

# Fast uplift in the Southern Patagonian Andes due to long and short term deglaciation and the asthenospheric window underneath

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**Abstract.** An asthenospheric window underneath much of the South American continent increases the heat flow in the Southern Patagonian Andes, where glacial-interglacial cycles drive the building and melting of the Patagonian Icefields since the latest Miocene. The Last Glacial Maximum (LGM) was reached ~26000 years Before Present (BP). Significant deglaciation onsets between 21000 and 17000 years BP, subject to an acceleration since the Little Ice Age (LIA), ~400 years BP. Fast uplift rates of up to  $41 \pm 3$  mm/yr are measured by GNSS around the Southern Patagonian Icefield and currently ascribed to post-LIA lithospheric rebound, but the possible longer-term post-LGM rebound is poorly constrained. These uplift rates, in addition, are one order of magnitude higher than those measured on other glaciated orogens (e.g., the European Alps), which raises questions about the role of the asthenospheric window in affecting the vertical surface displacement rates. Here, we perform geodynamic thermo-mechanical numerical modelling to estimate the surface uplift rates induced by post-LIA and post-LGM deglaciation accounting for temperature dependent rheologies and different thermal regimes in the asthenosphere. Our modelled maximum postglacial rebound matches the observed uplift rate budget only when both post-LIA and post-LGM deglaciation are accounted for and if a standard continental asthenospheric mantle potential temperature is increased by 150-200 °C. The asthenospheric window thus plays a key role in controlling the magnitude of presently observed uplift rates in the Southern Patagonian Andes.

## 33 1. Introduction

34 Vertical displacements of the Earth's surface with respect to the geoid occur in  
35 response to the motion of crustal and mantle rock masses due to mantle dynamics, plate  
36 tectonics, and the redistribution of sediments, water, and ice by surface processes (e.g.,  
37 Molnar and England, 1990; Watts, 2001; Champagnac et al., 2012; Sternai, 2023; Cloetingh  
38 et al., 2023). For instance, excess of topography in orogenic regions due to convergence,  
39 crustal shortening, and magmatism deflects the lithosphere downward, whereas surface  
40 unloading by erosion and ice melting causes upward deflection of the lithosphere, known as  
41 "isostatic" adjustment (e.g., Peltier and Andrews, 1976; Peltier, 1996, 2004; Mitrovica and  
42 Forte, 1997; Butler and Peltier, 2000; Kaufman and Lambeck, 2002; Watts, 2001; Turcotte  
43 and Schubert, 2002; Sternai et al., 2016a). Glacial isostatic adjustment (GIA) models study  
44 the visco-elastic response of the solid Earth to the building and melting of regional ice sheets  
45 and commonly use GNSS and/or satellite-measured rock uplift rates in regions subject to  
46 deglaciation to estimate, for instance, the regional mantle rheology and sea-level changes  
47 (e.g., Peltier and Andrews, 1976; Peltier, 1996, 2004; Mitrovica and Forte, 1997; Kaufman  
48 and Lambeck, 2002; Stuhne and Peltier, 2015; Van der Wal et al., 2015; Peltier et al., 2018;  
49 Whitehouse, 2018). Most of GIA studies address the post Last Glacial Maximum (LGM)  
50 around 21000 years Before Present (BP) deglaciation as a trigger to increasing uplift rates in  
51 glaciated regions (e.g., Peltier, 2004; Whitehouse, 2018). The magnitude of uplift rates is set  
52 primarily by the lithosphere and asthenosphere viscosities, which depend, amongst other  
53 factors, on the thermal field at depth (McKenzie and Richter, 1981; McKenzie and Bickle,  
54 1988; Gurnis, 1989; Ranalli, 1995, 1997; Kaufman et al., 1997; Watts, 2001; Turcotte and  
55 Schubert, 2002). While the GIA theory is well developed, few studies use thermo-mechanical  
56 visco-elasto-plastic (non-Newtonian Earth layers) geodynamic models to estimate uplift rates  
57 in response to surface load changes to be compared with GNSS data. Here, we use this  
58 approach to constrain the role of the solid Earth rheology in setting the rates of surface  
59 vertical displacement in Southern Patagonia, which hosts the biggest continental ice-sheets  
60 outside Antarctica and presents ongoing very high rock uplift rates (Ivins and James, 2004;  
61 Dietrich et al., 2010; Lange et al., 2014; Richter et al., 2016; Lenzano et al., 2023).

62 The Southern Patagonian Andes in the South American Continent are located above a  
63 transition zone between the subducting Antarctic and Nazca plates and a wide asthenospheric  
64 window, where hot buoyant asthenospheric mantle upwells (Fig. 1a; Cande and Leslie, 1986;  
65 Breitsprecher and Thorkelson, 2009; Russo et al., 2010, 2022; Dávila et al., 2018; Ávila and

66 Dávila, 2018, 2020; Mark et al., 2022; Ben-Mansour et al., 2022). The Chile Triple Junction  
67 (CTJ) at  $\sim 46^\circ\text{S}$  delimits the surface tip of the asthenospheric window, which opened during  
68 the last  $\sim 16$  Ma from south to north (Ramos and Kay, 1992; Breitsprecher and Thorkelson,  
69 2009). First order effects of the asthenospheric flow on the surface continental geology are  
70 the inhibition of arc volcanism in favour of retroarc magmatism, reduction of shortening to  
71 null or very minor, and rock uplift and exhumation (Ramos and Kay, 1992; Ramos, 2005;  
72 Lagabriele et al., 2004, 2010; Breitsprecher and Thorkelson, 2009; Guillaume et al., 2009,  
73 2013; Scalabrino et al., 2010; Lange et al., 2014; Georgieva et al., 2016, 2019; Ávila and  
74 Dávila, 2020; Mark et al., 2022; Ávila et al., 2023; Muller et al., 2023). Rock uplift due to  
75 asthenospheric upwelling occurs through dynamic and thermal effects (Guillaume et al.,  
76 2009, 2013; Conrad and Husson, 2009; Flament et al., 2013; Sternai et al., 2016b; Dávila et  
77 al., 2018; Ávila and Dávila, 2020; Faccenna and Becker, 2020; Mark et al., 2022). Dynamic  
78 uplift occurs above zones of viscous convection of the asthenospheric mantle, generating  
79 long wavelength ( $>300$  km) deformation of the lithosphere through vertical stresses (Hager  
80 and O'Connell, 1981; Flament et al., 2013, 2015; Sternai et al., 2016b; Ávila and Dávila,  
81 2020; Faccenna and Becker, 2020). This effect is difficult to measure because vertical  
82 stresses in the lithosphere occur also due to lithospheric tectonics and the surface mass  
83 redistribution of glaciers, lakes, and sediments (Lachenbruch and Morgan, 1990; Molnar and  
84 England, 1990; Watts, 2001). Dynamic uplift was estimated between 0.02 and 0.15 mm/yr in  
85 the last 3 Ma over an area of about  $100000\text{ km}^2$  around the CTJ latitude (Guillaume et al.,  
86 2009, 2013; Flament et al., 2015; Ávila and Dávila, 2020; Ávila et al., 2023). The thermal  
87 component is expressed by an increase of temperatures and shallowing of the lithosphere-  
88 asthenosphere boundary where asthenospheric mantle upwells (Ávila and Dávila, 2018;  
89 2020; Russo et al., 2010, 2022; Mark et al., 2022; Ben-Mansour et al., 2022), decreasing the  
90 integrated elastic lithospheric thickness and generating uplift and higher surface heat flow  
91 than in normal subduction zones (Ranalli, 1997; Flament et al., 2015; Ávila and Dávila, 2018,  
92 2020; Ávila et al., 2023). The heat flow was calculated as  $>100\text{ mW/m}^2$  near the CTJ,  $\sim 70\text{-}90$   
93  $\text{mW/m}^2$  in the centre of the asthenospheric window ( $\sim 50^\circ\text{S}$ ), and  $50\text{-}60\text{ mW/m}^2$  near its  
94 northern boundary ( $\sim 46^\circ\text{S}$ ) (Ávila and Dávila, 2018). Uplift due to lithospheric thinning was  
95 estimated as  $\sim 0.3$  mm/yr since the middle Miocene in the Southern Patagonian Andes (Pedoja  
96 et al., 2011; Ávila and Dávila, 2020; Ávila et al., 2023; Ding et al., 2023).

97 The Patagonian Ice Sheet covered the Southern Patagonian Andes between  $\sim 47000$   
98 and  $\sim 17000$  years BP, extending from latitudes  $38^\circ$  to  $55^\circ\text{S}$  with an estimated area of  
99  $\sim 490000\text{ km}^2$  (Fig. 1 a), volume of  $\sim 550000\text{ km}^3$ , and average and maximum thickness of

