



1 Total Column Ozone Trends from the NASA Merged Ozone Time Series 1979 to 2021 Showing Limited  
2 Recovery to 1979 Amounts after Declining into the Mid 1990s

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17 **Abstract**

18 Monthly averaged total column ozone data  $\Omega_{\text{MOD}}$  from the Merged Ozone Data set (MOD) were  
19 examined to show that the latitude-dependent ozone depletion turnaround dates  $T_A(\theta)$  range from 1994  
20 to 1998.  $\Omega_{\text{MOD}}$  used in this study was created by combining data from Solar Backscattered Ultraviolet  
21 instruments (SBUV/SBUV-2) and the Ozone Mapping and Profiler Suite (OMPS-NP) from 1979 to 2021.  
22  $T_A(\theta)$  is defined as the date when the zonally average ozone ceased decreasing. The new calculated  
23 systematic latitude-dependent  $T_A(\theta)$  shape should appear in atmospheric models that combine the  
24 effects of photochemistry and dynamics in their estimate of ozone recovery. Trends of zonally averaged  
25 total column ozone in percent per decade were computed before and after  $T_A(\theta)$  using two different  
26 trend estimate methods that closely agree, Fourier Series Multivariate Linear Regression and linear  
27 regression on annual averages. During the period 1979 to  $T_A(\theta)$  the most dramatic rates of SH ozone loss  
28 were  $P_D = -10.9 \pm 3\%$  per decade at  $77.5^\circ\text{S}$  and  $-8.5 \pm 0.9\%$  per decade at  $65^\circ\text{S}$ , which is about double the  
29 NH rate of loss of  $P_D = -5.6 \pm 4\%$ /decade at  $77.5^\circ\text{N}$  and  $4.4 \pm 1\%$ /decade at  $65^\circ\text{N}$  for the period 1979 to  
30  $T_A(\theta)$ . After  $T_A(\theta)$ , there has been an increase at  $65^\circ\text{S}$  of  $P_D = 1.6 \pm 1.4\%$  per decade with smaller increases  
31 from  $55^\circ\text{S}$  to  $25^\circ\text{S}$  and a small decrease at  $35^\circ\text{N}$  of  $-0.4 \pm 0.3\%$ /decade. Except for the Antarctic region,  
32 there only has been a small recovery in the Southern Hemisphere toward 1979 ozone values and almost  
33 none in the Northern Hemisphere.

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## 37 1.0 Introduction

38 Ozone is a photolytically produced, photochemically destroyed, and dynamically distributed  
39 atmospheric gas that plays a crucial role in protecting the planet from harmful ultraviolet (UV) radiation  
40 from the sun. The atmospheric presence of bromine and the release of chlorine from the UV  
41 dissociation of man-made chemicals, such as chlorofluorocarbons (CFCs), can break down the ozone  
42 layer. This is especially the case in the Antarctic region where heterogeneous chemistry on and within  
43 ice crystals in polar stratospheric clouds PSCs have a strong effect on the destruction of ozone during  
44 September and October (Solomon et al., 1986; 1999; 2016; Crutzen and Arnold, 1986; Khosrawi et al.,  
45 2011). As the sun rises in spring, chemically active nitrogen oxides, chlorine and bromine are released  
46 causing the ozone hole to develop within the region enclosed by the polar vortex winds. The weak levels  
47 of sunlight initiate and maintain the catalytic ozone loss photochemistry. In November and December,  
48 the isolating polar vortex winds break down and the Antarctic ozone hole region back fills from southern  
49 mid-latitudes. The recurring annual event triggered international action to limit the production and use  
50 of ozone-depleting substances (ODS) under the Montreal Protocol, which has been successful in  
51 reducing the emission of these substances, slowing down the depletion of the ozone layer globally,  
52 leading to a partial recovery in the Antarctic ozone hole region (Solomon et al., 2016; Strahan and  
53 Douglass 2018). After the mid-1990s, several studies have reported an increase in TCO, particularly in  
54 the mid to high latitudes of the Southern Hemisphere, as well as a reduction in the size and depth of the  
55 Antarctic ozone hole starting in the late 1990s (Solomon et al., 2016; Stone et al., 2018, Weber et al.,  
56 2022).

57 Ozone recovery was first observed in the mid-1990s when satellite data showed a stabilization and slight  
58 increase in ozone concentrations in the Antarctic ozone hole region. However, the recovery was not  
59 significant enough to be considered a trend at that time (Strahan and Douglass 2018). In the early 2000s,  
60 further analysis of satellite and ground-based data showed that the rate of ozone depletion had slowed  
61 down. After the mid-1990's, the cessation of ozone depletion has been most evident in the Southern  
62 Hemisphere SH polar region, where ozone depletion had been most severe, but the recovery has been  
63 slow or non-existent at other latitudes. Recently, Weber et al. (2022) reported positive statistically  
64 significant total column ozone (TCO) trends from 1996-2020 at southern middle and high latitudes, and  
65 over the SH polar cap in September. When dynamical terms were included in the regression, small  
66 positive trends were near the 2-standard deviation  $2\sigma$  threshold at northern mid- and high-latitudes,  
67 with no trend detected in the tropics or over the NH polar cap.

68 Despite the success of the Montreal Protocol (Velders and Andersen, 2018), ozone concentrations  
69 continue to fluctuate, driven by natural and anthropogenic factors, such as changes in solar radiation,  
70 stratospheric circulation, global warming, and changing emissions of ozone precursors (Dameris and  
71 Baldwin, 2012; Weber et al., 2022). The discussion by Dameris and Baldwin (2012) explores possible  
72 effects of climate change on the dynamics of the atmosphere affecting ozone as ODSs change, and  
73 particularly the change in the Brewer-Dobson circulation (Brewer, 1949; Dobson et al., 1926) that  
74 transports ozone to mid- and high-latitudes.



75 This study will calculate ozone trends based on new latitude dependent ozone recovery dates  $T_A(\theta)$   
 76 ranging from 1994 (equatorial region and  $60^\circ\text{N}$  to  $70^\circ\text{N}$ ) to 1998 ( $60^\circ\text{S}$  –  $80^\circ\text{S}$ ). Ozone data used in this  
 77 study are a subset of the Merged Ozone Data MOD set (1970 – 2021) starting in 1979 with the Nimbus-7  
 78 SBUV (Solar Backscattered Ultraviolet) satellite instrument. From 1979 to 2021, the MOD data set was  
 79 created by combining data from Solar Backscattered Ultraviolet instruments (SBUV/SBUV-2) and the  
 80 Ozone Mapping and Profiler Suite (OMPS-NP). Methods of calculating trends from time series data are  
 81 essential in the analysis of environmental and climate-related data. Here, we discuss two independent  
 82 methods to estimate linear trends: 1) linear regression of annual averaged data and 2) Fourier time  
 83 series decomposition or multivariate linear regression MLR (Ziemke et al., 2019). The two methods are  
 84 compared and shown to give nearly identical results over their mutual latitude range of validity,  $65^\circ\text{S}$  to  
 85  $65^\circ\text{N}$ . The MLR method is not used in the regions poleward of the Arctic and Antarctic circles that have  
 86 extended winter polar night.

