





29 **Non-technical short summary (max 500 char)**

30 StratoBayes is a novel Bayesian method for aligning stratigraphic data from multiple sites. It integrates
31 diverse information, such as geochemical signals and radiometric dates, and provides robust age estimates
32 with quantified uncertainty for all sites. We use StratoBayes to correlate lower Cambrian $\delta^{13}\text{C}$ records from
33 Morocco with an undated record from Siberia, and estimate the age of the world's oldest trilobites.

34 **1 Introduction**

35 Stratigraphic correlation works on the basis that rocks that were deposited under similar conditions or at
36 the same time tend to share characteristics that allow for their attribution to a stratigraphic or temporal
37 horizon. For example, insofar as temporal changes in the global $\delta^{13}\text{C}$ composition of seawater are reflected
38 in marine sedimentary rocks, matching trends of changing $\delta^{13}\text{C}$ in rock sections from different locations
39 can be used to place those sections on a relative time scale (Cramer and Jarvis, 2020; Saltzman et al., 2012).
40 Quantitative signals such as isotopic compositions, elemental concentrations or geophysical well-log data
41 present a particular challenge: in aligning those signals by eye, the stratigrapher has to make a large number
42 of intuitive decisions about which peaks and troughs are likely to line up. Trying to integrate all the
43 stratigraphic evidence from multiple sites often results in more than one potential alignment solution and
44 differing interpretations between different workers (Bowyer et al., 2022, 2023; Landing and Kruse, 2017;
45 Smith et al., 2016).

46 Computer algorithms have been designed to address the problems arising from visual correlation
47 (Agterberg, 1990; Lisiecki and Lisiecki, 2002; Rudman and Lankston, 1973). Algorithms designed for
48 aligning quantitative signals from two or more sites typically use a point-based approach, aligning each
49 point of site A with zero, one or multiple points from site B. This approach proposes variable sedimentation
50 rates between points. This flexibility in principle allows the most precise alignments, though potentially at
51 the cost of overfitting. Point-based algorithms commonly use dynamic time warping (DTW), a technique
52 that finds the optimal match between two time-series data by adjusting their alignment (Sakoe and Chiba,
53 1978). For a selection of recent approaches using dynamic time warping for stratigraphic alignment, see
54 Wheeler and Hale (2014); Hay et al. (2019); Baville et al. (2022); Sylvester (2023); and Hagen et al. (2024).
55 The limitations of DTW-based approaches are that they commonly require known section tops and bottoms
56 (Sylvester, 2023); and they are generally deterministic, providing only a single solution without any
57 indication of uncertainty or alternative alignments (but see Al Ibrahim, 2022; Hay et al., 2019). The



58 integration of additional stratigraphic information besides the quantitative signals tends to be difficult,
59 requiring extra steps outside of the core DTW-algorithm (e.g. Hagen and Creveling, 2024).

60 Probabilistic approaches overcome some of these limitations by estimating the probabilities of different
61 outcomes, rather than producing deterministic predictions. An effective probabilistic approach is offered
62 by the Bayesian framework, which integrates multiple sources of uncertainty by combining prior
63 knowledge, encapsulated mathematically as a prior probability distribution, with a custom likelihood
64 function that is used to evaluate the likelihood of observed data. Given an appropriate prior and likelihood
65 function it is straightforward to integrate different different types of stratigraphic information. Lee et al.
66 (2022) have implemented a Bayesian method that uses Gaussian process regression to match Cenozoic
67 oxygen isotope data from one site to an oxygen isotope stack, while simultaneously integrating age
68 estimates from radiocarbon dates to produce probabilistic age-depth models (i.e. the BIGMACS model).
69 This method improves upon earlier approaches by specifying uncertainty for tie points and integrates prior
70 knowledge on Cenozoic sedimentation rates with absolute age information from the aligned site. However,
71 age uncertainties from the reference site are not included, and varying sampling resolution or large
72 sedimentation rate changes may violate model assumptions and impede the broader adoption of this method
73 in its current form (Middleton et al., 2024). Edmonson and Dyer (2024) have developed a different
74 Bayesian method based on Gaussian process regression that works without prior knowledge of
75 sedimentation rates, but requires minimum and maximum age estimates for all sections, and the absence of
76 an explicit prior on sedimentation rates may risk overfitting.

77 Here, we introduce a versatile Bayesian method for stratigraphic correlation and age modelling that can
78 align quantitative signals from two or more sites without the need to specify tie points or top and bottom
79 ages, and with no restrictions on sampling frequencies. Possible sedimentation rates can be specified by the
80 user as priors, and the likelihood encompasses the alignment of the signals and, optionally, additional age
81 constraints such as dated horizons. The method requires only vague prior knowledge on the ages and on the
82 degree of overlap of the sections, along with order-of-magnitude estimates of sedimentation rates; it is not
83 necessary to specify matching section tops or bottoms. The model is able to integrate radiometric dates
84 from different sites, meaning that ages from well-dated sites can inform age estimates at sites with little or
85 no age information. Age estimates with uncertainty can thus be obtained for any point within any site, and
86 alternative alignments can be identified. Additional stratigraphic knowledge, such as hiatuses or tie points,
87 can be readily incorporated.

88 Our Bayesian model works by evaluating the fit of a single cubic spline (Heaton et al., 2020) to the
89 combined quantitative signal of all sites. If more than one type of signal is included, e.g. $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$, a
90 different spline is constructed for each signal type, and their joint likelihood is used to evaluate the



91 alignment. Different alignments are generated by shifting the sites relative to each other, and by scaling
92 segments of the sites using different sedimentation rates. Markov chain Monte Carlo methods are used to
93 obtain the posterior distributions of the unknown model parameters. Our method is implemented as an R
94 package, ‘StratoBayes’.

95 To demonstrate the potential of this method, we apply it to artificial stratigraphic data and to a real case
96 study using lower Cambrian $\delta^{13}\text{C}$ records from Morocco (Magaritz et al., 1991; Maloof et al., 2005, 2010;
97 Tucker, 1986) and Siberia (Kouchinsky et al., 2007). Integrating radiometric dates (Landing et al., 1998,
98 2021; Maloof et al., 2010), we provide age estimates for the studied sections of an interval spanning several
99 lower Cambrian carbon isotope excursions, and compare our algorithm-derived correlation with recent
100 stratigraphic models relying on visual expert-based interpretations (Bowyer et al., 2022, 2023). Our solution
101 also provides a fully quantitative age estimate for the appearance of the first Siberian trilobites, which are
102 thought to be the world’s oldest trilobites (Landing et al., 2021).

103 **2 Bayesian stratigraphic model**

104 StratoBayes generates and evaluates alignments of quantitative stratigraphic signals. A signal consists, for
105 example, of geochemical or geophysical measurements that vary across height or depth (Fig. 1a), obtained
106 from a contiguous sedimentary sequence which may be interrupted by hiatuses at known horizons.
107 Alignments are generated by shifting the sites containing the signals either a) against a fixed reference site,
108 or b) against each other on an absolute age scale. Additionally, the sites are scaled (“stretched” or
109 “squeezed”) assuming different sedimentation rates. The fit of different alignments, corresponding to
110 different shifts and sedimentation rates, is evaluated in the Bayesian framework.

111 Statistical analysis in the Bayesian framework starts by formulating a probabilistic model that includes
112 known data y and unknown model parameters θ . Instead of trying to identify a single estimate for θ ,
113 Bayesian inference involves estimating probability distributions for the model parameters, termed
114 “posterior probability distributions”. Posterior distributions are obtained by combining prior knowledge of
115 the parameters with the data via a likelihood function. Bayes’ theorem states that the probability of the
116 parameters given the data, $p(\theta|y)$, i.e. the posterior probability, is proportional to the probability of the
117 data given the model parameters (i.e. the likelihood), $p(y|\theta)$, times the prior probability of the model
118 parameters, $p(\theta)$:

$$119 \quad p(\theta|y) \propto p(y|\theta)p(\theta) \quad (1)$$



In our case, this approach requires specifying prior probability distributions for the unknown model parameters that control the shifting and scaling (Fig. 1b), and optionally for the duration of pre-determined hiatuses. Our model assumes that the measurements in each sedimentary sequence are samples (with noise) from a common underlying signal, whose value can be modelled by a smooth curve described by a cubic B-spline. Our likelihood function quantifies how well a cubic B-spline fitted to a given alignment explains the observed data (Fig. 1c). Additional likelihood components can integrate absolute age constraints such as radiometric dates or other tie points, e.g. index fossils. Using Bayes' theorem, the priors are combined with the likelihood to obtain the posterior probability for any alignment.

We obtain probability distributions for the parameters of the model by running a Markov chain Monte Carlo (MCMC) simulation. This involves repeatedly generating parameter values over a large number of iterations. To ensure thorough exploration of the parameter space, we run multiple chains in parallel but retain samples only from the primary (cold) chain. An initial portion of the samples is discarded (burn-in) to remove dependency on starting values, and only every n^{th} iteration is recorded to reduce autocorrelation. Details on the MCMC implementation are provided in Appendix A.

In the following, we will assume that measurements were taken on a height scale (increasing from the bottom to the top), but depth-scale measurements can be used interchangeably by inverting their sign.

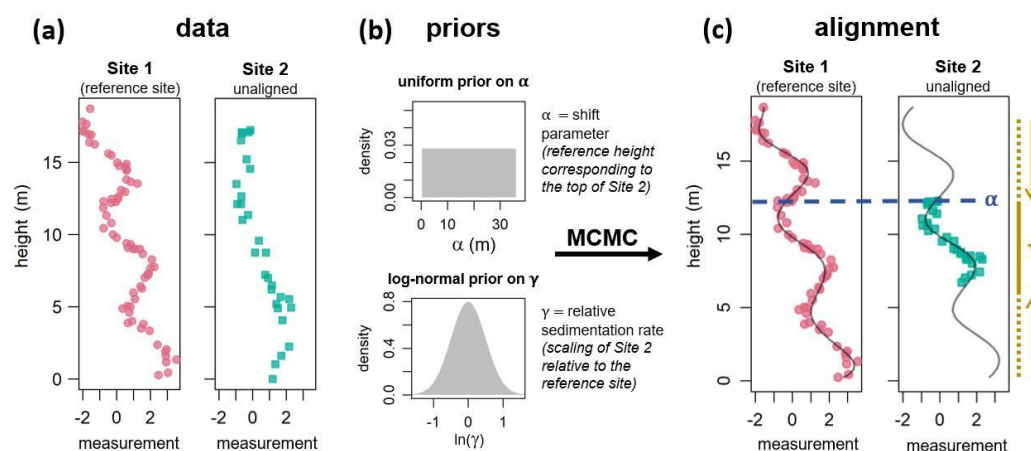


Figure 1: Schematic of the alignment algorithm. a) Input data: Quantitative stratigraphic measurements (e.g. geochemical data) from two sites recorded along their section height (here given in meters). b) Priors must be placed on the shift parameter α and on the relative sedimentation rate γ . Here, α determines the reference height (at Site 1) corresponding to the top of the height range of Site 2, and γ corresponds to the sedimentation rate of Site 2 relative to Site 1. c) An alignment corresponding to a



single sample from the posterior. The blue dashed line indicates the position of the top of the data from Site 2 at the reference height scale (α ; median: 12.5 m). The relative sedimentation rate γ has been estimated at a median of 2.8, corresponding to a shortening of the dataset from Site 2 relative to the reference site (indicated by the dashed and solid light brown line). The curved grey line shows the cubic B-spline corresponding to the alignment.

136 2.1 Evaluating alignments with a cubic B-spline

137 Identifying good alignment positions requires evaluating and comparing different potential alignments. In
138 the Bayesian framework, the measure used for this evaluation is the likelihood. We derive the likelihood of
139 an alignment from its fit to a single cubic B-spline (Eilers and Marx, 1996), fitted to the measurements from
140 all sites, including the reference site (see Fig. 1c).

141 We model each measured value y_i as normally distributed:

$$142 \quad y_i \sim \text{Normal}(\mu_i, \sigma), \quad (2)$$

143 where μ_i is the mean, and the standard deviation σ represents the scatter around the spline. μ_i is given by
144 the spline function

$$145 \quad \mu_i = \sum_{j=1}^{k+2} \beta_j B_j(h_i) \quad (3)$$

146 Here, k denotes the number of internal knots of the spline, with more knots implying that the spline can
147 potentially capture higher-frequency variations. β_j is the spline coefficient associated with the j -th basis
148 function, and $B_j(h_i)$ is the j -th B-spline basis function evaluated at a reference height h_i . A roughness
149 penalty controlled by a smoothing parameter λ is incorporated in the prior on $\boldsymbol{\beta}$, such that higher values of
150 λ serve to favour smoother splines (Appendix A). The number of knots and the roughness penalty each
151 influence spline flexibility in different ways: increasing k provides a finer resolution for fitting local
152 features, whereas increasing λ penalizes abrupt changes and yields smoother fits. The knots for the spline
153 can be distributed across the reference height range that the converted measurement heights occupy for a
154 specific combination of α and γ parameters. Our current model implementation uses evenly spaced knots,
155 but knot placement could also follow, for example, the density of measurements. Alternatively, the knots
156 can be fixed at specific heights on the reference scale, in which case combinations of α and γ that result in
157 converted measurement heights falling outside the knot range cannot be evaluated.



158 The likelihood of an alignment, given β , σ and λ , is determined by the residual deviations of the y_i values
159 from the corresponding μ_i values. The overall likelihood for n data points is obtained by taking the product
160 over all individual likelihoods for each pair of y_i and μ_i :

$$161 \quad L(\mathbf{y}|\beta, \sigma, \lambda) = \prod_{i=1}^n \frac{1}{\sqrt{2\pi\sigma^2}} \times e^{\left(-\frac{(y_i - \mu_i)^2}{2\sigma^2}\right)} \quad (4)$$

162 We thus assume that the deviations of the data from the spline are independently and identically distributed
163 according to a normal distribution with mean 0 and standard deviation σ .

164 Our model allows for using more than one type of measurement simultaneously. In this case, a separate
165 spline is fitted to all data, from all sites, for each type of measurement. The product of all likelihoods from
166 all measurement types gives the overall likelihood.