100 1100 and 2500 m, respectively, based on preserved glacial geomorphology, stratigraphy,  
101 paleoecology, and geochronological data (Moreno et al., 1999, 2005, 2015; McCulloch et al.,  
102 2000, 2005; Hulton et al., 2002; Rabassa, 2008; Glasser et al., 2004, 2005, 2008, 2016;  
103 Glasser and Jansson, 2008; Hein et al., 2010; Boex et al., 2013; Strelin et al., 2014; Bourgois  
104 et al., 2016; Martinod et al., 2016; Kaplan et al., 2016; Bendle et al., 2017; Thorndycraft et  
105 al., 2019; Reynhout et al., 2019; Davies et al., 2020; Yan et al., 2022). The LGM in Southern  
106 Patagonia is estimated around 26000 years BP, but the beginning of significant glacial retreat  
107 occurred between 21000 and 17000 years BP (Hulton et al., 2002; Hein et al., 2010; Glasser  
108 et al., 2011; Glasser and Davies, 2012; Moreno et al., 2015; Bendle et al., 2017; Reynhout et  
109 al., 2019; Davies et al., 2020). Long term ice loss rate is uncertain, but more than 75% of ice  
110 was certainly lost since the LGM, and some models predicted more than 95% of ice loss with  
111 separation between the Southern Patagonian Icefield (SPI) and the Northern Patagonian  
112 Icefield (NPI) in the first 5000 to 10000 years of post-LGM deglaciation (McCulloch et al.,  
113 2000; Hulton et al., 2002; Boex et al., 2013; Bourgois et al., 2016; Thorndycraft et al., 2019;  
114 Davies et al., 2020). A glacial minimum must have been attained around 13000 years BP, but  
115 several glacial advances were recorded since that time, and the last one was the Little Ice Age  
116 (LIA) with apex around 1630 AD, well dated by terminal moraines around the present-day  
117 NPI and SPI (Ivins and James, 1999, 2004; McCulloch et al., 2000; Glasser et al., 2004,  
118 2008, 2011; Davies and Glasser, 2012; Strelin et al., 2014; Kaplan et al., 2016; Reynhout et  
119 al., 2019; Davies et al., 2020). Recent mass balance measurements in the Patagonian icefields  
120 - e.g., Shuttle-Radar Topography Mission (SRTM) or Gravity Recovery and Climate  
121 Experiment (GRACE) - often present discrepancies, but consistently show an increasing ice  
122 loss from ~15 Gt/yr between ~1940-2000, to ~25 Gt/yr between ~2000-2012 (Aniya, 1996;  
123 Aniya et al., 1997; Rignot et al., 2003; Chen et al., 2007; Ivins et al., 2011; Jacob et al., 2012;  
124 Willis et al., 2012; Gómez et al., 2022). Currently, the SPI covers an area of ~13219 km<sup>2</sup> with  
125 a volume of  $3632 \pm 675$  km<sup>3</sup>, whereas the NPI covers an area of ~3976 km<sup>2</sup> with a volume of  
126  $1124 \pm 260$  km<sup>3</sup> (Fig. 1). The present-day ice thickness may reach up to ~2000 m in deep  
127 glacial valleys (Millan et al., 2019).

128 GNSS-measured data show ongoing vertical rock uplift rates between  $18 \pm 3$  and  $41 \pm 3$   
129 mm/yr in the northern part ( $18^\circ - 50.5^\circ$  S) of the SPI (Fig. 1b), decreasing to values between  
130  $2 \pm 6$  and  $17 \pm 5$  mm/yr in its southern part ( $50.5 - 51.5^\circ$  S) (Ivins and James, 2004; Dietrich  
131 et al., 2010; Lange et al., 2014; Richter et al., 2016; Lenzano et al., 2023). Such outstandingly  
132 high uplift rates, specially in the northern part of the SPI, are currently ascribed to

133 lithospheric viscoelastic GIA following the LIA, which was responsible for an ice loss of 503  
134  $\pm 101.1 \text{ km}^3$  in the SPI (Glasser et al., 2011). To match the very high observed uplift rate  
135 budget, previous GIA studies infer low asthenosphere viscosity (in the order of  $10^{18} \text{ Pa s}$ ) and  
136 thin elastic lithosphere ( $\sim 35 \text{ km}$  thick) (Ivins and James, 1999, 2004; Klemann et al., 2007;  
137 Dietrich et al., 2010; Lange et al, 2014; Richter et al., 2016; Ávila and Dávila, 2020; Mark et  
138 al., 2022; Lenzano et al., 2023). Although this is consistent with abnormally high  
139 asthenospheric mantle temperatures, viscosity estimates from these previous studies are  
140 untied to the regional thermal regime, which prevents a more thorough characterization of the  
141 role of the asthenospheric window underneath the SPI in affecting the observed uplift rates.  
142 In addition, the contribution of post-LGM deglaciation to present-day rock uplift rate was  
143 marginally addressed (Ivins and James, 1999, 2004; Klemann et al., 2007). Here, we perform  
144 fully coupled thermo-mechanical numerical geodynamic experiments forced by surface  
145 unloading scaled on post-LIA and post-LGM ice melting to evaluate their relative  
146 contribution to the observed regional uplift rates. Numerical experiments account for a range  
147 of positive thermal anomalies in the asthenosphere to further assess the role of the  
148 asthenospheric window in setting the mantle viscosity and associated postglacial rebound.  
149 Focusing on the magnitude rather than the pattern of the inferred surface uplift rates due to  
150 limited information on the spatial-temporal variations of the ice net mass balance and  
151 thickness since the LGM (e.g. Davies et al., 2020), we use the observed budget of rock uplift  
152 rate to constrain plausible thermal and viscosity structures at depth as well as the timing of  
153 postglacial rebound.

154

## 155 **2. Methodology**

156 We used as reference the GNSS-derived data from 31 GPS stations installed by 380  
157 km in north-south and 130 km in east-west directions around the SPI since 1996, published in  
158 Lange et al. (2014). The observed and estimated regional aseismic viscoelastic uplift rates  
159 presented in that study are shown in Fig. 1b. Details on the GPS data acquisition and analysis  
160 are given in the reference study (Lange et al., 2014).

### 161 **2.1. Numerical model**

162 We use a fully coupled thermo-mechanical, visco-elasto-plastic numerical  
163 geodynamic model to quantify the effect of thermal anomalies in the asthenospheric mantle  
164 on the magnitude of surface uplift rates due to deglaciation. We provide a short overview of

165 the governing equations hereafter, while a detailed description of numerical technique can be  
 166 found, for instance, in Gerya and Yuen (2007), Gerya et al. (2019), Sternai (2020), Sternai et  
 167 al. (2021), and Muller et al. (2022). The continuity equation allows for the conservation of  
 168 mass during the displacement of a geological continuum:

$$169 \quad (1) \quad \frac{\partial \rho}{\partial t} + \nabla(\rho v) = 0$$

170 where  $\rho$  is the local density,  $t$  is time,  $v$  is the velocity vector, and  $\nabla$  is the divergence  
 171 operator. The momentum equation describes the changes in velocity of an object in the  
 172 gravity field due to internal and external forces:

$$173 \quad (2) \quad \frac{\partial \sigma_{ij}}{\partial x_i} + \rho g_i = \rho \left( \frac{\partial v_i}{\partial t} + v_j \frac{\partial v_i}{\partial x_j} \right)$$

174 where  $\sigma_{ij}$  is the stress tensor,  $x_i$  and  $x_j$  are spatial coordinates, and  $g_i$  is the  $i$ -th component of  
 175 the gravity vector. The energy equation allows for the conservation of energy during  
 176 advective and conductive heat transfer in the continuum:

$$177 \quad (3) \quad \rho C_p \frac{DT}{Dt} - \text{div}(c \nabla T) + v \nabla T = H_r + H_s + H_a + H_l$$

178 where  $P$  is pressure,  $T$  is temperature,  $C_p$  is specific heat capacity at a constant  $P$ ,  $c$  is the  
 179 thermal conductivity,  $H_r$ ,  $H_s$ ,  $H_a$ , and  $H_l$  are the volumetric heat productions by radiogenic,  
 180 shear, adiabatic and latent heat, respectively.  $H_a \propto \frac{DP}{Dt}$ ,  $H_s = \sigma'_{ij} \dot{\epsilon}'_{ij(viscous)}$ , where  $\sigma'_{ij}$  is  
 181 the deviatoric stress tensor and  $\dot{\epsilon}'_{ij(viscous)}$  is the viscous deviatoric strain rate tensor. Ductile  
 182 deformation is thermally activated generating viscous flow, calculated according to the  
 183 material shear viscosity,  $\eta$ :

$$184 \quad (4) \quad \eta = \frac{1}{2A_d \sigma'_{II}{}^{n-1}} \exp\left(\frac{E_a + PV_a}{RT}\right)$$

185 where  $\sigma'_{II}$  is the second invariant of the deviatoric stress tensor,  $n$  is the stress exponent,  $A_d$  is  
 186 the pre-exponential factor,  $E_a$  is the activation energy,  $V_a$  is the activation volume, and  $R$  is  
 187 the gas constant.  $\dot{\epsilon}'_{ij(viscous)}$ , is computed as:

$$188 \quad (5) \quad \dot{\epsilon}'_{ij(viscous)} = \frac{\sigma'_{ij}}{2\eta} - \delta_{ij} \eta_{bulk} \dot{\epsilon}_{kk}$$

189 where  $\delta_{ij}$  is the Kronecker delta,  $\dot{\epsilon}_{kk}$  is the bulk strain rate in response to irreversible volume  
 190 changes, and  $\eta_{bulk}$  is the bulk viscosity. Recoverable deformation is defined by the elastic  
 191 deviatoric strain rate tensor,  $\dot{\epsilon}'_{ij(elastic)}$ , as:

$$192 \quad (6) \quad \dot{\epsilon}'_{ij(elastic)} = \frac{1}{2\mu} \frac{D\sigma'_{ij}}{Dt}$$

193 where  $\mu$  is the shear modulus and  $\frac{D\sigma'_{ij}}{Dt}$  is the objective co-rotational time derivative of the  
 194 deviatoric stress tensor. The plastic deformation, brittle and localised, occurs at low  
 195 temperature when the absolute shear stress limit,  $\sigma_{yield}$ , is reached, with

$$196 \quad (7) \quad \sigma_{yield} = C + \sin \sin(\varphi)P$$

197 where  $C$  is cohesion and  $\varphi$  is the effective internal friction angle. The plastic strain rate  
 198 tensor,  $\dot{\epsilon}'_{ij(plastic)}$ , is defined as:

$$199 \quad (8) \quad \dot{\epsilon}'_{ij(plastic)} = 0 \text{ for } \sigma_{II} < \sigma_{yield}, \dot{\epsilon}'_{ij(plastic)} = X \frac{\partial \sigma'_{ij}}{2\sigma_{II}} \text{ for } \sigma_{II} \geq \sigma_{yield}$$

200 where  $X$  is the plastic multiplier which satisfies the plastic yielding condition  $\sigma_{II} = \sigma_{yield}$ .  
 201 The bulk strain rate tensor,  $\dot{\epsilon}'_{ij(bulk)}$ , integrates the viscous, elastic and plastic deformation:

$$202 \quad (9) \quad \dot{\epsilon}'_{ij(bulk)} = \dot{\epsilon}'_{ij(viscous)} + \dot{\epsilon}'_{ij(elastic)} + \dot{\epsilon}'_{ij(plastic)}$$

