



# Measurements and Calculations of Enhanced Side/Back Scattering of Visible Radiation by Black Carbon Aggregates

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**Abstract.** Aerosol particles have both natural and anthropogenic origins and are ubiquitous in the atmosphere. One particularly important type is carbonaceous aerosol, including a specific subset, often termed ‘elemental carbon’ chemically or ‘black carbon’ (BC) radiatively. Carbonaceous aerosol particles have implications for atmospheric chemistry, human health, and climate both directly and via their ability to act as site of cloud droplet or ice crystal formation. Laboratory experiments and theory are needed to better understand these particles, specifically their radiative impact. We present here laboratory measurements of side/back scattering of visible radiation by analogues of atmospheric BC aggregates obtained using a depolarizing optical particle counter and accompanying theoretical calculations of scattering by compact and fractal theoretical BC aggregates. We show that with random-orientation, the theoretical calculations reproduce the qualitative behavior of the measurements but are unable to reproduce the highest values of the linear depolarization ratio; we are only able to obtain high values of the linear depolarization ratio using fixed orientation. Thus, we suggest that it is possible that models of scattering by BC aggregates that employ the random orientation assumption/option may underpredict the linear depolarization ratio of actual BC aggregates. Both our measurements and our theoretical calculations point to the possibility that bare (uncoated) BC aggregates, as opposed to the aged/coated BC or soot that was investigated in previous studies, can exhibit higher backscattering linear depolarization than previously assumed.

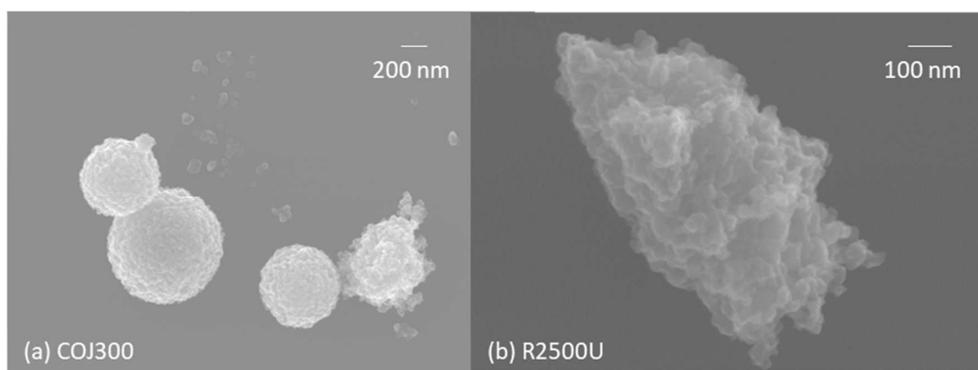
## 1 Introduction

Sela and Haspel (2021) presented theoretical calculations of scattering of visible radiation by pairs of aggregates comprised of spherical nano-scale primary particles. Each aggregate pair consisted of an ordered aggregate with a simple cubic (SC) configuration and a disordered aggregate with an ideal amorphous solid (IAS) configuration based on the model of Stachurski (2003; 2011; 2013), and the scattering was computed using the multiple sphere *T*-matrix (MSTM) model of



Mackowski and Mishchenko (1996). Sela and Haspel (2021) found that holding all other parameters constant, in most cases, the overall scattering and absorption and hence extinction of radiation by ordered aggregates is stronger than for disordered aggregates. At the same time, they found that holding all other parameters constant, disordered aggregates tend to side scatter and back scatter more strongly than ordered aggregates.

To further investigate the enhanced side/back scattering by disordered aggregates, in the present study, we perform a set of scattering calculations on ordered and disordered aggregates with varying degrees of disorder and compare our results with scattering measurements conducted on analogues of atmospheric black carbon (BC) aggregates whose microphysical and ice nucleation properties were presented in Zhang et al. (2020). The BC sample sets labelled “COJ300” and “R2500U 400 nm” in Zhang et al. (2020) exhibit similar primary particle diameters ( $d_{pp}$ ;  $\sim 35 \pm 10$  nm) and mobility diameters ( $D_m$ ; 400 nm) to one another. At the same time, the outer envelopes of the COJ300 samples appear more spherical, while the R2500U samples appear more fractal (see Fig. 1), which is consistent with the fact that the mean fractal dimension ( $D_f$ ) of the COJ300 samples (2.34 with a 95% confidence interval range of 2.12-2.56; see Table 1 of Zhang et al. (2020)) is higher than the mean fractal dimension of the R2500U samples (1.92 with a 95% confidence interval range of 1.68-2.16). See DeCarlo et al. (2004) for a comprehensive discussion of particle morphology parameters.



**Figure 1: SEM images of (a) some aggregates from the COJ300 sample set and (b) an aggregate from the R2500U sample set from the Zhang et al. (2020) study.**

In the present study, we examine the scattering measurements at an angular range of  $135 \pm 20^\circ$  obtained with the SPectrometer for Ice Nuclei (SPIN) (Garimella et al., 2016) instrumentation at 670-nm wavelength for the aforementioned two sets of samples, COJ300 and R2500U size-selected at 400 nm, from Zhang et al. (2020). We conduct new theoretical calculations for comparison to the measured scattering in a similar vein to Sela and Haspel (2021), where the aggregates in each set consist of the same number of primary particles,  $N_{pp}$ , of the same primary particle size ( $d_{pp}$ ) but differing degrees



of disorder, and now focusing on BC aggregates. Thus, we can examine whether the side/back scattering tendencies found in  
55 Sela and Haspel (2021) are reproduced in actual measurements and how the degree of disorder influences these tendencies.

In addition, the fact that the SPIN measurements are in situ measurements of scattering by individual particles rather than  
bulk scattering measurements, we have a unique opportunity to examine how the present set of measurements and  
calculations compare with previous measurements and calculations of side/back scattering by bare (uncoated) BC  
aggregates, such as those presented in Bohren and Kho (1985), Lu and Sorensen (1994), Gustafson and Kolokolova (1999),  
60 Liu and Mishchenko (2005), Liu and Mishchenko (2007), Liu et al. (2008), Burton et al. (2013; 2014), and Kahnert and  
Kanggießer (2020) and references therein.

## 2 Methods

### 2.1 SPIN scattering measurements

Optical measurements were performed using a linear depolarization optical particle counter (OPC) associated with the SPIN  
65 instrument (Garimella et al., 2016). The SPIN OPC is equipped with a continuous-wave 500-mW 670-nm wavelength laser  
(Osela ILS-640-250-FTH-1.5MM-100uM). Particle measurements are made with four optical detectors. Size is measured  
based on side scattering with a detector situated at a zenith angle of  $90^\circ$  (i.e.,  $90^\circ$  from the direction of propagation of the  
incident laser beam) using a Mangin mirror pair. Three backscattering detectors measure the scattered photon counts  
according to polarization. The incident radiation from the laser is polarized with its electric field vector parallel to the  
70 scattering plane. Detectors P1 and P2 measure scattered photons with parallel polarization (the same polarization as the  
incident radiation), while detector S1 measures scattered photons with perpendicular polarization (electric field vectors  
perpendicular to the plane of the scattering). As mentioned in Sect. 1, these three detectors are each situated at a scattering  
zenith angle,  $\theta_{\text{sca}}$ , of  $135^\circ$  with a half angle of acceptance of  $20^\circ$ . Detectors P1 and S1 collect photons from the same  
scattered photon stream after it passes through a 50/50 polarizing beam splitter, while detector P2 collects photons from a  
75 separate photon stream that propagates at a different azimuthal angle with respect to the direction of propagation of the  
incident laser beam (but still propagates at a scattering zenith angle of  $135 \pm 20^\circ$ ). For each of these two photon streams, the  
scattered laser light propagating at  $\theta_{\text{sca}} = 135 \pm 20^\circ$  first passes through a collimating lens, which transforms the scattered  
rays into parallel rays, followed by a focusing lens, which focuses the rays towards the detector. This lens configuration is  
intended to provide equal weight to each ray in the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  and approximately unit transmission. Scattering  
80 data for each particle is recorded in units of photon counts (photons per second). Given that the incident radiation from the  
laser is polarized parallel to the scattering plane, a higher photon count registered in detector S1 (in an absolute sense and/or  
relative to the photon counts registered in detectors P1 and P2) indicates some asymmetry/nonsphericity in the shape of the  
scattering particles or possibly birefringence or chirality in the scattering particle material. See Droplet Measurement  
Technologies, Inc. (2013) and Garimella et al. (2016) for more details on the SPIN instrumentation.



85 Particle generation and characterization of the BC samples followed the methodology outlined in Zhang et al. (2020). The size distributions of measured BC particles follow a Poisson/log-normal distribution. To avoid the influence of multiply charged BC particles, which could reach up to 16% of total BC population, size thresholds corresponding to the 90% quantile of diameter (1310.7 nm for COJ300 and 6769.4 nm for R2500J, respectively) were applied to the particle-by-particle data.

90 The relative humidity (RH) conditions of the SPIN experiments (62% at  $-50^{\circ}\text{C}$  to 68% at  $-40^{\circ}\text{C}$ ) were below liquid water saturation. If any water vapor molecules had condensed onto the surfaces of the particles, they would have frozen immediately, resulting in an observable ice crystal signal. Ice crystals were not observed, and we therefore assume that the BC particles examined were dry.

## 2.2 Theoretical calculations

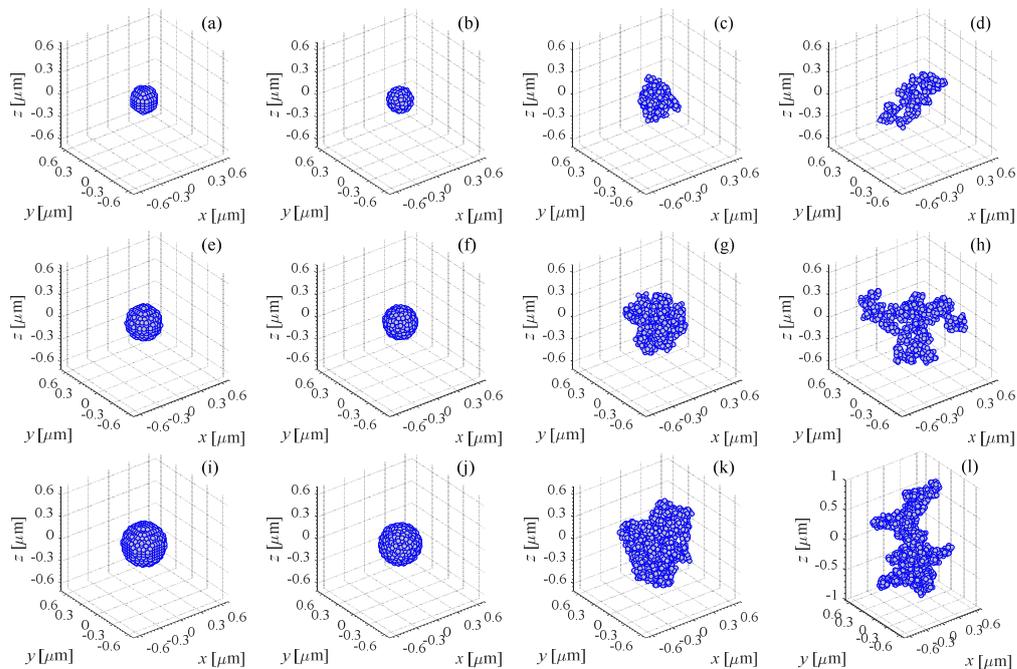
95 The theoretical aggregates are based on the mean microphysical properties of the COJ300 and R2500U 400 nm samples from Zhang et al. (2020), as listed in Sect. 1, but we also test the sensitivity of the results to variations in  $D_m$  and variations in  $d_{pp}$ . For each set of aggregates, first an SC aggregate with a roughly spherical outer envelope is constructed, where in each SC aggregate, the primary particles touch but do not overlap (point contact). Our default SC aggregate has an outer-envelope diameter ( $D_{\text{outer-envelope}}$ ) of 400 nm and a primary particle diameter of 35 nm.

100 Next, the IAS model of Stachurski (2003; 2011; 2013) is employed to construct a disordered but very compact and still roughly spherical aggregate with the same values of  $D_{\text{outer-envelope}}$ ,  $d_{pp}$ , and  $N_{pp}$  as the respective SC aggregate. (Refer to the description of the pairs of aggregates in Sela and Haspel (2021).) As with the SC aggregates, in the IAS aggregates, the primary particles touch one another but do not overlap, and each aggregate is monodisperse with respect to its primary particles.

