

1 **Integration of microseism, wavemeter buoy, HF Radar and hindcast data to analyze the**  
2 **Mediterranean cyclone Helios**

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27 monitoring sea state

## 28 **Abstract**

29 In this work, we study a Mediterranean cyclone, called Helios, took place during the period 9-  
30 11 February 2023 in the southeastern part of Sicily and Malta Island, by a multiparametric  
31 approach combining microseism results with sea state and meteorological data provided by  
32 wavemeter buoy, HF Radar, hindcast maps and satellite SEVIRI images. The sub-tropical  
33 system Helios caused heavy rainfall, strong wind gusts and violent storm surges with  
34 significant wave heights greater than 5 meters. We deal with the relationships between such a  
35 system and the features of microseism (the most continuous and ubiquitous seismic signal on  
36 the Earth) in terms of spectral content, space-time variation of the amplitude and source  
37 locations tracked by means of two different methods (amplitude-based grid search and array  
38 techniques). By comparing the location of the microseism sources and the area affected by  
39 significant storm surges, derived from sea state data, we note that the microseism location  
40 results are in agreement with the real position of the storm surges. In addition, we are able to  
41 obtain the seismic signature of Helios using a method that exploits the coherence of continuous  
42 seismic noise. Hence, we show how an innovative monitoring system of the Mediterranean  
43 cyclones can be designed by integrating microseism information with other techniques  
44 routinely used to study meteorological phenomena.

45

## 46 **1. Introduction**

47 Significant storm surges driven by intense low-pressure systems represent one of the main  
48 hazards to the Mediterranean coastal areas causing flooding, beach erosion and damage to  
49 infrastructures and cultural heritages (Flaounas et al., 2022; Lionello et al., 2019;).  
50 Occasionally, when there are favorable conditions like high sea temperature and high contrast

51 of temperature sea-air, the cyclones can acquire the characteristics of a MEDiterranean  
52 hurriCANE (hereinafter Medicane). Medicanes genesis is favored when an extratropical  
53 depression gets isolated from the polar jet stream. The "cut-off" feature, when situated above  
54 the Mediterranean Sea, remains relatively stable and takes advantage of the abundant heat and  
55 humidity from the sea to generate organized convection (Faranda et al., 2022).

56 The structure of a Medicane is characterized by the presence of a central free-cloud "eye", a  
57 strong rotation around the pressure minimum, an eyewall with convective cells, from which  
58 rain bands extend. It, that can be considered like a small-scale tropical cyclone, can lead to sea-  
59 level rise, storm surge and sea waves that can reach significant heights of about five meters  
60 (Miglietta and Rotunno, 2019). The typical Medicane lifetime is limited to a few days,  
61 generally from 2 to 5 days, as a result of the small areal extension of the Mediterranean Sea  
62 that represents their main energy source; for the same reason also the diameter is generally  
63 restricted to a range between 100 and 300 km (Comellas et al., 2021), and their intensity rarely  
64 exceeds the category 1 of the Saffir-Simpson hurricane wind scale (Miglietta and Rotunno.,  
65 2019). In addition, due to the geometrical and meteorological characteristics of the  
66 Mediterranean Sea, a Medicane reaches fully tropical characteristics (a symmetric, deep warm-  
67 core structure and convection in their development and maintenance) for a short time, while  
68 extratropical features (non-symmetrical structure and not well-developed convection around  
69 the core) prevail for most of their lifetime (Miglietta et al., 2011, 2013). As yet, there is no  
70 clear separation between tropical and extratropical cyclones, and the first approach to  
71 differentiate these cyclones was developed by Hart (2003). This method, called cyclone phase  
72 space analysis, relies on a large spectrum of different cyclone types by using several parameters  
73 such as the symmetry/asymmetry and the cold or warm core. In this way, as explained by the  
74 author, an objective classification of cyclone phase is possible, merging the basic structural  
75 description of tropical, extratropical, and hybrid cyclones into a continuum. The favorable

76 months to the Medicanes generations are the autumn and early winter months (from September  
77 to January). Indeed during these months, the Mediterranean Sea preserves high temperatures  
78 after the summer season, and the first cold upper-air troughs are observed, thus creating a high  
79 sea-air temperature gradient (Cavicchia et al., 2014; Nastos et al., 2018). Specifically, the  
80 occurrence of intense convective instability is initiated when the polar jet stream transports  
81 cold air masses over the warmer Mediterranean Sea (Cavicchia et al., 2014; Nastos et al., 2018).  
82 The Medicanes generation during the late-winter months (February and March) is possible but  
83 less common (Cavicchia et al., 2014, Tous and Romero, 2013).

84 These Mediterranean extreme weather events caused damages, floods, deaths, and injuries in  
85 several Mediterranean coastal areas (South France, Central and South Italy, Malta, Balearic  
86 islands, Greece, Crete, Turkey, and some African states; Androulidakis et al, 2022; Bouin and  
87 Brossier, 2020; Carrió et al., 2017; Dafis et al., 2018; Di Muzio et al., 2019; Faranda et al.,  
88 2022; Kerkmann and Bachmeier, 2011; Lagouvardos et al., 2022; Pravia-Sarabia et al., 2021;  
89 Portmann et al., 2020; Rumora et al., 2018; Varlas et al., 2020; Zimbo et al., 2022). As  
90 explained by Cavicchia et al. (2014), the most frequent genesis regions are the Balearic Islands  
91 and the Ionian Sea. In particular, during the last 12 years, the majority of the Medicanes have  
92 been developed over the Ionian Sea and this is probably linked to the sea surface temperature  
93 that, as shown by Shaltout and Omstedt (2014), in the Ionian Sea is constantly 1.0°-1.5° C  
94 higher than that in the Tyrrhenian Sea.

95 Although these extreme Mediterranean events showed significant wave heights (hereinafter  
96 SWH, defined as the average wave height of the highest one-third of the waves), comparable  
97 to the common seasonal storms, they caused greater coastal flooding (Scardino et al., 2022;  
98 Scicchitano et al., 2021). The strong winds, generated during a Medicane, cause the  
99 development of powerful wave motions and lead to an energy transfer from the sea waves to  
100 the solid Earth (Borzì et al., 2022). This energy transfer between the atmosphere, the

101 hydrosphere and the solid Earth is one of the generation mechanisms of the most continuous  
102 and ubiquitous seismic signal on the Earth, called microseism (e.g. Hasselmann, 1963;  
103 Longuet-Higgins, 1950). In connection with the spectral content and the source mechanism  
104 (e.g. Haubrich and McCamy, 1969), it is possible to divide this signal into: primary microseism  
105 (PM), that shows the same period as the oceanic waves (13 - 20 s) and low amplitudes, and is  
106 generated by the energy transfer of oceanic waves breaking against the shoreline (Ardhuin et  
107 al., 2015; Hasselmann, 1963); secondary microseism (SM), generated by sea waves with the  
108 same frequency traveling in opposite directions and exhibiting frequency about twice of the  
109 frequency of the oceanic waves (period of 5 - 10 s) and amplitude higher than the PM (e.g.  
110 Ardhuin and Roland, 2012; Ardhuin et al., 2015; Lepore and Grad, 2018; Longuet-Higgins,  
111 1950; Oliver and Page, 1963); short-period secondary microseism (SPSM), that has a period  
112 shorter than 5 seconds and is generated by the interaction between local wave motions near the  
113 coastline (Bromirski et al., 2005).

114 Several works deal with the relationship between microseism and the sea state (Ardhuin et al.,  
115 2019; Cannata et al., 2020; Guerin et al., 2022; Moschella et al., 2020), while others take into  
116 account specifically the relationship between microseism and cyclonic activity (e.g.,  
117 Bromirski, 2001; Bromirski et al., 2005; Gerstoft et al., 2006; Gualtieri et al., 2018; Lin et al.,  
118 2017; Retailleau and Gualtieri, 2019, 2021; Zhang et al., 2010) considering in particular  
119 typhoons (Lin et al., 2017), tropical cyclones (Zhang et al., 2010), and hurricanes (Gerstoft et  
120 al., 2006). Interestingly, Bromirski (2001) and Bromirski et al. (2005) showed that the  
121 microseism bands most affected by the presence of a cyclone are the SM and SPSM ones.

122 For the first time, the relationship between SM, SPSM, and Medicane was analyzed by Borzì  
123 et al. (2022), who considered the Medicane Apollo to reconstruct both the seismic variation in  
124 terms of power spectral density (PSD), root mean square (RMS) amplitude and the Medicane  
125 position during its lifetime by two different methods (array analysis and grid search method by

126 means of seismic amplitude decay). In this work, we explore the relationship between  
127 microseism and the Sub-Tropical system “Helios” that occurred in the Sicily Channel during  
128 the period 9-11 February 2023 (**Figure 1**). The microseism results are integrated with  
129 wavemeter buoy, HF Radar, hindcast and satellite data (SEVIRI Images) to perform an  
130 investigation as comprehensive as possible of this extreme Mediterranean meteo-marine event.

## 131 **2. Sub-tropical system “Helios” (9-11 February 2023)**

132 During the period 9-11 February 2023, a low-pressure system, later renamed Helios, developed  
133 over the Sicily Channel due to the strong contrast between the very cold air, coming from NE  
134 (Balkans area) and the relatively warm sea surface. From satellite data, the warm core anomaly  
135 of this cyclone is evident, requisite for the development of the Medicane. However, this storm  
136 failed to become a Medicane, for which it is necessary that the cyclone maintains well-  
137 developed convection around the eye, absent in this case probably due to a little interaction  
138 between sea-air caused by sea surface temperature not suitable for the development of a  
139 Medicane (<https://twitter.com/medcyclones/status/1623795373423620096?s=20>;  
140 <https://twitter.com/medcyclones/status/1623992335104081921?s=20>;  
141 <https://twitter.com/medcyclones/status/1624143740800536591?s=20>; last access 23/05/2023;  
142 D’Adderio et al., 2023).

143 In spite of this, Helios, thanks to its proximity to the Sicilian and Maltese coasts, was able to  
144 produce damage along these areas. The effects of the sub-tropical system Helios were  
145 significant, especially in Catania, Ragusa and Siracusa provinces (located in the south-eastern  
146 part of Sicily), the Sicilian Meteorological service (“Regione Siciliana—SIAS—Servizio  
147 Informativo Agrometeorologico Siciliano”, <http://www.sias.regione.sicilia.it/>) recorded heavy  
148 rainfall, more than 200 mm/48 h and peaks of about 500 mm/48 h near Noto for the days 9-10  
149 February 2023, heavy snowfall starting from 1200 m a.s.l. with accumulations of fresh snow  
150 on Etna thicker than two meters, strong wind gusts up to 90 km/h along the exposed coast

151 (Davies, 2023) and severe storm surge with SWH greater than 5 meters. Similar effects were  
152 also recorded in Malta. The minimum pressure value in the cyclone eye amounted to 1002 hPa.  
153 An overview of the positions and the extension of the sub-tropical system Helios is represented  
154 in **Figure 1**.

155 As a consequence of the damage caused by Helios, the Regional Sicilian Government decided  
156 to require a national state of emergency for 12 months for all the municipalities of Catania,  
157 Siracusa, and Ragusa provinces and some municipalities of Messina province  
158 ([https://www2.regione.sicilia.it/deliberegiunta/file/giunta/allegati/N.099\\_15.02.2023.pdf](https://www2.regione.sicilia.it/deliberegiunta/file/giunta/allegati/N.099_15.02.2023.pdf), last  
159 access 23/05/2023).

160

### 161 **3. Data and Methods**

162 We analyzed the data recorded in the period 8 to 13 February 2023 comprising the development  
163 of Helios, the climax in terms of minimum pressure value, wind velocity, precipitation intensity  
164 and SWH, and its decline.

