

# Surface dynamics and history of the calving cycle of the Astrolabe glacier (Adélie Coast, Antarctica) derived from satellite imagery

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**Abstract.** The recent calving of the Astrolabe glacier (Terre Adélie, East Antarctica) in November 2021 is an opportunity to better understand the processes leading to ice tongue fracturing. The archive of Sentinel-2 optical images is used to measure the ice motion and the ice strain rates for the period 2017-2021 in order to document fractures and rift evolution that lead to the calving. Additionally, the evolution of the Astrolabe ice tongue is mapped with satellite imagery from 1947 to November 2021 and used to understand the calving cycle of the Astrolabe ice tongue through time. These observations are compared with sea ice extent and concentration measurements. We found that a significant change in the sea ice periodicity surrounding the Astrolabe glacier occurred in the last decade (2011-2021) with respect to previous observations (1979-2011). Indeed, the duration of sea-ice free conditions significantly decreases after 2011 at the vicinity of the glacier and seems to have favored the ice tongue spatial extension. The analysis of strain rate time series revealed that the glacier dislocated suddenly in June 2021 in the middle of the winter before releasing an iceberg of about 20 km<sup>2</sup> in November 2021 at the onset of sea ice melting season. These observations suggest that sea-ice unbuttressing does not lead to instantaneous calving of the Astrolabe ice tongue, and that pre-existing opened fissures should first develop.

## 1 Introduction

Determining the contribution of polar ice sheets to sea level rise is a major concern for the society, and better understanding the processes and the factors controlling ice retreat is of paramount importance to simulate the ice-sheet response to global warming (Seroussi et al., 2020; Chambers et al., 2022). Coastal glaciers in polar regions differ from temperate mountain glaciers in terms of volume, catchment size and thermal state associated to complex interactions with the ocean. The presence of floating tongues with marine termini makes Antarctic glaciers more sensitive to the atmospheric and ocean dynamics (Gudmundsson et al., 2019; Olinger et al., 2019; Paolo et al., 2015; Pritchard et al., 2012; Christie et al., 2022). Monitoring of Antarctic glaciers remains heterogeneous Baumhoer et al. (2018). Studies focus either on continental scale monitoring, which usually lead to commenting the evolution of the largest glaciers of Antarctica (Walker et al., 2013; Rignot et al., 2019; Miles et al.,

2022; Millan et al., 2022; Baumhoer et al., 2023) or to specific glaciers or group of glaciers that concentrate most of the attention Baumhoer et al. (2018). In this study, we document and analyze the evolution of the Astrolabe glacier's ice tongue calving cycle, which has not been updated since Frezzotti and Polizzi (2002).

25 The Astrolabe glacier is located in Terre Adélie, (140°E, 67°S) near the Dumont d'Urville French research station. The glacier outlet is ca. 8 km wide (Figure 1a), while the drainage basin stretches as much as 200 kilometers inland. It is characterized by a tongue of ice developing on the water, presenting a calving front of 6 km wide (Figure 1a). Due to its proximity to the Dumont D'Urville research station, the glacier has been instrumented over the last decades with a focus on its grounding zone (Drouet, 2012; Le Meur et al., 2014). However, the last study documenting the calving cycle of the glacier's ice tongue  
30 covers the period 1940-2000 Frezzotti and Polizzi (2002) while recent observations show an unusual spatial extension of the ice tongue until November 2021 when a major calving event occurred (Figure 1f-i). Due to its small size, and rapid recent dynamic, the Astrolabe glacier ice tongue is not adequately monitored by global value-added products such as the NASA MEaSUREs ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7).

Ice calving is defined as the detachment of a smaller ice piece of ice from a larger one (Alley et al., 2023). Calving is mostly  
35 controlled by brittle processes (Alley et al., 2023) and results from the extensive opening of cracks or rifts within the ice shelf. Lateral spreading and thinning of the ice shelf can explain the formation and propagation of these fractures/rifts (Liu et al., 2015; Larour et al., 2021; Borstad et al., 2017; Alley et al., 2023). However, environmental forcing can also accelerate their propagation through hydrofracturing (Scambos et al., 2000), sub-glacial warm water intrusion and basal melting (Ritz et al., 2015; Rignot et al., 2019; Pritchard et al., 2012), bending of the ice due to flexural rebound after lake drainage (Banwell et al.,  
40 2013). Example of tsunamis that contributed to open rifts and trigger calving are also reported (Liang et al., 2023; Alley et al., 2023). Another important forcing is the influence of the sea-ice surrounding the ice tongues, which result from atmospheric and ocean dynamics (Campagne et al., 2015). Indeed, changes in atmospheric and ocean dynamics favoring the presence of sea-ice can act as a protection and allow glacier extension by either protecting the ice tongue from ocean swell and/or in the case of landfast sea-ice (i.e. sea-ice faster to the glacier/to the coastline) by buttressing the ice tongue (Massom et al., 2001; Walker  
45 et al., 2013; Robel, 2017; Wearing et al., 2020; Massom et al., 2010; Gomez-Fell et al., 2022; Massom et al., 2018; Wille et al., 2022; Christie et al., 2022). The thickness of sea-ice or ice mélange within a pre-existing rift may influence the acceleration of the rift opening (Larour et al., 2021), leading to complex calving cycles. In several cases, the disappearing of sea-ice around the ice tongue front has been reported to trigger the instantaneous calving (Massom et al., 2001, 2018; Robel, 2017; Wearing et al., 2020; Gomez-Fell et al., 2022; Xie et al., 2019). However, it remains unclear if the sea ice and in particular, landfast sea  
50 ice, is buttressing the ice tongue preventing fracture propagation or if it only holds the ice tongue parts together until calving is possible? All these processes are still poorly understood, as they exhibit strong spatial and temporal variability, which are highly difficult to document with direct observations in Antarctica.

In this study, we determine for the first time the ice tongue extension cycle of the Astrolabe glacier from high resolution satellite imagery (ERS, MODIS, Landsat, Sentinel-2 and ASTER) over the period 1947-2021. The archive of Sentinel-2 images  
55 is used to compute surface velocity of the ice for the entire area of the Astrolabe glacier from 2017 to 2021. We show the added-value of optical satellite images to monitor fractures propagation using ice velocity and strain rate calculated with optical image

correlation. We compare the front line evolution of the ice tongue with the sea-ice extent around the Astrolabe glacier from the NSIDC (National Snow and Ice Data Center; Fetterer and Windnagel2017). We show that calving events occur almost systematically when sea-ice extension decreases around the ice tongue terminus, but that the rift propagation can take place in the middle of the austral winter when the ice tongue is totally embedded in sea ice suggesting that sea ice buttressing may not be sufficient at the Astrolabe glacier to prevent calving.

