

# Energy Transfer from Internal Solitary Waves to Turbulence via High-Frequency Internal Waves: Seismic Observations in the Northern South China Sea

Linghan Meng<sup>1,2</sup>, Haibin Song<sup>1,2</sup>, Yongxian Guan<sup>3</sup>, Shun Yang<sup>1,2</sup>, Kun Zhang<sup>1,2</sup>, Mengli Liu<sup>1,2</sup>

<sup>1</sup> School of Ocean and Earth Science, Tongji University, Shanghai, 200092, China

<sup>2</sup> State Key Laboratory of Marine Geology, Tongji University, Shanghai, 200092, China

<sup>3</sup> MNR Key Laboratory of Marine Mineral Resources, Guangzhou Marine Geological Survey, China Geological Survey, Guangzhou, 510760, China

**Correspondence:** Haibin Song (hbsong@tongji.edu.cn) and Yongxian Guan (813385148@qq.com)

**Abstract.** The shoaling and breaking of internal waves (IW) are critical processes in the ocean's energy cascade and mixing. Using the seismic data, we observed high-frequency internal waves (HIWs), which were primarily distributed in the depth range of 79-184 m. Their amplitude scale is  $O(10\text{ m})$ , with half-height widths ranging from 154 to 240 m. The shoaling thermocline and gentle slope with a low internal Iribarren number suggest that high-frequency internal waves observed are likely a result of fission. The remote sensing data supports the point. Instability estimations showed that due to the strong vertical shear,  $Ri$  in the range of 20-30 km was less than  $1/4$ , and KH billows can be found in the seismic transect, suggesting that these waves were unstable and might dissipate rapidly. We use seismic data to estimate diapycnal mixing, and we found that the HIWs can enhance diapycnal mixing, averaging  $10^{-4}\text{ m}^2\text{s}^{-1}$ . The maximum mixing value is up to  $10^{-3}\text{ m}^2\text{s}^{-1}$ , and it is associated with the breaking of IWs caused by the strong shear. The results show a new energy cascade route from shoaling internal solitary waves (ISWs) to turbulence, i.e., the fission of ISWs into HIWs, which improves our knowledge of ISW energy dissipation and their roles in improved mixing in the northern South China Sea.

## 1 Introduction

Internal solitary waves (ISWs) are widely distributed in the global ocean, and the South China Sea is recognized as one of the best places to study large-amplitude ISWs (Klymak et al., 2006; Zheng et al., 2022; Cai et al., 2015). After years of research, there is a clearer understanding of their generation, propagation, and evolution: barotropic tides interact with complex topography to produce internal tides (baroclinic tides), and internal tides radiate out from the Luzon Strait, with those radiating westward crossing the northern South China Sea basin and propagating towards the slope and shelf areas. During propagation, nearly half of the internal tide energy generates ISWs through nonlinear steepening, and these waves deform due to shoaling effects, eventually breaking and dissipating on the shelf (Alford et al., 2015; Bourgault et al., 2007; Fu et al., 2012; Sinnett et

29 al., 2022; Liu et al., 2022; Zhang et al., 2023). These large-amplitude ISWs have significant impacts on ocean mixing, sediment  
30 suspension, nutrient transport, and offshore oil and gas engineering (Bogucki et al., 1997; Osborne et al., 1978; Wang et al.,  
31 2007; Xu and Yin, 2011).

32 When energy transfers from large and mesoscale motions to small-scale turbulence, internal waves (IWs) play a crucial  
33 role (Liu et al., 2022). Particularly in slope-shelf areas, the shoaling evolution and dissipation of IWs (primarily ISWs) are key  
34 processes driving turbulent mixing. As ISWs propagate from deep to shallow waters, they interact with the seabed topography,  
35 a process known as shoaling (Sinnott et al., 2022). Ocean instruments such as mooring and high-frequency acoustics can help  
36 us better record this process (e.g., Bourgault et al., 2007; Fu et al., 2012; Orr and Mignerey, 2003; Xu and Yin, 2012). Generally,  
37 at critical depths, the rear of a depression wave steepens and undergoes polarity reversal (the shape of the soliton is evolving  
38 from depression to elevation as the slope of the leading edge is less than the trailing slope) to form an elevation wave (e.g.,  
39 Shroyer et al., 2009). When the continental slope is relatively gentle, dispersion continues to form elevated wave trains, a  
40 process called fission (Bai et al., 2019; Djordjevic and Redekopp, 1978; Gong et al., 2021b; Liu et al., 1998; Zheng et al., 2001;  
41 Vlasenko et al., 2002). In the slope-shelf regions of the South China Sea, where the topography is relatively gentle, ISWs  
42 frequently appear as wave packets through fission, as confirmed by satellite images (e.g., Zhao et al., 2003). During this process,  
43 the energy of ISWs continuously dissipates into the seawater, enhancing turbulent mixing. Studies have shown that the mixing  
44 caused by fission is second only to that caused by plunging breakers, which have a longer time scale for mixing over the length  
45 of the gentle slope (Masunaga et al., 2019). Ultimately, these shoaling waves break due to shear instability or convective  
46 instability, generating turbulence and achieving the downscale transfer of ocean energy (Lamb et al., 2014). This represents  
47 our most common understanding of the dissipation mechanisms of ISWs.

48 However, Bai et al. (2013, 2019) discovered that ISWs on the South China Sea shelf also undergo fission to produce high-  
49 frequency internal waves (HIWs) during shoaling and proposed a new hypothesis for the energy dissipation. They suggested  
50 that the fission of ISWs into HIWs is also a key process for energy dissipation. Their subsequent theoretical analyses and  
51 numerical simulations confirmed this hypothesis. In their review, Rippeth and Green (2020) pointed out that this finding is  
52 significant for understanding the dissipation mechanisms of ISWs. And they think the fission of ISWs into HIWs on the  
53 continental shelf is an important pathway for the tidal energy cascade.

54 Therefore, the energy dissipation mechanisms of IWs (ISWs) are complex and the ocean is a multi-scale coupled system  
55 with various dynamic processes. However, most current studies consider the ideal scenario where only IWs are present in the  
56 ocean's dynamic system, with few considering or only considering simplified dynamic environmental impacts. This results in  
57 an incomplete understanding of the shoaling processes and ISWs' dissipation mechanisms. Additionally, whether HIWs,  
58 another product of fission, significantly enhance energy dissipation and mixing remains a critical question for comprehending  
59 the formation and evolution of wave packets and their role in ocean mixing and energy cascades.

60 In this paper, we use the seismic method to investigate these questions. Seismic oceanography, initially proposed by  
61 Holbrook (2003), has provided us with a novel high-resolution method to image the thermohaline fine structure in the water  
62 column. The image formed by the reflection seismic method in the water column is mainly the result of the convolution of the

63 seawater temperature gradient and the seismic source wavelet (Ruddick et al., 2009; Sallarès et al., 2009). Due to its advantages  
64 of fast acquisition, high horizontal resolution, and full-depth water column imaging, it has been widely proven to be an effective  
65 method for capturing ocean dynamic phenomena covering a multi-scale range from fine-scale  $O$  (10 m) to mesoscale  $O$  (100  
66 km) (Song et al., 2021a). Many ocean features are captured and imaged through these high-resolution seismic images, such as  
67 ocean currents (e.g., Tsuji et al., 2005), eddies (e.g., Yang et al., 2022), fronts (e.g., Gunn et al., 2021), internal solitary waves  
68 (e.g., Bai et al., 2017; Geng et al., 2019; Gong et al., 2021a; Song et al., 2021b; Tang et al., 2014 ; Tang et al., 2018),  
69 thermohaline staircases (e.g., Fer et al., 2010), and turbulence (e.g., Sheen et al., 2009). In recent years, this method has  
70 gradually become more quantitative, especially in estimating turbulent dissipation rate and diffusivity (e.g., Dickinson et al.,  
71 2017; Gong et al., 2021b; Tang et al., 2021; Yang et al., 2023). In this study, we applied this method to seismic line 25, where  
72 we sequentially observed HIWs generated by fission, bore-like nonlinear waves, and induced shear instabilities. Based on  
73 seismic data, we estimated turbulent dissipation rates and found that HIWs indeed enhance turbulent mixing. This indicates  
74 that the fission of ISWs into HIWs is also a key process for energy dissipation.

## 75 **2 Data and methods**

### 76 **2.1 Seismic data acquisition and processing**

77 From July to September 2009, the Guangzhou Marine Geological Survey Bureau conducted multi-channel reflection  
78 seismic data collection on the northern slope and shelf of the South China Sea, covering 43 lines (not shown). Our study  
79 focused on Line 25 (the blue solid line in Fig. 1a), which was collected on July 31. The acquisition vessel towed an air gun  
80 array with a total volume of 5080 in<sup>3</sup> (approximately 83 L), firing every 25 m (shot interval). The main frequency of the source  
81 was 35 Hz. A 6 km streamer with 480 hydrophones, spaced 12.5 m apart and with a minimum offset of 250 m (the distance  
82 from the air gun array to the nearest hydrophone), was towed to receive reflections, sampling every 2 ms.

83 To clearly image the water column, standard processes were applied to marine seismic data, including 1) defining the  
84 observation system; 2) denoising and direct wave suppression; 3) velocity analysis; 4) NMO correction; 5) stacking; 6) further  
85 denoising; and 7) migration. A detailed description of the seismic data processing can be found in Ruddick et al. (2009) and  
86 Holbrook et al. (2013).

