The Joint Effect of Mid-latitude Winds and the Westerly Quasi-Biennial Oscillation Phase on the Antarctic Stratospheric Polar Vortex and Ozone

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Abstract. The quasi-biennial oscillation (QBO) dynamically interacts with the extratropical atmosphere. However, the relationship between the QBO in austral winter and the Antarctic stratospheric polar vortex in spring remains unclear. Here, we proposed a joint predictor involving the QBO for the Antarctic polar vortex and ozone in austral spring. During the westerly QBO phase (WQBO), positive anomalies in the zonal-mean zonal wind at 20°S−40°S in the upper stratosphere in July, named as the extratropical positive mode, can lead to a stronger Antarctic stratospheric polar vortex and lower ozone concentration in November, with correlations reaching 0.75 and 0.60. The mechanism is summarized as follows: the positive extratropical mode triggers a secondary circulation, which further alters the environmental condition for wave propagation in the stratosphere, pushing the positive anomalous zonal-mean zonal wind towards the pole. While during the easterly QBO phase (EQBO), the correlation of the extratropical mode and the strength of polar vortex is only 0.1. Due to stronger upward motions in the tropics, which opposes the secondary circulation caused by the extratropical mode, the EQBO cannot sustain the positive anomalous zonal-mean zonal wind until November. Our results highlight that the extratropical mode during WQBO could serve as a reliable predictor of the Antarctic stratospheric polar vortex and Antarctic ozone hole with a five-month time lag.

1 Introduction

The quasi-biennial oscillation (QBO) is a dominant mode of interannual variability in the tropical stratosphere. It is well known to the alternating westerly and easterly zonal wind that periodically descend from the tropical upper stratosphere to the tropical tropopause (Lindzen and Holton, 1968; Andrews and McIntyre, 1976; Baldwin et al., 2001), as well as its influence of the atmospheric circulation and chemical species outside the tropical stratosphere (Holton and Tan, 1980; Ruti et al., 2006; Garfinkel and Hartmann, 2011; Gray et al., 2018; Rao et al., 2019; 2020a). Holton and Tan (1980) proposed that QBO can modify upward-propagating planetary waves by altering the zero-wind line in the stratosphere, which further affects the extratropical waveguide (Baldwin et al., 2001; Anstey et al., 2014; Zhang et al., 2019). Consequently, the Arctic stratospheric polar vortex and the Brewer-Dobson circulation (hereafter referred to as B-D circulation) tend to be weaker on average during the easterly QBO phase (EQBO) compared to the westerly QBO phase (WQBO). Additionally, Garfinkel et al. (2012) proposed that the secondary meridional circulation induced by the QBO also plays a crucial role in the Arctic stratosphere. This QBO-induced changes in the Arctic stratospheric polar vortex can further influence the distribution of Arctic ozone (Wang et al.,

2022). Zhang et al. (2021) revealed that dynamical processes contribute more to the Arctic ozone-QBO connection than chemical processes. Furthermore, observational and modeling evidence both show that the QBO's influence on the Arctic stratospheric polar vertex and ozone occurs within the extended winter (November–March; Rao et al., 2019; 2020b; Zhang et al., 2021).

In the Southern Hemisphere (SH), upward-propagating planetary waves are weak due to the less thermal contrast between land and sea. Consequently, the QBO-vortex coupling, which is closely related to planetary waves, has received less attention (Garcia and Solomon, 1987; Lait et al., 1989; Baldwin and Dunkerton, 1998; Naito, 2002; Hitchman and Huesmann, 2009). Naito et al. (2002) examined the QBO signal in the SH, finding that the zonal-mean zonal wind in the lower stratosphere decelerates more rapidly from September to October during the EQBO than during the WQBO. This deceleration is attributed to the stronger upward wave propagation from the troposphere and larger wave convergence in the easterly phase. Additionally, Anstey et al. (2014) demonstrated that the extratropical response to the QBO in late-winter SH can be interpreted as a modulation of the final warming by the QBO. On one hand, the responses in the SH stratospheric polar vortex appears to be much smaller than that in the NH, and it remains unclear whether the high-latitude signatures of the QBO are robust in the SH. On the other hand, the timing of the Antarctic stratospheric polar vortex's response to the QBO signal remains unclear. In addition to the Antarctic stratospheric polar vortex, Wang et al. (2022) examine the QBO signals in stratospheric ozone. They found that Antarctica is the only region where total column ozone (TCO) does not significantly respond to the QBO signal (Figure 2 in Wang et al., 2022). So far, the impacts of the QBO on the Antarctic stratospheric polar vortex and ozone have not been well documented.

The QBO of which period varies irregularly in the range from 17 to 38 months is considered as a reliable predictor of the stratospheric polar vortex, and further the near-surface climate and weather (Baldwin and Dunkerton, 2001; Zhang et al., 2020; Tian et al., 2023). In the SH, if the QBO-vortex coupling does exist, it could also serve as a predictor for the Antarctic stratospheric polar vortex and ozone. In this study, we aim to explore when and how the QBO influences the southern stratospheric polar vortex. Understanding this connection between the QBO and high-latitude circulation may help improve the forecasting accuracy of the southern stratospheric polar vortex, and, by extension, the Antarctic ozone hole. Section 2 presents the data and methods used in this research. In Section 3, we propose an improved predictor for the Antarctic stratospheric polar vortex and ozone based on the QBO. We also discuss the underlying mechanisms. The conclusions are provided in section 4.

2 Data and methods

2.1 Data

Monthly and daily meteorological data and ozone mass mixing ratio are derived from Modern-Era Retrospective analysis for Research and Applications, Version 2 (MERRA-2; GMAO, 2015) reanalysis for the period from 1980 to 2022. The reanalysis data has a resolution of 1.25°×1.25° and 42 pressure levels extending from 1000 to 0.1 hPa.

Previous studies have employed different methods to define the QBO index. Some are based on the tropical zonal-mean
zonal wind at a single pressure level (Holton and Tan, 1980; Gray et al., 1992; Baldwin et al., 2001; Garfinkel and Hartmann,
2007), while others use two QBO indices on different pressure levels (Andrews et al., 2019) or use empirical orthogonal
function (EOF) analysis applied on the tropical zonal-mean zonal wind (Randel et al., 1999; Anstey et al., 2010; Rao and Ren,
2018), which can better capture QBO's vertical structure. The QBO phase defined by the EOF method is similar to that defined
by the single pressure level QBO index (Baldwin et al., 2001; Rao et al., 2020b). Therefore, this study uses the standardized
zonal-mean zonal wind averaged over 10°S−10°N at 20 hPa to define the QBO phase (Baldwin et al., 2001). The EQBO phase
is defined as years when the tropical standardized zonal-mean zonal wind is less than −1, while the WQBO corresponds to
years when the tropical standardized zonal-mean zonal wind is greater than 1. The El Nino-Southern Oscillation (ENSO) index
is defined as the sea surface temperature (SST) anomalies averaged over the Niño 3.4 region (5°N−5°S; 170°−120°W). SST
is derived from the Monthly NOAA Extended Reconstruction SSTs Version 5 (ERSSTv5).

2.2 Method

2.2.1 Singular value decomposition (SVD)

The SVD analysis is performed between the zonal-mean zonal wind at latitudes ranging from 0° to 40°S in July (extratropical
mode) and the zonal-mean zonal wind at latitudes ranging from 50° to 70°S in November (polar mode). To validate the
robustness of the findings, we conduct a Monte Carlo test (Iwasaka and Wallace, 1995) by creating 1000 ensembles of the
SVD analysis. In this test, the tropical wind field is kept fixed, while the polar wind is randomly disordered in both time and
space. The total variance of the 1000 paired ensembles is calculated and sorted. If the observed SVD total variance exceeds
the 95th percentile of the ensemble distribution, the SVD mode is considered significant at the 95% confidence level.

2.2.2 E-P flux and stream function

The E-P flux (Andrews et al., 1987) is used to diagnose the propagation of waves, which is calculated as follows:

$$F_\phi \equiv \rho_0 a \cos\varphi \left(\frac{\overline{u_z}\,\overline{v'\theta'}}{\overline{\theta_z}} - \overline{u'v'} \right) \tag{1}$$

$$F_z \equiv \rho_0 \cos\varphi \left\{ \left[f - (a\cos\varphi)^{-1} \left(\overline{u}\cos\varphi \right)_\varphi \right] \frac{\overline{v'\theta'}}{\overline{\theta_z}} - \overline{w'u'} \right\} \tag{2}$$

$$\nabla \cdot F \equiv (a\cos\varphi)^{-1} \frac{\partial}{\partial\varphi} \left(F^{(\varphi)} \cos\varphi \right) + \frac{\partial F^{(z)}}{\partial z} \tag{3}$$

Where ρ_0 is the density; z is the altitude; a is the radius of the Earth; φ is the latitude; f is the Coriolis parameter; θ is the
potential temperature; u and v are the zonal and meridional winds, respectively; w is the vertical velocity. The overbars
represent the zonal average, and primes represent deviation from the zonal average.

