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Aggregation of Slightly Buoyant Microplastics in Three-Dimensional Vortex Flows

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Abstract

Although the movement and aggregation of microplastics at the ocean surface has been well studied, less is known about the subsurface. Within the Maxey-Riley framework governing the movement of small spheres with high drag in fluid, aggregation of buoyant particles is encouraged in vorticity-dominated regions. We explore this process in an idealized model of a three-dimensional eddy with an azimuthal and overturning circulation. In the axially symmetric state, particles that do not accumulate at the top boundary are attracted to a closed contour consisting of periodic orbits. Such a contour exists when drag on the particle is sufficiently strong. For small slightly-buoyant particles, this contour is located close to the periodic fluid trajectory. If the symmetric flow is perturbed by a symmetry-breaking disturbance, additional attractors arise near periodic orbits of fluid particles within the resonance zones created by the disturbance. Disturbances with periodic time dependence produce even more attractors, with a shape and location that recurs periodically, and which are composed of quasiperiodic orbits of rigid particles. Not all such contours attract, and particles released in the vicinity may instead be attracted to a nearby attractor. Examples are presented along with mappings of the respective basins of attraction.

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Significance statement

This paper investigates the phenomenon of aggregation of small, slightly buoyant, rigid body particles in a three-dimensional vortex flow. Our goal was to gain insights into the behaviour of slightly buoyant marine microplastic particles in a flow that qualitatively resembles ocean eddies. Attractors are mapped out for the steady axisymmetric, steady asymmetric, and non-steady asymmetric vortices over a range of flow and particle parameters. Simple theoretical arguments are used to interpret the results.

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I. Introduction

Marine microplastic pollution has been a rising concern for the ocean environmental and for human health. Microplastics (scales < 5mm) and nanoplastics (scales < 1 μm) have been found in the tissues of marine animals, some of which are consumed by humans (Landrigan, et al. 2023). This comes at a time when global production of plastics is projected to increase. Most observations of marine microplastics have occurred at or near the sea surface, where concentrations are largest. However, the density of many types of particles, including high-density polyethylene, is sufficiently close to that of sea water that suspension within the water column for long periods of time is feasible. Indeed, microplastics have been found well beneath the ocean surface, but less is known regarding their spatio-temporal distributions (Shamskhany et al. 2021).

A potentially important aspect of the movement of plastics and microplastics is aggregation, a process that occurs at the surface over large scales near the centers of the five major subtropical gyres and has been attributed to Ekman drift, windage and inertia (Beron-Vera, 2021). Many early models concentrated on the ocean surface, but Froyland et al. (2014) has highlighted the importance of resolving the full three dimensional circulation. If aggregation also occurs below the surface, well beneath the direct influence of Ekman layers, the dynamics is likely to be different. Indeed, modeling results by Wichmann et al. (2019), and based on a framework created by Lange and van Sebille (2017) and Delandmeter and van Sebille (2019), suggests that the large scale accumulation associated with the garbage patches disappears below 60m depth.

Typically the position *X(t)* of a non-fluid particle is tracked according to

$$X(t + \Delta t) = X(t) + \int_t^{t+\Delta t} v dt + dX_b,$$

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where v is the fluid velocity and dX_b is an extra displacement due the non-fluid nature of the particle. The user can introduce custom schemes for calculating contributions to dX_b due to factors such as windage and inertia (e.g. Beron-Vera et al. 2016), turbulent diffusion (e.g. Kulkulka, 2012), wave induced Stokes drift (Onink et al. 2019), etc. Eulerian schemes in which plastic particles are treated as concentrations, are rare, but Mountford and Morales Maqueda (2019) developed an Eulerian model in which concentrations are advected by the fluid and are subject to parameterized turbulence as well as sinking or rising according to a simple law involving buoyancy and friction.

An alternative approach would be to use the Maxey-Riley equation (discussed below) to solve for the particle velocity v in the above equation, then use the latter to compute the trajectory of that parcel. This equation would account for effects such as inertia and added mass in a deductive way, however the resulting 6^{th}-order system (for the three components of velocity and position) would be computationally challenging. To better understand the implications of the use of this approach while avoiding the computational burden and complexity, we have elected to analyze the movement and aggregation of individual particles using a Maxey-Riley framework in connection with an idealized, 3D flow field resembling the circulation in an ocean eddy. The aim is to develop a basic understanding of the circumstances that would lead to aggregation of rigid particles in ocean mesoscale and submesoscale eddies. We note that other idealized studies have been carried out in connection with 2D wave fields and vortex flows (e.g. DiBenedetto 2018a,b and Kelly et al., 2021).

Aggregation can be attributed to the presence of an attractor: here, an object with a dimension less than three that is somehow set up by the fluid circulation patterns and towards which rigid particle trajectories attract. As long as the fluid is incompressible, fluid parcels will not

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experience attraction and will not aggregate, but plastic particles with inertia, added mass, and drag may do so. In order to reach a better understanding of what leads to attraction and attractors in 3D flows, we explore a canonical example in geophysical fluid dynamics, namely the flow in a rotating cylinder. This flow has some of the characteristics of ocean eddies, including a horizontal swirl and an overturning component in the vertical. The Lagrangian properties of this circulation have been previously studied (Fountain, et al. 2000; Pratt et al. 2014; Rypina et al. 2015) allowing us to begin to investigate inertial particles from an established base of knowledge. A prior theory (Haller and Sapsis, 2008) governing the movement of particles with high drag indicates that accumulation is favored for slightly buoyant particles in flows dominated by vorticity, and this also motivates our choice of background flow. Identification of the attractors that can arise in this flow field, evaluating their reach and domains of attraction, and clarifying the circumstances that lead to their formation are the primary objectives of this work.

II. Methods

The physics of the motion of a small, rigid sphere that moves with velocity $\vec{v}(t)$ through a fluid with pre-existing velocity distribution $\vec{u}(\vec{x},t)$ has been the subject of investigation by Stokes (1851), Basset (1888), Boussinesq (1903), Faxen (1922), Oseen (1927), Tchen (1947) and many others, and was put in a unifying framework by Maxey and Riley (1983). More recent theoretical extensions include Beron-Vera et al. (2019) and Beron-Vera (2021). We will use a form of the Maxey-Riley equation that has been extended to include constant frame rotation with angular velocity $\vec{\Omega}^*$:

$$\frac{d\vec{v}}{dt} = \frac{\rho_f}{\rho_p}\frac{D\vec{u}}{Dt} + \frac{\rho_f}{2\rho_p}\left(\frac{D\vec{u}}{Dt} - \frac{d\vec{v}}{dt}\right) - \frac{9\nu\rho_f}{2\rho_p d^2}(\vec{v}-\vec{u}) + \left(1-\frac{\rho_f}{\rho_p}\right)\vec{g} + \frac{\rho_f}{\rho_p}\vec{\Omega}^* \times (\vec{u}-\vec{v}) \quad 123$$

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$$+\frac{\rho_f}{\rho_p}2\vec{\Omega}^* \times \vec{u} - 2\vec{\Omega}^* \times \vec{v} + \left(\frac{\rho_f}{\rho_p} - 1\right)\vec{\Omega}^* \times \vec{\Omega}^* \times r\,. \qquad (1)$$

In this statement of Newton's second law for the rigid particle, the right-hand side represents, in order, the effects of inertia, added mass, drag, buoyancy, Coriolis acceleration associated with the added mass, the Coriolis acceleration associated with the particle mass, Coriolis acceleration associated with the fluid motion, and centripetal acceleration. (See Beron-Vera, et al. 2019 for a derivation, though the centripetal acceleration appears to have been omitted.) We have omitted the lift force, the Basset history force, and the Faxen corrections (Gatignol,1983). Here ρ_p and ρ_f are densities of the particle and the fluid, d is the particle radius, ν is viscosity of the fluid, $\vec{g}$ is the gravity vector, and $\frac{D\vec{u}}{Dt} = \frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla\vec{u}$ is the fluid material derivative, evaluated for undisturbed fluid velocity at the position of the center of the particle. The position $x_p(t)$ of a particle is determined by

$$\frac{dx_p}{dt} = \vec{v}, \qquad (2)$$

and together (1) and (2) compose a coupled, 6th-order system for computation of the particle position and velocity as functions of time.

If the velocities and lengths are nondimensionalized using characteristic scales U and L for the background fluid flow, and L/U is used as a time scale, then (2) remains formally unchanged while the nondimensional form of (1) is

$$\frac{d\vec{v}}{dt} = \frac{3R}{2}\frac{D\vec{u}}{Dt} + \tilde{\varepsilon}^{-1}(\vec{v} - \vec{u}) + \left(1 - \frac{3R}{2}\right)\vec{g}_r + 3R\vec{\Omega} \times (\vec{u} - \vec{v}) + 2\left(\frac{3R}{2} - 1\right)\vec{\Omega} \times \vec{v}, \qquad (3)$$

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where $R = \frac{2\rho_f}{\rho_f + 2\rho_p}$, $\vec{g}_r = (g - \vec{\Omega}^* \times \vec{\Omega}^* \times \vec{r})/(\frac{U^2}{L})$, $\vec{\Omega} = \frac{\vec{\Omega}^* L}{U}$ and $\tilde{\varepsilon} = \frac{2}{9}\left(\frac{d}{L}\right)^2 \frac{UL}{\nu}\frac{1}{\mathrm{R}}$ is the Stokes number, the ratio of the adjustment time scale of a particle (due to drag) to the time scale of the background flow. For $\tilde{\varepsilon} \ll 1$, viscous drag is the dominant force acting on the particle, implying that a particle with an initial velocity differing by an amount > $O(\tilde{\varepsilon})$ from the local fluid velocity will be rapidly accelerated over a time scale $\tilde{\varepsilon}$ to a velocity proximal to that of the fluid. Thereafter the particle will undergo a slow evolution in which the weaker forces due to inertia, added mass, and buoyancy cause slight departures from the movement of the fluid itself.