#### 165 **3.1. Seismic data**

166 We used 105 seismic stations installed along the Italian and French coastal areas, in the Sicily  
167 channel coastlines (in Malta, Lampedusa and Linosa islands), in Corsica island and along the  
168 Greek coastal areas to perform spectral analysis, localization analysis by the grid search method  
169 based on seismic amplitude decay and to obtain the seismic signature of the analyzed event  
170 (**Figure 2a** and **Supplementary Table 1**). Three of these stations (IWAV2, IWAV3 and  
171 IWAV5) were installed as part of the i-waveNET “Implementation of an innovative system for  
172 monitoring the state of the sea in climate change scenarios” project, funded by the Interreg  
173 Italia-Malta Programme (<https://iwavenet.eu/>; notice 2/2019 Axis 3; project code C2-3.2-106).  
174 The aim of this project is to set up an innovative sea state monitoring network, integrating  
175 different measurement technologies, such as HF radars, seismic stations, sea level probes, wave

176 buoys and weather stations. Additionally, 15 seismic stations, installed in the Etnean area, were  
177 used to conduct array analysis (**Figure 2b** and **Supplementary Table 2**). The selected seismic  
178 stations show specific characteristics: they are i) installed near the coastal areas and ii) equipped  
179 with 3-component broadband seismic sensors.

## 180 **3.2. Sea state measures**

181 In this work, we use sea state data derived from four independent methodologies. In particular,  
182 we used: i) significant wave height (SWH-Hind), provided by the hindcast maps produced by  
183 Copernicus; ii) significant wave height (SWH-Buoy), period and direction of the waves  
184 measured by the wavemeter buoy installed near Mazara del Vallo; iii) significant wave height  
185 (SWH-HF), period and direction of the waves obtained by the HF Radar installed at the Marina  
186 di Ragusa harbor; iv) SEVIRI Images to spatially and temporally track the position of the  
187 cyclone. These four different data sources have been used to both describe the sea state  
188 evolution during the Helios event, and characterize the physical state of the sea which is strictly  
189 correlated to the microseism derived outputs.

### 190 **3.2.1. Copernicus Data**

191 Regarding the description of the Helios event in terms of spatio/temporal distribution of SWH,  
192 wave period and direction over the whole domain, we referred to the  
193 “MEDSEA\_HINDCAST\_WAV\_006\_012” product, provided by the Copernicus Marine  
194 Environment Monitoring Service (CMEMS) (Korres et al., 2019). The CMEMS product  
195 contains the hindcast maps of the Mediterranean Sea Waves forecasting system and is based  
196 on the third-generation wave model WAM Cycle 4.5.4 composed by hourly wave parameters  
197 at 1/24° horizontal resolution (Korres et al., 2019).



### 198 3.2.2 Wavemeter buoy data

199 Concerning the wave buoy, in **Figure 2c** we show the locations of this buoy, located offshore  
200 of Mazara del Vallo, at a depth of 85 m. The Mazara buoy is managed by ISPRA and is part of  
201 the National Wave Buoy Network (RON). The instrumental equipment consists of buoys  
202 allowing the acquisition of wave parameters in real time. The long time series represent an  
203 important heritage for the knowledge of marine phenomena affecting the Italian seas, both in  
204 terms of climatology and extreme events. The RON National Wave Network is now composed  
205 of 7 stations located off the Italian coasts for the continuous measurement of wave and  
206 meteorological parameters, such as wind direction and speed, atmospheric pressure, water  
207 surface and air temperatures, with real-time data transmission. Until 2014, the ISPRA wave  
208 buoy network was equipped with WatchKeeper™ weather wave meters manufactured by the  
209 Canadian company AXYS Ltd. (Bencivenga et al., 2012). The new meteo-marine buoys were  
210 developed, designed and built in Italy for the specific needs of ISPRA. Data are collected  
211 continuously for periods of 20-25 minutes and are provided every 30 minutes.

212 The parameters recorded by the wavemeter buoy and used in this study are: i) SWH (m), ii)  
213 wave mean period (s) and iii) wave mean direction (°).

### 214 3.2.3. HF Radar data

215 Sea state measures are also provided by the HF system located in the Marina di Ragusa harbor  
216 (**Figure 2c**) which is owned by the CNR-IAS (Consiglio Nazionale delle Ricerche - Istituto  
217 per lo Studio degli Impatti Antropici e Sostenibilità in Ambiente Marino). This HF radar is part  
218 of the CALYPSO HF network operating in the Malta-Sicily channel since 2013. The network  
219 is nowadays composed of seven HF Codar SeaSonde systems transmitting at 13.5 MHz (central  
220 frequency). The network provides sea surface current maps at 3 km of spatial resolution at  
221 hourly scale (Capodici et al., 2019). Each HF radar provides sea state variables (SWH-HF,

222 wave period, wave direction) every 15 minutes; these data are referred to 10 independent  
223 annular rings 3 km wide, centered at the HF site location. Data used in this work regard the last  
224 annular ring (30 km far from the HF site) showing the best temporal continuity of the  
225 measurements. The sea state derived by the HF technology has been deeply validated by several  
226 authors (e.g. Long et al., 2011; Lorente et al., 2021; Orasi et al., 2019; Saviano et al., 2019).

#### 227 **3.2.4. Satellite data**

228 The passage of the Helios cyclone in the study area was tracked by means of the High Rate  
229 SEVIRI Level 1.5 Image Data. The Level 1.5 image data represents the geolocated and  
230 radiometrically pre-processed images that are prepared for subsequent processing steps, e.g.  
231 the extraction of meteorological products helpful in our case study. For further information  
232 about this methodology you can see the EUMETSAT website  
233 (<https://navigator.eumetsat.int/product/EO:EUM:DAT:MSG:HRSEVIRI> last access  
234 25/05/2023). In particular, an image at 10.8  $\mu\text{m}$  each 15' was downloaded and analyzed.

#### 235 **3.3. Spectral Analysis of Microseism**

236 The seismic data were corrected for the instrument response and thereafter spectral and  
237 amplitude analyses were performed. For the spectral analysis, hourly spectra of the seismic  
238 signal were calculated by applying Welch's method (Welch, 1967) with time windows of 81.92  
239 s. The hourly spectra, thus obtained, were gathered and represented as spectrograms, with time  
240 on the x-axis, frequency on the y-axis, and the  $\log_{10}$  of the PSD indicated by a color scale.  
241 Some spectrograms obtained from the vertical component of 4 stations are shown in **Figure 3**.  
242 Concerning the amplitude, we estimated hourly RMS amplitude time series for the typical  
243 microseism frequency bands: 0.2-0.4 Hz (SPSM, **Supplementary Figure 1**), 0.1-0.2 Hz (SM,  
244 black lines in **Figure 3**) and 0.05-0.07 Hz (PM, **Supplementary Figure 2**).

245 In addition, to show the spatial and temporal distribution of the RMS amplitude during the  
246 period under investigation, we plotted the mean RMS amplitude computed on non-overlapped  
247 1-day-long moving windows for the three microseism bands (**Figure 4, Supplementary**  
248 **Figures 3 and 4**). Each dot, in **Figure 4** and in **Supplementary Figures 3 and 4**, represents a  
249 seismic station and the color of the dot relates to the corresponding RMS amplitude at that  
250 location, as specified in the color bar. Noteworthy, the colorbar of the PM band  
251 (**Supplementary Figure 3**) shows a different range of RMS amplitude highlighting a different  
252 response between the PM and the SM and SPSM bands to the sub-tropical system Helios  
253 (**Figure 4** and **Supplementary Figure 4**). Furthermore, in these figures, we compared the RMS  
254 amplitude with the SWH, represented by the contour lines. A detailed description of all the  
255 figures cited in this paragraph is given in section 4.2.

### 256 **3.4 Correlation analysis between microseism amplitude and significant wave height**

257 The calculation of correlation coefficients between the RMS amplitude time series and the  
258 significant wave height time series was conducted in accordance with previous studies (e.g.,  
259 Bromirski, 2001). This calculation was performed for each grid cell of the hindcast maps during  
260 the investigated period to obtain information about the spatial variability of the correlation  
261 coefficients. This kind of analysis provides information about the location of the main sources  
262 of the microseism recorded by the 4 selected stations. To explore the non-linear relationship  
263 between seismic RMS amplitudes and significant wave heights, we followed the approach  
264 proposed by Craig et al. (2016) and employed the Spearman correlation coefficient. The  
265 Spearman correlation coefficient is defined as a nonparametric measure of rank correlation  
266 (Craig et al., 2016). Correlation maps, gathering together the correlation values obtained in the  
267 nodes of the whole Mediterranean Sea, were obtained for the vertical component of each station  
268 and the 3 typical microseism frequency bands (Figure 5).

### 269 3. 5. Tracking Helios position by Microseism

270 Following Borzì et al. (2022), we used two different and complementary methods to track the  
271 position of the sub-tropical system Helios from a seismic point of view. In particular, we use  
272 i) a grid search method based on the seismic amplitude decay and ii) array analysis. These  
273 methods allow us to track the evolution over time of the location of the centroid of the seismic  
274 sources generated by the sea state variations induced by the cyclone. Such seismic data were  
275 compared with the cyclone trajectory which was identified by processing the High Rate MSG  
276 SEVIRI images. This latter task was accomplished by visually identifying the positions of the  
277 cyclone's eye which was clearly identifiable only between 01:00 and 23:00 of the 10 February  
278 2023.

#### 279 3.5.1. Grid search method

280 We used the seismic signals recorded by 105 seismic stations (**Figure 2a**) to map out the  
281 position of Helios during the analyzed period by employing a grid search approach (**Figure 6**).  
282 The region, where we executed the grid search, is a bi-dimensional area of 1760 km x 2400 km  
283 (minimum longitude: 5°; maximum longitude: 30°; minimum latitude: 30°; maximum latitude:  
284 46°) with a spacing of 1°. As shown by several authors who used seismic amplitude decay  
285 methods, both to locate microseism sources (Borzì et al. 2022) and seismo-volcanic sources  
286 (Battaglia and Aki, 2003; Cannata et al., 2013; Kumagai et al., 2011), the grid spacing is chosen  
287 as a compromise between good spatial resolution and reasonable computation time. The  
288 microseism source is localized based on the goodness of the linear regression fit (hereafter  
289 referred to as  $R^2$ ) computed for each node of the bi-dimensional (2D) grid previously  
290 mentioned. Specifically, the source was identified at the centroid position of all the grid nodes  
291 where the  $R^2$  values deviate by no more than 1% from the maximum  $R^2$  value. In this method,  
292 we used an RMS signal window of 4 hours and thus we were able to obtain 1 localization every

293 4 hours. In addition, following Borzì et al. (2022), we applied a method to evaluate the  
294 statistical significance of the retrieved maximum  $R^2$  value and to test the confidence of the  
295 location results. Specifically, we conducted 20 iterations by randomly rearranging the RMS  
296 amplitude values among the stations. Then, we calculated the 95th percentile and we obtained  
297 a value of 0.27. In accordance with this result, we consider reliable the localizations with  $R^2$   
298 values greater than 0.27. To retrieve the errors associated with each localization, the bootstrap  
299 technique is used (Johnson, 2001). This technique consists in recalculating each source location  
300 1000 times, by randomly resampling the data (amplitude-distance pairs) with repetition.  
301 It must be underlined that the grid search method used in this study shows various limits that  
302 in specific cases can invalidate the source locations. In particular, the first limit concerns the  
303 fact that in this method we consider the microseism source as a point-like source, while the  
304 microseism is produced in a wide area of the Mediterranean Sea. In this scenario, the  
305 localization of the point-like source is determined as the barycentric point of the extended  
306 source. However, it is important to consider a limitation of this method related to the presence  
307 of multiple sources with similar amplitude in the same frequency range. In such cases, the  
308 constrained source location shifts towards a position between the actual seismic source  
309 locations (Battaglia et al., 2005), resulting in a significant decrease in  $R^2$ . In our case, we  
310 neglect localization showing  $R^2$  values smaller than 0.27, to avoid unreliable localization.

### 311 **3.5.2. Array analysis**

312 In order to track the location of the sub-tropical system Helios using array techniques, we  
313 considered fifteen stations belonging to the Mt. Etna seismic permanent network and used them  
314 as a roughly circular array (**Figure 2b**).

315 The Array Response function (ARF) is a good tool to plan the array geometry required to  
316 investigate microseism signals or, in this case, to evaluate the performance of a pre-existing

317 array in microseism studies. The ARF, previously obtained by Borzì et al. (2022), exhibits that  
318 the roughly circular array has a good response for the PM and SM cases.