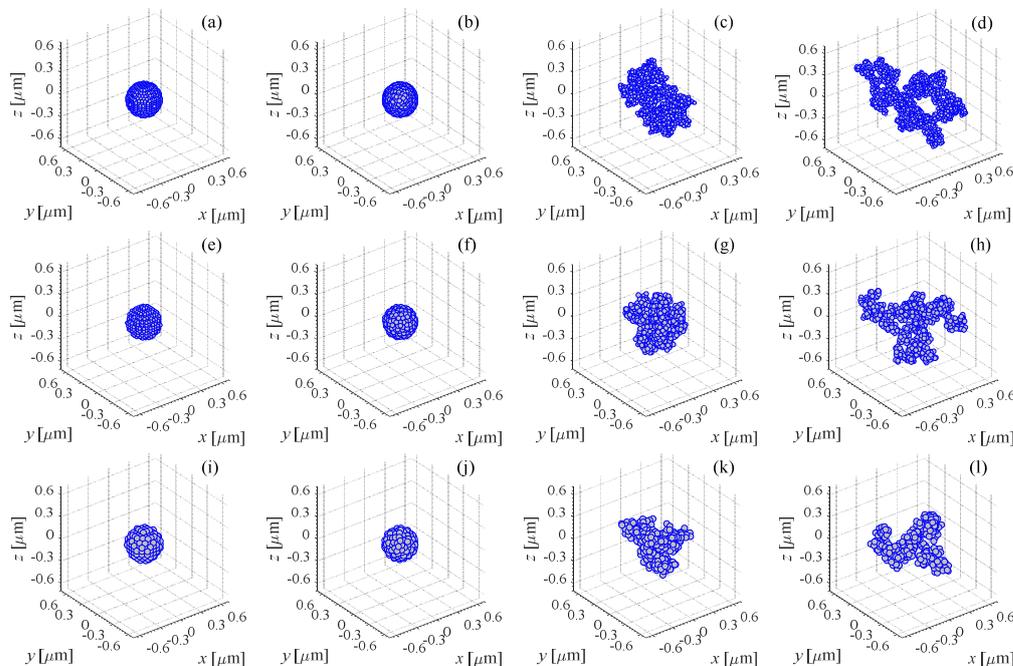
105 Next, the fractal aggregate generating code of Mackowski (1995; 2006) is employed to generate two more aggregates based on a sequential cluster-cluster aggregation (CCA) algorithm. One of these two fractal aggregates is more compact in order to mimic the COJ300 samples, while the second of these two fractal aggregates is more extended in order to mimic the R2500U samples. Once again, as with the SC and IAS aggregates, in the CCA aggregates, the primary particles may touch but do not overlap, and each aggregate is monodisperse with respect to its primary particles. The CCA aggregates have the same values  
110 of  $d_{pp}$  and  $N_{pp}$  as the SC and IAS aggregates but do *not* have the same outer envelope diameter (which in any case is not a meaningful diameter for such particles; see DeCarlo et al. (2004)). The more compact CCA aggregates have a significantly larger outer envelope and a significantly higher porosity than the SC and IAS aggregates. The more extended CCA aggregates have an outer envelope diameter that is even larger and a porosity that is even higher, as well as a more fractal appearance. (See, e.g., Figs. 2 and 3.)



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120 **Figure 2: Positions of the primary particles for: (a) aggregate SC\_300\_35, (b) aggregate IAS\_300\_35, (c) aggregate CCA\_2.34\_1.085\_300\_35, and (d) aggregate CCA\_300\_35, (e) aggregate SC\_400\_35, (f) aggregate IAS\_400\_35, (g) aggregate CCA\_2.34\_1.085\_400\_35, and (h) aggregate CCA\_1.92\_1.873\_400\_35, (i) aggregate SC\_500\_35, (j) aggregate IAS\_500\_35, (k) aggregate CCA\_2.34\_1.085\_500\_35, and (l) aggregate CCA\_500\_35. Note the different scale of the  $z$ -axis in plot (l). (Refer to Table 1 for the aggregate label scheme.)**



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**Figure 3: Positions of the primary particles for: (a) aggregate SC\_400\_25, (b) aggregate IAS\_400\_25, (c) aggregate CCA\_2.34\_1.085\_400\_25, and (d) aggregate CCA\_400\_25, (e) aggregate SC\_400\_35, (f) aggregate IAS\_400\_35, (g) aggregate CCA\_2.34\_1.085\_400\_35, and (h) aggregate CCA\_1.92\_1.873\_400\_35, (i) aggregate SC\_400\_45, (j) aggregate IAS\_400\_45, (k) aggregate CCA\_2.34\_1.085\_400\_45, and (l) aggregate CCA\_400\_45. Note the different scale of the  $x$ -axis in plot (d). (Refer to**

130 **Table 1 for the aggregate label scheme.)**

The input to the fractal aggregate generating code of Mackowski (1995; 2006) consists of the value of  $N_{pp}$ , the radius of the primary particle,  $a_{pp} = \frac{1}{2}d_{pp}$ , the 3D fractal dimension,  $D_f$ , and the fractal pre-factor, labelled here  $k_{Sorensen}$ , from the following relationship (Sorensen, 2001):

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$$N_{pp} = k_{Sorensen} \left( \frac{R_g}{a_{pp}} \right)^{D_f}, \quad (1)$$

where  $R_g$  is the radius of gyration. As mentioned in Sect. 1, for COJ300,  $D_f = 2.34$  with a 95% confidence interval range of 2.12-2.56, and for R2500U,  $D_f = 1.92$  with a 95% confidence interval range of 1.68-2.16 (Zhang et al., 2020). Regarding



the fractal pre-factor, by assuming that  $R_g = \frac{1}{3}L_{\max}$ , where  $L_{\max}$  is the length of longest dimension of the aggregate periphery, Zhang et al. (2020) wrote a similar relationship to that of Sorensen (2001):

$$140 \quad N_{pp} = k_{\text{Zhang et al.}} \left( \frac{L_{\max}}{d_{pp}} \right)^{D_f}, \quad (2)$$

where the fractal pre-factor as defined by Zhang et al. (2020) is labelled  $k_{\text{Zhang et al.}}$ . From this, we can obtain the following relationship between  $k_{\text{Sorensen}}$  and  $k_{\text{Zhang et al.}}$ :

$$k_{\text{Sorensen}} = k_{\text{Zhang et al.}} \times \left( \frac{2}{3} \right)^{-D_f}. \quad (3)$$

Based on the data from Zhang et al. (2020), e.g., as shown in their Fig. A6(c), for COJ300,  $k_{\text{Zhang et al.}} = 0.42$ , and for  
 145 R2500U,  $k_{\text{Zhang et al.}} = 0.86$ . By using Eq. (3), together with the respective values of  $D_f$  above, for COJ300, we obtain  
 $k_{\text{Sorensen}} = 1.085$ , and for R2500U, we obtain  $k_{\text{Sorensen}} = 1.873$ . Thus, the CCA aggregates with  $D_f = 2.34$  and  $k_{\text{Sorensen}} =$   
 $1.085$  appear more compact and mimic the COJ300 samples, while the CCA aggregates with  $D_f = 1.92$  and  $k_{\text{Sorensen}} =$   
 $1.873$  appear more extended and mimic the R2500U samples (refer again to Figs. 2 and 3). We also test different realizations  
 of these CCA aggregates, varying the values of  $D_f$  within the 95% confidence interval ranges stated above, as well as  
 150 different realizations of the IAS aggregates.

We note that the fractal aggregate generating code of Mackowski (1995; 2006) includes an option to generate aggregates  
 based on diffusion-limited particle-cluster aggregation (PCA). However, as discussed in Mackowski (1995; 2006) and in  
 Filippov et al. (2000), for given values of  $D_f$  and  $k_{\text{Sorensen}}$ , with the sequential CCA algorithm, Eq. (1) above is fulfilled  
 exactly at each step. Thus, the sequential CCA algorithm should generate more precise fractal aggregates. From preliminary  
 155 tests (not shown here), we find that on the whole, scattering calculations on aggregates generated using the sequential CCA  
 option better reproduce some of the tendencies in the measured results than scattering calculations on aggregates generated  
 with the PCA option. Thus, with respect to the fractal aggregates, by default, we present calculations for aggregates  
 generated using the sequential CCA option of the fractal aggregate generating code of Mackowski (1995; 2006). However,  
 when we vary the value of  $D_f$  to its highest value within the 95% confidence interval range of  $D_f$  for the COJ300 samples  
 160 (2.56), the sequential CCA algorithm gives repeated error messages of “clusters did not combine” and produces a list of  
 primary particle positions that partially overlap one another. Therefore, for this highest value of value of  $D_f$  only, we  
 employ the PCA option of the fractal aggregate generating code of Mackowski (1995; 2006).



Even though the SC and IAS aggregates are not expected to represent either of the Zhang et al. (2020) sample sets well, these two configurations are useful to test for two reasons: (1) By constructing the SC and IAS aggregates of a given aggregate set first, we can determine how many primary particles of a given value of  $d_{pp}$  fit compactly into sphere of a given value of  $D_{outer-envelope}$ . Then, as explained above, we use this same number of primary particles  $N_{pp}$  with the same  $d_{pp}$  to construct the CCA/PCA aggregates of the same set. By doing so, all of the aggregates of a given set possess the same mass equivalent diameter,  $D_{me}$ , but varying degrees of disorder, which allows us to isolate the effect of the degree of disorder of the primary particles, holding all other parameters constant. (2) Although the SC and IAS aggregates are the most spherical of each set, they have a roughness on the nanometer scale and are not perfectly symmetric. Thus, even the SC and IAS aggregates should provide a minimal perpendicularly polarized scattered intensity against which the perpendicularly polarized scattered intensity provided by the CCA/PCA aggregates can be compared.

In testing the sensitivity of the results to variations in  $D_{outer-envelope}$ , we hold  $d_{pp}$  constant at 35 nm and change the value of  $D_{outer-envelope}$  of the SC aggregate to 300 nm or 500 nm. In testing the sensitivity of the results to variations in  $d_{pp}$ , we hold the value of  $D_{outer-envelope}$  of the SC aggregate constant at 400 nm and change  $d_{pp}$  to 25 nm or 45 nm. Throughout the sensitivity studies, each individual aggregate is monodisperse with respect to its primary particles. The aggregate labels and aggregate parameters for the SC aggregate with  $D_{outer-envelope} = 300$  nm and  $d_{pp} = 35$  nm, for the SC aggregate with  $D_{outer-envelope} = 400$  nm and  $d_{pp} = 35$  nm, for the SC aggregate with  $D_{outer-envelope} = 500$  nm and  $d_{pp} = 35$  nm, and for the aggregates generated starting from an SC aggregate with  $D_{outer-envelope} = 400$  nm but with varying values of  $d_{pp}$  are given in Table 1. All other aggregates in this study are labelled similarly following this labelling scheme.

Aggregate label	Aggregate type	$N_{pp}$	$d_{pp}$	$D_f$	$k_{Sorensen}$
SC_300_35	SC	317	35	N/A	N/A
SC_400_35	SC	771	35	N/A	N/A
SC_500_35	SC	1529	35	N/A	N/A
SC_400_25	SC	2106	25	N/A	N/A
SC_400_35	SC	771	35	N/A	N/A
SC_400_45	SC	377	45	N/A	N/A
IAS_400_25	IAS	2106	25	N/A	N/A
IAS_400_35	IAS	771	35	N/A	N/A
IAS_400_45	IAS	377	45	N/A	N/A
CCA_2.34_1.085_400_25	CCA	2106	25	2.34	1.085
CCA_2.34_1.085_400_35	CCA	771	35	2.34	1.085
CCA_2.34_1.085_400_45	CCA	377	45	2.34	1.085



CCA_1.92_1.873_400_25	CCA	2106	25	1.92	1.873
CCA_1.92_1.873_400_35	CCA	771	35	1.92	1.873
CCA_1.92_1.873_400_45	CCA	377	45	1.92	1.873

**Table 1.** Aggregate labels and aggregate parameters for some of the aggregates in this study.

Once the aggregates are generated, we employ the MSTM model (Mackowski and Mishchenko, 1996) to calculate the extinction efficiency,  $Q_{\text{ext MSTM}}$ , the absorption efficiency,  $Q_{\text{abs MSTM}}$ , the scattering efficiency,  $Q_{\text{sca MSTM}}$ , and the normalized scattering phase function,  $P(\theta_{\text{sca}})$ , of the aggregate at the wavelength of measurement, 670 nm. The default value of the complex refractive index of BC at 670 nm is taken to be  $2.0+1.0i$  (Janzen, 1979; soot G of Fuller et al., 1999; Liu and Mishchenko, 2005, 2007; Liu et al., 2008; Moteki et al., 2010), where the real part,  $m_{\text{real}}$ , represents the refractive capability of the material, and the imaginary part,  $m_{\text{imag}}$ , represents the absorptive capability of the material, but the sensitivity to this choice is also investigated.

The random-orientation option of MSTM (Mackowski, 2013) is used as a proxy for averaging over many different realizations of each of the IAS and CCA aggregates (see, e.g., Mishchenko et al. (2007) for an explanation of this), but we also test the sensitivity of the results to the choice of realization, and we also conduct simulations with fixed orientation.

