



1 Climate Affects Global Basin-Related Metallogeny

2

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6

7 Abstract

8

9 The basin-related hydrothermal mineral deposits are the products of metal deposition in a relatively
10 small area from metal-rich saline brines that source from basins. Recent studies have confirmed that
11 the metal-rich ore-forming fluids were formed in semi-arid to arid environments, and are the
12 products of a complex system involving precipitation, weathering, groundwater, precipitation-
13 dissolution reactions, and evaporation. The evaporation is the main reason for the buildup of metals
14 in saline brines. The formation of metal-rich saline brines is commonly accompanied by the
15 formation of evaporites.

16

17 The statistical results of basin-related mineral deposits worldwide show that there are two
18 metallogenic periods after the great oxidation event: 2.1-1.4Ga (Period I) and 0.8Ga to present
19 (Period II), with few scattered between these two periods (metallogenic quiescence period). In
20 addition, Metallogenic Period II has five metallogenic peaks: ~380-340Ma (II-1), ~300-240Ma (II-
21 2), ~160-100Ma (II-3), 60-40Ma (II-4), and one specific stratiform Cu metallogenic concentration
22 period of ~580-500Ma (II-5). These two metallogenic periods and five peaks are coupled with the
23 widespread development of saline deposits in time. The basin-related ore deposits are mainly
24 symmetrically occurring in 10°-60° in paleo-latitudes, which is consistent with the occurring
25 latitudes of evaporites.

26

27 The metallogenic quiescence period corresponded to the scarcity of saline deposits and was
28 probably caused by the combination of a lack of hydrological closed basins and arid to semi-arid
29 environments during 1.4-0.8 Ga. This quiescence period was coupled with the booming of



30 stromatolites and the extremely thin continents, both of which suggest an Earth with flat continents
31 that were covered by a hot and wet climate, and the widely developed shallow marine environments
32 of the major continents at middle and low altitudes during 1.4-0.8 Ga.

33

34 **Keywords:**

35

36 Climate, Basin-related mineral deposits, Mesoproterozoic, Saline giants, Stromatolite

37

38 1. Introduction

39

40 For more than thirty years, geologists have strived to unlock the secrets of secular distribution of
41 many types of ore deposits in Earth's history (e.g., Meyer, 1981, 1988; Sawkins, 1984; Groves et al.,
42 2005a, b; Holland, 2005; Leach et al., 2010; Cuney, 2010), such as sediment-hosted Pb-Zn deposits
43 (e.g., Sangster, 1990; Goodfellow and Lydon, 2007), and sediment-hosted stratiform Cu deposits
44 (e.g., Kirkham, 1989; Kirkham et al., 1994; Hitzman et al., 2010). There are three principal factors
45 that have influenced the global pattern of metallogenesis: long-term global tectonic trends (e.g.,
46 Barley and Groves, 1992; Goldfarb et al., 2010), the evolution of the hydrosphere-atmosphere,
47 which is inextricably related to the tectonic recycling (e.g., Slack and Cannon, 2009; Bekker et al.,
48 2010), and the secular decrease in global heat production (e.g., Holland, 1984; Sawkins, 1990;
49 Barley and Groves, 1992; Farquhar et al., 2014). The plate tectonic evolution has been increasingly
50 emphasized as a factor that led to the formation of many famous mineral provinces (e.g., Rona, 1980;
51 Mitchell and Garson, 1981; Sawkins, 1984; Groves et al., 2005a; Robb, 2005, 2020). Nevertheless,
52 due to the lack of a reasonable understanding of the formation processes of ore-forming fluids for
53 basin-related ore deposits, there are ambiguities in the correlation between global metallogeny and
54 the complex Earth history of supercontinent assembly and breakup during the Phanerozoic (e.g.,
55 Leach et al., 2005, 2010; Hitzman et al., 2010).

56

57 The availability of information on important mineral deposits worldwide is steadily increasing,
58 significant advances have been achieved in understanding long-term secular changes in the Earth



59 System, and data gaps for defining broad temporal distributions of ore types are becoming smaller.
60 However, there continues to be one significant gap in our understanding of the secular distribution
61 of many types of ore deposits in time and space (e.g., Robb, 2005; Goldfarb et al., 2010). If we shift
62 our perspective from ore-deposit centric to one centered on ore-forming fluids (e.g., Richard et al.,
63 2013, 2014; Zhang et al., 2022, 2023a, 2023b), the global distribution pattern of basin-related
64 mineral deposits (e.g., SEDEX and MVT Pb-Zn deposits, sediment-hosted stratiform Cu deposits,
65 Unconformity related U deposits) in time and space would be understood more easily. Recently,
66 some researchers have begun to study the metal-rich fluids in developing basins, so as to reveal the
67 formation processes of paleo-ore-forming fluids for the basin-related ore deposits (e.g., Li and
68 Barnes, 2019; Zhang et al., 2022, 2023a, b). The goal of this paper is to provide a new perspective
69 into the climate controls on the formation of metal-rich fluids, and therefore controls the formation
70 of basin-related ore deposits.

71

72 The present is the key to the past. Metalliferous saline brines in basins are the ore-forming fluids for
73 basin-related ore deposits, such as unconformity-related U deposits, MVT Pb-Zn deposits, and
74 redbed Cu deposits (e.g., Leach et al., 1997, 2005; Zhang et al., 2022). This paper covers six main
75 subjects: (1) the description of major types of basin-related ore deposits, in particular SEDEX and
76 MVT Pb-Zn deposits, sediment-hosted stratiform Cu deposits, and basin-related hydrothermal type
77 U deposits; (2) the formation processes of metal-rich saline brines in hydrologically closed basins
78 that are responsible for basin-related ore deposits; (3) the global distribution of basin-related ore
79 deposits in time and space; (4) the factors that influence the formation of metal-rich saline brines;
80 (5) how climate influences the formation and distribution of basin-related ore deposits globally; and
81 (6) how major geological events influence the global climate and the formation and distribution of
82 basin-related ore deposits.

83

84 2. Description of Basin-Related Ore Deposits

85

86 Basin-related ore deposits are a diverse group of ores that are genetically related to basinal brines,
87 and have no direct genetic association with igneous activity. These include Unconformity-type U



88 deposits, SEDEX and MVT Pb-Zn deposits, and Redbed Cu deposits (e.g., Hitzman et al., 2010;
89 Leach et al., 2010), as well as vein type polymetallic deposits outside of basins but genetically
90 related to basinal brines (Kyser, 2014; Zhang et al., 2023a, 2023b). These deposits are typically
91 stratiform or strata-bound, although some discordant ores occur (e.g., Bradley and Leach, 2003;
92 Zhang et al., 2019, 2022). Vein ore is often important in many deposits, and the ores consist of
93 various metals such as Pb, Zn, Cu, As, Ag, Fe, and U (e.g., Selley et al., 2005; Kerrich et al., 2005,
94 2008; Farquhar et al., 2010). Comprehensive review articles are available for sediment-hosted Pb-
95 Zn deposits (Leach et al., 2005), basin-related hydrothermal type U deposits (Cuney, 2010; Kyser,
96 2014), and sediment-hosted stratiform Cu deposits (Hitzman et al., 2010). Therefore, the author will
97 not present a detailed description of the geological characteristics of basin-related ore deposits, but
98 rather a brief summary and overview.

99
100 The presence of laminated metallic minerals that parallel bedding is commonly accepted as
101 permissive evidence for exhalative ore (e.g., Goodfellow et al., 1993; Goodfellow and Lydon, 2007).
102 However, this kind of explanation is unsatisfactory (detailed discussion in Leach et al., 2010). In
103 addition to the laminated Pb-Zn deposits, there are many laminated deposits that did not form as
104 exhalites, such as the strata-bound scheelite mineralization near Halls Creek in Western Australia
105 (Todd Ririe, 1989), stratiform W-Sb-Au mineralization in South China (Gu et al., 2012), stratiform
106 Mn-Fe deposits in North Europe (Bostrom et al., 1979), the Nuheting U deposit in China (Bonnetti
107 et al., 2015), and partially the Bayan Obo REE deposit (e.g., Liu et al., 2004; Yang et al., 2017;
108 Zhang et al., 2017). Previous researchers proposed that these deposits with ore textures that mimic
109 synsedimentary textures were formed by replacement processes (e.g., Kelley et al., 2004a, b; Leach
110 et al., 2005). This explanation is also unsatisfactory, as the highly selective mineral replacement
111 after the deposition of host rocks without specific alteration is difficult to convince geologists.

112
113 The basins that host or are genetically related to ore deposits mainly occur in passive margins,
114 continental rifts, and sag basins (e.g., Leach et al., 2010; Hitzman et al., 2010; Cuney, 2010). The
115 tectonic setting of the basins determines the ore-hosting rock type, ore controls, as well as the
116 survivability of the deposit during tectonic recycling (e.g., Goldfarb et al., 2010). The author did not



117 provide detailed descriptions of these tectonic settings; however, the similarity between them is the
118 ubiquity of convergent sedimentary basins. In this paper, the author exemplified the SEDEX and
119 MVT Pb-Zn deposits, sediment-hosted Cu deposits, and basin-related hydrothermal type U deposits,
120 to find out their distribution characteristics in time and space, which could be used to test their major
121 influencing factors.

122

123 **2.1 SEDEX and MVT Pb-Zn deposits**

124

125 The most common basin-related Pb-Zn deposits are SEDEX and MVT Pb-Zn deposits, hosted by a
126 variety of carbonate and siliciclastic rocks (Leach et al., 2005). The major ore minerals are sulfides
127 such as sphalerite, galena, and pyrite, while the nonsulfide minerals are mainly dolomite, calcite,
128 barite, siderite, and chert (e.g., Cooke et al., 2000; Yardley, 2005). These deposits have formed in a
129 variety of geologic and tectonic environments over the last two billion years, as documented by the
130 compilation result of Leach et al. (2005), based on the World Minerals Geoscience Database Project
131 (Sinclair et al., 1999).

132

133 **2.2 Sediment-hosted Cu deposits**

134

135 Sediment-hosted stratiform copper deposits comprise disseminated to veinlet copper and copper-
136 iron sulfides in siliciclastic or dolomitic sedimentary rocks. Sulfides conform closely, but usually
137 not exactly, with the stratification of host rocks (Brown, 1989, 1992, 1993, 1997). These deposits
138 account for approximately 23% of the world's production and known reserves, and are also
139 important sources of silver and cobalt (Hitzman et al., 2005). Some deposits contain other metals
140 including lead, zinc, and uranium, and a few contain gold and platinum group elements. The basins
141 that host the stratiform copper deposits can be marine or continental, and usually contain evaporites
142 overlying redbeds or in isolated non-red units within the continental redbed sequences themselves
143 (e.g., Cox et al., 2003; Hitzman et al., 2010). The compilation of sediment-hosted stratiform copper
144 deposits was cited from the statistical work of Kirkham (1989) and Kirkham et al. (1994).

145



146 **2.3 Hydrothermal type U deposits**

147

148 The basin-related hydrothermal type U deposits include the sub-types of vein-type U deposits hosted
149 by various lithologies, unconformity-related U deposits hosted by metamorphic rocks, collapse-
150 breccia type deposits, and hydrothermal type U deposits hosted by Na- and/or K metasomatism
151 alteration (Cuney, 2011, 2013, 2014, 2016; Cuney et al., 2012; Zhang et al., 2022). The hydrothermal
152 U districts that the author has compiled include Proterozoic U deposits in the Nordic region,
153 metasomatism type U deposits in Ukraine (Kirovograd-Smolino region, Krivoy Rog region) and the
154 Russian Federation (Lake Ladoga and Onega districts, European Part), Proterozoic unconformity-
155 related and vein type U deposits in Canada, Australia, India, Brazil, China, and Africa, such as the
156 famous U deposits in Athabasca and Thelon basins, Cuddapah and Bhima U region, Singhbhum Cu-
157 U belt, Pink Creek Inlier U region, Franceville U region in Gabon, and so on (Dahlkamp, 2009,
158 2010, 2016). In addition, the Phanerozoic vein-type U deposits account for a large portion, such as
159 the Erzgebirge district in Germany, the La Crouzille district in France, the Holdita and Crucea
160 districts in Romania, South China U Province, Streltsovsk district in the Russian Federation, the
161 Spokane mountain area in the US, the Poços de Caldas Region in Brazil, and so on (Dahlkamp,
162 2009, 2010, 2016). The detailed statistics of U districts are in Appendix A.

163

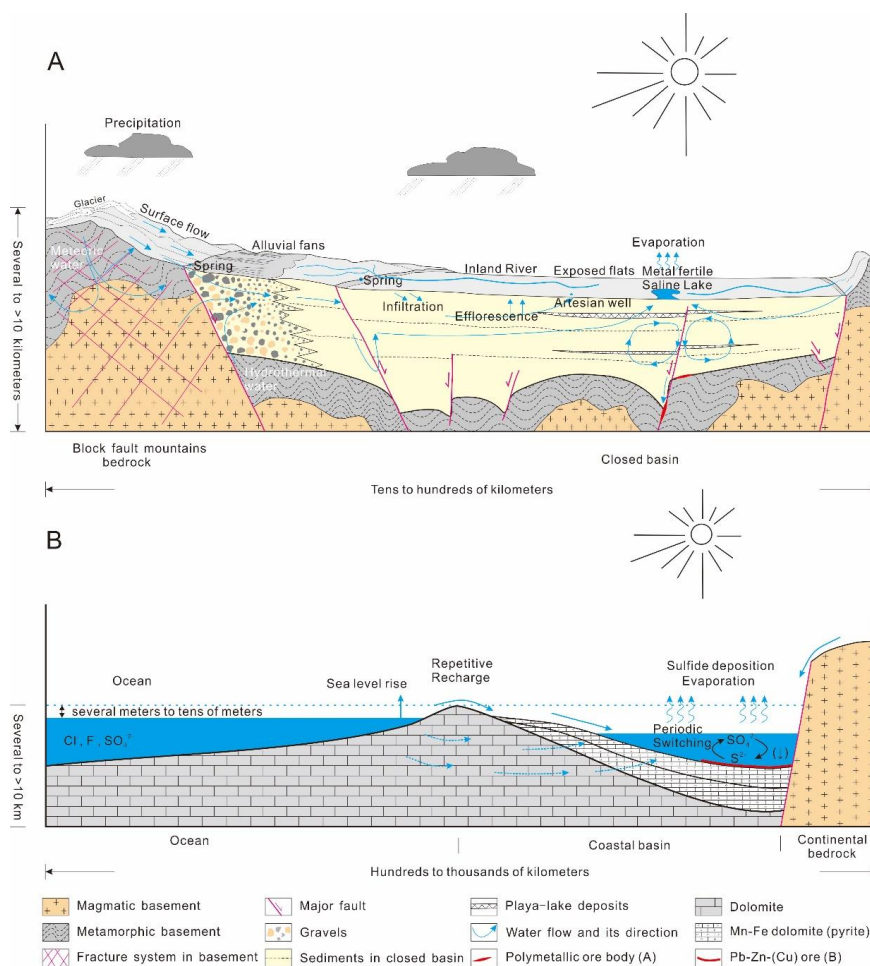
164 **3. Formation Processes of Ore-Forming Fluids of Basin-related Ore**
165 **Deposits**

166

167 Ore-forming fluids responsible for the formation of basin-related ore deposits are commonly metal-
168 rich, hot, and of medium to high salinities (Kesler, 2005). The homogenization temperatures of fluid
169 inclusions of ore-stage minerals from these deposits typically range from 50°C to 250°C, as high as
170 300°C; however, most temperatures are between 100 and 200°C, such as those of MVT Pb-Zn
171 deposits, unconformity-related U deposits, vein type Pb-Zn and U deposits, and so on (e.g., Kesler
172 and Reich, 2006; Stoffell et al., 2008; Richard et al., 2012; Zhang et al., 2017, 2019). Salinities
173 commonly range from 10% to over 30% wt NaCl equiv (Basuki and Spooner, 2004; Leach et al.,
174 2005). The geochemical composition of the paleo-ore-forming fluids is remarkably similar to the



175 compositions of present-day basin brines (Carpenter et al., 1974; Kesler et al., 1996; Viets et al.,
 176 1996), although there is a relatively large temperature difference between the developing basin
 177 brines and paleo-ore-forming fluids.
 178



179
 180 Fig.1 Sketch map showing the formation processes of metal fertile saline brines in continental closed
 181 basin (A, modified from Zhang et al., 2022) and coastal basin (B, modified from Large et al., 2002
 182 and Zhang et al., 2022).

183
 184 It is generally assumed, from mineralogical and geochemical considerations, that the paleo-ore-
 185 forming fluids were principally hot metalliferous basinal brines (Badham, 1981; Lydon, 1983;



186 Cooke et al., 2000). However, this does not mean that the initial ore-forming fluid reached high
187 levels of trace metal compositions at high temperature conditions. There are three possibilities for
188 the formation procedures of hot metalliferous basinal brines: I. the initial fluid was metal-rich and
189 then heated to a high temperature; II. the initial fluid was heated to a high temperature first, and then
190 obtained a large amount of trace metals from regional lithologies; III. the initial fluid obtained the
191 trace metals and was heated simultaneously.

192

193 The author first discusses the possibility of surficial saline brines acting as ore-forming fluids. Many
194 researchers have confirmed that the surficial saline brines in closed basins are metal-rich, such as
195 the ephemeral Lake Merouane Chott in North Africa ($U > 20$ ppm, $Cr > 100$ ppm, $Cu > 200$ ppm,
196 $Co > 10$ ppm, $Pb > 2$ ppm, $V > 2000$ ppm, and $Zn > 200$ ppm; e.g., Hacini and Oelkers, 2010), various
197 soda saline lakes in Mongolia ($U > 20$ ppm, $Mo > 5$ ppm; e.g., Volkova, 1998; Linhoff et al., 2011;
198 Shvartsev et al., 2014; Zhang et al., 2022), and saline lakes on the Tibet Plateau ($W > 5$ ppm, $Mo >$
199 5 ppm, $REE > 10$ ppm; Zheng et al., 1989). These metalliferous saline brines are the products of a
200 complex system involving precipitation, weathering, precipitation-dissolution reactions,
201 groundwater circulation and water-rock interactions, evaporations, and probably biological activity
202 (Carroll and Bohacs, 1999). Such metal-rich saline brines usually exist in the form of saline lakes,
203 and mainly develop in arid to semi-arid environments. The main influence on the formation of these
204 brines is hydrological closure; that is, evaporation is the primary or only way for water to leave the
205 converging area (Deocampo and Jones, 2014). The balance between water inflow and evaporation
206 over time is the major factor causing ions to accumulate (Jones and Bodine, 1987) by exceeding the
207 saturated solubility of various minerals, usually beginning with the alkaline-earth carbonates, then
208 progressing to gypsum, and then to halite (Warren, 2014). In this enriching process, trace metals
209 build up to an unusually high level (detailed examples in Zhang et al., 2022, 2023a, b). The
210 mineralogical composition and order of salts precipitation in a surficial saline brine are controlled
211 by the ionic make-up of the mother brine, which is commonly the surficial and underground runoff
212 (Hardie and Eugster, 1970; Hardie et al., 1978).

213

214 The surficial saline brines have high salinities with total dissolved solids of more than 35 g/L but



215 commonly less than 400 g/L (Deocampo and Jones, 2014; Zhang et al., 2022). The high salinities
216 mean higher density compared to fresh groundwater, which drives the surficial saline brines to
217 circulate into the deep area and thus changes the salinities and trace metals concentrations of
218 groundwater (Fig. 1A). During the circulation, these brines would be heated and pressurized, and
219 obtain and deposit some specific metal ions, and alter the basement through water-rock chemical
220 interactions (Hecht and Cuney, 2000; Mercadier et al., 2010, 2012, 2013; Kyser, 2014). Over time,
221 the groundwater near saline lakes would eventually be transformed into metalliferous, pressurized,
222 and hot salinized fluids.

