

1 **Comparison of the imaginary parts of the atmospheric** 2 **refractive index structure parameter and aerosol flux based** 3 **on different measurement methods**

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14 Abstract: The complexity of aerosol particle properties and the diversity of characterizations make
15 aerosol vertical transport flux measurements and analysis difficult. Although there are different
16 methods, such as aerosol particle number concentration flux and aerosol mass flux based on the eddy
17 covariance principle, and aerosol mass flux measurements based on the light-propagated large-
18 aperture scintillation principle, there is a lack of mutual validation among the different methods. In
19 this paper, a comparison of aerosol mass flux measurements based on the eddy covariance principle
20 and aerosol mass flux measurements based on the light-propagated large aperture scintillation
21 principle is carried out. The key idea of aerosol mass flux measurements based on the light-
22 propagated large-aperture scintillation principle is the determination of the imaginary part of the
23 atmospheric equivalent refractive index structure parameter (AERISP). In this paper, we first
24 compare the AERISPs measured by two different methods and then compare the aerosol mass
25 vertical transport fluxes obtained by different methods. The experiments were conducted on the
26 campus of the University of Science and Technology of China (USTC). A light propagation
27 experiment was carried out between two tall buildings to obtain the imaginary and real parts of the
28 AERISPs for the whole path, which in turn can be used to obtain the aerosol vertical transport flux.
29 An updated visibility meter is installed on the meteorological tower in the middle of the light path,
30 which is utilized to obtain the single-point visibility, which is then converted to the imaginary part
31 of the atmospheric equivalent refractive index (AERI). The imaginary parts of the AERISP were
32 obtained via spectral analysis with visibility data. The results show that the imaginary parts of the
33 AERISPs obtained by different methods are in good agreement. The imaginary part of the AERI
34 measured by the visibility meter and the vertical wind speed obtained by the ultrasonic anemometer
35 were used for covariance calculations to obtain the aerosol vertical transport flux. The trends in
36 aerosol vertical transport fluxes obtained by the different methods are consistent, and there are
37 differences in some details, which may be caused by the inhomogeneity of the vertical transport of
38 aerosol fluxes. The experimental results also showed that urban green land is a sink area for aerosol
39 particles.

40 Keywords: light propagation, scintillation, atmospheric equivalent refractive index, structure
41 parameter, eddy covariance, aerosol fluxes

1 Introduction

Atmospheric aerosols are solid or liquid particles suspended in the atmosphere that can affect public health, reduce near-surface visibility, decrease direct radiation from the air, and act as condensation nuclei affecting cloud structure and distribution (McNeill, 2017; Rosenfeld et al., 2014). Human activities have dramatically altered air quality, climate and the Earth system. The expansion of urban, agricultural and industrial areas and changes in the nature of land use have increased aerosol concentrations. Due to the complexity of aerosols, many observations have been carried out from different perspectives. However, most of the current observations only measure the state characteristics of aerosols, such as concentration, particle size distribution, and composition, and what is obtained is an average characterization of aerosol properties (Krieger et al., 2012).

Aerosol particles in the atmosphere follow atmospheric motion, which manifests as an uneven distribution of aerosol particle concentrations in space and time. On the one hand, unevenly distributed aerosol particles will have a corresponding effect on light wave propagation in the atmosphere, on the other hand, we can understand the distribution characteristics of aerosol particles based on the optical effect of aerosol particles and then obtain more information about the transportation of aerosols.

Previously, Yuan et al. (2016) introduced the concepts of the atmospheric equivalent refractive index (AERI) and the atmospheric equivalent refractive index structure parameter (AERISP). The term AERISP corresponds to the equivalent medium containing air and aerosol particles, relative to the commonly used atmospheric refractive index structure parameter (RISP) obtained from single-point measurements (Wyngaard et al., 1971). The AERI includes real and imaginary parts; accordingly, the AERISP also includes real and imaginary parts of the structure parameters. When the working wavelength is in the atmospheric transparent band, the light wave is almost not absorbed by the gas components in the atmosphere, and the attenuation of the light wave is caused mainly by the extinction of aerosol particles. Theoretical analysis revealed that, the imaginary part of the AERISP is determined by the fluctuation in aerosol concentration, and the real part of the AERISP corresponds to the atmospheric temperature variation (Yuan et al., 2021). Experiments have shown that aerosol particles follow the same theory of locally homogeneous isotropic turbulence as air molecules (Martensson et al., 2006; Vogt et al., 2011a; Ren et al., 2020); that is, the fluctuation in the particle concentration follows the ‘-5/3’ power law under unstable atmospheric stratification, and the concentration-velocity cospectra for particle number flux follow the ‘-4/3’ power law, similar to the temperature-velocity cospectra (Kaimal et al., 1972). Therefore, the distribution of small particles can be regarded as a passive conservative quantity, similar to the temperature field.

75 Then, it can be assumed that the aerosol mass concentration fluctuation also follows the locally
76 homogeneous isotropic turbulence theory; thus the aerosol mass concentration structure parameter
77 can be defined (Yuan et al., 2016). Based on the fact that the temperature structure function satisfies
78 the surface layer similarity theory and thus the surface layer sensible heat flux is obtained from the
79 temperature structure parameter (Wyngaard et al., 1971), it is analogous to that, based on the fact
80 that the aerosol mass concentration structure parameter satisfies the surface similarity theory, the
81 surface layer aerosol mass flux is obtained from the aerosol mass concentration structure parameter
82 (Yuan et al., 2016; Yuan et al., 2019). Using the relationship between the temperature-real part of the
83 AERI and aerosol mass concentration-imaginary part of the AERI, the temperature structure
84 parameter and aerosol mass concentration structure parameter are obtained from the real part of the
85 AERISP and imaginary part of the AERISP.

86 From this, the aerosol mass concentration flux can be obtained, utilizing large aperture
87 scintillometer (LAS) measurements. AERISP observations are carried out in many places, after
88 which the aerosol flux is obtained via the similarity theory (Yuan et al., 2016; Yuan et al., 2019).
89 However, there is a lack of experimental verification of the imaginary structure parameter and
90 aerosol flux observations. Currently, the imaginary part of the AERISPs is obtained using only LAS
91 measurements; therefore, it is necessary to carry out measurements of the imaginary part of the
92 AERISPs based on other methods, as well as measurements of aerosol fluxes based on different
93 methods.

94 At present, in addition to the previously mentioned measurements of the AERISP based on the
95 principle of long-range light propagation and the similarity theory of the surface layer to obtain the
96 vertical transport flux of aerosol mass in the surface layer, several studies utilize eddy covariance
97 (EC) techniques with fluctuations in aerosol particle number concentration and fluctuation in vertical
98 wind speed to obtain the flux of the number concentration of aerosol particles. Such measurements
99 are carried out in many places (Gordon et al., 2011; Vogt et al., 2011b; Ripamonti et al., 2013).
100 Measurements have revealed quantitative relationships for urban aerosol fluxes among urban vehicle
101 emissions and meteorological conditions (Jarvi et al., 2009), and characteristics of sea salt transport,
102 and aerosol properties (Nemitz et al., 2009). We followed this approach and conducted several
103 measurements in 2019 and 2020 to measure aerosol particle number concentration fluxes using an
104 eddy-correlation system combining a fast-response particle counter with an ultrasonic anemometer
105 with a response frequency of up to 10 Hz, and to calculate aerosol mass concentration fluxes by
106 simultaneously measuring aerosol particle size distribution, mass concentrations, forward scattering
107 coefficients, extinction coefficients, and other parameters. For half of the experimental period, the
108 trends of the measurements of the two methods were the same, while the other periods differed more
109 (unpublished experimental results). The main reason may be the very weak extinction of aerosol

110 particles at scales much smaller than the working wavelength. Second, the aerosol number
111 concentration flux must be combined with parameters such as particle size distribution, complex
112 refractive index of aerosol particles and aerosol particle density, which also need to be sampled in
113 real time. This also illustrates the complexity of aerosol particles.

114 Recently, Ren et al. (2020) improved upon the conventional visibility meter method to obtain 1
115 Hz visibility data, and subsequently utilized the EC method to obtain the aerosol vertical transport
116 flux based on the relationship between visibility and aerosol mass concentration. The visibility is
117 approximately inversely proportional to the atmospheric extinction coefficient, i.e., approximately
118 inversely proportional to the imaginary part of the AERI; therefore their theory of obtaining the
119 aerosol vertical transport flux by the EC method is close to theory of the aerosol vertical transport
120 flux based on the measurement of the long-path light propagation. Inspired by their work, we used
121 an improved visibility meter in this study to obtain visibility data at a higher frequency of 1 Hz and
122 cross-correlated the data with ultrasonic anemometer measurements to potentially utilize the
123 obtained aerosol vertical transport fluxes to achieve experimental validation of the imaginary part of
124 the AERISP and aerosol flux observations.

125 The theoretical and experimental introduction is given in the second part of the paper, the
126 experimental results are given in the third part, and the conclusions and discussion are given in the
127 fourth part.

128 **2 Theoretical methods and experiments**

129 The AERISP and aerosol vertical transport flux are the topics of interest in this paper. In this
130 section, definitions and theoretical expressions for these parameters are given, as well as how the
131 measurements were carried out.

