolution of the ice crystal radius. For the medium stratification scenario,  $r_p$  begins to increase because the plume is still located in supersaturated air with respect to ice. However, as the plume descends with the vortex pair, the plume is no longer supersaturated and the crystals begin to shrink. Eventually, every ice crystal in the primary wake is transformed back to water vapor while the crystals in the secondary wake return to flight altitude where they can grow. This results in an increase of the averaged value of  $r_p$  in the contrail. Contrastingly, for the strong stratification scenario, the temporary decrease is this time due to some ice crystals quickly moving back to flight altitude and beyond, where the ambient air is not supersaturated in the case of a thermal inversion phenomenon. Eventually, since most ice crystals are trapped in the primary wake, ice crystals begin to grow as the air becomes more supersaturated as altitude decreases. These explanations also apply to the evolution of the IWC plotted in Fig.15c, with the additional effect of plume dilution.

Part of the radiative properties for recently formed contrails can be quantified by computing the extinction  $E$  (see Unterstrasser and Gierens (2010)). To do so, we have to compute the optical depth  $\tau_v$  defined as  $I_{\text{tr}} = I e^{-\tau_v}$  with  $I$  the intensity of an incident wave and  $I_{\text{tr}}$  the intensity of the corresponding transmitted wave after traveling inside the contrail.  $\tau_v$  is given by Hulst and van de Hulst (1981):

$$\tau_v(x, y) = \int_{-\infty}^{+\infty} \pi r_p^2 N_s Q_{\text{ext}}(r_p) dz \quad (13)$$

$$Q_{\text{ext}} = 2 - \frac{4}{e} \left( \sin(e) - \frac{1 - \cos(e)}{e} \right) \quad (14)$$

with

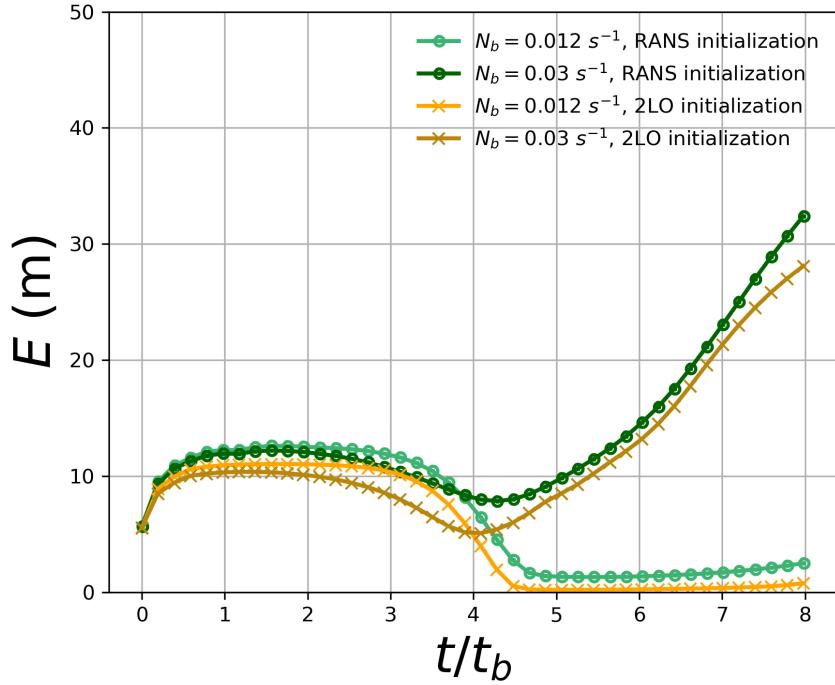
$$e = \frac{4\pi r_p(\mu_r - 1)}{\lambda} \quad (15)$$

where  $Q_{\text{ext}}$  is the Mie extinction coefficient.  $\mu_r = 1.31$  is the refractive index of ice and  $\lambda$  the wavelength of the incident light wave coming perpendicularly to the contrail. Extinction is then defined by:

$$E(x) = \int_{-\infty}^{+\infty} (1 - e^{-\tau_v}) dy \quad (16)$$

Physically speaking, the higher is  $E(x)$ , the more opaque the contrail will appear to an observer located perpendicularly to it, meaning that the contrail will absorb and diffuse more light. Thus,  $E$  and  $\tau_v$  are directly linked to the radiative properties of contrails.

$E(x)$  is averaged along  $x$  for different times  $t$  and is plotted in Fig.16 for the two stratification scenarios and the RANS/2LO initializations. We observe that for both stratification scenarios, the highest extinction is obtained for the RANS initialization case. The relative gap between the RANS initialization and the 2LO initialization is maximum for the medium stratification



**Figure 16.** Average extinction as a function of time  $t$  for the different stratification scenarios. Comparison between the RANS initialization strategy and the 2LO initialization strategy. ( $t_b = 25.7$  s).

scenario. These differences can be explained by the ice crystal number, ice crystal radius and contrail size differences that were observed and analyzed previously. We previously hypothesized that the vortex pair is more unstable for the RANS initialization in the strong stratification scenario. We can observe that this results in a higher extinction compared to the 2LO initialization. This result is consistent with Ferreira et al. (2024) findings where it was shown that forcing Crow instability, and

480 thus decreasing the lifetime of the wingtip vortices, increased the radiative forcing — which was estimated through radiative forcing parametrization techniques — of contrails in the first minutes after their formation.

It is important to note that extinction is the same at  $t = 0$  for every case. Thus, this shows the significant influence of near-field aerodynamics on the contrail evolution in the far-field. It would be valuable to use Paoli et al. (2017) methodology to extend the calculations to the diffusion regime and see if the differences persist after a few hours. Indeed, the differences in

485 extinction observed for the first few minutes may potentially decrease, or even vanish, over longer timescales owing to the effect of atmospheric turbulence and wind shear on the ice crystals spatial distribution. A larger extinction in the beginning does not necessarily imply a larger radiative impact at later times and over the full lifetime of the contrail.

