

1 **Role of atmospheric aerosols in severe winter fog over Indo Gangetic Plains of India: a**
2 **case study**

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10

11 **Abstract**

12 Winter fog and severe aerosol loading in the boundary layer over North India, particularly in the Indo-
13 Gangetic Plain (IGP), disrupt daily life of millions of people in the region. To understand better the
14 role of aerosol-radiation feedback on the occurrence, spatial extent, and persistence of winter fog; and
15 the associated aqueous chemistry in fog in the IGP, several model simulations have been performed
16 using the Weather Research and Forecasting model coupled with chemistry (WRF-Chem). While
17 WRF-Chem was able to represent the fog formation for the December 23-24, 2017 fog event over the
18 central IGP in comparison to station and satellite observations, the model underestimated PM_{2.5}
19 concentrations compared to the Central Pollution Control Board of India monitoring network. While
20 evaluating aerosol composition for fog events in IGP, we found that the WRF-Chem aerosol
21 composition was quite different from measurements obtained during the Winter Fog Experiment in
22 Delhi, with secondary aerosols, particularly chloride aerosol fraction being strongly underpredicted
23 (~66.6%). Missing emission sources (e.g., industry and residential burning of cow dung and trash) and
24 aerosol and chemistry processes need to be investigated to improve model-observation agreement. By
25 investigating a fog event on December 23-24, 2017 over central IGP, we found that the aerosol-
26 radiation feedback weakens turbulence, lowers the boundary layer height, and increases PM_{2.5}
27 concentrations and RH within the boundary layer. Factors affecting the feedback include loss of
28 aerosols through deposition of cloud droplets and internal mixing of absorbing and scattering aerosols.
29 Aqueous-phase chemistry increases the PM_{2.5} concentrations, which subsequently affects the aerosol-
30 radiation feedback by both increased mass concentrations and aerosol sizes. With aerosol-radiation
31 interaction and aqueous phase chemistry, fog formation began 1-2 hours earlier and caused a longer
32 fog duration than when these processes were not included in the WRF-Chem simulation. The increase
33 in RH in both the experiments is found to be important for fog formation as it promoted the growth of
34 aerosol size through water uptake, increasing the fog water content over IGP. The results from this

35 study suggest that the aerosol-radiation feedback and secondary aerosol formation play an important
36 role in the air quality and the intensity and lifetime of fog over IGP, yet other feedbacks, such as
37 aerosol-cloud interactions, need to be quantified.

38

39 **1 Introduction**

40 The Indo-Gangetic Plain (IGP; 21°35'-32°28'N latitude. and 73°50'-89°49'E longitude) in the
41 northern part of the Indian subcontinent is one of the most densely populated and heavily polluted
42 regions in South Asia. The rapid population and economic growth in the IGP region over the last
43 decade have increased air pollution over this region. This is evident from the increasing trend in AOD
44 and NO₂ column concentration over India reported in recent studies (Dey and Di Girolamo, 2011;
45 Ghude et al., 2013; Krishna Moorthy et al., 2013), which has slowed and reversed only recently
46 (Sarkar et al., 2019). The high concentration of aerosols along the IGP and their adverse effects on
47 human health and the environment are increasing (Ghude et al., 2016). Consequently, more than 500
48 million people living in the IGP breathe air that exceeds the National Ambient Air Quality Standards
49 (NAAQS), which has reduced the life expectancy of the people (Debnath et al., 2022; Lelieveld et al.,
50 2015). Lelieveld et al., (2015) estimated a very high number of premature deaths (0.716 million per
51 year) linked to aerosols (PM_{2.5}), thus making Southeast Asia one of the largest regions affected by
52 premature mortality globally.

53 One of the major environmental concerns in the IGP is the urban air quality during winter,
54 especially over the mega-cities, e.g., Delhi, located in the north-western part of IGP (Ghude et al.,
55 2020; Jena et al., 2021; Sengupta et al., 2022). Several urban air pollution hotspots along the IGP
56 extend from northwest to east with monthly average PM_{2.5} greater than 200 μgm^{-3} (NAAQS=60 μgm^{-3} ,
57 24 hr average) in the winter season (Bharali et al., 2019; Krishna et al., 2019). IGP is dominated
58 mainly by fine mode particulates, especially over central to eastern IGP, during post-monsoon and
59 winter (Kumar et al., 2018). Biomass burning (agricultural waste burning, domestic heating, etc.) is an
60 important contributor to the observed high PM_{2.5} loading over IGP during these seasons (Kulkarni et
61 al., 2020; Pant et al., 2015; Pawar and Sinha, 2022; Sharma et al., 2010; Yadav et al., 2020). Delhi is
62 affected substantially by the emissions from agricultural waste burning in the north-western states of
63 Punjab and Haryana during the post-monsoon (October-November) season (Badarinath et al., 2009;
64 Jethva et al., 2018; Kumar et al., 2021). Studies showed that PM_{2.5} increased from ~50 $\mu\text{g m}^{-3}$ to as
65 high as 300 $\mu\text{g m}^{-3}$ (Ojha et al., 2020), and AOD reached 0.98 with the presence of absorbing aerosols
66 (Singh et al., 2018) during the peak biomass burning in post-monsoon.

67 IGP experiences fog (both radiation and advection fog) every winter after the passage of the
68 synoptic wind system called the “Western Disturbances”. The majority of fog events in the IGP during
69 December-January are radiation fog (Deshpande et al., 2023; Ghude et al., 2023), formed due to
70 radiative cooling of the surface. The number of low visibility days due to haze/fog formation has been

71 increasing significantly (Ghude et al., 2017; Jenamani, 2007; Singh and Dey, 2012), impacting socio-
72 economic activities, e.g., aviation (Kulkarni et al., 2019). The increase in the intensity and regional
73 extent of fog over IGP is consistent with the increasing trend in aerosol concentration due to
74 increasing anthropogenic emissions (Sarkar et al., 2006; Syed et al., 2012).

75 Several factors control the formation and persistence of fog in the IGP, e.g., stable boundary
76 layer, low temperature, availability of moisture (supplied by the Western Disturbances and irrigation
77 activities), and the aerosol number and composition (Acharja et al., 2022; Dhangar et al., 2021). It has
78 also been suggested that the atmospheric rivers (moisture incursion from Arabian Sea) act as a source
79 of water vapor over IGP, which fuels the intensification of fog and haze (Verma et al., 2022) during
80 winter. The high aerosol concentration in the boundary layer influences fog formation (Gautam et al.,
81 2007; Safai et al., 2019) over the IGP by providing the needed cloud condensation nuclei (CCN) for
82 activation into fog droplets. In addition, aerosols induce surface cooling by reducing solar radiation at
83 the surface while warming the lower troposphere by absorption (Ding et al., 2016; Yu et al., 2002). A
84 reduction in surface-reaching solar radiation by ~19% has been reported during winter over Kanpur in
85 the IGP (Dey and Tripathi, 2007). The reduced solar flux affects the boundary layer stability and
86 depth by suppressing the thermals and thus further increasing the surface aerosol concentration via
87 aerosol-radiation feedback, which is very strong over the IGP (Bharali et al., 2019). Kumar et al.,
88 (2020) have shown that aerosol-radiation feedback significantly improves the accuracy of PM_{2.5} and
89 temperature forecasts in Delhi. Srivastava et al., (2018) reported that the direct aerosol forcing over
90 polluted regions is very large with values up to $-80.0 \pm 7.2 \text{ W m}^{-2}$ over the IGP in the winter season.

91 Aerosol-radiation interaction determines that the aerosol distribution is critical for the evolution
92 of fog (Bodaballa et al., 2022; Steeneveld et al., 2015), while microphysics is important for fog
93 formation and dispersal (Boutle et al., 2018; Maalick et al., 2016). Although the relationship between
94 the aerosol chemical composition and aerosol activation to CCN has not been fully understood yet,
95 studies have found that the chemical composition and mixing state of aerosols affect the
96 hygroscopicity (κ) of aerosols (Bodaballa et al., 2022; Ma et al., 2013; Moore et al., 2012; Zhang et
97 al., 2014a). Fog processes involve a complex interplay between local meteorology, radiation,
98 microphysics, and aerosol chemistry, making it difficult to understand the fog lifecycle (Acharja et al.,
99 2022; Maalick et al., 2016; Zhang et al., 2014b). There is considerable heterogeneity in the spatial and
100 temporal aerosol properties over IGP and the poor estimates of their mixing state. Therefore,
101 prediction of fog by weather models is still challenging with biases in fog's onset and dispersal
102 timings.

103 Previous studies have focussed on the impacts of meteorological conditions, topography, or
104 anthropogenic emissions on the poor air quality and intensification of fog during winter over IGP (e.g.
105 Hakkim et al., 2019). However, studies on the effect of feedback induced by the aerosols on the
106 meteorological conditions and thus on aerosol concentration are very limited over this region, except
107 for a few above-mentioned studies which discuss how the aerosol-radiation feedback favors haze and

108 fog during winter. Moreover, fog can provide a medium for aqueous-phase reactions. While several
109 earlier studies have reported an increase in secondary aerosols during fog over IGP, a sensitivity study
110 examining the impact of fog on aqueous phase chemistry has not yet been done over IGP.

111 In the present work, we aim to find the suitable chemistry/physics as well as the meteorology
112 initial/boundary conditions that lead to improved simulations of fog events in the Weather Research
113 and Forecasting model coupled with chemistry (WRF-Chem; (Fast et al., 2006; Grell et al., 2005;
114 Powers et al., 2017). We also explore the role of aerosol-radiation feedback on fog properties as the
115 high aerosol loadings in northern India can impact the heating rates, temperature inversions, and
116 boundary layer height. The role of aqueous chemistry on fog properties and vice-versa is also
117 investigated.

118

119 **2 Methodology**

120 Fog formed due to radiative cooling, based on the onset time of fog and low wind speeds (Deshpande
121 et al., 2023 and references therein), at the surface on both 23rd and 24th December 2017 over a
122 widespread region of the IGP (Fig. 1a, b). The fog region is located over an area with high PM_{2.5}
123 anthropogenic emissions (Fig. 1c). The IGP is a large region with varying meteorology and aerosol
124 characteristics, therefore, it is divided into three areas, northwest (NWIGP: latitude-longitude range,
125 27°N-32°N,75°E-79°N), central (CIGP: latitude-longitude range, 25°N-28°N,79°E-83°E), and east
126 (EIGP: latitude-longitude range, 24°N-27°N, 83°E-87°E) which are marked by the black rectangles in
127 Fig.1c. Although biomass burning and anthropogenic emissions dominate throughout the IGP during
128 post-monsoon and winter season, the north-westerly wind system results in the gradient distribution of
129 AOD over this region. The downwind regions, CIGP and EIGP are influenced by the long-range
130 transport from the NWIGP, resulting in high AOD with dominant fine particulates over CIGP and
131 EIGP, especially during post-monsoon and winter (Kedia et al., 2014; Kumar et al., 2018). Therefore,
132 representative stations from each region listed in section 2.2 are considered for the sensitivity
133 analyses.

134

135 **2.1 Modeling**

136 The WRF-Chem model version 4.0.3 is used for this study. Earlier studies have successfully
137 used WRF-Chem to predict fog (Pithani et al., 2019) and in the study of aerosol-radiation feedback on
138 air quality (Kumar et al., 2020; Bharali et al., 2019) and fog (Shao et al., 2023). The model domain is
139 centered at Delhi (77.1°E, 28.7°N) with 300 grid points in the east-west, 170 grid points in the south-
140 north direction (Fig. 1c), and 50 vertical eta levels with the model top at 50 hPa. The horizontal grid
141 spacing of the domain is 10 km, while the vertical grid spacing varies from higher resolution (~200 m)
142 in the boundary layer to coarser resolution (~1200 m) near the model top. We conduct three model
143 configurations (Table 1) for December 20-24, 2017 to identify the best configuration for
144 meteorological simulations. The three experiments have been designed with different combinations of

145 meteorological initial/lateral boundary conditions and planetary boundary layer (PBL) physics.
146 Experiment 1 (EXP1) uses the National Centers for Environmental Predictions (NCEP) Final Analysis
147 (GFS-FNL; $1^\circ \times 1^\circ$, 6 hourly) meteorology data for initial and boundary conditions and the YSU
148 (Yonsei University; (Hong et al., 2006) PBL scheme. Experiments 2 and 3 (EXP2, EXP3) use ERA-
149 Interim Project ($1.125^\circ \times 0.703^\circ$, 6 hourly) for meteorology initial and boundary conditions. EXP2
150 uses the YSU PBL scheme while EXP3 uses the ACM2 (Asymmetric Convective Model version 2)
151 PBL scheme. ACM2, is a hybrid of the original nonlocal closure (Pleim and Chang, 1992) and a local
152 closure eddy diffusion scheme (Pleim, 2007a, 2007b). The YSU PBL option was coupled with the
153 Noah LSM while ACM2 was coupled with Pleim-Xiu LSM. While YSU permits investigations of
154 both aerosol-radiation (AR) and aerosol-cloud interactions (ACI), ACI is not possible when using the
155 ACM2 PBL scheme because in WRF the ACM2 PBL scheme does not provide the exchange
156 coefficient for heat, which is required to calculate the maximum supersaturations and therefore the
157 activation fraction for aerosol mass and number for each bin/mode, which is based on the Abdul-
158 Razzak and Ghan, (2002) scheme. However, ACM2 has been shown to perform well for air quality in
159 the polluted regions (Mohan and Gupta, 2018; Gunwani and Mohan, 2017; Xie et al., 2012, Mohan
160 and Bhati, 2011). Mohan and Gupta (2018) tested the YSU and ACM2 schemes during the summer
161 time (1-15 June 2010) and focused on the evaluation of temperature, wind speed, PBL height, ozone,
162 and PM_{10} . Although these studies recommend using the nonlocal ACM2 PBL scheme for air quality
163 prediction for IGP, there are still seasonal, day-night, and regional biases in the PBL schemes.
164 Gunwani and Mohan (2017) showed that ACM2, QNSE (Quasi Normal Scale Elimination), and MYJ
165 (Mellor-Yamada-Janjić) schemes work well in predicting temperature, humidity, and wind speed in
166 different regions of India. Mohan and Bhati (2011) found that using the ACM2 PBL scheme with
167 Pleim Xiu surface physics improved wintertime meteorology estimates in Delhi indicating its potential
168 in fog predictions, whereas Pithani et al., (2019) recommend using the local PBL scheme MYNN2.5
169 (Mellor-Yamada-Nakanishi-Niino level 2.5). Shin and Hong (2011) found that a non-local (e.g.,
170 ACM2, YSU) scheme is favorable in unstable conditions and a local scheme (e.g., MYJ, Boulac) in
171 stable conditions. All these studies suggest the need for careful consideration of the above-mentioned
172 biases while selecting a PBL scheme. Therefore, to ensure that WRF captures all the relevant
173 meteorological parameters including relative humidity reasonably well during fog events in winter, we
174 designed EXP1, EXP2 and EXP3.

175 The advantage of Pleim-Xiu LSM (PX-LSM) is that it allows nudging of soil moisture and
176 temperature to improve the prediction of meteorology near the surface (Pleim and Gilliam, 2009;
177 Pleim and Xiu, 2003; Xiu and Pleim, 2001) which Noah LSM does not include. The PX-LSM
178 includes two-layer soil (0–1 and 1–100 cm) model, canopy moisture, and aerodynamic and stomatal
179 resistance. Ground surface (1 cm) temperature is calculated from the surface energy balance using a
180 force-restore algorithm for heat exchange within the soil. Although the two-layer approach in PX-
181 LSM is less detailed than the multilayer soil models such as the Noah LSM (four soil layers; Chen and

182 Dudhia 2001), it performs well with realistic initialization for soil moisture and through dynamic
183 adjustment in the model simulation where soil moisture is indirectly nudged according to differences
184 in 2-m temperature (T2) and 2-m relative humidity (RH) between the model and observation (Pleim
185 and Xiu, 2003). Soil moisture nudging adjusts the surface evaporation (direct soil surface evaporation,
186 vegetative evapotranspiration, and evaporation from wet canopies) which then affects the partitioning
187 of available surface energy into latent and sensible heat flux and thus reduces errors in T2 and 2-m
188 RH.