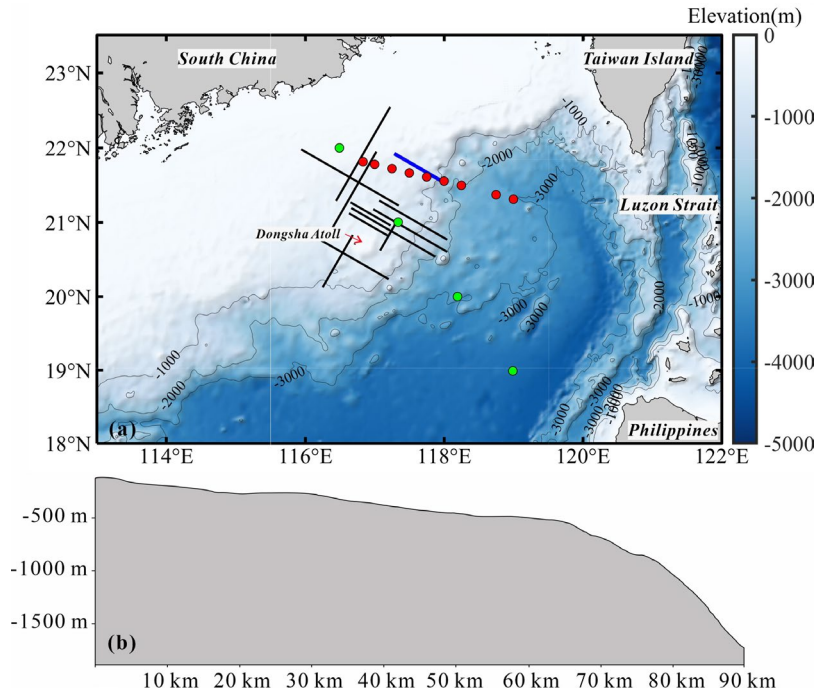
87 In the process, several processing steps are particularly critical. A key step is using median filtering and matched  
88 subtraction in shot gathers to suppress direct waves, avoiding their strong energy overshadowing the reflections and enhancing  
89 the imaging quality of the water column (Gong et al., 2021b). Secondly, to ensure the extracted reflectors accurately represent  
90 isopycnal displacement, further denoising (step 6) is required: band-pass filtering is applied to ambient noise, and notch  
91 filtering to harmonic noise, ensuring the seismic frequency band carries the best possible turbulence information. Holbrook et  
92 al. (2013) used the signal-to-noise ratio of adjacent traces to assess the optimal filtering range, calculated as

$$93 \quad \frac{S}{N} = \sqrt{\frac{|c|}{|a-c|}}, \quad (1)$$

94 where  $c$  is the maximum cross-correlation coefficient of adjacent traces, and  $a$  is the autocorrelation coefficient of the first  
 95 trace. The median S/N value across all traces determines the final result for the profile. Upon verification, we used an 8-12-  
 96 75-85 Hz band-pass filter to suppress ambient noise. Additionally, harmonic noise from shots appears periodically in the  
 97 horizontal wavenumber domain  $k=k_s$  ( $k_s=n/\eta$ , where  $n$  is an integer,  $\eta$  is the shot spacing) as pulses. **Given the 25 m shot**  
 98 **interval, appropriate notch filters were designed to suppress harmonic noise** near  $k = 0.04 \text{ m}^{-1}$ ,  $0.08 \text{ m}^{-1}$ , and so on. This  
 99 denoising process improved the S/N ratio to 9, significantly exceeding the minimum standard of 4 (Holbrook et al., 2013).

100 Finally, assuming a seawater speed of 1500 m/s (with variations from 1480-1540 m/s), the Stolt migration method quickly  
 101 reveals the true form and position of reflectors, facilitating subsequent reflector picking. **We could select a portion of the image**  
 102 **to compare the effects before and after filtering.** Figures 2a and 2b, respectively show the seismic images of the 69-81 km  
 103 section before and after filtering, and Figure 2c is the difference between Figures 2a and 2b. We observed that the seismic  
 104 image after filtering (Fig. 2b, 2e) allows for the identification of more reflectors (blue solid lines), with the average length  
 105 increasing from 2.1 km to 2.3 km (Fig. 2d, 2e), and it clearly reveals more turbulent features (Fig. 2f). A Seismic image  
 106 essentially represents a high-resolution snapshot of the ocean's vertical temperature gradient (Ruddick et al., 2009; Sallarès et  
 107 al., 2009). According to Fig. 2a, above 500 m depth (especially between 100-400 m), clear reflectors indicate strong  
 108 thermohaline gradients. Below 500 m, as depth increases, the reflection signals gradually weaken or even disappear, suggesting  
 109 more uniform seawater.

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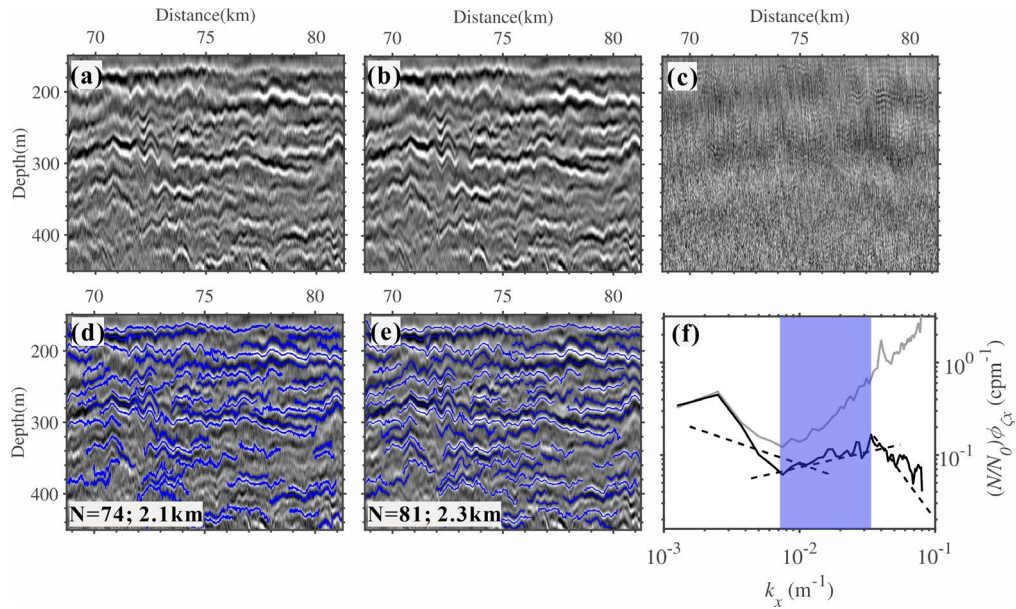
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**Figure 1 Bathymetry map of the South China Sea and seabed topography along seismic line 25. (a) Topographic map of the research area. The blue solid line represents the seismic line 25, red dots are XBT data (numbered from right to left, XBT1-XBT9), green dots**

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are CTD data (numbered from right to left as CTD1-CTD4). The black solid lines have been studied by Geng et al., (2019) and Gong et al. (2021a); (b) Seabed topography along Line 25. The slope  $\gamma = 0.018$ .

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117

118 **Figure 2 Example from Line 25 showing the effects of the band-pass filtering and notch filtering. (a) Original seismic data for 68-81**  
119 **km section; (b) After the band-pass filtering and notch filtering; (c) The difference between the (a) and (b); (d), (e) Tracked reflections**  
120 **(blue solid lines) from the (a) and (b), respectively. (f) Slope spectra. The gray (black) line calculated from the tracked reflectors**  
121 **before(after) filtering of the seismic data.**

122

## 2.2 Hydrographic data

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Hydrographic data can be utilized to analyze local hydrographic characteristics and estimate the buoyancy frequency. In this study, due to the lack of synchronous hydrographic data, we used historical CTD and XBT data collected in September 2009, which is close to the acquisition time of the seismic data. Figure 3 shows the vertical variations of temperature, salinity,

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density, and buoyancy frequency with depth, derived from 4 CTD. The buoyancy frequency  $N$  can be expressed as  $\sqrt{-\frac{\rho}{g} \frac{d\rho}{dz}}$ ,

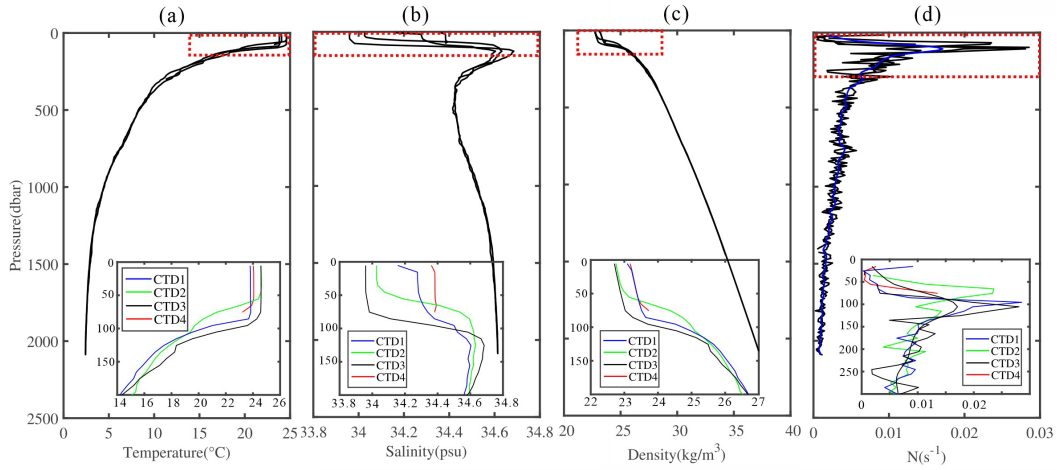
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where  $\rho$  is the seawater density,  $g$  is the gravitational acceleration, and  $z$  is the seawater depth. The alignment direction of the 9 XBT stations roughly corresponds to the survey line direction, allowing for an approximate assessment of changes in the thermocline depth with varying seawater depth.

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**Figure 3** Data for temperature (a), salinity (b), density (c), and buoyancy frequency (d) from 4 CTD stations (see station distribution in Fig. 1b). Insets in each graph (a, b, c, d) show data from a depth of 200-300 m within the red dashed-line boxes. The thick blue solid line in graph (d) represents the average buoyancy frequency of seawater.

