



1 **Evidence of a dual African and Australian biomass burning**  
2 **influence on the vertical distribution of aerosol and carbon**  
3 **monoxide over the Southwest Indian Ocean basin in early 2020**

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## 2 **Abstract**

3 The pristine atmosphere of the southwest Indian Ocean (SWIO) basin underwent significant  
4 perturbations during the 2020 austral summer. This study documents the complex variability of  
5 aerosols and carbon monoxide (CO) over this remote oceanic region and identifies the processes  
6 governing it in the upper troposphere – lower stratosphere (UT-LS). Aerosol profiles exhibit a  
7 multi-layer structure in the tropical UT-LS in January and February 2020. The numerical  
8 models (FLEXPART and MIMOSA) showed that the modulation of the aerosol content in the  
9 lower stratosphere is due to the intense and persistent stratospheric aerosol smoke layer  
10 generated during the 2019-20 extreme Australian bushfire events. One part of this stratospheric  
11 aerosol smoke layer was advected zonally by the prevailing easterly winds and its passage over  
12 Reunion was recorded by increased aerosol extinction profiles on 27<sup>th</sup> and 28<sup>th</sup> January. The  
13 analysis of the advected potential vorticity highlights an isentropic transport of air masses  
14 containing Australian biomass burning aerosol from extra-tropical latitudes to Reunion at the  
15 400 K isentropic level, on 28<sup>th</sup> January. Interestingly, our results show that the BB activity in  
16 eastern Africa, weak during this season, contributed to modulate (up to 90%) the vertical  
17 distribution of CO and aerosols in the upper troposphere over the SWIO basin. The  
18 simultaneous presence of African and Australian aerosols smoke layers has been recorded by  
19 ground-based observations at Reunion. This study highlights for the first time the influence of  
20 the African emissions from BB to the CO and aerosols distribution in the upper troposphere  
21 over the SWIO basin during the convective season. The results show that besides PyroCb -  
22 driven injection of BB products to the stratosphere, an alternative pathway may exist during the  
23 regular deep-convection season in the tropics.

## 24 **1. Introduction**

25 Large amounts of aerosols and active trace gases such as carbon monoxide (CO) are injected  
26 throughout the atmosphere during biomass burning (BB) events. The Southern American and  
27 Southern African regions are recognized to be significant primary sources of carbonaceous  
28 aerosols and active traces gases in the Southern Hemisphere through the BB season from July  
29 to November (Bencherif et al., 2020; Garstang et al., 1996; Holanda et al., 2020). Such BB  
30 activities have the potential to modulate the vertical distribution of trace gases and aerosols  
31 from the troposphere to the stratosphere (Andreae and Merlet, 2001; Dufлот et al., 2010; Héron  
32 et al., 2020). Under favorable meteorological conditions, pyro-convection events can take place  
33 and have the potential to inject soot and smoke directly into the stratosphere (Dowdy and Pepler,  
34 2018; Fromm et al., 2010). Radiative impact of the aerosols and traces gases is determined by



1 abundance, vertical distribution, , and atmospheric residence time (which, in turn, will affect  
2 the resultant horizontal distribution following advection (Darbyshire et al., 2018; Morgan et al.,  
3 2019)). High concentrations of trace gases and aerosols from these fires can be transported far  
4 from the source regions. This intercontinental transport has the potential to affect the  
5 atmospheric composition of regions considered as aerosol-free areas.

6 The southwest Indian Ocean (SWIO) basin is known to be one of the few pristine regions on  
7 Earth where the aerosol concentration is mainly governed by sea salts (Dufлот et al., 2022). The  
8 SWIO basin is characterized by a wet season (December to April) and a dry season (May to  
9 November). Previous works showed that the atmospheric composition over the SWIO region  
10 during the dry season is driven by the Southern Hemisphere BB activity (Clain et al., 2009;  
11 Dufлот et al., 2022; Edwards et al., 2006; Kaufman et al., 2003, Swap et al., 2003). These studies  
12 pointed out that BB plumes cross South Africa during the dry season. Edwards et al. (2006)  
13 revealed that southern African BB emissions mostly find their way into the SWIO basin and  
14 follow the five transportation modes identified by Garstang et al. (1996). Being located in the  
15 subtropical southern Indian Ocean at the crossroads of the transport pathway bringing air  
16 masses from southern Africa, Reunion Island (21,0°S, 55.5°E) is a favorable location to study  
17 the effect of this regional transport on atmospheric composition over the SWIO basin. Based  
18 on ozone radiosonde and ground-based lidar observations recorded at Reunion, Clain et al.  
19 (2009) highlighted a significant annual increase of tropospheric ozone over Reunion Island  
20 during August-November period, in phase with the BB season in southern Africa and  
21 Madagascar. High concentrations of ozone precursors from these fires are vented into the free  
22 troposphere by convection and are subsequently advected into the SWIO basin by westerly  
23 winds. In addition to regional transport, the tropical tropospheric composition over the SWIO  
24 basin can be modulated by the long-range transport of BB plumes from South America (Dufлот  
25 et al., 2010, 2022; Zhou et al., 2018). By combining ground-based observations of Carbon  
26 monoxide (CO) from a Fourier Transform Infrared (FTIR) spectrometer installed at Reunion  
27 and the FLEXPART model simulations, Dufлот et al. (2010) showed that southern African and  
28 southern American BB events have the potential to inject large amounts of ozone precursors  
29 such as CO and aerosols throughout the troposphere over the SWIO basin. The synergy of CO  
30 and aerosols observations are helpful in discussions of the influence of BB events on the  
31 evolution of the aerosol burden (Bègue et al., 2021; Bencherif et al., 2020; Jones et al., 2001).  
32 Most recently, the analysis of Aerosols Optical Depth (AOD) recorded from sun-photometer at  
33 Reunion over a period of 12-years has been undertaken by Dufлот et al. (2022). Dufлот et al.  
34 (2022) showed that the BB activity explains 67 % of the variability of the AOD, within which



1 the contributions of the BB activity in Southern Africa and southern America are estimated at  
2 22% and 20%, respectively. Although Australia is known for its intense BB events (Fromm et  
3 al., 2006; 2010; De Laat et al., 2012), the contribution of the Australian BB activity on the  
4 observed AOD variability over Reunion is relatively low (4.7 %).

5 The Australian BB activity takes mainly place in the northern part of the continent between  
6 September and January (Russel-Smith et al., 2007). Nevertheless, the most disastrous fires burn  
7 in the southeastern region of Australia. Extreme fires occurring over southeastern Australia  
8 during the austral summer can lead to the development of pyro-convection events (Dowdy and  
9 Pepler, 2018; Fromm et al., 2010). Southeastern Australia has a long history of severe pyro-  
10 convection events which have significantly impacted the composition of the stratosphere at  
11 regional and global scales. The last and largest event occurred during the 2019-20 fire season  
12 (referred to as “Black Summer”). Previous works reported that this fire season is unrivaled with  
13 a surface burnt estimated at 19 million hectares (Boer et al., 2020; Cai et al., 2022; Levin et al.,  
14 2021). Furthermore, this extreme fire season can be considered as unprecedented due to  
15 persistent planetary-scale perturbations induced in the stratosphere (Kablick et al., 2020;  
16 Khaykin et al., 2020; Kloss et al., 2021; Ohneiser et al., 2020, 2022; Yu et al., 2020).

17 Kablick et al. (2020) showed at least 18 pyro-convection outbreaks occurred between 29<sup>th</sup>  
18 December 2019 and 12<sup>th</sup> January 2020 with the largest event occurring around 1<sup>st</sup> January 2020.  
19 Based on satellite observations and CESM-CARMA model results, Yu et al. (2021) showed  
20 that Australian wildfires burning from December 2019 to January 2020 injected approximately  
21 0.9 Tg of smoke into the stratosphere. The analysis of the numerical simulations suggest that  
22 the smoke mass contained 2.5% black carbon which induced a 1 K warming in the stratosphere  
23 of the Southern Hemisphere mid-latitude for more than 6 months following its injection. The  
24 carbon-rich aerosol clouds were confined during their solar-driven rise by a persistent synoptic-  
25 scale anticyclone (Kablick et al., 2020; Khaykin et al., 2020). As a consequence, the highly-  
26 concentrated absorbing aerosols were lofted into the middle stratosphere, which prolongs their  
27 stratospheric residence time and radiative effect on climate. The combustion products injected  
28 into the stratosphere were advected by westerly winds and dispersed across all of the southern  
29 hemisphere extra-tropical latitudes (Khaykin et al., 2020; Ohneiser et al., 2020, 2022; Tencé et  
30 al., 2022; Yu et al., 2020). The stratospheric smoke layer was rapidly advected westward and  
31 its optical characteristics were measured by the lidar system installed at Punta Arena (53.2°S,  
32 70.9°E; Chile) and Rio Grandé (53.8° S, 67.7° W, Argentina) one week following its injection  
33 (Ohneiser et al., 2022). Ohneiser et al. (2022) pointed out that the presence of the smoke layers  
34 extended, on average, from 9 to 24 km in height, with one part ascending to more than 30 km



1 height as a result of self-lofting processes. The maximum smoke AOD was around 1.0 over  
2 Punta Arenas in January 2020 and thus 2 to 3 orders of magnitude above the stratospheric  
3 aerosol background of 0.005 (Ohneiser et al., 2022). Their results suggest an influence of this  
4 stratospheric smoke layer on the record breaking ozone hole over Antarctica in September-  
5 November 2020. This is consistent with the results reported by Tencé et al. (2022) from lidar  
6 and ozonesondes measurements at the French Antarctic station Dumont d'Urville  
7 (66.6°S,140.0°E). Tencé et al. (2022) pointed out the persistence of smoke aerosols in the  
8 southern high-latitude stratosphere following the pyro-convection events. They also reported  
9 that the 2020 stratospheric ozone depletion is above the decadal average at Dumont d'Urville.  
10 The common point among these studies is their interest for the perturbation induced by the  
11 2019-20 Australian fires on the stratospheric composition, and on the dynamical circulation  
12 over the extra-tropical latitudes. In contrast, relatively little attention has been paid to the  
13 perturbation induced by this Australian smoke layer over the tropical/subtropical latitudes.

14 This study has two aims: 1) document the transport of the Australian smoke layer in the southern  
15 subtropics over Indian Ocean during the January-February period which correspond to the wet  
16 season, and 2) to investigate the influence of Australian aerosol smoke layers on variability of  
17 the aerosol optical properties and CO in the Upper Troposphere-Lower Stratosphere (UT-LS)  
18 of the SWIO basin accounting for convective activity. Convective activity is more intense  
19 during this season as the Inter-Tropical Convergence Zone (ITCZ) is present over the whole  
20 basin (Lashkari et al., 2017). Furthermore, the convective activity is often synonym of tropical  
21 depression which can reach the stage of tropical cyclone (Barthe et al., 2021; Neuman et al.,  
22 1993). The Regional Specialized Meteorological Centre (RSMC) at Reunion reported that the  
23 cyclonic season 2019-20 was characterized by the development of 6 tropical cyclones and 4  
24 tropical storms in the SWIO basin.

25 The study is organized as follows: Section 2 describes the observations and the model used for  
26 the investigation of the transport of the aerosol smoke layer. A review of the formation and the  
27 transport of the Australian aerosol smoke layer over the Southern Hemisphere is provided in  
28 Section 3. An analysis of the influence of the Australian aerosol smoke layer on the aerosol and  
29 CO variability over the SWIO basin is given in Section 4. The discussion on the influence of  
30 the convective activity on the transport of the aerosol smoke layer over the SWIO basin is  
31 provided in Section 5. A summary and the perspectives of this study are given in Section 6.

## 32 **2. Instrumentation and Model description**

### 33 **2.1 Aerosols data sets**



1 The aerosol datasets used in this study resulted mainly from two ground-based observations  
2 sites from the Network for the Detection of Atmospheric Composition Change (NDACC,  
3 [www.ndacc.org](http://www.ndacc.org)) network as well as a suite of spaceborne sensors products.

#### 4 **2.1.1 Lauder ground-based lidar**

5 Aerosol optical properties measurements at Lauder (45.0°S; 169.7°E) have been performed  
6 using lidars since 1992. A detailed description of the lidar system operating at Lauder is given  
7 by Sakai et al. (2016), and it is summarized hereafter. The emitter system consists of a Nd:YAG  
8 laser beam at 532 nm and is linearly polarized. The lidar detects Rayleigh-Mie backscattering  
9 at 532 nm with parallel and perpendicular components. The methodology described by Fernald  
10 et al. (1984) is used to obtain the extinction and backscatter coefficient from a Rayleigh-Mie  
11 lidar. This methodology involves the use of an extinction-to-backscatter ratio for aerosol, also  
12 called lidar ratio (LR). For the January-May 2020 period, the values of LR derived with the  
13 lidar, are 88 and 60 sr for altitude above and below 23 km respectively. These values are  
14 obtained from the attenuation of the lidar signals by using the methodology described by Uchino  
15 et al. (1983) and by Young (1995). The aerosol depolarization is computed from the backscatter  
16 coefficient and the total linear volume depolarization ratio (Sakai et al., 2003). This latter is  
17 obtained by taking the ratio of the perpendicular to total components of the backscattered signal  
18 at 532 nm. To investigate the aerosol variability induced by the Australian fires, it is needed to  
19 define a background profile. The background profile is built from measurements performed  
20 when atmosphere is not undergoing major disturbance (e.g. volcanic eruptions, pyro-convection  
21 outbreaks). In the present work, the background extinction profile at Lauder is built from  
22 measurements performed between 1997 and 2004.

