

# CO<sub>2</sub> and CO temporal variability over Mexico City from ground-based total column and surface measurements

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## Abstract.

Accurate estimates of greenhouse gas emissions and sinks are critical for understanding the carbon cycle and identifying key drivers of anthropogenic climate change. In this study, we investigate the variability of CO and CO<sub>2</sub> concentrations and their ratio over the Mexico City Metropolitan Area (MCMA) from long-term time-resolved columnar measurements at three stations, using solar absorption Fourier transform infrared spectroscopy (FTIR). Using a simple model and the mixed layer height derived from a ceilometer, we determined the CO and CO<sub>2</sub> concentration in the mixed layer from the total column measurements and found good agreement with surface cavity ring-down spectroscopy measurements. In addition, we used the diurnal pattern of CO columnar measurements at specific time intervals to estimate an average growth rate that, when combined with the space-based TROPOMI CO measurements, allowed deriving annual CO and CO<sub>2</sub> MCMA emissions from 2016 to 2021. A decrease of more than 50% of the CO emissions was found during the COVID19 lockdown period with respect to the year 2018. These results demonstrate the feasibility of using long-term EM27/Sun column measurements to monitor the annual variability of anthropogenic CO<sub>2</sub> and CO emissions in Mexico City without recourse to complex transport models. This simple methodology could be adapted to other urban areas if the orography favours low ventilation for several hours per day, which allows that column growth rate to be dominated by emission flux.

## 1 Introduction

The greenhouse gas (GHG) mitigation strategies implemented in megacities following the 1997 Kyoto Protocol and the 2015 Paris Agreement play a crucial role in the global action plan to mitigate climate change, given that cities are accountable for more than 70% of the global anthropogenic emissions (Duren and Miller, 2012). With the recent progress in space-based and ground-based remote GHG measurements in terms of accuracy, spatial coverage, resolution and temporal frequency, GHG emissions can increasingly be constrained by comparing bottom-up and top-down estimates. Top-down approaches are generally based on ground- or space-based atmospheric measurements coupled with inverse modelling, using 3D-Eulerian (i.e: WRF-Chem) or Lagrangian and hybrid (i.e: X-STILT, Hysplit) approaches (Wu et al., 2018, Che et al., 2022; Lian et al., 2023). The quantification of anthropogenic CO<sub>2</sub> enhancements from cities using satellite data e.g: GOSAT (Wang et al.,

2019), OCO-2 (Ye et al., 2020) or TanSat (Liu et al., 2018) is still challenging due to the sparsity of the observations, the low signal from the anthropogenic contribution compared to the background levels and biogenic contribution, and some inconveniences inherent to space-measurements such as the non-negligible aerosol effects (Wang et al., 2020 and references therein). Some studies have estimated the urban enhancements of anthropogenic CO<sub>2</sub> concentrations along with CO and NO<sub>2</sub> from satellite measurements, as these air pollutants can serve as tracers of anthropogenic CO<sub>2</sub> (Silva et al., 2013; Park et al., 2021 and references therein). The CO/CO<sub>2</sub> ratio is often used to determine the combustion efficiency of the cities (Park et al., 2021 and references therein). With the development of a new generation of space-based observatories, such as Sentinel-5P and OCO-2,3, the evolution of GHGs at the city scale can now be characterised with a finer temporal and spatial resolution (Kiel et al., 2021) but more validation efforts are needed. As inverse modelling is likely undermined by the approximations used for defining the emission patterns, transport processes and meteorology, top-down approaches may lead to discrepancies in emissions estimates, in particular in sites with complex orography.

Ground-based total column FTIR instruments provide valuable long-time concentration measurements of GHG and pollutant reactive species, as well as anthropogenic tracers, constituting a key element to validate regional and local inventories. Some studies reported estimates of CO<sub>2</sub> and CH<sub>4</sub> emissions from large urban areas (Babenhauserheide et al., 2018 in Tokyo; Hedelius et al., 2018 in the California Southern Coast Air Basin California megacity), using data from high-resolution FTIR instruments (i.e: Bruker IFS120/5HR) contributing to the Total Column Carbon Observing Network (TCCON). Nevertheless, only a few TCCON stations are located in urban areas (Toon et al., 2009; Chevallier et al., 2011; Sussman et al., 2020). The development of the Collaborative Carbon Column Observing Network (COCCON, Frey et al., 2019), using a new generation of portable low spectral resolution FTIR spectrometers (EM27/SUN, Gisi et al., 2012; Hase et al., 2016) able to simultaneously measure the CO<sub>2</sub>, CO, H<sub>2</sub>O and CH<sub>4</sub> average total columns with a similar quality as TCCON, has considerably densified the number of measurements in urban environments. Some studies reported emission estimates for big cities by means of the deployment of several EM27/SUN instruments at strategic sites throughout the cities (Hase et al., 2015 and Zhao et al., 2019 in Berlin; Vogel et al., 2019 in Paris; Makarova et al., 2021 in St Petersburg; Zhou et al., 2022 in Beijing and Xianghe; Che et al. 2022, in Beijing; Rißmann et al., 2022 for Munich) coupling columnar measurements with inverse modelling. Most of these studies were based on short-term campaign observations, applying the Differential Column Methodology (DCM, Chen et al., 2016) or dedicated dispersion models (Hase et al., 2016), coupled with simple mass balance-based methods or inverse modelling to derive emissions. Most of these studies reported significant discrepancies between the estimates, depending on the models used (Viatte et al., 2017).

In this study, we aimed to determine the Mexico City Metropolitan Area (MCMA) CO<sub>2</sub> and CO emissions using ground-based FTIR and surface measurements, without resorting to complex dispersion and/or chemistry transport models. The MCMA, with a population around 22 million inhabitants, is in the top ten most populous cities in the world and ranks among the major emitters of GHGs in North America. The available information of GHGs emission estimates are mainly based on the inventories reported by the Ministry of the Environment of Mexico City (SEDEMA), which is updated every two years, but lagging several years behind. In the report based on 2018, the latest published before the COVID19-lock-down (2020), a total emission of 75.2 Mt CO<sub>2</sub>-eq is estimated for the MCMA, 87% of which is attributed to fossil fuel combustion and 58% originates from the transport sector (SEDEMA Inventory, 2018). The Mexico City government is actively engaged in the C40 Climate

85 Change Program and implemented significant policy measures since 2008, including promoting sustainable  
86 transportation systems, implementing energy efficiency measures, increasing the use of renewable energy sources,  
87 and adopting green building practices. On a national scale, the country is committed to reduce its GHGs emissions  
88 by 35% by 2030 with respect to its base level, as stated in the last Nationally Determined Contributions report  
89 (NDC-2022, UNFCCC). To assess the effect of the national and local mitigation policies, the installation of  
90 ground-based GHG measurement networks and the refinement of bottom-up estimates by comparing them with  
91 the top-down method (i.e: inverse modelling) is of critical importance to obtain a comprehensive GHGs database  
92 that can serve as follow-up of the mitigation actions.

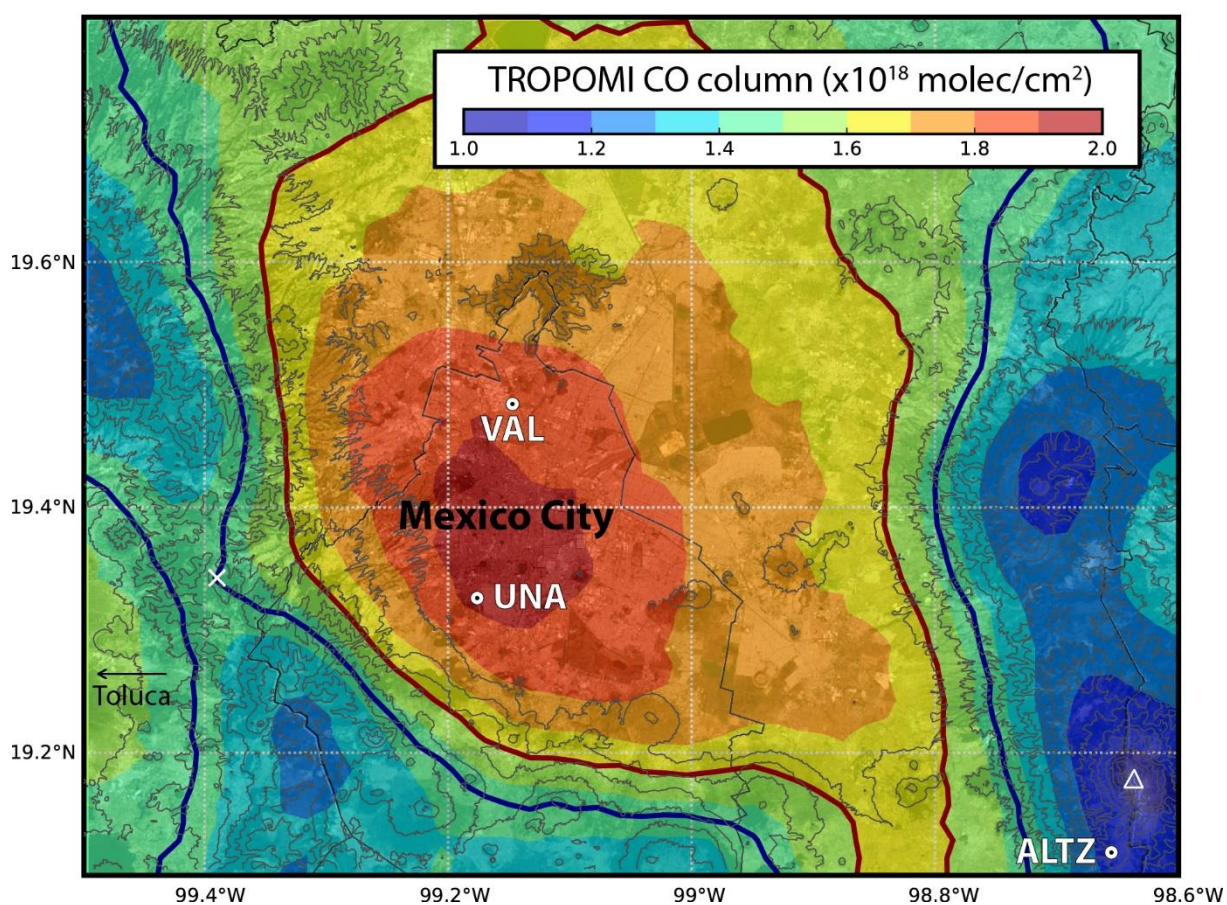
93 The Institute of Atmospheric Sciences and Climate Change (ICAyCC, Spanish acronym) at UNAM  
94 (Universidad Nacional Autónoma de México) deployed in the last decade a wide range of surface gas sensors and  
95 ground-based remote sensing instruments across the MCMA (Grutter, et al., 2003; Molina et al., 2010; Bezanilla  
96 et al., 2014; Stremme et al., 2009; 2013; Baylon et al., 2017) in the frame of research projects related to air quality  
97 assessment, **atmospheric monitoring and satellite products validation**. Since 2013, UNAM has contributed to the  
98 Network for the Detection of Atmospheric Composition Change (NDACC), performing continuous composition  
99 measurements of the free troposphere from the high altitude Altzomoni Atmospheric Observatory (ALTZ) station,  
100 located 60 km southeast of Mexico City at 3985 m a.s.l. Baylon et al., (2017) reported the background CO<sub>2</sub>  
101 variability and trend from this station between 2013 and 2016. Stremme et al., (2013) reported the first top-down  
102 estimate of carbon monoxide (CO) emissions for the MCMA, based on FTIR CO total column measurements and  
103 the Infrared Atmospheric Sounding Interferometer (IASI) data. These authors derived the CO<sub>2</sub> emissions for the  
104 MCMA using the CO emission estimates and the average CO/CO<sub>2</sub> ratio reported in Grutter (2003), using FTIR  
105 measurements. In 2018, the Mexican/French “Mexico City’s Regional Carbon Impacts (MERCICO2)” project  
106 (coordinated by UNAM and LSCE) was launched aiming to assess the CO<sub>2</sub> emissions from MCMA using  
107 EM27/SUN measurements and inverse modelling to evaluate the effectiveness of the mitigation strategies  
108 implemented by the local authorities. Xu et al., (submitted) examined the performance of a modelling system based  
109 on WRF-Chem to assess the whole-city emissions using the EM27/SUN measurements deployed in the frame of  
110 the MERCICO2 project. The complex orography of the region posed a challenge in the atmospheric transport  
111 simulations and thus for the top-down estimates using inverse modelling. Indeed, Mexico City is situated in a high  
112 altitude basin (~2300 m. a.s.l.), surrounded by mountains reaching up to 5.6 km a.s.l., and is prone to accumulate  
113 anthropogenic emissions, especially during the dry season, when the atmospheric boundary layer ventilation is  
114 limited (Burgos-Cuevas et al., 2023). The boundary layer dynamics in the basin and the wind surface circulation  
115 is complex, due to the temperature contrasts and rough topography.

116 In this study, we report the long-term (2013-2021) variability of the CO<sub>2</sub> and CO total columns and  
117 surface concentrations (from 2014) **over** the MCMA using ground-based FTIR and surface Cavity Ring-Down  
118 Spectroscopic (CRDS) measurements. Using the mixed layer height data from the continuous ceilometer  
119 measurements at UNAM, we examined the consistency of the surface and total column measurements of our  
120 network. We also determined an average CO/CO<sub>2</sub> ratio based on FTIR and surface measurements at different  
121 temporal resolutions (from daily to intraday). Then, using the spatial distribution of TROPOMI CO column  
122 measurements, we explore the potential of our FTIR network to capture the variability of the megacity CO and

123 CO<sub>2</sub> emissions using a simplified model, i.e.: without recourse to complex numerical simulations. Our estimates  
 124 are compared with the available bottom-up and previous top-down estimates.

## 125 2 Sites, instrumentation and measurement protocols

126 We used in this study the column-averaged dry-air mole fractions of CO<sub>2</sub> and CO (XCO<sub>2</sub> and XCO) from  
 127 three permanent FTIR stations distributed in a radius of 100 km around MCMA (Fig. 1), and the surface  
 128 measurements performed at UNA and ALTZ sites. The measurement periods for the different instruments at each  
 129 site are reported in Table 1. The VAL station is located at the northern part of the city in a highly industrialised  
 130 zone. The UNA station is situated at the south of the city in the main campus of UNAM. The third station is the  
 131 ALTZ background site (3985 m a.s.l.), located 60 km ESE from UNAM, within the Izta-Popo National Park. The  
 132 equipment of the different stations and measurement protocols are described in the following sub-sections.