## **2 Data and methods**

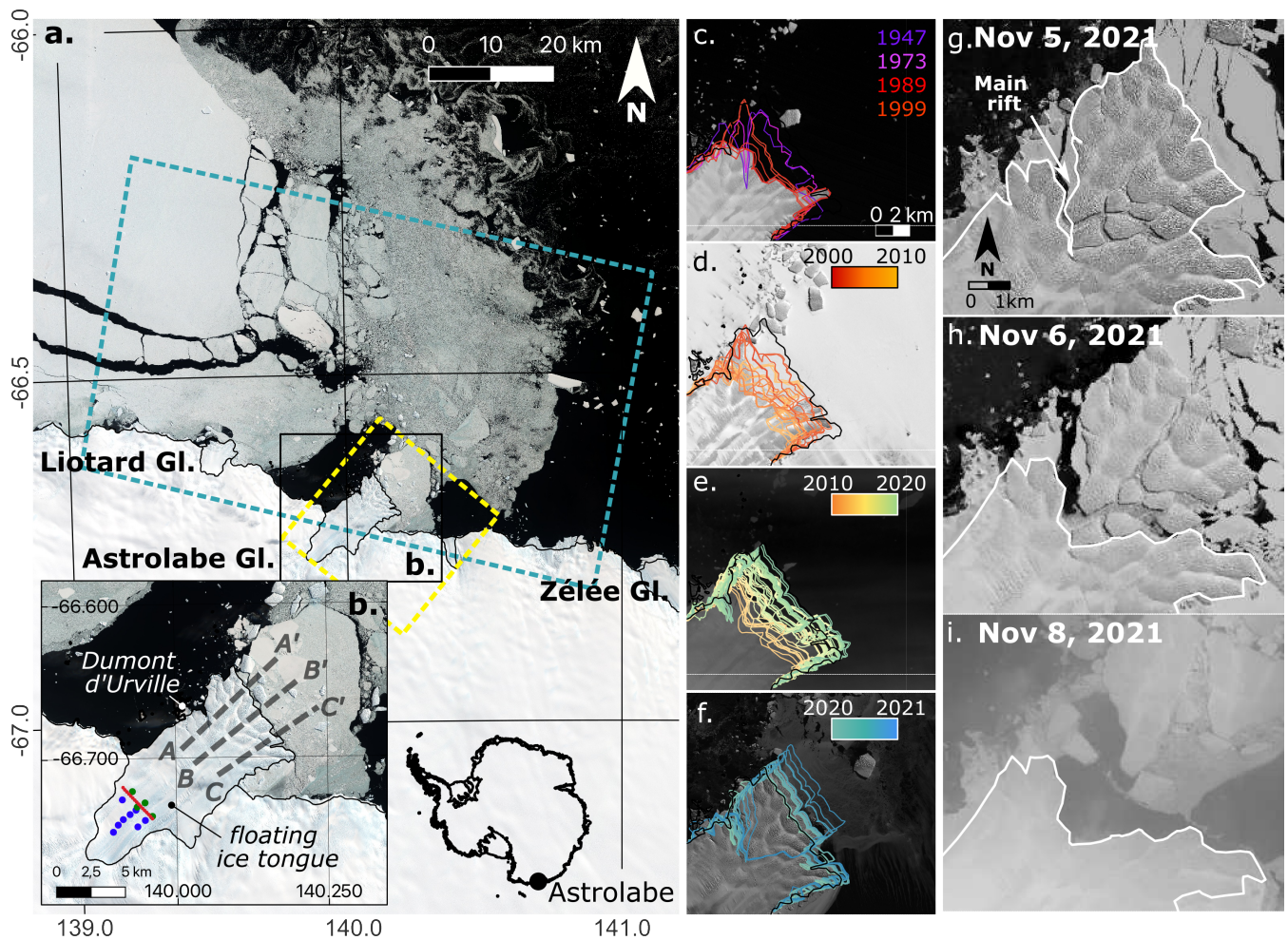
### **2.1 Satellite imagery**

#### **2.1.1 Mapping of the ice front position**

The ice front of the Astrolabe glacier was mapped using mainly optical satellite imagery at high resolution ( $< 50$  m) available in open access (i.e. Landsat, MODIS, ASTER and Sentinel-2). The first satellite image available has been acquired by Landsat-1 on Januray 29, 1973 (Figure 1c). The next available acquisitions are acquired in 1989 with Landsat-4/5 and then in 1999 with Landsat-4 (Figure 1c). In 1947, an aerial photography was taken of the Astrolabe. We used the sketch of the photography to extract the ice front position (Figure 1c) although important distortions are visible. In order to complete the optical dataset, we used radar acquisitions from ERS satellites and Radarsat RAMP product (Jezek and Barry., 2013) to map the ice front position between 1996 and 1999 (Figure 1c). From 2000 to 2010, Landsat and ASTER satellites provide around 1 to 3 images per year. We complete this dataset with the analysis of MODIS images (Figure 1d). From 2010, Landsat-8 and then, from 2017 Sentinel-2 provides regular acquisition at high resolution of the ice front evolution (Figure 1e). The combination of these two satellites allowed to monitor the calving of November 2021 with daily acquisition (Figure 1g-i). In total, 113 images are analyzed and the evolution of the ice front is mapped manually. Finally, the area of the floating tongue is estimated considering an arbitrary reference grounding line position (Bindschadler et al. 2011; Figure 1a). A more precise delineation of the grounding line has been proposed by Le Meur et al. (2014).

#### **2.1.2 Ice velocity monitoring from optical images**

Satellite imagery is commonly used to compute the ice velocity with image correlation techniques (Avouac et al., 2006; Lep-  
rince et al., 2007; Rignot et al., 2011; Mouginot et al., 2017; Millan et al., 2022). These techniques consist in matching pixels from one image to another to retrieve the shift in the position of a particular feature through time. Several studies have shown the interest of this technique to monitor ice surface velocity (Dehecq et al., 2015; Altena et al., 2019) especially in polar regions (Joughin et al., 2018; Millan et al., 2022). We used the GDM-OPT-ICE service (Provost et al., 2022; Stumpf et al., 2017) to compute ice displacement time series. The GDM-OPT-ICE service allows for the precise co-registration of the satellite imagery stack using the CO-REGIS algorithm (Stumpf et al., 2018), computes the shift between pairs of co-registered images with the open source stereo-photogrammetric library MicMac (Rosu et al., 2015; Rupnik et al., 2017) and inverts the displacement time series with the TIO algorithm (Doin et al., 2011; Bontemps et al., 2018).



