



# 1 **Stratospheric gravity waves excited by Hurricane Joaquin in 2015:** 2 **3-D characteristics and the correlation with hurricane intensification**

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## 16 **Abstract.**

17 Despite progress, accurately forecasting tropical cyclone (TC) intensity, especially rapid intensification, remains a  
18 significant challenge. The correlations between the stratospheric gravity waves (GWs) excited by TCs and TC intensity have  
19 been recognized. However, partly due to the limitations of conventional analysis methods and observational filters of current  
20 satellite instruments, the characteristics of stratospheric GWs that indicate TC intensification remain unclear. This study  
21 examined the specific characteristics of GWs and their linkage to hurricane intensification by high-resolution, realistic model  
22 simulations and 3-D wave analysis method. First, the stratospheric GWs excited by Hurricane Joaquin in 2015 were simulated  
23 using the Advanced Weather Research and Forecasting (WRF) model. Then, the GW characteristics were analyzed using the  
24 novel 3-D Stockwell transform method. The GWs excited by Hurricane Joaquin are in the mid-frequency range and propagate  
25 outward from the hurricane center counterclockwise while moving upward in a spiral. A high-level time-lagged correlation  
26 exists between the intensities of the hurricane and stratospheric GWs during hurricane intensification, making it possible to  
27 detect an increase in hurricane intensity by observing an increase in stratospheric GW intensities. Compared to the weakening  
28 period, the stratospheric GWs excited during hurricane intensification exhibit relatively higher frequencies, shorter horizontal  
29 wavelengths, and longer vertical wavelengths, with this contrast particularly evident near the center of the hurricane. This  
30 study provides further knowledge for potentially monitoring hurricane intensification by observing stratospheric GWs using  
31 satellite instruments in the infrared and microwave bands when it is difficult to use other measurement techniques.



## 32 **1 Introduction**

33 Tropical cyclones (TCs) are hazardous meteorological events that consistently result in loss of life and property. Despite  
34 a substantial reduction in track forecast errors over the past decade, there has been only a modest decrease in the forecast error  
35 of intensity (DeMaria et al., 2014). Particularly, predicting abrupt changes in TC intensity, such as rapid intensification (RI),  
36 remains challenging and is rarely successful (Cangialosi et al., 2020). In recent decades, with climate change, TCs have become  
37 more intense (Emanuel, 2017; Kang and Elsner, 2019), the frequency of TCs undergoing RI has increased (Bhatia et al., 2019;  
38 Balaguru et al., 2018), and TCs that experience rapid intensification have a higher probability of developing into stronger and  
39 potentially destructive TCs (Lee et al., 2016). While many studies have investigated changes in TC structure and the influence  
40 of environmental conditions before TC intensification (e.g., Chen et al., 2011; Kieu et al., 2014; Onderlinde and Nolan, 2014;  
41 Kaplan et al., 2015; Rogers et al., 2017), accurate forecasting of TC intensity changes still faces significant challenges (Doyle  
42 et al., 2017).

43 In addition to being catastrophic weather systems, tropical cyclones also serve as sources of excitation for stratospheric  
44 gravity waves (GWs) (Hoffmann et al., 2013). Oscillating drafts and thermal forcing due to latent heat release in TCs are two  
45 primary mechanisms for generating GWs (Beres et al., 2002). While GWs excited in the upper troposphere rapidly propagate  
46 from the source upward to the stratosphere and even higher levels, the wave amplitudes grow and become even more  
47 prominent (Liu et al., 2014). The characteristics of stratospheric GWs excited by TCs (TC-GWs hereafter) may reflect  
48 changes in TC intensity as the two mechanisms that excite GWs are also associated with changes in TC intensity. First,  
49 extremely strong updrafts (also known as convective bursts) typically intensify in the eyewall up to a few hours before TC  
50 intensification (Wang and Wang, 2014; Hazelton et al., 2017). Second, this process leads to deep latent heating in the upper  
51 troposphere and the formation of a warm core (Ohno and Satoh, 2015); the balanced dynamics of TC structure and flows  
52 towards thermal forcing in the eyewall contribute to TC intensification (Fudeyasu and Wang, 2011; Wang and Wang, 2014).

53 The connections between the stratospheric GWs and the TC intensities are found by idealized model simulations (Nolan  
54 and Zhang, 2017) and verified by satellite observations. Hoffmann et al. (2018) utilized approximately 14 years of  
55 Atmospheric Infrared Sounder (AIRS) observations of TC-GWs to investigate the correlation between TC-GWs and TC  
56 intensity. The study discovered a significant correlation, indicating that TC-GWs are more intense and active during TC  
57 intensification. The frequency of TC-GWs during TC intensification was twice as high as during TC weakening. Using 16  
58 years of multiple sources of satellite observations, Wright (2019) observed a similar phenomenon: GW amplitudes increase  
59 before the peak intensity of TCs, followed by a sudden decrease afterward. Due to the "observational filter" effect, none of  
60 the presently available instruments can comprehensively detect the entire spectrum of GWs (Alexander and Barnet, 2007;  
61 Wright et al., 2015). Therefore, it is essential to further investigate the connections between the stratospheric GWs and the  
62 TC intensities with realistic numerical simulations. Wu et al. (2022) conducted high-resolution realistic model simulations



63 focusing on a specific case, the Hurricane Joaquin, which is also the subject of the present study. They verified the  
64 significant correlations between TC-GWs and TC intensity that revealed by theoretical model simulations and satellite  
65 observations, and found that GW activity is more frequent and intense during hurricane intensification, particularly for the  
66 most intense GW activity.

67 Observing stratospheric GWs offers an advantageous method for inferring changes in TC intensity, particularly when  
68 cloud canopies obscure the TC eye and eyewall from instruments using visible and infrared bands. Motivated by the critical  
69 need to monitor and accurately predict TC intensification, recent research has increasingly focused on using satellite  
70 instruments in the infrared and microwave bands for observing TC-GWs, thereby enabling more precise inferences about TC  
71 intensity evolution (Miller et al., 2018; Tratt et al., 2018). However, the mechanism that cause the correlation between TC-  
72 GWs activity and TC intensification is not yet thoroughly understood, and the specific characteristics, such as wavelengths of  
73 the TC-GWs that indicate TC intensification are not clear.

74 Building upon Wu et al. (2022), this study extends beyond their emphasis on validating the correlations between the  
75 intensities of stratospheric GWs and the hurricane. It delves deeper into examining the characteristics of GWs, especially their  
76 linkage to the intensification phases of the hurricane. First, we conducted high-resolution numerical simulations of a  
77 challenging case, Hurricane Joaquin (2015). Hurricane Joaquin is a representative case of difficulties in hurricane intensity  
78 forecasts. The official intensity forecast errors for Hurricane Joaquin exceeded the mean official errors during the preceding  
79 five-year period at all forecast intervals (Berg, 2016). Then, our study examines the 3-D characteristics of the simulated TC-  
80 GWs. The characteristics of TC-GWs have been widely analyzed based on numerical simulations and observations (e.g., Kim  
81 and Chun, 2010; Chen et al., 2012; Jewtoukoff et al., 2013; Chane Ming et al., 2014; Yue et al., 2014; Wu et al., 2015).  
82 Nevertheless, the conventional 1-D and 2-D wave analysis method can only reveal the wavelengths and frequency distribution  
83 in a domain but lacks the ability to pinpoint the specific location or timing of distinct characteristics. In this study, a novel 3-  
84 D Stockwell transform method is employed to estimate the characteristics of the stratospheric GWs on 3-D grid points, namely  
85 the dominant wave frequencies, wavelengths, wave speed, and propagation directions. By investigating the features of the  
86 GWs and their connection with the intensification of Hurricane Joaquin, this study will identify the key characteristics and  
87 locations of the stratospheric TC-GWs that indicate hurricane intensification. The goal is to lay the groundwork for monitoring  
88 TC intensification through satellite observations of stratospheric GWs.

89 The remaining sections are organized as follows: Sections 2 and 3 describe the data and model configuration; Sect. 4  
90 introduces the novel 3-D Stockwell transform method; Sect. 5 presents the results; Sect. 6 discusses the results and gives the  
91 conclusions.



92 **2 Data**

93 **2.1 ERA5 reanalysis**

94 The initial and boundary conditions of the WRF simulation of Hurricane Joaquin were obtained from the ERA5 reanalysis.  
95 The ERA5 reanalysis (Hersbach et al., 2020), a fifth-generation product from the European Centre for Medium-Range Weather  
96 Forecasts (ECMWF), covers the period from 1940 to the present. The dataset is accessible at 137 vertical hybrid sigma-pressure  
97 levels, with the highest level at 0.01 hPa (approximately 80 km altitude) and at the surface level. For our simulations, we  
98 utilized hourly data with a horizontal sampling of  $0.25^{\circ} \times 0.25^{\circ}$  to establish initial and boundary conditions.

99 The ERA5 reanalysis demonstrates an enhanced representation of tropical cyclones (Hodges et al., 2017; Li et al., 2020)  
100 compared to the preceding ECMWF reanalysis. It provides improved resolution of convection over both oceanic and  
101 continental regions (Hoffmann et al., 2019). This enhancement significantly contributes to the precision of the hurricane  
102 intensity simulation.

