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Opinion: Beyond Global Means: Novel Space-Based Approaches to Indirectly Constrain the Concentrations, Trends, and Variations of Tropospheric Hydroxyl Radical (OH)

Bryan N. Duncan[1], (co-authors listed in alphabetical order) Daniel C. Anderson[1,2], Arlene M. Fiore[3], Joanna Joiner[1], Nickolay A. Krotkov[1], Can Li[1], Dylan B. Millet[4], Julie M. Nicely[1,5], Luke D. Oman[1], Jason M. St. Clair[1,2], Joshua D. Shutter[4], Amir H. Souri[1,6], Sarah A. Strode[1,6], Brad Weir[1,6], Glenn M. Wolfe[1], Helen M. Worden[7], Qindan Zhu[3]

[1] NASA Goddard Space Flight Center, Greenbelt, MD, USA

[2] GESTAR II, University of Maryland Baltimore County, Baltimore, MD, USA

[3] Massachusetts Institute of Technology, Cambridge, MA USA

[4] University of Minnesota, St. Paul, MN, USA

[5] University of Maryland College Park, College Park, MD, USA

[6]GESTAR II, Morgan State University, Baltimore, MD, USA

[7] National Center for Atmospheric Research, Boulder, CO, USA

Correspondence to: Bryan N. Duncan (Bryan.N.Duncan@nasa.gov)

Abstract. The hydroxyl radical (OH) plays a central role in tropospheric chemistry as well as influencing the lifetimes of some climate gases, such as methane. Because of limitations in our ability to observe OH, we have historically relied on indirect methods to constrain its concentrations, trends, and variations, but only as annual global or semi-hemispheric averages. Recent methods demonstrated the feasibility of indirectly constraining tropospheric OH on finer spatio-temporal scales (e.g., seasonal, 1º latitude x 1º longitude), using satellite observations as proxies of the photochemical drivers of OH (e.g., nitrogen dioxide, formaldehyde, isoprene, water vapor, ozone). We found that there are currently reasonable satellite proxies to constrain about 75% of the global source of tropospheric OH and about 50% of the global sink. With additional research and investment in observing various volatile organic compounds, there is the potential to constrain an additional 10% of the global source and 30% of the global sink. In addition, these novel methods could be refined and made more robust by improvements in the capabilities of satellite instruments (e.g., signal-to-noise, spatial resolution) and retrieval algorithms that are used to develop data products. Another benefit of more robust data products is that they may be used to better constrain the chemical and dynamical processes in atmospheric chemical transport models that simulate the spatio-temporal variations of OH and OH drivers. Therefore, we propose steps forward for the development of a strategic and comprehensive space-based and suborbital observing strategy, which will improve our ability to indirectly constrain OH on much finer spatio-temporal scales than previously achieved. We discuss the strengths and limitations of such an observing strategy and potential improvements to current satellite instrument observing capabilities that would enable better constraint of OH. These improvements include ones that are obtainable with current technologies (e.g., more observations, co-located

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observations) as well as ones requiring additional technology development (e.g., to obtain vertically-resolved observations). Suborbital observations, which are required for information difficult to obtain from space and for validation of satellite-based OH estimates, will be an integral part of a comprehensive observing strategy.

1 Introduction

The hydroxyl radical (OH) has a pivotal role in atmospheric chemistry and controls the lifetimes of methane (CH_4), hydrochlorofluorocarbons/hydrofluorocarbons (HCFCs/HFCs), carbon monoxide (CO), volatile organic compounds (VOCs), and numerous other gases relevant to climate and air quality. The relationship of OH to other atmospheric constituents is often complex and non-linear. For instance, the concentrations and lifetimes of ozone (O_3) and OH are convolved as the primary source of OH involves the destruction of O_3, OH is integral in many tropospheric chemical reactions (e.g., initiating the oxidation of VOCs) that lead to O_3 production, and the reaction of HO_2 with NO is a secondary source of OH that can also lead to O_3 production.

In situ measurements provide information about tropospheric OH at a particular point in time and space. However, these observations are sparse, and OH is highly variable in space and time as its lifetime is < 1 second. Therefore, in situ observational networks are not a viable way to map the variability and trends of OH across the globe. In addition, there are few in situ instruments that can quantify OH at typical ambient concentrations of 10^5 - 10^7 molecules cm^{-3} (e.g., Stone, Whalley, and Heard, 2012). Generally, in situ measurement uncertainties are ~30% for integration times of 30 - 60 seconds (Brune et al., 2018). Comparing OH observations from multiple airborne missions to steady-state box model simulations, Miller and Brune (2022) concluded that "oxidation chemistry in most of the free troposphere is understood to as well as current measurements can determine." This implies that, for some parts of the atmosphere, further refinements in our understanding of tropospheric oxidation are limited by inherent measurement uncertainties for both OH and related species. Nevertheless, the suite of other trace gases and aerosols observed during campaigns, both with and without the sparse in situ data of OH, have provided a detailed, albeit geospatially-limited, view of tropospheric composition (e.g., Nicely et al., 2016).

There are a number of methods to indirectly constrain tropospheric OH on global scales. Methods that use in situ data of methylchloroform or a combination of HCFCs observations constrain only global or, possibly, hemispheric OH (e.g., Liang et al., 2017; Thompson et al., 2024). For methylchloroform, recent declines in tropospheric abundance limit its utility (Liang et al., 2017). Nevertheless, the multi-decadal methylchloroform observations have provided an important constraint on the long-term trends and variations of tropospheric OH (e.g., Patra et al., 2021 and references therein). Shortwave infrared (SWIR) observations of CH_4 from the Japan Aerospace Exploration Agency's (JAXA) Greenhouse Gases Observing Satellite (GOSAT) have been proposed as a way to indirectly constrain global OH (Zhang et al., 2018), though additional surface and thermal infrared (TIR) observations may be required (Zhang et al., 2021).

In addition to the lack of data for spatio-temporal constraint, uncertainties in the atmospheric processes (e.g., chemistry, emissions) that influence tropospheric OH (e.g., as summarized in Table 2 of Fiore et al., 2024; Prather

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and Zhu, 2024) have resulted in a large range in simulated OH among chemical transport models (CTMs; e.g., Nicely et al., 2017; Nicely et al. 2020; Zhao et al., 2019; Murray et al., 2021). Consequently, there is considerable uncertainty in the sink of CH_4, which contributes significantly to the overall uncertainty in the budget, interannual variability, and trends of CH_4 (Saunois et al., 2020).

Tropospheric OH is not directly observable from space; however, recent efforts have demonstrated satellite-based methods to indirectly constrain its spatio-temporal concentrations, trends, and variations on local and regional scales (e.g., 1° latitude x 1° longitude; e.g., Valin et al., 2013; Valin et al., 2016; Wolfe et al., 2019; Wells et al., 2020; Pimlott et al., 2022; Zhu et al., 2022b; Anderson et al., 2023; Zhao et al., 2023; Anderson et al., 2024; Souri et al., in press; Shutter et al., 2024; Zhu et al., 2024). These methods use various satellite datasets as proxies for OH drivers (e.g., CH_4, water vapor ($H_2O_{(v)}$), nitrogen dioxide (NO_2), CO, formaldehyde (HCHO), isoprene, tropospheric O_3). Thanks to contemporary, multi-decadal data records, these approaches open new avenues for characterizing OH over long periods of time, which is important for understanding the impact of both anthropogenic activities and natural phenomena on tropospheric OH.

The purpose of this manuscript is to make recommendations for developing a strategic and comprehensive space-based and suborbital observing strategy, which will improve our ability to indirectly constrain OH on much finer spatio-temporal scales than previously achieved. Our recommendations are intended to inform efforts to prioritize observational needs, including the upcoming 2027-2037 Earth Science Decadal Survey (ESDS) for the National Aeronautics and Space Administration (NASA), National Oceanic and Atmospheric Administration (NOAA) and United States Geological Survey (USGS) and the broader effort as proposed by Waliser & KISS Continuity Study Team (2024) "for a greater and more impactful US contribution to the global satellite observing system." In *Section 2*, we describe the current space-based approaches to constrain tropospheric OH. In *Section 3*, we summarize suborbital observations, both in situ and remote-sensed, required for information currently unobtainable from space (e.g., fine vertical resolution of satellite proxies) and for validation of satellite-based OH estimates. In *Section 4*, we discuss potential refinements to retrieval algorithms as well as to instrument capabilities that would improve satellite data products. Finally, in *Section 5*, we discuss potential improvements to current satellite observing capabilities that would better enable constraint of OH. These improvements include ones that are obtainable with current technologies (e.g., more observations, co-located observations) as well as ones requiring additional technology development (e.g., to obtain finer vertically-resolved observations).

