

# The Weddell Gyre heat budget associated with the Warm Deep Water circulation derived from Argo Floats

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**Abstract:** The Weddell Gyre plays an important role in the global climate system by supplying heat to underneath the ice shelves, and to the formation of deep and bottom water masses, which have been subject to widespread warming over past decades. In this study, we investigate the redistribution of heat throughout the Weddell Gyre by diagnosing the terms of the heat conservation equation for a 1000 m thick layer of water encompassing the core of Warm Deep Water. The spatial distribution of the different advective and diffusive terms in terms of heat tendencies are estimated using gridded climatologies of temperature and velocity, obtained from Argo floats in the Weddell Gyre from 2002 to 2016. While the results are somewhat noisy on the grid scale, and the representation of the effects of eddy mixing is highly uncertain due to need to parameterise them by means of turbulent diffusion, the heat budget (i.e., the sum of all terms) nearly-closes (within the uncertainty range) when integrated over the open inflow region in the southern limb, whereas ~~and~~ the interior circulation cell remains unbalanced of the Weddell Gyre. There is an overall balance in the southern limb between the mean horizontal advection and horizontal turbulent diffusion of heat, whereas the vertical terms contribute comparatively little to the heat budget throughout the Weddell Gyre. Heat convergence due to mean horizontal advection balances with divergence due to horizontal turbulent diffusion in the open southern limb of the Weddell Gyre. In contrast, heat divergence due to mean horizontal advection nearly balance is much weaker with than convergence due to horizontal turbulent diffusion in the interior circulation cell of the Weddell Gyre, due to large values in the latter along the northern boundary, due to large meridional temperature gradients. Heat is advected into the Weddell Gyre along the southern limb, some of which is turbulently diffused northwards into the interior circulation cell, while some is likely turbulently diffused southwards towards the shelf seas. This suggests that horizontal turbulent diffusion plays a role in transporting heat both towards the gyre interior where upwelling occurs, as well as towards the ice shelves. Horizontal turbulent diffusion is also a mechanism by which heat can be transported into the Weddell Gyre across the open northern boundary. Temporal deviations from the mean terms are not included due to study limitations. In order to appreciate the role of transient eddying processes, a continued effort to increase the spatial and temporal coverage of observations in the eastern Weddell Sea is required.

**Plain Language Summary:** Ocean currents in the Weddell Sea are governed by a wind-driven clockwise circulating cyclonic gyre-circulation, which is connected to the Atlantic sector of the Antarctic circumpolar-Circumpolar Southern Ocean Current to its north. Cyclonic means that the horizontal circulation is clockwise, and deep waters are brought close to the surface in the centre by upwelling. Warm and salty deep water enters the Weddell Sea in its east, and is transported by the gyre circulation first southward, then westward, and back northwards again following the continental boundaries. During this circulation, some of this water becomes lighter by mixing with the overlying surface waters, thus shoaling as it circulates. It also ~~The eastern limb of this cyclonic gyre transports relatively warm water poleward, where it~~ loses heat to the atmosphere and by contact with the ice shelves. When this occurs, Thereby the water becomes denser/heavier, as also by salty brine released during the freezing of sea ice. The densest/heaviest waters masses then sink down wards along the Antarctic continental slope, eventually and after northward transport with western limb of the gyre filling the deepest abyssal ocean basins ~~of abyssal global ocean~~. The main

source water mass for these processes and also the main source of heat to the Weddell Sea is ~~the so-called~~ Warm Deep Water. Previous studies have shown the whole water column, especially in the deeper layers, is warming over recent decades in the Weddell Sea. ~~The temperature of Warm Deep Water, however, fluctuates too strongly to tease out long-term trends from the “snapshot” data that is available to us. Warm Deep Water, however, varies in its properties too strongly to tease out long-term trends.~~ To better understand how ~~Warm Deep Water heat is redistributed~~ ~~is heat throughout~~ in the Weddell Gyre ~~within the Warm Deep Water~~, we ~~use combine hydrography temperature and drift~~ velocity observations from a fleet of Argo floats freely drifting throughout the Weddell Gyre between 2002 and 2016. ~~Using these observations, w~~We estimate a heat budget in ~~the layer that extends a 1000 m thick layer deep from below the surface layer, where the upper boundary is defined as the mid-thermocline, which varies typically around 150 m.~~ This layer always includes ~~us, we ensure that~~ the core of Warm Deep Water, ~~characterised by a sub-surface temperature maximum, is always included in the 1000 m thick layer,~~ regardless of its vertical position in the water column. Overall, large uncertainty ~~and variability~~ prevents us from interpreting the results on a local scale, but ~~interesting interpretable features of heat flux divergence and convergence~~ features emerge when integrating the heat budget over large areas. ~~The large-scale currents carry heat into the westward-flowing southern limb from the east, and upwelling brings heat upwards from below the layer throughout the whole gyre. According to our results heat is advected into the westward-flowing southern limb from the east, but also upwards from below the layer throughout the whole gyre.~~ Turbulent ~~diffusion~~ ~~mixing, representing small scale processes,~~ removes heat from the Warm Deep Water core through the top of the layer ~~throughout (i.e., upwards through the thermocline into the upper-ocean surface throughout.), and~~ It also ~~removes~~ ~~transports~~ heat from the southern limb, ~~northwards~~ into the ~~interior of the central~~ gyre where Warm Deep Water recirculates and moves closer to the surface, as well as southwards towards the Antarctic coastline. Lastly, turbulent ~~diffusion-mixing~~ also brings heat into the gyre across the northern boundary.

## 1 Introduction

Understanding the drivers and pathways of large-scale ocean circulation is a fundamental component of climate science (Rhein et al., 2013). To comprehend the regulatory role of the oceans in the climate system, one can determine the ocean heat budget, which describes the redistribution of heat throughout the ocean by means of horizontal and vertical advection, turbulent diffusion, and surface heat fluxes to the atmosphere (e.g., Tamsitt et al., 2016; an adapted form of the heat budget equation is given in Eq. 1.1).

The Weddell Gyre is located south of 50° S in the Atlantic sector of the Southern Ocean, where Circumpolar Deep Water (CDW) predominantly enters the gyre’s southern limb in the east at about 30° E. The CDW that enters the Weddell Gyre ~~is becomes modified, and is~~ commonly referred to as Warm Deep Water (WDW). WDW circulates the cyclonic gyre, undergoing cooling and freshening ~~en-route~~, through interaction with the underlying and overlying water masses (e.g., Fahrbach et al., 2004; Klatt et al., 2005; Fahrbach et al., 2011; Leach et al., 2011). The core of WDW is identified as the sub-surface temperature maximum (hereon referred to as  $\Theta_{\max}$ ), which feeds heat into the Weddell Gyre (Fahrbach et al., 2004, 2011; Cisewski et al., 2011; Ryan et al., 2016). The distribution of temperature at the depth of the  $\Theta_{\max}$  is shown in Fig. 1, which is derived from in situ observations from Argo floats (from Reeve et al., 2016, 2019). The Weddell Gyre has been ascribed the role of a heat buffer (Fahrbach et al., 2011), in that it acts to store and redistribute heat and salt in the water column, effectively transferring heat to the deeper layers where it is ultimately exported northwards, spreading throughout the abyssal global ocean (e.g., Foster et al., 1987; Naveira Garabato et al., 2002, 2016; Fahrbach et al., 2011).

Warming trends over recent decades have been observed in the WSDW and WSBW (Weddell Sea –Deep and -Bottom Water respectively) (Fahrbach et al., 2011; Meredith et al., 2011; Strass et al., 2020). However, their primary source of heat, WDW, exhibits pronounced decadal variations and shows no significant long-term warming trend (Fahrbach et al., 2011; Kerr et al., 2017). While Strass et al. (2020) show that the whole water column below 700 m exhibits a significant long-term warming

trend, which would incorporate the lower part of the WDW layer, they also observed significant variability and even areas of cooling in the upper 700 m, which incorporates a significant chunk of the WDW layer.

85 In this study, we combine observations of the velocity field (Reeve et al., 2019) with the temperature field (Reeve et al., 2016), both derived primarily from Argo floats, to diagnose components of the heat budget of a fixed volume of water fully encompassing the core of WDW within the Weddell Gyre. Given ~~the Argo float data lacks the spatial-temporal coverage to resolve the near-surface~~ seasonal cycle ~~while also objectively mapping the entire region is unresolved throughout~~, a full basin analysis of the upper 50 m is unfeasible. By analysing the heat budget for a fixed volume encompassing the core of WDW, we can, however, determine the ways in which heat from WDW is redistributed throughout the Weddell Gyre. The rest of this paper is structured as follows. Section 2 describes the data sources used in this study, while Section 3 details the method of applying the heat budget, and its associated estimation of uncertainty. Section 4 presents the individual heat budget terms for the Weddell Gyre for the whole region as well as integrated over specific areas, while Section 5 interprets these results in context of the study limitations and the literature. Lastly, Section 6 provides our final summary.

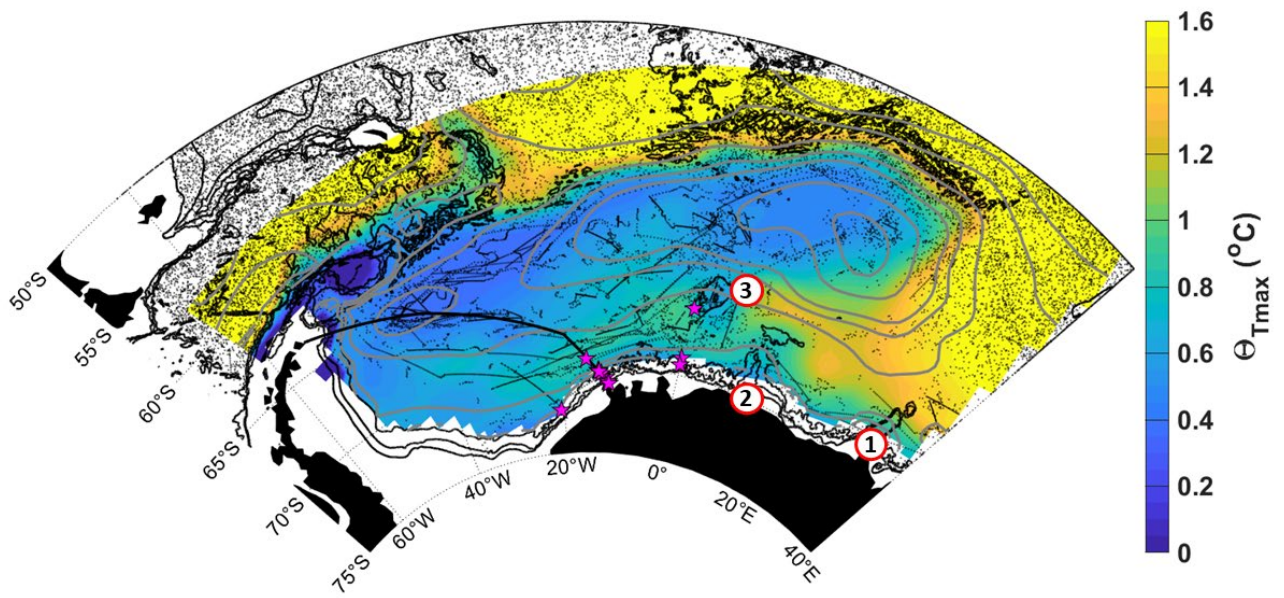


Figure 1: Sub-surface conservative temperature ~~at the depth of temperature~~ maximum ( $\Theta_{\max}$ ) with streamlines (grey contours) of the vertically integrated stream function for 50-2000 dbar with a spacing of 5 Sv, derived from in-situ observations from Argo floats (Reeve et al., 2019, 2016). Black dots show Argo float profile positions and ~~red-magenta~~ stars show mooring positions used in velocity field estimates. The thick black line shows the repeat ship-based transect from Kapp Norvegia to Joineville Island. The red circles labelled 1-3 show ~~(1) Gunnerus Ridge, (2) Astrid Ridge and (3) Maud Rise, respectively~~. The black contours show the 1000, 2000 and 3000 m isobaths, from the general bathymetric chart of the oceans (GEBCO, IOC et al., 2003).

## 95 2 Data sources

For this study we computed the heat budget of the entire Weddell Gyre using different sets of data from various sources, which will be described in more detail in the following subsections.

### 2.1 Data sources

100 Table 1 describes the data sources used for determining the heat budget presented in this paper. Each source is then described in further detail in Sections 2 and 3, as well as in the Supplement (S1-4).

Table 1. A summary of the associated data source, method and citation for each variable required in the heat budget (Eq. 1.2 in Section 3, and column 2 below).

Section for details	Variable	Data source	Methodology	Citation
2.2	Gridded Temperature ( $\theta$ )	Argo floats (2002-2016)	Objective mapping	Reeve et al., 2016
2.2	Gridded horizontal velocity ( $U$ )	Argo floats (2002-2016)	Objective mapping, cost function	Reeve et al., 2019
2.2	Long-term mean velocity at the shelf edge ( $U$ )	Moorings (1989-2016)		Le Piah et al., 2020
2.2	Bathymetry	GEBCO		IOC et al., 2003
S3	Vertical (Ekman) velocity ( $w_E$ )	ERA5-interim reanalysis and Polar Pathfinder (listed below)	Regridded using a distance-weighted mean	Hersbach et al., (2020) & Tschudi et al., (2019)
S3	Wind stress ( $\tau_{\text{air-ocean}}$ ) derived from wind field at 10 m above sea surface ( $U_{10}$ )	ERA5-interim reanalysis	Used to compute $w_E$	Hersbach et al., (2020)
S3	Sea-ice concentration ( $\alpha$ )	ERA5-interim reanalysis	Used to compute $w_E$	Hersbach et al., (2020)
S3	Sea-ice velocity ( $U_{\text{ice}}$ )	Polar Pathfinder Daily 25 km EASE-Grid Sea Ice Motion Vectors	Used to compute $w_E$	Tschudi et al., (2019)
3 & S4	Horizontal diffusivity ( $\kappa_H$ )	Sevellec et al., (2022), and estimates in the literature (see citation)	Where available (i.e., outskirts of gyre) the Sevellec dataset. Infilled within gyre with a constant value based on the literature	Sevellec et al., 2022; Donnelly et al. (2017); Zika et al. (2009) & Cole et al., (2015)
3	Vertical diffusivity ( $\kappa_V$ )	Based on estimates in the literature (see citation)		Donnelly et al. (2017), Zika et al. (2009) & Cole et al., (2015)
S2	Air-sea heat fluxes	ERA-interim reanalysis		Jones et al., 2016

## 2.2 Gridded velocity and temperature fields derived from Argo floats:

All Argo float data available for the Weddell Gyre region between 2002 and 2016 were used in this study to create gridded fields of velocity and conservative temperature. The latter was derived using the TEOS-10 program in MATLAB (McDougall et al., 2011). We applied objective mapping to the profile data resulting in a climatology of gridded conservative temperature on 41 pressure levels between 50 and 2000 dbar. The grid cell resolution varies slightly with changing latitude (Reeve et al. 2019), but is on the order of  $\sim 80 \times 60$  km. This method follows, and is an extension of the objective mapping provided in Reeve et al. (2016), and the reader is referred to that study for further details on the data quality control and mapping method used in this study. The absolute velocity field was derived from Argo float trajectory data at the depth of the float drift (800 dbar in the Weddell Gyre). This process required careful quality control assessments and surface drift corrections following Park et al. (2005), and, given that under-ice profiles have no geo-located position, all such interpolated trajectories were omitted from the study. Thus, the derived velocity field exhibits considerable bias to summer conditions. The velocities were objectively mapped to provide a grid of absolute velocity at 800 dbar. This process is detailed in Reeve et al. (2019).

Following Reeve et al. (2019), a stream function was fitted to the velocity field at 800 dbar through the application of a cost function. By applying a cost function, the resulting stream function provides a best fit for the entire Weddell Gyre, where the direction of mean flow at the boundaries are assumed to be parallel to the boundaries. The stream function at 800 dbar provides the reference level for the relative geostrophic velocity, derived from the gridded density field from Reeve et al. (2016), above. For full details on this methodology, refer to Reeve et al. (2019), which, in addition to Reeve et al. 2016, is the prelude to this study. In Reeve et al. (2019), a careful error analysis and a detailed comparison of volume transports to available

120 estimates in the literature justify the method as a reasonable solution to obtaining a large-scale observation-based estimate of the Weddell Gyre circulation.

There are, however, important improvements between the stream function provided in Reeve et al. (2019), and the stream function used in this study. The cost function was adapted to allow for a variable coastline, so that the stream function provides a solution which includes the southern ice shelf edge east of the Prime Meridian (though not the shelf edge currents, such as the Antarctic Slope Front, which are not resolved in this study as they are beyond the domain of the Argo floats), which is excluded in Reeve et al. (2019). To better estimate the velocity field along the southern coast of the Weddell Gyre, a few minor adjustments were made to the gridded velocity field prior to fitting a stream function. Firstly, long-term average velocities derived from mooring data were included at the coast near the Prime Meridian and Kapp Norvegia, to better resolve the flow which follows the coastline as it curves southwards towards the Filchner-Ronne ice shelves (the mooring positions are marked in Fig. 1. See Le Piah et al. 2020 for further information about the mooring data). Secondly, the velocity field at Gunnerus Ridge (also marked in Fig. 1) required special treatment. The trajectories of Argo floats show a tight, bathymetrically steered flow around Gunnerus Ridge, which is lost during the objective mapping process due to larger length scales. Also, while the potential vorticity values to either side of the ridge are the same, the direction of the flow is opposing (i.e., primarily northwards on the eastern flank of Gunnerus Ridge and southwards to the west of Gunnerus Ridge; see Fig. 11 in Reeve et al. 2019). This opposing direction is averaged out in the objective mapping. Therefore, after the objective mapping, the closest grid cells to Gunnerus Ridge are replaced with direct velocity measurements derived from the three Argo floats that drift along the ridge (the Argo floats have WMO numbers: 7900164, 7900166 and 7900168). There are caveats to this decision, in that these are data points from three floats with a limited time span (from February-May 2007, and then December 2007 until April 2008), during a period when the area is ice-free. However, by making this adjustment, we improve the performance of the cost function in providing a stream function representative of the large-scale circulation, which includes a more complete inflow (in comparison to Fig. 4 in Reeve et al. 2019). The resulting stream function for the vertically integrated flow from 50-2000 dbar is shown in Fig. 1, where the streamlines curve around Gunnerus Ridge, indicating the main inflow into the southern limb of the Weddell Gyre.

### 3 Methods

#### 3.1 The heat budget

145 Due to conservation, the rate of change of heat storage of a certain ocean volume needs to equal the sum of all fluxes comprising advection and turbulent diffusion, both vertical and horizontal. Thus, the heat budget integrated for a volume of water not in contact with the atmosphere is defined as:

$$\frac{d\theta}{dt} = -\nabla_H(U \cdot \theta) - w_E \frac{d\theta}{dz} + \kappa_H \nabla_H^2 \theta + \kappa_V \frac{d^2 \theta}{dz^2} + R \quad (1.1)$$

$$\rho_0 C_p \int_{m_T-1000}^{m_T} \frac{d\theta}{dt} dz = \rho_0 C_p \left\{ \int_{m_T-1000}^{m_T} (-\nabla_H(U \cdot \theta) + \kappa_H \nabla_H^2 \theta) dz - w_E (\theta_{m_T} - \theta_{m_T-1000}) + \kappa_V \left( \frac{\partial \theta}{\partial z} \Big|_{m_T} - \frac{\partial \theta}{\partial z} \Big|_{m_T-1000} \right) + R \right\} \quad (1.2)$$

150 where  $\nabla_H$  is the horizontal divergence operator,  $U$  is the horizontal geostrophic velocity,  $\theta$  is the conservative temperature,  $w_E$  is the vertical velocity (defined as the Ekman pumping velocity; see Supplements S3),  $z$  is depth and  $\kappa_H$  and  $\kappa_V$  are the horizontal and vertical diffusivity respectively (adapted from Tamsitt et al. 2016). R represents the unresolved transient processes excluded in this study (discussed below). For the vertical integration in Eq. 1.2, the subscript  $m_T$  describes the mid-point of the thermocline, which provides the upper boundary, while  $m_T + 1000$  describes the lower boundary. This is to avoid incorporating highly seasonally variable surface waters from the analysis whilst also fixing the volume of water; (detailed

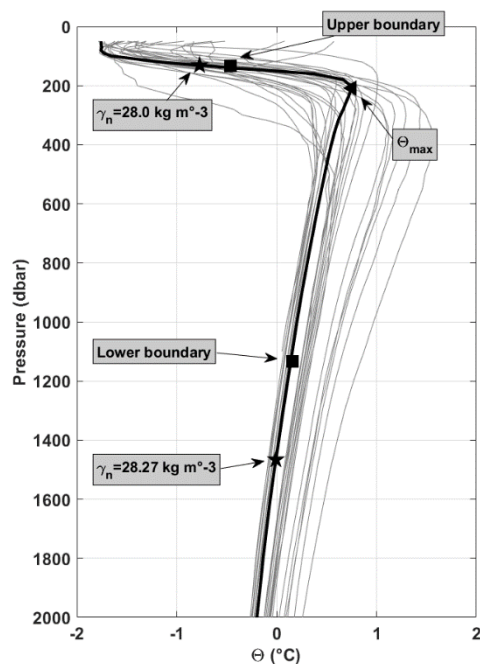
155 explanation of the vertical boundaries is provided in the Supplements S1). Figure 2 shows selected vertical profiles with the upper and lower boundaries marked (the corresponding position of the profiles is found in Fig. S1, selected at random to provide a broad coverage of the Weddell Sea). Note the upper boundary temperature is always less than the lower boundary temperature within the Weddell Sea (there are regions where the opposite is true to the north of the gyre, within the ACC).

