

# Surface dynamics and history of the calving cycle of the Astrolabe glacier (Adélie Coast, Antarctica) derived from **optical-satellite** imagery

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**Abstract.** The recent calving of the Astrolabe glacier (Terre Adélie, East Antarctica) in November 2021 ~~presents is~~ an opportunity to better understand the processes leading to ice ~~fracturing. Optical satellite imagery is used to retrieve the calving cycle of the glacier since 2000 by mapping the ice front location. A recent archive of high-resolution optical images from Sentinel-2 tongue fracturing. The archive of Sentinel-2 optical images~~ is used to measure the ice motion and the ice strain rates for the period 2017-2021 in order to document fractures and rift evolution ~~-that lead to the calving. Additionally, the evolution of the Astrolabe ice tongue is mapped with satellite imagery from 1947 to November 2021 and used to understand the calving cycle of the Astrolabe ice tongue through time.~~ These observations are compared with sea ice extent and concentration measurements. We found that a significant change in the sea ice ~~melting periodicity at the vicinity of the periodicity surrounding the~~ Astrolabe glacier occurred in the last decade (2011-2021) with respect to previous observations (1979-2011). ~~After 2011, the occurrence of consecutive years of high~~ ~~Indeed, the duration of~~ sea-ice ~~concentration free conditions significantly decreases after 2011~~ at the vicinity of the glacier ~~and~~ seems to have favored the ice tongue spatial extension. ~~This lead to an unprecedentedly observed extension of the ice tongue until November 2021.~~ The analysis of strain rate time series revealed that the glacier dislocated suddenly in June 2021 in the middle of the winter before releasing an iceberg of ~~around about~~ 20 km<sup>2</sup> in November 2021 at the onset of sea ice melting season. These observations suggest that ~~although the presence of sea ice favors glacier extension, its buttressing effect may not be sufficient to prevent fracture opening~~ ~~sea-ice unbuttressing does not lead to instantaneous calving of the Astrolabe ice tongue, and that pre-existing opened fissures should first develop.~~

## 1 Introduction

~~Defining-Determining~~ the contribution of polar ice sheets to ~~sea-level~~ sea level rise is a major concern for the society, and better understanding the processes and the factors controlling ice retreat is of paramount importance to simulate the ice-sheet response to global warming (Seroussi et al., 2020; Chambers et al., 2022). Coastal glaciers in polar regions differ from temperate moun-

tain glaciers in terms of ~~volumes, catchment sizes and thermal states~~ volume, catchment size and thermal state associated to complex interactions with the ocean. The presence of floating tongues with marine ~~terminus-termini~~ makes Antarctic glaciers more sensitive to the atmospheric and ocean dynamics (Gudmundsson et al., 2019; Olinger et al., 2019; Paolo et al., 2015; Pritchard et al., 2012; Christie et al., 2022). ~~Most of the studies focused on the largest ice shelves and ice tongues of Western~~ Monitoring of Antarctic glaciers remains heterogeneous ~~Antarctica (Walker et al., 2013; Liu et al., 2015; Massom et al., 2018; Rignot et al., 2019; Millan et al., 2022) and only a few studies have analyzed the behavior of smaller size glaciers in East Antarctica (Miles et al., 2017)~~ Baumhoer et al. (2018). Studies focus either on continental scale monitoring, which usually lead to commenting the evolution of the largest glaciers of Antarctica (Walker et al., 2013; Rignot et al., 2019; Miles et al., 2022; Millan et al., 2022) or to specific glaciers or group of glaciers that concentrate most of the attention Baumhoer et al. (2018). In this study, we document and analyze the evolution of the Astrolabe glacier's ice tongue ~~and, in particular, its most recent calving event of November 2021.~~ calving cycle, which has not been updated since Frezzotti and Polizzi (2002).

The Astrolabe glacier is located in Terre Adélie, (140°E, 67°S) near the Dumont d'Urville French research station. The glacier outlet is ca. 8 km wide (Figure 1a), while the drainage basin stretches as much as 200 kilometers inland. It is characterized by a tongue of ice developing on the water, presenting a calving front of 6 km ~~width-wide~~ (Figure 1a). Due to its proximity to the Dumont D'Urville research station, the glacier has been ~~extensively~~ instrumented over the last decades ~~(Le Meur et al., 2014)~~ with a focus on its grounding zone (Drouet, 2012; Le Meur et al., 2014). However, the ~~last study documenting the calving cycle of the glacier has never been documented and due~~ 's ice tongue covers the period 1940-2000 Frezzotti and Polizzi (2002) while recent observations show an unusual spatial extension of the ice tongue until November 2021 when a major calving event occurred (Figure 1f-i). Due to its small size, ~~and rapid recent dynamic,~~ the Astrolabe glacier ~~dynamic is not properly monitored with ice tongue is not adequately monitored by~~ global value-added products such as the NASA MEaSUREs ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7). ~~Recent observations show an unusual spatial extension of the ice tongue until November 2021 when a major calving event occurred (Figure 1b-d).~~

~~Ice calving events result from rift opening and/or fracture propagation that usually pre-exist in the ice tongue for several months to years (Benn et al., 2007; Walker et al., 2013). The release of an iceberg during a calving event results from the extensive opening~~ Ice calving is defined as the detachment of a smaller ice piece of ice from a larger one (Alley et al., 2023). Calving is mostly controlled by brittle processes (Alley et al., 2023) and results from the extensive opening of cracks or rifts within the ice shelf. Lateral spreading and thinning of the ice shelf can explain the formation and propagation of these fractures that may be triggered by hydro-fracturing/rifts (Liu et al., 2015; Larour et al., 2021; Borstad et al., 2017; Alley et al., 2023). However, environmental forcing can also accelerate their propagation through hydrofracturing (Scambos et al., 2000), subglacial warm water intrusion and basal melting (Ritz et al., 2015; Rignot et al., 2019; Pritchard et al., 2012), ice heterogeneity (Borstad et al., 2017), bending of the ice due to flexural rebound after lake drainage (Banwell et al., 2013), thinning of the ice mélange within pre-existing rifts (Larour et al., 2021) or/and in the ice shelf itself (Liu et al., 2015; Larour et al., 2021), and decrease. Example of tsunamis that contributed to open rifts and trigger calving are also reported (Liang et al., 2023; Alley et al., 2023). Another important forcing is the influence of the sea-ice buttressing (Massom et al., 2001, 2018; Wearing et al., 2020; Gomez-Fell et al., 2020) surrounding the ice tongues, which result from atmospheric and ocean dynamics (Campagne et al., 2015). Indeed,