203

## 204 **2.2. Reference model setup and modeling approach**

205 The model domain is 700 km wide and 120 km thick, to account for a region similar  
 206 to the South American continent at latitudes of the SPI, realistic thickness of the lithosphere  
 207 and asthenospheric mantle (van der Meijde et al., 2013; Ávila and Dávila, 2018, 2020), and  
 208 avoid boundary effects in the numerical results. From top to bottom, the model accounts for  
 209 10 km of ‘sticky’ air, 30 km of continental crust (with rheology of quartzite, Ranalli, 1995),  
 210 30 km of lithospheric mantle, and 50 km of asthenospheric mantle (with rheology of dry  
 211 dunite, Ranalli, 1995), in agreement with literature data (e.g., van der Meijde et al., 2013;  
 212 Ávila and Dávila, 2018, 2020). The initial geotherm is piece-wise linear resulting from an  
 213 adiabatic temperature gradient of 0.5 °C/km in the asthenosphere (Turcotte and Schubert,  
 214 2002) and thermal boundary conditions equal to 0 °C at the surface and 1327 °C at the  
 215 bottom of the lithosphere, with nil horizontal heat flux across the vertical boundaries. The  
 216 rheologic and thermal structure of the reference model give a lithospheric elastic thickness,  
 217  $Te$  (*sensu* Burov and Diament, 1995), of ~30 km, comparable to previous estimates  
 218 underneath the SPI based on GIA models (Ivins and James, 1999; Dietrich et al., 2010; Lange  
 219 et al., 2014), heat flow data (Ávila and Dávila, 2018), waveform inversion (Robertson  
 220 Maurice et al., 2003), and low-temperature thermochronology data (Thomson et al., 2010;  
 221 Guillaume et al., 2013; Georgieva et al., 2016, 2019; Stevens Goddard and Fosdick, 2019;  
 222 Ávila et al., 2023; Muller et al., 2023). Rocks rheological properties are listed in Table 1.

223 The numerical model uses the finite differences with marker-in-cell technique,  
224 resolved by  $51 \times 61$  nodes in the horizontal,  $x$ , and vertical,  $y$ , directions, respectively,  
225 distributed on a Eulerian grid that accounts for a maximum resolution of 1 km along the  $y$   
226 direction in the upper part of the model domain, and  $\sim 13$  km in the  $x$  direction.  $400 \times 400$   
227 Lagrangian markers are randomly distributed along the  $x$  and  $y$  dimensions and used for  
228 advecting the material properties (Gerya and Yuen, 2007; Gerya et al., 2019). The material  
229 properties carried by Lagrangian markers are then interpolated onto the Eulerian grid via a 4<sup>th</sup>  
230 order Runge-Kutta interpolation scheme. An internal free surface is simulated through the 10  
231 km thick layer of sticky air. The velocity boundary conditions are free slip at all boundaries  
232 ( $x = 0$  and  $x = 700$  km;  $y = 0$  and  $y = 120$  km).

233 On the top of the crust and in the middle of the model domain we impose a 2 km thick  
234 pseudo-icecap to simulate lithospheric unloading during deglaciation (Fig. 2a). The pseudo-  
235 icecap has an initial density,  $\rho_{ice}$ , of  $920 \text{ kg/m}^3$  (Harvey et al., 2017) (Table 1), and we  
236 compute the surface load through time,  $L$ , as

$$237 \quad (10) \quad L = \rho_{ice} g h_{ice},$$

238 where  $g$  is the gravity acceleration, and  $h_{ice}$  is the icecap thickness. The load change due to  
239 the deglaciation occurs by gradually and uniformly reducing  $h_{ice}$  in time (Fig. 2 b, c). We run  
240 two sets of experiments for the post-LGM deglaciation. In *Model set 1*, 75% of ice loss  
241 occurs in 20000 years (i.e., 1500 m drop of ice thickness, Fig. 2 b), thus assuming a  
242 conservative estimate of ice loss since the beginning of the LGM until the present-day,  
243 simplifying the several glacial retreats and re-advances since the LGM (e.g., Glasser et al.,  
244 2004, 2008, 2011; Davies and Glasser, 2012; Strelin et al., 2014; Kaplan et al., 2016;  
245 Reynhout et al., 2019). In *Model set 2*, 95% of ice loss occurs in 10000 years (i.e., 1900 m  
246 drop of ice thickness, Fig. 2 b), assuming faster deglaciation rates of the Patagonian Ice Sheet  
247 in the first half of post-LGM deglaciation (McCulloch et al., 2000; Hulton et al., 2002; Boex  
248 et al., 2013; Bendle et al., 2017; Thorndycraft et al., 2019; Davies et al., 2020). For the post-  
249 LIA deglaciation, we simulate 10% of ice loss in 400 years (i.e., 200 m drop of ice thickness,  
250 Fig. 2 c), using estimates of ice loss rates since the 19<sup>th</sup> century (Aniya, 1996; Aniya et al.,  
251 1997; Rignot et al., 2003; Ivins and James, 1999, 2004; Chen et al., 2007, Dietrich et al.,  
252 2010; Ivins et al., 2011; Jacob et al., 2012; Willis et al., 2012). The pseudo-icecap is 200 km  
253 wide for the post-LGM model sets 1 and 2, based on estimates of LGM maximum extent of  
254 the Patagonian Ice Sheet (e.g., McCulloch et al., 2000; Hein et al., 2010; Thorndycraft et al.,  
255 2018; Davies et al., 2020), and 70 km wide for the post-LIA model set, based on the



256 estimates of LIA maximum extent of the SPI (e.g., Glasser et al., 2011; Strelin et al., 2014;  
257 Kaplan et al., 2016; Reynhout et al., 2019) (Fig. 2 a).

258 In the models, the lateral extent of the pseudo-icecap does not change throughout the  
259 deglaciation. Although this simplification may affect the inferred pattern of postglacial  
260 rebound, it greatly facilitates the simulation of deglacial lithospheric unloading without  
261 significantly affecting the magnitude of postglacial rebound, which is the main focus here.  
262 All simulations account for some spin up time before the deglaciation begins, so that the  
263 lithosphere-asthenosphere system adjusts to the pseudo-icecap initial load. The uplift rate  
264 during the deglaciation is calculated through time as the surface elevation change resulting  
265 from the modelled strain field divided by the viscoelastic timestep (i.e.,  $U = (z_{curr} -$   
266  $z_{prev})/t$ , where  $z_{curr}$  is the modelled topography at the considered timestep,  $z_{prev}$  is the  
267 modelled topography at the previous timestep, and  $t$  is the viscoelastic timestep duration).  
268 Given the geologically short time window investigated here, we neglect deformation related  
269 to longer term tectonic forces (Breitsprecher and Thorkelson, 2009; Guillaume et al., 2013;  
270 Eagles and Scott, 2014; Muller et al., 2021). The parametric study focuses on the  
271 asthenospheric mantle potential temperature (*sensu* McKenzie and Bickle, 1988) which  
272 accounts for positive thermal anomalies, TA, of up to 200 °C in steps of 50 °C, added to the  
273 reference asthenospheric mantle potential temperature of 1265 °C (McKenzie and Bickle,  
274 1988; Currie and Hyndman, 2006; Ávila and Dávila, 2018, 2020; Sternai, 2020; Mark et al.,  
275 2022) to mimic the presence of a slab window at depth.

### 276 3. Results

277 Results are shown in Table 2 and Figs. 4-7. In agreement with the theory of  
278 lithospheric flexure (e.g., Turcotte and Schubert, 2002) the deglaciation triggers uplift in the  
279 region covered by the melting pseudo-icecap and subsidence in the neighbouring regions  
280 (Figs. 4-6). Overall, increasing the asthenospheric mantle potential temperature decreases the  
281 asthenospheric viscosity, with significant effects on the magnitude of the modelled surface  
282 velocity field. The asthenosphere viscosity ranges between  $10^{22}$ - $10^{19}$  Pa s in simulations with  
283 TA equal to 0 (reference model), 50 and 100 °C, and between  $10^{19}$ - $10^{16}$  Pa s in simulations  
284 with TA equal to 150 and 200 °C (Fig. 3 a-d). Lithospheric warming due to increasing  
285 asthenospheric mantle potential temperature also leads to a reduction of the lower lithosphere  
286 viscosity (from  $10^{22}$  to  $10^{20}$  Pa s), thereby decreasing the integrated lithospheric strength.

287 In *Model set 1* for Post-LGM deglaciation, when TA is 0 (reference model) the  
288 maximum uplift rates is < 1 mm/yr during the first 5000 years of the deglaciation, increasing  
289 gradually up to 9.5 mm/yr in the later stages of the deglaciation (i.e., 20000 years, Fig. 4).  
290 When TA equals 50, 100, 150 and 200 °C, the maximum uplift rates can reach up to ~2, ~5,  
291 ~12, and ~15 mm/yr, respectively, already in the first 1000 years of the deglaciation (Fig. 4  
292 a). When TA is 50 and 100 °C the maximum uplift rate is subject to a protracted increase in  
293 time, reaching up to ~12 and ~14 mm/yr after 20000 years of deglaciation (Figs. 4 b-d and 7  
294 a). For TA equal to 150 and 200 °C, the maximum uplift rate reaches a plateau between 11  
295 and 17 mm/yr during the 20000 years of deglaciation (Figs. 4 and 7 a, Table 2a). After the  
296 end of the deglaciation, the maximum uplift rate takes longer than about 5000 years to re-  
297 equilibrate to 0 mm/yr when  $TA \leq 100$  °C, whereas it drops to 0 mm/yr almost immediately  
298 when TA is 150 or 200 °C (Fig. 7 a).