**Figure 1.** a) Location of the Astrolabe glacier with the coastline and grounding lines from Gerrish (2022) and Sentinel-2 image of the February 7, 2020 in the background. The limits of the 4,000 km<sup>2</sup> box where the sea-ice extent is extracted is represented in dotted blue lines. Yellow dotted lines delineate the pixel extent and location of the sea-ice concentration grid from which the sea-ice concentration is extracted. The inset b) is a zoom over the Astrolabe glacier ice tongue and indicate the profiles where the evolution of the ice front position is presented (Figure 2). In-situ measurements are also represented: red dots for the location of the bamboo stakes, blue and green dots for the GNSS initial position of 2018 and 2021 campaigns respectively. Figures c) to f) show the ice front position at different dates. Figure g) to i) shows the calving of November, 6 2021 from the Sentinel-2 acquisition of November 5, 2021 (g), the Landsat-8 acquisition of November 6, 2021 (h), and the Sentinel-2 acquisition of November 8, 2021 (i).

The Copernicus Sentinel-2 mission provides acquisitions over the Astrolabe glacier every three to six days, during the austral summer (September to April). In total, from February 2017 to early November 2021, 59 Sentinel-2 images were acquired with no overcast over the Astrolabe glacier. The pairing network is set up to pair each image successively with the five next acquisitions, resulting in 280 pairs. The correlation is computed on a window of 5 by 5 pixels using sub-pixel refinement. The

displacement time series is inverted for each acquisition date with a spatial resolution of 1 by 1 pixel (i.e., 10 m x 10 m). The resulting displacement time series is interpolated at 30 days in order to compute the evolution of the ice velocity and reduce the noise.

### 95 2.1.3 Computation of the strain rates from the ice velocity fields

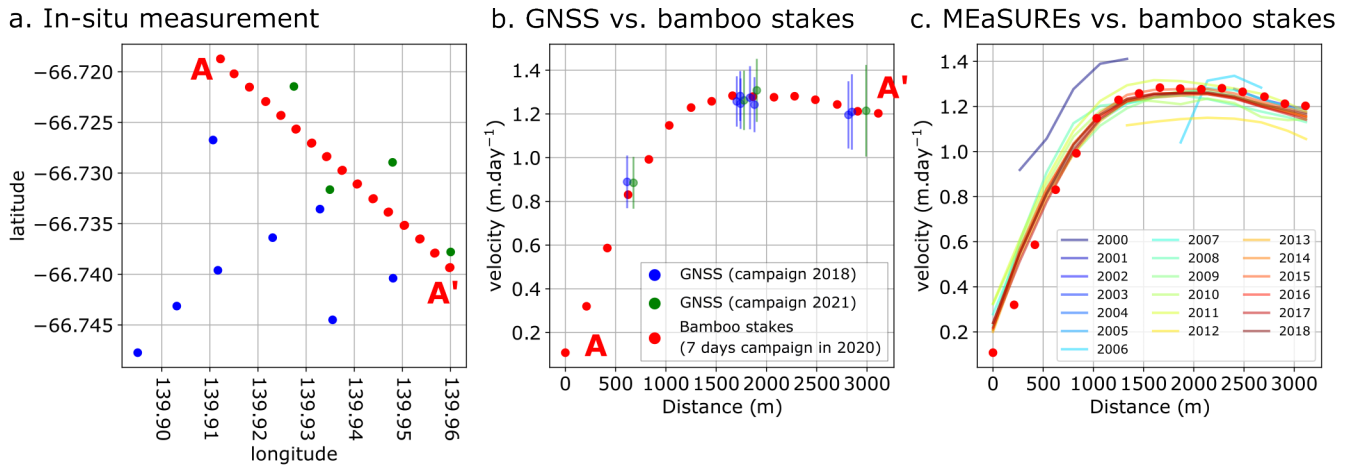
Strain is a measure of how much a medium (here ice) stretches, compresses and deforms in all directions as it flows, whereas strain rates represent how quickly these deformations occur. The strain rates can therefore be computed using satellite-derived velocity (Alley et al., 2018; Cheng et al., 2021). We used the method described in Alley et al. (2018) and Nye (1959) to compute the longitudinal, transverse and shear strain rates using the yearly estimation of the ice velocity derived from the  
100 GDM-OPT-ICE outputs (see section 3.2.1). The strain rates are computed at a spatial resolution of 20 meters.

## 2.2 In situ sensors

### 2.2.1 On-site GNSS observations and displacement measurements

A permanent GNSS network (<https://astrolabe.osug.fr/>) is maintained by the Institut des Géosciences de l'Environnement (IGE) on the Astrolabe glacier. It consists of 8 GNSS stations in 2018 and 4 stations in 2021 (mainly because of a lack of maintenance  
105 in 2019/2020 due to cancellation of the summer operations in Antarctica because of the COVID pandemic; Figure 1b). The GNSS receivers and antennas are mounted on beacons specifically designed to withstand harsh environmental conditions (strong winds, local wind-drifted accumulation of snow, ice motion, summer melting leading to beacon tilting or even collapse). These harsh conditions explain some gaps in the GNSS time series, mainly during the austral winters. The receivers are geodetic dual-frequency receivers (Trimble™NetR9) connected to Zephyr geodetic antennas. The GNSS observations consist  
110 of 3 two-hour measurement sessions per day, where positions are averaged from 30-s sampling measurements. The positions are calculated for 24h measurements in PPP mode (Precise Point Positioning) using the GipsyX geodetic software (Zumberge et al., 1997). The accuracy is 1.5 cm (standard deviation 0.9 cm) in the horizontal component and 3.8 cm (standard deviation 2.7 cm) in the vertical component.

A field campaign was conducted in 2020 to quantify the ice velocity in the vicinity of the grounding line position. It consisted  
115 of 16 bamboo stakes that were implanted in the ice during winter 2020 for one week between January 31, 2020 and February 7, 2020 (Figure 1b). The position of the stakes was measured the first day and then, one week later, with a GNSS dual-frequency receiver, allowing for an estimation of the ice velocity. The derived velocity is compared to the GNSS's velocity derived from the 2018 and 2021 campaigns and to the NASA MEaSUREs ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7) available in this part of the ice tongue. The result shows a good agreement between all dataset (Figure 2) meaning that the velocity remains locally  
120 constant through the years from 2000 to 2018 in this part of the ice tongue.

