



Stratospheric and upper tropospheric measurements of long-lived tracers and photochemically active species with GLORIA-B

Gerald Wetzel¹, Anne Kleinert¹, Sören Johansson¹, Felix Friedl-Vallon¹, Michael Höpfner¹, Jörn Ungermann², Tom Neubert³, Valéry Catoire⁴, Cyril Crevoisier⁵, Andreas Engel⁶, Thomas Gulde¹, Patrick Jacquet⁴, Oliver Kirner⁷, Erik Kretschmer¹, Thomas Kulesa³, Johannes C. Laube², Guido Maucher¹, Hans Nordmeyer¹, Christof Piesch¹, Peter Preusse², Markus Retzlaff², Georg Schardt³, Johan Schillings³, Herbert Schneider³, Axel Schönfeld², Tanja Schuck⁶, Wolfgang WoIWode¹, Martin Riese^{2,8}, and Peter Braesicke^{1,*}

¹Institute of Meteorology and Climate Research Atmospheric Trace Gases and Remote Sensing (IMKASF), Karlsruhe Institute of Technology, Karlsruhe, Germany

²Institute of Climate and Energy Systems - Stratosphere (ICE-4), Forschungszentrum Jülich, Jülich, Germany

³Institute of Technology and Engineering (ITE), Forschungszentrum Jülich, Jülich, Germany

⁴Laboratoire de Physique et Chimie de l'Environnement et de l'Espace (LPC2E/CNRS), Université Orléans, Orléans, France

⁵Laboratoire de Météorologie Dynamique, IPSL, CNRS, Palaiseau Cedex, France

⁶Institute for Atmospheric and Environmental Sciences, Goethe Universität, Frankfurt, Germany

⁷Scientific Computing Center (SCC), Karlsruhe Institute of Technology, Karlsruhe, Germany

⁸Institute for Atmospheric and Environmental Research, University of Wuppertal, Wuppertal, Germany

*now at: Business Area Research and Development, Deutscher Wetterdienst, Offenbach am Main, Germany

Correspondence to: Gerald Wetzel (gerald.wetzel@kit.edu)

Abstract

The Gimballed Limb Observer for Radiance Imaging of the Atmosphere (GLORIA) is a limb-imaging Fourier-Transform Spectrometer (iFTS) providing high-resolution mid-infrared spectra in the 780-1400 cm⁻¹ wavenumber range. Originally designed for aircraft, GLORIA has been deployed in eight research campaigns to date. To extend its observational range from the middle troposphere to the middle stratosphere, the instrument was adapted for a stratospheric balloon platform. GLORIA-B completed its first flight from Kiruna (Sweden) in August 2021 and a second from Timmins (Canada) in August 2022 as part of the EU Research Infrastructure HEMERA. The main objectives of these flights were technical qualification and the provision of a first imaging hyperspectral limb-emission dataset from 5 to 36 km altitude. This study evaluates the performance of GLORIA-B using vertical volume mixing ratio (VMR) profiles from the August 2021 flight, focusing on trace gases like O₃, CH₄, CFCs, HCFC-22, and SF₆.



35 Comparisons with in-situ measurements (ozonesonde, MegaAirCore, and cryosampler) show
agreement within 10 % for O₃, CH₄, SF₆, and CFC-12, and within 10-20 % for CFC-11, HCFC-
22, and CFC-113 up to 18 km, with larger deviations above this altitude. Another objective is
analyzing diurnal changes in photochemically active species (N₂O₅, NO₂, ClONO₂, BrONO₂).
Observed VMR variations align well with simulations from the EMAC (ECHAM5/MESSy
40 Atmospheric Chemistry) chemistry-climate model, though absolute concentrations differ to a
certain extent. Using the nighttime BrONO₂-to-Br_y ratio from simulations, the observed lower
stratospheric Br_y amount was estimated to 20.4 ± 2.5 pptv. In summary, the successful
deployment of GLORIA on a stratospheric balloon platform has demonstrated the capability of
the limb-imaging technique to provide high-quality vertical trace gas profiles up to the middle
45 stratosphere, contributing to a better understanding of the distribution and temporal evolution
of key atmospheric species.

1 Introduction

Balloon-borne mid-infrared limb-emission remote sensing measurements are an essential tool
50 to monitor the temporal and spatial distributions of many trace gases in the stratosphere and
upper troposphere with high vertical resolution below the flight altitude (Fischer et al., 2008).
Limb-emission spectrometers have the great advantage with respect to solar occultation
instruments that they can be used in a flexible way during day and night conditions, being
independent of any light source (Friedl-Vallon et al., 2004). However, cooled spectrometers
55 demand more technical effort compared to uncooled devices. This is a reason why the
stratospheric balloon missions of such instruments were always limited to a small number (Carli
et al., 1984; Murcray et al., 1984; Brasunas et al., 1988; Johnson et al., 1995). For instance,
Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) instruments were
developed at the Karlsruhe Institute of Technology (KIT). Beside an aircraft instrument, two
60 balloon versions of MIPAS (Fischer and Oelhaf, 1996; Friedl-Vallon et al., 2004) participated
in many scientific campaigns between 1989 and 2014 (e.g. Oelhaf et al., 1994; von Clarmann
et al., 1993; Wetzel et al., 1995; von Clarmann et al., 1997; Wetzel et al., 1997; Stowasser et
al., 1999; Waibel et al., 1999; Höpfner et al., 2002; Stowasser et al., 2002, 2003; Wetzel et al.,
2002; Wiegele et al., 2009; Wetzel et al., 2010, 2012, 2015, 2017). In addition, these MIPAS
65 instruments were precursors of the MIPAS Envisat satellite mission (Fischer et al., 2008) and
important components for the validation of MIPAS Envisat (e.g. Wetzel et al., 2022) and other



satellite instruments like Global Ozone Monitoring by Occultation of Stars (GOMOS; Renard
et al., 2008), the second Improved Limb Atmospheric Spectrometer (ILAS-II; Wetzell et al.,
2006), and the Superconducting Submillimeter-Wave Limb-Emission Sounder (SMILES;
70 Sagawa et al., 2013).

While limb-scanning spectrometers like MIPAS need some time to measure a full vertical trace
gas profile, the new generation of imaging spectrometers capture a vertical profile at once and
even with enhanced vertical resolution. This advantage led to the first aircraft-borne
development of the Gimballed Limb Observer for Radiance Imaging of the Atmosphere
75 (GLORIA; Friedl-Vallon et al., 2014; Riese et al., 2014) realized in a collaboration between
KIT and the Forschungszentrum Jülich (FZJ). In order to enhance the vertical range of GLORIA
to observations in the middle stratosphere albeit still reaching down to the middle troposphere,
the instrument was adapted to measure from a stratospheric balloon platform (called
GLORIA-B). The first flight was performed from Kiruna (northern Sweden) in August 2021.

80 Here we show the technical qualification and the provision of a first imaging hyperspectral
limb-emission dataset from 5 to 36 km altitude. We present examples of selected retrieval
results including uncertainty estimates, altitude resolution as well as long-lived tracer volume
mixing ratio (VMR) comparisons to external datasets. In addition, diurnal variations of
photochemically active gases are compared to simulations of the chemistry-climate model
85 ECHAM5/MESSy Atmospheric Chemistry (EMAC; Jöckel et al., 2010).

2 GLORIA-B instrument, data analysis and modelling

In the following subsections, we give an overview of the GLORIA balloon instrument, together
with the corresponding data analysis and a description of the atmospheric modelling performed
90 for this study.

2.1 GLORIA-B instrument

GLORIA-B is a modification of the airborne instrument (GLORIA-AB). It is a cryogenic limb-
imaging Fourier-Transform Spectrometer (iFTS) operating in the thermal infrared spectral
region in the wavelength range between 7 μm and 13 μm using a two-dimensional detector
95 array. It observes 128 vertical and 48 horizontal interferograms per measurement. The
instrument provides two-sided interferograms with a maximum optical path difference (MOPD)



of up to 8 cm, corresponding to a spectral sampling of 0.0625 cm^{-1} (Friedl-Vallon et al., 2014; Riese et al., 2014). The balloon borne version of GLORIA has been developed using the airborne version as basis. The balloon application calls for some fundamental modifications in the design of the instrument. Image rotation by roll movements of the carrier is small and the azimuth viewing direction can be controlled by the gondola, such that only the elevation needs to be steered and stabilised by the instrument. This allows for significant simplification by a replacement of the gimballed frame with a frame with an elevation axis only. On the other hand, the weight of the instrument should be as low as possible for the balloon application. Power consumption is not an issue on an aircraft but batteries on a balloon provide only a limited amount of power. At the high float altitude of a stratospheric balloon with pressure levels of only a few hPa, convective cooling does not work anymore, therefore the thermal management has to be performed by radiation only. The electronic system must be modified to take these requirements and the different temperature and pressure ranges into account. The major modifications with respect to the aircraft instrument are outlined in the following.

2.1.1 Mechanical modifications

The following mechanical modifications have been implemented to integrate the GLORIA spectrometer in the balloon gondola:

- The function of the outer frames of the gimbal frame is taken by the steering of the gondola. The two outer frames, azimuth frame and the image rotation frame, are thus omitted. The design of the innermost frame, the elevation frame, including bearings and direct drive, has been used in the new carrier for GLORIA-B.
 - The instrument mounting structure has been manufactured almost completely from aluminium and optimised regarding weight.
 - The junction between the mounting structure and the gondola is realised by six steel wire rope isolators.
 - The Inertial Measurement and Control Unit (IMCU) is thermally insulated from the mounting structure by a ring made from glass reinforced plastic.
 - The blackbody system has been completely redesigned (see calibration section).
- During the design process, iterative calculations with finite element analysis provided the optimal solution regarding material utilisation, required rigidity, and sufficient stability. This



approach ensures compliance with both the scientific specification regarding the stability of the line of sight and the safety requirements imposed by the balloon operator (CNES).

The final design results in a total mass of 128 kg, roughly half of the mass of the GLORIA-AB, with the centre of gravity centrally located within the structure. The dimensions in length x width x height are as follows: 1290 mm x 1200 mm x 580 mm. Figure 1 shows the carrier with all its components mounted. The instrument integrated into the balloon gondola is displayed in Fig. 2.

