# 1 Ozone trends in homogenized Umkehr, Ozonesonde, and COH

# 2 overpass records

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Abstract. This study presents an updated evaluation of stratospheric ozone profile trends at 19 20 Arosa/Davos/Hohenpeißenberg, Switzerland/Germany, Observatory de Haute Provence (OHP), France, Boulder, 21 Colorado, Mauna Loa Observatory (MLO) and Hilo, Hawaii, and Lauder, New Zealand with focus on the ozone 22 recovery period post 2000. Trends are derived using vertical ozone profiles from NOAA's Dobson Network via the 23 Umkehr method (with a recent new homogenization), ozonesondes, and the NOAA COHesive SBUV/OMPS satellite-24 based record (COH) sampled to match geographical coordinates of the ground-based stations used in this study. 25 Analyses of long-term changes in stratospheric ozone time series were performed using the updated version (0.8.0) of 26 the Long-term Ozone Trends and Uncertainties in the Stratosphere (LOTUS) Independent Linear Trend (ILT) 27 regression model. This study finds a consistency of the trends derived from the different observational records, which 28 is a key factor to the understanding of the recovery of the ozone layer after the implementation of the Montreal Protocol 29 and its amendments that control ozone-depleting substances production and release into the atmosphere. The Northern 30 Hemispheric Umkehr records of Arosa/Davos, OHP, and MLO all show positive trends in the mid to upper 31 stratosphere with trends peaking at ~+2%/decade. Although the upper stratospheric ozone trends derived from COH 32 satellite records are more positive than those detected by the Umkehr system, the agreement is within the two times 33 standard error uncertainty. Umkehr trends in the upper stratosphere at Boulder and Lauder are positive but not 34 statistically significant, while COH trends are larger and statistically significant (within 2 times standard error 35 uncertainty). In the lower stratosphere, trends derived from Umkehr and ozonesonde records are mostly negative 36 (except for positive ozonesonde trends at OHP), however the uncertainties are quite large. Additional dynamical proxies were investigated in the LOTUS model at five ground-based sites. The use of additional proxies did not 37

38 significantly change trends, but equivalent latitude reduced the uncertainty of the Umkehr and COH trends in the

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upper stratosphere and at higher latitudes. In lower layers, additional predictors (tropopause pressure for all stations,
 two extra components of Quasi-Biennial Oscillation at MLO, Arctic Oscillation at Arosa/Davos, OHP and MLO)

49 improve the model fit and reduce trend uncertainties as seen by Umkehr and sonde.

#### 50 1 Introduction

51 The WMO Ozone Assessments (WMO, 2018; WMO, 2022), indicate that for some geographical regions, the 52 stratospheric ozone layer is recovering in accordance with the reduction of ozone depleting substances (ODS) whose 53 production was restricted by the Montreal Protocol and its amendments. The US Clean Air Act requires NOAA to 54 monitor prohibited chemicals and the ozone layer to ensure the success of the Montreal Protocol. NOAA's long-term network of measurements helps to interpret total column and vertically resolved ozone changes and link ozone 55 56 recovery to the reduction of ODS levels in the stratosphere, changes in the lower stratosphere that are associated with 57 climate changes, and to the increases in the troposphere that are influenced by the stratosphere/troposphere exchange 58 and long-range transported pollution. The ongoing recovery of the stratospheric ozone layer is of great importance to 59 human health (i.e. cancer from enhanced UV exposures, Madronich et al., 2021), the sustained production of crops, 60 and the success of fisheries (dangerous algae blooms). For more information see the Environmental Effects 61 Assessment Panel 2022 Quadrennial Assessment (EEAP, 2023). 62 The Long-term Ozone Trends and Uncertainties in the Stratosphere (LOTUS) study was initiated under Stratosphere-63 troposphere Processes And their Role in Climate (SPARC, changed to APARC in 2024) project to reconcile the 64 differences in defining trend uncertainties between methods outlined in the WMO Assessment (WMO, 2014) and the 65 SPARC/IO3C/IGACO-O3/NDACC (SI2N) study (Harris et al., 2015). Phase 1 focused on developing best practices

66 for data merging, trend determination and error analyses. Results focused on analysis of broad latitudinal regions, 67 near global, Northern and Southern Hemisphere, and Tropics as were used in the SI2N studies. Results are found in 68 the 2019 report (SPARC/IO3C/GAW, 2019). Phase 2 refined the trend models, and extended the study to gridded, and GB ozone data sets. The development of methods used in trend detection is built on the community knowledge 69 gained during the Tiger Team project in early 1990s (Reinsel et al., 2005), collaborations through the SPARC, WMO 70 71 and IO3C supported LOTUS activity (Hassler et al., 2014; Harris et al., 2015; Godin-Beekmann et al., 2022) and the 72 most recent contributions to the WMO Ozone assessment analyses published in Chapter 3, "Update on Global Ozone: 73 Past, Present and Future" (Hassler et al., 2022).

74 Understanding the causes of the differences between GB and satellite records can create improvements not only in the 75 internal consistency of data sets, but also in the uncertainties of overall ozone trends. Further, development of techniques to directly assess uncertainties in the merged records resulting from discrepancies that cannot be completely 76 77 reconciled, such as small relative drifts and differences resulting from coordinate transformations and sampling 78 differences, allows for a more precise estimate of significance of the mean trends. For the GB and satellite data used 79 in the 2019 LOTUS Report, information on stability and drifts of the measurement was incomplete. The 80 homogenization of many ozonesonde records was recently addressed and data were reprocessed (Tarasick et al., 2016; Van Malderen et al., 2016; Witte et al., 2017; Sterling et al., 2018; Witte et al., 2018; Ancellet et al., 2022) while some 81 82 instrumental artifacts still need to be addressed (Smit, 2021).

Deleted: Studies of ozone recovery require long-term datasets often consisting of data merged from several instruments, or from a single instrument type on multiple statellite platforms, or at a groundbased (GB) station. 2011 saw the initiation of the SPARC/IO3C/IGACO-O3/NDACC (SI2N) activity to evaluate ozone trends in the depletion and recovery phases using both GB and merged satellite observations. The resulting report (Harris et al., 2015) emerged near the release of the 2014 WMO Ozone Assessment (WMO, 2014). The two studies resulted in broadly similar trend values in both the depletion and recovery phases. But the WMO report determined the recovery trends to be statistically significant, whereas the SI2N study did not. The discrepancy centered on differing error analysis techniques for the merged datasets: SI2N using distribution of the individual variances around the mean and WMO using weighted mean of the individual standard deviations.¶

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101 The first attempt to evaluate representativeness of the trends derived from GB station records in the middle and upper 102 stratosphere using SBUV data was done as a part of the LOTUS activity and was discussed in the 2019 LOTUS report. Comparisons of trends derived from satellite data sub-sampled at the station location (overpass) to those derived from 103 the relevant zonal average provide a measure of potential sampling errors when comparing satellite and GB trends 104 105 (Zerefos et al., 2018; Godin-Beekmann et al., 2022). This paper continues that work by comparing trends derived from 106 several GB and satellite records that are matched spatially. We further investigate the impact of temporal matching on 107 trends. 108 The common statistical linear regression trend model used in the 2019 LOTUS Report and the 2022 update (Godin-

109 Beekmann et al., 2022) was optimized for analyses of the zonally averaged satellite data sets. However, analyses of 110 the GB and satellite overpass ozone profile data may require reconsideration of additional proxies and optimization 111 methods to improve interpretation of the processes that impact ozone changes over limited geophysical regions and 112 reduce trend uncertainties. An assessment of model sensitivities to uncertainties in the volcanic aerosols, solar cycle, 113 QBO, El Nino Southern Oscillation (ENSO) and other mechanisms also need to be considered in the GB and satellite 114 overpass record trend analysis. The localized time series for the assessment of dynamical and chemical proxies can 115 improve attribution of ozone variability, especially in the lower stratosphere, thus reducing uncertainties in the derived 116 trends. This paper provides an assessment of uncertainties in the derived trends from the NOAA ground-based, 117 ozonesonde and SBUV/OMPS (zonally averaged and overpass) records and reports improvements in the Multiple 118 Linear Regression (MLR) trend uncertainties with addition of proxies representing interannual dynamical variability 119 or long-term changes in atmospheric circulation. Ability of the ground-based and ozonesonde records in capturing 120 semi-global ozone changes is evaluated by comparing trends derived from the satellite overpass and zonally averaged 121 records

In the LOTUS report, the ozone trends were analyzed at low and middle latitudes, with a focus on the upper and middle stratosphere. This paper includes middle and low latitude trends assessed in the lower stratosphere and thus offers an opportunity to test the additional proxy of the tropopause pressure (Thompson et al., 2021).

#### 125 2 Data

#### 126 2.1 Umkehr and Ozonesonde Records at NOAA

127 The Dobson Ozone Spectrophotometer has been used to study total ozone since its development in the 1920s 128 (Staehelin et al., 2018). Dobson records are regularly used in satellite record validation (Bai et al., 2015; Koukouli et 129 al., 2016; Boynard et al., 2018) and the development of global combined ozone data records (Fioletov et al., 2008; 130 Hassler et al., 2018). The NOAA Dobson ozone record was homogenized in 2017 to account for inconsistencies in 131 past calibration records, data processing methods and selection of representative data (Evans et al., 2017). NOAA 132 Dobson instruments at 4 stations and MeteoSwiss at Arosa/Davos also measure Umkehr profiles, which are derived as partial column ozone amounts in ~5 km layers. Profiles are derived using an optimum statistical inversion of Dobson 133 134 measurements taken continuously at different solar zenith angles (SZAs) (Petropavlovskikh, 2005; Hassler, 2014). 135 These Umkehr data were recently homogenized to assure the removal of small but significant instrumental artifacts 136 that can impact the accurate detection of stratospheric ozone trends (Petropavlovskikh et al., 2022, Maillard Barras et 137 al., 2022). This study focuses on Umkehr records from the MeteoSwiss station of Arosa/Davos, Switzerland, and on 138 Umkehr records from the NOAA stations of Boulder, Colorado, Mauna Loa Observatory (MLO), Hawaii, Lauder, 139 New Zealand, and the Umkehr record from Observatory de Haute Provence (OHP), France, NOAA/GML for Umkehr 140 data means that the NOAA optimization process was applied to the operational records (N-values) prior to the retrieval 141 of ozone profiles. The source data used in this study are available at 142 https://gml.noaa.gov/aftp/data/ozwv/Dobson/AC4/Umkehr/Optimized/. See Table 1 for details on the GB datasets, 143 locations, source of data and temporal extent of data used. Umkehr measurements are typically made twice per day 144 when there is no cloud obstruction.

145 The ozonesonde instrument has been flown at 4 NOAA stations since the 1980s. Evolving instrumentation and 146 standard operating procedures led to the development of data homogenization methods by NOAA and the international 147 community (i.e., ASOPOS-1, Smit, 2014) to resolve record inconsistencies in the NOAA (Sterling et al., 2018), 148 Canadian (Tarasick et al., 2016) and SHADOZ (Southern Hemisphere Additional Ozonesondes) networks (Witte et 149 al., 2017; Witte et al., 2018). The effort was extended in the ASOPOS-2 (Smit et al., 2021) activity and included a 150 larger group of stations that are part of the NDACC (Network for Detection of Atmospheric Composition Change) 151 and WMO GAW (World Meteorological Organization Global Atmosphere Watch program) networks. The error 152 budget for each profile is calculated and included in the archived files (Sterling, 2018). Modern ozonesonde 153 instruments measure ozone at high vertical resolution, on the order of 100 m (Thompson et al., 2019) depending on 154 the balloon accent velocity and the time response of the instrument, 155 The sondes constitute an essential component of satellite calibration and cross-calibration (Hubert et al., 2016), 156 verification and improvement of climate chemistry and chemistry-transport models (Wargan et al., 2018; Stauffer et

verification and improvement of chinae chemistry and chemistry-transport models (wargan et al., 2018, statuter et al., 2019). The Dobson total ozone, Umkehr and ozonesonde profile records provide key measurements for upper and
 middle stratospheric ozone trend calculations, and are part of the NOAA benchmark network for stratospheric ozone
 profile observations (SPARC/IO3C/GAW, 2019; Godin-Beekmann et al., 2022; WMO, 2022).

160 The ozonesonde data are used for trend analyses from OHP, Boulder, and Lauder stations where we have Umkehr 161 observations. Ozonesondes are launched at Hilo, Hawaii, which is nearly co-located with MLO. Ozonesonde data

161 observations. Ozonesondes are launched at Hilo, Hawaii, which is nearly co-located with MLO. Ozonesonde data 162 for the Arosa/Davos panel are selected from Hohenpeißenberg (HOH), Germany station that is in close vicinity to

163 Arosa/Davos station. Sonde measurements are typically measured once or twice per week, varying somewhat with

164 station operational procedures.

165 Data for the NOAA GML ozonesonde records are publicly available from the NOAA Global Monitoring Lab (GML)

 $166 \qquad at \ https://gml.noaa.gov/aftp/data/ozwv/Ozonesonde/. \ We \ use \ the \ `100 \ Meter \ Average \ Files' \ in \ each \ station \ directory.$ 

167 Other sonde datasets used in this study are also available at several other data centers including the World Ozone and

168 Ultraviolet Radiation Data Centre (WOUDC, www.woudc.org), Network for the Detection of Atmospheric

169 Composition Change (NDACC, www.ndacc.org) data centers, or at the Harmonization and Evaluation of Ground-

based Instruments for Free-Tropospheric Ozone Measurements (HEGIFTOM, <u>https://hegiftom.meteo.be/</u>) archive.
Table 1 denotes the source of each dataset used in this study.

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The ozonesonde data is of significantly higher vertical resolution (even when used as 100 m averages) than the Umkehr 176 data layers of approximately 5000 m. In order to create a dataset with comparable resolution, we use the Umkehr 177 178 averaging kernels (AK) to smooth the sonde data. Details appear in Appendix A. We cap the sonde profile at Umkehr 179 layer 5 (16-32 hPa) as there is not sufficient sonde information at higher altitudes to meet the requirements of the AKs 180for layers 6 and above. We further match the ozonesonde data to the dates when both Umkehr and sonde data are 181 available using  $\pm 24$  hours to find a match, then generate the ozonesonde monthly mean. Appendix D explores the 182 impact of temporal sampling on trends. The final matched dataset, with AK averaging, is publicly available at 183  $\underline{https://gml.noaa.gov/aftp/ozwv/Publications/2023\_Umkehr\_Ozone\_Trends\_Paper/.$ 

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Location	WOUD C Station #	Instrument	Date Range used in trend calculations	Source
Arosa/Davos Arosa, Switzerland (46.8° N, 9.7° E) Davos, Switzerland (46.8° N, 9.8° E)	035	Umkehr	1980 – 2018 2018 – 2020	Optimization by NOAA/GML
Hohenpeißenberg (HOH), Germany (47.8° N, 11.0° E)	099	Ozonesonde	1980 – 2020	NDACC
Observatory de Haute Provence (OHP), France (43.9°N, 5.8°	040	Umkehr	1983 - 2020	NOAA/GML
E)		Ozonesonde	1991 - 2020	NDACC <sup>=</sup> (same as HEGIFTOM)
Boulder, Colorado (40.0° N, 105.3° W)	067	Umkehr	1980 - 2020	NOAA/GML
(+0.0 h, 105.5 w)		Ozonesonde	1980 - 2020	NOAA/GML - 100m average data
Mauna Loa Observatory (MLO), Hawaii (19.5°N, 155.6°W)	031	Umkehr	1982 - 2020	NOAA/GML
Hilo, Hawaii (19.7° N,155.1° W)	109	Ozonesonde	1982 - 2020	NOAA/GML - 100m average data

	Lauder, New Zealand (45.0°S, 169.7°E)	256	Umkehr	1987 - 2020	NOAA/GML
			Ozonesonde	1987 – 2020	NDACC
5	*Note, data from the "Corrected	Ozone partial p	ressure" column is used	d for trend analyses	

Table 1: GB datasets, location, instrument type, temporal extent, data record source. For the trend calculations we remove data during volcanic periods from 1982–1984 and 1991–1994.

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## 8 2.2 The NOAA Cohesive (COH) Station Overpass Ozone Profile Datasets

190 NASA and NOAA have produced satellite measurements of ozone profiles through the Solar Backscatter Ultraviolet 191 (SBUV) on the sequence of Polar Orbiting Environmental Satellites (POES) since 1978. This measurement series is 192 extended with the related Ozone Mapping and Profiler Suite (OMPS) nadir profiler (NP) instruments using similar 193 measurement techniques and retrieval algorithms. These combine to provide nearly 45 years of continuous data (1978 194 - present). This single instrument type dataset eliminates many homogeneity issues including varying vertical resolution, or instrumentation differences. Version 8.6 SBUV data incorporates additional calibration adjustments 195 196 beyond the Version 8 release (McPeters et al., 2013). Small but evident biases remain (Kramarova et al., 2013a). 197 Several methods have been historically used to combine these datasets into a continuous series. The NASA MOD 198 version 1 dataset based on SBUV and OMPS v8.6 (Frith, 2014) combines data from all available satellites with no 199 modification or bias adjustments. NASA has developed an alternate processing for the SBUV and OMPS data (v8.7) 200 which incorporates new calibrations at the radiance level, and updated a priori with improved troposphere. 201 Additionally, the a priori is chosen to be representative of the local solar time of the measurement. MOD v2 is based 202 on the v8.7 processing (Frith et al., 2020), and further applies an adjustment to the v8.7 data to shift all measurements to a nominal measurement time of 1:30 PM local time. 203 204 The NOAA SBUV/2 and OMPS Cohesive dataset (further referred to as COH) combines data from the SBUV/2 and OMPS instruments using NASA's version 8.6 for the SBUV/2 data and NOAA/NESDIS version 4r1 for the OMPS 205 206 Suomi National Polar-orbiting Partnership (SNPP) data. This dataset uses correlation-based adjustments providing 207 an overall bias adjustment plus an ozone-dependent factor (Wild et al., 2016) to moderate the remaining biases 208 between instruments in the series. The resulting profile product is a set of daily or monthly zonal means and is publicly available at https://ftp.cpc.ncep.noaa.gov/SBUV\_CDR. Zones are 5° wide in latitude, identified by the central latitude 209  $(2.5^{\circ}, 7.5^{\circ}, \text{etc.})$ . Contributing satellites and their period of use is shown in Table 2. 210

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Satellite	Dates
Nimbus 07	10/1978 - 5/1989
NOAA 11 (ascending)	6/1989 - 12/1993
NOAA 09	1/1994 - 6/1997
NOAA 11 (descending)	7/1997 - 12/2000

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NOAA 16	1/2001 - 12/2003
NOAA 17	1/2004 - 12/2005
NOAA 18	1/2006 - 12/2010
NOAA 19	1/2011 - 12/2013
SNPP	1/2014 - present

212 Table 2: Satellite mapping for COH data series.

213 A previous version of this dataset using OMPS v3r2 has been used in climate reviews and trend studies (Godin-

214 Beekmann et al., 2022; Weber et al., 2022a, 2022b) including Chapter 3 of the WMO Ozone Assessment (Hassler et

al., 2022). Appendix B examines the differences between the data versions. The impact on trends is limited to lessthan 1% per decade, well within the precision of the trend results.

217 We create the overpass data at a ground station by collecting all profiles from a satellite within a  $\pm 2/20^{\circ}$ 

218 latitude/longitude box centred on the station. The box size is chosen to ensure that one to four points are found per

219 day. Fewer points are found if the orbit passes directly over the station; more points are found if the orbits straddle

the station. The collected profiles are inverse-distance weighted to the station location and averaged. COH style adjustments are applied (Wild et al., 2016) creating a COH overpass time series from 1978 to present. This dataset is

222 available on the NOAA website at https://ftp.cpc.ncep.noaa.gov/SBUV\_CDR/overpass.

