



# Uncertainties in cloud-radiative heating within an idealized extratropical cyclone

Behrooz Keshtgar<sup>1</sup>, Aiko Voigt<sup>2</sup>, Bernhard Mayer<sup>3</sup>, and Corinna Hoose<sup>1</sup>

<sup>1</sup>Institute of Meteorology and Climate Research – Department Troposphere Research, Karlsruhe Institute of Technology, Karlsruhe, Germany

<sup>2</sup>Department of Meteorology and Geophysics, University of Vienna, Vienna, Austria

<sup>3</sup>Meteorological Institute, Ludwig Maximilian University of Munich, Munich, Germany

**Correspondence:** Behrooz Keshtgar (behrooz.keshtgar@kit.edu)

**Abstract.** Cloud radiative heating (CRH) within the atmosphere affects the dynamics and predictability of extratropical cyclones. However, CRH is uncertain in numerical weather prediction and climate models, and this could affect model predictions of extratropical cyclones. In this paper, we present a systematic quantification of CRH uncertainties. To this end, we study an idealized extratropical cyclone simulated at a convection-permitting resolution of 2.5 km, and combine large-eddy simulations at 300 m resolution with offline radiative transfer calculations. We quantify four factors contributing to the CRH uncertainty in different regions of the cyclone: 3D cloud radiative effects, parameterization of ice-optical properties, cloud horizontal heterogeneity, and cloud vertical overlap. The two last factors can be considered to be essentially resolved at 300 m but need to be parametrized at 2.5 km resolution. Our results indicate that the parameterization of ice-optical properties and the cloud horizontal heterogeneity are the two factors contributing most to the uncertainty in CRH. On the other hand, 3D cloud radiative effects are much smaller, especially for stratiform clouds within the warm conveyor belt of the cyclone. Our analysis in particular highlights the potential to improve the simulation of CRH by better representing ice-optical properties. Future work should in particular address how uncertainty in ice-optical properties affects the dynamics and predictability of extratropical cyclones.

## 1 Introduction

Our ability to predict extratropical cyclones is crucial for midlatitude weather as they are often associated with strong winds, heavy rain, and snow. Cloud diabatic processes, such as latent heat release due to phase changes of water, play an important role in shaping the dynamics of extratropical cyclones (Ahmadi-Givi et al., 2004; Booth et al., 2013; Binder et al., 2016). Yet, recent studies have shown that the impact of cloud radiative heating and cooling (CRH) on the dynamics of extratropical cyclones can also be substantial (Schäfer and Voigt, 2018; Keshtgar et al., 2023; Voigt et al., 2023). Our study is motivated by the CRH impact.

The most common effects of cloud-radiation interaction are cooling at the cloud top and warming at the cloud base, which promote convective instabilities within the cloud (Webster and Stephens, 1980). By doing so, CRH can alter cloud microphysical heating and precipitation (Klinger et al., 2017) and the intensity and predictability of an idealized extratropical cyclone



(Keshtgar et al., 2023). Voigt et al. (2023) showed that the impact of CRH on extratropical cyclones differs between low-  
25 level and high-level clouds, implying that cyclones may respond differently to CRH depending on the vertical distribution of  
clouds. On the longer time scales of climate, modeling studies showed that CRH impacts the planetary-scale circulation of the  
atmosphere and its response to surface warming (Li et al., 2015; Voigt et al., 2020).

Small errors in the parameterization of subgrid-scale physical processes can quickly grow to the synoptic scale, leading to  
forecast errors of significant magnitude (e.g., Baumgart et al., 2019). The interaction between radiation and clouds is known to  
30 be uncertain due to many factors, such as simplified radiation schemes or poor representations of clouds in models. Keshtgar  
et al. (2023) showed that the modulation of latent heating by CRH within the ascending region of the cyclone changes the  
vertical motion and the divergent flow near the tropopause. After the divergent flow has changed, changes in potential vorticity  
are amplified by the rotational flow near the tropopause during the highly nonlinear stage of cyclone evolution. Keshtgar et al.  
(2023) further demonstrated that CRH continuously changes the latent heating during the growth phase of the cyclone. This  
35 result indicates that CRH uncertainties might affect the magnitude of latent heating and be relevant for error growth near the  
tropopause.

The correct simulation of CRH is challenging. Several previous studies have addressed the question of how the simulation of  
CRH depends on model assumptions, the representation of clouds in models, and the parameterization of radiative processes.  
Compared to models with coarse horizontal resolutions of 10-100 km and parameterized convection, clouds can be better  
40 simulated in kilometer-scale models and even better in large eddy models (Griewank et al., 2020; Stevens et al., 2020). In  
this paper, we make use of these modeling advances to present the first systematic assessment of CRH uncertainty within an  
extratropical cyclone by simulating the same idealized cyclone as in Keshtgar et al. (2023) at a convection-permitting resolution  
of 2.5 km. We further run large eddy model simulations at a horizontal resolution of 300 m for different regions of the cyclone  
and use their output for offline radiative transfer calculations. We assume that we have perfect knowledge of the clouds from  
45 the large eddy model simulations and this leads us to prioritize the following four factors as the potential sources of uncertainty  
in CRH in our study:

– 3D cloud radiative effects:

3D cloud radiative effects arise from horizontal photon propagation that is not taken into account in 1D radiation schemes,  
as currently operational in weather and climate models. The main 3D cloud radiative effects are cloud-side illumination  
50 by shortwave radiation especially for high solar zenith angles (Jakub and Mayer, 2015, 2016), cloud-side shortwave  
radiation leakage for low solar zenith angles (Hogan and Shonk, 2013), and cloud-side longwave absorption and emission  
(Klinger et al., 2017). Thus, the largest differences between 3D and 1D radiative transfer calculations are expected to  
occur at the interface between cloudy and clear areas, where the gradient of cloud optical properties is large. Strong  
horizontal variability of in-cloud optical properties might also lead to noticeable horizontal radiative exchanges that are  
55 neglected in 1D radiative transfer schemes.

– Ice-optical parameterization:



The representation of ice-optical properties in models is challenging due to the complexity of ice particles and insufficient observations. This creates an important source of uncertainty for simulating CRH (e.g., Zhao et al., 2018; Yi, 2022). Several parameterizations for the ice-optical properties have been developed based on different assumptions about ice crystal shapes or habits, size distributions, and surface roughness (Yang et al., 2013; Baran et al., 2014; Baum et al., 2014). It is therefore important to examine how the assumed ice-optical properties affect CRH.

– Cloud horizontal heterogeneity and vertical overlap assumption:

Clouds are not homogeneous over their horizontal and vertical extent, and this spatial variability significantly affects their interaction with radiation. Cloud horizontal heterogeneity refers to the horizontal variation of cloud optical properties within a grid box. The vertical overlap determines how cloud layers are stacked on top of each other. The ability of models to account for cloud horizontal heterogeneity and vertical overlap depends primarily on their horizontal resolution. Neglecting cloud horizontal heterogeneity increases longwave emissivity and shortwave absorption of clouds, an effect known as the plane-parallel problem of radiative transfer calculations (e.g., Črnivec and Mayer, 2019). In current models with resolutions on the order of 1-10 km, parameterizations are used to represent both cloud horizontal heterogeneity within grid boxes and cloud vertical overlap between adjacent grid boxes. As a result, assumptions in the parameterizations of these two subgrid effects can lead to errors in the simulation of CRH (e.g., Wang et al., 2021).

We organize the manuscript as follows. The model simulation setup and radiative transfer calculations are described in Sect. 2. In Sect. 3 we quantify the CRH uncertainty due to the four factors discussed above, and we compare their importance in Sect. 4. The manuscript closes with a discussion and conclusion in Sect. 5.

## 2 Methods

This section describes the model setup for the idealized baroclinic life cycle simulation, the large eddy model simulations, and the offline radiative transfer calculations that we use in combination to investigate the uncertainty in CRH within an extratropical cyclone.

### 2.1 Baroclinic life cycle simulation

We use the ICOSahedral Nonhydrostatic atmosphere model (ICON, version 2.6.2.2; Zängl et al. (2015)) to simulate a baroclinic life cycle. The simulation provides us with an idealized extratropical cyclone for which we can test the uncertainty in CRH. We use the same model setup developed and described by Keshtgar et al. (2023) and only briefly describe the model setup here. The planar channel grid extends 4000 km and 9000 km in the zonal and meridional directions, respectively. The horizontal resolution is 2.5 km. Periodic boundary conditions are applied in the zonal direction, and the Coriolis parameter is set to a uniform value valid at 45°N. In the vertical direction, 75 model levels are used. The initial and lateral boundary conditions follow the life cycle type 1 of Polvani and Esler (2007). The model surface is configured as an ice-free ocean with a prescribed sea surface temperature that is 0.5 K lower than the initial temperature of the lowest model level. To trigger the development



of the cyclone, we add a temperature perturbation with a wavelength equal to the domain length and an amplitude of 1 K at all levels. We run the ICON model with the numerical weather prediction (NWP) physics package, including the shallow  
90 convection scheme of Bechtold et al. (2008), the two-moment microphysics of Seifert and Beheng (2006), and the turbulence scheme of Raschendorfer (2001). Due to the high horizontal resolution of the model, deep convection is treated explicitly.