167 **2.2 Alignment and partitioning**

168 In order to generate alignments of stratigraphic signals from different sites, one site is picked as a fixed
169 reference site. The other sites are shifted and stretched (or squeezed) relative to the fixed reference site r .
170 This requires specifying a height α_s , which anchors an arbitrary, specified height of site s to a height in the
171 reference site r . Here, we anchor the top of site s , so we set $\alpha_s = \alpha_{top,s}$ meaning $\alpha_{top,s}$ will be the height
172 at site r that aligns with the top of site s . To stretch or squeeze site s , a relative sedimentation rate γ_s can
173 be specified, where γ_s is defined relative to the reference site. For any height $h_{x,s}$ at site s , the corresponding
174 height in the reference site r can then be calculated as

$$175 \quad h_r = \alpha_{top,s} - \frac{1}{\gamma_s} \times (h_{top,s} - h_{x,s}), \quad (5)$$

176 where $h_{top,s}$ is the height of the top of site s . Although we here chose the top of site r as the reference
177 horizon α for simplicity, any horizon at site r can be used as α . A $\gamma_s < 1$ implies that site s has a lower
178 sedimentation rate than site r , and consequently, s has to be stretched to match r . A $\gamma_s > 1$, i.e. a higher
179 sedimentation rate at site s will lead to s being squeezed to match r .

180 The model described here is simple in that the same γ is applied to all measurements of the same site. In
181 this scenario, any site may be used as the reference site. Below, we introduce more complex models with
182 more than one sedimentation rate per site, and with hiatuses. With these models, it is practical to select the
183 site with the most sedimentation rate changes and hiatuses as the reference site. This reduces the number
184 of unknown parameters in the model, making it easier to obtain a representative sample from the posterior.



185 2.2.1 Multiple sedimentation rates per site

186 Instead of having one sedimentation rate per site, sites can be partitioned, reflecting for example lithological
187 units, with each partition being modelled with a distinct sedimentation rate:

$$188 \quad h_r = \alpha_{top,s} - \sum_l^{n_{p,s}-1} \left(\frac{1}{\gamma_p} \times (h_{p,s} - h_{p+1,s}) \right) - \frac{1}{\gamma_{n_{p,s}}} \times (h_{n_{p,s}} - h_{x,s}), \quad p = 1 \dots n_{p,s} \quad (6)$$

189 Here, $n_{p,s}$ is the number of partitions encountered from h_{top} to $h_{x,s}$, $h_{p,s}$ is the top height of partition p at
190 site s , and $h_{p+1,s}$ is the top height of the partition below partition p at site s . If $h_{x,s}$ falls in the first partition
191 from the top, the calculation simplifies to the equivalent of Equation 5, with $h_{p,s}$, the top height of the first
192 partition being also the top height of site s . The relative sedimentation rates of partitions, γ_p , can differ for
193 each partition in each site, or partitions in different positions within a site or across sites may share
194 sedimentation rates.

195 2.2.2 Site-specific sedimentation rate multipliers

196 The sedimentation rate model above can be further expanded by adding an overall site-specific
197 sedimentation rate multiplier ζ_s :

$$198 \quad h_r = \alpha_{top,s} - \sum_l^{n_{p,s}-1} \left(\frac{1}{\zeta_s \gamma_p} \times (h_{p,s} - h_{p+1,s}) \right) - \frac{1}{\zeta_s \gamma_{n_{p,s}}} \times (h_{n_{p,s}} - h_{x,s}), \quad p = 1 \dots n_{p,s} \quad (7)$$

199 This may be useful in scenarios where sedimentation rates systematically differ between sites, perhaps due
200 to varying distances from a sediment source, but where the sedimentation rate ratios of different partitions
201 are assumed to be constant across sites.

202 2.2.3 Hiatuses

203 Known hiatuses (also referred to as unconformities or stratigraphic gaps) can be included at specific pre-
204 defined locations in a site. Expanding Equation 5 to include gaps of height δ , we obtain

$$205 \quad h_r = \alpha_{top,s} - \frac{1}{\gamma_s} \times (h_{top,s} - h_{x,s}) - \sum_g^{n_{G_s}} \delta_g, \quad g = 1 \dots n_{G_s} \quad (8)$$

206 where n_{G_s} is the number of gaps encountered from $h_{top,s}$ until height $h_{x,s}$.



2.2.4 Tie points

Tie points define specific heights within an aligned site and assign a probability distribution to indicate to which horizon these heights correspond on the reference scale. For example, a tie point might be a lithological boundary, a biostratigraphic horizon, or a radiometric date. If tie points are specified, the likelihood of an alignment is expanded to include not only the fit of the signal data to the spline, but also the positions of the ties on the reference height scale relative to the specified probability distribution.

For example, a point in an aligned section which is tied by observation to the reference section at a position m_t with a normally distributed uncertainty with standard deviation s_t that ends up being shifted to a reference height h_t (computed from the relevant α and γ parameters) contributes a likelihood of

$$L(m_t|h_t, s_t) = \frac{1}{\sqrt{2\pi s_t^2}} \times e^{\left(-\frac{(m_t-h_t)^2}{2s_t^2}\right)} \quad (9)$$

to the overall likelihood of the model.

2.2.5 Age-scale alignment

Data on an (absolute) age scale can be aligned using the methods introduced above by using ages instead of heights. However, height-scale data can be aligned on an age-scale if absolute age constraints (specified as ties) are provided from at least one site. In this case, all sites will be shifted to align on a common age scale, i.e., there is no reference site.

Analogous to the heights in the reference height scale in Equation 5, ages (a) can be calculated as:

$$a = \alpha_{top,s} + \frac{1}{\gamma_s} \times (h_{top,s} - h_{x,s}) \quad (10)$$

Here, $\alpha_{top,s}$ is the top age (minimum age), rather than top height (maximum height), of site s . Sedimentation rates γ_s need to be expressed on the common age scale, rather than relative to a reference site. Equations 6–8 can be modified accordingly for an analysis on the age scale.

It should be noted that due to sedimentation rates being fixed for an entire site or within partitions, our current model implementation does not necessarily result in increasing age uncertainty away from absolute age constraints. Potential sedimentation rate changes within sites or partitions could lead to our model underestimating age uncertainty with growing stratigraphic distance from absolute age constraints (see De Vleeschouwer and Parnell, 2014).



233 2.3 Priors

234 The Bayesian framework requires priors to be placed on all unknown model parameters. In our model, these
235 include the alignment parameters (e.g. α , γ), the smoothing parameter λ , the residual standard deviation σ
236 (if it is not fixed), and the spline coefficients β . The priors on the alignment parameters determine the range
237 of possible alignments and need to be chosen with care. For the other parameters, weakly informative priors
238 with minimal influence on the analysis are preferred (Appendix A). In addition to those priors, we penalise
239 a lack of overlap by specifying a prior probability of data points from different sites overlapping each other.

240 2.3.1 Alignment parameters

241 The priors on the alignment parameters should reflect the stratigraphic knowledge on the input data. The
242 user may specify different types of prior distributions (e.g., normal, uniform, exponential) for the alignment
243 parameters during model setup.

- 244 • α determines the reference site (site r) height or age that a specific position within the aligned site
245 (site s) corresponds to. In the absence of prior knowledge on how the sites are likely to align, a
246 uniform prior can be placed on α . For example, if α refers to the top of site s , a uniform prior on
247 α with min and max equal to the height or age range of site r implies that the top of site s will be
248 placed within the height range of site r .
- 249 • γ is either a relative (*height scale alignment*) or an absolute (*age scale alignment*) sedimentation
250 rate. In our model implementation, priors are placed on the natural logarithm of γ , $\ln(\gamma)$, rather
251 than on γ directly. Specifying rate parameters on the logarithmic scale ensures that their priors are
252 symmetric: a doubling or halving of a rate has equivalent distances on the logarithmic scale. If the
253 sedimentation rate is relative, $\ln(\gamma) < 0$ (i.e. $\gamma < 1$) results in “stretching”, and $\ln(\gamma) > 0$
254 (i.e. $\gamma > 1$) results in “squeezing” of site s relative to site r . In the absence of strong prior
255 knowledge about the relative sedimentation rate, a normal prior on $\ln(\gamma)$ with a mean of 0 places
256 equal prior probability on “stretching” or “squeezing” of site s relative to site r . The standard
257 deviation requires at least a broad guess of the potential magnitude of sedimentation rate
258 differences. For example, a standard deviation of $\frac{\ln(4)}{1.96}$ places 95% of prior probability on $\frac{1}{4} <$
259 $\gamma < 4$ for $\ln(\gamma) \sim \text{Normal}\left(0, \frac{\ln(4)}{1.96}\right)$. If γ is an absolute sedimentation rate, the range of plausible
260 prior sedimentation rates may be estimated from the absolute age constraints.



- 261 • ζ_s is a multiplier applied to all relative or absolute sedimentation rates γ corresponding to a single
262 site s . As with γ , $\ln(\zeta_s) < 0$ (i.e. $\zeta_s < 1$) causes additional “stretching”, and $\ln(\zeta_s) > 0$ (i.e. $\zeta_s >$
263 0) causes additional “squeezing” of site s .
- 264 • δ is the reference height range or duration of a hiatus. An exponential prior may be useful when
265 little is known about the extent of the hiatus, placing higher probabilities on short extents. The
266 rate needs to be chosen to make sense in the context of the height of the sections, or of the
267 anticipated age range of the sites.

268 2.3.2 Penalising a lack of overlap

269 Individual splines fitted to data from each site separately can almost always follow the data more closely
270 than a single spline fitted to aligned data from all sites. Given enough knots, alignments in which the data
271 do not overlap, or only overlap little, will thus generally result in a higher likelihood than alignments with
272 a partial or full overlap. This means that if the priors allow non-overlapping alignments, those will generally
273 be preferred in the model inference. To counteract this tendency, we impose a prior on the overlap of each
274 individual data point from all sites that penalises non-overlap with data from other sites.

275 The prior on overlap for data point i from site s is

$$276 \quad P(i_s) = e^{(-\sqrt{S-1} + \sqrt{S_{\text{overlap},s,i}}) \times c_{\text{overlap}}}, \quad (11)$$

277 where $S_{\text{overlap},s,i}$ is the number of other sites overlapping the reference height h_r or age a of point i_s , and
278 c_{overlap} is a constant. This formulation implies that the penalty for a point i_s that overlaps all other sites is
279 0, and the penalty is strongest (most negative) if i_s overlaps no other sites. To work effectively, the penalty
280 needs to be stronger for data sets with little noise (low residual σ), to offset the larger likelihood differences
281 resulting from fitting a spline with low σ . A range of c_{overlap} values may work in practice. A formulation
282 that we have found works well in many scenarios sets

$$283 \quad c_{\text{overlap}} = c \times \frac{1}{S} \sum_{s=1}^S \left(\frac{\sigma_{y,s}}{\sigma_s} \right)^q \quad (12)$$

284 where c is a constant determining the strength of the overlap penalty (set to a default of $c = \frac{1}{4}$), $q = 1$ if σ
285 is fixed, and $q = \frac{1}{2}$ if σ is variable (i.e. estimated in the model inference). Here, $\sigma_{y,s}$ is the standard deviation
286 of all data y from site s , and σ_s is the residual standard deviation of a Bayesian spline fitted to the data y
287 from site s , using the same priors as for the overall model inference.



288 3 Model illustration

289 We illustrate the performance of our stratigraphic alignment method with a simple, artificial dataset (Fig
290 2a). We generated measurements from a reference site (Site_{ref}) using a sine wave covering 3.5 periods,
291 where each period corresponds to 2π radians. To generate the signal data, we intercepted this sine wave at
292 heights h with 250 evenly spaced points per period, i.e. the number of data points (n) is $3.5 \times 250 = 875$.
293 Each signal value y_i was generated with random white noise $\sigma = \frac{1}{5}$ added, such that

$$294 \quad y_i \sim \text{Normal}\left(\eta_i \sin\left(h_i - \frac{1}{2}\pi\right), \sigma\right), \quad i = 1 \dots n \quad (13)$$

295 The factor η_i modulates the amplitude of the sine wave at each height h_i . It was set to $\eta = 1$ for the heights
296 ranging from -0.5π to 5π , and to $\eta = 0.75$ from heights 5π to 6.5π , which reduces the amplitude
297 beginning in the middle of the third period of the sine wave. The aligned signal was simulated as above,
298 but from a sine wave covering one period, sampling 250 data points, again with random noise using $\sigma = \frac{1}{5}$
299 and $\eta = 1$. To simulate a sedimentation rate twice as high as at the reference site, we multiplied the heights
300 of $\text{Site}_{\text{align}}$ by 2. The heights of $\text{Site}_{\text{align}}$ were then shifted to start at 0.

301 The aligned signal should thus match either the first or the second, but not the third period of the reference
302 signal. To align the two sites, we used a simple model with a site-specific shift α , referring to the top of
303 $\text{Site}_{\text{align}}$ and relative sedimentation rate γ as in Equation 5. From the data generation, we know that the
304 posterior of γ should be ≈ 2 , with $\ln(\gamma) \approx 0.69$, and α (defined as the reference height corresponding to
305 the top height of $\text{Site}_{\text{align}}$) should be $\approx 2\pi$ (top of first period) or $\approx 4\pi$ (top of second period).

306 To minimise the influence of the priors, we used a uniform prior on α that extends well beyond the
307 alignment positions known from generating the data, and a broad normal prior on $\ln(\gamma)$ that encompasses
308 the known sedimentation rate $\gamma = 2$ (Fig. 2b):

$$309 \quad P(\alpha) \sim \text{Uniform}(-\pi, 8\pi) \quad (14)$$

$$310 \quad P(\ln(\gamma)) \sim \text{Normal}(0, 1) \quad (15)$$

311 These priors place 95% of prior probability for the relative sedimentation rate of $\text{Site}_{\text{align}}$ between 0.14 and
312 7.1, and place the top of $\text{Site}_{\text{align}}$ anywhere from half a period below the start of the first period ($-\pi$) up to
313 one period above the third period (8π). For the cubic spline, we specify 20 evenly spaced knots, which is
314 more than enough to approximate the three periods of the sine wave.



315 We estimated the posterior of the model with three independent runs, each with 16 chains and 60,000
316 iterations. The first 10,000 samples were discarded as burn-in, and every 25th iteration was recorded,
317 resulting a total of 6000 samples after burn-in across all three independent model runs.