133  
 134 **Figure 1: Map of the ALTZ, UNA and VAL stations and average distribution (2018-2022) of carbon monoxide total**  
 135 **columns over the Mexico City Metropolitan Area (MCMA) calculated from the TROPOMI CO product. Red and blue**  
 136 **contour lines represent the inner and outer area used to calculate the effective area (see details in text). The cross symbol**  
 137 **indicates the smallest CO total column value observed upwind of the city at the elevation of the Mexican basin, which is**  
 138 **used to estimate the background. The average total column can be decomposed into two main contributors: i) a**  
 139 **background of around  $1.45 \times 10^{18}$  molec.cm<sup>-2</sup> (limits represented by blue contour lines) and ii) the local influence**  
 140 **corresponding to the carbon monoxide emitted on the same day. The total columns are highly influenced by the**  
 141 **topography which is clearly visible over the highest terrains of the region, near to the Popocatepétl and Iztaccíhuatl**  
 142 **volcanoes at the south east of Mexico City. The mountains of Ajusco are located southwest of Mexico City. The**  
 143 **enhancement in the center of the metropolitan area reflects the carbon monoxide locally emitted on the same day.**

144

145

146 **Table1: Instrumentation and measurement periods used in this study.**

Station	Instrument	Measurement period	Product
ALTZ (19.119°N, 98.655°W 3.99 km a.s.l.)	IFS120/5HR	01/01/2013 - 01/06/2021	XCO and XCO <sub>2</sub>
	EM27/SUN #038	21/10/2020 - 20/12/2020 & 10/02/2021 - 22/02/2021	XCO and XCO <sub>2</sub>
	EM27/SUN #104	07/02/2020 - 18/02/2020	XCO and XCO <sub>2</sub>
	CRDS G2401 Picarro	15/11/2015 - 01/06/2021	Surface CO and CO <sub>2</sub>
UNA (19.326°N, 99.176°W 2.28 km a.s.l.)	Vertex	15/11/2015 - 20/06/2017	XCO
	EM27/SUN #038	07/05/2021 - 25/05/2021	XCO and XCO <sub>2</sub>
	EM27/SUN #062	17/03/2016 - 01/06/2017	XCO <sub>2</sub>
	EM27/SUN #104	01/06/2017 - 01/06/2021 04/04/2019 - 19/09/2019	XCO and XCO <sub>2</sub> XCO and XCO <sub>2</sub>
	CDRS G2401 Picarro	15/11/2015 - 01/06/2021	Surface CO and CO <sub>2</sub>
CL31 Vaisala ceilometer	15/11/2015 - 01/06/2021	Mixed Layer Height	
VAL (19.484°N, 99.147°W 2.26 km a.s.l.)	EM27/SUN #104	23/09/2019 - 01/06/2021	XCO and XCO <sub>2</sub>

147

148 **2.1 The UNA station: Total columns, surface concentrations and mixed-layer height measurements**

149 Atmospheric total columns of several gas species, such as O<sub>3</sub>, NH<sub>3</sub>, CH<sub>4</sub>, CO, and HCHO have  
150 continuously been measured at UNA since 2010 (Bezanilla et al., 2014; Plaza-medina et al., 2017; Baylon et al.,  
151 2017; Rivera-Cardenas et al., 2021; Herrera et al., 2022) using solar absorption FTIR spectroscopy.

152 Measurements are performed in the mid-infrared (MIR) and near-infrared (NIR) spectral ranges using a Bruker  
153 model Vertex 80 spectrometer. The instrument has a Maximum optical Path Difference (MPD) of 12 cm  
154 (corresponding to a spectral resolution of 0.075 cm<sup>-1</sup>) and is equipped with two detectors, a liquid-nitrogen cooled  
155 mercury-cadmium-telluride (MCT) and InGaAs detectors. Solar absorption measurements are performed using a  
156 home-built solar tracker. A full description of the instrumental set-up and measurement protocols is given in  
157 Bezanilla et al. (2014) and Plaza-Medina et al. (2017). The CO measurements are routinely performed in the MIR  
158 spectral range with a spectral resolution of 0.1 cm<sup>-1</sup>, using the MCT detector.

159 In March 2016, an EM27/SUN spectrometer was implemented at UNA to continuously measure XCO<sub>2</sub>,  
160 XCH<sub>4</sub>, XH<sub>2</sub>O, XCO total columns from solar NIR spectra with a spectral-resolution of 0.5 cm<sup>-1</sup> (MPD of 1.8 cm).  
161 The spectrometer is equipped with its own solar tracker (Bruker CAMTracker; Gisi et al., 2011) capturing and  
162 redirecting the solar beam into a RockSolid™ pendulum interferometer equipped with a Quartz beamsplitter. The  
163 EM27/SUN, with serial number #62 installed at the UNA station (hereafter EM27-SUN\_62), was initially operated  
164 with a standard InGaAs-diode detector sensitive to the 5500-11000 cm<sup>-1</sup> spectral range, to which a second InGaAs  
165 detector with Ge filter was added in 2017 for CO measurements through a second channel (4000 – 5500 cm<sup>-1</sup>)  
166 (Hase et al., 2016). Further details on the technical characteristics and systematic performance evaluation of the  
167 EM27/SUN spectrometer are given in Frey et al., (2019) and Alberti et al., (2022). The spectrometer was installed  
168 in a home-made protective box, including a remotely-controlled dome cover, a GPS and a PCE-THB-40 data-

169 logger for precise timing and surface pressure measurements. Double sided forward-backward interferograms are  
170 routinely recorded with a scanner velocity of 10 kHz, so that the recording time of one measurement (averaging  
171 10 IFGs scans) is close to one minute.

172 Additionally, CO<sub>2</sub>, CO, CH<sub>4</sub> and H<sub>2</sub>O surface measurements are continuously performed at the UNA  
173 station using a Cavity Ring-Down Spectrometer (CRDS, model G2401 from Picarro Inc.). The CRDS spectrometer  
174 uses a laser to quantify the spectral features of gas-phase molecules in an optical cavity offering effectively of up  
175 to 20 km absorption path length. Frequency shifts are prevented with a high-precision-wavelength monitor and  
176 temperature and pressure are precisely controlled by the analyzer. The quantification is improved by the  
177 simultaneous spectral analysis of the measured gases. A calibration system using 3 gas standards provided by the  
178 National Oceanic and Atmospheric Administration Earth System Research Laboratory (NOAA ESRL), traceable  
179 to the WMO2007 scale, was set up in 2018 at UNA and in 2019 at ALTZ. Data collected before the installation of  
180 the calibration systems were corrected with calibration coefficients obtained in 2018. The sampling inlet using  
181 Synflex tubing was placed at 24 m a.g.l. at UNA station and includes a Nafion air dryer, as described in detail by  
182 González del Castillo et al. (2022). Data are continuously collected at 0.3 Hz rate and their uncertainties, calculated  
183 as the standard deviation of raw data over 1-minute intervals when measuring calibration gases, are equal to 0.03  
184 ppm at UNA (González del Castillo et al., 2022).

185 Finally, continuous mixed-layer height (MLH) measurements are performed since 2008 at UNA using a  
186 CL31 ceilometer instrument (Vaisala). This is a robust commercial instrument which emits light pulses at 10 kHz  
187 repeating frequency at 910 nm using an indium-gallium-arsenide diode laser. It detects the backscatters signal  
188 through a single lens with a silicon avalanche photodiode. The resulting backscattering profiles have a vertical  
189 resolution of 10 m and reach an altitude of 7,500 m. The profiles have been used to retrieve MLH above the city  
190 since 2011 (García-Franco et al., 2018).

## 191 **2.2 The ALTZ background station: Total columns and surface measurements**

192 The Altzomoni Atmospheric Observatory (ALTZ) was equipped with a high-resolution FTIR  
193 spectrometer (model IFS120/5HR, Bruker) in 2012, capable of measuring atmospheric spectra in the NIR and MIR  
194 spectral regions with 257 cm MPD, equivalent to a spectral resolution of 0.0035 cm<sup>-1</sup>. The instrument is installed  
195 into a container with a motorised dome cover on the roof and a microwave communication system (60 km line-of-  
196 sight to the university campus), which allows a fully-remote control of the instruments. When the dome is open, a  
197 solar tracker (CAMTracker; Gisi et al., 2012) collects the solar beam and orients it toward the spectrometer  
198 entrance. The spectrometer can be operated with KBr or CaF<sub>2</sub> beam splitters, 3 different detectors (MCT, InSb,  
199 and InGaAs) and a set of 7 optical filters is installed in a rotating wheel. The measurement routine consists in the  
200 acquisition of high (0.005 cm<sup>-1</sup>), medium (0.02 cm<sup>-1</sup> and 0.1 cm<sup>-1</sup>) and low (0.5 cm<sup>-1</sup>) resolution spectra in the NIR  
201 and MIR spectral ranges using the different NDACC filters (~40 min for a complete sequence).

202 The NIR CO and CO<sub>2</sub> spectra (0.02 cm<sup>-1</sup>) used in this study were recorded as the average of two scans  
203 taken for approximately 38 s with a scanner speed of 40 kHz. The MIR CO spectra (0.005 cm<sup>-1</sup>) are deduced from  
204 the coaddition of 6 scans (<200 s) with a scanner speed of 40 kHz. Due to a spectrometer laser replacement, the  
205 IFS120/5HR measurements were interrupted between November 2020 and February 2021 (Table 1). To avoid an  
206 important gap in the measurements, an EM27/SUN (EM27/SUN\_38) was temporarily installed at the station

207 during this period. The intercalibration factors used for combining the two types of measurements were determined  
 208 from previous side-by-side measurements performed during February 2021 (see Table S1 and section 3.1.3).

209 A CRDS (model G2401 from Picarro Inc.) instrument was implemented at the station in 2014 providing  
 210 continuous CO<sub>2</sub>, CO, CH<sub>4</sub> and H<sub>2</sub>O surface measurements (González del Castillo et al., 2022). The sampling inlet  
 211 using Synflex tubing was placed at 4 m a.g.l. and includes a Nafion air dryer (similar installation to UNA). A  
 212 calibration system similar to that implemented at UNA, using 3 NOAA ESRL gas standards, was set up in 2019.  
 213 The station also includes meteorological instruments, pressure and temperature sensors and visible cameras among  
 214 other instrumentation for atmospheric and environmental monitoring.

### 215 2.3 The VAL station: Total column measurements

216 The VAL station, located in Vallejo in the northern part of MCMA, is part of the city's air quality network  
 217 (RAMA) run by SEDEMA. An EM27/SUN spectrometer (EM27/SUN\_104) was installed at this station in 2019  
 218 together with a surface CO<sub>2</sub> sensor. The VAL spectrometer has been performing measurements with the two  
 219 detectors since November 2019. **Additionally, the VAL site included a low-cost medium precision CO<sub>2</sub> sensor, as  
 220 a part of a network implemented during the MERCI-CO2 campaign. It consists of a NDIR-type of sensor  
 221 (SenseAir, model HPP3) that can measure in the 0 to 1000 ppm range and after a calibration and target gas follow-  
 222 up procedure, can produce data with <1% accuracy (Porrás et al., 2023).**

### 223 3.1 FTIR data processing and analysis

224 In this study, we used the solar absorption measurements acquired by five different FTIR instruments  
 225 (i.e: three EM27/SUN, a Vertex 80 and a IFS120/5HR) to estimate the XCO<sub>2</sub>, and XCO total columns at each  
 226 station. The retrieval strategies were adapted as a function of the spectral resolution and averaging kernel of each  
 227 species. Table 2 summarises the different products used in this study, and their retrieval parameters.

228

229 **Table 2: FTIR analysis: Description of the different FTIR products, retrieval strategies and parameters used in this study.**

Instrument (spectral resolution)	Gas	Microwindows (cm <sup>-1</sup> )	Interfering gases	Retrieval code	Retrieval method
EM27/SUN and IFS-120/5HR LowRes (0.5 cm <sup>-1</sup> )	CO <sub>2</sub>	6173.0 - 6390.0	H <sub>2</sub> O, CH <sub>4</sub>	PROFFAST	Scaling VMR COCCON strategy
	CO	4208.7 - 4318.8	H <sub>2</sub> O, HDO, CH <sub>4</sub> , HF		
	O <sub>2</sub>	7765.0 - 8005.0	H <sub>2</sub> O, CO <sub>2</sub> , HF		
IFS-120/5HR (0.02 cm <sup>-1</sup> ) (TCCON-type)	CO <sub>2</sub>	6180.0 - 6260.0 6310.0 - 6380.0	H <sub>2</sub> O, CH <sub>4</sub> , HDO	PROFFIT9.6	Scaling VMR
	CO	4208.7 - 4257.3 4262.0 - 4318.8	CH <sub>4</sub> , H <sub>2</sub> O, HDO		
	O <sub>2</sub>	7765.0 - 8005.0	H <sub>2</sub> O, CO <sub>2</sub> , HF		
IFS-120/5HR (0.005 cm <sup>-1</sup> ) (NDACC-type)	CO	2057.70 - 2058.00 2069.56 - 2069.76 2157.50 - 2159.15	O <sub>3</sub> , N <sub>2</sub> O, H <sub>2</sub> O, OCS and CO <sub>2</sub>	PROFFIT9.6	Profile NDACC strategy

Vertex80 (0.1 cm <sup>-1</sup> )	CO	2056.70 – 2059.00 2068.56-2069.77 2156.50-2160.15	O <sub>3</sub> , N <sub>2</sub> O, H <sub>2</sub> O, OCS and CO <sub>2</sub>	PROFFIT9.6	Profile
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231 **3.1.1 EM27/SUN spectra analysis**

232 Double-sided interferograms from the EM27/SUN were analysed following the standardised COCCON protocol,  
 233 using PREPROCESS and PROFFAST codes, developed by the KIT and made freely available ([https://www.imk-  
 235 asf.kit.edu/english/COCCON.php](https://www.imk-<br/>
  234 asf.kit.edu/english/COCCON.php)). The codes and retrieval methods are fully described in Sha et al. (2020), Frey  
 236 et al. (2021) and Alberti (2023) and only briefly summarised here. The PREPROCESS algorithm generates the  
 237 required spectra by a Fast Fourier Transform. The processing incorporates various quality checks, as a signal  
 238 threshold, intensity variations during recording, requirement of proper spectral abscissa scaling, and generates  
 239 spectra only from raw measurements passing all checks (the remaining ones being flagged). We used the ILS  
 240 parameters (i.e: modulation efficiency amplitude and phase error) reported on the KIT-COCCON website  
 241 (<https://www.imk-asf.kit.edu/english/COCCON.php>) and in Alberti et al. (2022), corresponding to the initial KIT  
 242 calibration of the spectrometers (Frey et al., 2019, Alberti et al., 2022). The PROFFAST-PCXS module (i.e:  
 243 forward model of PROFFAST) pre-calculates daily lookup tables of the molecular absorption cross-sections  
 244 according to the meteorological parameters and gas trace VMR profiles priors. The latest PROFFAST-PCXS  
 245 version uses the HITRAN 2020 spectroscopic linelists (with some extensions, e.g., line mixing parameters added  
 246 for CH<sub>4</sub>). Here, we used the standard COCCON linelists as incorporated in the previous PROFFAST version, i.e:  
 247 HITRAN 2008 for CH<sub>4</sub>, HITRAN 2012 for CO<sub>2</sub>, a modified version of HITRAN 2009 by Toon (2014) for H<sub>2</sub>O, a  
 248 TCCON standard linelist for O<sub>2</sub>, and the same solar line list as previously used by TCCON (compiled by G.C.  
 249 Toon for GGG2014). The least-squares fitting code PROFFAST-INVERS retrieves the total columns by scaling  
 250 the prior VMR profiles iteratively until adjusting the fit to the measured spectra. The intraday variability of surface  
 251 pressure is considered in the retrieval, interpolated from the in-situ pressure measurements. For tying the column-  
 252 averaged abundances provided by COCCON to TCCON data, PROFFAST applies post-process Airmass-  
 253 Dependent (ADCF) and Independent (AICF) corrections, independent from the instrument, similar as used in the  
 254 TCCON process (Sha et al., 2020, and Alberti, 2023). The corrections and parameters used are reported in the  
 255 COCCON website and Alberti, (2023).