160 Each term is multiplied by the specific heat capacity of seawater,  $C_p$  ( $\sim 4000 \text{ J K}^{-1} \text{ Kg}^{-1}$ ), and seawater density, ( $\rho_0 = 1027 \text{ kg m}^{-3}$ ), and integrated for a 1000 m thick layer so that units of each component are given in  $\text{W m}^{-2}$ . The first term on the right-hand side in Eq. 1.1 describes the mean horizontal geostrophic heat advection, where  $U$  is derived from horizontal differentiation of the geostrophic stream function derived from Argo float data (i.e., where  $u = \partial\psi/\partial y$  and  $v = -\partial\psi/\partial x$ ; see Section 2.2 and Reeve et al. 2019). Since we derive velocity from a non-divergent stream function, we assume geostrophic flow conditions and omit ageostrophic advection from the first term. This is a pertinent assumption given that the Ekman Layer is excluded from the analysis. The second term on the right-hand side in Eq. 1.1 describes the mean vertical heat advection. The third and fourth terms in Eq. 1.1 (or, the second part of the first term and the third term in Eq. 1.2) describe the horizontal and vertical turbulent heat diffusion components respectively. The sum of these terms results in an estimate of heat tendency over time ( $d\theta/dt$ ), which can be used to determine mean temperature change for a column of water, although this method results in an accumulation of associated errors of the individual terms, and should therefore be treated with caution. Note that given the constraints of the method used (i.e., our data resources are an objectively mapped long-term mean temperature field, and horizontal velocity derived from a long-term mean stream function of the Weddell Gyre, derived from in situ observations), we are unable to look at deviations from the mean, i.e., transient processes. This means, that the meaning of advective and diffusive heat fluxes are different from the ones quantified by Tamsitt et al. (2016). Their underlying numerical model resolves large-scale variations of the flow and temperature field and (partly) mesoscale eddies; these processes are part of the advection, while turbulent diffusion refers to unresolved small-scale processes. In our study, advection is computed from time mean quantities while the effects of mesoscale eddies are parameterised by horizontal turbulent diffusion. Large-scale variations of the flow field are not accounted for in our study. Thus, in addition to the unknown time variability term,  $d\theta/dt$ , we also have an unknown 5<sup>th</sup> term in the heat budget in Eq. 1.1 and 1.2,  $R$ , which represents the unresolved transient processes excluded from the study. Increased spatial and temporal coverage of in situ observations within the Weddell Gyre would be required to address these gaps (further discussed in section 5.2.2). Maps of the different components are provided in the results section.

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**Figure 2: (a) A sample of vertical profiles of conservative temperature (the position of the profiles are marked by stars in Fig. S1, the red star marking the position of the example profile in black), where  $\Theta_{\max}$  is marked by a triangle, the upper boundary (i.e., mid-thermocline) and the lower boundary (mid-thermocline + 1000) are marked by squares, and the upper and lower WDW boundaries (defined by a neutral density range of 28 to 28.27 kg m<sup>-3</sup>) are marked by stars. This is to highlight that our method for the vertical boundary limits allows for the full inclusion of the core of WDW while also excluding Winter Water.**

There are two further unknowns in Eq. 1. These are the horizontal and vertical diffusivities,  $\kappa_H$  and  $\kappa_V$  respectively. For the northern and eastern outskirts of the Weddell Gyre, we define  $\kappa_H$  from the dataset provided by Sevellec et al. (2022), who derive horizontal diffusivities directly from Argo float trajectories by fitting a “pseudo-trajectory” to increase the spatial resolution required for the computation. Given this requires trajectory data without gaps in the record, estimates are missing for much of the Weddell Sea due to the presence of sea-ice. For the ~~rest of the remainder of the~~ Weddell Gyre, we define  $\kappa_H$  as  $400\text{-}247 \pm 200\text{-}63 \text{ m}^2\text{s}^{-1}$ , and  $\kappa_V$  as  $(2.39 \pm 2.83) \times 10^{-5} \text{ m}^2\text{s}^{-1}$  based on the estimates provided by Donnelly et al. (2017), which are derived from ship-based observations throughout the Weddell Sea in combination with velocity estimates from the ECCO live access server. varying values provided in the literature. (See Supplements S4 for our reasoning and maps of  $\kappa_H$ ). We define  $\kappa_V$  as  $2.6 \times 10^{-5} \pm 2.4 \times 10^{-5} \text{ m}^2\text{s}^{-1}$ , again based on values provided in the literature. The error ranges are to provide a range of reasonable estimates while acknowledging lack of consensus of appropriate values for the diffusivities. While  $\kappa_V$  is similar to Donnelly et al.’s estimate of  $2.4 \times 10^{-5} \text{ m}^2\text{s}^{-1}$ , our baseline horizontal diffusivity is slightly larger than the  $247 \text{ m}^2\text{s}^{-1}$  provided by Donnelly et al. (2017; though within range of our values provided by the uncertainty range) and is chosen in consideration of a range of values provided by Donnelly et al. (2017), Zika et al. (2009) and Cole et al. (2015), of  $247 \text{ m}^2\text{s}^{-1}$ ,  $300 \text{ m}^2\text{s}^{-1}$  and  $200\text{-}2000 \text{ m}^2\text{s}^{-1}$  respectively. Our chosen baseline  $\kappa_H$  is larger than the first two listed here and more closely aligned with Cole et al. (2015), because we focus on larger length scales, most similar to the latter study, owing to the resolution and sample density of our data set. The implications of defining the diffusivities in this way these decisions are discussed in Section 5.

### 3.2 Assessing the uncertainty

The errors for each heat budget term are calculated using the laws of propagation, as detailed in the Supplement to this paper (S7). The main sources of error are from the variables: temperature, horizontal velocity, vertical (Ekman) velocity, and unknown diffusivity, which is assumed constant in the horizontal for the Weddell Gyre interior and southern limb, and throughout the whole Weddell Sea in the vertical.

We used the objective mapping error to represent the error for temperature (Fig. S7), which is provided in Reeve et al. (2016), and assumed to be the dominating source of error for temperature. This means that the error is also representative of the length scales applied in the objective mapping (Reeve et al., 2016). In Reeve et al. (2016), the length scales were assigned based on an investigation which showed that 95 % of the grid points have at least 40 Argo profiles within a distance of 500 km (we were then able to reduce this to 400 km in Reeve et al. 2019), which was thus the length scale applied in the second stage of the objective mapping, along with a fractional scale on the effect of  $f/H$ , which alters the shape of the area of influence about a grid point from circular, when the bottom bathymetry is flat, to elongated, when a grid point is in close proximity to bathymetric gradients (Reeve et al., 2016). The fractional scale on the effect of  $f/H$  allows us to take into account the tendency of water parcels to follow paths of constant  $f/H$  in order to conserve vorticity –(for a visual demonstration of the impact of including  $f/H$  in the scaling, see Fig. 9 in Reeve et al. 2016). Since the focus is a climatological mean from 2002 to 2016, the applied length scales are chosen for the mapping to represent the large-scale field of the entire Weddell Gyre (Reeve et al., 2016). The resulting mapping errors are large in regions where bathymetry is complex and data coverage is sparse, and low in regions where the bathymetry is flat or where data density is high. Thus, the regions of largest uncertainty include the northern

periphery of the gyre, where data ~~is~~ are relatively abundant but the bathymetry is complex, and the eastern edge of the Weddell Gyre, where data ~~is~~ are sparse (Fig. S7).

The error for horizontal velocity was provided in Reeve et al. (2019), and is derived from the stream function, where a sensitivity study was implemented, perturbing the velocity field using a combination of factors (mapping errors and drift correction), to provide a range of possible stream function values, from which a standard error was estimated.

Since uncertainty is currently unavailable for the ERA-5 reanalysis data that provided the wind stress field from which Ekman vertical velocity was computed, we took the standard error of the mean of a monthly time series from January 2002 to December 2016 to represent the error for vertical velocity. This represents the natural temporal variability of the vertical velocity, which we assume to be dominating over other possible sources of error, given the large seasonal variability of the wind field.

Lastly, we ~~arbitrarily~~ define an uncertainty range for the diffusivity terms from Donnelly et al. (2017), to be  $\pm 200\text{--}63\text{ m}^2\text{s}^{-1}$  and  $2.4 \times 10^3 \times 10^{-5}\text{ m}^2\text{s}^{-1}$  for the horizontal and vertical diffusivity terms respectively (with the exception of the regions where the Sevellec et al. (2022) dataset is used, where a standard error of the mean is used, see Supplements S4). ~~These are large ranges to account for the possible range in values available in the literature, while acknowledging the lack of consensus and reliable data for this input (w~~ We discuss this further in Section 5.1.3). The derivation for the propagated errors is provided in the Supplement (S7), ~~along with~~ whilst the maps of the error for each heat budget term are provided alongside the heat budget terms in Fig 3. The errors are provided in the large-scale integrations of the IC and SL in Figs. 3-5-7 and -7-9 (as pale blocks of colour surrounding the lines).

## 4 Results

Each of the maps provided throughout this paper show streamlines (grey contours) indicating the horizontal circulation of the Weddell Gyre derived from Argo floats (see Section 2.2 and Reeve et al. 2019 for further details). Where the streamlines are closely spaced, flow is more intense than where they are loosely spaced. The streamlines describe a double-gyre structure whereby the western sub-gyre is weaker than the elongated, stronger eastern sub-gyre. While large uncertainties are associated with the eastern part of the eastern sub-gyre, numerical model simulations (e.g., Timmermann et al., 2002), historical observations (e.g., Orsi et al. 1993), and direct volume transport estimates support the concept of a double-gyre structure in the Weddell Sea (e.g., Fig. 7 in Reeve et al., 2019). The following section is presented in two parts. In part one, we provide maps of the vertically integrated heat budget terms (from Eq. 1.2), to obtain an overview of the large-scale heat field of the Weddell Gyre, ~~whereas~~ in part 2, we consider the larger-scale zonal variation of the heat budget, for two regions: (1) the southern limb (SL), which describes the westward flowing part of the gyre that extends from Gunnerus Ridge ( $\sim 33^\circ\text{E}$ ) to  $\sim 45^\circ\text{W}$  (i.e., where the stream function,  $\psi$ , is 16 to 30 Sv; the stippled area in Fig. 2b) and (2) the interior circulation cell (IC), where streamlines form a fully closed circuit west of Gunnerus Ridge, incorporating both the eastern and western sub-gyres, (i.e.  $\psi > 25.9\text{ Sv}$ ; the stippled area in Fig. 2d).

### 4.1 Part one: the large-scale investigation of heat within the Weddell Gyre

The heat budget contributions from the different terms in Eq. 1.2 (including the signs of each term, e.g.,  $-\nabla_H U \theta$  and  $+\kappa_H \nabla^2 \theta$ ) are provided in Fig. 23. While for the gyre at large, the mean horizontal geostrophic heat advection (Fig. 2a3a, left panel) shows a patchwork display of heat transport convergence (positive) and divergence (negative), the southern limb of the gyre is generally dominated by heat transport convergence, of about  $+20\text{ W m}^{-2}$  west of the Prime Meridian. While small areas of heat flux divergence are found throughout the southern limb, these small areas appear to be less prominent than the areas of



255 ~~positive heat flux and are unlikely to be significant given the high associated errors (Fig. 3a, right panel). A small patch of~~  
~~divergence is found over the western sub-gyre of about  $-10$  to  $-20$   $\text{W m}^{-2}$  ( $-35$   $45^\circ\text{W}$ ,  $-65^\circ\text{S}$ ), and over Maud Rise of about~~  
 ~~$-20$  to  $-30$   $\text{W m}^{-2}$  ( $-3^\circ\text{W}$ ,  $65^\circ\text{S}$ ).~~ The whole region east of the Prime Meridian is dominated by particularly strong patches of  
260 positive and negative values in excess of  $\pm 80$   $\text{W m}^{-2}$ . Along the northern limb of the gyre, the pattern is dominated by bands  
of alternating positive and negative values, of about  $\pm 60$ - $80$   $\text{W m}^{-2}$ , which are aligned in a manner that appears to follow the  
complex bathymetry in the region. Note associated errors become increasingly large directly to the north of the northern limb,  
related to the highly dynamic boundary.

The heat flux due to mean vertical advection ( $-w_E(\Theta_{mT} - \Theta_{mT-1000})$ , Fig. 2b3b, left panel) is positive throughout, and considerably  
weaker than that due to mean horizontal advection, in the range of  $\sim 0$ - $3$   $\text{W m}^{-2}$ . Vertical advection is weakest west of  $\sim 10^\circ\text{W}$ ,  
and strongest over the eastern sub-gyre region between  $0$  and  $30^\circ\text{E}$ . That the entire region shows positive vertical fluxes (with  
265 the exception of north of the Weddell Gyre) results from two factors: (1) the mean Ekman pumping velocity is positive  
(indicating upwelling) throughout the offshore Weddell Gyre (downwelling, i.e. negative Ekman pumping velocity, is found  
in regions shallower than  $2000$  m and thus outside of our region of Argo float data availability) (Fig. S3), and (2) the upper  
boundary displays lower temperatures than the lower boundary (Fig. S1a2), which is a consequence of the vertical boundaries  
applied in this analysis. Positive vertical advection implies that more heat is advected upwards into the core layer of WDW  
270 from below than is leaving by advection through the top, implying that the net effect of vertical advection is to warm the layer  
in question ~~convergence of heat within that layer, unless it is removed through other mechanisms such as mean horizontal~~  
~~advection or turbulent diffusion.~~

Horizontal turbulent diffusion (Fig. 2e3c, left panel) is characterised by a positive signal of about  $0$ - $40$   $\text{W m}^{-2}$  within the ICgyre  
interior, and a negative signal along the SL-southern limb in the range of  $20$ - $650$   $\text{W m}^{-2}$ , with the exception of local patches of  
275 heat flux convergence such as over Maud Rise and just north of Astrid Ridge at  $\sim 10^\circ\text{E}$  (topographic features are marked in  
Fig. 1). The northern limb of the Weddell Gyre is mostly positive, ( $40$ 20- $100$   $\text{W m}^{-2}$ ), though a strip of heat flux divergence  
sits directly north of this area of heat flux convergence, in the northern boundary zone between the Weddell Gyre and the  
Antarctic Circumpolar Current.

As with mean vertical advection, vertical turbulent diffusion (Fig. 32d, left panel) exhibits a uniform sign throughout the  
280 Weddell Gyre, with negative values in the range of  $-2$ 1 to  $-7$ 6  $\text{W m}^{-2}$ . The strongest values are found in the southern limb, at  
about  $\sim 20^\circ\text{E}$  ( $\sim -7$ 6  $\text{W m}^{-2}$ ), as well as just west of Maud Rise. ~~There also appears to be slightly enhanced vertical turbulent~~  
~~diffusion along the gyre axis between about  $10^\circ\text{E}$  and  $20^\circ\text{W}$ .~~

The heat tendency resulting from the sum of the heat budget terms (left-hand side of eq. 1.2). is provided in Fig. 32e (left  
panel, with the corresponding propagation of error provided in the right panel). There is a patchwork of negative (cooling) and  
285 positive (warming) values throughout, and overall Fig. 2e-3e is spatially mostly similar to the heat flux due to mean horizontal  
advection and horizontal turbulent diffusion in Figs. 2a3a and 3e (the latter especially to the north of the gyre where diffusivity  
is large). There is, however, a discernible cooling throughout the open southern limb in Fig. 2e, in comparison to Fig. 2a, and  
a clear warming along the northern limb, driven by horizontal turbulent diffusion (Fig. 2e3c). Any non-zero value in Fig. 2e  
3e should correspond to an area of warming (positive tendency) or cooling (negative tendency) of the water column. Given  
290 that data from a 15 years-long observation period have entered the calculation, we wouldn't expect the real ocean to have  
experienced such a patchy warming and cooling pattern, resulting – in particular – from the horizontal advection field. In  
chapter-section 4.2 we therefore perform spatial integration of the fields displayed in Fig. 2a3a-e over distinct areas defined  
by the circulation (SL and IC), in order to eliminate some of the noise, such that more robust statements on the main balances  
between the different heat flux terms can be made on a regional scale. The uncertainties will be discussed further in chapter  
295 section 5.



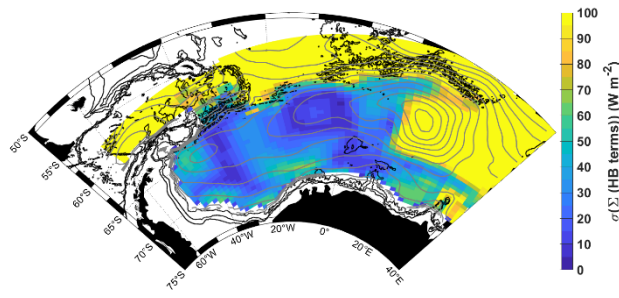
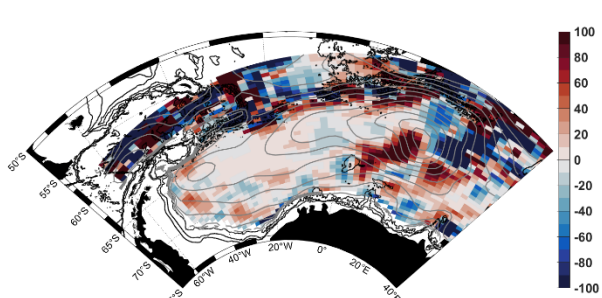
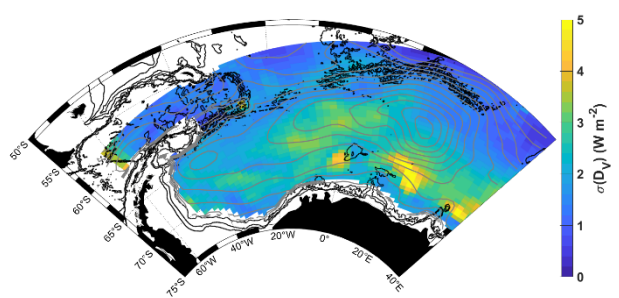
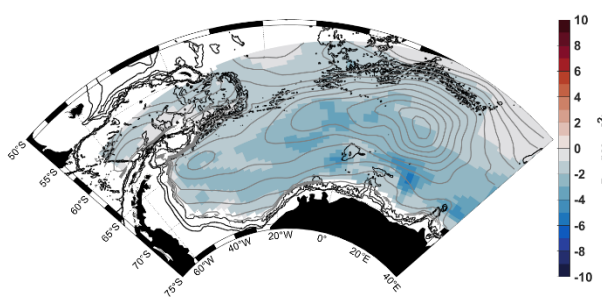
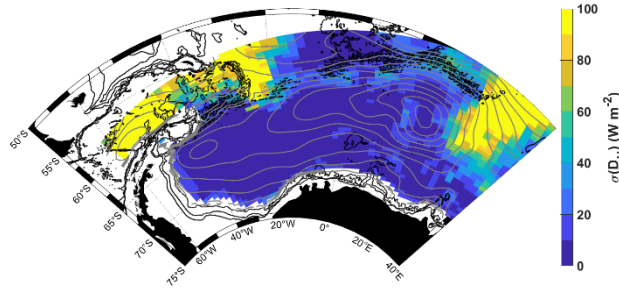
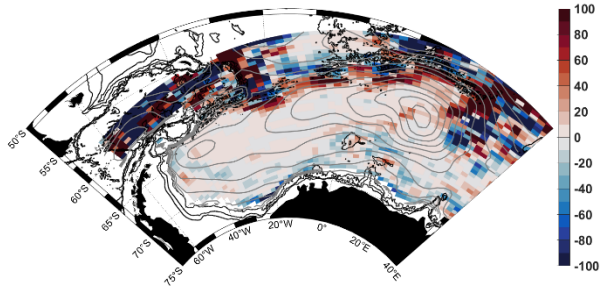
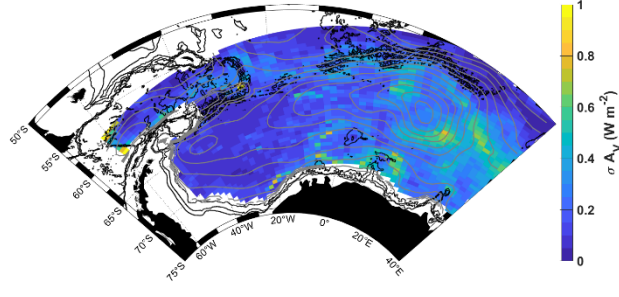
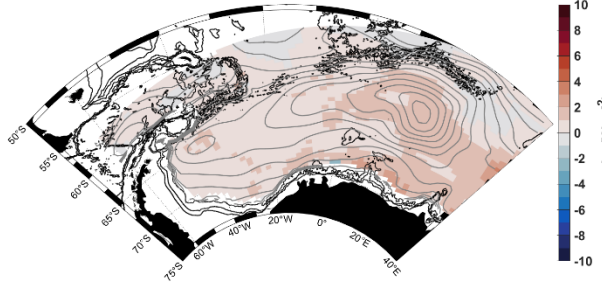
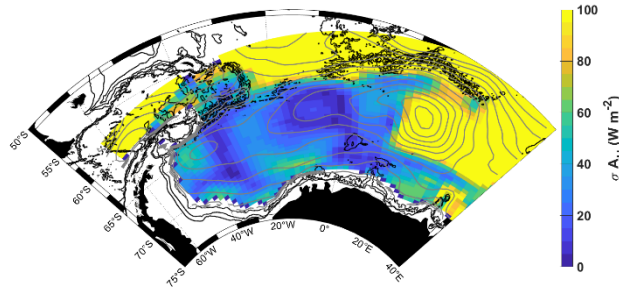
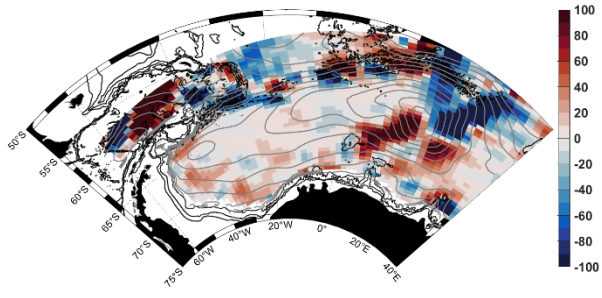
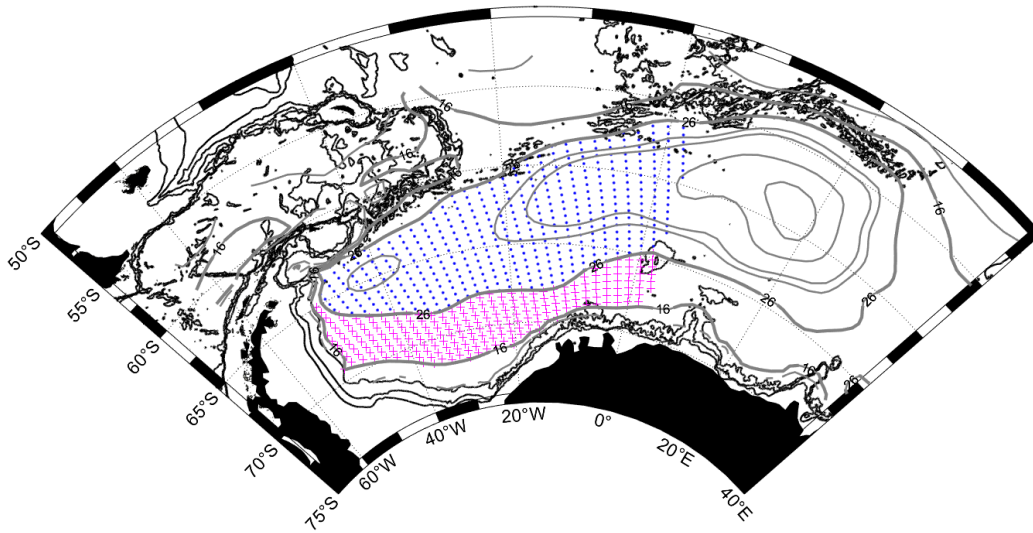


Figure 3: The panels on the left shows the heat budget terms in  $\text{W m}^{-2}$ , from Eq. 1.2, for a laver of water 1000 m thick from the depth of the mid-thermocline: (a) mean horizontal geostrophic heat advection, (b) mean vertical advection, (c) horizontal turbulent diffusion, (d) vertical turbulent diffusion, and (e) the sum of the terms in a-d. Positive values indicate warming, i.e. heat transport convergence, where more heat is entering the grid cell than is leaving it, whereas negative (blue) values indicate cooling, i.e. heat transport divergence, where more heat is leaving the grid cell than is entering it. The panels on the right show the corresponding total propagated errors of each term. Grey and black contours provide the horizontal streamlines and the 1000, 2000 and 3000 m isobaths respectively (as in Fig. 1). Note the different colour scales for the horizontal versus vertical fluxes.