103 **2.2 Tropical cyclone track and intensity archive**

104 This study utilized the International Best Track Archive for Climate Stewardship (IBTrACS) dataset (Knapp et al., 2010)  
105 to evaluate the track and intensity of Hurricane Joaquin in the model simulation. The dataset is constructed through subjective  
106 satellite-based Dvorak technique intensity assessments from the Tropical Analysis and Forecast Branch (TAFB) and the  
107 Satellite Analysis Branch (SAB), in conjunction with the objective Advanced Dvorak Technique (ADT). It provides the TC  
108 track and intensity estimates in terms of minimum sea level pressure and maximum sustained winds at intervals of 3 to 6 hours.

109 **3 WRF Model configuration**

110 The numerical simulation was performed using Version 3.9.1 of the Advanced Weather Research and Forecasting (WRF)  
111 model (Skamarock et al., 2008). The model configuration mainly follows the setup of Wu et al. (2022). The simulation adopted  
112 a concurrent one-way nested configuration comprising a stationary outer domain (D01) and an inner nested vortex-following  
113 domain (D02). The grid resolutions for D01 and D02 were 12 km and 4 km, respectively, with corresponding time steps of 12  
114 s and 4 s. D01 provided boundary conditions for D02, and no feedback from D02 to D01 was incorporated. The vertical domain  
115 covered 90 sigma levels from the surface to 1 hPa (~48 km). Vertical resolution exhibited a gradient, with the finest resolution  
116 near the surface gradually decreasing to approximately 500 m from the tropopause and above. A damping layer was  
117 implemented in the uppermost 5 km. The simulation spanned 100 hours, from 00 UTC on 30 September to 04 UTC on 4  
118 October 2015, with outputs recorded at 6-minute intervals.

119 The initial and boundary conditions of the simulations, including sea surface temperature, were derived from the ERA5  
120 reanalysis. In both D01 and D02, the following model physics schemes were selected: 1) the Kain-Fritsch (KF) convective



121 scheme (Kain, 2004) for cumulus parameterization; 2) the WRF single moment 6-class (WSM6) scheme (Hong and Lim,  
122 2006) for microphysics; 3) the updated version of the rapid radiative transfer model scheme (RRTMG) (Iacono et al., 2008)  
123 for longwave and shortwave radiation; and 4) the Yonsei University (YSU) planetary boundary layer scheme (Hong et al.,  
124 2006) for the vertical diffusion process. While cumulus parameterizations are theoretically recommended primarily for grid  
125 sizes exceeding 10 km (Skamarock et al., 2008), these schemes also initiate convection at grid sizes smaller than 10 km  
126 (Nasrollahi et al., 2012; Li and Pu, 2009). In the case study of Joaquin, the KF scheme was also applied to the inner domain  
127 to ensure that the simulated hurricane intensity aligned with the IBTrACS dataset.

## 128 4 3-D Stockwell transform (S-transform) wave spectral analysis

### 129 4.1 The extended S-transform method

130 The S-transform, as proposed by Stockwell et al. (1996), is a spectral analysis technique that localizes wave perturbations  
131 in the spatial domain through spectral localization in the frequency domain. The S-transform achieves this by applying a  
132 scalable localizing Gaussian window to the short-time Fourier transform, and it can be applied to any time series or distance  
133 profile to provide localized measurements of wave properties.

134 For a continuous one-dimensional function of time or distance  $h(x)$ , the generalized S-transform  $S(\tau, f)$  is given as:

$$135 S(\tau, f) = \int_{-\infty}^{\infty} h(x) \omega_g(x - \tau, f) e^{-i2\pi f x} dx, \quad (1)$$

136 where  $\tau$  represents the translation in the time (spatial) domain,  $f$  is the frequency or wavenumber, and  $\omega_g(x - \tau, f)$  is the  
137 normalized Gaussian window:

$$138 \omega_g(x - \tau, f) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{(x-\tau)^2}{2\sigma^2}}, \quad (2)$$

139 In Eq. (2), the normalization factor  $\frac{1}{\sigma\sqrt{2\pi}}$  ensures that the integral of the window function equals unity, a prerequisite for  
140 the windowing function employed in the S-transform. The standard deviation, denoted as  $\sigma$ , is scaled for each frequency using  
141 the formula  $\sigma = \frac{c}{|f|}$ , where  $c$  represents a scaling parameter.

142 Substituting Eq. (2) into Eq. (1) gives the explicit S-transform:

$$143 S(\tau, f) = \frac{|f|}{c\sqrt{2\pi}} \int_{-\infty}^{\infty} h(x) e^{-\frac{(x-\tau)^2 f^2}{2c^2}} e^{-i2\pi f x} dx, \quad (3)$$

144 The one-dimensional S-transform in Eq. (3) could be extended to three dimensions (Hindley et al., 2019) to localize  
145 wavelengths and directions at every grid point in the 3-D WRF outputs. For function  $h(x)$ , where  $x = (x_1, x_2, x_3)$  is a column  
146 vector describing a 3-D coordinate system,  $S(\tau, f)$  can be written as:

$$147 S(\tau, f) = \frac{1}{(2\pi)^{3/2}} \int_{-\infty}^{\infty} h(x) \left[ \prod_{n=1}^N \frac{|f_{n1}|}{c_n} e^{-\frac{(x_n - \tau_n)^2 f_{n1}^2}{2c_n^2}} \right] e^{-i2\pi f^T x} dx, \quad (4)$$



148 Here,  $\tau = (\tau_1, \tau_2, \tau_3)$  and  $f = (f_1, f_2, f_3)$  are column vectors representing spatial translations and spatial frequencies  
149 (inverse of wavelength) in the  $x_1, x_2, x_3$  directions, and the  $f^T$  denotes the transposed  $f$ . The scaling factors  $c_n = (c_1, c_2, c_3)$   
150 are tuned for each of the three dimensions independently to emphasize different wave properties (Hindley et al., 2016). The S-  
151 transform is used for spectral analysis of gravity waves in diverse geophysical datasets (e.g., Hindley et al., 2016; Wright et  
152 al., 2017; Hindley et al., 2019; Hindley et al., 2020). Here we apply the 3-D S-transform approach of Hindley et al., (2019)  
153 who provide a full description of its implementation in their Sect. 3.

#### 154 4.2 Wave properties derived by the 3-D S-transform

155 For each WRF output on a 3-D grid  $(x, y, z)$ , the S-transform in Eq. (4) produces  $S(\tau, f) \equiv S(\tau_x, \tau_y, \tau_z, f_x, f_y, f_z)$ . We  
156 then collapsed this 6-D object into a 3-D object by only considering the peak of the localized  $(f_x, f_y, f_z)$  spectrum for each  
157 location. To do this, for each location in  $(\tau_x, \tau_y, \tau_z)$ , we find the peak spectral amplitude in the localized spectrum  $(f_x, f_y, f_z)$ ,  
158 which corresponds to the largest amplitude wave at this location. The location of this peak in  $(f_x, f_y, f_z)$  tells us the dominant  
159 frequencies, which we record as  $F_x(\tau_x, \tau_y, \tau_z)$ ,  $F_y(\tau_x, \tau_y, \tau_z)$ , and  $F_z(\tau_x, \tau_y, \tau_z)$ , respectively.

160 In our application, the  $(\tau_x, \tau_y, \tau_z)$  domain corresponded to the regular grids  $(x, y, z)$  of the WRF outputs, i.e.,  
161  $(\tau_x, \tau_y, \tau_z) = (x, y, z)$ , so we got the spatial frequencies  $f_x \equiv F_x(x, y, z)$ ,  $f_y \equiv F_y(x, y, z)$ ,  $f_z \equiv F_z(x, y, z)$  which are the  
162 inverse of wavelength. Because our WRF grid is cartesian with axes aligned in the zonal, meridional, and vertical directions,  
163  $f_x, f_y$ , and  $f_z$  are simply the zonal, meridional, and vertical wavenumbers  $k, l$ , and  $m$ , respectively.

164 Then, the horizontal wavelength is:

$$165 \quad \lambda_H = \frac{1}{\sqrt{k^2 + l^2}}, \quad (5)$$

166 and the vertical wavelength is:

$$167 \quad \lambda_Z = \frac{1}{m}, \quad (6)$$

168 In Eq. (4), the  $c_n$  values of 0.25, 0.5, 0.75, and 1.0 were tested. We empirically selected  $c_n = (c_x, c_y, c_z) = (0.5, 0.5, 0.5)$ ,  
169 which is a compromise between spatial and spectral resolution. In our application, the spatial domain convolution in the S-  
170 transform equation in Eqn. 4 is computed as a frequency-domain multiplication using fast discrete Fourier transform for  
171 efficient computation. As a result, the computed outcomes are discrete.

