

Quantum Yields of CHDO above 300 nm

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Abstract : *The photolysis of mono-deuterated formaldehyde, CHDO, is a critical process in the deuterium-enrichment of stratospheric hydrogen formed from methane. In this work, a consistent description of the quantum yields of the molecular and radical channels of the CHDO photolysis is deduced from literature data. The fluorescence measurements of Miller and Lee (1978) provided a first data set to deduce the product quantum yields. An alternative analysis is based on the measured quantum yield spectrum for the radical channel of the CD₂O photolysis by McQuigg and Calvert (1969), which is corrected for wavelength dependency and combined with the CH₂O quantum yield spectrum to provide an approximation for CHDO. Both approaches provide consistent results. Finally, the findings of Troe (1984, 2007) enable the specification of the pressure dependence of the quantum yield for CH₂O and CD₂O and, hence, for CHDO. We find that the radical channel does not show a pressure dependence, whereas the molecular channel is dominated by tunneling and quenching processes. Simplified representations are given that are readily implemented in kinetic atmospheric models. As an example of their application, the altitude dependence of the ratio of $J(\text{CHDO} \rightarrow \text{HD} + \text{CO})$ and $J(\text{CH}_2\text{O} \rightarrow \text{H}_2 + \text{CO})$ is provided. Also, the importance of the photolysis of formaldehyde in the atmosphere is presenting the altitudinal dependence of the isotopic fractionation through the yield of the HD channel.*

1 **1. Introduction**

2

3 Measurements over the last decades showed that molecular hydrogen, H₂, in the stratosphere
4 is enriched in deuterium compared to H₂ in the troposphere (see e.g.: Ehhalt and Volz, 1976;
5 Gerst and Quay, 2001; Rahn et al., 2003; Rice et al., 2003; Röckmann et al., 2003; McCarthy
6 et al., 2004; Rhee et al., 2006). Gerst and Quay (2001) suggested that this enrichment could be
7 due to the differential isotope fractionation in the photo-oxidation of methane. Measurements
8 of the vertical profiles of the isotope content in H₂ and CH₄, available since 2003, allowed the
9 interpretation and modeling of the observed enrichment (see e.g. Pieterse et al., 2011). The
10 methane photo-oxidation consists of various reaction steps, each of which contribute kinetic
11 isotope effects, KIE, that have to be considered (e.g. Feilberg et al., 2005; Mar et al., 2007).
12 The last but critical step in the reaction chain to produce the hydrogen isotope HD from the
13 mono-deuterated isotopologue of formaldehyde, CHDO, is its photolysis.

14

15 Compared to CH₂O, the available data for the mono-deuterated isotopologue CHDO are
16 scarce. Only its spectrum was measured (c.f. Mainz Spectral Atlas, Keller-Rudek and
17 Moortgat, 2021). The quantum yields for the molecular and the radical fragmentation
18 branches of the CHDO photolysis, as well as the rate constants for the quenching reactions
19 were not measured at all or with insufficient accuracy. Thus, despite its importance for the
20 atmospheric production of HD, the photolysis of CHDO is still poorly defined; at this time, it
21 is the most uncertain factor in the overall fractionation of formaldehyde. For example, the
22 measured or estimated fractionation factors for the molecular channel range from 1.08 to 1.82
23 (e.g. Feilberg et al., 2005; Rhee et al., 2006, Mar et al., 2007; Nilsson et al., 2009; Röckmann
24 et al., 2010). Moreover, the measurements by Nilsson et al. (2009) are the only ones
25 considering the pressure dependence of the fractionation factor due to reactions R3, R4, and
26 R7 (see Table 1).

27

28 In this work, we aim to provide information for the modeling of CHDO photochemistry for
29 atmospheric conditions, i.e. for a limited domain of temperature and pressure, by deducing the
30 molecular and radical quantum yields Φ^{mol} and Φ^{rad} for CHDO from literature information,
31 based on the scant data available and supplemented by a number of plausible assumptions.
32 We do this based on two approaches: the first is based on the fluorescence measurements of
33 Miller and Lee (1978) and literature data on energy transitions (e.g. Yeung and Moore, 1973;
34 Chuang et al., 1987; Osborn, 2008; Fu et al. 2011). The second approach assumes that the

35 measurements of McQuigg and Calvert (1969) can be corrected via the comparison of the
 36 CH₂O measurement with later experiments (see e.g. the overview by Röth and Ehhalt, 2015).
 37 The photochemistry derived thus is then used to assess the altitudinal dependence of the
 38 CHDO isotopic fractionation.

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 40

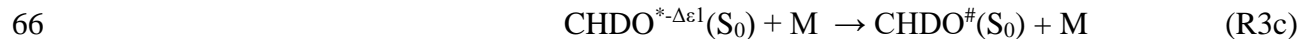
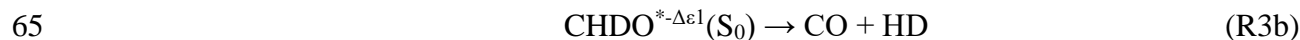
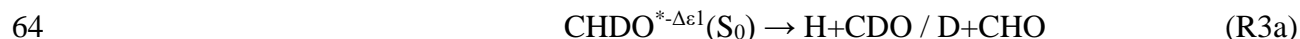
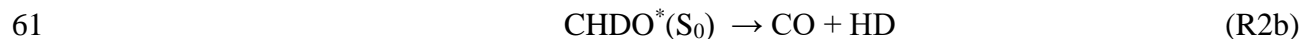
41 2. Photolysis reaction mechanism

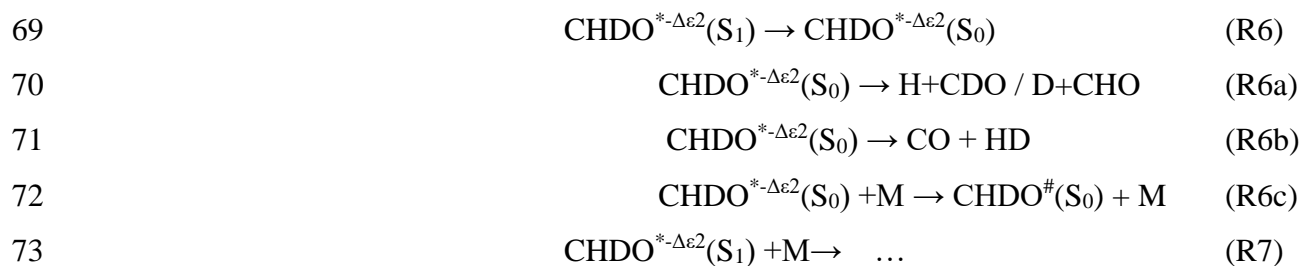
42

43 Based on the available literature (e.g.: Aràujo et al., 2009; Breuer and Lee, 1971; Chuang et
 44 al., 1987; Yamaguchi et al., 1998) we propose a photolytic reaction scheme of CHDO in
 45 Table 1, analogous to that of CH₂O (Röth and Ehhalt, 2015). The scheme involves a
 46 cascading series of fragmentation channels competing with stepwise quenching by collisional
 47 energy loss, starting at the excited singlet state S₁. Reactions via the triplet state of CHDO are
 48 not considered here, as they are only accessible at wavelengths below 300 nm (Aràujo et
 49 al.,2009), while we concentrate on wavelengths above this limit in this work. Under
 50 atmospheric conditions, which are considered here, the system is thermalized.

51

52 **Table 1** : Reaction scheme of the photolysis of CHDO occurring over the S₁ and S₀ electronic
 53 singlet states. $-\Delta\epsilon_1$, $-\Delta\epsilon_2$ indicate the collisional energy losses to bath gas M in the respective
 54 reactions. This quenching is stepwise but is represented here for energies above/below the
 55 threshold for dissociation; the asterix * stands for excitations able to lead to bond breaking,
 56 whereas the index # indicates lower energies and lead ultimately to thermalized CHDO.





75 After excitation of the ground state $\text{CHDO}(\text{S}_0)$ (R0) by a photon of a given wavelength, the
 76 excited reaction product $\text{CHDO}^*(\text{S}_1)$ decays by fluorescence (R1), or transitions to the S_0
 77 ground state surface as an excited CHDO^* molecule with either all available energy (R2) or
 78 with a variable amount of energy $-\Delta\epsilon_1$ lost by quenching (R3). These excited $\text{CHDO}^*(\text{S}_0)$ and
 79 $\text{CHDO}^{*\Delta\epsilon_1}(\text{S}_0)$ can in turn be quenched by the bath gas in a cascading series (R2c, R3c, R6c),
 80 at each energy level competing with fragmentation to radicals H+CDO/D+CHO (R2a,
 81 R3a,R6a) or to molecular products CO+HD (R2b, R3b,R6b), as described for CH_2O by
 82 Yeung and Moore, (1973). Alternatively, the excited $\text{CHDO}^*(\text{S}_1)$ can lose an amount of
 83 energy by quenching, but remain on the S_1 excited electronic surface (R4). This state can then
 84 undergo processes as above, i.e. decay by fluorescence (R5), transition to the S_0 ground state
 85 without (R6) or with (R7) energy loss by quenching, where once again it can undergo further
 86 quenching (R6c) in competition with fragmentation (R6a, R6b). Overall, this scheme
 87 represents a cascading series of quenching steps competing against decomposition and
 88 fluorescence. Only the first few steps in the cascade are represented, but more cascading steps
 89 are possible at lower internal energies. According to the analysis of the fluorescence
 90 measurements by Miller and Lee (1978), these lower-energy reactions are not critical and
 91 need not be considered in detail. Here, R7 simply represents the summation of all subsequent
 92 cascades, from which negligible channels such as *e.g.* the fluorescence channels are omitted.
 93 Schematic energy diagrams for this reaction mechanism were already depicted in the
 94 literature (*e.g.* Fig.3 in Aràujo et al., 2009 and Fig.1 in Chuang et al., 1987), and are not
 95 repeated here.

96

97 The quantum yield Φ^{rad} represents the combined fragmentation to radicals (R2a, R3a, R6a),
 98 while summed fragmentation through the molecular branches (R2b, R3b, R6b) is described by
 99 the quantum yield Φ^{mol} . The total photolysis quantum yield Φ^{tot} , i.e. the decay of excited
 100 formaldehyde into products other than its ground-state, can be experimentally derived from
 101 the observed CO production, where CDO and CHO radical fragments react with O_2 to form

102 CO and HO₂ / DO₂. The quantum yield of the fluorescence is always less than 1% (Miller
103 and Lee, 1978) and is omitted henceforth.

$$104 \quad \Phi^{tot} = \Phi^{mol} + \Phi^{rad} \quad (F1)$$

105 Obviously, the sum of Φ^{tot} and Φ^{quench} , the summed yield of the quenching reactions (R2c,
106 R3c, R6c), must equal 1 at any wavelength $h\nu$.

$$107 \quad \Phi^{tot} + \Phi^{quench} = 1 \quad (F2)$$

108
109

110 3. Analysis of fluorescence measurements

111

112 From the fluorescence measurements of Miller and Lee (1978) the quantum yields of both the
113 fluorescence and the total non-CHDO products can be derived. The contribution of the second
114 step in the reaction cascade is small at low pressure (see later), so we assume that Table X
115 provided by these authors directly gives the reaction rate constants k_1 and k_2 , where k_1 equals
116 the reciprocal lifetime $\tau_{radiation}$ listed and $1/k_2$ is the non-radiative lifetime. Similarly, the
117 constants k_5 and k_6 are determined by the lifetimes of the next lower vibrational level.

118

119 The reaction constants k_3 , k_4 , and k_7 can be deduced from the pressure dependence of the
120 CHDO fluorescence quantum yield in the Table II of Miller and Lee (1978). In the present
121 paper only the quantum yields at pressures above 1 Torr are considered, where the Ar bathgas
122 used is assumed to have similar collisional properties as air (Hirschfelder et al., 1954). For
123 each wavelength the pressure dependence of the data is fitted by a Simplex algorithm
124 according to Nelder and Mead (1965) by formula F3 for the fluorescence quantum yield Φ_F .

$$125 \quad \Phi_F(M) = \frac{k_1}{\alpha} + \frac{k_4[M] \cdot k_5}{\alpha \beta} \quad (F3)$$

126 with $\alpha = k_1 + k_2 + k_3[M] + k_4[M]$ and $\beta = k_5 + k_6 + k_7[M]$. For consistency we only used
127 the $2^1 4^j$ transitions, only.

128

129 The corresponding reaction constants are listed in Table 2. With this data set the experimental
130 fluorescence measurements are well fitted as shown in Figure 1 where, to improve the clarity
131 of the fit, only the pressure dependent part $\theta(M)$ of equation F3 is plotted vs pressure:

$$132 \quad \theta(M) = \frac{k_1}{\phi_F(M)} - (k_1 + k_2) \quad (F4)$$

133

134 **Table 2** : Results of the least square fit of the quantum yields of CHDO (Miller and Lee
 135 (1978). k_1 , k_2 and k_5 , k_6 are literature data (Miller and Lee, 1978), k_3 , k_4 , and k_7 are deduced
 136 from these data.