189 For EXP2, meteorological initial conditions were refreshed every 24 hours, while EXP3 was a
190 continuous run but soil moisture was nudged to the Era-Interim dataset to improve the prediction of
191 surface fluxes. All other physics and chemistry options are the same for all the experiments except the
192 surface physics option, which changes with the PBL scheme used. The deposition of cloud droplets is
193 an important moisture and aerosol sink during fog events. For all these simulations, the deposition
194 velocity of cloud droplets was reduced to 0.01 m s^{-1} based on Stoke's Law and previous studies
195 (Katata et al., 2015; Tav et al., 2018) because its default value (0.1 m s^{-1}), is large.

196 To examine the radiative effects of aerosols and aqueous phase chemistry additional
197 simulations have been done using the meteorological configuration in EXP3, with aerosol-radiation
198 (wFB) feedback plus aqueous chemistry (wAq.chem), without aerosol-radiation feedback (nFB) but
199 with aqueous chemistry, and without aqueous chemistry (noAq.chem) but with aerosol-radiation
200 feedback. The analysis has been done for the fog events on 23rd and 24th December 2017 as

201

$$202 \quad F_{\text{ARF}} = P(\text{wFB-nFB})$$

$$203 \quad F_{\text{Aq.chem}} = P(\text{wAq.chem-noAq.chem})$$

204 where F_{ARF} is the impact due to aerosol radiation feedback on the meteorological parameters/chemical
205 species (P) and $F_{\text{Aq.chem}}$ is the impact due to the inclusion of aqueous phase chemistry

206

207 Emissions used in the WRF-Chem simulations are from the EDGAR-HTAP v2 (*Emissions*
208 *Database for Global Atmospheric Research- Hemispheric Transport of Air Pollution; $0.1^\circ \times 0.1^\circ$*)
209 inventory for anthropogenic emissions and FINN v2.2 (*Fire INventory from NCAR; 1 km x 1 km*) fire
210 emission inventory (Wiedinmyer et al., 2011). Trash-burning emissions (Chaudhary et al., 2021) are
211 also included in the simulations. The model calculates the biogenic emissions online using MEGAN
212 v2.04 (*Model of Emissions of Gases and Aerosols from Nature*) (Guenther et al., 2006). The initial and
213 lateral boundary conditions for chemical constituents are from the global chemistry transport model
214 CAM-Chem (*Community Atmosphere Model with Chemistry*) (Emmons et al., 2020).

215 The MOZART (Model for Ozone and Related chemical Tracers) chemical mechanism
216 (Emmons et al., 2010) is used for gas-phase chemistry, which includes 85 gas-phase species, 39
217 photolysis, and 157 gas-phase reactions. It has been updated to include an explicit treatment of
218 aromatic compounds, HONO, C_2H_2 , and isoprene oxidation scheme (Knote et al., 2014). The lumped

219 toluene used by Emmons et al., (2010) has been speciated into benzene, toluene, and lumped isomers
220 of xylenes (Knote et al., 2014). For this study, HCl emissions, transport, dry, and wet deposition are
221 represented. However, HCl gas-phase reaction is not included in MOZART.

222 The Model for Simulating Aerosol Interactions and Chemistry (MOSAIC) with four size bins
223 (0.039–0.156, 0.156–0.625, 0.625–2.500, and 2.5–10.0 μm dry diameters) coupled with MOZART
224 gas-phase chemistry is used (Fast et al., 2006; Zaveri et al., 2008). The bin sizes are defined by their
225 lower and upper dry particle diameters, so there is no transfer of particles between bins during water
226 uptake or loss. It is assumed that aerosols in each bin are internally mixed with the same chemical
227 composition while they are externally mixed in different bins.

228 The aerosol composition includes sulfate (SO_4^{2-}), ammonium (NH_4^+), nitrate (NO_3^-), aerosol
229 water, sea salt (Na^+ , Cl^-), methanesulfonate (CH_3SO_3^-), carbonate (CO_3^{2-}), calcium (Ca^+), black carbon
230 (BC), organic mass (OC), and unspecified inorganic species such as silica, inert minerals, and trace
231 metals lumped together as other inorganic mass (OIN). For OC, primary OC and secondary OC are
232 represented separately, where the latter is simulated using the volatility basis set (VBS) approach.
233 Reactive inorganic species such as potassium (K^+) and magnesium (Mg^+) are usually present in much
234 smaller amounts and are equivalent to Na^+ since their sulfate, nitrate, and chloride salts are similar in
235 terms of their solubility in water.

236 MOSAIC treats condensation and evaporation of trace gases to/from particles, nucleation
237 (new particle formation), and coagulation. Aerosol coagulation (Brownian) is based on (Jacobson et
238 al., 1994) and nucleation is based on (Wexler et al., 1994) parameterization of $\text{H}_2\text{SO}_4\text{-H}_2\text{O}$
239 homogeneous nucleation. Sulfate, nitrate, chloride, and ammonium aerosols are mainly formed
240 through oxidation and neutralization/condensation of gas precursors. Gas-phase sulfuric acid (H_2SO_4)
241 is produced by the gas-phase oxidation of SO_2 by OH and nitric acid (HNO_3) formation is via the
242 oxidation of NO_2 by OH. HCl is a primary emission product. The neutralization/condensation of
243 H_2SO_4 , HCl, and HNO_3 with NH_3 produces ammonium such as ammonium sulfate ($(\text{NH}_4)_2\text{SO}_4$),
244 ammonium bisulfate (NH_4HSO_4), ammonium chloride (NH_4Cl) and ammonium nitrate (NH_4NO_3),
245 respectively. The thermodynamic modules in MOSAIC for the dynamic gas-particle partitioning of
246 aerosols MTEM (Multicomponent Taylor Expansion Method) and MESA (Multicomponent
247 Equilibrium Solver for Aerosols) calculate the activity coefficient in aqueous phase aerosols and
248 compute the intraparticle solid-liquid phase equilibrium respectively (Zaveri et al., 2005, 2008). The
249 Adaptive Step Time-split Euler Method (ASTEM) coupled with MESA-MTEM dynamically
250 integrates the mass transfer equations.

251 Aqueous-phase chemistry uses a bulk water approach employing the Fahey and Pandis (2001)
252 mechanism. It calculates sulfate formation, formaldehyde oxidation, and non-reactive uptake of nitric
253 acid, hydrochloric acid, ammonia, and other trace gases (Chapman et al., 2009; Pye et al., 2020).
254 Aqueous-phase sulfate is produced via oxidation of SO_2 by H_2O_2 , O_3 , TMI (Transition metal Ion:
255 Fe(III), Mn(II)) catalyzed O_2 and NO_2 . TMI concentrations are prescribed in the model to $0.01 \mu\text{g m}^{-3}$

256 for Fe(III) and $0.005 \mu\text{g m}^{-3}$ for Mn(II) (Martin and Good, 1991). The Fe(III) values are within the
257 range of water soluble iron in winter time aerosol reported in India (Kumar and Sarin, 2010). Wet
258 removal (scavenging), is represented by the (Neu and Prather, 2012) scheme for trace gases and Easter
259 et al., (2004) for aerosols.

260

261 **2.2 Observations**

262 To evaluate the model output, observations of aerosols and meteorology have been obtained
263 from several satellites as well as ground-based measurement platforms. To examine the aerosol
264 loading and spatial and temporal distribution, daily Level 2 Aerosol Optical Depth (AOD) retrievals
265 from the Moderate Resolution Imaging Spectroradiometer (MODIS) aboard Terra and Aqua satellites
266 are obtained at the spatial resolution of $10 \text{ km} \times 10 \text{ km}$ (at nadir) pixel array. It provides aerosol
267 properties from the Dark Target (DT) algorithm applied over the ocean and dark land (e.g.,
268 vegetation) and Deep Blue (DB) algorithms over the entire land areas, including both dark and bright
269 surfaces. Each MOD04_L2 (Terra) / MYD04_L2 (Aqua) products are available at a 5-minute time
270 interval with an output grid of 135 pixels in width by 203 pixels in length.

271 The Indian National Satellites (INSAT-3D) in the geostationary orbit at inclinations of 82°
272 longitude provide an imager fog product (3DIMG_L2C_FOG) with a spatial resolution of 4 km every
273 30 min (www.mosdac.gov.in). For daytime, the visible channel observation is used to detect fog,
274 whereas thermal infrared is used to reduce false alarms such as medium/high clouds and snow areas.
275 INSAT 3D's 'day microphysics' data component analyzes solar reflectance at three wavelengths: 0.5
276 μm (visible), $1.6 \mu\text{m}$ (shortwave infrared), and $10.8 \mu\text{m}$ (thermal infrared). Night-time fog is derived
277 from TIR-1 ($12.0 \mu\text{m}$ and $10.0 \mu\text{m}$) and MIR ($10.8 \mu\text{m}$ and $3.9 \mu\text{m}$) channel brightness temperature
278 over the Indian region. INSAT-3D provides fog intensity varying from 1 to 4 indicating SHALLOW
279 for visibility $> 600 \text{ m}$; MODERATE, DENSE, and VERY_DENSE, respectively for visibility varying
280 from 0 to 500 m (Banerjee and Padmakumari, 2020). If the visibility is greater than 700 m it indicates
281 no fog while visibility $> 1000 \text{ m}$ represents very clear skies. Validation of INSAT-3D fog products
282 over the IGP shows a 66%-68% probability of detection and a 10% false alarm rate. It also captures
283 the entire life cycle of fog from formation to dissipation. However, detecting fog during multilayer
284 clouds is still challenging with INSAT-3D (Arun et al., 2018; Chaurasia and Gohil, 2015; Chaurasia
285 and Jenamani, 2017).

286 Ground-based monitoring sites provide hourly data of relative humidity, surface temperature,
287 and wind speed measured by the Central Pollution Control Board, CPCB (<http://cpcb.nic.in>). Given
288 the data availability from CPCB stations, nine stations have been considered representing each region
289 of IGP, which include, Amritsar, IGI Airport (Indira Gandhi International Airport, Delhi), IHBAS
290 (Institute of Human Behaviour and Allied Sciences, Delhi), Dwarka (Delhi), RKP (Ramakrishna

291 Puram, Delhi) in the North-West IGP; Kanpur, Lucknow in Central IGP and Patna, Muzaffarpur in
292 East IGP.

293 In addition, measurements of several aerosols, trace gases, and meteorology at Delhi (IGI
294 Airport) from the Winter Fog Experiment (WiFEX) for the period December 10-31, 2017, have also
295 been used to validate the model output. The WiFEX, an initiative of the Ministry of Earth Sciences
296 (MoES), India, is a ground-based measurement campaign at the IGI Airport Delhi to understand fog's
297 physical and chemical features. Additional details of the WiFEX project and related publications can
298 be found in Ghude et al., (2017).

299

300 **3 Meteorology Evaluation**

301 Previous studies simulating fog highlight the importance of high model vertical resolution
302 (Pithani et al., 2019; Van Der Velde et al., 2010) for representing the fog formation and the growth of
303 the fog layer, model initialization (Yadav et al., 2022), initial relative humidity (Bergot and Guedalia,
304 1994; Pithani et al., 2020), and PBL schemes (Chen et al., 2020; Pithani et al., 2019). In the present
305 study, 2-m relative humidity (RH2), 2-m temperature (T2), and 10-m wind speed (WS) from WRF-
306 Chem have been evaluated using ground-based measurements from CPCB monitoring network and
307 WIFEX campaign for nine stations across the IGP. The comparison of WRF-Chem results with
308 observations shows that RH2 and T2 are sensitive to the choice of the meteorological initial and
309 boundary conditions as illustrated by six stations in major cities (Fig. S1). WRF-Chem compares
310 better with the observations for simulations driven by the ERA-Interim reanalysis than with GFS-FNL
311 reanalysis since ERA-Interim provides more realistic RH2 than GFS-FNL (Figs. S2 a-f). For example,
312 RH2 from EXP1 (GFS) varies from 10 to 50%, while RH2 from EXP2 and EXP3 varies from 30 to
313 100%, which is closer to observation, especially for NWIGP and CIGP. For EIGP, RH2 from EXP1
314 (GFS) compares better than ERA-Interim, which overestimates the observed RH2. ERA-Interim and
315 YSU PBL scheme showed damping of RH2 continuously increasing the bias in RH2 with time (not
316 shown), which was corrected in EXP2 by refreshing meteorology every day at 00h UT during the
317 model simulation. In addition, maps of surface RH2 and T2 (Figs. S2 g-j) show that the GFS-FNL
318 dataset has lower relative humidity throughout the domain as compared to ERA-Interim. There are
319 differences in simulated 2-m temperature between these two datasets which are of smaller relative
320 magnitude compared to the RH2.

321 The GFS-FNL driven meteorology EXP1 has a warm bias in NWIGP and CIGP, especially during
322 night-time, while over EIGP, the model prediction agrees well with observations. EXP2 with the
323 ERA-Interim driven meteorology and YSU PBL scheme also shows good agreement between
324 modeled and observed T2 in EIGP. The ERA-Interim driven meteorology with the ACM2 PBL
325 scheme in EXP3 has a cold bias of up to 7°C over EIGP during daytime from 22nd to 24th December.
326 The wind speed evaluation shows that WRF-Chem is over-predicting wind speed. However, it is also
327 possible that some CPCB stations (e.g., Amritsar and RK Puram) have a wind speed low bias due to

328 the low measurement height and obstructions such as tall trees near the monitoring station as shown in
 329 FigS3. WRF-chem in general overestimates wind speed and several earlier studies have reported this
 330 bias in wind speed (e.g., Mohan and Gupta 2018; Pithani et al.,2019). Moreover, WRF-Chem does not
 331 have the capability to represent building meteorology and parameterizes the effects of urban areas on
 332 meteorology through roughness length, which likely leads to overestimation of wind speed. Note that
 333 at other sites (e.g., over IGI-Delhi and Kanpur) the model measurement agreement is better.
 334 The WRF-Chem performance has been statistically assessed against observation using the Taylor
 335 Diagram (Taylor, 2001), which provides a statistical summary of how well the model output agrees
 336 with the observation in terms of the Pearson correlation, their centered root-mean-square error
 337 (RMSE) difference, and the ratios of their variances (Fig. 2). The centered RMS difference is
 338 proportional to the distance to point “OBS” in the x-axis which measures the extent to which the
 339 simulated and observed datasets match. The centered RMS difference (E'), the correlation (R), and the
 340 standard deviations, σ_m^2 (simulated) and σ_o^2 (observed) are calculated as:

$$341 \quad R = \frac{\frac{1}{N} \sum (M_n - \bar{M})(O_n - \bar{O})}{\sigma_m \sigma_o} \quad (1)$$

$$342 \quad E'^2 = \frac{1}{N} \sum [(M_n - \bar{M}) - (O_n - \bar{O})]^2 \quad (2)$$

$$343 \quad \sigma_m^2 = \frac{1}{N} \sum (M_n - \bar{M})^2 \quad (3)$$

$$344 \quad \sigma_o^2 = \frac{1}{N} \sum (O_n - \bar{O})^2 \quad (4)$$

345 where the overall mean of a field is indicated by an overbar.

346 Each point in the two-dimensional space of the Taylor diagram represents the above mentioned three
 347 different statistical metrics simultaneously, as they are related by the follow equation

$$348 \quad E'^2 = \sigma_o^2 + \sigma_m^2 - 2\sigma_o\sigma_m R \quad (5)$$

349 The diagram is constructed based on the similarity of the above equation and the Law of Cosines:

$$350 \quad c^2 = a^2 + b^2 - 2ab\cos\phi \quad (6)$$

351 The percentage bias has also been included to further evaluate the WRF-Chem results. In Fig.
 352 2, better agreement of WRF-Chem results with observations is shown by the marker's proximity to the
 353 “OBS” dashed black line. The WRF-Chem RH has a good correlation for all three experiments with r
 354 > 0.75 at all the locations in IGP for all the experiments. However, the RMSE (shown by red dashed
 355 contours) and the standard deviations are larger for EXP1 compared to EXP2 and EXP3 which lie
 356 closer to the dashed black line indicating that the simulated RH variations are similar to observations.
 357 The mean bias is also large ($>20\%$) for EXP1 (GFS-FNL) for all the stations, marked by red triangles.
 358 For example, simulated RH at Dwarka (4) and Lucknow (7) for EXP2, and IGI Airport (2), IHBAS
 359 (3), Lucknow (7), and Patna (8) for EXP3 show good agreement with observation, with $r > 0.7$,
 360 standard deviation within ± 0.25 and mean bias within 10%. Among these stations, the model performs
 361 better for Dwarka (4) and Lucknow (7) for EXP2, IGI Airport (2), and IHBAS (3) for EXP3 with a
 362 smaller centered RMSE (< 0.75).

