Technical Note: Nighttime OH and HO₂ chemical equilibria in the mesosphere – lower thermosphere

Mikhail Yu. Kulikov¹, Mikhail V. Belikovich¹, Aleksey G. Chubarov¹, Svetlana O. Dementyeva¹, and
 Alexander M. Feigin¹

¹A. V. Gaponov-Grekhov Institute of Applied Physics of the Russian Academy of Sciences, 46 Ulyanov
Str., 603950 Nizhny Novgorod, Russia

7 Correspondence to: Mikhail Yu. Kulikov (mikhail_kulikov@mail.ru)

8

9 Abstract. At the altitudes of the mesosphere – lower thermosphere, OH and HO_2 play a significant role in many physicochemical processes. Thus, monitoring-of their spatiotemporal evolution together with 10 11 other chemically active trace gases is one of the most important problems for this atmosphere region, in 12 which direct measurements are difficult. The paper studies the nighttime OH and HO_2 chemical equilibria 13 using the 3D chemical transport modeling within the general approach-including, which includes the 14 extractionidentification of the main sources and sinks in the equilibrium space-time areas and derivation 15 of analytical criteria for equilibrium validity. The presented analysis shows, that there are extended areas, where nighttime HO₂ and OH are close to their local equilibrium concentrations, determined mainly by 16 the reaction between $HO_x - O_x$ components <u>among</u> themselves and with H_2O_2 , N, NO, NO_{27} and CO. In 17 18 the upper mesosphere – lower thermosphere, the equilibrium expressions can be shortened, includingso 19 <u>that they include</u> the $HO_x - O_x$ chemistry only. These <u>conditions describes</u> expressions describe the HO_2 20 and OH equilibriumequilibria from the top down to the some lower bordersboundaries, the altitude 21 position positions of which vary in the interval between 72-73 and 85 km and depends depend essentially 22 on season and latitude. The developed analytical criteria almost everywhere well reproduce well the main 23 features of these borders.boundaries. Due to weak sensitivity to uncertainties of reaction rates and other 24 parameters, the criteria can be considered a robust instrument for HO₂ and OH equilibrium validation. The obtained results allow us to extend the abilities of previously proposed methods for the retrieval of 25 26 poorly measured components from measurement data and to develop new approaches.

28 **1 Introduction**

29 Monitoring the spatiotemporal evolution of chemically active trace gases is one of the most 30 important problems in atmospheric research. Despite the increase of the experimental data volume 31 nowadays, primarily due to the development of remote sensing methods, many important trace gases continue to be unavailable for direct and regular measurements. A well-known way to increase the 32 33 information content of experimental campaigns is to use chemical transport models and the available experimental data for deriving in conjunction with a certain chemical or physicochemical model to derive 34 35 unmeasured characteristics indirectly. Within the framework of this approach, the model acts as a priori relationship between directly measured and retrieved characteristics. The simplest model, that makes it 36 37 possible to implement this approach, is based on the condition of local (in both time and space) photochemical/chemical balance (local equilibrium) between sources and sinks of the so-called "fast" 38 components: trace gases with short lifetimes relatively, in particular, to the characteristic transport times. 39 40 Mathematically, this condition does not mean that the fast variables are at equilibrium, but when it is 41 fulfilled, the corresponding concentrations are close to their instantaneous equilibrium values. At the 42 same time, due to the strong dissipation, in most cases (except the special cases where the ensemble of fast components includes the slow family of these components), there is no need to follow the law of 43 matter conservation. It is possible to discarddisregard insignificant sinks and sources, including those 44 45 caused by transport, in the corresponding balance equations without the loss of accuracy, including those 46 caused by transport. The resulting algebraic equations are the simplest *a priori* local relations between measurable and retrieved trace gases. These relationships can be used to derive information about hard-to-47 48 measure atmospheric species, determine key atmospheric characteristics (for example, temperature), 49 (Marchand et al., 2007)), validate the data quality of simultaneous measurements of several atmospheric components, (Kulikov et al., 2018a), estimate reaction rate constants known with high 50 uncertainty. (Stedman et al., 1975; Avallone and Toohey, 2001), evaluate sources-(emissions/sinks 51 52 (Cantrell et al., 2003), etc.

For several decades, the photochemical/chemical equilibrium approximation havehas been used to 53 solve many atmospheric tasks. It is applied (see, e.g., the short review in Kulikov et al. (2018a) and 54 55 references therein) in investigations of the surface layer and free troposphere chemistry in different regions (over megalopolises megapolises, in rural areas, in the mountains, over the seas), in stratospheric 56 chemistry studies, including derivation of a critical parameters of the ozone destruction catalytic cycles, 57 and to study in studies of the HO_x – O_x chemistry and airglows (O(1 S) green-line, O₂ A-band, OH Meinel 58 band emissions) at the heights of the mesosphere – lower thermosphere. In the latter $case_{\overline{1}}$ the 59 distributions of unmeasured characteristics are determined from the data of daytime and nighttime rocket 60 and satellite measurements (e.g., Evans and Llewellyn, 1973; Good, 1976; Pendleton et al., 1983; 61

62 McDade et al., 1985; McDade and Llewellyn, 1988; Evans et al., 1988; Thomas, 1990; Llewellyn et al., 1993; Llewellyn and McDade, 1996; Russell and Lowe, 2003; Russell et al., 2005; Kulikov et al., 2006, 63 64 2009, 2017, 2022a, 2022b; Mlynczak et al., 2007, 2013a, 2013b, 2014, 2018; Smith et al., 2010; Xu et al., 65 2012; Siskind et al., 2008, 2015; Fytterer et al., 2019) with the use of equilibrium assumptions for ozone 66 and excited states of OH, O, and O₂. For example, such an approach is applied to the data of the SABER (Sounding of the Atmosphere using Broadband Emission Radiometry) instrument onboard the TIMED 67 (Thermosphere Ionosphere Mesosphere Energetics and Dynamics) satellite, which since 2002 continues 68 to measure simultaneous profiles of temperature, ozone, and volume emission rates of OH* transitions in 69 70 wide ranges of altitude, local time, and latitude with a rather high space-time resolution.

71 Note a number of general aspects of the application of equilibrium conditions in the above examples. First of all, there are no clear criteria-why, indicating the conditions under which the use of 72 equilibrium condition should be satisfied.approximation is justified. Usually, a certain component is taken 73 74 to be a fast variable, if its lifetime is much shorter, than the lifetimes of other components of studied 75 photochemical/chemical system or the duration of a day, daytime, nighttime, etc. For example, in the 76 papers on SABER data processing (Mlynczak et al., 2013a, 2013b, 2014, 2018)) it is assumed, that the 77 nighttime ozone chemical equilibrium in the mesopause is well fulfilled at altitudes of 80-100 km, since 78 the nighttime ozone lifetime at these altitudes is too short varying varies in the range from several minutes 79 to several tens of minutes. Note, that this assumption is quite popular and used in different tasks (e.g., 80 Swenson and Gardner, 1998; Marsh et al., 2006; Smith et al., 2009; Nikoukar et al., 2007; Xu et al., 2010, 81 2012; Kowalewski et al., 2014; Grygalashvyly et al., 2014; Grygalashvyly, 2015; Sonnemann et al., 2015; 82 Kulikov et al., 2021). Belikovich et al. (2018) and Kulikov et al. (2018b, 2019, 2023a) analyzed the 83 nighttime ozone chemical equilibrium numerically, analytically, and with the use of SABER/TIMED 84 data. It was revealed, that the short lifetime is not a sufficient condition, so, this equilibrium may be significantly disrupted disturbed above 80 km. Secondly, there is no detailed numerical examination of 85 this approximation validity, depending on altitude, latitude, <u>The</u> local time, <u>ratio between true</u> and season. 86 87 Correspondingly, there is noequilibrium concentrations may vary widely and reach up to several orders of magnitude (e.g., Figure 5 in Kulikov et al. (2018b)). Thus, without special restrictions the assessment of 88 possible errorserror in retrieved characteristics due to disturbance of the used equilibrium condition the 89 use of equilibrium approximation is uncontrollable and may significantly exceed all other errors in the 90 retrieval procedure due to, for example, uncertainties in the measurement data and rate constants. 91

Starting since our <u>Since the papers byof</u> Belikovich et al. (2018) and Kulikov et al. (2018b, 2019,
2023a), we <u>developdeveloped</u> the general approach to <u>correct search of correctly identify</u> fast components
<u>using, employing</u> the data from a global 3D chemical transport model. It includes:

95 1. Plotting of the equilibrium space-time maps of interested component. the components of interest.

96 2. Identification of the main sources and sinks in the found equilibrium areas.

97 3. Derivation and subsequent use of analytical criteria that make it possible to determine the fulfillment of
98 the equilibrium condition locally (in time and space) with the use of the measurement data only.

99 The last<u>latter</u> point is based on the theory of chemical equilibrium of a certain trace gas<u>based on</u>, 100 acquired from estimations of its lifetime and equilibrium concentration, and time dependences of these 101 characteristics (Kulikov et al., 2023a). Note, that when equilibrium condition is applied to measurement 102 data in the retrieval of unmeasured characteristics, the criterion allows controlling and limiting the 103 possible error caused by the equilibrium approximation.

104 The main goalsgoal of thisthe paper is to apply mentioned this approach forto the analysis of 105 nighttime OH and HO₂ chemical equilibriums equilibria in the mesosphere – lower thermosphere. Along with O and H, OH and HO₂ are important components of HO_x – O_x chemistry, participating (a) in 106 chemical heating through, in particular, $O+OH \rightarrow O_2+H$ and $O+HO_2 \rightarrow O_2+OH$ exothermic reactions, (b) 107 in formation of airglows, (c) in catalytic cycles of the ozone destruction. Moreover, the equilibrium 108 109 conditions of OH and HO₂ are additional *a priori* relationships, that can be used to retrieve these components or other characteristics from measured data. In particular, the method proposed by Panka et 110 al. (2021) proposed the method for nighttime total OH retrieval from SABER/TIMED data at 80-100 km, 111 112 which does not use the nighttime ozone chemical equilibrium, but nevertheless. However, the method applies the equilibrium between sources and sinks for all exited and ground not only to excited states of 113 OH (v=0-9) with ultrashort lifetimes, but also to the ground state. Therefore, this approach point is 114 testedverified in our paper. 115

In the next section, we present the used model and methods. In Section 3, the model data are used to plot HO₂ and OH equilibrium maps. In Sections 4-5, we <u>extractidentify</u> the main reactions, determining equilibria of these gases, and present their shortened equilibria conditions at the upper mesosphere and lower thermosphere altitudes. In Section 6, the criteria for HO₂ and OH equilibria validity are developed. In Section 7, we discuss the obtained results and their possible applications.

121

122 2 Used 3D model and Approaches

The analysis of OH and HO₂ nighttime chemical equilibria was carried out, using the data, obtained with calculation of 3D chemical transport model of the middle atmosphere, developed at the Leibniz Institute of Atmospheric Physics (e.g., Sonnemann et al., 1998; Körner & Sonnemann, 2001; Grygalashvyly et al., 2009; Hartogh et al., 2004, 2011) to investigate the mesosphere – lower thermosphere chemistry, in particular, in the extended mesopause region. A number of papers (e.g., Hartogh et al., 2004, 2011; Sonnemann, et al., 2006, 2008) validated the model with measurements, in
particular, for ozone and water vapourvapor.

The space-time distribution of temperature and winds were taken from the model of the dynamics 130 of the middle atmosphere COMMA-IAP (e.g., Kremp et al., 1999; Berger and von Zahn, 1999) with an 131 updated frequency of 1 day and linear smoothing between subsequent updates to avoid unrealistic jumps 132 in the calculated concentrations of trace gases. 3D advective transport is taken into account with the use 133 134 of the Walcek-scheme (Walcek, 2000). The vertical diffusive transport (turbulent and molecular) is calculated with the use of the implicit Thomas algorithm (Morton and Mayers, 1994). The model grid 135 includeshas 118 pressure-headheight levels (0–135 km), 16 latitudinal and 32 longitudinal levels. The 136 chemical module (see Table 1) comprises 25 constituents (O, O(¹D), O₃, H, OH, HO₂, H₂O₂, H₂O, H₂, N, 137 NO, NO₂, NO₃, N₂O, CH₄, CH₂, CH₃, CH₃O₂, CH₃O, CH₂O, CHO, CO, CO₂, O₂, N₂), 54 chemical 138 139 reactions between them, and 15 photo-dissociation reactions. The model utilizes the pre-calculated dissociation rates (Kremp et al., 1999) and their dependence on the altitude and solar zenith angle. 140

141 The model was used to calculate a one-year global evolution of the above mentioned trace gases. 142 For removing To remove the transition regions corresponding to sunset and sunrise, we took into account the<u>use only</u> local time<u>times</u> when the solar zenith angle $> 95\chi > 105^{\circ}$. As a result, we find the 143 spatiotemporal series of the OH/OH^{eq} and HO_2/HO_2^{eq} ratios. Here, OH and HO_2 are the local nighttime 144 values of hydroxyl and hydroperoxyl radicals, calculated by the model, OH^{eq} and HO_2^{eq} are 145 their local equilibrium values, corresponding to the instantaneous balance between production and loss 146 terms, respectively. Therefore, for determination of To determine each local value of OH^{eq} and HO_2^{eq} , 147 we used the local values of the parameters (temperature, O2, and N2) and the concentrations of other trace 148 gases, determining local chemical sources and sinks of OH and HO_2 . Then, the OH/OH^{eq} and $HO_2/$ 149 HO_2^{eq} series were averaged over the zonal coordinate and time during each month and were presented as 150 height-latitude maps, depending on the month. Each map contains lines, marking the boundaries of the 151 152 equilibrium areas, where the following conditions are satisfied:

153
$$\begin{cases} |< 0H/0H^{eq} > -1| \le 0.1 \\ \sigma_{0H/0H^{eq}} \le 0.1 \end{cases}, \begin{cases} |< HO_2/HO_2^{eq} > -1| \le 0.1 \\ \sigma_{HO_2/HO_2^{eq}} \le 0.1 \end{cases},$$
(1)

where the angle brackets are used to denote the values averaged in time and space, $\sigma_{OH/OH^{eq}}$ and $\sigma_{HO_2/HO_2^{eq}}$ are standard deviations of the OH/OH^{eq} and HO_2/HO_2^{eq} ratios from 1, respectively.