Mass stream function $\bar{\chi}^*$ associated with the residual circulation is defined as:

$$\frac{\partial \bar{\chi}^*}{\partial \varphi} = \rho_0 a \cos \varphi \bar{w}^* \tag{4}$$

Where $\bar{w}^*$ is represented as:

$$\bar{w}^* = \bar{w} + (a \cos \varphi)^{-1} (\frac{\cos \varphi \overline{v'\theta'}}{\bar{\theta}_z})_\varphi \tag{5}$$

In the SH, the climatological mean stream function is negative.

2.2.3 The refraction index

We use the method developed by Harnik and Lindzen (2001) to divided the traditional refraction index (n_{ref}^2) into vertical (m^2) and meridional components (l^2) by solving the quasi-geostrophic equations in the spherical coordinates:

$$\frac{\partial^2 \psi}{\partial z^2} + \frac{N^2}{f^2} \frac{\partial^2 \psi}{\partial y^2} + n_{ref}^2 \psi = 0 \tag{6}$$

Where ψ is the weighted wave geopotential height, N^2 is the Brunt-Väisälä frequency, and f is the Coriolis parameter. By solving Eq. (6), the vertical and meridional components of the refraction index can be expressed as:

$$\text{Re}(\frac{\psi_{zz}}{\psi}) = -m^2 \text{ and } \text{Re}(\frac{\psi_{yy}}{\psi}) = -l^2 \tag{7}$$

And regions with a large refraction index facilitate wave-propagation.

2.2.4 The Lorenz energy cycle

The Lorenz energy cycle is used here to measure the atmospheric circulation changes. As described by Lorenz (1967) and Holton (1968), the diabatic heating generates mean available potential energy (PM), which is described as (Hu et al., 2004):

$$PM = \frac{c_p}{2} \int \gamma [<T>]^{''2} dm \tag{8}$$

$$\gamma = -\{\frac{\theta}{T} [(\frac{R}{c_p})P]\} \tag{9}$$

And then the baroclinic eddies transport warm air poleward, cold air equatorward, and transform the PM to eddy available potential energy (PE). At the same time, the PE is transformed into eddy kinetic energy (KE) by the vertical motions of the eddies, and the KE can be described as:

$$KE = \frac{1}{2} \int [<u^2> + <v^2>] dm + \frac{1}{2} \int [<u>^{*2} + <v>^{*2}] dm \tag{10}$$

The zonal kinetic energy (KM), defined as the integration of $\overline{u}^2$, can be maintained primarily by the conversions from KE due to the correlation $\overline{u'v'}$ (Eq. 11). Additionally, it can also be converted into the PM (Eq. 12).

$$C(KE,\ KM) = \int [<u'v'>+<u>^*<v>^*]\cos\varphi\{\frac{\partial([<u>]/\cos\varphi)}{a\partial\varphi}\}dm$$

$$+ \int [<v'^2>+<v>^{*2}]\frac{\partial[<v>]}{a\partial\varphi}dm$$

$$+ \int [<u'w'>+<u>^*<w>^*]\frac{\partial[<u>]}{\partial p}dm \tag{11}$$

$$+ \int [<u'w'>+<u>^*<w>^*]\frac{\partial[<v>]}{\partial p}dm$$

$$- \int [<v>][<u'^2>+<u>^{*2}]\frac{[\tan\varphi]}{a}dm$$

$$C(KM,\ PM)=\int [<v>]g\frac{\partial[<z>]}{a\partial\varphi}dm \tag{12}$$

Where u and v are the zonal and meridional winds, respectively, w is the vertical wind speed, T is the temperature, a is the radius of the earth, c_p is the specific heat of air at constant pressure, φ is the latitude, p is pressure, R is the gas constant for dry air, and g is the gravitational acceleration. The $<>$ and [ ] are represent the time and zonal average. The * and ' denote the departure from zonal mean and time average. The " represents departure from meridional average. *C(A, B)* represents the conversion from A to B.

2.3 Model simulation

In this study, we use the Whole Atmosphere Community Climate Model (WACCM) in the Community Earth System Model version 2 (CESM2) to investigate the connection between the QBO signal and the Antarctic polar vortex state. The model has a horizontal resolution of 1.9° × 2.5° and a hybrid vertical coordinate with 70 levels from the surface to approximately 140 km (Gettelman et al., 2019). To avoid the impact of initial fields and stochastic low-frequency climate fluctuations on the experiment results (Zhang et al., 2024), 20 ensemble experiments were conducted using the same annual cycle of external forcing in 2000 (e.g., sea ice concentration, sea surface temperature, stratospheric ozone and other chemistry species, greenhouse gas, aerosols, solar radiation) but with different initial fields. The nudging technology uses a dynamical-core-independent scheme (Davis et al., 2022).

$$F = -W\frac{(X-X_{ref})}{\tau} \tag{13}$$

Where X_{ref} is the meteorology value from reanalysis datasets at the next update step, τ is the relaxation timescale (6 hours), and W is the nudging coefficients from 0 to 1. Stratospheric zonal wind, meridional wind, and temperature from 1.2 to 103.3 hPa are nudged with a nudging coefficient of 1.0. This nudging coefficient decreases linearly to 0 from 103.3 hPa to 143.0

135 hPa. Tropical stratosphere (22°S to 22°N) is nudged to a coefficient of 1.0, with a transitional latitude of 22°N/S to 35°N/S. The stratospheric conditions mentioned above are nudged to the JRA-55 reanalysis, and each nudging experiment runs from 1980 to 2022. Note that different sets of reanalysis data are used to force the CESM and perform other analysis. The consistency of the results across both datasets indicates the robustness of the findings.

3 Results

3.1 A better predictor for the polar vortex and ozone

Figure 1a indicate the time-lag correlation between the QBO index and the zonal-mean zonal wind at 60°S and 70 hPa, representing the strength of the Antarctic stratospheric polar vortex. Anstey and Shepherd (2014) reviewed the impact of QBO on high latitudes, showing stronger westerlies during WQBO, though with weak statistical significance. Lecouffe et al. (2022) suggested that the QBO can modulate the polar vortex strength, characterized by a stronger vortex and longer vortex duration

145 during WQBO. Therefore, our analysis focuses on the positive correlations in Figure 1a, which represents a stronger polar vortex with westerly QBO phase. We find that the QBO in austral winter begins to influence the Antarctic stratospheric polar vortex. Additionally, the austral winter and spring QBO signals in the Antarctic stratospheric polar vortex can persist for several months, consistently reaching their maximum in November (e.g., July QBO at a 4-month time lag, August QBO at a 3-month lag and so on) and then gradually decline. Here, we focus on the earliest QBO signal in the Antarctic stratospheric

150 polar vortex to determine the longest possible prediction lead time. Note that there is a maximum positive correlation between July QBO index and polar vortex strength 4 months later, suggesting that the QBO signal in austral winter influence the Antarctic stratospheric polar vortex in austral spring. During austral late spring, the duration of the Antarctic ozone hole is highly correlated with the strength of the polar vortex. Consequently, the QBO in winter potentially serves as a predictor of the Antarctic polar vortex and ozone hole in austral spring. However, this QBO-vortex signal is not statistically significant.

155 Additionally, we also examine the response of the Antarctic stratospheric polar vortex in austral spring to the winter QBO (Fig. 1b). The WQBO in winter (QBO index in July greater than 1) does not consistently lead to a stronger zonal-mean zonal wind in polar regions in spring, and the correction between them is only 0.23. These results suggest that the direct impact of the QBO on the Antarctic polar vortex is weak or potentially non-existent.