Figure 1 shows the ozone anomaly time series for the 40–45° N zonal average data, and for the data at 3 stations in or near the zone. Anomalies are calculated with respect to the zonal average climatology. The series shown are for the layer data with the bottom pressure of the layer displayed on the right side of the graph. This depiction retains the information about the relative differences between the stations and the zonal average. In the mid-stratosphere (25–10 hPa) the biases between the stations are most pronounced with Arosa/Davos usually showing less ozone and Boulder

usually showing more ozone. At the uppermost layers (1 and 4 hPa), and the lowest layer (41 hPa) the bias between

stations is reduced. The anomalies for Arosa/Davos and OHP, which are geographically closer than Boulder, are often

230 nearly anticorrelated with the Boulder anomalies especially in the second half of the year. Indeed at 16 hPa in

particular, one can see that often the Boulder anomalies are positive when the Arosa/Davos and OHP anomalies are negative.

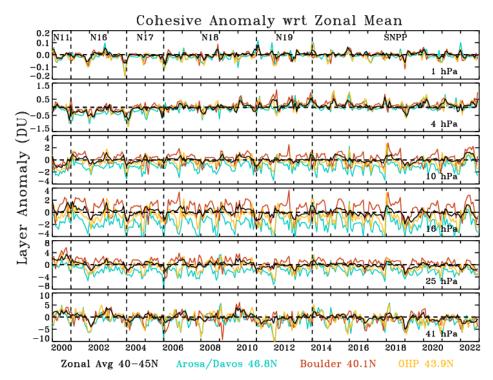
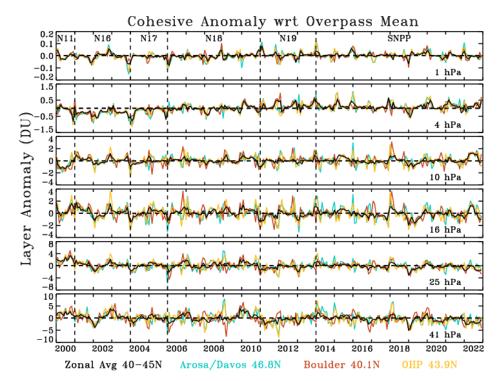


Figure 1. Monthly ozone anomaly relative to the zonal mean monthly averages. This process leaves intact the trend for each site and the zone, and accentuates the differences between the station values since all anomalies are referenced to the zonal product. Evident at 4 hPa is a positive trend from 2002 to 2013, then a levelling out after.

238 Figure 2 also shows the anomalies for the  $40-45^{\circ}$  N zonal average with the station anomalies, but each anomaly is 239 now created using the climatology derived from each separate dataset. This removes the bias between the stations and 240 the zones. At 1 hPa, Arosa/Davos appears to display the most variation (largest peaks and dips) in the anomalies. 241 Since the anomaly for each site is now based on the seasonality of each site's data the structure in the anomalies is 242 more uniform. For example, now at 16 hPa, the difference between Boulder and the two sites Arosa/Davos and OHP in the latter half of the year is removed. In 2012, where the Boulder anomaly was positive with respect to the zonal 243 average seasonal value, and the Arosa/Davos and OHP sets were negative with respect to the zonal seasonal average, 244 245 all are now of the same sign with respect to their own seasonal averages. Nonetheless, there are events where one station shows opposite anomalies to the other two, for example early 2009 at 41 hPa, when the Boulder anomaly is 246 247 negative, and Arosa/Davos and OHP are positive. Thus, it is noted that when comparing daily or monthly data values 248 from GB and satellite data, the overpass data will reveal a different structure than the zonal data. The trend calculations 249 in this paper are based on the datasets of Fig. 2, where the seasonal behavior is removed using the station seasonality. 250



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Figure 2. Monthly ozone anomaly relative to the monthly climatology for each station overpass dataset. This process leaves intact the trend for each site and the zone, and shows the consistency among the stations when each station climatology is removed. This dataset is used for the trends calculations. Evident at 4 hPa is a positive trend from 2002 to 2013, then a leveling out after. Trends are run on this dataset.

256 The COH overpass and zonal datasets have a similar vertical granularity as the Umkehr dataset, but use somewhat 257 different pressures for the demarcation of the top and bottom of each layer. Since no additional smoothing is required, 258 we simply use interpolation and integration to convert the COH layer profiles to the Umkehr layers. We exclude layers 259 1 to 4 since there is little sensitivity in SBUV and OMPS NP in these layers (Kramarova et al., 2013b). The overpass 260 monthly-mean dataset in this study uses all COH data matched to dates when Umkehr data also exists. This dataset is 261 publicly available at https://gml.noaa.gov/aftp/ozwv/Publications/2023\_Umkehr\_Ozone\_Trends\_Paper/. Appendix 262 D explores the impact of temporal sampling on trends. 263 This study also uses a specialized zonal monthly-mean COH product which is the average of all daily profiles with an Umkehr match at the associated GB station. Zones used for most stations are the 5° wide zone which includes the 264 265 geographic station latitude (Arosa/Davies: 47.5° N, OHP: 42.5° N, MLO: 17.5° N). Boulder and Lauder, however,

are located directly on the border of two zones, so the zonal product in this study is the mathematical average of the

 $267 \qquad \mbox{two adjacent zones (Boulder: $37.5^\circ$ N and $42.5^\circ$ N, Lauder $42.5^\circ$ S and $47.5^\circ$ S)}.$ 

#### 3 Methods 268

#### 3.1 LOTUS Model overview - the Reference Model 269

270 The Long-term Ozone Trends and Uncertainties in the Stratosphere (LOTUS) activity is a project of SPARC (Stratosphere-Troposphere Processes and their Role in Climate) and has produced a statistical Multiple Linear 271

272 Regression (MLR) model called the LOTUS model (https://usask-arg.github.io/lotus-regression/index.html).

273 The 2019 LOTUS report (SPARC/IO3C/GAW, 2019) and update (Godin-Beekmann et al., 2022) have quantified

274 stratospheric ozone trends and evaluated their uncertainties. The LOTUS model is a general-least-squares approach

275 MLR model. This study uses version 1 (v 0.8.0) with the independent linear trends (ILT) configuration. The

276 independent linear terms represent the ozone depletion period (pre-1997), the ozone recovery period (post-2000) and

an optional gap period (1997-2000). We will call the terms "pre", "post" and "gap" for short. The version 0.8.0 adds 277

278 an option to enforce continuity across the gap period which is used in this study. The regression uses an interactive

279 procedure (Cochrane and Orcutt, 1949) and the autocorrelation coefficient is adjusted with each iteration. The

280 covariance matrix is modified accordingly to account for measurement gaps (Savin and White, 1978).

281 The LOTUS model (further referred as reference model in this study) is written here:

$$\frac{y(t,z) = \beta_0(t,z)C_{pre}(t) + \beta_1(t,z)C_{post}(t) + \beta_2(t,z)Linear_{pre}(t) + \beta_3(t,z)Linear_{post}(t) + \sum_{i=4}^{n}\beta_iX_i(t,z) + \epsilon(t,z)$$
(1)

where  $\hat{y}(t, z)$  is the estimated ozone at time t and altitude z;  $\beta$  are the fitted coefficients of the model; the residual 283

term,  $\varepsilon(t, z)$  is the difference between the LOTUS model and the input data.  $C_{pre}$  and  $C_{post}$  are the constant terms as 284 285 defined by:

$$Constant_{pre} = \begin{cases} 1 & \text{for } t < 1997\text{-Jan} \\ 1 - mt & \text{for } 1997\text{-Jan} \leq t < 2000\text{-Jan} \\ 0 & \text{for } t \geq 2000\text{-Jan} \end{cases}$$

$$Constant_{post} = \begin{cases} 0 & \text{for } t < 1997\text{-Jan} \\ mt & \text{for } 1997\text{-Jan} \le t < 2000\text{-Jan} \\ 1 & \text{for } t \ge 2000\text{-Jan} \end{cases}$$
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288 where  $\underline{m=0.029135}$  and t = month starting in January 1980 and ending in December 2020. Indeed, the constant terms 289 are only constant in the "pre" and "post" periods. The 3-year "gap" period is represented by a line of slope m 290 connecting the two constant (pre and post period bias) terms.

291 The linear terms of the model are defined as:

292
$$linear_{pre} = egin{cases} mt-b & ext{for} \ t < 1997 ext{-Jan} \\ 0 & ext{for} \ t \geq 1997 ext{-Jan} \end{cases}$$

$$_{293} linear_{post} = egin{cases} 0 & ext{for } t < 2000 ext{-Jan} \ mt & ext{for } t \geq 2000 ext{-Jan} \end{cases}$$

294 where m=0.008487, b = -1.700240, and t = month starting in January 1980 and ending in December 2020.

for  $t \geq 1997$ -Jan

295 Natural variability is a complicating factor in deriving trends associated with the changes in the ozone depleting

296 chemistry. LOTUS fits predictor variables as proxies for natural variability to the ozone data so that one can interpret

297 the resulting linear trend as a trend due to the changes in chemistry. The summation term is the summation of the 298 predictors used as a proxy for the dynamical induced ozone variability.

The natural variability proxies in the LOTUS model v 0.8.0 are Aerosol Optical Depth (AOD), El Nino/ Southern 299 300 Oscillation (ENSO), and the Quasi-Biennial Oscillation (QBO) in the form of the first two principal components (also known as an empirical orthogonal function analysis). The data sources for each are described in Table 3. 301 302 Large sulfur dioxide (SO2) Jevels reaching the lower stratosphere following major volcanic eruptions (i.e. El Chichon, Pinatubo or Hunga) can impact the validity of ozonesonde, measured values (Yoon et al., 2022). However, SO2 is not 303 304 long-lived gas and is soon converted to sulphate aerosols that can alter observations by ozone remote sensing systems. 305 Both Umkehr and satellite ozone profiles from SBUV and OMPS are highly uncertain and/or biased because of high 306 aerosol load during volcanic eruptions (DeLuisi et al, 1989; Petropavlovskikh et al., 2005, 2022; Bhartia et al, 1993, 307 Torres et al., 1995, Bhartia et al, 2013). It is recommended that the data for 2 to 3 years after the El-Chichon and 308 Pinatubo large volcanic eruptions should not be used in trend analyses. Therefore, we exclude data during the volcanic 309 periods (1982-1983 and 1991-1993) from the analyzed time series. Moreover, this study is focused on the linear trend 310 analyses after 2000 when there are no large stratospheric aerosol perturbations that significantly influence 311 stratospheric ozone variability over the middle latitudes and therefore impact trend and uncertainty estimates. Since 312 we have eliminated the data during the volcanic period, this study does not include the AOD proxy in the calculations. We define the 'reference' model (RM) as the proxies most commonly used for the dynamical proxies which is 313 314 equivalent to the LOTUS model v 0.8.0 minus the AOD term. The representative equation is:

 $\sum_{i=4}^{n} \beta_i(t,z) X_i(t) = \beta_4(t,z) QBO_A(t) + \beta_5(t,z) QBO_B(t) + \beta_6(t,z) ENSO(t) + \beta_7(t,z) Solar(t)$ (2)

The Quasi-biennial Oscillation (QBO) is derived from the Singapore radiosonde profiles (1979–2020) that detect variability in the direction of the tropical winds in the lower stratosphere. It also shows that zonal wind variation propagates downward with an average period of ~28 months [Wallace, 1973]. The principal component analysis of the 100–10 hPa zonal winds can describe the majority of the wind variability. The reference model (and LOTUS v 0.8.0) use the two leading modes of the calculated empirical orthogonal functions (EOF) for trend analyses [Wallace et al., 1993].

322 The El Niño/ Southern Oscillation (ENSO) is a periodic mode of climate variability of the atmosphere and sea surface 323 temperatures associated with the equatorial Pacific Ocean with periods ranging from 2-8 years. The Multivariate 324 ENSO index (MEI) is produced by the NOAA Physical Sciences Laboratory and is derived from the EOF analysis of 325 sea surface temperature, sea level pressure, outgoing terrestrial radiation, and surface winds in the area of the Pacific 326 basin from 30° S to 30° N and from 100° E to 70° W (Wolter and Timlin, 2011). Temperature anomalies in the 327 troposphere with corresponding stratospheric temperature anomalies during El Niño/ La Niña events modulate the 328 tropical upwelling of the Brewer-Dobson circulation (BDC) and thus the meridional transport of ozone in the 329 stratosphere. (Diallo et al., 2018).

The solar cycle is the 11-year periodic cycle of solar activity and solar irradiance that reaches the Earth's atmosphere.
The change in UV radiation that is absorbed by the atmosphere, most notably in the upper stratosphere, leads to
changes in atmospheric temperature and the photochemistry which produces ozone. (Lee and Smith, 2003). The 10.7
cm solar radio flux data is used as the proxy for the solar cycle in the LOTUS model.

334 Seasonal components in the form of Fourier harmonics were added into the LOTUS model with version 0.8.0. Godin-

Beekman et al. (2022) showed in their Fig. 7 that the model fit for the ozone profile satellite and model records is

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340 improved by adding seasonal components to the proxies, increasing the adjusted R-squared ( $R_{s}^{2}$ ) from 0.3 or less to

341 0.3 to 0.5. The seasonality and relevant contributions of some predictor's variables are compensated in this study by

342 adding the seasonal components to the fitted predictors. Seasonal components are represented in the model by sine

- 343 and cosine functions with periods of 12 and 6 months that describe the variability of the proxies on these timescales.
- 344 So, for each fitted predictor in the model

 $eta_i X_i(t,z) ext{ where } i>1$ 

345

346 a seasonal variation in the form of Fourier components is added as follows:

$$\beta_m(t,z) = \beta_{m,0}(z) + \sum_{i=1}^2 \beta_{m,1,i}(z) \sin(\frac{2\pi i t}{12}) + \sum_{i=1}^2 \beta_{m,2,i}(z) \cos(\frac{2\pi i t}{12})$$

#### 348 3.2 The Extended Model - Adding Predictors

349 Recent publications (i.e. Petropavlovskikh et al., 2019; Szelag et al., 2020; Godin-Beekmann et al, 2022; Millan et al. 2024) highlight the need to reduce the trend uncertainties in the lower stratosphere (LS). There is still a discrepancy 350 351 between modeled and observed ozone trends in the LS but large uncertainties make comparisons difficult. In this 352 study, we test additional predictors in the model to account for dynamical variability of ozone in the stratosphere, thus 353 improving the model performance and reducing the uncertainty of the trends. The argument for additional predictors 354 is that the LOTUS model was developed for the regression of zonally averaged ozone data, which reduces some 355 variability that might be impacting the ground-based records on regional bases. Impact of additional proxies in trend 356 analyses were reported in other publications (Weber et al, 2022a, Bernet, 2023 and references therein), and were mostly found to improve the statistical model fit at high latitudes where the impact of the descending branch of the 357 Brewer-Dobson circulation and Arctic/Antarctic oscillations has contributed to additional variability in stratospheric 358 359 ozone records.

360 In what we define as the 'extended' model, we add single additional predictors (one at a time) in the model as such:

 $\sum_{i=4}^{n} \beta_i(t,z) X_i(t,z) = \beta_4(t,z) QBO_A(t) + \beta_5(t,z) QBO_B(t) + \beta_6(t,z) ENSO(t) + \beta_7(t,z) Solar(t) + \beta_8(t,z) X_{predictor}(t,z) + \beta_8(t,z) X_{p$ 

The fitted predictors contain Fourier components, like in the reference model, to allow for seasonal variation.
 We test the following additional predictors as described below to assess the impact on trends and uncertainties:

- Quasi-Biennial Oscillation (QBO): Two notable disruptions to the otherwise relatively periodic QBO have
   occurred during the study period: 2015–2016 and in 2020 (Diallo, et. al 2022). Two additional leading modes
   of the calculated empirical orthogonal functions (EOF) are tested to improve the trend model fit during the
   anomalous QBO years.
- Arctic/Antarctic Oscillation (AO/AAO): the pattern of surface air pressure anomalies in the polar region and
   certain mid-latitude regions. The AO/AAO has strong correlations (Lawrence et al 2020) with stratospheric
   ozone through the strength of the polar vortex. The positive phase of the AO or AAO in the winter months
   is associated with low activity in the vertically propagating planetary Rossby waves, a strong polar vortex, a

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low vortex wavenumber, and low stratospheric temperatures. Thus, the positive (negative) phase of the 373 374 AO/AAO is correlated to low (high) ozone anomalies especially in the winter months (Lawrence et al, 2020). North Atlantic Oscillation (NAO): Similar to the Arctic Oscillation, this is a pattern of surface air pressure 375 ٠ 376 anomalies between certain regions in the high altitudes of the North Atlantic Ocean. This index is calculated 377 by the pressure difference between the Azores high and the subpolar low. 378 Eddy Heat Flux (EHF): The flux of heat through a zonal plane by transport due to the Brewer-Dobson • 379 circulation, here averaged from 45-75° N/S (use EHF S for Lauder only). This represents the planetary wave 380 activity that drives transport of ozone. 381 Tropopause Pressure (TP): Pressure level of boundary between the troposphere and the stratosphere. In this ٠ 382 study, we use the monthly mean pressure level of the tropopause from the NOAA National Centers for 383 Environmental Prediction (NCEP) reanalysis product. As the troposphere warms due to release of GHGs 384 and the stratosphere cools due to ODSs destroying stratospheric ozone, the tropopause is rising (Meng et al., 385 2021). Thompson et al, (2021) and Stauffer et al., (2023) found that the lower stratospheric ozone trends in 386 tropics become slightly positive when recomputed with respect to the tropopause height (which has its own 387 trend). This finding indicates that ozone depletion in the lower stratosphere (i.e. Ball et al., 2020) is driven

- by climate-change-related changes in transport and mixing in the lower stratosphere (i.e. Dan et al., 2020) is invenilist
  by climate-change-related changes in transport and mixing in the lower stratosphere. Therefore, we are
  testing the TP proxy in the model to account for non-chemical ozone losses in order to assess chemical
  attribution of ozone trends.
  Equivalent Latitude (EqLat): Geographical latitude of the isoline encircling the area of equal Potential
- Vorticity (PV) (Lary et al, 1995). The EqLat normalizes the range of PV values that change with season and interannually and makes it convenient for interpretation of ozone variability and trends (i.e. Wohltmann et al 2005). The dataset was generated from GMI CTM analyses (private communications with Susan Strahan, June 2021) for each ground-based station overpass criteria (latitude and longitude envelope, see above) and at several altitude levels coincident with Umkehr ozone profile layers. Appendix C discusses a COH dataset based on EqLat instead of geometric latitude. No advantage was found by using the EqLat coordinate system for the COH zonal dataset.
- 399 Source datasets for all predictors in the reference and extended models are shown in Table 3.

