

1 **Solar radiation estimation in West Africa: impact of dust conditions during**
2 **2021 dry season**

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15 Abstract

16 The anticipated increase in solar energy production in West Africa requires high-quality solar
17 ~~irradiance~~radiation estimates, which is affected by meteorological conditions and in particular
18 the presence of desert dust aerosols. This study examines the impact of incorporating desert
19 dust into solar ~~irradiance~~radiation and surface temperature estimations. The research
20 focuses on a case study of a dust event in March 2021, which is characteristic of the dry
21 season in West Africa. Significant desert aerosol emissions at the Bodélé depression are
22 associated with a Harmattan flow that transports the plume westwards. Simulations of this
23 dust event were conducted using the WRF meteorological model alone, as well as coupled
24 with the CHIMERE chemistry-transport model, using three different datasets for the dust
25 aerosol initial and boundary conditions (CAM5, GOCART, MERRA2). Results show that
26 considering desert dust reduces estimation errors in global horizontal irradiance (GHI) by
27 about 75%. The dust plume caused an average 18% reduction in surface solar
28 ~~irradiance~~radiation during the event. Additionally, the simulations indicated a positive bias in
29 aerosol optical depth (AOD) and PM10 surface concentrations. The choice of dataset for
30 initial and boundary conditions minimally influenced GHI, surface temperature, and AOD
31 estimates, whereas PM10 concentrations and aerosol size distribution were significantly
32 affected. This study underscores the importance of incorporating dust aerosols into solar
33 forecasting for better accuracy.

34

35 Short summary

36 Solar energy production in West Africa is set to rise, needing accurate solar
37 ~~irradiance~~radiation estimates, which is affected by desert dust. This work analyses a March
38 2021 dust event using a modelling strategy incorporating desert dust. Results show that
39 considering desert dust cut errors in solar ~~irradiance~~radiation estimates by 75% and reduces
40 surface solar radiation by 18%. This highlights the importance of incorporating dust aerosols
41 into solar forecasting for better accuracy.

42

43 1. Introduction

44 The West African region is facing significant development challenges due to global change.
45 One of these challenges is related to access to electricity, particularly through the use of
46 renewable energy. West African countries have committed to reduce their greenhouse gas
47 emissions as part of the Paris Agreement (2015). Furthermore, assessments of solar
48 resources in West Africa demonstrate the region's substantial potential, as shown by
49 Diabaté et al. (2004), Plain et al. (2019) and Yushchenko et al. (2018). The International
50 Energy Agency (IEA) projects that the installed capacity for photovoltaic (PV) power
51 generation will increase by almost 20 times from 2020 to 2030 under its Sustainable Africa
52 Scenario (~~Africa Energy Outlook~~, IEA, 2022). PV energy is expected to experience
53 significant growth due to its competitiveness and low-carbon nature. However, solar
54 production is highly dependent on weather conditions (Dajuma et al., 2016).

55 The growth of solar energy in West Africa calls for the development of tailored tools to
56 facilitate its integration into power grids and ensure optimal operational maintenance.
57 Accurate production forecasts are required by solar power plant operators, spanning various
58 timescales, ranging from a few hours to several days. This is essential for maximising
59 production, reducing penalties linked to predicted deliverable energy, and optimising plant
60 maintenance to minimise production losses. High-quality forecasts are also crucial for
61 electricity grid operators to maintain supply-demand equilibrium and ensure system stability.
62 Therefore, the variability of energy production significantly affects them. The key

63 meteorological variables that influence photovoltaic production are the Global Horizontal
64 Irradiance (GHI) and the air temperature. These factors, which directly impact electricity
65 production and cell efficiency, often reach high levels in this region as demonstrated by
66 Dajuma et al. (2016) and Ziane et al. (2021). Their findings indicate that solar
67 irradiance radiation is the primary factor influencing PV production, as the generated current
68 by the photoelectric effect is proportional to the irradiance. Furthermore, they demonstrate
69 that, at the second order, the air temperature affects the efficiency of solar cells, as both
70 parameters are inversely correlated.

71
72 Clauzel et al. (2024) identified desert dust aerosol as a significant source of GHI forecast
73 errors for the only two solar power plants in the Sahel region of Sococim (Senegal) and
74 Zagtoui (Burkina Faso), particularly during the dry season. Dust aerosols are a key element
75 in the West African climate and strongly influence solar farm production through their direct
76 effect (aerosol-radiation interaction (ARI), Briant et al., 2017) and indirect effects (aerosol-
77 cloud interaction (ACI), Tuccella et al., 2019) on radiation, and also through their deposition
78 on solar panels (fouling effect, Diop et al., 2020, Aidara et al., 2023). As mentioned by Kok et
79 al. (2021), the West African desert aerosol load is the highest in the world and occurs mainly
80 during the dry season. In fact, North Africa, including the Sahara, is the world's largest
81 contributor to desert dust emissions (Prospero et al., 2002), and 60% of this dust is
82 transported to the West African region (D'Almeida, 1986; Kok et al., 2021). Most dust
83 emissions are associated with synoptic-scale atmospheric dynamics such as the Harmattan
84 flow during the dry season (Klose et al., 2010). Engelstaedter and Washington (2007)
85 pointed out the importance of small-scale wind events associated with the large-scale flow,
86 especially in the Bodele depression, which is a hotspot for dust emissions (Engelstaedter et
87 al., 2006). Analysing satellite observations, Schepanski et al. (2009) show that 65% of the
88 activation of the dust source area occurred in the early morning, demonstrating the important
89 role of the breakdown of the nocturnal low-level jet. Washington and Todd (2005) confirmed
90 the importance of the Bodele low-level jets during the dry season in initiating dust emissions
91 that can be transported to the West African coast within a few days. Dust aerosol emissions
92 are also highly linked to Mesoscale Convective Systems (MCS, Marsham et al., 2008 ;
93 Bergametti et al., 2017) and to strong near-surface winds in the intertropical discontinuity
94 zone during the rainy season (Bou Karam et al., 2009).

95
96 Some studies intend to model dust events in West Africa such as Ochiegbu (2021) who
97 implemented a back-trajectories model to understand the dust event reaching Nigeria. This
98 work revealed that most of the aerosols coming to Nigeria between 2011 and 2014 were
99 originating from the Bodele Depression. Menut (2023) focused on dust forecasting during the
100 Cloud-Atmospheric Dynamics-Dust Interactions in West Africa (CADDIWA) campaign during
101 summer 2021 (Flamant et al., 2024) using the CHIMERE regional chemistry-transport model
102 (Menut et al., 2021). The model was coupled online with the Weather Research and
103 Forecasting (WRF) meteorological model (Briant et al., 2017; Tuccella et al., 2019) to
104 perform dust aerosol concentration forecasts. The results of this work provide confidence in
105 the model coupling in the region as the dust forecast quality does not decrease with time
106 over a few days. In addition, only a limited number of studies have been conducted on the
107 prediction of GHI in the West African region. Sawadogo et al. (2024) conducted an
108 evaluation of WRF-solar GHI forecast (Jimenez et al., 2016) in Ghana for the year 2021. In
109 their work, a version of the model coupled offline with Copernicus Atmosphere Monitoring
110 Service (CAMS) Aerosol Optical Depth (AOD) forecasts was considered to integrate

111 information on aerosol load. They showed that WRF-Solar outperforms in predicting GHI
112 under clear sky conditions while its performance under high aerosol levels remains poor, that
113 was mainly attributed to uncertainties in the input AOD during data assimilation within the
114 model. Close to the region of interest, for the northern Morocco area, El Alani et al. (2020)
115 compared the performance of global models (Global Forecast System, Integrated Forecast
116 System, McClear) and demonstrated their proficiency in capturing GHI hourly temporal
117 variability.

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119 As far as our knowledge is concerned, no studies have been conducted to assess online
120 coupled simulations between a meteorological model and an aerosol life cycle model
121 representing the emissions, the transport and the deposition in West Africa to estimate solar
122 irradiance radiation. This is despite the significant presence of desert dust, characterised by
123 high concentrations in the region. Additionally, scarce attention has been given to the
124 significance of initial and boundary conditions for conducting the aerosol model on the
125 performance of analysis simulations, and to our knowledge, investigating these aspects
126 would represent a novel contribution to research in the West African region.

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128 Within this general context, the objectives of this study are two folds i) to evaluate the ability
129 to reproduce a dust event using a meteorological and dust life cycle model coupling
130 configuration, and ii) to investigate whether the performance of the simulations can be
131 enhanced by modifying the aerosol initial and boundary conditions employed, and to
132 estimate the uncertainty associated with this dataset selection with regard to the errors made
133 by the model. Section 2 introduces the case study, the simulation configuration, the data and
134 models selected for this work. In Section 3, the results are presented, beginning with the
135 variables of interest for solar production (GHI and surface air temperature), followed by the
136 variables associated with the desert aerosols (AOD, concentration, size distribution,
137 emissions). Section 4 gives main conclusions and draws some perspectives for this study.

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139 **2. Material and methods**

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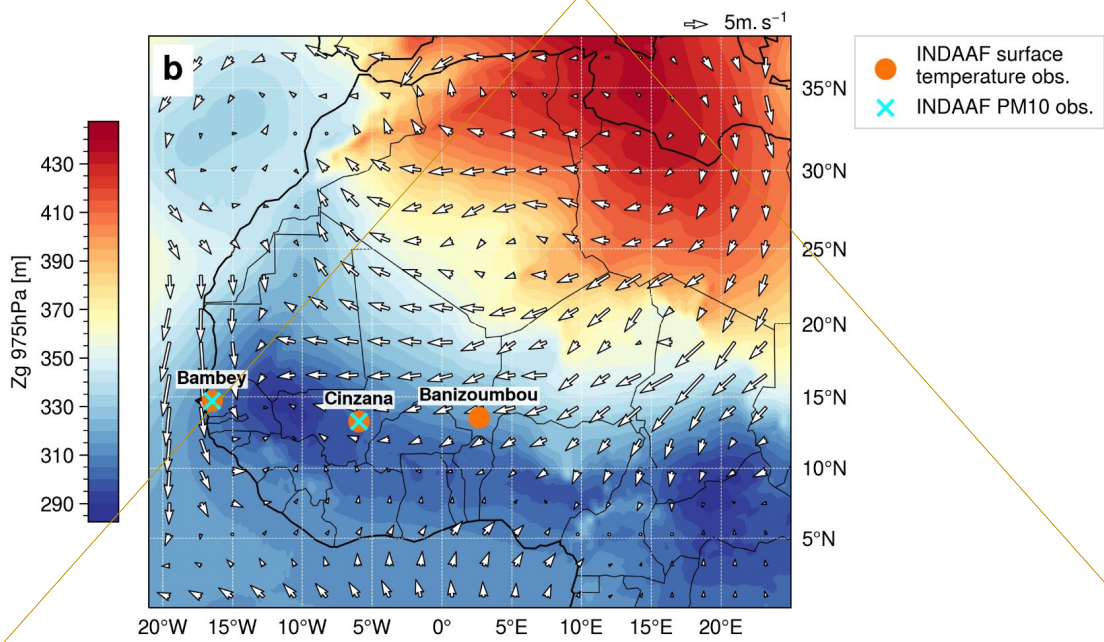
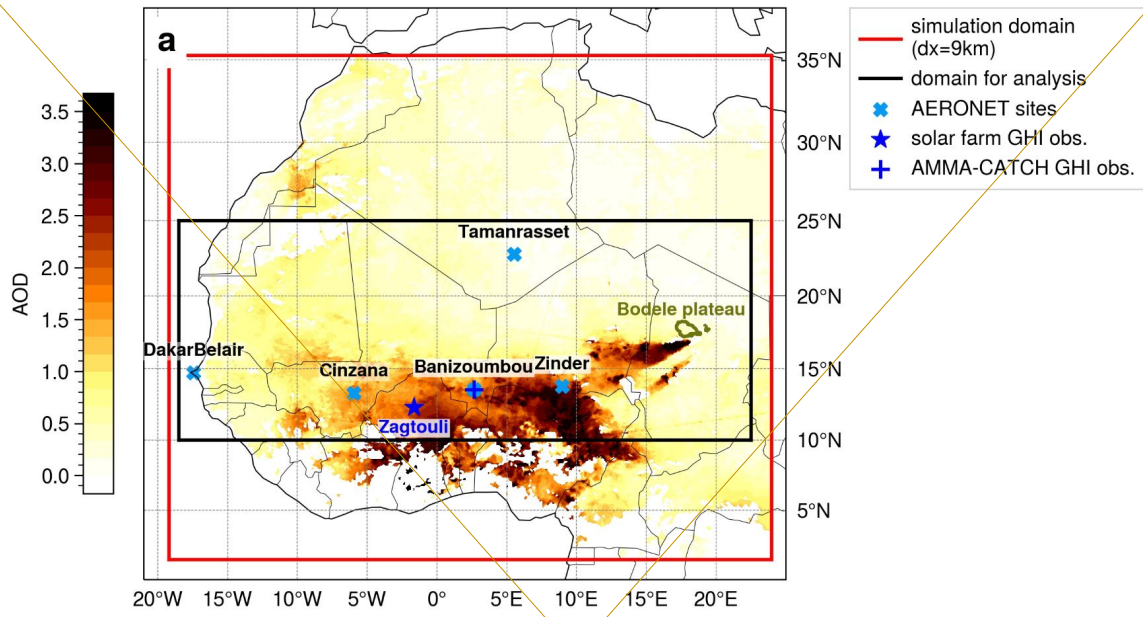
140 **2.1. Case study**

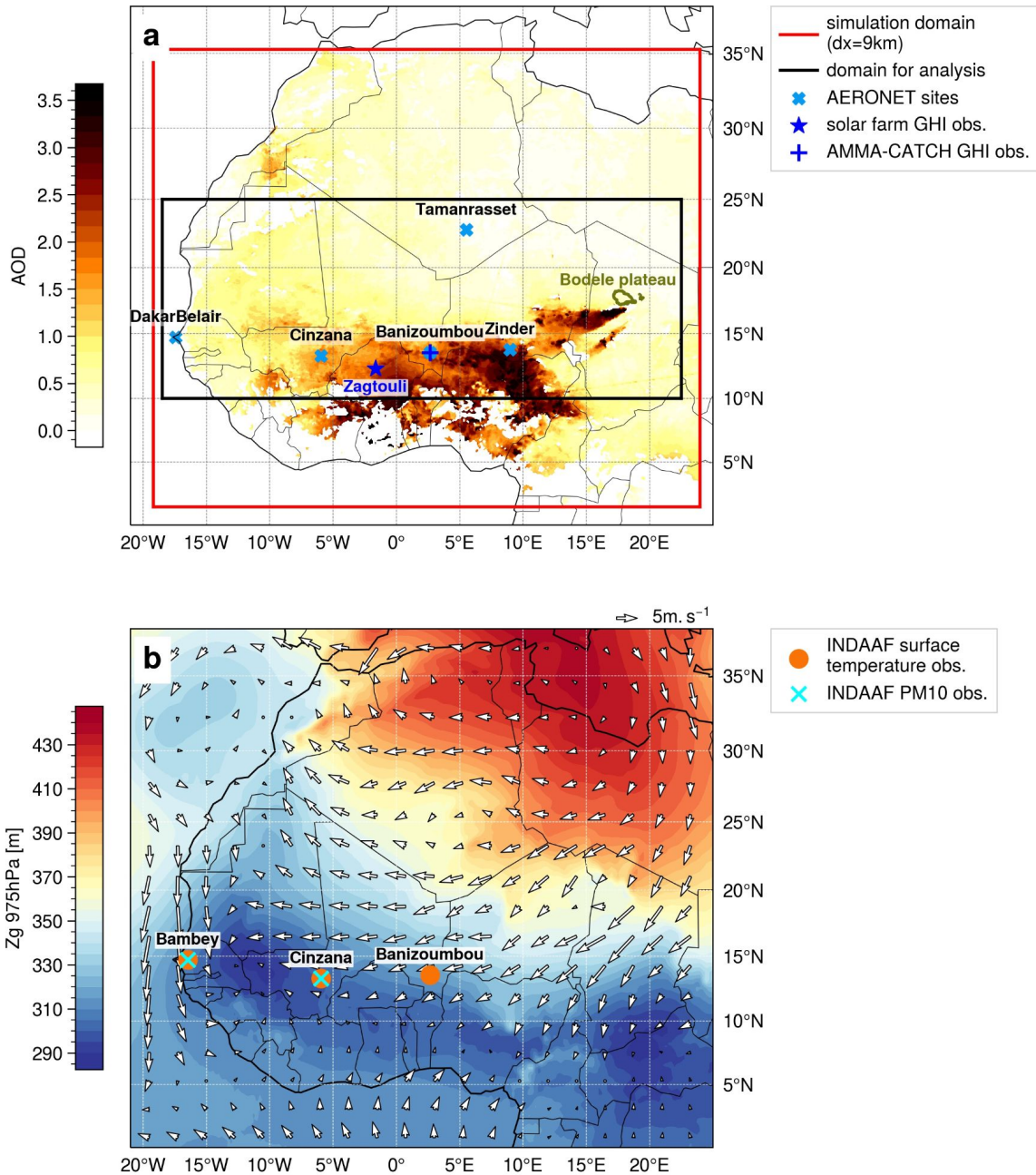
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142 The case study is a dust event that occurs in West Africa from March 26th-00 UTC to April
143 2nd-00 UTC, 2021, i.e., during the dry season. High dust emissions occur at the Bodélé
144 Depression (Chad), the plume being then transported westward. The dust plume reached its
145 maximum intensity in terms of AOD and dust concentration over West Africa, and in
146 particular over the Zagtouli solar farm (Burkina-Faso, Fig. S11a), on March 30th. The event
147 was also chosen because it was not predicted in the solar forecast currently implemented for
148 the Zagtouli solar farm, leading to solar forecast errors during the passage of the dust plume
(Clauzel et al., 2024).