172 By simultaneously characterizing the wavenumber  $k, l$ , and  $m$ , we estimated GWs intrinsic frequencies  $\omega$  and phase  
173 speed  $c_{ph}$ . The GWs intrinsic frequency  $\omega$  is calculated from the GW dispersion relation (Fritts and Alexander, 2003):

$$174 \quad \omega \equiv \sqrt{\frac{N_B^2(k^2 + l^2) + f^2(m^2 + 1/4H^2)}{k^2 + l^2 + m^2 + 1/4H^2}}, \quad (7)$$

175 where  $N_B$  is the Brunt-Väisälä frequency,  $H$  is the scale height ( $\sim 7$  km in the stratosphere).  $f = 2\Omega \sin(\phi)$  is the Coriolis  
176 frequency, where  $\Omega$  is the Earth's rotation rate and  $\phi$  is the latitude.

177 GW intrinsic phase speed  $c_{ph}$  is given by:

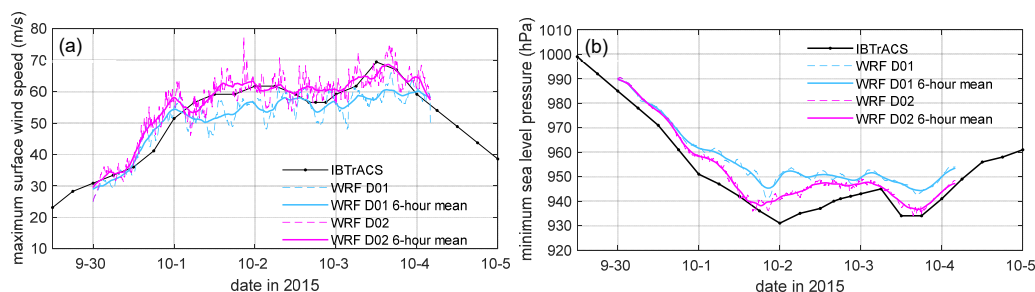


178 
$$c_{ph} = \frac{\omega}{k^2+l^2+m^2}(k, l, m) . \tag{8}$$

179 **5 Results**

180 **5.1 WRF simulations of GWs excited by Hurricane Joaquin (2015)**

181 The accurate simulation of hurricane intensity is crucial for simulating and analyzing TC-GWs. Figure 1 compares the  
 182 simulation results from D01 and D02 with the IBTrACS dataset. Consistent with prior studies (e.g., Jin et al., 2014; Wu et al.,  
 183 2018), Hurricane Joaquin's intensity is sensitive to horizontal grid spacing, and the hurricane strengthens as grid spacing  
 184 decreases. The maximum surface wind speed (MSFCW) of D01 (12-km grid) underestimated the intensity from 00 UTC, 1  
 185 October, and failed to replicate the secondary intensification from 00 UTC, 3 October. The minimum sea level pressure (MSLP)  
 186 of D01 also underestimated the intensity but could follow the trend of intensity change.



187 **Figure 1.** (a) Comparison of maximum sustained wind speed (MSW) from IBTrACS with maximum surface wind speed (MSFCW) from  
 188 WRF simulation of the outer domain (D01, in blue) and inner domain (D02, in magenta). (b) Comparison of minimum sea level pressure  
 189 (MSLP) from IBTrACS with MSLP from WRF simulation of D01 and D02. The simulation outputs every 6 minutes are indicated with thin  
 190 dashed lines, and the 6-hour running mean of the outputs every 6 minutes is shown with thick solid lines.  
 191

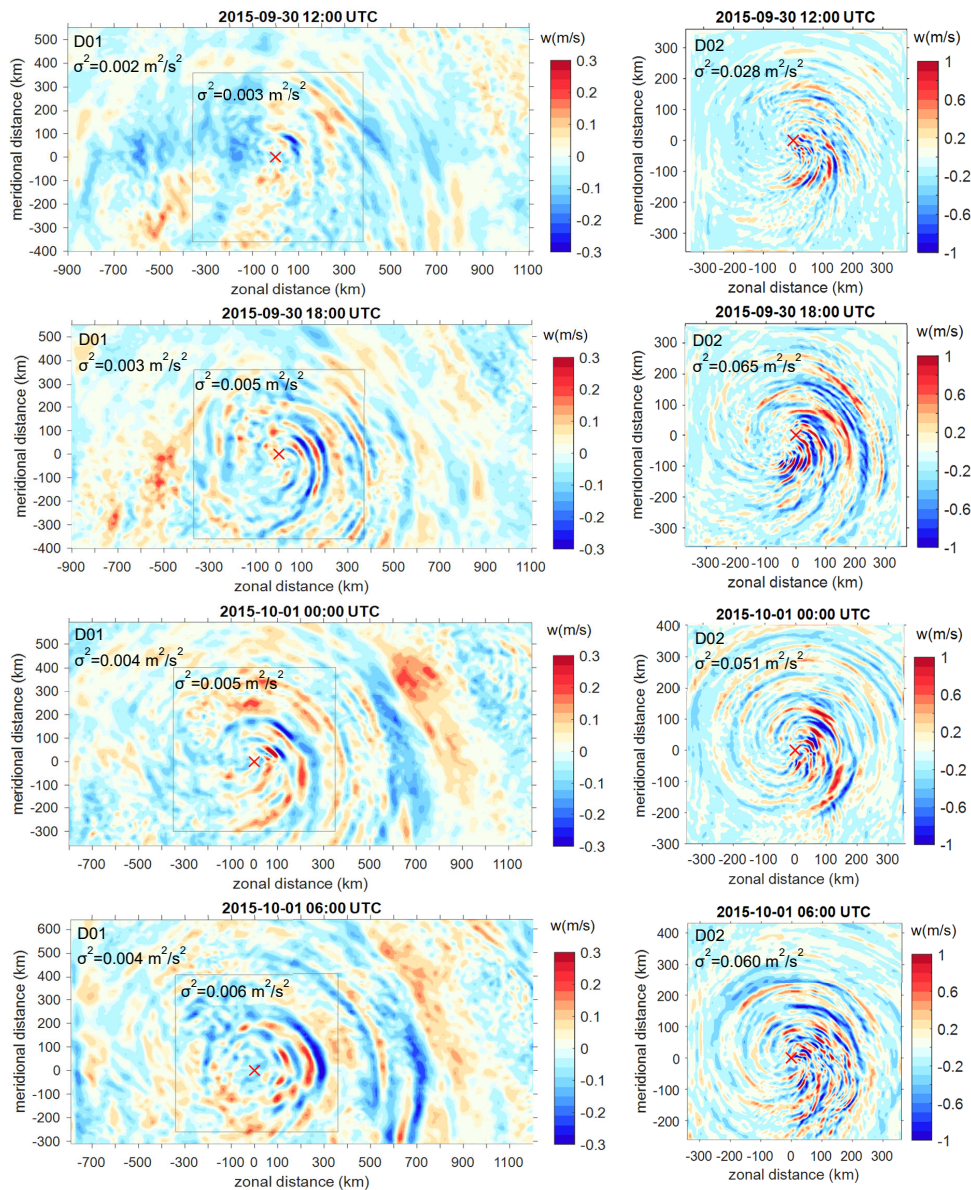
192 In contrast, the simulated intensity in D02 (4-km grid) exhibits good agreement with the IBTrACS intensity after an initial  
 193 spin-up period of approximately 12 hours. D02 successfully replicated the hurricane's rapid intensification until 18 UTC, 1  
 194 October, and accurately captured the subsequent weakening, re-intensification, and a second weakening of the hurricane. For  
 195 the sake of conciseness, the evaluation of the simulated track and convection strength of the hurricane is presented in the  
 196 appendix.

197 The examples of GWs from D01 and D02, represented by vertical velocities, are compared in Fig. 2, depicting GWs at  
 198 30 km altitude during the rapid intensification of Hurricane Joaquin. The GWs simulated with fine and coarse horizontal grid  
 199 spacings exhibit notable differences in features and intensities. At 12 UTC on 30 September, GW patterns in D02 are clear,  
 200 whereas those in D01 are much less organized, suggesting insufficient spin-up in D01 compared to D02. Afterward, the waves  
 201 in both D02 and D01 become clear and organized. Influenced by the easterly flow (not shown), the waves are inhibited  
 202 downstream on the western side of the TC, while the wavefronts become more closely packed on the eastern side. In D02,  
 203 GWs form tight spirals emanating from the center of the hurricane, with radial horizontal wavelengths of about 30 km at the





204 depicted time slots in Fig. 2. These spiral waves resemble GWs observed in prior studies (e.g., Chow et al., 2002; Kim and  
205 Chun, 2010; Hima Bindu et al., 2016; Nolan and Zhang, 2017). However, the small-scale waves in D02 are not fully  
206 represented in D01, where radial horizontal wavelengths are approximately 60 km. Unlike D02, D01 did not produce the spiral  
207 GWs that are typically caused by the asymmetry of instabilities rotating along the eyewall (Chow et al., 2002; Nolan, 2020)  
208 or unsteady vortical motions in the inner-core region of TCs (Hendricks et al., 2010). The spiral wave structures in D02 persist  
209 when resampling or averaging the 4-km grid results to the 12-km grid (not shown).