## 4.2 Results part 2: zonal variation in the heat budget

In this section, we consider the zonal variation of the heat budget, for two regions: (1) the open southern limb (SL) and (2) the interior circulation cell (IC). The open SL region (i.e., the magenta stippled area in Fig. 4) is defined by the stream function as  $16 < \Psi < 26$  Sv, which describes the open inflow zone where water masses enter the gyre, and spans the entire zonal extent of the double gyre system, from just west of Gunnerus Ridge ( $\sim 33^\circ$ E) to  $\sim 50^\circ$  W, where the streamlines veer northwards to follow the coastline of the Antarctic Peninsula. The southern boundary of the SL is the southernmost streamline that does not intersect with the coastline (16 Sv). This definition of the SL enables us to focus on the water that enters the gyre from the east, and circulates the entire zonal extent of the gyre, thus reaching into the south-western interior. The IC region (i.e., the blue stippled area in Fig. 4) is defined as  $\Psi \geq 26$  Sv, which is the largest streamline that spans the entire zonal extent of the double gyre system, this time forming a fully enclosed circuit. This definition of the IC allows us to focus on the recirculating waters of the gyre, from just west of Gunnerus Ridge to near the continental shelf edge of the northern tip of the Antarctic Peninsula ( $\sim 50^\circ$  W). For both regions, the area east of Maud Rise ( $3^\circ$  E) is omitted, due to large uncertainties east of Maud Rise (discussed in Section 5). In this section, we provide an analysis of the zonal variation of the heat budget in the Weddell Gyre of the two regions described at the beginning of Section 4 (highlighted in Figs. 2b and d). This analysis is carried out as follows: we integrate compute the mean meridionally for each zonal band in our data grids, from just west of Gunnerus Ridge Maud Rise ( $\sim 33^\circ$  E) in the westward direction towards the Antarctic Peninsula (Figs. 35-46, upper panel). The latitudinal range of the band is defined using the stream function, focusing on the SL and the IC. Lastly, we plot the cumulative zonal integration from east to west (Figs. 35-46, lower panel) and provide the zonally integrated heat budget terms in Table 3. We also take the sum of the heat budget terms and divide by the time period to get the temperature change for SL and IC, also listed in Table 3. For both regions, the zonal variation in the heat budget terms (Fig. 35-46, upper panel) show large local imbalances in the overall heat budget, which are physically implausible. However, much of the local (grid-scale) imbalances (i.e., the random noise part) cancels out useful information is provided in the net (zonally integrated) heat budget terms, allowing regional patterns not affected by the differentiation at the grid scale to emerge (Fig. 35-64, lower panel). The resulting volume integration describes the heat flux divergence (negative) or convergence (positive), owing to heat fluxes across the boundaries of the volume of water in question. By considering the 4 different heat budget terms with respect to each other, we can build up a picture of how and where heat is redistributed throughout the Weddell Gyre. Table 2 provides a list of the abbreviations for the terms presented in Figs. 35-810. The method for computing the associated errors is detailed in Section 3.2 as well as the Supplement (S7). We also provide a zonal and meridional analysis of the entire region marked by both blue and magenta stippling in Fig. 4 in the supplements (Figs. S7 and S8). These analyses provide results that agree with the analyses presented in this section and are described in section S8.



**Figure 4: The stippled areas show the regions defined as SL (magenta crosses) and IC (blue dots) respectively, encased in the streamlines used to define the SL and IC regions ( $16 < SL < 26$  Sv and  $IC > 26$  Sv). These regions are horizontally integrated across in the following Section 4.2. Grey and black contours provide the horizontal streamlines and the 1000, 2000 and 3000 m isobaths respectively (as in Fig. 1).**

### Region 1: SL (southern limb)

The SL region (Fig. 2b) spans the entire zonal extent of the double-gyre system, with the northern boundary being the largest streamline that spans both sub-gyres, from just west of Gunnerus Ridge to  $\sim 45^\circ$  W (30 Sv), and the southern boundary being the southernmost streamline that does not break into the coastline (16 Sv). This enables us to focus on the water that circulates the entire zonal extent of the gyre, thus reaching into the south-western interior.

Overall, the primary sources of heat (heat flux convergence) are from the advection terms, mean horizontal advection ( $A_H$ ) and mean vertical advection ( $A_V$ ), whereas the primary heat sinks (heat flux divergence) are from the turbulent diffusion terms, horizontal turbulent diffusion ( $D_H$ ) and vertical turbulent diffusion ( $D_V$ ) (Fig. 35, lower panel). The vertical terms are spatially invariant and contribute to the heat budget to a much lesser extent than the horizontal terms. This complements the findings in Fig. S2 of the net air-sea heat fluxes, which shows a relatively small heat loss through the surface of the ocean in the SL region. The heat flux convergence due to vertical advection is driven by upwelling, which leads to an advection of warm water upwards into the layer, whereas heat flux divergence due to vertical turbulent diffusion removes heat from the layer through the upper boundary (the mid-thermocline). Thus, heat enters the SL primarily through horizontal advection, and to a lesser extent by vertical advection due to upwelling, and is removed from the layer primarily by horizontal turbulent diffusion and vertical turbulent diffusion (a small amount of which may escape through the surface of the ocean). Thus, the turbulent diffusion terms redistribute the heat that is advected into the gyre layer.

The largest regional fluctuations (Fig. 35, upper panel) in the sum of the terms are caused by large fluctuations in mean horizontal advection, with the largest fluctuation occurring downstream of Maud Rise, from about  $-20$  at  $0^\circ$  E to  $+20$   $\text{Wm}^{-2}$  at  $\sim 5^\circ$  W, in particular upstream of Maud Rise (i.e., mean horizontal advection increases by  $\sim 55$   $\text{Wm}^{-2}$  just east of the Maud Rise seamount, from  $10^\circ$  to  $7^\circ$  E, and then sharply decreases by  $\sim 80$   $\text{Wm}^{-2}$  from  $7^\circ$  E to the Prime Meridian). There is also an initial removal of  $\sim 2$  TW of heat from the SL by mean horizontal advection east of  $0^\circ$  E, possibly related to the influence of Maud

rise. Overall, there is a net heat gain due to mean horizontal advection, of  $\sim 8$  TW, although taking into account the initial heat sink, the amount advected into the layer west of  $0^\circ\text{W}$  is  $\sim 10$  TW -effect of Maud Rise is also visible in the net contribution (lower panel of Fig. 35), where it results in an increase in mean horizontal advection of  $\sim 7$  TW from  $\sim 10^\circ$  to  $\sim 5^\circ$  E, after which a small decrease of  $\sim 2$  TW occurs from  $5^\circ$  E to the Prime Meridian. This influences the net heat tendency (sum of the heat budget terms), which, due to a combination of both the horizontal terms, initially shows a heat sink of  $\sim 4$  TW between  $3^\circ$  E and  $30^\circ$  W, followed by a gradual build-up of heat west of  $30^\circ$  W, so that the overall net heat flux convergence due to the sum of the heat budget terms for the SL is  $0.3 \pm 3$  TW, increases from 0 to  $\sim 7$  TW, and then decreases to  $\sim 4$  TW over the same longitudinal range. The perturbation in mean horizontal advection at Maud Rise is an artefact of being unable to resolve the regional circulation around Maud Rise, whereas the gridded temperature captures the presence of the relatively cold Taylor Column-Cap situated over Maud Rise (clearly visible in Fig. 1). This creates strong horizontal temperature gradients on the eastern and western flanks of Maud Rise, i.e., in the direction of the dominant flow. This is discussed further in Section 5.2.2.

## Region 2: IC (Interior Circulation cell)

The IC region (Fig. 2d) spans the entire zonal extent of the double gyre system, this time forming a fully enclosed circuit to focus on the recirculating waters of the gyre, from just west of Gunnerus Ridge to near the continental shelf edge of the northern tip of the Antarctic Peninsula ( $\sim 50^\circ$  W).

Overall, the magnitudes means of the heat budget terms (Fig. 6, upper panel) are similar in the IC to the SL, with the exception of the mean horizontal advection term, which has a larger magnitude mean in the SL ( $10 \pm 15$   $6 \pm 13$   $\text{Wm}^{-2}$ ) than in the IC ( $-2 \pm 24$   $0.2 \pm 11$   $\text{Wm}^{-2}$ ) (Table 3). Note the differences in the magnitudes of the net heat budget terms are due to the larger area of the IC (Table 3). As with the SL, vertical terms in the IC remain zonally spatially uniform and are have much smaller in total magnitude means than the horizontal terms, with heat being vertically advected into the IC domain (via upwelling) and removed from the IC domain by vertical turbulent diffusion. In contrast, the horizontal terms in IC change sign in comparison to the SL. Horizontal turbulent diffusion ( $D_H$ ) is the primary heat source (i.e., heat flux convergence), whereas horizontal mean advection ( $A_H$ ) is the primary heat sink (i.e., heat flux divergence) (Fig. 46, lower panel). Additionally, in contrast to the SL, the surface heat fluxes in this region (Fig. S2) are positive, implying a positive heat flux into the ocean through the surface. This heat does not appear to cross the thermocline (i.e., the upper boundary of our domain), given the convergence of heat due to vertical advection and the positive vertical velocity values in Fig. S3 imply that, in the vertical, heat is entering the IC domain from below (due to upwelling), and vertical turbulent diffusion provides a heat sink (removal of heat), likely upwards through the thermocline, owing to the strong vertical temperature gradients associated with the thermocline. Thus, the heat entering through the ocean surface is likely to be redistributed by other processes (such as surface transports or ice melt).

The largest regional fluctuations within the IC (Fig. 64, upper panel) in the sum of the terms are caused by two large peaks in mean horizontal advection, east just west of Maud Rise, where mean horizontal advection fluctuates from  $-65$   $7$   $\text{Wm}^{-2}$  to  $50$   $20$   $\text{Wm}^{-2}$  and then to  $6$   $-25$   $\text{Wm}^{-2}$  over  $\sim 1400$  km (from  $28$   $-3$   $^\circ\text{E}$  to  $12$   $-0$  to  $7$   $-5$   $^\circ\text{E}$ ). We assume these peaks are somewhat an artifact of the mean circulation interacting with strong lateral temperature gradients in the eastern sub gyre region (i.e., the red and blue ellipses in Fig. 2e over the eastern sub gyre coincide with strong temperature gradients in Fig. 1 and S5b), which is further discussed in Section 5.2.2.

**Table 2: Explanations of the abbreviations used in Figs. 3-8.**

Term	Description
$A_H$	Mean horizontal geostrophic heat advection
$A_V$	Mean vertical heat advection
$D_H$	Horizontal turbulent diffusion

$$\overline{\sum AD}$$

The sum of the heat budget terms in Eq. 1, listed above, where A stands for the horizontal and vertical advection terms and D stands for the horizontal and vertical diffusion terms.

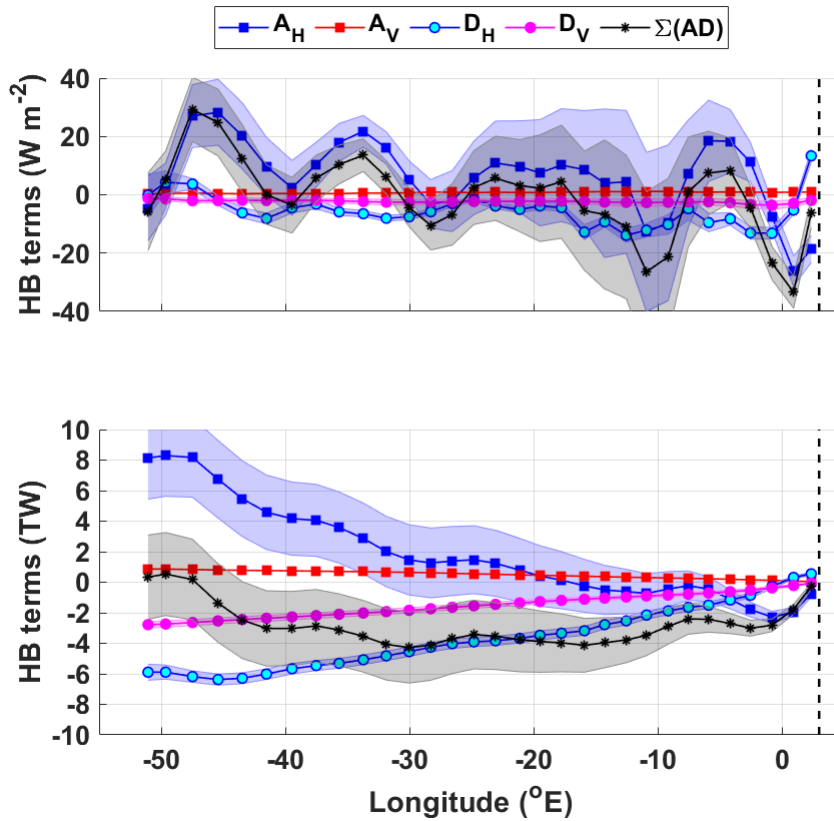


Figure 53: upper panel: the heat budget terms for the open southern limb (SL) of the gyre in  $\text{Wm}^{-2}$ ; lower panel: the cumulative heat budget terms in Terawatts (TW). The key for the legend is listed in Table 2. The dashed vertical line marks the approximate longitude of Maud Rise, at  $3^\circ \text{ E}$ . The shaded errors provide the associated propagated error (detailed in section 3.2 and the supplement). The SL region is marked by magenta stippling in Fig. 4.



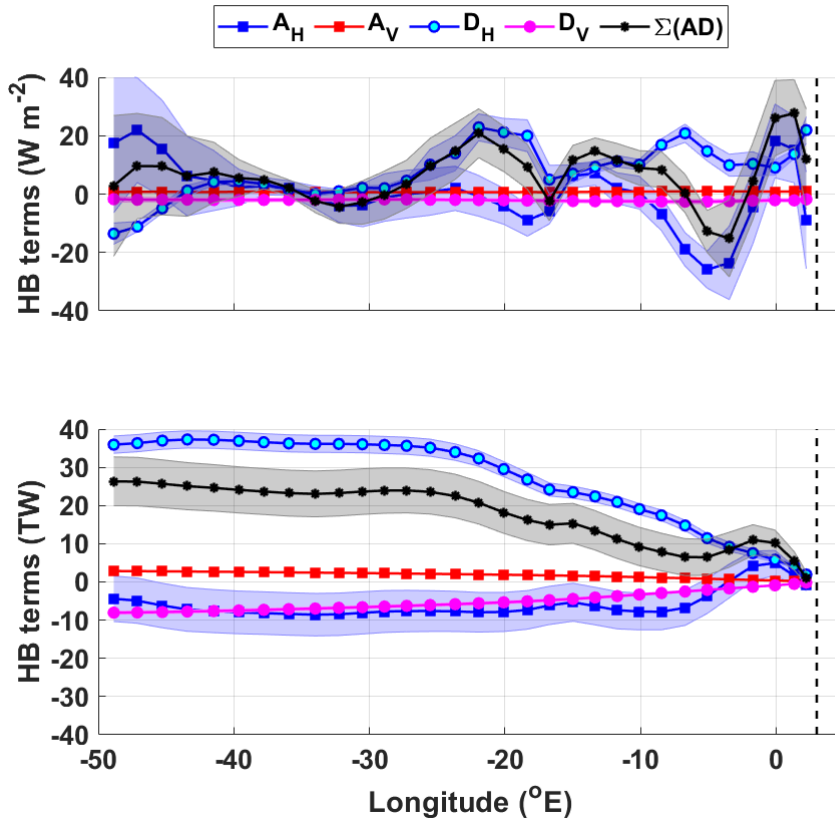


Figure 46: As in Fig. 35, but for the interior circulation cell (IC) of the Weddell Gyre. The IC region is marked by blue stippling in Fig. 4.

390

Table 3: Zonal mean ( $\text{W m}^{-2}$ ) and net (TW) heat budget contribution of the different terms in Eq. 1.2 for the SL and IC zones of the Weddell Gyre (Fig. 35 and 46 respectively), and an estimate of temperature change ( $dT/dt$ ), using a time period of 14 years. The uncertainty provided for the net heat budget terms (in TW) are the sum in quadrature of the propagated errors. For the mean heat budget terms (in  $\text{W m}^{-2}$ ), the provided uncertainty is the standard deviation of the zonal mean heat budget term, and the mean error (in  $\text{W m}^{-2}$ ) is the mean of the propagated error and the standard deviation of the mean propagated error. See Section 5.2 for further information.

