



Real-time plotting and evaluation of the data quality control from the CSIR- NGRI Magnetic observatories

Vengala. Pavan Kumar, Nelapatla. Phani Chandrasekhar and Potharaju. Sai Vijay Kumar
CSIR-National Geophysical Research Institute, Hyderabad, India
Corresponding author: phaninelapatla@gmail.com

Abstract

Earth's magnetic field, a dynamic shield influenced by internal and external forces, holds critical insights into space weather forecasting and the planet's core dynamics. The Choutuppal (CPL) and Hyderabad (HYB) magnetic observatories in India are pioneering this field by delivering high-resolution geomagnetic data to INTERMAGNET with unprecedented speed and precision. Utilizing a novel, low-cost protocol, CPL transmits 1s resolution data and HYB provides 1 min data, both achieving a latency of less than 300s making them among the first observatories worldwide to accomplish this feat. This rapid data transmission enhances global collaboration in space weather prediction, safeguarding critical infrastructure like satellites and power grids from solar storms.

To further elevate data utility, we developed a Python based software for real-time visualization and quality control at both observatories. This tool generates plots, performs initial quality checks, and computes first differences at 1s and 1min intervals, with a latency under 300s. By enabling daily evaluation of data quality, the software facilitates the identification of anomalies and noise, supporting the preparation of quasi-definitive data essential for geomagnetic research. Our Python server and web applications are designed with the future in mind, integrating artificial intelligence (AI) and machine learning (ML) capabilities. These advancements at CPL and HYB are set to transform the processing, forecasting, and visualization of geomagnetic data. By improving both the accuracy and accessibility of this data, we aim to revolutionize geomagnetic research, making it more precise, accessible, and actionable.

1. Introduction

Geomagnetic observatories are fixed locations on the Earth's surface that continuously monitor the geomagnetic field. The continuous time series data from these observatories reflect various physical processes associated with solar and Earth activities. The reliability and accuracy of the data are crucial for both scientific research and commercial applications [Matzka et al. 2010; Zhang et al. 2016]. Each observatory is equipped with at least two tri-axial fluxgates to record variations along with Overhauser magnetometers. Additionally, it has one fluxgate theodolite and Overhauser / Proton precession magnetometer for absolute measurements.

The Hyderabad Magnetic Observatory (HYB) of CSIR-NGRI has maintained 60 years of uninterrupted and stable recordings of magnetic variations. In 2009, with upgraded instruments, HYB became an INTERMAGNET observatory. Due to rapid urbanization and the introduction of the Hyderabad Metro Rail project nearby, it became essential to establish an alternate observatory to ensure the continuity of the geomagnetic data series. The campus of the former Choutuppal (CPL) Geo-electric Observatory provided an ideal location for recording magnetic measurements at one-second intervals. Preliminary 1 min observations began in 2012 and 1s data recording was initiated in 2015, and the ongoing data collection has



44 led to CPL being recognized as an INTERMAGNET observatory in 2019. The high-quality 1s
45 definitive data from CPL is currently being submitted to INTERMAGNET, making it one of
46 the first observatories in this region to achieve this status [Arora et al. 2016]. The HYB and
47 CPL observations have made significant contributions to global data, alongside other
48 observatories worldwide, for the main model of the Earth's magnetic field. These observations
49 have also supported various studies of low-latitude magnetic phenomena [Dwivedi and
50 Chandrasekhar, 2024] and regional induction anomalies [Edara and Arora, 2023].

51 For any observatory, the quality of the data is very important to achieve and maintain
52 INTERMAGNET status [Clarke et al. 2013]. Many scholars have developed various quality
53 control methods for geomagnetic data [Curto and Marsal 2007 and references therein]. Clarke
54 et al. [2013] developed an automated data processing software that integrates daily extrapolated
55 baseline values of H, D, and Z, derived from baseline functions, with H, D, and Z variometers
56 data. 1min data are delivered to the Edinburgh INTERMAGNET Geomagnetic Information
57 Node (GIN) in near real-time and on the following day. After implementing a few procedures,
58 the quality data are prepared and delivered to GIN by running the data processing software in
59 manual mode on the next working day. Chandrasekhar et al. [2017] discussed the challenges
60 involved in measuring 1s variations in the geomagnetic field to meet the standards set by
61 INTERMAGNET for quality and data transmission at observatories over extended periods.
62 They also provided a detailed account of the progressive steps that led to the successful
63 establishment of these measurements at the CPL observatory. Khumotov et al. [2017]
64 developed a new method for the noise identification in the identification in the data at
65 observatories of IKIR FEB RAS (Russia) and CSIR-NGRI (India). They also presented a
66 review of commonly used methods for noise identification in practical situations, highlighting
67 the potential for reducing the impact of noise on data through various examples. He et al. [2022]
68 proposed a method that combines genetic algorithms and linear regression to evaluate
69 geomagnetic data quality. Their approach considers factors such as observational data, attitude
70 angle, scale factor, long-term drift, and temperature. They highlight that agreement among
71 geomagnetic vector observations is crucial for assessing data quality and utilize Bland-Altman
72 plots, applying a 95% confidence interval to evaluate this agreement quantitatively and
73 qualitatively. Lingala et al. [2022] discussed the observed noticeable differences in the noise
74 levels present in vector and scalar variation data, due to the vehicular noise observed before
75 and during the COVID lockdown period and also discussed the details of increased data quality
76 in the absence of traffic-generated noise sources. da Silva et al. [2023] developed the Magnetic
77 Observatories and Stations Filtering Tool (MOSFiT), a Python package designed to visualize
78 and filter data from magnetic observatories and magnetometer stations. This tool can also be
79 utilized for quality control of geomagnetic observatory data, similar to the methods
80 implemented by the British Geological Survey as described by Macmillan and Olsen [2013].
81 Several studies have discussed the quality of geomagnetic observatory data and improved
82 protocols for addressing noise in the data [Zhang et al. 2024 and references therein].