319 In this study, we employed the f-k (frequency-wavenumber) analysis technique on microseism  
320 signals (e.g. Rost and Thomas, 2002). This approach involves a spectral domain beamforming  
321 method that utilizes a grid search of slowness to determine the back azimuth and apparent  
322 velocity values that maximize the amplitude of the combined array traces. The result of the f-  
323 k analysis is the PSD as a function of slowness. In accordance with Borzì et al. (2022), we  
324 followed the subsequent processing steps to implement array analysis on microseism data: (i)  
325 demeaning and detrending; (ii) applying a specific frequency band filter for microseism; (iii)  
326 segmenting the data into tapered windows of 120 seconds each; (iv) excluding windows  
327 containing seismo-volcanic amplitude transients (such as volcano-tectonic earthquakes, long-  
328 period events, and very long-period events) identified using the STA/LTA technique (e.g.,  
329 Trnkoczy, 2012); (v) performing f-k analysis for each window by conducting a slowness grid  
330 search (ranging from -1 to 1 s/km in the east and north components of the slowness vector)  
331 with a spacing of 0.05 s/km. An illustrative example of the outcomes is presented in Figure 6.  
332 For further details about these two methods used in this work, you can see Borzì et al. (2022).

### 333 **3.6. Seismic signature of the Medicanes**

334 In a way to obtain the seismic signature and the main spectral characteristics of the sub-tropical  
335 system Helios, we use a method developed by Soubestre et al. (2018). This method was initially  
336 developed as a network-based method to detect and classify seismo-volcanic tremors. The  
337 proposed method utilizes the coherence of tremor signals within the network, which is  
338 determined based on the array covariance matrix. Using this technique, as explained by  
339 Soubestre et al. (2018), it is possible to highlight both volcanic tremors and other types of  
340 seismic sources such as tectonic earthquakes (local, regional, and teleseismic), and oceanic

341 seismic noise (microseism). This method allows the identification of the spatially coherent  
342 individual noise source within a specific network, identified as small spectral width, as opposed  
343 to other noises such as local effects that would generate multiple individual sources. For further  
344 details about the method, see Soubestre et al. (2018).

345 In this study, we are interested in detecting the microseism produced by the sub-tropical system  
346 Helios. Since we are interested in such a microseism, we filtered the signal in the band 0.1-1  
347 Hz and resample it to 25 Hz in a way to reduce the computation time. To compute the  
348 covariance matrix we use only the vertical component of the seismic signal and a window  
349 length of 60 seconds. The analyses were performed using the data recorded by a station set  
350 comprising both stations installed near the south Sicilian coast (CLTA and IWAV5) and in the  
351 Malta (MSDA) and Linosa (LINA) islands in order to have the microseism source within the  
352 selected station set.

353

#### 354 **4. Results and Discussion**

355 We analyze the sea state, derived by four independent techniques, and the seismic data recorded  
356 by the i) 105 seismic stations represented in **Figure 2a** and by ii) the 15 Etnean seismic stations  
357 (**Figure 2b**) during the period 8-13 February 2023. We chose a period longer than the real  
358 lifetime of the sub-tropical system Helios in a way to include the development, the climax in  
359 terms of meteorological events that occurred during the days 9-10 February, and the following  
360 loss of intensity.

#### 361 **4.1. Sea state**

362 In **Figure 7**, we show the SWH, the mean wave period and the direction time series recorded  
363 during the period 8-13 February 2023 by the buoy of Mazara del Vallo (**Figure 7b, d and f**)  
364 and the HF radar installed in Marina di Ragusa harbor (**Figure 7a, c and e**).

365 The buoy data indicates that the sea storm reached its maximum significant wave height of 3.1  
366 m at 20:00 on 9 February, with a mean direction of  $140^\circ$  indicating that waves were generated  
367 by Scirocco wind, which is the main wave direction for the period under investigation and a  
368 period that varies from 5 to 9 s. The data from HF Radar, installed closer to the cyclone position  
369 than the buoy, show a maximum SWH of about 6 m recorded on 9 February at 22:30, the mean  
370 wave direction was of  $100^\circ$  approximately and the period varies in the range of 5-12 s. Both  
371 datasets allowed defining the time interval of the wave storm which spanned between 8-13  
372 February 2023. The higher SWH measured by the HF radar agrees with the shorter distance  
373 from the cyclone eye ( $\sim 90$  km) of this instrument compared to that of the wave buoy ( $\sim 190$   
374 km).

375 The spatio-temporal distribution of the SWH in the whole area is shown by the hindcast maps  
376 in **Figure 8**. Noticeably that the stronger effect of the sub-tropical system Helios on the sea  
377 state of the Malta-Sicily channel was the increase of the SWH during 9 - 10 February. On 9  
378 February the higher SWH values were recorded mainly in the patch of the sea at the east of  
379 Sicily Island and of the Maltese archipelago, whereas the wave storm invaded the whole Malta-  
380 Sicily channel on 10 February. Finally, on 11 February the wave storm started moving to the  
381 south, reaching the northern part of the African coasts.

382 The comparison between SWH-Hind and SWH-HF revealed a good agreement ( $R^2 \sim 0.85$ ) even  
383 if an underestimation of the event by the CMS model (slope of  $\sim 0.7$ ) was observed  
384 (**Supplementary figure 6**).

#### 385 **4.2. Spectral Analysis and RMS spatial distribution**

386 To perform the spectral analysis, we used the 105 seismic stations installed along the Italian,  
387 Greek and Maltese coastal areas. In **Figure 3**, we plot the spectrograms and the RMS amplitude  
388 time series (black lines), obtained by analyzing the vertical component of the seismic signals  
389 recorded by four stations installed in Malta (MSDA **Figure 3a**), on Linosa Island (LINA



390 **Figure 3b**), near Pozzallo (IWAV5 **Figure 3c**) and in Central Italy (CELB **Figure 3d**). We  
391 chose these four stations to compare the results obtained from the stations installed near the  
392 sub-tropical system (LINA, MSDA and IWAV5), with the result obtained from the far station  
393 (CELB).

394 As shown by Borzì et al. (2022), in the spectrograms a great part of the energy is focused in  
395 the 0.1-1 Hz band, corresponding with the SM and SPSM bands. In addition, it is also evident  
396 how the considered stations show a different behavior, in terms of both spectrograms (**Figure**  
397 **3**) and RMS amplitude time series (**Figure 3** and **Supplementary Figures 1** and **2**), that  
398 depended on their position. In particular, spectrograms and RMS amplitude time series  
399 obtained from the data recorded by MSDA (**Figure 3a** and **Supplementary Figure 1a** and **2a**),  
400 LINA (**Figure 3b** and **Supplementary Figures 1b** and **2b**) and IWAV5 (**Figure 3c** and  
401 **Supplementary Figures 1c** and **2c**) stations, installed close to Helios (**Figure 1** and **Figure**  
402 **2**), show the maximum PSD and RMS amplitude values during the time interval 9-11 February  
403 2023, highlighted by the vertical dashed lines in **Figure 3**. On the other hand, the spectrogram  
404 and RMS amplitude time series of the station CELB (**Figure 3d** and **Supplementary Figures**  
405 **1d** and **2d**), installed in the Tyrrhenian area, exhibited the maximum PSD and RMS amplitude  
406 values a few days before Helios at the same time as a local storm surge.

407 To show the space-time distribution of the RMS amplitude, we calculate the daily RMS average  
408 for the three main microseism bands (PM and SPSM in **Supplementary Figures 3** and **4**, SM  
409 in **Figure 4**) during the period 8-13 February 2023. All the three analyzed microseism bands  
410 show a relationship with the position of Helios indicated by the five-pointed red star (**Figure**  
411 **4c** and **Supplementary Figures 3c** and **4c**). In particular, on 10 February 2023, when Helios  
412 reached its climax, the maps in **Figure 4c** (SM) and **Supplementary Figures 3c** and **4c** (PM  
413 and SPSM respectively) show a cluster of high RMS values for the stations installed near the

414 sub-tropical system Helios, highlighting a good match between the SM, SPSM, PM and Helios  
415 positions.

416 Comparing the RMS amplitude time series obtained for the sub-tropical system Helios with  
417 those obtained for the Medicane Apollo (Borzi et al., 2022), we observe a similar trend for the  
418 SM and SPSM bands and a different behavior for the PM band. In particular, during the  
419 Medicane Apollo, the analysis showed a significant amplitude increase for the SM and SPSM  
420 bands while no significant changes for the PM band were observed. Instead, during the sub-  
421 tropical system Helios, the RMS amplitude time series for the PM (**Supplementary Figure 2**),  
422 although with two orders of magnitude smaller, showed a trend similar to the SM (**Figure 3**)  
423 and SPSM (**Supplementary Figure 1**) ones. Also the space-time distribution shows a good  
424 match between the PM (**Supplementary Figure 3**) and Helios position (**Figure 1, 4c** and  
425 **Supplementary Figures 3c** and **4c**), even if for this band there are stations installed in central  
426 Italy that show high RMS amplitude values related to a very local storm surge that occurred at  
427 the same time as Helios.

428 The involvement of the PM band in the case of the sub-tropical system Helios can be explained  
429 by considering the position of this low-pressure system (**Figure 1**) and comparing it with the  
430 Apollo position. Indeed, the Medicane Apollo develops and moves in the Ionian Sea (**Figure**  
431 **2**) in an area with a sea depth greater than 2000 m, while the sub-tropical system Helios  
432 develops and moves largely in the Sicily Channel, a shallower sea with a depth that reaches a  
433 maximum depth of 500 m b.s.l. and an average depth of 316 m b.s.l. In addition, the average  
434 depth between Sicily and Malta is about 65 m b.s.l. As outlined in the literature (Bromirski et  
435 al., 2005), the production of the PM is hindered in large water depths due to the attenuation of  
436 pressure fluctuations, which generate the signal, as a result of depth-dependent amplitude  
437 decay. More specifically, the PM is generated solely in depths less than half of the wavelength  
438 (where  $\lambda$  represents the wavelength of the oceanic waves generating the pressure fluctuations).

439 If we consider, for the Catania and Mazara areas, a mean waves period of 6.1 and 5.6 s and a  
440 peak waves period of 9.7 and 9 s respectively (Agenzia per la Protezione dell’Ambiente e per  
441 i Servizi Tecnici Dipartimento Tutela Acque Interne e Marine Servizio Mareografico - Atlante  
442 delle onde nei mari italiani), by utilizing the correlation that connects period and wavelength  
443 ( $\lambda = gT^2/2\pi$  with  $g$  acceleration of gravity that is  $9.8 \text{ m/s}^2$  and  $T$  the period of the waves;  
444 Sarpkaya and Isaacson, 1981) we obtain a wavelength, for this part of the Mediterranean sea,  
445 ranging between  $\sim 45 \text{ m}$  and  $\sim 150 \text{ m}$ . Considering these wavelengths, the shallow depth of the  
446 Sicily Channel, especially in the Malta Channel where the average depth is about  $65 \text{ m b.s.l.}$ ,  
447 and the fact that the generation of the PM occurs only for depths less than  $\frac{1}{2} \lambda$  (Bromirski et  
448 al., 2005), we can remark that the generation of the PM is possible in the Sicily Channel, while  
449 can not occur in the Ionian Sea except in limited areas near the coastline. This is evident in our  
450 analysis, and in particular in the RMS amplitude time series. Indeed in the sub-tropical system  
451 Helios case, which occurred in the Sicily Channel, we can note a similar trend between the  
452 three analyzed microseism bands (**Figure 3** and **Supplementary Figures 1** and **2**), while in the  
453 case of the Medicane Apollo, that developed in the Ionian Sea (Borzì et al., 2022), we observe  
454 an RMS amplitude increase only for the SM and SPSM bands and no significant variations in  
455 the PM bands.