### 2.3 Geostrophic shear estimation

To estimate the geostrophic velocity in the density field, it is necessary to assume that the Coriolis force and the horizontal pressure gradient force are balanced. The Rossby number  $R_\theta = U/fL$  should be much less than 1, where  $U$  is the characteristic velocity and  $L$  is the characteristic length. Under such geostrophic balance conditions, the density layers within the water column tend to tilt. In two-dimensional transects, the vertical shear  $S$  of the horizontal velocity perpendicular to the density field is determined using the standard thermal wind equation based on the slope of the isopycnal (McWilliams 2006). Thus,

$$S = \frac{g}{\rho_0 f} \frac{\partial \rho}{\partial z} \tan \gamma, \quad (2)$$

where  $g = 9.8 \text{ m s}^{-2}$  is the gravitational acceleration,  $\rho_0$  is the mean density of two layers,  $\frac{\partial \rho}{\partial z}$  is the vertical gradient of density across the isopycnal,  $f$  is the Coriolis parameter, and  $\gamma$  is the slope of the isopycnal. Krahnmann et al. (2009) demonstrated that the slope of the isopycnal measured using a yoyo CTD probe matches the slope of the reflections from the seismic data within 4 h. Given that the average speed of the vessel is 2.5 m/s, and within a certain length range ( $\sim 36 \text{ km}$ ), these reflections coincide with the isopycnal. Therefore, we can use seismic data to obtain the slope of the reflections as an approximation for the isopycnal slope  $\gamma$ .

In the study area,  $U \approx 0.1 \text{ m s}^{-1}$ ,  $f \approx 5 \times 10^{-5} \text{ s}^{-1}$ , and for  $R_\theta=1$ ,  $L = 2 \text{ km}$ . To ensure  $R_\theta \ll 1$ , we can grid the seismic transect with each window sized at  $10 \text{ km} \times 75 \text{ m}$  (length  $\times$  width), i.e.,  $L=10 \text{ km}$ , and an automatic picking algorithm is used to identify the seismic reflections within each window. The specific estimation method can be referred to in previous works (Sheen et al., 2011; Tang et al., 2020).

### 2.4 Turbulent dissipation and diapycnal diffusivity estimates from seismic data

153 Klymak and Moum (2007) used horizontal wavenumber spectra from oceanic horizontal towed measurements to estimate  
 154 the diapycnal diffusivity ( $K_\rho$ ) in seawater. In the open ocean, the horizontal wavenumber spectrum ( $\phi_\zeta$ ) can be clearly divided  
 155 into two parts: the low wavenumber part is related to internal waves ( $\phi_\zeta^{IW}$ ), with spectral characteristics consistent with the  
 156 Garrett and Munk (1975) model (GM75), and proportional to the  $-2.5$  power of the wavenumber; the high wavenumber part is  
 157 dominated by turbulence ( $\phi_\zeta^T$ ), exhibiting Kolmogorov-like behavior (proportional to the  $-5/3$  power of the wavenumber).

158 Compared with the internal wave subrange, diapycnal diffusivity values computed from the turbulent subrange are  
 159 probably more robust (Klymak & Moum, 2007a; Sheen et al., 2009). Therefore, we choose the turbulent subrange of the slope  
 160 spectra to estimate diapycnal mixing in this study. The horizontal wavenumber spectra in the turbulent subrange can be  
 161 represented by a simplified Batchelor (1959) model (equation 4), so the turbulence kinetic energy dissipation rate ( $\varepsilon$ ) can be  
 162 obtained from the horizontal wavenumber spectra. According to equation 5, the diffusivity  $K_\rho$  can be obtained (Osborn, 1980):

$$\phi_\zeta^T = \frac{4\pi\Gamma}{N^2} C_T \varepsilon^{2/3} (2\pi k_x)^{-5/3}, \quad (3)$$

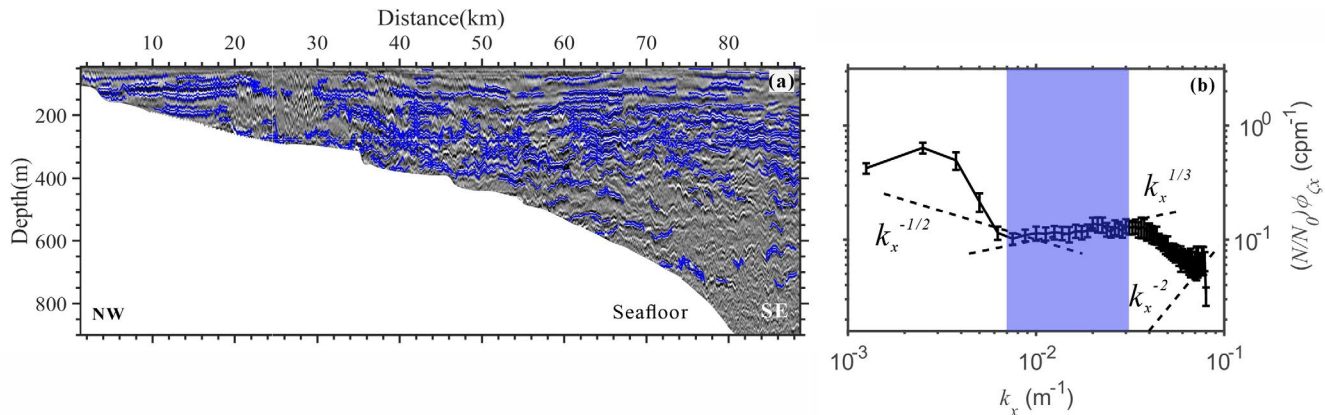
$$K_\rho = \Gamma \varepsilon / N^2, \quad (4)$$

163 where  $\Gamma \approx 0.2$  is the mixing efficiency of seawater,  $C_T \approx 0.4$  is considered a constant, and  $N$  is the average buoyancy frequency  
 164 at the corresponding water depth.

165 Assuming that the isopycnals of seawater coincide with the seismic reflectors (Holbrook et al., 2013; Sheen et al., 2009;  
 166 Krahmann et al., 2009), the horizontal wavenumber spectra obtained from the vertical displacement of reflectors can replace  
 167 those obtained from horizontal towed measurements; thus the turbulence dissipation rate and diapycnal diffusivity can also be  
 168 calculated from seismic data. (e.g., Dickinson et al., 2017; Gong et al., 2021b; Tang et al., 2021; Yang et al., 2023). **The steps**  
 169 **are as follows: Firstly, an automatic picking algorithm was used to pick seismic reflectors in the seismic data, and the length**  
 170 **of reflectors should be no less than 1 km, totaling 410 (Fig. 4a). And then the vertical displacement of the picked reflectors is**  
 171 **obtained by subtracting their linear fit curves, and the power spectral density of the displacement curve, i.e., the horizontal**  
 172 **wavenumber spectrum ( $\phi_\zeta^T$ ), is calculated using a Fourier transform. The spectral calculation process uses a 128-point sampling**  
 173 **point width and a non-overlapping Hanning window. To better distinguish internal wave subrange from turbulence subrange,**  
 174 **the displacement spectrum ( $\phi_\zeta^T$ ) is usually multiplied by  $(2\pi k_x)^2$  to obtain the slope spectrum ( $\phi_{\zeta_x}^T$ ), so the slope of the**  
 175 **turbulence subrange changes from  $-5/3$  to  $1/3$ , and the slope of the internal wave subrange changes from  $-5/2$  to  $-1/2$  (Holbrook**  
 176 **et al., 2013). After obtaining the slope spectra, we use the least square method to fit the spectra in the turbulent subrange**  
 177 **( $0.0075$ - $0.0378 \text{ m}^{-1}$ , the blue shaded area in Fig. 4c) and obtain fitted spectra to ensure the stability of the results. Finally, the**  
 178 **fitted spectrum is substituted into equations 3 and 4 to obtain the turbulent dissipation rate and diapycnal mixing. Additionally,**  
 179 **to eliminate the influence of seawater stratification ( $N$ ) on the internal wave field, we need to normalize the slope spectra**  
 180 **according to the local average buoyancy frequency, i.e., multiply by  $N/N_0$ , where  $N$  is derived from Fig. 3d,  $N_0=3 \text{ cph}$ .**

181 To obtain the spatial distribution of the mixing parameters, we gridded the seismic transect of line 25, with each window  
 182 size being  $5 \text{ km} \times 75 \text{ m}$  (length  $\times$  width), the lateral step size  $2.5 \text{ km}$ , and the vertical step size  $37.5 \text{ m}$ . The average value within  
 183 each window is taken as the dissipation rate and diffusivity for that window, following the previously mentioned method.  
 184  
 185

186 Moreover, we interpolated and smoothed the data appropriately to ensure the spatial distribution's continuity.  
 187



188  
 189 **Figure 4 Picked reflections and the slope spectrum. (a) Picked Reflectors from the seismic Line 25 (blue solid lines); (b) Slope**  
 190 **spectrum calculated from the reflectors in Fig. 4a, with 95% confidence interval. The black dashed lines with -1/2 and 1/3 slopes**  
 191 **correspond to the theoretical slopes of GM75 for internal wave subrange and Batchelor model for turbulence subrange, respectively.**  
 192 **The black dashed line with -2 slope is the noise subrange. The blue shaded area represents the turbulent subrange for fitting the**  
 193 **spectrum.**