132 **2.1 The imaginary part of the AERISP**

133 Normally, the atmosphere consists of gas molecules and aerosol particles. When a light beam
134 propagates in the atmosphere, due to the inhomogeneous distribution of the atmospheric gas
135 refractive index, the beam will be refracted and diffracted, which results in an inhomogeneous spatial
136 distribution of the beam energy. Due to the existence of aerosol particles in the atmosphere, the beam
137 will be scattered and absorbed, and the energy of the beam will be weakened. Therefore, the
138 atmospheric gas molecules and aerosol particles can be taken as a whole, called the equivalent
139 medium; thus, the atmospheric equivalent refractive index (AERI) n_{equ} concept is introduced (van
140 de Hulst, 1957; Yuan et al., 2021),

$$141 \quad n_{equ} = n_m + i \frac{2\pi}{\eta^3} \int_0^\infty S(0) \frac{dN}{dD} dD \quad (1)$$

142 where n_m is the refractive index of atmospheric molecules, η is the wavenumber of light waves, and
 143 i denotes an imaginary number. $S(0)$ is the forward scattering function (0 in parentheses is the
 144 scattering angle). N is the number of aerosol particles per unit volume, and dN/dD is the size
 145 distribution of aerosol particles.

146 The AERI consists of real and imaginary parts denoted by n_{Re} and n_{Im} , respectively; i.e., n_{equ}
 147 $=n_{Re} +in_{Im}$. The real part is the refractive index of the molecule, and the imaginary part is.

$$148 \quad n_{Im} = \frac{2\pi}{\eta^3} \int_0^\infty Re[S(0)] \frac{dN}{dD} dD \quad (2)$$

149 The atmospheric extinction coefficient has a similar form(Liou, 2002):

$$150 \quad \beta_{ext} = \frac{4\pi}{\eta^2} \int_0^\infty Re[S(0)] \frac{dN}{dD} dD \quad (3)$$

151 From Eqs (2) and (3), we can see that

$$152 \quad n_{Im} = \lambda\beta_{ext}/4\pi \quad (4)$$

153 where λ is the working wavelength ($\lambda=2\pi/\eta$).

154 Due to the dependence of the reduction in contrast on atmospheric absorption and scattering,
 155 the following relationship between visibility V and extinction coefficient β_{ext} can be obtained:
 156 $V=3.912/\beta_{ext}$ (Middleton, 1957;Charlson, 1969). Thus, β_{ext} in the relationship ($V=3.912/\beta_{ext}$)
 157 represents the extinction by all compositions in the air, e.g., absorption and scattering of aerosols
 158 and atmospheric molecular extinction. In other words, the visibility-based extinction coefficient is
 159 sum of the extinction coefficient from aerosol absorption and scattering and the atmospheric
 160 molecular extinction coefficient. However, in the urban atmosphere, the extinction effect of aerosols
 161 is much greater than that of atmospheric molecules. Therefore, the contribution of extinction by
 162 atmospheric molecules can be neglected. Therefore, the aerosol extinction coefficient can be
 163 deduced from visibility measurements, and the imaginary part of the AERI can be obtained based
 164 on Eq. (4).

165 Experiments show that the temperature fluctuation satisfies the turbulence "2/3" law(Liu et al.,
 166 2017), and due to small relative changes in pressure and air temperature (unit K) occurring over a
 167 short period, the change in the real part of the AERI has a good linear relationship with the
 168 temperature change, and the fluctuation in the real part of the AERI also satisfies the turbulence
 169 "2/3" law; thus, we can define the structure parameter of temperature, C_T^2 , and the real part of the
 170 AERISP $C_{n,Re}^2$. Therefore, general scalars can be extended, such as the fluctuation of the imaginary
 171 part of the AERISP and the fluctuation of the atmospheric extinction coefficient. Thus, we can
 172 assume that the imaginary part of the AERI satisfies the turbulence "2/3" law; that is, the structure
 173 function of the imaginary part of the AERI $D_{n,Im}(r)$ (r is the separation) can be defined as

$$174 \quad D_{n,Im}(r) = \overline{[n_{Im}(\vec{x}) - n_{Im}(\vec{r} + \vec{x})]^2} = C_{n,Im}^2 r^{2/3} \quad (5)$$

175 where $\vec{x}, \vec{r} + \vec{x}$ are the coordinates of two points in space, \vec{r} is the separation vector, $C_{n,Im}^2$ is the
 176 imaginary part of the AERISP, and the overbar indicates the mean.

177 Thus, we can introduce the imaginary part of the AERISP $C_{n,Im}^2$, a parameter used to describe
 178 the fluctuation intensity of the imaginary part of the AERI ($C_{n,Im}^2$ should be the structure parameter
 179 for the imaginary part of the AERI, conveniently denoted as the imaginary part of the AERISP).
 180 Correspondingly, we can introduce the structure parameter of the atmospheric extinction
 181 coefficient $C_{\beta_{ext}}^2$ and the structure parameter of the fluctuation of the aerosol mass concentration C_M^2 .

182 2.2 Two methods of AERI measurement

183 From the definition of the AERISP in the last part and the relationship between the AERI and
 184 the extinction coefficient, it can be seen that the AERISP has an important influence on light
 185 propagation in the atmosphere; thus, the AERISP can be estimated from the fluctuations in light
 186 propagation intensity and the monitoring of the extinction coefficient. This section describes how to
 187 measure the AERISP via two methods.

188 2.2.1 Long-Path Light Propagation Methods

189 When an approximately collimated light beam in the transparent band of the atmosphere is
 190 selected and propagated over a distance, the light intensity at the receiving end fluctuates. The
 191 fluctuation in light intensity has two causes: one is the uneven distribution of the real part of the
 192 AERI caused by the temperature fluctuation, and the other is the uneven distribution of the imaginary
 193 part of the AERI caused by the uneven distribution of aerosol particles. Assuming that the above two
 194 causes are not related, they can be decomposed. The power spectral density is usually used to
 195 characterize the fluctuation in light intensity. Through spectral analysis, the power spectral density
 196 of light intensity fluctuations can be decomposed into the contribution of the imaginary part of the
 197 AERISP and the contribution of the real part of the AERISP. The contribution of the inhomogeneous
 198 distribution of the imaginary part of the AERISP to the light intensity fluctuation is expressed as the
 199 temporal spectrum $W_{lnl,lm}(f)$ (Yuan et al., 2015),

$$200 \quad W_{lnl,lm}(f) = 64\pi^2\eta^2 \int_0^L dx \int_{2\pi f/v}^{\infty} \Phi_{n,lm}(\kappa) \cos^2\left[\frac{\kappa^2 x(L-x)}{2\eta L}\right] [(\kappa v)^2 - (2\pi f)^2]^{-1/2} \cdot$$

$$201 \quad \left[\frac{2J_1\left(\frac{D_r \kappa x}{2L}\right)}{D_r \kappa x / 2L}\right]^2 \left[\frac{2J_1\left(\frac{D_t \kappa(L-x)}{2L}\right)}{D_t \kappa(L-x) / 2L}\right]^2 \kappa d\kappa \quad (6)$$

202 where f is the frequency of the log-intensity spectrum, η is the wavenumber of the spherical wave
 203 ($\eta=2\pi/\lambda$, λ is the light wavelength), x is the position of the propagating wave, L is the length of the
 204 propagation path, κ is the wavenumber of the two-dimensional log-intensity spectrum, and $\Phi_{n,lm}$ is
 205 the spectrum of the imaginary parts of the refractive index, where the subscript n denotes the
 206 refractive index and the subscript lm denotes the imaginary parts of the refractive index, D_t is the
 207 transmitting aperture diameter, D_r is the receiving aperture diameter (D_t and D_r are usually identical
 208 for an LAS), v is the transverse wind speed and J_1 is the first-order Bessel function. The widely used
 209 von Karman spectral form for $\Phi_{n,lm}$ is adopted in this study (Andrews and Phillips, 2005) and can
 210 be expressed as follows:

$$211 \quad \Phi_{n,lm}(\kappa) = 0.033C_{n,lm}^2 \left(\kappa^2 + \frac{1}{L_0^2}\right)^{-\frac{11}{6}} e^{-\frac{\kappa^2 l_0^2}{5.92^2}} \quad (7)$$

212 Here, L_0 is the outer scale of turbulence, and l_0 is the inner scale of turbulence.

213 Substituting Eq. (7) into Eq. (6) and integrating the right-hand side of Eq. (6) yields,

214
$$W_{\ln I, \text{Im}}(f) = 0.129 C_{n, \text{Im}}^2 \eta^2 L v^{5/3} [f^2 + (\frac{v}{2\pi L_0})^2]^{-4/3} \quad (8)$$

215 Using Eq. (8), the imaginary part of the AERISP can be determined based on the shape of the
 216 spectrum while being constrained by the low-frequency variance in the light intensity fluctuation
 217 from the imaginary part of the AERISP.