### 87 1.1 The Merged Ozone Data Set MOD

88 Figure 1 (left panel) shows the MOD zonally averaged  $\Omega_{\text{MOD}}$  TCO data (Frith et al., 2014; 2020) set as a  
 89 function of latitude ( $5^\circ$  latitude bands from  $77.5^\circ\text{S}$  to  $77.5^\circ\text{N}$ ) and time (January 1979 to December  
 90 2021). Part of the Antarctic ozone hole ( $75^\circ\text{S}$  to  $80^\circ\text{S}$ ) is shown (blue color), and the high latitude  
 91 maxima, North and South, (red color), with low values in the equatorial region. Figure 1 (right panel)  
 92 shows the 42-year zonal and time averaged ozone amounts and the maxima and minima annual  
 93 envelopes as a function of latitude. Figure 1 shows the asymmetry in the monthly and zonally averaged  
 94 ozone data between the hemispheres, with the Northern Hemisphere NH having more ozone than the  
 95 Southern Hemisphere SH at corresponding latitudes. Part of the asymmetry is driven by the Spring  
 96 Antarctic ozone hole backfilling in the SH summer.

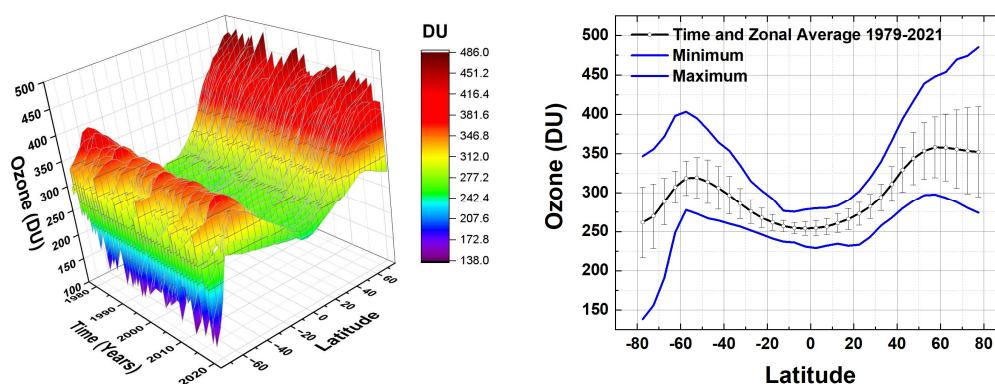


Fig. 1 Left: The zonally and monthly averaged  $\Omega_{\text{MOD}}$  data set 1979 – 2021 and  $-77.5^\circ$  to  $77.5^\circ$ . Right: Time and zonal averaged ozone and its maxima and minima 1979 – 2021. Error bars are 1 standard deviation  $\pm 1\sigma$ .

97  $\Omega_{\text{MOD}}(t, \theta)$  provides a global view of ozone levels needed to track changes in ozone concentrations over  
 98 time  $t$  for each latitude band  $\theta$ . The SBUV and OMPS-NP series of satellite instruments form the longest  
 99 (1979 to 2022) continuous global ozone  $\Omega_{\text{MOD}}(t, \theta)$  data record from a single instrument type. Merged



100 ozone retrievals from the individual instruments use the version 8.7 retrieval algorithm (described by  
101 Weber et al., 2022) as an extension of the version 8.6 algorithm (Bhartia et al., 2013; McPeters et al.,  
102 2013; DeLand et al., 2012; Frith et al., 2017) specifically designed to improve cross calibrations between  
103 the later SBUV-type instruments in MOD starting from NOAA-16 in 2000. There were no external  
104 adjustments made to the ozone retrieval except for small high-altitude diurnal corrections to account  
105 for different measurement times between satellites and varying measurement time of day as individual  
106 satellite orbits slowly drift in equator crossing time. These adjustments are very minor in TCO (Frith,  
107 personal communication). Data from each instrument are selected based on quality criteria outlined in  
108 Frith et al. (2014; 2020) and then data are averaged during periods when more than one instrument was  
109 operational. The  $\Omega_{\text{MOD}}(t, \theta)$  are available as a function of latitude and month,  
110 [https://acd-ext.gsfc.nasa.gov/Data\\_services/merged/](https://acd-ext.gsfc.nasa.gov/Data_services/merged/).

111  
112 Analysis of the long-term ozone time series has been looked at extensively with references given in  
113 Weber et al., 2022. Methods for estimating trends from an oscillating time series with several distinct  
114 periodicities are well known (Ziemke et al., 2019; Stolarski et al. 1991;1992, Herman et al., 1993). For  
115 ozone, one of the difficulties in trend estimation is that the early part of the time series shows a strong  
116 ozone decrease at all latitudes that continued until the mid-1990s and then flattens out and shows  
117 almost no recovery thereafter toward 1979 values. The  $\Omega_{\text{MOD}}$  time series has been used extensively in  
118 ozone assessments and State of the Climate reports and was recently compared to several other merged  
119 total ozone records in Weber et al., 2022. The validity of the  $\Omega_{\text{MOD}}$  time series for estimating ozone  
120 trends was further checked (See Appendix Figs. A1 and A2) in this study by showing detailed  
121 comparisons between the deseasonalized  $\Omega_{\text{MOD}}$  time series with the deseasonalized MLS (Microwave  
122 Limb Sounder) overlapping stratospheric ozone time series (2005 to 2023).

## 123 2.0 Trend Estimates from the MOD Ozone Data

124 Multivariate Linear Regression MLR is a Fourier based method for analyzing atmospheric time series  
125 data that decomposes the time series into its component parts, including trend, quasi-biennial  
126 oscillation QBO, solar cycle, ENSO (El Nino Southern Oscillation), seasonality, and noise resulting in a  
127 trend estimate and 2-standard deviation  $2\sigma$  uncertainty estimates (Ziemke et al., 2019). Calculated  $2\sigma$   
128 uncertainties for the MLR trends included a first order autoregressive adjustment applied to the derived  
129 residuals (Weatherhead et al., 1998).

130 Linear trend estimates for the long-term changes in  $\Omega_{\text{MOD}}(t, \theta_i)$  globally and as a function of latitude  $\theta_i$   
131 have been obtained using a generalized multivariate linear regression (MLR) model (e.g., Randel and Cobb,  
132 1994, and references therein). Trends  $B(\theta_i)$  were determined for  $\Omega_{\text{MOD}}(t, \theta_i)$  using Eqns. 1 and 2.