Predictor	Description	Source				
ENSO	El Nino/Southern Oscillation	Monthly https://psl.no	Mean aa.gov/enso/n	Multivariate nei.old/ <sup>1</sup>	ENSO	Index
Solar	Solar 10.7cm flux	https://spacev solaire/solarf	0	/forecast-prevision/s hp	olar-	

QBO	Quasi-Biennial Oscillation	Principal Component Analysis of the Monthly Mean Zonal Wind https://www.geo.fu-berlin.de/met/ag/strat/produkte/qbo/qbo.dat
AOD	AOD is included in t	he LOTUS model, but not used in this study
AO	Arctic Oscillation, Monthly Mean index	http://www.cpc.ncep.noaa.gov/products/precip/CWlink/daily_ao_index/m onthly.ao.index.b50.current.ascii
AAO	Antarctic Oscillation, Monthly Mean index	https://www.cpc.ncep.noaa.gov/products/precip/CWlink/daily_ao_index/a ao/aao.shtml
NAO	The North Atlantic Oscillation, monthly mean index	https://www.cpc.ncep.noaa.gov/products/precip/CWlink/pna/norm.nao.mo nthly.b5001.current.ascii.table
EHF	Eddy Heat Flux	Cumulative Mean (from September to April) of Heat Flux at 100 hPa from MERRA2 reanalysis averaged over 45–75° N (45–75° S for Lauder), deseasonalized. It is kept constant from April to Sep. https://acd- ext.gsfc.nasa.gov/Data_services/met/ann_data.html
TP	Tropopause Pressure	Monthly Mean NCEP-NCAR reanalysis (Kalnay et al., 1996); Tropopause pressure at the lat/lon of each station, deseasonalized. ftp://ftp.cdc.noaa.gov/Datasets/ncep.reanalysis.derived/tropopause/pres.tro pp.mon.mean.nc
EqLat	Equivalent Latitude	Monthly Mean equivalent latitude derived from MERRA2 -GMI CTM potential vorticity (PV) contours on 31 potential temperature surfaces

	[Susan Strahan, private communication, 8/24/2022]. The PV at each station	
	is determined by a 1/distance weighted average of the values in a $\pm$ 2 lat, $\pm$	
	2 lon grid, then converted to EqLat on the Umkehr layers.	

400 Table 3: List of predictors either previously used (bolded) in the LOTUS 0.8.0 (reference) model and additional predictors 401 evaluated in this study for a future use in the extended LOTUS trend regression model. Note, two components of the QBO 402 predictors were used in the reference model (i.e. Godin-Beekmann et al., 2022). We added two more components in the 403 extended model for tests described in this paper.

<sup>1</sup> Since the incorporation of the ENSO index into the LOTUS model, NOAA GSL has updated the index to v1.2.
 <u>https://psl.noaa.gov/enso/mei/</u>. However, for consistency with results from the Godin-Beekmann (2022) paper we use the old MEI index that is part of the LOTUS v 0.8.0 package.

407 All proxies are used as is. No de-trending (removal of the long-term trend in proxy) is applied to the proxies. Therefore,

408 we interpret any changes to the trends derived with additional proxies as approximations of trends driven by chemistry

409 and transport related to climate change. These are rough approximations as some feedbacks are known to impact

410 chemistry (e.g. changes in stratospheric temperature).

#### 411 **3.3 The Full Model - Combining Additional Predictors**

412 After we have determined the impact of the additional predictors singly, we discern which predictors should be 413 combined to constitute the 'Full Model'. Prior to selecting additional predictors for the 'Full Model', we perform

414 correlation tests to identify any cross correlations between predictors. We select predictors that are not highly

415 correlated (less than +/- 0.2) to ensure that all predictors are largely independent. We use the square of the Pearson

416 correlation coefficient  $R_1^2$  for each pair of the predictors to test our assumptions. We find that ENSO, Solar, QBO

417 (1,2,3,4), AAO, AO, EHF (N and S), and TP (at each station) have correlations less than +/- 0.2 (with the exception

418 of  $R_a^2 = 0.3$  for EHF (N) and AO). Therefore, any of these predictors can be combined in the 'Full Model'. We find

419 that NAO has a correlation of .38 with AO so we do not use these two predictors in the same model.

420 We also test the independence of EqLat proxies calculated at several geographic locations (defined by the latitude and

421 longitude of each Umkehr station) and by selecting a proxy at several altitude levels centered in the middle of Umkehr

- 422 layers 3–9. We find that the R<sup>2</sup> between the TP and EqLat in the lower stratosphere (Umkehr layer 3) can be large but
- 423 anticorrelated -0.7 (Boulder), moderate 0.4 (MLO and Lauder), while close to zero at Arosa/Davos and low at OHP
- 424 (-0.2). In the middle and upper stratosphere, the  $R_a^2$  varies from -0.5 to -0.4 (MLO), 0.2 to 0.3 (Arosa/Davos and OHP),
- 425 0.5 to 0.6 (Boulder), and 0.4 to 0.7 (Lauder). EqLat has mostly low correlations ( $< \pm 0.3$ ) with all other proxies except
- 426 for higher correlations with QBO B in layers 5 (-0.3) and 6 (-0.4), and QBO A in layer 7 (0.3) at MLO; and with AO
- 427 in layer 8 (0.3) at OHP and Arosa/Davos. Also, EqLat has no correlation with the TP proxy in layer 4 in Boulder, in

428 layer 9 at Lauder, and in layers 8 and 9 at OHP. Since there are occasional high correlations between EqLat and TP

429 proxies, we do not use them together in the 'Full trend Model'.

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#### 430 4 Results

#### 431 4.1 Reference Model Trend Results

432 First, we discuss the reference model trends derived from the COH overpass, Umkehr and ozonesonde records at 5 433 geographic locations. All datasets are deseasonalized with a climatology computed from a subset of data taken from 434 1998-2008 prior to the trend analysis. Trend results are presented in Fig. 3 and organized in 5 panels. Each panel shows trends at selected pressure/altitude levels detected from Umkehr (green), COH (orange) and ozonesonde (blue) 435 436 records at Arosa/Davos, OHP, Boulder, MLO/Hilo and Lauder ground-based stations. Ozonesonde data for the 437 Arosa/Davos panel are selected from Hohenpeißenberg, Germany station that is in close vicinity to Arosa/Davos 438 station. We show trends for layers where the measurement is of highest quality: Umkehr (layer 3 through 8), COH 439 (layers 5 through 9) and ozonesonde (layers 3 through 5) records.

440 The Umkehr data used in this analysis is the monthly mean of all available Umkehr data (one or two measurements

441 per day). The sonde and COH monthly means use only those profiles that have corresponding Umkehr measurements

442 on that date. We explore the impact of temporal sampling on trends in Appendix D. For COH with the Umkehr

443 matched data (see Figure A12), trends are slightly larger at OHP but well within the error bars. At all other stations

444 the COH trends are not impacted by sampling. At OHP the ozonesonde trends matched to Umkehr (see Figure A13)

are slightly larger at layer 4 only and well within the error bars; while at Lauder in layers 4 and 5 trends are smaller,but barely within the error bars.

Figure 3 shows that in the upper (above 10 hPa) stratosphere, Umkehr (black) and COH (orange) trends are positive and agree within the error bars (+/- 2 standard errors). The exception is found at 8–2 hPa pressure level over the Lauder station, where Umkehr trends are near zero and COH trends are  $\sim$  +3–4 %/decade. The error bars show +/- 2 standard errors, and the fact that they do not overlap suggests that the differences in trends are statistically significant. This could be related to the relatively large uncertainties in the instrumental corrections applied to homogenize the Umkehr record (Petropavlovskikh et al, 2022). Björklund (2023) discusses relative drifts in Umkehr, ozonesonde, FTIR and MW ozone records over Lauder. The authors are not able to identify instrumental artifacts that may have caused the

454 discrepancies in the co-located records, but point out that it is not related to the sampling biases.

455 In the middle stratosphere (60–10 hPa) agreement between Umkehr and COH is within uncertainty of the trend except

456 at Arosa/Davos where COH trends are statistically different from Umkehr trends at 16–8 hPa. COH trends at 32–16

457 hPa are mostly negative (-2-3 %/decade) with the exception of Lauder where trends are near zero and similar to

458 Umkehr trends. Umkehr trends between 32–16 hPa are close to zero. The ozonesonde trends (blue) agree with COH

459 (orange) and Umkehr (black) trends in layer 32-16 hPa at Arosa/Davos, Boulder and MLO. However, at OHP

460 (Lauder) the ozonesonde trends are found to be positive at  $+2\pm2,2$  %/decade (negative at  $-3\pm1.5$  %/decade) and 461 significantly different from the near-zero trends seen in the COH and Umkehr results.

462 In the lower stratosphere (125–63 hPa), Umkehr trends vary between small positive (+1–2 %/decade at Hilo and

463 Lauder) and negative (-2-3 %/decade at Arosa/Davos, OHP and Boulder); however, trend uncertainties are the largest

464 (2 standard errors are 2–3 %/decade, see Table 4 below) in comparison to the middle and upper stratospheric trends.

465 Ozonesonde trends at OHP station are positive (+2, %/decade), and negative over Lauder (-2 %/decade). They also

466 feature large uncertainties (±4.2,%/decade at OHP) that are larger than the uncertainties found in Umkehr trends which

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could be caused by the limited sampling (see Appendix D<u>, Figure A11</u>). Sonde trends at Hilo show negative trend
values with large uncertainties. But the data in this study at Hilo is not corrected for the ozonesonde drop off after
2014 known to occur at this station (Stauffer, 2022), so the deviation from the Umkehr results at these levels may be
misleading.

480 Figure 3 also shows trends derived from the zonal-mean COH data associated with each station (orange dashed line).

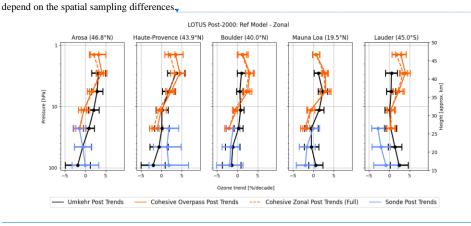
481 These are shown for comparison with the overpass COH data (solid line) to study the impact of the spatial sampling

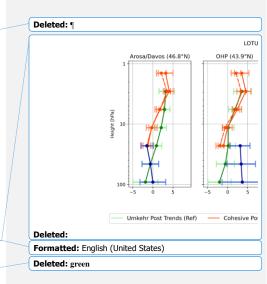
482 biases on the trends. Though Figs. 1 and 2 show clear interannual differences between the records from the individual

483 stations, and the associated zonal average, we find very small differences in trends (mostly in the upper stratosphere

484 at middle latitude stations). Therefore, the station overpass sampling provides trends that are representative of the

zonal averaged trends (Zerefos, 2018) and the discrepancies in trends between GB and satellite records do not strongly
 depend on the spatial sampling differences.





487

488Figure 3: The 2000–2020 ozone trends are shown at 7 altitude/pressure levels. The LOTUS model v 0.8.0 is used for trend489analyses. Umkehr trends (black), COH (orange) and ozonesonde (blue) are shown for 5 ground-based stations:490Arosa/Davos, OHP, Boulder, MLO and Lauder (panels left to right). Ozonesonde data for the Arosa/Davos panel are491selected from Hohenpeißenberg, Germany that is in close vicinity to Arosa/Davos. Trends from the zonal-mean COH data492(orange dashed line) are shown for comparison with the overpass COH data (solid line). The error bars indicate ± 2 standard493errors.

#### 494 4.2 Standard Error of Reference Model

	LOTUS Model Proxy Tests: Standard Error for Reference Model															
Height	Height Umkehr Arosa/Davos		OHP			Boulder				MLO		Lauder				
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		0.92			0.91			0.62			0.43			0.63	
2-4	8	0.85	0.59		1.06	0.68		0.51	0.52		0.52	0.37		0.72	0.57	
4-8	7	0.69	0.59		0.77	0.54		0.41	0.52		0.58	0.62		0.57	0.66	
8-16	6	0.66	0.68		0.75	0.59		0.42	0.43		0.55	0.49		0.61	0.56	
16-32	5	0.66	0.75	0.76	0.89	0.68	<u>1.10</u>	0.54	0.51	0.77	0.82	0.55	0.75	0.73	0.54	0.73

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32-6	3	4	1.05	1.04	1.13	1.55	0.90	 1.04	0.90	 0.94	0.83	1.16
63-12	7	3	1.55	1.60	0.15	2,10	1.15	1.63	0.87	1.07	1.11	1.50

504

505

500	Table 4: Standard Error (SE) for the Reference model 2000-2020 trend for five ground-based station locations
501	(Arosa/Davos, OHP, Boulder, MLO and Lauder). Results are provided for trend analyses of the Cohesive satellite (COH),
502	Dobson Umkehr (UMK) and ozonesonde (SND) records and for Umkehr. The layers are selected to represent the best
503	quality of data. Values of SE shown are the actual errors in DU/decade.

rstanding of trand values. The standard arr

We will use the standard error of the linear (trend) term in Equation 1, to evaluate the success of the additional proxies

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505	to improve understanding of rend values. The standard error is an output of the regression code, and indicates the
506	uncertainty in the trend value. Smaller Standard Errors indicate increased confidence in the trend result. We use the
507	standard error as a metric instead of standard deviation to reduce dependence on the number of points in the trends
508	model. The Table 4 provides the Standard Errors for the Reference Model fit and represents uncertainty of the trend
509	in DU of the mean ozone in each layer at the station. The standard errors of the trend detected in three co-located
510	ozone records at each station (or in the nearby location as in case of Arosa/Davos or MLO comparisons) do not
511	significantly differ, although in general ozonesonde errors are slightly larger than Umkehr errors most likely due to
512	the larger sampling errors in ozonesonde monthly mean record. Also, the errors in trends detected in COH layers 5-8
513	are on average smaller than for Umkehr trends (with the exception of layer 7 at Boulder, MLO and Lauder) which
514	could be explained by an overpass method that averages several satellite profiles from adjacent orbits and therefore
515	reduces meteorological scale variability in averaged ozone data.
<b>C1</b> C	

#### 516 4.3 Adjusted R<sup>2</sup>

517 The adjusted R<sub>4</sub><sup>2</sup> values of the 2000–2020 trends are shown in Fig. 4 and Table 5 for the data fit using the Reference 518 model. The adjusted  $\frac{R^2}{r}$  is a modified version of  $\frac{R^2}{r}$  that adjusts for the number of predictors in a regression model and 519 represents the 'goodness' of the model fit to the data. For COH adjusted R<sup>2</sup><sub>2</sub> is shown for both the overpass and the 520 zonal datasets.

521 Though values are significantly less than the high values usually seen when comparing data that includes the prevalent

seasonal variation, the adjusted  $\frac{R^2}{r}$  values for the COH zonal mean record are similar in magnitude and vertical shape

- 523 to the results of the (60°S–60°N) broadband trend analyses published in Godin-Beekmann (2022), Fig. 7 varying
- between 0.1 and 0.5. We designate the average values (0.3) as a threshold for satisfactory fit indicating conformance
- 525 with prior LOTUS results. We indicate in bold in Table 5 adjusted  $\frac{R^2}{4}$  values of 0.3 or greater to note achievement of
- 526 that threshold and include vertical dashed line in Fig. 4 for reference.

527 The adjusted  $\frac{R^2}{r}$  for the Reference model fit is slightly better for the zonal mean COH data than for the COH overpass

528 over the Northern middle latitude stations. This is expected as much of the variability of the time series is reduced in

529 the zonal average as compared to the station overpass data as shown in Fig. 2, and more easily explained by the

530 typically used predictors. Indeed, the goal of this study is to determine if the additional predictors help to explain the

531 additional variation as measured at point locations.

532 The model fit to the GB data is similar to the COH overpass results in the middle stratosphere (layers 5 and 6), but the

533 model explains less ozone variability in the Umkehr records in the upper stratosphere (layers 7 and 8). In the lower

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546 stratosphere (layers 3, 4 and 5), the model fit to the ozonesonde and Umkehr records is similar with the exception of

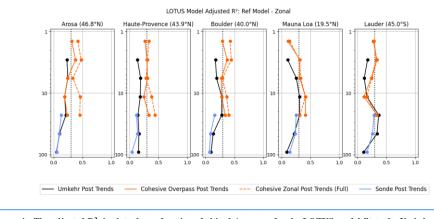
547 Lauder (Umkehr has larger adjusted  $\mathbb{R}^2_{\psi}$  in layers 4 and 5). The adjusted  $\mathbb{R}^2_{\psi}$  for COH overpass in layer 5 is similar to

548 Umkehr and sonde with a larger difference at OHP. The adjusted R2, in the lower stratosphere is less than in the middle

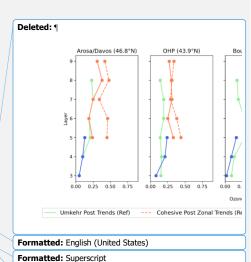
549 stratosphere, which points to other processes (e.g., transport) that drive ozone variability. In this paper we investigate

550 improvement to the trend model fit by introducing additional proxies that can improve representation of the

551 dynamically-driven ozone variability in the stratosphere.







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Figure 4: The adjusted R<sup>2</sup><sub>a</sub> is plotted as a function of altitude/pressure for the LOTUS model fit to the Umkehr (<u>plack</u>),
 ozonesonde (blue), COH overpass (orange, solid), and COH zonal-mean (orange, dashed). Results are shown in 5 panels
 that represent trend analyses of ozone records over Arosa/Davos (Hohenpeißenberg for sondes), OHP, Boulder, MLO (Hilo
 for sondes) and Lauder ground-based stations.

				LO	TUS Mo	del Prox	y Tests:	Adjusted	R2 for	Referenc	e Model					
Height	Umkehr	A	rosa/Dav	'05		OHP			Boulder			MLO			Lauder	
(hPa)	Layer	UMK	СОН	SND	UMK	сон	SND	UMK	сон	SND	UMK	СОН	SND	UMK	сон	SNE
1-2	9		0.31			0.27			0.29			0.11			0.29	
2-4	8	0.23	0.38		0.14	0.30		0.17	0.37		0.11	0.32		0.17	0.32	
4-8	7	0.25	0.25		0.19	0.31		0.19	0.27		0.26	0.32		0.12	0.24	
8-16	6	0.19	0.19		0.19	0.25		0.28	0.28		0.31	0.41		0.16	0.11	
16-32	5	0.21	0.24	0.13	0.14	0.33	0. <u>14</u>	0.28	0.34	0.16	0.31	0.41	0.25	0.37	0.34	0.26
32-63	4	0.10		0.10	0.16		0.24	0.13		0.09	0.22		0.24	0.34		0.20
63-127	3	0.05		0.05	0.15		0.25	0.09		0.02	0.09		0.14	0.11		0.10

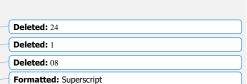
557 Table 5: Similar to Table 4, but for the adjusted R<sup>2</sup>. Values of 0.30 and above are indicated in Bold as a threshold to 558 indicate a satisfactory fit.

# 559 4.4 Reference Model P-Values:

560 The p-values are often used to evaluate statistical significance of predicted results and results labelled "significant" if

they remain below a threshold of 0.05. However, Chang et al. (2021) argued as Wasserstein et al. (2019) does that all

trends should be reported with their associated p-values and a thorough discussion of the certainty of trend detection



## 572 as described by the p-values. Therefore, the p-values can be used for understanding the certainty of the trend. Under

573 the IGAC TOAR activity, p-values are scored to define a consistent scale for comparison of the trends between

574 different analyses (see Table 3, Chang et al., 2023).

