



# Retrieval of top-of-atmosphere fluxes from combined EarthCARE LiDAR, imager and broadband radiometer observations: the BMA-FLX product

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**Abstract.** The Earth Cloud, Aerosol and Radiation Explorer (EarthCARE) satellite mission is expected to provide new insights into aerosols, clouds, and radiation. The satellite's payload include four instruments design to synergistically retrieve vertical profiles of clouds and aerosols, along with the atmospheric radiation data. This will enable the determination of atmospheric heating rates and top-of-atmosphere radiances and fluxes. This paper focuses on the BMA-FLX processor, an algorithm specifically created, developed, and validated to retrieve thermal and solar top-of-atmosphere radiative fluxes from longwave and shortwave radiances, measured along-track by the EarthCARE Broadband Radiometer (BBR) instrument. These measurements are co-registered either at the surface or at the radiatively most significant vertical layer of the atmosphere in cloudy condition (reference level). The Multi-Spectral Imager (MSI) and Atmospheric LiDAR (ATLID) on-board EarthCARE support cloud identification, while meteorological data from the European Centre for Medium-Range Weather Forecasts provide the surface and atmospheric necessary information. In the BMA-FLX processor, flux is estimated independently for each BBR view using different approaches for the longwave and shortwave radiances. A combined flux, derived from co-registered radiances at the reference level, is provided as the best estimate for a given scene. The radiance-to-flux conversion algorithms have been successfully validated through end-to-end verification using L1 and L2 synthetic data for three EarthCARE orbits. The BMA-FLX's objective is to achieve radiative closure for EarthCARE with solar and thermal fluxes within  $10 \text{ Wm}^{-2}$ .

## 15 1 Introduction

The accurate representation of the complex interplay between aerosols, clouds, and radiation in climate models remains a significant uncertainty in climate projections (Boucher et al., 2013). The Earth Cloud Aerosol and Radiation Explorer (EarthCARE) mission, a collaboration between the European Space Agency (ESA) and the Japan Aerospace Exploration Agency (JAXA), aims to enhance our understanding of the interaction between clouds and aerosols and their impact on Earth's radiation balance (Illingworth et al., 2015; Wehr et al., 2023; Eisinger et al., 2024). EarthCARE is the first platform to measure both vertical structure and horizontal distribution of cloud and aerosol fields simultaneously from one platform using one passive



and two active instruments. Additionally, the emitted longwave and reflected shortwave radiation are measured. The four scientific instruments observe the Earth from a sun-synchronous polar orbit crossing the equator in the early afternoon to optimize observation of convective cloud systems in the tropics. The two active instruments, cloud profile radar (CPR) and atmospheric  
25 LiDAR (ATLID), provide information on the vertical structure of the micro- and macro-physical properties of clouds, aerosols, and hydro meteors (Kollias et al., 2022; Irbah et al., 2023; Mroz et al., 2023; Donovan et al., 2023; van Zadelhoff et al., 2023b). The passive instrument, the multi-spectral imager (MSI), gives horizontal context to the observed scene (Wandinger et al., 2023; Docter et al., 2023; Hünerbein et al., 2023b, a; Haarig et al., 2023) and is used to create 3D scenes (Qu et al., 2023). The 3D information of the atmosphere based on measurements of these instruments (Mason et al., 2023) will be used as input for  
30 1D and 3D-radiative transfer calculations (Cole et al., 2022). The calculated top-of-atmosphere (TOA) solar and thermal fluxes leaving Earth to space are finally compared to the short and longwave fluxes, estimated from radiance measurements from the broad-band radiometer (BBR) to achieve radiative closure (Tornow et al., 2015; Barker et al., 2024).

The BBR aboard EarthCARE will measure accurate shortwave (SW) and total-wave (TW) radiances in a push broom along-track configuration at three fixed viewing angles (nadir, 55° fore, and 55° aft) (Caldwell et al., 2017). The telescopes measure  
35 TW radiances from 0.25  $\mu\text{m}$  to beyond 50  $\mu\text{m}$ . The shortwave channel covers radiances from 0.25  $\mu\text{m}$  to 4  $\mu\text{m}$  by applying an uncoated synthetic quartz filter mounted on a rotating drum to the telescopes. Synthetic longwave (LW) radiances are obtained by subtracting the SW from the TW channel. The SW-LW inter-channel contamination and the instrument spectral response effects are removed in the BM-RAD processor (Velázquez Blázquez et al., 2023). The obtained "unfiltered" SW and LW radiances are then used as input for the BMA-FLX processor, described hereafter to estimate top-of-atmosphere fluxes.  
40 BM-RAD reports radiances for the three viewing angles at fixed spatial resolutions, being the "standard resolution" of 10 x 10  $\text{km}^2$ , and the configurable "assessment domain resolution" of 21 km along-track by 5 km across-track employed in the radiative closure (Barker et al., 2024).

Due to the highly anisotropic character of some physical phenomena, like the reflection of solar radiation by clouds, the estimation of the radiative fluxes from measured radiances at a single Sun-observer geometry are challenging. In recent years,  
45 various approaches have been developed to estimate the anisotropy of the observed scenes using the so-called angular distribution models (ADM). An overview of different approaches is described in Gristey et al. (2021).

Several studies have analyzed different radiance-to-flux approaches for the BBR over the past years of EarthCARE development (Domenech et al., 2011a; Domenech et al., 2011b; Domenech and Wehr, 2011; Tornow et al., 2019, 2020, 2021). This paper presents the final implementation of the mission's operational flux retrieval algorithm. The BMA-FLX SW algorithm is  
50 created per scene type and constructed from six years of Clouds and the Earth's Radiant Energy System (CERES) and Moderate Resolution Imaging Spectroradiometer (MODIS) Terra and Aqua measurements using an artificial neural network approach. The BMA-FLX LW algorithm follows the approach described in Clerbaux et al. (2003a), and is based on correlations between BBR radiance anisotropy and the spectral information provided by the EarthCARE's MSI radiances derived from radiative transfer calculations (RTC). As the processor name indicates, inputs from BBR (radiances), MSI (radiances, cloud top height,  
55 and cloud fraction) and ATLID (cloud top height) are used to estimate fluxes.



Providing instantaneous thermal and solar flux estimates, BMA-FLX allows a final comparison with calculated TOA fluxes using radiative transfer models and evaluates if the mission goal of  $10 \text{ Wm}^{-2}$  has been reached (Wehr et al., 2023).

## 2 Algorithm description

Building on the experience gained from previous Earth Radiation Budget missions such as CERES (Wielicki et al., 1996), 60 GERB (Harries et al., 2005), and ScaRaB (Kandel et al., 1998), an angular distribution model (ADM) methodology is considered an appropriate candidate to meet EarthCARE's radiative requirements.

The flux leaving the Earth-atmosphere system is obtained by integrating the radiance field,  $L(\theta_0, \theta, \phi)$ , which depends on the solar zenith angle (SZA,  $\theta_0$ ), the viewing zenith angle (VZA,  $\theta$ ), and the relative azimuth angle between the Sun and the satellite view (RAA,  $\phi$ ) as follows:

$$65 \quad F(\theta_0) = \int_0^{2\pi} \int_0^{\pi/2} L(\theta_0, \theta, \phi) \cos(\theta) \sin(\theta) d\theta d\phi. \quad (1)$$

To estimate the instantaneous top-of-atmosphere radiative flux from a single radiance measurement, some information of the angular variation of the radiation field that constitutes the flux is needed. A primary error in deriving flux arises from the lack of knowledge of the target's anisotropy. The ADMs have accurately represented this variation (Su et al., 2015), and so are used as the basis for flux retrieval.

70 An ADM estimates a radiative flux from a radiance observation from a single viewing angle by taking the mean behaviour of the anisotropy of the radiance field. It is defined by the anisotropic factor ( $R$ ), the ratio between the equivalent Lambertian flux and the actual flux:

$$R(\theta_0, \theta, \phi) = \frac{\pi L(\theta_0, \theta, \phi)}{F(\theta_0)}. \quad (2)$$

The SW flux retrieval algorithm is defined independently from the LW design since the anisotropy of the radiance field in the 75 SW and LW regimes depends physically on different geophysical parameters.