Then, we plotted spatiotemporal maps, showing the relative contribution of each reaction to a summarized source or sink at all altitudes and latitudes. These maps helped us to identify the main sources and sinks, describing the chemical equilibrium of nighttime OH and HO_2 in the equilibrium areas to an accuracy of better than a few percent. 160 At final stageFinally, we obtained and verified the analytical criteria of OH and HO₂ nighttime 161 chemical equilibria according to Kulikov et al. (2023a). The paper considered the poorlypure chemical 162 evolution of a certain trace gas n. It was shown:

163
$$\frac{dn}{dt} = I_n - S_n = -\frac{1}{\tau_n} (n - n^{eq})_a$$
164 $\tau_n = \frac{s}{s_n} n^{eq} = \frac{n \cdot t_n}{s_n}$ (2)165where t is time. I_n and S_n are total photochemical/chemical sources and sinks of n respectively. τ_n is the166n.lifetime and n^{eq} is its equilibrium concentration, corresponding to the condition $I_n = S_n$. The lifetime167determines the characteristic time scale, for which n approaches n^{eq} , when $n^{eq} = const.$ In general case168 τ_n and n^{eq} is its equilibrium concentration n^{eq} are close to each other $(n(t) \approx n^{eq}(t))$, when $\tau_n \ll$ 169values of n and its equilibrium-concentration n^{eq} are close to each other $(n(t) \approx n^{eq}(t))$, when $\tau_n \ll$ 170 τ_n^{eq} , where τ_n is the n lifetime and $\tau_{n^{eq}}$ is the local time scale of n^{eq} :171 $\tau_n^{eq} = \frac{n^{eq}}{|dn^{eq}/dt|^2}$ 172(2)173 $\frac{n^{eq}}{|dn^{eq}/dt|^2}$ 174The expression for τ_n is found from the total sink of n. The expression for τ_n^{eq} is derived from Eq. (3)175with the use of differential equations, describing chemical evolution of other reacting components, which176 $suff(dn^{eq}) \cdot \frac{\tau_n}{\tau_n^{eq}}$ in the first order approximation. Thus, the criterion177 $suff(dn^{eq}) \cdot \frac{\tau_n}{\tau_n^{eq}}$ in the first order approximation. Thus, the criterion178 $\tau_n/\tau_n^{eq} \leq 0.1$ 179is sufficient, in order to the possible relative difference between n and n^{eq} to be no more than 0.1.180**3 Nighttime HO2 and OH chemical equilibriumsequilibria**181**3 Nighttime HO2 and OH chemical equilibriumsequilibria**<

184 $H+O_2+M \rightarrow HO_2+M$ (R20), $OH+O_3 \rightarrow O_2+HO_2$ (R22), $H_2O_2+OH \rightarrow H_2O+HO_2$ (R29), $H_2O_2+O \rightarrow$ 185 $OH+HO_2$ (R19), $CHO+O_2 \rightarrow HO_2+CO$ (R40), $CH_3O+O_2 \rightarrow CH_2O+HO_2$ (R37),

186 whereas chemical sinks of this component are as follows:

187 $HO_2+O \rightarrow OH+O_2$ (R18), $HO_2+O_3 \rightarrow OH+2O_2$ (R23), $OH+HO_2 \rightarrow H_2O+O_2$ (R28), $H+HO_2 \rightarrow 2OH$ 188 (R14), $H+HO_2 \rightarrow H_2O+O$ (R15), $H+HO_2 \rightarrow H_2+O_2$ (R16), $NO+HO_2 \rightarrow NO_2+OH$ (R50), $HO_2+HO_2 \rightarrow$ 189 $H_2O_2+O_2$ (R30), $HO_2+HO_2+M \rightarrow H_2O_2+O_2+M$ (R31).

190 Thus, HO₂ local equilibrium concentration is described by the following equation:

191
$$HO_2^{eq} = \frac{k_{20} \cdot H \cdot M \cdot O_2 + k_{22} \cdot O H \cdot O_3 + k_{29} \cdot H_2 O_2 \cdot O H + k_{19} \cdot H_2 O_2 \cdot O + k_{40} \cdot C HO \cdot O_2 + k_{37} \cdot C H_3 O \cdot O_2}{k_{18} \cdot O + k_{23} \cdot O_3 + k_{28} \cdot O H + (k_{14} + k_{15} + k_{16}) \cdot H + k_{50} \cdot NO + 2 \cdot (k_{30} + k_{31} \cdot M) \cdot HO_2}$$
(45)

The Figure 1 plots height-latitude cross sections for the $< HO_2/HO_2^{eq} >$ ratio for each month. The 192 dashed area corresponds to $\chi < 95^{\circ}$. The white area represents the ratio outside the [0.5, 1.5] interval. The 193 194 black solid lines mark the bordersboundaries of equilibrium areas, where, according to condition (1), 195 local values of HO₂ are close to their equilibrium values with a possible bias of less than 10%. At low and 196 middle latitudes, one can see the presentpresence of the main equilibrium area, which extends from the 197 top of the analyzed altitude range to the lower boundary. The height of this equilibrium borderboundary, $z_{HO_2^{eq}}$, depends on the season and latitude and varies in the interval between 73 and 85 km. It is the 198 highest and the lowest during the summer and winter respectively at the middle latitudes. Near the 199 equator, $z_{HO_2^{eq}}$ demonstrates the weakest annual variations and varies in the range of 81-83 km range. 200 201 There are local areas below the upper longest black line, but they are small and irregular and can be omitted from our consideration. Note only that the maps in many months show the existence of 202 203 equilibrium near 50 km, which can be assumed to be the beginning of the main equilibrium area in the stratosphere. At high latitudes, there is the main equilibrium area as at low and middle latitudes, but this 204 area above 70-75° of latitude can extend down to 50 km-with small exceptions. 205

In accordance to the Table 1, OH chemical sources are determined by the following reactions:

208 NO+HO₂ \rightarrow NO₂+OH (R50), H₂O₂+O \rightarrow OH+HO₂ (R19), H +NO₂ \rightarrow OH+NO (R51), O(¹D)+H₂O \rightarrow

209 2OH (R7), $O(^{1}D)+H_{2} \rightarrow H+OH (R8), CH_{4}+O(^{1}D) \rightarrow CH_{3}+OH (R9),$

210 whereas chemical sinks of this component are as follows:

206

211
$$OH+O \rightarrow H+O_2$$
 (R17), $OH+O_3 \rightarrow O_2+HO_2$ (R22), $OH+HO_2 \rightarrow H_2O+O_2$ (R28), $OH+OH \rightarrow H_2O+O_2$

212 (R26), $OH+OH+M \rightarrow H_2O_2+M$ (R27), $H+OH+N_2 \rightarrow H_2O+N_2$ (R24), $H_2O_2+OH \rightarrow H_2O+HO_2$ (R29),

213 OH+CO \rightarrow H+CO₂ (R32), CH₄+OH \rightarrow CH₃+H₂O (R33), OH+H₂ \rightarrow H₂O+H (R25), N+OH \rightarrow NO+H 214 (R49).

215 Thus, OH local equilibrium concentration is described by the following equation:

 $\begin{array}{ll} 216 & OH^{eq} = (k_{21} \cdot H \cdot O_3 + k_{18} \cdot O \cdot HO_2 + k_{23} \cdot HO_2 \cdot O_3 + 2 \cdot k_{14} \cdot H \cdot HO_2 + k_{50} \cdot HO_2 \cdot NO + k_{19} \cdot \\ \\ 217 & H_2O_2 \cdot O + k_{24} \cdot H \cdot N_2 + k_{51} \cdot NO_2 \cdot H + 2 \cdot k_7 \cdot O(^1D) \cdot H_2O + k_8 \cdot O(^1D) \cdot H_2 + k_9 \cdot O(^1D) \cdot \\ \\ 218 & CH_4)/(k_{17} \cdot O + k_{22} \cdot O_3 + k_{28} \cdot HO_2 + 2 \cdot (k_{26} + k_{27} \cdot M) \cdot OH + k_{29} \cdot H_2O_2 + k_{32} \cdot CO + k_{33} \cdot CH_4 + \\ \\ 219 & k_{25} \cdot H_2 + k_{25} \cdot N) \end{array}$

Figure 2 shows height-latitude cross sections for the $\langle OH/OH^{eq} \rangle$ ratio for each month. In this case, the equilibrium covers up to 70-80% of the presented ranges of heights and latitudes, so that the black solid lines mark the external bordersboundaries of non-equilibrium areas. In March and September, this area is almost symmetrical to the equator. In April-August, it is shifted towards the northern hemisphere. In October-February, this area is higher in the southern hemisphere. In all months, it is below 85-86 km. In the polar regions, there are latitudinal ranges, where OH is close to equilibrium throughout the entire range of heights.

227

228 4 The main reactions, determining HO₂ and OH equilibriumsequilibria

229 The Figure 3 showspresents height-latitude contour maps, showing the relative contribution of a certain reaction to the total source and sink of HO₂ in different monthsJanuary, taken foras an example. 230 To increase the information content of the panels, the altitude range is cut off everywhere to $\frac{100 \text{ km}10^{-3}}{100 \text{ km}10^{-3}}$ 231 hPa, since there are no significant changes above. Note, first firstly, that reaction $H+O_2+M \rightarrow HO_2+M$ 232 determines a major (up to 95% and more) contribution in the main equilibrium area almost everywhere, 233 except for the polar regions above 70-75° of latitude and below 75-80 km, where the reactions $OH+O_3 \rightarrow$ 234 O_2+HO_2 and $H_2O_2+OH \rightarrow H_2O+HO_2$ become important and should be taken into account. Second, other 235 <u>Other</u> reactions $(H_2O_2+O \rightarrow OH+HO_2, CHO+O_2 \rightarrow HO_2+CO, CH_3O+O_2 \rightarrow CH_2O+HO_2)$ together 236 contribute less than 2-3% to the total source of HO_2 in the main equilibrium area and may be omitted. 237

The Figure 4 presents height-latitude contour maps showing the relative contribution of a certain 238 239 reaction to the total sink of HO₂ in the same months as in Figure 3. Firstly, it should be noted that Secondly, the reaction HO₂+O \rightarrow OH+O₂ determines a major (up to 95% and more) contribution to 240 the total sink in the main equilibrium area almost everywhere, except for the same small polar areas, as in 241 the considered case with the sources, where the reactions $HO_2+O_3 \rightarrow OH+2O_2$ and $NO+HO_2 \rightarrow NO_2+OH$ 242 are important and should be taken into account. Secondly, the <u>The</u> reactions $OH+HO_2 \rightarrow H_2O+O_2$, 243 $H+HO_2 \rightarrow 2OH, H+HO_2 \rightarrow H_2O+O, and H+HO_2 \rightarrow H_2+O_2$ give together contribute cumulatively up to 244 10-15% of the total sourcesink near the boundary of the main equilibrium area. Thirdly, the The remaining 245 reactions $(HO_2+HO_2 \rightarrow H_2O_2+O_2, HO_2+HO_2+M \rightarrow H_2O_2+O_2+M)$ are not important in the main 246 247 equilibrium area and can be omitted.

Therefore, the expression for HO₂ local equilibrium concentration can be simplified as follows:

249
$$HO_2^{eq} = \frac{k_{20} \cdot H \cdot M \cdot O_2 + k_{22} \cdot O H \cdot O_3 + k_{29} \cdot H_2 O_2 \cdot O H}{k_{18} \cdot O + k_{23} \cdot O_3 + k_{28} \cdot O H + (k_{14} + k_{15} + k_{16}) \cdot H + k_{50} \cdot N O}$$
(67)

250 Figures 5-6 showFigure 4 presents height-latitude contour maps, showing the relative contribution of a certain reaction to the total source and sink of OH in the same months as in Figure 3January, taken 251 foras an example in Figure 3. As in the previous case, the altitude range is cut off at 100 km10⁻³ hPa, 252 253 because only the panels for the reactions $H+O_3 \rightarrow OH+O_2$ and $HO_2+O \rightarrow OH+O_2$ consist of interesting 254 variations at the 100-130 km altitudes.above. Note, that firstly these reactions are the main OH sources in 255 the upper part of the presented distributions down to 70-75 km, where they jointly provide up to a 95% contribution into the equilibrium concentration. Also, the reaction $HO_2+O_3 \rightarrow OH+2O_2$ is major source in 256 the lower part of the presented distributions from 50 to 60-70 km, depending on the month. 257 The reaction NO+HO₂ \rightarrow NO₂+OH is important around non-equilibrium areas of OH and should be taken 258 into account, whereas the reaction H +NO₂ \rightarrow OH+NO is important in compact altitude-latitude areas 259 near the poles, the reaction $H+HO_2 \rightarrow 2OH$ gives up to 10-15% contribution in small areas near the 260 equilibrium boundary. Other reactions $(O(^{1}D)+H_{2}O \rightarrow 2OH, O(^{1}D)+H_{2} \rightarrow H+OH, CH_{4}+O(^{1}D) \rightarrow O(^{1}D)+H_{2} \rightarrow O(^{1}D)$ 261 $CH_3+OH, H_2O_2+O \rightarrow OH+HO_2$) together contribute less than 2-3% of the total source of OH in the main 262 equilibrium area and can be omitted. 263

264 Figures 7-8 present height-latitude contour maps showing the relative contribution of a certain reaction to the total sink of OH. First, note that Secondly, the reaction $OH+O \rightarrow H+O_2$ is the main OH 265 sink in the upper part of the presented distributions down to 70-80 km, depending on the month, where it 266 provides up to 95% of the equilibrium concentration. The reactions $OH+CO \rightarrow H+CO_2$ and $OH+O_3 \rightarrow$ 267 O_2+HO_2 are major in the loverlower part of the presented distributions from 50 to 70-80 km, depending 268 on the month. The reaction $OH+HO_2 \rightarrow H_2O+O_2$ is remarkable significant enough around non-269 equilibrium areas of OH, whereas the reaction $H_2O_2+OH \rightarrow H_2O+HO_2$ is important in the compact 270 altitude-latitude area near the poles. Other reactions (OH+OH \rightarrow H₂O+O, OH+H₂ \rightarrow H₂O+H, N+OH \rightarrow 271 NO+H, $CH_4+OH \rightarrow CH_3+H_2O$, $H+OH+N_2 \rightarrow H_2O+N_2$, $OH+OH+M \rightarrow H_2O_2+M$) together contribute 272 less than 2-3% to the total sourcesink of OH in the main equilibrium area and can be omitted. 273

Therefore, the expression for OH local equilibrium concentration can be can be simplified as follows:

276
$$OH^{eq} = \frac{k_{21} \cdot H \cdot O_3 + k_{18} \cdot O \cdot H O_2 + k_{23} \cdot H O_2 \cdot O_3 + 2 \cdot k_{14} \cdot H \cdot H O_2 + k_{24} \cdot H \cdot N_2 + k_{50} \cdot H O_2 \cdot N O + k_{51} \cdot N O_2 \cdot H}{k_{17} \cdot O + k_{22} \cdot O_3 + k_{28} \cdot H O_2 + k_{29} \cdot H_2 O_2 + k_{32} \cdot C O}$$
(78)

277

248

T

5 Shortened equilibrium conditions of HO₂ and OH in the upper mesosphere and lower thermosphere

280 The above analysis revealed, that the reactions describing the equilibrium conditions (6-7-8) in the lower and middle mesosphere are mainly different from those- in the upper mesosphere and lower 281 282 thermosphere. This means that the task of applying these conditions can be divided into two parts 283 depending on the selected altitude range. At the upper mesosphere and lower thermosphere altitudes, we can consider only the $HO_x - O_x$ chemistry, excluding the reactions with participation of H_2O_2 , N, NO, 284 NO₂, and CO. In addition, we can omit the reactions HO₂+O₃ \rightarrow OH+2O₂, OH+O₃ \rightarrow O₂+HO₂, and 285 $OH+HO_2 \rightarrow H_2O+O_2$ due to their insignificance here. As the result, the shortened equilibrium conditions 286 287 of HO₂ and OH for this altitude range are as follows:

288
$$HO_{2sh}^{eq} = \frac{k_{20} \cdot H \cdot M \cdot O_2}{k_{18} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H},$$
(89)

289
$$OH_{sh}^{eq} = \frac{k_{21} \cdot H \cdot O_3 + k_{18} \cdot O \cdot H O_2 + 2 \cdot k_{14} \cdot H \cdot H O_2}{k_{17} \cdot O}$$
(910)

The Figure 95 shows height-latitude cross sections for the $\langle HO_2/HO_{2sh}^{eq} \rangle$ ratio for each month. 290 291 In each panel, the upper longest black line marks the lower borderboundary of the main equilibrium area, where, according to <u>condition</u> (1), $HO_2 \approx HO_{2sh}^{eq}$ with possible bias of less than 10%. As in the case of 292 Figure 1, this area extends from the top of the analyzed altitude range. There are also very small 293 equilibrium areas below, which can be omitted from our consideration. The height of the lower 294 border<u>boundary</u> of the main equilibrium area, $\frac{z_{HO_{exh}}}{z_{sh}} z_{HO_{exh}} z_{hO_{exh}}$, depends essentially on the season and 295 latitude. Comparing with Figure 1, one can see, <u>that it repeats well reproduces</u> many features of $z_{HO_2 ch}^{eq}$ at 296 low and middle latitudes. In particular, $z_{HO_{2sh}}^{eq}$ varies in the interval between 73 and 85 km, as in the case 297 of $z_{HO_2^{eq}}$. In the middle latitudes, $\frac{z_{HO_2^{eq}}}{z_{sh}} Z_{HO_2^{eq}}$ in summer is several km higher than in winter. Near the 298 equator; $z_{HO_{2sh}}^{eq}$ demonstrates the weakest annual variations and varies in the range of 81-83 km. So, one 299 can conclude, that the exclusion of a number of reactions does not lead to significant changes in the 300 space-time distributions of the HO₂ equilibrium. 301

The Figure 106 plots height-latitude cross sections for the $\langle OH/OH_{sh}^{eq} \rangle$ ratio for each month. As in the previous case, this is the lower borderboundary of the equilibrium area, where, according to condition (1), $OH \approx OH_{sh}^{eq}$ with good precision. The dependence of the borderboundary height, $z_{OH_{sh}^{eq}}$, on the season and latitude mainly repeats mainly the behavior of $z_{HO_{2sh}}^{eq}$. In particular, $z_{OH_{sh}}^{eq}$ varies in the interval between 73 and 85 km. At middle latitudes, $z_{OH_{sh}}^{eq}$ in summer is several km higher than in winter. Near the equator, $z_{OH_{sh}}^{eq}$ also demonstrates the weakest annual variations and varies in the range of 81-83 km. Nevertheless, in some cases, the OH equilibrium borderboundary lies slightly higher than the HO₂ borderboundary. In particular, it can be seen in April-August above 50°S, which can be explained by the difference between HO₂ and OH lifetimes ($\tau_{HO_2} < \tau_{OH}$), mainly, due to $k_{18} > k_{17}$. Comparing with Figure 2, one can see that the exclusion of the mentioned reactions from consideration results in the absence of the OH equilibrium areas at the low and middle mesosphere altitudes, as expected.