133

$$\Omega_{\text{MOD}}(t, \theta_i) = A(\theta_i, t) + B(\theta_i, t) \cdot t + C(\theta_i, t) \cdot \text{QBO}_1(t) + D(\theta_i, t) \cdot \text{QBO}_2(t) + E(\theta_i, t) \cdot \text{ENSO}(t) + F(\theta_i, t) \cdot \text{Solar}(t) + R(\theta_i, t) \quad (1)$$

134

135 where  $t$  is the month index ( $t=1$  to 504 months with data for 1979–2021),  $A(\theta_i, t)$  is the seasonal cycle  
136 coefficient,  $B(\theta_i, t)$  is the trend coefficient,  $C(\theta_i, t)$  is the first empirical orthogonal function (EOF) QBO



137 coefficient,  $D(\theta_i, t)$  is the second EOF QBO coefficient, both representing the major components of the  
 138 QBO variability,  $E(\theta, t)$  is the ENSO coefficient,  $F(\theta, t)$  is the solar cycle coefficient, and  $R(t)$  is the  
 139 residual error time series. The F10.7 cm solar flux monthly time series is used for the Solar(t) proxy, first  
 140 and second leading EOF QBO monthly time series proxies  $QBO_1(t)$  and  $QBO_2(t)$  are used for the QBO  
 141 component (Wallace et al., 1993), and Nino 3.4 is used for ENSO(t).  $A(\theta, t)$  involves 7 fixed constants while  
 142  $B(\theta, t)$  (and all other remaining coefficients) involves 5 fixed constants for each  $\theta_i$ . The harmonic expansion  
 143 for  $A(t)$  (similar for the other coefficients) is (Eqn. 2)  
 144

$$A(t) = a(0) + \sum_{p=1}^3 [a(p) \cos(2\pi p t / 365) + b(p) \sin(2\pi p t / 365)] \quad (2)$$

145 where  $a(p)$  and  $b(p)$  are constants. Statistical uncertainties for  $A(t)$  and  $B(\theta_i)$  were derived from the  
 146 calculated statistical covariance matrix involving the variances and cross-covariances of the constants  
 147 (e.g., Guttman et al., 1982; Randel and Cobb, 1994). The linear deseasonalized trend results  $B(\theta_i)$  are  
 148 obtained for 14 latitude bands  $\theta_i$  (centered on  $65^\circ\text{S}$  to  $65^\circ\text{N}$ ). The latitudinal trends  $P_D(\theta_i)$  are expressed  
 149 in %/Decade given by Eq. 3, where the denominator  $D$  is either the time average of the area weighted  
 150 global ozone average (Fig 3) or the time average  $D(\theta_i) = \langle \Omega_{\text{MOD}}(t, \theta_i) \rangle$  for each latitude band over the  
 151 considered period. The whole year period considered is 1979 – 2021.  
 152

$$P_D(\theta_i) = 1000 B(\theta_i) / D(\theta_i) \quad (\% / \text{Decade}) \quad (3)$$

153  
 154  
 155 In the second method, the trend is estimated using annual integrals (annual averages) that remove the  
 156 seasonality and other short-term oscillations but ignores longer term oscillations such as the 2-to-3-year  
 157 QBO cycle and the 11-year solar cycle. A comparison of the two trend estimating methods is shown in  
 158 Fig.2 for the entire 1979 to 2021 period showing that they agree quite closely, but that the annual  
 159 average method has slightly larger two standard deviations  $2\sigma$  than the MLR method.

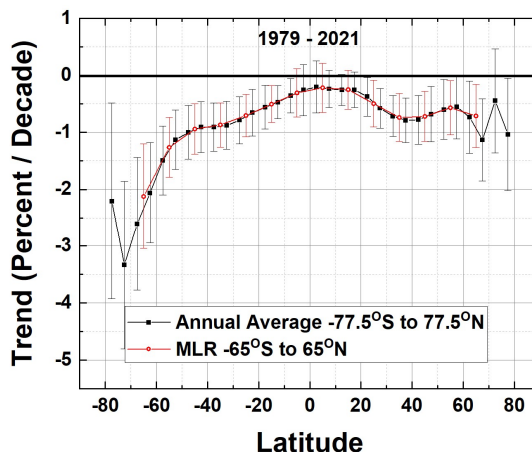


Fig. 2 The ozone trend  $P_D(\theta)$  for the entire period 1979 – 2021 for two methods, MLR and Annual Average. The latitude grids for the two methods are offset to show the agreement in the trends and  $2\sigma$  error bars.



160 The Fig. 2 estimation of linear long-term trends since 1979 is misleading, since ozone showed significant  
 161 annual declines until the mid-1990s and then increased slightly thereafter, meaning the average long-  
 162 term time series is non-linear. The usual procedure is to determine linear trends separately before and  
 163 after the turnaround dates  $T_A$  (Zhang and Thompson, 2021; Weber et al., 2022). However, as is shown  
 164 later, there is no single turnaround date applicable to all the latitudes between  $80^\circ\text{S}$  and  $80^\circ\text{N}$ . Instead,  
 165 there is a range spanning 1994 to 1998.

166 In this study the Locally Weighted Scatterplot Smoothing Lowess( $f$ ) technique is used to find the turn-  
 167 around dates (Figs. 3 and 4), where  $f$  = the fraction of data averaged together (Cleveland, 1979 and  
 168 Cleveland and Devlin, 1988).