The BBR instrument observes each target on Earth from three different directions almost simultaneously, providing a detailed view of scene anisotropy. However, the BBR's multi-pointing capability is not directly utilized in the flux retrieval. This approach aligns with the EarthCARE contingency plan, which mandates that flux estimates must be retrievable from each individual telescope to address potential critical failures. The radiances acquired by the three telescopes for a single scene 80 are not converted into a flux estimate; instead, different ADMs are applied to the radiances and three fluxes are derived for the observed scene. As the outgoing flux is only dependent on the viewing geometry and the radiometric properties of the atmospheric-surface domain, the three fluxes estimated by the ADMs should result in a similar flux assuming perfect instrument and retrieval responses. This is not the case in a real scenario, where discrepancies in ADM-based fluxes obtained from different viewing geometries for the same surface-atmosphere scene can still be significant (Domenech et al., 2012).



## 85 2.1 BMA-FLX SW algorithm

### 2.1.1 SW radiance-to-flux description

Domenech et al. (2011b) developed angular models for the BBR from Monte-Carlo RT simulations to construct a synthetic ADM specifically defined for the multi-pointing capability of the BBR. Two main conclusions were extracted: methodologies entirely based on simulated data are biased and difficult to extrapolate for real use; and multi-view flux conversion algorithms  
90 improve performance of single-view ADMs in highly anisotropic scenes (Domenech et al., 2012).

The artificial neural network (ANN) method described by Loukachine and Loeb (2003, 2004) and adopted for ScaRaB-3 (Viollier et al., 2009) using CERES data to train the models, was modified to incorporate simultaneous use of along-track radiance measurements in Domenech and Wehr (2011); Tornow et al. (2019). These studies are further developed in the BMA-FLX SW algorithm, where the radiance-to-flux conversion algorithm employs a feed-forward backpropagation ANN to model  
95 the ADM-based fluxes from the CERES instrument (Loeb et al., 2005; Su et al., 2015). Backpropagation is the generalization of the Widrow-Hoff learning rule (Widrow and Lehr, 1990) to multiple-layer networks. Input vectors and the corresponding target networks are used to train a network until it can approximate a function that associates the input vectors with specific output vectors. The backpropagation employed here is a gradient descent algorithm in which the network weights are moved along the negative of the performance function gradient.

100 The BMA-FLX processor retrieves fluxes for each BBR telescope, checks the consistency of the estimates, and combines them according to their estimated errors,  $\varepsilon_F$ , and the corresponding radiance error provided by the BM-RAD processor,  $\varepsilon_R$ .

### 2.1.2 Scene definition

By scene definition we understand the target classification based on surface and cloud properties, and viewing geometry that determines the ADM model applied to the BBR observation. During the ADM development, the scene definition concludes the  
105 number of ANN training sets, and hence the number of trained networks that build the ADM.

Our scene definition relies on four cloud fractions (cloud-free,  $> 0.1$  and  $< 50$ ,  $\geq 50$  and  $< 99$ , and overcast) and six static surface type (ocean, forests, savannas, grasslands, shrubs, desert, and permanent snow), plus two dynamic ones (fresh snow and sea-ice) for each scattering direction (backward and forward). Cloud fractions retrieved from the CERES/MODIS algorithm and based on MODIS pixel-level measurements (Minnis et al., 2021) are used in the training datasets. The static categories  
110 considered for training are obtained from the International Geosphere–Biosphere Programme (IGBP) land cover, while the fresh snow and sea-ice surface types are derived from a combination of the microwave snow/ice map from the National Snow and Ice Data Center (NSIDC) and the snow/ice map from the National Environmental Satellite, Data and Information Service (NESDIS), available in the Single Scanner Footprint TOA/Surface Fluxes and Clouds (SSF) Edition 4 product.



### 2.1.3 Model training

115 A geophysical database for model training has been created using six years of Level 2 CERES radiance and flux measurements. For cloudy scenes, MODIS radiances and cloud masks matching the CERES measurements are collected. Additionally, climatology data on surface albedo and aerosol optical depth, along with meteorological reanalysis data on atmospheric gases, wind speed, and vegetation for cloud-free scenes, are gathered.

The geophysical database is designed to match the scene definition. For each scene class, pairs of input and output vectors  
120 are selected as training datasets. The output values represent CERES anisotropic factors, while the input vectors are chosen from a predefined list of parameters that influence scene anisotropy. These parameters may or may not effectively characterize anisotropy for each scene class. To determine the most significant input parameters for each class and BBR-received scattering direction, we assess the variables' importance in reproducing the anisotropic factors. This assessment is conducted using two independent tests: a Random Forest regression-based permutation test and a Genetic Algorithm applied to a Linear Model,  
125 both of which largely agree on the optimal subsets of input parameters. SW radiances at the BBR viewing directions and the illumination/viewing geometry from CERES are consistently included as inputs. Since SW ADMs are developed separately for nadir, forward scatter, and backward scatter directions, the CERES observations are matched with either forward or backward scattering directions based on the relative azimuth angle during the input selection process.

To consider the anisotropic characteristics over cloud-free land surfaces, we use RossThick-LiSparse bidirectional re-  
130 flectance factor (BRDF) parameters that are taken from a climatology derived from the MODIS MOD43B product (Qu et al., 2022). Additionally, we employ the leaf-area index (LAI) for high and low vegetation, and the aerodynamic surface roughness length from meteorological reanalysis (Poli et al., 2016). We also consider the hotspot effect as input to describe the enhanced reflectivity of the surface for an observational geometry close to the solar illumination geometry. Over ocean, aerosol optical depth (AOD) and wind-speed are selected as parameters for the network training. AOD is based on an AeroCom climatology  
135 (Koffi et al., 2016), and wind-speed from meteorological reanalysis (Poli et al., 2016). In addition, the glitter radiance normalized by the incident irradiance (sun-glint reflectance) given by Jackson and Alpers (2010) is also used in the training. To further consider anisotropy changes due to atmospheric characteristics, we also considered the total-column water-vapour and total-column ozone from reanalysis. Snow and ice surfaces are assessed independently.

The inputs to create the training sets for cloud scenes include MSI-like MODIS radiances and cloud cover. The assumption  
140 behind this choice is that the non-linear combination of narrow-band (NB) radiances provides sufficient information on the cloudy scene anisotropy, so that the use of imager retrieved cloud properties is not necessary. Imager radiances are considered independently over the cloud free part and the cloudy part of the observed scene. The best combination of NB radiances includes the 0.67, 0.865, 10.8, and 12.0  $\mu\text{m}$  MSI bands.

Each training dataset is employed to train a network. Since the networks depend on the bias/weight initialization, a Nguyen-  
145 Widrow initialization method is used to create additional models for each scene class. Resulting ANN models are crosschecked against a validation dataset (20% of the training data) and the networks with better performance are selected for each scene class of the ADM. After training, the resulting networks are evaluated against a CERES validation dataset. The root mean



square errors (RMSE) of the ANN-based flux estimates and original CERES derived fluxes define the theoretical uncertainties for each scene class.

#### 150 **2.1.4 Scene identification**

The scene identification is the first step in the flux retrieval. The illumination and viewing geometry of the scene classify the fore and aft BBR observations into either the forward or backward scattering regime based on their relative azimuth angle.

155 Surface types at a 1-km resolution are collected from a simplified version of the Global Land Cover Characterization (GLCC) dataset (Belward and Loveland, 1996) at the BBR pixel level, and the predominant surface type is selected for cloudy scenes. In clear-sky scenarios, the retrieval algorithm employs the two surface types with the highest coverage within the BBR pixel. The mixed flux is then calculated by weighting the average of the flux estimates for these two types by their surface area and albedo, following the methodology outlined in Bertrand et al. (2005). The dynamic surface types, sea ice and fresh snow, override the GLCC selection, and are obtained from the European Centre for Medium-Range Weather Forecasts (ECMWF) high-resolution forecasts contained in the X-MET product (Eisinger et al., 2024).

160 In clear-sky conditions, without cloud parallax and at sea-level locations, the default surface co-registration is used. However, if a new co-registration is needed (i.e., surface elevation is greater than 0 and/or a cloud is observed in the nadir view, see 2.1.5), the ADM scene classes for the BBR oblique observations are reconstructed. When the digital elevation model (DEM) indicates elevations above 0, the scene identified for the nadir view is also used in the BBR oblique views in the new co-registration. To reconstruct the scene classes for the oblique views in cloudy nadir conditions, the cloud properties observed in the nadir view and the surface definition from the displaced oblique views (new fore and aft BBR views intersecting the cloud observed in the nadir view) are used. The cloud mask and cloud top height are derived from the MSI operational retrievals (Hünerbein et al., 2023b).

