

Opinion: Stratospheric Ozone – Depletion, Recovery and New Challenges

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Abstract. We give a personal perspective on ~~recent current important and open~~ issues related to the depletion of stratospheric ozone and ~~some discuss related well-established and~~ newly emerging challenges. A common theme for these issues is the need for continued global observations of the stratosphere, coupled with process-based laboratory and modelling work. We first provide a brief review of historic work on ~~understanding the ozone layer where we highlight some work from stratospheric ozone research with a focus on the contributions of~~ the late Paul Crutzen as ~~a contribution to the part of this~~ special issue in his honour. We then review the status of ozone recovery from the effects of halogenated source gases and discuss the undoubted ~~continued~~ effectiveness of the Montreal Protocol ~~and its challenges from despite some recent~~ renewed production of controlled substances and ~~substantial abundances of some~~ short-lived uncontrolled substances. ~~We then discuss~~ The increasing atmospheric abundance of greenhouse gases such as carbon dioxide (CO₂), methane (CH₄) and nitrous oxide (N₂O) have a large potential to perturb the ozone layer. This is partly through radiative-dynamical effects which could cause a change in the speed of the stratospheric circulation. While the atmospheric chemical impacts of N₂O and CH₄ are well understood, there is uncertainty in their future evolution and hence the overall effect on the ozone layer. We summarise, in some detail, the recent ~~observations newly observed phenomenon~~ of ozone depletion through injection of smoke particles ~~from following the~~ Australian ~~fires wildfires~~ in early 2020. Further ~~unexpected~~ perturbations to the ozone layer are currently occurring at the moment through ~~the unexpected~~ injection of very large amounts of water vapour (and some sulphur dioxide) from the Hunga Tonga-Hunga Ha'apai volcano in January 2022. ~~We conclude with some thoughts on the urgent~~ The open research questions in these topics illustrate the critical need to ensure continuity in at least maintain, if not expand, the observational network and therefore to address the impending 'satellite data gap' in global, height-resolved observations and on how to exploit ever more complex and expensive of trace gases and aerosols in the stratosphere. This data gap will, in effect, make us largely blind to the stratospheric effects of, for instance, large wildfires and volcanic eruptions in the near future. Complex Earth System Models (ESMs) are being constructed by groups worldwide for coupled climate projections with the stratosphere as an important component. However, it is essential that the huge computational requirement of these models. Overall, does not result in an oversimplification of many stratospheric processes needed for an accurate projection of the ozone layer. Regardless of the complexity of ESMs, a hierarchy of process models will continue to be important tools for testing and developing our evolving understanding of the ozone layer and for providing policy-relevant information. In summary, study of the stratospheric ozone layer continues to produce uncover novel research challenges and to reveal yet more processes and mechanisms that threaten can perturb this essential component of the Earth system. Global monitoring of stratospheric ozone and of its gaseous and particulate drivers, combined with numerical modelling, remains absolutely vital if these challenges are to be met including providing reliable future projections.

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1 Introduction

Depletion of the stratospheric ozone layer has been a major environmental issue of the past few decades, especially since the discovery of the Antarctic ozone hole in 1985 (Farman et al., 1985). The observed depletion at middle and high latitudes has been caused by increasing abundances of chlorine and bromine species, which are derived from long-lived surface-emitted halogenated gases, so-called ozone-depleting substances (ODSs) (see e.g. Solomon, 1999). A primary reason for concern is that the ozone layer prevents harmful, biologically damaging ultraviolet (UV) radiation (wavelengths below about 300 nm) from reaching the surface. This UV radiation can, among other impacts, cause skin cancer in humans and can be damaging to plants (Barnes et al., 2019). Ozone not only absorbs UV radiation, heating up the stratosphere, but also interacts with terrestrial infrared (IR) radiation (e.g. Riese et al., 2012). As such, it plays a key role in determining the temperature structure of the atmosphere. Hence, changes in the ozone layer can also affect surface climate, and moreover the long-lived ODSs, such as chlorofluorocarbons (CFCs), themselves are also potent greenhouse gases (Velders et al., 2007).

The Montreal Protocol on Substances that Deplete the Ozone Layer was signed in 1987 and ratified two years later. With several subsequent amendments, the Protocol now controls (limits) the production and consumption of all major long-lived ODSs, which are ultimately emitted to the atmosphere. The atmospheric abundances of these species have responded to these controls; the stratospheric levels of chlorine and bromine peaked in the 1990s and are now slowly declining (e.g. Newman et al., 2007; Engel et al., 2018). In consequence, an increase ('recovery') of stratospheric ozone has been detected in the upper stratosphere and the Antarctic, although the signal is currently small and is difficult to separate from other atmospheric influences (e.g. Chipperfield et al., 2017). Nevertheless, the Protocol can therefore be considered on track in its aim of protecting the ozone layer from the effects of halogenated ODSs (see Section 3). A common measure of recovery is the date at which stratospheric ozone values are predicted to return to 1980 levels, before the occurrence of large depletion. This return will also be affected by other factors other than ODSs, notably climate change (see Section 4). Models predict that this will occur around the middle of this century (e.g. Dhomse et al., 2018), although there are limitations using this simple measure of the timing of a specific event for quantifying the ongoing process of recovery (e.g. Pyle et al., 2022). Accordingly, the Montreal Protocol (MP) is arguably the most successful international environmental treaty to date. However, recent discoveries related to increased emissions of controlled ODSs (Monzka et al., 2018) and uncontrolled shorter-lived halogenated source gases (e.g. Hossaini et al., 2017) had raised some concerns on the continued success of the treaty and the outlook for ozone recovery. However, the success of dealing with the CFC-11 issue (see Section 3.1) has demonstrated the resilience of the protocol, the effectiveness of its provisions, and the importance of continued vigilance regarding atmospheric trace gases.

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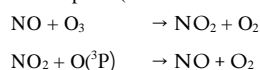
This Opinion paper gives our personal view of some selected current long-standing and future recently emerging issues in ozone layer science. It is not a review of the subject; there are many excellent text books and the 4-yearly WMO/UNEP assessments (e.g. WMO, 2022) which serve that purpose. Section 2 gives a brief summary of ozone layer research, with emphasis on the contribution of Paul Crutzen, to whom selected papers in this issue are dedicated. Section 3 addresses the long-standing issue of ozone depletion and the role of driven by halogenated species. Section 4 discusses the impact increasing greenhouse gas loadings on stratospheric ozone and the new research areas of wildfire smoke and the ozone layer, and the expanding topic of volcanic impacts. Sections 5 and 6 give some personal thoughts on issues related to the availability of observations necessary to follow the evolution of the ozone layer and understand its changes. Section 6 discusses the range of modelling tools available, some further developments that are still needed, and how these tools can be best employed. Finally, a summary and outlook is provided in Section 7.

2 A Century of Ozone Layer Research

Active research into stratospheric ozone dates back around 100 years. Dobson pioneered the detection and quantification of ozone in the stratosphere using a UV spectrometer (Dobson and Harrison, 1926) following earlier work by Fabry and Buisson (see historical summary in Brasseur, 2020). A theoretical model for creation of a stratospheric ozone layer, based solely on oxygen chemistry, was first proposed by Chapman (1930). This was ~~formulated~~ based on the slow production and destruction of 'odd oxygen' ($Ox = O_3 + O(^3P)$) along with fast interconversion of O_3 and $O(^3P)$ within the Ox family. This oxygen-only model appeared to suffice until the 1960s when improved observations and laboratory measurements of key rate coefficients revealed a major quantitative discrepancy. The Chapman cycle included the only significant chemical source of Ox, i.e. photolysis of O_2 , but ignored around 80% of stratospheric Ox loss via catalytic cycles that destroy ozone through reactions involving HOx: (e.g. Nicolet, 1970), NOx and halogen radicals.

Here, ~~as part of this special issue~~, we ~~should~~ highlight the contribution of the late Paul Crutzen (1933 – 2021) to ozone layer science. For a comprehensive summary of his whole career please see Müller et al. (2022), and references therein, ~~and~~ Solomon (2021a) ~~and~~ Fishman et al. (2023). Paul Crutzen started contributing to our understanding of the ozone layer very early in his scientific career. In 1965, at Stockholm University, Crutzen helped visiting US scientist J.R. Blankenship to develop a numerical model of different forms of oxygen in the stratosphere, mesosphere and lower thermosphere. This marked the start of his scientific career and gave him his first paper (Blankenship and Crutzen, 1965). Following this work Crutzen chose to study for a PhD in stratospheric ozone as it appeared, at that time, to be a topic of "pure science related to natural processes" rather than one about human impact. Clearly, that situation later changed! In due course Crutzen submitted his PhD thesis 'On the photochemistry of ozone in the stratosphere and troposphere and pollution of the stratosphere by high-flying aircraft' to Stockholm University in May 1973.

In his PhD work (Crutzen 1970, 1971, 1972, 1973), Crutzen was the first to suggest that reactions catalysed by NO and NO₂ control the abundance of ozone in the middle stratosphere (around 25-35 km). This is summarised by the cycle:



Where the sum of NO and NO₂ is termed NOx. This discovery was a major achievement and helped to pave the way for a quantitative understanding of stratospheric ozone whereby catalytic cycles ~~due to~~ driven by radical species from various ~~chemical~~ families (HOx, NOx, Clx, Brx) are added to the original oxygen-only model of Chapman (1930). This work formed part of the basis for Crutzen being awarded the 1995 Nobel Prize for Chemistry jointly with Mario J. Molina and F. Sherwood Rowland "for their work in atmospheric chemistry, particularly concerning the formation and decomposition of ozone".