To further explore the robust relationship between the winter QBO and the Antarctic stratospheric polar vortex in spring,

160 we first regressed the zonal-mean zonal wind in winter against the strength of the polar vortex in November during the WQBO. Interestingly, the winter zonal-mean zonal wind around 30°S in the upper stratosphere highly correlates with the Antarctic polar vortex in November (Figures 2a–c). Additionally, the winter mid-latitude wind also shows a high correlation with the Antarctic ozone in November (Figs. 2d–f). These results suggest that the extratropical zonal-mean zonal wind in the upper stratosphere (hereafter, extratropical mode), along with the WQBO signal in winter, might serve as a joint predictor of the

165 Antarctic polar vortex and ozone in austral spring.

Here, the SVD analysis is further used here to measure the relationship between the winter extratropical mode (Fig. 2a) and the Antarctic polar vortex in spring. Figures 3a and b show the first paired mode of zonal mean zonal wind in the extratropic and polar, respectively. The extratropical mode is characterized by a positive zonal wind anomaly centered around 30°S, which is closely resembles the regression patterns in Fig. 2. In the tropical regions, it displays a weak WQBO signal at 20 hPa (right panel of Fig. 3a). In austral spring polar region (Fig. 3b), it features a positive zonal wind south of 50°S indicating a strong polar vortex. The first paired mode explains 98.2% of the total variance and is significant at the 95% confidence level according to the Monte Carlo test. Since the first paired mode represents the dominant coupled variance, our analysis mainly focuses on the first SVD paired mode. The correlation coefficient between the time series of the two modes is 0.75, which is also significant at the 95% significance level (Fig. 3c). Note that almost all years represented by circles fall into the first and third quadrants in Fig. 3c, suggesting that when the extratropical mode is in its positive (negative) phase, a stronger (weaker) polar vortex will occur five months later. It is worth noting that the blue and yellow circles representing the phase and strength of ENSO show no consistent pattern with the strength of polar vortex in November, suggesting that the strong correlation between the two modes appears to have little connection with ENSO. Moreover, the Antarctic stratospheric ozone in austral spring, which is closely related to the stratospheric polar vortex, also exhibits a high correlation of 0.6 with the extratropical mode in July (Fig. 3c). Thus, we can conclude that the winter extratropical mode, in conjunction with the WQBO, is closely linked to the spring Antarctic polar vortex and ozone.
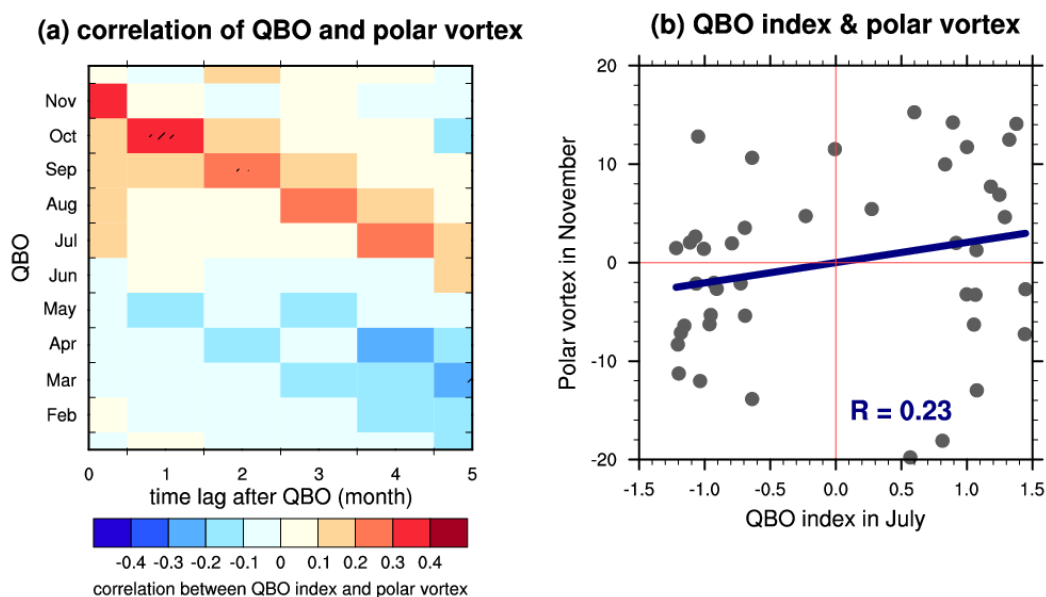
Figure 1: (a) Correlations between the zonal-mean zonal wind at 60ºS and 70 hPa and the QBO index at different time lag. The Y-axis represents the month in which using the QBO index is used for the correlation analysis. The X-axis indicates the time lags. The shadings indicate that the correlations are statistically significant at 90% confidence level. (b) Monthly mean zonal-mean zonal wind at 60ºS and 70 hPa in November (with a 4-month lag after the QBO in July) plotted against the July QBO index according to MERRA-2 reanalysis dataset from 1980 to 2022. The solid blue line represents the linear regression of the QBO index and the strength of Antarctic stratospheric polar vortex, with their correlation coefficient shown in the bottom right-hand corner.
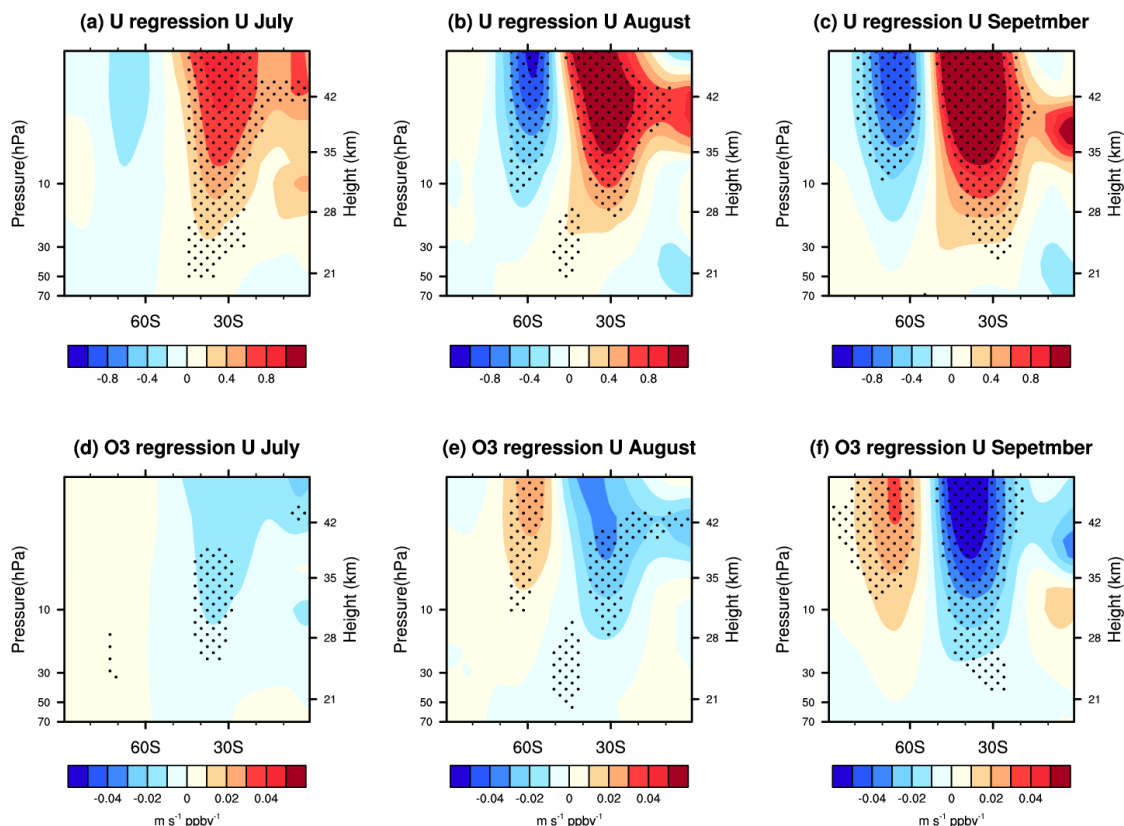
Figure 2. (a)−(c) Regression patterns of the zonal-mean zonal wind in (a) June, (b) July, and (c) August against the zonal-mean zonal wind at 60ºS and 70 hPa in November. (d)−(f) Regression patterns of the zonal-mean zonal wind in (d) June, (e) July, and (f) August against the ozone concentration in the south of 60ºS and 70 hPa in November. The shadings indicate that the regression coefficients are statistically significant at 95% confidence level.

3.2 The underlying mechanisms responsible for the relationship between the extratropical mode and the polar vortex.

Note that the winter extratropical mode appears early in June and gradually develops (Fig. 2). In July, the extratropical mode is fully formed, and the correlation between the winter extratropical mode and the polar vortex in spring reaches its peak (0.74 in June, 0.75 in July, and 0.73 in August). Therefore, we primarily focus on the relationship between the extratropical mode in July and the Antarctic stratospheric polar vortex in November. As mentioned above, when the extratropical mode in July is in its positive phase (i.e., when the extratropic mode is greater than 0), there is a stronger polar vortex five months later, and vice versa (Fig. 3c). Thus, first we categorize the WQBO into WQBO-strong polar vortex (W-SPV; 1980, 1990, 2006, 2008, 2015, 2022) and WQBO-weak polar vortex (W-WPV; 1985, 1997, 2004, 2013, 2016) based on the phase of the extratropical mode in July. The W-SPV corresponds to a positive phase of the extratropical mode in July, while W-WPV corresponds to a negative phase (Fig. 3c). A composite analysis is then conducted to compare these two categories.