363 For all the experiments, WRF-Chem T2 agrees well with observations with a correlation
364 between 0.8 and 0.95. The points are concentrated near the dashed line showing a low RMSE and
365 standard deviation for T2, signifying a good agreement of simulated T2 with observation in terms of
366 temporal variation but the T2 relative bias is large for EXP1 (>20%). The RMSE and relative bias for
367 EXP1 are larger for several of the stations. For example, simulated T2 agrees best with observation at
368 IHBAS (3) for EXP1 and IGI Airport (2) for EXP2, with smaller centered RMSE and standard
369 deviation, and bias <5%. The temporal variability of T2 and RH is predicted well for all the
370 combinations of inputs (Fig. S1), however, the accuracy of simulated T2 and RH is sensitive to the
371 choice of meteorological initial/boundary conditions. WRF-Chem predicted RH and T2 agree better
372 with observations when initialized with ERA-Interim meteorology than with GFS-FNL.

373 The WRF-Chem runs driven by ERA-Interim with YSU (EXP2) and ACM2 PBL (EXP3)
374 schemes predicted the surface meteorology better over the IGP than the WRF-Chem run driven by
375 GFS (EXP1). By examining the modeled cloud water content (averaged over all the grids in the
376 analysis region) in the lowest model level with the INSAT-3D satellite fog intensity for the 23rd and
377 24th December 2017 (Fig. 3), it is apparent that WRF-Chem with the ACM2 PBL scheme compared
378 qualitatively well with observations obtained from INSAT-3D satellite in terms of fog coverage over
379 CIGP, while the WRF-Chem run with the YSU PBL scheme did not produce widespread fog.
380 However, there is also fog over EIGP in WRF-Chem with the ACM2 PBL scheme although it is not
381 observed by the satellite. This is because the model has a cold bias in T2 and a high surface RH over
382 East IGP with ACM2 PBL and Pleim-Xiu surface scheme as discussed earlier, which favors the
383 formation of fog in this region. The time series in Fig. 4 shows that EXP3 is capable of predicting the
384 duration of fog on 23rd and 24th December. There is a data gap from INSAT 3D observations because
385 it is unable to capture fog during daytime in the presence of mid and high-level clouds. Although fog
386 LWC data is available for Delhi from the WiFEx campaign, the WRF-Chem simulation does not
387 produce fog in NWIGP, therefore, the WiFEx observations are not included in this evaluation.

388 In conclusion, EXP3 is the best configuration for predicting fog formation where the ERA-
389 Interim meteorology, the ACM2 PBL and surface schemes, and soil moisture nudging is used in the
390 WRF-Chem simulation. Therefore, the evaluation of predicting AOD, surface aerosol concentrations,
391 and aerosol composition as well as analysis of the impact of aerosols on fog formation uses the EXP3
392 configuration.

393

394 **4 Aerosol Evaluation**

395 Aerosol is an important factor in correct prediction of fog (Maalick et al., 2016; Stolaki et al.,
396 2015) as the number of fog droplets depends on the aerosol size distribution and concentration.
397 Aerosols as CCN can affect the liquid water content in fog and therefore an increase in aerosol
398 concentration can significantly affect fog lifetime (Stolaki et al., 2015; Zhang et al., 2014b). AOD
399 retrievals from the MODIS satellite have been used to validate the modeled AOD (Fig. 5). It is

400 observed that the model captures several important features of the MODIS retrieved AOD spatial
401 distribution but at the same time somewhat struggles to reproduce the observed AOD magnitude in
402 some parts of the domain. One possible reason for the underestimation would be the EDGAR-HTAP
403 emission inventory, which has a low bias for residential sector PM_{2.5} emissions in India (Sharma et al.,
404 2022). For instance, the model successfully predicts high aerosol loading seen by MODIS on 20 and
405 21 December over CIGP and EIGP. This is the region with dense fog both in model and observation.
406 Higher AOD (>0.5) over CIGP and EIGP can be attributed to the accumulation of aerosols that are
407 transported by north-westerly winds to these regions from NWIGP (Dey and Di Girolamo, 2011; Jain
408 et al., 2020; Jethva et al., 2018; Kumar et al., 2018; Yadav et al., 2020). However, WRF-Chem
409 underestimates AOD over the NWIGP (AOD<0.3) throughout the simulation period and during 23-24
410 December over CIGP and EIGP where the latter may be related to enhanced scavenging of aerosols by
411 fog droplets.

412 The west to east gradient in aerosol loading over IGP is consistent with surface PM_{2.5}
413 distribution (Fig. 6a). Surface PM_{2.5} concentration is highest in EIGP (>100 µg m⁻³) and it decreases
414 gradually towards NWIGP (~60-80 µg m⁻³). The time series of PM_{2.5} from CPCB measurements and
415 the model at stations representative of each region in IGP shows that simulated PM_{2.5} compares well
416 with observation in terms of day-to-day variation over most of the locations in the IGP (Fig. 6 b-e).
417 The comparison is good over Amritsar (an NWIGP location), where PM_{2.5} is mostly primary aerosols
418 from local emissions e.g., residential heating related biomass burning. Agricultural waste burning is at
419 its peak during post monsoon months (Oct-Nov), whereas during winter burning for residential
420 heating increases and the stable boundary layer confines these emissions near the surface (Kumar et
421 al., 2021; Pawar and Sinha, 2022). PM_{2.5} at Amritsar shows a bimodal distribution with morning and
422 evening peaks whereas it is absent in the model likely due to the absence of diurnal variations in the
423 WRF-Chem anthropogenic emissions.

424 At Delhi, the daily variations are predicted well although WRF-Chem underestimates PM_{2.5}
425 observations during the first 4 days. Delhi experiences severe air pollution and haze with high PM
426 loading (> 500 µg m⁻³) (Bharali et al., 2019). The model is successful in predicting the high PM_{2.5}
427 episode on the 24th of December, but WRF-Chem underpredicts the SO₄²⁻, NH₄⁺, NO₃⁻ and Cl⁻
428 concentrations (Fig. 7). Although simulated SO₂ and NH₃ are comparable with observation, sulfate,
429 and ammonium are underestimated in the model. SO₄²⁻ is underestimated by ~ 9 µg m⁻³, while NH₄⁺,
430 NO₃⁻ and Cl⁻ are underestimated by ~30 µg m⁻³, ~19 µg m⁻³ and ~40 µg m⁻³ on average, respectively.
431 In addition, the WRF-Chem model results show that a large percentage of PM_{2.5} is classified as “other
432 inorganics”, which is usually dominated by PM_{2.5} other than BC and OC. The WiFEx observations
433 show that secondary aerosols contribute ~50% of the PM_{2.5} concentration, which is in line with other
434 studies that found secondary aerosols contribute 15-50% to PM_{2.5} mass (Sharma and Mandal, 2017;
435 Behera and Sharma, 2010; Nagar et al., 2017) and NO₃⁻ constituted 9-13% of PM_{2.5} mass (Lalchandani
436 et al., 2021; Sharma and Mandal, 2017). This leads to the underestimation of PM_{2.5} over Delhi.

437 Studies report very high chloride over the IGP with values exceeding $100 \mu\text{g m}^{-3}$ (Lalchandani et al.,
438 2021) during winter emitted from increased trash burning and industrial emissions (Pant et al., 2015;
439 Patil et al., 2013). WRF-Chem incorporates trash-burning emissions which include HCl emissions
440 from Chaudhury et al.,(2021) for this study however, the inventory contains annual emissions and fails
441 to resolve the seasonality of trash-burning emissions as identified by Nagpure et al., (2015). They
442 suggested almost all the waste-burning emissions in neighbourhoods with higher socioeconomic status
443 in Delhi occur due to the use of waste as cheap heating fuel by individuals such as night watchmen
444 and pavement dwellers. Chaudhary et al., (2021) considers waste burning that occurs due to lack of
445 collection infrastructure, and at landfills and, therefore, shows a concentration of waste burning
446 emissions around the periphery of Delhi but low waste burning emissions in the relatively prosperous
447 city centre. In addition, emissions from other sources (e.g., industries) are unaccounted for in the
448 model which likely leads to the underestimation in modeled chloride.

449 Over the CIGP and EIGP, the underestimation in $\text{PM}_{2.5}$ is mostly observed during the dense
450 fog. It is well known that the hygroscopic aerosols grow in size and are deposited to the surface during
451 fog (Gupta and Mandariya, 2013; Kaul et al., 2011). $\text{PM}_{2.5}$ shows an increase initially with the onset
452 of fog and then it decreases as the aerosols grow and get deposited through fog droplets. A two order
453 higher deposition rate (Fig. 6 f, g) during fog compared to the deposition rate of dry aerosol results in
454 the lower $\text{PM}_{2.5}$ over CIGP and EICP during fog events.

455 A statistical analysis (Table S1) shows a minimum normalized mean bias (NMB) for $\text{PM}_{2.5}$ at
456 Amritsar (2.2%), while in other stations, the normalized mean bias ranges from 48 to 53%, similar to
457 the reported range of model bias (underestimated by 40–60%) in winter over IGP by earlier studies
458 (Bran and Srivastava, 2017; Ojha et al., 2020). This was accomplished by incorporating trash-burning
459 emissions in the model simulation, which improved the $\text{PM}_{2.5}$ prediction, increasing NMB by ~4-8%
460 in IGP. RMSE values range from 41 to $138 \mu\text{g m}^{-3}$ (normalized RMSE~0.4 to 0.7) comparable to the
461 reported values by these studies. The Pearson correlation coefficient (r) for the simulated and observed
462 day-to-day variation in $\text{PM}_{2.5}$ lies between 0.4 and 0.7 for all the stations in Fig. 6 except at Patna and
463 Muzaffarpur which lies within the range in these studies. The poor correlation at Patna and
464 Muzaffarpur is due to the low modeled $\text{PM}_{2.5}$ concentrations which are caused by increased dry
465 deposition of aerosol particles activated as fog droplets during fog periods, as discussed earlier.
466 Furthermore, the fog events in WRF and observations have somewhat different time periods causing
467 WRF-predicted and the observed $\text{PM}_{2.5}$ concentrations to decrease at different times.

468 Previous studies have reported that models tend to underestimate the AOD observation (David
469 et al., 2018; Pan et al., 2015) during the post-monsoon and winter when agricultural waste burning and
470 anthropogenic emissions dominate. While anthropogenic emissions include a contribution from the
471 residential sector, the emissions from small-scale burning for residential heating over IGP especially
472 during winter are likely underestimated in the current emission inventory (Sharma et al., 2022). This
473 leads to an underestimation of aerosol concentration in the model. Other possible causes for the

474 underestimation are the biases in the simulated meteorology (Govardhan et al., 2015; Kumar et al.,
475 2015; Pan et al., 2015) which affects the aerosol concentration. We corrected some of the biases in
476 meteorology as discussed earlier however there are still residual biases in the simulated meteorology
477 e.g., overestimation of wind speed by WRF-Chem. We also observe underestimation of secondary
478 aerosols over NWIGP which contribute significantly to the aerosol loading over IGP. Secondary
479 aerosol formation is substantial over CIGP and EIGP in the model compared to NWIGP which will be
480 discussed in a later section. The underestimation of $PM_{2.5}$ could also be linked to the uncertainty in the
481 model's chemistry scheme to simulate the secondary aerosols due to missing chemical processes or
482 due to underestimation of sulfur oxidation at different RH levels (Acharja et al., 2022; Pawar et al.,
483 2023; Ruan et al., 2022). Moreover, several modeling studies have shown significant improvements in
484 forecasting surface $PM_{2.5}$ by assimilation of satellite AOD and $PM_{2.5}$ (Ghude et al., 2020; Jena et al.,
485 2020; Kumar et al., 2020) suggesting the importance of correct initialization of the model in
486 simulating aerosols over IGP.

487

488 **5 Effect of Aerosol Radiation feedback**

489 Interactions of aerosols with radiation affects temperature and surface heat fluxes, thereby
490 weakening the turbulence in the PBL and stabilizing the boundary layer height (Fig. 8b) compared to
491 the clean environment (Fig. 8a). In the presence of well mixed aerosols within the PBL, the radiative
492 effect of aerosols lowers the noontime PBL height (Fig. 8b). However, the presence of absorbing
493 aerosols in the PBL warms the air and changes the thermodynamics. Three cases are shown in Fig.
494 8(c-e) where increases of scattering aerosol concentrations at the top of PBL (Fig. 8c) increases
495 scattering of radiation by the aerosol layer and reduces the surface reaching solar radiation similar to
496 Fig. 8b. Higher concentrations of absorbing aerosols at the top of PBL (Fig. 8d) warms the air above
497 the boundary layer and strengthens the capping inversion stabilizing the PBL and suppressing its
498 growth. The shallow PBL and weakened daytime vertical mixing confines aerosols and water vapor
499 near the surface and worsens the air quality of a region. The aerosols trapped in the stagnant PBL
500 further affects the radiation flux at the surface and creates a positive feedback loop wherein the PBL is
501 continually suppressed until interrupted by some synoptic weather phenomenon, such as the western
502 disturbances in the IGP. On the other hand, higher concentration of absorbing aerosols within the PBL
503 (Fig. 8e) warms the air in the PBL and this results in the higher PBL height. The raised PBL decreases
504 the aerosol concentration near the surface which is termed as a negative feedback effect.

505 The aerosol radiation feedback can affect shortwave heating rates (SWHR). The high aerosol
506 loading over the IGP (Fig. 6 and Fig. 7) allows the AR feedback to reduce the PBL height by more
507 than 140 m throughout the IGP compared to the surrounding region with AR feedback (Fig. 9a). The
508 difference in PBL height with and without aerosol radiation feedback is largest during noontime. The
509 suppressed PBL is due to the decrease in the surface heating flux and the consequent weakening of
510 turbulence in the PBL. The surface solar radiation flux (SWF) decreases by 5-35 % while the surface

511 latent heat (LH) and sensible heat (HFX) fluxes decrease by 5-35 % and 10-60 %, respectively (Fig.
512 S3). The stable, shallow PBL reduces the vertical mixing of aerosols and moisture and confines them
513 near the surface, resulting in increased $PM_{2.5}$ concentrations and RH near the surface with AR
514 feedback (Fig. 9). Although T2 should decrease with the reduction in surface SWF, T2 shows mixed
515 signals with both cooling and warming over IGP. While surface cooling is observed over NWIGP and
516 EIGP, T2 increases with AR feedback over most of CIGP. The response of AR feedback to T2 varies
517 in these three regions probably due to differences in the distribution and types of aerosols and the
518 presence of fog. Increase in surface concentration of $PM_{2.5}$ occurs more over NWIGP and EIGP with
519 increase in BC and OIN over NWIGP, and sulfate aerosol over EIGP which results in the surface
520 cooling due to positive AR feedback in these two regions. Over the CIGP, the AR feedback causes a
521 depletion of surface $PM_{2.5}$ (Fig. 9d), which is likely due to their hygroscopic growth, and then dry
522 deposition (average $PM_{2.5}$ dry deposition flux of $331 \mu\text{g m}^{-2} \text{hr}^{-1}$ with AR feedback and $282 \mu\text{g m}^{-2} \text{hr}^{-1}$
523 without AR feedback) in dense fog. The increase in RH with AR feedback favours the growth of
524 aerosols in size by the uptake of water

525 Examining further, the time variation of the changes in PBL height, T2, and RH between the
526 simulations with and without aerosol-radiation feedback (Fig. 9g) shows an increase in T2 while the
527 surface fluxes, sensible heat flux, latent heat flux, and downward shortwave radiation flux decrease
528 over CIGP (Fig. 9h). AR feedback affects mostly the lower atmosphere at multiple levels; however,
529 our finding suggests that the decreased shortwave radiation flux decreases the surface fluxes and thus
530 the turbulence in the boundary layer resulting in a reduced PBL height on both days. Figure 9 g and h
531 clearly show a decrease in HFX and LH following the decrease in SWF. Moreover, we observe that
532 the PBL height is sensitive to latent heat flux likely due to its strong dependence on moisture
533 availability (Xiu and Pleim, 2001; Zhang and Anthes, 1982).