313

6 The <u>criterionscriteria</u> for HO₂ and OH equilibrium validity in the upper mesosphere and lower thermosphere

316 LetFirstly we determine HO₂ and OH lifetimes and the local time scales of HO_{2sh}^{eq} and OH_{sh}^{eq} , 317 according to Section 2.

318 From (8Eqs. (2-3) and (9), HO₂ lifetime and the local time scales of
$$HO_{2sh}^{eq}$$
 are as follows:

$$\begin{aligned} \tau_{HO_{2}} &= \frac{+}{k_{18} \cdot \theta + (k_{14} + k_{15} + k_{16}) \cdot H^{2}} \\ & (10) \\ \hline \\ 1 \\ \hline \\ k_{18} \cdot \theta + (k_{14} + k_{15} + k_{16}) \cdot H^{2} \\ \hline \\ 11) \end{aligned}$$

$$\begin{aligned} & (11) \\ & (12) \\ & (12)$$

329 Kulikov et al. (2023a) analyzed <u>analytically</u> the local nighttime evolution of O and H within the 330 framework of pure $HO_x - O_x$ chemistry and found the expression for $\frac{d}{dt} \left(\frac{O}{H}\right)$:

$$331 \quad \left| \begin{array}{c} \frac{d}{dt} \left(\frac{0}{H} \right) = -2 \cdot k_{20} \cdot M \cdot O_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}} \right) - k_{21} \cdot O_3 - k_{12} \cdot M \cdot O_2 \cdot \frac{\Theta \cdot O}{H \cdot H} \right|$$

$$332 \quad (1314)$$

$$333 \quad)$$

334 Thus, the expression (12Eq. (13) can be rewritten in the following form:

$$335 \quad \frac{dHO_{2}{}^{eq}_{sh}}{dt} = \frac{k_{18} \cdot k_{20} \cdot M \cdot O_{2} \cdot H^{2} \cdot (2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H})}{(k_{18} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H)^{2}}.$$

$$(14)$$

$$336 \quad \frac{k_{18} \cdot k_{20} \cdot M \cdot O_{2} \cdot H^{2} \cdot (2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H})}{(k_{10} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H)^{2}}.$$

$$(15)$$

By combining (8), (11Eqs. (9), (12), and (1415) we obtain the expression for the local time scales of HO_{2sh}^{eq} :

$$\begin{aligned} 340 \quad \left(\tau_{HO_{25h}^{eq}} \tau_{HO_{25h}^{eq}} = \frac{(k_{18} \cdot \theta + (k_{14} + k_{15} + k_{16}) \cdot H)}{k_{18} \cdot H \cdot (2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{\theta}{H})} \frac{(k_{18} \cdot \theta + (k_{14} + k_{15} + k_{16}) \cdot H)}{k_{18} \cdot H \cdot (2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{\theta}{H})} \\ 341 \quad (15\underline{16}) \quad ($$

343 Thus, taking into account (3), (10Eqs. (4), (11)) and (1516), the criterion for HO₂ equilibrium validity is 344 written in the form:

345
$$Crit_{HO_{2}} = \frac{\tau_{HO_{2}}}{\tau_{\tau_{HO_{2}sh}}^{eq}} = \frac{\frac{k_{18}\cdot H \cdot (2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H})}{(k_{18} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H)^{2}} \leq 0.1.$$
346
$$(16)$$

$$347 \quad \frac{k_{18} \cdot H \cdot (2 \cdot k_{20} \cdot M \cdot O_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}}\right) + k_{21} \cdot O_3 + k_{12} \cdot M \cdot O_2 \cdot \frac{O}{H})}{(k_{18} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H)^2} \le 0.1.$$
(17)

We calculated $Crit_{HO_2^2}$ using the global 3D chemical transport model, and included the zonally and monthly averaged lines $\langle Crit_{HO_2} \rangle = 0.1$ in Figure 95 (see magenta lines). One can see that, depending on the month, each redmagenta line well-reproduces well the lower borderboundary of the main OHHO₂ equilibrium area and repeatsfollows almost all its features and variations. Note, that, in zerothe zeroth order approximation, the criterion (1617) can be simplified as

353
$$\begin{vmatrix} Crit_{HO_2} \approx \left(2 \cdot k_{20} \cdot M \cdot O_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}}\right) + k_{21} \cdot O_3 + k_{12} \cdot M \cdot O_2 \cdot \frac{\Theta}{H}\right) \left(2 \cdot k_{20} \cdot M \cdot O_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}}\right) + k_{21} \cdot O_3 + k_{12} \cdot M \cdot O_2 \cdot \frac{\Theta}{H}\right) \cdot \frac{H}{k_{18} \cdot O^2} \leq 0.1.$$
355
$$(1718)$$
356
$$(1718)$$
357 From (9Eqs. (2-3) and (10). OH lifetime and the local time scales of OH^{eq} are as follows:

358
$$\tau_{OH} = \frac{1}{k_{17} \cdot 0} \frac{1}{k_{17} \cdot 0},$$

359 (1819
360)

$$361 \quad \tau_{OH_{sh}^{eq}} = \frac{OH_{sh}^{eq}}{|dOH_{sh}^{eq}/dt|}.$$

$$362 \qquad (1920)$$

)

Before determining the expression for dOH_{sh}^{eq}/dt , first of all, one should to keep in mind, that the 364 expression (910) depends on the HO₂ concentration. Above As previously mentioned, that near and above 365 the OH equilibrium border, boundary HO₂ is in equilibrium $(HO_2 \approx HO_{2sh}^{eq})$ and we can use expression 366 (8<u>Eq. (9</u>). In view of $k_{18} \cdot 0 \gg (k_{14} + k_{15} + k_{16}) \cdot H$, 367

368
$$HO_{2sh}^{eq} \approx \frac{k_{20} \cdot H \cdot M \cdot O_2}{k_{18} \cdot O} \left(1 - \frac{(k_{14} + k_{15} + k_{16}) \cdot H}{k_{18} \cdot O}\right).$$
369 (2021)

)

The substitution of (20Eq. (21) into (9Eq. (10) yields: 371

$$\begin{array}{cccc} 372 & OH_{sh}^{eq} = k_{20} \cdot H \cdot M \cdot O_{2} \cdot \frac{(1 + \frac{2 \cdot k_{14} \cdot H}{k_{18} \cdot O}) \cdot (1 - \frac{(k_{14} + k_{15} + k_{16}) \cdot H}{k_{18} \cdot O})}{k_{17} \cdot O} + \frac{k_{21} \cdot H \cdot O_{3}}{k_{17} \cdot O} \approx \\ 373 & \left| \begin{array}{c} \frac{k_{20} \cdot H \cdot M \cdot O_{2}}{k_{17} \cdot O} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + \frac{k_{21} \cdot H \cdot O_{3}}{k_{17} \cdot O}}{k_{17} \cdot O} \right| + \frac{k_{21} \cdot H \cdot O_{3}}{k_{17} \cdot O} \end{array} \right.$$

$$\begin{array}{c} (2122) \\ 375 \end{array} \right)$$

Thus, the expression for
$$dOH_{sh}^{eq}/dt$$
 is:

$$377 \qquad \frac{dOH_{sh}^{eq}}{dt} = \frac{d}{dt} \left(\frac{H}{O} \right) \cdot \left(\frac{k_{20} \cdot M \cdot O_2}{k_{17}} \cdot \left(1 + \frac{2 \cdot (k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O} \right) + \frac{k_{21} \cdot O_3}{k_{17}} \right) + \frac{k_{21} \cdot H}{k_{17} \cdot O} \frac{dO_3}{dt}.$$

$$(22223)$$

$$(22223)$$

Taking into account (13Eq. (14)) and the differential equation for O_3 time evolution: 380

381
$$\frac{dO_3}{dt} = k_{12} \cdot M \cdot O_2 \cdot O - k_{21} \cdot H \cdot O_3$$

the expression ($\frac{2123}{2}$) can be rewritten in following form: 382

$$\frac{dOH_{sh}^{eq}}{dt} = \frac{(2 \cdot k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H}) \cdot \frac{H^{2}}{O^{2}} (\frac{k_{20} \cdot M \cdot O_{2}}{k_{17}} \cdot (1 + \frac{2 \cdot (k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + \frac{k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H}) \cdot \frac{H^{2}}{O^{2}} (\frac{k_{20} \cdot M \cdot O_{2}}{k_{17}} \cdot (1 + \frac{2 \cdot (k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + \frac{k_{21} \cdot O_{3}}{k_{17}}) + \frac{k_{21} \cdot O_{3}}{O^{2}} \cdot (\frac{k_{20} \cdot M \cdot O_{2} \cdot (1 - \frac{k_{15} + k_{16}}{k_{18}}) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{O}{H}) \cdot H^{2}}{k_{17}} \cdot (1 + \frac{2 \cdot (k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + \frac{k_{21} \cdot O_{3}}{k_{17}}) + \frac{k_{21} \cdot H \cdot (k_{12} \cdot M \cdot O_{2} \cdot O - k_{21} \cdot H \cdot O_{3})}{k_{17} \cdot O}.$$

(2324

)

389 Thus, by combining (3), (18Eqs. (4)), (19), (2120), (22), and (2324) we obtain the expression for the 390 criterion for OH equilibrium validity:

397 We calculated $Crit_{OH}$, using the global 3D chemical transport model, and included the zonally and monthly averaged lines $\langle Crit_{OH} \rangle = 0.1$ in Figure <u>106</u> (see magenta lines). One can see that, depending 398 399 on the month, redthe magenta line almost everywhere reproduces the lower borderboundary of the OH 400 equilibrium area and repeats mainly its features and variations. Nevertheless, there are a few (by latitude) 401 narrow areas (in April-August near 70°S and in October-December near 70°N), where the criterion gives 402 a few km lower position of the OH equilibrium boundary, that is going to be these are discussed below.in 403 the next section. Note, that our numerical analysis shows that, in zeroth order approximation, the 404 criterion $(\frac{2425}{2})$ can be simplified as:

409

410 7 Discussion

411 Let's<u>We now</u> discuss obtained results and their possible applications.

412 Pay attention to the fact, that the presented results were plotted, using the lower threshold at 105° for the 413 nighttime solar zenith angle (χ) to exclude the twilight transition processes. Nevertheless, our additional

414 analysis revealed, that OH and HO₂ equilibrium conditions are fulfilled at $\chi > 95^{\circ}$. Evidently, during the

415	processing of the measurement data, taking twilight χ in (95°,105°) range into account extends the
416	latitude range of OH and HO ₂ equilibria application and allows us to include a noticeable part of the data
417	into consideration. However, in this case one should check for additional condition (Kulikov et al.,
418	<u>2023a):</u>
419	$e^{\int_{lt_{bn}}^{lt} \tau_{HO_2}^{-1} dt} \gg 1, e^{\int_{lt_{bn}}^{lt} \tau_{OH}^{-1} dt} \gg 1, $ (27)
420	where τ_{HO_2} and τ_{OH} are the HO ₂ and OH lifetimes, determined by Eqs. (11) and (19), lt is local time of
421	data, lt_{bn} is the local time at the beginning of the night. Mind, that at night O and H tend to decrease due
422	to the shutdown of the O_x and HO_x family photochemical sources, so τ_{HO_2} and τ_{OH} increase. Thus,
423	analyzing the measurement data one can apply more stringent conditions:
424	$e^{\frac{lt-lt_{bn}}{\tau_{HO_2}}} \gg 1, e^{\frac{lt-lt_{bn}}{\tau_{OH}}} \gg 1. $ (28)
425	The main results were obtained using a 3D model, where temperature and wind distributions are
426	updated every 24 hours. This excluded the influence of the atmospheric wave motion, in particular,
427	associated with tides, which is one of the main dynamical drivers in the tropical mesopause. We carried
428	out additional modeling with the distributions of the main characteristics, calculated by the Canadian
429	Middle Atmosphere Model for the year 2009 (Scinocca et al., 2008) with a 6-hourly frequency for
430	updating. The analysis of the time-height evolution of OH and HO ₂ , especially at low latitudes, showed
431	that our criteria reproduce quite well the local variations of the OH and HO ₂ equilibrium boundaries in
432	such conditions.
433	We evaluated the sensitivity of the presented HO ₂ and OH criteria ($Crit_{HO_2}$ and $Crit_{OH}$) to the
434	uncertainties of characteristics, involved in the expressions (17) and (25). The local heights of the OH and
435	<u>HO₂ equilibrium boundaries $(z_{HO_2}^{crit} \text{ and } z_{OH}^{crit})$ according to the criteria are determined as the</u>
436	<u>altitudes, at which $Crit_{HO_2} = 0.1$ and $Crit_{OH} = 0.1$ respectively. We considered the whole dataset of</u>
437	nighttime profiles, obtained by the numerical simulation of a one-year global evolution of mesosphere –
438	lower thermosphere, and estimated total uncertainties to determination of $z_{HO_2}^{crit}$ and z_{OH}^{crit} from each
439	local (in time and space) dataset (profiles of O, H, O ₃ , M, O ₂ and temperature). Following the typical
440	analysis presented, for example in Mlynczak et al. (2013a, 2014), each uncertainty was calculated as a
441	root sum square of the sensitivities to the individual perturbations of certain variables or parameters in the
442	expressions (17) and (25). The following uncertainties of the variables were used: 5K in the temperature
443	and 30% in O ₃ , O, and H. The uncertainties in reaction rates and their temperature dependencies were
444	taken from Burkholder et al. (2020). As the result (see Figure 7), the monthly and longitudinally mean of
445	total uncertainties in determination of $z_{HO_2}^{crit}$ and z_{OH}^{crit} were found varying in the range 0.02-1 km,
446	depending on altitude and season. Note, that these values are comparable with the typical height

447 resolution of satellite data. The latter allows us to consider our criteria as a robust instrument for 448 equilibrium condition validation. The main reason of relatively low sensitivity of $z_{HO_2}^{crit}$ and z_{OH}^{crit} is 449 the strong height-dependence of $Crit_{HO_2}$ and $Crit_{OH}$ near the value of 0.1.