83 Apart from the data quality, the other important aspect of any Geomagnetic observatory is
84 remote site data transfer, which is crucial for various applications, including environmental
85 monitoring, scientific research, and etc. All the INTERMAGNET Observatories data world-
86 wide is collected by the Geomagnetic Information Nodes (GINs) serve as collection points for
87 real-time data and are connected to the INTERMAGNET Observatories through satellite,
88 computer, and telephone networks. GINs operate in five different countries: the UK, USA,



89 Japan, Canada, and France. They utilize four satellites: GOES-E, GOES-W, METEOSAT, and
90 GMS to receive real-time data from INTERMAGNET Observatories worldwide
91 (<https://intermagnet.org/gins.html>)

92 Numerous observatories around the globe transfer 1-minute to 1-second data in real time to
93 various GINs using different technologies. These technologies include satellite
94 communication, ISDN telephone links, FTP, VPN router servers, in-house built NDL HSS,
95 MQTT, and various third-party software and tools [Torta et al. 2009; Clarke et al. 2013;
96 Chulliat and Chambodut, 2014; Thomson, 2014; Gvishiani et al. 2016; Reda and Neska, 2016;
97 Zhang et al. 2016]

98 Recently, Potharaju and Nelapthla [2023] addressed the challenges of data transmission for both
99 1s and 1 min intervals. They detailed the step-by-step development process, algorithm creation,
100 function libraries, and the implementation of real-time data transmission from a remote
101 observatory. Using the Python programming language, they developed an algorithm to
102 automate the transmission of high-resolution real-time magnetic data from CPL and HYB to
103 Edinburgh GIN, all while relying on minimal internet connectivity. The automation system
104 securely transfers data in an encrypted manner using SSH keys, while also saving the same
105 dataset on a local server at CSIR-NGRI. Data from both observatories is sent to GIN in real-
106 time within a timeframe of less than 300 seconds. After successfully transmitting 1-minute
107 geomagnetic data in real-time from the CPL and HYB observatories to Edinburgh GIN, the
108 transmission of CPL's 1-second real-time data have also commenced. This achievement marks
109 CPL as one of the first Indian observatories to send 1-second real-time data to GIN.

110 In this paper, we outline the processes involved in upgrading the Python tool package that
111 facilitates real-time data quality control checks at the CSIR-NGRI magnetic observatories,
112 HYB and CPL, for both 1-second and 1-minute data intervals. Additionally, we discuss the
113 various options for installing and implementing our package, ensuring it integrates smoothly
114 with the available resources for real-time data transmission and quality checks.

115 **2. Real-time data transfer to INTERMAGNET GIN**

116 The real-time geomagnetic data, both vector and scalar, is initially collected by the MAGREC-
117 4B logger. This data is then transmitted to a local machine running the CentOS operating
118 system, which is deployed at the observatory (CPL/HYB). A secure communication is
119 established via SSH (Secure Shell) using key-based authentication. Initially, an SSH key pair
120 is generated, and to enhance security, this key is changed every two weeks to prevent
121 unauthorized access.

122 Once the data is received on the CentOS machine, it is processed and prepared for transfer to
123 the centralized server located at the NGRI-HYB Observatory. The secure transfer to the
124 observatory server is conducted using the same SSH protocol, ensuring a robust and encrypted
125 handshake for optimal data integrity and confidentiality. Upon arrival at the NGRI-HYB
126 Observatory server, the collected data is methodically organized based on its temporal
127 granularity. The data is categorized and stored in specific directories, namely "Minute Data"
128 and "Second Data," facilitating organized data management and easy access for analysis.

129 The segregated data from the HYB server is then prepared for transmission to the
130 INTERMAGNET Geomagnetic Information Node (GIN) located in Edinburgh, UK. This
131 transmission process is automated using Python scripts and daemon processes that run in the
132 background. These scripts are designed to execute every 300 seconds (5 minutes), ensuring



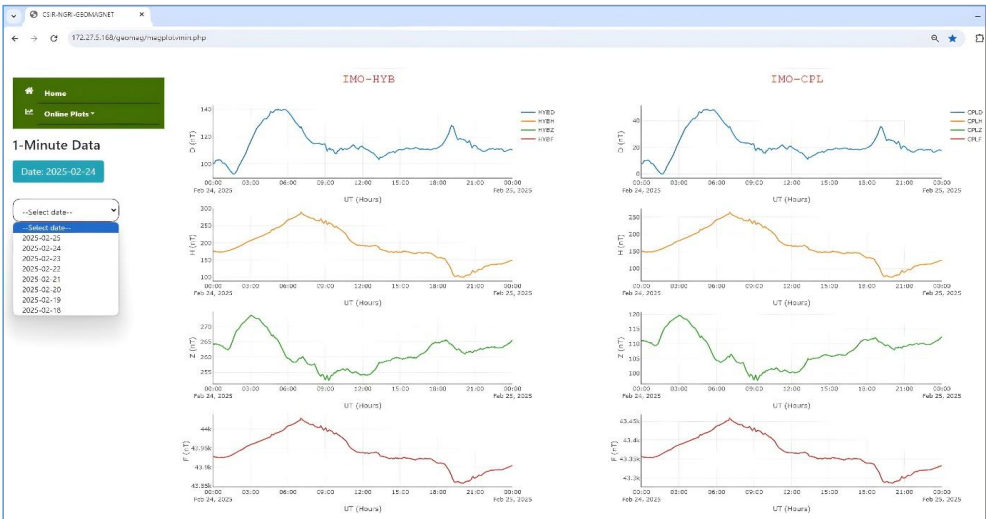
133 timely and regular data updates without duplication. The Python code handles data packaging,
134 error checking, and retransmission logic to ensure reliable data delivery.