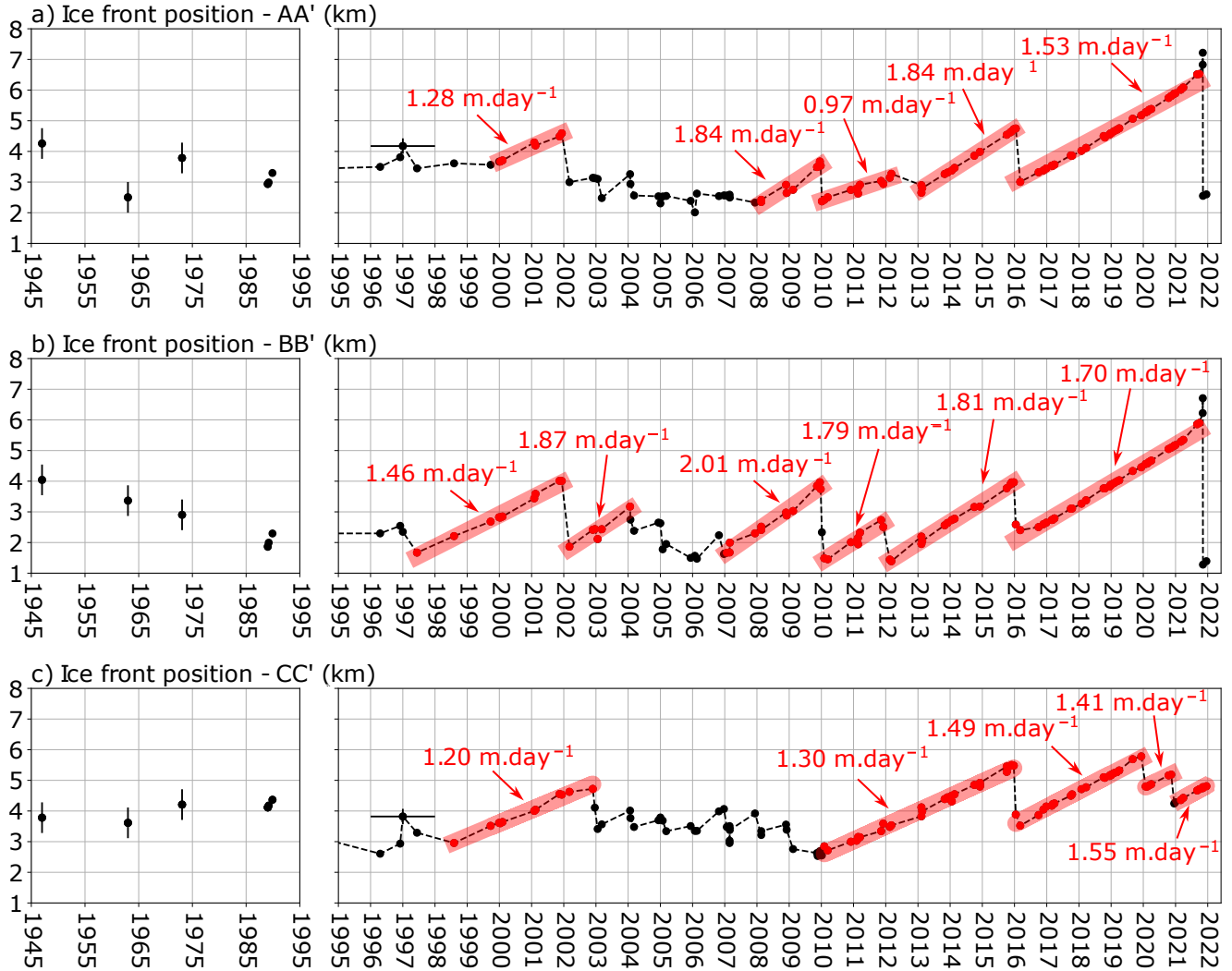


**Figure 2.** Comparison between yearly velocity measured with in-situ measurements from year-long GNSS campaigns and week-long bamboo stakes campaign: a) location of the GNSS's and bamboo stakes (see Figure 1 for the location on the ice tongue); b) comparison between the different in-situ dataset. c) comparison between the estimation of the velocity from the bamboo stake campaign and the yearly estimation of the velocity from satellite imagery from the NASA MEaSURES ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7).

### 3 Results

#### 3.1 Ice front position: 2000-2021

Changes in the frontal position are presented in Figure 3. The evolution of the ice front position varies from one profile to another. Between 1945 and 1995, historical images are sparse but show a maximal position of 4.2, 4.0 and 4.1 km for profiles AA', BB' and CC' respectively. In 2016 and 2019, the ice front terminus reached this maximal position simultaneously for all three profiles. From 2019 to 2021, an unprecedentedly observed position of 7.2 km and 6.7 km is reached for profiles AA' and BB' respectively (Figure 3a, b). Conversely, on profile CC' the ice front position decreases progressively after 2020 due to successive calving events (Figure 3c). It should be noted that the central profile BB' reaches regularly its maximum position before the calving events of 2002 and 2010 (Figure 3b) while on profile AA' and CC', the maximum position is only reached in 2002 or late 2002 (Figure 3a). From 2002 to 2010, the ice front position experience periods of yearly calving of different lengths depending on the considered profiles: 2002-2008 for profile AA', 2004-2007 for profile BB' and 2003-2010 for profile CC'. A linear regression is performed to retrieve the velocity of the ice front progression in between the successive calving events (Figure 3). The velocity varies highly from one period to another, but one can observe that the velocity are significantly lower for profile CC' ( $1.17 \text{ m.day}^{-1}$ - $1.55 \text{ m.day}^{-1}$ ) than for profile AA' ( $0.96 \text{ m.day}^{-1}$ - $1.79 \text{ m.day}^{-1}$ ) and BB' ( $1.75 \text{ m.day}^{-1}$ - $2.12 \text{ m.day}^{-1}$ ).