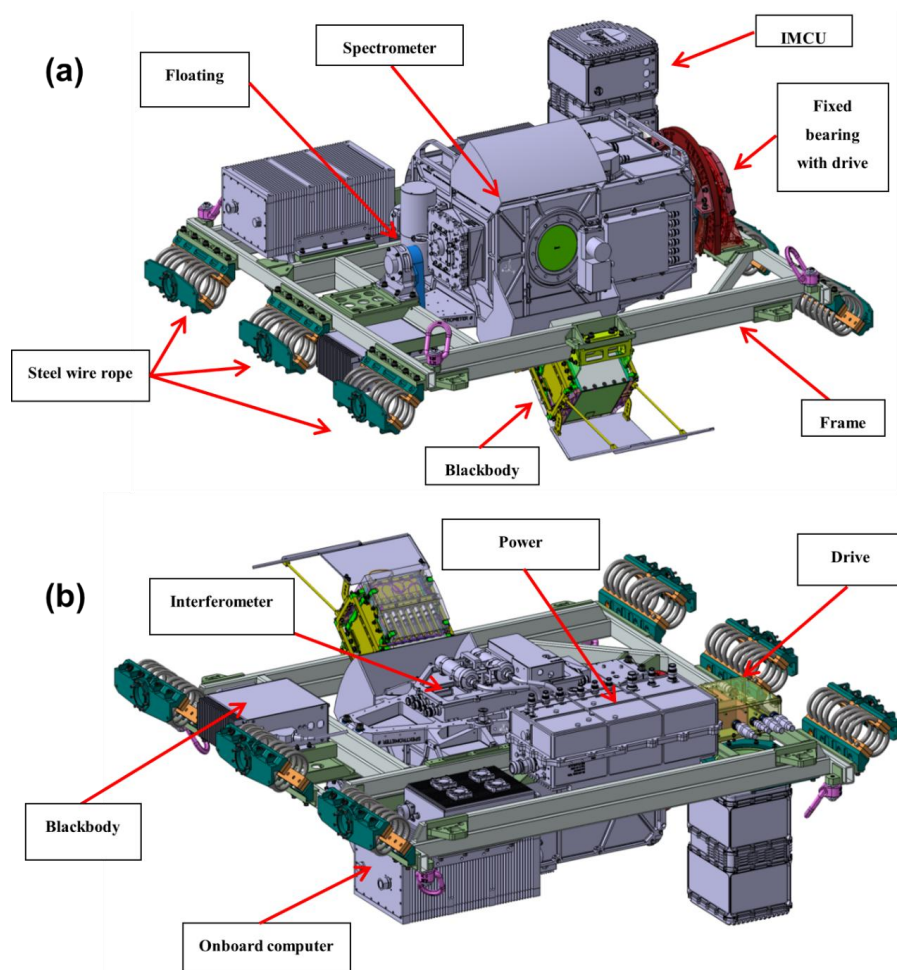


Figure 1. GLORIA-B instrument carrier with different mounted components as seen from top (a) and bottom (b).

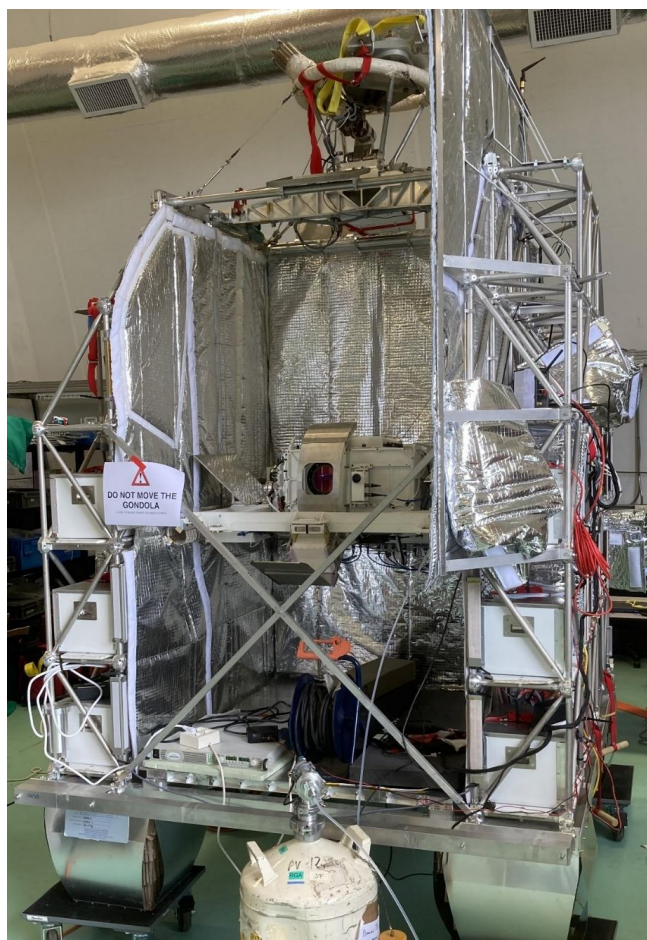


Figure 2. GLORIA-B instrument after integration into the balloon gondola.

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2.1.2 Modifications of the pointing system:

The new elevation frame allows the turning of the viewing axis within an angular range from -105° to $+75^\circ$. It allows pitch stabilization and pointing to different targets: limb, deep space, calibration blackbody, and nadir. In the park position at -40° the interferometer's view is directed into the black body, enabling calibration. The azimuth of the whole gondola is directed and stabilized by an independent azimuth control system operated by a dedicated Centre National d'Études Spatiales (CNES) team. The amplitude of the oscillations in roll and pitch on float height amount only to some tenths of a degree. Therefore, the roll angle is not

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150 stabilized, however measured by the navigation system, such that it can be considered in the retrieval. The software for the IMCU was adapted to this single axis task by the developer iMAR Navigation GmbH, Germany. A new additional communication interface was implemented to have the possibility for direct control of the azimuth stabilization system of CNES by GLORIA-B for special measurement options.

2.1.3 Modifications of the thermal concept:

155 In contrast to the situation on the aircraft, there is no convective heat transfer at balloon float altitude (35 to 40 km), therefore only radiative cooling can be utilized. During ascent, however, the instrument is exposed to strong air flow, and care has to be taken that the electronics do not become too cold.

160 In order to design a thermal concept for the instrument, a ray tracing model has been used to calculate, which part of the surfaces of the components is exposed to space and which part is exposed to Earth or other parts of the gondola. Based on these results, thermal fluxes have been modelled and thermal coupling, insulation, radiative surfaces and reflection shields have been designed such that electronics and instrument stay within their respective operation ranges during ascent and float.

165 The spectrometer together with the elevation electronics and the interferometer electronics is mounted on the elevation plate, which serves as heat sink. The spectrometer housing itself is painted with white Nextel Suede Coating 3101 428-04, reflecting in the visible and emitting in the infrared, such that radiative cooling is possible whereas sun radiation onto the instrument does not lead to too much heating. The elevation electronics and the IMCU are equipped with
170 an insulation cover to avoid too much cooling during ascent. During float, heat from these components is conducted to the elevation plate.

In the aircraft instrument, detector and compressor are cooled with fans. In the balloon configuration, two heatpipes have been mounted to the detector finger and the compressor, respectively, which are connected to the elevation plate.

175 The onboard computer, which is also painted in white, has been mounted onto an extra plate to take up the heat. Both the plate and the onboard computer itself are exposed to space for radiative cooling. A reflective plate is mounted next to the spectrometer to allow for further radiation into space.



2.1.4 Electrical design

180 The electrical design of the GLORIA instrument is based on a distributed embedded system architecture, where individual components are coordinated and controlled via a communication network. By integrating autonomous control units within each functional module, significant advantages in terms of scalability, fault tolerance and real-time capability have been achieved.

A key feature of this architecture is the decentralized distribution of the computational load,
185 where multiple embedded control units independently process local operations, thereby relieving the on-board computer. This contributes to increased system reliability, as the failure of individual components does not necessarily affect the entire system. In addition, the modular structure allows efficient maintenance and expansion: defective modules can be selectively replaced and new functionalities can be integrated into the existing system without extensive
190 modifications.

In adapting GLORIA for use as a balloon instrument, a significant proportion of the existing system components could be retained unchanged. Essential modules such as the detector electronics, the visual camera, the interferometer electronics in the spectrometer, the pointing control unit (IMCU) and the direct control of the elevation drive remained largely in their
195 original design. Only the firmware of these components was reconfigured to meet the changed operational requirements.

In order to successfully integrate GLORIA-B into the balloon gondola, it was necessary to make specific modifications to the electrical interfaces and to optimize the power consumption in order to efficiently use the limited battery capacity. As part of these adaptations, a new compact
200 Power Distribution Unit (PDU) with extensive housekeeping data acquisition was developed, specifically tailored to the remaining sub-components of GLORIA-B. The implementation of highly efficient, isolated voltage converters and an integrated power sequencing system for individual functional components enabled intelligent power management. This ensures that power consumption remains largely constant across different measurement configurations,
205 avoiding additional peak loads on the batteries. The high energy efficiency of up to 90% of the new customized PDU has led to a significant reduction in the average total power consumption to 320 watts for the remaining sub-components. This represents a reduction of more than a factor of 3 compared to the original aircraft configuration.



Another modification concerns the on-board computer, which was previously located in the
210 passenger cabin of the aircraft and has now been integrated directly into the instrument carrier
(Neubert et al., 2021). The computing unit and the uninterruptible power supply (UPS) for
GLORIA-B are now in white, thermally conductive enclosures. Internal modules/circuit boards
have been partially sealed, with thermal hotspots managed by copper strips and heat sink plates.
The modified computing system is based on a 3U VPX-REDI single-board computer powered
215 by a low-power Intel® Core™ i7 processor, which controls the GLORIA-B instrument through
Gigabit Ethernet ports (Neubert et al., 2021). The storage capacity has been increased to 16 TB
through the implementation of four solid state drives to enable flights with an extended duration
of three to five days. The data storage is organized in a redundant RAID system. The real-time
measurement data is streamed via an FPGA-based accelerator board directly to the storage
220 system, where it can be efficiently downloaded through a 10 Gbit Ethernet downlink after each
measurement flight. This facilitates the timely execution of post-processing operations.

