

Improving estimation of a record breaking East Asian dust storm emission with lagged aerosol Ångström Exponent observations

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15 **Abstract.** A record-breaking East Asian dust storm over recent years occurred in March 2021. Ångström Exponent (AE) which, measures the wavelength-dependence of aerosol optical thickness (AOT), is significantly sensitive to large aerosol such as dust. Due to lack of observation during dust storm, and given that the accuracy of satellite-retrieved AE depends on the instrument and retrieval algorithm, it is possible to estimate the dust storm emission using the lagged ground-based AE observations. In this study, the hourly AEs observed by Aerosol Robotic Network (AERONET) are assimilated with the
20 ensemble Kalman smoother (EnKS) and Weather Research and Forecasting model coupled with Chemistry (WRF-Chem) to optimize the simulated dust emission from 14 to 23 March 2021. Our results demonstrate that the additional inclusion of AE can optimize the size distribution of dust emission and the associated total flux depending on the covariance between time-lagged AE observations and the simulated dust emission in each bin. Compared to assimilating only AOT, the additional inclusion of AE induces the extra 17% (61%) improvement of root mean square error (RMSE) between the simulated AOT
25 (AE) and the independent Skynet Observation NETwork (SONET) observations. The temporal variation of simulated AOT and AE can be both improved through assimilating additional AE information. The assimilation of AOT and AE also makes the magnitude and variations of aerosol vertical extinctions more comparable to the independent Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) observations both in the westward and eastward pathways of dust transport. The optimized dust emission in Gobi desert during this period is estimated as 52.63 Tg and reaches the peak value of 3837 kt h⁻¹
30 at 07:00 UTC on 14 March.

1 Introduction

Mineral dust is the most abundant atmospheric aerosol component in terms of aerosol dry mass in the atmosphere. It affects the climate system by scattering and absorbing longwave and shortwave radiation and also contributes to the formation of cloud condensation nuclei (CCN) and ice-nucleating particles (INP) (Huang et al., 2006; IPCC AR6, Kok et al., 2018; Liu et al., 2020). Dust also carries organic matter and transports iron to the ocean that is vital to ocean productivity and ocean-atmosphere CO₂ exchange, inducing the impacts on the cycles of dust and carbon (Shao et al., 2011). Dust deposition on snow surface can influence the snow albedo and modify the water cycle and energy budget (Wu et al., 2018; Kang et al., 2019; Wang et al., 2020). Moreover, severe dust storm can induce air pollution and affect human health (Chen et al., 2020).

East Asia, including the Taklimakan and Gobi deserts, is the world's second largest dust source, which occupies approximately 40% of the dust emissions globally (Satake et al., 2004; Kok et al., 2021) and accounts for 88% of the dust in China and neighbouring seas (Han et al., 2022). Although the dust activities in East Asia have declined recently (Wu et al., 2022), the unexpected extreme dust storm event occurred during 14 to 23 March 2021 reawakened widespread concern. Numerical model are important tools to study severe dust storm, and the dust emission is a significant quantity to characterize dust activity. However, due to insufficient knowledge of the actual dust mechanisms, more than ten fold diversity exists in simulated East Asian dust emissions among different models (Uno et al., 2006, Kok et al., 2021), indicating dust emission is a significantly uncertain process in dust simulation. Data assimilation, which feeds the observation information into numerical model, provides a top-down method for the optimization of the estimates of dust emissions. Yumimoto et al. (2008) assimilated the dust extinction coefficients derived from ground-based lidar network using four-dimensional variational (4D-Var) method, increasing the East Asian dust emissions by approximately 10 times. Sekiyama et al. (2010) developed an Ensemble Kalman Filter (EnKF) data assimilation system to jointly correct the global dust emissions and aerosol mixing ratios by assimilating the vertical observations from the Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO). Wang et al. (2012) constrained the amount and location of dust emissions in Taklimakan and Gobi deserts with the GEOS-Chem adjoint model by assimilating aerosol optical thicknesses (AOTs) from Moderate Resolution Imaging Spectroradiometer (MODIS). Based on the ensemble Kalman smoother technique, Schutgens et al. (2012) assimilated observations from AEROSOL ROBOTIC NETWORK (AERONET) and MODIS to estimate emissions for dust, sea salt, and carbonaceous aerosols. Yumimoto and Takemura (2015) performed inverse modeling of Asian dust with four-dimensional variational (4D-Var) data assimilation system and MODIS-retrieved AOT over ocean. Escribano et al. (2016) estimated dust emission and reduced their uncertainty over the Sahara desert and the Arabian Peninsula by assimilating MODIS AOT retrievals. Di Tomaso et al. (2017) used the four dimensional local ensemble transform Kalman filter (LETKF) to assimilate MODIS Dark Target and Deep Blue AOTs for improving dust analyses and forecasts in the global domain. AOTs with high frequency from Himawari-8 geostationary satellite have been recently used for the optimization of dust storm emission with a reduced tangent linearization 4D-Var technique (Jin et al., 2019).

Except AOT, Ångström Exponent (AE) which measures the wavelength-dependence of AOT and is significantly sensitive to size of aerosol particle, may have a positive impact on data assimilation (Tsikerdekis et al., 2022, 2023). The estimated emission may be misrepresented by not including observations related to size (Tsikerdekis et al., 2021). However, most of the abovementioned studies have estimated new dust emissions based on the assimilation of AOT, while few studies have explored the potential benefits of aerosol size information like AE observations on improving the estimate of dust emission. Therefore, how will the assimilation of the AE observations affects the optimization of the dust emission? It becomes an important scientific question. Due to the accuracy of satellite-retrieved AE depends on the instrument and retrieval algorithm, the ground-based AE is better than satellite-based AE and can be more useful for optimizing dust emissions. Therefore, the AOT and AE observations from the ground-based AEROSOL ROBOTIC NETWORK (AERONET) are assimilated to investigate the sensitivity of dust emission to observed size information in this study. The additional benefit of aerosol size information in estimating dust emission is explored by only assimilating AOT and simultaneously assimilating AOT and AE. The experiments are conducted with the Ensemble Kalman smoother (EnKS) assimilation framework (Dai et al., 2019) to constrain the extreme dust storm emission over East Asia from 14 to 23 March 2021. The Sun-Sky Radiometer Observation Network (SONET) and CALIPSO observations are used for independent validation. In Sect. 2, we describe the assimilated and independent validation observations. Our dust emission optimization system and experimental design are presented in Sect. 3. Section 4 presents the optimized emission results and the validations by multi-sensor independent observations. The conclusions are given in Sect. 5.

2 Observation Data

2.1 Assimilated AERONET observations

AEROSOL ROBOTIC NETWORK (AERONET) is a federation of ground-based remote sensing aerosol network that collects aerosol optical observations with sun photometers from various stations in global (<https://aeronet.gsfc.nasa.gov>) (Holben et al., 1998; Giles et al., 2019). Version 3 AOT data are divided into three levels according to the data quality procedures: Level-1 (unscreened), Level-1.5 (cloud-screened and quality-controlled), and Level-2 (quality-assured). In this study, the version 3 level 2 AOT at 550 nm and AE at 440-870 nm from AERONET are assimilated. AOT (τ) at 550 nm is obtained by logarithmic interpolation of the ones at 440 nm and 675 nm. AE (α) at 440-870 nm is calculated with the AOT at 440 nm and 870 nm from the following equation: $\alpha_{440-870} = -\ln(\tau_{870}/\tau_{440})/\ln(870/440)$. To ensure the accuracy, AE value is considered to be valid only when the AOTs at 440 nm and 870 nm both exceeds 0.05 (Giles et al., 2019). The instantaneous observations are averaged every 1 hour, centering on the assimilation time slot. The observation error (ϵ) attributed to this averaged observation is calculated by a representation error (ϵ_r) and an instrument error (ϵ_o) as $\epsilon^2 = \epsilon_r^2 + \epsilon_o^2$. Due to the representation error is related to the WRF-Chem grid resolution, the representation error in AERONET AOT and AE is calculated depending on the AOT and AE temporal variability of AERONET and WRF-Chem with 45 km horizontal resolution (Schutgens et al., 2010). By averaging results at all AERONET sites in March 2021, the relative AOT temporal

variations of AERONET and WRF-Chem in 1 h interval are 0.11τ and 0.1τ , while the AE temporal variations of AERONET and WRF-Chem in 1 h interval are 0.05 and 0.02, respectively. Therefore, the representation errors in AERONET AOT (τ) and AE in the 1 h interval are $\epsilon_{\tau} = 0.01\tau$ and $\epsilon_r = 0.03$, respectively. The instrument error in AOT is defined as 0.015 and the instrument error of AE is estimated by propagating the instrument error in AOT at 440 and 870 nm (Schutgens et al., 2010).

