

1 **Response to Anonymous Referee #2**

2 We thank the reviewer for his/her valuable and helpful comments on the manuscript. We propose to  
3 implement the following changes in a revised version.

4 Black = reviewer comment / blue = author's response / "*italic*" = revised text.

5 Review of Bouchet et al., 2023 – AICC2023

6 Bouchet et al., 2023 present an update to the AICC ten years after the first AICC. The update is focused  
7 on the EDC ice core and the older portion of the timescale that is based on orbital tuning. The increased  
8 density of measurements of  $\delta^{18}O_{atm}$ , TAC,  $\delta O_2/N_2$  and  $\delta^{15}N$  are welcome and represent a significant  
9 improvement.

10 This manuscript describes a useful update the AICC. I remain confused by the exclusion of all(?)  
11 US, British, New Zealand, and Australian ice cores from the AICC. This is of relatively minor  
12 importance to this manuscript given that the different age ranges and the focus here on ages older than  
13 100 ka and almost exclusively on EDC. I will urge that this chronology is names the EAICC – the East  
14 Antarctic Ice Core Chronology – given that there are more West Antarctic cores excluded from this  
15 chronology than East Antarctic cores that are included.

16 **Author's response:** Many thanks for this comment. We acknowledge that the aim of this study was  
17 perhaps not made very clear. The objective was indeed to focus on the long timescale (before 60 ka BP)  
18 to present the numerous new data available on the EDC ice core and using them to update AICC2012  
19 with a special focus on deep time scales.

20 Adding the ice cores not yet included in the Paleochrono tool and mainly covering the last 60 kyr would  
21 be a different study which requests a lot of resources for the implementation of the ice cores in the  
22 Paleochrono tool. It was thus not possible to include everything in this study and we decided to focus  
23 more on the deep timescales with a particular focus on EDC. This will be better explained in the new  
24 manuscript.

25 As for the name of the chronology, after discussion with co-authors, we acknowledge that the name  
26 EAICC could have been a more suitable choice in the first place.

27 However, we would prefer to keep the name AICC for several reasons. It is less confusing and allows  
28 to show that AICC2023 is an update with respect to AICC2012 and that it should replace it. AICC  
29 provides an age model mainly for multiple glacial cycles where only the East Antarctic cores provide  
30 information. AICC uses age constraints from a set of cores (including NGRIP, so not just East Antarctic)  
31 and can be used as a template for West Antarctic cores as well (as for the Skytrain Ice Rise, Mulvaney  
32 et al. 2023). Finally, as mentioned above, one important future development would indeed be to include  
33 the high-resolution information from WAIS Divide and other cores.

34 The authors describe a large range of atmospheric gas measurements. The improvement in resolution of  
35 the many records is impressive. The orbital tuning of these records remains quite challenging and thus  
36 requires a myriad of subjective choices to develop both the timescale and the uncertainty. The orbital  
37 tuning, and the tuning to speleothem calcite, suffer from a lack of understanding in either cause of the  
38 variations in the measured parameter, the orbital parameter to tune to, or both. In particular, both  $O_2/N_2$   
39 and TAC have no process-driven explanation for why they vary based on the orbit characteristics and  
40 the variations are not produced by firm models. While this highlights the need for better understanding,  
41 particularly as great effort is going to extracting multiple >1Ma ice cores that reach the 40 ka world, it  
42 should not prevent doing the best that can be done with current understanding. And Bouchet et al. do  
43 this. They have produced a thoughtful chronology and while the manuscript is dense, it is also clearly  
44 written.