The intensity of parallel polarized scattered radiation for parallel polarized incident radiation,  $I_{\text{sca}\parallel\rightarrow\parallel}$ , is obtained from the elements of the  $4\times 4$  scattering matrix outputted from MSTM,  $\mathbf{S}$ , as:

$$I_{\text{sca}\parallel\rightarrow\parallel}(\theta_{\text{sca}}) = \frac{1}{2} \frac{k^2 \sigma_{\text{sca}}}{4\pi} \left[ (S_{11}(\theta_{\text{sca}}) + S_{12}(\theta_{\text{sca}})) + (S_{21}(\theta_{\text{sca}}) + S_{22}(\theta_{\text{sca}})) \right], \quad (4)$$

where  $k = 2\pi/\lambda$  is the wave number,  $\lambda$  is the wavelength, and  $\sigma_{\text{sca}}$  is the scattering cross section of the aggregate with respect to unpolarized incident radiation (see also Sect. 3.2 regarding the scattering cross section). Similarly, the intensity of perpendicularly polarized scattered radiation for parallel polarized incident radiation,  $I_{\text{sca}\parallel\rightarrow\perp}$ , is obtained from the

elements of  $\mathbf{S}$ , as:

$$I_{\text{sca}\parallel\rightarrow\perp}(\theta_{\text{sca}}) = \frac{1}{2} \frac{k^2 \sigma_{\text{sca}}}{4\pi} \left[ (S_{11}(\theta_{\text{sca}}) + S_{12}(\theta_{\text{sca}})) - (S_{21}(\theta_{\text{sca}}) + S_{22}(\theta_{\text{sca}})) \right], \quad (5)$$

and the total intensity of scattered radiation as a function of scattering angle is given by the sum,

$$I_{\text{sca tot}}(\theta_{\text{sca}}) = I_{\text{sca}\parallel\rightarrow\parallel}(\theta_{\text{sca}}) + I_{\text{sca}\parallel\rightarrow\perp}(\theta_{\text{sca}}). \quad (6)$$



In using MSTM and Eqs. (4)-(6), we implicitly assume that the incident laser light is a 100% coherent plane wave that is 100% polarized parallel to the scattering plane. The scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  is calculated as:

$$\begin{aligned}
 I_{\text{sca}\parallel\rightarrow\parallel}(135 \pm 20^\circ) &= \frac{k^2 \sigma_{\text{sca}}}{4\pi} \int_{115^\circ}^{155^\circ} d\mu_{\text{sca}} I_{\text{sca}\parallel\rightarrow\parallel}(\theta_{\text{sca}}) \\
 I_{\text{sca}\parallel\rightarrow\perp}(135 \pm 20^\circ) &= \frac{k^2 \sigma_{\text{sca}}}{4\pi} \int_{115^\circ}^{155^\circ} d\mu_{\text{sca}} I_{\text{sca}\parallel\rightarrow\perp}(\theta_{\text{sca}}), \\
 I_{\text{sca tot}}(135 \pm 20^\circ) &= \frac{k^2 \sigma_{\text{sca}}}{4\pi} \int_{115^\circ}^{155^\circ} d\mu_{\text{sca}} I_{\text{sca tot}}(\theta_{\text{sca}})
 \end{aligned} \tag{7}$$

where  $\mu_{\text{sca}} \equiv \cos \theta_{\text{sca}}$ . In implementing Eq. (7) numerically, for each discrete value of scattering zenith angle,  $\theta_{\text{sca}_i}$ ,  $d\mu_{\text{sca}}$  is calculated explicitly as  $\left| \cos(\theta_{\text{sca}_i} - 0.5^\circ) - \cos(\theta_{\text{sca}_i} + 0.5^\circ) \right|$ , i.e., with a span of  $1^\circ$ .

### 3 Results

#### 3.1 SPIN measurements

A summary of the scattering measurements from the SPIN OPC is given in Table 2.

Percentiles	COJ300					R2500U				
	5	25	50	75	95	5	25	50	75	95
P1 [photons s <sup>-1</sup> ]	205.0	527.0	716.0	924.0	1318.0	126.6	567.0	1187.0	2327.8	5538.8
P2 [photons s <sup>-1</sup> ]	167.0	436.0	612.0	808.0	1163.0	80.0	418.8	964.0	1866.0	4356.8
P=(P1+P2)/2 [photons s <sup>-1</sup> ]	237.5	506.0	676.0	849.0	1159.0	170.1	569.6	1147.0	2147.4	4675.9
S=S1 [photons s <sup>-1</sup> ]	0.0	0.0	0.0	33.0	118.0	0.0	26.0	231.0	767.2	2408.9
P+S [photons s <sup>-1</sup> ]	252.0	523.0	693.0	876.0	1218.0	201.8	672.4	1462.0	2946.4	6848.0
S/P*	0.0	0.0	0.0	0.039	0.102	0.0	0.046	0.201	0.357	0.515

\*Obtained by dividing the “S” percentiles respectively by the corresponding “P” percentiles.

**Table 2.** Percentiles of filtered photon counts from the SPIN instrumentation.

From Table 2, the photon counts for the more fractal sample set, R2500U, are significantly higher and exhibit more variation than the photon counts for the more spherical sample set, COJ300. This is true both for each polarization individually and for the total P+S. Thus, from the SPIN measurements, we find a stronger side/back scattering signal from the more fractal sample set.



Likewise, from Table 2, we see that the S/P ratio (also known as the linear depolarization ratio) for R2500U is significantly higher than the S/P ratio for COJ300. For the R2500U sample set, the median value of S/P is 0.201, with 25th and 75th percentile values of 0.046 and 0.357, respectively, and a 95th percentile value greater than 0.5. In contrast, for the COJ300 sample set, the median value of S/P is 0.0, with 25th and 75th percentile values of 0.0 and 0.039, respectively. Overall, from our measurements, more than half (~60.4%) of the COJ300 particles have undetectably low S scattering signals and therefore S/P values.

As mentioned in Sect. 2.1, a nonzero value of S when the incident radiation is polarized parallel to the scattering plane indicates some asymmetry/nonsphericity in the shape of the scattering particle (or possibly chirality or birefringence in the scattering particle material). Thus, the higher median S/P for the R2500U sample set corresponds with it being the more fractal sample set, exhibiting more irregular and extended shapes, while the zero median S/P ratio for the COJ300 sample set corresponds well with it being the less fractal sample set, exhibiting shapes that are closer to spherically symmetric. At the same time, the nonzero 75th and 95th percentile values of S/P for COJ300 indicate that some of the COJ300 particles are non-spherical, albeit less so than the R2500U, which also corresponds with the fact that the mean fractal dimension of the COJ300 samples ( $D_f = 2.34$ ) is lower than 3.

Below, we examine whether these tendencies are also reproduced in our theoretical calculations.

### 3.2 Theoretical calculations – Sensitivity to $D_{\text{outer-envelope}}$ of the SC aggregate

The results for the aggregates shown in Fig. 2 (i.e., for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm (our default set), and for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm) are given in Table 3, in Fig. 4, and in Table 4.

In columns 2, 4, and 6 of Table 3, the values of  $Q_{\text{ext MSTM}}$ ,  $Q_{\text{abs MSTM}}$ , and  $Q_{\text{sca MSTM}}$  of each aggregate as given by MSTM are presented. The efficiencies as given by MSTM are with respect to the volume-mean radius,  $R_{\text{volume-mean}}$ , which is the radius of a sphere that has the same ratio of volume to surface area. For a monodisperse aggregate:

$$R_{\text{volume-mean}} = \left( \sum_{i=1}^{N_{\text{pp}}} a_{\text{pp}}^3 \right)^{1/3} = \left( N_{\text{pp}} a_{\text{pp}}^3 \right)^{1/3} = N_{\text{pp}}^{1/3} a_{\text{pp}}, \quad (7)$$

such that:



$$Q_{\text{ext/abs/sca}} = Q_{\text{ext/abs/sca MSTM}} \frac{\pi (R_{\text{volume-mean}})^2}{\sigma_{\text{geometric}}} = Q_{\text{ext/abs/sca MSTM}} \frac{\pi (N_{\text{pp}}^{2/3} a_{\text{pp}}^2)}{\sigma_{\text{geometric}}}, \quad (8)$$

where  $\sigma_{\text{geometric}}$  is the actual geometric cross section of the aggregate. Accordingly, the extinction, absorption, and scattering cross sections, respectively ( $\sigma_{\text{ext}}$ ,  $\sigma_{\text{abs}}$ ,  $\sigma_{\text{sca}}$ ) are given by:

$$\sigma_{\text{ext/abs/sca}} = Q_{\text{ext/abs/sca}} \times \sigma_{\text{geometric}} = Q_{\text{ext/abs/sca MSTM}} \times \pi (N_{\text{pp}}^{2/3} a_{\text{pp}}^2). \quad (9)$$

The cross sections calculated based on Eq. (9) are given in columns 3, 5, and 7 of Table 3.

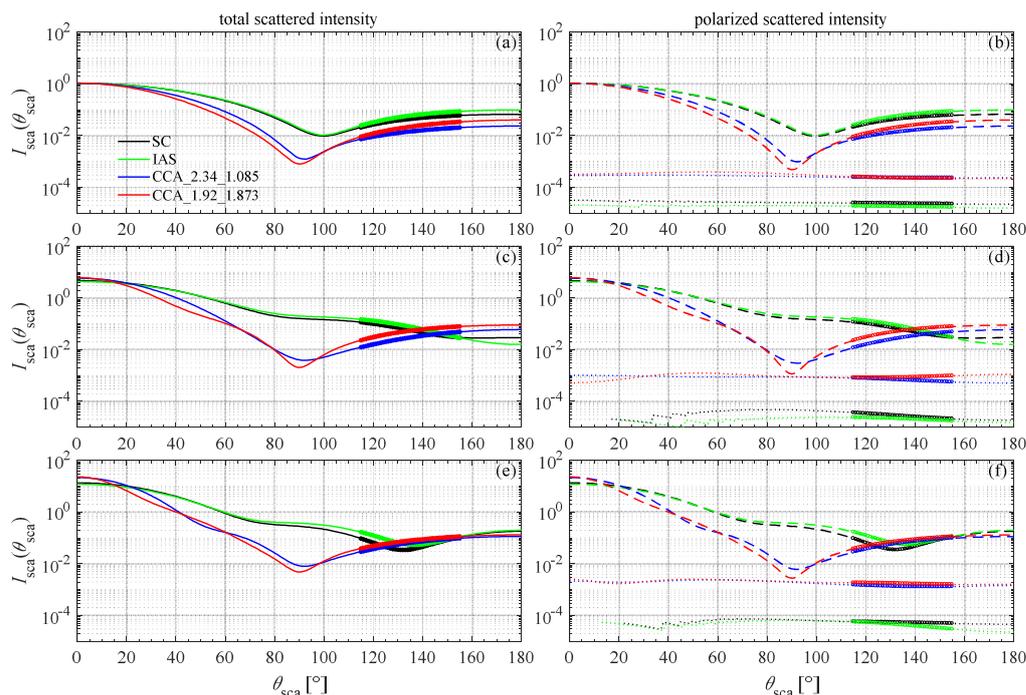
From Table 3, we see that the SC and IAS aggregates tend to have higher extinction, absorption, and scattering cross sections than the CCA aggregates in the same set. This agrees with the findings of Liu and Mishchenko (2005) and Liu et al. (2008), who found that the extinction and scattering cross sections of soot aggregates increase as the aggregates become more compact. (See also the review in Kahnert and Kanngießer (2020).) For aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm, and for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm, the SC aggregate exhibits the highest extinction cross section of the set, which corresponds with the results in Sela and Haspel (2021). For aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, the IAS aggregate exhibits the highest extinction cross section of the set.

Aggregate label	$Q_{\text{ext MSTM}}$	$\sigma_{\text{ext}}$ [m <sup>2</sup> ]	$Q_{\text{abs MSTM}}$	$\sigma_{\text{abs}}$ [m <sup>2</sup> ]	$Q_{\text{sca MSTM}}$	$\sigma_{\text{sca}}$ [m <sup>2</sup> ]
SC_300_35	2.741	$1.226 \times 10^{-13}$	1.792	$8.016 \times 10^{-14}$	0.949	$4.246 \times 10^{-14}$
IAS_300_35	2.780	$1.244 \times 10^{-13}$	1.764	$7.890 \times 10^{-14}$	1.016	$4.546 \times 10^{-14}$
CCA_2.34_1.085_300_35	2.157	$9.649 \times 10^{-14}$	1.629	$7.287 \times 10^{-14}$	0.528	$2.362 \times 10^{-14}$
CCA_1.92_1.873_300_35	2.069	$9.254 \times 10^{-14}$	1.639	$7.331 \times 10^{-14}$	0.430	$1.923 \times 10^{-14}$
SC_400_35	3.516	$2.844 \times 10^{-13}$	2.065	$1.671 \times 10^{-13}$	1.451	$1.174 \times 10^{-13}$
IAS_400_35	3.448	$2.789 \times 10^{-13}$	1.978	$1.600 \times 10^{-13}$	1.471	$1.190 \times 10^{-13}$
CCA_2.34_1.085_400_35	2.988	$2.417 \times 10^{-13}$	2.091	$1.692 \times 10^{-13}$	0.897	$7.257 \times 10^{-14}$
CCA_1.92_1.873_400_35	2.845	$2.301 \times 10^{-13}$	2.144	$1.734 \times 10^{-13}$	0.701	$5.667 \times 10^{-14}$
SC_500_35	3.981	$5.083 \times 10^{-13}$	2.193	$2.800 \times 10^{-13}$	1.788	$2.283 \times 10^{-13}$
IAS_500_35	3.805	$4.859 \times 10^{-13}$	2.076	$2.651 \times 10^{-13}$	1.728	$2.207 \times 10^{-13}$
CCA_2.34_1.085_500_35	3.776	$4.821 \times 10^{-13}$	2.515	$3.212 \times 10^{-13}$	1.260	$1.609 \times 10^{-13}$
CCA_1.92_1.873_500_35	3.612	$4.612 \times 10^{-13}$	2.614	$3.338 \times 10^{-13}$	0.998	$1.274 \times 10^{-13}$

**Table 3.** Values of extinction, absorption, and absorption efficiency and values of extinction, absorption, and scattering cross section for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, for



260 aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm, and for aggregates  
 generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm.