218 To carry out a comparative analysis with the results of the real part of the AERISP, the
 219 expression for the power spectral density of the logarithmic light intensity fluctuation $W_{\ln I, \text{Re}}(f)$ due
 220 to the real part of the AERI is also given here as (Clifford, 1971; Nieveen et al., 1998),

221
$$W_{\ln I, \text{Re}}(f) = 64\pi^2 \eta^2 \int_0^L dx \int_{2\pi f/v}^{\infty} \Phi_{n, \text{Re}}(\kappa) \sin^2 \left[\frac{\kappa^2 x(L-x)}{2\eta L} \right] [(\kappa v)^2 - (2\pi f)^2]^{-1/2} \cdot$$

222
$$\left[\frac{2J_1(\frac{D_r \kappa x}{2L})}{D_r \kappa x / 2L} \right]^2 \left[\frac{2J_1(\frac{D_t \kappa(L-x)}{2L})}{D_t \kappa(L-x) / 2L} \right]^2 \kappa d\kappa \quad (9)$$

223 Integrating Eq. (9) yields the fluctuation variance of the log light intensity as

224
$$\sigma_{\ln I, \text{Re}}^2 = \int_0^{\infty} W_{\ln I, \text{Re}}(f) df = 0.89 C_{n, \text{Re}}^2 L^3 D_t^{-7/6} D_r^{-7/6} \quad (10)$$

225 The real part of the AERISP is usually calculated using Equ. (10) (Wang et al., 1978).

226 The calculation steps for the real and imaginary parts of AERISP are as follows: first, power
 227 spectrum analysis or correlation analysis of the irradiance fluctuation data are performed; then, the
 228 irradiance fluctuation data are decomposed into high-frequency and low-frequency parts; the high-
 229 frequency part corresponds to the contribution of the real part of the AERI; and the low-frequency
 230 part of the fluctuation corresponds to the contribution of the imaginary part of the AERI; finally, the
 231 real part of the AERISP $C_{n, \text{Re}}^2$ can be obtained from Eq. (10); and the imaginary part of the AERISP
 232 $C_{n, \text{Im}}^2$ can be obtained from the low-frequency part of the irradiance fluctuation.

233 2.2.2 Spectral analysis methods for single-point 234 measurements

235 Aerosol particles experience atmospheric motion, which is consistent with general atmospheric
 236 motion characteristics, and the "-5/3" law can be used to characterize fluctuations in aerosol-related
 237 properties. Therefore, in the inertial subregion, the extinction coefficient power spectral density is

238
$$S_{\beta_{\text{ext}}}(f) = (2\pi / U) S_{\beta_{\text{ext}}}(\kappa) = 0.25 C_{\beta_{\text{ext}}}^2 (2\pi / U)^{-2/3} f^{-5/3} \quad (11)$$

239 The extinction coefficient structure parameter $C_{\beta_{\text{ext}}}^2$ can be converted to the imaginary part of
 240 the AERISP according to equation (4). The coefficient in Eq. (11) is 0.25 (Wyngaard et al., 1971). It
 241 has been suggested in the literature that the coefficient for the spectral density should be 0.125 (Gibbs
 242 and Fedorovich, 2020). The difference between the two coefficients 0.25 and 0.125 is whether the
 243 integral of the spectral density is equal to the variance or half of the variance. If the integral of the
 244 spectral density is equal to the variance, a coefficient of 0.25 is taken; if the integral of the spectral
 245 density is equal to half of the variance, the coefficient is taken as 0.125. According to the spectral
 246 density curve, the coefficients are determined within the inertial subregion, and the structure

247 parameters $C_{\beta_{ext}}^2$ can be obtained. According to the relationship between the extinction coefficient
 248 and the imaginary part of the AERI in Eq. (4), the imaginary part of the AERISP can be obtained as
 249 $C_{n,Im}^2$.

250 Similarly, power spectral density profiles with temperature fluctuations that

$$251 \quad S_T(f) = (2\pi/U)S_T(\kappa) = 0.25C_T^2(2\pi/U)^{-2/3}f^{-5/3} \quad (12)$$

252 The actual temperature turbulence spectral density profile often takes the form of a von Karman
 253 spectrum as

$$254 \quad S_T(f) = 0.25C_T^2(2\pi/U)^{-2/3} \left(f^2 + \left(\frac{U}{2\pi L_0} \right)^2 \right)^{-5/6} \quad (12')$$

255 Based on the relationship between the temperature and the real part of the AERI, we have

$$256 \quad C_{n,Re}^2 = C_T^2/R_{TN}^2 \quad (13)$$

257 where R_{TN} denotes the coefficient of proportionality between the change in the real part of the AERI
 258 and the change in atmospheric temperature (Tatarskii, 1961; Zhou et al., 1991).

$$259 \quad R_{TN} = \frac{dT}{dn_{Re}} = -1.29 \times 10^4 \times \left(1 + \frac{7.52 \times 10^{-3}}{\lambda^2} \right)^{-1} \frac{\bar{T}^2}{\bar{P}} \quad (14)$$

260 where the wavelength λ is in microns, the atmospheric pressure P is in hectopascals, and the
 261 temperature T is in K.

262 The real part of the AERISP can be obtained by fitting the experimental data using Eqs. (12) or
 263 (12').

264 **2.3 Flux estimation**

265 The method for estimating the AERISP was given in the former sections. The purpose of
 266 estimating the AERISP in this paper is to estimate the aerosol flux in the near-surface layer. Here,
 267 the method of estimating the aerosol flux based on the AERISP is given first, and then the method
 268 of estimating the aerosol vertical transport flux based on the EC technique is introduced.

269 **2.3.1 Light propagation method**

270 Experiments have shown that the AERISPs satisfy the theory of surface layer similarity; thus,
 271 (Yuan et al., 2019)

$$272 \quad F_{a_LAS} = \left(\frac{C_{n,Im}^2}{C_{n,Re}^2} \right)^{1/2} \frac{R_{MN}}{R_{TN}} u_* |T_*| \quad (15)$$

273 where u_* is the friction velocity and T_* is the characteristic potential temperature. These two
 274 parameters can be determined from the wind speed and temperature profiles. The real and imaginary
 275 parts of the AERISP are determined from LAS measurements. The R_{MN} can be obtained from aerosol
 276 mass concentration and visibility measurements ($R_{MN} = M/n_{Im}$, where M is the aerosol mass
 277 concentration approximated as PM_{10} and n_{Im} can be determined from visibility measurements)(Yuan
 278 et al., 2021), and the R_{TN} can be calculated from the mean air temperature and other measurements
 279 using Eq. (14) again.

280 When turbulence in the surface layer develops, Eq. (15) can be approximated as(Yuan et al.,
281 2019),

$$282 \quad F_{a_LAS} = a\left(\frac{g}{T}\right)^{1/2} R_{TN}^{1/2} (C_{n,Re}^2)^{1/4} R_{MN} (C_{n,lm}^2)^{1/2} (z-d) \quad (16)$$

283 Here, a is the scale factor with a theoretical value of 0.567 (which needs to be determined by
284 comparative experiments), g is the gravitational acceleration, z is the scintillator height, and d is the
285 zero-plane displacement. Equation (16) does not require measurements of the u^* and T^* data.
286 Generally, the measurement heights are high, and the assumption of developed turbulence in the
287 surface layer is easily met during the day under unstable conditions.

288 **2.3.2 Based on single-point eddy covariance**

289 Eddy covariance is a commonly used method for the measurement of Earth air exchange fluxes
290 in the near-surface layer. Using rapid measurements of the vertical wind speed and extinction
291 coefficient to obtain the ups and downs of the vertical wind speed and extinction coefficient, the
292 expression for the vertically transported aerosol flux calculated by the eddy-covariance method with
293 a mean vertical velocity close to zero is given by(Wilczak et al., 2001)

$$294 \quad F_{a_EC} = R_{MN} \frac{\lambda}{4\pi} \overline{w' \beta'_{ext}}$$

295 The prime' in Eq. (17) denotes fluctuation.

296 **2.4 Introduction to the experiment**

297 The experiments were performed on the campus of the University of Science and Technology
298 of China (USTC) in Hefei, Anhui Province, China. The campus of the USTC is located in downtown
299 Hefei. Figure 1a shows part of the Hefei city area, where the red rectangle corresponds to Fig. 1b,
300 the campus of the USTC. The campus is surrounded by four highways, and the two highways in the
301 west and north have more vehicles, especially viaducts in the west. The campus is composed of
302 vegetation, roads and teaching buildings. As shown in Fig. 1b, green vegetation covers most of the
303 campus. The roofs of the school buildings are almost on a plane with the tree canopy and are
304 approximately 17 meters above the ground ($z_H=17$ m). Thus, the zero-plane displacement was 11.4
305 m ($17 \times 0.67=11.4$) (Shao et al., 2021;Grimmond and Oke, 1999;Leclerc and Foken, 2014). There
306 are two tall buildings (T and R in Fig. 1b) at the southernmost and northernmost parts of the campus,
307 and the distance between the two buildings is approximately 960 meters. The experiment consists of
308 two parts: one part consists of carrying out the light propagation experiment using a self-developed
309 large aperture scintillator (LAS), and the other part consists of carrying out the measurement using
310 the instruments on the meteorological tower in the middle of the beam (the details of the instruments
311 are listed in Table 1). The transmitting end of the LAS was installed on the 12th floor of the
312 southernmost building (T in Fig. 1b), the receiving end was installed on the 12th floor of the
313 northernmost building (R in Fig. 1b), and the distance of the beam from the ground was
314 approximately 35 meters. The apertures of the transmitting and receiving ends were 250 mm. The
315 sampling frequency of the receiving end was 500 Hz, and a data file was saved every 30 minutes.
316 The height of the meteorological tower is 18 meters above the roof of the teaching building (P in Fig.
317 1b). The height of the top of the meteorological tower is equal to the height of the beam. The
318 meteorological tower is equipped with 5 layers of wind speed, wind direction, temperature and

319 humidity measurement sensors. At the top of the tower, there is a radiation quadrature sensor, and at
320 the bottom of the tower, there is a rainfall measurement sensor. In this paper, we use data from the
321 top 18 meters of height of the meteorological tower with sensors installed for conventional
322 meteorological parameters, including temperature, humidity, wind speed, wind direction and
323 radiation. Conventional meteorological data were collected at 1-second intervals, average data were
324 obtained every half hour after data collection, and precipitation data were recorded every half hour.
325 A three-dimensional sonic anemometer thermometer was installed at the top of the tower, and the
326 high-frequency sampling visibility sensor CS120A (Campbell, 2012) was upgraded to obtain 1-Hz
327 visibility (Ren et al. 2020). A three-dimensional sonic anemometer thermometer can obtain a
328 sampling frequency of 10 Hz and is a common instrument used in atmospheric turbulence research;
329 as such, we will not introduce it in depth. To correlate the vertical wind speed with the extinction
330 coefficient to obtain the aerosol flux, the data collected by the sonic anemometer-thermometer at 10
331 Hz were averaged to obtain 1-Hz data, which were saved in a data file. By doing so, the aerosol flux
332 only contains eddies with a frequency lower than 1 Hz; in other words, any turbulent eddy, whose
333 frequency is higher than 1 Hz, is automatically eliminated. By comparing the T-w correlations
334 calculated from the 10 Hz data and the 1 Hz data, it can be seen that the error due to this high-
335 frequency neglect is less than 5% (details in Appendix).