2.1.5 Calibration system

The GLORIA instrument deployed on aircraft is equipped with two calibration blackbodies,
one of them operated slightly below ambient temperature and one operated about 30 K above
225 ambient temperature. Both of them can be heated or cooled and temperature stabilised using
Peltier elements (Olschewski et al., 2013). Due to the restricted power budget on the balloon, a
new blackbody with passive cooling had to be developed. The high flight altitude and an
elevation angle of 40° for deep space measurements make it possible to derive the instrument
self-emission directly from these measurements, therefore one calibration blackbody for gain
230 determination is sufficient.

The pyramid structure of the backplane and the coating with Nextel Velvet Coating 811-21
have been inherited from GLORIA-AB, as well as the mounting of the temperature sensors
inside the backplane. The same type of temperature sensors and the same temperature
acquisition electronics as for GLORIA-AB have been used. The backplane measures 110 mm
235 x 110 mm, and the cavity has a length of 130 mm. The blackbody is almost completely made
of aluminium, and the parts are connected with closely spaced screws to provide good thermal
contact.

The top side of the blackbody is exposed to deep space, while the bottom side is exposed to
Earth (or the lower parts of the gondola), and the outer side of the backplane is partly exposed



240 to Earth and partly to space. In order to obtain a homogeneous temperature across the
backplane, the top and the backplane are painted with white Nextel Suede Coating 3101 428-
04. In addition, a blank radiation shield made from stainless steel has been mounted below and
behind the backplane to shield radiation from Earth and to reflect the radiation from the
backplane into space (Fig. 1). The left and right side of the blackbody are kept with blank
245 aluminium, and the bottom side has been insulated to avoid heating from Earth radiation. The
blackbody temperature during flight ranged from about 240 K to 270 K, and although a small
systematic gradient over the backplane can be observed, the overall homogeneity is still within
+/- 100 mK, which meets the requirement.

2.1.6 Measurement scenario and automation

250 Since the vertical angular coverage of the instrument does not cover the whole altitude range
of interest between flight altitude (about 36 km) and 5 km, two alternating elevation angles
were chosen to cover this altitude range. This leads to an overlap of the vertical coverage of the
two alternating measurements in the altitude range of about 26 to 29 km, varying with flight
altitude.

255 Interferograms are Fourier transformed into the spectral domain and calibrated using in-flight
measurements of a blackbody and deep space measurements. During the first flight a set of
sixteen blackbody measurements and eight deep space measurements was taken about every 31
minutes for radiometric calibration. In the second flight in August 2022, an additional sequence
of eight Nadir-looking measurements was included additionally in every measurement cycle.

260 An automated script ensures this sequence is followed without operator input. This script also
ensures that the two alternating elevation angles are interchanged every four measurements. In
specific measurement scenarios, the GLORIA-B system has the possibility to take over the
command and control of the azimuthal orientation of the gondola by sending commands to an
on-board interface provided by CNES. Such special measurement scenarios include constantly
265 turning the gondola at a slow, but specific and constant rate while running the standard
measurement sequence.



2.2 Generation of calibrated spectra

The level 1b processing from raw interferograms to calibrated spectra basically follows the procedure outlined by Kleinert et al. (2014) and Ungermann et al. (2022). Some important
270 modifications are outlined below.

Instrument offset determination:

In case of GLORIA-AB, deep space spectra are calibrated in a first step using the hot and the cold onboard blackbody. These pre-calibrated spectra are then used to perform a forward calculation of synthetic spectra representing the atmospheric contribution in the measurement.
275 This contribution is then subtracted from the measurement in order to correct for the instrument offset (Ungermann et al., 2022). In case of GLORIA-B, only one onboard blackbody is flown, but deep space spectra are much less affected by atmospheric radiation due to the higher flight altitude. Therefore, it is possible to perform a pre-calibration using the blackbody and the deep space measurement with low spectral resolution, such that the atmospheric lines are strongly
280 smoothed. The following fit and calculation of synthetic spectra is the same as for GLORIA-AB.

Spectral calibration:

In case of GLORIA-AB, spectral calibration and the determination of the off-axis angle for each pixel is based on the spectral position of the CO₂ laser lines around 950 cm⁻¹ using deep
285 space measurements. In case of GLORIA-B, the higher flight altitude and the larger elevation angle leads to much weaker line intensities in the spectra, such that the quality of the spectral calibration suffers from the low signal to noise ratio. Therefore, nominal atmospheric measurements taken with the upper elevation angle are used for this purpose. This proved to work well, because line intensities are sufficiently strong and even the lower pixel rows are not
290 affected by clouds.

Pixel binning and co-addition of spectra:

In case of GLORIA-AB, the spectra of all pixels of one row are averaged after filtering of bad pixels, such that the final level 1b product consists of one set of 128 vertical pixels. The same binning is done for the spectra of GLORIA-B, but in addition, eight consecutive measurements
295 (four each with the upper and lower elevation angle, respectively) are combined to one dataset. In order to obtain the combined dataset on a regular angular grid, the grid of the measurements that were taken lower down is extended vertically, and the spectra of the measurements that



were taken higher up are interpolated onto this grid, using linear interpolation in the vertical direction. In the overlapping range, a linear weighting factor is introduced, such that a smooth transition is established between the two datasets. For the binned and combined datasets, the noise equivalent spectral radiance (NESR) finally amounts to about $5 \times 10^{-9} \text{ W}/(\text{cm}^2 \text{ sr cm}^{-1})$. In the overlapping range, the NESR is reduced by a factor of 0.7.

2.3 Data analysis of measured spectra

In this study, we report results from the first GLORIA-B balloon flight carried out over northern Scandinavia on 21/22 August 2021 in the framework of the European Union Research Infrastructure HEMERA. Spectra were recorded with a maximum optical path difference of 8.0 cm which corresponds to a spectral sampling of 0.0625 cm^{-1} . They were apodized with the Norton and Beer (1976) “strong” function. The high spectral resolution allows for the separation of individual spectral lines from continuum-like emissions such that many species even with minor emission contributions to the spectra can be retrieved with high spatial resolution. Flight path of the balloon gondola and the observed geolocations are displayed in Fig. 3. The gondola was launched from Esrange near Kiruna (67.9° N , 21.1° E). The float of up to 36 km lasted more than 12 hours and covered some time before sunset and after sunrise of the following day (between 17:11 UTC and 05:39 UTC). Due to weak stratospheric winds, the gondola did not move far away from the launch pad. Azimuth rotations were carried out continuously to probe the air masses around the gondola.

In order to analyse measured spectra, forward radiances are calculated with the Karlsruhe Optimized and Precise Radiative transfer Algorithm (KOPRA; Stiller et al., 2002) based on spectroscopic parameters from the high-resolution transmission molecular absorption database (HITRAN; Gordon et al., 2017). The procedure KOPRAFIT (Höpfner et al., 2002) uses derivatives (Jacobians) of the radiance spectrum with respect to atmospheric state and instrument parameters as calculated by KOPRA. The inverse problem is solved with a Gauss-Newton iterative method (Nocedal and Wright, 2006) in combination with a Tikhonov-Phillips regularization (Phillips, 1962; Tikhonov, 1963) using a constraint with respect to a first derivative of the a priori profile of the target species. The retrieval grid was set to 0.25 km up to 40 km while increasing gradually above this altitude.

Cloud-affected spectra were filtered out by a cloud index (in general larger than 2) taking into account the colour ratio of the mean radiance between $788.20\text{--}796.25$ and $832.30\text{--}834.40 \text{ cm}^{-1}$



(Spang et al., 2004). In a further step, the pointing elevation angle was retrieved to check for a possible misalignment of the line of sight (LOS). The LOS error was estimated to 0.01° according to the method shown by Johansson et al. (2018). Prior to the trace gas analysis, the temperature retrieval was performed in spectral windows around 801, 811, 942, and 957 cm^{-1} containing appropriate CO_2 transitions. The total error of the temperature retrieval was estimated to about 1.5 K (Johansson et al., 2018). Table 1 gives an overview of the spectral windows used for the GLORIA-B target gas retrieval performed in connection with this study. A careful selection of appropriate spectral windows is essential to perform retrievals with good accuracy. Main interfering species were fitted simultaneously together with the target gas. Minor species were either pre-fitted in a different spectral interval, or climatological profiles were used during the retrieval process. The error estimation of the retrieved molecules includes random noise, temperature errors, calibration (multiplicative gain), pointing, and field of view (FOV) inaccuracies, errors of non-simultaneously fitted interfering species, as well as spectroscopic data errors. All errors refer to the 1- σ confidence limit.

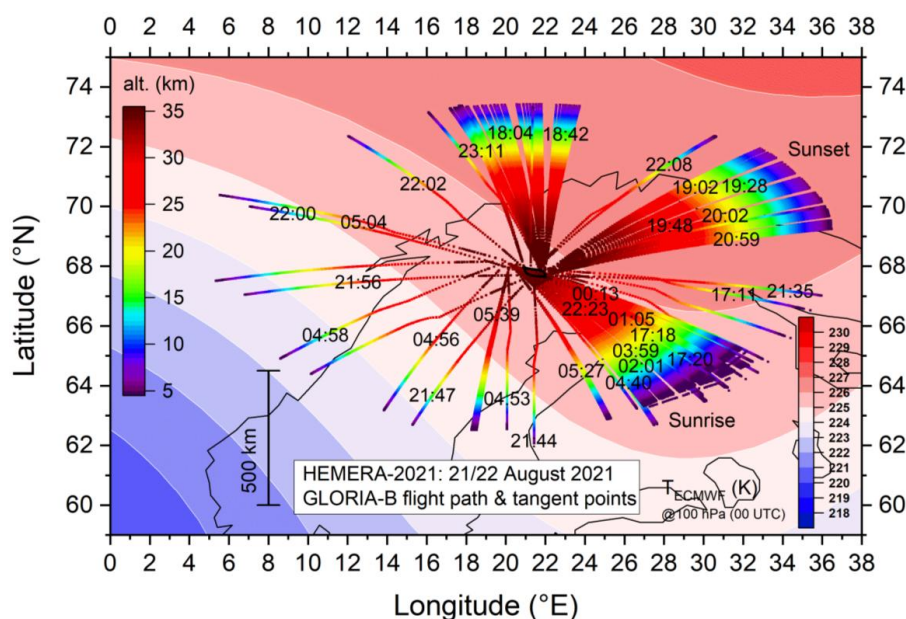


Figure 3. Flight path of the balloon gondola on 21/22 August 2021 (black line) and GLORIA-B tangent points (altitude colour coded). Measurement times are given in UTC. European Centre for Medium-Range Weather Forecasts (ECMWF) 100 hPa temperature field at 00 UTC is indicated in the background.