265 **Figure 4: Scattered intensity as a function of scattering angle as obtained from MSTM for aggregates generated starting with an**  
**SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, for aggregates generated starting with an SC aggregate with**  
 **$D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm (our default set), and for aggregates generated starting with an SC aggregate with**  
 **$D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm: (a),(b) aggregates SC\_300\_35, IAS\_300\_35, CCA\_2.34\_1.085\_300\_35, and**  
**CCA\_300\_35; (c),(d) aggregates SC\_400\_35, IAS\_400\_35, CCA\_2.34\_1.085\_400\_35, and CCA\_1.92\_1.873\_400\_35; (e),(f)**  
 270 **aggregates SC\_500\_35, IAS\_500\_35, CCA\_2.34\_1.085\_500\_35, and CCA\_500\_35. Plots (a), (c), and (e) contain the total scattered**  
**intensity (solid curves), while plots (b), (d), and (f) contain the scattered intensity polarized parallel (dashed curves) and**  
**perpendicular (dotted curves) to the scattering plane.**

In Figs. 4a and b, we show the scattered intensity as a function of scattering angle as obtained from MSTM for SC\_300\_35 (black curves), IAS\_300\_35 (green curves), CCA\_2.34\_1.085\_300\_35 (blue curves), and CCA\_1.92\_1.873\_300\_35 (red  
 275 curves), where the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  is highlighted on each curve with a thicker curve. In Fig. 4a, the total scattered  
 intensity is shown, while in Fig. 4b, the scattered intensity is separated according to polarization. From Fig. 4a, we see that  
 the SC aggregate exhibits a slightly higher scattered intensity in the direct forward direction ( $\theta_{\text{sca}} = 0^\circ$ ) than the IAS



aggregate, while the IAS aggregate exhibits a higher side/back scattered intensity than the SC aggregate, both of which agree with the results of Sela and Haspel (2021). In addition, the more extended CCA\_1.92\_1.873\_300\_35 aggregate exhibits a higher side/back scattered intensity than the more compact CCA\_2.34\_1.085\_300\_35 aggregate, which agrees with the SPIN measurements. However, the two CCA aggregates exhibit lower side/back scattered intensities than both the SC aggregate and the IAS aggregate, and this is due to the fact that the very compact SC and IAS aggregates exhibit higher scattering cross sections (refer to Table 3) and scatter more overall than the two CCA aggregates. Thus, we find that the continued increase in disorder/fractalness in going from the IAS aggregate to the CCA aggregates does not translate into a continued increase in side/back scatter for aggregates of this size.

From Fig. 4b, we see that, as expected, all of the aggregates exhibit significantly more parallel polarized scattered intensity (the same polarization as the incident radiation; dashed curves) than perpendicularly polarized scattered intensity (dotted curves). Also as expected, we see that the SC and IAS aggregates exhibit a minimal but nonzero perpendicularly polarized scattered intensity (dotted black curve and dotted green curve, respectively); as mentioned in Sect. 2.2, these two aggregates are the most spherical of each set but contain a roughness on the nanometer scale and are not perfectly symmetric. (Refer also to Figs. 2 and 3.) From Fig. 4b, we also see that as with the total intensity, the more extended fractal aggregate (CCA\_1.92\_1.873\_300\_35; red dashed curve) exhibits a higher parallel polarized side/back scattered intensity than the more compact fractal aggregate (CCA\_2.34\_1.085\_300\_35; blue dashed curve), but at the same time, these two CCA aggregates exhibit lower parallel polarized side/back scattered intensities than the SC aggregate (SC\_300\_35; black dashed curve) and the IAS aggregate (IAS\_300\_35; green dashed curve). Finally, from Fig. 4b, we see that at this size, the two CCA aggregates exhibit very similar perpendicularly polarized side/back scatter intensities to one another (dotted blue curve and dotted red curve, respectively).

In Figs. 4c and 4d, we show the scattered intensity as a function of scattering angle as obtained from MSTM for our default set of aggregates: SC\_400\_35, IAS\_400\_35, CCA\_2.34\_1.085\_400\_35, and CCA\_1.92\_1.873\_400\_35. In Fig. 4c, the total scattered intensity is shown, while in Fig. 4d, the scattered intensity is separated according to polarization. We can see that the tendencies exhibited in Figs. 4c and 4d are similar to the tendencies exhibited in Figs. 4a and 4b, but with several distinctions. (1) There is more of a difference in the side/back scattered intensity between the more extended CCA\_1.92\_1.873\_400\_35 aggregate and the more compact CCA\_2.34\_1.085\_400\_35 aggregate, with the more extended CCA\_1.92\_1.873\_400\_35 aggregate exhibiting a clearly higher total side/back scattered intensity, a clearly higher parallel polarized side/back scattered intensity, and a clearly higher perpendicularly polarized side/back scattered intensity than the more compact CCA\_2.34\_1.085\_400\_35 aggregate, which agrees with the results from the SPIN measurements. In fact, we find that in this way, this default set of aggregates mimics the results from the SPIN measurements better than any set of aggregates that we tested. (2) In the range  $\theta_{sca} = 135 \pm 20^\circ$ , the curves of parallel polarized scattered intensity for the two CCA curves cross the curves of parallel polarized scattered intensity for the SC and IAS aggregates, which means that their values in that range are more comparable to those of the SC and IAS aggregates.



In Figs. 4e and 4f, we show the scattered intensity as a function of scattering angle as obtained from MSTM for SC\_500\_35, IAS\_500\_35, CCA\_2.34\_1.085\_500\_35, and CCA\_1.92\_1.873\_500\_35. We can see that the tendencies exhibited in Figs. 4e and 4f are similar to the tendencies exhibited in Figs. 4a and 4b and in Figs. 4c and 4d, respectively. However, from Figs. 4e and 4f, we see that there is less of a difference in the side/back scattered intensity between the more extended CCA\_1.92\_1.873\_500\_35 aggregate and the more compact CCA\_2.34\_1.085\_500\_35 aggregate, as compared with the difference in side/back scattered intensity exhibited by our default set of CCA aggregates.

In Table 4, we list the values of scattered intensity over the range  $\theta_{sca} = 135 \pm 20^\circ$  corresponding to the curves in Fig. 4. From Table 4, we see the same tendencies as exhibited in Fig. 4 but now quantified. For example, the value of  $I_{sca \parallel \rightarrow \parallel} (135 \pm 20^\circ)$  is higher for the more extended CCA\_1.92\_1.873\_400\_35 aggregate ( $2.423 \times 10^{-2} \text{ W m}^{-2}$ ) than for the more compact CCA\_2.34\_1.085\_400\_35 aggregate ( $1.344 \times 10^{-2} \text{ W m}^{-2}$ ). Likewise, the value of  $I_{sca \parallel \rightarrow \perp} (135 \pm 20^\circ)$  is higher for the more extended CCA\_1.92\_1.873\_400\_35 aggregate ( $4.265 \times 10^{-4} \text{ W m}^{-2}$ ) than for the more compact CCA\_2.34\_1.085\_400\_35 aggregate ( $3.400 \times 10^{-4} \text{ W m}^{-2}$ ), and the value of  $I_{sca \text{ tot}} (135 \pm 20^\circ)$  is higher for the more extended CCA\_1.92\_1.873\_400\_35 aggregate ( $2.465 \times 10^{-2} \text{ W m}^{-2}$ ) than for the more compact CCA\_2.34\_1.085\_400\_35 aggregate ( $1.378 \times 10^{-2} \text{ W m}^{-2}$ ). These tendencies resemble the tendencies from the SPIN measurements.

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Aggregate label	$I_{sca \parallel \rightarrow \parallel} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	$I_{sca \parallel \rightarrow \perp} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Ratio of perpendicular to parallel $\frac{I_{sca \parallel \rightarrow \perp} (135 \pm 20^\circ)}{I_{sca \parallel \rightarrow \parallel} (135 \pm 20^\circ)}$	$I_{sca \text{ tot}} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]
SC_300_35	$2.022 \times 10^{-2}$	$1.174 \times 10^{-5}$	$5.808 \times 10^{-4}$	$2.023 \times 10^{-2}$
IAS_300_35	$2.686 \times 10^{-2}$	$8.778 \times 10^{-6}$	$3.268 \times 10^{-4}$	$2.687 \times 10^{-2}$
CCA_2.34_1.085_300_35	$6.571 \times 10^{-3}$	$1.159 \times 10^{-4}$	$1.764 \times 10^{-2}$	$6.687 \times 10^{-3}$
CCA_1.92_1.873_300_35	$9.985 \times 10^{-3}$	$1.139 \times 10^{-4}$	$1.141 \times 10^{-2}$	$1.010 \times 10^{-2}$
SC_400_35	$3.425 \times 10^{-2}$	$1.423 \times 10^{-5}$	$4.155 \times 10^{-4}$	$3.426 \times 10^{-2}$
IAS_400_35	$4.444 \times 10^{-2}$	$1.054 \times 10^{-5}$	$2.371 \times 10^{-4}$	$4.445 \times 10^{-2}$
CCA_2.34_1.085_400_35	$1.344 \times 10^{-2}$	$3.400 \times 10^{-4}$	$2.530 \times 10^{-2}$	$1.378 \times 10^{-2}$
CCA_1.92_1.873_400_35	$2.423 \times 10^{-2}$	$4.265 \times 10^{-4}$	$1.761 \times 10^{-2}$	$2.465 \times 10^{-2}$
SC_500_35	$2.882 \times 10^{-2}$	$2.656 \times 10^{-5}$	$9.215 \times 10^{-4}$	$2.885 \times 10^{-2}$
IAS_500_35	$4.416 \times 10^{-2}$	$2.199 \times 10^{-5}$	$4.979 \times 10^{-4}$	$4.419 \times 10^{-2}$
CCA_2.34_1.085_500_35	$2.938 \times 10^{-2}$	$6.699 \times 10^{-4}$	$2.280 \times 10^{-2}$	$3.005 \times 10^{-2}$
CCA_1.92_1.873_500_35	$3.573 \times 10^{-2}$	$8.581 \times 10^{-4}$	$2.402 \times 10^{-2}$	$3.659 \times 10^{-2}$



**Table 4.** Scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm, and for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm.

However, referring to Table 2, the relative differences in scattered photon counts between the R2500U samples and the COJ300 samples are larger than the relative differences in scattered intensity between the CCA\_1.92\_1.873\_400\_35 aggregate and the CCA\_2.34\_1.085\_400\_35 aggregate shown in Table 4. For example, from the values in Table 2, the ratio of the median value of P+S for R2500U to the median value of P+S for COJ300 is 2.11, whereas from the values in Table 4, the ratio of the value of  $I_{\text{sca tot}}(135 \pm 20^\circ)$  for CCA\_1.92\_1.873\_400\_35 to the value of  $I_{\text{sca tot}}(135 \pm 20^\circ)$  for CCA\_2.34\_1.085\_400\_35 is 1.79. In addition, referring to Table 2, the highest ratio of perpendicularly polarized scattered intensity to parallel polarized scattered intensity listed, i.e., the value of S/P corresponding to the 95th percentile, is 0.055 for the COJ300 sample set and is 0.375 for the R2500U sample set; both of these values of S/P are higher than the values of  $\frac{I_{\text{sca} \parallel \rightarrow \perp}(135 \pm 20^\circ)}{I_{\text{sca} \parallel \rightarrow \parallel}(135 \pm 20^\circ)}$  for the CCA samples in Table 4, which range from  $1.141 \times 10^{-2}$  to  $2.530 \times 10^{-2}$ . This indicates that there were some samples measured in the SPIN measurements, especially in the R2500U sample set, that exhibit higher linear depolarization ratios than the theoretical aggregates shown in Fig. 2.

### 3.3 Theoretical calculations – Sensitivity to $d_{\text{pp}}$

Results for SC\_400\_25, IAS\_400\_25, CCA\_2.34\_1.085\_400\_25, and CCA\_1.92\_1.873\_400\_25, i.e., for aggregates generated starting with an SC aggregate with our default value of  $D_{\text{outer-envelope}} = 400$  nm but with a smaller primary particle diameter of  $d_{\text{pp}} = 25$  nm, and results for SC\_400\_45, IAS\_400\_45, CCA\_2.34\_1.085\_400\_45, and CCA\_1.92\_1.873\_400\_45, i.e., for aggregates generated starting with an SC aggregate with our default value of  $D_{\text{outer-envelope}} = 400$  nm but with a larger primary particle diameter of  $d_{\text{pp}} = 45$  nm (refer to Table 1) are shown in Table 5, in Fig. 5, and in Table 6. The tendencies shown in Table 5, in Figs. 5a and 5b, in Figs. 5e and 5f, and in Table 6 are similar to those for our default set of aggregates. (Note that the scattered intensity as a function of scattering angle for our default set of aggregates from Figs. 4c and 4d are repeated as Figs. 5c and 5d for ease of comparison.) However, again, there is less of a difference in the side/back scattered intensity between the more extended CCA\_1.92\_1.873\_400\_25 aggregate and the more compact CCA\_2.34\_1.085\_400\_25 aggregate, as compared with the difference exhibited by our default set of aggregates. Likewise, there is less of a difference in the side/back scattered intensity between the more extended



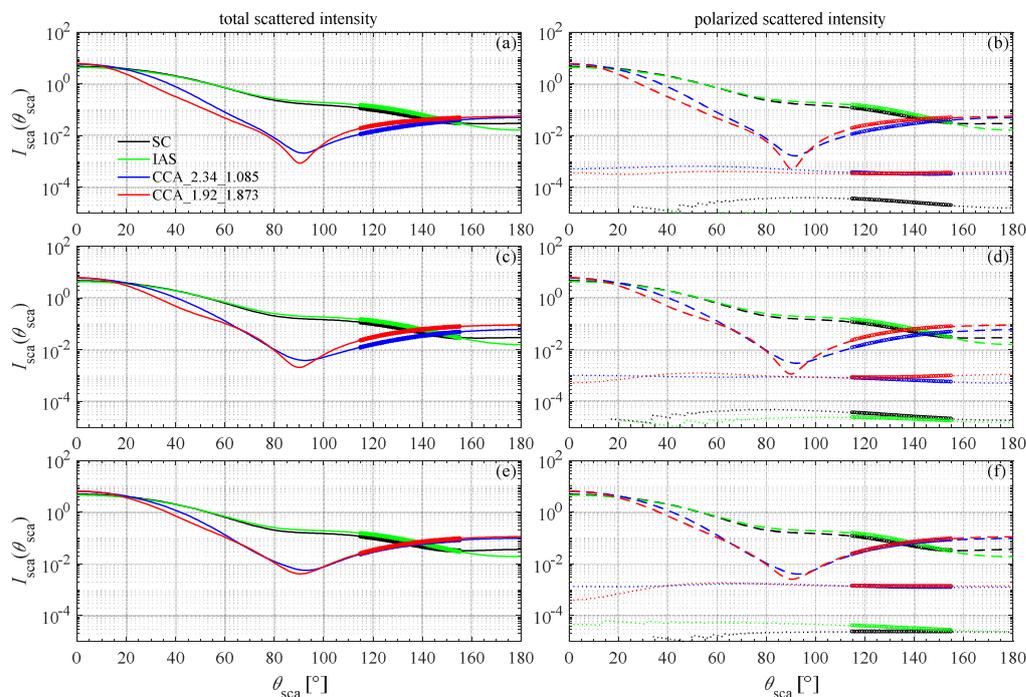
CCA\_1.92\_1.873\_400\_45 aggregate and the more compact CCA\_2.34\_1.085\_400\_45 aggregate, as compared with the difference in side/back scattered intensity exhibited by our default set of CCA aggregates. Thus, once again, we find that our default set of aggregates mimics the results from the SPIN measurements better than any set of aggregates that we tested.