2 Current Approaches to Indirectly Constrain Local and Regional Concentrations, Trends, and Variations of Tropospheric OH with Satellite Observations

Here we provide a brief overview of the various space-based approaches that have been used to constrain tropospheric OH as partial and/or total vertical column densities (VCDs) or surface concentrations on monthly to seasonal timescales. Each of these approaches has its strengths and limitations that need to be considered when applying it to constrain tropospheric OH.

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2.1 Process-based Approaches

These approaches exploit photochemical relationships between satellite-observable species and tropospheric OH, which aid understanding of the causes of trends and variations in inferred OH (e.g., Valin et al., 2016; Nicely et al., 2018; Baublitz et al., 2023).

- Wolfe et al. (2019) found a linear relationship between OH and HCHO using in situ observations from NASA's Atmospheric Tomography (ATom) field campaign, which is consistent with the modeling work of Valin et al. (2016) who found that HCHO VCDs primarily depend on OH production rates at low OH concentrations. This indicates that HCHO VCDs, such as from the Ozone Monitoring Instrument (OMI) on NASA's Aura satellite, could be used to infer tropospheric OH VCDs on local scales. This approach is valid *over remote oceanic environments*, where CH_4 oxidation is the primary HCHO source; there are broad regions of the troposphere that are characterized by this photochemical environment.

- Pimlott et al. (2022) developed an approach that employs a simplified steady-state approximation with multiple trace gas observations (i.e., O_3, CO, CH_4, and $H_2O_{(v)}$) from the Infrared Atmospheric Sounding Interferometer (IASI) on the European Space Agency's (ESA) MetOp-A satellite. This approach is most representative of OH *from 600-700 hPa*, though the authors argue that their approach could be reasonably applied from 400-800 hPa.

- Shutter et al. (2024) developed an approach to constrain tropospheric OH trends and variability that uses the ratio of VCDs of isoprene from the NOAA Suomi NPP Cross-track Infrared Sounder (CrIS) and HCHO from the Ozone Mapping Profiler Suite (OMPS) *over isoprene-emitting forest ecosystems*. They take advantage of the fact that the ratio of the VCDs of isoprene and HCHO scales with 1/[OH] in such environments (Wells et al., 2020).

2.2 Machine Learning (ML) Approaches

These approaches use a combination of an ML model, output from a CTM, and satellite observations. The utility of these approaches for constraining tropospheric OH depends on the accuracy of the training dataset. For example, there are known knowledge gaps in our understanding of the processes that influence OH (Fiore et al., 2024). In addition, these approaches reproduce OH only for photochemical environments that are represented in the training dataset. For example, the following two approaches were designed for polluted urban environments and cleaner tropical/subtropical ones, respectively.

- Zhu et al. (2022b) developed an approach that takes advantage of the known response of surface OH to NO_x (= NO_2 + NO) concentrations to infer trends and variations in summertime, surface OH *over selected cities*. Their six input parameters, on which their ML model was trained, explain 76% of the variance in simulated surface OH. Two of the input parameters are satellite data (i.e., OMI NO_2 and HCHO), though the other parameters are simulated by a CTM.

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- Anderson et al. (2023) developed a multi-satellite approach, testing the utility of datasets of numerous trace gases (i.e., Table 1 of Anderson et al., 2023) that influence OH through chemical, radiative and meteorological effects. They demonstrated the potential of their approach for inferring trends of tropospheric OH VCDs *over tropical and subtropical oceans (Figure 1)*, and their approach is being expanded over land and the extra-tropics. Preliminary results indicate that this approach, which uses tropospheric and total VCDs, may be used to infer OH for various layers of the troposphere; information of the vertical distribution of OH is important for estimating, for instance, CH_4 lifetime given that CH_4 loss is weighted toward the lower troposphere.

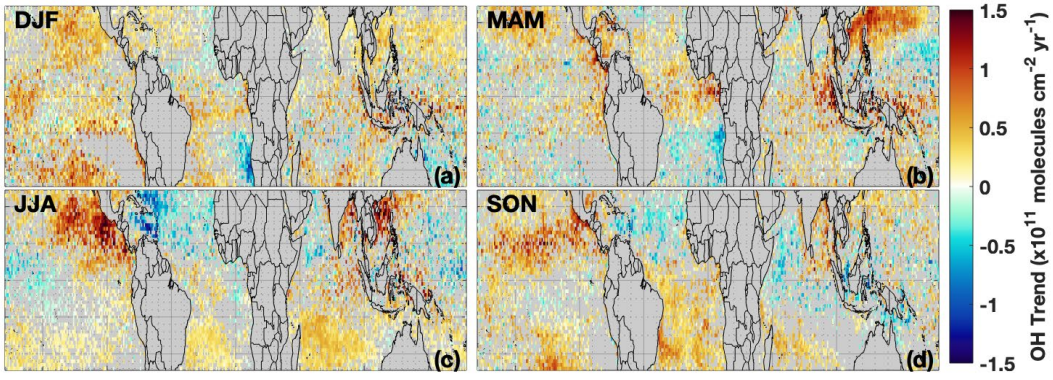

Figure 1: There is significant seasonal and spatial variability in the trend of the tropospheric VCD of OH over ocean from 2005 to 2019. We calculated the trend using the OH product described in Anderson et al. (2023) and a multiple linear regression, described in Anderson et al. (2024). Colored areas indicate those grid boxes where the trend is statistically significant and with at least 10 years of data.

2.3 Chemical Data Assimilation Approaches

Historically, these approaches incorporate satellite observations of OH drivers into CTMs to constrain tropospheric OH. They use inverse modeling and/or chemical data assimilation methods (Bocquet et al., 2015) to reconcile modeled estimates with observations while relying on the underlying chemical mechanisms of the model to infer OH concentrations consistent with the assimilated observations. A notable benefit of data assimilation is that it not only provides an improved estimate of concentrations, but the resulting increments indicate where model parameterizations are deficient and/or require improvement. Overall, the benefit of data assimilation depends on the magnitude of the difference between the observed and simulated variable.

- Miyazaki et al. (2012, 2020) and several other papers from the same group developed multi-species (notably NO_2, O_3, and CO) assimilation systems for estimating tropospheric composition and emissions. In particular, assimilation reduced the multi-model spread and annual biases of key OH drivers in four CTM frameworks. The multi-model spread for tropospheric mean OH was reduced by 24 – 58 % over polluted areas.
- Gaubert et al. (2016, 2017) assimilated Measurements of Pollution in the Troposphere (MOPITT) CO observations into the Community Atmosphere Model with Chemistry (CAM-Chem) of the Community Earth

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System Model (CESM). They showed, for example, that increases in CO because of assimilation can lead to decreases in OH and finally longer lifetimes of CH_4. In contrast, decreases in CO can lead to decreased OH and longer CH_4 lifetimes. Highlighting nonlinear effects, the study also demonstrated that assimilation led to changes in O_3 and photochemical activity due to HO_x recycling.

2.4 Simplified Hybrid Approaches

Several recent, simplified approaches allow for the quantification of the impacts of constraining various OH drivers with satellite observations, which could inform decision-making (e.g., prioritization of observables) in inverse model and data assimilation efforts (*Section 2.3*).

- Zhao et al. (2023) developed a method to post-process a CTM simulation to improve simulated tropospheric OH. Their method adjusts the monthly-averaged, simulated concentrations of OH drivers (CO, CH_4, O_3, HCHO, NO_2, total column O_3, and $H_2O_{(v)}$) with satellite observations for each model column and then uses a photochemical box model to recalculate the corresponding concentration of tropospheric OH.
- Zhu et al. (2024) used an ML technique to develop an emulator of simulated tropospheric OH VCDs in the CESM Whole Atmosphere Community Climate Model (WACCM), which facilitated their assessment of the roles of satellite VCDs of NO_2, HCHO, and CO on tropospheric OH concentrations, trends, and variations.
- Souri et al. (in press) developed a Bayesian data fusion method that adjusts monthly-averaged distributions of OH drivers from a CTM simulation using satellite observations. This method considers the quality of the datasets and the a priori used in the retrieval. The adjusted model fields are then used in a parameterization of OH (Anderson et al., 2022) in a CH_4-CO-OH cycle model (ECCOH, Elshorbany et al., 2016), allowing for easy evaluation of the impact of adjustments on OH, CH_4, and CO.

3 Recommendations for Suborbital Needs

The utility of any space-based OH estimate (*Section 2*) hinges on an adequate understanding of how OH responds to its drivers. Therefore, long-term measurements of key atmospheric properties, supplemented with occasional air and ground-based intensive campaigns, would build confidence far beyond what has been achieved with previous intermittent efforts, such as the NASA Atmospheric Tomography Mission (AToM) that provided a comprehensive suite of observations, primarily in remote oceanic environments. In addition, a comprehensive suite of co-located observations, both in situ and remote-sensed, is desired for validating CTMs, developing process-based diagnostics (*Section 5.3*), interpreting trends and variations in inferred OH, and improving satellite retrievals (*Section 4*).