170 MSI can separate snow from most obscuring clouds, but it does not consistently discern optically thin cirrus clouds from snow. Snow and ice surfaces not only share similar reflectance and brightness temperatures with the overlying clouds but can also be colder than clouds themselves due to inversions in the atmospheric temperature profile (Liu and Key, 2003). Techniques commonly used to detect inversions cannot be employed with the MSI due to the lack of water vapor absorption bands (6.7 and 7.2  $\mu\text{m}$ ) and the carbon dioxide band (13.3  $\mu\text{m}$ ). Instead, the cloud top height derived from the EarthCARE's LiDAR (ATLID) is used to verify the cloud top height retrieved by MSI over cold surfaces. In cases where there is inconsistency between the CTHs provided by ATLID and MSI over cold surfaces, ATLID prevails. The cloud cover and CTH used in the processing are adjusted accordingly.

#### **2.1.5 Determination of the solar flux at reference level**

Once the scene is identified, the anisotropic factors required to convert BBR radiances into fluxes are selected accordingly. Given that cloud parallax can affect this selection—and consequently the flux calculation—the radiance-to-flux procedure is integrated with a process of radiance co-registration.



180 BBR's three telescopes are oriented towards nadir and  $50^\circ$  fore and aft along the ground-track. Projecting the oblique views  
to Earth's surface results in viewing zenith angles around  $55^\circ$ . BBR radiance measurements are typically co-registered at  
surface level by default. In clear-sky conditions, the primary emission or reflection observed is from the surface, making this  
default radiance collocation adequate, provided there is no significant elevation. However, in the presence of clouds, the most  
radiatively significant reflection or emission occurs in a layer between the surface and the top of the cloud, not necessarily at the  
185 surface. Therefore, to ensure accurate flux estimates from the three BBR radiances for each BBR sample, these measurements  
must be co-registered to this vertical layer, known as the reference level altitude. Without this co-registration, the flux estimates  
cannot be accurately compared.

To accurately determine the co-registration for the aft and fore views, the nadir view is used as the reference. Fore and  
aft radiances are matched to the nadir sample at a reference level altitude. Flux measurements for nadir and aft observations  
190 across several orbits of the Advanced Along-Track Scanning Radiometer (AATSR) were calculated using a radiance narrow-  
to-broadband conversion. This approach was employed to identify the optimal reference level height for both SW and LW  
radiances. The findings indicated that using a variable reference level is essential to fully exploit the multi-angular viewing  
capabilities of the BBR in the flux retrieval algorithm. The reference level algorithm was tested in a 3D environment generated  
with a Monte Carlo radiative transfer code, providing satisfactory results as reported in Barker et al. (2024).

195 The 90th percentile of the CTH derived from the MSI brightness temperature (BT), referred to as M-COP (Hünerbein et al.,  
2023a), is a reliable estimator for co-registering the BBR radiances in the thermal regime. In the SW regime, co-registering  
BBR radiances using BT-derived CTH results in higher errors compared to the default surface co-registration. A more effective  
approach for SW radiances involves calculating the fluxes for oblique BBR measurements in 1-kilometer increments, i.e.,  
BBR sampling. These calculations consider the atmospheric path of each radiance that intersects the vertical levels between  
200 the surface and the 90th percentile CTH derived from MSI data.. The reference level,  $H_{sw}$ , is then selected by minimizing the  
flux differences between nadir, fore, and aft fluxes as follows:

$$H_{sw} = \arg_x \min (|F_{sw,fore}^x - F_{sw,aft}^x| + |F_{sw,fore}^x - F_{sw,nad}^x| + |F_{sw,nad}^x - F_{sw,aft}^x|), \quad (3)$$

where  $x$  represents the vertical layers from the surface to the cloud top height, and  $F_{sw,nad}^x$ ,  $F_{sw,fore}^x$ , and  $F_{sw,aft}^x$  are the nadir,  
fore, and aft SW fluxes calculated for each layer. The new co-registration of the oblique views is given by  $d = x \tan(\theta_{obl})$  for  
205 the  $x$  value that minimizes the flux differences.

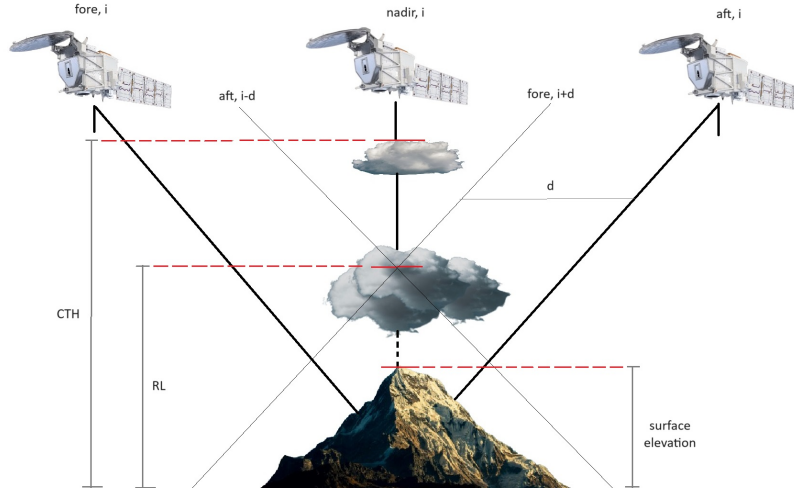
The fore, aft, and nadir SW flux estimates,  $F_{sw,j}^i$ , are calculated from the anisotropic factors,  $R_{sw,j}^i$ , using the following  
equation:

$$\bar{F}_{sw,j}^i(\theta_0) = \frac{\pi L_{sw,j}^i(\theta_0, \phi)}{R_{sw,j}^i(\theta_0, \phi)}, \quad (4)$$

where the index  $i$  represents the nadir, fore, and aft views, and the index  $j$  represents the observed scene.

210 Figure 1 shows an example of co-registration of fore and aft SW measurements for a scene with two cloud layers. The opti-  
cally thin high cloud is semi-transparent to SW radiation, with the main reflection occurring at the core of the lower thick cloud.





**Figure 1.** Diagram illustrating the co-registration applied to SW radiance measurements in a two-layer cloud system. Satellite image provided courtesy of ESA.

While the CTH derived from MSI would set the reference level, RL, at the top of the high cloud, the minimization technique redefines the reference level to the center of the lower thick cloud. The oblique views that minimize the flux differences are displaced by a distance  $d$  from the default surface co-registration. These oblique observations (ideally reflected from the same atmospheric domain) are then co-registered with the nadir observation of sample  $i$ .

### 2.1.6 Combining resulting SW fluxes

To provide the most reliable flux estimate for the target observed by the BBR telescopes during operations, the three fluxes retrieved from the radiances co-registered at  $H_{sw}$  are used in an internal consistency check and are then merged using the uncertainties obtained in the ADM construction.