223

224 The nearby sedimentary strata may become metalliferous fluid-hosting aquifers, such as the
225 Mississippi Salt Dome (Kharaka and Thordsen, 1992; Kesler et al., 1996). These brines could be
226 expelled from the aquifers during physical and/or chemical compaction and/or diagenesis, and
227 migrate into structural conduits such as faults, creating opportunities for vein-type mineralization
228 (Frape et al., 2014; Heinrich and Candela, 2014). If the basins are coastal, the formation of metal-
229 fertile brines sourced from seawater requires (1) a steady influx of seawater, (2) multiple intervals
230 of isolation from the ocean, (3) a saline-lake elevation lower than sea level, in addition to an arid,
231 evaporative environment (Fig.1B; e.g., Messinian salinity crisis in the Mediterranean region,
232 Krijgsman et al., 1999; Gillet et al., 2007).

233

234 In contrast, more researchers prefer Possibilities II and III, which suggest that the initial fluid
235 obtained trace metals at high temperature conditions, such as the “Reflux brine model for Pb-Zn
236 deposits in Passive Continental Margin” (Leach et al., 2005) and “Seawater extracting metals from
237 redbeds model for sediment-hosted stratiform copper deposits” (e.g., Hitzman et al., 2005; Robb,
238 2005). The trace metals were extracted from the surrounding rocks, such as redbeds, marine
239 sedimentary rocks, and wall rocks of regional faults (e.g., Hitzman et al., 2005; Kyser, 2014). Some
240 of these models require the existence of evaporites. Direct dissolution of evaporites in the flow path
241 of meteoric/marine waters has been the preferred source for dissolved salts (Goodfellow et al., 1993;
242 Lydon, 1995; Emsbo, 2009), and then extract the trace metals from regional lithologies during
243 groundwater circulation to become the metal-rich fluid (e.g., Hitzman et al., 2005; Kyser, 2014). It



244 is difficult to verify the authenticity of these models (Possibilities No. II and III) because the
245 extraction of trace metals occurs several kilometers underground and is unobservable directly.
246 However, the authenticity of these models can be indirectly verified by statistics of paleo-latitude
247 of the basin-related ore deposits during their mineralization. For Possibility I, the initial ore-forming
248 fluid formed in arid to semi-arid environments, which is dominantly controlled by the latitude.
249 However, for Possibilities II and III, the formation of initial ore-forming fluid was independent of
250 latitude, although some of these models need the existence of evaporites. The detailed latitude
251 statistics of basin-related ore deposits are below.

252

253 4. Global Distribution of Metal-rich Saline Lakes and Its Major Controlling

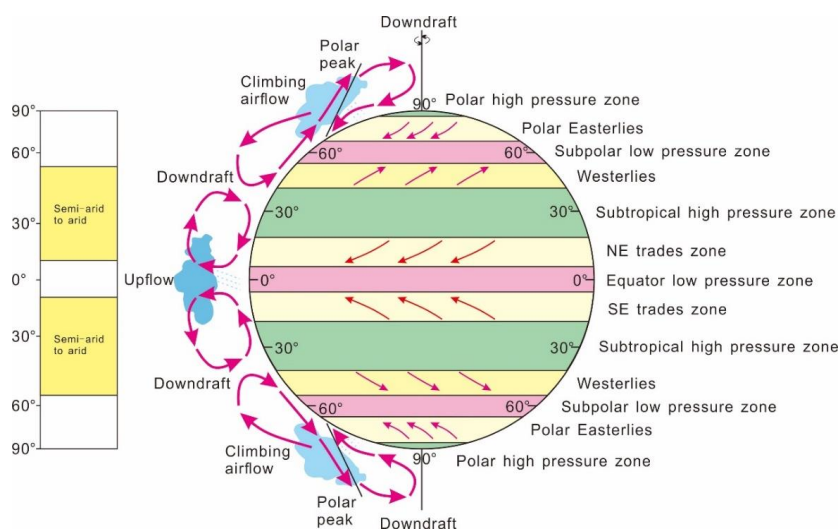
254 Factors

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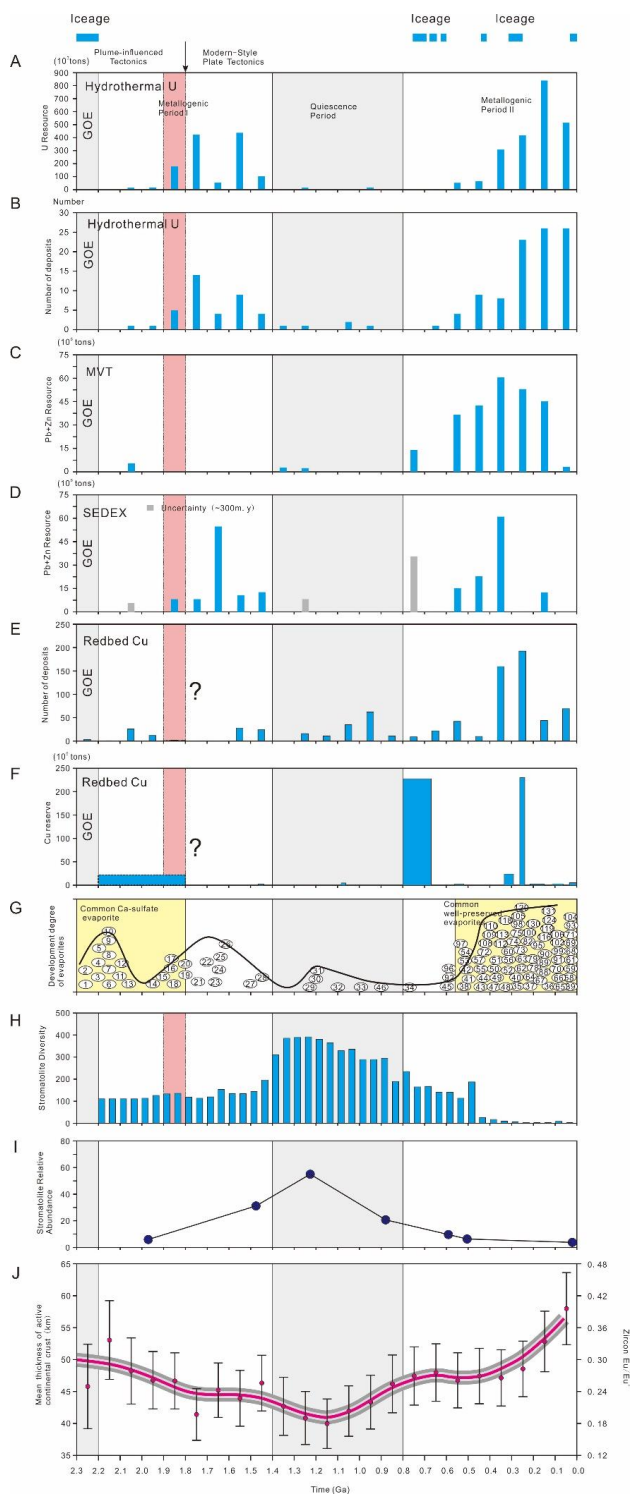
256 The current distribution of metal-rich saline lakes and their controlling factors could reveal the
257 factors that control the global distribution pattern of basin-related metallogenesis. Climate plays a
258 critical role in the formation of saline lakes; the amount of inflow to a hydrologically closed basin,
259 mainly precipitation, must be closely balanced by evaporative loss, meaning that saline lakes are
260 mainly found in semi-arid to arid environments (Deocampo and Jones, 2014). In the Quaternary, the
261 saline lakes are mainly zonally distributed, as revealed by the global distribution pattern of
262 Quaternary evaporites (Kottek et al., 2006; Warren, 2014). In terms of latitude, the saline lakes are
263 dominantly located in the range of 5° to 50° in the Northern and Southern Hemispheres (Cooke and
264 Warren, 1973; Smoot and Lowenstein, 1991). This zonal distribution pattern of saline lakes is the
265 comprehensive result of many factors, including the zonal distribution of solar radiation over the
266 surface of the Earth, the distribution of land masses and their elevations, and ocean and atmospheric
267 circulation and interaction, among other causes (Mackenzie, 2003). These factors combine to
268 control the global patterns of winds, the balance between precipitation and evaporation, and
269 humidity, cloudiness, air pressure, and temperature (Kump et al., 2010). However, the primary factor
270 is the zonal difference of solar radiation over the Earth. The high flux of solar radiation at the equator
271 results in air rising in the equatorial zone and moving toward the poles, and cooler air from the poles
272 moving toward the equator. The warm, moist, rising air forms large cumulus up into the high



273 troposphere, and is the source of precipitation in the equatorial zone. On the way to the poles, the
 274 dry air cools and begins to sink, forming the subtropical high pressure zone, which is commonly dry
 275 and rainless. This is the reason that saline lakes are commonly found in the low-mid latitudes (Fig.2).
 276 Then, the cooler air descends and moves toward the equator, creating the trade winds of both
 277 hemispheres, also known as the “Hadley cell” (Fig. 2). Due to the rotation of the Earth (termed the
 278 “Coriolis Effect”), the wind systems in the subtropical high pressure zone are prevailing westerlies.
 279 In contrast, very cold and dry air that forms in the poles moves toward the equator. On the way, the
 280 cold air uplifts the warm air in the lower latitudes. As warm air rises, it carries water vapor upward
 281 into the atmosphere, creating precipitation (Schneider and Londer, 1989). With consideration of
 282 temperature and associated evaporation, the arid to semi-arid zones are mainly located in the low
 283 and middle latitudes (Fig. 2). In addition, the distance to the ocean is another important factor;
 284 rainfall is much less in the interior of the continent than in coastal areas, which is caused by being
 285 far away from the water vapor source. The saline lakes are also located on the leeward slope of
 286 mountains, and may also unexpectedly appear on the west coast of the continent, where there is a
 287 middle latitude low pressure system towards the equator (Warren, 2014).



288
 289 Fig.2 Distribution of modern climatic belts showing the Earth’s atmosphere circulation cells across
 290 90°N to 90°S. Belts of cool dry descending air around 15–45°N and S of the equator create the main
 291 arid zones of the world (e.g., Trewartha and Horn, 1980; Schneider and Londer, 1989; Darnell et al.,
 292 1992).



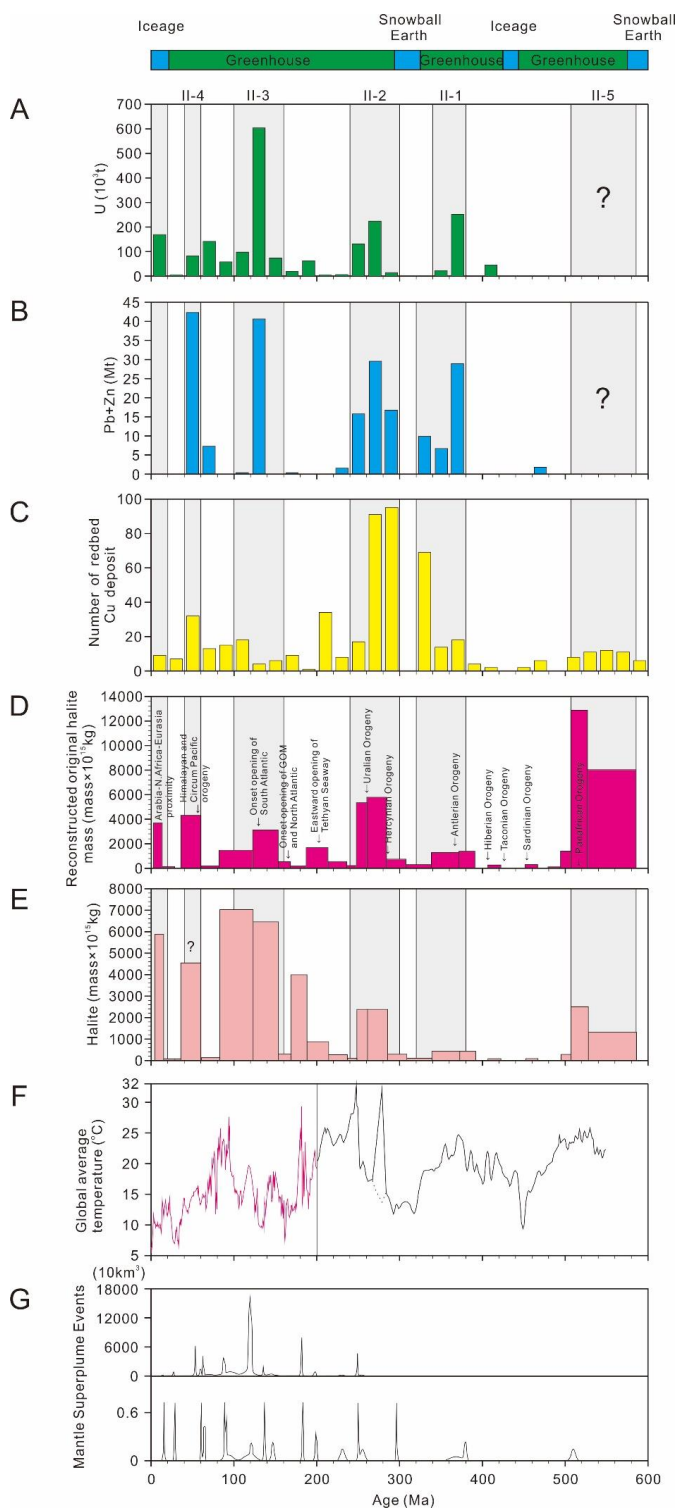


294 Fig.3 Occurrence (in terms of relative abundances) of basin related hydrothermal type U deposits
295 (U resource for A, and Number of deposit for B), MVT (Pb+Zn resource, C, Leach et al., 2005) and
296 SEDEX (Pb+Zn resource, D, Leach et al., 2005) Pb-Zn deposits, sediment-hosted stratiform Cu
297 deposits (Number of deposits for E, and Cu reserve for F, Kirkham, 1989; Kirkham et al., 1994;
298 Hitzman et al., 2010), saline giants (G, Bekker et al., 2006; Evans, 2006; Bekker and Holland, 2012;
299 Ba, bel and Schreiber, 2014), Stromatolite relative diversity (H, Awramik and Sprinkle, 1999) and
300 abundance (I, Walter and Heys, 1985; Sheldon, 2013), and reconstructed thickness of active
301 continental crust since 2.3Ga (J, Tang et al., 2021a, b). The basin-related mineral resources are
302 mainly distributed in two stages which are 2.2-1.4 Ga, and 0.8-0.6 Ga, with one basin-related
303 metallogenic gap period of 1.4-0.8 Ga. While the metallogenic gap period has witnessed the great
304 prosperity of stromatolite in abundance and diversity. The global distribution pattern of basin-related
305 ore deposits are showing similar shape with the saline giants, and reconstructed thickness of active
306 continental crust since 2.3Ga.

307

308 The saline lakes that existed during geological history have disappeared, leaving evidence of their
309 existence, such as evaporite salts and associated redbeds (Warren, 2010). Evaporites are believed to
310 be important in the genesis of metal-fertile basin brines, as they are a key feature of basins hosting
311 supergiant deposits. However, the exact genetic relationship between evaporites and ore deposits is
312 controversial (e.g., Robb, 2005; Groves et al., 2010; Leach et al., 2010; Zhang et al., 2022). Ancient
313 evaporites are varying combinations of gypsum or anhydrite, alkaline earth carbonates, and halite,
314 with or without potash-bearing minerals such as sylvite, carnallite, and epsomite (e.g., Pope and
315 Grotzinger, 2003; Bekker et al., 2006; Schroder et al., 2008). With the improvement of the
316 availability of information in recent years, the global statistics of evaporites are becoming
317 increasingly convincing. The author has compiled evaporite sediments of the world based on
318 previous studies of Pope and Grotzinger (2003), Bekker et al. (2006), Schroder et al. (2008), Warren
319 (2010, 2014), and others (Fig.3, 4, and 5).

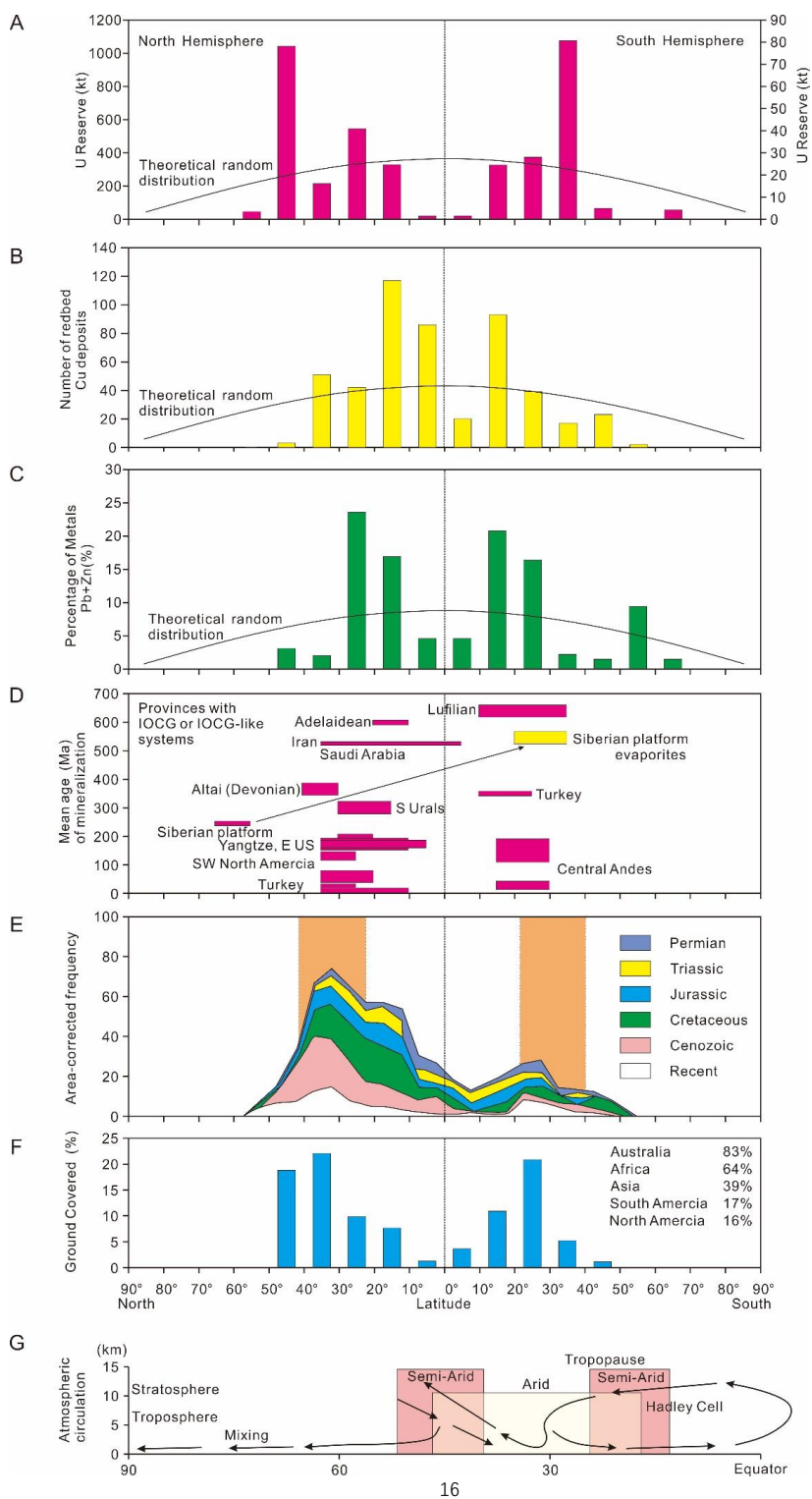
320





322 Fig.4 Temporal distribution (in terms of relative abundances) of basin related hydrothermal type U
323 deposits (A, Dahlkamp, 2009, 2010, 2016; Zhang et al., 2022), MVT and SEDEX Pb-Zn deposits
324 (B, Leach et al., 2005, 2010), sediment-hosted stratiform Cu deposits (C, Kirkham, 1989; Hitzman
325 et al., 2005), saline deposits (D and E, Hay et al., 2006; Warren, 2010), global paleo-average
326 temperature (Prokoph et al., 2008; Donnadieu et al., 2008; Ernst, 2014; Spray, 2020; Scotese et al.,
327 2021), and large igneous provinces (Ernst and Buchan, 2001; Prokoph et al., 2013) since ~600Ma.
328