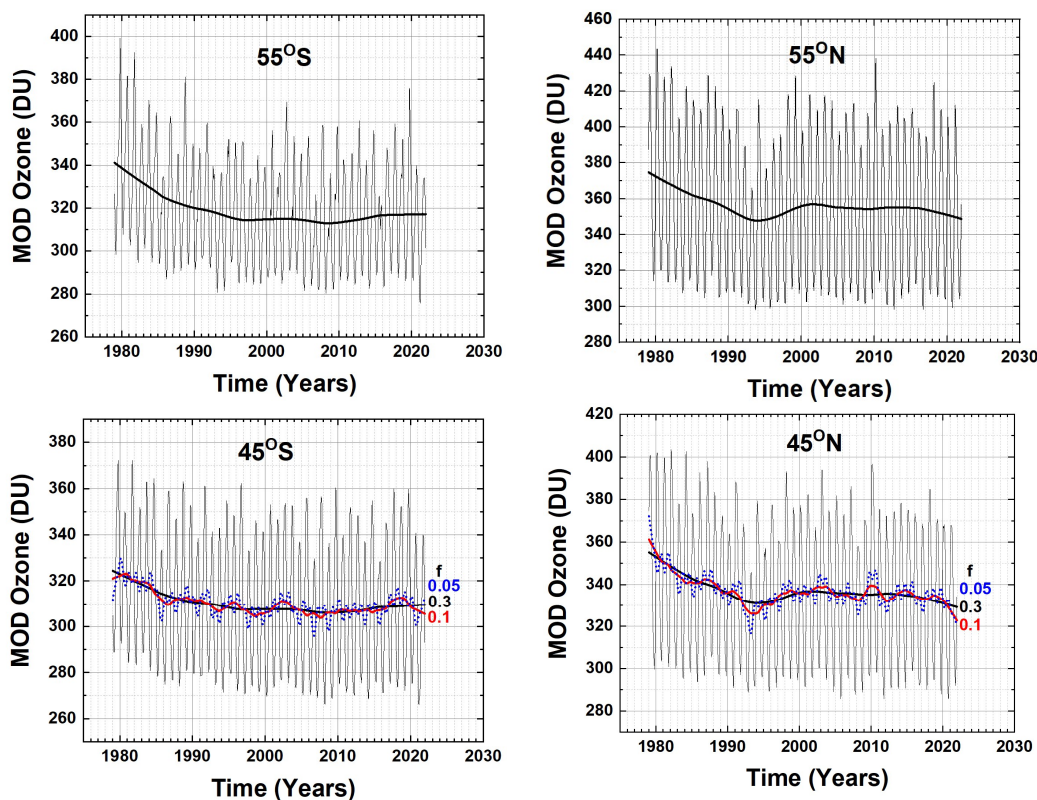


Fig. 3  $\Omega_{\text{MOD}}$  in four latitude bands and Lowess(0.3) fitting functions ( $f = 0.3$ , black lines). Examples of different  $f = 0.1$  (Red) and  $f = 0.05$  (blue dots) are shown at  $45^\circ\text{S}$  and  $45^\circ\text{N}$ . Note the slight downturn in the Lowess(0.3) at  $45^\circ\text{N}$  and  $55^\circ\text{N}$ .

169 Figure 3 shows the Lowess(0.3) fits (black curves) to the  $\Omega_{\text{MOD}}$  data for four sample latitude bands  $55^\circ\text{S}$ ,  
 170  $45^\circ\text{S}$ ,  $55^\circ\text{N}$ , and  $45^\circ\text{N}$  that tracks the longer-term changes in the  $\Omega_{\text{MOD}}$  time series. Also shown for  $45^\circ\text{N}$   
 171 and  $45^\circ\text{S}$  are examples of  $f = 0.1$  (red) and  $f = 0.05$  (blue dots). The values of  $T_A$  for  $f = 0.1$  and  $0.05$  are  
 172 approximately the same but are harder to see. The Lowess(0.3) degree of smoothing removes all short-  
 173 term effects on ozone such as volcanic eruptions from El Chichon (1982) and Mt. Pinatubo (1991), both



174 well before the earliest estimated turnaround time  $T_A$  in 1994. The Lowess(0.05) fit (blue dots) shows  
 175 considerable structure with a minimum in 1993 that is likely related to the Mt. Pinatubo eruption and a  
 176 modest El Nino effect in 1991-1992. In the equatorial zone at  $5^\circ\text{S}$  and  $5^\circ\text{N}$ , there are two minima,  
 177  $f=0.05$ , occurring in 1992 and 1994.

178 Figure 4 shows the Lowess(0.3) fits to the  $\Omega_{\text{MOD}}$  data (1979 to 2021) for 16 latitude bands,  $-75^\circ < \theta < 75^\circ$   
 179 on an expanded ozone scale. Each of the Lowess(0.3) plots for the various latitudes shows different  
 180 periods of ozone decrease and subsequent turnaround  $T_A(\theta)$  after the mid-1990's. Use of expanded  
 181 ozone scales appear to show a sharp downturn at some latitudes ( $25^\circ\text{N}$  to  $75^\circ\text{N}$ ). As shown later, the  
 182 apparent downturn in the Lowess(0.3) fit to  $\Omega_{\text{MOD}}$  after 2010 is not yet statistically significant as an  
 183 indicator of long-term decrease.

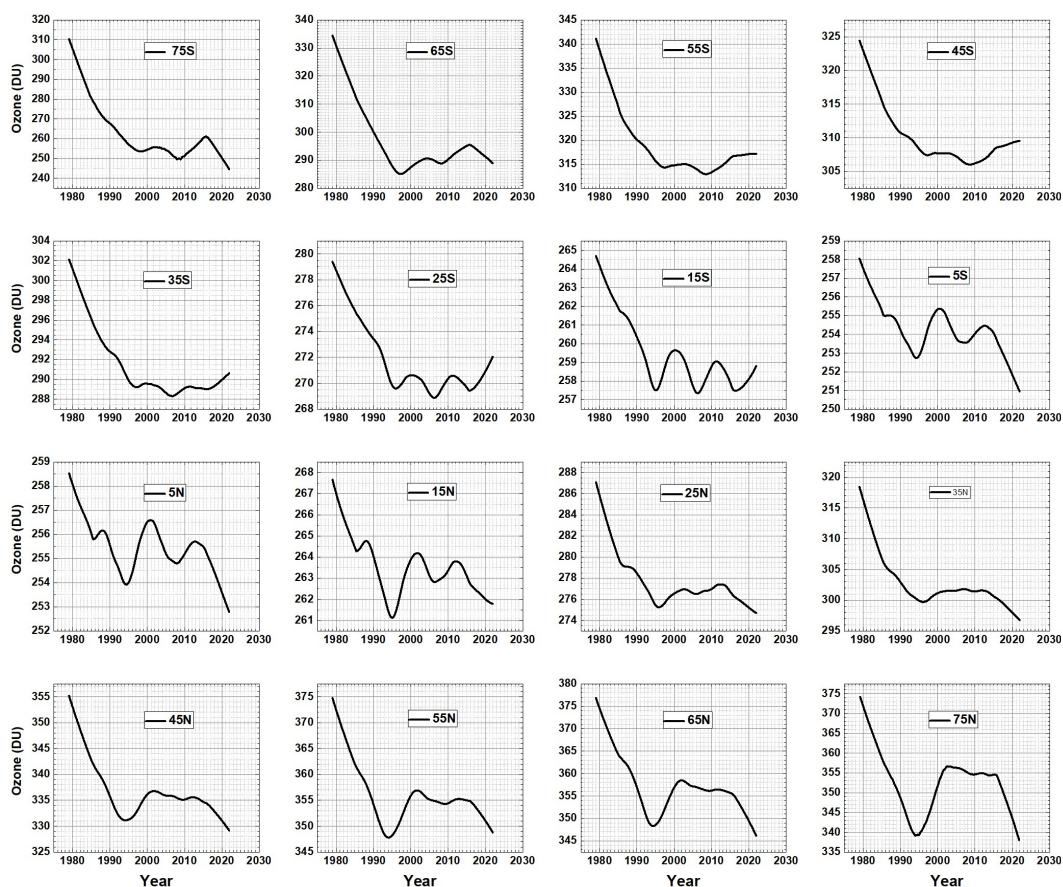


Fig. 4 Lowess(0.3) fits to the  $\Omega_{\text{MOD}}$  data for 16 latitude bands Used to determine the turnaround dates.