### 456 **4.3. Comparing the RMS amplitude with the SWH**

457 As mentioned before, microseism is a continuous seismic signal linked to the hydrosphere-  
458 solid Earth energy transfer and, as demonstrated by several authors (e.g. Ardhuin et al., 2012;  
459 Bromirski et al., 1999; Bromirski et al., 2005; Cutroneo et al., 2021; Ferretti et al., 2013, 2018),  
460 the amplitude of this signal is strictly related to the sea state and in particular to the SWH. To  
461 analyze the relationship between microseism and SWH, we plot in **Figure 9a** the RMS  
462 amplitude time series for the station IWAV5 and the SWH time series derived from both HF  
463 Radar (SWH-HF) and hindcast maps (SWH-Hind), and in **Figure 5** the results of the

464 correlation analysis between the RMS amplitude time series and the significant wave height  
465 (the position of the seismic stations IWAV5, LINA, MSDA and CLTA and HF radar are  
466 respectively shown in **Figure 2a** and **2c**). The sea state information provided by the Mazara del  
467 Vallo buoy was not taken into account in this analysis because of the long distance between  
468 this instrument and the cyclone eye. We chose stations IWAV5, LINA, MSDA and CLTA both  
469 because they are some of the nearest stations to the cyclone eye and since these stations were  
470 used for the seismic signature analysis. In addition, as it is possible to see in **Figure 3a, b** and  
471 **c**, all the seismic stations installed in the Sicily Channel area show very similar microseism  
472 amplitude patterns. The time series of SWH-Hind was obtained by computing the median value  
473 of the SWH data within a wide area of the Sicily Channel shown in **Figure 9d**. In general, the  
474 three datasets exhibit a good agreement among them. In particular, we observe an increase that  
475 occurs almost simultaneously for SWH-HF, SWH-Hind and the RMS amplitude. However,  
476 from the first hours of 10 February, the SWH-HF started showing a different behavior  
477 compared to SWH-Hind; indeed, the SWH-HF decreased while the SWH-Hind continued to  
478 show high values up to the end of 10 February. This difference can be explained by considering  
479 that the HF Radar provides information about the SWH for a limited area while SWH-Hind  
480 gives median information about a wide area of the Sicily Channel. It is interesting to note that  
481 the microseism amplitude follows more closely the areal sea state, shown by the SWH-Hind  
482 series, than the punctual one, shown by the SWH-HF. We can note this in both a qualitative  
483 way in the time series of **Figure 9a** and a quantitative way in the cross-plots of **Figures 9b** and  
484 **c**. For both the cross-plots we calculated the  $R^2$  value, to evaluate the goodness of the linear  
485 regression, and we obtain  $R^2$  values equal to 0.68 and 0.85, for the cross-plot RMS amplitude  
486 vs SWH-HF and RMS amplitude vs SWH-Hind, respectively. The higher value of  $R^2$  for the  
487 RMS amplitude - SWH-Hind relationship can be explained by considering that microseism  
488 recorded by a seismic station is generated by multiple extended sources distributed on a wide

489 portion of the sea. The correlation analysis confirms a good match between the spatial  
490 distributions of SWH and RMS amplitudes recorded by the 4 aforementioned stations during  
491 the period under investigation (**Figure 5**). Values of the correlation factor higher than 0.85  
492 were observed for the SM and SPSM bands for the area of the Sicily Channel and Ionian Sea  
493 affected by the storm surge due to the cyclone Helios.

494

#### 495 **4.4. Localization analysis**

496 As concerns the microseism source location, we performed the location analysis for the PM,  
497 SM and SPSM using both the array techniques and the grid search method based on the seismic  
498 amplitude decay. Concerning the array techniques, we chose to focus on PM and SM bands  
499 since, according to the information from the ARF, we expect reliable localizations only for  
500 these two bands, while for the SPSM band appears spatial aliasing. As for the grid search  
501 method, we obtained reliable locations only for the SM and SPSM bands, while for the PM  
502 band, although we got localizations in agreement with the cyclone position, the associated  $R^2$   
503 values turned out to be slightly lower than the threshold from which we can consider the  
504 locations reliable.

505 By the grid search method, we obtained reliable locations from 9 February 2023 at 8:00 to 11  
506 February 2023 at 00:00 (**Supplementary figure 7**). In agreement with satellite images (**Figure**  
507 **1**), on 9 February 2023, the sub-tropical system Helios was not yet well-developed and did not  
508 show the cyclone eye. Indeed, during this day, we are able to locate a storm surge, probably  
509 linked to the primitive formation of Helios, that occurred in the Ionian Sea. In particular, our  
510 localizations, on 9 February from 8:00 to 20:00, indicate the source position near to the south-  
511 eastern part of Sicily in agreement with the hindcast data showing for this area SWH greater  
512 than 5 meters (**Figure 4b and 7**). During the following hours, our localizations  
513 (**Supplementary figure 7**) show a small but continuous shift of the source toward the Malta

514 area (**Figure 6**), in agreement with the relatively stable position of the cyclone eye for the first  
515 hours of 10 February 2023, retrieved from SEVIRI data (**Figure 1**). Successively, the sub-  
516 tropical system rapidly loses its strength and runs out completely on the day of 11 February  
517 2023, making landfall against the Libyan coast. The errors associated with the localization  
518 show an anti-correlation with the  $R^2$  (**Supplementary Figure 5**). In particular, we can observe  
519 errors that vary from 535 to 210 km for the longitude and from 245 to 165 km for the latitude  
520 In particular, we obtained lower errors during the climax of Helios, when the  $R^2$  reached the  
521 highest values, and higher during the initial and final phases of the phenomenon  
522 (**Supplementary Figure 5 and Table 1**).

523 In **Table 1** we summarised the main features of the microseism source located by using the  
524 grid search method during the days 9-10 February 2023 and compared these with the results  
525 obtained during the period 01:00 - 23:00 of 10 February 2023 in terms of the coordinates of the  
526 cyclone eye retrieved by SEVIRI data. During the first hours of the cyclone life, the coordinates  
527 of the cyclone eye and the microseism source show only a small offset that, as explained in  
528 section 3.4.1, can be due to the fact that the point-like microseism source corresponds with the  
529 barycentric position of an extended source and it is expected that the microseism source  
530 location could differ from the cyclone eye. During the following hours, as shown in **Figure 1**,  
531 the cyclone moved southward until the time when the cyclone made landfall against the Libyan  
532 coast. This shift is not visible in the microseism location results, probably due to the lack of  
533 seismic stations in Africa, that would help locate more accurately seismic sources placed close  
534 to the African Coastlines. In **Supplementary figure 8** we plot the temporal distribution of the  
535  $R^2$  values and compare this with the SWH-Hind time series. These two datasets show a good  
536 agreement and highlight that the higher  $R^2$  values obtained from the grid search method are  
537 influenced by the presence of the sub-tropical system Helios.

538 Concerning the array analysis, the result obtained for the PM band shows that, for the period  
539 9-11 February, the back azimuth values pointed toward the Ionian Sea (**Supplementary figure**  
540 **9**) with apparent velocity values of  $\sim 3.0$  km/s (**Supplementary figure 10**). For the same days,  
541 the back azimuth values for the SM band pointed toward the same region (**Supplementary**  
542 **figure 11**), with apparent velocity values of  $\sim 2.0$  km/s (**Supplementary figure 12**).

543 For the PM and the SPSM bands, we obtained reliable locations only by one of the two used  
544 methods (array technique for the PM and grid search method for the SPSM), while for the SM  
545 we obtained reliable locations from both the aforementioned methods and we can compare the  
546 results obtained independently from these two methods. We can observe that the grid search  
547 method based on the seismic amplitude decay is able to detect both the storm surge that  
548 occurred in the Ionian sea and the following formation of the sub-tropical system Helios  
549 (**Supplementary figure 7**), and in particular by this method, we localize the cyclone as a point-  
550 like source, considered as a barycentre of an extended source. The array technique instead  
551 provides back azimuth values pointing toward the Ionian Sea for the entire period of 9-11  
552 February (**Supplementary figure 11**). From these results, we can note that the two methods  
553 are influenced by different sources. This different result can be explained based on the spatial  
554 station distribution. Indeed, if we consider the grid search method we use a wide station  
555 network including stations installed near the sub-tropical system Helios (MSDA, CLTA,  
556 IWAV5, LINA and other stations), while the Etnean array includes 15 stations clustered in a  
557 small area. In addition, the distance array center-Ionian Sea ( $\sim 20$  km) is smaller than the  
558 distance array center-Sicily Channel ( $\sim 90$  km). During the period taken into account, we have  
559 the coexistence of two strength sources, the first in the Ionian Sea (storm surge with SWH  
560 greater than 5 m) and the other, probably the strongest, in the Sicily Channel (Helios), both  
561 represented by the red contour line in **Figure 4c**. Hence, the coexistence of two sources and  
562 the great difference in distance between the array center and the Ionian Sea and the array center

563 and the Sicily channel does not allow us to locate the sub-tropical system Helios with the array  
564 technique, which gives us back azimuth values pointing always toward the nearest source.

#### 565 **4.5. Seismic signature**

566 To show the main spectral characteristics of the sub-tropical system Helios, we used the method  
567 developed by Soubestre et al. (2018). From our analysis, the evolution in time of the  
568 microseism spectral characteristics between the first 24 hours and the following ones is evident  
569 (**Figure 10**). In particular, with the vertical dashed lines in **Figure 10**, we show the time interval  
570 when Helios develops, reaches the climax and loses power to run out. The first 24 hours are  
571 characterized by the lack of a clear predominance of a particular frequency range. Starting from  
572 early 9 February 2023 (first vertical white dashed line), we can observe high coherence values  
573 clustered on a narrow frequency range between about 0.14 and 0.25 Hz, and this frequency  
574 range is constant until the 80 hours (vertical black dashed line) corresponding to the morning  
575 of 11 February. Successively, the frequency with the highest coherence starts to increase  
576 reaching a value of about 0.35 Hz before the exhaustion of the phenomenon. This increase in  
577 the frequency could be related to the landfall of Helios against the Lybian coast  
578 (<https://twitter.com/medcyclones/status/1624143740800536591?s=20>, last access 23/05/2023  
579 ). Indeed, as described in the literature (Gerstof et al., 2006; Lin et al., 2017; Sun et al., 2013),  
580 there exists a relationship between the frequency, the sea depth and the development of local  
581 wave motion near the coastline. In particular, Gerstof et al. (2006) show an increase in the  
582 microseism frequency during the two landfalls of Hurricane Katrina. Similar results are shown  
583 by Sun et al. (2013), who highlighted an increase in the intensity of the microseism for the SM  
584 and SPSM and the frequency increase from the SM band toward the SPSM one during the  
585 approaching of the three analyzed typhoons against the Chinese coast. Finally, Lin et al. (2017)  
586 show an increase in the microseism frequency during the first landfall of typhoon Megi.  
587 Similarly, we obtained an increase in frequency during the time interval when Helios



588 approached the Lybian coastline and made landfall on 11 February  
589 (<https://twitter.com/medcyclones/status/1624143740800536591?s=20>, last access  
590 23/05/2023). We were also able to observe the gradual loss of power of the sub-tropical system  
591 highlighted by ever-lower coherence values until its disappearance.

## 592 **5. Conclusions**

593 Several works have dealt with the relationship between microseism and several meteorological  
594 phenomena, considering in particular storm surges (Ardhuin et al., 2019; Cannata et al., 2020;  
595 Guerin et al., 2022; Moschella et al., 2020) and different types of cyclones (Borzì et al., 2022;  
596 Bromirski, 2001; Bromirski et al., 2005; Gerstoft et al., 2006; Gualtieri et al., 2018; Lin et al.,  
597 2017; Retailleau and Gualtieri, 2019, 2021; Zhang et al., 2010) that affect various parts of the  
598 world (hurricanes, typhoons, tropical cyclones and medicanes). In this work, we analyzed the  
599 relationship between the three main microseism bands (PM, SM and SPSM) and the sub-  
600 tropical system Helios that occurred in the Mediterranean Sea during the time interval 9-11  
601 February 2023. Although all the meteorological parameters suggest that Helios has not been  
602 able to reach the fully Medicanne characteristics, remaining a rather weak sub-tropical system,  
603 the proximity of this cyclone to the southeastern Sicilian and Maltese coastal area has caused  
604 heavy rainfall, strong wind gusts and violent storm surge in the two above mentioned areas.

605 To obtain information about the sea state we take into account the data of hindcast maps,  
606 wavemeter buoy, HF radar and SEVIRI images. The first three methodologies show an increase  
607 in the SWH during the period under investigation, with a climax during the days 9-10 February  
608 2023, due to the presence of the sub-tropical system Helios. In particular, the HF radar and the  
609 hindcast maps exhibit an SWH of about 6 meters while the wavemeter buoy shows an SWH of  
610 about 3 meters. This difference is only linked to the position where the two instruments are  
611 installed, indeed the HF radar is installed at about 90 km from the cyclone eye while the  
612 wavemeter buoy is installed at about 190 km. The last method based on SEVIRI images provide

613 information about the location of the cyclone eye that is clearly visible between 01:00 and  
614 23:00 on 10 February 2023.