The limit $\tilde{\varepsilon} \to 0$ constitutes a singular perturbation of (3), a problem that can be addressed using an approach due to Fenichel (1979) that was originally formally developed for a steady background flow, but that has been extended by Haller and Sapsis (2008) to include a time-varying background flow. In either case, it can be shown that following the initial viscous adjustment, the particle position and velocity tend toward a subspace or "slow manifold" on which the particle velocity is determined directly by the fluid velocity through an "inertial" equation, here extended to include frame rotation:

$$\vec{v} = \vec{u} + \tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right)\left[\frac{D\vec{u}}{Dt} + 2\vec{\Omega} \times \vec{u} - \vec{g}_r\right] + O(\tilde{\varepsilon}^{\,2}). \qquad (4)$$

This version with rotation was written down in Beron-Vera et al. 2019, though with $\vec{g}_r$ replaced by the non-generalized gravity vector $\vec{g}$. The same authors also present more general cases, including those with the lift force. In Supplementary Material we present a simple, alternative derivation of Eq. (4) based on a multiple-scale expansion instead of the Fennochel approach.

A chief advantage of the slow manifold reduction is that the 6$^{\text{th}}$ order system (2) and (3) is reduced to a 3$^{\text{rd}}$ order system (2) and (4) in which the particle velocity is known in advance.

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The bracketed expression in (4), which determines the velocity of the rigid particle relative to the fluid, is nothing more than $\frac{\partial}{\partial x_j}\tau_{ij}$, where τ_{ij} is the stress tensor for the fluid. Thus the relative velocity of a rigid particle on the slow manifold is in the same direction as the net force that would act on a fluid particle occupying the same space. Ordinarily, for a fluid particle, that force would equate with an acceleration, but on the slow time scale, the relative particle velocity points in the same direction as the net fluid force and its magnitude is proportional to $\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)$ $=\frac{2}{9}\frac{d^2}{L^2}\frac{UL}{\nu}\frac{(\rho_f-\rho_p)}{\rho_f}$. Since the aggregation of rigid particles requires departures of the particle velocity from the (divergence free) velocity field of the fluid, one can expect that aggregation will occur more slowly if d and $(\rho_f-\rho_p)/\rho_f$ are small, or if ν is large. At the same time, the existence of attractors internal to the fluid may depend on $(\rho_f-\rho_p)/\rho_f$ being small: for example, a large density difference may mean that rigid particles simply sink to the bottom or rise to the surface.

As pointed out by Haller and Sapsis (2008) (also see Beron-Vera et al. 2019), we can consider a continuous concentration of rigid particles with the like properties, and with smoothly varying velocity (4). The aggregation of such a concentration would appear to require that the divergence of that velocity be negative (though see an apparent counterexample, presented later). Following Haller and Sapsis (2008), consider the evolution of a material volume of rigid particles. The time rate of change of this volume is

$$\frac{dV}{dt} = \oiint \vec{v}\cdot\vec{n}\,dA_V = \iiint(\nabla\cdot\vec{v})\mathrm{d}V = \iiint \nabla\cdot\left[\vec{u}+\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)\left(\frac{D\vec{u}}{Dt}+2\vec{\Omega}\times\vec{u}-\vec{g}_r\right)\right]\mathrm{d}V \qquad (5)$$

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where $\nabla \cdot \vec{u} = 0$ for an incompressible fluid. Shrinking V to an infinitesimal size allows the right-hand side to be approximated by V times the local value in the integrand, and the result may be integrated in time, yielding

$$V(t) = V_0 \exp\left(\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) \int_{t_0}^{t} \nabla \cdot \left(\frac{D\vec{u}}{Dt} + 2\vec{\Omega} \times \vec{u} - \vec{g}_r\right) ds\right)$$

$$= V_0 exp\left(-2\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) \int_{t_0}^{t} [Q_r(x(s), s) + \vec{\Omega} \cdot \vec{\zeta}_r + \left|\vec{\Omega}\right|^2]\, ds\right)$$

$$= V_0 exp\left(-2\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) \int_{t_0}^{t} Q_a(x(s), s) ds\right). \qquad (6)$$

Here $Q_r = \frac{1}{2}\left(\left|\vec{\zeta}_r\right|^2 - |S|^2\right)$ is the three-dimensional Okubo-Weiss parameter (Okubo, 1970; Weiss, 1991), $\vec{\zeta}_r$ represents the relative vorticity vector for the fluid, $S = 1/2(\nabla\vec{u} + (\nabla\vec{u})^T)$ is the strain tensor, and $|S|$ is its Frobenius norm. The final step in (6) follows from introduction of the absolute vorticity vector

$$\vec{\zeta}_a = \vec{\zeta}_r + \overline{2\Omega} \qquad (7)$$

and the corresponding function $Q_a = \frac{1}{2}\left(\left|\vec{\zeta}_a\right|^2 - |S|^2\right)$. We note that for a volume V of any size:

$$\frac{dV}{dt} = 2\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) \iiint Q_a\, \mathrm{dV} = \tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) \iiint \frac{\partial^2}{\partial x_i \partial x_j} \tau_{ij}\, \mathrm{dV} = \frac{2}{9} \frac{d^2}{L^2} \frac{UL}{\nu} \frac{(\rho_f - \rho_p)}{\rho_f} \oiint \frac{\partial}{\partial x_j} \tau_{ij} n_i\, dA_V,$$

(8)

where n_j denote the components of the outward unit vector normal to the bounding surface A_V. Thus for buoyant particles, a volume $V(t)$ of any size will contract if the force normal to A_V due to the fluid stresses, integrated around A_V, is inward. In many cases the stress tensor is

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dominated by pressure, i.e., $\frac{\partial}{\partial x_j}\tau_{ij} \cong -\frac{1}{\rho_f}\nabla p$, so the tendency to aggregate is determined entirely by the pressure field.

In general, Q_a can change sign along a particle trajectory, making it hard to predict whether the surrounding volume shrinks or expands with time. If a buoyant particle is trapped in a region in which Q_a is predominatly positive, then this region is a good candidate for aggregation. Persistent ocean eddies and other vortical structures are possibilities, not only because vorticity tends to dominate over strain, but also because such features have the ability to trap fluid for long periods of time. For dense particles, contraction occurs in areas dominated by strain, and it has been shown that aggregation of heavy particles can occur in strain-dominated filaments that arise in particle-laden turbulent flows, though the considered particle-to-fluid density differences tend to be quite large (see Brandt and Coletti, 2022 for a review). In our study, we will focus on eddies, and on lower dimension objects within eddies that can act as attractors for buoyant particles.

A simple example of aggregation is given by Haller and Sapsis (2006), who argue that the elliptical center of a steady, non-divergent 2d eddy, with $\vec{g} = |\vec{\Omega}|$=0, acts as an attractor for buoyant particles. Here Q_a (now =Q_r), is ostensibly positive near the elliptical center of the eddy, corresponding to contraction of the phase space of the rigid particle motion. Since the central fixed point of the velocity field of the eddy is also a fixed point of the slow manifold particle velocity (4), buoyant particles initiated about the center should migrate towards the center. If the eddy is inviscid and its streamlines are circular, then the pressure and azimuthal velocity are related by the cyclostrophic balance $\frac{1}{\rho_f}\frac{\partial p}{\partial r} = \frac{u_\theta^2}{r}$ so that $2Q_r = \frac{1}{\rho_f}\left(\frac{1}{r}\frac{\partial p}{\partial r} + \frac{\partial^2 p}{\partial r^2}\right)$, and

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for an eddy in solid body rotation ($u_\theta = \Gamma_s r$), $2Q_r = \frac{1}{\rho_f}\left(\frac{1}{r}\frac{\partial p}{\partial r} + \frac{\partial^2 p}{\partial r^2}\right) = 2\Gamma_s^2$. As suggested in Figure 1a, a small concentration of particles indicated by the cross hatched area shrinks as it moves towards the center of the eddy. The contraction is partially due to the geometric effect of movement towards smaller radius (term $\frac{1}{r}\frac{\partial p}{\partial r}$) but also due to the fact that the pressure gradient decreases to zero as the center is approached and thus the inner edge of the path moves more slowly inward than the outer part (term $\frac{\partial^2 p}{\partial r^2}$). In the case of solid body rotation the two terms contribute equally. A second example (Fig. 1b) is of an eddy with an azimuthal velocity given by $u_\theta = \Gamma_C r^{1/2}$. Here $\frac{\partial^2 p}{\partial r^2} = 0$ and $2Q_r = \frac{1}{\rho_f}\left(\frac{1}{r}\frac{\partial p}{\partial r}\right) = \Gamma_C^2/r > 0$, so the contraction of the patch is entirely due to the geometric effect of its movement towards smaller radius. The most curious case is that of a point vortex: $u_\theta = \Gamma_P r^{-1}$, for which $2Q_r = \frac{1}{\rho_f}\left(\frac{1}{r}\frac{\partial p}{\partial r} + \frac{\partial^2 p}{\partial r^2}\right) = \frac{\Gamma_P^2}{r^4} - \frac{3\Gamma_P^2}{r^4} < 0$. Here the vorticity is zero away from the eddy center and the velocity field is dominated by strain. The pressure gradient *increases* as the center of the vortex is approached, meaning that the inner part of the patch moves towards the center more rapidly than the outer portion (Fig. 1c) and this tendency (quantified by the factor $-\frac{3\Gamma_P^2}{r^4}$) surpasses the tendency towards geometrical contraction (quantified by the factor $\frac{\Gamma_P^2}{r^4}$). The phase space of the particle motion thus expands as particles are drawn towards the center of the vortex. This behavior is made possible by the singularity at the center, and although this feature is artificial, point vortices are often used in idealized models of fluid flow and will act as sinks or "black holes" for buoyant particles even though $2Q_r < 0$.

The sign of Q_a is clearly not the whole story and does not encompass the effects of boundaries. For example, consider the fate of heavy ($\rho_f < \rho_p$) particles in the eddy show in Fig. 1a. The

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particles will migrate outward in each case, and no interior attraction will occur unless the eddy is surrounded by a rigid boundary, which would then act as an attractor.