149

150 Figure 1 illustrates that this event is characterised by a strong Harmattan flow, with surface
151 winds from the South/South-West sweeping across the Bodélé Depression (Chad), where
152 the potential for desert dust emissions is very high (Prospero et al., 2002; Washington et al.,
153 2006). Additionally, this event is characterised by a westward flow between Chad and the
154 Atlantic coast, which facilitates the transportation of the dust plume. Fig. 1a shows
155 MODerate-resolution Imaging Spectroradiometer (MODIS) satellite observations of the AOD,
156 identifying the initial dust source area on the Bodélé Depression, as well as the westward
157 movement of the plume. This event is characteristic of the West African dry season
158 climatology, with a dominant Harmattan flow as described in the introduction. Figure S1
159 provides further insight into the dust plume transport during the case study.





161
 162 **Figure 1** - a) Mean aerosol Optical Depth at 550nm from MODIS satellite observations over
 163 the period 28 March-00 UTC to 02 April-00 UTC 2021. The Global Horizontal Irradiance
 164 (GHI) observations and AERONET aerosol measurement network, introduced in 2.4, are
 165 presented, as well as the boundaries of the simulated domain (red rectangle) and the area of
 166 interest for analysis (black rectangle). b) Mean synoptic conditions of the geopotential height
 167 (Zg) at 975hPa and the 10m-wind (white arrows - in m/s) over the period 28 March-00 UTC
 168 to 02 April-00 UTC 2021 from ERA5 reanalysis. The surface temperature and aerosol
 169 concentration observations from the INDAAF network, introduced in 2.4, are presented.

170 171 2.2. Modelling tools

172 2.2.1. WRF model

173 The meteorological Weather and Research and Forecasting model (WRF) model version
 174 3.7.1 is taken for compatibility with the CHIMERE coupling procedure. It is used in its non-

175 hydrostatic configuration (Skamarock et al., 2008) and is forced at the boundaries of the
 176 domain every hour by the meteorological reanalysis data of ERA5 (ECMWF) provided on a
 177 regular 0.25° x 0.25° grid.

178 The model is run with a 9 km horizontal resolution, a 45s integration time step and 50
 179 vertical levels, from the surface to 50 hPa. The updated Rapid Radiative Transfer Model
 180 (RRTMG) radiation scheme (Iacono et al., 2008), which is mandatory for the aerosol optical
 181 properties feedback, is employed for both long- and short-wave radiations. Additionally, the
 182 Thompson aerosol-aware microphysics scheme (Thompson and Eidhammer, 2014) is
 183 applied. The Yonsei University planetary boundary layer's surface layer scheme (Hu et al.,
 184 2013) is also used, and the cumulus parameterisation is based on the Grell-Freitas scheme
 185 (Arakawa, 2004). The Revised MM5 surface layer scheme (Jiménez et al., 2012) is
 186 employed, while the Noah-MP Land Surface Model (Niu et al. 2011) is implemented for the
 187 land surface physics scheme.

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189 **2.2.2. CHIMERE model**

190 The chemistry-transport model CHIMERE version v2020r3 (Menut et al., 2021) is used in
 191 conjunction with the WRF model. Both models have a 9 km horizontal grid. The CHIMERE
 192 model has 30 pressure-dependent vertical levels from the surface up to 200 hPa, with a first
 193 layer thickness of 3 hPa. The model is configured for dust-only, with no chemistry and only
 194 considering dust aerosols (details in section 2.3). The threshold friction velocities for dust
 195 emission are estimated using the Shao and Lu scheme (2000) and the 6-km spatial
 196 resolution GARLAP (Global Aeolian Roughness Lengths from ASCAT and PARASOL)
 197 dataset from Prigent et al. (2012). Mineral dust emission fluxes were calculated employing
 198 the Alfaro and Gomes (2001) scheme on 10 aerosol size bins ranging from 0.01 to 40 µm.
 199 The Fécan et al. (1999) parametrization is employed to account for the inhibitory effect of
 200 soil moisture on dust emission. Dry deposition is treated as described in Zhang et al. (2001).
 201 Wet scavenging for aerosol is computed following the Willis and Tattelman scheme (1989).
 202 The CHIMERE model includes the Fast-JX module, version 7.0b (Wild et al., 2000; Bian et
 203 al., 2002) for the calculation of radiative processes. It considers the radiative properties for
 204 each aerosol species and each aerosol size bin independently to compute the aerosol
 205 optical depths, the single scattering albedo and the aerosol asymmetry factor. More details
 206 on the dust aerosol radiative properties are given in Tables S1 and S2. Finally, we test three
 207 different initial and boundary condition datasets for mineral dust load (see 2.2.3).

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209 **Table 1 - Parameterizations used in WRF and CHIMERE**

WRF	
<u>microphysics</u>	<u>Thompson aerosol-aware (Thompson and Eidhammer, 2014)</u>
<u>radiation</u>	<u>RRTMG scheme for LW and SW (Iacono et al., 2008)</u>
<u>land surface</u>	<u>Noah-MP land surface scheme (Niu et al., 2011)</u>
<u>planetary boundary layer</u>	<u>Yonsei University scheme (Hu et al., 2013)</u>

<u>surface layer</u>	<u>Revised MM5 surface layer scheme (Jimenez et al., 2012)</u>
<u>cumulus</u>	<u>Grell-Freitas scheme (Arakawa, 2004)</u>
CHIMERE	
<u>threshold friction velocities</u>	<u>Shao and Lu (2000) scheme</u>
<u>soil moisture</u>	<u>Fécan et al. (1999) scheme</u>
<u>dust emission fluxes</u>	<u>Alfaro and Gomes (2001) scheme</u>
<u>radiative processes</u>	<u>Fast-JX model, version 7.0b (Wild et al., 2000; Bian et al., 2002)</u>
<u>aerosol size distribution bins (diameters in μm)</u>	<u>0.010 - 0.022</u>
	<u>0.022 - 0.048</u>
	<u>0.048 - 0.107</u>
	<u>0.107 - 0.235</u>
	<u>0.235 - 0.516</u>
	<u>0.516 - 1.136</u>
	<u>1.136 - 2.500</u>
	<u>2.500 - 5.000</u>
<u>5.000 - 10.00</u>	
<u>10.00 - 40.00</u>	
WRF	
<u>microphysics</u>	<u>Thompson aerosol-aware (Thompson and Eidhammer, 2014)</u>
<u>radiation</u>	<u>RRTMG scheme for LW and SW (Iacono et al., 2008)</u>
<u>land surface</u>	<u>Noah-MP land surface scheme (Niu et al., 2011)</u>
<u>planetary boundary layer</u>	<u>Yonsei University scheme (Hu et al., 2013)</u>
<u>surface layer</u>	<u>Revised MM5 surface layer scheme (Jimenez et al., 2012)</u>
<u>cumulus</u>	<u>Grell-Freitas scheme (Arakawa, 2004)</u>
CHIMERE	
<u>threshold friction velocities</u>	<u>Shao and Lu (2000) scheme</u>
<u>soil moisture</u>	<u>Fécan et al. (1999) scheme</u>

dust emission fluxes	Alfaro and Gomes (2001) scheme
radiative processes	Fast-JX model, version 7.0b (Wild et al., 2000; Bian et al., 2002)
	0.010—0.022
	0.022—0.048
	0.048—0.107
	0.107—0.235
aerosol size distribution bins (diameters in μm)	0.235—0.516
	0.516—1.136
	1.136—2.500
	2.500—5.000
	5.000—10.00
	10.00—40.00

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2.2.3. Dust aerosol initial and boundary condition datasets

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In this study, the uncertainty in the solar estimate associated with the initial and boundary conditions of the dust aerosol load is evaluated. Three datasets were used: a climatology derived from the Global Ozone Chemistry Aerosol Radiation and Transport (GOCART, Ginoux et al., 2001), the Modern-Era Retrospective analysis for Research and Applications Version 2 (MERRA2) reanalysis (Gelaro et al., 2017) and the CAMS reanalysis (Inness et al., 2019).

The GOCART climatology is provided with the distribution of the CHIMERE model. It is a monthly climatology on a coarse horizontal grid ($2^\circ \times 2.5^\circ$), which is corrected by applying a factor of 0.3 as in Vautard et al. (2005).

The MERRA2 reanalysis combines the Goddard Earth Observing System (GEOS) and GOCART models, which are online coupled and implemented with a data assimilation system. It has a 3-hour temporal resolution and is presented on a $0.5^\circ \times 0.635^\circ$ horizontal grid. The observational data considered in the data assimilation process are AOD satellite observations from MODIS, Advanced Very High Resolution Spectroradiometer (AVHRR), Multi-angle Imaging SpectroRadiometer (MISR) and ground observations from the AERosol RObotic NETwork (AERONET).

The CAMS reanalysis was constructed using 4DVar data assimilation in ECMWF's Integrated Forecast System (IFS). It has a temporal resolution of 3 hours and is computed on a regular 0.75° horizontal grid. The AOD data from the Visible Infrared Imaging Radiometer Suite (VIIRS), the MODIS and the Infrared Atmospheric Sounding Interferometer (IASI) satellite observations are used as observational information in the data assimilation process. [The version 48R1 of CAMS is used in this study.](#)

These three dust aerosol initial and boundary datasets differ in type (climatological or reanalysis), in horizontal, vertical and temporal resolution, and in the resolution and range of their aerosol size distribution. While GOCART has the highest number of aerosol classes with 7 bins, CAMS covers a wider size spectrum despite a lower size resolution with only 3 classes. MERRA2 has an intermediate resolution with 5 classes, but covers a smaller particle size spectrum than CAMS. The CHIMERE model pre-processes these dust aerosol size distributions by applying a transfer coefficient δ to compute the dust aerosol concentration on the 10 aerosol size bin defined for the simulations :

$$c_j = \sum_i \delta_{i,j} \times c_i \quad (1)$$

242 where c_i is the dust aerosol concentration of the i^{th} size bin from the initial and boundary
 243 condition dataset considered, c_j is the dust aerosol concentration of the j^{th} size bin in the
 244 CHIMERE simulation, and $\delta_{i,j}$ is the transfer coefficient. This transfer coefficient is derived
 245 as :

- 246 - $\delta_{i,j}=0$ if the i^{th} size bin from the initial and boundary condition dataset is found to be
 247 wholly outside the j^{th} size bin in the CHIMERE simulation;
- 248 - $\delta_{i,j}=1$ if the i^{th} size bin from the initial and boundary condition dataset is wholly
 249 encompassed by the j^{th} size bin in the CHIMERE simulation;
- 250 - $\delta_{i,j} = \frac{\log(r_{j,max}) - \log(r_{j,min})}{\log(R_{i,max}) - \log(R_{i,min})}$ if the i^{th} size bin from the initial and boundary condition
 251 dataset wholly encompasses the j^{th} size bin in the CHIMERE simulation;
- 252 - $\delta_{i,j} = \frac{\log(R_{i,max}) - \log(r_{j,min})}{\log(R_{i,max}) - \log(R_{i,min})}$ if the i^{th} size bin from the initial and boundary condition
 253 dataset partially overlaps the j^{th} size bin in the CHIMERE simulation, but extends
 254 below the start of this size bin;
- 255 - $\delta_{i,j} = \frac{\log(r_{j,max}) - \log(R_{i,min})}{\log(R_{i,max}) - \log(R_{i,min})}$ if the i^{th} size bin from the initial and boundary condition
 256 dataset partially overlaps the j^{th} size bin in the CHIMERE simulation, but extends
 257 beyond the end of this size bin;

258 where $R_{i,min}$ and $R_{i,max}$ are respectively the radius of the lower and upper limit of the i^{th} size
 259 bin from the initial and boundary condition dataset, and $r_{j,min}$ and $r_{j,max}$ are respectively the
 260 radius of the lower and upper limit of the j^{th} size bin in the CHIMERE simulation.

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 262 For the sake of simplicity, throughout this article, we will refer to the WRF-CHIMERE
 263 simulations runned with the GOCART, the MERRA2, and the CAMS dust aerosol initial and
 264 boundary conditions as *wrf_chimere-G*, *wrf_chimere-M*, and *wrf_chimere-C* simulations
 265 respectively.

266 Table 2 summarises the characteristics of the three dust aerosol datasets and their
 267 associated size distributions.

268
 269 **Table 2.** Summary of the characteristics of the dust initial and boundary condition products.

	GOCART	MERRA2	CAMS
type	climatology	reanalysis	reanalysis
temporal resolution	monthly	3h	3h
vertical levels	20	72	60
horizontal resolution (lat x lon)	2°x2.5°	0.5°x0.635°	0.75°x0.75°

	0.20 - 0.36 μm	0.1 - 1.0 μm	0.03 - 0.55 μm
	0.36 - 0.60 μm	1.0 - 1.8 μm	0.55 - 0.90 μm
dust aerosol size distribution (radius in μm)	0.60 - 1.20 μm	1.8 - 3.0 μm	0.90 - 20.00 μm
	1.20 - 2.00 μm	3.0 - 6.0 μm	
	2.00 - 3.60 μm	6.0 - 10.0 μm	
	3.60 - 6.00 μm		
	6.00 - 12.00 μm		
	GOCART	MERRA2	CAMS
type	climatology	reanalysis	reanalysis
temporal-resolution	monthly	3h	3h
vertical-levels	20	72	60
horizontal resolution (lat x lon)	2°x2.5°	0.5°x0.635°	0.75°x0.75°
	0.20 - 0.36 μm	0.1 - 1.0 μm	0.03 - 0.55 μm
	0.36 - 0.60 μm	1.0 - 1.8 μm	0.55 - 0.90 μm
aerosol size distribution (radius in μm)	0.60 - 1.20 μm	1.8 - 3.0 μm	0.90 - 20.00 μm
	1.20 - 2.00 μm	3.0 - 6.0 μm	
	2.00 - 3.60 μm	6.0 - 10.0 μm	
	3.60 - 6.00 μm		
	6.00 - 12.00 μm		

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2.3. Modelling strategy

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The domain of simulation extends from 2° to 35°N and from 19°W to 24°E, as illustrated by the red box in Figure 1b. The domain is large enough to represent the primary atmospheric flows, including the Harmattan North/North-West flow and the monsoon South flow, as well as the transport of the emitted aerosol plumes. A horizontal resolution of 9 km has been selected in order to ensure that the grid ratio is approximately 3 with the ERA5 meteorological forcing. This choice is also motivated by the a priori intention to achieve a resolution higher than that of previous CHIMERE simulations performed in this region and compared to the operational solar forecast model used for the Zagtoui solar farm, which are based on global forecast models (see 2.4.1). The CHIMERE model is configured in a “dust only” model, which models only the mineral dust type. This hypothesis is supported for this dust case study by Fig. S2, as desert dust is the dominant aerosol during the event, particularly above 10°N. [This hypothesis is also reinforced by the dust optical depth \(DOD\) to AOD ratio derived from the CAMS reanalysis, which exceeds 80% during this case study and for the domain of interest \(not shown\).](#) It is notable that biomass burning, which

286 represents the other principal aerosol source in this region, is no longer a significant
287 contributor to aerosol levels at that time of the year (Evans et al., 2018).

288 The WRF and CHIMERE models are coupled online through the OASIS3 MCT coupler. A
289 two-way coupling strategy is selected, in which WRF sends meteorological variables to
290 CHIMERE which in turn exchanges aerosol information such as AOD, Single Scattering
291 Albedo (SSA) and Asymmetry Factor. This coupling strategy imposes most of the WRF
292 parameterisations. The exchange frequency is set to 15 minutes. The WRF model computes
293 fields on 50 levels, which are linearly interpolated over the 30 CHIMERE vertical levels via
294 the OASIS coupler. The coupling includes the feedbacks of aerosol-radiation interactions
295 (ARI, direct aerosol effect) and aerosol-cloud interactions (ACI, indirect aerosol effects)
296 simultaneously.