299 In the *Model set 2* for Post-LGM deglaciation, the maximum uplift rate is less than 2  
300 mm/yr during the first 1000 years of deglaciation when TA is 0, 50 and 100 °C, whereas it  
301 reaches up to ~22 and ~30 mm/yr during in the first 1000 years of deglaciation when TA is  
302 150 and 200 °C (Fig. 5 a, 7 b, and Table 2). Between 5000 and 10000 years of deglaciation,  
303 the maximum uplift rate increases to ~19, ~25 and ~36 mm/yr, respectively when TA is 0, 50  
304 and 100 °C, whereas it reach up to between 36 and 41 mm/yr between 50000 and 1000 years  
305 of deglaciation when TA equal to 150 and 200 °C. The maximum uplift rate decreases slower  
306 if TA is 0, 50 and 100 °C, taking longer than 5000 year after the deglaciation to drop to  
307 values <5 mm/yr (Fig. 7 b and Table 2b), whereas it quickly drops to < 2 mm/yr when the  
308 deglaciation is over and TA is 150 and 200°C (Figs. 5 b-d and 7b). Overall, a warmer and

309 less viscous asthenosphere generates a higher magnitude and fast changing postglacial  
310 rebound than a cooler and more viscous asthenosphere.

311 In the post-LIA model set, the maximum uplift rate is  $\sim 1.4$ ,  $\sim 2.3$  and  $\sim 2.2$  mm/yr  
312 during the first 100 years of deglaciation when TA is respectively 0, 50, and 100 °C, whereas  
313 it reaches  $\sim 8.3$  and  $\sim 23$  mm/yr during the same interval when TA is respectively 150 and 200  
314 °C (Figs. 6 a, 7 c, and Table 2c). Between 200 and 400 years of deglaciation, the maximum  
315 uplift rate reaches  $\sim 1.9$ ,  $\sim 2.5$  and  $\sim 3$  mm/yr when TA equal to 0, 50 and 100 °C, and  $\sim 14$  and  
316  $\sim 25.5$  mm/yr when TA is 150 and 200 °C, respectively (Figs. 6 c-d, 7 c, and Table 2c). When  
317 the deglaciation ends, the maximum uplift rate drops to  $\sim 0$  mm/yr in  $\sim 100$  years when TA  $\leq$   
318 100 °C, whereas it takes longer than 1000 years when TA equals 150 °C or 200 °C (Fig. 7 c).  
319 Overall, a warmer and less viscous asthenosphere generates a higher magnitude postglacial  
320 rebound which, however, takes much longer to re-equilibrate to 0 mm/yr after the end of the  
321 deglaciation than a cooler and more viscous asthenosphere.

322

#### 323 **4. Discussion**

324 Our modelling is simplistic in that we impose a linear and uniform ice loss instead of  
325 a more realistic ice-sheet melting pattern in space and time (Fig. 2b,c). Although the  
326 stratigraphic and geochronologic record is fairly precise for the post-LGM ice extent (e.g.,  
327 Lagabrielle et al., 2004; Rabassa, 2008; Glasser et al., 2011; Davis and Glasser, 2012; Strelin  
328 et al., 2014; Kaplan et al., 2016; Martinod et al., 2016; Bendle et al., 2017; Reynhout et al.,  
329 2019; Davies et al., 2020), information about melting velocities and associated ice thickness  
330 and redistribution of the surface masses are limited for the time windows investigated here.  
331 GNSS, SRTM, and GRACE data constraining the net ice mass balance only during the last  
332 few decades, still showing some discrepancies (e.g., Aniya, 1996; Aniya et al., 1997; Rignot  
333 et al., 2003; Ivins and James, 1999, 2004; Chen et al., 2007; Dietrich et al., 2010; Ivins et al.,  
334 2011; Jacob et al., 2012; Lange et al., 2014; Willis et al., 2012; Richter et al., 2016; Gómez et  
335 al., 2022; Lenzano et al., 2023). Tracing back the post-LGM or Holocene ice loss rate from  
336 current measurement is difficult, considering that climate was at least 6 °C colder during the  
337 LGM (Hulton et al., 2002; Sugden et al., 2002; Seltzer et al., 2021; Yan et al., 2022). As a  
338 result, previous models have assumed simple deglaciation histories as well (e.g., Ivins and  
339 James, 1999, 2004; Hulton et al., 2002; Klemann, 2007; Ivins et al., 2011; Boex et al, 2013).  
340 Measurements of regional erosion rates since the LGM are between 0.02 to 0.83 mm/yr  
341 (Fernandez et al., 2016). However, given the short time intervals investigated here, it seems

342 reasonable to assume that the eroded material is still in the transport zone and therefore does  
343 not significantly contribute to unloading the surface of the orogen. If one refers to erosion  
344 rates from low-temperature thermochronology, although these measures quantify erosion  
345 rates over Myrs and not millennia, Fosdick et al. (2013), Herman and Brandon, 2015,  
346 Fernandez et al., (2016), and Muller et al. (2023), suggests values between 0.1 and 1 mm/yr  
347 from 7 to 4 Ma, followed by a period of erosional quiescence ( $<0.1$  mm/yr), and a possible  
348 increase to 1 mm/yr in the last  $\sim 2$  Ma in the SPI region (Muller et al., 2023). Supposing that  
349 these erosion rates still apply in the last  $\sim 20000$  years, this would translate into 2-20 m of  
350 rocks eroded on average since the LGM, leading to local unloading of approximately 60-600  
351 kPa if one assumes a crustal density of  $3000 \text{ kg/m}^3$ . Such stress change is approximately  
352 equivalent to the melting of about 6-60 m of ice, whereas we simulate the melting of 200-  
353 1500 m of ice in our simulations. The forcing of Quaternary cooling on increasing erosion  
354 rates is, however, debated, and not widely quantified in Patagonia nor worldwide (Valla et  
355 al., 2012; Champagnac et al., 2014; Herman et al., 2013, 2018; Herman and Brandon, 2015;  
356 Fernandez et al., 2016; Georgieva et al., 2019; Yan et al., 2022). Even if long term erosion  
357 rates contribute to present-day uplift rate (Herman et al., 2018), since they are comparable to  
358 those of e.g., the European Alps, we assume a similar contribution to regional uplift rates  
359 (i.e., generally a fraction of a mm/yr; Sternai et al., 2019), that is a negligible contribution in  
360 the context of the Southern Patagonian Andes. We also assume a homogeneous lithosphere  
361 and neglect lateral viscosity variations in the asthenosphere, despite the long-term southern  
362 Andean orogenic history (Cande and Leslie, 1986; Ramos, 2005; Breitsprecher and  
363 Thorkelson, 2009; Muller et al., 2021) and suggested contribution from lateral rheological  
364 heterogeneities (Klemann et al., 2007; Richter et al., 2016). Overall, notwithstanding these  
365 limitations in the model, our fully coupled numerical thermo-mechanical geodynamic  
366 experiments provide realistic uplift rates (Figs. 4-7) that one can compare to current geodetic  
367 observations. Following the example of previous studies (Ivins and James, 1999, 2004;  
368 Klemann et al., 2007; Dietrich et al., 2010; Lange et al., 2014; Richter et al., 2016; Lenzano  
369 et al., 2023), we discuss our results assuming that GNSS-measured rock uplift rates are  
370 mostly related to the deglaciation history and only marginally controlled by the longer term  
371 geodynamics (e.g., Ramos, 2005; Breitsprecher and Thorkelson, 2009; Eagles and Scott,  
372 2014; Muller et al., 2021).

373         The elastic thickness of the lithosphere ( $T_e$ ) varies between the simulations according  
374 to the imposed asthenospheric thermal anomaly, but it is generally lower than 30 km,  
375 resulting in a decoupled lithospheric rheology (*sensu* e.g., Burov and Diament, 1995), as

376 shown by the yield stress envelope in Fig. 2a. This results in predominant elastic deformation  
377 in the upper crust (below the  $\sim 300$  °C isotherm) and upper mantle lithosphere (below the  
378  $\sim 700$  °C isotherm) and viscous deformation in the lower crust, lower lithospheric mantle and  
379 asthenosphere (Fig. 3). We remark that, when we impose higher temperatures in the  
380 asthenospheric mantle, shallower 300 °C and 700 °C isotherms decreases  $Te$  and increases  
381 the isostatic surface uplift rates. Lithospheric thinning due to the asthenospheric window  
382 underneath Southern Patagonia thus affects the regional uplift rates as previously suggested  
383 (Avila and Davila, 2018, 2020; Mark et al., 2022; Ben-Mansour et al., 2022; and Avila et al.,  
384 2023).

385 The inferred maximum post-LIA uplift rate of up to a few mm/yr from experiments  
386 without or with a low asthenospheric thermal anomaly ( $TA \leq 100$  °C, Fig. 7c) are within the  
387 same order of magnitude of maximum uplift rates measured in collisional orogens such as the  
388 European Alps (Sue et al., 2007; Serpelloni et al., 2013; Walpersdorf et al., 2015; Sternai et  
389 al., 2019) and the Himalayas (Larson et al., 1999). Since these collisional orogens are  
390 characterised by a thicker lithosphere (Geissler et al., 2010; Ravikumar et al., 2020), they are  
391 likely less sensitive to mantle dynamics than the Southern Patagonian Andes. When we  
392 consider lithospheric unloading due to post-LGM deglaciation of a wider ice sheet, however,  
393 the inferred maximum uplift rate via *Model set 1* and *Model set 2* reaches up to 10 mm/yr for  
394 and 20 mm/yr, respectively, even without asthenospheric thermal anomaly (Fig. 7a,b). This  
395 suggests a likely contribution from long-term postglacial rebound to the present-day uplift  
396 rates measured in the SPI.