Figures 4a−e show the difference in composite monthly mean zonal-mean zonal wind between W-SPV and W-WPV from July to November. In the upper stratosphere, a positive center of zonal wind anomalies is located at 20°S−40°S, while a negative center is located at higher latitudes in July. From July to November, the positive zonal-mean zonal wind anomalies exhibit a downward and poleward movement, resulting in a strong polar vortex in austral spring. The negative anomalies also shift towards higher latitudes and gradually weaken, eventually being replaced by positive anomalies in November. The evolution of zonal-mean zonal wind anomalies confirms the connection between the extratropical mode in winter and the polar vortex in November. We further revealed that the positive anomalies in zonal-mean zonal wind can be reasonably explained by the E-P flux divergence anomalies and their poleward shift from July to November. The composite differences in E-P flux and its divergence between W-SPV and W-WPV are shown in Figs. 4f−j. In July and August, positive E-P flux divergence anomalies (Figs. 4f−g) are established between 30°S and 60°S in the upper stratosphere, corresponding to the weaker poleward and upward propagation of planetary waves in this region than normal. These E-P flux divergences lead to the poleward shift of positive zonal-mean zonal wind anomalies from July to August (Figs. 4a−b). Note that the E-P flux divergence anomalies shift to 60°S in the mid and lower stratosphere in September, with these positive anomalies persisting until November.
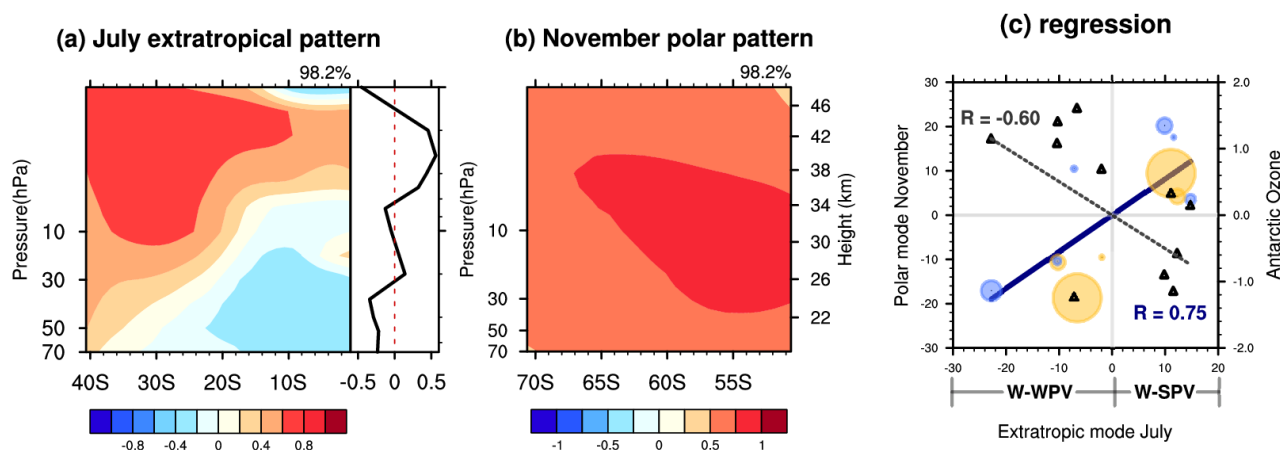
Figure 3. Spatial patterns for the first paired mode of the (a) monthly mean zonal-mean zonal wind over 0−40°S and 1−70 hPa in July, i.e., extratropical mode (b) zonal-mean zonal wind over 50°S−70°S and 1−70 hPa in November by the singular value decomposition (SVD) analysis during WQBO years, based on the MERRA-2 reanalysis dataset from 1980 to 2022. The right panel of (a) is the profile of the extratropical mode averaged over 0−5°S. The variance explained by the first mode is shown in the top right-hand corner. (c) The corresponding time series for the paired mode, with their correlation coefficient shown in the bottom right-hand corner (text in blue). The solid blue line represents the linear regression of the extratropical mode and polar mode. The size and color of the circle markers in panel (c) are proportional to the Niño 3.4 index, with yellow dots indicating a positive Niño 3.4 index and blue dots indicating a negative Niño 3.4 index. The standardized ozone mixing ratios, averaged over 60°S−90°S at 70 hPa in November, against the extratropical mode time series are shown with triangular marker (right Y-axis), with their correlation coefficient displayed in the top left-hand corner (text in black). The dashed black line represents the linear regression of the extratropical mode and ozone in November.
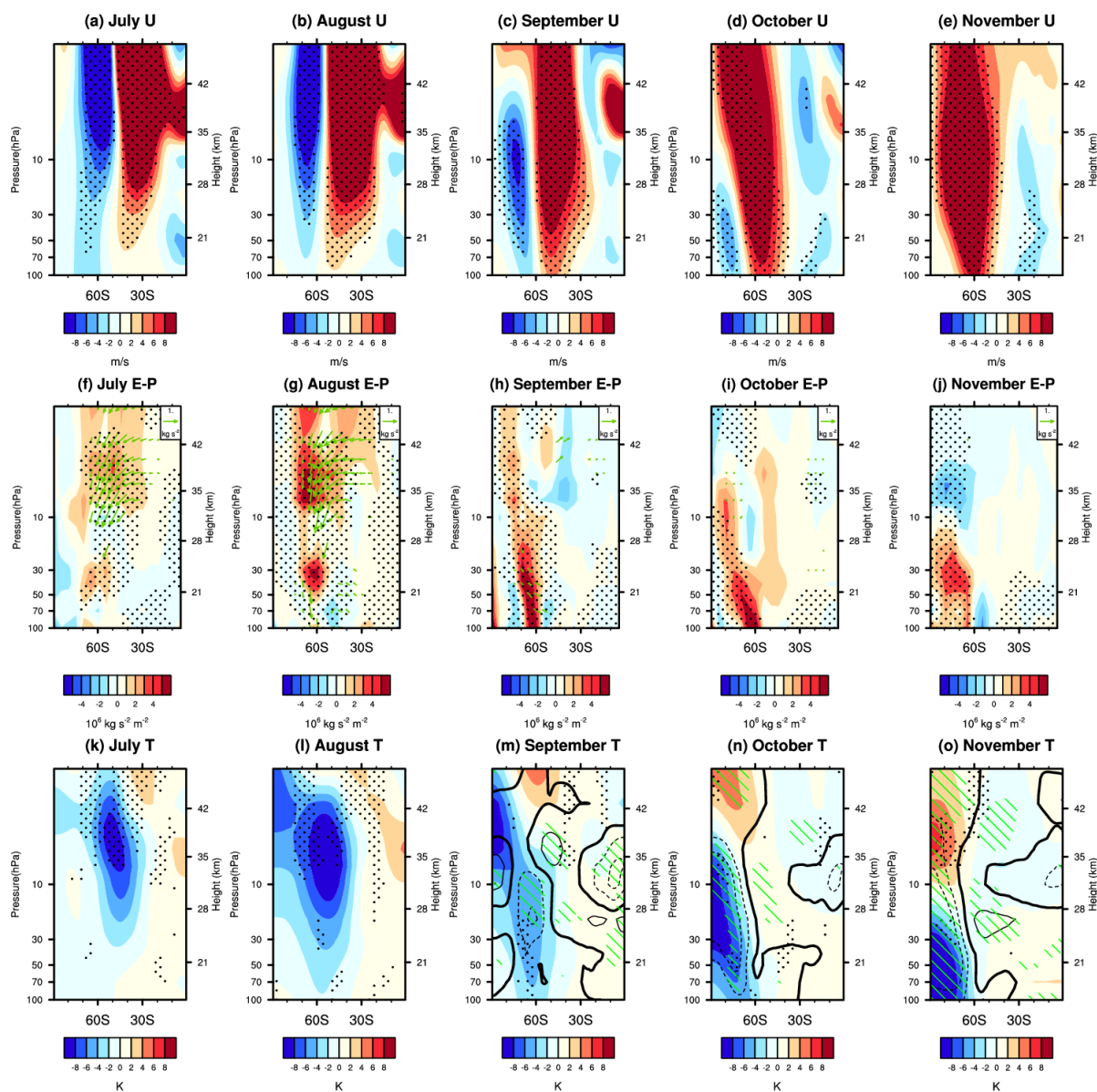
Figure 4. Composite differences in (a)−(e) the zonal-mean zonal wind anomalies, (f)−(j) the scaled E-P flux anomalies (green vectors; horizontal component unit: 107 kg s−2; vertical component unit: 105 kg s−2) and the E-P flux divergence anomalies (shadings), (k)−(o) the zonal-mean temperature anomalies (shadings) and the zonal-mean ozone volume mixing ratio anomalies (contours; dashed lines are negative, and thick lines are zero contours. The contour intervals are 200 ppbv) from July to November between the W-SPV and W-WPV according to MERRA-2 reanalysis dataset. The E-P flux and its divergence are calculated from wave 1 to 3. E-P flux vectors are scaled by the factor cosφ and multiplied by the square root of 1000.0/p in both the vertical and horizontal directions, where p is pressure in hPa. The dotted regions mark the differences in zonal-mean zonal wind, zonal temperature, and E-P flux divergence are statistically significant at the 90% confidence level. Green shading marks the regions where the differences in ozone volume mixing ratio between the W-SPV and W-WPV are statistically significant at the 90% confidence level. Only the significant (at the 90% confidence level) E-P flux vectors have been plotted in panels (f)−(j).