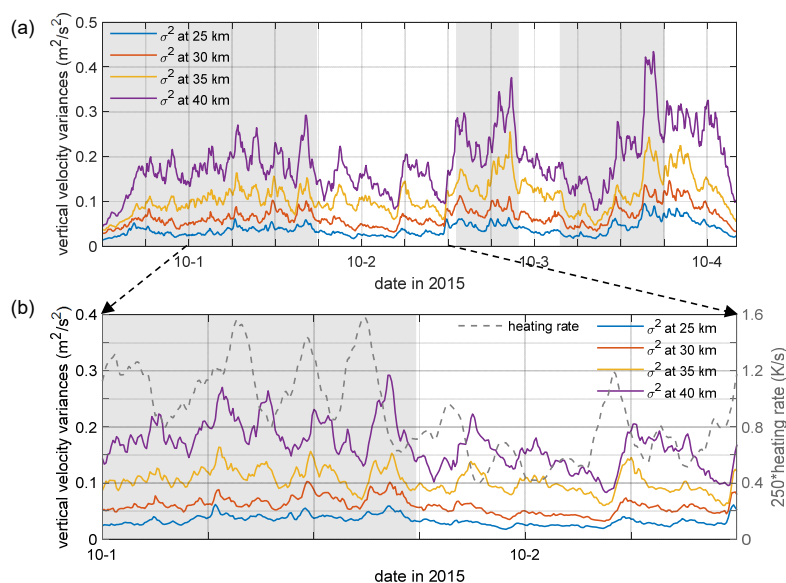


211 **Figure 2.** Vertical velocities ( $w$ , unit: m/s) at 30 km altitude in D01 (left panel) and D02 (right panel). The red crosses indicate the hurricane  
 212 center. The gray rectangles on the left panel denote the area of D02. The values of the intensity of GWs represented as the variances of  
 213 vertical velocities ( $\sigma^2$ ) in the domains are labeled. Please notice that different color bar ranges are used for D01 and D02.

214 The GW intensity (GWI), represented by the variances of vertical velocities ( $\sigma^2$ ), increased as the hurricane intensified.  
 215 The GWI is much larger in D02 than that in D01. Before D01 fully spun up, the GWI in D02 was ten times as large as that in  
 216 the same area in D01. Afterward, the GWI in D02 is about four to five times larger than in D01. In the following sections, we  
 217 focus on analyzing the simulation results from D02 because they are more accurate in terms of hurricane intensity and intensity  
 218 change and may produce more reliable results of the GWs features associated with the intensity tendency.

### 219 5.2 GW intensity associated with hurricane intensity

220 Figure 3 shows the time series of GWI at altitudes of 25, 30, 35, and 40 km. Generally, GWs exhibit higher intensity  
 221 during the intensification period compared to the weakening period, aligning with previous statistical analyses based on  
 222 satellite observations (Hoffmann et al., 2018; Wright, 2019). In Fig. 3b, the GWI from 00 UTC, 1 October to 12 UTC, 2  
 223 October, is magnified and overlaid with the time series of maximum heating rate ( $\partial T/\partial t$ ) in the upper troposphere (5–15 km)  
 224 denoted as HR. As depicted in Fig. 3b, considering thermal forcing in convection as one of the primary triggers for GWs,  
 225 phases of increased HR precede phases of increased GWI, with varying time lag.

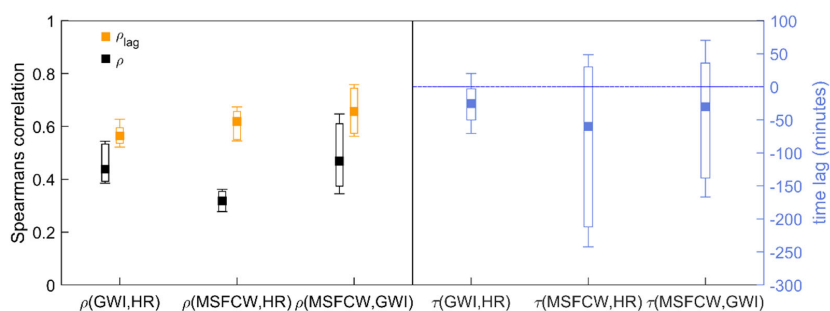


226 **Figure 3.** (a) Time series of GW intensity (vertical velocity variances  $\sigma^2$ ) on levels of 25, 30, 35, and 40 km from 12 UTC, 30 September to  
 227 04:00 UTC, 4 October; (b) Time series of GW intensity overlaid with time series of the maximum heating rate ( $\partial T/\partial t$ , gray dashed line) at  
 228 5–15 km from 00 UTC, 1 October to 12 UTC, 2 October. The intensification periods in (a) and (b) are marked with gray shading.  
 229

230 Using HR as a proxy for thermal forcing, we examine the correlations between thermal forcing and GWI, as well as  
 231 between thermal forcing and hurricane intensity. Following the approach outlined by Wu et al. (2022), we selected 6-hour



232 segments at 6-minute intervals, excluding the initial 12 hours of the spin-up period, resulting in 820 6-hour segments. These  
 233 segments encompass thermal forcing (represented by HR), hurricane intensity (represented by MSFCW), and stratospheric  
 234 GWI (represented by the mean variances of vertical velocities estimated between 20–43 km). Subsequently, Spearman's rank-  
 235 order correlation coefficients  $\rho$  were calculated between HR and GWI, as well as between HR and MSFCW. To account for  
 236 the time shift between the phases of these three variables, we systematically identified the time shift that yielded the largest  
 237 time-lagged  $\rho_{lag}$  in the 6-hour time series, recording it as the "best" time lag  $\tau$ .



238  
 239 **Figure 4.** Spearman's rank-order correlation coefficients  $\rho$  between the GWs intensity (GWI) and the heating rate (HR), the maximum  
 240 surface wind (MSFCW) and the HR, and the MSFCW and the GWI are presented on the left. The instantaneous  $\rho$  values are depicted in  
 241 black, while the time-lagged  $\rho_{lag}$  values are shown in orange. All displayed values have undergone a significance test with a confidence level  
 242 of 95%. The corresponding time lags ( $\tau$ ) are displayed on the right. The box plot illustrates the minimum, 25th percentile, median, 75th  
 243 percentile, and maximum values.

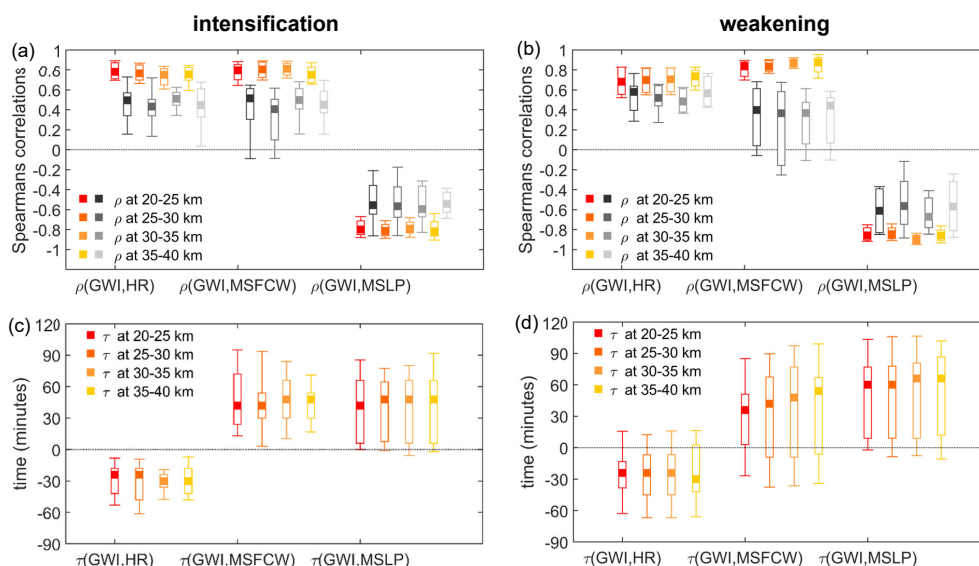
244 Figure 4 shows the Spearman's correlation coefficients  $\rho$  marked in black and the time-lagged  $\rho_{lag}$  marked in orange. The  
 245 "best" time lag  $\tau$  is presented on the right panel of Fig. 4. As depicted in Fig. 4, a moderate correlation is observed among the  
 246 three variables, even when not accounting for the time shift. The median values of  $\rho(GWI, HR)$ ,  $\rho(MSFCW, HR)$ , and  $\rho(GWI,$   
 247  $MSFCW)$  are 0.42, 0.33, and 0.50, respectively. Similar to the correlation levels between GWI and HR, and between GWI and  
 248 MSFCW at varying altitudes in Wu et al., (2022), all the correlation level increases when considering the time shift between  
 249 the three variables. The median values of  $\rho_{lag}(GWI, HR)$ ,  $\rho_{lag}(MSFCW, HR)$ , and  $\rho_{lag}(GWI, MSFCW)$  increase to 0.57, 0.63,  
 250 and 0.67, respectively. The time lag  $\tau$  between GWI and HR, and MSFCW and HR align with expectations. The negative  
 251 median time lag  $\tau(GWI, HR)$  suggests that GWI follows the variations of HR, which is sensible as GWI depends on the strength  
 252 of the thermal forcing.