534 The impact of AR feedback on T2 depends on factors such as the presence of absorbing
535 aerosols and their vertical distribution via heating or increased SWF (as observed in CIGP, Fig. S4).
536 Absorbing aerosols in WRF-Chem include BC and OIN (other inorganic aerosols), which both
537 increase near the surface (Fig. 9e, Fig. S5) due to their confinement in the stable PBL. Some areas in
538 the fog-affected region show a decrease in BC as well as SO_4^{2-} likely due to increased dry deposition
539 in fog water as discussed earlier in this section for $PM_{2.5}$. As a result, AR feedback changes the
540 absorbing to scattering ratio of aerosols over IGP indicated by the decrease in SSA (Single Scattering
541 Albedo; Fig. S6). In EIGP, sulfate concentrations are larger with AR feedback than without AR
542 feedback with time periods where the difference is $>1 \mu\text{g m}^{-3}$ (Fig.11). The BC concentration changes
543 are small ($<0.5 \mu\text{g m}^{-3}$) in the EIGP, resulting in a higher SSA near the surface with AR feedback in
544 EIGP. In the CIGP, BC concentrations increase while sulfate aerosols decrease within the PBL with
545 AR feedback (Fig.11) compared to the simulation without AR feedback. A decrease in SSA is seen for
546 the CIGP throughout the boundary layer while in EIGP the decrease occurs near the top of the PBL;
547 difference in SSA due to AR feedback is negligible in NWIGP. Also contributing to the higher SSA in

548 EIGP is the increase in RH (Fig. 9) due to AR feedback favoring the growth of aerosols in size by
549 uptake of water and the production of secondary aerosols such as SO_4^{2-} and NH_4^+ .

550 A similar observation has been made by Ramachandran et al., (2020) where SSA decreases
551 with increasing altitude due to absorbing carbonaceous aerosols at higher elevations which contributes
552 $\geq 75\%$ to the aerosol absorption over IGP. Increased shortwave heating (Fig. 10) is probably caused by
553 the increased absorbing aerosols near the surface which overwhelms the surface cooling due to
554 reduced shortwave radiation at the surface.

555 The increase in 2-m RH is substantial over CIGP on 24th December (Fig. 9g) compared to the
556 previous day following the decrease in PBL height which constrains the moisture near the surface.
557 The decrease in RH by 2% or more when aerosol-radiation feedback is included compared to no
558 aerosol-radiation feedback is likely due to increase in T2. However, the increase in RH in the
559 afternoon associated with a decrease in LH and PBL height is important for the air to saturate which
560 then favors the formation of fog in a polluted environment. Note that the increase in T2 with AR
561 feedback is very small ($< 0.5^\circ\text{C}$) which reduces further after noon ($\sim 12:30$ pm IST) on both days.

562 Another important factor that can affect the extent of change in PBL height is the distribution
563 of aerosols in the vertical (illustrated in Fig. 8). The pressure-time cross-sections of differences in T,
564 $\text{PM}_{2.5}$, BC, and SO_4^{2-} between aerosol radiation (AR) feedback (wFB) and no aerosol radiation
565 feedback (nFB) for three regions, NWIGP, CIGP, and EIGP are shown in Fig. 11. The difference in
566 the PBL height reaches a maximum with the AR feedback during midday (12:30-15:30 IST). Increase
567 in temperature in the boundary layer is observed with AR feedback particularly at the upper PBL in all
568 the regions of IGP. This induces a temperature inversion resulting in a stable and suppressed PBL. In
569 all the regions the decrease in PBL height (100-200 m) is larger on 24th December compared to 23rd
570 December. The difference in the PBL height on 23rd and 24th December with AR feedback on these
571 days is possibly controlled by the aerosol distribution during the previous day or early morning on the
572 same day. For example, in all the regions an increase in $\text{PM}_{2.5}$ is observed the previous night (23:30
573 onwards) till $\sim 11:30$ of December 24, with increased BC over NWIGP and CIGP whereas both BC
574 and SO_4^{2-} over EIGP. The increased $\text{PM}_{2.5}$ concentrations suppress the development of the PBL after
575 sunrise with AR feedback on December 24 compared to that on December 23, leading to the observed
576 differences in ΔPBL height on these two days. Increase in BC concentrations in NWIGP and CIGP are
577 found above the PBL on 24th December whereas BC concentrations decrease within the PBL. This BC
578 concentration gradient creates a temperature inversion, for example between 10:30-14:30 IST. The
579 increase in BC warms the air in the PBL; however, the warming is not strong enough to cause
580 negative feedback over CIGP. On 23rd December a small increase in BC is uniform throughout the
581 PBL, while there is a decrease in SO_4^{2-} concentrations, resulting in a warmer PBL (Fig. 11) with AR
582 feedback.

583 In EIGP, BC distribution is similar to that in CIGP with AR feedback while there is a
584 substantial increase in sulfate aerosol in the PBL. This results in the strongest extinction in EIGP as

585 evident from the largest difference in PBL height and surface cooling with AR feedback among the
586 three regions. Although Δ PBL is small on 23rd December, it still results in the accumulation of
587 aerosols during night-time (~23:30 pm onwards) which further strengthens the AR feedback effect the
588 next day in NWIGP and CIGP. Thus, AR feedback stabilizes the PBL, increases PM_{2.5} and RH in the
589 PBL making conditions favourable for persistence of fog over IGP.

590

591 **6 Effect of Aqueous phase chemistry**

592 In this section we discuss the impact of aqueous phase chemistry on aerosol composition and
593 its interaction with meteorology. There is a considerable difference in the surface concentration of
594 PM_{2.5} (>16 $\mu\text{g m}^{-3}$) in the absence of aqueous chemistry over CIGP and EIGP where fog occurs (Fig.
595 12a) while the difference is negligible over NWIGP where fog does not occur. This is due to the
596 formation of secondary aerosols through aqueous phase chemistry and the hygroscopic growth of
597 aerosols during fog in these regions with the inclusion of aqueous chemistry in the model. In the
598 region between CIGP and EIGP (83E-84E; marked by the box in Fig. 12a), PM_{2.5} concentration is less
599 in the simulation with aqueous-phase chemistry than without aqueous-phase chemistry because
600 deposition of fog water aerosols to the surface increases as the fog thickens (Fig. 13, Fig. S7). Figure
601 13 shows the relation between formation of secondary aerosols, deposition flux of PM_{2.5}, and fog with
602 and without aqueous phase chemistry. During the fog event, the secondary aerosols (SO₄²⁻, NH₄⁺)
603 increase significantly by 4-10 $\mu\text{g m}^{-3}$ due to aqueous phase chemistry adding to the PM_{2.5} burden over
604 IGP. The intensity of fog is high around midnight December 24-25 compared to that on 23rd and 24th
605 (1:30-11:30 IST)) which increases the dry deposition flux of PM_{2.5} causing a sharp drop in the PM_{2.5}
606 concentration on 24th December compared to the previous night's fog event. The observed change in
607 PM_{2.5} over a region is the net result of the formation of secondary aerosols and its deposition with fog
608 droplets.

609 The composition distribution of PM_{2.5} (Fig. 12b) has a similar distribution for the simulations
610 with and without aqueous phase chemistry over NWIGP where fog did not occur. The primary
611 aerosols are higher (BC > 9%, OC ~ 16-30%, OIN > 50%), than the secondary aerosols (<5%). While
612 the model requires fog for accelerated formation of secondary inorganic aerosol, experimental data
613 (Fig. 7) supports significant formation of secondary inorganic aerosol at elevated RH levels even in
614 haze aerosol (Acharja et al., 2022). On the other hand, the central and east IGP stations are fog-
615 covered and therefore, there is an increase in secondary aerosols especially SO₄²⁻ and NH₄⁺ when
616 aqueous phase chemistry is included in the simulation. SO₄²⁻ is chemically produced via aqueous
617 phase chemistry in cloud water, hence the abrupt increase whereas NH₄⁺ maintains a gas-aerosol and
618 gas-cloud equilibrium with NH₃ and SO₄²⁻ via neutralizing the drop or aerosol. NO₃⁻ is high in the
619 model compared to SO₄²⁻ and NH₄⁺ and it decreases by ~1-2 % with aqueous phase chemistry. We
620 observe a small increase in NO₃⁻ during fog, however it drops as fog intensifies, more rapidly than that
621 without aqueous phase chemistry likely due to increase in dry deposition. This results in lower

622 average NO_3^- to $\text{PM}_{2.5}$ ratio with aqueous phase chemistry. Moreover, NO_3^- is high over the fog
623 covered CIGP and EIGP compared to NWIGP suggesting that transport and chemistry of NO_x in
624 CIGP and EIGP produce more HNO_3 . Aerosol NO_3^- is also in equilibrium with HNO_3 and it is formed
625 only if excess NH_3 is available beyond the sulfate neutralization. Thus, NH_4^+ and NO_3^- changes are
626 likely due to changing the partitioning between gas and liquid based on the production of sulfate.

627 $\text{PM}_{2.5}$ is mostly composed of organic aerosols (OA) over CIGP and EIGP (Lalchandani et al.,
628 2021; Srinivas and Sarin, 2014) whereas $\text{PM}_{2.5}$ is OIN (dust) and OA over NWIGP (Ram et al., 2012a;
629 Sharma and Mandal 2023). Although observational studies report Cl^- as one of the largest contributors
630 (12-17%) to $\text{PM}_{2.5}$ after the organics (Lalchandani et al., 2021; Pant et al., 2015) during winter, Cl^- is
631 largely underestimated by the model as discussed in section 4 and contributes only ~3%. A small
632 increase (2-4%) in secondary organic aerosols (SOA) from glyoxal production in aerosols occurs for
633 the simulation with aqueous phase chemistry included during intense fog, suggesting there are
634 feedbacks between cloud chemistry (without glyoxal aqueous chemistry) and aerosol chemistry.
635 However, similar to NO_3^- , average SOA (ASOA (anthropogenic)+BSOA (biogenic) + GlySOA)
636 shows a decrease when aqueous phase chemistry is included. SOA contributes significantly to organic
637 aerosol loading over IGP (Kaul et al., 2011; Mandariya et al., 2019).

638 The WRF-Chem results on aerosol composition during fog behave similarly to those of
639 previous observational studies. For example, Ram et al., (2012a) reported an increase of EC, OC, and
640 WSOC concentrations by ~30% during fog and haze events at Allahabad, a location in the Central
641 IGP, and a marginal increase of these constituents at Hisar (NWIGP). Several studies report an
642 increase in inorganic ions (NH_4^+ , NO_3^- , and SO_4^{2-}) during fog over IGP and elsewhere (Gundel et al.,
643 1994; Ram et al., 2012a). Recent studies suggest that a significant fraction of atmospheric particulate
644 matter in the IGP is comprised of carbonaceous aerosol (~30–35% of the PM) and water-soluble
645 inorganic species (~10–20% of the PM) during October–January when emissions from biomass
646 burning (including residential heating) are dominant over IGP (Ram et al., 2012b; Rengarajan et al.,
647 2007; Tare et al., 2006).

648 Both the simulations with and without aqueous-phase chemistry include the AR feedback. The
649 aqueous chemistry increases the mass of $\text{PM}_{2.5}$ and the size of the aerosols, both of which contribute to
650 AR feedback, thus increasing RH and PBL stability. The increase in RH also saturates the air,
651 promotes aerosol growth by water uptake, and thus favors fog formation. Since the secondary
652 inorganic aerosols are scattering aerosols, the increased scattering of radiation further reduces the
653 solar radiation reaching the surface (Fig. 14a). Over CIGP the presence of higher aerosol loading
654 reduces the T2 during daytime, particularly on the 24th of December which then reduces the PBL
655 height and increases RH near the surface (Fig. 14b). These conditions favor fog formation over the
656 CIGP. Further, the fog water content with aqueous-phase chemistry is higher than that without
657 aqueous-phase chemistry on 24th December post-midnight (Fig. 13b). This is likely due to saturation
658 of air due to increase in RH and lower T2, induced by the AR feedback caused by the increase in

659 PM_{2.5}. Although the difference in T2 is small (<0.4), favourable conditions mentioned above are
660 conducive to fog formation. Because aqueous chemistry increases sulfate concentrations, the size of
661 the aerosols also increase. The increased aerosol size, which can grow further by water uptake, also
662 impacts the solar radiation reaching the surface, affecting fog formation and dissipation.

663

664 **7 Effect of AR feedback and aqueous chemistry on the duration of fog**

665 Aerosol and its radiative effects impact fog characteristics, including the fog liquid water content
666 (LWC), the fog lifetime over a region and hence its spatial and temporal distribution. Variations of fog
667 LWC in WRF-Chem contrast the fog in the CIGP and EIGP (Figure 15) as well as among the three
668 experiments (with aqueous chemistry plus AR feedback, with aqueous chemistry without AR
669 feedback, and without aqueous chemistry but with AR feedback). WRF-Chem does not simulate fog
670 over NWIGP in the model for the study period. In Figure 15, only foggy grid points are considered for
671 the first fog event on 23-24 December. The LWC is 5-15% higher with AR feedback than without AR
672 feedback and without aqueous phase chemistry for both CIGP and EIGP. The interquartile range is
673 larger for the simulation with and without AR feedback than without aqueous phase chemistry in
674 CIGP showing large variability in the LWC. On the other hand, in EIGP the variability in LWC is
675 greater in the simulation with AR feedback compared to the other two experiments.

676 The formation and dissipation times of the two fog events for the three experiments are listed
677 in Tables 2 and 3 for CIGP and EIGP. The 23-24 December fog starts forming two hours earlier and
678 the 24-25 December fog forms one hour earlier in both CIGP and EIGP with AR feedback than
679 without AR feedback. In the simulation without aqueous phase chemistry, fog formation is delayed by
680 an hour or two compared to the simulation with aqueous chemistry plus AR feedback in CIGP. In
681 EIGP the 23-24 December fog forms at the same time with AR feedback and without aqueous phase
682 chemistry while the 24-25 December fog is delayed by an hour without aqueous phase chemistry. Fog
683 dissipation usually occurs after sunrise when the shortwave radiative warming at the surface warms
684 the air, which results in PBL mixing. In addition, absorbing aerosols like BC affect fog dissipation by
685 increasing the radiative heating in and above the fog. We find an increase in BC and shortwave
686 heating in the PBL with AR feedback (Fig. 10,11) and warming over CIGP with AR feedback. Fog
687 intensity starts to decrease after 01:00 UTC (06:30 IST), however, in our study, we find that the fog
688 dissipates completely in the afternoon (~10:00 UTC or 15:30 IST) for both the simulations with AR
689 feedback and no aqueous chemistry while an hour later without AR feedback in CIGP. Fog dissipation
690 is delayed in EIGP with AR feedback compared to that without AR feedback and without aqueous
691 phase chemistry. In both the regions, fog lifetime increases with AR feedback. All the stations,
692 however do not show the same pattern, for example, the 23-24 December fog in Lucknow forms and
693 dissipates at the same time for simulations with AR feedback and without aqueous phase chemistry,
694 and the 24-25 December fog forms later with AR feedback than without AR feedback. Patna shows no
695 difference in the 24-25 December fog formation in all the three experiments. To gain better insights on

696 the fog timing, we recommend that simulations at higher spatial and temporal resolutions be
697 performed to represent better the fog dynamics at point locations. Furthermore, there are other
698 important factors to consider, e.g., improved emissions, better simulations of aerosol chemical
699 composition, and evaluation of aerosol deposition.