### 4.3 Analytic initialization strategy with HTP vortices

In order to better understand the interaction between stratification and short wavelength vortex instabilities, another temporal  
 490 LES simulation is performed for the medium stratification scenario  $N_b = 0.012 \text{ s}^{-1}$ . This new simulation is similar to the 2LO initialization strategy. The only difference is that the vortex wake is now initialized with four vortices instead of two. To do so, a new Lamb-Oseen vortex pair is added and aims to model horizontal tailplane (HTP) vortices. The circulation, location and core radius of the HTP vortices are deduced from the RANS calculation and are respectively equal to  $-77 \text{ m}^2/\text{s}$  and  $2.5 \text{ m}$ . This new simulation will now be referred to as 4LO.

495 Fig.17 shows for  $t/t_b = 3.89$  and  $t/t_b = 7.78$  a soot number density isosurface colored by ice crystal mean radius for the RANS initialization and 4LO initialization. We observe that the contrail obtained with the 4LO initialization is qualitatively much more similar to the one obtained with the RANS initialization than the one obtained with the 2LO initialization. Indeed, the contrail exhibits the same turbulent aspect, both in the primary and secondary wake. This can be understood by looking at Fig.18, where a  $Q$ -criterion isosurface is plotted at  $t/t_b = 3.89$  for both cases to visualize the two vortex tubes of the vortex  
 500 pair. For the two cases, we observe the development of short-wave instabilities on the vortex tubes that are responsible for the wavy aspect of the vortex tubes. These instabilities are typical of quadripolar vortex systems in a counter-rotating configuration theoretically (see Fabre et al. (2002) theoretical study) and were not observed for the 2LO initialization as it can be seen in Fig.21 of Bouhafid et al. (2024). However, Widnall instability for vortex pairs, theorized in Tsai and Widnall (1976), may also occur as observed in Holzapfel et al. (2001), despite not being observed in a comparable case studied by Picot et al. (2014).  
 505 The absence of Widnall instability on the vortex pair for the 2LO initialization can be explained as follows.

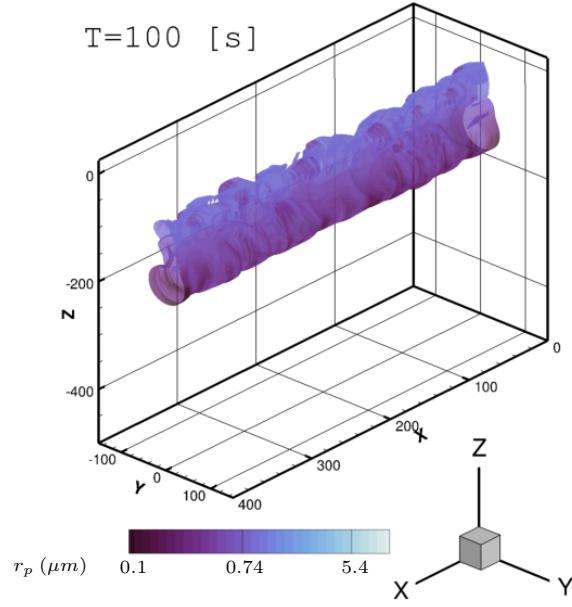
Widnall instability occurs in vortex pairs and is due to a resonance phenomenon between the transverse shear caused by one vortex on the other, and the existence of Kelvin waves with the same wavenumber  $k$  and azimuthal wavenumbers  $m$  and  $m+2$ . For a pair of Lamb-Oseen vortices, Sipp and Jacquin demonstrated in Sipp and Jacquin (2003) that a narrow band of Widnall instabilities exists only for Kelvin waves with  $m = \pm 1$ . The width of this instability band thus depends on the wavenumber  
 510  $k_w$  of the Kelvin modes, the ratio  $r_c/b_0$  between the vortex core radius and the distance between the vortices, and two real constants. The instability bands for the first three modes are given by:

$$2.26 - 8.68r_c^2/b_0^2 < kr_c < 2.26 + 8.68r_c^2/b_0^2 \quad (17)$$

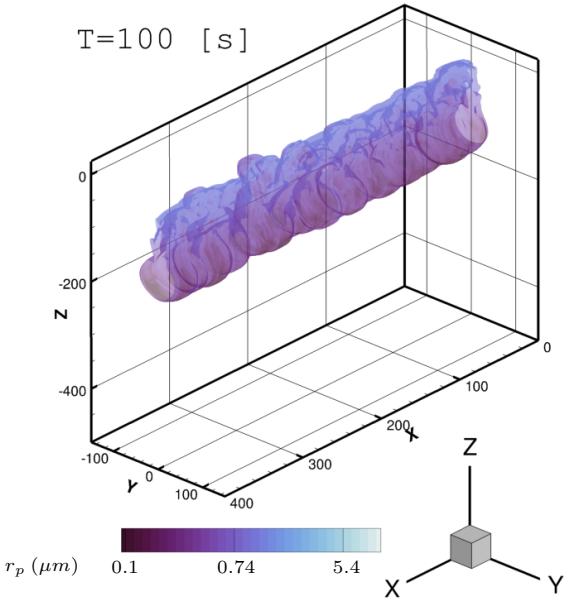
$$3.96 - 14.4r_c^2/b_0^2 < kr_c < 3.96 + 14.4r_c^2/b_0^2 \quad (18)$$