329 The compilation of evaporite sediments has shown that there are mainly two development stages
330 (Fig.3). The first stage was about 2.3-1.4Ga and was represented by pseudomorphs after gypsum,
331 anhydrite, and halite (Zharkov, 2005). The second stage is Phanerozoic and is represented by well-
332 preserved gypsum, anhydrite, and halite. During most of the Phanerozoic, the Earth was in a
333 greenhouse state, except for part of the Permo-Carboniferous. In the greenhouse times, there were
334 no continental or polar ice sheets, and the planet's meridional temperature gradient was much less
335 steep than it is at present (Warren, 2006, 2010). In this context, the Hadley-related circulation belts
336 would be pushed away from the equator, global water recycling would be sped up, and many
337 continental deserts would enter relatively humid environments than the Earth in the glacial age (Yan
338 et al., 1998; Baker et al., 2001; Hesse et al., 2004). Therefore, saline lakes would be much more
339 numerous and distributed more widely in the greenhouse stage than the Earth in the glacial stage.
340 On the whole, the saline lakes during Phanerozoic were produced in the low to middle latitudes, as
341 shown by the existence of evaporites (Fig.5). With estimates of the volume of salt being recycled
342 back into the oceans added to the volumes of actual salt in various megahalite beds of the world, the
343 original halite volume during the Phanerozoic can be reconstructed (Fig.4; Hay et al., 2006; Warren
344 et al., 2010). This reconstructed original halite mass distribution can reveal the volumes of saline
345 lakes during Phanerozoic generally. There are mainly seven periods of widely development of saline
346 lakes on the Earth during Phanerozoic, which are 580-510Ma, 390-340Ma, 300-240Ma, 210-190Ma,
347 160-80Ma, 60-40Ma, and ~20Ma, respectively, corresponding to the greenhouse stages (Fig.4).
348





350 Fig.5 Paleo-latitude distribution (in terms of relative abundances) of basin related hydrothermal type
351 U deposits (A, Dahlkamp, 2009, 2010, 2016; Zhang et al., 2022), sediment-hosted stratiform Cu
352 deposits (B, Kirkham, 1989; Hitzman et al., 2005), MVT and SEDEX Pb-Zn deposits (C, Leach et
353 al., 2005, 2010), IOCG deposits (D, Barton, 2014), Phanerozoic saline deposits (E, Hay et al., 2006;
354 Warren, 2010), and Quaternary deserts (F, Prokoph et al., 2008; Donnadiou et al., 2008; Ernst, 2014;
355 Spray, 2020; Scotese et al., 2021) since ~600Ma.

356

357 5. Uneven Distribution of Basin-Related Ore Deposits in Time and Space

358

359 The previous compilations of basin-related ore deposit types as a function of geological time by
360 Meyer (1981, 1988), Veizer et al. (1989), Barley and Groves (1992), and Leach et al. (2005) have
361 revealed that these deposits developed during particular periods of Earth's evolution, and the global
362 pattern of basin-related mineral deposits is extremely uneven in time and space (Garnett and Bassett,
363 2005; Large et al., 2005; Cuney, 2010). With the improvement and increase of information on basin-
364 related ore deposits, it is necessary to conduct a more comprehensive statistical analysis. Here, the
365 author has compiled statistics for the world's major hydrothermal-type U ore deposits related to
366 basins (Appendix A), the previous compilation of SEDEX and MVT Pb-Zn deposits during the
367 Proterozoic and Phanerozoic (Leach et al., 2005; Zhang et al., 2022), sediment-hosted Cu deposits
368 (Hitzman et al., 2005), and the world's major saline deposits (Appendix B) since the global oxidation
369 event (GOE). The temporal and spatial distribution of these basin-related ore deposits is highlighted
370 (Fig. 3, 4, and 5).

371

372 5.1 Metallogenic Peak of 1.8-1.5 Ga

373

374 The compilation results of basin-related ore deposits show that these types of ore deposits increased
375 gradually in quantity and resource to a peak of about 1.8-1.5 Ga, with the onset of the Great
376 Oxidation Event (Fig. 2). During this peak, a large number of basin-related ore deposits developed
377 worldwide, such as the world-class unconformity-related U deposits in Canada (e.g., Athabasca
378 basin, 1.8-1.7 Ga), Australia (e.g., Pink Creek Inlier U region, 1.8-1.7 Ga), and India (e.g.,



379 Singhbhum Cu-U belt, Cuddapah and Bhima U region, ~1760 Ma), world-class metasomatism type
380 U deposits in Europe (e.g., Kirovograd-Smolino U region, Krivoy Rog U region, 1.8-1.7 Ga, 1750-
381 1770 Ma; Lake Ladoga and Onega U districts, 1.8-1.7 Ga), China (e.g., metasomatism type U
382 deposits North China Craton, 1740-1800 Ma), Australia (Mount Isa region, 1750-1730 Ma), and the
383 famous vein type U deposits around the world (e.g., Beaverlodge area in Canada, Sierra
384 Ancha/Apache Proterozoic Basin in Arizona, Central Ceará U Region in Brazil, Longshoushan U
385 region in Northwest China; e.g., Dahlkamp, 2009, 2010, 2016). For the sediment-hosted Pb-Zn
386 deposits, the SEDEX Pb-Zn deposits show a similar variation trend and reach a peak at 1.7-1.6 Ga,
387 such as the world-class Pb-Zn deposits of eastern Australia (Mount Isa, Broken Hill, and McArthur
388 River, Bodon, 1998; Chapman, 2004; Large et al., 2005), South Africa (Aggeneys and Gamsberg
389 Pb-Zn districts, Rozendaal, 1980; Rozendaal and Stumpfl, 1984; Stalder and Rozendaal, 2004),
390 India (Zawar Pb-Zn district, Bhattacharya and Bull, 2010), and China (Yanliao Pb-Zn region, Zhong
391 et al., 2012; Duan et al., 2017). In contrast, during this period, the famous sediment-hosted stratiform
392 Cu deposits were much less than the basin-related Pb-Zn and U deposits. The statistics of sediment-
393 hosted stratiform Cu deposits are mainly based on the ages of ore-hosting rocks, and the distribution
394 pattern of the sediment-hosted stratiform Cu deposits during 2.2-1.4 Ga is different from that of the
395 basin-related Pb-Zn and U deposits. Post the Great Oxidation Event, the first famous supergiant
396 stratiform Cu deposits were contained in the Paleoproterozoic Kodaro-Udokan basin of Siberia. The
397 age of the ore-hosting sequence is poorly constrained, ranging from 2200 to 1800 Ma (Abramov,
398 2008), and the mineralization age is unknown (Volodin et al., 1994). The second famous ore-hosting
399 sequence of stratiform Cu deposits is the Revett Formation, which hosts deposits such as Spar Lake
400 in the NW United States. The Revett Formation is about 1460 Ma, constrained by the underlying
401 1468 Ma lower Belt sequence (Anderson and Davis, 1995) and the overlying 1454 Ma carbonate
402 sequence (Evans et al., 2000).

403

404 **5.2 Metallogenic Quiescence during 1.4-0.8 Ga**

405

406 Irregular decline of basin-related ore deposits occurred during 1.5-1.4 Ga, and then the Earth entered
407 a basin-related metallogenic quiescence period until about 0.8 Ga. This metallogenic quiescence



408 was also stressed by Bekker et al. (2010) to reflect a gap of Superior-type iron deposits and
409 manganese ore deposits (Maynard, 2010). During this quiescence period, the global basin-related
410 metallogeny was maintained at a low level, and a much smaller number of basin-related ore deposits
411 were formed in this period in contrast to the period 1.8-1.5 Ga. For the hydrothermal type U deposits,
412 most of the reported U mineralization ages during this period were from the Paleoproterozoic
413 unconformity-related U deposits, and were interpreted to be the reformation age of the existing
414 deposits or the new mineralization events that superimposed on the old deposits, such as the vein
415 type U deposits in the Great Bear Batholith in Canada (e.g., ~1076Ma, Chi et al., 2018; Dahlkamp,
416 2016), unconformity-related U deposits in the Franceville region of Africa (e.g., 890-860Ma, Cuney
417 and Mathieu, 2000), and vein type U deposits in the Kombolgie basin in Australia (e.g., ~1040Ma,
418 Rawlings and Page, 1999; Polito et al., 2006a, b; Cuney and Kyser, 2008). The sediment-hosted
419 stratiform Cu deposits were represented by the Mesoproterozoic Keweenaw basin of the Mid-
420 Continent Rift of the USA (White Pine, Brown, 1971; White, 1971; Mauk et al., 1992). Moreover,
421 the reported MVT and SEDEX Pb-Zn deposits were fewer during the quiescence period (Fig. 3).

422

423 **5.3 Metallogenic Period of Phanerozoic**

424

425 The quiescence period ended at about 0.8-0.7 Ga. From 0.8 to 0.5 Ga, the basin-related ore deposits
426 began to increase in quantity and resource (Fig. 3), and the Earth entered a basin-related
427 metallogenic period again. The second metallogenic period was mainly the Phanerozoic. Due to the
428 lack of absolute values of resources of sediment-hosted stratiform Cu deposits, the author compiled
429 the number of Cu deposits that differed from the basin-related Pb-Zn and U deposits, both of which
430 were compiled in their resources (Fig. 4 and 5). The histogram of resource of basin-related U
431 deposits and Pb-Zn deposits, and number of Cu deposits that formed within 600 Ma, referring to
432 their formation or ore-hosting rock ages, show that these types of ore deposits are strikingly similar
433 to each other (Fig. 4). The metallogenic peaks of basin-related U deposits are nearly coincident with
434 the Pb-Zn deposits, and partially coincident with the stratiform Cu deposits. There are five
435 metallogenic peaks (Fig. 4) that include ~380-340 Ma (II-1), ~300-240 Ma (II-2), ~160-100 Ma (II-
436 3), 60-40 Ma (II-4), and one stratiform Cu metallogenic concentration period of ~580-500 Ma (II-



437 5).

438

439 During the metallogenic peak of II-1, the basin-related U deposits were represented by the
440 Kokshetau and Kendyktas-Chuily-Betpak Dala U region of Kazakhstan, Eastern Karamazar-
441 Northeastern Fergana U Region of Kyrgyzstan, and Schwarzach Valley U region of Germany (e.g.,
442 Dahlkamp, 2010). The Pb-Zn deposits were represented by the East Tennessee and Pine Point Pb-
443 Zn district of the United States, and Lennard Shelf Pb-Zn region of Australia (e.g., Leach et al.,
444 2005, 2010). The sediment-hosted stratiform Cu deposits are represented by the famous Cha-Sarysu
445 Cu region of Kazakhstan, and numerous Cu deposits in Europe and the United States (e.g., Kirkham,
446 1989; Kirkham et al., 1994; Hitzman et al., 2010). The second metallogenic peak (II-2) was
447 represented by the supergiant stratiform Cu deposit of Kupferschiefer in Europe, and the East
448 Tennessee Pb-Zn region, Southeast Missouri Pb-Zn region, Tri-State Pb-Zn region, North Arkansas
449 Pb-Zn region, and Upper Mississippi Valley Pb-Zn region of the United States, and numerous U
450 regions in Europe (e.g., Dahlkamp, 2009, 2016; Hitzman et al., 2010; Leach et al., 2010). The
451 metallogenic peak of II-3 and II-4 were represented by the Upper Silesia Pb-Zn region in Europe,
452 Cevennes Pb-Zn region, Pine Point Pb-Zn region in the United States, South China metallogenic
453 province, Streltsovsk U district, Elkon U district, Bureinsky U District in the Russian Federation,
454 and numerous stratiform Cu deposits around the world (e.g., Dahlkamp, 2009, 2016; Hu et al., 2008,
455 2017; Hitzman et al., 2010). The temporal distributions of Phanerozoic hydrothermal type U
456 deposits, SEDEX and MVT Pb-Zn deposits, and stratiform Cu deposits show a similar pattern with
457 the Phanerozoic distribution of global evaporites (Warren, 2010; Fig.4).

458

459 The author has also compiled the paleo-latitudes of the Phanerozoic basin-related ore deposits
460 during their mineralization (Fig. 5). The paleo-latitudes of basin-related U deposits, SEDEX and
461 MVT Pb-Zn deposits, and stratiform Cu deposits are similar to each other. Most of the Phanerozoic
462 basin-related ore deposits are formed at low to middle latitudes, with few deposits formed at middle-
463 high latitudes (e.g., Barton and Johnson, 1996; Scotese et al., 1999; Soloviev, 2010a, b; Leach et al.,
464 2010; Boucot et al., 2013). The formation latitudes of the basin-related ore deposits generally
465 coincide with the broadly contemporaneous evaporitic settings (Fig. 2 and 5). Previous researchers



466 have often emphasized the Earth's evolving tectonic setting and geochemical systems; many studies
467 have proposed that the large-scale formation of basin-related ore deposits is related to global tectonic
468 processes and patterns (orogenic processes, Wilson cycle, e.g., Nance et al., 1986; Larson, 1991;
469 Barley and Groves, 1992; Barley et al., 1998; Kerrich et al., 2000, 2005; Goldfarb et al., 2005).
470 However, the global tectonic evolution cannot explain the uneven distribution of basin-related ore
471 deposits in time and space well. Admittedly, many basin-related ore deposits have been destroyed
472 by subduction and erosion, or modified by post-ore metamorphism and tectonism, so that they are
473 no longer recognizable (e.g., Veizer et al., 1989; Bradley, 2008). However, some basic distribution
474 characteristics should still be retained; for example, if there were basin-related ore deposits formed
475 at high latitudes (e.g., $>70^\circ$), it is impossible to destroy them all; if there were a significant amount
476 of basin-related ore deposits formed during 1.4-0.8 Ga, it is impossible that all the hydrothermal
477 type U deposits, SEDEX and MVT Pb-Zn deposits, and stratiform Cu deposits showed one
478 mineralization quiescence during this period (Fig.2).

479

480 6. Discussion

481

482 6.1 Another explanation for the genesis of basin related Pb-Zn deposits

483

484 The metal-fertile saline brines that are contemporaneous with evaporites are probably the initial ore-
485 forming fluid of basin-related ore deposits (e.g., Zhang et al., 2022, 2023a, 2023b). This view comes
486 from current research on saline lakes. The buildup of metals in basinal saline brines was mainly
487 caused by the balance between water inflow and evaporation over time (Jones and Bodine, 1987).
488 During the evaporation, the commonly enriched ions include Na^+ , K^+ , Ca^{2+} , Mg^{2+} , Cl^- , HCO_3^- , SO_4^{2-}
489 (S^{2-} and/or HS^-), and sometimes Li^+ (e.g., Zheng, 1997; Dugamin et al., 2021). Due to the
490 evaporation over time, some dissolved metals can accumulate by exceeding the saturated solubility
491 of various minerals, usually beginning with alkaline-earth carbonates, then progressing to gypsum,
492 and then to halite and sylvite (Warren, 2014). These chemical deposition processes generally limit
493 the salinities of saline brines to not exceed 40%. In this enriching process, some trace metals build
494 up to unusually high levels, such as U concentrations of Lake Shuiquanzi (China) reaching 35.0



495 mg/L, Pb and Zn concentrations of the central Mississippi Salt Dome Basin (US) reaching 100.0
496 mg/L and 250.0 mg/L, respectively (Hagni, 1983; Kharaka et al., 1987; Kharaka and Hanor, 2014),
497 and Cu concentrations of Lake Merouane Chott (northern Africa) reaching 800.0 mg/L (Hacini and
498 Oelkers, 2010), and so on.

499

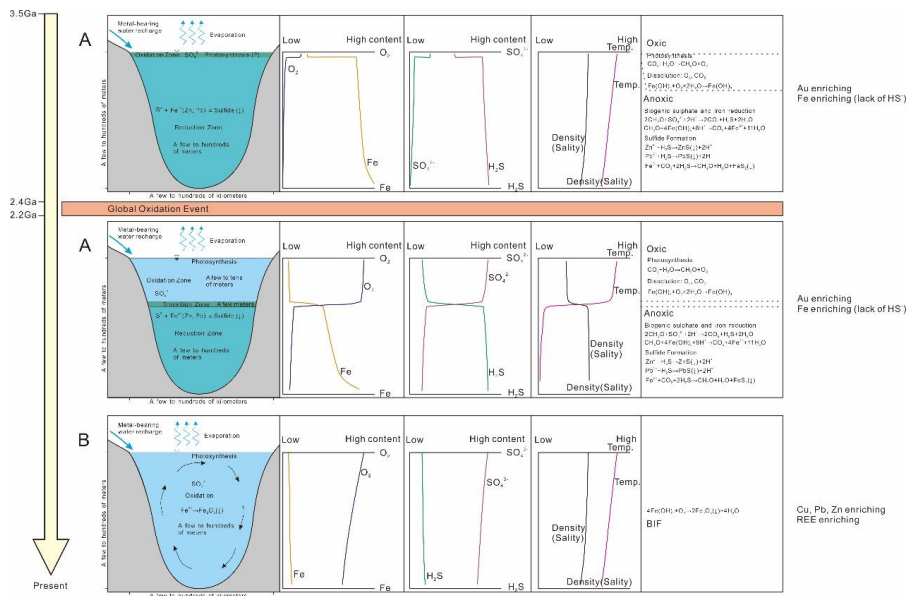
500 The ores of basin-related Pb-Zn deposits (MVT and SEDEX) throughout the world are commonly
501 fine-grained, banded, or colloform and dendritic in texture, which highly suggests sulfide deposition
502 in an open space (McLimans et al., 1980; Leach and Sangster, 1993). The parallel bedding sulfides
503 are assumed to be permissive evidence for exhalative ores (Leach et al., 2005). It is true that metal-
504 rich exhalative fluids have been found in seafloor spreading centers, such as the Atlantis II Abyss
505 of the Red Sea (Thisse et al., 1982, 1983; Barrett et al., 2021) and the Galapagos Spreading Center
506 in the Eastern Pacific Ocean (Perfit et al., 1983, 1999). However, most of the Pb-Zn deposits
507 classified as SEDEX lack unequivocal evidence of an exhalite in the ore or alteration component
508 (Leach et al., 2005, 2010). In addition, the bedrocks of the Pb-Zn ores found in seafloor spreading
509 centers include pillow lavas, pelagic sediments, and hyaloclastites. In contrast, most basin-related
510 Pb-Zn deposits are thought to have been formed in extensional basins associated with orogeny,
511 intracontinental and intracratonic failed rifts (Ervin and McGinnis, 1975), and continental marginal
512 basins (Leach et al., 2005; Large et al., 2005). The bedrocks are commonly shoal and/or neritic
513 carbonate rocks and continental clastic rocks (Leach et al., 2005). There are few basin-related Pb-
514 Zn deposits thought to have been formed in ocean spreading centers. Therefore, researchers should
515 look for another reasonable explanation. There is no denying that the exhalative process can result
516 in the formation of laminated Pb-Zn ores. Here, the author proposes another simple explanation,
517 referring to the studies of modern saline lakes.

