

Luminescence and a New Approach for Detecting Heat Treatment of Geuda Sapphire

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Abstract. For decades, unravelling heat treatment of sapphire has been a challenging issue. The present study offers new aspects that support the detection of heat treatment of sapphire. Natural geuda sapphire exhibits orange to red luminescence under long-wave ultraviolet (LWUV, 365 nm) light, while heated geuda sapphire shows blue luminescence under short-wave ultraviolet (SWUV, 225 nm) light. The presence of melt inclusions in dissolved silks serves as an indicator of sapphire heat treatment. Fourier-transform infrared (FTIR) spectroscopy alone is insufficient for distinguishing unheated from heated sapphire. By combining orange to red luminescence with blue luminescence and melt inclusions, we provide a practical method for accurately differentiating natural gem-quality sapphire and heated gem-quality sapphire.

Keywords: gem; sapphire; heat treatment; luminescence

1 Introduction

Since the 1970s, Sri Lanka has been renowned for its gemstone resources, particularly geuda sapphire, a milky or silky corundum variety that is frequently heat-treated to enhance color and clarity (Ediriweera and Perera, 1989; Perera et al., 1991; Soysa and Fernando, 1992). High-temperature treatment of corundum (including ruby and sapphire) can significantly alter its milkiness, asterism, color, and internal features such as mineral inclusions (Nassau, 1981; Ediriweera and Perera, 1989; Hughes, 1997, 2017; Kyi et al., 1999;

38 Pisutha-Arnond, 2017; Themelis, 2018). Key factors in these transformations include the temperature,
39 duration, and atmospheric conditions of the heating process (Nassau, 1981; Emmett and Douthit, 1993;
40 Peiris, 1993; Emmett et al., 2003; Hughes, 2017; Pisutha-Arnond, 2017; Soonthorntantikul et al., 2019).

41 One of the first-rank challenges encountered by gemologists nowadays is the precise and
42 reliable identification of heat-treated ruby and sapphire. Blue luminescence under SWUV light, observed
43 in heated sapphire for over 50 years (Crowningshield 1966, 1970), can extend into the green region
44 (Nassau, 1981) and has been extensively studied (Evans, 1994; Wong et al., 1995a; Wong et al., 1995b;
45 Hughes 1997; McClure and Smith, 2000; Page et al., 2010; Alombert-Goget et al., 2016a; Alombert-Goget
46 et al., 2016b; Hughes, 2017; Vigier et al., 2021a, b, 2023). This luminescence may relate to rutile inclusions
47 commonly found in natural blue sapphire (Hughes, 2017). During heating at around 1600 °C, rutile
48 decomposes, incorporating Ti^{4+} ions into the sapphire structure (Sutthirat et al., 2006). Blue luminescence
49 serves as a key indicator of heat treatment in sapphire (Crowningshield, 1966; McClure and Smith, 2000;
50 Hughes, 2017; Hughes and Perkins, 2019). However, changes in original brown silk inclusions and orange
51 luminescence in natural unheated sapphire, relative to blue luminescence in heated sapphire, have not been
52 thoroughly explored. The present study provides a novel approach by observing the transformations in silk
53 inclusions and luminescence in sapphire before and after heating.

54

55 **2 Materials and methods**

56

57 Natural unheated geuda sapphire (A-type) samples (as described by Vertriest et al., 2019) were separated
58 based on the appearance of silk inclusions into three distinctive groups, i. e., high-density-silk (HS), low-
59 density-silk (LS), and silk-free (SF) specimens (Fig. 1). Using an Enraf-Nonius Kappa single-crystal X-ray
60 diffractometer with a charge-coupled device (CCD) area detector, these samples were oriented (based on
61 10 frames at a crystal detector distance of 35 mm), cut and polished into wafers with surfaces parallel to
62 the *c*-axis. For chemical analysis using an electron probe micro-analyser (EPMA) system, slabs were
63 coated with carbon for conductivity.

64 Heating experiments were conducted using a high-temperature electric furnace, Linn-
65 HT-1800-Vac. Heating was performed under ambient atmospheric conditions without any additional
66 oxygen buffer. Experimental conditions involved the maximum temperature of 1650 °C, which was
67 maintained for 10 hours, prior to natural cooling down in the furnace. A heating rate of 300 °C per hour
68 was set to reach the maximum temperature. To minimize surface contamination, the samples were placed
69 into a highly purified alumina (Al_2O_3) crucible.

70 The refractive index of samples was measured by a gemological refractometer (Krüss,
71 model ER605) with 1.81 refractive index liquid. Specific gravity was determined by a hydrostatic weighing
72 balance, by weighing samples in water (with a drop of dishwashing detergent added to reduce surface
73 tension) and in air.

74 Micro-inclusions in all samples were imaged using an Olympus BX-series microscope
75 equipped with Olympus DP27 digital camera. The camera was operated using the Olympus Stream micro-
76 imaging software. Raman spectra of inclusions were acquired using a confocal Raman spectrometer Horiba

77 Jobin Yvon LabRAM-HR Evolution. Using 473 nm laser excitation (15 mW at the sample) and a 50×/0.50
78 objective lens, a spectral range of 100–1350 cm⁻¹ Raman shift was recorded. Wavenumber calibration was
79 done using the Rayleigh line, resulting in wavenumber accuracy of better than 0.5 cm⁻¹. A spectral
80 resolution of ca. 1.2 cm⁻¹ resulted from 800 mm focal length and an 1800 grooves/mm optical grating in
81 the monochromator system. For more details see Zeug et al. (2018).

82 Chemical compositions of the samples were determined using a JEOL JXA 8100 EPMA.
83 Analytical conditions were set to 15 kV accelerating voltage and a probe current of about 2.5×10^{-8} A with
84 electron beam focussed to <1 μm. Natural mineral and synthetic oxide references were selected suitably
85 for calibration, including fayalite (Fe₂SiO₄) for Fe, wollastonite (CaSiO₃) for Ca, synthetic corundum
86 (Al₂O₃) for Al, synthetic periclase (MgO) for Mg, synthetic quartz (SiO₂) for Si, potassium titanyl
87 phosphate (KTiOPO₄) for K and Ti, synthetic manganosite (MnO) for Mn, synthetic eskolaite (Cr₂O₃) for
88 Cr, synthetic gadolinium gallium garnet (Gd₃Ga₅O₁₂) for Ga, and synthetic lead vanadium germanium
89 oxide for V. Counting times were 600 s peak and 300 s background for all elements. The K-α line was
90 analysed for all elements except for Ga where the L-α line was measured. Analytical crystals were selected
91 appropriately including thallium acid phthalate (TAP) crystal for Si and Al; pentaerythriol (PET) crystal
92 for Ti, Mg, K, and Ca; lithium fluoride (LIF) crystal for V, Cr, Ga, Fe, and Mn. The detection limit
93 (estimated from threefold background noise) is approximated at 0.005 wt% or 50 ppm. Three spots in each
94 sample were analyzed.

95 Polarized optical absorption (also referred to as UV–VIS–NIR) spectra of samples were
96 recorded on double-sided polished crystal slabs in the spectral range of 35000–3500 cm⁻¹, covering the
97 near ultraviolet (UV), the visible (VIS) and the near infrared (NIR) ranges. The measurements were
98 performed in the sample chamber of a Bruker Vertex 80 FTIR spectrometer at 2 mm measuring spot, using
99 a calcite Glan-prism polarizer and appropriate combinations of light sources (Xe or W lamp), beam splitters
100 (CaF₂–VIS/UV or CaF₂–NIR), and detectors (GaP, Si or InGaAs diodes) to cover the desired spectral
101 range. Hence, each full spectrum was combined from three partial spectra: 1) 35000–18000 cm⁻¹ with 40
102 cm⁻¹ spectral resolution and averaged from 256 scans; 2) 18000–9500 cm⁻¹ with 20 cm⁻¹ resolution and
103 256 scans; 3) 9500–3500 cm⁻¹ with 10 cm⁻¹ resolution and 128 scans.

104 Fourier-transform infrared spectra were acquired by means of a Bruker Tensor 27 FTIR
105 spectrometer attached to a Bruker Hyperion microscope in the spectral range from 4000 cm⁻¹ to 1600 cm⁻¹.
106 A glowbar light source, a KBr beamsplitter, and a deuterated L-alanine doped triglycene sulphate
107 (dLATGS; Tensor27) or Hg–Cd–telluride (MCT) detector (Hyperion) were employed. The spectral
108 resolution was 4 cm⁻¹, sample and reference spectra were averaged from 128 scans.

109 Luminescence phenomena were observed and photo-captured both before and after heat
110 treatment. The images were obtained under LWUV illumination using ZEISS microscope model stemi 508
111 with 0.63× magnification. The images were captured in a darkened room using CANON digital single lens
112 reflex (DSLR) camera model EOS 80D (24.2 MP resolution), which was mounted on top of the microscope.
113 The Superfire UV (365 nm) mini flashlight model S11–H, 3W (max), DC 3.7 V, was held approx. 15 cm
114 above the samples. The camera settings involved an exposure time of 5 s, an exposure bias of 0 steps, and
115 an ISO speed of 200. The aperture was adjusted to f/0, and the focal length of 0 nm. For SWUV illumination

116 (approximately 225 nm), a DiamondView™ device was used. The parameter settings for DiamondView™
117 were established as follows: Integration duration: 2.83 s; minimum excitation status: Off. Power settings
118 ranged from 50% to 80%, contingent upon the intensity of luminescence. A gain of 13.85 dB is measured.
119 The aperture was set to 80% and the field stop to 67%. Gamma was disabled. Photoluminescence (PL)
120 spectra in the visible and near-infrared ranges were acquired using a confocal Horiba Jobin Yvon
121 LabRAM-HR 800 spectrometer. Spectra were excited using the 325 nm emission of a He–Cd laser (ca. 10
122 mW at the sample surface). The system was calibrated using emission lines of a Kr lamp. The spectral
123 resolution was in the range 0.07 nm (violet) to 0.02 nm (NIR range). All the spectra were acquired at the
124 same position both before and after heating experiments.