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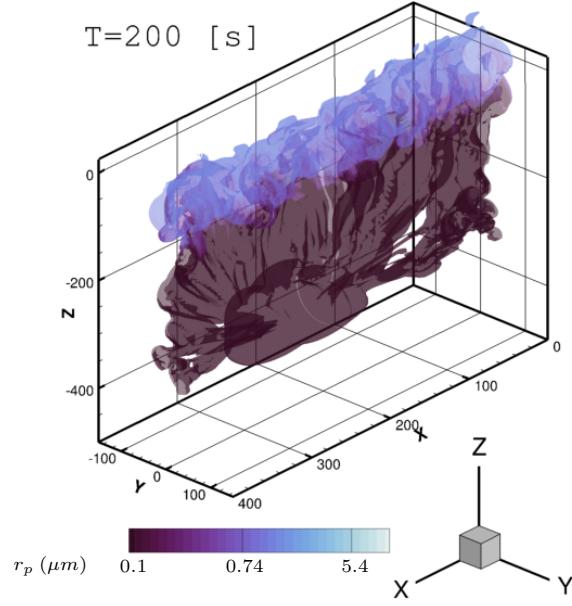
$$5.61 - 20.1r_c^2/b_0^2 < kr_c < 5.61 + 20.1r_c^2/b_0^2 \quad (19)$$



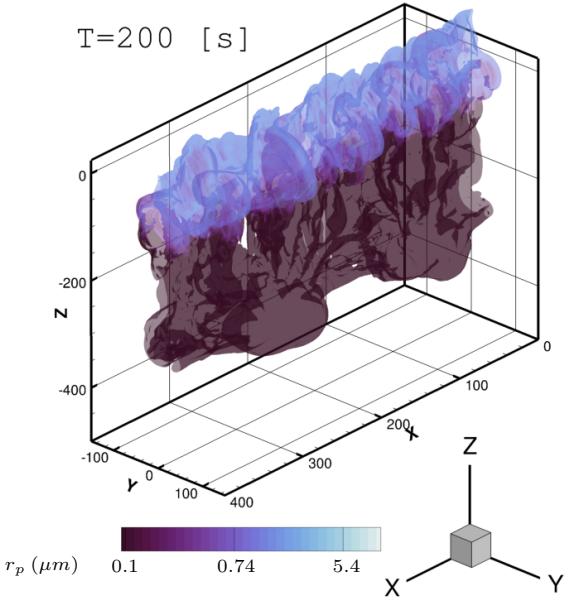
(a)  $t/t_b = 3.89$ , RANS initialization



(b)  $t/t_b = 3.89$ , 4LO initialization



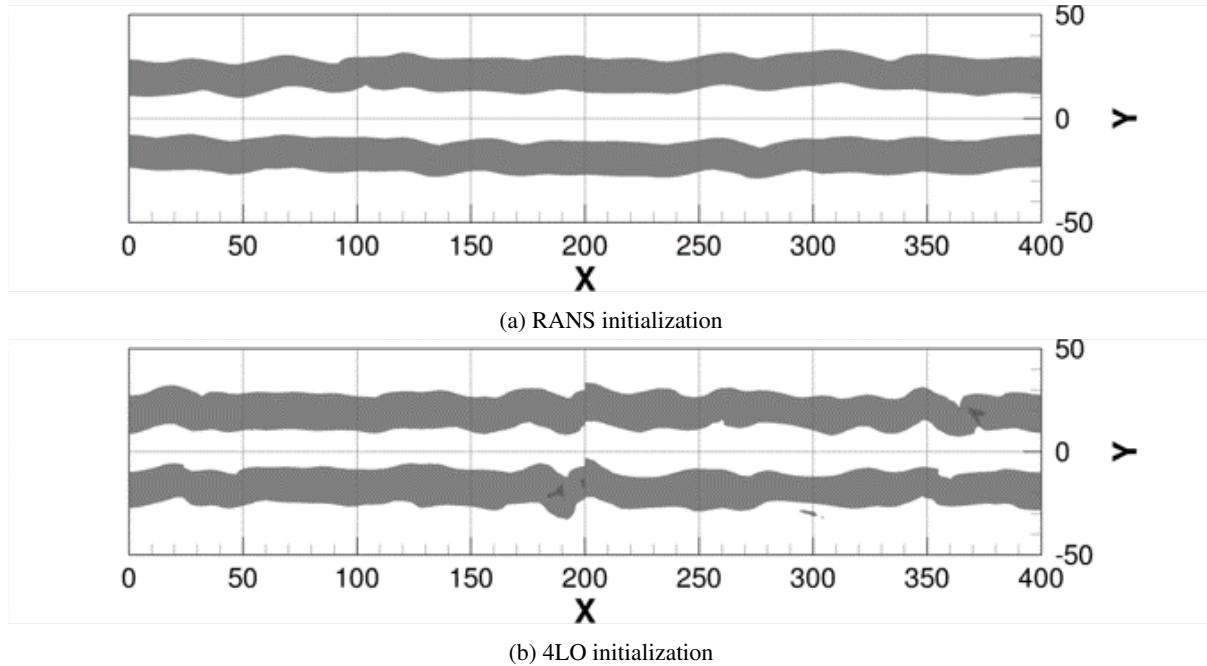
(c)  $t/t_b = 7.78$ , RANS initialization



(d)  $t/t_b = 7.78$ , 4LO initialization

**Figure 17.** Soot number density isosurface  $N_s = 10^7 \text{ m}^{-3}$  colored by mean crystal radius  $r_p$  in logarithmic scale.  $N_b = 0.012 \text{ s}^{-1}$ . Comparison between RANS initialization strategy and 4LO initialization strategy. Distances are in meters. ( $t_b = 25.7 \text{ s}$ ).

