

Real-Time Pollen Identification using Holographic Imaging and Fluorescence Measurements

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Abstract. Over the past few years, a diverse range of automatic real-time instruments has been developed to respond to the needs of end users in terms of information about atmospheric bioaerosols. One of them, the SwisensPoleno Jupiter, is an airflow cytometer used for operational automatic bioaerosol monitoring. The instrument records holographic images and fluorescence information for single aerosol particles, which can be used for identification of several aerosol types, in particular different pollen taxa. To improve the pollen identification algorithm applied to the SwisensPoleno Jupiter and currently based only on the holography data, we explore the impact of merging fluorescence spectra measurements with holographic images. We demonstrate, [using manually generated data](#), that combining information from these two sources results in a considerable improvement in the classification performance compared to using only a single source (balanced accuracy of 0.992 vs. 0.968 and 0.878). This increase in performance can be ascribed to the fact that often classes which are difficult to resolve using holography alone can be well identified using fluorescence and vice versa. We also present a detailed statistical analysis of the features of the pollen grains that are measured and provide a robust, physically-based insight into the algorithm's identification process. The results are expected to have a direct impact on operational pollen identification models, particularly improving the recognition of taxa responsible for respiratory allergies.

1. Introduction

Over the past decades a considerable increase in aeroallergen-related diseases such as asthma or allergic rhinitis has been observed (Ring et al. 2001; Woolcock et al. 2001; Woolcock et Peat 2007). This has resulted in a rise in associated direct and indirect health costs in terms of hospitalisation, medication costs and absence from work (Zuberbier et al. 2014; Greiner et al. 2011). Currently, the prevalence of pollen allergy ranges between 10 to 30% of the population in Westernised countries and up to 40% of children in high-income countries (Pawankar et al. 2011). In future, the relevance of pollen as an allergen may increase further as a result of climate change, which perturbs the life cycle of plants through drier environmental conditions and increased temperatures. Stressed plants tend to have an earlier and/or longer blooming season (Ziello et al. 2012) and

33 produce more pollen with higher concentrations of allergens (Damialis et al. 2019; Beggs 2016; D'Amato et al. 2016), possibly
34 further contributing to the increase and severity of allergic diseases. For these reasons, systems to measure airborne pollen
35 concentrations are essential to meet public health challenges associated with respiratory allergies. Through real-time
36 measurements and the development of forecast models (Chappuis et al. 2020), they can help reduce health costs with better
37 diagnosis and prevention, thus helping patients to better manage their symptoms.

38
39 Most European countries started monitoring pollen in the second half of the 20th century using Hirst-type instruments (Hirst
40 1952) with manual identification and counting part of the process (Clot 2003; Spieksma 1990). However, this method provides
41 data at low time resolution, typically daily mean values, after a processing time of up to 10 days. The spread of pollen grains
42 on the collection band and the limited sampling (Oteros et al. 2017) mean that data at higher temporal resolutions, or at low
43 concentrations (below 10 pollen grains/m³), have considerably increased uncertainty (Adamov et al. 2021). Although little
44 data is available to study atmospheric pollen phenomena at high temporal resolutions, it is widely expected that pollen
45 production and dispersal processes take place at sub-daily scales since they are highly influenced by local meteorological
46 environmental conditions (Rojo et al. 2015; Rantio-Lehtimäki 1994). Provision of real-time pollen data is also crucial for
47 forecasting purposes, since models can then integrate these real-time data to deliver considerably improved forecasts (Sofiev
48 2019).

49
50 Over the past few years, several instruments designed for real-time pollen monitoring have come onto the market (Crouzy et
51 al. 2016; Oteros et al. 2015), as comprehensively reviewed in previous work (Huffman et al. 2020; Buters et al. 2022; Maya-
52 Manzano et al. 2023). Among the most promising instruments are airflow cytometers which allow the characterisation of
53 particles almost in real-time as they pass through the instrument and enable continuous monitoring with high temporal
54 resolution (10 minutes as for weather parameters or below) over a whole season. In particular, the SwisensPoleno Jupiter
55 (developed by Swisens AG, Switzerland) is an instrument for bioaerosol identification which can take in-flight holographic
56 images of particles and measure their fluorescence (FL hereafter) (Sauvageat et al. 2020; Tummon et al. 2021; Lieberherr et
57 al. 2021). Coupled with a machine learning (ML) algorithm, it has been shown to perform well for pollen monitoring even if
58 the algorithm uses just the holographic data (Sauvageat et al. 2020; Crouzy et al. 2022; Maya-Manzano et al. 2023).

59
60 The FL data has to date not been used for pollen identification with the SwisensPoleno Jupiter. Sauvageat et al. 2020 reached
61 an accuracy above 96% for eight of the main allergenic pollen species in central Europe (*Ambrosia artemisiifolia*, *Corylus*
62 *avellana*, *Dactylis glomerata*, *Fagus sylvatica*, *Fraxinus excelsior*, *Pinus sylvestris*, *Quercus robur* and *Urtica dioica*) using
63 only holographic images. However, some species have similar morphologies which can cause misclassifications and thus lower
64 the algorithm performance, as previously identified in Sauvageat et al. 2020. In this paper, we investigate whether FL helps
65 discriminate single pollen grains between different allergenic taxa based on their chemical compositions to reduce the level of

66 confusion resulting from their similar shapes. Moreover, we also verify whether the FL measurements are consistent for each
67 species when using different SwisensPoleno units.