We use the modeling technique developed by Keshtgar et al. (2023) to simulate the cyclone with only the contribution of CRH; the clear-sky radiative heating is not used for stepping the temperature forward in time. The CRH in the model is given by

$$95 \quad \left. \frac{\partial T}{\partial t} \right|_{\text{radiation}}^{\text{cloud}} = \left. \frac{\partial T}{\partial t} \right|_{\text{radiation}}^{\text{all-sky}} - \left. \frac{\partial T}{\partial t} \right|_{\text{radiation}}^{\text{clear-sky}} = \frac{1}{\rho c_v} \cdot \frac{\partial}{\partial z} (F^{\text{all-sky}} - F^{\text{clear-sky}}), \quad (1)$$

where  $\rho$  is the air density and  $F$  is the radiative flux in all-sky and clear-sky conditions defined as positive downward. Since ICON is based on height levels, the conversion of radiative fluxes to heating rate requires the heat capacity of air at constant volume,  $c_v$ . As described in Keshtgar et al. (2023), the technique eliminates the impact of strong clear-sky radiative cooling early in the simulation.

100 Different from Keshtgar et al. (2023), we use ecRad (Hogan and Bozzo, 2018), the new operational radiation scheme in ICON instead of the Rapid Radiative Transfer Model (Mlawer et al., 1997). The advantage of ecRad is that we can configure the offline radiative transfer setup in a manner that is very close to the radiative transfer setup used in ICON (Sect. 2.3). From the various radiation solvers available in ecRad, we use the homogeneous solver. The solver homogenizes the cloud properties over the entire grid box, i.e., it ignores the cloud fraction and does not require any assumption about cloud vertical overlap.  
105 Thus, for the purpose of radiative transfer calculation, clouds are treated as grid-box clouds. This treatment is the same as in the large eddy model simulations described later. For the cloud optical parameterization, we use the scheme from the Suite Of Community RAdiative Transfer codes (SOCRATES) based on Edwards and Slingo (1996) for liquid water droplets. For ice crystals, we use the scheme of Fu (Fu, 1996; Fu et al., 1998).

Overall, this model configuration results in a typical wintertime extratropical cyclone (Fig. 1 a). Note that for the latitude-  
110 longitude plots, we use the geographic coordinates assigned during the generation of the planar channel grid, but the model simulation is based on the Cartesian coordinate of the grid. The baroclinic simulation provides the initial and lateral boundary conditions for the large eddy model simulations described in the following section.

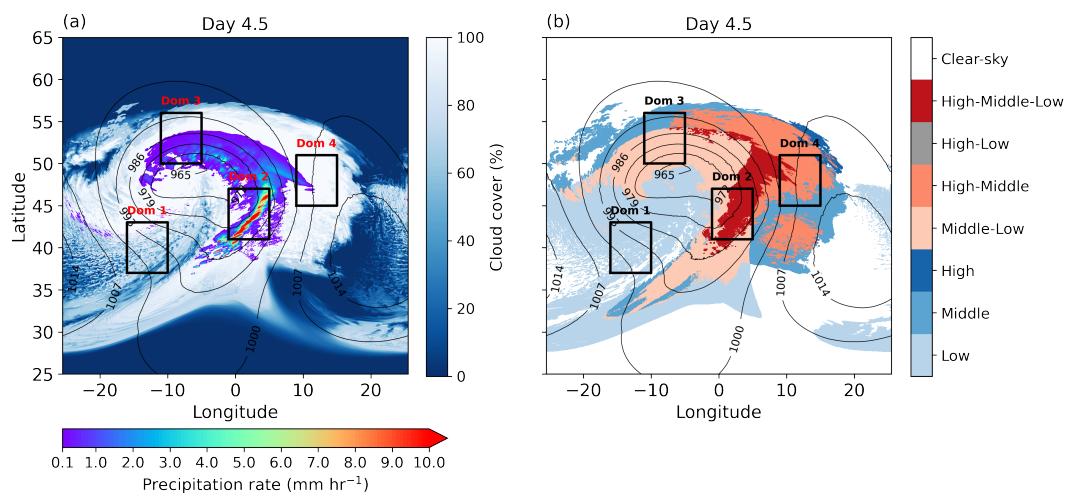
## 2.2 Large eddy model simulations

We target four different regions of the cyclone by the large eddy model (LEM) simulations, shown by black rectangles in Fig. 1.  
115 To characterize the 3D structure of clouds, panel b of Fig. 1 shows cloud classes. We derive the cloud classes from the clouds' vertical extent (Sullivan et al., 2023) using a 50% threshold for low-, mid-, and high-level cloud cover that are diagnosed within ICON during the model simulation. For example, if cloud cover exceeds the 50% threshold at low, middle, and high altitudes, the cloud class is High-Middle-Low.



Domain 1 targets low-level shallow cumulus and stratocumulus clouds southwest of the cyclone center within the cold  
 120 sector. These clouds are formed by the passage of cold, dry air over the warm ocean behind the cold front. Domain 2 targets the  
 extensive vertical clouds within the warm conveyor belt (WCB) of the cyclone, where strong latent heating and precipitation  
 occur. Domain 3 captures mid- and low-level clouds in the cyclonic branch of the WCB outflow, and Domain 4 captures mostly  
 mid- and high-level clouds in the anticyclonic branch of the WCB outflow. We will refer to domains 1, 2, 3 and 4 as shallow  
 cumulus, WCB ascent, WCB cyclonic outflow and WCB anticyclonic outflow, respectively.

125 For the LEM simulations, we use the ICON model in a limited area setup. The LEM simulations use a planar grid with  
 perfectly uniform triangular grid-cell areas. Unlike the planar channel grid used in the NWP simulation, the planar grid does  
 not apply zonal periodic and fixed meridional boundary conditions. Instead, hourly lateral boundary conditions are provided  
 by the baroclinic life cycle simulation. The planar grid extends 471 km and 667 km in the zonal and meridional directions,  
 respectively. This is approximately equivalent to a  $6^\circ \times 6^\circ$  domain at  $45^\circ\text{N}$ . The horizontal resolution is 300 m. We use 150  
 130 model levels.



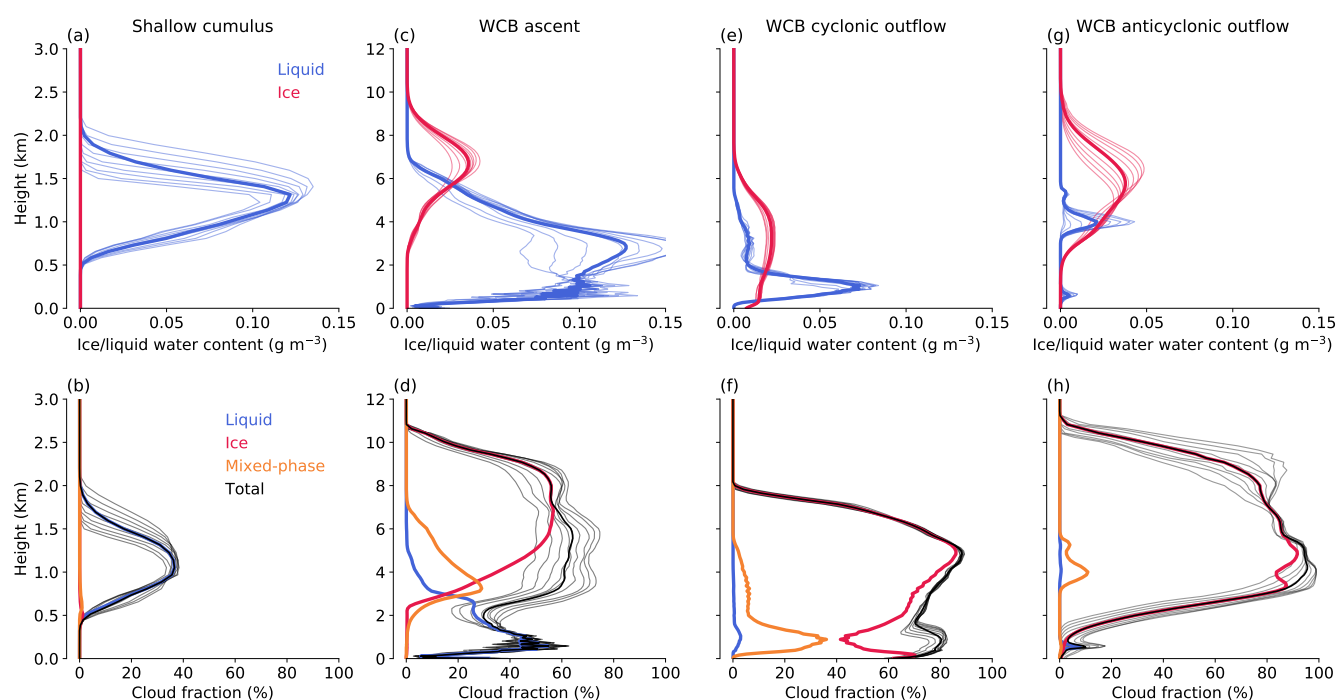
**Figure 1.** (a) Surface pressure (black contours, hPa), cloud cover, and precipitation at day 4.5 in the baroclinic life cycle simulation with ICON-NWP. Panel (b) shows cloud classes. The rectangles in both panels indicate the location of the domains for the LEM simulations.