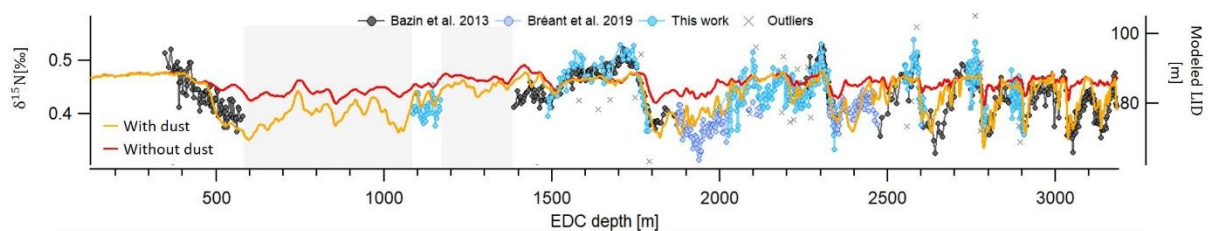
45 There are a couple of areas that stand out as areas of concern:

46 1) The firm modeling

47 This sentence is particularly confusing: “To obtain a coherent scenario, the firm modeling estimates have  
48 been adjusted, by standard normalization, to the scale of LID values derived from  $\delta^{15}\text{N}$  data (later  
49 referred to as experimental LID).”

50 This seems to be hiding a major limitation in the methodology. If I understand correctly, the authors cannot  
51 get the firm model to match the  $\delta^{15}\text{N}$ -inferred firm thicknesses, so they just give up on the actual values  
52 and instead seek to match the variations. Whether this is due to an inappropriate firm model (Breant) or  
53 outdated forcing (the forcing isn't shown but I suspect the authors are using the classical isotope-  
54 temperature scaling that Buizert et al. 2021 showed to be too cold at the LGM). The firm modeling should  
55 really be done with multiple models – which is actually relatively easy to do thanks to the Community  
56 Firm Model – and with a range of climate forcings. I think the authors' efforts would be better served  
57 employing other firm models and forcings rather than the impurity scenarios which the author reject.

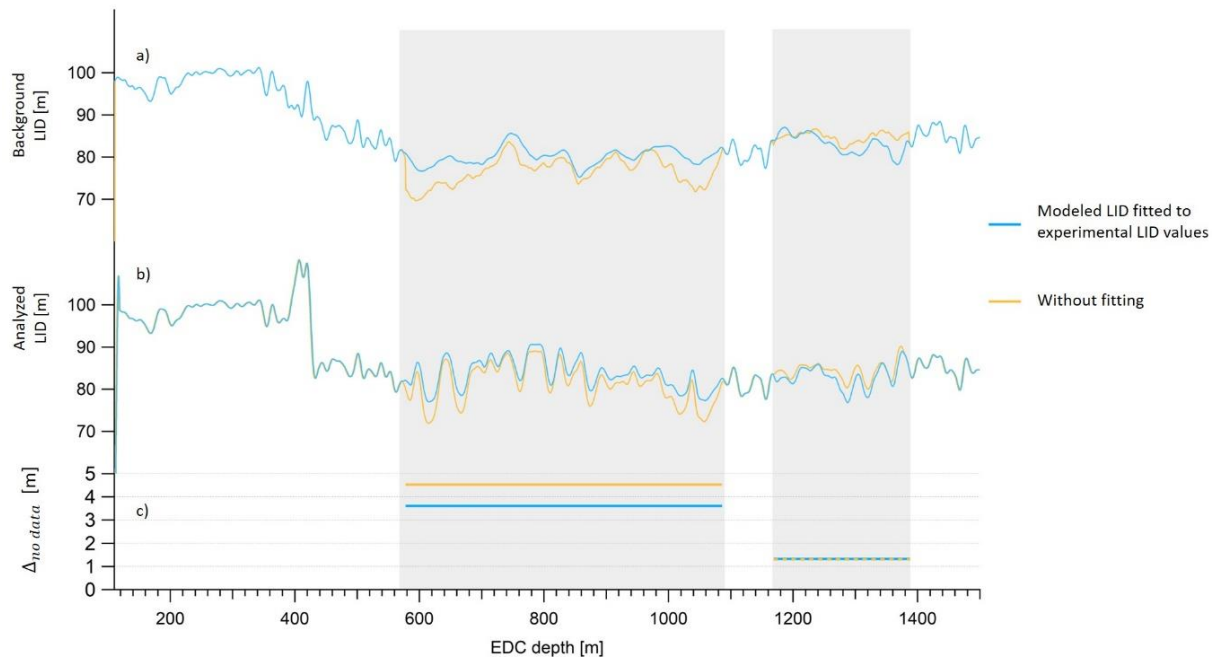
58 **Author's response:** Thank you for raising this contradiction. The idea behind this proposition of fitting  
59 the modeled LID (orange curve on Fig. 1) to experimental LID values was to avoid any discontinuity  
60 when switching from experimental to modeled values when no data are available (grey rectangles on  
61 Fig. 1).