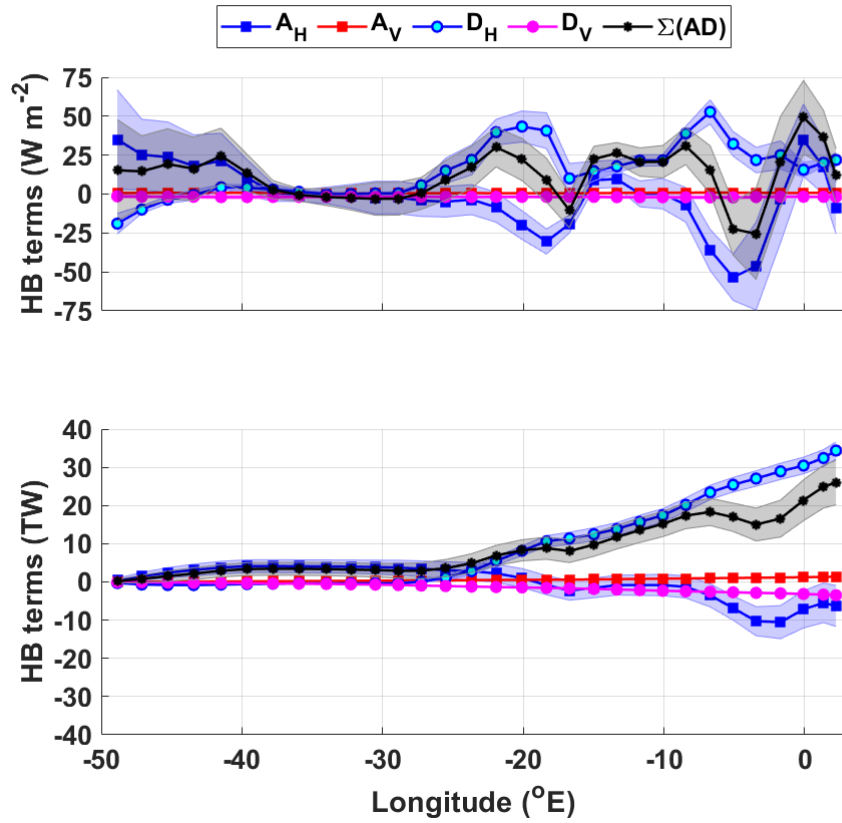
Heat Budget Term	SL			IC		
	Mean ( $\text{W m}^{-2}$ )	Mean Error ( $\text{W m}^{-2}$ )	Net (TW)	Mean ( $\text{W m}^{-2}$ )	Mean Error ( $\text{W m}^{-2}$ )	Net (TW)
Mean Air-Sea flux	$-1.8_{-2.5}$		$-4.6_{-3}$	+8		$+60_{+27}$
Mean horizontal advection	$+10_{\pm 1.5} +6_{\pm 1.3}$	$10_{\pm 5} 13_{\pm 7}$	$+26_{\pm 4} +8_{\pm 3}$	$-2_{\pm 2.4} +0.2_{\pm 1.1}$	$17_{\pm 2.2} 9_{\pm 5}$	$-37_{\pm 2.7} 4_{\pm 6}$
Mean vertical advection	$+0.9_{\pm 0.3} +0.7_{\pm 0.3}$	$0.1 \pm 0.04$	$+2.5_{\pm 0.03} +0.9_{\pm 0.02}$	$+0.8 \pm 0.2_{\pm 1}$	$0.05_{\pm 0.04} 0.04_{\pm 0.01}$	$+6.5_{\pm 0.1} +2.8_{\pm 0.03}$
Horizontal turbulent diffusion	$-6.7_{\pm 5.7} +8.6_{-5.5}$	$4_{\pm 2.2} \pm 1$	$-19.0_{\pm 2.6} \pm 0.5$	$+8.7_{\pm 9} +14_{\pm 7.8}$	$5_{\pm 6.3} \pm 2$	$+75_{\pm 9} +36_{\pm 2}$
Vertical turbulent diffusion	$-2.7_{-2.3} \pm 0.5$	$0.9_{\pm 1} \pm 0.2$	$-7.2_{\pm 0.4} 2.8_{\pm 0.2}$	$-2.3_{-2.1} \pm 0.3$	$0.6 \pm 0.1$	$-17_{\pm 0.6} 8_{\pm 0.4}$
Heat tendency (i.e., sum of the heat budget terms)	$+1.6_{\pm 1.4} +16_{-0.9}$	$11_{\pm 5} 13_{\pm 7}$	$+2.2_{\pm 5} +0.3_{\pm 3}$	$+5.5_{\pm 10} +26_{\pm 6.7}$	$17_{\pm 2.2} 9_{\pm 5}$	$+28_{\pm 2.9} +26_{\pm 6}$
Temperature tendency over 14 years ( $^{\circ}\text{C}$ )			$\pm 0.09_{\pm 0.03} \pm 0.2$ $^{\circ}\text{C}$ or $+0.006_{\pm 0.01}$ $0.002_{\pm 0.02} ^{\circ}\text{C/yr}$			$+0.4_{\pm 0.4} +0.8_{\pm 0.2}$ $^{\circ}\text{C}$ or $0.03_{\pm 0.03}$ $0.030_{\pm 0.01}$ $^{\circ}\text{C/yr}$

In order to learn more about how the heat budget terms act to redistribute heat throughout the IC, we further split the IC into IC-north (Fig. 57) and IC-south (Fig. 79), where the interface between the two regions is defined as the central gyre axis (i.e., the zonal maximum stream function, north of which the flow is predominantly eastward, south of which the flow is predominantly westward). ~~When integrating over these smaller regions, the heat budget does not close. We can, however, learn more about the processes which provide heat sources and sinks to these different regions. The first challenge that emerges is the non-objective decision making required regarding whether to include the easternmost part of the IC, where the horizontal mean advection may be unrealistic. We remove the eastern most values, east of 25°E in the IC north, and focus only on west of Maud Rise (inclusive of Maud Rise) for the IC south. We choose these limits in order to remove the anomalous ellipses that extend diagonally north eastwards from Maud Rise.~~ To further investigate horizontal diffusion, we computed horizontal turbulent diffusion heat fluxes across three zonal boundaries: (1) across the northern boundary of the gyre (defined as  $\Psi = 26$  Sv for the northern limb of the gyre, Fig. 4); ~~the northernmost streamline of the filled region in Fig. 2d i.e., this is~~ the streamline that, in Fig. 2e3c, marks the boundary between heat flux convergence within the Weddell Gyre (red) and heat flux divergence to the north (blue)); (2) across the interface between the divergence zone within the SL and the convergence zone in the IC; and (3) across the central gyre axis from IC-south into IC-north, ~~for west of Maud Rise~~. The zonal variation in heat flux for each boundary is provided in Fig. 5b-7b and 97b-c respectively, whereas the zonal integrations of the fluxes are provided in the Supplements (Fig. S6).

Horizontal turbulent diffusion is the dominating heat source to the IC-north, providing a net ~~78 ± 434 ± 2~~ TW of heat, whereas horizontal mean advection is a much weaker net heat sink, removing ~~68-6 ± 45-5~~ TW (Fig. 5a7a, lower panel). Half the horizontal turbulent diffusive flux of heat occurs along the northern boundary of the gyre (~~32-17 ± 1~~ TW; Fig. S6a), with particularly large southward heat fluxes ( $>200 \text{ Wm}^{-2}$ ) directly downstream of the South Sandwich Trench, ~~and in the region of the southwest Indian Ridge east of 10°E~~ (Fig. 57b), indicating that the rest of this heat flux occurs in the easternmost part of the IC-north ~~(the heat transfer occurring from the IC-south across the central gyre axis into the IC-north is  $0.5 \pm 0.4$  TW, and so can be ruled out; Fig. S6c)~~. The large fluxes are most likely due to the strong meridional temperature gradients characteristic of the boundary between the warmer ACC to the north and the colder Weddell Gyre to the south.

Overall, when looking at the zonal variation in the heat budget terms in Fig 5a-7a (upper panel), the positive peaks in horizontal turbulent diffusion are synchronous with the troughs in horizontal mean advection, implying that as heat is turbulently diffused across the boundary into the IC-north, horizontal advection “carries” this heat away along the eastward-flowing northern limb of the gyre. We can also demonstrate this by following the evolution of the sub-surface temperature maximum along a single streamline (in this case,  $\Psi=26$  Sv, which is the outer boundary of the IC, as shown in Fig. 2d4). ~~There is an overall decrease in temperature along the southern limb of the Weddell Gyre and -which shows-~~ an overall increase in temperature along the northern limb of the Weddell Gyre, ~~as well as an overall decrease in temperature along the southern limb of the Weddell Gyre~~ (Fig. 68).

(a)



(b)

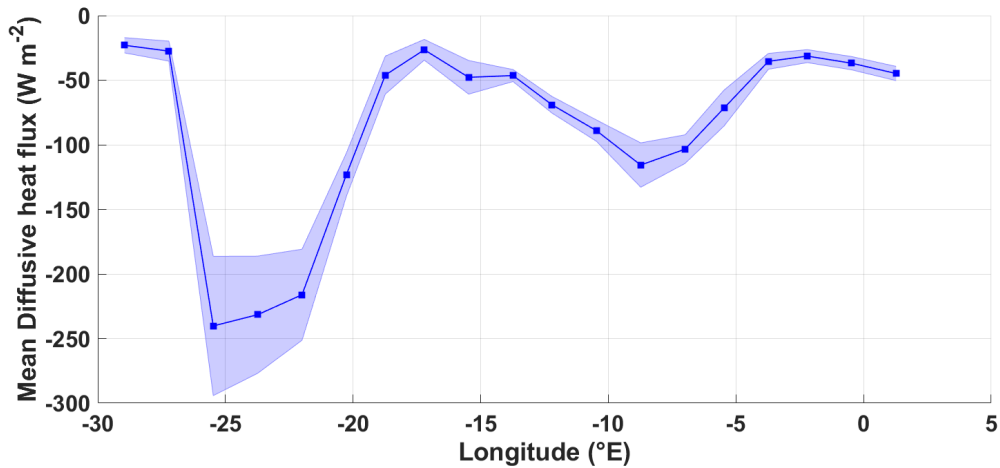
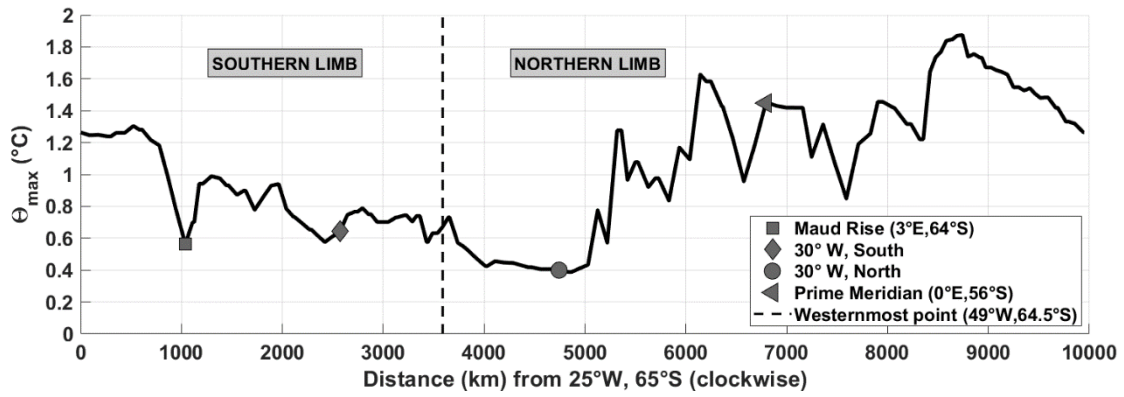


Figure 75: (a) Heat budget terms for the IC-north, west of 25°E: upper panel: zonal means in  $W m^{-2}$ ; lower panel: the cumulative heat budget terms from west to east in Terawatts (TW). The key for the legend is listed in Table 2. The dashed vertical line marks the approximate longitude of Maud Rise, at 3° E. Panel b shows the zonal variation of the diffusive horizontal heat flux across the northern boundary of the northern limb of the Weddell Gyre, in  $W m^{-2}$ , defined by the streamline that equals 25-26 Sv. Negative values indicate a southward flux of heat into the eastward-flowing northern limb of the Weddell Gyre from north of the northern Weddell Gyre boundary (the subsequent cumulative horizontal diffusive heat flux across the northern boundary is provided in the Supplements in Fig. S6a).



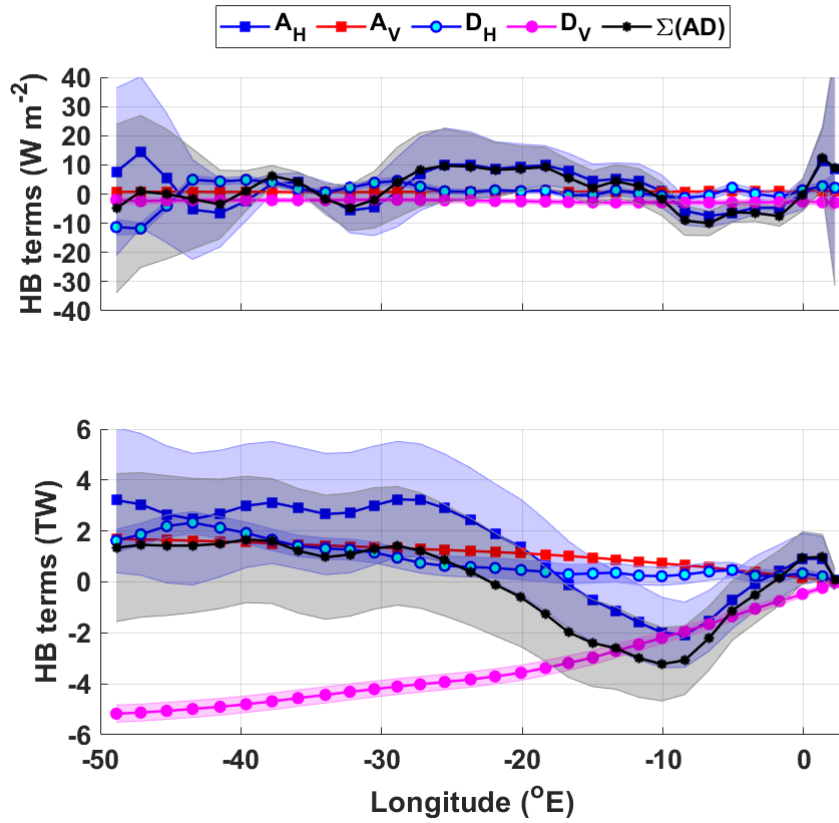
**Figure 68:**  $\theta_{\max}$  ( $^{\circ}\text{C}$ ) along the streamline  $\Psi = 26$  Sv (i.e., the outermost boundary of the IC). The distance in km is along the streamline in a clockwise direction from  $25^{\circ}\text{W}$ ,  $65^{\circ}\text{S}$ , with key locations marked using the legend.

425 Horizontal mean advection provides a heat sink of  $3 \pm 3$  TW in the IC-South directly west of the Prime Meridian, up until  $\sim 10^{\circ}\text{W}$  (Fig. 79a), after which it provides an overall heat source of  $\sim 5$  TW, so that the net heat flux convergence due to mean horizontal advection is  $(3.27 \pm 32.8)$  TW. The initial heat sink could be indicative of heat advection following the mean circulation about the eastern sub-gyre, as indicated by the streamlines in Fig. 23; west of this, heat is advected into the area that circulates the western sub-gyre. In contrast to the SL, horizontal turbulent diffusion is also a source of heat to this area

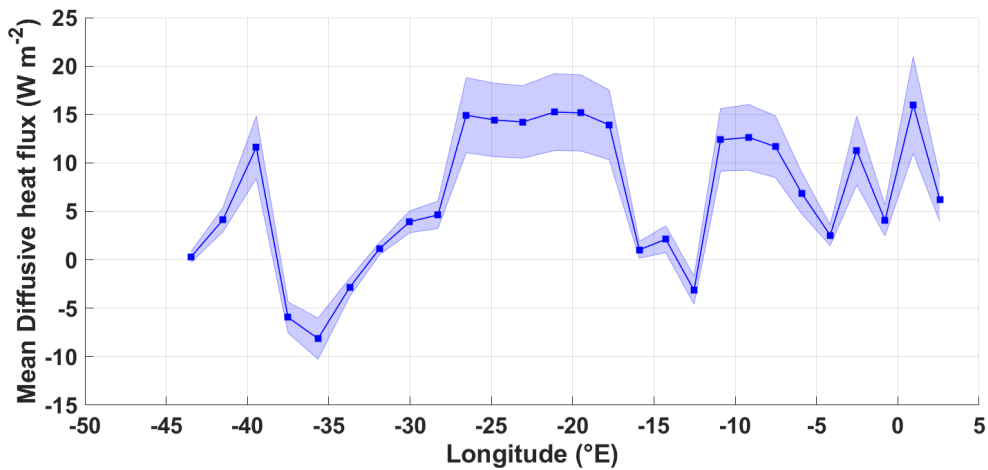
430  $(3.31.6 \pm 1.0.4)$  TW, which is indicative of a northward diffusive heat flux, which removes heat from the SL, transporting it northwards into the IC-south. This is indeed the case when computing the diffusive heat flux across the northern boundary of the SL in Fig. 7b9b, which gives a net heat flux of  $6.41.1 \pm 0.6.1$  TW, of which  $3.4 \pm 0.6$  TW occurs downstream of Maud Rise (Fig. S6b). Additionally, a net diffusive heat flux of  $0.8.5 \pm 0.1.4$  TW occurs northwards from the IC-south across the central gyre axis into the IC-north, mainly occurring  $10$ - $28^{\circ}\text{W}$ , i.e., at the interface between the eastern and western sub-gyres (Fig. 7e.9c & S6c). The diffusive flux of heat northwards from the SL into the IC occurs at a mean rate of  $22 \pm 12$   $\text{W m}^{-2}$  (Fig. 7b), with a maximum at  $10^{\circ}\text{E}$ , of  $\sim 50 \pm 25$   $\text{W m}^{-2}$ , where the streamline along which we are integrating starts to curve northwards to meander around the northern flank of Maud Rise. The streamlines indicate a strong flow in this region, with a sharp meridional gradient between the colder recirculated water close to the central gyre axis, and the relatively warm advected inflow of the southern limb. The mainnet heat sink in the IC-south is the vertical turbulent diffusion term (Fig. 7a9a), which

440 removes  $6.5 \pm 0.3$  TW upwards through the thermocline.

(a)



(b)



(c)

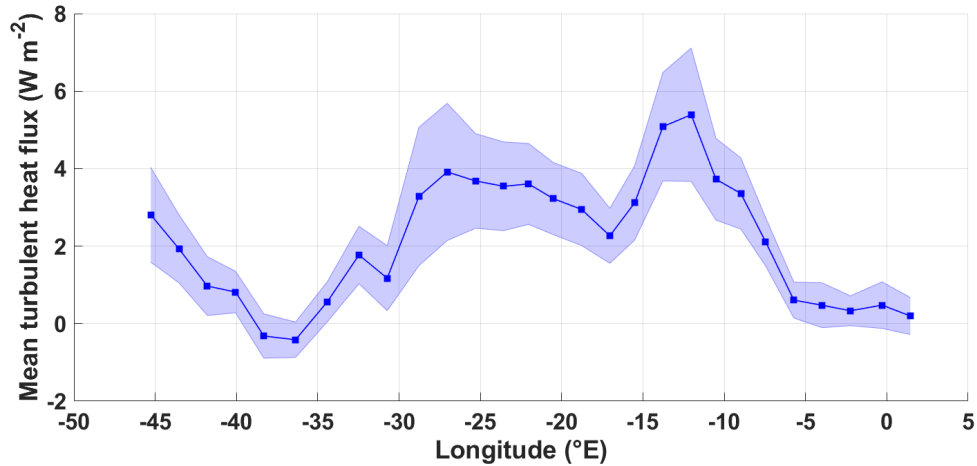


Figure 97: (a) Heat budget terms for the IC-south, west from  $-3^{\circ}\text{E}$  (Maud Rise): upper panel: the zonal mean heat budget terms in  $\text{W m}^{-2}$ ; lower panel: the cumulative heat budget terms in Terawatts (TW). The key for the legend is listed in Table 2. The dashed vertical line marks the approximate longitude of Maud Rise, at  $3^{\circ}\text{E}$ . Panels b and c show the zonal variation of the diffusive horizontal heat flux in  $\text{W m}^{-2}$  for (b) across the boundary between the SL and the IC, and (c) across the central gyre axis from the

IC-south to the IC-north, ~~west of Maud Rise~~. Positive values indicate a northward flux of heat from the SL into the IC, and from the IC-south northwards into the IC-north, respectively (the subsequent cumulative horizontal diffusive heat flux across the northern boundary is provided in the Supplements in Fig. S6b-c).

## 5 Discussion

In this study, velocity and temperature derived from Argo floats were used to determine the heat budget of the Weddell Gyre, for a 1000 m thick layer of water extending from the mid-point of the thermocline, encompassing the core of the WDW layer. In the following section, we provide an in-depth discussion of the study limitations, before ultimately discussing and interpreting the results in Section 5.2.

### 5.1 Study limitations

Before interpreting the results, several limitations of the study require discussion.

#### 5.1.1 Vertical boundary limits

The first limitation relates to the omission of the upper 50 dbar from the profiles prior to mapping (Supplements S1), to avoid the highly seasonally varying surface waters. Ideally, we would include the upper ocean layer, in order to explicitly apply the net air-sea heat flux term ( $Q_{net}$ ) as a surface boundary condition in the heat budget equation. Since we are interested in the re-distribution of heat throughout the Weddell Gyre (with the main heat source being the WDW), we applied boundary conditions which ensure that the Winter Water layer (defined by its temperature-salinity minimum) is not included within the subsurface layer of interest, while the core of WDW (i.e.,  $\Theta_{max}$ ) is fully included. The vertical boundary conditions allow us to consider the varying depths of the upper boundary, while also avoiding bias by fixing the thickness of the vertical layer. However, this, along with the non-uniform resolution of the vertical grid, may introduce some noise into the analysis from grid cell to grid cell, owing to the different depths of the water column the heat budget is integrated over. This may partly explain why there is considerable noise in the maps in Section 4.1, while the noise partially cancels out in Figs. 3-5 and 46. The implication is that our heat budget analysis is reasonable on large-scales, but would introduce considerable noise when assessed on local-scales. We return to this when discussing regional imbalances in Section 5.2.2, and find that the lateral gradients in the upper boundary depth are unlikely to be a major source of the noise in Fig. 2-3 (Fig. S5a).

#### 5.1.2 Estimating vertical velocity

Another limitation in the study is a suitable estimate for vertical velocity. Vertical Ekman pumping velocity is used to represent vertical velocity, which is held constant with depth (Supplements S3) throughout the 1000 m thick layer, thus assuming a perfect geostrophic (thus non-divergent) flow. However, reanalysis-based wind products from which  $w_E$  is computed, are also limited in their accuracy, largely due to lack of in-situ wind measurements in the Southern Ocean. Additionally, vertical velocities resulting from eddies, fronts and along steeply sloping bathymetry are not taken into account in this study, which might be important sources of vertically non-uniform vertical velocities in the upper ocean.

#### 5.1.3 Diffusivity

The work presented within this paper is based on a long-term mean gridded dataset derived from Argo floats between 2002 and 2016 (Section 2.2). Thus, we are unable to directly incorporate an eddy component into the analysis, which requires fluctuations from the mean temperature and velocity (i.e.,  $u'T'$ ). Unresolved eddying fluctuations are most likely to be an important factor when assessing the heat budget east of the Prime Meridian in the Weddell Gyre, a region we know to be dominated by a mesoscale eddy field (e.g., Schröder & Fahrbach, 1999; Leach et al., 2011; Ryan et al., 2016). We therefore exclude the region east of Maud Rise in the regional analysis in section 4.2. Meanwhile, within this study, we rely on the parameterization of eddy-influences through the estimation of horizontal turbulent diffusion, and the parameterization of vertical instabilities through the estimation of vertical turbulent diffusion. Horizontal and vertical diffusivity ( $\kappa_H$  and  $\kappa_V$  in Eq.