575

							el Refer									
Pressure			osa/Dav			e-Prov			Boulder			auna Lo			Lauder	
(hPa) 1-2	Layer 9	UMK	COH 0.00	SND	UMK	COH 0.00	SND	UMK	COH 0.03	SND	UMK	COH 0.10	SND	UMK	COH 0.00	SNE
2-4	8	0.00	0.00		0.00	0.00		0.05	0.03		0.02	0.10		0.47	0.00	
4-8	7	0.00	0.01		0.02	0.00		0.12	0.00		0.00	0.00		0.43	0.00	
8-16	6	0.00	0.62		0.84	0.98		0.05	0.66		0.01	0.17		0.85	0.50	
16-32 32-63	5 4	0.17	0.08	<b>0.05</b>	0.87	0.15	0.12	0.58	0.00	0.21	0.56	0.00	0.93	0.61	0.62	<b>0.0</b>
63-127	3	0.55		1.00	0.62		0.41	0.21		0.12	0.61		0.01	0.10		0.0
	Similar to bolded nu															
	low certa															
Fable 6 n	rovides p	-values	s for th	e Refe	rence l	Model	These	e are fi	urther u	used as	a hase	line fo	r comp	arison	to mor	lel fi
1	itional pre												1			
									00	U		• · ·	ŕ			
	data in la								•		U	•	,			
ound at 1	MLO in 1	ayer 9.	Also,	high c	ertaint	y in de	rived t	rends i	s reach	ned for	COH	record	s in lay	er 5 at	Bould	er ar
MLO.																
Jmkehr (	trend anal	lyses al	lso sho	w high	confi	dence i	in trend	l detec	tion at	Arosa	/Davos	and N	ILO st	ations	in laye	rs 6,
und 8, at	OHP in la	ayers 7	6 and	8, and	in Bou	lder in	layers	6 and	8. For	the oz	onesoi	nde da	ta the h	igh co	nfiden	ce (i.
ow unce	rtainty) is	s found	for He	ohenpe	ißenbe	rg, and	d Boul	der trei	nds det	acted :	in lave	r 5, an	d at La	uder ir		
5.										ecteu i	in naje.			ader m	layers	4 ai
										ecteur	in huj e.				layers	s 4 ai
The medi	ium level	of the	certain	nty (0.0	)5 <p≤(< td=""><td>0.10) i</td><td>s found</td><td>l in tre</td><td>nds de</td><td></td><td>2</td><td>er 5 of</td><td>СОН о</td><td></td><td></td><td></td></p≤(<>	0.10) i	s found	l in tre	nds de		2	er 5 of	СОН о			
	ium level				1 -					tected	in laye			ozone 1	time se	
Arosa/Da		r 3 of c	ozones	onde a	MLO	, and i	n layer	4 of <u>o</u>	zonesc	tected	in laye	<u>r 3 of </u>	Umkeh	ozone 1 r at La	time se uder.	ries
Arosa/Da Low cert	ivos, laye	er 3 of c detecte	ozones d tren	onde at ds at p	MLO -value	, and i of 0.	n layer 10 (no	4 of <u>o</u> t inclu	<u>zonesc</u> sive) t	tected onde an o 0.33	in laye ad laye is fou	r <u>3 of 1</u> nd in	Umkeh Umkel	ozone 1 r at La nr laye	time se uder. er 3 an	ries d 5
Arosa/Da Low cert Arosa/Da	ivos, laye ainty in	er 3 of c detecte COH lay	ozones d trend yer 5 <u>a</u>	onde at ds at p nd,Um	MLO -value	, and i of 0. of 3	n layer 10 (no at OHF	4 of <u>o</u> t inclu	<u>zonesc</u> sive) t	tected onde an o 0.33	in laye ad laye is fou	r <u>3 of 1</u> nd in	Umkeh Umkel	ozone 1 r at La nr laye	time se uder. er 3 an	ries d 5
Arosa/Da Low cert Arosa/Da n ozones	avos, laye ainty in avos; in C sonde laye	er 3 of c detecte COH lay er 4 and	ozones d trend yer 5 <u>a</u> d COH	onde at ds at p nd,Um [ layer	MLO -value kehr, la 6 recoi	, and in of 0. ayer 3 and at M	n layer 10 (no at OHF ILO.	4 of <u>o</u> t inclu P; in Ui	zonesc sive) t mkehr	tected onde an o 0.33 layers	in laye ad laye is fou 3,4, an	r <u>3 of</u> and in ad ozor	Umkeh Umkel nesonde	ozone f r at La nr laye e layer	time se uder. er 3 an s <sub>v</sub> at Bo	eries d 5 oulde
Arosa/Da Low cert Arosa/Da n ozones Highest (	ainty in ainty in avos; in C sonde laye lowest ce	er 3 of c detecte COH lay er 4 and ertainty	ozones d tren yer 5 <u>a</u> d COH	onde at ds at p nd,Um [ layer lues (>	MLO -value kehr la 6 recor 0.33) v	, and in of 0. ayer 3 and rd at M were fo	n layer 10 (no at OHF ILO. pund ir	4 of <u>o</u> t inclu P; in U n layer	zonesc sive) t mkehr 6 of C	tected onde an o 0.33 layers OH ov	in laye id laye is fou 3,4, an verpass	r <u>3 of</u> and in ad ozor record	Umkeh Umkel nesonde ls at m	ozone f r at La nr laye e layer ost sta	time se uder. er 3 an s <sub>v</sub> at Bo tions (	ries d 5 oulde exce
Arosa/Da Low cert Arosa/Da n ozones Highest ( For MLO	avos, laye ainty in avos; in C sonde laye lowest ce where p-	er 3 of c detecte COH lay er 4 and ertainty values a	ozonesi od trend yer 5 <u>a</u> d COH ) p-val are me	onde a ds at p nd Um [ layer lues (> dium h	t MLO -value kehr la 6 recor 0.33) v igh). V	, and in of 0. ayer 3 a rd at M were for We note	n layer 10 (no at OHF ILO. pund ir e that th	t inclu r inclu r; in Un n layer ne COF	zonesc sive) t mkehr 6 of C I trend	tected onde an o 0.33 layers OH ov s are cl	in laye is fou 3,4, an verpass ose to 2	r 3 of ind in d ozor record zero ar	Umkeh Umkel nesonde ds at m nd the u	ozone f r at La nr laye e layer ost sta ncerta	time se uder. er 3 an s <sub>v</sub> at Bo tions ( inty en	eries d 5 oulde exce veloj
Arosa/Da Low cert Arosa/Da n ozones Highest ( for MLO crosses th	ainty in ainty in wos; in C sonde laye lowest ce where p- ne zero lin	er 3 of c detecte COH lay er 4 and ertainty values a ne. <u>The</u>	ozoneso od treno yer 5 <u>a</u> d COH ) p-val are me <u>erefore</u>	onde at ds at p nd, Um [ layer lues (> dium h e, the st	t MLO -value kehr, la 6 recor 0.33) v igh). V atistica	, and in of 0. a of 0. a of 0. a of 0. a of 0. a of 0. b of 0.	n layer 10 (no at OHF ILO. pund ir e that th d mode	t inclu r inclu r; in U r layer ne COF	zonesc sive) t mkehr 6 of C I trend ot sepa	tected onde an o 0.33 layers OH ov s are cl arate tre	in laye is fou 3,4, an verpass ose to a ends fr	r 3 of 1 nd in d ozor record zero ar om zer	Umkeh Umkel hesonde ds at m hd the u	ozone f r at La nr layer e layer ost sta ncerta	time se uder. er 3 an s <sub>v</sub> at Bo tions ( inty en cplaine	d 5 ouldd exce velo d hig
Arosa/Da Low cert Arosa/Da n ozones Highest ( For MLO crosses th pzone va	avos Jaye ainty in avos; in C sonde laye lowest ce where p- ne zero lin riability i	er 3 of c detecte COH lay er 4 and ertainty values a ne. The n this 1	ozoneso d trend yer 5 <u>a</u> d COH ) p-val are me erefore layer.	onde av ds at p nd Um layer lues (> dium h e, the st , Simil	t MLO -value kehr, la 6 recor 0.33) v igh). V atistica arly, n	, and in of 0. ayer 3 a ad at M were for We note al trend ear-zer	n layer 10 (no at OHF ILO. ound ir e that th d mode ro Uml	4 of o t inclu ; in U h layer he COF	zonesc sive) t mkehr 6 of C I trend ot sepa ends w	tected onde an o 0.33 layers OH ov s are cl arate tra-	in laye is fou 3,4, an verpass ose to 2 ends fr atively	r 3 of 1 nd in d ozor record zero ar om zer large	Umkeh Umkel nesonde ds at m nd the u SE in 1	ozone f r at La nr laye e layer ost sta ncerta to une layer 6	time se uder. er 3 an s, at Bo tions ( inty en cplaine	ries d 5 ouldd exce velo d hig IP ar
Arosa/Da Low cert Arosa/Da n ozones Highest ( For MLO crosses th pzone va	ainty in ainty in wos; in C sonde laye lowest ce where p- ne zero lin	er 3 of c detecte COH lay er 4 and ertainty values a ne. The n this 1	ozoneso d trend yer 5 <u>a</u> d COH ) p-val are me erefore layer.	onde av ds at p nd Um layer lues (> dium h e, the st , Simil	t MLO -value kehr, la 6 recor 0.33) v igh). V atistica arly, n	, and in of 0. ayer 3 a ad at M were for We note al trend ear-zer	n layer 10 (no at OHF ILO. ound ir e that th d mode ro Uml	4 of o t inclu ; in U h layer he COF	zonesc sive) t mkehr 6 of C I trend ot sepa ends w	tected onde an o 0.33 layers OH ov s are cl arate tra-	in laye is fou 3,4, an verpass ose to 2 ends fr atively	r 3 of 1 nd in d ozor record zero ar om zer large	Umkeh Umkel nesonde ds at m nd the u SE in 1	ozone f r at La nr laye e layer ost sta ncerta to une layer 6	time se uder. er 3 an s, at Bo tions ( inty en cplaine	ries d 5 ouldd exce velo d hig IP ar
Arosa/Da Low cert Arosa/Da n ozones Highest ( For MLO crosses th ozone va Lauder, 1	avos Jaye ainty in avos; in C sonde laye lowest ce where p- ne zero lin riability i	contraction of the second seco	d trend d trend yer 5 <u>a</u> d COH ) p-val are me <u>erefore</u> layer.	onde at ds at p nd Um [ layer lues (> dium h c, the st c, the st rosa/D	t MLO -value kehr, la 6 recor 0.33) v igh). V atistica arly, n	, and in of 0. ayer 3 a rd at M were for Ve note al trend ear-zen station	n layer 10 (no at OHF ILO. ound ir e that th d mode ro Uml us, and	4 of o t inclu r; in U h layer he COH t cann kehr tr in laye	zonesc sive) t mkehr 6 of C I trend ot sepa ends w er <u>s</u> 3 a	tected onde an o 0.33 layers OH ov s are cl arate tra- rith rela- nd 4 a	in laye is fou 3,4, an verpass ose to a ends fr atively t MLC	r 3 of 1 ind in d ozor record zero ar om zer large ) show	Umkeh Umkel aesonde ds at m ad the u so due t SE in the sa	ozone f r at La nr laye e layer ost sta ncerta to unes layer 6 me lev	time se uder. er 3 an s_at Bo tions ( inty en cplaine 5 at OF rel of h	d 5 oulde exce veloj d hig IP ar

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	herefore, these results point to the trend model's tetect non-zero trends and account for all ozone this layer

617 It is also important to note that the reference trend model fit to ozone in Umkehr layers 7 and 8 at Lauder has high p-

618 values, which is related to the near-zero trends that shows large disagreement with COH trend. This difference could

619 be caused by remaining instrumental step changes that were not fully removed during the record homogenization

620 (Petropavlovskikh et al., 2022).

621

622 While near-zero trends and high p-values are found in the fit of the Hilo ozonesonde record in layer 5, the p-values in

623 layer\_4 show only medium p-values for near zero trends. It is possible that infrequent launches of ozonesonde

624 observations at Hilo could create the temporal sampling bias and appear noisy. The ozonesonde record at

625 Hohenpeißenberg has sufficiently frequent sampling (3 times per week) for successful trend analyses (Chang et al.,

2020; Chang, 2023 preprint), but the p-values remain high in layers 3–4. The p-values for Umkehr fit at Arosa/Davos

are in the medium to high range for layers 3, 4, 5, but somewhat smaller which could be due to non-zero trends in

628 layers 3 and 5. The p-value difference could be also related to the different location of the ozonesonde (HOH) and

629 Umkehr (Arosa/Davos) observations, thus the records could contain different atmospheric variability that might

630 impact the model fit.

631 We will discuss changes to the p-values in the next section after we add more proxies to the trend model in an attempt

632 to improve confidence in trend detection.

#### 633 5 Trends with the Extended Model - testing the addition of single predictors

634 The LOTUS styled Reference Model is developed and optimized for zonal average datasets. Modeling and trend 635 analysis for GB and satellite overpass data may improve by the addition of other proxies not used in the reference 636 model to improve capturing processes that impact ozone changes over limited geographical regions. The Extended 637 Model tests the addition of single predictors to see if fit statistics can be improved for GB and overpass datasets. We 638 judge success of the Extended Model by examination of the reduction in the Standard Error of the trend term, and by evaluation of the impact on the adjusted  $R_{k}^{2}$  of the model fit. Table 7 displays the change in the Standard Error of the 639 post 2000 trend for each proxy tested determined as SEref - SEext as a percent of SEref. As such positive values 640 641 correspond to the desired reduction of SE, and are highlighted in the table in <u>blue</u>. Low impact changes in the SE are 642 highlighted in white, and increases in SE (negative values) are highlighted in red. It may seem unusual for the addition 643 of proxies to increase the SE (negative values in the table) which indicates less confidence in the fit. But these SE are 644 the uncertainty in the trend term, not in the overall model fit. The new proxies considered each have a possible trend 645 and associated error budget for that trend. Whether the additional proxy increases trend uncertainty can depend on how well the trend of the new proxy can be characterized. The adjusted  $R_{k}^{2}$  is a better indicator of the overall model 646 647 improvement. Table 8 displays the adjusted  $R_{k}^{2}$  for the Extended Model for each proxy tested. Values of 0.30 and 648 above are indicated in bold as a threshold to indicate a satisfactory fit.

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		a) LO	rus Mo	del Test	: Differ	ence [%	6] in Sta	ndard E	Frror: Tr	opopau	se Pres	ssure vs	s Refer	ence Mo	odel	
Pressure				Haute-Provance			Boulder			Mauna Loa				Lauder		
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	COH	SND	UMK	СОН	SND	UMK	COH	SND
1-2	9		0.3			0.1			0.5			1.4			3.0	
2-4	8	-0.7	-0.5		-0.1	-0.4		-0.2	0.4		-0.6	-0.3		1.3	2.6	
4-8	7	-0.3	0.0		0.3	1.3		0.3	-0.2		2.6	0.3		3.7	1.4	
8-16	6	-1.1	-0.7		0.0	0.3		0.7	-0.2		0.6	0.8		3.1	5.4	
16-32	5	-0.2	2.1	-0.9	1.1	5.3	2.4	-0.4	0.6	0.6	4.5	9.3	2.7	0.0	0.7	2.4
32-63	4	6.6		6.1	5.9		9.9	3.4		7.5	7.0		6.1	8.0		9.4
63-127	3	12.8		10.2	12.8		10.7	6.8		6.0	5.8		4.6	9.8		7.9

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<u>b)</u>																
	L	OTUS M	odel Te	st: Diff	erence	[%] in S	tandaro	Error:	Equival	ent Lati	tude vs	Refere	nce Mo	del		
Pressure	Umkehr	Are	Arosa/Davos			te-Prova	ance	e Boulder			М	auna Lo	ba	Lauder		
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		8.4			2.9			1.9			-7.2			2.9	
2-4	8	-0.5	0.7		0.1	1.2		-0.4	1.5		-3.5	-5.4		1.0	3.1	
4-8	7	3.8	3.2		2.1	0.6		5.4	4.1		-2.6	-3.9		0.5	1.2	
8-16	6	6.1	8.3		2.5	10.9		2.4	7.8		5.3	7.8		3.4	7.7	
16-32	5	7.9	10.6	5.9	1.9	13.4	8.7	-1.9		1.4	0.3	0.7	0.7	0.8	3.9	-1.1
32-63	4	-1.4		-1.8	3.2		0.6	-0.2		-0.5	0.3		1.0	-0.2		-0.6
63-127	3	1.3		2.0	-1.4		-3.3	-0.8		-0.4	9.6		2.3	1.4		0.6

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<u>c)</u>

	c) LOTUS Model Test: Difference [%] in Standard Error: QBO C/D vs Reference Model															
Pressure	Umkehr	Are	Arosa/Davos			Haute-Provance			Boulder			auna Lo	ba	Lauder		
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		-1.6			-0.2			0.5			-0.5			-3.3	
2-4	8	-0.8	3.1		-0.3	9.1		2.9	4.6		-3.5	-0.3		-1.7	-0.4	
4-8	7	-0.1	1.5		0.3	3.3		-2.7	-1.2		-6.1	-4.2		1.8	1.4	
8-16	6	0.5	-1.3		1.1	-0.3		-2.4	0.7		-0.4	0.8		-2.5	-2.9	
16-32	5	-0.8	1.3	0.0	-0.9	3.0	2.8	0.6	0.6	0.1	7.1	8.4	10.1	-3.1	-0.6	1.1
32-63	4	0.9		0.2	2.0		-0.9	2.7		-1.8	2.9		6.5	-1.6		-1.9
63-127	3	-0.3		-0.8	5.7		-0.2	0.3		-4.2	0.4		3.0	-2.8		-3.2

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<u>d)</u>

	d) LOTUS Model Test: Difference [%] in Standard Error: AO/AAO vs Reference Model															
Pressure	Umkehr	Arc	Arosa/Davos			Haute-Provance			Boulder			auna Lo	ba			
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		1.2			-1.6			0.3			-1.9			-0.5	
2-4	8	-0.8	0.0		-3.8	-1.2		-0.8	-0.4		-2.1	-2.4		0.8	-1.9	
4-8	7	-0.7	1.7		-4.2	-2.6		3.2	4.7		1.2	-3.4		1.2	-1.2	
8-16	6	-0.2	-0.6		-2.4	-3.9		1.2	0.5		1.6	-1.6		-0.3	2.5	
16-32	5	-1.2	-0.4	-1.2	0.5	-2.1	-2.4	0.2	-0.6	-2.1	3.9	1.8	0.7	-1.6	-1.5	1.4
32-63	4	5.8		7.8	0.4		4.6	-1.2		-1.7	7.6		1.7	-0.7		2.5
63-127	3	13.1		12.9	5.5		6.8	-1.4		-3.2	4.4		1.3	-1.1		2.4

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<u>e)</u>

		e) LO	rus Mo	del Test	: Differ	ence [%	6] in Sta	ndard I	rror: N	AO vs R	eferen	ce Mode	el			
Pressure	Umkehr	Are	Arosa/Davos			Haute-Provance			Boulder			auna Lo	ba	Lauder		
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		0.5			-2.5			-0.2			-3.7			-1.3	
2-4	8	-0.2	0.0		-3.1	-1.9		-0.4	-1.9		-1.7	-3.8		-2.0	-1.8	
4-8	7	-0.6	0.7		-2.0	-2.0		0.0	3.9		2.6	-1.1		-2.8	-2.4	
8-16	6	0.2	-0.9		-1.7	-3.4		-2.2	-2.8		2.4	-0.4		-2.0	-0.5	
16-32	5	-0.5	-1.2	-1.1	0.7	-2.2	-4.0	-0.4	-1.4	-1.2	-1.4	-0.7	-3.0	-2.5	-4.3	-1.5
32-63	4	2.6		3.1	-0.6		-0.9	-0.2		0.4	1.5		-0.8	-2.6		-4.8
63-127	3	10.6		6.7	2.7		1.0	0.4		-2.7	1.7		-0.5	-2.3		-4.9

670 671

f)

		LOTUS	Model	Test: D	ifferend	ce [%] ii	n Stand	ard Erro	or: Eddy	Heat Fl	ux vs R	eferenc	e Mode	el		
Pressure	Umkehr	Are	osa/Dav	/0S	Haute-Provance			Boulder			М	auna Lo	ba	Lauder		
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	COH	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		5.0			4.5			4.4			-3.2			0.2	
2-4	8	-1.4	4.6		2.6	6.0		3.1	8.8		-1.6	-3.3		0.7	1.9	
4-8	7	-2.7	-3.4		-0.4	-3.9		-3.0	-2.3		5.0	-4.4		1.8	4.5	
8-16	6	-3.1	-3.2		-2.5	-4.8		-2.4	-3.5		-1.1	0.4		-0.2	1.1	
16-32	5	-3.4	-2.8	-3.2	-2.2	-3.7	-2.5	-2.6	-2.4	-2.5	9.3	-0.4	4.3	0.7	2.4	0.7
32-63	4	-1.9		-1.6	-2.0		-1.8	-2.7		-3.5	8.8		3.1	1.9		1.6
63-127	3	1.5		1.4	-0.9		-1.6	-2.5		-3.8	0.9		0.9	2.1		1.1

672 675

676Table 7: Change in Standard Error (SE) of the post-2000 trend estimate, in percent of SE of Reference Model for adding677single predictors. Panel a: Tropopause Pressure; b: Equivalent Latitude; c: QBO terms C and D; d: AO/AAO; e: NAO; f:678Eddy Heat Flux. Cells with reduced (increased) SE have blue (red) background, while cells with low impact changes (<0.5</td>679%) have no colours.