Additionally, negative anomalous temperatures are located south of the positive anomalous zonal wind center according

240 to the thermal wind balance. There is also a poleward and downward shift in temperature anomalies along with the anomalous

zonal wind from July to November (Figs. 4k−o). Finally, the lower Antarctic stratosphere is much colder in W-SPV than W-

WPV. In austral spring (September to November), temperature plays a crucial role in influencing the Antarctic stratospheric

ozone. Lower temperature in the lower stratosphere favours heterogeneous chemistry and results in increased chemical ozone

depletion in the polar regions. Consequently, the temperature decrease induced by the positive extratropical mode leads to

245 negative anomalous ozone in the Antarctic lower stratosphere from September to November (Figs. 4m−o).
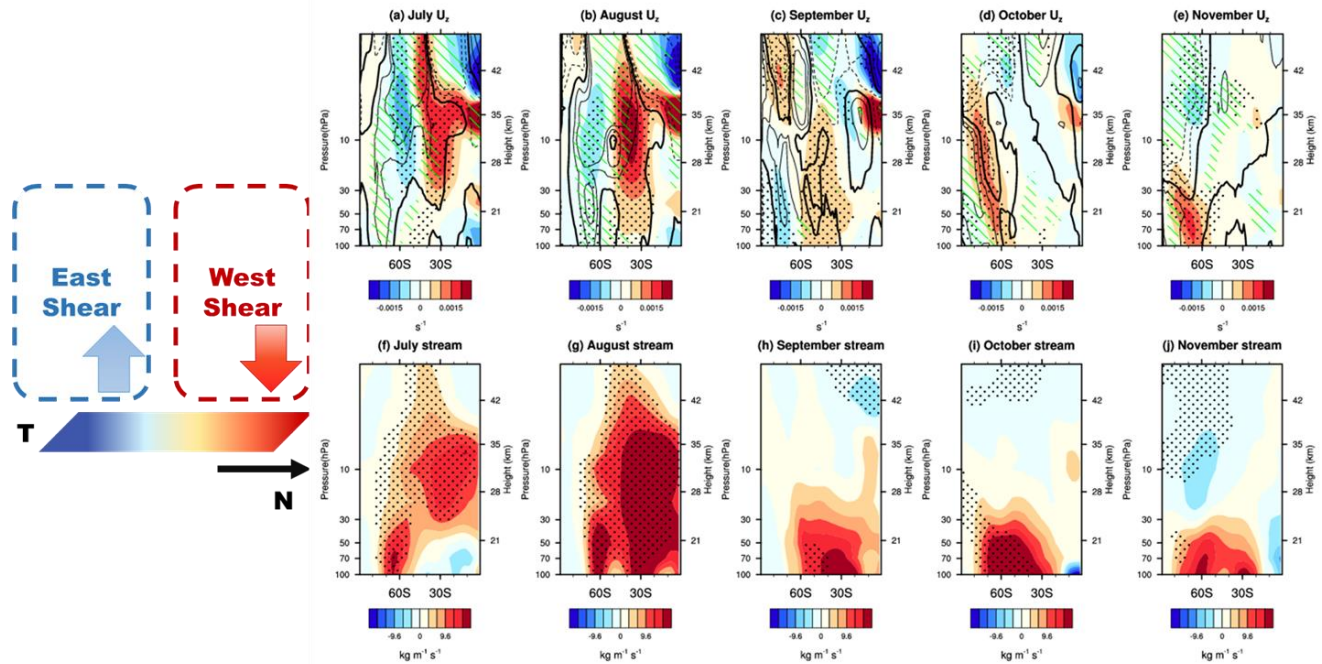
Figure 5. Composite differences in the (a)−(e) vertical gradient of zonal-mean zonal wind anomalies (shadings) and vertical component of residual mean (W*) circulation anomalies (contours; dashed lines are negative, and thick lines are zero contours. The contour intervals are 0.1 mm/s), and (f)−(j) stream function anomalies from July to November between the W-SPV and W-WPV according to MERRA-2 250 reanalysis dataset. The dotted regions mark the differences in (a)−(e) vertical gradient of zonal-mean zonal wind and (f)−(j) the stream function between the W-SPV and W-WPV are statistically significant at the 90% confidence level. Green shading marks the regions where the differences in w* between the W-SPV and W-WPV are statistically significant at the 90% confidence level. The schematic diagram on the left illustrates how the secondary circulation is triggered and sustained.

There are three main processes responsible for the extratropic-polar connection during the WQBO. The first process,

255 named as the thermal wind balance, is described as follows. The positive extratropical mode defined in Figure 3a exhibits a

positive zonal-mean zonal wind anomalies, as well as a positive vertical gradient around 30°S (Figure 5a). Meridional

temperature anomalies must exist to maintain thermal wind balance for the vertical shear of the zonal-mean zonal wind in the

basic state flow. These meridional temperature anomalies are sustained by the secondary vertical motion. Specifically, a

positive zonal-mean zonal wind shear anomaly (shadings in Fig. 5a), accompanied by a warm center in north of 30°S (Fig. 4k),

260 is maintained by the downward motion (contours in Fig. 5a), while a negative zonal-mean zonal wind shear anomaly in higher

latitudes accompanies a cold center and upward motion (Figs. 4k and 5a). Note that the climatological basic state meridional flow in austral spring is characterized by a negative stream function, with upward motion in the tropic and downward motion at higher latitudes (i.e., B-D circulation). In July and August, positive stream function anomalies appear at lower latitudes in the upper stratosphere, indicating clockwise secondary circulations opposite to the B-D circulation induced by the extratropical
265 mode during WQBO (Figs. 5f−g).
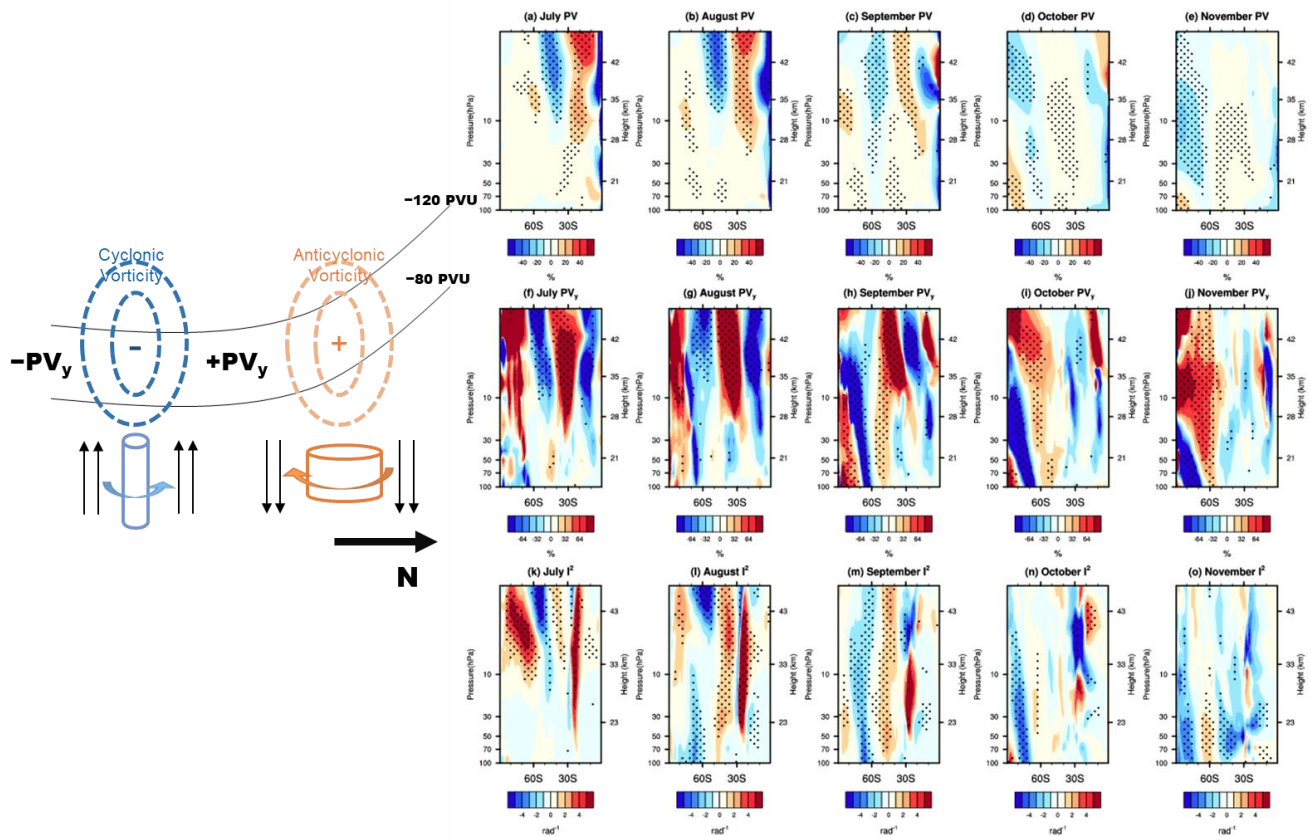
Figure 6. Composite differences in the (a)−(e) potential vorticity (PV) anomalies, (f)−(j) meridional gradient of PV anomalies, and (k)−(o) meridional component of refraction index anomalies from July to November between the W-SPV and W-WPV according to MERRA-2
270 reanalysis dataset. The dotted regions mark the differences in (a)−(e) PV, (f)−(j) meridional gradient of PV, and (k)−(o) meridional component of refraction index between the W-SPV and W-WPV are statistically significant at the 90% confidence level. The schematic diagram on the left illustrates the mechanism of the PV adjustment. The black lines indicate the climatological PV contours in the Southern Hemisphere.