Prior to submitting his PhD thesis Crutzen spent two years (1969-1971) as a visitor to the University of Oxford. Here he developed his ideas on the importance of NOx in controlling ozone in order to address the issue of human-induced perturbations to the ozone layer caused by emissions from high flying supersonic transport (SST) aircraft. The debate on the atmospheric impacts of SST had begun in the early 1970s when it was envisaged that large fleets of around 500 aircraft such as the Anglo-French Concorde might be flown within the lower stratosphere: (e.g. Johnston, 1971). Through his modelling work Crutzen was aware of inherent model uncertainties which prompted him to make the statement that the "minimum requirement is therefore that extensive supersonic air traffic should not take place in the stratosphere before reliable predictions can be made of the possible environmental consequences of such operations" (Crutzen, 1972). This is an insightful lesson that would be equally applicable to many other past and present areas of atmospheric science and therefore one ~~well~~ worth remembering.

120 During the 1970s Crutzen's scientific interests extended into other areas, though he did maintain a link with the stratosphere
121 through the study of the impact of NO_x produced from solar proton events on the ozone layer (Solomon and Crutzen, 1981).
122 He also addressed the budget of stratospheric NO_y (reactive odd nitrogen) from the perspective of surface sources, highlighting
123 the human impact on stratospheric ozone of increased fertilizer and associated increased emissions of N₂O, the main source of
124 stratospheric NO_y (Crutzen and Ehalt, 1977). Following the same reasoning as for N₂O, he also worked out that the dominant
125 non-volcanic source of stratospheric sulfur was the surface emissions of carbonyl sulfide (COS), the long-lived atmospheric
126 sulfur compound (Crutzen, 1976).

127
128 Before his works on N₂O and COS, Crutzen had also keenly followed the publication of ~~the~~ seminal ~~paper~~ papers on
129 chlorofluorocarbons (CFCs) ~~by~~ (Molina and Rowland, 1974) and stratospheric chlorine (Stolarki and Cicerone, 1974),
130 prompting him to publish a contribution on this topic (Crutzen, 1974). Similarly, following the surprise discovery of the
131 Antarctic Ozone Hole (Farman et al., 1985) Crutzen was quick to think about the possible implications of co-condensation of
132 HNO₃ and H₂O (at temperatures above that at which pure ice clouds form) as a mechanism for widespread ~~occurrence~~ formation
133 of polar stratospheric clouds (PSCs) and the initiation of key ozone-destroying halogen chemistry (Solomon et al., 1986) via
134 heterogeneous reactions on PSCs (Crutzen and Arnold, 1986). In this way Crutzen made important scientific contributions to
135 the early research into the causes of polar ozone depletion. His later work on describing the epoch of the Anthropocene still
136 has ongoing relevance to protecting the Earth's ozone layer shield (see Solomon et al., 2021b). His multi-faceted scientific
137 legacy for stratospheric ozone is assured.

139 3 Ozone Depletion and the Montreal Protocol

141 3.1 Montreal Protocol

142 The signing of the Montreal Protocol on Substances that Deplete the Ozone Layer in 1987 and its subsequent amendments
143 have had a major impact on the anthropogenic halogen source to the stratosphere. The Protocol now controls (limits) the
144 production and consumption of all major long-lived ODSs, which are ultimately emitted to the atmosphere. ~~This control~~
145 ~~of~~ Controls on ODS production is leading to have caused a net reduction in the tropospheric source gases (**Figure 1a**) ~~which that~~
146 transport chlorine and bromine to the stratosphere.

147
148 A very important recent development in the Montreal Protocol was the inclusion of hydrofluorocarbons (HFCs) in the Kigali
149 Amendment of 2016 (WMO, 2018). HFCs do not contain any chlorine or bromine and hence do not lead directly to ozone
150 depletion. However, they are potent greenhouse gases and are only present in the atmosphere as replacements for CFCs and
151 HCFCs, hence the need to control these gases and to do so within the MP.

152
153 The majority of long-lived halocarbon source gases are now controlled by the Protocol. Further, or more rapid, reductions in
154 stratospheric chlorine (and bromine) would depend on extension of the Protocol to chlorinated very short-lived substances
155 (VSLS), defined as having an atmospheric lifetime of less than 6 months. The prime example of this is dichloromethane
156 (CH₂Cl₂) (Hossaini et al., 2017) which is mainly of anthropogenic origin and, although largely removed in the troposphere,
157 does deliver a large fraction of the estimated 130 ± 20 pptv of VSLS chlorine to the stratosphere both directly (source gas
158 injection, SGI) and through decay products (product gas injection, PGI) (see WMO, 2022). Although this is only around 4%
159 of the current stratospheric chlorine loading, its contribution is expected to increase (Section 3.3).

160
 161 The history of the MP since its signing in 1987 (and ratification in 1989) ~~has been very successful~~ is one of continued success
 162 – as evidenced by the decreasing loading of ODSs and stratospheric chlorine and bromine; (WMO, 2022). Indeed, the former
 163 UN Secretary General, Kofi Annan, described the Protocol to be not only “the most successful environmental treaty in history”,
 164 but also “perhaps the most successful international agreement to date” of any kind. However, that success appeared to be
 165 challenged for the first time by the observation of an unexpected ~~reductions~~ lowdown in the atmospheric CFC-11 decay ~~rate~~
 166 (Montzka et al., 2018), which implied renewed emissions. A large fraction (at least) of these emissions (~~at least~~) were traced
 167 to eastern China (Rigby et al., 2019). It ~~should~~ must be emphasised that this detection of apparent contravention of the MP was
 168 only possible through continued observations by the distributed ground-based monitoring networks (see Section 5). Following
 169 this discovery, alarm was raised by policy makers involved in the MP. ~~Updates to process that these renewed emissions could~~
 170 cause a delay in recovery of the ozone layer (e.g. Dhomse et al., 2019). Extensions of these observations for a further three
 171 years later (Montzka et al., 2021; Park et al., 2021) show that these renewed emissions of CFC-11 appeared to have greatly
 172 declined. Therefore, we can argue that this episode has been further evidence of the success of the MP and of the effective
 173 combination of monitoring observations, science and policy. ~~We emphasise that a key component of these interconnected~~
 174 activities is communicating ODS and ozone layer science and findings to policy makers to guide future decision making to
 175 protect ozone and climate. Despite the undoubted progress in our understanding of the atmospheric abundance of the major
 176 ODSs, some important ~~to note that and~~ persistent uncertainties remain ~~in~~. In particular, this is the budget of case for carbon
 177 tetrachloride (CCl₄) ~~another of the major controlled ODS (Sherry et al., 2018) CCl₄), which~~ is produced in large quantities for
 178 feedstock use (e.g. Chipperfield et al., 2020) and ~~it~~ also has soil and oceanic sinks (e.g. Butler et al., 2016). ~~Therefore, it is~~ has
 179 proved challenging to pin down the atmospheric budget of this species and explain the apparently slower atmospheric decay
 180 than expected based on its estimated lifetime (e.g. Park et al., 2018). ~~Continued observations~~ The continued observation of
 181 these controlled ODSs, and further improved understanding of their atmospheric budgets, is important to ensure the
 182 continued ongoing success of the MP.
 183

184 3.2 Ozone Recovery

185 The undoubted success of the Montreal Protocol in halting and turning around the increasing trend in stratospheric chlorine
 186 and bromine is clearly expected to lead to ozone recovery, e.g. an increase in global ozone. However, the detection of ozone
 187 recovery, and even maintaining consistency on the definition of what recovery is within the community, has proven difficult.
 188 There is now a general consensus that recovery means ‘recovery from the effects of depletion caused by halogen (chlorine and
 189 bromine) species’ (e.g. WMO, 2011). Stratospheric ozone amounts clearly depend on many other varying factors (e.g. solar
 190 radiation, temperature, dynamics) which can also lead to an increase or a decrease in ozone-its concentration. These ‘non-
 191 halogen’ influences need to be removed if the ozone recovery from halogens is to be quantified. ~~Therefore~~ Thus, recovery
 192 cannot generally be detected directly from observations of ozone alone; and a statistical or physical model is needed to isolate
 193 the effects of halogen chemistry from other effects in the ozone evolution.
 194

195 Given that recovery is from the effects of halogen-catalysed chemical ~~loss~~ depletion, the clearest signal of recovery ~~would~~ might
 196 be expected in regions where this chemistry exerts ~~a strong~~ the strongest influence on ozone. Newchurch et al. (2003) first
 197 claimed the detection of ozone recovery in the upper stratosphere where the classical ClO + O cycle (Molina and Rowland,
 198 1974; Stolarski and Cicerone 1974) has its maximum efficiency. In this region ~~the contributing~~ there are also non-halogen
 199 effects; in particular the contribution of ozone increase from stratospheric cooling need to be removed, which ~~can be~~ is done
 200 by model attribution studies of the different processes (noting that the stratospheric temperature changes can be due to both ~~to~~
 201 increased longwave cooling by CO₂ and ~~less~~ reduced shortwave heating by O₃ itself). It proved more elusive to detect recovery

202 in the other atmospheric regions subject to large halogen-catalysed loss - namely the polar lower stratospheres. Solomon et al.
 203 (2016) succeeded in detecting Antarctic recovery (or 'healing') by focussing on the period of rapid chemical loss in September,
 204 rather than the period of lowest ozone in October-~~which is subject to saturation of the ozone loss and variability in breakdown~~
 205 ~~of the polar vortex~~. As expected, the larger interannual variability in Arctic ozone loss has made detection of any trends in this
 206 region difficult. However, using long-term ground-based UV-visible observations, Pazmino et al. (2023) recently claimed
 207 some measure of Arctic ozone recovery. ~~These studies show that when searching for the signal of ozone recovery in a variable~~
 208 ~~atmosphere it is important to bear in mind that the different metrics used for the same phenomenon may indicate different~~
 209 ~~behaviours for the recovery.~~