Table 1. Overview of GLORIA-B spectral windows for atmospheric species used in this study together with altitude resolution (alt. reso.), typical noise errors and total errors, valid for a single vertical profile.

Target molecule	Spectral range (cm ⁻¹)	Fitted species	Alt. reso. (km)	Noise error (%)	Total error (%)
O₃	763.5000 – 764.5625	O ₃	< 2	1 – 5	8 – 10
	766.6250 – 767.2500				
	780.6250 – 781.6875				
	782.1875 – 783.5000				
	787.0000 – 787.6250				
	788.4375 – 789.5625				
	806.8125 – 807.5000				
	822.0625 – 823.0625				
	964.8750 – 965.9375				
CH₄	968.0625 – 968.9375	N ₂ O, CH ₄	< 2	2 – 5	20 – 22*
	1215.8125 – 1217.0000				
	1227.8125 – 1228.5000				
	1231.1250 – 1232.5625				
CFC-11	1234.0000 – 1235.2500	H ₂ O, O ₃ , HNO ₃ , COCl ₂ , CFC-11	< 2	1 – 5	5 – 15
CFC-12	840.5000 – 859.6875	H ₂ O, HNO ₃ , CFC-12	< 2	1 – 5	7 – 15
HCFC-22	918.6250 – 923.5000	CO ₂ , O ₃ , HCFC-22	< 2	1 – 10	6 – 20
	803.5000 – 804.7500				
	808.2500 – 809.3750				
	820.5000 – 821.1250				
CFC-113	828.0000 – 829.5000	H ₂ O, O ₃ , C ₂ H ₆ , HCFC-22, CFC-113, ClONO ₂	< 4	3 – 20	7 – 30
	801.0000 – 819.5000				
SF₆	819.5000 – 830.0000	H ₂ O, O ₃ , COF ₂ , SF ₆	< 3	2 – 6	5 – 8
HNO₃	947.0000 – 948.7500	HNO ₃	< 2	1 – 5	15 – 20*
	862.0000 – 863.4375				
	866.1250 – 867.4375				
ClONO₂	901.3125 – 901.8125	O ₃ , ClONO ₂	< 2	3 – 8	4 – 9
N₂O₅	779.8750 – 781.0000	H ₂ O, O ₃ , N ₂ O, CH ₄ , HOCl, COF ₂ , N ₂ O ₅	< 2	2 – 4	5 – 8
NO₂	1227.2500 – 1241.3750	H ₂ O, O ₃ , NO ₂ , HNO ₃ , COF ₂ , HCFC-22, CCl ₄ , CFC-113, ClONO ₂ , HNO ₄ , PAN, BrONO ₂	< 5	2 – 5	5 – 15
HNO₄	801.0000 – 820.0000	H ₂ O, O ₃ , NO ₂ , HNO ₃ , COF ₂ , HCFC-22, CCl ₄ , CFC-113, ClONO ₂ , HNO ₄ , PAN, BrONO ₂	< 4	3 – 8	5 – 25
BrONO₂	801.0000 – 820.0000	H ₂ O, O ₃ , NO ₂ , HNO ₃ , COF ₂ , HCFC-22, CCl ₄ , CFC-113, ClONO ₂ , HNO ₄ , PAN, BrONO ₂	< 30	9 – 15	10 – 30
PAN	780.0000 – 790.0000	H ₂ O, O ₃ , HCFC-22, CCl ₄ , PAN	< 2	5 – 15	10 – 30
	794.0000 – 805.0000				



2.4 Model simulations

A multi-year simulation with the chemistry climate model ECHAM5/MESSy Atmospheric Chemistry (EMAC) was performed. This Eulerian model includes a number of submodels describing processes in the troposphere and middle atmosphere (Jöckel et al., 2010). The core
355 model is the fifth generation European Centre Hamburg general circulation model (ECHAM5; Roeckner et al., 2006) that is connected to the submodels using the interface Modular Earth Submodel System (MESSy). For this study, an EMAC (MESSy version 2.55.0) multi-year simulation from 1979 until 2022 was used with a time step of 10 minutes and a spherical truncation of T42L90 (corresponding to a horizontal resolution of approximately 2.8° by 2.8°
360 in latitude and longitude) with 90 hybrid pressure levels from the ground up to 0.01 hPa. The model was nudged towards the ECMWF ERA5 reanalysis data (Hersbach et al., 2020) to simulate realistic synoptic conditions. A comprehensive chemistry setup is included from the upper troposphere to the lower mesosphere. Rate constants of gas-phase reactions originate from Atkinson et al. (2007), Sander et al. (2011), and Burkholder et al. (2015). Photochemical
365 reactions of short-lived bromine containing organic compounds are integrated into the model set-up (Jöckel et al., 2016) together with surface emissions of these species taken from scenario 5 of Warwick et al. (2006). The boundary condition of the greenhouse gases (including nitrogen oxides) was chosen according to the Coupled Model Intercomparison Project Phase 6 (CMIP6; Eyring et al., 2016). The model output was interpolated in time and space to the GLORIA-B
370 measurement geolocations during the first flight.

3 GLORIA-B measurements and external data comparison

In this section mean vertical profiles of O₃ and longer-lived trace gases retrieved from GLORIA-B measurements are compared to coincident in-situ ozonesonde, AirCore (Engel et al., 2017; Membrive et al., 2017; Laube et al., 2020), and cryosampler (Engel et al., 2016;
375 Schuck et al., 2025) instruments. An AirCore is a passive sampling device based on a long coil of extremely thin-walled stainless-steel tubing (Karion et al., 2010). The newly developed MegaAirCore consists largely of two connected pieces of coiled tubing (Laube et al., 2025). In contrast, the well-established whole air cryosampler holds 15 stainless steel sample flasks
380 which were evacuated before the launch and opened and closed interactively during flight through telemetry commands from ground (Schuck et al., 2025). For CH₄ we also compared GLORIA-B data to SPECIES (SPECtromètre Infrarouge à lasErs in Situ) in-situ observations



(Catoire et al., 2023). In the case of O₃, an additional comparison to Microwave Limb Sounder (MLS; e.g. Livesey et al., 2008; Hubert et al., 2016) satellite observations was performed. For this intercomparison, all GLORIA-B data points were averaged on a common altitude grid. This has the advantage that the noise part of the retrieval error is reduced by a factor $1/n^{0.5}$ where n is the number of averaged data points. To determine whether a bias between the instruments is significant or not, combined errors which are calculated as the square root of the sum of the squares of the errors of each measurement are calculated according to Wetzel et al. (2022). To allow for a more realistic comparison in terms of altitude resolution, the data of the in-situ measurements were smoothed with the averaging kernel of the GLORIA-B retrieval.

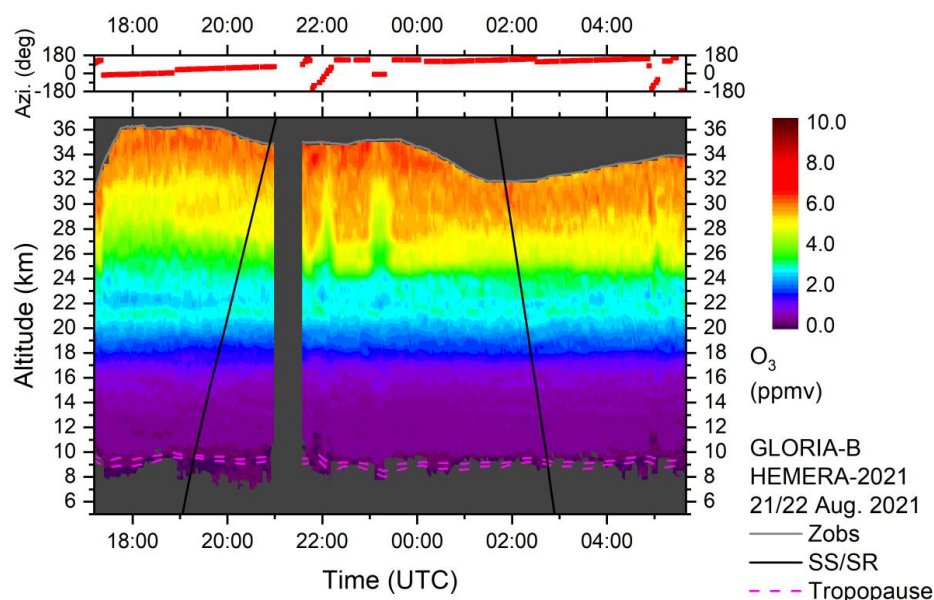


Figure 4. Top: Azimuth orientation of the horizontal line of sight of the GLORIA-B instrument (0° = north, $\pm 180^\circ$ = south). Bottom: Temporal evolution of ozone volume mixing ratio (ppmv) as seen from GLORIA-B over northern Scandinavia. The grey solid line shows the observer altitude (z_{obs}). Black solid lines mark sunset and sunrise terminators. Dynamical tropopause (2 and 4 potential vorticity units from ECMWF) is plotted as dashed magenta lines. Changes in the ozone VMR around 17:18 UTC, 22:05 UTC, 23:13 UTC, and 05:02 UTC correlate with azimuthal rotations which indicate a gradient in the horizontal ozone amount.



3.1 O₃

The monitoring of the expected recovery of the stratospheric ozone layer is still of scientific interest (Chipperfield and Bekki, 2024; WMO, 2022). Continuous ozone measurements are also important because this molecule is known to be an effective greenhouse gas in the upper troposphere and lowermost stratosphere (UTLS) region since it affects the radiation budget in this altitude range (Forster and Shine, 1997; Gettelman et al., 2011; Hansen et al., 1997; Riese et al., 2012; Xie et al., 2008). Figure 4 shows the ozone measurements performed with GLORIA-B between 17:11 UTC on 21 August 2021 and 05:39 UTC on the following day during float of the gondola. The vertical ozone distribution reflects the expected increase of ozone VMR with altitude reaching maximum values of about 6 ppmv near 35 km. Short-term variations in the ozone mixing ratio at 17:18 UTC, 22:05 UTC, 23:13 UTC, and 05:02 UTC are due to the changes in the azimuthal viewing direction and reflect the horizontal variation in the amount of ozone.