Aggregate label	$Q_{\text{ext}}$ MSTM	$\sigma_{\text{ext}}$ [m <sup>2</sup> ]	$Q_{\text{abs}}$ MSTM	$\sigma_{\text{abs}}$ [m <sup>2</sup> ]	$Q_{\text{sca}}$ MSTM	$\sigma_{\text{sca}}$ [m <sup>2</sup> ]
SC_400_25	3.436	$2.771 \times 10^{-13}$	1.939	$1.564 \times 10^{-13}$	1.496	$1.207 \times 10^{-13}$
IAS_400_25	3.434	$2.769 \times 10^{-13}$	1.950	$1.573 \times 10^{-13}$	1.484	$1.197 \times 10^{-13}$
CCA_2.34_1.085_400_25	2.787	$2.248 \times 10^{-13}$	2.018	$1.627 \times 10^{-13}$	0.769	$6.205 \times 10^{-14}$
CCA_1.92_1.873_400_25	2.591	$2.089 \times 10^{-13}$	2.064	$1.664 \times 10^{-13}$	0.527	$4.251 \times 10^{-14}$
SC_400_45	3.544	$2.942 \times 10^{-13}$	2.075	$1.723 \times 10^{-13}$	1.468	$1.219 \times 10^{-13}$
IAS_400_45	3.476	$2.885 \times 10^{-13}$	1.991	$1.653 \times 10^{-13}$	1.485	$1.232 \times 10^{-13}$
CCA_2.34_1.085_400_45	3.122	$2.591 \times 10^{-13}$	2.130	$1.768 \times 10^{-13}$	0.992	$8.229 \times 10^{-14}$
CCA_1.92_1.873_400_45	3.011	$2.499 \times 10^{-13}$	2.184	$1.813 \times 10^{-13}$	0.826	$6.859 \times 10^{-14}$

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**Table 5.** Values of extinction, absorption, and absorption efficiency and values of extinction, absorption, and scattering cross section for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 25$  nm and 45 nm, respectively.

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**Figure 5: Scattered intensity as a function of scattering angle as obtained from MSTM for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 300$  nm and  $d_{\text{pp}} = 35$  nm, for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm (our default set), and for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 500$  nm and  $d_{\text{pp}} = 35$  nm: (a),(b) aggregates SC\_400\_25, IAS\_400\_25, CCA\_2.34\_1.085\_400\_25, and CCA\_400\_25; (c),(d) aggregates SC\_400\_35, IAS\_400\_35, CCA\_2.34\_1.085\_400\_35, and CCA\_1.92\_1.873\_400\_35; (e),(f) aggregates SC\_400\_45, IAS\_400\_45, CCA\_2.34\_1.085\_400\_45, and CCA\_400\_45. (Refer to Table 1 for the aggregate labels.) Plots (a), (c), and (e) contain the total scattered intensity (solid curves), while plots (b), (d), and (f) contain the scattered intensity polarized parallel (dashed curves) and perpendicular (dotted curves) to the scattering plane.**

Aggregate label	$I_{\text{sca}\parallel\rightarrow\parallel}(135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	$I_{\text{sca}\parallel\rightarrow\perp}(135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Ratio of perpendicular to parallel $\frac{I_{\text{sca}\parallel\rightarrow\perp}(135 \pm 20^\circ)}{I_{\text{sca}\parallel\rightarrow\parallel}(135 \pm 20^\circ)}$	$I_{\text{sca tot}}(135 \pm 20^\circ)$ [W m <sup>-2</sup> ]
SC_400_25	$3.418 \times 10^{-2}$	$1.394 \times 10^{-5}$	$4.077 \times 10^{-4}$	$3.420 \times 10^{-2}$
IAS_400_25	$4.513 \times 10^{-2}$	$3.513 \times 10^{-6}$	$7.785 \times 10^{-5}$	$4.513 \times 10^{-2}$
CCA_2.34_1.085_400_25	$1.142 \times 10^{-2}$	$1.644 \times 10^{-4}$	$1.440 \times 10^{-2}$	$1.158 \times 10^{-2}$
CCA_1.92_1.873_400_25	$1.662 \times 10^{-2}$	$1.652 \times 10^{-4}$	$9.939 \times 10^{-3}$	$1.678 \times 10^{-2}$



SC_400_45	$3.443 \times 10^{-2}$	$1.166 \times 10^{-5}$	$3.386 \times 10^{-4}$	$3.445 \times 10^{-2}$
IAS_400_45	$4.542 \times 10^{-2}$	$1.703 \times 10^{-5}$	$3.750 \times 10^{-4}$	$4.544 \times 10^{-2}$
CCA_2.34_1.085_400_45	$2.537 \times 10^{-2}$	$6.307 \times 10^{-4}$	$2.486 \times 10^{-2}$	$2.600 \times 10^{-2}$
CCA_1.92_1.873_400_45	$2.774 \times 10^{-2}$	$6.838 \times 10^{-4}$	$2.465 \times 10^{-2}$	$2.842 \times 10^{-2}$

**Table 6.** Scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 25$  nm and 45 nm, respectively.

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### 3.4 Theoretical calculations – Sensitivity to complex refractive index

To test the sensitivity of our results to the assumed complex refractive index, we repeat the calculations on our default aggregates (aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm) with three additional complex refractive indices that have been tabulated for non-graphitic light absorbing carbon: (1)  $m=1.75+0.63i$ , the lowest complex refractive index from Table 5 of Bond and Bergstrom (2006); (2)  $m=1.85+0.71i$ , the complex refractive index in the middle of the range from Table 5 of Bond and Bergstrom (2006) and that adopted by Bond et al. (2006); and (3)  $m=2.26+1.26i$ , the complex refractive index retrieved by Moteki et al. (2010). These complex refractive indices were not necessarily tabulated at the identical wavelength of 670 nm, but they bracket a reasonable range of possible values of refractive indices of black carbon at wavelengths relevant to incident solar radiation (500-1064 nm) (Janzen, 1979; Fuller et al., 1999; Bond and Bergstrom, 2006; Bond et al., 2006; Liu and Mishchenko, 2005, 2007; Liu et al., 2008; Moteki et al., 2010). (See also the review in Kahnert and Kanngießer (2020).) For this sensitivity test, each of these three additional complex refractive indices in turn are set to be the complex refractive index of the primary particles in the aggregate. Results for our default aggregates, SC\_400\_35, IAS\_400\_35, CCA\_2.34\_1.085\_400\_35, and CCA\_1.92\_1.873\_400\_35, but with primary particle complex refractive indices of  $m=1.75+0.63i$ ,  $m=1.85+0.71i$ ,  $m=2.26+1.26i$ , respectively, are shown in Table 7, in Fig. 6, and in Table 8.

From Table 7, we see that the higher the complex refractive index of the primary particles, the higher the extinction, absorption, and scattering cross sections of the aggregates, respectively, as would be expected. (See, also, Liu et al. (2008).) From Fig. 6, we see that the higher the complex refractive index, the farther towards the end of the  $\theta_{\text{sca}} = 135 \pm 20^\circ$  range the curves of parallel polarized scattered intensity for the two CCA curves cross the curves of parallel polarized scattered intensity for the SC and IAS aggregates. From Table 8, we also see that the ratio of perpendicularly polarized to parallel polarized scattered radiation in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  increases with the magnitude of the refractive index for the CCA aggregates, which agrees with the findings of Bescond et al. (2013) regarding the direct backscatter depolarization caused by BC aggregates. Aside from that, the tendencies shown in Table 7, in Fig. 6, and in Table 8 are quite similar to those for our default set of aggregates.

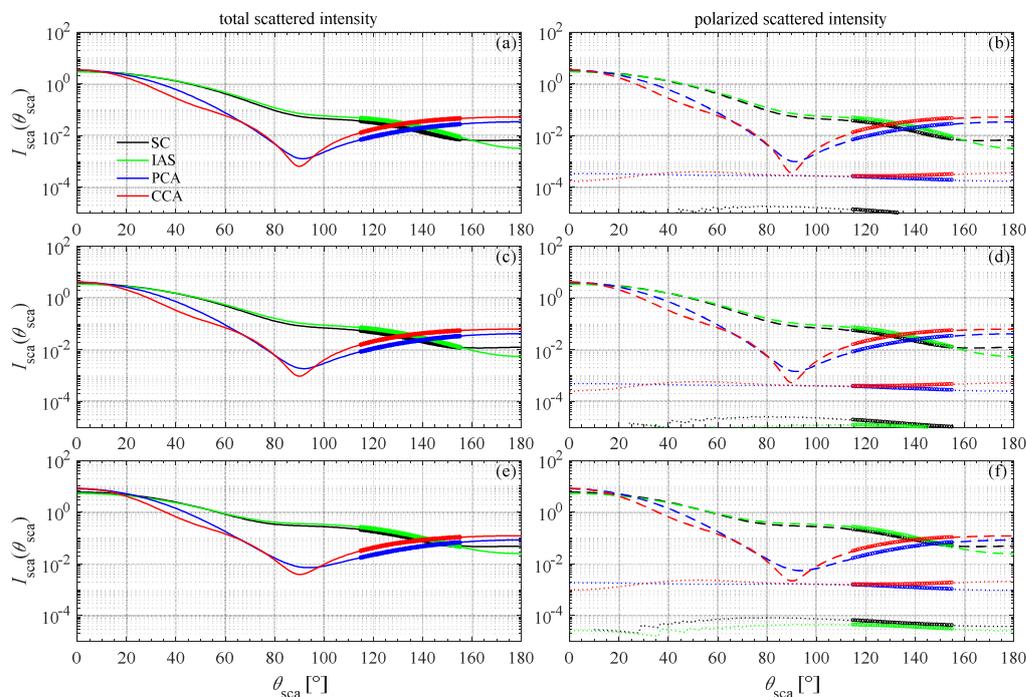


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Aggregate label	$Q_{\text{ext}}$ MSTM	$\sigma_{\text{ext}}$ [m <sup>2</sup> ]	$Q_{\text{abs}}$ MSTM	$\sigma_{\text{abs}}$ [m <sup>2</sup> ]	$Q_{\text{sca}}$ MSTM	$\sigma_{\text{sca}}$ [m <sup>2</sup> ]
SC_400_35 $m=1.75+0.63i$	2.627	$2.125 \times 10^{-13}$	1.659	$1.342 \times 10^{-13}$	0.968	$7.831 \times 10^{-14}$
IAS_400_35 $m=1.75+0.63i$	2.650	$2.144 \times 10^{-13}$	1.629	$1.318 \times 10^{-13}$	1.021	$8.260 \times 10^{-14}$
CCA_2.34_1.085_400_35 $m=1.75+0.63i$	2.148	$1.737 \times 10^{-13}$	1.626	$1.315 \times 10^{-13}$	0.522	$4.221 \times 10^{-14}$
CCA_1.92_1.873_400_35 $m=1.75+0.63i$	2.060	$1.666 \times 10^{-13}$	1.660	$1.343 \times 10^{-13}$	0.400	$3.236 \times 10^{-14}$
SC_400_35 $m=1.85+0.71i$	2.912	$2.356 \times 10^{-13}$	1.770	$1.432 \times 10^{-13}$	1.143	$9.243 \times 10^{-14}$
IAS_400_35 $m=1.85+0.71i$	2.923	$2.365 \times 10^{-13}$	1.730	$1.399 \times 10^{-13}$	1.193	$9.653 \times 10^{-14}$
CCA_2.34_1.085_400_35 $m=1.85+0.71i$	2.348	$1.899 \times 10^{-13}$	1.718	$1.390 \times 10^{-13}$	0.629	$5.090 \times 10^{-14}$
CCA_1.92_1.873_400_35 $m=1.85+0.71i$	2.237	$1.809 \times 10^{-13}$	1.753	$1.418 \times 10^{-13}$	0.484	$3.914 \times 10^{-14}$
SC_400_35 $m=2.26+1.26i$	4.056	$3.281 \times 10^{-13}$	2.225	$1.800 \times 10^{-13}$	1.831	$1.481 \times 10^{-13}$
IAS_400_35 $m=2.26+1.26i$	3.898	$3.154 \times 10^{-13}$	2.104	$1.702 \times 10^{-13}$	1.794	$1.451 \times 10^{-13}$
CCA_2.34_1.085_400_35 $m=2.26+1.26i$	3.419	$2.766 \times 10^{-13}$	2.201	$1.781 \times 10^{-13}$	1.218	$9.852 \times 10^{-14}$
CCA_1.92_1.873_400_35 $m=2.26+1.26i$	3.196	$2.586 \times 10^{-13}$	2.242	$1.814 \times 10^{-13}$	0.954	$7.721 \times 10^{-14}$

**Table 7.** Values of extinction, absorption, and absorption efficiency and values of extinction, absorption, and scattering cross section for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm with varying primary particle refractive indices.

