



1 **Large Spatiotemporal Variability in Aerosol Properties over Central Argentina during the**  
2 **CACTI Field Campaign**

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9 **Abstract**

10 Few field campaigns with extensive aerosol measurements have been conducted over continental areas in  
11 the southern hemisphere. To address this data gap and better understand the interactions of convective  
12 clouds and the surrounding environment, extensive in situ and remote sensing measurements were  
13 collected during the Cloud, Aerosol, and Complex Terrain Interactions (CACTI) field campaign  
14 conducted between October 2018 and April 2019 over the Sierras de Córdoba range of central Argentina.  
15 This study describes measurements of aerosol number, size, composition, mixing state, and cloud  
16 condensation nuclei (CCN) collected at the ground and from a research aircraft during seven weeks of the  
17 campaign. Large spatial and multi-day variations in aerosol number, size, composition, and CCN were  
18 observed due to transport from upwind sources controlled by mesoscale to synoptic-scale meteorological  
19 conditions. Large vertical wind shears, back trajectories, single particle measurements, and chemical  
20 transport model predictions indicate that different types of emissions and source regions, including  
21 biogenic emissions and biomass burning from the Amazon and anthropogenic emissions from Chile and  
22 eastern Argentina, contribute to aerosols observed during CACTI. Repeated aircraft measurements near  
23 the boundary layer top reveal strong spatial and temporal variations in CCN and demonstrate that  
24 understanding the complex co-variability of aerosol properties and clouds is critical to quantify the impact  
25 of aerosol-cloud interactions. In addition to quantifying aerosol properties in this data-sparse region, these  
26 measurements will be valuable to evaluate predictions over the mid latitudes of South America and  
27 improve parameterized aerosol processes in local, regional, and global models.

28 **1. Introduction**

29 Earth system models (ESMs), high-resolution models, and observations are key tools for improving our  
30 understanding of the natural and human-influenced atmospheric processes affecting Earth's climate.  
31 Despite recent scientific advances, models still contain biases arising from knowledge gaps and imperfect  
32 parameterizations of important atmospheric processes. The impacts of these biases are multifaceted, but  
33 they make important contributions to uncertainties in the net change of the Earth system energy balance  
34 between preindustrial and present-day periods. Much of this uncertainty has been attributed to current  
35 understanding and/or representation of aerosol–cloud interaction (ACI) processes and the magnitude of  
36 this uncertainty among ESM predictions has remained unchanged for Intergovernmental Panel on Climate  
37 Change (IPCC) assessments since 1995 (Seinfeld et al., 2016; Carslaw et al., 2018).

38 Aerosols are known to perturb cloud hydrometeors, albedo, growth, dissipation, lifetime, and  
39 precipitation (Twomey 1974; Albrecht 1989; Rosenfeld et al., 2014) that subsequently influence climate  
40 over long time scales. ACI processes depend on the co-variability of aerosol and cloud properties. While  
41 the impact of aerosols on shallow marine stratocumulus has been studied extensively (e.g., Twohy et al.,  
42 2005; Wood et al., 2011; Feingold et al., 2024), there are few measurements that characterize the  
43 spatiotemporal variability of key aerosol properties in the presence of convective cloud populations that  
44 can be observed routinely by satellites. One challenge for models is that the spatiotemporal variability of  
45 aerosols and convective clouds is often subgrid-scale (e.g. Fast et al., 2022). While models often contain  
46 parameterizations of subgrid-scale variability for certain cloud types they usually assume aerosols are  
47 constant within a grid cell which could lead to erroneous estimates of the impact of ACI. In addition to



48 model resolution of intersecting aerosol and cloud properties, there are complex ACI pathways for  
49 convective clouds that are still highly uncertain (Fan et al., 2016; Varble et al., 2023) for many reasons.

50 Aircraft in situ and remote sensing data provide key spatiotemporal measurements of aerosol properties  
51 needed to develop an improved understanding of aerosol processes and evaluate and improve models.  
52 Aircraft campaigns are usually relatively short (i.e. a few weeks or less) and do not have the global  
53 coverage of satellite and ground measurements. Most aircraft field campaigns that include aerosol  
54 measurements have been conducted over the northern hemisphere (Reddington et al., 2017; Watson-Parris  
55 et al., 2019). Field campaigns in the southern hemisphere have been conducted in the tropics, such as the  
56 Green Ocean Amazon Experiment, (GoAmazon, Martin et al., 2017) or over the southern hemisphere  
57 ocean, such as the VAMOS Ocean-Cloud-Atmosphere-Land Study Regional Experiment (VOCALS-Rex,  
58 Wood et al., 2011), Southern Ocean Clouds, Radiation, Aerosol Transport Experimental Study  
59 (SOCRATES, McFarquhar et al., 2021), and the combined CLOUDS–Aerosol–Radiation Interaction and  
60 Forcing for Year 2017 (CLARIFY-2017), Observations of Aerosols above Clouds and their  
61 interactions (ORACLES), and Layered Atlantic Smoke and Interactions with Clouds (LASIC)  
62 experiments conducted over the southeast Atlantic Ocean (Barrett et al., 2022). Global aircraft campaigns,  
63 such as the HIPER Pole-to-Pole Observations (HIPPO, Wofsy et al., 2011) and the Atmospheric  
64 Tomography Mission (AToM, Brock et al., 2019) obtained snapshots of aerosol measurements over both  
65 hemispheres, primarily over the ocean. Few aerosol field campaigns with extensive aerosol  
66 measurements, however, have been conducted over subtropical and midlatitude continental areas in the  
67 southern hemisphere.

68 To address this data gap and better understand the interactions of convective clouds and the surrounding  
69 environment, extensive in situ and remote sensing measurements were collected during the Cloud,  
70 Aerosol, and Complex Terrain Interactions (CACTI) field campaign conducted between October 2018  
71 and April 2019 over the Sierras de Córdoba range of central Argentina (Varble et al., 2021). The U.S.  
72 Department of Energy’s Atmospheric Radiation Measurement (ARM) mobile facility (AMF, Mather and  
73 Voyles, 2013) and G-1 research aircraft (Schmid et al., 2014) were deployed during CACTI. The field  
74 campaign was designed to address science questions related to how orographic convective clouds interact  
75 with and depend on environmental conditions, thermodynamics, aerosols, and surface properties. CACTI  
76 occurred at the same time and in the same region as the National Science Foundation (NSF)-led Remote  
77 Sensing of Electrification, Lightning, and Mesoscale/Microscale Processes with Adaptive Ground  
78 Observations (RELAMPAGO) field campaign (Nesbitt et al., 2021). Observations from CACTI have  
79 been used to characterize the range of deep convective cloud life cycles and their relationships with  
80 ambient environments (Feng et al., 2022, 2023), determine the environmental conditions favorable for  
81 deep convection initiation (Marquis et al., 2021, 2023; Nelson et al., 2022), identify factors affecting rain  
82 rates in warm clouds (Borquez et al., 2022), examine the relationship between depth of convective cores  
83 and aerosol concentrations (Veals et al., 2022), evaluate the ability of kilometer scale simulations to  
84 represent the characteristics of mesoscale convective systems (Zhang et al., 2021), and determine the  
85 influence of the South American low-level jet on the convective environment (Sasaki et al., 2023).

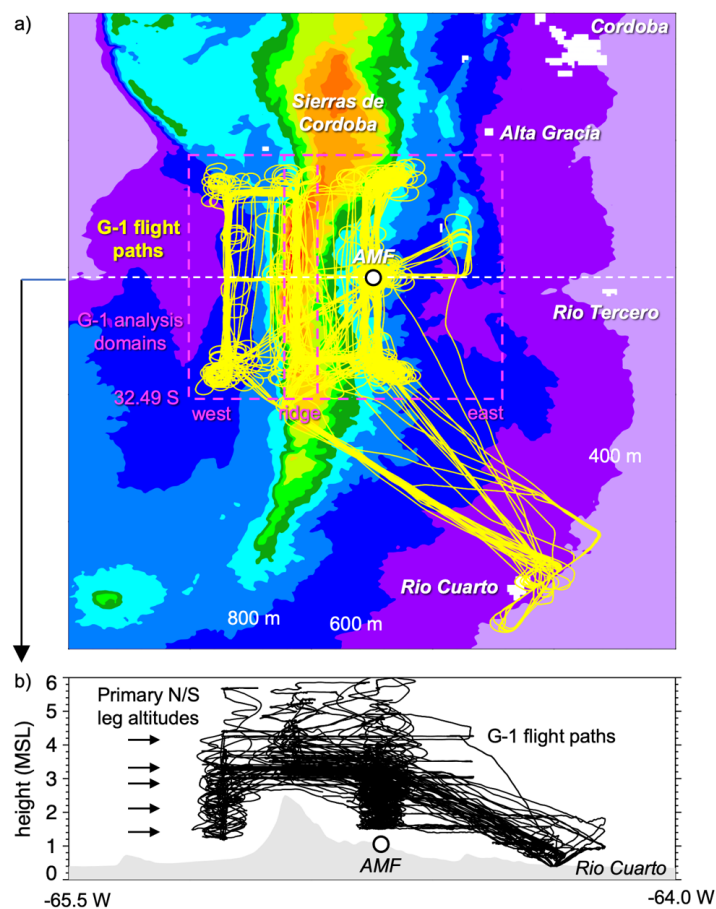
86 In contrast with the previous studies that focus on clouds, this paper describes measurements of aerosol  
87 number, size, composition, mixing state, and cloud condensation nuclei (CCN) collected by the AMF and  
88 G-1 platforms during the CACTI campaign. It is important to understand the spatiotemporal variation of  
89 aerosol properties before assessing how those properties influence convective clouds forming along the  
90 Sierras de Córdoba (SDC) range. CACTI measurements are analyzed to demonstrate the multi-day and  
91 diurnal variations in aerosol properties at the ground as well as the vertical and horizontal variations of  
92 aerosols aloft. As will be shown later, large multi-day variations in aerosol number, size, composition,  
93 and CCN was observed due to regional to long-range transport from upwind sources that is controlled by  
94 mesoscale to synoptic-scale meteorological processes. Since repeated aircraft measurements near the  
95 boundary layer top reveal strong spatial and temporal variations in CCN, inferring the impact of aerosols  
96 on convective clouds over the region will be challenging.



97 **2. Measurements and Models**

98 2.1 Ground measurements

99 A wide range of continuous meteorological, radiation, and aerosol measurements were collected at the  
 100 ground by the ARM mobile facility (AMF, Mather and Voyles, 2013) during CACTI between 15 October  
 101 2018 and 30 April 2019. As shown in Fig. 1a, the AMF was located along the eastern slope of the SDC  
 102 at an elevation of 1141 m MSL. The mountains along the continuous ~300-km long north-south crest of the  
 103 SDC are as high as ~2790 m MSL. About 100 km northeast of the AMF site is Córdoba, the largest city  
 104 in the region with an urban population of ~2.1 million. Rio Cuarto, ~130 km southeast of the AMF site, is  
 105 the second largest city in the region with a population of nearly 180,000. Alta Gracia and Rio Tercero are  
 106 ~60 km to the northeast and east, respectively, each having a population of ~45,000, while Villa Carlos  
 107 Paz 80 km to the north has a population of 75,000. Villa Dolores, 50 km to the west on the other side of  
 108 the SDC crest, has a population of 30,000. The SDC and the lower slopes surrounding mountain range,  
 109 however, are much less populated with no more than a handful of towns having populations near 10,000.



110

111 Figure 1. a) Topography of Sierras de Córdoba and the location of AMF ground site and horizontal G-1  
 112 flight paths and b) west-east cross section of topography at the AMF latitude and the altitude  
 113 of the G-1 flight paths. In a), the white cells denote urban areas and magenta dashed lines  
 114 denote three analysis domains for the G-1 data.



115

116 Near-surface winds (Kyrouac et al., 2018) during the CACTI campaign were predominately from the  
 117 northeast; otherwise, the winds were usually from the east to southeast (Fig. S1a). At about 2.5 km MSL  
 118 which is often within the afternoon convective boundary layer, winds obtained from the radiosondes (3 –  
 119 5 per day; Keeler et al. 2018) were mostly from the north to northeast (Fig. S1b). Winds were usually  
 120 from the northwest in the transition zone between the daytime convective boundary layer and free  
 121 troposphere at 3 km MSL (Fig. S1c) and were predominantly from the west at higher altitudes such those  
 122 at 4.5 km MSL (Fig. S1d). The lower frequency of easterly winds suggests that aerosols originating from  
 123 the most populated areas of Argentina near Buenos Aires are not often directly transported to the AMF  
 124 site. However, they may be transported to the site by more complex circulations. The common directional  
 125 vertical wind shear indicates that aerosols measured at the AMF site could originate from different  
 126 locations at the same time, as will be discussed later.