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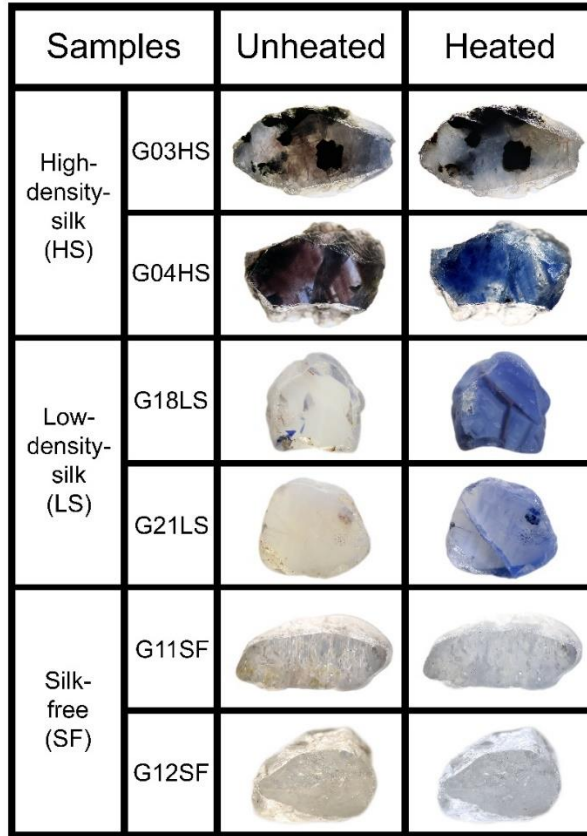
126 **3 Results**

127

128 **3.1 Heating-induced property changes and alteration**

129 Representatives of natural unheated and their heated counterparts of all groups are shown in Fig. 1. All the
130 samples ranged from a specific gravity of 3.83 to 4.08, and refractive indices of 1.760 to 1.770 falling well
131 within the range of corundum properties. Before heating, samples showed varying natural appearances
132 based on inclusion density. Geuda samples with HS inclusions (e.g., G03HS and G04HS, Fig. 1) exhibited
133 brown silk and brown color banding or zoning, while a few samples also displayed a natural blue color.
134 Samples with LS inclusions (e.g., G18LS and G21LS, Fig. 1) generally appeared milky with yellowish or
135 brownish tints. After heating, most samples turned blue, ranging from pale to dark shades, with the milky
136 appearance and yellowish or brownish tints significantly reduced. On the other hand, the SF group usually
137 showed a slightly yellowish appearance (Fig. 1, samples G11SF and G12SF). After the heating experiment,
138 they had changed slightly to a very pale blue color.

139



140

141 **Figure 1.** Representatives of natural unheated geuda sapphire samples within three separate groups, i. e.,
 142 HS (G03HS, G04HS), LS (G18LS, G21LS), and SF (G11SF, G12SF) groups, and their appearances after
 143 heating. Sizes of stones range between 4 and 12 mm.

144

145 3.2 Mineral chemistry

146 Chemical compositions of samples in the three distinct groups are summarized in Tables 1 to 3. The Al_2O_3
 147 contents range between 98.2 and 99.5 wt%. Other elements are found as trace contents only, particularly
 148 Fe, Ti, and Ga. Fe and Ti are essential coloring elements in sapphire. The HS group contained the highest
 149 Fe contents of 0.32–0.36 wt% FeO, together with 0.02–0.04 wt% TiO_2 and <0.7 wt% Ga_2O_3 . The LS group
 150 had a high Ti content of 0.02–0.51 wt% TiO_2 with ≤ 0.06 wt% FeO and ≤ 0.8 wt% Ga_2O_3 . The SF group
 151 contained 0.06–0.26 wt% FeO, ≤ 0.04 wt% TiO_2 and <1 wt% Ga_2O_3 .

152

153 **Table 1.** Representative chemical compositions (EPMA results) and calculated mineral formulae of HS
 154 sapphire samples.

Samples	G01HS	G02HS	G03HS	G04HS
Major oxides (wt%):				
SiO ₂	0.00	0.00	0.45	0.40
TiO ₂	0.02	0.02	0.04	0.03
Al ₂ O ₃	99.0	98.7	98.9	98.7
V ₂ O ₃	0.01	0.00	0.00	0.03
Cr ₂ O ₃	0.00	0.02	0.03	0.00
Ga ₂ O ₃	0.62	0.66	0.00	0.39
FeO _{total} *	0.32	0.33	0.36	0.36
MnO	0.02	0.00	0.00	0.02
MgO	0.00	0.00	0.01	0.01
K ₂ O	0.00	0.00	0.00	0.00
CaO	0.01	0.01	0.02	0.01
Total	100.0	99.8	99.8	100.0
Mineral formulae (apfu)**:				
Si	0.000	0.000	0.008	0.007
Ti	0.000	0.000	0.001	0.000
Al	1.990	1.989	1.985	1.982
V	0.000	0.000	0.000	0.000
Cr	0.000	0.001	0.000	0.000
Ga	0.007	0.007	0.000	0.004
Fe	0.013	0.014	0.005	0.005
Mn	0.000	0.000	0.000	0.000
Mg	0.000	0.000	0.000	0.000
K	0.000	0.000	0.000	0.000
Ca	0.000	0.000	0.000	0.000
Sum	2.010	2.011	1.999	2.000

155 * FeO_{total} = total Fe oxide, assuming all Fe to be ferrous

156 ** Calculated based on 3 O atoms per formula unit

157

158 **Table 2.** Representative chemical compositions (EPMA results) and calculated mineral formulae of LS
 159 sapphire samples.

Samples	G06LS	G16LS	G18LS	G20LS	G21LS
Major oxides (wt%):					
SiO ₂	0.00	0.00	0.11	0.00	0.00
TiO ₂	0.04	0.27	0.37	0.51	0.02
Al ₂ O ₃	99.0	98.5	98.4	98.3	98.6
V ₂ O ₃	0.02	0.03	0.01	0.02	0.01
Cr ₂ O ₃	0.00	0.06	0.00	0.02	0.00
Ga ₂ O ₃	0.10	0.37	0.58	0.81	0.59
FeO _{total} *	0.00	0.06	0.06	0.06	0.05
MnO	0.01	0.00	0.00	0.00	0.01
MgO	0.02	0.02	0.01	0.01	0.02
K ₂ O	0.00	0.00	0.00	0.00	0.00
CaO	0.00	0.00	0.01	0.01	0.00
Total	99.2	99.3	99.6	99.7	99.3
Mineral formulae (apfu)**:					
Si	0.000	0.000	0.002	0.000	0.000
Ti	0.001	0.003	0.005	0.007	0.000
Al	1.997	1.989	1.984	1.981	1.992
V	0.000	0.000	0.000	0.000	0.000
Cr	0.000	0.002	0.000	0.001	0.000
Ga	0.001	0.004	0.006	0.009	0.007
Fe	0.000	0.002	0.002	0.002	0.002
Mn	0.000	0.000	0.000	0.000	0.000
Mg	0.001	0.001	0.001	0.000	0.001
K	0.000	0.000	0.000	0.000	0.000
Ca	0.000	0.000	0.000	0.000	0.000
Sum	2.001	2.003	2.000	2.000	2.002

160 * FeO_{total} = total Fe oxide, assuming all Fe to be ferrous

161 ** Calculated based on 3 O atoms per formula unit

162

163 **Table 3.** Representative chemical compositions (EPMA results) and calculated mineral formulae of SF
 164 sapphire samples.

