



1 **Marine cloud base height retrieval from MODIS cloud properties using**
2 **machine learning**

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12 **Abstract**

13

14 Clouds are a crucial regulator in the Earth's energy budget through their radiative properties, both at the top-of-the-atmosphere
15 and at the surface, hence determining key factors like their vertical extent is of essential interest. While the cloud top height is
16 commonly retrieved by satellites, the cloud base height is difficult to estimate from satellite remote sensing data. Here we present
17 a novel method leveraging spatially resolved cloud properties from the MODIS instrument to retrieve the cloud base height over
18 marine areas. A machine learning model is built with two components to facilitate the cloud base height retrieval: the first
19 component is an autoencoder designed to learn a representation of the data cubes of cloud properties and reduce their
20 dimensionality. The second component is developed for predicting the cloud base using ground-based ceilometer observations
21 from the lower dimensional encodings generated by the aforementioned autoencoder. The method is then evaluated based on a
22 collection of co-located surface ceilometer observations and retrievals from the CALIOP satellite lidar. The statistical model
23 performs well on both datasets, exhibiting accurate predictions in particular for lower cloud bases and a narrow distribution of
24 the absolute error, namely 379 m and 328 m for the mean absolute error and the standard deviation of the absolute error
25 respectively for cloud bases in the test set. Furthermore, cloud base height predictions are generated for an entire year over
26 ocean, and global mean aggregates are also presented, providing insights about global cloud base height distribution and offering
27 a valuable dataset for extensive studies requiring global cloud base height retrievals. The global cloud base height dataset and the
28 presented models are available from Zenodo (Lenhardt et al., 2024).



29 **1 Introduction**

30

31 Clouds play a key role in the Earth's energy budget through their interactions with incoming shortwave and outgoing longwave
32 radiation fluxes. It is thus critical to adequately quantify cloud radiative properties and their changes under global climate
33 change. However, cloud radiative properties remain a large uncertainty in estimating anthropogenic climate change and possible
34 impacts in the future (Boucher et al., 2013; Forster et al. 2021). Radiative properties of clouds are related to numerous quantities
35 that can be used to characterise them. For instance, the cloud base height (CBH) is a crucial radiative property through its impact
36 on the surface longwave radiation. Furthermore, the cloud geometrical thickness (CGT), defined as the difference between the
37 cloud top height (CTH) and the CBH, links to the adiabatic cloud water content allowing the quantification of the cloud's
38 subadiabaticity. Additionally, deriving the CBH is of practical use for pilots, providing crucial information during flights.

39 However, while the CTH can be rather easily obtained through passive satellite observations, the CBH retrieval remains
40 problematic due to the fact that it is only indirectly accessible to satellites, and due to retrieval errors related to satellite remote
41 sensing such as instrument shortcomings or noisy measurements. Since the difference between the CTH and the CBH quantifies
42 the vertical extent of a cloud, one way to retrieve the CBH from passive satellites is by making heavy assumptions on the vertical
43 distribution of the cloud water path inside the cloud profile. It is thus a challenging retrieval with passive satellites data that
44 provide information about the cloud top (e.g. cloud top temperature (CTT), pressure (CTP) or height (CTH)) or about the entire
45 column (e.g. cloud optical thickness (COT)) assuming the cloud's adiabaticity. For example, Noh et al. (2017) rely on a
46 semiempirical approach to link the CGT to the CTH and the cloud water path (CWP, includes both ice and liquid water paths). In
47 a different approach, Böhm et al. (2019) retrieve the CBH from triangulation of a multi-angle spectroradiometer. However, in
48 this case, assumptions were required on the distribution of convective clouds. On the other hand, active satellite remote sensing
49 retrieves information with vertical resolution which greatly helps resolving the clouds vertical distribution. However, active
50 satellite measurements can display attenuated signals close to the surface (Tanelli et al., 2008; Marchand et al., 2008) particularly
51 in the presence of thick clouds or precipitation, rendering the retrieval of the CBH difficult even for radar and lidar. Among
52 others, Mülmenstädt et al. (2018) and Lu et al. (2021) present methods focusing on low clouds which use the CBH from active
53 satellite retrievals of neighbouring thin clouds as representative of the surrounding cloud field. Active remote sensing
54 additionally suffers from the sparse sampling that is confined to a narrow swath below the satellite. Finally, Goren et al. (2018)
55 combine information from both passive and active satellite remote sensing and rely upon an adiabatic cloud model to derive the
56 CBH. More generally, remote sensing retrievals of the CBH rely on the assumed homogeneity of the cloud field in the vicinity of
57 its base.

58 The retrieval of the CBH using satellite remote sensing data relies on a number of simplifying assumptions and is, consequently,
59 prone to errors. Subsequent uncertainties in the estimation of the CBH can then relate to uncertainties in the overall cloud
60 radiative effect (CRE) (Kato et al., 2011; Trenberth et al., 2009).

61 The method presented here leverages passive satellite retrievals of cloud properties in combination with marine surface
62 observations to derive the CBH of a cloud scene using an innovative machine learning (ML) model. Our developed ML model
63 aims to draw on the spatial information present in a cloud scene in combination with relevant cloud properties to inform the CBH
64 prediction. As the CBH is typically derived from the surface, we focus on lower clouds in particular as the retrieval quality is
65 generally higher for those clouds, and as it is the lowest cloud that often matters most (e.g. for the surface radiation budget). The
66 combination of satellite remote sensing and surface-based CBH retrievals has the potential to provide robust global CBH
67 estimates.

68 Section 2 firstly introduces the datasets and the co-location between ground-based observations and satellite retrievals. Secondly,
69 the ML method is described. In section 3 we evaluate our predictions against other methods including Noh et al. (2017) and other
70 products from active satellite measurements like the 2B-CLDCLASS-LIDAR product (Sassen et al., 2008). Section 4 presents
71 the global dataset of the CBH which is derived from the ML approach. We discuss the benefits and remaining challenges of our
72 method in section 5. Further details about the spatial distribution of the observations and the ML method are included in the
73 appendices A-E. Additional links to available data outputs and codes are listed in the corresponding sections.

74

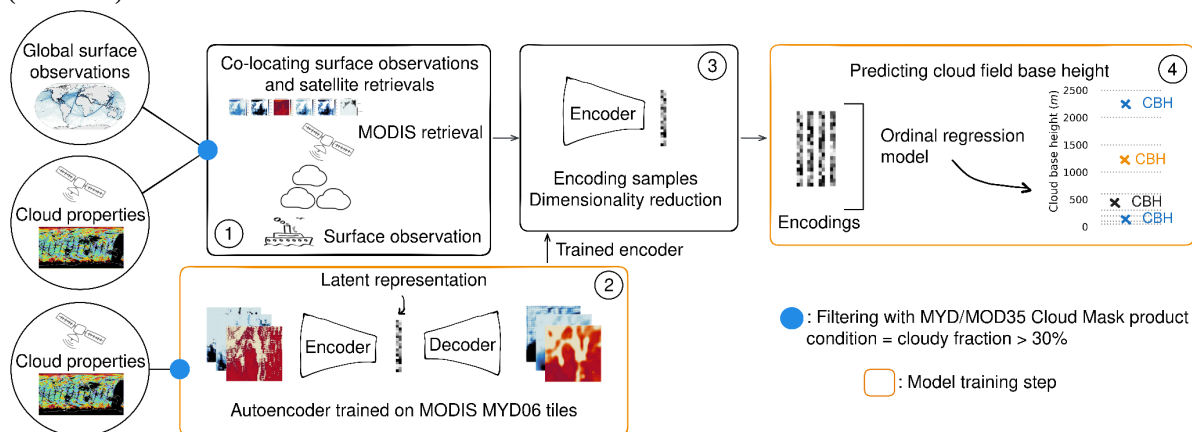
75 **2 Data and methods**

76

77 In this study we approach the retrieval of the CBH of a cloud scene by combining marine surface-based observations of the CBH
78 and passive satellite retrievals of relevant cloud properties. The cloud scenes are defined within a tile of size 128 km x 128 km,
79 which incorporates different satellite-retrieved cloud properties at a 1 km horizontal resolution from the MODerate Resolution
80 Imaging Spectroradiometer (MODIS, Platnick et al. (2017)). The satellite retrievals concern the CTH, the COT and the CWP,



81 which are related to the ground-based CBH observations (cf. [Table 1](#)). We focus on marine regions to remove the impact of
 82 orography on surface observations especially for low level clouds. The approach is based on the assumption that the CBH is
 83 homogeneous in the considered cloud scenes (similar to e.g., Böhm et al., 2019). To leverage the spatial extent of the cloud scene
 84 and derive relevant features from the input channels, we rely on convolutional neural networks (CNNs, LeCun et al., 1989;
 85 LeCun et al., 1995). This type of artificial neural network has been widely used in computer vision (Krizhevsky et al., 2012;
 86 LeCun et al., 2010) but also more recently in various applications in climate science (Reichstein et al., 2019; Watson-Parris et al.,
 87 2022). CNNs typically require a large number of labelled training samples due to their high number of parameters. However, the
 88 co-location step between surface-based observations and satellite retrievals limits the number of available data samples to train
 89 the prediction model. We overcome this hurdle by introducing an unsupervised step using unlabeled satellite data.
 90 Hence, the novel method we present here can be summarised in four main steps ([Fig. 1](#)) and are further elaborated on in the
 91 following sections: Firstly, we co-locate ground-based CBH observations and corresponding satellite-retrieved cloud properties
 92 from MODIS (cf. sections 2.1, 2.2, 2.3 for more information on ground-based observations, satellite retrievals and co-location,
 93 respectively). Secondly, we train an autoencoder (AE) with a CNN backbone solely on MODIS data in order to extract relevant
 94 features from the cloud scenes (section 2.4). Thirdly, we project the cloud properties tiles from the co-located dataset to the latent
 95 feature space constructed by the encoder. Ultimately, we predict the CBH from the encodings using an ordinal regression model
 96 (section 2.5).