To minimize the influences of anthropogenic emissions, only the AERONET AOT and AE dominated by dust are assimilated to optimize the dust emissions, which are chosen with AE at 440-870 nm less than 0.4 (Huneeus et al., 2011). In addition, due to the uncertainties of modelled covariance between dust emission and aerosol optical properties increase with the distance far away from the source region, only the observations within 2190 km (3.65 times of the localization length) of the East Asian dust source region are used for data assimilation. As shown in Fig. 1(a), there are 5 AERONET sites with available observations from 14 March to 23 March 2021 for data assimilation, including 4 sites named as Beijing-CAMS (39.93°N, 116.32°E), Beijing (39.98°N, 116.38°E), Beijing_PKU (39.99°N, 116.31°E), and Beijing_RADI (40.00°N, 116.38°E) in the downwind area and a site near the dust source region named as Dalanzadgad (43.58°N, 104.42°E). The assimilated AOT and AE values at the AERONET sites are also given in Fig. 1(b,d). For Dalanzadgad site, the AOT values from 14 March to 17 March 2021 are significantly higher than those from 18 March to 23 March, while AE values show the opposite features.

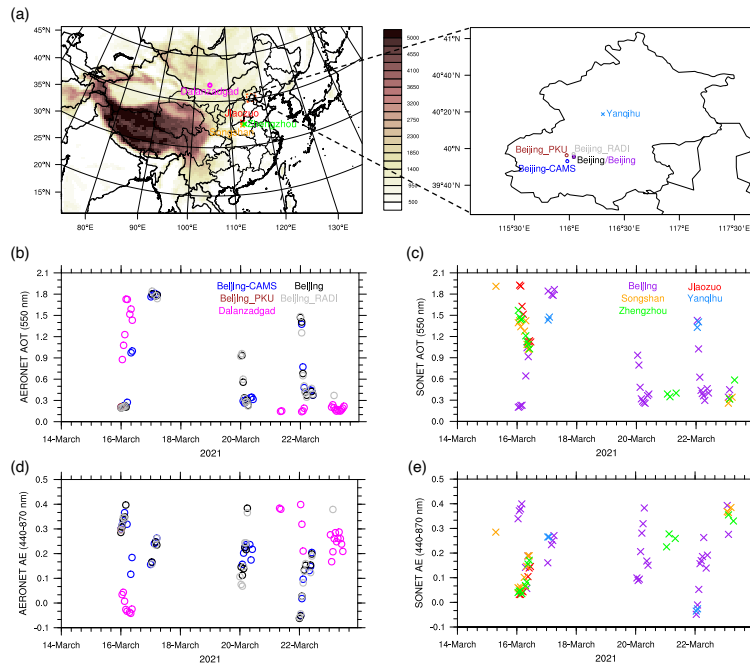


Figure 1. Locations of selected Aerosol Robotic Network (AERONET) sites for data assimilation and Skynet Observation NETWORK (SONET) sites for independent validation in this study (a). AERONET sites are illustrated by circle and SONET sites

115 are illustrated by cross. The hourly values of aerosol optical thickness (AOTs) at 550 nm and Angstrom Exponent (AEs) at 440-870
nm at AERONET (b,d) and SONET sites (c,e) during 14-23 March 2021.

2.2 Independent SONET and CALIOP observations

Sun-Sky Radiometer Observation Network (SONET) is a ground-based aerosol network employing Cimel radiometer and
multiwavelength polarization measurement to provide long-term columnar atmospheric aerosol properties over China
120 (www.sonet.ac.cn) (Li et al., 2018). The aerosol optical related products including AOT, AE, fine-mode fraction (FMF) are
graded into three levels: Level-1 (no triplet), Level-1.5 (cloudy), and Level-2 (no cloud). In this study, the level 2 AOT at
550 nm and AE at 440-870 nm from SONET are used for independent validation and the instantaneous observations are also
averaged every hour to generate hourly AOT and AE datasets. As shown in Fig. 1(a), there are 5 SONET sites with available
observations from 14 March to 23 March 2021, including: Yanqihu (40.40°N, 116.67°E), Beijing (40.00°N, 116.37°E),
125 Jiaozuo (35.18°N, 113.20°E), Songshan (34.53°N, 113.09°E), and Zhengzhou (34.70°N, 113.66°E). The AOT and AE
values at the SONET sites are also given in Fig. 1 (c,e). Similar with Dalanzadgad site, Jiaozuo, Songshan, and Zhengzhou
sites experience a stronger dust process from 14 March to 17 March 2021 with higher AOTs and lower AEs.

Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP), aboard the CALIPSO satellite, is a dual-wavelength
polarization lidar that performs observations of vertical structures of aerosols and clouds on the global scale ([http://www-
130 calipso.larc.nasa.gov](http://www-calipso.larc.nasa.gov)) (Winker et al., 2007). In this study, the aerosol extinction coefficients at 532 nm in the CALIPSO
lidar level 2 (L2) version 4.20 aerosol profile products over the altitude range below 12 km are also used for evaluation. The
CALIPSO lidar level 2 version 4.20 Vertical Feature Mask (VFM) products, which include the feature types and subtype
information, are used for aerosol discrimination (Omar et al., 2009; Cheng et al., 2019).

3 Model and data assimilation methodology

135 3.1 Forward model

Weather Research and Forecasting model coupled with Chemistry (WRF-Chem) version 4.4, which served as the forward
model, is configured with the domain covering China with 45 km horizontal resolution and 28 vertical levels. The Goddard
Global Ozone Chemistry Aerosol Radiation and Transport (GOCART) aerosol scheme (chem_opt =300) is adopted to
simulate the aerosols including both the dust and non-dust species (Chin et al., 2000, 2002). The Air Force Weather Agency
140 (AFWA) dust emission scheme (dust_opt=3) is chosen for dust simulation. The mass mixing ratio of main aerosol
components, including dust, sea salt, organic carbon, black carbon, and sulfate, are predicted. Other main selected physics
are identical to those of Dai et al. (2019). To match the characteristics of East Asian dust, the fractions of dust emissions in
AFWA scheme are changed to 0.034, 0.187, 0.327, 0.163, and 0.309 for 0.2-2 μm (bin 1), 2-3.6 μm (bin 2), 3.6-6 μm (bin 3),
6-12 μm (bin 4), and 12-20 μm (bin 5) dust size bins in diameter (Su and Fung, 2015). To reduce the underestimation of dust
145 emission in AFWA scheme and start from a relatively unbiased simulation, the adjustable dust emission factor is calibrated

and selected as 21 based on the AERONET-observed AOT and AE. This rescaling of dust emission can benefit the data assimilation since the Kalman filter assumes that the model is unbiased (Tsikerdekis et al., 2021). The aerosol optical properties are calculated with the Mie parameterization using the ‘‘Aerosol chemical to Aerosol Optical Properties’’ module (Ghan and Zaveri, 2007, Barnard et al., 2010), which is based on the sectional approach. The 8 dust size bins in Mie subroutine are as same as the Model for Simulating Aerosol Interactions and Chemistry (MOSAIC) module. The calculation of the dust optical properties is improved with three corrections: (1) remap the fractions of AFWA bin 1 dust in 0.2-2 μm into Mie calculation bins as Ukhov et al. (2021); (2) redistribute fractions of the dust mass based on the assumption that bin concentration is a function of natural logarithm radius as Ukhov et al. (2021); (3) increase the 8 dust size bins in Mie subroutine to 9 as 0.039-0.078, 0.078-0.156, 0.156-0.312, 0.312-0.625, 0.625-1.25, 1.25-2.5, 2.5-5.0, 5.0-10.0 μm , and 10-20 μm to distribute the AFWA bin 5 dust in 12-20 μm into bins for Mie calculation. To compare with AERONET observed aerosol optical properties, the simulated ones are calculated assuming that the particles are spherical and internally mixed with all the simulated aerosol components (Barnard et al., 2010). The initial and lateral boundary meteorological conditions are from the NCEP Final (FNL) analysis. To reduce the uncertainties associated with the meteorological fields, the predicted wind (u , v), temperature (t), and specific humidity (q) by the WRF dynamical core are also nudged to the NCEP FNL analysis every 6 h for all layers (Dai et al., 2018).