#### 685 5.1 Tropopause pressure (TP)

Adding the TP proxy to the standard LOTUS model produces the most consistent results between different techniques 686 687 (COH, Umkehr and ozonesonde) and also have similar magnitude of standard error changes among different latitudes (i.e. Arosa/Davos, OHP, Boulder, MLO, Lauder). The most significant impact in improving the SE is found in the 688 689 lower stratosphere (layers 3, 4) and in the middle stratosphere (layer 5) at the MLO tropical station. The impact of the TP proxy on the COH trend uncertainty in the model stratosphere (layer 5) is somewhat larger, likely due to the 690 691 satellite AK extending into the lower stratosphere. Similarly, larger reduction of the standard error in the Umkehr 692 trends in the lowermost stratosphere (layer 3) in comparison to the AK-smoothed ozonesonde record could be due to 693 sampling biases in the ozonesonde record. Adding the TP proxy to the Reference Model improves the adjusted  $R_{A}^{2}$  in 694 layers 3-5, whereas the SE improvements are also consistent across geo-locations and measurement techniques. The 695 TP proxy only explains ozone variability near the tropopause because changes in both parameters are linked to the 696 same dynamical processes (i.e. irreversible mixing). In the middle and upper stratosphere ozone variability is not 697 linked to the processes that change TP, thus using this proxy add error to the model fit. Several improvements resulted in adjusted R<sup>2</sup> to exceed the 0.3 threshold (Umkehr at OHP in layer 3, sonde and Umkehr at Lauder and MLO in layer 698 699 4) and in many cases the adjusted  $R_{a}^{2}$  increased by more than 0.02.

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LOTUS Model Proxy Tests: Adding NAO (% difference in Std. Error of Model)

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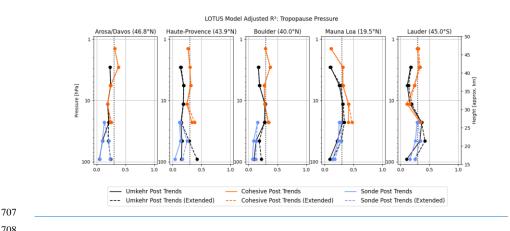
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709 Figure X a) Similar to Figure 4, but adjusted R<sup>2</sup><sub>4</sub> results are shown for both Reference model (solid line) and the 710 Extended model (Dashed line, Full) for COH overpass (orange), Umkehr (black) and ozonesonde (blue) trend. 711 Extended model includes additional TP proxy.

#### 712 5.2 Equivalent Latitude (EqLat)

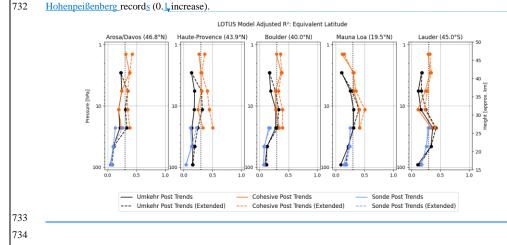
713 In the mid-latitudes, the addition of EqLat as a predictor shows consistent results across measurement techniques and 714 stations with few exceptions. The reduction in the SE of the model fit is evident in the COH data in the upper stratosphere (above 4 hPa or ~ 40 km), but is less pronounced in Umkehr profiles. The impact on MLO SE of the trend 715 716 fit in the upper stratosphere is negative (in both COH and Umkehr records) which can be explained by the fact that 717 the EqLat is much closer to geometric latitude near the equator than at the middle/high latitudes and therefore its use 718 as a proxy would not provide any additional information in interpretation of the tropical upper stratospheric ozone 719 variability. It could also suggest that the addition of EqLat will overfit the record. 720 The ozone record trend fits in the middle stratosphere (32-4 hPa or 25-40 km) benefit from adding the EqLat proxy

721 at most locations. Improvement in the SE of the trends in the lower stratosphere (127-63 hPa or ~15-20 km) is 722 minimal, limited to some locations and instrumental records (Arosa/Davos Umkehr and HOH sonde, MLO Umkehr 723 and sonde, and Lauder Umkehr and ozonesonde), which could be related to the location of subtropical jet that 724 modulates mixing of tropical and subtropical (and occasionally polar) air masses and influences the strat/trop

- 725 exchange. Unexpectedly, the addition of the EqLat proxy to the MLR statistical model for trend detection in Boulder 726 Umkehr and ozonesonde lower stratospheric ozone records increases the uncertainties of the fit, while the influence 727 of subtropical jet on Boulder lower stratosphere is well known (Manney et al, 2018). Perhaps, the data analyses also
- 728 need to consider the tropopause variability.
- 729 In terms of the impact on the adjusted  $R_{a}^{2}$ , the EqLat proxy significantly improves model fit for multiple instruments, 730 mostly in layers 5–7, and in COH fit in layer 9. The adjusted  $R_a^2$  improvements also often exceeded 0.3 threshold. No
- 731 significant improvement is found in the ozonesonde model fit in layer 5 with the exception of the OHP and

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## 735 Figure X b) the same as a), but Extended model includes Equivalent Latitude proxy.

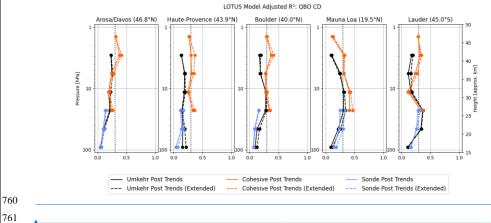
## 736 5.3 Extra QBO terms C and D

737 QBO is an important driver of ozone variability at tropical stations. Based on the results of adding 2 extra terms of the 738 QBO to the standard model, the recommendation could be to exercise this option only for the tropical station trends. 739 At the Northern middle latitudes (i.e in Arosa/Davos, OHP and, to a lesser degree, in Boulder) an improvement to the 740 trend SE uncertainties in layer 8 is noted. There seems to be a similar pattern for the upper stratosphere in trends derived with Heat Flux. Tweedy et al. (2017) show that the first two EOFs of the QBO did not describe the anomalous 741 742 QBO behavior, while Anstey et al. (2021) show that the addition of two more EOFs of the QBO could capture the 743 effect of the disruptions on the zonal winds. Therefore, including additional QBO EOFS could benefit attribution of 744 ozone variability in the middle stratosphere (layers 4 and 5) in the tropical latitudes (reduced errors in MLO/Hilo 745 trends) and in the upper stratosphere (layer 8 in COH and in some Umkehr trends) in the NH middle latitude stations 746 (Arosa/Davos, OHP, Boulder) related to the global circulation pattern that are also represented by the Heat Flux proxy. 747 A slight reduction in the errors at SH middle latitude (sonde at Lauder, New Zealand) could be invoked by the EqLat 748 variability that has a small correlation with the QBO-D proxy and sampling bias. Reduction of SE in the trend fit of 749 the layer 5 ozonesonde (up to 2.8 %) and COH (up to 3.0%) records at OHP is not found in the Umkehr results, which 750 suggests overfitting and sampling bias (see results in Appendix D). 751 The addition of extra QBO terms slightly improves the adjusted  $R_{a}^{2}$  model fit (see Figure X, c) for all COH station 752 overpass records in layer 8 (except at MLO) and occasionally improves Umkehr adjusted R<sup>2</sup> (Boulder and Luader).

753 The most significant improvement is found at MLO in layers 3–5 in all three instrument records.

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762 Figure X c) the same as a), but Extended model includes 2 extra QBO terms as an additional proxy.

#### 763 5.4 Arctic and Antarctic Oscillations (AO/AAO)

#### 764 AO/AAO proxies reduce SE (<u>blue</u> colored cells) in the lower stratosphere (layers 3 and 4) at Arosa/Davos, OHP, and 765 MLO, although the reduction somewhat differs between the Umkehr and ozonesonde records. At the same time, at 766 Boulder and Lauder the SE does not show an improvement after the addition of the AO/AAO proxy (AAO is used 767 instead of AO at Lauder). In the middle stratosphere (layer 7), a reduction in SE is found over Boulder in both COH and Umkehr records. The addition of AO/AAO proxies improves the SE of the trend at MLO and Lauder but only in 768

Umkehr records, while it worsens the COH SE. At Lauder, the COH SE in layer 6 shows an improvement, but not in 769

770 Umkehr record. Since results in the middle stratosphere (layers 5-7) are not always consistent among different

771 techniques (reductions are not in the same layers) it could indicate statistical model overfit into the record's noise, or

772 vertical smoothing of the Umkehr or COH technique that combines ozone variability in the layer with a portion of

773 ozone variability in the adjacent layers, thus partially or completely reducing the correlation with the proxy.

774 The addition of the AO predictor increases the adjusted  $R_s^2$  in the lower stratosphere at Arosa/Davos, OHP and MLO.

775 Also, a small enhancement of the adjusted  $R_{a}^{2}$  is seen in the middle and upper stratosphere, including in Umkehr layers

776 6 and 7 and COH layers 6, 7 and 9 over Boulder, as well as in Umkehr fit in layers 5-7 at MLO, and at Lauder (AAO)

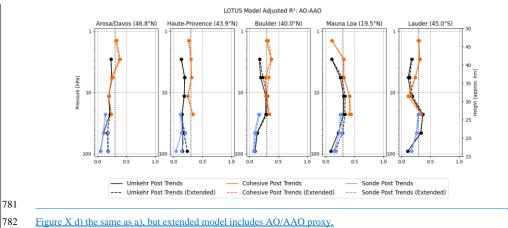
777 for Umkehr and COH records in layer 6. These results are not very consistent across different geolocations, but seem

778 to be consistent across instrumental records at some stations (Umkehr and ozonesonde in the lower layers, and COH

779 and Umkehr in the upper layers). Formatted: English (United States)

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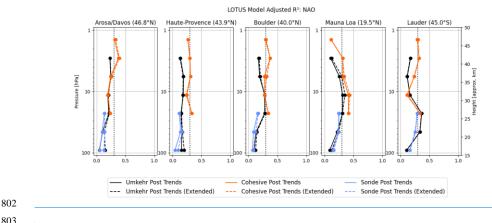
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#### 783 5.5 North Atlantic Oscillation (NAO)

784 Including the NAO proxy in the trend model appears to have a similar pattern (i.e., in latitude and altitude) of changes 785 in the standard error as compared to the result of inclusion of the AO/AAO proxy. It is not a surprise, since indices of 786 the NAO and AO are highly correlated in time due to their common link to the downward propagation of stratospheric anomalies. Standard errors are somewhat reduced in the lower stratospheric layers at the middle NH latitude and 787 788 tropical Umkehr records, but the change is less significant than in AO/AAO cases. The impacts on ozonesonde trend 789 uncertainties are very minimal and inconclusive at Boulder (layers 5, and 4), Hohenpeißenberg (layer 3 and 4), and 790 OHP (layer 3), The impacts on Lauder are similar or stronger (SE is increased for both Umkehr and sonde records) to 791 the impacts of the AO/AAO. In the middle and upper stratosphere, the standard errors are typically increased. The 792 exception is found in layer 7 of the COH record at Boulder and Arosa/Davos, and in layers 6 and 7 of the Umkehr 793 record at MLO. Similar negative results are found when AO/AAO proxies are added, which suggests that the observed 794 time series are overfitted and potentially some instrumental or sampling anomalies are misinterpreted with addition

#### 795 of these proxies.

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$\neg$	Deleted: 5
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$\mathcal{T}$	Deleted: 4
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804 Figure X e) the same as a) but extended model includes NAO as an additional proxy.

#### 805 5.6 Eddy Heat Flux (EHF)

806 The EHF represents a dynamical proxy for assessment of the impact of the Brewer Dobson Circulation (BDC). It is 807 expected to have an impact on the upper stratospheric ozone by accelerating the transport in the upper branch that 808 brings more ozone at higher latitudes (i.e. Arosa/Davos) and middle latitudes (i.e. OHP, Boulder, and Lauder). It could 809 possibly represent changes in the lower branch of the BDC circulation and the expansion of the tropical band, thus 810 modulating ozone in the lower stratosphere at tropics (i.e. MLO). In the Southern middle latitudes (i.e. Lauder), the correlations could be related to the shift in the subtropical wave activities to the higher latitudes in response to the 811 812 ozone hole healing.

813 The addition of the EHF predictor leads to the reduced SE uncertainties in the upper stratosphere in COH and Umkehr

814 trends at OHP and Boulder, and in COH only trends at Arosa/Davos. It has a much smaller reduction of SE for the

- 815 Lauder trend and even an increase in uncertainties if used to fit upper stratospheric ozone time series at MLO. At the
- 816 same time, the SE in the Umkehr and ozonesonde middle stratosphere (layers 4-5) at MLO is substantially reduced,
- 817 including smaller improvements at Lauder. In the lower stratospheric (layer 3) ozone trend SE in Umkehr and sonde
- 818 records at MLO, Lauder and Arosa/Davos are somewhat reduced when using the EHF proxy.
- 819 Addition of the EHF predictor seems to have an impact in the upper stratosphere increasing the adjusted  $R_{a}^{2}$  for COH
- 820 records in layers 8 and 9 in all but MLO or Lauder records, which indicates impact of the BDC upper branch on the 821 middle  $\underline{NH}$  latitudes. In contrast to the COH, the Umkehr adjusted  $R_{a}^{2}$  has not changed significantly, which possibly