The numerical application for the case considered in this work gives  $r_c^2/b_0^2 = 5.75 \times 10^{-3}$  at  $t = 0$ , leading to very narrow instability bands. Consequently, the absence of Widnall instabilities in the LES using a dipolar analytical initialization from



**Figure 18.** 2D view of  $Q$ -criterion isosurface  $Q = 0.03 \text{ s}^{-1}$  at  $t/t_b = 3.89$  for RANS initialization and 4LO initialization. Distance are in meters. ( $t_b = 25.7 \text{ s}$ ).

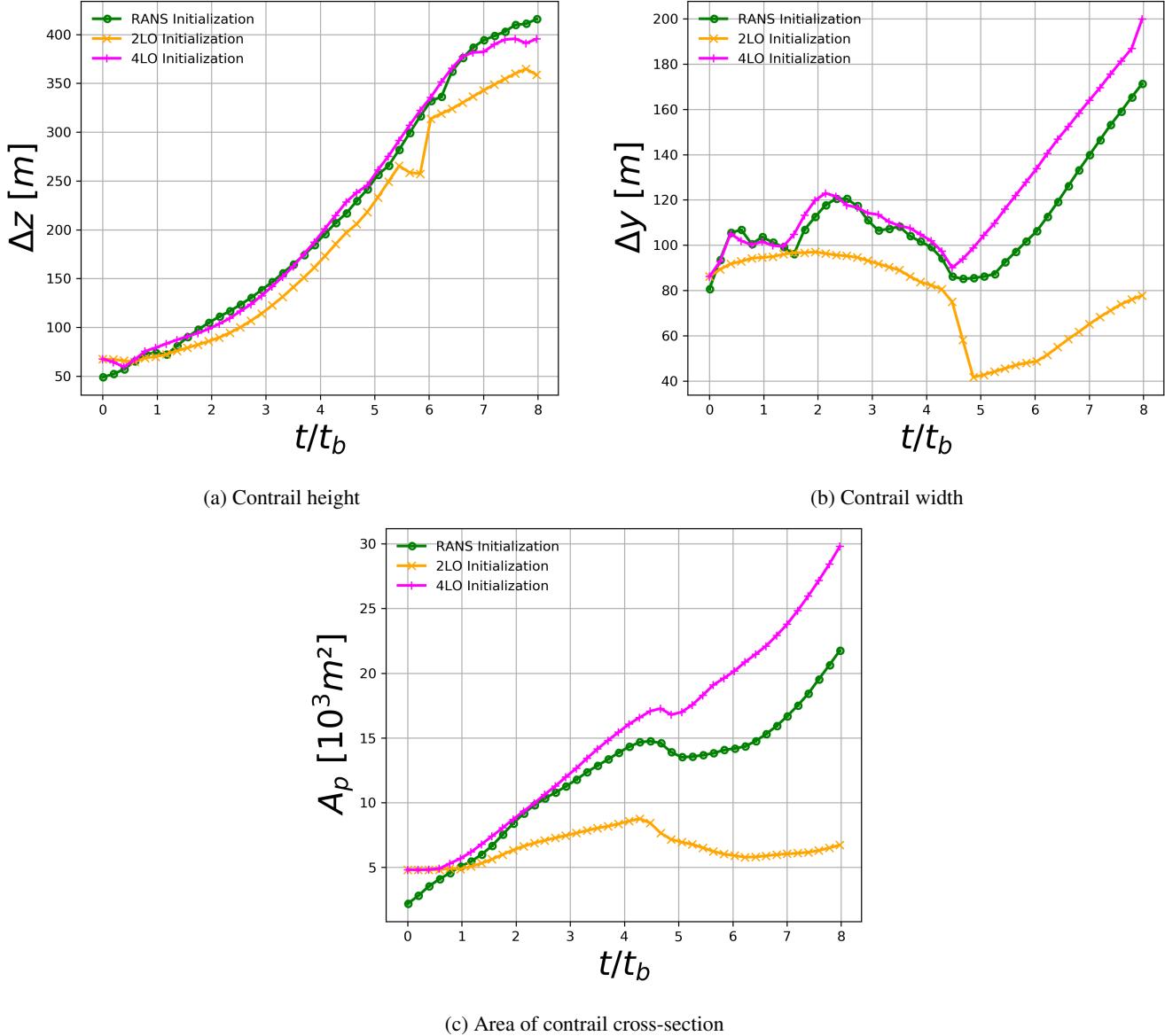
two Lamb–Oseen vortices could be explained by the fact that the perturbation wavenumber lies outside these instability bands.

520

The temporal LES simulation performed by Holzäpfel et al. (2001) using a dipolar analytical initialization strategy displayed the Widnall instability on the vortex pair. The appearance of these instabilities can be explained by the synthetic turbulence approach used to model wake turbulence and atmospheric turbulence. First, wake turbulence was added directly in the vortices while atmospheric turbulence was added in the whole domain. Second, the corresponding synthetic wake turbulence 525 fluctuations generated in the simulation by Holzäpfel et al. (2001) were distributed according to a normal distribution and are, therefore, not spatially correlated. As a result, their perturbation energy does not depend on their size and short-wave Widnall instability modes can potentially be excited with sufficient energy.

By contrast, the temporal LES simulation by Picot in Picot et al. (2014) also used a pair of Lamb–Oseen vortices to initialize the vortex regime but did not introduce synthetic wake turbulence inside the vortices. In addition, atmospheric turbulence 530 was generated through an Ornstein-Uhlenbeck forcing technique, which resulted in correlated turbulence following a realistic turbulence spectrum in a stratified fluid. No short-wavelength instability was then observed on the vortices.