700 The AR feedback and aqueous-phase chemistry have the potential to impact aerosol-fog
701 interactions. We can learn about the effect of the aerosol-radiation interactions on CCN concentrations
702 because the WRF-Chem model calculates the CCN concentrations at different supersaturations as a
703 diagnostic output. We compare CCN at 0.02% supersaturations, a value typical of fog, among the
704 three experiments. For the 23-24 December fog in CIGP, hourly CCN concentrations are ~10% higher
705 for the simulations with AR feedback with or without aqueous chemistry than with no AR feedback
706 (Figure S8) during the first 8 hours of the fog event (16:00-24:00 IST 23 December). Surprisingly, the
707 simulation with no aqueous chemistry has higher CCN concentrations than the simulation with
708 aqueous chemistry, as more CCN are expected with aqueous chemistry, because aqueous chemistry
709 adds sulfate to the aerosol mass increasing the mass of PM_{2.5}. Increased PM_{2.5} further contributes to
710 AR feedback, increasing RH which favours the growth of the aerosol size, categorizing more aerosols
711 as CCN. However, the dry deposition flux (ddmass) also increases in dense fog which causes rapid
712 loss in CCN and activated aerosols during fog events with the AR feedback (Fig. S7) and more so
713 without aqueous-phase chemistry. Shao et al. (2023) examined aerosol-fog interactions for two
714 consecutive fog events by comparing WRF-Chem results with current emissions strengths to those
715 with low emission strengths. They show that the first fog event promotes formation of the second fog
716 event leading to wider fog distribution, and longer fog lifetime favoured by multiple feedbacks
717 including AR feedback i.e., low temperature, high humidity and high stability similar to our study.
718 While Shao et al. (2023) observe a delay in dissipation of the first event and early formation of second
719 fog event, we find an early dissipation and early formation of fog with AR feedback as discussed
720 earlier in this section. In summary, aqueous phase chemistry together with AR feedback promotes
721 early formation of fog while AR feedback alone promotes early dissipation of fog and plays a critical
722 role in the formation and evolution of the fog over IGP.

723

724 **8 Conclusions**

725 The effects of aerosol-radiation (AR) feedback and aqueous chemistry in air quality and fog
726 have been assessed over IGP. We carried out three experiments using WRF-Chem testing different
727 combinations of PBL schemes and meteorology initial and boundary conditions. The best
728 representation of surface meteorology for the IGP region for the case study (December 20-24, 2017)
729 used ERA-Interim reanalysis to drive the meteorology and ACM2 PBL scheme with soil moisture
730 nudging to ERA-Interim. With this meteorology configuration for WRF-Chem, evaluation of aerosol
731 concentrations with measurements and the impact of aerosols on atmospheric processes during fog
732 were examined. Further, we included trash-burning emissions to represent anthropogenic chloride

733 aerosols in our configuration. Incorporation of trash burning emissions did improve the model
734 simulations of $PM_{2.5}$ and better captured the day-to-day variability of $PM_{2.5}$ in IGP, yet underestimated
735 its magnitude compared to CPCB observations. Moreover, secondary aerosols particularly, chloride
736 aerosols are underestimated in the model. This underestimation is likely caused by a low bias in the
737 residential burning emission inventory and a failure of the emission inventory to represent residential
738 sector emissions from the use of trash as cheap heating fuel properly. AOD regional distribution is
739 predicted well by the model for most of the IGP. However, AOD is underestimated over NWIGP
740 likely due to an underestimation of fugitive emissions during wintertime cold spells.

741 The AR interactions showed a significant impact on meteorology and air quality over IGP. A
742 WRF-chem simulation with AR interactions resulted in a lower PBL height by ~50-270 m compared
743 to a simulation without AR interactions leading to accumulation of aerosols and moisture near the
744 surface. Reduced surface shortwave radiation flux and the surface sensible and latent heat fluxes due
745 to aerosol radiative effect suppressed the turbulence resulting in a stable PBL. The shallow PBL
746 further increased surface $PM_{2.5}$ ($> 8 \mu g m^{-3}$) and RH (2-8%) over IGP and this positive feedback
747 mechanism promoted thickening of fog over IGP. However, an increase in absorbing aerosols in the
748 PBL gave negative feedback, increasing the shortwave heating and temperature particularly over
749 CIGP. Fog forms when air is saturated which occurs when the surface temperature is reduced or the
750 moisture content increases causing saturation of air. This study suggests that increase in RH saturated
751 the air and the increase in aerosols favoured fog formation as depicted by the thickening of fog
752 intensity. Aqueous phase chemistry on the other hand contributed significantly to secondary aerosols
753 in the fog, especially sulfate aerosols, indicating substantial formation of secondary aerosols in the
754 cloud. The underpredicted secondary aerosols over NWIGP where no fog occurred implies
755 underestimation of formation of aerosols through gas and aerosol chemistry in the model. This
756 underestimation could also be linked to an underestimation of pH in the default MOSAIC scheme
757 (Ruan et al., 2022) which slows the secondary aerosol formation, or an underestimation of the aqueous
758 sulfur oxidation in haze aerosol at $> 80\%$ RH before the onset of fog (Acharja et al., 2022), or missing
759 multiphase oxidation processes (Wang et al., 2022). Nevertheless, we find that the model successfully
760 simulates the same changes in the inorganic composition during fog in IGP as reported by
761 observational studies referred earlier in section 6. We also observed that AR feedback with aqueous
762 chemistry initiated the fog formation 1-2 hours earlier than the initiation time in the simulation
763 without AR feedback and without aqueous phase chemistry whereas AR feedback alone led to early
764 dissipation of fog. In addition, fog acted as an important sink of aerosols in a polluted environment
765 with increased dry deposition with cloud water. Thus, AR feedback and aqueous chemistry play a
766 significant role in modulating the distribution and concentration of aerosols and evolution of fog in the
767 PBL. Aerosol-cloud interactions were not investigated in this study due to the limitation of the ACM2
768 PBL scheme in providing necessary information with other modules in WRF. Previous studies of
769 aerosol-fog interactions have found that ACI also promotes early onset of fog formation and increases

770 fog duration (Maalick et al., 2016; Yan et al.,2021). While these previous studies were applied to
771 midlatitude fog events, it is likely that ACI also plays a dominant role in IGP fogs, suggesting that
772 future studies are needed to fully understand aerosol effects on IGP fog events.

773 The large emission of aerosols and trace gases in the IGP makes the atmospheric dynamics as
774 well as chemistry complex, suggesting the need for more studies using both models and ground-based
775 measurements to better understand the processes. While all aerosol types interact with solar radiation
776 and reduce the surface reaching flux, presence of absorbing aerosols in the boundary layer and its
777 vertical distribution plays an important role in modulating the meteorology over IGP. It is therefore
778 crucial to improve the simulation of absorbing aerosols e.g., BC in the vertical as well as at the surface
779 to increase the accuracy in predicting formation as well as the dissipation of fog in this region.
780 Emissions from burning for residential heating are an important source of aerosols in IGP during post-
781 monsoon and winter and the inclusion of these sources in the emission inventory would improve the
782 prediction of wintertime aerosols. For example, the underestimation of chloride aerosol in the model
783 indicates unaccounted emission sources over IGP and the need for more work on better quantifying
784 trash burning emissions, which may not only improve particulate chloride in the model but also
785 improve simulations of other aerosol chemical components through aerosol thermodynamics.
786 Additionally, more detailed modeling studies are required to understand the missing chemical
787 processes if any in the model which leads to biases in sulfate, nitrate and ammonium partitioning
788 between gas and aerosol phases. We find that the change in PBL height with AR feedback is sensitive
789 to changes in LH, signifying the role of soil moisture in PBL dynamics. Several studies have reported
790 cooling over IGP due to an increase in irrigation (Kumar et al., 2017; Mishra et al., 2020). Further
791 investigations into the role of irrigation in the increasing fog events over NWIGP would help in better
792 understanding the formation and persistence of fog over this region. It can be concluded that fog
793 forecasting is a complex process due to the multiple factors involved and this work suggests that AR
794 feedback is important in fog forecasting while aqueous phase chemistry plays an important role in
795 defining the composition of aerosols over IGP.

796

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812

813 **Data availability:** All the model simulations are archived on the NCAR campaign storage
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815 corresponding author. WIFEX data can be made available by contacting Dr S.D. Ghude. Trash
816 Burning emission data is available on Mendeley data (doi- <http://dx.doi.org/10.17632/t2tn4t9473.1>).
817 MODIS AOD retrievals can be downloaded from <https://earthdata.nasa.gov/>. Meteorology and
818 aerosol data from the Central Pollution Control Board (CPCB) is available at <http://cpcb.nic.in>.

819

820 **Author contributions:**

821 CB: Conceptualization, Formal Analysis, Writing

822 MB: Conceptualization, Supervision, Writing-review and editing, Funding acquisition

823 RK: Conceptualization, Supervision, Writing-review and editing

824 SDG: provided ground-based observation data, Writing-review and editing

825 VS and BS: provided trash burning emission data, Writing-review and editing

826

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828

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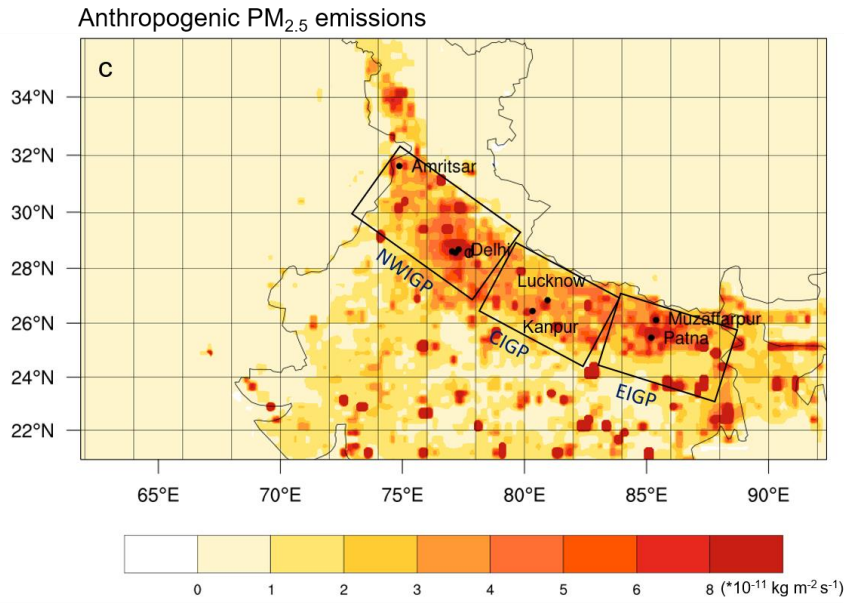
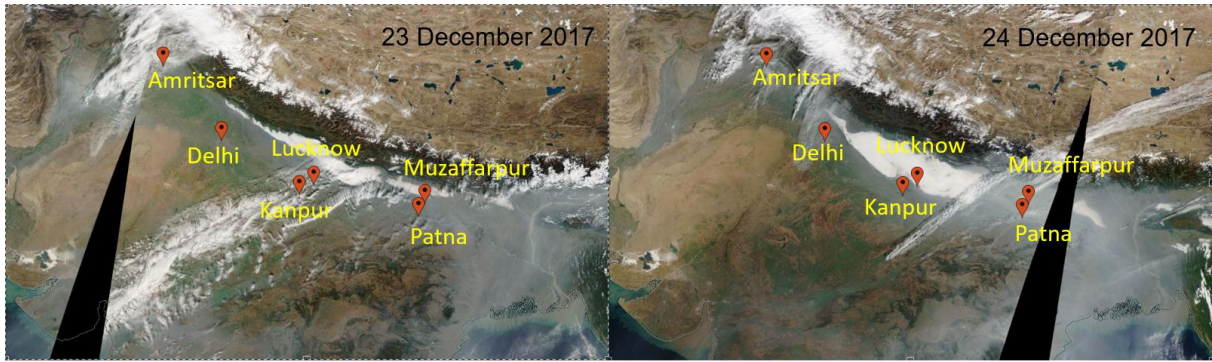
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Table 1 Experiment set-up for the study. Numbers in parentheses for the physics options denote the namelist settings of the WRF-Chem model.

	EXP 1	EXP 2	EXP 3
Meteorology Initial /lateral boundary Condition:	<i>NCEP Final Analysis (GFS-FNL), 1° x 1°, 6 hourly</i>	<i>ERA-Interim Project, 1.125° x 0.703°, 6 hourly</i>	<i>ERA-Interim Project, 1.125° x 0.703°, 6 hourly</i>
Physics Options			
Cloud Physics	<i>Morrison 2-mom (10)</i>	<i>Morrison 2- mom (10)</i>	<i>Morrison 2- mom (10)</i>
Longwave Radiation	<i>RRTMG scheme (4)</i>	<i>RRTMG scheme (4)</i>	<i>RRTMG scheme (4)</i>
Shortwave Radiation	<i>Goddard shortwave (2)</i>	<i>RRTMG scheme (4)</i>	<i>RRTMG scheme (4)</i>
Surface Layer Physics	<i>Revised MM5 Monin-Obukhov scheme (1)</i>	<i>Revised MM5 Monin-Obukhov scheme (1)</i>	<i>Pleim-Xiu (7)</i>
Surface Model	<i>unified Noah land-surface model (2)</i>	<i>NoahMP (4)</i>	<i>Pleim-Xiu (7)</i>
PBL Scheme	<i>YSU scheme (1)</i>	<i>YSU (1)</i>	<i>ACM2 (7)</i>
Convective Parameterization	<i>Grell-Freitas (3)</i>	<i>Grell-Freitas (3)</i>	<i>Grell-Freitas (3)</i>
	<i>Continuous simulation</i>	<i>*Meteorology refreshed every 24 hr</i>	<i>**Continuous simulation: Soil nudging included</i>

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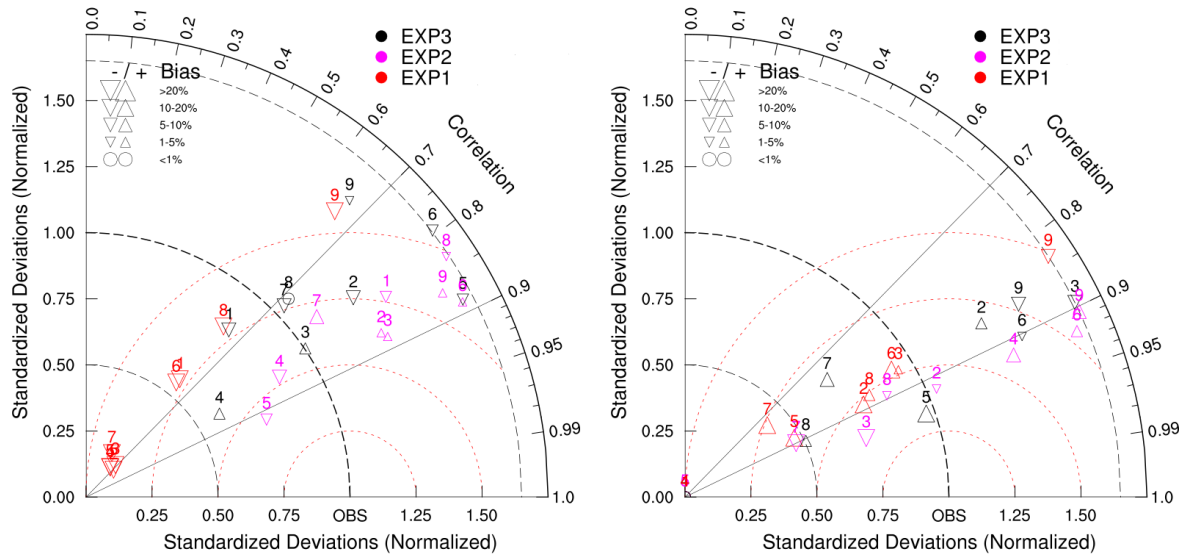


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1276 Figure 1 The MODIS reflectance (true color) map representing low cloud over Indo Gangetic Plains,
 1277 India (study region) indicative of likely fog and haze on 23rd December (a) and 24th December (b)
 1278 2017. (c) Anthropogenic emission of PM_{2.5} over IGP for December 2017 obtained from EDGAR-
 1279 HTAP. The boxes represent the regions Northwest IGP (NWIGP), Central IGP (CIGP), and East IGP
 1280 (EIGP).