We recommend the development of a long-term, robust suborbital observing network, composed of instruments at stationary sites as well as ones deployed in field campaigns. The establishment of OH "supersite" observatories would allow for the collection of a comprehensive suite of co-located measurements of key atmospheric constituents (discussed below) to constrain tropospheric OH production and loss (*Section 5.2.1*). Focusing first on augmenting existing efforts (e.g., NOAA; In-service Aircraft for a Global Observing System, IAGOS; Network for the Detection of Atmospheric Composition Change, NDACC; Advanced Global Atmospheric Gases Experiment, AGAGE) would

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keep early development efforts feasible by leveraging existing facility infrastructure. In addition, some existing efforts (e.g., four NOAA baseline observatories, IAGOS program) already observe many of the variables that reasonably constrain OH, as discussed below. Periodic field campaigns would supplement the data collected at OH supersites, especially vertical composition profiles and in photochemical environments not represented by the ground and routine aircraft networks. Continuous and co-located surface and VCD observations would be ideal, especially over a range of photochemical environments (e.g., remote to urban) and seasons. Planetary boundary layer height data would be useful for interpreting these observations. *The placement and density of OH supersites and locations/timing of field campaigns would benefit from Observing System Simulation Experiments (OSSEs) in order to maximize the value of the suborbital data for constraining OH.* For instance, a simpler strategy would likely be satisfactory over the relatively homogeneous remote oceans (e.g., fewer stations and observations of individual VOCs) as compared to more complex land environments. Therefore, the OSSE studies would also need to consider the required suite of observations at individual sites.

We recommend a focus on the collection of suborbital observations of OH drivers, given the lower uncertainties typically associated with these measurements as compared to those for OH (*Section 1*). Such observations would allow for the constraint of OH within photochemical box models, though OH budget closure remains an active area of research (Stone, Whalley, and Heard, 2012), and it is unclear whether existing in situ observations can adequately explain OH in all cases (e.g., Fuchs et al., 2017; Lew et al., 2020; Hansen et al., 2021; Yang et al., 2022; Bottorff et al., 2023; Yang et al., 2023).

The following variables reasonably constrain tropospheric OH in photochemical box models and should be priority measurements: $H_2O_{(v)}$, NO, NO_2, CO, O_3, actinic flux (or filter radiometers for wavelength ranges relevant for NO_2 and $O(^1D)$), HCHO, and temperature. CH_4 observations could be useful given that there are space-based observations. In regions closer to strong emissions sources (over and near land), chemical controls on OH are more complex and variable and a more comprehensive measurement suite (e.g., total OH reactivity, VOC composition, nitrous acid (HONO), and peroxy radicals (HO_x, RO_x); e.g., Lelieveld et al., 2016; Yang et al., 2016; Murray et al., 2021) may be needed to constrain all drivers. In remote marine environments, information on VOC speciation and reactivity is less critical, but not unimportant (Travis et al., 2020; Baublitz et al., 2023), as the background atmosphere is (relatively) uniform, and large-scale OH variability primarily reflects variability in OH production from photolysis and HO_x cycling (Wolfe et al., 2019; Brune et al., 2020).

Occasional deployment of in situ instruments that observe tropospheric OH at OH supersite observatories and in field campaigns would allow for an assessment of the consistency between observed OH and OH inferred from the satellite-based approaches (*Section 2*). As mentioned above, there are few in situ instruments and technology development is necessary before such observations could become routine. New technology development could possibly lead to an OH lidar instrument (Pan et al., 2022).

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4 Recommendations for Retrieval Algorithm Refinement

Here, we discuss potential refinements to retrieval algorithms as well as to instrument capabilities that would improve data products for the satellite-based approaches (*Section 2*).

4.1 Ultraviolet/Visible (UV/Vis)

The quality of UV/Vis satellite-based data products is determined by a number of issues, including signal-to-noise (SNR), spectral fitting uncertainties, systematic and random measurement errors, and uncertainties associated with creating VCDs (e.g., Duncan et al., 2014; Lorente et al., 2017). For example, to extract the spectral signature of HCHO used in retrievals, several interfering geophysical processes need to be accounted for, including absorption by O_3, NO_2, bromine monoxide (BrO), and oxygen dimer (O_2-O_2), as well as rotational Raman scattering (Ring effect; e.g., González Abad et al., 2016). Coupled with its small atmospheric loading, especially over remote oceanic areas, the HCHO VCDs are "noisy," requiring temporal and/or spatial averaging to reduce random errors so that the geophysical signal becomes clear (e.g., Liao et al., in review). As another example, NO_2 VCDs are dominated by the stratospheric component over remote areas, leading to relatively large uncertainties in the corresponding tropospheric portion of VCDs (e.g., Lamsal et al., 2021). Improvements in data quality, whether through data product refinement (below), oversampling techniques, or enhancements in instrument capabilities (*Section 5*), will be required to use the satellite-based approaches (*Section 2*) at finer spatial and temporal resolutions than achievable with current instruments.

Work is ongoing to improve the accuracy and precision of satellite data products through retrieval algorithm refinement and also using, for example, data-driven ML data analysis methods that show promise for reducing noise in data products of low abundance, weak absorbers (e.g., Li et al., 2022; Joiner et al., 2023). For instance, Joiner et al. (2024) have tested their technique on NASA Tropospheric Emissions: Monitoring of Pollution (TEMPO) HCHO retrievals. Preliminary results (*Figure 2*) show visibly less noise while keeping overall consistency with the original noisy data and there is a reduction in the number of pixels with negative slant column densities (SCD) values in the noise-reduced data.

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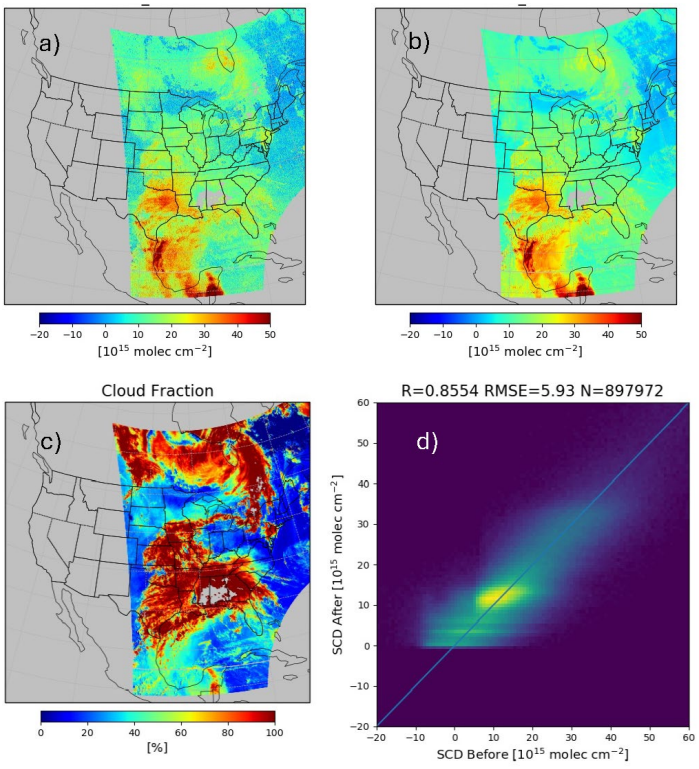

Figure 2. Preliminary results of noise reduction applied to HCHO retrieved SCD from version 3 TEMPO data on 5 May 2024 scan 5: a) original data before noise reduction used as training; b) after noise reduction applied; c) effective cloud fraction supplied in the L2 data; and d) scatter diagram showing overall agreement between noise-reduced and original SCDs relative to the 1:1 line with fit statistics. R is correlation, RMSE is root mean squared error, and N is the number of pixels used in the comparison.

Additionally, different retrieval methods applied to the same instrument often produce large differences in data products, which impact inferred OH (e.g., Anderson et al., 2023). Discrepancies created by retrieval algorithm assumptions could be resolved with more robust suborbital data for validation (*Section 3*), especially for a priori profiles used in retrievals.

Additional potential areas of retrieval algorithm refinement include aerosol correction and stratosphere-troposphere separation that will benefit from improvements in radiative transfer forward model simulations as implemented in L2 retrieval algorithms. For example, to reduce computation time, many retrieval algorithms use pre-computed lookup tables and do not explicitly represent important geophysical processes, such as aerosol scattering/absorption. With sufficient acceleration, radiative transfer models or ML proxies can be used directly in Level 2 (L2) algorithms, reducing interpolation errors and providing greater flexibility to use input data that are more representative of the atmospheric state (e.g., aerosol profiles from MERRA-2 reanalysis used in OMI NO_2 retrievals; Vasilkov et al., 2021);

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satellite data processing levels indicate the degree of data processing of satellite data, where L2 indicates that a geophysical parameter has been derived.