Wavelength [nm]	k_1 [10^5s^{-1}]	k_2 [10^8s^{-1}]	k_3 [$10^{-11}\text{cm}^3\text{s}^{-1}$]	k_4 [$10^{-11}\text{cm}^3\text{s}^{-1}$]	k_5 [10^5s^{-1}]	k_6 [10^8s^{-1}]	k_7 [$10^{-12}\text{cm}^3\text{s}^{-1}$]
314.0	3.03	1.79	29.7	4.59	2.78	0.50	0.57
318.3	2.50	1.32	15.4	3.48	2.50	0.40	1.15
325.7	2.78	0.50	10.9	1.77	3.57	0.22	1.79
330.8	2.50	0.40	4.81	1.05	2.44	0.13	1.35
338.6	3.57	0.22	4.89	0.84	3.45	0.07	0.77
344.4	2.44	0.13	5.95	2.78	2.40*	0.06*	1.39
352.9	3.45	0.07	2.38	0.76	4.00*	0.03*	1.24

137 * estimated by extrapolation of the other values

138
 139 The energy transferred in reaction R2 is either quenched to form a stable molecule
 140 CHDO[#](S₀) or used to drive fragmentation to molecular (CO + HD) or radical products
 141 (H+CDO / D+CHO). Hence, the reactions R2a and R2b form part of the product-forming
 142 channel. Analogously, the secondary reactions of the pressure dependent reactions R3 and R4
 143 lead to products via the reactions R3a and R3b, respective R6a and R6b. With this, the total
 144 product quantum yield of the photolysis of CHDO is the sum of the individual product
 145 quantum yields across all channels k, where the index k=2, 3, 6 stands for the non-radiative
 146 reactions R2, R3, and R6.

147 The individual product quantum yield can be approximated by

$$148 \quad \Phi_k^{tot} = \frac{1}{1+a \cdot \exp\left(\frac{\varepsilon_k - \varepsilon_0}{b}\right) \frac{[M]}{[M_0]}} \quad (\text{F5})$$

149 analog to the publication by Röth and Ehhalt (2015) on CH₂O.

150 In equation F5, ε_2 is the excitation energy of the photolysis reaction. The energies ε_3 and ε_6
 151 are related to ε_2 by the approximated energy transfer in a collision, respective by the averaged
 152 width of the band intervals, given by $\varepsilon_3 = \varepsilon_2 - 0.0124$ eV (Troe, 2007) and $\varepsilon_6 = \varepsilon_2 - 0.13$ eV
 153 (Miller and Lee, 1978). The pivot wavelength $1/\varepsilon_0$ is 348.6 nm, as published in Nilsson et al.
 154 (2014) from quantum chemical calculations of the barriers to dissociation of H-CHO, H-
 155 CDO, D-CHO, and D-CDO.

156
 157 The total quantum yield of the products (molecules plus radicals) can be deduced from the
 158 rate constants of Table 2 and the measurements of Nilsson et al. (2010, 2014), who

159 investigated the pressure dependence of the kinetic isotope effect KIE of the photolysis
 160 frequencies of CH₂O and CHDO.

$$161 \quad KIE = \frac{j_{CH_2O}}{j_{CHDO}} \quad \text{with} \quad j = \int \Phi_{CH_2O/CHDO}^{tot} \sigma F d\lambda \quad (F6)$$

162 As the quantum yield of CH₂O is known from the literature (see e.g. Röth and Ehhalt, 2015)
 163 Φ_{CHDO}^{tot} remains the only unknown factor in formula F6. With the actinic flux density F of the
 164 lamp used by Nilsson et al. (2014) and the absorption spectra σ_x of CH₂O and CHDO from
 165 Gratien et al. (2007) the ratio KIE can be calculated with optimized values for *a* and *b* in eq.
 166 F5. Comparing the results of the simulation with the measured data by Nilsson et al. (2010,
 167 2014) the constants *a* and *b* can be determined via a least square fit. Figure 2 presents the
 168 result with optimized values *a*=2.94 and *b*=6.5×10⁻⁵ nm⁻¹ together with the measurements.
 169 The data at 1000 hPa is included in the fit as its mean value to accommodate the large
 170 variation of the data.

171

172 The total product quantum yield, deduced from the reaction scheme R0 to R7 is

$$173 \quad \Phi^{tot} = \frac{k_2}{\alpha} \cdot \Phi_2^{tot} + \frac{k_3[M]}{\alpha} \cdot \Phi_3^{tot} + \frac{k_4[M]}{\alpha} \cdot \frac{k_6}{\beta} \cdot \Phi_6^{tot} \quad (F7)$$

174 with α and β as defined in formula F3, and Φ_k^{tot} , the sub-product yield, according to formula
 175 F5. The measured wavelength dependence of Φ^{tot} at 1000 hPa pressure is depicted in Figure 3,
 176 where the total quantum yield is calculated with the rate constants from Table 2. The pressure
 177 dependence of the three terms of Φ^{tot} is illustrated in Figure 4.

178

179 To obtain a continuous and smooth wavelength dependence, the rate constants *k*₁ through *k*₇
 180 can be represented by an approximation function

$$181 \quad k = A \exp (B (\lambda - 300\text{nm})) \quad (F8)$$

182 The values for the parameters A and B are obtained from a least square fit to the data in Table
 183 2 and listed in Table 3. Where-ever the value of B was less than 0.001 it was set to 0, and A
 184 then corresponds directly to the mean of the respective rate constant. The wavelength
 185 dependence of Φ^{tot} at 1000 hPa with these functions is presented by the solid line in Figure 3.
 186 The comparison to the experimental data by Miller and Lee (1978) suggests a variance of the
 187 data of around 15%.

188

189 **Table 3:** Parameters of the rate constants according to equation F8, B in nm⁻¹ and A in s⁻¹,
 190 respective in cm³ s⁻¹, derived from least square fits.

	k₁	k₂	k₃	k₄	k₅	k₆	k₇

A	2.90×10^5	6.10×10^8	7.70×10^{-10}	1.30×10^{-10}	3.00×10^5	1.50×10^8	1.2×10^{-12}
B	0	0.086	0.069	0.071	0	0.075	0

191

192 For CHDO the only quantitative indication for the quantum yield of the radical channel in the
 193 literature are measurements of the kinetic isotope effect KIE (Feilberg et al., 2007, Rhee et al.,
 194 2008, Röckmann et al., 2010, and Nilsson et al., 2014). Following eq. F5, simulating these
 195 KIE-measurements requires three parameters for the individual radical quantum yield
 196 Φ_k^{rad} expressed in eq. F9: the maximum value Φ^{max} of the wavelength dependence, its
 197 curvature b , and the pivot wavelength λ_0 . The parameter a is set to 1, as for the radical
 198 quantum yield no pressure dependence is assumed, cancelling the $[M]/[M_0]$ factor.

$$199 \quad \Phi_k^{rad} = \frac{\Phi^{max}}{1 + a \exp\left(\frac{\epsilon_k - \epsilon_0}{b}\right)} \quad (F9)$$

200 Analog to the analysis for CH₂O (Röth and Ehhalt, 2015), where the curvatures of the
 201 wavelength dependence of Φ^{tot} and Φ^{rad} are similar, b can be set to $6.5 \times 10^{-5} \text{ nm}^{-1}$ for the
 202 radical quantum yield of CHDO. The maximum Φ^{max} was varied in the interval [0.70 / 0.78]
 203 around the corresponding value for CH₂O, but the resulting scattering is very small (see
 204 shaded area in Fig. 5). Consequently, parameter Φ^{max} is set to 0.74, matching the value also
 205 used for CH₂O (Ehhalt and Röth, 2015).

206

207 With these parameters the KIE of 1.63 as measured by Röckmann et al. (2010) was fitted with
 208 the actinic flux density given by Röckmann et al. and the optical spectra by Gratien et al.
 209 (2007). The best fit gave a pivot wavelength λ_0 of 327 nm. This value lies in the middle of the
 210 bond energies of 362.63 kJ/mol for C-H and 369.6 kJ/mol for C-D, calculated by Chuang et
 211 al. (1987). With the constants $\Phi^{max} = 0.74$, $a=1$, $b = 6.5 \cdot 10^{-5} \text{ nm}^{-1}$ and $1/\epsilon_0 = 327.1 \text{ nm}$ the
 212 quantum yield function Φ^{rad} of the radical channel of CHDO is analog to F7:

$$213 \quad \Phi^{rad} = \frac{k_2}{\alpha} \cdot \Phi_2^{rad} + \frac{k_3[M]}{\alpha} \cdot \Phi_3^{rad} + \frac{k_4[M]}{\alpha} \frac{k_6}{\beta} \cdot \Phi_6^{rad} \quad (F10)$$

214 where the radical quantum yields of the individual channels is given by function F9 and with
 215 α and β as defined in F3. Figure 5 depicts the wavelength dependence of the total quantum
 216 yield together with that for the radicals. At atmospheric pressures, as considered in this paper,
 217 the contributions of the individual quenching processes are insignificant with respect to the
 218 overall radical quantum yield.

219

220 To provide a more handy tool for atmospheric modeling, we introduce an exponential
 221 function (F11), with only one term and three parameters for the total and the radical quantum

222 yields of CHDO, similar to those deduced by Ehhalt and Röth (2015) for CH₂O, as a proxy
 223 for the three-term functions F7 and F10:

$$224 \quad \Phi = \frac{a}{1 + \exp\left(\frac{-\left(\frac{1}{\lambda} - \frac{1}{\lambda_0}\right)}{b}\right) \frac{[M]}{[M_0]}} \quad (\text{F11})$$

225 The corresponding parameters for the total quantum yield of CHDO are a=1.0, b=7.7×10⁻⁵ s⁻¹,
 226 and λ₀=336.2 nm. For the radical channel the factor [M]/[M₀] is set to 1, as the photolysis
 227 leading to the radicals is nearly pressure independent. The respective parameters are a=0.74,
 228 b=7.7×10⁻⁵ s⁻¹, and λ₀=325.0 nm. Both approximation curves are depicted in Figure 5, and
 229 Figure 6 shows the pressure dependent comparison with the measured data by Miller and Lee
 230 (1978).

231

232

233 4. Analysis of the CHDO photo-decomposition

234 Our second approach to estimate the quantum yields for the photolysis of CHDO is based on
 235 the experiments of McQuigg and Calvert (1969) who measured the photo-decomposition of
 236 CH₂O, CHDO, and CD₂O. Unfortunately, the authors only presented the quantum yields for
 237 the two radical reaction channels of CH₂O and CD₂O. They further assumed that the total
 238 quantum yield equals 1, independent of wavelength. It appears, however, that these data have
 239 a bias which becomes evident when the data for CH₂O are compared to more recent
 240 measurements.

241

242 In Figure 7 the dependence on the wavelength of Φ^{rad} of CH₂O by McQuigg and Calvert
 243 (1969) is depicted together with a curve for CH₂O, averaged over measured data from the
 244 paper by Röth and Ehhalt (2015). The latter evaluation showed no pressure dependence, but
 245 indicated a weak temperature effect which is neglected here. The curve is represented by the
 246 following function:

247

$$248 \quad \Phi_{CH_2O}^{rad} = \frac{0.74}{1 + \exp\left(\frac{-\left(\frac{1}{\lambda} - \frac{1}{327.4}\right)}{5.4 \times 10^{-5}}\right)} - \frac{0.40}{1 + \exp\left(\frac{\left(\frac{1}{\lambda} - \frac{1}{279.0}\right)}{5.2 \times 10^{-5}}\right)} \quad (\text{F12})$$

249 Equation F12 exhibits a maximum in Φ^{rad} around 310 nm, independent of the small
 250 temperature shift, whereas the earlier values of McQuigg and Calvert exhibit a monotonic
 251 decay with increasing wavelength above 280 nm, which points to a bias in the latter. The
 252 second summand in F12 is less than 1% at wavelengths above 300 nm and, hence, can be

253 omitted in the present paper. Figure 7 also includes the data of McQuigg and Calvert (1969)
 254 for CD₂O which show a quite similar wavelength dependency as the data for CH₂O.

255
 256 Our first assumption is that the bias in the experiments of McQuigg and Calvert extends
 257 equally to both isotopologues (CD₂O and CH₂O), and that, therefore, the ratio R of their
 258 quantum yields is correct. This ratio is displayed in Figure 8 and shows a mostly monotonic
 259 decrease with increasing wavelength. In this context, it is interesting to note that the ratio of
 260 the rate constants for the decomposition of excited CH₂O* and CD₂O* into the respective
 261 radical channels, as calculated by Troe (1984) from theory, result in a curve with a monotonic
 262 decrease with increasing wavelength similar to that of the quantum yield ratio (see Figure 8).

263 Using ratio R together with the fit function F12 for $\Phi_{CH_2O}^{rad}$ allows to estimate $\Phi_{CD_2O}^{rad}$ for the
 264 radical channel of CD₂O, as shown in Figure 8.