336 The time period of the experiment is January 9-23, 2022, a total of 15 days. The winter period
337 was chosen, because it is considered to be typical of this period, with mainly sunny days, weak
338 rainfall, and relatively high pollution in winter.

339 **2.5 Data quality control**

340 The quality of the data obtained from field observations needs to be controlled before further
341 processing (Foken and Wichura, 1996). This study involves several types of data, mean variables,
342 cumulative variables, and fluctuating variables. The mean variables included 30-minute averages of
343 temperature, humidity, wind speed, wind direction, and global radiation. Data quality control for
344 mean variables was performed by comparing measurements at different heights or different sites.
345 The same variables with the same trend at different heights and different locations were considered
346 high-quality data. All the measured mean data were determined to be satisfactory. The cumulative
347 variables refer to 30-minute rainfall data. Rainfall data were qualified with reference to relative
348 humidity, total radiation and air temperature. The fluctuating data included 10-Hz ultrasonic
349 anemometer data and 1-Hz visibility data, as well as high-frequency intensity fluctuation data
350 measured by the LAS, the real and imaginary parts of the AERISP, and calculated aerosol fluxes.
351 Quality control consisted mainly of eliminating spikes and replacing missing data.

352 The reason for the spike points in the light intensity fluctuation data is that the received signal
353 jumps when there are flying birds and other obstructions to the optical signal on the propagation path.
354 This situation is automatically determined by the program. When this occurs, the data for that time
355 period are not processed. The AERISP and aerosol flux data are judged according to (a) three times
356 the standard deviation (SD) from the mean value and (b) three times the standard deviation from the
357 mean of differences between adjacent moment data. To determine the three times the SD from the
358 mean value, the trend is obtained by averaging over a two-hour period, then calculating the difference
359 between the measured value and the trend at each moment, calculating the mean and variance of the
360 difference, and considering a spike point if the difference is outside 3 times the SD. The 3 times the
361 SD of adjacent differences is determined by first calculating the difference between adjacent

362 moments and then calculating the mean and SD of the difference. Any data that deviates from the
363 mean by more than 3 times the SD is considered a spike point.

364 The data judged to be spikes will be supplemented by the average of adjacent moments. Of
365 course, the data processed according to this method appear to be completely missing for longer time
366 periods. For such cases, no further methods to realize supplementation are considered in this paper.
367 There are other errors in measurements made with the LAS due to specific reasons (Moene et al.,
368 2009); for example, the effect of spectral shape deviations using the von Karman model and
369 intermittent variations in the properties of this spectrum on the LAS signal are not considered in this
370 study.

371 Like for CO₂ flux calculations, EC calculations for aerosol flux were performed to obtain
372 aerosol fluxes, and several data quality control studies were conducted, such as coordinate system
373 rotations(Wilczak et al., 2001;Yuan et al., 2011), and WPL corrections(Webb et al., 1980).

374 **3 Experimental results**

375 In the following, the variation curves of conventional meteorological parameters during the
376 experimental period, individual examples of AERISPs, a comparison of the two methods for the
377 results of multiday continuous observations and a comparison of the two methods for the results of
378 flux measurements are presented to verify the reliability of the means of light propagation
379 measurements.

380 **3.1 General meteorological parameters and extinction** 381 **coefficients**

382 The variation curves of conventional meteorological parameters during the experiment,
383 including temperature, humidity, wind speed, wind direction, radiation and precipitation, and
384 extinction coefficient are shown in Fig. 2, where the extinction coefficient is calculated from the
385 visibility ($\beta_{ext} = 3.912/V$, V denotes visibility). Seven days during the experiment were sunny, and
386 four of the remaining eight days had rainfall. The temperatures on sunny days were characterized by
387 significant daily variations, with a minimum temperature of 0.4°C, and the maximum diurnal
388 temperature difference could reach more than 9°C. The relative humidity exceeded 80% for only a
389 few periods during sunny days. The wind speed was generally less than 3 m/s, and there were very
390 few periods of north wind with a speed greater than 3 m/s. There was no obvious prevailing wind
391 direction during the experimental period, and only the north wind was equivalent to the other
392 directions with a slight predominance. The meteorological conditions during the experiment were
393 similar to those of the local winter season. The extinction coefficient curve with time during the
394 experiment is given in Fig. 2(g). The pollution gradually increased from the 9th to Jan. 13th and
395 decreased on the 13th; from the 14th to the 20th, the pollution gradually increased and decreased on
396 the 20th. The meteorological conditions during the experimental period can be considered typical.

397 **3.2 Example results from measurements of the imaginary part** 398 **of the AERISP**

399 Before carrying out the comparison of the measurement results of the two methods for obtaining
400 the AERISP, the comparison of the measurement results of an individual example is given. The
401 experimental data measured from 2022-01-16 13:00-13:30 will be used here as an example to
402 illustrate the calculation of the AERISP, and the results will be given. This time period is midday on
403 a clear day (shown in Fig. 2e), and both the total radiation and sensible heat fluxes are large, so this
404 time period can be taken as a good typical example.

405 **3.2.1 Structure parameters obtained by light propagation**

406 The AERISP is first described using the light propagation method. The sequence of light
407 intensity signals obtained at the receiving end is shown in Fig. 3a. The time duration is 2022-01-16
408 13:00-13:30, and the sampling frequency is 500 Hz, thus there are 900000 data points in the time
409 series of light intensity fluctuations in Fig. 3a. The curve has both low- and high-frequency
410 fluctuations. Using spectral analysis and correlation analysis, the variance in the low-frequency part
411 of the logarithmic light intensity is $1.08e-4$, and the variance in the high-frequency part is $5.06e-4$.
412 The solid dots in Fig. 3b are the measured spectral densities of the logarithmic light intensity
413 fluctuations, and the black dashed lines and solid lines represent the results calculated by Eqs. (6)
414 and (9), respectively, and represent the contributions of the imaginary part and the real part. As seen
415 from the power spectral density curves of the logarithmic light intensity fluctuations in Fig. 3b, the
416 high-frequency part and the low-frequency part have different characteristics.

417 In the logarithmic plot, the low-frequency part is prominent with a much higher spectral density
418 than the high-frequency part. Theoretical analysis revealed that the low-frequency part corresponds
419 to the contribution of the imaginary part of the AERISP. The high-frequency part is flat plus high-
420 frequency attenuation. The high-frequency part corresponds to the contribution of the real part. The
421 part greater than 100 Hz is noise.

422 Based on the previous theoretical approach, the spectral density fitting for the low-frequency
423 part, while constrained by the low-frequency variance, yields an equivalent refractive index structure
424 parameter of $1.14 \times 10^{-25} m^{-2/3}$. Correspondingly, the structure parameter of the real part of the
425 refractive index, based on the high-frequency variance, is obtained as $2.54 \times 10^{-14} m^{-2/3}$.

426 **3.2.2 Obtaining the imaginary part of the AERISP based on** 427 **the spectrum**

428 The coefficients of the power spectral density curves are proportional to the refractive index
429 structure parameters from which they can be determined. The extinction coefficient structure
430 parameter can be deduced from the power spectral density of the extinction coefficient fluctuation,
431 and the temperature structure parameter can be deduced from the power spectral density of the
432 temperature fluctuation. The fluctuations in the extinction coefficient (Fig. 4a) and temperature (Fig.
433 4b) with time for the period 2022-01-16 13:00-13:30 are shown in Fig. 4. As shown in Fig. 4, the

434 extinction coefficient curve has more noise, while the temperature curve has less noise. On the
435 temperature fluctuation curve, there are five distinct ramp structures.

436 Power spectral analysis of the data in Fig. 4 was carried out to obtain the power spectral density
437 in Fig. 5. From the extinction coefficient power spectral density curve in Fig. 5a, it can be seen that
438 spectral densities greater than 0.05 Hz exhibit noise, and spectral densities less than 0.05 Hz have
439 inertial subregions. According to practical analysis, the inertial subregion ranges from 0.002 Hz to
440 the noise onset frequency. The motion of aerosol particles in the atmosphere conforms to the "-5/3"
441 law of turbulence. The extinction coefficient structure parameter was obtained by fitting the data in
442 the inertial subregion using Eq. (11) with a value of $3.9 \times 10^{-11} \text{m}^{-2} \text{m}^{-2/3}$, which was then
443 converted to the structure parameter of the imaginary part of the refractive index of $1.04 \times$
444 $10^{-25} \text{m}^{-2/3}$.