4.2 Thermal Infrared (TIR)

Trace gas measurements in the TIR often target species with relatively weak absorption, and low instrument noise is thus a key requirement for useful retrievals (Fu et al., 2019). High spectral resolution is also pivotal for resolving relevant spectral features and distinguishing them from atmospheric or surface interferences (Clarisse et al., 2011). In addition to providing spectral coverage encompassing the chemical species of interest, future TIR sounders aiming to constrain atmospheric composition therefore need to prioritize these instrumental characteristics.

New retrieval developments employing ML and species detection based on a hyperspectral range index (Walker, Dudhia, and Carboni, 2011) have led to increased sensitivity and expanded the suite of species that can be measured in the TIR (e.g., acetic acid, acetone, ethane, ethene, ethyne, formic acid, glycoaldehyde, hydrogen cyanide, isoprene, methanol, peroxyacetyl nitrate (PAN); e.g., Wells et al., 2020; Franco et al., 2022; Franco et al., 2024; Brewer et al., accepted). Since TIR observations are based on the thermal contrast between the Earth surface and the absorber, uncertainties increase when that temperature difference is small. For the same reason, profile shape uncertainties can be an important error source in TIR retrievals. Validation data are often limited for species targeted in this spectral range, and future space-based measurements would benefit from an expanded portfolio of surface and airborne data to validate retrievals and refine profile shape assumptions. Observational errors can also arise from uncertainties in the absorption cross section and surface emissivity datasets employed in retrievals; advancing the fidelity of these resources would therefore benefit the reliability of TIR trace gas products.

4.3 Short-wave Infrared (SWIR)

Measurements of trace gases in the SWIR rely on reflected solar radiance and can provide a true total column observation through the atmosphere under cloud-free conditions and when the atmospheric pathlength is characterized. Here we assume the SWIR range includes wavelengths from approximately 0.7-3 μm. This range is especially useful for measuring total columns (surface to top-of-atmosphere) of CO_2 (Greenhouse gases Observing SATellite, GOSAT, and, Orbiting Carbon Observatory-2, OCO-2), CH_4 (GOSAT and TROPOspheric Monitoring Instrument, TROPOMI) and CO (MOPITT, GOSAT-2 and TROPOMI). In many cases, the O_2-A band around 0.76 mm is used for a reference observation of atmospheric path and cloud detection due to the constant and well-known concentration of molecular oxygen. SWIR observations of total column CO and CH_4 are especially important for quantifying surface source emissions that have a significant contribution to OH sinks (*Section 5.2*)

4.4 Enhancing Vertical Resolution

As is the case for suborbital observation needs (*Section 3*), vertically-resolved data of OH drivers are a priority for future satellite instruments to improve retrieval algorithms, especially for a priori profiles. In addition, they will advance the ability of the satellite-based approaches (*Section 2*) to constrain tropospheric OH at finer vertical

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resolution than a tropospheric VCD measurement can provide. Such data are also useful for validating CTMs, developing process-based diagnostics (*Section 5.3*), and interpreting the causes of trends and variations in inferred OH throughout the tropospheric column (e.g., Baublitz et al., 2023). For example, the efficiency and quality of the ML approaches (*Section 2.2*) would benefit from vertically-resolved input variables.

We recommend that studies be done to maximize the vertical information of the observations of OH drivers, which may be achieved through a number of approaches, including developing satellite instrument technology (e.g., multi-angle and/or multispectral instruments) and investing in the development of multispectral data products that use radiances from one or more instruments (e.g., CH_4, Zhang et al., 2018). Multispectral data products were developed for O_3 (TES/OMI, Fu et al., 2013, Colombi et al., 2021; Atmospheric Infrared Sounder (AIRS)/OMI, Fu et al., 2018; TROPOMI/Cross-track Infrared Sounder (CrIS), Mettig et al., 2022) and CO (MOPITT NIR/TIR, Worden et al., 2010, Gaubert et al., 2017; CrIS/TROPOMI, Fu et al., 2016); development of these and other products is occurring through the NASA TRopospheric Ozone and Precursors from Earth System Sounding (TROPESS) project.

In addition to gaining vertical information through the development of satellite instrument technology and multispectral algorithms, one may combine multiple data products of an OH driver from instruments that observe different tropospheric vertical layers. As an example, Oman et al. (2013) constrained the response of O_3 to ENSO using vertically-resolved data from the NASA Tropospheric Emission Spectrometer (TES) and Microwave Limb Sounder (MLS), both on the NASA Aura satellite. In this analysis, MLS and TES data were treated separately and not merged. (TES operations ended in 2018 and MLS operations are predicted to end in mid-2026 with no follow-on instruments being currently planned.)

5 Recommendations for Satellite Needs

In this section, we discuss considerations and make recommendations for an observing strategy that would advance our ability to quantify the concentrations, trends, and variations of tropospheric OH using atmospheric remote-sensing observations of its drivers.

5.1 Optimizing an Observing Strategy through Tradespace Analyses

As discussed in *Section 2*, there are a number of approaches in the literature to indirectly constrain OH with satellite proxies of OH drivers. Each of the approaches applies to specific regions and/or photochemical environments or were devised for specific applications, and each of them has its strengths and limitations that should be considered. Therefore, *some combination of these approaches, as well as the development of new approaches for specific photochemical environments, may be required to maximize spatio-temporal coverage of tropospheric OH around the globe for individual applications (as discussed below).*

The design process of a comprehensive, multi-satellite observing strategy would benefit from tradespace analyses (i.e., assessing the impact of changing one or more variables while simultaneously changing one or more other variables in the opposite direction). For example, such analyses of spatial, temporal, and spectral resolutions for

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specific instruments would inform design decision making. *The design process would also benefit from cost-benefit analyses.* For instance, what observations are achievable using more cost-effective suborbital instruments (*Section 3*) than satellite instruments?

A few considerations for optimizing an observing strategy include:

5.1.1 Continuity

The fidelity of a sustained, long-term record of inferred tropospheric OH from satellite-based approaches (*Section 2*) depends on continuity of satellite proxies of OH drivers. Therefore, *a priority for a robust observing strategy is for continuity of satellite instruments with similar or enhanced capabilities as well as inter-agency coordination to maintain self-consistent and trend-quality Level 1B (L1B) radiances for the generation of multi-instrument data records*; L1B data, which are used in scientific algorithms to derive geophysical parameters, have been calibrated and geo-located. The transition from one instrument to its successor instrument ideally would include a period of observational overlap (e.g., 2-5 y) so that self-consistent, multi-instrument data records may be achieved. As an example, near daily, global coverage of the NO_2 VCD is currently observed by OMI (since 2004) and its successor, TROPOMI (since 2017) as well as OMPS on Suomi NPP (since 2011), NOAA-20 (since 2017), and NOAA-21 (since 2022). All five instruments have overpass times near 13:30 local time. However, the OMPS instruments are missing visible wavelengths (e.g., < 380 nm for Suomi NPP) as compared to OMI or TROPOMI and have lower spectral resolutions (~1 nm), which result in lower quality NO_2 data products than those from either OMI or TROPOMI. The planned successor to TROPOMI, Sentinel-5, will have similar capabilities as TROPOMI, but will have an equator overpass time near 9:30 local time. Therefore, there will be discontinuities in the long-term record given that NO_x emissions, NO_x partitioning between NO and NO_2, and the vertical distribution of NO_2 within the VCD vary throughout the day. Depending on TROPOMI's lifespan, potential successor instruments, with overpass times near 13:30 local time, are the NOAA series of OMPS instruments being planned for the Joint Polar Satellite System (JPSS)-3 and JPSS-4 satellites. JPSS-3 and JPSS-4 OMPS will observe additional visible wavelengths relative to those observed by earlier OMPS instruments, but they will not have comparable instrument capabilities (e.g., SNR, spectral coverage) as OMI or TROPOMI. Therefore, the data quality of a self-consistent, multi-instrument data record for NO_2 at an overpass time near 13:30 local time from OMI, TROPOMI and OMPS will be determined by the instrument parameters (e.g., spectral resolution, wavelength range) of the least capable OMPS instrument.