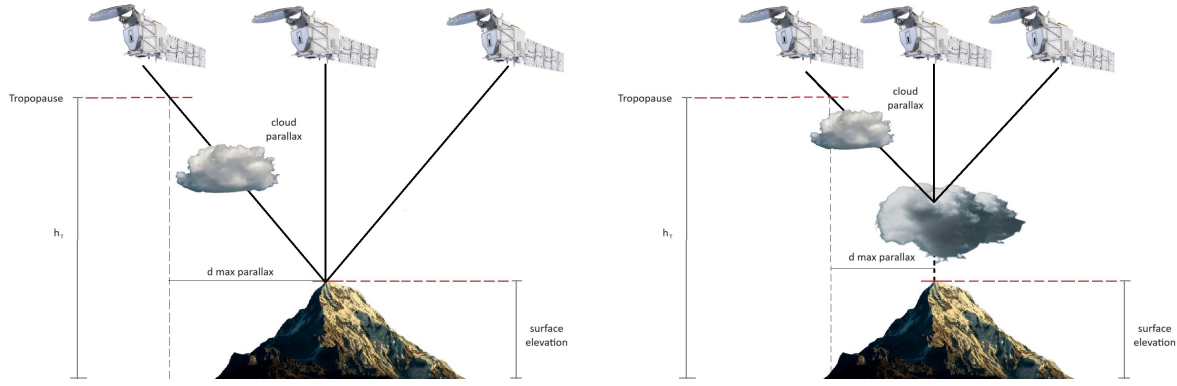
The presence of clouds along the optical path create inconsistencies between the atmospheric domains observed by nadir and oblique views (see Figure 2). Therefore, the cloud presence must be checked before combining the different fluxes. The MSI CTH is employed to detect either clouds in the surroundings of the nadir clear-sky domain or higher clouds those observed in the nadir domain. Given a tropopause height,  $h_T$ , the oblique observation is considered affected by cloud parallax if the CTH of clouds between the coordinates defined by the nadir observation and the coordinates corresponding to the distance  $d_{maxparallax} = h_T \tan(\theta_{obi})$  are located at the height  $h_i = \frac{d_i}{\tan(\theta_{obi})}$ , where  $d_i$  ranges from 1 to  $d_{maxparallax}$ .

In absence of parallax, the discrepancies between the fluxes derived from the nadir, aft, and fore observations are calculated as follows

$$\Delta_{y-z} = 100 \frac{|F_y - F_z|}{\frac{1}{2}(F_y + F_z)}, \quad (5)$$

where  $\Delta_{y-z}$  is the fractional error (0-200%) between the SW fluxes obtained from BBR views  $y$  and  $z$ .





**Figure 2.** Diagram illustrating the presence of cloud parallax in a co-registered fore observation under nadir-observed clear and cloudy sky conditions. Satellite image provided courtesy of ESA.

230 The discrepancies are used to combine the derived fluxes for each view  $i$  as

$$\bar{F}_{sw} = \left[ \sum_{i=1}^3 \frac{\delta^i}{\varepsilon_{F_{sw}}^i \pi \varepsilon_{L_{sw}}^i} \right] \left[ \sum_{i=1}^3 \frac{\delta^i F_{sw}^i}{\varepsilon_{F_{sw}}^i \pi \varepsilon_{L_{sw}}^i} \right], \quad (6)$$

with the weights defined by the uncertainties.  $\varepsilon_{F_{sw}}^i$  are the flux uncertainties arising from the ADM, and  $\varepsilon_{L_{sw}}^i$  the unfiltered radiance errors provided by the BM-RAD processor (Velázquez Blázquez et al., 2023).

When all  $F_{sw}^i$  show discrepancies  $< 10\%$ ,  $\delta^i = 1$  for all  $i$ . When two  $F_{sw}^i$  agree to within  $\pm 10\%$  agree,  $\delta^i = 1$  with the outlier getting  $\delta^i = 0$ . If all  $F_{sw}^i$  show fractional errors  $> 10\%$ , only the lowest  $\varepsilon_{F_{sw}}^i \pi \varepsilon_{L_{sw}}^i$  uses  $\delta^i = 1$ .

## 2.2 BMA-FLX LW algorithm

### 2.2.1 LW radiance-to-flux description

The methodology proposed is based on Stubenrauch et al. (1993), Domenech et al. (2011b) and on the operational GERB LW flux estimation (Clerbaux et al., 2003a, b). The algorithm estimates the anisotropy as a function of the MSI thermal channels information through polynomial second order regressions on the MSI channels brightness temperatures.

Even though in theory a multi-spectral model should be able to correctly handle all scene types, in practice it was demonstrated from previous GERB studies that a large bias was introduced in the case of semi-transparent clouds, fact that was indeed verified in the early stages of the selection and validation of the algorithm. To overcome this problem, it was proposed to use a set of different regressions, instead of one, in line with the approach adopted by the EUMETSAT Central application facility (CAF) for the SEVIRI Outgoing Longwave Radiation (OLR) estimation (EUMETSAT, 2010). Different regressions, and consequently different anisotropy models, have been developed every 5 degrees in VZA and in bins of  $20 \text{ W m}^{-2} \text{ sr}^{-1}$ .

To reduce the instability due to the collinearity between the MSI radiances, the model is constructed using as predictors the brightness temperatures in the  $10.8 \mu\text{m}$  channel and the differences between brightness temperatures ( $T_b$ ) in the  $10.8 \mu\text{m}$  and



12.0  $\mu\text{m}$  channels. Therefore, the predictor variables have been defined as  $z_1 = T_b$  (10.8  $\mu\text{m}$ ) and  $z_2 = T_b(12.0 \mu\text{m}) - T_b$  (10.8  
250  $\mu\text{m}$ ) and  $a_i$  are the coefficients of the regression, dependent on the viewing zenith angle  $\theta$ .

$$R(\theta) = a_0 + a_1 z_1 + a_2 z_2 + a_3 z_1^2 + a_4 z_1 z_2 + a_5 z_2^2. \quad (7)$$

The lower anisotropy in the LW domain, with lower errors expected in the inversion process, allows to obtain anisotropy  
models from radiative transfer (RT) simulations. The theoretical anisotropic factors  $R(\theta)$  have been estimated from theoretical  
simulated thermal radiances and fluxes computed using LibRadtran 1.4 (Mayer and Kylling, 2005) and SBDART (Ricchiazzi  
255 et al., 1998) radiative transfer models, following the standard approach, adapted from Eq. 2.

A detailed description of the radiative transfer databases used in both unfiltering of the BBR and the LW radiance-to-flux  
conversion can be found in Velazquez et al. (2010).

The proposed methodology using level-1 brightness temperatures from the M-RGR product (Eisinger et al., 2024) as pre-  
dictors has the advantage of not using a bin classification for the estimation of the flux, avoiding potential errors due to a  
260 misidentification of the scene.

In summary, the LW flux retrieval algorithm provides instantaneous TOA thermal radiative flux estimates for the BBR  
measurements co-registered at the reference level, converting the broadband radiance measurements (4  $\mu\text{m}$  - 500  $\mu\text{m}$ ) into flux  
estimates. It is assumed that the combination of the off-nadir and nadir BBR fluxes improves the estimation of the flux, and  
the correlation between the broadband radiances and the spectral signature for the radiation field can be exploited to reduce the  
265 LW flux error.

### 2.2.2 Combining resulting LW fluxes

The three BBR thermal unfiltered radiances are co-registered at a reference level defined by the percentile 90th of the MSI  
CTH as described in 2.1.5. The only difference between the SW and LW domains regarding the co-registration of the views is  
the altitude and definition of the reference level for the co-registration of the off-nadir views.

270 In a standard limb-darkening situation, an oblique BBR observation, done at VZA close to 55 degrees, would lead to a highly  
accurate flux estimation (Bodas-Salcedo et al., 2003; Smith et al., 1994). To capitalize on this, the LW merging algorithm would  
assign a greater weight to the fore and aft views for plane-parallel scenes. However, for non plane-parallel scenes, the average  
of the three views would maximize the sampling of the radiance field and statistically reducing the error on the flux estimation.

The approach followed in the LW regime consists in using the weighted average of the LW fluxes obtained from the three  
275 BBR radiances to optimize the use of the nadir view:

$$F = \frac{(1 - \alpha)\pi L_{fore}}{2R_{fore}} + \frac{(1 - \alpha)\pi L_{aft}}{2R_{aft}} + \alpha \frac{\pi L_{nad}}{R_{nad}}, \quad (8)$$

where the anisotropic factors are obtained from the NB observations of the MSI and  $\alpha$  is the weighting factor of the nadir view.  
The latter is important to discriminate the scene types where the LW flux accuracy is limited by the angular sampling. Note  
that as in the SW case, views affected by parallax will not be used in the combined flux calculation.

280 The best value for the  $\alpha$  parameter has been evaluated using CERES along-track data to be very close to  $\alpha=1/3$ . This means  
that in the real world the thermal flux is in average more dependent on 3-dimensional effects than on plane parallel ones.



Therefore the current version of the BMA-FLX thermal flux retrieval algorithm assigns the same weight to each of the 3 views. A methodology to discriminate scenes with standard limb-darkening behaviour is expected to improve this approach and this will be further consolidated using real EarthCARE data during the commissioning.