445 Correspondingly, as seen from the temperature fluctuation power spectrum density curve in Fig.
446 5b, almost no noise appears, which is mainly due to the small amount of noise in the temperature
447 signal itself, while the 1 Hz temperature data here are obtained by averaging the data collected at 10
448 Hz. The temperature structure parameter of $0.0218^\circ\text{C}^2 \text{m}^{-2/3}$ is obtained by fitting using Eq. (12),
449 which is converted to a refractive index real part structure parameter of $2.1 \times 10^{-14} \text{m}^{-2/3}$.

450 The imaginary part of the AERISP obtained by using a visibility meter and the real part of the
451 AERISP obtained by an ultrasonic anemometer are in good agreement with the previous results given
452 by using optical propagation methods.

453 **3.3 Comparison of all the results for the AERISP**

454 The previous section gives an individual example. A comparison of all the data during the
455 experiment is given below, as shown in Figs 6 and 7.

456 A comparison of the time series of AERISPs measured by the two methods is given in Fig. 6,
457 where Fig. 6a shows the time series of the imaginary part of the AERISP and Fig. 6b shows the time
458 series of the real part of the AERISP. There are large fluctuations in the imaginary part of the AERISP
459 during the experimental period. This trend is close to that of the aerosol extinction coefficient. Figure
460 6a shows that there is no obvious daily variation characteristic. The trend agreement of the results
461 obtained by the two methods is very good. From Fig. 6b, it can be seen that the real part of the
462 AERISP on sunny days has obvious daily variation characteristics; these characteristics are large
463 during the day and small at night. The agreement of the results obtained by the two methods is good
464 during the day (8:00-17:00), and at night, the results obtained by the light propagation method are
465 greater than those of the large point measurements.

466 Scatter plots of the results of the measurements of the two methods are given in Fig. 7. Figure
467 7a shows the scatter plot of the results of the two methods for the imaginary part of the AERISP with
468 almost the same correlation coefficient R^2 for daytime and nighttime, while Fig. 7b shows the scatter
469 plot of the results of the two methods for the real part of the AERISP with a correlation coefficient
470 of real R^2 of 0.74 for daytime and 0.15 for nighttime. This shows that the correlation coefficients of
471 the imaginary part of the AERISP obtained by the two methods are almost equal during both daytime
472 and nighttime, and the correlation coefficient of the real part of the AERISP obtained by the two
473 methods is smaller at night than during the daytime. This shows that the spatial distribution of aerosol
474 at night may be more homogeneous than the temperature distribution. The reason for this difference
475 may be that the temperature distribution in the overlying surface of the campus at night is not uniform,

476 and weak turbulence does not produce strong mixing, resulting in a nonuniform distribution of the
 477 real part of the AERISP. There are no strong aerosol emission sources on the night-time campus, so
 478 the distribution of the imaginary part of the AERISP behaves more uniformly.

479 **3.4 Velocity-extinction coefficient correlation for a single point**

480 To calculate aerosol fluxes using EC techniques, a delayed correlation of the vertical velocity
 481 and extinction coefficient is needed. The delayed correlation curves of the vertical velocity and
 482 extinction coefficient are given in Fig. 8.

483 The horizontal coordinate of the delay correlation curve in Fig. 8 is the delay time τ , and the
 484 vertical coordinate is the delay correlation. From Fig. 8, it can be seen that at $\tau = -2$ s, the correlation
 485 curve has an obvious extreme value, which is also the minimum value of the delay time for a duration
 486 of 300 s. The minimum value is $-5.22e-6 \text{ m}^{-1}$. The extreme value of the correlation curve does not
 487 appear at 0 s because there is a distance of approximately 0.20 m between the sensing element of the
 488 visibility meter and that of the ultrasonic anemometer. Here, we present the cases with obvious
 489 extremes, and there are some cases where no obvious extremes appear. In such cases where there are
 490 no significant extremes, the value associated with a delay time of 0 seconds is taken.

491 **3.5 Flux**

492 The AERISP was given in the former part, and the aerosol vertical transport flux can be
 493 estimated for the duration of 2022-01-16 13:00-13:30 according to Eq. (16),

$$494 \quad F_{a_LAS} = 0.567 * \left(\frac{9.8}{283}\right)^{\frac{1}{2}} * (1.01 \times 10^6)^{\frac{1}{2}} * (2.54 \times 10^{-14})^{\frac{1}{4}} * 6216 * (1.14 \times 10^{-25})^{1/2} * 18 * 10^9 =$$

$$495 \quad 1.60 \mu\text{gm}^{-2} \text{ s}^{-1} \quad (18)$$

496 where $a=0.567$, $T=283 \text{ K}$, $g=9.8 \text{ m/s}^2$, $R_{TN} = 1.01 \times 10^6 \text{ K}$, $C_{n,Re}^2 = 2.54 \times 10^{-14} \text{ m}^{-2/3}$, $R_{MN} =$
 497 $6216 \text{ Kg} \cdot \text{m}^{-3}$ (Yuan et al., 2015), $C_{n,Im}^2 = 1.14 \times 10^{-25} \text{ m}^{-2/3}$, $z=35 \text{ m}$, and $d=17 \text{ m}$.

498 Similarly, the aerosol flux is obtained from the eddy covariance method according to Eq.(17)

$$499 \quad F_{a_EC} = -0.522 \times 10^{-6} * 6216 * 10^9 * \frac{0.65 \times 10^{-6}}{4\pi} = -1.67 \mu\text{gm}^{-2} \text{ s}^{-1} \quad (19)$$

500 where $\overline{w'\beta_{ext}}' = -0.522 \times 10^{-6} \text{ s}^{-1}$, $R_{MN} = 6216 \text{ Kg} \cdot \text{m}^{-3}$, and $\lambda = 0.65 \times 10^{-6} \text{ m}$.

501 From the previous calculations, we can see that during the half hour from 2022-01-16 13:00-
 502 13:30, the absolute values of the aerosol fluxes obtained by the two methods are very close but of
 503 opposite signs. Since the LAS method based on light propagation cannot determine the direction of
 504 flux transport, only the magnitude of the flux can be determined. This is similar to the fact that the
 505 estimation of surface sensible heat fluxes using an LAS provides information about only the
 506 magnitude but not the direction. There are some judgments for estimating the direction of sensible
 507 heat flux using a LAS, such as those based on sunrise and sunset times and atmospheric stability
 508 (Zhao et al., 2018). Here, a negative flux indicates the deposition of aerosol particles. Because the
 509 experimental site is a campus, there is almost no source of aerosol particle emission in the overlying
 510 surface, which is manifested as a sink of aerosol particles inside the city. Therefore, the direction of
 511 aerosol flux measurements based on the LAS needs to be judged based on the nature of the surface.

512 The results of aerosol flux calculations throughout the experiment, except for two days of rain,
 513 the 22nd and 23rd days. are given in Fig. 9. Figure 9a shows the absolute values of the aerosol
 514 vertical transport fluxes measured by the two methods based on the imaginary part of the AERISP

515 and EC methods, and Fig. 9b shows the aerosol vertical transport fluxes with signs for transport
516 direction measured, which correspond to the rectangular-point line in Fig. 9a. The trend of aerosol
517 fluxes obtained by the two methods given in Fig. 9a is consistent with the diurnal variation in aerosol
518 fluxes on sunny days, with larger values of aerosol fluxes at noon. At night, the aerosol flux values
519 are lower. As shown in Fig. 9(a), the absolute value of the aerosol flux obtained by the LAS is greater
520 than that obtained by the EC at noontime on 10-11 Jan, 2022. This is because the imaginary parts of
521 the AERISP obtained by the LAS are larger than those obtained by the EC, as shown in Fig. 6a.
522 Another possible reason is that it was a cloudy day during both the 10th and 11th days, there was a
523 weak rainfall process on the 10th day at 16:00, and the winds on the 10th and 11th days were lighter
524 and had a greater change in direction. The turbulence during noontime on 10-11 is weaker, resulting
525 in an inhomogeneous horizontal distribution and a large difference in measurements between the
526 two methods.

527 A comparison of Fig. 9a and Fig. 9b, reveals that the aerosol flux is negative at noon on clear
528 days, indicating that the turbulence is strong at noon, which enhances the downward transport of
529 aerosol particles. This study was conducted on a campus with no emission sources, and the
530 downward flux was reasonable; in fact, there was an upward flux measured by the EC method if
531 there were emission sources in the observation area (Ren et al. 2020).

532 **4 Conclusion and discussion**

533 To validate the previously developed method of measuring the AERISP and aerosol mass flux,
534 this paper theoretically organizes the concept of the AERISP, introduces two methods for measuring
535 the AERISP and estimating the aerosol vertical transport flux by using the AERISP and EC methods,
536 and carries out field observation experiments in an urban area. The experimental results show that
537 the AERISPs estimated by the two methods are in good agreement, and the aerosol vertical transport
538 fluxes obtained by the two methods based on the AERISP and EC are in good agreement.

539 According to the experimental results, the imaginary part of the AERISP expresses the intensity
540 of the fluctuation in the attenuation of light during transmission. When the air-transparent band is
541 used, the imaginary part of the AERISP characterizes the intensity of the fluctuation in the extinction
542 coefficient of the aerosol.