518

519 What economic geologists who focus on basin-related ore deposits easily ignore is that condensed
520 seawater and saline brines can be significantly enriched in metals and sulfates (e.g., Volkova, 1998;
521 Hacini and Oelkers, 2010; Linhoff et al., 2011; Shvartsev et al., 2014). Previous studies of saline
522 lakes have identified that many of them show vertical stratification of their water masses over
523 extended periods (e.g., Lake Walker, Lake Mono, Lake Great Salt, Domagalski et al., 1990; Boehrer



524 and Schultze, 2008; Deocampo and Jones, 2014; Fig.6). This stratification includes density,
 525 temperature, dissolved substances, oxygen fugacity, and even living organisms. It can be caused by
 526 many factors, such as higher atmospheric temperatures during the warm season (e.g., Peeters et al.,
 527 2000) and external inflow of river water with lower density into saline lakes (e.g., Boehrer and
 528 Schultze, 2008). In some cases, stratification of saline lakes is temporarily destroyed by full or
 529 partial overturning, which may occur as a result of seasonal winds or seasonal temperature change
 530 (especially in the transition from the warm season to cold season; e.g., Steinhorn, 1985). The Dead
 531 Sea in the Middle East is a representative example of the occurrence of switching between
 532 stratification and full circulation in saline lakes; this switch occurs once a year, with stratification
 533 being weak in spring, autumn, and winter, but fully developed in summer (Steinhorn, 1985; Bartov
 534 et al., 2002; Arnon et al., 2019). The Jordan River was the water recharge source for the Dead Sea.
 535



536
 537 Fig. 6 Schematic diagrams showing redox status of saline lakes since the 3.5 Ga (A, appearance of
 538 photosynthetic algae), and 2.4-2.2 Ga (Global Oxidation Event, stratification of B and full overturn
 539 of C) to present (Boehrer and Schultze, 2008; Deocampo and Jones, 2014; Saleh et al., 2018; Arnon
 540 et al., 2019).
 541



542 After the GOE, SO_4^{2-} is a common ion in saline lakes, and the stratification of saline lakes would
543 transform the SO_4^{2-} into S^{2-} in the deeper reducing layer, following which the enriched sulfurophiles
544 (e.g., Pb, Zn, Fe) would be deposited as sulfides, including galena, sphalerite, and pyrite (Fig. 6;
545 Domagalski et al., 1990; Stoffell et al., 2008; Deocampo and Jones, 2014). When the stratification
546 is terminated (i.e., the saline lakes experience vertical circulation), the water mass becomes
547 oxidizing, SO_4^{2-} again becomes one of the major ions (Boehrer and Schultze, 2008), and sulfurophile
548 metal ions start to become enriched again (Steinhorn, 1985; Boehrer and Schultze, 2008). The
549 switching between stratification and full (or partial) overturn of saline lakes (condensed seawater)
550 would form alternating-layered sedimentary structures of sulfurophile sulfides and other sediments.
551 In this case, the laminated Pb-Zn ores would be formed (Schmalz, 1969; Fig. 6).

552

553 If the concentrations of S ($\text{SO}_4^{2-} + \text{S}^{2-}$) are relatively low, or the concentrations of sulfurophile ions
554 are much greater than those of S ($\text{SO}_4^{2-} + \text{S}^{2-}$) ions, the stratification of saline lakes would allow for
555 the enrichment of sulfurophile metal ions, commonly Fe^{2+} , in the deeper reducing layer. The full (or
556 partial) overturn of saline lakes would lead to the oxidation of Fe^{2+} to Fe^{3+} , which is then deposited
557 as hematite. In this case, the switch between stratification and full (or partial) overturn may form
558 the alternating-layered sedimentary structure of hematite and other sediments, which is likely the
559 origin of the structure of partial banded iron formations (Fig. 6C; Gutzmer et al., 2006; Boehrer and
560 Schultze, 2008; Deocampo and Jones, 2014). This is the reason that the SEDEX Pb-Zn deposits and
561 BIF deposits have similar sedimentary textures.

562

563 The surficial saline brines and evaporated seawater were commonly located above regional faults,
564 such as Lake Maharlou in the central Shiraz Basin of Iran (Fayazi et al., 2007), Lake Olduvai in
565 Tanzania (Blumenschine et al., 2012), and various saline lakes on the Tibet Plateau (Zheng, 1997).
566 Due to their high salinities, these brines have higher densities than those of fresh and brackish
567 groundwater and seawater. Therefore, these saline brines migrate downward into the underlying or
568 basal sequence of sedimentary basins. During the migration, the saline brines are heated and
569 pressurized, in some cases, by high heat flow and/or igneous activity, and there is an exchange of
570 elements between fluids and surrounding rocks, and/or metal deposition from metal-fertile saline



571 brines (Hoeve and Sibbald, 1978; Komninou and Sverjensky, 1996; Kyser, 2014; Chi et al., 2017).
572 These saline brines may utilize fault systems within and connect to the basin to escape to other
573 suitable regions and precipitate some specific minerals when they react with certain rocks. This is
574 the reason that the researchers can observe Pb-Zn ores with vein-like texture (Fig. 1; Zhang et al.,
575 2022). If the saline lakes are always in a non-stratified oxidation state, sulfurophile metal ions, such
576 as Pb^{2+} , Zn^{2+} , Cu^{2+} , Mo^{2+} , and so on, are enriched.

577

578 **6.2 The factors that influence the formations of Ore-forming fluid**

579

580 The factor that controls the formation of initial ore-forming fluid also influences the formation of
581 basin-related ore deposits (Zhang et al., 2022, 2023a, 2023b). The ancient metal-rich saline lakes
582 have disappeared, but they have left evidence of their existence, such as evaporites, fossils, and
583 residual metal-fertile brines (Babel and Schreiber, 2014). The metal-rich saline brines in
584 hydrologically closed basins are controlled jointly by arid-semiarid climate, tectonic setting, and
585 circulation processes underground (Carroll and Bohacs, 1999). However, the climatic setting is the
586 dominant factor, even if seawater was the starting fluid (Deocampo and Jones, 2014). The
587 distribution of global arid-to-semiarid zones is firstly controlled by the global atmospheric
588 circulation model, which means the arid-to-semiarid zones are closely related to the latitudes (Fig.
589 2; Trewartha and Horn, 1980; Schneider and Londer, 1989).

590

591 Geologists working with continental saline lakes have reached a consensus that these lakes, as well
592 as evaporites, form mostly in hot, arid to semi-arid zones within world-scale climatic regions known
593 as the horse latitudes (Warren, 2010). The term "horse latitudes" encompasses the regions beneath
594 the north and south Subtropical High atmospheric pressure belts (Oliver, 2005); see Warren (2010)
595 for a detailed explanation of the formation of the horse latitudes. For marine saline brines, at
596 particular times in the Earth's past, the hydrologies that facilitated the deposition of widespread
597 sulfate or halite evaporites across large parts of ancient salt basins were tied to platform or basin-
598 wide settings that have no same-scale modern counterparts (Fig. 5; Warren, 2010). The location of
599 marine Phanerozoic evaporites in zones of appropriate adiabatic aridity and continentality extended



600 well into the equatorial belts; therefore, in the Earth's past, the equatorial belts could also have been
601 the location of saline lakes (Ziegler et al., 2003). This is because evaporation rates of ocean waters
602 were at their maximum if the seawater were drawn into isolated evaporitic depressions (Kendall et
603 al., 2003; Warren, 2010). This is the reason that the paleo-latitudes of basin-related ore deposits
604 range from 60° north to 60° south, and there are hardly any basin-related ore deposits formed at
605 latitudes higher than 60° (Fig. 5). The paleo-latitudinal distribution of basin-related ore deposits
606 should not be explained by the selective preservation of basin-related ore deposits.

607

608 The factors that influence global climate include CO₂ concentrations in the atmosphere (the primary
609 driver of Phanerozoic climate, with large igneous provinces being the main source, according to
610 Royer et al., 2004 and Royer, 2016), plate tectonics and continental weathering (which influence
611 the long-term global temperature, Berner, 1994; van der Meer, 2017; Brune et al., 2017), the amount
612 of solar radiation, biological influences (e.g., C₄ grasses can reduce the atmospheric CO₂ contents
613 faster than C₃ plants, Lunt et al., 2007; Jiang, 2019), and ocean currents, which influence the global
614 climatic pattern (Holland et al., 1986; Hoffman and Schrag, 2002; Huybers and Langmuir, 2009;
615 Gnanaseelan and Deshpande, 2018). Of these factors, atmospheric CO₂ concentration might be the
616 most important (Anagnostou et al., 2016; Wally, 2018; Scotese et al., 2021).

617

618 High concentrations of atmospheric CO₂ result in a greenhouse effect, accelerating evaporation
619 globally (Linacre et al., 1970; Gilman, 1994). More importantly, these high concentrations result in
620 greater volumes of precipitation (Walker et al., 1981; Lackner et al., 2012), making surface waters
621 more acidic, leading to enhanced dissolution of calcium and silicates and accelerating the leaching
622 of specific metals from rocks, thereby enhancing the degree of enrichment of trace metals in saline
623 lakes under semi-arid climates (Almendinger, 1990; White and Buss, 2014). Greenhouse eustasy
624 (with associated epeiric seaways) favors the formation of surface saline brine as well as platform
625 evaporites (Warren, 2010).

626

627 Except for the planet-scale atmospheric circulation mainly caused by the varying intensity of solar
628 irradiation from the equator to the poles, there are three interrelated factors which influence the



629 broad scale distribution of the world's arid climate belts in time and their associated saline brines
630 and evaporitic sediments (Warren, 2010). These three factors are also controlled by the latitudinal
631 position of continental plates and variations in styles of plate-plate interaction. How these three
632 factors influence the distribution of arid belts was discussed in Warren (2010). Meanwhile, the
633 orogenic factor is also crucial for the development of arid to semi-arid regions. Likewise, orographic
634 deserts are better developed at times when two continental plates collide or a transcontinental rift
635 valley forms (Kendall et al., 2003; Warren, 2006). A rising mountain range, which is the direct result
636 of tectonic activity, is a barrier to atmospheric circulation, especially if perpendicular to the
637 incoming circulation (Sun et al., 2008). Tectonic activity is one dominant factor that influences the
638 formation of hydrological closed basins which are home to metal-rich saline brines (Zhang et al.,
639 2022, 2023a, 2023b). Therefore, in addition to climate, tectonic activity is the second prime factor
640 that influences the formation and distribution of continental and marine saline brines (Warren, 2010).

641

642 The global tectonics are governed by the Wilson Cycle of plate–plate interactions, which
643 encapsulates the notion that as new ocean basins open (rift of continents) via the formation of new
644 seafloor, oceans close due to the processes of subduction of oceanic crust, and ultimately by
645 continent–continent collision (Wilson, 1966; Nisbet and Fowler, 1983). Compilations of the time-
646 related distribution of ancient evaporites suggest there are four plate tectonic situations where
647 significant amounts of metal-rich saline brines could be produced at particular times in the past
648 (Jackson et al., 2000; Warren, 2010). These four tectonic settings are: (1) rifting basins where
649 continental plates are beginning to move apart, (2) foreland basins where continental plates are
650 coming close to each other, (3) intracontinental or intraplate sagging (intracontinental basins), which
651 are caused by the far-field effects of continental orogeny, (4) rapidly subsiding continental
652 depressions in transform or strike-slip settings where plates are sliding horizontally past one-another
653 (e.g., Kingston et al., 1983; Gasse and Fontes, 1989; Bosworth et al., 2005; Warren, 2006; Jordan et
654 al., 2007). These saline brines occur in both continental and marine realms (Warren, 2010). The
655 existence of metal-fertile saline brines is evidenced by the large volume of evaporites; see the
656 detailed reasons and formation processes of surficial saline brines in different tectonic settings in
657 Warren (2010). In summary, the factors that can affect global or regional climates indirectly affect



658 the formation of global or regional basin-related ore deposits.

659

660 **6.3 How the major geological events influence the formation and distribution of basin**
661 **related ore deposits globally?**

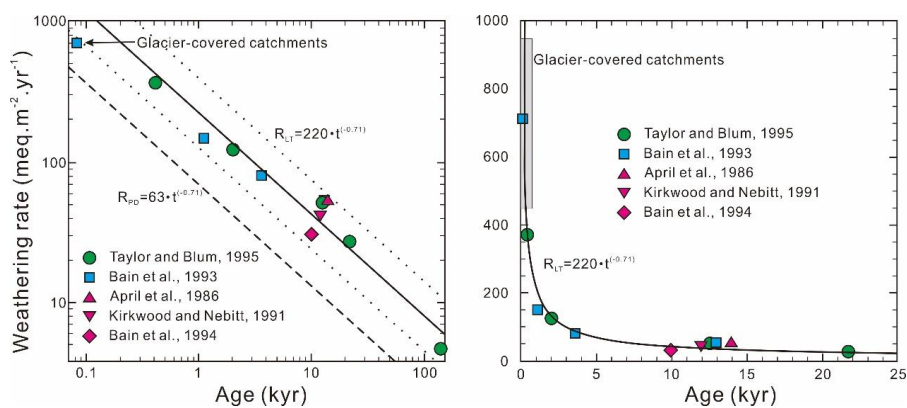
662

663 Major geological events that can significantly influence climates include mantle super-plume events
664 (or large igneous provinces), snowball Earth, and plate-plate tectonic events (Zhang et al., 2022,
665 2023a, 2023b). Mantle super-plume events create large volumes of igneous rocks (large igneous
666 province, Coffin and Eldholm, 1994; Ernst and Buchan, 2001; Courtillot and Renne, 2003),
667 triggering the input of large volumes of gas, including CO₂, into the ocean-atmosphere system
668 (Prokoph et al., 2008, 2013). As a result, CO₂ concentrations in the atmosphere increase to high
669 levels; for example, the Siberian large igneous province (ca. 251 Ma) is estimated to have released
670 $0.7\text{--}4.8 \times 10^7$ Tg CO₂, which could have raised the global temperature by 1.5–4.5 °C (Wignall, 2001;
671 Berner, 2002; White and Saunders, 2005; Tang et al., 2013). Mantle super-plume events thus
672 influence both atmosphere and ocean chemistry (Coltice et al., 2007; Ernst and Youbi, 2017).
673 Furthermore, the increase in global temperature during mantle super-plume events would increase
674 the rates of rainfall and evaporation, as well as the rate of weathering of silicate minerals, giving
675 rise to bursts of metal-fertile saline brines and evaporites (Fig. 4). In addition, snowball Earth events
676 could also result in a burst of global metallogenesis. Such an event dramatically slows the normal
677 consumption of CO₂ through chemical weathering of silicate rocks. However, shifting tectonic
678 plates would continue to generate volcanoes and supply the atmosphere with CO₂ (Hoffman et al.,
679 1998; Hoffman and Schrag, 2000, 2002). In this situation, CO₂ could accumulate to an extremely
680 high level (roughly 350 to 1000 times the present-day concentration of CO₂ in the atmosphere;
681 Hoffman and Schrag, 2000; Higgins and Schrag, 2003; Le Hir et al., 2009). The heat-trapping
682 capacity of CO₂ would warm the planet and begin to melt the ice that wrapped the planet. The thaw
683 would take only a short time (a few hundred years), after which Earth would enter a brutal
684 greenhouse stage (Allen and Hoffman, 2005; Font et al., 2010). Surface temperatures would soar
685 to >50 °C, driving an intense cycle of evaporation and rainfall (Hoffman and Schrag, 2000). Torrents
686 of carbonic-acid rain would erode exposed rocks, with the weathered metal ions being transported



687 into the oceans and (locally) lakes. In this case, there would most likely be more numerous and
 688 larger metal-bearing brine bodies on Earth's surface compared to the present. This explains the
 689 global burst of basin-related ore deposits in the period 295-260 Ma (e.g., Hein and Lehmann, 2009;
 690 Dolníček et al., 2009, 2014), following the snowball Earth event of the early Permian.
 691
 692 During Icehouse times (not Snowball Earth times), the oceanic eustatic response to rapidly waxing
 693 and waning polar ice sheets is excessively high amplitude and too short in time to allow continental
 694 (or periplatform) margin biogenic buildups to grow into continuous barriers (more than 100m of sea
 695 level oscillation every 100,000 years, Fischer, 1981; Warren, 2010). In contrast, the less than 5-10m
 696 of sea level change every 100,000 years during Greenhouse times are much more beneficial to the
 697 formation of barriers of platform-edge reefs and shoals (Fischer, 1981; Warren, 2010). These barriers
 698 are the key to the periodical hydrographic disconnection of back-barrier lagoons (periplatform) from
 699 the open ocean. This facilitated the creation of metal-fertile saline brines that originated from
 700 seawater, and therefore the formation of related SEDEX and MVT Pb-Zn deposits, stratiform Cu
 701 deposits, and so on (Warren, 2006, 2010).

702



703

704 Fig.7 A Log-log plot of base cation weathering rates vs. soil age (Blum, 1997). Data points are long-
 705 term weathering rates (RLT) from the references listed on the figure. The dark solid line is a power-
 706 law curve fit to the data from Taylor and Blum (1995, R = 0.99). Light dotted lines bracket curve fit
 707 to data and represent estimated maximum uncertainties of a factor of two in weathering rates.
 708 Question mark indicates two points from Bain et al. (1993) that are suspect on the basis of high Zr



709 contents of parent materials. Dashed line represents present-day weathering rates (RPD) calculated
710 from solid curve (see text). The shaded area represents the range of fluxes in glacier-covered
711 catchments compiled by Sharp et al. (1995) and are plotted at an arbitrary age. Fig.7 B Linear plot
712 of the same data and power law function as in Fig. 7A, but with the age axis truncated at 25 thousand
713 years.

714

715 For the mantle Super-plume event (or Large Igneous Provinces), Snowball Earth, and Plate-Plate
716 tectonic events, there is one collective result: these geological events can produce a large area of
717 exposed fresh bedrocks. Especially for the aftermath of Snowball Earth Events, chemical weathering
718 rates (or metal and base cation releasing rates) should be at their peak, as all the continents are
719 exposed with fresh bedrock (Fig. 7; Allen and Hoffman, 2005; Font et al., 2010). Because the
720 weathering rates for all of the base cations decrease with increasing soil age, such as the weathering
721 rates of soil age of 0.4 thousand years being at least two orders of magnitude higher than the soils
722 with age of 138 thousand years (Fig. 7; e.g., Bain et al., 1993; Sharp et al., 1995; Taylor and Blum,
723 1995). When the soil age is relatively large, such as bigger than 138kyr, the weathering rate would
724 be relatively low and stable (e.g., 1-5 meg/m².yr.; Blum, 1997). Actually, there are many factors
725 other than soil age that influence the chemical weathering rates, including human activities,
726 precipitation, temperature, vegetation, relief, silicate rock mineralogy, uplift rate, and glacial history
727 (Raymo and Ruddiman, 1992; Richter et al., 1992; Blum, 1997). However, it is extremely difficult
728 to confidently isolate the influence of the various factors affecting weathering rates. On the whole,
729 the relation between soil age and silicate weathering rates suggests one well-defined power law
730 relation for young soils (≤ 138 thousand years old), and the chemical weathering rates are
731 significantly accelerated in fresh bedrock outcropping environments (Fig. 7; Bain et al., 1993; Sharp
732 et al., 1995; Taylor and Blum, 1995). Therefore, the releasing rate of metals of continents is the
733 fastest following Snowball Earth Events, and the accumulation rate of metals elements in saline
734 lakes is the fastest, with consideration of the Greenhouse Effect of Snowball Earth Events.