265 To calculate Φ_{CHDO}^{rad} we need one further assumption. Our hypothesis is suggested by the
 266 results of Feilberg et al. (2004), who found that the KIE of the reactions of CHDO with OH,
 267 Cl and Br are arithmetic means of the KIE of the reactions of CH₂O and CD₂O with those
 268 radicals. This in turn implies that the C-H bond strengths are similar in the isotopologues, and
 269 the same is true for the C-D bond strength. We, therefore, assume that Φ_{CHDO}^{rad} can be
 270 calculated from the average of $\Phi_{CH_2O}^{rad}$ and $\Phi_{CD_2O}^{rad}$ at each wavelength:

$$271 \quad \Phi_{CHDO}^{rad}(\lambda) = \left(\Phi_{CH_2O}^{rad}(\lambda) + \Phi_{CD_2O}^{rad}(\lambda) \right) / 2 \quad (F13)$$

272 The resulting radical quantum yields are compared in Figure 9. Φ_{CHDO}^{rad} does not depend on
 273 pressure since neither $\Phi_{CH_2O}^{rad}$ nor $\Phi_{CD_2O}^{rad}$ are pressure dependent. The respective maxima in
 274 Φ^{rad} , on the other hand, decrease from 0.72 over 0.70 to 0.65 for increasing deuteration.

275 Moreover, there is a blue shift of 5 nm, resp. 10 nm in the decreasing part of the quantum
 276 yield spectra of CHDO and CD₂O, i.e. at wavelengths above 315 nm. These blue shifts have
 277 the same tendency but do not quite match the measured threshold energies of 362.3 kJ/mol,
 278 368.4 kJ/mol, and 370.6 kJ/mol for CH₂O, CHDO, and CD₂O, respectively (Chuang et
 279 al.,1987), which correspond to the wavelengths 330.9 nm, 325.5 nm, and 323.5 nm.

280

281 The one-term fit function for the radical channel of CHDO is:

$$282 \quad \Phi_{CHDO}^{rad} = \frac{0.72}{1 + \exp\left(\frac{-\left(\frac{1}{\lambda} - \frac{1}{323.0}\right)}{7.7 \times 10^{-5}}\right)} \quad (F14)$$

283 and is also show in Figure 9. In Figure 10 the result of the interpretation of the measured
 284 photo-decomposition of CHDO by McQuigg and Calvert (1969) is compared to the radical

285 quantum yield deduced from the fluorescence measurements of Miller and Lee (1978). Both
 286 estimations lead to a wavelength dependence of Φ_{CHDO}^{rad} which lie in each others uncertainty
 287 range. This is a strong hint that the deduced results are robust and represent the true quantum
 288 yield of the radical channel of the photolysis of CHDO.

289
 290

291 **5. The isotope fractionation during the photolysis of formaldehyde**

292

293 The photolysis frequency J_i of the isotopologues CH_2O and $CHDO$ is given by the integration
 294 of quantum yield Φ , absorption cross section σ , and spectral actinic photon flux density $F_\lambda(\lambda)$
 295 over the λ wavelength domain:

$$296 \quad J_i = \int \Phi_{i,j}(\lambda) \cdot \sigma_i(\lambda) \cdot F_\lambda(\lambda) d\lambda \quad (F15)$$

297 where the quantum yield $\Phi_{i,j}(\lambda)$ depends on the product channel j , either molecular or radical,
 298 of isotopologue i , and the absorption cross section $\sigma_i(\lambda)$ is specific to the isotopologues i . For
 299 our calculations the absorption spectra of CH_2O and $CHDO$ from Gratien et al. (2007) were
 300 applied. We used these values instead of the JPL-recommendation (Burkholder, 2020) for
 301 consistency with the calculations in section 2 and 3. The solar spectral actinic flux density F_λ
 302 was calculated from a quasi-spherical 1-D radiation transfer model (Röth, 2002); the $\Phi(\lambda)$ are
 303 those from section 2. An example of the terms $\Phi^{mol}(\lambda)$, $\sigma(\lambda)$, $F_\lambda(\lambda)$ for the molecular channel
 304 of $CHDO$ is given in Figure 11 for the pressure and temperature at an altitude of 20 km. The
 305 product of these terms, integrated over 5 nm intervals for better visibility, is also displayed to
 306 demonstrate the spectrally resolved contributions to the photolysis frequency of the molecular
 307 channel of $CHDO$.

308

309 The kinetic isotope effect for the molecular channel is given by

$$310 \quad KIE_{mol} = \frac{J_{CH_2O}^{mol}}{J_{CHDO}^{mol}} \quad (F16)$$

311 and correspondingly for the radical channel

$$312 \quad KIE_{rad} = \frac{J_{CH_2O}^{rad}}{J_{CHDO}^{rad}} \quad (F17)$$

313 For a quick overview the dependence of KIE_{rad} and KIE_{mol} on altitude for globally averaged
 314 conditions (equinox, 30°N) are depicted in Figures 12a and 12b. KIE_{mol} decreases
 315 monotonically with decreasing pressure from 1.59 at 1000 hPa to 1.06 at 1 hPa. The radical
 316 channel in contrast shows hardly any pressure dependency as the rate of this reaction is not
 317 influenced by the quenching process. The marginal variation of the kinetic isotope effect with

318 altitude is caused by the altitudinal increase of the photon flux and its differing contribution to
319 the photolysis frequency integrals of CH₂O and CHDO.

320

321 To examine whether the quantum yield functions for CHDO deduced above are applicable for
322 modeling purposes, additional sensitivity studies were carried out, varying the main features
323 of the quantum yield functions. With respect to the fractionation factor, only the variations of
324 those parameters are relevant which alter the relation of the entire photolysis frequency
325 integrals (eq. F15) of the molecular and the radical channels. In Figures 12a and 12b we
326 additionally show the variances of the photolysis frequencies as well as of the fractionation
327 factors. The shaded area is produced by varying one parameter of the CHDO quantum yield
328 as indicated below. The photolysis frequency of CH₂O remained unchanged.

329

330 The sensitivity of the molecular branch of the photolysis frequency of CHDO to the
331 preexponential factor of the quantum yield function is roughly 10 % throughout the
332 atmosphere if this value is varied by 10%. All other parameters do not alter the integral
333 equation F15 significantly and produce only variances less than 1 % when changed by 10%. It
334 can thus be concluded that the estimated equation parameters are good representations of the
335 actual values. At higher altitudes (<10 hPa) Φ_{CHDO}^{mol} and $\Phi_{CH_2O}^{mol}$ are close to unity in the
336 wavelength regime 330 nm to 360 nm (see e.g. Fig. 6). So, the photolysis frequency in the
337 stratosphere does not change much if the parameters of the respective functions are varied.
338 Therefore, the variance of the fractionation factor does not much decrease above 30 km
339 altitude. Here, measurements at tropospheric pressures could be much more informative as
340 becomes evident from Figure 12.

341

342 The photolysis frequency of the radical channel of CHDO is only sensitive to the maximum of
343 the quantum yield and to the threshold wavelength 323 nm. Shifting the latter value by ± 3 nm
344 produces changes of about 20 % in the troposphere, decreasing to 10 % at 50 km altitude as
345 shown in Figure 12. This variation of the threshold produces an error bar of the fractionation
346 factor of the same magnitude.

347

348

349 **6. Discussion**

350

351 Due to consecutive reactions only the molecular channel contributes to the HD production.
352 Up to now there had been a handicap in the interpretation of stratospheric measurements of
353 the concentration of deuterated hydrogen HD due to the lack of exact knowledge of the
354 photolysis frequencies of deuterated formaldehyde, resulting in an uncertainty on the
355 fractionation factor. There have been a number of experimental approaches to deduce the
356 fractionation factor, where e.g. Feilberg et al. (2005) measured a value of 1.82 ± 0.07 for α_{mol} ,
357 while Röckmann et al. (2010) found a value of 1.63 ± 0.03 for that ratio. In their modeling
358 paper, Mar et al. (2007) varied the fractionation factor between 1.2 and 1.5 for stratospheric
359 conditions.

360

361 In all these studies the pressure dependence of the photolysis frequencies could not be
362 investigated. An interesting experiment by Nilsson et al. (2009) addressed this problem.
363 Unfortunately, the spectral radiance of the light source used did not resemble the sun light
364 well enough, and their findings could not be transferred to the real atmosphere without
365 information on the quantum yield of CHDO.

366

367 Beside its pressure dependence the variation of the photolytic fractionation factors can also be
368 caused by different actinic fluxes at the times and sites of the experiments. The actinic flux in
369 the numerator and denominator of the fractionation factor in equations F16 and F17 do not
370 cancel out, and, therefore, the factor is depending on the local insolation conditions.

371 Calculations of the solar zenith angle (SZA) dependency with the complex radiation transfer
372 model ART (Röth, 2002) result in values from 1.47 at overhead sun to 1.95 at $\text{SZA}=83^\circ$ for
373 clear sky and free horizon at ground level. This zenith angle dependency is less expressed at
374 20 km altitude and disappears at 50 km, as depicted in Figure 13. This effect may explain the
375 differences in the measurements of the fractionation factors. To check the variance with the
376 solar zenith angle the measured fractionation factor KIE_m (eq. F16) is compared to model
377 calculations. The factor 1.63 ± 0.03 (Röckmann et al., 2010) was derived from experimental
378 studies in the atmospheric simulation chamber SAPHIR between 60° and 70° SZA
379 (Röckmann et al., 2010). The absorption cross sections by Gratien et al. (2007) and the
380 quantum yields derived above together with the radiation spectra result in a fractionation
381 factor of 1.54 for 60° SZA and 1.70 for 70° SZA are in good agreement with the measured
382 value.

383

384 **Conclusions**

385
386 The current work derives a framework and set of equations for describing the CHDO
387 photolysis, based on two different approaches building on the available literature data, finding
388 a consistent result across all data sets. It could be shown that the most influential parameters
389 of the rates of photolysis of CHDO are the absolute value and the threshold of the quantum
390 yield of the radical channel. Simplified parametrized equations (F11 and F14) that are readily
391 implemented in kinetic models are provided for these quantities. Measurements around 300
392 nm and 325 nm could help to further reduce the uncertainty on the fractionation factor.
393 Additional measurements of the pressure dependence of the total quantum yield, i.e. the
394 quenching rate of excited CHDO^{*}, would be valuable to further test the assumptions made in
395 this paper.

396

397 **Competing interests**

398 The authors declare they have no competing interests

399

400

401

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404 suggestions to improve the clarity and readability of the paper.

405

406 **References**

- 407
408 Araùjo, M., Lasome, B., Magalhaes, A. L., Worth, G. A., Bearperk, M. J., and Robb, M. A.
409 *The molecular dissociation of formaldehyde at medium photoexcitation energies: A quantum*
410 *chemistry and direct quantum dynamics study*
411 J. Chem. Phys. 131, 144301-1 – 144301-8, 2009
412
413 Breuer, G. M., and Lee, E. K. C.
414 *Fluorescence decay times of Cyclic Ketones, Acetone, and Butanal in the gas phase*
415 J. Phys. Chem. 75, 989 – 990, 1971
416
417 Burkholder, J. B., Sander, S. P., Abbatt, J. P. D., Barker, J. R., Cappa, C., Crous, D., Dipple,
418 T. S., Huie, R. E., Kolb, C. E., Kurylo, M. J., Orkin, V. L., Percival, C. J., Wilmouth, D. M.,
419 and Wine, P. H.
420 *Chemical kinetics and photochemical data for use in atmospheric studies*
421 JPL-Publication 19-5, Pasadena, 2020
422
423 Chuang, M-C., Foltz, M. F., and Moore, C. B.
424 *T₁ barrier height, S₁-T₁ intersystem crossing rate, and S₀ radical dissociation threshold for*
425 *H₂CO, D₂CO, and HDCO*
426 J. Chem. Phys. 87, 3855 – 3864. 1987
427
428 Ehhalt, D. H., Volz, A.
429 *Coupling of the CH₄ with the H₂ and CO cycle : isotopic evidence*
430 In : Symposium on Microbiol. Production and Utilization of Gases (H₂, CH₄, and CO)
431 (eds.: Schlegel, H. G., Gottschalk, G., Pfenning, N.), Akad. Wiss. Göttingen, Germany, 1976
432
433 Feilberg, K. L., Johnson, M. S., Nielsen, C. J.
434 *Relative Reaction Rates of HCHO, HCDO, DCDO, H¹³CHO, and HCH¹⁸O with OH, Cl, Br,*
435 *and NO₃ Radicals.*
436 J. Phys. Chem. A 108, 7393 – 7398, 2004
437
438 Feilberg, K. L., D'Anna, B., Johnson, M. S., Nielsen, C. J.
439 *Relative Tropospheric Photolysis Rates of HCHO, H¹³CHO, HCH¹⁸O, and DCDO Measured*
440 *at the European Photoreactor Facility*
441 J. Phys. Chem. 109A, 8314 – 8319, 2005
442
443 Feilberg, K. L., Johnson, M. S., Bacak, A., Röckmann, T., Nielsen, C. J.
444 *Relative Tropospheric Photolysis Rates of HCHO and HCDO Measured at the European*
445 *Photoreactor Facility*
446 J. Phys. Chem. 111A, 9034 – 9046, 2007
447
448 Fu, B., B. Shepler, B. C., and Bowman, J. M.
449 *Three-State Trajectory Surface Hopping Studies of the photodissociation dynamics of*
450 *formaldehyde on ab initio potential energy surfaces*
451 J. Am. Chem. Soc. 133, 7957 – 7968, 2011
452
453 Gerst, S., Quay, P.
454 *Deuterium component of the global molecular hydrogen cycle*
455 J. Geophys. Res., 106, 5021 – 5031, 2001
456