68 **2. Material and Methods**

69 In this work we investigate the impact of including the set of FL measurements, constituting the particle FL spectra, as input
70 for pollen identification using artificial neural networks. We trained and assessed the performance of three neural networks
71 with the same dataset but using different inputs: only holographic images (holo), only FL spectra (FL), or both (combined).
72 The performance of each model is evaluated using classical metrics, here the balanced accuracy, the F1-score, and Matthew's
73 Correlation Coefficient (MCC) as defined in Chicco et al. 2020, as well as the (relative) error rate derived from the accuracy.

74 **2.1. Pollen holography and fluorescence dataset**

75 The SwisensPoleno Jupiter measures particles in flight, in the size range from 0.5 to 300 μm , as they pass through the
76 instrument. When a particle triggers the detector, holographic images are taken by two cameras which are both orthogonal to
77 the direction of flight and at 90° to each other. These images are greyscale with a resolution of 200 by 200 pixels after numerical
78 reconstruction and cropping, with each pixel representing a square of $0.595 \times 0.595 \mu\text{m}$ in the physical domain. Right after the
79 holographic images, FL is measured using the Laser Induced Fluorescence (LIF) method. FL is then sequentially induced by
80 three excitation sources and captured in five different wavelength channels, for a total of 15 measured FL intensities. For each
81 source, the FL is induced by shooting at the particle at the moment it passes the detector and the FL subsequently emitted by
82 the particle is captured by Silicon Photomultipliers (SiPM). The FL lifetime is also measured but is not used in the present
83 work. The three different excitation wavelengths are 405, 365, and 280 nm, while the reception wavebands are 333-381, 411-
84 459, 465-501, 539-585, and 658-694 nm. In the following, we will refer to each waveband by its central wavelength, i.e., 357,
85 435, 483, 562, and 676 nm. Note that the first measurement channel is saturated by scattered light when the 365 nm excitation
86 source is activated. Also, for single-photon excitation, we expect to measure no signal in the first measurement channel when
87 the 405 nm source is active. This effectively reduces the useful intensity measurements to 13. The FL data requires additional
88 pre-processing to simplify its usability and improve robustness. More details on these steps are provided in Section 2.2. Finally,
89 the SwisensPoleno Jupiter also performs polarised scattered light measurements, which are however not used in the present
90 work. We therefore limit the analysis to characterisation of particle morphology using digital holography and chemical
91 composition with FL intensity measurements. From hereon, we refer to the set of holographic images and FL measurements
92 for each individual particle as "an event". A more extensive description of the data collection process is provided in Sauvageat
93 et al. 2020.

94
95 This study is based on a pollen dataset created by aerosolising freshly collected pollen at the Swiss Federal Office of
96 Meteorology and Climatology MeteoSwiss (hereafter MeteoSwiss) station in Payerne, Switzerland. In total, the dataset
97 consists of measurements from 57'300 pollen grains distributed among seven different wind-pollinated and allergy relevant

98 plant taxa as reported in Table 1. For simplicity, we will refer to these taxa also as "classes" and only the genus name will be
 99 used to refer to each of them. In Figure 1, we present examples of reconstructed images for the different classes considered in
 100 this work. To compare results across different instruments (of the same type), all measurements were performed using two
 101 SwisensPoleno Jupiter systems denoted P4 and P5. The counts for each pollen taxa and SwisensPoleno are also given in Table
 102 1.

103
 104 The pollen samples were collected from a single tree for *Alnus*, *Betula*, *Corylus*, *Fagus*, and *Quercus*, from two different trees
 105 for *Fraxinus*, and from a few neighbouring stems for the grass *Cynosurus*. After ~~outdoor~~ collection, pollen was brought to the
 106 outdoor measurement site and aerosolised. This was achieved using a SwisensAtomizer which disperses particles using a
 107 vibrating membrane and an airstream. Samples are thus scattered in a chamber and drawn into the instrument, producing a
 108 regular flow of pollen grains. To prevent the pollen from drying out, plants that were not more than 15 km away from the
 109 MeteoSwiss station were selected, which means it was possible to aerosolise samples soon after collection (usually within one
 110 hour). Pollen samples were analysed using two instruments one after another implying a time lag between the data for P4 and
 111 P5, which ranges from just 35 minutes for *Alnus* to 80 minutes for *Quercus* (the mean time lag is 60 minutes). For *Fraxinus*
 112 there is no such lag since the data come from two different samples that were measured on different days. Datasets for all the
 113 considered pollen taxa were created in 2020, except *Alnus* and *Corylus* which are from early 2021.

114

<u>Class E</u> (common name)	Latin scientific name	Number of events for P4	Number of events for P5
Alder	<i>Alnus glutinosa</i>	8416	2643
Birch	<i>Betula pendula</i>	6128	5458
Hazel	<i>Corylus avellana</i>	4714	4444
Crested Dog's-Tail (Grass)	<i>Cynosurus cristatus</i>	5895	2117
Beech	<i>Fagus sylvatica</i>	2178	2827
Ash	<i>Fraxinus excelsior</i>	2557	4837
Oak	<i>Quercus robur</i>	3036	2050
	TOTAL	32924	24376

115 **Table 1: Distribution of pollen counts per taxa and Poleno.**

116

117 2.2. Data pre-processing

118 The datasets required to train the algorithms were generated as follows. First, the holographic data for each class were cleaned
 119 to eliminate any non-pollen events or events associated with other pollen taxa. This was achieved with additional filters on
 120 shape properties (image features computed after binarisation as described in Sauvageat et al. 2020), which were appropriately
 121 selected for every class by heuristic visual inspection of the holographic images. Thereafter, for each event the background