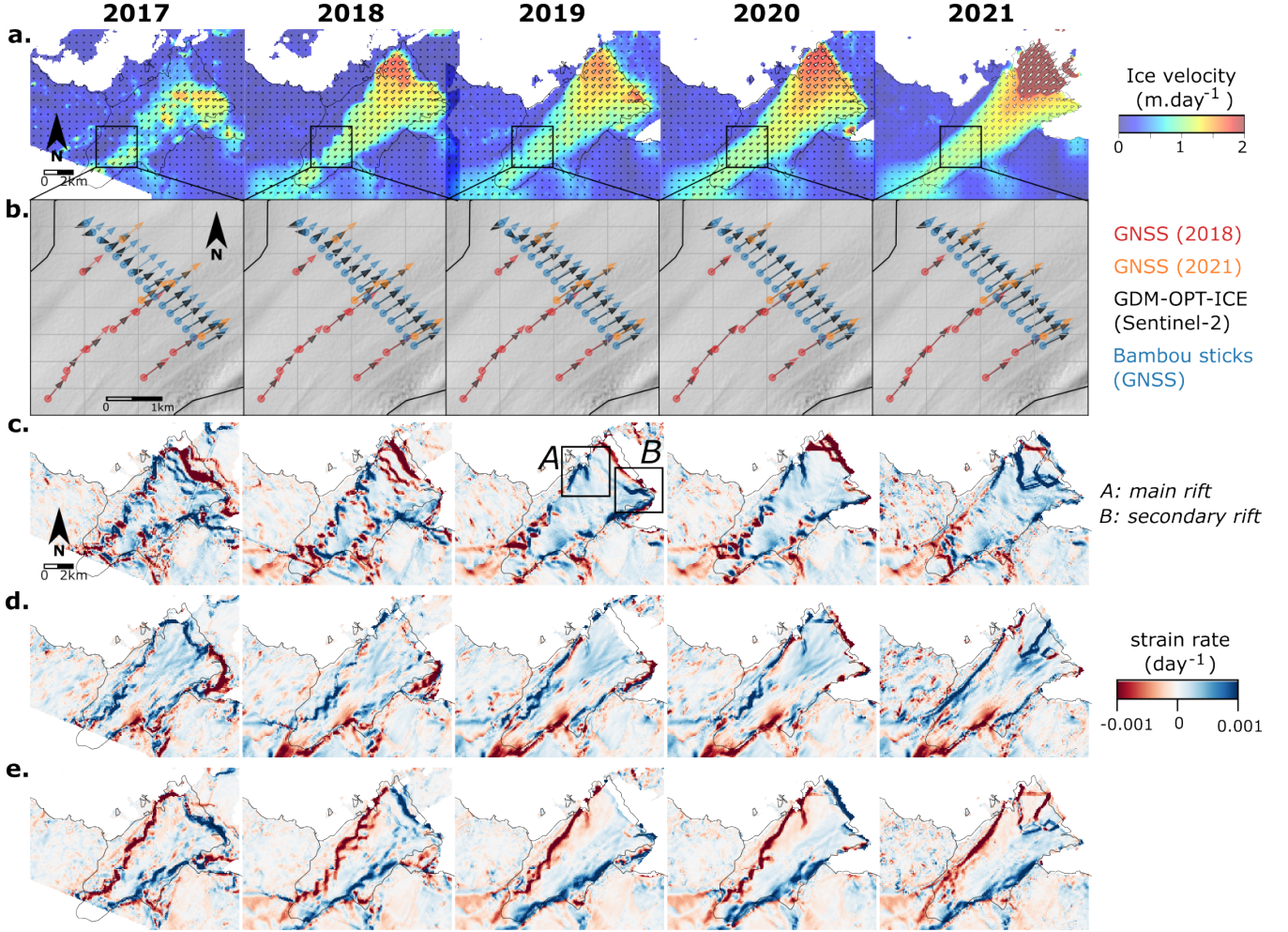


**Figure 3.** Evolution of the front position of the glacier terminus along profiles AA' (a), BB' (b) and CC' (c); (see Figure 1b for the location of the profiles). The velocity of the ice front motion is indicated for the periods of ice front progression.

### 3.2 Ice velocity: 2017-2021

Ice velocity is plotted for each year from 2017 to 2021 (Figure 4a) together with the derived longitudinal, transverse and shear strain rates (Figure 4c, d, e respectively). The yearly estimation obtained with GDM-OPT-ICE is compared to the one measured with in situ instrumentation (GNSS's and bamboo stakes campaigns). The in-situ data shows that the velocity in this part of the glacier is very constant though time (Figure 2b, c) allowing for comparison between different years. Figure 4b presents the ice velocity computed with GDM-OPT-ICE and the velocity measured with in situ instrumentation (i.e., GNSS's and bamboo

stakes campaign). The comparison shows that the estimation of the velocity from GDM-OPT-ICE improves with time, with a poor accuracy in 2017 ( $\text{RMS} = 0.76 \text{ m.day}^{-1}$ ) and a much better one from 2019 ( $\text{RMS} < 0.25 \text{ m.day}^{-1}$ ). One can observe that the gradient of velocity from the western border to the center of the glacier is well retrieved with the GDM-OPT-ICE products of 2019-2021 (Figure 3b) while in 2017 and 2018, the limit between stable ice and the flowing ice tongue is retrieved in the wrong position with the GDM-OPT-ICE products. The small number of cloudless Sentinel-2 acquisitions for those years may explain the low RMS error of these two years, as well as the wrong estimation of the ice tongue limit.



**Figure 4.** a) Yearly estimation of ice velocity for the Astrolabe glacier for the period 2017-2021, b) comparison of the velocity magnitude and direction as measured by the in situ instrumentation (GNSS's and bamboo sticks campaign) and as measured by GDM-OPT-ICE. Figures c, d, e present the longitudinal, transversal and shear strain rates derived from the ice velocity fields.

The velocity field shows a smooth gradient with lower velocity of circa  $1 \text{ m.day}^{-1}$  in the south-eastern part of the glacier and faster velocity of  $1.2\text{-}1.5 \text{ m.day}^{-1}$  in the north-western part of the glacier. In 2019, a small block of ice accelerated in the