735

736 In contrast to the Snowball Earth Event, Plate-Plate Tectonic Events, particularly plate convergence,
737 can result in the rapid uplifting of mountainous areas, which undergo high rates of mechanical



738 denudation, often predominantly by mass wasting (Derry and France-Lanord, 1997). In addition,
739 mountainous areas are more likely to experience earthquakes and mass failures, as well as, for
740 example, heavy rains and snow melts, all of which facilitate erosion. Mechanical erosion processes
741 periodically regenerate fresh mineral surfaces and expose them to the agents of chemical weathering
742 (Berner and Berner, 1997; Taylor and Blum, 1995). Furthermore, one of the consequences of
743 tectonic uplift is that mountain ranges are often raised to elevations that can support alpine glaciation.
744 The available data on chemical weathering rates of silicate catchments containing active alpine
745 glaciers has shown that they are elevated by at least an order of magnitude compared to nonglaciated
746 catchments, although the exact magnitude of this increase is difficult to estimate (Hallet et al., 1996;
747 Blum, 1997). Both glaciation and rapid uplift are accompanied by rapid physical denudation, which
748 exposes fresh bedrock minerals to the agents of chemical weathering. This means that the rapid
749 uplifting of mountains caused by plate convergence can quickly release dissolved metals into nearby
750 hydrological basins. Other indirect outcomes of uplift are the changes in mean annual temperatures
751 that accompany high elevation, and changes in precipitation patterns owing to orographic effects
752 (e.g., rain shadow effect, An et al., 2014; Caves, 2016). Therefore, metal-rich saline lakes are likely
753 to be produced in this geological context.

754

755 The Large Igneous Provinces are massive crustal emplacements of predominantly iron- and
756 magnesium-rich (mafic) rock that formed by processes other than normal seafloor spreading (Coffin
757 and Eldholm, 2005). They are the dominant form of near-surface magmatism on the terrestrial
758 planets, with areal extents greater than 0.1 million km² (Ernst and Jowitt, 2013). These occurrences
759 on the continents, such as the continental flood basalt provinces and volcanic passive margins well
760 preserved in the Mesozoic and Cenozoic periods, created a cover of fresh bedrocks. In this situation,
761 the large igneous provinces could also rapidly release base and metal cations into regional
762 hydrological closed basins. These are the reasons that the Snowball Earth events, large igneous
763 provinces, and plate-plate tectonic events are commonly related to global and regional basin-related
764 ore deposits in time and space (e.g., Ernst and Jowitt, 2013; Zhang et al., 2022).

765

766 Global tectonism controls the distribution of lithologic assemblages (both supracrustal rocks and



767 intrusive rocks), styles of deformation, and the grade, duration, and location of metamorphic events.
768 Rock assemblages that form in specific plate tectonic settings and which are available for weathering
769 exert control on the initial chemistry of water flowing into closed basins and, therefore, influence
770 subsequent brine chemistry (Deocampo and Jones, 2014; Zhang et al., 2022). In addition, the
771 geology of watersheds (drainage divides) is commonly affected by tectonic processes, which
772 directly affect the topography and lithologic assemblages (Pietras et al., 2003; Singh and Jain, 2009;
773 Sangma and Balamurugan, 2017). As a consequence, certain patterns of major ions and enriched
774 metals correspond to particular tectonic settings around the world (Deocampo and Jones, 2014). For
775 example, calc-alkaline volcanic activity has dominated East Africa since the Miocene and has
776 produced widespread sodium carbonate water (Jones et al., 1977; Cerling, 1996). In contrast, the
777 northern plains of North America are underlain by Paleozoic to Mesozoic evaporitic sedimentary
778 rocks and Quaternary till, which have helped to generate strongly sulfatic waters (Last, 1999). Trace-
779 metal compositions of brines, such as the brines of Inner Mongolia, are enriched in U–Mo, whereas
780 brines in the Tien Shan mountains and northern Qinghai-Tibet Plateau are enriched in Cu–Cr and
781 Co–Mn, respectively (Zhang et al., 2022). These differences suggest that heterogeneity in the
782 evolution and nature of continental crust may be an additional large-scale influence on basin-related
783 mineralization (O’Nions and Oxburgh, 1988; Rudnick and Gao, 2014; Tang et al., 2016).

784

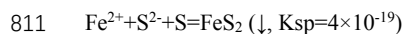
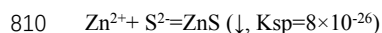
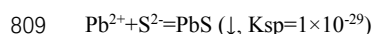
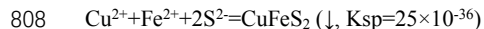
785 Another important role of tectonism is to provide the geological and structural setting for metal-
786 bearing brines, including the generation of fractures or fold-related structures that can act as fluid
787 conduits and points/zones of flow convergence (e.g., Sibson, 1985, 1987; Cox et al., 1991). This
788 would favor the convergence of large volumes of ore-forming fluids in a relatively small space
789 (several kilometers) over a relatively short time (<1Ma; Cox and Knackstedt, 1999; Skinner, 1997),
790 thus creating the opportunity for mineralization to occur (Cox et al., 2001).

791

792 From a metallogenic viewpoint, the emergence of basin-related ore deposits in the rock record
793 between 2.02 Ga and 1.85 Ga (Hitzman et al., 2010; Leach et al., 2010) shows that these deposits
794 are unevenly distributed in two periods: 2.1-1.4 Ga (Period I) and 0.8 Ga to present (Period II), with
795 few scattered between these two periods and barely before the Great Oxygenation Event (GOE; Fig.



796 3 and 6). The Period I occurred following the GOE, which is very significant due to the major
797 changes of the atmosphere, especially the rise of oxygen contents in the atmosphere at around 2.2-
798 2.4 Ga (Catling, 2014). Before the GOE, the surficial water bodies were mainly in a reducing state,
799 and sulfur existed in the form of S^{2-} (e.g., Roscoe, 1969; Prasad and Roscoe, 1996; Williford et al.,
800 2011). The existence of S^{2-} limited the contents of sulfophilic elements (e.g., Cu, Zn, Pb, Fe) in the
801 surficial water bodies. The Fe, S, Zn, Pb and Cu contents in the crust are ~58000 ppm, 200-500 ppm,
802 60-95 ppm, 10-16 ppm, and ~55 ppm, respectively (Liu et al., 1984). The solubility product of
803 galena is 1×10^{-29} , of chalcopyrite is 25×10^{-36} , of sphalerite is 8×10^{-26} , and of pyrite is 4×10^{-19} (the
804 chemical reaction equations are as follows; Liu et al., 1984). This means that S^{2-} would consume
805 the Pb^{2+} and Zn^{2+} first. In addition, the abundance of sulfur is much higher than that of Pb and Zn.
806 In this situation, it is difficult to accumulate Zn, Pb, and Cu (sulphophile elements) but can
807 accumulate Fe in surficial water bodies.



812 The solubility of uranium in surficial water is primarily controlled by its oxidation state. Generally,
813 U is more soluble in oxidizing, alkaline, and carbonate-rich water than in acidic, reducing water
814 (Cuney, 2010). This is due primarily to the tendency of U^{6+} to form strong complexes in oxidizing
815 fluids, regardless of the temperature. Before the GOE, Earth's surficial waters were dominantly
816 reducing and acidic, both of which restricted the uranium content in surficial brines at that time (e.g.,
817 Fig.6A; Cuney and Kyser, 2008; Kyser, 2014). Therefore, there were barely any basin-related Pb-
818 Zn, Cu, and U deposits before the GOE.

819

820 After the GOE, the relatively high oxygen content resulted in the oxidation of surficial waters, and
821 sulfur existed in the form of SO_4^{2-} , which was different from the time before the GOE (e.g., Roscoe,
822 1969; Prasad and Roscoe, 1996; Williford et al., 2011). In this environment, some metals in the
823 surficial saline brines were significantly enriched to an extreme high level, such as Cu (~800 ppm),
824 Zn (lack of S^{2-} , locally up to 250 ppm), Pb (lack of S^{2-} , locally up to ~100 ppm), and so on (Hagni,



825 1983; Kharaka et al., 1987; Hacini and Oelkers, 2010; Kharaka and Hanor, 2014; Zhang et al., 2022).
826 This means that metal-fertile saline brines were present on the Earth's surface since this time point,
827 which was the prerequisite for the formation of basin-related Pb-Zn, and Cu deposits. In addition,
828 after the GOE, the surficial brines had the ability to render greater solubilities of uranium by several
829 orders of magnitude (over tens of ppm), particularly when phosphate and carbonate were present
830 (Cuney, 2010; Kyser, 2014). Therefore, basin-related U deposits began to emerge.

831
832 The period of 1.9-1.8 Ga was the transformational period from plume-influence tectonics to modern-
833 style plate tectonics (e.g., de Wit, 1998; Goldfarb et al., 2001a, b; Condie, 2002a, 2002b, 2014), and
834 Cordilleran-style tectonics began to dominate Earth by 1.7 Ga (Goldfarb et al., 2010). The start-up
835 of modern-style plate tectonics and the widespread development of evaporite sediments meant the
836 extensive development of hydrological closed basins covered by arid-semi arid climates similar to
837 the Phanerozoic (e.g., Condie, 2005; Warren, 2006, 2010). This is the main reason that the period of
838 1.8-1.5 Ga was the peak period of basin-related mineralization and characterized by a global basin-
839 related U and Pb-Zn metallogeny (Fig. 3; e.g., Rozendaal, 1980; Rozendaal and Stumpfl, 1984;
840 Bodon, 1998; Chapman, 2004; Stalder and Rozendaal, 2004; Large et al., 2005; Dahlkamp, 2009,
841 2010, 2016). For the famous U and Pb-Zn districts, contemporaneous saline deposits (direct
842 evidence of the existence of metal-fertile saline brines) were developed in their region, such as the
843 Lower Proterozoic saline giants in Australia, Proterozoic redbeds and saline giants in Canada, Lower
844 Proterozoic saline giants in Papaghni Group of India, Lower Proterozoic redbeds in North China
845 Cratons, and Lower Proterozoic evaporites in Europe (detailed comparison in Appendix A and B,
846 e.g., Pirajno and Grey, 2002; Pope and Grotzinger, 2003; Bekker et al., 2006; Evans, 2006; Aspler
847 and Chiarenzelli, 2002; Walker et al., 1977). Basin-related mineralization could last for tens of
848 millions to hundreds of millions of years within the evolving basins (Brown, 2009; Hitzman et al.,
849 2010), because as long as it was covered by semi-arid climatic environments, there would be (metal-
850 rich oxidizing) saline brines existing in the hydrological closed basins, as well as basin-related
851 mineralization.

852

853 In contrast to Period I, the distribution of evaporites and evolution of paleoclimate of the Earth is



854 relatively well understood (e.g., Prokoph et al., 2013; Godderis et al., 2014; Wang et al., 2020;
855 Scotese et al., 2021), and there is a greater degree of confidence that the semi-arid sediments
856 accompanied the basin-related ore deposits during Period II, which is mainly the Phanerozoic Eon
857 (a detailed comparison is shown in Fig.4 and 5). There are many famous basin-related mineral
858 regions during Period II, such as the Central Massif of France (U-W, 290-260Ma), Bohemian Massif
859 of the Czech Republic (U-W, 290-260Ma), Pyrenees Peninsula (U-W, 290-260Ma), stratiform red-
860 bed hosted Cu–Ag ores of the Permian Kupferschiefer in Poland and Germany, Devonian to Permian
861 and Cretaceous to Paleogene SEDEX and MVT Pb-Zn deposits in the USA, and South China
862 Polymetallic Province (Leach et al., 2005; Zhang et al., 2022). During the mineralization period of
863 these basin-related deposits, they were generally located in semi-arid to arid zones of low to middle
864 latitudes, and accompanied by the development of evaporites (detailed locations are in Zhang et al.,
865 2022).

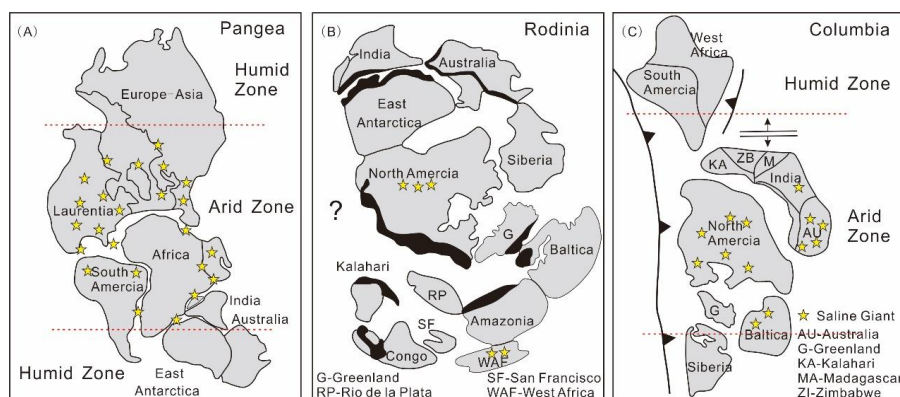
866

867 **6.4 One Warm, Humid and Flat Era during 1.4-0.8 Ga**

868

869 The statistical results of basin-related ore deposits show one quiescence period during 1.4-0.8 Ga
870 (Fig. 3). This mineralization gap has been addressed by Meyer (1981, 1988), and is now referred to
871 as the "Boring Billion". During this quiescence period, the saline giants were much less than in
872 Metallogenic Periods I and II, which means the surficial metal-fertile saline brines were also much
873 less than in both periods. There are several possibilities for the cause of this Metallogenic
874 Quiescence Period: I. there were few hydrologically closed basins; II. the climate zoning was weak,
875 and the arid-semi arid climatic zones or "horse latitudes" during this period were relatively small;
876 III. both I and II.

877



878

879 Fig.8 Three supercontinents and the major saline giants in Earth's history: (A) Pangea formed 300-
 880 250Ma ago (Rogers et al., 1995); (B) Rodinia formed ~1.0 Ga ago (Dalziel, 1997; Dalziel et al.,
 881 2000); and (C) Columbia formed ~1.8Ga ago (Rogers and Santosh, 2002).

882

883 For Possibility I, Tang et al. (2021a) proposed that the Earth was in an Orogenic Quiescence Period
 884 in its middle age, using a recently calibrated zircon-based crustal thickness proxy (Tang et al.,
 885 2021b). The zircon-based crustal thickness proxy uses pressure-sensitive Eu systematics during
 886 magmatic differentiation, which is recorded as Eu anomalies in crystallizing zircons (see Tang et al.,
 887 2021a, b for details). The reconstructed thickness of active continental crust averaged 50 to 60 km
 888 in Metallogenic Periods I and II (Fig. 3J). In contrast, the average thickness of continental crust was
 889 smaller than 45 km and even close to 40 km during the metallogenic quiescence period. This thin
 890 continental crust suggests that the continents were likely flat and lacked hydrological closed
 891 continental basins during the metallogenic quiescence period.

892

893 For Possibility No. II, much less saline deposits were reported during the metallogenic quiescence
 894 period than Metallogenic Periods I and II (Fig.3 and 8; e.g., Kozary et al., 1968; Boucot et al., 2013).
 895 The metallogenic quiescence period witnessed the breakup of the Nuna supercontinent (1.6-1.2 Ga)
 896 and the subsequent amalgamation of the Rodinia supercontinent during 1.2 to 0.9 Ga (e.g., Zhao et
 897 al., 2002; Roberts, 2013; Cawood et al., 2016). Published evidence suggests that the breakup of
 898 Nuna was limited and transitioned to Rodinia with only minor reconfiguration (Bradley, 2008;
 899 Cawood and Hawkesworth, 2014; Cawood, 2020). The Rodinia supercontinent consisted of most of



900 the continents and was characterized by multiple Grenvillian continent–continent collisions
901 (Bradley, 2008). These collision belts are generally located inboard of the supercontinental
902 boundaries (Hoffman, 1991; Condie and Rosen, 1994; Pelechaty, 1996; Frost et al., 1998; Rainbird
903 et al., 1998). Geodynamically, the mainlands were at low to mid latitudes during this period (Robb
904 and Hawkesworth, 2020). In this context, the Earth was probably in a warm and humid greenhouse
905 stage during this gap, which ended with glaciations of ~850-630 Ma, despite a 7-14% reduction in
906 solar output compared to modern during this interval (Fiorella and Sheldon, 2017). Elevated
907 greenhouse gas concentrations have been invoked to explain the warmth of this period, such as high
908 contents of CO₂ (10-300 × preindustrial CO₂ level, ~280ppmv., Kah and Riding, 2007; Sheldon,
909 2006; Kanzaki and Murakami, 2015) and/or CH₄ (~28-140ppmv., Fiorella and Sheldon, 2017). CIA
910 and Eu/Eu* values of the Mesoproterozoic argillites from India and North America signify the
911 intensity of secular weathering similar to CIA values for the Orinoco, Nile, and Amazon (Bose et
912 al., 2008). These sedimentary researches also suggest that the Earth was generally in a hot, wet
913 climate and the mainlands were experiencing aggressive chemical weathering during the basin-
914 related metallogenic gap, since Proterozoic rivers had less sediment residence times due to a lack of
915 vegetation cover and hydrological closed basins (Bose et al., 2008). This warm and humid climate
916 was also supported by the fact that Grenvillian-aged zircon grains are everywhere, which means the
917 sediments during this period were laid down by a colossal river system that emanated from the same
918 Grenvillian mountain chain (Rainbird and Young, 2009). Such river systems may have covered a
919 large part of the supercontinent Rodinia (Rainbird and Young, 2009). In this context, there were
920 barely any hydrological closed basins.

921

922 The climatic conditions can also be roughly estimated by the life on Earth. During the period of the
923 Precambrian, stromatolites were the most prominent and widespread life forms, as suggested by the
924 widely preserved Precambrian benthic microbial carbonates (Walter, 1994). Several ecological
925 factors combine to permit stromatolite formation and preservation, regardless of geologic age (Pratt,
926 1982). These include sufficient light and oxygen, and favorable water temperature and chemistry
927 (Peters et al., 2017). The constraints of oxygen concentrations during the Proterozoic were poor, but
928 were estimated to be within 1-10% of today's value (e.g., Buick et al., 1995; Canfield, 2014). The



929 Sun was about 85% of its present value at 2.2 Ga, and was gradually getting brighter to 95% at the
930 Late Precambrian. Therefore, the light and oxygen can be regarded as stable factors. The water
931 temperature and chemistry are the remaining major factors, which could influence the stromatolites'
932 abundance and diversity.