297 The simulation starts on March 14th-00 UTC and ends on April 2nd-00 UTC, 2021. The first
298 two weeks served as the spin-up period. The simulation outputs are analysed for the period
299 of March 28th-00 UTC to April 2nd-00 UTC, which corresponds to the passage of the
300 dust plume in the Sahel region, in particular around the Zagtouli solar farm in Burkina Faso.
301 Four simulations were conducted: a meteorological simulation using WRF model alone, and
302 dust simulations with the coupled WRF-CHIMERE models using as initial and boundary
303 conditions the GOCART climatology, the MERRA2 reanalysis and the CAMS reanalysis. The
304 simulation using only WRF allows for the evaluation of the impact of taking into account dust
305 aerosols in estimating solar irradiance radiation. This is compared to the other three
306 simulations, which are also used to evaluate the uncertainties associated with the choice of
307 the aerosol initial and boundary condition dataset. A domain of interest, spanning 10°N to
308 25°N (Fig. 1a), was selected for analysis and comparisons. This choice was guided by the
309 dust plume trajectory (Fig. S1) and the “dust only” hypothesis (Fig. S2).

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311 2.4. Evaluation datasets

312 This section presents the local and regional data that are employed in the evaluation of the
313 simulations.

314 2.4.1. GHI

315 The Global Horizontal Irradiance (GHI) is the total shortwave irradiance from the Sun on a
316 horizontal surface on Earth. It is the sum of direct irradiance, which takes into account the
317 solar zenith angle, and diffuse horizontal irradiance. It is measured in $W.m^{-2}$ for the
318 wavelength range 0.3 - 3.0 μm .

319

320 The national electricity company of Burkina-Faso, Sonabel, operates a solar farm in Zagtouli
321 (12.31°N;1.64°W; Fig. 1a), approximately 15 km west of the capital, Ouagadougou. It has an
322 installed capacity of 34 MWp and contributes up to 4% of Burkina Faso's annual electricity
323 production. Ground GHI measurements from pyranometers are available at a temporal
324 resolution of 15 minutes for the Zagtouli solar plant and undergo pre-processing to ensure
325 quality control. This involves removing outliers and days with missing data, visually checking
326 the consistency of the measured values and selecting data corresponding to production
327 hours (positive values for solar irradiance radiation at the top of the atmosphere). Operational
328 GHI forecasts for this solar farm are computed by the French company Steadysun. These
329 forecasts are based on a multi-model, multi-member and multi-mesh grid aggregation, which
330 is derived from the NCEP Global Ensemble Forecast System and the ECMWF Integrated
331 Forecast System (Clauzel et al., 2024).

332 In-situ measurements of GHI from pyranometers (Fig. 1a) are also available at a 15-minutes
333 temporal resolution for the Banizoumbou (Niger) surface station, installed as part of the

334 AMMA-CATCH observatory (Analyse Multidisciplinaire de la Mousson Africaine - Couplage
335 de l'Atmosphère Tropicale et du Cycle Hydrologique, AMMA-CATCH (2005)).
336 The two measurement sites were selected because they are the only locations where GHI
337 observations have been made available along the dust plume transport for the case study,
338 with the Zagtoui power station being one of the first large solar farms in West Africa and the
339 AMMA-CATCH observatory being the only one to offer continuous GHI measurements for
340 the region and period of interest.

341
342 The CAMS gridded solar radiation dataset (CAMS solar radiation services v4.6, Schroedter-
343 Homscheidt et al., 2022), based on the Heliosat-4 method (Qu et al., 2017), provides several
344 variables related to solar [irradiance radiation](#), such as clear-sky and all-sky GHI. It has a
345 horizontal resolution of $0.1^{\circ} \times 0.1^{\circ}$ and provides data every 15 minutes. The clear sky model
346 includes aerosols through the CAMS chemical transport model (Inness et al., 2019), which
347 integrates data assimilation of AOD and is coupled online to a numerical weather prediction
348 model. Cloud information for the all-sky model is derived from MeteoSat Second Generation
349 (MSG) satellite observations using the AVHRR Processing scheme Over cLOUDs, Land and
350 Ocean (APOLLO) Next Generation cloud processing scheme (Klüser et al., 2015). The
351 dataset was selected for comparison with the simulations as it integrates a description of
352 aerosol processes. While Yang and Bright (2020) [and Sawadogo et al. \(2023\)](#) show that it is
353 the best performing product for estimating surface solar [irradiance radiation](#) in the West
354 African region among several satellite-based gridded irradiance products, this dataset still
355 has a negative bias of about 10% [for all-sky solar irradiance estimates](#) at desert stations in
356 North Africa (CAMS solar radiation regular validation report, [Lefèvre, 20222020](#)).

357

358 **2.4.2. Surface temperature**

359 In-situ surface temperature measurements are available for three stations of the
360 International Network to study Deposition and Atmospheric composition in Africa (INDAAF) :
361 Banizoumbou (Niger, 13.54° N, 2.66° E, 6.2m above surface; Rajot et al, 2010a; Marticorena
362 et al, 2010; Kaly et al., 2015), Cinzana (Mali, 13.28° N, 5.93° W, 2m above surface; Rajot et
363 al, 2010b; Marticorena et al, 2010; Kaly et al., 2015) and Bambey (Senegal, 14.70° N,
364 16.47° W, 5.2m above surface; Marticorena et al, 2021a) (Fig. 1b). The measurement sites
365 were selected since they are almost aligned around $13-15^{\circ}$ North, which represents the main
366 pathway of Saharan and Sahelian dust towards the Atlantic Ocean during the case study.
367 The ERA5 atmospheric reanalysis (Hersbach et al., 2020) provides spatially continuous
368 hourly values of surface temperature at 2 metres and has a horizontal resolution of $0.25^{\circ} \times$
369 0.25° .

370

371 **2.4.3. Aerosol**

372 The INDAAF network also provides data on aerosol concentration through ground
373 measurements of PM_{10} , i.e. the concentration of atmospheric particles having an
374 aerodynamic diameter less than $10 \mu m$. For this case study, hourly PM_{10} measurements are
375 available for two stations (Fig. 1b): Cinzana (Rajot et al, 2010c; Marticorena et al, 2021; Kaly
376 et al, 2015) and Bambey (Marticorena et al, 2021b).

377 The CAMS atmospheric reanalysis (Inness et al., 2019) is also used to evaluate regional
378 surface PM_{10} concentration and AOD. It provides 3-hourly data with a horizontal resolution of
379 $0.75^{\circ} \times 0.75^{\circ}$, with a surface layer thickness of 2.4 hPa.-

380

381 Local ground measurements of AOD are retrieved from the AEROSOL ROBOTIC NETWORK level
 382 1.5 dataset (AERONET, Holben et al., 1998; Giles et al., 2019). AOD is calculated from sun
 383 photometer recordings, along with Ångström Exponent, and is only available during clear sky
 384 conditions in daylight hours, with a resolution of 1 minute. The AOD at 400 nm simulated
 385 with the WRF-CHIMERE model is converted to 440 nm for comparison with AERONET,
 386 using the Ångström formula :

$$\frac{AOD_{\lambda}}{AOD_{\lambda_0}} = \left(\frac{\lambda}{\lambda_0} \right)^{-\alpha} \quad (2)$$

~~$$\frac{AOD_{\lambda}}{AOD_{\lambda_0}} = \left(\frac{\lambda}{\lambda_0} \right)^{-\alpha} =$$~~ (2)

387 where AOD_{λ} is the AOD at the desired wavelength, $\lambda=440\text{ nm}$ here ; AOD_{λ_0} is the AOD at
 388 the wavelength simulated in the model, $\lambda_0=400\text{ nm}$ here ; α is the Ångström exponent,
 389 derived from the simulated AOD at different wavelengths and here given for the range from
 390 400 nm to 600 nm.

391
 392 AERONET also provides an aerosol size distribution dataset estimated through inversion of
 393 the photometers data, as described in Dubovik and King (2000). The algorithm for inversion
 394 provides a volume particle size distribution for 22 bins, which are logarithmically distributed
 395 for radii between 0.05 μm and 15 μm . For comparison with the modelled aerosol size
 396 distribution, this distribution is interpolated on the CHIMERE simulated aerosol size
 397 distribution which is composed of 10 bins ranging from 0.01 μm to 40.00 μm in diameter (see
 398 Table 1). Given that the coarsest bin (10.00-40.00 μm) is at the limit of the capabilities of the
 399 inversion method, and the two thinnest bins (0.010-0.022 μm and 0.022-0.048 μm) are out of
 400 the range of the inversion product, the AERONET dataset size sections are interpolated on
 401 the CHIMERE size sections ranging from 0.048 to 10.0 μm . Consequently, only comparisons
 402 between the three simulations can be made for the three size sections which are out of the
 403 range of AERONET product. The column aerosol volume size distribution simulated by the
 404 model is calculated for each bin "i" as in Menuet et al. (2016) :

$$\frac{dV(r_i)}{d \ln(r_i)} = \sum_{k=1}^{nlevels} \frac{m_{k,r_i} \times \Delta z_k}{\rho_{dust} \times \ln(r_{i,max}/r_{i,min})} \quad (3)$$

405 where r_i is the mean mass median radius (in μm) and $r_{i,min}$ and $r_{i,max}$ the boundaries of the
 406 i^{th} bin. m_{k,r_i} is the dust aerosol mass concentration (the mass of aerosol in one cubic metre
 407 of air, in $\mu\text{g} \cdot \text{m}^{-3}$). ρ_{dust} is the dust aerosol density (the mass of the particle in its own volume,
 408 $\rho_{dust}=2300\text{ kg} \cdot \text{m}^{-3}$). Δz_k is the model layer thickness (in metres), for a total of n levels (here
 409 30 vertical levels).

410
 411 ~~AERONET provides an aerosol size distribution dataset estimated through inversion of the~~
 412 ~~photometers data, as described in Dubovik and King (2000). The algorithm for inversion~~
 413 ~~provides a volume particle size distribution for 22 bins, which are logarithmically distributed~~
 414 ~~for radii between 0.05 μm and 15 μm .~~

415 The locations of the five AERONET sites used for comparison in this study are illustrated in
 416 Figure 1a.

417

418 The spatially continuous AOD is also derived from level 2 aerosol products of MODIS Terra
 419 and Aqua satellites (combined Dark Target, Deep Blue AOD at 0.55 micron, Collection 6.1,
 420 Platnick et al., 2015). It provides a measure of the AOD at 550 nm during daytime for clear
 421 sky conditions, with a spatial resolution of 10 km. To compare simulated AOD from WRF-
 422 CHIMERE models with AOD from MODIS, the former is converted from 600 nm to 550 nm.
 423 The conversion is performed using the Ångström formula (eq. 2).

424

425 Table 3 provides a general overview of the data used to evaluate the simulations in this
 426 study.

427

428 **Table 3** - Summary of data used to evaluate the simulations.

	product	type	resolution
GHI	Zagtouli solar farm monitoring system	pyranometer GHI measurement	local
	AMMA-CATCH observational network	pyranometer GHI measurement	local
	CAMS gridded solar radiation	atmospheric reanalysis	0.01°x0.01°
temperature	INDAAF network	ground measurements	local
	ERA5	atmospheric reanalysis	0.25°x0.25°
	INDAAF network	ground measurements	local
	CAMS (v48R1, EAC4)	atmospheric reanalysis	0.75°x0.75°
Aerosol Size Distribution	AERONET network	inversion product	local
Aerosol Optical Depth	AERONET network	sunphotometer ground measurements	local
	MODIS	satellite observations	10km
	product	type	resolution
GHI	Zagtouli solar farm monitoring system	pyranometer GHI measurement	local
	AMMA-CATCH observational network	pyranometer GHI measurement	local
	CAMS gridded solar radiation	atmospheric reanalysis	0.01°x0.01°
	INDAAF network	ground measurements	local

temperature	ERA5	atmospheric reanalysis	0.25°x0.25°
	INDAAF-network	ground-measurements	local
PM₁₀	CAMS (EAC4)	atmospheric reanalysis	0.75°x0.75°
Aerosol Size Distribution	AERONET-network	inversion-product	local
Aerosol Optical Depth	AERONET-network	sunphotometer-ground-measurements	local
	MODIS	satellite-observations	±10km

429

430

3. Results

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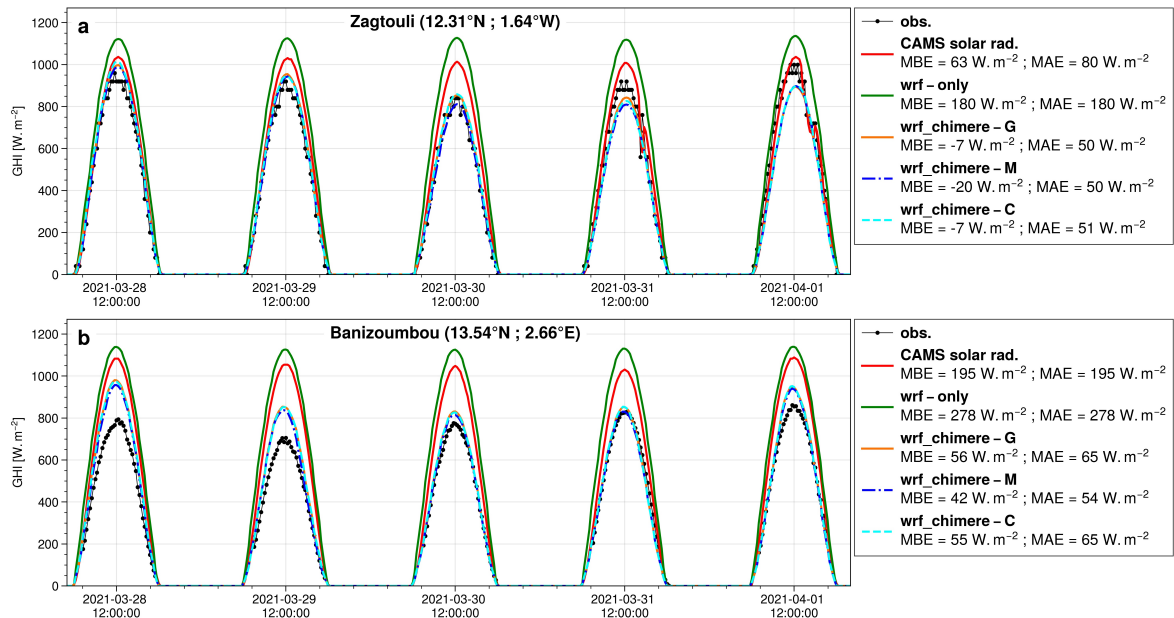
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438

439

The analysis starts by assessing the errors and uncertainties associated with the dust aerosol initial and boundary condition dataset employed to estimate the variables of interest for solar production, i.e. GHI and surface temperature. Subsequently, we investigate the potential causes of these uncertainties by evaluating the AOD, aerosol size distribution, and surface aerosol concentration (PM₁₀), as well as by examining mineral dust emissions and the flux of these aerosols at the boundaries of the domain. The metrics used to assess the quality of the simulations are described in Supplementary Materials.