397 In the Southern Patagonian Andes, GIA models estimated the regional asthenosphere  
398 viscosity between  $1.6$  and  $8 \times 10^{18}$  Pa s (Ivins and James, 1999, 2004; Dietrich et al., 2010;  
399 Willis et al., 2012; Lange et al., 2014; Richter et al., 2016; Lenzano et al., 2023). Similarly,  
400 the asthenosphere viscosity from our models when  $TA > 100$  °C is  $< 10^{19}$  Pa s, with the  
401 lowest viscosity value of  $10^{16}$  Pa s imposed where partial melting, supported by the regional  
402 Holocene volcanism (Stern and Kilian, 1996) and by geophysical data (e.g., shear wave  
403 velocity data by Mark et al., 2022), occurs. Under these conditions, however, our experiments  
404 provide max uplift rates between 14 and 26 mm/yr toward the end of the LIA deglaciation  
405 (Fig. 7c). Even with a very low viscosity asthenosphere, the rebound due to short-term post-  
406 LIA deglaciation does not reach the presently observed maximum uplift rates of  $41 \pm 3$   
407 mm/yr. Experiments that account for a low viscosity asthenosphere and long-term post-LGM  
408 deglaciation lasting for 20000 years and 10000 years reach up to  $\sim 25$  and  $\sim 42$  mm/yr of uplift  
409 rate during the final stages of the deglaciation (Fig. 7a-b), respectively, comparable to

410 present-day values. Results, therefore, indicate that the outstanding observational budget of  
411 rock uplift in the SPI is matched only when accounting for higher-than-normal  
412 asthenospheric mantle temperatures, thereby highlighting the relevance of the regional  
413 asthenospheric window. Consistently, although the higher heat flow is currently further north  
414 from our study region, near the CTJ (46-48 °S) (Ramos, 2005; Breitsprecher and Thorkelson,  
415 2009; Avila and Davila, 2018, 2020; Ben-Mansour et al., 2022), increased asthenospheric  
416 temperatures beneath the Southern Patagonia is highly supported by the geophysical data  
417 (e.g., Russo et al., 2010, 2022; Mark et al., 2022; Avila and Davila, 2018, 2020; Ben-  
418 Mansour et al., 2022).

419 Because of the limited knowledge regarding the timing and amount of ice loss since  
420 the LGM (e.g., Ivins and James, 1999, 2004; Hulton et al., 2002; Klemann, 2007; Boex et al.,  
421 2013; Davies et al., 2020), it is difficult to position in time present-day uplift rate  
422 measurements within the investigated deglaciation scenarios to assess the contribution of  
423 post-LGM, post-LIA, and present-day deglaciation to the maximum uplift rate budget. In the  
424 faster post-LGM deglaciation scenario (*Model set 2*) the observed maximum uplift rate  
425 budget is attained in about 10000 years of deglaciation, but only minor residual rebound  
426 could be observed today regardless of the amount of ice loss (Fig. 7 b). If post-LGM  
427 deglaciation occurred slower (*Model set 1*), this event may contribute up to 40% to the  
428 present-day uplift rate budget (Fig. 7a). Although it is difficult to reconcile this scenario with  
429 the geomorphological and geochronological evidences (Hulton et al., 2002; Boex et al., 2013;  
430 Davis and Glasser, 2012; Martinod et al., 2016; Bendle et al., 2017; Thorndycraft et al., 2019;  
431 Davies et al., 2020), it appears that post-LIA rebound alone cannot cover the entire budget of  
432 the observed uplift rates even with the highest tested TA, which points to a non-negligible  
433 contribution from post-LGM deglaciation. This latter conclusion is reinforced by estimates of  
434 the mantle relaxation time,  $\tau_r$ , as (Turcotte and Schubert, 2002):

$$435 \quad (11) \quad \tau_r = \frac{4\pi\nu}{g\lambda},$$

436 where  $\nu$  is the asthenosphere viscosity,  $\lambda$  is the width of the ice sheet, and  $g$  is the  
437 gravity acceleration. Using  $10^{16} < \nu < 10^{18}$  Pa s and  $\lambda = 200$  km leads to  $\sim 2000 < \tau_r < \sim 200000$   
438 years, a time range considerably longer than the post-LIA deglaciation and including full  
439 Pleistocene glacial-interglacial cycles (Ruddiman et al., 1986). Although increasingly  
440 negative ice mass balance in the last  $\sim 50$  years contribute to the elastic lithospheric uplift  
441 rates (Dietrich et al., 2010, Lange et al., 2014), a longer term contribution from the viscous

442 lithosphere is necessary to explain the GNSS-measured uplift rates and (Ivins and James,  
443 2004; Dietrich et al., 2010; Lange et al., 2014; Richter et al., 2016; Lenzano et al., 2021).

444 As a final consideration, our models suggest that we shall measure regional uplift  
445 rates in the order of the tens of cm/yr in the next century if the currently observed ice loss rate  
446 of at least -20 Gt/yr in the SPI (Willis et al., 2012) will continue until the total meltdown of  
447 the ice sheet in ~200 years.

448

## 449 **5. Conclusions**

450

451 We propose that rock uplift rates of up to 40 mm/yr in the Southern Andes are due to  
452 both post-LIA and long-term post-LGM lithospheric rebound, as postulated for other  
453 glaciated orogens (e.g., the European Alps, Fennoscandia, and North America, Peltier et al.,  
454 2018). We also propose that currently observed uplift rates in the Southern Andes are  
455 enhanced by a mantle thermal anomaly of at least 150 °C due to the regional asthenospheric  
456 window. Asthenospheric thermal anomalies higher than 200 °C are unlikely and would  
457 decrease the asthenospheric viscosities to unrealistic values (less than  $10^{16}$  Pa s). Our thermo-  
458 mechanical visco-elasto-plastic forward modelling approach thus helps constraining the  
459 increase in temperature in geodynamic asthenospheric upwelling contexts such as in Southern  
460 Patagonia (Russo et al., 2010, 2022; Avila and Davila, 2018, 2020; Mark et al., 2022; Ben-  
461 Mansour et al., 2022).

462

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471

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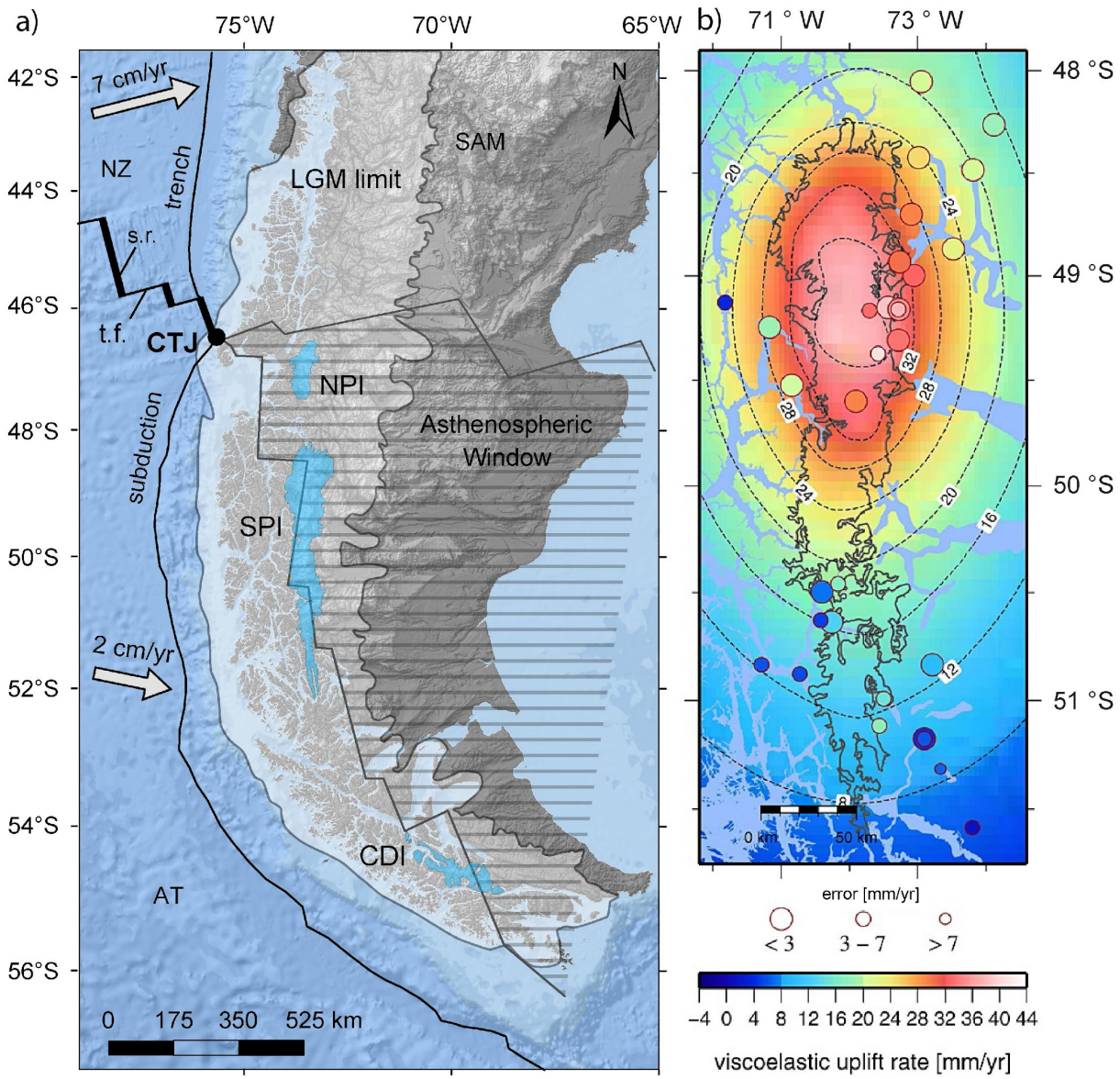
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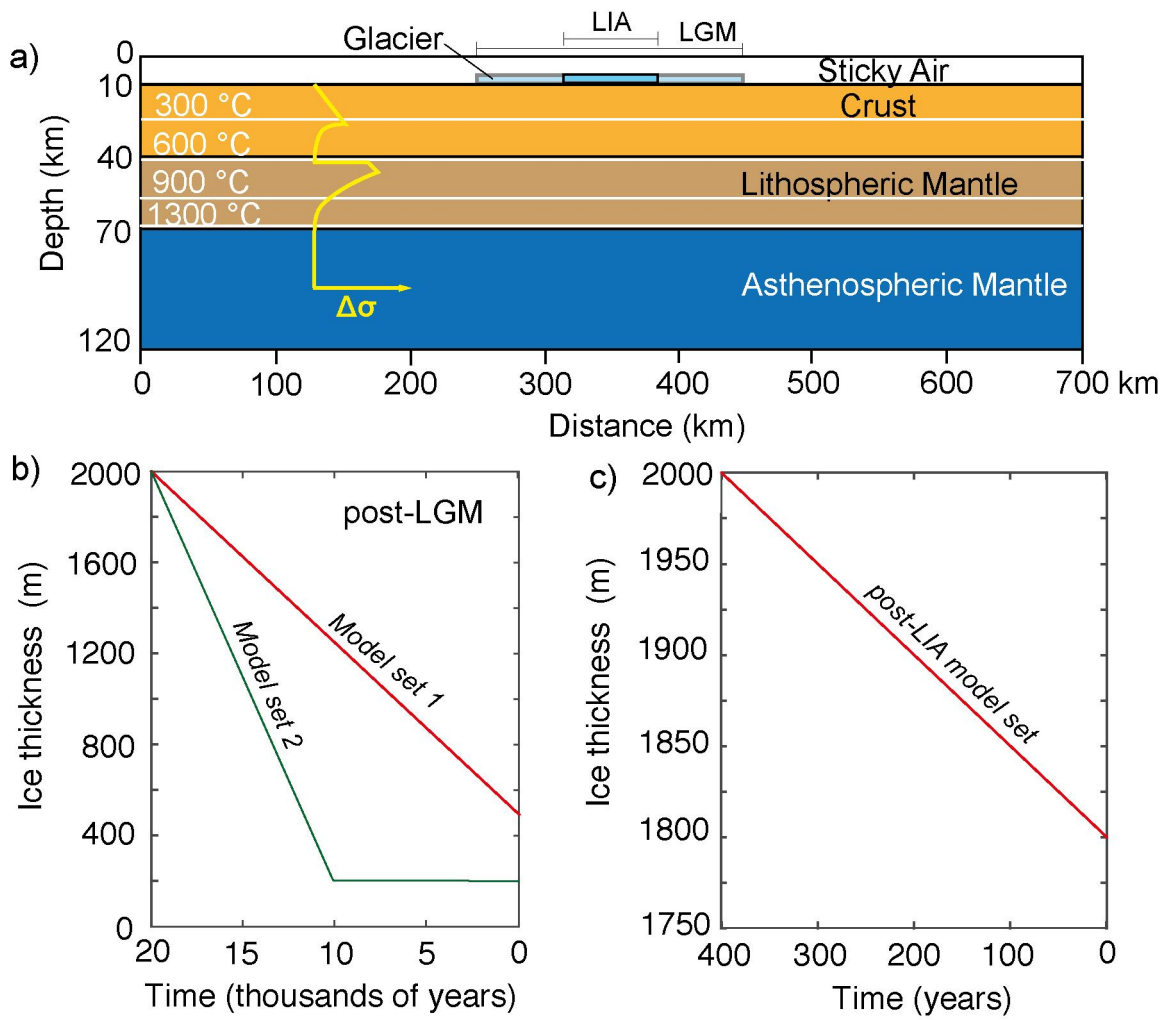