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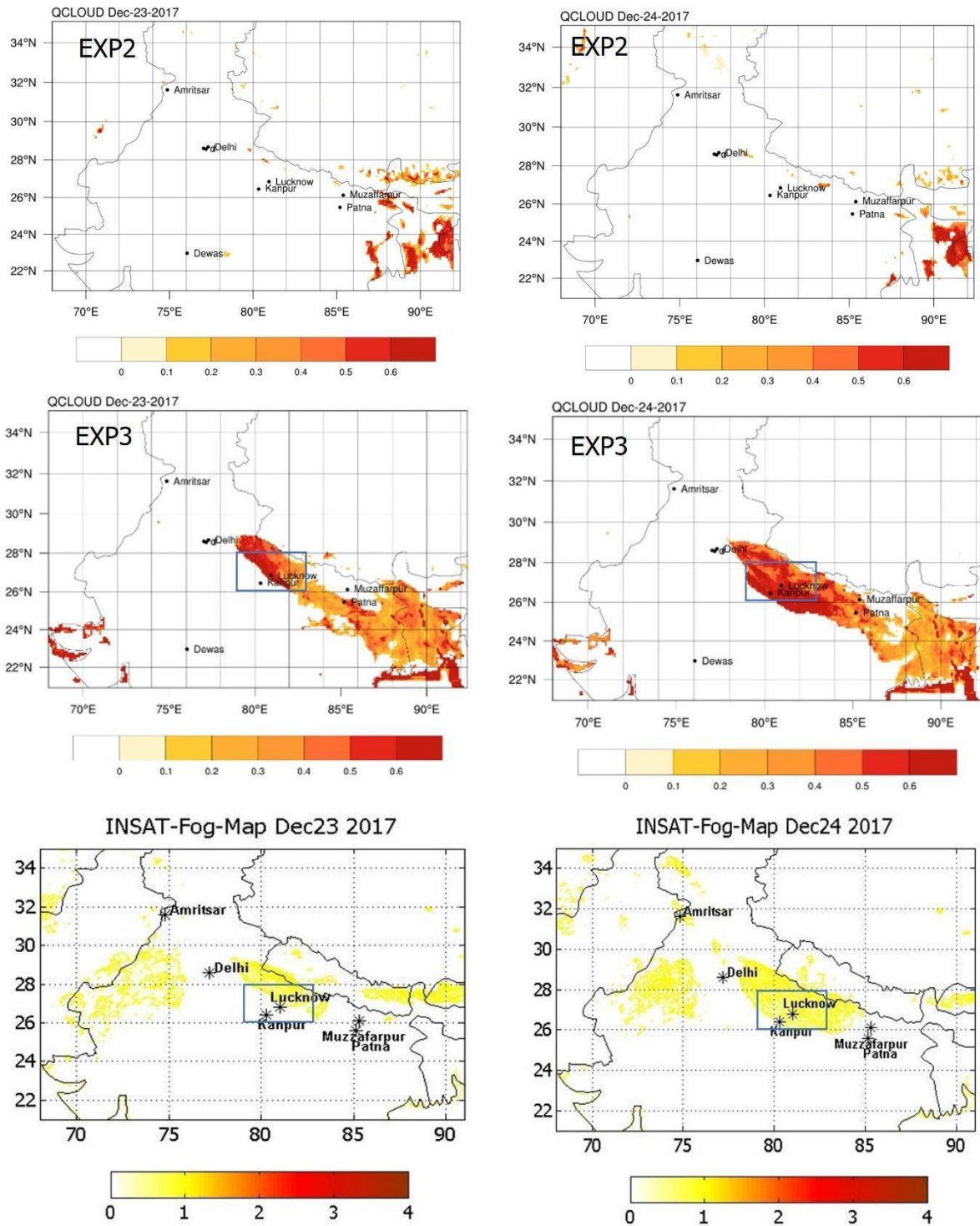
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1284 Figure 2 Taylor Diagram of simulated (WRF-Chem) and observed (CPCB) relative humidity (left) and
 1285 2-m temperature (right) over IGP. The colors indicate the experiments. The red dotted contours
 1286 represent RMS values. The marker (triangles) size varies with a mean bias between the experiments
 1287 and observation. Upside-down triangles represent positive bias (exp-obs) and vice versa. The stations
 1288 over IGP are denoted by number 1. Amritsar, 2. IGI Airport (Delhi), 3. IHBAS (Delhi), 4. Dwarka
 1289 (Delhi), 5. RKP (Delhi), 6. Kanpur, 7. Lucknow, 8. Patna, 9. Muzaffarpur. The locations are marked
 1290 in Fig.1a.

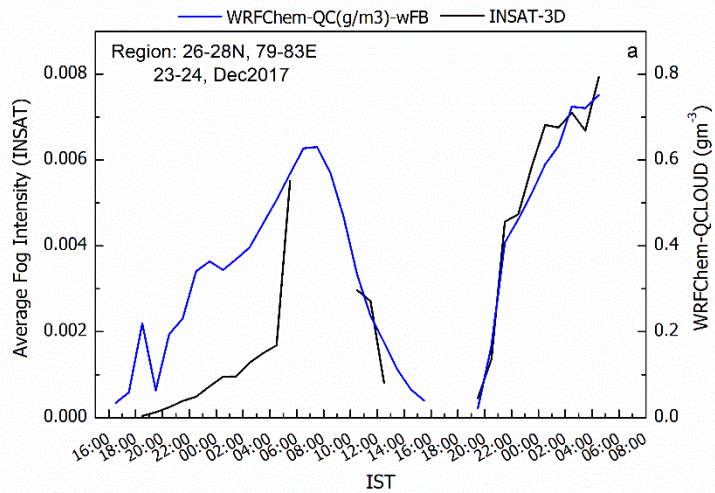
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1293 Figure 3 Average WRF-Chem surface layer cloud water mixing ratios (g m^{-3}) for EXP2 and EXP3
 1294 (top four panels) and INSAT-3D satellite fog intensity (bottom two panels) on 23 and 24 December
 1295 2017. INSAT-3D satellite fog intensity varies from 0 to 4 indicating SHALLOW, MODERATE,
 1296 DENSE, and VERY_DENSE, respectively. The rectangle in central IGP is the region for the time
 1297 series analysis.

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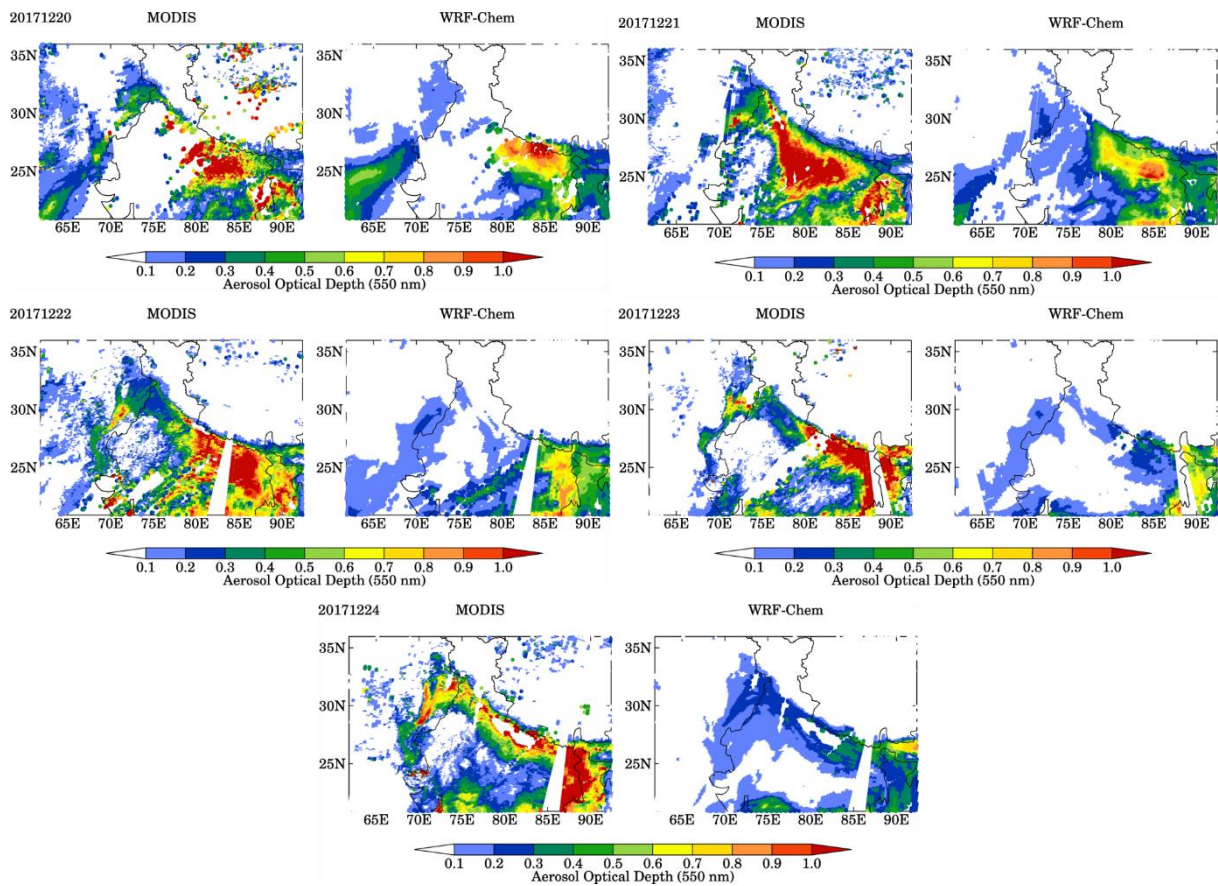


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1300 Figure 4 Average Hourly variation of fog on 23 and 24 December 2017 from WRF-Chem EXP3
 1301 simulation and INSAT-3D satellite between 26°N-28°N,79°E-83°E (region shown in Fig 3). The time
 1302 is in IST (Indian Standard Time; IST is 5.5 hours ahead of Universal Time Coordinate (UTC).

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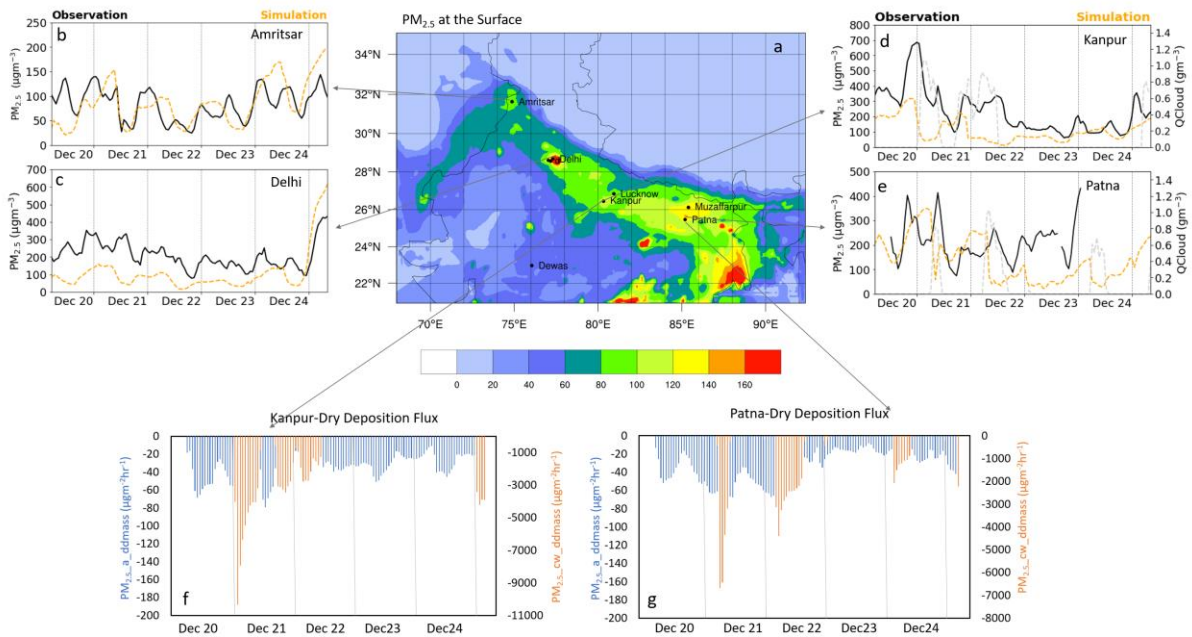
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1306 Figure 5 Comparison of WRF-Chem AOD with MODIS observation over the model domain on 20,
 1307 21, 22, 23, and 24 December 2017.

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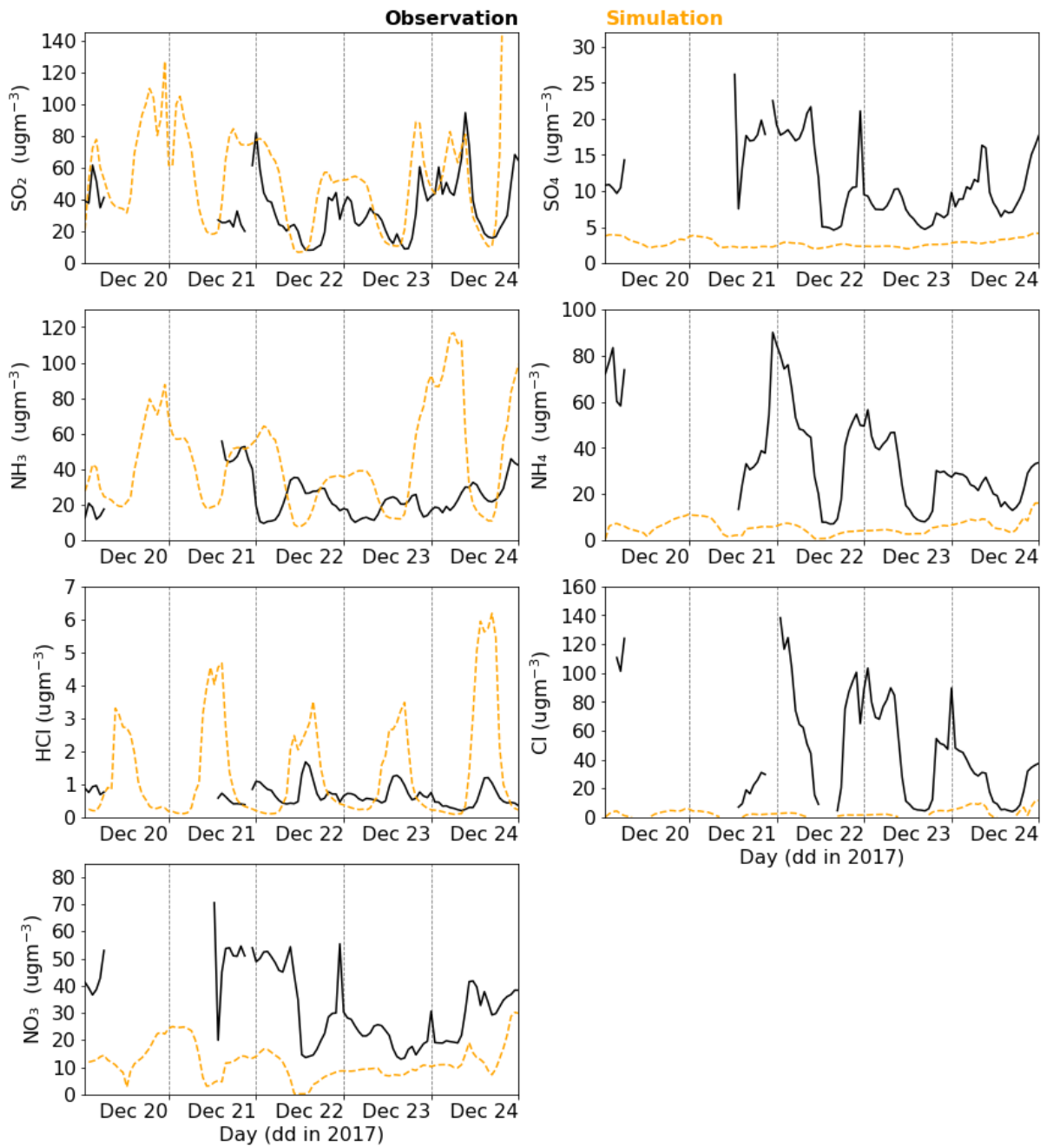
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1312 Figure 6 WRF-Chem simulated surface PM_{2.5} map over IGP (a); comparison of WRF-Chem PM_{2.5}
1313 with CPCB observation for the period 20-24 Dec 2017 for (b) Amritsar, (c) Delhi, (d) Kanpur and (e)
1314 Patna. Dry Deposition rate of PM_{2.5} for (f) Kanpur and (g) Patna. The grey dotted line in (d) Kanpur
1315 and (e) Patna is fog (QCloud) present during the study period.