3.1. GHI



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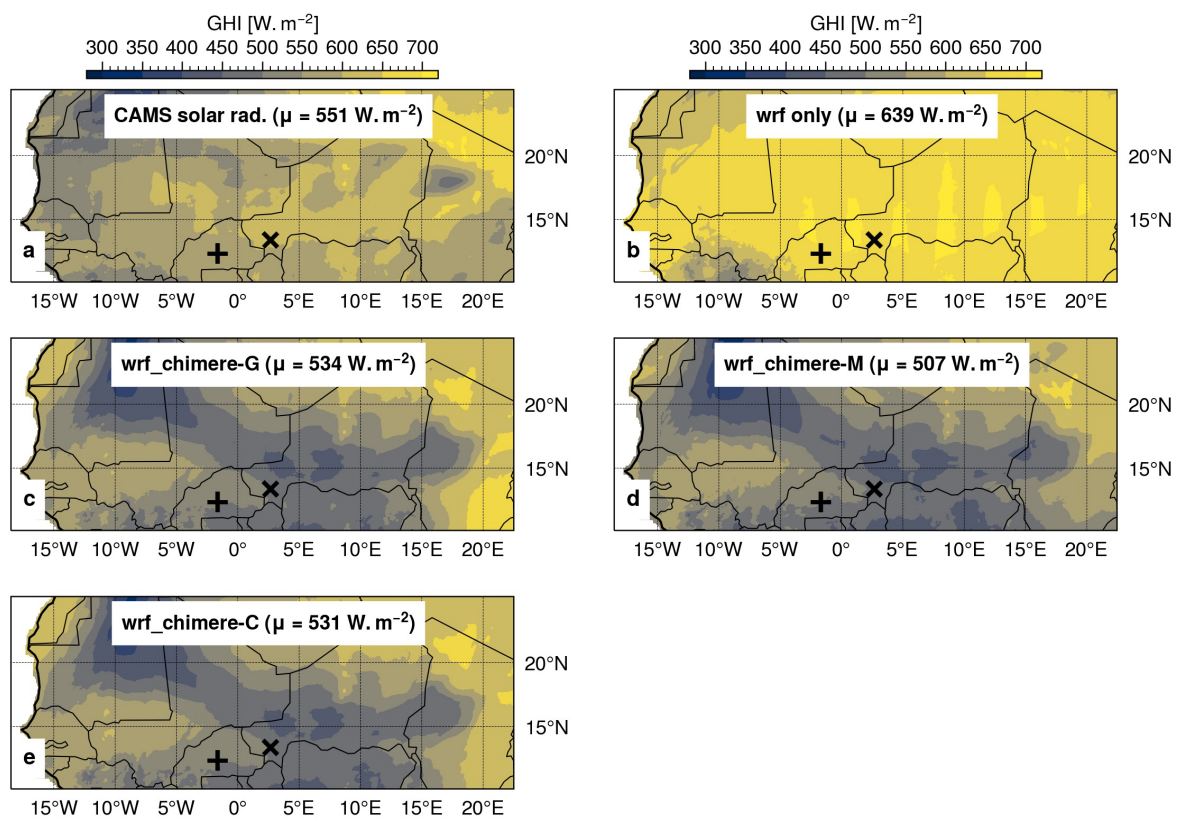
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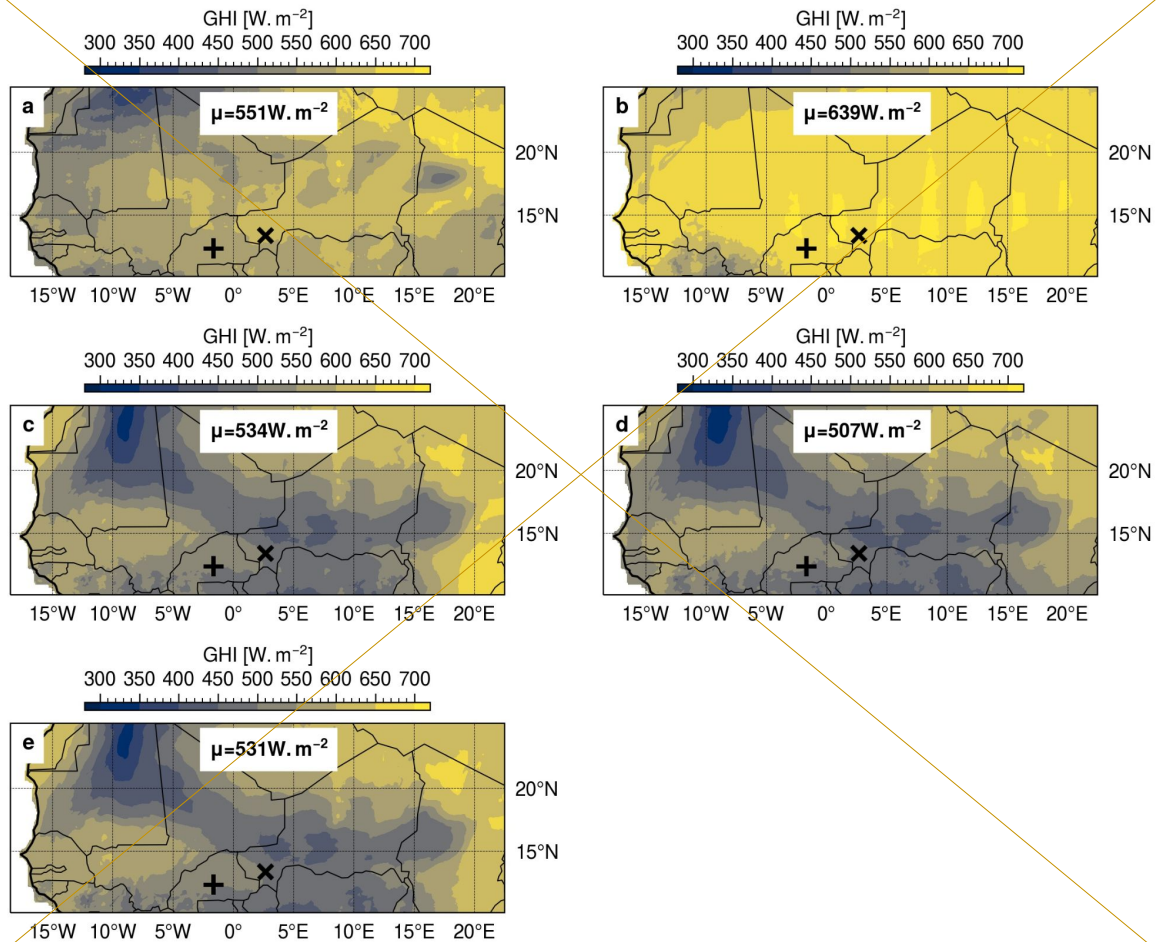
Figure 2 - Local comparison of CAMS gridded solar radiation product and simulated GHI against a) the Zagtouli solar farm observations and b) the Banizoumbou AMMA-CATCH observations. *wrf_chimere-G*, *wrf_chimere-M* and *wrf_chimere-C* refer to the WRF-CHIMERE simulations using GOCART, MERRA2 and CAMS as dust aerosol initial and boundary condition dataset respectively.

447 In Fig. 2, the local evaluation demonstrates the effect of taking into account dust aerosol for
 448 GHI estimation with the WRF-CHIMERE coupling over the WRF meteorological model
 449 alone. The coupling reduces the MAE by a factor of 3.6 at Zagtouli and by a factor of 4.6 at
 450 Banizoumbou on average. The simulations accurately represent the reduction in GHI
 451 intensity caused by the dust plume at both stations. However, the reduction persists
 452 compared to the observations at Zagtouli. At Banizoumbou, the simulations overestimate
 453 GHI at the beginning and end of the case study.

454 Figure 2 also indicates that the CAMS gridded solar radiation product fails to fully reproduce
 455 the dust event, with only a small reduction in GHI during the passage of the dust plume and
 456 an intermediate MAE between the WRF only and the WRF-CHIMERE simulations. This point
 457 serves to highlight the advantages of using a regional model in comparison to a global
 458 product for the simulation of dust conditions and the estimation of solar irradiance radiation.
 459 Furthermore, the uncertainty in GHI estimation related to the choice of the dust aerosol initial
 460 and boundary condition dataset is limited, particularly when compared to the errors. This is
 461 evidenced by the fact that the mean standard deviation between the three WRF simulations
 462 is only 7% of the average MAE of these simulations at Zagtouli, and only 5% at
 463 Banizoumbou.
 464



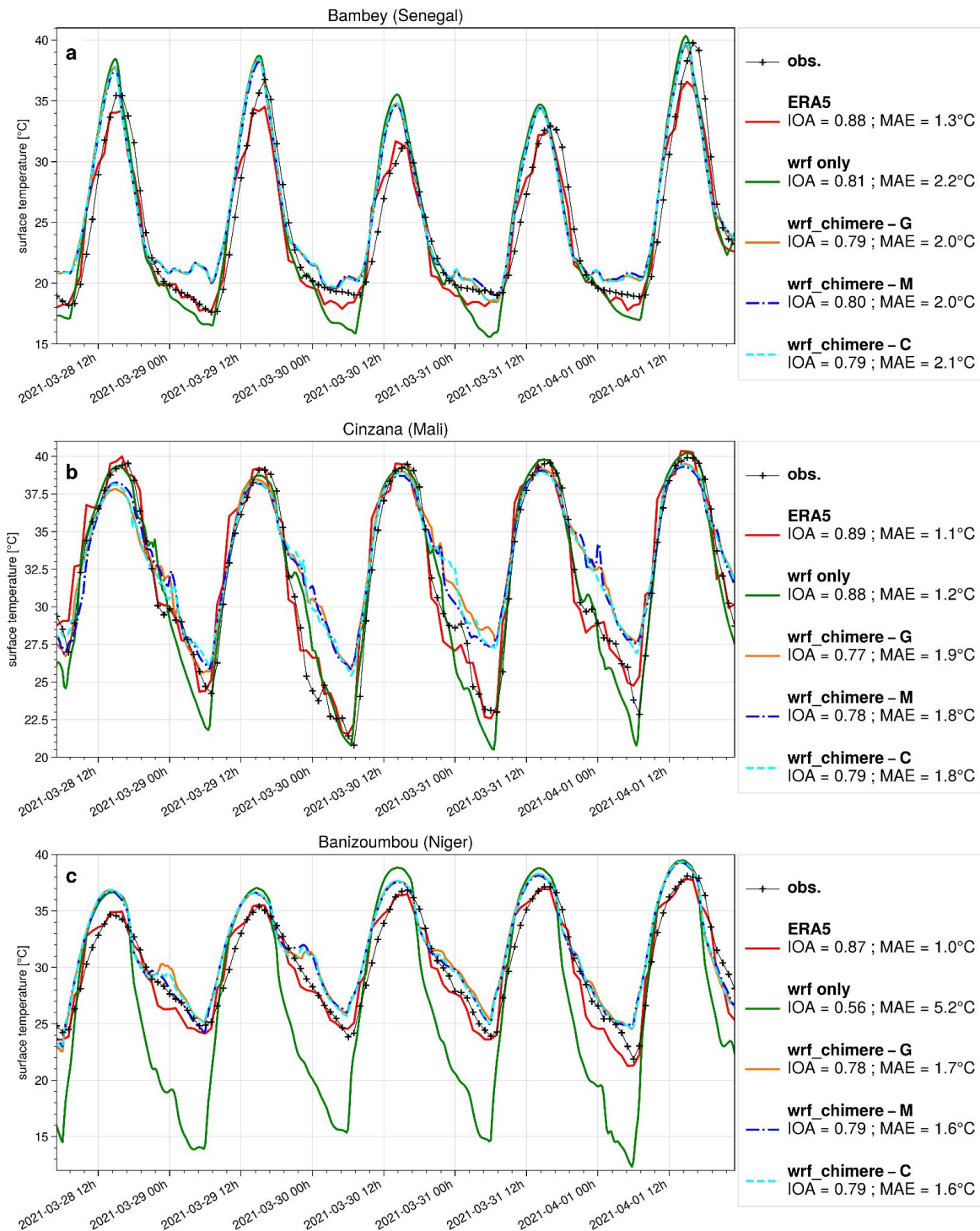
465



466
 467 **Figure 3** - Mean day-time GHI during the period of 28 March-00 UTC to 02 April-00 UTC
 468 2021 as estimated by a) the CAMS gridded solar radiation dataset, b) the WRF only
 469 simulation, and the WRF-CHIMERE simulations with c) GOCART, d) MERRA2 and e)
 470 CAMS as dust aerosol initial and boundary condition dataset; + is the Zagtouli solar farm
 471 and x is the Banizoumbou site. μ is the mean GHI estimates over the domain.

472
 473 The regional comparison presented in Fig. 3 provides more insight into the impact of
 474 incorporating dust on GHI estimation with the WRF-CHIMERE coupling, when compared to
 475 the WRF meteorological model alone. As anticipated the WRF-only simulation has the
 476 highest GHI estimates. The WRF-CHIMERE simulations indicate that dust aerosols reduce
 477 the mean GHI estimation by approximately $115 \text{ W} \cdot \text{m}^{-2}$ (-18%) as compared to the WRF-only
 478 simulation, while the CAMS gridded solar radiation global product shows a reduction of
 479 $88 \text{ W} \cdot \text{m}^{-2}$ (-14%). The three WRF-CHIMERE simulations exhibit identical regional patterns,
 480 with lower mean GHI values observed on the dust plume trajectory from the Bodélé
 481 Depression to the West, and also in the South Atlas region. In contrast, the CAMS gridded
 482 solar radiation dataset does not show this regional pattern, which may indicate that this
 483 global product does not fully capture the dust event.
 484 Furthermore, the uncertainty in GHI estimation associated with the choice of the dust aerosol
 485 initial and boundary conditions dataset is limited, particularly when compared to the changes
 486 brought by the taking of dust aerosol into account. Indeed, the standard deviation between
 487 the three WRF-CHIMERE simulations represents only 5% of the mean difference between
 488 these three simulations and the WRF-only simulation without dust.

3.2. Temperature



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Figure 4 - Local comparison of ERA5 and simulated surface temperature with the INDAAF observations for a) Bambey (Senegal), b) Cinzana (Mali) and c) Banizoumbou (Niger) measurement sites. *wrf_chimere-G*, *wrf_chimere-M* and *wrf_chimere-C* refer to the WRF-CHIMERE simulations using GOCART, MERRA2 and CAMS as dust aerosol initial and boundary condition dataset respectively. *IOA* is the Indicator of Agreement and *MAE* is the Mean Absolute Error.

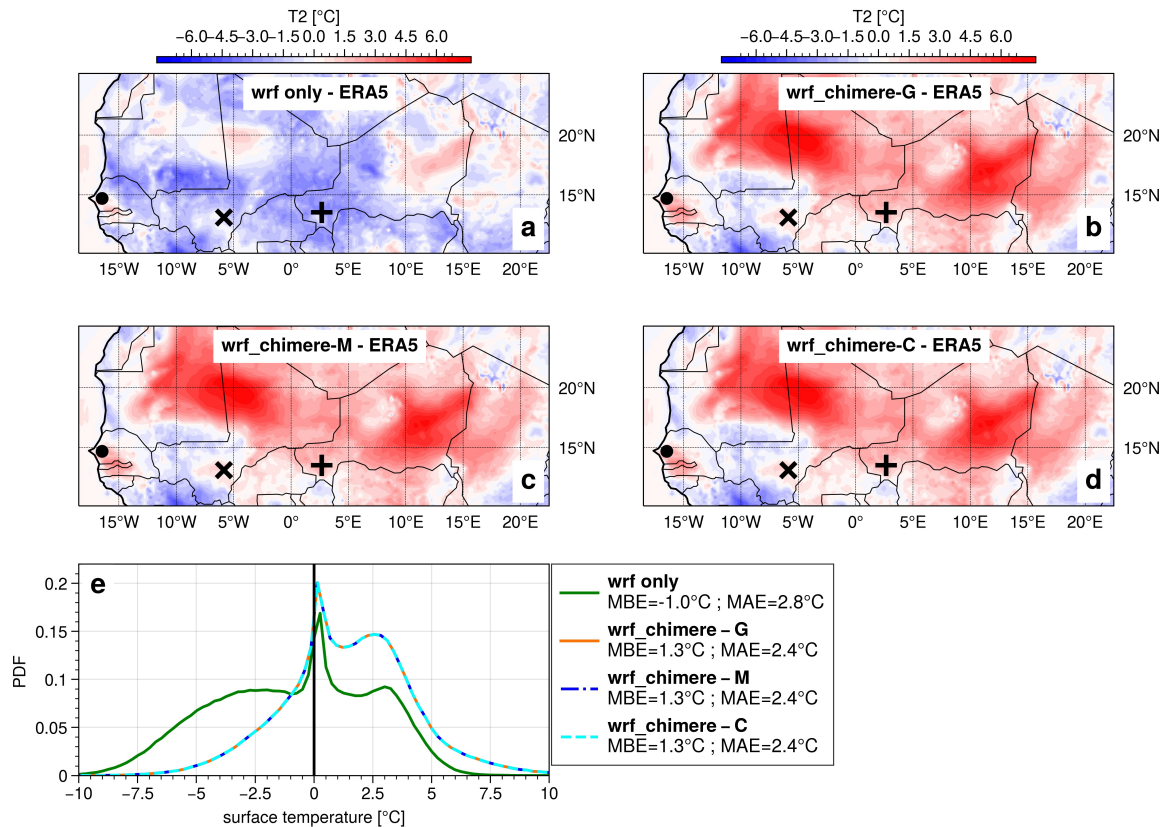
499 Figure 4 illustrates the contrasting outcomes of taking into account dust aerosols into the
500 WRF-CHIMERE coupling in comparison to the WRF meteorological model alone for the
501 estimation of surface temperature. At Bambey (Fig. 4a), which is far from the dust source
502 areas, the coupling has no effect on daytime temperatures but does affect night-time
503 temperatures. The WRF-CHIMERE and WRF-only simulations have IOA and MAE of the
504 same order of magnitude. At Cinzana (Fig. 4b), the WRF-only simulation performed better,
505 with a MAE 0.6°C lower than the coupled simulations, especially for night-time temperatures
506 but also for estimating the daily temperature peak. Finally, at Banizoumbou (Fig. 4c), which
507 is near the dust source areas, the coupling leads to a significant improvement in surface
508 temperature estimation, with an IOA of approximately 0.79 compared to 0.56 for the WRF-
509 only simulation and a MAE reduced by around 3.6°C. The impact of dust aerosols on
510 temperature is particularly pronounced at night-time. However, dust also affects the daily
511 temperature peak, with a reduction of 1.1°C of the daily maximum temperature observed on
512 the 30th of March.