97 **Figure 1: Schematic of the cloud base height retrieval method. 1) Co-location of surface-based cloud base height**
 98 **observations and satellite retrievals. 2) Autoencoder training on satellite cloud properties. 3) Encoding of co-located**
 99 **samples using the trained encoder. 4) Prediction of the cloud field base height.**

102 2.1 Surface observations

104 The CBH labels used in this study are part of a global marine meteorological observation dataset maintained by the UK Met
 105 Office (Met Office, 2006), which provides observational data ongoing from 1854. The observations are conducted from
 106 measuring stations that were located on ships, buoys or platforms. As a consequence, this study largely relies on observational
 107 data representing the areas along the corresponding ship routes ([Fig. 2a](#)). At the beginning of meteorological and weather reports,
 108 surface-based cloud observations were retrieved manually or visually by human observers, but they have been gradually replaced
 109 by automated systems.

110 The CBH is derived using a ceilometer, an instrument based on a laser pointing upright and measuring the backscatter from the
 111 cloud base, and is then reported following the current standards from the World Meteorological Organisation (WMO; WMO,
 112 2019). The CBH observations are sorted into bins of increasing width (from 50 m to 500 m bin width) corresponding to the
 113 altitude ([Fig. 2b](#)). As a result, the binning process can lead to an underestimation of the actual CBH, especially for a higher CBH
 114 for which the bin size is larger. In addition, the surface-based observations specify quantities like temperature, humidity and
 115 wind speed at a given time and location.

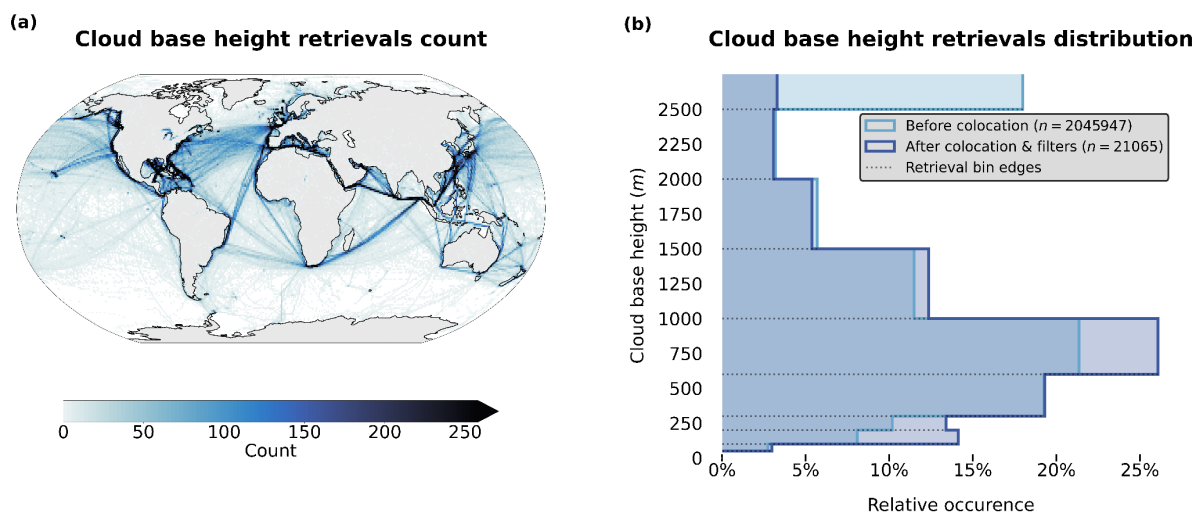
116 Despite their coarse resolution, the reported cloud base observations provide valuable information of clouds in remote marine
 117 areas. The distribution of CBH observations and corresponding bins are shown in [Figure 2](#).



118

Data product	Description	Variables	Resolution	Usage
Global marine meteorological observations (Met Office, 2006)	Surface observations	Cloud base height (m)	Latitude/longitude coordinates 0.1° Hourly/daily observations	Labels
MODIS Atmosphere L2 Cloud Product (MYD06) (Platnick et al., 2017)	Cloud-top properties, cloud optical and microphysical properties	Cloud top height, CTH (m) Cloud optical thickness, COT (a.u.) Cloud water path, CWP (g.m ⁻²)	1 km pixel resolution Daily overpass	Input features
MODIS Atmosphere L2 Cloud Mask Product (MYD35) (Ackerman et al., 2017)	Cloud pixel flag	Cloud mask	1 km pixel resolution Daily overpass	Used for cloud scene filtering

119 **Table 1 : Dataset description. The MODIS data are derived from the collection 6.1 of the datasets (Platnick et al., 2017;**
 120 **Ackerman et al., 2017; cf. section 2.1). The surface observations are provided by a worldwide station network available**
 121 **from the UK MetOffice (Met Office, 2006; cf. section 2.2).**
 122



123

124

125 **Figure 2: (a) Spatial distribution of cloud base retrievals count (1° grid) and (b) distribution of the retrieved cloud base**
 126 **height before and after the co-location and filtering process, for observations from the years 2008 and 2016.**

127

128 2.2 Satellite data

129

130 In this study we use MODIS products from the AQUA satellites as input data that is later combined with the CBH labels derived
 131 from the surface-based observations to train the prediction model. We choose MODIS satellite retrievals as they provide a large
 132 amount of data with kilometre-scale resolution and daily overpasses. The spatial coverage of one swath is around 2330 km x
 133 2000 km. We make use of the CUMULO dataset (Zantedeschi et al., 2019) since it provides already preprocessed satellite data
 134 from the A-train with daily full coverage of the Earth for the years 2008 and 2016. In particular out of the available variables we
 135 use two aligned products (cf. [Table 1](#)), namely the MODIS06 level 2 cloud product (hereafter MYD06; Platnick et al., 2017)



136 which provides relevant cloud properties and the MODIS35 level 2 cloud flag mask (hereafter MYD35; Ackerman et al., 2017)
137 which allows us to filter scenes and screen for clouds.

138 The MYD06 product contains various cloud top properties (temperature, pressure, height) and cloud optical and microphysical
139 properties (optical thickness, effective radius, water path). Level 2 data are derived from calibrated radiances through various
140 algorithms and physical relations detailed in Platnick et al. (2017). The cloud top quantities are derived from radiance data of
141 several channels. Wavelengths in the CO₂ absorption range are particularly used to identify the cloud top pressure (CTP) of high
142 clouds because of the opacity of CO₂. For thicker or low boundary layer clouds, infrared bands are additionally required. The
143 cloud optical thickness (COT) and cloud effective radius (CER) are simultaneously derived from multispectral reflectances,
144 cloud masks, CTP data and surface type characteristics. The cloud water path (CWP) is additionally retrieved as part of the cloud
145 optical properties algorithm described in Platnick et al. (2017).

146 In general, the MYD06 level 2 product offers the advantage that the statistical model can be built relying on cloud properties and
147 it can thus allow the study of relationships between the CBH and other cloud properties. Calibrated radiances, one step ahead in
148 the data processing pipeline, would also provide insightful information but would require inputs of larger dimensionality since
149 key information about clouds would be scarcer. Furthermore, using MYD06 level 2 data allows us to compare our method to
150 others which in most cases use cloud properties to retrieve the CBH. It is to be noted that the level 2 product provides pre-
151 processed data on top of the calibrated radiances and reflectances of level 1 data, which might introduce biases in the statistical
152 model. From the entirety of available MYD06 retrievals, we select three cloud properties in particular, namely the CTH, COT,
153 and CWP. The CTH is used as it provides key information about the CBH in the cloud field, as seen in Böhm et al. (2019).
154 Vertically integrated cloud quantities like the COT and CWP further help the statistical model by providing key information
155 about the cloud's vertical extent, lacking in cloud top only properties, making them commonly used for retrieving the CBH (e.g.
156 Noh et al., 2017). The CWP as computed from COT and CER, and, in consequence, also the CBH are built on adiabatic
157 assumptions (Grosvenor et al., 2018) and therefore cannot be used to constrain subadiabaticity as also highlighted in
158 Mülmenstädt et al. (2018).

159

160 2.3 Datasets co-location

161

162 We proceed to match our two data sources over the two years of data available. To obtain the cloud properties of the cloud scene
163 corresponding to the surface retrieval of CBH we select a square tile of 128 km from the *closest* MODIS swath available around
164 the observation location. Here *closest* means that the MODIS swath contains the (latitude, longitude) coordinate of the CBH
165 observation and the satellite retrieval was made during a one hour time-window before/after the CBH observation time. The
166 spatial scale of the extracted satellite retrieval was chosen in order to give enough spatial information to the AE while ensuring
167 the measured CBH is representative of the observed satellite retrieval. This spatial scale corresponds to using information from
168 clouds in approximately a 60 km radius around the observation location. Such a threshold is an adequate compromise between
169 considering all the relevant information while not discarding too many samples which might fall outside of the distance limit.
170 These spatial and temporal thresholds for the co-location are in line with other similar studies (Mülmenstädt et al. (2018) 100 km
171 and 1 hour; Lu et al. (2021) 150 km and 30 minutes; Böhm et al. (2019) 20 km and 15 minutes; Noh et al. (2017) 0.1 degree and
172 5 minutes) even though the data products are partially different here. We furthermore add a condition that the corresponding tile
173 is fully located inside of the swath to avoid any missing data in the cloud scene. The extracted tile is then filtered using the
174 MOD35 product to only keep the cloud scenes with at least a 30% cloud cover. The latter condition is primarily aimed at
175 retrievals of poor quality leading to missing pixels which is predominantly the case for the COT and CWP channels for which
176 the retrieval fails more frequently. However it leads to a higher rate of removal for higher CBH observations (Fig. 2). Lowering
177 the cloud cover filter led to a higher number of usable samples but ultimately did not improve the model's performance.