Samples	G07SF	G11SF	G12SF	G14SF	G22SF	G23SF
Major oxides (wt%):						
SiO ₂	0.01	0.13	0.06	0.00	0.00	0.00
TiO ₂	0.03	0.04	0.03	0.00	0.04	0.01
Al ₂ O ₃	98.7	98.7	98.8	99.5	98.2	98.7
V ₂ O ₃	0.00	0.02	0.00	0.00	0.00	0.00
Cr ₂ O ₃	0.00	0.00	0.01	0.00	0.00	0.00
Ga ₂ O ₃	0.71	0.00	0.00	0.15	0.78	0.94
FeO _{total} *	0.06	0.26	0.08	0.10	0.22	0.13
MnO	0.01	0.02	0.00	0.01	0.02	0.00
MgO	0.00	0.02	0.00	0.00	0.00	0.00
K ₂ O	0.00	0.01	0.00	0.00	0.00	0.00
CaO	0.02	0.00	0.02	0.01	0.01	0.01
Total	99.6	99.2	99.0	99.8	99.3	99.8
Mineral formulae (apfu)**:						
Si	0.000	0.002	0.001	0.000	0.000	0.000
Ti	0.000	0.001	0.000	0.000	0.001	0.000
Al	1.991	1.993	1.997	1.997	1.988	1.988
V	0.000	0.000	0.000	0.000	0.000	0.000
Cr	0.000	0.000	0.001	0.000	0.000	0.000
Ga	0.008	0.000	0.000	0.002	0.009	0.010
Fe	0.002	0.011	0.003	0.004	0.009	0.005
Mn	0.000	0.000	0.000	0.000	0.000	0.000
Mg	0.000	0.001	0.000	0.000	0.000	0.000
K	0.000	0.000	0.000	0.000	0.000	0.000
Ca	0.000	0.000	0.000	0.000	0.000	0.000
Sum	2.002	2.009	2.003	2.003	2.007	2.004

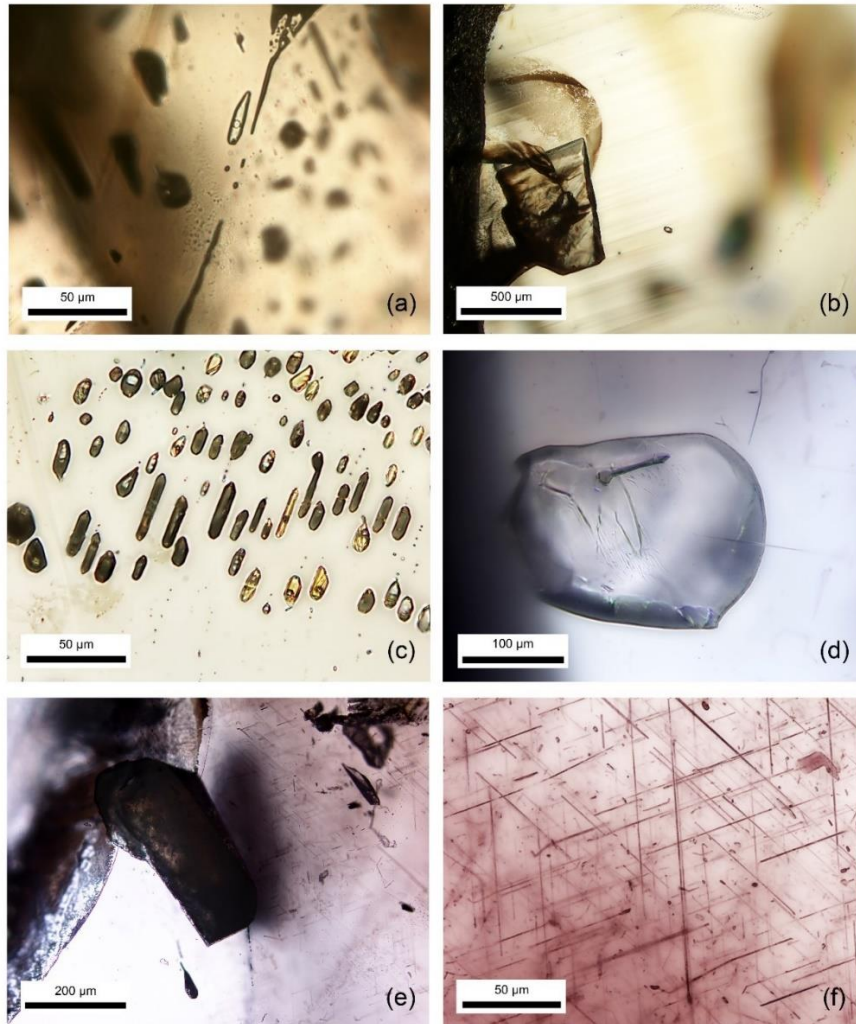
165 * FeO_{total} = total Fe oxide, assuming all Fe to be ferrous

166 ** Calculated based on 3 O atoms per formula unit

167

168 3.3 Microscopic features

169 Negative crystals, with or without CO₂ gas bubbles, were commonly observed alongside mineral inclusions,
 170 such as oligoclase feldspar, calcite, and muscovite, in these sapphire samples. Brown silk inclusions were
 171 prominent in both HS and LS groups, as shown in Fig. 2 (G22SF, G18LS, and G04HS for Fig. 2a, Fig. 2b–
 172 c, and Fig. 2d–f, respectively). Micro-Raman spectroscopy was used to identify CO₂ and mineral
 173 inclusions. Although brown silk inclusions, typically needle-shaped and aligned with color banding (Fig.
 174 3a), were often less than 1 μm in diameter and difficult to identify, irregular or flaky platelet forms (Fig.
 175 3c) were also noted.

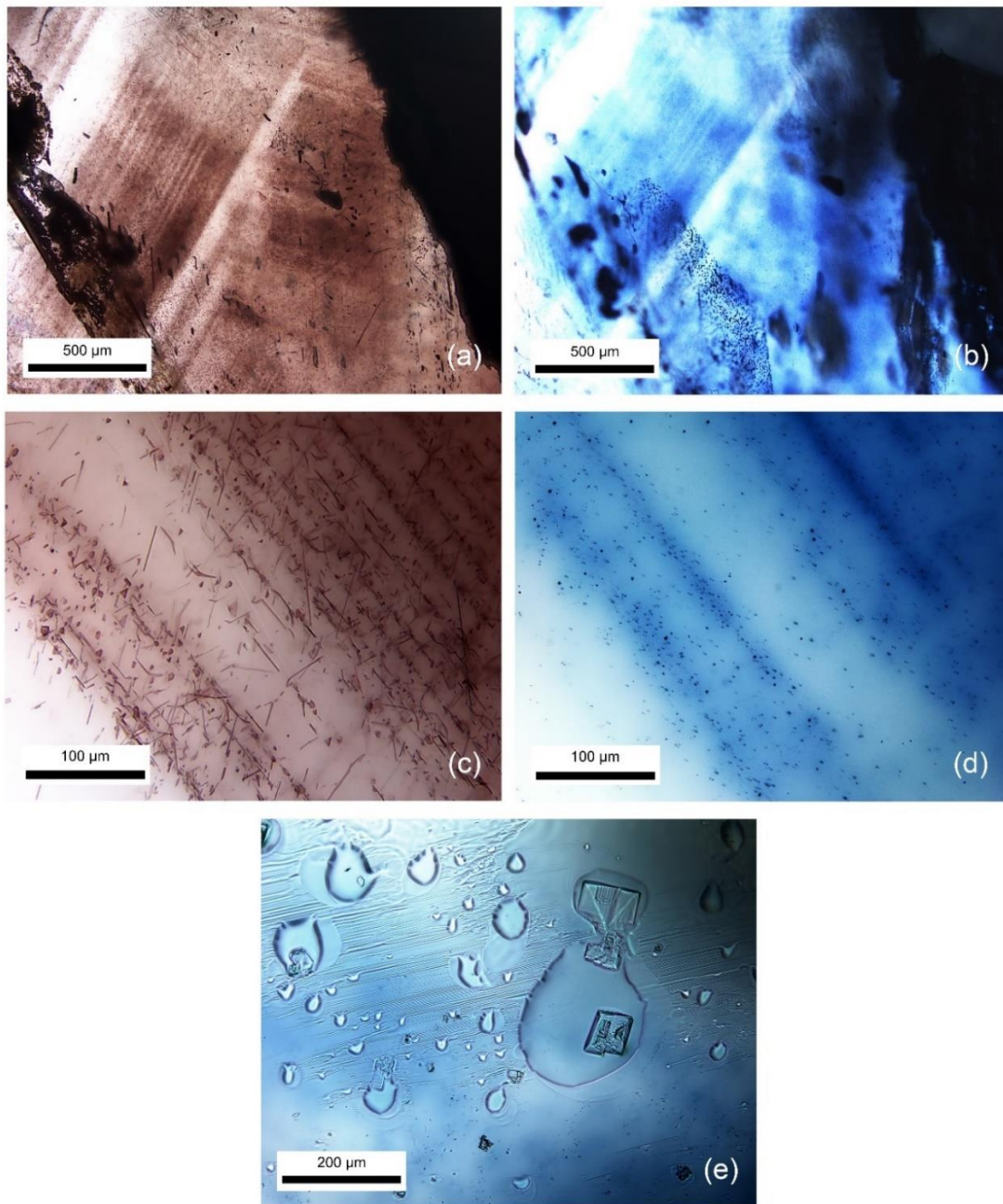


176
 177 **Figure 2.** Transmitted-light photomicrographs of inclusions including CO₂-containing negative crystals
 178 (a), calcite (b), cluster of negative crystals (c), oligoclase (d), muscovite (e), and brown silks (f) in natural
 179 unheated sapphire.

180

181 After high-temperature heating, molten surfaces (Fig. 3e) and decomposed crystal
 182 inclusions were commonly observed in these samples. The most notable alteration was also detected in the
 183 initial area of brown silks (Fig. 3a), which exhibited distinct bluish color banding/zoning (Fig. 3b) after
 184 heating.