127 The AMF aerosol measurements used in this study are listed in Table 1. Measurements of aerosol optical  
 128 depth at five wavelengths, total aerosol number concentration, aerosol chemical composition, aerosol size  
 129 distribution, and CCN concentrations at six supersaturations were collected during the 6.5-month period.  
 130 Measurements of scattering (nephelometer) and absorption (particle soot absorption photometer) by  
 131 aerosols were also collected but are not used in this study. Aerosol optical depth at five wavelengths is  
 132 derived from the multifilter rotating shadowband radiometer (MFRSR) during the day when the sky is  
 133 relatively free of clouds (Koontz et al. 2018a). Two particle condensation counters (CPC) obtained total  
 134 aerosol number concentration for particle diameters greater than 3 (Koontz et al. 2018b) and 10 nm  
 135 (Koontz et al. 2018c). The aerosol chemical speciation monitor (ACSM, Ng et al. 2011) measures bulk  
 136 non-refractory organic matter, sulfate, nitrate, ammonium, and chloride for particle sizes less than 1 µm  
 137 (Zawadowicz et al., 2018). Refractory black carbon (rBC) concentrations and size distributions are  
 138 measured by a single particle soot photometer (SP2, Schwarz et al., 2006). The scanning mobility particle  
 139 sizer (SMPS) obtained aerosol concentration for 106 size bins ranging from 10.9 to 495.8 nm (Kuang et  
 140 al., 2018). The ultrahigh sensitivity aerosol spectrometer (UHSAS) collected aerosol concentrations for  
 141 299 size bins between 55.81 and 985.5 nm (Uin et al., 2018). The SMPS and UHSAS size distribution  
 142 was combined to obtain a single aerosol size distribution, with the SMPS and UHSAS values merged  
 143 around diameters of 260 nm. The first column of the CCN counter cycles through six supersaturations  
 144 with measurements at each supersaturation lasting ~10 min (Koontz et al. 2018d) and the second column  
 145 continuously samples CCN at 0.4% supersaturation (Koontz et al., 2018e). Trace gas measurements  
 146 consisted of carbon monoxide (CO; Koontz et al., 2018f) and ozone (O<sub>3</sub>, Springston et al., 2018) and did  
 147 not include aerosol precursor gas-phase species. Additional details on these instruments and a description  
 148 of the meteorological and radiation measurements are presented in Varble et al. (2019, 2021).

149 Table 1. Aerosol measurements, instruments, and sampling rates between 15 October 2018 and 30 April  
 150 2019 at the AMF site.

Measurement	Instrumentation	Sampling Rate
aerosol optical depth	multifilter rotating shadowband radiometer (MFRSR) at 415, 500, 615, 673, 870 nm	20 s
aerosol number concentration	ultrafine (> 3 nm) and fine (> 10 nm) condensation particle counters (CPC)	1 s
aerosol chemical composition	aerosol chemical speciation monitor (ACSM), single-particle soot photometer (SP2)	~30 min, ~ 1 min
aerosol size distribution	ultrahigh sensitivity aerosol spectrometer (UHSAS), scanning mobility particle sizer (SMPS), best estimate aerosol size distribution (BEASD)	1 s, 64 s interpolated to 1 s 1 s
cloud condensation nuclei (CCN) concentration	dual column CCN counter (0.1, 0.2, 0.4, 0.6, 0.8, 1% supersaturation)	1 s
trace-gas concentrations	O <sub>3</sub> , CO monitoring systems	1 s



151 2.2 Aircraft measurements

152 The G-1 aircraft (Schmid et al. 2014) collected a wide range of meteorological, radiation, trace gas, and  
 153 aerosol measurements around the AMF site over 79.4 h on 22 days between 4 November and 8 December  
 154 2018. Flight durations were usually ~4 h, although flight durations on a few days were closer to ~2 h. All  
 155 flights were conducted between 0915 and 1715 LT and flights were either during the morning, mid-day,  
 156 or afternoon. All the flight paths that originated at the Rio Cuarto airport where aircraft operations were  
 157 based are depicted in Fig. 1a. Most of the science flight time was spent along north-south transects at  
 158 constant altitudes over the AMF site, over the crest of the mountain range, and over the western slope of  
 159 the of the SDC. The constant altitude flight legs were conducted below ~4 km MSL although some  
 160 profiles were made up to 6 km MSL as shown in Fig. 1b. These aircraft measurements describe the  
 161 spatiotemporal variability of aerosol properties in the boundary layer and free troposphere over the same  
 162 region where deep convection frequently forms (Feng et al., 2022).

163 One of the objectives of CACTI was to determine how environmental conditions (including aerosol  
 164 properties) influence convective cloud life cycles and how those convective clouds in turn alter aerosol  
 165 properties; therefore, many of the constant altitude flight legs were at or just below cloud base. Cloud  
 166 sampling was usually done within shallow cumulus or cumulus congestus, most frequently observed to  
 167 form along the crest of the SDC. Three out of the 22 G-1 flights were conducted on clear-sky days to  
 168 sample boundary layer and lower free troposphere aerosol properties.

169 Profiles of aerosol properties in the vicinity of the AMF site were likely influenced by vertical wind  
 170 shears, boundary layer mixing, convective updrafts and downdrafts, and cloud processing. However, few  
 171 true profiles through the depth of the boundary layer and lower free troposphere were made since the  
 172 flight paths usually consisted of constant altitude legs connected by short vertical ascents/descents  
 173 between those legs. As described later, the aircraft sampling strategy combined with the large  
 174 spatiotemporal variability of aerosol properties complicates the interpretation of vertical profiles of  
 175 aerosol properties within a short time window.

176 Table 2 lists the G-1 aircraft aerosol measurements used in this study. Two CPC instruments, with 3 and  
 177 10 nm lower cutoffs, were deployed on the aircraft to provide the same type of total aerosol number  
 178 concentrations as the ground site (Koontz et al. 2018g,h). The miniSPLAT (Zelenyuk et al., 2015)  
 179 instrument was deployed to measure the chemical composition of thousands of individual particles. While  
 180 miniSPLAT does not produce bulk concentrations of aerosol composition like a high-resolution time-of-  
 181 flight aerosol mass spectrometer (HR-ToF-AMS, DeCarlo et al. 2006), it does provide information on  
 182 size and composition of individual aerosol particles, which determine aerosol activation into cloud  
 183 droplets (e.g., Saliba et al., 2023). miniSPLAT measures the size, nonrefractory composition, and  
 184 refractory composition (e.g., soot, sea salt, dust) of several hundreds of individual particles per minute to

185 Table 2. G-1 aircraft aerosol measurements, instrumentation, and sampling rate between 4 November  
 186 and 8 December 2018.

Measurement	Instrumentation	Sampling Rate
aerosol number concentration	ultrafine (> 3 nm) and fine (> 10 nm) condensation particle counters (CPC)	1 s
aerosol chemical composition	single-particle mass spectrometer (miniSPLAT), derived particle class information	0.05 s 300 s
aerosol size distribution	scanning mobility particle sizer (SMPS), ultrahigh sensitivity aerosol spectrometer (UHSAS), cloud aerosol spectrometer (CAS), fast cloud droplet probe (FCDP), best estimate aerosol size distribution (BEASD)	1 s
cloud condensation nuclei (CCN) concentration	dual column CCN counter (0.2, 0.5% supersaturation)	1 s
trace-gas concentrations	O <sub>3</sub> , CO, SO <sub>2</sub> monitoring systems	1 s



187 obtain information on aerosol mixing state. The single particle mass spectra were classified into hundreds  
188 of clusters, which for simplicity have been subsequently combined into 14 distinct, physically  
189 meaningful, aerosol types or classes (Zelenyuk et al. 2015). Aerosol size distribution was obtained from  
190 SMPS (Mei and Pekour, 2018), UHSAS (Tomlinson, 2018), passive cavity aerosol spectrometer (PCASP;  
191 Marinovici and Tomlinson, 2018), cloud aerosol spectrometer (CAS; Cromwell et al., 2018), and fast  
192 cloud droplet probe (FCDP; Mei et al., 2018) instruments. CAS and FCDP are usually used to obtain  
193 droplet size distribution, but they also provide coarse mode aerosol number concentration outside of  
194 clouds with additional data processing. The merged aerosol size distribution ARM data product (Pekour  
195 and Ermold, 2023) combines data from these four instruments to produce 35 size bins from 15 nm to 9.69  
196  $\mu\text{m}$  and is used for our analyses of aircraft aerosol size distributions. In contrast to the ground CCN  
197 instrument, continuous measurements of CCN were obtained at 0.2 and 0.5% supersaturations (Koontz et  
198 al., 2018i,j). Trace gas measurements included sulfur dioxide ( $\text{SO}_2$ ; Burk et al., 2018a) in addition to CO  
199 (Burk and Ermold, 2018) and  $\text{O}_3$  (Burk et al., 2018b). The  $\text{SO}_2$  monitor was able to detect large sulfur  
200 dioxide plumes greater than  $\sim 1$  ppb.  $\text{SO}_2$  data below  $\sim 1$  ppb was quite noisy and cannot be used to  
201 examine spatial variability at ppt levels. As with the ground instrumentation, there were no measurements  
202 of trace gas aerosol precursors. The 1-s sampling rate combined with G-1 flight speed of  $100 \text{ m s}^{-1}$  results  
203 in aerosol measurements over 100 m distances.

### 204 2.3 CAM-Chem description

205 The Community Atmosphere Model with chemistry (CAM-chem), a component of the Community Earth  
206 System Model (Danabasoglu et al., 2020) is used to illustrate transport pathways of smoke during  
207 CACTI. This version of the model uses the MOZART-TS1 chemical mechanism (Emmons et al. 2020)  
208 and the Modal Aerosol Model version 4 (MAM4, Liu et al., 2016; Tilmes et al., 2019). The  
209 meteorological fields are nudged to the Modern-Era Retrospective analysis for Research and  
210 Applications, Version 2 (MERRA-2) meteorological analyses. Simulation output with a horizontal grid  
211 spacing of  $1.25 \times 0.94$  degrees is available (Tilmes et al., 2022) at 6-hour intervals from 2001 – 2020.

## 212 3. Results

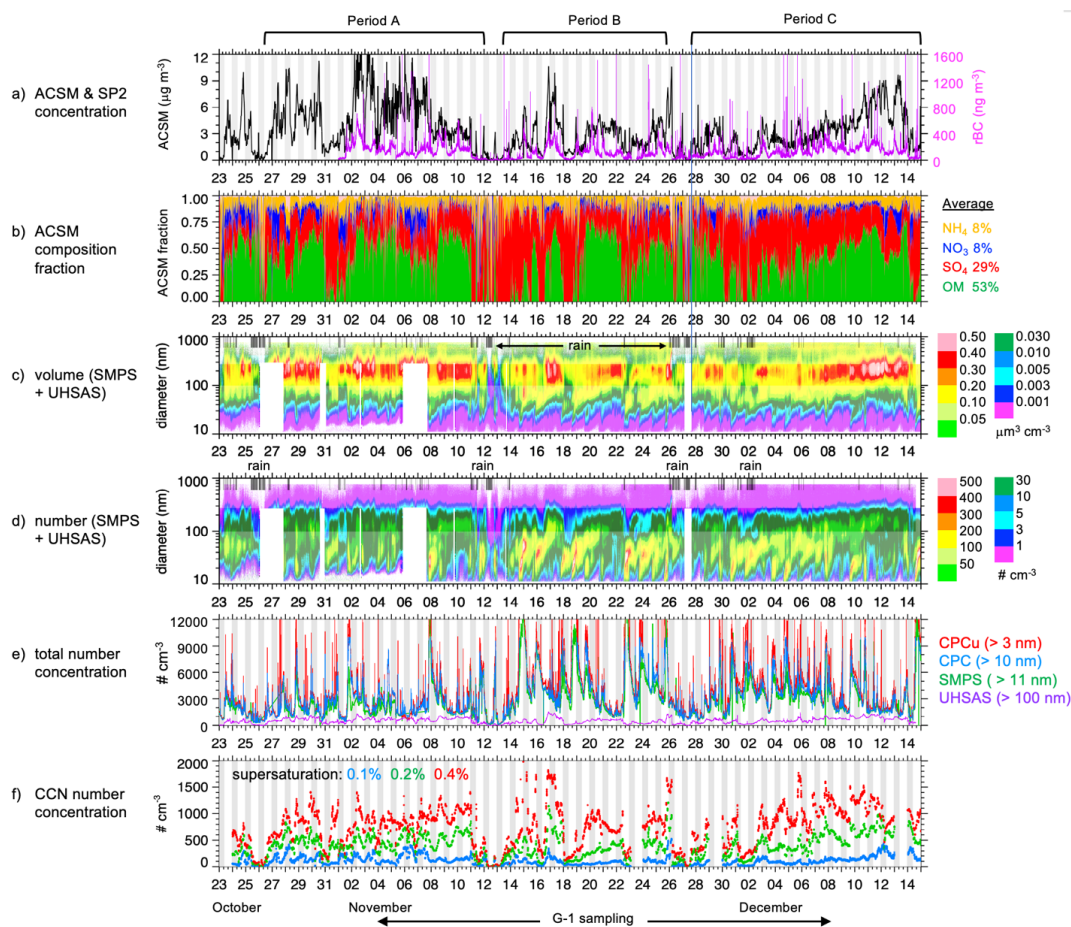
### 213 3.1 Multi-day and diurnal variability of ground aerosol measurements

214 Even though the SDC is not heavily populated, Fig 2a shows large multi-day variations in aerosol mass  
215 for particles less than  $1 \mu\text{m}$  in diameter (PM1) from the ACSM and black carbon from the SP2. For  
216 example, PM1 concentrations were as high as  $15.8 \mu\text{g m}^{-3}$  on November 2 and close to zero on days with  
217 rain (October 25-26, November 11-12, 26-27). The highest PM1 concentrations occurred between the rain  
218 events on October 25-26 and November 11-12 (period A, Figs. S2a and S2b). Between the November 11-  
219 12 and 26-27 rain events (period B), the lowest PM1 concentrations occurred. After November 27 (period  
220 C), PM1 concentrations increase somewhat but are still lower than during period A. In addition to wet  
221 scavenging, the large day-to-day variability is likely due to interactions of synoptic and terrain-induced  
222 circulations that transport biogenic, anthropogenic, and biomass burning aerosols from different source  
223 regions to the site. The temporal variation of PM1 is somewhat correlated ( $r = 0.56$  and  $0.52$  for 415 and  
224 500 nm, respectively) with AOD (Fig. S3a), suggesting that aerosols in the boundary layer contribute to a  
225 large fraction of the column integrated extinction. Note that AOD measurements are missing on 43% of  
226 the daytime periods due to partly to mostly cloudy conditions as indicated by the KAZR-ARSCL (Ka-  
227 band ARM Zenith Radar, Active Remote Sensing of CLouds) ARM product (Johnson et al., 2018; Fig.  
228 S3b). Relatively high and low PM1 concentrations and AOD are often associated with northerly and  
229 southerly winds, respectively, as will be shown later. While multiday trends in PM1 and AOD are similar,  
230 they differ over several hours on some days (e.g., November 23) indicating the presence of aerosol layers  
231 in the free troposphere that significantly contribute to column extinction.