As in the NWP simulation, we run the LEM simulations with the two-moment microphysics scheme, the homogeneous  
 radiation solver of ecRad, the same cloud optical parameterization, and with only the CRH contribution from the radiation  
 scheme. In contrast to the NWP simulation, we use the 3D Smagorinsky diffusion model (Smagorinsky, 1963) for the turbu-  
 135 lence scheme and switch off the shallow convection scheme. We configure the cloud cover scheme to treat clouds as grid-scale  
 quantities, i.e. grid boxes are either fully cloudy or clear.

We run LEM simulations for 12 hours starting on day 4.25 of the baroclinic life cycle simulation when the cyclone is in its  
 main growth phase. We exclude the first 3 hours to remove spin-up effects. Keshtgar et al. (2023) showed that between days 4  
 and 5, the modulation of latent heating by the CRH is strong enough to affect the dynamics of the cyclone.



The offline radiative transfer calculations described in Sect. 2.3 require as input the 3D fields of cloud ice and liquid water  
 140 content and effective particle radii simulated by the large eddy model. We briefly characterize the clouds from the LEM  
 simulations in Fig. 2, which shows domain and time averages of cloud ice and liquid water contents and cloud fractions for the  
 four LEM domains. We calculate the time average over nine snapshots that are each separated by a 30-minute time interval  
 between hours 10 and 14. We use the same snapshots for the offline radiative transfer calculation. During this time interval,  
 the solar zenith angle changes from low to high values in all regions, which is an important factor for the shortwave 3D cloud  
 145 radiative effects (Sect. 3.2).



**Figure 2.** Profiles of domain-averaged cloud hydrometeor contents and total cloud fraction, decomposed into the contributions from liquid, ice, and mixed-phase clouds for all four LEM domains. Thin lines show profiles for nine snapshots between hours 10 to 14. The thick lines show time-averaged profiles. The threshold used to determine cloudy boxes for both ice and liquid water contents is  $10^{-8}$  kg kg<sup>-1</sup> (Costa-Surós et al., 2020). Note the different y-axes for panels (a) and (b).

In the shallow cumulus region, low-level liquid clouds prevail in the boundary layer between 0-2 km, with cloud fraction  
 peaking at 40% at an altitude of 1.5 km (Fig. 2 a, b). In the WCB ascent region, extensive vertical clouds contain both ice and  
 liquid water. The liquid water content in the lower and middle levels is much higher and more variable in time than the ice  
 water content in the upper levels (Fig. 2 c). The decomposition of cloud fraction shows that liquid clouds are concentrated in the  
 150 lower levels between 0-4 km, where their cloud fraction reaches a maximum of 60% at around 1 km. Mixed-phase clouds are  
 located at intermediate levels between 2-7 km. Above the boundary layer, the fraction of ice clouds increases rapidly, reaching  
 a maximum of nearly 60% at about 8 km (Fig. 2 d).





Clouds in the WCB cyclonic outflow region also contain both ice and liquid water but are located at lower altitudes compared to the WCB ascent region (Fig. 2 e). The total cloud fraction is dominated by ice clouds and reaches nearly 90% at around 6  
155 km. Mixed-phase clouds exist only in the boundary layer (Fig. 2 f). In the WCB anticyclonic outflow, mid- and high-level clouds contain both ice and liquid water, with ice water content dominating over liquid water content (Fig. 2 g). The total cloud fraction is dominated by ice clouds (Fig. 2 h).

The WCB ascent and outflow regions are much cloudier than the shallow cumulus region. In the WCB outflow regions, stratiform ice clouds cover almost the entire domain.

### 160 2.3 Offline radiative transfer calculations

For the offline radiative transfer calculations, we use the LibRadtran software (Mayer and Kylling, 2005; Emde et al., 2016), which includes several radiation solvers and different cloud optical parameterizations. In order to achieve the best possible match between the offline calculations and the CRH calculated with ICON, we configure the radiative transfer setup of LibRadtran as similar as possible to ICON.

165 For gas absorption, we use the parameterization of Fu (Fu and Liou, 1992, 1993), which employs the correlated-k method with 6 and 12 spectral bands in the solar and thermal spectrums, respectively. This parameterization is optimized for use in climate models. The optical properties of cloud droplets are prescribed according to the parameterization of Hu and Stamnes (1993), which assumes spherical droplets. The optical properties of ice crystals are prescribed according to the parameterization of Fu (Fu, 1996; Fu et al., 1998), which is the same scheme used in the ICON simulations. The scheme of Fu uses an  
170 approximate phase function and assumes a pristine hexagonal column habit for the ice crystals. For the radiative transfer solver, we use the classical Delta-Eddington two-stream solver (Zdunkowski, 2007), which is suitable for horizontally homogeneous clouds. This solver is similar to the ecRad homogeneous solver used in the ICON simulations.

In LibRadtran, inputs are required to have normal Cartesian rectangular coordinates. Therefore, we remap the LEM simulation output from the unstructured triangular grid to a regular Cartesian grid with the same resolution using nearest neighbor  
175 interpolation. For the clear-sky background, we use pressure, temperature, density, specific humidity, and trace gas concentration from the LEM simulations. We set other inputs such as surface type, albedo, and emissivity to the same values used in the ICON simulations. We use the solar zenith angles used in the LEM simulations for the offline solar radiative transfer calculations.

The offline radiative transfer calculations are very time-consuming, especially for the 3D Monte Carlo radiative transfer calculation (Mayer, B., 2009). Ideally, one would run the calculations over the entire LEM domain using the 3D cloud and  
180 clear-sky fields. However, this is not feasible because LibRadtran cannot be run in parallel and the wall-clock time of the computing nodes of the Levante DKRZ supercomputer that we are using for our study is limited to 8 hours. To overcome the computational challenge, we divide the LEM domains into  $36 1^\circ \times 1^\circ$  subdomains for which we run the radiative transfer calculations individually. For each subdomain, we use a 1D clear-sky background profile that we derive by horizontally averaging  
185 the clear-sky fields. After the radiative transfer calculations for all subdomains are finished, we merge the results to obtain the radiative heating rate over the entire LEM domain. While this means that we neglect horizontal variability of the clear-sky



background within the subdomains, we show in Sect. 3.1 that this has a negligible impact and we are able to achieve a very good agreement between the CRH calculated offline with LibRadtran in the reference setup using the 1D radiation calculation and online in ICON-LEM.

190 The partitioning of the LEM domains into subdomains allows us to perform expensive 3D Monte Carlo radiative transfer calculation, but can introduce errors in the radiative transfer calculation due to the discontinuity of the horizontal photon transport at the lateral boundaries of the subdomains. As an example, consider a shortwave radiative transfer calculation with a cumulus cloud near the northern boundary of a subdomain. The shadow of this cloud should be present at the southern boundary of the subdomain to the north of it, yet the periodic boundary conditions applied in the 3D solver mean that the shadow is placed at the southern boundary of the cloud's own subdomain. The gravity of such an error depends on the solar zenith angle, cloud top height, and cloud cover. We minimize this issue by making the subdomains overlap in all directions by approximately 12 km. The overlap is sufficient for shallow cumulus clouds with a maximum height of 2 km. The overlap is too small for clouds at higher altitudes in the WCB, but since WCB clouds are mostly stratiform the resulting errors tend to be small and do not affect the generality of our results.

**Table 1.** List of radiative transfer calculations to quantify the uncertainty in CRH. The first column shows the purpose of the calculations, the second column the radiation solvers, and the third column the ice-optical parameterization. The last column lists the sections in which the calculations are discussed.