210
 211 At extrapolar latitudes, observations confirm that the ozone decline in 1990s and earlier, caused by increasing atmospheric
 212 concentrations of ODSs, has now transitioned to a slow ozone increase in both hemispheres (**Figure 2**, WMO (2022)). This is
 213 consistent among the ground- and satellite-based measurements and chemistry-climate model simulations in the middle and
 214 upper stratosphere, despite the larger variability of the ground-based measurements. This is apparent in the evolution of
 215 observed and modelled annual mean deseasonalized ozone anomalies, relative to the 1998–2008 climatology ~~off~~ for each
 216 individual dataset in **Figure 2**, in the upper stratosphere (42 km or 2 hPa) and in the lower stratosphere (19 km or 70 hPa).
 217 ~~Ozone~~ Upper stratospheric ozone anomalies averaged over 2017–2020 ~~in the upper stratosphere~~ from most datasets are positive
 218 relative to the 1998–2008 average, which is consistent with expectations from the chemistry-climate model (CCM)
 219 simulations. In contrast, lower-stratospheric ozone anomalies over 2017–2020 continue to be about the same as for the 1998–
 220 2008 average. ~~Interestingly, in~~ 2019 and 2020, stratospheric ozone values were lower than in previous years and below the
 221 level expected from model simulations (Weber et al., 2020). The particularly low 2020 annual mean is the result of a very
 222 weak ~~BDC~~ Brewer Dobson circulation (BDC) and a large and stable Antarctic ozone hole (Klekociuk et al., 2021; Weber et
 223 al., 2021). Such large interannual variability, driven by variations in meteorology and transport (e.g. Chipperfield et al., 2018),
 224 is typical for the lower stratosphere and limits our ability to drawing definite conclusions about long-term trends, especially
 225 for the mid-latitudes (30°–60°) in both hemispheres (see WMO, 2022). Evidently, longer observational time series ~~are needed~~
 226 ~~should~~ reduce the uncertainty due to this variability, again reinforcing the need for continued atmospheric monitoring.

227
 228 While we can see that stratospheric halogen levels are decreasing, and therefore their impact on ozone is decreasing, there are
 229 a number of concerns about the extent and rate of ozone recovery. Clearly, ongoing emissions of chlorine and bromine from
 230 ODSs or VSLS ~~that are not already accounted for~~ will act to slow down this recovery (Sections 3.1 and 3.3). However, there
 231 are other factors which are not controlled by the MP and which may also lead to decreases in column ozone, ultimately the
 232 parameter of primary concern for protecting the biosphere. There are many studies (e.g. Ball et al., 2018, see also **Figure 2**)
 233 which point to an ongoing decrease in ozone in the mid-latitude lower stratosphere. This may be related to dynamical changes,
 234 which are predicted to decrease tropical column ozone in the future (Section 6). The model simulations of Chipperfield et al.
 235 (2018) supported this cause and showed a negligible impact of assumed trends on VSLS bromine and chlorine. In contrast,
 236 Villemayor et al. (2023) have suggested a role for the combined effects of chlorine, bromine and iodine VSLS acting together.
 237 This is a region where further work is needed to determine the extent of ozone depletion/recovery and to quantify its driving
 238 factors.

240 3.3 Other Issues ~~related~~ Related to Halogen Chemistry

241 As noted in Section 2.2, VSLS deliver important amounts of chlorine and bromine to the stratosphere. VSLS bromine is largely
 242 of natural oceanic origin and contributes 5 ± 2 pptv to stratospheric bromine, which is around 27% of the total (WMO, 2022).
 243 There is currently no suggestion of a trend in this VSLS bromine contribution but this could potentially occur due to climate

244 feedbacks on the strengths of the emission sources. In contrast, VSLS chlorine is largely of anthropogenic origin. Although
 245 the total VSLS chlorine injection of 130 ± 20 pptv is only 4% of the total stratospheric chlorine (WMO, 2022), it is showing a
 246 small increasing trend notably through increases in the atmospheric ~~abundance~~ abundances of CH_2Cl_2 and CHCl_3 (e.g. Fang et
 247 al., 2019; Claxton et al., 2020). Far larger local stratospheric chlorine inputs from VSLS have recently been observed in regions
 248 where strong convection and emissions co-locate, notably the Asian Summer Monsoon (Adcock et al., 2020), ~~pointing to the~~
 249 ~~importance of observing chlorine species directly in the lower stratosphere.~~

250
 251 Solomon et al. (1994) pointed out that iodine depletes ozone more efficiently than chlorine, and thus could be responsible for
 252 significant contribution to past and future ozone changes. However, there are still large uncertainties in the main gas- and
 253 condensed-phase iodine photochemical processes (see e.g. Saiz-Lopez et al., 2012; Feng et al. 2023) and observations of
 254 inorganic iodine (Iy) species in the upper troposphere – lower stratosphere (UTLS) are sparse. So far, only a few global 3-D
 255 models have included iodine chemistry (e.g., atmospheric chemistry-climate models such as CAM by Ordóñez et al., 2012;
 256 SOCOL-AERv2-1 by Karagodin-Doyennel et al., 2021; WACCM by Cuevas et al., 2022; LMDZ-INCA by Caram et al., 2023;
 257 Chemical transport models MOZART by Youn et al., 2010; TOMCAT/SLIMCAT by Hossaini et al., 2015 and GEOS-Chem
 258 by Sherwen et al., 2016). These models have included the major sources of iodine from the ocean, including short-lived
 259 iodocarbons (e.g. CH_3I , CH_2I_2) and primary HOI and I_2 emissions (e.g., Carpenter, 2003; Jones et al., 2010, Saiz-Lopez et al.,
 260 2012; Carpenter et al., 2013). Recent measurements have indicated that up to 0.77 ± 0.10 parts per trillion by volume (pptv)
 261 total inorganic iodine ~~is injected to~~ reaches the stratosphere from ~~the oceans~~ ocean emissions (Koenig et al., 2020).
 262 ~~These~~ Modelling studies have indicated that iodine may play an important role in stratospheric ozone depletion. However,
 263 ~~significant~~ large uncertainties remain over the ~~magnitude and impact~~ contribution of iodine ~~on~~ to stratospheric ozone levels,
 264 ranging from a few percent ~~reduction~~ (e.g., Hossaini et al., 2015; Karagodin-Doyennel et al., 2021) to 10% (Cuevas et al.,
 265 2022) and up to 30% (e.g., Ordóñez et al., 2012). Indeed, the contribution of iodine could become more ~~important~~ pronounced
 266 ~~in the future~~ (Cuevas et al., 2022; Villemayor et al., 2023) ~~due to~~ with the decreasing amounts of stratospheric chlorine and
 267 bromine brought about by the Montreal Protocol (Feng et al., 2021).

268
 269 It is worth pointing out that volcanoes are also a potentially significant source of halogens to the atmosphere (Bobrowski et
 270 al., 2003; Pyle and Mather, 2009). Large halogen-rich eruptions could in principle inject large amounts of halogens, notably
 271 bromine, directly into the stratosphere, causing massive ozone destruction (Kutterolf et al., 2013; Cadoux et al., 2015).
 272 However, this phenomenon has not been observed during the ~~current~~ satellite era.

274 4 Other Challenges

275 The MP has been focused on reducing ozone depletion by anthropogenic halogens. However, there are other well-known
 276 causes of global ozone perturbations, notably natural ones such as the 11-year solar variability (for which the recent solar cycle
 277 23 showed decreased flux) and stratospheric sulfur injections by large volcanic eruptions (e.g. El Chichon in 1982 and ~~Mt~~
 278 Pinatubo ~~eruption~~ in 1991) (WMO, 2022). So far, since the start of satellite observations around 1980, these natural factors
 279 have had a relatively limited impact on global ozone and, unlike the anthropogenic halogen ~~threat~~ emissions, are only expected
 280 to cause short-term (decadal ~~timescales~~ timescale at most) fluctuations in stratospheric ozone.

281
 282 Climate change represents a pressing and long-term issue for stratospheric ozone. The overall impacts of climate change
 283 (largely driven by the increase in CO_2 levels) on stratospheric ozone are complex, ~~with~~ uncertainties ranging from transport to
 284 chemistry effects ~~and their couplings~~ (e.g. changes in the strength of the stratospheric ~~general circulation~~ BDC, changes in the

285 tropospheric water flux into the stratosphere, temperature-dependent chemistry effects, chemistry changes linked to the
286 increasing levels in stratospheric source gases such as CH₄ and N₂O that are also major ~~the~~ greenhouse gases) (e.g. WMO,
287 2018; 2022). Many of these effects are coupled and some of the resulting stratospheric perturbations can, in return, influence
288 the surface climate. For example, the projected increasing speed of the stratospheric BDC will decrease column ozone in the
289 tropics – a region which has so far not been subject to substantial column depletion (Eyring et al., 2007). Increasing levels of
290 N₂O will lead to enhanced NO_x-catalysed ozone depletion in the middle atmosphere (Revell et al., 2012). The impact of
291 increasing CH₄ is more complex: it could lead to increased ozone depletion through increased HO_x but less chlorine-catalysed
292 depletion through deactivation of Cl to HCl (Revell et al., 2012). These effects will increase as GHG levels increase but the
293 details will depend on the relative changes in CO₂, N₂O and CH₄. Therefore, the chemical details of the different prescribed
294 scenarios are important for the ozone impact. Understanding and forecasting the effects of climate change on stratospheric
295 ozone has been a major ~~challenge for several decades now~~ outstanding challenge for several decades now and will remain one
296 for years to come. While our knowledge of relevant atmospheric processes will improve, there will remain the issue of
297 uncertainty in GHG scenarios which are based on societal decisions.