318 The results show that the analysis identified both matching alignments, corresponding to the first and
319 second period of the reference site (Fig. 2b). The posterior probability for (Site_{align}) matching period 1 is
320 50.1%, and 49.9% for matching period 2. A density plot of the posterior of α and $\ln(\gamma)$ shows that α has a
321 bimodal posterior, corresponding to the two alignments (Fig. 2c). The trace plots indicate good mixing of
322 the chains (Fig. 2d), suggesting that the posterior estimates are robust.

323 It is notable that the model estimate for the relative sedimentation rate γ is lower at 1.90 (95% credible
324 interval: 1.82 to 1.99) than the value used for the data generation (2.00). Reported values, here and
325 throughout, represent the posterior median, with 95% credible intervals – given in brackets – refer to the
326 interval between the 2.5% and 97.5% points of the posterior distribution. This deviation of the posterior
327 from the known sedimentation rate estimate arises because the priors favour greater overlap (see Sect.
328 2.3.2). The posterior alignment tends to “compress” the data from Site_{align} slightly less than expected,
329 leading to an increased overlap of points (see also Fig. 5b).

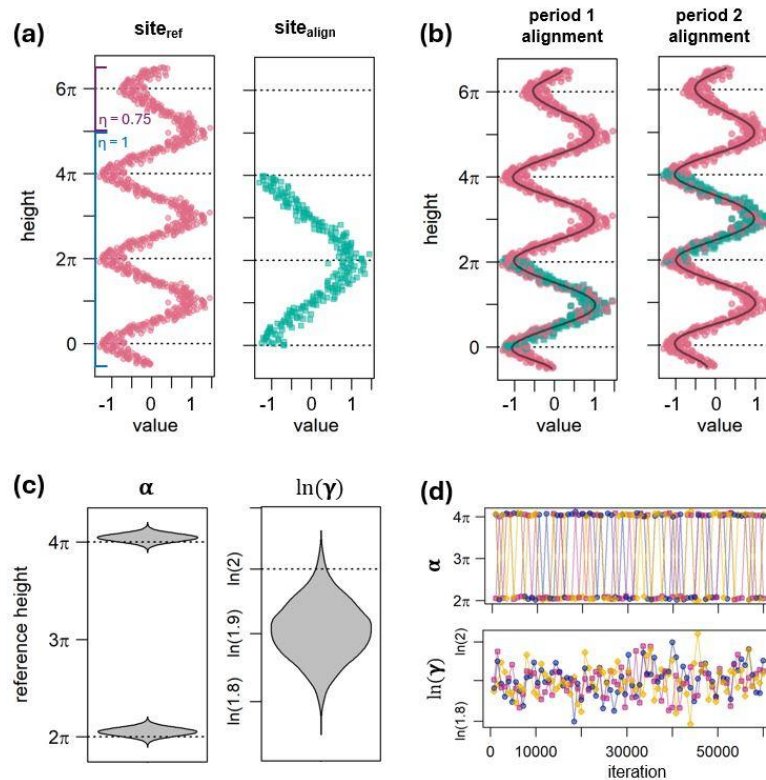


Figure 2: Model illustration using artificial data. a) Input data: Quantitative stratigraphic data from two sites. The blue line indicates the range in which $Site_{ref}$ was created with $\eta = 1$, and the purple line above indicates the range for which η was set to 0.75 to lower the amplitude. b) Two alignments identified by the inference, with ($Site_{align}$; blue squares) matching the first or second period of ($Site_{ref}$; red points). The alignments shown here correspond to two distinct samples from the posterior; other samples will result in slightly different positions of ($Site_{align}$). The curved dark lines show the cubic spline corresponding to each alignment. c) Posterior densities of α and $\ln(\gamma)$. The two modes of α correspond to the two distinct alignments in b). The dotted lines indicate the γ values with which ($Site_{align}$) was simulated, and the two plausible α values. d) Trace plots of α and $\ln(\gamma)$. The three distinct colours correspond to the three independent model runs. For visual clarity, only 75 selected samples are shown from each run.



330 4 Case study: Lower Cambrian $\delta^{13}\text{C}$ records

331 To demonstrate the utility of this method, we use it to align stable carbon isotope records ($\delta^{13}\text{C}$) from lower
332 Cambrian marine shelf carbonates (Fig. 3). We integrate a combination of radiometric dates, $\delta^{13}\text{C}$ and
333 astrochronological information from four sites to obtain age estimates for the sampled intervals from all
334 sites, and use this age model for dating the first documented occurrence of Siberian trilobites.

335 4.1 Data

336 We selected three records from the Anti-Atlas mountains in Southern Morocco, corresponding to the Oued
337 Sdas, the Tiout and the Talat n' Yissi sections, which were part of West-Gondwana during the early
338 Cambrian (Magaritz et al., 1991; Maloof et al., 2005, 2010; Tucker, 1986). Oued Sdas and Tiout harbour
339 multiple precise U-Pb radiometric ages (Landing et al., 2021; Maloof et al., 2010). Talat n' Yissi has no
340 radiometric dates, but a radiometric date exists from the stratigraphically equivalent Lemdad syncline
341 (Landing et al., 1998) that has been correlated biostratigraphically to Talat n'Yissi with the *Antatlasia gutta-*
342 *pluviae* zone (Maloof et al., 2005); we include this date in the analysis. We will align these sites with each
343 other, and with a $\delta^{13}\text{C}$ record from the Sukharikha section from the northwestern Siberian platform
344 (Kouchinsky et al., 2007), corresponding to the palaeocontinent Siberia. There are no radiometric dates
345 available for the Siberian section for this stratigraphic interval. Data that was inferred to be below the lower
346 leg of the prominent “5p” excursion (lowest peak in Fig. 3a and d) was excluded to simplify the correlation,
347 reducing the number of modelled sedimentation rates unconstrained by radiometric dates. This cropping of
348 data affects the Oued Sdas and Sukharikha sections; Fig. 3 shows all data that was included in the analysis.

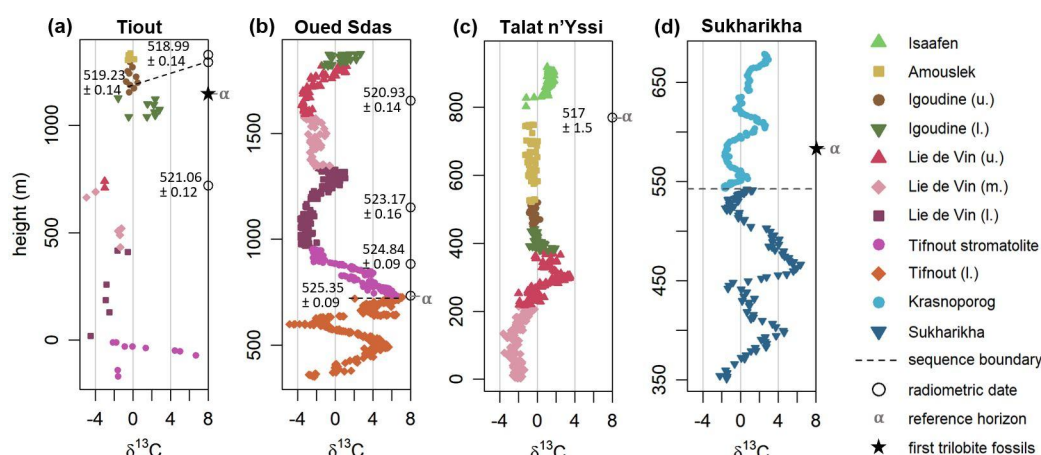




Figure 3: Cambrian $\delta^{13}\text{C}$ data and radiometric dates from Morocco (a - c) and Siberia (d). Different colours, in conjunction with different symbols, delineate different lithological units or formations. Circles indicate the position of radiometric dates, with mean age and 2 standard deviations denoting the uncertainty. Stars denote the positions where the oldest trilobite remains are found in Morocco (a) and Siberia (d). The dashed line in (d) indicates a hiatus. α indicates the reference horizon chosen for specifying the prior on the shift parameter α for each site.

349 4.2 Model specification

350 To align the four sites on the age scale, we specify an α parameter on the absolute age scale (Ma) for each
 351 site, and use absolute, rather than relative sedimentation rates (expressed in mMyr^{-1}). We encapsulate
 352 variation in sedimentation rates (γ) by partitioning sites into members, formations or lithological units,
 353 leading to multiple sedimentation rates per site. As there are few radiometric dates to constrain
 354 sedimentation rates, partitions shared between the Moroccan sites are set to have the same relative
 355 sedimentation rate across sites. To account for potentially faster or slower sedimentation rates at different
 356 sites, a site-specific sedimentation rate multiplier ζ is added for Oued Sdas and Talat n'Yissi that is
 357 multiplied with the γ from those sites. The γ for a partition applies to all sites at which this partition occurs;
 358 for Tiout, they are used unaltered, and no ζ is needed for Sukharikha as there are no shared partitions with
 359 other sites. We partition the Moroccan data based on the lithostratigraphy from Maloof et al. (2005). We
 360 divide the Adoudounian Tifnout Member into a lower part (Tifnout l.), and an upper stromatolitic part
 361 (Tifnout stromatolite), as preliminary results suggested pronounced sedimentation variability between those
 362 parts. We subdivide the Lie de Vin Formation into three members; the Igoudine Formation is subdivided
 363 into two members. The Amouslek and Isaafen formations are not subdivided. The Sukharikha section is
 364 divided into two formations, which we assign separate sedimentation rates. At the boundary, a substantial
 365 hiatus is evinced by the truncation of the “7p” $\delta^{13}\text{C}$ peak (Kouchinsky et al., 2007). We include the duration
 366 of this hiatus (δ) as an additional unknown parameter in the model.

367 The model requires priors to be specified for each of its 18 alignment parameters: Four α , eleven γ , two ζ
 368 and one δ (Fig. 4). These priors are broadly guided by the radiometric dates and by previous work (Bowyer
 369 et al., 2023; Landing et al., 2021; Sinnesael et al., 2024). The α for the Tiout and Sukharikha sites are placed
 370 at the height positions of the first trilobite fossil remains found at Tiout (Sinnesael et al., 2024), and the first
 371 appearance of Siberian trilobites correlated to Sukharikha (Landing et al., 2021; Varlamov et al., 2008).
 372 Here, we place normal distributed priors with mean age 520 Ma and a wide standard deviation of 2 Myr on
 373 the α parameters at Tiout and Sukharikha. This prior reflects the notion that first appearance dates of



374 trilobites may be broadly similar at ≈ 520 Ma, but not necessarily identical, and the data is allowed to
375 determine the exact age of each α . The α priors for Oued Sdas and Talat n'Yissi are placed at the position
376 of the lowest or the only available radiometric date, respectively, consisting of normal distributions with
377 mean age equal to the mean age estimate of the radiometric data and a wide standard deviation of 2 Myr.

378 For the sedimentation rates, priors informed by an astrochronology of the Tiout section (Sinnesael et al.,
379 2024) are used for the following five stratigraphic partitions: The lower, middle and upper members of the
380 Lie de Vin Formation, and for the lower and upper (Tiout Member) members of the Igoudine Formation.
381 Those priors are chosen such that the 95 percentile interval of γ spans the minimum and maximum of the
382 astrochronological sedimentation rate estimates when using an uncertainty of ± 1 short eccentricity cycle
383 for each partition, with an estimated duration of short (≈ 100 kyr) eccentricity cycles ranging from 92.5 to
384 100.5 kyr (two standard deviations, following Lantink et al., 2022).

385 To specify priors for the remaining Moroccan partitions (lower part of Tifnout Fm., Tifnout stromatolite,
386 Amouslek Fm., and Isaafen Fm.), sedimentation rates between the radiometric dates from Oued Sdas and
387 Tiout are calculated using the mean ages of the dates. The prior on $\ln(\gamma)$ is defined as a normal distribution
388 with a mean of 5.39, corresponding to the mean of the empirical sedimentation rates from Oued Sdas and
389 Tiout, calculated on the logarithmic scale. A wide standard deviation of 0.75 is set, resulting in the 95
390 percentile interval of γ spanning 50.3 to 951 m Myr $^{-1}$. This interval significantly exceeds the range of
391 sedimentation rates inferred from the radiometric dates at Oued Sdas and Tiout, 147 to 314 m Myr $^{-1}$,
392 allowing for the possibility of lower or higher sedimentation rates in some partitions.

393 Prior sedimentation rate estimates for the Siberian formations are estimated in the absence of radiometric
394 dates, very broadly based on global correlations by Bowyer et al. (2023). These correlations suggest average
395 sedimentation rates on the order of 20 to 30 m Myr $^{-1}$; we place a normal prior on $\ln(\gamma)$ with a mean of
396 3.30 and a standard deviation of 0.75, resulting in a 95 percentile interval of γ spanning 6.23 to
397 117.9 m Myr $^{-1}$, which allows for the possibility of significantly different sedimentation rates from those
398 inferred by Bowyer et al. (2023).

399 Finally, a prior needs to be placed on the duration of the hiatus δ between the Sukharikha and the
400 Krasnoporog formations. Kouchinsky et al. (2007) do not give an indication of the potential duration of this
401 hiatus, but if the under- and overlying $\delta^{13}\text{C}$ peaks are correlated as indicated by previous work (Bowyer et
402 al., 2022; Landing et al., 2021), a relatively short hiatus of ≈ 1 Myr is likely. To express considerable
403 uncertainty about the duration of the hiatus, we place an exponential prior on δ with a rate of 1, which
404 places 95% of prior probability on the duration being < 3 Myr, with 5% probability accounting for the
405 possibility of a longer gap.



406 The cubic spline comprises 40 evenly spaced knots, allowing it to closely follow trends in the $\delta^{13}\text{C}$ records
 407 while keeping the MCMC runtime manageable, as a higher knot count increases computational cost. For
 408 the smoothing parameter λ , we applied a gamma prior with StratoBayes' default values of $a_\lambda = 1$ and $b_\lambda =$
 409 1000. We fixed σ , which is the residual standard deviation of the overall spline, at 0.66, which is the
 410 average residual standard deviation of individual cubic splines fitted to each $\delta^{13}\text{C}$ record from the four
 411 respective sites. These individual splines were constructed with 40 knots evenly spaced across the height
 412 range of each respective site and fitted with Gibbs sampling using 2000 iterations, discarding 25% of
 413 samples as burn-in. The same default λ priors as described above were applied, while the prior for these
 414 splines' standard deviations was specified as a gamma prior on the precision τ , with $a_\tau = b_\tau = 0.01$ (see
 415 Appendix A for details).