In the next section, we will consider a more general, 3D, eddy-like circulation: one that has both vertical and horizontal components of vorticity, time dependence, and a variety of vortical structures that act as candidates for attraction. Our model is based the incompressible flow in a rotating cylinder (Greenspan, 1986), which has been studied in many configurations by numerous authors as a models of ocean circulation, ocean eddies, and industrial processes, and can be easily set up in the laboratory setting. It its original configuration the cylinder rotates about a vertical axis at a constant (positive) angular velocity ($\vec{\Omega} = \Omega\vec{k}$), and the lid, which is in contact with the fluid, rotates with a slightly greater angular speed. The differential rotation sets up an azimuthal circulation in the horizontal and an overturning circulation in the vertical. (Overturning is observed in ocean eddies as well and Ledwell et al. (2008) present an example.) The steady, axially symmetric state that is established will be our first object of investigation. A steady but asymmetric perturbed variant can be established by moving the axis of rotation of the lid away from the axis of rotation of the cylinder, and this offset can also be varied in order to induce time dependence. Fountain et al. (2000) set a similar situation up in a laboratory cylinder using a submerged impeller that can be tilted, rather than the differentially rotating lid that can be shifted, to establish an asymmetric disturbance flow. The authors discussed the Lagrangian characteristics of the undisturbed flow and demonstrated the existence of secondary vortical structures generated when the flow is perturbed. Pratt et al. (2014) reproduced similar structures using a primitive equation simulation and explored the rich assembly of chaotic regions and non-chaotic vortical structures as function of the Ekman and Rossby numbers of the flow. The time-dependent version of the rotating cylinder flow and a theory describing the resulting vortical

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structures were discussed by Rypina et al. (2015), who based their examples on a phenomenological model that reproduced many of the qualitative features of the numerically-obtained velocity field. In dimensionless Cartesian coordinates, the model velocity field is given by

$$u^{(x)} = -bx(1-2z)\frac{r_o - r}{3} - ay(c+z^2) + \varepsilon\left[y\big(y - y_o + \gamma cos(\sigma t)\big) - \frac{r_o^2 - r^2}{2}\right](1-\beta z)\,, \qquad \text{(9a)}$$

$$u^{(y)} = -by(1-2z)\frac{r_o - r}{3} + ax(c+z^2) - \varepsilon x\big(y - y_o + \gamma cos(\sigma t)\big)(1-\beta z), \qquad \text{(9b)}$$

$$u^{(z)} = bz(1-z)\frac{2r_o - 3r}{3}, \qquad \text{(9c)}$$

in which $r = (x^2 + y^2)^{1/2}$ and r_o is the cylinder radius. The velocity field consists of a steady, axially symmetric flow of strength a with an overturning circulation of strength b. To this symmetric state one can add an asymmetric, possibly unsteady and depth dependent, perturbation of amplitude ε (not to be confused with the Stokes number $\tilde{\varepsilon}$). The perturbation is quantified by an offset parameter y_o that introduces axial asymmetry in the velocity field, a frequency σ, and an amplitude β for linear depth dependence and an amplitude γ for the time dependence. For the case of axially symmetric, steady flow ($\varepsilon = 0$) the horizontal velocity field, in cylindrical coordinates, becomes

$$u^{(r)} = -br(1-2z)\frac{r_o - r}{3} \qquad \text{(10a)}$$

and

$$u^{(\theta)} = ar(c+z^2), \qquad \text{(10b)}$$

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where θ is the azimuthal angle. Table 1 lists the parameter values used for each numerical experiment.

In the steady, symmetric configuration, each fluid trajectory is confined to the surface of a torus as it winds around the cylinder. The typical torus is associated with quasi-periodic trajectories and any such trajectory, followed for a sufficient length of time, will sketch out the torus in 3D. Fig. 2b contains several examples of such tori and Fig. 2a shows the corresponding Poincare map, made by marking the crossing points of trajectories through a vertical slice through the cylinder. After a large number of crossings each quasi-periodic trajectory traces out the cross section of the torus on which it lives. The tori are nested within each another, with a single, horizontal, periodic trajectory located at the center of the nest. Certain tori contain periodic trajectories, and these will show up as a finite number of dots on the Poincare map. Because of this geometry, the motion of fluid parcels is most naturally described in terms of action-angle-angle variables, where the action, I, acts a label for a particular torus and is constant following each trajectory, and the two angle variables, $\tilde{\theta}$ and ϕ, define the location of a parcel on the torus. Here $\tilde{\theta}$ is an azimuthal angle that differs from the above cylindrical coordinate θ in how its origin is defined, while the 'poloidal' angle ϕ wraps around the cross-section of each torus. The coordinate are non-orthogonal but are defined in such a way that the angular velocities, $\Omega_{\tilde{\theta}}$ and Ω_{ϕ}, are also constant following a trajectory. The explicit transformations to the action-angle-angle variables are given in Mezic and Wiggins (1994).

When the symmetric RC flow is perturbed by a small, steady, symmetry-breaking perturbation, as controlled by the parameters ε and y_o in (9), the tori that are populated by periodic orbits potentially become resonant and break up, resulting in chaotic motion of fluid parcels in the

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vicinity (Fig. 2d-i). Tori with quasiperiodic orbits deform but stay intact. Examples are discussed by Fountain et al. (2000) and Pratt et al. (2013), and the latter found that chaos generally dominates in a large region that includes the central axis of the cylinder and extends around the boundaries of the cylinder. Away from this region the space is occupied by tori that have survived the perturbation, and these are sandwiched between tori that have broken up and created braided regions of chaos. The breakup of a torus also gives rise to new tori that appear as islands in the Poincare maps (Fig. 3d and 3g) and these contain non-chaotic trajectories. The number of islands can be predicted by a theory that decomposes the symmetry-breaking perturbation into Fourier modes, written in the $(I, \tilde{\theta}, \phi)$ coordinates, with wave numbers *n* and *m* in the $\tilde{\theta}$ and ϕ direction. If the angular velocities $\Omega_{\tilde{\theta}}$ and Ω_{ϕ} characterizing the trajectories on a particular torus satisfy the resonance condition $n\Omega_{\tilde{\theta}} + m\Omega_{\phi} = 0$ for some *n* and *m*, equivalent to the trajectories on that torus being periodic, then that torus will break up and a new set of invariant tori (islands) will form. Running through the center of the islands will be a periodic trajectory that will execute *n* azimuthal cycles to every *m* poloidal (overturning) cycles. In the case shown in Fig. 3a, *n*=*m*=1, so the periodic trajectory circles the cylinder horizontally once for each overturning cycle: a so-called 1:1 resonance.

If the symmetry breaking perturbation is quasi-periodic in time, with underlying frequencies σ_i, the resonance condition for the breakup of a torus becomes $n\Omega_{\tilde{\theta}} + m\Omega_{\phi} + l_i\sigma_i = 0$, where l_i's are integers (Rypina, et al. 2015). Unlike the resonance condition for the steady perturbation, which is only satisfied on tori foliated by periodic trajectories, this new resonant condition may be satisfied on tori that have quasi-periodic orbits, and the resonant islands that form will have a shape and location that vary in time. An example (Fig. 2g,h) of the case of a resonance with a single-frequency (i.e., time-periodic) perturbation shows a number of resonant islands. These

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features vary in time, recovering their shape and location periodically, and the snapshots shown are obtained by strobing the trajectories in 3D and at the forcing frequency. The green and blue islands in Fig. 2h have resulted from the breakup of tori with quasiperiodic trajectories, and center of the island corresponds to a closed material curve that is populated with quasiperiodic trajectories.

III. Results

Aggregation of rigid particles will occur in presence of an attractor, an object with a dimension <3 to which particles tend asymptotically in time. We are most interested in attractors that occur in the interior of the rotating cylinder, and are set up by the background circulation, as opposed to the physical boundaries of cylinder. We will see that a closed material contour consisting of periodic orbits near the core of the nested tori in the steady symmetric case act as an attractor for slightly buoyant particles, and that similar material contours consisting of periodic or quasiperiodic orbits near the centers of the resonant islands in the asymmetric cases can play the same role. We will explore three cases in increasing complexity, beginning with steady flows with axial symmetry, and proceeding to steady, asymmetric flows and finally unsteady asymmetric flows.

The search for attractors is motivated by the hypothesis that for cases of strong drag, where the particle velocity lies close to the fluid velocity, a periodic orbit for the particle motion will exist in the vicinity of a periodic trajectory for the fluid motion, and that if $Q_a>0$ in a region surrounding the latter, that it should attract particles. For the time-dependent case, we extend the search to included closed material contours that contain recirculating particles and that vary periodically in time.

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(a) steady, axially-symmetric 3D flows

The fluid velocity field for this case is given by Eqs. 9c and 10, and these indicate that the location of the horizontal, periodic trajectory living at the center of the nested tori, is given by $r = 2r_o/3$ and $z = \frac{1}{2}$. It is natural to ask whether a periodic trajectory for rigid particles also exists nearby. In the slow-manifold approximation, the steady radial, azimuthal and vertical particle velocities are obtained by writing (4) in cylindrical coordinates, leading to

$$v^{(r)} = u^{(r)} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\left[\left(u^{(r)}\frac{\partial}{\partial r} + u^{(z)}\frac{\partial}{\partial z}\right)u^{(r)} - u^{(\theta)}\left(2\Omega + \frac{u^{(\theta)}}{r}\right) - \Omega^2 r\right] \tag{11a}$$

$$v^{(\theta)} = u^{(\theta)} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\left[\left(u^{(r)}\frac{\partial}{\partial r} + u^{(z)}\frac{\partial}{\partial z}\right)u^{(\theta)} + u^{(r)}\left(2\Omega + \frac{u^{(\theta)}}{r}\right)\right] \tag{11b}$$

$$v^{(z)} = u^{(z)} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\left[\left(u^{(r)}\frac{\partial}{\partial r} + u^{(z)}\frac{\partial}{\partial z}\right)u^{(z)} + \mathrm{g}\right] \tag{11c}$$

<u>Position of attracting periodic orbit; approximate analytical expression on a slow manifold</u>

Searching for points $r = r_c$ and $z = z_c$ for which $v^{(r)} = v^{(z)} = 0$, and that lie in the proximity of the horizontal trajectory of the flow, we introduce

$$r_c = \frac{2r_o}{3} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\tilde{r} \text{ and } z_c = \frac{1}{2} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\tilde{z}.$$

Substituting into the right-hand sides of (11a,c) and setting both to zero results, after neglect of $O(\tilde{\varepsilon}^2)$ terms, in

$$r_c = \frac{2r_o}{3} + \tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\frac{g}{b}r_o \tag{12a}$$

and

$$z_c = \frac{1}{2} + \frac{9}{2br_o}\tilde{\varepsilon}\left(\frac{3R}{2} - 1\right)\left[\Omega^2 + a\left(c + \frac{1}{4}\right)\left(2\Omega + a\left(c + \frac{1}{4}\right)\right)\right]. \tag{12b}$$

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For the parameters a>0 and b>0, circulation is cyclonic with upwelling in the center of the cylinder, and (3R/2)-1>0 for buoyant particles, so the $O(\tilde{\varepsilon})$ corrections are positive and the periodic particle orbit lies at larger radius and elevation than the periodic fluid orbit. Note also from Eq. (11b) that the azimuthal velocity component of the rigid particle on the periodic orbit is equal to that of the fluid.