513 Depending on the position of the measurement station, the results show a contrast, with a
514 significant improvement with the model coupling close to the source zones at Banizoumbou.
515 However, this improvement is reversed with increasing distance at Cinzana. This suggests
516 errors in the simulation of the transport of the dust plume from the source zones (Bodélé
517 Depression) towards the West. Overall, the main differences between WRF only and WRF-
518 CHIMERE coupled simulations occur at night time when there is no solar production. These
519 differences highlight the warming effect due to the dust aerosol interaction with the longwave
520 earth radiation.

521 In general, the uncertainty associated with the choice of the dust aerosol initial and boundary
522 condition dataset for the WRF-CHIMERE simulations is negligible compared to the errors in
523 temperature estimation or the difference with the WRF-only simulation.

524 The value of the ERA5 reanalysis for surface temperature evaluation is also reinforced in
525 Fig. 4, since it shows the lowest MAE and highest IOA. This dataset can therefore be
526 considered reliable for a regional evaluation of surface temperature.

527



528
 529 **Figure 5** - Mean difference in surface temperature as compared to the ERA5 reanalysis for
 530 a) the WRF only simulation, the WRF-CHIMERE simulations with b) GOCART, c) MERRA2
 531 and d) CAMS as dust aerosol initial and boundary condition dataset, during the period of 28
 532 March-00 UTC to 02 April-00 UTC 2021; the black point is the Bambey, x is the Cinzana and
 533 + is the Banizoumbou INDAAF sites. e) Probability Density Function for the differences in
 534 surface temperature between simulations and the ERA5 reanalysis.

535
 536 The regional surface temperature evaluation in Fig. 5 also reveals a contrast benefit of the
 537 coupling approach for the surface temperature estimation. While the WRF alone simulation
 538 (Fig. 5a) underestimates the surface temperature all over the domain, WRF-CHIMERE
 539 simulations are overestimating surface temperature in the dusty areas (Saharan region, Fig.
 540 5bcd). Overall, taking into account dust aerosol in the estimation of surface temperature
 541 reduces the MAE by 14% (Fig. 5e) when comparing the surface temperature estimates from
 542 simulations with the ERA5 reanalysis.

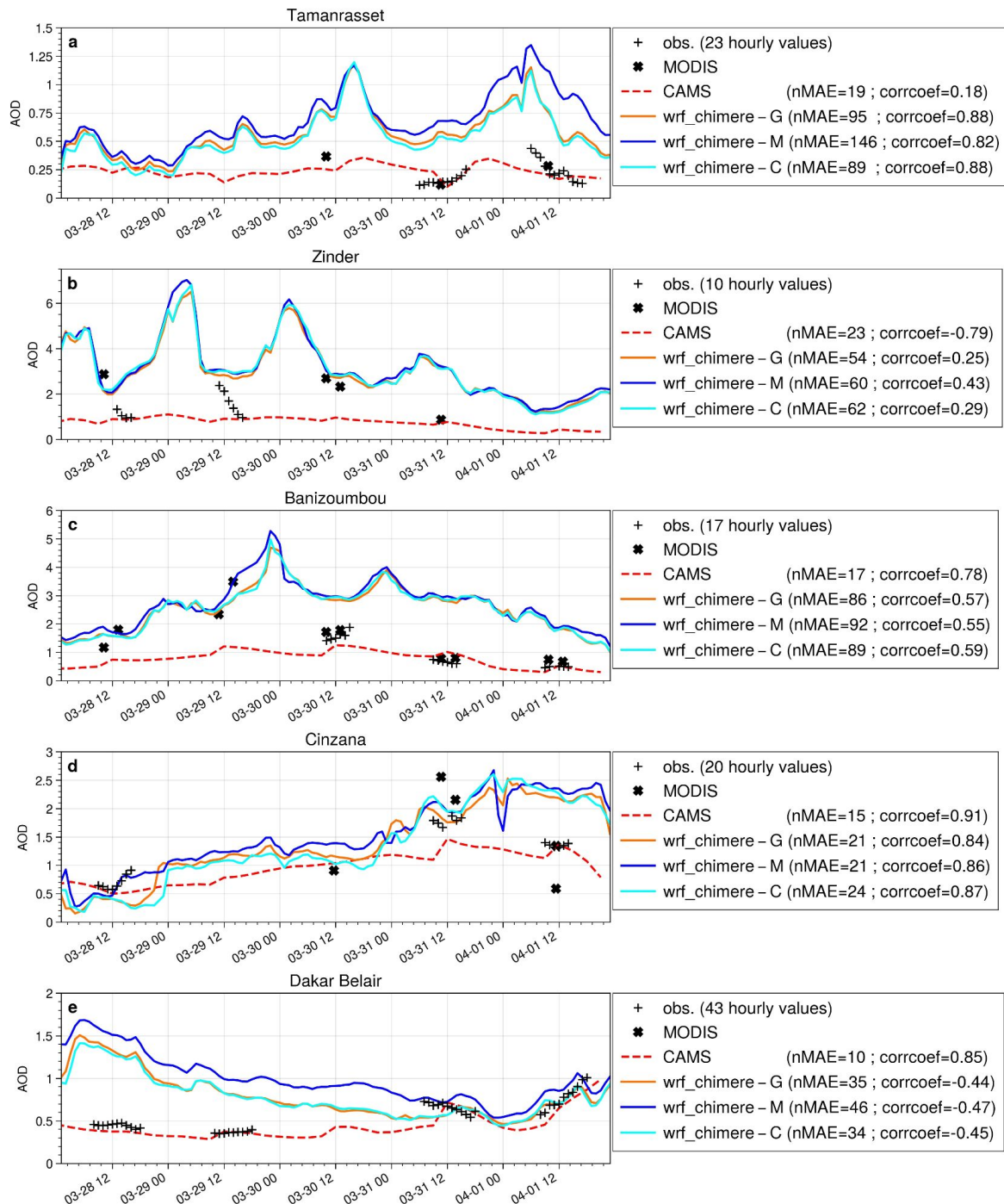
543 Furthermore, the uncertainty associated with the choice of the dust aerosol initial and
 544 boundary conditions dataset is limited. This is demonstrated by the fact that the standard
 545 deviation between the three WRF-CHIMERE simulations averaged over the period of
 546 analysis is 12% of the mean bias of those three simulations in comparison to ERA5
 547 reanalysis, and only 7% of the difference between the coupled simulations and the WRF-
 548 only simulation without dust.

549
 550 Finally, the incorporation of dust aerosol into the estimation of GHI appears to be a crucial
 551 element in this case study. However, the value of this approach is more debatable in the
 552 context of surface temperature estimation. Furthermore, the uncertainty related to the dust
 553 aerosol initial and boundary condition dataset selection is limited, particularly when

554 compared to the simulation errors, and to the differences between including dust in the
 555 simulation and not including it. The following sections will examine the simulated dust
 556 aerosol condition during the case study in order to explain the discrepancies observed in
 557 GHI and surface temperature, which are key parameters for solar production.
 558

559

3.3. Aerosol Optical Depth



560

561 **Figure 6** - Local comparison of simulated AOD with AERONET in-situ measurements at 440
 562 nm for a) Tamanrasset, b) Zinder, c) Banizoumbou, d) Cinzana and e) Dakar Belair stations.
 563 *wrf_chimere-G*, *wrf_chimere-M* and *wrf_chimere-C* refer to the WRF-CHIMERE simulations
 564 using GOCART, MERRA2 and CAMS as dust aerosol initial and boundary condition dataset

565 respectively; *MODIS* and *CAMS* refer to the AOD at 440 nm from the MODIS satellite
566 observations and the CAMS atmospheric reanalysis respectively. *nMAE* is the normalised
567 mean absolute error in % and *corrcoef* is the Pearson correlation coefficient, both derived
568 with AERONET measurements as the reference.

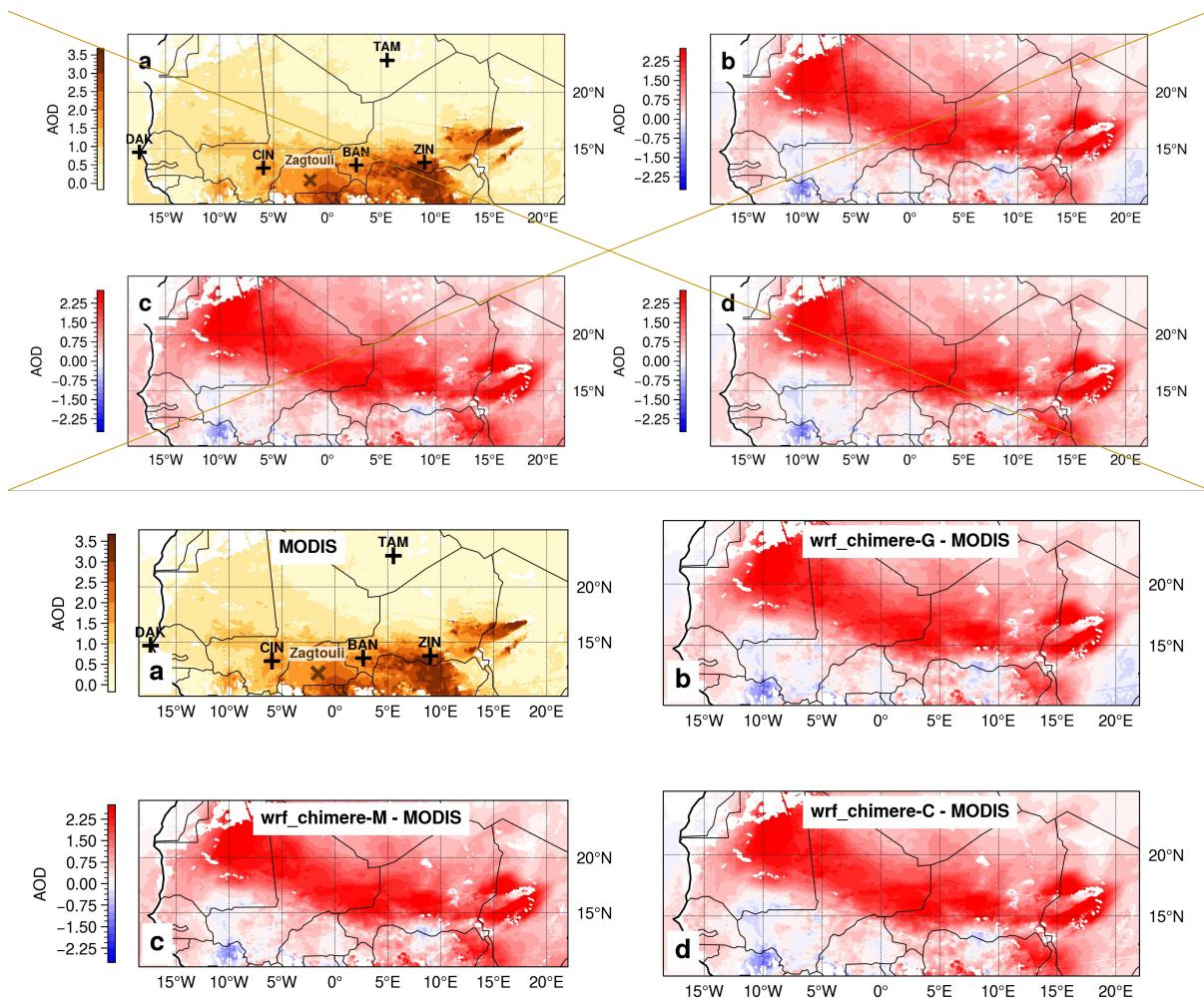
569 ~~The local evaluations presented in Figure 6 reveal an overestimation of the AOD for stations~~
570 ~~close to dust sources such as Tamanrasset (Fig. 6a), Zinder (Fig. 6b) and Banizoumbou~~
571 ~~(Fig. 6c). This overestimation is more limited with increasing distance from the dust source~~
572 ~~at Cinzana (Fig. 6d) and Dakar (Fig. 6e). The order of magnitude of the dispersion between~~
573 ~~the three simulations is small when compared to the errors of the simulation in representing~~
574 ~~the observed AOD. As a consequence, the uncertainty associated with the choice of the dust~~
575 ~~aerosol initial and boundary condition dataset is limited. Overall, the AERONET AOD~~
576 ~~measurements appear to be very scarce, particularly close to the dust aerosol sources~~
577 ~~(Zinder, Tamanrasset, Banizoumbou, Cinzana). The AOD measurements are performed by~~
578 ~~sun photometers which give recording by pointing at the sun. Thus these recordings are only~~
579 ~~available during daytime and with clear sky conditions. In some cases of intense dust~~
580 ~~plumes with very high concentration, leading to strong solar radiation absorption, the sun~~
581 ~~photometers are technically limited and cannot produce any record or, sometimes, the~~
582 ~~AERONET quality control system removes them. This may be the reason for the scarcity of~~
583 ~~observations in this case study, which focuses on an intense dust event, increasing the~~
584 ~~perceived overestimation of the simulations. To compensate for this, the AOD estimates~~
585 ~~from MODIS satellite observations have been added to Figure 6 to complete the data.~~

586 ~~Furthermore, the CAMS reanalysis appears to be a reliable dataset for dust AOD estimation,~~
587 ~~as it has no overestimation and has the lowest *nMAE* for all sites. Although it does not~~
588 ~~reproduce the AOD dynamics close to the dust source at Tamanrasset and Zinder, it has the~~
589 ~~highest correlation coefficient for the other sites.~~

590
591 The local evaluations presented in Figure 6 reveal an overestimation of the AOD for stations
592 close to dust sources such as Tamanrasset (Fig. 6a), Zinder (Fig. 6b) and Banizoumbou
593 (Fig. 6c). This overestimation is more limited with increasing distance from the dust source
594 at Cinzana (Fig. 6d) and Dakar (Fig. 6e). The order of magnitude of the dispersion between
595 the three simulations is small when compared to the errors of the simulation in representing
596 the observed AOD. As a consequence, the uncertainty associated with the choice of the dust
597 aerosol initial and boundary condition dataset is limited. Overall, the AERONET AOD
598 measurements appear to be very scarce, particularly close to the dust aerosol sources
599 (Zinder, Tamanrasset, Banizoumbou, Cinzana). The AOD measurements are performed by
600 sun photometers which give recording by pointing at the sun. Thus these recordings are only
601 available during daytime and with clear sky conditions. In some cases of intense dust
602 plumes with very high concentration, leading to strong solar radiation absorption, the sun
603 photometers are technically limited and cannot produce any record or, sometimes, the
604 AERONET quality control system removes them (Mueller et al., 2015 ; Giles et al., 2019).
605 This may be the reason for the scarcity of observations in this case study, which focuses on
606 an intense dust event, increasing the perceived overestimation of the simulations. To
607 compensate for this, the AOD estimates from MODIS satellite observations have been
608 added to Figure 6 to complete the data.

609 Furthermore, the CAMS reanalysis appears to be a reliable dataset for dust AOD estimation,
610 as it has no overestimation and has the lowest *nMAE* for all sites. Although it does not
611 reproduce the AOD dynamics close to the dust source at Tamanrasset and Zinder, it has the
612 highest correlation coefficient for the other sites. Nevertheless, this result should be

613 interpreted with caution, given the limited data available for calculating the dataset
 614 evaluation metrics. More research is needed to substantiate this conclusion.
 615



616

617 **Figure 7** - a) Mean from March 28th-00 UTC to April 2nd-00 UTC 2021 of MODIS AOD at
 618 550 nm satellite observations; x is the Zagtoui solar farm and + corresponds to AERONET
 619 stations. For panels b, c and d, AOD at 550 nm mean differences from March 28th-00 UTC
 620 to April 2nd-00 UTC 2021 between each of the WRF-CHIMERE simulations driven by
 621 GOCART, MERRA2 and CAMS, respectively, and the MODIS satellite observations.
 622

623

624 The AOD differences shown in Fig. 7bcd show that the simulations significantly overestimate
 625 the AOD as compared to the MODIS satellite observations, particularly in the Saharan and
 626 North Sahelian zones and in the South Atlas, with an average overestimation of +1.25
 627 between 15°N and 20°N. It is important to note that this overestimation is localised close to
 628 the desert aerosol source zones. The simulated AOD error in the Sahel zone, particularly
 629 around the Zagtoui solar power plant, is more limited with an average of +0.51 between
 630 10°N and 15°N. The mean standard deviation between the three WRF-CHIMERE
 631 simulations is only 10% of the mean error and 5% of the mean simulated AOD.
 632 Consequently the uncertainty in the AOD estimate associated with the selection of the dust
 633 aerosol initial and boundary condition dataset is small.