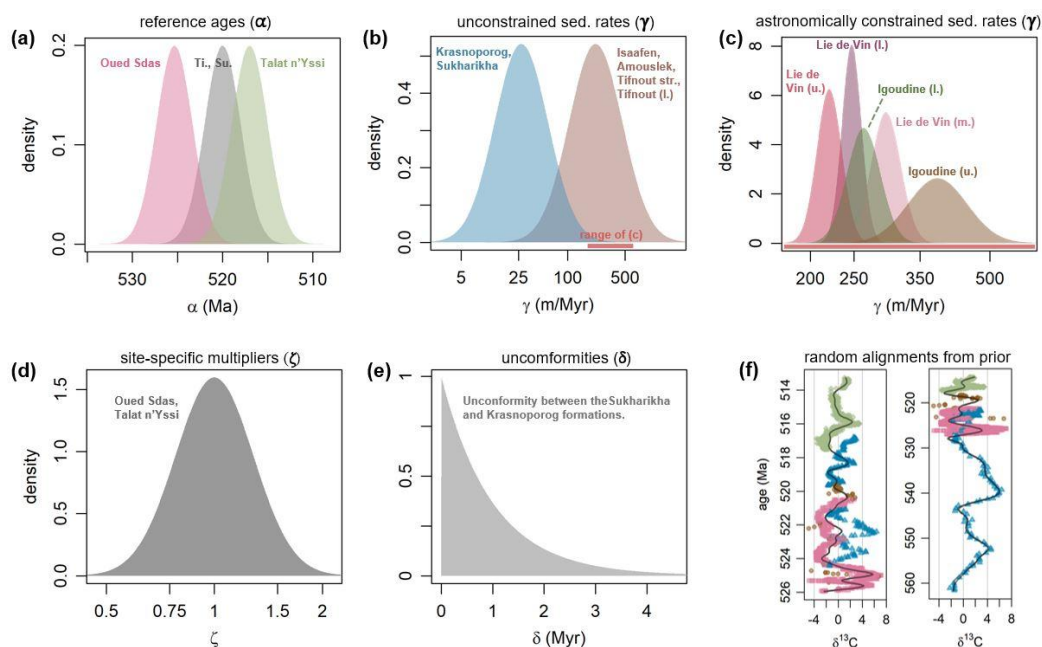


Figure 4: Priors on the 18 alignment parameters for the Cambrian model. Prior probability density is shown (a) for four α parameters corresponding to one site each (priors for Tiout and Sukharikha in grey are identical), (b) for six γ (sedimentation rate) parameters with little prior knowledge, (c) for five γ parameters from Morocco with tight priors based on astrochronology, (d) for ζ parameters (site-specific sedimentation rate multipliers) for Oued Sdas and Talat n'Yissi (identical), and (e) for the duration of the hiatus between the Sukharikha and the Krasnoporog formations. The width of the red bar in (b) visualises the range of sedimentation rates spanned by (c). Panel (f) visualises two alignments



generated by randomly drawing parameter values from their respective priors, to give an indication of the broad range of alignments that the priors on the alignment parameters allow; colours correspond to the four sites (see Fig. 6). Panels (b), (c), and (d) are depicted with a logarithmic x-axis as the priors were specified on $\ln(\gamma)$ and $\ln(\zeta)$.

416 4.3 Parameter estimation

417 This model is more complex than our earlier examples, and hence requires longer runs with more chains.
418 We conducted four independent model runs, each with 750,000 iterations and 24 chains. The runs were
419 executed in parallel using four workers on a desktop computer (Intel i7-10700 CPU, 8 cores / 16 threads,
420 40 GB RAM) and completed within 5 days. The first 150,000 iterations were discarded as burn-in. From
421 the remaining 600,000, every 50th iteration was retained, resulting in 12,000 samples per run and 48,000
422 samples in total.

423 Inspection of trace plots of the model runs indicates stationarity and good mixing of the chains with the
424 exception of infrequent visits of secondary posterior modes (Appendix B, Fig. B1). The potential scale
425 reduction factor (using eq.4 in Vats and Knudson, 2021) is between 1.00 and 1.05 for all alignment
426 parameters, suggesting approximate convergence of the MCMC. The multivariate effective sample size
427 (Vats et al., 2019) of the 48,000 samples is 4161.

428 4.4 Results

429 To identify distinctly different alignments in the posterior, a hierarchical density-based cluster analysis
430 (Campello et al., 2015) was conducted using the inferred ages of all partition boundaries of the four sites
431 (Fig. 4a,b). We specified 1% of samples (480) as the minimum number of points per cluster, resulting in
432 three distinct clusters with 93%, 2.8% and 2.6% of posterior samples, respectively, and 1.5% of samples
433 not being assigned to any cluster. These alignment clusters also differ in the prior probabilities and
434 likelihoods associated with individual posterior samples. On average, samples from alignment 1 tend to
435 exhibit a lower degree of overlap, but a higher likelihood (Fig. 4c), indicating a better fit to the data.

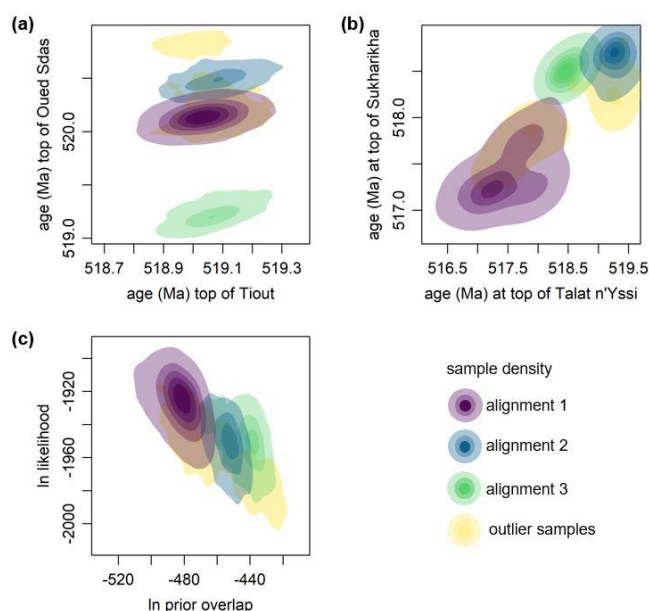


Figure 5: (a, b) 2D density plots of the inferred top ages of the four sites, representing some of the ages used for obtaining alignment clusters from posterior samples. (c) 2D density plot of the ln prior probability of overlap against the overall ln likelihood. Areas with more opaque shadings correspond to a higher density of individual posterior samples. Colours correspond to alignment clusters: alignment 1 - violet; alignment 2 - blue; alignment 3 - green; outlier samples not assigned to any cluster - yellow.

Using samples from the posterior of the model parameters, alignments can be generated. Fig. 6 visualises three alignments drawn from the three alignment clusters identified in the posterior. For each alignment cluster, the iteration with partition boundary ages that are, on average, closest to the median ages of the partition boundaries within that cluster is selected for displaying. All three alignments exhibit a good match between the long-term trends of the $\delta^{13}\text{C}$ curves from the four sites and the common spline curve, although many shorter-term deviations are visible (Fig. 6a-c). The spline curve notably follows the more densely sampled sites (Oued Sdas, Talat n'Yissi) more so than the thinly sampled sites (Tiout, Sukharikha), resulting in greater deviations of the latter two sites.

The posterior age estimates for the stratigraphic positions of the radiometric dates broadly match the age estimates that were used as inputs in the analysis (Fig. 6d). The deviations are greatest for the Talat n'Yissi date (Ta_1), which has large uncertainty and therefore less influence on the analysis, and the second date from Oued Sdas (Ou_2). The first appearances of trilobites are visualised alongside the dates in Fig. 6d, and

are dated to 519.46 Ma (519.25 to 519.68 Ma) at Tiout. The age estimate for the first Siberian trilobites differs considerably between the different alignment solutions: For the most likely alignment 1, the age estimate is 520.79 Ma (520.98 to 520.61 Ma), and for alignment 2 the estimate is somewhat higher at 521.05 Ma (521.19 to 520.91 Ma). Alignment 3 suggests a significantly later appearance of Siberian trilobites at 519.98 Ma (520.15 to 519.84 Ma). All three alignments place the appearance of the first Siberian trilobites before their appearance at Tiout, with the temporal gap (computed directly from the posterior distribution) being estimated at 1.33 Myr (1.09 to 1.54 Myr) for alignment 1, 1.71 Myr (1.54 to 1.87 Myr) for alignment 2, and 0.63 Myr (0.53 to 0.74 Myr) for alignment 3.

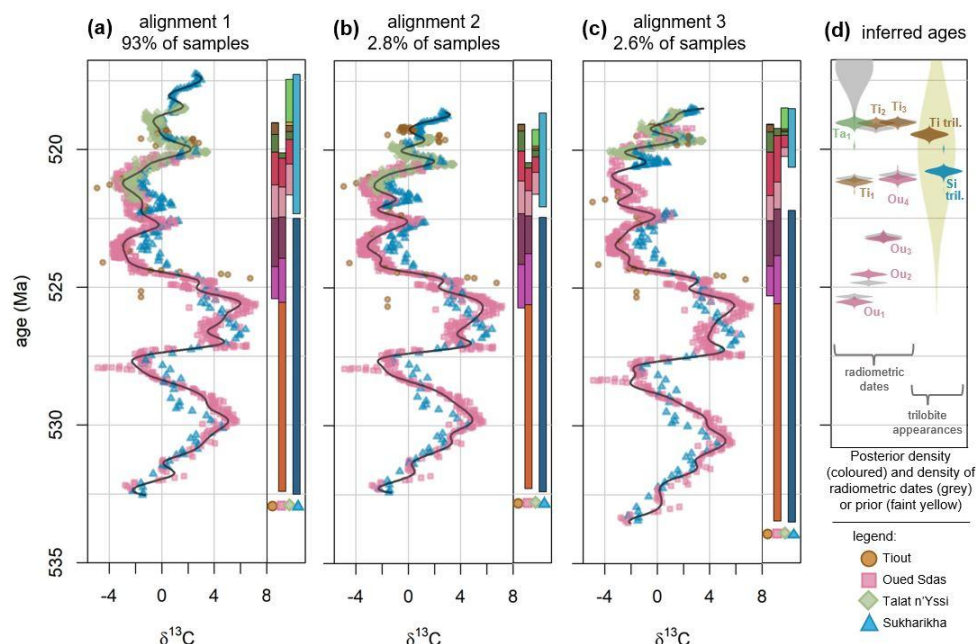


Figure 6: Three possible alignments identified by the inference with Cambrian data. (a) Exemplary sample from the cluster of the most likely alignment (93% of posterior samples). (b, c) Exemplary samples from a second and third identified alignment cluster (2.8% and 2.6% of posterior samples, respectively). Each shown alignment corresponds to a single sample from the posterior; other samples will result in slightly different alignments. 1.5% of samples were not assigned to any cluster (see Fig. 5). The curved dark lines show the cubic B-splines corresponding to each visualised sample. The coloured bars to the right of each alignment show the median duration of the stratigraphic partitions under each respective alignment cluster, based on the median ages of partition boundaries, with colours repeating the colour scheme of Fig. 2. (d) Posterior density of the inferred ages corresponding to the



radiometric dates to the left (3 from Tiout, 4 from Oued Sdas, and 1 from Talat n'Yissi) and the first occurrences of trilobites at Tiout (Ti tril.) and Sukharikha (Sh tril.) to the right, in colours. All samples from all alignment clusters were included. Greater width corresponds to higher posterior density; all densities are scaled to have the same maximum for better visibility. Densities representing the uncertainties of radiometric dates based on their mean and standard deviation are shown in grey (left). The faint yellow shading to the right shows the prior density on α , i.e. the first appearance of trilobites at Tiout and Siberia based on a mean age of (520 Ma) and a standard deviation of 2 Myr (identical for Tiout and Siberia). Colours and shapes of the points correspond to the four sites: Tiout - brown circles; Oued Sdas - pink squares; Talat n'Yissi - green diamonds; Sukharikha - blue triangles.

456 The posterior of the model runs allows the construction of age models that span the entire height of each
457 site (Fig. 7). As sedimentation rates are constrained to be constant within the pre-defined partitions,
458 sedimentation rate changes are visible as inflections at the boundaries of these partitions. Age uncertainties
459 are relatively low at Tiout and most of Oued Sdas, which are relatively well constrained by radiometric
460 dates in the top (Tiout) and middle (Oued Sdas) parts of the sections, as well as by astronomical priors on
461 sedimentation rates. Uncertainty noticeably increases towards the top and bottom of Oued Sdas. The lowest
462 partition of Oued Sdas is constrained only by its match to the lower part of the Sukharikha Fm., their age
463 estimates are thus varying considerably between different alignments (Fig. 6). Differences in the positioning
464 of the $\delta^{13}\text{C}$ curves between alignments are greatest at Talat n'Yissi and the Siberian Krasnoporog Fm. (Fig.
465 6), which results in large uncertainties in the age models (Fig. 7c, d).

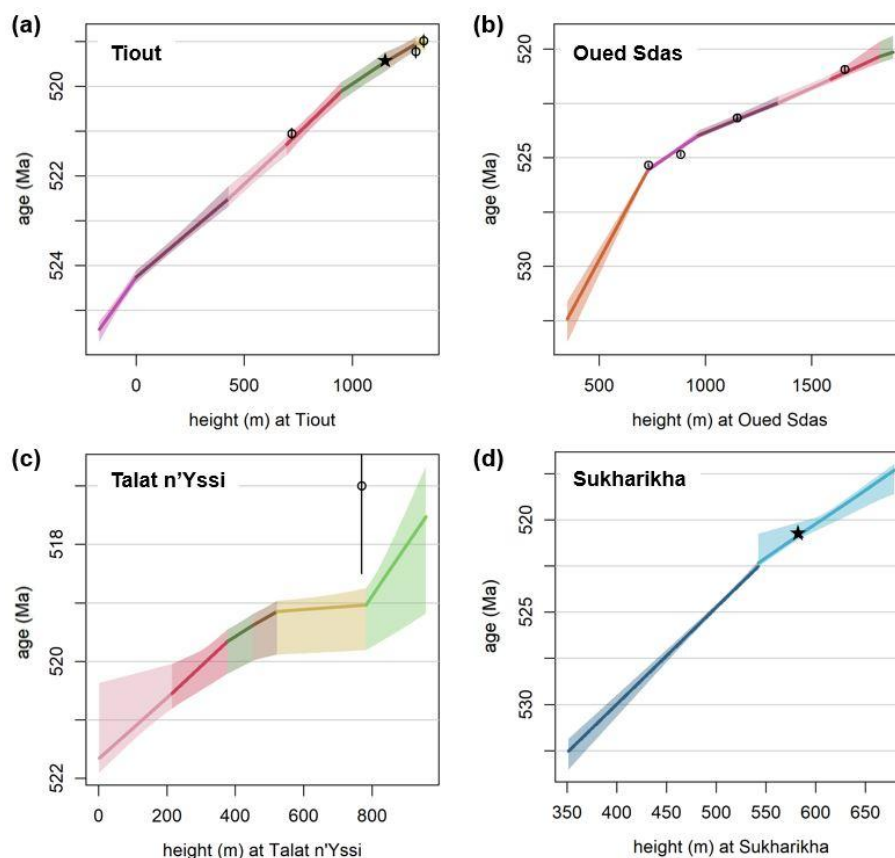


Figure 7: Age-depth model for each of the four sites. The solid lines indicate the median posterior ages corresponding to the respective heights; the shaded interval denotes the 95% credible interval of posterior ages. Colours correspond to the colours of partitions introduced in Fig. 3. Circles indicate the mean age estimates of radiometric dates, with vertical lines spanning two standard deviations around the mean of these age estimates. Stars denote the first appearances of trilobites in Morocco and Siberia.