An explanatory sketch (Fig. 3) shows the position of the periodic orbit of the rigid particle relative to that of the periodic orbit of the fluid. Since the rigid particle is buoyant, it can maintain its level z only if it is situated in a region where the vertical fluid velocity is <0, here to the right of the fluid periodic orbit. Also, the horizontal pressure gradients associated with the centripetal acceleration associated with the frame rotation (term $\Omega^2 r$), the Coriolis acceleration (term $2\Omega u^{(\theta)}$), and the centripetal acceleration due to the azimuthal velocity ${u^{(\theta)}}^2/2r$ are all positive for this flow, so that low pressure exists at r=0 and the rigid particle is forced horizontally inward. To remain stationary the particle must sit in a region where the radial velocity of the fluid is outward. In this manner, the periodic trajectory exists at a location where the forces of inertia, buoyancy and added mass can be countered by the drag due to the background flow. If we fix all other parameters and increase Ω through positive values, the term multiplying $\tilde{\varepsilon}$ in (12b) will become dominated by the Ω^2 term and will grow without bound and the periodic trajectory may cease to exist. At the same time, a periodic orbit for the rigid particle can always be found close to that of the fluid, regardless of the magnitudes of the parameters Ω, a, b etc., provided that the relative particle size d/L (and thus $\tilde{\varepsilon}$), and/or the relative density difference $\frac{(\rho_f-\rho_p)}{\rho_f}$ (and thus $\frac{3R}{2}-1$) are made sufficiently small.

Position of attracting periodic orbit; conditions for the loss of periodic orbit

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We have suggested that periodic orbits for rigid particles are encouraged when the $\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)$ <<1, and in the case of Run 1 the value is .0066. A cross-sectional plot of the radial and vertical components of the slow manifold particle velocity in a vertical section through the cylinder (Fig. 4a) shows that the periodic orbit lies at r=.369 and z=.504 (as compared to the values $r_c = .338$ and $z_c = .502$ predicted by (12). (The convergence of the surrounding velocity field is too weak to be seen in the graphic.) If $\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)$ is raised to the moderately small value .02, the position of periodic trajectory migrates towards larger radius (Fig. 4b), the reason being that the greater buoyancy (larger value of $\frac{3R}{2}-1$) or smaller drag (larger $\tilde{\varepsilon}$) requires a larger downward fluid velocity for equilibrium. Since the maximum downward fluid velocity occurs at the outer cylinder wall (see Eq. 9c) the position of the periodic orbit continues to migrate outward and is lost (Fig. 4c) when $\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)$ exceeds a value close to 0.3.

Position of periodic orbit in numerical simulations:

The slow-manifold reduction yields to prediction (Eq. 12) of the position of the attracting material contour for slightly buoyant particles. We can compare this prediction to what is observed in numerical simulations using the Maxey-Riley equations (1) and (2) over a range of particle size d (and thus $\tilde{\varepsilon}$) and frame rotation Ω. As shown in Fig. 5, qualitative agreement with the slow-manifold prediction, and the sketch in Fig. 3, holds for a very small d (when $\tilde{\varepsilon}$ is small). Here the attractor in Fig. 5 is located close to the central periodic fluid parcel trajectory that lives at mid-depth, $z = 0.5$ and $r = \frac{2R}{3} \approx 0.33$. As d (and $\tilde{\varepsilon}$) increases, the attractor moves increasingly up and outward, and although the theory captures the trends, quantitative agreement with the numerical results worsens. Also, when frame rotation Ω is increased (panel c), the

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attractor responds by shifting up from mid-depth, again in qualitative but not quantitative agreement with the slow-manifold prediction in eq. (12b).

Geometry of particle trajectories and evidence of attraction in numerical simulations:

If in the neighborhood of the period trajectory the Q_a function >0, the phase space for buoyant particles will contract and the periodic trajectory becomes a candidate for an attractor of such particles. An example of the attraction towards the periodic orbit is shown in Figure 2c, where a set of slightly buoyant particles ($\frac{\rho_p}{\rho_f} = 0.97$) has been initialized over the volume of the cylinder, and eqs. (1) and (2) have been integrated forward in time to determine their subsequent trajectories. Each trajectory is shown using a unique color. It can be seen that the particles aggregate within a ring-like structure of decreasing thickness in the general vicinity of the periodic orbit of the fluid flow.

Basin of attraction – relationship to Q_a:

To map out the basin of attraction for the particle periodic orbit we first consider the region over which phase space contraction for the buoyant particles (i.e. Q_a>0) occurs. This region is shown in Fig. 6a for the current example, along with the streamlines of the fluid overturning stream function. Much of the fluid flow recirculates entirely within the region of positive Q_a, whereas some of the outer streamlines cross the boundary (thick contour) between positive and negative Q_a. If it were the case that rigid particles exactly followed streamlines of the fluid overturning circulation, then net contraction or expansion of phase space along a rigid particle trajectory would depend on the sign of the time-integrated value of Q_a along streamlines. The Q_a=0 contour, shown by a bold contour in each frame of Fig. 6, might then approximately delineate the

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basin of attraction for rigid particles. In the slow-manifold approximation, where rigid particle velocities lie close to the fluid velocities, the Q_a=0 contour might continue to do so.

To test this conjecture, we locate the basin of attraction in the numerical simulations by releasing buoyant particles at various locations in the cross-section $0<x<r_o$ and $0<z<1$, integrating the subsequent trajectories over many overturning cycles, and recording the position (x_{final} and z_{final}) of each particle where it crosses the same plane the final time (i.e., recording final crossing with the Poincare section). The values of z_{final} as a function of initial particle position are mapped in Fig. 7a, where the large green area corresponding to $z_{final} \cong 0.5$ indicates the region from which particles are attracted. Only particles initiated near the central axis of the cylinder, and close to the cylinder boundaries lie outside this region, and these rise to the surface of the cylinder, contact the upper lid, and are no longer followed. It can be seen that the green area in Fig. 7a has an oval shape that somewhat resembles the overturning streamlines at small x in the central part of the cylinder, but extends to near the top, bottom and outer cylinder boundaries at larger x. Thus the Q_a=0 contour provides a rough indication of the size and shape of the basin of attraction, but misses some important details.

Basin of attraction – dependence on Ω

We have seen that the location of the periodic orbit that acts as an attractor for buoyant particles shifts up and out in response to increasing frame rotation Ω (Fig. 5c). In Fig. 8 we indicate the corresponding changes in the extent of the basin of attraction with respect to changing Ω by re-computing Fig. 8a with $\Omega = 0.3$, 1, and 10. The two smaller Ω values (0.3 and 1) correspond roughly to Rossby numbers $a/2\Omega$ of about 1 and 0.2, i.e., are representative of the ocean submesoscale and mesoscale flows. The Q_a-functions for these cases are plotted in Fig. 6b-c.

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Most submesoscale eddies are going to tend to have $u^{(\theta)}/r$ about the same magnitude as Ω (except on the equator) and mesoscale eddies will have $u^{(\theta)}/r \ll \Omega$. The results in Fig. 8 suggest that, while the basin of attraction does shrink slightly with increasing Ω, this dependence is weak. The main difference between the three numerical runs in Fig. 8 is in the associated attraction time, which gets significantly shorter for larger values of Ω. This is explored in more detail below.

Attraction time:

It follows from Eq. (6) that the attraction time towards the periodic orbit should scale as $T_a = \left[2\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) Q_a\right]^{-1}$ where $Q_a = \frac{1}{2}\left(\left|\vec{\zeta}_a\right|^2 - |S|^2\right)$ with $\vec{\zeta}_a = \vec{\zeta}_r + \overrightarrow{2\Omega}$. Thus, for $\vec{\zeta}_r \geq 0$, as in most of our numerical runs (except Experiment 1e), attraction time decreases with increasing Ω for positive $\Omega \geq 0$. For negative $\vec{\zeta}_r$, which corresponds to the reversed direction of the flow in our simulations (Experiment 1e), an increase in Ω will initially slow the attraction by decreasing the magnitude of $\vec{\zeta}_a$ all the way to 0, at which point the periodic orbit will lose its attraction properties, but then will speed up the attraction as Ω is further increased. This trend is confirmed numerically in Fig. 9, where for the flow parameters corresponding to the "reversed flow" run in Table 1 (Experiment 1e, with $\vec{\zeta}_r < 0$), we release a sample trajectory within the basin of attraction and plot its z-coordinate as it winds around the can and eventually approaches the attracting periodic orbit. As anticipated, the attraction time initially increases as Ω is increased from 0 to 0.6, but then decreases as Ω is further increased to 2.