232 Figure 2b shows that most of the PM1 mass is comprised of organic matter (OM, 53% on average) and  
233 sulfate ( $\text{SO}_4$ , 29% on average). The relative contribution of OM is larger on days with relatively high  
234 PM1 concentrations, while the relative contribution of  $\text{SO}_4$  is usually larger on days when PM1



235 concentrations are less than  $1 \mu\text{g m}^{-3}$ . Nitrate ( $\text{NO}_3$ ) and ammonium ( $\text{NH}_4$ ) each comprise 8% of the PM1  
 236 concentrations on average between October 23 and December 15. While the fraction of  $\text{NH}_4$  does not  
 237 vary significantly during the campaign, the fraction of  $\text{NO}_3$  varies substantially. Though  $\text{NO}_3$  is very  
 238 small on many days, it contributes to over 20% of the total mass during 6% of this measurement period.  
 239 OM,  $\text{SO}_4$ ,  $\text{NH}_4$ , and  $\text{NO}_3$  concentrations all exhibited similar variations as PM1 among periods A, B, and  
 240 C (Figs. S2a - S2e).



241

242 Figure 2. Time series of a) PM1 and rBC concentration, b) ACSM composition fraction, c) aerosol  
 243 volume distribution, d) aerosol number distribution, e) total number concentration, and f)  
 244 CCN concentration at three supersaturations, focusing on G-1 flights during the first two  
 245 months of CACTI.

246 Temporal variations in the volume size distribution (Fig. 2c) are consistent with the variability in PM1,  
 247 with volume decreasing dramatically during rain events. Outside of rain events, peak aerosol volume  
 248 usually occurs between 200 and 300 nm. The 25<sup>th</sup> to 75<sup>th</sup> percentiles of accumulation mode particles (>  
 249 100 nm) are the largest during periods A and C and the lowest during period B (Fig. S2f) consistent with  
 250 the variations in PM1 from the ACSM. Nevertheless, there are a few times during period B with the  
 251 highest accumulation mode number concentrations as indicated by the largest 95<sup>th</sup> percentile. Temporal  
 252 variations in aerosol number distributions (Fig. 2d) and total aerosol number concentrations (Fig. 2e)



253 reveal that new particle formation (NPF) events occur on many days that produce large numbers of  
254 ultrafine particles (UFP, diameter < 50 nm). While rain removes a large fraction of accumulation mode  
255 particles, smaller particles are not removed entirely by wet scavenging. In addition to differences in the  
256 meteorology and trace gas precursors between the rain events, NPF and growth is mostly likely controlled  
257 by the presence or absence of accumulation mode aerosols. Since gas-phase aerosol precursors  
258 preferentially condense on the largest particles, the higher PM<sub>1</sub> concentrations during period A  
259 suppresses the formation and growth of UFP (Fig. S2g). Conversely, the lower PM<sub>1</sub> concentrations  
260 during period B permit more particles to form and grow by condensation of trace gas precursors. While  
261 NPF events resume a few days after the rain on November 26 and 27, accumulation mode aerosols and  
262 PM<sub>1</sub> concentrations slowly increase after December 6 and suppress the formation and growth of UFP.  
263 Thus, UFP and Aiken mode number concentrations during period C are somewhat lower than during  
264 period B but higher than during period A.

265 CCN is a function of aerosol number, size, and hygroscopicity (Kohler, 1936; Petters and Kriendenweis,  
266 2007). As with PM<sub>1</sub> and accumulation mode aerosols, CCN concentrations decrease dramatically during  
267 rain events (Fig. 2f). The temporal variations in CCN are similar to variations in number concentrations of  
268 particles larger than 100 nm with a correlation coefficient (*r*) of 0.68, 0.88, and 0.86 for supersaturations  
269 0.1, 0.2, and 0.4%, respectively. Therefore, differences in CCN concentrations (Figs. S4a) among the  
270 three periods are similar to those from the accumulation mode aerosols (Fig. S2e); however, the 95<sup>th</sup>  
271 percentile for CCN at 0.4% supersaturation is the highest during period B because particles with smaller  
272 diameters (usually with higher concentrations) can activate. The ARM CCNKAPPA product (Kulkarni et  
273 al., 2018), that combines Kohler theory with CCN and SMPS measurements, is used to examine the  
274 critical diameter for activation and aerosol hygroscopicity (*κ*) between October 23 and December  
275 15. As expected, critical diameters for activations decrease with increasing supersaturation. The median  
276 critical diameters for 0.1, 0.2, and 0.4% supersaturation are, 195 to 210, 118 to 126, and 76 to 82 nm,  
277 respectively (Fig. S4b); however, the interquartile variations among periods A, B, and C are usually small  
278 (less than 7%). Hygroscopicity is also relatively low and invariant during the campaign period (Fig. S4c),  
279 with median values ranging from 0.19 to 0.29. As indicated by the 95<sup>th</sup> percentiles, hydrophilic aerosols  
280 occurred most often during period B followed by periods A and C. Determining the relative importance of  
281 aerosol number, size, hygroscopicity, and mixing state factors on the temporal variations in CCN seen in  
282 Fig. 2f will require a closure study (e.g. Kulkarni et al., 2023).

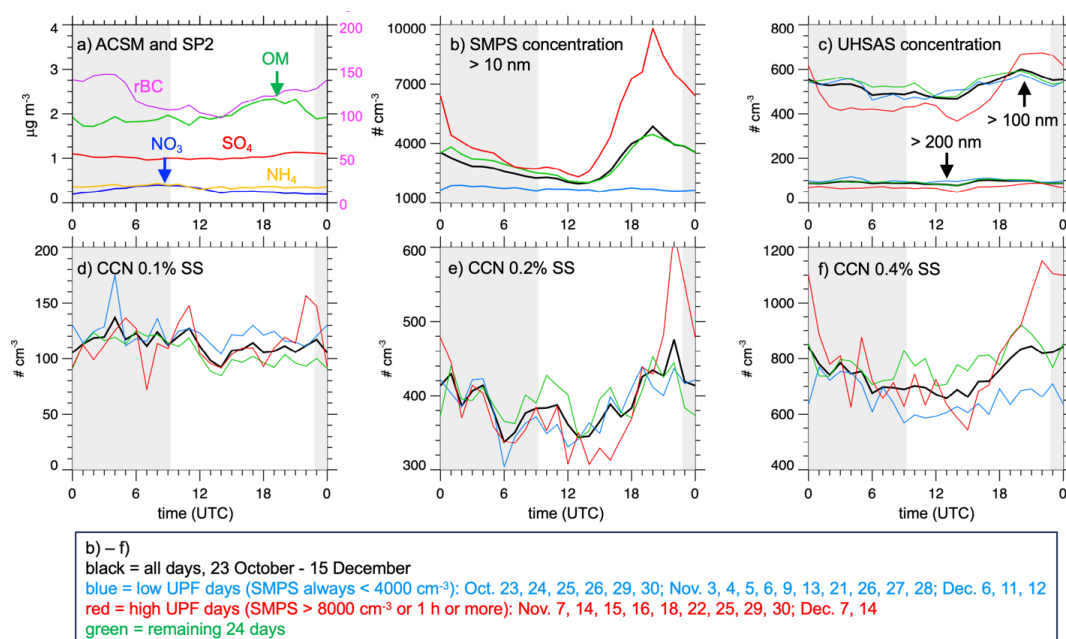
283 Figure 3a shows the average diurnal variations in aerosol composition, which reflect the impact of local  
284 processes such as boundary layer mixing and photochemistry. SO<sub>4</sub> and NH<sub>4</sub> have little diurnal variability  
285 during this seven-week period, suggesting multi-day transport may be the dominate process influencing  
286 those concentrations over the AMF site. NO<sub>3</sub> has peak concentrations around sunrise that are twice as  
287 high (0.4 μg m<sup>-3</sup>) as those during the late afternoon (0.2 μg m<sup>-3</sup>). This diurnal variability is likely due to  
288 temperature dependent condensation/evaporation processes because the lowest temperatures occur at  
289 sunrise and NO<sub>3</sub> tends to evaporate during warmer temperatures. Peak OM concentrations usually occur  
290 during the afternoon, probably due to secondary organic aerosol (SOA) formation that depends on a wide  
291 range of photochemical processes. The net increase in surface OM suggests that photochemistry and/or  
292 entrainment of higher OM from the free troposphere more than compensates for the effects of growing  
293 boundary layer dilution. The peak rBC concentration at night is more difficult to explain since there are  
294 no local emissions of black carbon in the vicinity of the AMF site. Since rBC is chemically inert and can  
295 be treated as a passive tracer, the diurnal variability is due to the combination of horizontal transport and  
296 the effects of boundary layer mixing.

297 Ultrafine particle number concentrations from the SMPS (Fig. 3b) show maximum concentrations  
298 occurring during the late afternoon at 20 UTC (17 LT). The diurnal variations in the UHSAS (Fig. 3c)  
299 also indicate modest concentration increases during the later afternoon for particle diameters greater than  
300 100 nm; however, there is no diurnal variability for particle diameters greater than 200 nm. This indicates  
301 that most of the diurnal variability in aerosol number is driven by ultrafine and Aiken mode aerosols.





302 Since NPF events vary during the sampling period (Fig. 2d), we divide the diurnal variability in aerosol  
 303 number concentration into 19 days with concentrations always lower than  $4000\text{ cm}^{-3}$  (“low-UFP”), 11  
 304 days (1 during period A, 6 during period B, 4 during period C) with number concentrations greater than  
 305  $8000\text{ cm}^{-3}$  for at least one hour (“high-UFP”), and the 24 remaining days that fall in between the low-UFP  
 306 and high-UFP days. The high-UFP days exhibit the largest diurnal variability, the low-UFP days exhibit  
 307 no diurnal variability, and the remaining 24 days have modest diurnal variability that is closer to the  
 308 average over the entire period.



309

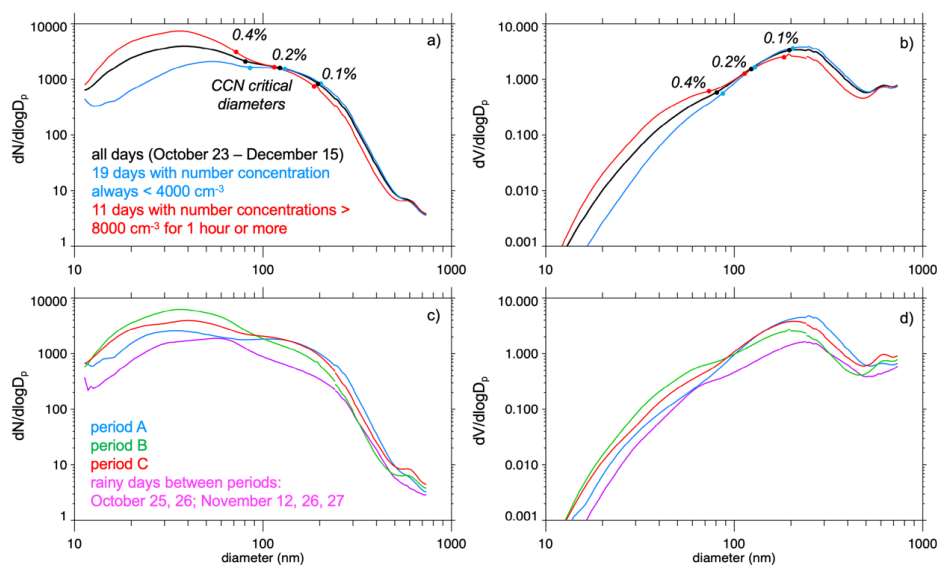
310 Figure 3. Diurnal variability in a) aerosol composition, b) aerosol number concentration from SMPS, c)  
 311 aerosol number concentration from UHSAS, and CCN at d) 0.1%, e) 0.2%, and f) 0.4%  
 312 supersaturation.