480 1) yield considerable uncertainty to this study, especially where they are assumed constant in space. Horizontal diffusivities  
have been shown to be length scale-dependent (Okubo, 1971; Ledwell et al., 1998). For example, Ledwell et al (1998)  
demonstrate the importance of length scales in the Eastern North Atlantic by releasing and following the dispersal of a patch  
of sulphur hexafluoride, estimating values of  $2 \text{ m}^2 \text{ s}^{-1}$  for length scales of 1-10 km, and as much as  $1000 \text{ m}^2 \text{ s}^{-1}$  for length scales  
of 30-300 km. For the entire Southern Ocean, Zika et al. (2009) estimate a value of  $300 \pm 150 \text{ m}^2 \text{ s}^{-1}$ . Within the Weddell Gyre,  
485 Leach et al. (2011) provide estimates of  $\kappa_H$  and  $\kappa_V$  derived from observations in the Maud Rise region, and estimate  $\kappa_H$  to be  
in the range of  $70\text{-}140 \text{ m}^2 \text{ s}^{-1}$ . Leach et al. (2011) consider their estimate for  $\kappa_H$  appropriate, albeit on the low side, for mesoscale  
eddies (although the station spacing of the observations in their analysis is 55 km; not the  $\sim 10$  km corresponding to the  
baroclinic Rossby radius in the region). For the Weddell Gyre, Donnelly et al. (2017) provide estimates of  $\kappa_H$  derived from  
observations (WOCE CTD stations from 37 hydrographic sections) on a larger scale than that of individual eddies, where  $\kappa_H$   
490 is estimated as  $247 \pm 63 \text{ m}^2 \text{ s}^{-1}$ . Cole et al. (2015) use salinity anomalies from Argo floats and velocity fluctuations from the  
ECCO2 product, which is an ocean state estimate that assimilates data (Menemenlis et al., 2005, 2008) to investigate horizontal  
mixing and the associated length scales. In Fig. 4 of Cole et al. (2015), at about  $50$  to  $60^\circ$  S in the Atlantic Ocean (the most  
southerly extent of the authors' analysis), the estimated values of  $\kappa_H$  range from  $200$  to  $2000 \text{ m}^2 \text{ s}^{-1}$  in the depth range of  $0$  to  
 $2000$  m. In Sevellec et al., (2022), the authors provide a useful dataset of  $\kappa_H$  at  $1000$  m, derived using Argo float trajectories:  
495 this is only possibly when a sequence of positions are available, from which the authors derived a "pseudo-trajectory". Thus,  
most of the ~~seasonally sea-ice covered~~ Weddell Gyre resembles a data gap, with the exception of the Prime Meridian and the  
north and eastern peripheries of the gyre. These values vary between  $0$  and  $8000 \text{ m}^2 \text{ s}^{-1}$ ; and are close to zero around Maud  
Rise, and can be  $4000\text{-}8000 \text{ m}^2 \text{ s}^{-1}$  within the eastern and northern gyre periphery. In this study, we are focused on the large-  
scale, and have a grid of data representative of the long-term mean from 2002-2016. Our grid cell resolution varies slightly  
500 with changing latitude (Reeve et al., 2019), but is on the order of  $\sim 80 \times 60$  km, which is nearly double the station spacing of  
 $55$  km in Leach et al. (2011). The length-scales applied in the objective mapping are  $800$  km in stage 1 and  $400$  km in stage 2,  
with a skew on the shape of the radius of influence based on the range in potential vorticity (See Section 2.2 in Reeve et al.,  
2019), and therefore the length scales upon which the analysis falls would actually be considerably greater than in Leach et al.  
(2011), but with similarities to the length-scale assumed in Cole et al. (2015) of  $300$  km. We make the decision to use the  
Sevellec et al. (2022) dataset where numerous data are available, and, for within the Weddell Gyre, we ~~use the values provided~~  
505 ~~by arbitrarily provide a baseline range of potential values for horizontal diffusivity, in the range of that provided by~~ Donnelly  
et al. (2017), ~~given, these are situated within the area of interest, namely the western Weddell Gyre~~ ~~Zika et al. (2009) and more~~  
~~specifically Cole et al. (2015), of  $400 \pm 200 \text{ m}^2 \text{ s}^{-1}$  (the ranges are incorporated into the error propagation computation)~~. A  
detailed explanation of our method and reasoning regarding the Sevellec et al (2022) dataset is provided in the Supplements  
510 (S4).

Referring now to the vertical diffusivities, Leach et al. (2011) provide particularly small estimates for  $\kappa_V$ , of  $3 \times 10^{-6} \text{ m}^2 \text{ s}^{-1}$  for  
the core of WDW, which they suggest is due to the timing of the survey – having taken place late spring/early summer, when  
sea-ice had just melted, and the wind had not yet had time to stir up the water column. Other studies provide larger estimates  
for  $\kappa_V$ , for various regions throughout the global ocean. Donnelly et al. (2017) estimate  $\kappa_V$  as  $2.4 \pm 2.8 \times 10^{-5} \text{ m}^2 \text{ s}^{-1}$ . Over rough  
515 topography, such as  $500$  m above abyssal sea mounts on the flanks of the mid-Atlantic Ridge in the Brazil basin, Ledwell et  
al. (2000) provide an estimate for  $\kappa_V$  of  $3 \times 10^{-4} \text{ m}^2 \text{ s}^{-1}$ . Over flat bathymetry of the ocean's abyss, however, Polzin et al. (1997)  
and Ledwell et al. (1998) estimate  $\kappa_V$  to be about  $1 \times 10^{-5} \text{ m}^2 \text{ s}^{-1}$ . Naveira Garabato et al. (2004a, 2004b, 2007) provide larger  
numbers for deep water in the Scotia Seas of  $3 \times 10^{-4}$  to  $1 \times 10^{-2} \text{ m}^2 \text{ s}^{-1}$ , which they attribute to breaking internal waves. Also  
within the Southern Ocean, Cisewski et al. (2005, 2008) acquired  $7 \times 10^{-4} \text{ m}^2 \text{ s}^{-1}$  in the upper pycnocline of the Antarctic  
520 Circumpolar Current at  $20^\circ$  E and Forryan et al. (2013) obtained a value of  $1.9 \times 10^{-5} \text{ m}^2 \text{ s}^{-1}$  based on observations in close  
proximity to a vigorous frontal system between  $60$  and  $80^\circ$  E, at the northern edge of the Kerguelen Plateau.

Initially, we attempted to estimate  $\kappa_v$  through the use of the Richardson Number, following Forryan et al. (2013), defined as the ratio of buoyancy frequency (N) squared to vertical shear squared, and often used to describe the stability of stratified shear flow. For large Richardson Numbers, the resulting diffusivity tends towards a background diffusivity coefficient provided by the authors. In the Weddell Gyre, the circulation is mostly barotropic, and so the vertical shear is small, leading to large Richardson Numbers, especially when one is focused on the long-term mean. Thus, with this method, the resulting diffusivity tends back towards a background diffusivity parameter. Hence, based on the range of values available from the literature, we made the decision to again use the values provided by Donnelly et al. (2017)-a range of potential values for vertical diffusivity, of  $(2.6-39 \pm 2.483) \times 10^{-5} \text{ m}^2\text{s}^{-1}$  (the “uncertainty” is incorporated into the error propagation computation).

A final consideration regarding uncertainty is related to using a simple differencing scheme at the grid cell resolution in order to compute spatial gradients in the heat budget equation (Eq. 1.2). Alternatively, we could have extracted curves from the grid from which to determine the gradients. However, this is dependent on further length scales, which could introduce additional uncertainty, and may also result in additional smoothing of the observed fields. Thus, we applied the former method, which extracts the gradients directly from the gridded datasets at the grid cell resolution, with the expectation that this would result in large regional fluctuations owing to the coarse grid resolution.

## 5.2 The Weddell Gyre heat budget

### 5.2.1 Overall findings

While the heat budget does not close on regional scales, it does approximately close when integrating over the open southern limb west of Maud Rise (SL), but not over the interior west of Maud Rise (IC). Nonetheless large areas, thus, important and useful information can be provided from comparing the four resulting heat budget terms.

In most regions of the Weddell Gyre, for the core of WDW, the terms dominating the heat budget are mean horizontal geostrophic advection (Fig. 2a3a), and horizontal turbulent diffusion (Fig. 2e3c). Horizontal mean advection appears patchy, though generally implying warming (convergence) in the southern limb of the gyre (Fig. 2a3a). Horizontal turbulent diffusion (Fig. 2e3c) acts as a heat source (heat flux convergence) within the eastward flowing northern limb of the gyre and to a lesser extent the gyre interior. In contrast, horizontal turbulent diffusion is associated with a divergent heat flux (or loss of heat) in the open southern limb of the gyre, as well as directly north of the gyre (downstream of the South Sandwich Trench). For the most part, mean vertical advection and vertical turbulent diffusion are spatially relatively uniform, where upwelling advects heat upwards into the WDW layer, and vertical turbulent diffusion removes heat through the top of that layer, upwards through the thermocline to the overlying Winter Water (WW). Additionally, the sum of the heat budget terms in Fig. 2e-3c imply a general cooling no clear signal in the SL (indeed the heat budget nearly closes when integrating over the region, providing a net gain of 0.3 TW), and a general warming in the IC, synchronous with the surface heat fluxes in Fig. S2, where the IC experiences a net heat flux into the ocean from the atmosphere, while in the SL the ocean loses heat to the atmosphere.

The spatial patterns in the heat budget terms become distinct when integrating over large areas, namely, the SL and IC (Section 4.2). While the vertical terms are spatially uniform throughout, providing a heat source (mean vertical advection) and a heat sink (vertical turbulent diffusion) throughout both regions, the horizontal terms “switch roles” in the SL versus the IC. Horizontal mean advection is a mechanism which brings heat into the SL, while horizontal turbulent diffusion removes much of that heat from the SL. In contrast, horizontal turbulent diffusion brings heat into the IC, whereas mean horizontal advection provides a heat sink in the IC. Thus, while both mean horizontal and vertical advection are responsible for bringing heat into the core of the WDW layer of the Weddell Gyre, turbulent diffusion is the mechanism which then redistributes that heat throughout the gyre interior. Surface heat fluxes (Fig. S2) also imply that the upward diffusive heat flux through the top of the



layer (i.e., the mid-thermocline) may represent a source of heat for the observed ~~negative air to sea heat flux~~ radiative heat loss in the SL. In the IC, where air-to-sea heat fluxes are positive, the upward diffusive heat flux across the upper boundary of the WDW layer may be horizontally redistributed by the upper ocean flow field and additionally may provide heat required to melt sea ice in this area.

### 5.2.2 Regional (im-)balances

~~When integrating over smaller regions, the~~ The heat budget ~~shown in section 4~~ does not close, ~~especially when integrated over smaller areas~~ (see Figs. 2e3c, 3 and 4), indicating that noise is largely cancelled when integrating over larger areas. The noise is likely a result of (1) discrepancies in the depth range from grid cell to grid cell, (2) the nature of differentiating across grid cells, (3) due the presence of mesoscale eddies unresolved by the data and methods used, and (4) aliased observations that result in a distortion of the mean state. Here, we discuss the influence of (1) unresolved mesoscale eddies and narrow currents in the eastern Weddell Gyre, (2) Maud Rise and (3) the open northern boundary, and discuss what might be missing that prevents the closure of our heat budget analysis.

***Unresolved mesoscale eddies and narrow currents*** - The presence of unresolved mesoscale eddies is particularly important east of the Prime Meridian. In Fig. 1, it is clear there is a misalignment between the streamlines and temperature distribution in the eastern part of the gyre; horizontal mean advection hugely dominates this region which is partially compensated by horizontal turbulent diffusion (Fig. 23). Given the domed shape of the isopycnals characteristic of a cyclonic gyre (e.g., Strass et al., 2020; Fahrbach et al., 2011), we hypothesised that a larger bias due to the horizontal gradient of the upper boundary depth (i.e., mid-thermocline, Fig. S1b) was occurring at the gyre periphery (i.e., where the slopes of the isopycnals are largest), which may be contributing to the large ~~positive and negative ellipses-values that extend diagonally outwards from the centre of the eastern sub-gyre (i.e. forming roughly shaped ellipses)~~ in horizontal mean advection ~~at the eastern periphery of the gyre~~ in Fig. 2a3a. This appears not to be the case, given there is no clear large horizontal gradient in the mid-thermocline depth in the east; indeed, the largest horizontal gradient in mid-thermocline depth occurs in the very west of the Weddell Gyre, along the Antarctic Peninsula, and to a lesser extent over Astrid Ridge (Fig. S5a). The large ellipses do, however, appear to be related to large horizontal temperature gradients (Fig. S5b). We found the largest horizontal temperature gradients unsurprisingly along the northern boundary of the Weddell Gyre (owing to the considerably warmer ACC to the north), but also in a diagonal line spanning from directly east of Maud Rise, northwards to the northeast of the eastern sub-gyre (~30°E, 55°S); in other words, synchronous with the north-easternmost ellipse of heat flux divergence and south-westernmost ellipse of heat flux convergence (Fig. 2a-3a & 2e3e). The strong horizontal temperature gradients therefore likely accounts for the two strongest ellipses in Fig. 23. This, along with the fact that the computation (which relies on differencing across grid cells) is sensitive to the alignment of the temperature field to the velocity field, which is imperfect, especially on such a coarse resolution grid. In particular, temperature is a more conservative (slow changing) variable in comparison to the velocity field. Thus, in capturing the mean state, our depiction of the mean velocity field may be especially distorted due to aliasing of observations, which we know are biased to summer conditions, and are particularly varying on shorter timescales in the dynamic region east of the Prime Meridian. ~~The summertime bias and coarse resolution likely results in an underestimate of the mean gyre strength, and thus the velocity field derived from the stream function (Neme et al., 2021).~~ In addition to the aliasing of the observations in representing the mean state, it is ~~possible-likely~~ that there is a significant additional component (in addition to mean horizontal advection), ~~such as transient eddying processes~~, which distributes heat, ~~particularly~~ in the eastern part of the gyre. The question is, how much of this is a result of unresolved mesoscale activity driving large temperature and velocity gradients on much shorter length scales than can be appreciated here, resulting in thus aliased observations, and whether the additional component is a real measurable yet unobserved (in this study) component, or an artefact of the method and data sources used in this study. ~~This is a key consideration, which means that the sum of the 4 heat budget terms, especially in the east, cannot be viewed as the heat tendency (which should be zero in a closed system), but that it additionally consists of unresolved processes (i.e., “R”~~

605 in Eq. 1.1 & 1.2). These include mesoscale eddy activity that is not fully represented by parameterization via turbulent diffusion. We know that the eastern part of the eastern sub-gyre is dominated by an intense mesoscale eddy field (Ryan et al., 2016, Leach et al., 2011 and Gordon and Huber 1984). Wilson et al. (2022) show, using idealised models, that transient eddies are responsible for most of the southward heat transport in the eastern limb of the Weddell Gyre. In addition, as discussed in section 3, a process not accounted for is regional-scale variations in the temperature and flow fields. This process might be particularly important in the eastern sub-gyre, where the boundary to the gyre is poorly-defined due to the openness of the topography. Indeed, Schröder and Fahrbach (1999) suggest that there is no continuous current marking the eastern boundary, and that baroclinic shear instabilities lead to a breakdown of the eastward-flowing current in the northern limb of the gyre, and that the current “reforms” in the westward-flowing southern limb. The geometry of the eastern sub-gyre might therefore be sensitive to interannual to decadal variations in the wind forcing, potentially affecting the time mean heat flux convergence in this area. ~~This~~ These findings in Schröder and Fahrbach (1999) also aligns with the findings of Sonnewald et al. (H review2023), who use machine learning in a climate model to diagnose the dominating dynamic regimes in the Southern Ocean, leading them to propose a circumpolar “super-gyre” which connects the Weddell and Ross sub-gyre systems. In the far-eastern sub-gyre region, recirculated “cold-regime” WDW (modified primarily through heat loss) comes into contact with incoming “warm-regime” WDW (Gordon and Huber, 1984). The “warm-regime” WDW represents relatively warm WDW advected into the gyre at the eastern inflow zone at about 30° E, driven by mesoscale eddies (Deacon 1979; Orsi et al. 1993; Orsi et al. 1995; Gouretski and Danilov 1993, 1994; Ryan et al., 2016). When comparing the two terms in Fig. 3a and 3c, while the magnitude is much larger in Fig. 3a (horizontal mean advection), horizontal turbulent diffusion displays the opposite signs and partially compensates in the eastern sub-gyre region in Fig. 3c. The terms do not cancel, and thus imply that the missing “R” term is significant in this region (although, large errors associated with mean horizontal advection also imply that mean advection is poorly represented in this region). This is not easily remedied, since the eastern Weddell Gyre is a region with poor data coverage, including from Argo floats, though at the time of writing, efforts are underway to close this key observational gap. It is therefore possible that horizontal turbulent diffusion, supposedly representative of turbulent dynamic processes within this study, is underestimated in this region. This is also implied when comparing the two terms in Fig. 2a and 2c: while the magnitude is much larger in Fig. 2a (horizontal mean advection), horizontal turbulent diffusion displays the opposite signs and partially compensates in the eastern sub-gyre region in Fig. 2c.

630 We’d like to acknowledge that our framework of inferring the heat budget is rather traditional, in which we parameterize the effects of eddies by means of horizontal diffusion. A more advanced approach is represented by the temporal-residual-mean framework (McDougall and McIntosh, 2001), in which the effects of eddies are decomposed into eddy-induced advection (adding an eddy-induced velocity to the time mean velocity in the advection term of the tracer equation) and eddy-induced diffusion. The latter can be decomposed into isopycnal and diapycnal diffusion (Groeskamp et al., 2016). We acknowledge this framework to be more physics-based than our classical approach, yet, given the limitation of our dataset, the estimation of the eddy-induced velocities is not straightforward. At the same time, Sevellec et al. (2019) demonstrated the usefulness of the temporal-residual-mean framework when applied to interpreting eddy-driven horizontal buoyancy transports from mooring-based observations acquired in Drake Passage. In particular, they highlight importance of eddy-driven horizontal transports in the direction perpendicular to the mean flow. For future work it would therefore be intriguing to demonstrate, whether the application of this framework to our data set would represent a major step forward towards closing the heat budget in the eddy-rich eastern part of the Weddell Gyre and around Maud Rise, where our approach does not lead to satisfactory results.

645 **Entrainment** - It could also be that there is a missing process, such as entrainment, that may account for the non-closure of the heat budget terms in Fig. 2e-3e (though entrainment would constitute a heat flux divergence which may help close the heat budget of the IC, but not of the SL). Indeed, Schlosser et al. (1987) used Helium-3 tracers within the north-western Weddell Sea to estimate a vertical diffusivity that is twice the value that we use here ( $5 \times 10^{-5} \text{ m}^2 \text{ s}^{-1}$ ) along with a rate of entrainment

of WDW into the overlying WW of 15-35 m yr<sup>-1</sup>. Behrendt et al. (2011) argue that increasing WW salinity during 1992-1996 is caused primarily by entrainment of WDW, and lists entrainment as one of three dominating causes of salinity changes to WW (the others being sea-ice formation related salt release and horizontal advection). Brown et al. (2015) also highlight the important role of entrainment in the carbon dynamics of the Weddell Gyre, by bringing dissolved inorganic carbon and salt upwards into the WW from the underlying WDW. Given this work is based entirely on observations, and that we lack sufficient observations to effectively resolve the complex dynamics of the eastern Weddell Gyre, the findings here illustrate the need for better observational coverage of the Southern Ocean in the high latitudes east of the Prime Meridian, a region thus far often omitted from observation campaigns.

Maud Rise - Maud Rise is a prominent feature in temperature (Fig. 1), mean horizontal advection (Fig. 2a3a), horizontal turbulent diffusion (Fig. 2e3c) and to a lesser extent, vertical diffusion (Fig. 2d3d). The effect of Maud Rise on WDW temperatures is due to the presence of a Taylor ~~column-cap~~ directly over Maud Rise, which has been previously observed as a ~~stagnant~~ column of relatively cold water surrounded by a warm halo on the flanks of Maud Rise (e.g., Muench et al., 2001; Leach et al., 2011). Regarding the heat budget terms shown in Fig. 3, mean horizontal advection results in a heat flux convergence upstream of Maud Rise, which is partially balanced by heat flux divergence due to horizontal turbulent diffusion. In contrast, downstream of Maud Rise, heat flux divergence occurs due to both mean horizontal advection and horizontal turbulent diffusion in Fig. 3 (although a convergence peak in horizontal turbulent diffusion does occur directly over Maud Rise).

The effect of mean horizontal heat advection on the flanks of Maud Rise is probably mainly an artefact of the velocity field. Since the water overlying Maud Rise is cold, and the velocity field does not adequately resolve the flow circulating the seamount (instead the streamlines in Fig. 1 cut directly East-West across Maud Rise), the heat convergence upstream of Maud Rise and divergence downstream of Maud Rise are caused by strong lateral temperature gradients between the cold water column overlying Maud Rise and the warm halo surrounding it. However, ~~the effects of mean horizontal advection upstream and downstream of Maud Rise partially cancel each other when integrating zonally in Fig. 3. Furthermore,~~ horizontal turbulent diffusion, which responds in a similar manner (since it is determined exclusively from observed horizontal temperature gradients, but not the velocity field, and therefore the response to Maud Rise is much smaller), acts to balance the influence of mean horizontal advection upstream of Maud Rise (Fig. 3c). Indeed, the overall heat loss that occurs downstream of Maud Rise in Fig. 3a is about 3040 W m<sup>-2</sup>, which is similar to previous estimates in Muench et al. (2001) and also in McPhee et al. (1999; 52 W m<sup>-2</sup> west of Maud Rise and 23 W m<sup>-2</sup> over Maud Rise), although these authors focus on the surface heat flux, both using surface drifting buoys.