## 285 2.3 Convolution of MSI radiances and cloud properties

MSI radiances, brightness temperatures, and cloud retrievals, which serve as input variables for flux retrieval models, are averaged over the different BBR resolutions (standard, small, full and assessment). This averaging process must account for the signal response characteristics of the BBR instrument to ensure a consistent comparison between the averaged imager values and the energy measured by the BBR.

290 MSI retrievals are integrated over the BBR sample using the instrument's point-spread function (PSF). Specifically, solar and thermal radiances, cloud mask, and cloud top height (used in the reference level height) at MSI pixel level are convolved within the BBR sample to produce PSF-weighted mean and standard deviation at the BBR resolution. These values are then used as inputs in the BBR SW and LW ADMs.

Let  $x$  represent a general input parameter over the 95% energy of the BBR sample. The weighted average value of  $x$  is  
295 calculated as:

$$\bar{x} = \frac{\sum_i \omega_i x_i}{\sum_i \omega_i}, \quad (9)$$

where the index  $i$  denotes each position within the BBR sample, and the  $\omega_i$  represents the PSF weight at position  $i$ . The denominator in Eq. 9 is the sum of the PSF weights over the sample, which should approximate 95% of the total measured energy.

## 300 3 Evaluation of the BMA-FLX processor

### 3.1 EarthCARE test frames

The EarthCARE mission features a simulator multi-instrument framework (ECSIM) (level-0 to level-2b) to test operational algorithms and validate the entire product processing chain. The BMA-FLX processor is chained, within EarthCARE processing scheme, to M-CLD to receive the MSI-based CTH and cloud mask from M-COP and M-CM, to M-RGR to receive MSI  
305 regrided radiances and brightness temperatures, to A-LAY to receive the ATLID-based cloud top height from A-CTH, to BM-RAD to receive unfiltered radiances, to X-MET to receive high-resolution forecasts, and to ACM-RT (Cole et al., 2023) to provide co-registered radiances and fluxes.

3D atmosphere-surface data is produced by the Global Environmental Multi-scale (GEM) Numerical Weather Prediction (NWP) model for three EarthCARE frames (1/8 of an EarthCARE orbit) (Qu et al., 2022). (Donovan et al., 2023) describes  
310 the radiative-transfer models and instrument models to prepare simulated Level 1 data for testing L2 EarthCARE processor interfaces and evaluating processor performances in these test frames. Level 2 data derived from the EarthCARE forward processors and the RTC radiative fluxes are available for testing BMA-FLX in the test frames. The L1 simulated data and



model truth fields for three GEM-derived scenes are available from <https://doi.org/10.5281/zenodo.7117115> (van Zadelhoff et al., 2023a).

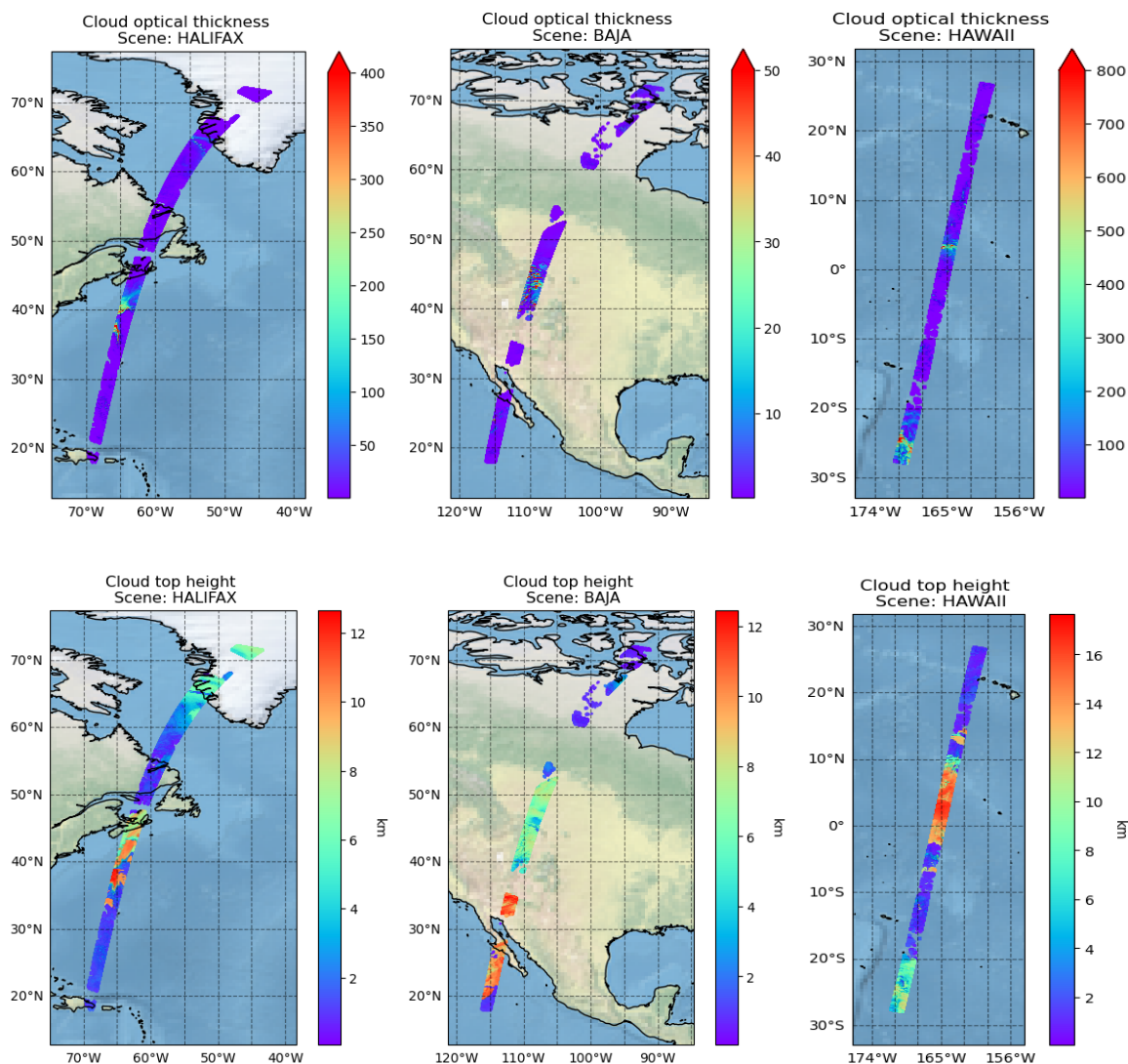
315 The test frames correspond to orbits over Halifax (Canada), Baja California and Hawaii (USA). In the high-latitude portion of the Halifax scene, nighttime mixed-phase clouds transition from deeper formations, featuring supercooled liquid, to mixed-phase clouds. Further south, the clouds become more broken and shallower. Near the center, a storm system with supercooled layers, convective precipitation, and ice clouds is present, followed by cloud-free areas and shallow, low-altitude water clouds. The Baja scene features significant topographical variation compared to the Halifax scene and contains large regions of thinly distributed aerosols. High-level ice clouds are present south of 35° N. Near the center of the scene, above the Rocky Mountains, 320 there are extensive regions of optically thick ice and water clouds. The Hawaii scene is almost entirely over ocean. The nadir track is completely over ocean, with a few of the smaller Hawaiian islands appearing within the MSI track. This scene includes areas of clear sky, upper-level cirrus clouds, and a tropical convective system near the center of the scene.

### 3.2 Processor uncertainty assessment

325 The errors in the L2 retrieval algorithms from ATLID (A-CTH), MSI (M-RGR, M-CM and M-COP) and BBR (BM-RAD) introduce uncertainties that impact both the solar and thermal fluxes retrieved using ADMs in the BMA-FLX processor. Before evaluating the performance of the BMA-FLX algorithm and the accuracy expected in the retrieval, it is necessary to get rid of the uncertainties in the data inputs. The analysis of how the processor deals with non-ideal inputs will be addressed in section 3.3.