changes in atmospheric and ocean dynamics favoring the presence of sea-ice ~~may-can~~ act as a protection and allow glacier extension (Gomez-Fell et al., 2022; Massom et al., 2018; Christie et al., 2022). These processes are still poorly understood, as they exhibit strong spatial and temporal variability, which are highly difficult to document with direct observations in Antarctica. Moreover, the effect of landfast ~~by either protecting the ice tongue from ocean swell and/or in the case of landfast~~ sea-ice ~~on~~ (i.e. sea-ice faster to the glacier/to the coastline) by buttressing the ice tongue (Massom et al., 2001; Walker et al., 2013; Robel, 2017). The thickness of sea-ice or ice mélange within a pre-existing rift may influence the acceleration of the rift opening (Larour et al., 2021), leading to complex calving cycles. In several cases, the disappearing of sea-ice around the ice tongue ~~remains unclear: is front~~ has been reported to trigger the instantaneous calving (Massom et al., 2001, 2018; Robel, 2017; Wearing et al., 2020; Gomez-Fell et al., 2021). However, it remains unclear if the sea ice and in particular, landfast sea ice, is buttressing the ice tongue ~~and hence~~ preventing fracture propagation or ~~is landfast sea-ice buttressing insufficient to inhibit the fracture and only hold if it only holds~~ the ice tongue parts together until calving is possible? All these processes are still poorly understood, as they exhibit strong spatial and temporal variability, which are highly difficult to document with direct observations in Antarctica.

In this study, we determine for the first time the ice tongue extension cycle of the Astrolabe glacier from high resolution optical satellite imagery (ERS, MODIS, Landsat, Sentinel-2 and ASTER) over the period 2000-2021. We also use 1947-2021. The archive of Sentinel-2 images is used to compute surface velocity of the ice for the entire area of the Astrolabe glacier from 2017 to 2021. We show the added-value of optical satellite images to monitor fractures propagation using ice velocity and strain rate calculated with optical image correlation. We compare these data to the the front line evolution of the ice tongue with the sea-ice extent around the Astrolabe glacier from the NSIDC (National Snow and Ice Data Center; Fetterer and Windnagel 2017). We show that although calving events occur in majority during periods of landfast almost systematically when sea-ice melting, the propagation of the fissure extension decreases around the ice tongue terminus, but that the rift propagation can take place in the middle of the austral winter when the ice tongue is totally embedded in sea-ice sea ice suggesting that sea ice buttressing may not be sufficient at the Astrolabe glacier to prevent calving.

## 2 Data and methods

### 2.1 Satellite imagery

#### 2.1.1 Mapping of the ice front position

The ice front of the Astrolabe glacier was mapped using Landsat, ASTER and Sentinel-2 optical satellite images mainly optical satellite imagery at high resolution ( $< 50$  m) available in open access (i.e. Landsat, MODIS, ASTER and Sentinel-2). The first satellite image available has been acquired by Landsat-4 on January 14, Landsat-1 on January 29, 1973 (Figure 1c). The next available acquisitions are acquired in 1989 with Landsat-4/5 and then in 1999 with Landsat-4 (Figure 1bc). In total, 54 images are analyzed and the evolution of the ice front is mapped manually for four periods: 1989-1990-1947, an aerial photography was taken of the Astrolabe. We used the sketch of the photography to extract the ice front position (Figure 1b), 2000-2010-c) although important distortions are visible. In order to complete the optical dataset, we used radar

acquisitions from ERS satellites and Radarsat RAMP product (Jezek and Barry., 2013) to map the ice front position between 1996 and 1999 (Figure 1c), ~~2010-2020~~. From 2000 to 2010, Landsat and ASTER satellites provide around 1 to 3 images per year. We complete this dataset with the analysis of MODIS images (Figure 1d), ~~and 2020-November 2021~~. From 2010, Landsat-8 and then, from 2017 Sentinel-2 provides regular acquisition at high resolution of the ice front evolution (Figure 1e). ~~No optical image is found for the period 1990-2000. The~~ The combination of these two satellites allowed to monitor the calving of November 2021 with daily acquisition (Figure 1g-i). In total, 113 images are analyzed and the evolution of the ice front is mapped manually. Finally, the area of the floating tongue is estimated considering ~~a~~ an arbitrary reference grounding line position (Bindschadler et al. 2011; Figure 1a). A more precise delineation of the grounding line has been proposed by ~~Le Meur et al., 2014.~~ Le Meur et al. (2014).

### 2.1.2 Ice velocity monitoring from optical images

Satellite imagery is commonly used to compute the ice velocity with image correlation techniques (Avouac et al., 2006; Leprince et al., 2007; Rignot et al., 2011; Mouginot et al., 2017; Millan et al., 2022). These techniques consist in matching pixels from one image to another to retrieve the shift in the position of a particular feature through time. Several studies have shown the interest of this technique to monitor ice surface velocity (Dehecq et al., 2015; Altena et al., 2019) especially in polar regions (Joughin et al., 2018; Millan et al., 2022). We used the GDM-OPT-ICE service (~~Provost et al., 2022~~) (~~Provost et al., 2022; Stumpf et al., 2017~~) to compute ice displacement time series. The GDM-OPT-ICE service allows for the precise co-registration of the satellite imagery stack using the CO-REGIS algorithm (Stumpf et al., 2018), computes the shift between pairs of co-registered images with the open source stereo-photogrammetric library MicMac (~~Rosu et al., 2015~~) (~~Rosu et al., 2015; Rupnik et al., 2017~~) and inverts the displacement time series with the TIO algorithm (Doin et al., 2011; Bontemps et al., 2018).

The Copernicus Sentinel-2 mission provides acquisitions over the Astrolabe glacier every three to six days, during the austral summer (September to April). In total, from February 2017 to early November 2021, 59 Sentinel-2 images were acquired with no overcast over the Astrolabe glacier. The pairing network is set up to pair each image successively with the five next acquisitions, resulting in 280 pairs. The correlation is computed on a window of 5 by 5 pixels using sub-pixel refinement. The displacement time series is inverted for each acquisition date with a spatial resolution of 1 by 1 pixel (i.e., 10 m x 10 m). The resulting displacement time series is interpolated at 30 days in order to compute the evolution of the ice velocity and reduce the noise.

### 2.1.3 Computation of the strain rates from the ice velocity fields

Strain is a measure of how much a medium (here ice) stretches, compresses and deforms in all directions as it flows, whereas strain rates represent how quickly these deformations occur. The strain rates can therefore be computed using satellite-derived velocity (Alley et al., 2018; Cheng et al., 2021). We used the method described in (~~Alley et al., 2018~~) and (~~Nye, 1959~~) (~~Alley et al. (2018) and Nye (1959)~~) to compute the longitudinal, transverse and shear strain rates using the yearly estimation

120 of the ice velocity derived from the GDM-OPT-ICE outputs (see section 3.2.1). The strain rates are computed at a spatial resolution of 20 meters.