253 Compared with Wu et al. (2022), the correlation and time lag between MSFCW and HR,  $\rho(MSFCW, HR)$ ,  $\rho_{lag}(MSFCW,$   
 254  $HR)$ , and  $\tau(GWI, HR)$ , are newly incorporated to facilitate comparison with existing research. Time lag  $\tau(MSFCW, HR)$   
 255 indicates that, in most cases, hurricane intensity changes occur after the HR. Hazelton et al. (2017) noted that strong updrafts  
 256 (convective bursts) frequently occur up to 3 hours before TC intensification. These updrafts lead to intense heat release. We  
 257 confirmed that heat release is correlated with hurricane intensity change, and the time lag is up to about 240 minutes (4 hours),  
 258 consistent with the findings of Hazelton et al. (2017), although the median lag in our Joaquin simulation is about 60 minutes.



259 The  $\tau$ (MSFCW, HR) spans both positive and negative values, potentially due to the intricate relationship between hurricane  
 260 intensity and heat release.

261 The correlation between hurricane intensity and GWI is of particular interest in this study. The  $\rho_{\text{lag}}(\text{GWI}, \text{MSFCW})$   
 262 indicates a high correlation between hurricane intensity and GWI. However, the time lag  $\tau(\text{GWI}, \text{MSFCW})$  exhibits a  
 263 significant spread, obscuring the sequence of the two activities. Considering that updrafts and thermal forcing are more intense  
 264 and stronger during TC intensification (e.g., Guimond et al., 2010, 2016; Hazelton et al., 2017), the GWs triggered during the  
 265 intensification period could have distinct properties. Consequently, we calculated the Spearman's correlation coefficients  
 266 between GWI and HR, and between GWI and hurricane intensity represented by MSFCW and MSLP separately for the  
 267 intensification and weakening periods, as illustrated in Fig. 5a and Fig. 5b. The "best" time lags for the intensification and  
 268 weakening periods are presented in Fig. 5c and Fig. 5d.



269  
 270 **Figure 5.** Spearman's correlation coefficients and the time lag between variable series. (a) illustrates Spearman's correlation coefficients  $\rho$   
 271 with instantaneous  $\rho$  values marked in black and gray, and time-lagged  $\rho_{\text{lag}}$  marked in orange. Only values have undergone a significance  
 272 test with a confidence level of 95% are shown. (b) displays the corresponding time lag  $\tau$ . The box plot illustrates the minimum, 25th  
 273 percentile, median, 75th percentile, and maximum values. The left panels correspond to the intensification period, while the right panels  
 274 correspond to the weakening period.

275 Same as for the entire hurricane lifetime in Fig. 4, there are moderate correlations among these variables in the  
 276 intensification and weakening scenarios, and the correlation increases to a higher level after considering the time shift. In the  
 277 intensification scenario, the spread of time lags is significantly reduced (Fig. 5c), and there is a clear sequence: GWI changes  
 278 after HR, and the hurricane intensity changes after GWI. In contrast, in the weakening scenario, the sequence is not clear.  
 279 Comparing the time lags with Fig. 4b in Wu et al. (2022), the large spread in the time lag seems induced during the weakening  
 280 period. The hypothesis regarding this contrast could be that during the intensification period, the intensive updraft and strong

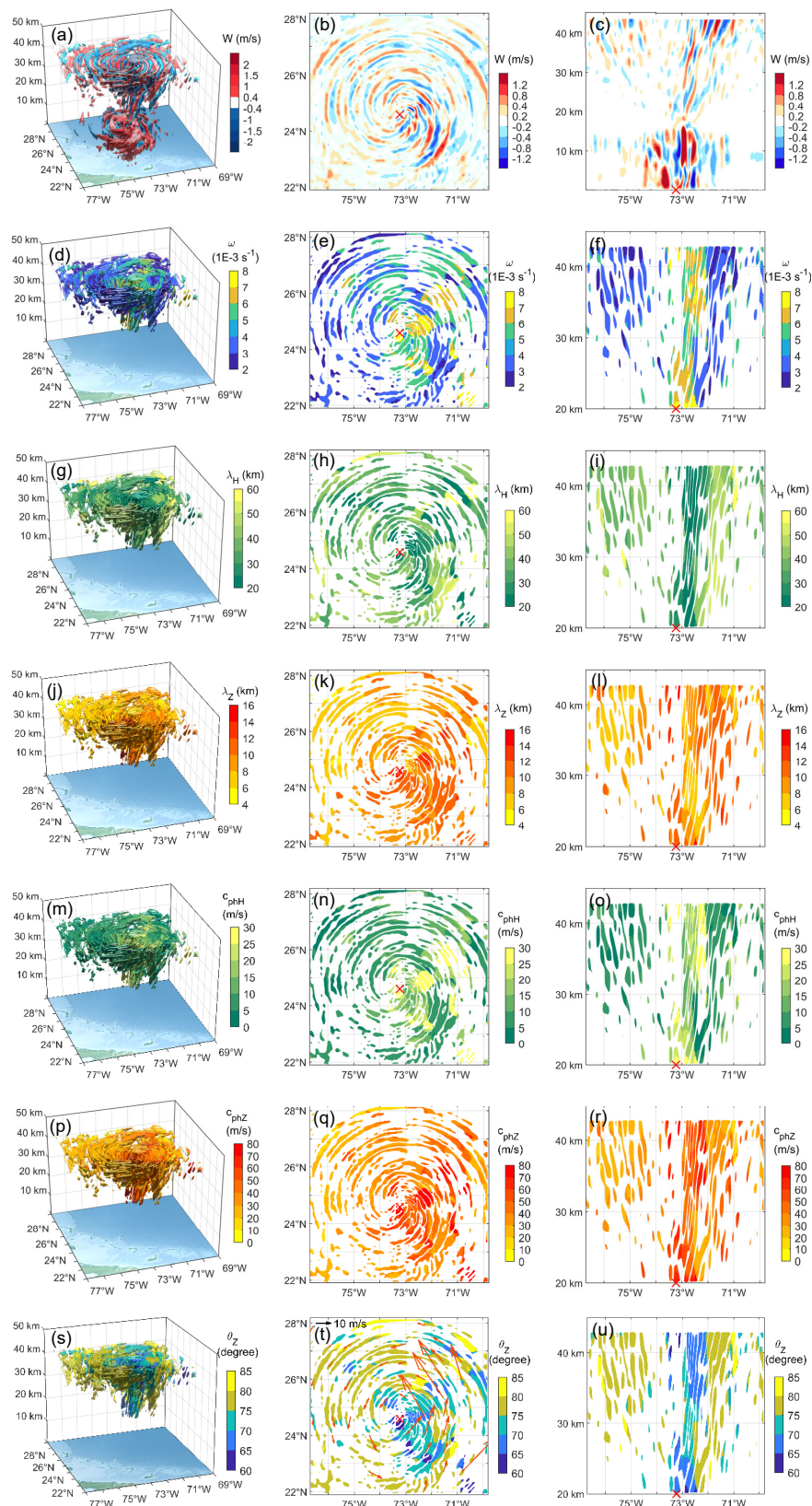


281 thermal forcing would generate GWs that propagate quickly to the stratosphere, possibly before the hurricane intensifies. On  
282 the contrary, during the weakening scenario, the vertical propagation of GWs is relatively slow so that the hurricane may  
283 intensify before or after the GWs propagate to the stratosphere, making the time lag between the change of hurricane intensity  
284 and GW intensity unclear.

### 285 **5.3 Gravity wave characteristics associated with hurricane intensification**

286 To explicitly investigate the characteristics of stratospheric GWs during the intensification period, we employed the 3-D  
287 S-transform to estimate the GW spectral properties. As an illustrative example, Fig. 6 presents snapshots of the 3-D GWs,  
288 represented by vertical velocities, and the estimated intrinsic frequency, wavelengths, and intrinsic phase speed at 00 UTC on  
289 1 October 2015. These properties are observed at 35 km altitude and along west-east vertical sections crossing the hurricane  
290 center.

291 The asymmetric GW patterns, suppressed on the west and compressed on the east side of the hurricane center, are  
292 influenced by the background easterlies. The frequency of the stratospheric GWs excited by Hurricane Joaquin is  
293 approximately one order of magnitude larger than the Coriolis frequency and one order of magnitude smaller than the Brunt-  
294 Väisälä frequency (Fig. 6d–f), indicating a frequency range consistent with mid-frequency GWs. The GWs with relatively  
295 higher frequency correspond to the inner core region, including the eyewall of the hurricane and the region just outside of it  
296 where deep convection occurs actively. These GWs also exhibit shorter horizontal wavelengths (about 20–40 km). The vertical  
297 phase speed exceeds twice the horizontal phase speed, resulting in an upward tilt of the vertical propagation angle (greater  
298 than 45°). The vertical propagation angle in Fig. 6t, overlaid with the horizontal phase velocity represented by the red arrows,  
299 illustrates the outward counterclockwise propagation of waves from the center while simultaneously moving upward, creating  
300 an upward spiral pattern. These waves exhibit faster vertical speeds, enabling them to reach the upper stratosphere without  
301 extensive horizontal propagation from the source. It can be anticipated that wave packets with distinct phase speeds and  
302 complex propagation directions may easily lead to wave superposition. The observed wavy structures, extending from the  
303 inner core to the outer region and from the lower stratosphere to the upper stratosphere, likely reflect the transiency of the  
304 wave sources. This phenomenon is expected and aligns with the characteristics observed in a realistic hurricane case.