Purpose	Radiation solvers	Ice-optical parameterization	Section
Reference	$\delta$ -Eddington two-stream	Fu	3.1, 3.3
3D cloud radiative effects	MYSTIC	Fu	3.2
	MYSTIC-ICA	Fu	3.2
Ice-optical parameterization	$\delta$ -Eddington two-stream	Baum-v36, general habit mixture	3.3
	$\delta$ -Eddington two-stream	Baum-v36, solid column	3.3
	$\delta$ -Eddington two-stream	Baum-v36, rough-aggregated	3.3
Cloud horizontal heterogeneity and overlap at 2.5 km NWP resolution (Fig. 3)	$\delta$ -Eddington two-stream	Fu	3.4
	$\delta$ -Eddington two-stream with maximum-random vertical overlap	Fu	3.4

200 To investigate the uncertainty in CRH, we perform several radiative transfer calculations that are listed in Table. 1. For 3D cloud radiative effects, we use the "Monte Carlo code for physically correct tracing of photons in cloudy atmospheres" (MYSTIC, Mayer, B. (2009)). MYSTIC can also be run in the independent column approximation mode (MYSTIC-ICA), which neglects horizontal photon transport between model columns and is equivalent to a 1D radiation scheme. We use the standard forward photon tracing method. The azimuth angle is set to a constant value of zero, which directs the solar radiation

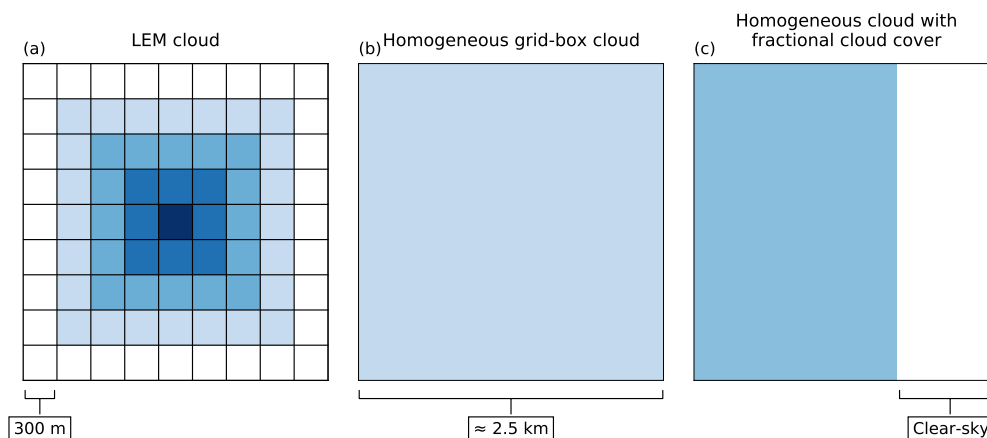




205 from south to north. To obtain low Monte Carlo noise, we run MYSTIC with 720 million photons (nearly 5000 photons per LEM column) for each subdomain in the LEM domains.

To assess the impact of the ice-optical parameterization, we use the more complex scheme of Baum-v36 (Yang et al., 2013; Baum et al., 2014) in addition to the scheme of Fu. The scheme of Baum includes the full phase function and three different ice habits: general habit mixture, solid column, and rough-aggregated. Ice crystal surfaces are treated as severely roughened in  
210 the Baum scheme. Zhao et al. (2018) compared the radiative fluxes calculated with the schemes of Fu and Baum and showed that in the longwave spectrum, the mass absorption coefficient is smaller in the Baum scheme than in Fu, resulting in weaker longwave radiative cooling at the cloud top and weaker warming below (Fig. 1 of Zhao et al. (2018)). In the shortwave spectrum, the backward scattering of radiation by ice crystals is higher in the Baum scheme due to the reduced parameterized asymmetry factor and forward peak. This reduces the radiative flux gradient between cloudy and clear sky layers, resulting in a weaker  
215 CRH with the ice scheme of Baum.

To study the impact of cloud horizontal heterogeneity on CRH, it would be misleading to directly compare the CRH between the NWP and LEM setups since the two setups simulate different clouds. Therefore, we instead create homogeneous clouds at the 2.5 km NWP resolution by coarse-graining the LEM clouds. The method is illustrated in Fig. 3. We derive two types of homogeneous NWP clouds. For the first type, we average the cloud fields over all LEM grid boxes that fall within the  
220 corresponding NWP grid box with a resolution of 2.5 km. We call this type "homogeneous grid-box clouds". The type assumes that clouds are resolved at the NWP resolution so that grid boxes are either entirely cloudy or clear and no vertical overlap assumption is needed. For the second type, we average the cloud fields only over those LEM boxes that are cloudy, and we use the number of clear-sky boxes to determine the cloud fraction. We refer to this type as "homogeneous clouds with fractional cloud cover". The type requires an overlap assumption, for which we use the Delta-Eddington two-stream solver that applies  
225 maximum-random overlap (Črnivec and Mayer, 2019).



**Figure 3.** Illustration of the method to derive homogeneous NWP clouds from LEM clouds. The plot shows a layer of grid boxes. (a) Clouds in the LEM simulation. (b) Homogeneous grid-box cloud at a resolution of 2.5 km. (c) Homogeneous cloud with fractional cloud cover at a resolution of 2.5 km.

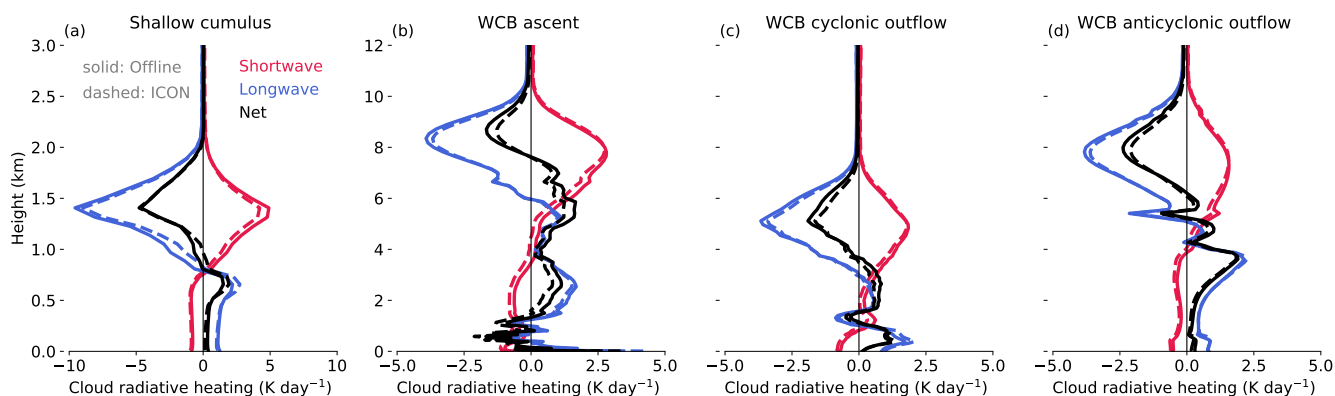


### 3 Results

In this section, we assess CRH in the four regions of the cyclone. We derive CRH from all-sky and clear-sky radiative heating in the shortwave and longwave spectrum for each of the radiative transfer settings listed in Table. 1. To quantify the impact of the four factors described in Sects. 1 and 2, we compare the time and domain averages of the CRH from the different radiative transfer calculations.

#### 3.1 Average profiles of CRH

Fig. 4 shows domain and time averages of longwave, shortwave, and net CRH in the four regions of the cyclone for the ICON and the reference offline radiative transfer calculation. Despite the differences in the radiative transfer setup between ICON and LibRadtran, the CRH profiles agree very well (cf. dashed and solid lines in Fig. 4). In all four regions, longwave radiation leads to cooling at the cloud tops and weak warming at the cloud bases (Fig. 4, blue), resulting in radiative destabilization of the cloud. In contrast, shortwave radiation warms the cloud tops and slightly cools the lower parts of the clouds, resulting in radiative stabilization of the cloud (Fig. 4, red). The longwave CRH is about twice as large as the shortwave CRH. Thus, the net CRH profiles are dominated by the stronger longwave CRH (Fig. 4, black).



**Figure 4.** Domain and time averages of shortwave, longwave, and net CRH. Profiles are shown for the (a) shallow cumulus, (b) WCB ascent, (c) WCB cyclonic outflow, and (d) WCB anticyclonic outflow regions. The ICON CRH is shown in dashed lines, and the CRH derived from the offline reference 1D radiative transfer calculation is shown in solid lines. Note the different x- and y-axes between panels (a) and (b-d).