135 Throughout the entire process, various security measures are implemented, including regular
136 updates to SSH keys and continuous monitoring of data transfer processes. Logs are maintained
137 to track data transfer activities and to quickly identify and rectify any anomalies or issues. By
138 employing this comprehensive workflow, the system guarantees secure, efficient, and reliable
139 data transfer from the MAGREC-4B logger to the INTERMAGNET GIN, thereby facilitating
140 continuous monitoring and analysis of geomagnetic data [Potharaju and Nelapla, 2023].

141 3. Customization of the PHP server for real-time data visualization

142 PHP (Hypertext Preprocessor) (<https://www.php.net/>) is a widely-used server-side scripting
143 language primarily designed for web development. It is embedded within HTML and executed
144 on the server, generating dynamic content that can be displayed on web browsers. PHP offers
145 flexibility, simplicity, and compatibility with various databases, making it a popular choice for
146 developing interactive and data-driven web applications. Real-time data visualization is
147 essential in various domains, including geomagnetic observatories, seismic monitoring and IoT
148 applications. PHP, in combination with Plotly (a JavaScript-based visualization library)
149 (<https://plotly.com/javascript/>), facilitates the rendering of real-time plots by handling server-
150 side data processing and sending the results to the client.

151 Since data from both observatories is available in real-time, we have developed a PHP server
152 that simultaneously plots the data from both locations. The screen refreshes every 300s to
153 display the updated trends for each component and continues till the end of the day. This server
154 is designed to store initially a weekday's data at one-minute sampling rate, as illustrated in
155 Figure 1.



156
157 Figure 1. Real-time plotting of Vector and Scalar data at 1 min sampling interval from the HYB
158 (left) and CPL (right) observatories of CSIR-NGRI

159



160 **4. Upgrading the PHP server to a Python server**

161 Real-time data visualization is crucial in various fields, including geomagnetism, seismology,
162 seismic, and many other Geophysical applications, where continuous monitoring and dynamic
163 plotting are essential. Traditionally, PHP has been a popular choice for web-based data
164 visualization due to its simplicity and widespread usage. However, with the increasing demand
165 for scalability, performance, and flexibility, Python Django (<https://www.djangoproject.com/>)
166 has emerged as a more powerful alternative.

167 Upgrading from PHP to Django for real-time plotting provides significant advantages in terms
168 of performance, scalability, and maintainability. Django's powerful framework, combined with
169 Python's rich ecosystem, allows efficient handling of large datasets, real-time updates, and
170 seamless integration with machine learning models. Although the migration process involves
171 challenges such as database compatibility and code refactoring, the long-term benefits in
172 flexibility, performance, and extensibility make Django a superior choice for real-time plotting
173 applications.

174 Django is a high-level Python web framework that promotes rapid development and clean
175 design. Python's libraries (e.g., Pandas, NumPy) enable complex data analysis and enhances
176 the efficient backend processing. Django channels and WebSockets provide low-latency, real-
177 time data streaming. It's ORM (Object-Relational Mapping) simplifies database interactions.
178 Django provides a modular architecture, making the application easier to scale and maintain.

179 Django, combined with **Bokeh** (<https://bokeh.org/>), a powerful visualization library, offers an
180 efficient solution for rendering real-time plots. Bokeh's interactive plotting capabilities, when
181 integrated with Django's backend, enable dynamic and responsive data visualizations. Bokeh
182 is a Python library for creating interactive and real-time visualizations, it generates JavaScript-
183 powered visualizations directly in the browser. It supports streaming data sources for dynamic,
184 live-updating plots. It provides interactive tools like panning, zooming, and hovering
185 functionalities for better data exploration.

186 With the increasing availability of geomagnetic data from observatories and satellites,
187 Artificial Intelligence (AI) and Machine Learning (ML) techniques are transforming the field
188 by enabling: a) Real-time monitoring and forecasting of geomagnetic events, b) Anomaly
189 detection for space weather and magnetic storms, c) Data-driven insights for understanding
190 magnetic field variations, d) Predictive models for geomagnetic hazards.

191
192 To achieve above mentioned applications, Python offers a robust ecosystem of AI/ML libraries
193 that can be directly integrated into Django-based geomagnetic applications. Django serves as
194 the API layer or backend framework for delivering AI/ML model predictions as web services.
195 Django + Python + AI/ML provides a future-proof, scalable, and efficient framework for
196 geomagnetic data processing, visualization, and prediction.