As noted, Figures 9-10 presentFigs. 5-6 represent an interesting peculiarity. At the middle latitudes, 450 summer $z_{HO_{2}_{sh}}^{eq}$ and $z_{OH_{sh}}^{eq}$ are several kmremarkably higher than winter ones. Recently (, Kulikov et al.,) 451 452 (2023b) found such a feature was found in the evolution of nighttime ozone chemical equilibrium boundary, derived from SABER/TIMED data, which was accompanied by the same variation of 453 the transition zone-dividing, separating deep and weak photochemical oscillations of O and H, caused by 454 the diurnal variations of solar radiation. Kulikov et al. (2023b) The authors analyzed this effect 455 analytically and near and below the transition zone. It was shown firstly, that nighttime O decreases with 456 the characteristic time scale $\tau_0 = 0/|d0/dt|$ proportional to the 0/H ratio at the beginning of the night. 457 Secondly, during the summer the daytime O/H at the middle latitudes is remarkably less than the one in 458 winter. Consequently, summer values of τ_0 are significantly shorter than winter ones, so summer O 459 during the night decreases much faster than in winter. In our case lifetimes of HO₂ and OH are 460 proportional mainly to $\frac{1}{0}$ (see Eqs. (11) and (19)), so the summer rise of $z_{HO_{2sh}}^{eq}$ and $z_{OH_{sh}}^{eq}$ can be 461 explained by the markedly lower values of the O and H nighttime evolution times in summer than in 462 winter by virtue, mainly, of the lower values of the O/H-ratio during the night, which, in turn, is 463 464 determined by the daytime photochemistry. At middle, the ozone boundary varies within 4-5 km interval above 80 km, whereas the range of OH and HO₂ boundaries variations is 72-85 km (see Figures 9-10). In 465 the case of ozone, its criterion for equilibrium validity (see (5) in Kulikov et al. (2023b)) is as 466 follows:season difference in O diurnal evolution at these latitudes. 467

468
$$Crit_{\theta_3} = 2 \frac{k_{12} \cdot \theta_2 \cdot M}{k_{21}} (k_{20} \cdot M \cdot \theta_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}}\right) + k_{21} \cdot \theta_3) \cdot \frac{1}{k_{21} \cdot H \cdot \theta_3} \le 0.1.$$
(26)

469 At $\theta_3 \approx \theta_3 \frac{eq}{e_1} = \frac{k_{\pm2} \cdot M \cdot \theta_2 \cdot \theta}{k_{\pm2} \cdot H}$, one can see that $Crit_{\theta_3} \sim \frac{4}{\theta}$. It is follows from simplified expressions (17) and 470 (25) that $Crit_{H\theta_2}$ and $Crit_{\theta_H}$ are proportional to $\frac{H}{\theta^2}$. Such dependence leads to a stronger annual variation 471 of OH and HO₂-equilibrium boundaries than in the case of O₃.

As noted, there are a few narrow areas near 70°S₇/N (Figure 10)6), where the criterion (25) does not agree<u>correspond</u> well withto the OH equilibrium boundary. Our analysis revealed, that the main reason is neglecting the reaction OH+CO \rightarrow H+CO₂ as the source of H in the corresponding differential equation of its chemical balance. In order to improve the criterion₇ we revised the derivation of expression (1714) for $\frac{d}{dt} \left(\frac{\theta}{H} \frac{\theta}{H}\right)$ following to Kulikov et al. (2023a):

$$477 \quad \frac{d}{dt} \left(\frac{O}{H} \right) = -2 \cdot k_{20} \cdot M \cdot O_2 \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}} \right) - k_{21} \cdot O_3 - k_{12} \cdot M \cdot O_2 \cdot \frac{O}{H} - \frac{k_{32} \cdot CO}{k_{17} \cdot H} \cdot (k_{20} \cdot M \cdot O_2 \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_3)).$$

$$479 \quad (2729)$$

$$480 \quad)$$

481 As the<u>a</u> result, the corrected criterion for OH equilibrium validity is as follows:

$$482 \quad \begin{array}{l} \mathcal{C}rit_{\mathcal{OH}} \stackrel{m}{=} \frac{1}{k_{17}\cdot \theta} \cdot \left(\left(2 \cdot k_{2\theta} \cdot M \cdot \theta_{2} \cdot \left(1 - \frac{k_{15} + k_{36}}{k_{18}} \right) + k_{21} \cdot \theta_{3} + k_{12} \cdot M \cdot \theta_{2} \cdot \frac{\theta}{H} + \frac{k_{32}\cdot \theta}{k_{17}\cdot H} \cdot \left(k_{20} \cdot M \cdot \theta_{2} \cdot \left(1 - \frac{k_{15} + k_{36}}{k_{18}} \right) + k_{21} \cdot \theta_{3} + k_{21} \cdot \theta_{3} + k_{21} \cdot k_{12} \cdot M \cdot \theta_{2} \cdot \frac{\theta}{H} + \frac{k_{32}\cdot \theta}{k_{17}\cdot H} \cdot \left(k_{20} \cdot M \cdot \theta_{2} \cdot \left(1 + \frac{k_{14} - k_{15} - k_{16} \cdot Hk_{18} \cdot \theta + k_{21} \cdot \theta_{3} + k_{21} \cdot \theta_{3} + k_{21} \cdot k_{12} \cdot M \cdot \theta_{2} \cdot \theta_{3} + k_{21} \cdot k_{12} \cdot M \cdot \theta_{2} \cdot \left(1 + \frac{(k_{14} - k_{15} - k_{16} \cdot Hk_{18} \cdot \theta + k_{21} \cdot \theta_{3} + k_{21} \cdot \theta_{3} + k_{21} \cdot \theta_{3} + k_{21} \cdot \theta_{3} \right) \\ \end{array}$$

$$485 \quad Crit_{OH}{}^{m} = \frac{2 \cdot k_{20} \cdot M \cdot O_{2} \cdot \left(1 - \frac{k_{15} + k_{16}}{k_{18}}\right) + k_{21} \cdot O_{3} + k_{12} \cdot M \cdot O_{2} \cdot \frac{H}{H} + \frac{k_{32} \cdot CO}{k_{17} \cdot H} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} \cdot \frac{H}{O} \cdot 486 \quad \left(k_{20} \cdot M \cdot O_{2} \cdot \left(1 + \frac{2 \cdot (k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}\right) + k_{21} \cdot O_{3}\right) + \frac{k_{21} \cdot (k_{12} \cdot M \cdot O_{2} \cdot O - k_{21} \cdot H \cdot O_{3})}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} + \frac{k_{21} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} + \frac{k_{21} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3}}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} + \frac{k_{21} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} + \frac{k_{21} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})}{k_{17} \cdot O \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{14} - k_{15} - k_{16}) \cdot H}{k_{18} \cdot O}) + k_{21} \cdot O_{3})} + \frac{k_{21} \cdot (k_{20} \cdot M \cdot O_{2} \cdot (1 + \frac{(k_{20} \cdot M \cdot$$

We calculated this criterion, using the global 3D chemical transport model and included the zonally and monthly averaged lines $\langle Crit_{OH}{}^m \rangle = 0.1$ on the OH equilibrium maps (see Figure 118). One can see that including, the inclusion of this additional term actually eliminates the noted discrepancy between OH boundary and criterion. But, the application of this criterion requires CO data.

491 As noted in the Introduction, the conditions of nighttime OH and HO₂ equilibriumsequilibria together with one for O₃ equilibrium and their analytical criteria constitute thea useful tool for to-retrieval 492 493 of these components or other characteristics (for example, O and H) from measured data. At the altitudes 494 of upper mesosphere – lower thermosphere, these conditions can be applied, for example, to MLS/Aura database (measured characteristics: OH, HO₂, O₃, and CO), SMILES (HO₂ and O₃), SCIAMACHY 495 (O(¹S) green-line, O₂ A-band, and OH Meinel band emissions), SABER/TIMED (O₃, OH Meinel band 496 emissions at 2.0 μ m (9 \rightarrow 7 and 8 \rightarrow 6 bands) and at 1.6 μ m (5 \rightarrow 3 and 4 \rightarrow 2 bands)) and other, including to 497 improve improvement of existing retrieval approaches. In particular, Panka et al. (2021) proposed the 498 499 method of simultaneous derivation of O and OH at the levels v=0-9 from SABER data (volume emission rates at 2.0 and 1.6 µm), VER_{2µm} and VER_{1.6µm}) at 80-100 km, taking into account the equilibrium 500 condition for all states of OH. Such approach is valid for exited excited states due to itstheir very lowshort 501 lifetimes determined by radiative transitions and quenching with O₂, N₂, and O. In the case of the OH 502 ground state, its <u>lifetimes</u> determined by the reaction $OH+O \rightarrow H+O_2$ only. It means, that 503

504 Panka et al. (2021) used an equilibrium condition for total OH, which, as one can see from Figure 6, may be significantly disrupted above 80 km-in certain. On the other hand, there are latitude ranges and 505 seasons, as one can see from Figure 8. In order to check this assumption, we processed months, when the 506 507 OH equilibrium boundary drops remarkably below 80 km. Moreover, the Panka et al. (2021)method requires external data (about HO₂ since the reaction HO₂+O₇ \rightarrow OH(ν =9), +O₂, N₂, and temperature 508 509 profiles at 80-100 km) becomes the important source for 2009 and calculated local profiles of Crit_{rue} according to criterion (24). One can see this criterion depends of O, H, and O₃. Thus, the O₃ data was 510 taken from SABER data collocated (via the orbit number) with the OH below 87 km (Panka et al. profiles 511 in time and space. The H data was derived with the use of the equilibrium equation for OH(v=9):., 2021; 512 see also Figure 4 in our paper). 513

 $a_{7} \cdot \theta_{3} \cdot H = \theta H(9) \cdot (a_{1} + a_{2} \cdot N_{2} + a_{3} \cdot \theta_{2} + (a_{4} + a_{5} + a_{6}) \cdot \theta),$ 514 -(29)

The results of our paper allow modifying the Panka et al. method to extend its capabilities. The 515 simplest development of this method seems to be the following. First of all, note that the HO₂ equilibrium 516 condition (9) depends on H and O only and can be used within the self-consistent retrieval procedure, 517 considering the following system of equations: 518

519
$$OH(v = 1 - 9) = \frac{k_{12} \cdot H \cdot O_3 \cdot M \cdot f(v) + \sum_{v' > v} (a_1(v', v) + a_2(v', v) \cdot O_2 + a_3(v', v) \cdot N_2 + (a_4(v', v) + a_5(v', v)) \cdot O) \cdot OH(v')}{a_6(v)) \cdot O + \sum_{v > v'} (a_1(v, v') + a_2(v, v') \cdot O_2 + a_3(v, v') \cdot N_2 + (a_4(v, v') + a_5(v, v')) \cdot O)},$$
520
$$OH(0) = \frac{\sum_{v' > 0} (a_1(v', 0) + a_2(v', 0) \cdot O_2 + a_3(v', 0) \cdot N_2 + (a_4(v', 0) + a_5(v', 0)) \cdot O) \cdot OH(v') + k_{18} \cdot O \cdot HO_2 + 2 \cdot k_{14} \cdot H \cdot HO_2}{k_{17} \cdot O},$$

 $HO_{2} = \frac{k_{20} \cdot H \cdot M \cdot O_{2}}{k_{10} \cdot O + (k_{14} + k_{15} + k_{16}) \cdot H^{2}}$ 521

 $VER_{2\mu m} = a_1(9,7) \cdot OH(9) + a_1(8,9) \cdot OH(8),$ 522

 $VER_{1.6\mu m} = a_1(5,3) \cdot OH(5) + a_1(4,2) \cdot OH(4),$ 523

where $a_{1-7}a_{1-6}$ are the constant rates of the processes $OH(9v) \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, $OH(9v) + O_2 \rightarrow OH(v' \le v \le 8) + hv$, OH(9v) + O(9v) + O(9v524 $\underline{OH(v \le v') + O_2, OH(v)} + N_2 \rightarrow OH(\underbrace{\$v \le v'}) + N_2, OH(\underbrace{9} + O_2 \rightarrow OH(v \le \$) + O_2, OH(\underbrace{9} v) + O({}^3P) \rightarrow OH(v \le v') + O(\underbrace{1}_{v \le v'}) + O(\underbrace{1}_{v \ge v'}) + O(\underbrace{1}_{v \le v'}) + O(\underbrace{1}_{v \ge v'})$ 525 $OH(\underline{\nu' \leq \nu \leq 4-5}) + O(^{1}D), OH(\underline{9\nu}) + O(^{3}P) \rightarrow OH(\underline{\nu' \leq \nu \leq 8}) + O(^{3}P), and OH(\nu) + O(^{3}P) \rightarrow H + O_{2}, and H + O_{2}$ 526 $O_3 \rightarrow O_2 + OH(9)$, respectively. The values of a_{2-7} correspond to the Panka et al. (2021) model (see 527 528 Table 1 there), the Einstein coefficients for OH(v=9) were taken from Brooke et al. (2016). Due to the strong air concentration dependence Crit_{ar} decreases rapidlyTake into consideration, that this system 529 includes 13 equations with the height. From each Crit_{an} profile13 unknown variables. Therefore, the 530 solution to the system for a single set of the SABER measurements (simultaneously measured profiles of 531 532 O_3 , T, pressure, VER_{2um} , and $VER_{1.6um}$) gives one simultaneously retrieved profiles of O, H, OH(v=0-9), and HO₂. By applying the criteria (17) and (25) to obtained O and H profiles, we determined the local 533 height position of the OH equilibrium boundary $(z_{OH^{eq}})$ according to the condition $Crit_{OH} = 0.1$. It was 534 revealed that $Crit_{OH} < 0.1$ throughout the entire altitude range for most profiles. The Figure 12 plots the 535

found values of z_{OHer} above 80 km in different months. One can see that, in accordance of the Panka et 536 al. data, verify the local height position of the OH equilibrium boundary can rise up to 87 km. Moreover, 537 the Panka et al. method requires external data about HO₂ since the reaction HO₂+O \rightarrow OH+O₂ become 538 the important source for OH below 87 km (Panka et al., 2021; see also Figure 5 in our paper). Note that 539 thefulfillment of OH and HO₂ equilibrium condition (8) depends on Hconditions and O only and can be 540 used withindetermine the generalheight, below which the resulting profiles should be cut. More advanced 541 retrieval procedure of O, H, OH(v=0.9), and HO_2 would be statistical, based on Bayesian theorem, taking 542 543 into account the criteria (16) and (24).uncertainties in measurement data and rate constants. Similarly, for example, to Kulikov et al. (2018a), it should include a derivation of posterior conditional probability 544 545 density function of retrieved characteristics and numerical analysis of this function. Detailed development of this retrieval method is outside of this paper and should be carried out in a separate extended work. 546