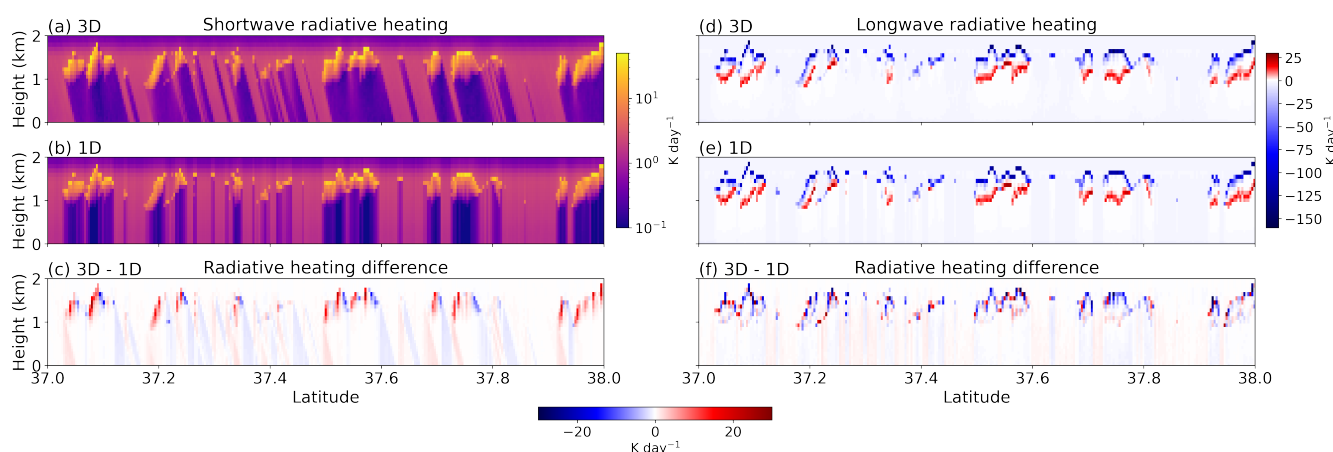
The CRH in the shallow cumulus region is almost twice as strong as in the WCB regions (Fig. 4 a). Grise et al. (2019) and Voigt et al. (2023) showed that the strong cloud-radiative cooling of the boundary layer has a weakening effect on the intensity of extratropical cyclones. The CRH in WCB regions is strongly influenced by the radiative heating of the ice clouds (Fig. 4 b, c and d). Although the CRH in the WCB regions is weaker than the CRH in shallow cumulus region, previous work showed it to be important for near-tropopause dynamics of the cyclone (Li et al., 2015; Keshtgar et al., 2023; Voigt et al., 2023). Thus, it is important to quantify CRH uncertainties in all four regions of the extratropical cyclone.



### 245 3.2 3D cloud radiative effects

We quantify 3D cloud radiative effects as the difference in CRH calculated with the MYSTIC solver (3D, Sect. 2.3) and the MYSTIC solver run in the independent column approximation mode (1D). The latter is a 1D radiation scheme and neglects horizontal radiative transfer between model columns. To visualize the 3D cloud radiative effects, Fig. 5 shows cross-sections of 3D and 1D all-sky radiative heating and their differences for shallow cumulus clouds southwest of the cyclone center at hour 14. At this location and time, the solar zenith angle is approximately  $65^\circ$ .

In the shortwave spectrum, neglecting horizontal radiative transfer leads to incorrect positions of illuminated and shadowed areas in the atmosphere and on the surface (Fig. 5 a, b and c). Črnivec and Mayer (2019) showed that this can have a substantial impact on surface radiative fluxes and hence the surface energy balance. Our focus here, however, is on the 3D radiative effects of clouds within the atmosphere. In Fig. 5 a, the southern sides of the clouds facing the sun receive more radiative energy in the 3D calculation than the 1D calculation (red colors around 1.5 km altitude in Fig. 5 c). In some cases, the northern sides of the clouds are shadowed and receive less energy (blue colors around 1.5 km altitude in Fig. 5 c). The shortwave cloud side leakage is small in this example due to the low position of the sun. In the longwave spectrum, radiative cooling from the cloud tops and modest warming from the bottom are clearly visible in both the 3D and 1D calculations (Fig. 5 d, e). However, the radiative cooling is stronger in the 3D calculation (blue colors around 1.5 km altitude in Fig. 5 f) due to the emission of longwave radiation from the cloud sides, which is neglected in the 1D calculation.



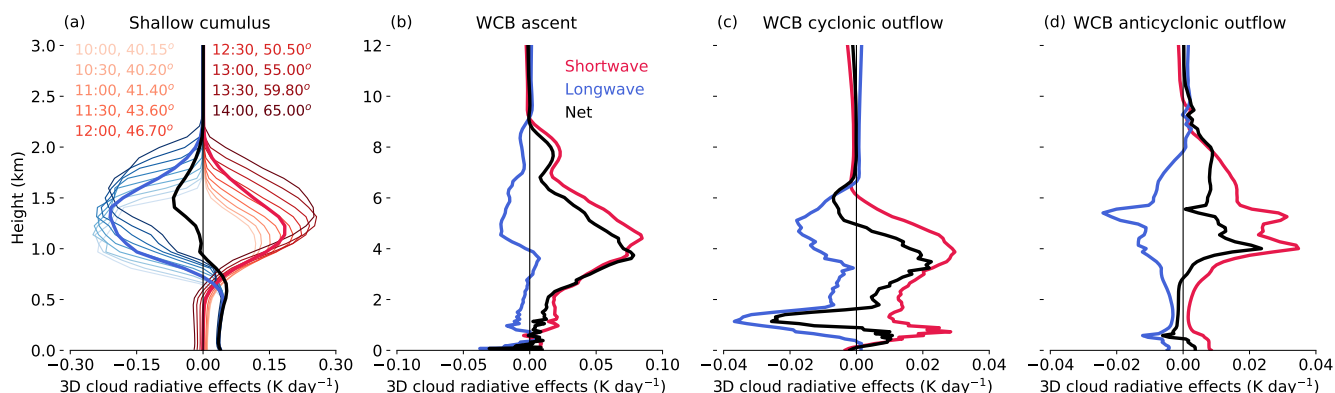
**Figure 5.** Cross-sections of shortwave (a, b, c) and longwave (d, e, f) all-sky radiative heating for shallow cumulus clouds southwest of the cyclone center. The upper row shows 3D calculations, the middle row shows 1D calculations, and the lower row shows the differences between the 3D and 1D calculations. The cross sections are at hour 14 and  $-11.5^\circ$  longitude and between  $37^\circ$  and  $38^\circ$  north.

Fig. 6 shows the average profiles of 3D cloud radiative effects for the four regions of the cyclone. In all regions, shortwave and longwave CRH are stronger in the 3D radiation calculation due to shortwave cloud-side illumination and longwave cloud-side cooling, respectively. There is a direct relationship between cloud-side illumination and solar zenith angle (Fig. 6 a, thin lines). In the 3D calculation, clouds receive more radiative energy from their sides at higher solar zenith angles, which increases



265 shortwave cloud radiative warming. The magnitude of the cloud-side cooling depends on the magnitude of the longwave CRH, such that the stronger the longwave CRH, the stronger the longwave cooling from the cloud sides.

The 3D cloud radiative effects are much weaker for WCB regions which are dominated by stratiform clouds, than the shallow cumulus region, which contains a lot of broken clouds (Fig. 6). This is expected due to the weak horizontal gradients of cloud optical properties and the small aspect ratio of the stratiform clouds in the WCB regions. For shallow cumulus  
 270 clouds, the net 3D cloud radiative effect is dominated by cloud-side longwave cooling (Fig. 6 a). This could lead to a stronger radiative destabilization of clouds during both day and night. In contrast, for stratiform clouds in the WCB, shortwave cloud-side illumination dominates the daytime 3D cloud radiative effects, except for shallow liquid clouds in the boundary layer of the WCB cyclonic outflow region, where longwave cloud-side cooling becomes dominant (Fig. 6 b, c and d). The stronger shortwave cloud-side illumination in the WCB regions is most likely due to the higher solar zenith angle at higher latitudes  
 275 compared to the shallow cumulus region at lower latitudes. Although the incoming shortwave radiative fluxes are weaker at higher latitudes, shortwave cloud-side warming becomes stronger at higher zenith angles.



**Figure 6.** Average profiles of shortwave, longwave, and net 3D cloud radiative effects for the four cyclone regions. 3D cloud radiative effects are calculated as the difference in domain mean CRH between 3D (MYSTIC) and 1D (MYSTIC-ICA) radiative transfer calculations. The thin lines in panel (a) show the 3D radiative effects for nine snapshots between hours 10 to 14, with the legend indicating the time and domain mean solar zenith angles for all snapshots. Note the different x-axes in panels.