#### 23 **2.1.2 Reunion Island ground-based lidars**

24 The Atmospheric Physics Observatory of La Réunion (OPAR) is a permanent station for long  
25 term atmospheric observations (Baray et al., 2013). In particular, two lidar systems operating  
26 in the UV (355 nm) and visible part of the light spectrum (532 nm) are used to retrieve ozone  
27 and aerosols profiles at the Maïdo Observatory situated at 2200 m above mean sea level  
28 (AMSL). These systems, named LiO3T (532 nm) and LiO3S (355 nm) are described by Dufлот  
29 et al. (2017) and by Baray et al. (2006), respectively. These lidars measure aerosol optical  
30 properties (extinction, backscatter ratio) from ~15 km to the middle stratosphere, up to ~35 km,  
31 with a high vertical resolution (15 m). Operating at two distinct wavelengths, the synergy of  
32 Reunion lidar profiles allows us to assess the Angström exponent of aerosols between 355 nm  
33 and 532 nm. In this study, the inversion process used the Klett method (Klett, 1985) with an  
34 assumed lidar ratio of 60 sr, typical of aged BB aerosols (Müller et al., 2007). In the present



1 study we used 9 lidar profiles recorded during the January-March 2020 period. The background  
2 extinction profile at Reunion is built from measurements performed between 2017 and 2019  
3 excluding the perturbation induced by the Calbuco eruption in April 2015 (Bègue et al., 2017).

### 4 **2.1.3 CALIOP**

5 Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) is a nadir pointing lidar orbiting  
6 the Earth onboard the Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation  
7 (CALIPSO) satellite since 2006. CALIOP is a two-wavelength polarization-sensitive lidar (532  
8 and 1064 nm) that measures total attenuated backscatter vertical profiles with altitude-varying  
9 vertical (30–300 m) and horizontal (300–5000 m) resolution. In the present study, we used  
10 CALIOP product version 3.3 level 1B which includes calibrated attenuated backscatter along  
11 with collocated meteorological information provided by the National Aeronautics and Space  
12 Administration Global Modeling and Assimilation Office (GMAO). These data are  
13 postprocessed using a treatment described and validated by Vernier et al. (2009). The Scattering  
14 Ratio (SR) profiles used for the detection of the smoke plume are calculated following the  
15 methodology described by Khaykin et al. (2018). As a first step, the collocated GMAO data is  
16 used to correct the backscatter profiles of molecular attenuation and ozone absorption. Then,  
17 the SR is calculated as the ratio of total and molecular backscatter coefficients, with the latter  
18 derived from GMAO air density. The SR profiles are recalibrated at 36-39 km, following the  
19 methodology given by Vernier et al. (2009). The data with depolarization larger than 30 % are  
20 discarded to the treatment in order to avoid aliasing cirrus clouds above the thermal tropopause.  
21 The CALIOP data were obtained from the ACDISC data archive (<ftp://acdisc.gsfc.nasa.gov>)  
22 hosted by NASA Goddard Space Flight Center.

### 23 **2.1.4 Sun photometry measurements: Sky radiometer**

24 A sky radiometer is a scanning sun-sky photometer able to perform measurements of direct sun  
25 and diffuse sky irradiance under clear sky conditions, at seven wavelengths (between 315 and  
26 1020 nm) and at several scattering angles. The direct solar extinction and diffuse sky radiance  
27 measurements are used to derive the aerosol optical properties such as Aerosol Optical Depth  
28 (AOD), Single Scattering Albedo (SSA) and aerosol size-distributions using the algorithm  
29 developed by Nakajima et al. (1996). A detailed description of the sky radiometer and the  
30 associated data retrieval is given by Hashimoto et al. (2012). In the present study, we used the  
31 sky radiometer observations performed at Lauder in the framework of the SKYNET network.  
32 SKYNET is a ground-based network of sky radiometers with observation sites spread over Asia  
33 and other areas. Previous works have shown that the AOD from SKYNET is obtained with high  
34 accuracy similar with that of the standard Langley method and with those from AERONET





1 (Campanelli et al., 2007; Che et al., 2008). At Lauder, the sky radiometer measurements have  
2 been made since 2011. These observations are used in the present study to investigate the  
3 aerosol variability induced by the passage of the Australian BB plume over Lauder. The  
4 background evolution of AOD is built with measurements performed between 2011 and 2018.  
5 The sky radiometer data used in this work are available on the SKYNET platform:  
6 <https://www.skynet-isdc.org/>

### 7 **2.1.5 OMPS-LP**

8 The Ozone Mapper and Profiler Suite Limb profiler (OMPS-LP) has been flying on the Suomi  
9 National Polar Partnership (NPP) satellite platform since October 2011. In the present study,  
10 we use aerosol extinction profile from the NASA OMPS data product version 2.0 (Taha et al.,  
11 2021). The aerosol extinction profiles are retrieved from the limb scattering solar radiation. The  
12 V2.0 algorithm uses OMPS-LP measurements at wavelengths 510, 600, 675, 745, 869 and 997  
13 nm, selected to minimize the effect of gaseous absorption (Taha et al., 2021). Aerosol extinction  
14 measurements are provided from 10 to 40 km altitude on a 1 km vertical grid. A near-global  
15 coverage is produced within 3-4 days. The OMPS data are used in the present study to  
16 investigate the global transport of aerosol BB plume and its influence on the aerosol variability  
17 over Reunion. As recommended by Taha et al. (2021), we use aerosol extinction measurements  
18 at 745 nm. The background extinction profile is built with measurements performed from 2012  
19 to 2014 and from 2016 to 2018. These periods are chosen in order to discard the perturbation  
20 induced by the Calbuco eruption (Bègue et al., 2015). The OMPS data are downloaded from:  
21 <https://ozoneaq.gsfc.nasa.gov/>.

22 The aerosol absorbing index (AAI) data from OMPS are also used to describe the transport of  
23 the aerosol BB plume. This index allows the detection of absorbing aerosols through the  
24 spectral difference between a given pair of UV wavelength. When its value is positive, it  
25 indicates the presence of UV-absorbing aerosols such as dust and smoke. Conversely, a  
26 negative value indicates the presence of non-absorbing aerosols while values close to zero are  
27 found in the presence of clouds. The AAI data used in this work are available on the NASA  
28 Earth Data platform: <https://earthdata.nasa.gov/earth-observation-data>

## 29 **2.2 CO and water vapor measurements**

### 30 **2.2.1 FTIR**

31 The total columns and volume mixing ratio profiles of trace gases such as CO are retrieved with  
32 high accuracy and precision with ground-based Fourier Transform Infrared (FTIR)  
33 spectrometers (Clerbaux et al., 2008; Vigouroux et al., 2015; Zhou et al., 2019). In the present  
34 study, we use the FTIR observations performed at Lauder and Reunion sites in the framework





1 of the NDACC and Total Carbon Column Observing Network (TCCON) networks  
2 respectively. A detailed description of the FTIR systems involved in both networks and the  
3 associated data retrieval is given by de Mazière et al. (2018) and Wunch et al. (2015). A brief  
4 description is given hereafter. The CO measurements from FTIR at Lauder have been made  
5 since the early 1990s. The measurements are made using a Bruker high-resolution spectrometer  
6 over a wide spectral range (around 600–4500  $\text{cm}^{-1}$ ). The CO dataset used in this study is the  
7 same as that used by Bègue et al., (2021) and Kloss et al., (2019). Details on the spectral  
8 measurements, CO retrieval strategy and derived CO column abundances can be found within  
9 these references. The CO total columns and volume mixing ratio profiles on 48-layer  
10 atmosphere (0.37–100 km asl) used in this study for the Lauder site are downloaded from the  
11 NDACC website (<http://www.ndacc.org>).

12 The FTIR measurements at Reunion have been routinely performed in the framework of the  
13 TCCON network since 2011. A Bruker high-resolution spectrometer over a wide spectral range  
14 is also used at Reunion. The CO and O<sub>2</sub> total columns are simultaneously retrieved by using the  
15 GGG2014 code (Wunch et al., 2015). Column-averaged dry-air mole fraction of CO are  
16 obtained from the retrieval of these two components. TCCON uses the O<sub>2</sub> total to determine  
17 the total column of the dry air. Then, the column-averaged dry-air mole fraction of CO is  
18 calculated as the ratio between the retrieved CO total columns and the total columns of the dry  
19 air. The abundance of CO used in the study for Reunion are downloaded from the TCCON  
20 database ([http://www.ndaccdemo.org/about/cooperating-networks/total-carbon-column-  
21 observing-network](http://www.ndaccdemo.org/about/cooperating-networks/total-carbon-column-observing-network)). In the present study, the background evolution of CO is built with  
22 measurements performed between 2015 and 2018 (De Mazière et al., 2017).

### 23 **2.2.2 IASI**

24 The Infrared Atmospheric Sounding Interferometer (IASI) measures chemical species such as  
25 CO by using a Fourier Transform spectrometer (Clerbaux et al., 2009; Coheur et al., 2009). It  
26 is flying onboard the three Metop satellites. Retrieval of CO total and partial columns occurs in  
27 near real-time from the nadir radiances measured by the instrument in the thermal infrared  
28 radiation covering wavelengths from 6.62 to 15.5  $\mu\text{m}$ . Global distributions are obtained for day  
29 and night measurements, with a vertical range covering the troposphere and the lower  
30 stratosphere. The total and partial column of CO are retrieved by using the Fast-Optimal  
31 Retrievals on Layers for IASI (FORLI-CO, Hurtmans et al., 2012). Furthermore, the  
32 contamination of the data by clouds is checked and flagged. In the present study, we used CO  
33 columns from IASI instruments on Metop-A and Metop-B, which have been operating since



1 2006 and 2012, respectively. The IASI products used in this work are available on the AERIS  
2 platform: <https://iasi.aeris-data.fr/CO>.

### 3 **2.2.3 MLS**

4 The Microwave Limb Sounder (MLS) performs vertical profile measurements of multiple trace  
5 gases in the UT-LS onboard Aura satellite since 2004 (Waters et al., 2006). In the present study,  
6 the CO and water vapor observations (version 5) from January 2017 to January 2020 over a  
7 global domain extending between 10°S and 25°S in latitude and 30°E and 60°E in longitude have  
8 been used. All MLS version 5 retrieval quality flags (quality, status, convergence, and  
9 precision) were properly adhered to for all of our analyses (Livesey et al., 2020). Generally,  
10 recommended pressure levels for science applications with CO and water vapor MLS data range  
11 from 0.0215 to 215 hPa ([http://mls.jpl.nasa.gov/data/v4-2\\_data\\_quality\\_document.pdf](http://mls.jpl.nasa.gov/data/v4-2_data_quality_document.pdf)). The  
12 CO and water vapor profiles from MLS are obtained from the Atmospheric Composition Data  
13 and Information Services Center (ACDISC) archive (<ftp://acdisc.gsfc.nasa.gov>) hosted by the  
14 NASA Goddard Space Flight Center.

## 15 **2.3 Numerical Modelling**

### 16 **2.3.1 FLEXPART Model**

17 The Lagrangian transport and diffusion model FLEXPART version 10.4 is used to simulate  
18 long-range transport of atmospheric tracers (Pisso et al., 2019; Stohl et al., 2005). This version  
19 of FLEXPART includes improvements in different aspects such as physico-chemical  
20 parameterizations (Pisso et al., 2019). Source identification occurs via the release of particles  
21 from a receptor location and the simulation of backward trajectories. Model calculations are  
22 based on ERA5 (Hersbach et al., 2020) meteorological data from the European Center for  
23 Medium-Range Weather Forecasts (ECMWF) extracted at 3-hourly intervals with a horizontal  
24 resolution of  $0.5^\circ \times 0.5^\circ$  and a vertical resolution of 137 hybrid model levels (from the ground  
25 to 0.01 hPa pressure altitude).

26 The model simulations are run with the aerosol (BC and OC) and CO tracers assuming removal  
27 by dry and wet deposition for aerosols and OH reactions for CO. Each simulation consists of  
28 20,000 particles released over Reunion during one day at altitudes between 15 and 19 km every  
29 0.5 km and followed backward in time during one month. The simulation of backward  
30 trajectories with FLEXPART for a long period (1-2 months) were previously explored in  
31 previous studies (Aliaga et al., 2021; Eckhardt et al., 2017; Xu et al., 2021). The simulations  
32 include the parameterization of turbulence and the activation of the convection. FLEXPART  
33 model outputs are distributed over a regular vertical grid of  $0.5^\circ \times 0.5^\circ$  from ground to 25 km in  
34 altitude. The model outputs are used to discuss the residence time of the BB aerosols and CO,



1 and their contributions on the variability of the aerosol optical properties and CO over the SWIO  
2 basin. Discussions are based on the analysis of the emission sensitivity obtained from backward  
3 simulations. The residence time of particles are integrated over the entire atmospheric column  
4 and over the latitude to create averaged-map and longitudinal cross-section map and to provide  
5 information on the geographical and vertical dispersion of BB aerosols in the atmosphere. The  
6 BB contributions on the vertical distribution of CO and the aerosol optical properties can be  
7 calculated by combining the potential emission sensitivity (PES) with an emission inventory.  
8 PES represents FLEXPART particles only in the layer/altitude at which the emissions are  
9 injected. Pyro-convection is not taking into account in the model (nor ECMWF data). Then, a  
10 mass concentration profile of BB aerosol and CO is extracted by summing all the output grid-  
11 points. A layer between 0 and 3 km is used for the BB in Africa (as observed in GFAS “top  
12 altitude of plume”, Kaiser et al., 2012) and between 9 and 16 km for the Australian fires (as  
13 observed by CALIOP observations)

14 In the present study, the Global Fire Assimilation System (GFAS) version 1.2 emission (Kaiser  
15 et al., 2012) has been used for the calculation. In addition, the Global Air Pollutant Emissions  
16 - EDGAR v6.1 emission inventory (Kaiser et al., 2012, <http://edgar.jrc.ec.europa.eu>) is used for  
17 CO for the year 2018 with a  $0.1^\circ \times 0.1^\circ$  grid. These emissions represent the total CO emissions  
18 by anthropogenic activities by excluding large scale BB with Savannah burning, forest fires.  
19 As for the BB emissions, multiplying the CO emission flux from this inventory with the  
20 FLEXPART emission sensitivity for a layer between 0 and 1 km gives access to the contribution  
21 of anthropogenic sources to the total CO abundance. At the end of the process, in the case of  
22 aerosols, the mass concentration profile is converted in extinction profile in order to discuss on  
23 the contribution of BB aerosols on the vertical distribution of the aerosol extinction observed  
24 from Lidar over Reunion. The conversion is performed by the use of Mie scattering model  
25 assuming spherical particles with a density of  $2 \text{ g.cm}^{-3}$  and a refractive index of  $2.0 + 0.64i$   
26 adapted to optically absorbing aerosols.