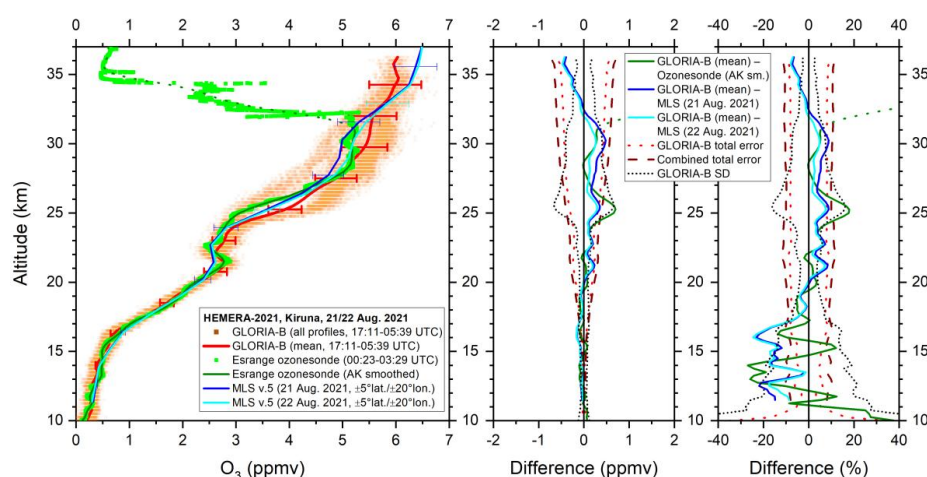


Figure 5. Left: All measured GLORIA-B O₃ data points (orange squares) and mean vertical profile (red solid line) compared to high-resolution ozonesonde data (green squares) and these data smoothed with the averaging kernel of GLORIA-B (solid dark green line, dotted outside altitude region of reliability). Satellite measurements (version 5) from the Microwave Limb Sounder (MLS), averaged in ± 5 deg. latitude and ± 20 deg. longitude bins around the GLORIA-B mean observation are shown, too. Middle and right: Absolute and relative differences together with combined error bars and GLORIA-B standard deviation (SD).



Figure 5 shows the mean ozone profile derived from the measured GLORIA-B spectra. As previously mentioned, the gondola did not travel far horizontally from the launch pad, maintaining a close horizontal distance of less than 100 km to the ozonesonde. This proximity, along with the circular rotating line of sight of GLORIA-B, allows for a direct comparison between both measurements. The ozonesonde data were smoothed using the averaging kernel of GLORIA-B. Above 30 km, the ozonesonde values gradually become unrealistic, a typical behavior of ozonesondes (Stauffer et al., 2022). VMR differences between the mean GLORIA-B ozone profile, the smoothed ozonesonde, and MLS remain within 10% for most altitudes. A slight positive bias in the GLORIA-B profile is observed above 22 km (compared to the ozonesonde) and between 20 and 32 km (compared to MLS). However, these biases are not significant when considering the combined error bars of each measurement.

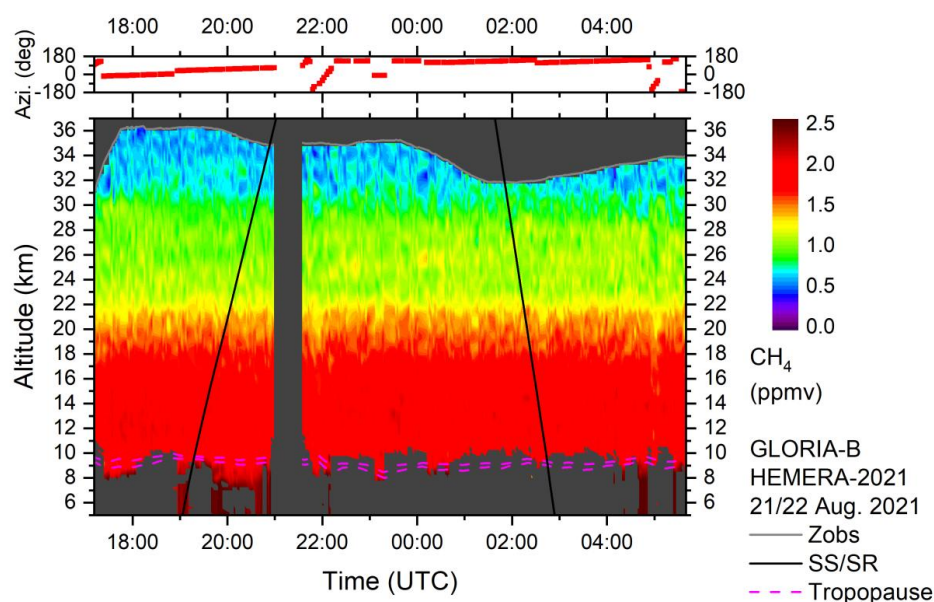


Figure 6. Same as Fig. 4 but for CH₄.

3.2 CH₄

Methane is an important greenhouse gas and further increasing concentrations are expected during this century (WMO, 2022). Furthermore, the oxidation of CH₄ in the stratosphere produces two H₂O molecules per CH₄ molecule, making CH₄ an important source of stratospheric H₂O (e.g. Brasseur and Solomon, 2005).



The CH₄ cross section as measured by GLORIA-B is given in Fig. 6. The CH₄ VMR exhibits the expected behavior with high tropospheric values close to 1.9 ppmv and decreasing amounts in the stratosphere. The overall structure appears a bit noisy. CH₄ variations due to changes in the azimuth are barely visible since the horizontal gradient of CH₄ concentration is small.

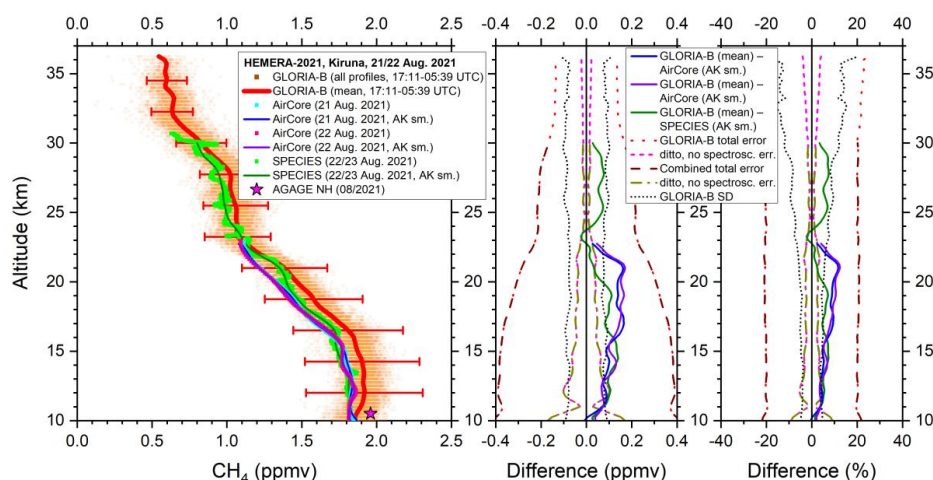


Figure 7. Left: All measured GLORIA-B CH₄ data points (orange squares) and mean vertical profile (red solid line) compared to in-situ AirCore (cyan and pink squares) and SPECIES data (green squares) and these data smoothed with the averaging kernel of GLORIA-B (solid blue, violet and dark green line). For comparison, a northern hemispheric lower tropospheric value from the AGAGE network is shown, too (magenta star). Middle and right: Absolute and relative differences together with combined error bars and GLORIA-B standard deviation (SD).

The comparison of the mean GLORIA-B CH₄ profile to in-situ measurements is displayed in Fig. 7. The overall shape of the vertical profiles of all observations is consistent. The difference between GLORIA-B and the in-situ measurements is within 10 %. However, there is a small positive bias visible in the GLORIA-B data that is not significant since the deviations are clearly within the combined error bars. Note that the CH₄ uncertainty in the GLORIA-B observation is dominated by the large spectroscopic error (mainly line strength and line width) given in the HITRAN database (Gordon et al., 2017). It seems that the spectroscopic error is overestimated, since the difference between GLORIA-B and the in-situ instruments is clearly smaller than the combined error itself. All measurements at 10 km are close to the northern hemispheric lower



tropospheric value from the Advanced Global Atmospheric Gases Experiment (AGAGE) network (Prinn et al., 2018).

3.3 CFC-11

The molecule CFC-11 (CCl_3F) is a rather long-lived chlorofluorocarbon that is removed from the atmosphere mainly by stratospheric photolysis (e.g. Ko and Dak Sze, 1982). Since it belongs to the main ozone depleting substances it is necessary to monitor the concentration of this trace gas to detect unexpected emissions (WMO, 2022).

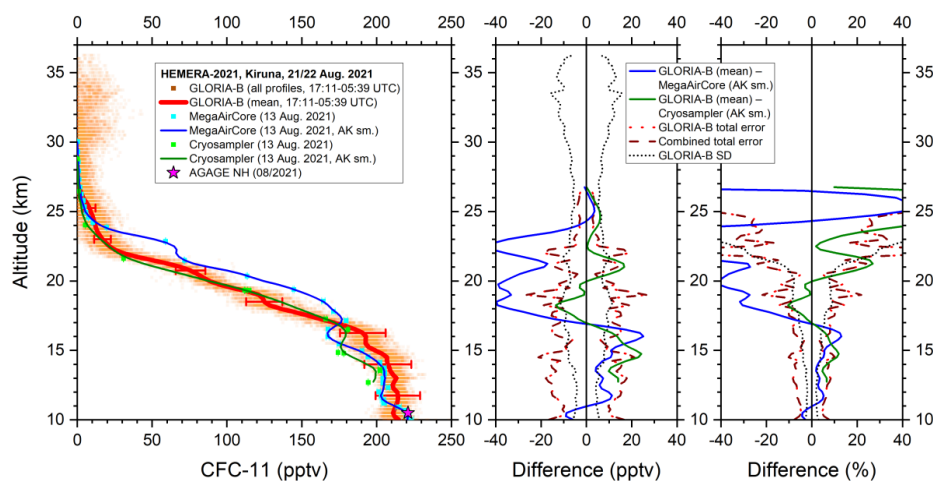


Figure 8. Left: All measured GLORIA-B CFC-11 data points (orange squares) and mean vertical profile (red solid line) compared to in-situ MegaAirCore (cyan squares) and cryosampler data (green squares) and these data smoothed with the averaging kernel of GLORIA-B (solid blue and dark green line). For comparison, a northern hemispheric lower tropospheric value from the AGAGE network is shown, too (magenta star). Middle and right: Absolute and relative differences together with combined error bars and GLORIA-B standard deviation (SD).