### 3.3 Impact of ice-optical parametrization

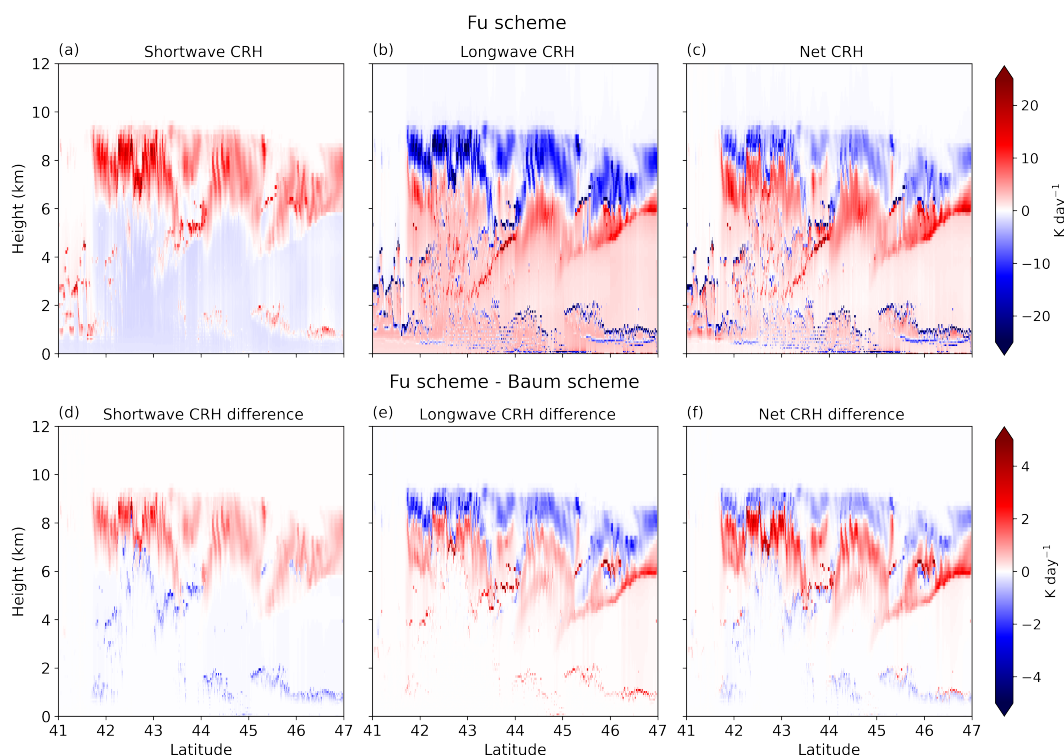
We now quantify the impact of ice-optical parameterization by comparing the CRH calculated with the ice-optical parameterization of Fu and the more complex ice-optical parameterization of Baum. The ice scheme of Baum includes three ice habits,  
 280 which allows us to also assess the effect of different ice habits. We only consider clouds in the three WCB domains, where ice crystals are abundant.

The top row in Fig. 7 shows cross-sections of shortwave, longwave, and net CRH in the WCB ascent region for the ice scheme of Fu. The bottom row in Fig. 7 shows the CRH differences between the ice schemes of Fu and Baum with the general habit mixture. Fu results in stronger longwave and shortwave CRH than Baum (Fig. 7 d, e), with CRH differences reaching up to



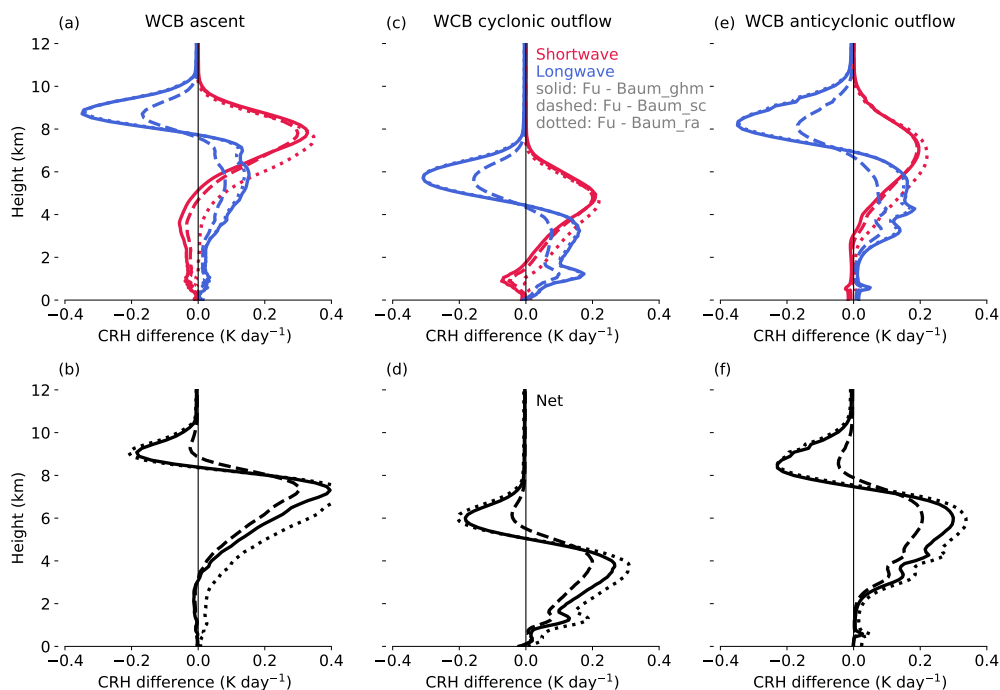
285 20% of the absolute values. As a result, Fu leads to stronger radiative destabilization in the longwave but stronger stabilization in the shortwave.

Fig. 8 shows CRH differences between Fu and Baum. In all three WCB regions, longwave and shortwave CRH are amplified in Fu compared to Baum. The magnitude of the CRH differences depends on which ice crystal habit is chosen in the Baum scheme, with different impacts in the longwave and shortwave. In the longwave, the differences between Fu and Baum are the same for the general mixture and rough-aggregated habits, but smaller by a factor of two when the solid column habit is used in Baum. In the shortwave, the impact of the ice habit is much smaller. The stronger impact of the ice habit on the longwave compared to the shortwave CRH is an interesting aspect of our results.



**Figure 7.** Cross-sections of (a) shortwave, (b) longwave, and (c) net CRH calculated with the ice-optical parameterization of Fu in the WCB ascent region. Panels (d-f) show CRH differences between the ice schemes of Fu and Baum with the general habit mixture (Fu - Baum). Cross-sections are shown at hour 12 and at 3° longitude.

Compared to Baum, Fu leads to stronger shortwave as well as longwave cloud-radiative warming in the middle and lower parts of the clouds, and therefore to net cloud warming (Fig. 8, lower row). Although Fu leads to some increase in cloud-radiative cooling of the upper cloud parts, this suggests that the radiative destabilization of clouds within the WCB regions is smaller in Fu than Baum.



**Figure 8.** Impact of ice-optical parameterization on time- and domain-averaged CRH in the three WCB regions. The CRH differences are calculated between the radiative transfer calculations with the ice schemes of Fu and Baum (Fu - Baum). The differences between Fu and Baum with general habit mixture (ghm), solid column (sc), and rough-aggregated (ra) habits are shown with solid lines, dashed lines, and dotted lines, as is indicated in the legend.

### 3.4 Cloud horizontal heterogeneity and vertical overlap

We continue by quantifying the impact of cloud horizontal heterogeneity and vertical overlap. Both factors are not resolved at 2.5 km and therefore need to be parameterized. We calculate the CRH differences between the radiation calculation using the two types of homogeneous NWP clouds and the LEM clouds (Sect. 2.3, Fig. 3). For homogeneous grid-box clouds, an overlap assumption is not needed. For the homogeneous clouds with fractional cloud cover, we use a maximum-random overlap assumption.

The solid lines in the first row of Fig. 9 show CRH differences between homogeneous grid-box clouds and the LEM clouds, decomposed into longwave and shortwave components. In all regions, the magnitude of the longwave and shortwave CRH profiles are overestimated when homogeneous grid-box clouds are used, with the longwave CRH being more strongly affected than the shortwave CRH and the largest differences occurring for shallow cumulus clouds.

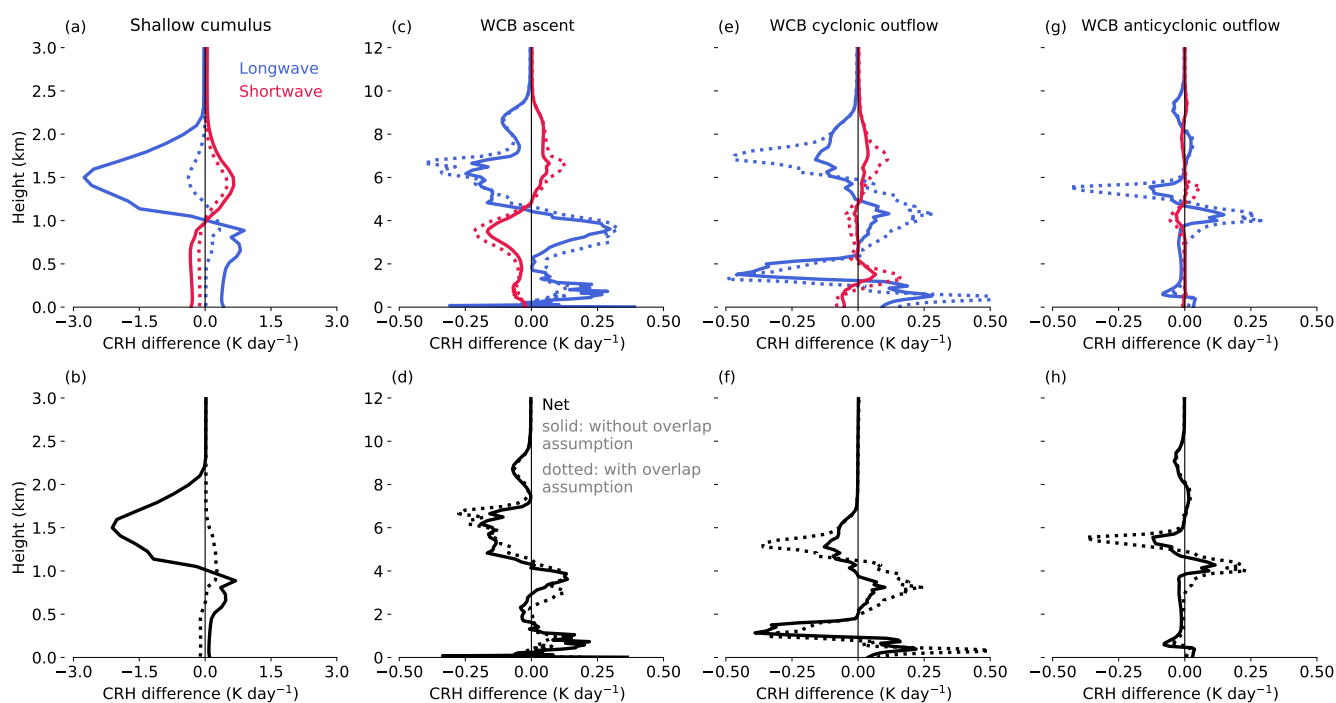
The second row of Fig. 9 compares the impact of the two types of homogeneous clouds. When the overlap scheme is used for the homogeneous clouds with fractional cloud cover, the differences in CRH between the NWP and LEM clouds are strongly reduced in the shallow cumulus region, from  $-2 \text{ K day}^{-1}$  to  $0.5 \text{ K day}^{-1}$ . This illustrates a clear positive impact of taking