330 The inputs for BMA-FLX derived from the EarthCARE forward models are replaced by data directly taken from geospatial simulations, used as inputs for the L1 processors. New inputs are generated using the model truth cloud profiles, thus avoiding the errors introduced by the other algorithms. The starting point was the extinction profiles at 680 nm obtained from the Global Environmental Multi-scale (GEM) Numerical Weather Prediction (NWP) model (Qu et al., 2022). The cloud optical thickness (COT) of the three-dimensional scene is regridded and resampled to the MSI grid. Different COT thresholds were considered 335 to calculate both the cloud mask and the cloud top height, being 0.1 the one with better performance in our analysis of the flux results. Once the MSI-like model truth cloud profile is obtained, we perform a PSF-weighted average of the cloud properties on the BBR footprint, obtaining the "true" cloud information for the three test frames. These new cloud inputs are ingested in the BMA-FLX processor to assess the model performance against the RTC-based flux derived for the test frames. The simulated fluxes are obtained from the RTC employed to produce the L1 EarthCARE radiance data from the GEM results as part of the 340 test frames (Donovan et al., 2023). Figures 3 and 4 show the surface types and cloud properties in the test frames derived from model truth fields at the MSI grid.

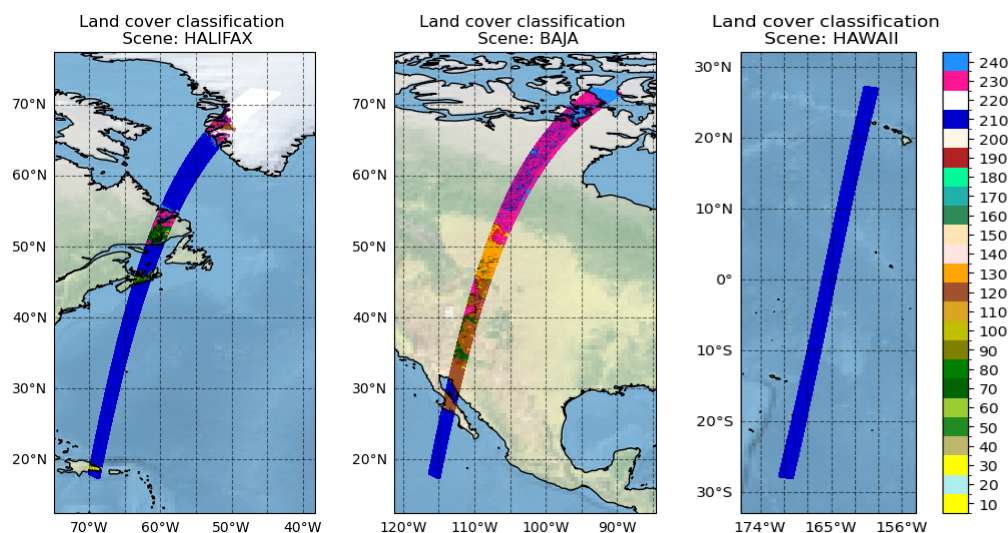
Table 1 presents the results of the radiative flux retrieved by the BMA-FLX processor compared to the model truth flux for two BBR spatial resolutions available in the product. The table includes the RMSE, standard deviation (stddev), and Mean Bias Error (MBE) for the solar and thermal fluxes obtained in the three test frames. Results are further divided into fore, nadir, aft 345 BBR viewing angles, and the combined flux. Across all scenes, BMA-FLX SW flux results for the aft view and the combined approach generally show better agreement with the model truth flux, as indicated by lower RMSE and stddev. The combined



**Figure 3.** MSI-like model truth cloud optical thickness and cloud top height for the three scenes. Colorbars ending with an arrow mean that for cloud optical thickness values greater than the upper tick label have been plotted red. The cloud properties have been calculated using a COT threshold of 0.1.

approach shows the greatest advantage in terms of biases, with lower values compared to the view-based flux estimations. This is particularly relevant for the Hawaii scene with the highest optically thick clouds, where the combined approach has the lowest error metrics, indicating better overall agreement. As expected, the aft and fore views present the lowest differences with respect to the model truth in the LW fluxes. The combined approach tends to have even lower RMSE and stddev, indicating the best overall performance. Biases are consistently negative across all scenes, but are relatively small. The combined view

350



**Figure 4.** Land use land cover (LULC) classification from the ESA Climate Change Initiative (CCI) Land Cover project. Land cover codes are defined in the Land Cover CCI Product User Guide v2 (ESA, 2017). In addition, 230 and 240 correspond to fresh snow and sea-ice, respectively.

consistently provides the best overall agreement with the model truth flux for both solar and thermal fluxes across different scenes.

Figure 5 showcases a comparison between combined SW and LW radiative fluxes derived from the BMA-FLX processor, which has ingested model truth cloud fields, and the RT model truth flux for the BBR standard resolution. The analysis is performed for the different scenes. The time series plots (top row for SW and third row for LW) show the model truth flux (black line) and the BMA-FLX derived flux (red dots) for each scene. The cloud cover from the model truth (ranging from clear to overcast) and surface classification from GLCC employed both in the ADM scene identification (e.g., water bodies, forests, savannahs, etc.) is presented at the bottom of the plots. The inclusion of these classes enhances the understanding of how these factors might influence flux retrievals. These comparisons are complemented by statistical measures (mean, standard deviation, and RMSE) to quantify the agreement between the model and the truth data. Additionally, the plots highlight discrepancies exceeding the  $10 \text{ Wm}^{-2}$  threshold defined by the EarthCARE radiative closure experiment. This comparison is shown as a "pre-launch" numerical assessment experiment for validating the performance and reliability of the BMA-FLX processor in diverse environmental conditions. Figure 5 demonstrates a generally good agreement between the retrieved fluxes and the model truth fluxes. Deep convective clouds, along with optically thick ice and water clouds, are well-represented in both LW and SW retrieved fluxes. However, the greatest discrepancies are observed in cloud-free and partially covered regions. Broken clouds pose significant challenges for both SW and LW retrieval algorithms due to the complex interaction of reflected and emitted radiation. Comparisons over clear-sky scenes are constrained by inherent differences between the surface definitions used in the RTC model derived from the ESA Climate Change Initiative Land Cover and the BMA-FLX surface definition





| Assessment Domain resolution |         |       |      |      |      |       |       |      |        |       |      |      |
|------------------------------|---------|-------|------|------|------|-------|-------|------|--------|-------|------|------|
| Scene                        | Halifax |       |      |      | Baja |       |       |      | Hawaii |       |      |      |
| SW                           | fore    | nadir | aft  | comb | fore | nadir | aft   | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 23.1    | 12.5  | 6.8  | 7.2  | 36.6 | 39.5  | 17.8  | 18.1 | 18.7   | 16.3  | 17.4 | 12.3 |
| stddev ( $Wm^{-2}$ )         | 19.6    | 11.3  | 6.4  | 6.8  | 36.2 | 38.7  | 14.5  | 17.1 | 16.6   | 15.3  | 15.1 | 12.0 |
| MBE ( $Wm^{-2}$ )            | 12.2    | -5.2  | -2.3 | -2.3 | -5.5 | 7.5   | -10.4 | -5.9 | -8.7   | -5.8  | 8.6  | -2.6 |
| LW                           | fore    | nadir | aft  | comb | fore | nadir | aft   | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 2.6     | 4.1   | 2.1  | 2.2  | 4.1  | 6.2   | 3.7   | 3.2  | 4.1    | 5.6   | 4.1  | 4.1  |
| stddev ( $Wm^{-2}$ )         | 2.4     | 3.3   | 2.0  | 1.8  | 3.9  | 6.1   | 3.6   | 3.0  | 4.0    | 5.2   | 3.9  | 3.7  |
| MBE ( $Wm^{-2}$ )            | -0.8    | -2.4  | -0.7 | -1.3 | -1.1 | -1.1  | -1.0  | -1.1 | -1.0   | -2.2  | -1.2 | -1.6 |
| Standard resolution          |         |       |      |      |      |       |       |      |        |       |      |      |
| Scene                        | Halifax |       |      |      | Baja |       |       |      | Hawaii |       |      |      |
| SW                           | fore    | nadir | aft  | comb | fore | nadir | aft   | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 23.2    | 12.3  | 6.2  | 6.7  | 39.0 | 40.0  | 17.1  | 17.2 | 16.9   | 14.8  | 15.2 | 9.3  |
| stddev ( $Wm^{-2}$ )         | 20.0    | 11.1  | 5.7  | 6.3  | 38.5 | 39.5  | 13.7  | 16.2 | 14.2   | 13.4  | 12.9 | 8.8  |
| MBE ( $Wm^{-2}$ )            | 11.7    | -5.4  | -2.4 | -2.5 | -6.5 | 6.1   | -10.2 | -5.8 | -9.2   | -6.2  | 8.1  | -3.1 |
| LW                           | fore    | nadir | aft  | comb | fore | nadir | aft   | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 2.9     | 4.1   | 2.1  | 2.3  | 4.6  | 6.7   | 4.2   | 3.9  | 5.6    | 7.0   | 5.5  | 5.6  |
| stddev ( $Wm^{-2}$ )         | 2.8     | 3.4   | 2.0  | 1.9  | 4.4  | 6.6   | 4.1   | 3.7  | 5.5    | 6.6   | 5.4  | 5.3  |
| MBE ( $Wm^{-2}$ )            | -0.7    | -2.4  | -0.7 | -1.3 | -1.1 | -1.2  | -1.1  | -1.2 | -1.0   | -2.2  | -1.2 | -1.6 |