197

198 **5. First difference tool for real-time data quality checks**

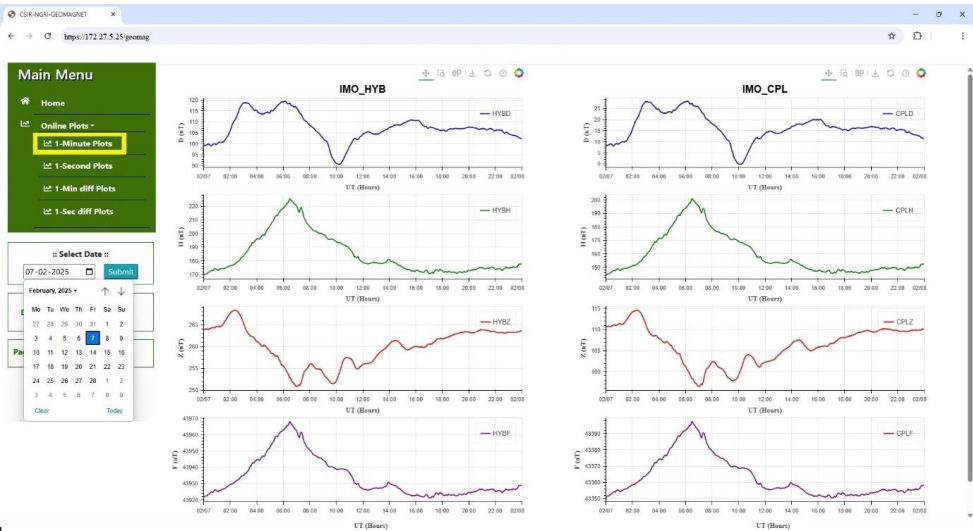
199 The "first difference" (FD) is a key analytical tool used in time series analysis. It refers to the
200 difference in values between consecutive observations in a dataset. This method transforms a
201 time series dataset to make it stationary, which helps in identifying patterns, trends, and other
202 dynamic aspects over time more easily of a signal. It is particularly useful for analysing



203 geomagnetic time series data and understanding its evolution. If the difference between two
204 consecutive time periods of a signal is abnormal, it may indicate the presence of noise in the
205 data, often caused by anthropogenic / environmental factors.

206 We have developed a real-time FD tool in Python that can calculate data for each component
207 of both observatories at 1s and 1 min intervals, allowing for a quick and hassle-free assessment
208 of data quality. This computation refreshes every 300s and displays the differences. Here is an
209 example that illustrates the 1-minute plot (Figure 2) and the 1-second plot (Figure 3) for
210 February 7, 2025, for the HYB (left panel) and CPL (right panel) observatories across each
211 component. We have upgraded the server to include several months of data, enabling users to
212 access the desired day instantly as needed. One example is illustrated in Figures 2 and 3.

213



214

215

216 Figure 2: Real-time plotting of Vector and Scalar data at 1 min sampling interval from the HYB
217 (left) and CPL (right) observatories of CSIR-NGRI, with the updated services in the server.

218

219

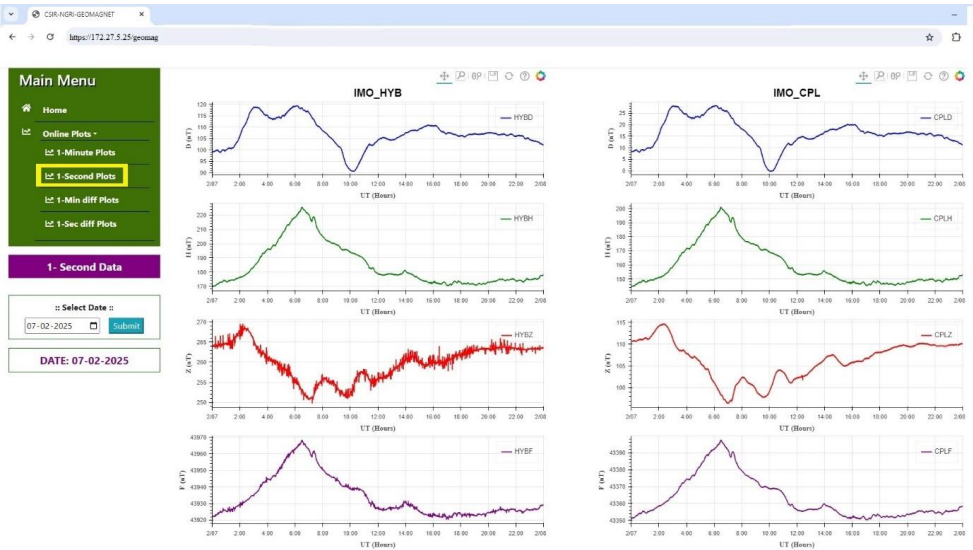


Figure 3: Real-time plotting of Vector and Scalar data at 1s sampling interval from the HYB (left) and CPL (right) observatories of CSIR-NGRI.

Upon reviewing Figures 2 and 3, we observe that the noon hours of the day began with a sudden storm commencement (SSC). This term refers to an abrupt increase or decrease in the northward component of the geomagnetic field and indicates the onset of a geomagnetic storm. The SSC event is noted around 07:14 UT at both observatories. In comparison to 1s, 1 min data exhibits more noise in the Z and F components at HYB than at CPL (Figures 2 and 3).