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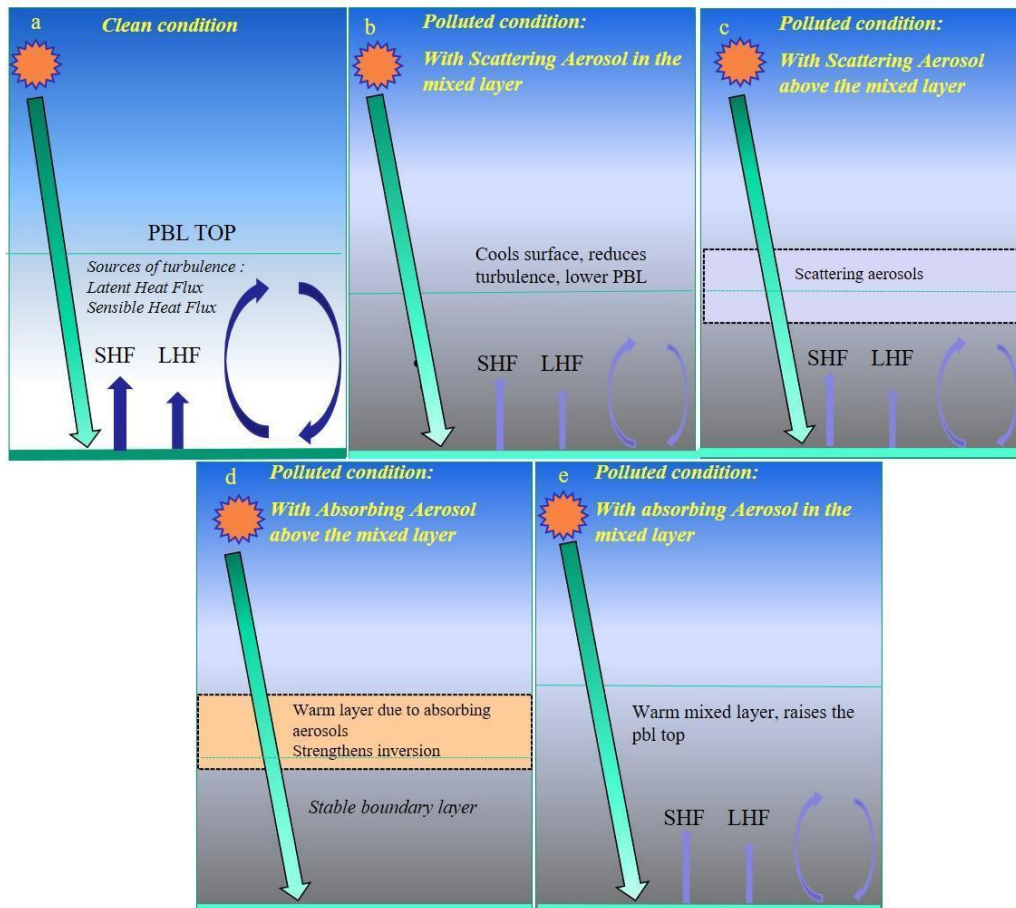
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1318 Figure 7 Comparison of WRF-Chem simulated ions (SO_4^{2-} , NH_4^+ , NO_3^- , Cl^-) and trace gases (SO_2 ,
 1319 NH_3 & HCl) with the observation from WIFEX campaign at Delhi.

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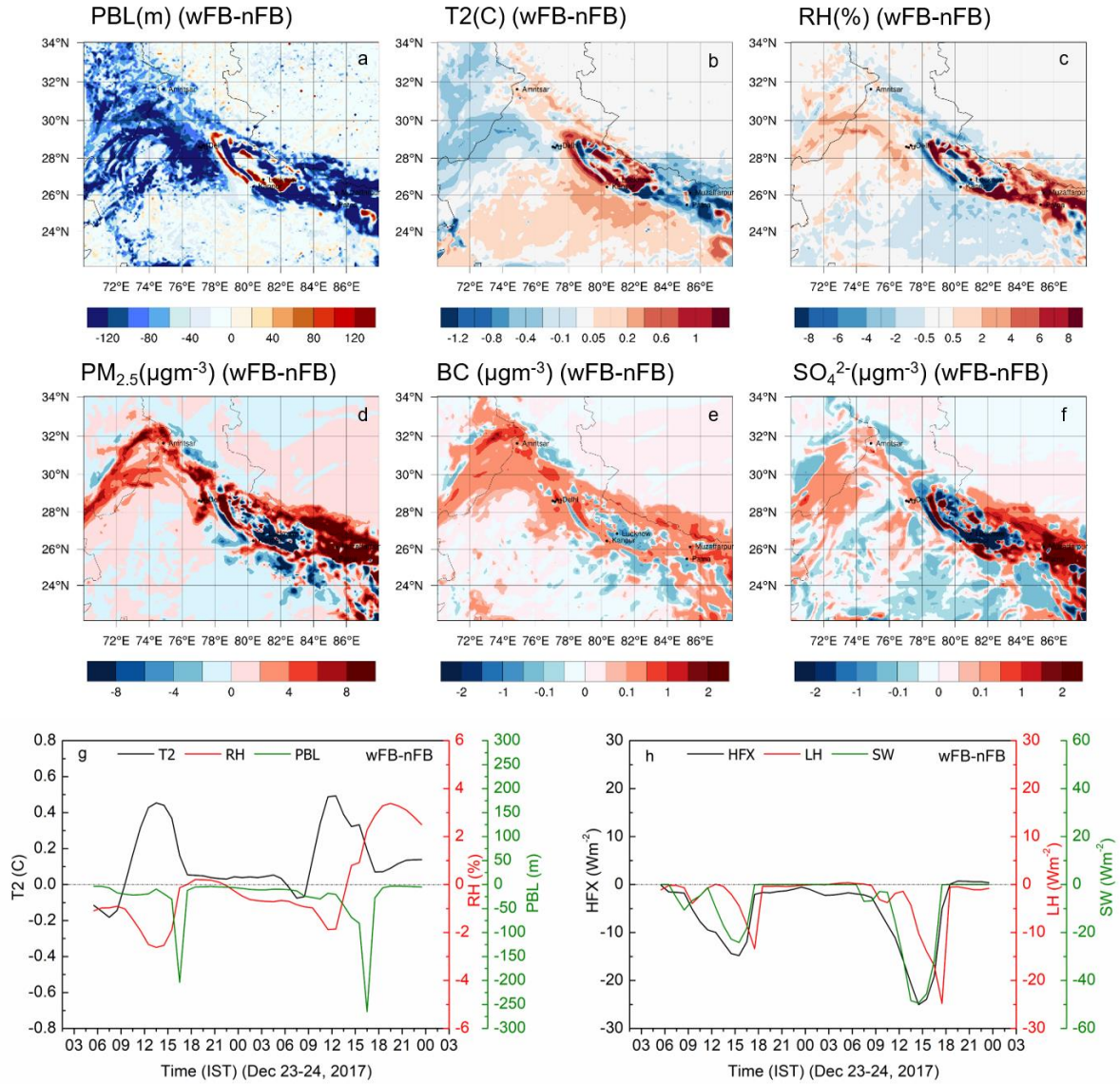
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1324 Figure 8 Schematic diagram of Aerosol Radiation Feedback.

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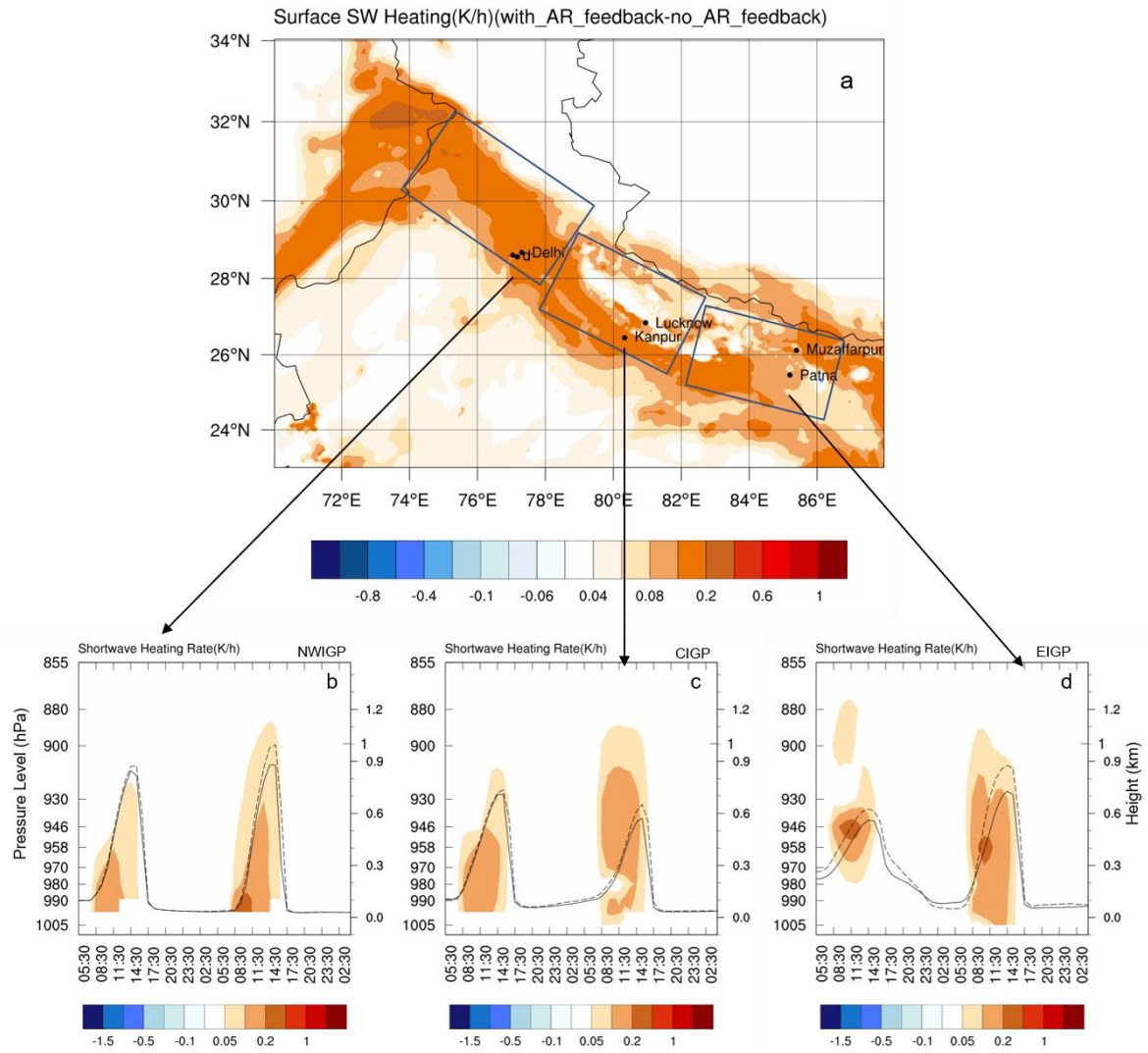
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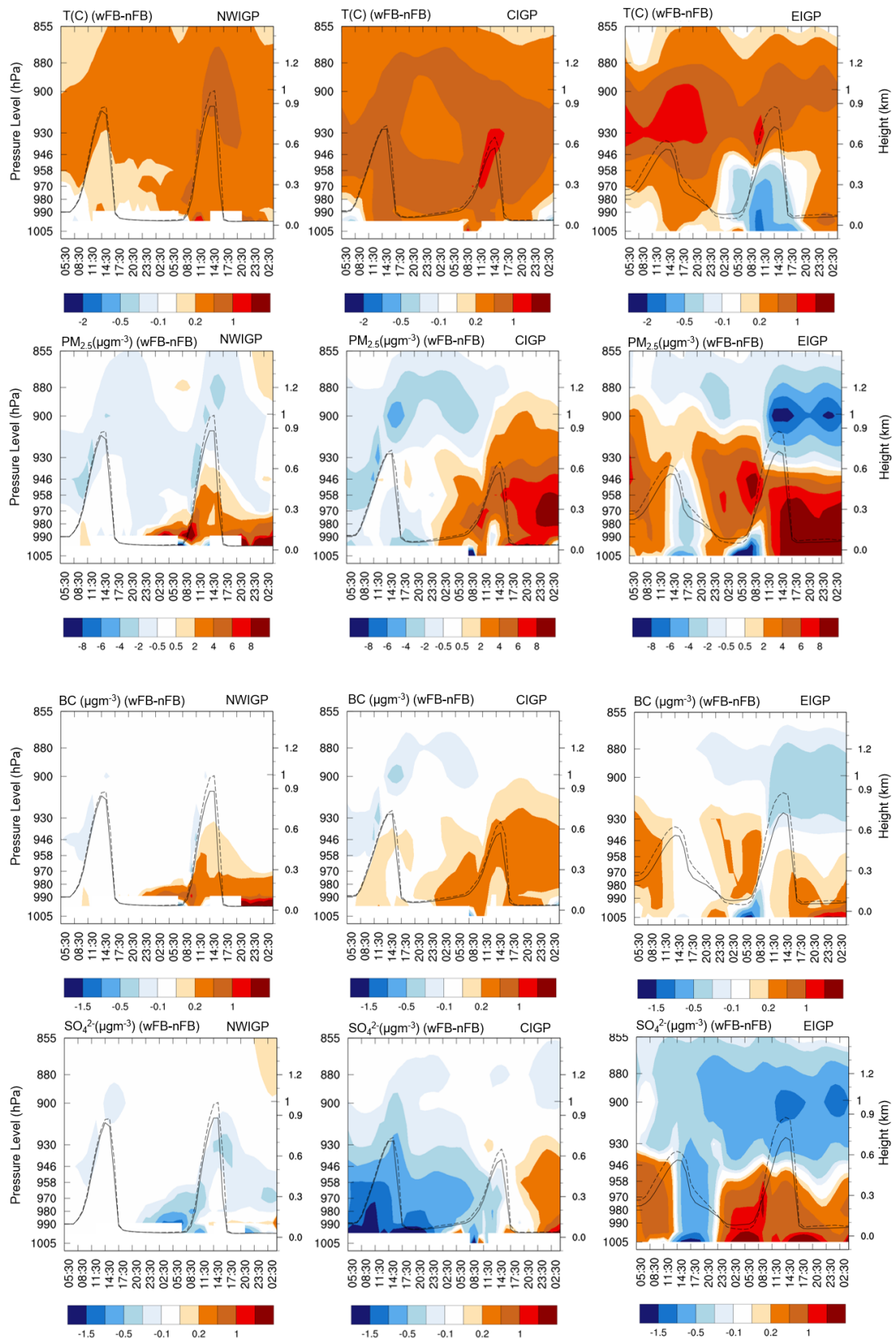
1329 Figure 9 Effect of Aerosol Radiation feedback (wFB-nFB) on (a) PBL height, (b) 2-m temperature, (c)
 1330 2-m relative humidity, (d) surface PM_{2.5}, (e) surface BC and (f) surface SO₄ for December 24 at local
 1331 noon (13:30-15:30 IST). (g) The time series of ΔPBL, ΔT2, and ΔRH; (h) ΔHFX (sensible heat flux),
 1332 ΔLH (latent heat flux), and ΔSWF (downward shortwave flux) over CIGP for December 23 and 24. Δ
 1333 denotes the difference between with and without AR feedback (wFB-nFB).