185



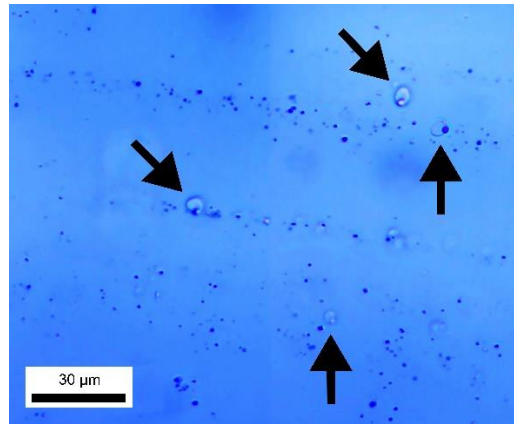
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187 **Figure 3.** Transmitted-light photomicrographs showing that brown banding (a) with irregular platy
 188 brownish flakes and tiny needles (c) in natural unheated sapphire sample G04HS turned into blue
 189 color banding (b) with blue dots (d) upon heating. Melted surface (e) was also observed after heat treatment.
 190

191

The brown silks (Fig. 3c) experienced a transformation upon heating into blue dots (Fig.
 192 3d). Additionally, melt inclusions among blue dots were likely developed by melting of brown silks of the
 193 sapphire host, which have never been reported elsewhere, becoming significantly noticeable and useful for
 194 indicating heat treatment of sapphire (Fig. 4).

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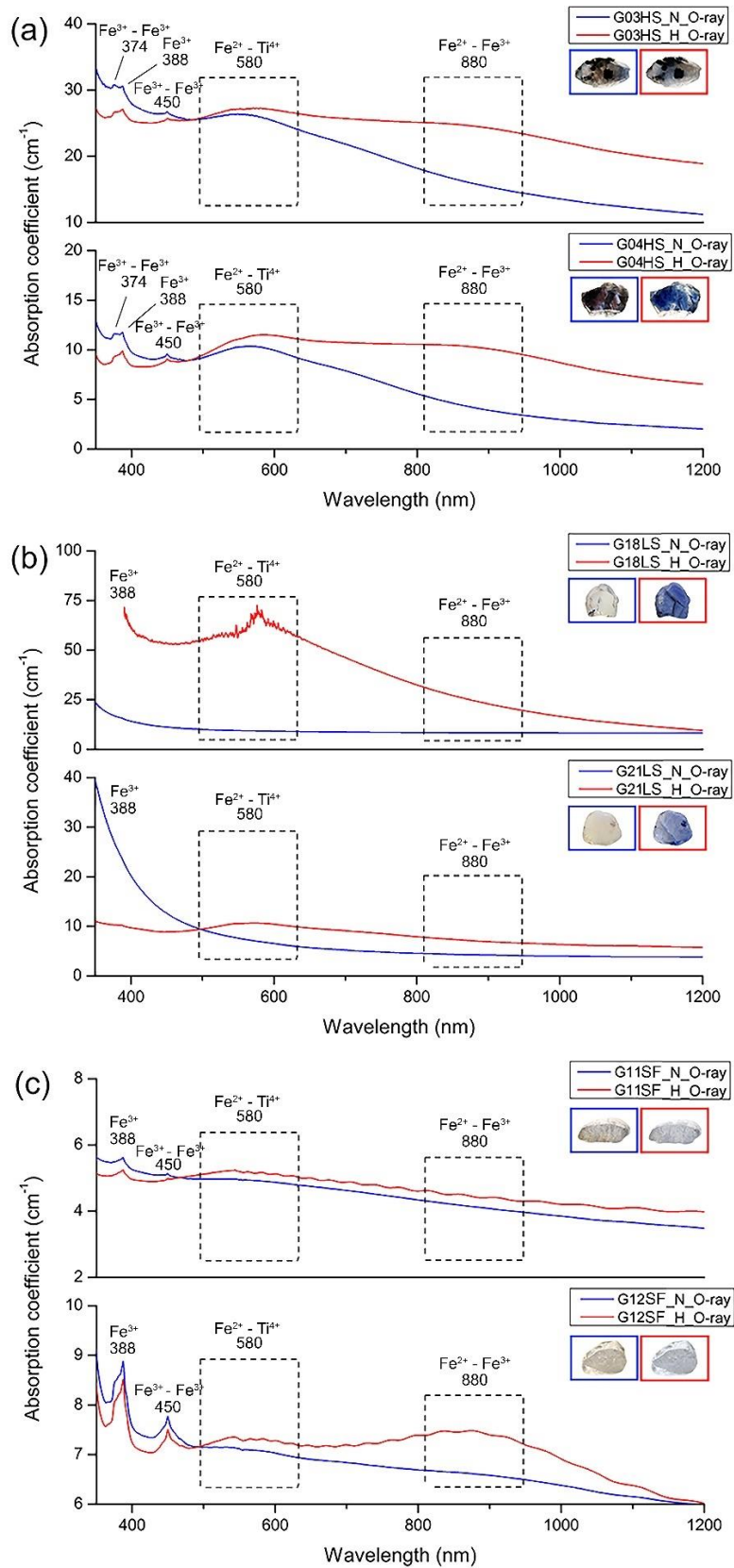
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197 **Figure 4.** Transmitted-light photomicrograph showing melt inclusions (arrows) among blue dots
198 transformed from silk inclusions in sapphire after heating (sample G04HS).

199

200 **3.4 Optical (UV-VIS-NIR) spectroscopy**

201 The optical spectra of representative sapphire samples are presented in Fig. 5. Absorption peaks at 374,
202 388, and 450 nm, as well as (broad) bands around 580 and 880 nm, were observed. Optical spectra have
203 been studied on unheated and heated sapphire by numerous previous researchers (e.g., Ediriweera and
204 Perera, 1989; Perera et al., 1991; Emmett and Douthit, 1993; Hughes, 1997; Kyi et al., 1999; Emmett et al.,
205 2003; Sripoonjan et al., 2014; Hughes, 2017; Pisutha-Arnond, 2017; Themelis, 2018; Palke et al., 2019;
206 Soonthorntantikul et al., 2019; Dubinsky et al., 2020). The 374, 388, and 450 nm peaks as well as the 880
207 nm band were proposed to be attributed to Fe, the 580 nm band to the Fe-Ti pair. After heating, all samples
208 showed a significant increase in the main Fe-Ti pair related absorption band at around 580 nm (Figs. 5a-
209 c), whereas Fe-Fe related absorption at around 880 nm was only increased in some samples (i.e., Figs. 5a
210 and 5c). The intensified absorption of the 580 nm band in these samples is referred to an increase of Fe-Ti
211 pairs after heating which leads to enhanced blue coloration in heated sapphires.



212

213 **Figure 5.** Optical absorption spectra of untreated (blue lines) and heated (red lines) samples: (a) HS group

214 (G03HS, G04HS); (b) LS group (G18LS, G21LS); (c) SF group (G11SF, G12SF). Sizes of stones range

215 between 4 and 12 mm.

216

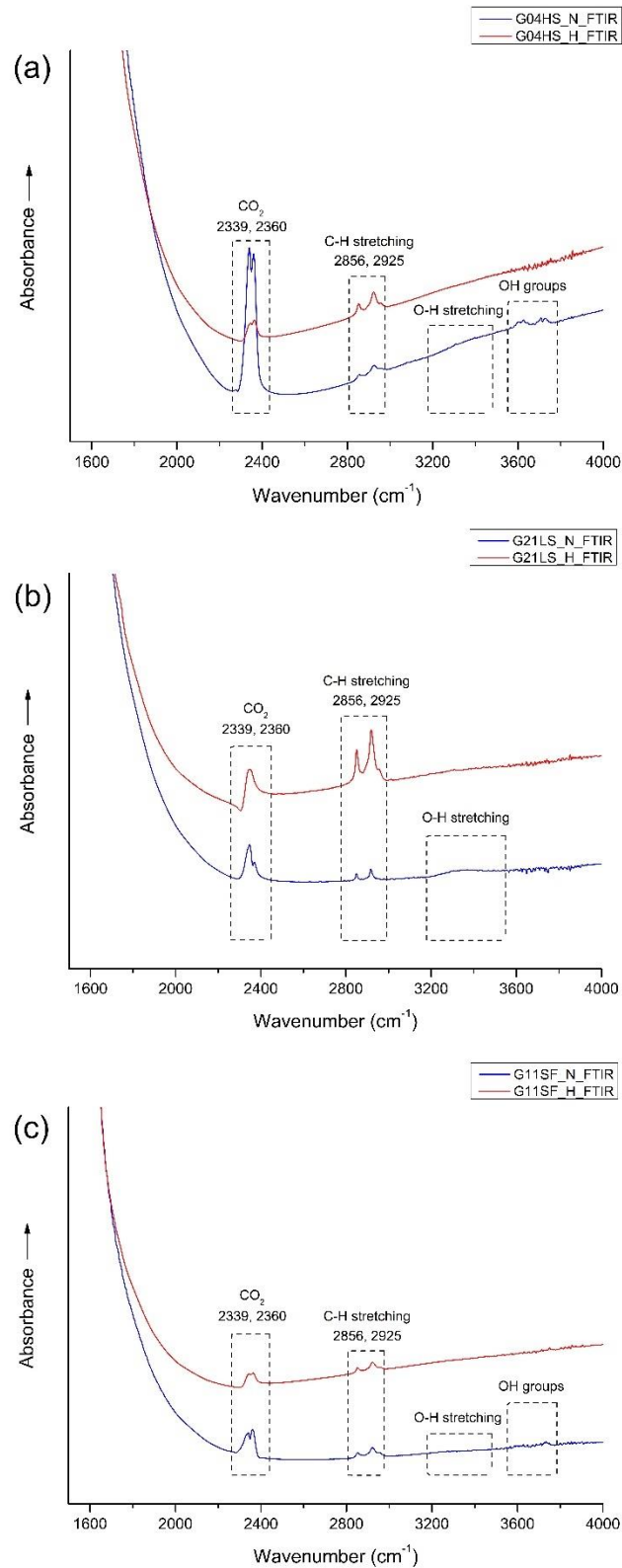
217 Spectral characteristics of corundum containing Fe^{3+} ions exhibit a high degree of
 218 complexity. It is noteworthy that Fe^{3+} has electron configuration d^5 resulting in a crystal field spectrum
 219 with ground state ${}^6\text{A}_1$ (Ferguson and Fielding, 1972). Small peaks at 374 nm (${}^4\text{E}^b$) and 450 nm (${}^4\text{A}_1, {}^4\text{E}^a$)
 220 should be attributed to the enhanced absorption of Fe^{3+} – Fe^{3+} pairs (McClure, 1962; Ferguson and Fielding,
 221 1971; Krebs and Maisch, 1971; Ferguson and Fielding, 1972) as well as a weak broadband absorption at
 222 540 nm (${}^4\text{T}_2$) which could not be seen in this work. The distinct peak observed at a wavelength of 388 nm
 223 (${}^4\text{T}_2^b$) (Krebs and Maisch, 1971) is linked to the individual Fe^{3+} ions. This, however, does not rule out the
 224 possibility of a higher-order cluster with extra ions or other point defects (Emmett et al., 2003).
 225 Additionally, there is also a broad band at a wavelength of 330 nm (${}^4\text{T}_1^b$) which is interpreted as a Fe^{3+} –
 226 Fe^{3+} pair absorption (Ferguson and Fielding, 1972). As they are in low concentration (traces), this is also
 227 present in the spectra of heated samples G03HS and G04HS, as well as in all spectra of sample G12SF in
 228 this study. In trace contents both Fe^{2+} (d^6) and Ti^{4+} (d^0) ions alone do not exhibit any absorption in corundum
 229 in the visible range (Townsend, 1968); on the other hand, Fe^{2+} – Ti^{4+} pairs ($t_2 \rightarrow {}^2\text{E}$) (Ferguson and Fielding,
 230 1971) may yield a broad band absorption around 580 nm ($\text{E}\perp\text{c}$), or 700 nm ($\text{E}\parallel\text{c}$) (Dubinsky et al., 2020).
 231 The Fe^{2+} – Fe^{3+} pair gives rise to the broad absorption band at ca. 880 nm (Fig. 5; Ferguson and Fielding,
 232 1972).