Disappearance of the subsurface attractor when $\tilde{\varepsilon}$ becomes too large:

Finally, to illustrate the disappearance of the subsurface attractor when $\tilde{\varepsilon}$ becomes too large, in Fig. 10, we contrast 2 numerical simulations with the same flow parameters (corresponding to

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the "slow overturn" run 1c in Table 1) but different particle diameters, $d = 10^{-3}$ vs $d = 5 \times 10^{-4}$. For larger d, the subsurface periodic orbit for rigid particles is no longer present within the can, leading to all particles rising up to the surface (Fig. 10b). For smaller d, the periodic orbit is still present and acts as an attractor for rigid particles over a significant portion of the can (green region in Fig. 10a). We note that this run is more typical of oceanic mesoscale or submesoscale eddies, where the overturning component of circulation is weak in comparison to the horizontal swirl.

(b) steady non-symmetrically perturbed case

We now consider a case in which the axial symmetry of the steady flow has been broken, here through a change in the perturbation amplitude parameter ε from zero to 0.25**,** and in the offset parameter y_o from 0 to -0.2 in the Eqs. 9a,b. The fluid velocity field now contains something like a stationary, "mode-1" azimuthal wave in the horizontal velocity field.

The resulting Lagrangian structure (Fig. 2d and e) has a sea of chaos that covers the near-axial and outer regions of the cylinder, where no unbroken tori anymore exist. Within this chaotic sea is a region containing a nest of unbroken tori that surround a central periodic orbit. This orbit has evolved from the central periodic orbit of the symmetry case and is now tilted. Within the nest of unbroken tori there exist resonant layers, in which new tori have arisen, and the most prominent is the "island" that is centered near x=0.4 and z=0.2 in the right-half (and near x=0.4 and z=0.2 in the right half) of Fig. (2d). We further note that this center lies within the region of positive Q_a (Figure 6b). The island corresponds to the yellow tori in Fig. 3e and is produced by a 1:1 resonance, so that the periodic trajectory running through its center executes one complete azimuthal cycle and one overturning cycle before connecting back onto itself. Thus, in this steady asymmetric configuration, we now have 2 periodic orbits of the fluid flow – the central

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slightly-tilted periodic orbit near mid-depth (that evolved from the central horizontal periodic orbit of the axisymmetric flow) and a new periodic orbit running through the center of the resonant island (resulting from the break-up of the resonant torus satisfying $\Omega_{\tilde{\theta}} + \Omega_{\phi} = 0$).

We speculate that for sufficiently small $\tilde{\varepsilon}$ a periodic orbit for the rigid particle motion exists in the vicinity of each of the 2 periodic orbits of the fluid flow. This conjecture is difficult to prove due to a complex geometry, leading to centrifugal forces that act in different directions at different locations along the particle path. For now we simply search for the supposed attractors by releasing particles and following their trajectories.

As shown in Fig. 2f, separate attractors arise in the vicinity of two periodic orbits. The first appears as a ring-like structure (purple core) lying near the center of the original nested tori and the second is a similar feature with a red core near the center of the resonant island. The two are chained together and each has its own basin of attraction (Fig. 7c): the first consisting of a roughly elliptical patch (inner green region) in the x-z-plane, which corresponds of a slice through a tube-like structure in 3D, and the second consisting on an annular (blue) region that surrounds the green region and that occupies a relatively larger volume.

In order to check that attraction of slightly buoyant, rigid particles towards periodic orbits located near the centers of the resonant islands in the perturbed flow is not limited to the case of the 1:1 resonance, in an additional simulation (Fig. 11, experiment 2c in Table 1), we adjusted the background flow parameter b in Eqs. (9), which is responsible for the overturning strength, to create a 2:1 resonance instead of a 1:1 resonance, as in the original run. In this case, the resonant torus breaks down giving rise to a 2-island chain on the corresponding Poincare section (Fig. 11a), and the periodic orbit that goes through the centers of both islands completes 2 full cycles

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in azimuth and 1 complete cycle in vertical before connecting onto itself. Also, as in the original run, a second slightly-tilted periodic orbit still exists near mid-depth of the can. When buoyant particles are released into this flow, two attractors arise, corresponding to the 2 periodic orbits – one near mid-depth (purple core in Fig. 11c) and another in red near the center of the 2:1 resonant island.

Shift in position of the periodic orbit associated with a resonant island as a function of flow and particle parameters, and frame rotation

The position of the attracting periodic orbit for rigid particles that is located within the resonant islands (we will refer to it as the resonant periodic orbit) in the asymmetrically-perturbed flow depends both on the perturbation strength (via ε), on the flow and particle parameters (via $\tilde{\varepsilon}$), and on the frame rotation Ω. Specifically, this resonant periodic orbit for the rigid particles will shift away from the corresponding periodic trajectory of the fluid flow as $\tilde{\varepsilon}$ and Ω are increased. The same is true for the slightly-tilted central attracting periodic orbit near mid-depth. This is qualitatively similar to the shifting of the central periodic orbit up and out from $z = 0.5$, $r = 0.34$ in the axisymmetric flow in response to changing $\tilde{\varepsilon}$ and Ω, which we explored in detail the previous section both analytically (Eqs. 12) and numerically (Fig. 3-5).

In order to numerically illustrate the shift in the position of the attracting periodic orbits, we present (Figs 12 and 13) numerical simulations in the steady perturbed flow configuration for 3 values of d (and thus $\tilde{\varepsilon}$) and 3 values of Ω. As both parameters increase, the attractors move away from the corresponding periodic orbits of the fluid flow. This shift is evident from the change in the color of the attraction basins in (a,d,g) and from the location of the yellow cloud of dots in (c,f,i) in Figs. 12-13. Increases in $\tilde{\varepsilon}$ and Ω also lead to the shrinkage of the attraction

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basins for both attractors and to a faster convergence rate, as is evident from the tighter cloud of yellow dots in (c,f,i), as discussed in more detail below. The basin of attraction for the central attractor – the green region in Fig. 12 – seems to shrink faster than the basin of attraction for the resonant attractor (the blue-ish region) as d increases, so when d is increased from 2×10^{-3} to 3×10^{-3}, the central attractor vanishes, whereas the resonant attractor is still present (Fig. 12g). On the other hand, the increase in Ω (Fig. 13) causes a faster shrinkage of the basin of attraction for the resonant attractor than for the central attractor, so when Ω is increased from 2 to 5 in Fig. 13g, the resonant attractor disappears, whereas the central attractor is still present. Figs. 12g,h,i (and Fig. 13g,h,i) show cases where this threshold has been exceeded, and one of the attractors has been lost, whereas the other is still present.

Attraction time:

Similar to the unperturbed flow, the attraction time for attractors in the steady, perturbed flow may still scale as $T_a = \left[2\tilde{\varepsilon}\ \left(\frac{3R}{2} - 1\right) Q_a\right]^{-1}$, provided that Q_a is regarded as a typical value within the corresponding basin of attraction. The predicted decrease in attraction time with increasing $\tilde{\varepsilon}$ and Q_a is evident from the numerical simulations in Figs. 12-13, where in (c,f,i) we color-coded trajectory crossings with the x-z Poincare plain by time, with blue/yellow corresponding to initial/final time. For smaller values of $\tilde{\varepsilon}$ and Ω, we observe a wider and more diffuse cloud of dots (because trajectories wind around the can many times before approaching the attractor), whereas as $\tilde{\varepsilon}$ and Ω increase, the clouds at comparable times become denser and more compact around the attractors.

Basin of attraction

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For the slightly-tilted central periodic orbit located within the central non-chaotic region near mid-depth in Fig. 2f, we observe that the basin of attraction – green region in Fig. 7b – extends roughly from the location of the periodic orbit to the edge of the central non-chaotic region (that is foliated by discretely sampled closed curves in Fig. 2d). Note that as $\tilde{\varepsilon}$ increases, the attracting periodic orbit moves away from the center of this non-chaotic region towards its edge, leading to the shrinkage and eventual disappearance of the corresponding basin of attraction, shown by the green regions in Fig. 12a,d,g).

Similarly, in all of our numerical simulations, we observe that for the resonant attracting periodic orbit running through the resonant islands, the basin of attraction seems to cover the region between the orbit and the edge of the corresponding resonant island. An analytical expression for the width of the (non-degenerate) resonant island in the fluid flow (Pratt et al., 2014) predicts that $\Delta I = \sqrt{\frac{\epsilon F_{nm}^{0}(I_0)}{\left(n\frac{d^j\Omega_\phi}{dI^j}+m\frac{d^j\Omega_\theta}{dI^j}\right)_{I_0}}}$, where ΔI is the deviation in the action coordinate away from I_0, the value of action at the resonant torus (i.e., at the center of the island). This width depends on the strength of the perturbation ϵ, the order of the resonance (via n and m in the resonance condition), the background flow (via $\frac{d^j\Omega_{\phi/\theta]}}{dI^j}$), and the structure of the perturbation (via $F_{nm}^{0}(I_0)$). This expression could be used as an upper limit on the extent of the basin of attraction. However, because the attracting periodic orbit will move away from the center of the island towards its edge as $\tilde{\varepsilon}$ and Ω increase, the basin of attraction for the resonant attractor (blue region in Figs. 12a,d and 13a,d) becomes increasingly smaller than ΔI. One might speculate, then, that the attractor will completely disappear when the attracting periodic orbit reaches the edge of the resonant island. This is the case in Figs. 13g where the resonant attractor is no longer present.

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(c) non-steady, non-symmetrically perturbed case

The final case that we will consider is one in which the perturbation is asymmetric and varies periodically in time. The chosen perturbation frequency, $\sigma = 2\pi/9.1$, causes 2 strong additional resonances (compared to the steady perturbed case) – one with n=0, m=1, and l=1 (i.e., with a torus whose overturning frequency is equal to the perturbation frequency) that is shown in blue in Fig. 2g,h and is located near the outer edge of the central non-chaotic region, and another resonance, shown in green in Fig. 2g,h, with n=1, m=1, and l=1, which is located between the central non-chaotic region and the larger n=1, m=1 resonant island (that was present in the steady case as well). Both of these new resonant structures are time dependent, their shape and position recurring periodically. For example, the blue island, which looks like a crescent moon pointing upward on the Poincare section at t=0, becomes a crescent moon pointing downward at time 4.55. The movement of the green island is a more complex, as it turns both in azimuth and vertical, making one complete loop over 9.1 time units. Because of the time-dependence, trajectories must be strobed at the forcing frequency σ in order to capture 'snapshots' of their forms as they recur at a particular phase in the time cycle. At the center of each feature is a closed material curve that also varies periodically. Where the island has emerged from the breakup of a torus with quasiperiodic orbits, the individual trajectories that populate the material curves are themselves quasiperiodic.