### 194 3 Results

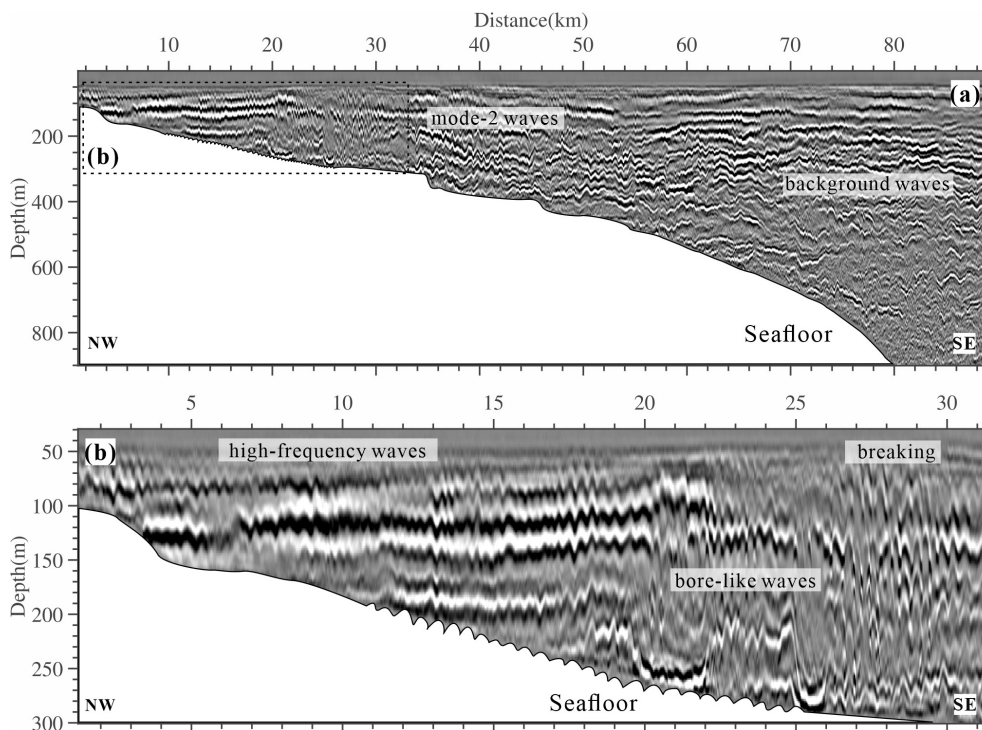
#### 195 3.1 Seismic observations

196 **Figure 5a** shows the seismic image obtained through the aforementioned standard processing procedure. We observe that  
 197 in the shallow water region above 420 m, the seismic reflection signal is strong, and the reflections are laterally continuous. In  
 198 contrast, in deeper waters (>500 m), the reflection amplitude is relatively weaker, the reflections lose their continuity, and both  
 199 their number and length are relatively reduced, especially in the deeper water between 70 and 89 km. Horizontally, the seismic  
 200 reflections also exhibit some differences. Likely influenced by the seafloor topography, the reflections on the left side of the  
 201 seismic image (1.25-55 km) show significant spatial morphological changes, with waveforms displaying high-frequency  
 202 oscillations and breaking. Some reflections develop in an inclined manner. The reflections on the right side of the image (60-  
 203 89 km) are mostly horizontally distributed, with the frequency and amplitude of internal waves far less pronounced than those  
 204 on the left.

205 **Figure 5b** is an enlarged view of the 1.25-33 km section of the image. Within 20 km, the seismic image shows a series of  
 206 high-frequency internal waves, characterized by small amplitude and high-frequency oscillations. The reflections gradually  
 207 increase from left to right, with more evident seawater stratification. Adjacent to the high-frequency internal waves (20-26  
 208 km), we can see two bore-like waves with relatively larger amplitude and wave width. These convex-shaped reflective  
 209 structures also remind us of high-mode internal waves, which are discussed in Sect 4. After the first bore-like wave (21-24  
 210 km), the seismic reflections temporarily return to the original stratification in the ocean. However, following the appearance



211 of the second wave, the trailing edge becomes unstable, with the right side of the waveform incomplete, and the reflections  
 212 become intermittent and inclined (26.5-30 km). This indicates that internal wave breaking has occurred, which disrupts the  
 213 original density stratification.



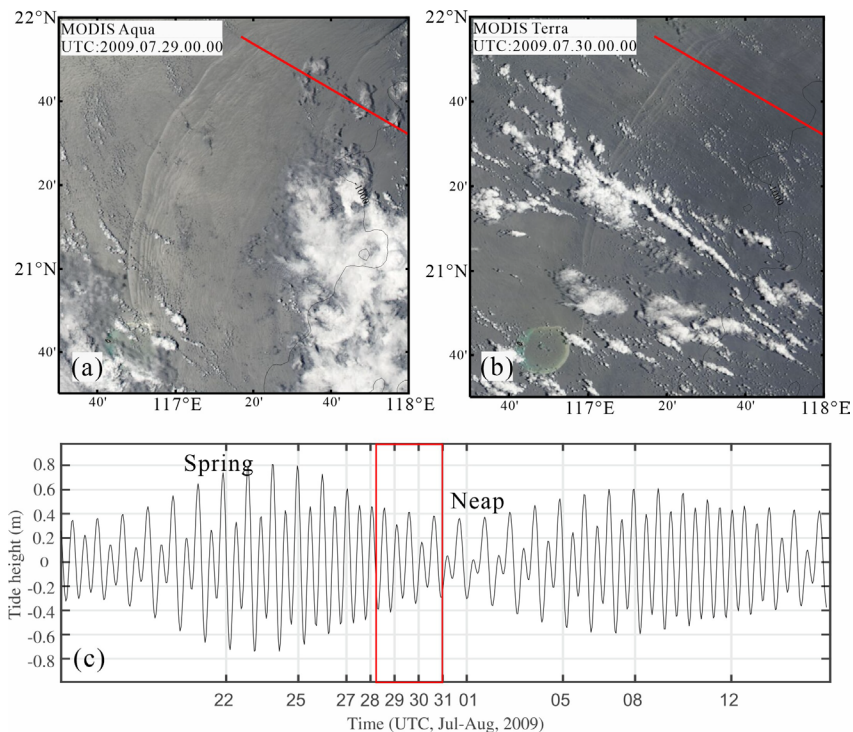
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 215 **Figure 5 Seismic image of the internal waves. (a) Seismic image of line25. Background waves and mode-2 waves are denoted; (b) The**  
 216 **1.25-33 km section of the image (a). High-frequency waves, bore-like waves and breaking waves are denoted.**

### 217 3.2 Satellite images

218 The actual data acquisition time for Line 25 was July 31, 2009. We collected recent MODIS images of the Dongsha waters,  
 219 with Figure 6a captured by the Aqua satellite on July 29 and Figure 6b by the Terra satellite on July 30. The alternating bright  
 220 and dark stripes on these images represent the sea surface manifestations of ISW packets. Both wave packets exhibit long wave  
 221 crests, and the convex curvature of these crests indicates their northwestward propagation on the shoaling continental shelf,  
 222 sequentially passing through Line 25 (indicated by the red solid line). Each pair of adjacent bright and dark stripes corresponds  
 223 to the leading and trailing edges of the wave, with the distance between them representing the wave width. The satellite images  
 224 show that the wave width of both packets gradually decreases from the leading wave until it becomes indistinguishable. Notably,  
 225 the number of identifiable solitons is significantly greater in the first packet compared to the second. This suggests that ISW  
 226 packets near Line 25 are well-developed and propagate northwestward along the line direction on the continental shelf. The  
 227 HIWs observed in the seismic image may result from the fission of these wave packets. However, due to the limited resolution  
 228 (250 m), these features cannot be fully identified from the remote sensing images.

229 Figure 6c shows the total tidal height time series near the Luzon Strait (20.6° N, 121.9° E) from July 14 to August 14,

230 2009. This data is derived from the global ocean tide model TPXO9 developed by Oregon State University (Egbert and  
 231 Erofeeva, 2002), which provides 15 tidal constituents (M2, S2, N2, K2, K1, O1, P1, Q1, Mm, Mf, M4, Mn4, Ms4, 2n2, S1).  
 232 The propagation speed of internal waves is influenced by various factors, including background currents, water stratification,  
 233 and changes in bottom topography, resulting in different speeds at different depths. Generally, internal waves propagate fastest  
 234 in deep sea areas, exceeding  $3 \text{ m s}^{-1}$ . As they propagate towards the continental shelf, their speed decreases to  $1\text{-}2 \text{ m s}^{-1}$  (Alford  
 235 et al., 2010; Cai et al., 2014). The distance from this point to the HIWs observed in the seismic image is approximately 500  
 236 km. Assuming an average speed of  $2 \text{ m s}^{-1}$ , it would take about 69 hours to cover this distance. Therefore, we infer that these  
 237 ISWs passing Line 25 may have originated during the neap tide period (indicated by the red box in Figure 6c).  
 238



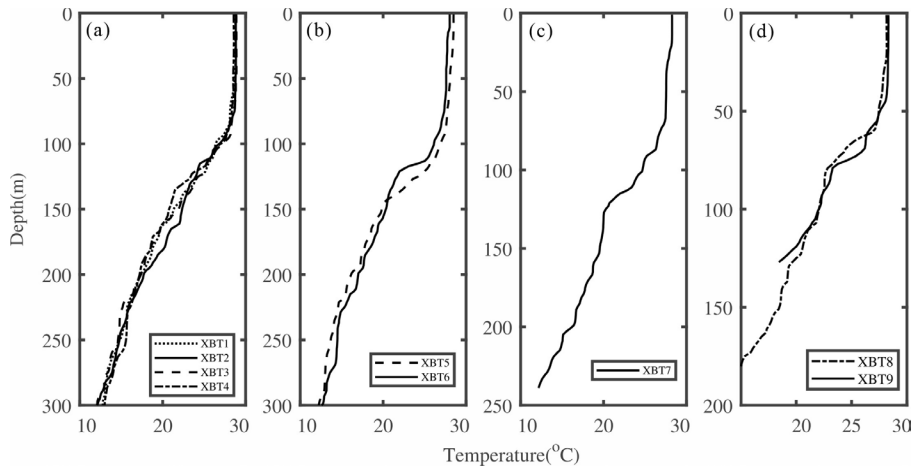
239 **Figure 6** Satellite images and tidal model. (a) MODIS image collected by the Aqua satellite in the northern South China Sea on July  
 240 29, 2009; (b) MODIS image collected by the Terra satellite in the northern South China Sea on July 30, 2009; (c) Tidal height time  
 241 series near the Luzon Strait ( $20.6^\circ \text{ N}$ ,  $121.9^\circ \text{ E}$ ) from July 14 to August 14, 2009. The red box indicates the time period used to infer  
 242 the generation of ISWs.  
 243

### 244 3.3 High-frequency internal waves

245 In fact, the thermocline may be shoaling, with the bottom shoaling on the continental slope or shelf. The density difference  
 246 between the upper and lower layers of seawater also changes with the shoaling topography. Therefore, considering these  
 247 environmental factors is crucial for explaining the evolution of ISWs. In September 2009, 9 XBT data were collected, with  
 248 their alignment direction roughly corresponding to the direction of line 25 (Figure 1a, Figure 7). This allows for a rough

249 estimation of the thermocline depth variation with changing seawater depth. We determine the thermocline depth by taking the  
 250 mid-depth between two inflection points on the temperature profile, where temperature changes significantly. Figure 8 shows  
 251 the variations of thermocline depth and the temperature difference with water depth. We can observe that, except for XBT-7,  
 252 the thermocline shoals with the shoaling topography, with the depth decreasing from 127 m to 57 m and the temperature  
 253 difference between the upper and lower layers decreasing from 8.2°C to 1.8°C. From the perspective of wave dynamics, these  
 254 changes can influence these waves' propagation characteristics in certain cases. Along the track, as the water depth decreases  
 255 from 1800 m to 20 m, Zheng et al. (2001) found that the thermocline shoals from 36 m in deep water to 13 m in shallow water.  
 256 Simultaneously, the density difference between the upper and lower layers of seawater is reduced to half of its original value,  
 257 which facilitates the shoaling of ISWs. The phase speed decreases from 0.85 m/s in deep water to 0.27 m/s in shallow water.  
 258 Therefore, if we are interested in the shoaling of internal waves in the South China Sea, we cannot neglect the changes in the  
 259 ocean dynamic environment.