313 The diurnal variations in CCN shown in Figs. 3d – 3f are closely related to the diurnal variability in  
 314 aerosol number. Both particles with diameters greater than 200 nm and CCN at 0.1% supersaturation  
 315 exhibited no diurnal variability (Fig. 3c and 3d), Smaller particles start to activate at 0.2% supersaturation  
 316 (Fig. 3e); therefore, CCN at this supersaturation exhibits a modest diurnal variation on average with  
 317 afternoon concentrations  $\sim 28\%$  higher than earlier in the day. However, CCN concentrations during the  
 318 late afternoon on high UFP days are twice as high as those around 15 UTC (12 LT). Finally, particles  
 319 smaller than 100 nm start to activate at CCN at 0.4% supersaturation so that CCN concentrations at this  
 320 supersaturation (Fig. 3f) are about double those at 0.2% supersaturation with a similar late afternoon peak.  
 321 Note that for both 0.2 and 0.4% supersaturations, peak CCN concentrations on high UFP days occur  $\sim 2$  h  
 322 later in the day and closer to sunset than aerosol number concentrations. This suggests growth in the  
 323 aerosol size distribution during the day influences CCN, consistent with the daily growth seen in Fig. 2d.

324 Figure 3c also shows that concentrations of particles with diameters greater than 200 nm are higher on  
 325 low-UFP days and conversely high-UFP days tend to have lower concentrations of larger particles.  
 326 Figures 4a and 4b illustrate the average aerosol number and volume size distribution for 23 October to 15  
 327 December along with the averages for the low-UFP and high-UFP days. The differences in the number  
 328 and volume distribution for particle diameters  $< 100\text{ nm}$  reflect the definition of these days and are



329 consistent with the differences in CCN among the three supersaturations. Peak number concentrations for  
 330 high-UFP days occur between 30 – 40 nm, while those on low-UFP days occur between 50 – 60 nm. In  
 331 addition, high-UFP days have lower concentrations for particle diameters between 150 – 600 nm, likely  
 332 reflecting the time between days to grow particles from UFP to accumulation mode size as indicated in  
 333 Fig. 2d. Differences in the aerosol size distribution leads to differences in the critical particle diameter for  
 334 CCN activation, which are largest for 0.4% supersaturation. The average aerosol number and volume size  
 335 distributions for periods A, B, and C are shown in Figs. 4c and 4d. Periods A and B have the lowest and  
 336 highest average UFP concentrations, respectively, since period B also contains the most high-UFP days.  
 337 Conversely, periods A and B have the highest and lowest accumulation mode concentrations,  
 338 respectively. The average aerosol size distribution for period C is between those for periods A and B,  
 339 and similar to the average size size distribution between the whole October 23 and December 15 period in Figs. 4a  
 340 and 4b. As expected, the rainy days at the AMF site have the lowest aerosol concentrations for all particle  
 341 sizes.



342

343 Figure 4. Average size distributions at the ground site between 23 October and 15 December for days  
 344 with high and low UFP concentration (a and b) and periods A, B, and C (c and d).

345 Understanding the aerosol-cloud interactions in the region will depend on the intersection of cloud  
 346 formation and growth with the growth rates of the aerosol size distribution and diurnal variability in CCN  
 347 concentrations such as those shown in Figs. 2 and 3. Over the crest of the SDC, shallow convection  
 348 typically forms by the late morning. Feng et al. (2022) show that deep convective cells preferentially form  
 349 east of the crest between 15 and 19 UTC (12 and 16 LT), growing and intensifying until about 21 UTC  
 350 (18 LT). This coincides with the increase in CCN concentrations at 0.2 and 0.4% supersaturations which  
 351 is most pronounced on high UFP days, although peak CCN concentrations occur after most of these deep  
 352 convective events.

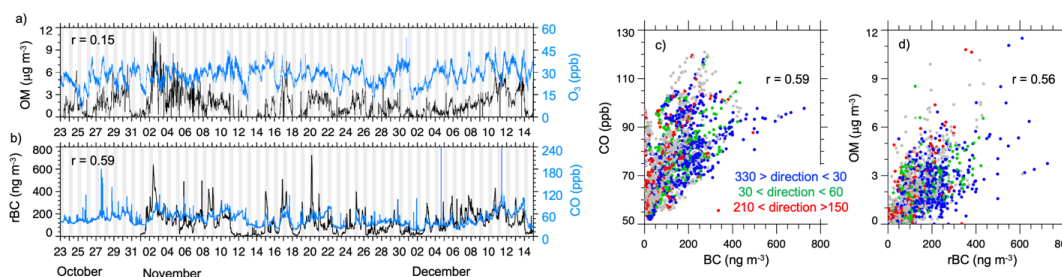
### 353 3.2 Sources of aerosols and trace gases observed at the ground

354 To explore the possible sources of aerosols over the AMF, we next compare aerosol composition with  
 355 trace gases measured at the AMF site. Not surprisingly, O<sub>3</sub> concentrations are relatively low (always < 45  
 356 ppb) at this remote site as shown in Fig. 5a. OM and O<sub>3</sub> are very weakly correlated ( $r = 0.15$ ) even though  
 357 both depend on photochemical production. In contrast, the temporal variability in CO and rBC  
 358 concentrations over the sampling period are moderately correlated ( $r = 0.59$ ) as shown by the time series



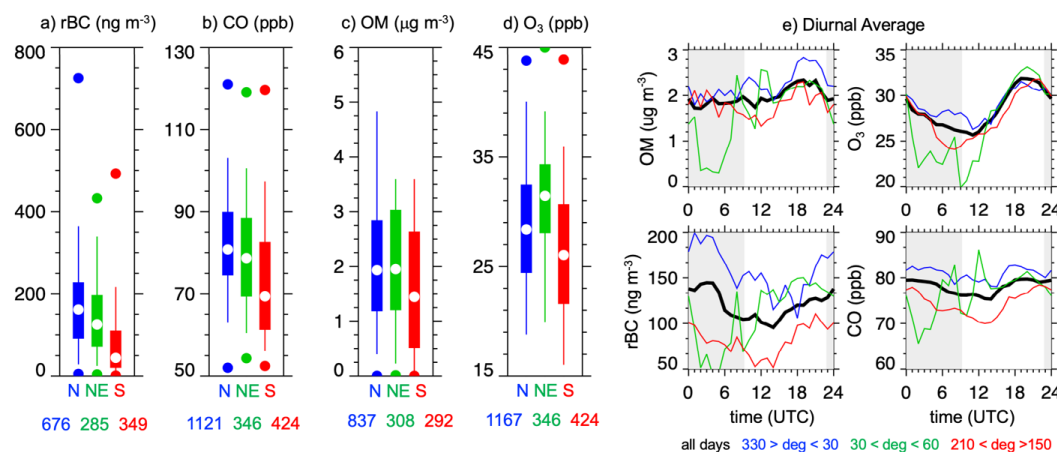
359 (Fig 5b) and scatter plot of 30-min averages (Fig. 5c). Temporal variability of CO and rBC is similar  
 360 because they are usually co-emitted from the same anthropogenic and/or biomass burning sources,  
 361 although the ratio of CO to rBC mass emitted varies from source to source. Correlations between CO and  
 362 rBC often approach one near the emission source; therefore, the lower correlations suggest mixing of  
 363 many emission sources during long-range transport and/or cloud processing of rBC containing particles  
 364 that leads to wet removal of rBC. OM is also moderately correlated with rBC ( $r = 0.56$ , Fig. 5d) for the  
 365 same reason as CO and rBC; however, biogenic sources that do not emit BC can also contribute to a  
 366 significant fraction of OM. Since the near-surface winds are usually from the north to northeast (Figs. S1a  
 367 and S1b), it is possible that a large fraction of OM, rBC, and CO originate from Córdoba and/or fires in  
 368 the Amazon and La Plata basins (Fig. S5a) that are transported by the low-level jet to AMF site. The  
 369 scatter plots in Figs. 5c and 5d are color coded by northerly (330 to 30 degrees), northeasterly (30 to 60  
 370 degrees), and southerly (150 to 210 degrees) wind directions show that the highest BC concentrations  
 371 occur during northerly winds.

372



373 Figure 5. a) Time series of a) OM and O<sub>3</sub> and b) CO and rBC between 23 October and 15 December  
 374 along with scatter plot of c) CO vs rBC and d) OM vs rBC.

375



376 Figure 6. Percentiles of a) rBC, b) CO, c) OM, and d) O<sub>3</sub> as a function of wind direction along with e)  
 377 the diurnal variability of these quantities as a function of wind direction.

378 The effects of wind direction are further illustrated in Fig. 6. The interquartile range of rBC and CO is  
 379 highest for northerly winds (Figs. 6a and 6b), followed closely by northeasterly winds. Southerly winds  
 380 have the lowest concentrations and a substantially lower interquartile range. For OM, the interquartile  
 381 ranges for northerly and northeasterly winds are similar while southerly winds have somewhat lower  
 382 concentrations (Fig. 6c). In contrast, O<sub>3</sub> has the highest interquartile range for northeasterly winds,



383 possibly due to transport of O<sub>3</sub> produced by urban emissions from Córdoba (Fig. 6d). Despite the low  
384 correlation between OM and O<sub>3</sub>, they have a similar average diurnal variability (Fig. 6e). This suggests  
385 the low correlation is due primarily to multi-day variability. Maximum OM and O<sub>3</sub> concentrations during  
386 the afternoon are ~25% and ~23% higher than at sunrise on average, respectively. OM diurnal variability  
387 is also similar to the diurnal variability of CCN at 0.2% supersaturation (Fig. 3e). The average diurnal  
388 variations for rBC and CO are similar with the highest values at night and lowest values a few hours after  
389 sunrise. While the diurnal variability for northerly and southerly winds is similar, the concentrations are  
390 significantly higher for northerly winds. OM, O<sub>3</sub>, rBC and CO concentrations are all lowest at night for  
391 northeasterly winds.

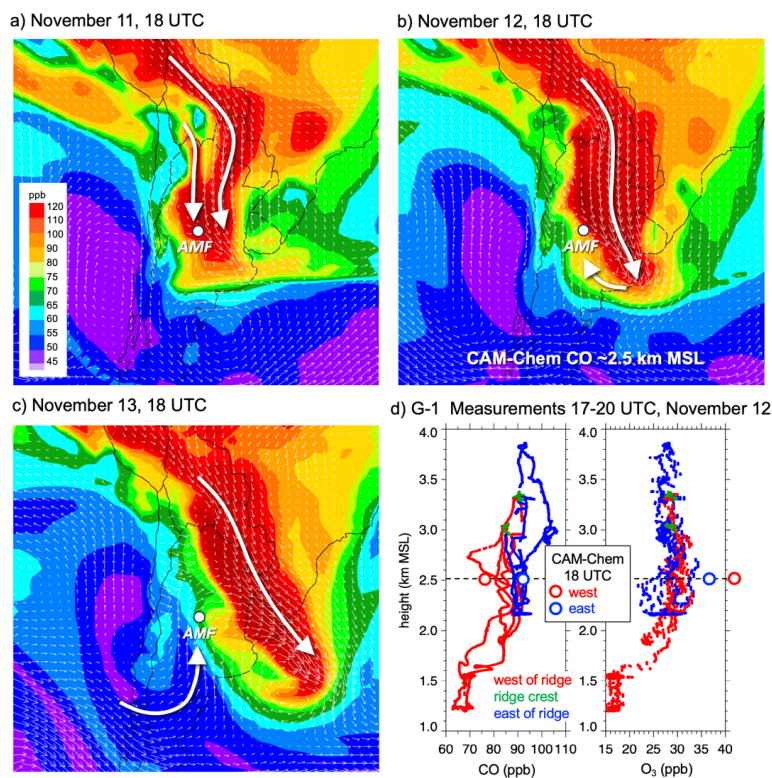
392 The CAM-Chem model is now used to illustrate transport pathways of biomass burning aerosols  
393 suggested by the relationships in Figs. 5 and 6. The largest CO emissions from smoke during the  
394 campaign period occur over the western Amazon basin as well as southern Brazil (Fig. S5a). Smaller fires  
395 occur across most regions outside of Patagonia and the Atacama Desert. In addition to primary  
396 carbonaceous particles, fires emit gas-phase aerosol precursors, including SO<sub>2</sub> (Fig. S5b), with emissions  
397 rates that are a factor of ~2 lower than CO. These SO<sub>2</sub> emissions ultimately become sulfate and affect  
398 CCN concentrations since sulfate is more hygroscopic than OM and rBC.

399 CAM-Chem predictions of CO and wind direction are evaluated with G-1 measurements as shown in  
400 Figs. S6 and S7, respectively. The vertical and temporal variations in the vicinity of the AMF site are in  
401 reasonable agreement with G-1 data, even though the coarse spatial grid does not resolve local measured  
402 variations. High CO concentrations are simulated during periods of northerly and northwesterly winds,  
403 suggesting that smoke from fires north of the AMF were transported over the site. Conversely, CO  
404 concentrations are lower when the winds were from the south. Fires and anthropogenic emissions of CO  
405 are both lower over the less populated areas of southern Argentina (Fig. S5a).

406 Since CAM-Chem is in reasonable agreement with the G-1 CO measurements, Fig. 7 shows horizontal  
407 cross sections of CO at ~2.5 km MSL to illustrate transport during on episode between November 11 and  
408 13. Figure 7a shows the transport of CO from the western Amazon to the AMF site by a low-level jet. A  
409 trough pushes the low-level jet and plume of CO towards the east on November 12 (Fig. 7b), somewhat  
410 reducing the CO concentrations. While the winds have become southerly near the AMF site, CAM-Chem  
411 suggests local recirculation on the western side of the CO plume keeps the concentrations relatively high.  
412 As the trough continues to propagate to the east on November 13 (Fig. 7c), lower CO from the Pacific  
413 Ocean is transported over the AMF. Vertical profiles of CO from the G-1 on November 12 in Fig. 7d  
414 show that the highest concentrations occur east of the SDC. This is similar to the horizontal gradient in  
415 CO at 2.5 km MSL produced by CAM-Chem that shows the CO plume east of the AMF site (Fig. 7b).