634 The observed overestimation of AOD by the WRF-CHIMERE simulations could be due to an
 635 overestimation of the aerosol concentration, or to an inaccurate estimation of the size

636 distribution of the dust plume, or to excessive aerosol emissions within the domain, or to an
 637 excessive inflow of desert aerosols at the domain boundaries. These hypotheses are
 638 investigated below. Another potential explanation may also be the uncertainties in the
 639 radiative properties of the dust aerosol incorporated in the CHIMERE model, or an
 640 underestimation of the aerosol deposition flux; these aspects are not investigated here.

641

642 3.4. Aerosol size distribution

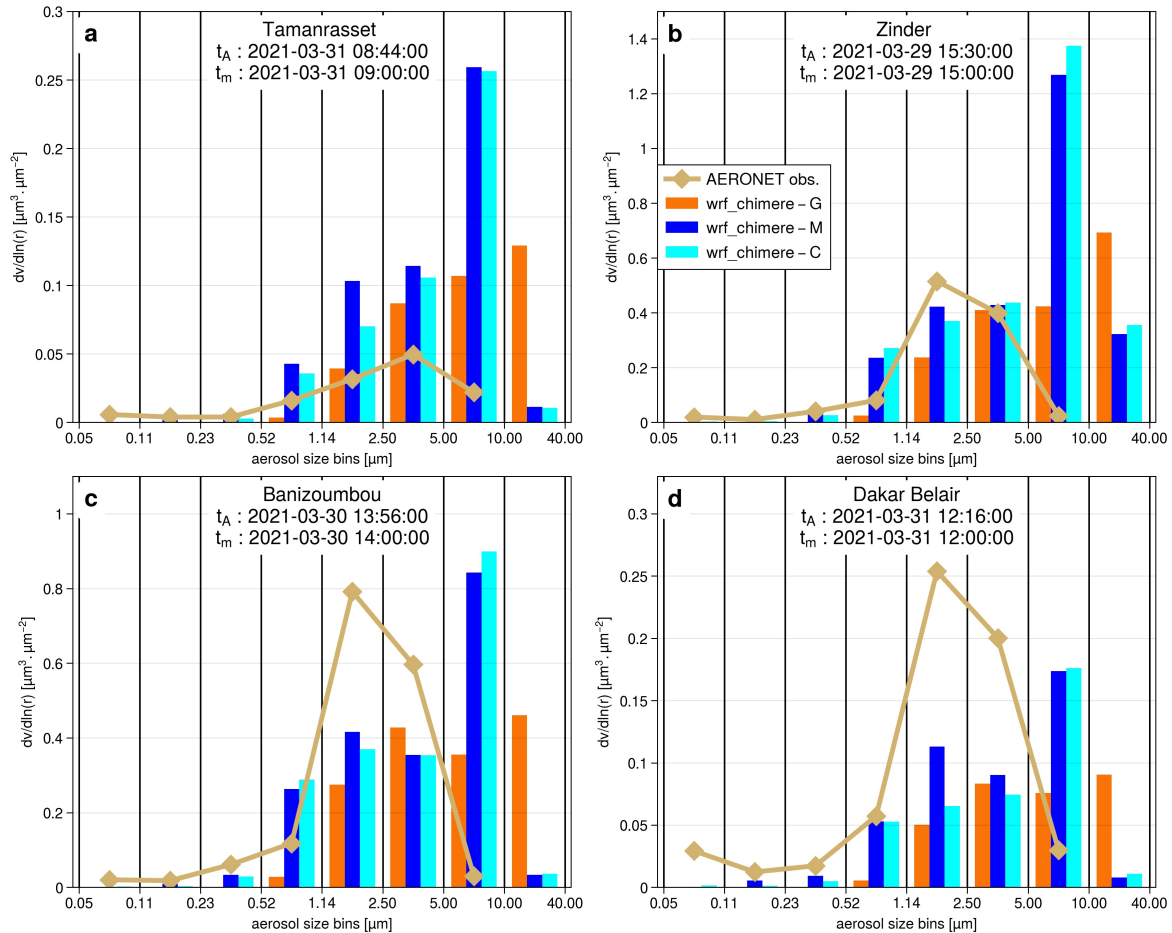
643 ~~As presented in section 2, the AERONET inversion products provide aerosol size distribution~~
 644 ~~for 22 bins logarithmically distributed ranging from 0.05 to 15 μm . For comparison with the~~
 645 ~~modelled aerosol size distribution, this distribution is interpolated on the CHIMERE~~
 646 ~~simulations aerosol size distribution which is composed of 10 bins ranging from 0.01 μm to~~
 647 ~~40.00 μm in diameter (see Table 1). Given that the last bin (10.00-40.00 μm) is at the limit of~~
 648 ~~the capabilities of the inversion method, with a maximum wavelength at which the AOD is~~
 649 ~~measured of 875 nm, it is not shown for the AERONET dataset. Consequently, only~~
 650 ~~comparisons between the three simulations can be made for the bigger size section. The~~
 651 ~~column aerosol volume size distribution simulated by the model is calculated for each bin “i”~~
 652 ~~as in Menut et al. (2016) :-~~

$$\frac{dV(r_i)}{d \ln(r_i)} = \sum_{k=1}^{nlevels} \frac{m_{k,r_i} \times \Delta z_k}{\rho_{dust} \times \ln(r_{i,max}/r_{i,min})} \quad (3)$$

653 where r_i is the mean mass median radius (in μm) and $r_{i,min}$ and $r_{i,max}$ the boundaries of the
 654 i^{th} bin. m_{k,r_i} is the dust aerosol mass concentration (the mass of aerosol in one cubic metre
 655 of air, in $\mu\text{g} \cdot \text{m}^{-3}$). ρ_{dust} is the dust aerosol density (the mass of the particle in its own volume,
 656 $\rho_{dust} \equiv 2300 \text{ kg} \cdot \text{m}^{-3}$). Δz_k is the model layer thickness (in metres), for a total of n levels (here
 657 30 vertical levels).

658

659

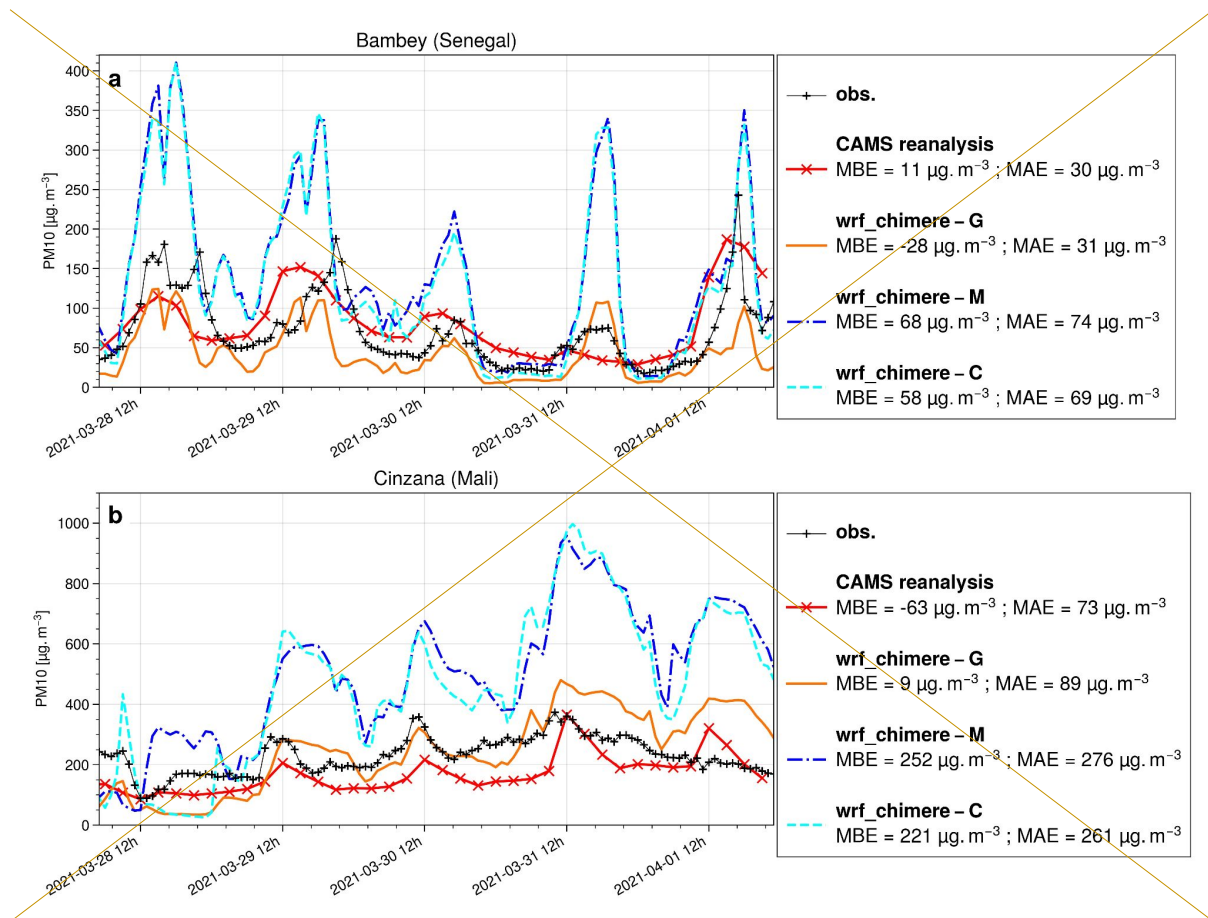


660
 661 **Figure 8** - Aerosol volume size distribution for the AERONET station located in a)
 662 Tamarasset, b) Zinder, c) Banizoumbou and d) DakarBelair. t_A and t_m indicate the
 663 times The time indicated corresponds to the time of the AERONET inversion product and the
 664 WRF-CHIMERE model respectively used for the comparison used for the comparison with
 665 the simulated aerosol size distribution. wrf_chimere-G, wrf_chimere-M and wrf_chimere-C
 666 refer to the WRF-CHIMERE simulations using GOCART, MERRA2 and CAMS as dust
 667 aerosol initial and boundary condition dataset respectively.

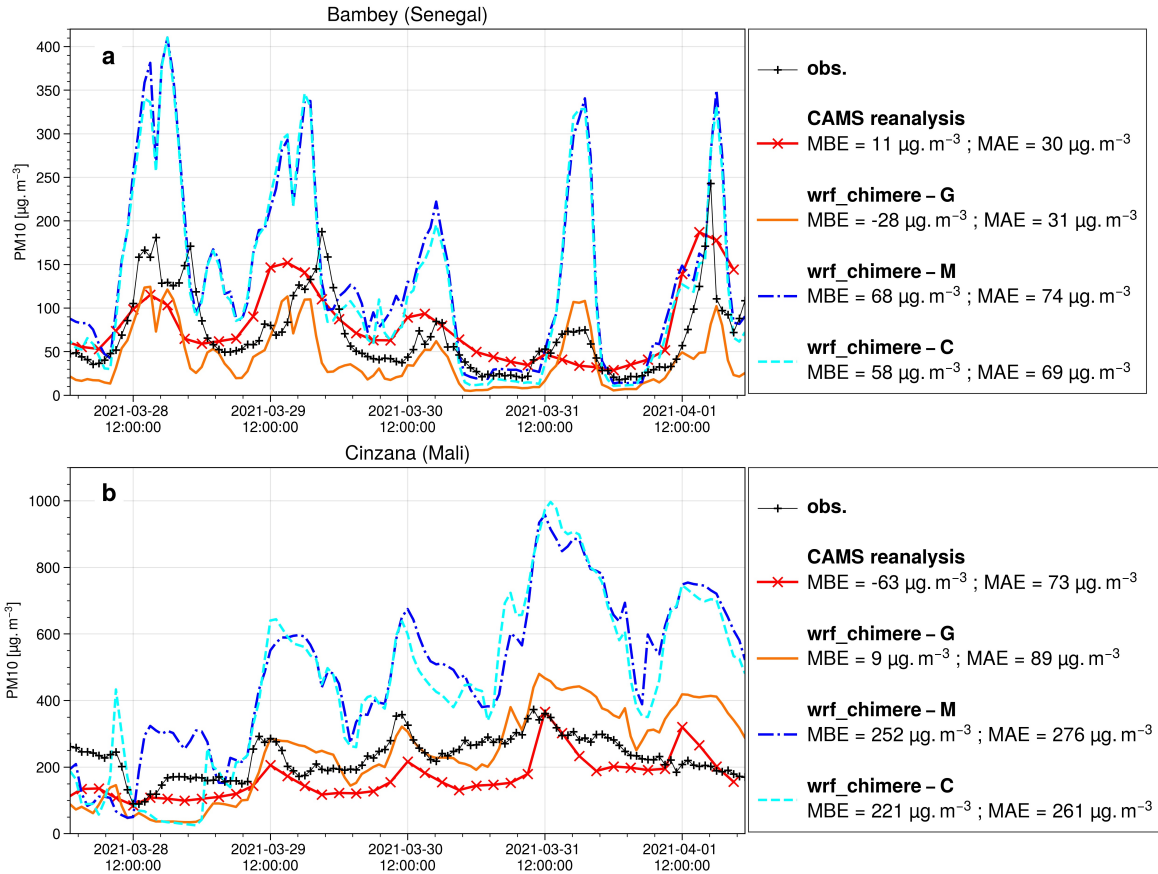
668
 669 The evaluation of the aerosol size distribution in Fig. 8 shows that the simulations generally
 670 have a dominant aerosol size mode shifted towards coarser sizes compared to the
 671 AERONET inversion product. The ground-based size distribution has a strong peak between
 672 1.14 μm and 5.00 μm , whereas the size distributions estimated by the WRF-CHIMERE
 673 simulations peak for coarser aerosol. For the Dakar Belair station (Fig. 8d), the AERONET
 674 inversion product indicates a first peak of lower intensity between 0.05 and 0.11 μm , which
 675 suggests the presence of aerosols other than desert dust. These aerosols may be of
 676 anthropogenic origin, given the proximity of the measurement site to the Senegalese capital.
 677 When comparing the size distributions between the three simulations with different dust
 678 aerosol initial and boundary condition dataset, it can be seen that the simulations driven with
 679 CAMS and MERRA2 reanalysis are relatively close and well separated from the one driven
 680 with the GOCART climatology. Notably, the dominant size bin in the simulation using
 681 GOCART dataset is consistently the largest particles, whereas with the aerosol from
 682 reanalyses, it is the aerosols between 5 μm and 10 μm . Consequently, the uncertainty

683 associated with the selection of the dust aerosol initial and boundary condition dataset is
 684 high when examining the aerosol size distribution, particularly for particles exceeding 5.00
 685 μm in diameter. The aforementioned uncertainties in the aerosol size distribution, which are
 686 linked to the choice of the dust aerosol initial and boundary conditions dataset, may be
 687 attributed to differences in the flow of desert dust entering the domain, as well as
 688 uncertainties in the transfer method carried out by the CHIMERE model to match the aerosol
 689 classes of these datasets to its own size distribution, described in section 2.2.3.
 690 As a result, the shift in the WRF-CHIMERE size distribution towards coarser particles
 691 compared to AERONET observations would result in a simulated AOD smaller than
 692 AERONET measurements. However, the opposite is observed (section 3.3). This suggests a
 693 positive bias in the simulated aerosol concentration, which would explain the positive bias in
 694 the AOD, while the coarser size distribution would tend to compensate.
 695
 696