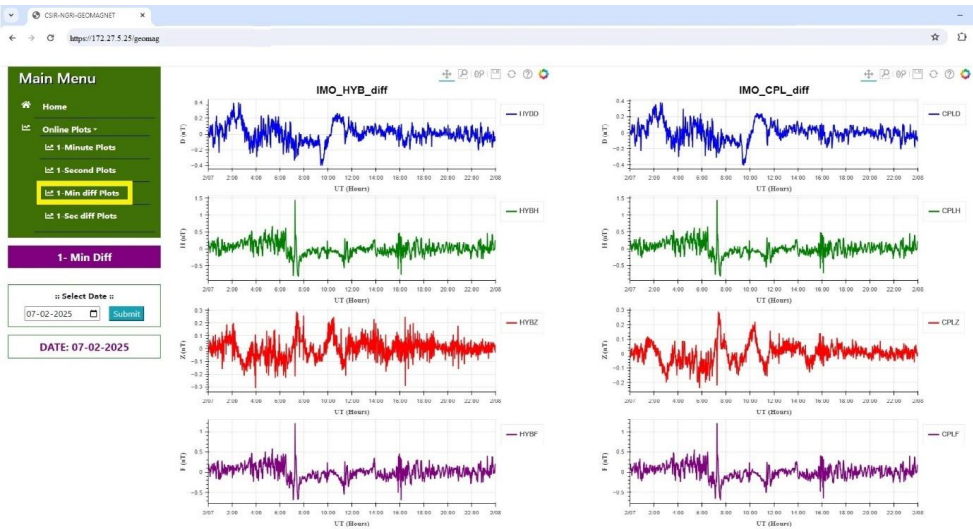


Figure 4: Real-time plotting of the first differences of vector and scalar data, sampled every minute from the HYB (left) and CPL (right) observatories

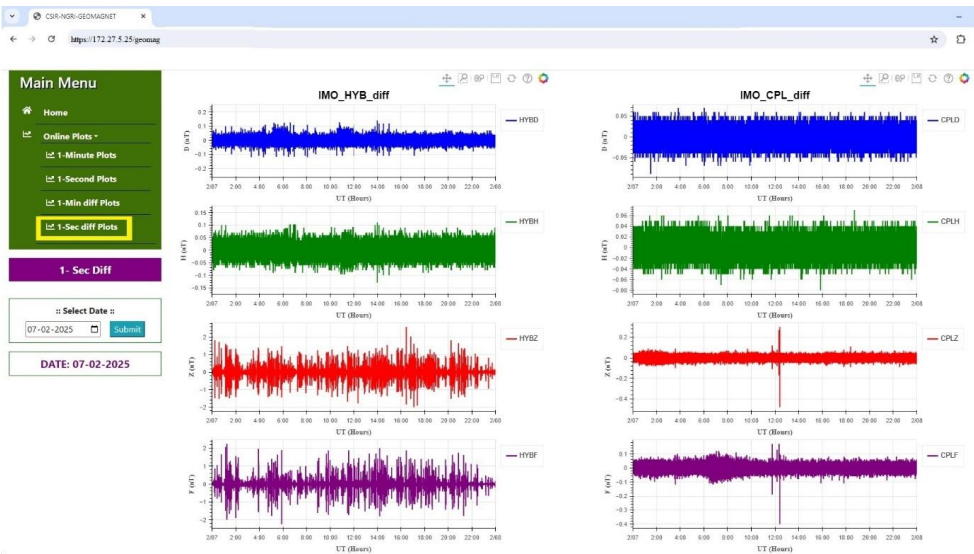


Figure 5: Real-time plotting of the first differences of vector and scalar data, sampled every second from the HYB (left) and CPL (right) observatories

Figures 4 and 5 present the calculated FD of each component at both the HYB and CPL observatories, measured in 1 min and 1s intervals. From these figures, it is apparent that small amplitude spikes were observed at the CPL observatory, particularly in the Z and F components around (12:21:07UT, and 12:24:57UT) during that day. The amplitude of these spikes at CPL are about: 0.4 nT and 1.5 nT. Further, the amplitude of the FD's of all the components are in the range of ± 0.5 nT in D, ± 1.5 nT in H, ± 0.3 in Z and ± 1.5 nT in F for 1 min data and ± 0.1 nT in D, ± 0.1 nT in H, ± 0.5 nT in Z at CPL, ± 2 nT in Z at HYB, ± 0.5 nT in F at CPL and ± 2 in F at HYB for 1s data. The next step is treating the spikes by evaluating the source behind the signal. After checking the logs, the spikes recorded at CPL are result from human intervention related to data collection from the flashcard of the spare fluxgate magnetometer deployed in the sensor hut. Hence the observed spikes at the components (D, H, Z and F are removed from the data set. In contrast, a greater number of spikes were recorded in the Z and F components at the HYB observatory on 7th February 2025. It is important to note that the spikes recorded in the Z and F components at HYB are attributed to vehicular traffic and metro rail operations, as discussed by Lingala et al. [2022]. These activities occur regardless of the geomagnetic conditions. Therefore, this data requires a treatment of noise removal before it can be submitted as quasi-definitive data. After treating the noise in the data at both the observatories, the final data for HYB and CPL for 1 min are shown in Figure 6.