298
299 More recently, other new challenges have emerged. The stratosphere contains aerosol particles which are mostly ~~found~~ located
300 in its lower altitude region. This stratospheric aerosol load is ~~overwhelmingly~~ usually dominated by supercooled sulfuric acid
301 particles whose main sources are stratospheric oxidation of volcanic SO₂ and of OCS, a long-lived sulfur species emitted at
302 the surface (Crutzen, 1976). Sulfuric acid aerosols play an important role in stratospheric chemistry and in the radiative balance
303 of the atmosphere, notably when it is enhanced volcanically. They provide surfaces for key heterogeneous reactions (Hofmann
304 and Solomon, 1989). ~~They also~~, cool the surface by scattering incident sunlight back to space and can heat the stratosphere
305 by absorbing near-infrared radiation (Stenchikov et al., 1998; Roberock, 2000). Until quite recently, almost all the observed
306 global enhancements in stratospheric aerosols and resulting ozone perturbations were linked to sulfur injections by large
307 volcanic eruptions (e.g. El Chichon in 1983, Mt Pinatubo in 1991). As the stratospheric aerosol variability appeared to be
308 essentially driven by volcanic sulfur inputs, only sulfur-induced perturbations of stratospheric aerosols have usually been
309 considered significant for the global stratosphere and climate. This focus on sulfur has also led to the development of
310 sophisticated stratospheric sulfate aerosol microphysical modules which are now implemented in several global climate models
311 (e.g. Zanchettin et al., 2016). These models are able to reproduce observed features of the stratospheric aerosol layer rather
312 well, especially the large enhancements by volcanic eruptions (Zanchettin et al., 2022) and associated ozone losses (e.g. Bekki
313 and Pyle, 1994; Mills et al., 2017). These models are also used to assess the impacts of other stratospheric sulfur injections on
314 stratospheric ozone, for example from aircraft or potential stratospheric geoengineering (Pitari et al., 2014).

315
316 ~~Two recent events have~~ Aircraft measurements in the lowermost stratosphere have already revealed that the nature and
317 composition of stratospheric aerosols are more variable and complex than assumed in most stratospheric aerosol-climate
318 models. These usually only consider sulfur and ignore the substantial components of meteoritic and organic material, dust, and
319 metallic particles from space activities (Murphy et al., 2014; Martinsson et al., 2019; Murphy et al., 2021; Schneider et al.,
320 2021; Murphy et al., 2023). Two recent events have further challenged the dominant view that sulfur is the only aerosol
321 component relevant for the global stratosphere, ozone layer and climate. The first event was the massive Australian wildfires
322 at the turn of 2020, the so-called Australian New Year's (ANY) event (Khaykin et al., 2020; Peterson et al., 2021), ~~and~~ the
323 second event was the eruption of the Hunga Tonga – Hunga Ha'apai (HTHH) volcano in January 2022 (Carr et al., 2022; Zuo
324 et al., 2022). The nature and magnitude of the various stratospheric impacts of these two events have been unexpected and
325 sometimes unprecedented in the historical records. After ~~intensive~~ extensive research on the stratosphere since the discovery
326 of the Antarctic ~~ozone hole phenomenon~~ Ozone Hole in 1985, these two recent events ~~have represented~~ represent extreme but
327 valuable testbeds of our understanding and modelling of stratospheric physics and chemistry.

4.2.1 Australian Wildfires

4.2.1.1 Injections of Carbonaceous Particles and Resulting Aerosol Changes

Wildfires can trigger the formation of pyrocumulonimbus (PyroCb) towers that can ~~rise high,~~ depending on ~~the~~ meteorological conditions and intensity of the fires, ~~and rise high enough to~~ transport biomass-burning material into the UTLS (Peterson et al., 2018). The Australian ‘Black Summer’ wildfires of 2019–2020 were exceptional in terms of scale, intensity and stratospheric impacts according to historical records (Damany-Pearce et al., 2022). The strongest set of PyroCb outbreaks (~~ANY~~) occurred at the turn of 2020, injecting massive amounts of gaseous and particulate biomass-burning products above the tropopause, ~~resulting in a sharp increase in global stratospheric aerosol optical depth (SAOD) which was on a par with the SAOD increases observed after the strongest volcanic eruptions in the last 27 years, Calbuco in 2015 and Raikoke in 2019 (see Figure 3). For example, of the order of,~~ ~~For instance,~~ ~1 Tg of carbonaceous aerosols and ~25 Tg of H₂O were released into the lower stratosphere during the ~~ANY event (Khaykin et al., 2020; Damany-Pearce et al., 2022). Most of the impacts of ANY on the stratosphere resulted from aerosol perturbations.~~ ~~main ANY event (Khaykin et al., 2020; Damany-Pearce et al., 2022), resulting in a sharp increase in global stratospheric aerosol optical depth (SAOD). The rise in SAOD was comparable to the increases produced by the strongest volcanic eruptions since Mt Pinatubo in 1991, namely Calbuco in 2015 and Raikoke in 2019 (see Figure 3). Stratospheric aerosol levels remained enhanced in the Southern Hemisphere throughout 2020. Note that the radiative properties and heterogeneous chemistry of carbonaceous aerosols are different from those of sulfate aerosols (Yu et al., 2022). As a result, the impacts of ANY aerosols on the stratosphere and surface climate are expected to differ from those of volcanic sulfate aerosols.~~

4.1.2.2 Aerosol Changes

~~A very peculiar effect of the strong solar heating by the highly absorptive ANY aerosol plumes was the generation of self-maintained anticyclonic vortices, including one with a diameter of ~1000 km, which contained extremely high concentrations of wildfire gases and aerosols. The massive and remarkably compact vortex persisted for several months while rising to ~35 km (Khaykin et al., 2020). The aerosol lofting opposed the effect of gravitational settling, extending the residence time of ANY aerosols in the stratosphere. Stratospheric aerosol levels remained enhanced in the southern hemisphere throughout 2020. The increase in stratospheric aerosol optical depth (SAOD) was comparable to the increase produced by the Raikoke volcanic eruption in 2019 which was then the greatest aerosol increment since the massive volcanic eruption of Mt Pinatubo in 1991.~~

4.2.3 Gaseous Composition Changes

ANY stratospheric aerosol changes were accompanied by very unusual large-scale perturbations in gaseous composition. For example, in the months following the ANY aerosol dispersion, unexpected partitioning between radicals and reservoir species in the chlorine and nitrogen families were observed at southern mid-latitudes at relatively warm stratospheric temperatures (Santee et al., 2022). The main stratospheric chlorine reservoir species HCl was found to be largely depleted while the other chlorine reservoir, ClONO₂, and the ozone-destroying chlorine radical ClO, were enhanced. The anomalous partitioning is somewhat reminiscent of the effects of ozone-~~destroying/depleting~~ heterogeneous chemistry on other stratospheric aerosols (sulfuric acid particles, PSCs) and ~~is likely caused/was probably initiated~~ by some heterogeneous processing on ~~the ANY~~ particles (Bernath et al., 2022; Solomon et al., 2023). ~~However, the physical state and chemical reactivity of such wildfire~~

carbonaceous particles in the conditions prevailing in the stratosphere needs to be characterised, notably through laboratory studies and field campaigns.

Overall, the enhanced ClO concentrations ~~must have resulted in~~ likely caused some, albeit weak, chemical ozone depletion. A mini ozone hole (depletion of up to 100 DU) was also apparent ~~early on~~ within the largest plume vortex ~~early on~~ (Section 4.1.3; Khaykin et al., 2020) and the Antarctic ozone hole was particularly long-lasting in 2020. ~~Different~~ (Klekociuk et al., 2022). Several aerosol-driven mechanisms have been proposed to explain these ozone changes, invoking changes in stratospheric dynamics and/or heterogeneous chemistry.

~~It has to be stressed that, at this stage, we do not know exactly the physical state (e.g. liquid, glassy, solid) and composition of such wildfire particles in the conditions prevailing in the stratosphere, all the more so when internally mixed with sulfuric acid (Solomon et al., 2023). As a result, the types and rates of heterogeneous reactions occurring on them can only be hypothesised. Further laboratory studies and, as importantly, detailed chemical composition measurements are certainly our best means to characterise unequivocally the physico-chemistry of these aerosols.~~

4.2.4.3 Dynamics and Radiative Forcing

In the early phase when aerosol concentrations within the ANY plumes were extremely high, the intense solar heating by ANY aerosols led to the formation of rising, confined vortex-like plumes through dynamical feedbacks. ~~the highly absorptive ANY aerosol plumes led to very peculiar dynamical feedbacks and the formation of self-maintained anticyclonic vortices. This included one with a size of ~1000 km, which contained extremely high concentrations of wildfire gases and aerosols. The massive and remarkably compact vortex persisted for several months while rising diabatically to ~35 km (Khaykin et al., 2020). The aerosol lofting opposed the effect of gravitational settling, prolonging the residence time of ANY aerosols in the stratosphere. Interestingly, after the discovery of heating and self-lofting by ANY carbonaceous aerosols, an analysis of high-resolution satellite observations has showed that the Raikoke volcanic eruption in 2019 also generated a stratospheric anticyclonic vortex which rose to 27 km and persisted for more than 3 months (Khaykin et al., 2022). Since sulfate aerosols absorb radiation only weakly, the heating must have been generated by absorption from another volcanic aerosol component, likely to be volcanic ash. Currently, most stratospheric aerosol models only consider sulfate aerosols and hence cannot reproduce the observed dynamical confinement and ascent of concentrated carbonaceous plumes or ash-rich plumes and hence the extended residence time in the stratosphere.~~ Once the ANY aerosol plumes were dispersed and spread, the aerosol heating led to a pronounced large-scale warming of the southern lower stratosphere (Stocker et al., 2021; Damany-Pearce et al., 2022) which was stronger than any warmings from recent volcanic eruptions.