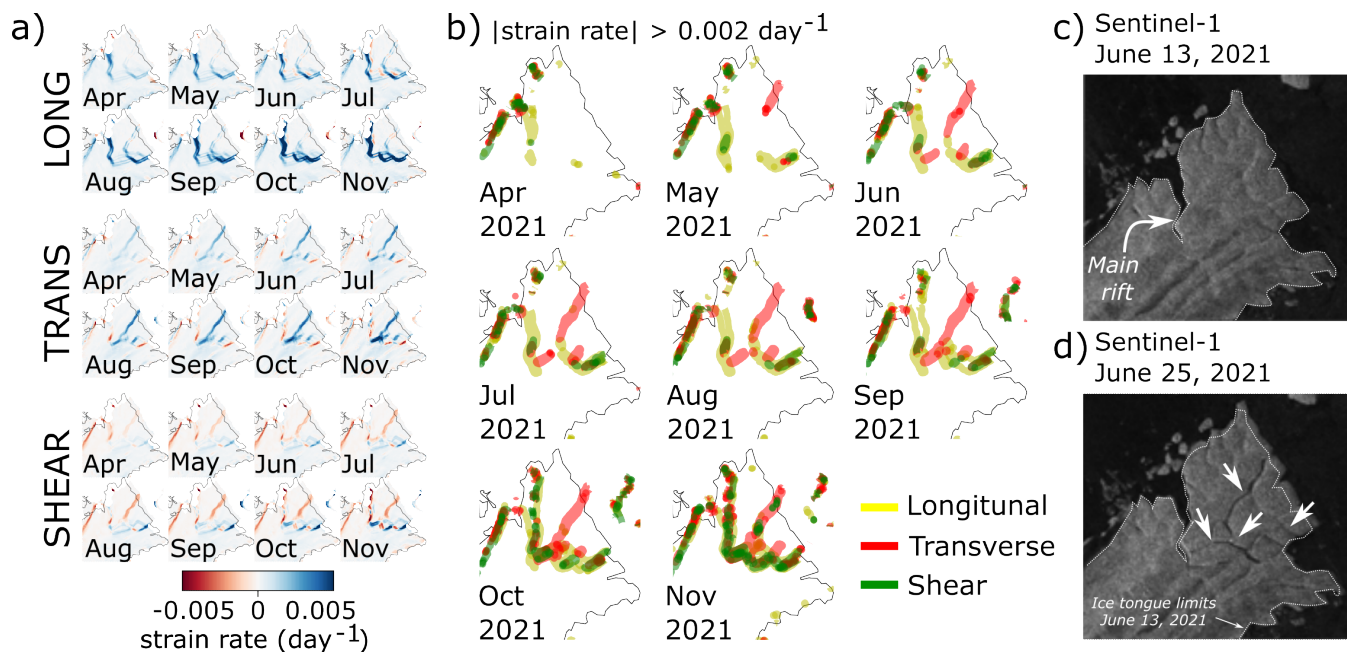
150 eastern part of the ice front (Figure 4a), also visible in the longitudinal strain rate field (Figure 4c, box B). This block disappears in both the mean velocity of 2020 (Figure 4a) and in the strain rate field (Figure 4c) due to the calving of this part of the glacier in December 2020 (Figure 3d, profile CC'). In 2019, an extensive fracture appeared in the western part of the ice tongue in front of the Dumont D'Urville station, clearly visible in the longitudinal and shear components of the strain rate fields (Figure 4c, e; box A). The northern-western part of the ice tongue starts simultaneously to exhibit larger velocities in 2020 and 2021  
155 (Figure 4a). In 2021, a complex network of localized increase of strain rates appears on the western-northern part of the glacier delimiting the potential area of the future iceberg to be calved (Figure 4c, d, e). This complex network delimits the fractures that were observed on the ice in the first available summer acquisition in September 2021 and that remained the same until the ice calving (Figure 1f). Beside the evolution of the fractures on the ice, one can also observe the high strain rates ( $> 0.002 \text{ day}^{-1}$ ) that are clearly identifiable through time along the lateral limits of the glacier (Figure 4e).

### 160 3.3 Ice tongue break off: 2021

The displacement time series is linearly interpolated over a time step of 30 days from the first acquisition in 2017 to November 5, 2021. Mean monthly velocity and strain rate fields are derived from this interpolation. We investigate the evolution of the strain rates for the period January and November 2021 to understand the dynamic of the recent calving (Figure 5). Strain rates maps show high concentration of strain localized along linear structures, which grow progressively from April 2021 to  
165 November 2021. We set a threshold on the strain rates in order to analyze the evolution of these localized concentration of strain, as well as the occurrence of the spatial connection between them (Figure 5a). The evolution of their growth is complex, with transitions from one component to another. For example, the main rift exhibits a clear longitudinal strain rate from April 2021 to September 2021 that evolves toward a shear strain rate in October-November 2021 (Figure 5a). From May 2021, a large concentration of strain appears in the transverse component along another fracture oriented in the North-East/South-West  
170 direction (Figure 5a). Similarly, a third fracture appears in the longitudinal component on the eastern side (Figure 5a). These fractures grow rapidly and connect together in June 2021 (Figure 5a). One can also observe that from October 2021, most of the fractures exhibit a shear strain rate, likely due to the rotation of the blocks. We analyzed the Sentinel-1 SAR images from May 2021 to August 2021 in order to validate these observations. We observe that the network of fracture opened suddenly between June 13, 2021, and June 25, 2021 (Figure 5b, c) which is coherent with the timing of the connection derived from the  
175 time series of strain rate (Figure 5a). It can be noted that compressional strain rates are measured from 2017 to 2020 at the terminus of the glacier tongue with strain rate larger than  $0.001 \text{ day}^{-1}$  while it is not observed anymore in 2021 (Figure 4c).

### 3.4 Sea-ice forcing

We analyze time series of sea-ice extent and concentration in the region of the Astrolabe glacier (Figure 1a). The data for sea-ice extent and concentration are downloaded on the NSDIC repository (Fetterer and Windnagel, 2017) and cover the periods  
180 from 1979 to the end of 2021. We cropped the data to analyze the monthly variation of the sea-ice extend over an area of  $4000 \text{ km}^2$  around the Astrolabe glacier (dotted blue lines on Figure 1a). This area is chosen arbitrary to represent the influence of regional sea ice variation on the Astrolabe ice tongue. The daily variation of the sea ice concentration is taken at the pixel (25



**Figure 5.** a) map of the strain rates larger than  $0.002 \text{ day}^{-1}$ . The three strain rate components (longitudinal, transverse and shear) are plotted together with different color. Subsets c) and d) are showing the occurrence of the fractures detected with Sentinel-1 acquisitions of June 13 and June 25, 2021. Arrows indicate the location of the main rift (b) and of the secondary fractures (c).

km x 25 km) encompassing the Astrolabe glacier (dotted yellow lines on Figure 1a) to focus on the conditions at the Astrolabe ice tongue.

185 We observe a significant change in the periodicity of sea-ice around the Astrolabe glacier in the last decade (2011-2021). From 1979 to 2011, the extent of the sea-ice decreases significantly every year during the summer (Figure 6a). From 2011 to 2021, the annual disappearing of sea ice does not occur every year (Figure 6a). Indeed, during two successive periods: 2012-2016 and 2016-2021, the extent of the sea-ice remains maximal during summer (Figure 6a). In detail, one can see that during those two periods, the sea-ice extent can drop occasionally (e.g., early 2015, 2018) or during larger periods such as for

190 austral summer 2018-2019 (Figure 6a). However, the reduced length or absence of periods of sea ice free conditions during 2012-2016 and 2016-2021 is notably different from the previous decades. The time series of daily sea ice concentration shows similar observations (Figure 6b). Before 2011, the sea ice concentration drop below 15% for periods of 2 to 3-4 months from November to mid-March, with small variations on the length of sea-ice free periods (Figure 6b). From 2008 to 2011, one can observe a decrease in the duration of low sea ice concentration to 2 months (Figure 6b, d), corresponding to a shift of the free