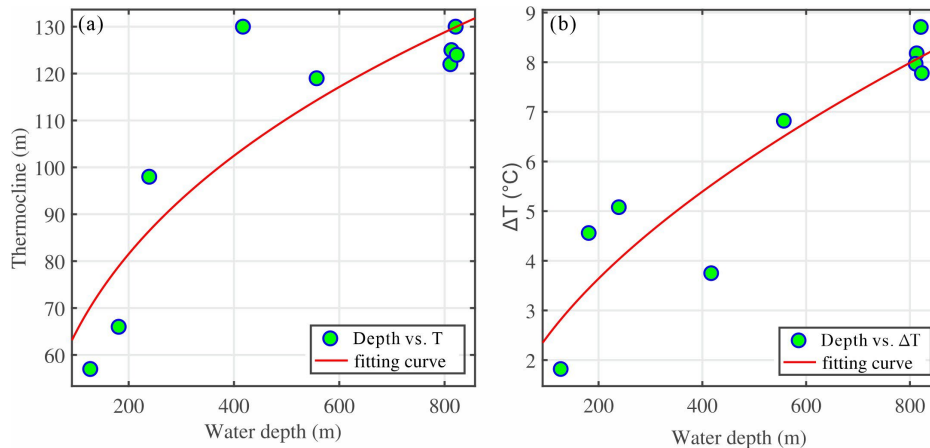
260



261

262 **Figure 7 Temperature profiles from XBT. The distribution of XBT stations is listed in Fig. 1b**

263



264

265 **Figure 8 Variations of (a) thermocline depth and (b) Temperature difference between the upper layer and lower layer with the**  
266 **shoaling water depth**  
267

268 Geng et al. (2019) utilized the same multichannel seismic data to observe large-amplitude ISWs and analyze the  
269 relationship between the vertical structure of the waves and water depth. Gong et al. (2021a) conducted a detailed statistical  
270 analysis of the dynamic characteristics and waveform information of these waves from this cruise, including amplitude and  
271 wavelength (Tables 1 and 2 in Gong et al. (2021a)). Unlike our observations, these waves are large-amplitude solitons that  
272 maintain original shapes during propagation in deeper water (>350 m) and have not yet strongly interacted with the seabed.  
273 While we observed a series of small-amplitude, high-frequency internal waves on one of the seismic lines, which primarily  
274 developed near the topography. Here, we consider these symmetric, large-amplitude ISWs as initial waves. Research indicates  
275 that internal wave breaking is directly related to both topography and waveform, commonly characterized by the internal  
276 Iribarren number ( $\xi_{in} = S/\sqrt{a/L_w}$ ), where  $S$  is the seabed slope,  $a$  is the initial wave amplitude, and  $L_w$  is the initial wave half-  
277 wavelength (Aghsaee et al., 2010). The seabed slope corresponding to line 25 is known to be  $S = 0.018$  (Fig. 1b). Calculations  
278 show that the range of wave steepness  $a/L_w$  is 0.01-0.36, and the range of  $\xi_{in}$  is 0.03-0.17. Therefore, ISWs may undergo  
279 fission under conditions of gentle slope ( $S \leq 0.05$ ) and low internal Iribarren number (Orr and Mignerey, 2003; Sinnett et al.,  
280 2022; Zheng et al., 2001). Additionally,  $a/L_w$  is generally proportional to the Froude number, used to describe the degree of  
281 nonlinearity of internal waves. A ratio greater than 1 would result in seawater overturning (Masunaga et al., 2019), while here  
282 the value is less than 1. Therefore, the seabed slope is one of the key factors determining the shoaling and breaking of internal  
283 waves, consistent with other research findings (e.g., Boegman et al., 2005; Terletska et al., 2020; Vlasenko et al., 2002).

284 The above observations can be summarized as follows. The thermocline depth is shoaling, and the temperature difference  
285 between the upper and lower layers also varies with bottom shoaling. The seabed slope is gentle enough to contribute to the  
286 fission. Bai et al. (2013, 2019) discovered that ISWs also undergo fission in the South China Sea shelf region, generating HIWs,  
287 which provides a new pathway for the energy dissipation of these waves. Therefore, the HIWs seen in Fig. 5b might also be  
288 the result of the fission. The high-frequency internal waves (packets) produced during the shoaling are mainly distributed  
289 within a depth range of 79-184 m, with an average amplitude of 7-9 m and a maximum amplitude of about 13 m. The half-  
290 height width ranges from 154 m to 240 m, as detailed in Table 1. It is similar to the size of high-frequency internal waves  
291 observed by Bai et al. (2019). Near the thermocline, the reflections exhibit high-frequency oscillations and good horizontal  
292 continuity. Closer to the seabed, due to strong interactions between internal waves and topography, the reflections become  
293 distorted and gradually lose their horizontal continuity (horizontal distance 18-20 km, water depth 200-250 m).

294 **Table 1 Waveforms' Statistics of High-Frequency ISWs.**

HIWs#	Num	$\bar{A}$ (m)	$A_0$ (m)	$\bar{L}$ (m)	$\bar{D}$ (m)
1	39	8.2	12.8	205.4	79.1

2	26	9.3	12.3	213.4	87.9
3	7	6.9	11.7	201.9	104.2
4	47	6.7	11.5	212.1	120.6
5	28	7.9	12.0	240.0	144.4
6	29	7.3	10.2	191.5	173.6
7	24	8.4	11.4	153.9	184.0

295 **HIWs#**, the number of high-frequency internal waves (packets); **Num**, the number of elevation waves in the packet;  $\bar{A}$ , average  
 296 amplitude;  $A_0$ , maximum amplitude in the packet;  $\bar{L}$ , average half-height width;  $\bar{D}$ , average depth of the packet location.

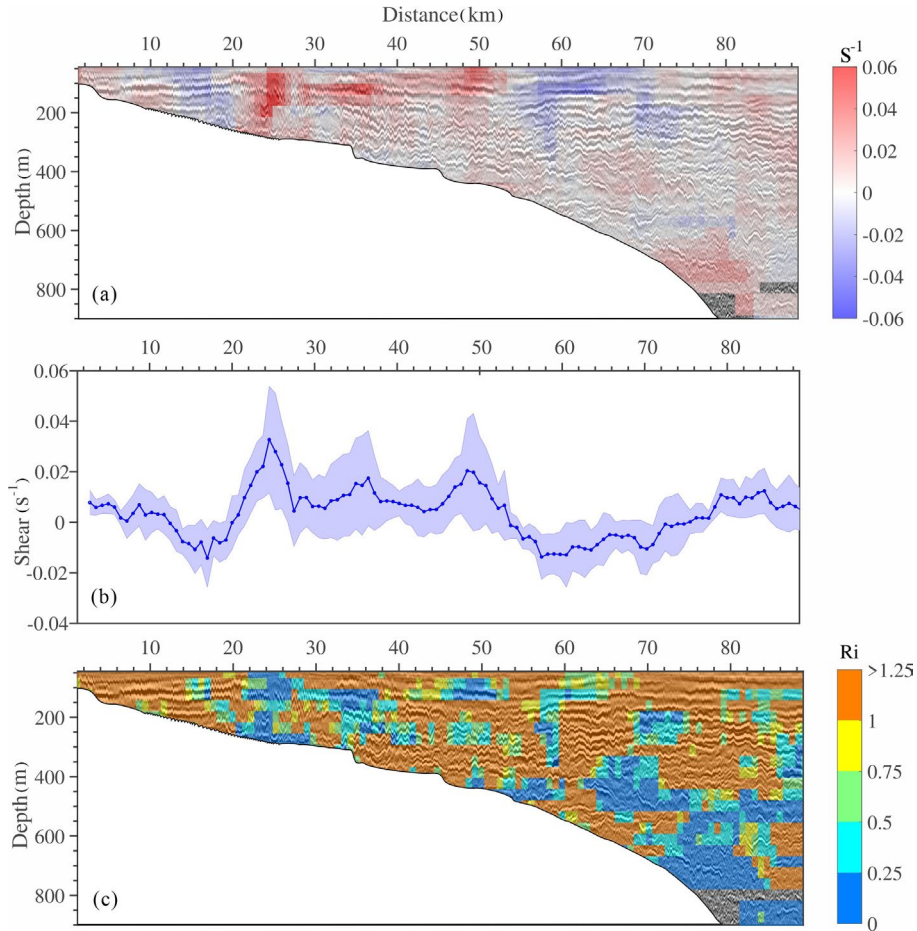
### 297 3.4 Geostrophic shear and shear instability

298 Based on the geostrophic balance theory, we can estimate the vertical shear of horizontal velocity by tracking seismic  
 299 reflectors. Vertical shear (in color) is superimposed on the seismic images (in gray) to better analyze the spatial distribution of  
 300 shear. In Figure 9a, the shear on the left side of the seismic transects is significantly stronger than on the right side, and its  
 301 strength often correlates positively with the inclination of seismic reflectors. In the high-frequency internal wave region (<20  
 302 km), the reflectors are horizontally distributed but exhibit high-frequency oscillations, resulting in an average shear of  $5 \times 10^{-3}$   
 303  $s^{-1}$ . In regions with bore-like waves and internal wave breaking (20-30 km), the reflectors show large inclinations or even  
 304 break, resulting in the highest shear across the transect, up to  $0.03 s^{-1}$ . The reflectors are slightly inclined at 30-45 km due to  
 305 seabed topography, which also causes shear. The shear caused by the background waves gradually weakens from left to right  
 306 (>70 km).