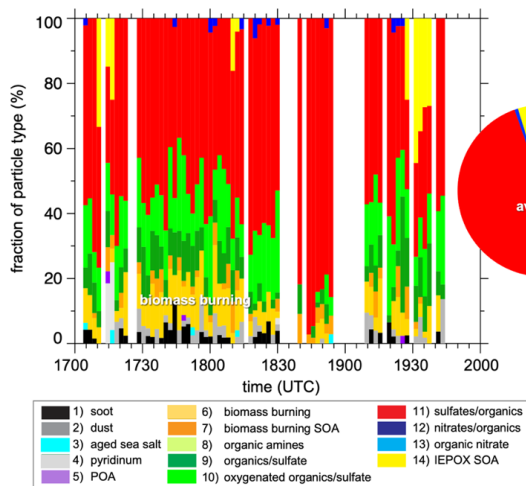
416 Single particle measurements from miniSPLAT from the November 12 G-1 flight shown in Fig. 8 reveal  
417 that ~12% of the particles originate from fires. The rest of the particles are composed of organics mixed  
418 with varying amounts of sulfate. A large portion of the organic material not identified as biomass burning  
419 aerosols is oxygenated organics, likely from biogenic sources in the Amazon that are transported by the  
420 same winds to the AMF site. As noted earlier, PM<sub>1</sub> measured by the ground ACSM was relatively low  
421 between November 11 and 12 because of wet scavenging by rain and changing synoptic conditions. PM<sub>1</sub>  
422 concentrations from the ACSM increased from 0.3 to 2.5  $\mu\text{g m}^{-3}$  during the three-hour G-1 flight. Average  
423 PM<sub>1</sub> concentrations were 1.2  $\mu\text{g m}^{-3}$ , comprised of 54% OM, 19% SO<sub>4</sub>, 11% NO<sub>3</sub>, and 16% NH<sub>4</sub>. While  
424 bulk sulfate mass was 19% on average, the miniSPLAT data sheds light on the aerosol mixing state that  
425 suggests it was mixed with ~80% of the particles. The three particle classes that have the largest fraction  
426 of particles are made up of organics and sulfate at various ratios. This variability in aerosol mixing state  
427 could impact CCN concentrations and aerosol optical properties (e.g., Ching et al. 2017; Saliba et al,  
428 2023).

429 The circulations depicted in Fig. 7 are repeated on many days during CACTI, indicating that the AMF site  
430 is periodically influenced by biomass burning. Thus, the sources of aerosols change from day to day.



431

432 Figure 7. CAM-Chem simulations of CO at ~2.5 km MSL at 18 UTC on a) November 11, b)  
 433 November 12, and c) November 13 along with d) vertical profiles of CO measured by the G-1  
 434 aircraft between 17 and 20 UTC on November 12.



435

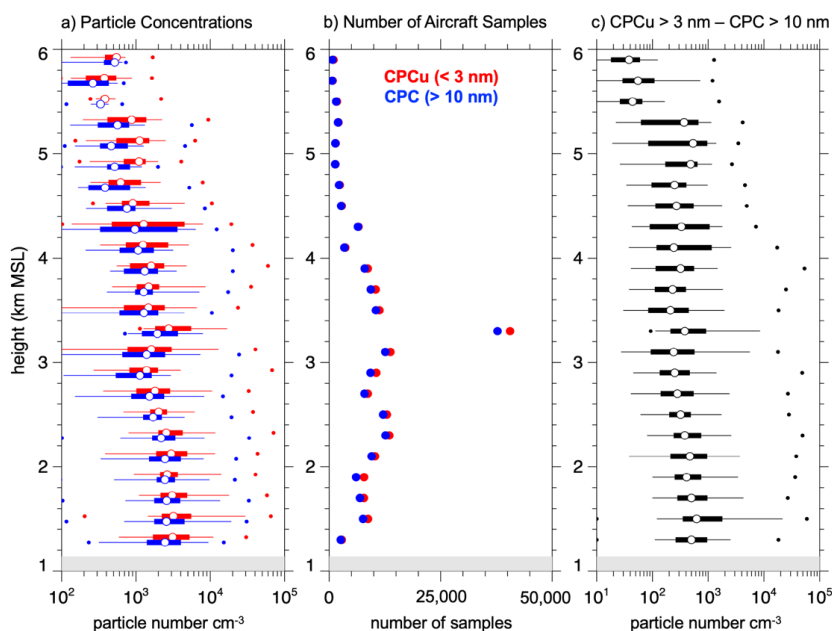
436 Figure 8. Particle classes derived from the aircraft miniSPLAT mixing state measurements on  
 437 November 12.



438 3.3 Aircraft Measurements

439 The evolving size distribution of ground measured particles suggest that NPF events are followed by  
 440 multi-day growth to larger sizes. However, it is not clear whether these NPF events occur near the surface  
 441 or aloft (either in the upper boundary layer or lower free troposphere) with mixing to the surface as has  
 442 been observed at other locations (Chen et al. 2018; O’Donnell et al. 2023; Wang et al. 2023). Therefore, it  
 443 is important to examine the vertical variations in aerosol properties that are available from the G-1  
 444 aircraft.

445 The overall particle number concentrations observed by the two CPC instruments on flight days between  
 446 November 4 and December 8 are shown in Fig. 9. Median CPC concentrations are highest within 1 km of  
 447 the ground and then gradually become smaller with height, up to a factor of ~2.2 smaller by 3 km MSL  
 448 (Fig. 9a). Median concentrations increase somewhat at 3.3 km MSL and then continue to gradually  
 449 decrease with height so that values at 6 km MSL are almost an order of magnitude smaller than at the  
 450 ground. Similar trends are produced for the interquartile and 5<sup>th</sup> to 95<sup>th</sup> percentile ranges; however,  
 451 maximum concentrations that are between 10,000 and 100,000 cm<sup>-3</sup> occur up to ~4 km MSL before  
 452 decreasing with height. This suggests that high concentrations of UFP could occasionally be entrained  
 453 into the growing daytime boundary layer and mixed to the surface. Note that number concentration  
 454 percentiles at 3.3 km MSL are slightly different than those at adjacent altitudes because the number of  
 455 samples at that altitude are much larger than any other altitude (Fig. 9b). While there are large temporal  
 456 variations in UFP concentrations within the G-1 flight periods as will be shown later, under-sampling is  
 457 not likely to affect the overall vertical variations. In addition, the number of samples above 4.5 km MSL  
 458 are much lower than at other altitude, so caution is warranted in interpreting the vertical variations at  
 459 those altitudes. Particle concentrations smaller than 10 nm, obtained by computing the differences  
 460 between the two CPC instruments, are similar up to 5.5 km MSL (Fig. 9c).



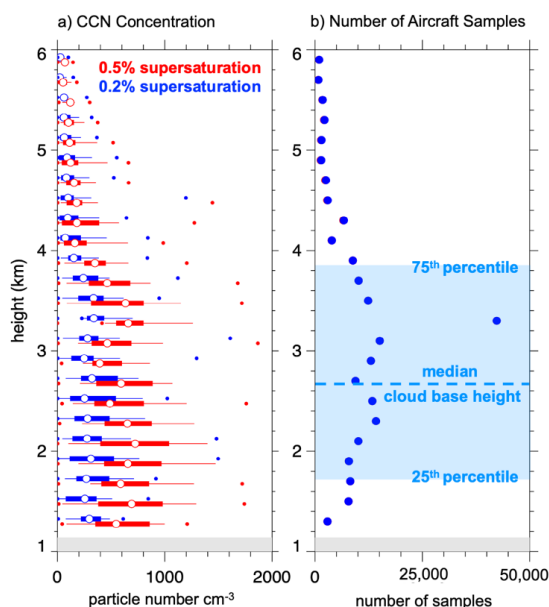
461

462 Figure 9. a) Particle number concentration percentiles as a function of height from all the G-1  
 463 measurements during CACTI along with b) the number of samples at each altitude bin. c) is  
 464 the same as a), except for particle diameters between 3 and 10 nm.



465 Next, we examine the vertical profiles of particle concentrations in relation to the low and high UFP days  
 466 as determined by the surface CPC measurements. The median particle number concentrations for the  
 467 seven G-1 flights associated with high UFP days were higher than the median among all the G-1 flights  
 468 up to ~ 3 km MSL (Fig. S8). Conversely, the four G-1 flights on the low UFP days were lower than the  
 469 median among all the G-1 flights up to ~3 km MSL. When the G-1 flights are divided into periods A, B,  
 470 and C determined by the surface CPC measurements, period B has the highest median concentration near  
 471 the surface, followed by periods C, A, and the rainy days. The results in Fig. S8 suggest that the boundary  
 472 layer particle number concentrations are consistent with the broad variations seen at the ground. However,  
 473 it is not surprising that the trends in the surface number concentrations are not representative of the  
 474 temporal variability in the upper boundary layer and free troposphere.

475 A summary of the vertical profiles of CCN concentration for all the aircraft flights is shown in Fig. 10a.  
 476 The median CCN at 0.2% supersaturation between the surface and 3.5 km MSL varies between 255 and  
 477 335  $\text{cm}^{-3}$ . At 0.5% supersaturation, CCN concentrations are about twice as high as those at 0.2% and vary  
 478 between 546 and 724  $\text{cm}^{-3}$ . Above 3.5 km MSL, concentrations drop to below 100  $\text{cm}^{-3}$  at 4 and 5.5 km  
 479 MSL for 0.2 and 0.5% supersaturations, respectively. The median cloud base height of 2.6 km MSL and  
 480 the interquartile range from KAZR-ARSCL computed during the G-1 flight periods (Fig. 10b) illustrates  
 481 the relevant altitudes in which CCN can be entrained into clouds. The median CCN concentrations at the  
 482 median cloud base height are similar to those at the surface suggesting a well-mixed boundary layer.  
 483 However, additional analyses are needed to determine whether surface CCN measurements are  
 484 representative of the conditions at cloud base at the same time as when the G-1 flew directly over the  
 485 AMF site.



486 Figure 10. a) CCN concentration percentiles as a function of height from all the G-1 measurements  
 487 during CACTI along with b) the number of samples at each altitude bin. Light blue shading in  
 488 b) denotes the 25<sup>th</sup> to 75<sup>th</sup> percentile of cloud base height observed over the AMF site.  
 489

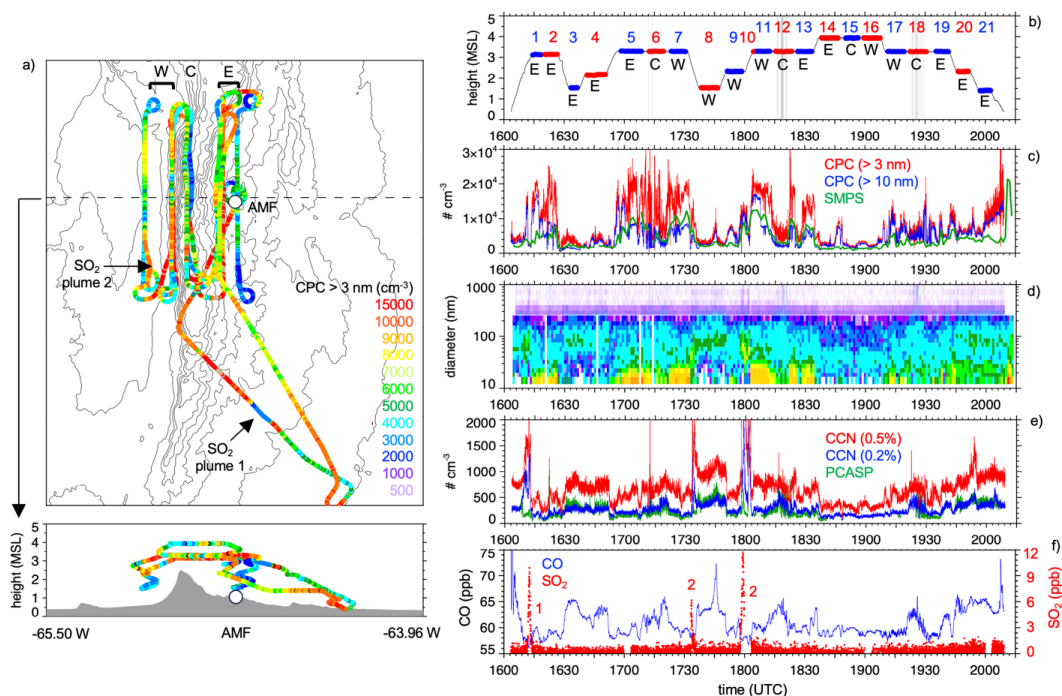
490 Median CCN concentrations for low UFP days below 4.5 km MSL are 20 – 25% lower than the median  
 491 for all the flights at 0.2% supersaturation (Fig. S9a) and up to 70% lower at 0.5% supersaturation (Fig.  
 492 S9b). Differences in the median CCN concentrations are also produced among periods A, B, and C;  
 493 however, the vertical variations are noisy and suggest that additional flights are needed to account for the



494 strong spatiotemporal variations in CCN aloft. While there are only two aircraft flights on rainy days, they  
 495 indicate CCN concentrations are lower than all other flights within 0.3 and 1 km of the surface for 0.2%  
 496 and 0.5% supersaturations, respectively. Above that altitude, median CCN concentrations are similar to  
 497 the median of all the aircraft flights.