62 **Figure 1.** Modeled LID and  $\delta^{15}\text{N}$  data over the 0-3200 m depth interval. Grey rectangles indicate depth intervals  
63 where  $\delta^{15}\text{N}$  data are not available (either between 578 and 1086 m or between 1169 and 1386 m).

64 However, adjusting the modeled LID to experimental LID values induces a modification of 4.7 m at  
65 most which remains within the background relative uncertainty (20%) so that the adjustment is small  
66 and probably not really needed. This was already shown by the good comparison between modeled and  
67  $\delta^{15}\text{N}$ -inferred LID (L.91-140). To check this, we performed several Paleochrono runs to assess the  
68 credibility of the two modeled LID scenarios (with and without adjustment).

69 On the depth interval from 578 to 1086 m, the raw background modeled scenario (orange curve, Fig. 2)  
70 is almost as credible as the one that was adjusted (blue curve, Fig. 2) (i.e., close  $\Delta_{\text{no data}}$  values). On the  
71 second depth interval of interest, from 1169 to 1386 m, both scenarios show equal  $\Delta_{\text{no data}}$  values, hence  
72 equally credible.



73 **Figure 2.** Background and analyzed LID scenarios at EDC. a) Background LID as per AICC2023 (blue) and  
 74 without fitting of the modeled LID to experimental LID values (orange). b) Analyzed LID. c) The averaged value  
 75 of the misfit,  $\Delta_{no\ data}$ , is calculated for the two LID over the two depth intervals where  $\delta^{15}N$  data are not available  
 76 (either between 578 and 1086 m or between 1169 and 1386 m, see intervals shown by grey rectangles).

77 For more coherence, we believe that we should use the raw firm thickness predicted by the firm model,  
 78 rather than fitting it to experimental LID values. This modification will be considered in the revised  
 79 manuscript.

80 Finally, we would like to emphasize the facts that such modifications of the background LID scenario  
 81 (less than 20% of the LID value) do not significantly affect the final age model and that the major  
 82 improvement of the LID background scenario with respect to AICC2012 is the use of new highly-  
 83 resolved  $\delta^{15}N$  data over the 100-800 ka BP period.

84 We also noted the comment on the use of other firm models. Actually, we tested other firm models in a  
 85 first instance (in particular the simple Herron and Langway model used also by Buizert et al., 2021) but  
 86 we chose to keep the firm model outputs giving the best agreement with the  $\delta^{15}N$  data over the last 800  
 87 kyr at EDC to fill the few gaps existing in the data series. The reason why we did not use the Buizert et  
 88 al. (2021) approach is that it would require (i) a new EDC temperature scenario over the last 800 kyr  
 89 while Buizert et al. (2021) only provided the temperature scenario over the last Termination as well as  
 90 (ii) a new adapted temperature scenario for Vostok (which would be confusing for the readers since our  
 91 goal is not to revise the Antarctic temperature reconstructions over the last climatic cycle). Indeed, to  
 92 use the Buizert approach, we would have needed to adjust the temperature scenario so that the Herron  
 93 and Langway model reproduces best the  $\delta^{15}N$  data. We thus do think that testing the Breant model with  
 94 different parameterizations (all of them published) and keeping the outputs resembling the most the  
 95  $\delta^{15}N$ -inferred LID was the simplest approach (and less confusing) to fill the few gaps in our  $\delta^{15}N$ -  
 96 inferred LID.