233

234 3.5 FTIR spectroscopy

235 FTIR spectra of most samples yielded identical patterns within the range of 1600–4000 cm^{-1} (Fig. 6). They
 236 usually showed CO_2 peaks (at 2339 and 2360 cm^{-1}), as well as C–H stretching related peaks (at 2856 and
 237 2925 cm^{-1}), likely from artefacts (Fig. 6, blue lines), in accordance with Hughes (2017) and
 238 Soonthorntantikul et al. (2021). However, O–H stretching of boehmite/diaspore peaks (at 1975 and 2105
 239 cm^{-1}) (Delattre et al., 2012; Sun et al., 2015; Choi et al., 2017; Filatova et al., 2021; Soonthorntantikul et
 240 al., 2021) was only observed in sample G03HS (Fig. 7a, blue line). Weak absorption features of O–H
 241 stretching from H_2O (broad band at ca. 3400 cm^{-1}) and OH groups (ca. 3600–3700 cm^{-1}) were only found
 242 in the untreated samples (blue lines), see Fig. 6a.

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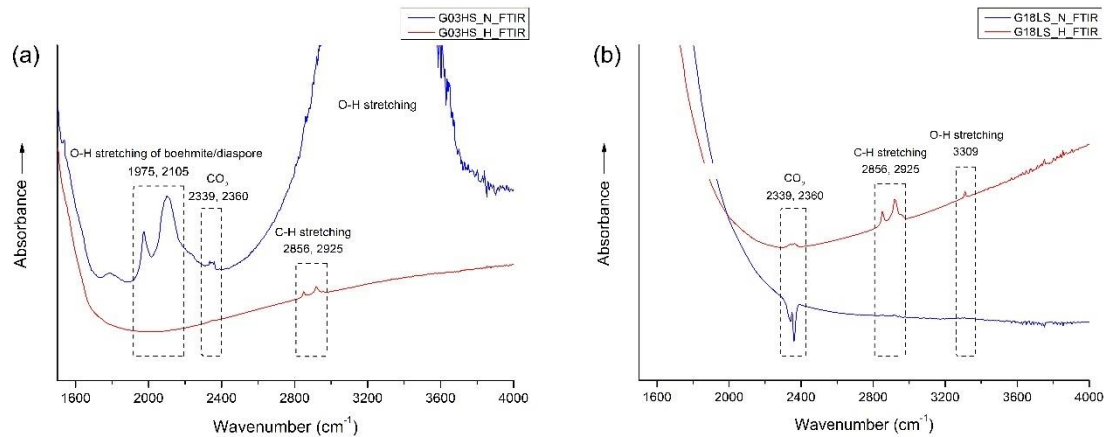


244

245 **Figure 6.** FTIR spectra obtained before (blue lines) and after (red lines) heating experiments of
 246 representative samples G04HS (a), G21LS (b), and G11SF (c), respectively.

247

248 After heating, boehmite/diaspore-related absorption peaks (only observed in sample
 249 G03HS, Fig. 7a) at 1975 and 2105 cm^{-1} disappeared. In contrast, the 3309 cm^{-1} hydroxyl (O–H) absorption,
 250 which was not present in any natural sample before heating, appeared only in sample G18LS after heating
 251 (Fig. 7b, red line).
 252



253
 254 **Figure 7.** FTIR spectra obtained before (blue lines) and after (red lines) heat treatment of samples G03HS
 255 (a) and G18LS (b).
 256

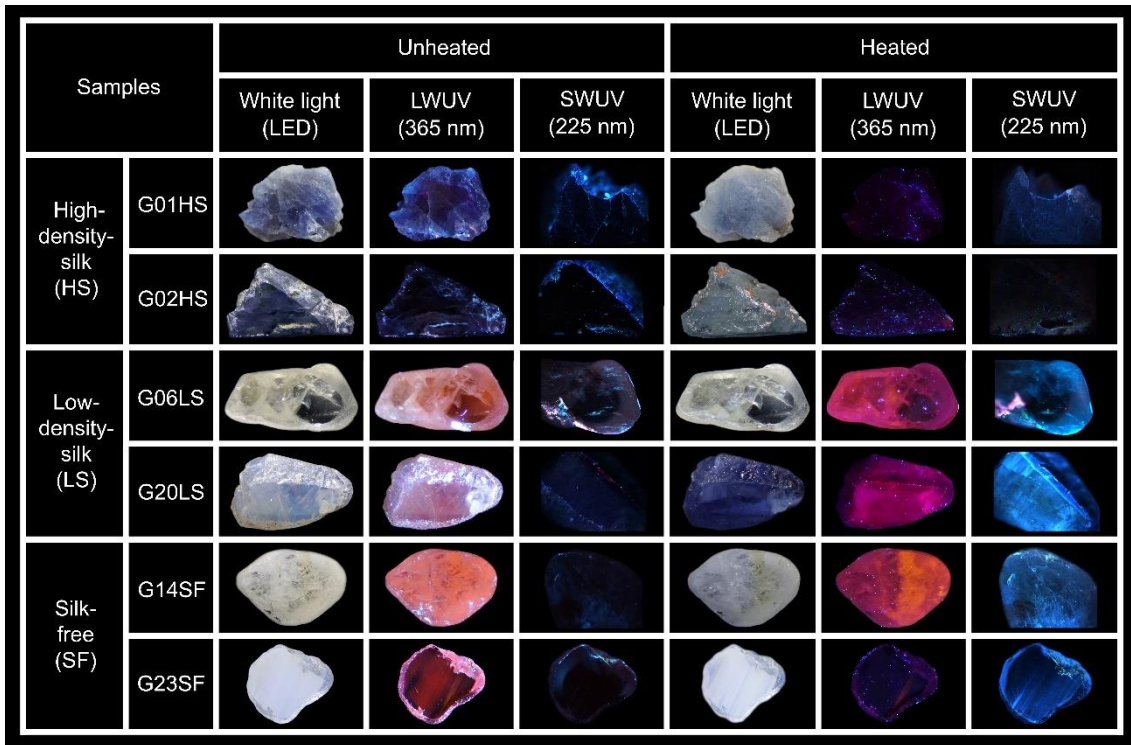
257 In this study, the non-systematic occurrence of O–H absorption in the 3100–3600 cm^{-1}
 258 range in all natural geuda sapphire samples, together with the development of a weak absorption at 3309
 259 cm^{-1} upon heating in only one of the samples (see Fig. 7b, red line), address the limitation to differentiate
 260 unheated and heated sapphire by FTIR spectroscopy. Furthermore, heat treatment employed in this study
 261 did not involve the use of any additional gases, such as hydrogen, to create a reducing atmosphere within
 262 the furnace. Despite this, the 3309 cm^{-1} absorption band was seen after the heating process. This might be
 263 in accordance with an explanation proposed earlier by Notari et al. (2018).