5.1.2 Accuracy/Precision and Signal-to-Noise (SNR)

As discussed in the scientific papers describing the satellite-based approaches (*Section 2*), *additional improvements to accuracy/precision and SNR of current observing capabilities are required to advance our ability to constrain OH drivers.* For instance, the design of new instruments should emphasize capabilities for in-flight instrument calibration and monitoring, such as solar diffusers and internal light sources. During the manufacturing and integration phase, adequate time and resources should be allocated to allow for thorough pre-flight calibration and characterization. As

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another example, technology improvements, such as better thermal design and detector cooling, can lead to reduced noise and enhancements in SNR.

5.1.3 Spatio-Temporal Coverage

In addition to enhanced instrument capabilities, *improvements to current spatio-temporal coverage of satellite proxies of OH drivers could be achieved with additional satellites*. For instance, the datasets used in the satellite-based approaches (*Section 2*) are from satellites in low Earth orbit (LEO), so most global locations, where OH levels are highest (i.e., tropics and subtropics), are observed daily or every few days. Therefore, a suite of instruments could be placed on multiple satellites, providing observations at various times throughout daylight hours. This strategy would also have the benefit of providing more opportunities to observe non-cloudy conditions (*Section 5.4*), especially if the instruments have finer spatial resolutions than current satellite instruments to allow more observations between clouds.

Instruments on geostationary (GEO) satellites provide an opportunity to assess the value of hourly data to constrain OH. For instance, several new UV/Vis instruments, Korea's Geostationary Environment Monitoring Spectrometer (GEMS; Kim et al., 2020) and NASA's Tropospheric Emissions: Monitoring of Pollution (TEMPO; Zoogman et al., 2017), observe VCDs of NO_2, HCHO and ozone over East Asia and North America, respectively. ESA's Sentinel-4 GEO satellite with the Ultraviolet-Visible-Near Infrared (UVN) instrument will make similar observations over Europe and northern Africa (Courrèges-Lacoste et al., 2017). The spatial coverages of these instruments are primarily limited to the mid-latitudes and subtropics of the northern hemisphere. For hyperspectral TIR instruments on GEO satellites, there already exists the Chinese Fengyun Geostationary Interferometric Infrared Sounder (GIIRS; Yang et al., 2017); similar instruments are tentatively planned by space agencies in Japan (Okamoto et al. 2020), Europe (Holmlund et al., 2020) and the US (Lindsey et al., 2024). The spatial coverages of these instruments are more comprehensive, covering most latitudes (i.e., below ~60°) of both hemispheres, as compared to the coverages of the UV/Vis instruments. *We recommend coordinated deployment of UV, Vis, NIR, and TIR instruments on GEO satellites to provide co-located observations of satellite proxies of key OH drivers (Section 5.2).*

5.1.4 Case Studies

Ultimately, the design of a space-based observing strategy to constrain tropospheric OH depends on the specific science need or needs to be addressed. For instance, is the need to constrain OH concentrations, trends, and/or variations, which could require different designs? For what regions and seasons? For illustrative purposes, we provide design considerations for two different potential applications.

In *Case 1*, we wish to constrain the spatio-temporal variations of methane's loss by reaction with OH, which occurs primarily in the lower troposphere of the tropics and subtropics. Given methane's long lifetime, one might think that the spatial and temporal resolutions of inferred tropospheric OH could be low (e.g., several degrees latitude by longitude, seasonal in photochemically homogeneous environments). However, one issue that requires consideration is that the number of pixels with sufficiently low cloud contamination would decrease as the horizontal resolution of

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observations degrades in cloud-prone areas (e.g., in much of the tropics and subtropics; *Section 5.4*). Therefore, an observational strategy for this case would require finer spatial resolutions for measurements of OH drivers than required for the spatial resolution of inferred OH. In addition, much of the loss of CH_4 occurs over relatively remote ocean environments so that improving the SNR of OH drivers is particularly important, especially since the signals for some important species (e.g., NO_2, HCHO) are on par with noise for current instruments (e.g., Zhu et al., 2020). Potentially, new satellite missions can be designed to take advantage of sun glint, where specular reflection over water bodies can enhance the SNR. Over land in the tropics and subtropics, observations of additional VOC constituents, especially isoprene (e.g., Fu et al., 2019; Shutter et al., 2024), are desired to constrain tropospheric OH (*Section 5.2*). Given the strong spatial gradients of OH drivers, spatial resolution should also be prioritized over land as well as over ocean affected by offshore flow. Satellite observations of dynamical variables (e.g., meteorological, climate indices) may also be useful for this case as they can influence OH (*Section 5.3*). *For this case, design objectives may be met by multiple instruments with different design objectives (e.g., fine spatial resolution for over land vs. improved SNR for over ocean) or instruments with multiple sampling modes.*

In *Case 2*, we wish to constrain tropospheric OH in polluted urban areas. This environment is typically heterogeneous and complicated, requiring fine spatio-temporal resolutions (e.g., kilometers, hourly-daily) to capture spatial gradients in OH drivers. Additional VOC constituents are highly desired as the urban environment has thousands of VOCs; therefore, observations of VOCs that represent classes of VOCs (e.g., isoprene, fast-reacting alkenes, alkanes, aromatics) are desired and will require multispectral satellite observations (i.e., TIR, NIR, UV/Vis) for the development of these data products (*Section 5.2*). Depending on the timescale of interest, photochemical variables may be sufficient to constrain OH, which can be assumed to be in photostationary state given its short lifetime (< 1 s). Nevertheless, meteorological variables that influence OH (e.g., temperature, convection) may be desired for studying the photochemical evolution of the urban plume by hour, for instance. *For this case, design objectives may be met by observations of UV/Vis (e.g., NO_2, HCHO) and TIR (VOCs) wavelengths. Instruments on GEO satellites would provide fine temporal resolution as opposed to ones on LEO satellites.*

5.2 Assessment of Current Satellite Data to Constrain Chemical Processes that Influence OH

We recommend studies to determine the optimal combination of input variables/satellite observables for the satellite-based approaches (Section 2) to maximize the interpretability (in a process-based sense) of the causes of trends and variations in tropospheric OH. ML approaches (*Section 2.2*) are able to reproduce OH quite well from the training datasets using various satellite observations; however, the utility of the approaches for attribution studies is determined by the choice of input variables. That is, the suite of input variables must adequately represent the chemical processes of interest that determine OH. For instance, attribution of OH variations and trends to the influence of CH_4 is not possible in an ML study if CH_4 concentration is not included as an input variable (e.g., Anderson et al., 2024). In addition, ML techniques may not reproduce OH in a process-based sense if the input variables to the ML model are not independent of one another, which is an inherent characteristic of atmospheric chemistry problems (e.g., O_3 and NO_2; temperature-dependence of most species' lifetimes; co-emitted species, such as CH_4 and VOCs; and tropospheric

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species impacted by the same meteorology). Therefore, the causes of trends and variations in tropospheric OH inferred from ML approaches should be carefully interpreted. This is also true for process-based approaches (*Section 2.1*). For example, Pimlott et al. (2022) concluded that their approach "suggests that O_3 and CO were the key drivers of variability in the production and loss of OH" for their study period. However, their approach uses input variables observed from a single satellite instrument that does not collect data of NO_2, which is responsible for a substantial portion of the variations of tropospheric OH (e.g., Zhu et al., 2022a; Anderson et al., 2023; Baublitz et al., 2023).

5.2.1 A Thought Experiment

A key question is as follows: What chemical processes are not adequately represented in the current suite of satellite observations used in the satellite-based approaches (*Section 2*) to indirectly constrain local and regional concentrations, trends, and variations of tropospheric OH? To answer this question, it is instructive to assess our current ability to constrain individual sources and sinks for tropospheric OH (*Table 1*). For this thought experiment, we use the annual tropospheric OH budget presented in Table 1 of Lelieveld et al. (2016), noting that the relative importance of the sources and sinks for various regions and seasons can vary substantially from the annual, global distribution. In addition, we acknowledge that OH budgets can vary significantly from one CTM to another (*Section 1*). As we will show, the results of our thought experiment, as summarized in Table 1, reveal that the ideal observing strategy for the indirect constraint of the chemical processes that control OH requires co-located observations of UV/Vis, NIR, and TIR wavelengths.