122 signal caused by scattered light was subtracted from the raw FL measurement. This background especially disturbs the low FL
123 intensity measurements where the scattered light dominates relative to the particle signal. The background signal was obtained
124 by conducting measurements with no particles present in the measurement chamber, leaving just the scattered light induced
125 by the excitation source. If the subtraction caused the final signal to be negative due to noise, the resulting value was set to
126 zero to avoid numerical instabilities that our ML model would not be able to deal with. Finally, since the absolute FL
127 compensated by the scattered light is still dependent on the measuring system, the particle size, and the particle position within
128 the measurement volume, we transformed it into relative FL. Namely, the relative fluorescence intensity r_{ij} for measurement
129 channel i and excitation source j is obtained by dividing the absolute FL intensity a_{ij} by the sum of the FL intensities on all
130 channels k for the same excitation source j :

$$r_{ij} = \frac{a_{ij}}{\sum_k a_{kj}}$$

132 Using relative FL, although we lose the absolute FL intensities, allows measurement systems to be compared without specific
133 data modification. The inter-compatibility aspect is especially important when considering a measurement network. Thanks to
134 this standardisation, the same algorithm can be used for all systems in the network rather than adjusting the classification
135 algorithm individually for each measurement system.

136 **2.3. Data exploration**

137 Before applying any ML algorithm, it is important to explore the data to better understand their characteristics. In the following,
138 the distributions of the various holographic image features as well as typical relative FL spectra for the different pollen types
139 are investigated. We also explore the structure of the data using dimension reduction.

140
141 To get other characteristic features from the reconstructed holographic images, further image processing steps are conducted
142 using the Python package "Scikit" (Van der Walt et al. 2014). Physically-based particle features, such as the minor and major
143 axes, the area, the eccentricity, and the particle brightness (mean intensity of the pixels reproducing the particle) are computed
144 for each image separately. Other statistics were calculated based on image features, e.g., the equivalent area diameter defined
145 as the diameter of a circle with the same area as the particle. The distribution of these features for each pollen class and each
146 measurement system were analysed separately and are presented in the Results section.

147
148 As previously discussed, alongside the holography images, relative FL spectra are used for enhanced characterisation of the
149 pollen grains. During data exploration, we observed inconsistent results for the 405 nm laser excitation, which upon further
150 inspection revealed a misalignment of this laser in one of the measurement systems. For this reason, we will only use the 280
151 nm and 365 nm excitation throughout the rest of the present work. The distributions of the valid FL spectra are presented and
152 discussed in the Results section.

154 As a way to explore all the features of the dataset at once, we performed dimensionality reduction. We used the Uniform
155 Manifold Approximation and Projection (McInnes et al. 2018~~20~~), called UMAP, on the input data of each model (holo, FL
156 and combined). This technique allows us to plot multidimensional data as points on a plane, therefore it gives an insight on
157 how similar/different data points are depending on how far from another they are in the plane.

158 **2.4. Machine Learning Model**

159 To handle the classification task, we randomly split the data into training (75%) and test (25%) sets and chose a multi-layer
160 "deep" artificial neural network to learn how to identify pollen grains based on the training set. This network maps input data
161 from the holographic images and relative FL spectra to the different pollen classes. The full network, built using the ML
162 framework Keras (Chollet et al. 2015), is shown in Figure 2. To handle the image input, an EfficientNet B0 model pre-trained
163 on ImageNet is used (Tan et Le 2019). It achieves state-of-the-art performance for classification tasks. For treating the spectral
164 information, a single fully connected (denoted FC hereafter) hidden layer with 255 neurons is used. As a pre-trained model,
165 the parameters of EfficientNet B0 are frozen and therefore not modified in the training on the pollen dataset. However, the
166 parameters of the layers after it are optimised according to the training data. The results of the two feature extraction networks
167 are concatenated, then dropout is added and finally the result is passed to the decision layer. The width of this FC decision
168 layer matches the number of classes (seven in this case). Lastly, the output is normalised by a softmax layer to obtain a
169 probability distribution. To compute the loss, we used the cross-entropy function between the predicted and reference classes.
170 To ensure a fair comparison, each model was trained for exactly 200 epochs. In training runs where only images or only relative
171 FL spectra were used, the path not used was removed from the model graph (Figure 2). The figure shows the model with both
172 features active.

173 The models were evaluated using a test set consisting of 25% of the data from both instruments, sampled randomly. We used
174 balanced accuracy, F1 score and Matthew's Correlation Coefficient (MCC) as metrics to assess the model performance. For
175 accuracy, the corresponding confidence intervals were calculated via normal approximation as explained in Raschka 2020. It
176 is important to note that the model used here is a baseline and has not undergone hyper-parameter optimisation, therefore no
177 validation set has been defined in order to keep a maximum of data for training. This means that a degradation of scores is
178 possible when applying the model to operational data as all sorts of pollen taxa can be encountered considering that other
179 particles are filtered out before the classification. Nonetheless, the present study does not aim to provide an operational model
180 but simply investigate the potential of using FL as a complement to holography for single particle identification.