543 The aerosol flux is related to both the fluctuations in aerosol concentration and the intensity of
544 atmospheric turbulence. When there is an aerosol emission source on the overlying surface, the
545 aerosol flux is positive, transporting aerosol particles upwards. When there is no aerosol emission
546 source in the overlying surface, the overall performance is aerosol particle deposition and
547 downwards flux transport. In general, urban green lands are areas of aerosol particle deposition,
548 while ocean and desert surfaces can often be viewed as source areas for aerosols. The large difference
549 in the real part of the AERISP measured by the two methods at night also contributes to the large
550 difference in the aerosol fluxes obtained by the two methods at night.

551 From the experimental results, we can also see that, as a comparison, this paper also gives
552 results for the temperature refractive index structure parameters, and as shown in Fig. 6, the trends
553 for the structure parameters in the real and imaginary parts of the AERISP are different, indicating
554 that temperature fluctuations and aerosol concentration fluctuations are uncorrelated. The purpose
555 of this paper is to illustrate the physical significance of the imaginary part of the AERISP obtained
556 using the LAS technique and to obtain the aerosol vertical transport flux based on the AERISPs.
557 When inverting the imaginary part of the AERISP using the light propagation principle, it is assumed

558 that the aerosol concentration fluctuations are not correlated with the temperature fluctuations. This
559 assumption cannot be proven theoretically. From the experimental results, as shown in Fig. 6, the
560 trends of the real and imaginary parts of the AERISP are different, indicating that the temperature
561 fluctuations and the aerosol concentration fluctuations are uncorrelated. This phenomenon shows
562 that the two sources are different and are basically consistent with the actual situation. This also
563 shows that the assumptions of the theory for obtaining the imaginary part of the AERISP are
564 reasonable.

565 To compare with aerosol transport fluxes obtained based on the AERISPs, this paper uses a
566 delay correlation between the visibility meter and vertical wind speed to obtain aerosol vertical
567 transport fluxes. Currently, a modified visibility meter is utilized to obtain 1 Hz visibility data, after
568 which the extinction coefficient is obtained. The extinction coefficient power spectrum in Fig. 5a
569 shows that there is a large amount of noise in the high-frequency part. The signal-to-noise ratio of
570 the extinction coefficient data is too low compared to the temperature fluctuation or velocity
571 fluctuation, which introduces a large error in the calculation of the aerosol flux. Although the overall
572 trend magnitude agreement of the fluxes obtained by the two methods is good enough to show that
573 the two methods can be corroborated with each other, there are still differences in the details;
574 however, technical methods are required to improve the performance of the instrument and to obtain
575 high-quality aerosol extinction coefficient data to carry out measurements of vertical aerosol
576 transport fluxes based on the EC method at a single point.

577 **Data availability.** Requests for data that support the findings of this study can be sent to
578 rmyuan@ustc.edu.cn.

579 **Competing interests.** The authors declare that they have no conflict of interest.

580 **Author contributions.** Renmin Yuan and Hongsheng Zhang designed experiments and wrote the
581 manuscript; Renmin Yuan, Jiajia Hua, Hao Liu, Xingyu Zhu and Peizhe Wu carried out
582 experiments; Renmin Yuan analyzed experimental results. Jianning Sun revised the manuscript and
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587 **Appendix: Comparison of fluxes between 10** 588 **Hz and 1 Hz**

589 To determine the high frequency loss due to the use of 1 Hz data for flux calculations, the T-w
590 covariance was used to perform an analytical comparison between the fluxes obtained by sampling
591 the data at 10 Hz and the fluxes obtained by gaining the data at a frequency of 1 Hz. The data from
592 January 9 and 23, 2022 were processed, and the fluxes corresponding to different sampling
593 frequencies were compared and are shown in Fig.10. There are two ways to obtain 1 Hz data: one is
594 directly obtained at 1 Hz sampling frequency (shown in Fig. 10a), and the other is 1 Hz data obtained
595 by averaging 10 Hz data over 10 data points (shown in Fig. 10b). In comparison, the flux calculated
596 from the 1 Hz data obtained by averaging 10 data points is smaller (slope of 0.97). This indicates a
597 slower response of the instrument. This is the case for the visibility meter, for which a slower

598 response was used in this study. Based on the linear fit results and the root mean square error (RMSE)
599 in Fig. 10, the difference in the fluxes between 10 Hz and 1 Hz is less than 5%.

600 Overall, the error due to the lower sampling frequency of 1 Hz is much smaller than the
601 difference between the two methods discussed in this study.

602

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Table 1 details of all the instruments

Meteorological elements	Manufacturer type	Sampling frequency (Hz)	Height (m) Above building top
LAS	Self-developed	500	18.0
3-D sonic anemometer	Campbell CSAT3	10	18.0
Visibility	Campbell CS120	1	18.0
Wind speed and direction	03001 R.M. Young	1	2.0, 4.5, 8.0, 12.0,18.0
Temperature and humidity	Vaisala HMP155A	1	2.0, 4.5, 8.0, 12.0,18.0
Radiation	Kipp&Zonen CNR4	1	16.0
Precipitation	TE525 Tipping Bucket	1	1.0

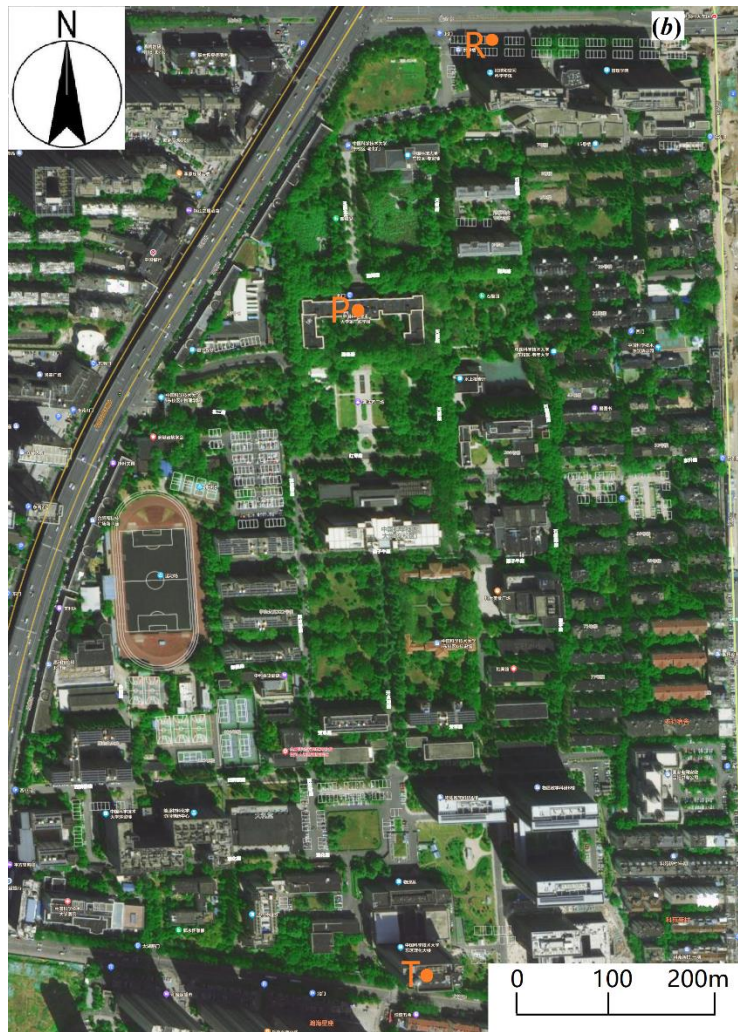
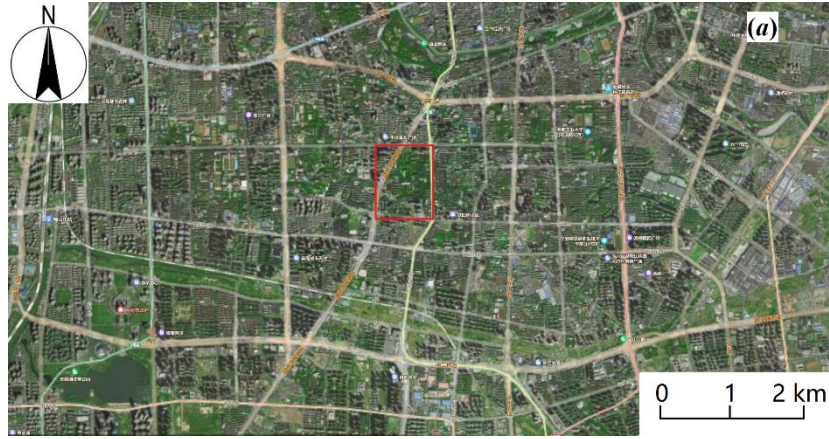
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695 **Figures**