307 Shear instability can be quantified using the Richardson number  $Ri = N^2 / S^2$ , where  $N$  is the buoyancy frequency and  $S$  is  
 308 the geostrophic shear. Generally, an area is considered unstable when  $Ri < 0.25$  (Lamb et al., 2014). Here, we use the average  
 309 buoyancy frequency obtained from historical CTD data as  $N$  and the shear estimated from seismic data as  $S$ . As shown in  
 310 Figure 9b, regions where  $Ri < 0.25$  account for about 8% of the seismic image. In the 20-30 km range at depths of 150-250  
 311 meters, weak stratification and strong shear cause  $Ri$  to be less than 0.25, indicating shear instability in the ocean. At depths of  
 312 500-800 m,  $Ri$  also drops below 0.25, likely due to sufficiently weak stratification or vertical oscillations of the reflectors.  
 313 Additionally, the low signal-to-noise ratio in the deep water may also cause some errors.

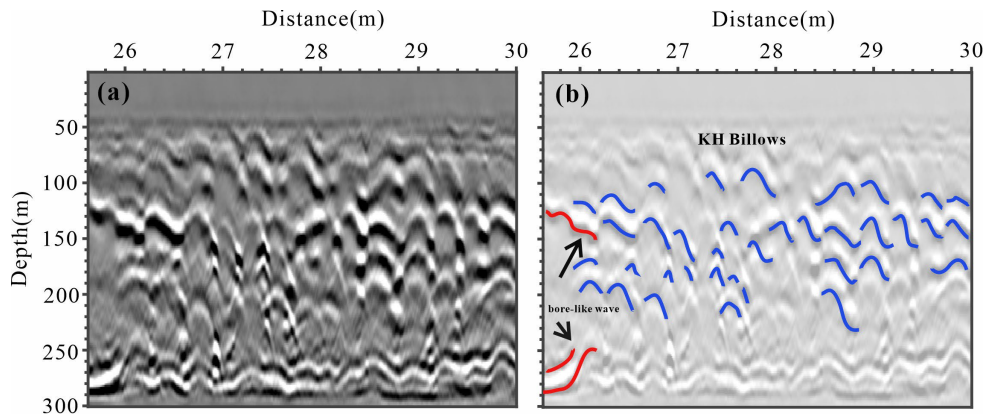
314 K-H billows typically have an alternating braid-core structure, resembling "cat's eyes" and often exhibit small-scale  
 315 secondary instabilities on the braids (Thorpe, 1987). It is rare to see a complete "cat's eye" structure in the ocean; more often,  
 316 they manifest as the braid structures (Acabado et al., 2021; Chen et al., 2022; Chang et al., 2016; Geyer et al., 2010; Tu et al.,  
 317 2022). In Fig. 10, we can clearly observe KH billows caused by instability on the seismic transect. These billows are distributed  
 318 in the range of 26.5-30 km, mostly inclined inverted "S" shapes, i.e., braid structures (blue solid line in Fig. 10b). Table 2 lists  
 319 the details of some clearer billows: The amplitude is about 26-45 m, and the apparent wavelength is about 212-322 m. When

320 less than 28 km, the reflectors are shorter and incomplete, which may be caused by high shear. At greater than 28 km, the braid  
 321 structure is more evident. We can also extract the waveform information of KH billows from the seismic transect to estimate  
 322 Aspect ratio and minimum Richardson number  $Ri_o$  ( $= 0.25-0.39 h_{es}/\lambda$ , where  $h_{es}/\lambda$  is the ratio of wave height to wavelength),  
 323 thereby roughly assessing the instability. Through calculation, the average Aspect ratio is 0.13, and the  $Ri_o$  averages 0.2,  
 324 indicating conditions favorable for the occurrence of instability, which is consistent with the result of Figure 9. Tu et al. (2022)  
 325 estimated the turbulence dissipation rate  $\varepsilon$  ( $= C^{-2} h_{es}^2 N^3$ , where  $C \approx 12.5$ ) caused by KH billows using waveforms extracted  
 326 from acoustic data. Here, we can apply this parameterization scheme to estimate the turbulence dissipation rate in the transect.  
 327 The average dissipation rate  $\varepsilon = 10^{-6.7 \pm 0.1} m^2 s^{-3}$ , indicating that shear instability plays an important role in the energy  
 328 dissipation of ISWs.  
 329



330  
 331 **Figure 9 Geostrophic shear and Richardson number ( $Ri$ ) estimated from seismic data. (a) Geostrophic shear (in color) are**  
 332 **superimposed on the seismic images (in gray). (b) represents the mean vertical shear (blue line) and standard deviation (blue shading)**  
 333 **for the depth of 50-200 m of the seismic transect. (c) Richardson number ( $Ri$ ) are superimposed on the seismic images (in gray).**

334



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338

**Figure 10 Structure of KH billows. (a) Enlarged view of the transect containing KH billows; (b) Interpretative diagram of the reflectors in the seismic image, where blue solid lines represent KH billows, and red solid lines represent the rear wing of bore-like waves, whose waveform is not very complete due to instability.**

339

340

**Table 2 Statistics of KH Billow Waveforms.**

Billows#	Amplitude (m)	Wavelength (m)	Distance (m)	Depth (m)	Aspect Ratio	$Ri_o$
1	31.6±0.8	212.6±2.9	29.6	128.5	0.15	0.19
2	41.8±0.1	241.0±8.7	29.4	130.3	0.17	0.18
3	33.3±0.3	258.6±3.6	29.1	135.3	0.13	0.20
4	31.7±0.3	328.1±7.7	28.7	139.9	0.10	0.21
5	34.0±0.7	221.0±14.2	28.4	134.7	0.15	0.19
6	34.7±0.3	318.9±24.8	28.1	139.2	0.11	0.21
7	32.8±0.3	284.2±12.3	28.9	172.4	0.12	0.20
8	33.9±1.7	322.1±4.1	28.7	166.9	0.11	0.21
9	26.2±0.5	236.9±0.5	28.3	157.0	0.11	0.21
10	44.9±1.4	373.1±1.4	28.7	200.7	0.13	0.20
11	28.8±1.4	234.5±12.3	29.1	195.6	0.12	0.20
12	43.2±2.5	303.2±29.6	27.7	211.6	0.14	0.19

341

**Billows#, the number of KH billows; Amplitude, the amplitude (wave height) of the billows; Wavelength, the wavelength of the**

342 **billows; Distance, the horizontal distance where the billows are located; Depth, the depth of the seawater where the billows are**  
343 **located.**

344

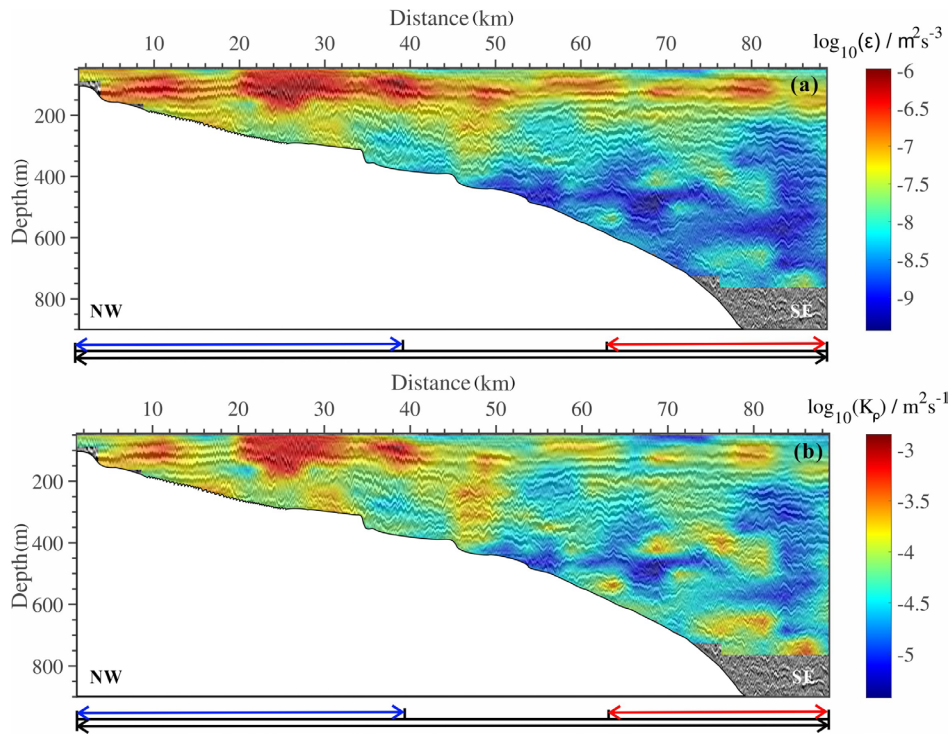
### 345 3.5 Diapycnal diffusivity maps

346 We used the method mentioned in Sect 2.4 to interpolate and smooth the average dissipation rate (diapycnal mixing)  
347 calculated for each window with the values from the surrounding windows. We obtain Fig. 11 by overlaying the results on the  
348 seismic image. This not only makes it easier to grasp the spatial distribution of the calculation but also to understand the  
349 correspondence between the results and the seismic image. Figure 12 is a histogram of the dissipation rate (diapycnal mixing)  
350 over different horizontal ranges. An analysis of the fitting errors is detailed in Appendix A.