264 The controversy of the presence of an O–H peak in the FTIR spectrum in unheated and
 265 heated sapphire could be attributed to an inherent hydrogen content of the corundum. Hydrogen was found
 266 in corundum, primarily in the form of alumina hydrates (Notari et al., 2018). These hydrates could release
 267 hydrogen through de-hydroxylation at temperatures as low as approx. approx. 450 °C. Additionally,
 268 hydrogen was present in the air as H₂O, which can be split at temperatures around 900 °C to produce
 269 hydrogen gas (H₂) and oxygen gas (O₂) through the reaction $2\text{H}_2\text{O} \rightarrow 2\text{H}_2 + \text{O}_2$ (Notari et al., 2018).
 270

271 3.6 Photoluminescence imaging and spectroscopy

272 Photos presenting luminescence of some samples both before and after heat treatment are shown in Fig. 8.
 273 Before heating, all natural sapphire samples were inert to SWUV light; moreover, all LS and SF samples
 274 exhibited orange to red luminescence under LWUV light (Fig. 8). After heating, all LS and SF samples,
 275 exhibited intense blue luminescence under SWUV light whereas their initial orange to red luminescence
 276 under LWUV light turned into a strong purplish red luminescence (Fig. 8, samples G06LS and G20LS in
 277 particular). In summary, the HS samples were all inert to SWUV and LWUV light both before and after

278 heating. Notably under LWUV light, an initial orange to red luminescence of a few samples from the SF
 279 group was drastically reduced after heating (e.g., G23SF in Fig.8).
 280



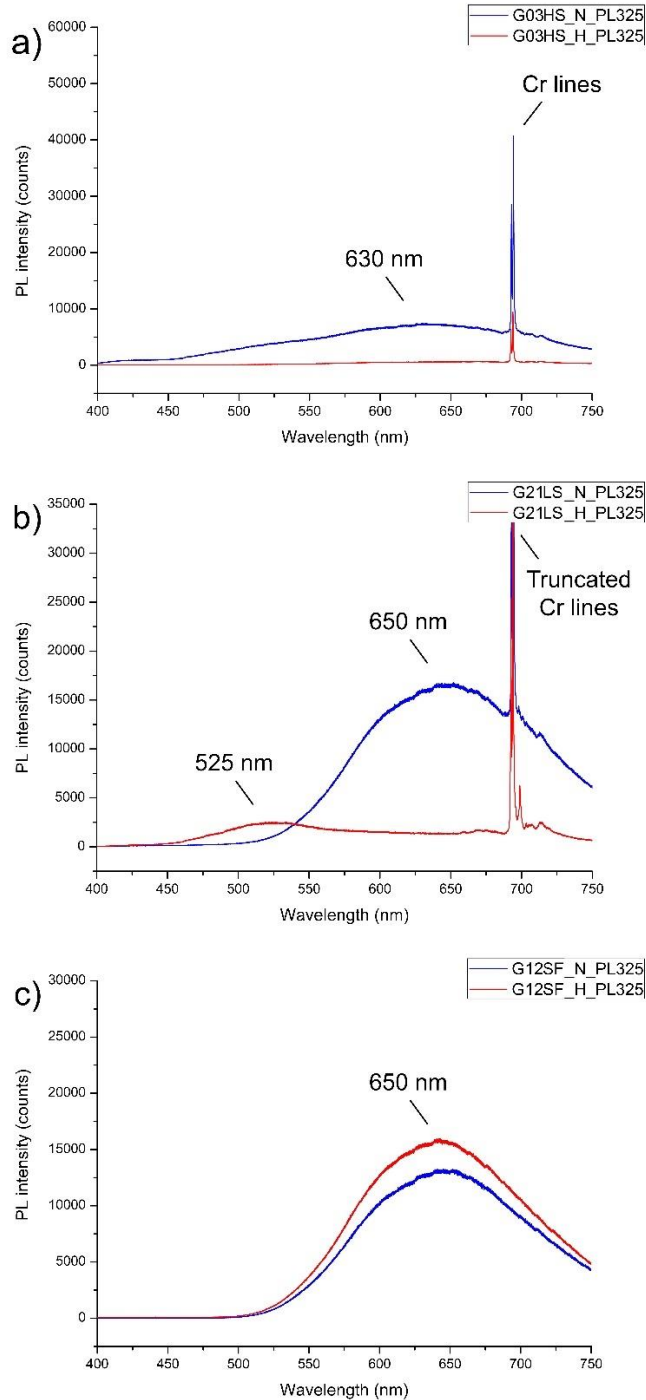
281
 282 **Figure 8.** Representative images of HS (G01HS, G02HS), LS (G06LS, G20LS), and SF (G14SF, G23SF)
 283 groups obtained under LWUV and SWUV illumination before and after heating. Sizes of stones range
 284 between 4 mm and 12 mm.

285
 286 The UV-excited PL spectra showed that all the unheated and heated sapphire samples
 287 have an identical feature of two narrow peaks of trace Cr^{3+} lines at around 692.8 and 694.2 nm (Fig. 9) that
 288 are assigned to the spin-forbidden ${}^2E \rightarrow {}^4A_2$ relaxation of trace Cr^{3+} (Nelson and Sturge, 1965). However,
 289 the Cr^{3+} lines of some samples (Fig. 9c) are too weak to be visible within the noise of a broad and strong
 290 emission band. All unheated sapphire samples showed a similar emission band in the orange to red region
 291 centered around 630–650 nm (Fig. 9a–c, blue line). Remarkably, this appears to be associated with orange
 292 to red luminescence under LWUV light, as noted by (Segura, 2013; Vigier et al., 2021a, b, c; Vigier and
 293 Fritsch, 2022). Despite having the emission band around 630–650 nm, only unheated sapphire from HS
 294 group appeared inert under LWUV illumination while the others revealed orange to red luminescence.

295 After heating, significant alteration in the emission band was observed, as depicted by
 296 the red lines in Fig. 9a–c. The PL spectra of sample G03HS exhibited a notable reduction in the emission
 297 band through the visible region (Fig. 9a, red line). This went along with a lack of luminescence both under
 298 SWUV and LWUV excitation, whereas sample G12SF demonstrated a slight increase of the emission band
 299 in the orange to red region (Fig. 9c, red line). More details are given in the discussion part below.

300 In contrast to the other groups, after heating, sample G21LS (Fig. 9b, red line) exhibited
 301 a significant emission band in the green region at around 525 nm. Note that this broad emission is excited
 302 with the 325 nm laser (Fig. 9) but does not seem to affect significantly the emission colors observed under

303 SWUV (225 nm) and LWUV (365 nm) excitation (Fig. 8). For a discussion of the possibly strong
 304 dependence of emission intensity (and color) on the excitation wavelength see for instance Zeug et al.
 305 (2022). Likewise heated sapphire has been proposed to have an emission band in the blue region, which
 306 corresponds to blue luminescence under SWUV light (Nassau, 1981; Hughes, 2017; Vigier et al., 2023).
 307



308
 309 **Figure 9.** Representative PL (UV-excited) spectra obtained before (blue lines) and after (red lines) heating
 310 of sample G03HS (a), sample G21LS (b), and sample G12SF (c).
 311

312 4 Discussion

313

314 4.1 Generalities

315 After Ti^{4+} ions are being exposed to SWUV light, they yield luminescence (Nasdala and Fritsch, 2024).
 316 However, blue luminescence was not observed in both unheated and heated basaltic sapphire, possibly due
 317 to the abundant presence of Fe^{2+} of basaltic origin that may strongly quench such blue luminescence
 318 (Soonthorntantikul et al., 2019). More details will be discussed in this report. Furthermore, even though
 319 microscopic inclusions have been the distinguishing characteristics of heated sapphire, identifying heat-
 320 treated sapphire remains challenging (Crowningshield, 1966; Hughes, 2017). FTIR spectroscopy has also
 321 been applied to detect heated sapphire. In some cases, the presence or absence of specific FTIR features in
 322 the O–H absorption region ($3100\text{--}3600\text{ cm}^{-1}$) may serve as an indicator of heat treatment (Smith, 1995;
 323 Beran and Rossman, 2006; Saeseaw et al., 2018); however, it is probably not a conclusive evidence
 324 (Ediriweera and Perera, 1989; Perera, 1993; Sutthirat et al., 2006; Cartier, 2009; Jaliya et al., 2020). For
 325 example, the presence of the 3309 cm^{-1} FTIR absorption peak was used as an indicator of heated corundum
 326 (Hughes and Perkins, 2019; Soonthorntantikul et al., 2021). However, recent discoveries show that this
 327 peak may also be found in unheated sapphire, suggesting that it is not a reliable indication of heat treatment
 328 (Hughes, 1997, 2017; Hughes and Perkins, 2019; Soonthorntantikul et al., 2021).