615 To analyze this meteorological phenomenon from a seismic point of view, we selected 120  
616 seismic stations installed along the Italian, Maltese and Greek coastal areas: i) 105 were used  
617 in spectral and amplitude analysis, in the grid search method and 4 of these to obtain the seismic  
618 signature of Helios using the method of the covariance matrix; ii) 15 were employed in array  
619 analysis. The results, obtained from the spectral analysis, highlight that the seismic signals, in  
620 particular the PM, SM and SPSM bands, are affected by the storm surge and by the sub-tropical  
621 system Helios. This is evident considering the RMS amplitude time series, the spectrograms  
622 and the RMS amplitude space-time distribution, in which it is possible to observe that the  
623 amplitude of the microseism signal, in the three main bands above mentioned, shows a similar  
624 trend and it is greater during the period 9-10 February 2023 for the stations installed close to  
625 the Sicily Channel (for example the stations MSDA, LINA and IWAV5), while the stations  
626 installed farther from Helios (for example CELB) show a different behavior conditioned by  
627 local sources. Furthermore, the increase of the RMS amplitude for the PM band can be  
628 explained on the basis of the position where Helios develops. Indeed, as explained in the  
629 literature (Bromirski et al., 2005), the generation of the PM is limited in the areas where the  
630 depth of the sea bottom is smaller than  $\frac{1}{2}\lambda$  (where  $\lambda$  is the wavelength of the sea waves) as a  
631 consequence of the amplitude decay of the pressure fluctuations. The Sicily Channel and in  
632 particular the area between Malta and Sicily (Malta Channel), where Helios stood still for the  
633 entire 10 February 2023, shows an average depth of about 65 m b.s.l., and considering  
634 wavelength between  $\sim 45$  m and  $\sim 150$  m it is possible the generation of the PM for this area.  
635 By comparing the SWH recorded by HF Radar, SWH retrieved by hindcast data and the RMS  
636 amplitude recorded at the station IWAV5, we observe that the RMS amplitude time series  
637 shows a trend more similar to that of hindcast data than to the HF Radar data. This could be

638 explained by considering that the microseism is a seismic signal generated by multiple and  
639 extended sources in a large area of the sea and hence its amplitude is related to the state of a  
640 wide portion of the sea. Also, the correlation analysis confirms a good match between the  
641 spatial distributions of significant wave heights and seismic RMS amplitudes.

642 We used two different methods to track the position of the sub-tropical system Helios during  
643 its lifetime and, through the performed analysis, we were able to locate both the storm surge  
644 that occurred in the Ionian Sea and Helios. In particular, using the grid search method based on  
645 the seismic amplitude decay we located the storm surge in the Ionian Sea on 9 February 2023  
646 and the sub-tropical system Helios on the next day, while with the array technique, we located  
647 only the storm surge in the Ionian Sea as a consequence of the position of the array closer to  
648 the Ionian Sea than to the Sicily Channel. The location obtained both for the storm surge and  
649 for the sub-tropical system Helios, during the first hours of its life, is in agreement with the real  
650 position of the two meteorological phenomena shown by the hindcast maps and satellite  
651 images. Successively, the sub-tropical system Helios moved southward until the time when the  
652 cyclone made landfall against the Libyan coast while the microseism source continued to show  
653 a stable position near Malta Island. This difference between the two positions is probably due  
654 to the lack of seismic stations in Africa, that does not locate accurately microseism sources  
655 placed close to the African Coastlines. In addition, using the method of the covariance matrix,  
656 we obtained the seismic signature of the sub-tropical system Helios. In particular, during the  
657 development and climax of the cyclone Helios, we can observe high coherence values clustered  
658 on a narrow frequency range between about 0.14 and 0.25 Hz that, as described in the literature,  
659 corresponds to the microseism bands (SM and SPSM) most affected by cyclonic activity. This  
660 narrow frequency range stays constant until the time when Helios makes landfall when we  
661 observe an increase in the frequency until a value of about 0.35 Hz probably linked to the  
662 decrease of the sea depth and the development of local wave motion near the coastline.

663 Starting from Borzì et al. (2022), this work aims at studying and monitoring the Mediterranean  
664 cyclones through microseism and its integration with sea state data. It underlines that it is  
665 possible to extract information about these meteorological phenomena by an innovative system  
666 for the sea state monitoring that includes not only the classical instruments (such as wavemeter  
667 buoys, radar HF and geostationary satellites) but also seismometers. In particular, the large  
668 number of broadband seismic stations, installed for earthquake and volcanic monitoring, can  
669 compensate for the lack of data of the classical instruments mentioned above, more often  
670 affected by instrument breakage.

671 Finally, since we obtained the seismic signature of this particular Mediterranean cyclone we  
672 can compare it with the seismic signature of other Mediterranean events (Medicane and  
673 common storms) to identify the similarities and differences in the spectral content of different  
674 Mediterranean cyclones and other more common events. The characterization of the seismic  
675 signature of these events could be useful to identify Mediterranean cyclones by old  
676 seismograms, and hence to reconstruct the temporal variability (in terms of occurrence rate and  
677 intensity) of these extreme meteo-marine events whose evolution seems to be strictly linked to  
678 the global warming (e.g. Emanuel, 2005; Reguero et al., 2019).

679

680

#### 681 **Data availability**

682 The seismic data, in the miniseed format used in this study, can be downloaded through the  
683 ORFEUS-EIDA database (<http://www.orfeus-eu.org/data/eida/>). All the seismic stations used  
684 in this study, along with their main features, are reported in Supplementary Tables 1 and 2. The  
685 hindcast data are available online on the Copernicus site  
686 (<https://resources.marine.copernicus.eu/products>). The Italian buoy data (Mazara del Vallo) are  
687 available on the ISPRA website (<https://dati.isprambiente.it/>) and on the MAREOGRAFICO

688 website ([www.mareografico.it](http://www.mareografico.it)). Satellite data are available on the EUMETSAT website  
689 (<https://navigator.eumetsat.int/product/EO:EUM:DAT:MSG:HRSEVIRI>). The HF Radar data  
690 are available on the CALYPSO project website by compiling the form indicated on the web  
691 page ([https://www.calypsosouth.eu/index.php/welcome/open\\_page/50/0](https://www.calypsosouth.eu/index.php/welcome/open_page/50/0)) or by contacting  
692 [giuseppe.ciraolo@unipa.it](mailto:giuseppe.ciraolo@unipa.it) or [salvatore.aronica@cnr.it](mailto:salvatore.aronica@cnr.it).

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#### 974 **Authors Contribution**

975 A.M.B., A.C. and F.C. conceived the idea. A.M.B. analyzed seismic data and V.M. performed  
976 array analysis. C.L.R, G.N., A.O. and M.P. analyzed buoys data. S.A., D.C., G.D.G., G.L.,  
977 S.D., R.D. and T.L. helped perform seismic analysis. F.C., G.C., S.A., I.F., G.G. G.M.,  
978 analyzed HF Radar Data. A.M.B. and V.M. wrote the paper. S.A., F.C., D.C., G.L. and V.M.  
979 dealt with the new seismic installation. G.C. lead one of the projects funding this research, G.S.



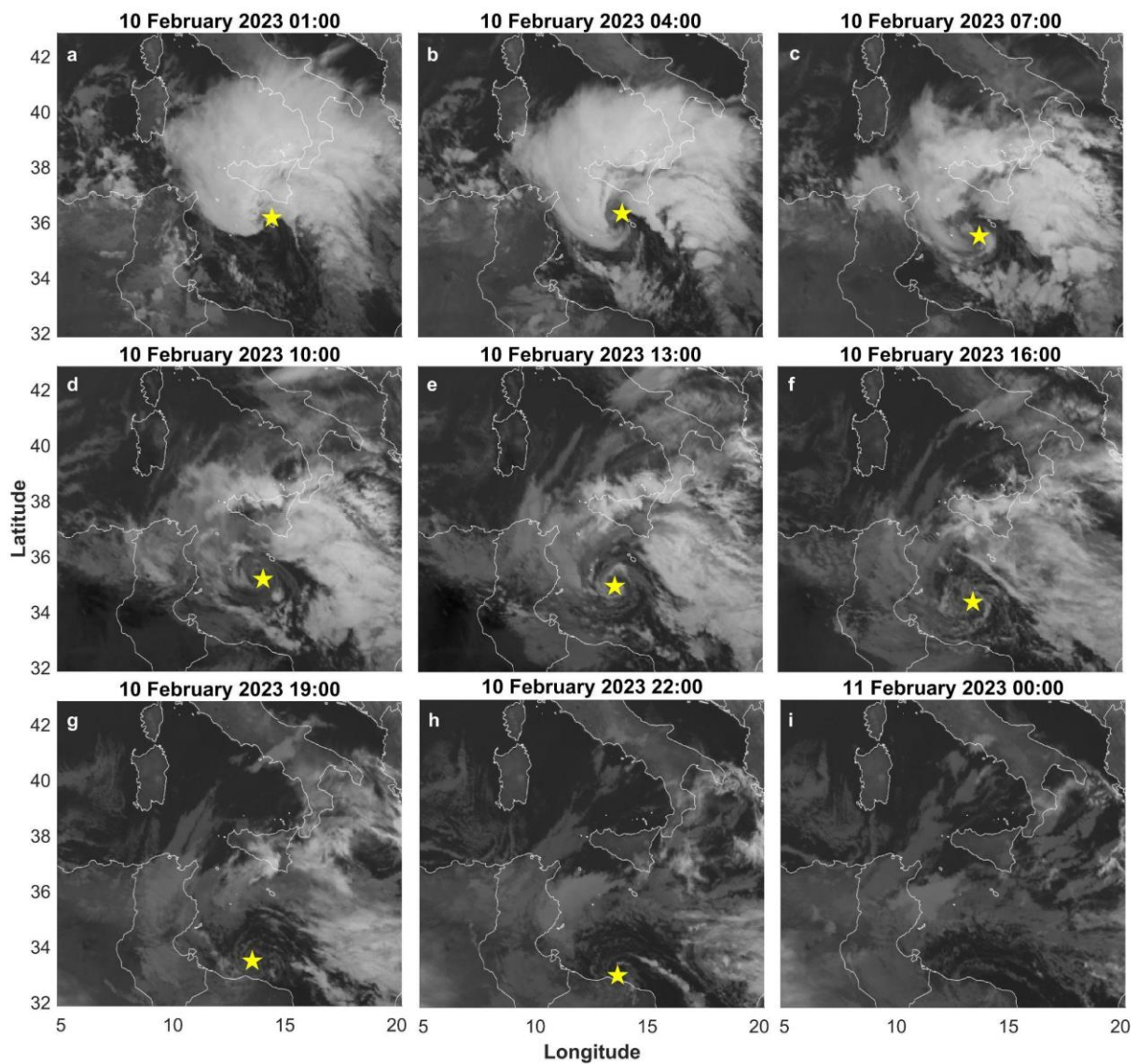
980 helped to interpret the sea state data. All the authors discussed the results, edited the paper and  
981 gave consent for this publication under the supervision of A.C.