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88 **Fig. 1. Regional context and uplift rate data.** a) Map of southern Patagonia with the Southern Patagonian Icefield (SPI), Northern  
 89 Patagonian Icefield (NPI), and the Cordillera Darwin Icefield (CDI) in light blue, the approximate extension of the Patagonian Ice Sheet at  
 90 the Last Glacial Maximum (LGM) (adapted from Thorndycraft et al., 2019), and the approximate extension of the present-day  
 91 asthenospheric window (dashed region) beneath the South American Continent (SAM) (adapted from Breitsprecher and Thorkelson,  
 92 2009). In the Pacific Ocean, the spreading ridges (s.r., thick black lines) and transform faults (t.f., thin black lines) separate the Nazca (NZ)  
 93 and the Antarctic (AT) plates. The subduction trench is also highlighted in black. The arrows show the approximate rate and direction of  
 94 subduction of the oceanic plates (adapted from DeMets et al., 2010). b) Zoom on the SPI with GNSS-measured rock uplift rates (color-  
 95 coded disks) used to estimate the viscoelastic uplift rates in Lange et al. (2014).



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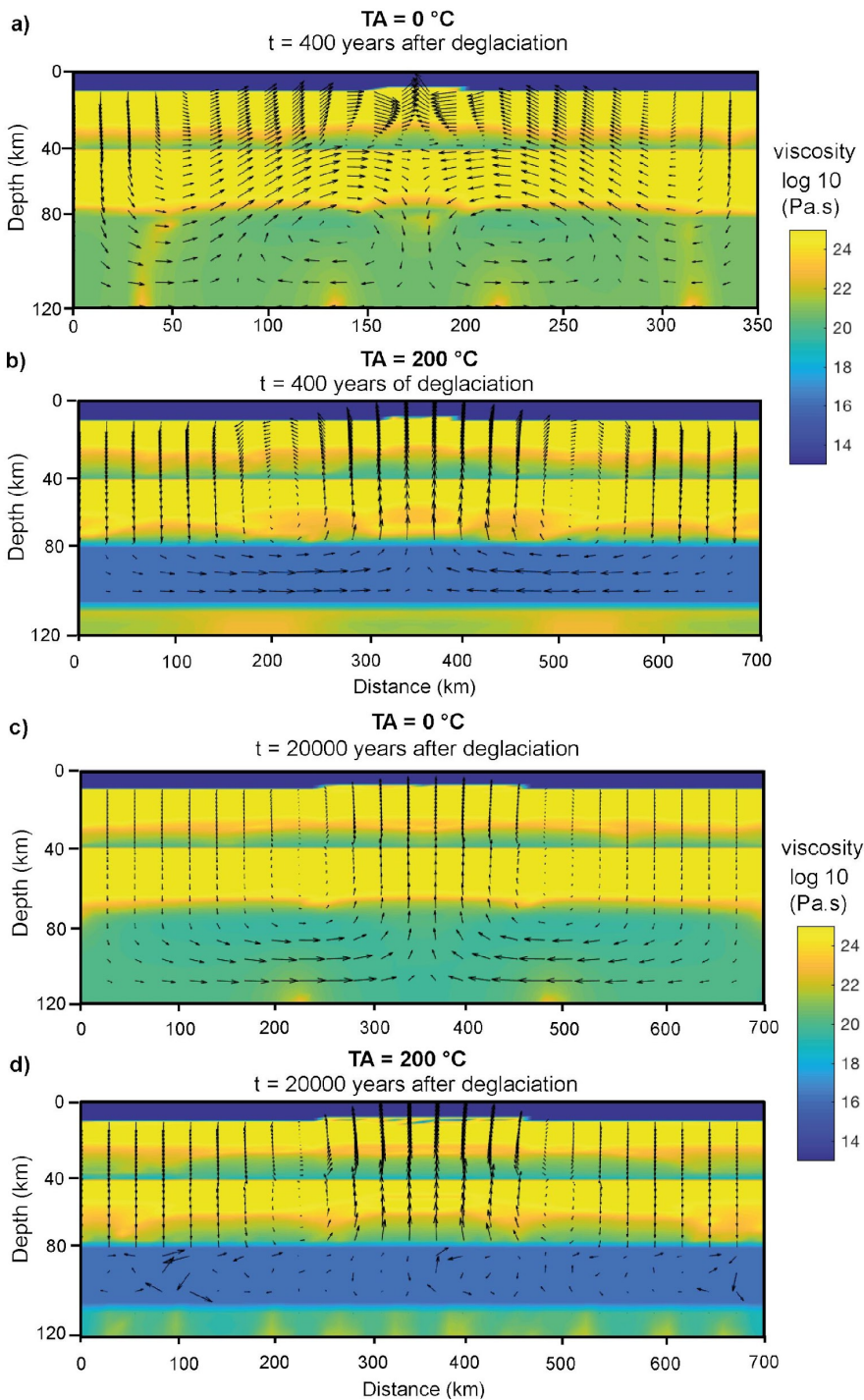
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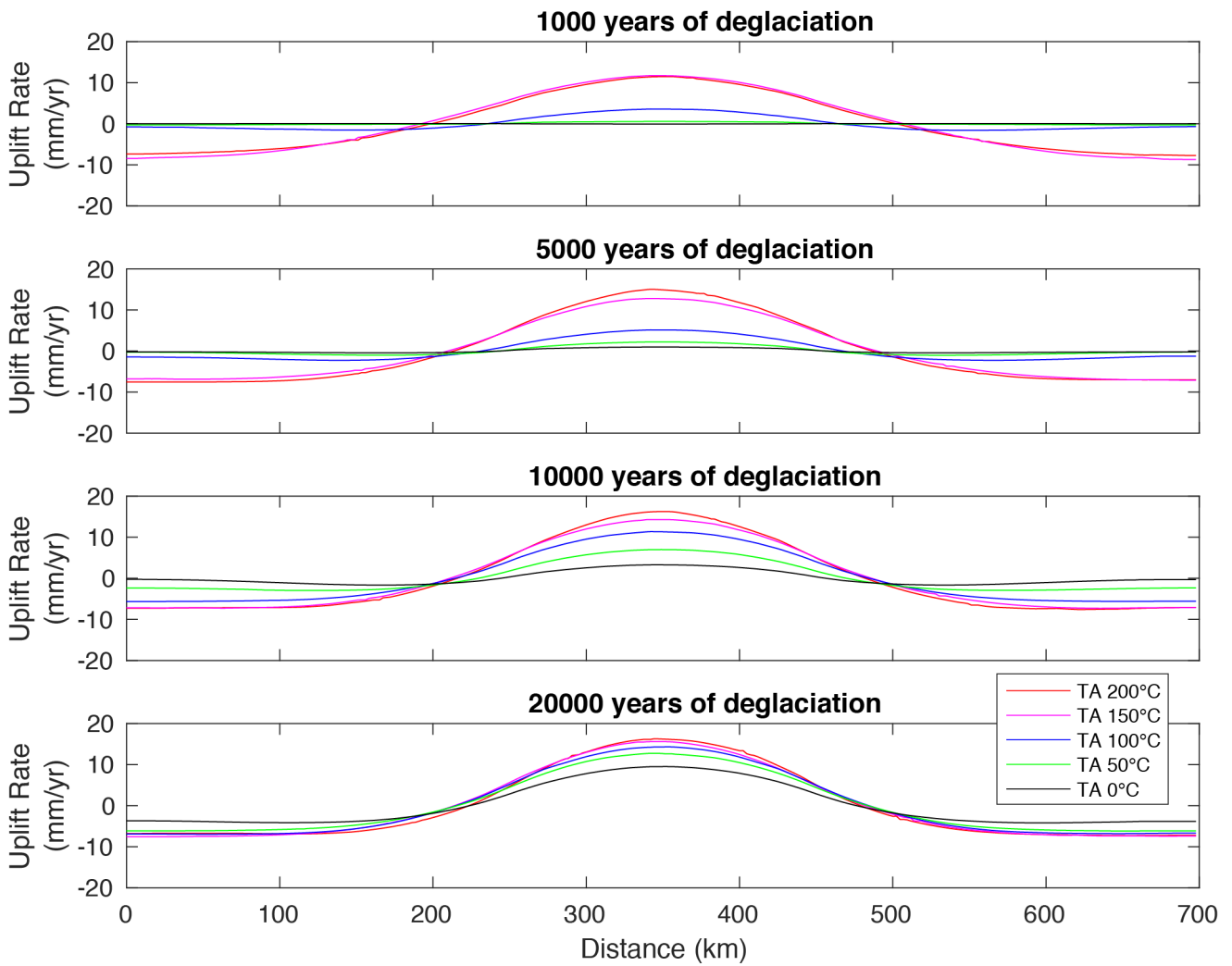
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**Fig. 2. Reference numerical model setup.** a) Thermo-mechanical numerical model domain with rheological layers (Table 1), isotherms (white lines), and yield strength ( $\Delta\sigma = \sigma_1 - \sigma_3$ ) profile (yellow line). The yield strength ( $\Delta\sigma$ ) profile is not scaled and aims to show the proportionality of the yield strength amongst the layers, dependent on the temperature and composition (Eq. 4). (b, c) Ice thickness vs. time used in the numerical models to simulate the post-LGM deglaciation in two model sets (b), and the post-LIA deglaciation (c).