457 Gratien, A., Nilsson, E., Doussin, J-F., Johnson, M. S., Nielsen, C. J., Stenstrom, Y., and
458 Picquet-Varrault, B.
459 *UV and IR Absorption Cross-sections of HCHO, HCDO, and DCDO*
460 *J. Phys. Chem.*, 111A, 11506 – 11513, 2007
461
462 Hirschfelder, J. O., Curtiss, C. F., and Bird, R. B.
463 *Molecular Theory of Gases and Liquids*
464 Wiley, London, 1249pp., 1954
465
466 Keller-Rudek, H., G. Moortgat, G. K., Sander, R., and Sörensen, R.
467 *MPI-Mainz UV-VIS Spectral Atlas of Gaseous Molecules*
468 <http://www.atmosphere.mpg.de/enid/2295>
469 Last access 01.06.2022
470
471 Mar, K. A., McCarthy, M. C., Connell, P., K. Boering, K. A.
472 *Modelling the photochemical origins of the extreme deuterium enrichment in stratospheric H₂*
473 *J. Geophys. Res.* 112, D19302, doi: 10.1029/2006JD007403, 2007
474
475 McCarthy, M. C., K. Boering, A., Rahn, T., J. Eiler, J. M., Rice, A. L., Tyler, S. C.,
476 Schaufli, S., Atlas, E., D. Johnson, D. G.
477 *The hydrogen isotopic composition of water vapor entering the stratosphere inferred from*
478 *high-precision measurements of $\delta D-CH_4$ and $\delta D-H_2$*
479 *J. Geophys. Res.* 109, D07304, doi:10.1029/2003JD004003, 2004
480
481 McQuigg, R. D., Calvert, J. G.
482 *The Photodecomposition of CH₂O, CD₂O, CHDO, and CH₂O-CD₂O Mixtures at Xenon Flash*
483 *Lamp Intensities*
484 *J. Am. Chem. Soc.* 91, 1590 – 1599, 1969
485
486 Miller, R. G. and Lee, E. K. C.
487 *Single vibronic level photochemistry of formaldehydes in the A ¹A₂ state: Radiative and*
488 *nonradiative processes in H₂CO, HDCO, and D₂CO*
489 *J. Chem. Phys.* 68, 4448 – 4464, 1978
490
491 Nelder, J. A. , and Mead, R.
492 *A simplex method for function minimization*
493 *Computer Journal*, 7, 308 -313, 1965
494
495 Nilsson, E. J. K., Bache-Andreassen, L., Johnston, M. S., and Nielsen, C. J.
496 *Relative Tropospheric Photolysis Rates of Acetaldehyde and Formaldehyde Isotopologues*
497 *Measured at the European Photoreactor Facility*
498 *J. Phys. Chem.* 113A, 3498 – 3504, 2009
499
500 Nilsson, E. J. K., Andersen, V. F., Skov, H., and Johnson, M. S.
501 *Pressure dependence of the deuterium isotope effect in the photolysis of formaldehyde by*
502 *ultraviolet light*
503 *Atmos. Chem. Phys.* 10, 3455 – 3463, 2010
504
505 Nilsson, E. J. K., Schmidt, J. A., and Johnson, M. S.
506 *Pressure dependent isotopic fractionation in the photolysis of formaldehyde-d₂*
507 *Atmos. Chem. Phys.* 14, 551 – 558. 2014

508
509 Osborn, D.L.
510 *Exploring multiple reaction paths to a single product channel*
511 *Adv. Chem. Phys.* 138, 213 – 265, 2008
512
513 Pieterse, G., Krol, M. C., Batenburg, A. M., Steele, L. P., Krummel, P. B., Langenfels, R. I.,
514 and Röckmann, T.
515 *Global modelling of H₂ mixing ratios and isotopic compositions with the TM5 model*
516 *Atmos. Chem: Phys.*, 11, 7001 – 7026, 2011
517
518 Rahn, T., Eiler, J. M., Boering, K. A., Wennberg, P. O., McCarthy, M. C., Tyler, S.,
519 Schaufli, S., Donnelly, S., Arlas, E.
520 *Extreme deuterium enrichment in stratospheric hydrogen and the global atmospheric budget*
521 *of H₂*
522 *Nature* 424, 918 – 921, 2003
523
524 Rhee, T. S., Brenninkmeijer, C. A. M., Braß, M., Brühl, C.
525 *Isotopic composition of H₂ from CH₄ oxidation in the stratosphere and the troposphere*
526 *J. Geophys. Res.* 111, D23303, doi:10.1029/2005JD00670, 2006
527
528 Rhee, T. S., Brenninkmeijer, C. A. M., and Röckmann, T.
529 *Hydrogen isotope fractionation in the photolysis of formaldehyde*
530 *Atmos. Chem. Phys.* 8, 1353 – 1366, 2008
531
532 Röckmann, T., Rhee, T. S., Engel, A.
533 *Heavy hydrogen in the stratosphere*
534 *Atmos. Chem. Phys.* 3, 2015 -2023, 2003
535
536 Röckmann, T., Walter, S., Bohn, B., Wegener, R., Spahn, H., Brauers, T., Tillmann, R.,
537 Schlosser, E., Koppmann, R., and Rohrer, F.
538 *Isotope effect in the formation of H₂ from H₂CO studied at the atmospheric Simulation*
539 *chamber SAPHIR*
540 *Atmos. Chem. Phys.* 10, 5343 – 5357, 2010
541
542 Röth, E.-P.
543 *Description of the Anisotropic Radiation Transfer Model ART to Determine*
544 *Photodissociation Coefficients*
545 *Ber. Forschungszentrum Jülich, Jül-3960*, 2002
546
547 Röth, E.-P. and Ehhalt, D. H.
548 *A simple formulation of the CH₂O photolysis quantum yields*
549 *Atmos. Chem. Phys.* 15, 7195 – 7202, 2015
550
551 Troe, J.
552 *Specific rate constants k(E,J) for the unimolecular dissociations of H₂CO and D₂CO*
553 *J. Phys. Chem.*, 88, 4375 – 4380, 1984
554
555 Troe, J.
556 *Analysis of quantum yields for the photolysis of formaldehyde at > 310 nm*
557 *J. Phys. Chem. A* 111, 3868 – 3874, 2007
558

559 Yamaguchi, Y., Wesolowski, S., Van Huis, T. J., and Schaefer III, H. F.
560 *The unimolecular dissociation of H₂O on the lowest triplet potential-energy surface*
561 *J. Chem. Phys.* 108, 5281 – 5288, 1998
562
563 Yeung, E.S. and Moore, C. B.
564 *Photochemistry of single vibronic levels of formaldehyde*
565 *J. Chem. Phys.* 58, 3988 – 3998, 1973
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569 **Figures**

570

571 **Fig01** : Comparison of the fluorescence quantum yield measured by Miller and Lee (1978)
572 (full dots) with the fitted function $\Phi_F(M)$ (see Eq. F3) for different wavelengths in nm as
573 indicated. To emphasize the quality of the fit, we depict only the pressure dependent part
574 $\theta(M)$ as defined in Eq. F4

575

576 **Fig02** : The pressure dependence of the kinetic isotope effect KIE (i.e. the ratio of the
577 $\text{CH}_2\text{O}/\text{CHDO}$ photolysis frequencies, see Eq. F6) is compared to the measured data of Nilsson
578 et al., 2010 (blue squares), and to Feilberg et al., 2007, Rhee et al., 2008, and Röckmann et al.,
579 2010 (red squares, 'others'). The solid curve at 1000 hPa is included to accommodate the
580 variation of the data.

581

582 **Fig03**: The total product quantum yields Φ^{tot} derived from the measured rate constants of
583 Miller and Lee (1978) at 1000 hPa through Eq. F7 (full circles) is well reproduced by the
584 continuous Φ^{tot} function obtained after fitting the rate coefficients to function Eq. F8 (solid
585 curve).

586 **Fig04**: Wavelength dependence of the contributions of the three terms for Φ_2^{tot} , Φ_3^{tot} and
587 Φ_6^{tot} of equation F7 to the total quantum yield Φ^{tot} of the CHDO photolysis at 10 hPa (a) and
588 1030 hPa (b).

589

590 **Fig05**: The total quantum yields Φ^{tot} of the photolysis of CHDO and that of the radical
591 channel, Φ^{rad} , calculated with the three-term functions Eq. F7 and F10 (black curves). The
592 blue shaded area indicates the variation of parameter a within the interval $[0.70 / 0.78]$. The
593 red curves, derived using the one-term approximation (eq. F11), and the black curves fall
594 within the variance of each other.

595

596 **Fig06**: Comparison of the one-termfit function F11 (open circles on the solid line) with the
597 measured data (Miller and Lee, 1978) of the total photolytic quantum yields Φ^{tot} (full circles)
598 at 1, 10, 200, and 1000 hPa.

599

600 **Fig07**: The original data of McQuigg and Calvert (1969) for CH_2O (full red squares) and
601 CD_2O (open squares) for the photolytic quantum yields of the radical channel are compared to
602 the averaged function for CH_2O by Röth and Ehhalt (2015).

603

604 **Fig08:** The ratio $\Phi_{CD_2O}/\Phi_{CH_2O}$ of the McQuigg and Calvert (1969) data shown in Fig.7 and
605 the corrected radical quantum yield of CD_2O , $\Phi_{CD_2O}^{rad}$ (black squares). The ratio of the
606 respective reaction constants (triangles) derived theoretically by Troe (1984) shows the same
607 tendency as the quantum yield ratio.

608

609 **Fig09:** Wavelength dependency of the quantum yields Φ^{rad} for the radical channel of the three
610 isotopologues of formaldehyde. The curves for CH_2O (Röth and Ehhalt, 2015) and that for
611 CD_2O (corrected data of McQuigg and Calvert, 1969) are used to calculate the quantum yield
612 of $CHDO$ (black dots) by their mean values. Φ_{CHDO}^{rad} is then fitted by the one-term function Eq.
613 F14 (black line).

614

615 **Fig10:** The $CHDO$ quantum yield Φ_{CHDO}^{rad} fit function deduced from the fluorescence
616 measurements (blue line, Eq. F11) of Miller and Lee (1978), and that from the interpretation
617 of the photo-decomposition (red line, Eq. F14) measurements of McQuigg and Calvert
618 (1969). These lie within the uncertainty range of each other. Also depicted is the function for
619 the total quantum yield Φ^{tot} (Eq. F11)

620

621 **Fig11 :** The photolysis rate is the combination of the actinic photon flux, the absorption cross
622 section and the quantum yield. Depicted are the contributions to the molecular channel of the
623 photolysis rate of $CHDO$, J_{CHDO}^{mol} , at 20 km altitude and integrated over 5 nm wavelength: the
624 actinic photon flux $F_\lambda(\lambda)$, the absorption cross section $\sigma(\lambda)$ (Gratien et al., 2007), and the
625 quantum yield for the molecular channel $\Phi^{mol}(\lambda)$. The photolysis rate, the photon flux and the
626 cross section are multiplied by 2.5×10^5 sec, 2.5×10^{-15} photons⁻¹ nm sec, and 1.5×10^{19} cm⁻¹,
627 respectively, to achieve comparability.

628

629 **Fig12 :** The altitudinal dependence of the photolysis frequencies J_{mol} of the molecular
630 channels (a) of CH_2O and $CHDO$ is important for the atmospheric production of HD. For
631 comparison the radical channels (b) are also depicted. The dependence on altitude of the
632 kinetic isotope effect, $KIE = J_{CH_2O}/J_{CHDO}$, is more pronounced for the molecular channel than
633 the radical channel. The shaded area indicates the variance upon changing (a) the quantum
634 yield Φ^{mol} by 10% and (b) the radical threshold wavelength by 3 nm.