498 3.4 December 3 Case Study

499 The percentiles in Figs. 9 and 10 reflect both large spatial and temporal variations in CPC and CCN  
 500 concentrations around the AMF site. Measurements during the afternoon on December 3 are shown in  
 501 Fig. 11 to demonstrate the spatiotemporal variability of aerosol number, size distribution, CCN, and trace  
 502 gases for a particular flight. The flight path color coded by CPC concentrations indicates that particle  
 503 number varied by over an order of magnitude both horizontally and vertically during the 4-h flight period  
 504 (Fig. 11a). By comparing the flight altitude (Fig. 11b), the time series of particle number concentration  
 505 (Fig. 11c), and the aerosol size distribution (Fig. 11d), one can see that the highest UFP number  
 506 concentrations with diameters < 30 nm occurred along ~3.3 km MSL flight legs (1, 2, 5-7, 11-13, 17-19).  
 507 While clear skies were observed directly over the AMF site on this day, a line of orographic cumulus  
 508 formed over the crest of the SDC. While there was widespread cumulus over the ridge, the time spent  
 509 flying through clouds was less than 2.7% total time at 3.3 km MSL. There were few ultrafine particles  
 510 (diameters < 30 nm) for flight legs at lower (legs 3, 4, 8, 9, 20, 21) and higher (legs 14-16) altitudes. For  
 511 the lower and higher altitude legs, the highest particle concentrations occurred at diameters of ~80-100  
 512 nm and ~30-60 nm, respectively. The spatiotemporal variability in CCN concentrations (Fig. 11e) is  
 513 similar to the accumulation mode aerosols with the highest CCN concentrations occurring for the lowest



514

515 Figure 11. a) Spatial variations in particle number concentration (> 3 nm) for the December 3 aircraft  
 516 flight along with temporal variations in b) altitude, c) particle number concentrations from  
 517 CPC and SMPS instruments, d) aerosol number distribution from BEASD, e) CCN and  
 518 particle number concentrations from PCASP, and f) CO and SO<sub>2</sub> concentrations. E, C, and W  
 519 in b) denote legs east of the crest, over the crest, and west of the crest as shown in a).

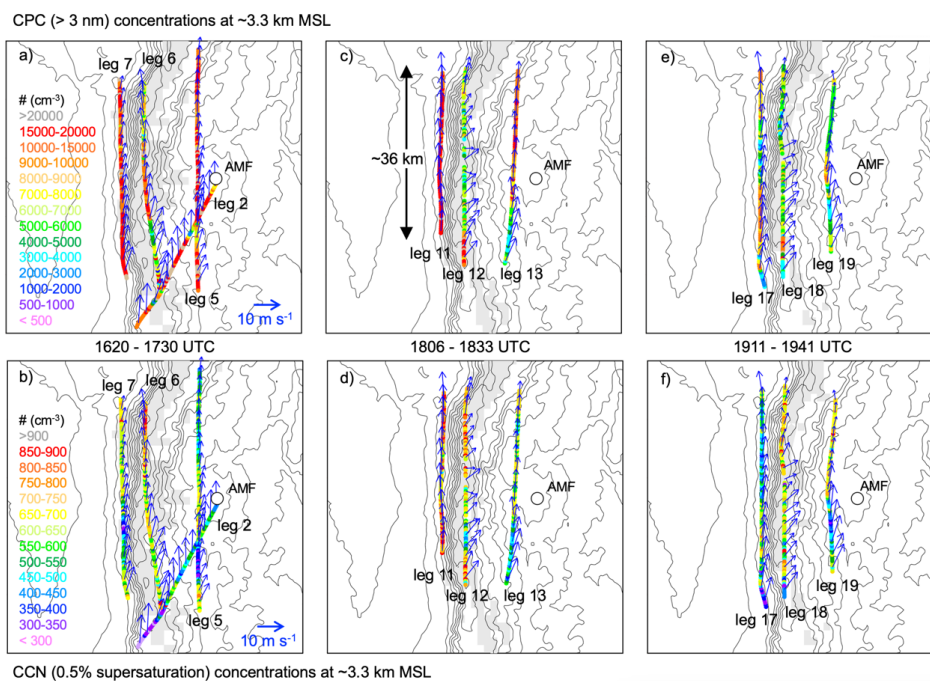




520 flight legs where the highest concentrations of larger particles occur. Outside of the Rio Cuarto airport,  
 521 the highest CO concentrations occur along the lowest flight legs (Fig. 11f) and spatiotemporal variations  
 522 in CO are similar to spatiotemporal variations in accumulation mode aerosols.

523 On this day, the G-1 encountered two SO<sub>2</sub> plumes with concentrations as high as 12 ppb denoted as  
 524 plume #1 and #2 in Figs. 11a and 11f. The G-1 passed through plume #2 twice at different altitudes. The  
 525 narrow plume width and high concentrations suggest that a local source is responsible for these two  
 526 plumes. Local maxima in particle number and CCN concentrations occurred at the same location as the  
 527 SO<sub>2</sub> plume, suggesting that SO<sub>4</sub> was also present. SO<sub>4</sub> is more hydrophilic than other aerosol species,  
 528 which may be why CCN concentrations are higher within the plume. Other than one other small plume of  
 529 SO<sub>2</sub> (~3 ppb) on December 3, no other SO<sub>2</sub> plumes of this magnitude were observed by the G-1 during  
 530 the campaign. The detection limit of the instrument is too high to describe variability of SO<sub>2</sub> at the ppt  
 531 level.

532 The advantage of the multiple constant altitude flight legs is that they can be compared to determine how  
 533 aerosol properties aloft evolve in time. Figure 12 depicts the variations in total particle number and CCN  
 534 concentrations for three periods: 1620-1730 UTC, 1806-1833 UTC, and 1911-1914 UTC, each with flight  
 535 legs west, over, and east of the crest of the SDC. During the first period between 1620 and 1730 UTC,  
 536 particle concentrations exceed 15000 cm<sup>-3</sup> along most of the four flight legs (Fig. 12a). CCN  
 537 concentrations vary by a factor of three (300 to 900 cm<sup>-3</sup>) over the region during the same period (Fig.  
 538 12b). The observed winds at this altitude are southerly and usually between 8 to 10 m s<sup>-1</sup>.



539

540 Figure 12. Spatial variations in a) particle number concentrations > 3 nm (a, c, and e) and CCN  
 541 concentrations (b, d, and f) for all the ~3.3 km MSL constant altitude transects on December  
 542 3. Blue arrows depict wind speed and direction at 30 s intervals. Gray shading denotes cloud  
 543 optical depth > 2 obtained from GOES satellite (~Δx = 2 km).

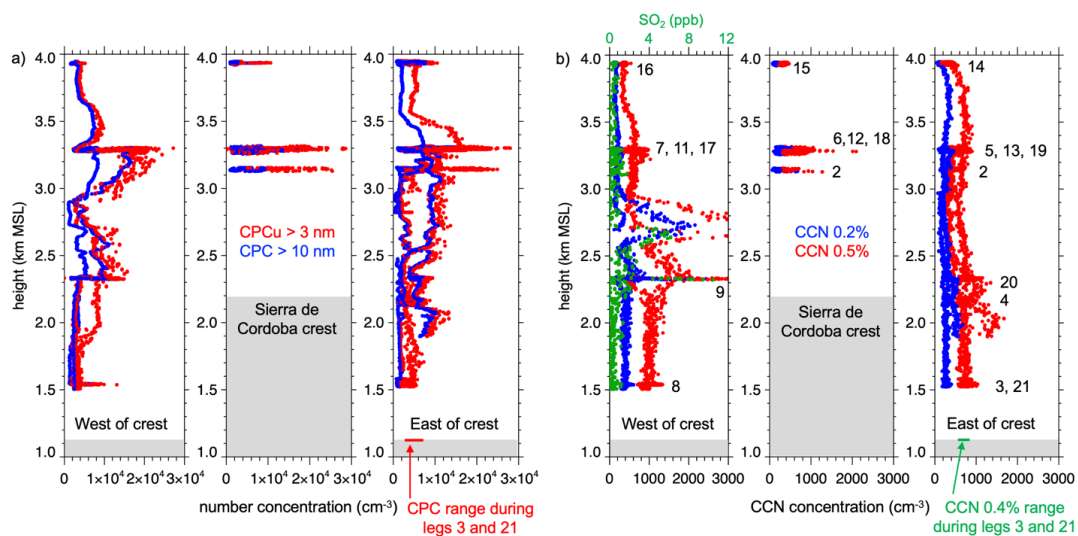
544



545 The second period between 1806 and 1833 UTC took place 36 to 133 minutes after the first period;  
 546 therefore, aerosols sampled along legs 2, 5, 6, and 7 would be transported ~17-80 km to the north by the  
 547 time the aircraft conducted legs 11 to 13. Since the flight legs are ~40 km long, aerosols measured during  
 548 the second period are not likely to be the same as those sampled during the first period. During the second  
 549 period, most of the particle concentrations remain above 15000 cm<sup>-3</sup> for the flight legs west and east of  
 550 the SDC crest (Fig. 12c). However, lower concentrations between 2000 to 5000 cm<sup>-3</sup> are being  
 551 transported by southwesterly winds towards the AMF site along the southern third of leg 13. At this time,  
 552 CCN concentrations are the highest for legs 11 and 12 and lower along leg 13 closer to the AMF site (Fig.  
 553 12d). Interestingly, winds over the crest along legs 6 and 12 switched from southerly to southwesterly and  
 554 particle number concentrations became a factor of two lower between during periods one and two. As will  
 555 be shown in more detail later, this is likely due to both the growing boundary layer and subsidence that  
 556 transports lower particle number concentrations to this altitude. While winds in the lower boundary layer  
 557 along legs 3 and 8 and from the AMF radiosonde at 18 UTC (not shown) are from the north to northeast,  
 558 southwesterly winds observed near the top of the boundary layer top along leg 12 and from the 18 UTC  
 559 AMF radiosonde at 3.3 km reflect the wind direction shear between the boundary layer and free  
 560 troposphere.

561 The same wind pattern for period two persists through period three between 1911-1941 UTC. Strong  
 562 gradients in total aerosol number and CCN concentrations are observed for all three flight legs during this  
 563 period (Figs. 12e and 12f). CCN concentrations are the highest along leg 18 over the crest during period  
 564 three. It is possible that recycling of aerosols through clouds change the size and hygroscopicity of  
 565 aerosol populations and thus CCN in this region, but that requires further analysis.

566 While Fig. 12 illustrates the strong spatial and temporal variations in aerosol properties at 3.3 km MSL on  
 567 December 3, Fig. 13 shows the vertical variations in number and CCN concentrations that are divided into  
 568 flight paths west, over, and east of the crest of the SDC. West of the crest, the highest number  
 569 concentrations occur in two layers, one between 2.3 and 2.7 km MSL and the other between 2.9 and 3.3



570

571 Figure 13. Vertical variations in a) particle number concentrations and b) CCN concentrations on  
 572 December 3 divided into transects that are west, over, and east of the Sierras de Córdoba  
 573 crest. Green dots in b) denote SO<sub>2</sub> concentrations west of the crest on December 3.



574 km MSL (Fig. 13a). Note that the smallest particles (CPC > 3 nm) concentrations vary for a given altitude  
575 due to both spatial and temporal variability along the flight legs. The differences for the larger particles >  
576 10 nm are much smaller at these altitudes; however, there are larger differences below 2.3 km MSL.

577 Thus, the spatial variability for UFP and larger particles is not necessarily the same. East of the crest, a  
578 layer of high number concentrations occurred between 3.1 and 3.5 km MSL; however, there is not a  
579 distinct second layer as seen west of the crest and the two CPC instruments have similar spatiotemporal  
580 variability below 3.1 km MSL. While the spatial variability in aerosol number is similar west, over, and  
581 east of the crest at 3.3 km MSL, the variability at 3.9 km MSL west of the crest is lower than over and  
582 east of the crest. In general, Fig. 13 illustrates that aerosol number concentrations and variability can be  
583 different west and east of the SDC crest.

584 As shown in Fig. 13b, there are also differences in CCN concentrations between the west and east sides of  
585 the mountain range. There are two layers of high CCN concentrations that were sampled by the aircraft  
586 west of the crest, one between 2.6 and 2.9 km MSL and the other between 2.3 and 2.5 km MSL. The  
587 highest SO<sub>2</sub> concentrations also occur within these layers, suggesting that SO<sub>4</sub> produced by the SO<sub>2</sub>  
588 plume #2 (Figs. 11a and 11f) lead to higher number concentrations and more hydrophilic aerosols. SO<sub>2</sub>  
589 concentrations were low at all altitudes east of the crest. Note that the layer between 2.6 and 2.9 km MSL  
590 occurs between the constant altitude transects; therefore, the spatial extent of this layer cannot be  
591 determined, and other layers could be missed between the constant altitude transects. In addition, CCN  
592 concentration profiles at both 0.2 and 0.5% supersaturations are somewhat lower east of the crest below  
593 2.5 km MSL and somewhat higher east of the crest above 3.5 km MSL. Thus, CCN also exhibits  
594 differences west and east of the crest.