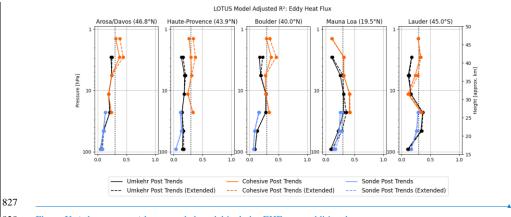
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822
        suggests a high measurement noise in the station records. There is, however, a small increase in adjusted R_{a}^{2} in the
```

- 823 Umkehr record in layer 7 at MLO (whereas COH does not show a change).
- 824 The increase in adjusted  $R_{\lambda}^{2}$  is found at MLO in Umkehr and sonde layers 4 and 5, including a small increase in layer
- 825 3, which probably is related to the EHF-driven changes in the middle stratosphere . Ozone variability in Umkehr and
- 826 sonde records at MLO appears to contain information about the circulation changes in the shallow BDC branch.

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828 Figure X e) the same as a) but extended model includes EHF as an additional proxy

#### 829 6 The Full Model - adding multiple predictors

830 In this paper we seek to develop an improved model and thus trend estimates for point located measurements of ozone

831 through modifications of a model optimized for zonal data. Our criteria for model improvement are based on reduction

832 of the SE of the trend with either improvement (at best) or moderate impact (at worst) on the model adjusted  $R_s^2$ . From

833 the results of the previous section, we see several opportunities to improve the model and improve confidence in the

834 trend estimates. This section examines if the gains of the above are improved while adding several predictors together.

835 As stated above the TP as a predictor exhibits the most consistent results for all stations and measurement techniques.

836 The other predictors have successes in SE reduction, but only at some layers, and some stations. Some results are 837 instrument dependent.

838 Based on the tests above we expect combining predictors can improve the model fit and trend SE reduction, but it is

clear that the predictor selection should vary by station and level. Appendix E details the choices made for the Full 839

840 Model which combines 1 to 3 additional proxies beyond the Reference Model.

#### 841 6.1 Predictors added for the Full Model

842 Reduction of the SE of the trend while improving (or at least not impacting) model adjusted  $R_{A}^{2}$  is the basis of predictor

843 choice for the Full Model. To qualify a predictor should exhibit consistent results for all measurement techniques.

844 Improvement at multiple stations is preferred to single station improvements. In general, we avoid combining highly

correlated predictors. Table 9 shows final choices for the Full Model. 845

846

		LOTUS F	ull Model predictor	selection	
	Arosa/Davos	OHP	Boulder	MLO	Lauder
Layer					

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9	EqLat	EqLat	EqLat	Reference only	EqLat
8	EqLat	EqLat	EqLat	Reference only	EqLat
7	EqLat	EqLat	EqLat	Reference only	EqLat
6	EqLat	EqLat	EqLat	EqLat	EqLat
5	EqLat	EqLat	EqLat	EqLat, QBO CD, AO	EqLat
4	TP, AO	TP, AO	TP	TP, QBO CD, AO	TP
3	TP, AO	TP, AO	TP	TP, QBO CD, AO	TP

848 Table 9: Added predictors for the Full model are tuned for each layer and each station. For layers 7 to 9 the SE and adjusted R<sup>2</sup><sub>x</sub> parameters at MLO are not improved by additional predictors, and the original LOTUS based Reference
 850 Model is used. Appendix E explains the logic of the predictor selection.

#### 851 6.2 Impact of the Full Model on trends,

852 Figure 5a shows the trends for the stations (with COH overpass) for the Reference and Full Models. An impact of the

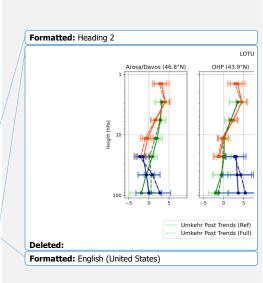
853 Full Model on ozone trends derived in the upper stratosphere (above 16 hPa) is neutral. Addition of proxies to the

LOTUS model does not change trends which remain the same magnitude as those derived using the Reference Model,

855 <u>i.e.</u> positive and statistically significant at the SH and MH middle latitudes and over tropics. The largest difference

(outside of the SE uncertainty) between upper stratospheric Umkehr and COH trends is found over Boulder, MLOand Lauder.

<sup>&</sup>lt;figure><figure>



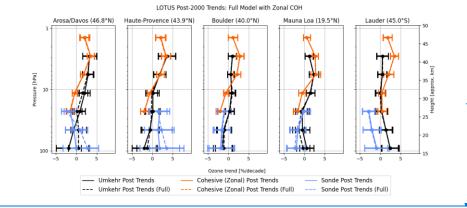
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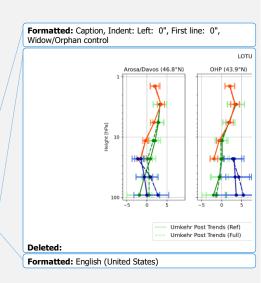
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Figure 5a: Post 2000 trends for the Full and Reference Model. In this figure the COH data shown in orange is the overpass
 data. Solid lines depict Reference Model values (unchanged from Fig. 3). Dashed lines depict Full Model values for all 3
 instrument types.







865

# Figure 5b: Post 2000 trends for the Full and Reference Model. In this case the orange lines are with the zonal data instead of the COH overpass data. Dashed lines depict Full Model values for all 3 instrument types. The Umkehr and sonde trends are unchanged from Fig. 5a.

869 In the middle stratosphere, additional proxies do not change trend values across locations and instrumental records

870 (outside of the SE). At OHP, Boulder and Lauder Umkehr trends in layer 6 (8–16 hPa) are barely positive while COH

trends are negative. At Arosa/Davos and MLO, COH trends in layer 6 are barely negative and Umkehr trends are
 significantly positive. Most COH trends in layer 5 (16–32 hPa) are statistically negative (except at Lauder), while

873 Umkehr trends are near zero.

In the lower stratosphere, Umkehr and sonde trends Arosa/Davos and MLO change after the Full model is used. However, Umkehr and sonde trend changes at MLO are within the SE and therefore can be deemed not significant. Ozonesonde trends at Arosa/Davos in layer 3 (125–63 hPa) change from zero to positive. Umkehr trends at Arosa/Davos in layer 3 change from negative to near zero. Large differences between ozonesonde and Umkehr trends at Lauder and OHP remain unchanged after the Full model is applied although respective SE envelopes overlap.

879 Figure 5b also shows the trends for the Reference and Full Models, but the COH data shown is the associated zonal

data relevant to each station. Incorporation of the additional proxies does not change the trend values for the zonal
 COH data. Impact on error estimates for the trends are discussed next.

#### 882 6.2 Impact of the Full Model on the Trend SE

883 Table 10 summarizes the reduction in the SE for the Full model. Selection of the EqLat predictor for the Full model 884 in the layers 5-9 and for all stations (except MLO/Hilo, to be discussed later) shows the improvement in the SE (as 885 discussed in the previous section). Also, the TP predictor is selected for inclusion to the Full model for trend analyses at Boulder and Lauder stations in layers 3 and 4. The combination of several predictors are used for individual stations 886 887 based on the additional reduction in the SE. For the Arosa/Davos and OHP stations we select a combination of the TP 888 and AO to reduce the SE almost twice as much in some layers. Inclusion of AO proxy is in support of the interpretation 889 of seasonal and interannual ozone variability recorded over stations in Europe that are north of 40 degrees latitudes 890 and are exposed to the seasonal events of ozone depleted air masses transported from the Polar region during the

spring season (Steinbrecht et al., 2011; Manney et al., 2011; Knudsen and Grooss, 2000; Fioletov and Shepherd, 2003;

893 Zhang et al., 2017; Weber et al., 2022a). The strong impact of AO/AAO on the lower stratosphere ozone variability

are not detected in Boulder or Lauder and we choose not to include it in the Full model for trend analyses at these stations.

LOTUS Model Test: Difference [%] in Standard Error: Full Model vs Reference Model Pressure Umkehr Arosa/Davos Haute-Provance Boulder Mauna Loa Lauder UMK COH SND (hPa) Layer 1-2 9 8.4 2.9 1.9 0.0 2.9 2-4 8 -0.5 0.7 0.1 1.2 -0.4 1.5 0.0 0.0 1.0 3.1 4-8 7 3.8 3.2 2.1 0.6 5.4 4.1 0.0 0.0 0.5 1.2 7.8 8-16 6 6.1 8.3 2.5 10.9 2.4 7.8 5.3 7.7 3.4 1.9 13.6 -1.1 16-32 5 7.9 10.6 5.9 13.4 8.7 -1.9 1.4 13.0 13.3 0.8 3.9 32-63 10.0 4 8.7 6.1 3.4 2.7 17.3 10.3 8.0 7.4 9.4 63-127 3 20.3 18.5 13.5 12.8 6.8 2.2 8.0 5.6 6.8 9.8

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896

# Table 10: Change in post 2000 trend SE in the Full Model as a % difference of the Reference Model. Color coding is the same as introduced in Table 7.

899 The MLO/Hilo location is close to the Tropical belt and therefore has different processes impacting stratospheric 900 ozone variability as discussed in the previous section. We find that EqLat proxy can be added to the Full model in 901 layer 6 and 5 (similar to other stations); however, above layer 6, EqLat or TP is not useful for interpretation of tropical 902 ozone variability and therefore we believe the trend model in these layers should remain as it currently is used in 903 Godin-Beekmann et al. (2022) analyses. The EqLat and TP are mildly correlated (-0.4) in the stratosphere, and 904 therefore we decided against combining both of these proxies in the Full model. However, we also found that adding 905 AO and QBO C/D proxies in layers 3, 4 and 5 improved the model fit and reduced the SE. These combined additional 906 proxies are not correlated and reduce SE more than when using them separately. 907 The Full Model showed impacts on the SE in the upper stratosphere (above 8 hPa). The trend errors were reduced 908 with the exception of Umkehr trends at 4-2 hPa over Boulder and Arosa/Davos where errors did not change. No changes in SE are found at MLO with additional proxies, thus the Full Model is kept the same as the Reference Model 909 910 for this station in the upper stratosphere.

911 Similarly, in the middle stratosphere SE were mostly reduced after the Full Model was applied (except for slightly

912 larger SE in trends derived from ozonesonde at OHP and from Umkehr at Boulder).

913 After applying the Full Model in the lower stratosphere, we still found high uncertainty due to higher ozone variability

914 (natural variability), but SE were reduced. Arosa/Davos and MLO Umkehr and sonde trends changed after Full Model

915 was used. Change in ozonesonde trends at HOH in layer 3 (125-63 hPa) goes from zero to positive and trend detection

916 becomes highly confident (p-value <0.05). Umkehr trends at Arosa/Davos in layer 3 changed from negative to near

917 zero but results have low certainty (p-value >0.1). Larger trend differences remain between ozonesonde and Umkehr

918 at Lauder and OHP after the Full Model is applied.

919

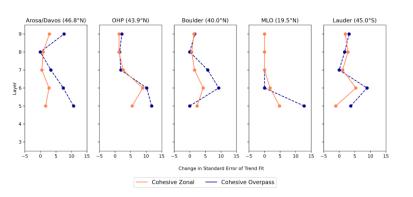
LOTUS Model Proxy Tests: (% difference of SE of Trend): overpass and zonal COH												
Height	Umkehr	Arosa/I	Davos	ОН	Р	Bould	ler	ML	0	Lauder		
(hPa)	Layer	Overpass	Zonal									

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8-16 16-32	6 5	7.35 10.67	2.75 1.74	10.17 11.76	8.98 5.54	9.30 0.00	4.30 2.36	0.00	1.79 4.81	8.93 3.70	5.34 -1.11
8-16	6	7.35	2.75	10.17	8.98	9.30	4.30	0.00	1.79	8.93	5.34
4-8	7	3.39	0.47	1.85	2.55	5.77	1.53	NA	NA	0.00	1.11
2-4	8	0.00	0.90	1.47	1.26	0.00	0.63	NA	NA	1.75	2.76
1-2	9	7.61	2.89	2.20	1.30	1.61	1.34	NA	NA	3.17	1.97

<sup>923</sup> 924

23	Table 11: Change in Standard Error of Trend, as percent of Reference Model SE, for the COH overpass data and zonal
24	data at the 5 ground stations. MLO Full Model in layers 9-7 is the same as the Reference Model (change is marked as NA).



#### 926

# 927Figure 6: Change in Standard Error of Trend, as percent of Reference Model SE, for the COH overpass data (blue) and928COH zonal data (red) at the 5 ground stations.

It is instructive to ponder if the addition of proxies that yield improvements via reduction of the standard error in the localized GB or overpass measurements also have the potential to improve uncertainties in the zonal data. To explore this Table 11 and Fig. 6 show the percent change in SE of the trend when adding the proxies for the Full model. Values are shown for both the COH overpass and the COH zonal data. In general, except when the improvement in the SE for the overpass COH is small (3% or less), addition of proxies has much less impact on the zonal results than on overpass results. This suggests that indeed the Reference LOTUS model is well tuned for zonal datasets, but can be improved with select addition of proxies for overpass or localized GB data.

## 936 **6.3 Impact of the Full Model on adjusted R**<sup>2</sup>

Table 12 shows the adjusted R<sup>2</sup> for the Full Model. In the upper stratosphere, the Full Model increases the adjusted
R<sup>2</sup> above 8 hPa (except in Umkehr at 4–2 hPa). Over MLO there is no change because the Full Model is kept the same
as the Reference Model for layers 7, 8 and 9.

	LOTUS Model Proxy Tests: (Adjusted <u>R</u> of <u>the Full Model</u> )																	
H	leight	Umkehr	Arosa/Davos			OHP				Boulder			MLO			Lauder		
0	hPa)	Layer	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND	
	1-2	9		0.42			0.37			0.36			0.11			0.32		

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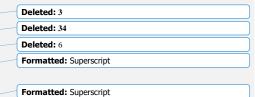
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2-4	8	0.23	0.39		0.14	0.31		0.17	0.39		0.11	0.32		0.18	0.34	
4-8	7	0.35	0.35		0.31	0.41		0.27	0.33		0.26	0.32		0.17	0.27	
8-16	6	0.31	0.35		0.33	0.45		0.33	0.40		0.40	0.51		0.25	0.23	
16-32	5	0.34	0.38	0.26	0.25	0.51	0. <u>2</u> 3	0.31	0.40	0.18	0.44	0.53	0.39	0.42	0.41	0.29
32-63	4	0.23		0.25	0.29		0. <u>27</u>	0.19		0.18	0.42		0.38	0.42		0.31
63-127	3	0.31		0.31	0.44		0.21	0.22		0.11	0.19		0.24	0.25		0.21



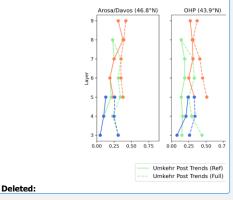
Table 12: Adjusted R<sup>2</sup> of the Full Model. Values of 0.30 and above are indicated in Bold as a threshold to indicate a 941 satisfactory fit. Compare to Table 4 containing values for the Reference Model. 942 In the middle stratosphere (32–8 hPa) adjusted  $R_{a}^{2}$  increases are found in all records (although smaller increases are 943 found in ozonesonde and Umkehr records at OHP, Boulder and Lauder at 32-64 hPa). At Arosa/Davos, Boulder and 944 Lauder the adjusted R<sup>1</sup><sub>a</sub> in the COH and Umkehr trend models increase and continue to be very close in value. The 945 COH adjusted R<sup>2</sup><sub>4</sub> is larger at OHP and MLO than in Umkehr and sonde records thus suggesting that overpass 946 conditions might have smoothed some natural variability observed in the GB records. In general, the adjusted  $R_{a}^{2}$  is 947 the largest at the 32-64 hPa level. This suggests that the Full Model shows an improvement for regional trend analyses 948 in the middle stratosphere.

949 Although Umkehr and sonde trend changes at MLO in the lower stratosphere are within the SE and therefore can be 950 deemed not significant, the adjusted R<sup>2</sup><sub>2</sub> is increased which suggests a better model fit in the Full Model. The adjusted 951  $R_s^2$  increases in both Umkehr and ozonesonde data, while the largest increases are found in the Arosa/Davos, OHP and 952 MLO records. 953



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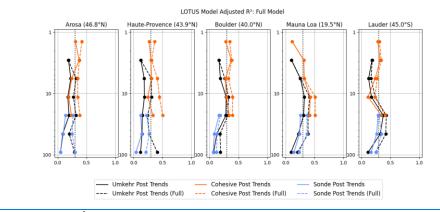
In the lower stratosphere, the adjusted  $R_{k}^{2}$  remains low in both Umkehr and sonde records at Boulder (only TP is added 954 for the Full model). While the p-values at 63-32 hPa are significantly reduced (see discussion in the next section), 955 they still remain relatively high. These results suggest that additional research is needed to identify the best set of 956 proxies for Boulder records in the lower stratosphere. At Lauder, the ozonesonde record shows smaller adjusted  $R_{A}^{2}$  as 957 compared to Umkehr partially due to low sampling biases. 958 It is valuable to further explore the impact of the Full Model on the adjusted  $R_{a}^{2}$  for the zonal and overpass COH data. 959 Fig. 7a shows the adjusted  $R_{k}^{2}$  for the Reference and Full Models at each of the 5 stations using the COH overpass 960 data. In all cases the Full Model improves the adjusted  $R_{a}^{2}$  except for MLO layers 7, 8 and 9 where the Full and 961 Reference Model are identical. The most significant improvements are seen by Umkehr at layers 3 to 7, COH overpass 962 at Layers 5, 6 and 7, and sonde layers 3-5. Figure 7b shows similar results using COH zonal data instead of overpass. 963 There is practically no further improvement in the adjusted  $R_{a}^{2}$  for the zonally averaged COH results (except for a

964 small increase for MLO layer 5). Comparison of results reveals that for OHP the implementation of the Full model 965 for the COH overpass data (Fig. 7a, dashed line) improves the adjusted  $R_{a}^{2}$  to values nearing that of the Reference Model zonal data in layer 7 and below (Fig. 7b, solid line). For MLO and Lauder the use of the Full Model on the 966

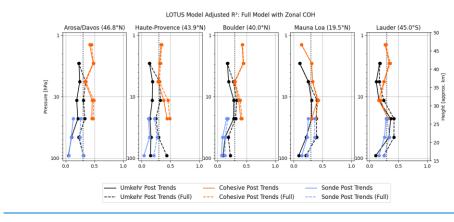
967 COH overpass data improves the adjusted  $R_k^2$  over the Reference Model beyond the improvement seen in the COH

zonal results for layers 5 and 6. At Arosa/Davos and Boulder the implementation of the Full Model does not fully 968 969 reach the magnitude of the COH zonal adjusted R<sup>2</sup>.

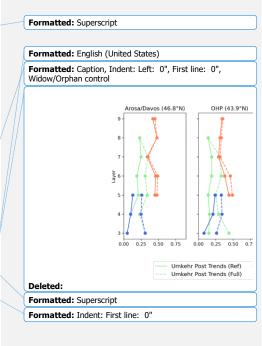
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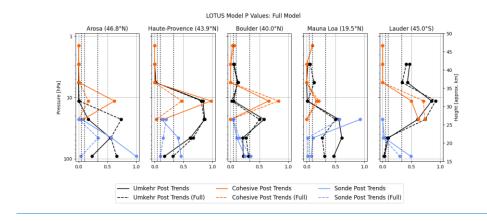






981 Figure 7b: Adjusted R<sup>2</sup> for the Full Model (dashed lines) and Reference Model (solid lines) at 5 stations. The COH data in this figure is the zonal data for each station. The Umkehr and sonde lines are identical to those in Fig. 7a.





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# Figure 7c: the same as b, but for the p-values. Vertical dotted lines indicate limits for the high (<0.05), medium-high</li> (between 0.05 and 0.1), medium (between 0.1 and 0.3) and low confidence (>0.3).

#### 987 6.3 Examination of the p-values of the Full Model

988In the upper stratosphere (above 8 hPa), the confidence in Umkehr trends remained high (see Figure 7c) for most989stations except at Boulder (medium to low) and Lauder (very low, although some improvement was found). COH

990 trends confidence was very slightly degraded over Boulder at 1–2 hPa, but mostly has not changed.

991 In the middle stratosphere (between 32 and 8 hPa), p-values were significantly reduced in COH records. At 8–16 hPa

remained high, but at 16–32 hPa the confidence improved (continued) to high over Arosa/Davos and OHP (Boulder

and MLO). In case of Umkehr analyses in layer 8–16 hPa at Arosa/Davos, Boulder and MLO the confidence remained

high. However, at 16–32 hPa the Umkehr trend detection confidence was degraded over Arosa/Davos and Lauder.

For the ozonesonde record, the p-values remained low (<0.05) except at MLO where some improvement was found after the Full Model was used, but the p-value remained high. It suggests that some instrumental records have either

after the Full Model was used, but the p-value remained high. It suggests that some instrumental records have either high atmospheric or instrumental noise and therefore perhaps high certainty in trend detection cannot be achieved with

- 998 linear trend models. For near zero trends with high variability, the p-values are not a good criterion for trend 999 detectability.
- 1000 In the lower stratosphere (between 125 and 32 hPa), analyses of p-values for the Full Model fit show significant
- 1001 improvement for Umkehr trends at MLO between 63–32 hPa (while the p-value was increased at other stations at this

1002 level). In addition, improvement in p-values was found for ozonesonde trends at all stations. Specifically, very low p-

values for the Full model were reached at Arosa/Davos (125–63 hPa), OHP (125–63 and 63–32 hPa), MLO (125–63

1004 and 63–32 hPa), and Lauder (63–31 hPa).

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### 007 7 Summary of the Full Model findings.

We find that upper stratospheric trends in COH overpass and Umkehr records detect ozone recovery with high