548

547

549 8 Conclusions

550 The presented analysis shows, that there are extended areas in mesosphere and lower thermosphere, 551 where nighttime HO₂ and OH are close to their local equilibrium concentrations, determined mainly by the <u>reaction reactions</u> between $HO_x - O_x$ components <u>among</u> themselves and with H_2O_2 , N, NO, NO₂, and 552 CO. In upper mesosphere – lower thermosphere, the shortened expressions for their local equilibrium 553 concentrations are valid, including the $HO_x - O_x$ chemistry only. These conditions describes describe the 554 HO₂ and OH equilibrium from the top to the some lower bordersboundaries, the altitude position of 555 which vary in the interval between 73 and 85 km and depends essentially on the season and latitude. We 556 557 proposed analytical criteria, which almost everywhere well-reproduces quite well the main features of these bordersboundaries. Due to weak sensitivity to uncertainties of reaction rates and variables, these 558 criteria can be considered a robust instrument for HO₂ and OH equilibrium validation. The obtained 559 results allow extending the abilities of the Panka et al. (2021) method of retrieval ofto retrieve 560 unmeasured components from SABER data. The simultaneous application of OH and HO₂ equilibrium 561 562 conditions to the SABER data (O₃, volume emission rates at 2.0 and 1.6 µm) together with the OH and HO₂ criteria (16) and (24) to control this equilibrium validity is goingallows us to retrieve all unknown 563 564 $HO_x - O_x$ components (O, H, OH, and HO_2), extending) and to extend the altitude range of retrieval downward below 80 km and without external information. 565

567 568 569	Data availability. The Panka et al. <u>CMAM</u> data are obtained from the <u>SABER</u> website (https://saber.gats-inc.com). climate-modelling.canada.ca/climatemodeldata/cmam/cmam30/, last access: <u>18 May 2024).</u>						
570							
571	Code availability . Code is available upon request.						
572							
573	Author contributions. <u>Conceptualization:</u> MK-and, MB-carried out the data processing and analysis and						
574	wrote the manuscript., AC, SD, and AM contributed to reviewing the articleAF. Methodology: MK, AF.						
575	Investigation: MK, MB, AC, SD. Software: MB, AC. Visualization: MB, AC. Funding Acquisition: MK.						
576	Writing – original draft preparation: MK, MB. Writing – review & editing: AC, SD. Supervising: AF.						
577							
578	Competing interests. The authors declare that they have no conflict of interest.						
579							
580	Acknowledgements. The paper is in the memory of Prof. G.M. Fraiman. The authors are grateful to						
581	reviewers for providing valuable recommendations to improve the paper.						
582							
583	Financial support. This work was supported by The main results presented in Sects. 3-6 were obtained						
584							
585	with the support of the Russian Science Foundation under grant No. 22-12-00064 (https://rscf.ru/project/22-12-00064/) and/, last access: 18 May 2024). The analysis in Discussion was						
586	carried out at the expense of state assignment noNo. 0729-2020-0037.						
	<u>eurred out ut the expense of</u> state assignment no <u>rto</u> . 0725 2020 0057.						
587							
588	Supplement link:						
589							
590	References						
591	Avallone, L. M. and Toohey, D. W.: Tests of halogen photochemistry using in situ measurements of ClO						
592	and BrO in the lower polar stratosphere, J. Geophys. Res., 106, 10411–1042,						
593	https://doi.org/10.1029/2000JD900831, 2001.						
594	Belikovich, M. V., Kulikov, M. Yu, Grygalashvyly, M., Sonnemann, G. R., Ermakova, T. S., Nechaev,						
595	A. A., and Feigin, A .M.: Ozone chemical equilibrium in the extended mesopause under the nighttime						
596	conditions, Adv. Space Res., 61, 426–432, https://doi.org/10.1016/j.asr.2017.10.010, 2018.						
	20						
	20						

- Berger, U., and U. von Zahn (1999), Two level structure of the mesopause: A model study, *J. Geophys. Res.*, 104, 22,083–22,093.
- Brooke, J. S., Bernath, P. F., Western, C. M., Sneden, C., Afşar, M., Li, G., and Gordon, I. E.: Line
 strengths of rovibrational and rotational transitions in the X² II ground state of OH, Journal of Quantitative
 Spectroscopy and Radiative Transfer, 168, 142–157, https://doi.org/10.1016/j.jqsrt.2015.07.021, 2016.
- Burkholder, J. B., Sander, S. P., Abbatt, J., Barker, J. R., Cappa, C., Crounse, J. D., Dibble, T. S., Huie,
 R. E., Kolb, C. E., Kurylo, M. J., Orkin, V. L., Percival, C. J., Wilmouth, D. M., and Wine, P. H.:
 Chemical Kinetics and Photochemical Data for Use in Atmospheric Studies, Evaluation No. 19, JPL
 Publication 19-5, Jet Propulsion Laboratory, Pasadena, http://jpldataeval.jpl.nasa.gov, 2020.
- Cantrell, C. A., Mauldin, L., Zondlo, M., Eisele, F., Kosciuch, E., Shetter, R., Lefer, B., Hall, S., Campos, 606 607 T., Ridley, B., Walega, J., Fried, A., Wert, B., Flocke, F., Weinheimer, A., Hannigan, J., Coffey, M., Atlas, E., Stephens, S., Heikes, B., Snow, J., Blake, D., Blake, N., Katzenstein, A., Lopez, J., Browell, E. 608 V., Dibb, J., Scheuer, E., Seid, G., and Talbot, R.: Steady state free radical budgets and ozone 609 during TOPSE, J. Geophys. Res., photochemistry 108, TOP9-1-TOP9-22, 610 https://doi.org/10.1029/2002JD002198, 2003. 611
- Evans, W. F. J., and Llewellyn, E. J.: Atomic hydrogen concentrations in the mesosphere and the
 hydroxyl emissions, J. Geophys. Res., 78, 323–326, https://doi.org/1010.1029/JA078i001p00323, 1973.
- Evans, W. F. J., McDade, I. C., Yuen, J., and Llewellyn, E. J.: A rocket measurement of the O_2 infrared atmospheric (0-0) band emission in the dayglow and a determination of the mesospheric ozone and atomic oxygen densities, Can. J. Phys., 66, 941–946, https://doi.org/10.1139/p88-151. 1988.
- Fytterer, T., von Savigny, C., Mlynczak, M., and Sinnhuber, M.: Model results of OH airglow
 considering four different wavelength regions to derive night-time atomic oxygen and atomic hydrogen in
 the mesopause region, Atmos. Chem. Phys., 19, 1835–1851, https://doi.org/10.5194/acp-19-1835-2019,
 2019.
- Good, R. E.: Determination of atomic oxygen density from rocket borne measurements of hydroxyl
 airglow, Planet. Space Sci., 24, 389–395, https://doi.org/10.1016/0032-0633(76)90052-0, 1976.
- Grygalashvyly, M., Sonnemann, G. R., and Hartogh, P.: Long-term behavior of the concentration of the
 minor constituents in the mesosphere <u>Aa</u> model study, Atmos. Chem. Phys., 9, 2779–2792,
 https://doi.org/10.5194/acp-9-2779-2009, 2009.
- 626 Grygalashvyly, M., Sonnemann, G. R., Lübken, F.-J., Hartogh, P., and Berger, U.: Hydroxyl layer: Mean
- 627 state and trends at midlatitudes, J. Geophys. Res. Atmos., 119, 12,391–12,419, 628 https://doi.org/10.1002/2014JD022094, 2014.
- Grygalashvyly, M.: Several notes on the OH* layer, Ann. Geophys., 33, 923-930,
 https://doi.org/10.5194/angeo-33-923-2015, 2015.

- Hartogh, P., Jarchow, C., Sonnemann, G. R., and Grygalashvyly, M.: On the spatiotemporal behavior of
- ozone within the upper mesosphere/mesopause region under nearly polar night conditions, J. Geophys.
- 633 Res., 109, D18303, https://doi.org/10.1029/2004JD004576, 2004.
- Hartogh, P., Jarchow, Ch., Sonnemann, G. R., and Grygalashvyly, M.: Ozone distribution in the middle
- latitude mesosphere as derived from microwave measurements at Lindau (51.66°N, 10.13°E), J. Geophys.
- 636 Res., 116, D04305, https://doi.org/10.1029/2010JD014393, 2011.
- Körner, U., and Sonnemann, G. R.: Global 3D-modeling of water vapor concentration of the
 mesosphere/mesopause region and implications with respect to the NLC region, J. Geophys. Res., 106,
 9639–9651, https://doi.org/10.1029/2000JD900744, 2001.
- 640 Kowalewski, S., v. Savigny, C., Palm, M., McDade, I. C., and Notholt, J.: On the impact of the temporal
- variability of the collisional quenching process on the mesospheric OH emission layer: a study based on
 SD-WACCM4 and SABER, Atmos. Chem. Phys., 14, 10193-10210, https://doi.org/10.5194/acp-14-
- 642 SD-WACCM4 and SABER, Atmos. Chem. Phys., 14, 10193-10210, https://doi.org/10.5194/acp-14-643 10193-2014, 2014.
- Kremp, C., Berger, U., Hoffmann, P., Keuer, D., and Sonnemann, G. R.: Seasonal variation of middle
 latitude wind fields of the mesopause region—A comparison between observation and model calculation,
 Geophys. Res. Lett., 26, 1279–1282, https://doi.org/10.1029/1999GL900218, 1999.
- Kulikov, M. Y., Feigin, A. M., and Sonnemann, G. R.: Retrieval of the vertical distribution of chemical
 components in the mesosphere from simultaneous measurements of ozone and hydroxyl distributions,
 Radiophys. Quantum Electron., 49, 683–691, https://doi.org/10.1007/s11141-006-0103-4, 2006.
- Kulikov, M. Yu., Feigin, A. M., and Sonnemann, G. R.: Retrieval of water vapor profile in the
 mesosphere from satellite ozone and hydroxyl measurements by the basic dynamic model of mesospheric
 photochemical system, Atmos. Chem. Phys., 9, 8199–8210, https://doi.org/10.5194/acp-9-8199-2009,
- 6532009.
- Kulikov, M. Y., Belikovich, M. V., Grygalashvyly, M., Sonnemann, G. R., Ermakova, T. S., Nechaev, A.
- A., and Feigin, A. M.: Daytime ozone loss term in the mesopause region, Ann. Geophys., 35, 677-682
- 656 https://doi.org/10.5194/angeo-35-677-2017, 2017.
- 657 Kulikov, M. Y., Nechaev, A. A., Belikovich, M. V., Ermakova, T. S., and Feigin, A. M.: Technical note:
- 658 Evaluation of the simultaneous measurements of mesospheric OH, HO₂, and O₃ under a photochemical
- 659 equilibrium assumption a statistical approach, Atm. Chem. Phys., 18, 7453-747, 660 https://doi.org/10.5194/acp-18-7453-2018, 2018a.
- 661 Kulikov, M. Y., Belikovich, M. V., Grygalashvyly, M., Sonnemann, G. R., Ermakova, T. S., Nechaev, A.
- 662 A., and Feigin, A. M.: Nighttime ozone chemical equilibrium in the mesopause region. J. Geophys.
- 663 Res.,123, 3228–3242, https://doi.org/10.1002/2017JD026717, 2018b.
- 664 Kulikov, M. Yu., Nechaev, A. A., Belikovich, M. V., Vorobeva, E. V., Grygalashvyly, M., Sonnemann,
- G. R., and Feigin, A. M.: Border of nighttime ozone chemical equilibrium in the mesopause region from

- 666 saber<u>SABER</u> data: implications for derivation of atomic oxygen and atomic hydrogen, Geophys. Res.
 667 Lett., 46, 997–1004, https://doi.org/10.1029/2018GL080364, 2019.
- Kulikov, M. Y., Belikovich, M. V., Feigin, A. M.: The 2-day photochemical oscillations in the mesopause
 region: the first experimental evidence? Geophys. Res. Lett., 48, e2021GL092795,
 https://doi.org/10.1029/2021GL092795, 2021.
- Kulikov M.Yu., Belikovich, M.V., Grygalashvyly, M., Sonnemann, G. R., and Feigin, A.M.: Retrieving 671 daytime distributions of O, H, OH, HO₂, and chemical heating rate in the mesopause region from satellite 672 observations of ozone and OH* volume emission: The evaluation of the importance of the reaction 673 $H+O_3 \rightarrow O_2+OH$ Adv. Res., 69(9), 674 in the ozone balance. Space 3362-3373, 675 https://doi.org/10.1016/j.asr.2022.02.011, 2022a.
- Kulikov, M. Y., Belikovich, M. V., Grygalashvyly, M., Sonnemann, G. R., and Feigin, A.M.: The revised
 method for retrieving daytime distributions of atomic oxygen and odd-hydrogens in the mesopause region
- 678 from satellite observations, Earth, Planets and Space, 74, 44, https://doi.org/10.1186/s40623-022-01603679 8, 2022b.
- 680 Kulikov, M. Yu., Belikovich, M. V., Chubarov, A. G., Dementeyva, S. O., Feigin, A. M.: Boundary of 681 nighttime ozone chemical equilibrium in the mesopause region: improved criterion of determining the boundary from satellite Adv. Space Res., 71 (6), 2770-2780, 682 data, https://doi.org/10.1016/j.asr.2022.11.005, 2023a. 683
- Kulikov, M. Yu., Belikovich, M. V., Chubarov, A. G., Dementyeva, S. O., and Feigin, A. M.: Boundary
 of nighttime ozone chemical equilibrium in the mesopause region: long-term evolution determined using
 20-year satellite observations, Atmos. Chem. Phys., 23, 14593–14608, https://doi.org/10.5194/acp-2314593-2023, 2023b.
- Llewellyn, E. J., McDade, I. C. Moorhouse, P. and Lockerbie M. D.: Possible reference models for
 atomic oxygen in the terrestrial atmosphere, Adv. Space Res., 13, 135–144, https://doi.org/10.1016/02731177(93)90013-2, 1993.
- Llewellyn, E. J., and McDade, I. C.: A reference model for atomic oxygen in the terrestrial atmosphere,
 Adv. Space Res., 18, 209–226, https://doi.org/10.1016/0273-1177(96)00059-2, 1996.
- Manney, G. L., Kruger, K., Sabutis, J. L., Sena, S. A., and Pawson, S.: The remarkable 2003–2004 winter
- and other recent warm winters in the Arctic stratosphere since the late 1990s. J. Geophys. Res., 110,
 D04107, https://doi.org/10.1029/2004JD005367, 2005.
- 696 Marchand, M., Bekki, S., Lefevre, F., and Hauchecorne, A.: Temperature retrieval from stratospheric O3
- 697 and NO3 GOMOS data, Geophys. Res. Lett., 34, L24809, https://doi.org/10.1029/2007GL030280, 2007.
- 698 Marsh, D. R., Smith, A. K., Mlynczak, M. G., and Russell III, J. M.: SABER observations of the OH
- Meinel airglow variability near the mesopause, J. Geophys. Res., 111, A10S05,
 https://doi.org/10.1029/2005JA011451, 2006.