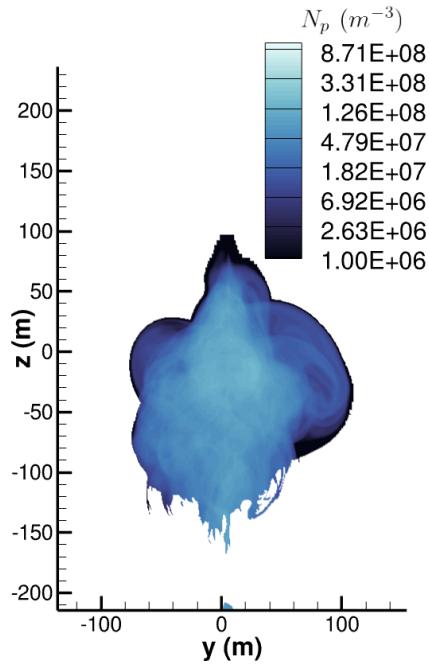
Not adding synthetic turbulence in the vortices for a far-field flow calculation can be justified by two main reasons. First, the jet plume concentrates the vast majority of the turbulent kinetic energy. This is actually a consequence of the second point, which is that turbulence generated in a Lamb–Oseen vortex is rapidly destroyed in the jet regime because of rotation effects 535 that were theoretically studied in Jacquin and Pantano (2002). Finally, in this work computations, the atmospheric turbulence is



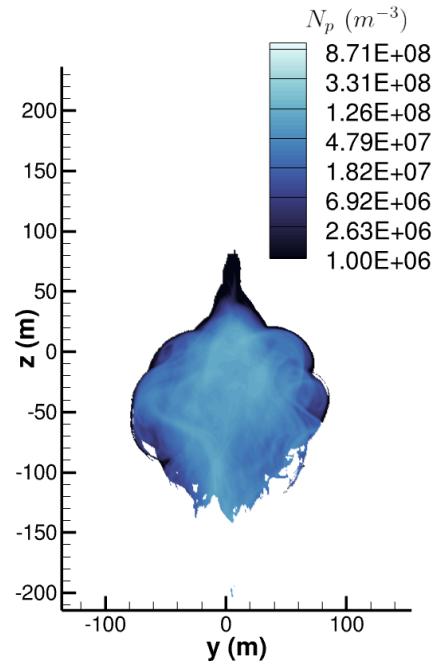
**Figure 19.** Contrail length scales for the different initialization strategies at  $N_b = 0.012 \text{ s}^{-1}$ . ( $t_b = 25.7 \text{ s}$ ).

much closer to the one in Picot et al. (2014) work than the one from Holzäpfel et al. (2001) in terms of turbulent kinetic energy spectra and spatial correlations. This too may justify the absence of Windall instability in the 2LO case. Thus, it is believed that it is not the Windall instability that is observed in the RANS and 4LO cases, but the shortwave instabilities resulting from the four-vortex nature of the wake (see Fabre et al. (2002)).

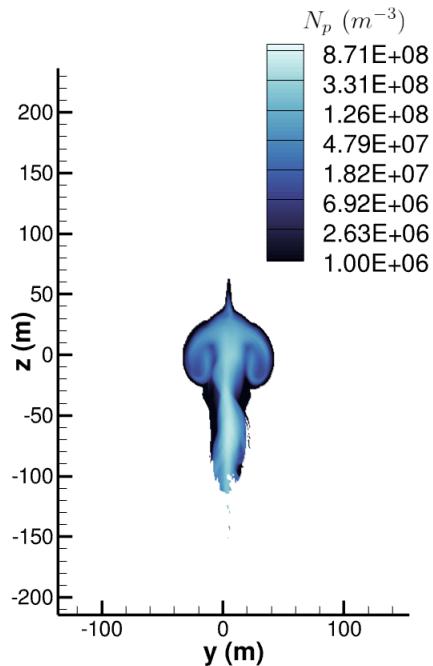
540 Fig.19 plots the relevant contrail dimensions as a function of time for the three different initialization strategies: RANS initialization, 2LO initialization and 4LO initialization. The corresponding spatially averaged cross-sections are plotted in