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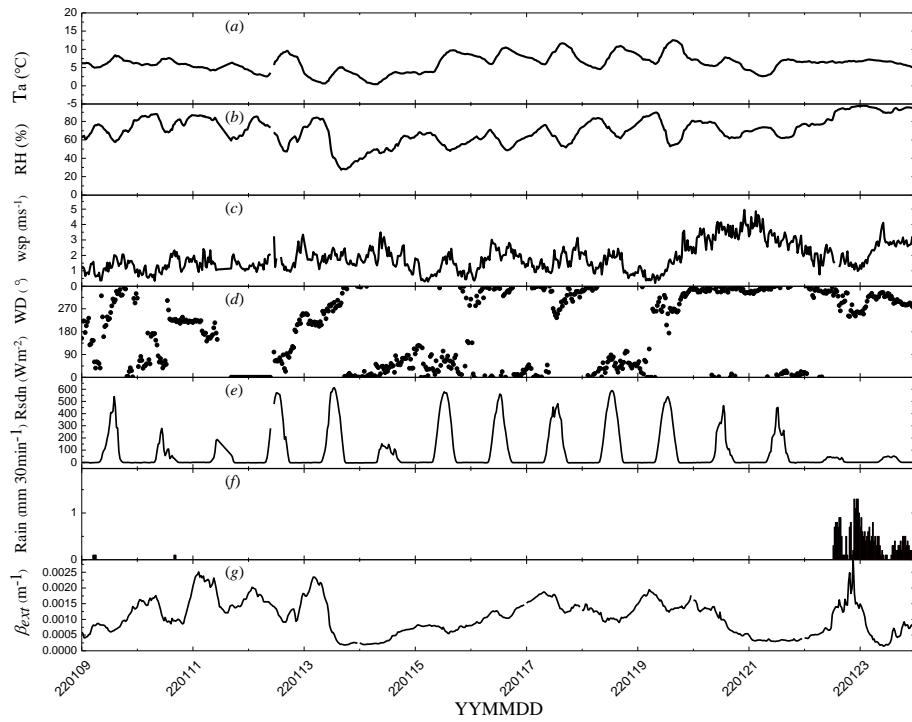


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701 Figure 1. Photographs of the measurement site. (a) Map of Hefei city and (b) expanded view of the
 702 measurement site on the USTC campus, which is marked by the red rectangle in (a). Points T and R in
 703 (b) show the locations of the transmitter and receiver, respectively. Point P in (b) marks the meteorological
 704 tower position. There are four heavy traffic roads surrounding the measurement site. Figure 1a and b @
 705 Baidu are from the following website:

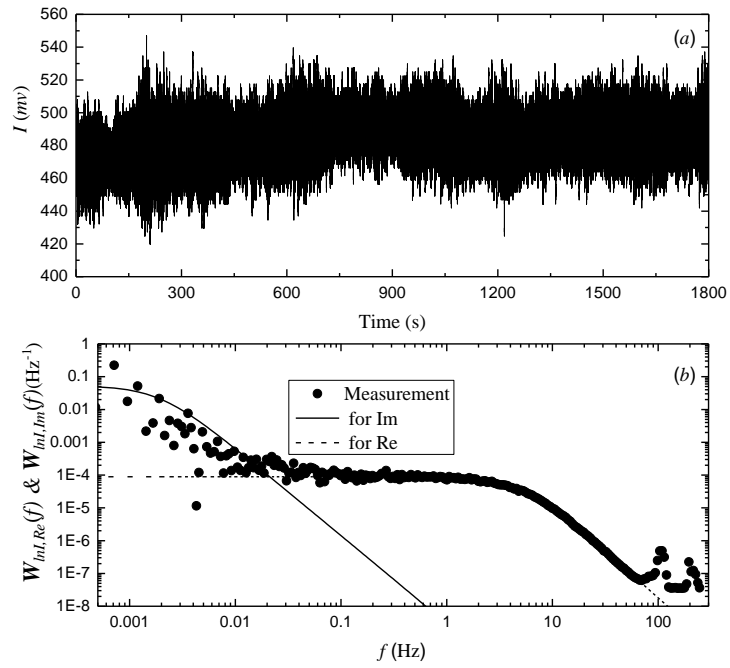
706 https://map.baidu.com/@13055953.105500832,3719556.851423825,15.3z/maptype%3DB_EART
 707 H_MAP



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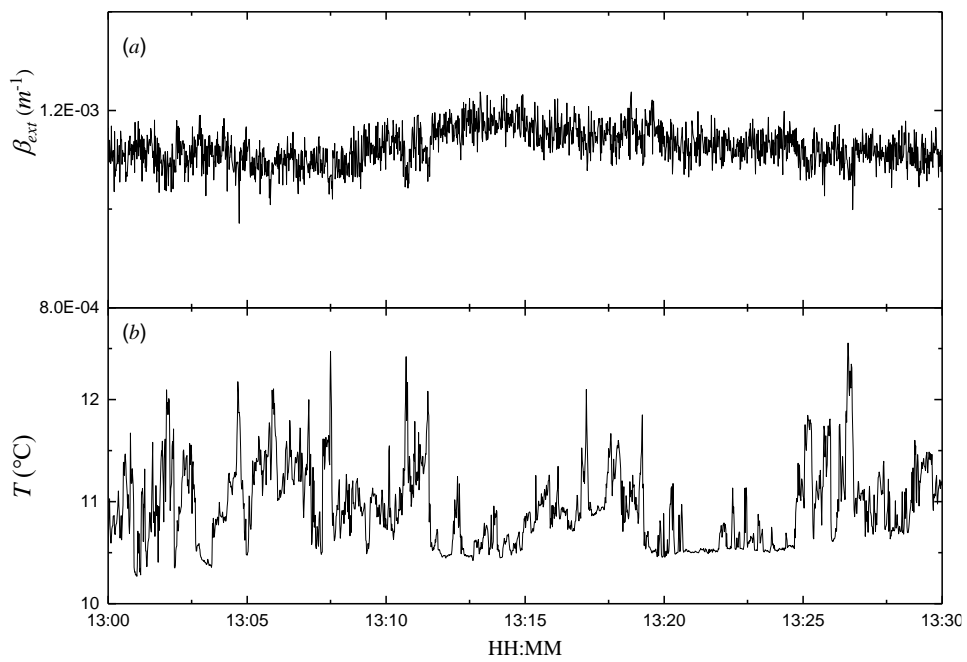
710 Figure 2. Temporal variations in the (a) air temperature (T), (b) relative humidity (RH), (c) wind
 711 speed (wsp), (d) wind direction (WD), (e) total radiation (Rsdn), (f) precipitation (Rain), and (g)
 712 extinction coefficient (β_{ext}). The details can be found in the text.



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714 Figure 3 Temporal variations in the light intensity received by the LAS and (b) power spectral

715 density of the logarithm of the light intensity during 2022-01-16 13:00-13:30.

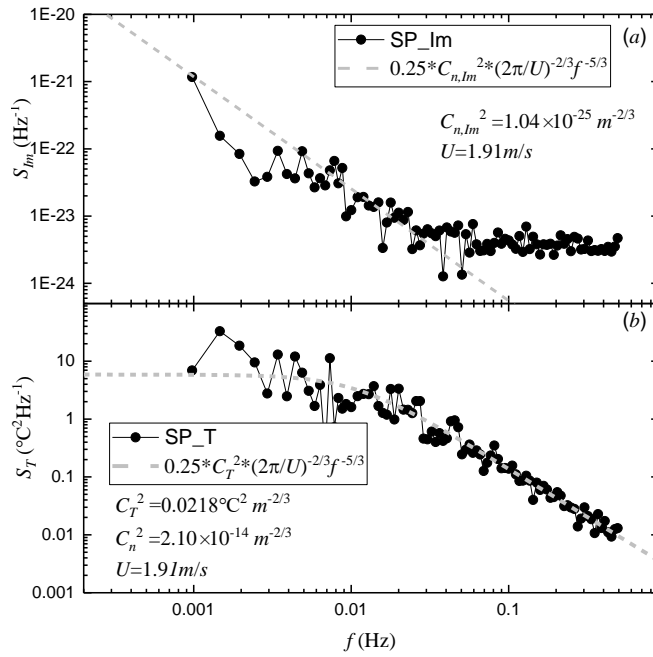


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717 Figure 4 Temporal variations in the (a) extinction coefficient and (b) air temperature during 2022-

718 01-16 13:00-13:30.

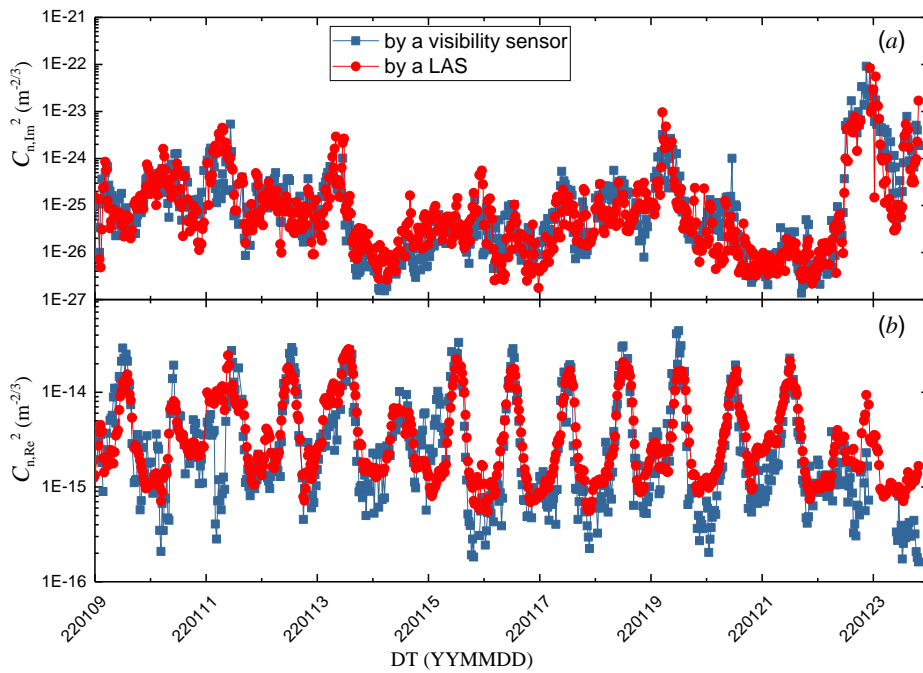
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721 Figure 5. Power spectral density of the (a) extinction coefficient and (b) air temperature during 2022-