184

185 Figure 5 shows the turnaround dates  $T_A(\theta)$  that are obtained by taking the 1<sup>st</sup> derivative and finding the  
 186 zero-crossing time corresponding to the appropriate minimum value in Fig. 4. The exact turnaround





187 dates determined have a precision of  $\pm 0.1$  years. The uncertainty does not affect the calculation of  
 188 trends before and after the estimated  $T_A(\theta)$ . What is interesting is that some of the turnaround dates in  
 189 Fig. 5 are separated by over 4 years. Figure 5 shows a near symmetry for early turnaround dates 1994-  
 190 1996 for low latitudes between  $\pm 25^\circ$  that corresponds to the Brewer-Dobson ozone upwelling region  
 191 (Brewer et al., 1926; Dobson, 1949; Butchart, 2014) where most of the ozone is created by sunlight and  
 192 then transported poleward. At poleward latitudes, the turnaround dates are quite different, with a  
 193 delayed date, 1997, at high SH latitudes ( $45^\circ\text{S} - 65^\circ\text{S}$ ), 1998 at  $75^\circ\text{S}$ , and 1994 at high NH latitudes ( $45^\circ\text{N}$   
 194 to  $75^\circ\text{N}$ ). The SH delay to 1997 is caused by the presence of the Spring Antarctic Ozone Hole and  
 195 backfilling during the summer months after the polar vortex winds break down in October - November.

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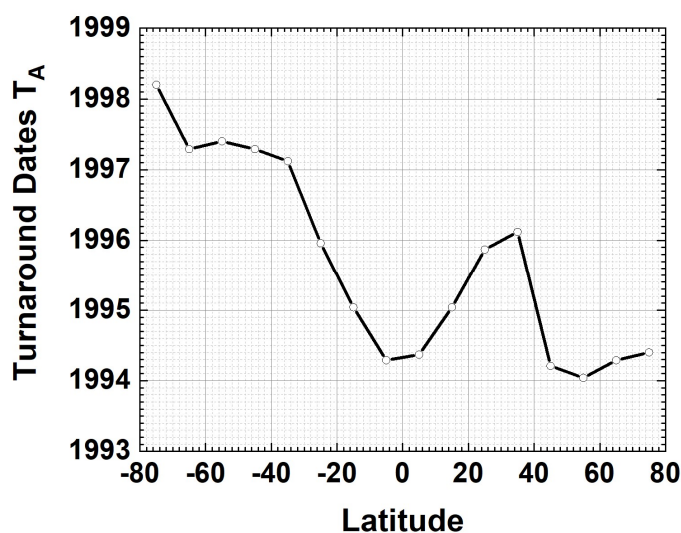


Fig 5 Turnaround dates  $T_A(\theta)$  as a function of latitude from Fig.4

197

Table 1 Turnaround Dates  $T_A(\theta)$

Latitude	$T_A$	Latitude	$T_A$
-75	1998.20	5	1994.37
-65	1997.29	15	1995.04
-55	1997.46	25	1995.87
-45	1997.29	35	1996.12
-35	1997.12	45	1994.21
-25	1995.96	55	1994.04
-15	1995.04	65	1994.29
-5	1994.29	75	1994.40

198 The general  $T_A(\theta)$  pattern shown in Fig. 5 should appear in model calculations as a signature of the  
 199 combined effects of photochemistry and dynamics on the cessation of decreasing ozone in the mid-  
 200 1990s.



201 Trends (linear slopes)  $P_D(\theta)$  in percent per decade are estimated (Eqn. 3) for the separate periods before  
 202 and after  $T_A(\theta)$  in each latitude band (Fig. 6) and for the entire period (Fig. 2). The linear slopes obtained  
 203 by the two methods, MLR and annual average closely agree (Figs. 2 and 6). Table 2 contains the data  
 204 from Figs. 6a and 6b.

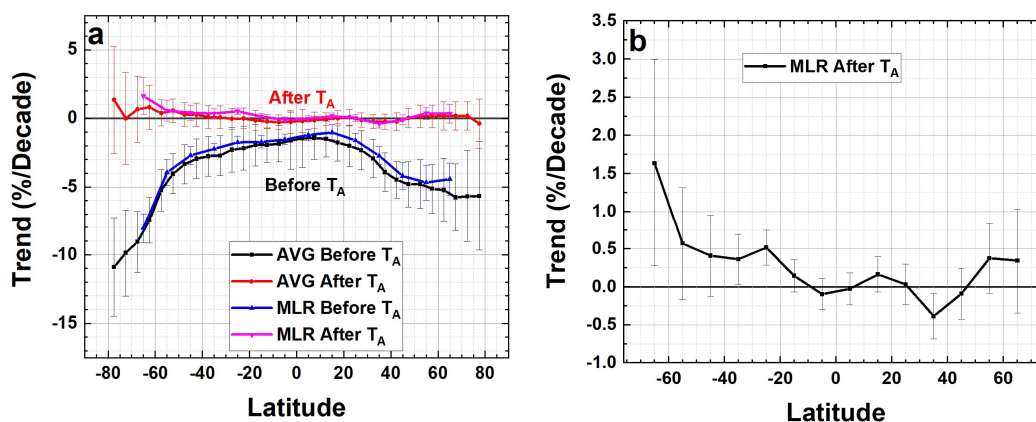


Fig. 6a Ozone trends  $P_D(\theta)$  (percent per decade) using the MLR and Annual Average methods before and after the turn-around times  $T_A(\theta)$  in Fig. 5. 6b A magnified version of the MLR estimated trends after  $T_A$  with  $2\sigma$  uncertainties.