(a) 4LO initialization

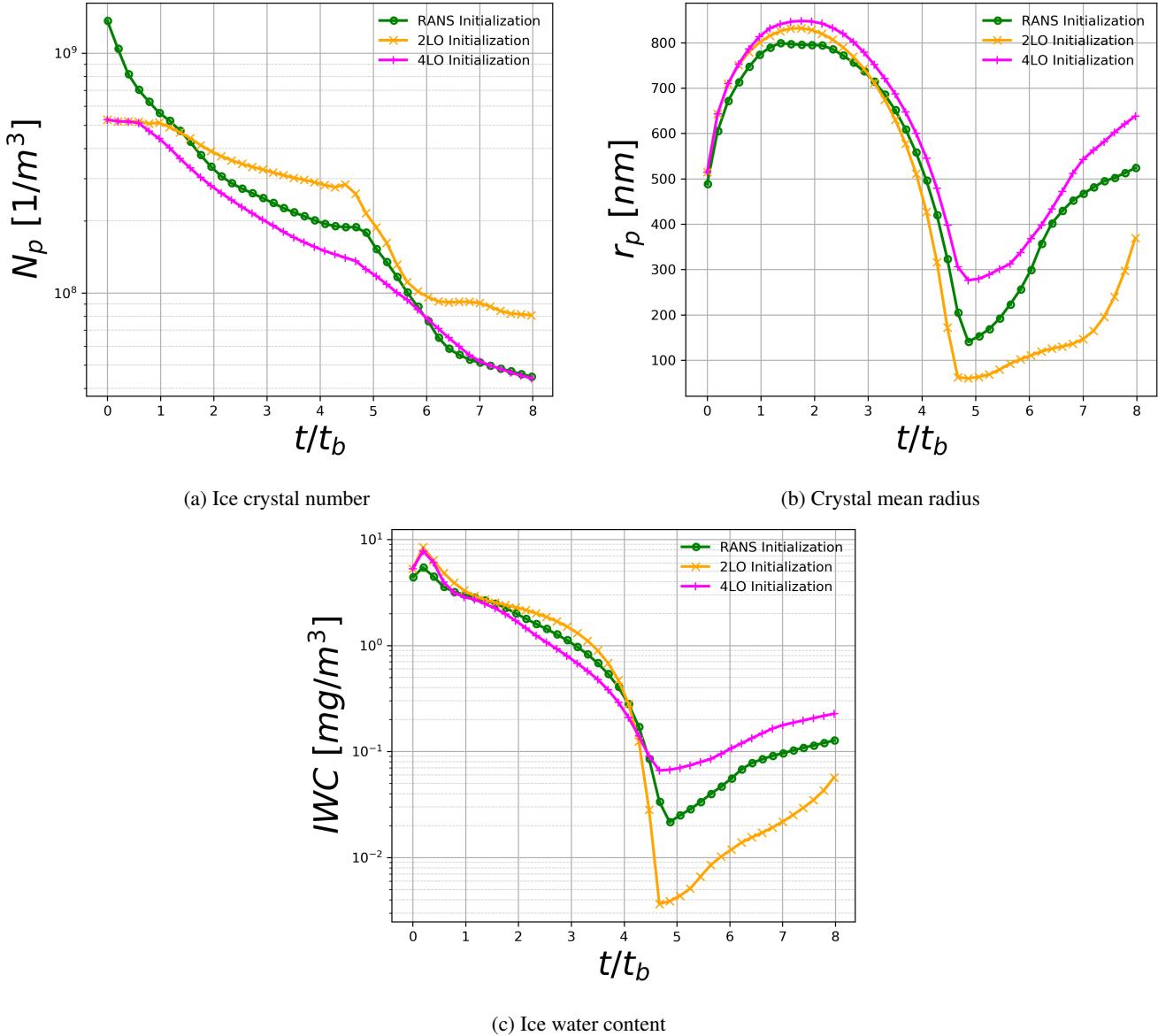


(b) RANS initialization



(c) 2LO initialization

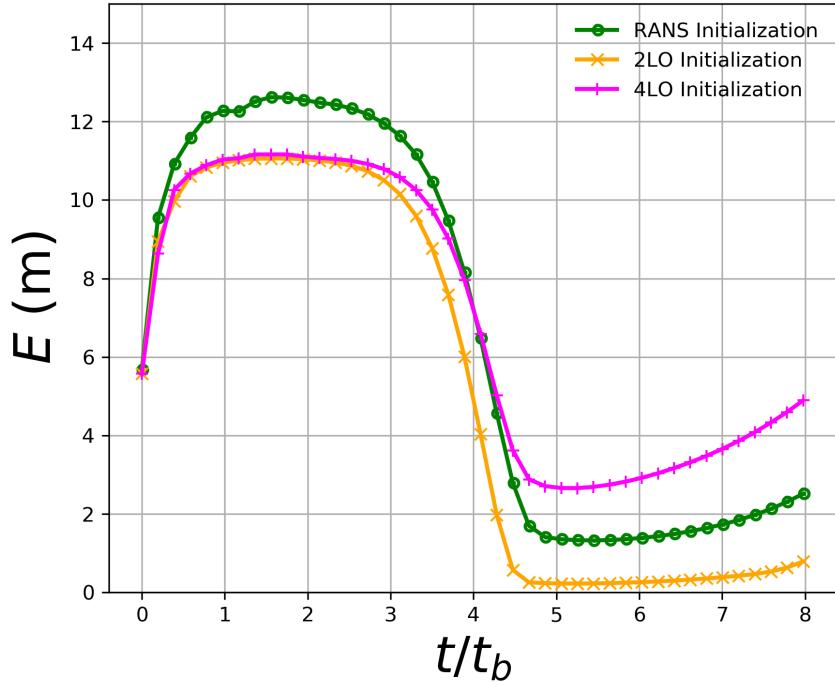
**Figure 20.** Spatially averaged ice crystal number field  $N_p$  in  $\text{m}^{-3}$  at  $t/t_b = 8$  ( $t_b = 25.7$  s) and for  $N_b = 0.012 \text{ s}^{-1}$ . Comparison between RANS, 2LO and 4LO initializations.



**Figure 21.** Average contrail microphysics quantities for the different LES initialization strategies at  $N_b = 0.012 \text{ s}^{-1}$ . ( $t_b = 25.7 \text{ s}$ ).

Fig.20 for the ice crystal number field. In terms of height and width, 4LO initialization gives a very similar contrail in terms of size compared to the RANS initialization. The contrail has a larger cross-sectional area for the 4LO initialization but is still much closer to the RANS initialization cross-sectional area than the 2LO initialization.