256 We automatized and adapted the data processing to obtain a preliminary “real-time” hourly-updated  
 257 analysis (hereafter, AN1) for each site, additionally to the off-line treatment (hereafter, AN2) applying the standard  
 258 COCCON procedure. The meteorological data used in the AN1 retrieval were derived from the daily-available  
 259 radiosonde data, provided by Servicio Meteorologico Nacional (SMN) from measurements performed in the early  
 260 morning (6 AM LT) at the Mexico City International Airport. The AN1 strategy adopted fixed VMR priors for  
 261 each species, consisting in the averaged profile of 41 years (1980-2020) run of the Whole Atmospheric Community  
 262 Climate Model (WACCM), as commonly used in the NDACC community. The AN2 processing, generating the  
 263 COCCON standard products, used the daily TCCON meteorological data and priors (GGG2014 version of MAPS  
 264 files), downloaded from the Caltech server, which are based on National Centers for Environmental Prediction  
 265 (NCEP) reanalysis. For both AN1 and AN2 processing, we used the in situ intraday surface pressure measurements  
 from the PCE-THB-40 sensors. A correction factor was applied to the pressure measurements to take into account



266 the bias between the different pressure sensors used, previously intercompared by a few days of side-by-side  
267 measurements.

268  $\text{CO}_2$ ,  $\text{O}_2$  and  $\text{CO}$  were analysed in the 6173.0 - 6390.0  $\text{cm}^{-1}$ , 7765.0 - 8005.0  $\text{cm}^{-1}$  and 4208.7 - 4318.8  $\text{cm}^{-1}$   
269 spectral windows, respectively. The  $\text{XCO}_2$  and  $\text{XCO}$  column-averaged dry air mole fractions were calculated  
270 using the  $\text{O}_2$  retrieved total columns, according to Wunch et al. (2009):

$$271 \quad X_{gas} = 0.2095 (C_{gas} / C_{O_2}) \quad (1)$$

272 where  $C_{gas}$  and  $C_{O_2}$  are the target gas and  $\text{O}_2$  total columns, respectively.

273 The real-time (AN1) and COCCON (AN2)  $\text{XCO}_2$  and  $\text{XCO}$  products showed relative differences lower than 0.05%  
274 and 5%, respectively. The results presented hereafter are based on the official COCCON products (AN2 analysis).

### 275 3.1.2 Vertex80 and IFS120/5HR spectra analysis

276 High (0.005  $\text{cm}^{-1}$ ) and medium (0.02  $\text{cm}^{-1}$  and 0.1  $\text{cm}^{-1}$ ) resolution solar-absorption spectra are processed using  
277 the PROFFIT9.6 code (Hase et al., 2004).

278  $\text{XCO}_2$  is retrieved from the NIR 0.02  $\text{cm}^{-1}$  resolution spectra applying the procedure described in Baylon et al.  
279 (2017), in which two independent  $\text{CO}_2$  and  $\text{O}_2$  VMR-scaling retrievals are performed using fixed WCCAM VMR  
280 priors and NCEP-derived meteorological data. Spectral windows and interfering gases (Table 2) are similar to  
281 those used in the standard TCCON procedure.  $\text{XCO}_2$  is then calculated from the retrieved  $\text{CO}_2$  and  $\text{O}_2$  total columns  
282 by applying Eq. (1).

283 For the ALTZ analysis,  $\text{CO}$  was retrieved from the high (0.005  $\text{cm}^{-1}$ ) resolution spectra in the MIR region, applying  
284 the standard NDACC procedure (Pougatchev et al., 1994; Rinsland et al. 1998; Table 2). It uses a profile retrieval  
285 strategy with fixed WACCM VMR priors and NCEP meteorological data. Since the  $\text{O}_2$  specie is not analysed in  
286 the MIR region, the  $\text{XCO}$  was determined using the dry air columns ( $C_{dryair}$ ):

$$287 \quad XCO = \frac{C_{CO}}{C_{dryair}} \quad (2)$$

288 with:

$$290 \quad C_{dryair} = \left( \frac{P_g}{g} m_{dryair} \right) - \left( C_{H_2O} \frac{m_{H_2O}}{m_{dryair}} \right) \quad (3)$$

291 where  $C_{CO}$  and  $C_{H_2O}$  are the retrieved  $\text{CO}$  and  $\text{H}_2\text{O}$  total columns,  $g$  the column-averaged gravity acceleration,  $P_g$   
292 the ground pressure and  $m_{dryair}$  and  $m_{H_2O}$ , the dry air and  $\text{H}_2\text{O}$  molecular masses respectively. In addition, we  
293 analysed  $\text{XCO}$  from the NIR spectral region to complement the MIR time-series, occasionally interrupted when  
294 the liquid nitrogen was missing at the station. The  $\text{CO}$  and  $\text{O}_2$  columns in the NIR region were analysed using  
295 scaling retrievals in the same spectral windows as that used by TCCON (Table 2), but with fixed WACCM VMR  
296 priors and NCEP meteorological data.  $\text{XCO}$  was calculated from the  $\text{CO}$  and  $\text{O}_2$  retrieved total columns applying  
297 Eq. (1). To minimise the air mass dependence effect (likely low for  $\text{CO}$ ), we filtered out data with a SZA  $>60^\circ$ .  
298  $\text{XCO}$  NIR and MIR products were compared and intercalibrated (section 3.1.3).

299 For UNA, we used the  $\text{XCO}$  total columns calculated from the Vertex80 measurements to complement the  
300 EM27/SUN time series during the period when it was operating with a single detector (between March 2016 and  
301 September 2017).  $\text{CO}$  was analysed from the 0.1  $\text{cm}^{-1}$  resolution spectra in the MIR spectral range, using a standard  
302

303 NDACC profile retrieval strategy and the PROFFIT9.6 retrieval program with constant WACCM VMR priors and  
304 NCEP meteorological data. Spectral windows (Table 2) were adapted following Pougatchev and Rinsland (1995).  
305 Previous CO total columns time series retrieved from the same method at UNA were presented in Garcia-Franco  
306 et al. (2018) and Borsdorff et al. (2018, 2020). Only the constraint of these CO retrievals were adjusted for the  
307 Megacity and allowed in addition a free fitting of the mixing layer concentration, following the work by Stremme  
308 et al. (2009) in which low resolution MIR spectra with a different retrieval program have been analysed.

### 309 **3.1.3 Measurement precision and FTIR product intercomparison**

310 Side-by-side measurements were performed at the ALTZ and UNA stations on several occasions (Table1)  
311 to assess the FTIR measurement precisions, to characterise the bias between the different products and to define  
312 the inter-calibration factors for the XCO<sub>2</sub> and XCO products. We used the EM27/SUN\_62 products as reference,  
313 for which we previously applied the standard XCO<sub>2</sub> and XCO calibration factors reported in Alberti et al. (2022),  
314 to inter-calibrate our results with the COCCON network and the Karlsruhe TCCON station operated by KIT. The  
315 linear regression parameters from the different measurement pairs and the calibration factors are presented in the  
316 Supplementary data (Table S1 and S2).

317 We found a bias lower than 0.2% and 1.0% between the three EM27/SUN, for XCO<sub>2</sub> and XCO respectively, and  
318 a coefficient of determination (R<sup>2</sup>) higher than 0.99.

319 On the other hand, the precision of the EM27/SUN measurements was assessed by calculating the standard  
320 deviation over a 5 min-interval period, and found to be on average 2.7 ppb and 0.3 ppm for XCO and XCO<sub>2</sub>,  
321 respectively.

322 The intercomparison of the IFS120/5-HR high resolution (0.02 cm<sup>-1</sup>) products and the EM27/SUN XCO<sub>2</sub> products  
323 was performed for the daily average data used in this study. The calibration factors were determined using i) the  
324 EM27/SUN XCO<sub>2</sub> products and ii) the IFS 120/5-HR low resolution (0.5 cm<sup>-1</sup>) product (Fig. S2), processed in the  
325 same way as the COCCON EM27/SUN data but having the advantage of being measured even outside the  
326 campaigns carried out with the EM27/Sun. We finally found a bias around 0.4% (slope=0.996), and a coefficient  
327 of determination R<sup>2</sup> of 0.92. This bias is of order of that expected when comparing TCCON and COCCON products  
328 (Frey et al., 2019), when no empirical calibration is applied. On the other hand, a bias of 2% (and R<sup>2</sup>=0.92) was  
329 found comparing the XCO from the EM27/SUN and the Vertex (MIR) products at UNA.

330 One of the main contributions of the apparent bias observed when comparing products from different instruments  
331 and using different retrieval strategies can be due to their respective Averaging Kernel (AK) which characterise  
332 the smoothing error. It is especially the case in the comparison of XCO from the EM27/SUN (i.e: NIR scaling  
333 retrieval product, Degree Of Freedom (DOF) =1) and from the Vertex (MIR profile-product, DOF > 2). To assess  
334 this effect, we refined the comparison after smoothing the vertically resolved Vertex profiles with the EM27/SUN  
335 AK (following Rodgers, 2000; Borsdorff et al., 2014, 2018) and re-calculating the smoothed Vertex total columns.  
336 After this smoothing, the bias is reduced to 0.2% instead of 4.1% for the CO total columns. For the XCO product,  
337 which includes the use of the surface pressure for the MIR product and the retrieved O<sub>2</sub> column for the NIR product  
338 the bias is reduced to 0.4% instead of 3.5%.

### 339 3.2 Surface CRDS data analysis

340 The surface CO<sub>2</sub> and CO data acquired with the CRDS analysers were processed and averaged following  
 341 the procedure described in González del Castillo et al. (2022). Data were averaged and their standard deviation  
 342 calculated, per minute, then per hour. To extract the trend and seasonal CO and CO<sub>2</sub> variability, data were filtered  
 343 by discarding hours generally affected by transient and very local effects. Data recorded between 13 and 17h with  
 344 standard deviations lower than 6.0 ppm were selected for the UNA station, while nighttime data (19 to 5h) with  
 345 standard deviations lower than 2.0 ppm were selected for the ALTZ station, according to González del Castillo et  
 346 al. (2022).

### 347 3.3 Mixed Layer height from the Lidar measurements

348 The MLH is retrieved using a combined algorithm based on the gradient method and a wavelet-covariance  
 349 transformation as described in detail by García-Franco et al. (2018). These results were compared with radiosonde  
 350 data and MLH values derived from surface and vertical column densities of trace gases, and more recently Burgos-  
 351 Cuevas et al. (2022) used the variance of the vertical velocity from a Doppler Lidar (Wind Cube 100, Leosphere)  
 352 and compared with the ceilometer results at the same location. These studies show that the ceilometer retrieved  
 353 MLHs compare well with other techniques during the daytime (they agree within 15% with the trace gas method),  
 354 which are relevant for this study, whereas late afternoon and nighttime retrieved values might be affected by  
 355 aerosol residual layers at higher altitudes.

### 356 3.4 Mixed layer CO and CO<sub>2</sub> concentrations from FTIR measurements

357 Pollutant concentrations within the mixed layer are often estimated using surface measurements, although  
 358 surface concentrations are very sensitive to the air mass vertical transport, unlike the total columns. It is especially  
 359 the case within the Mexico City basin where the mixed layer has a strong diurnal dynamics controlling the vertical  
 360 distribution of the emitted pollutants (Stremme et al., 2009; Garcia-Franco et al., 2018). An estimate of the CO<sub>2</sub>  
 361 and CO vertically averaged concentrations across the mixed layer can be made using the total columns measured  
 362 at the UNA and ALTZ stations. The dry air mole fraction measured at the UNA station ( $XCO_2^{UNA}$ ) is the weighted  
 363 mean of that measured in the mixed layer ( $CO_2^{ML}$ ) and in the free troposphere at the ALTZ station ( $XCO_2^{ALTZ}$ ):  
 364

$$365 \quad XCO_2^{UNA} = w_1 \times CO_2^{ML} + w_2 \times XCO_2^{ALTZ} \quad (4)$$

$$366 \quad CO_2^{ML} = \frac{XCO_2^{UNA} - w_2 \times XCO_2^{ALTZ}}{w_1} \quad (5)$$

367  
 368 The weights ( $w_1$  and  $w_2$ ) depend on the pressure difference between the mixed-layer height (MLH) and the UNA  
 369 station, the pressure on top of the mixed layer is calculated assuming an exponential decay and an effective scale  
 370 height  $H_{scale}$  (assumed to be 8.0 km):

$$371 \quad w_1 = \left(1 - e^{-\frac{MLH}{H_{scale}}}\right) \text{ and } w_2 = \left(e^{-\frac{MLH}{H_{scale}}}\right) \quad (6)$$

373 The MLH above Mexico City was estimated using the hourly-averaged measurements of the ceilometer at the  
 374 UNA station. The hourly-averaged  $\text{CO}_2^{\text{ML}}$  and  $\text{CO}^{\text{ML}}$  products were calculated applying the same strategy for the  
 375 entire time series and are reported in Fig. 7, concurrently to the surface data.

#### 376 4 Results

377 The FTIR  $\text{XCO}_2$  and  $\text{XCO}$  daily-averaged time series and  $\text{CO}_2$  and  $\text{CO}$  surface concentrations obtained at the  
 378 UNA, VAL and ALTZ stations between November 2015 and June 2021 are shown in Fig. 2. Trends and seasonal  
 379 variabilities were fitted using a Fourier series analysis (Eq. (7) and black and red solid lines in Fig. 2), following  
 380 Wunch et al. (2013):

$$381 \quad f(x) = ax + \sum_{k=0}^n a_k \cos(2\pi kx) + b_k \sin(2\pi kx), \text{ with } n = 2 \quad (7)$$

382 where  $x$  is the time (decimal year),  $a$  the mean growth rate (ppm/year), and  $a_k$  and  $b_k$  the Fourier coefficients  
 383 modulating the annual cycles. The coefficients for each gas species and station are reported in Table 3.  
 384

385  
 386  
 387 *Table 3: Fourier series fitting parameters for the UNA, VAL and ALTZ  $\text{XCO}_2$  and  $\text{XCO}$  time series presented in Fig. 2, and  
 388 calculated from Eq.(7).*