410



415 **Figure 6:** (a),(b) Scattered intensity as a function of scattering angle as obtained from MSTM for our default aggregates SC\_400\_35, IAS\_400\_35, CCA\_2.34\_1.085\_400\_35, and CCA\_1.92\_1.873\_400\_35: (a),(b) with primary particle refractive index  $m=1.75+0.63i$ ; (c),(d) with primary particle refractive index  $m=1.85+0.71i$ ; (e),(f) with primary particle refractive index  $m=2.26+1.26i$ . (Refer to Table 1 for the aggregate labels.) Plots (a), (c), and (e) contain the total scattered intensity (solid curves), while plots (b), (d), and (f) contain the scattered intensity polarized parallel (dashed curves) and perpendicular (dotted curves) to the scattering plane.

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Aggregate label	$I_{\text{sca}\parallel\rightarrow\parallel}(135\pm 20^\circ)$ [W m <sup>-2</sup> ]	$I_{\text{sca}\parallel\rightarrow\perp}(135\pm 20^\circ)$ [W m <sup>-2</sup> ]	Ratio of perpendicular to parallel $\frac{I_{\text{sca}\parallel\rightarrow\perp}(135\pm 20^\circ)}{I_{\text{sca}\parallel\rightarrow\parallel}(135\pm 20^\circ)}$	$I_{\text{sca tot}}(135\pm 20^\circ)$ [W m <sup>-2</sup> ]
SC_400_35 $m=1.75+0.63i$	$1.054\times 10^{-2}$	$5.121\times 10^{-6}$	$4.861\times 10^{-4}$	$1.054\times 10^{-2}$
IAS_400_35 $m=1.75+0.63i$	$1.469\times 10^{-2}$	$3.760\times 10^{-6}$	$2.560\times 10^{-4}$	$1.469\times 10^{-2}$
CCA_2.34_1.085_400_35 $m=1.75+0.63i$	$7.719\times 10^{-3}$	$1.098\times 10^{-4}$	$1.422\times 10^{-2}$	$7.829\times 10^{-3}$
CCA_1.92_1.873_400_35 $m=1.75+0.63i$	$1.403\times 10^{-2}$	$1.354\times 10^{-4}$	$9.652\times 10^{-3}$	$1.416\times 10^{-2}$



SC_400_35				
$m=1.85+0.7li$	$1.590 \times 10^{-2}$	$7.240 \times 10^{-6}$	$4.554 \times 10^{-4}$	$1.590 \times 10^{-2}$
IAS_400_35				
$m=1.85+0.7li$	$2.131 \times 10^{-2}$	$5.424 \times 10^{-6}$	$2.546 \times 10^{-4}$	$2.131 \times 10^{-2}$
CCA_2.34_1.085_400_35				
$m=1.85+0.7li$	$9.361 \times 10^{-3}$	$1.612 \times 10^{-4}$	$1.722 \times 10^{-2}$	$9.522 \times 10^{-3}$
CCA_1.92_1.873_400_35				
$m=1.85+0.7li$	$1.692 \times 10^{-2}$	$1.991 \times 10^{-4}$	$1.177 \times 10^{-2}$	$1.712 \times 10^{-2}$
SC_400_35				
$m=2.26+1.26i$	$6.122 \times 10^{-2}$	$2.603 \times 10^{-5}$	$4.252 \times 10^{-4}$	$6.125 \times 10^{-2}$
IAS_400_35				
$m=2.26+1.26i$	$7.623 \times 10^{-2}$	$1.885 \times 10^{-5}$	$2.473 \times 10^{-4}$	$7.625 \times 10^{-2}$
CCA_2.34_1.085_400_35				
$m=2.26+1.26i$	$1.863 \times 10^{-2}$	$6.415 \times 10^{-4}$	$3.443 \times 10^{-2}$	$1.927 \times 10^{-2}$
CCA_1.92_1.873_400_35				
$m=2.26+1.26i$	$3.302 \times 10^{-2}$	$8.047 \times 10^{-4}$	$2.437 \times 10^{-2}$	$3.382 \times 10^{-2}$

**Table 8.** Scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm with different primary particle refractive indices.

### 425 3.5 Theoretical calculations – Sensitivity to realizations of aggregate generating algorithms

Liu and Mishchenko (2007) found that varying the geometrical configuration of the primary particles in a soot cluster for fixed values of  $D_f$ ,  $k_{\text{Sorensen}}$ ,  $N_{\text{pp}}$ , and  $d_{\text{pp}}$  has a weak effect on scattering and absorption in the visible part of the spectrum. Here we test the sensitivity to aggregate realization in a similar vein, but with respect to the Zhang et al. (2020) experimental configuration and our associated theoretical aggregate parameters. We test additional realizations of the IAS  
430 aggregate of default size (IAS\_400\_35), the more compact CCA aggregate of default size (CCA\_2.34\_1.085\_400\_35), and the more extended CCA aggregate of default size (CCA\_1.92\_1.873\_400\_35), respectively, where the values of  $N_{\text{pp}}$  (771; refer to Table 1) and  $d_{\text{pp}}$  (35 nm) are identical for all of the realizations. First, we create two additional realizations of IAS\_400\_35, five additional realizations of CCA\_2.34\_1.085\_400\_35, and five additional realizations of CCA\_1.92\_1.873\_400\_35, all with identical realization parameters to the original aggregates, respectively. Then we create  
435 six additional realizations of the more compact CCA aggregate with the minimum value of  $D_f$  within the 95% confidence interval range mentioned in Sects. 1 and 2.2 for the COJ300 samples ( $D_f = 2.12$ ; CCA\_2.12\_0.992\_400\_35), six additional realizations of the more compact CCA aggregate with the maximum value of  $D_f$  within the 95% confidence interval range



for the COJ300 samples ( $D_f = 2.56$ ; PCA\_2.56\_1.186\_400\_35), six additional realizations of the more extended CCA aggregate with the minimum value of  $D_f$  within the 95% confidence interval range for the R2500U samples ( $D_f = 1.68$ ; 440 CCA\_1.68\_1.700\_400\_35), and six additional realizations of the more extended CCA aggregate with the maximum value of  $D_f$  within the 95% confidence interval range for the R2500U samples ( $D_f = 2.16$ ; CCA\_2.16\_2.065\_400\_35). In the deriving the 95% confident interval range of  $D_f$ , the regression parameter  $k_{\text{Zhang et al.}}$  was held constant. Accordingly, for each new value of  $D_f$ , a new value of  $k_{\text{Sorensen}}$  is calculated from Eq. (3) using the value of  $k_{\text{Zhang et al.}}$  for the corresponding sample set. Note that as stated in Sect. 2.2, when we vary the value of  $D_f$  to its highest value within the 95% 445 confidence interval range for the COJ300 samples, which is  $D_f = 2.56$ , we employ the PCA option of the fractal aggregate generating algorithm rather than the CCA option, while for all of the other realizations, we employ the CCA option of the fractal aggregate generating algorithm.

In Table 9, we list the ranges of values of extinction, absorption, and scattering cross section for all of the realizations of the aggregate generating algorithms. From Table 9, we see again that the most compact aggregates (SC, IAS, 450 PCA\_2.56\_1.186\_400\_35, and CCA\_2.16\_2.065\_400\_35) tend to have the highest extinction cross sections and scattering cross sections, which again agrees with the results of Liu and Mishchenko (2005) and Liu et al. (2008).

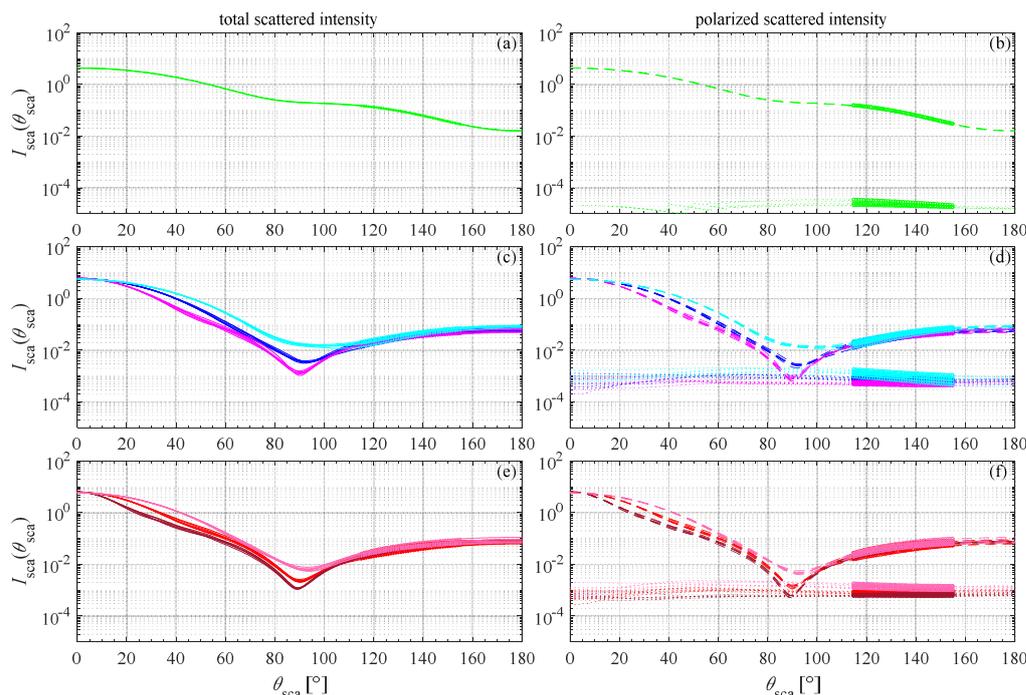
Aggregate label	Range of $\sigma_{\text{ext}}$ values [m <sup>2</sup> ]	Range of $\sigma_{\text{abs}}$ values [m <sup>2</sup> ]	Range of $\sigma_{\text{sca}}$ values [m <sup>2</sup> ]
SC_400_35	$2.844 \times 10^{-13}$	$1.671 \times 10^{-13}$	$1.174 \times 10^{-13}$
IAS_400_35	$(2.788 - 2.790) \times 10^{-13}$	$(1.598 - 1.600) \times 10^{-13}$	$(1.190 - 1.190) \times 10^{-13}$
CCA_2.34_1.085_400_35	$(2.404 - 2.417) \times 10^{-13}$	$(1.684 - 1.695) \times 10^{-13}$	$(7.144 - 7.290) \times 10^{-14}$
CCA_2.12_0.992_400_35	$(2.241 - 2.258) \times 10^{-13}$	$(1.733 - 1.747) \times 10^{-13}$	$(4.975 - 5.202) \times 10^{-14}$
PCA_2.56_1.186_400_35	$(2.600 - 2.635) \times 10^{-13}$	$(1.682 - 1.689) \times 10^{-13}$	$(9.186 - 9.470) \times 10^{-14}$
CCA_1.92_1.873_400_35	$(2.287 - 2.321) \times 10^{-13}$	$(1.727 - 1.742) \times 10^{-13}$	$(5.551 - 5.932) \times 10^{-14}$
CCA_1.68_1.700_400_35	$(2.152 - 2.180) \times 10^{-13}$	$(1.784 - 1.791) \times 10^{-13}$	$(3.618 - 3.961) \times 10^{-14}$
CCA_2.16_2.065_400_35	$(2.500 - 2.531) \times 10^{-13}$	$(1.700 - 1.710) \times 10^{-13}$	$(7.950 - 8.309) \times 10^{-14}$

**Table 9.** Ranges of values of extinction, absorption, and absorption cross section for different realizations of aggregates 455 generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm. The values for the single SC realization from Table 3 are also included here for reference.

In Fig. 7, we show the scattered intensity as a function of scattering angle as obtained from MSTM for all of the realizations of the aggregate generating algorithms. From Figs. 7a and b, we see that there is hardly any discernable difference in the



460 scattering patterns of the three IAS realizations (green curves), with just a small amount of discernable spread only in the very low values of perpendicularly polarized scattered intensity.



465 **Figure 7: Scattered intensity as a function of scattering angle as obtained from MSTM for: (a),(b) different realizations of IAS\_400\_35 (green curves); (c),(d) different realizations of CCA\_2.34\_1.085\_400\_35 (magenta curves:  $D_f = 2.12$ ,  $k_{\text{Sorensen}} = 0.992$ ; blue curves:  $D_f = 2.34$ ,  $k_{\text{Sorensen}} = 1.085$ ; cyan curves: PCA rather than CCA;  $D_f = 2.56$ ,  $k_{\text{Sorensen}} = 1.186$ ); (e),(f) different realizations of CCA\_1.92\_1.873\_400\_35 (maroon curves:  $D_f = 1.68$ ,  $k_{\text{Sorensen}} = 1.700$ ; red curves:  $D_f = 1.92$ ,  $k_{\text{Sorensen}} = 1.873$ ; pink curves:  $D_f = 2.16$ ,  $k_{\text{Sorensen}} = 2.065$ ). (Refer to Table 1 for the aggregate labels.) Plots (a), (c), and (e) contain the total scattered intensity (solid curves), while plots (b), (d), and (f) contain the scattered intensity polarized parallel (dashed curves) and perpendicular (dotted curves) to the scattering plane.**

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From Figs. 7c-f, where each color represents a group of realizations of fractal aggregates with identical values of  $D_f$ ,  $k_{\text{Sorensen}}$ ,  $N_{\text{pp}}$ , and  $d_{\text{pp}}$ , we see that as Liu and Mishchenko (2007) found, there is indeed a similarity to the scattering patterns of each group of curves. However, at the same time, there is some discernable spread in the scattering patterns, including in the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$ . From Figs. 7c-f, we see that *within each graph*, the more compact the fractal aggregate, the higher the values of scattered intensity over nearly the entire range of scattering angles, including over the

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range  $\theta_{\text{sca}} = 135 \pm 20^\circ$ . In Figs. 7c and 7d, which are different realizations of the more compact fractal aggregate (CCA\_2.34\_1.085\_400\_35), the cyan curves ( $D_f = 2.56$ ,  $k_{\text{Sorensen}} = 1.186$ ) lie largely above the blue curves ( $D_f = 2.34$ ,  $k_{\text{Sorensen}} = 1.085$ ), which in turn lie largely above the magenta curves ( $D_f = 2.12$ ,  $k_{\text{Sorensen}} = 0.992$ ). In Figs. 7e and 7f, 480 which are different realizations of the more extended fractal aggregate (CCA\_1.92\_1.873\_400\_35), the pink curves ( $D_f = 2.16$ ,  $k_{\text{Sorensen}} = 2.065$ ) lie largely above the red curves ( $D_f = 1.92$ ,  $k_{\text{Sorensen}} = 1.873$ ), which in turn lie largely above the maroon curves ( $D_f = 1.68$ ,  $k_{\text{Sorensen}} = 1.700$ ). This is true of the total scattered intensity, as well as of the parallel polarized and perpendicularly polarized scattered intensities, and corresponds with the fact that the scattering cross sections of the more compact fractal aggregates as calculated from the output of MSTM are higher than the scattering cross sections of the 485 less compact fractal aggregates. (Refer to Table 9.)