The second process is the potential vorticity (PV) adjustment. The secondary circulation mentioned above causes vertical
275 compression and stretching of air columns in the lower and higher latitudes, respectively. As a results, the air column will acquire anticyclone vorticity (positive PV anomalies) in lower latitudes and cyclone vorticity (negative PV anomalies) in higher latitudes (schematic diagram of Figure 6). In July, positive stream function anomalies are centered around 30°S (Fig. 5f), resulting in positive and negative PV anomalies at 20°S and 40°S in the upper stratosphere, respectively (Fig. 6a). This

redistribution of PV leads to a negative meridional gradient of PV (PV_y) anomalous around 60ºS (Fig. 6f), which dominates
280   the wave refractive index and results in anomalous E-P flux divergence around 50ºS in the upper stratosphere in July (Fig. 4f).
In the following months, the positive center of the stream function gradually shifts toward the lower stratosphere and polar
regions (Figs. 5g−j). Consequently, the induced negative PV_y anomalies and wave refractive index also exhibit a poleward and
downward shift (Fig. 6). Especially from August to September, there is a notable transition in E-P flux divergence anomalies
from the upper stratosphere around 60ºS to the middle and lower stratosphere (30−100 hPa; Figs. 4g−h). At the same time,
285   both the anomalous centers of the stream function and PV_y also shift downward to the lower stratosphere (Figs. 5g−h, 6b−c,
6g−h, and 6l−m). This downward and poleward procession can be interpreted as wave-mean flow interaction, as described
below.
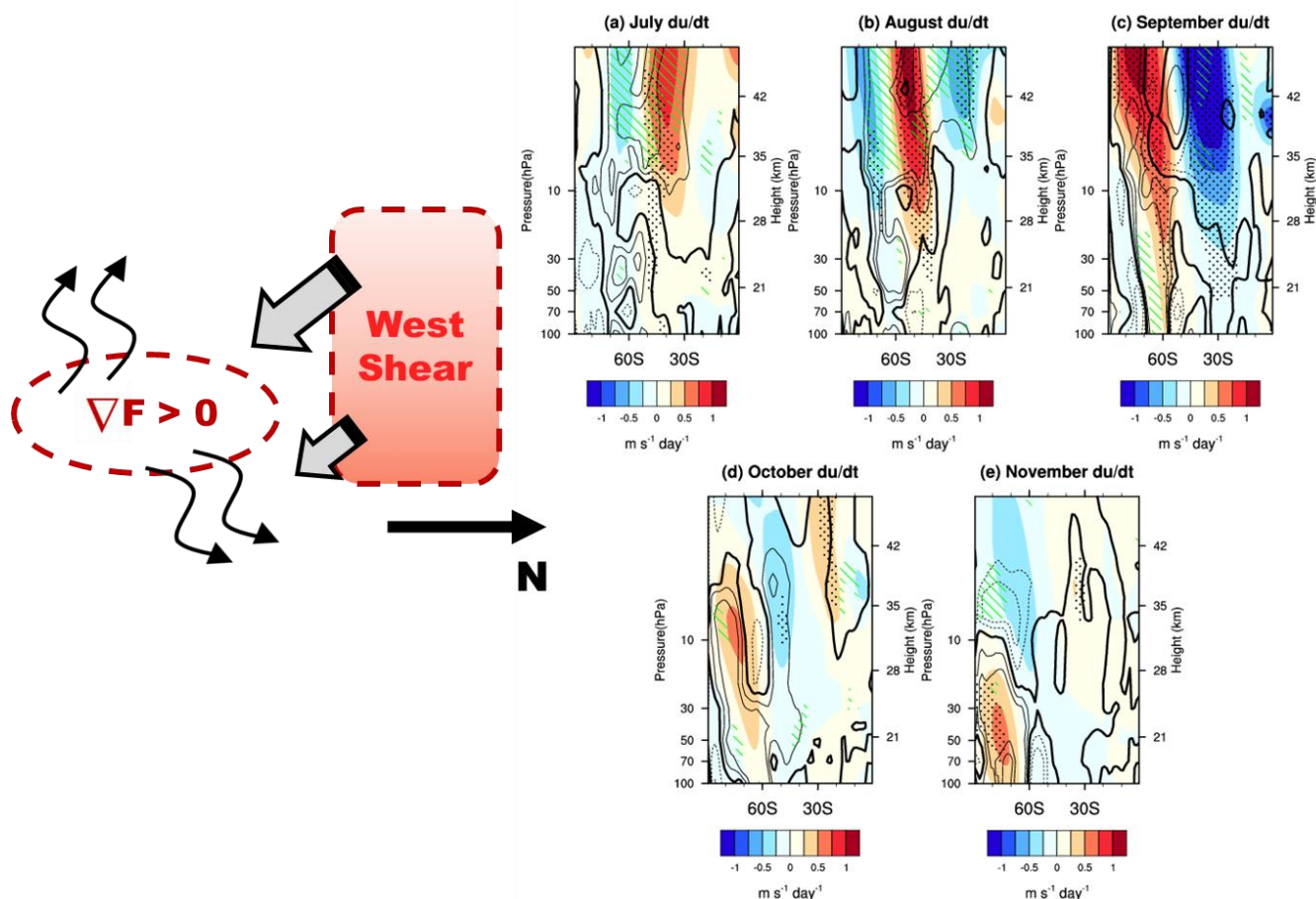
Figure 7. Composite differences in the du/dt anomalies (color shadings) and E-P flux divergence anomalies (contours; dashed lines are
290   negative, and thick lines are zero contours. The contour intervals are 5×10^5 kg s^{-2} m^{-2}) from July to November between the W-SPV and W-
WPV according to MERRA-2 reanalysis dataset. The dotted regions and green shadings mark the differences in du/dt and the differences in
E-P flux divergence between the W-SPV and W-WPV are statistically significant at the 90% confidence level. The schematic diagram
illustrates how these anomalies propagate towards the Antarctic lower stratosphere from July to November.