178 The overall filtering and co-location process yields around 21 000 samples. This only represents around 1% of the initial CBH
179 observations mainly due to the co-location process both in time and space with the MODIS overpasses. Missing values and cloud
180 cover filters are an additional factor in the reduced number of co-located samples. The presented co-located dataset is the basis to
181 build our cloud scene CBH prediction.

182 Classical semi-supervised pipelines, like the one presented here, characterised by a small labelled dataset and a vast unlabelled
183 dataset, necessitate this kind of co-location or matching process. However, future avenues of research could consider directly
184 modelling unmatched datasets, as in e.g. Lun Chau et al. (2021), which could additionally make use of other variables present in
185 the surface observations.

186



187 2.4 Autoencoder

188

189 To circumvent the lack of labelled samples from which the relevant features are extracted, and to work in a lower dimensional
190 space, we add a dimensionality reduction step to our method through an unsupervised learning model. AEs (Kramer, 1991;
191 Hinton et al., 2006) offer a wide application spectrum, ranging from preprocessing to the generation of new outputs. AEs are
192 commonly used in unsupervised learning settings for reducing the dimension of the input data to leverage the latent
193 representations learned by the model to perform clustering, classification or regression in a lower dimensional space (Baldi et al.,
194 2012). We use classical AEs for their simplicity and versatility, but it should be noted that other approaches to unsupervised
195 latent representation learning, such as variational AEs and its many variants, can be used in a similar fashion.

196 In general, AEs learn to encode the given input data to produce a latent representation of lower dimension. From the latent
197 representation, the input data is then reconstructed. The learning process is driven by what is called the reconstruction loss that
198 minimises the difference between the input and the reconstructed output.

199 Here we use a convolutional AE architecture which is based on a CNN (LeCun et al., 1989; LeCun et al., 1995) backbone in
200 order to leverage the spatial structure of our input data (Pu et al., 2016). However, this unsupervised step requires a large amount
201 of data that the AE can learn from. Therefore, we select one full year of data of MODIS swaths from the CUMULO dataset
202 (from the year 2008, cf. section 2.2) and randomly sample tiles following the same criteria as during the co-location process (cf.
203 section 2.3). We thus create around 500 000 tiles to train our model. We sample a maximum of 20 tiles from a single swath and
204 this for only a single year of data in order to avoid possible spatial and temporal auto-correlation in the data used for training and
205 testing leading to a non-representative performance of the mode (Kattenborn et al., 2022). Further details on the study of the
206 generalisation performance of the model for new observations in space and time are given in appendix B.

207 Using the relevant MYD06 retrievals as input data (cf. section 2.2), we define several convolution layers grouped into a total of
208 five blocks for both the encoder and the decoder. The architecture of the decoder is thereby being mirrored to the encoder. Each
209 block consists of three convolutional layers with a kernel-size of 3 and Leaky Rectified Linear Units (LeakyReLU; Maas et al.,
210 2013) as activation functions. At the end of each block, a maximum pooling layer is added with a kernel-size and a stride of 2.
211 The model code was developed following implementations from the packages *PyTorch* (Paszke et al., 2019) and *TorchVision*
212 (*TorchVision*, 2016).

213 The main goal of the AE is to minimise the loss function during the optimization or learning process, and to reproduce the input
214 data with the highest fidelity. We denote the sampled tiles used for training the AE by $B = \{b_i \in \mathbb{R}^{3 \times 128 \times 128}\}_{i \in [1, N]}$, with

215 $N \sim 500\,000$. A common choice for the reconstruction metric is the ℓ_2 - norm: $\mathcal{L}_{reconstruction} = \sum_{b \in B_i} \|b - D_\theta(E_\theta(b))\|_2^2$ where

216 B_i represents a batch of samples and θ the parameters of the encoder E and decoder D models. Details of the AE architecture,
217 training and performance are provided in appendix C.

218

219 2.5 Cloud base height ordinal regression

220

221 Once the AE's optimization process is completed (cf. appendix C), the next step is to predict the corresponding CBH for the
222 observed scene. As seen in [Figure 2](#), the retrieved CBH observations are binned into different categories following WMO
223 standards (WMO, 2019). This leads to a prediction problem at the intersection of regression (i.e. predicting numerical values)
224 and classification (i.e. predicting the object class) called ordinal regression (OR; Winship et al., 1984). The labels from the target
225 variable are defined by classes following a certain order, in this case the increasing CBH. A wide array of methods stems from
226 this field with diverse applications for example in computer vision (e.g. Niu et al., 2016; Shi et al., 2023). Different methods
227 exist to tackle such problem setups either via modification of the target variable, ordinal binary decomposition or threshold
228 modelisation (Gutiérrez et al., 2016; Pedregosa et al., 2017). Threshold models were shown to be able to perform better than the
229 ones designed for regression or multi-class classification on OR tasks (Rennie et al., 2005). Further details on threshold OR
230 models are added in appendix D. We use the OR implementation of threshold models from the *mord* Python package (based on
231 Pedregosa, 2015). A ℓ_2 regularisation term is also added during the optimization process. We adopt the macro-averaged mean
232 absolute error (MA-MAE) as our reference metric during hyperparameter tuning. This metric is in particular useful for OR
233 problems when faced with imbalanced datasets (Bacianella et al., 2009). Using a macro-averaged metric prevents us from
234 choosing a trivial model which might always predict the dominating class. We additionally reported the macro-averaged root
235 mean square error (MA-RMSE) during training and validation of the models as it puts a larger penalty than the MAE on higher
236 errors and is also a useful performance indicator.



237 3 Results, evaluation, and comparison to previous retrieval approaches

238

239 3.1 Cloud base height retrieval, evaluation and comparison to previous retrievals

240

241 In this section, we present the results of the retrieval, evaluate it using the ground-based observations, and investigate how our
242 method fares by comparing it to a method assuming an adiabatic cloud model (adapted from Goren et al. (2018), cf. appendix E
243 for implementation) and to the method from Noh et al. (2017). It is to be mentioned that, for the former the sources of the CTH
244 retrievals differ, and for the latter a different method was used for retrieving the CTH from the available MODIS CTP. For these
245 two methods we first compute a CBH value for each cloudy pixel of the scene that is then averaged. The analysis is performed
246 for the co-located scenes where ground-based observations are available. To be able to compare the relevant metrics for the
247 different methods we proceed to a binning of the data following the WMO standard presented in section 2.1. In [Table 2](#) we report
248 several metrics including the MAE, the mean error (bias), the RMSE and the standard deviation of the absolute error. The latter
249 helps us characterise the spread and uncertainty in the overall predictions with respect to the surface observations. Furthermore,
250 we do not report quantities such as the correlation coefficient or the regression line on the 2-dimensional histograms of [Figure 3](#)
251 and [Figure 4](#), as the stratified and categorical aspects of the data would make reporting these not clearly informative. It is to be
252 noted that later on we refer to the overall conceived method including the AE (cf. section 2.4) and the OR prediction model (cf.
253 section 2.5), listed in [Table 2](#) as OR + AE, interchangeably as OR or as the *prediction model*.

254 We first note that the OR method with an immediate-threshold setup fails at predicting with good accuracy the cloud scene base
255 height, producing large errors (double-fold in comparison to the all-threshold setup). On the other hand the OR method with an
256 all-threshold setup performs well with satisfying error measures and uncertainty in the predictions. Compared to the method from
257 Noh et al. (2017), our method succeeds in decreasing on average the error, displaying a reduction of 100 m for the MAE. The
258 method also effectively diminishes the uncertainty in the CBH retrievals, bringing down the absolute error standard deviation
259 200 m lower. Our method thus provides accurate retrievals with comparatively low general uncertainty levels. Even though on
260 average the predictions exhibit a slight positive bias, we find that the CBH values above 2000 m are systematically
261 underestimated ([Fig. 3](#)). In consideration of the low representation of such observations in the dataset, due to data filtering and
262 surface observations being less reliable for higher clouds, the method still struggles to properly quantify the cloud scene base
263 height of these samples. These samples also make up for most of the measurement uncertainty in the labels considering that
264 ceilometers face challenges for retrieving cloud signals higher up in the boundary layer. Focusing on lower cloud scene base
265 height retrievals, the predictions demonstrate even lower errors: the MAE is lowered to 379 m while the absolute error standard
266 deviation is narrowed down to 328 m. Achieved accuracy levels and uncertainty measures attest to a certain trustworthiness of
267 the cloud scene base height estimates, in particular in the context of product requirements for example the ones outlined by the
268 Joint Polar Satellite System (JPSS; Goldberg et al. (2013); 2 km accuracy threshold). However, the cloud scene base height
269 retrieval method presented here does not aim at constituting a product on its own as it is not operational with the processing of
270 daily new data available from the MODIS instrument, but rather at providing robust estimates of CBH for lower level clouds.
271 Therefore, it is expected and reasonable that the accuracies and uncertainties presented here are below such thresholds. However,
272 the available method code (Lenhardt et al., 2024) easily allows the processing of new data for users, in addition to the available
273 dataset for the year 2016.