Only in the direct forward scattering direction does the scattered intensity of the less compact fractal aggregates in each graph increase above the scattered intensity of the more compact fractal aggregates, and this is only to a small extent that is difficult to discern by eye from the graphs. This is despite the fact that the extinction cross sections of the less compact aggregates are lower than the extinction cross sections of the more compact aggregates (again, refer to Table 9) and is 490 probably due to the larger overall outer envelopes of the less compact aggregates (refer to Figs. 2 and 3). Due to the larger overall outer envelopes of the less compact aggregates, their normalized phase functions (normalized according to Eq. (4)), exhibit stronger and narrower forward scattering peaks (see, e.g., Bohren and Kho (1985), Gustafson and Kolokolova (1999), and Liu and Mishchenko (2005), their Fig. 2). Even though  $P(\theta_{\text{sca}})$  is multiplied by  $\sigma_{\text{sca}}$  in converting from  $P(\theta_{\text{sca}})$  to  $I_{\text{sca}}(\theta_{\text{sca}})$  via Eq. (5), the multiplication by  $\sigma_{\text{sca}}$  is not enough to increase the directly forward scattered intensity in the 495 broader forward scattering peak exhibited by the more compact aggregates to values greater than the directly forward scattered intensity in the narrower forwarding scattering peak exhibited by the less compact aggregates.

In Table 10, we list the ranges of values of scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  corresponding to Fig. 7. As was evident from Fig. 7, we can see from Table 10 that *within each category*, the more compact the fractal aggregate, the higher the values of scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$ . However, the overall range of perpendicularly 500 polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more extended fractal aggregates ( $(2.883 - 7.706) \times 10^{-4}$ ) is higher than the overall range of perpendicularly polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more compact fractal aggregates ( $(2.153 - 6.899) \times 10^{-4}$ ), and both of these ranges of intensity values encompass values that are more than an order of magnitude higher than the values of perpendicularly polarized scattered intensity for the SC aggregate ( $1.423 \times 10^{-5}$ ) and for the IAS aggregates ( $9.524 \times 10^{-6} - 1.391 \times 10^{-5}$ ), all of which agrees with 505 the direction of the SPIN measurements. On the other hand, the highest ratio of perpendicularly polarized scattered intensity



to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  among all of the fractal aggregates is  $3.133 \times 10^{-2}$ , which is lower than the 95th percentile value of the ratio S/P of either of the sample sets from the SPIN measurements (0.102 and 0.515, respectively; refer to Table 2).

Aggregate label	Range of values of $I_{\text{sca} \parallel \rightarrow \parallel} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Range of values of $I_{\text{sca} \parallel \rightarrow \perp} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Range of values of ratio of perpendicular to parallel $\frac{I_{\text{sca} \parallel \rightarrow \perp} (135 \pm 20^\circ)}{I_{\text{sca} \parallel \rightarrow \parallel} (135 \pm 20^\circ)}$	Range of values of $I_{\text{sca tot}} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]
SC_400_35	$3.425 \times 10^{-2}$	$1.423 \times 10^{-5}$	$4.155 \times 10^{-4}$	$3.426 \times 10^{-2}$
IAS_400_35	$(4.421 - 4.506) \times 10^{-2}$	$9.524 \times 10^{-6} - 1.391 \times 10^{-5}$	$(2.154 - 3.087) \times 10^{-4}$	$(4.422 - 4.507) \times 10^{-2}$
CCA_2.34_1.085_400_35	$(1.344 - 1.907) \times 10^{-2}$	$(3.050 - 4.047) \times 10^{-4}$	$(1.751 - 2.530) \times 10^{-2}$	$(1.378 - 1.940) \times 10^{-2}$
CCA_2.12_0.992_400_35	$(1.348 - 1.854) \times 10^{-2}$	$(2.153 - 2.848) \times 10^{-4}$	$(1.324 - 1.864) \times 10^{-2}$	$(1.370 - 1.880) \times 10^{-2}$
PCA_2.56_1.186_400_35	$(1.484 - 2.202) \times 10^{-2}$	$(3.371 - 6.899) \times 10^{-4}$	$(2.271 - 3.133) \times 10^{-2}$	$(1.518 - 2.271) \times 10^{-2}$
More compact fractal aggregates, overall	$(1.344 - 2.202) \times 10^{-2}$	$(2.153 - 6.899) \times 10^{-4}$	$(1.324 - 3.133) \times 10^{-2}$	$(1.370 - 2.271) \times 10^{-2}$
CCA_1.92_1.873_400_35	$(1.480 - 2.695) \times 10^{-2}$	$(3.047 - 4.314) \times 10^{-4}$	$(1.427 - 2.213) \times 10^{-2}$	$(1.512 - 2.733) \times 10^{-2}$
CCA_1.68_1.700_400_35	$(1.870 - 2.163) \times 10^{-2}$	$(2.883 - 3.301) \times 10^{-4}$	$(1.435 - 1.681) \times 10^{-2}$	$(1.899 - 2.194) \times 10^{-2}$
CCA_2.16_2.065_400_35	$(1.871 - 2.894) \times 10^{-2}$	$(5.176 - 7.706) \times 10^{-4}$	$(2.321 - 3.089) \times 10^{-2}$	$(1.923 - 2.971) \times 10^{-2}$
More extended fractal aggregates, overall	$(1.480 - 2.894) \times 10^{-2}$	$(2.883 - 7.706) \times 10^{-4}$	$(1.427 - 3.089) \times 10^{-2}$	$(1.512 - 2.971) \times 10^{-2}$

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**Table 10.** Ranges of values of scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for different realizations of aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm. The values for the single SC realization from Table 4 are also included here for reference.

515 In Sect. 3.6, we explore the range of values obtained with these same aggregate realizations, but with fixed aggregate orientation in the MSTM model calculations.

### 3.6 Theoretical calculations – Fixed orientation versus random orientation

As described in Sect. 2.2, by default, all of the theoretical calculations presented up to this point were obtained using the random orientation option of MSTM. On the one hand, we do not expect a particular orientation of the particles in the SPIN  
520 OPC to have been dominant; as a whole, the particles would have been more or less randomly oriented during the measurement. On the other hand, as an individual particle passed through the SPIN system, it would have been in some individual orientation. While we cannot assure that the fixed orientation of an individual realization that we generated would



be the same as the orientation that a particular aggregate had as it passed through the SPIN system, it is still worthwhile examining how removing the random orientation option in the MSTM calculations changes the range of calculated scattered intensity values. To this end, in this section, we conduct calculations on the same realizations as in Sect. 3.5, but now with each aggregate in fixed orientation. In Table 11, we list the ranges of values of extinction, absorption, and scattering cross section for all of the realizations of the aggregate generating algorithms with each aggregate in fixed orientation. From Table 11, we see once again that the most compact aggregates (SC, IAS, PCA\_2.56\_1.186\_400\_35, and CCA\_2.16\_2.065\_400\_35) tend to have the highest extinction cross sections and scattering cross sections, which again agrees with the results of Liu and Mishchenko (2005) and Liu et al. (2008). We also see that the ranges of values in Table 11 are a little broader than the ranges of values in Table 9, as expected.

In Table 12, we list the ranges of values of scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  with each aggregate in fixed orientation. As expected, overall, with fixed orientation, the ranges of the values of scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  are much broader, with the lowest value of each range significantly lower and the highest value of each range significantly higher than the respective values in Table 10, but the tendencies are the same as those seen in Table 10. As in Table 10, we see from Table 12 that the overall range of perpendicularly polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more extended fractal aggregates ( $5.601 \times 10^{-5} - 1.965 \times 10^{-3}$ ) is higher than the overall range of perpendicularly polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more compact fractal aggregates ( $5.858 \times 10^{-6} - 8.439 \times 10^{-4}$ ), and both of these ranges of intensity values encompass values that are more than an order of magnitude higher than the values of perpendicularly polarized scattered intensity for the SC aggregate in fixed orientation ( $2.030 \times 10^{-6}$ ) and for the IAS aggregates in fixed orientation ( $(1.065 - 7.332) \times 10^{-6}$ ).

In addition, from Table 12, we can see that the highest ratio of perpendicularly to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  among all of the more compact fractal aggregates is  $1.974 \times 10^{-1}$ , and the highest ratio of perpendicularly to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  among all of the more extended fractal aggregates is  $5.103 \times 10^{-1}$ . These values are comparable to the 95th percentile values of the ratio S/P of the sample sets from the SPIN measurements (again, 0.102 and 0.515, respectively; refer to Table 2). Thus, we find that individual aggregates in fixed orientation can reproduce the highest ratios of perpendicularly to parallel polarized scattered intensity exhibited by the samples from the SPIN measurements.

Aggregate label	Range of $\sigma_{\text{ext}}$ values [m <sup>2</sup> ]	Range of $\sigma_{\text{abs}}$ values [m <sup>2</sup> ]	Range of $\sigma_{\text{sca}}$ values [m <sup>2</sup> ]
SC_400_35	$2.849 \times 10^{-13}$	$1.676 \times 10^{-13}$	$1.173 \times 10^{-13}$
IAS_400_35	$(2.788 - 2.792) \times 10^{-13}$	$(1.599 - 1.601) \times 10^{-13}$	$(1.189 - 1.191) \times 10^{-13}$
CCA_2.34_1.085_400_35	$(2.326 - 2.565) \times 10^{-13}$	$(1.651 - 1.746) \times 10^{-13}$	$(6.663 - 8.181) \times 10^{-14}$



CCA_2.12_0.992_400_35	$(2.145 - 2.319) \times 10^{-13}$	$(1.717 - 1.769) \times 10^{-13}$	$(4.250 - 5.498) \times 10^{-14}$
PCA_2.56_1.186_400_35	$(2.557 - 2.765) \times 10^{-13}$	$(1.646 - 1.742) \times 10^{-13}$	$9.102 \times 10^{-14} - 1.024 \times 10^{-13}$
CCA_1.92_1.873_400_35	$(2.213 - 2.524) \times 10^{-13}$	$(1.695 - 1.805) \times 10^{-13}$	$(4.735 - 7.189) \times 10^{-14}$
CCA_1.68_1.700_400_35	$(2.131 - 2.190) \times 10^{-13}$	$(1.744 - 1.821) \times 10^{-13}$	$(3.347 - 4.463) \times 10^{-14}$
CCA_2.16_2.065_400_35	$(2.359 - 2.734) \times 10^{-13}$	$(1.645 - 1.813) \times 10^{-13}$	$(7.122 - 9.213) \times 10^{-14}$

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**Table 11.** Ranges of values of extinction, absorption, and absorption cross section for different realizations of aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm when the aggregates are in fixed orientation rather than random orientation. The values for the single SC realization in fixed orientation are also included here for reference.