## 2.2 In situ sensors

### 2.2.1 On-site GNSS observations and displacement measurements

A permanent GNSS network (<https://astrolabe.osug.fr/>) is maintained by the Institut des Géosciences de l'Environnement (IGE) on the Astrolabe glacier. It consists of 8 GNSS stations in 2018 and 4 stations in 2021 (mainly because of a lack of maintenance in 2019/2020 due to cancellation of the summer operations in Antarctica because of the COVID pandemic; [Figure 1b](#)). The GNSS receivers and antennas are mounted on beacons specifically designed ~~for to withstand~~ harsh environmental conditions (strong winds, local wind-drifted accumulation of snow, ice motion, summer melting leading to ~~beacon~~ tilting or even collapse ~~of the beacons~~). These harsh conditions explain some gaps in the GNSS time series, mainly during the austral winters. The receivers are geodetic dual-frequency receivers (Trimble™NetR9) connected to Zephyr geodetic antennas. The GNSS observations consist of 3 two-hour ~~measuring measurement~~ sessions per day, where positions are averaged from 30-s sampling measurements. The positions are calculated for 24h measurements in PPP mode (Precise Point Positioning) using the GipsyX geodetic software (Zumberge et al., 1997). The accuracy is 1.5 cm (standard deviation 0.9 cm) in the horizontal component and 3.8 cm (standard deviation 2.7 cm) in the vertical component.

135 A field campaign was ~~realized conducted~~ in 2020 to quantify the ice velocity in the vicinity of the grounding line position. It ~~consists of 15 bamboo sticks consisted of 16 bamboo stakes~~ that were implanted ~~on the ice in January 2020 during one week in the ice during winter 2020 for one week between January 31, 2020 and February 7, 2020 (Figure 1b)~~. The position of the ~~sticks stakes~~ was measured the first day and then, one week later, with a GNSS dual-frequency receiver, allowing for an estimation of the ice velocity. ~~The derived velocity is compared to the GNSS's velocity derived from the 2018 and 2021 campaigns and to the NASA MEaSUREs ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7) available in this part of the ice tongue. The result shows a good agreement between all dataset (Figure 2) meaning that the velocity remains locally constant through the years from 2000 to 2018 in this part of the ice tongue.~~

## 3 Results

### 3.1 Ice front position: 2000-2021

145 Changes in the ~~ice tongue surface Figure 2a-d, and ice front position along three profiles are presented in Figure 2e-g. Before 2007, the number of cloudless acquisitions per year is very limited (1-2 images). From 2006, Landsat-7 (and then Landsat-8 and Sentinel-2) provides 5 to more than 10 cloudless images per year, allowing to resolve the calving cycle of the Astrolabe glacier. Evolution of Profiles AA' (western profile) and BB' (central profile) exhibit a similar temporal evolution (Figure 2c, f) .In particular, along those two profiles, the most advanced position frontal position are presented in Figure 3. The evolution of the ice tongue is 5 and 4 km for profile AA' and BB' respectively (Figure 2) with calving occurring regularly every 2 to 4~~

years. From front position varies from one profile to another. Between 1945 and 1995, historical images are sparse but show a maximal position of 4.2, 4.0 and 4.1 km for profiles AA', BB' and CC' respectively. In 2016 to 2021 and 2019, the ice tongue experienced a particularly long period of growth and reached front terminus reached this maximal position simultaneously for all three profiles. From 2019 to 2021, an unprecedentedly observed position at of 7.2 km and 6.7 km for profile is reached for profiles AA' and BB' respectively (Figure 3a, b). Conversely, the evolution of the ice front position on profile CC' remains from 2000 to 2021 at a maximum distance of 5.8 km (Figure 2g). Moreover, the occurrence of calving of this part of the ice tongue is significantly different from the central and western parts. Indeed, on profile the ice front position decreases progressively after 2020 due to successive calving events (Figure 3c). It should be noted that the central profile BB' reaches regularly its maximum position before the calving events of 2002 and 2010 (Figure 3b) while on profile AA' and CC', the ice front progresses regularly during 6 years from 2010 maximum position is only reached in 2002 or late 2002 (Figure 3a). From 2002 to 2016 when it regresses once in 2012 on profiles AA' and BB'. The ice tongue also experienced two calving events in the eastern side of the tongue in winter 2019-2020, and winter 2020-2021, while this is not observed on the other two profiles (Figure 2e-g) 2010, the ice front position experience periods of yearly calving of different lengths depending on the considered profiles: 2002-2008 for profile AA', 2004-2007 for profile BB' and 2003-2010 for profile CC'. A linear regression is also performed to retrieve the velocity of the ice front progression in between the successive calving events (Figure 2e-g3). The velocity varies highly from one period to another, but one can observe that the velocity are significantly lower for profile CC' ( $1.17 \text{ m.day}^{-1}$ - $1.55 \text{ m.day}^{-1}$ ) than for profile AA' ( $0.96 \text{ m.day}^{-1}$ - $1.79 \text{ m.day}^{-1}$ ) and BB' ( $1.75 \text{ m.day}^{-1}$ - $2.12 \text{ m.day}^{-1}$ ).

### 3.2 Ice velocity: 2017-2021

Ice velocity is plotted for each year from 2017 to 2021 (Figure 34a) together with the derived longitudinal, transversal-transverse and shear strain rates (Figure 34c, d, e respectively). Figure 3b presents the ice velocity computed with GDM-OPT-ICE and the velocity measured with in situ instrumentation (i.e., GNSS's and bambou stieks campaign). The yearly estimation obtained with GDM-OPT-ICE is compared to the one measured with in situ instrumentation -The (GNSS's and bamboo stakes campaigns). The in-situ data shows that the velocity in this part of the glacier is very constant though time and do not exhibit seasonal variations (Figure 2b, c) allowing for comparison between different years. Figure 4b presents the ice velocity computed with GDM-OPT-ICE and the velocity measured with in situ instrumentation (i.e., GNSS's and bamboo stakes campaign). The comparison between the two datasets shows that the estimation of the velocity from GDM-OPT-ICE improves with time, with a poor accuracy in 2017 ( $\text{RMS} = 0.76 \text{ m.day}^{-1}$ ) and a much better one from 2019 ( $\text{RMS} < 0.25 \text{ m.day}^{-1}$ ). One can observe that the gradient of velocity from the western border to the center of the glacier is well retrieved with the GDM-OPT-ICE products of 2019-2021 (Figure 3b). In while in 2017 and 2018, a large portion of the glacier is estimated to be stable from the results of the the limit between stable ice and the flowing ice tongue is retrieved in the wrong position with the GDM-OPT-ICE whereas the GNSS's campaign of this year show that the position of the western glacier border is stable through time products. The small number of cloudless Sentinel-2 acquisitions for those years may explain the low RMS error of these two years, as well as the wrong estimation of the ice tongue limit.