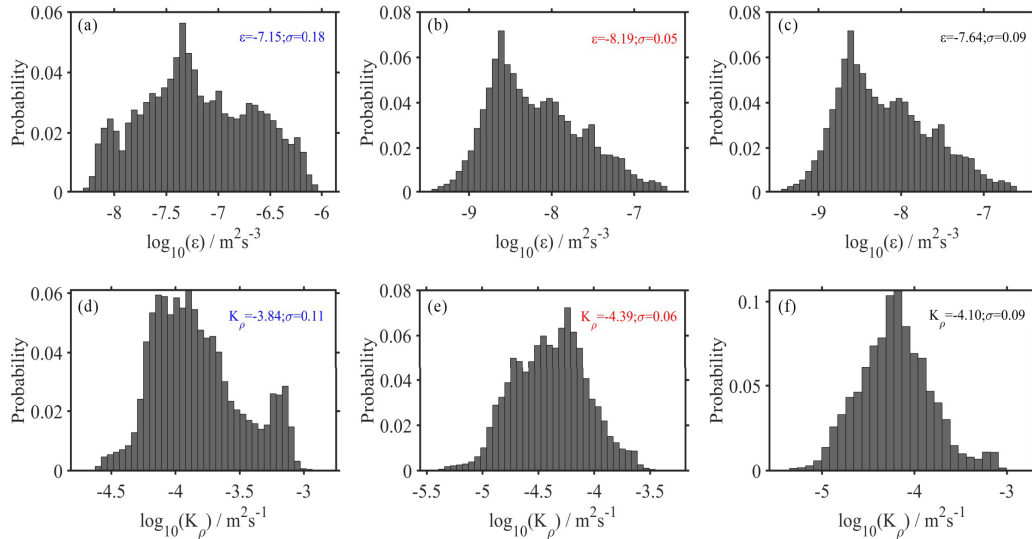
351 According to Fig. 11 and Fig. 12, we can find that the mean diapycnal mixing is  $10^{-4.10 \pm 0.09} \text{ m}^2\text{s}^{-1}$  (-4.10 is the average  
352 and 0.09 is the standard deviation), which is an order of magnitude greater than the average in the open ocean ( $10^{-5} \text{ m}^2\text{s}^{-1}$ ). In  
353 the upper layer of seawater (above 200 m), especially near the thermocline, the dissipation rate (diapycnal mixing) is high,  
354 while below 200 m, the dissipation rate (diapycnal mixing) decreases with increasing depth. However, the dissipation rate  
355 (diapycnal mixing) is also high near the seabed. The calculated diapycnal mixing for the 1.25-38.75 km range is  $10^{-3.84 \pm 0.11}$   
356  $\text{m}^2\text{s}^{-1}$ , which is 3.5 times greater than the calculation for the 62.5-88.75 km range ( $10^{-4.39} \text{ m}^2\text{s}^{-1}$ ). As we discussed earlier, a  
357 series of shoaling events occurred within the 1.25-38.75 km range: the HIWs can enhance the diapycnal mixing, averaging  $10^{-4}$   
358  $\text{m}^2\text{s}^{-1}$ ; In the range of 20-30 km, a marked increase in  $K_p$  suggests that enhanced mixing is associated with the breaking of  
359 IWs caused by the strong shear. The red patches have a slightly deeper color than the other area, indicating that this is a high  
360 mixing area in the transect, up to  $10^{-3} \text{ m}^2\text{s}^{-1}$ . And when away from the topography (62.5-88.75 km), the mixing weakens.

361 The results suggest that HIWs generated from the shoaling ISWs can also enhance turbulent dissipation and mixing. And  
362 the elevated velocity shear during the shoaling may cause the occurrence of shear instability. The results of numerical  
363 simulations by Bai et al. (2019) support this point, and they indicate that the shear within the thermocline intensifies  
364 significantly when ISWs fission into HIWs, and the shear instability by strong shear may cause the enhanced diapycnal mixing.  
365 Therefore, the fission of ISWs into HIWs is one of the key processes for the energy dissipation of these waves.





366  
 367 **Figure 11** Distribution map of mixing parameters. (a) Spatial distribution of turbulence dissipation rate; (b) Spatial distribution of  
 368 diapycnal mixing, where the blue arrow represents the horizontal range of 1.25-38.75 km, the red arrow represents the horizontal  
 369 range of 62.5-88.75 km, and the black arrow represents the horizontal range of 1.25-88.75 km.



371  
 372 **Figure 12** (a) and (d) are distribution histograms of the dissipation rate and diapycnal mixing, respectively, for the range indicated  
 373 by the blue arrow (1.25-38.75 km) in Fig. 11; (b) and (e) are distribution histograms of the dissipation rate and diapycnal mixing,  
 374 respectively, for the range indicated by the red arrow (62.5-88.75 km) in Fig. 11; (c) and (f) are distribution histograms of the

375 dissipation rate and diapycnal mixing, respectively, for the range indicated by the black arrow (1.25-88.75 km) in Fig. 11.

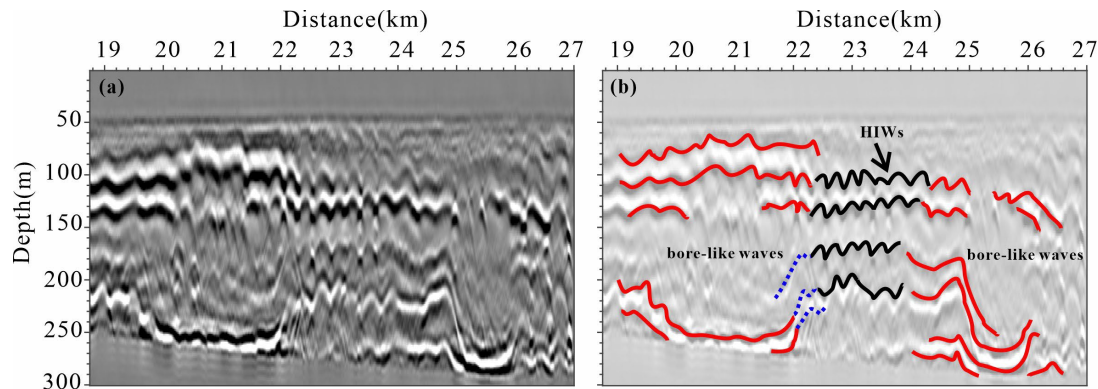
376

## 377 4 Discussion

### 378 4.1 Internal structure of the bore-like nonlinear waves

379 As shown in Figure 13, we observe the nonlinear waves (NIWs) with complex structures and their shapes like bores. The  
380 upper boundary of these two "convex" structure reflectors is convex upwards, while the lower boundary is concave downwards.  
381 Compared to the strong reflectors at the boundaries, the core is transparent or weakly reflective, with a central depth of around  
382 180 m and a thickness of about 150 m, indicating well-mixed water.

383 At around 20 km, the reflectors no longer remain horizontal. The reflectors of the upper boundary gradually become  
384 convex upwards with about a 70 m displacement, and the lower boundary becomes concave downwards with about a 50 m  
385 displacement. In contrast, the upper and lower boundaries at the wave rear (around 22 km) have opposite displacements,  
386 forming the bore-like NIW. In 22-24 km, the seawater at the rear gradually returns to its original stratification due to the  
387 hydraulic jump, and HIWs reappear. However, another similar-shaped NIW appears about 2 km behind. The upper boundary  
388 of this wave is more diffuse, and the lower boundary layer is concave down to near a depth of 290 m with about a 100 m  
389 displacement. Unlike the former, the reflectors at the rear (~ 26 km) do not return to the original stratification but become  
390 disordered and breaking, representing the enhanced turbulence state in the ocean.



391

392 **Figure 13 Bore-like NIWs. (a) Enlarged view of the 19-27 km transects in Fig. 5; (b) Interpretative diagram of the reflectors in the**  
393 **seismic image, where red solid lines represent the boundaries of NIWs, black solid lines represent HIWs, and blue dashed lines**  
394 **indicate hydraulic jump.**

### 395 4.2 Mode-2 ISWs or shoaling bores?

396 Utilizing field observations and numerical simulations, Scotti et al. (2008) found that strong nonlinear HIWs interacting  
397 strongly with shoaling topography in Massachusetts Bay can result in "Seabed Collision Events". Specifically, they found that

398 under conditions of gentle slope and moderate amplitude, these waves would undergo high-energy collision events. The  
399 waveforms generated in these collisions are remarkably similar to those observed in Figure 13 and are often accompanied by  
400 the occurrence of hydraulic jumps. Thus, we speculate that such complex wave packets may represent a form of internal wave  
401 deformation during "seabed collision events".

402 There are two speculations about what they are exactly. Firstly, shoaling internal waves usually form bores with bottom-  
403 enclosed structures carrying cold water masses that continue to propagate shoreward after breaking (Jones et al., 2020). The  
404 strength and structure of bores are influenced by the background stratification and the depth of the thermocline (if  $(A+z_{th})/h >$   
405  $1/2$ , shoaling internal waves will form bores, where  $A$  is the amplitude of ISW,  $z_{th}$  is the thermocline depth, and  $h$  is the water  
406 depth) (Scotti et al., 2008; Sinnett et al., 2022; Walter et al., 2014). Bores typically begin to dominate the structure of internal  
407 waves near depths less than 50 m, becoming a primary feature of the wave during the shoaling process (McSweeney et al.,  
408 2020). However, the two complex wave packets in this paper are located at a depth of 300 m. Therefore, we believe that  
409 although the structure of the two wave packets is somewhat similar to that of bores, their distribution depth range does not  
410 align with the typical distribution depth of common bores, which are often in shallower regions at the run-up phase. Another  
411 explanation is that they are the high-mode NIWs. It is worth noting that the apparent wave width range of these two NIWs is  
412 1.5-2 km, with an amplitude range of 50-100 m, which is clearly larger than the size of the mode-2 internal solitary waves  
413 found by Yang et al. (2010) in the South China Sea shelf and slope areas (with amplitudes of 20-30 m and an average time  
414 scale of 6.9-8 minutes). According to simulations by Brandt and Shipley. (2014), we think that they might belong to  
415 exceptionally large-amplitude internal solitary waves.