205

Table 2 MLR Trends (%/decade)  $\pm 2\sigma$

Latitude	$P_D$ Before $T_A$	$P_D$ After $T_A$
-65	-8.04 $\pm$ 1.1	1.64 $\pm$ 1.4
-55	-3.93 $\pm$ 1.0	0.57 $\pm$ 0.7
-45	-2.69 $\pm$ 0.7	0.41 $\pm$ 0.5
-35	-2.22 $\pm$ 0.4	0.36 $\pm$ 0.3
-25	-1.75 $\pm$ 0.5	0.52 $\pm$ 0.2
-15	-1.71 $\pm$ 0.4	0.15 $\pm$ 0.2
-5	-1.54 $\pm$ 0.4	-0.10 $\pm$ 0.2
5	-1.21 $\pm$ 0.4	-0.03 $\pm$ 0.2
15	-1.01 $\pm$ 0.6	0.16 $\pm$ 0.2
25	-1.61 $\pm$ 0.5	0.03 $\pm$ 0.3
35	-2.71 $\pm$ 0.6	-0.39 $\pm$ 0.3
45	-4.20 $\pm$ 1.0	-0.09 $\pm$ 0.3
55	-4.67 $\pm$ 1.3	0.38 $\pm$ 0.5
65	-4.43 $\pm$ 1.2	0.35 $\pm$ 0.7

206

207 The latitude dependent trends derived by Weber et al., 2022 (their Fig. 3) agree closely with the trends  
 208 shown in Fig. 6 with Weber et al. using 1996.5 as the approximate  $T_A$  for all latitudes but does suggest  $T_A$   
 209 = 2000 for the polar regions. As mentioned earlier, the trend estimates are not very sensitive to the  
 210 exact  $T_A$ , but the shape of  $T_A(\theta)$  should be a marker contained in model calculations.



211 Because of the severity of the Antarctic ozone hole ozone before  $T_A$ , total column ozone decreased at a  
 212 rate of  $P_D = -10.9 \pm 3.6\%$  at  $77.5^\circ\text{S}$  and  $-8.0 \pm 1.1\%$  per decade at  $65^\circ\text{S}$ , near the edge of the nearly  
 213 continuous Spring polar vortex wind, during the period from 1979 to 1997 with smaller decreases at  
 214  $55^\circ\text{S}$  to  $25^\circ\text{S}$ . The seasonal breakup of the Antarctic polar vortex winds led to an inflow of ozone from  
 215 other SH latitudes contributing to depletion in most of the SH. In the NH the decreases were smaller  
 216 before  $T_A$  because of the absence of continuous polar vortex winds and an Arctic ozone hole. At  $77.5^\circ\text{N}$   
 217 was  $P_D = -5.6 \pm 4\%$ /decade and at  $65^\circ\text{N}$   $P_D = -4.4 \pm 0.35\%$ /decade After the turnaround period  $T_A$ , ozone at  
 218  $65^\circ\text{S}$  increased at  $P_D = 1.6 \pm 1.4\%$ /decade based on the MLR method. After  $T_A$ , most other latitudes (Fig.  
 219 6b) show stationary ozone amounts within  $2\sigma$ .

220 The Lowess(0.3) plots in Fig. 4 suggest that  $\Omega_{\text{MOD}}$  has been declining since approximately 2010 from  $5^\circ\text{S}$   
 221 to  $65^\circ\text{N}$  but still increasing from  $45^\circ\text{S}$  to  $65^\circ\text{S}$  (Fig. 7). However, computing the trends from either the  
 222 MLR or annual average methods suggest a small decline from  $15^\circ\text{S}$  to  $65^\circ\text{N}$  of  $1.5 \pm 2\%$  per decade that  
 223 is not significant at the  $2\sigma$  level.

224

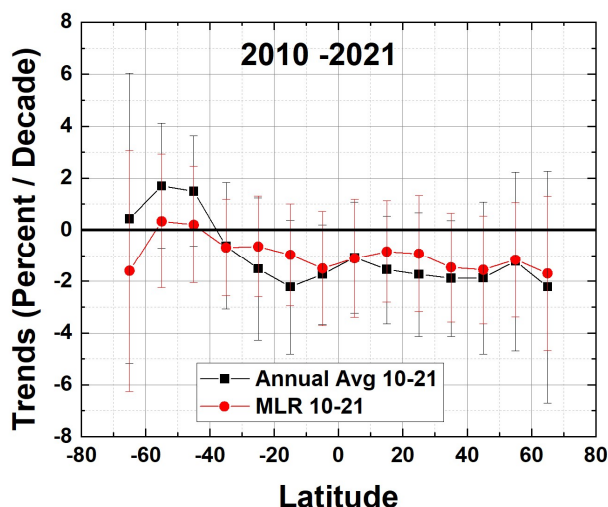


Fig. 7 Ozone trends  $P_D(\theta)$  (Percent per Decade) for the period 2010 – 2022 for the Annual Average and MLR methods.

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226 A comparison of deseasonalized  $\Omega_{\text{MOD}}(t, \theta)$  with deseasonalized Microwave Limb Sounder MLS (see  
 227 Appendix Figs. A1, A2, and A3) Stratospheric Ozone from 2005 to 2021 shows small average  
 228 (Lowess(0.3)) differences that are not statistically significant. For 2016 to 2018,  $\Omega_{\text{MOD}}$  is obtained from  
 229 NOAA-19 SBUV plus OMPS-NP and from OMPS-NP since 2018. The largest Lowess (0.3) differences,  
 230 deseasonalized  $\Omega_{\text{MOD}}$  – deseasonalized MLS, since 2016 are at  $65^\circ\text{S}$  of  $-4 \pm 9$  DU,  $5^\circ\text{S}$  of  $-1 \pm 2$  DU,  $65^\circ\text{N}$  of -  
 231  $2.5 \pm 7$  DU,  $35^\circ\text{N}$  of  $-1 \pm 2.6$  DU,  $25^\circ\text{N}$  of  $-0.5 \pm 2$  DU,  $15^\circ\text{N}$  of  $-0.5 \pm 2$  DU, and  $5^\circ\text{N}$  of  $-1 \pm 2$  DU. Other latitudes



232 55°S to 15°S and 55°N have no difference or are slightly positive. This suggests that the calibrations of  
233 the later SBUV-2 and OMPS-NP instruments are stable.