The velocity field shows a smooth gradient with lower velocity of circa  $1 \text{ m.day}^{-1}$  in the south-eastern part of the glacier and faster velocity of  $1.2\text{-}1.5 \text{ m.day}^{-1}$  in the north-western part of the glacier. In 2019, a small block of ice accelerated in the eastern part of the ice front (Figure 3a, 4a), also visible in the longitudinal strain rate field (Figure 4c, box B). This block disappears in both the mean velocity of 2020 (Figure 34a) and in the strain rate field (Figure 4c) and due to the calving of this part of the glacier occurred in December 2020 (Figure 23d, profile CC'). In 2019, an extensive fracture appeared in the longitudinal and shear components of the strain rate field in the western part of the ice tongue in front of the Dumont D'Urville station, clearly visible in the longitudinal and shear components of the strain rate fields (Figure 4c, e; box A). The northern-western part of the ice tongue starts simultaneously to exhibit larger velocities in 2020 and 2021 (Figure 34a). In 2021, a complex network of localized increase of strain rates appears on the western-northern part of the glacier delimiting the potential area of the future iceberg to be calved (Figure 34c, d, e). This complex network delimits the fractures that were observed on the ice in the first available summer acquisition in September 2021 and that remained the same until the ice calving (Figure 1f). Beside the evolution of the fractures on the ice, one can also observe the high strain rates ( $> 0.002 \text{ day}^{-1}$ ) that are clearly identifiable through time along the lateral limits of the glacier (Figure 34e).

### 3.3 Glacier acceleration and ice Ice tongue break off: 2021

The displacement time series is linearly interpolated over a time step of 30 days from the first acquisition in 2017 to November 5, 2021. Mean monthly velocity and strain rate fields are derived from this interpolation. We investigate the evolution of the strain rates for the period January and November 2021 to understand the dynamic of the recent calving (Figure 45). Strain rates maps show high concentration of strain localized along linear structures, which grow progressively from April 2021 to November 2021 (Figure 4a). These linear structures are not visible at the surface of the glacier (Figure 4e) until June 2021 except for the main rift in front of the Dumont D'Urville station. 2021. We set a threshold on the strain rates in order to analyze the evolution of these localized concentration of strain, as well as the occurrence of the spatial connection between them (Figure 4b5a). The evolution of their growth is complex, with transitions from one component to another. For example, the main rift exhibits a clear longitudinal strain rate from April 2021 to September 2021 that evolves toward a shear strain rate in October, November-October-November 2021 (Figure 4b5a). From May 2021, a large concentration of strain appears in the transverse component along another fracture oriented in the North-East/South-West direction (Figure 4a, b5a). Similarly, a third fracture appears in the longitudinal component on the eastern side (Figure 4a, b5a). These fractures grow rapidly and connect together in June 2021 (Figure 4b)-5a). One can also observe that from October 2021, most of the fractures exhibit a shear strain rate, likely due to the rotation of the blocks. We analyzed the Sentinel-1 SAR images from May 2021 to August 2021 in order to validate these observations. We observe that the network of fracture opened suddenly between June 13, 2021, and June 25, 2021 (Figure 4e, d5b, c) which is coherent with the timing of the connection derived from the time series of surface displacement. One can also observe that from October 2021, most of the fractures exhibit a shear strain rate, likely due to the rotation of the blocks strain rate (Figure 5a). It can be noted that compressional strain rates are measured from 2017 to 2020 at the terminus of the glacier tongue with strain rate larger than  $0.001 \text{ day}^{-1}$  while it is not observed anymore in 2021 (Figure 34c).



### 3.4 Landfast sea-ice Sea-ice forcing

We analyze time series of sea-ice extent and concentration in the region of the Astrolabe glacier (Figure 1a). The data for sea-ice extent and concentration were are downloaded on the NSDIC repository (Fetterer and Windnagel, 2017) and cover the periods from 1979 to the end of 2021. We cropped the data to analyze the monthly variation of the sea-ice extent over an area of 4000 km<sup>2</sup> around the Astrolabe glacier from 1979 to the end of 2021 (Figure 5a) and the (dotted blue lines on Figure 1a). This area is chosen arbitrary to represent the influence of regional sea ice variation on the Astrolabe ice tongue. The daily variation of the sea ice concentration at is taken at the pixel (25 km x 25 km) encompassing the border of the Astrolabe glacier from 2000 to the end of 2021 (Figure 5b) (dotted yellow lines on Figure 1a) to focus on the conditions at the Astrolabe ice tongue.

We observe a significant change in the periodicity of sea-ice melt around the Astrolabe glacier in the last decade (2011-2021). Indeed, from From 1979 to 2011, the extent of the sea-ice decreases significantly every year during the summer while from (Figure 6a). From 2011 to 2021, the annual disappearing of sea ice does not occur every year (Figure 6a). Indeed, during two successive periods: 2012-2016 and 2016-2020, where 2016-2021, the extent of the sea-ice remains maximal (Figure 5 during summer (Figure 6a). These periods of Multi-Year Fast sea-Ice (MYFI) are well correlated with periods of extension of the Astrolabe ice tongue (Figure 5). In detail, one can see that during those two periods, the sea-ice extent can drop occasionally (e.g., early 2015, 2018) or during larger periods such as for austral summer 2018-2019 (Figure 6a).

However, the reduced length or absence of periods of sea ice free conditions during 2012-2016 and 2016-2021 is notably different from the previous decades. The time series of daily sea ice concentration shows similar observations (Figure 6b). Before 2011, the sea ice concentration drop to 0 below 15% for periods of 2 to 3-4 months from November to mid-March with small variation, with small variations on the length of sea-ice free periods (Figure 6b). From 2008, the onset of the sea-ice free periods shifts to 2011, one can observe a decrease in the duration of low sea ice concentration to 2 months (Figure 6b, d), corresponding to a shift of the free sea ice period onset from November to mid-December/January, while the end of the sea-ice free periods remain stable through time (Figure 5b): beginning to mid-March. From 2012 to 2016, the sea ice free periods disappear or are shortened to one month and delayed to the month of February (Figure 5b). From 2017 less than one month (i.e. February 2015; Figure 6d). Between 2016 and 2021, the regime of sea ice concentration is highly variable, with years with no to very short periods of sea ice melt (2017 free conditions (austral summer 2016-2017, February 2020; Figure 5b, 6b, d) and years with prolonged sea ice melting (December 2018 - March free conditions (December 2018-March 2019; November 2020 - March 2021). The date of the different calving events cannot be determined with daily precision, but one can observe that all calving occur when sea-ice concentration decreases to 0% at the end of the Austral fall (Figure 5b). From austral summer 2020-2021, the duration of sea ice free conditions seems to resume as before 2012 with a duration of 3-4 months from mid-November to March (Figure 6b, d).