The radiative climate forcing ~~of~~ by ANY aerosols is more difficult to estimate ~~by comparison to that from~~ than the forcing by sulfate aerosols. Sulfate aerosols cool the surface by efficiently scattering incoming sunlight back to space ~~which and this effect readily~~ dominates the surface-warming tendency from their absorption of longwave radiation. Carbonaceous aerosols not only scatter solar radiation but also absorb it, and this absorption is strongly dependent on the aerosol composition. ANY aerosols are thought to have been mostly composed of a small fraction of black carbon (BC, soot-like component) and a vastly dominant fraction of organic material (OM, including the so-called brown carbon (BrC) component) (Liu et al., 2022). BC absorbs across the entire solar spectrum and hence is by far the most efficient source of heating. Most OM compounds absorb strongly in the IR and UV wavelengths, but are relatively transparent in the visible and near-IR wavelengths. This is not the case for BrC which can also absorb in the blue and near-UV spectral regions, albeit with a much weaker efficiency than BC (Laskin et al., 2015; Yu et al., 2021). ~~Since~~ Given the exact-poor observational constraints on the composition, ~~mixture~~ physical and mixing

state, and size distribution of ANY carbonaceous aerosols ~~are poorly constrained, the ANY (all key parameters of aerosol radiative properties), the radiative impact of ANY aerosol remains as difficult to quantify assess as their heterogeneous chemistry.~~ Estimations of ANY aerosol surface radiative forcing (RF) vary from negligible to about -1 Wm^{-2} ; this range can be compared to the RF of small to moderate volcanic eruptions during the last 3 decades, estimated at between -0.1 and -0.2 Wm^{-2} (Sellitto et al., 2022a; Liu et al., 2022). An additional complication in the ANY RF estimation is the effect of the aerosol-driven stratospheric warming on the longwave radiation budget (Liu et al., 2022).

It is worth pointing out that, as global surface warming intensifies, massive wildfires and associated pyro-convective injections of carbonaceous particles in the stratosphere are expected to become more frequent. Pyro-convection could turn into a significant source of large-scale perturbations of stratospheric aerosols, ozone, and climate. Therefore, it might be necessary to account for stratospheric wildfire particle processes in CCMs and comprehensive Earth system models (ESMs) in the future.

4.32 Hunga Tonga – Hunga Ha’apai ~~volcanic eruption~~ Volcanic Eruption of the January 15th, 2022

4.32.1 Injection of H₂O and Sulfur

The eruption of the Hunga Tonga – Hunga Ha’apai volcano with an underwater caldera occurred on January 15th 2022. Several features of this eruption were unique in the record of stratospheric observations. First, it generated a very powerful blast that injected volcanic material up to an altitude of nearly 58 km (Proud et al., 2022; Carr et al., 2022). A volcanic plume from a moderate eruption reaching the lower mesosphere was barely conceivable until this event, especially when the plume of the Mt Pinatubo eruption in 1991 with an explosivity index larger than the HTHH eruption reached at most an altitude of ~40 km (McCormick et al., 1995). Second, the HTHH eruption injected a very small amount of SO₂ (0.4-0.5 Tg) but a very large massive quantity of H₂O into the middle atmosphere, between 120 and 150 Tg (Carr et al., 2022; Millan et al., 2022; Xu et al., 2022; Khaykin et al., 2022), into the middle atmosphere, resulting in unprecedented very large increases in stratospheric water vapour (see **Figure 4**). Again, such a volcanic emission scenario had not been generally considered previously. H₂O isotopic ratio data strongly suggest indicate that sea water was a major source of stratospheric hydration by the HTHH eruption (Khaykin et al., ~~2022~~2022), which is consistent with the high concentrations of sea salts found in HTHH tephra (volcanic ash) collected shortly after deposition at the surface (Colombier et al., 2023).

4.32.2 H₂O and Sulfate Aerosol Changes

~~Following the~~ The HTHH eruption, increased the mean global stratospheric water content ~~increased~~ by approximately 10%, which is unprecedented in the entire observational record dating back to 1985. Note that, as there are no significant sinks of H₂O within the stratosphere, this excess H₂O is expected to last over should persist at least several years- during which time the water vapour is slowly transport to the troposphere. In contrast, volcanic sulfate particles have a shorter residence time in the stratosphere, with typically an e-folding decay time of a year, because of the effect of gravitational sedimentation. This difference between the gaseous and aerosol components has led to an increasing vertical decoupling of the HTHH enhanced water vapour and aerosol layers in the stratosphere (Millan et al., 2022; Khaykin et al., 2022).

Most of the HTHH SO₂ was oxidised to sulfate aerosols within a month because of the H₂O-driven OH enhancement (Zhu et al., 2022). As a result, the The SAOD (averaged between 60°S and 60°N above 380 K) increased rapidly and reached a peak 5 months after the eruption (Khaykin et al., 2022). Surprisingly, the magnitude of the SAOD increment did not follow at all the

usual common relationship between SAOD and volcanic SO₂ mass. The ANY SAOD peak exceeded by about a factor 2 the SAOD perturbations caused by injected in the stratosphere. For instance, the HTHH SAOD enhancement, which easily outweighed all the volcanic and wildfire events aerosol perturbations in the last three decades, including exceeded by a factor 4 the SAOD peak caused by 2015 Calbuco eruption that injected roughly the same amount of sulphur, and by a factor 2 the SAOD peak caused by the 2019 Raikoke eruption that injected two times more SO₂ than the Toga eruption. This 4 fold enhancement in SAOD for the HTHH-THHH eruption compared to the (Khaykin et al., 2022). This unexpected increase in SAOD expected for a 0.4-0.5 Tg SO₂ injection, such as the 2015 Calbuco eruption, in the case of the HTHH could not be linked to the possible presence of volcanic ash because that was apparently removed within days after the eruption and, according to satellite data, ANY HTHH aerosols were essentially liquid sulfate droplets (Legras et al., 2022; Bernath et al., 2023). The highly enhanced SAOD must have its origin in been due to the excess humidity in the stratosphere, possibly through aerosol hygroscopic growth or coagulation. Indeed, in sulfate aerosol microphysical model simulations of the HTHH eruption, the SAOD generated by the ~0.4 Tg SO₂ injection is approximately doubled by the co-injection of 150 Tg of water (Zhu et al., 2022). However, Nonetheless, the model still underestimates the observed SAOD by a factor 2, suggesting that the effect of water vapour on sulfate aerosols is yet not fully understood, or that the HTHH aerosols were not just composed of sulfuric acid and water, possibly with sea salts affecting the aerosol hygroscopicity.

Satellite observations of trace gases have also provided strong evidence for heterogeneous chemical processing on HTHH aerosols with unambiguous signatures of substantial chlorine and nitrogen repartitioning in the regions of aerosol enhancements almost immediately after the eruption (Santee et al., 2023). Model simulations indicate that stratospheric ozone has been significantly impacted by the eruption following several mechanisms (e.g. through not only heterogeneous chemistry on diluted sulfate aerosols, but also other chemical and dynamical mechanisms (e.g. H₂O-enhanced gas-phase radical chemistry, and circulation changes) (Lu et al., 2023).

4.3.3 Dynamics and Radiative Forcing

It is worth recalling The first radiative forcing (RF) model calculations for HTHH took only into account the sulfur injection, ignoring the water injection, and, as expected, concluded that an increased the HTHH sulfate aerosols would slightly cool the surface (Zuo et al., 2022). However, enhancements in lower stratospheric H₂O and enhanced stratospheric sulfate aerosols generally have opposite radiative impacts. The H₂O increment tends to cool the stratosphere and warm the surface while a sulfate aerosol increment tends to warm the stratosphere and cool the surface.

The water vapour content within the HTHH plume was initially so large high that the H₂O radiative cooling led to a descent of the volcanic plume during the first weeks after the eruption (Sellitto et al., 2022b). After this initial phase, negative temperature anomalies were found to be correlated with H₂O-rich layers and cooling rate (Schoeberl et al., 2022). The decrease in global temperatures were was rather extreme in the mid-stratosphere during 2022, deviating markedly from all the previous 42 years of meteorological data (Coy et al., 2022). The sign of the stratospheric temperature response confirms that the H₂O cooling clearly dominated the sulfate aerosol warming in the stratosphere. These temperature perturbations were also accompanied by circulation adjustments.

The effect of the HTHH event on surface climate is not as clear-cut as in the stratosphere. RF model calculations suggest that ultimately the eruption warmed the surface; i.e. that the H₂O warming was slightly dominant over sulfate cooling (Sellitto et al., 2022b; Jenkins et al., 2023). Interestingly, the initial RF model calculations only took into account the amount of sulfur injected and hence concluded that the HTHH eruption would slightly cool the surface (Zuo et al., 2022). It is also worth

491 ~~pointing out that, as global warming intensifies, massive wildfires and of pyro-convective injections of carbonaceous particles~~
492 ~~in the stratosphere are expected to become more frequent. Pyro-convective could turn into a significant source of large-scale~~
493 ~~perturbations for stratospheric ozone and climate. As such, it will be necessary to account for it in Earth system models (ESMs)~~
494 ~~in the future.~~

496 5 Maintaining Observational Capacity

497 Our understanding of the ozone layer, and ~~of~~ the processes which control its evolution including those outlined here,
498 ~~depend~~depends on the availability of high-quality observations. In recent years we have benefitted from a wealth of
499 observations from instruments in ground-based networks, and on balloon and aircraft flights and satellites satellite platforms.
500 However, there are several indications that ~~we future progress~~ will be ~~hampered~~impeded by fewer observations in the future.