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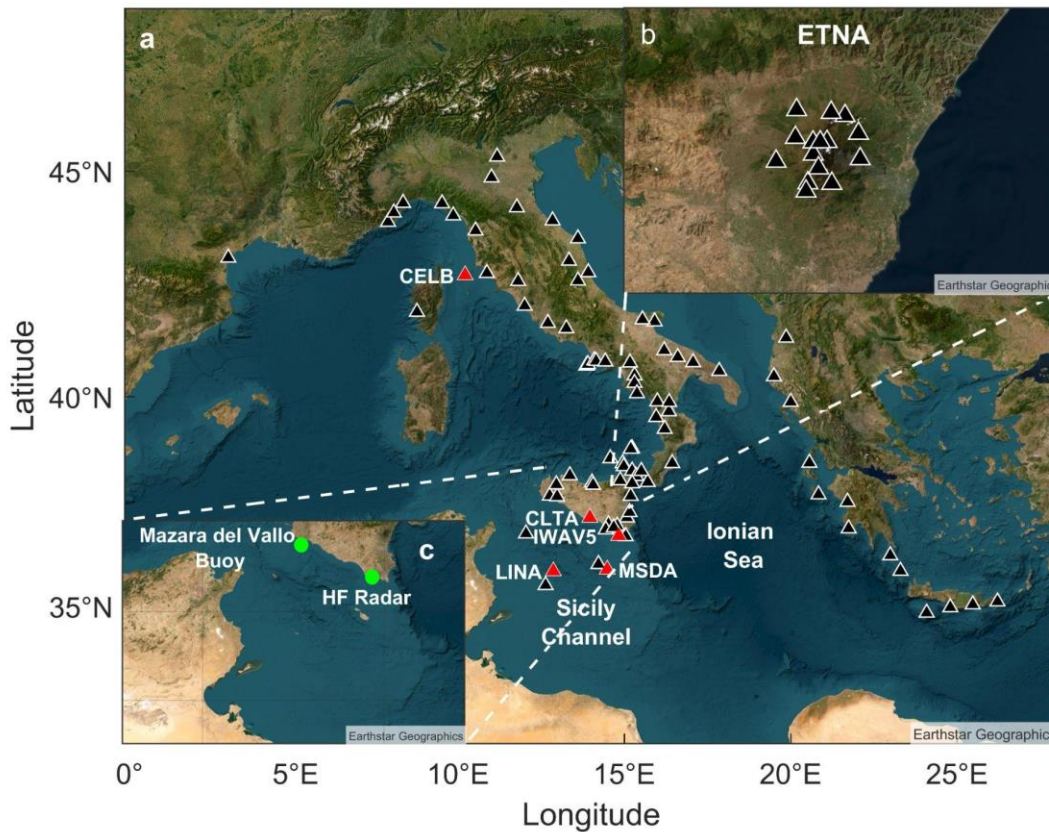
985 **Figures**



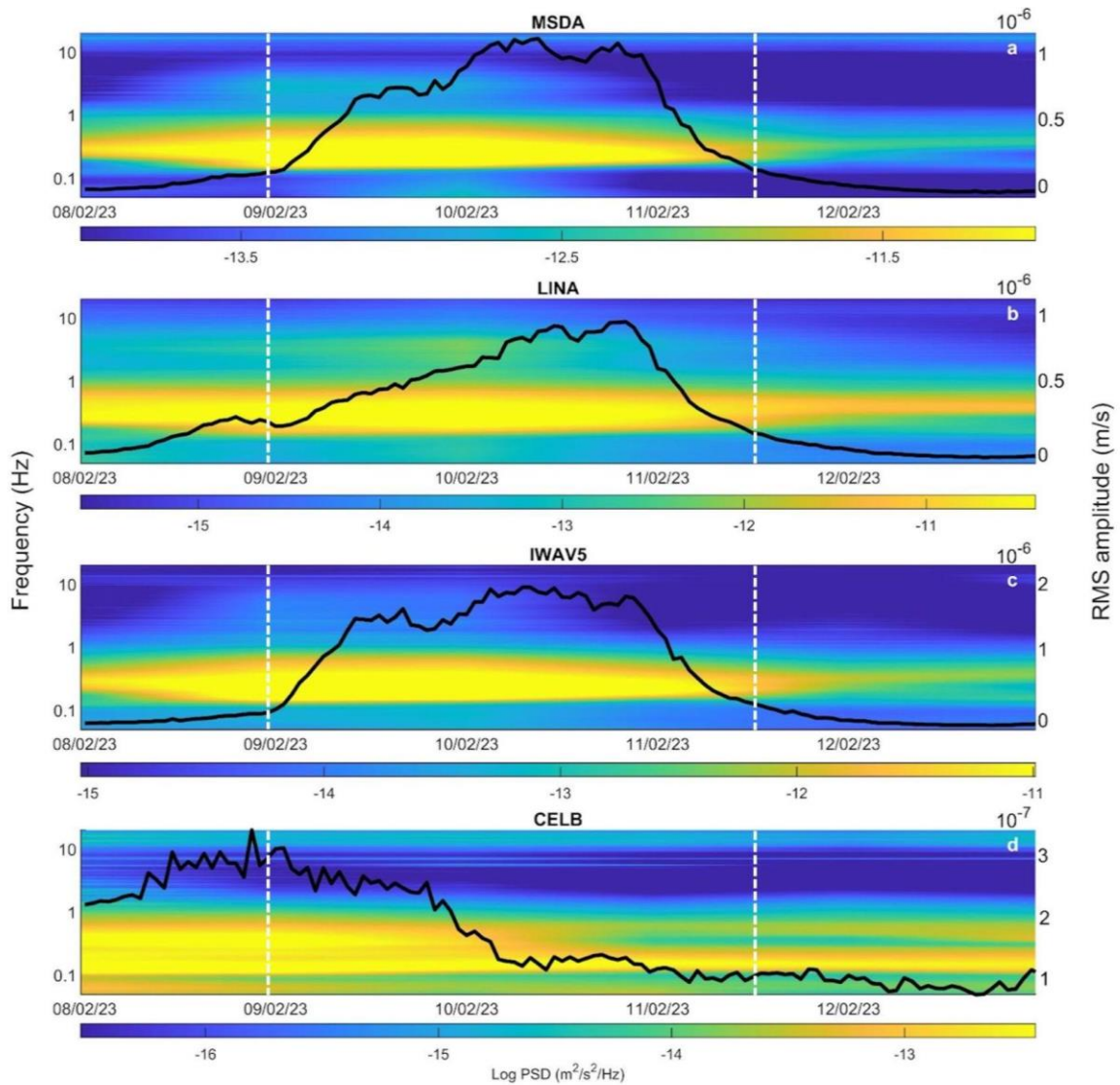
986

987 **Figure 1:** Satellite images of the Mediterranean area and of the sub-tropical system Helios  
988 during the period 10/02/2023 01:00 - 11/02/2023 00:00. The yellow stars in (a-h) show the

989 position of the cyclone eye. The lack of the star in (i) is due to the dissipation of Helios after  
990 the landfall against the Libyan coast (h). (©EUMETSAT SEVIRI Images)



991  
992 **Figure 2:** Satellite image of the Mediterranean area with a selection of the broadband seismic  
993 stations available in the ORFEUS and INGV databases and used in the spectral analysis and in  
994 the grid search method (a) and selection of the broadband seismic stations in the Etna area  
995 maintained by INGV-OE (b), used in the array analysis (base image source ©Earthstar  
996 Geographic). The red triangles indicate the stations used in the detailed analysis shown in  
997 Figures 3 and 9 and in **Supplementary Figures 1 and 2**. The green dots in (c) indicate the  
998 position of the wavemeter buoy (Mazara del Vallo) and of the HF Radar (Marina di Ragusa)  
999 used for the sea state monitoring.

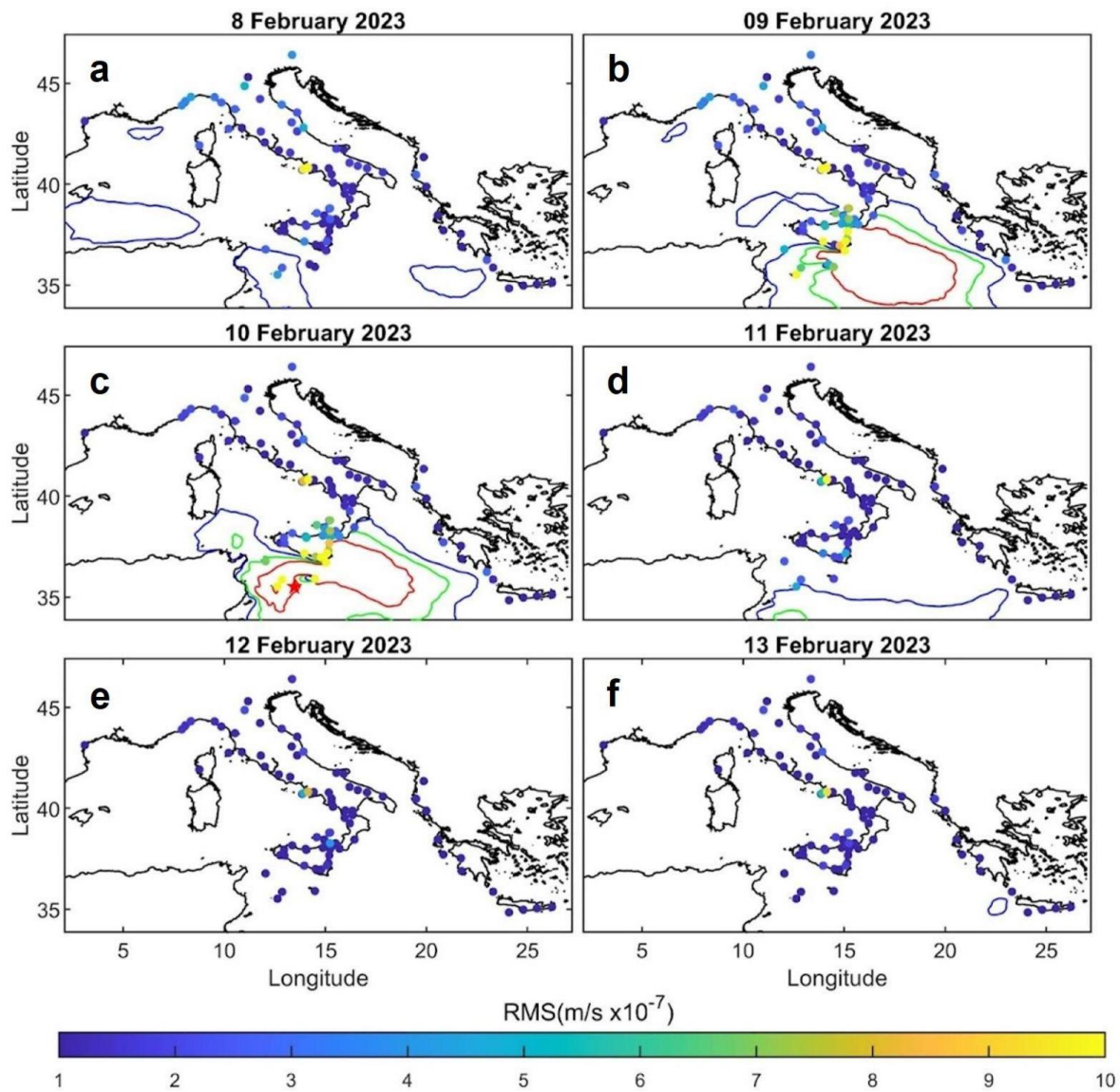


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1001 **Figure 3:** Spectrograms and RMS amplitude time series (black lines) for the SM band (0.1-0.2  
 1002 Hz) of the seismic signal recorded by the vertical component of 4 stations located along the  
 1003 Maltese coastline (a), in Linosa Island (b), in the southern part of Sicily (c) and in Central Italy  
 1004 (d) (see **Figure 2a** for the station locations).