635

636 **Fig13** : At 50 km altitude the solar zenith angle dependency of the photolysis frequency ratio
637 of the molecular channel is nearly constant. In contrast, at 0 km the ratio increases
638 significantly for solar zenith angles above 30 degrees.
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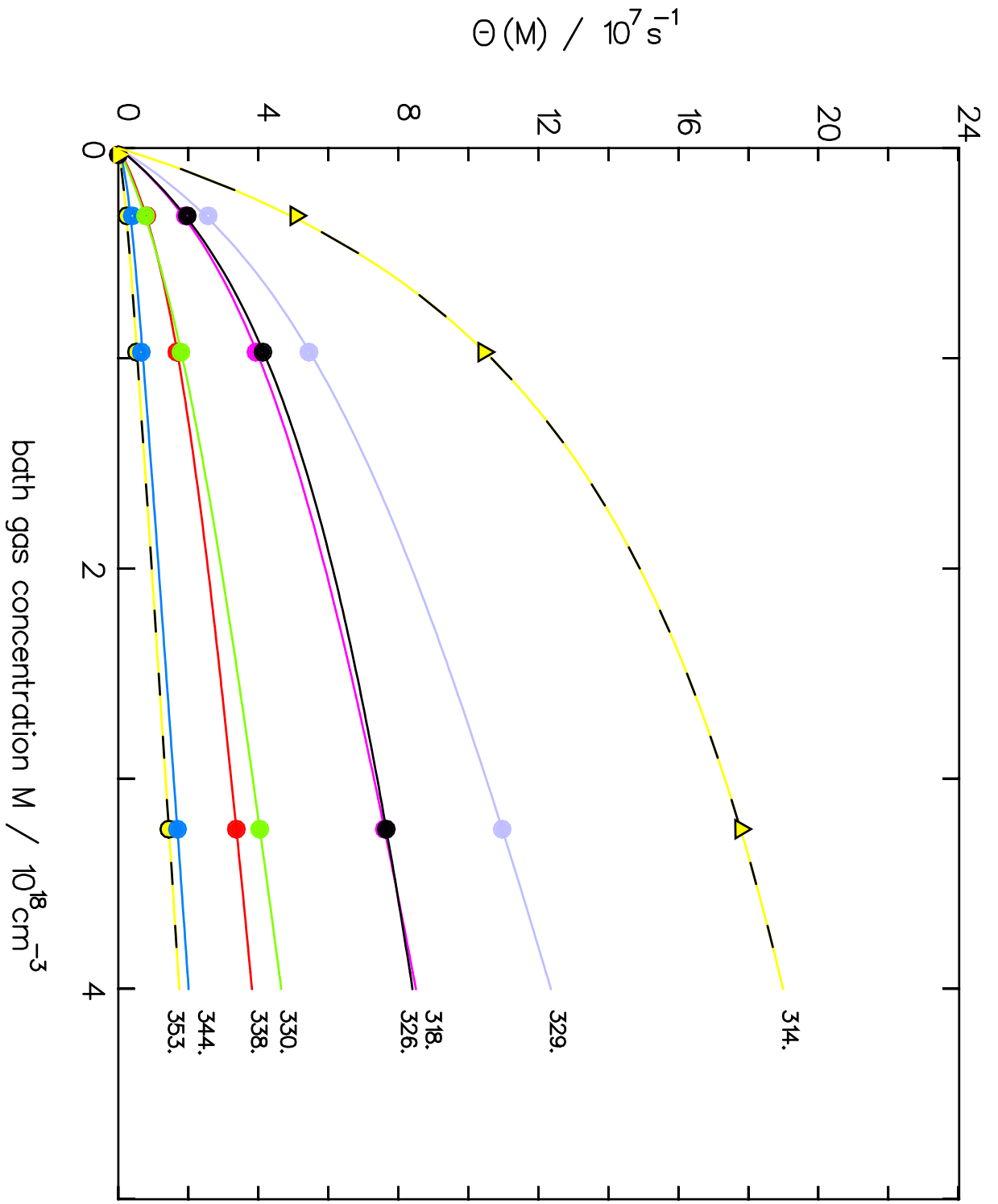


Fig. 1

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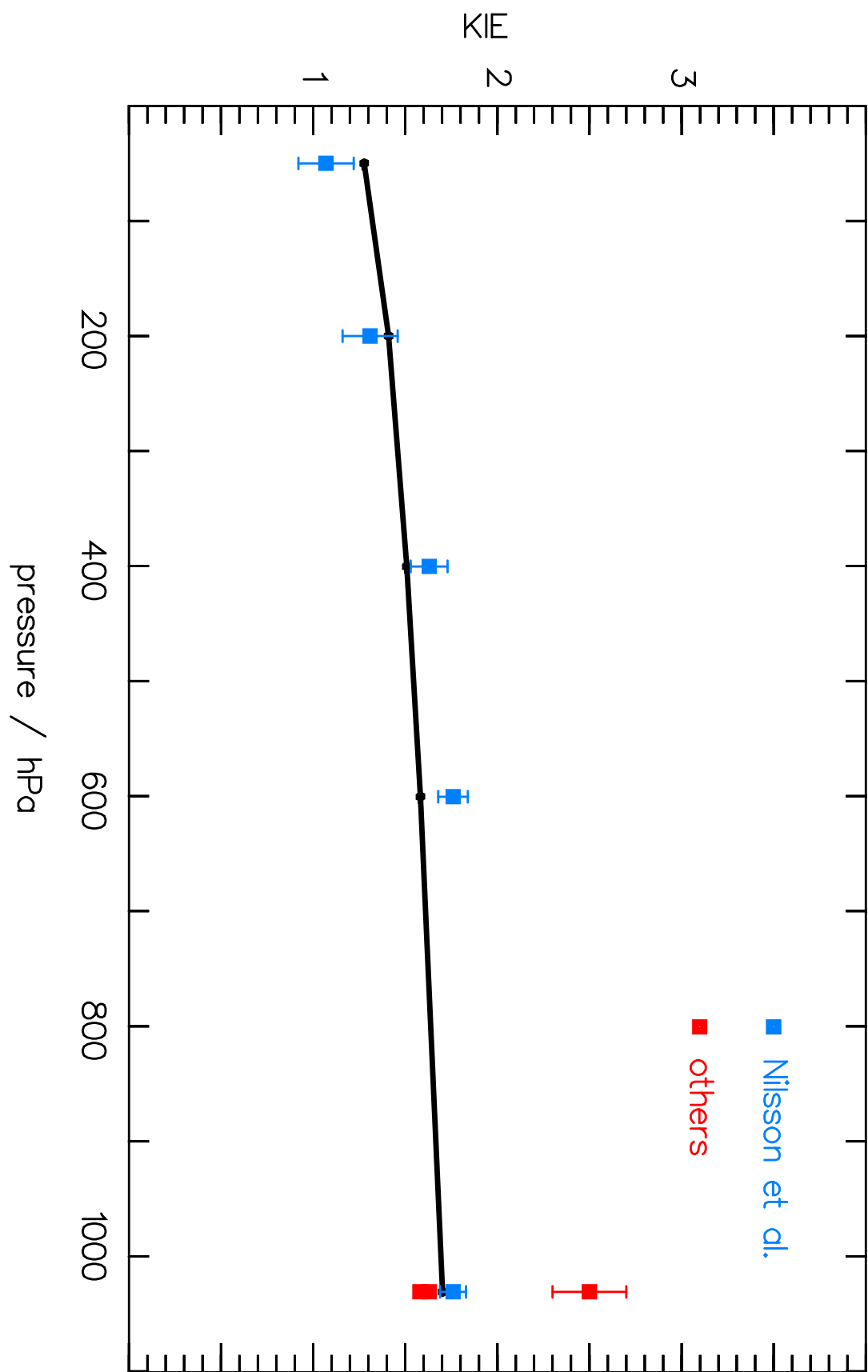


Fig. 2

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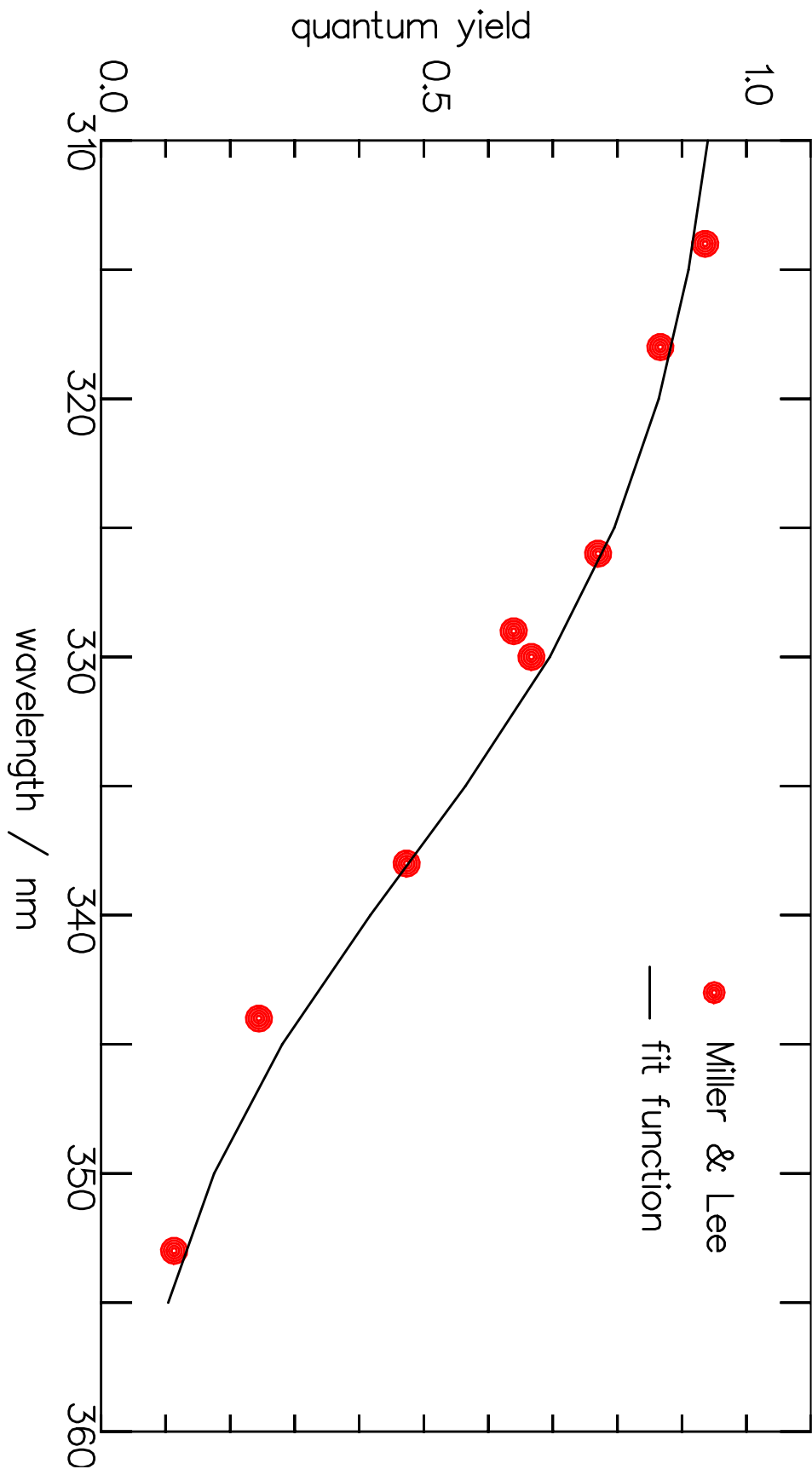