The wave-mean flow interaction sustains the anomalies mentioned above propagating towards the Antarctic lower
295   stratosphere from July to November (schematic diagram of Figure 7). Specifically, in July, the anomalous E-P flux divergence
centered around 50ºS in the upper stratosphere (Fig. 7a) induces a poleward and downward shift of the positive anomalous
zonal-mean zonal wind (positive anomalies of du/dt), along with other anomalous centers. In the following months, the positive
anomalous E-P flux divergence, as well as the positive du/dt anomalies, continue to move towards the lower stratosphere in
the polar region. By November, the center of positive zonal-mean zonal wind anomaly is located in the polar region in the
300   mid-stratosphere (Fig. 4e), indicating a stronger and colder polar vortex.
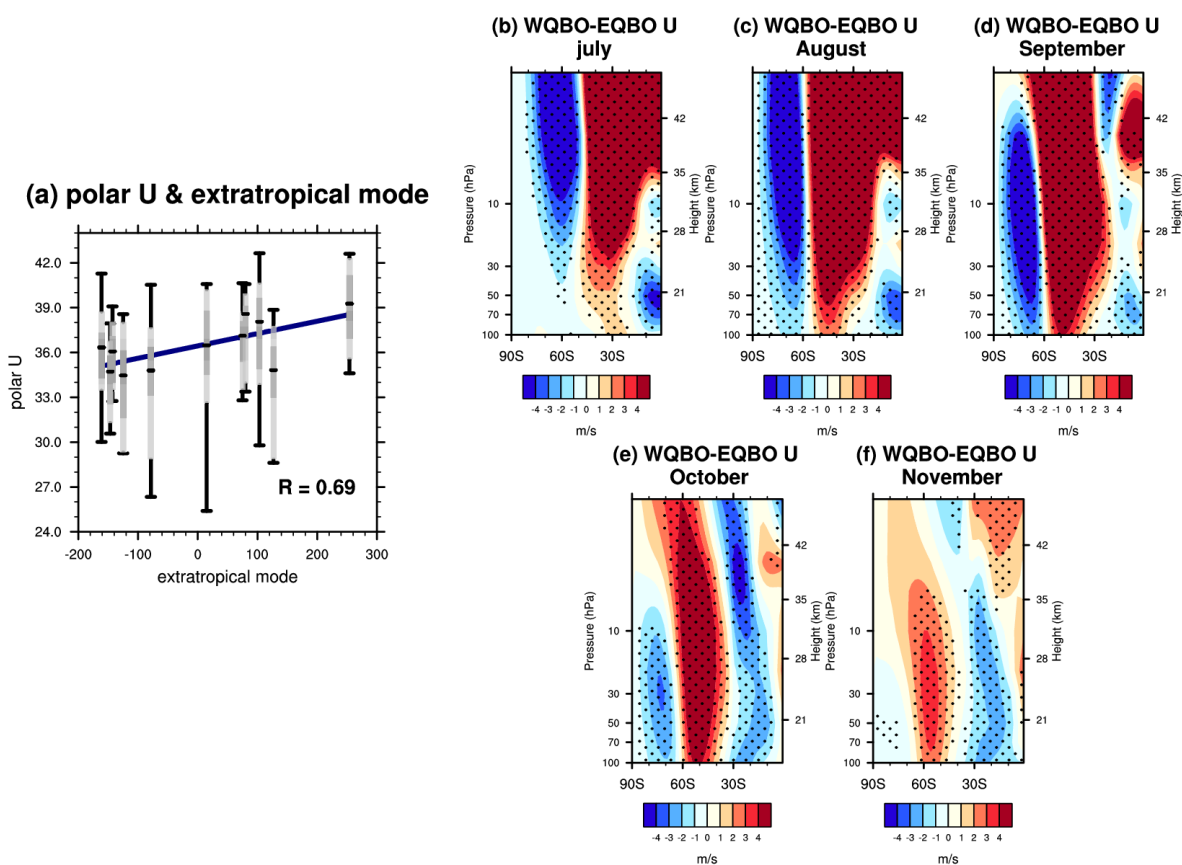
Figure 8. (a) The zonal-mean zonal wind at 60ºS and 100 hPa in November is plotted against the extratropical mode's PC in July, derived
305   from the 25 ensembles. The boxplot describes a summary of these ensembles. The light grey box spans from the lower decile to the upper
decile, and the dark grey spans the lower quartile to upper quartile. The lines inside the dark grey box marks the median zonal wind. The
lower and upper whiskers indicate the minimum and maximum zonal wind among the 25 ensembles. Additionally, the blue line represents
the linear fit between the extratropical mode and the median zonal wind, with their correlation coefficient displayed in the bottom right-hand
corner. (b−f) Composite differences in zonal-mean zonal wind from July to November between the W-SPV and W-WPV according to CESM
simulations. The dotted regions mark the differences between the W-SPV and W-WPV are statistically significant at the 95% confidence
310   level.

Based on the abovementioned analyses, we proposed a new predictor for the Antarctic stratospheric vortex in spring. During the WQBO, the correlation between the extratropical mode and polar vortex can reach 0.75 with a five-month time lag. To further verify these results, additional model experiments are conducted by the CESM2. First, the relation between the extratropical mode in July and the stratospheric Antarctic polar vortex in November has been validated. Figure 8a shows the

315   zonal wind at 60ºS in November as a function of the extratropical mode in July, derived from the 20 ensembles during the WQBO. Similar to the reanalysis dataset (Fig. 3c), the strength of the polar vortex increases with a larger extratropical mode, with a correlation reaching 0.69. We further display the composite results of the zonal-mean zonal wind (Figs. 8b−f). For each ensemble, the years in W-SPV and W-WPV are selected, and the difference in zonal-mean zonal wind between W-SPV and W-WPV is calculated across all ensembles. Initially, positive anomalies in zonal-mean zonal wind appear around 30ºS in the

320   upper stratosphere in July (Fig. 8b). These anomalies gradually shift to the lower stratosphere and polar regions due to wave-mean flow interactions (Figs. 8c−e), ultimately resulting in a stronger polar vortex in the Antarctic lower stratosphere by November (Fig. 8f). The continuous evolution of these positive anomalous zonal-mean zonal wind from the extratropical upper stratosphere to the Antarctic lower stratosphere indicates a robust relationship between the extratropical mode and the polar vortex during the WQBO.

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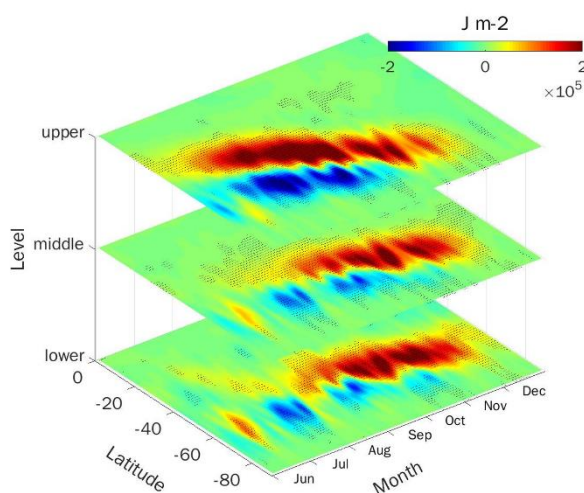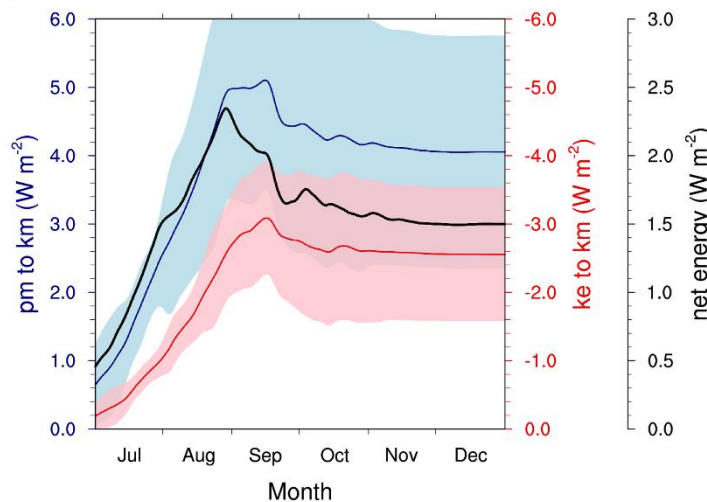
Figure 9. (a) Composite differences in mean kinetic energy (KM) from July to November between the W-SPV and W-WPV according to MERRA-2 reanalysis dataset. The KM at "upper", "middle", and "lower" levels are integrated 1−20 hPa, 30−50 hPa, and 70−100 hPa, respectively. The dotted regions mark the differences between the W-SPV and W-WPV are statistically significant at the 90% confidence

330   level. (b) Composite differences in time-integrated energy transformations averaged over 1−100 hPa and 40ºS−70ºS between the W-SPV and W-WPV. The blue and red shadings indicate significant intervals. If these shadings do not cross the X-axis, the difference is significant at the 90% confidence level. The blue line, corresponding to the left Y-axis, represents the energy transformed from the mean potential energy (PM) to KM. The red line, corresponding to the first right Y-axis, represents the energy transformed from the eddy kinetic energy (KE) to KM. The black line, corresponding to the second right Y-axis, represents the net energy transformed to KM.

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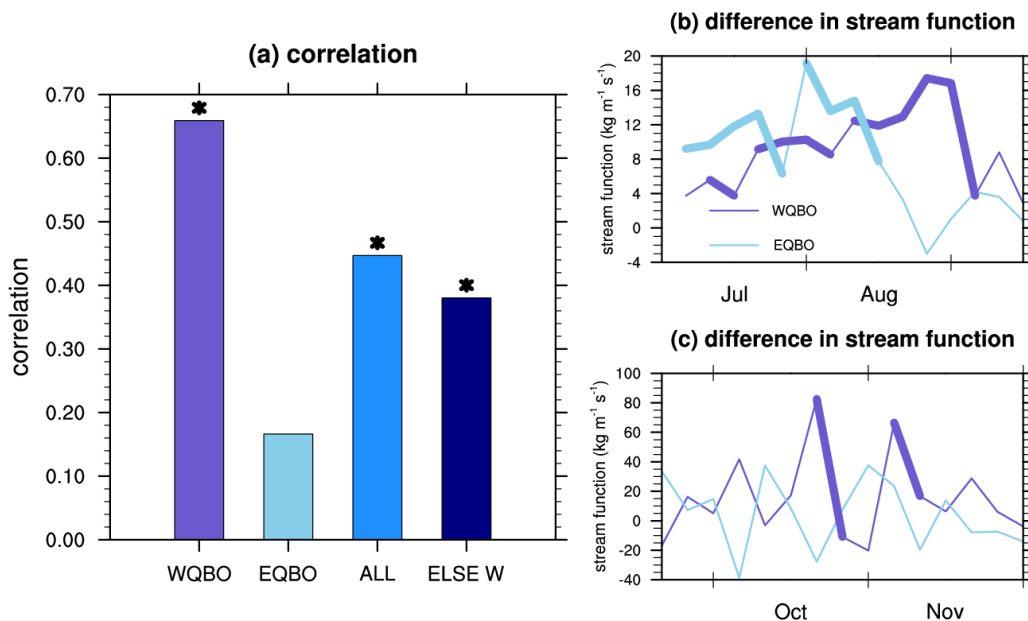
Figure 10. (a) The correlations between zonal-mean zonal wind in extratropical and polar regions under WQBO, EQBO, all years, and years except WQBO. The asterisks indicate that the correlations are statistically significant at the 95% confidence level. (b) Composite differences in stream function averaged over 0−60ºS and 1−30 hPa between the strong and weak polar vortex during the WQBO (purple lines) and EQBO (blue lines) from July to September. (c) Same as panel (b), but the stream function averaged over 30ºS −60ºS and 50−100 hPa from September to November.