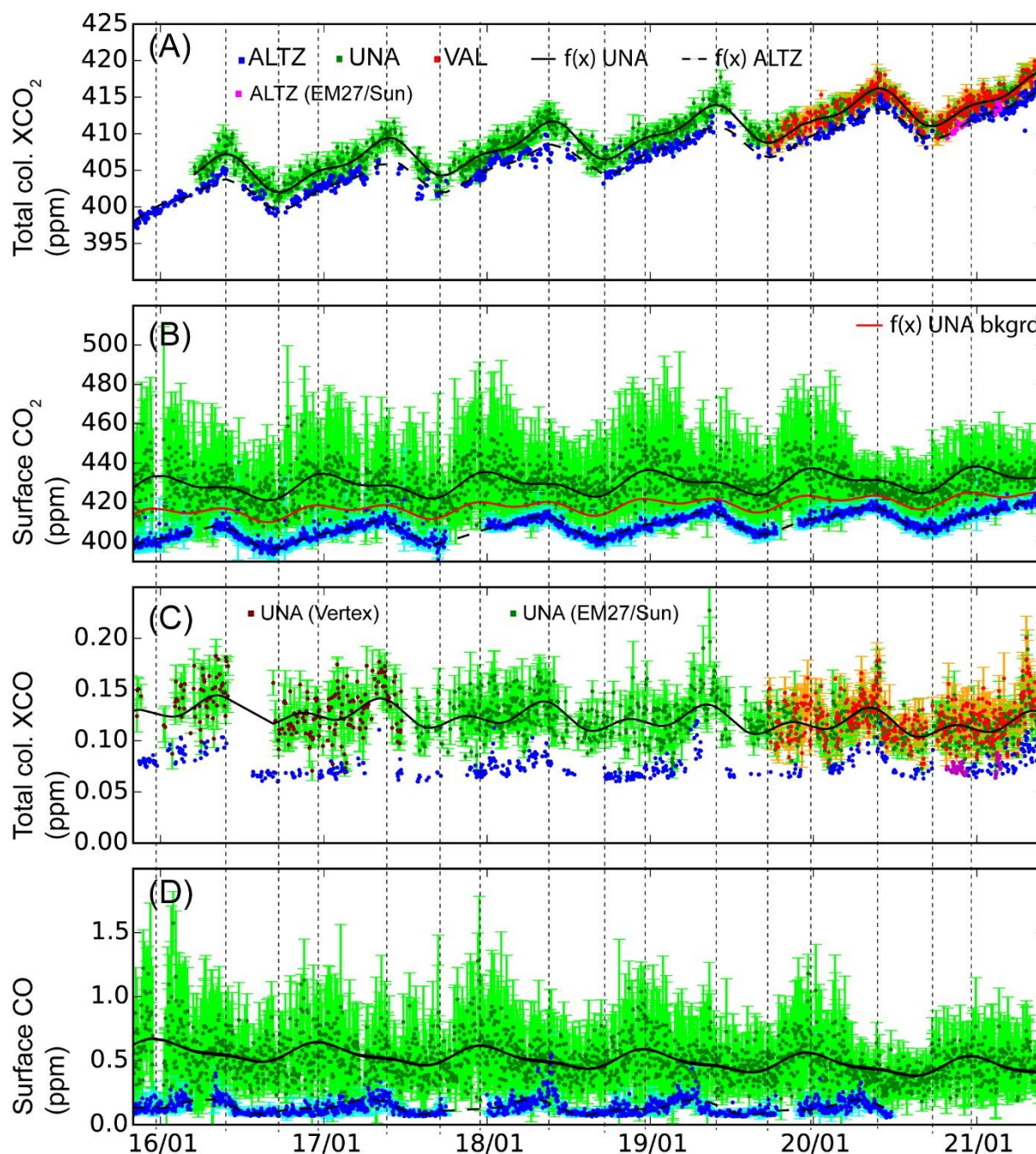
Fitting parameters (ppm/year)	$\text{XCO}_2$ , UNA Tot. Col.	$\text{XCO}_2$ ALTZ Tot. Col.	$\text{CO}_2$ UNA Surface	$\text{CO}_2$ ALTZ Surface	$\text{XCO}$ UNA Tot. Col.	$\text{CO}$ UNA Surface
$a$	$2.25 \pm 0.02$	$2.40 \pm 0.01$	$1.6 \pm 0.1$	$2.48 \pm 0.02$	$(-4.0 \pm 0.8) \times 10^{-3}$	$(-2.7 \pm 0.1) \times 10^{-2}$
$a_1$	$-1.06 \pm 0.04$	$-0.78 \pm 0.04$	$1.7 \pm 0.2$	$-0.39 \pm 0.05$	$(-2.4 \pm 0.7) \times 10^{-3}$	$(6.5 \pm 0.4) \times 10^{-2}$
$a_2$	$2.11 \pm 0.04$	$1.93 \pm 0.04$	$1.1 \pm 0.2$	$-0.36 \pm 0.05$	$(-3.2 \pm 0.8) \times 10^{-3}$	$(1.5 \pm 0.4) \times 10^{-2}$
$b_1$	$0.71 \pm 0.04$	$0.64 \pm 0.04$	$2.1 \pm 0.2$	$4.62 \pm 0.05$	$(8.6 \pm 0.8) \times 10^{-3}$	$(6.5 \pm 4.0) \times 10^{-3}$
$b_2$	$-0.78 \pm 0.04$	$-0.45 \pm 0.04$	$-2.1 \pm 0.2$	$-1.69 \pm 0.05$	$(-7.9 \pm 0.7) \times 10^{-3}$	$(-2.2 \pm 0.4) \times 10^{-2}$

#### 389 4.1 Trends and interannual variability

390 The total column  $\text{XCO}_2$  time series (Fig. 2A) at ALTZ and UNA show a similar mean growth rate around 2.4  
 391 ppm/year (2.4 and 2.3 ppm/year for ALTZ and UNA, respectively, Table 3) over the whole measurement period.  
 392 A similar mean growth rate is also found for the surface  $\text{CO}_2$  time series (Table 3 and Fig. 2 B) in ALTZ (2.5  
 393 ppm/year). These values are consistent with those estimated at the Mauna Loa Observatory (MLO) reference  
 394 station for the 2016-2021 period (average of  $2.5 \pm 0.5$  calculated from surface data available in the NOAA site  
 395 <https://gml.noaa.gov/ccgg/trends>).

396 At the UNA station a surface mean growth rate of 1.6 ppm/year is found, lower than that observed from the total  
 397 column measurements. Comparing the surface mean growth rates with those reported by González del Castillo et

398 al. (2022) for the 2014-2019 period, we observe a significant difference for the UNA station (2.3 ppm/year in  
 399 González del Castillo et al., 2022) but very similar values for the ALTZ station (2.6 ppm/year in González del  
 400 Castillo et al., 2022). The difference observed at UNA could stem from (i) starting our new time series at the end  
 401



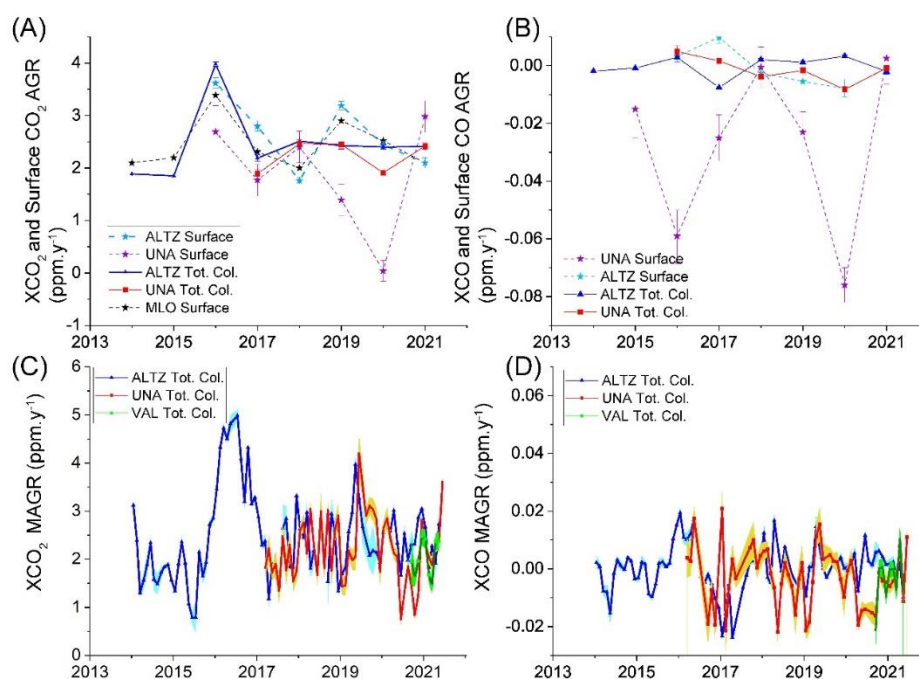
402  
 403 **Figure 2:** Time series of (A) the total column  $XCO_2$  from the FTIR measurements (B) the  $CO_2$  surface concentration  
 404 from the CRDS measurements, (C) the total column  $XCO$  from the FTIR measurements (D) the  $CO$  surface  
 405 concentration from the CRDS measurements for the UNA (in green), VAL (in red) and ALTZ (in blue) stations. For  
 406 each time series, the daily average data are presented as dots with their daily standard deviations. Black traces show  
 407 the annual fit calculated from the Fourier series (Eq. (7)). In (A) and (C), we distinguished between ALTZ data obtained  
 408 from the IFS120/5HR (in blue) and from the EM27/Sun (in magenta) and in (C), between the  $CO$  total columns obtained  
 409 from the VERTEX instrument (in brown) and the EM27/Sun (in green) at the UNA station. In (B) the red curve  
 410 corresponds to the background fit, calculated following González del Castillo et al. (2022), to determine the annual trend  
 411 and seasonal cycles. Dash lines highlight the minimum and maximum of the annual cycles for the different products.

412 of 2015, when the annual growth rate is maximum (González del Castillo et al., 2022) and (ii) the inclusion of the  
 413 2019-2021 period, when the mean growth rate clearly decreased. At the VAL station, the total column XCO<sub>2</sub> time  
 414 series are found very similar to those observed at UNA stations (Fig. 2A). Figure S1 shows that 86% of the daily  
 415 average data at VAL and UNA have a difference lower than 1.0 ppm, although a large part of the comparison was  
 416 done during the COVID19 lock-down period (Table1), for which lower gradients are expected due to the decrease  
 417 of the anthropogenic emissions.

418 The interannual variability can be explored through the time series of the mean annual growth rate (AGR) and the  
 419 monthly-sampled annual growth rate (MAGR), according to Buchwitz et al. (2018). The MAGR is calculated by  
 420 month, as the difference between the monthly-average Xgas data of a year *i* and the monthly-averaged data of the  
 421 previous year (*i-1*). The AGR is obtained for each year, averaging all of the MAGR. The AGR and MAGR for  
 422 total column and surface measurements are presented in Fig. 3. We include data from the MLO in Fig. 3A, for  
 423 which the AGR (dashed black curve) was derived from the surface data available in the NOAA site.

424 At ALTZ, the interannual variability of the total column XCO<sub>2</sub> AGR (Fig. 3A) was found similar to that obtained  
 425 from both the ALTZ and MLO surface data, with a coincident peak in 2016, reaching an AGR value of 3.5 (surface  
 426 data) and 4.0 (total column data) ppm/year. Surface data AGR time series show a second peak in 2019, which is  
 427 not apparent for the total column XCO<sub>2</sub> time series. The time series of the MAGR (Fig. 3C) allows better  
 428 identifying and characterising the period and duration of the anomalies. The 2016 XCO<sub>2</sub> anomaly has a duration  
 429 up to 15 months (from October 2015 to March 2017), reaching a maximum value (around 5.0 ppm/year) between  
 430 March and July 2016, corresponding to a factor of 2.8 higher than the 2013-2015 base level (1.8 ppm/year).

431

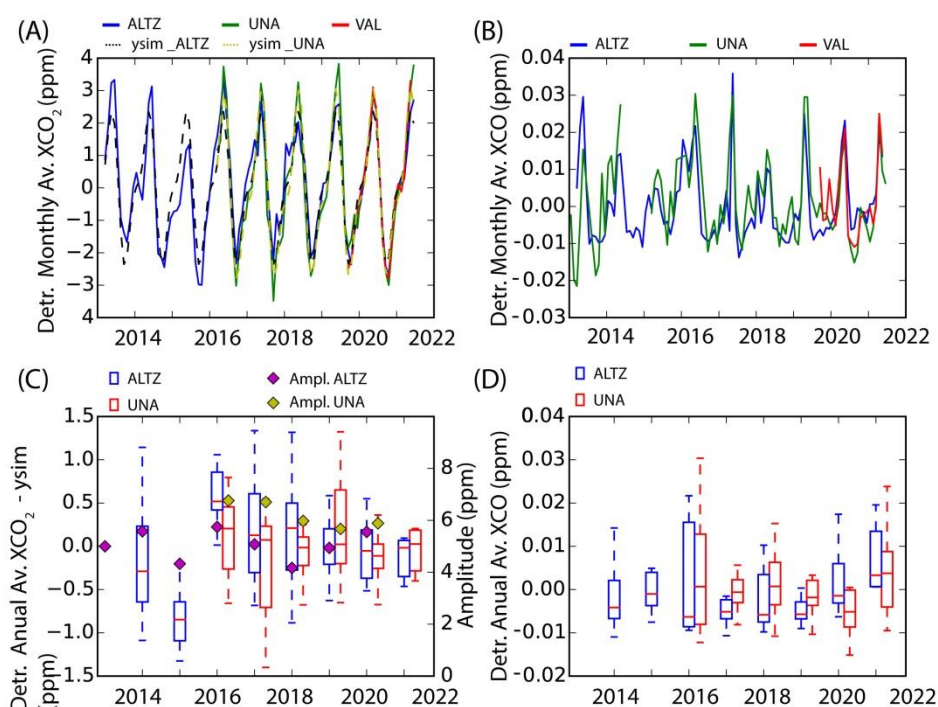


432

433 **Figure 3:** XCO<sub>2</sub> (A) and XCO (B) annual growth rates (AGR) and XCO<sub>2</sub> (C) and XCO (D) monthly-sampled annual  
 434 growth rate (MAGR) obtained from total column and surface measurements for UNA, VAL, and ALTZ stations. In  
 435 (A), the Mauna Loa (MLO) AGR trend was added in black dash-line. In (A) and (B) errors bars represent the standard  
 436 error after removing annual cycles, reflecting the data sample quality. The standard error for the MAGR is shown as  
 437 shaded area in (C) and (D).

438 At UNA, the XCO<sub>2</sub> AGR and MAGR time series (Fig. 3A and 3C) are very similar to those observed at the ALTZ  
 439 station, except for the year 2020. During this year, the AGR dropped by ~20% at UNA before returning in 2021  
 440 to the level of the previous two years. This behaviour contrasts with the AGR observed at ALTZ, which remains  
 441 nearly constant between 2017 and 2021. The MAGR time series at UNA (Fig. 3C) shows that this drop is  
 442 dominated by the exceptionally low June and October growth rates, representing the lowest MAGR values of the  
 443 UNA time series. This observation is supported by the VAL MAGR, although the time series is much shorter. The  
 444 surface CO<sub>2</sub> AGR at UNA shows a much higher interannual variability, with the strongest anomaly observed in  
 445 2020, where the AGR is close to zero. A very clear decrease of the day-to-day and intraday CO<sub>2</sub> surface variability  
 446 is observed in Fig. 2B from April to mid-September 2020, consistent with the XCO<sub>2</sub> MAGR anomaly.

447



448

449 **Figure 4: Interannual and annual variability of the detrended XCO<sub>2</sub> and XCO total column data at the UNA, VAL and**  
 450 **ALTZ stations. In (C) and (D) the whisker diagrams are calculated from the monthly average detrended data. The**  
 451 **amplitude is determined as the max-min values.**

452 Upon examining CO, the UNA XCO time series (Fig. 2C) has daily averages ranging between 0.10 and  
 453 0.23 ppm with a mean and standard deviation of 0.12 and 0.02 ppm, respectively, but shows a decreasing rate ( $-$   
 454  $4.0 \times 10^{-3}$  ppm/year) over the whole measurement period. The VAL XCO time series show a very similar baseline  
 455 to UNA, with a daily average difference lower than 0.02 ppm for 85% of the coincident dataset (Fig. S1). At the  
 456 ALTZ background site, the XCO baseline and day-to-day variability are lower than at UNA and VAL, as expected  
 457 (mean and standard deviation equal to 0.08 and 0.01 ppm, respectively). The surface CO time series (Fig. 2D)  
 458 shows a more significant decreasing trend ( $-2.68 \times 10^{-2}$  ppm/year) than the total column data at UNA, while the  
 459 baseline at ALTZ remains constant around 0.11 ppm. The CO AGR and MAGR at ALTZ and UNA are shown in  
 460 Fig. 3B and D. Generally, the XCO AGR and MAGR oscillate around their base level at the ALTZ and UNA  
 461 stations, with short-term anomalies. At ALTZ, a strong negative XCO AGR anomaly is observed in 2017, which  
 462 was not observed for XCO<sub>2</sub>, likely resulting from the exceptionally high XCO columns measured during 2016.  
 463 This is supported by the increase of the XCO MAGR from October 2015 to July 2016 (Fig. 3D), coinciding with

464 the first 10 months of the highest XCO<sub>2</sub> anomaly and followed by the lowest XCOMAGR values of the time series  
 465 (around -0.02 ppm/year in April 2017). At the UNA station, the AGR slightly decreases between 2016 and 2020  
 466 and increases again in 2021. The most significant and prolonged (>5 months) MAGR anomaly (Fig. 3D) occurred  
 467 between April and September 2020, with negative values. Some short-term additional anomalies are observed, but  
 468 only a few of them (in May 2018 and January 2019) are not affected by the limited number of available  
 469 measurements.