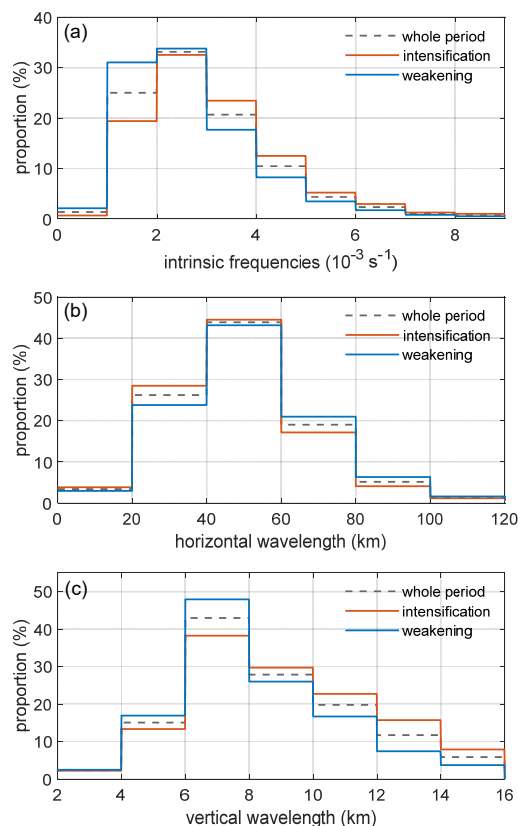


306 **Figure 6.** Properties of the stratospheric GWs excited by Hurricane Joaquin (2015) at 00 UTC on 1 October 2015: (a–c) the simulated  
307 vertical velocities  $w$ , (d–f) the estimated intrinsic frequency  $\omega$ , (g–i) horizontal wavelength  $\lambda_H$ , (j–l) vertical wavelength  $\lambda_Z$ , (m–o) intrinsic  
308 horizontal phase speed  $c_{phH}$ , (p–r) intrinsic vertical phase speed  $c_{phZ}$ , and (s–u) vertical propagation angle  $\theta_Z$ . The vertical propagation angle  
309 ( $\theta_Z$ ) is overlaid with the horizontal phase velocity (the red arrows). Figures in the left column show the 3D features, figures in the middle  
310 column depict the features at 35 km, and those in the right column present a west-east vertical cross-section. Only waves of amplitudes larger  
311 than 0.2 m/s and their properties are shown. The red crosses in the middle column indicate the hurricane center. The red crosses in the right  
312 column mark the longitude of the hurricane center.

313 Furthermore, we analyzed the dominant characteristics of stratospheric GWs during the intensification and weakening  
314 periods based on the 3-D wave properties from 12 UTC, 30 September, to 04 UTC, 4 October 2015. The analysis is confined  
315 to wave properties between 20 and 35 km for two reasons. Firstly, upon comparing the background zonal and meridional winds  
316 from the WRF simulations to those from ERA5, we found that the background winds in the stratosphere from the two sources  
317 agreed very well below 35 km. However, discrepancies emerged above approximately 35 km due to the absence of some wind  
318 reversals in the WRF simulations that were present in the ERA5 winds. These wind reversals could potentially induce wave  
319 filtering, implying that wave patterns in the WRF simulations above 35 km might not represent the actual atmosphere.  
320 Therefore, waves above 35 km are excluded from the subsequent analysis. Please refer to the appendix for a detailed  
321 comparison of simulated background winds and winds from the ERA5 reanalysis. Secondly, we restrict the analysis to  
322 estimated wave properties between 20 and 35 km to mitigate potential boundary effects in the wave property analysis.

323 Figure 7 summarizes the occurrence probability of the intrinsic frequency, horizontal wavelengths, and vertical  
324 wavelengths during the whole period of the hurricane from 12 UTC, 30 September to 04 UTC, 4 October 2015, and separately  
325 during the intensification and weakening period. The occurrence probability of the intrinsic frequency peaks at  $2\text{--}3 \times 10^{-3} \text{ s}^{-1}$ .  
326 The hurricane tends to generate GWs with relatively higher intrinsic frequency during the intensification period and relatively  
327 lower intrinsic frequency during the weakening period. Specifically, during hurricane intensification, the waves with intrinsic  
328 frequency lower than  $2 \times 10^{-3} \text{ s}^{-1}$  are approximately 14% less compared to the weakening period, whereas the waves with  
329 intrinsic frequency between  $3\text{--}7 \times 10^{-3} \text{ s}^{-1}$  show an increase of about 10% compared to the weakening period. The horizontal  
330 and vertical wavelengths peak at approximately 40–60 km and 6–8 km, respectively. During the intensification period, the  
331 GWs tend to have slightly shorter horizontal wavelengths and longer vertical wavelengths. Figure 8 illustrates this trend from  
332 the perspective of wavenumber versus frequency.

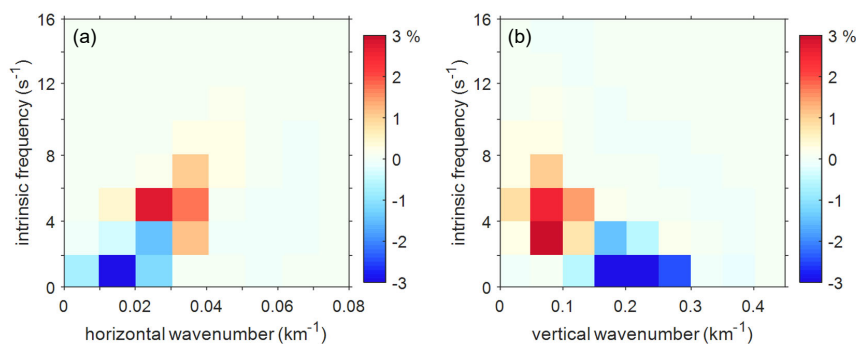




333

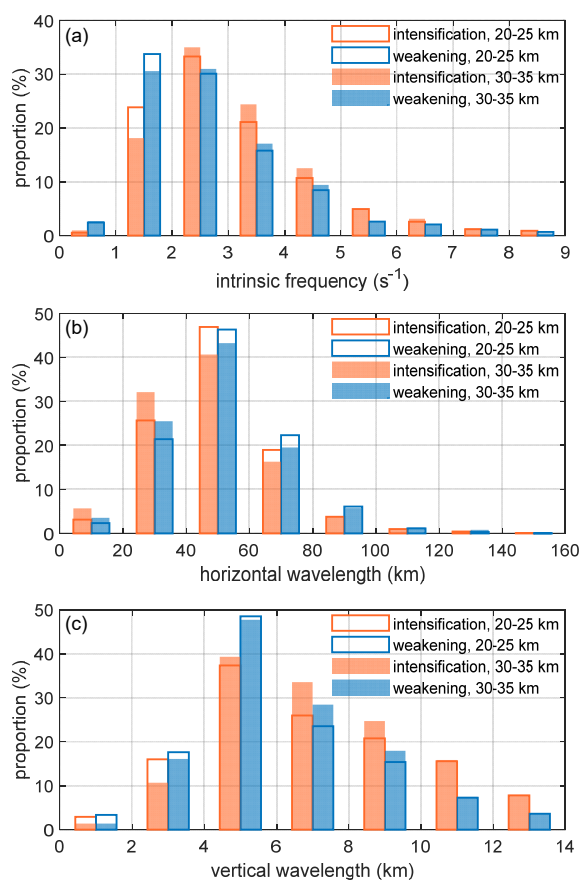
334 **Figure 7.** The occurrence frequency of stratospheric gravity waves (GWs) properties regarding the change in hurricane intensity. (a) The  
 335 occurrence probability of intrinsic frequency, (b) horizontal wavelengths, and (c) vertical wavelengths during the entire lifetime and the  
 336 intensification and weakening periods.

337 The differences in the occurrence probability of wavenumber versus frequency (intensification minus weakening) indicate  
 338 that, compared with GWs generated during hurricane weakening, GWs generated during hurricane intensification are  
 339 characterized by higher frequency ( $> 3 \times 10^{-3} \text{ s}^{-1}$ ), shorter horizontal wavelengths (20–40 km), larger horizontal wavenumbers  
 340 ( $0.025\text{--}0.05 \text{ km}^{-1}$ ), longer vertical wavelengths ( $> 8 \text{ km}$ ), and smaller vertical wavenumbers ( $0.06\text{--}0.12 \text{ km}^{-1}$ ). As observed  
 341 from Fig. 9, ascending from the lower stratosphere (20–25 km) to the middle stratosphere (30–35 km), the occurrence  
 342 possibilities of intrinsic frequency and wavelengths slightly shift to higher frequencies, shorter horizontal wavelengths, and  
 343 longer vertical wavelengths in both intensification and weakening scenarios. This shift is expected since these waves propagate  
 344 to higher altitudes faster. As GWs propagate from the lower stratosphere to the middle stratosphere, the contrast between the  
 345 intensification and weakening scenarios in the above three attributes becomes more prominent.