~~In Fig. 3, regardless of the perturbation resulting from the influence of Maud Rise, there is a build-up of heat due to mean horizontal advection across the entire southern limb, even if the region over Maud Rise is removed from the integration. There is also a continual removal of heat by horizontal turbulent diffusion regardless of Maud Rise, which would be the case even if we increased the diffusivities on the flanks of Maud Rise, where deep convection has been previously observed (e.g., Akitomo, 2006), since the turbulent diffusion would cancel itself out when integrated zonally across Maud Rise (unless the diffusivities themselves were significantly different on each flank). Thus, it is important that we integrate over large areas and view these results as a representation of the processes influencing large-scale heat distribution of the Weddell Gyre (such as in Section 4.2), and avoid focusing on localised regions.~~ Ultimately, we show that heat flux divergence due to horizontal turbulent diffusion occurs on the flanks of Maud Rise, in agreement with previous studies, which show for instance that baroclinic instabilities on the flanks of Maud Rise are the source of recurrent eddies (Akitomo et al., 2006). Furthermore, Leach et al. (2011) and Ryan et al. (2016) suggest a mixing of WDW with modified recirculating WDW downstream of Maud Rise which would explain the initial heat sink between Maud Rise and the Prime Meridian features observed in Fig. 35a, ~~where the zonal variation in horizontal turbulent diffusion is smaller west of the Prime Meridian.~~

***The open northern boundary*** - Regarding the northern limb of the Weddell Gyre, our findings imply that horizontal turbulent diffusion is a mechanism by which heat enters the region (Fig. [2e3c](#)). This is in some agreement with Jullion et al. (2014), who use an inverse model based on ship-based sections along 30° E to the coast and also along the northern periphery of the gyre at about 55-60° S in order to diagnose the heat budget of the full water column. They suggest that most of the heat advected into the Weddell Gyre occurs along the northern gyre periphery, rather than from the eastern periphery, and reaches the southwestern Weddell Gyre through recirculation in the interior Weddell Gyre, leading to an entrainment of heat into the Antarctic Slope Front. This analysis is not able to resolve localised features such as the Antarctic Slope Front, but there is an indication, especially from the streamlines, that recirculation of the eastern sub-gyre plays a role in the distribution of heat in the Weddell Gyre (Fig. 1). Furthermore, in Jullion et al. (2014), eddy-induced transport contributes significantly to the heat budget of the Weddell Gyre, with a heat flux of  $5 \pm 1$  TW, out of a net heat flux of  $36 \pm 13$  TW from the ACC into the Weddell gyre, which is primarily due to mean circulation. The results provided in Figs. [23c](#), [5-7](#) & [9](#) show agreement with this finding; the Weddell Gyre's northern limb and the IC are dominated by a convergence of heat due to horizontal turbulent diffusion. Both the southern limb and the area directly north of the Weddell Gyre in the ACC, as well as east of  $\sim 30^\circ$ E, in contrast, exhibit heat flux divergence ( $\ominus$  cooling) due to horizontal turbulent diffusion (Fig. [2e3c](#)). This suggests that horizontal turbulent diffusion constitutes an important role in transporting heat into the Weddell Gyre along the open northern boundary as well as from the East- (although large uncertainty in the east requires some caution), in agreement with Jullion et al. (2014). The heat fluxes provided here are somewhat more than the eddy induced heat transport from Jullion et al. (2014) across both the northern and eastern boundary, which, however, accounts for the whole water column, across 2 hydrographic sections which represent the entire open boundary of the Weddell Gyre. We expect their eddy heat flux estimate to exhibit major uncertainties, as it is firstly based on a station spacing of the temperature and velocity profiles that are not nearly eddy resolving (especially along the northern boundary of the gyre) and secondly represents a one-time snapshot.

According to Tamsitt et al. (2016) and Naveira Garabato et al. (2011), major topographic features result in a divergence of horizontal and vertical eddy heat fluxes, leading to substantial warming in association with regions of enhanced mesoscale energy. Thompson & Sallée (2012) use particle advection experiments to show that the enhancement of eddy kinetic energy occurs downstream off topographic obstacles, which may explain cross-front exchange associated with jets in the lee of topographic features. This may explain the heat flux divergence due to horizontal turbulent diffusion that occurs in the lee of Maud Rise (Fig. [2-and-3c](#)). It may also help to explain the fluctuations along the northern limb of the Weddell Gyre in Figs. [2a-3a](#) and [23e](#), where the topography is complex, creating an open-ocean northern boundary to the Weddell Gyre. Indeed, the alternating bands of convergence and divergence along the northern limb of the gyre between 30° W and 20° E in Figs. [2a-3a](#) and [2e-3e](#) appear to reflect the underlying bathymetry, though the alternating bands are also likely due to the effects of meandering of the northern boundary on a coarse resolution grid.

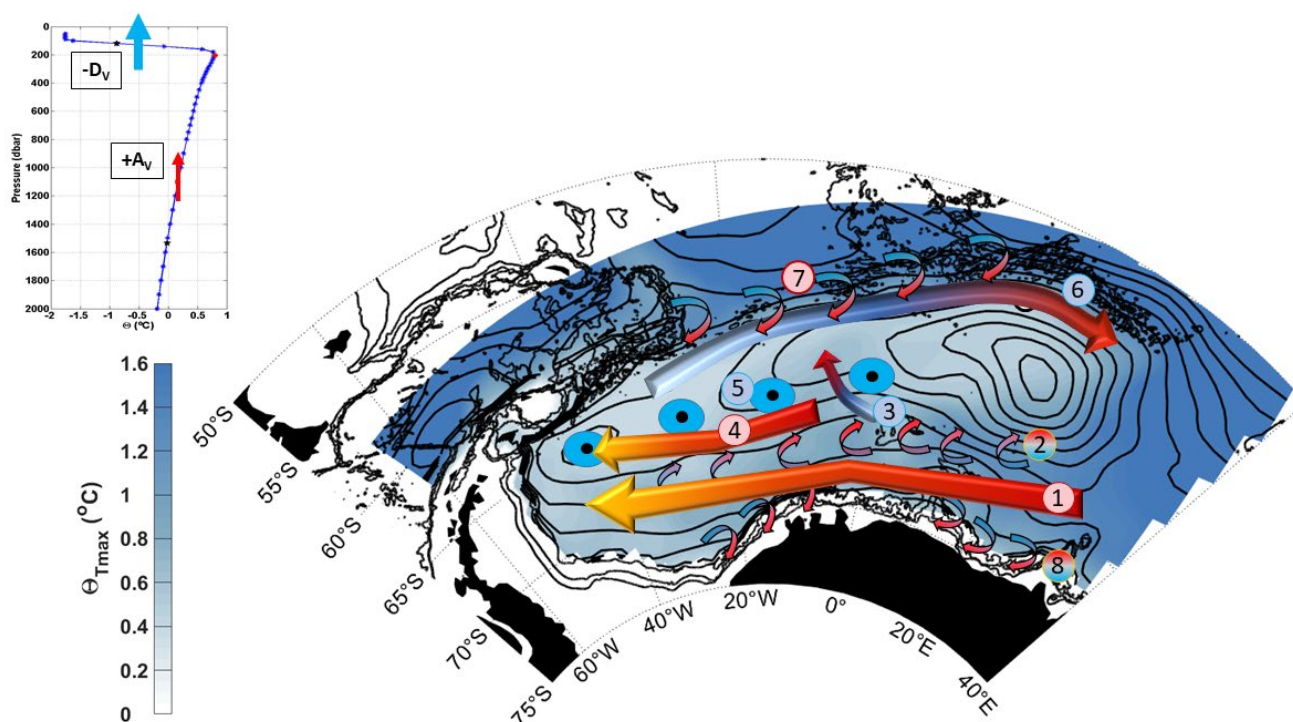
***Turbulent diffusion*** - There is also a relatively strong heat flux divergence due to horizontal turbulent diffusion along the southern boundary of the Weddell Gyre towards the coastline in Fig. [23c](#), especially between 40 and 10° W, and between Astrid ridge and Gunnerus ridge (between 10 and 30° E), indicating that horizontal turbulent diffusion may also constitute an important role in transporting heat towards the shelves along the southern coastline. Indeed, enhanced diffusive mixing over the continental slope (region not covered by this study) has been observed in the southwestern Weddell Sea (Fer et al., 2016; Daae et al., 2009). We are unable to directly compute turbulent heat fluxes across the southern boundary of the southern inflow limb toward the shelf edge due to the requirement of differencing across grid cells, and caution should be made with any attempt at inferring fluxes due to the large uncertainty of computing the heat budget at the boundary. Yet, we may estimate the shelf-ward heat flux indirectly. Fig. S6b reveals the net diffusive heat flux across the northern boundary of the southern inflow limb to amount to  $1.1 \pm 0.1$  TW  ~~$6.4 \pm 0.6$  TW~~. If we subtract this value from the total heat flux divergence due to horizontal turbulent diffusion of the SL of  $19 \pm 2.6 \pm 0.5$  TW (Table 3), then we can cautiously infer that most of the remaining

730 horizontal turbulent heat flux occurs southwards towards the ice-covered shelf seas west of Maud Rise (Fig. 2e3c), of  $12.65 \pm$   
735  $3.1$  TW (southwards having a much larger length with larger negative values in Fig. 2e-3c in comparison to the western end of  
the SL region, (i.e., the magenta stipled area in Fig. 2b4).

These results indicate that horizontal turbulent diffusion may play an important role in transporting heat southwards across the  
open northern boundary of the Weddell Gyre (Fig. 5b-7b & S6a), and also southwards towards the continental shelves along  
the Antarctic coast (Fig. 2e3c). Furthermore, the horizontal turbulent diffusion of heat may allow for the removal of some heat  
from the southern limb of the Weddell Gyre (Fig. 79b & S6b), before it is able to advect westwards towards the southwestern  
corner of the gyre, where the fragile large Filchner-Ronne ice shelf and the ice shelves of the Antarctic Peninsula are located  
(Hellmer et al., 2012), and where recent advance in the understanding of ocean heat fluxes have been made (Ryan et al., 2020).  
740 Since the turbulent diffusive heat fluxes are dependent on horizontal temperature gradients (related to geostrophic shear), this  
implies a complex interaction between the strength of the Weddell Gyre, thus mean horizontal advection, and the rate of  
meridional turbulent diffusion. Potentially, up to a certain point, meridional turbulent diffusion may provide a buffer, protecting  
the southwestern gyre from increased advective heat fluxes resulting from an intensified Weddell Gyre, by also increasing in  
intensity (due to stronger lateral temperature gradients and velocity shear). This mechanism requires careful understanding if  
we are to understand the role of the Weddell Gyre in the redistribution of heat in a changing climate.

## 745 6 Conclusions

Gridded climatologies of temperature and velocity derived from Argo floats spanning 2002-2016 were used to determine the  
heat budget of a 1000 m thick layer encompassing the core of WDW within the Weddell Gyre. This investigation was to  
establish the mechanisms by which heat is distributed throughout the Weddell Gyre, implicitly assuming non-divergent,  
geostrophic flow conditions. The mechanisms are summarised in the form of a basic schematic in Fig. 810, and interpreted  
750 below.



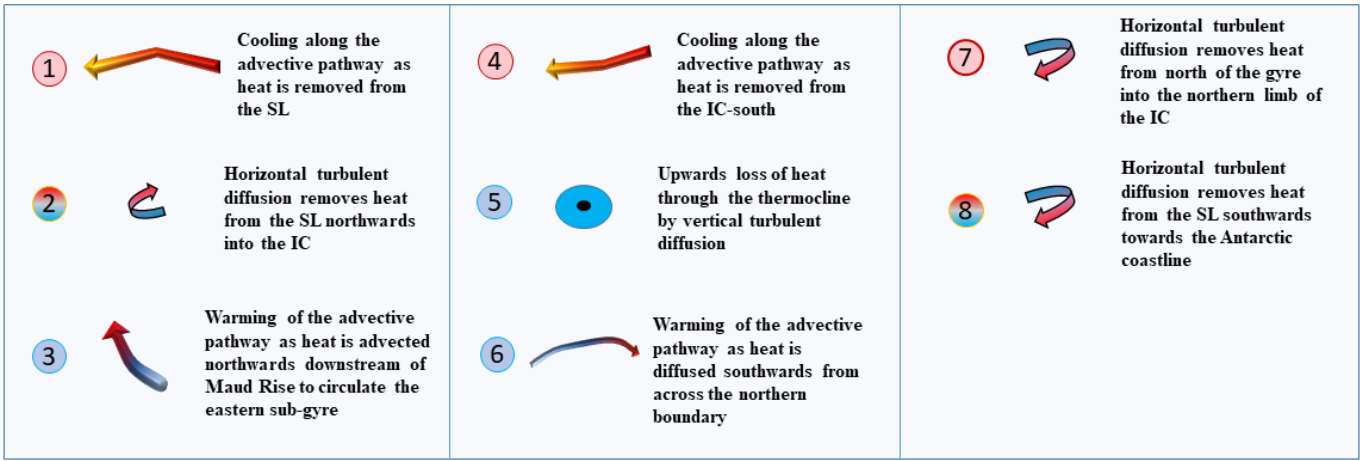


Figure 810: Schematic of proposed mechanisms by which heat is transported throughout the Weddell Gyre, based on interpretation of results in Sections 4 & 5.32.3. The blue-scale colour shading shows  $\Theta_{\max}$  and the black contours show the stream function with 5 Sv spacing, as in Fig. 1. The numbered keys assigned to each feature (arrows and circles) are described in panel b and Table 4. Where the colour of the number and feature are red, heat flux convergence (heat source) is indicated, and where they are blue indicates heat flux divergence (heat sink). The horizontal turbulent diffusive fluxes, indicated by small curved arrows (2,7,8), change from blue to red to indicate a direction of heat flux, (i.e., in 2, horizontal turbulent diffusion removes heat from the SL northwards into the IC, and southwards to the Antarctic coastline (8), and in 7, removes heat from north of the gyre, into the northern limb of the IC). The change in colour from red to orange of the arrow in the SL (1) and IC-south (4) indicates cooling along the advective pathway as heat is removed from both areas, whereas the change in colour from pale blue to red of the arrow in the northern limb (6) indicates warming of the advective pathway as heat is diffused southwards from across the northern boundary, and to a lesser extent as heat is advected northwards downstream of Maud Rise to circulate the eastern sub-gyre (3). The blue circles with the central black dot (5) demonstrates the upwards loss of heat through the thermocline by vertical turbulent diffusion. The upper left inset shows an example of the vertical temperature profile, with  $-D_v$  indicating a vertical diffusion of heat upwards out of the layer through the thermocline, and  $+A_v$  indicating a vertical advection of heat upwards into the layer from below.

Table 4: Summary of the key heat budget terms that are shown in the schematic in Fig. 8. The numbers correspond to the number key in Fig. 108, and the associated features in the schematic. Abbreviations used for the heat budget terms are defined in Table 2.

Key	Process	Net heat budget contribution (TW)
1	$A_H$ into SL	$+26 \pm 4 + 8 \pm 3$ TW
2	$D_H$ out of SL (into IC-south)	$-19.0 \pm 2.6 \pm 0.5$ TW (into IC-south: $-6.4 \pm 0.6 \sim 1.1 \pm 0.1$ TW, with a net heat flux convergence in IC-south of $1.6 \pm 0.4$ TW)
3	$A_H$ out of IC-south (West of MR and East of $10^\circ W$ ; i.e., circulating eastern sub-gyre)	$3 \pm 3.2 \pm 1.4$ TW
4	$A_H$ into IC-south (downstream of MR west of $10^\circ W$ )	$5 \pm 3$ TW
5	$D_v$ out of IC-south	$-6.5 \pm 0.3$ TW
6	$A_H$ out of IC-north	$-68 \pm 15$ TW
7	$D_H$ into IC-north from ACC	$78 \pm 434 \pm 2$ TW (direct flux across northern boundary is $\sim 32.17 \pm 1$ TW, suggesting over half exits the eastern end of the IC-north layer)

If we accept the potential sources of uncertainty discussed in Section 5, we can interpret the results presented in this study as follows:

- 755
1. In the SL of the Weddell Gyre, heat *convergence* due to mean horizontal and vertical advection is balanced by a *divergence* of heat due to horizontal and vertical turbulent diffusion.  $26 \pm 48 \pm 3$  TW of heat is advected into the southern limb of the Weddell Gyre (Fig. 810, no. 1), while  $6 \pm 0.5$   $19 \pm 2$  TW of heat is removed from the same region by horizontal turbulent diffusion (Fig. 810, no. 2).  $1.1 \pm 0.1$   $6.4 \pm 0.6$  TW of heat is directly removed from the southern limb by horizontal turbulent diffusion northwards into the IC, ~~half of which occurs west of Maud Rise which accounts for most of the net heat flux convergence due to horizontal turbulent diffusion in the IC-south~~ ( $1.6 \pm 0.4$  TW).
  - 760 2. In contrast, in the IC, heat *convergence* due to horizontal turbulent diffusion is ~~partly not~~ balanced by heat *divergence* due to mean horizontal advection.  $37 \pm 274 \pm 6$  TW of heat is advected out of the IC (Fig. 810, no. 6), while  $7436 \pm 92$  TW of heat is turbulently diffused into the IC (Fig. 810, no. 7).
  - 765 3. When we consider the eastward flowing northern limb and the westward flowing southern limb as separate regions, we see that the heat flux *convergence* due to horizontal diffusion mostly occurs in the northern limb of the IC. ~~Half~~ ~~The~~ net heat flux *convergence* into the IC-north of  $78-34 \pm 42$  TW is ~~(mainly)~~ due to a southward heat flux across the northern boundary of the gyre ( $32-17 \pm 1$  TW, Fig. 810, no. 7), ~~and the remainder mostly also~~ from the eastern gyre ~~periphery~~. A net heat flux *convergence* due to horizontal turbulent diffusion also occurs in the IC-south of  $1.1 \pm 0.1$   $3.3 \pm 1$  TW (Fig. 810, no. 2). Horizontal mean advection acts to remove heat from the northern limb of the IC ( $-68 \pm 15$  TW; Fig. 810, no. 6).
  - 770 4. There is an advective pathway of heat northwards from the gyre's southern limb between 10 and 20° W, which has an impact on the spatial distribution of  $\Theta_{\max}$ , and removes  $2 \pm 1.4$   $3 \pm 3$  TW from the IC-south (i.e., north of Maud Rise), between 0°E and 10° W to circulate the eastern sub-gyre (Fig. 810, no. 3). West of 10° W, heat is advected into the IC-south ( $5 \pm 3$  TW) (Fig. 810, no. 4). The main mechanism that removes heat from the IC-south is vertical turbulent diffusion (heat is lost upwards through the thermocline with a net flux of  $-65 \pm 0.3$  TW; Fig. 810, no. 5).
  - 775 5. At the southern boundary, there appears to be a diffusive flux of heat towards the shelf seas along the southern coastline, especially between 10 and 30° W, and 10 and 30° E. This may account for the remaining heat flux divergence due to horizontal turbulent diffusion along the southern limb of the Weddell Gyre (i.e.,  $19-6 \pm 20.5$  TW in table 3, of which  $6-41.1 \pm 0.16$  TW is diffused northwards (Fig. 810 and Fig. S6b), leaving a divergence of  $12-65 \pm 32$  TW to be diffused southwards along the southern boundary in Fig. 108, no. 8). While errors are high at the boundary since Argo floats cannot fully resolve shelf edge boundary currents, this may indicate that horizontal turbulent diffusion plays an important role in delivering heat towards the ice-covered shelf seas.
  - 780 6. In contrast to the horizontal terms, the vertical terms are spatially uniform, where the vertical advection term leads to heat flux convergence throughout both the SL and IC while the vertical turbulent diffusion term results in heat flux divergence. The terms are nearly negligible in comparison to the horizontal terms. Given that Ekman pumping velocity is mostly positive upwards throughout (Fig. S3), and that the lower boundary of the layer is warmer than the upper boundary throughout (based on the vertical boundary conditions in 2.1, also see example of temperature profile ~~inserted~~ in Fig. 810), we can interpret the vertical terms as:
    - a. heat is uniformly advected upwards through Ekman upwelling into the layer from below (SL:  $0.79 \pm 0.3$   $\text{Wm}^{-2}$ , or a total of  $+0.92-5 \pm 0.023$  TW, and IC:  $0.8 \pm 0.12$   $\text{Wm}^{-2}$ , or a total of  $6-52.8 \pm 003.1$  TW).
    - 790 b. heat is diffused upwards out of the top of the layer, due to the relatively strong vertical temperature gradient at the thermocline, and due to vertical instabilities at the thermocline (SL:  $-2.37 \pm 0.5$   $\text{Wm}^{-2}$ , or a total of  $-2.87-2 \pm 0.24$  TW, and IC:  $-2.13 \pm 0.3$   $\text{Wm}^{-2}$ , or a total of  $-817 \pm 0.46$  TW). This heat may eventually be lost to the atmosphere, especially in the SL.

795 7. East of the Prime Meridian, 2 ellipses of strong heat flux divergence and 2 ellipses of strong heat flux convergence  
are found to the north and south of the eastern sub-gyre respectively. This is directly related to horizontal mean  
advection and is likely an artifact of a misalignment between horizontal circulation and strong horizontal temperature  
gradients on a coarse resolution grid. However, the “misalignment” may be linked to the occurrence of unresolved  
mesoscale eddies that are not represented by turbulent heat flux diffusion, that may be skewing the mean state  
representation of the temperature and flow fields. This is possibly-in part due to poor data coverage in a region where  
instabilities are likely generated from the interaction between “cold” regime recirculating WDW and incoming  
800 “warm” regime WDW (Gordon and Huber, 1984 and Fig. 1). Unresolved horizontal circulation around Maud Rise  
adds to the uncertainty in the region. Thus, to improve estimates in the eastern Weddell Gyre, more observations are  
required to resolve the complex ocean dynamics on smaller length-scales in the eastern sub-gyre region.

805 From using Argo floats, we have described the heat budget of a 1000 m thick layer encompassing the core of WDW within  
the Weddell Gyre. The role of mean horizontal advection is evident in feeding heat towards the southwestern Weddell Gyre,  
where the Filchner-Ronne ice shelves and Antarctic Peninsula are located. What is also important, however, is understanding  
the respective roles of mean horizontal advection and horizontal turbulent diffusion in removing some of that heat from the  
southern limb of the Weddell Gyre before it is able to reach the southwestern interior. This is crucial since Hellmer et al. (2012)  
suggest that under future climate scenarios, a redirection of the coastal current toward the Filchner-Ronne ice shelf could lead  
810 to increased advection of waters into the ice-shelf cavity, leading to increased basal ice melt from 0.2 to 4 m/year. While  
Naughten et al. (2021) argues that global temperatures would need to reach 7°C for warm water to intrude to the ice-shelf  
cavity, which exceeds the pledges in the Paris Agreement, nonetheless, the authors argue that unless global temperatures level  
out, melting of the ice shelf will at some point prevail.

815 **Data availability:** The Argo float data were collected and made freely available by the International Argo Program and the  
national programs that contribute to it (<http://www.argo.ucsd.edu>, <http://argo.jcommops.org>). The horizontal diffusivities were  
provided by Florian Sevellec (Sevellec et al., 2020: <https://doi.org/10.17882/91335>). The ERA5 dataset is available at  
<https://www.ecmwf.int/en/forecasts/datasets/reanalysis-datasets/era5>. Sea-ice data are available at <https://nsidc.org/data/nsidc-0116/versions/4>. Objectively mapped temperature from Reeve et al. (2016) is available at  
<https://doi.pangaea.de/10.1594/PANGAEA.842876>.