**Table 1.** Statistics of the BBR flux estimation compared to the model truth flux using the model truth cloud fields as inputs for the BMA-FLX processor in the three test frames (Halifax, Baja, Hawaii) and each of the BBR views (fore, nadir, aft) and the combined approach for the Assessment Resolution (5 JSG x 21 JSG pixels) and the Standard Resolution (10 x 10 km<sup>2</sup>). A COT threshold of 0.1 has been used to derive the cloud properties.

370 based on GLCC. This issue is particularly pronounced for land surfaces covered by fresh snow and ocean sea ice, which show the highest discrepancies. The discrepancies arise from the differing attributions of fresh snow and sea ice between the X-MET and RTC simulations (Figure 4). RMSE values are generally within the 10  $Wm^{-2}$  threshold defined by the EarthCARE radiative closure, except for the Baja region. In regions above 50° N, the simulations indicate fresh snow cover, which is not reflected in the X-MET data. When the analysis is limited to latitudes below 50° N, the RMSE results fall below 10  $Wm^{-2}$ .

### 375 3.3 End-to-end uncertainty assessment

The end-to-end verification of the BMA-FLX processor was evaluated using three test frames created for EarthCARE. In this analysis, the BMA-FLX processor ingests the products retrieved by the L2 EarthCARE processors. A detailed description of





**Figure 5.** Comparison between BMA-FLX flux retrievals incorporating model truth cloud fields and the corresponding model truth fluxes, across the three scenes. The plots illustrate the combined TOA solar and thermal fluxes retrieved by the BMA-FLX processor and their alignment to model truth fluxes. Each time series includes the cloud cover and surface classification information. The cloud properties have been calculated using a COT threshold of 0.1.



the L2 products derived from the L1 data (Donovan et al., 2023), produced for these scenes, is available in the respective papers of this special issue.

380 The results of the comparison between the BMA-FLX and the model truth fluxes are presented in Table 2 for the Assessment Domain resolution (5 x 21 JSG pixels) and the Standard resolution (10x10 km<sup>2</sup>). The LW results remain consistent with the previous section because the LW algorithm leverages broad and narrow-band radiances, independent of L2 cloud retrievals. This independence from MSI retrievals is a key strength of the LW algorithm, ensuring its ADM estimates are not affected by potential errors in other retrieval processes. In SW, the impact of using cloud properties retrieved by M-COP and M-CM  
385 is minimal in the Halifax scene. However, the flux results for the aft and nadir views in the Hawaii scene are significantly higher compared to the analysis using model truth cloud fields (Table 1). The results for Baja do not differ significantly from the previous analysis. It is important to note that this is influenced by the lack of cloud information retrieved by M-CLD above 50° N. Overall, the algorithm for combining the view-based fluxes performs exceptionally well in mitigating the impact of incorrect retrievals from the nadir, aft, and fore models.

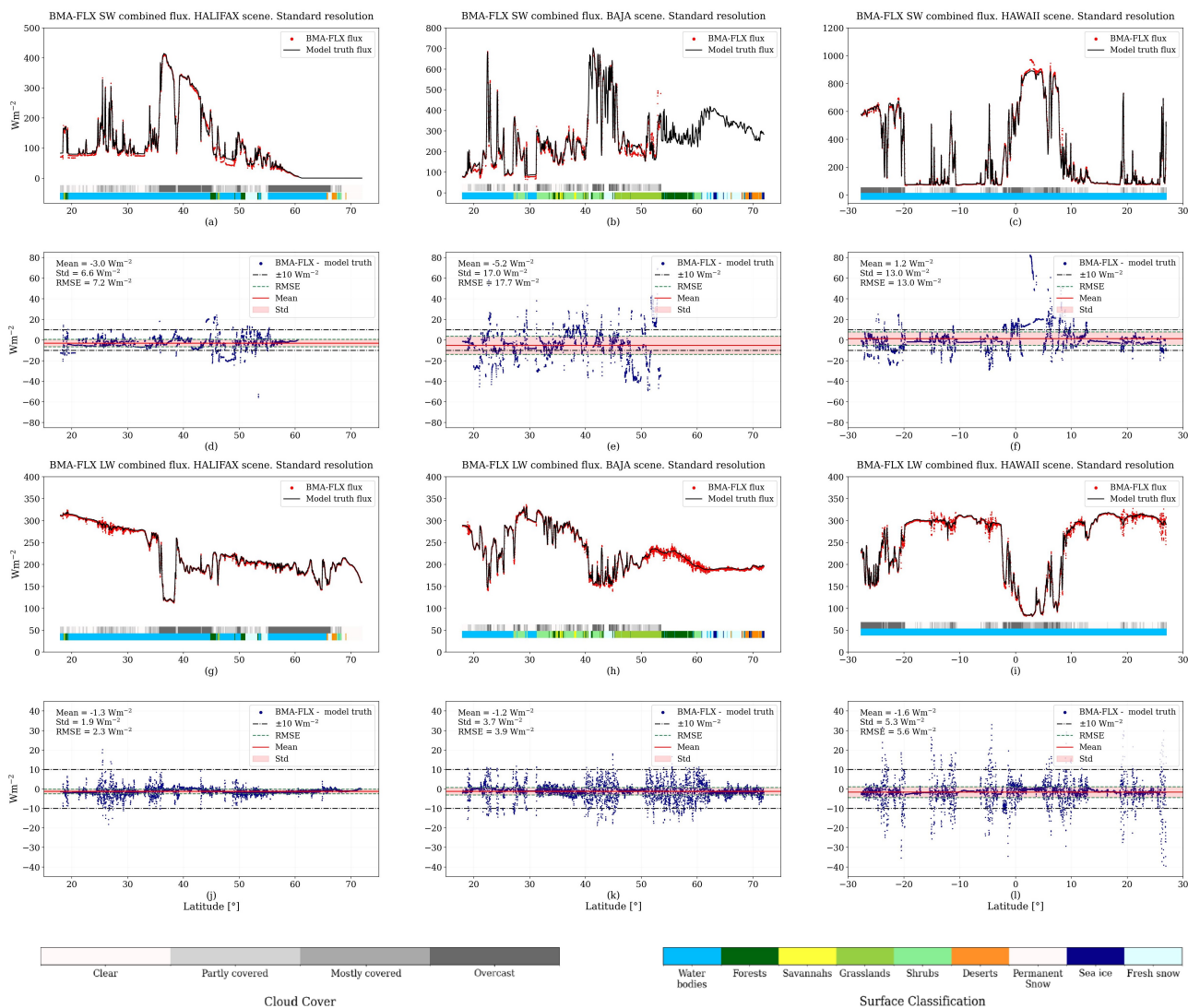
390 Figure 6 shows the combined fluxes retrieved with the BMA-FLX processor alongside the model truth fluxes for the Halifax, Baja, and Hawaii scenes. Unlike Figure 5, the plots in this section present fluxes derived from the L2 EarthCARE operational processors. To facilitate interpretation, both cloud cover and surface classification are included. The results in this figure are significantly worse than those in Figure 5.