- McDade, I. C., Llewellyn, E. J., and Harris, F. R.: Atomic oxygen concentrations in the lower auroral
 thermosphere, Adv. Space Res., 5, 229–232, https://doi.org/10.1029/GL011I003P00247, 1985.
- McDade, I. C., and Llewellyn, E. J.: Mesospheric oxygen atom densities inferred from night-time OH
 Meinel band emission rates, Planet. Space Sci., 36, 897–905, https://doi.org/10.1016/00320633(88)90097-9, 1988.
- 706 Mlynczak, M. G., Marshall, B. T., Martin-Torres, F. J., Russell III, J. M., Thompson, R. E., Remsberg, E.
- E., and Gordley, L. L.: Sounding of the Atmosphere using Broadband Emission Radiometry observations
- of daytime mesospheric $O_2(^1D)$ 1.27 µm emission and derivation of ozone, atomic oxygen, and solar and
- chemical energy deposition rates, J. Geophys. Res., 112, D15306, https://doi.org/10.1029/2006JD008355,
 2007.
- 711 Mlynczak, M. G., Hunt, L. A., Mast, J. C., Marshall, B. T., Russell III, J. M., Smith, A. K., Siskind, D. E.,
- Yee, J.-H., Mertens, C. J., Martin-Torres, F. J., Thompson, R. E., Drob, D. P., and Gordley, L. L.: Atomic
- oxygen in the mesosphere and lower thermosphere derived from SABER: Algorithm theoretical basis and
- 714 measurement uncertainty, J. Geophys. Res., 118, 5724–5735, https://doi.org/10.1002/jgrd.50401, 2013a.
- 715 Mlynczak, M. G., Hunt, L. H., Mertens, C. J., Marshall, B. T., Russell III, J. M., López-Puertas, M.,
- Smith, A. K., Siskind, D. E., Mast, J. C., Thompson, R. E., and Gordley, L. L.: Radiative and energetic
- constraints on the global annual mean atomic oxygen concentration in the mesopause region, J. Geophys.
- 718 Res. Atmos., 118, 5796–5802, https://doi.org/10.1002/jgrd.50400, 2013b.
- 719 Mlynczak, M. G., Hunt, L. A. Marshall, B. T. Mertens, C. J. Marsh, D. R. Smith, A. K. Russell, J. M.
- 720 Siskind D. E., and Gordley L. L.: Atomic hydrogen in the mesopause region derived from SABER:
- Algorithm theoretical basis, measurement uncertainty, and results, J. Geophys. Res., 119, 3516–3526,
- 722 https://doi.org/10.1002/2013JD021263, 2014.
- 723 Mlynczak, M. G., Hunt, L. A., Russell, J. M. III, and Marshall, B. T.: Updated SABER night atomic
- oxygen and implications for SABER ozone and atomic hydrogen, Geophys. Res. Lett., 45, 5735–5741,
- 725 https://doi.org/10.1029/2018GL077377, 2018.
- Morton, K. W., and Mayers, D. F.; Numerical Solution of Partial Differential Equations, Cambridge
 University Press, 1994.
- Nikoukar, R., Swenson, G. R., Liu, A. Z., and Kamalabadi, F.: On the variability of mesospheric OH emission profiles, J. Geophys. Res., 112, D19109, https://doi.org/10.1029/2007JD008601, 2007.
- 730 Panka, P. A., Kutepov, A. A., Zhu, Y., Kaufmann, M., Kalogerakis, K. S., Rezac, L., et al.: Simultaneous
- retrievals of nighttime $O({}^{3}P)$ and total OH densities from satellite observations of Meinel band emissions.
- 732 Geoph. Res.Lett., 48, e2020GL091053, https://doi.org/10.1029/2020GL091053, 2021.
- Pendleton, W. R., Baker, K. D., Howlett, L. C.: Rocket-based investigations of $O({}^{3}P)$, O_{2} ($a^{1}\Delta_{g}$) and OH*
- 734 (v=1,2) during the solar eclipse of 26 February 1979, J. Atm. Terr. Phys., 45(7), 479-491, 1983.

- Siskind, D. E., Marsh, D. R., Mlynczak, M. G., Martin-Torres, F. J., and Russell III, J. M.: Decreases in
 atomic hydrogen over the summer pole: Evidence for dehydration from polar mesospheric clouds?
 Geophys. Res. Lett., 35, L13809, https://doi.org/10.1029/2008GL033742, 2008.
- Russell, J. P., and Lowe, R. P.: Atomic oxygen profiles (80-94 km) derived from Wind Imaging
 Interferometer/Upper Atmospheric Research Satellite measurements of the hydroxyl airglow: 1.
- 740 Validation of technique, J. Geophys. Res., 108(D21), 4662, https://doi.org/10.1029/2003JD003454, 2003.
- Russell, J. P., Ward, W. E., Lowe, R. P., Roble, R. G., Shepherd, G. G., and Solheim, B.: Atomic oxygen
- profiles (80 to 115 km) derived from Wind Imaging Interferometer/Upper Atmospheric Research Satellite
- measurements of the hydroxyl and green line airglow: Local time–latitude dependence, J. Geophys. Res.,
- 744 110(D15), D15305, https://doi.org/10.1029/2004JD005570, 2005.
- Siskind, D. E., Mlynczak, M. G., Marshall, T., Friedrich, M., Gumbel, J.: Implications of odd oxygen
- observations by the TIMED/SABER instrument for lower D region ionospheric modeling, J. Atmos. Sol.
- 747 Terr. Phys., 124, 63–70, https://doi.org/10.1016/j.jastp.2015.01.014, 2015.
- Smith, A. K., Lopez-Puertas, M., Garcia-Comas, M. and Tukiainen, S.: SABER observations of
 mesospheric ozone during NH late winter 2002–2009, Geophys. Res. Lett., 36, L23804,
 https://doi.org/10.1029/2009GL040942, 2009.
- Smith, A. K., Marsh, D. R. Mlynczak, M. G. and Mast, J. C.: Temporal variations of atomic oxygen in the
 upper mesosphere from SABER, J. Geophys. Res., 115, D18309, https://doi.org/10.1029/2009JD013434,
 2010.
- Scinocca, J. F., McFarlane, N. A., Lazare, M., Li, J., and Plummer, D.: The CCCma third generation
 AGCM and its extension into the middle atmosphere, Atmos. Chem. Phys., 8, 7055–7074,
 https://doi.org/10.5194/acp-8-7055-2008, 2008.
- Sonnemann, G., Kremp, C. Ebel, A. and Berger U.: A three-dimensional dynamic model of minor
 constituents of the mesosphere, Atmos. Environ., 32, 3157–3172, https://doi.org/10.1016/S13522310(98)00113-7, 1998.
- Sonnemann, G. R., Grygalashvyly, M., Hartogh, P., and Jarchow, C.: Behavior of mesospheric ozone
 under nearly polar night conditions, Adv. Space Res., 38, 2402–2407,
 https://doi.org/10.1016/j.asr.2006.09.011, 2006.
- Sonnemann, G. R., Hartogh, P., Grygalashvyly, M., Li, S., and Berger, U.: The quasi 5-day signal in the
 mesospheric water vapor concentration at high latitudes in 2003-a comparison between observations at
- ALOMAR and calculations, J. Geophys. Res., 113, D04101, https://doi.org/10.1029/2007JD008875,
 2008.
- Sonnemann, G. R., Hartogh, P., Berger, U., and Grygalashvyly, M.: Hydroxyl layer: trend of number
 density and intra-annual variability, Ann. Geophys., 33, 749–767, https://doi.org/10.5194/angeo-33-749-
- 769 2015, 2015.

- Stedman, D. H., Chameides, W., and Jackson, J. O.: Comparison of experimental and computed values
 for J(NO2), Geophys. Res. Lett., 2, 22–25, https://doi.org/10.1029/GL002i001p00022, 1975.
- Swenson, G. R., and Gardner C. S.: Analytical models for the responses of the mesospheric OH* and Na
 layers to atmospheric gravity waves, J. Geophys. Res., 103(D6), 6271–6294,
 https://doi.org/10.1029/97JD02985, 1998.
- Thomas, R. J.: Atomic hydrogen and atomic oxygen density in the mesosphere region: Global and
 seasonal variations deduced from Solar Mesosphere Explorer near-infrared emissions, J. Geophys. Res.,
 95, 16,457–16,476, https://doi. org/ 10. 1029/ JD095 iD10p 16457, 1990.
- Walcek, C. J.: Minor flux adjustment near mixing ratio extremes for simplified yet highly accurate
 monotonic calculation of tracer advection, J. Geophys. Res., 105, 9335-9348,
 https://doi.org/10.1029/1999JD901142, 2000.
- Xu, J., Smith, A. K., Jiang, G., Gao, H., Wei, Y., Mlynczak, M. G., and Russell III, J. M.: Strong 781 782 longitudinal variations in the OH nightglow, Geophys. Res. Lett., 37. L21801. 783 https://doi.org/10.1029/2010GL043972, 2010.
- Xu, J., Gao, H. Smith, A. K. and Zhu Y.: Using TIMED/SABER nightglow observations to investigate
 hydroxyl emission mechanisms in the mesopause region, J. Geophys. Res., 117, D02301,
 https://doi.org/10.1029/2011JD016342, 2012.

- **Table 1.** List of reactions included in 3-d chemical transport model with the corresponding reaction rates
- taken from Burkholder et al. (2020).

$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$						
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	1	$O(^{1}D)+O_{2} \rightarrow O+O_{2}$	24	$H+OH+N_2 \rightarrow H_2O+N_2$	47	$NO+O_3 \rightarrow NO_2+O_2$
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	2	$O(^{1}D)+N_{2} \rightarrow O+N_{2}$	25	$OH+H_2 \rightarrow H_2O+H$	48	$NO_2+O_3 \rightarrow NO_3+O_2$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	3	$O(^{1}D)+O_{3} \rightarrow O_{2}+2O$	26	$OH+OH \rightarrow H_2O+O$	49	$N+OH \rightarrow NO+H$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	4	$O(^{1}D)+O_{3} \rightarrow 2O_{2}$	27	$OH+OH+M \rightarrow H_2O_2+M$	50	$NO+HO_2 \rightarrow NO_2+OH$
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	5	$O(^{1}D)+N_{2}O \rightarrow 2NO$	28	$OH+HO_2 \rightarrow H_2O+O_2$	51	$H + NO_2 \rightarrow OH + NO$
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	6	$O(^{1}D)+N_{2}O \rightarrow N_{2}+O_{2}$	29	$H_2O_2+OH \rightarrow H_2O+HO_2$	52	$NO_3+NO \rightarrow 2NO_2$
9 $O(^1D)+CH_4 \rightarrow CH_3+OH$ 32 $OH+CO \rightarrow H+CO_2$ 55 $O_2+hv \rightarrow 2O$ 10 $O(^1D)+CH_4 \rightarrow H_2+CH_2O$ 33 $CH_4+OH \rightarrow CH_3+H_2O$ 56 $O_2+hv \rightarrow O+O(^1D)$ 11 $O+O+M \rightarrow O_2+M$ 34 $CH_3+O_2 \rightarrow CH_3O_2$ 57 $O_3+hv \rightarrow O_2+O$ 12 $O+O_2+M \rightarrow O_3+M$ 35 $CH_3+O \rightarrow CH_2O+H$ 58 $O_3+hv \rightarrow O_2+O(^1D)$ 13 $O+O_3 \rightarrow O_2+O_2$ 36 $CH_3O_2+NO \rightarrow CH_3O+NO_2$ 59 $N_2+hv \rightarrow 2N$ 14 $H+HO_2 \rightarrow 2OH$ 37 $CH_3O+O_2 \rightarrow CH_2O+HO_2$ 60 $NO+hv \rightarrow N+O$ 15 $H+HO_2 \rightarrow H_2O+O$ 38 $CH_2O \rightarrow H_2+CO$ 61 $NO_2+hv \rightarrow NO+O$ 16 $H+HO_2 \rightarrow H_2+O_2$ 39 $CH_2O \rightarrow H+CHO$ 62 $N_2O+hv \rightarrow N_2+O(^1D)$ 17 $OH+O \rightarrow H+O_2$ 41 $O_3+N \rightarrow NO+O_2$ 64 $NO_3+hv \rightarrow NO_2+O$ 18 $HO_2+O \rightarrow OH+O_2$ 41 $O_3+N \rightarrow NO_2+O_2$ 65 $H_2O+hv \rightarrow H+OH$ 20 $H+O_2+M \rightarrow HO_2+M$ 43 $O+NO+M \rightarrow NO_2+M$ 66 $H_2O_2+hv \rightarrow CH_2+H_2$ 21 $H+O_3 \rightarrow OH+O_2$ 44 $NO_2+O \rightarrow NO+O_2$ 67 $CH_4+hv \rightarrow CH_2+H_2+H_2$ 22 $OH+O_3 \rightarrow O_2+HO_2$ 45 $NO_2+O+M \rightarrow NO_3+M$ 68 $CH_4+hv \rightarrow CH+H_2+H_2+H_2$	7	$O(^{1}D)+H_{2}O \rightarrow 2OH$	30	$HO_2+HO_2 \rightarrow H_2O_2+O_2$	53	$N+NO \rightarrow N_2+O$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	8	$O(^{1}D)+H_{2} \rightarrow H+OH$	31	$HO_2+HO_2+M \rightarrow H_2O_2+O_2+M$	54	$N+NO_2 \rightarrow N_2O+O$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	9	$O(^{1}D)+CH_{4} \rightarrow CH_{3}+OH$	32	$OH+CO \rightarrow H+CO_2$	55	$O_2 + hv \rightarrow 2O$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	10	$O(^{1}D)+CH_{4} \rightarrow H_{2}+CH_{2}O$	33	$CH_4+OH \rightarrow CH_3+H_2O$	56	$O_2+hv \rightarrow O+O(^1D)$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	11	$O+O+M \rightarrow O_2+M$	34	$CH_3+O_2 \rightarrow CH_3O_2$	57	$O_3+hv \rightarrow O_2+O$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	12	$O+O_2+M \rightarrow O_3+M$	35	$CH_3+O \rightarrow CH_2O+H$	58	$O_3+hv \rightarrow O_2+O(^1D)$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	13	$O+O_3 \rightarrow O_2 + O_2$	36	$CH_3O_2+NO \rightarrow CH_3O+NO_2$	59	$N_2 + hv \rightarrow 2N$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	14	$H+HO_2 \rightarrow 2OH$	37	$CH_3O+O_2 \rightarrow CH_2O+HO_2$	60	$NO+hv \rightarrow N+O$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	15	$H+HO_2 \rightarrow H_2O+O$	38	$CH_2O \rightarrow H_2+CO$	61	$NO_2+hv \rightarrow NO+O$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	16	$H+HO_2 \rightarrow H_2+O_2$	39	$CH_2O \rightarrow H+CHO$	62	$N_2O+hv \rightarrow N_2+O(^1D)$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	17	$OH+O \rightarrow H+O_2$	40	$CHO+O_2 \rightarrow HO_2+CO$	63	$N_2O+hv \rightarrow N+NO$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	18	$HO_2+O \rightarrow OH+O_2$	41	$O_3+N \rightarrow NO+O_2$	64	$NO_3+hv \rightarrow NO_2+O$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	19	$H_2O_2+O \rightarrow OH+HO_2$	42	$NO_3+O \rightarrow NO_2+O_2$	65	$H_2O+hv \rightarrow H+OH$
$22 OH+O_3 \rightarrow O_2+HO_2 \qquad 45 NO_2+O+M \rightarrow NO_3+M \qquad 68 CH_4+h\nu \rightarrow CH+H_2+H$	20	$H+O_2+M \rightarrow HO_2+M$	43	$O+NO+M \rightarrow NO_2+M$	66	$H_2O_2 + hv \rightarrow 2OH$
	21	$H+O_3 \rightarrow OH+O_2$	44	$NO_2+O \rightarrow NO+O_2$	67	$CH_4 + hv \rightarrow CH_2 + H_2$
$23 HO_2+O_3 \rightarrow OH+2O_2 \qquad 46 N+O_2 \rightarrow NO+O \qquad 69 CO_2+hv \rightarrow CO+O$	22	$OH+O_3 \rightarrow O_2+HO_2$	45	$NO_2+O+M \rightarrow NO_3+M$	68	$CH_4+hv \rightarrow CH+H_2+H$
	23	$HO_2+O_3 \rightarrow OH+2O_2$	46	$N+O_2 \rightarrow NO+O$	69	$CO_2 + hv \rightarrow CO + O$