504 5.1 Gaseous Composition

505 Several currently operational spaceborne instruments are well beyond their design lifetimes, and some are scheduled to be
506 decommissioned in the next few years. Instruments whose data have been cited ~~hereabove~~ or regularly used as part of the 4-
507 yearly WMO/UNEP Ozone Assessments (e.g. WMO, 2022) will likely cease operations within the next few years, including
508 the Aura Microwave Limb Sounder (MLS), the SciSat Atmospheric Chemistry Experiment Fourier Transform Spectrometer
509 (ACE-FTS), the Odin Optical Spectrograph and Infrared Imager System (OSIRIS), and the Odin Sub-Millimetre Radiometer
510 (SMR). With the loss of these current limb-viewing capabilities, vertically resolved global measurements of many trace gases
511 relevant for studies of stratospheric chemistry and dynamics will no longer be available. These trace gases include ozone-
512 destroying reactive (ClO) and reservoir (HCl, ClONO₂) chlorine species, water vapour, nitric acid (HNO₃), and long-lived
513 tracers of transport~~tracers~~ (e.g., nitrous oxide, N₂O; methane, CH₄; carbon monoxide, CO). As noted in WMO (2022), the
514 2021 Report of the Ozone Research Managers of the Parties to the Vienna Convention (ORM, 2021a) identified the need to
515 “continue limb emission and infrared solar occultation observations from space” that are “necessary for global vertical profiles
516 of many ozone- and climate-related trace gases” as one of the “key systematic observations recommendations.” Indeed, the
517 impending loss of these measurements, many of which have been taken continuously over the last several decades, will hamper
518 our ability to ~~reduce key appreciate and address key gaps and~~ uncertainties in our understanding of stratospheric ozone
519 depletion, including the lack of emergence of a clear signature of recovery in the Arctic, the ~~potential influence~~influences of
520 volcanic and wildfire emissions, the role of VLSL, and the impact of strengthening of the Brewer-Dobson circulation.
521 Ultimately, this will risk weakening the scientific framework of the Montreal Protocol including the decision-making process.
522 It may take many years for the next generation of improved limb sounders to become operational and provide us with the
523 observational capacity that ~~we have been used to~~has been so essential to understanding the evolution of the ozone layer over
524 the past three decades. For example, the novel, high resolution Changing Atmosphere Infra-Red Tomography Explorer
525 (CAIRT) (<https://www.cairt.eu/>) is currently a candidate mission for ~~at~~ the European Space Agency Earth Explorer 11 mission
526 but, if selected, will not start operating before 2033 at best.

527
528 Ground-based networks have also proved essential for ~~our continued study of~~ the ozone layer ~~and processes that affect~~
529 ~~research~~. Examples are the Network for the Detection of Atmospheric Composition Change (NDACC, De Mazière et al.,
530 2018), and the National Oceanic and Atmospheric Administration (NOAA, e.g. Montzka et al., 2018) and Advanced Global

531 Atmospheric Gases Experiment (AGAGE, e.g. Rigby et al., 2019) surface networks. While these networks have an important
532 monitoring function, the data acquired have proved central to trend analyses, to the validation of satellite measurements, and
533 to the identification of many of the new scientific challenges discussed here. The benefit of these data sets increases greatly as
534 the time series extend so that longer term variations can be characterised and studied. Therefore, it is ~~very important~~crucial to
535 maintain their continuity- as discussed, for example, in ORM (2021a, b).

537 5.2 Aerosol Composition

538 As for trace gases, the much reduced availability of satellite limb-viewing observations in the future is concern for research on
539 stratospheric aerosols, an important driver of stratospheric ozone. Several spaceborne instruments arriving towards the end of
540 their lifetimes have been providing critical information on stratospheric aerosol properties. It is not ideal to reduce stratospheric
541 aerosol observations with their global coverage, especially when it is becoming increasingly clear that large-scale ozone
542 perturbations from stratospheric aerosol changes are not limited to volcanic sulfur injections. The chemical composition of
543 stratospheric aerosols is more variable and complex than often assumed. In addition to sulfuric acid and water, stratospheric
544 aerosols contain significant fractions of carbonaceous, meteoritic, and space activity material but the impacts of some of these
545 components on stratospheric ozone are more or less unknown. As a result of poor observational constraints, large uncertainties
546 pertain to the sources, size distribution, heterogeneous chemical reactivity, possibly polar stratospheric cloud activation ability,
547 or/and radiative properties of these components. This incomplete knowledge hinders our ability to foresee the state of the
548 ozone layer under the effect of a range of potential aerosol perturbations such as massive wildfires expected to be favoured by
549 global warming, the anticipated increase in space activities within the next few decades (e.g. Ryan et al., 2022), the impact of
550 meteoritic particles (Plane et al., 2023) or stratospheric geoengineering (i.e. deliberate injection aerosols or/and gaseous
551 precursors in the stratosphere in order to counteract climate warming (e.g. Tilmes et al., 2022)).

552
553 It is worth pointing out that satellite observations cannot alone constrain unambiguously key aerosol parameters, in particular
554 chemical composition and size distribution. Satellite data have to be confronted and combined with in-situ detailed composition
555 and size measurements from balloon and aircraft; in addition, laboratory studies help to characterise the primary processes
556 relevant to the aerosol physico-chemistry (Burkholder et al., 2017). All these types of measurements are needed to advance
557 our understanding of stratospheric aerosol processes and impacts, and thus improve their representations in models.

559 **6 Chemistry-Climate Modelling and Ozone Projections**

560 Our understanding of the chemical, dynamic and radiative processes and of their couplings which control stratospheric ozone
561 is encapsulated in mathematical form in numerical models. These models are powerful tools ~~and simulators~~ in tackling a range
562 of scientific and societal challenges. Obviously, they can only include known processes (as the surprise discovery of the
563 Antarctic ozone hole clearly demonstrated) and even for these there can be significant uncertainties. Overall, progress in our
564 understanding of the ozone layer will depend on the improvement and careful application of a hierarchy of models from
565 detailed chemical-aerosol box models, through 3-D chemical transport models (CTMs) to complex chemistry-climate models
566 (CCMs).

567
568 Regarding Computationally inexpensive 3-D CTMs will continue to play an important role in interpreting observations on a
569 range of spatial and temporal scales, testing our understanding and developing parameterisations for new processes. These
570 models contain detailed chemistry-aerosol schemes but are forced by off-line meteorological analyses making them ideal tools

571 ~~for comparing with observations and for many sensitivity studies.~~ CCMs ~~are needed to study chemical-radiative-dynamical~~
572 ~~interactions but~~ these models ~~can be~~ relatively very computationally expensive to run. ~~Huge~~ Continuing advances in
573 computing resources allow ever more complex processes to be added ~~to models~~ which can help understanding of feedback
574 pathways but can mean that simulations are often at the limit of what is practical. ~~A set of ensemble~~ For example, the
575 stratospheric impact of halogenated VSLS will ideally require detailed tropospheric chemistry in order to accurately model the
576 transport of product gases to the stratosphere. This ‘whole atmosphere’ chemistry is also desirable for many other reasons, but
577 it adds to the cost of all stratospheric simulations, and to the amount of model output generated. A set of ensemble CCM
578 simulations (needed to characterise the model internal variability) can take many months of real time even on a powerful High
579 Performance Computing (HPC) system. The ~~Moreover, the~~ costs increase greatly as other modules, such as ocean, crvosphere
580 and biosphere, are added to build a full ESM. Within an ESM. ~~Therefore, framework there is, we think, a danger that the~~
581 treatment of the stratosphere is simplified to such an extent that the model will not capture many of the important processes
582 discussed above (e.g. VSLS, wildfire smoke) and thus will not produce the best estimate for processes such as ozone layer
583 recovery. For example, the standard UKESM (Archibald et al., 2020) only treats three ODSs (2 CFCs and CH₃Br) with other
584 simplifications for PSCs. In practice, other versions of the ESM may be available (in effect a ‘CCM’ if other modules are not
585 used to save time) but these will likely not be used for flagship climate simulations in major international assessments. This is
586 important not only for simulating the stratosphere itself but because changes in the stratosphere are known to exert important
587 impacts on the troposphere and the surface (e.g. Thompson et al., 2011).

588
589 Given the computational challenges simulations with CCMs (and ESMs) need to be planned carefully. Results from any given
590 CCM (or ESM) model will have various causes of uncertainty: (1) ~~scenario uncertainty – related for example to the ODS and~~
591 GHG scenarios used to force the model (1) internal variability; (2) structural uncertainty – related to the model grid and
592 parameterisations used to represent known processes and (3) internal variability-scenario uncertainty – related for
593 example to the ODS and GHG scenarios used to force the model. To address ~~(2) CCMs need~~ (1) each CCM needs to perform
594 an ensemble of simulations. To address (2) a selection of models are needed to perform a given experiment in order to obtain
595 a robust result (in the sense that the result is not, or at least only weakly, model-dependent). ~~For example~~ To address (3) the
596 models must be computationally cheap enough to simulate a range of possible scenarios. For example, as discussed above, an
597 important use of CCMs is to predict recovery of the ozone layer from chlorine and bromine-catalysed loss, and the dependence
598 of that recovery on climate change. These results are obtained from projects such as the Chemistry-Climate Modelling Initiative
599 (CCMI, <https://igacproject.org/activities/CCMI>) and feed into the WMO/UNEP Assessments. It is important that the
600 participating models have been thoroughly evaluated and that they perform sufficient experiments (with ensemble members).
601 For example, as noted by Dhomse et al. (2018), robust estimates of sensitivity to GHG scenarios are better achieved when all
602 (well evaluated) models perform all experiments and these results from around 20 models fed into the projections used in
603 WMO (2018) (see Figure 1b). In comparison, projections used in WMO (2022) were based on only 5 or 6 models (depending
604 on region) and from simulations that were performed for the wide-ranging Coupled Model Intercomparison Project 6 (CMIP6,
605 <https://pcmdi.llnl.gov/CMIP6>) which were not focussed on the stratosphere. While all models used for assessment purposes
606 should ideally have comprehensive stratospheric processes, their projections of ozone recovery will always have some caveats.
607 Clearly the models cannot contain unknown processes – and the recent example of chlorine activation on wildfire smoke
608 particles (Section 4.1) is one example. Even for known processes, we do not know how external forcings, such as volcanic
609 eruptions, will vary. Therefore, there will always be an important role for additional, focussed studies of chemistry-climate
610 interactions and projections outside of the main assessment process in order to explore detailed interactions and accommodate
611 new knowledge.