#### 470 **4.2 Seasonal variability and short-term cyclic events**

471 Annual cycles are observed for both total column XCO<sub>2</sub> and CO<sub>2</sub> surface measurements at ALTZ, UNA and VAL  
 472 stations (Fig. 2). The maximum and minimum of the total column XCO<sub>2</sub> cycles are observed in May-June and  
 473 September, respectively, with an average amplitude around 5 (ALTZ) and 6 (UNA) ppm.

474 To examine the temporal changes in amplitude and shape of the annual cycles, total column data were monthly-  
 475 averaged, detrended by subtracting the linear part of the fit ( $f(x) = ax$ , in Eq. (7)), and compared to the detrended  
 476 mean annual cycle ( $f(x) - ax$ ) in Fig. 4. To obtain a longer-term view, we included the 2013-2015 period from the  
 477 ALTZ station, previously published in Baylon et al. (2017), after applying the inter-calibration factors (section  
 478 3.1.3). At ALTZ, two periods significantly deviated from the average XCO<sub>2</sub> seasonal cycle, i.e.: (i) the year 2015,  
 479 where all the monthly averaged XCO<sub>2</sub> are below the fit and with one of the lowest seasonal amplitudes (~4.0 ppm,  
 480 Fig. 4A and 4C) of the whole time series, and (ii) the year 2016, with higher monthly averages than the mean  
 481 XCO<sub>2</sub> seasonal cycle and the highest amplitude (~5.8 ppm, Fig. 4A and 4C). At UNA, the difference with respect  
 482 to the average XCO<sub>2</sub> seasonal cycle is not significant, except for the year 2020, where all the monthly averages  
 483 are below the mean annual cycle (Figure 4C). During this period, the UNA and VAL XCO<sub>2</sub> monthly-averaged  
 484 data fit exceptionally well with those of the ALTZ station between March 2020 and March 2021 in terms of shape  
 485 and amplitude, while the UNA and VAL annual cycle amplitudes are slightly higher than those of ALTZ for the  
 486 other years.

487 Regarding the CO<sub>2</sub> surface data (Fig. 2B), annual cycles are observed with maxima and minima reached mid-  
 488 December and mid-September, respectively. As also reported in González del Castillo et al. (2022), the maximum  
 489 occurred during winter, when shallower boundary layer prevails and the summer-autumn minimum can be  
 490 explained by the dilution of trace gases in a deeper convective boundary layer and more active urban vegetation.  
 491 XCO peaks every year in April-May at the three stations (Fig. 2C and Fig. 4B) and then shows minimal annual  
 492 values in August, preceding by 1 month the minimum and maximum values of the XCO<sub>2</sub> time series. The April-  
 493 May maximal annual values, also confirmed by TROPOMI measurements (Borsdorff et al., 2020), coincide with  
 494 the biomass burning season and the periods during which the mixed layer reaches its maximum altitude (García-  
 495 Franco et al., 2018). During 2015, the XCO time series show a very low maximum reached in February instead of  
 496 May (Fig. 4B), contrasting with 2016, where high total column XCO values are reached in January and maintained  
 497 for a period of at least 5 months. 2016 also corresponds to the year with the highest XCO variability of the time  
 498 series (Figure 4D). Additionally, in 2018, the XCO annual cycles differ from the other years with lower values  
 499 and a flat shape during the first semester of the year (January-May).

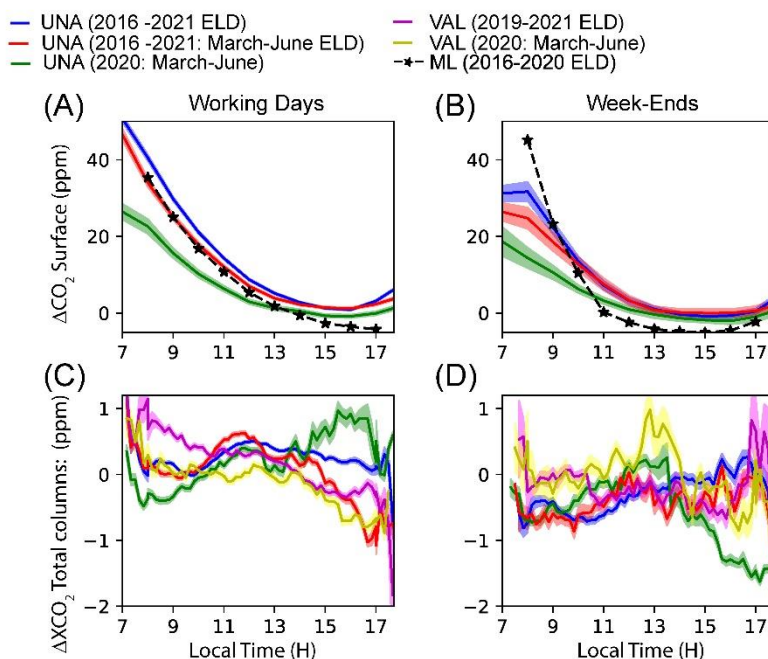
500 Surface CO data (Fig. 2D) also show periodic increases at the ALTZ station with maxima reached during April-  
 501 May, coinciding with the maxima observed from total column XCO measurements. They confirm the increase of  
 502 the CO emissions during the biomass burning season, at least dominant in the ALTZ measurements. However, at



503 the UNA station, cycles are also observed in the surface data but with a maximum coinciding with that of the CO<sub>2</sub>  
 504 surface data, and lagging behind the XCO<sub>2</sub> total columns. These cycles are likely dominated by other processes  
 505 affecting both CO and CO<sub>2</sub> species such as the mixed layer seasonal dynamic.

### 506 4.3 Intraday variability

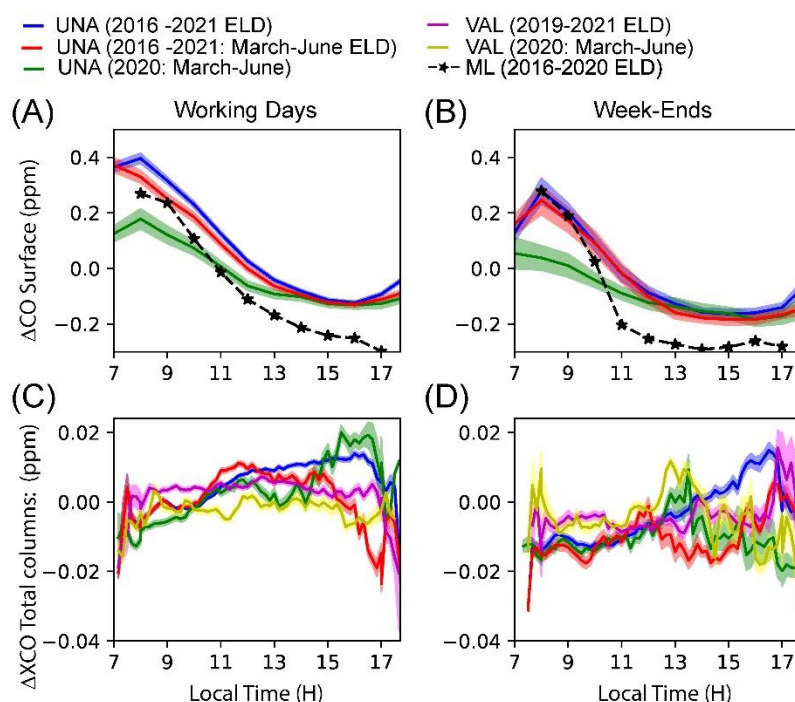
507 The intraday variability of the total columns and surface data are depicted in Fig. 5 and Fig. 6. Since the ALTZ  
 508 total column data do not present a significant diurnal pattern (the hourly variability remains lower than the standard  
 509 error of the time series), they are not presented in these plots.



510  
 511 **Figure 5: Diurnal patterns of the detrended surface CO<sub>2</sub> mole fractions (A and B) and XCO<sub>2</sub> total columns (C and D)**  
 512 **measured at UNA and VAL stations. For each panel, the different curves represent different time periods: in blue, the**  
 513 **whole measurement period excluding the lock-down period (March-June 2020 ELD), in green the lock-down period**  
 514 **(March-June 2020) and in red the whole measurement periods from March to June, excluding the lock-down. The**  
 515 **standard errors are presented as shaded areas. Black curves represent the diurnal pattern of CO<sub>2</sub> in the Mixed Layer**  
 516 **(ML) calculated from the total columns data for the UNA station.**

517 Total column data were detrended by removing the seasonal fit (black traces in Fig. 2A and Fig. 2C), and averaged  
 518 over 10 min. To avoid a possible bias due to strong ventilation periods, a filter based on a ventilation index (VI)  
 519 was applied, following recommendations in Hardy (2001), Su et al. (2018) and Storey and Price (2022). The VI is  
 520 calculated as the product of average wind speed velocity (between the surface and 100 m height), and the planetary  
 521 boundary layer height for UNA and VAL locations. The wind velocity and the MLH were estimated with the U  
 522 and V wind components and the PBL height fields from the hourly ERA5 reanalysis product (Hersbach et al.,  
 523 2020). In the MCMA, the surface wind speed presents a diurnal pattern, generally reaching a maximum during the  
 524 afternoon between 14 and 15h LT (Fig. S4). The filter selects the days complying with the following criteria (i) a  
 525 maximum wind velocity (average 10-100m height) between 10h and 12h LT lower than 1.5 m.s<sup>-1</sup> (threshold based  
 526 on Stremme et al., 2013) and (ii) a daily VI lower than 2350 m<sup>2</sup>.s<sup>-1</sup>, which represents a commonly used threshold  
 527 for selecting poor ventilation conditions (Hardy, 2001; Storey and Price, 2022). About 60% of the original XCO<sub>2</sub>  
 528 and XCO dataset is selected by applying the filter, and will be considered in the following analysis. We note that  
 529 about 70% of discarded data corresponds to the January-May period of the year. Filtered total column XCO<sub>2</sub> and

530 XCO data were averaged by 10 min and presented in Fig. 5C-D and Fig. 6C-D, distinguishing between the working  
 531 days (WD) and the week-end (WE) periods. To explore the 2020 lock-down influence on the diurnal pattern, three  
 532 different periods were distinguished for each plot, the first one (blue trace: 2016 - 2021) corresponding to the  
 533 whole measurement period excluding the interval between March and June 2020 corresponding to the lock-down  
 534 period (hereafter, called “ELD” for “excluding the lock down period”), where a significant MAGR decrease was  
 535 observed; the second (green trace: March- June 2020) only includes the lock-down period, and additionally  
 536 excludes the rainy season to avoid bias due to incomplete daily time series; and the third period (red trace) is the  
 537 same as the first one, but only considering the March to June months to be compared with the lock-down period.



538  
 539 **Figure 6: Same as Figure 5, but from surface CO and total column XCO measurements.**

540 Surface data from the CRDS analyzers were detrended by removing the background fit following the methodology  
 541 described in the section 3.2, and filtered to be coincident with the filtered total column measurements (selection of  
 542 data between 7 and 18 h LT and only including the days with low ventilation conditions). They were finally  
 543 averaged by hours and presented in Fig. 5A-B and Fig. 6A-B for the WD and WE periods, respectively, for which  
 544 each curve represents the periods mentioned above.

545  
 546 The surface CO<sub>2</sub> diurnal pattern at UNA station for the whole measurement period (2016-2021, Fig. 5A and 5B in  
 547 blue) is consistent with the one previously described in Gonzalez del Castillo et al. (2022) for the 2014-2019  
 548 period, with a maximum observed during the early morning (reached before 7h LT), a minimum during the  
 549 afternoon (between 15 and 16h LT) and an average amplitude around 45 ppm. A lower amplitude of these cycles  
 550 is observed at WE (average amplitude of 28 ppm) with respect to the WD periods. During the 2020 lock-down  
 551 period (green curve), the WD surface CO<sub>2</sub> diurnal profile has a comparable amplitude (average amplitude of 26  
 552 ppm) to those of the WE for the whole measurement period, and slightly higher than that observed during the lock-  
 553 down WE periods (average amplitude of 22 ppm). The surface CO diurnal profile (Fig. 6: 2016-2021, blue curve)  
 554 peaks at 8h and then decreases until 16 h LT during any day of the week. The WD and WE data shows amplitudes

555 of up to 0.5 ppm and 0.3 ppm, respectively. During the lock-down period the WD and WE amplitudes are much  
556 lower (0.3 and 0.2, respectively), consistently with the CO<sub>2</sub> surface observations.