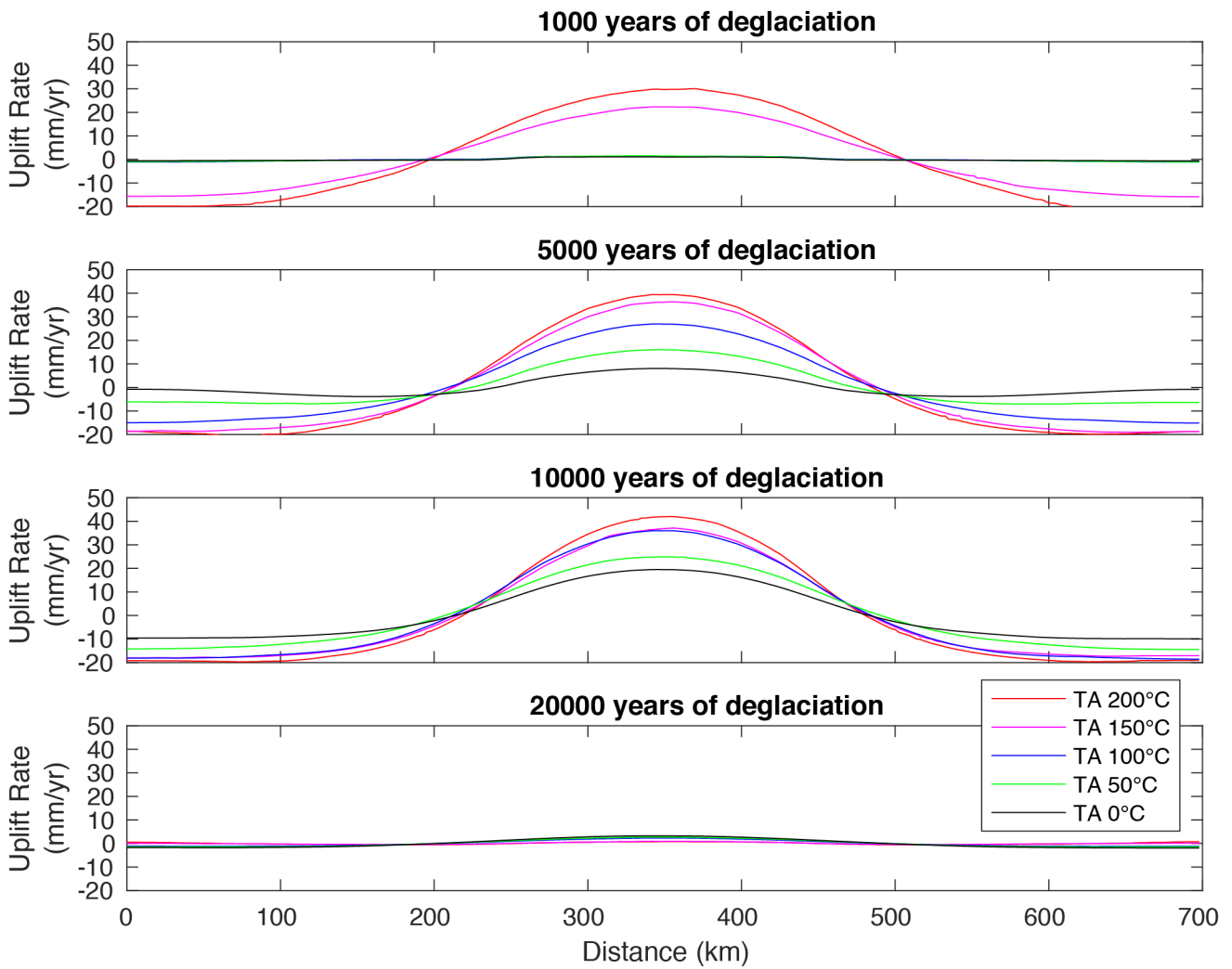
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**Fig. 3. Distribution of viscosity and velocity vectors in the numerical models.** a, c) Reference model without an asthenospheric thermal anomaly,  $TA = 0\text{ }^{\circ}\text{C}$ , in the last timestep of post-LIA deglaciation (a) and of *Model set 1* of post-LGM deglaciation (c). b, d) Model with the higher simulated asthenospheric thermal anomaly,  $TA = 200\text{ }^{\circ}\text{C}$ , in the last timestep of post-LIA deglaciation (b) and of *Model set 1* of post-LGM deglaciation. *Model set 2* has a very similar viscosity and velocity vectors distribution with *Model set 1* in the last deglaciation timestep. Velocity vectors do not have the same scaling and are only meant for visualization purpose.



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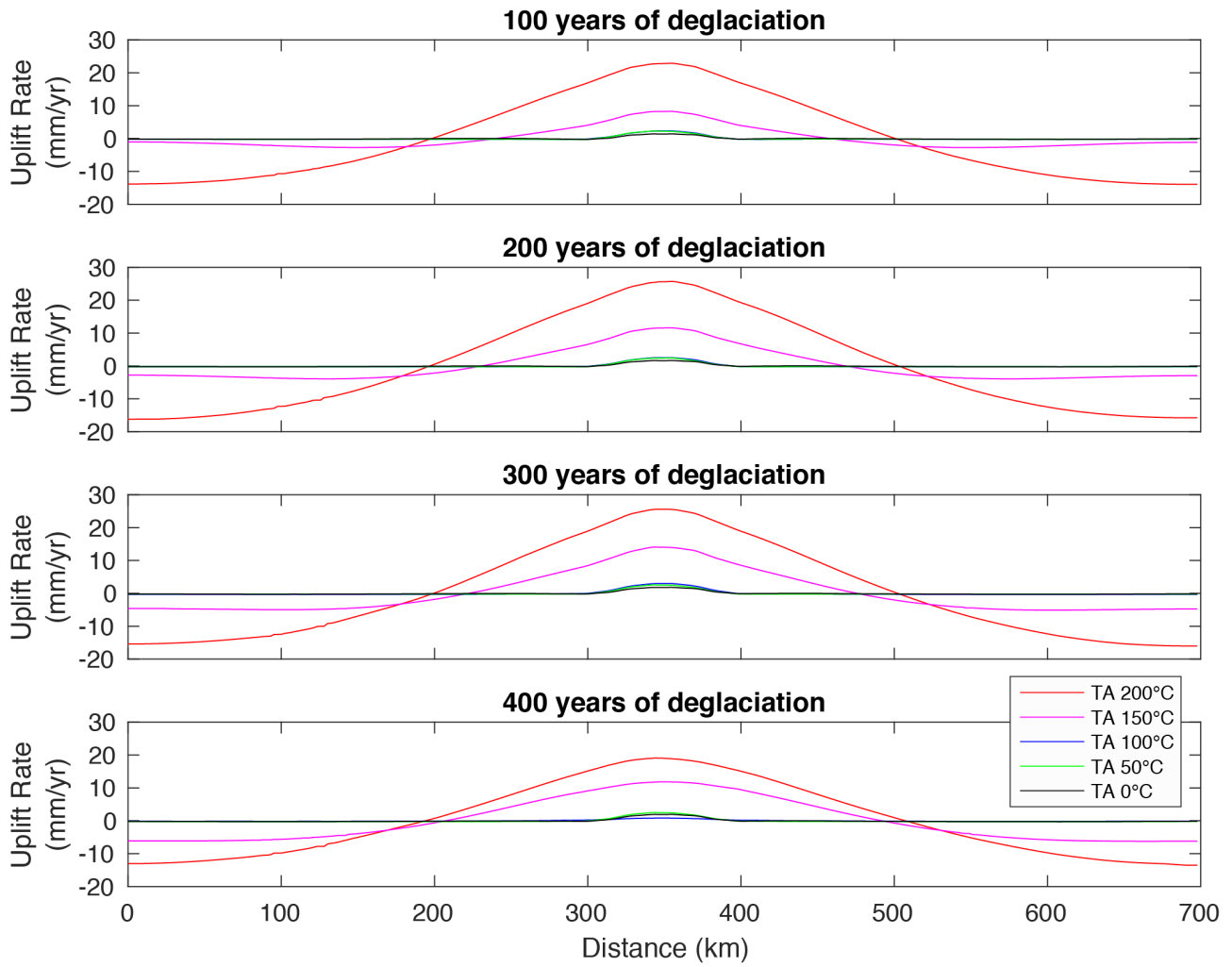
**Fig. 4. Surface uplift rates vs. distance for Model set 1 of post-LGM deglaciation.** a)  $t = 1000$  years of deglaciation, b)  $t = 5000$  years of deglaciation, c)  $t = 10000$  of deglaciation, d) 20000 years of deglaciation. Different line colours correspond to different TA.