274 We performed further sensitivity studies on our retrieval method trying to improve the quality of the predictions. An attempt to
275 balance the dataset by oversampling the higher CBH values (cloud base retrievals falling into the 2500 m bin), however, did not
276 yield better results overall but also posed a higher risk of overfitting to these specific samples. Furthermore, any spatial
277 information about the location of the satellite retrieval was not included as to prevent possible overfitting to the latitude and
278 longitude coordinates of the observations present in the training data. Since the observations are sparsely distributed especially in
279 the southern hemisphere (cf. figures from appendix A), the goal is to avoid any kind of induced spatial bias and sensitivity in the
280 model's predictions. Accordingly we can then ensure proper generalisation skill to new spatial areas, but not only based on
281 known retrieval distributions at similar locations. Correspondingly, the generalisation skill of the model requires further
282 assessment to guarantee meaningful and representative predictions. Spatial generalisation is rather challenging as the co-located
283 samples are so sparsely distributed ([Fig. A.3](#), [Fig. A.4](#)). Limiting the training dataset to a selected area would greatly hinder the
284 representativeness notably because the different labels display diverse spatial patterns. As a consequence, the choice was made to
285 evaluate the potential generalisation skill of the prediction model by establishing a geographic distribution of the mean predicted
286 cloud scene base height for a whole year's worth of MODIS overpasses. This is discussed in more detail in section 4. On the
287 other hand, the temporal aspect of the model's generalisation skill was intrinsically ensured by building a test set temporally
288 distinct from the training set, including co-located samples only from the last months of 2016.

289



Method	MAE (m)	Bias (m)	RMSE (m)	Absolute error standard deviation (m)
Goren et al. (2018)	457	- 262	689	515
Noh et al. (2017)	578	- 35	860	638
OR (IT) + AE	991	+ 595	1296	836
OR (AT) + AE	447	+ 58	614	420

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Table 2: Performance on the test set of different CBH retrieval methods. OR models are either built with the immediate threshold (IT) or all threshold (AT) variant. The method on which the rest of the study is based has been highlighted in bold.

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3.2 Comparison to spaceborne radar-lidar retrievals of the CBH

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The combined datasets which are part of CUMULO (Zantedeschi et al., 2019), in particular the radar and lidar retrievals, facilitate the joint evaluation of our method with both ceilometer surface observations and active satellite retrievals. Specifically we leverage the 2B-CLDCLASS-LIDAR product (Sassen et al., 2008) which is derived from the combination of CloudSat’s Cloud Profiling Radar (CPR; Stephens et al., 2008) and CALIPSO’s Cloud-Aerosol Lidar with Orthogonal Polarisation (CALIOP; Hunt et al., 2009). The base height of the lowest cloud layer retrieved by the instruments in each scene is considered the scene CBH and then averaged over the available pixels along the track, preserving the same spatial extent as the associated cloud properties from the MODIS instrument. For the co-located samples of the year 2008, we thus jointly retrieve the obtained CBH from the 2B-CLDCLASS-LIDAR product, only considering cases where a surface observation was in the vicinity of the satellite track (inside a disc with a ~60 km radius around the surface observation, cf. section 2.3). For the samples fulfilling these conditions, we then compare how the different retrievals fare. In Figure 4, the joint histograms for the surface observations, the 2B-CLDCLASS-LIDAR retrieval and the method’s corresponding predictions are documented, representing a total of around 800 samples.

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Investigating the joint histogram between the surface observations and the 2B-CLDCLASS-LIDAR retrievals (Fig. 4a) allows to identify shortcomings of the active satellite retrievals in particular close to the surface (Tanelli et al., 2008; Marchand et al., 2008). Indeed, the 2B-CLDCLASS-LIDAR retrievals closer to the surface are not well captured as partially expected, due to thick clouds attenuating the lidar signal, and due to ground clutter and lack of sensitivity to small droplets near cloud base for the radar signal. A similar explanation can eventually be articulated as a whole for the co-located retrievals, considering that the mean bias between the two retrievals is greater than + 600 m. Concurrently, it is fruitful to compare the 2B-CLDCLASS-LIDAR retrievals with the predictions from the developed method (Fig. 4b). As seen previously, the OR method struggles at higher CBHs, but agrees here reasonably well with the active satellite retrievals, especially for retrievals between 500 m and 1500 m. Focusing on retrievals under 1.5 km, the prediction model achieves similar performance as presented in Table 2, even though the subset here is much smaller.

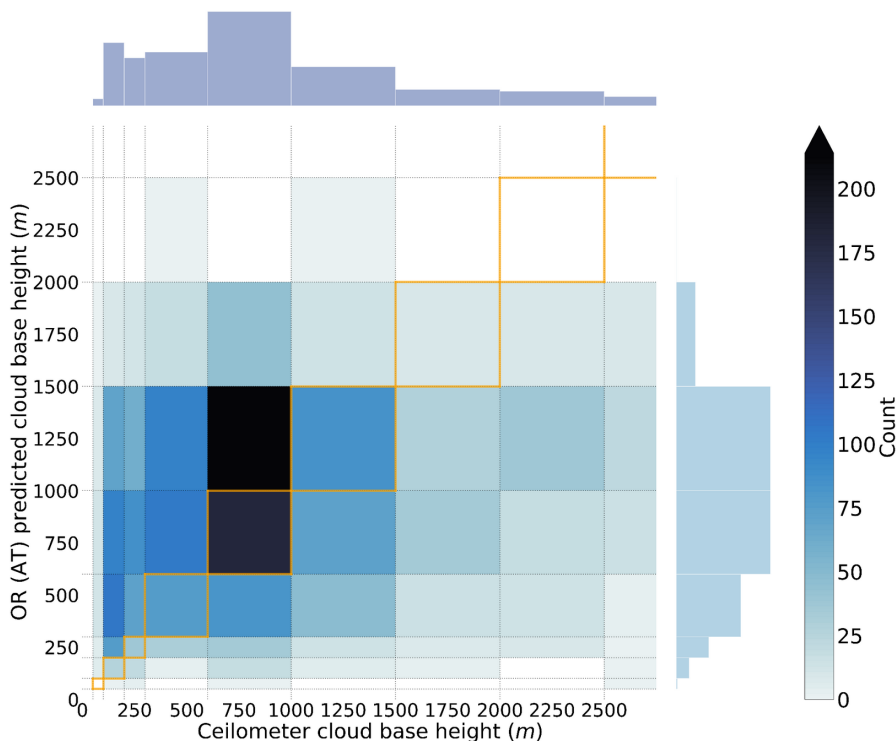
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Furthermore, we created a more extensive dataset using only 2B-CLDCLASS-LIDAR retrievals and the cloud scene predictions with the aim of obtaining a more complete view of the relationship between these two retrievals. To this extent, we collated around 160 000 samples of aligned cloud scene base height predictions and the 2B-CLDCLASS-LIDAR retrievals over the year 2016. For this dataset, the performance metrics exhibit similar values as on the previously presented subset, displaying even lower values for the MAE and the absolute error standard deviation (around a 50 m decrease for both). Similarly to the previous co-located subset, limiting the evaluation to lower cloud base retrievals yields performance metrics close to a 450 m MAE and a 270 m absolute error standard deviation, both of these being mainly impacted by agreeing retrievals in the 500 m to 1500 m range.

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Joint histogram - Surface observations and model predictions



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Figure 3: Joint histogram over the test set of the surface observations and the predicted cloud scene base height with the ordinal regression all-threshold model. The 1:1 boxes are highlighted in orange in the figure.

331 4 Global distribution

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333 To further evaluate the method, we also apply the prediction model on global MODIS data for the whole year of 2016. The
334 sampling process yields approximately 700 000 CBH retrievals for the corresponding cloud properties tiles. The final prediction
335 model was beforehand re-trained on the whole co-located dataset including the test set of section 3.1. We then spatially aggregate
336 the predictions over the year and consider the spatial mean and median absolute deviation (MAD). The MAD constitutes a useful
337 metric to quantify the variability while removing the effects of outliers. For more robust evaluation and statistics, only grid cells
338 with more than 100 CBH retrievals over the year are displayed thus impacting mostly coastal and polar regions. The spatial
339 distribution of the mean cloud base (Fig. 5, top) is similar to the outlined global distributions from other studies using different
340 instruments and methods (Böhm et al., 2019; Lu et al., 2021; Mülmenstädt et al., 2018). It is to be noted that the illustrated global
341 quantities were established using MODIS overpasses which happen at a practically constant local time (13:30 h, early afternoon
342 for AQUA). The MAD pattern exhibits similar characteristics (Fig. 5, bottom), even though variability slightly increases in the
343 vicinity of land masses. These interpretations still remain valid when looking at relative deviations. Typical features are lower
344 cloud bases towards polar regions and the mid-latitudes, and higher ones in the tropical regions. One can further observe regions
345 like the pacific coast of South America or the Namibian coast which display lower cloud bases concurrently with lower
346 variability (also highlighted in Lu et al. (2021)). It is however impossible to follow up the study for nighttime retrievals, as some
347 MODIS cloud properties are not retrieved then.