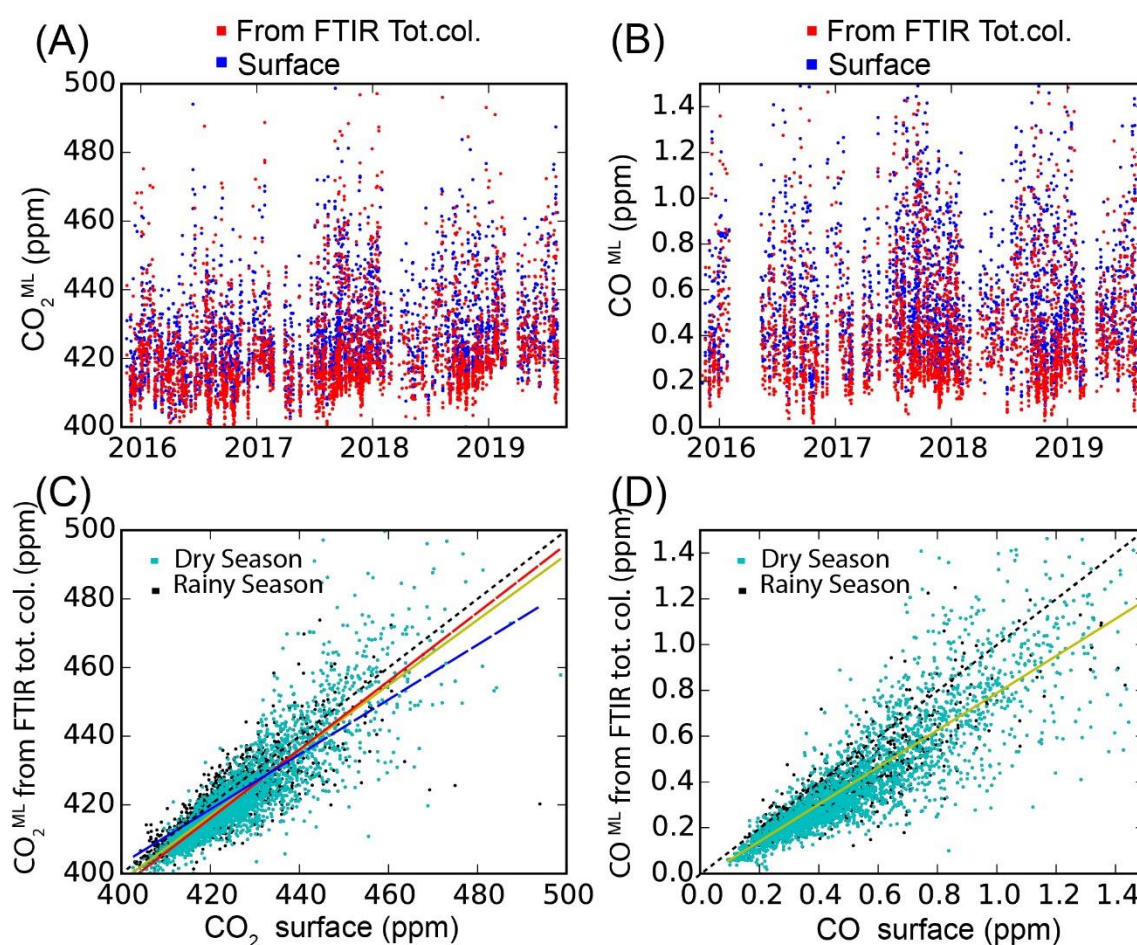
557 The XCO<sub>2</sub> and XCO diurnal patterns (Fig. 5C-D and Fig. 6C-D) have very different shapes than those of the surface  
558 data, with amplitudes one order of magnitude lower. The variability observed between 7 and 8h is likely due to  
559 the low number of measurements during this time interval, and will not be taken into account in the following  
560 analysis. The UNA and VAL XCO<sub>2</sub> diurnal patterns significantly differ in shape. The VAL WD curve (magenta)  
561 continuously decreases from 8h to 17h (amplitude around 2 ppm) during both the whole measurement and lock-  
562 down periods, but during the lockdown period, lower values are generally recorded with higher intra-hour  
563 variability between 11h and 14h. The general WD decreasing trend suggests a maximum reached during the early  
564 morning (before 7h LT). This observation is supported by the CO<sub>2</sub> surface measurements performed with the low-  
565 cost medium precision CO<sub>2</sub> sensors (Porrás et al., 2023), recording a maximum between 6h and 7h LT. The UNA  
566 XCO<sub>2</sub> WD diurnal pattern (blue trace) is almost constant until 10h, then increases until reaching a maximum around  
567 12h, slightly decreases until 17h LT and finally shows an abrupt decrease after that. The amplitude of the diurnal  
568 variability is around 1 ppm. During the lock-down period, the diurnal profile is different, increasing until 12h LT,  
569 slightly decreasing until 13h LT and then increasing again until reaching a maximum at 16h, and finally abruptly  
570 decreasing until 17h LT. The lock-down WD XCO<sub>2</sub> profile shows lower values than the other periods until 13h,  
571 but the peak observed at 16h is not apparent for the other periods. Variability is generally lower at WE (<1ppm),  
572 except for the lock-down period, for which an important decrease is observed after 14h LT, but it is likely affected  
573 by a low number of measurement days. For XCO, the diurnal profiles also have different shapes at UNA and VAL.  
574 At UNA, the March-June XCO diurnal profiles (red and green curves) resemble that of XCO<sub>2</sub> for both the lock  
575 down and whole measurement periods. When considering the twelve months of the year (blue trace), the maximum  
576 curve slightly increases between 12h and 16h, when it reaches its maximum. It contrasts with the variability of the  
577 March to June months curves during this time interval, for which an increase is observed during the lock-down  
578 period or a decrease if considering the whole measurement period. At VAL, the diurnal profile is fairly constant  
579 until 17h with slightly lower values during the lock-down period.

580 The total column XCO diurnal profiles at WE are less reliable with larger standard errors, likely due to the low  
581 number of considered measurements. An increase is nevertheless observed at UNA where the considered day's  
582 number is statistically more reliable, with a peak around 17h LT, which was not observed for XCO<sub>2</sub>.

583 The difference observed between the diurnal pattern of the XCO and XCO<sub>2</sub> at VAL and UNA is likely due to the  
584 different advection drivers in the region mainly controlled by the topography. A Northern surface wind direction  
585 (Figure S6) is generally dominating over the Mexican valley but is locally highly influenced by the mountainous  
586 barriers. The West-northwest wind component at UNA is likely to be the effect of down-slope flows from the  
587 mountain ridge in the early morning (6 – 9 LT mostly), while at VAL, the plateau-to-basin winds are the main  
588 influx into the basin coming from the northwest in the morning. There can also be an influence from an up-valley  
589 flow in the mornings (de Foy et al., 2006). More generally the VAL station is likely influenced by the north  
590 mountain, generating a significant gradient in the CO distribution upwind of the VAL station (Figure 1). In  
591 contrast, near the UNA station, the flat ground allows a more efficient mixing and due to the dominant North-  
592 Northeast wind component in the late morning, the captured airmasses likely often reflects the MCMA plume  
593 emissions.

#### 594 4.4 CO and CO<sub>2</sub> within the mixed layer from FTIR and surface data.

595 Figure 7 shows the hourly-averaged CO<sub>2</sub> and CO concentration within the mixed layer (CO<sub>2</sub><sup>ML</sup> and CO<sup>ML</sup>  
 596 products), calculated from the FTIR measurements (see section 3.4), concurrently to the surface data. The CO<sub>2</sub><sup>ML</sup>  
 597 and CO<sup>ML</sup> products are in agreement with the surface observation, with a slope of  $0.95 \pm 0.02$  ( $R^2=0.74$ ) for CO<sub>2</sub>  
 598 (Fig. 7C) and  $0.81 \pm 0.02$  ( $R^2=0.74$ ) for CO (Fig. 7D). For CO<sub>2</sub>, the slope was found closer to 1.0 ( $1.00 \pm 0.02$ ) with  
 599 an offset of  $-2.9 \pm 0.2$  and a better  $R^2$  (0.77) when discarding the data corresponding to the rainy season. This effect  
 600 is likely due to the removal of the incomplete daily time series frequently interrupted at the beginning of the  
 601 afternoon during the rainy season.  
 602



603  
 604 **Figure 7: Comparison between (A) the CO<sub>2</sub><sup>ML</sup> and (B) CO<sup>ML</sup> products (red) derived from the ALTZ and UNA total**  
 605 **column measurements and the surface measurements at the UNA station (blue). (C) and (D) represent the correlation**  
 606 **plots for CO<sub>2</sub> and CO, respectively. In (C) and (D), we distinguished between data corresponding to the dry (November**  
 607 **to May: cyan) and rainy (June to October: black) seasons. In (C), yellow, red and blue linear regression curves**  
 608 **correspond to the whole measurement period (yellow: slope= $0.95 \pm 0.02$ ; Offset=  $17.9 \pm 0.2$ ;  $R^2=0.74$ ), the dry season (red:**  
 609 **slope= $1.00 \pm 0.02$ ; Offset:  $-2.9 \pm 0.2$ ;  $R^2=0.77$ ) and the rainy season (blue: slope= $0.80 \pm 0.03$ ; Offset:  $83.7 \pm 0.39$ ;  $R^2=0.66$ ). In**  
 610 **(D), since no significant difference was found for the different period, the regression line (yellow: slope= $0.81 \pm 0.02$ ;**  
 611 **offset:  $-0.021 \pm 0.004$ ;  $R^2=0.74$ ) represent the whole measurement. The black dash line represents  $y=x$ .**

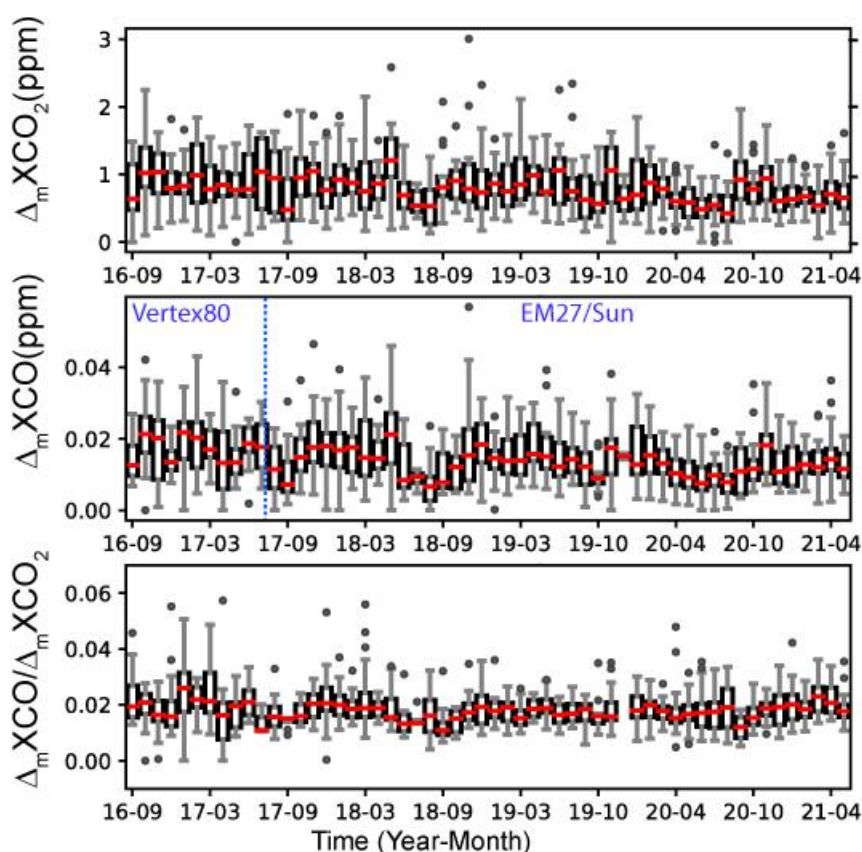
612  
 613 The CO<sub>2</sub><sup>ML</sup> and CO<sup>ML</sup> diurnal patterns are presented in Fig. 5 and Fig. 6 (dash lines) together with those of surface  
 614 measurements, after a similar filtering. The CO<sub>2</sub><sup>ML</sup> and surface CO<sub>2</sub> diurnal patterns (Fig. 5A and 5B) are very  
 615 similar in shape and amplitude, especially during the WD, although a small difference is observed at the end of

616 the afternoon (<5 ppm). This difference is likely due to the increase of the uncertainties of the MLH estimate when  
 617 it is more diluted. The  $\text{CO}^{\text{ML}}$  and surface CO diurnal profiles (Fig. 6A and 6B) also have similar amplitudes and  
 618 shape for both WD and WE, although the  $\text{CO}^{\text{ML}}$  diurnal profile shows lower values (offset around 0.1 ppm at WD).  
 619 Despite this very simplified model, these results show that the total column and surface measurements are mutually  
 620 very consistent when the seasonal and diurnal variability of the ML expansion above Mexico City is taken into  
 621 account.

#### 622 4.5 $\text{XCO}_2$ to $\text{XCO}$ enhancements ratios

623 The  $\text{XCO}$  and  $\text{XCO}_2$  correlated enhancements and their ratio can give insights into the combustion efficiency of  
 624 the sources in a city, and therefore on their contributions. In this study we explored the variability of the  
 625  $\text{XCO}/\text{XCO}_2$  ratios at both long-term and intraday scales.

626 For the long-term analysis, the  $\text{XCO}_2$  “background” level was calculated using a statistical method, using the lower  
 627 5<sup>th</sup> percentile of the measured  $\text{Xgas}$  over a 1-day running window (You et al., 2021). We did not use the ALTZ  
 628 measurements because of (i) the periodic influence of the wildfires in the region during the dry season, and (ii) the  
 629 discontinuity of our daily averaged time series. The enhancements above background  $\Delta_m\text{XCO}_2$  and  $\Delta_m\text{XCO}$   
 630 measured at UNA and averaged by months and their ratios are presented in Fig. 8, as whisker diagrams.



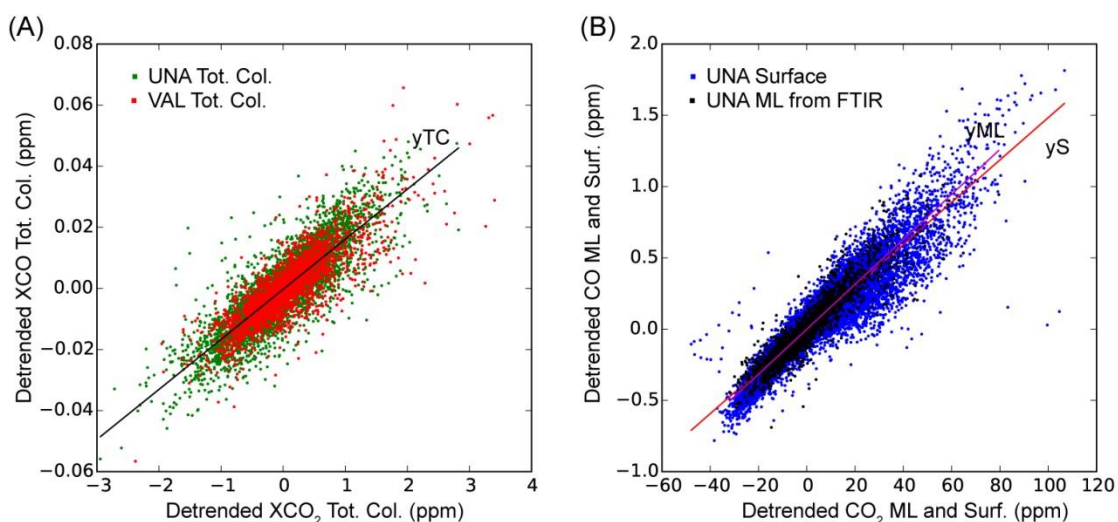
631  
 632 **Figure 8: Whisker diagram representing by month the variability of  $\Delta\text{XCO}_2$ ,  $\Delta\text{XCO}$  and their ratio from the UNA**  
 633 **measurements.**

634 Both  $\Delta_m\text{XCO}_2$  and  $\Delta_m\text{XCO}$  time series show a slight decrease over time (around 0.05 ppm/year and 0.001  
 635 ppm/year, respectively). Although the  $\Delta_m\text{XCO}/\Delta_m\text{XCO}_2$  ratio displays a variability around its mean value  
 636 ( $0.018 \pm 0.003$ ), there are no discernible cyclic or long-term trend in the time series, except for the rainy periods of  
 637 2017, 2018 and 2020 when low ratios (and low  $\Delta_m\text{XCO}$  and  $\Delta_m\text{XCO}_2$  values) were observed. The  $\Delta_m\text{XCO}$  and

638  $\Delta_m XCO/\Delta_m XCO_2$  ratio show a higher variability at the beginning of the time series (until July 2017) likely due to  
 639 the use of the CO Vertex products. The long term  $\Delta_m XCO$  decrease, also observed in other studies (Garcia-Franco,  
 640 et al., 2019; Molina, 2021, Hernández-Paniagua et al., 2021) likely reflect the effect of the successive air quality  
 641 management programs implemented in the CDMX since the 1990s to improve the air quality, including  
 642 technological advancements and fuel quality enhancements as well as refinery closures, industrial relocation, or  
 643 fuel substitution. Regarding the low seasonal variability observed for the CO/CO<sub>2</sub> ratios, it is likely related to mass  
 644 burning episodes and high-pressure weather conditions that occur during the dry season.