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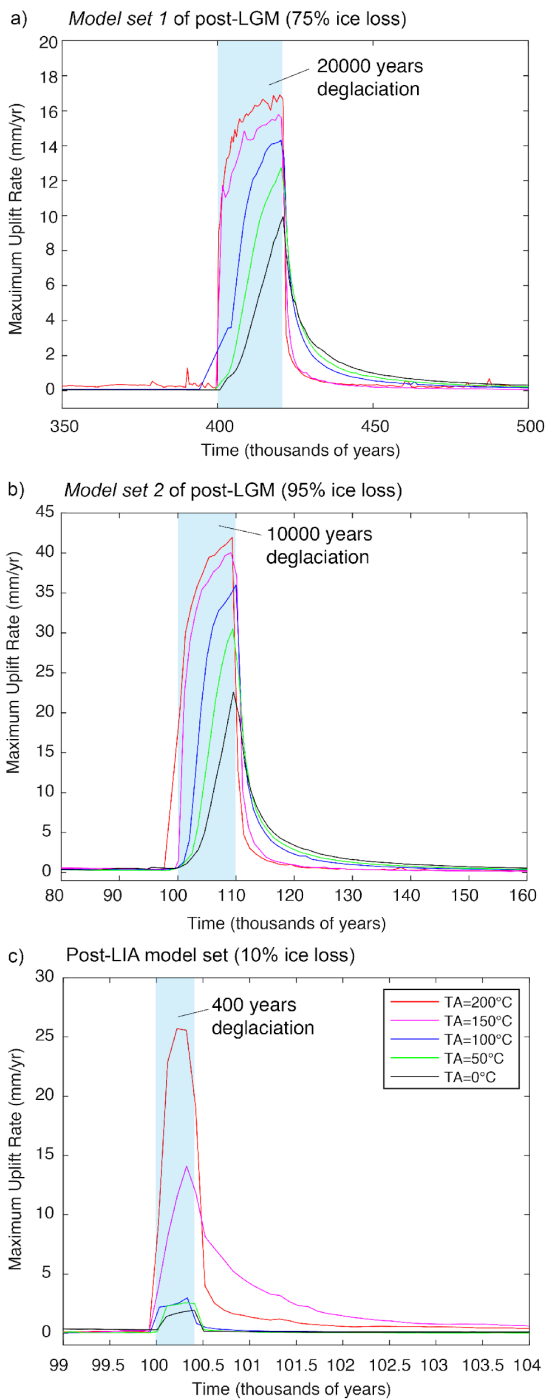
14 **Fig. 5. Surface uplift rates vs. distance for Model set 2 of post-LGM deglaciation.** a)  $t = 1000$  years of deglaciation, b)  $t = 5000$  years  
 15 of deglaciation, c)  $t = 10000$  of deglaciation, d) 20000 years of deglaciation. Different line colours correspond to different TA.

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**Fig. 6. Surface uplift rates vs. distance for post-LIA deglaciation model set.** a)  $t = 100$  years of deglaciation, b)  $t = 200$  years of deglaciation, c)  $t = 300$  years of deglaciation, d) 400 years of deglaciation. Different line colours correspond to different TA.



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**Fig. 7. Maximum uplift rates vs. time for model sets of deglaciation with different TA.** a) *Model set 1* of post-LGM deglaciation accounting of 75% of ice loss in 20000 years, deglaciation starts at 400000 years. b) *Model set 2* of post-LGM deglaciation accounting of 95% of ice loss in 10000 years, deglaciation starts at 100000 years and c) Post-LIA deglaciation model set accounting 10% of ice loss in 400 years (blue-shaded region), deglaciation starts at 100000 years. Blue-shaded regions highlight the modelled deglaciation intervals. Please note that the time axis in a, b, and c are different and post-LGM models account for longer timescales.



26 **Table 1** – Material properties used in the numerical experiments.

	$\rho_0^s$ (km/m <sup>3</sup> )	$E_a$ (kJ/mol)	$V_a$ (m <sup>3</sup> /mol)	$n$	$C$ (Mpa)	$Visc.$ flow law	$Sin$ ( $\phi_{eff}$ )	$c$ (W/m/K)	$\mu$ (Gpa)	$C_p$ (J/kg/K)	$H_r$ ( $\mu$ W/m <sup>3</sup> )	$H_l$ (kJ/kg)	$\alpha$ (1/k)	$\beta$ (1/Pa)
<b>Crust</b>	2800	154	0	2.3	10	Wet Qz.	0.2	$0.64+807/(T+77)$	10	1000	1	300	$3 \times 10^{-5}$	$1 \times 10^{-11}$
<b>Lithospheric mantle</b>	3250	532	10	3.5	10	Dry Ol.	0.6	$0.73+1293/(T+77)$	67	1000	0.022	400	$3 \times 10^{-5}$	$1 \times 10^{-11}$
<b>Asthenospheric mantle</b>	3250	532	10	3.5	10	Dry Ol.	0.6	$0.73+1293/(T+77)$	67	1000	0.022	400	$3 \times 10^{-5}$	$1 \times 10^{-11}$
<b>Ice</b>	920	154	0	2.3	10	-	0	$0.73+1293/(T+77)$	67	1000	0.022	400	$3 \times 10^{-5}$	$1 \times 10^{-11}$

27  $\rho_0^s$  is the standard densities of solid rocks;  $E_a$  is the activation energy;  $V_a$  is the activation volume;  $n$  is the stress exponent;  $C$  is cohesion;  
 28  $\phi_{eff}$  is the effective internal friction angle;  $c$  is thermal conductivity;  $\mu$  is the shear modulus;  $C_p$  is the specific heat capacity;  $H_r$  and  $H_l$   
 29 are the radiogenic and latent heat productions, respectively;  $\alpha$  is thermal expansion;  $\beta$  is compressibility. Qz and Ol are quartzite and  
 30 olivine, respectively. All rheological and partial melting laws/parameters are based on experimental rock mechanics and petrology  
 31 (Ranalli, 1995; Hirschmann, 2000; Johannes, 1985; Turcotte and Schubert, 2002).  
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**Table 2 – Maximum uplift rates derived from the numerical models with a thermal anomaly (TA) of 0, 50, 100, 150 and 200 °C for the Model set 1 (a) and Model set 2 (b) of post-LGM deglaciation, and the post-LIA deglaciation model set (c). The t = 0 is the timestep immediately before the beginning of deglaciation, and other selected timesteps show how the uplift rates change during the deglaciation until it is over for the post-LGM (a,b) and post-LIA(c) deglaciation intervals. Fig. 7 is a plot of the maximum uplift rate vs. time calculated for each timestep in all numerical models.**

<b>a) Model set 1 of post-LGM deglaciation (20000 years)</b>							
<b>TA (°C)</b>	<b>Maximum uplift rate (mm/yr)</b>						
<b>0</b>	0.04	0.04	0.98	3.28	6.43	9.50	4.98
<b>50</b>	0.05	0.56	2.21	6.10	10.76	12.75	4.66
<b>100</b>	0.07	3.58	5.14	11.37	13.63	14.31	4.07
<b>150</b>	0.05	11.72	12.79	14.32	15.18	15.59	1.39
<b>200</b>	0.15	11.48	15.02	16.26	16.46	16.26	0.90
	<b>t = 0</b>	<b>t = 1000 yr</b>	<b>t = 5000 yr</b>	<b>t = 10000 yr</b>	<b>t = 15000 yr</b>	<b>t = 20000 yr</b>	<b>t = 25000 yr</b>
<b>b) Model set 2 of post-LGM deglaciation (10000 years)</b>							
<b>TA (°C)</b>	<b>Maximum uplift rate (mm/yr)</b>						
<b>0</b>	0.50	1.09	8.03	19.48	5.69	3.12	2.15
<b>50</b>	0.25	1.52	15.93	24.87	5.24	2.72	1.73
<b>100</b>	0.33	1.29	26.94	36.02	4.94	2.30	1.41
<b>150</b>	0.43	22.30	36.33	37.11	1.93	0.93	0.60
<b>200</b>	0.37	30.05	39.46	41.98	1.48	0.75	0.50
	<b>t = 0</b>	<b>t = 1000 yr</b>	<b>t = 5000 yr</b>	<b>t = 10000 yr</b>	<b>t = 15000 yr</b>	<b>t = 20000 yr</b>	<b>t = 25000 yr</b>
<b>c) post-LIA deglaciation model set (400 years)</b>							
<b>TA (°C)</b>	<b>Maximum uplift rate (mm/yr)</b>						
<b>0</b>	0.43	1.412	1.67	1.84	1.95	0.18	0.10
<b>50</b>	0.03	2.28	2.43	2.57	2.45	0.30	0.23
<b>100</b>	0.03	2.20	2.32	2.52	2.99	0.49	0.38
<b>150</b>	0.09	8.27	11.57	14.03	11.83	8.11	7.15
<b>200</b>	0.10	22.89	25.70	25.57	18.97	4.00	2.55
	<b>t = 0</b>	<b>t = 100 yr</b>	<b>t = 200 yr</b>	<b>t = 300 yr</b>	<b>t = 400 yr</b>	<b>t = 500 yr</b>	<b>t = 600 yr</b>

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