We also plot compare the evolution of the ice surface velocity along profile AA' and CC' from 2000 to November 2021 (Figure 5c, d). Yearly estimation of the velocity provided by the NASA MeaSURES ITS\_LIVE project (Gardner et al., 2018) are plotted for years 2000-2017. From 2017 onwards, the monthly velocity estimated from the GDM-OPT-ICE chain is



plotted. Over the two last decades, we do not see significant variations in the ice velocity from 2000 to 2021 except for local accelerations (Figure 5c, d) sea ice extent and concentration to the evolution of the ice tongue area (Figure 6c). We observe that the surface seems to flow slightly faster on the western part of the ice tongue (Profile AA', Figure 5e) than on the eastern part (Profile CC', Figure 5d) which is also observed in the velocity of the terminus position (Figure 2e-g). Local acceleration are visible in 2021 on profile AA' and in 2019-2012-2016 and 2020 for profile CC' (Figure ??e, d) and correspond to the area of the glacier that calved in 2019, 2020-2016-2021 periods corresponds to periods of significant extension of the Astrolabe ice tongue (Figure 6c) with an increase of 15 km<sup>2</sup> between 2012 and 2021 (Figure ??b). 2016 and of almost 20 km<sup>2</sup> between 2016 and 2021. For the period 2002-2012, the ice tongue extension is much more limited due to the regular calving at different location of the ice front (Figure 3). Before 2017, the yearly measures are cropped before the end of the ice tongue and the area where the fissures usually initiate is not mapped by the dataset. Moreover, the seasonal variations remain difficult to infer as optical data are not acquired during winter. 2002, the satellite images are scarcer, but the ice tongue seems to have reach a rather advance position in 2002 with an area of almost 81 km<sup>2</sup>. This advance can not be linked to significant variation in the sea ice seasonal cycle. We report the calving event that can be observed with the analyzed satellite images (Figure 6d) with the incertitude on the date of the different calving events. One can observe that all detected calving occur when sea-ice concentration decreases at the end of the Austral fall (Figure 5d) except for austral summer 2006-2007 where multiple calving events are reported and do not occur necessarily at the onset of the sea ice concentration decrease.

#### 4 Discussion

To understand the recent evolution of the Astrolabe glacier, we investigated the evolution of sea-ice in the vicinity of the Astrolabe ice tongue. Sea-ice and in particular, landfast sea-ice evolution is usually assumed to delay ice tongue break off and favor its extension by buttressing the ice tongues and protecting them from ocean swells (Massom et al., 2010, 2018; Rott et al., 2018; Gomez . At the Astrolabe glacier, we observe a significant change in the periodicity of sea-ice in the recent decade (2011-2021) in comparison with the previous observations (1979-2011; Figure 6). The recent periods of multi-year sea-ice presence are well correlated with the ice tongue spatial extension (Figure 6) and seems to validate the assumption that sea-ice protects the ice tongue and favors its extension. Moreover, the disappearing of such protection has been reported to initiate rift propagation which, in some cases, lead to rapid calving (Miles et al., 2017; Cassotto et al., 2021; Gomez-Fell et al., 2022; Christie et al., 2022) . In the case of the Astrolabe glacier, we also observe that the calving, when it occurs, systematically takes place at the beginning of the sea-ice disappearing (Figure 6d) which seems to confirm the potential of sea-ice disappearing as a triggering factor of calving. However, the analysis of satellite images from 2017 to 2021 at the Astrolabe ice tongue shows that the rifts or crevasses that lead to the 2021 calving event are forming several months to years before calving (Figure 4, 5) suggesting a different mechanism. The presence of rift and fracture networks in ice tongue several years to several months prior to calving has also been reported in other glaciers (Fricker et al., 2005; Walker et al., 2015; Cheng et al., 2021; Larour et al., 2021). However, we observe at the (Fricker et al., 2005; Walker et al., 2013, 2015; Cheng et al., 2021; Larour et al., 2021; Gomez-Fell et al., 2022) . In most cases, the growth of the rifts or their (re-)activation are observed during the austral summer (Fricker et al., 2005; Walker et al., 2013,

and few rifts propagation are reported during Austral winter (Walker et al., 2013; Larour et al., 2021). At the Astrolabe glacier's  
285 ice tongue ~~that a~~, the main rift is located in front of the Dumont D'Urville Research station and initiated in 2019, a year with  
almost two consecutive month of low sea ice concentration. The absence of a significant rift in 2017 and 2018 suggests that  
sea ice may have had an effect in inhibited rift growth on the Astrolabe ice tongue and delayed the calving in 2017-2018 and,  
possibly, in 2012-2016. We note that in 2019, no calving occurred despite the long period of low concentration of sea ice while  
in 2020, the eastern part of the ice tongue calved at the onset of the short (one month) period of sea ice concentration decrease  
290 (Figure 6d). These observations suggest that sea ice acts like a glue to hold together the ice tongue, and that deep opened rifts  
must pre-exist for calving to occur.