820 **Author Contribution:** K.A. Reeve performed the data analysis, figure preparation and wrote the manuscript with contributions  
from all authors.

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# Supplements to “The Weddell Gyre heat budget associated with the Warm Deep Water circulation derived from Argo Floats”

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## S1. Vertical boundary conditions for heat budget integrals

Since we are unable to integrate for the full ocean depth (Argo floats profile the upper 2000 dbar), and most importantly, to the surface of the ocean, due to the high seasonal variability at the surface (Section 2.2), we need to decide on suitable vertical boundaries for the heat budget integrals. However, if we simply select uniform depths for the vertical boundaries, we may include certain features in the TS-profile (e.g., such as the temperature minimum) in some areas of the Weddell Gyre, while excluding them in other regions. Hence, we need to ensure the upper boundary is defined such that  $\Theta_{\max}$  (i.e., the core of WDW) is always fully included in the vertical layer, given that the core of WDW is the primary source of heat to the Weddell Gyre. We also need to avoid incorporating seasonally variable Surface Water and Winter Water (a surface water mass characterised by a sub-surface temperature minimum, e.g., Behrendt et al., 2011). Otherwise, the vertical advection component will be stronger in regions where the profile includes the temperature minimum layer due to a larger vertical temperature gradient, in contrast to regions where the Winter Water is altogether omitted. This will result in bias in the spatial distribution of the heat budget terms. It is also necessary to ensure the same thickness of water is analysed throughout the Weddell Gyre. We resolve this by defining the upper boundary as the mid-point of the thermocline, and the lower boundary as 1000 m below the thermocline mid-point (Fig. 2 & S1). That way, regardless of location, we exclude Winter Water, but always include the core of WDW, and maintain equal layer thickness throughout. An example vertical temperature profile in Fig. 2 illustrates the vertical boundary conditions, while a map of the upper boundary depth is shown in Fig. S1.

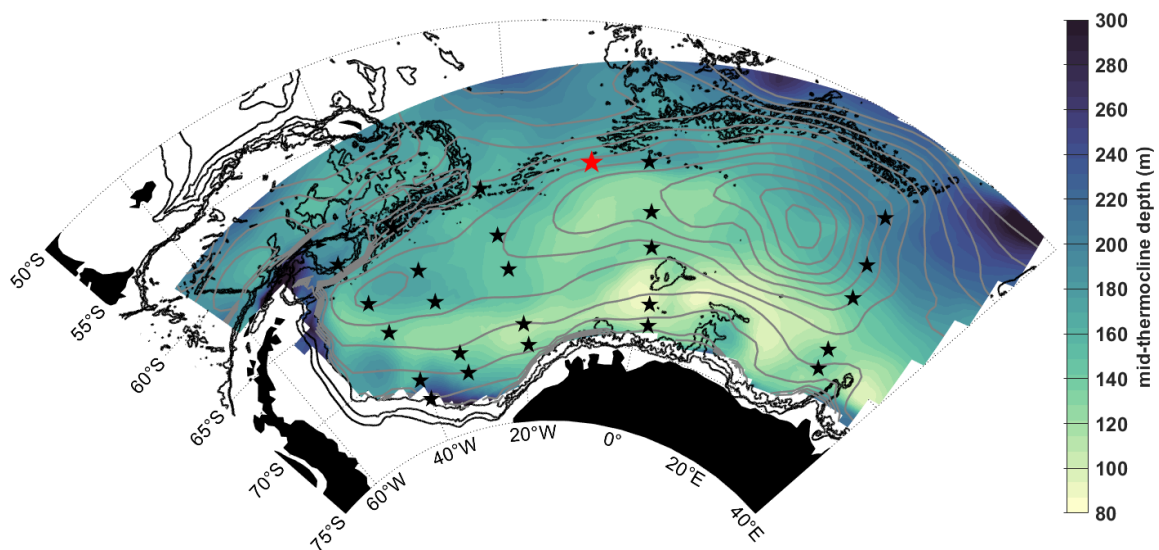


Figure S1: The mid-thermocline, defined as the upper boundary depth (m), with stars marking the positions of profiles plotted in Fig. 2, and streamlines (grey contours) of the vertically integrated stream function for 50-2000 dbar with a spacing of 5 Sv, derived from in-situ observations from Argo floats (Reeve et al., 2019, 2016), as in Fig. 1. The black contours show the 1000, 2000 and 3000 m isobaths, from the general bathymetric chart of the oceans (GEBCO, IOC et al., 2003).

## S2. Air-sea-heat fluxes

The net air-sea heat flux is the sum of the following components: shortwave solar radiation, longwave thermal radiation, latent heat flux and sensible heat flux. The air-sea heat fluxes are not directly relevant in the calculation of the heat budget in this study, since the heat budget calculations are restricted to depths exceeding 80 m (i.e., the minimum depth of the mid-thermocline is below 50 m, Fig. S1), and downward shortwave radiation decreases exponentially with depth (only 1% of the radiation reaches past 75 m; Tamsitt et al., 2016). However, gaining perspective on where we expect heat loss through the surface of the ocean may help to understand the mechanisms through which WDW loses its heat as it circulates the Weddell Gyre. Long-term means between 2002 and 2016 of surface heat fluxes are derived from ERA-interim reanalysis (Hersbach et al., 2020), based on previous convictions that it provides one of the most reliable estimates for the Southern Ocean (Jones et al., 2016). We also use this data source to derive vertical Ekman pumping velocity, as detailed in the following Section S3.

The net air-sea heat flux into the ocean in Fig. S2 (here on referred to as  $Q_{\text{net}}$ ) is positive along the northern limb of the gyre and negative over the southern limb, creating a north-south divide (the thick black line in Fig. S2) parallel to the central gyre axis (the dashed black line in Fig. S2). There is, however, some latitudinal offset between  $Q_{\text{net}}$  and the central axis of the depth-integrated circulation. Thus, the IC is dominated by positive  $Q_{\text{net}}$  fluxes, and the zero contour in  $Q_{\text{net}}$  (thick black line in Fig. S2) appears south of the central gyre axis, roughly between the IC and the SL, instead of directly along the central axis of the gyre.

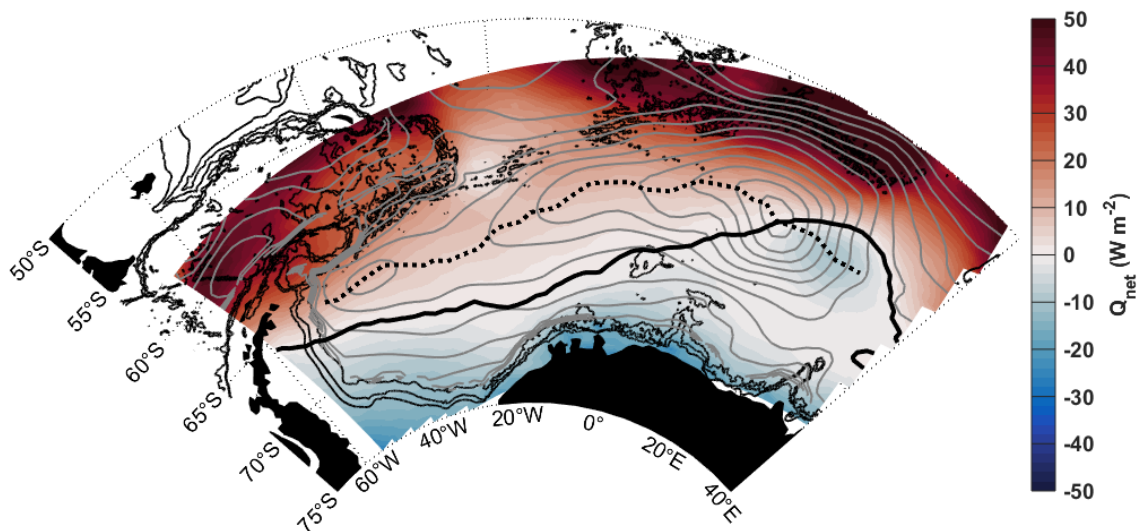


Figure S2: Net air-sea heat flux (i.e., the sum of long/short-wave radiation, latent and sensible heat fluxes, in  $\text{W m}^{-2}$ ): long-term mean from 2002 to 2016, from ERA-interim reanalysis data. Positive values indicate heat flux into the ocean. The thick black line indicates the zero-contour line for  $Q_{\text{net}}$ . The central gyre axis is indicated by the dashed black line, defined as the meridional maximum stream function value. South of the dashed line the flow of the water column is westward and flow north of the dashed line is eastward. Streamlines (grey contours) show the vertically integrated stream function for 50-2000 dbar with a spacing of 5 Sv, derived from in-situ observations from Argo floats (Reeve et al., 2019, 2016), as in Fig. 1. The black contours show the 1000, 2000 and 3000 m isobaths, from the general bathymetric chart of the oceans (GEBCO, IOC et al., 2003).

## S3. Vertical Ekman pumping velocity derived from the wind stress field

Since the flow is assumed geostrophic (non-divergent), vertical velocity is assumed constant with changing depth (outside of the turbulent surface and bottom boundary layers, which are excluded in this study, see Section S1). As an estimate of the

vertical velocity is required in the vertical advection term (the second term on the right-hand side of Eq. 1.1), we therefore use the Ekman vertical velocity,  $w_E$ . The Ekman velocity,  $w_E$ , was derived from the mean wind stress field, sea ice concentration ( $\alpha$ ) and sea ice velocity ( $U_{ice}$ ) for the period 2002-2016, following Le paih et al. (2020) and Dotto et al. (2018):

$$w_E = \vec{\nabla} \times \left( \frac{\vec{\tau}}{\rho_0 f} \right) \quad (2.1)$$

where

$$\vec{\tau} = \alpha \vec{\tau}_{ice-ocean} + (1 - \alpha) \vec{\tau}_{air-ocean} \quad (2.2)$$

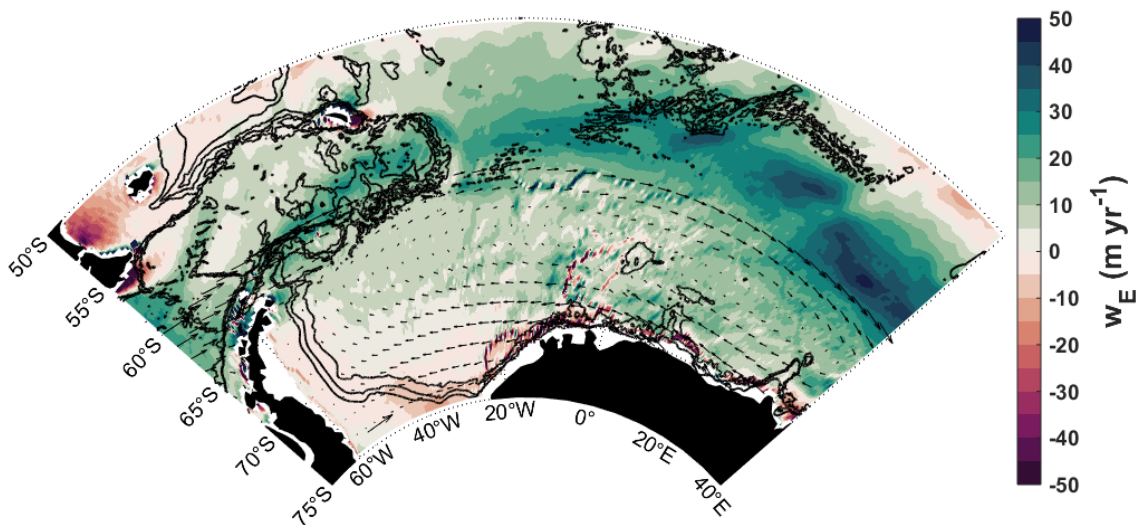
$$\tau_{air-ocean} = \rho_{air} C_{dw} |\overline{U_{air}}| \overline{U_{air}} \quad (2.3)$$

and

$$\tau_{ice-ocean} = \rho_0 C_{id} |\overline{U_{ice}}| \overline{U_{ice}} \quad (2.4)$$

where  $\rho_0$  is seawater density ( $1027 \text{ kg m}^{-3}$ ) and  $f$  is the Coriolis parameter;  $\vec{\tau}$  is the total surface stress, which takes into account sea ice concentration,  $\alpha$ , by combining stress over the ocean by wind ( $\vec{\tau}_{air-ocean}$ ) and stress over the ocean by ice ( $\vec{\tau}_{ice-ocean}$ ).  $\rho_{air}$  is the density of air ( $1.29 \text{ kg m}^{-3}$ ),  $C_{dw}$  is the wind drag coefficient ( $1.5 \times 10^{-3}$ ),  $\overline{U_{air}}$  is the wind velocity at 10 m above sea level;  $C_{id}$  is the ice-ocean drag coefficient ( $5.5 \times 10^{-3}$ ) and  $\overline{U_{ice}}$  is the sea ice velocity. The data sources for the variables in Eq. 2 are listed in Table 1, in Section 2.1.

Eq. 2.1 relies on a formulation for wind stress (air-sea momentum flux) of an ice-free ocean. Le Paih et al. (2020) showed that different stress formulations, including sea ice-sensitive ones, in the Weddell Gyre area yield momentum input into the ocean within  $\pm 10\%$  relative to the formulation used here. Since  $w_E$  is assumed to be constant with depth below the Ekman layer (i.e., assuming the horizontal flow below the Ekman layer to be in perfect geostrophic balance), and the surface mixed layer is omitted from the objectively mapped grids of velocity and temperature,  $w_E$  is a suitable estimate of the vertical velocity within the scope of this study. The resulting  $w_E$  was averaged over 2002-2016 (Fig. S3) and re-gridded to the same grid as the Argo data, using a simple distance-weighted mean.  $w_E$  is positive throughout most of the Weddell Gyre (indicative of upwelling), whereas downwelling is found along the Antarctic coastline and on the continental shelf in the south-west, which is excluded from this study since Argo float data are sparse for the shelf seas.



**Figure S3: Mean 2002-2016 vertical Ekman pumping velocity ( $w_E$ : m/year), from Era-interim ECMWF reanalysis (ERA5 monthly mean output; Hersbach et al., 2020) and sea-ice velocity from Polar Pathfinder Daily 25 km EASE-Grid Sea Ice Motion Vectors (Tschudi et al., 2019), where positive values indicate upwelling. Black arrows show surface wind stress ( $\text{N m}^{-2}$ ), south of  $60^\circ\text{S}$  (in**

order to emphasise the weaker surface wind stress within the Weddell Gyre region, which would disappear if we were to include the much stronger wind stress over the ACC); the arrow at 50°W, 74°S is 0.05 N m<sup>-2</sup>. The black contours show the 1000, 2000 and 3000 m isobaths, from the general bathymetric chart of the oceans (GEBCO, IOC et al., 2003).

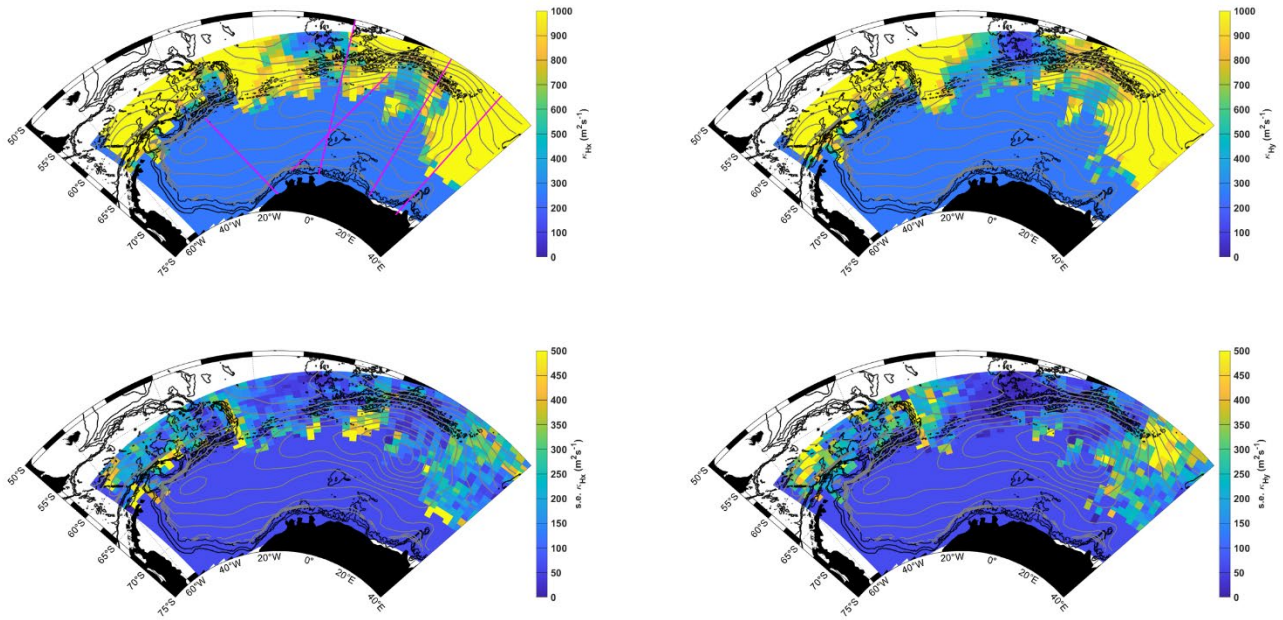
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#### S4. Horizontal diffusivity

In the north and east, we use horizontal diffusivity estimates from Sevellec et al. (2022; 2020). For all data within 1° lat/lon of a grid cell, we take the mean and standard error of the mean (s.e.). For grid cells where no data is available, we use  $247 \pm 63$  m<sup>2</sup>s<sup>-1</sup>. The Sevellec et al (2022) dataset might underestimate values within the interior gyre where sea ice dominates, which is why we replaced these values with  $247 \pm 63$  m<sup>2</sup>s<sup>-1</sup>, from Donnelly et al. (2017). This is because the authors need ice-free consecutive float positions for an adequate length of time for the implementation of their method, suggesting the data are biased to conditions where the sea ice has melted and the wind has not yet stirred up the water column, leaving a layer of fresh water over the surface (as occurred, for example, in Cisewski et al., 2011). Also, given the Sevellec dataset is representing 1000 m, whereas the depth range we are looking at varies from 87 to 1530 m, with a mean upper boundary of 166 m, and a mean middle depth of 666 m, we can expect higher values overall for the shallower layer. Maps of the resulting zonal and meridional horizontal diffusivities used in this paper are provided in Fig. S4, along with their corresponding standard errors.

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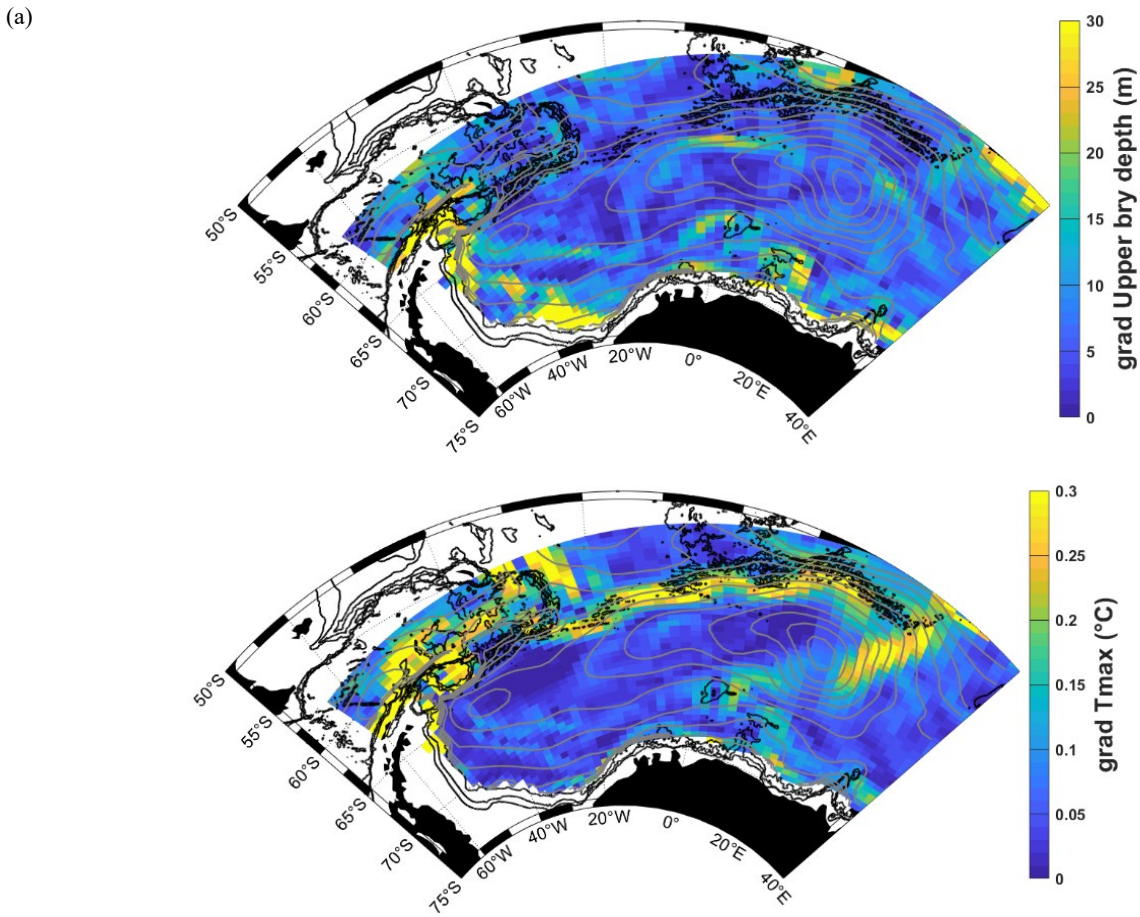
**Figure S4: Horizontal diffusivity as a combination of the Sevellec et al. (2022) dataset and infilling with a background value of  $247 \pm 63$  m<sup>2</sup>s<sup>-1</sup>, from Donnelly et al. (2017). The upper panels show the zonal and meridional components from left to right, whereas the lower panels show their corresponding standard errors respectively. In the upper left panel, magenta lines mark the approximate positions of the ship-based data used in Donnelly et al., (2017).**

#### S5. Investigating uncertainty: horizontal gradients

To look into the causes of the large ellipses in the pattern of horizontal heat advection, spread diagonally north-eastwards from Maud Rise (Fig. 3a), we checked the horizontal gradients in the upper boundary depth and the sub-surface temperature maximum, shown in Figs. S5a-b respectively. These have been discussed in Section 5.2.2.