### 27 **2.3.2 MIMOSA Model**

28 The Modèle Isentropique de transport Mésoéchelle de l’Ozone Stratosphérique par Advection  
29 (MIMOSA) model is a potential vorticity (PV) advection model running on isentropic surfaces  
30 at a resolution of  $0.3^\circ \times 0.3^\circ$  (Hauchecorne et al., 2002). The advection scheme is semi-  
31 Lagrangian with a time step of 1 h and driven by ECMWF meteorological analyses. The model  
32 can be run continuously in order to follow the evolution of PV filaments for several months.  
33 The accuracy of the model has been evaluated by Hauchecorne et al. (2002) and validated  
34 against airborne lidar ozone measurements using a correlation between PV and ozone, a quasi-



1 conserved chemical tracer on timescales of a week or so within most of the lower stratosphere  
2 (Heese et al., 2001). The MIMOSA model can also be used to determine the origin of air masses  
3 influencing a given site, similar to an isentropic Lagrangian trajectory model (Bencherif et al.,  
4 2011; Hauchecorne et al., 2002; Portafaix et al., 2003; Bègue et al., 2017).

### 5 **3. Transport of the Australian BB plume over the Southern** 6 **Hemisphere**

#### 7 **3.1 Formation of an intense stratospheric smoke layer**

8 Following the strongest outbreak during New Year's Eve, a wide plume of BB aerosol with  
9 large values of AAI (higher than 12) is transported toward the Tasman Sea on 1<sup>st</sup> January 2020  
10 (Fig. 1a). Figure 1b depicts the CALIOP attenuated scattering ratio (SR) profiles on 1<sup>st</sup> January  
11 2020 above New-Zealand. The CALIOP attenuated SR profiles are calculated along the  
12 CALIOP track (blue line in Fig. 1a) crossing the absorbing aerosol plume. CALIOP  
13 observations reveal a broad region of high values (ranging from 10 to 25) between 36° S and  
14 46°S centered at 16.5 km altitude (Fig. 1b). Khaykin et al. (2018) reported similar values during  
15 the passage of the 2017 Canadian smoke plume over Observatoire de Haute-Provence (OHP,  
16 France). Figure 1 reveals that New Zealand was significantly exposed to the transport of the  
17 Australian BB plume. The ground-based measurements recorded in New-Zealand are helpful  
18 to investigate the evolution of the aerosol smoke. Figure 2a illustrates the daily extinction  
19 profiles at 532 nm derived from lidar measurements over Lauder between 1<sup>st</sup> December 2019  
20 and 1<sup>st</sup> April 2020. Note that a strong convective activity prevented lidar operations between  
21 mid-December 2019 and the 1<sup>st</sup> January 2020. Figure 2a reveals a sharp increase in the  
22 extinction in the stratosphere over Lauder starting from mid-January 2020 with values ranging  
23 from  $3 \times 10^{-3} \text{ km}^{-1}$  to  $9 \times 10^{-3} \text{ km}^{-1}$ , one order of magnitude above the typical stratospheric  
24 aerosol background (Vernier et al., 2012). The vertical extent of the plume increased  
25 significantly between mid-January and 1<sup>st</sup> April 2020 with an aerosol layer spanning from 11.5  
26 to 20 km. The ascent of the aerosol plume could be attributed to efficient adiabatic heating as a  
27 result of the strong absorption of solar radiation by this black carbon rich plume. This is  
28 consistent with the results found by De Laat et al. (2012) during the Australian Black Saturday  
29 fires occurring on February 2009. They showed a strong ascent of the BB aerosol due to  
30 adiabatic heating effect. Figure 2b depicts the daily evolution of CO mixing ratio profile  
31 obtained from FTIR measurements over Lauder between 1<sup>st</sup> December 2019 and 1<sup>st</sup> April 2020.  
32 Prior to the convective period, the maximum of CO mixing ratio (120-130 ppbv) is observed in  
33 the troposphere. An increase of CO mixing ratio in the lower stratosphere is visible from mid-



1 December 2019. One can observe that the maximum of the CO mixing ratio (50-90 ppbv) is  
2 mainly observed in the UT-LS (9-13 km) during the February-April 2020 period.  
3 The injection of BB aerosols and CO in the stratosphere induced significant disturbance evident  
4 in the total columns over Lauder (Fig. 3a). Figure 3a depicts the monthly mean evolution of  
5 AOD and total columns of CO (TCO) at Lauder between 1<sup>st</sup> December 2019 and 1<sup>st</sup> April 2020  
6 obtained from sky-radiometer and FTIR measurements respectively. AOD reaches its  
7 maximum value (0.17, 3 times higher than background value) in January 2020, decreasing to  
8 background values in February 2020 (Fig. 3a). The similar evolution is also observed for the  
9 TCO values. An abrupt increase in TCO (~9 % of the pre-event levels) is observed in January  
10 2020 and the return to pre-event values is observed as of February 2020 as also already shown  
11 in Kloss et al. (2021). Conversely, the perturbation on the stratospheric columns still persisting  
12 after February 2020. Figure 3b illustrates the monthly mean evolution of stratospheric AOD  
13 (sAOD) and CO columns (sCO) at Lauder between 1<sup>st</sup> December 2019 and 1<sup>st</sup> April 2020. The  
14 sAOD and sCO are calculated between 12 and 30 km from lidar and FTIR measurements,  
15 respectively. Figure 3b indicates that the evolution of sAOD and sCO is fairly similar. A  
16 statistically significant increase of sAOD is observed in January 2020 (2.5 times higher than  
17 background value) and still visible in April 2020 with same amplitude. sCO reaches its  
18 maximum values (~24 % of the pre-event levels) in January 2020 and slightly decreases in April  
19 2020 (~14% of the pre-event). This result is consistent with results found by Khaykin et al.  
20 (2020). Based on the MLS observations, Khaykin et al. (2020) showed that the stratospheric  
21 mass of CO within the southern extra-tropics increased abruptly by around 20% compared to  
22 the observed value during the pre-event period.  
23 Our results agree with previous studies (Kablick et al., 2020; Khaykin et al., 2020, Kloss et al.,  
24 2021) and confirm a persistent perturbation on the atmospheric composition of the extra-  
25 tropical stratosphere. In order to extend the discussion, the spatial dispersion of the smoke layer  
26 will be discussed in the next section.

### 27 **3.2 Spatial dispersion of sAOD and CO**

28 Figure 4 depicts a time-averaged map of partial columns of aerosols (between 16 and 30 km,  
29 sAOD) and CO (between 9 and 30 km) obtained from OMPS and IASI observations between  
30 9<sup>th</sup> and 16<sup>th</sup> January 2020, respectively. One can observe the correlation in the shape and spatio-  
31 temporal distribution between CO and sAOD over the Southern Pacific (Figs 4a and 4b). It is  
32 worthwhile mentioning that the transport of the aerosol (with values ranging from  $6 \times 10^{-3}$  to  $1$   
33  $\times 10^{-2}$  km<sup>-1</sup>) and CO (with values ranging from  $6$  to  $8 \times 10^{17}$  molecules. cm<sup>-2</sup>) plume over the  
34 Southern Pacific occurred mainly within the 18°S–60°S latitudinal band. The plume's structure



1 over the Southern Pacific can be explained by the fact that the core plume started to encapsulate  
2 into a compact bubble-like structure (with 4 km vertically extended and 1000 km horizontally  
3 wide) on 7<sup>th</sup> January 2020 (Khaykin et al., 2020). Figures 4a and 4b reveals that the aerosol and  
4 CO plume cross the Southern Pacific fairly quickly. Based on CALIOP and Lidar observations  
5 at Punta Arenas (53.2°S, 70.9°E, Chile), Ohneiser et al. (2020) showed that aerosols injected  
6 into the lower stratosphere over southeastern Australia crossed the Pacific Ocean within 10  
7 days. On 10<sup>th</sup> January 2020, the aerosol plume is already visible over Punta Arenas and  
8 structured into two layers. One can observe an aerosol band (with values ranging from  $5 \times 10^{-3}$   
9 to  $9 \times 10^{-3} \text{ km}^{-1}$ ) across the Southern Hemisphere between 40°S and 60°S during the 9-16<sup>th</sup>  
10 January 2020 period (Figure 4a). Therefore, the Australian aerosol plume has already circled  
11 the Southern Hemisphere during the first two weeks of January 2020. Based on MLS and  
12 CALIOP observations, Khaykin et al. (2020) showed that the smoke layer dispersed across the  
13 whole Southern Hemisphere extra-tropics in less than two weeks whereas the carbon-rich core  
14 remained bounded within the mid-latitudes (30°S-50°S). The fastest patches of smoke returned  
15 over Australia by 13<sup>th</sup> January 2020 (Khaykin et al., 2020). The same conclusion cannot be  
16 made for CO from space-borne observations (Figure 4b). One can observe weak values of CO  
17 (less than  $5 \times 10^{17} \text{ molecule.cm}^{-2}$ ) over southern Atlantic and without real link with the large  
18 plume observed over southern Pacific (Fig. 4b). Except for the bubble structure, it seems that  
19 the main part of the Australian BB plume has been transported bounded within the extra-tropical  
20 latitudinal band. Further discussions the dispersion of the Australian smoke layer over the  
21 SWIO basin will be given in the next section.

## 22 **4. Presence of aerosols smoke layer over the SWIO basin**

### 23 **4.1 Aerosol and CO variability over a subtropical site: Reunion**

24 Figure 5a depicts the evolution of the sAOD at 532 nm calculated between 15 and 30 km from  
25 the ground-based lidars (LiO3T and LiO3S) and OMPS observations over Reunion from 1<sup>st</sup>  
26 January to 1<sup>st</sup> March 2020. The Angström exponent for the 532–745 nm wavelength pair is  
27 adopted from methodology in Taha et al. (2021) and set to 1.9. lidar observations over Reunion  
28 are also used to calculate the aerosol Angstrom exponent (using 355 nm and 532 nm pairs). An  
29 abrupt increase in the aerosol loading is clearly observed over Reunion as of 16<sup>th</sup> January 2020  
30 according to satellite observations. This increase of the aerosol loading, three times above the  
31 typical background, was still visible until 1<sup>st</sup> March 2020. It is worthwhile mentioning that the  
32 sAOD values observed between 16<sup>th</sup> January and 1<sup>st</sup> March 2020 are higher than those observed  
33 during the passage of the Calbuco plume over Reunion site, which did not exceed 0.013 (Bègue  
34 et al., 2017). The increase of sAOD in mid-January coincided with an increase of CO, as shown



1 in Figure 5b based on the use of total columns and CO abundance from IASI and FTIR at the  
2 same site and over the same period. The evolution of CO obtained from IASI and FTIR  
3 measurements correlate. The ground-based and satellite observations clearly indicate a  
4 significant increase of CO with a peak observed in mid-January. The ground-based observations  
5 show that the CO abundance observed during this increased phase is on average 20% higher  
6 than the values observed during the background period (Fig. 5b). The evolution of sAOD and  
7 CO observations in mid-January suggests that Reunion, and its surrounding, have been  
8 influenced by the transport of smoke aerosol layers.

9 Figures 6a and 6b illustrate the night-averaged extinction profiles at 355 nm derived from lidar  
10 measurements over Reunion on January and February 2020, respectively. The two first weeks  
11 of January 2020 are representative of the January typical background (shaded area), as  
12 illustrated on 13<sup>th</sup> January 2020 (Fig. 6a). Conversely, the extinction profiles at the end of  
13 January 2020 (27<sup>th</sup> and 28<sup>th</sup>) are marked by a significant increase (4 times higher than the  
14 background values) located in the lower stratosphere between 16.8 and 18 km altitude  
15 (equivalent to potential temperature levels 380-404 K). One can observe that the structure of  
16 the extinction profile in the lower stratosphere has changed between these two days. On 28<sup>th</sup>  
17 January, the extinction profile exhibits a sudden increase at 17.4 km (~400 K) and quickly  
18 decreased afterwards to values observed the previous day (Fig. 6a). It is worthwhile mentioning  
19 that the values of extinction ( $10$  to  $17 \times 10^{-3} \text{ km}^{-1}$ ) observed in the lower stratosphere on these  
20 two days are of the same order as those observed at Lauder a few days after the pyro-convective  
21 event (Fig. 2). Figure 6a reveals a statistically significant increase (4 times higher than  
22 background values) in aerosol extinction between 15 and 16.5 km altitude (361-375 K), on 27<sup>th</sup>  
23 and 28<sup>th</sup> January 2020. Over Reunion, the lidar observations hence confirm the presence of a  
24 significant aerosol layer in the UT-LS by the end of January. In February, the extinction profiles  
25 clearly exhibit two significant aerosol layers with the first one located between 16 and 19.5 km  
26 (370-440 K) and the second one between 20 and 22.5 km (465-500 K) (Fig. 6b).