310 into account cloud fraction and partial overlap for these clouds, and that treating these clouds as grid-box clouds at 2.5 km resolution is problematic.

In the three WCB regions, cloud vertical overlap assumption has a smaller impact on CRH because the cloud fraction is higher. However, in contrast to the shallow cumulus region, taking into account cloud overlap assumption has a detrimental effect and in fact increases the CRH differences for shortwave, longwave and net (Fig. 9 c to h). This supports the idea that  
 315 clouds in the WCB regions can be treated as grid-box clouds at 2.5 km resolution for radiative purposes because there is no obvious benefit by taking into account cloud fraction and vertical overlap assumption.



**Figure 9.** Impact of cloud horizontal heterogeneity and vertical overlap on CRH that are not resolved at 2.5 km horizontal resolution. The CRH differences are calculated between the LEM reference calculation and the radiative transfer calculations for the homogeneous NWP clouds (NWP - LEM), with the solid lines for the homogeneous grid-box clouds (without overlap assumption) and the dotted lines for the homogeneous clouds with fractional cloud cover (with overlap assumption). Note the different x and y-axes for panels (a) and (b).

#### 4 Relative importance of uncertainties

In the previous section, we have assessed the individual impacts of the four factors on CRH: 3D cloud radiative effects, ice-optical parametrization, cloud horizontal heterogeneity and cloud vertical overlap. In this section, we characterize the  
 320 relative importance of these four factors, with the hope that this can guide efforts to improve CRH in models. As part of this characterization, we distinguish between mean uncertainties on spatial scales of around 500 km, which is approximately the



size of the LEM domain, and local uncertainties, which arise on the scale of the horizontal grid resolutions from 300 m to 2.5 km.

For the mean uncertainties, we compute the *absolute mean* difference,

$$325 \quad \delta CRH_{mean} = |\overline{CRH}_a - \overline{CRH}_b|, \quad (2)$$

where the bars represent the average over time and domain and the subscripts  $a$  and  $b$  indicate different radiative transfer calculations. This means that the mean uncertainties are obtained from the absolute values of the figures shown in Sect. 3.

For the local uncertainties, we compute the *mean absolute* difference by first computing the absolute CRH differences at all grid points and then averaging over time and domain,

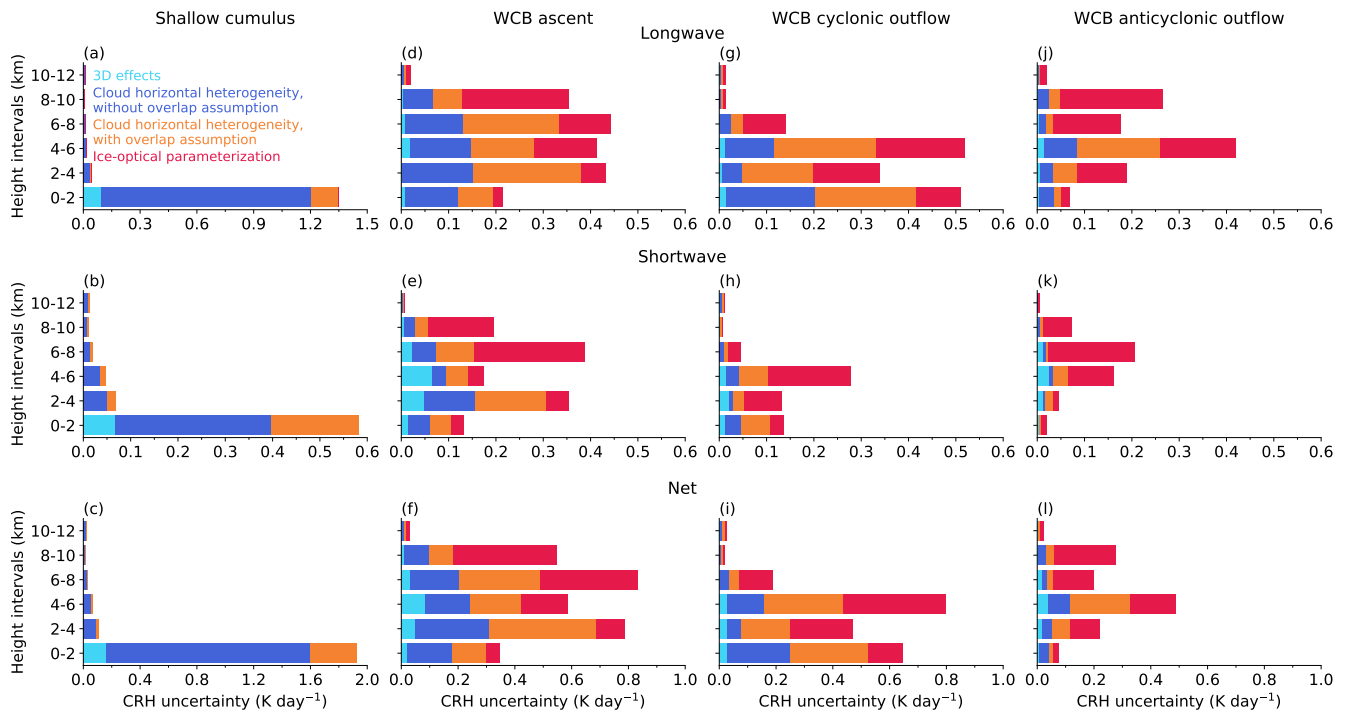
$$330 \quad \delta CRH_{local} = \frac{1}{n \cdot t} \sum_{j=1}^t \sum_{i=1}^n |CRH_a(i, j) - CRH_b(i, j)|. \quad (3)$$

$n$  is the number of horizontal grid points  $i$  at each vertical layer, and  $t$  is the number of time steps  $j$ . We do not compute the local uncertainty due to 3D cloud radiative effects. The latter would require to first quantify the Monte Carlo noise of the MYSTIC solver, which if substantial can on its own create strong local uncertainties. To derive the local uncertainties due to horizontal cloud heterogeneity and vertical overlap, we coarse-grain the LEM CRH to the same horizontal resolution as the  
335 NWP CRH.

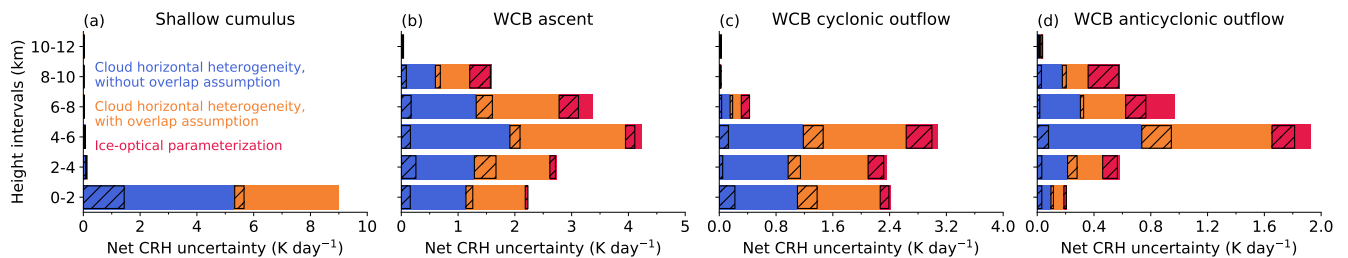
Fig. 10 shows mean uncertainties by means of stacked bar plots in altitude intervals of 2 km calculated based on Eq.2. For shallow cumulus clouds in the boundary layer between 0-2 km, the largest source of mean uncertainty is cloud horizontal heterogeneity (panels a, b and c; blue bars). Allowing for partial overlap for homogeneous clouds with cloud fraction significantly reduces the mean uncertainty (cf. orange and blue bars in panels a, b and c). The mean uncertainty due to 3D cloud radiative  
340 effects is about half of the uncertainty due to cloud horizontal heterogeneity when vertical overlap is considered.