645 To perform the intraday analysis, the hourly-averaged data were first detrended by subtracting the daily average.  
 646 The resulting  $\Delta XCO_2$  vs.  $\Delta XCO$  datasets are plotted in Fig. 9A. The entire  $\Delta XCO_2$  and  $\Delta XCO$  datasets showed a  
 647 good correlation at both the UNA and VAL stations, with similar linear regression slopes around  $0.0164 \pm 0.0003$ ,  
 648 which is consistent with that found from the surface measurements and the ML product (Fig. 9B). Although there  
 649 is an actual difference in the emission types of the southern and northern parts of the city, the North hosting  
 650 industrial and commercial sources and the South being largely residential and commercial, the common and  
 651 dominant source of CO in the MCMA (at UNA and VAL stations) could incriminate motorised vehicles. The data  
 652 dispersion around the regression line likely reflects more punctual and local influence of other sources with an  
 653 important week-to-week variability.

654



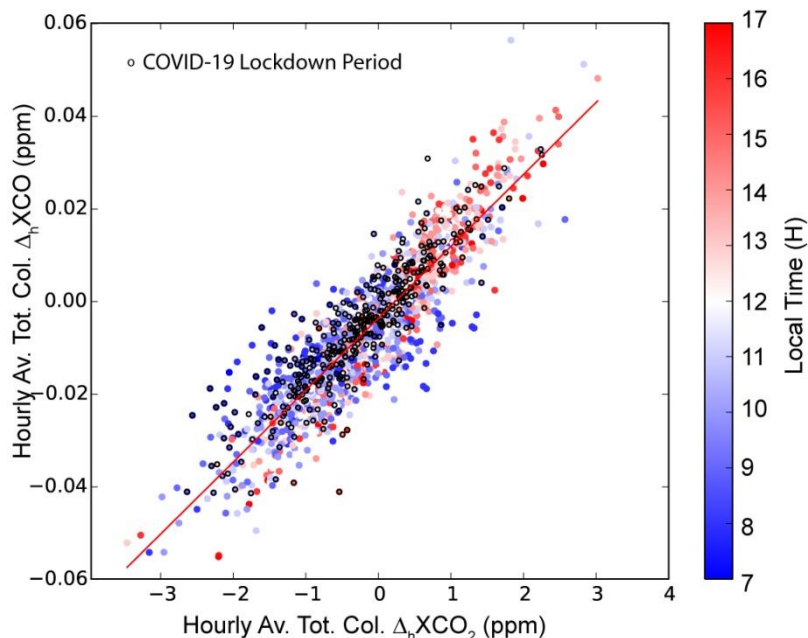
655

656

657 **Figure 9: A: Correlation plot of (A) the detrended (by removing the daily averages) hourly-average total column XCO<sub>2</sub>**  
 658 **vs. XCO data, and (B) the detrended hourly average Mixing Layer (ML) and surface CO<sub>2</sub> vs. CO products. Solid lines**  
 659 **represent the linear regression lines, with the following parameters: TC slope= $0.0164 \pm 0.0003$ ,  $R^2=0.72$  for the total**  
 660 **columns at UNA and VAL; yS slope= $0.0148 \pm 0.0001$ ,  $R^2=0.87$  for the surface products and yML slope= $0.0158 \pm 0.0002$ ,**  
 661  **$R^2=0.88$  for the Mixing Layer products.**

662 On the other hand, the total column (UNA-VAL) differences, presented in Fig. S3 can also be used to calculate  
 663 the  $\Delta XCO/\Delta XCO_2$  ratio, with a more precise subtraction of a common background (which assumes a  
 664 homogeneous background across the entire city) from the two stations. Figure 10 shows the hourly-average  $\Delta XCO_2$   
 665 (UNA-VAL) vs.  $\Delta XCO$  (UNA-VAL) correlation plot for the coincident measurement period. A well-defined  
 666 linear correlation is observed with a slope of  $0.015 \pm 0.001$  and a coefficient of determination of  $R^2=0.80$ , highly  
 667 consistent with that found in Fig. 9. The use of the (UNA-VAL) total columns difference notably improved the  
 668 coefficient of determination, by removing the regional long-term and short-term perturbations affecting the two

669 sites. The intraday variability of the  $\Delta XCO$  (UNA-VAL)/ $\Delta XCO_2$  (UNA-VAL) ratio (Fig. 10: colour scale),  
 670 showing higher columns at VAL during the morning and at UNA during the afternoon likely reflect the North to  
 671 South transport of air across the city. We note that the ratio remains the same during the lock-down period. We  
 672 would expect lower intraday (UNA-VAL)  $\Delta XCO$  and  $\Delta XCO_2$  amplitudes during the lock-down period, but it is  
 673 not clearly apparent in this correlation plot.



674  
 675 **Figure 10: Correlation plot of the  $\Delta XCO$  (UNA - VAL) vs.  $\Delta XCO_2$  (UNA - VAL) hourly averages (colour scale**  
 676 **depending on the time is shown to the right) for the coincident measurement period (September 2019 - June 2021). Dots**  
 677 **with black edges highlight the measurements during the COVID19 lock-down period (March-June 2020). Regression**  
 678 **line (in red): Slope:  $0.015 \pm 0.001$ ,  $R^2=0.80$ .**

#### 679 4.6 Estimate of CO and CO<sub>2</sub> MCMA emissions.

680 The variability of the long-term CO emissions in the MCMA can be estimated, following the method detailed in  
 681 Stremme et al. (2013). In that study, they assumed that, since the XCO emissions in the MCMA are mainly due to  
 682 traffic pollution, the rapid changes observed in the XCO total column (less affected by the air mass vertical  
 683 distribution) should reflect the CO fresh emissions under certain meteorological conditions. Low ventilation,  
 684 strong turbulence in the mixed layer and limited zenithal angle of measurements are critical criteria to avoid  
 685 enhancement due to horizontal transport or local heterogeneity. XCO growth rates can be estimated at specific  
 686 time intervals complying with these conditions from long-term time series. Further details on the method and  
 687 estimates of uncertainties due to these assumptions are given in Stremme et al. (2013). Here, we determined an  
 688 optimised time interval for estimating the mean CO growth rate using (i) the diurnal surface wind speed patterns  
 689 and (ii) the MLH growth rate, the latter reflecting the turbulence within the mixed layer (Fig. S4). The time interval  
 690 complying with a rapid growth of the mixed layer and low surface wind speed ( $< 2 \text{ m}\cdot\text{s}^{-1}$ ) was found between 10  
 691 and 12h, which is in agreement with the requirements mentioned in Stremme et al. (2013). Growth rates and their  
 692 uncertainties were determined by year, based on the linear regression (with 95% confidence interval) of the 10-  
 693 min averaged detrended CO total columns over the 10-12h interval. For example, for the year 2018, we found a  
 694 CO growth rate of  $52 \pm 5 \text{ kg}\cdot\text{km}^{-2}\cdot\text{h}^{-1}$ .

695 To extrapolate the growth rate over the MCMA, we used the TROPOMI CO total column data that we averaged  
 696 over the 2018-2022 period (Fig. 1), following the same method as described in Stremme et al. (2013). We assume  
 697 that the total amount of fresh CO is proportional to the total emission of the MCMA and to the total column  
 698 enhancement at the UNA site, which reflects the CO accumulated at this site. The ratio of the total accumulated  
 699 CO in the MCMA to the accumulated CO at UNA is therefore the same as the emission ratio of the whole Megacity  
 700 to the emission flux at UNA. Therefore this ratio is the extrapolation factor and represents an effective area, defined  
 701 as Eq. (8):

$$702 \quad Eff\_Area = \frac{\int(CO_{MCMA} - CO_{bgrd})}{CO_{UNA} - CO_{bgrd}} \quad (8)$$

703 In Eq. (8),  $(CO_{MCMA} - CO_{bgrd})$  is integrated over the area where the CO TROPOMI total columns are higher than  
 704 a predefined background value. As the TROPOMI overflight time is around 13h30 LT, we cannot neglect the  
 705 ventilation and slight advection is smoothing out the distribution, so that both the background and the column at  
 706 UNA have to be chosen carefully. The background column was therefore estimated in two ways (i) from the  
 707 smallest value observed upwind of the city (cross symbol in Fig. 1) at the elevation of the Mexican basin (contour  
 708 line separating Mexico City from the Toluca area in the west in Fig. 1) and found to be  $1.45 \times 10^{18}$  molec.cm<sup>-2</sup> and  
 709 (ii) from the Tecamac site, where the border of MCMA was assumed in Stremme et al. (2013) and where the  
 710 column was found to be  $1.60 \times 10^{18}$  molec.cm<sup>-2</sup>.

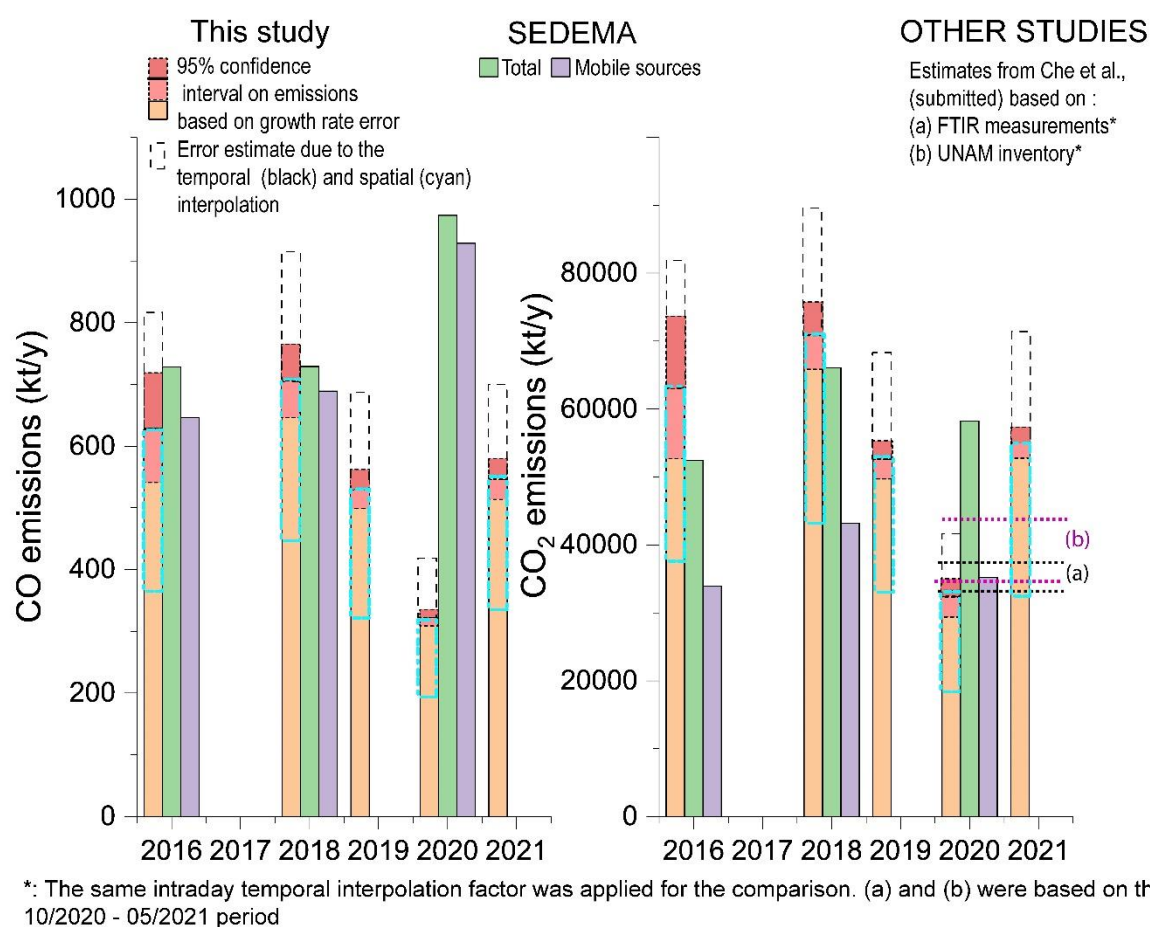
711 Due to advection, even locations slightly out of the megacity are presenting enhanced CO columns and it is not  
 712 clear which is the background column in the Mexican basin. Figure S5 illustrated the sensitivity of the effective  
 713 area to the background uncertainties. A 10% higher background leads to a 40% smaller extrapolation factor and a  
 714 40% emission underestimate. The fresh CO was estimated from the TROPOMI data by removing the background  
 715 ( $1.45 \times 10^{18}$  molec.cm<sup>-2</sup>) to the average total columns found at UNA ( $1.93 \times 10^{18}$  molec.cm<sup>-2</sup>) and was found to be  
 716  $4.79 \times 10^{17}$  molec.cm<sup>-2</sup>. In cases where the CO total column is lower than the background, likely due to the  
 717 topography effect, we set the difference column to zero for the integration. This topographic effect is important  
 718 for the considered area, as there are plenty of mountains around the basin, like the mountain ridge in the west  
 719 (including Ajusco, Desierto de Leones, etc.), some mountains in the mountain ridge on the eastern part of the area  
 720 including in the south the two volcanoes Popocatepetl and Iztaccihuatl.

721 Finally, we found effective areas of ~2017 km<sup>2</sup> (outer area, blue contour line in Fig.1) and ~1178 km<sup>2</sup> (inner area,  
 722 red contour line in Fig.1) considering the two background values given above. The “inner area” reflects conditions  
 723 without ventilation effect, therefore the outer area is more appropriate for the emission estimates given that the  
 724 TROPOMI measurements occurred at 13:30 when the ventilation cannot be neglected. The other estimates  
 725 calculated from the inner area will be thereafter only indicated within brackets and considered to estimate the  
 726 sensitivity of the result.

727 Since the measured growth rate corresponds to a time interval of only 2 hours in the middle of the day, the CO  
 728 intraday fluctuations have to be taken into account. Stremme et al. (2013) used a factor which was taken from the  
 729 available bottom-up inventories and described that the CO emissions per/day are roughly 18.5 times the emission  
 730 per hour at noon. Assuming the same factor, we estimate a CO rate around  $0.71 \pm 0.06$  ( $0.42 \pm 0.04$ ) Tg/year for  
 731 2018. If no information about the diurnal distribution of the emission rate is available, we should assume a uniform  
 732 distribution and an upper value of the CO rate could be estimated using an intraday time interpolation factor of 24



733 hours instead of 18.5, finally resulting in ~30% higher estimates. Despite the significant uncertainties introduced  
 734 by spatial and temporal interpolation, their impact on the relative variability, trends and anomalies of the emission  
 735 rates is less important if the same method and assumptions are consistently applied across the entire time series.  
 736 CO<sub>2</sub> emissions could not be directly estimated using the same method, given its complex diurnal pattern, which is  
 737 a cumulative result of both natural and anthropogenic contributions and likely been influenced by additional  
 738 factors, related to instrumental and retrieval effects (i.e: air mass dependence error with a sub-percentage  
 739 contribution for CO<sub>2</sub>, non-ideal column sensitivity of the retrieval which represent near 25% overestimation for  
 740 CO<sub>2</sub> anomaly and 5% underestimation for CO anomaly in the PBL). Instead, we based our CO<sub>2</sub> estimates on the  
 741 measured XCO/XCO<sub>2</sub> ratio. The average XCO/XCO<sub>2</sub> molec. ratio ( $0.0164 \pm 0.0003$ ) determined from the UNA  
 742 and VAL total column measurement (Fig. 9) was converted to a mass ratio (multiplying it by the molecular weight  
 743 ratio) and found to be  $0.0100 \pm 0.0002$ . Considering this ratio, we estimated the CO<sub>2</sub> annual emission at  $71 \pm 6$  ( $42 \pm 4$ )  
 744 Tg/year for 2018. Our estimates of CO and CO<sub>2</sub> emissions by year and their average over the whole time series,  
 745 applying the same method, are presented in Fig. 11 and Table S3, concurrently with the SEDEMA inventories for  
 746 the MCMA. We obtained a 2016-2021 CO and CO<sub>2</sub> average emissions of  $0.55 \pm 0.02$  ( $0.32 \pm 0.01$ ) and  $46 \pm 2$  ( $32 \pm 1$ )  
 747 Tg/year, respectively, when excluding the lockdown period (Table S3). Here, the given uncertainties are solely  
 748 those stemming from the propagation of errors in growth rate estimates. Uncertainties on absolute values are much  
 749 higher when considering spatial and temporal extrapolations errors, but they do not influence the interpretation of  
 750 relative values.