181 **3. Results**

182 **3.1. Feature observations**

183 Important observations can already be made by looking at basic geometrical features derived from holographic images. As an
184 example, we consider the distributions of equivalent area diameter and eccentricity in Figure 3 (a) and (b). Note that for

185 geometrical features, the value associated with each particle is the largest result obtained for the pair of holographic images.
186 Regarding the equivalent area diameter, its distribution provides information about the size of the pollen grains for a given
187 class. As illustrated in Figure 3 (a), *Fagus* pollen grains are typically large with a maximum equivalent area diameter of 45-55
188 μm , which corresponds to the literature (Halbritter et al. 2021) and is clearly ~~superior-larger than~~ all other classes we
189 considered in our study. Conversely, the distribution of the eccentricity gives an insight on how round the pollen grains are. In
190 that case, *Cynosurus* pollen grains have the roundest shape with a maximal eccentricity between 0.4 and 0.55 (0 representing
191 a circle and 1 an ellipse), whereas *Quercus*' values are in the range 0.8-0.9 due to its more elliptical shape. These characteristics
192 can also be observed on the holographic images in Figure 1. While the eccentricity is used to give a hint on the symmetry of
193 the pollen grain, further metrics could be introduced to further quantify symmetry. This was not implemented in the present
194 study as feature extraction is done automatically by the convolutional neural network.
195

196 The distributions of the relative FL spectra allow us to identify some classes that have distinct FL signatures. Figure 3 (c) and
197 (d) show the distribution of the relative FL for the two excitation-emission combinations where the differences between taxa
198 are the largest. The excitation sources are at 280 and 365 nm with emission channels at 357 and 435 nm respectively. In Figure
199 3, we observe, for both plot (c) and (d), clear differences in relative FL for *Cynosurus*, which presents considerably higher
200 values compared to the other taxa. In addition, differences between instruments show that P4 and P5 have similar
201 measurements in the 280/357 nm but P5 has significantly lower measurements for *Corylus* and *Cynosurus* in the 365/435 nm.
202 Overall, all combinations of excitation sources and emission channels provide relevant information for pollen characterisation
203 and the ones presented in Figure 3 (c) and (d) represent well the type of patterns that can be observed.
204

205 Finally, the UMAP plots, given in the left column of Figure 4, show how different or similar are the image and FL features of
206 each taxon. We observe a clear distinction based on morphology (Figure 4 (a)) for *Fagus* and *Quercus*, with *Cynosurus* also
207 having only little overlap with *Corylus*. However, the latter and especially *Betula* and *Alnus* are clearly mixed up. In Figure 4
208 (b), the UMAP on FL spectra does not exhibit the same group structure as for morphology. Here, *Fagus* and *Cynosurus* are
209 plainly detached from the remaining groups which are themselves imbricated. Ultimately, all groups are fully separated when
210 building the UMAP on both morphology and FL features. We observe a correspondence between the separation of groups on
211 the UMAPs and the capacity of the ML model to classify those classes correctly.

212 3.2. Classification performance

213 The classification results for each model are given as confusion matrices in Figure 4 and summarised in Table 2. We observe
214 in these results that the holo model globally performs better than the FL model when training on a single modality. The FL
215 model indeed encounters difficulties distinguishing some classes such as *Quercus* and *Fraxinus* or *Betula* and *Corylus* (Figure
216 4 (b)) which exhibit similar relative FL spectra. When considering the morphology of *Quercus* and *Fraxinus* (Figure 3 (a) and
217 (b)), it is not surprising that the holography model performs better at differentiating these classes as they present significantly

218 distinct shapes. As the performance for the single-input models here is already (very) high, minor dips in performance can
 219 make a notable difference. Combining holography and FL improves the performance compared to the single input models for
 220 every taxon considered, except for *Fagus* and *Cynosurus* that already obtain perfect scores with single input models. The
 221 performance gain is noteworthy as the combined model achieves an overall balanced accuracy of 99.2% compared to either
 222 96.8% or 87.8% for the individual holography or FL models respectively. As a complement, the confidence intervals associated
 223 with the accuracy of each model for each taxon are displayed in Figure 5. The non-overlapping of the confidence intervals
 224 indicates a statistical difference between accuracies. The combined model outperforms both single-input models for five of the
 225 seven taxa, namely, *Alnus*, *Betula*, *Corylus*, *Fraxinus* and *Quercus*. Thus, logically, the balanced accuracies of the holo and
 226 FL models are significantly lower than that of the combined model (see Table 2). It follows that the absolute error rates, defined
 227 as 1 minus the accuracy, of the holo- (3.2%) and FL-only (12.2%) models are respectively 4 and 15 times higher than that of
 228 the combined model (0.8%). This indicates that mistakes in particle identification occur for roughly 3 particles over 100 for
 229 the holo model, 12 particles over 100 for the FL model but less than 1 particle over 100 for the combined model.
 230

Model	Balanced accuracy	F1-score	MCC
Holography only	0.968, [0.965; 0.970]	0.964	0.958
FL only	0.878, [0.874; 0.882]	0.890	0.874
Combined	0.992, [0.991; 0.993]	0.992	0.991

231 **Table 2: Classification performance of each model. The balanced accuracy, with its associated 95% confidence interval,**
 232 **represents the average of the recalls (ratio of correct prediction over total countevents for each class), ranging from 0**
 233 **to 1. The F1-score is the harmonic mean of the precision and recall, ranging from 0 to 1 and MCC stands for Matthew's**
 234 **Correlation Coefficient and is a robust metric for classification performance, ranging from -1 to 1.**