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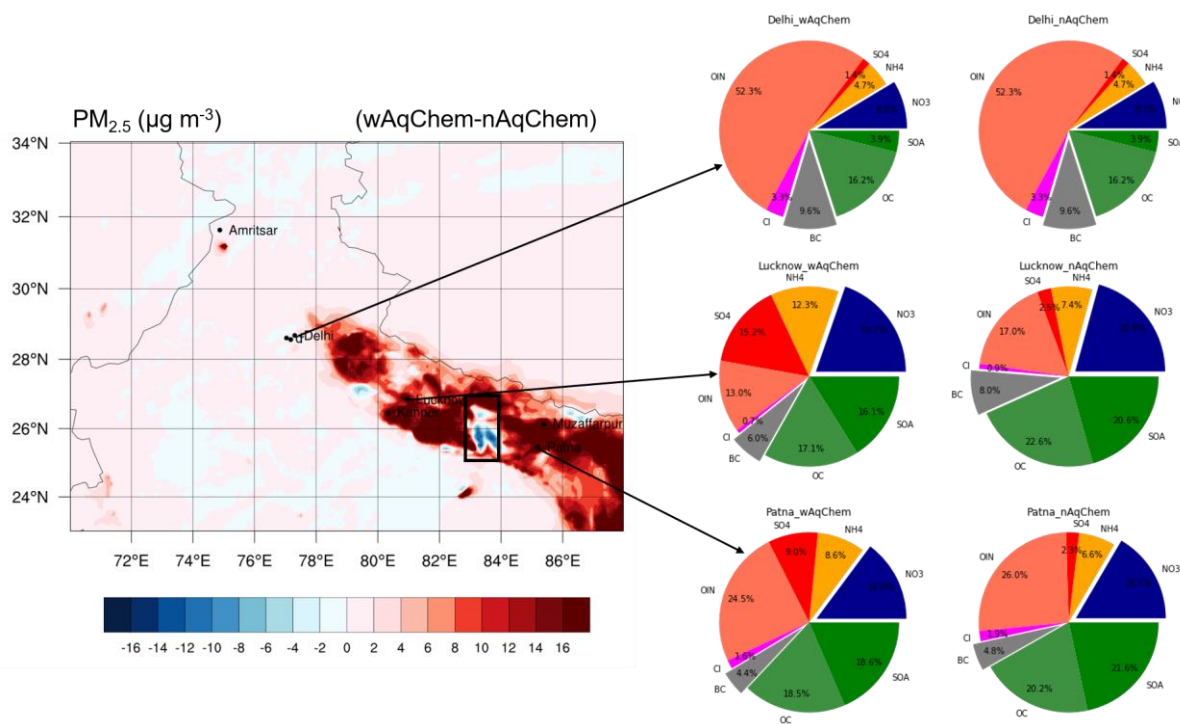
1337 Figure 10 Differences in shortwave heating rates ($K h^{-1}$) between simulations with and without aerosol
 1338 radiation feedback (a) at the surface, and for pressure-time cross-sections over (b) NWIGP, (c) CIGP.
 1339 And (d) EIGP for December 23 and 24. The solid and dashed lines are the PBL height with and
 1340 without AR feedback respectively. The time is in IST.



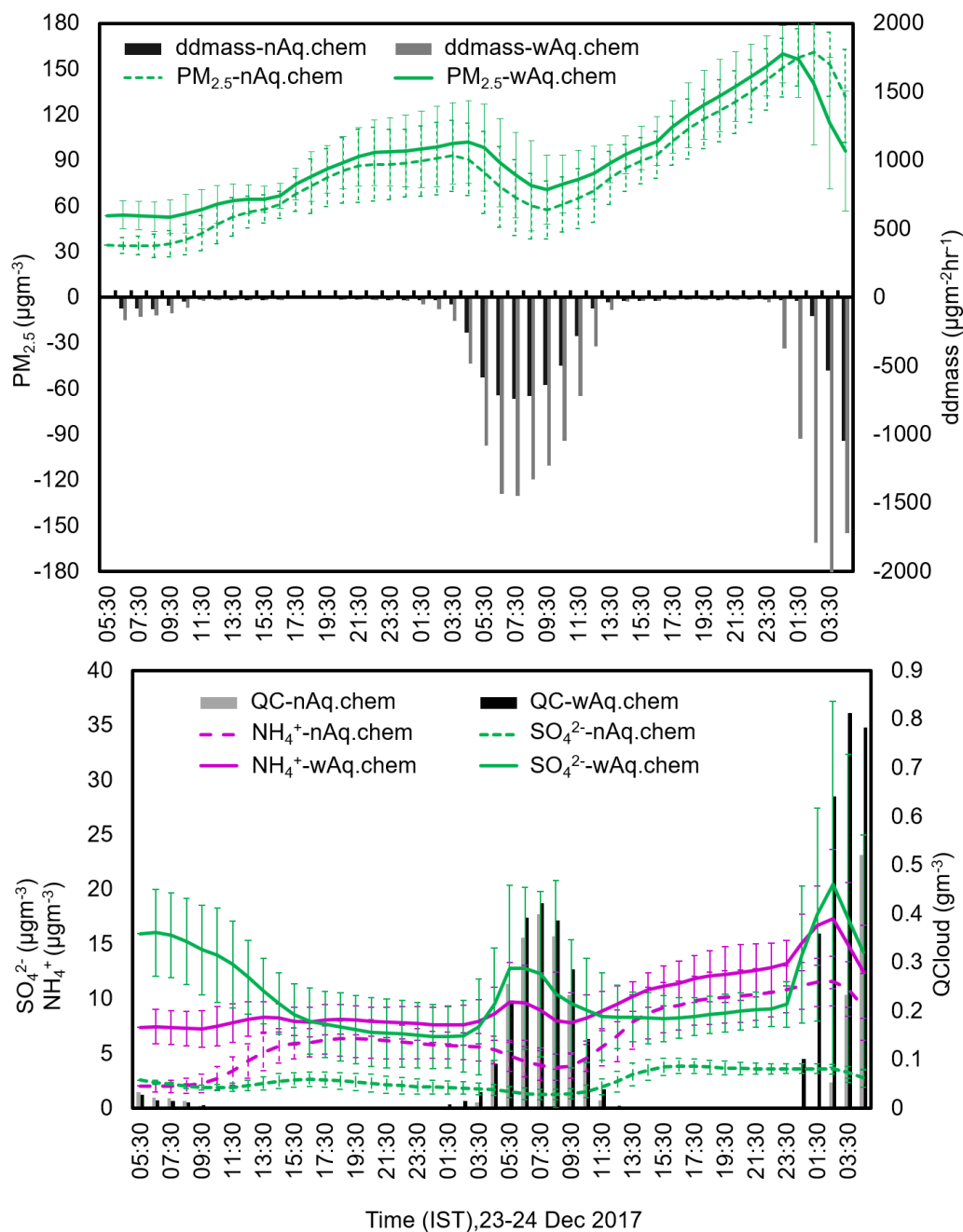
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 1344 Figure 11 Pressure-time cross-section of the differences in T, PM_{2.5}, BC and SO₄²⁻ between
 1345 simulations with and without the AR feedback for December 23 and 24. The solid and dashed lines
 1346 are the PBL height with and without AR feedback respectively. The time is in IST.
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 1350 Figure 12 (a) Surface ΔPM_{2.5} (wAq.chem-noAq.chem) and (b) pie charts of PM_{2.5} composition
 1351 distribution for the two cases, with and without Aqueous phase Chemistry for 24 Dec 2017. The
 1352 stations Delhi, Lucknow (LKN), and Patna are representative of NWIGP, CIGP, and EIGP regions
 1353 respectively.
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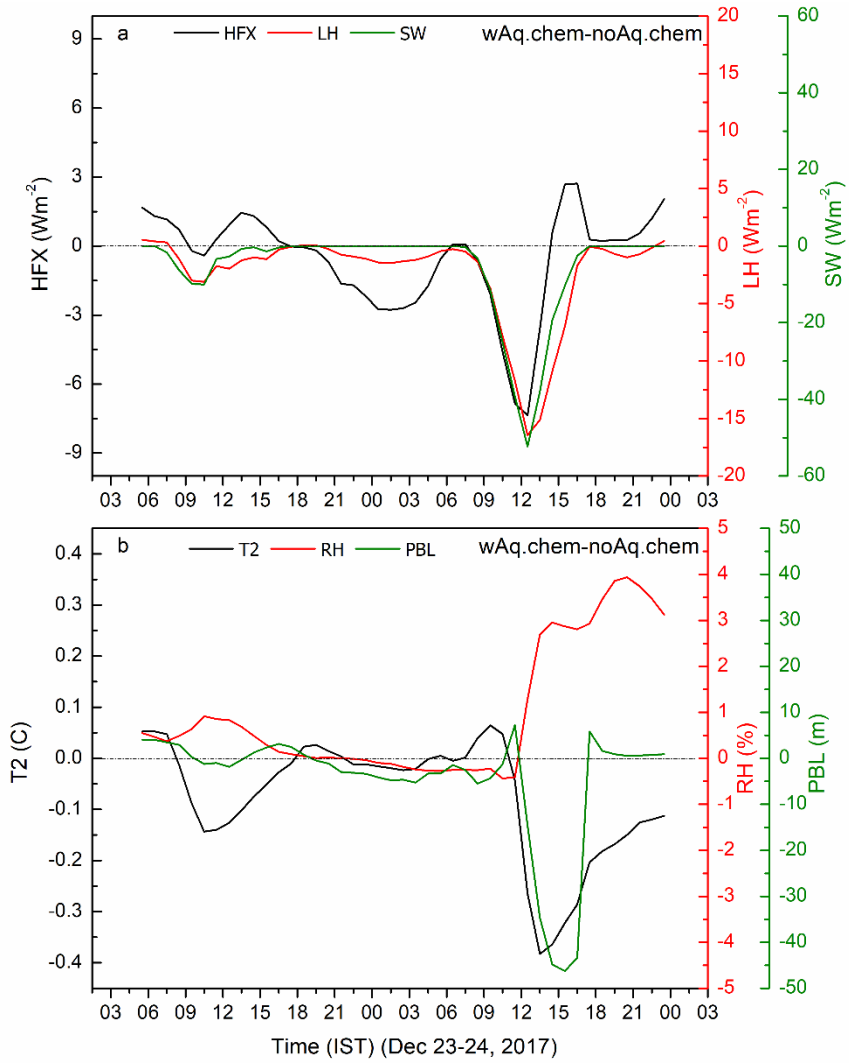


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1358 Figure 13 Time series of (a) PM_{2.5} and its dry deposition (ddmass) flux change, (b) SO₄²⁻, NH₄⁺ and
 1359 LWC (QCloud) with and without aqueous phase chemistry included in the model, averaged over the
 1360 region bounded by a black rectangle in Fig. 12, for 23 and 24 December, 2017.

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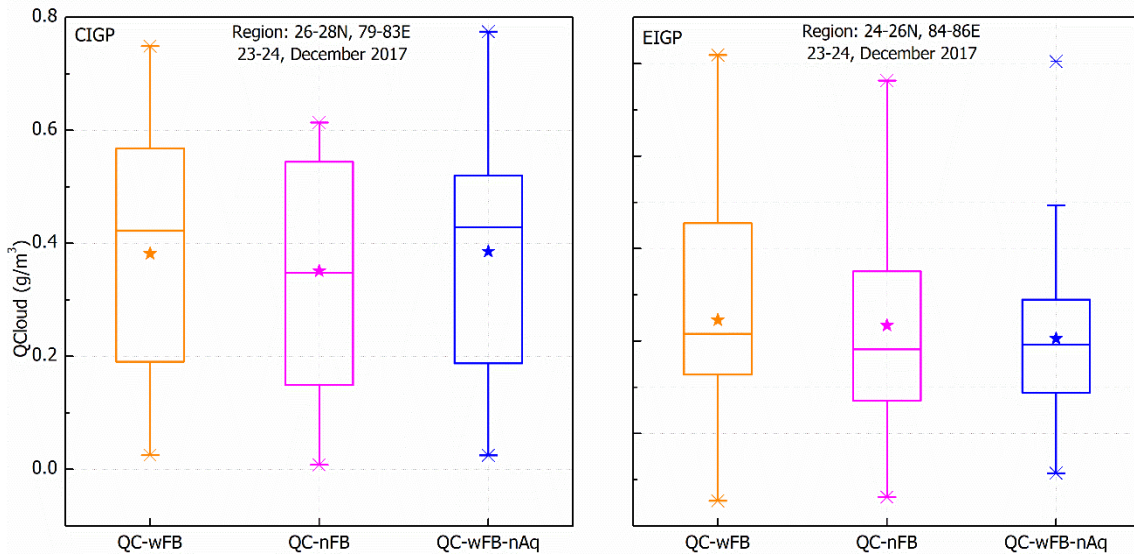
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1365 Figure 14 Time series of (a) Δ HFX (sensible heat flux), Δ LH (latent heat flux), and Δ SWF
 1366 (downward shortwave flux); (b) Δ T2, Δ RH, and Δ PBL over CIGP (79E-83E,26N-28N), for
 1367 23 and 24 December, 2017. Δ denotes the difference between with and without aqueous phase
 1368 chemistry.

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1373 Figure 15 Averages (stars), medians (horizontal lines), quartiles (boxes), maxima, and minima for
1374 LWC (QCloud) averaged over CIGP (left panel) and EIGP (right panel) for the fog event on 23-24
1375 December 2017. Gold is for the simulation with AR feedback and aqueous chemistry, magenta for the
1376 simulation with no AR feedback but includes aqueous chemistry, and blue for the simulation with AR
1377 feedback but no aqueous chemistry. WRF-Chem does not produce fog in the NWIGP during the study
1378 period.

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Table 2: Table showing the start and end time of fog1 on 23-24 December 2017 with LWC for the sensitivity experiments, with AR feedback, no AR feedback and no Aqueous phase chemistry

Fog 1 (December 23-24, 2017)											
	EXP-wFB			EXP-nFB			EXP-nAq.Chem				
	Start time (IST)	End time (IST)	Duration of Fog	Start time (IST)	End time (IST)	Duration of Fog	Start time (IST)	End time (IST)	Duration of Fog		
CIGP	16:30	15:30	23h	18:30	17:30	23h	18:30	15:30	21h		
LWC (g m ⁻³)	0.034	0.036±0.032		0.141±0.154	0.068±0.005		0.184±0.138	0.034 ±0.021			
Kanpur	05:30	13:30	8h	05:30	12:30	7h	05:30	12:30	7h		
LWC (g m ⁻³)	0.334±0.487	0.017		0.458±0.357	0.173±0.071		0.533	0.025±0.0123			
Lucknow	23:30	14:30	15h	00:30	14:30	14h	23:30	14:30	15h		
LWC (g m ⁻³)	0.269±0.145	0.087±0.040		0.232±0.132	0.029±0.024		0.139±0.084	0.025±0.012			
EIGP	21:30	12:30	15h	23:30	10:30	11h	21:30	10:30	13h		
LWC (g m ⁻³)	0.099±0.092	0.007		0.198±0.188	0.084±0.060		0.026±0.008	0.153±0.119			
Patna	00:30	12:30	12h	04:30	10:30	6h	02:30	10:30	8h		
LWC (g m ⁻³)	0.100±0.090	0.007		0.009±0.005	0.038±0.041		0.196±0.198	0.166±0.130			
Muzzafarpur	05:30	11:30	6h	06:30	10:30	4h	06:30	09:30	3h		
LWC (g m ⁻³)	0.112±0.146	0.043±0.057		0.051±0.041	0.003		0.142±0.151	0.157±0.064			

1396 Table 3: Table showing the start time of fog 2 on 24 December 2017 with LWC for the sensitivity
 1397 experiments, with AR feedback, no AR feedback and no Aqueous phase chemistry. Fog2 end time
 1398 could not be noted as simulation ended on 25 December 2017, 00UT (5:30 IST) before fog2
 1399 dissipates.

Fog 2 (December 24, 2017)			
Start time			
(IST)			
	EXP-wFB	EXP-nFB	EXP-nAq.Chem
CIGP	19:30	20:30	21:30
LWC (g m ⁻³)	0.025	0.008±0.007	0.025
Kanpur	21:30	22:30	23:30
LWC (g m ⁻³)	0.041±0.007	0.298±0.218	0.482±0.398
Lucknow	21:30	20:30	00:30
LWC (g m ⁻³)	0.203±0.165	0.005	0.229±0.209
EIGP	00:30	01:30	01:30
LWC (g m ⁻³)	0.024±0.030	0.072±0.088	0.014±0.009
Patna	03:30	03:30	03:30
LWC (g m ⁻³)	0.030±0.046	0.018	0.060
Muzzafarpur	04:30	No fog	No fog
LWC (g m ⁻³)	0.159±0.038		

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