Figure 8 presents the mean retrieved GLORIA-B CFC-11 profile in comparison to the in-situ MegaAirCore and cryosampler observations. The rapid decrease in volume mixing ratio (VMR) with altitude due to photolysis is evident in all measured CFC-11 profiles, approaching values close to zero above approximately 26 km. Below 17 km, measured CFC-11 values of all three instruments lie close together (within 10 %) and inside the combined error limits. Above this altitude we see an agreement between GLORIA-B and the cryosampler data with respect to the



combined errors but a significant positive bias of the MegaAirCore observation. This bias is caused by a combination of internal MegaAirCore air mixing and averaging of air masses due to refilling into subsampling canisters (Laube et al., 2025). It is therefore especially pronounced for species with a large concentration gradient between 19 and 24 km. At an altitude of 10 km, all measurements align closely with the northern hemispheric lower tropospheric AGAGE value.

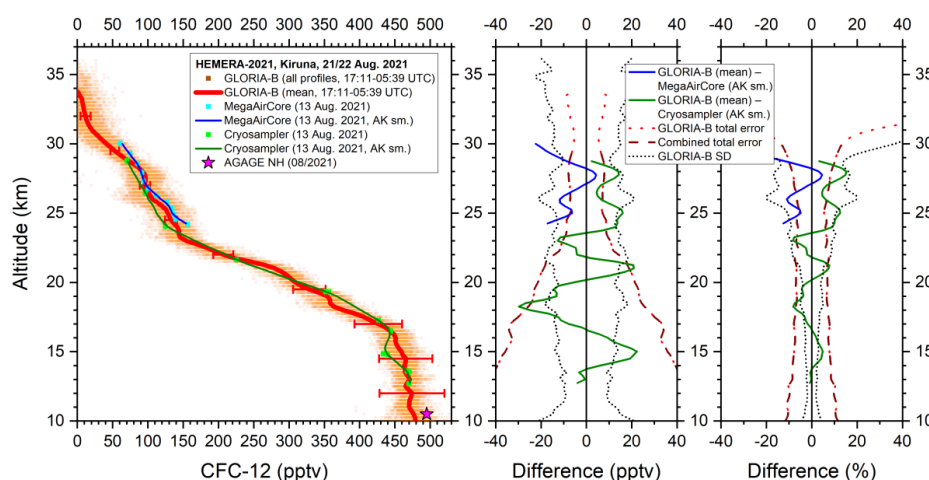


Figure 9. Same as Fig. 8 but for CFC-12.

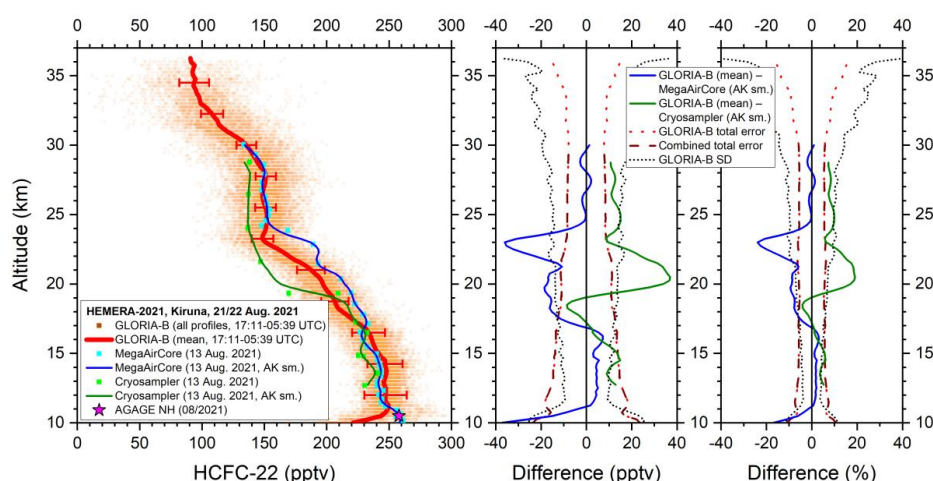
3.4 CFC-12

The chlorofluorocarbon CFC-12 (CCl_2F_2) is, like CFC-11, removed from the stratosphere mainly by photolysis (e.g. Ko and Dak Sze, 1982). Since the atmospheric lifetime of CFC-12 (around 102 years) is about twice as long as for CFC-11 (Hoffmann et al., 2014), the stratospheric decrease of the CFC-12 VMR with altitude is less pronounced than in the case of CFC-11. Figure 9 illustrates the comparison of vertical CFC-12 profiles observed by GLORIA-B and the in-situ instruments. Due to technical problems, the MegaAirCore device collected measurements only within a limited altitude range between 23 and 26 km which are in line with the GLORIA-B observations. The cryosampler data show agreement (within 10 %) with the mean GLORIA-B profile taking into account the combined error limits. At an altitude of 10 km, the GLORIA-B data points are close to the northern hemispheric lower tropospheric AGAGE value.



3.5 HCFC-22

505 The longer-lived hydrofluorocarbon HCFC-22 (CHClF_2) is often used as an alternative to the highly ozone-depleting substances CFC-11 and CFC-12 such that its tropospheric concentration increased until the early 2020s (e.g. Chirkov et al., 2016; WMO, 2022). The comparison between GLORIA-B and the in-situ observations is shown in Fig. 10. The difference between both in-situ instruments and the retrieved GLORIA-B data is less than 10 % below about 19 km. However, above this altitude, a clear negative bias is evident in the cryosampler VMR, which becomes significant between 19.5 and 23 km. In contrast, a positive bias is evident in the MegaAirCore data up to 24 km altitude. This positive bias is caused by the same problems as described in section 3.3.



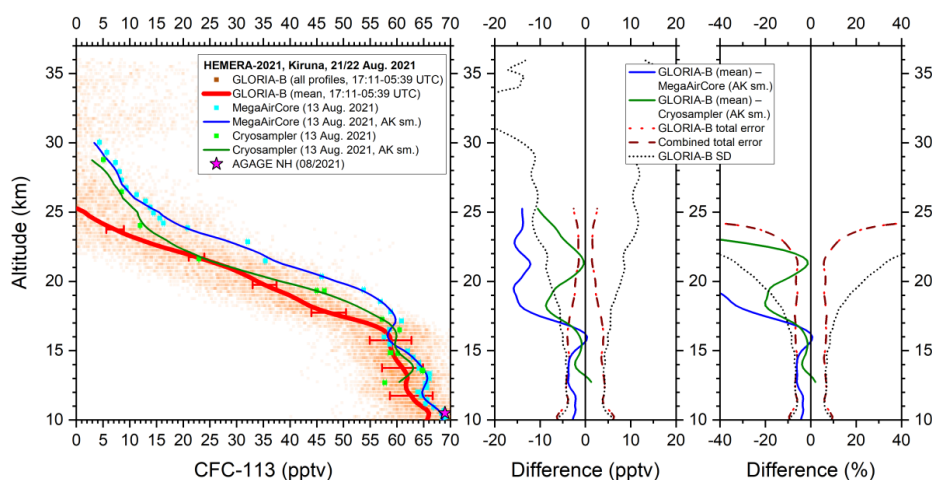
515 **Figure 10.** Same as Fig. 8 but for HCFC-22.

3.6 CFC-113

520 The chlorofluorocarbon CFC-113 ($\text{CCl}_2\text{FCClF}_2$) is also very unreactive in the troposphere and is photolyzed in the stratosphere (e.g. Sen et al., 1999; Dufour et al., 2005). Due to the significant spectral overlap from many interfering species, retrieving the VMR of CFC-113 is more challenging compared to the previously discussed target gases. Figure 11 illustrates the intercomparison of vertical profiles from GLORIA-B and the in-situ instruments. Up to 17 km altitude, the VMR differences are less than 10 % and fall within the combined error limits. However, above this altitude, MegaAirCore values are significantly larger than those retrieved



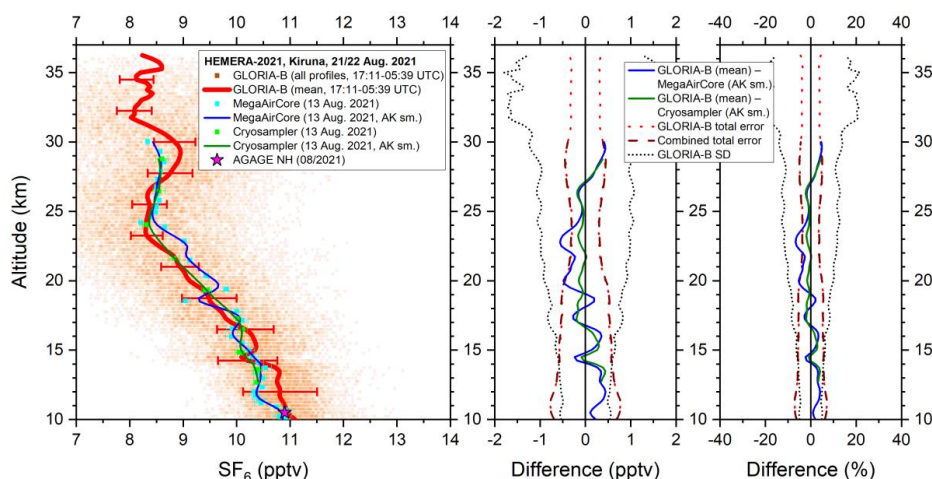
525 by GLORIA-B. In this upper altitude region, the cryosampler values are mostly somewhere between GLORIA-B and MegaAirCore. It should be noted that the standard deviation of the mean GLORIA-B profile is large, as the CFC-113 signal in the GLORIA-B spectra is weak at higher altitudes.