545 The evolution of averaged microphysical quantities is plotted in Fig.21. Despite the initial differences, the ice crystal number is almost identical at the end of the dissipation regime between the RANS and 4LO initialization. This emphasizes the additional mixing of the plume with ambient air caused by the turbulent nature of the wake. In terms of mean ice crystal radius and IWC,



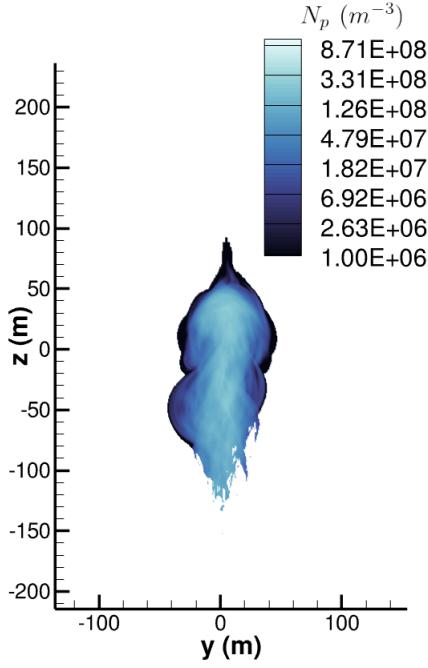
**Figure 22.** Average extinction as a function of time  $t$ . Comparison between the different initialization strategies for  $N_b = 0.012 \text{ s}^{-1}$ .

4LO initialization gives slightly higher values than RANS initialization at the end of the simulation. However, we observe that for the complete simulation, 4LO initialization results are qualitatively more similar to RANS initialization compared to 2LO initialization.

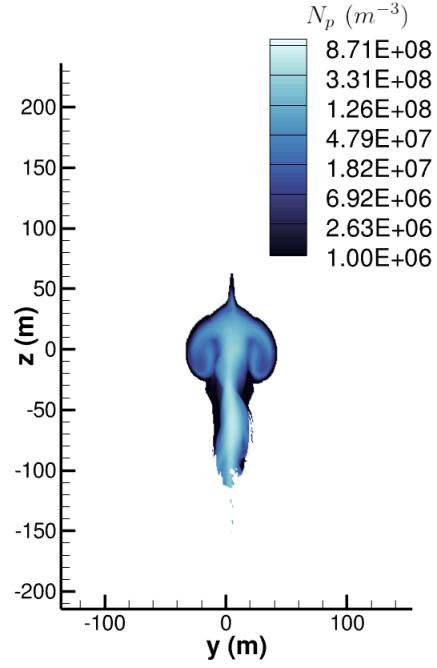
Fig.22 displays the evolution of the spatially averaged extinction  $E$  with time. The largest extinction is obtained for the 4LO initialization since the corresponding contrail was predicted to be wider with larger ice crystals on average. Contrary to 2LO initialization, the extinction of recently formed contrails will be overestimated. In light of that result, it appears that accurately estimating the radiative properties of recently formed contrails strongly depends on how the early aerodynamic of the near-field influences the contrail in the far-field. Horizontal tailplane vortices, which are commonly neglected in most temporal LES of contrails and aircraft wakes, seem to play an important role. The differences observed between the RANS initialization and 4LO initialization could possibly be explained by the difference in strength for the shortwave instabilities. Indeed, it appears that the vortex tube deformation amplitude is higher for the 4LO initialization as seen in Fig.18. This might result in a more turbulent secondary wake and a wider contrail.

#### 560 4.4 Influence of plume initialization

As previously indicated, a final LES simulation is performed where the far-field calculation is initialized using two Lamb-Oseen vortices and the plume obtained from the RANS calculation instead of a circular plume as done classically (2LO initialization



(a) 2LO initialization with RANS plume



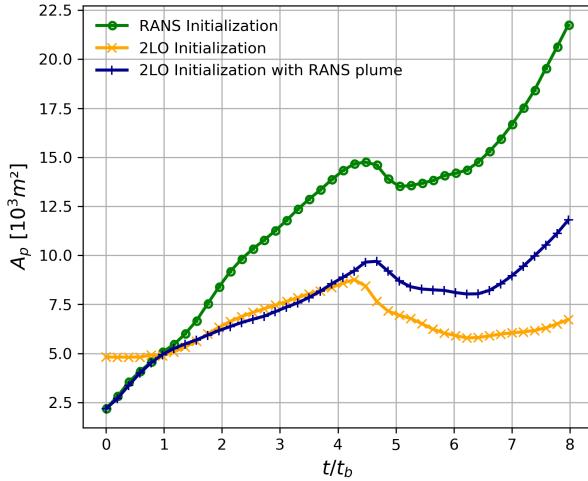
(b) 2LO initialization

**Figure 23.** Spatially averaged ice crystal number field  $N_p$  in  $\text{m}^{-3}$  at  $t/t_b = 8$  ( $t_b = 25.7$  s) and for  $N_b = 0.012 \text{ s}^{-1}$ . A logarithmic scale is used for the color levels.

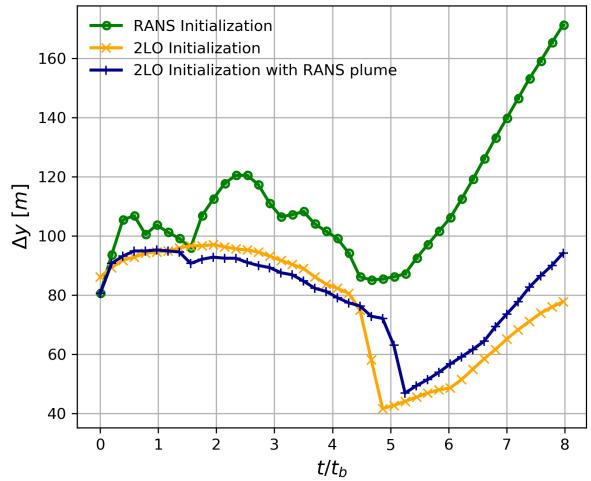
strategy). Such simulation allows us to investigate the effect of plume initialization for a contrail evolution. It will also enable a comparative analysis between the impact of plume initialization and the effect of vortex initialization, thereby elucidating their 565 relative contributions to contrail development.