Sources: The first and second sources (i.e., $O(^1D)+H_2O_{(v)}$ and $NO+HO_2$; *Table 1*) account for 63% of the total source in Lelieveld et al. (2016) and are represented relatively well with current satellite datasets. For the second source, satellite observations of NO_2 can be used to estimate NO using a CTM; this conversion is a function of a number of variables, such as ozone concentrations and photolysis (e.g., de Foy et al., 2015). In addition, lightning is an important source of NO_x in the middle and upper troposphere (Allen et al., 2021), which modulates OH there (e.g., Fiore et al., 2006). Satellite observations of lightning flash counts (e.g., over the Americas from the Geostationary Lightning Mapper, GLM, aboard the NOAA Geostationary Operational Environmental Satellite-16, GOES-16; over Europe, Africa and Middle East from EUMETSAT Meteosat Third Generation – Imager 1 (MTG-I1) Lightning Imager) may be useful in combination with NO_2 VCDs, especially to constrain the vertical distribution of NO_2. While we do have estimates of tropospheric O_3 VCDs, representation of the third source (i.e., O_3+HO_2) would benefit from an accurate separation of the stratospheric and tropospheric portions of the total O_3 VCD (e.g., Ziemke et al., 2006; Orfanoz-Cheuquelaf et al., 2024) in addition to information on the vertical distribution of tropospheric O_3 (*Section 4.4*). We are unaware of satellite proxies for tropospheric HO_2, which occurs in both the second and third sources and has a lifetime that is slightly longer than the lifetime of OH. Field campaign data from remote ocean regions indicate that HO_2 variation does modulate the rates of the second and third OH source reactions (Figure 3a, b), but whether possible proxies for HO_2 (Figure 3c) are robust in more complex photochemical regimes is a question that requires further research. A proxy for HO_2 may not be necessary for the ML approaches (*Section 2.2*) given that OH and HO_2 (i.e., HO_x family) share many of the same chemical drivers. There are satellite observations from Atmospheric Chemistry

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Experiment (ACE) for H_2O_2, the fourth source (i.e., H_2O_2+hν), in the mid to upper troposphere; these observations are sparse so multi-year averages are required to obtain seasonal, zonally-averaged distributions (Allen et al., 2013). For satellite proxies that are not yet well supported by current satellite instruments, such as H_2O_2 and many VOCs (discussed below), simulated distributions may be used in the satellite-based approaches (*Section 2*), especially those
475 distributions which have been well validated with suborbital datasets (*Section 3*). Similar to the argument for HO_2, OH and H_2O_2 share many of the same chemical drivers, so we may not need a satellite proxy for H_2O_2 in ML approaches (*Section 2.2*). We will discuss the fifth source, which involves VOCs, below. We added the source of OH from the photolysis of nitrous acid (HONO; HONO + h*v*) to *Table 1*, though it is not included in Table 1 of Lelieveld et al. (2016); the overall importance of this source on a global scale is uncertain (e.g., Wu et al., 2022; Ha et al., 2023;
480 Zhang et al., 2023), but is thought to be an important source in urban and agricultural environments. Satellite observations of HONO (e.g., TROPOMI, IASI) are primarily limited to intense wildfire plumes (e.g., Theys et al., 2020; Fredrickson et al., 2023; Franco et al., 2024).

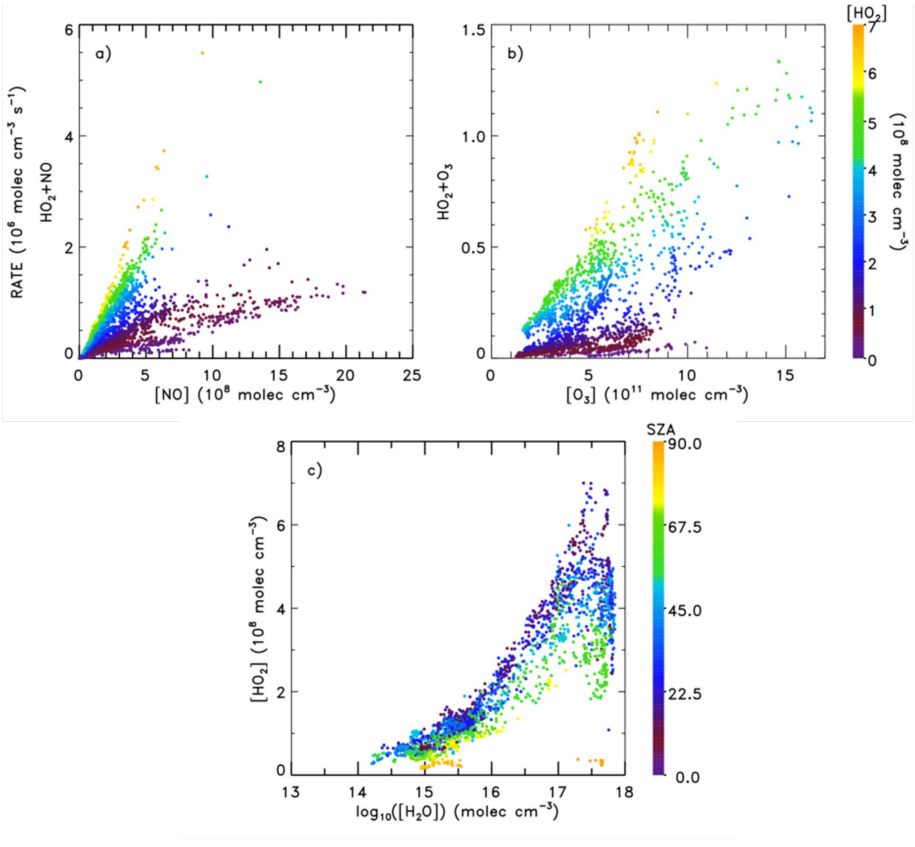

Figure 3. Data shown are for over-ocean, tropical samples from the ATom-1 field campaign. Box modeled rates are
485 for the HO_2+NO reaction plotted against NO concentration (panel a) and the HO_2+O_3 reaction versus O_3 concentration

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(b), both colored by concentration of HO_2. Panel c shows HO_2 concentration plotted against $\log_{10}$ of water vapor concentration and colored by solar zenith angle.

Sinks: The sinks that have straightforward satellite proxies are the second and third sinks (i.e., OH+CH_4; OH+CO; e.g., Gaubert et al., 2017; Nguyen et al., 2020; Jacob et al., 2022; Worden et al., 2022), which together represent 51% of the total tropospheric OH sink from Leliveld et al. (2016). For the first sink (i.e., OH+HO_y, where HO_y = H_2, O_3, H_2O_2, radical–radical reaction), there are satellite data for tropospheric O_3 and limited data for H_2O_2 as discussed above. In addition, H_2 is relatively well mixed throughout the troposphere, therefore we could assume a tropospheric distribution informed by in situ observations. We will discuss the fourth and fifth sinks, which involve VOCs, below.

Overall, we currently have reasonable satellite proxies to constrain 77% of the global source and 51% of the sink (*Table 1*) in the annual tropospheric OH budget of Leliveld et al. (2016). With development of new VOC data products (e.g., De Longueville et al., 2021; Franco et al., 2022; Brewer et al., accepted; Wells et al., submitted), which may require new technology development, we may well be able to represent a larger portion of the total sources (up to 90%) and sinks (up to 80%). To do this would require that we leverage numerous spectral regions (UV/Vis, NIR, TIR) to maximize our ability to represent key source and sink reactions for tropospheric OH with satellite proxy data.

VOCs: We recommend additional research to be a priority to explore the potential of these observable VOCs to constrain tropospheric OH. In *Table 1*, there is one source (i.e., OVOCs, ROOH+hν; 13% of total source) and two sinks (i.e., OH+other C_1 VOC; OH+C_{2+} VOC; 29% of total sink) that would benefit from the development of space-based observations of VOCs, including oxygenated VOCs (OVOCs). Several VOCs are currently observable from space, including in the UV/Vis (HCHO, glyoxal) and TIR (acetic acid, acetone, ethane, ethene, ethyne, formic acid, glycoaldehyde, hydrogen cyanide, isoprene, methanol, peroxyacetyl nitrate (PAN); e.g., Franco et al., 2018, 2019, 2020, 2022; Wells et al., submitted). Given that it is not feasible to observe the hundreds of individual VOCs in the troposphere and potentially thousands in the polluted urban atmosphere, we recommend that observable VOCs be identified that could represent 1) important classes of compounds (e.g., alkanes, alkenes, aromatics) that contribute to total OH reactivity and 2) markers of various VOC emission sources (e.g., biogenic, anthropogenic, pyrogenic), which have characteristic VOC mixes. Additional constraint of these sources and sinks may benefit from other satellite proxies, such as CO (e.g., Baublitz et al., 2024), that covary with VOCs under certain conditions.

Rate constants: We have satellite proxies for the dependence of reactions on temperature and sunlight. The temperature dependence of chemical reactions is represented by satellite observations of temperature (TIR) and continuity is assured with CrIS, for example. The stratospheric O_3 VCD (UV/Vis) may serve as a proxy for sunlight (hν) where photochemically important wavelengths are modulated by the thickness of the stratospheric O_3 layer. Continuity is assured with several current and upcoming UV/Vis instruments (e.g., TROPOMI, OMPS). Aerosols also modulate the amount of sunlight in the troposphere, affecting tropospheric oxidants (e.g., Martin et al., 2002; Mok et al., 2016; Madronich et al., 2024). There are several potential proxies, such as total and absorption aerosol optical depth and single scattering albedo, that are observed by a number of satellite instruments. Additional research is

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required to develop a combined satellite proxy for spectral solar planar and scalar (i.e., actinic flux) irradiances (e.g., Vasilkov et al., 2022).