The discrepancies stem from differences between the model truth cloud fields and the M-CLD retrieved cloud properties, which affect the selection of the SW ADM and, consequently, the accuracy of the flux estimation. For instance, in Hawaii, the  
395 tropical convective system near the center of the scene (0-5° N) is classified as overcast using the model truth cloud fields, whereas the M-CM cloud mask reports partly or mostly covered scenes, resulting in flux differences exceeding 80 Wm<sup>-2</sup>. Another significant discrepancy is observed in Baja, around 50° N, where there is a considerable difference between the model truth and retrieved clouds. Similarly, in Halifax (48° N), the SW retrieval algorithm classifies the region as partly clear based  
400 on the MSI retrievals, while the model truth indicates clear conditions. The ADM selected for clear-sky reports SW flux values much closer to the model truth than the MSI-based retrieval.

#### 4 Conclusions

This paper describes the algorithm used by the BMA-FLX processor to estimate top-of-atmosphere solar and thermal fluxes using radiance measurements from the BBR and MSI, along with cloud properties from L2 retrievals. Fluxes are independently  
405 estimated for each of the three BBR views, and a combined flux is provided by integrating all three estimates in a common atmospheric layer. For SW flux, an artificial neural network model is trained with ADM-based fluxes from the CERES instrument and MODIS radiances and retrievals. LW flux estimates use ADMs constructed from MSI simulated measurements, employing multiple second-order polynomial regressions.

An end-to-end evaluation of the processor was conducted using three synthetic EarthCARE frames, which included L1 data  
410 from instrument models and L2 data from EarthCARE retrieval models, along with corresponding RTC-derived fluxes. The



**Figure 6.** Comparison between BMA-FLX flux retrievals ingesting EarthCARE products and the corresponding model truth fluxes, across the three scenes. The plots illustrate the combined TOA solar and thermal fluxes retrieved by the BMA-FLX processor and their alignment to model truth fluxes. Each time series includes the cloud cover and surface classification information.



| Assessment Domain Resolution |         |       |      |      |      |       |      |      |        |       |      |      |
|------------------------------|---------|-------|------|------|------|-------|------|------|--------|-------|------|------|
| Scene                        | Halifax |       |      |      | Baja |       |      |      | Hawaii |       |      |      |
| SW                           | fore    | nadir | aft  | comb | fore | nadir | aft  | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 23.4    | 12.2  | 7.3  | 7.9  | 26.3 | 26.7  | 17.6 | 17.5 | 17.9   | 46.8  | 63.4 | 15.4 |
| stddev ( $Wm^{-2}$ )         | 19.1    | 11.0  | 6.8  | 7.3  | 26.2 | 26.6  | 15.1 | 16.6 | 17.6   | 43.9  | 62.1 | 15.3 |
| MBE ( $Wm^{-2}$ )            | 13.4    | -5.2  | -2.7 | -2.9 | -2.0 | 1.7   | -9.1 | -5.5 | -3.0   | -16.0 | 12.6 | 1.7  |
| LW                           | fore    | nadir | aft  | comb | fore | nadir | aft  | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 2.6     | 4.1   | 2.1  | 2.2  | 4.1  | 6.2   | 3.7  | 3.2  | 4.1    | 5.6   | 4.1  | 4.1  |
| stddev ( $Wm^{-2}$ )         | 2.4     | 3.3   | 2.0  | 1.8  | 3.9  | 6.1   | 3.6  | 3.0  | 4.0    | 5.2   | 4.0  | 3.7  |
| MBE ( $Wm^{-2}$ )            | -0.8    | -2.4  | -0.7 | -1.3 | -1.1 | -1.1  | -1.0 | -1.1 | -1.0   | -2.2  | -1.2 | -1.6 |
| Standard Resolution          |         |       |      |      |      |       |      |      |        |       |      |      |
| Scene                        | Halifax |       |      |      | Baja |       |      |      | Hawaii |       |      |      |
| SW                           | fore    | nadir | aft  | comb | fore | nadir | aft  | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 22.9    | 12.0  | 6.7  | 7.2  | 25.6 | 24.1  | 16.5 | 17.7 | 16.1   | 47.0  | 64.7 | 13.0 |
| stddev ( $Wm^{-2}$ )         | 19.0    | 10.7  | 6.0  | 6.6  | 25.5 | 24.1  | 14.0 | 17.0 | 15.7   | 44.0  | 63.3 | 13.0 |
| MBE ( $Wm^{-2}$ )            | 12.8    | -5.5  | -2.8 | -3.0 | -1.1 | 0.9   | -8.7 | -5.2 | -3.4   | -16.6 | 13.3 | 1.2  |
| LW                           | fore    | nadir | aft  | comb | fore | nadir | aft  | comb | fore   | nadir | aft  | comb |
| RMSE ( $Wm^{-2}$ )           | 2.9     | 4.1   | 2.1  | 2.3  | 4.6  | 6.7   | 4.2  | 3.9  | 5.6    | 7.0   | 5.5  | 5.6  |
| stddev ( $Wm^{-2}$ )         | 2.8     | 3.4   | 2.0  | 1.9  | 4.4  | 6.6   | 4.1  | 3.7  | 5.5    | 6.6   | 5.4  | 5.3  |
| MBE ( $Wm^{-2}$ )            | -0.7    | -2.4  | -0.7 | -1.3 | -1.1 | -1.2  | -1.1 | -1.2 | -1.0   | -2.2  | -1.2 | -1.6 |

**Table 2.** Statistics of the BBR flux estimation errors using the outputs of the processor chain for the three test frames (Halifax, Baja, Hawaii) and each of the three views of the BBR (fore, nadir, aft) for the SW and LW fluxes and the combined flux for the Assessment Resolution (5 JSG x 21 JSG pixels) and the Standard Resolution (10x10 km<sup>2</sup>).

assessment also included an evaluation of the processor’s uncertainty, using cloud inputs derived directly from the radiative transfer computations.

Using cloud properties from EarthCARE processors introduces additional uncertainties. Despite this, the combined flux estimates demonstrated stable results. While the retrieved fluxes from each of the BBR views exhibited significant differences when compared to the model truth fluxes, the combined approach significantly reduced these discrepancies. The combined flux consistently provided the best overall agreement with the model truth flux for both solar and thermal fluxes across all scenes. This approach is particularly advantageous in reducing biases and improving accuracy, especially under challenging conditions, such as the super tropical convective systems in the Hawaii scene.

For all three scenes, the LW fluxes showed good agreement with the model truth fluxes, benefiting from reduced anisotropy in the oblique BBR views. The RMSEs of the scenes were generally below 6  $Wm^{-2}$ . In contrast, the SW fluxes showed higher



deviations from the model truth. This is due to the more complex anisotropy of the solar radiation footprints and the greater dependency on the cloud retrieved fields. The RMSEs for the combined fluxes ranged from  $7 \text{ Wm}^{-2}$  for the Halifax scene to  $18 \text{ Wm}^{-2}$  for the Baja scene at the BBR standard resolution.

425 In conclusion, the BMA-FLX processor demonstrates a significant advancement in flux estimation. By combining flux estimates from multiple BBR views, it effectively reduces uncertainties present in individual retrievals. This combined flux approach offers a robust and reliable method for calculating fluxes across various atmospheric conditions. The BMA-FLX processor's ability to handle the complexities of the anisotropic radiance measurements provides accurate top-of-atmosphere flux estimations for both SW and LW radiation. This leads to reliable data for the EarthCARE's radiative closure and brings the Mission closer to achieving its goals.



430 *Data availability.* The EarthCARE L2 demonstration products from the simulated scenes, including the BMA-FLX product discussed in this paper are available from <https://doi.org/10.5281/zenodo.7728948> (van Zadelhoff et al., 2023a)

*Author contributions.* A.V.B. and C.D. conceived the idea for this manuscript. A.V.B. and N.C designed the LW flux retrieval algorithm. A.V.B. N.C and E.B. developed the LW algorithm and the BMA-FLX LW processor. A.V.B. validated the LW algorithm. C.D. developed the theory of the SW algorithm, created and validated the SW algorithm and coded the BMA-FLX SW processor. E.B is responsible for  
435 the pre-processing and PSF weighting for the SW and LW processors. C.D. supervised the projects funding this activity. N.M. helped in the writing and proofreading of the manuscript. C.S.M. prepared graphical material of the manuscript, and contributed to the work done for obtaining the model truth fields. All authors discussed the results and contributed to the final manuscript.

*Competing interests.* The authors declare that they have no conflict of interest.

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