Fig. 3

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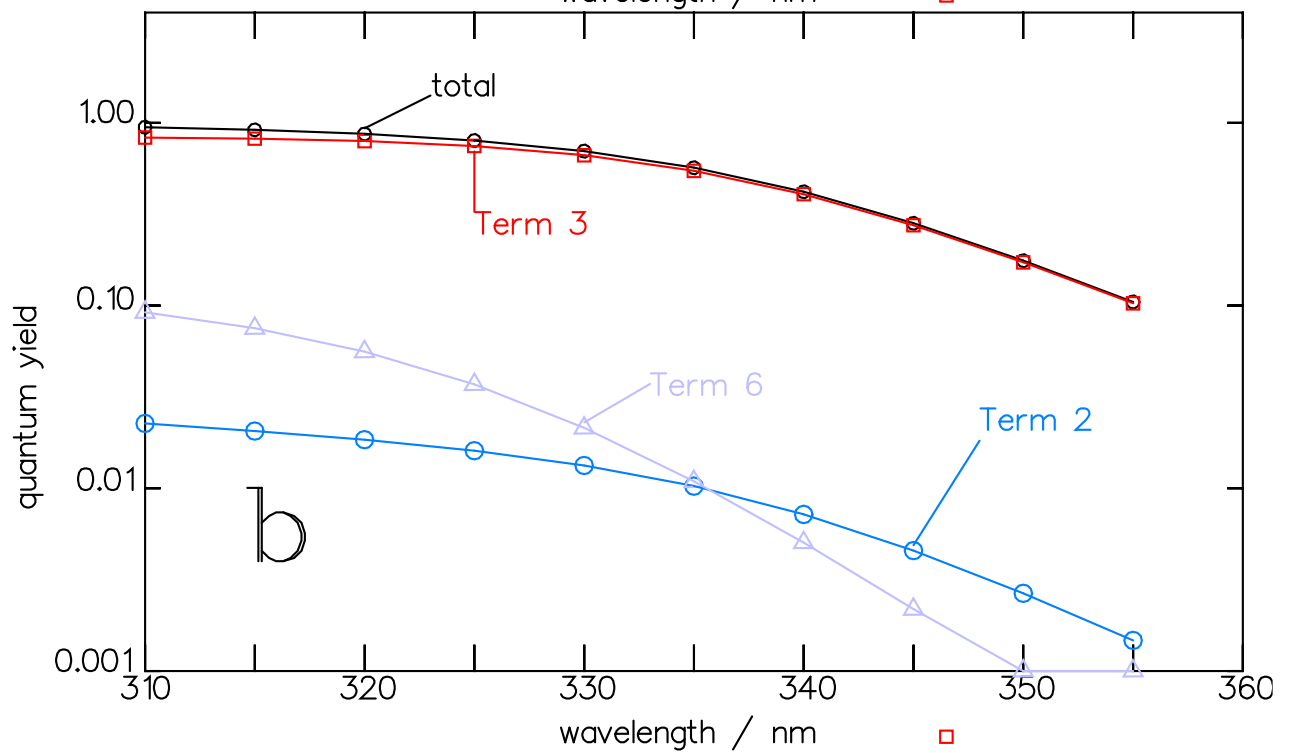
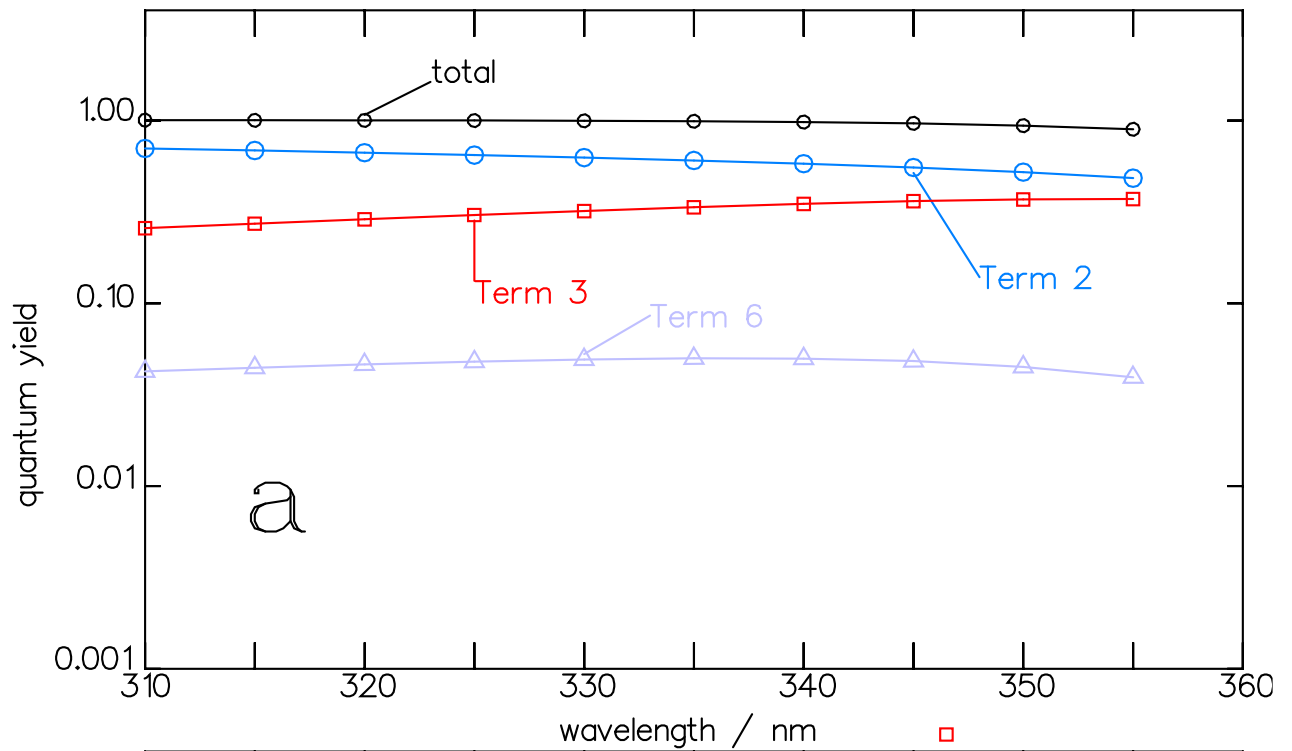


Fig. 4

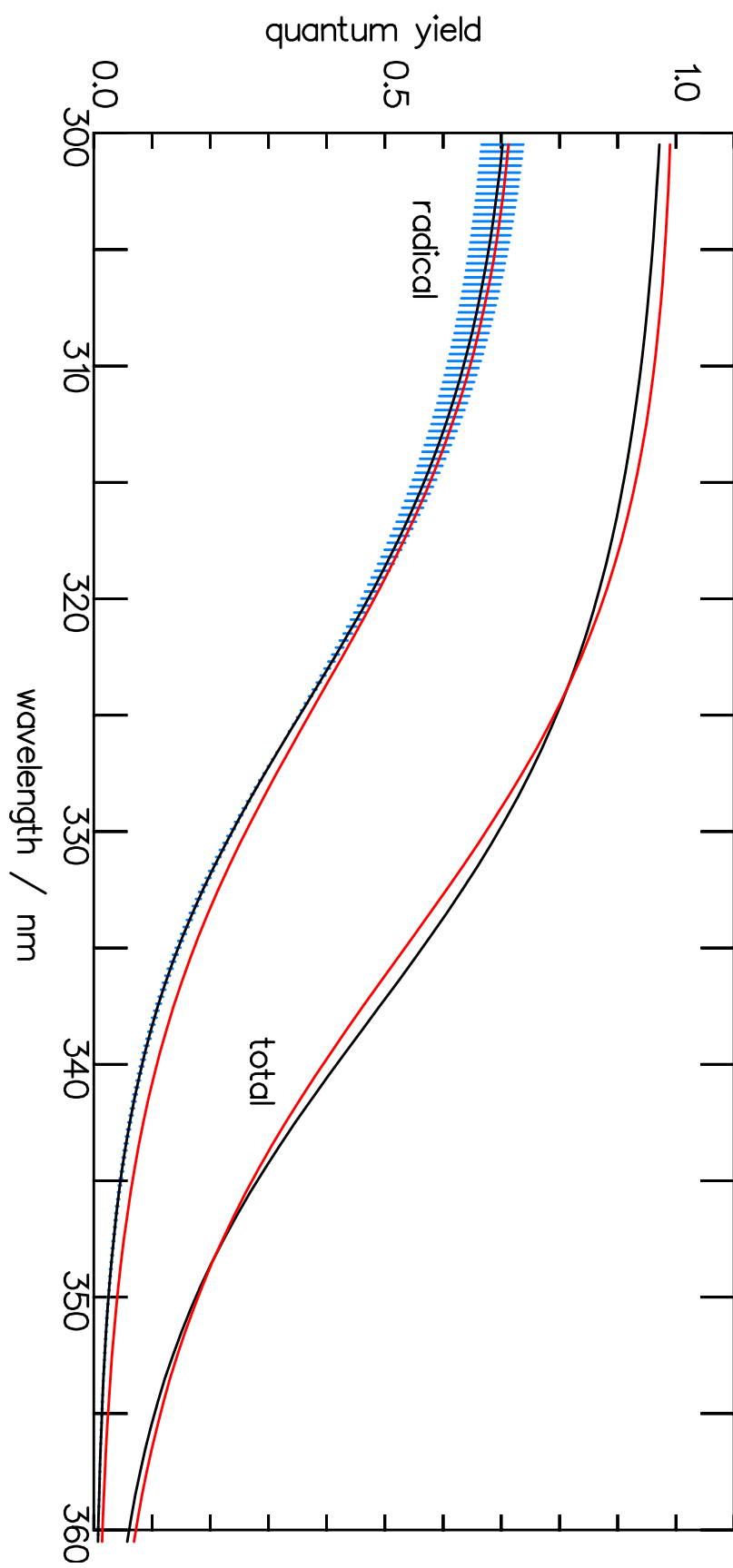


Fig. 5

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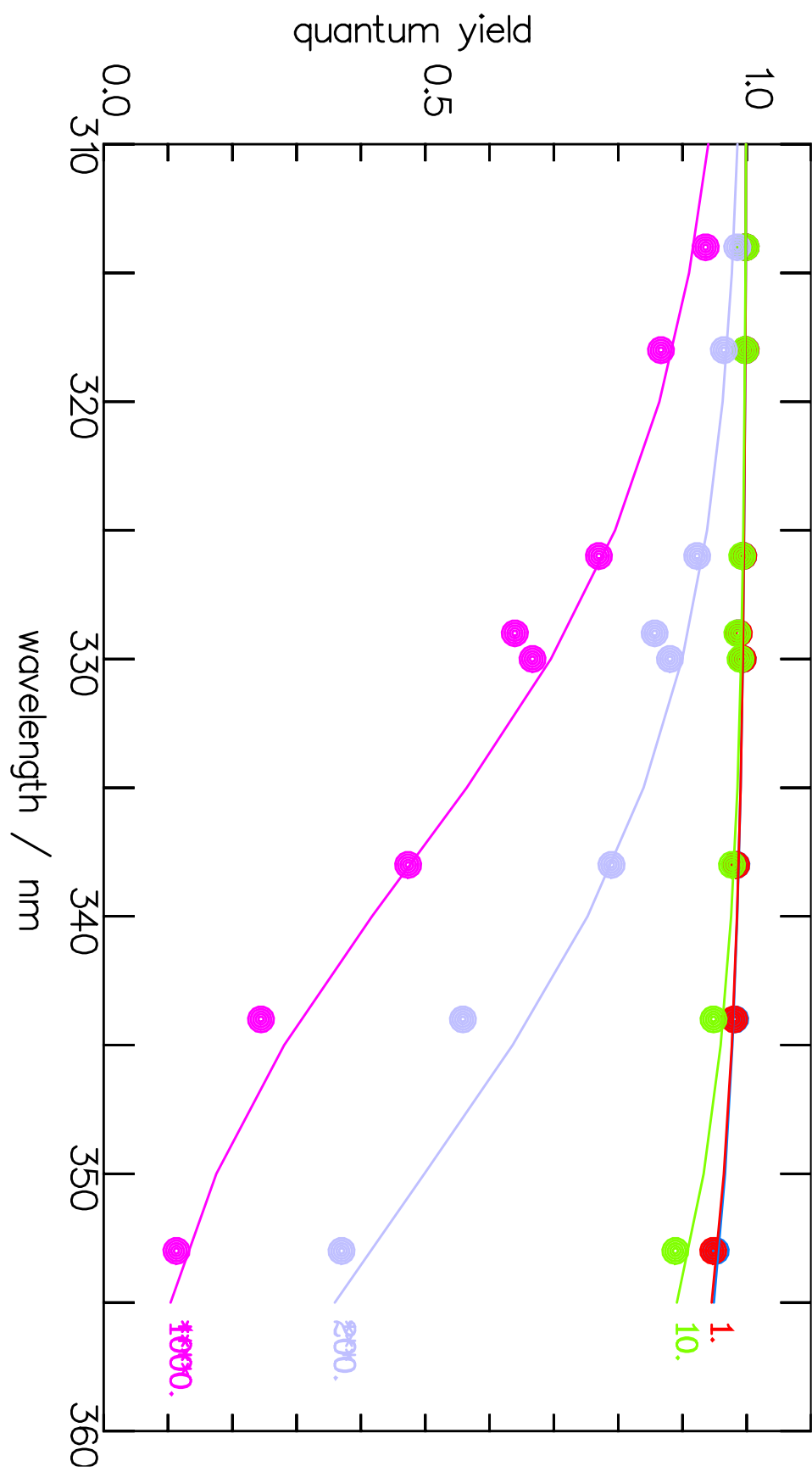