3.2 Data assimilation framework

The adopted assimilation system, integrating measurements with model simulations, is based on the Ensemble Kalman Smoother (EnKS) with WRF-Chem (Dai et al., 2019). The Kalman Smoother is in essence a Kalman Filter that iteratively estimates emissions (Schutgens et al., 2012). As shown in Fig. S1, based on the EnKS, the dust emission of WRF-Chem ensemble is optimized every 12 h, which corresponding to the assimilation time window of 12 h. Each assimilation cycle advances a time step of 12 h, and the dust emissions for 6 time steps are optimized using the observations in the 6th time step. After each assimilation cycle, the dust emission for the first time step is the final optimized result, which has been optimized 6 times and will no longer be optimized in the next cycle. The finally optimized dust emissions therefore serve as the forced dust emissions for advancing the system one time step, and they provide the initial conditions for the next assimilation cycle. The posterior dust emission \bar{x}^a is obtained from the solution to the Kalman equations, as the following formula:

$$\bar{x}^a = \bar{x}^b + X^b \bar{w}^a, \quad (1)$$

where \bar{x}^b and X^b represent the ensemble mean of prior dust emission and the first guess ensemble perturbation, respectively. The weight matrix \bar{w}^a determines the increment between the analysis and first guess as

$$\bar{w}^a = \tilde{P}^a (Y^b)^T R^{-1} (y^o - \bar{y}^b), \quad (2)$$

where the matrix $\tilde{P}^a (Y^b)^T R^{-1}$ is called the ‘‘Kalman gain’’; R is the observation error covariance matrix; the y^o and \bar{y}^b represent the assimilated hourly AOT and AE observations from AERONET and the first guess of the simulated AOT and AE observations averaged over the ensemble members; the WRF-Chem serves as the observation operator H to relate the

prior dust emission to the first guess of the simulated observations, $\bar{y}^b = H(\bar{x}^b)$; the first guess ensemble perturbation matrix in observation space Y^b is calculated as $y^{b(i)} - \bar{y}^b$, $\{i = 1, 2, \dots, k\}$ with k ensemble members. The analysis error covariance is obtained as

$$\tilde{P}^a = \left[(k-1)I + Y^{bT} R^{-1} Y^b \right]^{-1}, \quad (3)$$

where I is the identity matrix. The analysis ensemble perturbations X^a are obtained as

$$X^a = X^b W^a, \quad (4)$$

whose i th column is $x^a(i) - \bar{x}^a$, $\{i = 1, 2, \dots, k\}$. In this study, the analysis ensemble by adding \bar{x}^a to each of the columns of X^a forms the optimal dust emission for the ensemble forecast to produce the initial conditions for the next analysis. W^a is calculated as

$$W^a = [(k-1)\tilde{P}^a]^{1/2}. \quad (5)$$

This assimilation scheme offers the advantage of selectively assimilating observations for a given grid point by employing localization in the horizontal, vertical, and temporal scales (Hunt et al., 2007; Gaspari and Cohn, 1999; Miyoshi et al., 2007). The horizontal localization factor is calculated as $f(r) = \exp(-r^2/2\sigma^2)$, the factor is truncated at the 3.65 times of the localization length $\sigma = 600$ km in this study, and r is the distance between the observation and the grid centroid. The vertical and time localization are not applied in this study.

3.3 Experimental design

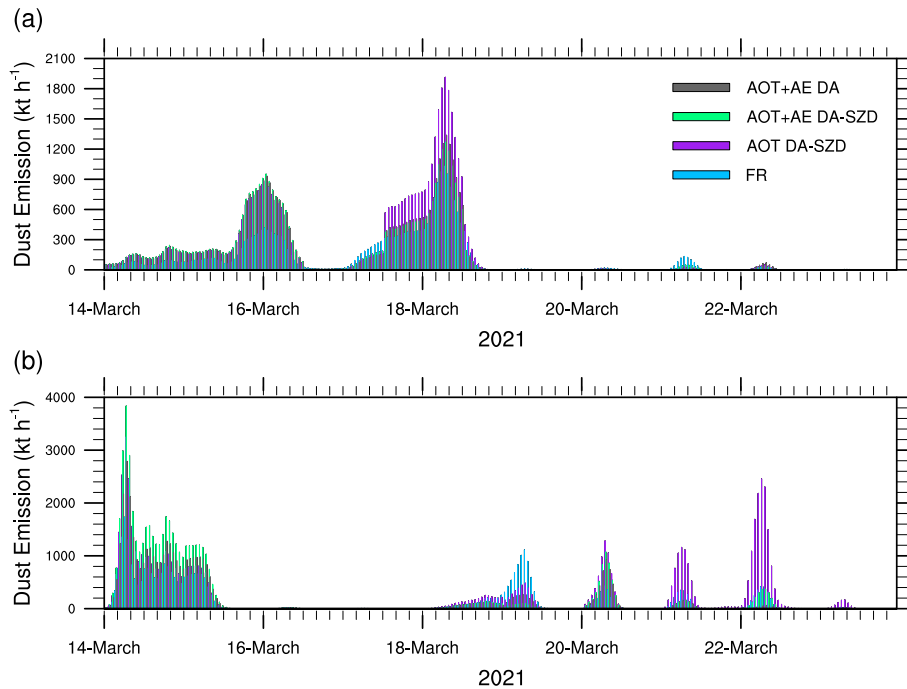
To investigate the influences of AERONET AOT and AE assimilation on the dust emission optimization, three assimilation experiments are conducted from 12:00 UTC on 11 March 2021 to 00:00 UTC on 24 March 2021. Due to WRF-Chem model has uncertainties not only on dust emission but also on dust deposition (Huang et al., 2020) and dust optical properties (Di Biagio et al., 2019) in simulation, two assimilation experiments with perturbation of dust emission and size distribution are conducted. One assimilation experiment named AOT DA-SZD only assimilates AERONET AOT observations, and the other assimilation experiment named AOT+AE DA-SZD assimilates both AERONET AOT and AE observations. 20 ensemble members are generated by perturbing the dust emission in each bin, and the perturbation factor of each bin has a mean of 1 and a spread of 0.6 followed the lognormal distribution. Correlated noise is used across the dust size bins in the perturbation, and the noise correlation decreases with increased difference of the diameter among the bins (Di Tomaso et al., 2017). The ensemble prediction dynamically estimates the covariance between the dust emission in each bin and the aerosol optical properties. The comparison between AOT DA-SZD and AOT+AE DA-SZD experiments shows the effects of the additional AE information on dust emission optimization. The effects of dust emission size distribution perturbation are investigated by one additional assimilation experiment named AOT+AE DA, which is conducted as same as the AOT+AE DA-SZD experiment except the 20 ensemble members are generated by perturbing the dust emission in each bin with same perturbation factor. The results from 12:00 UTC on 11 March 2021 to 23:59 UTC on 13 March 2021 are excluded in the

analysis as the spin-up. The baseline experiment named FR does not assimilate any observations but otherwise share the
210 same configuration with the assimilation experiments.