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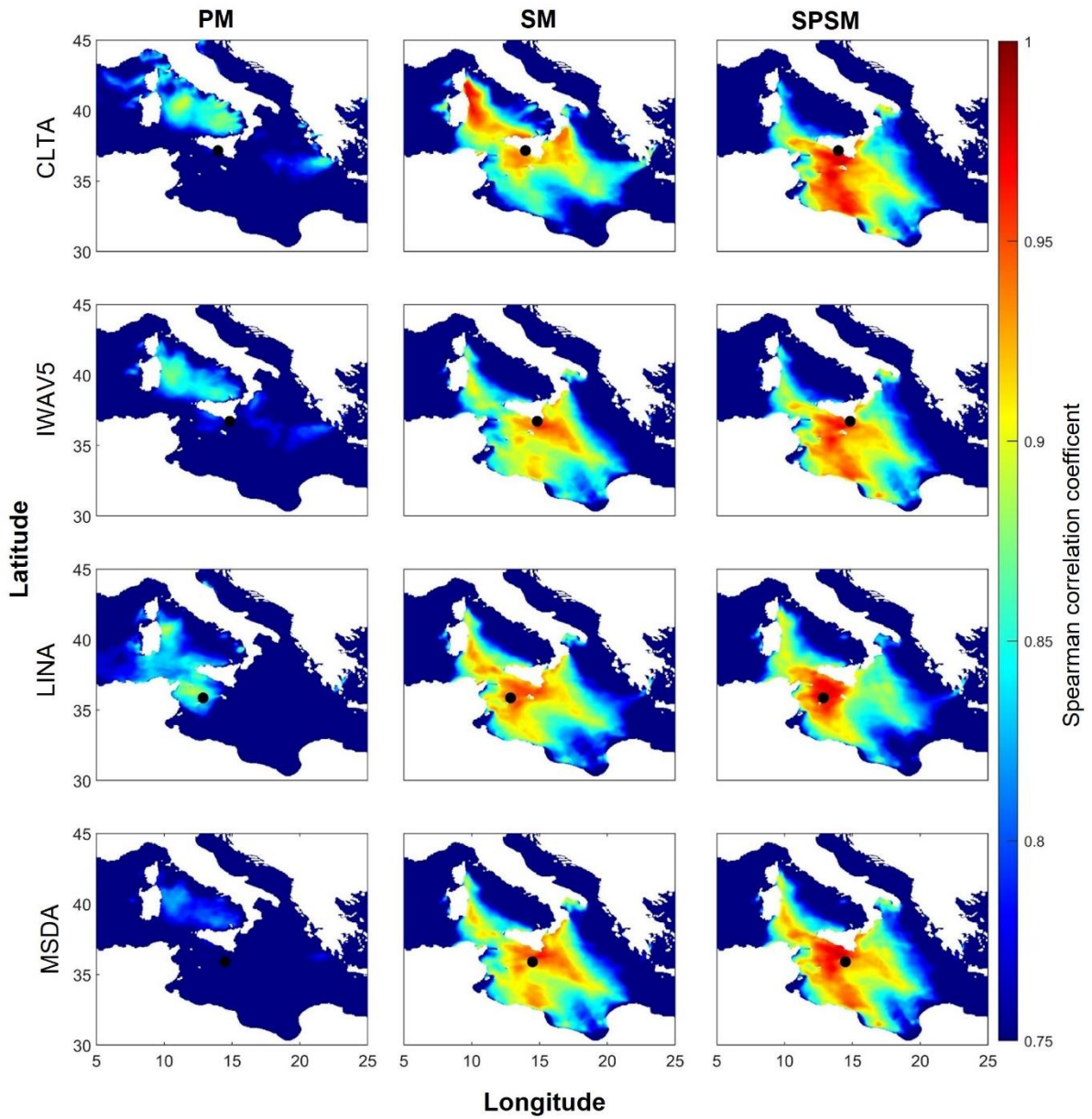
1007 **Figure 4:** Spatial and temporal distribution of the RMS amplitude for the SM band computed  
 1008 at the 105 stations considered (dots). The colors of dots represent the RMS amplitude as  
 1009 specified in the color bar. The blue, green and red contour lines represent significant wave  
 1010 heights of 3, 4 and 5 m, respectively, while the red five-point star in (c) indicates the eye  
 1011 position of the sub-tropical system Helios obtained from satellite images.

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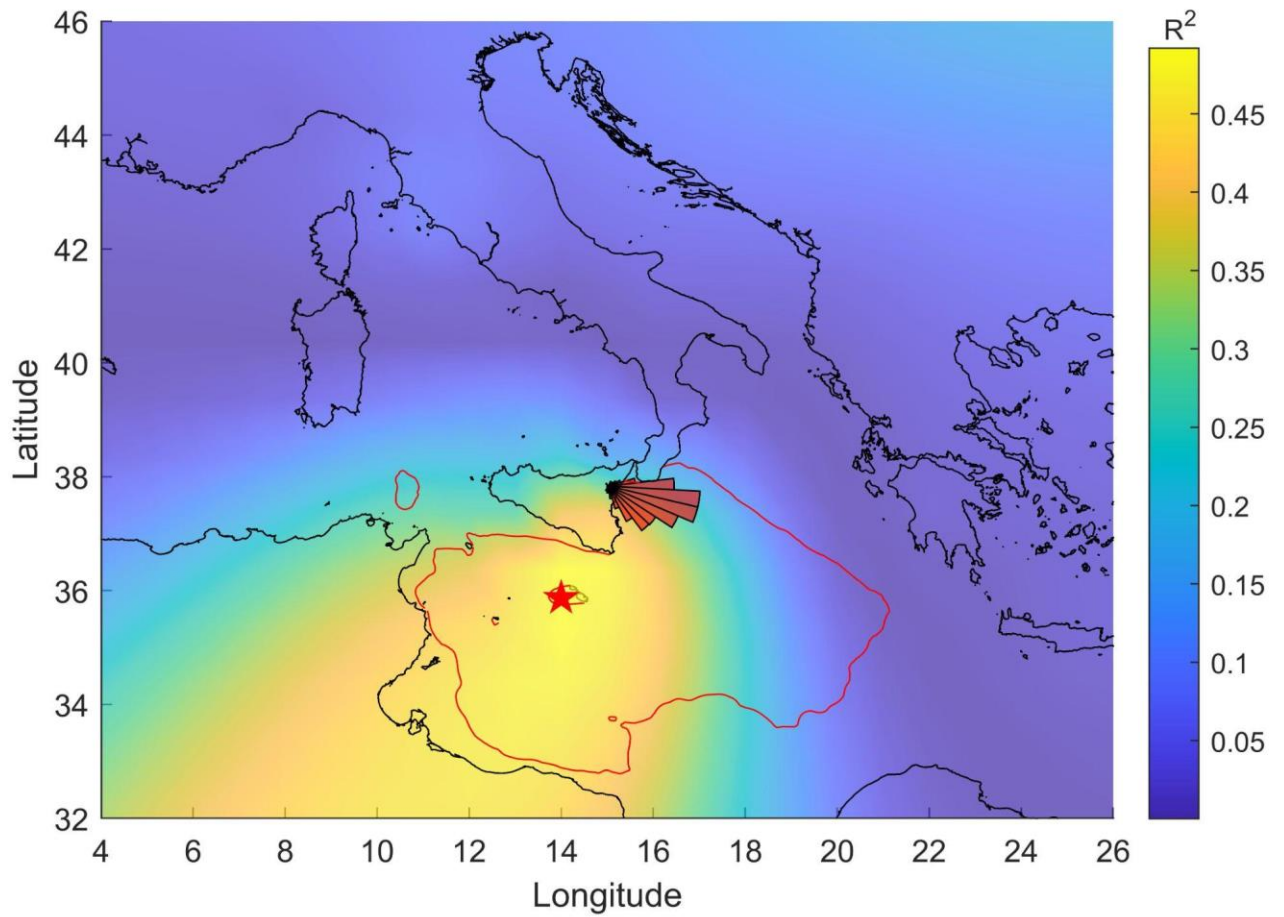
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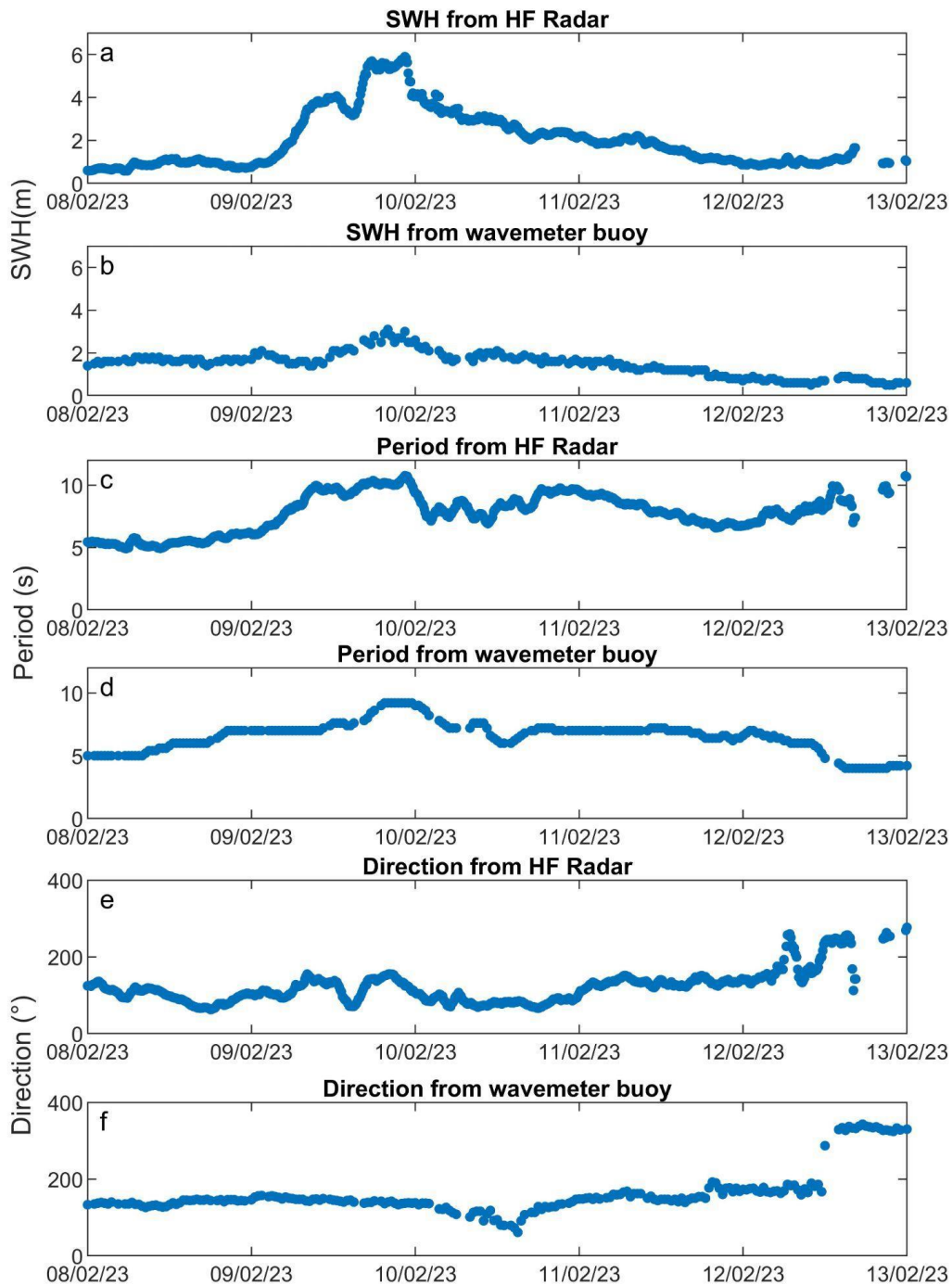
1017 Figure 5: Correlation maps obtained for the vertical component of the seismic stations MSDA,  
 1018 LINA, CLTA and IWAV5 for the PM, SM and SPSM frequency bands during the period under  
 1019 investigation. The black dots indicate the position of the seismic station.

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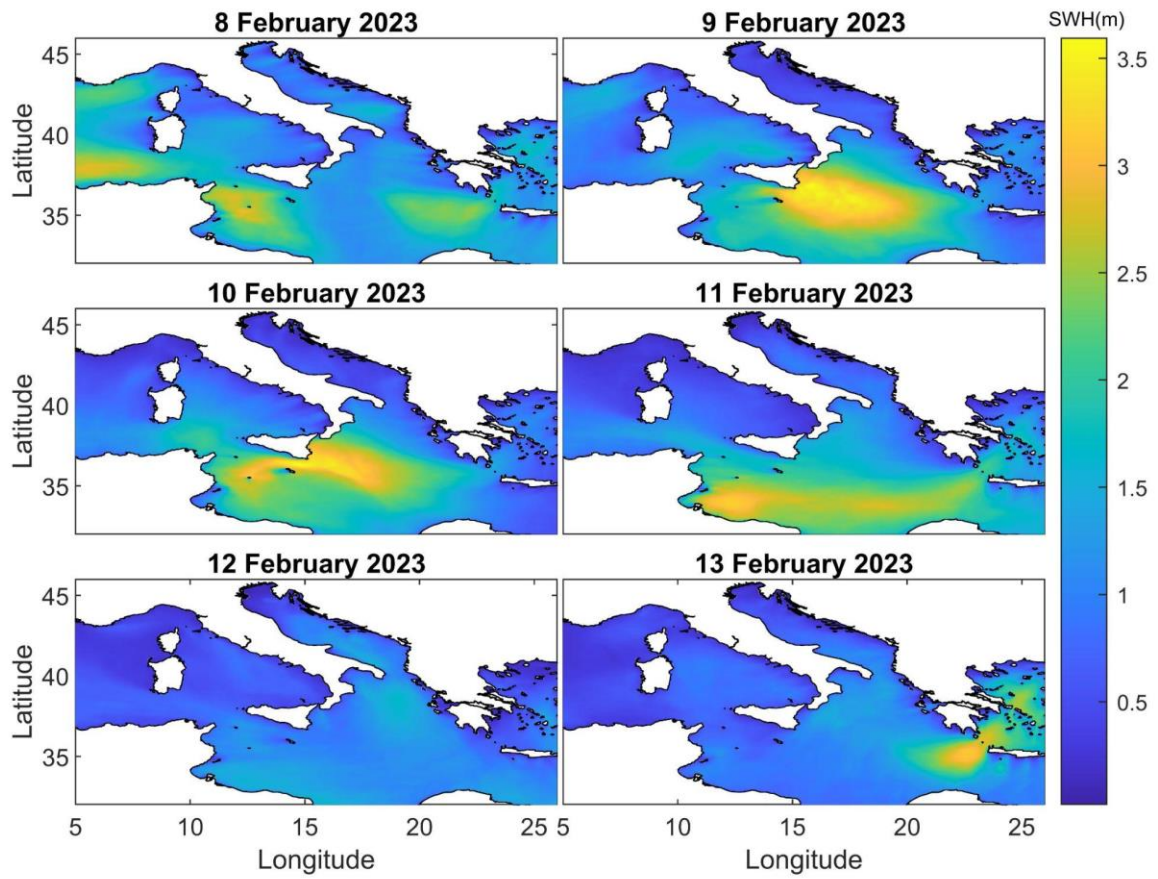
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1022 **Figure 6:** Localization of the microseism source for 10 February 2023 at 16:00. The red five-  
 1023 point star indicates the centroid position of all the grid nodes whose  $R^2$  values do not differ by  
 1024 more than 1% from the maximum  $R^2$  value obtained with the grid search method, while the  
 1025 rose diagram, located at the center of the summit area of Mt. Etna (see **Figure 2b**), shows the  
 1026 distribution of the back azimuth values on the same day. The red contour line represents  
 1027 significant wave heights of 4 m obtained from the Copernicus product  
 1028 MEDSEA\_HINDCAST\_WAV\_006\_012 during the same time interval.