A question arises as to why the abovementioned wave mean-flow interactions could persist for almost five months. We attempt to explain it from the perspective of the Lorenz energy cycle. Lorenz (1967) proposed that the solar energy is converted into the available potential energy, which then drives kinetic energy to maintain atmospheric circulation against both thermal and mechanical dissipations. As a result, the Lorenz energy cycle reflects changes in atmospheric circulation. Figure 9a displays the evolution of KM from June to December, which is proportional to the square of the zonal-mean zonal wind (u^2) and, therefore, effectively captures the transition of atmospheric circulation depicted in Figs. 4a−e. Positive anomalies in KM originated in the mid-latitudes in upper stratosphere at the end of June and then propagate poleward and downward until November. This evolution of KM suggests a continuous anomalous energy transfer from the mid latitudes in upper stratosphere to the polar regions, sustaining the positive zonal wind anomalies observed in W-SPV. Next, we explain why the positive anomalies in KM can persist for almost five months (Fig. 9b). Note that the KM is primarily maintained by conversion from KE (Eq. 11) and PM (Eq. 12). There is a continuous anomalous conversion from PM to KM since July (blue line, Y-axis in the left), which is corresponding to the secondary circulation in Figs. 5f−j. In July, the positive anomalous stream function around 30ºS at 20 hPa (Fig. 5f), indicating a clockwise secondary circulation, is associated with the uplift of relative cold air at higher latitudes (Fig. 5a). The relative cold air is subsequently replaced by warmer air from north of 30°S, lowing the center of mass and reducing potential energy. The following moths are similar to this process. Consequently, more potential energy from July to November is converted into kinetic energy in W-SPV (Fig. 9b). In addition, KM is also maintained by the

Note that this relationship only exists under WQBO conditions. The correlations between the extratropical mode in July and polar vortex in November can reach 0.65 during WQBO, but it is as low as 0.15 in EQBO (Figure 10a). This is due to the intrinsic difference in B-D circulation between WQBO and EQBO, driven by the variation in the vertical gradient of the zonal-mean zonal wind in the tropical upper stratosphere (Figure 11). Compared with EQBO, there are anomalous descending
365   motions around 15ºS at 10 hPa in WQBO, which facilitate the formation and maintenance of the secondary circulation (Figs. 5f–j). Therefore, during the WQBO, the secondary circulation, as well as other anomalous centers, propagate downward and poleward, especially from September to November. In contrast, during EQBO, these secondary circulations weaken and gradually dissipate. Figs. 10b and c illustrate the composite difference in the spatial integral stream function. Note that at the beginning of July, the anomalous secondary circulation is stronger in EQBO than in WQBO. However, the extratropic-polar
370   connection gradually disrupts from September in EQBO.
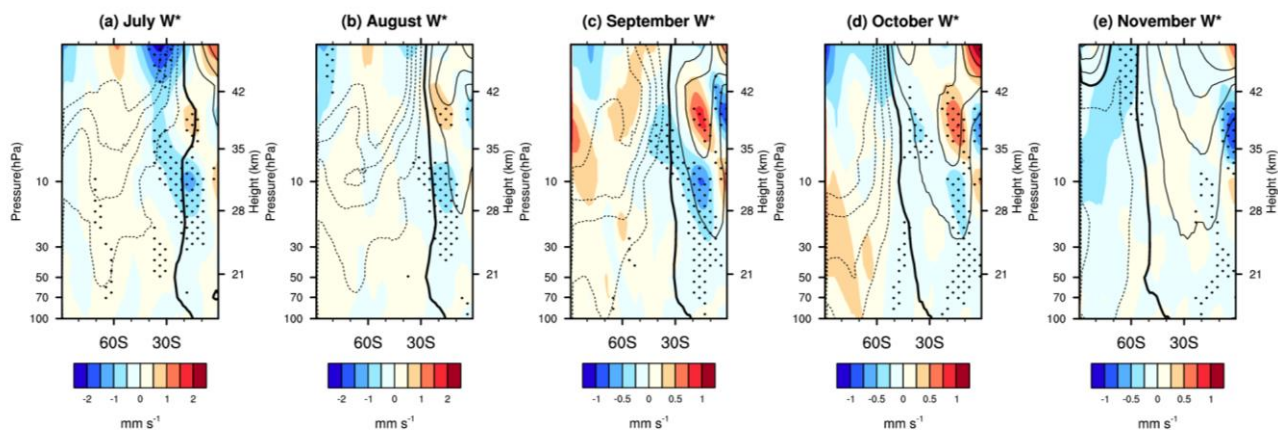
Figure 11. Composite differences (color shadings) in w* between WQBO and EQBO from July and November according to the MERRA-2 reanalysis dataset. The dotted regions mark the differences w* between the WQBO and EQBO are statistically significant at the 90% confidence level. The contours indicate the climatological w* from 1980 to 2022. The dashed lines are negative. Thick lines are zero contours,
375   and the contour intervals are 0.5 mm/s.

4 Conclusion and discussion

The dynamical coupling between high and low latitudes has been widely discussed. However, no studies have pointed out a robust connection between the QBO and the Antarctic stratospheric polar vortex, as well as the Antarctic ozone. In this study, we used the MERRA-2 reanalysis data and CESM model simulations to investigate the relationship between the QBO
380   and the Antarctic stratospheric polar vortex.

During the WQBO, positive anomalies in the zonal-mean zonal wind at 20°S−40°S in the upper stratosphere in July, named as the positive extratropical mode (Figs. 2 and 3a), lead to a stronger polar vortex in November, with a correlation reaching 0.75 (Fig. 3c). The positive extratropical mode can trigger a secondary circulation (Figs. 5f−j). The anomalous downward motions in lower latitudes and upward motions in higher latitudes cause anomalous meridional gradient of PV (Figs. 6a−j), which further alters the environmental condition for wave propagation in the stratosphere (Figs. 6k−o), pushing the positive anomalous zonal-mean zonal wind towards the pole (Fig. 7). It takes nearly five months for the positive anomalous zonal-mean zonal wind to propagate from 30°S to 60°S. Additionally, the anomalous Antarctic stratospheric polar vortex induced by the extratropical mode, along with the anomalous temperature in Antarctic stratospheric regions, can influence the Antarctic ozone hole in austral spring (Figs. 4m−o). This results in a high correlation between the extratropical mode in July and the Antarctic ozone concentration in November (Fig. 3c). Therefore, based on the relation between the extratropical mode and the strength of polar vortex, the extratropical mode can be regarded as a predictor of the Antarctic stratospheric polar vortex and ozone in austral spring during the WQBO.

While in EQBO, the correlation of the extratropical mode and the strength of polar vortex is only 0.15 (Fig. 10a). Due to stronger upward motions in the tropics, which oppose the secondary circulation caused by the extratropical mode, the EQBO can only sustain the positive anomalous zonal-mean zonal wind until September. Therefore, further analysis is needed to find a realistic connection between the EQBO and the polar vortex.

As mentioned in the literature review, QBO through the middle and lower stratospheric pathway impacts the stratospheric polar vortex (Naoe and Shibata, 2010; Yamashita et al., 2011). In this study, we propose an upper stratospheric pathway for the QBO's impact on the stratospheric polar vortex, which can persist for nearly five months. This suggests that the Antarctic stratospheric polar vortex and ozone concentration in spring can be predicted up to five months in advance. Additionally, it is generally recognized that the EQBO has a greater influence on the stratospheric polar vortex than the WQBO. However, through the upper stratospheric pathway, the WQBO during austral winter, combined with the extratropical zonal-mean zonal wind in the upper stratosphere, are highly correlated with the Antarctic stratospheric polar vortex and ozone in spring.

Data availability: MERRA-2 data are available at https://disc.gsfc.nasa.gov/datasets?project=MERRA-2. The code used in this article is accessible from the corresponding author.

Author contributions: All authors designed the study. ZW analysed and prepared the data for the paper. JZ, ZW, and SZ contributed to data interpretation and writing of the paper. The authors declare that they have no conflict of interest.

Competing interests: The authors declare that they have no conflict of interest.

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