235 4. Discussion

236 The results based on manually generated data show that combining FL with holography leads to a substantial identification
 237 performance gain. The differences between the combined model accuracy and both single-input models confirm the findings
 238 from the UMAPs. This demonstrates that by combining the two inputs, the complementary morphological and biochemical
 239 properties of pollen grains can be used for a better classification. Although it seems small, the gain in accuracy is important
 240 for the field of aerobiology and specifically pollen monitoring since pollen grains only represent a minor part of all the particles
 241 in the air. Since pollen concentrations typically range from a few grains (< 10) to a few hundred grains per cubic metre, and
 242 the thresholds for allergy symptoms are usually around a few tens of grains per cubic metre (Gehrig Bichsel et al. 2017;
 243 Pollen.lu 2003), misclassifications can have an impact on the information provided to allergic people. Above all, high
 244 identification accuracy is particularly important for plants with highly allergenic pollen such as *Ambrosia artemisiifolia*
 245 (common ragweed) as a few grains are sufficient to cause allergy symptoms.
 246

247 Not only is the combined model's accuracy superior to the other models, but this gain is specifically important for some key
248 pollen taxa. Indeed, the group of *Alnus*, *Betula* and *Corylus*, all from the Betulaceae family, is known to be difficult to classify
249 accurately and presents a very high allergic potency with possible cross-reactivity in central and northern Europe (Puc et
250 Kasprzyk 2013). Thus, the excellent classification performance obtained here opens the gate towards better monitoring by
251 using holography together with fluorescence data. In addition, the consistent FL signal in between instruments and the available
252 excitation sources and measurement channels characterise single pollen grains precisely even though the 405 nm excitation
253 source was set aside. Also, the combinations of excitation and emission wavelengths used in the Poleno correspond to the most
254 prominent fluorescence modes for a variety of dry pollen studied in (Pöhlker et al. 2013). Then, ~~the coherence~~ between our
255 results and those from Pöhlker et al. 2013 of the fluorescence spectra obtained here with the measurements of the latter study,
256 ~~brings confidence into our measurements and the measurement instrument~~ stability of the Poleno per se. In future work, the
257 405 nm excitation source needs to be included to verify its potential for improvement.

258
259 When working with images, choosing neural networks for classification is the obvious solution to be sure not to lose
260 information by using the image itself as input. However, the discrimination of pollen taxa using the UMAP dimension
261 reduction method shows that working with features derived from the holographic images is also a possibility for pollen
262 classification. Future work testing other machine learning methods on image features and fluorescence spectra needs to be
263 conducted as other classifiers may perform similarly while being cheaper in terms of computational resources. In addition, the
264 main limitation of this study, focusing on a reduced number of pollen taxa and manually generated data, should be overcome
265 in following work by gathering more data to train a broader model and test it on operational data.

266
267 In the end, we expect the benefit of combining holography with FL measurements for pollen classification to have a positive
268 impact on the capacity of models to discriminate different pollen taxa. Moreover, in an operational setup, the benefit of using
269 FL in addition to holography could be even higher as it would allow for an easy distinction between biological and non-
270 biological particles (e.g. water droplets, sand particles or dust) assuming that they do not fluoresce. Yet, the extent of the gain
271 in the real case scenario remains to be quantified as the dataset used in this study probably does not catch all the environmental
272 variability. For example, in ambient air, pollen can break into fragments also impacting allergy sufferers but not currently
273 monitored.

274 5. Conclusion

275 The present study demonstrates the ~~benefit-potential~~ of using FL measurements as a complementary input to holographic
276 images for single-grain pollen identification using the SwisensPoleno and ML algorithms for the most important allergy
277 causing pollen taxa in Central Europe. The capacity of the ML model to identify pollen grains depends on both inputs and they
278 compensate each other when one does not provide enough information for accurate identification. As a result, the performance

279 of the combined model is systematically higher than either of the models trained with a single input. The restricted and
280 manually created dataset used in this study has several limitations, but it still provides strong evidence for the complementary
281 role of FL and holography.

282

283 In conclusion, we recommend the use of relative FL as a secondary input for automatic pollen identification using the
284 SwisensPoleno Jupiter. In this study, we tested its contribution on a restricted dataset, showing that the contribution of FL is
285 of great value for operational networks where similar pollen taxa can be encountered. Finally, the use of relative FL for
286 automatic pollen identification further opens the door towards a larger and more precise monitoring of bioaerosols. For
287 example, objects which are challenging to identify using holographic imaging only, such as fungal spores, could be added to
288 the panel of particles.

289

290 **Author contribution**

291 EG, SE and YZ conducted the study and contributed equally as main authors. SL guided the machine learning aspects and
292 supervised PW in his work on the relative fluorescence. AB, BCl, GL and FT contributed to writing, and BC_r supervised the
293 study and contributed to writing.

294

295 **Competing interests**

296 EG and YZ are employees of Swisens AG, and AB is a member of the editorial board of AMT. The investigations were carried
297 out in compliance with good scientific practices and the declared relationships have no effect on the results presented. The
298 peer-review process was guided by an independent editor, and the authors have also no other competing interests to declare.