Fig. 6

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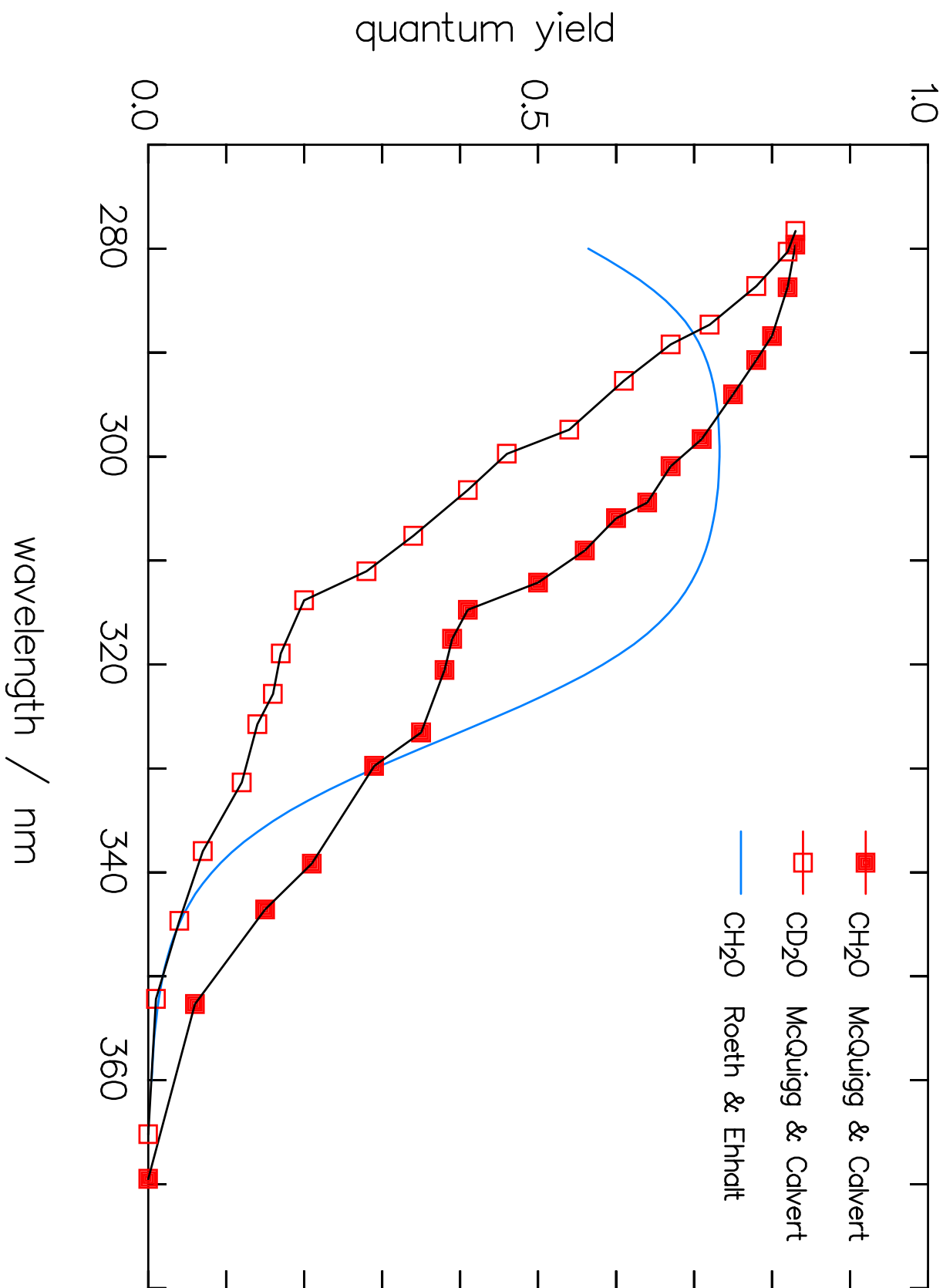


Fig 7

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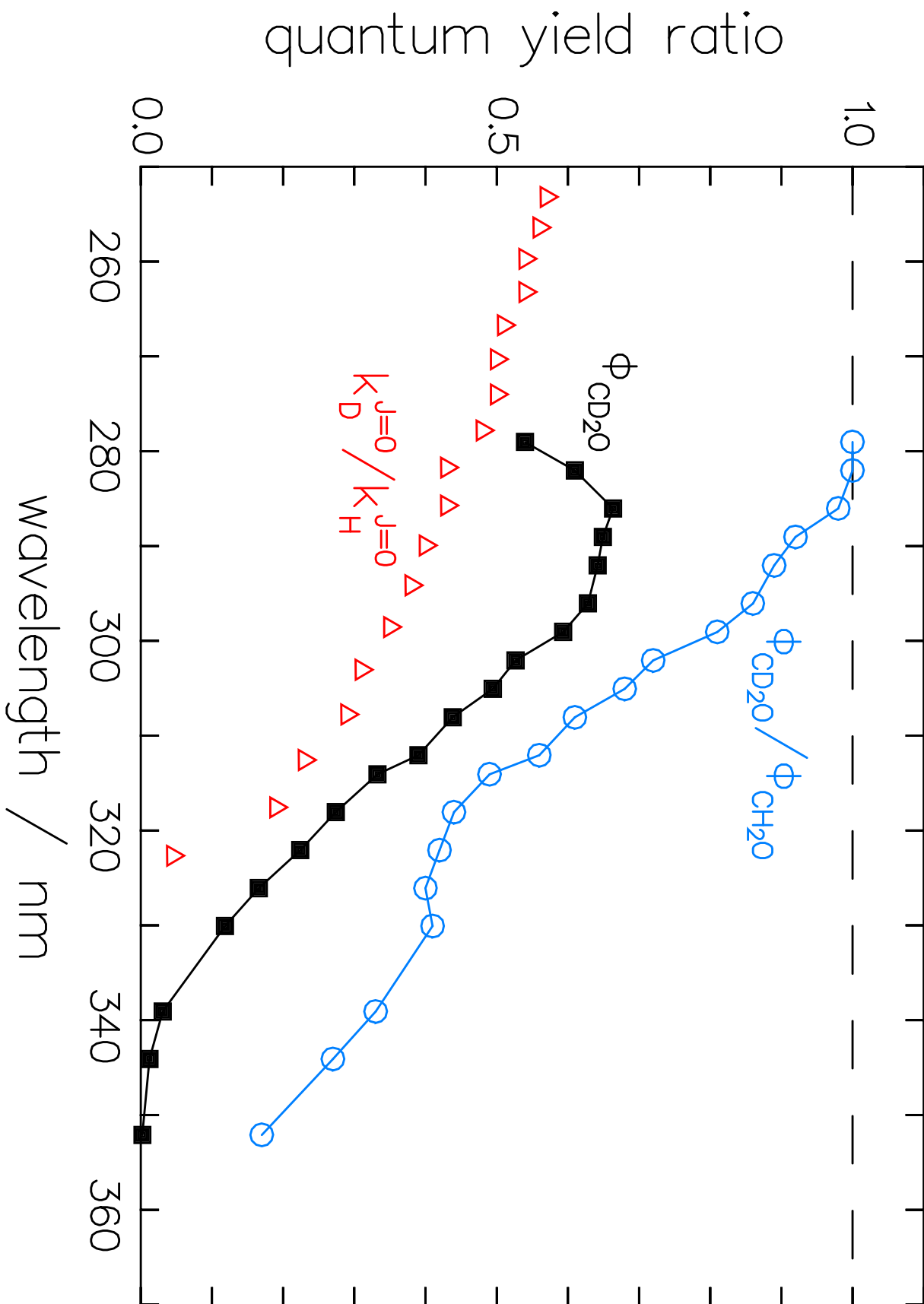


Fig. 8

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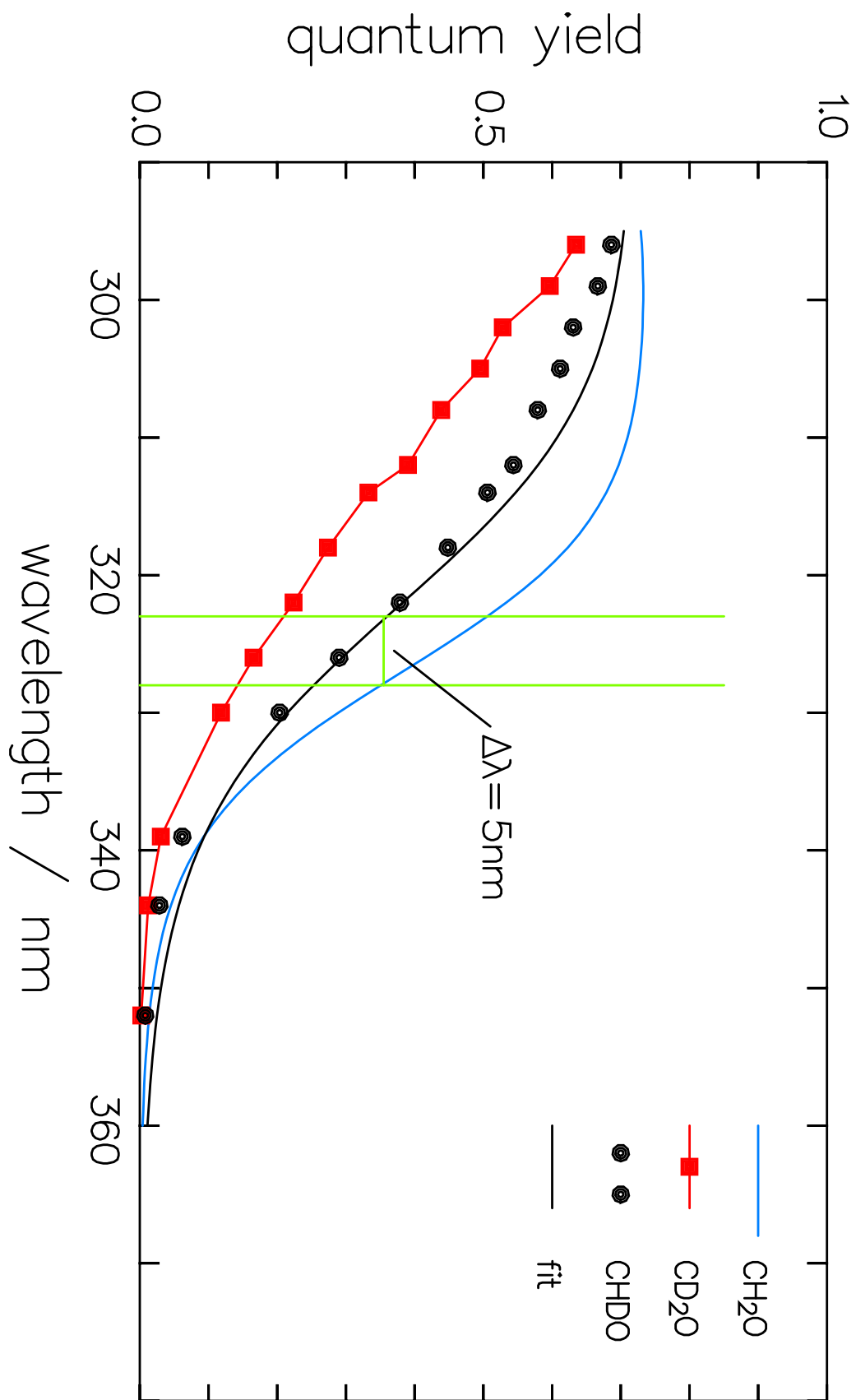
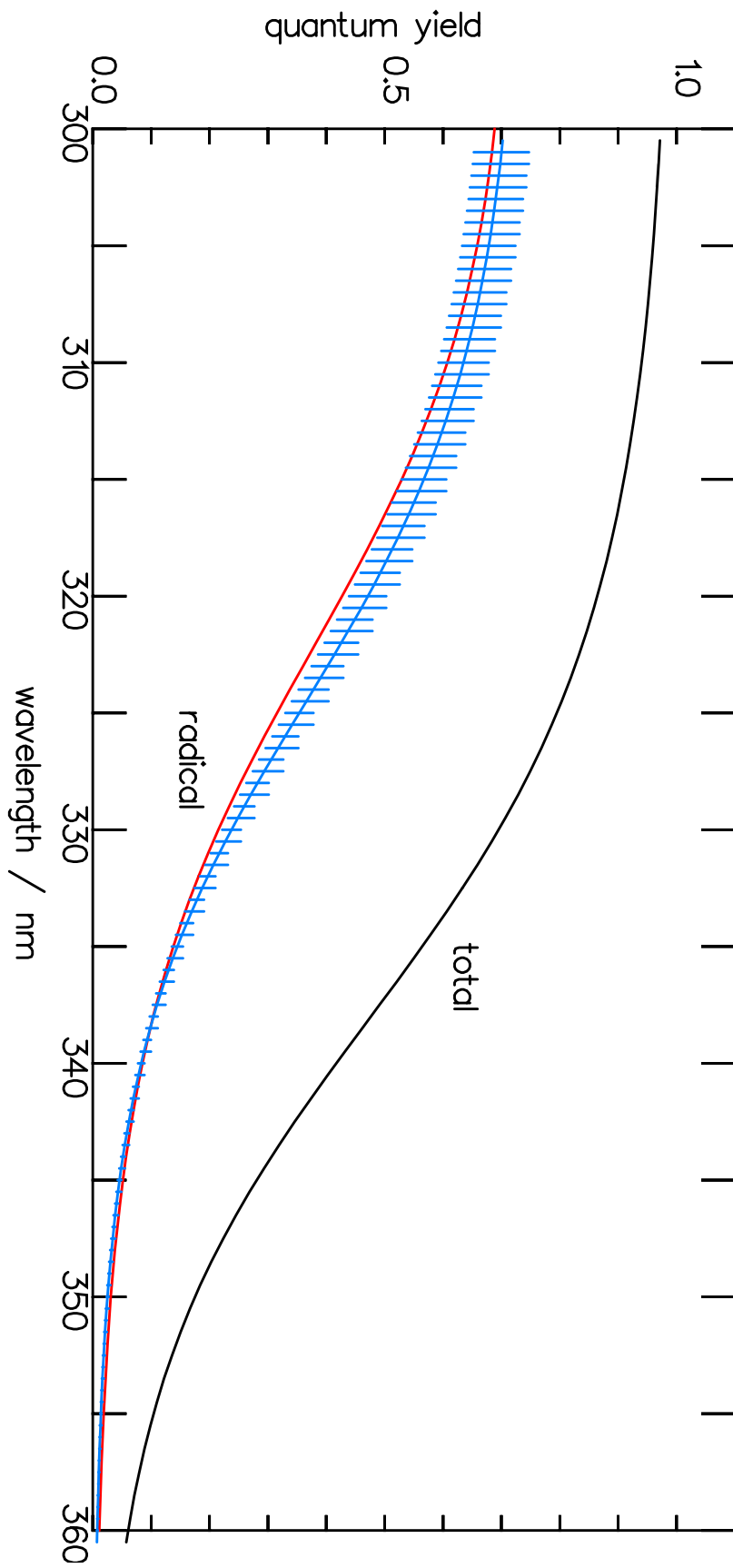