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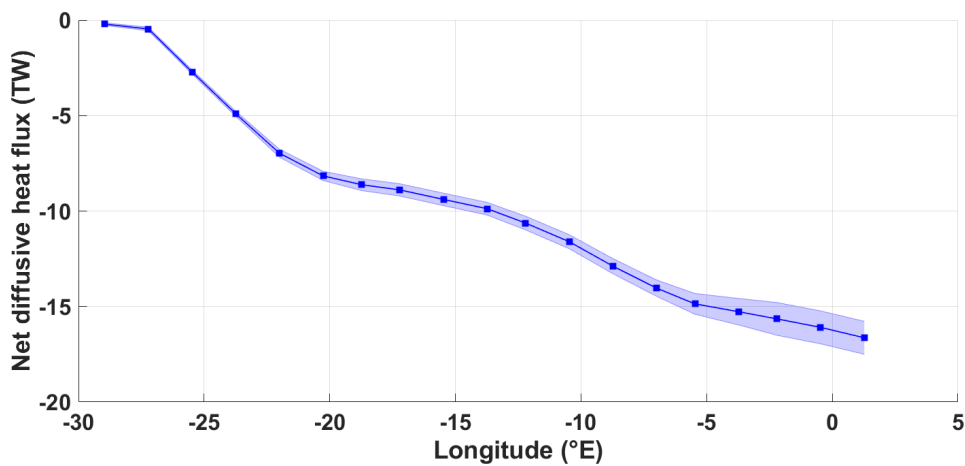




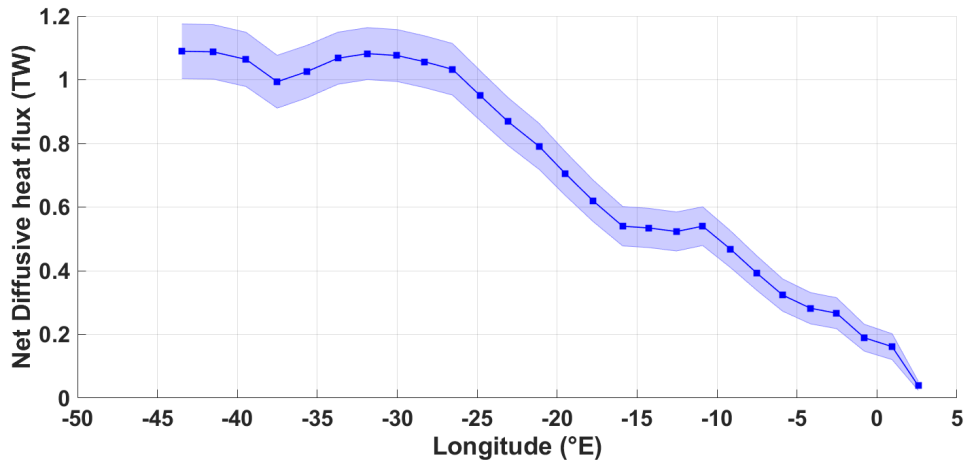
(b)  
**Figure S5: (a) Horizontal gradient in the upper boundary depth (i.e. mid-thermocline depth; m), and (b) horizontal gradient in the sub-surface temperature maximum (°C).**

### S6. Diffusive heat fluxes across boundaries

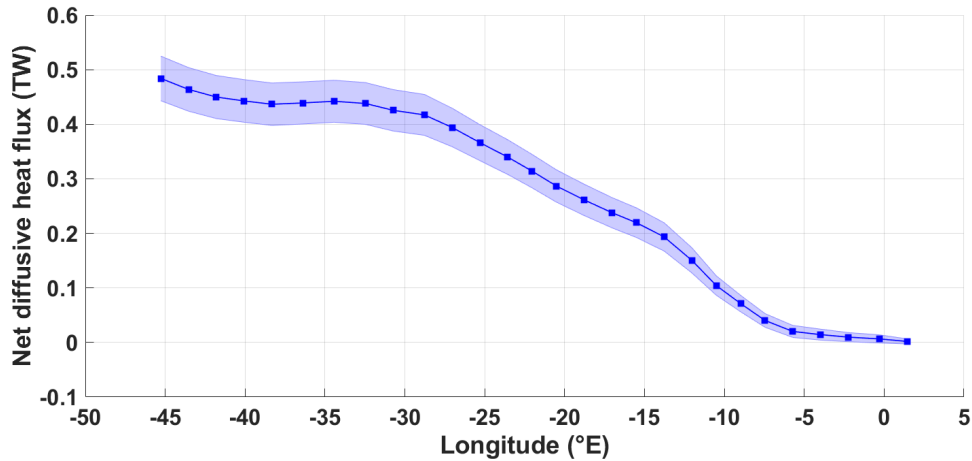
1090 In Section 4.2 we investigate the role of horizontal turbulent diffusion in redistributing heat across zonal boundaries: across the northern boundary of the Weddell Gyre from the ACC (Fig. 7b), across the interface between the open SL and the IC-south (Fig. 9b), and across the central gyre axis between IC-south and IC-north (Fig. 9c). Provided below in Fig. S6 are the zonally integrated diffusive heat fluxes corresponding to each of these regions respectively, from which we derive our net diffusive heat fluxes in Section 4.2.



(a)



(b)



(c)

**Figure S6: the cumulative sum of the diffusive heat flux in TW. (a) across the northern gyre boundary from west to east, where negative values indicate a southward flux of heat into the eastward-flowing northern limb of the Weddell Gyre from north of the northern Weddell Gyre boundary; (b) along the boundary between the southern inflow limb and IC-south, from east to west, where positive values indicate a removal of heat from the open southern limb of the gyre into the interior; in other words, a northward flux of heat from the southern limb into the interior circulation cell; (c) across the central gyre axis from IC-south to IC-north, where positive values indicate a diffusive heat flux northwards across the gyre axis. Errors are computed following the same method outlined in Section 3.2 and S7, with the cumulative net error summed in quadrature (i.e., the square root of the cumulative sum of the squared error).**

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### S7. Error propagation computation

The errors for each heat budget term are calculated using the laws of propagation, where the error for temperature is defined as the objective mapping error (Fig. S7; Reeve et al., 2016), and the error for horizontal velocity is provided from Reeve et al. (2019), which contains a sensitivity study, perturbing the original stream function by using mapping errors and uncorrected drift velocities, to get a maximum range of values (see Reeve et al., 2019 for more details). For further details refer to Sections 2.2 and 3.2. The error for  $w_E$  are the standard error of the mean of the timeseries variability for monthly values from January 2002 to December 2016, since uncertainties associated with the wind field are as of yet unavailable for ERA-5 reanalysis data. Lastly, we use the error ranges for horizontal and vertical diffusivity as provided in Donnelly et al. (2017), for all regions within the Weddell Gyre, or, where available, the Sevellec et al. (2022) dataset, which also provides error estimates (Section S4). With all this in mind, the derivation for the propagated errors for each term are provided below.

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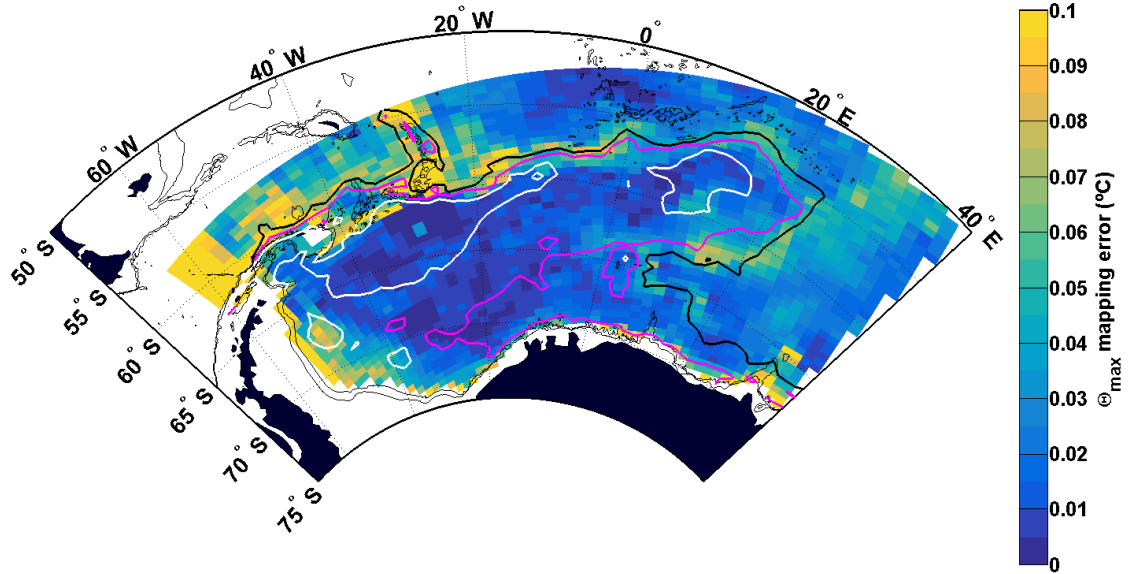


Figure S7: Mapping error for  $\Theta_{\max}$  (°C). The gridded  $\Theta_{\max}$  field is shown in Fig. 1. The white, magenta and black contours show the 0.5, 0.8 and 1.1 °C isotherms respectively. The grey contours are as in Fig. 1.

For any one grid cell, the uncertainty for mean horizontal advection is the propagation of error as applied to two variables multiplied together:

$$\delta A_H = |U\theta| \cdot \sqrt{\left(\frac{\delta\theta}{\theta}\right)^2 + \left(\frac{\delta U}{U}\right)^2 + \frac{2COV(U\theta)}{|U\theta|}} \quad (S2.1)$$

where  $\delta$  is the uncertainty of a variable, and  $COV(U\theta)$  is the covariance between  $U$  and  $\Theta$ , which is  $COV(U\theta) = \rho_{\theta U}\delta U\delta\theta$ , where  $\rho_{\theta U}$  is the correlation between  $\Theta$  and  $U$ . The resulting uncertainties are propagated using the root-sum-square (RSS) law of error propagation, which is first applied to the differential:  $\nabla_H(U \cdot \theta) = \frac{d(u\theta)}{dx} + \frac{d(v\theta)}{dy}$  (the length scales  $\Delta x$  and  $\Delta y$  are also applied). RSS is then applied to the vertical integration for the 1000 m thick layer, and also applied when integrating horizontally over larger areas. RSS is applied when integrating vertically and horizontally in the same way for all terms in Eq.1.1, and the resulting uncertainty of all terms are scaled by  $\rho_0$  and  $C_p$ , as well as the associated length scales ( $\Delta x$  and  $\Delta y$  in the horizontal terms,  $\Delta z$  for the vertical terms) as in the heat budget in Eq. 1.2.

For the mean vertical advection term, RSS is applied to the uncertainties of the temperature difference between the upper and lower boundaries, and then this error is propagated with the uncertainty associated with  $w_E$  as follows:

$$\delta A_V = |A_V| \cdot \sqrt{\left(\frac{\delta w_E}{w_E}\right)^2 + \left(\frac{\delta T}{T}\right)^2} \quad (S2.2)$$

where  $T = \theta_{mT} - \theta_{mT-1000}$ , i.e., the temperature difference between the upper and lower boundary (Eq. 1.2), and the resulting uncertainty of  $T$  is:

$$\delta T = \sqrt{\delta^2\theta_{mT} + \delta^2\theta_{mT-1000}} \quad (S2.3)$$

Horizontal turbulent diffusion (i.e., the third term on the right-hand side of Eq. 1.1) consists of a double differential of  $\Theta$  multiplied by  $\kappa_H$ , both of which have an associated uncertainty. The double differential is computed by taking the temperature differences between grid cells (i.e.,  $\tau = (\Theta_{i-1} + \Theta_{i+1} - 2\Theta_i) / \Delta x^2$ , where  $i$  represents the grid index for one direction; this is

repeated for both x and y directions). Thus, we apply the propagation of error to the uncertainty of the double differential using RSS (i.e., we sum together the square of 12 uncertainties for temperature, which include  $2^2$  x grid cell  $i$ , plus its surrounding grid cells,  $i-1$  and  $i+1$ , which is repeated in both  $x$  and  $y$  directions, and then scaled by  $\Delta X^2 = \Delta x^2 = \Delta y^2$ ; Eq. S2.5). We then take that uncertainty and apply the propagation of error for two variables multiplied together to incorporate the uncertainty for  $\kappa_H$  (Eq. S2.4). The vertical and horizontal integrations are treated using RSS, as above.

$$\delta D_H = |\kappa_H \nabla_H^2 \theta| \cdot \sqrt{\left(\frac{\delta \kappa_H}{\kappa_H}\right)^2 + \left(\frac{\delta \tau}{\tau}\right)^2} \quad (\text{S2.4})$$

$$\text{Where } \delta \tau = \sqrt{\frac{12 \delta \tau^2}{\Delta X^2}} \quad (\text{S2.5})$$

Lastly, vertical turbulent diffusion is treated in the same manner as horizontal turbulent diffusion, except that it is applied in the vertical (i.e., the horizontal operator is not required when computing this term), and thus vertical integration is already accounted for:

$$\delta B = \frac{\sqrt{4 \delta \theta^2}}{|B|} \quad (\text{S2.6})$$

$$\text{where } B = \left(\frac{d\theta}{dz}\right)_{mT} - \left(\frac{d\theta}{dz}\right)_{mT-1000} \quad (\text{S2.7})$$

The resulting propagated error is then:

$$\delta D_V = |\kappa_V B| \cdot \sqrt{\left(\frac{\delta \kappa_V}{|\kappa_V|}\right)^2 + \left(\frac{\delta B}{B}\right)^2} \quad (\text{S2.8})$$

Lastly, the net uncertainty of the sum of the heat budget terms is the RSS of the uncertainty of all 4 terms described above:

$$\delta_{HB} = \sqrt{\delta A_H^2 + \delta A_V^2 + \delta D_H^2 + \delta D_V^2} \quad (\text{S2.9})$$

The panels on the right side in Fig. 3 show the total propagated errors for the 4 heat budget terms and the sum of all terms respectively. The overall error is dominated by the mean horizontal advection term (Fig. 3a), and to a slightly lesser extent the horizontal turbulent diffusion term (Fig. 3c), with largest errors occurring to the north of the gyre and also over the eastern sub-gyre, particularly east of about  $20^\circ$  E. This is to be expected as this region is dominated by a mesoscale eddy field (Ryan et al., 2016; Schröder & Fahrbach, 1999), while also lacking in observations to adequately represent the long-term mean of the region's hydrography and circulation. The gyre periphery has the largest errors for the horizontal turbulent diffusion term, which corresponds to regions with sharp lateral temperature gradients and large objective mapping errors (Reeve et al., 2016), such as towards the southern coastline, where data are also lacking, and at the northern boundary of the gyre, where high temporal variability is known and sharp meridional temperature gradients are found. For further discussion regarding uncertainty, refer to Sections 5.1 & 5.2.2.

### S8. Zonal and meridional variation of the heat budget terms for the whole western Weddell Gyre

The zonal and meridional means and integrations for the whole Weddell Gyre region west of Maud Rise (i.e., SL + IC) were also computed for the heat budget terms. These figures are shown below. The zonal mean in Fig. S7 is very similar to the IC analysis, although it shows a net zero contribution of mean advection, which makes sense because of the gyre-characteristics of the circulation, where westward flowing southern and eastward flowing northern limbs cancel with each other: heat is advected into the gyre east of 10 °E, and then advected out of the gyre west of 40 °W. Horizontal turbulent diffusion dominates as a heat source east of ~25 °W, dominated by high values along the northern boundary. The mean contributions in the upper panel show three zonal peaks at 0 °E, 20 °W and 45 °W. These are related to the recirculation about the eastern sub-gyre, and the western sub-gyre respectively. The peak at 20 °W is particularly interesting as this is where the bottom bathymetry transitions from complex in the east to smooth in the west, which is known to impact diffusivity (Whalen et al., 2012) and overall circulation dynamics (Sonnewald et al., 2023).

The meridional mean in Fig. S8 also shows the closure of the mean advection component where the main heat source due to mean advection occurs in the southern limb (south of ~63°S) and the main heat sink due to mean advection occurs in the northern limb (north of ~63°S). Again, this agrees with all previous findings in the paper. Horizontal turbulent diffusion removes heat in the southern limb, and becomes a source of heat in the northern limb, with peaks occurring at ~61°S and ~59°S. These peaks might be related to the meridional change in the northern boundary across the area we are averaging over.

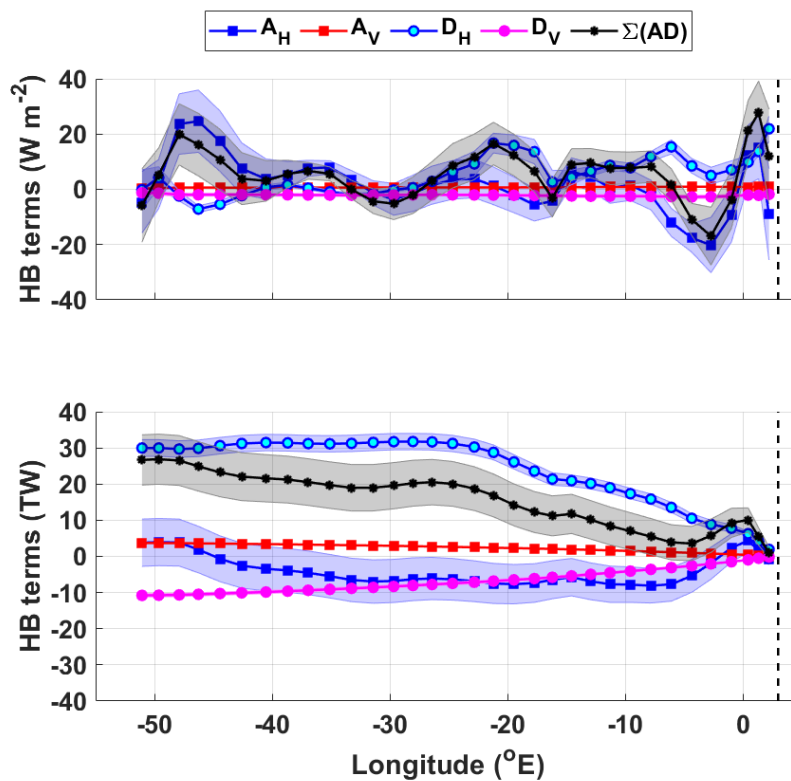


Figure S7: upper panel: the zonal mean heat budget terms, in  $\text{W m}^{-2}$ , for the whole Weddell Gyre west of 3 °E; lower panel: the corresponding cumulative heat budget terms in Terawatts (TW). The key for the legend is listed in Table 2. The dashed vertical line marks the approximate longitude of Maud Rise, at 3° E. The shaded errors provide the associated propagated error (detailed in section 3.2 and the supplement). The total region is marked by both blue and magenta stippling in Fig. 4.

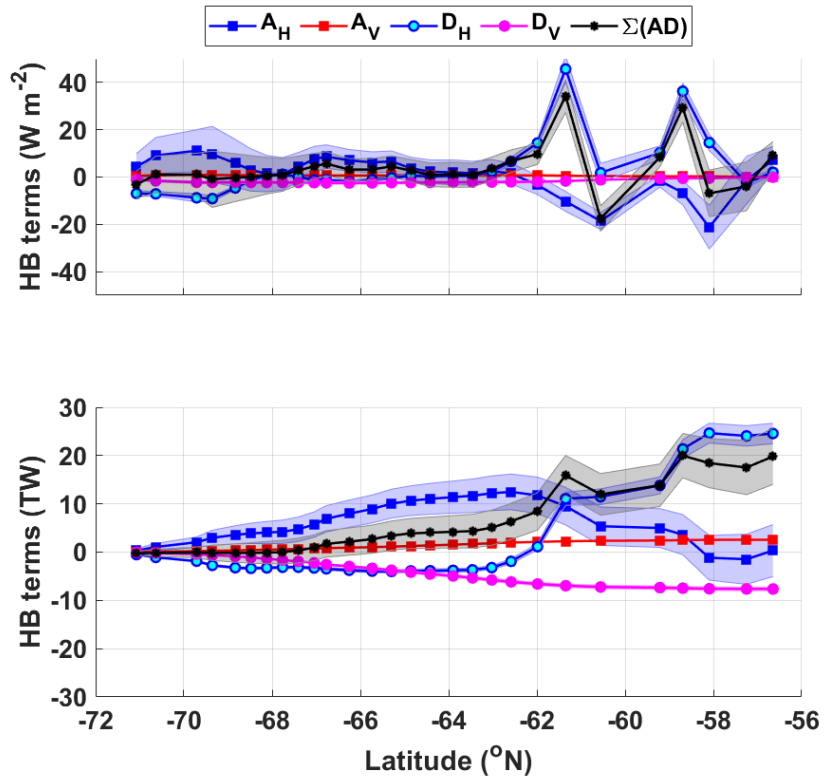


Figure S8: upper panel: the meridional mean heat budget terms, in  $\text{W m}^{-2}$ , for the whole Weddell Gyre west of  $3^\circ \text{E}$ ; lower panel: the corresponding cumulative heat budget terms in Terawatts (TW). The key for the legend is listed in Table 2. The dashed vertical line marks the approximate longitude of Maud Rise, at  $3^\circ \text{E}$ . The shaded errors provide the associated propagated error (detailed in section 3.2 and the supplement). The total region is marked by both blue and magenta stippling in Fig. 4.