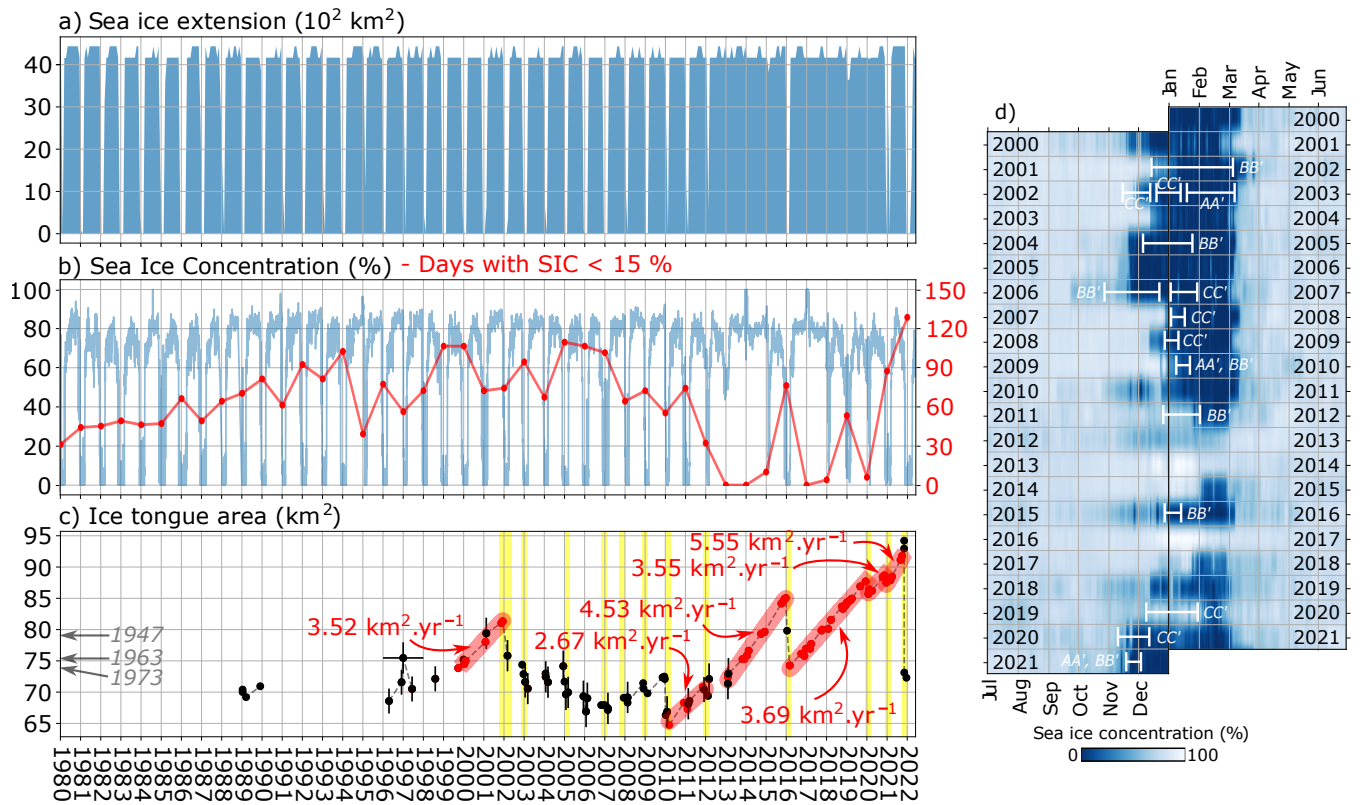
195 sea ice period onset from November to mid-December/January, while the end of the sea-ice free periods remain stable through time : beginning to mid-March. From 2012 to 2016, the sea ice free periods disappear or are shortened to less than one month (i.e. February 2015; Figure 6d). Between 2016 and 2021, the regime of sea ice concentration is highly variable, with years with no to very short periods of sea ice free conditions (austral summer 2016-2017, February 2020; Figure 6b, d) and years with

prolonged sea ice free conditions (December 2018-March 2019). From austral summer 2020-2021, the duration of sea ice free conditions seems to resume as before 2012 with a duration of 3-4 months from mid-November to March (Figure 6b, d).

We compare the evolution of the sea ice extent and concentration to the evolution of the ice tongue area (Figure 6c). We observe that 2012-2016 and 2016-2021 periods corresponds to periods of significant extension of the Astrolabe ice tongue (Figure 6c) with an increase of  $15 \text{ km}^2$  between 2012 and 2016 and of almost  $20 \text{ km}^2$  between 2016 and 2021. For the period 2002-2012, the ice tongue extension is much more limited due to the regular calving at different location of the ice front (Figure 3). Before 2002, the satellite images are scarcer, but the ice tongue seems to have reach a rather advance position in 2002 with an area of almost  $81 \text{ km}^2$ . This advance can not be linked to significant variation in the sea ice seasonal cycle. We report the calving event that can be observed with the analyzed satellite images (Figure 6d) with the incertitude on the date of the different calving events. One can observe that all detected calving occur when sea-ice concentration decreases at the end of the Austral fall (Figure 5d) except for austral summer 2006-2007 where multiple calving events are reported and do not occur necessarily at the onset of the sea ice concentration decrease.

## 4 Discussion

To understand the recent evolution of the Astrolabe glacier, we investigated the evolution of sea-ice in the vicinity of the Astrolabe ice tongue. Sea-ice and in particular, landfast sea-ice evolution is usually assumed to delay ice tongue break off and favor its extension by buttressing the ice tongues and protecting them from ocean swells (Massom et al., 2010, 2018; Rott et al., 2018; Gomez-Fell et al., 2022). At the Astrolabe glacier, we observe a significant change in the periodicity of sea-ice in the recent decade (2011-2021) in comparison with the previous observations (1979-2011; Figure 6). The recent periods of multi-year sea-ice presence are well correlated with the ice tongue spatial extension (Figure 6) and seems to validate the assumption that sea-ice protects the ice tongue and favors its extension. Moreover, the disappearing of such protection has been reported to initiate rift propagation which, in some cases, lead to rapid calving (Miles et al., 2017; Cassotto et al., 2021; Gomez-Fell et al., 2022; Christie et al., 2022). In the case of the Astrolabe glacier, we also observe that the calving, when it occurs, systematically takes place at the beginning of the sea-ice disappearing (Figure 6d) which seems to confirm the potential of sea-ice disappearing as a triggering factor of calving. However, the analysis of satellite images from 2017 to 2021 at the Astrolabe ice tongue shows that the rifts or crevasses that lead to the 2021 calving event are forming several months to years before calving (Figure 4, 5) suggesting a different mechanism. The presence of rift and fracture networks in ice tongue several years to several months prior to calving has been reported in other glaciers (Fricker et al., 2005; Walker et al., 2013, 2015; Cheng et al., 2021; Larour et al., 2021; Gomez-Fell et al., 2022). In most cases, the growth of the rifts or their (re-)activation are observed during the austral summer (Fricker et al., 2005; Walker et al., 2013; Cheng et al., 2021; Gomez-Fell et al., 2022) and few rifts propagation are reported during Austral winter (Walker et al., 2013; Larour et al., 2021). At the Astrolabe glacier's ice tongue, the main rift is located in front of the Dumont D'Urville Research station and initiated in 2019, a year with almost two consecutive month of low sea ice concentration. The absence of a significant rift in 2017 and 2018 suggests that sea ice may have had an effect in inhibited rift growth on the Astrolabe ice tongue and delayed the calving in 2017-2018 and, possibly, in