In June 2021, we observe the opening of a complex network of fractures ~~opened~~ suddenly in the middle of ~~June 2021 (Figure~~  
~~4)at the tip, and in the transversal direction, of the pre-existing main rift that has grown regularly for 3 years (Figure 3). Few~~  
~~rift or fracture propagation are reported to occur during Austral winter (Walker et al., 2013; Larour et al., 2021). The Austral~~  
295 ~~winter (Figure 5). Larour et al. (2021) proposed a mechanism to explain the winter propagation at the Larsen C Ice Shelf,~~  
~~Antarctica, prior to the calving of iceberg A68 based on the critical thinning of the ice shelf and of the ice mélange within the~~  
~~rifts. Here, the~~ critical thinning of the ice tongue due to its exceptional extension may explain this timing (Robel, 2017; Larour  
et al., 2021; Åström and Benn, 2019) although it would likely favor the propagation of the rift along the same direction as the  
pre-existing rift, which is not observed at the Astrolabe glacier (Figure 45). Instead, the main fracture propagating in June 2021  
300 is oriented along the flow direction and opened in extension (Figure 45a, b). ~~A~~ ~~Another~~ possibility to explain the development  
of these fractures could be a transition from a ductile to a brittle behavior with the decrease of temperature during winter that  
may favor fractures along the flow resulting from the differential compressive load due to sea-ice buttressing and rift opening  
(Figure 4). This mechanism might be possible as the compressive longitudinal strain seems to disappear in 2021 at the glacier  
terminus (Figure 34c) and the rift opens progressively. However, such a scenario ~~remain to be validated as it~~ would maintain  
305 the ice tongue terminus at the same position due to the effect of sea-ice buttressing, which is not what is observed (Figure 4c,  
d). ~~Moreover, and because~~ the compressive strength of the ice is much higher than extensive strength (Benn et al., 2007). ~~A~~  
~~possible mechanism could that the rupture results from~~ ~~The presence of~~ extensive circumferential stress that appear when the  
unconfined part of the ice tongue reaches a certain extension (Wearing et al., 2020). ~~Basal should be also considered, as well~~  
~~as the presence of basal~~ channels and basal melt ~~may as well that may~~ play a role in the dislocation of the ice tongue (Vaughan  
310 et al., 2012; Alley et al., 2023). The difference in the calving cycle and ice velocity between the eastern part and the western  
part of the glacier terminus also suggest that the bathymetry underneath the ice tongue controls the location and evolution of  
the rifts. ~~Further~~ ~~Our analysis remains limited, and further~~ modelling is necessary to understand the mechanisms that lead to  
the apparition of these fractures at this time of the year ~~and with this geometry~~ (Åström and Benn, 2019; Crawford et al., 2021;  
Alley et al., 2023).

315 ~~To understand the recent evolution of the glacier, we investigated the evolution of landfast sea-ice in the vicinity of the ice~~  
~~tongue. Landfast sea-ice evolution is usually assumed to delay ice tongue break off and favor its extension by buttressing the~~  
~~ice tongues and protecting them from ocean swells (Massom et al., 2010, 2018; Rott et al., 2018; Gomez-Fell et al., 2022). We~~  
~~observed a significant change in the periodicity of landfast sea-ice in the recent decade (2011-2021) in comparison with the~~

320 previous observations (1979-2011; Figure 5a). Indeed, in the period 2011-2021, multi-year landfast sea-ice occurred with no sea-ice melt, or very short and/or episodic periods of melting (Figure 5b) at the vicinity of the Astrolabe ice tongue. The recent periods of multi-year sea-ice presence are well correlated with the ice tongue spatial extension (Figure 5a) and seems to validate the assumption that sea-ice protect the ice tongue and favor its extension. We also observe that the calving systematically occur at the beginning of the sea-ice melting (Figure 5b) and is preceded by local accelerations of the areas that calved then (Figure 5c-d). These observations together with the presence of rifts visible several months/years before calving suggest that landfast sea-ice acts like a glue to hold together the ice tongue and have little effect to prevent rift opening at the Astrolabe ice tongue.

325 ~~Landfast sea-ice is~~ Sea-ice is deeply connecting to regional and local atmospheric and oceanic states (Fogt et al., 2022). At the scale of the continent, records in Antarctica show a ~~similar trend with a general~~ positive increase of the sea ice extent from 1979 to 2016 with a minimum of global sea-ice extent recorded in summer 2017 (Fogt et al., 2022). In the region of the Astrolabe, ~~Miles et al. 2022 reports similar observations with the~~ encompassing the Adélie Coast and George V Land,  
330 the calving of the ~~Commandant glacier in 2016 after multi-year expansion of this ice tongue between 2008 and 2016. These observations suggest that the calving of 2016 could be related to~~ Mertz Ice tongue in 2010 (Massom et al., 2018) lead to severe modifications of the sea ice production and location, traducing regional changes in the oceanic and atmospheric currents ~~maybe similar to the ones observed in the Antartecia Peninsula (Christie et al., 2022). In 2021, the Astrolabe ice tongue reached an unprecedentedly observed advanced position with a total surface of 95 km<sup>2</sup>. Due to the limited temporal extent of the datasets and sparse availability of the satellite imagery before the 2000s, it is impossible to conclude if these changes are exceptional or~~  
335 ~~part of a longer cycle. However, the timing of the fracture suggests that the mechanisms that lead to~~ (Campagne et al., 2015). This event is likely at the origin of the calving of November 2021 might have been different from the previous calving at the Astrolabe glacier. The recent increase in the frequency of satellite observations has significantly improved the glacier monitoring and recent calving (like the one of November 2021 at the Astrolabe) could be documented almost day by day  
340 ~~with open access optical satellite imagery (Figure 1f, g, h) while before~~ transition of the sea ice seasonal cycle at the Astrolabe glacier (Figure 6). Miles et al. (2022) reports similar observations further west on the Adélie Coast, with the continuous growth of the Commandant glacier (Adélie Coast) from 2010 to 2018 due to the presence of persistent sea ice. This illustrates how one calving event such as the Mertz Ice tongue calving in 2010, ~~1 to 2 images per year at most are available during summer. The dataset gathered in this study could be completed with the processing of radar imagery to monitor the glacier seasonal velocity variations and document more precisely the evolution of the rift propagation during winter, where optical acquisitions are not available. Radar interferometry (InSAR) could also be used to map~~ may significantly modify the calving cycle of neighboring ice tongues, which are difficult to account in current models (Edwards et al., 2021; Miles et al., 2022). The extent of the regional impact of the Mertz Ice tongue 2010 calving is not clearly known, as most studies focus on the Georges V land area (Kusahara et al., 2011, 2017; Campagne et al., 2015; Cougnon et al., 2017). Moreover, the evolution  
350 ~~of the glacier elevation and the position of the grounding line~~ ice shelves of the Adélie coast and Georges V land remains limited (Frezzotti et al., 1998; Frezzotti and Polizzi, 2002), preventing a better understanding of the environmental forcing Massom et al. (2018); Christie et al. (2022).