### 234 **3.0 Summary**

235 The monthly averaged Merged Ozone Data set  $\Omega_{\text{MOD}}$  (5° latitude bands, 77.5°S to 77.5°N) from 1979 to  
236 2021 were also averaged into 10° latitude bands  $75^\circ\text{S} < \theta < 75^\circ\text{N}$ . A smoothed  $\Omega_{\text{MOD}}$  version based on  
237 Lowess(0.3) was used to determine the approximate dates of the latitude dependent ozone turnaround  
238  $T_A(\theta)$  ranging from 1994 to 1998. The systematic latitude dependent pattern  $T_A(\theta)$  should appear in  
239 atmospheric models that combine the effects of photochemistry and dynamics in their estimate of  
240 ozone recovery. Trends of ozone  $P_D(\theta)$  in percent per decade were computed before and after the  
241 latitude dependent  $T_A(\theta)$  using two different methods, MLR and annual averages, that closely agree over  
242 their mutual latitude range of validity, 65°S to 65°N. The most dramatic rates of ozone loss were  $P_D =$   
243  $-10.9 \pm 3.6\%$  decade at 77.5°S and  $-8.0 \pm 1.1\%$ /decade at 65°S, which is about double the rate of loss of  $P_D$   
244  $= -5.7 \pm 4\%$ /decade at 77.5°N and  $-4.4 \pm 1.2\%$  per decade at 65°N. During the period after the start of  
245 recovery  $T_A$  to 2021, there has been a small increase at latitudes in the SH from 25°S to 65°S with the  
246 largest value being  $1.6 \pm 1.4\%$  per decade at 65°S. Aside from the small increases in the SH region there  
247 has been no statistically significant ozone recovery toward 1979 values, just an almost constant ozone  
248 amount after  $T_A(\theta)$ . During the period 2010 to 2021, there has been a small apparent decrease in ozone  
249 amount in  $\Omega_{\text{MOD}}$  that is not statistically significant at the 2-standard deviation level. A comparison  
250 between  $\Omega_{\text{MOD}}$  and MLS stratospheric column ozone shows some small differences both positive and  
251 negative after 2016 when the Lowess (0.3) smoothing of the total ozone shows the largest decrease.

252



253 **Appendix**

254 The MOD TCO data record since 2018 is obtained from OMPS-NP, which appears to show decreasing  
255 TCO (Fig. 4). Because of this, the deseasonalized  $\Omega_{MOD}$  are compared with MLS (Microwave Limb  
256 Sounder) deseasonalized stratospheric column ozone for the period 2004 to 2021 to look for calibration  
257 drifts in the  $\Omega_{MOD}$  time series. The question addressed here is not the absolute agreement between  $\Omega_{MOD}$   
258 and the MLS mostly stratospheric ozone column, but rather if there is a systematic drift between the  
259 two data sets after 2016. Figures A1 and A2 show that the difference between the two deseasonalized  
260 time series for latitudes from 65°S to 65°N and for the entire period 2005 – 2021. Of interest is the  
261 period 2016 to 2021 when  $\Omega_{MOD}$  was derived using NOAA-19 SBUV plus OMPS-NP 2016 – 2018 and from  
262 OMPS-NP since 2018.

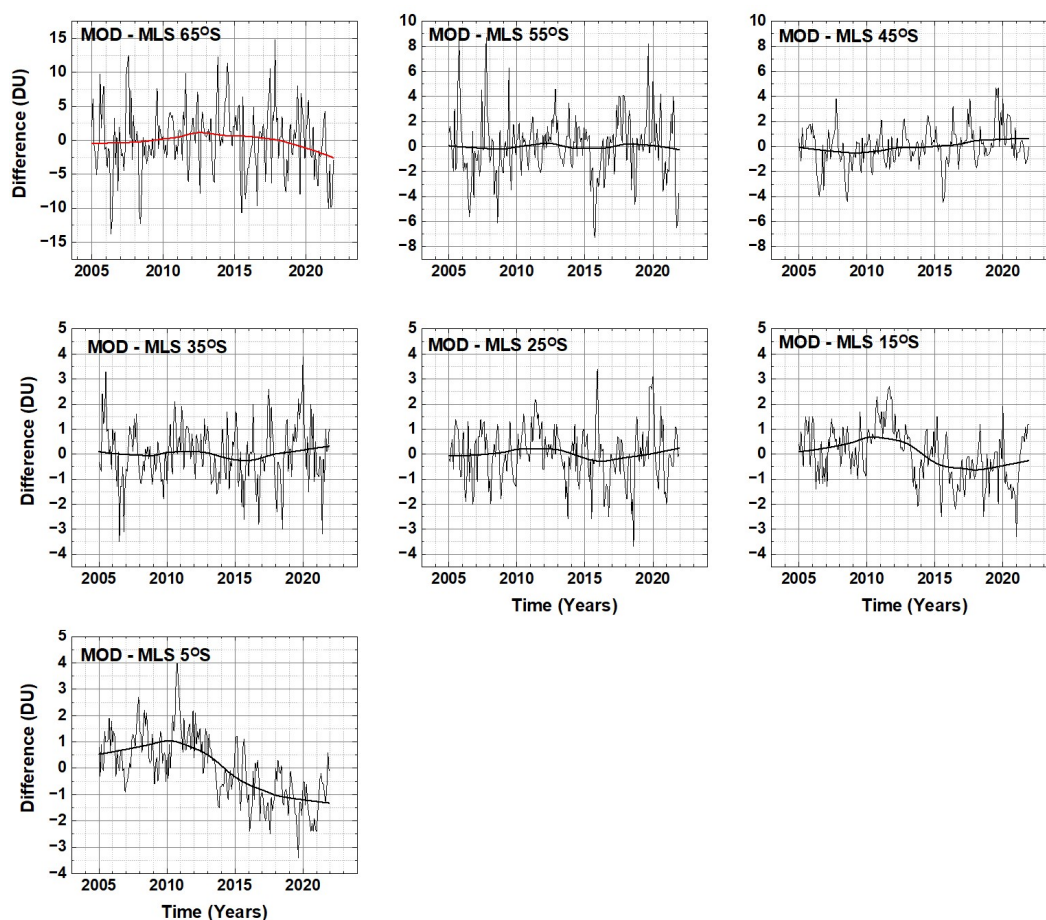


Fig. A1 A comparison of deseasonalized  $\Omega_{MOD}$  with deseasonalized MLS stratospheric column ozone for 65°S to 5°S.