299

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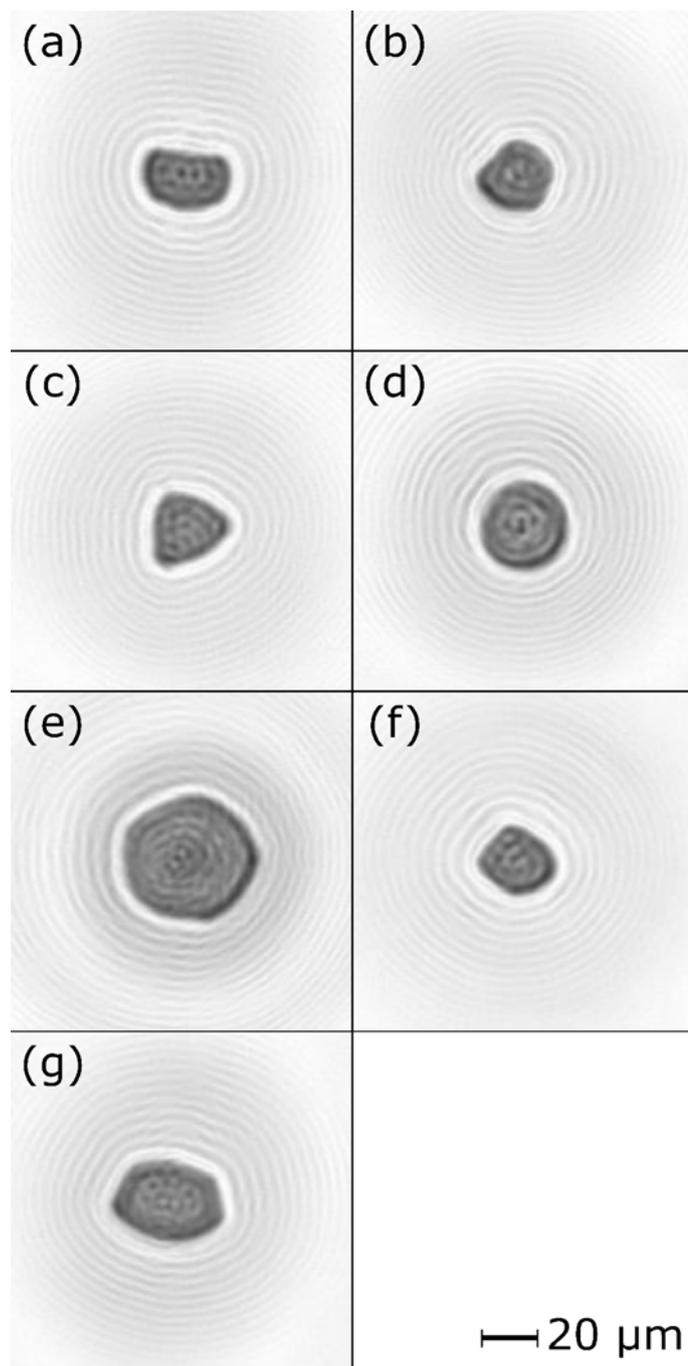
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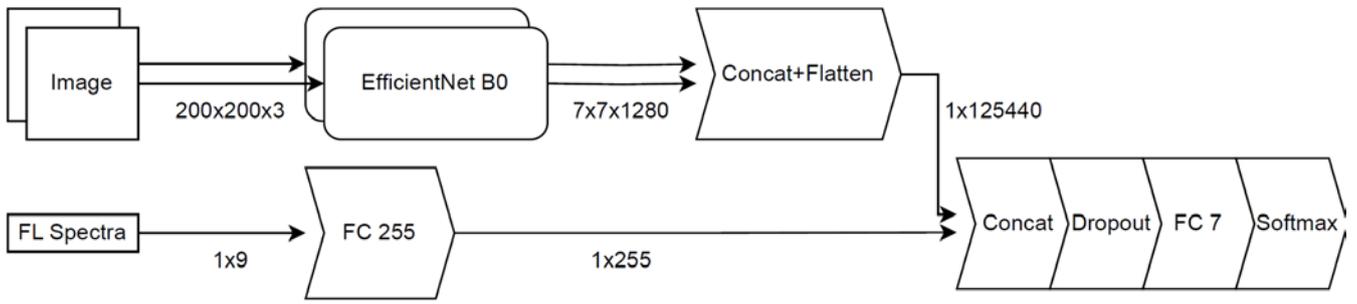
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406 **Figure 1: Holographic images of pollen after numerical reconstruction: (a) *Alnus glutinosa* (b) *Betula pendula*, (c)**
407 ***Corylus avellana*, (d) *Cynosurus cristatus*, (e) *Fagus sylvatica*, (f) *Fraxinus excelsior*, (g) *Quercus robur***