**Figure 6.** Evolution of monthly sea ice extent (a), daily sea ice concentration and number of days with sea ice concentration lower than 15% (b). Monthly sea ice extent is computed for a wide region of  $4000 \text{ km}^2$  around the Astrolabe glacier (dotted blue line in Figure 1a) while the daily sea ice concentration is taken for the pixel of  $25 \text{ km}$  by  $25 \text{ km}$  at the Astrolabe ice tongue (dotted blue line in Figure 1a) and both are extracted from Fetterer and Windnagel (2017). The evolution of the ice tongue area is presented in (c) with in red the period of extension and the ice tongue growth speed. Calving event that could be observed with satellite imagery are plotted in yellow. In years 2003-2005, calving likely occurred, although no observation can confirm the date. Figure d) also presents the evolution of the daily sea ice concentration from 2000 to 2021 and calving events are reported.

2012-2016. We note that in 2019, no calving occurred despite the long period of low concentration of sea ice while in 2020, the eastern part of the ice tongue calved at the onset of the short (one month) period of sea ice concentration decrease (Figure 6d). These observations suggest that sea ice acts like a glue to hold together the ice tongue, and that deep opened rifts must pre-exist for calving to occur.

In June 2021, we observe the opening of a complex network of fractures suddenly in the middle of Austral winter (Figure 5). Larour et al. (2021) proposed a mechanism to explain the winter propagation at the Larsen C Ice Shelf, Antarctica, prior to the calving of iceberg A68 based on the critical thinning of the ice shelf and of the ice mélange within the rifts. Here, the critical thinning of the ice tongue due to its exceptional extension may explain this timing (Robel, 2017; Larour et al., 2021; Åström and Benn, 2019) although it would likely favor the propagation of the rift along the same direction as the pre-

existing rift, which is not observed at the Astrolabe glacier (Figure 5). Instead, the main fracture propagating in June 2021 is oriented along the flow direction and opened in extension (Figure 5a, b). Another possibility to explain the development of these fractures could be a transition from a ductile to a brittle behavior with the decrease of temperature during winter that may favor fractures along the flow resulting from the differential compressive load due to sea-ice buttressing and rift opening (Figure 4). This mechanism might be possible as the compressive longitudinal strain seems to disappear in 2021 at the glacier terminus (Figure 4c) and the rift opens progressively. However, such a scenario remain to be validated as it would maintain the ice tongue terminus at the same position due to the effect of sea-ice buttressing, which is not what is observed (Figure 4c, d) and because the compressive strength of the ice is much higher than extensive strength (Benn et al., 2007). The presence of extensive circumferential stress that appear when the unconfined part of the ice tongue reaches a certain extension (Wearing et al., 2020) should be also considered, as well as the presence of basal channels and basal melt that may play a role in the dislocation of the ice tongue (Vaughan et al., 2012; Alley et al., 2023). The difference in the calving cycle and ice velocity between the eastern part and the western part of the glacier terminus also suggest that the bathymetry underneath the ice tongue controls the location and evolution of the rifts. Our analysis remains limited, and further modelling is necessary to understand the mechanisms that lead to the apparition of these fractures at this time of the year (Åström and Benn, 2019; Crawford et al., 2021; Alley et al., 2023).

Sea-ice is deeply connecting to regional and local atmospheric and oceanic states (Fogt et al., 2022). At the scale of the continent, records in Antarctica show a positive increase of the sea ice extent from 1979 to 2016 with a minimum of global sea-ice extent recorded in summer 2017 (Fogt et al., 2022). In the region of the Astrolabe, encompassing the Adélie Coast and George V Land, the calving of the Mertz Ice tongue in 2010 (Massom et al., 2018) lead to severe modifications of the sea ice production and location, traducing regional changes in the oceanic and atmospheric currents (Campagne et al., 2015). This event is likely at the origin of the transition of the sea ice seasonal cycle at the Astrolabe glacier (Figure 6). Miles et al. (2022) reports similar observations further west on the Adélie Coast, with the continuous growth of the Commandant glacier (Adélie Coast) from 2010 to 2018 due to the presence of persistent sea ice. This illustrates how one calving event such as the Mertz Ice tongue calving in 2010 may significantly modify the calving cycle of neighboring ice tongues, which are difficult to account in current models (Edwards et al., 2021; Miles et al., 2022). The extent of the regional impact of the Mertz Ice tongue 2010 calving is not clearly known, as most studies focus on the Georges V land area (Kusahara et al., 2011, 2017; Campagne et al., 2015; Cougnon et al., 2017). Moreover, the evolution of the ice shelves of the Adélie coast and Georges V land remains limited (Frezzotti et al., 1998; Frezzotti and Polizzi, 2002), preventing a better understanding of the environmental forcing Massom et al. (2018); Christie et al. (2022).

## 270 5 Conclusions

In this study, we analyzed the evolution of the Astrolabe glacier located in Terre Adélie/Adélie Coast, Antarctica. We used open access optical satellite imagery (MODIS, ASTER, Landsat and Sentinel-2) completed by ERS and RADARSAT images to map the evolution of the ice front from 1947 to 2022. We also measure the surface velocity and derived strain rate fields between

2017 and 2022, using image correlation of Sentinel-2 images. The recent evolution of the glacier shows an unprecedentedly  
275 documented extension of 95 km<sup>2</sup> favored by the concomitant high concentration of the sea ice in the region during 2011-2021  
in comparison with previous records (2000-2011). The early melt of the sea ice in November 2021 favored the released of a 20  
km<sup>2</sup> iceberg in the north-western part of the Astrolabe glacier. This is the first time a calving of this magnitude is documented  
at the Astrolabe glacier. We also observed that a complex network of fractures opened during the austral winter in June 2021  
several months before the iceberg calving. This study shows the importance of ice velocity and strain rates fields time series  
280 derived from optical satellite imagery at high resolution to document fracture opening and raises further questions on the  
mechanism of rift propagation.

*Data availability.* We acknowledge the use of imagery from Copernicus Sentinel-1 and 2 data and NASA ASTER and Landsat images  
(through <https://earthexplorer.usgs.gov/>). The GNSS observations are accessible on the Astrolabe repository: <https://astrolabe.osug.fr/>.

*Author contributions.* FP designed the experiments with contributions from DZ, ELM, JPM and CH. ELM provided the GNSS data and  
285 JPM processed them. FP processed the satellite data. All co-authors participated in the writing and/or revision and approval of the submitted  
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