Particle trajectory computations in this case confirm that the purple, red and green islands give rise to attractors (Fig. 3i), whereas the blue island does not. In fact, particles that are released in the blue region converge towards the attractor that lies near the purple region. This is also indicated by the basin of attraction of the central attractor extending across the space occupied by the blue resonant island in Fig. 7c.

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IV. Discussion

We have considered attraction phenomena for small, spherical, buoyant, rigid particles in a three-dimensional rotating cylinder flow with azimuthal rotation and overturning. The aim has been to gain insights into the behavior of slightly buoyant microplastic particles in 3D vortex flows that qualitatively resemble ocean eddies. The particle motion is governed by a simplified version of the Maxey-Riley equations (accounting for inertia, buoyancy and simplified quantification of drag and added mass), and, approximately, by the slow-manifold reduction of these equations.

We have explored a steady axisymmetric rotating cylinder flow and a steady flow with its axial symmetry broken. In all cases, we have observed emergence of subsurface attracting structures that lead to the aggregation of buoyant particles towards them. We have linked these attractors to the periodic orbits of rigid particles that exist in a region of net contraction of the phase space of the particle motion. The slow manifold equations suggest that periodic orbits for rigid particles exist near periodic orbits of the underlying fluid flow, provided the drag is sufficiently strong (Stokes number <<1).

We have also explored one case of an axially asymmetric and time-periodic flow, with focus on the resonant “islands” that arise due to the time-dependence. At the center of such islands are closed material contours composed of quasi-periodic orbits of the fluid flow. One such structure has nearby attractor, also a closed material contour of quasiperiodic orbits for rigid particles, while a second example does not. A detailed explanation awaits formulation of a quantitative theory, something that is beyond the scope of the present paper and that will be presented in a future work.

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We have observed that the disappearance of an attractor, which can occur as the result of increasing particle size or frame rotation, coincides roughly with the displacement of the position of the attractor to the outer edge of the resonant island from which it sprang. Whether this purely geometric observation forms the basis for a general criterion for the loss of attraction is unknown, as a dyamical justification is needed.

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Experiment	a	b	ε	y_o	σ	γ	β	Ω	d
1 – steady symmetric	0.62	7.5	0	0	0	0	0	0	10^{-3}
1a (small Ω)	0.62	7.5	0	0	0	0	0	0.3	10^{-3}
1b (large Ω)	0.62	7.5	0	0	0	0	0	1	10^{-3}
1c (slow overturn)	0.62	0.25	0	0	0	0	0	1	10^{-3} $vs.$ 5×10^{-4}
1d ($z_{attractor}$ vs Ω)	0.62	7.5	0	0	0	0	0	Sweep 0 to 10	10^{-3}
1e (reversed flow)	-0.62	-7.5	0	0	0	0	0	0, 0.6, 2	10^{-3}
2 – steady asymmetric	0.62	7.5	0.25	-0.2	0	0	0	0	10^{-3}
2a (small Ω)	0.62	7.5	0.25	-0.2	0	0	0	0.3	10^{-3}
2b (large Ω)	0.62	7.5	0.25	-0.2	0	0	0	1	10^{-3}
2c (2:1 resonance)	0.62	3.8	0.25	-0.2	0	0	0	0	10^{-3}
3 - non-steady asymmetric	0.62	7.5	0.25	-0.2	$\frac{2\pi}{9.1}$	0.2	1	0	10^{-3}

Table 1: Dimensionless parameter values for numerical experiments. Fixed parameters in the kinematic model (Eqs. 9a-c) are c=0.69, and r_o=1/2 in all cases. Parameters that appear in the nondimensional Maxey-Riley equation (3) are also nondimensional, with *L*, *U L/U* as length, velocity and time scales. Fixed parameter values based on *L*=1m and *U=1m/s* include $\frac{\rho_p}{\rho_f} = 0.97$, $R = \frac{2\rho_f}{\rho_f+2\rho_p} = 0.680$, $\frac{3R}{2} - 1 = .020$ $\vec{g}_r = \frac{gL}{U^2} = 10.0$, $\tilde{\varepsilon} = \frac{2}{9}\left(\frac{d}{L}\right)^2 \frac{UL}{\nu}\frac{1}{R}$=0.33, and $\tilde{\varepsilon}\left(\frac{3R}{2} - 1\right) = .0067$. Note that $\vec{\Omega} = \Omega\vec{k} = \frac{\vec{\Omega}^* L}{U}$.

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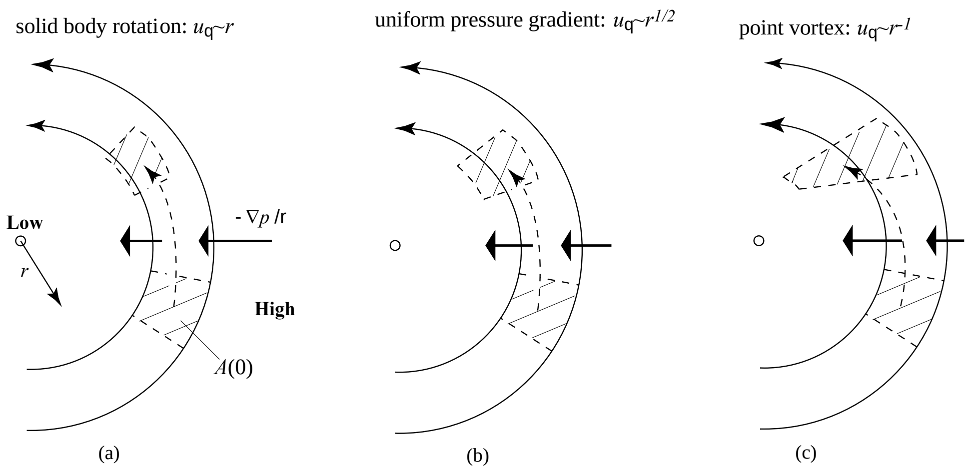

Figure 1. Three types of two-dimensional eddies with zero frame rotation and for which gravity is imagined to be zero: solid body rotation (a), constant pressure gradient (b), and point vortex (c). In each case, the cross hatched area represents a concentration of rigid particles with area $A(t)$.

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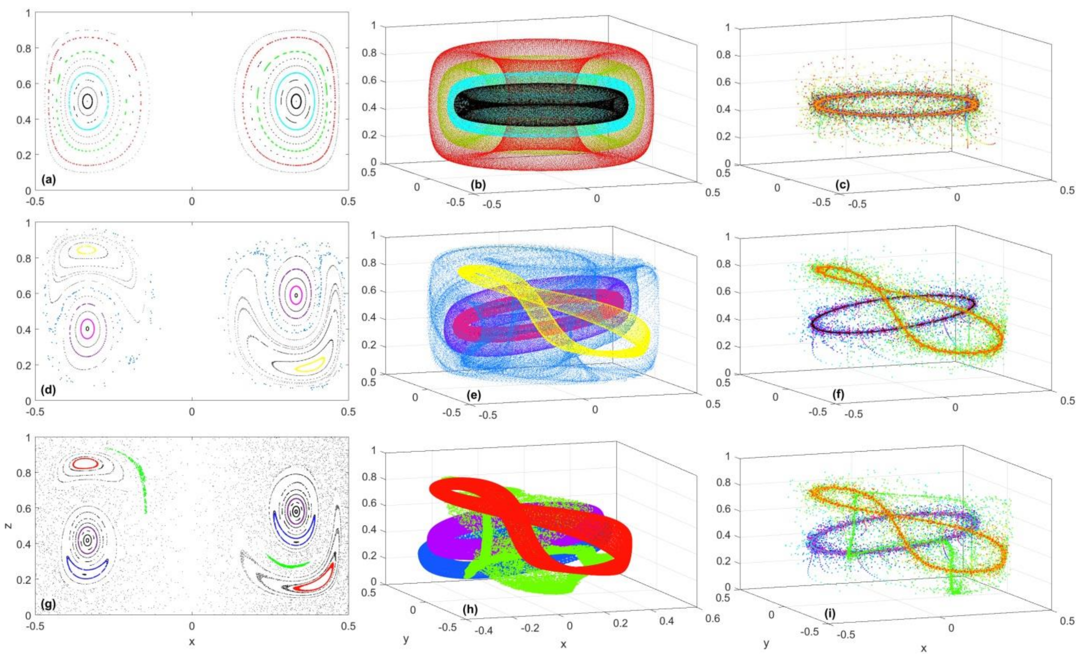

Figure 2. (left) Poincare section, (middle) fluid parcels trajectories in 3D, (right) buoyant particle trajectories in 3D for a steady symmetric fluid flow (top row), steady asymmetric flow (middle row), and non-steady, asymmetric flow. Parameter setting are listed under Experiments 1, 2 and 3 in Table 1. Colors in the left column of panels match the corresponding panel in the middle column, but the colors in the right column indicated time after release of the particles. Note the attraction of buoyant particles to a single attractor at mid-depth in panel (c), to 2 attractors in panel (f), and to 3 attractors in panel (i). Particles are released along a vertical line $x = 0.334$, $y = 0$, $0 < z \leq 0.6$ with initial velocity equal to that of the co-located fluid parcels.