329

330 4.2 Silk inclusions and coloration of sapphire

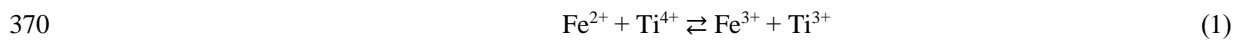
331 Studies addressing brown silk inclusions in corundum are scarce. Soonthorntantikul et al. (2021) reported
 332 a mix of whitish silk and irregular/flaky/platelet-like brownish silk inclusions in corundum from Mogok.
 333 Brown silk was ascribed as presumable ilmenite (FeTiO_3) which is noticeable in high-Fe sapphire, whereas
 334 white silk was suggested to consist of rutile (TiO_2). The brown silks seen in our sapphire samples are likely
 335 ilmenite, which is supported by their irregular/flaky/platelet-like brownish appearance and their high Fe
 336 and Ti contents (note that the highest quantity of Fe was found in the HS group). Ilmenite decomposition
 337 upon heat treatment does result in Fe and Ti migration into the host sapphire and subsequently causes blue
 338 coloration. In particular, the decomposition of brown silks during heat treatment induces the formation of
 339 blue dots, which is a result of the $\text{Fe}^{2+}\text{--Ti}^{4+}$ pairing formation. Upon closer inspection using a high-
 340 resolution microscope, these blue dots reveal distinct micro-inclusions as melt inclusions (with size of ≤ 1
 341 μm , see Fig. 4), which have never been documented before. However, it should be noted that these melt
 342 inclusions are possibly derived from the decomposition of silks.

343 This work focuses only on blue coloration in sapphire which mainly relates to the $\text{Fe}^{2+}\text{--}$
 344 Ti^{4+} pair, as initially noted by Townsend (1968), followed by Mattson and Rossman (1988), Moon and
 345 Phillips (1994), and Emmett et al. (2003). Ti exhibits electron-donor properties, whereas Fe may function
 346 as an electron acceptor. When occupying neighboring Al^{3+} positions, absorption due to intervalence charge
 347 transfer between such donor-acceptor pairs may occur (details reported by Emmett et al., 2003 and
 348 Monarumit et al., 2023).

349 It should also be mentioned that Ti^{4+} ions do not exhibit any absorption characteristics in
 350 the visible spectrum when considered individually. The Ti^{4+} ion has a closed-shell electron configuration,

351 whereas the Fe^{2+} ion mainly absorbs wavelengths within the near infrared and low-energy visible regions.
 352 In contrast, when Fe^{2+} and Ti^{4+} ions are situated on neighboring structural sites, notable absorption bands
 353 develop across the visible and near-infrared spectral regions. These Fe^{2+} - Ti^{4+} pairs exhibit a band center at
 354 around 580 nm (see Fig. 5) when the electric field vector E is perpendicular to the crystallographic c -axis
 355 ($E \perp c$), but a peak at 700 nm is seen when the electric field vector E is parallel to the crystallographic c -axis
 356 ($E \parallel c$) (Dubinsky et al., 2020). Although the theory of the energy levels of an individual transition metal ion
 357 inside a crystal has been extensively explored, the corresponding theory for ion pairs or clusters within a
 358 crystal remains underdeveloped (Dubinsky et al., 2020).

359 In the present study, the natural unheated geuda sapphire samples were placed in
 360 atmospheric conditions and subjected to a maximum temperature of 1650 °C for a duration of 10 h.
 361 According to the examples presented in Fig. 1, samples G03HS and G04HS exhibited a noticeable increase
 362 in blue coloration, particularly around the area of brown silks and brown color banding/zoning, after
 363 heating. On the other hand, the initial blue patch (e.g. samples G01HS and G02HS) became paler blue after
 364 heating, which might be due to the breakage of initial Fe-Ti pairs in those areas. The other groups, which
 365 have yellowish and/or milky appearances, revealed an increase in blue color after heating (samples G12SF
 366 and G18LS, Fig. 1). This blue coloration is attributed to two distinct factors, notably the decomposition of
 367 silk inclusions and a subsequent charge transfer mechanism (Emmett and Douthit, 1993; Hughes, 1997,
 368 2017; Nassau, 1980, 1981; Themelis, 2018). The process of charge transfer (Ferguson and Fielding, 1972;
 369 Nassau, 1981) is described as:



371 It is important to note that the blue color observed in sapphire could also be produced
 372 with the application of heat in oxidizing conditions at high temperature. Heat treatment can be classified as
 373 high- or low-temperature according to the decomposition of rutile silks in corundum (Nassau, 1981;
 374 Emmett and Douthit, 1993; Emmett et al., 2003; Hughes, 2017; Hughes and Perkins, 2019). The term low-
 375 temperature heat treatment has been used (typically referred to as below 1000 °C) when rutile particles still
 376 reveal their original structures. On the other hand, temperatures beyond 1350 °C denote high-temperature
 377 heat treatment when rutile silks start to decompose and dissolve within the corundum host (Hughes, 2017;
 378 Themelis, 2018). Consequently, internal diffusion (indicated by a colored halo surrounding the crystal
 379 inclusion), molten or altered inclusions, and/or broken silk are strong indicators of high-temperature heat
 380 treatment. However, low-temperature heat treatment can also produce various altered mineral inclusions
 381 (Kammerling et al., 1990; McClure and Smith, 2000; McClure et al., 2010; Pisutha-Arnond, 2017;
 382 Soonthorntantikul et al., 2019).

383 In recent studies, the possibility of employing Fe^{2+} - Fe^{3+} charge transfer as an alternate
 384 method for blue coloration has also been mentioned (Nikolskaya et al., 1978; Schmetzer and Kiefert, 1990;
 385 Häger, 1992, 2001; Sripoonjan et al., 2014; Pisutha-Arnond, 2017). However, it is necessary to emphasize
 386 that this approach was considered highly improbable (Nassau, 1981). Nevertheless, previous studies have
 387 indicated that a minor proportion of geuda sapphire from Sri Lanka and geuda-like sapphire from Mogok
 388 in Burma revealed an alteration in color to blue when subjected to heating in an oxidizing environment
 389 (Hughes, 1997, 2017; Kyi et al., 1999), which is in complete contradiction to the treatment method
 390 employed for the geuda sapphire in a reducing condition. The appearance of certain stones displaying a

391 blue coloration under oxidizing conditions might be attributed to the presence of ilmenite silks, which is
392 composed of Fe and Ti, with Fe in its reduced Fe^{2+} state (Hughes, 1997). Therefore, it is unnecessary to
393 reduce Fe^{3+} to Fe^{2+} ions to generate the $\text{Fe}^{2+}\text{-Ti}^{4+}$ pairs that are responsible for the manifestation of the blue
394 color. Hence, the blue areas have a substantial concentration of Fe ions in form of Fe-Ti pairs, derived from
395 the decomposed ilmenite silk inclusions.

396 According to Nassau (1981) and Koivula (1987), the presence of blue dots in heated
397 sapphire is attributed to remains of dissolved silk inclusions and internal cation diffusion. The diffusion
398 process is positively correlated with temperature and duration of heat treatment (Nassau, 1981). Despite
399 the slow diffusion rates of Fe and Ti, the distances across are extremely short, i.e., just a few micrometers
400 (Nassau, 1981). Consequently, a potential Fe-Ti combination within the corundum's lattice may generate
401 the blue dots.

402 The presence of melt inclusions among the blue dots after high-temperature heating
403 might be due to the decomposition of brown silk and its solubility into the host sapphire as demonstrated
404 by Jung et al. (2009). They predicted a phase relationship within the $\text{Al}_2\text{O}_3\text{-Ti}_2\text{O}_3\text{-TiO}_2$ system based on
405 experimental data and thermodynamic calculation. Consequently, they suggested that a liquid phase (the
406 composition of the liquid inclusion phase varies significantly between Al_2O_3 and Ti_3O_5) could possibly be
407 present at a temperature of 1600 °C and slightly below, which is close to the heating temperature (1650 °C)
408 of our experiment. Silk inclusions as represented by $\text{Ti}_2\text{O}_3\text{-TiO}_2$ components may have dissolved into the
409 host sapphire (Al_2O_3 component), and produced a proper composition of solution which could be melted
410 partially at ≤ 1650 °C. Some of these melts can be preserved as inclusions after cooling down.

411

412 **4.3 Luminescence of sapphire**

413 Luminescence of corundum may be assigned to two types, namely (a) emissions of impurity-related centers
414 such as Ti^{4+} (commonly known) and (b) emissions of defect-related centers, which typically involve either
415 vacancies, such as oxygen (O) or aluminum (Al) vacancies known as F center (color center; from the
416 German "*Farbzentrum*"), or interstitials (Al_i and O_i), possibly trapped at impurities (less known), or both
417 (Vigier et al., 2021a-c). This means that defect-related emission centers in corundum refer to an
418 inconsistency in the atomic arrangement limited to one or a few atoms (often called color-centers). O
419 vacancies (or electron holes) are sometimes called hole centers, because the holes simply designate the
420 absence of an electron. The holes are sometimes filled with one or two electrons in order to maintain
421 electroneutrality (Vigier et al., 2021a).

422 As presented in Fig. 8, a notable orange to red luminescence is easily noticeable under
423 LWUV excitation in most unheated sapphire samples, except for those of the HS group, which appear inert.
424 After heat treatment, the orange to red luminescence that is initially observed in all samples of the LS group
425 and many samples of the SF group turns into a purplish-red luminescence. In contrast, no orange to red or
426 purplish red luminescence is observed in any sample of the HS group both before and after heating.

427 The origin of orange to red luminescence in sapphire remained controversial, with
428 varying ideas among researchers (Vigier et al., 2021a, b, 2023). The occurrence of orange luminescence
429 has been documented in some previous studies (e.g. Spencer, 1927; Kane, 1982; Emmett et al., 2003;
430 Fritsch et al., 2003; Nasdala and Fritsch, 2024). In the beginning, it was hypothesized that this luminescence

431 is associated with the geographic origin of yellow sapphire from Sri Lanka (Webster, 1984). Subsequently,
432 Segura (2013) presented an alternative argument to this notion, suggesting the presence of orange
433 luminescence in various colors of corundum, regardless of treatment or synthetic origin, might be attributed
434 to the existence of some defects. However, the orange to red luminescence observed in our study
435 (characterized by a broad emission band) seems to be associated with complex defect-related centers.