466 5 Discussion

467 5.1 Lower Cambrian stratigraphy

468 We used StratoBayes to correlate and date four lower Cambrian carbonate sections using $\delta^{13}\text{C}$ records,
 469 radiometric dates and astrochronological sedimentation rate estimates. From a large space of possible
 470 alignment configurations (Fig. 4), the software identified alignment solutions that visibly match the large-



471 scale features in the $\delta^{13}\text{C}$ records from multiple sites, while simultaneously achieving an approximate fit to
472 the radiometric dates (Fig. 6).

473 The most likely alignment solution from the posterior, alignment 1 (probability = 93%), results in a
474 correlation of the three Moroccan sites that has much in common with that proposed by Maloof et al. (2005).
475 In our model, we used common sedimentation rates for the stratigraphic partitions (members, formations)
476 shared between the sites, whilst allowing sedimentation rates to systematically differ from the reference
477 sedimentation rates at Tiout by adding a site-specific multiplier. This multiplier, ζ , is 1.02 (95% credible
478 interval: 0.97 to 1.08) for Oued Sdas, meaning the model estimates very similar sedimentation rates for
479 Tiout and Oued Sdas (Fig. 6a), consistent with their close geographical proximity. Sedimentation rates for
480 the shared partitions at Talat n'Yissi are lower by a factor of 0.86 (0.76 to 0.96), which would be consistent
481 with a moderately lower accommodation space at Talat n'Yissi relative to Tiout and Oued Sdas (as
482 suggested by Fig. 3B in Maloof et al., 2005). We deliberately chose broad priors that did not explicitly
483 enforce a relationship between sedimentation rates and palaeogeography; nonetheless, the model identified
484 a geologically plausible solution. In contrast, the higher $\zeta_{\text{Talat n'Yissi}}$ of alignment 2 (probability = 2.8%,
485 1.07 to 1.37) and alignment 3 (probability = 2.6%, 2.07 to 2.45) are harder to reconcile with the
486 palaeogeographic context.

487 Alignments 2 and 3 also suggest different sedimentation rates between Tiout and Oued Sdas, with a higher
488 value of $\zeta_{\text{Oued Sdas}}$ (1.13 to 1.26) being estimated by alignment 2, and a lower value of $\zeta_{\text{Oued Sdas}}$ (0.83 to
489 0.88) by alignment 3. The most consistent lithostratigraphic alignment between Tiout and Oued Sdas is
490 achieved by alignment 1, meaning that the age estimates for partition boundaries (based on members or
491 formations) are most similar (Fig. 6). For the more distant Talat n'Yissi, age estimates of partition
492 boundaries differ to varying degrees across all three alignments.

493 Breaking down the posterior probability into individual components – likelihood (fit of $\delta^{13}\text{C}$ measurements
494 to the spline, fit of age estimates to the radiometric dates) and prior probability from the overlap penalty –
495 reveals that samples from alignment 1 have a higher likelihood, on average (Fig. 5c). In contrast, alignments
496 2 and 3 have a greater number of overlapping $\delta^{13}\text{C}$ points, which results in higher overlap prior probabilities
497 (Fig. 5c). The overlap prior reflects the prior belief that substantial parts of the sections involved in the
498 correlation should be overlapping. However, the weight of that prior is somewhat arbitrary and reflects the
499 technical requirement to facilitate overlap despite non-overlap allowing for closer fit to the spline, similar
500 to the role of the “edge value” in some DTW implementations (Hay et al., 2019). A lower prior weight on
501 overlap would thus have caused alignments 2 and 3 to receive lower posterior probabilities relative to



502 alignment 1. Taken together, the evidence from above leads us to strongly favour alignment 1, and we will
503 focus further discussion on that most likely alignment solution.

504 A radiometric date of 517.0 Ma (± 2 SD: 515.5 – 518.5 Ma) has been recovered from the Lemdad
505 Syncline in the Atlas mountains (Landing et al., 1998), and has been correlated biostratigraphically to a
506 horizon in the lower Isaafen Fm. at Talat n'Yssi (Maloof et al., 2005). In our alignment 1, this horizon has
507 a posterior age estimate of 519 Ma (519.2 to 518.8 Ma) – ≈ 2 Myr older than the mean of the radiometric
508 date. This date has informed the age estimates for Talat n'Yssi in Maloof et al. (2005) and Maloof et al.
509 (2010), whereas alignment 1 produces age estimates close to those of Bowyer et al. (2022) and Bowyer et
510 al. (2023). Age estimates deviating from radiometric dates are not necessarily incorrect: Although
511 radiometric dates are sometimes treated as “absolute truth” within the stratigraphic community, they are
512 the result of various sources of technical uncertainties (Condon et al., 2024) and geological interpretations
513 like the actual zircon crystallisation versus eruption age (Keller et al., 2018). This is illustrated by the
514 recalculation of the radiometric date from Landing et al. (1998) to 515.56 Ma (± 2 SD: 514.40 –
515 516.72 Ma) in the Geological Time Scale 2012 (Schmitz et al., 2012).

516 The two radiometric dates measured at Tiout at the bottom of and within the Amouslek Formation suggest
517 a sedimentation rate of 146 m Myr⁻¹ (± 2 SD: 78.7 to 613 m Myr⁻¹) for the Amouslek formation. However,
518 the posterior estimates for the sedimentation rate in the Amouslek formation are poorly constrained and
519 high compared to the sedimentation rates of all other partitions, at 3030 m Myr⁻¹ (800 to 17,300 m Myr⁻¹).
520 It appears that the model has overestimated the Amouslek sedimentation rate in aligning the $\delta^{13}\text{C}$ record of
521 the overlying Isaafen formation with a part of the Siberian Krasnoporog formation which has similar $\delta^{13}\text{C}$
522 values (Fig. 6a). The alignments of Bowyer et al. (2022) imply significant sedimentation rate changes
523 within the Krasnoporog formation, allowing the $\delta^{13}\text{C}$ records to be better reconciled with the radiometric
524 dates. We didn't allow for sedimentation rate changes within the Krasnoporog formation because the
525 stratigraphic log of Kouchinsky et al. (2007) indicates a uniform facies. Additional sedimentation rate
526 changes might lead to a closer alignment with the radiometric dates, at the cost of greater model complexity.

527 The alignment of the Siberian Sukharikha section with the Moroccan sites is relatively precise in the lower
528 half of the records: The prominent positive $\delta^{13}\text{C}$ excursions interpreted as the “5p” and “6p” excursions
529 have a similar magnitude both at Oued Sdas and Sukharikha, and are readily aligned visually (Bowyer et
530 al., 2022) and by our model (Fig. 6). Our model aligns the main 6p peak of Sukharikha with the first subpeak
531 of the second large excursion at Oued Sdas, as in model C in Bowyer et al. (2022). The lesser, positive
532 excursion below the hiatus at the top of the Sukharikha formation lines up with the positive excursion in
533 the lower Lie-de-Vin formation, representing the “II” peak as in model C in Bowyer et al. (2022). The upper



534 parts of the Moroccan records and the Siberian Krasnoporog formation appear to be aligned primarily by
535 matching the prominent positive excursion interpreted as excursion “IV” (Bowyer et al., 2022; Kouchinsky
536 et al., 2007). The “III” peak below is only weakly expressed at Oued Sdas, leading to uncertainty in the
537 alignment with the corresponding part of the Krasnoporog formation, and in the inferred duration of the
538 hiatus even within alignment solution 1 (Fig. B2a-c). Similarly, considerable uncertainty exists in how the
539 top of Talat n’Yissi corresponds to the Krasnoporog formation. This is evident from variations between
540 samples in alignment solution 1 (Fig. B2a-c) and in the wide credible intervals of those parts of the age
541 models (Fig. 7). The relatively small magnitude of $\delta^{13}\text{C}$ changes limits the model’s ability to identify a
542 definitive alignment solution for that part of the record.

543 Our estimate for the Moroccan first appearance of trilobites at Tiout from alignment 1, 519.47 Ma (519.68
544 to 519.26 Ma), is slightly younger and somewhat less precise than the recent, astrochronological estimate
545 of 519.62 Ma (95% highest posterior distribution: 519.70 to 519.54 Ma) by Sinnesael et al. (2024). We
546 attribute this difference to our model simultaneously combining different data types from multiple sites.
547 Additionally, Sinnesael et al. (2024) allowed sedimentation rates to vary between cycles, whereas our model
548 assumed a single sedimentation rate per member. In our alignment 1 solution, the highest $\delta^{13}\text{C}$ values of
549 Tiout correlate to shortly after the peak of the IV $\delta^{13}\text{C}$ excursion. This correlation suggests that the actual
550 peak of the excursion at Tiout has not been sampled by Magaritz et al. (1991) and Tucker (1986), which
551 may result in misalignments when correlating the record to other sections. Further $\delta^{13}\text{C}$ samples from the
552 lower Igoudine and upper Lie-de-Vin formation at Tiout are required to improve correlation with other
553 sections, including the correlation presented herein.

554 Our model successfully reconstructs the first appearance of trilobites at Tiout, within error, despite using a
555 simpler astrochronology and enforcing a less variable sedimentation rate history than Sinnesael et al.
556 (2024). It also provides the first fully quantitative estimate for the first appearance of trilobites in Siberia
557 based on chemostratigraphic correlation and the Moroccan radiometric dates and astrochronology, at
558 520.79 Ma (520.98 to 520.61 Ma). This refines earlier estimates of ≈ 521 Ma (Landing et al., 2021), and
559 quantifies the temporal gap between the appearance of trilobites in Siberia and Morocco as 1.33 Myr (1.09
560 to 1.54 Myr). We do not suggest that these estimates are definitive; indeed, we anticipate that the
561 incorporation of additional $\delta^{13}\text{C}$ data from Tiout, the inclusion of astrochronological estimates of individual
562 short eccentricity cycles, and the relaxation of the assumption of constant sedimentation rates within
563 partitions may update the estimate. A high-resolution temporal sequence of trilobite first occurrence dates
564 could be used to delineate trilobite evolutionary rates and dispersal; to evaluate evolutionary hypotheses on
565 the origins and biomineralisation of trilobites (Holmes and Budd, 2022; Paterson et al., 2019); and to inform
566 the definition of the base of the Cambrian Series 2 (Zhang et al., 2017).



567 **5.2 Statistical alignment and age modelling**

568 **5.2.1 Advantages of Bayesian stratigraphic alignment**

569 As shown above, our algorithm can identify the correct alignment positions in scenarios with one (Fig. 1)
570 or more than one (Fig. 2) known solution. In scenarios where more than one distinctly different alignment
571 is identified, the probability of each solution, given the specified data and model, is identified. This can be
572 used to evaluate the likelihood of competing models for the alignment of stratigraphic records found by
573 visual (e.g. Bowyer et al., 2023; Landing and Kruse, 2017) or algorithmic (e.g. Hay et al., 2019) correlation.
574 The requirement to specify priors for the alignment parameters can be leveraged to provide information
575 beyond that which is contained in the signals: for example, information on sedimentation rates may be
576 expressed in the prior.

577 Because our model can integrate absolute age constraints such as radiometric dates, a user is able to
578 correlate stratigraphic records and construct probabilistic age models in a single step. In our Cambrian
579 example, the posterior alignment and the posterior age model are thus influenced by the priors, the
580 quantitative signals and the radiometric dates. In contrast, age models constructed in a separate step after
581 identifying alignments do not reflect uncertainty arising during the alignment stage (Hagen and Creveling,
582 2024).

583 In our integrated approach, discrepancies between radiometric dates and signal alignment are resolved
584 probabilistically, with the model weighting the available evidence based on its likelihood and prior
585 information. This means that posterior age estimates may diverge from the age information provided by
586 radiometric dates, as seen with the Ou_2 date in Fig. 6d. This is not necessarily a deficiency of the model;
587 rather, it indicates that the priors and non-radiometric data provide sufficiently strong evidence to suggest
588 that the actual age of the horizon associated with the radiometric date falls toward the tails of its confidence
589 interval, or that the radiometric uncertainty may be underestimated. Some degree of discrepancy is expected
590 when integrating multiple data types rather than relying on a single proxy (see also Lee et al., 2022).

591 If, on the other hand, the user wishes to increase the influence of radiometric dates on the posterior age
592 estimates, this can potentially be achieved by introducing additional sedimentation rate changes to allow
593 more flexible alignment of the proxy signals, reducing the weight of the proxy signal records – such as by
594 imposing a larger σ for the cubic spline – or by weakening priors.



5.2.2 Model choice and priors

Stratigraphic alignment using algorithms has the advantage of removing some of the inherent subjectivity of visual alignment (Sylvester, 2023). Yet, somewhat subjective decisions are still explicitly or implicitly made with every alignment algorithm. In the case of DTW, subjectivity is introduced e.g. with restrictions on the warping path (i.e. relative sedimentation rates, Sakoe and Chiba, 1978), with the amount of overlap required between sections (Hay et al., 2019), or with the choice of an exponent controlling the weight of outlier values (Wheeler and Hale, 2014). All of those settings can alter the outcome of DTW-based alignments. Likewise, our Bayesian approach comes with a number of subjective choices. The appropriate model structure can be readily determined when the data-generating process is known (Sect. 3), but has to be carefully considered and potentially revised when dealing with complex real-world data (Sect. 4). Lithological data may guide the partitioning of data and can inform somewhat objective choices of horizons with likely sedimentation rate changes (Sect. 4.2), but such information may not be readily available with some datasets, such as with well logs.