4 Results

4.1 Dust emission and simulated dust field

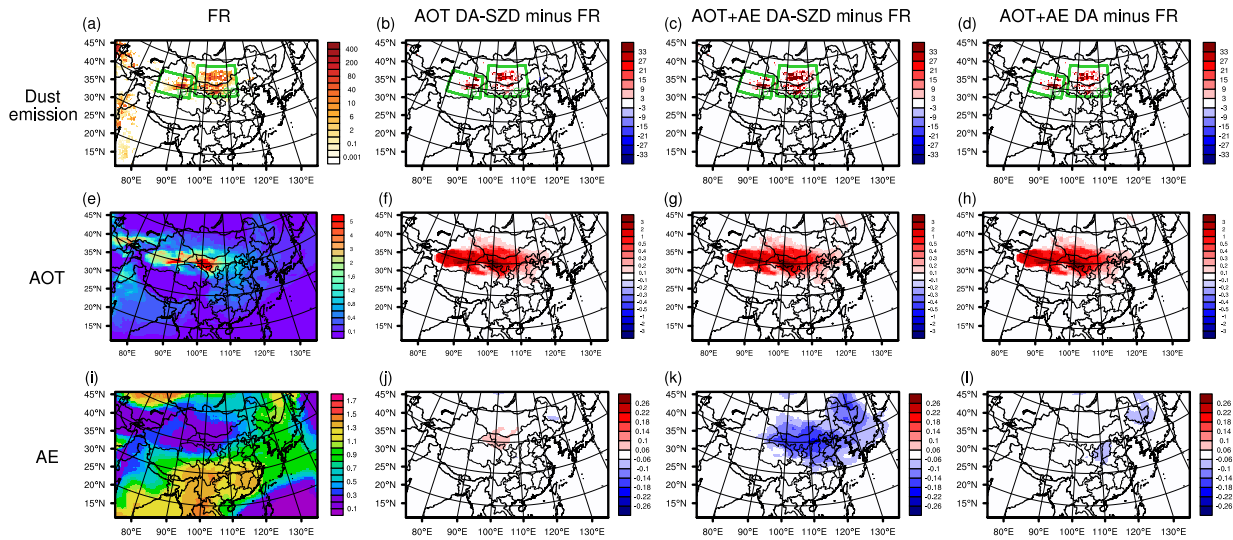
Figure 2 shows the temporal variations of hourly accumulated dust emissions in Taklamakan desert and Gobi desert
simulated by FR, AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments. The first dust storm in Taklamakan
215 desert mainly emits from 00:00 UTC on 14 March to 12:00 UTC on 16 March, and reaches the peak value of 420, 887, 952,
and 933 kt h⁻¹ at 01:00 UTC on 16 March for FR, AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments. The
second dust storm in Taklamakan desert mainly emits from 00:00 UTC on 17 March to 20:00 UTC on 18 March, and
reaches the peak value of 1033, 1914, 1272, and 1338 kt h⁻¹ at 07:00 UTC on 18 March for FR, AOT DA-SZD, AOT+AE
DA-SZD, and AOT+AE DA experiments. There are almost no dust emissions in Taklamakan desert after 20:00 UTC on 18
220 March. With respect to Gobi desert, the strongest dust storm generally emits from 00:00 UTC on 14 March to 18:00 UTC on
15 March, and reaches the peak value of 1735, 3253, 3837, and 2791 kt h⁻¹ at 07:00 UTC on 14 March for FR, AOT DA-
SZD, AOT+AE DA-SZD, and AOT+AE DA experiments. After 18 March, there are five relatively weak dust processes in
Gobi desert. Depend on the temporal variations of dust emission in the Gobi desert, the whole period is divided into two dust
processes: the strong dust storm from 14 March to 17 March 2021 and the weak dust storm from 18 March to 23 March 2021.
225 In general, the total dust emissions during 14-17 (18-23) March 2021 in Gobi desert are 21.45 (13.03), 32.03 (18.84), 44.41
(8.22), and 33.10 Tg (8.57 Tg) for FR, AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments. The dust
emission in each dust bin is given in Table S1.



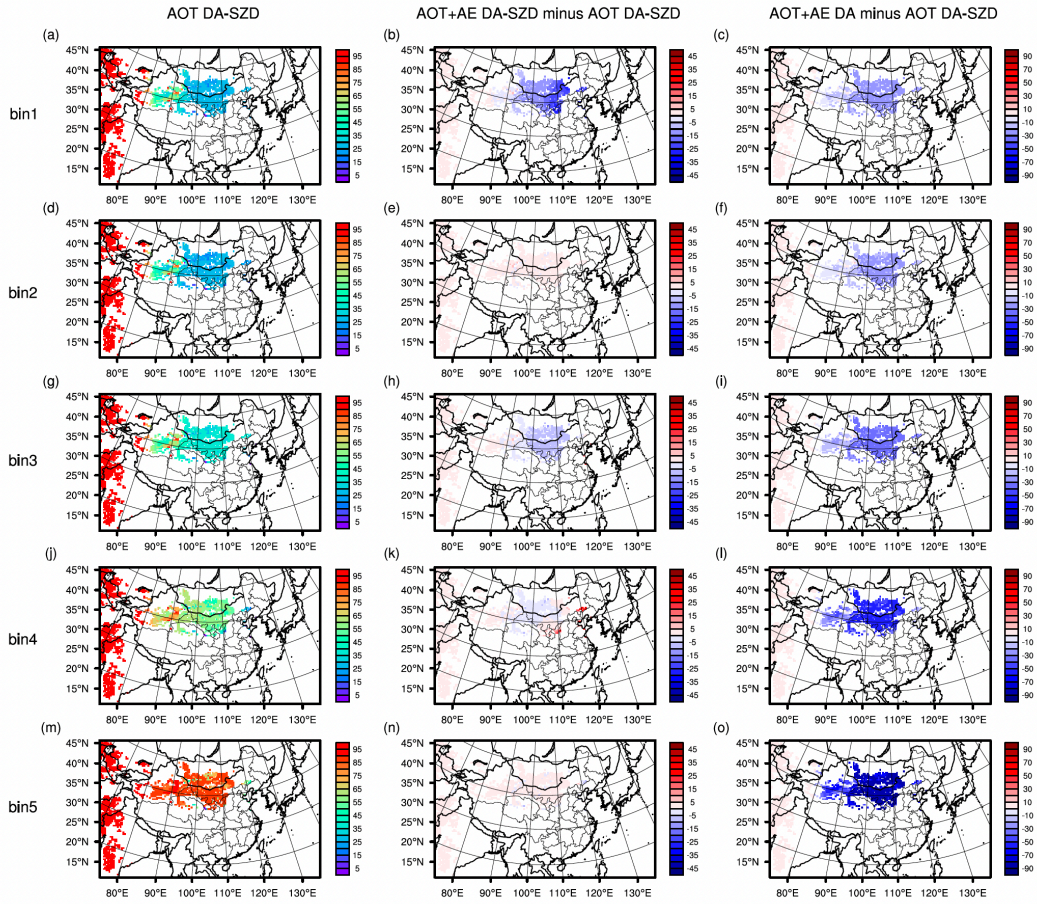
230 **Figure 2. Time series of hourly accumulated dust emissions (units: kt h^{-1}) for FR, AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments during 14-23 March 2021 summed across the Taklimakan desert (a) and Gobi desert (b) (marked in Fig. 3).**

The simulated accumulated dust emission, averaged AOT and AE for the four experiments during 14-17 March 2021 are given in Fig. 3. During 14-17 March 2021, the dust storm in Taklamakan desert and Gobi desert both transports eastward and mainly affects Northern China. The dust emissions in the three assimilation experiments are significantly higher than that in FR experiment, especially AOT+AE DA-SZD experiment. The differences of dust emission between FR and
 235 that in FR experiment, especially AOT+AE DA-SZD experiment. The differences of dust emission between FR and assimilation experiments are mainly concentrated in the Taklamakan desert and Gobi desert, while there is generally no difference in India due to the distance truncation of horizontal localization. The increased dust emissions in all the assimilation experiments induce the higher AOTs over the deserts and the associated downwind regions, whereas only the AEs over the Gobi desert and its downwind region in the AOT+AE DA-SZD experiment are significantly reduced. The
 240 ratios between posterior error of dust emission and the prior one in each dust bin for the three assimilation experiments during 14-17 March 2021 are shown in Fig. 4. It helps visualize how adding AOT and AE for the data assimilation reduces the posterior errors of simulated dust emissions. Due to the same prior error of the dust emission, the difference of ratio represents the difference in posterior error. The ratio with value lower than 100% indicates the assimilation decreasing the uncertainties of the dust emission, and lower value represents higher constraint. The posterior error increases with the dust
 245 size bin increased in AOT DA-SZD experiment, and this is due to the time-lagged AOT observations in the downwind areas are more relevant with the fine-mode dust emission because of the stronger gravity settling of coarse-mode dust and the

higher extinction efficiency of fine-mode dust (Fig. S2 and Fig. S3). The additional AE observations in AOT+AE DA-SZD experiment further adjust the dust emission size distribution over Gobi desert through decreasing dust emission in bin 1 and increasing dust emission in bin 3 (Fig. S4 and Fig. S5), inducing the obviously reductions of the posterior error in bin 1 and bin 3 in Gobi desert. This induces the significant decrease of the AE over the Gobi desert and its downwind region. Although AOT+AE DA experiment includes the AE observations, there is no changes in the dust emission size distribution due to the same perturbation parameter (Fig. S4 and Fig. S5). This leads to the simulated AEs are similar as the FR experiment. It is found that dust emission in bin 3 over Gobi desert in AOT+AE DA-SZD experiment is obviously higher than that in other assimilation experiments (Table S1), however, there is no significant difference in AOT among the three assimilation experiments. This is due to the effect of increased emission in bin 3 on AOT is offset by the effect of decreased emission in bin 1, since the coarse-mode dust is generally removed by gravitational sedimentation (Fig. S2) and the coarse-mode dust has lower extinction efficiency (Fig. S3).