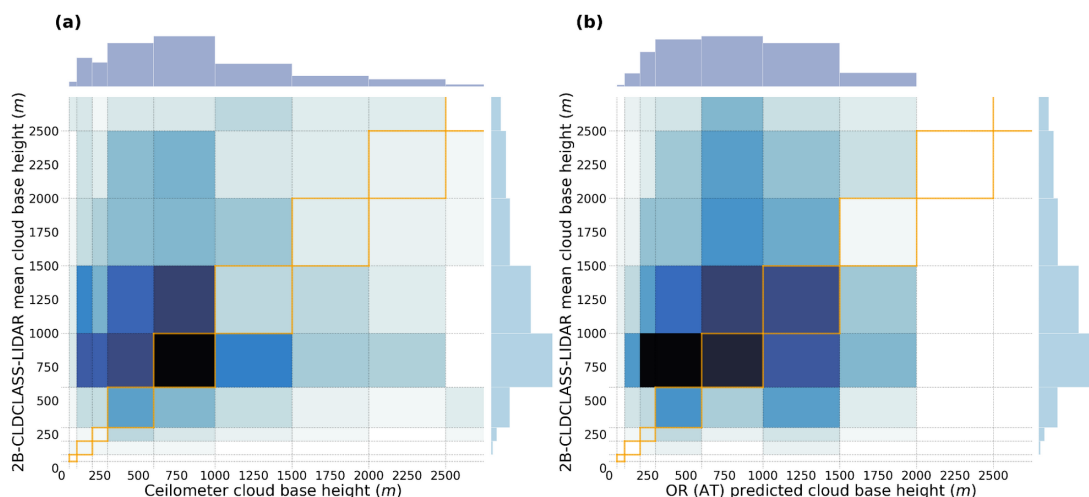


Figure 4: Joint histogram of (a) surface observations and 2B-CLDCLASS-LIDAR retrievals, and (b) ML-model predictions and 2B-CLDCLASS-LIDAR retrievals, for the co-located cloud scenes during the year 2008. The 1:1 boxes are highlighted in the figure in orange.

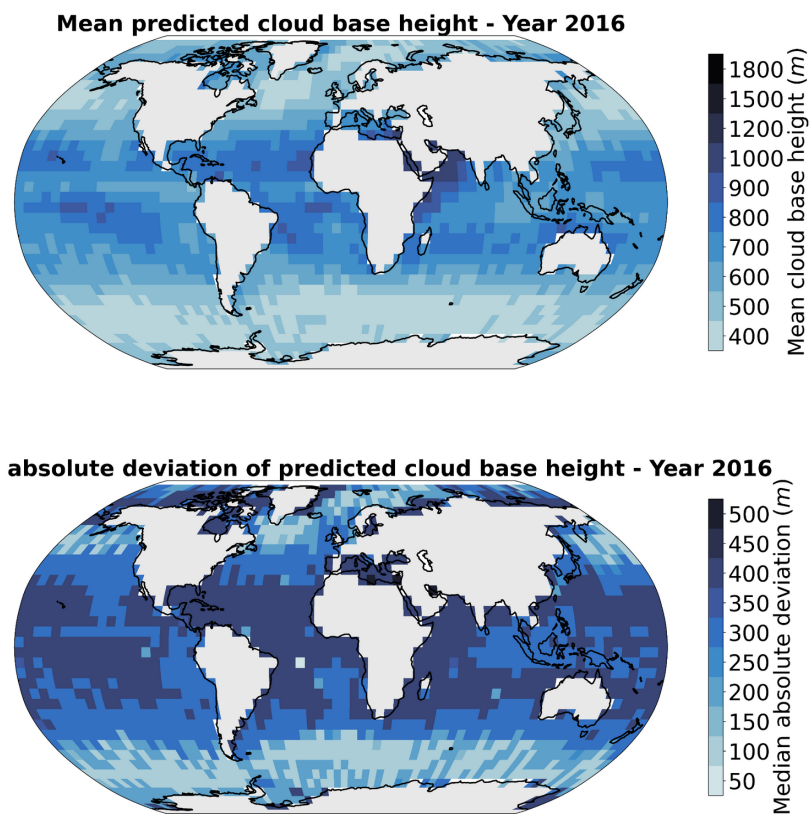
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5 Conclusion

355 We have presented here a novel method which retrieves the cloud scene base height over marine areas from MODIS cloud
356 properties, specifically CTH, COT and CWP. This method can produce robust CBH estimates for cloud scenes in particular for
357 lower cloud bases (MAE of 379 m and absolute error standard deviation of 328 m for up to 2 km cloud bases), based on the
358 assumption of a homogeneous cloud base across the considered cloud field. The statistical model was built on surface
359 observations of cloud bases with ceilometers (section 2.1), and then evaluated in comparison to other methods using passive
360 satellite instruments (section 3.1) and active satellite retrievals (section 3.2). Analysis of the yearly averaged CBH (section 4)
361 helped to further make sense of the predicted cloud bases and variability. The global dataset for the year 2016 is available from
362 Zenodo (Lenhardt et al., 2024).

363 Using the spatially-resolved information of cloud fields with passive satellites allows to properly quantify lower cloud bases,
364 more specifically avoiding the noisy retrievals of active satellites closer to the surface. A CNN proves to be valuable to leverage
365 spatial information without making any assumption with respect to how the cloud quantities are related to the CBH. The OR
366 modelisation helps bridging the gap between regression and classification, facilitating the use of the binned cloud base
367 observations provided by the surface observation dataset. Overall, our prediction model achieves low error in the retrievals,
368 around 400 m, and concurrently a narrow absolute error distribution, more precisely around 400 m absolute error standard
369 deviation. Both of these performance metrics are additionally reduced when focusing on cloud bases lower than 2 km.
370 Application to data over land areas has not been processed yet but would certainly require adding surface observations from land
371 during the training process (e.g. Böhm et al., 2019; Lu et al., 2021; Mülmenstädt et al., 2018).

372 The main benefit of producing better cloud base estimates is to gain accuracy in the overall retrieval of cloud geometry,
373 impacting in particular radiation estimates (Kato et al., 2011) like the surface downwelling longwave radiation (Mülmenstädt et
374 al., 2018). Our method can thus prove to be useful by helping to produce CBH with enhanced confidence at a global scale.



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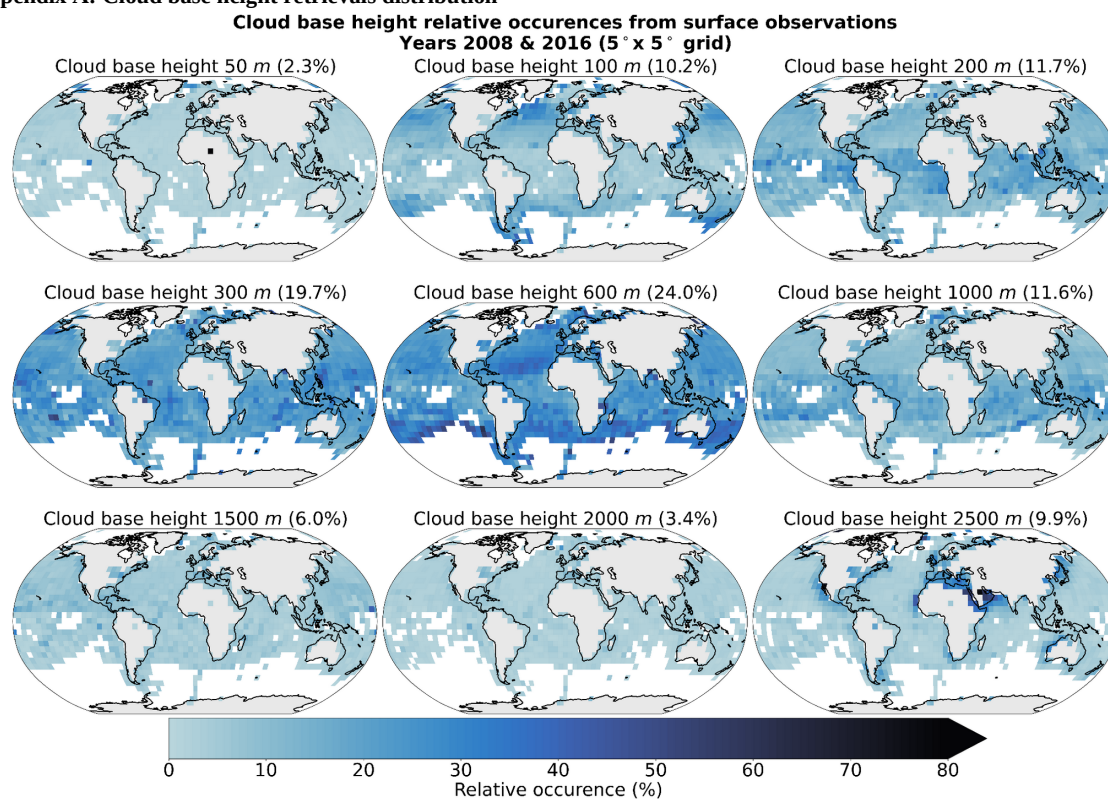
Figure 5: Spatial distribution of (top) mean and (bottom) median absolute deviation of predicted cloud base height for the MODIS data of the year 2016 aggregated on a 5 ° grid.