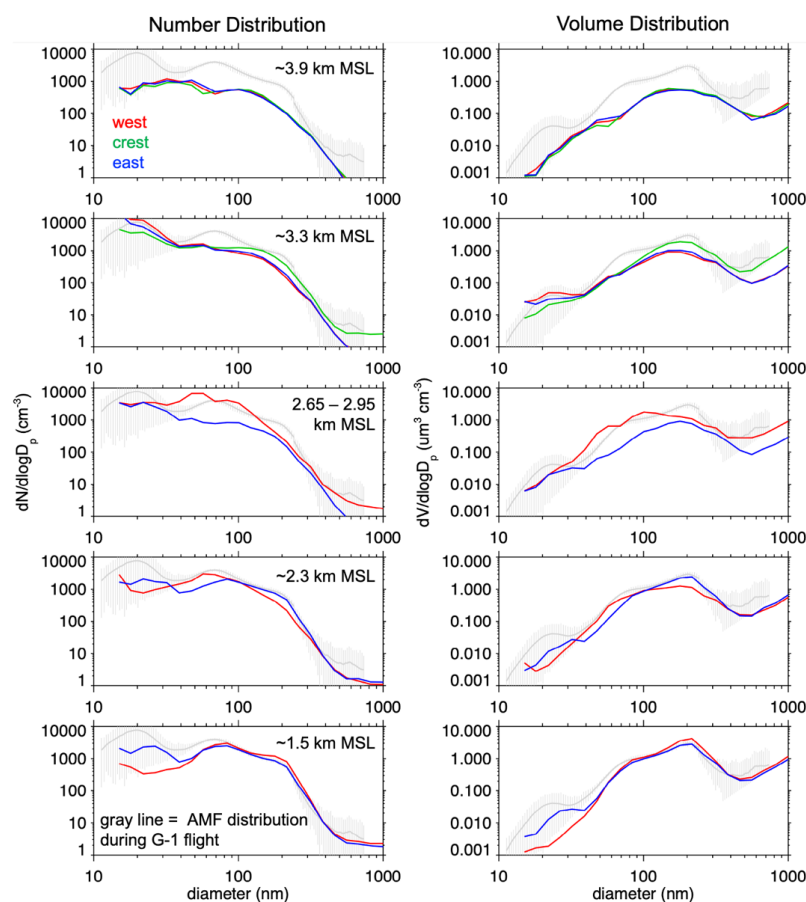
595 To examine whether the aerosol size distribution is different across the SDC, the average number and  
596 volume distributions as functions of altitude and location relative to the crest are shown in Fig. 14. For the  
597 highest aircraft constant altitude flight legs at 3.9 km MSL, the number and volume distributions west,  
598 over, and east of the crest are very similar. At 3.3 km MSL, particle number concentrations and volumes  
599 for diameters less than 40 nm are the lowest over the crest and highest west of the crest. But for particles  
600 greater than 80 nm in diameter, particle number concentrations are highest over the crest and the size  
601 distributions for the flight legs west and east of the crest are similar. For flight legs below 2.65 km MSL,  
602 there are significant differences in the number and volume distributions west and east of the crest and  
603 those differences vary with height. At the lowest flight legs at 1.5 km MSL (~0.4 km above ground east of  
604 the crest), the particle number concentrations and volumes for particles diameters greater than 80 nm is  
605 very similar to those measured at the ground AMF site. The differences for particles smaller than 80 nm  
606 indicate large spatial variability in UFP concentrations in the region. Differences between the AMF and  
607 aircraft size distributions at higher altitudes suggest that the ground measurements are not representative  
608 of conditions aloft. The spatiotemporal variability in size distribution likely contributes to variability in  
609 CCN concentrations, similar to the differences in critical diameters shown in Fig. 4.

610 To further understand the role of boundary layer growth over the crest, we next examine variations in  
611 potential temperature and vertical velocity among the 3.3 km MSL aircraft flight legs in relation to the  
612 AMF radiosonde profiles as shown in Fig. 15. Between 12 and 21 UTC, the convective boundary layer at  
613 the AMF site grows from 1.9 to 2.3 km MSL (Fig. 15a,b). At 21 UTC, the inflection of potential  
614 temperature at ~3.3 km MSL and the higher relative humidity and southwesterly winds just below that  
615 level reflect the advection of the boundary layer air from the higher terrain towards the AMF site that  
616 produces a layer of constant potential temperature between 2.3 and 3.2 km MSL. The increase in potential  
617 temperature between 15 to 18 UTC just above 3.3 km MSL is likely due to subsidence as southwesterly  
618 air is transported across the leeward side of the crest.

619 For the flight legs between 1658 and 1731 UTC at 3.3 km MSL, variability in vertical velocity (Fig. 15c)  
620 and potential temperature (Fig. 15f) is small in the free troposphere west and east of the crest. During this  
621 time period, the aircraft flew just above most of the growing clouds along the ridge (leg 6). Larger  
622 variations in both quantities are measured over the crest reflecting turbulent motions generated by clouds

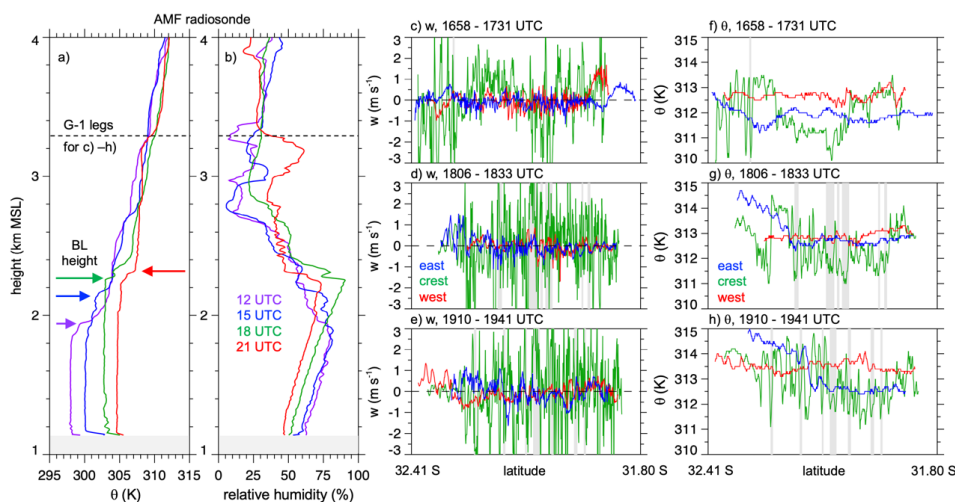


623 just above growing boundary layer and terrain variability. By the second period between 1806 and 1833  
 624 UTC, spatiotemporal variations in vertical velocity (Fig. 15d) and potential temperature (Fig. 15g)  
 625 increase over the crest as the boundary layer grew and increased the intensity of turbulent eddies. Some of  
 626 the anomalies with lower potential temperatures within clouds are due to updrafts that reduce the potential  
 627 temperatures near the boundary layer top. High-frequency variability in vertical velocity and potential  
 628 temperature remains low west and east of the crest suggesting the convective boundary layer does not  
 629 reach 3.3 MSL for those transects over lower terrain elevations. The variability during the third period  
 630 between 1910-1941 UTC (Fig. 15e and 15h) is similar to the previous period. While there are few high  
 631 frequency variations in potential temperature west and east of the crest, there are larger-scale variations  
 632 likely due to larger-scale horizontal advection. For example, the southerly winds along the southern third  
 633 of leg 12 (Fig. 12c) between 1806 and 1833 UTC coincide with both higher potential temperature and  
 634 lower particle number concentrations, suggesting a different air mass. The higher potential temperatures  
 635 progress northward by the third period between 1920-1941 UTC.



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637 Figure 14. Number and volume aerosol distributions at five altitudes that are west, over, and east of the  
 638 Sierras de Córdoba crest on December 3. Gray line denotes average ground measurements  
 639 during the aircraft flight.

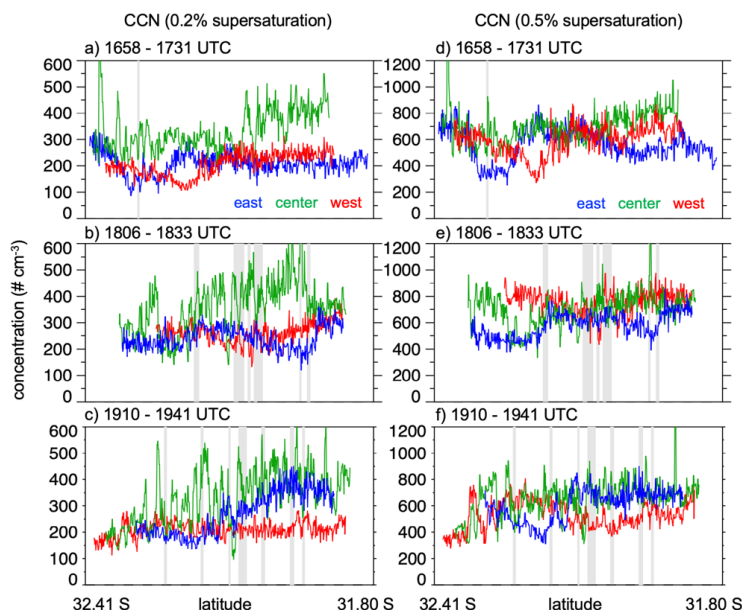


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642 Figure 15. Radiosonde profiles of a) potential temperature and b) relative humidity at the AMF site on  
643 December 3 along with the spatial variations in vertical velocity c) – e) and potential  
644 temperature f) – h) for G-1 flight legs at ~3.3 km MSL divided transects that are west, over,  
645 and east of the Sierras de Córdoba crest. Gray shading in c) – h) denotes aircraft sampling  
646 within clouds.

647 CCN concentrations shown in Fig. 16 further illustrate spatial variabilities within and among the legs  
648 west, over, and east of the mountain range crest. During the first period between 1658-1731 UTC, CCN  
649 concentrations at 0.2% supersaturation are higher over the crest than those along the flight legs west and  
650 east of the crest (Fig. 16a). There are also larger spatial fluctuations along portions over the crest that  
651 might be tied to the larger vertical velocity variations (Fig. 15c). High frequency fluctuations in CCN  
652 increase as the boundary layer grows over the crest and envelopes the flight leg over the crest between  
653 1806-1833 UTC (Fig. 16b). CCN concentrations at 0.2% supersaturation are still the highest with the  
654 largest amount of spatial variability during the third period between 1910-1941 UTC (Fig. 16c). In the  
655 free troposphere, the overall CCN concentrations west of the crest remain stable while there are large-  
656 scale variations east of the crest. CCN at 0.5% supersaturation still exhibits differences between the legs  
657 west, over, and east of the crest; however, CCN concentrations are not consistently higher or lower across  
658 the crest during the three time periods (Figs. 16d – 16f). CCN concentrations are also lower along the  
659 southern third of legs 13 and 19 east of the crest, consistent with the lower particle number concentrations  
660 transported northward along legs 13 and 19 during the second and third time periods.

661 The previous figures demonstrate that local meteorological processes affect the variability in observed  
662 aerosol properties. Back trajectories are used next in Fig. 17 to illustrate long-range transport pathways  
663 and possible sources of aerosols transported over the AMF site on December 3. The HYSPLIT model  
664 (Rolph et al., 2017; Stein et al., 2015) is used to compute 4-day back trajectories originating over the  
665 AMF site at 1.5, 2.3, 2.8, and 3.3 km MSL levels and at hourly intervals between 16 and 20 UTC during  
666 the G-1 aircraft flight period. Back trajectories are also computed at the corners of a 1-degree wide box  
667 around the AMF site at the same altitudes and times. The trajectories are driven by winds from the  
668 National Center for Environmental Prediction’s Global Data Assimilation System at 0.5-degree grid  
669 spacing.



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Figure 16. Spatial variations in 0.2% CCN a) – c) and 0.5% CCN d) – f) for G-1 flight legs at ~3.3 km MSL that are west, over, and east of the Sierras de Córdoba crest. Gray shading denotes aircraft sampling within clouds.

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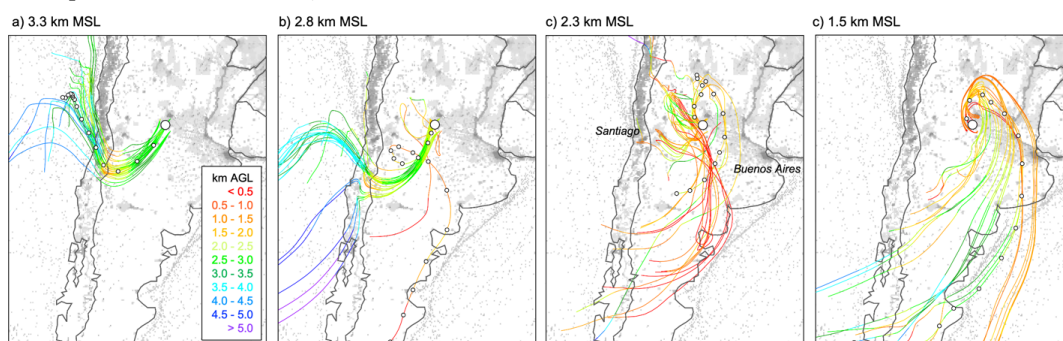
Back trajectories arriving over the AMF site at 3.3 km MSL (Fig. 17a) suggest that air in the lower free troposphere over the Pacific Ocean is transported by northerly to northwesterly winds toward the coast of Chile. As the trajectories pass over the Andes, they become closer to the terrain (within 1.5 km MSL). It is possible that anthropogenic aerosols and aerosol precursors emitted over Chile and transported by upslope flows could be mixed with this air mass over the Andes. Westerly winds then transport these aerosol and precursors over the Andes into Argentina and are subsequently transported by southwesterly winds towards the AMF site. A similar transport pattern is produced for back trajectories originating at 2.8 km MSL (Fig. 17b), except that the air pass over Chile further to the south and farther from the largest anthropogenic emissions. Some trajectories suggest that low emission regions over Argentina could be lifted to that altitude. In contrast, back trajectories starting at 2.3 km MSL near the top of the growing boundary layer are more complex (Fig. 17c). Some trajectories passing over the low anthropogenic emission regions of southern Argentina are transported over the eastern slopes of the Andes where they are lofted to higher altitudes and then transported by westerly winds to the AMF site. Other trajectories exhibit a counterclockwise circulation so that air from southern Argentina is transported east and north of the AMF site northerly winds finally transport the air masses to over the AMF site. Finally for trajectories arriving near the surface at 1.5 km MSL (Fig. 17d), almost all the back trajectories exhibit the counterclockwise circulation, suggesting that the lower atmosphere over the AMF is a mixture of cleaner air passing over southern Argentina and higher concentrations of aerosols originating from more populated regions along the Paraná River between Sante Fe and Buenos Aires.