436 Orange to red luminescence in sapphire is not due to impurities (Vigier et al., 2021a, b).
437 HS sapphire (e.g., with ilmenite, FeTiO_3) lack noticeable luminescence, likely because Fe^{2+} suppresses
438 luminescence, contrasting with LS and SF sapphire, which display stronger luminescence both before and
439 after heating. While sample G23SF shows decreased purplish-red luminescence after heating, most display
440 increased purplish-red luminescence, potentially due to complex, defect-related centers in the sapphire
441 lattice. Observations suggested that Fe^{2+} acts as a luminescence quencher in the orange to red range
442 (Andrade et al., 2008; Norrbo et al., 2016; Vigier et al., 2021a, b, c; Vigier and Fritsch, 2022); Orange
443 luminescence generally appears in colorless, low-Fe areas (Segura, 2013; Notari et al., 2003). However, a
444 definitive explanation remains unresolved. Regarding blue luminescence, it has been observed that upon
445 exposure to SWUV light, all natural unheated samples appeared inert. After heating, apart from HS
446 sapphire, a distinct blue luminescence has been detected throughout most heated sapphire samples (Fig. 8).
447 Previous studies suggest that luminescence in sapphire becomes noticeable only at heating temperatures of
448 1000 °C (Hughes and Perkins, 2019), at which point blue luminescence is linked to heat treatment detection.
449 This luminescence is believed to arise from silk inclusions composed of TiO_2 , commonly found in natural
450 blue sapphire. Notably, despite the relatively low Ti concentration (0.02–0.03 wt% oxide) in comparison
451 to Fe (0.05–0.08 wt% oxide) in some samples (e.g., G12SF and G21LS), blue luminescence remains
452 detectable. In contrast, HS samples (e.g., G02HS) show an absence of blue luminescence, likely due to the
453 presence of ilmenite, supporting findings by Norrbo et al. (2016), Andrade et al. (2008), as well as Vigier
454 et al. (2021a–c; 2023) that Fe^{2+} acts as a luminescence quencher. Blue luminescence has been associated
455 with the interaction between O^{2-} and Ti^{4+} ions (Evans, 1994; Wong et al., 1995b; Nasdala and Fritsch,
456 2024), followed by a later hypothesis of a charge transfer process involving Ti^{4+} ions and certain defect-
457 related centers (Lacovara et al., 1985; Mikhailik et al., 2005). However, it was widely accepted that the
458 blue luminescence (characterized by a broad emission band at blue to green region) observed in sapphire
459 under SWUV illumination is associated with the presence of Ti impurities, which are classified as element-
460 related defects (Vigier et al., 2021a, b). Thus, it is likely that the blue luminescence reported in this work
461 is associated with Ti impurities, whereas orange to red luminescence seems to be associated with complex
462 defect-related emission centers.

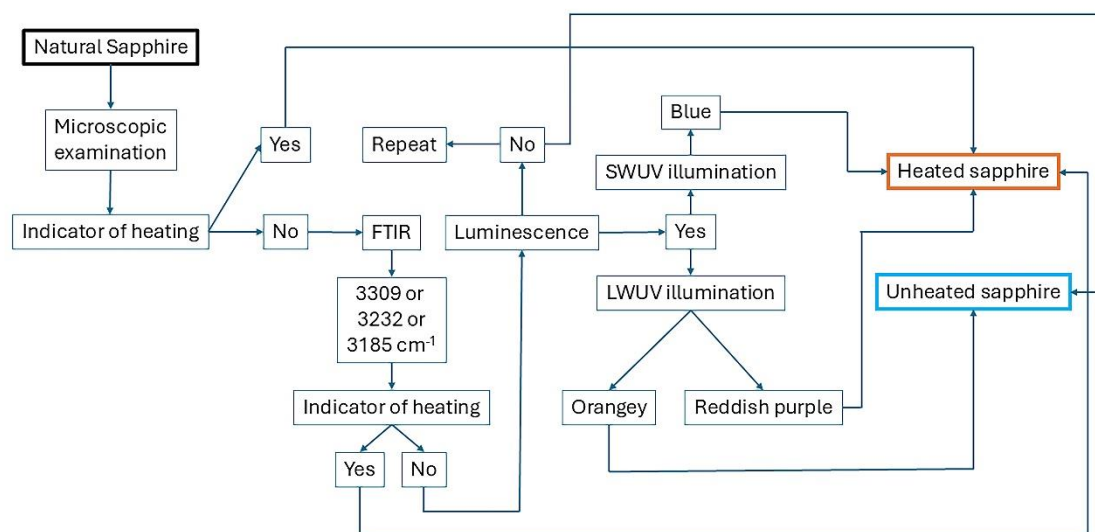
463 The correlation between the orange to red PL emission band (approx. 650 nm, Fig. 9 blue
464 lines) and orange to red luminescence in unheated sapphire (Fig. 8), as well as the emission band (approx.
465 525 nm, Fig. 9 red lines) and blue luminescence in heated sapphire (Fig. 8), is particularly evident in the
466 LS group (Fig. 9b). In contrast, the HS group shows a reduction in emission across the visible spectrum
467 after heating, indicating inertness under LWUV and SWUV excitation (Fig. 9a). The SF group exhibits a
468 notable increase in the red emission band (Fig. 9c) and intense purplish-red luminescence under LWUV
469 excitation after heating. Interestingly, this group also displays strong blue luminescence under SWUV
470 excitation despite the absence of a corresponding blue emission band, likely due to the 325 nm excitation

471 laser used in our PL investigation. Variations in excitation wavelengths significantly affect observed
472 emissions, as noted by Wong et al. (1995a) and Vigier et al. (2023), who showed that their sapphire
473 emission band at 425 nm was only visible with a 254 nm excitation laser. Utilizing distinct SWUV (225
474 nm) and LWUV (365 nm) lasers, or conducting excitation spectroscopy, may yield more accurate results
475 compared to relying solely on a 325 nm laser. Thus, the presence of orange to red luminescence at approx.
476 650 nm and blue luminescence at around 525 nm are vital indicators for differentiating unheated and heated
477 sapphire.

478

479 **5 Conclusions**

480 The present study demonstrates that melt inclusions (~1 μm) serve as indicators of heat treatment in
481 sapphires, and highlights the critical role of luminescence in distinguishing unheated from heated geuda
482 sapphire. Under LWUV light, orange luminescence may arise from defect-related F centers, while blue
483 luminescence under SWUV light likely correlates with Ti impurities. Geuda sapphires with low Fe
484 concentrations exhibit distinct luminescence, whereas those with HS inclusions show minimal
485 luminescence due to Fe^{2+} quenching effects. The presence of orange luminescence may be a helpful
486 indicator for unheated geuda sapphires, while blue luminescence is generally absent in unheated samples,
487 confirming its utility for identifying heated geuda sapphire. Although the 3309 cm^{-1} O–H stretching band
488 from FTIR analysis alone is insufficient for differentiation, increased intensity around the 580 nm of an
489 optical spectra effectively indicates heat treatment, as it corresponds to higher Fe–Ti pair concentrations
490 from silk inclusion decomposition. Combining blue and/or purplish–red luminescence with additional
491 analytical techniques provides a promising strategy for accurately distinguishing between unheated and
492 heated geuda sapphires (Fig. 10). Future research should acquire emission and excitation spectra on the
493 samples before and after heat treatment. Further investigation of luminescence characteristics from various
494 sapphire origins (and colors) and clarify the specific Fe and Ti concentrations impacting luminescence is
495 also recommended. Finally, using as low as 200–254 nm laser excitation may enhance the detection of
496 emission shifts towards the blue region in heated sapphires, improving gemological identification criteria.



497

498 **Figure 10.** Flowchart proposing the combined strategy for heated sapphire identification criteria.

499

500 *Author contributions.* T.P., C.S., B.W., L.N. conducted conceptualization, E.G.Z. acquired samples, T.P.,
 501 C.S., B.W., L.N., C.C.N., M.W., E.L., G.G., T.S. conducted analyses and evaluation, T.P. wrote the
 502 manuscript, all co-authors reviewed and edited the manuscript.

503

504 *Competing interests.* The authors declare that they do not have any conflict of interest.

505

506 *Financial support.* This research project is supported by the Second Century Fund (C2F), Chulalongkorn
 507 University, Bangkok (postdoctoral researcher number 80004543).

508

509 *Data availability statement.* Not applicable

510

511 *Acknowledgments.* We thank Andreas Wagner (Universität Wien) for sample preparation and Sopit
 512 Poompeang (Chulalongkorn University, Bangkok) for assistance in EPMA analysis. Constructive
 513 comments of two anonymous reviewers are greatly appreciated. Author T.P. acknowledges the use of
 514 QuillBot's artificial intelligence to facilitate grammatical verification. This research project was supported
 515 by the Second Century Fund (C2F), Chulalongkorn University, Bangkok (postdoctoral researcher number
 516 80004543). Finally, this research was an initiative of international network, Hub of Talents in Gem and
 517 Jewelry Industries supported by National Research Council of Thailand (Ministry of Higher Education,
 518 Science, Research and Innovation).

519

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