confidence (p<0.05) above 8 hPa (with the exception of near-zero positive Umkehr trends over Lauder and Boulder).</li>
We note the largest difference between Umkehr and COH trends (outside of the SE uncertainty) at Boulder, Mauna
Loa and Lauder.

1012 Confidence for the middle stratosphere (between 32 and 8 hPa) trends vary between high, medium and low. Although 1013 most of the trends are narrowly different from zero (especially when error bars are considered), there are some 1014 differences in results across instrumental groups: trends in COH and sonde (except at OHP) between 32 and 16 hPa 1015 tend to be small negative, while Umkehr trends are slightly positive. Some trends are statistically different from zero. 1016 However, instrument-specific error bars often overlap and thus making differences in trends not significant.

1017 Confidence in lower stratosphere trends is highly variable and even lower than in the middle stratosphere due to higher 1018 ozone variability unaccounted for by Solar, QBO and ENSO proxies used in the Reference Model. However, high 1019 confidence (p<0.05) is still found in ozonesonde trends at Arosa/Davos, OHP, MLO and Lauder (although not at all 1020 layers). Umkehr trends in the lower stratosphere show lower confidence than ozonesonde trends (except at Lauder 1021 and Arosa/Davos in the lowermost altitudes). The low confidence levels could be related to the near-zero trends 1022 derived from Umkehr data, whereas ozonesonde trends are often different from zero lines. Also, we apply AK-1023 smoothing to the sondes to account for the wide AKs in the Umkehr retrieval. We tested the impacts of the AK on 1024 ozonesonde trends (see Appendix A) and did not find any significant impacts. Most notably, ozonesonde and Umkehr 1025 trends significantly disagree in the lower stratosphere at OHP and Lauder and therefore require further investigation. 1026 The instrumental drifts and differences in Lauder trends are also discussed in Bjorkland et al. (2023 preprint) and are 1027 consistent with our findings.

### 028 8 Conclusions.

This paper is a follow up to Godin-Beekmann et al. (2022) with a focus on the GB record trend assessment. Therefore, our trend analyses focus on the questions:

- Do proxies for evaluating trends of GB stations need to be different from those of the optimized set for zonal data?
- 1033 2) Are station records representative of the small geophysical region or semi-global changes?
- 1034 3) Do uncertainties of the zonal averaged trends improve with additional proxies?
- 1035

1036The Full Model developed in this paper for station and overpass data adds proxies to the LOTUS models of Godin-1037Beckmann (2022). Our trend analysis of stratospheric ozone records from the Umkehr, ozonesonde and COH station1038overpass data at 5 geographical regions using the Full Model (LOTUS v 0.8.0) show similar trends to those published1039in Godin-Beekmann et al. (2022) paper. We analyze trends for instrumental records converted to 7 Umkehr layers that1040represent ozone changes in the upper, middle and lower stratosphere over NH and SH middle latitudes and over high

1041 tropics of the NH. We also analyze GB station records at Arosa/Davos, Hohenpeißenberg and OHP separately in 1042 contrast to the "European regional" trend analyses presented in Godin-Beekmann et al. (2022) and included COH 1043 overpass records for comparisons with the GB records. Our analyses include evaluation of the adjusted  $R_{k}^{2}$  (aka 1044 goodness of the model fit), standard error and p-values. 1045 We also investigate differences between satellite trends as detected in the records sampled for individual geographical

locations (spatial and temporal overpass criteria) versus zonal average datasets. We find that COH overpass ozone
records capture ozone variability of the ground-based station records (Umkehr and sonde) better than COH zonal data.
We do not find that the COH zonal record is improved by using EqLat instead of geometric latitude to construct the
dataset (see Appendix C), but EqLat can be an important additional proxy at some levels for GB data. To determine
the improvement to the model fit we use the Standard Error and adjusted R<sup>2</sup> for the Full and Reference model fit.

Using the Reference model for the zonal mean COH data we find slightly better adjusted R<sup>2</sup>

than for the COH overpass data fit over the Northern middle latitude stations. This is expected as much of the variability of the overpass time series is reduced in the zonal average data. Therefore, we also explore the impact of additional predictors in the trend model fit applied to the more variable GB and satellite COH overpass data to

determine if that will reduce the SE and improve the adjusted  $R_{k}^{2}$ . We also apply the Full model to the zonally averaged data to assess the benefits of additional proxies to further reduce trend uncertainties.

1057 We find that adding predictors (with few exceptions) does not change the trends but often reduces SEs and increases 1058 the adjusted  $R_a^2$  (with the exception of the upper stratospheric ozone trends at MLO). We also find that the p-values 1059 are useful for interpretation of improvements of the model fit in the data, although improvements in the SE do not 1060 always result in improved confidence in derived trends, especially when the trends are close to zero. In these cases we 1061 conclude that either longer records are needed to discern trend information outside of the atmospheric noise or further 1062 research into the inconsistencies between instrumental records and homogenization procedures is required. We also 1063 find the small changes in trends in the lower stratosphere and improvements in the model fit after additional proxies 1064 are used. However, the sampling tests indicate that trends can depend on the temporal selection of the records when 1065 AK are used to smooth ozonesonde high resolution profiles (see discussion in Appendix D).

1066 This paper concludes that additional proxies bring improvements to trend detectability for GB and gridded satellite 1067 data analyses and better agreement is achieved between satellite overpass and GB trends. We also find that zonally 1068 averaged and gridded satellite records produce comparable trends over the studied middle latitudes and subtropical 1069 regions. Therefore, the GB trends are representative of the stratospheric ozone changes over the semi-global area. 1070 Finally, zonally averaged data do not benefit from addition of proxies beyond what LOTUS model uses for global 1071 trend detection whereas the uncertainties in GB and gridded trends are significantly reduced and sometimes (Boulder, 1072 MLO, Lauder) become comparable to the uncertainties of the zonally averaged trends in the upper and middle 1073 stratosphere. Based on analyses presented in this paper we strongly recommend using additional proxies for trend 1074 analyses of GB and gridded satellite stratospheric ozone records. Additional proxies should be selected based on the 1075 latitude and altitude of the observational ozone record to adequately represent stratospheric transport and mixing 1076 processes impacting interannual and seasonal ozone variability.

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### 1078 Appendices

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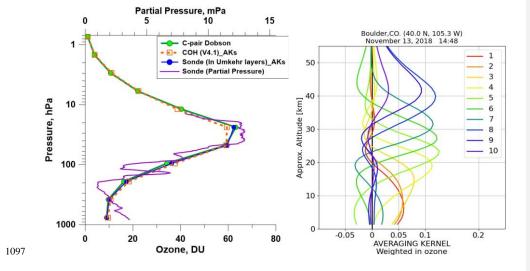
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## 1079 Appendix A: AK Smoothing for ozonesondes

1080 Ozonesonde profiles have high vertical resolution (purple line in Fig. A1) in comparison to the Umkehr (green solid 1081 line) or COH (orange dashed line) ozone profiles. Each Umkehr layer is referenced to the atmospheric pressure at the 1082 bottom of the layer, which is constructed using half of the pressure in the layer below. Averaging Kernels (AK) as 1083 shown in Fig. A1, panel b, define the granularity of the Umkehr vertical grid. In order to compare trends from three 1084 instrumental records in the same vertical system, we convert the ozonesonde and COH profiles to the Umkehr layers 1085 and DU. The COH overpass data is in units of DU, but on different layers than the defined Umkehr layers, so only 1086 vertical grid modification is required. The sonde profiles (purple thin line) are in units of partial pressure and are first 1087 converted to DU, then converted to the Umkehr grid (blue solid line in panel a). Conversion to the Umkehr grid can 1088 be done either by interpolation, or by AK smoothing. The equation describing the process of applying AK smoothing 1089 is

 $Ozone_{smoothed}(i) = Ozone_{apriori}(i) + \sum_{j} AK_{ij} \{Ozone_{true}(j) - Ozone_{apriori}(j)\}$ 

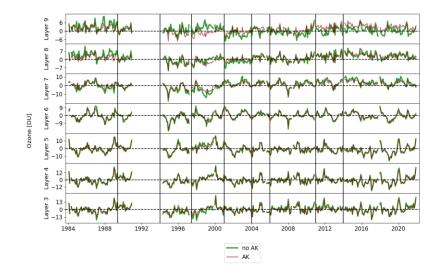
where AK is the Averaging Kernel for layer i,  $Ozone_{smoothed_j}$  is the smoothed ozone result,  $Ozone_{true}$  is the ozonesonde profile, and  $Ozone_{apriori}$  is the Umkehr a priori (climatological) profile. The AK for each Umkehr layer is used as a weighting function applied to the ozonesonde profile ( $Ozone_{true}$ ) prior to the integration which simulates the Umkehr optimal estimation method used for estimating the ozone content in the targeted layer (Rodgers, 2000).



 $Ozone_{smoothed}(i) = \sum_{j} \{AK_{i,j} * Ozone_{tru}$ Deleted: 1100Figure A1: a) An example of ozone observations over the Boulder, CO station. The purple line is 100-m averaged ozone1101partial pressure (hPa) vertical profile measured by sonde on 13 November, 2018. The green line with solid circles is the1102ozone profile derived from Dobson Umkehr observations on the same day. The blue line with blue dots is the ozonesonde1103profile converted to the Umkehr layers and smoothed with the Umkehr AK. The orange dashed line with open squares is1104the COH ozone profile derived from observations in Boulder on 13 November, 2018. Each line represents the1105smoothing function for one of 10 Umkehr layers (see color legend).

1107 Although the ozonesonde measurement typically reaches altitudes between 32 and 10 hPa, the balloon often bursts 1108 before reaching the top of layer 6 (16 hPa), therefore only partially covering the ozone content in that layer. We also 1109 note that Umkehr AKs are relatively wide and therefore will incorporate (weight in) ozone variability from the layer 1110 above and layer below of the targeted Umkehr layer. (See layer 6, green line in Fig. A1, panel b.) Therefore, there 1111 are two sources of error in ozonesonde comparisons with Umkehr ozone in layer 6: a) burst level for ozonesonde does 1112 not reach the top of the layer 6, thus the integrated ozone is smaller than expected. b) the Umkehr AK for layer 6 is 1113 relatively wide and therefore the Umkehr layer partially contains information from above the burst altitude of the 1114 ozonesonde, thus making smoothed ozonesonde concentration lower than expected. In order to avoid these errors, we 1115 only show ozonesonde results up to layer 5. 1116 Similarly, we explored smoothing COH profiles with Umkehr AKs. Figure A2 demonstrates the time series of the 1117 COH ozone over the Mauna Loa station. The trend model was fitted to the COH record with and without AK applied. 1118 The reference trend model included proxies and trends. To focus on ozone variability that contributes to the trends we 1119 subtracted the modeled ozone variability from the COH data and then added the trend component back. The COH 1120 record residuals in Fig. A2 are shown in Umkehr layers where COH is either smoothed with AK (red lines) or not 1121 (green lines). We notice that the AK-smoothing of the COH profile in layer 9 does not have a lot of independent 1122 information from layer 8. In this example it clearly shows that the trends in layer 8 are embedded in the COH layer 9 1123 ozone time series, which was confirmed when we compared trends derived from the AK-smoothed COH in layers 8 1124 and 9. In case of the integrated COH ozone record, the trends in layers 8 and 9 differed. In order to avoid biasing the 1125 COH trends at layer 9 we decided to not apply Umkehr AKs for COH smoothing and only use COH profiles 1126 interpolated into the Umkehr layers. This result makes sense since COH overpass data are derived from UV 1127 backscatter radiances also using an Optimal Estimation technique. COH overpass data has a comparable vertical 1128 resolution to Umkehr, simply with different layer definitions. Interpolation makes the most sense for rendering COH 1129 data in the Umkehr vertical coordinate system.





### 1130

1131Figure A2: Modified residuals (seasonal cycle, Solar, QBO, and ENSO are removed, but trend is retained) of COH overpass1132data at Mauna Loa (20N, 156W). Red: AK smoothed to Umkehr layers; Green: Interpolated to Umkehr layers. Vertical1133lines show the dates of satellite records in COH. The largest impact of the AK is seen between 1997 and 2001 where two1134curves separate in layers 7, 8 and 9, and also after 2001 in layer 9.

1135 Figure A3 demonstrates time series of monthly mean ground-based records the lower stratosphere at 5 stations. The

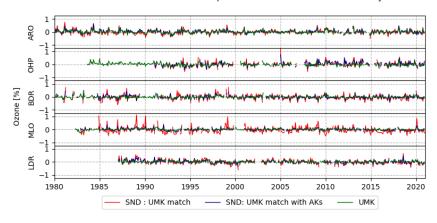
1136 Umkehr data (blue) are compared with the ozonesonde anomalies either interpolated to the Umkehr layer 3 (green),

1137 or ozonesonde profiles matched with Umkehr profiles in time and smoothed using the Umkehr averaging kernels

1138 (crimson). All three datasets have been deseasonalized using their respective climatological (using 1998-2008

1139 climatology) average monthly mean ozone. The application of the Averaging Kernels has the effect of smoothing the

1140 temporal variability.



Deseasonalized Time Series: Comparisons between UMK and SND : Layer 3

### 1141

1142Figure A3: Time series of monthly averaged and de-seasonalized (in %) ozone anomalies of Umkehr (green) and1143ozonesondes records are compared at 5 ground-based stations. Ozonesonde data are either calculated using only profiles1144that are interpolated in Umkehr layer 3 (blue) or matched with Umkehr profile in time and smoothed with the Umkehr

1145 averaging kernels (crimson).

# 1146 Appendix B: COH using OMPS v3r2 vs OMPS v4r1

1147 OMPS SNPP v4r1 uses updated SDRs as input which incorporate unified and consistent calibration algorithms

1148 removing artificial jumps caused by operational changes, instrument anomalies, or contamination for anomaly views

1149 of the environment or spacecraft. Also included are new interpolated band-passes, and updated soft calibration based

1150 on the new input SDR's.

1151 Differences between the v3r2 and v4r1 versions of the resulting COH dataset are typically less than 1 percent (Fig.

1152 A4 and A5). Small seasonal variation is apparent at all levels. Larger differences are visible in 2020 when the soft

1153 calibration for v3r2 is extended beyond its period of relevance. Figure A6 shows the drift between the two versions.

1154 Drift between the datasets is less than +- 1% at all levels. This is a reasonable estimate of the resulting expected trend

1155 difference in using the newest COH version as compared to the v3r2 results used in Godin-Beekmann (2022).

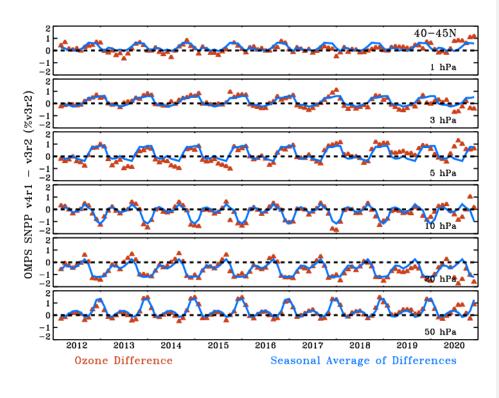
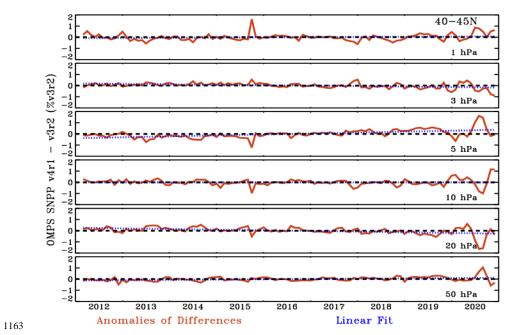
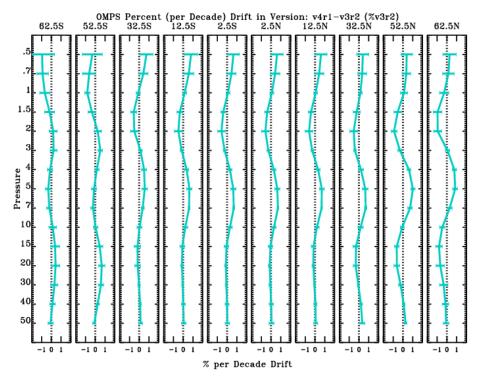




Figure A4: Differences in the COH monthly average zonal product as generated from SNPP v4r1 and v3r2 processing. Also shown is the annual cycle in this difference as depicted by the average over all years for each month. Exhibited at 40-45N is a less than 2% difference with an annual cycle. A somewhat different pattern is seen in 2020 where the soft calibration for v3r2 is extended beyond its period of relevance. 1159 1160 1161



1164 1165 1166 Figure A5: Anomalies of the differences in version (v4r1 vs v3r2) in the COH monthly average zonal product at 40-45N. Anomalies are enhanced in 2020. Also shown as a blue dotted line is a linear least square fit to the anomalies representing the drift between the two versions.

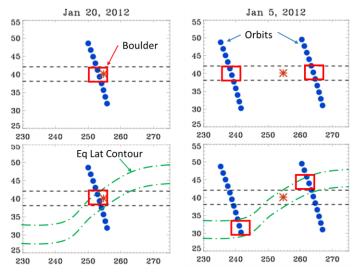


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#### 1169 Appendix C: Impact of using equivalent latitude in generation of the COH product

1170 The COH overpass data used in this paper collects all profiles during the day within a latitude and longitude box of 1171 +/- 2 degrees by +/- 20 degrees, then generates a 1/distance averaged value for the station. The box is based on 1172 geometric latitude and longitude. With 15 orbits per day, the chosen box size guarantees 2 to 4 possible profiles within 1173 the box depending on whether the orbit overpasses or straddles the site as shown in Fig. A7. Also shown is a scenario 1174 when the equivalent latitude (EqLat) near the site is particularly non-zonal. In such cases the profiles selected using 1175 a geometric coordinate box will select SBUV profiles from an Eq Lat that is different from that of the measurement 1176 station. 1177



1179Figure A7: Shows orbits of SNPP and positions of OMPS NP ozone profiles on January 20, 2012 and January 5, 2012. The1180second row displays a possible EqLat contour overlaid.

1181 It is informative to create an overpass product using boxes based on EqLat and determine the impact on the data.

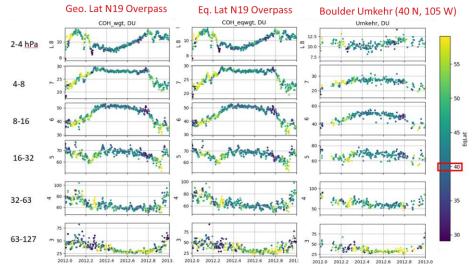
1182 Since EqLat is layer dependent, the included profiles must be selected independently for each layer. Figure A8 shows

1183 COH overpass data for Boulder using geometric coordinates, EqLat based coordinates, and the associated Umkehr

1184 data. Color coding shows the EqLat at Boulder for each measurement day with dark blue and yellow indicating days

1185 with extreme variation from 40N.

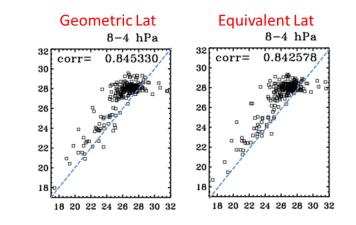
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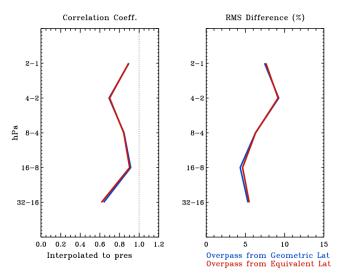
1189Figure A8: COH overpass data generated with geometric coordinates, EqLat based coordinates, and the associated Umkehr1190dataset at Boulder for 2012. Data points are color coded for the EqLat at the measurement site. Boulder is at 40 N.

1191 Variation in EqLat is most apparent in Winter months and transitional Fall and Spring, less so in Summer. Yet the 1192 value of the COH ozone is not dramatically altered in the time series. Figure A9 shows correlation plots of the COH 1193 overpass to Umkehr for the data at layer 7 (4-8 hPa). The pattern of the scatter and the value of the correlation 1194 coefficient are not substantially altered for overpass determination using geometric latitude (left) and EqLat (right). 1195 Figure A10 shows the vertical distribution of the Correlation coefficient and the RMS Difference for the two COH 1196 datasets vs Umkehr. These two metrics are minimally impacted for this sample year in the layers where COH is valid. 1197



### 1199 Figure A9: Correlation between Umkehr and COH overpass using Geometric Latitude (left) and EqLat (right) to select 1200 included profiles for layer 7 (4-8 hPa).

1201



# 1202

1203Figure A10: Profiles of Correlation coefficients and RMS differences between COH overpass data at Boulder for 2012 using1204Geometric Latitude (blue) and EqLat (red) to select data points included in the average.

1205

# 1206 The use of geometric latitude appears to be sufficient in the choice of included data points in the overpass COH product

# 1207 at the layers used in this paper. Likely this is a ramification of the smooth horizontal resolution of the sate llite product.

# 1208 Appendix D: Temporal Sampling and Impact on Trends

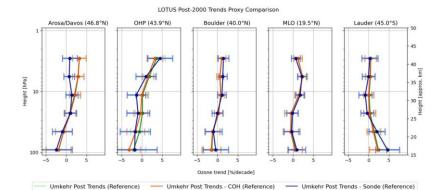