97 2) I would like to see an analyses of the thinning function. The EDC AICC2012 thinning function does  
 98 not decrease monotonically as expected from ice flow modeling (i.e. the input background scenario). If  
 99 AICC2023 results in a smoother thinning function, this would provide significant support for the  
 100 methodology.

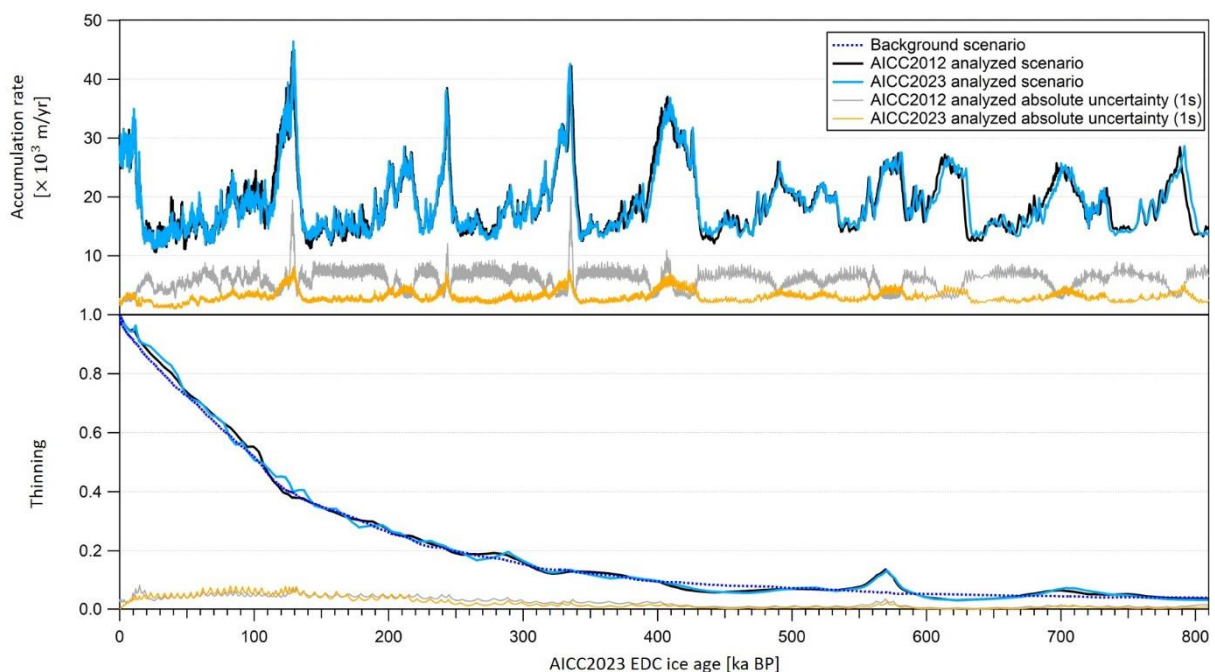
101 **Author's response:** Although the new AICC2023 chronology reduces the absolute uncertainty of the  
102 thinning function compared to AICC2012, it does not provide a smoother and strictly monotonous  
103 scenario (see Figure).

104 However, we believe that this is not a problem for the following reasons:

105 (i) In a tube flow model, like Vostok's (Parrenin et al., 2004), the thinning function is not monotonous  
106 since ice thickness variations are reflected in the thinning function. If the location of the dome at Dome  
107 C shifted over the past 800 kyr, the same effect could have affected the thinning function.

108 (ii) There may also be non-laminar flow effects such as deformation due to more or less hard ice layers.

109 For instance, Dreyfus et al. (2007) described such a particularly complex thinning scenario at Dome C  
110 over the MIS 15 (~580-560 ka BP).



111 **Figure.** Analyzed accumulation and thinning functions of EDC provided by AICC2012 and AICC2023 (black and  
112 blue plain lines respectively) along with their absolute uncertainties (gray and yellow respectively). The  
113 background thinning function is the same for AICC2012 and AICC2023 (dark blue dotted line).