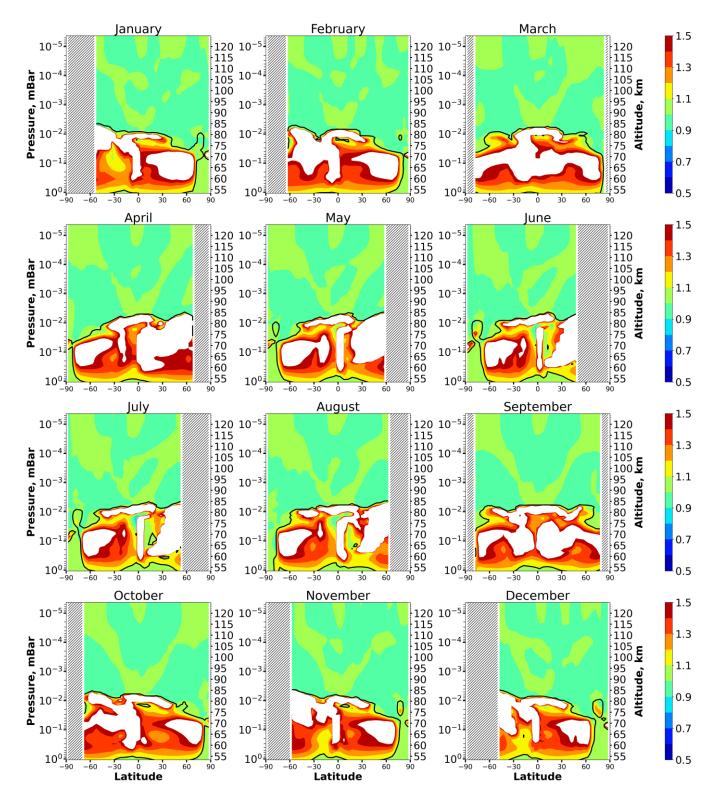


Figure 1. Nighttime mean and monthly averaged HO_2/HO_2^{eq} , where HO_2^{eq} is equilibrium concentration determined by Eq. (5). Black line shows the borderboundary of HO₂ equilibrium according to condition (1). The stippling corresponds to $\chi < 105^{\circ}$. The white area represents the $< HO_2/HO_2^{eq} >$ ratio outside the [0.5, 1.5] interval.

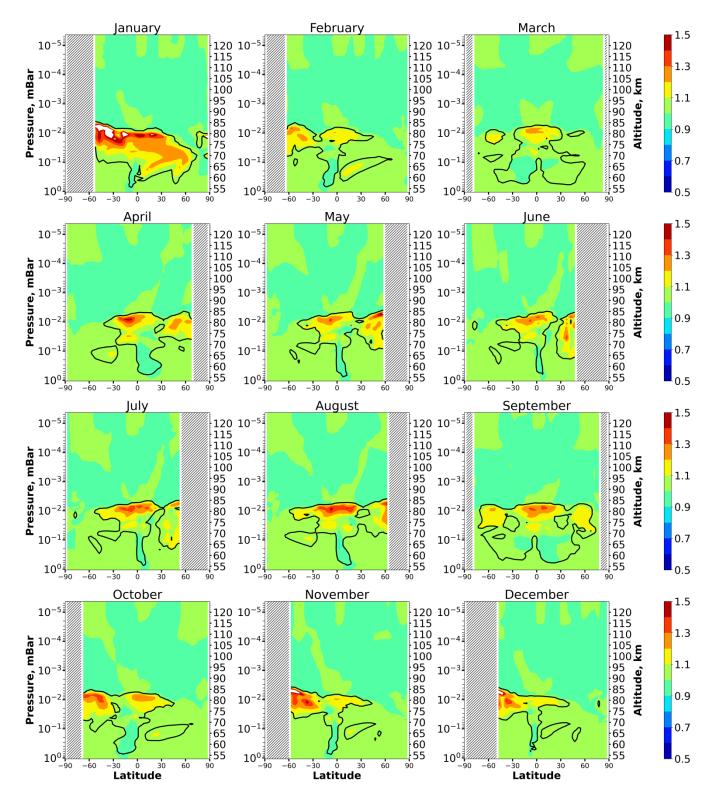
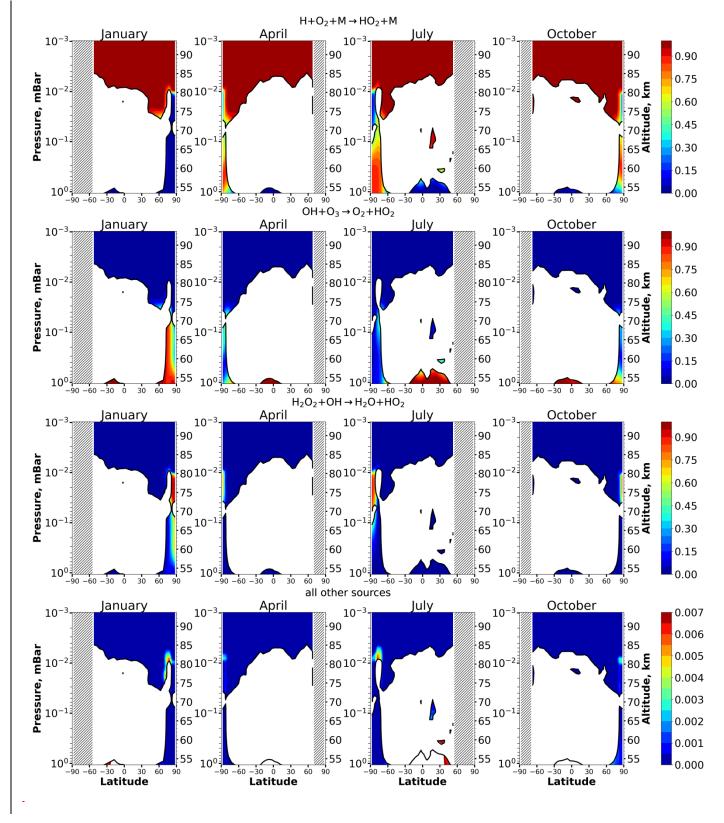


Figure 2. Nighttime mean and monthly averaged OH/OH^{eq} , where OH^{eq} is equilibrium concentration determined by Eq. (6). Black line shows the borderboundary of OH equilibrium according to condition (1). The stippling corresponds to $\chi < 105^{\circ}$. The white area represents the $< OH/OH^{eq} >$ ratio outside the [0.5, 1.5] interval.



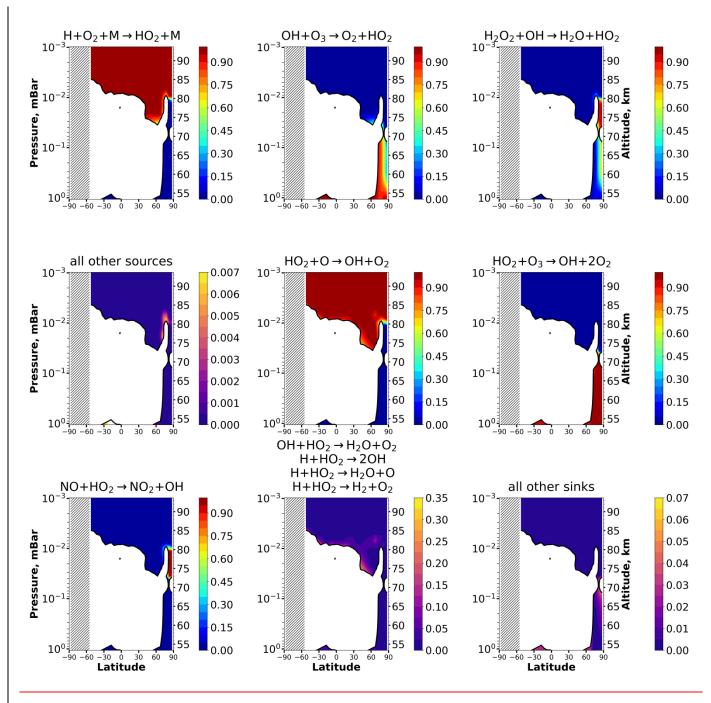
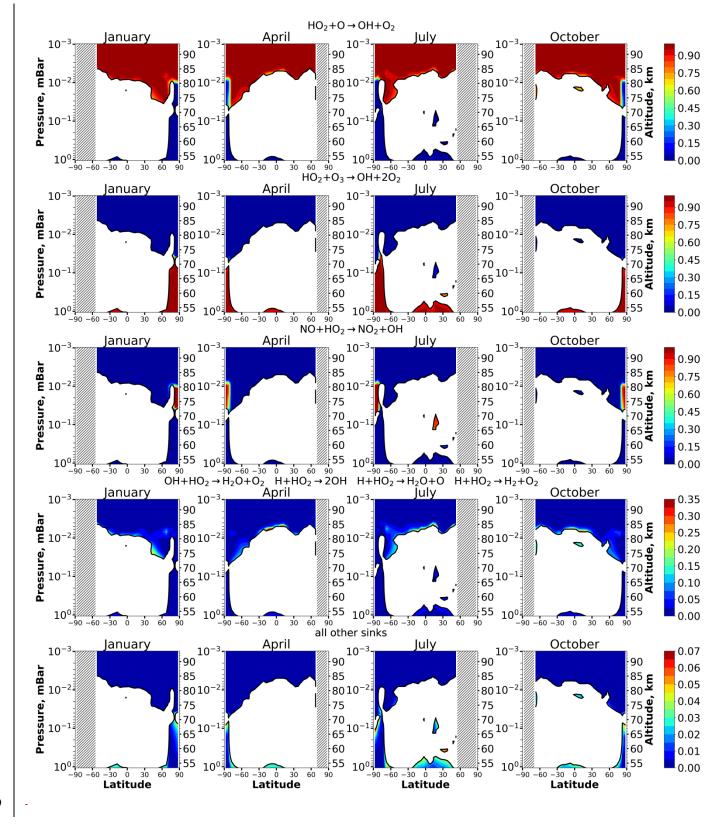
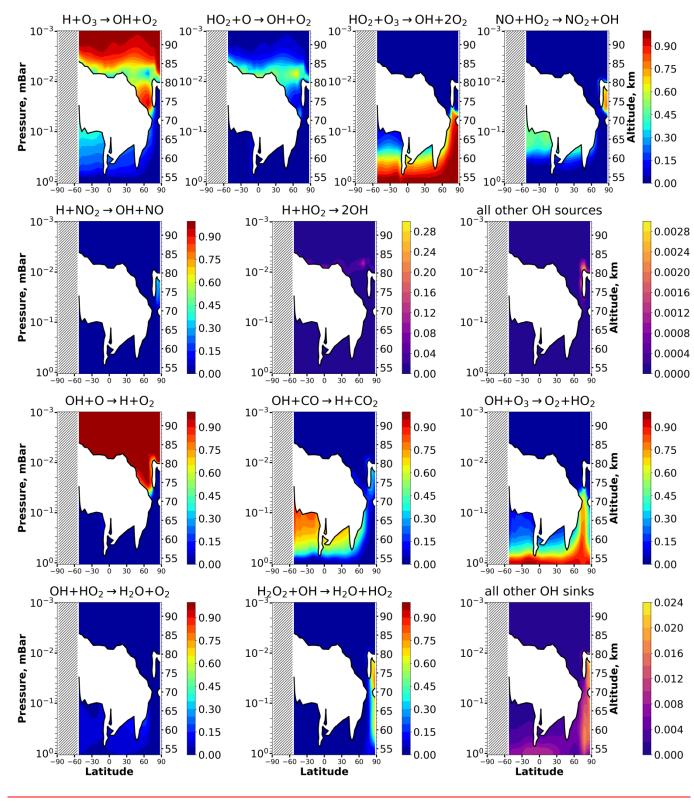


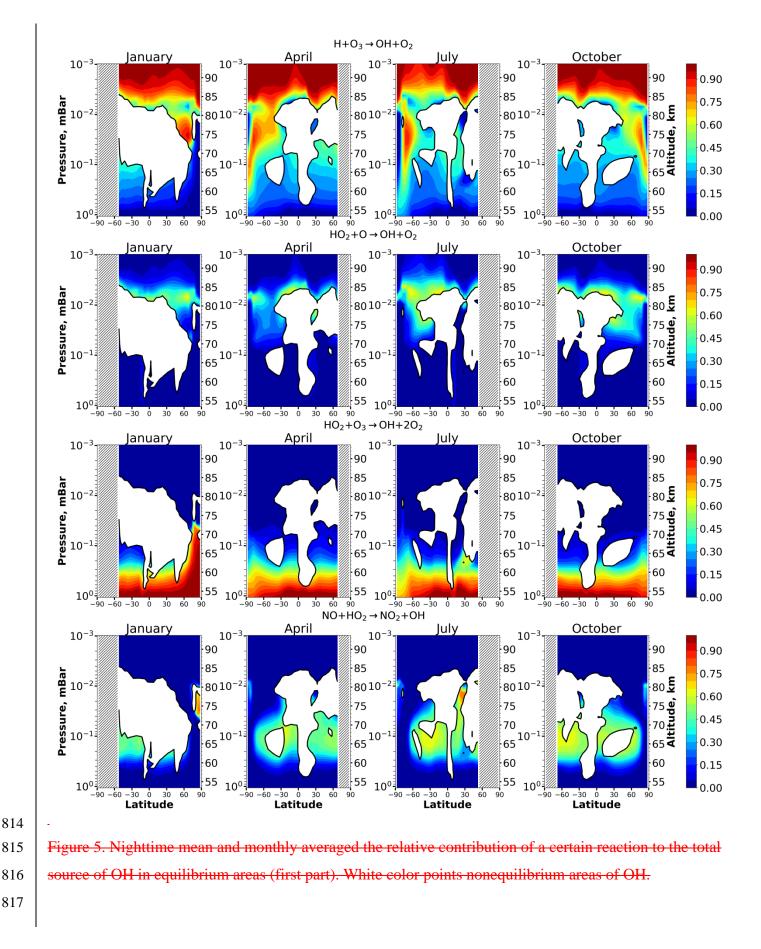
Figure 3. Nighttime mean and monthly averaged the relative contribution of a certain reaction to the total source or sink of HO₂ in equilibrium areas. The stippling corresponds to $\chi < 105^{\circ}$. White color points indicates nonequilibrium areas of HO₂.