## 5 Conclusions

In this study, we analyzed the evolution of the Astrolabe glacier located in Terre Adélie/Adélie Coast, Antarctica. We used  
355 open access optical satellite imagery (MODIS, ASTER, Landsat and Sentinel-2) completed by ERS and RADARSAT images  
to map the evolution of the ice front from 2000-1947 to 2022. We also measure the surface velocity and derived strain rate  
fields between 2017 and 2022, using image correlation of Sentinel-2 images. The recent evolution of the glacier shows an  
unprecedentedly documented extension of 95 km<sup>2</sup> favored by the concomitant high concentration of the ~~landfast~~-sea ice in  
the region ~~of Antarctica~~ during 2011-2021 in comparison with previous records (2000-2011). The early melt of the sea ice in  
360 November 2021 favored the released of a 20 km<sup>2</sup> iceberg in the north-western part of the Astrolabe glacier. This is the first  
time a calving of this magnitude is documented at the Astrolabe glacier. We also observed that a complex network of fractures  
opened during the austral winter in June 2021 several months before the iceberg calving. This study shows the importance of  
ice velocity and strain rates fields time series derived from optical satellite imagery at high resolution to document fracture  
opening and raises further questions on the mechanism of rift propagation.

365 *Data availability.* We acknowledge the use of imagery from Copernicus Sentinel-1 and 2 data and NASA ASTER and Landsat images  
(through <https://earthexplorer.usgs.gov/>). The GNSS observations are accessible on the Astrolabe repository: <https://astrolabe.osug.fr/>.

*Author contributions.* FP designed the experiments with contributions from DZ, ELM, JPM and CH. ELM provided the GNSS data and  
JPM processed them. FP processed the satellite data. All co-authors participated in the writing and/or revision and approval of the submitted  
manuscript

370 *Competing interests.* We have no competing interests.

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Mesocentre) allowing optimized computation. The service is accessible on-demand through the ForM@Ter data hub ([en.poleterresolide.fr/](http://en.poleterresolide.fr/services-en/gdm-en/#/optic)  
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Figure1.png

**Figure 1.** a) Map-Location of the Astrolabe glacier with the coastline and grounding lines from (Gerrish, 2022) Gerrish (2022) and the topography Sentinel-2 image of the area (from the 2021 GLO30 Copernicus DEM at 30m resolution February 7, <https://doi.org/10.5270/ESA-e5d3d65>) 2020 in the background. The limits of the 4,000 km<sup>2</sup> box where the sea-ice extent is extracted is represented in dotted blue lines. Yellow dotted lines delineate the pixel extent and location of the sea-ice concentration grid from which the sea-ice concentration is extracted. The inset b) is a zoom over the Astrolabe glacier ice tongue and indicate the profiles where the evolution of the ice front position is indicated in presented (Figure 2). In-situ measurements are also represented: red dots for the small inset on location of the left upper corner bamboo stakes, blue and green dots for the GNSS initial position of 2018 and 2021 campaigns respectively. Figures bc) to df) show the ice front position at different dates. Figure g) to i) shows the calving of November, 6 2021 visible on b) from the Sentinel-2 acquisition of November 5, 2021 -e(e) the Landsat-8 acquisition of November 6, 2021 and d(h) and the Sentinel-2 acquisition of November

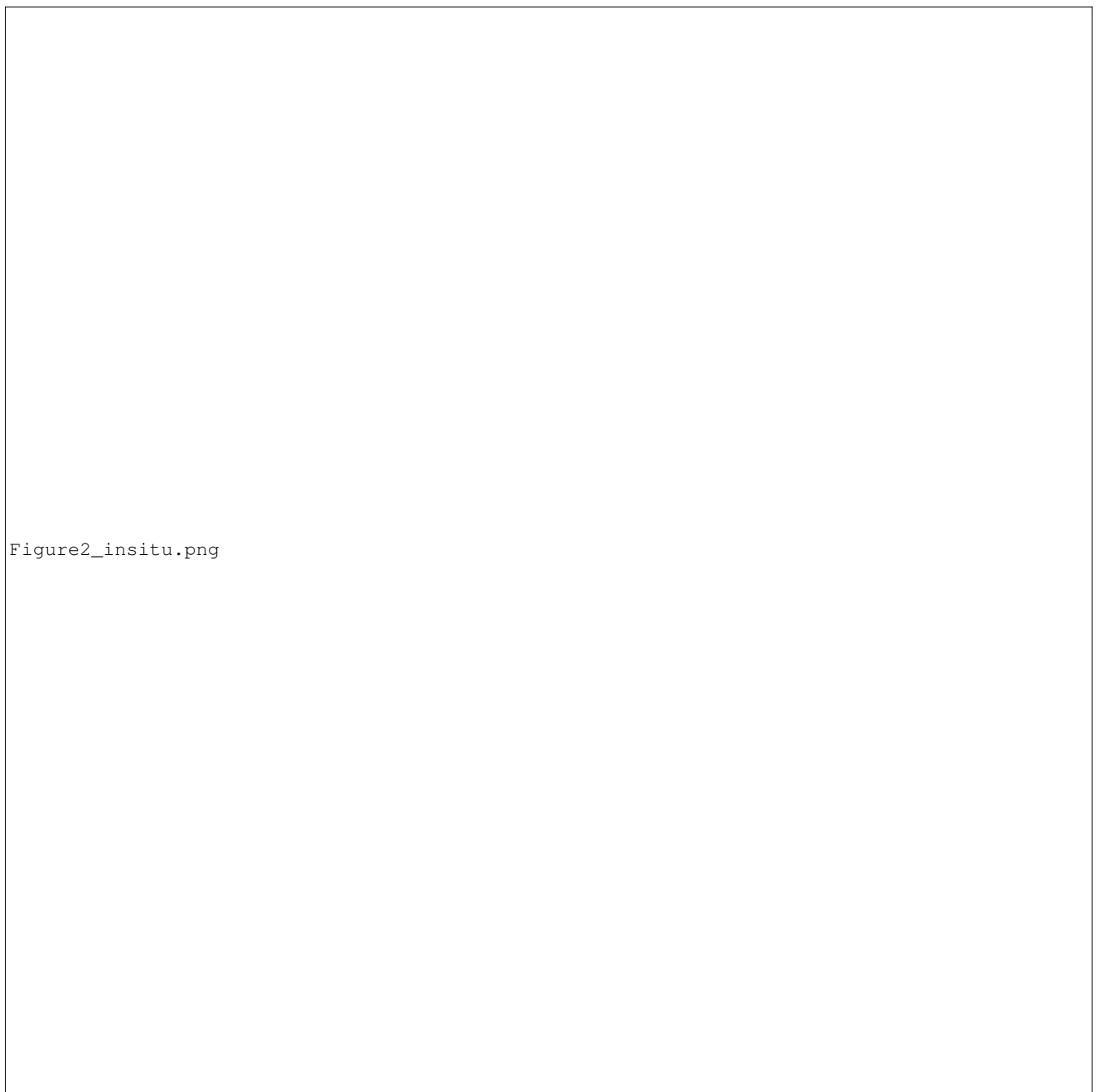


Figure2\_insitu.png

**Figure 2.** Comparison between yearly velocity measured with in-situ measurements from year-long GNSS campaigns and week-long bamboo stakes campaign; a) location of the GNSS's and bamboo stakes (see Figure 1 for the location on the ice tongue); b) comparison between the different in-situ dataset. c) comparison between the estimation of the velocity from the bamboo stake campaign and the yearly estimation of the velocity from satellite imagery from the NASA MEaSUREs ITS\_LIVE (doi:10.5067/6II6VW8LLWJ7).