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Even though back trajectories do not pass over the Amazon, miniSPLAT measurements on this day suggest that ~14% of the particles originate from biomass burning at all altitudes (Fig. S10). This is nearly the same percentage as on November 12 when smoke was transported from more distant sources in the Amazon. However, the combination of back trajectories, wind directions (Fig. S7), and low CO concentrations (Figs. 11f), suggest that biomass burning from the Amazon did not contribute significantly to aerosols on this day. Fire emissions inventories indicate smaller fires occurred over eastern Argentina



700 and along the coast of Chile three days prior to December 3 (not shown); therefore, the source of biomass  
 701 burning aerosols over the AMF site was different on November 12 and December 3. The transport time  
 702 also suggests that primary and secondary aerosols originating from the populated regions over eastern  
 703 Argentina would be aged by up to 1.5 days. In addition, the types of organic mixtures for these days were  
 704 very different. On December 3, the fraction of particle class 10 containing oxygenated organics and  
 705 sulfate mixtures was far less, particle class 11 containing the highest ratio of sulfate to organics was  
 706 higher on December 3, and particle class 14 containing IEPOX SOA was very small relative to November  
 707 12. While many of the back trajectories passed over southern Argentina, dust contributed to less than 1%  
 708 of the particles smaller than 1  $\mu\text{m}$  in diameter on December 3.



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710 Figure 17. Back trajectories originating at a) 3.3 km MSL, b) 2.8 km MSL, c) 2.3 km MSL, and d) 1.5  
 711 km MSL over the AMF site on December 3 during the G-1 flight between 16 and 20 UTC.  
 712 Dots denote 6-h periods along one trajectory for each height.

713 Figure 17 illustrates the impact of vertical wind shear and mountain venting processes on the transport of  
 714 aerosols and their precursors. The back trajectories are consistent with AMF radiosonde wind directions  
 715 that are northerly in the boundary layer, westerly near the boundary layer top, and southwesterly in the  
 716 lower free troposphere (Fig. S7). Mountain venting of aerosols and precursors into the lower free  
 717 troposphere results from combination of upslope flows and detrainment from the boundary layer over the  
 718 highest terrain (De Wekker and Kossmann, 2015) that subsequently produce layers in the lower free  
 719 troposphere transported far downwind. Mountain venting processes and their impact on pollutant  
 720 transport has been observed in many regions of the world, including the Alps (e.g. Henne et al., 2005;  
 721 Kossmann et al., 1999; Nyekel et al., 2000), central Mexico (Fast and Zhong, 1998), and California (e.g.  
 722 Fast et al., 2014; Lu and Turco, 1994; Langford et al., 2010), and Andes (e.g. Lopere et al., 2021). It is  
 723 important to note that the 0.5-degree grid spacing in GDAS is not likely to resolve local terrain driven  
 724 circulations and around the SDC and the Andes, which introduces uncertainties in these trajectories.

#### 725 4. Summary and Conclusions

726 In situ measurements of aerosol properties are needed to evaluate and improve air quality, chemical  
 727 transport, and climate model predictions and to better understand complex aerosol-cloud interaction  
 728 processes. While surface monitoring networks and aircraft field campaigns have collected aerosol  
 729 measurements in the northern hemisphere over the past several decades, few field campaigns with  
 730 extensive aerosol measurements have been conducted over subtropical and midlatitude continental areas  
 731 in the southern hemisphere. This study analyzes a wide range of surface and aircraft measurements  
 732 collected over a seven-week period during the recent CACTI field campaign in central Argentina.

733 CACTI surface observations show large multi-day variations in aerosol number, mass, composition, and  
 734 size distribution. On average, PM<sub>1</sub> aerosol mass obtained from the ACSM instrument ( $3.7 \mu\text{g m}^{-3}$ ) was  
 735 comprised of 53% OM, 29% SO<sub>4</sub>, 8% NO<sub>3</sub>, and 8% NH<sub>4</sub>; however, the fraction of SO<sub>4</sub> was significantly  
 736 larger when PM<sub>1</sub> mass was low ( $< 1 \mu\text{g m}^{-3}$ ). While multiday trends in PM<sub>1</sub> and AOD are qualitatively



737 similar and suggest most aerosol mass occurs in the boundary layer, on some days these trends differ over  
738 several hours indicating that aerosol layers in the free troposphere can significantly contribute to column  
739 extinction. As expected, days with the highest precipitation (October 25 - 26, November 11 - 12, and  
740 November 26 - 27) have the lowest PM<sub>1</sub> mass concentrations due to wet removal and/or changing  
741 regional air masses. These rainy days divide the campaign into three periods: period A between October  
742 27 and November 10 with the highest aerosol mass and lowest aerosol number concentrations, period B  
743 between November 13 and 25 with the lowest aerosol mass and highest aerosol number concentrations,  
744 and period C after November 28 with mass concentrations similar to period A but with aerosol number  
745 concentrations between those from periods A and B. The high ultrafine particle concentrations during  
746 period B, and to a lesser extent period C, suggest that new particle formation occurred over or upwind of  
747 the AMF site. Average size distributions are also different among these three periods. All these  
748 measurements suggest that changing mesoscale to synoptic-scale meteorology alters transport patterns as  
749 well as local aerosol formation and growth processes.

750 Diurnal variation in aerosol composition, number, and CCN can be attributed to local meteorological and  
751 chemical processes. While SO<sub>4</sub> and NH<sub>4</sub> showed no diurnal variability on average, peak OM, NO<sub>3</sub>, and  
752 rBC concentrations occurred during mid-afternoon, around sunrise, and at night, respectively. The daily  
753 peak in OM is likely due to photochemistry associated with SOA and the daily peak of NO<sub>3</sub> at sunrise  
754 when the temperatures are the coldest likely inhibits partitioning of NO<sub>3</sub> to gas phase. Particles greater  
755 than 200 nm in diameter exhibit no diurnal variations which is consistent with the weak diurnal variability  
756 in PM<sub>1</sub> mass. Conversely, smaller particles exhibit diurnal variations with a peak at ~20 UTC (17 LT),  
757 and the strongest diurnal variations are for the smallest particles less than 10 nm in diameter. No diurnal  
758 variability in CCN at 0.1% supersaturation was observed, consistent with little diurnal variability in  
759 accumulation mode aerosols. CCN at higher supersaturations have diurnal variations with peak  
760 concentrations at 22 UTC reflecting the growth of aerosols during the afternoon.

761 Potential sources of aerosols can be determined by correlating aerosol composition with trace gases and  
762 wind directions. Since rBC and CO are often co-emitted, observed rBC and CO exhibits similar diurnal  
763 variability and are temporally correlated ( $r = 0.59$ ). The concentrations of both quantities are largest  
764 during northerly winds, followed by northeasterly, with the lowest concentrations during southerly winds.  
765 OM concentrations are largest for northerly and northeasterly winds with somewhat lower concentrations  
766 during southerly winds. O<sub>3</sub> is the largest during northeasterly winds and therefore may originate from  
767 anthropogenic emissions in Córdoba. The CAM-Chem global chemical transport model and single  
768 particle instrument data show that biomass burning aerosols from the Amazon are frequently transported  
769 by the South American low-level jet to the AMF region. Single particle measurements show that ~12% of  
770 particles during one of the transport events are from biomass burning particles. Aged biogenic OM may  
771 also be transported by the same winds. While SO<sub>4</sub> is 29% of the PM<sub>1</sub> mass on average at the ground, it is  
772 important to note that single particle measurements reveal it is mixed with organics at various ratios,  
773 illustrating a more complex mixing state than from bulk measurements.

774 The aircraft measurements show that the largest total aerosol number concentrations usually occur at the  
775 surface and decrease by an order of magnitude by 5 km MSL, on average. Statistics of all the flights  
776 illustrate that aerosol number concentrations vary significantly (order of magnitude or more for 5<sup>th</sup> to 95<sup>th</sup>  
777 percentiles) both temporally (multi-day and within 4-h flight periods) and horizontally within 50 km of  
778 the AMF site at all altitudes. In contrast, average CCN concentrations remain relatively constant from the  
779 surface up to 3.5 km MSL and then gradually decrease to small values by 5 km MSL. The percentiles of  
780 CCN also show temporal and horizontal variability (factor of ~2 for 5<sup>th</sup> to 95<sup>th</sup> percentiles), but that  
781 variability becomes small above 5 km MSL.

782 Since the aircraft data reveals large variations in aerosol properties aloft on many days, we focus on the  
783 December 3 measurements as a case study. On this day, repeated constant altitude flights reveal large  
784 spatial gradients in aerosol number and CCN concentrations that change from hour to hour. Two small  
785 plumes of SO<sub>2</sub> with high aerosol number concentrations were observed; however, it is possible that the





876 aircraft could easily miss other small, transient plumes even with repeated patterns at multiple altitudes.  
877 While the aerosol size distribution from the lowest flight transects east of the mountain crest and within  
878 the boundary layer was similar to the ground measurements, it changed with height and was also different  
879 west, over, and east of the crest of the Sierras de Córdoba range. Some flight tracks over the mountain  
880 crest occurred just as the growing boundary layer intersected those altitudes, illustrating the effects of  
881 aerosol entrainment across the top of the boundary layer. It is possible that cloud processing also affects  
882 the size, composition, and hygroscopicity of aerosols over the mountain crest, but further analysis is  
883 needed to examine this process.

884 Strong vertical wind shear at the AMF site results in transport pathways of aerosol sources that vary with  
885 height. Back trajectories on December 3 indicate that aerosols within the boundary layer likely originated  
886 from the more populous regions of eastern Argentina with anthropogenic and biomass burning  
887 contributions. At higher altitudes mountain venting processes could produce lofting of aerosols and  
888 aerosol precursors emitted from Chile or western Argentina on either side of the Andes, that is  
889 subsequently transported by westerly winds over the AMF site. A similar, smaller scale process may  
890 operate along the Sierras de Córdoba crest since the aircraft measurements indicate aerosols entrained into  
891 the free troposphere are transported eastward over the AMF site.

892 In addition to quantifying aerosol properties in a data sparse region, the aerosol property measurements  
893 presented in this study will be valuable to evaluate predictions over the mid latitudes of South America  
894 and improve parameterized aerosol processes in local, regional, and global models. For aerosol-cloud  
895 interaction studies, the measurements clearly show that accounting for the co-variabilities of aerosol  
896 properties and convective cloud populations over the Sierras de Córdoba range will be critical. Knowing  
897 aerosol properties just below and surrounding clouds is important because aerosols are entrained into the  
898 base of clouds by convective updrafts. In addition to cloud base entrainment, aerosols with different  
899 properties at higher altitudes are also entrained into the sides of clouds as they grow vertically. As clouds  
900 evaporate, the size and composition of aerosol particles that were within cloud droplets can be different  
901 than aerosol populations surrounding clouds because of cloud chemistry and coalescence of cloud  
902 droplets. Thus, the properties of aerosol populations near the top of the boundary layer and around clouds  
903 will change in time as aerosols are cycled through clouds multiple times over several hours. Given the  
904 large observed variations in both aerosol and cloud properties and complexity of their interactions, it will  
905 be challenging to develop a robust statistical signal of aerosol-cloud interactions from all the  
906 measurements. Thus, studying impacts of aerosols on cloud properties and impacts of clouds on aerosol  
907 properties on a case-by-case basis will provide critical insights.

#### 898 **Data Availability**

899 In addition to citations given in the text, all CACTI data are available through links provided at  
900 [www.arm.gov/research/campaigns/amf2018cacti](http://www.arm.gov/research/campaigns/amf2018cacti). Citations for individual datasets are provided  
901 throughout the text.

#### 902 **Author Contributions**

903 JF performed the data analyses and wrote the manuscript. AC was the PI of the CACTI campaign,  
904 assisted in the design and interpretation of the data analyses, and contributed to the manuscript. FM, MP,  
905 JT, AZ, AS, and MZ were instrument mentors during CACTI and provided comments on the manuscript  
906 when they are first mentioned.

#### 907 **Competing Interests**

908 The contact author has declared that none of the authors has any competing interests.

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