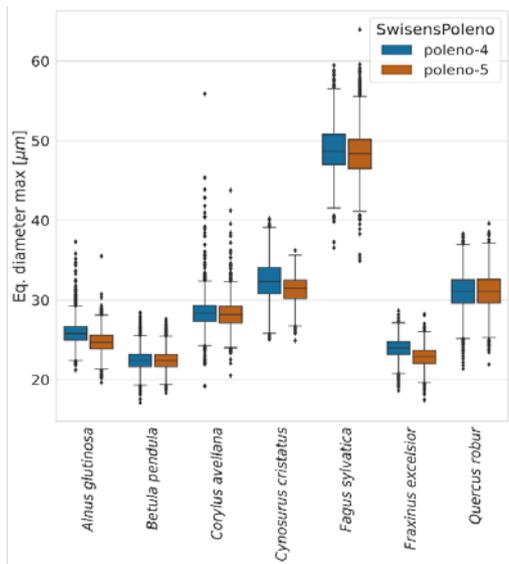
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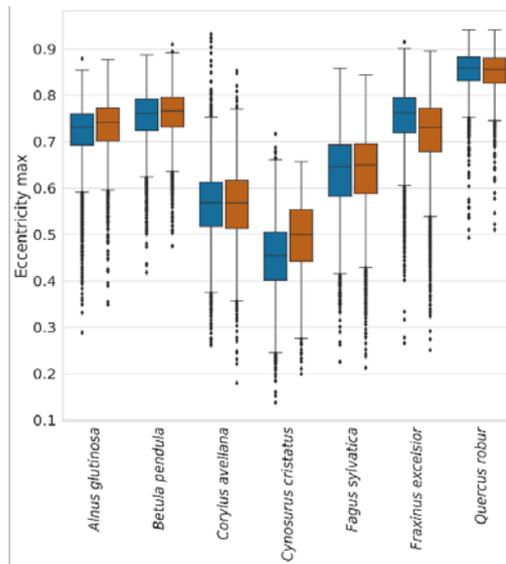
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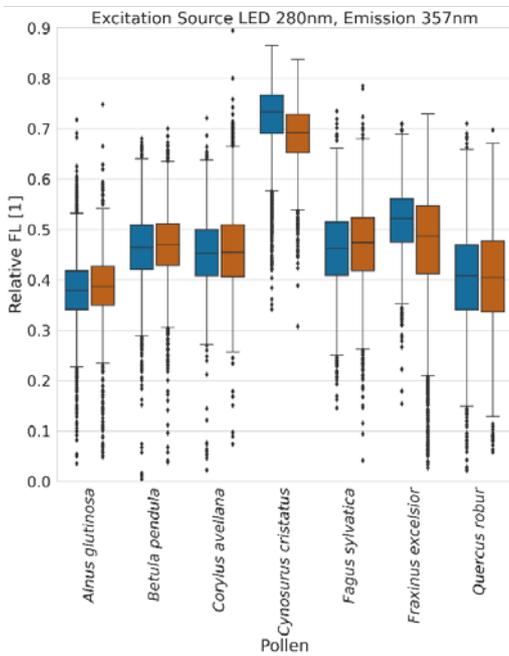
Figure 2: ML model structure used to classify the pollen data. The top path handles the holographic image data while the bottom path processes the relative FL spectra data. The numbers on the connecting lines denote the dimensions of the data.



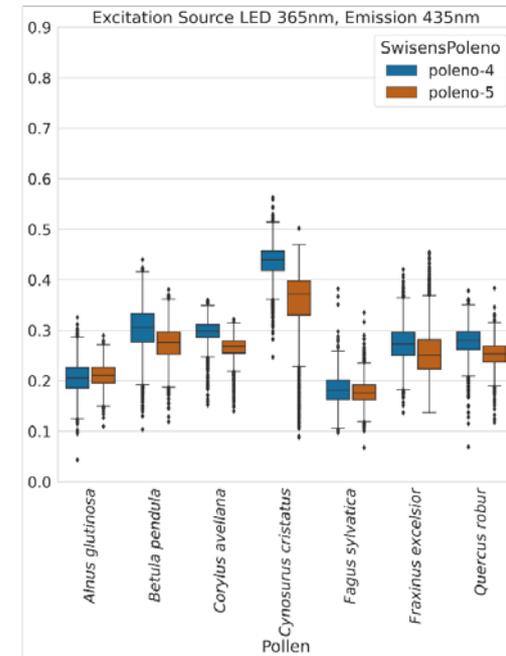
(a)



(b)



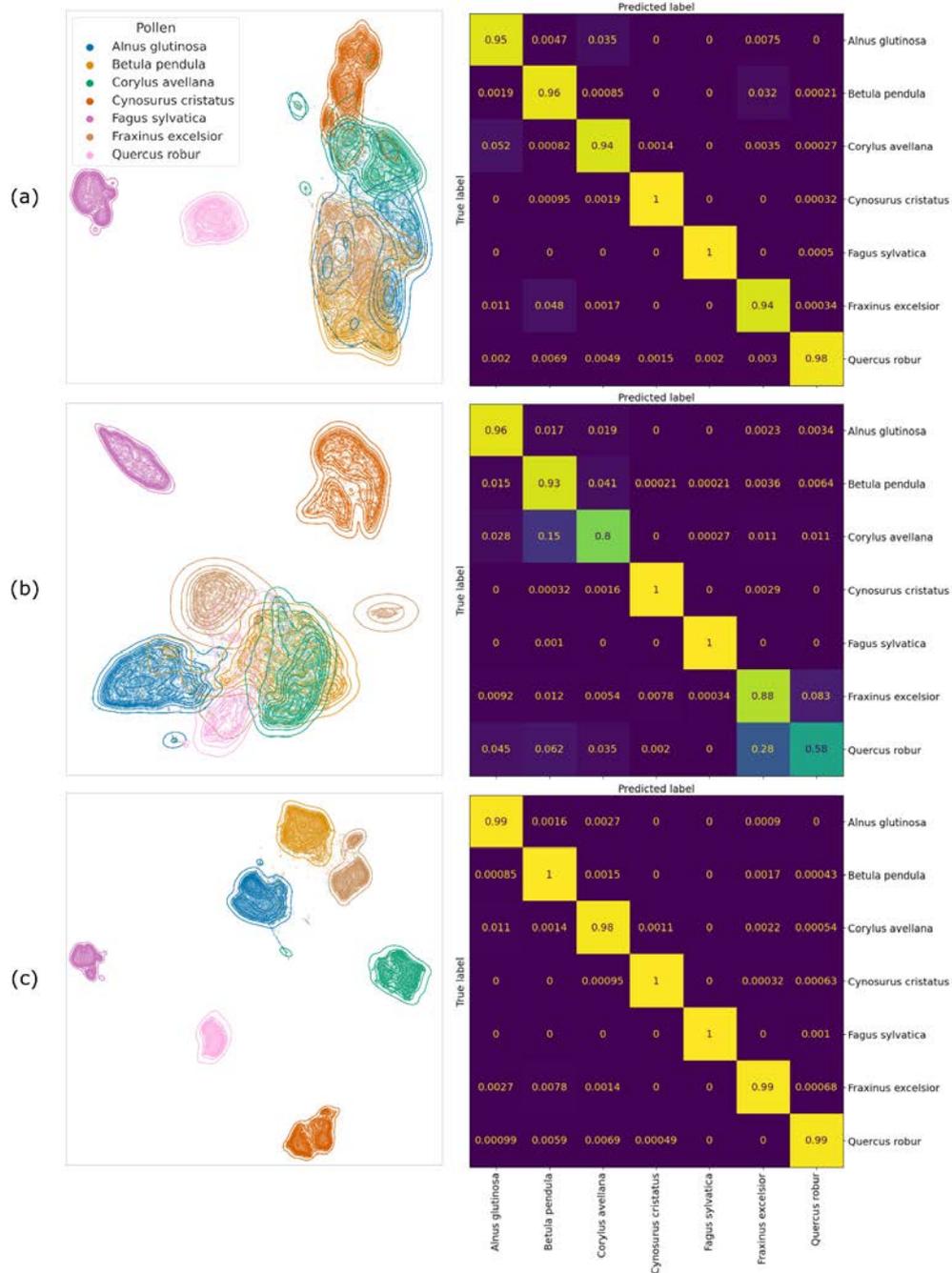
(c)