260 **Figure 3. Spatial distributions of accumulated dust emission, aerosol optical thickness (AOT), and Angstrom Exponent (AE) for FR experiment during 14-17 March 2021 (a,e,i). Differences of accumulated dust emission (b,c,d), AOT (f,g,h), and AE (j,k,l) between AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments minus FR experiment. The unit of dust emission is g m^{-2} . The green boxes represent the Gobi desert (GD) and Taklimakan desert (TD).**

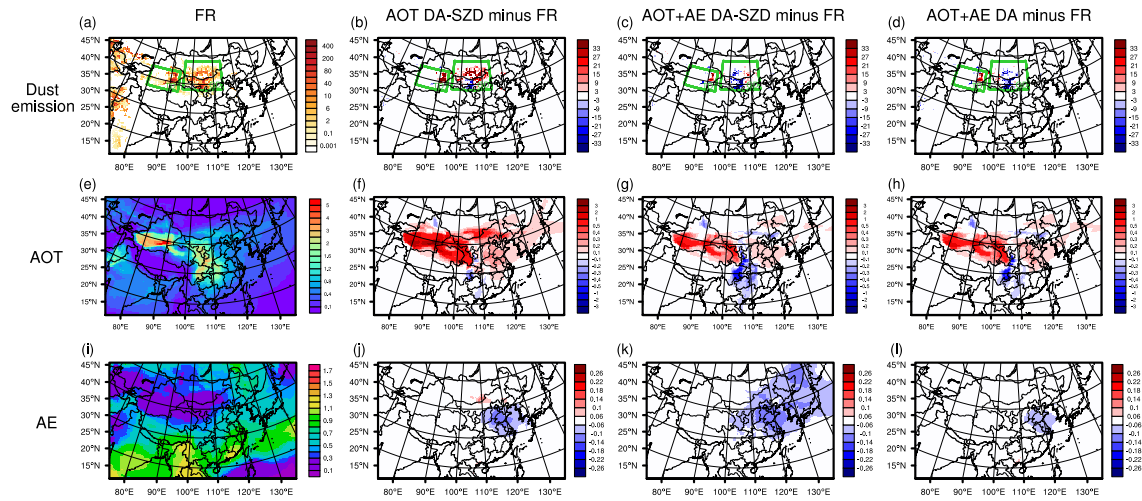


265 **Figure 4. The ratio between posterior error of FR simulated dust emission and prior error of optimized dust emission in the 5 dust size bins for AOT DA-SZD experiment (a,d,g,j,m) during 14-17 March 2021. Differences of the ratios between AOT+AE DA-SZD (b,e,h,k,n) and AOT+AE DA (c,f,i,l,o) experiments minus AOT DA-SZD experiment.**

The simulated accumulated dust emission, averaged AOT, and AE for the four experiments during 18-23 March 2021 are given in Fig. 5. Compared with FR experiment, the dust emission in AOT DA-SZD experiment is increased in Gobi desert of Mongolia and decreased in Gobi desert of China while the dust emission in AOT+AE DA-SZD and AOT+AE DA 270 experiments are decreased in most part of Gobi desert. Due to part of the dust emitted from Gobi desert transports southward, the decreased dust emission induces the lower AOTs in Southern China. It is also found only the AEs over the Gobi desert downwind region in AOT+AE DA-SZD experiment are significantly reduced. As given in Table S1, AOT DA-SZD experiment significantly increases the dust emission in bin 5 and slightly decreases the dust emissions in bin 1, bin 2, bin 3, and bin 4 over Gobi desert, whereas AOT+AE DA-SZD and AOT+AE DA experiments with additional AE observations 275 decrease the dust emission in each bin over Gobi desert. The ratios between posterior error of dust emission and the prior one in the each dust size bin for the three assimilation experiments during 18-23 March 2021 are shown in Fig. 6. It is interesting

that the ratios in AOT DA-SZD and AOT+AE DA-SZD experiments during 18-23 March 2021 are lower than those during 14-17 March 2021. This is due to the lower AOT observations and associated observation errors at dust source site Dalanzadgad during 18-23 March 2021 generate more constraints on dust emission than those during 14-17 March 2021.

280 AOT DA-SZD experiment significantly increases the dust emission in bin 5 and leads to its emission fraction even more than 70% (Fig. S4), however, this phenomenon is not found in AOT+AE DA-SZD experiment. The obviously lower posterior error in bin 5 in AOT+AE DA-SZD experiment indicates the additional AE observations can eliminate the sharply increase of dust emission in bin 5 and constrain the dust emission size distribution with higher confidence.



285 **Figure 5.** Same as Fig. 3 but during 18-23 March 2021.

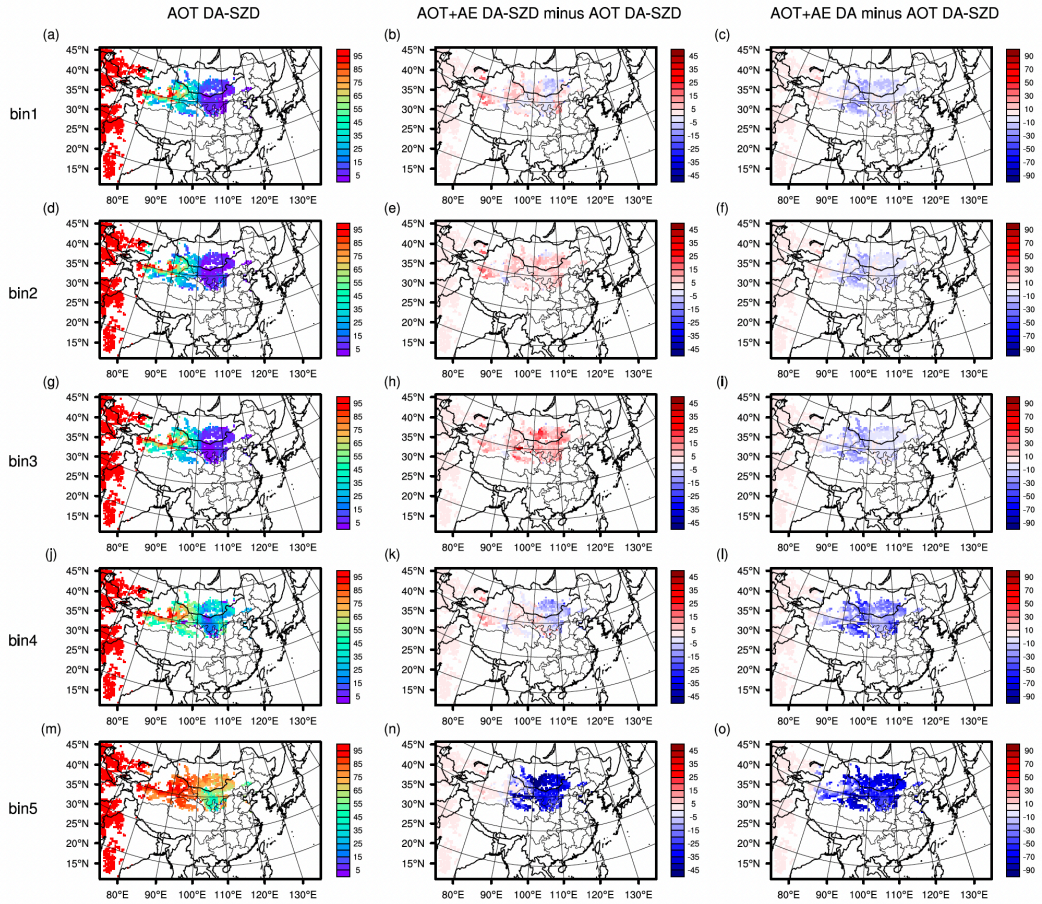


Figure 6. Same as Fig. 4 but during 18-23 March 2021.