Besides the model structure, StratoBayes requires the user to specify priors for several model parameters: relative or absolute sedimentation rates (γ, ζ), the shifts of sections relative to one another (α), the duration of hiatuses (δ), the degree of smoothing of the spline (λ), the extent to which overlap of signal points should be favoured ($C_{overlap}$), and optionally the residual standard deviation of the spline (σ). Although the choice of any of those parameters has the potential to affect posterior alignments and age models, they also offer a chance to explicitly include geological information that could otherwise only be incorporated by discarding or modifying alignment solutions after the algorithmic alignment.

5.2.3 Challenges with the $\delta^{13}\text{C}$ and sedimentary record

Chemostratigraphy, and, more broadly, correlating geological sections based on proxy data relies on the proxies accurately reflecting a common, underlying signal. Several processes may disrupt this assumption. For example, $\delta^{13}\text{C}$ recorded in carbonates differs between different depositional environments, water depths, and grain types (Geyman and Maloof, 2021), while the $\delta^{13}\text{C}$ recorded in restricted basins may be offset significantly relative to contemporary carbonates elsewhere (Uhlein et al., 2019). Where known, such offsets could be accounted for by subtracting or adding the estimated offset relative to global values. Alternatively, anticipated offsets could be modelled as additional unknown variables, as in Edmonson and Dyer (2024). This approach will likely require substantial prior knowledge on the potential magnitude and direction of offsets; otherwise, the combination of variation along the height or time axis and along the proxy value axis may result in a large range of mathematically feasible alignments.



626 Several challenges arise from the variability of sedimentation and the incompleteness of the sedimentary
627 record. Sediment accumulation rates vary with measurement scale (Sadler, 1981): closer spacing between
628 measurements allows more variability to be identified, with actual sedimentation rate histories displaying
629 fractal properties (Miall, 2015). This implies that depositional ages tend to vary non-linearly along a
630 vertically sampled sedimentary section, with substantial incompleteness in shallow-water records (Curtis
631 et al., 2025). These discontinuities can lead to drastically altered shapes of proxy curves from different
632 depositional settings, and cycles from periodic proxy fluctuations may be missed due to insufficient
633 preservation or sampling (Curtis et al., 2025). This issue is evident in the Sukharikha section, where it is
634 somewhat ambiguous whether the hiatus represents a fraction of a $\delta^{13}\text{C}$ excursion (alignment 1 and 2) or
635 extends over more than one full cycle (alignment 3, Fig. 6). For correlations within sedimentary basins, the
636 method of Bloem and Curtis (2024) could help resolve ambiguous alignments by reconstructing
637 depositional histories through geological process modelling, but this method requires exceptionally high-
638 resolution sampling and its utility has yet to be demonstrated with real-world data sets.

639 Besides the completeness, the sampling density of proxy records may influence correlations. In
640 StratoBayes, densely sampled sections or parts of sections exert more influence on the shape of the spline
641 than those that are thinly sampled, which can be seen in the spline curve primarily following the densely
642 sampled Oued Sdas and Talat n'Yissi records in Fig. 6. Despite this, our Cambrian case study demonstrates
643 that sections with differing sampling densities – both between and within sites – can still be effectively
644 aligned. Varying sampling density would, however, pose a challenge for reconstructing a global average
645 proxy curve from local records, as the global curve would primarily reflect the more densely sampled sites.

646 StratoBayes introduces a simplification in modelling sedimentary histories by forcing uniform
647 sedimentation rates within pre-defined segments of a stratigraphic section. An effect of this simplification
648 can be seen in the age-depth plots in Fig. 7: Due to sedimentation rates being modelled as uniform within
649 stratigraphic partitions, the uncertainty of age estimates does not necessarily decrease away from the
650 radiometric dates. We acknowledge that this may underestimate the uncertainty associated with potential
651 sedimentation rate variability (De Vleeschouwer and Parnell, 2014), especially when allowing for few
652 sedimentation rate changes.

653 In principle, our method could be used to divide stratigraphic sections into an arbitrary number of segments
654 with differing sedimentation rates. In practice, estimating the parameters of a model with more than a low
655 double-digit number of alignment parameters (shift parameters, sedimentation rates, hiatuses) represents a
656 challenge for the current implementation of the MCMC algorithm within StratoBayes, as finding and
657 exploring the posterior becomes increasingly difficult as more parameters are added. This limitation could
658 be alleviated by incorporating MCMC methods suited for higher dimensional problems and difficult



posterior geometries. Alternatively, a continuous process model such as the compound Poisson-gamma process of BChron (Haslett and Parnell, 2008) might be integrated with our model for the proxy signal, but again the complexity of the MCMC would increase. Another approach would be to divide the alignment problem into sub-problems, e.g. by multiple pairwise correlation of sites (e.g. Hagen et al., 2024; Sylvester, 2023), or by correlating shorter subsections.

5.3 Towards quantitative stratigraphy

Quantitative stratigraphic correlation and age modelling of diverse geological data represent a long-standing challenge in stratigraphic research. Although many algorithms exist for correlating stratigraphic data (e.g. Baille et al., 2022; Bloem and Curtis, 2024; Hay et al., 2019; Sylvester, 2023); few can readily provide uncertainty estimates or incorporate different types of data simultaneously (e.g. Al Ibrahim, 2022; Edmondson and Dyer, 2024; Lee et al., 2022). Consequently, integrated statistical approaches have only rarely been applied to complex real-world stratigraphic problems (Hagen et al., 2024; Lee et al., 2022).

Our new method has the potential to be applied to diverse datasets; examples range from shallow borehole data from the Holocene (Finlay et al., 2022) to Proterozoic carbonates (Halverson et al., 2010). The ability of our model to incorporate multiple proxy records simultaneously opens new possibilities for refining stratigraphic correlations. For instance, correlations involving both $\delta^{13}\text{C}$ and $\delta^{87}\text{Sr}$ records could benefit from a probabilistic framework that accounts for their respective uncertainties (Bowyer et al., 2022). The integration of multiple proxies, e.g. multiple element ratios, in the StratoBayes framework could allow correlations based on the entire record of all proxies, rather than a few visually distinct transitions (Craigie, 2015).

Beyond geochemical records, our approach could also be applied e.g. to geophysical well-logs such as gamma ray or density logs, and magnetostratigraphic records could be correlated directly rather than relying on visually interpreted polarity reversals (Langereis et al., 2010). While index fossils can currently be integrated as tie points, the modelling framework could be expanded to explicitly model first and last occurrences to better incorporate biostratigraphic uncertainty. Similarly, astrochronological constraints can be expressed as priors on sedimentation rates, but an additional model component would be needed to incorporate all astrochronological information from a given site (Sinnesael et al., 2024).

Conclusions

StratoBayes is a Bayesian modelling framework for the probabilistic alignment of stratigraphic proxy records and age modelling. It correlates quantitative proxy signals such as isotope ratios, and integrates



689 additional stratigraphic information such radiometric dates, to construct probabilistic age models. Applying
690 our model to both simulated data and real-world stratigraphic records from the lower Cambrian of Morocco
691 and Siberia, we have demonstrated its ability to account for uncertainty from all model components and to
692 identify multiple plausible alignment solutions. Our lower Cambrian case study provides a fully
693 probabilistic estimate for the first appearance of trilobites in Siberia, and quantifies the temporal gap
694 between their first occurrence and the oldest Moroccan trilobites. While our results remain dependent on
695 model assumptions, they represent a step towards a more objective and reproducible approach to early
696 Palaeozoic stratigraphy; they also highlight sources of uncertainty and identify targets for future research.
697 Beyond this case study, StratoBayes has broad applicability to stratigraphic problems across all time
698 intervals that involve the correlation of quantitative proxy records.

699 **Appendix A: Markov chain Monte Carlo sampling scheme**

700 Appendix A details the Metropolis-within-Gibbs sampling scheme and the parallel tempering framework
701 that are used within the StratoBayes software to sample from the posterior of the unknown model
702 parameters.

703 The MCMC sampling includes an adaptive phase. During this phase, proposal distributions and the
704 probabilities with which different proposal types are selected for the Metropolis-Hastings updates are
705 adjusted based on the history of the MCMC chains to improve acceptance rates and mixing. Additionally,
706 the temperature ladder of the parallel tempering framework is updated to improve the swap rates of chains.
707 After the adaptive phase, the proposal distributions and probabilities, as well as chain temperatures, remain
708 fixed for the remainder of the run to ensure proper sampling from the posterior.

709 In the current implementation, the length of the adaptive phase is pre-determined by the user, specified as
710 a fixed number of iterations. However, the user has the option to extend the adaptation period by continuing
711 the run if needed. More generally, adaptation could also be stopped automatically based on criteria such as
712 mixing within chains (Yang and Rosenthal, 2017) or convergence criteria.

713 Adaptive MCMC algorithms do not always preserve the stationarity of the target distribution during the
714 adaptive phase (Roberts and Rosenthal, 2009). Therefore, all samples from the adaptive phase are discarded
715 as burn-in. Additionally, if diagnostic checks suggest that the MCMC has not converged by the end of the
716 adaptive phase, further samples may need to be discarded.



717 **Gibbs sampling scheme for the cubic B-splines**

718 The following sampling scheme was adapted from Heaton et al. (2020). The spline coefficients are sampled
719 from a multivariate normal distribution of the form:

$$720 \quad \beta \sim MVN(b\mathbf{Q}, \mathbf{Q}), \quad (16)$$

721 where b is given by .

$$722 \quad b = (\mathbf{B}(h))^T \frac{\mathbf{y}}{\sigma^2}, \quad (17)$$

723 $\mathbf{B}(h)$ are cubic B-splines (Eilers and Marx, 1996) at a set of k knots evaluated at heights h at which $\mathbf{y}$, the
724 composite stratigraphic signal of all sites, was observed. Here, σ is the residual standard deviation.

725 The other element needed for sampling from the posterior of b is $\mathbf{Q}$, given by

$$726 \quad \mathbf{Q} = (\mathbf{H} + \lambda \mathbf{D})^{-1}, \quad (18)$$

727 where λ is a smoothing parameter, $\mathbf{D}$ is a penalty matrix to prevent the spline from overfitting the data, and

$$728 \quad \mathbf{H} = \left(\frac{\mathbf{B}(h)}{\sigma} \right)^T \frac{\mathbf{B}(h)}{\sigma} \quad (19)$$

729 The standard deviation σ can be fixed as

$$730 \quad \sigma = \frac{1}{S} \sum_{s=1}^S \sigma_s, \quad (20)$$

731 where S is the number of sites, and σ_s is the standard deviation of individual splines fitted to the data of site
732 s . This often provides a good approximation of σ , while removing an unknown model parameter, potentially
733 facilitating quicker convergence of the model run.

734 Alternatively, σ can be estimated within the Gibbs sampling scheme from the data, by placing a conjugate
735 gamma prior on the inverse of the variance (precision, $\tau = 1/\sigma^2$):

$$736 \quad \sigma^{-2} \sim \text{Gamma} \left(a_\sigma + \frac{n_y}{2}, b_\sigma + \frac{1}{2} \sum_{i=1}^{n_y} (y_i - \beta \mathbf{B}(h))^2 \right) \quad (21)$$

737 The smoothing parameter λ is estimated by placing a gamma prior on λ :



$$\lambda \sim \text{Gamma} \left(a_\lambda + \frac{k}{2}, \frac{1}{\frac{1}{b_\lambda} + \frac{1}{2} \sum^k \beta \mathbf{D} \times \beta} \right) \quad (22)$$

Metropolis-Hastings step

The starting heights or ages α , sedimentation rates γ , site multipliers ζ and gaps δ are updated in a Metropolis-Hastings step. For each unknown parameter, a new value is randomly sampled from a proposal distribution. Initially, proposals are sampled independently for each parameter from its respective prior, or alternatively from a custom proposal distribution.

In the following, the current set of parameter values is labelled θ , and the proposed set is labelled θ' . To decide whether to accept or reject the new set of parameters, an acceptance probability A is calculated, and the proposal is randomly accepted or rejected with a probability of A . This probability is calculated as

$$A = \min \left(1, \frac{\pi(\theta')}{\pi(\theta)} \right), \quad (23)$$

where $\pi(\theta)$ is the unnormalised posterior probability of the current values, and $\pi(\theta')$ is the unnormalised posterior probability of the proposed values. These can be calculated as

$$\pi(\theta) = p(\theta) \times L(\text{data}|\theta), \quad (24)$$

where $p(\theta)$ is the prior probability of θ , and $L(\text{data}|\theta)$ the likelihood of the data given θ .

We calculate the likelihood of the data given θ as a product of the probability densities of each data point of the signal y (recorded at two or more sites) and of all absolute age information. For the signal, we assume that the observed values y are normally distributed and centred around the values predicted by the splines, μ , at height h , with a standard deviation σ which has been introduced earlier. The likelihood of a data point i from the signal y is thus

$$L(y_i|\theta) = \frac{1}{\sqrt{2\pi\sigma^2}} \times e^{\left(\frac{(y_i - \mu_i)^2}{2\sigma^2} \right)} \quad (25)$$

and the log-likelihood for all data points of the signal is calculated as

$$\ln L(y|\theta) = \sum_i \ln L(y_i|\theta) \quad (26)$$



If more than one type of signal is used, the log-likelihood of additional signals can be calculated analogously and added in Equation 29.

Age constraints are incorporated by using an age estimate from radiometric dates d with, for example, mean ages a_{mean} and uncertainties given by standard deviations a_{sd} . The probability density of a date d_i is then calculated as

$$L(d_i|\theta) = \frac{1}{\sqrt{2\pi a_{sd,i}^2}} \times e^{\left(-\frac{a_{mean,i}-a_{predicted,i}}{2a_{sd,i}^2}\right)} \quad (27)$$

where $a_{predicted,i}$ is the age predicted by the age-height transform at the height $h_{s,i}$, the height at the site at which date d_i was obtained.