379 **Appendix**

380

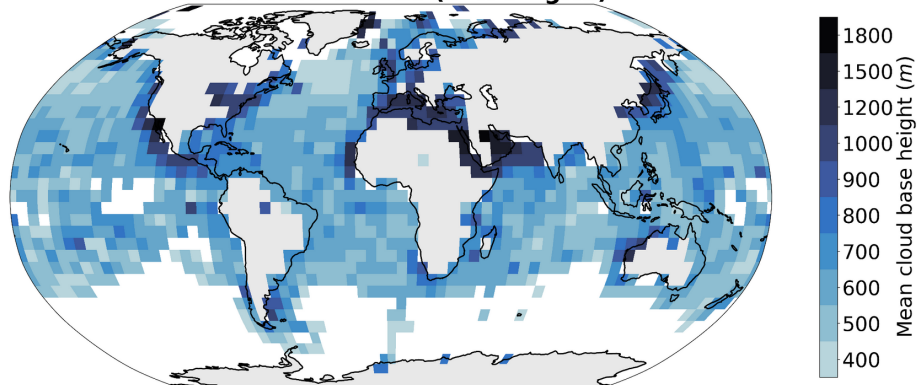
381 **Appendix A: Cloud base height retrievals distribution**



382

383 **Figure A.1: Spatial distribution of cloud base height retrievals (Met Office, 2006) for the years 2008 and 2016 on a 5°**
384 **grid. Overall percentage of each label in the total observations is indicated in brackets. Only grid cells with more than 50**
385 **retrievals are displayed.**

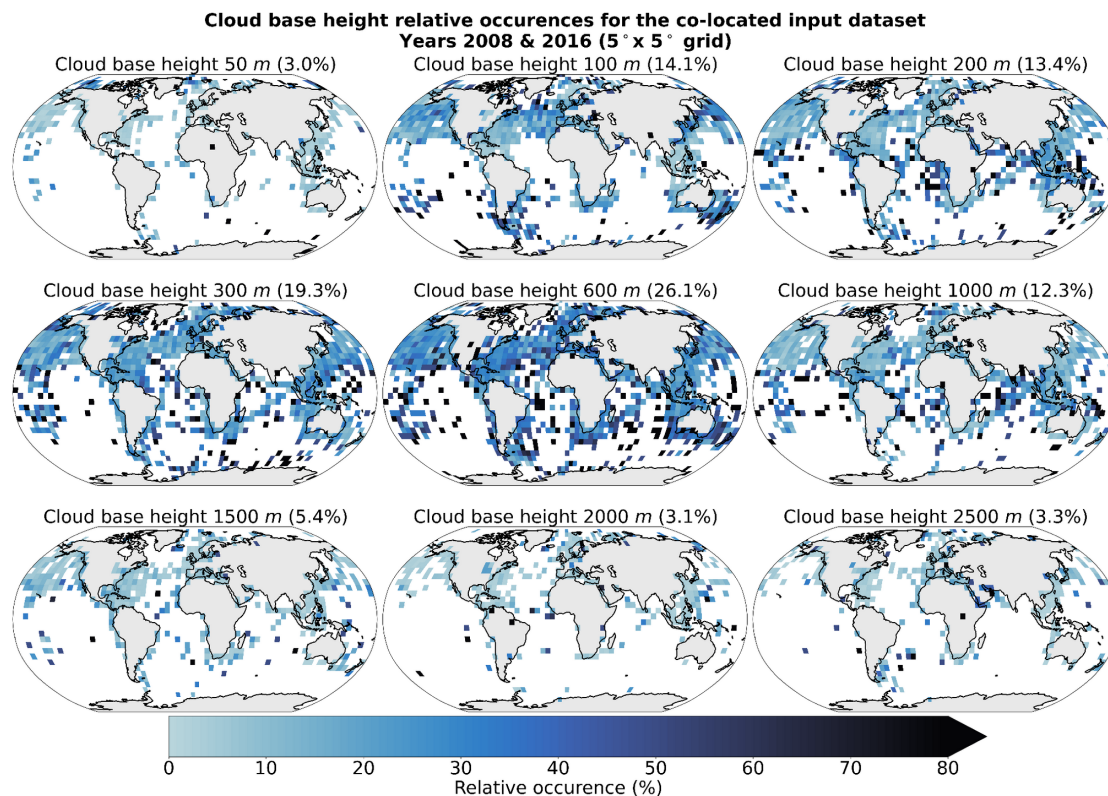
Mean cloud base height from surface observations
Years 2008 & 2016 (5° x 5° grid)



386

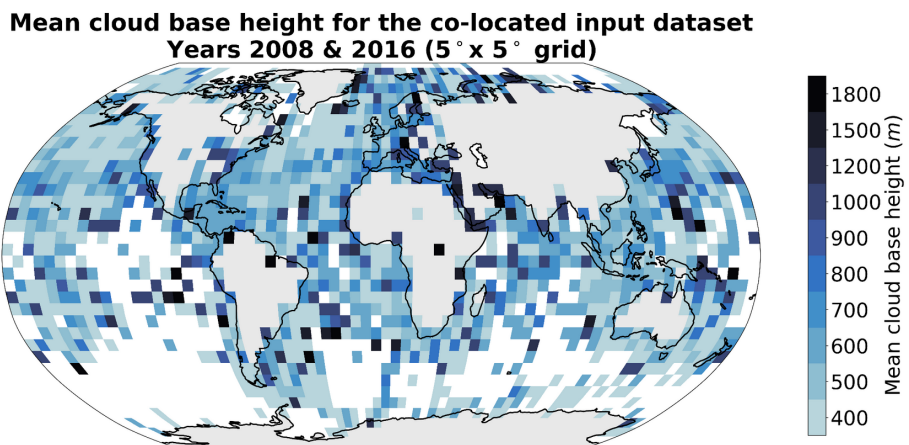
387 **Figure A.2: Mean cloud base height from retrievals (Met Office, 2006) for the years 2008 and 2016 on a 5° grid. Only**
388 **grid cells with more than 50 retrievals are displayed.**

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Figure A.3: Spatial distribution of the co-located cloud base height retrievals (Met Office, 2006) and the satellite cloud properties used for training the prediction model for the years 2008 and 2016 on a 5° grid. Overall percentage of each label in the total dataset is indicated in brackets.



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Figure A.4: Mean cloud base height from the co-located retrievals (Met Office, 2006) and the satellite cloud properties used for training the prediction model for the years 2008 and 2016 on a 5° grid.



399 **Appendix B: Spatio-temporal correlation study**

400

401 We create five different datasets to evaluate how the chosen AE architecture is capable of generalising to new data while trying
 402 to remove some possible autocorrelation biases which might inflate the performance scores. We also use this study to analyse
 403 how the AE model behaves when trained with our input data. We define two splits for space and time in order to build the
 404 training and testing datasets, namely the South-western (SW) quadrant and the period from March to October, respectively. The
 405 swaths used to build the datasets span across the whole year of 2016. The *random* data split is the basis for the training of the
 406 model and consists of tiles sampled in the aforementioned quadrant and time period. These tiles are then split randomly between
 407 training, validation and testing datasets. This split represents the common way of splitting data when building a ML model. In
 408 contrast, we build 3 other datasets which vary through their respective spatial and time spans. The *spatial* split is built
 409 considering tiles spanning across a distinct time period, here between November and February, regardless of their spatial
 410 location. The *temporal* split is built considering tiles located anywhere but in the South-western quadrant regardless of the time
 411 at which the retrieval occurred. Finally the *spatio-temporal* split combines the previous two conditions in order to build a dataset
 412 in which the tiles come from an independent location and time as the ones used for training. Additionally, we create a global data
 413 split using data from a different year, here 2008, without any spatial restriction for the tiles. Furthermore, only a limited number
 414 of tiles was extracted from each swath file while only swaths from non-consecutive days were used in order to limit possible
 415 correlation between the extracted scenes.

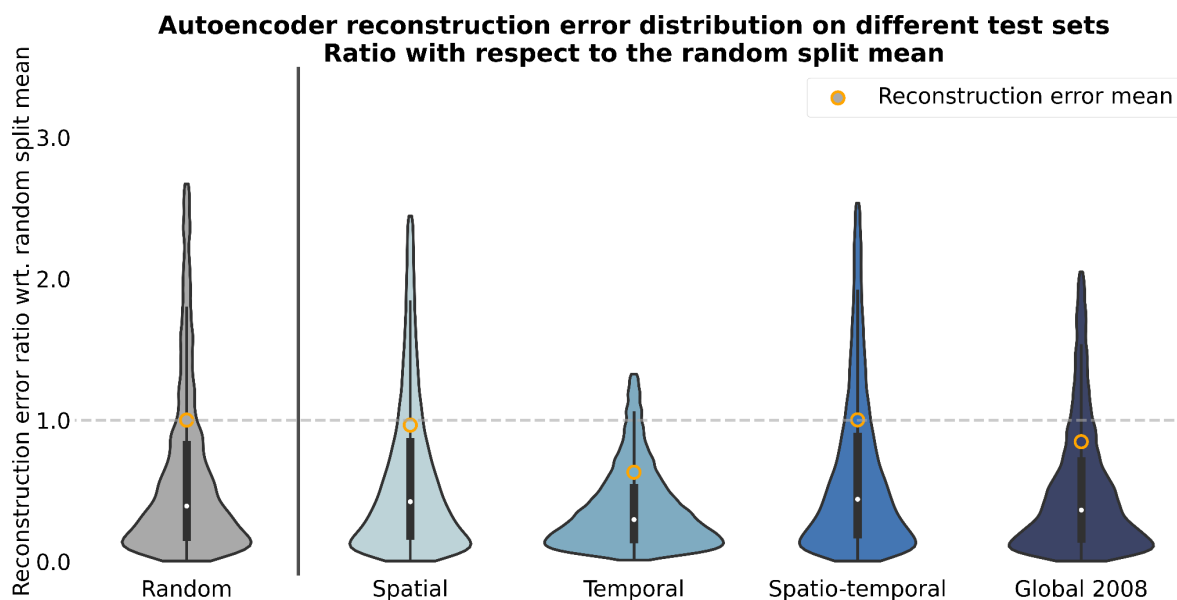
Data split	Time period	Spatial extent	
Random	03-10.2016	SW quadrant	Train: 14 691 Validation: 4 198 Test: 2 099
Spatial	03-10.2016	Global except SW quadrant	107 736
Temporal	01-02 and 11-12.2016	SW quadrant	12 420
Spatio-temporal	01-02 and 11-12.2016	Global except SW quadrant	30 659
Global	12.2008	Global	7 111