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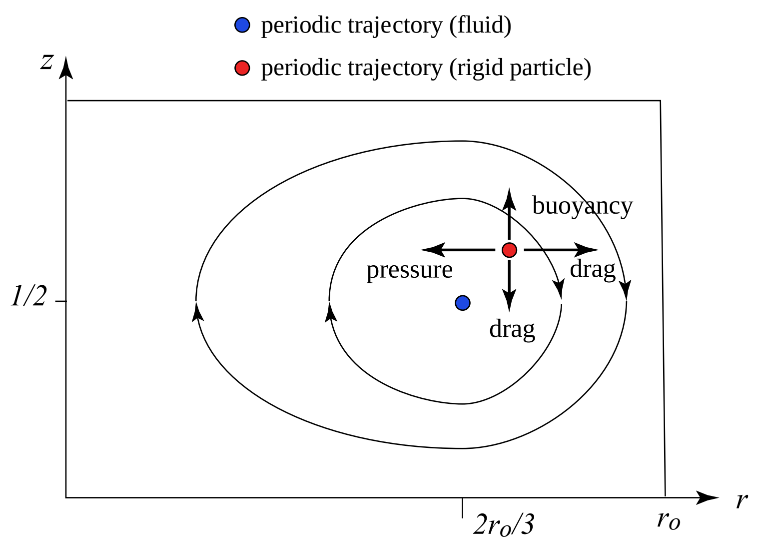

Figure 3. Sketch showing the position in a vertical section of the periodic orbit (red dot) of the rigid particle relative to the periodic orbit (blue dot) of the fluid flow. The viewer sees one half of a vertical slide though the cylinder, with the azimuthal flow directed away from the viewer and the cylinder center at the left edge.

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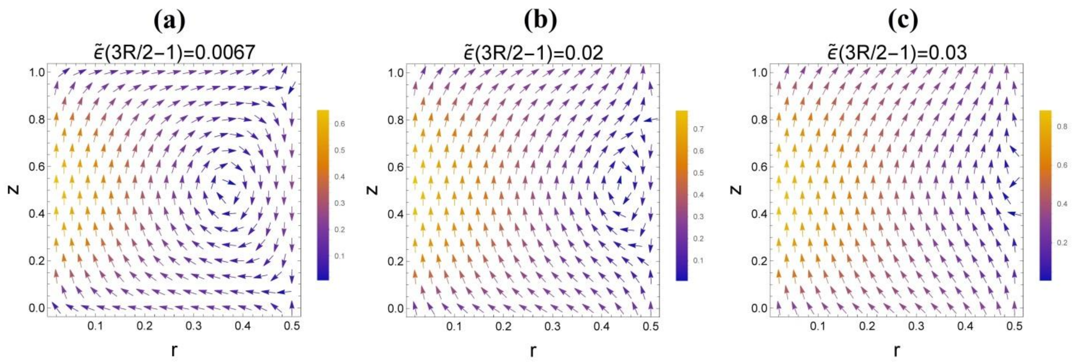

Figure 4. The slow-manifold radial and vertical velocity components for the rigid particles, plotted in the (r,z) plane for (a) $\tilde{\varepsilon}\left(\frac{3R}{2}-1\right)$ =.0067, (b) =.02, and (c) =.03. Other parameters are as listed for Experiment 1a in Table 1.

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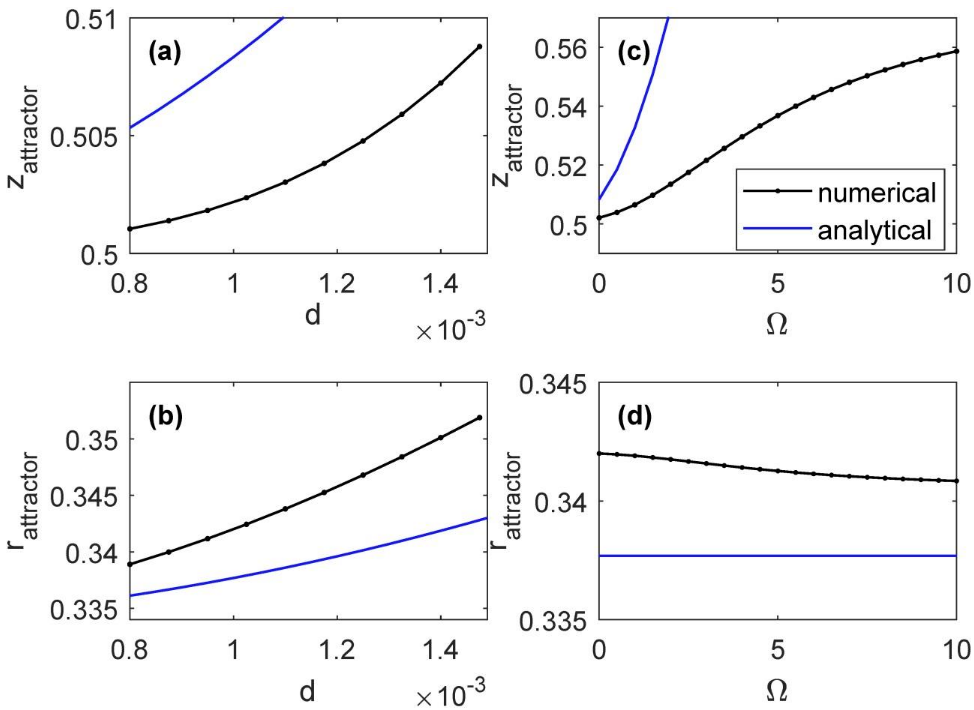

Figure 5. For the steady symmetric rotating cylinder flow, the coordinates of the periodic orbit that acts as an attractor for buoyant particles as a function of particle diameter (a-b) and frame rotation (c). Flow parameters are listed in Table 1 and correspond to Experiment 1 for (a-b) and Experiment 1d for (c-d).

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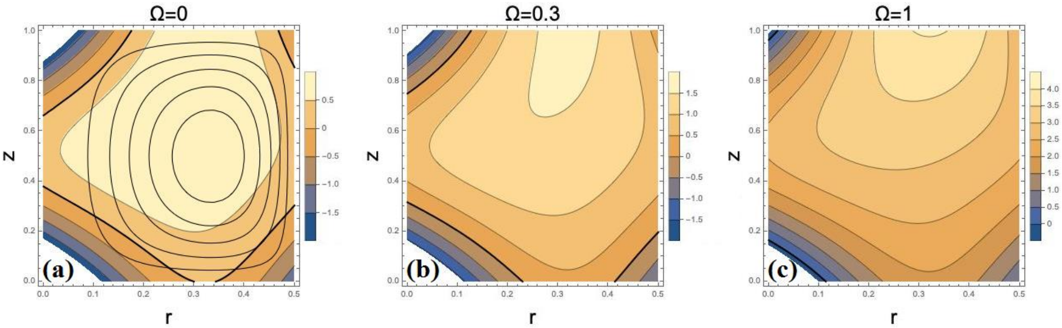

Figure 6. (a): The Q_a function for the steady, axisymmetric, cylinder flow with the same parameter setting (see Experiment 1a) as for Figure 3a-c, and plotted in (x,z) along with the streamlines of the overturning circulation. The thick rigid curve corresponds to Q_a=0. (b): The same parameter settings, except Ω has been raised from 0 to 0.3 (Rossby number ≅ 1) (c): Ω=1.0. (Rossby number ≅ 0.2).

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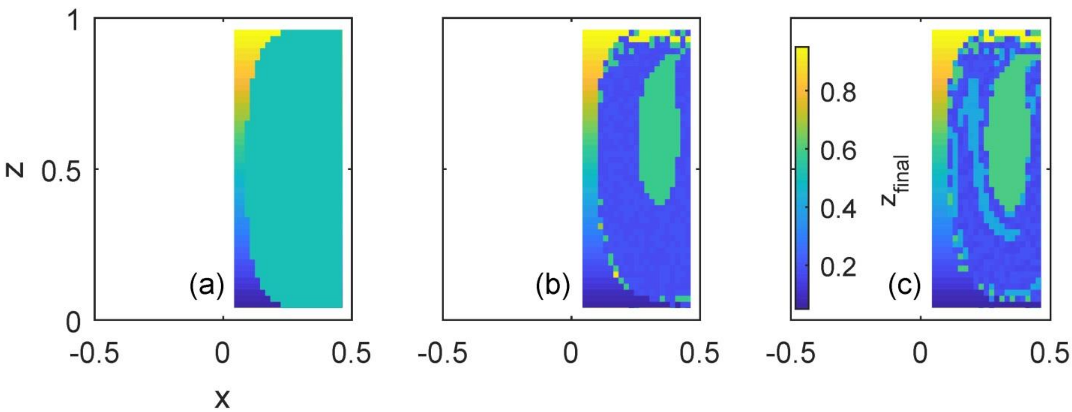

Figure 7. Domain of attraction for the attractors in (a) steady symmetric (Experiment 1 in Table 1), (b) steady asymmetric (Experiment 2 in Table 1), and (c) time-periodic asymmetric rotating cylinder flow (Experiment 3 in Table 1). (These are the same 3 experiments that were used to produce Fig. 2.) The color indicates the height (i.e., value of z-coordinate) of the final crossing of a trajectory with the Poincare section, as a function of particle's release location. Particles attracted to the same attractor thus correspond to same color.

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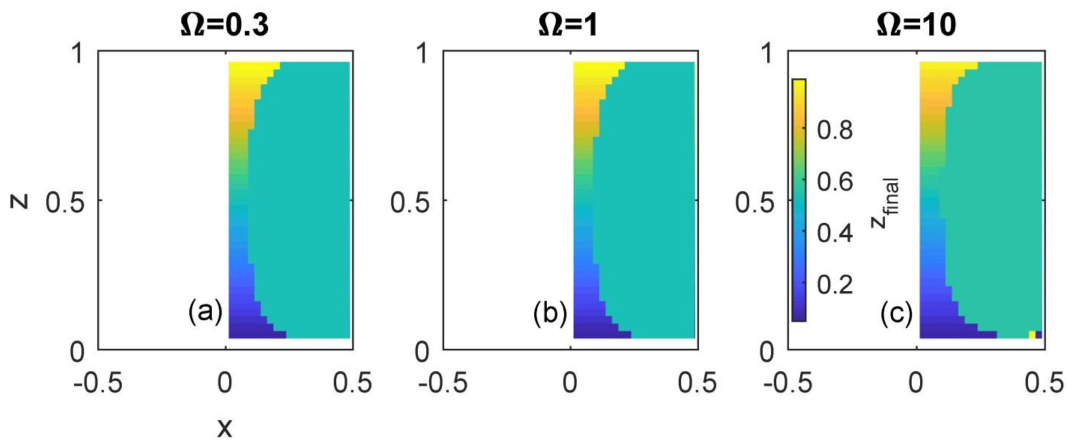

Figure 8. Same as in Fig. 7a but with frame rotation.