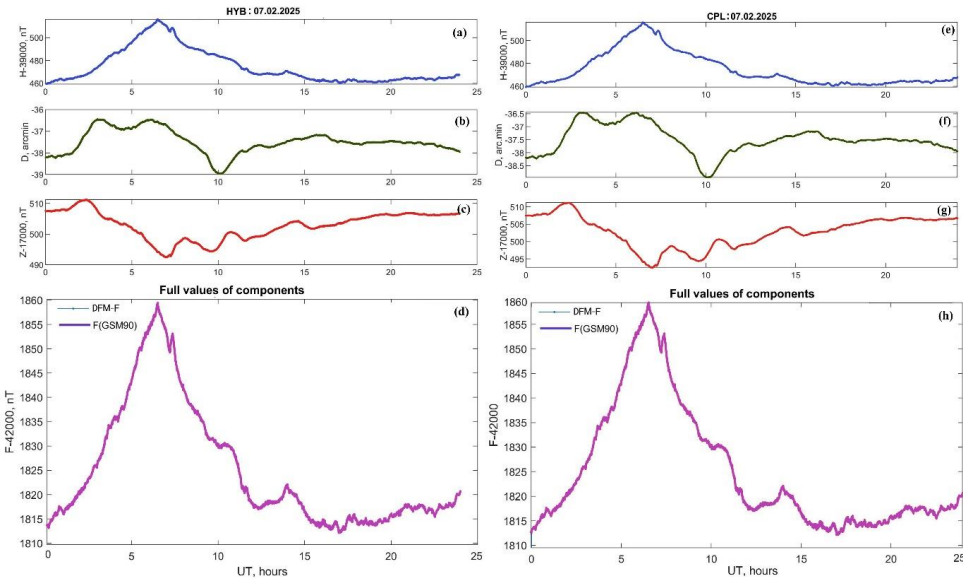


Figure 6: 1 min vector and scalar data from the HYB (left) and CPL (right) observatories after removing spikes identified through real-time first differences.

The plotting page also offers the following tools:

- (a) **PAN:** This tool allows you to move the view of an image or document around on the screen while maintaining the same zoom level. Essentially, it enables you to "slide" the image to view different parts without zooming in or out.
- (b) **Box Zoom:** This zoom method lets you click and drag to create a rectangular area on the screen, specifying which part of the content you want to zoom in on. This tool is particularly useful for identifying noise in the data.
- (c) **Wheel Zoom:** This tool lets you zoom in and out using the mouse scroll wheel.
- (d) **Reset:** This tool restores the view to the normal plotting window, removing any Box Zoom or Wheel Zoom adjustments.
- (e) **Save:** This function allows you to download the current plot as a ".png" image format.

6. The essence of extending the Python based facility to Remote Sites

CSIR-NGRI currently operates remote sites at various locations: Himalayas, Andaman & Nicobar Islands, and Ladakh. Extending the existing data transfer facility to these remote locations offers several key benefits:

- **Consistency:** Ensures uniform data collection and transmission across multiple sites, resulting in consistent and reliable datasets.



- 285 • **Scalability:** Facilitates easy expansion of the network to include more remote locations
286 without incurring significant additional costs.
- 287 • **Resource Optimization:** Maximizes the use of existing infrastructure and resources,
288 reducing the necessity for redundant systems.
- 289 • **Enhanced Data Insights:** Provides a broader range of data, leading to more
290 comprehensive analyses and improved decision-making.
- 291 • **Cost Efficiency:** By utilizing the same setup, organizations can save on costs associated
292 with deploying and maintaining separate systems for each site.

293 All remote site recording systems are equipped with maintenance-free batteries that are
294 regularly charged by solar panels. We have identified several potential systems for installing
295 the developed Python-based package, which is designed to implement low-cost and effective
296 data transfer techniques and tools, along with real-time quality control checks in the next few
297 months.

298 **A. Techniques:**

299 **1. LoRaWAN:** A low-power, long-range wireless communication technology that is ideal for
300 remote areas with minimal infrastructure. It is cost-effective and suitable for IoT applications.

301 **2. Satellite Communication:** Using Low Earth Orbit (LEO) satellites can provide reliable
302 internet connectivity in remote locations. Companies like Soracom and Sateliot offer affordable
303 satellite IoT solutions.

304 **3. Store-and-Forward Systems:** Data loggers can store data locally and transfer it when a
305 connection is available. This method is beneficial in areas with intermittent connectivity.

306 **4. High-Frequency Radio Communication:** HF radio can be employed for long-distance
307 communication in remote areas, requiring a clear line of sight, and can be a cost-effective
308 solution.

309 **B. Tools:**

310 **1. Airbyte:** This tool supports incremental data synchronization, making it efficient and
311 resource-effective. It primarily runs on Linux-based operating systems, specifically Ubuntu
312 and Debian. Although it can also be deployed using Docker on Windows and macOS, Linux is
313 the preferred operating system.

314 **2. Raspberry Pi with SIM Card:** This versatile and cost-effective option is equipped with a
315 SIM card to collect and transmit data. It typically runs on Raspberry Pi OS (formerly known
316 as Raspbian), which is based on Debian Linux, but it also supports Ubuntu, Alpine Linux, Arch
317 Linux, and other Linux distributions.

318 **3. Omega2 LTE:** This single-board computer (SBC) provides high-speed cellular connectivity
319 with LTE Cat 4 support, making it ideal for remote data transfer. It operates on OpenWRT, a
320 lightweight Linux-based operating system optimized for networking and IoT applications.

321 **4. Libre Computer Board Le Potato:** While primarily focused on multimedia applications, this
322 board can be adapted for data transfer using SIM card technology. It mainly runs on Ubuntu,
323 Debian, and Armbian (a lightweight Linux distribution designed for ARM-based SBCs).