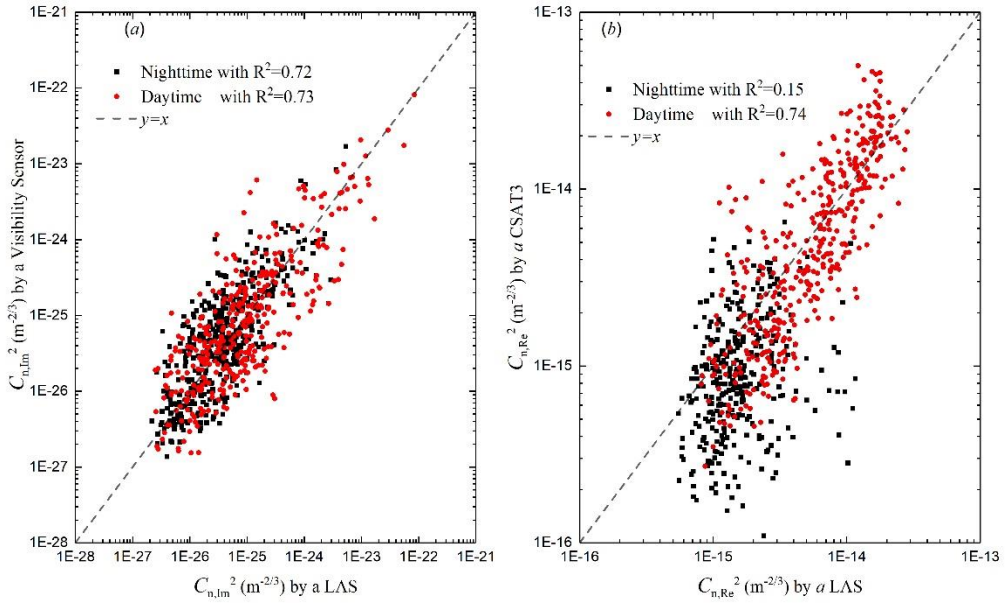
722 01-16 at 13:00-13:30.



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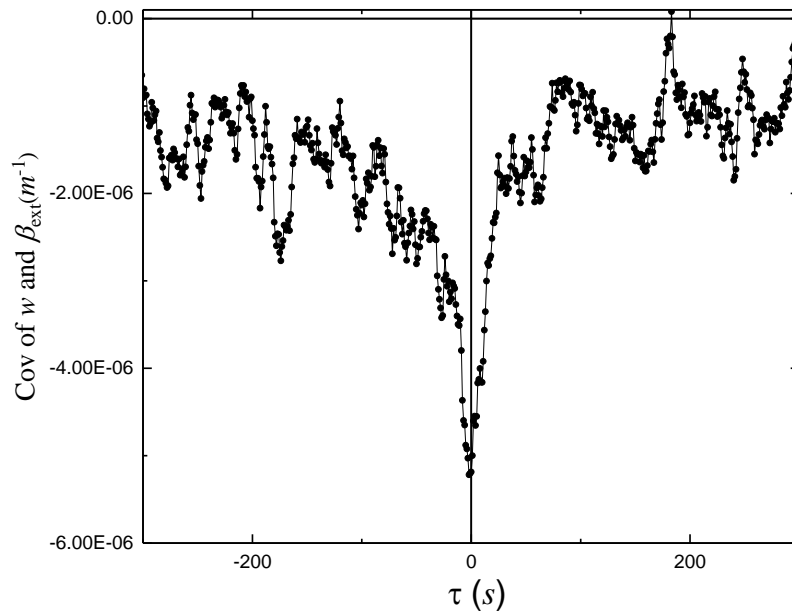
724 Figure 6. Temporal variations in (a) the imaginary part and (b) real part of the AERISP during 09-

725 23 Jan. 2022.



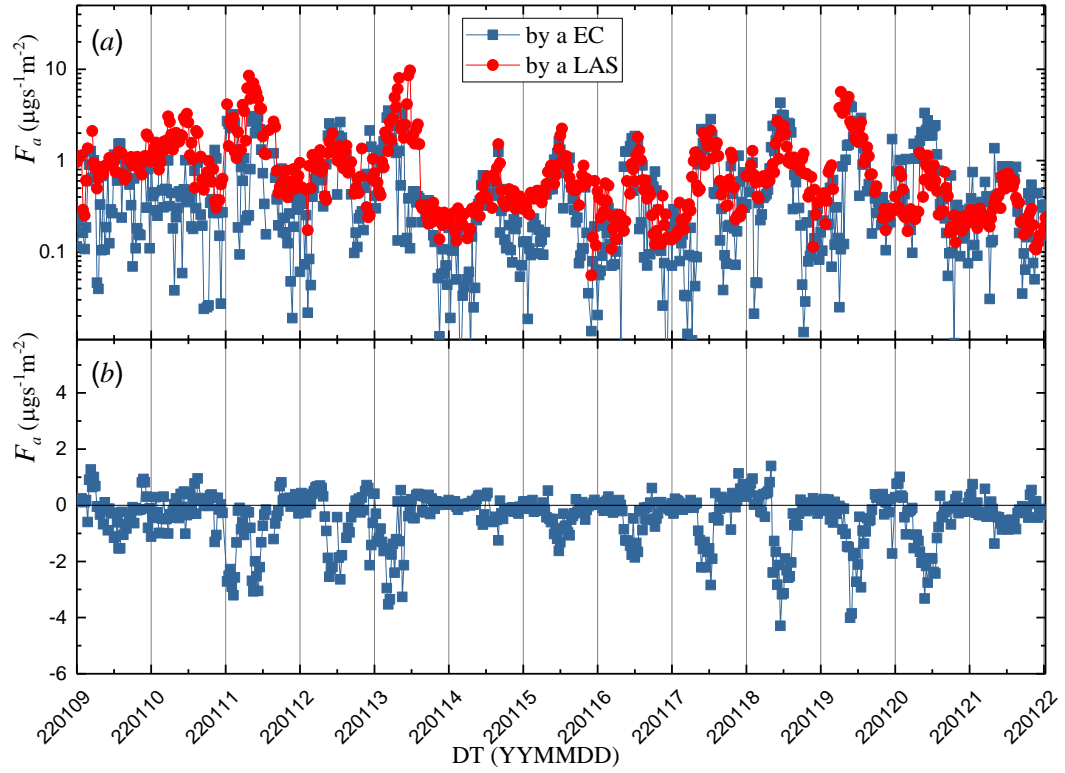
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Figure 7. Comparison of (a) the imaginary part and (b) real part of the AERISP during 09-23 Jan. 2022. The red solid circles indicate daytime and the black solid rectangles indicate nighttime.



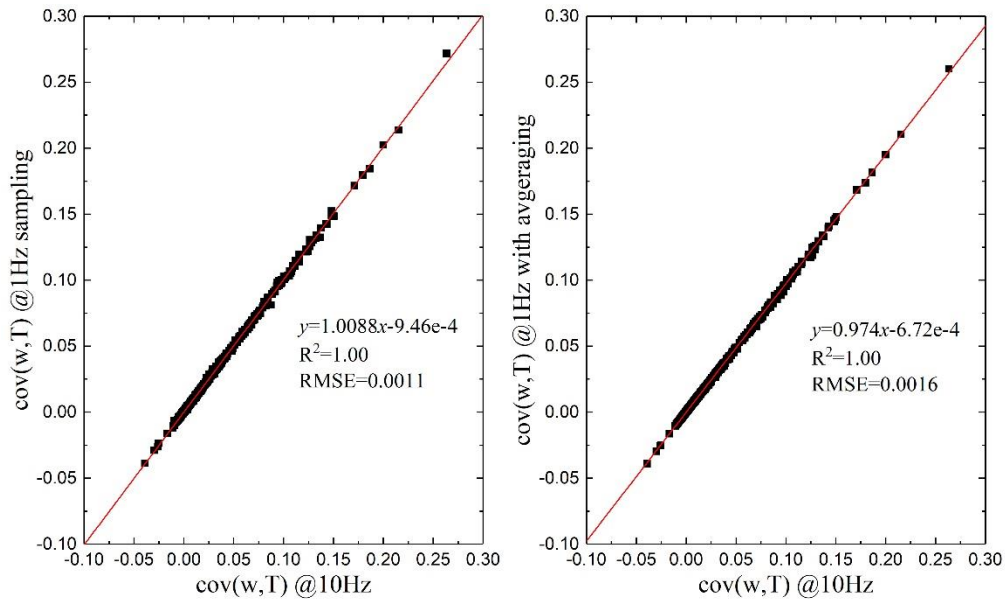
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Figure 8. Delay covariance between the extinction coefficient and vertical velocity during 2022-01-16 13:00-13:30.



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Figure 9. Temporal variations in (a) absolute value of aerosol flux based on the AERISP and EC methods and (b) aerosol flux based on the EC methods during 09-21 Jan. 2022.



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Figure 10 Comparison of covariance of w and T between 10 Hz and 1 Hz, with a 1 Hz sampling rate (a) and 1 Hz data obtained by averaging 10 Hz data over 10 data points (b)