The log-likelihood for all age constraints is calculated as

$$\ln L(d|\theta) = \sum_i \ln L(d_i|\theta) \quad (28)$$

and the overall likelihood, if absolute age constraints are included, is

$$\ln L(y, d|\theta) = \ln L(y|\theta) + \ln L(d|\theta) \quad (29)$$

Proposal types

In order to allow for a broad search of the parameter space, proposals are initially selected independently for each parameter, and are selected independently of the current parameter values. These proposals lead to a decreasing acceptance rate over time, and the chain tends to arrive at a single set of values with high posterior probability, $\pi(\theta)$, remaining there for many iterations due to frequent rejections. Therefore, different types of proposals are used after an initial period:

- 1) Proposing from the prior or a custom distribution: This proposal is used exclusively for a small number of initial iterations and is alternated with other proposals later on.
- 2) Adaptive independent (univariate) proposals: Proposals for each parameter are selected independently from other parameter values. Proposals are dependent on the current state of the parameter θ_i , and sampled from a normal distribution $N(\theta_i, \sigma_i)$, where σ_i is a standard deviation that is estimated based on the history of the MCMC chain, i.e. based on the sampled θ_i from previous iterations.



785 3) Adaptive dependent (multivariate) proposals (Roberts and Rosenthal, 2009): Proposals for the
786 parameters are selected jointly and are dependent on the current state of the parameters θ .
787 Proposals are sampled from a multivariate normal distribution $MVN(\theta, \Sigma)$, where Σ is a
788 covariance matrix that is estimated based on the history of the MCMC chain, i.e. based on the
789 sampled θ_i from previous iterations.

790 4) Shifting some or all α and or δ parameters while keeping the other parameters constant. This can
791 accelerate the convergence of the MCMC in cases where some sites are aligned with each other,
792 but offset relative to other sites.

793 Proposal types are chosen with a probability that broadly corresponds to the relative acceptance probability
794 of the respective proposal type, i.e. proposal types that are rejected often are chosen less frequently.
795 Adaptation for types 2) and 3), and the adjustment of proposal type probabilities ends after the adaptive
796 phase. Posterior samples from the adaptive phase have to be discarded as burn-in, to ensure the correct
797 convergence of the chain.

798 **Parallel tempering**

799 To avoid the MCMC chain becoming trapped at isolated peaks of the posterior probability distribution, we
800 implement a parallel tempering framework, following Sambridge (2014). This involves running multiple
801 chains in parallel. The target chain, the chain from which the posterior samples will be taken, is left
802 unaltered (“cold chain”). The other chains are tempered, i.e. their unnormalised log posterior probabilities
803 are raised to the power of $1/T$, with T being the temperature. The higher T , the more “flattened” the posterior
804 probability landscape becomes, and the easier it is for the chain to explore the landscape. Frequently, chain
805 swaps are proposed, during which the model parameter values of different chains are exchanged with a
806 Metropolis-Hastings acceptance probability based on the ratios of posterior probabilities of the states of the
807 two chains, evaluated at both temperatures as in Appendix A2 of Sambridge (2014).

808 The initial temperatures for a number of chains n_c are selected using a geometric spacing, with $T_1 = 1$ (cold
809 chain) and $T_{n_c} = \infty$ (hottest chain). The infinite temperature of the hottest chain implies that all proposals
810 during the MCMC will be accepted, and we let that chain sample from the prior probability distributions of
811 the parameters. If $n_c > 2$, intermediate chain temperatures are selected as

812
$$T_c = 10^{\sum_{j=2}^c d_j}, \quad (30)$$

813 where



814
$$d_c = \frac{(n_c - 1)^{(2/3)}}{n_c - 2} + \frac{c - 1 - (n_c - 1)/2}{1.5 * n_c}, \quad c = 2 \dots n_c - 1 \quad (31)$$

815 This leads to the spacing of temperatures decreasing with increasing number of chains, and temperature
816 spacing is narrower for lower temperatures on the log scale. A small amount of white noise from a normal
817 distribution with zero mean and a standard deviation of $(5 \times n_c)^{-1}$ is added to each d_c to vary the initial
818 temperature ladders between independent model runs. Temperatures are updated in the adaptive phase of
819 the MCMC to increase the swap rates of chains (Vousden et al., 2016).

820 **Appendix B: Inspecting the posterior of the lower Cambrian case**

821 **study**

822 Appendix B provides additional details on the posterior of the inference with lower Cambrian $\delta^{13}\text{C}$ data and
823 radiometric dates.

824 **Trace plots**

825 Trace plots visualise the evolution of chains from an MCMC and, together with tools such as the potential
826 scale reduction factor (Gelman and Rubin, 1992; Vats and Knudson, 2021), allow for assessing convergence
827 of model runs. The trace plot indicative of a well-behaved model run should be stationary after the burn-in
828 phase, with different chains mixing well (Gelman et al., 1995). An example of a well-behaved trace plot is
829 the first panel of Fig. B1. Inspecting the trace plots of the 18 model parameters of the lower Cambrian case
830 study reveals that all parameters seem to have reached stationarity, this said; some chains occasionally visit
831 distinctly different values (e.g. Fig. B1, column 1, row 2). The chains are not mixing well in those regions
832 of the parameter space. Running the model for considerably more iterations is likely to overcome this
833 problem. However, this affects only the less likely alignments; the most likely alignment (alignment cluster
834 1) is well explored across all parameters.

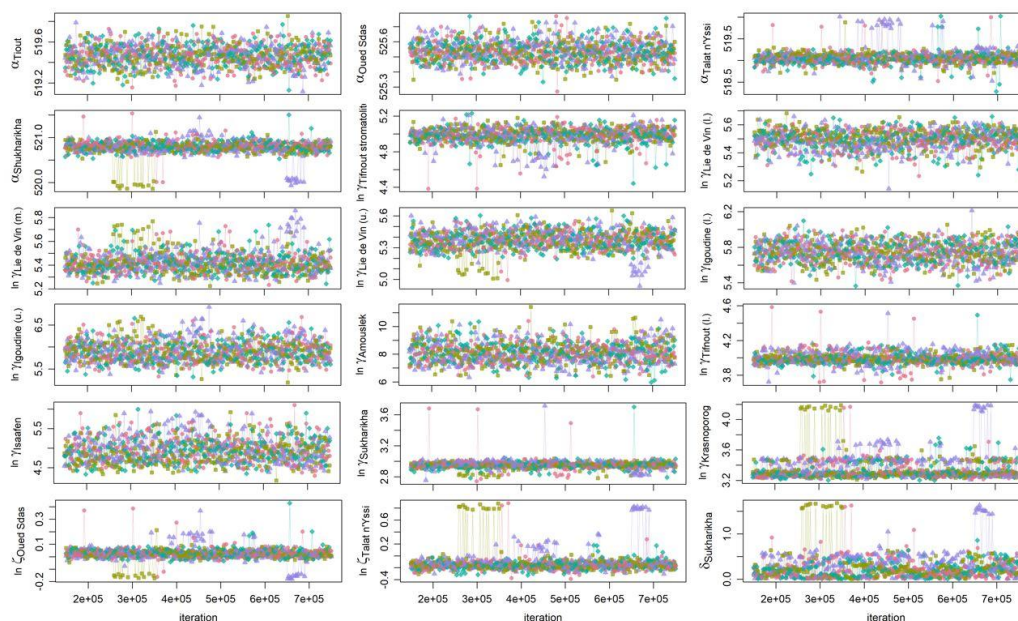


Figure B1: Trace plots of the 18 alignment parameters. Each colour corresponds to a distinct run. For visual clarity, only 250 samples are displayed per run. The burn-in phase (the first 150,000 iterations) is omitted.

Alignment clusters

Summarising the posterior by grouping samples into clusters of similar alignments facilitates discussion of the results but risks oversimplifying the variation within each cluster. Each cluster represents a set of posterior samples that share similar inferred ages for the partition boundaries, but differences still exist between individual samples within the same cluster. As an example, three distinct alignments from cluster 1 are visualised in Fig. B2. An alignment from a sample not assigned to any cluster is shown in Fig. B2d.

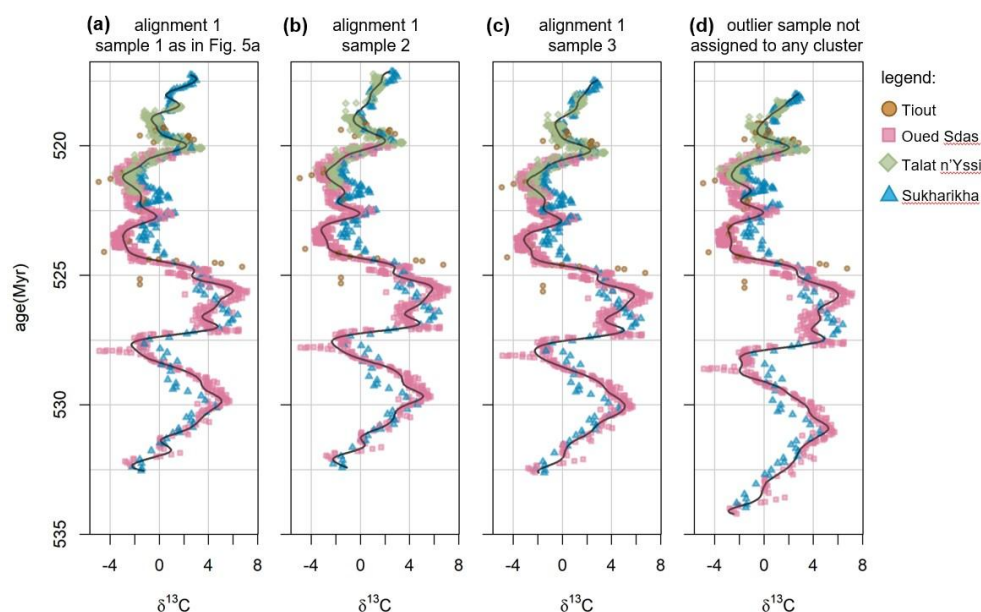


Figure B2: Alternative alignments, each corresponding to a single sample from the posterior. (a) A sample from the most likely cluster 1, corresponding to that shown in Fig. 6a. (b, c) Alignments corresponding to other samples from cluster 1. (d) Alignment corresponding to an outlier sample that was not assigned to any cluster. The curved dark lines show the cubic B-splines corresponding to each alignment.

841 Posterior of alignment parameters

842 The posterior distributions of the alignment parameters are summarised in histograms in Fig. B3.

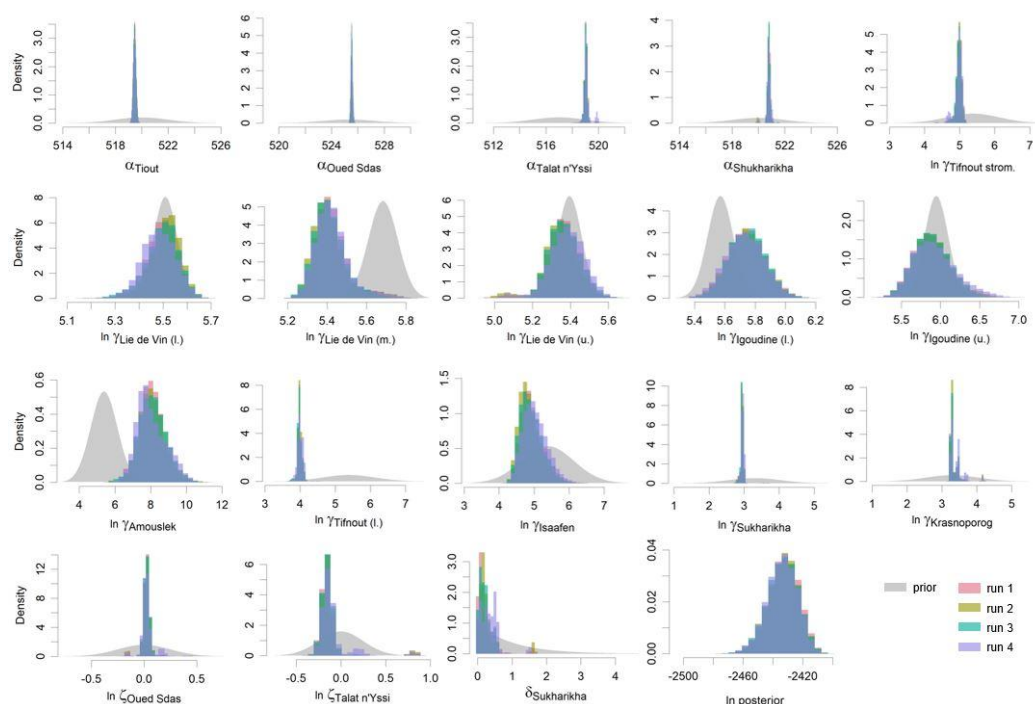


Figure B3: Comparison of prior and posterior probability densities. Histograms in colour denote the posterior probability densities of the 18 alignment parameters; the grey, smooth shadings represent prior probability densities. The four colours correspond to the four independent model runs.

843 Code and data availability

844 The StratoBayes R package can be installed from binaries by following the instructions at
845 <https://stratobayes.github.io>. The data and code used to generate the results, and copies of the software
846 binaries, are available at <https://github.com/StratoBayes/StratoBayes-Manuscript> and the linked Zenodo
847 archive <https://zenodo.org/records/15065336>.

848 Author contribution

849 MRS and ARM conceived the idea. MRS, ARM and MS designed the initial model version, and all authors
850 contributed to the design of the final model. MS, ARM and MRS coded the software prototype. KE, MRS
851 and ARM modified and expanded the software. MS curated the lower Cambrian data. KE conducted the



852 analyses and visualised the results. KE wrote the initial manuscript draft, and all authors contributed to the
853 final version. MRS coordinated the project. MRS and ARM acquired the funding.

854 **Competing interests**

855 The authors declare that they have no conflict of interest.

856 **Acknowledgements**

857 We thank the participants of the StratoBayes workshop (Durham, 2024) for feedback on an early version
858 of the StratoBayes software. We thank Tim Heaton for sharing his code for spline fitting. We thank Andrew
859 Valentine for discussions on the model. We thank Prof. Nasrddine Youbi and Kamal Mghazli for hosting
860 an exploratory field visit to the lower Cambrian sections in Morocco.