324 **5. Orange Pi 5:** This board supports multiple Linux-based operating systems, including
325 Ubuntu, Debian, Armbian, and Android. It offers better performance than the Raspberry Pi at
326 a similar price point and is compatible with SIM card technology for data transmission.

327 By selecting any of the above approaches will enhance the capabilities of CSIR-NGRI's remote
328 operations, ensuring effective data management and communication across challenging
329 environments

330 **7. Summary**

331 Raw magnetic observatory data often contains noise and artifacts that need to be removed or
332 corrected. Pre-processing steps include the removal of spikes, correction for temperature
333 effects, time synchronization, and baseline adjustments. Data quality control is essential to
334 ensure the reliability of the information, involving visual inspections, statistical analyses, and
335 data flagging.

336 Many researchers prepare quasi-definitive data on a monthly, weekly, or daily basis by
337 incorporating these pre-processing steps and submit their findings to INTERMAGNET GIN.
338 For instance, the IGP quasi-definitive data method is a monthly process focusing on obtaining
339 the most accurate results for the recent past [Peltier and Chulliat, 2010]. The BGS method,
340 while similar, also aims to produce next-day quasi-definitive data using predicted baseline
341 values [Clarke et al. 2013]. Both methods are valid for meeting the quasi-definitive data
342 definitions set by INTERMAGNET and offer distinct strengths to benefit various data users.
343 da Silva et al. [2023] introduced the Python package MOSFiT, designed to work with 1-minute
344 INTERMAGNET definitive and quasi-definitive data, but it can also be applied to any
345 geomagnetic observatory or magnetometer data. The CSIR-NGRI geomagnetic observatories
346 at HYB and CPL participate in and contribute to INTERMAGNET and submitting quasi and
347 definitive data. Although both observatories are equipped with the same magnetometers, the
348 CPL observatory operates in a noise-free environment, while HYB does not. CPL is India's
349 first observatory to provide 1-second data, whereas HYB offers 1-minute real-time data to GIN.

350 Before processing data to produce quasi-definitive results, it is vital to validate its quality. To
351 address this need, we have developed a Python-based plotting service tool as the first step in
352 our data quality control process. This tool not only monitors real-time trends and continuity in
353 observational data but also includes a dedicated review process to rigorously assess data quality
354 regularly. This ensures the accurate preparation of quasi-definitive data for the observatories.
355 Additionally, the tool has indicators that flag data when the FD values exceed specified
356 thresholds, ensuring accuracy and completeness.

357 Our Python-based tool can be installed on various client-side devices, including LoRaWAN,
358 data loggers, Airbyte, Raspberry Pi, Omega2 LTE, Libre Computer Board Le Potato, and
359 Orange Pi 5. These devices are suitable for deployment in remote locations with limited power
360 availability. On the server side, the system can be configured to connect to a workstation or
361 server to receive data in real-time. The establishment of this real-time quality control system
362 significantly enhances the data quality from both permanent and remote observatory sites,
363 providing reliable support for related scientific research.

364 Further, our Python-based server is designed to provide a robust ecosystem of AI/ML libraries
365 that can be seamlessly integrated into Django-based geomagnetic applications. Django serves
366 as the API layer and backend framework for delivering AI/ML model predictions as web



367 services. The combination of Django, Python, and AI/ML creates a future-proof, scalable, and
368 efficient framework for processing, visualizing, and predicting geomagnetic data.

369 **8. Conclusions**

370 High-quality magnetic observatory data is vital for understanding the Earth's magnetic field.
371 Observatories in exceptional locations provide valuable long-term data, but assessing data
372 quality requires expertise and can be time-consuming. Producing reliable geomagnetic data is
373 challenging, especially for institutions with limited staff. Rising operational costs can make it
374 difficult to secure necessary funding. Despite these hurdles, many scientific publications rely
375 heavily on high-quality absolute magnetic observatory data.

376
377 We have developed a Python-based tool to assist observatory staff in identifying noise in data,
378 which will help reduce their workload. In the future, we plan to create a scalable framework
379 for processing, visualizing, and predicting geomagnetic data using Django, Python, and AI/ML
380 technologies. Additionally, we have established a cost-effective data transfer system that
381 enables reliable data collection and analysis from remote locations without imposing
382 significant financial burdens, thereby benefiting organizations with limited budgets.

383

384

385 **Acknowledgements**

386

387 The authors thank the Director of CSIR-National Geophysical Research Institute (Hyderabad)
388 for the support and permission to publish this work (NGRI/Lib/2025/Pub-45). The authors wish
389 to thank Dr. Kusumita Arora for her constant encouragement and permission to carry out this
390 research work.

391

392 **Author contributions**

393

394 VPK (V. Pavan Kumar), NPCS (N. Phani Chandrasekhar) and PSVK. (P. Sai Vijay Kumar)
395 wrote the manuscript and reviewed the article, NPCS prepared figures.

396

397 **Competing interests**

398

399 The authors declare no competing interests.

400

401

402 **References**

403 Arora, K., Chandrashakhar Rao, K., Manjula, L., Suraj Kumar and Nagarajan, N.: The new
404 magnetic observatory at Choutuppal, Telangana, India, J. Ind. Geophys. Union (January 2016),
405 Special Volume-2/ 2016 pp: 67-75, 2016.

406 Curto, J. J. and Marsal, S.: Quality control of Ebro magnetic observatory using momentary
407 values, Earth Planets Space, 59, 655–1187–1196, 2007.