4.2 Evaluation of simulated AOTs and AEs

To validate the posterior dust emission, the simulated hourly AOTs and AEs with the prior and posterior dust emissions are compared with the assimilated AERONET observations as the sanity check in Fig. 7. The comparisons with independent SONET observations are further given in Fig. 8. To quantify the model performances, statistical criteria (Boylan and Russell, 2006; Willmott et al., 2012; Yumimoto et al., 2017), including the mean bias (BIAS), the mean fractional error (MFE), the root mean square error (RMSE), and the index of agreement (IOA) are calculated between the simulated results and observations. It is apparent that all the three assimilation experiments can optimize the dust emissions to better simulate AOTs and AEs closer to the assimilated AERONET and independent SONET observations. Compared to the independent SONET observations, the AOT DA-SZD experiment reduces the AOT BIAS and RMSE of the FR experiment by 85% and 14%, however, it can only reduce the AE BIAS and RMSE of the FR experiment by 4% and 1%. This indicates the assimilation of AERONET AOT observations can only optimize the dust emission but not the dust emission size distribution.

The AOT+AE DA-SZD experiment reduces the AOT BIAS and RMSE of the FR experiment by 92% and 17%, and it can also reduce the AE BIAS and RMSE of the FR experiment by 68% and 62%. Although the AOT+AE DA experiment assimilates the AE observations, however, it has limited improvement of AE due to the uncertainty of dust emission size distribution is not considered. Those indicate the additional assimilation of AE observations with consideration of the dust emission size distribution uncertainty are helpful to the optimization of dust emission through better adjustment of dust size distribution. The similar conclusions are also found in the comparison with assimilated AERONET observations.

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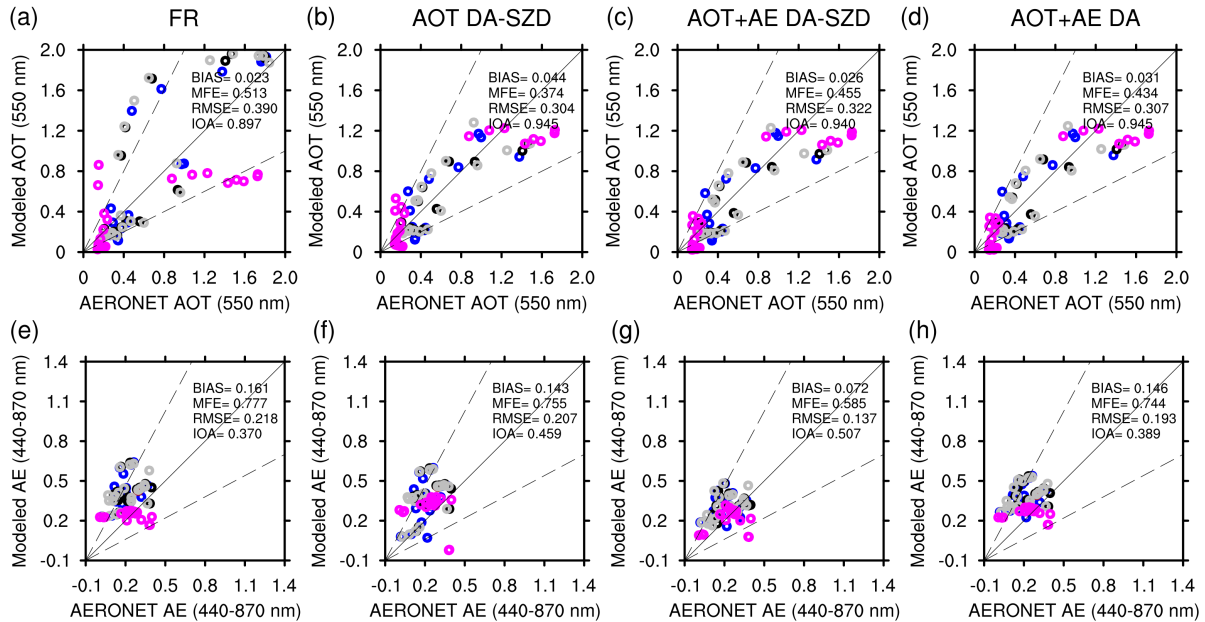


Figure 7. Scatter plots of assimilated AEROSOL ROBOTIC NETWORK (AERONET) hourly aerosol optical thicknesses (AOTs) versus the simulated ones at 550 nm for FR (a), AOT DA-SZD (b), AOT+AE DA-SZD (c), and AOT+AE DA (d) experiments from 14 March to 23 March 2021. The colours of the different sites are illustrated in Fig. 1(a). The solid black line is the 1:1 line and the dashed black lines correspond to the 1:2 and 2:1 lines. BIAS, MFE, RMSE, and IOA represent the mean bias, the mean fractional error, the root mean square error, and the index of agreement. (e,f,g,h) Same as (a,b,c,d) but for Angstrom Exponents (AEs) in the wavelength 440-870 nm.

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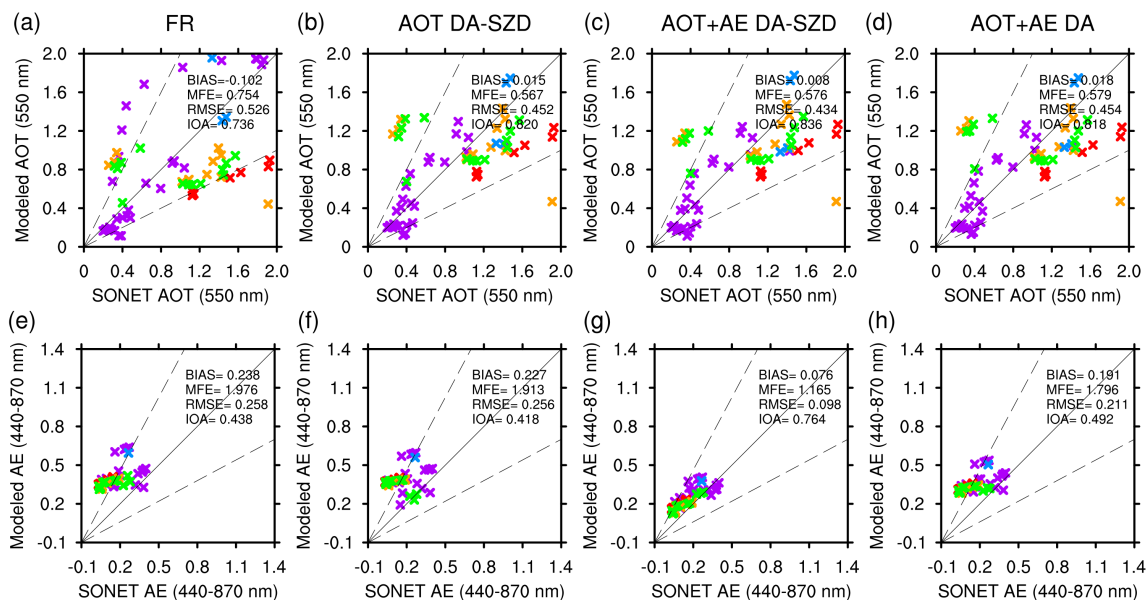
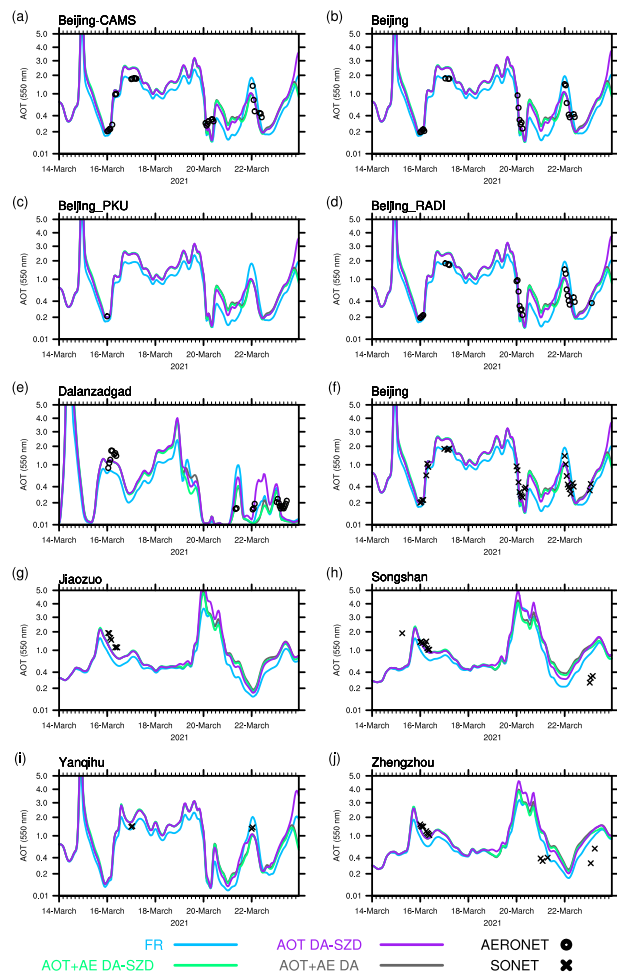


Figure 8. Same as Fig. 7 but for Skynet Observation NETwork (SONET) hourly observations for independent validation.