Fig. 9

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Fig. 10

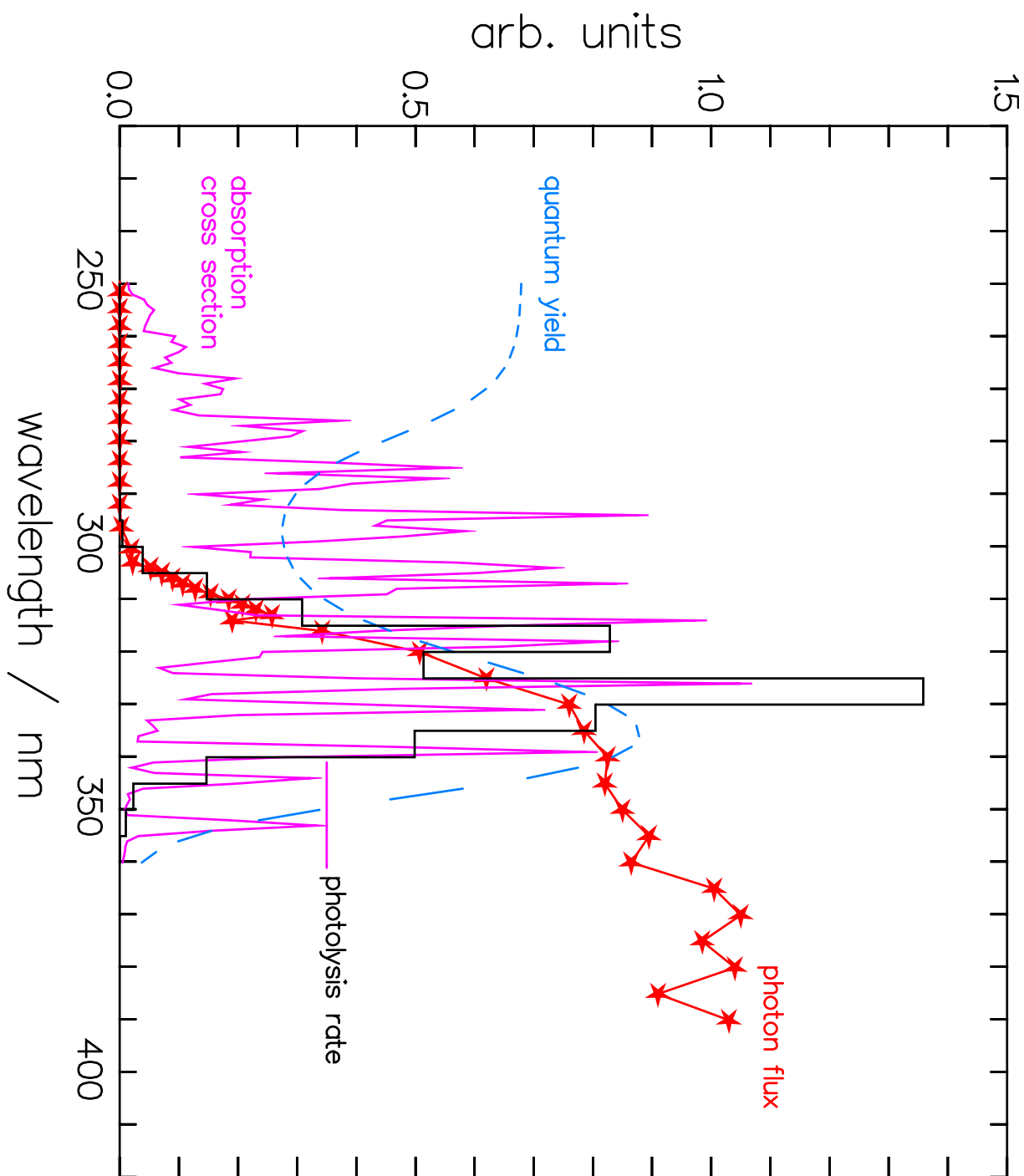
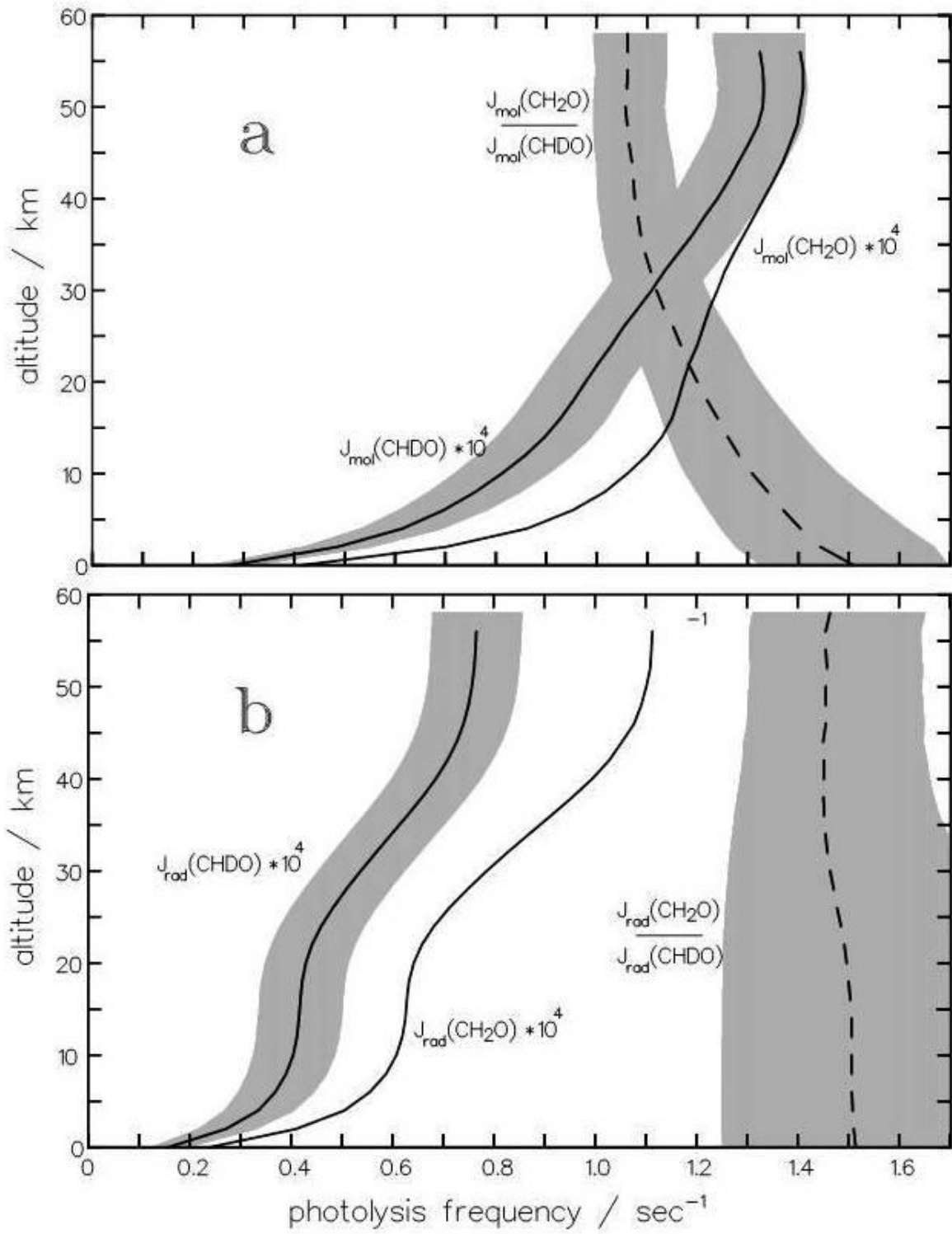


Fig. 11

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Fig. 12

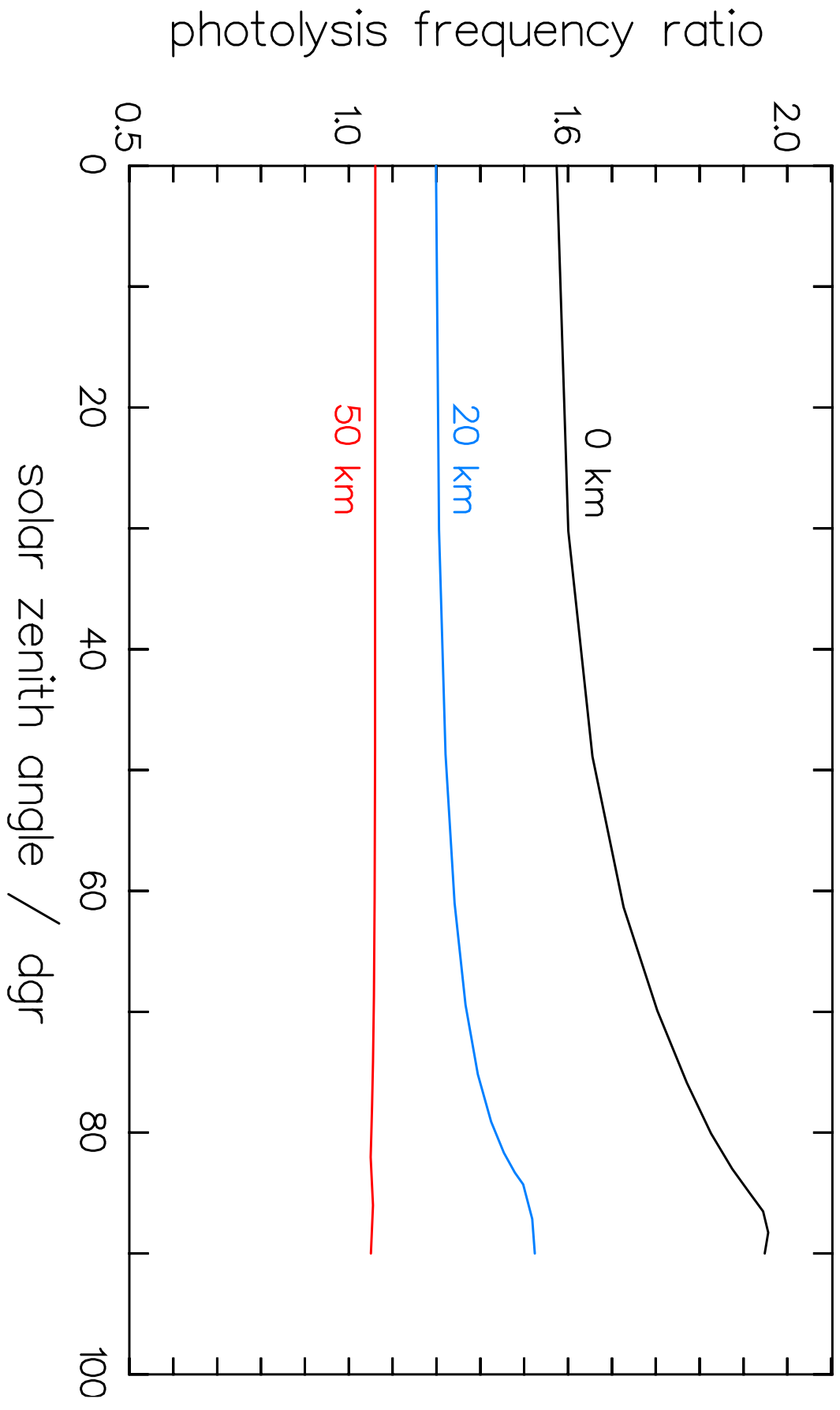
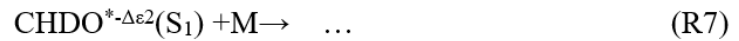
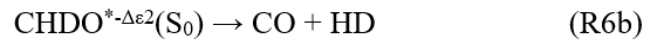
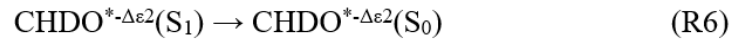
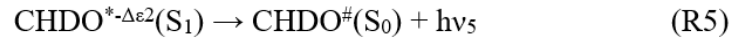
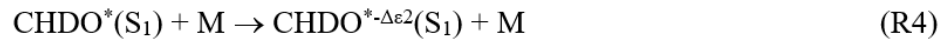
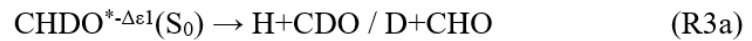
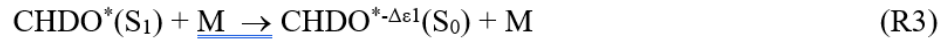
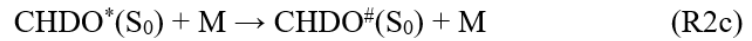
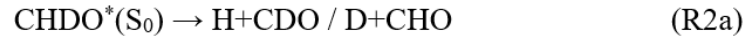


Fig. 13

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Reaction scheme (Table 1)