3.5. Aerosol concentrations



697

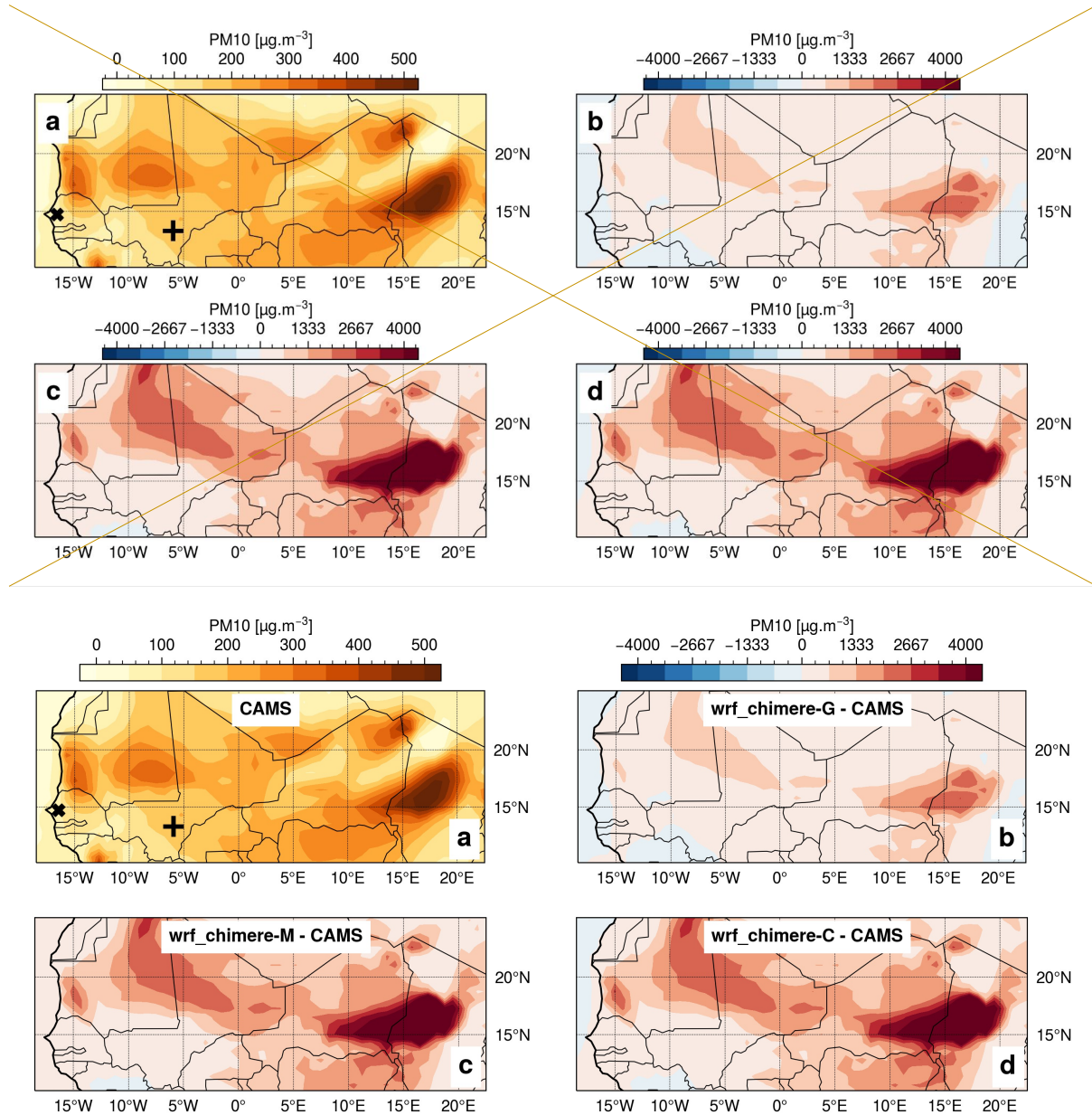


698
 699 **Figure 9** - Local comparison of CAMS reanalysis and simulated PM₁₀ surface concentrations
 700 with INDAAF network observations for a) Cinzana and b) Bambeey stations. *wrf_chimere-G*,
 701 *wrf_chimere-M* and *wrf_chimere-C* refer to the WRF-CHIMERE simulations using GOCART,
 702 MERRA2 and CAMS as dust aerosol initial and boundary condition dataset respectively.
 703 MBE is the mean bias error and MAE refers to the mean absolute error.

704
 705 The three simulations properly capture the dynamics of the PM₁₀ surface concentration with
 706 respect to the INDAAF ground measurement (Fig. 9) as correlation coefficients are around
 707 0.6 at Cinzana and close to 0.7 at Bambeey. The WRF-CHIMERE simulations driven with
 708 MERRA2 and CAMS dust aerosol datasets overestimate the surface PM₁₀ concentration
 709 peaks for Bambeey (Fig. 9a) and Cinzana (Fig. 9b), with high positive bias values of around
 710 63 g.m⁻³ at Bambeey and 247 g.m⁻³ at Cinzana. The latter station is closer to the dust
 711 aerosol sources. In contrast, the simulation using the GOCART dust aerosol dataset
 712 demonstrates superior performance in representing this variable, with an MAE that is
 713 approximately 60% and 70% lower than the two other simulations at Bambeey and Cinzana,
 714 respectively.

715 Furthermore, the uncertainty associated with the selection of initial and boundary condition
 716 dataset for dust aerosols is of a comparable magnitude to the simulation errors observed for
 717 surface PM₁₀ concentrations. Section 3.4 partly explains these discrepancies in surface PM₁₀
 718 concentration estimates between the simulation driven with the GOCART climatology and
 719 those driven with CAMS or MERRA2 reanalysis in terms of aerosol size distribution. These
 720 differences may also be attributed to variations in the size distribution of dust aerosol
 721 emissions or in the inflow of dust into the simulation domain and its aerosol size distribution.

722 Furthermore, Fig. 9 indicates that the CAMS reanalysis provides reliable estimates of
 723 surface PM₁₀ concentration, as evidenced by the fact it has the lowest MAE values.
 724 However, the Bambey and Cinzana ground measurements, which are the only two available
 725 for the case study, are situated at a considerable distance from the dust sources, limiting our
 726 ability to assess the accuracy of the CAMS reanalysis in capturing the dust event. Moreover,
 727 the CAMS reanalysis exhibits a negative bias at Cinzana, which is the closest site to the dust
 728 sources.
 729



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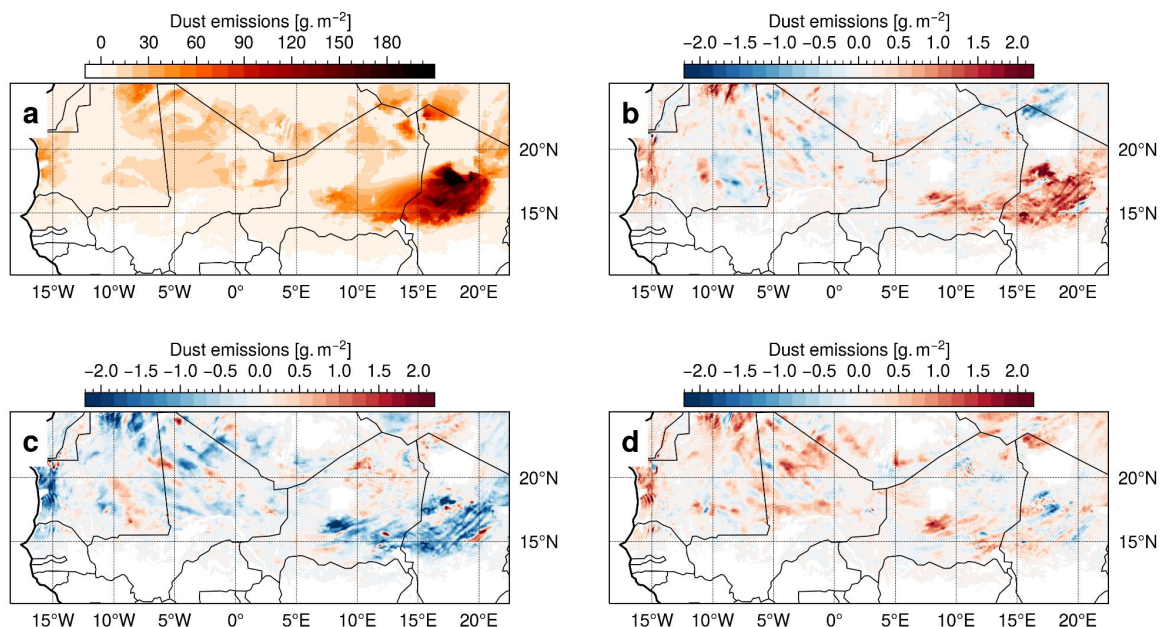
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Figure 10 - a) Mean from March 28th-00 UTC to April 2nd-00 UTC 2021 of CAMS reanalysis PM₁₀ surface concentration; x refers to the Bambey and + corresponds to Cinzana INDAAF stations. For panels b, c and d, PM₁₀ surface concentration mean differences from March 28th-00 UTC to April 2nd-00 UTC 2021 between each of the WRF-CHIMERE simulations driven by GOCART, MERRA2 and CAMS, respectively, and the CAMS reanalysis.

738 Figure 10 illustrates an overestimation of the PM₁₀ concentrations as compared to the CAMS
 739 reanalysis. This is particularly evident in dust source areas such as the Bodélé Depression.
 740 The WRF-CHIMERE simulation driven with the GOCART dataset is the closest to the CAMS
 741 reanalysis, with a mean estimate 3.6 times higher. However, this ratio reaches 8.6 for the
 742 simulations driven with the CAMS and MERRA2 reanalysis dataset.
 743 The mean standard deviation between the three WRF-CHIMERE simulations is 35% of their
 744 mean PM₁₀ surface concentration estimate. Consequently the uncertainty in the estimation of
 745 dust PM₁₀ surface concentration associated with the selection of the dust aerosol initial and
 746 boundary condition dataset is significant. The discrepancies between the simulation using
 747 the GOCART climatology and the two other ones using CAMS or MERRA2 reanalysis can
 748 be partly explained by the differences in the simulated aerosol size distribution, as shown in
 749 section 3.4.

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3.6. Dust emissions



752

753 **Figure 11 - a)** Total dust emissions flux from March 28th-00 UTC to April 2nd-00 UTC 2021,
 754 averaged between the three WRF-CHIMERE simulations. For panels b, c and d, total dust
 755 emissions individual differences between each of the WRF-CHIMERE simulations driven by
 756 GOCART, MERRA2 and CAMS, respectively, and the mean of the three WRF-CHIMERE
 757 simulations.

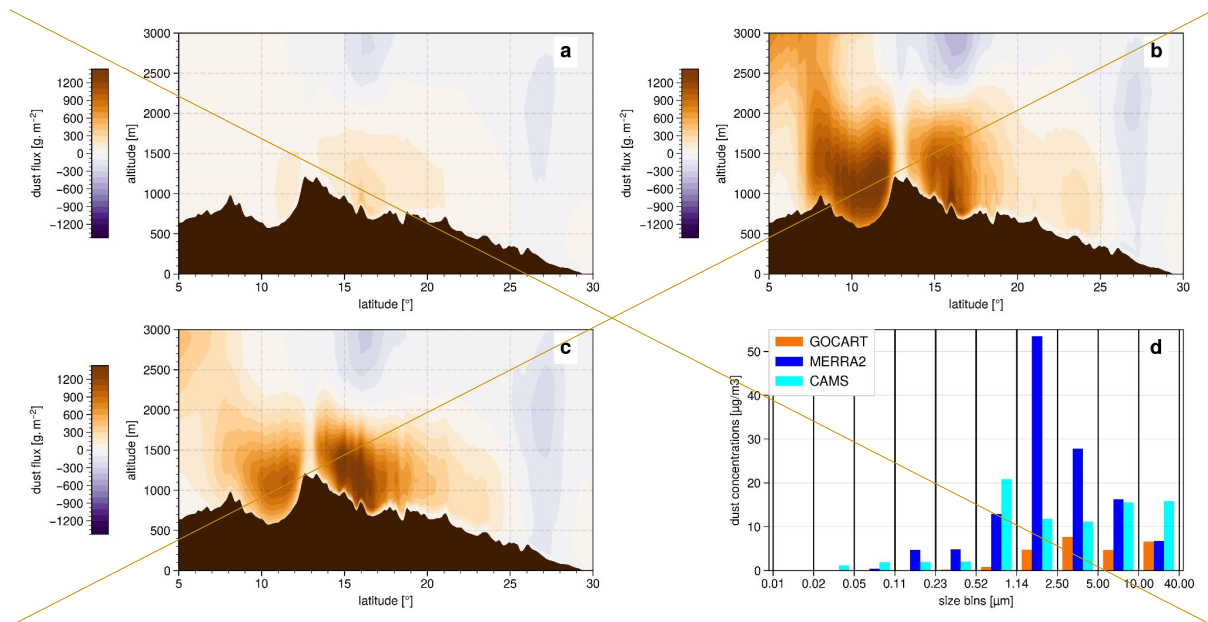
758

759 In terms of dust emissions (Fig. 11), the Bodélé Depression is, as expected, identified as
 760 the primary dust source area, with emissions reaching up to 244 g/m². The differences of the
 761 simulations with each of the three dust aerosol initial and boundary conditions dataset,
 762 relative to their mean, exhibit highest values in the source zones located at the Bodélé
 763 Depression and the South Atlas. Nevertheless, it is worth noting that there is a factor of 100
 764 in between the emissions in the Bodélé area (approximately 200g/m²) and the observed
 765 differences between the three simulations. Consequently, the uncertainties in dust emissions
 766 resulting from the choice of the dust aerosol initial and boundary conditions dataset can be
 767 considered negligible. As emissions are primarily influenced by surface wind, it can be
 768 inferred that the uncertainty generated by the dust aerosol driving dataset on the surface

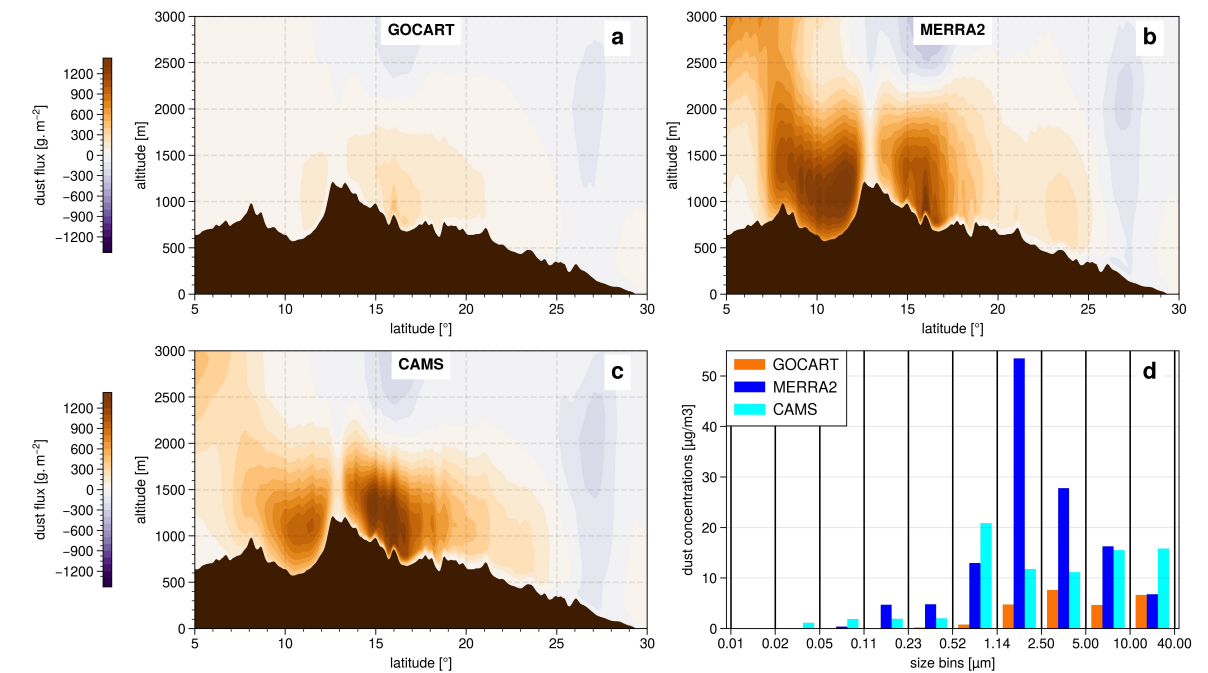
769 wind is negligible too, which is confirmed by Fig. S4. Additionally, the size distributions of the
 770 aerosols emitted during the case study are found to be identical (not shown). Therefore, the
 771 differences in dust surface concentration AOD and dust aerosol size distribution may be
 772 partly concentration may be attributed to the dust flows at the boundaries of the domain and
 773 are not linked to differences in simulated dust emissions within the domain. However, there
 774 is no observational data available to enable a quantitative evaluation of the accuracy of the
 775 emissions computed within the WRF-CHIMERE simulations.

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3.7. Dust boundary flux



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Figure 12 - Cumul of the dust flux at the eastern boundary of the simulation from March 28th-00 UTC to April 2nd-00 UTC 2021 for the WRF-CHIMERE simulation with a) GOCART, b) MERRA2 and c) CAM5 as dust aerosol initial and boundary conditions dataset; d) Dust

784 size distribution at the eastern boundary limit average during the case study period, from the
785 surface to 200hPa and over latitude. In panel abc, the dust flux is derived as the product
786 between the dust aerosol concentration and the zonal wind, and positive values of the dust
787 flow indicate a flow entering the simulation domain.