#### 416 4.3 Dissipation mechanism of ISWs in the South China Sea

417 Previous extensive studies have shown that shoaling ISWs enhance the mixing of seawater (Gong et al., 2021b; Moum et  
418 al., 2007). Lamb (2014) summarized four scenarios in which shoaling ISWs enhance mixing: 1) Vertical shear caused by waves  
419 leads to shear instability, usually occurring in areas where  $Ri < 0.25$ ; 2) Convective instability occurs when the velocity of the  
420 water mass exceeds the phase velocity, often accompanied by enclosed vortex cores, within which the water is unstable and in  
421 a turbulent state; 3) *Instability near the seabed boundary layer causes sediment resuspension and transport*; 4) Direct breaking  
422 occurs when encountering steep terrain during the shoaling process.

423 In this study, our observational results demonstrate a new energy cascade route from shoaling ISWs to turbulence,  
424 deepening our understanding of the energy dissipation process of ISWs and their roles in enhanced mixing in the northern SCS.  
425 We find that the fission of ISWs into HIWs on the continental slope or shelf is also an important pathway for the tidal energy  
426 cascade. And the HIWs generated during the shoaling can also enhance turbulent dissipation and mixing, on average causing  
427 diapycnal mixing of the order of  $10^{-4} \text{ m s}^{-1}$ . The results are consistent with previous research by Bai et al. (2013, 2019).

#### 428 5 Conclusions

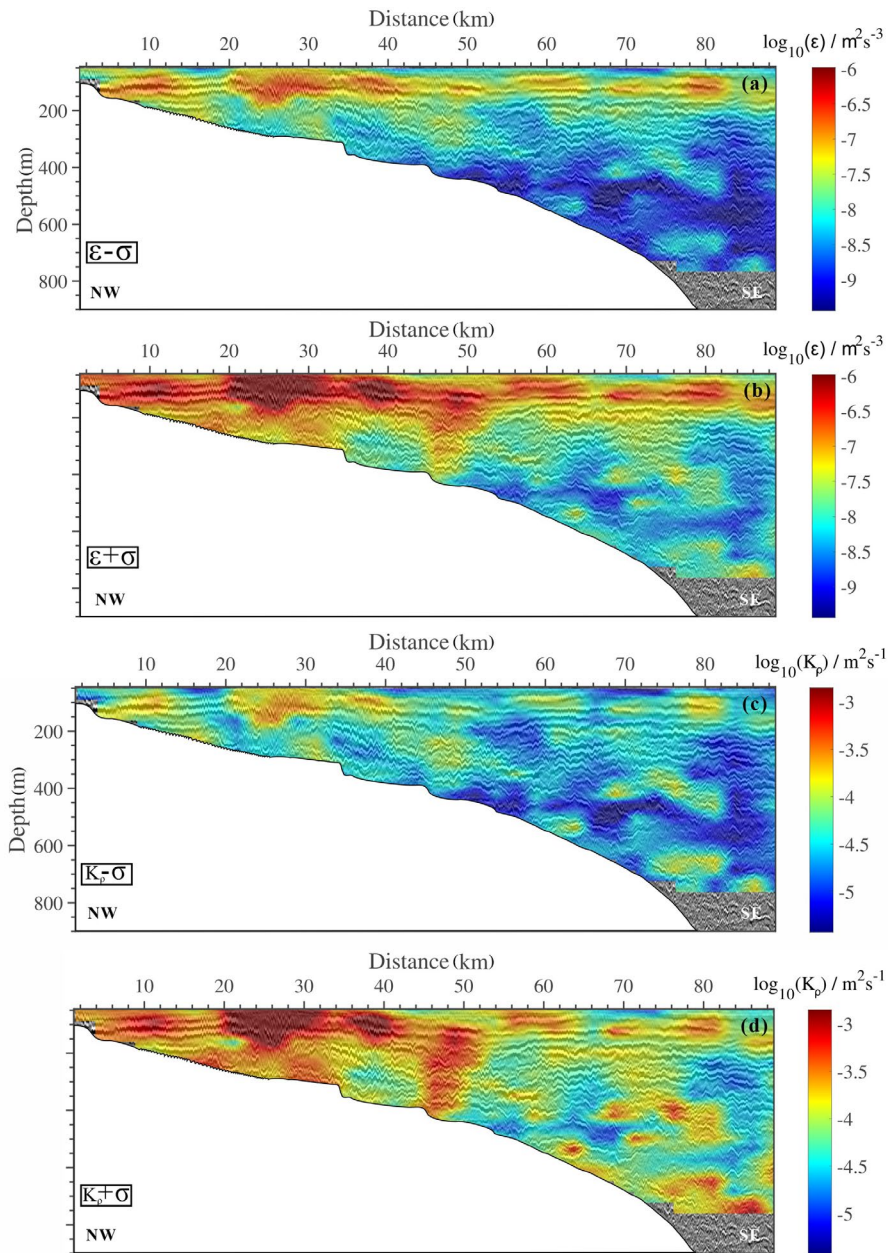
429 In this study, we employed seismic methods to investigate the shoaling evolution and energy dissipation mechanisms of  
430 ISWs. We observed HIWs along seismic line 25, which were primarily distributed in the depth range of 79-184 m. Their  
431 amplitudes are  $O(10\text{ m})$ , with half-widths ranging from 154 to 240 m. The MODIS image revealed that ISWs from the Luzon  
432 Strait typically appeared as wave packets near Line 25 during the data collection period, with tidal models indicating these  
433 waves originated during the neap tide. Fission generally applies to gentle slopes with low  $\xi_{in}$ . Calculations show that the  
434 seabed slope corresponding to Line 25 is  $S = 0.018 (\leq 0.05)$ , with  $\xi_{in}$  ranging from 0.03 to 0.17. Therefore, we infer that HIWs  
435 are products of the fission of shoaling ISWs. Within the 20-30 km range of the transect, there are two complex bore-like  
436 nonlinear waves, resulting from strong interactions between nonlinear waves and the topography. At the trailing edge of the  
437 second nonlinear wave, the reflectors appeared disordered and even fragmented, with KH billows forming. The amplitudes of  
438 these billows ranged from 26 to 45 m, with wavelengths between 212 and 322 m. By combining seismic and hydrological  
439 data, we estimated the geostrophic shear and Richardson number. We found that within the 20-30 km range, vertical shear  
440 could reach up to  $0.3\text{ s}^{-1}$ , with  $Ri$  less than 0.25, indicating the occurrence of shear instability. Therefore, when shoaling ISWs  
441 undergo fission into HIWs, the enhanced shear within the seawater contributes to shear instability.

442 We used seismic data to estimate the mixing parameters of seawater and found that diapycnal mixing for the 1.25-38.75  
443 km range is  $10^{-3.84 \pm 0.11} \text{m}^2 \text{s}^{-1}$ , which is 3.5 times greater than the calculation for the 62.5-88.75 km range ( $10^{-4.39} \text{m}^2 \text{s}^{-1}$ ). A  
444 series of shoaling events occurred within the 1.25-38.75 km range: the HIWs can enhance the diapycnal mixing, averaging  $10^4$   
445  $\text{m}^2 \text{s}^{-1}$ ; In the range of 20-30 km, a marked increase in  $K_p$  suggests that enhanced mixing associated with breaking of IWs  
446 caused by the strong shear, up to  $10^3 \text{m}^2 \text{s}^{-1}$ . When away from the topography (62.5-88.75 km), the mixing weakens. Our  
447 observational results show a new energy cascade route from shoaling ISWs to turbulence, i.e., the fission of ISWs into HIWs,  
448 which improves our knowledge of ISW energy dissipation and their roles in improved mixing in the northern SCS.

## 449 **Appendix A: Error analysis**

450 Generally speaking, there is a certain error between the fitted spectrum obtained by the least square method and the actual  
451 spectrum, which we represent here using the standard deviation  $\sigma$ . Therefore, we can obtain the upper and lower limits of the  
452 turbulence dissipation rate (diapycnal mixing) (Fig. A1).

453 Additionally, the selection of different parameter values in formulas 3 and 4 will also result in errors in the outcome. We  
454 typically take the 95% confidence interval of the average buoyancy frequency (Fig. 3d), so the error in the diapycnal mixing  
455 in the log domain is about  $0.02 \text{m}^2 \text{s}^{-1}$ . The range of the mixing efficiency  $\Gamma$  is 0.1-0.4 (Mashayek et al., 2017), and here we  
456 take 0.2, causing an error in the diapycnal mixing in the log domain of about  $0.15 \text{m}^2 \text{s}^{-1}$ . The range of  $C_T$  values is 0.3-0.5  
457 (Sreenivasan, 1996), and here we take 0.4, resulting in an error in the diapycnal mixing of about  $0.14\text{-}0.17 \text{m}^2 \text{s}^{-1}$  in the log  
458 domain.



459

460 **Figure A1 (a)  $\epsilon$ - $\sigma$ ; (b)  $\epsilon$ + $\sigma$ ; (c)  $K_p$ - $\sigma$ ; (d)  $K_p$ + $\sigma$ .  $\epsilon$  is the turbulence dissipation rate,  $K_p$  is the diapycnal mixing, and  $\sigma$  is the standard**  
 461 **deviations for fitting spectrum.**

462

463 **Code and data availability.** The seismic data was processed using Seismic Unix developed by the Center for Wave  
 464 Phenomena (CWP) at the Colorado School of Mines. XBT data comes from the National Oceanic and Atmospheric  
 465 Administration's World Ocean Database 2018 (WOD18, <https://www.ncei.noaa.gov/OCL/>). CTD data is sourced from the

466 Climate and Ocean Project and Carbon Hydrographic Data Office (CCHDO, <https://cchdo.ucsd.edu/>). We acknowledge the  
467 use of imagery from the NASA Worldview application (<https://worldview.earthdata.nasa.gov/>), part of the NASA Earth Science  
468 Data and Information System (ESDIS).

469  
470 **Author contribution.** LM completed this paper under the guidance of Professor HS and YG. LM processed and analyzed the  
471 data, and drafted the manuscript. SY, KZ, and ML discussed the results and revised the manuscript.

472  
473 **Competing interests.** The authors declare that they have no conflict of interest.

474  
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483

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