315 Time series of the simulated AOT and AE at AERONET and SONET sites for the four experiments are further given in Fig. 9 and Fig. 10, respectively. The BIAS and RMSE of simulated and observed AOT and AE over each site are given in Table S2 and Table S3. It is found that FR experiment can generally reproduce the time variations of the observed aerosol optical properties especially AOT over all sites, indicating the dust processes simulated by WRF-Chem are reasonable. The reasonable simulated dust processes demonstrate the covariance between the simulated dust emissions and aerosol optical properties is reliable to optimize the dust emission. Due to FR experiment underestimates AOTs and overestimates AEs over both the dust source and downwind sites during 16-17 March 2021, the assimilation of only AOT observation leads to increase the dust emission in each bin except bin 4 over the Gobi desert during 14-15 March 2021 (Table S1) and the additional AE observation leads to significantly decrease of dust emission in bin 1 and increase of dust emission in bin 3. The latter induces the simulated AEs on March 16 are comparable to the observed ones especially over the independent 325 SONET sites named Songshan and Zhengzhou, and this proves that not only the simulated dust emissions but also their size distributions over Gobi desert in AOT+AE DA-SZD experiment is optimized. The superiority of the adjustment of dust emission size distribution is further demonstrated by the limited effects of AE observation on the model in AOT+AE DA experiment due to the uncertainty of dust size distribution is not considered. Due to the simulated AOTs during 20-23 March 2021 in FR experiment is underestimated except on March 21 in dust source site and overestimated in downwind sites, the assimilation of AOT leads to slightly decreases of the dust emissions in bin 1, bin 2, bin 3, and bin 4 and significantly increase of that in bin 5 (Table S1). This induces the increase of the total dust emission in Gobi desert of Mongolia and decrease of the total dust emission in Gobi desert of China (Fig. 5). Due to the simulated AEs during 20-23 March 2021 in 330

FR experiment is slightly underestimated except on March 23 in dust source site and significantly overestimated in downwind sites, the assimilation of additional AE induces the decrease of dust emission in each bin.



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Figure 9. Hourly time series of the simulated aerosol optical thicknesses (AOTs) for the FR, AOT DA-SZD, AOT+AE DA-SZD, and AOT+AE DA experiments and the observed ones over AERONET sites (a-e) and SONET sites (f-j) during 14-23 March 2021.

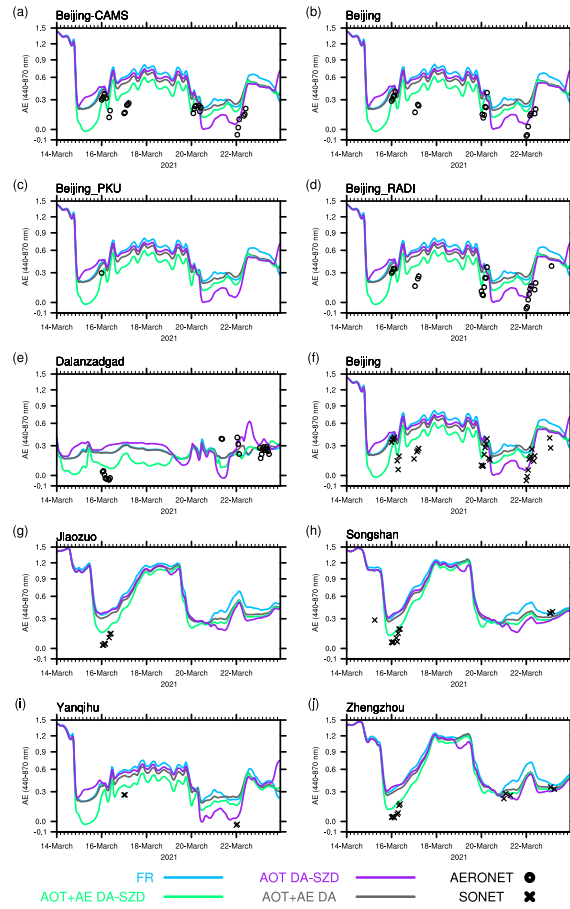
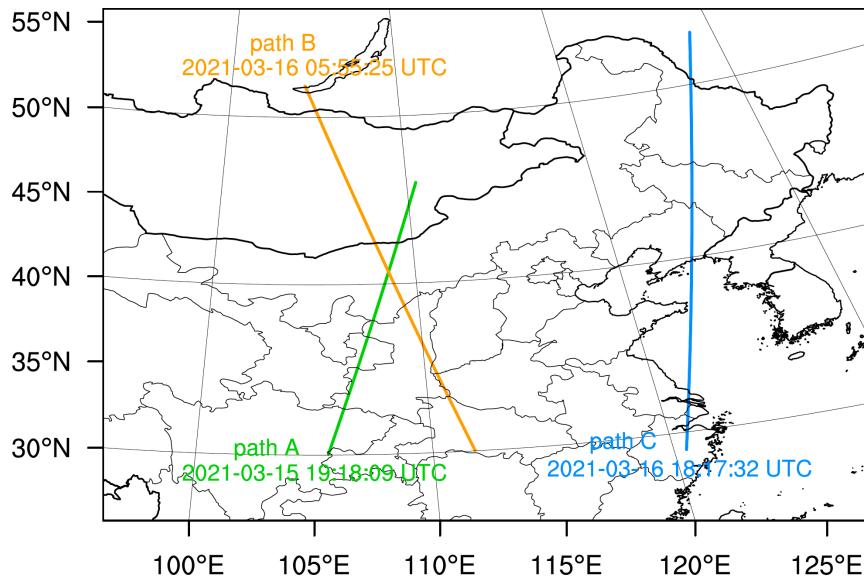


Figure 10. Same as Fig. 9 but for Angstrom Exponents (AEs) in the wavelength 440-870 nm.

340 4.3 Evaluation of simulated aerosol vertical extinctions

To further evaluate the dust emission optimization, the simulated aerosol extinction coefficients are compared with the independent CALIOP observed ones in the three CALIPSO orbit paths (Fig. 11). Due to AOT+AE DA-SZD experiment has the best performance among the three assimilation experiments, the results for FR and AOT+AE DA-SZD experiments in the three CALIPSO orbit paths are given in Fig. 12. Two paths near dust source region crosses the westward pathway of dust transport at 19:18:09 UTC on 15 March (path A) and the eastward pathway of dust transport at 05:55:25 (UTC) on 16 March (path B), and one path is far away from the dust source region at 18:17:32 UTC on 16 March (path C) (Fig. S8). As shown in Fig. 12, the vertical aerosols in path A are dominated by dust from 1 km to 12 km altitude. FR experiment can capture the observed aerosol vertical patterns, whereas it significantly underestimates the aerosol extinction coefficients near surface. AOT+AE DA-SZD experiment reduces the underestimation and performs more reasonable magnitude of aerosol extinction coefficients with values higher than 1 km^{-1} around surface from 35°N to 41°N and 0.1 km^{-1} from 2 km to 4 km around 40°N . It indicates that the assimilation of AOT and AE observations can better reproduce the features of aerosol vertical variations

during dust transportation near the dust source region. In addition, it should be noted that the improvements of the aerosol extinctions with posterior dust emission on March 15 benefit from assimilating the time-lagged observations from downwind areas. In path B, the dust transported from west to east is mainly concentrated in 4 km and FR generally reproduce this dust vertical structure with significant underestimations. AOT+AE DA-SZD experiment can improve the underestimations and the simulated aerosol extinctions are further consistent with the observed ones. CALIOP observed aerosol extinctions from 6 to 7 km around 41°N are higher than 1 km⁻¹, while the simulated aerosol extinctions for FR experiment are around 0.3 km⁻¹. AOT+AE DA-SZD experiment successfully reproduce the magnitude and variations of aerosol extinctions around 41°N. The vertical aerosols in path C are also dominated by dust in all heights. Although FR experiment successfully reproduces the vertical structure of double dust-layers observed by CALIOP between 33°N to 43°N, the aerosol extinction coefficients are significantly underestimated around 4 km. AOT+AE DA-SZD experiment increases the transported dust, diminishing the underestimation of aerosol extinction coefficients. The aerosol extinction coefficients in AOT+AE DA-SZD experiment are more comparable to the CALIOP observations. This proves that the assimilation of AOT and AE observations can also better reproduce the features of dust vertical distributions in areas far away from the dust source region.