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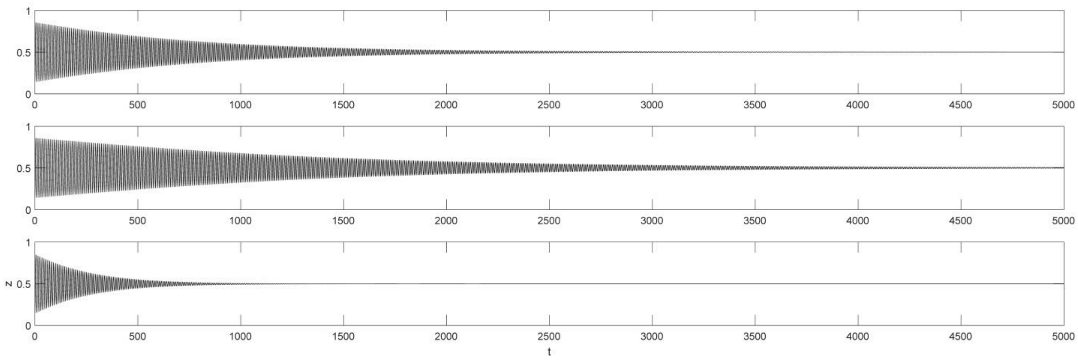

Figure 9. For the “reversed flow” experiment (Experiment 1e in Table 1), z-position of a sample particle trajectory as function of time for 3 values of Ω: 0 (top), 0.6 (middle), and 2 (bottom). Time t is in dimensionless units (but since our scaling coefficient for time is equal to 1 sec, the numbers on the x-axis can also be read as dimensional time in sec.)

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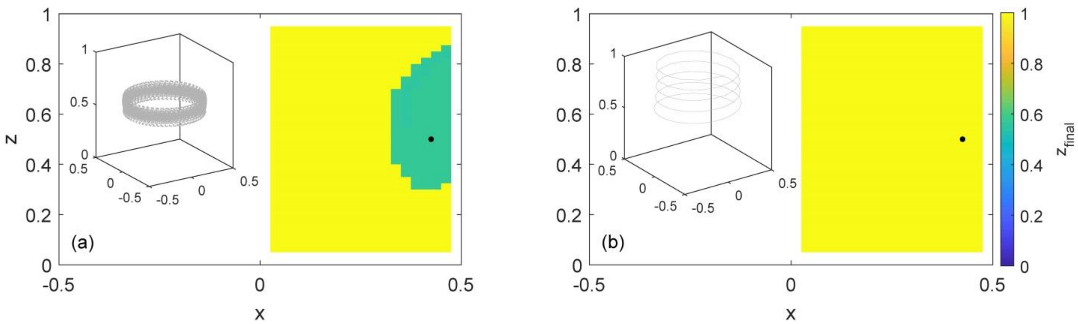

Figure 10. For the "slow overturn" Experiment 1c from Table 1, color indicates the final z-coordinate of a particle's trajectory at the end of integration time as a function of particle's release location for 2 values of d: (a) 5×10^{-4} and (b) 10^{-3}. Yellow corresponds to particles rising up to the top, whereas green indicates the basin of attraction of the subsurface attracting periodic orbit. The insets at the left side of each frame show a sample trajectory whose release location is indicated by the black dot.

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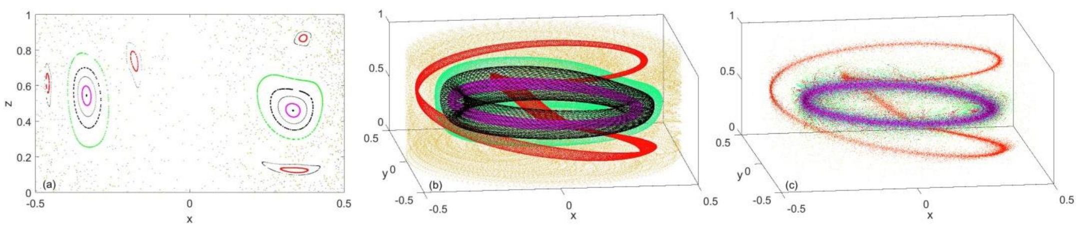

Figure 11. Same as Fig. 2(d-f) but with b=3.8.

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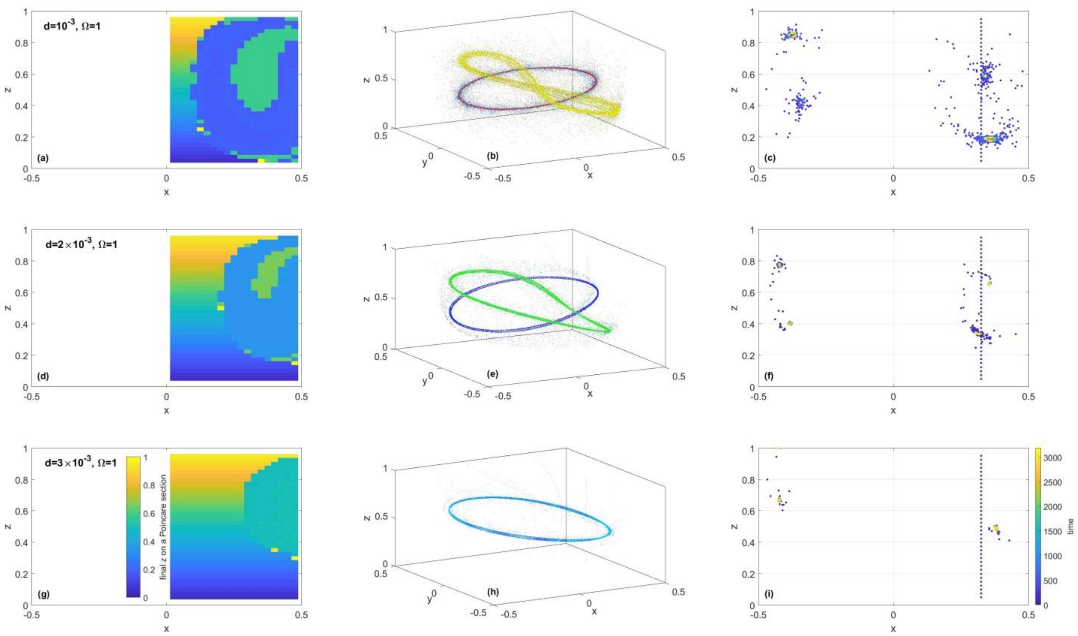

Figure 12. For the steady perturbed system (Experiment 2 in Table 1), changes in the location of the attracting periodic orbits, basins of attractions, and time of attraction as a function of particle diameter d (and thus $\tilde{\varepsilon}$). (a,d,g) show z-coordinate of the last crossing of trajectory with the x-z Poincare plane as a function of release location; flat regions are basins of attraction for the 2 attractors. (b,e,h) show 20 trajectories in 3d released along a vertical line at y=0, x=0.334, 0.05<z<0.95; denser cores indicate attractors. (c,f,i) show crossing of the same select 20 trajectories with the x-z Poincare plane, color coded by time; blue corresponds to release location, yellow corresponds to final positions.

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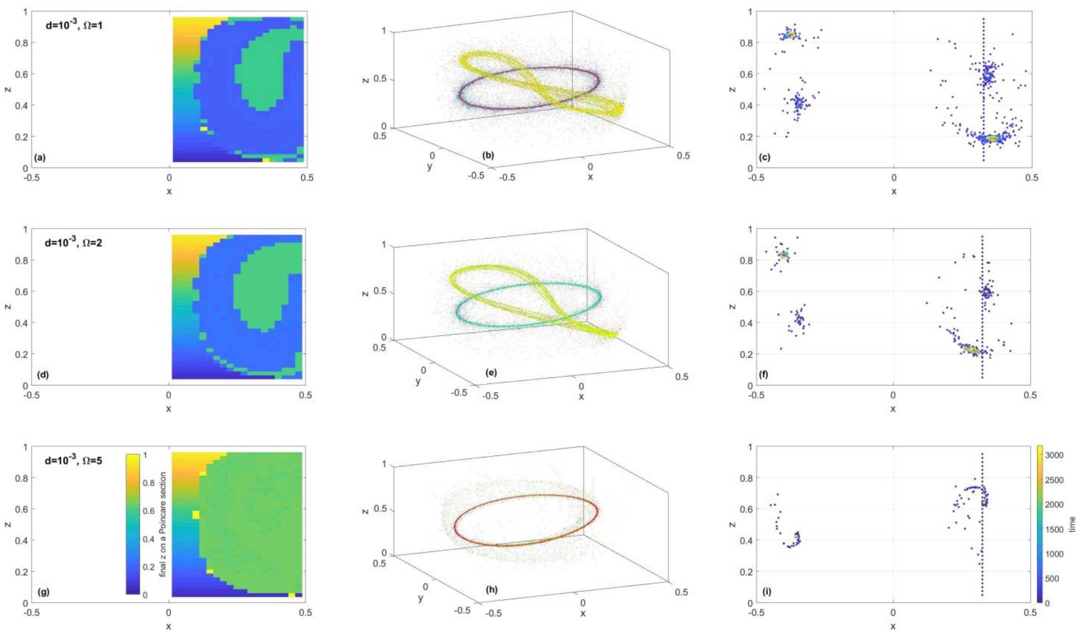

Figure 13. For the steady perturbed system (Experiment 2 in Table 1), changes in the location of the attracting periodic orbits, basins of attractions, and time of attraction as a function of frame rotation Ω. (a,d,g) show z-coordinate of the last crossing of trajectory with the x-z Poincare plane as a function of release location; flat regions are basins of attraction for the 2 attactors. (b,e,h) show 20 select trajectories in 3d released along a vertical line at y=0, x=0.334, 0.05<z<0.95; denser cores indicate attractors. (c,f,i) show crossing of the same 20 trajectories with the x-z Poincare plane, color coded by time; blue corresponds to release location, yellow corresponds to final positions.