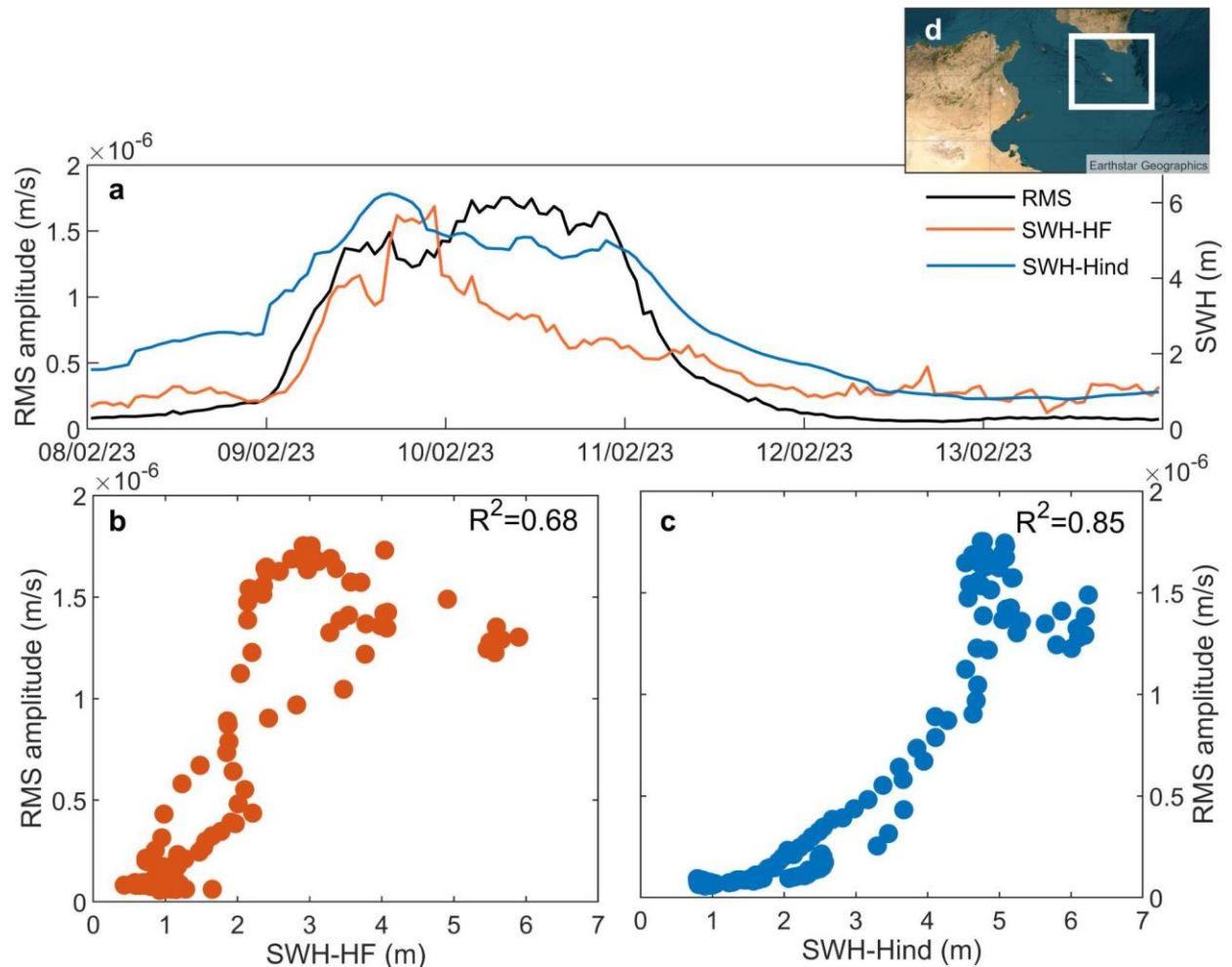
1029  
 1030 **Figure 7:** Wave features in terms of SWH, period and mean direction time series retrieved by  
 1031 using the HF Radar (a, c and e) and Mazara del Vallo buoy (b, d and f) data. For the instruments  
 1032 location see **Figure 2c**.





1033  
 1034 **Figure 8:** Hindcast maps, obtained from the Copernicus product  
 1035 MEDSEA\_HINDCAST\_WAV\_006\_012, showing the spatio-temporal variations of SWH  
 1036 during the days taken into account.





1037

1038 **Figure 9:** (a) RMS amplitude time series, recorded at the station IWAV5, and SWH time series

1039 retrieved by HF Radar (SWH-HF; orange line) and by hindcast data (SWH-Hind; light-blue).

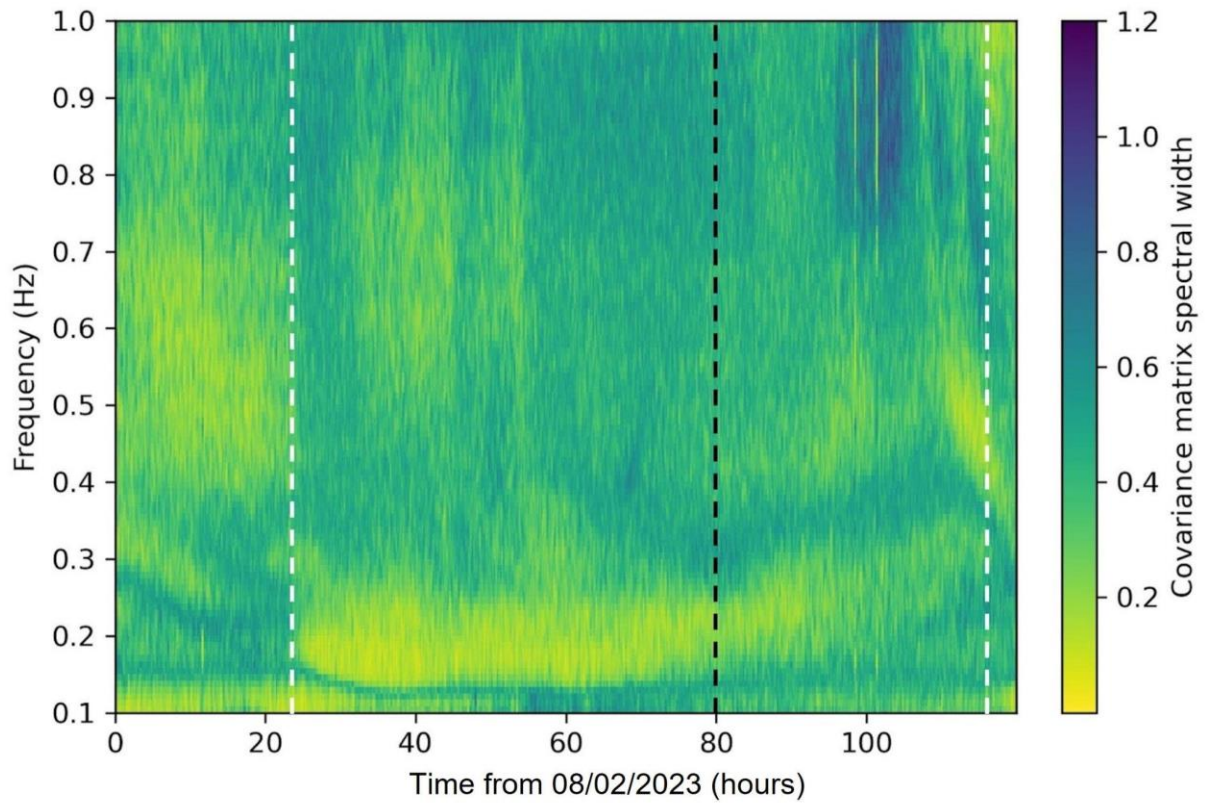
1040 Cross-plot showing the relation between SWH-HF and RMS amplitude (b) and between SWH-

1041 Hind and RMS amplitude (c). The value of the determination coefficient ( $R^2$ ) is reported in the

1042 upper right corner of the plots (b) and (c). In (d) the area of the Sicily Channel used to calculate

1043 the SWH-Hind time series is shown (base image source ©Earthstar Geographic). For the

1044 instruments location see **Figure 2a** and **2c**.



1045

1046 **Figure 10:** Covariance matrix spectral width obtained for the period 8-12 February 2023 using  
 1047 the stations IWAV5, LINA, MSDA and CLTA. The white dashed lines represent the  
 1048 development and the ending of the sub-tropical system Helios, while the black dashed line  
 1049 indicates the time when Helios makes landfall against the Libyan coast. For stations location  
 1050 see **Figure 2a**.

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1055 Table

Date	Hour	Longitude of microseim source (degrees)	Latitude of microseis m source (degrees)	R <sup>2</sup> Value	Errors latitude (km)	Errors longitude (km)	Longitude of cyclone eye from satellite data (degrees)	Latitude of cyclone eye from satellite data (degrees)
09/02/2023	08:00	15.0000	37.3681	0.3233	<b>243</b>	<b>535</b>	/	/
09/02/2023	12:00	15.0000	36.3681	0.3459	<b>217</b>	<b>412</b>	/	/
09/02/2023	16:00	15.0000	36.3681	0.3554	<b>181</b>	<b>364</b>	/	/
09/02/2023	20:00	15.0000	36.3681	0.3865	<b>173</b>	<b>339</b>	/	/
10/02/2023	00:00	14.5000	36.3681	0.4725	<b>185</b>	<b>298</b>	/	/
10/02/2023	04:00	14.5000	36.3681	0.5066	<b>166</b>	<b>214</b>	14.01584	36.24165
10/02/2023	08:00	14.3030	36.0348	0.5107	<b>178</b>	<b>228</b>	13.79612	35.57149
10/02/2023	12:00	14.3030	36.0348	0.5091	<b>170</b>	<b>211</b>	13.91697	35.13203
10/02/2023	16:00	14.0000	35.8681	0.4920	<b>165</b>	<b>228</b>	13.59836	34.5937
10/02/2023	20:00	14.0000	35.8681	0.4762	<b>165</b>	<b>245</b>	13.52146	33.59395
11/02/2023	00:00	14.0000	35.8681	0.3864	<b>206</b>	<b>321</b>	13.74118	32.9897

1056 **Table 1.** Main features of the microseism sources and comparison with the coordinates of the  
1057 cyclone eye retrieved from satellite data. The coordinates of the cyclone eye between 8:00 of  
1058 9 February and 00:00 of 10 February are absent since the cyclone eye is clearly visible between  
1059 01:00 and 23:00 of 10 February 2023.