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Figure 11. The Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) 3 orbit paths during 14-23 March 2021.

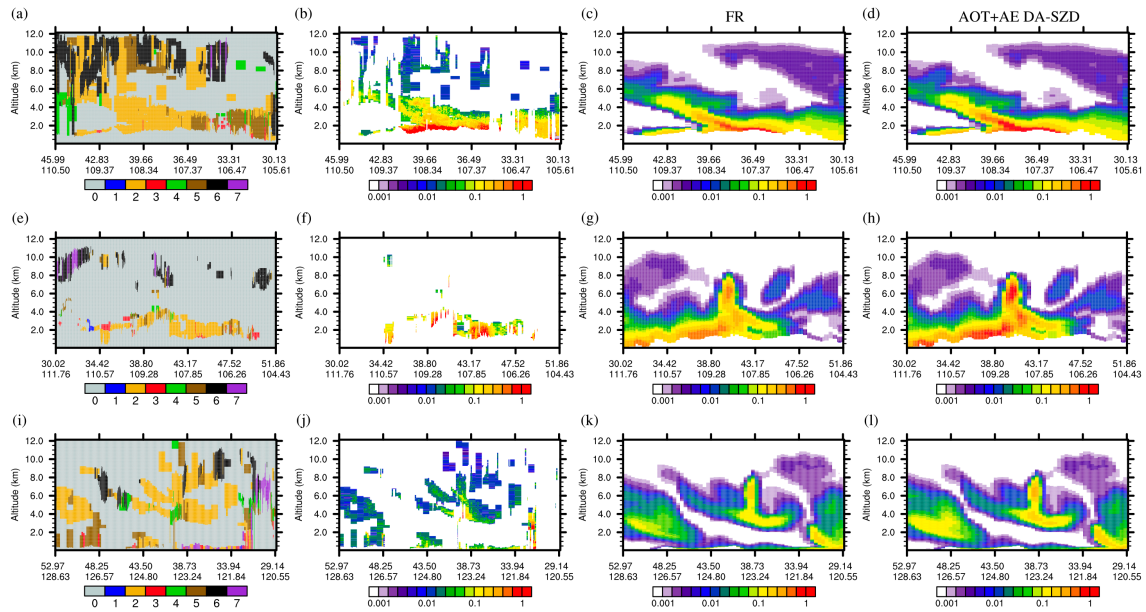


Figure 12. (a) Time-height cross section of CALIPSO-derived vertical aerosol subtypes in path A. The Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP)-observed aerosol extinction coefficients at 532 nm (km^{-1}) (b) and the simulated ones at 550 nm in the FR (c) and AOT+AE DA-SZD (d) experiments. (e-h) Same as (a-d) but for path B. (i-l) Same as (a-d) but for path C.

5 Conclusions

To investigate the additional benefit of aerosol size information in dust emission optimization, the Aerosol Robotic Network (AERONET) ground-based aerosol optical thickness (AOT) and Ångström Exponent (AE) time-lagged observations are assimilated during the severe East Asian dust storm outbreak in March 2021 in this study. The Ensemble Kalman smoother (EnKS) assimilation framework (Dai et al., 2019) with the Weather Research and Forecasting model coupled with Chemistry (WRF-Chem) version 4.4 is applied for the dust emission optimization.

Three assimilation experiments are conducted during 14-23 March 21. The first one named “AOT DA-SZD” only assimilates AERONET AOT observations with perturbation of dust emission and size distribution, the second one named “AOT+AE DA-SZD” assimilates is conducted as same as the first one except assimilating both AERONET AOT and AE observations, and the third one named “AOT+AE DA” is conducted as same as the second one except the ensemble members are generated by perturbing the dust emission in each bin with same perturbation factor. The baseline experiment “FR” without assimilation is used for comparison.

Our results demonstrate that the additional assimilation of AE observations with consideration of the dust emission size distribution uncertainty are helpful to the optimization of dust emission through better adjustment of dust size distribution. AOT assimilation can only optimize the dust emission flux depending on the covariance between time-lagged AOT observations and the simulated total dust emission, while the additional inclusion of AE assimilation can optimize the size

distribution of dust emission and the associated total flux depending on the covariance between time-lagged AE observations and the simulated dust emission in each bin.

390 All the three assimilation experiments can optimize the dust emissions to better reproduce the assimilated AERONET and independent SONET AOT and AE observations. Although the assimilation of AOT observations can only optimize the dust emission but not the dust emission size distribution, the assimilation with additional AE observations can not only reduce the AOT BIAS and RMSE of the FR experiment by 92% and 17%, but also reduce the AE BIAS and RMSE of the FR experiment by 68% and 62% through optimizing both the dust emission size distribution and the associated total flux. The
395 temporal variation of simulated AOT and AE can be both improved through assimilating additional AE information. The assimilation of AOT and AE also makes the magnitude and variations of aerosol vertical extinctions more comparable to independent Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) observations both in the westward and eastward pathways of dust transport.

This study emphasizes the additional AE assimilation is useful in the dust emission optimization. To further explore the roles
400 of the assimilated observations on the dust emission optimization and accurate simulation of dust life cycle, sensitivity experiments should be taken to quantify the influences of observation uncertainties and frequencies on the assimilation efficiency. The assimilation parameters such as spatial and temporal localization length are also important for dust emission optimization. In addition, the coarse-mode AERONET AOT from the spectral deconvolution algorithm (SDA) is also useful for dust emission optimization since all fine-mode aerosols are truncated and only dust/sea-salt remains.

405 **Code and data availability**

All data used in this study is freely available from public data repositories. AERONET products are available from https://aeronet.gsfc.nasa.gov/new_web/download_all_v3_aod.html. SONET products are available from <http://www.sonet.ac.cn/en/cpin/html/?194.html>. CALIOP products are available from the NASA Langley Research Center–Atmospheric Sciences Data Center (ASDC).

410 **Author contributions**

YC conceived the study and designed the dust storm data assimilation. YC and TD performed the control and assimilation tests and carried out the data analysis. JC, DG, JJ, TN, and GS provided useful comments on the paper. YC prepared the manuscript with contributions from TD and all others co-authors.

Acknowledgments

415 This study was financially supported by the National Natural Science Funds of China (grant nos. 42175186, 42305088, 42375190), the China Postdoctoral Science Foundation (grant no. 2022M723091), the Special Research Assistant Project of the Chinese Academy of Sciences, the Open fund by Jiangsu Key Laboratory of Atmospheric Environment Monitoring and Pollution Control (KHK 2206), the Key Laboratory of Atmospheric Chemistry, China Meteorological Administration,

LAC/CMA (grant no. 2022B05), the Youth Innovation Promotion Association CAS (grant no. 2020078), and the
420 International Partnership Program of Chinese Academy of Sciences (grant no. 134111KYSB20200006). Model simulations
were performed using NEC SX-Aurora TSUBASA supercomputers at NIES, Japan. We thank to the relevant researchers
who provided AERONET, SONET, and CALIOP observations. We also thank the anonymous reviewers for their valuable
comments and suggestions that improved the manuscript.

Competing interests

425 The contact author has declared that none of the authors has any competing interests.

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