In summary, the results of our thought experiment, as summarized in Table 1, reveal that the ideal observing strategy for the indirect constraint of the chemical processes that control OH would require co-located observations of multiple wavelengths in the UV/Vis, NIR, and TIR.

5.2.2 Looking Backward

This manuscript is primarily focused on looking forward - i.e., what is required/desired to improve our ability to constrain tropospheric OH with satellite observations? However, a natural question is "How far back in time can we adequately constrain tropospheric OH?" Addressing this question and comparing estimates of past OH abundance would increase confidence in our capability to project OH into the future. Answering this question depends on the level of explanatory power that is desired as discussed above. For instance, one of the first satellite proxies listed in *Table 1* is the stratospheric O_3 VCD. As shown by Rohrer and Berresheim (2006), tropospheric OH correlates linearly with solar UV irradiance, which is modulated by stratospheric O_3. Therefore, one could partially constrain tropospheric OH back to 1979, when satellite measurements of stratospheric O_3 became routine (e.g., Stolarski et al., 1986; Weber et al., 2018). As another example, discussion surrounding continuity of the NO_2 VCD (*Section 5.1.1*) started with OMI (launched in 2004), though such observations actually began in 1996 with ESA's Global Ozone Monitoring Experiment (GOME; 1995-2011; Burrows et al., 1999), which was followed by SCanning Imaging Absorption SpectroMeter for Atmospheric CHartographY (SCIAMACHY; 2002-2012), GOME-2 instruments on the Meteorological Operational (METOP) satellites (METOP-A, 2006-2021; -B, since 2012; and -C, since 2018). All these ESA instruments were/are in morning orbits. Therefore, depending on one's tolerance for data that have relatively coarser resolution, poorer SNR, or a longer time period to obtain global coverage, partial constraint of tropospheric OH with satellite data of NO_2 from a morning orbit could extend back to 1996. Continuity of the morning orbit is planned with Sentinel-5 (anticipated launch in 2025).

Comprehensive observations of satellite proxies (*Table 1*) of tropospheric OH began with the launches of Terra (1999), Aqua (2002), and Aura (2004), the flagship missions of NASA's Earth Observing System (EOS). All three satellites are still currently operating but nearing the end of their operating lifetimes. Continuity is assured for most satellite proxies (*Table 1*) from Terra, Aqua, and Aura. Consequently, *we currently have the ability to constrain adequately or partially most of the sources and sinks of tropospheric OH (Table 1) with the satellite-based approaches (Section 2) from late 2004 to the present.* As discussed in *Section 5.2.1*, the primary exception is related to VOCs (e.g., isoprene, which became available only in 2012).

Table 1. Satellite proxies for tropospheric source and sink fluxes of OH.

	Sources [a]	% of	Satellite Proxies	Satellite Proxies Limitations	Data Continuity? [h]

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		Total Sources [a]	(wavelengths)		
1	$O(^1D)+H_2O_{(v)}$	33%	Stratospheric O_3 VCD (UV/Vis) for $O(^1D)$. Water vapor (IR).	No major limitations.	Stratospheric O_3 VCD (yes) - e.g., TROPOMI. Water vapor (yes) - e.g., CrIS.
2	$NO+HO_2$	30%	NO_2 VCD (UV/Vis). Lightning flash counts. Research is needed to identify a proxy for HO_2.	Needs improved SNR where NO_2 VCDs are low (Buscela et al., 2013).	NO_2 VCD (yes) - e.g., TROPOMI. Lightning flash counts (yes, over some continents) - e.g., NOAA GeoXO Lightning Mapper (LMX); MTG-I1 Lightning Imager
3	O_3+HO_2	14%	Tropospheric O_3 VCD (UV/Vis). Research is needed to identify a proxy for HO_2.	Needs accurate stratosphere-troposphere separation of total column O_3 VCD (Ziemke et al., 2006).	Tropospheric O_3 VCD (yes) - e.g., TROPOMI, OMPS.
4	$H_2O_2+h\nu$	10%	H_2O_2 (IR).	Instrument and/or retrieval algorithm development required. ACE data of H_2O_2 are sparse; zonal averages of multiple years are required to obtain near-global coverage in the mid to upper troposphere (Allen et al., 2013).	H_2O_2 (no).
5	OVOCs[b], ROOH[c]+hν	13%	HCHO, glyoxal (UV/Vis). Numerous potential	Needs improved SNR where HCHO VCDs are low. VOC instrument (TIR) and	HCHO, glyoxal VCD (yes) - e.g., TROPOMI.

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			VOCs (TIR).	retrieval algorithm development required.	Numerous VOCs (yes for some, but not others) - e.g., CrIS.
	HONO[d] + hν	–	HONO (UV/Vis; TIR).	Observations primarily for intense wildfire plumes that reach altitude. Data are very noisy.	HONO (yes) - e.g., TROPOMI.
	Sinks [a]	% of Total Sinks [a]			
1	OH+HO_y[e]	18%	Tropospheric O_3 VCD (UV/Vis). H_2O_2 (IR). Assume constant distribution of H_2.	See source #4 above.	See source #3 above.
2	OH+CH_4	12%	CH_4 (IR).	Needs (1) better SNR to detect variations in high background concentration and (2) improved sensitivity to near-surface.	CH_4 (partially, yes) - e.g., TROPOMI (over land, glint over ocean).
3	OH+CO	39%	CO (IR).	Improved sensitivity to near-surface required.	CO VCD (yes) - TROPOMI, CrIS.
4	OH+other C_1 VOC [f]	15%	Methanol.	VOC instrument (TIR) and retrieval algorithm development required.	Numerous VOCs (yes for some, but not others) - e.g., CrIS.
5	OH+C_{2+} VOC [g]	14%	Isoprene. PAN, etc.	VOC instrument (TIR) and retrieval algorithm development required.	Numerous VOCs (yes for some, but not others) - e.g., CrIS.

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[a]Reproduced from Table 1 of Lelieveld et al. (2016), except neglecting two minor sinks. [b]OVOCs = oxygenated VOCs, such as acetone and acetaldehyde. [c]ROOH = organic peroxides, such as CH_3OOH. [d]HONO is not explicitly listed as a source of OH in Leliveld et al. (2016), though it is an important source in some environments (Theys et al., 2020; Fredrickson et al., 2023). [e]HO_y = H_2, O_3, H_2O_2, radical–radical reaction. [f]VOC with one C atom (excluding methane), including methanol, C1-reaction products. [g]VOC with ≥2 C atoms, C2+ reaction products. [h]Defined in *Section 5.1.1*. We focus on satellite instruments that provide near complete spatial coverage of the troposphere, neglecting instruments that provide data of, for instance, the upper troposphere (e.g., MLS) and geostationary orbits.

5.3 Influence of Tropospheric Dynamics on OH

We recommend additional research to determine the full potential of using satellite-observable meteorology (e.g., temperature) and climate variables in space-based approaches (Section 2) to better understand the dependence of tropospheric OH on dynamics. For instance, climate indices, such as the Multivariate El Niño–Southern Oscillation (ENSO) Index (MEI) and indicators of drought, correlate well with long-term (e.g., monthly, seasonal) variations in tropospheric constituents (e.g., CO, NO_2, O_3, isoprene, water vapor) that influence OH (e.g., Oman et al., 2011; Oman et al., 2013; Wells et al., 2020; Anderson et al., 2021; Anderson et al., 2023; Shutter et al., 2024) as well as OH itself, primarily through NO_x emissions associated with lightning (Turner et al., 2018). Anderson et al. (2021) estimated that ~30% of simulated variability in global OH in boreal winter is associated with ENSO alone. However, additional proxies may be required to fully capture the complex impact of meteorological variations on the tropospheric constituents that influence OH, including weather-dependent emissions (e.g., CH_4, isoprene, lightning NO_x) and anthropogenic activities (e.g., Shutter et al., 2024). For instance, Field et al. (2016) reported a nonlinear sensitivity of CO abundance to emissions from fires in Indonesia during El Niño-induced drought, when many fires escape or are intentionally set to clear land on areas of degraded peat. Building on this research, the locations of peat deposits convolved with satellite observations of fire-counts could be used to estimate pollution emitted from these fires. Alternatively, a total sum of pollutant emission estimates (e.g., Global Fire Emissions Database, GFED, van Wees et al., 2022) from Indonesia or the preceding dry season total rainfall in conjunction with the locations of degraded peat could be used to capture the impact of ENSO variations on tropospheric OH.