933

934 Two categories of abundance data have been compiled for the stromatolites since GOE (Walter,
935 1994). Morphotype diversity data counts the number of stromatolite form taxa through time. Forms
936 are based on macroscopic features such as external shape and internal lamina arrangement (Awramik,
937 1971; Walter and Heys, 1985; Awramik and Sprinkle, 1999; Semikhatov and Raaben, 2000).
938 Stromatolites increased in abundance from the Paleoproterozoic to ~1450Ma ago, at the beginning
939 of the basin-related metallogenic gap, and then rapidly increased to a peak between ~1350Ma and
940 1100Ma ago. An irregular decline from the peak occurred between ~900 and ~700Ma ago,
941 corresponding to the global glaciation of the Cryogenian and also the end of the basin-related
942 metallogenic quiescence period. In addition, the stromatolites diversity shows similar variation
943 characteristics to the stromatolites abundance (Fig.3; Walter and Heys, 1985; Sheldon, 2013), with
944 the stromatolite diversity increasing rapidly at about ~1.4 Ga and reaching a peak at about ~1350Ma.
945 After the peak, the stromatolite diversity declined gradually to 550 Ma ago. The stromatolite
946 diversity of the basin-related quiescence period was generally at least two times that of metallogenic
947 period I (Fig.3). After the Precambrian biological explosion, the stromatolites were restricted to
948 locations where the activity of metazoans was limited, so the stromatolites abundance and diversity
949 of the Phanerozoic were not used to reflect the climate conditions (Awramik and Sprinkle, 1999).
950 The peak of stromatolites diversity and abundance during 1.4-0.8 Ga suggests that there should have
951 been widely developed shallow marine environments on the continents at middle and low latitudes.
952 The environments were unfavorable for the formation of saline brines on the continents and their
953 margins (Fig.8). Therefore, this Metallogenic Quiescence Period was probably caused by the
954 combination of the lack of hydrological closed basins and a relatively warm and humid climate
955 during 1.4-0.8 Ga.

956

957 **7. Conclusion**



958

959 For the basin-related ore deposits, the ore-forming fluids were probably metal-rich saline brines
960 formed in hydrologically closed basins. These saline brines usually exist in the form of saline lakes
961 and mainly develop in arid to semi-arid environments. The main influence on the formation of these
962 brines is evaporation, which is the primary or only way for water to leave the convergent area.
963 Therefore, climate is an important factor influencing global metallogeny.

964

965 The statistical results of basin-related mineral deposits worldwide show that there are two
966 metallogenic periods after the great oxidation event: 2.1-1.4Ga (Period I) and 0.8Ga to present
967 (Period II), with a few scattered between these two periods (the metallogenic gap period). In addition,
968 Metallogenic Period II has five metallogenic peaks: ~380-340Ma (II-1), ~300-240Ma (II-2), ~160-
969 100Ma (II-3), 60-40Ma (II-4), and one stratiform Cu metallogenic concentration period of ~580-
970 500Ma (II-5). These two metallogenic periods and five metallogenic peaks are coupled with the
971 widespread development of saline deposits in time and space. The metallogenic quiescence period
972 corresponded to the scarcity of saline deposits, and was probably caused by the combination of the
973 lack of hydrologically closed basins and arid to semi-arid environments during 1.4-0.8 Ga.

974

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978

979 Author contributions

980

981 Chuang Zhang wrote the paper and drafted the figures.

982

983 Competing interests

984

985 The author has declared that there isn't any competing interests.



986

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991

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2154 Appendix A Major hydrothermal type uranium districts of the world, with showing their major
 2155 enriched metal affinities and regional contemporaneous saline deposits and/or redbeds (detailed
 2156 reference in Dahlkamp, 2009, 2010, 2016).

No.	Continent	Country	Major U District (Belt)	Ore deposit type	Representative U Deposit	Resource	Mineralization Age	Contemporaneous Saline Giant (or Redbeds) in the Region?
1	Europe	Austria	Land Salzburg	Strata-bound and structure-controlled	Mitterberg	>700 t U, ~0.1% U	290–260Ma, 90±5Ma	Permian and Jurassic evaporites in Europe
2	Europe	Bulgaria	Bukhovo district and other deposits	Structure-controlled, Vein type	Bortshe, Goten, Kamiko	>25000 t U, 0.02–3% U	4Ma, 1–2Ma?	Messinian saline giant
3	Europe	Czech Republic	Northwest Bohemian U region	Structure-controlled, Vein type	Horni Rozmysl, Prisenice	>10000 t U, 0.1 to 1% U	280–260Ma	Permian evaporites in Europe
4	Europe	Czech Republic	Horni Slavkov (Schlaggenwald) district	Structure-controlled, Lignite, Coal type	Barbora, Nadlesi, Zdar Buh	1000 t U	280–260Ma	Permian evaporites in Europe
5	Europe	Czech Republic	West Bohemian district	Structure-controlled, Vein type	Tachova, Vitkov I and II	>12000 t U, 0.05–0.2% U	185±15Ma	Saline aquifers of Central Europe
6	Europe	Czech Republic	Central Bohemian U region	Structure-controlled, Vein type	Pribram district	>80000 t U, 0.15% U	280–240 (265±15) Ma	Permian evaporites in Europe



7	Europe	Czech Republic	Southwest Bohemian U region	Structure-controlled, Vein type	Dametice, Ustalec, Zelenov	> 3000 t U, 0.1-0.2% U	280-240 (265±15) Ma	Permian evaporites in Europe
8	Europe	Czech Republic	West Moravian region	Structure-controlled, Vein type	Dancovice, Slavkovice	>25000 t U, 0.08-0.15% U	270±15 Ma; 190±10Ma	Permian and Jurassic evaporites in Europe
9	Europe	Czech Republic	Other Major U deposits	Structure-controlled, Vein type	Rychlebskehor y, Orlickehory	>5000 t U, 0.1-0.2% U	280-250 (265) Ma	Permian evaporites in Europe
10	Europe	Finland	Proterozoic U deposits in Finland	Structure-controlled, Vein type, Polymetallic	Juomasuo, Hangaslampi	>1500 t U, >20 g/t Au, 0.03% Co, 0.18% Cu, 0.13% Mo, 0.20% U	1900-1800Ma	Early Proterozoic evaporites in Baltica
11	Europe	France	La Crouzille district, Limousin	Structure-controlled, Vein type	Margnac-Pen-y-ore field	25000 t U, 0.1-0.3% U	280-260 (270)Ma; 190-150Ma; 30Ma	Permian and Jurassic evaporites in Europe
12	Europe	France	La Marche	Structure-controlled, Vein type	Bernardan, Mas Grimaud, Piegut	10000 t U, 0.03-1.5% U	280-260 (270)Ma; 190-150Ma; 30Ma	Permian and Jurassic evaporites in Europe
13	Europe	France	Plateau de Millevaches	Structure-controlled, Vein type	Northern Millevaches	2000 t U, 0.15-0.3% U	270Ma; Oligocene-Early Miocene	Permian and Miocene evaporites in Europe
14	Europe	France	Forez District	Structure-controlled, Vein type	Bois Noirs-Limouzat	7500 t U, 0.05-0.5% U	270Ma; Oligocene-Early Miocene	Permian and Miocene evaporites in Europe
15	Europe	France	Morvan District	Structure-controlled, Vein type	Le Cartelet, Mazille, Niallin	1500 t U, 0.1-0.4% U	175±5Ma	Jurassic evaporites
16	Europe	France	Margeride district	Structure-controlled, Vein type	Le Vigan, Les Bondons	4024 t U, 0.05-0.3% U	188±12Ma	Jurassic evaporites
17	Europe	France	Rouergue-Levezou Massifs and adjacent regions	Structure-controlled, Vein type	Entraygues, Saint Leger de Peyre	759.2 t U; 0.15% U	180-170Ma	Jurassic evaporites
18	Europe	France	Mortagne Massif	Structure-controlled, Vein type	Le Chardon ore field	15600 t U, 0.03-0.7% U	420-390Ma; 345-310Ma; 280-260Ma (290-270Ma); 180Ma; 70Ma	Middle Devonian, Carboniferous, Permian, Jurassic, Late Cretaceous Saline deposits
19	Europe	Germany	Schneeberg-Schlema-Alberoda ore field	Structure-controlled, Vein type	Schneeberg, Bernsbach	96603 t U;	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits



20	Europe	Germany	Schwarzenberg ore field	Structure-controlled, Vein type	Bermsgrun, Weiber Hirsch	1051.9 t U	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits
21	Europe	Germany	Johannegeorgenstadt ore field	Structure-controlled, Vein type	Rabenberg, Seifenbach	4537 t U	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits
22	Europe	Germany	Pohla-Tellerhauser ore field	Structure-controlled, Vein type	Ehrenzipfel, Tellerhauser	7918.6 t U	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits
23	Europe	Germany	Central Erzgebirge	Structure-controlled, Vein type	Barenstein-Niederschlag	800 t U	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits
24	Europe	Germany	East Erzgebirge	Structure-controlled, Vein type	Barenhecke, Freiberg	32000 t U	278-270Ma, 250-240Ma, 185-180Ma, 160-140Ma	Middle Devonian, Carboniferous, Permian, Jurassic Saline deposits
25	Europe	Germany	Königstein district, a northern extension of the North Bohemian Basin	Sandstone type	?	27812 t U	74Ma, 49.5Ma, 24Ma	Late Cretaceous – Early Paleogene, Miocene evaporites
26	Europe	Germany	Dohlen Basin, Freital district	Lignite type U deposits	Heidenschanze, Gittersee	3977 t U	240Ma	Triassic evaporites
27	Europe	Germany	Kyhna-Schenkenberg	Structure-controlled, Vein type	Schenkenberg, Qüering	6660 t U	250-280Ma, 76-40Ma	Permian, Late Cretaceous to Early Paleogene evaporites
28	Europe	Germany	Gera-Ronneburg Region, East Thuringia	Black Shale	Ronneburg, Stolzenberg	200 000 t U	280-260Ma, 240Ma, 120-90Ma	Permian, Triassic, Late Cretaceous to Early Paleogene evaporites
29	Europe	Germany	East Thuringian Anticline	Black Shale	?	15 300 t U	280-260Ma, 240Ma, 120-90Ma	Permian, Triassic, Late Cretaceous to Early Paleogene evaporites
30	Europe	Germany	Grobschloppen-Hebanz/Weibenstadt Granitic Massif	Structure-controlled, Vein type	Grobschloppen, Rudolfstein	>2000 t U	~268Ma, ~233Ma	Permian and Triassic evaporites
31	Europe	Germany	Poppenreuth-Mahring area	Structure-controlled, Vein type	Hohensteinweg, Waldel	?	~336±17Ma, 295±5Ma	Carboniferous and Permian evaporites
32	Europe	Germany	Schwarzach Valley/Neunburg Massif	Structure-controlled, Vein type	Altfalter	?	~360Ma?, ~210Ma?	Carboniferous and Triassic evaporites



33	Europe	Germany	Other deposits in Northeast Bavaria	Structure-controlled, Vein type	Schonthan, Wolsendorf-Nabburg	?	~295±14Ma?	Permian evaporite
34	Europe	Germany	Menzenschwand district	Structure-controlled, Vein type	Krunkelbach Mine	?	310±3Ma, 65Ma, 295±7Ma, 50±8Ma	Triassic, Permian, Late Cretaceous to Early Paleogene evaporites
35	Europe	Germany	Upper Kinzigtal district	Structure-controlled, Vein type	Stammelbach Valley, Wittichen	?	230-245Ma; 240Ma	Triassic evaporites
36	Europe	Germany	Other deposits in Schwarzwald region	Structure-controlled, Vein type	Sulzburg, St. Ulrich, Fischerhof	?	~240Ma, 160±20Ma	Triassic and Jurassic evaporites
37	Europe	Poland	Kowary Ore Field, Kletno Ore Field	Structure-controlled, Vein type	?	760 t U	265Ma, 70Ma	Triassic, Late Cretaceous to Early Paleogene evaporites
38	Europe	Romania	Major deposits in Romania	Structure-controlled, Vein type	Arieseni, Avram Iancu	16500-17000 t U	280Ma, 230Ma, 110Ma	Permian, Triassic, and Cretaceous evaporites
39	Europe	Russian Federation	Lake Ladoga district	Unconformity-related, Hydrothermal metasomatic type	Salmi, Shotkusa	35 000 t U	1.8-1.75Ga; 1.5Ga; 412±11Ma	Paleo-Proterozoic, Mesoproterozoic, and Late Silurian to Early Devonian evaporite
40	Europe	Russian Federation	Lake Onega district	Metasomatism type	Srednaya Padma, Tsarevskoye,	9 000 t U	1740±30Ma	Paleo-Proterozoic evaporite
41	Europe	Russian Federation	Northern Caucasus region	Structure-controlled, Volcanic type	Naratinskoye, Dakhovskoye	?	190Ma, 140Ma, 115Ma, 40Ma?	Jurassic, Cretaceous, and Early Paleogene evaporites
42	Europe	Slovak Republic	Gemer zone	Structure-controlled, Volcanic type	Cierna hora, Haniskova	25000 t U; >3000 t Mo	240±30Ma, 130±20Ma	Triassic and Cretaceous evaporites
43	Europe	Slovak Republic	Kozie Mountains and Northern Lower Tatra	Stratiform type	Vikartov-Cierny Vah zone	>3000 t U, >0.10% U	160Ma, 105-70Ma	Jurassic and Cretaceous evaporite
44	Europe	Slovak Republic	Major deposits in Slovak Republic	Structure-controlled, Volcanic type	Kalnica, Krajna dolina, Selec	>2000 t U, 0.1% U	160Ma	Jurassic evaporite
45	Europe	Portugal	Beiras (Guarda-Visou) U region	Structure-controlled, Vein type		>10 000 t U, 0.1-0.25% U	100-60Ma, 57-37Ma	Late Cretaceous to Early Paleogene evaporite
46	Europe	Spain	South Spain region	Structure-controlled, Vein type	Aguila, Alameda, Villar, Retortillo ore field	36700 t U, 0.05% U	57-37Ma	Early Paleogene evaporites



47	Europe	Sweden	Petrozoic U deposits in Sweden	Structure-controlled, Vein type, Metasomatism	Arvidsjaur, Arjeplog, and Sorsele region	10000 t U, 0.03-0.1% U	1738±10Ma, 1736-1767Ma	Paleo-Proterozoic evaporites
48	Europe	Sweden	Other U deposits in Sweden	Structure-controlled, Vein type	Flistjam, Klappibacken	>2000 t U	420±1Ma	Late Silurian evaporites
49	Europe	Switzerland	Les Marecottes-La Creusa area	Structure-controlled, Vein type	Gisiger, Juillard	> 1000 t U	270Ma, 240Ma	Permian and Triassic evaporites
50	Europe	Ukraine	Kirovograd-Smolino region	Structure-controlled, Metasomatism	Michurinsk ore field	250 000 t U, 0.08-0.14% U	1812±42 to 1753±42Ma; 1.84-1.80Ga; 1.8-1.7Ga	Paleo-Proterozoic evaporites
51	Europe	Ukraine	Krivoy Rog region	Structure-controlled, Metasomatism	Annovskoye, Krasnogvardei skoye	>25000 t U, 0.1% U	1770±50Ma	Paleo-Proterozoic evaporites
52	Europe	Ukraine	Pobuzhsky region	Structure-controlled, Metasomatism	Kalinovskoye, Lozovatskoye	>15000 t U, 0.5-0.22% U	1770±50Ma	Paleo-Proterozoic evaporites
53	Europe	Ukraine	Ingul Megablock	Structure-controlled, Vein type	Geikovskoye, Lagodovskoye	>3000 t U, >0.10% U	1770±50Ma	Paleo-Proterozoic evaporites
54	Europe	Ukraine	Dnieper Block	Structure-controlled, Vein type	Sergeevskoye	>3000 t U, >0.10% U	1770±50Ma	Paleo-Proterozoic evaporites
55	Europe	Ukraine	Priazov Block	Structure-controlled, Vein type	Barbasovskoye, Dibrovskoye	>3000 t U, >0.10% U	1770±50Ma	Paleo-Proterozoic evaporites
56	Europe	United Kingdom	Southwest U region in U.K.	Structure-controlled, Vein type	Restormel, Royal, South Terras	>2000 t U	280-270Ma, 244-212Ma, 175-160Ma, 75-55Ma	Permian, Triassic, Jurassic and Late Cretaceous to Early Paleogene evaporites
57	Asia	China	Wuyishan Belt	Structure-controlled, Vein type	Maoyangtuo	>50000 t U	127-137Ma, ~81Ma, 71-75Ma	Cretaceous evaporites
58	Asia	China	Taoshan-Zhuguang Belt	Structure-controlled, Vein type	Juntian, Dabu	>50000 t U	~147-152Ma, ~101-110Ma, ~82-89Ma; ~65-67Ma, ~52-57Ma, ~42-45Ma	Late Jurassic to Early Paleogene evaporites
59	Asia	China	Chenzhou-Qinzhou U Belt	Structure-controlled, Vein type	Jiuyishan	>50000 t U	~70Ma, ~55-64Ma	Late Cretaceous to Early Paleogene evaporites
60	Asia	China	Gan-Hang Belt	Structure-controlled, Vein type	Xiangshan district	>50000 t U	~141-148Ma, ~120-130Ma, ~84-94Ma	Late Jurassic to Early Paleogene evaporites
61	Asia	China	Xixia-Luzong Belt	Structure-controlled, Vein type	Xixia	>20000 t U	110-120MaMa, 60Ma	Late Cretaceous to Early Paleogene evaporites
62	Asia	China	Mufushan-Hengshan Belt	Structure-controlled, Vein type	Jinguanhong, Baofengyuan	>20000 t U	~5-14Ma, ~24-26Ma	Miocene evaporite



63	Asia	China	Xuefengshan- Jiuwandashan Belt	Structure-controlled, Vein type	Silihe	>20000 t U	~316-360Ma, ~60-71Ma, ~73- 88Ma, ~44Ma	Carboniferous, Cretaceous to Paleogene evaporites	Late to Early
64	Asia	China	Jungar-Tien Shan	Structure-controlled, Vein type	Baiyanghe	>1000 t U	~230Ma	Triassic evaporites and redbeds	
65	Asia	China	Liaoning Region/Yinshan- Liaohe Belt	Structure-controlled, Vein type	Lianshanguan	>3000 t U	1894Ma; 1829Ma; 1810Ma; 1823Ma	Paleo-Proterozoic evaporites and redbeds	
66	Asia	China	Yanshan Belt	Structure-controlled, Vein type	Dashiqiao	4 000 t U	1763-1794Ma; 226Ma	Paleo-Proterozoic, Triassic evaporites and redbeds	
67	Asia	China	Taihangshan Belt	Structure-controlled, Vein type	Wutaishan	1500 t U	1765-1746Ma; 103Ma; 122Ma	Paleo-Proterozoic, Cretaceous evaporites and redbeds	
68	Asia	China	Qinglong Region/Yinshan- Liaohe Belt	Structure-controlled, Vein type	Qinglong	8000 t U; 0.03 to 0.1% U	124Ma	Cretaceous evaporites	
69	Asia	China	Guyuan Region/Yinshan- Liaohe Belt	Structure-controlled, Vein type	Zhangmajing	> 3000 t U	~90Ma; ~24Ma	Late Cretaceous evaporites	
70	Asia	China	Hongshanzi Region/Yinshan- Liaohe Belt	Structure-controlled, Vein type	Hongshanzi	~3000 t U	~90Ma	Late Cretaceous evaporites	
71	Asia	China	Longshoushan region/Qilian- Qinling	Structure-controlled, Vein type	Hongshiquan	1500-3000 t U, 0.03- 0.1% U	~1760Ma; 700- 600Ma	Paleo-Proterozoic redbeds,	
72	Asia	China	Longshoushan region/Qilian- Qinling	Structure-controlled, Vein type	Jiling	1500-3000 t U, 0.03- 0.1% U	430-400Ma; 383- 357Ma	Devonian evaporites	
73	Asia	China	North Qinling region/Qilian- Qinling	Structure-controlled, Vein type	Lantian	~8000 t U, 0.171 % U	96Ma	Late Cretaceous	
74	Asia	China	North Qinling region/Qilian- Qinling	Structure-controlled, Vein type		<3000 t U	280-260Ma	Permian evaporites	
75	Asia	China	Ruoergai region/Qilian- Qinling	Structure-controlled, Vein type	Luojungou	>3000 t U	130-45Ma	Late Cretaceous to Early Paleogene evaporites	
76	Asia	India	Singhbhum Belt	Cu-U Structure-controlled, Vein type	Khadandungri- Purandungri ore field	56000 t U	1766+82Ma; 1.5- 1.6Ga	Paleo-Proterozoic redbeds	