408 Clarke, E., Baillie, O., Sarah J. Reay, and Chris W. Turbitt.: A method for the near real-time
409 production of quasi-definitive magnetic observatory data, Earth Planets Space, 65, 1363–1374,
410 2013.



- 411 Chulliat, A. & Chambodut, A.: Bureau Central de Magnétisme Terrestre Strategic Plan 2014–
412 2018 1–23 (Bureau Central de Magnétisme Terrestre), 2014.
- 413 Chandrasekhar, N.P, P.Sai Vijay Kumar, Kusumita Arora, KCS. Rao, Leonid Rakhlin, Sergey
414 Tymoshyn, Laszlo Merenyi, Ansuha, CH, Jayashree Bulusu, Sergey Khumotov.: One second
415 scalar and vector measurements at the low-latitude Choutuupal Observatory (India),
416 Copernicus publications, 6, 547-560, <https://doi.org/10.5194/gi-6-547-2017>, 2017.
- 417 da Silva, M. V., Pinheiro, K. J., Ohlert, A., and Matzka, J.: Analysis of geomagnetic observatory
418 data and detection of geomagnetic jerks with the MOSFiT software package, Geosci. Instrum.
419 Method. Data Syst., 12, 271–283, <https://doi.org/10.5194/gi-12-271-2023>, 2023.
- 420 Dwivedi, D and Chandrasekhar, N. P.: Geomagnetic field variations due to solar tides at the
421 Indian Observatories, Earth, Planets and Space, 76:61, [https://doi.org/10.1186/s40623-024-](https://doi.org/10.1186/s40623-024-01996-8)
422 [01996-8](https://doi.org/10.1186/s40623-024-01996-8), 2024.
- 423 Edara,A., and Arora, K.: Annual/Seasonal Variation in Induction Vectors at Different
424 Geological Locations in an Indian Sector, Pure and Applied Geophysics, Volume 180, pages
425 3527–3543, 2023.
- 426 Gvishiani, A., Soloviev, A., Krasnoperov, R. & Lukianova, R.: Automated hardware and
427 software system for monitoring the Earth’s magnetic environment. Data Sci. J. 15(18), 1–24.
428 <https://doi.org/10.5334/dsj-2016-018>, 2016.
- 429 He, Z., Hu, X., Teng, Y. et al. Data agreement analysis and correction of comparative
430 geomagnetic vector observations. Earth Planets Space 74, 29,
431 <https://doi.org/10.1186/s40623-022-01583-9>, 2022.
- 432 Khomutov, S.Y.; Mandrikova, O.V.; Budilova, E.A.; Arora, K.; Manjula, L.: Noise in raw data
433 from magnetic observatories. Geosci.Instrum. Methods Data Syst. 6, 329–343, 2017.
- 434 Lingala, M.; Nelapatla, P.C.; Arora, K.: Evaluating the Effect of Noise from Traffic on HYB
435 Magnetic Observatory Data during COVID-19 Lockdown. Appl. Sci. 2022, 12, 2730.
436 <https://doi.org/10.3390/app12052730>, 2022.
- 437 Matzka, J., Chulliat, A., Manda, M., Finlay, C., and Qamili, E.: Geomagnetic observations for
438 main field studies: from ground to space, Space Sci. Rev., 155, 29–64,
439 <https://doi.org/10.1007/s11214-010-9693-4>, 2010.
- 440 Macmillan, S. and Olsen, N.: Observatory data and the Swarm mission, Earth Planets Space,
441 65, 1355–1362, <https://doi.org/10.5047/eps.2013.07.011>, 2013.
- 442 Peltier, A. and Chulliat, A.: On the feasibility of promptly producing quasi-definitive magnetic
443 observatory data, Earth Planets Space, 62(2), e5–e8,doi:10.5047/eps.2010.02.002, 2010.
- 444 Potharaju SVK, Nelapatla P.C.: Development of a robust real-time synchronized data
445 transmission technique from a magnetic observatory to an INTERMAGNET GIN. Scientific
446 Reports, 12:10277. <https://doi.org/10.1038/s41598-022-13820-y>, 2022.
- 447 Reda, J. & Neska, M.: The one second data collection system in Polish geomagnetic
448 observatories. J. Ind. Geophys. Union Spec. 2, 62–66, 2016.



- 449 Torta, J. M. et al.: An example of operation for a partly manned Antarctic Geomagnetic
450 Observatory and the development of a radio link for data transmission. *Ann. Geophys.* 52(1),
451 45–56, 2009.
- 452 Thomson, A. W. P. (ed.): *Geomagnetism Review*, 1–34 (British Geological Survey, 2015),
453 2014.
- 454 Zhang, S., Changhua Fu, Yufei He, Dongmei Yang, Qi Li, Xudong Zhao and Jianjun Wang.:
455 Quality Control of Observation Data by the Geomagnetic Network of China, *Data Science*
456 *Journal*, 15: 15, pp. 1–12, DOI: <http://dx.doi.org/10.5334/dsj-2016-015>, 2016.
457
- 458 Zhang, S, Fu, C, Zhao, X, Zhang, X, He, Y, Li, Q, Chen, J, Wang, J and Zhao, Q.: Strategies
459 in the Quality Assurance of Geomagnetic Observation Data in China. *Data Science*
460 *Journal*, 23: 9, pp. 1–11, <https://doi.org/10.5334/dsj-2024-009>, 2024.