1209 This paper compares trends for three instrument types each with differing measurement frequency. From each set of 1210 measurements а monthly average is constructed. See the data files at 1211 https://gml.noaa.gov/aftp/ozwv/Publications/2023\_Umkehr\_Ozone\_Trends\_Paper/ for the data and the number of data points in each monthly average with the sampling variations. Umkehr measures once or twice per day depending 1212 1213 on cloud interference with the measurement. At Arosa/Davos and Lauder, Umkehr measurements are sparser than 1214 the other GB stations, often less than 10 per month. At Boulder beginning in 1983 measurements number 20 or more per month. At OHP the Umkehr record begins in 1983 with a strong 20 or more measurements per month. From 1215 1999 to 2016, however, measurements per month are often less than 15 per month. The most Umkehr measurements 1216 1217 at MLO are the most abundant, especially after 1985 measuring multiple times in a day, resulting in 50-70 data points 1218 contributing to the monthly average. The COH overpass dataset is typically available once per day at each station with 1219 occasional misses, contributing usually 27-30 data points per month. Since Umkehr can measure multiple times per 1220 day, the COH data matched to Umkehr can contain more profiles in the monthly average than the original full COH data, since the COH overpass data will appear twice in the monthly average, once per each Umkehr measurement. 1221

1222 This occurs often at MLO. Ozonesonde launches are typically one to three times per week depending on the station. 1223 At Arosa/Davos, sonde measurements are typically about 15 per month. Sonde measurements at the other stations 1224 usually have approximately 5 measurements per month, with some periods of up to 10 per month. As with COH 1225 overpass measurements, the sonde dataset matched to Umkehr can have more contributions to the monthly average 1226 resulting from dates with more than one Umkehr measurement, resulting in multiple sonde matches.

1227 The trend results in this paper use all available Umkehr data to generate the monthly means. The COH and sonde data 1228 are matched to Umkehr to use the Umkehr temporal sampling for COH, and to be able to use the Umkehr averaging 1229 kernels for sonde. It is important to determine how the temporal sampling within the monthly mean data may impact 1230 trend results. To aid this understanding, we create three subsets of Umkehr data each with different temporal sampling 1231 and create the corresponding monthly mean: 1) all observations in Umkehr record; 2) Umkehr matched to the COH 1232 dataset; and 3) Umkehr matched to the sonde dataset. In this way we use the same data, but only vary the temporal 1233 sampling. Since the COH is measured every day, except in the rare case that the satellite data is missing due to 1234 instrument issues, sampling 1 and 2 should provide nearly identical results. We expect a strong change in the monthly 1235 mean and resulting trends for Umkehr record when it is matched with infrequent sampling of ozonesonde profiles 1236 (especially in Boulder, Hilo and Lauder).

1237 Figure A11 summarizes the results. Each line in Fig. A11 is trend derived from Umkehr data, but with sampling of all 1238 data, data matched to COH dates, and data matched to sonde dates. In general, the differences are within the envelope 1239 of trend uncertainty (+/- 2 std errors). As expected, the trends and standard errors for all (green) and COH-matched 1240 subsampled (orange) Umkehr records are nearly the same. The largest differences in all Umkehr and COH matched 1241 Umkehr lines are apparent at OHP. We have determined that this arises from occasional months when there is a short 1242 satellite outage coupled with sparse Umkehr observations at the station. However, trends derived from sonde-matched 1243 Umkehr data (blue) show deviations from other observations. This is especially clear at Arosa/Davos in the upper 1244 stratosphere (~2-3 % above 10 hPa). But since this is above the measurement capability of the ozonesonde, this will 1245 not impact the ozonesonde trend results at Arosa/Davos. At Lauder the most significant differences are seen in layer 1246 3 (2.5%), but unfortunately not in the direction to explain sonde differences in the Lauder trend curves as compared to 1247 Umkehr. Smaller differences are seen at other layers (very small, less than 1 %, differences in layers 6 and 4). At

1248 OHP small differences of less than 1 % are seen between 50 and 10 hPa, well within error estimates.

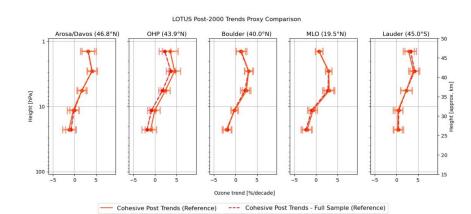




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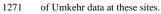
1250Figure A11: Trend results for the Reference Model using Umkehr data mimicking the temporal sampling of COH and1251sonde. Green is all available Umkehr data; orange is Umkehr data matched to COH measurements dates; blue is Umkehr1252data matched to sonde measurement dates.

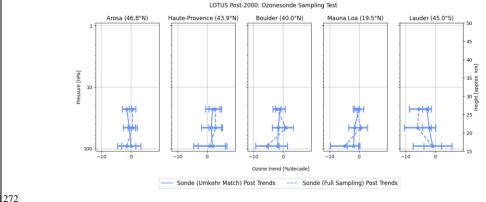
Figure A12 further explores sampling differences by examination of trends of COH data using the full COH dataset, and data sampled to the Umkehr dates in generation of the monthly mean datasets. As with Fig A11, the trend lines are nearly identical at all stations except OHP. At OHP in the early 2000's there are significantly fewer COH points matched to Umkehr because of the drop in Umkehr measurements. This likely impacts the post-2000 trend estimate. The differences remain below 2%, and are within the error estimate of the trends. In summary, the sampling biases between COH overpass and Umkehr data cannot explain the difference in the derived trends (see Fig. 3, most notable in layers 7 and 8 at Boulder and Lauder).



1262Figure A12: Trend results for the Reference Model exploring variations in sampling of the COH data. Solid orange is COH<br/>data matching Umkehr sampling; dotted orange is all available COH data.

Figure A13 explores the impact on trends from sampling differences of the sonde data. Shown are trends with all sonde data, and trends with Umkehr matched data. In this figure only, the sonde data is not AK smoothed since the Umkehr AK are only available on dates when there is an Umkehr measurement. So shown here are trends from sonde data integrated to the Umkehr levels. As with Fig. A11 the only visible impact is seen at OHP and Lauder, though both are within error estimates. At Lauder the trends remain negative for both samplings, but sonde sampled to Umkehr moves closer to the zero line. At OHP the sonde trends are positive, but sonde sampled to Umkehr moves slightly closer to zero. The sampling impact on trends for both OHP and Lauder are likely due to the reduced number







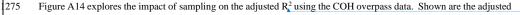
eight [hPa]

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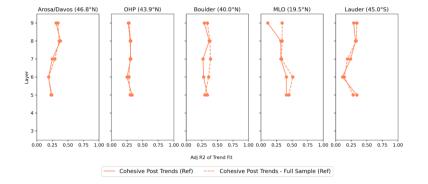
OHP (43.9°N)

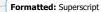
Arosa/Davos (46.8°N)

Figure A13: Trend results for Reference model exploring sampling of the sonde data. Solid blue is all sonde data; dashed is Umkehr matched sonde data.



- $R_{\star}^{2}$  for all available COH overpass data, and the same using only COH overpass with matches to the Umkehr data. For
- Arosa/Davos, OHP and Lauder the differences are small. For Boulder and MLO at some layers (Boulder, layers 6,7;
- MLO layers 6,9), the impact is more apparent with the Full COH exhibiting higher adjusted  $R_{a}^{2}$  at these stations.





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 1282
 Figure A14: Adjusted R<sup>2</sup> for the Reference Model exploring variations in sampling of the COH data. Solid orange is COH data matching Umkehr sampling; dotted orange is all available COH data.

### 1284 Appendix E: Decision process for the Full Model

1285 The LOTUS styled Reference Model is developed and optimized for zonal average datasets. The Extended Model

1286 tests the addition of single predictors to see if fit statistics can be improved for GB and overpass datasets. For 1287 Tropospheric Pressure (TP), improvements are consistent among layers and among instrument types. The addition of

1288 EqLat also yields consistent results for instrument types and at most stations, though not Mauna Loa. Addition of

1289 other predictors gives mixed results depending on level and station. The potential for improving confidence in trend

- 1290 results exists by combining predictors using different choices depending on layer and station. We choose additional
- 1291 predictor combinations with consideration of three criteria: 1) combined predictors should not have a high correlation

1292 with each other (usually .2 or less); 2) predictors should reduce the SE of the trend consistently for all instrument

types; 3) addition of the predictor should not greatly reduce the adjusted  $R_{A}^{2}$  of the model fit, but preferentially increase

it. As seen in Tables 7e and 7f the NAO and the EHF predictors do not make a significant improvement when added

1295 to the Reference Model, so we do not include either in the Full Model.

### 1296 Mixed Model:

- 1297 We have noted a high correlation between the TP and EqLat predictors at all levels especially for Boulder, Mauna 298 Loa and Lauder with correlation adjusted  $R_s^2$  of .4 to .7, and somewhat less correlated at Arosa/Davos and OHP with 299 adjusted  $R_{k}^{2}$  of .2 and .3. Subsequently, we choose to not use these two predictors together (at the same station/layer 1300 combination). The addition of TP at all stations for layers 3 and 4 uniformly decreases the standard errors at all 1301 stations for both Umkehr and sonde. The addition of EqLat (with the exception of Umkehr at Boulder, level 5) almost 1302 uniformly decreases the standard errors at all stations for layers 5 and 6. There is additional reduction in the SE for 1303 layers 7 to 9 for all stations except at Mauna Loa. Thus, we choose TP and EqLat as additional predictors at these 1304 layers. QBO C and D, have significant impact in decreasing the SE in layers 4 and 5 for both Umkehr and sonde, and 1305 layer 3 for sonde with only a small degradation for Umkehr. QBO-CD shows an improvement in layer 8 at OHP, both COH and Umkehr, and Arosa/Davos and Boulder for COH only. We have tested adding both QBO and EqLat for 1306 1307 layer 8 at these 3 stations. For Umkehr measurements, there is no improvement beyond EqLat only with QBO-CD 1308 also included. For COH there is additional improvement, but not to the extent of QBO-CD alone. Since the 1309 improvement is limited to one layer, and for only COH, we choose to only add the additional QBO-CD for the tropical 1310 MLO. Table A1 shows the resulting combination of additional predictors for this Mixed Model. 1311
  - LOTUS Mixed Model Arosa/Davos OHP MLO Layer Boulder Lauder 9 EqLat EqLat EqLat Ref EqLat 8 EqLat EqLat EqLat Ref EqLat 7 EqLat EqLat EqLat Ref EqLat

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6	EqLat	EqLat	EqLat	EqLat	EqLat
5	EqLat	EqLat	EqLat	EqLat, QBO CD	EqLat
4	TP	TP	TP	TP, QBO CD	TP
3	TP	TP	TP	TP, QBO CD	TP

1312

### 1313 Table A1: Details of additional predictor combinations for each level and station in the Mixed Model

1314 The resulting change in SE from the Reference Model is shown in Table A2. For most stations/layers this is simply a

1315 composite of the values from the single EqLat or TP Extended Model results. There remain a few

316 instrument/station/layers where the SE is slightly increased - Arosa/Davos Umkehr layer 8 and Boulder Umkehr layer

8, but these are negligible. At Boulder Layer 5 Umkehr the increase in SE is somewhat more at 1.85% difference, but

1318 this is still small enough to not be of great concern. For Mauna Loa at layers 3,4 and 5 the model is rerun adding two

319 predictors together and the results are new. Indeed, in these cases the SE is improved beyond the single predictor

1320 results of either QBO alone, or TP or EqLat alone with the exception of Sonde layer 5 where the change in SE is just

slightly degraded from QBO alone (13.42% vs 13.69% reduction in SE).

		LOTU	S Mode	l Test: D	Differen	ice [%]	in Stand	dard Eri	or: Mix	ed Mod	el vs Re	ference	e Model			
Pressure	Umkehr	Are	osa/Dav	/0S	Haut	te-Prov	ance		Boulder		М	auna Lo	ba		Lauder	
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND
1-2	9		8.4			2.8			1.9			0.0			2.9	
2-4	8	-0.5	0.7		0.1	1.0		-0.4	1.5		0.0	0.0		1.0	3.1	
4-8	7	3.8	3.0		2.1	1.9		5.4	4.1		0.0	0.0		0.5	1.2	
8-16	6	6.1	8.4		2.5	10.8		2.4	7.8		5.3	0.8		3.4	7.7	
16-32	5	7.9	10.8	5.9	1.9	13.3	8.7	-1.9	0.0	0.9	6.6	11.1	13.4	0.8	3.9	
32-63	4	6.6		6.1	5.9		9.9	3.4		3.0	17.3		9.7	8.0		4.1
63-127	3	12.8		10.2	12.8		10.7	6.8		2.5	8.0		4.2	9.8		4.4

322

 Table A2: Change in the SE of the trend using the Mixed Model.

Table A3 shows the adjusted  $R_{\star}^2$  for the proposed Mixed Models. Similarly to the change in SE (Table A2), the

adjusted  $R_{k}^{2}$  is a composite of the individual EqLat or TP results from the extended model with the exception of the

results for layers 3,\_4, and 5 at Mauna Loa where both predictors are included concurrently. At these layers the

adjusted  $R_{k}^{2}$  in some cases matches the higher Adj  $R_{k}^{2}$  values of the two predictors, and in others improves with the

328 combination of QBO and TP or EqLat

				I	LOTUS	Model I	Proxy T	ests: (A	djusted	R <sup>2</sup> of M	(odel)					
Height	Umkehr	Ar	osa/Dav	/05		OHP			Boulder	•		MLO			Lauder	
(hPa)	Layer	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND	UMK	сон	SND
1-2	9		0.42			0.37			0.36			0.11			0.32	
2-4	8	0.23	0.39		0.14	0.31		0.17	0.39		0.11	0.32		0.18	0.34	
4-8	7	0.35	0.35		0.31	0.41		0.27	0.33		0.26	0.32		0.17	0.27	
8-16	6	0.31	0.35		0.33	0.45		0.33	0.40		0.40	0.51		0.25	0.23	
16-32	5	0.34	0.38	0.26	0.25	0.51	0.23	0.31	0.40	0.19	0.40	0.35	0.37	0.42	0.41	0.24
32-63	4	0.21		0.22	0.29		0.25	0.19		0.13	0.34		0.35	0.42		0.25

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63-127	3	0.24		0.23	0.42	0. <u>17</u>	0.22	0.13	0.14	0.21	0.25	0.19	 	Deleted: 26
able A3:	Adjuste	d R <sup>2</sup> fo	r the M	fixed M	Iodel									Formatted: Superscript

### 1337 Augmented Mixed Model

1338 It is hard to ignore the substantial reduction of SE when adding the AO/AAO predictor especially for layers 3,4 and 5

1339 at Mauna Loa, and for layers 3 and 4 at Arosa/Davos. The results for OHP layers 3 and 4 are still compelling, though

1340 somewhat less so. So we explore the addition of AO/AAO at these three stations only, for the layers specified. Table

1341 A4 summarizes the predictor choices for this Augmented Mixed Model.

1342 1343

> LOTUS Augmented Mixed Model Arosa/Davos OHP Boulder MLO Lauder Layer 9 EqLat EqLat EqLat Ref EqLat 8 EqLat EqLat EqLat Ref EqLat 7 EqLat EqLat EqLat Ref EqLat 6 EqLat EqLat EqLat EqLat EqLat 5 EqLat EqLat EqLat EqLat, QBO, AO/AAO EqLat 4 TP, AO/AAO TP, AO/AAO TP TP, QBO CD, AO/AAO TP 3 TP, AO/AAO TP, AO/AAO TP TP, QBO CD, AO/AAO TP

1344 1345

 Table A4:
 Details of additional predictor choices for each level and station in the Augmented Mixed Model . This differs from Table A1 by adding AO/AAO at some levels for Arosa/Davos, OHP and Mauna Loa,

		LOT	US Mod	el Test:	Differe	nce [%]	in Star	ndard Ei	ror: Ful	l Model	vs Refe	erence	Model			
Pressure	Umkehr	Are	osa/Dav	/OS	Haut	te-Prov	ance		Boulder	r	М	auna Lo	ba		Lauder	
(hPa)	Layer	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	СОН	SND	UMK	COH	SND
1-2	9		8.4			2.9			1.9			0.0			2.9	
2-4	8	-0.5	0.7		0.1	1.2		-0.4	1.5		0.0	0.0		1.0	3.1	
4-8	7	3.8	3.2		2.1	0.6		5.4	4.1		0.0	0.0		0.5	1.2	
8-16	6	6.1	8.3		2.5	10.9		2.4	7.8		5.3	7.8		3.4	7.7	
16-32	5	7.9	10.6	5.9	1.9	13.4	8.7	-1.9		1.4	13.6	13.0	13.3	0.8	3.9	-1.1
32-63	4	8.7		10.0	6.1		9.4	3.4		2.7	17.3		10.3	8.0		7.4
63-127	3	20.3		18.5	13.5		12.8	6.8		2.2	8.0		5.6	9.8		6.8

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Table A5 (the same as Table 10): Change in the SE of the trend using the <u>Augmented</u> Mixed Model.

1349Table A5 displays the change in the SE from the Reference Model now for the Augmented Mixed Model. Adding

AO/AAO at Arosa/Davos (layers 3 and 4) and Mauna Loa (layers 3 to 5) greatly reduces the SE beyond that of the Mixed Model results in **Table A2**. For OHP (layers 3 and 4) the impact is less dramatic for Umkehr. For sonde

Mixed Model results in **Table A2**. For OHP (layers 3 and 4) the impact is less dramatic for Umkehr. For sonde measurements at layer 4 the AO/AAO addition has no impact beyond the Mixed Model; for layer 3 the addition of

AO/AAO results in less reduction of the SE.

<u>A</u> <u>UMK</u> <u>0.23</u> <b>0.35</b>	COH           0.42           0.39           0.35	vos SND	<u>UMK</u> 0.14	OHP           COH           0.37           0.31	SND	<u>UMK</u> 0.17	Boulder           COH           0.36           0.39	<u>SND</u>	<u>UMK</u>	MLO COH 0.11 0.32	SND	<u>UMK</u> 0.18	Lauder <u>COH</u> <u>0.32</u>	<u>SND</u>
<u>0.23</u>	<u>0.42</u> <u>0.39</u>	SND	0.14	0.37	SND		0.36	SND		<u>0.11</u>	SND		0.32	SND
	0.39					<u>0.17</u>			<u>0.11</u>			0.18	_	
				<u>0.31</u>		<u>0.17</u>	<u>0.39</u>		<u>0.11</u>	0.32		0.18		
0.35	0.35	1	0.04									0.10	<u>0.34</u>	1
0.00			<u>0.31</u>	<u>0.41</u>		<u>0.27</u>	<u>0.33</u>		<u>0.26</u>	<u>0.32</u>		<u>0.17</u>	<u>0.27</u>	
<u>0.31</u>	<u>0.35</u>		<u>0.33</u>	<u>0.45</u>	ĺ	<u>0.33</u>	<u>0.40</u>		<u>0.40</u>	<u>0.51</u>		<u>0.25</u>	<u>0.23</u>	
<u>0.34</u>	<u>0.38</u>	<u>0.26</u>	<u>0.25</u>	<u>0.51</u>	<u>0.23</u>	<u>0.31</u>	<u>0.40</u>	<u>0.18</u>	<u>0.44</u>	<u>0.53</u>	<u>0.39</u>	<u>0.42</u>	<u>0.41</u>	0.29
<u>0.23</u>		<u>0.25</u>	<u>0.29</u>		<u>0.27</u>	<u>0.19</u>		<u>0.18</u>	<u>0.42</u>		<u>0.38</u>	<u>0.42</u>		0.31
<u>0.31</u>		<u>0.31</u>	<u>0.44</u>		<u>0.21</u>	<u>0.22</u>		<u>0.11</u>	<u>0.19</u>		<u>0.24</u>	<u>0.25</u>		0.21
	0.31	0.31	<u>0.31</u> <u>0.31</u>	0.31         0.31         0.44	0.31         0.31         0.44	0.31         0.31         0.44         0.21	0.31         0.31         0.44         0.21         0.22	0.31         0.31         0.44         0.21         0.22	0.31         0.31         0.44         0.21         0.22         0.11	0.31         0.44         0.21         0.22         0.11         0.19	0.31         0.44         0.21         0.22         0.11         0.19	0.31         0.44         0.21         0.22         0.11         0.19         0.24	0.31         0.31         0.44         0.21         0.22         0.11         0.19         0.24         0.25	

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Table A6 displays the Adj  $R_{\star}^2$  for the Augmented Mixed Model. Adding AO/AAO improves the Adj  $R_{\star}^2$  results for Arosa/Davos and MLO and has little to no impact at OHP. Based on the criteria outlined at the beginning of this appendix, we assign the Augmented Mixed Model as the 'Full Model' in the body of this paper.

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 1362
 Code/Data availability:
 All dataset used in this study are publicly available at the website

 1363
 https://gml.noaa.gov/aftp/ozwv/Publications/2023\_Umkehr\_Ozone\_Trends\_Paper/.

1364 **Competing interests:** The authors declare that they have no conflict of interest.

1365 Author contributions: IP and JW conceptualized the paper, and IP led the paper preparation. PE, KA, and JW performed the data analysis. KM is responsible for the production of the spatial and temporally matched ground-based 1366 1367 and satellite ozone profile data. JW is responsible for producing COH zonally averaged and station overpass ozone 1368 profile records. LF is responsible for the retrieval and calibration of the OMPS data. GM, PE, KM and KA are responsible for NOAA Umkehr measurements. EMB is responsible for measurements in Arosa/Davos. RQ is 1369 1370 responsible for Umkehr and ozonesonde observations in Lauder, New Zealand. BJ and PC are responsible for 1371 ozonesonde observations in Boulder and Hilo. GA is responsible for the ozonesonde observations in OHP. RVM is 1372 responsible for HEGIFTOM ozonesonde records and data analyses. RD, SGB, DZ provided context of the LOTUS 1373 model use and interpretation of trend analyses. All authors contributed to the writing of the paper.

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