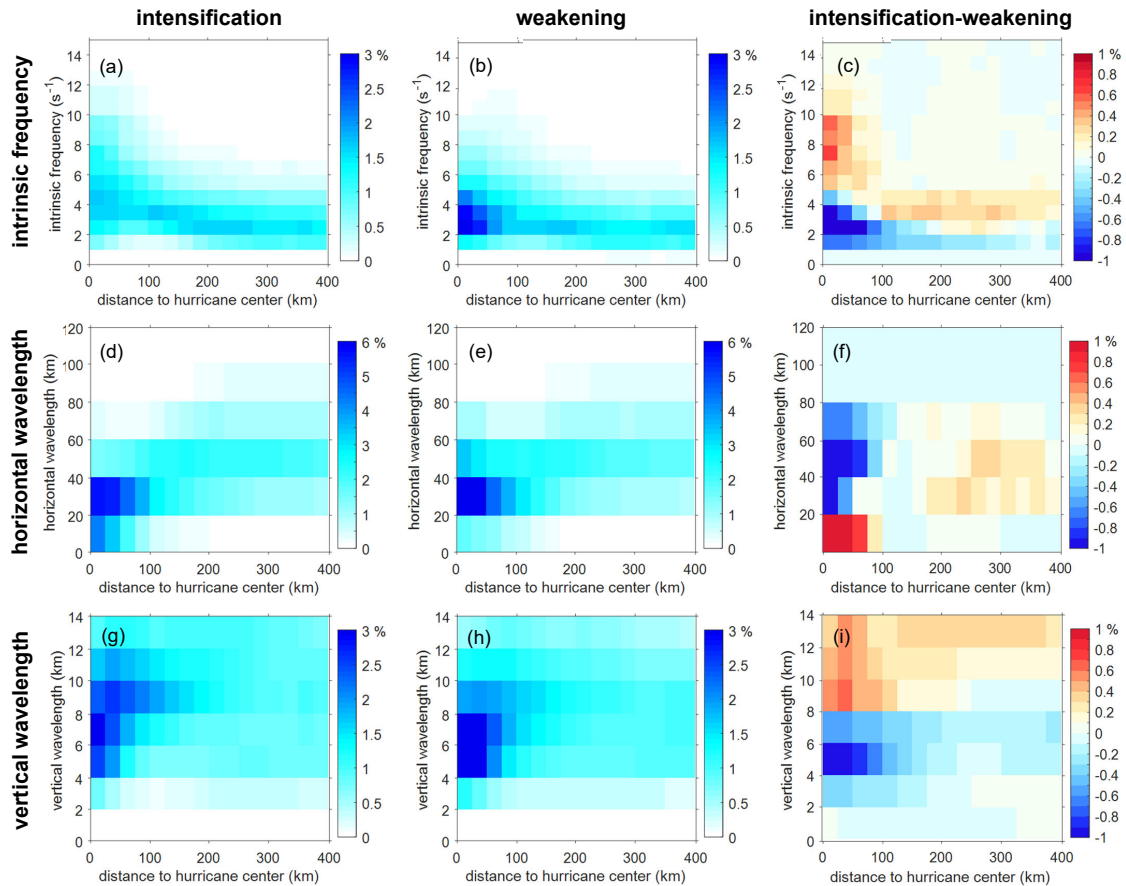
346

347 **Figure 8.** Differences in the occurrence probability of wavenumber versus intrinsic frequency during the intensification and weakening  
 348 period (intensification minus weakening). (a) Horizontal wavenumber versus intrinsic frequency. (b) Vertical wavenumber versus intrinsic  
 349 frequency.



350

351 **Figure 9.** Occurrence probability of stratospheric gravity waves (GWs) properties: (a) intrinsic frequency, (b) horizontal wavelengths, and  
 352 (c) vertical wavelengths separately during hurricane intensification and weakening at the altitude range of 20–25 km and 30–35 km.



354

354 **Figure 10.** The probability density of intrinsic frequency, horizontal wavelengths, and vertical wavelengths during the intensification and  
 355 weakening period concerning the distance to the hurricane center, and the differences in the occurrence probability density (intensification  
 356 minus weakening).

357 Figure 10 illustrates the probability density distribution of intrinsic frequency, horizontal and vertical wavelengths, and  
 358 their differences concerning the distance to the hurricane center. The waves with relatively higher frequency ( $\geq 3 \times 10^{-3} \text{ s}^{-1}$ ),  
 359 shorter horizontal wavelengths ( $\leq 40 \text{ km}$ ), and longer vertical wavelengths ( $\geq 4 \text{ km}$ ) display a higher probability density in the  
 360 inner core region (approximately 200 km to the hurricane center for the case of Hurricane Joaquin) for both intensification and  
 361 weakening scenarios. The distinction in the probability density of intrinsic frequency becomes more pronounced in the inner  
 362 core, particularly for the frequencies exceeding  $5 \times 10^{-3} \text{ s}^{-1}$ . Despite the general tendency for GWs to exhibit shorter horizontal  
 363 wavelengths during the intensification period, this characteristic diminishes beyond the 100 km radius. GWs generated during  
 364 the intensification period feature relatively longer vertical wavelengths, a trend that persists within the 400 km radius and is  
 365 particularly evident in the inner core region as well.

366 In summary, based on the analyses of the characteristics of the GWs of Hurricane Joaquin, we identified distinctive  
 367 features in the GWs generated during the intensification of hurricanes. Specifically, the GWs associated with TC intensification



368 exhibit relatively higher intrinsic frequencies, shorter horizontal wavelengths, and longer vertical wavelengths across the lower  
369 and middle stratosphere. Notably, the differences in wave characteristics are particularly pronounced in the inner core region  
370 of the hurricane, within a radius of approximately 200 km around the center. These GWs that can help separate the hurricane  
371 intensification and weakening periods may indicate the dynamics associated with deep convection in the eyewall and its  
372 immediate surroundings. The identified features of the GWs offer insights that contribute to the interpretation of the findings  
373 presented in Fig. 5. During the intensification period, the time required for the GWs of long vertical wavelengths and high  
374 frequencies to be generated and reach the stratosphere may be shorter than the time needed for the hurricane to intensify in  
375 response to increased convective activities. So, observing the stratospheric GWs with these characteristics could help monitor  
376 the intensification of the hurricane.

## 377 **6 Discussion and conclusions**

378 This study characterized the stratospheric GWs induced by Hurricane Joaquin in 2015 through realistic simulations  
379 using the Weather Research and Forecasting (WRF) model. The simulations encompass wave generation and interactions  
380 with background winds in the actual atmosphere based on ERA5 initial and boundary conditions. Additionally, we employed  
381 the novel 3-D S-transform method to estimate the wave properties. This study emphasizes the specific GW characteristics  
382 and the correlations with hurricane intensification, aiming to provide basic knowledge that could potentially facilitate the  
383 monitoring of hurricane intensification by observing the stratospheric GWs using satellite instruments operating in the  
384 infrared and microwave bands.

385 Similar as Wu et al. (2022), this study first confirmed the statistically significant correlation between the intensity of the  
386 GWs triggered by TCs and the intensity of TCs observed in long-term satellite data (Hoffmann et al., 2018; Wright, 2019) is  
387 also valid in a specific hurricane case. Additionally, by separating the lifetime of the hurricane into intensification and weakening  
388 period, we found that the pronounced time-lagged correlation between the hurricane intensity and stratospheric GWs primarily  
389 exists during hurricane intensification. This finding suggests the feasibility of monitoring hurricane intensity increases by  
390 observing increased stratospheric GW intensity. Secondly, based on the 3-D wave analyses, we identified distinct  
391 characteristics of the GWs generated during the intensification of Hurricane Joaquin. These features include relatively higher  
392 intrinsic frequency ( $> 3 \times 10^{-3} \text{ s}^{-1}$ ), shorter horizontal wavelengths ( $\leq 40 \text{ km}$ ), and longer vertical wavelengths ( $\geq 8 \text{ km}$ ). It  
393 should be note that the Hurricane Joaquin generated spiral GWs during the period of analysis, which may show relatively  
394 shorter horizontal wavelengths than concentric GWs (Kim and Chun, 2010). The wave packets exhibit rapid vertical speeds,  
395 and the time required for them to propagate up to the stratosphere may be comparable to or even shorter than the time it takes  
396 for the hurricane to intensify in response to increased convection activities. These wave characteristics offer supplementary  
397 information for monitoring and warning of hurricane intensification.



398        Considering the rapid vertical speeds of these waves, they ascend to the upper stratosphere without extensive horizontal  
399        propagation from the source. A discernible contrast in GW frequencies and wavelengths between hurricane intensification and  
400        weakening scenarios is primarily observed within the inner core ( $\leq 200$  km around the center). Therefore, observing these  
401        distinct GWs, which manifest during hurricane intensification, might necessitate a satellite instrument with extensive coverage  
402        of the inner core region, ideally with high horizontal resolution, such as provided by infrared nadir sounders. Additionally, the  
403        high temporal resolution of the measurements would allow us to better characterize the time evolution of GW and TC intensity  
404        and facilitate the characterization of the GW spectral characteristics. Such high temporal resolution could be achieved by a  
405        satellite instrument operating in a geostationary orbit. Meanwhile, the distinction in GW frequencies and wavelengths between  
406        hurricane intensification and weakening scenarios extends across the lower and middle stratosphere, but it gets even more  
407        evident at higher altitudes. Consequently, a spectral channel focused on higher altitudes may have advantages in identifying  
408        hurricane intensification, provided the GWs can effectively propagate to these elevated altitudes.