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Aggregate label	Range of values of $I_{\text{sca} \parallel \rightarrow \parallel} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Range of values of $I_{\text{sca} \parallel \rightarrow \perp} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]	Range of values of ratio of perpendicular to parallel	Range of values of $I_{\text{sca tot}} (135 \pm 20^\circ)$ [W m <sup>-2</sup> ]
			$\frac{I_{\text{sca} \parallel \rightarrow \perp} (135 \pm 20^\circ)}{I_{\text{sca} \parallel \rightarrow \parallel} (135 \pm 20^\circ)}$	
SC_400_35	$3.143 \times 10^{-2}$	$2.030 \times 10^{-6}$	$6.461 \times 10^{-5}$	$3.143 \times 10^{-2}$
IAS_400_35	$(4.400 - 4.730) \times 10^{-2}$	$(1.065 - 7.332) \times 10^{-6}$	$2.250 \times 10^{-5} - 1.666 \times 10^{-4}$	$(4.401 - 4.730) \times 10^{-2}$
CCA_2.34_1.085_400_35	$4.232 \times 10^{-3} - 5.003 \times 10^{-2}$	$3.420 \times 10^{-5} - 8.341 \times 10^{-4}$	$1.739 \times 10^{-3} - 9.018 \times 10^{-2}$	$4.337 \times 10^{-3} - 5.086 \times 10^{-2}$
CCA_2.12_0.992_400_35	$(1.209 - 3.089) \times 10^{-2}$	$6.258 \times 10^{-5} - 4.309 \times 10^{-4}$	$4.579 \times 10^{-3} - 3.447 \times 10^{-2}$	$(1.224 - 3.109) \times 10^{-2}$
PCA_2.56_1.186_400_35	$1.567 \times 10^{-3} - 4.613 \times 10^{-2}$	$5.858 \times 10^{-6} - 8.439 \times 10^{-4}$	$2.786 \times 10^{-4} - 1.974 \times 10^{-1}$	$1.876 \times 10^{-3} - 4.690 \times 10^{-2}$
More compact fractal aggregates, overall	$1.567 \times 10^{-3} - 5.003 \times 10^{-2}$	$5.858 \times 10^{-6} - 8.439 \times 10^{-4}$	$2.786 \times 10^{-4} - 1.974 \times 10^{-1}$	$1.876 \times 10^{-3} - 5.086 \times 10^{-2}$
CCA_1.92_1.873_400_35	$6.188 \times 10^{-3} - 4.526 \times 10^{-2}$	$5.601 \times 10^{-5} - 5.430 \times 10^{-4}$	$2.122 \times 10^{-3} - 4.991 \times 10^{-2}$	$6.419 \times 10^{-3} - 4.565 \times 10^{-2}$
CCA_1.68_1.700_400_35	$2.212 \times 10^{-3} - 2.380 \times 10^{-2}$	$6.399 \times 10^{-5} - 7.694 \times 10^{-4}$	$1.066 \times 10^{-2} - 2.406 \times 10^{-1}$	$2.551 \times 10^{-3} - 2.405 \times 10^{-2}$
CCA_2.16_2.065_400_35	$2.161 \times 10^{-3} - 4.644 \times 10^{-2}$	$2.830 \times 10^{-4} - 1.965 \times 10^{-3}$	$6.094 \times 10^{-3} - 5.103 \times 10^{-1}$	$3.264 \times 10^{-3} - 4.672 \times 10^{-2}$
More extended fractal aggregates, overall	$2.161 \times 10^{-3} - 4.644 \times 10^{-2}$	$5.601 \times 10^{-5} - 1.965 \times 10^{-3}$	$2.122 \times 10^{-3} - 5.103 \times 10^{-1}$	$2.551 \times 10^{-3} - 4.672 \times 10^{-2}$

**Table 12.** Ranges of values of scattered intensity over the range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for different realizations of aggregates generated starting with an SC aggregate with  $D_{\text{outer-envelope}} = 400$  nm and  $d_{\text{pp}} = 35$  nm when the aggregates are in fixed orientation rather than random orientation. The values for the single SC realization from Table 4 are also included here for reference.

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#### 4. Discussion

As mentioned in Sects. 2.1 and 3.1, a larger value of perpendicularly polarized scattered intensity for parallel polarized incident intensity indicates some asymmetry/nonsphericity in the shape of the scattering particles or possibly birefringence or chirality in the scattering particle material. This might lead one to expect that the more fractal/extended the aggregate, the larger the value of perpendicularly polarized scattered intensity obtained. However, in computing the absolute value of scattered intensity (rather than the normalized phase function), each phase function is weighted by the total scattering cross section of the aggregate (refer to Eq. (7)). Thus, the higher scattering cross sections exhibited by the more compact aggregates of each set of realizations (refer to Sect. 3.5) gives more weight to their calculated scattered intensity.

We find that combining these two facts, the aggregates that possess a relatively high porosity but that are not too extended in shape are those that exhibit the highest perpendicularly polarized scattered intensity. Indeed, the realizations of aggregate PCA\_2.56\_1.186\_400\_35 exhibit the highest ranges of values of perpendicularly polarized scattered intensity of all of the more compact fractal aggregates, and the realizations of aggregate CCA\_2.16\_2.065\_400\_35 exhibit the highest ranges of values of perpendicularly polarized scattered intensity of all of the more extended fractal aggregates (refer to the discussion of Fig. 7 and to Tables 10 and 11).

In addition, as presented in Sects. 3.2-3.5, we find that using the random-orientation option of MSTM on our theoretical aggregates, we are able to reproduce the qualitative behavior of the SPIN measurements when we compare to the *median values* of those measurements. Namely, the overall range of perpendicularly polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more extended theoretical fractal aggregates is consistently higher than the overall range of perpendicularly polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  for the more compact theoretical fractal aggregates.

Although the measurements and theory agree qualitatively, quantitative agreement is not always observed. As described in Sects. 3.2-3.5, we found that using the random-orientation option of MSTM on our theoretical aggregates, the *highest values* of the ratio of perpendicularly polarized scattered intensity to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  exhibited by our theoretical aggregate realizations are not as high as the highest S/P ratios exhibited by the COJ300 and R2500U 400 nm samples from the SPIN measurements.

As shown in Sect. 3.6, only with fixed orientation do some values of the ratio of perpendicularly polarized scattered intensity to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  resemble the ratios in the 95th percentile of the measured S/P values. In fact, for individual aggregates, an even higher value of the S/P ratio is possible. The bottom row of Table 2 was obtained by dividing the row labelled “S” by the row labelled “P”, but if we were to present the different



percentiles of S/P based on the value of S/P for individual aggregates, the 95% percentile value of S/P would actually be

~1.0. We did not obtain a value of  $\frac{I_{\text{sca}\parallel\rightarrow\perp}(135\pm 20^\circ)}{I_{\text{sca}\parallel\rightarrow\parallel}(135\pm 20^\circ)}$  close to 1.0 for any of our theoretical aggregates.

A number of reasons for the lack of quantitative agreement are possible. Foremost, we note that there could be differences between the specifications of our theoretical aggregates and the actual chemical and physical properties of the measured aggregates. As described in Sects. 1 and 2.2, in our simulations, the aggregates in each set consist of the same number of primary particles of the same primary particle size but differing degrees of disorder. Thus, we are able to make an apples-to-apples comparison, in which all of the parameters in each set of aggregates are held constant except for the degree of disorder. However, from observations (e.g., Fig. 1), the primary particle size can vary within the same aggregate, as well as from aggregate to aggregate, and the number of primary particles can vary from aggregate to aggregate even within the same sample set. As found by Bescond et al. (2013) and Liu and Mishchenko (2005) and as reviewed in Kahnert and Kanngießer (2020), the direct backscatter depolarization ratio can vary with primary particle size and with the number of primary particles. On the other hand, Paulien et al. (2020) (also reviewed in Kahnert and Kanngießer (2020)) found that the number of primary particles does not have a significant impact on the direct backscatter depolarization ratio. For the cases we tested,

we found that when all other parameters are held constant, the ratio  $\frac{I_{\text{sca}\parallel\rightarrow\perp}(135\pm 20^\circ)}{I_{\text{sca}\parallel\rightarrow\parallel}(135\pm 20^\circ)}$  increases with  $d_{\text{pp}}$  and with  $N_{\text{pp}}$  for the more extended fractal aggregate ( $D_{\text{f}} = 1.92$ ;  $k_{\text{Sorensen}} = 1.873$ ) but not for the more compact fractal aggregate ( $D_{\text{f}} = 2.34$ ;  $k_{\text{Sorensen}} = 1.084$ ). (Refer to Tables 4 and 6.) Thus, we cannot say for certain whether further variations in  $d_{\text{pp}}$  and  $N_{\text{pp}}$  beyond what we already tested would reconcile the quantitative discrepancies.

Aside from further variations in  $d_{\text{pp}}$  and  $N_{\text{pp}}$ , there could be additional differences in the configuration of the primary particles within the aggregates beyond what our various realizations of the aggregate generating algorithm covered. When inspecting the SEM images, such as our Fig. 1, the viewing angle can mask additional asymmetry in the overall structure. Perhaps there is some chirality of shape (a slight helicity or handedness of some other form) in the aggregates examined in the SPIN measurements that the theoretical aggregate generating algorithms we employed do not fully reproduce. Alternatively, as investigated in Lu and Sorensen (1994) and Bescond et al. (2013) and as reviewed in Kahnert and Kanngießer (2020), effects such as overlapping of primary particles and “necking” can increase the linear depolarization ratio in the direct backscattering direction to as high as 0.03, but this value is still significantly lower than the highest S/P ratios we measured. (Interestingly, Lu and Sorensen (1994) suggested necking in an attempt to reconcile the fact that their calculations underestimated the depolarization of *forward* scattered radiation.) It seems less likely, but there could also be a measure of intrinsic chirality or birefringence in the BC material used to generate the SPIN measurement samples itself; such possible intrinsic chirality or birefringence was not considered in our theoretical calculations.



620 We believe the orientation of the particles throughout our experimental setup, and specifically in the detection region of the optical particle counter, is random. There remains a possibility that we do not fully understand the flow in this region and that it could lead to an organized orientation.

Yet another possibility concerns the contribution of Rayleigh scattering due to the presence of air in the chamber in the SPIN measurements. However, this is likely to be a minor effect, due to both the weak depolarizing ability of air molecules ( $S/P =$   
625  $\sim 0.02$ ; see, e.g., Sassen (2000)) and the low intensity of scattered radiation from Rayleigh scattering as compared to the intensity of scattered radiation from the aggregates, which would give the Rayleigh depolarization signal only a small weight in the overall depolarization signal. Likewise, while carbonaceous particles, such as soot, can exhibit Raman scattering (see, e.g., Le et al. (2022)), the Raman scattered signal is by nature very weak and only exhibits depolarization if the new vibrational mode to which the molecules transition is asymmetric enough. Other technical aspects of the measurements, such  
630 as deviations of the incident wave from being a 100% coherent plane wave that is 100% polarized parallel to the scattering plane would also likely have only a minor effect.

## 5. Summary and Conclusions

Carbonaceous aerosol particles are ubiquitous in the atmosphere. Their ability to impact atmospheric chemistry, human health, and climate have led to numerous studies of their morphological, chemical, cloud formation, and radiative properties  
635 (see, e.g., Bond et al. (2006), Bond and Bergstrom (2006), Liu and Mishchenko (2018), Kahnert and Kanngießer (2020), and references therein). In this study, we analyzed laboratory measurements of side/back scattering of visible radiation by analogues of bare (uncoated) atmospheric BC aggregates obtained with the SPIN instrumentation, and using the MSTM model, we conducted theoretical calculations of scattering of visible radiation by compact and fractal theoretical BC aggregates constructed based on the measured morphological parameters of the laboratory generated aggregates. As  
640 discussed in Sect. 4, we found that using the random-orientation option of MSTM on our theoretical aggregates, we are able to reproduce the qualitative behavior of the SPIN measurements when we compare to the *median values* of those measurements. However, using the random-orientation option of MSTM on our theoretical aggregates, the *highest values* of the ratio of perpendicularly polarized scattered intensity to parallel polarized scattered intensity in the angular range  $\theta_{\text{sca}} = 135 \pm 20^\circ$  exhibited by our theoretical aggregate realizations are not as high as the highest  $S/P$  ratios exhibited by the COJ300  
645 and R2500U 400 nm samples from the SPIN measurements. We found that only with fixed orientation do some values of the ratio of perpendicularly polarized scattered intensity to parallel polarized scattered intensity resemble the ratios in the 95th percentile of the measured  $S/P$  values.

We note that relatively high values of backscattering linear depolarization ratio were also obtained in the field measurements of Burton et al. (2015) (original and corrigendum). Liu and Mishchenko (2018) demonstrated that they were able to  
650 reproduce such high values of backscattering linear depolarization ratio only by simulating aged soot containing large



amounts of refractory materials along with black carbon, not with bare soot. Similarly, Kahnert and Kanngießer (2020) state in their review that in most cases typical for atmospheric BC, the depolarization ratio of bare BC aggregates rarely exceeds the range 0.01-0.03. In this work, we demonstrated that even bare black carbon can exhibit high values of S/P at side/backscattering directions and that we can reproduce such high values in calculations of single scattering by bare black carbon aggregates if we use fixed orientation. Given this, it is possible that models of scattering by BC aggregates that employ the random orientation assumption/option may underpredict the linear depolarization ratio of actual BC aggregates, and this should be investigated further in the future. Additional realizations further varying the three-dimensional aggregate structure, such as varying the internal size distribution of primary particles in each aggregate, further varying the number of primary particles, and further varying the chemical properties of the black carbon could be investigated. Likewise, a wider range of experimental BC analog samples, more in-depth imaging, and more sophisticated size selection would provide even better experimental statistics. It is important to note that the angles we inspected in the backscattering hemisphere are not the exact direct backscattering direction considered in Burton et al. (2015) and Liu and Mishchenko (2018). Nonetheless, our results might have important implications for remote sensing of soot aerosol via lidar backscattering. Furthermore, such direct comparison of theory to laboratory experiments of light scattering by BC aggregates represents an additional step towards a better overall understanding of the impact aerosol particles have on our environment and our climate system.

**Funding:** This work was supported by the MIT International Science and Technology Initiatives (MISTI)-Israel Seed Fund (2017-2018); Carynelisa Haspel is supported by the Israel Science Foundation (1872/17 and 2187/21); Cuiqi Zhang is supported by National Natural Science Foundation of China (No. 42205101) and LAC/CMA (2022B07).

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**Data availability:** The data and model output from this study are available from the authors by request.

**Author contribution:** CZ and MJW carried out the laboratory experiments under the supervision of DJC. CH performed the calculations and prepared the manuscript with contributions from all co-authors. DJC, CH, and CZ acquired the funding for the project leading to this publication.

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**Competing interests:** The authors declare that they have no conflict of interest.

**Acknowledgements:** We thank Daniel Mackowski for supplying the aggregate generating algorithm and for consultation on its use, as well as for helpful discussions on MSTM in general; Ynon Hefets for helpful discussions on polarization; Ofir Shoshanim and Jonas Gienger for helpful discussions on scattering experiments; Droplet Measurement Technologies and Sarvesh Garimella for useful information on the SPIN instrument; and Yue Zhang and Leonid Nichman for assistance with the black carbon measurements.

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<https://doi.org/10.5194/egusphere-2023-556>

Preprint. Discussion started: 24 April 2023

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