27 To further discuss the optical properties of these aerosol layers, the Angstrom exponent has  
28 been calculated between 355 nm and 532 nm from the LiO3S and LiO3T measurements (Figs.  
29 6c and 6d). In February, the Angström exponent values reveal that the two aerosol layers consist  
30 mainly of small aerosol particles (Fig. 6d), consistent with a stratospheric smoke layer (Haarig  
31 et al., 2018; Hu et al., 2019; Ohneiser et al., 2021). In January, the profile of Angström exponent  
32 exhibits more variability in the UT-LS (Fig. 6c) with values ranging from 0.6 to 1.9, on 27<sup>th</sup>  
33 and 28<sup>th</sup> January. Based on the literature, the largest proportion of the observed Angström  
34 exponent range (up to ~1.5) within the plume, points to the presence of an aerosol smoke layer





1 (Das et al., 2021; Haarig et al., 2018; Hu et al., 2019; Kloss et al., 2019). The wide range of  
2 Angström exponent values suggests that the aerosol layer is not homogeneously distributed at  
3 this stage and might be interpreted as a mixture of fresh and aged smoke layers (Fig. 6c). This  
4 may indicate growth and removal processes (e.g., coagulation, condensation, sedimentation)  
5 which can modulate the morphology and mixing state of the aerosol layer during its transport  
6 (Burton et al., 2015; Hamil et al., 1997). The residence time of the aerosol particles in the  
7 atmosphere depends on the balance between the growth processes and the removal processes,  
8 which are likely to be controlled by the dynamical context. Previous works showed that the  
9 dynamical context can modulate the structure and optical properties of the aerosol layer over a  
10 given site from day to day (Bègue et al., 2017; Kremser et al., 2016). Fresh aerosols can be  
11 rapidly transported and mixed with pre-existing aged aerosol. Given the fact that the Angström  
12 exponent values decrease with the duration of transport, we cannot exclude that the vertical  
13 distribution of optical properties of aerosol over Reunion may also be explained by the regional  
14 transport of air masses.

#### 15 **4.2 Origin of the air masses**

16 To analyze the origin of air masses at Reunion on 27<sup>th</sup> and 28<sup>th</sup> January, one-month backward  
17 trajectories are calculated using FLEXPART (Aliaga et al., 2021; Eckhardt et al., 2017; Xu et  
18 al., 2021). A period of one month is chosen because it refers to time separating the pyro-  
19 convective outbreak event and the day of the measurement at Reunion. The representation of  
20 the PES (potential emission sensitivity) from back-trajectories simulations initialized at 18 km  
21 originating from Reunion on 27<sup>th</sup> and 28<sup>th</sup> January 2020 are presented in Figure 7.

22 Figures 7A-1 and 7B-1 display the horizontal trajectory paths, whilst vertical movement is  
23 shown in Figures 7A-2 and 7B-2, respectively. The vertical transect of FLEXPART back  
24 trajectories in Figure 7A-2 confirms a high probability of air mass contribution from the  
25 Australia if the fires emissions are directly injected into the stratosphere by convection (black  
26 rectangle in the figure), (i.e. layer of 9 to 16 km of injection taken for the PES, see section  
27 2.3.1). Then, according to FLEXPART results, the air masses at 18 km over Australia moved  
28 westward and reached Reunion on 27<sup>th</sup> January. One can observe that the same pattern is  
29 observed on 28<sup>th</sup> January 2020 (Figs. 7B-1 and 7B-2). One part of the Australian smoke layer  
30 is advected zonally by the prevailing easterly winds and is observed over Reunion on 27<sup>th</sup> and  
31 28<sup>th</sup> January 2020 at 18 km. It is worthwhile mentioning that the same pattern has been observed  
32 during the volcanic eruption of the Hunga Tonga on January 2022 (Baron et al, 2023, Kloss et  
33 al., 2022; Sellitto et al., 2022). The FLEXPART simulations also suggest that Reunion is also  
34 influenced by eastward transport of air masses. This pathway is clearly visible on 28<sup>th</sup> January



1 2020 (Figs. 7B-1 and 7B-2). Figure 7B-1 reveals that air masses coming from the South  
2 American region and Australia region both reach the SWIO basin by passing over southern  
3 Africa. Furthermore, air masses from high latitudes seem to cross the subtropical latitudes  
4 following a wave shape and reach the SWIO basin by passing over the Cape of Good Hope  
5 (Figure 7B-1).

6 In order to improve the discussion on this eastward transport of air masses over the SWIO basin,  
7 the MIMOSA model has been used to produce a continuous evolution of PV fields for the period  
8 from 1<sup>st</sup> to 31<sup>st</sup> January 2020. Two advected PV maps derived for the 400 K isentropic level  
9 from the MIMOSA model are depicted in Figure 8. The localization of the aerosol plume  
10 obtained from OMPS observations at the 400 K  $\pm$  5 K isentropic level are also superimposed  
11 (Fig. 8). The 400 K isentropic level is chosen according to the layers observed in the extinction  
12 profiles over Reunion between 390 and 404 K isentropic level on 27<sup>th</sup> and 28<sup>th</sup> January 2020  
13 (Fig. 6a). Figure 8 reveals significant wave activity during these two days. It is clearly shown  
14 that air masses from mid-latitudes (40-60°S) cross the subtropical latitudes (20-40°S) and are  
15 advected eastward between South Africa and Madagascar following a wave shape. Given the  
16 Australian BB aerosol are mainly located in the mid-latitudes (Fig. 4a), we can reasonably  
17 conclude that the filament reaching the SWIO basin contains aerosol from Australian BB event.  
18 On 27<sup>th</sup> January, air masses containing aerosol are observed at Madagascar and its surroundings  
19 (Fig. 8a). These air masses are advected eastward following the displacement of the wave shape  
20 and reach Reunion, on 28<sup>th</sup> January (Fig. 8b). These results demonstrate that the unusual aerosol  
21 load observed in the UT-LS above SWIO is to be linked to the Australian fires. Parts of the  
22 smoke plume underwent an isentropic transport from the mid to tropical latitudes following two  
23 distinct pathways to reach Reunion.

24 Nevertheless, the contribution of the Australian BB event on the variability of aerosols in the  
25 upper troposphere (15-16.5 km; 361-375 K) remains ambiguous. The representation of the PES  
26 from back-trajectories simulations initialized at 16 km originating from Reunion for the 27<sup>th</sup>  
27 and 28<sup>th</sup> January 2020 are presented in Figure 9. Figure 9A-1 and 9B-1 reveal that the  
28 trajectories from Reunion at 16 km pass over southern Africa and Madagascar. One can observe  
29 that the highest values of PES are located over southern Africa and Madagascar (Figs. 9A-1  
30 and 9B-1). Furthermore, the FLEXPART simulations suggest that air masses from southern  
31 Africa and Madagascar might have reached altitudes up to 16 km between 25°E and 55°E in  
32 longitude and reached Reunion on 27<sup>th</sup> and 28<sup>th</sup> January 2020 (Figs. 9A-2 and 9A-2). The results  
33 of Figure 9 show an influence of air masses coming from Africa and reaching the SWIO basin  
34 at 16 km. Thus, the moderate increase in aerosol extinction observed in the upper troposphere



1 (between 16 and 17 km altitude) on 27<sup>th</sup> and 28<sup>th</sup> January in Figure 6a may be attributed to air  
2 masses from regional sources, namely southern Africa and Madagascar.

3 In summary, our analysis suggests that the variability of CO and aerosols in the UT-LS over  
4 the SWIO basin in January is explained both by long-range and regional transport of air masses.

5 The extent to which regional sources have contributed to the variability of the atmospheric  
6 compounds over the SWIO is now investigated.

## 7 **5. Discussion on the influence of the regional sources**

8 We investigate the contribution of the African and Australian BB activity on the atmospheric  
9 composition in the UT-LS over the SWIO basin. Because a significant simultaneous increase  
10 of CO and sAOD is observed over Reunion and its surroundings from 16<sup>th</sup> to 29<sup>th</sup> January 2020,  
11 the discussion will focus on this period. In order to assess the contribution of African and  
12 Australian BB activity, FLEXPART is coupled with the GFAS inventory.

### 13 **5.1 Presence of African air masses in the UT-LS**

14 The location of fire-flagged pixels and the associated Fire Radiation Power (FRP) values from  
15 MODIS between 16<sup>th</sup> and 29<sup>th</sup> January 2020 are reported in Figure 10a. FRP gives quantitative  
16 information on combustion rates and its intensity. The sparse activity of the African fires in  
17 January is clearly illustrated in Figure 10a with moderate values of FRP ranging from 20 to 200  
18 MW.m<sup>-2</sup>. These values are ten times lower than those observed over the southeastern Australia  
19 between 30<sup>th</sup> December 2019 and 12<sup>th</sup> January 2020 (Bègue et al., 2021). One can observe that  
20 the African BB activity in January 2020 is mainly located over the northwestern (near the  
21 Equator) and southeastern side of southern Africa. The most intense values (100-200 MW.m<sup>-2</sup>)  
22 are observed over the southeastern side.

23 Despite this sparse activity of BB, the amount of CO injected into the atmosphere is fairly  
24 significant and ranging from 5 to  $6 \times 10^{17}$  molecules.cm<sup>-2</sup>. Time-averaged map of the partial  
25 column of CO (between 9 and 30 km) from IASI observations between 16<sup>th</sup> and 29<sup>th</sup> January  
26 2020 over the SWIO basin is reported on Figure 10b. The partial column of CO over the  
27 southern Africa is characterized by two regions of high values (higher than  $5 \times 10^{17}$   
28 molecules.cm<sup>-2</sup>). The first region stretches between the eastern side of southern Africa and  
29 western side of Madagascar which corresponds to a domain extending between 10°S and 25°S  
30 in latitude and 30°E and 45°E in longitude (Fig. 10b). The second region is located on the  
31 opposite side, over a domain extending between 10°S and 15°S in latitude and 5°E and 15°E in  
32 longitude. These regions of high values of CO seem to be mainly related to BB activities located  
33 over the northwestern side of southern Africa (Fig. 10a). Although these BB activities are less  
34 intense than those observed over the southeastern side, they seem to have the potential to impact



1 the variability of CO between 9 and 30 km over the SWIO basin. This could be explained by  
2 convective activity which can rapidly transport air masses from the boundary layer to the upper  
3 troposphere (Pfister et al., 2010; Blamey and Reason, 2012). Previous works showed that,  
4 during the southern summer, the convective activity over southern Africa and the SWIO basin  
5 is triggered either from the ITCZ or tropical storms (Barimala et al., 2018; Héron et al., 2020;  
6 Lashkari et al., 2017).

7 Figure 10c depicts a time-averaged map of OLR (Outgoing Longwave Radiation) anomalies  
8 from NCEP analysis between 16<sup>th</sup> and 29<sup>th</sup> January 2020. One can observe that the main  
9 convective regions (region of negative OLR anomalies) are located in mainland Africa between  
10 12° S and 25° S and the northern side of the SWIO basin. It is worthwhile mentioning that the  
11 daily brightness temperature values obtained from MODIS between 16<sup>th</sup> and 29<sup>th</sup> January 2020  
12 (not shown) are ranging from 195 to 210 K over the eastern side of southern Africa and the  
13 Mozambique Channel (around the north tip of Madagascar). These values of brightness  
14 temperature can be attributed to deep convection clouds (Héron et al., 2020; Young et al., 2013).

15 In most tropical basins, the deep convection can be separated into local convection and tropical  
16 storms (Bovalo et al., 2012). In our case, the deep convection activity could be explained by  
17 the ITCZ activity moving southward, near the subtropical latitudes, during the austral summer  
18 (Barimala et al., 2018; Lashkari et al., 2017). Through the analysis of ERA-Interim data over a  
19 period of 66 years, Lashkari et al. (2017) investigated the annual and seasonal displacement of  
20 the ITCZ. On average, the motion of ITCZ over southern Africa on January is characterized by  
21 southward move from 5°N to 20°S in latitude occurring between 20° E and 35° E in longitude  
22 (Lashkari et al., 2017). Over the Mozambique Channel, the deep convection can be explained  
23 by the tropical storm activity. Indeed, a tropical depression has been formed in the east side of  
24 the Mozambique Channel (near the northwestern side of Madagascar) between 20<sup>th</sup> and 22<sup>nd</sup>  
25 January 2020. This tropical depression reached the stage of strong tropical storm on 24<sup>th</sup> January  
26 2020 and was called Diane by the RMSC (Regional Specialized Meteorological Center) of  
27 Reunion. The intensification of the tropical depression into strong tropical storm occurred  
28 around the north tip of Madagascar. Diane passed near Reunion on 25<sup>th</sup> January 2020 (Fig. 10c).

29 In the present study, the convective activity over southern Africa and the SWIO basin may  
30 hence be due to both ITCZ proximity and Diane activity. This convective activity also explains  
31 the gap of CO observations from satellites over Madagascar and the northern side of southern  
32 Africa (Fig. 10b).