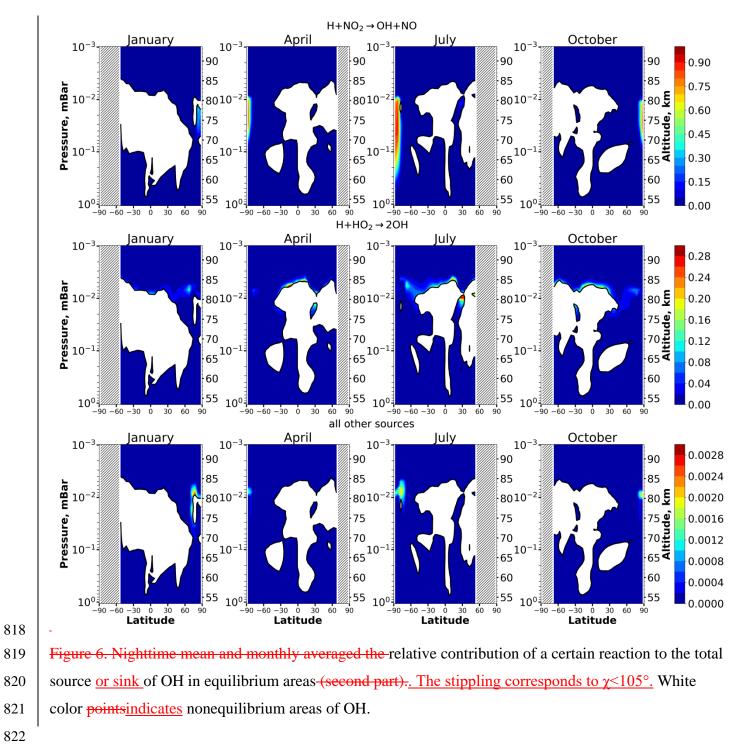


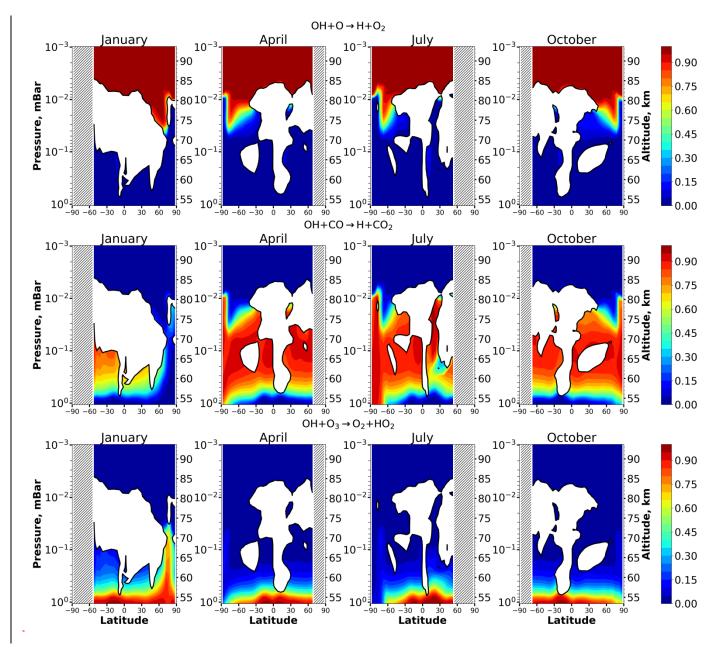


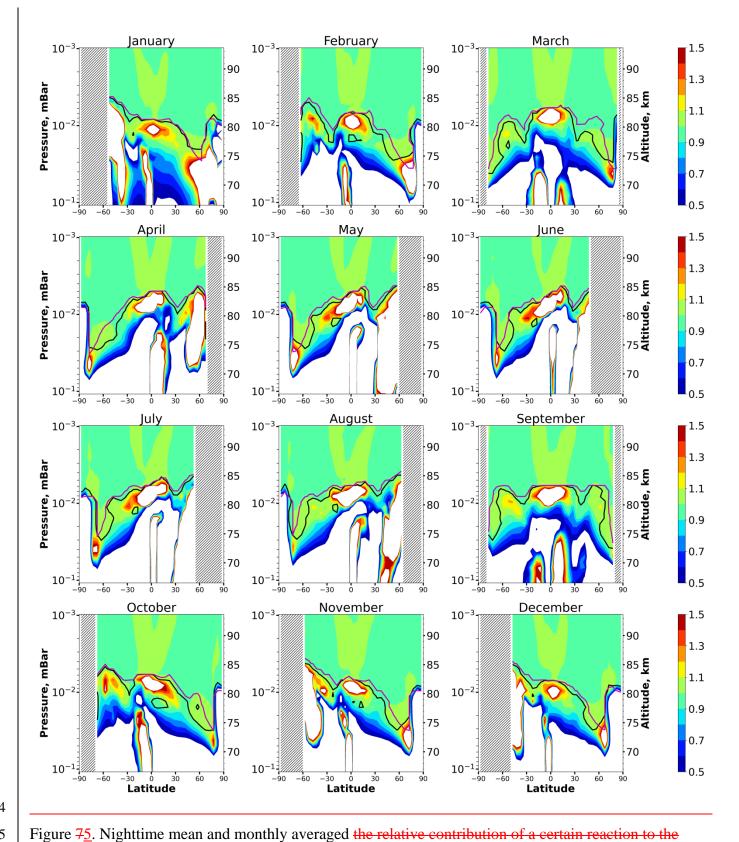


811Figure 4. Nighttime mean and monthly averaged the relative contribution of a certain reaction to the total812sink of HO_2 in equilibrium areas. White color points nonequilibrium areas of HO_2 .

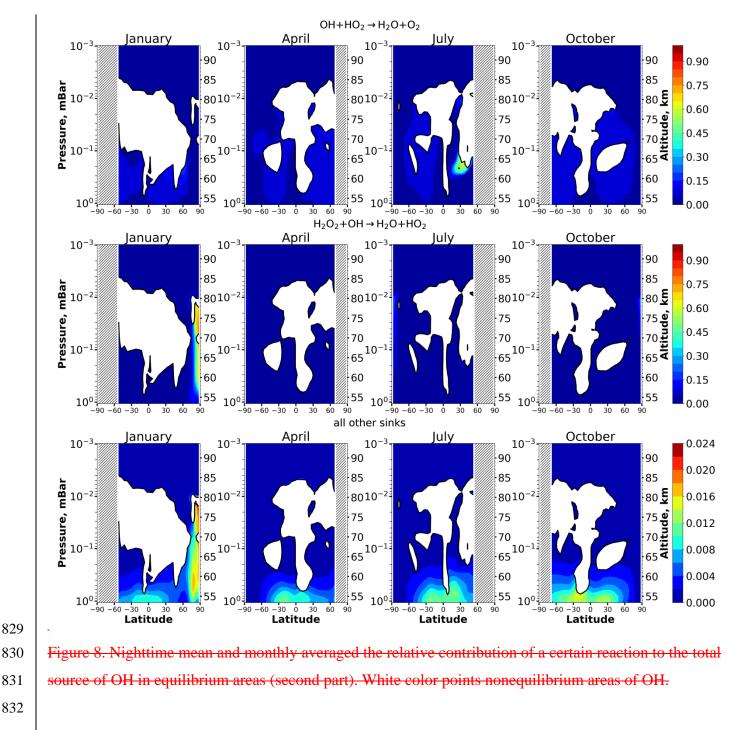


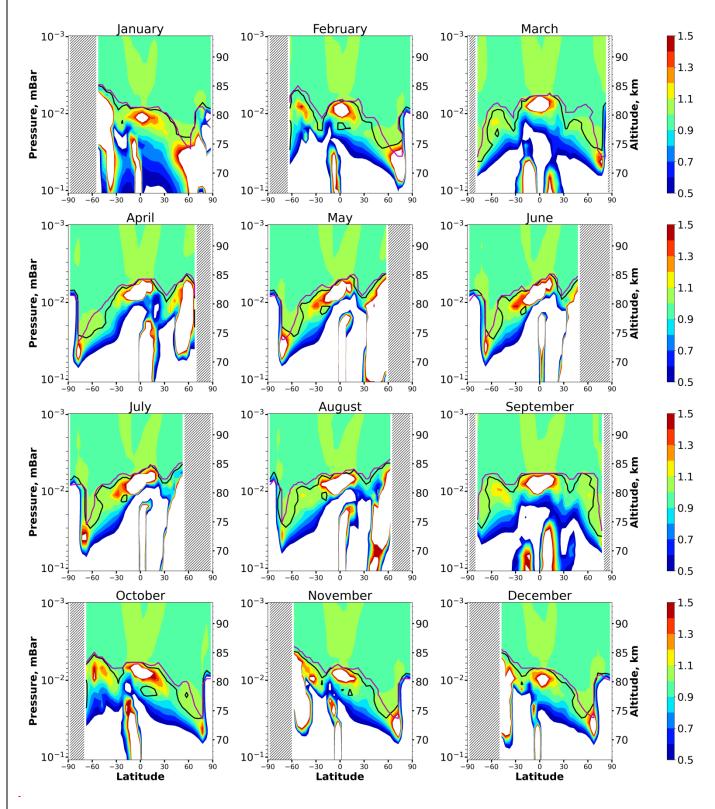






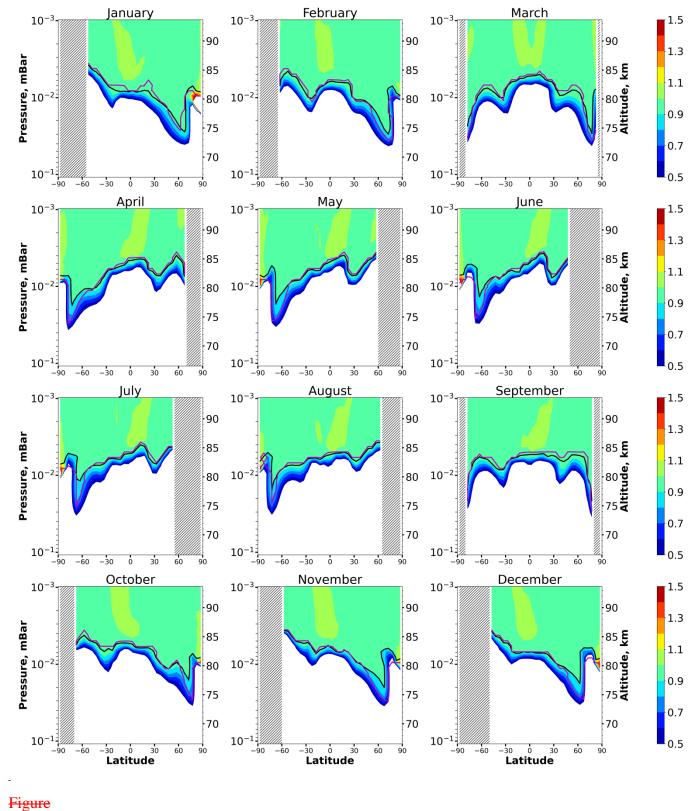
total sink of OH in equilibrium areas (first part). White color points nonequilibrium areas of OH.





833

834 Figure 9. Nighttime mean and monthly averaged HO_2/HO_{2sh}^{eq} , where $HO_2_{sh}^{eq}$ is shortened equilibrium 835 concentration determined by Eq. (9). Black line shows the borderboundary of HO₂ equilibrium according 836 to condition (1). Magenta line shows $< Crit_{HO_2} >= 0.1$. The stippling corresponds to $\chi < 105^\circ$. The white 837 area represents the $< HO_2/HO_2^{eq} >$ ratio outside the [0.5, 1.5] interval.



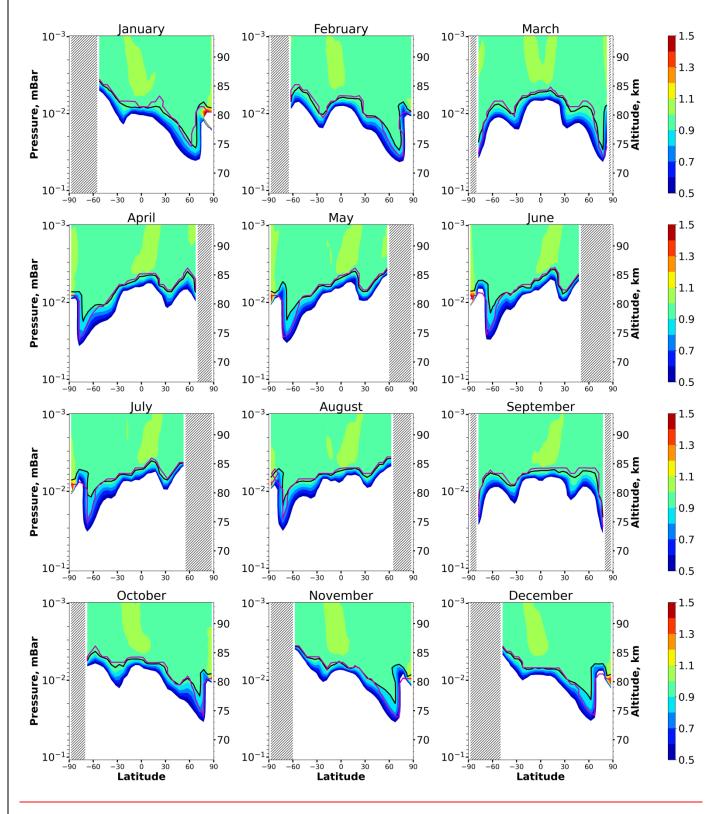
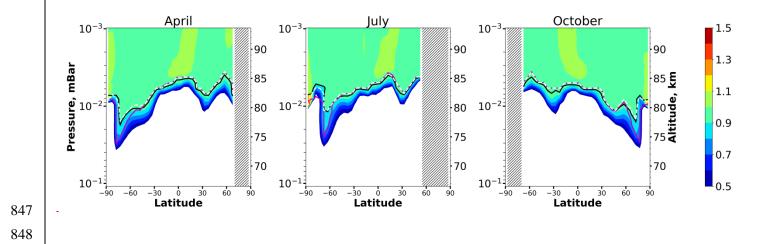


Figure 6. Nighttime mean and monthly averaged OH/OH_{sh}^{eq} , where OH_{sh}^{eq} is shortened equilibrium concentration determined by Eq. (10. Nighttime mean and monthly averaged OH/OH_{sh}^{eq} .). Black line shows the borderboundary of OH equilibrium according to condition (1). Magenta line shows < $Crit_{OH} \ge 0.1$. The stippling corresponds to $\chi < 105^{\circ}$. The white area represents the $< OH/OH^{eq} > ratio$ outside the [0.5, 1.5] interval.



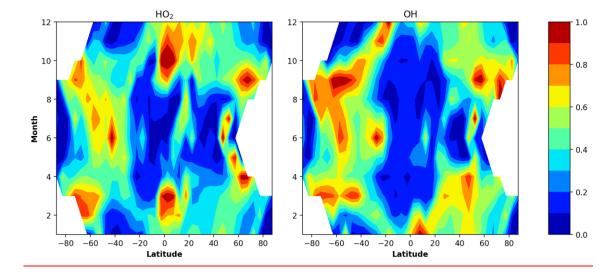
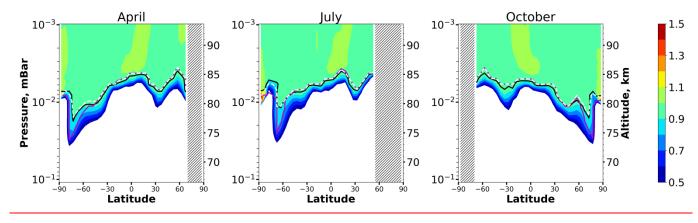
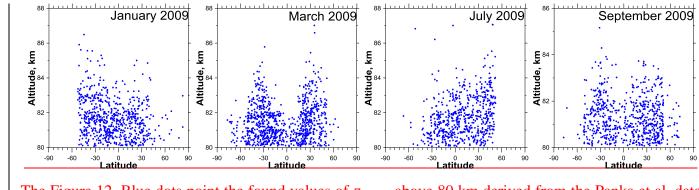


Figure 117. Monthly and longitudinally mean of total uncertainties in determination of the local heights
of the OH and HO₂ equilibrium boundaries according to the criteria (17) and (25). The white color
indicates the absence of data due to polar day.



855 Figure 8. Nighttime mean and monthly averaged OH/OH_{sh}^{eq} . Black line shows the borderboundary of OH 856 equilibrium according to condition (1). Magenta line shows $\langle Crit_{OH} \rangle = 0.1$, dotted white line shows 857 $\langle Crit_{OH}^{m} \rangle = 0.1$.

854





861 The Figure 12. Blue dots point the found values of $z_{OH_{sh}^{eq}}$ above 80 km derived from the Panka et al. data

in different months of 2009.