Figure2.pdf

**Figure 3.** Evolution of the front position of the glacier for the periods a: 1980-1989, b: 2000-2010, c: 2010-2020 and d: January 2020–November 2021. The background images are band-4 of Landsat-4, -7, -8 and Sentinel-2 chosen as representative of the period. Figures e-g present the position of the terminus of the glacier along profiles AA' (ea), BB' (fb) and CC' (gc); (see Figure 1a) for the location of the profiles). The velocity of the ice front motion is indicated for the periods of ice front progression and calving events are presented with boxes.



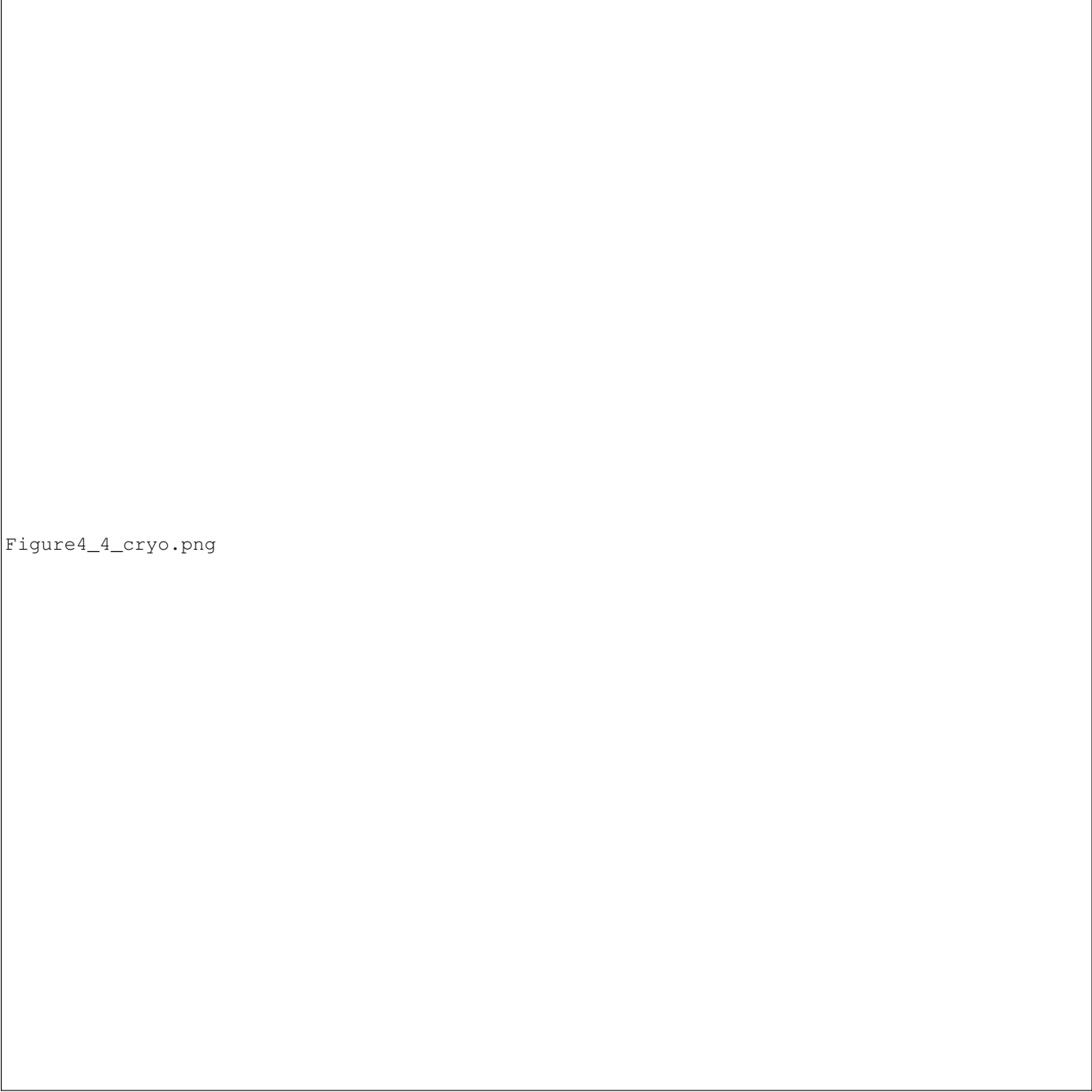


Figure4\_4\_cryo.png

**Figure 4.** a) Yearly estimation of ice velocity for the Astrolabe glacier for the period 2017-2021, b) comparison of the velocity magnitude and direction as measured by the in situ instrumentation (GNSS's and bamboo sticks campaign) and as measured by GDM-OPT-ICE. Figures c, d, e present the longitudinal, transversal and shear strain rates derived from the ice velocity fields.

Figure8\_cryo\_v2.png

**Figure 5.** a) ~~Monthly strain field derived from interpolated displacement time series~~, b) map of the strain rates larger than  $0.002 \text{ day}^{-1}$ . The three strain rate components (longitudinal, transverse and shear) are plotted together with different color. Subsets c) and d) are showing the occurrence of the fractures detected with Sentinel-1 acquisitions of June 13 and June 25, 2021. Arrows indicate the location of the main rift (b) and of the secondary fractures (c).

Evolution of monthly sea ice extent versus ice tongue extension (a), daily sea ice concentration (b) and ice tongue velocities along profile AA' (c) and CC' (d). Monthly sea ice extent is computed for a wide region of 4000 km<sup>2</sup> around the Astrolabe glacier (a) while the daily sea ice concentration is taken for the pixel of 25 km by 25 km at the Astrolabe ice tongue. Interval where observed calving have been observed are reported on the sub-figure (b) in white. The intervals usually span several days to weeks, depending on the availability of cloudless optical imagery. In years 2003-2005, calving likely occurred (see sub-figure a), although no observation can confirm the date. For the period 2000-2017, yearly estimation of the ice velocity from NASA MeaSURES ITS\_LIVE project (Gardner et al., 2018) are plotted while after 2017, monthly estimation of the velocity from the GDM-OPT-ICE processing chain are represented.

Figure6\_review1.pdf