33 To further discuss the vertical distribution of CO in the UT-LS within the region of deep  
34 convection, the vertical cross section of CO and water vapor mixing ratio anomalies calculated



1 from MLS observations between 16<sup>th</sup> to 29<sup>th</sup> January 2020 are analyzed (Figs. 11a and 11b).  
2 The CO and water vapor mixing ratio anomalies are calculated as a relative difference by  
3 considering the monthly background means as the reference values. The calculations are  
4 performed over a domain extending between 10°S and 25°S in latitude and 30°E and 60°E in  
5 longitude (black box in Fig. 10b). The altitude-longitude cross-section is averaged for all  
6 latitudes covering the study domain. This domain includes both the region of deep convection  
7 and the first region of high values of CO. The monthly background is calculated from available  
8 MLS observations between 2017 and 2019. The altitude-longitude cross-section is averaged for  
9 all latitudes covering the study domain. Figure 11a exhibits two regions of high values of CO  
10 mixing ratio anomalies (higher than 15%) centered at 37°E and 50°E in longitude. The  
11 maximum CO mixing ratio anomalies in the first region are centered at 215 hPa (~12 km) with  
12 anomalies ranging from 25% to 30%. The values of the anomalies decreased rapidly with  
13 altitude. Indeed, the anomalies values obtained at 146 hPa (~15 km) and 100 hPa (~17 km)  
14 ranged from 20% to 25% and from 15% to % respectively. In the second region, the maximum  
15 CO mixing ratio anomalies are also centered in the middle troposphere (215 hPa). One can  
16 observe that these regions of CO mixing ratio anomalies are in coincidence with two regions of  
17 high values (higher than 20%) of water vapor mixing anomalies (Fig. 11b). The maxima of  
18 water vapor mixing ratio anomalies are centered at 146 hPa with values ranging from 40% to  
19 50%. This coincidence suggests a significant perturbation of the vertical distribution of CO and  
20 water vapor mixing ratio induced by the convective activity over southern Africa and the SWIO  
21 basin. Given the localization of the Diane tropical storm over the SWIO basin, we can  
22 reasonably assume that it contributed significantly to impact the vertical distribution of CO and  
23 water vapor in the second region aforementioned. One can observe that fire flagged pixels (with  
24 FRP values range between 10 and 30 MW.m<sup>-2</sup>) are detected by MODIS over the East of  
25 Madagascar (Fig. 10a). We can assume that the convective activity induced by Diane near  
26 Madagascar may have contributed to lift air masses enriched in CO from the lower troposphere.  
27 This could explain the high values of CO mixing ratio anomalies observed in the second region.  
28 This is consistent with the FLEXPART simulations which highlight a lift of air masses from  
29 the lower troposphere to lower stratosphere between 25°E to 55° E in longitude (Figure 9). One  
30 can observe that the most significant anomalies of the vertical distribution of CO and water  
31 vapor mixing ratio stretch from the middle troposphere (215 hPa) up to the tropopause layer  
32 (100 hPa.). The values observed in the lower stratosphere are less than those observed in the  
33 troposphere. Indeed, the CO and water vapor mixing ratio anomalies at 68 hPa (~19 km) are  
34 less than 10 % and 20 % respectively. At 68 hPa, it can be observed that high values of water



1 vapor mixing ratio anomalies are not in coincidence with high values of CO mixing ratio  
2 anomalies, but located along the vertical extent of the maxima of CO mixing ratio anomalies in  
3 the troposphere.

4 Our analysis corroborates the results found by Héron et al. (2020). Based on radiosonde and  
5 satellite observations, Héron et al. (2020) showed that convective activity over the SWIO basin  
6 has the potential to influence the variability of ozone and water vapor in the upper-troposphere  
7 during the austral summer. In the present study, the satellite observations and FLEXPART  
8 simulations strongly suggests that African air masses enriched in CO and aerosols were lifted  
9 from the lower troposphere to the lower stratosphere due to convective activity. Our results  
10 demonstrate that the variability of CO and aerosol over the SWIO basin can be explained both  
11 by the influence of long-range transport of the Australian fires plumes, together with regional  
12 transport from southern Africa, enhanced by convective activity due to the passage of a tropical  
13 storm.

## 14 **5.2 Discussion on the variability of the total and partial column of CO**

15 The African and Australian BB contributions to CO measured concentrations over Reunion are  
16 assessed by coupling the FLEXPART model with gases emissions from the GFAS inventory.  
17 The emission sensitivity from FLEXPART, only at the altitude where the emissions are injected  
18 (0-3 km for Africa fires and 9-16 km for Australian fires), is combined with CO emission  
19 inventory. The CO emissions due to anthropogenic activity are also taken considered by  
20 coupling the FLEXPART model with the EDGAR inventory. The evolutions of the total and  
21 partial column of CO observed by IASI and simulated by FLEXPART from 15<sup>th</sup> to 29<sup>th</sup> January  
22 over Reunion are depicted in Figure 12.

23 Despite the underestimation of the simulated total and partial column of CO, their overall  
24 temporal evolution is fairly well reproduced by FLEXPART (Fig. 12). The discrepancies  
25 between FLEXPART and IASI may be attributable to several possible caveats. A possible  
26 source of error can be due to the fact that the vertical motion induced by pyro-convection is not  
27 included in FLEXPART. We address this issue by applying an injection height in agreement  
28 with CALIOP observations (9-16 km, Fig. 1) for the Australian plume (Khaykin et al., 2020).  
29 An injection height ranging up to 3 km was chosen for the African fires (Labonne et al., 2007).  
30 The injection height of the plume plays a key role in its long-range transport (Sofiev et al.,  
31 2012). An inappropriate or unrealistic injection height can lead to either a dilution or an  
32 overestimation of the plume. The injection height depends on the intensity of the fire, as well  
33 as on the meteorological conditions. Another possible explanation in these differences can come  
34 from the duration of the backward calculation (1 month) (Brocchi et al., 2018). Underestimation



1 of the total and partial columns of CO may also be the result of an underestimation of the CO  
2 emissions by GFAS. In the framework of their FLEXPART simulation, Brocchi et al. (2018)  
3 reveal that an amplification factor of 2 has been applied to CO emissions from GFAS to get  
4 similar CO quantities to observations. Nevertheless, these simulations can reasonably be used  
5 to discuss the contribution of CO emissions from Africa and Australia on the variability of the  
6 total and partial columns of CO over the SWIO basin.

7 The contribution of the CO emissions from Africa and Australia on the total and partial columns  
8 of CO over Reunion is isolated and depicted in Figure 12. On average, the CO emissions from  
9 Africa contribute up to 90% of the enhancement of the total and partial column of CO from 15<sup>th</sup>  
10 to 29<sup>th</sup> January. The transport of the CO plume induced by the Australian sources have not been  
11 efficient over the SWIO basin. This could be explained by the fact that CO decays quickly due  
12 to photochemical oxidation whose efficiency increase with altitude. Khaykin et al. (2020)  
13 showed that the decay of CO within the bubble occurred during its transport over the Southern  
14 Pacific on its way to the Indian ocean. The variability of CO over the SWIO basin is hence  
15 mainly explained by the regional transport of air masses.

### 16 **5.3 Discussion on the variability of the aerosols**

17 The FLEXPART simulations for the contribution of aerosol (BC and OC) emissions from  
18 Africa and Australia on the sAOD over Reunion use the same parameters as for CO, in  
19 particular the injection height. The emission sensitivity from FLEXPART is combined with BC  
20 and OC from the GFAS emission inventory. Figure 13 depicts the evolution of the sAOD  
21 observed by OMPS and simulated by FLEXPART from 15<sup>th</sup> to 29<sup>th</sup> January over Reunion.  
22 Unfortunately, few data have been recorded during the period aforementioned. Nevertheless,  
23 the simulated sAOD compare fairly well with the available OMPS observations during this  
24 period, and the sAOD peak observed on 19<sup>th</sup> January is acceptably well reproduced. Figure 13  
25 reveals that the variability of the sAOD between 15 and 30 km altitude is mainly driven by the  
26 aerosol emissions from Australia. On average, the aerosol emissions from Australia contributed  
27 up to 95 % of the sAOD variability over Reunion from 15<sup>th</sup> to 29<sup>th</sup> January. One can observe  
28 that the evolution of sAOD is not correlated to the evolution of the aerosol emission from Africa  
29 which is marked by an increase from 21<sup>th</sup> to 29<sup>th</sup> January. The weak contribution of African  
30 component on the sAOD can be explained by the fact that the amounts of African BB aerosols  
31 injected in the atmosphere by the convective activity decrease with altitude. Moreover, it is  
32 likely that aerosol would be scavenged by cloud droplets (taken into account in FLEXPART)  
33 in a strongly convective environment such as tropical storm Diane. Overall, FLEXPART  
34 simulations clearly suggested the contribution of African and Australian BB aerosols.





## 1 **6. Summary and Conclusion**

2 The complex aerosol and CO variabilities over the SWIO basin during the 2020 austral summer  
3 have been investigated. The meteorological context and the extensive fires over southeastern  
4 Australia were favorable to trigger pyro-convective events between 29<sup>th</sup> December 2019 and  
5 12<sup>th</sup> January 2020. These pyro-convective events led to massive injection of combustion  
6 products in the stratosphere. The ground-based and space-borne lidars revealed the presence of  
7 an intense stratospheric smoke layer over the southeastern Australia region. Over the Lauder  
8 site in New Zealand, this smoke layer was detected in the stratosphere (centered at 16 km) until  
9 April and beyond. The analysis of the spatial and temporal dispersion of the Australian smoke  
10 layer highlighted its quick transport circling the entire Southern Hemisphere in less than two  
11 weeks. Furthermore, the satellite observations revealed that the transport of the Australian  
12 smoke layer has been mainly bounded within an extra-tropical latitudinal band.

13 Nevertheless, the numerical models clearly showed the influence of the Australian smoke layer  
14 on the variability of aerosol over the SWIO basin. Over Reunion, the aerosol extinction profiles  
15 exhibited a significant increase in the lower stratosphere during the end of January. The  
16 MIMOSA simulations highlighted the isentropic transport of the Australian BB aerosols from  
17 extra-tropical latitudes to Reunion at 400 K isentropic level, on 28<sup>th</sup> January. As a consequence,  
18 the corresponding aerosol extinction profile exhibited a sudden increase by drawing a structure  
19 similar to a laminae at the 400 K isentropic level. The aerosol extinction profiles also exhibited  
20 a moderate increase in the upper troposphere.

21 This paper investigates for the first time the possibility of the African emissions from BB to  
22 influence the CO and aerosol distribution in the UT-LS during the convective season. Despite  
23 the fact that African BB activity is usually sparse in January, it contributed to modulate the  
24 vertical distribution of CO and aerosols in the upper troposphere over the SWIO basin. The  
25 analysis of satellite observations and FLEXPART simulations suggested that because of the  
26 convective activity, air masses enriched in CO and aerosols have been lifted from the lower  
27 troposphere to the lower stratosphere. Air masses from Africa contributed up to 90% of the total  
28 and partial column (between 9 and 30 km) of CO variability over Reunion and its surroundings.  
29 The simulations shows that the modulation of the aerosol extinction in the upper troposphere  
30 and the lower stratosphere, over Reunion, was driven by the transport of air masses from both  
31 Africa and Australia, respectively. Our findings suggest simultaneous presence of African and  
32 Australian aerosol smoke layers at Reunion.

33

34



## 1 **Acknowledgements**

2 We thank Isamu Morino and his staff for their effort in establishing and maintaining the sun-  
3 sky radiometer at Lauder. Data used in this publication were obtained from NDACC network  
4 and are available through its website (<http://www.ndacc.org/>). The work of S. Khaykin has been  
5 supported by the Agence Nationale de la Recherche PyroStrat project (21-CE01- 335 0007-01).  
6 IASI was developed and built under the responsibility of the “Centre National d’Etudes  
7 Spatiales” (CNES, France) and flown on board the MetOp satellites as part of the EUMETSAT  
8 Polar System. The authors thank the AERIS infrastructure (<https://iasi.aeris-data.fr/>) for  
9 providing access to the IASI data, the National Aeronautical and Space Administration (NASA)  
10 for providing CALIOP data, MLS and MODIS fire products. We would especially like to thank  
11 the staff of the team working on the lidar systems at the Maïdo observatory. The French  
12 Research Infrastructure ACTRIS-FR, CNES and the EU project REALISTIC (GA 101086690)  
13 are also acknowledged for their support in the upgrade and operation of the Maïdo observatory  
14 lidars. Lauder observations are funded by the New Zealand Government’s Strategic Science  
15 Investment Fund (SSIF), administered by the Ministry of Business, Innovation and  
16 Employment (MBIE). The TCCON site at Réunion Island has been operated by the Royal  
17 Belgian Institute for Space Aeronomy with financial support since 2014 by the EU project  
18 ICOS-Inwire, the ministerial decree for ICOS (FR/35/IC1 to FR/35/C6), ESFRI-FED ICOS-  
19 BE project and local activities supported by LACy/UMR8105 and by OSU-R/UMS3365 –  
20 Université de La Réunion. The lidar measurements are supported by the funding of GOSAT  
21 series project.

## 22 **Data availability**

23 The data used for this study are available and open access by request to scientist mentioned or  
24 through the link hereafter: Lidar measurements ([tetsu@mri-jma.go.jp](mailto:tetsu@mri-jma.go.jp), [nelson.begue@univ-  
25 reunion.fr](mailto:nelson.begue@univ-reunion.fr)), radiometer measurements ([ben.liley@niwa.co.nz](mailto:ben.liley@niwa.co.nz)), FTIR measurements from  
26 TCCON network ([mahesh.sha@aeronomie.be](mailto:mahesh.sha@aeronomie.be)); Lauder FTIR data available on the NDACC  
27 public access database (<https://www-air.larc.nasa.gov/missions/ndacc/data.html>); The satellite  
28 observations and emission inventory used are available on-line from the sources as stated in the  
29 manuscript.