788

789 As shown in Fig. 1b, the dust event is associated with a strong Harmattan flow,
790 characterised by a northeasterly flow in the lower layer. It is thus interesting to quantify the
791 dust inflow associated with each of the dust aerosol initial and boundary conditions dataset
792 for the eastern domain boundary. The lowest dust flux is observed with GOCART (Fig. 12a),
793 with a maximum of approximately 480 g/m². In contrast, MERRA2 and CAMS (Fig. 14 b
794 and c respectively) exhibit higher dust fluxes, with maximum values of around 1650 g/m².
795 The maximum flow is around 10°N for MERRA2, while for CAMS, it is closer to 16°N. Given
796 that GOCART is a climatology, it is reasonable to expect a lower dust flux compared to the
797 CAMS and MERRA2 reanalyses, which are real case simulations incorporating data
798 assimilation of AOD. This is particularly true for the presented case study, which involves an
799 intense dust event associated with a Harmattan flow.

800 There are also significant differences in both quantity and distribution by aerosol size bin
801 (Fig. 12d). MERRA2 exhibits a strong dominant mode for the class between 1.14 µm and
802 2.50 µm, while CAMS shows significant values from 0.52 µm to 40 µm, with a maximum for
803 the size class between 0.52 µm and 1.14 µm. Finally, the GOCART model displays a lower
804 variability between 1.14 µm and 40.00 µm, with the maximum occurring for the size class
805 between 2.55 µm and 5.00 µm.

806 The eastern dust fluxes at the boundary significantly vary depending on the dataset used as
807 dust aerosol initial and boundary conditions, both in terms of quantity and size distribution.
808 The reanalysis dataset, CAMS and MERRA2, are expected to provide a more accurate
809 representation of dust flux in terms of quantity as they are real case simulations assimilating
810 observational data in their calculations, as compared to GOCART which is a climatology.
811 However, GOCART provides a more comprehensive description of aerosol size distribution
812 with seven classes, in comparison to CAMS, which has only three classes but proposes a
813 higher horizontal resolution. While GOCART considers the effect of aerosol size to be
814 essential, CAMS assumes the horizontal resolution to be a key parameter. MERRA2 is the
815 most comprehensive of the three datasets, with the highest horizontal resolution, and an
816 aerosol size distribution that is close to the GOCART one with five classes. ~~Despite the
817 absence of observational data that would permit a quantitative evaluation of the eastern dust
818 fluxes, the aforementioned elements suggest that the MERRA2 dataset might be more
819 accurate.~~

820 As a result, and in consideration of the negligible uncertainty in dust emissions within the
821 simulation domain related to the choice of the dataset for dust aerosol initial and boundary
822 conditions (see 3.6), these differences in eastern dust fluxes appear to account for the
823 uncertainties of the simulated surface dust aerosol concentrations (see 3.5) and dust aerosol
824 size distribution AODs (see 3.43).

825

826

3.8. Discussions

827 The evaluation of the simulated GHI at the Zagtouli solar power plant and the Banizoumbou
828 site (Fig. 2) ~~indicates a significant enhancement in surface solar irradiance shows a clear
829 improvement in its estimation when WRF is coupled with CHIMERE. Specifically, to
830 CHIMERE rather than not as~~ the local MAE is reduced by approximately around 75%. This
831 confirms the relevance of incorporating the dust radiative effect with a coupling approach, in

832 comparison with the operational forecasts currently employed based on meteorological
833 models alone. During the dry season, dust events similar to the one presented here, with
834 emissions at Bodélé and then transport of the plume westwards, are common. This work
835 therefore calls for forecasters in the photovoltaic sector to better account for the desert dust
836 cycle in their forecast products. This local evaluation also highlights the potential benefits of
837 using a regional model rather than a global product, as the WRF-CHIMERE simulations
838 outperform the CAMS gridded solar radiation product with an average MAE reduced by
839 approximately 38% at the Zagtouli solar farm and by 70% at the Banizoumbou site, which is
840 closer to dust sources. These discrepancies are corroborated by the regional comparison
841 presented in Figure 3, which reveals that the mean WRF-CHIMERE GHI estimate is 5%
842 lower than the CAMS solar radiation dataset. Additionally, the latter does not exhibit a
843 geographical pattern with lower GHI estimation along the dust plume trajectory, in contrast to
844 the WRF-CHIMERE simulations. These results confirm those from Sawadogo et al. (2023)
845 who recently showed that the CAMS reanalysis have low performances in estimating solar
846 irradiance during high AOD episodes like the one studied here. Furthermore, the comparison
847 reveals that incorporating indicates that the incorporation of dust in the simulation reduces
848 surface solar irradiance radiation by 18% infor this case study. This reduction is notably
849 higher but remains within the same order of magnitude as previous studies that integrated
850 dust aerosol information for solar estimation. For example, Masoom et al. (2021) in India and
851 Mostamandi et al. (2023) in the Arabian Peninsula reported GHI reductions due to dust of
852 approximately 5-10%. This discrepancy underscores the potential variability of the dust
853 impact on solar irradiance depending on the method used to account for dust effects in the
854 simulations. In light of the anticipated expansion of PV production in West Africa, this point
855 underscores the potential consequences of such dust events if they are not accurately
856 predicted.

857
858 The evaluation of local surface temperature (Fig. 4) reveals contrasting results regarding the
859 effectiveness of the coupled approach. It demonstrates an average local MAE reduction of
860 approximately 10% compared to the WRF-only simulation. However, the main differences
861 occur mainly at night, when no photovoltaic is produced, as previously observed by Yue et
862 al. (2010) and Briant et al. (2017). It can be attributed to the opposing radiative forcing
863 effects of dust aerosols across different wavelength ranges. In the case of longwave, which
864 corresponds to terrestrial radiation, the presence of dust aerosols has a warming effect.
865 Conversely, for shortwave, which corresponds to solar radiation, the presence of dust
866 aerosols induces a cooling effect. Consequently, during night-time when solely terrestrial
867 radiation is present, there is an increase in surface temperature. During day-time a
868 competition between the warming effect of terrestrial radiation and the cooling effect of solar
869 radiation ensues. The net impact is a decrease in surface temperature, indicating that the
870 effect of solar radiation dominates, with the cooling effect exceeding the warming effect
871 (Sokolik and Toon, 1999). ~~The regional evaluation in Fig. 5 confirms these contrasting~~
872 ~~results and indicates a reduction of regional MAE by about 14% with the coupling rather than~~
873 ~~WRF alone. The overestimation of surface temperature in dusty areas with the coupling, not~~
874 ~~present in the WRF only simulation, reveals the dominant aerosol warming effect during~~
875 ~~night time as compared to the cooling effect during daytime. These statements strongly~~
876 ~~depend on the accuracy of the ERA5 reanalysis which serves as reference. ERA5 integrates~~
877 ~~data assimilation but does not consider aerosol information in its calculation. Due to the~~
878 ~~limited ground measurements in the Saharan region to constrain the reanalysis, it is possible~~
879 ~~that ERA5 underestimates the aerosol effect in dusty areas.~~

880 [The regional evaluation in Fig. 5 confirms these contrasting results and indicates a reduction](#)
881 [of regional MAE by about 14% with the coupling rather than WRF alone. The overestimation](#)
882 [of surface temperature in dusty areas with the coupling, not present in the WRF only](#)
883 [simulation, reveals the dominant aerosol warming effect during night time as compared to](#)
884 [the cooling effect during daytime. These results align with those of Briant et al. \(2017\), who](#)
885 [estimated dust-induced warming of up to +5°C during nighttime and cooling of approximately](#)
886 [-1°C during daytime in a 2012 dust event in West Africa. These statements strongly depend](#)
887 [on the accuracy of the ERA5 reanalysis which serves as reference. ERA5 integrates data](#)
888 [assimilation of temperature and incorporates aerosol radiative effects through prescribed](#)
889 [monthly climatologies from the GOCART model, but does not dynamically simulate aerosols.](#)
890 [Due to the limited ground measurements in the Saharan region to constrain the reanalysis,](#)
891 [and to the significant biases that can come when considering a coarse climatology for the](#)
892 [radiative effects of aerosols to represent an intense dust event, it is possible that ERA5](#)
893 [underestimates the aerosol effect in dusty areas.](#)

894
895 Nevertheless, despite the improvements demonstrated in solar ~~irradiance~~ [radiation](#) and
896 surface temperature estimation, the WRF-CHIMERE simulations exhibit a notable positive
897 bias in terms of AOD, as evidenced by the local and regional evaluations presented in Figs.
898 6 and 7. This overestimation cannot be attributed solely to differences in aerosol
899 concentrations, as the simulations yield markedly disparate surface concentrations of PM10,
900 depending on the dust aerosol initial and boundary condition dataset chosen (Fig. 10)-, while
901 this discrepancies do not appear in the AOD estimates. [However, the results from Yahi et al.](#)
902 [\(2013\) and Léon et al. \(2020\) emphasized the importance of considering dust plume height](#)
903 [when linking surface PM10 concentrations to AOD. Therefore, differences in the vertical](#)
904 [distribution of the dust plume, not evaluated in this study due to the lack of quantitative](#)
905 [observational data, could account for part of the observed discrepancies between simulated](#)
906 [AODs and surface PM10 concentrations.](#) This excess of aerosol load may be attributed to an
907 overestimation of emissions within the domain, but this cannot be verified as there is not any
908 such measurement. The incoming flux of dust in the domain plays a minor role as shown in
909 Fig. 12 where the flux significantly also varies depending on the dust aerosol initial and
910 boundary condition dataset employed, while these differences are not any more present in
911 the simulated AOD estimates. Additionally, the underestimation of aerosol deposition, by
912 sedimentation (not studied in this research) could be at the origin of the overestimation of the
913 simulated dust loads. Finally, another potential explanation for these AOD biases may be the
914 inaccuracies in the dust radiative properties incorporated in the CHIMERE model calculation
915 (see Table S1 and S2). These depend on the mineralogical composition of the desert dust
916 particles emitted, which are considered uniform in this work. [The radiative properties of](#)
917 [aerosols also depend on their granulometry. In the CHIMERE model, dust aerosols are](#)
918 [treated as spherical particles in the calculation of their radiative properties using Mie theory,](#)
919 [which introduces biases. Adbiyi et al. \(2023\) showed that ellipsoidal dust particles have a](#)
920 [slightly higher mass extinction efficiency compared to spherical particles. As a result,](#)
921 [accounting for ellipsoidal dust aerosols would lead to a slight increase in AOD associated](#)
922 [with a small decrease in GHI. This study further indicates that dust particles with radii](#)
923 [smaller than 20.0 \$\mu\text{m}\$ are the primary contributors to dust AOD for shortwave radiation, with](#)
924 [the contribution from larger particles being an order of magnitude lower. Therefore, including](#)
925 [particles larger than 40.0 \$\mu\text{m}\$ in the CHIMERE model would not significantly affect AOD and](#)
926 [GHI estimates. This is corroborated by Mostamandi et al. \(2023\), who demonstrated that](#)
927 [dust particles with radii smaller than 3 \$\mu\text{m}\$ are primarily responsible for the reduction in solar](#)

928 | irradiance, while particles larger than 10 μm mainly contribute to dust deposition, which was
929 | not examined in this study.—————

930

931 The uncertainty associated with the choice of the large scale dust aerosol initial and
932 boundary condition dataset is very low when considering the variables of interest for solar
933 production, namely GHI and surface temperature (Fig. 3 and 5). This uncertainty is also low
934 compared to the performance of simulations for AOD estimation (Fig. 7). This result is similar
935 when examining dust emissions within the domain, which are nearly identical for the three
936 coupled simulations (Fig. 11). This can be explained by the fact that dust emissions depend
937 | on the cubed of surface wind speed (Marticorena and Bergametti, 1995) which present
938 no significant signature of the selection of the dust aerosol initial and boundary conditions
939 (Fig. S4). The aerosols emitted within the chosen domain are much greater than those
940 entering, as the domain accounts for the main source zones. This is why the simulations are
941 not that sensitive to dust aerosol large-scale dataset employed. The results regarding the
942 uncertainty associated with the choice of the dust aerosol initial and boundary condition
943 dataset differs when examining various elements of the dust life cycle. Indeed, aerosol size
944 distributions vary significantly between the simulation driven with GOCART on one hand,
945 and simulations driven with CAMS and MERRA2 on the other hand. GOCART climatology
946 over-represents aerosols larger than 10 μm compared to the CAMS and MERRA2
947 reanalyses. These differences partially account for the significant deviation in surface PM_{10}
948 concentration estimates (Fig. 10), indicating that reanalysis-type datasets result in much
949 higher values, up to 3 times higher, compared to climatological-type data which is closer to
950 ground observations. The dust flux entering the domain may also partly explain these
951 differences. In fact, this flux is very low with GOCART, with values up to 3.5 times lower than
952 CAMS and MERRA2 (Fig. 12). The size distribution of this incoming aerosol flux is also a
953 determining factor.

954

955 | **4. Conclusion and perspectives**

956 This study aims to evaluate the ability of the WRF-CHIMERE coupling to simulate GHI
957 during a typical dust event in the dry season in West Africa. This event is characterised by a
958 Harmattan flux associated with significant desert dust emissions over the Bodélé
959 Depression, with the dust plume subsequently transported westward. This work
960 demonstrates the utility of coupling a meteorological model with a desert aerosol life cycle
961 model to represent such events, particularly for improving solar forecasts. Indeed, GHI
962 estimations are markedly enhanced with this approach compared to using a meteorological
963 model alone with a 75% reduction of local MAE. Nevertheless, the performance of the WRF-
964 CHIMERE simulations in representing the aerosol load of this event is more controversial.
965 There is an overall overestimation of AOD and PM_{10} surface concentration by the coupled
966 model in the North Sahelian-Saharan zone.

967 This work also aims at investigating whether the performance of the simulations can be
968 improved by changing the dust aerosol initial and boundary condition dataset, and to
969 estimate the uncertainty associated with this choice. The results show that this selection has
970 | almost no influence on the estimation of the solar irradiance~~radiation~~, surface temperature
971 and AOD. On the contrary, the choice of the dust aerosol initial and boundary condition
972 dataset has a significant impact on the surface PM_{10} concentration and the aerosol size
973 distribution.

974

975 This work outlines new research perspectives. Firstly, we observe the difficulty of evaluating
976 simulations in West Africa due to the scarcity of available observations. Establishing a
977 denser measurement network or conducting observation campaigns, particularly for GHI,
978 would help research on solar estimation and forecasting in this region. Additionally, the
979 WRF-CHIMERE simulations demonstrate significant biases in terms of AOD and PM₁₀
980 surface concentration which are not fully explained here. One potential explanation for this is
981 an overestimation of dust emission, for which no evaluation is possible. Furthermore,
982 studying aerosol deposition (not conducted in this work) would complement the study of the
983 desert aerosol life cycle. On the one hand, an underestimation of deposition might be a
984 contributing factor to the overestimation of the simulated aerosol load. On the other hand,
985 dust deposition on solar panels affects solar production by masking the available solar
986 irradiance radiation (soiling effect), and this should be taken into account in forecasting
987 systems to conduct optimised cleaning operations. Finally, the study focuses on a typical
988 dust event during the dry season, presenting essentially aerosol-radiation interaction. It
989 could be beneficial to test such simulation configuration for more complex cases involving
990 cloud presence. Indeed, the interaction between aerosols and clouds have a significant
991 impact on solar forecasting by increasing albedo, extending cloud lifespan, and promoting
992 cloud formation through increased condensation nucleus concentration (indirect aerosol
993 effects).

994

995 **Code and data availability**

996 WRF namelist configuration files, CHIMERE parameter files, Python codes exploited in this
997 study and GOCART climatology data can be found on the following Zenodo repository:
998 <https://zenodo.org/records/10808476>

999 ERA5 data can be found on the Copernicus Climate Data Store service :
1000 <https://cds.climate.copernicus.eu/cdsapp#!/home>

1001 CAMS data were downloaded on the Copernicus Atmosphere Data Store service :
1002 <https://ads.atmosphere.copernicus.eu/cdsapp#!/home>

1003 MERRA2 data can be found on the dedicated platform from NASA :
1004 <https://goldsmr5.gesdisc.eosdis.nasa.gov/data/MERRA2/>

1005 Data from AMMA ground measurements stations can be accessed from the dedicated
1006 website : <https://amma-catch.osug.fr/-jeux-de-donnees->

1007 INDAAF web page allows access to the data : <https://indaaf.obs-mip.fr/catalogue/>

1008 AERONET data measurements and inversion products are available through the following
1009 link: <https://aeronet.gsfc.nasa.gov/>

1010 The MODIS satellite observations are available on the “Level-1 and Atmosphere Archive &
1011 Distribution System Distributed Active Archive Center” platform from NASA :
1012 <https://ladsweb.modaps.eosdis.nasa.gov/>

1013

1014 **Author contributions**

1015 LC, SA, CL conceptualised the study. LC performed the simulations, the analysis and the
1016 editions of the figures. LC, SA, CL, GB, BM, GS, CB, RL and JT discussed the results. LC
1017 wrote the paper

1018

1019 **Competing interest**

1020 The contact author has declared that none of the authors has any competing interests.

1021

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1036

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