613 Given the increasing computational cost of the CCM ~~(or ESM)~~ simulations then it ~~seems~~ desirable that ~~some~~ other
614 approaches are used to update projections of ozone layer recovery which do not depend on extensive new sets of model runs.
615 A commonly used metric is the 'ozone return date' (see **Figure 1b**). This is the date at which modelled ozone levels return to
616 a reference value, which is often taken to be 1960 or 1980. These return dates are, for example, typically around 2040 for
617 global mean column ozone and 2066 for the Antarctic in October, but with large uncertainty due to e.g. GHG scenarios (WMO,
618 2022). ~~Performing simulations to update these estimates is~~ Although an apparently simple metric, there are a number of obvious
619 shortcomings with return dates. The return date measures recovery as a single event and does not take account the trajectory
620 of ozone prior to that date, noting that the impact of increased surface UV will depend on time history of ozone depletion.
621 Furthermore, small shifts in the extent of ozone depletion around the return date can cause a large 'delay' in when ozone
622 recovery is deemed to have occurred. Performing simulations to update these estimates is also expensive with possibly only a
623 small benefit if the ODS and GHG scenarios have only changed slightly. Therefore, alternative approaches should be
624 investigated for estimating, for example, the dependence of the ozone return date on the chlorine and bromine return dates,
625 and the sensitivity of this to different GHGs.

626
627 Recently, Pyle et al. (2022) proposed the Integrated Ozone Depletion (IOD) metric and showed how it applies to similar long-
628 lived ODSs. IOD is an absolute measure of the time-integrated column ozone depletion for different halocarbon scenarios
629 which, for long-lived ODSs, reduces to a simple empirical formula with a model-derived scaling factor. As noted above,
630 application of ODPs to VSLS depends on the distribution of the surface emissions, which leads to a range of IOD values.
631 Because VSLS can cause ozone changes in the troposphere, Zhang et al. (2020) proposed the use of 'stratospheric ODP'
632 (SODP) as a simpler and more direct measure of only stratospheric column changes. Further work from the modelling
633 community is needed to derive a robust range of (S)ODPs for VSLS, and to also extend the work of Pyle et al. (2022) to
634 investigate how to apply the IOD metric to VSLS. In particular, we need to test the sensitivity of ~~modelled/simulated~~ ozone
635 depletion to emissions (i.e. IOD scaling factor) ~~in a range of models.~~

637 **7 Summary-Future Outlook**

638 This Opinion article demonstrates that after 100 years of research, and nearly 4 decades after the discovery of the Antarctic
639 ozone hole, the stratospheric ozone layer is still ~~throwing up~~ producing surprises and new research challenges, ~~showing that~~
640 Clearly we cannot lower our guard on this global environmental issue. The great progress that we have made in ozone layer
641 science has been achieved through the combination of laboratory studies, observations from a range of platforms, and
642 modelling. All of these components are essential for continued progress in research and policymaking concerning the
643 preservation of the ozone layer.

644
645 Our reflections of the long-standing and new challenges presented in this paper can, we think, be summarised in the following
646 overarching research needs:

- 647 • Maintaining and expanding the observational monitoring networks to ensure compliance with the Montreal Protocol
648 for the controlled gases and to understand the distribution and emissions of important uncontrolled gases. This
649 monitoring should cover the important ODSs, the replacements of the ODSs and VSLS, and be of high enough
650 coverage that emissions can be traced to specific regional sources.
- 651 • Addressing the critical issue of the impending satellite gap in observational capacity of the stratosphere which will
652 greatly reduce our ability to study processes globally. In order to understand changes in stratospheric ozone, for
653 example to track recovery or understand new perturbations, we need height-resolved profiles of related chemical and

654 aerosol species. Other targeted observational campaigns from aircraft and balloons in the low-mid stratosphere are
655 critical for increasing the observational database to more species and providing observations at high spatial resolution.

- 656 • Supporting the development, testing and application of process models. This will also require relevant laboratory
657 studies to measure key parameters. CTMs will continue to be important tools to test understanding and interpret
658 observations. This development can feed into the chemistry-aerosol modules used in more complex CCMs.
- 659 • Ensuring that ESMs being developed worldwide treat the stratosphere in sufficient detail. Use of ESMs for assessment
660 simulations should be based on well-tested models, a sufficiently large number of ensemble members to account for
661 model internal variability and include provision for a range of scenarios and sensitivity runs. If the full ESM is too
662 costly for this then regular CCMs should be used. New metrics need to be explored, e.g. to quantify ozone recovery,
663 which provide direct measure of the impact of the process being considered and to reduce the need for a large number
664 of repeated expensive model runs as external forcings change only slightly.

665
666 Our personal experience has also convinced us of the great importance of collaborative international programmes and
667 campaigns which have been truly instrumental in advancing our knowledge on the topic. Ultimately, society's interest in the
668 ozone layer is due to the impact of ozone depletion on surface UV and climate. As this article has shown, although the ozone
669 layer is demonstrating recovery from the effects of long-lived ODSs, other processes such as uncontrolled short-lived species,
670 changing dynamics, and wildfire smoke, ~~threaten could cause~~ further ~~depletion perturbations~~. We need to continue to observe,
671 understand and model these processes; *Atmospheric Chemistry and Physics* will continue to ~~provide a primary be an important~~
672 journal for our community's major advances in these areas.
673

674 **List of Acronyms**

675 ANY – Australia New Year
676 CCM – Chemistry-climate model
677 [CMIP – Coupled Model Intercomparison Project](#)
678 [CTM – Chemical Transport Model](#)
679 EEC1 – Equivalent effective chlorine
680 ESM – Earth system model
681 GHG – Greenhouse gas
682 HT HH – Hunga Tunga – Hunga Ha’apai
683 IOD – Integrated ozone depletion
684 [IR - Infrared](#)
685 MP – Montreal Protocol
686 ODP – Ozone depletion potential
687 ODS – Ozone depleting substance
688 PGI – Product gas injection
689 PSC – Polar stratospheric cloud
690 RF – Radiative forcing
691 SAOD – Stratospheric aerosol optical depth
692 SGI – Source gas injection
693 SODP – Stratospheric ODP
694 UTLS – Upper troposphere – lower stratosphere
695 UV – Ultraviolet
696 VSLS – Very short-lived substance
697 WMO – World Meteorological Organisation

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700 Both authors contributed to the writing of this article.

702 **Competing Interests**

703 The authors declare that they have no conflicts of interest.

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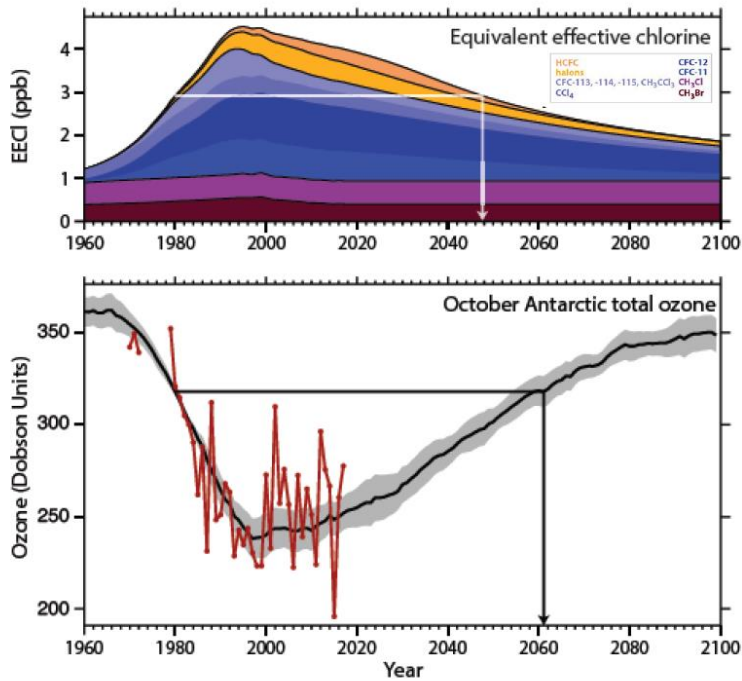
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1086 **Figures**
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Figure 1: (a) Past observations and projections of the equivalent effective chlorine (EECI; total chlorine + $65 \times$ total bromine at surface, ppb) from different long-lived ODSs between 1960 and 2100. After the signing of the Montreal Protocol and subsequent ~~phase-out~~phaseout of many long-lived ODSs, the EECI began to decline and is expected to return to 1980 levels by around 2050, as indicated by the horizontal and vertical dashed lines. ~~(b) Measured~~Note that more recent estimates of EECI would give a slightly later return date (WMO, 2022). (b) ~~Measured~~ (red line) and predicted (black line, with uncertainty shown as grey shading) October Antarctic column ozone (Dobson units) between 1960 and 2100. In this simulation the Antarctic ozone layer is expected to return to 1980 levels around 2061, around a decade later than the EECI (horizontal and vertical dashed lines). CFC, chlorofluorocarbon; HCFC, hydrochlorofluorocarbon. Note that this Antarctic October return date is slightly earlier than the most recent estimate given in WMO (2022) but still within the model uncertainty range. Figure adapted from ~~WMO (2018) and~~ Chipperfield et al. (2020) ~~and based originally on WMO (2018).~~