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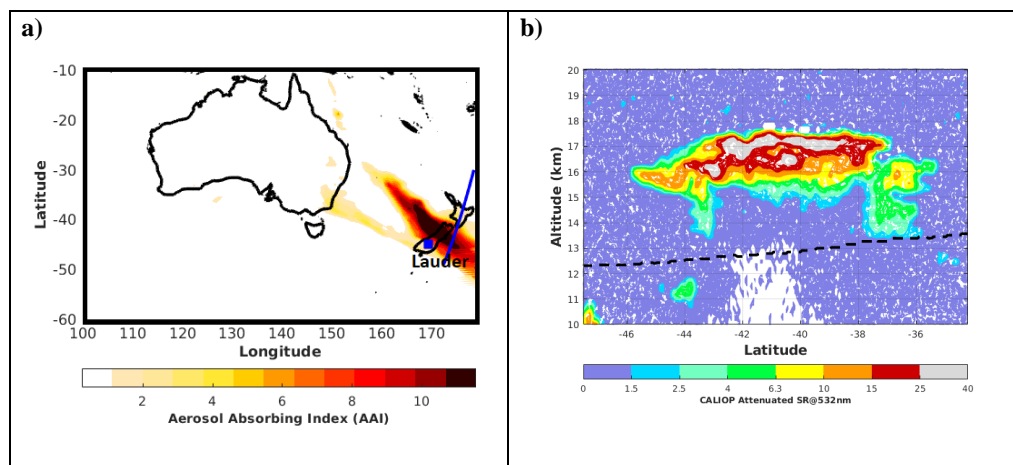
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2 **FIGURES AND TABLE**

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6 **Figure 1:** (a) Map of Aerosol Absorbing Index obtained from OMPS observations and (b)

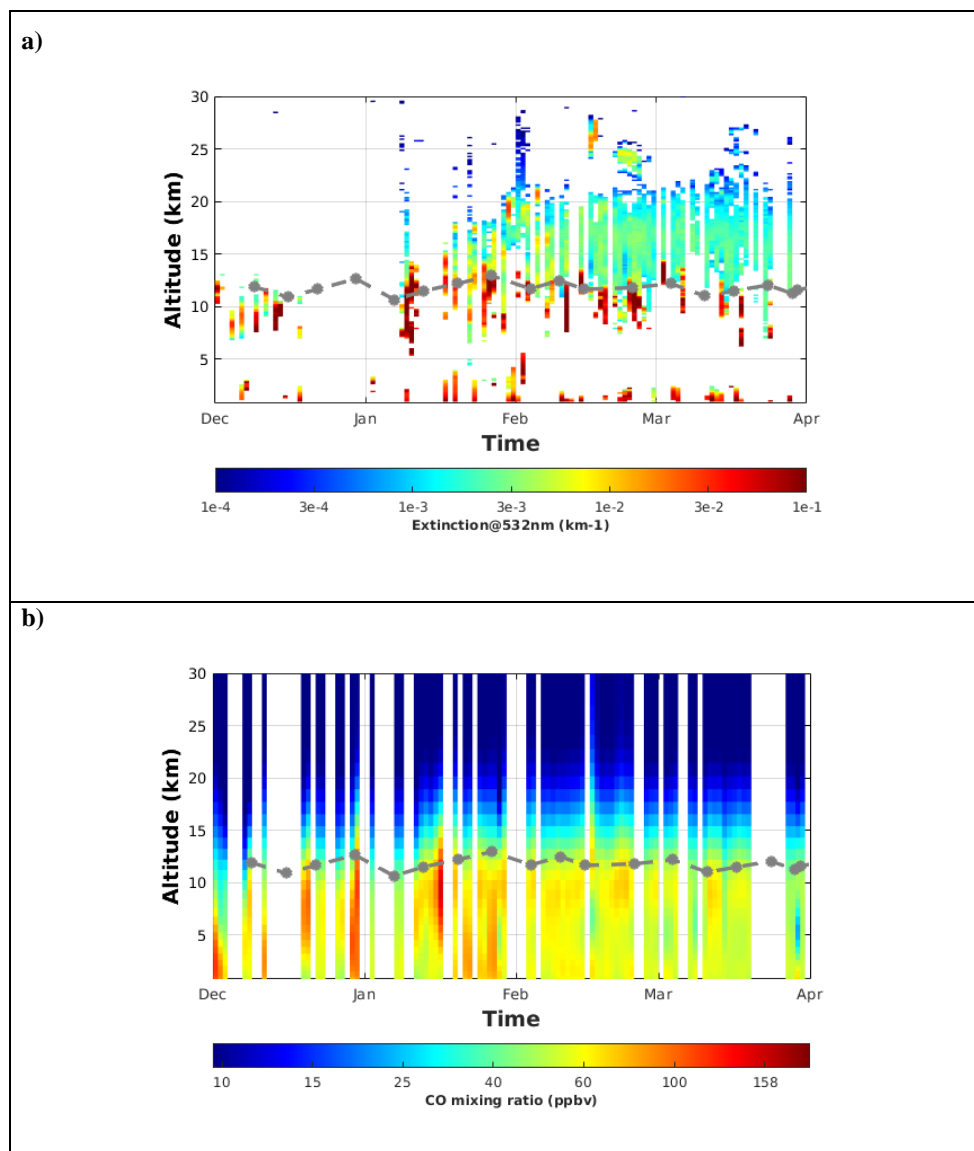
7 scattering ratio profiles at 532 nm obtained from CALIOP observations on 1<sup>st</sup> January 2020.

8 The orbit overpass of CALIOP is indicated by the blue curve, while the blue square corresponds

9 to the Lauder site in plot (a). The black dashed line in (b) corresponds to the 380 K isentropic

10 level calculated from CALIOP observations.

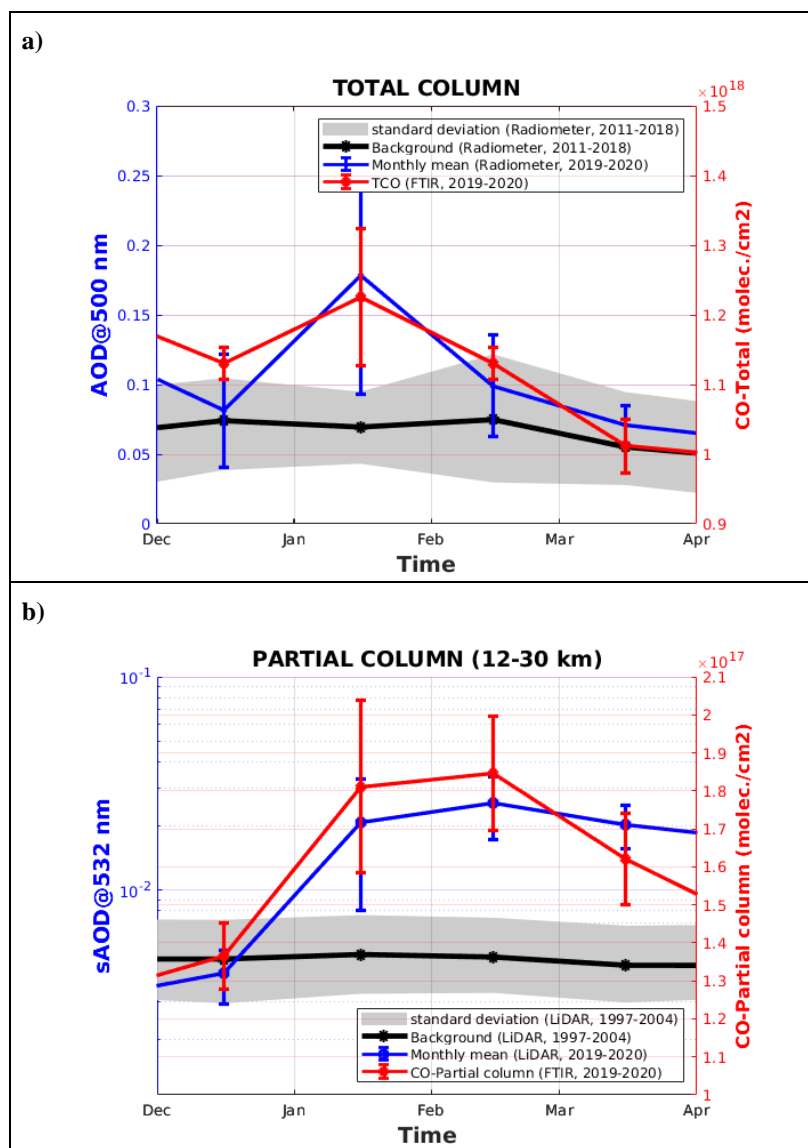
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1 **Figure 2:** Time series of daily profiles at Lauder of (a) aerosol extinction at 532 nm obtained  
2 from lidar and (b) CO mixing ratio obtained from FTIR between 1<sup>st</sup> December 2019 and 1<sup>st</sup>  
3 April 2020. In order to screen non-aerosol contributors (such as clouds) on the extinction  
4 measurements, a mask based on the method reported by Nicolae et al. (2013) that includes  
5 consideration on plausible aerosols properties was used. Namely, we only kept profile parts  
6 with positive depolarization values, and Angström exponent ranges from 0.1 to 4. The grey line  
7 indicates the tropopause height obtained from radiosonde measurements.  
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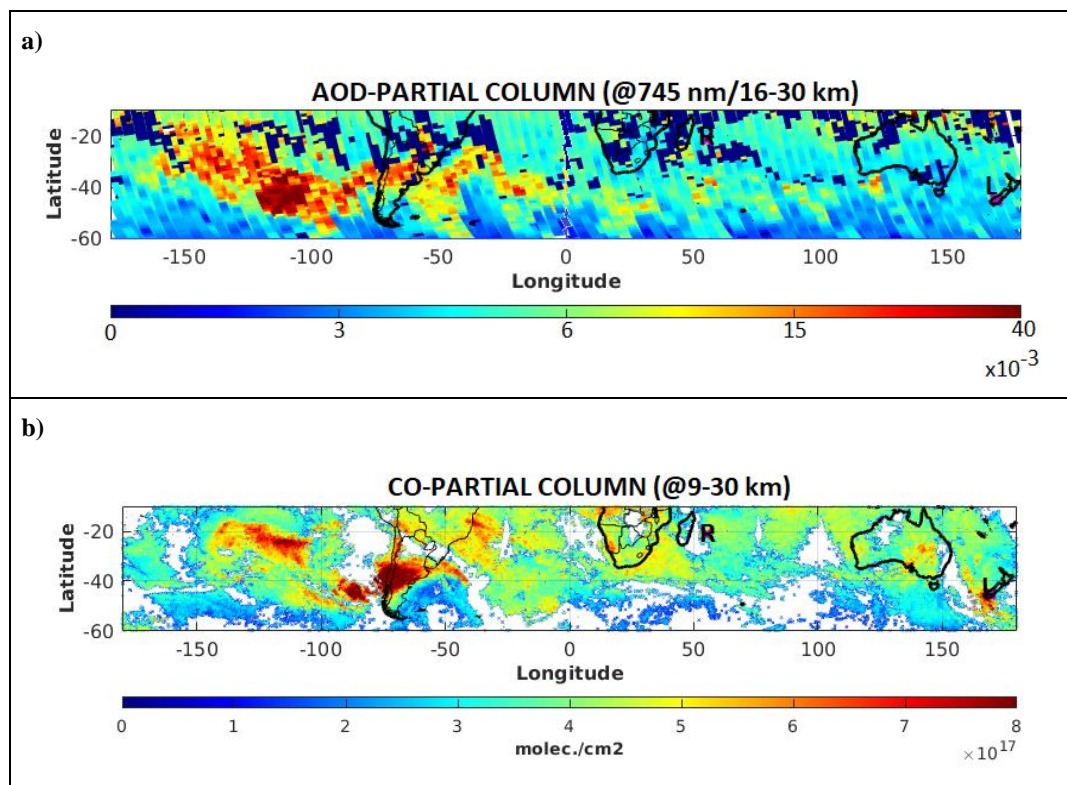
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2 **Figure 3:** (a) Monthly mean evolution of total column of aerosol (AOD at 500 nm) and CO  
 3 obtained from SKYNET radiometer and FTIR respectively over Lauder from 1<sup>st</sup> December  
 4 2019 to 1<sup>st</sup> April 2020. (b) Monthly mean evolution of the stratospheric AOD (sAOD) and CO  
 5 (sCO) columns at Lauder between 1<sup>st</sup> December 2019 and 1<sup>st</sup> April 2020. The sAOD and sCO  
 6 are calculated between 12 and 30 km from lidar and FTIR measurements respectively. The  
 7 background evolution of aerosol data (AOD: 2011-2018 and sAOD: 1997-2004) and the  
 8 associated standard deviation are given in black lines and grey areas, respectively.



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3 **Figure 4:** Time-averaged map (from 9<sup>th</sup> to 16<sup>th</sup> January 2020) of (a) sAOD (between 16 and 30  
4 km at 745 nm) obtained from OMPS observations and (b) partial column of CO (averaged  
5 between 9 and 30 km) obtained from IASI observations. The location of Reunion and Lauder  
6 sites are indicated by R and L respectively.