409        Based on the analysis of wave properties outlined above, the  $4.3 \mu\text{m}$   $\text{CO}_2$  fundamental band of AIRS on board Aqua, the  
410        Infrared Atmospheric Sounding Interferometer (IASI) instruments on board the European MetOp satellites (Hoffmann et al.,  
411        2013, 2014), or the Cross-track Infrared Sounder (CrIS) instrument on board Suomi-NPP, NOAA-20 and NOAA-21  
412        (Eckermann et al., 2019) hold the potential for observing stratospheric GWs to identify TCs intensification. However,  
413        uncertainties exist in establishing a clear relationship between the observed GWs and the intensity of the TC that triggered the  
414        waves in the real atmosphere. The amplitudes and spectra of the observed waves are significantly influenced by background  
415        winds. Changeable background wind conditions may obscure the distinction between the characteristics of GWs during TC  
416        intensification and weakening. Moreover, the complex thermodynamics of a hurricane are treated as a "black box" in this study.  
417        While the dynamical and thermal processes and changes in TC structure before intensification have been extensively studied  
418        (e.g., Wang and Wang, 2014; Miyamoto and Nolan, 2018), this topic is beyond the scope of our current investigation. The  
419        uncertainties in TC intensity changes resulting from convective activities may introduce additional challenges.

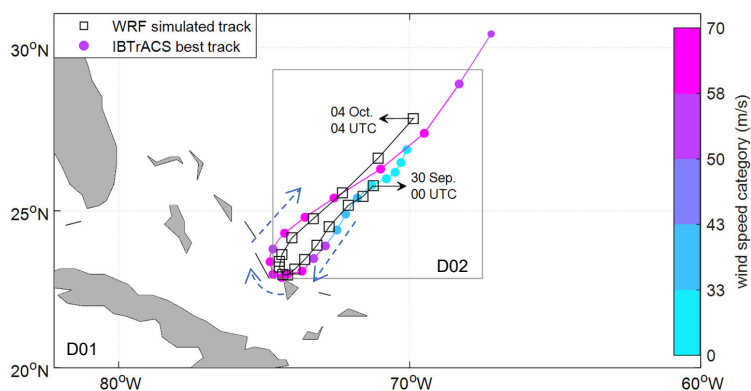
420        In summary, this study found a high-level time-lagged correlation exists between the stratospheric GWs amplitudes and  
421        the hurricane intensity during the intensification period. Moreover, the stratospheric GWs during hurricane intensification  
422        exhibit relatively higher frequencies, shorter horizontal wavelengths, and longer vertical wavelengths in the inner core region.  
423        The findings of this study further support the feasibility of estimating the intensification of TCs by observing stratospheric  
424        GWs. They may provide further knowledge for optimal utilization of current observation techniques and for planning new  
425        instruments tailored for these specific "target waves." This approach can provide valuable insights for estimating TC  
426        intensification in instances where the top of TCs is obscured by clouds. However, future research, particularly concerning the  
427        influence of background wind conditions, is still needed to better specify the applicable scenarios of this approach.



428 **Appendices**

429 **Appendix A Evaluation of the simulated track of Hurricane Joaquin (2015)**

430 The WRF simulation was conducted for 100 hours from 00 UTC, 30 September to 04 UTC, 4 October 2015. The  
431 simulated hurricane track is depicted and compared with the IBTrACS dataset in Fig. A1. An accurate track simulation is  
432 necessary to ensure a suitable background for the hurricane's development. As illustrated in Fig. A1, the simulated hurricane  
433 track agrees well with the IBTrACS data. The simulation successfully captured the gradual southwestward movement before  
434 the track reversed and the subsequent relatively faster northeastward movement. The simulated hurricane center moves slightly  
435 slower than the IBTrACS hurricane center in the first 12 hours and faster afterward. The hurricane center moves further toward  
436 the northwest compared with the IBTrACS hurricane track after 18 UTC, 3 October.



437  
438 **Figure A1.** The IBTrACS hurricane centers from 00 UTC, 29 September to 12 UTC, 4 October 2015, are shown every 6 hours with colored  
439 dots, and the color is assigned according to the IBTrACS maximum sustained wind speed. The WRF simulated hurricane centers during the  
440 simulation period (00 UTC, September 30–04 UTC, 4 October 2015) are shown every 6 hours with black squares. The grey rectangle  
441 indicates the initial location of the inner domain (D02).

442 **Appendix B Comparison of WRF simulated background winds with winds from ERA5 reanalysis data**

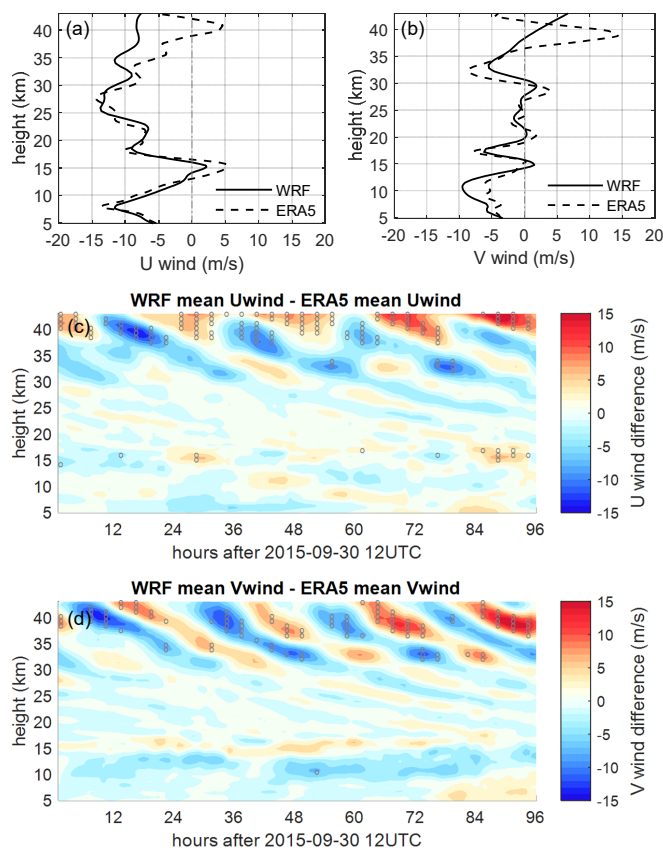
443 In order to assess the background winds in the WRF simulation, the mean zonal and meridional winds in the inner domain  
444 are compared with those averaged from ERA5 data in the same area. Examples of the comparison at 12 UTC on 1 October  
445 2015 are presented in Figure B1a–b. The background winds from the two sources exhibit good agreement below 30 km, but  
446 the discrepancies become more prominent from about 30 km upward. Above 35 km, the meridional wind exhibits a reversal  
447 in the ERA5 reanalysis, which is not present in the WRF simulations. Figure B1c–d summarizes the comparison of the winds  
448 during the simulation period. Above approximately 35 km, significant differences emerge in the background winds, with  
449 frequent instances of opposite wind directions between the simulation and ERA5 reanalysis. The wind reversal observed in  
450 the ERA5 reanalysis might induce wave filtering, potentially leading to discrepancies in the wave features between the WRF  
451 simulations and the ERA5 reanalysis or the real atmosphere. To ensure the reliability of our analysis, we focused exclusively





452 on GWs in the lower and middle stratosphere, specifically within the altitude range of 20–35 km, thereby excluding any wave

453 features questionable in the real atmosphere.



454

455 **Figure B1.** Comparison of the mean zonal (U) and meridional (V) winds from the WRF simulation and the ERA5 reanalysis. Examples of

456 the mean (a) zonal and (b) meridional winds in D02 from WRF simulations and ERA5 reanalysis on 1 October 2015, 12 UTC and time

457 series of differences of the mean (c) zonal and (d) meridional winds between WRF simulations and ERA5 reanalysis. The gray circles in (c)

458 and (d) mark opposite wind directions between the WRF simulated and ERA5 winds.

#### 459 **Data availability.**

460 The ERA5 reanalysis data (C3S, 2017) were retrieved from the ECMWF Meteorological Archival and Retrieval

461 System (10.24381/cds.adbb2d47; last accessed: 1 September 2023). IBTrACS data were acquired from the National Centres

462 for Environmental Information, National Oceanic and Atmospheric Administration (10.25921/82ty-9e16; last accessed: 1

463 September 2023).



464 **Author contribution.**

465 XW and LH drafted the framework of the study. NPH and CJW developed the 3-D Stockwell transform analysis and  
466 adapted it to the WRF outputs on 3-D grids. XW, LH, CJW, and BC processed the data, produced all figures, and drafted most  
467 of the manuscript. MJA and SK offered key scientific guidance. All authors participated in the editing of the final manuscript.

468 **Competing interests.**

469 The authors declare that they have no conflict of interest.

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