530 **Figure 11.** Same as Fig. 8 but for CFC-113.

3.7 SF₆

The long-lived anthropogenic tropospheric source gas sulfur hexafluoride (SF₆) is a trace gas with a high global warming potential (Hu et al., 2023). Since this molecule is chemically inert in the troposphere with a monotonic increase in atmospheric mixing ratio, it is suitable to study tropospheric and stratospheric transport processes based on the concept of age of air determination (e.g. Hall and Waugh, 1998; Engel et al., 2009; Stiller et al., 2012; Ray et al., 2014; Schuck et al., 2024; Saunders et al., 2025; Voet et al., 2025). The comparison of GLORIA-B retrieved SF₆ and the in-situ instruments is displayed in Fig. 12. The vertical shape of the observed profiles is largely as expected. Differences between all measurements are less than 10 % and stay within the combined error bars. The VMR peak seen in the GLORIA-B data around 29 km appears not to be significant since the standard deviation is large in this altitude region. At an altitude of 10 km, the GLORIA-B data points correspond to the northern hemispheric lower tropospheric AGAGE value.



545

Figure 12. Same as Fig. 8 but for SF₆.

4 GLORIA-B measurements and model comparison

In this section we demonstrate that GLORIA-B is able to measure temporal changes in the
550 mixing ratios of photochemically active nitrogen, chlorine, and bromine containing compounds
with high spatial and temporal resolution. Measurements of ClONO₂, N₂O₅, NO₂, and BrONO₂
are compared to the EMAC chemical model simulations.

4.1 ClONO₂

Chlorine nitrate (ClONO₂) is an important compound in stratospheric chemistry, particularly in
555 the context of ozone layer depletion (e.g. Brasseur and Solomon, 2005; von Clarmann and
Johansson, 2018). Since it serves as a key reservoir species for reactive chlorine in the
stratosphere, it temporarily stores chlorine atoms that could otherwise participate in ozone-
depleting reactions. The diurnal variation of ClONO₂ as observed by GLORIA-B is shown in
Fig. 13a. ClONO₂ undergoes minor diurnal variations at higher altitudes as a result of the
560 interaction between the ClONO₂ build-up via NO₂ and ClO and the ClONO₂ decay by
photolysis, resulting in somewhat higher values during night compared to day. Changes in the
VMR around 17:18 UTC, 22:05 UTC, 23:13 UTC, and 05:02 UTC correlate with azimuthal
rotations which indicate a gradient in the horizontal distribution of the species. The EMAC



simulation reproduces the observation (Fig. 13b) such that differences between GLORIA-B and
565 EMAC are not larger than ± 0.3 ppbv (Fig. 13c).

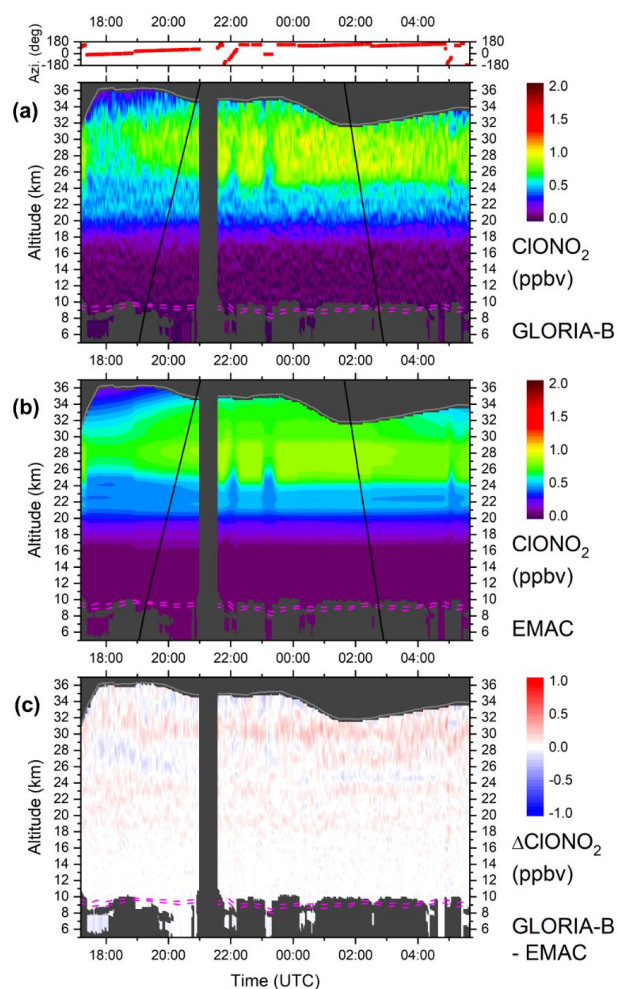


Figure 13. Temporal evolution of ClONO_2 as seen by GLORIA-B on 21/22 Aug. 2021 above northern Scandinavia (a). The variation of the azimuthal viewing direction is given on top (red dots). The observer altitude is shown as grey line. Black solid lines mark the sunset (SS) and sunrise (SR) terminators. The dynamical tropopause using 2 and 4 potential vorticity (PV) units is given as dashed magenta lines. The corresponding simulation by EMAC (b) and the difference (c) between the measurement and the chemical model are also shown.

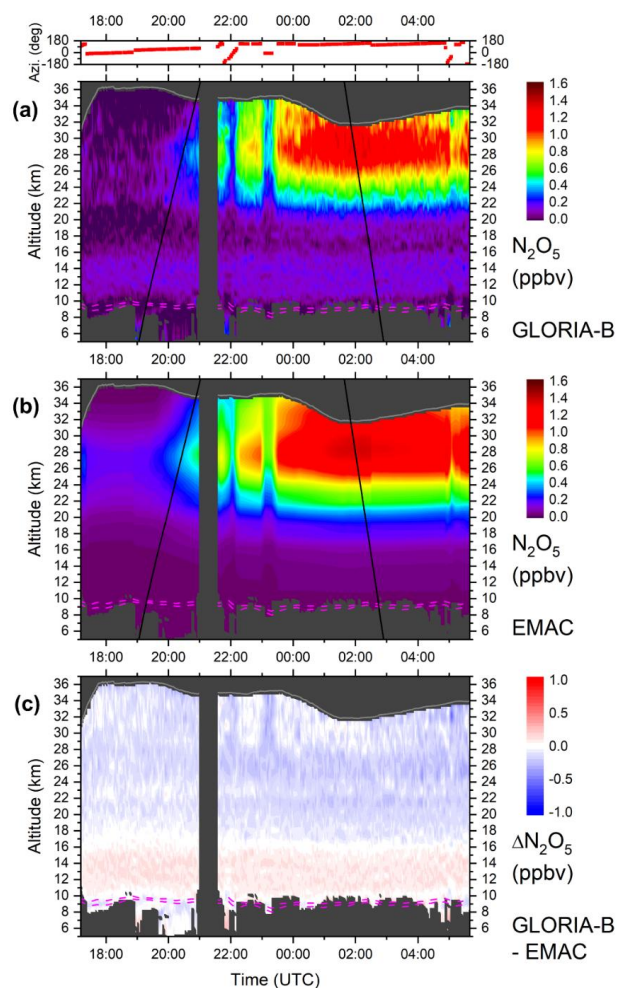


Figure 14. Same as Fig. 13 but for N_2O_5 . A pronounced diurnal variation of the amount of N_2O_5 is evident both in the measured GLORIA-B data and in the EMAC simulation. The strong correlation between azimuthal line of sight rotations and the changing N_2O_5 VMR due to different times of illumination is clearly visible.

4.2 N_2O_5

Dinitrogenpentoxide (N_2O_5) is, besides HNO_3 and ClONO_2 , a stratospheric reservoir species for NO_2 and NO (e.g. Brasseur and Solomon, 2005). It is produced during the night from reactions with O_3 , NO_2 and NO_3 . After sunrise, N_2O_5 gradually photolyzes back into NO_2 and



NO₃. Therefore, the amount of N₂O₅ shows a clear daily cycle in the stratosphere (Wiegele et al., 2009; Wetzel et al., 2012). A significant increase in the N₂O₅ VMR after local sunset is visible in the GLORIA-B measurements (see Fig. 14a). Azimuthal rotations of the line of sight around 22:05 UTC and 23:13 UTC correlate clearly with the observed amount of N₂O₅. The EMAC simulation reproduces the measured N₂O₅ VMR close to perfect (Fig. 14b). Consequently, deviations between measurement and model are small despite minor systematic differences (Fig. 14c).

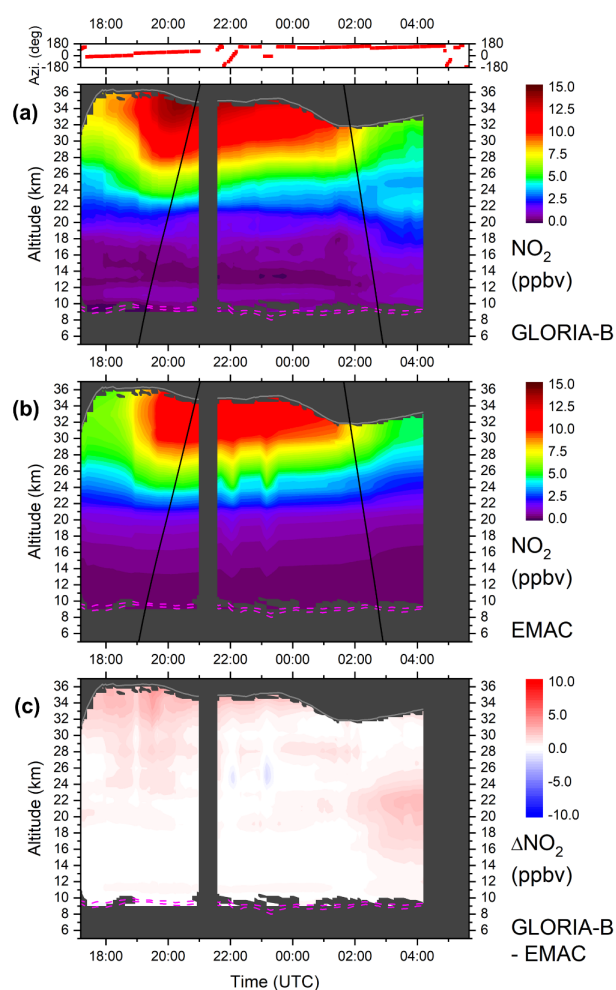


Figure 15. Same as Fig. 13 but for NO₂. The measured data (a) have been temporally smoothed with a 17-point adjacent averaging routine to attenuate noisy structures. Therefore, no changes in the measured VMR due to azimuthal line of sight rotations are visible.