264

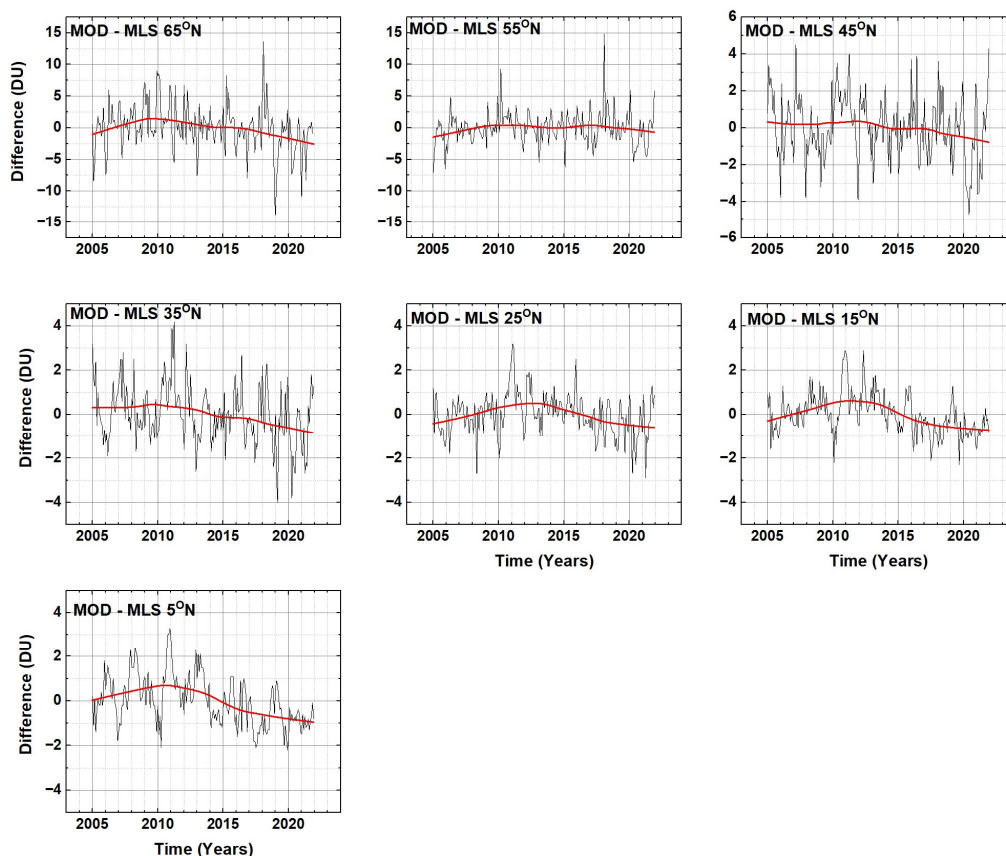


Fig. A2 A comparison of deseasonalized MOD total ozone with deseasonalized MLS stratospheric column ozone for 5°N to 65°N. Variations of  $\pm 3$ DU are within the MOD merged record uncertainties.

265

266 The differences in Figs A1 and A2 between  $\Omega_{\text{MOD}}$  and MLS since 2016 are not statistically significant at  
 267 the  $2\sigma$  level. Variations of  $\pm 3$ DU are within the  $\Omega_{\text{MOD}}$  merged record uncertainties.

268

269 The differences in 2021 average of Lowess(0.3) are summarized in Fig. A3 along with the  $2\sigma$  standard  
 270 deviations estimated from the average of the deseasonalized time series. One latitude shows a  
 271 statistically significantly difference, 5°N. The conclusion is that there is no statistically significant drift  
 272 between  $\Omega_{\text{MOD}}$  and MLS stratospheric ozone.

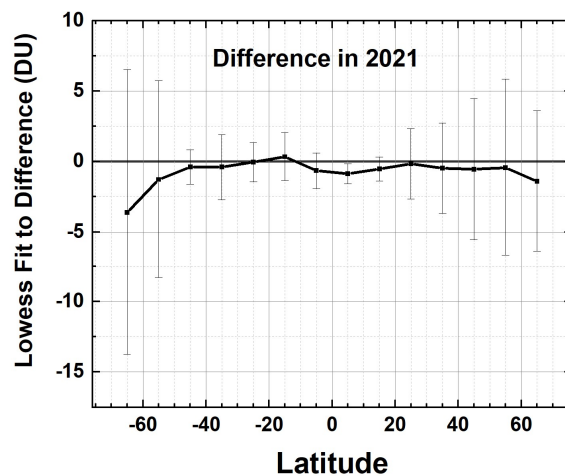


Fig. A3 The difference between the 2021 average Lowess (0.3) fit to  $\Omega_{\text{MOD}} - \text{MLS O}_3$  during 2021. The  $2\sigma$  standard deviations are estimated from the average of the deseasonalized time series during 2021 shown in Figs. A1 and A2.

273

274



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367

#### 368 **Author contribution:**

369 Jay Herman is responsible for writing the text, the annual integral trend calculations, and all the  
370 figures. Jerald Ziemke supplied the MLR trend calculations and the comparison with MLS. Richard  
371 McPeters supplied the MOD ozone as a continuous function of time from 1979 to 2021 for each  
372 latitude band.

#### 373 **Data Availability**

374 The original data used are publicly available in an ASCII format.

375 [https://acd-ext.gsfc.nasa.gov/Data\\_services/merged/](https://acd-ext.gsfc.nasa.gov/Data_services/merged/)

376 and processed data in Excel format

377 [https://avdc.gsfc.nasa.gov/pub/DSCOVr/JayHerman/MOD\\_Ozone\\_Trends/](https://avdc.gsfc.nasa.gov/pub/DSCOVr/JayHerman/MOD_Ozone_Trends/)

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#### 380 **Competing interests:**

381 The authors declare that they have no conflict of interest.

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385



386 **Acknowledgements:**

387 The authors want to acknowledge the contribution and help of Stacey Frith for compiling the SBUV  
388 and OMPS-NP data sets to produce the long ozone data record. She also reviewed the paper and  
389 added some important corrections.

390

391 **Figure Captions**

392 Fig. 1 Left: The zonally and monthly averaged  $\Omega_{\text{MOD}}$  data set 1979 – 2021 and  $-77.5^\circ$  to  $77.5^\circ$ . Right: Time  
393 and zonal averaged ozone and its maxima and minima 1979 – 2021. Error bars are 1 standard deviation  
394  $\pm 1\sigma$ .

395

396 Fig. 2 The ozone trend  $P_D(\theta)$  for the entire period 1979 – 2021 for two methods, MLR and Annual  
397 Average. The latitude grids for the two methods are offset to show the agreement in the trends and  $2\sigma$   
398 error bars.

399 Fig. 3  $\Omega_{\text{MOD}}$  in four latitude bands and Lowess(0.3) fitting functions ( $f = 0.3$ , black lines). Examples of  
400 different  $f = 0.1$  (Red) and  $0.05$  (blue dots) are shown at  $45^\circ\text{S}$  and  $45^\circ\text{N}$ . Note the slight downturn in the  
401 Lowess(0.3) at  $45^\circ\text{N}$  and  $55^\circ\text{N}$ .

402 Fig. 4 Lowess(0.3) fits to the  $\Omega_{\text{MOD}}$  data for 16 latitude bands Used to determine the turnaround dates.

403 Fig 5 Turnaround dates  $T_A(\theta)$  as a function of latitude from Fig.4

404 Fig. 6a Ozone trends  $P_D(\theta)$  (percent per decade) using the MLR and Annual Average methods before and  
405 after the turn-around times  $T_A(\theta)$  in Fig. 3. 6b. A magnified version of the MLR estimated trends after  $T_A$   
406 with  $2\sigma$  uncertainties.

407 Fig. 7 Ozone trends  $P_D(\theta)$  (Percent per Decade) for the period 2010 – 2022 for the Annual Average and  
408 MLR methods.