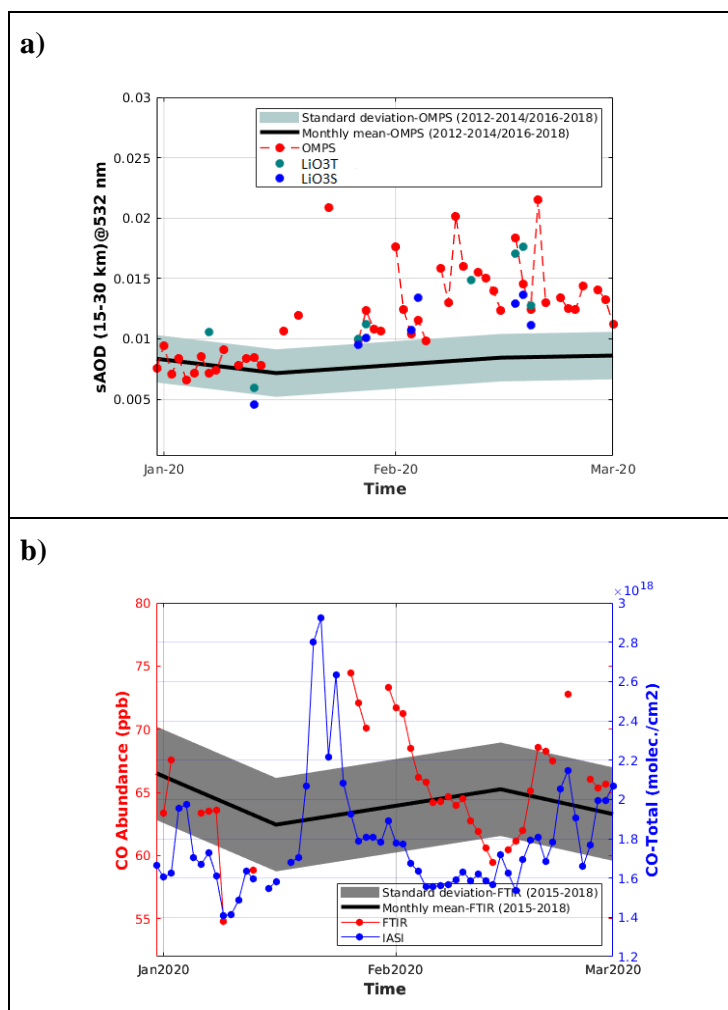
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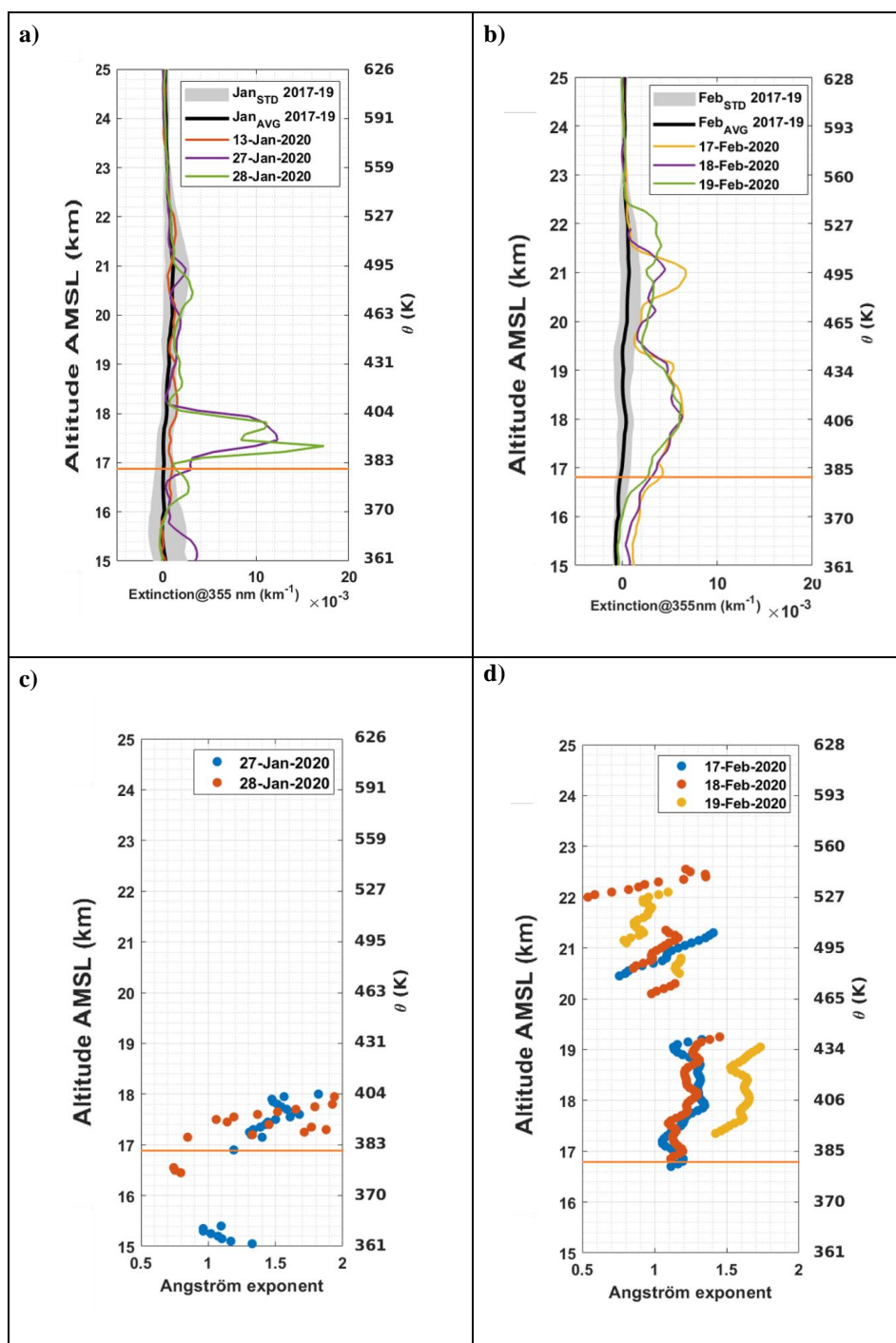
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3 **Figure 5:** Daily mean evolution of aerosol (a) and CO (b) abundances obtained from ground-  
 4 based and satellite observations at Reunion between 1<sup>st</sup> January and 1<sup>st</sup> March 2020. Total  
 5 column (molecule.cm<sup>-2</sup>) and abundance (ppb) of CO obtained from IASI (blue line) and FTIR  
 6 (red line) respectively are given in the lower panel (b), while sAOD obtained from OMPS (red  
 7 line) and Lidar (blue and green dots) are given in the upper panel (a). The black and dashed  
 8 lines correspond to monthly mean and the associated standard deviation calculated during the  
 9 background period.

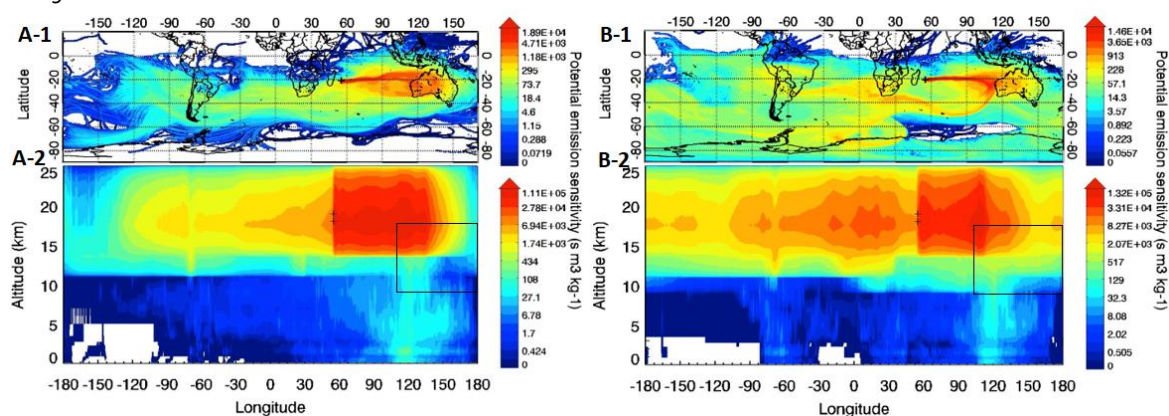
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1 **Figure 6:** Aerosol extinction (at 355 nm) (a, b) and Angström exponent (355-532 nm) (c, d)  
 2 obtained from lidar observations at Reunion in the months January and February 2020. The  
 3 tropopause height is indicated by the orange horizontal lines.



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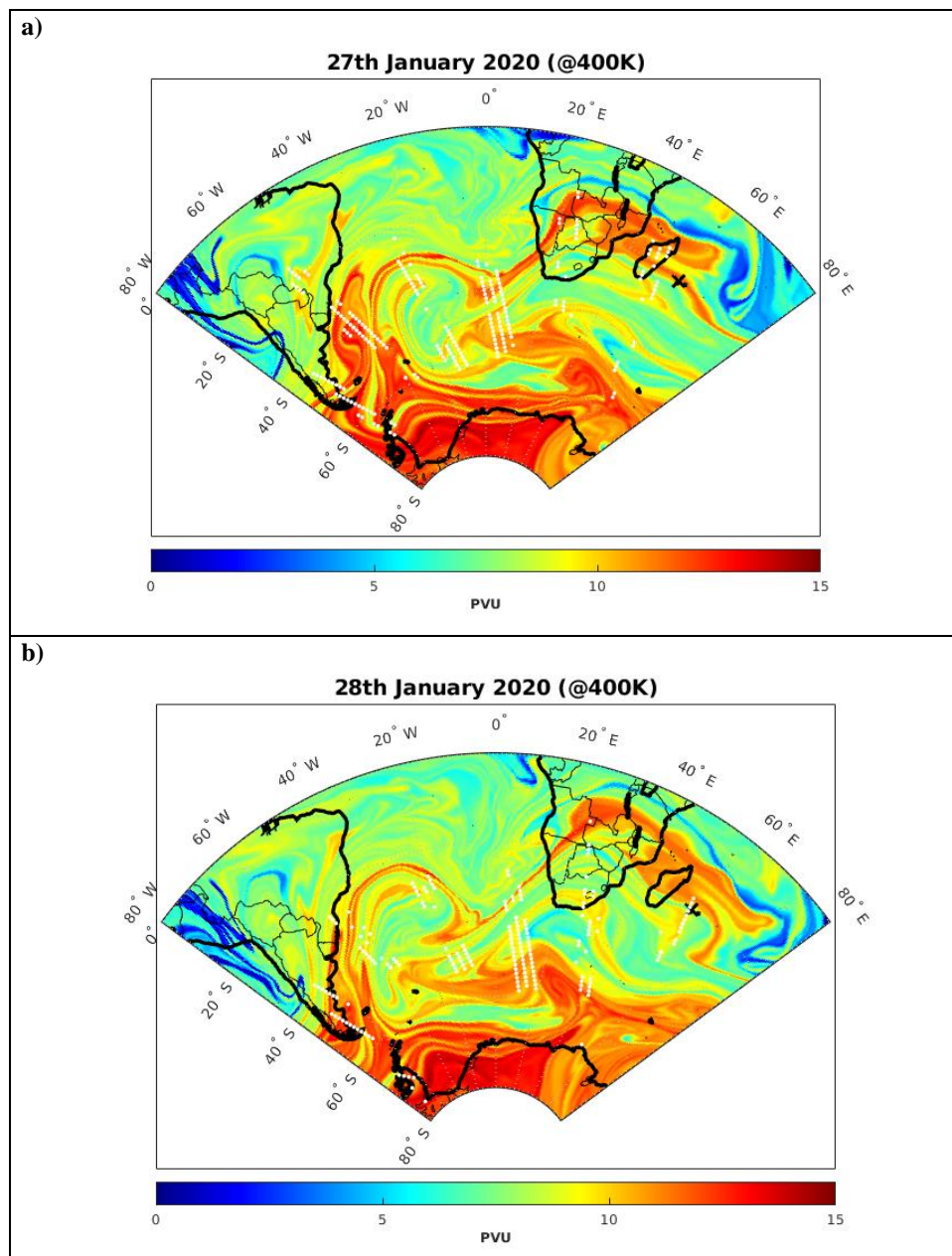


11 **Figure 7:** FLEXPART 30-day back trajectories initialized from Reunion (black cross) at 18 km  
 12 on 27<sup>th</sup> January 2020 (A-1-2) and 28<sup>th</sup> January 2020 (B-1-2). A-1 and B-1 correspond to an  
 13 integration of the trajectory positions over the whole altitude range. A-2 and B-2 are the vertical  
 14 view integrated over the whole latitude range of the back trajectories A-1 and B-1. The black  
 15 rectangle represented the injection height of the biomass burning aerosols.