4.3 NO₂

Nitrogen dioxide (NO₂) is not only involved in the temporal cycle of N₂O₅ but also shows a strong interaction with NO (e.g. Brasseur and Solomon, 2005). This results in a pronounced diurnal variation of the NO₂ VMR (Wiegele et al., 2009; Wetzel et al., 2012). This is also visible
585 in the GLORIA-B measurements (Fig. 15a) and the EMAC simulation (Fig. 15b). The start of the NO₂ VMR build-up at 34 km in the measurement is reproduced slightly delayed in the model. The small secondary maximum near 22 km (after sunrise) seen in the measurement is not visible in the model calculation. However, it should be emphasized that in this part of the measurement the retrieval noise error is enhanced such that this secondary maximum may also
590 be attributed to a retrieval artefact. In general, the NO₂ values from GLORIA-B are slightly higher in most regions as simulated with EMAC. All VMR differences between GLORIA-B and EMAC are displayed in Fig. 15c.

4.4 BrONO₂

Bromine nitrate (BrONO₂) is the most abundant inorganic bromine (Br_y) compound in the stratosphere (e.g. Brasseur and Solomon, 2005). It is produced from BrO and NO₂ during night
595 and it is destroyed during the day by photolysis and reaction with atomic oxygen which results in a pronounced diurnal cycle of BrONO₂. The first stratospheric measurements of BrONO₂ were reported by Höpfner et al. (2009) who derived vertical profiles of BrONO₂ from atmospheric infrared emission spectra recorded by MIPAS aboard Envisat. The first
600 observations of stratospheric diurnal variations of this substance were enabled by the analysis of spectra measured by MIPAS-B, the balloon version of MIPAS (Wetzel et al., 2017). The temporal evolution of BrONO₂ as seen by GLORIA-B is depicted in Fig. 16a. The nighttime increase of the amount of BrONO₂ and its daytime decay (towards BrO) are clearly visible. This variation is largely reproduced by EMAC (Fig. 16b). Differences between measurement
605 and model above about 16 km (Fig. 16c) are obvious in the absolute amount of available bromine in the model that is dependent on the surface emission scenario of short-lived bromine-containing organic compounds (Warwick et al., 2006). Differences below this altitude are difficult to interpret since GLORIA-B errors there exceed the 100 % limit.

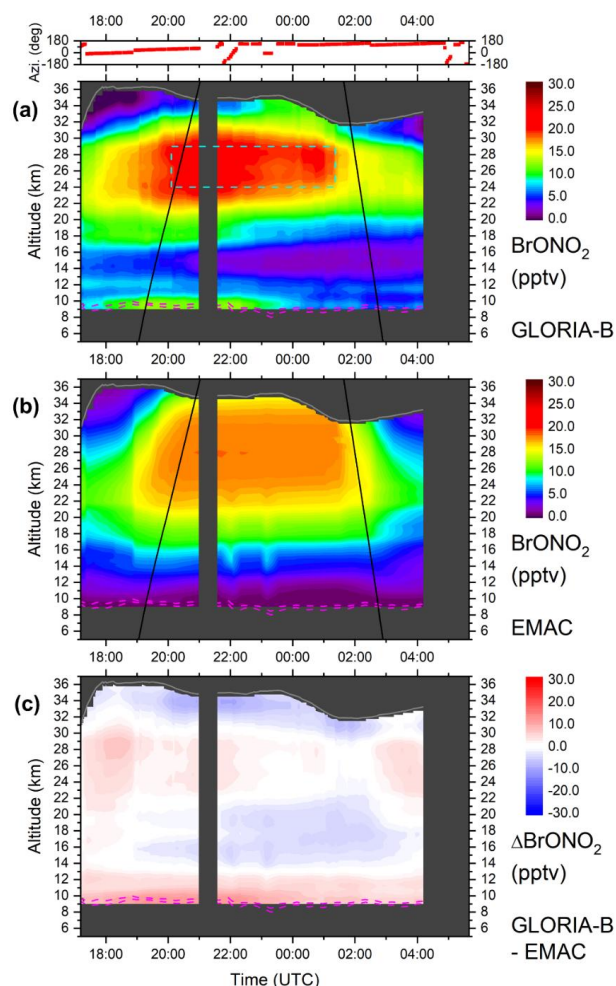


Figure 16. Same as Fig. 13 but for BrONO_2 . The measured data (a) have been temporally smoothed with a 39-point adjacent averaging routine to attenuate noisy structures. Therefore, no changes in the measured VMR due to azimuthal line of sight rotations are visible. The cyan coloured box marks the region for estimation of ‘measured’ Br_y .

As mentioned, BrONO_2 is the dominant species of stratospheric inorganic bromine. In our case, simulations of EMAC indicate that more than 90 % of nocturnal (20:06 UTC - 01:21 UTC) Br_y is in the form of BrONO_2 between 24 and 29 km. According to the method described by Wetzel et al. (2017), we can estimate ‘measured’ inorganic bromine [$\text{Br}_y(\text{meas})$] in this nocturnal altitude region by scaling measured [$\text{BrONO}_2(\text{meas})$] with the modeled ratio



620 $[\text{Br}_y(\text{mod})]/[\text{BrONO}_2(\text{mod})]$ from EMAC. We then calculate mean $[\text{Br}_y(\text{meas})]$ to 20.4 ± 2.5 pptv. The resulting error bar represents the $1-\sigma$ total error originating from $[\text{BrONO}_2(\text{meas})]$. Voss et al. (2024) recently determined stratospheric $[\text{Br}_y]$ of 17.5 ± 2.2 pptv from the ratio $[\text{BrO}]/[\text{Br}_y]$ using differential optical absorption spectroscopy (DOAS) BrO observations from a balloon flight carried out from Timmins (Ontario, Canada) on 23 August 2022, in combination
625 with chemical modelling. Although the Br_y values estimated from measurements of the two different instruments differ slightly, there is still an overlap between 17.9 and 19.7 pptv taking into account the error bars. A similar positive bias of Br_y deduced from mid-infrared BrONO_2 measurements in comparison to UV-visible Br_y inferred from BrO observations was also recognized and discussed by Wetzel et al. (2017). Since molecular spectroscopy plays an
630 important role in the accuracy of mid-infrared and UV-visible (DOAS) measurements, the observed Br_y difference might be explained at least partly by unquantified inaccuracies in the spectroscopic parameters (mainly line intensity).

5 Conclusions

635 Vertical profiles of a number of trace gases were retrieved from limb-imaging infrared emission spectra measured during the first flight of the GLORIA balloon instrument above northern Scandinavia. First, long-lived tropospheric source gases and stratospheric tracer species were compared to external measurements to assess the quality and the reliability of the GLORIA-B observations. Note that the accuracy of the measured gases depends mainly on the strength of
640 spectral signatures of the target gas and the separability of these signatures from those of interfering gases. For the gases O_3 , CH_4 , SF_6 , and CFC-12 we find an overall agreement within 10 % between GLORIA-B retrieved vertical profiles and the in-situ MegaAirCore and cryosampler instruments. No unexplained biases between GLORIA-B and the in-situ instruments are detectable, the differences stay within the combined error limits. Bearing in
645 mind this agreement, the large estimated spectroscopic CH_4 error (see Table 1) originating from the HITRAN database appears to be overestimated. For the molecules CFC-11, HCFC-22, and CFC-113 we find a level of agreement within 10 % to 20 % up to about 18 km, and differences are inside the combined error limits. Above this altitude region, deviations between GLORIA-B and also between the MegaAirCore and cryosampler instruments become larger and at least
650 partly significant with respect to the combined errors. However, a bias in the MegaAirCore altitude assignment due to diffusion cannot be ruled out. This could partly explain a bias



between MegaAirCore and GLORIA-B, and also a bias between MegaAirCore and the cryosampler instrument, especially in the vicinity of strong concentration gradients.

Second, photochemically active species (N_2O_5 , NO_2 , ClONO_2 , and BrONO_2) are compared to
655 simulations performed with the chemistry climate model EMAC. Calculations largely reproduce the temporal evolution of the observed gases while some differences appear in the absolute VMR values of the gases. While for N_2O_5 , NO_2 , and ClONO_2 the absolute amount differs only slightly, the difference for BrONO_2 is significantly larger. However, it should be
660 emphasized that the absolute VMR of BrONO_2 in the model simulation is strongly dependent on the assumed available inorganic bromine, which was chosen according to scenario 5 by Warwick et al. (2006).

In conclusion, the first flight of the GLORIA balloon instrument demonstrated its capability to retrieve accurate vertical profiles of trace gases, showing agreement with both external measurements (taking into account the combined error limits) and chemical model simulations.
665 Hence, GLORIA-B and, in general the limb-imaging technique, can contribute to a better understanding of the distribution and temporal evolution of key atmospheric species from the upper troposphere to the middle stratosphere, and supports climate relevant process understanding that allows continuous atmospheric model improvements.

670 *Data availability.* GLORIA-B measurements will be available on the KITopen repository. EMAC data are available upon request. MegaAirCore CH_4 and part of cryosampler data is on the Zenodo repository.

Author contributions. GW wrote most parts of the manuscript and performed the bulk of the
675 data analysis, with input from all co-authors. AK managed the instrumental part of the manuscript. SJ, AK, JU, MH, PP, and WW performed the GLORIA data processing and were involved in many helpful scientific discussions. TK, TN, HN, JS, HS, GS, and AS developed the necessary modifications to bring the GLORIA instrument to the balloon platform. FFV, EK, GM, HN, CP, TG, MRe, and TN, operated GLORIA during the balloon campaign in Kiruna.
680 CC and JL provided MegaAirCore data. AE, TS, and JL analyzed cryosampler measurements. VC and PJ provided SPECIES observations. OK performed the EMAC simulations. PB and MRi directed the research. All authors commented on and improved the manuscript.



Competing interests. One author is a member of the editorial board of AMT.

685

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