861 **Financial support**

862 This research was funded by Leverhulme Research Project Grant 2019-223 to MRS.

863 **References**

- 864 Agterberg, F. P.: Automated stratigraphic correlation, Elsevier, 1990.
- 865 Al Ibrahim, M. A.: Uncertainty in Automated Well-Log Correlation Using Stochastic Dynamic Time
866 Warping, *Petrophysics*, 63, 748–761, 2022.
- 867 Baviile, P., Apel, M., Hoth, S., Knaust, D., Antoine, C., Carpentier, C., and Caumon, G.: Computer-
868 assisted stochastic multi-well correlation: Sedimentary facies versus well distality, *Marine and Petroleum*
869 *Geology*, 135, 105371, 2022.
- 870 Bloem, H. and Curtis, A.: Bayesian geochemical correlation and tomography, *Scientific Reports*, 14,
871 9266, 2024.
- 872 Bowyer, F. T., Zhuravlev, A. Y., Wood, R., Shields, G. A., Zhou, Y., Curtis, A., Poulton, S. W., Condon,
873 D. J., Yang, C., and Zhu, M.: Calibrating the temporal and spatial dynamics of the Ediacaran-Cambrian
874 radiation of animals, *Earth-Science Reviews*, 225, 103913, 2022.



- 875 Bowyer, F. T., Zhuravlev, A. Y., Wood, R., Zhao, F., Sukhov, S. S., Alexander, R. D., Poulton, S. W.,
876 and Zhu, M.: Implications of an integrated late Ediacaran to early Cambrian stratigraphy of the Siberian
877 Platform, Russia, *Bulletin*, 135, 2428–2450, 2023.
- 878 Campello, R. J., Moulavi, D., Zimek, A., and Sander, J.: Hierarchical density estimates for data
879 clustering, visualization, and outlier detection, *ACM Transactions on Knowledge Discovery from Data*
880 (TKDD), 10, 1–51, 2015.
- 881 Condon, D., Schoene, B., Schmitz, M., Schaltegger, U., Ickert, R. B., Amelin, Y., Augland, L. E.,
882 Chamberlain, K. R., Coleman, D. S., Connelly, J. N., et al.: Recommendations for the reporting and
883 interpretation of isotope dilution U-Pb geochronological information, *Geological Society of America*
884 *Bulletin*, 2024.
- 885 Craigie, N. W.: Applications of chemostratigraphy in Middle Jurassic unconventional reservoirs in
886 eastern Saudi Arabia, *GeoArabia*, 20, 79–110, 2015.
- 887 Cramer, B. and Jarvis, I.: Carbon isotope stratigraphy, in: *Geologic time scale 2020*, Elsevier, 309–343,
888 2020.
- 889 Curtis, A., Bloem, H., Wood, R., Bowyer, F., Shields, G. A., Zhou, Y., Yilales, M., and Tetzlaff, D.:
890 Natural sampling and aliasing of marine geochemical signals, *Scientific Reports*, 15, 760, 2025.
- 891 De Vleeschouwer, D. and Parnell, A. C.: Reducing time-scale uncertainty for the Devonian by integrating
892 astrochronology and Bayesian statistics, *Geology*, 42, 491–494, 2014.
- 893 Edmonson, S. and Dyer, B.: A Bayesian framework for inferring regional and global change from
894 stratigraphic proxy records (StratMC v1. 0), *EGUsphere*, 2024, 1–45, 2024.
- 895 Eilers, P. H. and Marx, B. D.: Flexible smoothing with B-splines and penalties, *Statistical science*, 11,
896 89–121, 1996.
- 897 Finlay, A., Bates, R., Bensharada, M., and Davies, S.: Applying chemostratigraphic techniques to shallow
898 bore holes: Lessons and case studies from Europe’s lost Frontiers, in: *Europe’s Lost Frontiers: Volume 1*
899 *Context and Methodology*, Archaeopress, 137–153, 2022.
- 900 Gelman, A. and Rubin, D. B.: Inference from iterative simulation using multiple sequences, *Statistical*
901 *science*, 7, 457–472, 1992.



- 902 Gelman, A., Carlin, J. B., Stern, H. S., and Rubin, D. B.: Bayesian data analysis, Chapman; Hall/CRC,
903 1995.
- 904 Geyman, E. C. and Maloof, A. C.: Facies control on carbonate $\delta^{13}\text{C}$ on the Great Bahama Bank,
905 *Geology*, 49, 1049–1054, 2021.
- 906 Hagen, C., Creveling, J., and Huybers, P.: Align: A User-Friendly App for Numerical Stratigraphic
907 Correlation, *GSA Today*, 34, 4–9, 2024.
- 908 Hagen, C. J. and Creveling, J. R.: An algorithm-guided Ediacaran global composite $\delta^{13}\text{C}_{\text{carb}}$ –Bayesian
909 age model, *Palaeogeography, Palaeoclimatology, Palaeoecology*, 112321, 2024.
- 910 Halverson, G. P., Wade, B. P., Hurtgen, M. T., and Barovich, K. M.: Neoproterozoic chemostratigraphy,
911 *Precambrian Research*, 182, 337–350, 2010.
- 912 Haslett, J. and Parnell, A.: A simple monotone process with application to radiocarbon-dated depth
913 chronologies, *Journal of the Royal Statistical Society Series C: Applied Statistics*, 57, 399–418, 2008.
- 914 Hay, C. C., Creveling, J. R., Hagen, C. J., Maloof, A. C., and Huybers, P.: A library of early Cambrian
915 chemostratigraphic correlations from a reproducible algorithm, *Geology*, 47, 457–460, 2019.
- 916 Heaton, T. J., Blaauw, M., Blackwell, P. G., Ramsey, C. B., Reimer, P. J., and Scott, E. M.: The IntCal20
917 approach to radiocarbon calibration curve construction: a new methodology using Bayesian splines and
918 errors-in-variables, *Radiocarbon*, 62, 821–863, 2020.
- 919 Holmes, J. D. and Budd, G. E.: Reassessing a cryptic history of early trilobite evolution, *Communications*
920 *Biology*, 5, 1177, 2022.
- 921 Keller, C. B., Schoene, B., and Samperton, K. M.: A stochastic sampling approach to zircon eruption age
922 interpretation, *Geochemical Perspectives Letters (Online)*, 8, 2018.
- 923 Kouchinsky, A., Bengtson, S., Pavlov, V., Runnegar, B., Torssander, P., Young, E., and Ziegler, K.:
924 Carbon isotope stratigraphy of the Precambrian–Cambrian Sukharikha River section, northwestern
925 Siberian platform, *Geological Magazine*, 144, 609–618, 2007.
- 926 Landing, E. and Kruse, P. D.: Integrated stratigraphic, geochemical, and paleontological late Ediacaran to
927 early Cambrian records from southwestern Mongolia: Comment, *Bulletin*, 129, 1012–1015, 2017.



- 928 Landing, E., Bowring, S. A., Davidek, K. L., Westrop, S. R., Geyer, G., and Heldmaier, W.: Duration of
929 the Early Cambrian: U-Pb ages of volcanic ashes from Avalon and Gondwana, *Canadian Journal of Earth*
930 *Sciences*, 35, 329–338, 1998.
- 931 Landing, E., Schmitz, M. D., Geyer, G., Trayler, R. B., and Bowring, S. A.: Precise early Cambrian U–Pb
932 zircon dates bracket the oldest trilobites and archaeocyaths in Moroccan West Gondwana, *Geological*
933 *Magazine*, 158, 219–238, 2021.
- 934 Langereis, C. G., Krijgsman, W., Muttoni, G., Menning, M., et al.: Magnetostratigraphy-concepts,
935 definitions, and applications, *Newsletters in Stratigraphy*, 43, 207, 2010.
- 936 Lantink, M. L., Davies, J. H., Ovtcharova, M., and Hilgen, F. J.: Milankovitch cycles in banded iron
937 formations constrain the Earth–Moon system 2.46 billion years ago, *Proceedings of the National*
938 *Academy of Sciences*, 119, e2117146119, 2022.
- 939 Lee, T., Rand, D., Lisiecki, L. E., Gebbie, G., and Lawrence, C. E.: Bayesian age models and stacks:
940 Combining age inferences from radiocarbon and benthic $\delta^{18}\text{O}$ stratigraphic alignment, *EGUsphere*,
941 2022, 1–29, 2022.
- 942 Lisiecki, L. E. and Lisiecki, P. A.: Application of dynamic programming to the correlation of
943 paleoclimate records, *Paleoceanography*, 17, 1–1, 2002.
- 944 Magaritz, M., Kirschvink, J. L., Latham, A. J., Zhuravlev, A. Y., and Rozanov, A. Y.:
945 Precambrian/Cambrian boundary problem: Carbon isotope correlations for Vendian and Tommotian time
946 between Siberia and Morocco, *Geology*, 19, 847–850, 1991.
- 947 Maloof, A. C., Schrag, D. P., Crowley, J. L., and Bowring, S. A.: An expanded record of Early Cambrian
948 carbon cycling from the Anti-Atlas Margin, Morocco, *Canadian Journal of Earth Sciences*, 42, 2195–
949 2216, 2005.
- 950 Maloof, A. C., Ramezani, J., Bowring, S. A., Fike, D. A., Porter, S. M., and Mazouad, M.: Constraints on
951 early Cambrian carbon cycling from the duration of the Nemakit-Daldynian–Tommotian boundary $\delta^{13}\text{C}$
952 shift, Morocco, *Geology*, 38, 623–626, 2010.
- 953 Miall, A. D.: Updating uniformitarianism: stratigraphy as just a set of “frozen accidents”, *Geological*
954 *Society, London, Special Publications*, 404, 11–36, 2015.



- 955 Middleton, J. L., Gottschalk, J., Winckler, G., Hanley, J., Knudson, C., Farmer, J. R., Lamy, F., and
956 Lisiecki, L. E.: Evaluating manual versus automated benthic foraminiferal $\delta^{18}\text{O}$ alignment techniques
957 for developing chronostratigraphies in marine sediment records, *Geochronology*, 6, 125–145, 2024.
- 958 Paterson, J. R., Edgecombe, G. D., and Lee, M. S.: Trilobite evolutionary rates constrain the duration of
959 the Cambrian explosion, *Proceedings of the National Academy of Sciences*, 116, 4394–4399, 2019.
- 960 Roberts, G. O. and Rosenthal, J. S.: Examples of adaptive MCMC, *Journal of computational and*
961 *graphical statistics*, 18, 349–367, 2009.
- 962 Rudman, A. J. and Lankston, R. W.: Stratigraphic correlation of well logs by computer techniques, *Aapg*
963 *Bulletin*, 57, 577–588, 1973.
- 964 Sadler, P. M.: Sediment accumulation rates and the completeness of stratigraphic sections, *The Journal of*
965 *Geology*, 89, 569–584, 1981.
- 966 Sakoe, H. and Chiba, S.: Dynamic programming algorithm optimization for spoken word recognition,
967 *IEEE transactions on acoustics, speech, and signal processing*, 26, 43–49, 1978.
- 968 Saltzman, M., Thomas, E., and Gradstein, F.: Carbon isotope stratigraphy, *The geologic time scale*, 1,
969 207–232, 2012.
- 970 Sambridge, M.: A parallel tempering algorithm for probabilistic sampling and multimodal optimization,
971 *Geophysical Journal International*, 196, 357–374, 2014.
- 972 Schmitz, M., Gradstein, F., Ogg, J., et al.: Radiogenic isotope geochronology, *The geologic time scale*, 2,
973 115–126, 2012.
- 974 Sinnesael, M., Millard, A. R., and Smith, M. R.: A Bayesian astrochronology for the Cambrian first
975 occurrence of trilobites in West Gondwana (Morocco), *Geology*, 52, 205–209, 2024.
- 976 Smith, E. F., Macdonald, F. A., Petach, T. A., Bold, U., and Schrag, D. P.: Integrated stratigraphic,
977 geochemical, and paleontological late Ediacaran to early Cambrian records from southwestern Mongolia,
978 *Geological Society of America Bulletin*, 128, 442–468, 2016.
- 979 Sylvester, Z.: Automated multi-well stratigraphic correlation and model building using relative geologic
980 time, *Basin Research*, 35, 1961–1984, 2023.
- 981 Tucker, M. E.: Carbon isotope excursions in Precambrian/Cambrian boundary beds, Morocco, *Nature*,
982 319, 48–50, 1986.



- 983 Uhlein, G. J., Uhlein, A., Pereira, E., Caxito, F. A., Okubo, J., Warren, L. V., and Sial, A. N.: Ediacaran
984 paleoenvironmental changes recorded in the mixed carbonate-siliciclastic Bambuí Basin, Brazil,
985 Palaeogeography, palaeoclimatology, palaeoecology, 517, 39–51, 2019.
- 986 Varlamov, A., Rozanov, A. Y., Khomentovsky, V., Shabanov, Y. Y., Abaimova, G., Demidenko, Y. E.,
987 Karlova, G., Korovnikov, I., Luchinina, V., Malakhovskaya, Y. E., et al.: The Cambrian System of the
988 Siberian Platform. Part 1: The Aldan-Lena Region, in: XIII Field Conference of the Cambrian Stage
989 Subcommittee Working Group, 2008.
- 990 Vats, D. and Knudson, C.: Revisiting the Gelman–Rubin diagnostic, Statistical Science, 36, 518–529,
991 2021.
- 992 Vats, D., Flegal, J. M., and Jones, G. L.: Multivariate output analysis for Markov chain Monte Carlo,
993 Biometrika, 106, 321–337, 2019.
- 994 Vousden, W., Farr, W. M., and Mandel, I.: Dynamic temperature selection for parallel tempering in
995 Markov chain Monte Carlo simulations, Monthly Notices of the Royal Astronomical Society, 455, 1919–
996 1937, 2016.
- 997 Wheeler, L. and Hale, D.: Simultaneous correlation of multiple well logs, in: SEG International
998 Exposition and Annual Meeting, SEG–2014, 2014.
- 999 Yang, J. and Rosenthal, J. S.: Automatically tuned general-purpose MCMC via new adaptive diagnostics,
1000 Computational Statistics, 32, 315–348, 2017.
- 1001 Zhang, X., Ahlberg, P., Babcock, L. E., Choi, D. K., Geyer, G., Gozalo, R., Hollingsworth, J. S., Li, G.,
1002 Naimark, E. B., Pegel, T., et al.: Challenges in defining the base of Cambrian Series 2 and Stage 3, Earth-
1003 Science Reviews, 172, 124–139, 2017.