416 **Table B.1 : Name, time period, spatial extent and number of samples for each of the five described data splits.**

417

418 We then train an AE model using the training data from the first data split (*random*). Each test data split is then used to evaluate
 419 the trained model through the reconstruction error (Fig. B.1). Spatial distribution of the mean reconstruction errors is shown in
 420 [Figure B.2](#).

421 We first notice that the reconstruction power of the model is consistent regardless of the test split considered with mean
 422 reconstruction error ratios ranging from 0.63 to 1.0, dividing the split's reconstruction error by the random data split mean
 423 reconstruction error. Ratios around 1 or below indicate that the model's performance is not inflated when considering a random
 424 data split, highlighting that the model did not only learn from possible spatial and/or temporal correlations between samples
 425 present in the training set. The distribution of the error is also very similar throughout the test splits with most of the samples
 426 located below an error ratio of 0.5. However, one of the main aspects regarding the performance of the model across test splits is
 427 the presence of a heavy tail in the distribution showcasing that for some samples the reconstruction error can be greater than 3
 428 times the mean error. Looking at the spatial patterns of the reconstruction error, we note that overall the error comes from the
 429 COT and CWP predictions, the average reconstruction errors across test sets being 0.15, 0.32 and 0.25 for CTH, COT and CWP
 430 respectively ([Table B.2](#)). For the CTH, the error is concentrated in the zones with frequent convection around the equator and
 431 could be explained by local convection cells exhibiting a larger spread in CTH values. Another source of error could be that
 432 higher CTH values are also less represented in the training data. On the contrary, the error for COT and CWP is prevailing in
 433 high-latitude regions. Overall, the performance skill of the AE model seems to hold through the different test data splits. One
 434 could argue that the training dataset already retains enough variability in the data which could explain why the model still
 435 performs well regardless of the test set split. However, this consistent skill also shows that the performance reported in appendix
 436 C on the test set can be trusted to hold for other datasets and supports the data generation process to train the AE (cf. section 2.4).



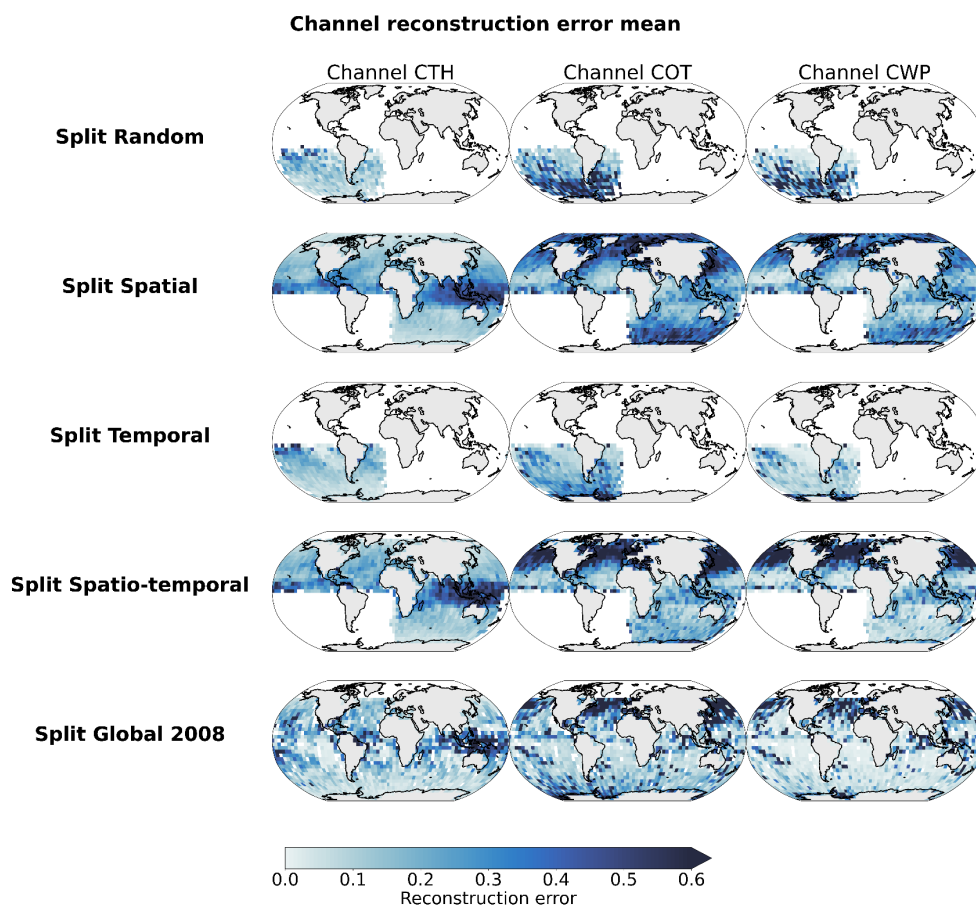
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Figure B.1: Reconstruction error ratios of an AE on different test datasets. The quartiles are indicated with the barplot inside each violin plot while the mean is indicated with an orange circle. Extreme values were removed before plotting. Each sample's reconstruction error is divided by the mean reconstruction error of the random data split.

Data split	Channel			Average
	CTH	COT	CWP	
Random	0.117	0.369	0.333	0.273
Spatial	0.171	0.344	0.276	0.263
Temporal	0.114	0.253	0.150	0.172
Spatio-temporal	0.202	0.332	0.286	0.274
Global	0.154	0.318	0.221	0.231
Average	0.152	0.323	0.253	0.243

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 443

Table B.2 : Average channel reconstruction relative error for each of the five described data splits.



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Figure B.2: Distribution of mean channel reconstruction errors aggregated on a 5 ° grid.



446 **Appendix C: Autoencoder architecture, training and performance**

447

448 The two components of the AE model, namely the encoder and the decoder, consist of five convolution blocks. Each block is
 449 then made of three convolution operators followed by LeakyReLU activation functions (Maas et al., 2013). After the last
 450 convolution of each block, batch normalisation is added to help convergence (Ioffe et al., 2015) followed by a maximum pooling
 451 layer. We then add linear layers to enforce the desired dimension of the latent space. The decoder architecture follows the same
 452 principles with transposed convolution layers (Zeiler et al., 2010) replacing the pooling layers of the encoder. This is summarised
 453 in [Table C.1](#). Details about the training of the AE are included in [Table C.2](#) and the loss history during training is shown in
 454 [Figure C.1](#).

Layer	Hyperparameters	Output shape
Input		(None, 3, 128, 128)
Encoder		
Conv2d	(kernel = 3, stride = 2)	(None, 3, 64, 64)
ConvBlock x 5	Conv2d (kernel = 3, stride = 1) LeakyReLU Conv2d (kernel = 3, stride = 1) LeakyReLU Conv2d (kernel = 3, stride = 1) BatchNorm2d LeakyReLU MaxPool2d (kernel = 2, stride = 2)	(None, 256, 2, 2)
Flatten + Linear		(None, 256)
Decoder		
Linear + Unflatten		(None, 256, 2, 2)
ConvTranspose2d	(kernel = 2, stride = 2)	(None, 256, 4, 4)
ConvTransposeBlock x 5	Conv2d (kernel = 3, stride = 1) LeakyReLU Conv2d (kernel = 3, stride = 1) LeakyReLU Conv2d (kernel = 3, stride = 1) BatchNorm2d LeakyReLU ConvTranspose2d (kernel = 2, stride = 2)	(None, 3, 128, 128)

455

Table C.1 : Autoencoder model specifications.