The ice crystal number field on the slice obtained from spatially averaged the contrail along flight direction is plotted in Fig.23 at  $t/t_b = 8$  ( $t = 200$  s) for the 2LO initialization and the 2LO initialization with RANS plume. Quantitative and qualitative differences can be observed. Indeed, the contrail is larger when the RANS plume is used instead of a simple circular plume. Moreover, a slightly smaller value for the ice crystal number is obtained in the case of the RANS plume. In other words, 570 the contrail is more diluted when the plume is initialized through the RANS calculation of the near-field. However, it appears that the contrail height is very similar between the two cases. Initializing the plume through a circular jet or the deflected jet obtained from the RANS calculation has little effect on how far the ice crystals descent. More precisely, the vortex dynamics and the stratification exert a stronger influence than the choice of initializing the plume with a circular jet or with a more complex jet obtained from a precursor RANS simulation.

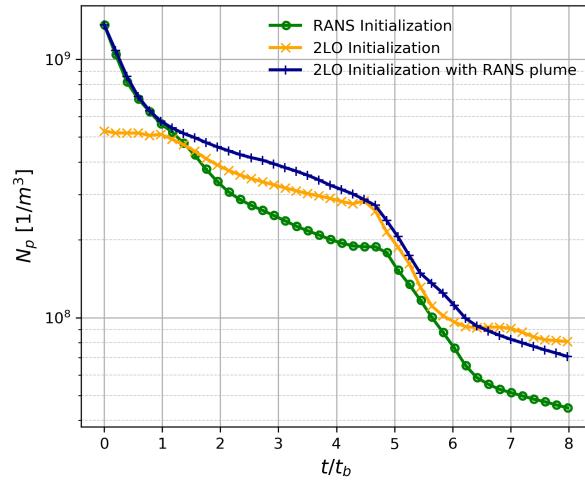
575 More quantitative comparisons can be made by plotting the evolution of the contrail spatial dimensions and the ice crystal number  $N_p$  in the spatially averaged contrail. These quantities are plotted in Fig.24 for the 2LO initialization, 2LO initialization with RANS plume and the RANS initialization which is our reference case. We observe that for the vortex regime, that is



(a) Contrail cross-sectional area



(b) Contrail width



(c) Ice crystal number

**Figure 24.** Average contrail spatial dimensions and ice crystal numbers. Comparison between the 2LO initialization and the 2LO with RANS plume initialization strategies.

$t/t_b < 4$ , the contrail cross-sectional area and the contrail width are very similar between the 2LO initialization and the 2LO initialization with RANS plume despite the initial discrepancies. Some differences arise after the vortex regime. The contrail obtained with the 2LO initialization with the RANS plume is larger than the one obtained with a circular plume. However, the cross-sectional area and width evolution show very similar trends between the two initializations. Moreover, the differences

observed are significantly smaller between the 2LO and the 2LO with RANS plume than between the 2LO and the RANS initialization strategies, indicating a greater impact of the vortex initialization.

Concerning the evolution of the ice crystal number  $N_p$ , the 2LO initialization with RANS plume show slightly larger values for  $N_p$  during the vortex regime compared to the classical 2LO initialization. However, these differences decrease for  $t/t_b > 4$  and the values for  $N_p$  become very close at the end of the simulation between the two 2LO initialization strategies. However, significant differences can be observed between the two 2LO initialization strategies and the RANS initialization strategy. The vortex initialization appears to have a greater impact on the contrail microphysics than the plume initialization.

## 5 Conclusions

This work extended the methodology of Bouhafid et al. (2024) to simulate the formation and early evolution of contrails for a realistic aircraft—from the formation of initial ice crystals to several minutes thereafter—without relying on analytical initialization classically used in the literature. Instead, we directly solved the near-field aerodynamic wake and contrail formation using a RANS approach with a Reynolds-Stress turbulence model and then transitioned to far-field evolution with a temporal LES augmented by synthetic turbulence.

Applied to the B777-like CRM geometry under two stratification scenarios, our method revealed significant differences compared to classical analytical initialization using Lamb-Oseen vortices. In particular, RANS initialization produced a more turbulent contrail with greater size, larger ice crystals, and larger extinction. This behavior is attributed to short-wave instabilities inherent to a quadripolar vortex system in a counter-rotating configuration. A subsequent LES using four Lamb-Oseen vortices (representing two main wing-tip vortices and two HTP vortices) yielded contrail characteristics qualitatively similar to those from the RANS initialization. Moreover, a LES simulation was performed using two Lamb–Oseen vortices together with the plume obtained from the RANS calculation, and it was found that vortex initialization exerted a stronger influence on the early contrail evolution than plume initialization.

It is important to note that these results are valid only for recently formed contrails under moderately ice-supersaturated conditions. Future work should assess whether these differences persist during the diffusion regime using Paoli et al. (2017) methodology. The differences observed in the early part of a contrail life cycle may potentially decrease, or even vanish, in the diffusion regime as atmospheric processes become predominant. Moreover, for calculations of the diffusion regime, it will be essential to incorporate microphysical models that account for more accurate soot activation, freezing, ice particle inertia, and background aerosol effects. Additionally, further investigation into the characteristics of HTP vortices and the impact of engine positioning (see Ramsay et al. (2024); Saulgeot et al. (2023)) on contrail development could inform more effective contrail mitigation strategies through optimized aircraft design.

*Author contributions.* **YB:** Writing – review & editing, Writing – original draft, Visualization, Validation, Methodology, Investigation, Formal analysis, Data curation, Conceptualization. **NB:** Writing – review & editing, Validation, Methodology, Supervision, Formal analysis, Conceptualization.

*Competing interests.* The authors declare that they have no known competing financial interests or personal relationships that could have  
615 appeared to influence the work reported in this paper.

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