For clouds in the WCB ascent region, cloud horizontal heterogeneity has a significant impact on the mean uncertainty at all levels, and dominates the CRH uncertainty in the boundary layer (panels d, e and f; blue bars). Apart from the boundary layer and in contrast to shallow cumulus clouds, considering vertical overlap assumption increases the mean uncertainty (cf., orange and blue bars in panels d, e and f). Ice-optical parameterization dominates the mean uncertainty of net CRH at middle and  
345 upper levels above 6 km (panel f). 3D cloud radiative effects contribute relatively little, except at altitudes between 2-6 km in the shortwave spectrum.

The relative importance of the four factors for the mean uncertainties in the WCB cyclonic outflow region is similar to that in the WCB ascent region (panels g, h and i). The ice-optical parameterization contributes most to the mean uncertainty in the middle and upper levels, while cloud horizontal heterogeneity dominates the mean uncertainty in the boundary layer. The  
350 vertical overlap assumption increases the mean uncertainty, and the impact of the 3D cloud radiative effects is small. The mean uncertainty in the WCB anticyclonic outflow region is dominated by the ice-optical parameterization, and the impact of cloud horizontal heterogeneity and vertical overlap is much weaker compared to the other regions of the cyclone, except between 4-6 km (panels j, k and l).



**Figure 10.** Mean uncertainties in CRH diagnosed as the absolute difference of domain and time mean CRH between different radiative transfer calculations. Uncertainties are decomposed into shortwave, longwave, and net. Uncertainties are computed as mass-weighted averages over 2 km altitude intervals. For the uncertainty due to the ice-optical parameterization, the difference between the ice schemes of Fu and Baum with the general habit mixture is used. The contribution of each factor is given by the horizontal length of its colored bar. Note the different x-axes in the panels.



**Figure 11.** Same as Fig.10 but for local uncertainties of net CRH. For comparison, the mean values of net CRH uncertainty from Fig.10 are superimposed as hatched bars. Note the different x-axes in the panels.

Now we turn to local uncertainties in CRH. These are illustrated by means of Fig.11 for net CRH. In all four regions of the cyclone, cloud horizontal heterogeneity dominates local uncertainties. As for the mean uncertainties, taking into account the vertical overlap assumption reduces the local uncertainties for shallow cumulus clouds, but increases them for stratiform



clouds in the WCB regions. For comparison, Fig.11 includes the mean uncertainties from Fig. 10 as hatched bars. This shows that cloud horizontal heterogeneity and vertical overlap have a much stronger impact on CRH locally than on the domain mean. In contrast, the ice-optical parameterization has similar impacts on local and mean uncertainties. This suggests that uncertainty due to the ice-optical parameterization is less important for km-scale circulations and more strongly imprints on circulations on larger-scales of hundreds of kilometers. On the other hand, uncertainty from cloud horizontal heterogeneity and vertical overlap should also affect circulations on smaller as well as larger scales.

## 5 Discussion and conclusion

Recent studies have demonstrated that cloud-radiative heating and cooling of the atmosphere (hereafter referred to as CRH) is important for the dynamics and predictability of extratropical cyclones. However, CRH is uncertain in numerical weather prediction (NWP) and climate models, and these uncertainties may have systematic effects on model predictions of extratropical cyclones. This motivates us to here study CRH and its uncertainty within extratropical cyclones. To this end, we simulate an idealized extratropical cyclone with the ICON atmosphere model at a horizontal resolution of 2.5 km and focus on CRH within four regions of the cyclone with distinct cloud patterns: shallow cumulus clouds southwest of the cyclone center, deep clouds in the ascending region of the warm conveyor belt (WCB), stratiform low- and mid-level clouds in the WCB cyclonic outflow, and high stratiform ice clouds in the WCB anticyclonic outflow. For these four regions, we perform large eddy model (LEM) simulations at a horizontal resolution of 300 m, which we use to drive offline radiative transfer calculations. This provides a framework to for the first time assess and compare uncertainty in CRH due to four factors within an extratropical cyclone: 3D cloud radiative effects, ice-optical parameterization, cloud horizontal heterogeneity, and cloud vertical overlap.

We find that 3D cloud radiative effects are overall small and have a noticeable impact only for shallow cumulus clouds. Horizontal radiative transfer within the cloud and from the cloud sides is small in the three WCB regions, where stratiform clouds prevail. These results are consistent with previous findings that 3D cloud radiative effects are small for stratiform clouds (Črnivec and Mayer, 2021), but important for the dynamics of shallow cumulus clouds (Jakub and Mayer, 2016). Thus, while 3D cloud radiative effects have been shown to affect the organization of subtropical low-level clouds, our study indicates that they do not strongly affect extratropical cyclones.

The uncertainty in CRH due to different ice-optical parameterizations is substantial in the WCB of the cyclone. We show that using the more complex ice-optical parameterization of Baum leads to weaker shortwave and longwave CRH. Our analysis shows that the ice-optical parameterization is the largest source of uncertainty on spatial scales of 500 km in the upper troposphere where ice clouds prevail. The uncertainty due to ice-optical parameterization is also rather uniform in space, which suggests that it can affect the cyclone by modulating the large-scale radiative heating. Our findings are consistent with the results of Zhao et al. (2018), who showed that the Baum ice scheme has a weaker effect on shortwave cloud radiative effects compared to the ice scheme of Fu. Zhao et al. (2018) showed that the Baum scheme decreases static stability and increases vertical motion in the midlatitudes. An important potential implication of this result that should be tested in future work is to



what extent this might enhance latent heating in the WCB ascent region and alter the dynamics of the cyclone (Keshtgar et al.,  
390 2023).

Cloud horizontal heterogeneity and cloud vertical overlap are not resolved at 2.5 km resolution, which leads to cloud subgrid  
variability that can affect CRH. We show that this variability contributes significantly to CRH uncertainty in all regions of the  
cyclone. This is consistent with previous findings that when cloud subgrid variability is neglected, both longwave and shortwave  
CRH are overestimated (e.g., Črnivec and Mayer, 2019). The impact is strong for shallow cumulus clouds, while it is much  
395 smaller for the stratiform clouds in the WCB. Including the vertical overlap assumption significantly improves the simulation  
of CRH for shallow cumulus clouds, but in fact slightly degrades CRH for clouds in the WCB since the maximum-random  
overlap assumption misrepresents the vertical arrangement of cloud layers in sheared flows and a more complex form of the  
overlap assumption would be needed (e.g., Giuseppe and Tompkins, 2015). An interesting corollary of our work is that for  
extratropical cyclones, treating clouds as grid-box quantities in km-scale models with resolutions of around 2 km appears a  
400 reasonable choice, because schemes for cloud cover and vertical overlap themselves can act as sources of uncertainty that are  
rarely quantified. This supports the approach taken by Hohenegger et al. (2023) for the ICON-Sapphire km-scale model, who  
decided to remove the cloud cover scheme.

Our study indicates that for extratropical cyclones, improving the representation of ice-optical properties is particularly  
important for reducing CRH uncertainty. This is because i) ice-optical properties represent an important source of uncertainty,  
405 ii) 3D radiative effects are small, and iii) cloud subgrid variability can be taken into account in existing state-of-the-art radiation  
schemes such as TripleClouds (Črnivec and Mayer, 2021). Although advances in radiation solvers can account for both 3D  
radiative effects and cloud subgrid variability, the lack of knowledge about the shape and surface roughness of ice particles  
continues to introduce large variability in the CRH and remains a critical challenge to better represent ice-optical properties  
in models. Future work in the direction of ice-optical properties should also address the problem that in many current models  
410 including ICON, the ice crystal effective radii are not treated consistently between the microphysics and radiation schemes  
(Sullivan and Voigt, 2021), and longwave cloud scattering is neglected (Fan et al., 2022). In view of our results, we believe  
that future studies should focus on the impact of uncertainties in the ice-optical properties on the dynamics and predictability  
of extratropical cyclones.

*Code and data availability.* The code repository containing scripts for ICON simulations, offline radiative transfer calculations, and the  
415 analysis is available in the Gitlab repository (<https://gitlab.phaidra.org/climate/keshtgar-et-al-2023-crh-uncertainty>) hosted by the University  
of Vienna. The simulation raw output will be archived on the High-Performance Storage System at the German Climate Computing Center  
(DKRZ). The post-processed data used in the analysis along with a copy of the Git repository will be published at the LMU open data server.



*Author contributions.* The study was designed by AV. The ICON simulations and offline radiative transfer calculations were carried out by BK. The data analysis was done by BK with inputs from AV, BM, and CH. BK led the writing process of the manuscript, with editing from  
420 AV and input from all authors.

*Competing interests.* At least one of the (co-)authors is a member of the editorial board of Atmospheric Chemistry and Physics, and the authors have also no other competing interests to declare.

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