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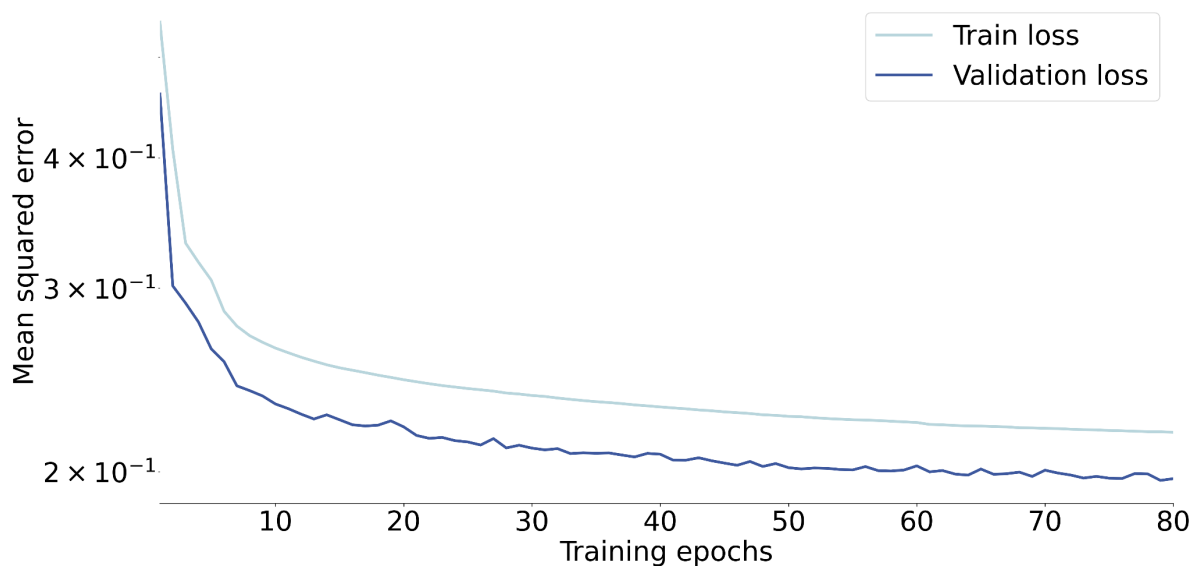
Hyperparameter	Value
Batch size	64
Epochs	80
Optimizer	Stochastic Gradient Descent (SGD), momentum = 0.9, learning rate = 0.0001
Metric	MSE
Early stopping	patience = 20

457

Table C.2 : Autoencoder model training specifications.

458

Autoencoder train/validation loss



459

Figure C.1 : Training and validation losses during model optimization.

460



461 **Appendix D: Ordinal regression**

462

463 We define our labels y which can take values in $K=9$ classes from $\{50 \text{ m}, 100 \text{ m}, \dots, 2500 \text{ m}\}$. We introduce $K-1$ thresholds
 464 α_y to define the separation of our K classes which actually correspond here to the classes too. For each labelled sample (s, y)
 465 the output of our model is $z=z(s)$. The correct interval for this this sample is then (α_{y-1}, α_y) . We consider a generic
 466 nonnegative penalisation function $f(\cdot)$ (eg. hinge loss, squared error loss, Huber loss). There are then different ways to represent
 467 threshold violations and thus to penalise the predictor. While immediate-threshold setup only considers the thresholds of the
 468 correct interval, all-threshold setup takes into account all the threshold violations. In the case of an immediate-threshold setup the
 469 loss function would look like:

470
$$\mathcal{L}(z, y) = f(z - \alpha_{y-1}) + f(\alpha_y - z).$$

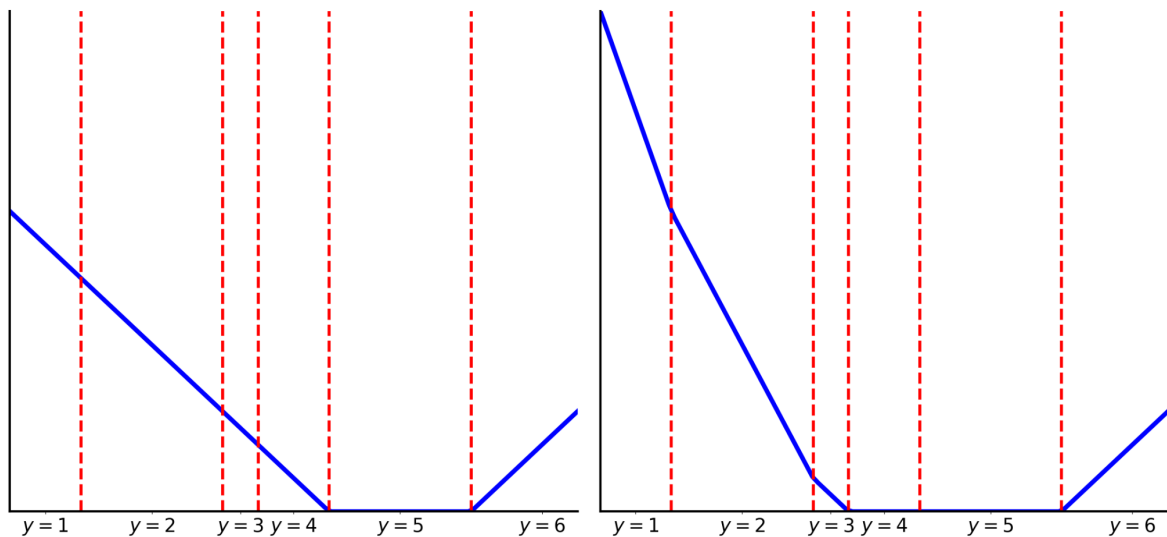
471 Here we can see that the loss is not aware of how many thresholds are actually violated. In the case of an all-threshold setup the
 472 loss function is a sum of violations across all thresholds:

473
$$\mathcal{L}(z, y) = \sum_{i=1}^{K-1} f(t(i, y)(\alpha_y - z)),$$

474 where $t(i, y) = -1$ if $i < y$ or $+1$ if $i \geq y$. Thus predictions are encouraged to violate the least amount of thresholds.

475 We give in [Figure D.1](#) an example of what the loss function would look like in the case of $K=6$ labels and using a hinge
 476 penalisation.

477



478

479 **Figure D.1: Threshold-based setups loss function representation for a hinge penalisation, $K=6$ labels and target label $y=5$.**

480

(left) Immediate-threshold and (right) All-threshold setup loss function. (figure adapted from Rennie et al. (2005))



481 **Appendix E: Cloud base height retrieval method assuming adiabatic cloud**

482

483 Algorithm adapted from Goren et al. (2018). We use the retrieved CTH, CTT, CTP and CWP from MODIS MYD06 (Platnick et
484 al., 2017).

485

Algorithm: Cloud base height retrieval

Data: CTH, CTT, CTP, LWP, look-up tables

Result: CBH

if CTT < 263.13 **then**

return NaN

$T \leftarrow CTT - 273.13$

 LWP obs \leftarrow LWP

 LWP adi \leftarrow 0.

$\delta z \leftarrow$ 0.

 Set corresponding cloud top indexes for temperature T_{ind} and pressure p_{ind} look-up tables.

 Read-in the water mixing ratio w at the corresponding indexes.

if w out of look-up table **then**

return NaN

while LWP adi < LWP obs **then**

$\rho_{tmp} \leftarrow$ density look-up table with T_{ind} and p_{ind}

$\delta_{tmp} \leftarrow$ layer depth look-up table with T_{ind} and p_{ind}

$\delta z \leftarrow \delta z + \delta_{tmp}$

$w_{tmp} \leftarrow$ mixing ratio look-up table with T_{ind} and p_{ind}

 LWP adi \leftarrow LWP adi + $(w_{tmp} - w) \times \delta z_{tmp} \times \rho_{tmp}$

 Adjust temperature T given the saturated lapse rate using look-up table with T_{ind} and p_{ind}

 Update indexes T_{ind} and p_{ind}

return CTH - δz

486

487 **Table E.1: Pseudo code for cloud base height retrieval algorithm assuming adiabatic cloud, adapted from Goren et al.**

488

(2018).



489 **Code availability**

490

491 The code used for the method and producing the plots is available on Zenodo (Lenhardt et al., 2024).

492 **Data availability**

493

494 The global dataset of the cloud base height predictions for the year 2016 is available on Zenodo (Lenhardt et al., 2024). The
495 dataset is available as a csv file with corresponding coordinates, MODIS swath file, time of retrieval and predicted cloud base
496 height or in a netCDF file as daily aggregates on a regular grid with a resolution of 1° or 5°. The meteorological observations
497 from the UK MetOffice (Met Office, 2006) are available through the CEDA archive at
498 <https://catalogue.ceda.ac.uk/uuid/77910bcec71c820d4c92f40d3ed3f249>. The files from the CUMULO dataset (Zantedeschi et
499 al., 2019) are available at <https://www.dropbox.com/sh/i3s9q2v2jjyk2it/AACxXnXfMF5wuIqLXqH4NJOra?dl=0>.

500 **Author contribution**

501

502 JL, JQ and DS designed the study. JL wrote the code. JL conducted the analysis and JL, JQ, DS interpreted the results. JL
503 prepared the manuscript, JQ and DS reviewed the manuscript and provided comments.

504 **Competing interests**

505

506 The authors declare that they have no conflict of interest.

507 **Acknowledgements**

508

509 This work was supported by the European Union's Horizon 2020 research and innovation programme under Marie Skłodowska-
510 Curie grant agreement No. 860100 (iMIRACLI). We thank the Leipzig University Scientific Computing cluster for computing
511 and data hosting. We further thank Tom Goren for providing access to code snippets from Goren et al. (2018) and thank Olivia
512 Linke for helping review the manuscript. We acknowledge the contributors of the CUMULO dataset (Zantedeschi et al., 2019)
513 for providing access to the data files hosted at
514 <https://www.dropbox.com/sh/i3s9q2v2jjyk2it/AACxXnXfMF5wuIqLXqH4NJOra?dl=0>. Additionally, we acknowledge the
515 MODIS L2 Cloud product data set from the Level-1 and Atmosphere Archive and Distribution System (LAADS) Distributed
516 Active Archive Center (DAAC), located in the Goddard Space Flight Center in Greenbelt, Maryland
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