752 **Figure 11: Comparison of CO and CO<sub>2</sub> emission estimations from UNA FTIR diurnal growth rates and from SEDEMA**  
 753 **inventories. For CO<sub>2</sub> (right), the estimates from Che et al. (submitted) are also reported, although it was based on the 10/2020**  
 754 **to 05/2021 period, after applying the same intraday temporal factor as used for our study to convert the Gg/hour to kt/year.**

## 755 5 Discussion

### 756 5.1 Long term variability

757 In this contribution, we characterised the seasonal and inter-annual variability and trends of the CO and CO<sub>2</sub> total  
758 column and surface concentrations from two urban and one background stations. The average total column 2013-  
759 2019 growth rate obtained at ALTZ (~2.5 ppm/year) and its inter-annual variability are in accordance with that  
760 typical of the Northern Hemisphere measurements from TCCON stations (hereafter, NH-TCCON) (Sussman et  
761 al., 2020: AGR of 2.4 ppm/year for the 2012-2019 period).

762 Both the NH-TCCON and ALTZ stations captured an important increase of the AGR in 2016 (+1.1 ppm/year for  
763 the TCCON stations and +2.1 ppm/year for the ALTZ station with respect to 2015), coinciding with the most  
764 intense ENSO (El Niño Southern Oscillation) event since the 1950s'. The impact of "El Niño" events on the carbon  
765 cycle is not yet fully understood, although they are consistently accompanied by a global increase of XCO<sub>2</sub> due to  
766 increasing drought in many regions and a decrease in global land carbon uptake. In 2016, an increase of 1.3  
767 ppm/year was observed in the Mauna Loa in situ AGR with respect to 2015 (Betts et al., 2018), for which the  
768 contribution of the 'El Niño' event was estimated at about 25%, the rest ascribed to an increase of the  
769 anthropogenic emissions. In Mexico, the "El Niño" events are generally associated with a decrease in  
770 precipitations, with deficits which can reach up to 250 mm in the South-Western area of the country, causing  
771 drought and a higher occurrence of wild and forest fires (Bravo-Cabrera et al., 2018, González del Castillo et al.,  
772 2020). Our observations from the ALTZ measurements highlight a much higher XCO<sub>2</sub> increase (+2.1 ppm/year)  
773 during 2016 with respect to 2015 than that observed at the NH-TCCON stations. During this period a small increase  
774 in the XCO MAGR (~ +0.02 ppm) is also observed at both ALTZ and UNA stations, maintaining the highest  
775 values of the whole time series over more than 4 months. Assuming that the CO MAGR variability captured at the  
776 ALTZ station during 2016 rather reflects a change in the global MCMA's emissions, we attempt to delineate the  
777 global and local contributions in the 2016 XCO<sub>2</sub> ALTZ AGR increase. Adopting a molecular CO/CO<sub>2</sub> ratio of ~  
778 0.016, a hypothetical increase of the XCO<sub>2</sub> MAGR over the 09/2015 - 09/2016 period due to the local emissions  
779 would be around +1.2 ppm/year, thus about 60% of the observed increasing rate during this period (+2.1 ppm/year).  
780 This gross estimate suggests that the El Niño regional effect only contributed at about 25% (0.9 ppm) to the  
781 observed AGR increase, which is close to the estimate from the NH-TCCON stations (~ +1.1 ppm) and from in  
782 situ data.

783 On the other hand, our long-term FTIR and surface time series allows examining the effect of the COVID-19 lock-  
784 down on the tropospheric CO<sub>2</sub> and CO concentration above the MCMA at local and regional scales. The reduction  
785 of the surface CO and CO<sub>2</sub> AGR at UNA (CO<sub>2</sub> AGR to a value close to zero, and CO AGR ~ -0.1 ppm/year) with  
786 respect to the other years (Fig. 3), and the strong diminution of their amplitude in the mean diurnal cycles clearly  
787 reflect a significant decrease of the local emissions near the UNA station, likely due to a drastic reduction of the  
788 urban traffic (the average annual congestion level decreased from 52% in 2019 to 36% in 2020 in Mexico City,  
789 from TomTom available estimates <https://www.tomtom.com/traffic-index/mexico-city-traffic/>).

790 The FTIR total column XCO<sub>2</sub> and XCO time series at UNA did not capture such a drastic change, only a small  
791 punctual decrease of the MAGR lower than the standard deviation of the whole time series was observed between  
792 April and October 2020. These results are in accordance with previous studies in other parts of the world. Although  
793 a reduction of 8.8% of the global CO<sub>2</sub> emissions was observed during the first five months of 2020 (Liu et al.,

794 2020; Jones et al., 2020) and an annual reduction from 4 to 7% (Le Quéré et al., 2020), the atmospheric total  
795 column XCO<sub>2</sub> showed a less clear effect (Sussman et al., 2020).

## 796 **5.2 CO/CO<sub>2</sub> ratio and MCMA emission estimates**

797 In this study, we robustly determined the CO/CO<sub>2</sub> ratio characterising the combustion efficiency of the city  
798 (0.016±0.01) from both surface and total column measurements at two urban stations. We found the same ratio for  
799 the UNA and VAL stations, and this ratio is very consistent with that found using the (UNA-VAL) gradients and  
800 using the surface measurements. This ratio is also consistent with that reported by MacDonald et al. (2023),  
801 calculated from TROPOMI and OCO-2/3 measurements (0.019) and slightly higher than that obtained from the  
802 EDGAR, FFDAS and ODIAC inventories (ratio ~0.012) reported in the same study.

803 Our estimate of CO emissions from the UNA measurements is based on a simplified approach, limited to  
804 days with low ventilation and time intervals corresponding to the late morning hours. It assumes a homogeneous  
805 area in the footprint and averages selected days without discrimination. Given that the temporal and spatial  
806 extrapolation introduces large uncertainties, only the relative and interannual behaviour of the emission can be  
807 discussed here, but the approach demonstrates how close column growth rate can be related to emission flux, if  
808 meteorological conditions allow neglecting advection. Our estimated range of CO emissions are consistent with  
809 the SEDEMA inventories at least for the year 2016 (factor 0.98) and 2018 (factor 1.04) if considering that they are  
810 dominated by the mobile sources. However, it is not the case for 2020, for which our estimate is much lower than  
811 SEDEMA by a factor of 0.3. During the lock-down period we estimated a decrease of about 55% compared to  
812 2018 while in the SEDEMA report, 2020 is the year with the maximum CO emissions (increase of 35% compared  
813 to 2018 considering the mobile sources). Both of these estimates contrast with Kutralam-Muniasamy et al. (2021),  
814 which reported an increase of 1.1% during the lock-down using the RAMA surface data. The large difference  
815 observed between these different studies can be due to i) the different methods used for extrapolating in space and  
816 time the emissions, ii) higher uncertainties of the FTIR-based estimates due to an important reduction the selected  
817 days of measurements and iii) an over-estimation of the SEDEMA inventory due to a lack of data during the lock-  
818 down period. Our estimate is based on the extrapolation of data from only one station (UNA), for which the  
819 dominant source is mainly the UNAM traffic activity. During the lockdown, the UNAM was closed and a  
820 significant reduction of the local traffic was recorded, but this traffic reduction was likely not representative of the  
821 whole MCMA. However, the decrease of the **MAGR** at both VAL and UNA stations does not support the increase  
822 of the CO emissions estimated by the SEDEMA inventory. Interestingly, it was not possible to apply the same  
823 method to calculate CO emissions at VAL because the average growth rate was close to zero (Fig. 6). This  
824 behaviour at VAL is likely due to the fast dispersion of the pollutant at this site, weakening the link between the  
825 diurnal pattern and the emissions.

826 Regarding CO<sub>2</sub>, our estimates also agree with the SEDEMA's inventory, especially if we consider the total  
827 emissions instead of mobile sources (factor of 1.2 and 1.1) for the years 2016 and 2018. For 2020, we estimated a  
828 decrease of 55% while the SEDEMA inventory indicates a decrease of about 10%. **The CO/CO<sub>2</sub> ratios calculated  
829 from the SEDEMA data for total emissions are similar to ours (0.014 and 0.011 in 2016 and 2018, respectively),  
830 suggesting that our average CO/CO<sub>2</sub> ratio is actually representative of the global mixing of the different sources  
831 of the MCMA, and not only dominated by the road traffic. Interestingly, according to the SEDEMA inventory,  
832 road traffic, the main anthropogenic CO source is identified by ratios (0.019 and 0.016 in 2016 and 2018,**

833 respectively) only slightly higher than our global average; whilst the industrial and domestic burning sectors, which  
834 represent the second main CO<sub>2</sub> anthropogenic sources, produces a one order of magnitude lower ratio. In any case,  
835 our measurements are well representative of the main source of the CO and CO<sub>2</sub> anthropogenic emissions. Indeed,  
836 if we consider the 2018 SEDEMA ratio for mobile sources (0.016), we find CO<sub>2</sub> emissions of the order of 43,100  
837 kt/year for this year, within ~5% of the SEDEMA estimates.

838 Our results were also compared with the estimates reported in Che et al. (submitted), based on an intensive FTIR  
839 measurement campaign performed during the 10/2020 to 05/2021 period and using a Column-Stochastic Time-  
840 Inverted Lagrangian Transport model (X-STILT) and a bayesian inversion (Fig. 11). Considering the same  
841 measurement period, our method leads to CO<sub>2</sub> emission estimates ranging between 29,000 and 49,800 kt/year  
842 using inner and outer effective area, respectively, which is consistent with the estimates obtained in Che et al.  
843 (submitted), ranging between 32,700 and 37,200 kt/year when applying the same intraday temporal extrapolation  
844 factor. Although the method we used for estimating the MCMA emissions is coarse and contains large  
845 uncertainties, mainly due to the temporal and space extrapolation, it shows the ability to use one station capturing  
846 the variability of the anthropogenic emissions of the MCMA and providing a year-by-year follow-up emission  
847 information without using complex dispersion models.

## 848 **6 Summary and conclusion**

849 We have analysed the variability of the total column XCO and XCO<sub>2</sub> above the MCMA from two urban and one  
850 background stations. The long-term XCO<sub>2</sub> data at the ALTZ station shows an average annual growth rate of ~2.5  
851 ppm/year, similar to what has been reported from TCCON stations in the northern hemisphere, and captured the  
852 perturbation driven by the 2015-2016 El Niño event. The urban stations show a similar growth rate (~2.3 ppm/year)  
853 and unlike at ALTZ, a slight decrease of XCO<sub>2</sub> and XCO during the COVID19 lock-down period could be  
854 observed. The CO<sub>2</sub> and CO concentrations within the mixed layer, estimated from the FTIR total column  
855 measurements and ceilometer data, were found to be consistent with the surface measurements. These findings  
856 confirm that the concentrations near the surface are mainly controlled by the emissions and the daily behaviour of  
857 the mixed layer in MCMA. Our long-term total column and surface time series from both urban stations allowed  
858 us to determine with great confidence an average CO/CO<sub>2</sub> ratio, indicative of the Mexico City combustion  
859 efficiency. The CO/CO<sub>2</sub> ratio over our long-term measurement period seems to be fairly constant and equals  
860 ~0.016 (mass ratio: 0.010). This value is consistent with other studies such as from satellite measurements (OCO-  
861 2/3 and TROPOMI) and the bottom-up inventories reported by MacDonald et al. (2023). Finally, we estimated the  
862 CO emissions using the average daily growth rate determined from measurements at the UNA station. Although  
863 this method likely leads to an under-estimate of the emissions due to the non-negligible effects of advection, our  
864 results were found to be very consistent with the 2016 and 2018 SEDEMA inventories. The same strategy could  
865 not be applied at the VAL station, likely because of dominant southward advection of the airmass, due to the  
866 complex topography in this part of the MCMA. In contrast, the UNA station is located in a flat ground downwind  
867 of the main anthropogenic source of the MCMA which likely allows establishing a direct relationship between the  
868 columnar measurements and the MCMA CO and CO<sub>2</sub> emissions. We finally estimated the CO<sub>2</sub> emissions using  
869 the CO growth rate and the CO/CO<sub>2</sub> ratio. The finding that our CO<sub>2</sub> emission estimates are within 20% of those of  
870 SEDEMA for total emissions show that our ratio reflects not only the traffic sources but is also affected by other  
871 sources such as industrial activities and domestic burning. The UNA station, with its advantageous orography, is

872 therefore a good site to capture well-mixed emissions from the city and serves as a site to follow the interannual  
873 variability and trends of the emissions in this urban environment. Finally, this study showed the feasibility to  
874 monitor the long-term evolution of anthropogenic CO<sub>2</sub> and CO emissions in Mexico City by deploying only a few  
875 EM27/SUN instruments. **The methodology employed here for monitoring the long-term temporal variability of**  
876 **CO emission fluxes is likely to be adapted to other urban areas where the topography damps the ventilation down**  
877 **for several hours each day, thereby establishing that the column growth rate is dominated by the emission flux.**  
878 **Although the straightforward model presented here is not intended to replace a complex transport/chemical model**  
879 **for a precise estimate of city emissions, the results obtained demonstrate that it is nevertheless possible to track**  
880 **their temporal evolution with a high degree of reliability.**

## 881 **7 Author contribution**

882 All the co-authors contributed in the discussion of concepts, and to the preparation of the manuscript. NT, WS and  
883 MG were responsible of FTIR measurements and the data analysis. MG and WS lead the ALTZ station  
884 development and its long-term operation. AB and EGC were responsible of the maintenance of the instruments at  
885 the Alzomoni station. VA helped to classify the days and hours with low ventilation and strong turbulence and  
886 provided the UNAM emission inventory. EGC was in charge of the in-situ measurements, with the support of OL.  
887 MG and MR led the MERCI-CO<sub>2</sub> project. FH lead at KIT the German-Mexican collaboration for the deployment  
888 of the high resolution FTIR spectrometer and supports its long-term operation as part of NDACC. FH has helped  
889 in the design and setup of the spectrometer and solar tracker before it was shipped to Mexico. He has developed  
890 the retrieval code PROFFIT and gives continuously support to the UNAM group for its use and in operating the  
891 spectrometer. FH and CA lead the German-Mexican collaboration and give precious help for the EM27/Sun  
892 measurements in the frame of the COCCON network. All the co-authors contributed of the **writing** of the  
893 manuscript.

## 894 **8 Competing interests**

895 The authors declare that they have no conflict of interest.

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