(d)

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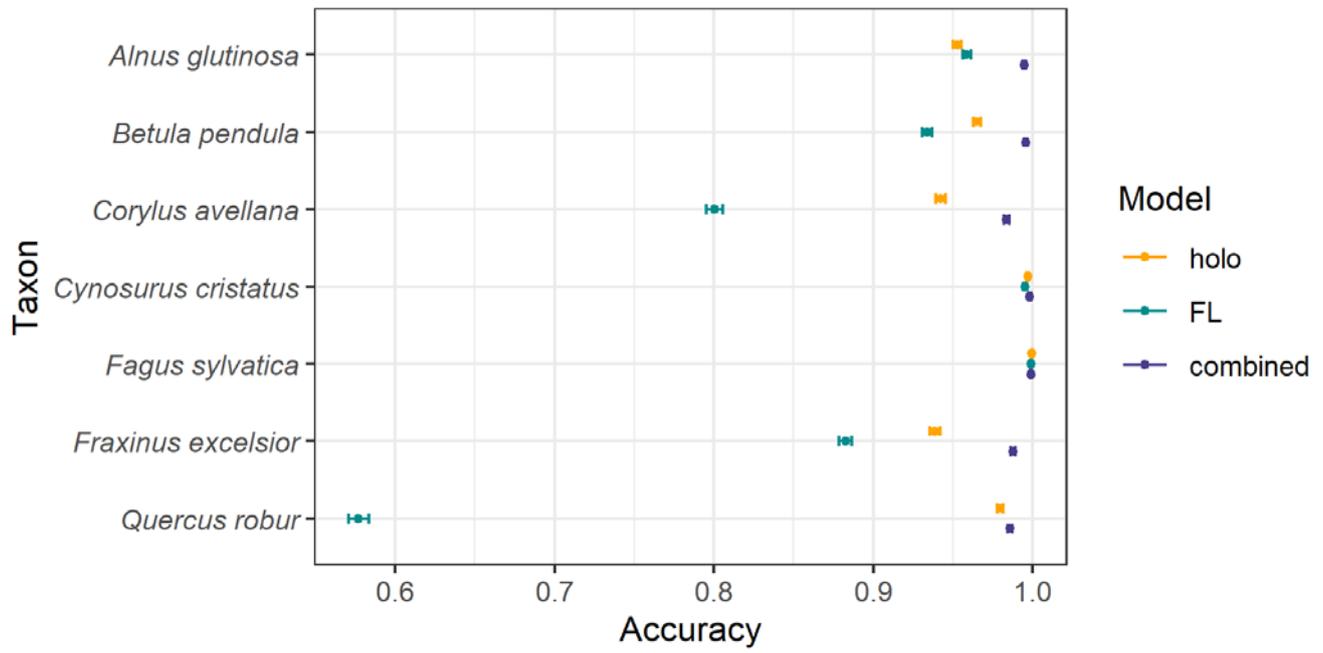
413 **Figure 3: Distribution of holographic image features (upper plots) and relative FL (bottom plots) for each pollen class**
 414 **and measurement system. (a) Maximum equivalent area diameter in μm , defined as the diameter of a circle with the**
 415 **same area as the particle, (b) Maximum eccentricity, defined as the deviation of the ellipse fitted to the particle from a**
 416 **perfect circle, ranging from 0 for a circle to close to 1 for an ellipse. (c) Measured relative FL intensity with 280 nm**
 417 **excitation and detector with centre wavelength 357 nm and (d) with 365 nm excitation and detector 562 nm.**



418

419 **Figure 4: Left side: Uniform Manifold Approximation and Projection (UMAP) of particle event features (morphology**
 420 **or/and FL features) of all the data. Right side: Confusion matrices indicating the performance of each model on the**
 421 **test set. Line (a) holography only, line (b) relative FL only and line (c) combined relative FL and holography. UMAP**
 422 **settings: neighbours = 15, minimum distance = 0.001, random state = 42.**

423



424 **Figure 5: Accuracy of each model for each taxon. The error bars represent the 95% confidence intervals.**