We recommend additional research to develop process-based diagnostics using satellite and suborbital observations to improve the representation of key atmospheric processes in CTMs that influence OH. Though many of the current knowledge gaps in CTMs (e.g., Table 2 of Fiore et al., 2024) require new laboratory studies, enhancements to current satellite capabilities as well as a more comprehensive suite of suborbital datasets (*Section 3*; e.g., Murray et al., 2021) could be used to develop process-based diagnostics to better constrain the factors that influence OH drivers, such as to constrain the response of O_3 to ENSO (Oman et al., 2013). Such diagnostics would inform development priorities of CTMs with the goals of improving the simulation of tropospheric OH as well as aiding interpretation of inferred OH from satellite-based approaches (*Section 2*).

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5.4 Addressing "Clear-Sky" Bias of Inferred OH from Satellite-Based Approaches

The presence of clouds limits a satellite's ability to observe tropospheric composition, especially in seasonally cloudy regions. King et al. (2013) used MODIS cloud fraction data to estimate that ~67% of Earth's surface is covered by clouds at any given time. Consequently, temporal averaging (e.g., over a season) of the data is often employed to achieve statistical significance, though the averages are biased toward clear-sky/low-cloud conditions. We found that these biases in simulated OH (from one CTM) relative to all-sky conditions are substantial (e.g., +/-25% in the tropospheric OH VCDs in the tropics and higher elsewhere; *Figure 4*). One way to reduce these biases is to improve the spatial resolutions of satellite instruments, which would increase the chances of collecting observations between broken clouds and, thus, minimize the number of cloud-contaminated pixels (e.g., Frankenberg et al., 2024). There may be viable ways to infer some of the OH drivers over clouds (e.g., NO_2; Marais et al., 2018; Marais et al., 2021), but further study is needed.

We recommend additional research to identify ways to gap fill OH inferred from the satellite-based approaches (Section 2). Potential ways are to (1) use simulated OH, which would benefit from model development priorities as informed by satellite-constrained, process-based diagnostics (*Section 5.3*), (2) scale inferred OH by the ratio of observed OH in clear and cloudy conditions from tropospheric composition field campaigns (*Section 3*), and (3) use a photolysis model with satellite-observable cloud properties (e.g., cloud optical depth, cloud top pressure, and cloud fraction) to scale inferred OH. The problem with the second way is that cloudy regions are typically avoided during field campaigns. Therefore, a priority of future field campaigns is to collect observations of OH and its drivers in both clear and cloudy environments. Finally, there is the potential to develop an ML model using simulated tropospheric OH with meteorological variables and climate indices (*Section 5.3*), which are observable from space, to understand the relationship of OH concentrations, trends and variations in cloudy and mostly cloudy environments. All of these potential methods to gap fill inferred OH require further study.

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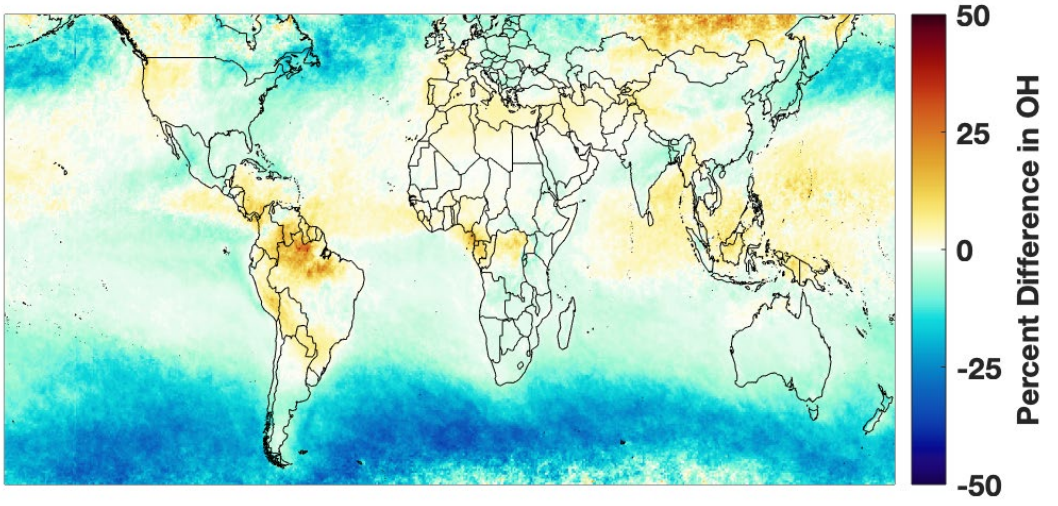

Figure 4: Simulated percent difference in tropospheric column OH between low-cloud (cloud fraction less than 30%) and all-sky conditions. Red indicates higher OH abundance in near clear sky conditions; blue indicates lower OH abundance under such conditions. Simulated data are from the MERRA2 GMI simulation and averaged over 2005 – 2019 for October.

6 Summary of Recommendations

The satellite-based approaches (*Section 2*) offer the promise that we may achieve a constraint of tropospheric OH on spatio-temporal scales that has been previously unobtainable. Given this promise, we make recommendations to facilitate the design of a robust observing strategy of satellite and suborbital observations of atmospheric constituents and dynamical variables to indirectly constrain the concentrations, trends, and variations of tropospheric OH. The increased spatio-temporal information of tropospheric OH that could be gained by following these recommendations would advance our ability to attribute the causes of trends and variations of the concentrations of atmospheric constituents that are influenced by OH. For instance, Turner et al. (2016) concluded that "the current surface observing system does not allow unambiguous attribution of the decadal trends in methane without robust constraints on OH variability, which currently rely purely on methyl chloroform data and its uncertain emissions estimate."

We recommend the development of a comprehensive satellite-based approach to maximize the indirect constraint of tropospheric OH that may be composed of a combination of existing satellite-based approaches (*Section 2*) and new satellite-based approaches developed for specific photochemical environments. Studies are needed to determine the optimal combination of input variables/satellite observables for some of the satellite-based approaches to maximize the interpretability (in a process-based sense) of the causes of trends and variations in tropospheric OH. The design process would benefit from tradespace and cost-benefit analyses, including consideration of incorporating more cost-effective suborbital observations, where feasible.

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Priorities (in no particular order) for this observing strategy include:

- continuity of satellite instruments with similar or enhanced capabilities,
- improvements to accuracy/precision and SNR of current observing capabilities through instrument design,
- additional LEO and GEO satellites to obtain greater spatio-temporal coverage,
- development of a robust suborbital observing network, composed of instruments at stationary sites (OH "supersites") as well as ones deployed in field campaigns, to collect suborbital observations of OH drivers, including in cloudy environments, and
- coordinated deployment of UV, Vis, NIR, and TIR instruments to provide co-located observations of satellite proxies of key OH drivers.

We also recommend additional investment in research (in no particular order) to:

- maximize the vertical information of the observations of OH drivers (e.g., with multispectral data products),
- maintain self-consistent and trend-quality L1B radiances for the generation of multi-agency, multi-instrument data records,
- determine the full potential of using satellite-observable meteorology (e.g., temperature, clouds, aerosols) and climate variables in space-based approaches to better understand the dependence of OH concentrations, trends, and variations on dynamics,
- develop process-based diagnostics using satellite and suborbital observations to improve the representation of key atmospheric processes in CTMs that influence OH,
- explore the potential of observable VOCs to aid constraint of tropospheric OH, and
- address the "clear-sky" bias of inferred OH from satellite-based approaches.

Data Availability

Figure 1: Data are available upon request from D. Anderson (daniel.c.anderson@nasa.gov).
Figure 2: Satellite retrievals for TEMPO HCHO V3 data can be found at https://doi.org/10.5067/IS-40e/TEMPO/HCHO_L2.003, the radiance v3 data at https://doi.org/10.5067/IS-40e/TEMPO/HCHO_L2.003, and the irradiance v3 data at https://doi.org/10.5067/IS-40e/TEMPO/IRR_L1.003 (González Abad et al., 2024).
Figure 3: Data from the ATom campaign are located at https://doi.org/10.3334/ORNLDAAC/1925 (Wofsy, 2021).
Figure 4: Output from the MERRA-2 GMI simulation is publicly available at https://acd-ext.gsfc.nasa.gov/Projects/GEOSCCM/MERRA2GMI/ (NASA Goddard Space Flight Center, 2024).

Author Contributions

BND initiated the opinion piece and wrote the first draft. All co-authors contributed to the overall content and direction of the manuscript through many collaborative discussions. Many co-authors contributed some text and all figures.

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Competing Interests

BND is a member of the editorial board of Atmospheric Chemistry and Physics. The authors also have no other competing interests to declare.

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