77	Asia	India	Cuddapah Basin	Unconformity-proximal, and fracture-controlled types	Nalgonda and Guntur Districts/Srisailam Subbasin	22 000 t U	1756 ± 29 Ma	Paleo-Proterozoic redbeds
78	Asia	India	Bhima Basin	Unconformity-proximal, and fracture-controlled types		1300 t U; 0.16% U	1756 ± 29 Ma;	Paleo-Proterozoic redbeds
79	Asia	Iran	Saghand Ore Field	Metasomatic and hydrothermal vein type		1400 t U	?	?
80	Asia	Kazakhstan	Kokshetau region	Vein-stockwork type	Ishimsky Ore Field, Forty uranium deposits	230000 t U; 0.1-0.2% U	~380-360Ma	Devonian evaporites
81	Asia	Kazakhstan	Kendyktas-Chuily-Betpak Dala region	Vein-stockwork type	Dzhideli Ore Field	65000 t U; 0.1-0.3% U	370-350Ma; 285-265Ma	Devonian and Permian evaporites
82	Asia	Kyrgyzstan	Eastern Karamazar-Northeastern Fergana Region	Vein-stockwork type	Charkasar, Mayлуу-Suu, Shakaptar, and Malisay	12000 t U; 0.1-0.3% U	370-350Ma; 285-265Ma	Devonian and Permian evaporites
83	Asia	Mongolia	Mardai, Ugtam, Turgen, Engershand	Vein type, volcanic type	Dornod, Gurvan bulag	57 000 t U; 0.16% U	138-136Ma	Cretaceous evaporites
84	Asia	Russian Federation, Asian Territory	Yenisey Region	Vein type, volcanic type	Labyshkoye, Solonechnoye	>2000 t U, 0.1% U	350-370Ma	Devonian evaporites
85	Asia	Russian Federation, Asian Territory	Streltsovsk district	Vein type, volcanic type	19 deposits, Streltsovskoye, Tulukuyevskoye, and Ochyabrskoye	28 0000 t U; 0.1-0.3% U	136-134Ma; 18-17Ma	Cretaceous, Miocene evaporites
86	Asia	Russian Federation, Asian Territory	Central Transbaykal region	Vein type, volcanic type	Olovskoye and Imskoye	40000 t U; 0.01-0.1% U	110-100Ma	Cretaceous evaporites
87	Asia	Russian Federation, Asian Territory	Elkon district	vein-stockwork-type	Agdinskaya, Severmoye, Sokhsolookhsk	342 000 t U; 0.1-0.15% U	135-130Ma	Cretaceous evaporites



88	Asia	Russian Federation, Asian Territory	Bureinsky District	vein-stockwork type	Lastochka, Kamenushinskoye, Skalnnoye, Svetloye	29000 t U; 0.03-0.2% U	136-134Ma	Cretaceous evaporites
89	Asia	Russian Federation, Asian Territory	Tas-Kastabyt	Volcanic type, vein type		>3000 t U; 0.1-0.3% U	119-89Ma; 127Ma; 72Ma; 60-50Ma	Cretaceous to Early Paleogene evaporites
90	Asia	Turkmenistan	Karamazar Uranium Region	vein stockwork type	Chauli, Alatanga	20 000 t U; 0.1-0.3% U	275-267Ma; 280-270±10Ma	Permian evaporites
91	America	USA	Arizona Strip Area	Collapse Breccia Pipes	Orphan Lode	15000 t U; 0.3-0.7% U	12000-260Ma; 200±20Ma; 141Ma	Permian and Jurassic evaporites
92	America	USA	Spokane Mountain Area	Vein type	Midnite and Sherwood	15000 t U;	51Ma	Late Cretaceous to Early Paleogene evaporites
93	America	USA	Ralston Buttes	Vein type	Schwartzwalde r	>7500 t U; 0.408% U	70-52Ma; 69.3 ± 1.1 Ma	Late Cretaceous to Early Paleogene evaporites
94	America	USA	Marshall Pass District	Vein type	Pitch deposit	>11 000 t U; >0.1% U	Laramide to Mid-Tertiary age	?
95	America	USA	McDermitt Caldera District, Nevada-Oregon;	Vein type	Aurora	>10000 t U; 0.04% U	13.3 ± 2 Ma; 12.3 ± 7 Ma	Miocene redbeds and evaporites
96	America	USA	Lakeview District, Oregon	Vein type	White King	>200 t U		?
97	America	USA	Spor Mountain/Thomas Caldera, Utah	Vein type	?	?	21Ma	Miocene redbeds and evaporites
98	America	USA	Marysvale Volcanic Complex, Utah	Vein type	?	540 t U	~20Ma; 19-18Ma	Miocene redbeds and evaporites
99	America	USA	Date Creek Basin, Arizona	Vein type	?	12000 t U; 0.06% U	~20Ma; 19-18Ma	Miocene redbeds and evaporites
100	America	USA	Sierra Ancha/Apache Proterozoic Basin, Arizona	Vein type	?	38 00 t U; 0.15-0.32% U	1.8-1.7Ga?	Paleo-Proterozoic redbeds and evaporites
101	America	USA	Coles Hill	Vein type, Disseminated impregnation-type	?	45770 t U; 0.05% U	562 ± 5 Ma; 417Ma	Neo-Proterozoic redbeds and evaporites; Late Silurian
102	America	Mexico	Sierra de Peña Blanca	Vein type	Nopal-1	4000 t U; 0.1% U	43.8-37.3Ma	Early Paleogene
103	America	Brazil	Caetité Massif	Metasomatite type	Ten deposits	85 000 t U; 0.15% U	1400Ma?	?
104	America	Brazil	Central Ceará Region, Itaitaia	Vein type	Santa Quitéria Deposit	12 0000 t U; 0.08% U	1.8-1.7Ga	Paleo-Proterozoic redbeds and evaporites



105	America	Brazil	Poços de Caldas Region	Vein type	Cercado and Agostinho	27000 t U; 0.07% U	80-60Ma;	Late Cretaceous to Early Paleogene
106	America	Brazil	Seridó Region	Metasomatite type	Espinharas Deposit	4240 t U; 0.04% U	450Ma?	?
107	America	Brazil	Parana Basin			11000 t U;	Triassic	?
108	America	Canada	Great Bear Batholith	Vein type, Metasomatite type		20000 t U	1860±20Ma, 1424±29Ma, 1076±96Ma, 457±26Ma, 511±86Ma, 415±29Ma, 536±66Ma, 339±22Ma, 125±48Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
109	America	Canada	Athabasca Basin	Unconformity type		150000 t U	1770-1730Ma (1740Ma), 1590Ma, 1500- 1520Ma, 1440Ma, 1350- 1330Ma, 1290Ma, 1250Ma, 1190Ma, 1150Ma, 1120Ma, 1075Ma, 1030Ma, 1000Ma, 920Ma, 825Ma, 750Ma, 671Ma, 643Ma, 530Ma, 250Ma, 116Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
110	America	Canada	Thelon	Unconformity type, vein type		5 0000 t U	1860Ma, 1667- 1640Ma, 1520- 1500Ma, 1284Ma, 1200Ma, 1131- 1100Ma, 982Ma, 577Ma, 532Ma, 513Ma, 489Ma, 284Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
111	America	Canada	Beaverlodge	Vein type		100000 t U	2058±34Ma, 1875-1855Ma, 1740Ma,	Episodic evaporites and redbeds of Paleo-



						1450Ma, 1200Ma, 422±38Ma, 340±4Ma	Proterozoic to Neo-Proterozoic
112	Australia	Australia	Kombolgie Basin	Unconformity type, vein type	3 0000 t U	1690-1680Ma, 1650-1600Ma, 1521±8Ma, 1445±20Ma, 1348±16Ma, 1040Ma, 474±6Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
113	Australia	Australia	North Olary Province	Vein type	3 0000 t U	1705Ma, 1580Ma, 580Ma, 513Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
114	Australia	Australia	Pink Creek Inlier	Unconformity type	5 0000 t U	1740Ma, 1723Ma, 1680Ma, 1650- 1600Ma, 1560Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
115	Australia	Australia	Mounte Isa	Na-Metasomatite type	3 0000 t U	1750-1730Ma, 1640Ma, 1534Ma, 1523- 1505Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
116	Australia	Australia	South Australia	Olympic Dam	30 0000 t U	1590Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic
117	Africa	Gabon, Africa	Franceville	Unconformity type	5 0000 t U	1950±40Ma, 890-860Ma, 500Ma	Episodic evaporites and redbeds of Paleoproterozoic to Neoproterozoic

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2158 Appendix B Inferred and direct evidence of evaporites from Proterozoic to Cenozoic, based on the

2159 compilation by Pope and Grotzinger (2003), Bekker et al. (2006), Schroder et al. (2008).

No.	Location	Age (Ga)	Units	Evaporite evidence	Thickness	Notes	Reference
1	Canada	Ca. 2.22-2.3	Gordon Lake Formation, Huronian Supergroup	Ba (as beds), silicified and pristine anhydrite and gypsum nodules and layers, Si-tr-An, beds of anhydrite	Multiple horizons in >300m	SG, marine, passive margin, supratidal and sabkha zone, correlated with	Cameron, 1983; Bekker et al., 2006



				nodules			Chocolay Group	
2	USA	Ca. 2.22-2.3	Kona Dolomite, Chocolay Group	Si-ps-Gy, ps-Ha (moulds), Si-ps-An, sc-breccias	30-1000m		SG, marine, associated with volcanics, intracratonic basin, open to passive margin, correlated with Gordon lake Formation	Bekker et al., 2006
3	Australia	Ca. 2.2	Bartle Member, Killara Formation, Yerrida Group	Si-ps-An, Kao-ps-Gy or (relics), Si-ps-Gy, An, An, ps-Sh, ps-Tro	?		Playa lake (Alkaline)	Pirajno and Gray, 2002
4	Africa	2.2	Pretoria Group, South Africa	Ps-Mir	< 2m		Sodic lake deposits in a playa setting	Pope and Grotzinger, 2003
5	Australia	Ca. 2.15 (2.2)	Bubble Well Member, Juderina Formation, Yerrida Group	Si-Evp, Q-ps-Gy, Q-ps-An	Ca. 100m		SG, Marine or marginal marine, associated with volcanics	El-Tabakh et al., 1999
6	Africa	Ca. 2.15 (2.10-2.20)	Lucknow Formation, Olifanshoek Group and Transvaal Supergroup, South Africa	Q-ps-Gy, molds after anhydrite and gypsum	Q-ps-An, ?		Marine, passive margin	Bekker et al., 2006; Schröder et al., 2008
7	Canada	Ca. 2.15	Laparré Formation, Peribonca Group, Otish Supergroup	Dol-ps-Gy, (after crystals and nodules) Dol-ps-An	?		Passive margin	Bekker et al., 2006
8	USA	Ca. 2.15	Lower part of the Nash Fork Formation, Snowy Pass Supergroup	Molds after anhydrite nodules and gypsum crystals	?		Passive margin	Bekker et al., 2006
9	Africa	Ca. 2.2-2.0	Francevillian C Formation, Francevillian	Ca-ps-An, Ca-ps-Gy	?		Marine, supratidal-sabkha	Preat et al., 2011



			Group, Gabon			environment	
10	Africa	Ca. 2.15	Norah Formation, Deweras Group, Zimbabwe	An, as layers	?	Intracratonic rift basin	Bekker et al., 2006
11	Russia	Ca. 2.1	Fedorovka (Fedorov) Formation (Aldan Shield)	An, as layers and veins	?	Passive margin	Zharkov, 2005; Bekker et al., 2006
12	Russia	Ca. 2.09	Tulomozero Formation, Upper Jatulian Group	Ca-ps-Gy, Dol-ps-Gy, Si-ps-Gy, An, pseudomorphs after anhydrite and gypsum crystals and nodules, ps-Ha, sc-breccias, enterolithic and chicken wire structures	Multiple units >20m, within ca. 500m of total thickness	SG, passive margin, playa lake, marine sabkha, intertidal flats	Melezhik et al., 2005; Brasier et al., 2011; Reuschel et al., 2012
13	Africa	2.06	Dewaras Group, South Africa	Gy		Lacustrine environment in rift	Pope and Grotzinger, 2003
14	Canada	Ca. 1.95	Rocknest Formation, Coronation Supergroup, Slave craton	Dol-ps-Gy, ps-An, ps-Ha	Traces of evaporites dispersed within carbonates	SG, marine, lagoon on inner shelf, passive margin	Evans, 2006
15	Canada	1.8-2.0	Kasegalik and Mc-Leary formations, Belcher Group	Ps-Gy, ps-Ha	Ca. 150m	Marine	Pope and Grotzinger, 2003
16	Canada	Ca. 1.87	Stark and Hearne formations, Great Slave Lake Supergroup	Ps-Ha, silicified hopper casts, and pagoda halite, ps-Gy, sc-megabreccia	200-600m, reconstructed thickness of evaporites are ca. 100m	SG, Marine to non-marine, halite>>gypsum	Pope and Grotzinger, 2003; Evans, 2006
17	Canada	1.81-1.91	Tavani Formation, Hurwitz Group	Q-ps-Gy, Dol-ps-Gy, halite moulds	?	Coastal pans, marine to non-marine	Aspler and Chiarenzelli, 2002
18	Russia	1.8-1.9		Ps-Gy, ps-An	?	Associated with barite, sabkha	Pope and Grotzinger, 2003
19	Canada	1.8	Brown Sound	Ps-Ha, ps-Gy, sc-	Ca. 300m	Marine to non-	Pope and



			Formation	breccias			marine halite>>gypsum	Grotzinger, 2003
20	Canada	1.8	Cowles Lake Formation	Ps-Ha, ps-Gy, sc- breccias	>200m		Marine to non- marine halite>>gypsum	Pope and Grotzinger, 2003
21	India	>1.7	Vempalle Formation, Papaghi Group	Ps-Ha, ps-Gy	?		Marine, associated with lava flows	Pope and Grotzinger, 2003
22	Australia	1.74 (1.54- 1.74)	Corella Formation, McArthur-Mt Isa basins	Ps-Sh, ps-Gy, quartz replacing anhydrite nodules	Ca. 500m		SG, alkaline lake suggested by shortite	Walker et al., 1977; Muir, 1987; Evans, 2006
23	Australia	1.66	Mallapunyah, Paradise Greek, Esperanza, Staveley formations, McArthur-Mt Isa basins	Ps-Gy, ps-Ha, botryoidal quartz nodules after anhydrite, massive replacement by gypsum	>10m		SG, marine sabkha	Walker et al., 1977; Evans, 2006
24	Australia	1.645	Myrtle, Emmerugga, and other formations, McArthur-Mt Isa basins	Ps-Gy, ps-Ha	Ca. 200m		SG	Walker et al., 1977; Evans, 2006
25	Australia	1.635	Lynott Formation, McArthur-Mt Isa basins	Ps-Gy, ps-Ha, cauliflower cherts	Ca. 300m		SG, marine sabkha	Walker et al., 1977; Evans, 2006
26	Australia	1.61	Balbirini Formation, McArthur-Mt Isa basins	Ps-Ha, pseudomorphs after sulfates, ps-Sh, Cauliflower cherts	?		SG, alkaline lake suggested by shortite	Walker et al., 1977; Evans, 2006
27	Australia	1.5	Discovery Formation, Edmund Group, Bangemall Supergroup, Bangemall basin	Ps-Gy, or ps-An	70		SG, marine, several evaporite horizons	Evans, 2006
28	USA, Canada	1.46	Waterton, Altyn, Prichard and Wallace formations, Belt Supergroup	Ps-Evp, ps-Gy, ps-An, ps-Ha, length-slow chalcedony, chicken- wire textures, scapolite	100 m		SG, marine, two evaporite horizons	Evans, 2006



29	USA	1.15-1.3	Upper Marble, Grenville Series	An, as lenses and beds	Beds (or lenses) > 40m thick	Metamorphosed evaporites	Whelan et al., 1990
30	Canada	1.2	Society Cliffs Formation, Victor Bay Group, Borden Basin	Gy, ps-Ha, sc-breccias	Multiple beds a few cm's to meter's thick, >100 m	SG, restricted marine	Kah et al., 2001; Evans, 2006
31	Africa	Ca. 1.2	Char Group, Mauritania, Douik Group, Algeria, Mauritania	ps-Ha	Ca. 50m	SG, marine, possibly correlate with evaporites in Atar and EI Mreiti Groups	Evans, 2006
32	Africa	Ca. 1.1	Qued Tarioufet Formation, Atar Group, Gouamir and Tenoumer formations, EI Mreiti Group, Taoudeni Basin, Mauritania	Ca-Evp, Ca-ps-Gy or Ar, ps-Ha, sc-breccias, chicken-wire texture	Ca. 50 m	SG, marine	Kah et al., 2012
33	Canada	0.7-1.2	Minto inlet and Kilian formations, Shaler Group	Gy	Multiple gypsum beds up to 30m thick	Marine	Pope and Grotzinger, 2003
34	Australia	0.80-0.83	Centralian superbasin	Gy, Ha	SG, ~140 000km ³	intracratonic basin	Babel and Schreiber, 2014
35	Africa	0.30-0.27	Dwyka Formation, Karoo Basin, South Africa	Ha, Gy	?	Intracratonic basin	Visser and Kingsley, 1982
36	Africa	0.16-0.13	Late Jurassic to Early Cretaceous, Morocco, Casamance River, Old Bahama Channel	Gy, Ha	SG	marine	Burke, 1975
37	Africa	0.23-0.20	Upper Triassic, basins in North Africa	An, Gy, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014