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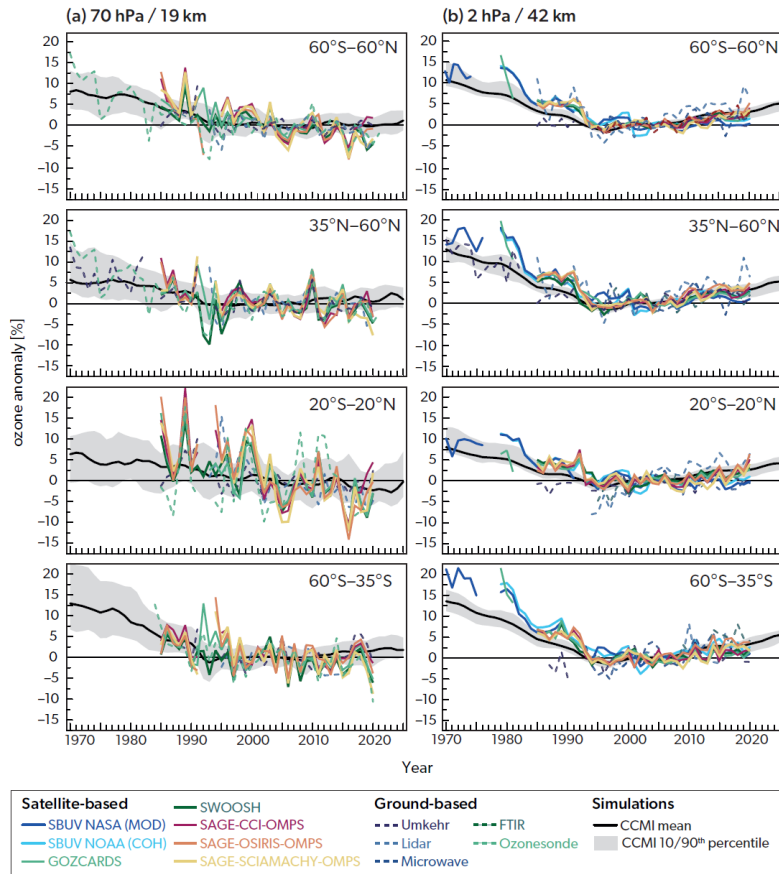


Figure 2. Annual mean anomalies of ozone in (a) the lower stratosphere, near 19 km altitude (70 hPa pressure) and (b) the upper stratosphere, near 42 km (2 hPa), for four latitude bands: 60°S–60°N, 35–60°N, 20°S–20°N (tropics), and 60–35°S. Anomalies are referenced to a 1998–2008 baseline. Coloured lines are long-term records obtained by merging data from different nadir (SBUV NASA (MOD) and SBUV NOAA (COH)) or limb-viewing (GOZCARDS, SWOOSH, SAGE-CCI-OMPS, SAGE-OSIRIS-OMPS, SAGE-SCIAMACHY-OMPS) satellite instruments. Dashed coloured lines are long-term records from ground-based observations (Umkehr, lidar, microwave, FTIR and ozonesondes); see Steinbrecht et al. (2017), WMO (2018), and Arosio et al. (2018) for details on the various datasets. The gray shaded areas show the range (10th and 90th percentiles) of 16 CCM simulations performed as part of the CCM1-1 REF-C2 experiment (see Morgenstern et al., 2017) with the black line indicating the median. Taken from Figure 3-9 in WMO (2022).

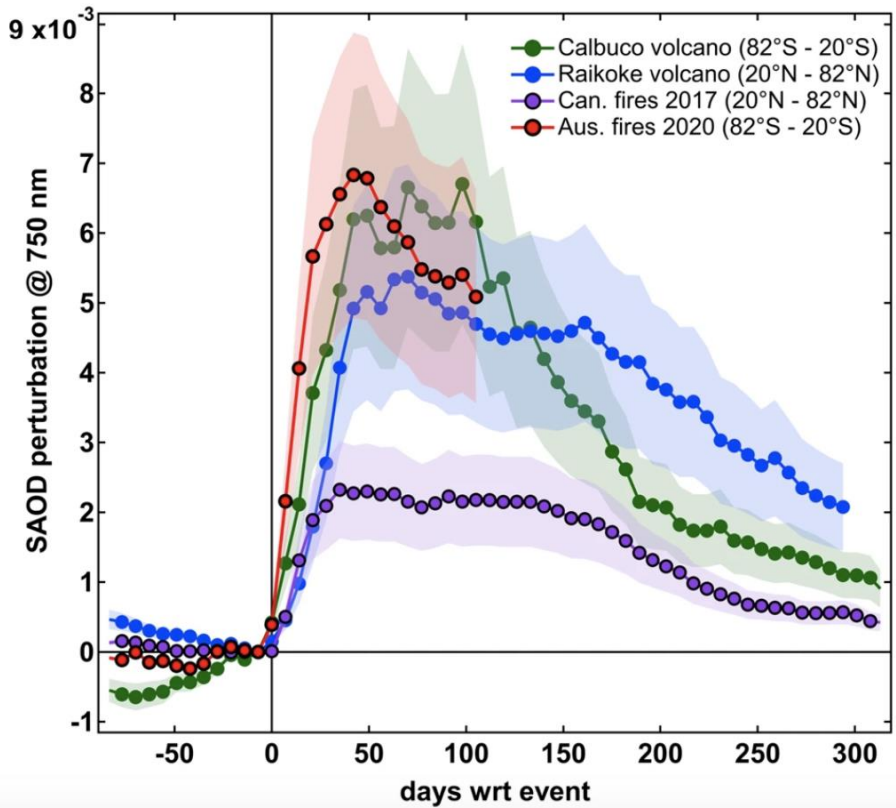


Figure 3. Perturbation of the stratospheric aerosol optical depth (SAOD) due to Australian fires and the strongest events since 1991. The curves represent the SAOD perturbation at 746 nm following the Australian wildfires, the previous record-breaking Canadian wildfires in 2017, and the strongest volcanic eruptions in the last 29 years (eruptions of Calbuco volcano in 2015 and Raikoke volcano in 2019). The time series are computed from OMPS-LP aerosol extinction profiles as weekly-mean departures of aerosol optical depth above 380 K isentropic level from the levels on the week preceding the ANY event. The weekly averages are computed over equivalent-area latitude bands roughly corresponding to the meridional extent of stratospheric aerosol perturbation for each event. The shading indicates a 30% uncertainty in the calculated SAOD, as estimated from SAGE III coincident comparisons. See Khaykin et al. (2020) for more details.

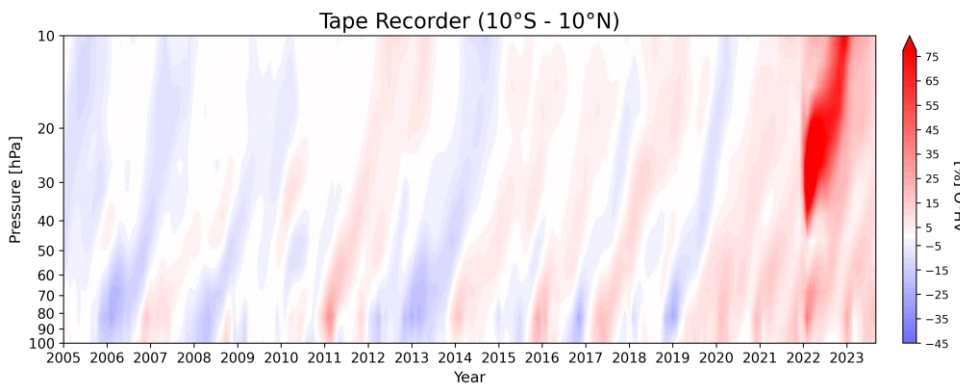
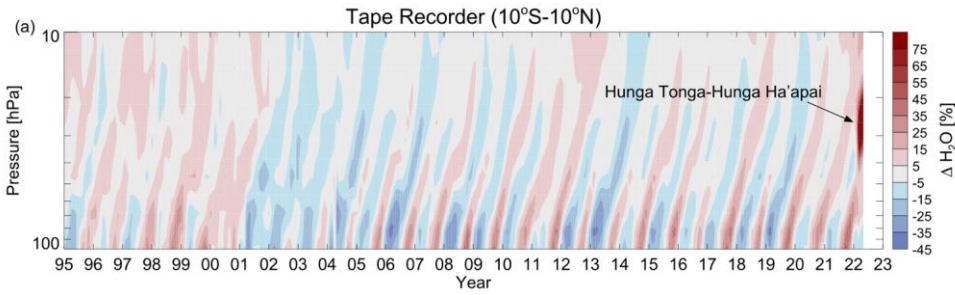


Figure 4. Zonal mean H₂O anomalies (%) in the tropics, between 10°S and 10°N (the so-called atmospheric tape recorder) for January 2005 to August 2023. H₂O abundances are based on GOZCARDS (Froidevaux et al., 2015) and version 5 Microwave Limb Sounder data. See Based on Figure 5a in Millan et al. (2022) for more details. Figure courtesy of Xin Zhou (University of Chengdu).