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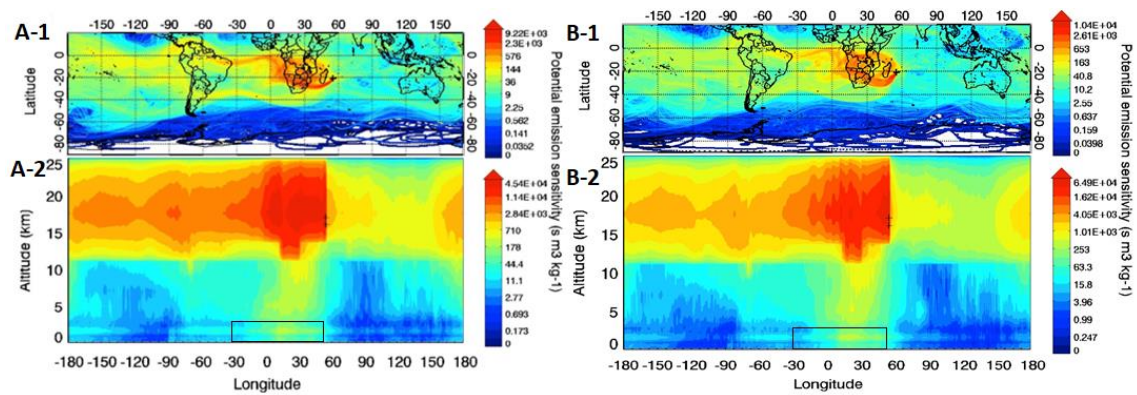
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2 **Figure 8:** Advected PV map at the 400 K level obtained from the MIMOSA model (a) on 27  
3 January 2022 and (b) on 28 January 2022. The white dots represent the localization of the  
4 aerosol plume at  $400 \text{ K} \pm 5 \text{ K}$  obtained from OMPS observations, while the black cross indicates  
5 Reunion.  
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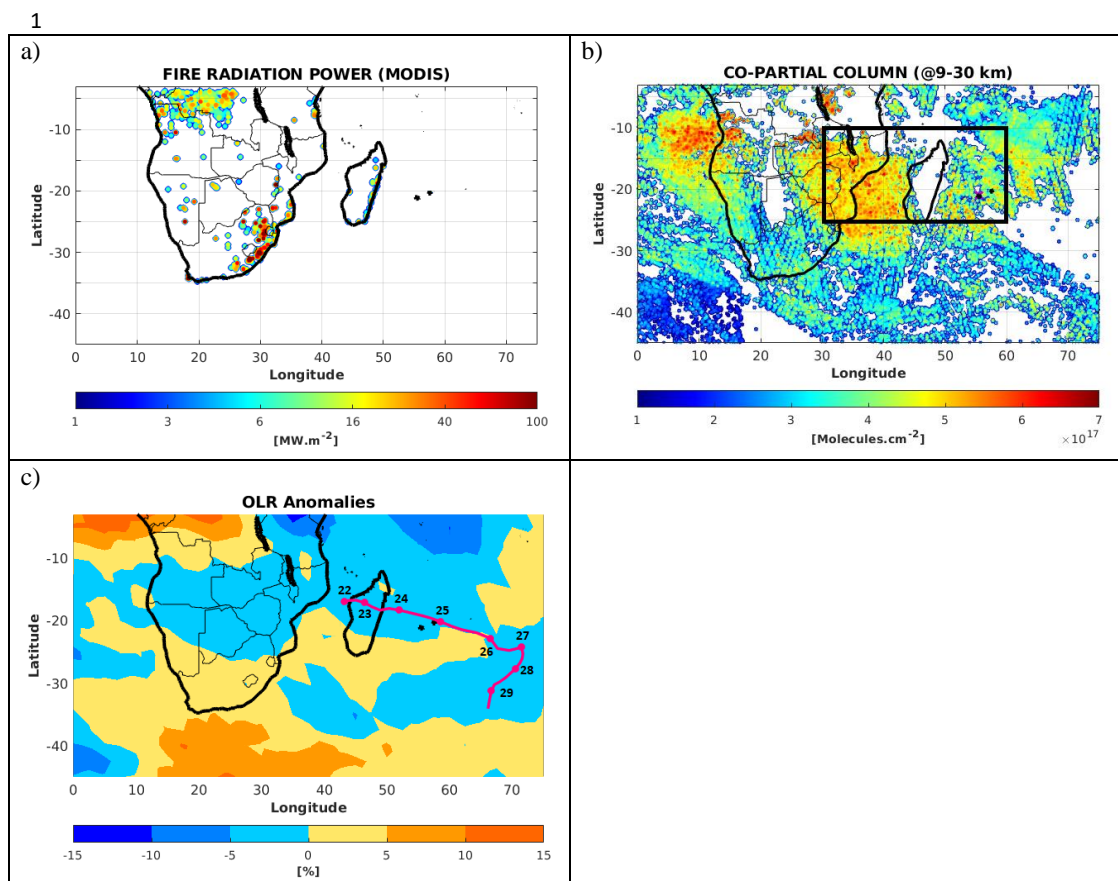
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10 **Figure 9:** Same as figure 7 with an injection height initialized at 16 km.

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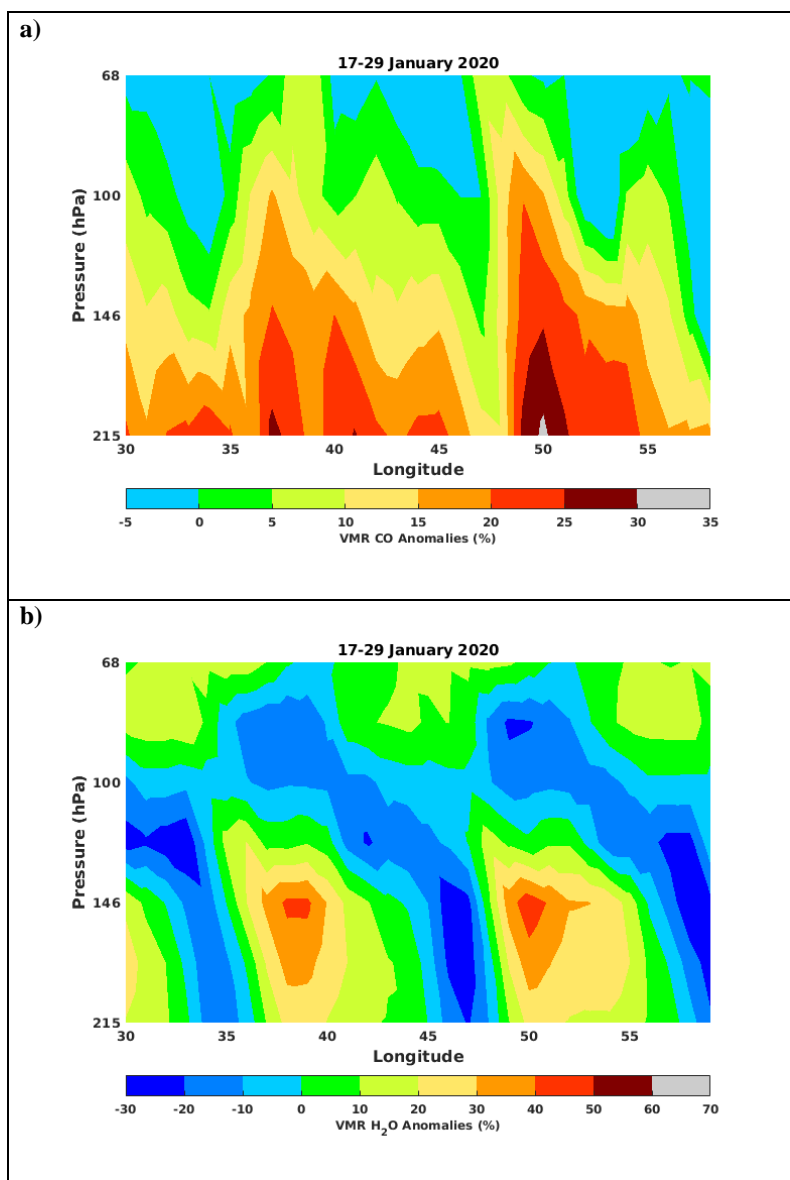
2 **Figure 10:** a) The total number of fire pixel and the associated fire radiative power obtained  
 3 from MODIS observation between 16<sup>th</sup> and 29<sup>th</sup> January 2020. b) Time-average map of partial  
 4 column of CO (calculated between 9 and 30 km) obtained from IASI observations averaged  
 5 between 16<sup>th</sup> and 29<sup>th</sup> January 2020. The black square corresponds to the study domain where  
 6 the vertical cross-section of CO and water vapor mixing ratio are calculated and reported in  
 7 Figure 11. c) Time-average map of outgoing longwave radiation anomalies obtained from  
 8 NCEP between 16<sup>th</sup> and 29<sup>th</sup> January 2020. The red curve corresponds the trajectory followed  
 9 by the Diane strong tropical storm from 22<sup>nd</sup> to 29<sup>th</sup> January 2020. This trajectory is obtained  
 10 from the RSMC (Regional Specialized Meteorological Center) of Reunion best-track database.

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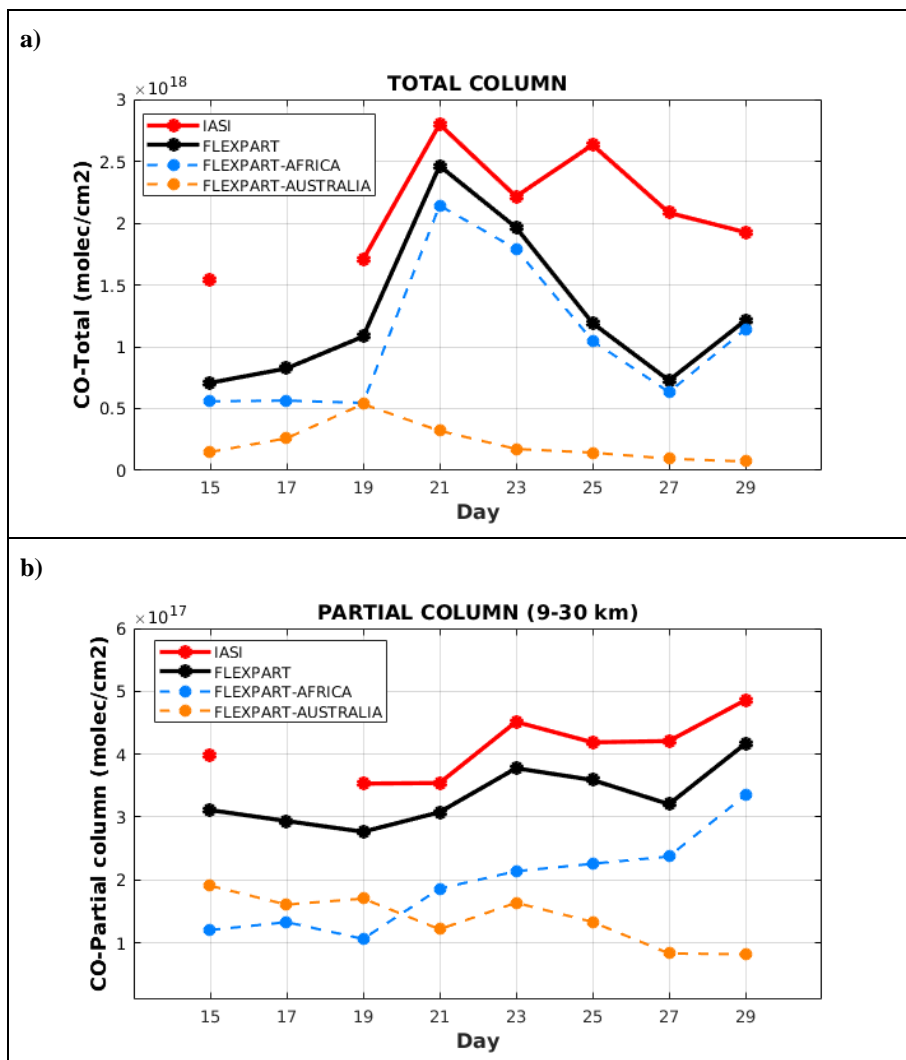
3 **Figure 11:** Vertical cross section of (a) CO and (b) water vapor mixing ratio anomalies obtained  
4 from MLS observation over southern Africa and the SWIO basin (black box in Figure 10b)  
5 between 16<sup>th</sup> and 29<sup>th</sup> January 2020.

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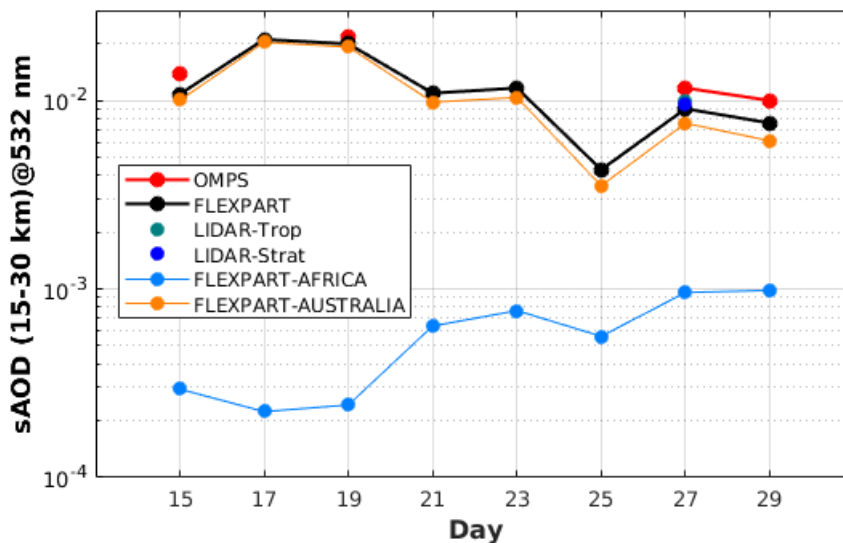


2 **Figure 12:** Daily evolution of (a) total column and (b) partial column (calculated between 9  
3 and 30 km) of CO observed by IASI (red line) and simulated by FLEXPART (black line) over  
4 Reunion from 15<sup>th</sup> to 29<sup>th</sup> January 2020. The CO evolution is simulated by FLEXPART in  
5 considering only the CO emission (including BB and anthropogenic activity) from African  
6 (cyan line) and Australian (orange line) BB sources.

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**Figure 13:** Daily evolution of sAOD (calculated between 15 and 30 km at 532 nm) obtained from OMPS (red line), lidar (blue dots) and simulated by FLEXPART (black line) over Reunion from 15<sup>th</sup> to 29<sup>th</sup> January 2020. The simulated sAOD are calculated in considering only the aerosol emission (BC and OC) from African (cyan line) and Australian (orange line) BB sources.