38	Asia	0.54-0.52	Salt Mountain Formation, Lower Cambrian, Salt Range, Pakistan	Ha, Gy, An, as layers; Dol-ps-Gy	Multiple horizons in > 100m	SG, marine	Vysotsky et al., 2014
39	Asia	0.02-0.01	Badenian, Middle Miocene, Carpathian Region	Ha, Gy	SG, Multiple horizons in > 10m	Intracratonic basin	Galamay et al., 2021
40	Asia	0.30-0.25	Kungurian Salt, Permian, Caspian Sea	Gy, Ha, An	SG	Intracratonic basin	Vysotsky et al., 2014
41	Australia	0.51-0.49	Saline River Formation, Northwest Territories	Ha, An; Dol-ps-An	Multiple horizons in > 200m	SG, marine to lagoonal, sub-tidal to peritidal environment, similar to Sabkhas	Vysotsky et al., 2014
42	Australia	0.54-0.52	Ouldburra Formation, Lower Cambrian, Officer basin, South Australia	Ha, An	Multiple horizons in > 20m	SG, marginal to restricted marine environments, mudflat	Kovalevych, 1988, 1990; Kovalevych et al., 2006
43	Australia	0.46-0.44	Carribuddy Group, Late Ordovician to Early Silurian, Canning basin, Western Australia	Ha, An, Dol-ps-An	Multiple horizons in > 100m	SG, marginal marine to ephemeral saltpan and saline mudflat	Williams, 1991
44	Australia	0.45-0.41	Dirk Hartog Group, Silurian, Southern Carnarvon Basin	An, Gy, Evaporite karst	?	Subtidal, peritidal, mudflat	El-Tabakh et al., 2004
45	Australia	0.63-0.54	Gillen Member, Undoolya Sequence, Late Proterozoic, Amadeus Basin	An, Gy, Ha	?	Intracratonic basin, marginal marine	Kennedy, 1993
46	Australia	0.9	Bitter Spring Formation	Gy, Ha	SG	Marine	Preiss, 1972; Marjoribanks



								and Black, 1974
47	Canada	0.42-0.41	Salina Formation, Upper Silurian, Great Lake	An, Gy, Dol-ps-An	?	intracratonic basin		Vysotsky et al., 2014
48	Canada	0.36-0.30	Carboniferous, Sverdrup Basin, Arctic Canada	Gy, An		SG, evaporite diapirs	Intracratonic basin	Balkwill, 1978
49	Canada	0.40-0.38	Fort Vermilion Formation, Western Canada basins	An, Gy, Ha, Syl, Car		SG, Multiple horizons in > 100m	Sabkham marginal marine	Klingspor, 1969; Wittrup and Kyser, 1990
50	Canada	0.40-0.38	Elk Point Group, Middle Devonian, Saskatchewan, Western Canada	Ha, Gy, An,		SG, Multiple horizons in > 10m	Intracratonic basin	Horita et al., 1996
51	Canada	0.39-0.36	Wabamun Group, Alberta Basin, West Canada	An, Dol-ps-An	?		Intracratonic basin	Anfort et al., 2001
52	Canada	0.35-0.33	Windsor Group, Early Middle Visean, Nova Scotia, New Brunswick, Atlantic Canada	Ha, Gy, An		SG, Multiple horizons in > 100m	Intracratonic basin	Schenk et al., 1994a, b; Ravenhurst et al., 1989
53	China	0.53-0.50	Wusonggeer Formation, Sayilike Formation, Awatage Formation, Lower-Middle Cambrian, Tarim basin, NW China	An, Dol-ps-An		Multiple horizons in > 100m	SG, lagoonal to tidal flat,	Wang et al., 2014
54	China	0.53-0.50	Meishucun Formation, Canglangpu Formation, Qingxu Formation,	Gy, Ha, An; Dol-ps-An		Multiple horizons in > 100m	SG, marine to lagoonal, to intracontinental, tidal flat	Huang, 1993; Goldberg et al., 2005



			Qiongzhusi Formation, Shilengshui Formation, Lower-Middle Cambrian, Yangtz Block					
55	China	0.46-0.44	Tanjianshan Group, Qaidam basin	Gy	?	Marginal marine, restricted basin	Wu et al., 2022	
56	China	0.36-0.30	Carboniferous, Tarim	Gy,	?	Marginal marine to restricted basin	Chen et al., 2016	
57	China	0.47-0.46	Majiagou Formation, Middle Ordovician carbonate-evaporite sequence, Ordos basin	Gy, Ha, Ca-ps-An	SG, Multiple horizons in > 20m	Inter-tidal	Meng et al., 2019; Xiong et al., 2020	
58	China	0.06-0.03	Paleogene, Tarim	Gy	?	Marginal marine to restricted basin	Chen et al., 2016	
59	China	0.02-0.005	Suwiyi Formation, Miocene, Tarim Basin	Gy, An, Ha	SG, Multiple horizons in > 10m	Intracratonic basin	Tang, 1992	
60	China	0.36-0.32	Lower Carboniferous, Tarim Basin	Gy, An, Ha	SG, Multiple horizons in > 10m	Intracratonic basin	Tang, 1992	
61	China	0.01-present	Quaternary, Qaidam Basin	Gy, Ha, An	SG, Multiple horizons in > 10m	Foreland basin	Schubel and Lowenstein, 1997	
62	China	0.28-0.26	Lucaogou Formation, Fengcheng Formation, Permian, Junggar Basin	Ha, Gy	?	Intracratonic basin	Zhang et al., 2021; Yu et al., 2018, 2019	
63	China	0.28-0.26	Pusige Formation,	Ha, Gy	?	Intracratonic basin	Wang et al., 2021	



			Permian, Tarim Basin				
64	China	0.23-0.20	Jialingjiang Formation, Leikoupo Formation, Triassic, Sichuan Basin	Gy, An	?	Intracratonic basin, Tidal flats, lagoon	Liu et al., 2018
65	China	0.10-0.06	Mengyejing Formation, Yunlong Formation, Late Cretaceous, Simao Basin,	Gy, An, Ha, Syl	?	Marginal marine	Liu et al., 2018
66	China	0.10-0.03	Tuyiluke Formation, Altash Formation, Kumugelimu Formation, Late Cretaceous to Oligocene, Tarim Basin	Gy, An	?	Marginal marine	Liu et al., 2018
67	China	0.18-0.16	Qiangtang basin	Gy, An	?	Intracratonic basin, lagoonal, subkha	Li et al., 2008a, b
68	China	0.01-present	Late Miocene to Pliocene, Qaidam Basin, China	Gy, Ha	?	Intracratonic basin	Guo et al., 2018
69	China	0.06-0.02, 0.01-present	Eocene, Qaidam, Tarim, Junggar, Bohai Bay, Subei, Nanxiang, Jiangnan Basin, South China Grabens	Gy, Ha	?	Intracratonic basin	Guo et al., 2020; Jiang et al., 2013
70	China	0.15-0.03	South China graben, Middle-Lower reaches of Yangtze River	GY, Ha, An	SG	Intracratonic basin	Teng et al., 2019
71	China	0.06-0.02	Paleogene, Tarim basin	Gy, An	Multiple horizons	SG, marine	Zhu et al., 2012



in >200m							
72	Europe	0.48-0.43	Middle Ordovician to Late Silurian red beds, Baltic basin	Red beds	?	Intracratonic sedimentary basin	Kiipli et al., 2009
73	Europe	0.26-0.25	St. Bees evaporites, Upper Permian, England	An, Gy	?	Intracratonic basin	Arthurton and Hemingway, 1972; Kiersnowski et al., 1995
74	Europe	0.32-0.25	Late Carboniferous to Permian, Lodeve Basin, Autun Basin, Bourbon l'Archambault basin, Saar-Nahe Basin, Thuringian Forest Basin, Saale Basin, North German-Polish Basin, Central and Western Bohemian Basins, Intra-Sudetic Basin, Boskovice Graben, Moroccan Basins	Ha, Gy, An	SG, Multiple horizons in > 100m	Intracratonic basin	Roscher and Schneider, 2006
75	Europe	0.27-0.25	Permian, Zechstein Basin	Ha, Gy, An	SG, Multiple horizons in > 10m	Isolated deep basin	Regenspurg et al., 2016; Lüders et al., 2010
76	Europe	0.28-0.25	Rotliegend Supergroup, South Permian Basin, Europe	Ha, Gy	SG, Multiple horizons in > 100m	Intracratonic basin	Kiersnowski et al., 1995; Gaupp et al., 2000
77	Europe	0.28-	Unidad Roja	Ha, Gy	?	Intracratonic	Valero Garces



		0.25	Superior Group, Permian, Aragon-Bearn Basin, Pyrenees basins			basin	and Aguilar, 1992
78	Europe	0.23-0.20	Late Triassic, Foreland basins of pyrenees	Gy, An	?	Intracratonic rift basin	Puigdefabregas and Souquet, 1986
79	Europe	0.23-0.20	Keuper Formation, Late Triassic, Central Europe	Gy, An, Ha	?	Intracratonic basin	Barth, 2000
80	Europe	0.02-0.01	Miocene, Ebro Basin	Gy, An, Ha	?	Intracratonic basin	Sanchez et al., 1999
81	Europe	0.23-0.20	Mercia Group, Upper Triassic, Bristol Channel Basin, U.K.	Gy, An, Ha	?	Intracratonic basin	Nemcok and Gayer, 1996
82	Europe	0.30-0.20	Permian, Triassic, Upper Triassic, Bay of Biscay, Europe basins	Gy, An, Ha	SG	Intracratonic basin	Bourquin et al., 2011
83	Europe	0.06-0.02	Paleocene to Miocene, Ebro, Tajo, Calatayud and Teruel basin, Spain	Gy, Ha	?	Intracratonic basin	Utrilla et al., 1992
84	Europe	0.40-0.38	Middle Devonian, Morsov Basin	Gy, An, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014
85	Europe	0.39-0.36	Upper Devonian, Dnieper Donets basin	Gy, An, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014
86	Europe	0.39-0.36; 0.30-0.25	Upper Devonian, Pripjat depression	Gy, An, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014
87	Europe	0.30-0.25	Permian, Dnieper Donets basin	Gy, An, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014
88	Latin America	0.16-0.14	Minas Viejas Formation, Late	Gy, An, Ha	SG	Marine to lagoonal	Guzman, 1962



			Jurassic, Sierra Madre Oriental, Mexico					
89	Latin America	0.16-0.14	Upper Jurassic, Mexican Gulf Coastal Plain, Mexico	Gy, An, Ha	SG	Marginal basin	Guzman, 1962	
90	Latin America	0.16-0.13	Late Jurassic to Early Cretaceous, Caribbean	Gy, Ha	SG	marine	Iturralde-Vinent, 2006	
91	Mediterranean and Black Sea	0.07-0.05	Messinian Saline Late Miocene, Mediterranean	Gy, Ha, An, Dol-ps-An	10 ⁶ km ³	SG, Desiccating deep basin	Mu"ller and Hsu", 1987	
92	Middle-East	0.63-0.50	Birba Formation, U/Athel Formation, Noor Formation, South Oman Salt Basin	Al Ha, An, Ca-ps-Ha	Multiple horizons in > 200m	SG, marginal to restricted marine environments	Ramseyer et al., 2013	
93	Middle-East	0.02-present	Miocene, Pleistocene, Danakil depression, Ethiopian rift, Red Sea Basin	Ha, Gy, Eva	Multiple horizons in > 200m	SG, Marine	Hutchinson and Engels, 1970	
94	Middle-East	0.04-0.01	Tuzluca Formation, Kagizman and Tuzluca basin, Anatolian basins	Gy, Ha	SG	Marginal marine	Palmer et al., 2004; Varol et al., 2015	
95	North Atlantic-Canada	0.23-0.18	Argo Formation, Late Triassic to Early Jurassic, Scotian Basin	Ba, Gy, An	SG	Intracratonic basin	Pe-Piper et al., 2015	
96	Russia	0.68-0.57	Motaya Formation, Eastern Siberia	An, as layers; Ca-ps-Gy; Ps-Ha	Multiple horizons in > 20m	SG, non-marine to marine	Vysotsky et al., 2014	
97	Russia	0.54-0.51	Jonov Formation, Huchtuisik Formation,	Ha, Gy, as layers; Ca-ps-Gy	Multiple (16) horizons in > 400m	SG, non-marine to marine, >60000 km ³	Vysotsky et al., 2014	



			Eastern Siberia, Ussoli Formation, Lower Cambrian					
98	Russia	0.30- 0.25	Permian, Ural depression	Gy, An, Ha, Syl	Multiple horizons in >20m	SG, marine	Vysotsky et al., 2014	
99	South America	0.15- 0.08	Caqueza Group, Villeta Group, Cretaceous, Cordillera Oriental, Columbia	An, Gy	SG, Multiple horizons in > 5m	foreland basin	McLaughlin, 1972	
100	South America	0.28- 0.25	Rio Bonito Formation, Permian, Parana Basin, Brazil	Ha, Gy	?	Intracratonic basin	Ketzer et al., 2009	
101	South America	0.28- 0.25	Motuca Formation, Parnaiba Basin, Brazil	Ha, Gy	SG, ?	Intracratonic basin	Abrantes Jr et al., 2019	
102	South America	0.13- 0.06	Codochedo Formation, Quebrada Monardes Formation, Atacama region	Gy, An	?	Marginal marine	Bell and Suarez, 1993	
103	South America	0.04- present	Purilactis Group, Oligocene- Pliocene, Salar de Atacama basin	Gy, An	SG	Intracratonic basin	Mpodozis et al., 2005	
104	South America	0.06- present	Cenozoic, Central Andes	Gy, Ha, An	SG	Intracratonic basin	Diaz, 1988	
105	South America	0.30- 0.25	Permian, Amazon Basin, Brazil	Gy	?	?	Vysotsky et al., 2014	
106	Southeast Asia	0.10- 0.06	Late Cretaceous, Laos-Khorat basin	Gy, An	?	Marginal marine	Liang and Xu, 2022	
107	Southeast ern Asia	0.10- 0.06	Upper Cretaceous,	Gy, An, Syl	Multiple horizons	SG	Vysotsky et al., 2014	



		Nam Kam Basin		in >20m			
108	USA	0.46-0.44	Ordovician Galena Group, Illinois basin	Gy, An	?	Marginal basin	Stueber and Walter, 1994
109	USA	0.43-0.42	Niagara and Salina Group, Silurian, Michigan basin	Ha, Gy, Car, Ca-ps-Gy	SG, >500m thick	Intracratonic sedimentary basin	Wilson and Long, 1993; Elliott and Aronson, 1993
110	USA	0.42-0.40	Detroit River Group, Lower Devonian, Michigan basin	Gy, Ha, An	SG, Multiple horizons in > 30m	Intracratonic basin	Wilson and Long, 1993
111	USA	0.43-0.41	Salina Group, Appalachian basin	Ha, Gy, Car, Ca-ps-Gy	SG, >500m thick	Intracratonic sedimentary basin	Johnson, 1997
112	USA	0.36-0.32	Mississippian Michigan Formation, Michigan basin	Gy	SG, Multiple horizons in > 1-10m	Intracratonic sedimentary basin	Johnson, 1997
113	USA	0.33-0.32	St. Louis Formation, Illinois basin	Evaporite karst	?	Intracratonic basin	McGrain and Helton, 1964
114	USA	0.41-0.36	Wapsipinicon Group, Cedar Valley Group, Devonian, Forest City basin	Gy, An, Evaporite karst	Multiple horizons in > 1-10m	Intracratonic basin	Witzke et al., 1988; Cody et al., 1997
115	USA	0.20-0.15	Fort Dodge Formation, Jurassic, Forest City basin	Evaporite karst	Multiple horizons in > 1-10m	Intracratonic basin	Witzke et al., 1988; Cody et al., 1997
116	USA	0.17-0.15	Louann Salt and Salt Domes, Gulf Coast basin	Ha, Gy, Evaporite karst	Multiple horizons in > 500m	SG, marine	Babel and Schreiber, 2014
117	USA	0.27-0.25	Schneibly Hill Formation, Holbrook basin	Ha, Gy, Evaporite karst	?	SG	Neal and Johnson, 2002
118	USA	0.32-0.30	Hermosa Formation, Paradox basin	Ha, Evaporite karst	?	Salt anticline	Hite and Lohman, 1973; Doelling, 1988
119	USA	0.17-0.15	Jurassic Arapien Formation,	Ha, Evaporite karst	?	Salt anticline	Willis, 1986; Witkind, 1994



Paradox basin							
120	USA	0.26-0.25	Castile Formation, Salado Formation, Rustler Formation, Delaware basin	Gy, Ha, An, Evaporite karst	SG, Multiple horizons in > 500m	Intracratonic basin	Dean and Johnson, 1989; Kirkland and Evans, 1980
121	USA	0.26-0.25	Castile Formation, Salado Formation, Rustler Formation, Midland basin	Gy, Ha, An, Evaporite karst	SG, Multiple horizons in > 500m	Intracratonic basin	Dean and Johnson, 1989; Kirkland and Evans, 1980
122	USA	0.27-0.25	Artesia Group, San Andres Formation, New Mexico	Gy, Ha, An, Evaporite karst	SG, Multiple horizons in > 100m	Intracratonic basin	Forbes and Nance, 1997
123	USA	0.27-0.25	Blaine Formation, Cloud Chief Formation, Texas and Oklahoma	Gy, Ha, An, Evaporite karst	SG, Multiple horizons in > 5m	Intracratonic basin	Johnson, 1990, 1992
124	USA	0.14-0.10	Terrett Formation, Cretaceous, Central Texas	Gy, Evaporite karst	>10m	Intracratonic basin	Warren et al., 1990
125	USA	0.30-0.25	Permian, Palo Duro Basin, Texas	Gy, An	SG	Intracratonic basin	Gu and Eastoe, 2021
126	USA	0.25-0.20	Spearfish Formation, South Dakota,	Gy, Evaporite karst	~5m	Intracratonic basin	Rahn and Davis, 1996; Davis and Rahn, 1997
127	USA	0.36-0.32	Madison Formation, Wyoming	Gy, Evaporite karst	?	Intracratonic basin	Sando, 1988
128	USA	0.28-0.25	Permian, Appalachian basin	An, Gy	?	Intracratonic basin	Rowen et al., 2015



129	USA	0.28-0.26	Opeche Shales, Nippewalla Group, Ochoan Series, Summer Group, Chase Group, Wolfcampian Formation, Salado Formation, Permian, Palo Duro Basin, Denver Basin, Paradox Basin, Central Basin, Delaware Basin, Midland Basin, North America	Ha, Gy, An, Syl, Ca	SG, Multiple horizons in > 100m	Intracratonic basin	Benison et al., 1998; Engle et al., 2016
130	USA	0.23-0.19	Newark Supergroup, North America grabens	Gy	?	Intracratonic basin	Smoot et al., 1991
131	USA	0.14-0.10	Edwards Group, Lower Cretaceous, Gulf Basin, Texas	Gy, An	SG	Marginal marine	Land and Prezbindowski, 1981; Hanor and Mcintosh, 2007
132	USA	0.02-present	Late Great Salt Lake, Picacho basin, Safford basin, Tucson basin, SE North America	Gy, Ha	?	Intracratonic basin	Kowalewska and Cohen, 1998

- 2160 Saline Giants (SG), with volume > 1000km³;
- 2161 An, anhydrite;
- 2162 Ank-ps-Gy, ankerite pseudomorphs after gypsum;
- 2163 Ba, barite; Ba-ps-Gy, barite pseudomorphs after gypsum;
- 2164 Ca-Evp, calcitized evaporites; Ca-ps-Ar, carbonate pseudomorphs after aragonite;
- 2165 Ca-ps-An, carbonate pseudomorphs after anhydrite;
- 2166 Ca-ps-Gy, carbonate pseudomorphs after gypsum;
- 2167 Ca-ps-Gy or -Ar, carbonate pseudomorphs after gypsum or aragonite;
- 2168 Dol-ps-An, dolomite pseudomorphs after anhydrite;
- 2169 Dol-ps-Gy, dolomite pseudomorphs after gypsum; Gy, gypsum;
- 2170 Kao-ps-Gy or An, kaolinite pseudomorphs after gypsum or anhydrite;



2171 ps-An, pseudomorphs after anhydrite;
2172 ps-Evp, pseudomorphs after evaporites;
2173 ps-Gy, pseudomorphs after gypsum;
2174 ps-Mir, pseudomorphs after mirabilite;
2175 ps-Nah, pseudomorphs after nahcolite;
2176 ps-Nat, pseudomorphs after natron;
2177 ps-Ha, pseudomorphs after halite;
2178 ps-Sh, pseudomorphs after shortite;
2179 ps-Tro, pseudomorphs after trona;
2180 Q-ps-An, quartz pseudomorphs after anhydrite;
2181 Q-ps-Gy, quartz pseudomorphs after gypsum;
2182 sc-breccias, solution collapse breccias;
2183 Si-Evp, silicified evaporites;
2184 Si-ps-An, silicified pseudomorphs after anhydrite;
2185 Si-ps-Ar, silicified pseudomorphs after aragonite;
2186 Si-ps-Gy, silicified pseudomorphs after gypsum;
2187 Si-ps-Nah, silicified pseudomorphs after nahcolite;
2188 Si-tr-An, quartz filled traces after anhydrite;
2189 Si-tr-Gy, quartz filled traces after gypsum.