114 General comments on Figure

115 For all figures, the timescale that each parameters is plotted on should be stated explicitly. It gets really  
116 confusing when match points are connected with lines which are not vertical but the two parameters are  
117 plotted on the same age x-axis.

118 **Author's response:** The changes will be made.

119 Vertical lines corresponding the major axes ticks would be really helpful in assessing the alignment of  
120 features

121 **Author's response:** The changes will be made.

122 It would be really helpful to see the uncertainty assigned to each tie point. Presumably this could be  
123 done with a horizontal bar on the match (on the EDC record)

124 **Author's response:** The changes will be made.

125 Specific comments

126 L36 – The introduction could really use subheadings.

127 **Author's response:** We agree and suggest the following subheadings:

128 1.1 Building age scales for deep polar ice cores

129 1.1.1 Motivation

130 1.1.2 Glaciological modeling

131 1.1.3 Chronological constraints derived from measurements

132 1.1.4 Bayesian dating tools

133 1.2 The AICC2012 chronology

134 1.3 The new AICC2023 chronology

135 L43 – “zipped” I don't think this is the right translation to English. I'm not sure what you are going for.

136 I think you are trying to say that a large amount of time is stored in a thin amount of ice.

137 **Author's response:** “zipped” will be changed to “stored”.

138 L44 – need to make community possessive > community's

139 L44 – “core” not “cores”

140 L46 – add “the” before surface

141 **Author's response:** The changes will be made.

142 L53 – what about Nye?

143 **Author's response:** We suggest to change to: “*chronologies of ice cores at low-accumulation sites are commonly established using ice flow and accumulation models (Nye, 1959; Schwander et al., 2001), later on tied up with chronological and glaciological constraints (Bazin and Veres et al., 2013; Parrenin et al., 2017).*”

147 L95 – I think it's worth emphasizing that Bender found no causal link between  $\delta O_2/N_2$  and insolation

148 and was quite forthright about that.

149 **Author's response:** We believe that the quote from Bender (2002): “*We assert that insolation influences snow metamorphism and grain properties in shallow firn. The insolation signature in these properties is retained throughout the firn, and influences  $O_2/N_2$  fractionation during bubble closeoff*” is

151 coherent with what we wrote in the introduction: “*observations led Bender (2002) to assert that local*

152 *summer solstice insolation affects near-surface snow metamorphism and that this imprint is preserved*

153 *as snow densifies in the firn and, later on, affects the ratio  $\delta O_2/N_2$  measured in air bubbles formed at*

154 *the lock-in-zone.*”

155

156 We agree to come back to this point if you believe it still needs modifications.

157 I don't expect that the authors will agree to incorporate WAIS Divide, but the introduction should have

158 a paragraph that acknowledges the exclusion and points readers to the timescales for these cores that are

159 tied to WAIS Divide as the best ones to use for past ~60 ka.

160 **Author's response:** We agree to designate the WD2014 chronology as the best candidate for the past  
161 60 kyr. For greater coherence within the manuscript, we suggest to add a paragraph at the end of sect.  
162 2.1 at L.206: *"We acknowledge the exclusion of the WAIS Divide ice core from the construction of the*  
163 *AICC2023 age scale. Over the last 60 kyr, though, we recommend the use of timescales tied to the WAIS*  
164 *Divide 2014 age model (WD2014, Buizert et al., 2015; Sigl et al., 2016). WD2014 hands over to*  
165 *AICC2023 for the period older than 60 ka BP (that is for the section below the depth of 950 m for the*  
166 *EDC ice core)."*

167 As mentioned above, the depth-depth correspondence between AICC2023 and WD2014 age models  
168 will be given in supplement.

169 L118 – "peculiar" I think you mean "particular"

170 **Author's response:** We suggest to change *"peculiar"* to *"singular"* as *"particular"* is not exactly what  
171 we meant.

172 L308 – shouldn't you reference Tison et al. 2015 here?

173 **Author's response:** The reference will be added.

174 More general comments

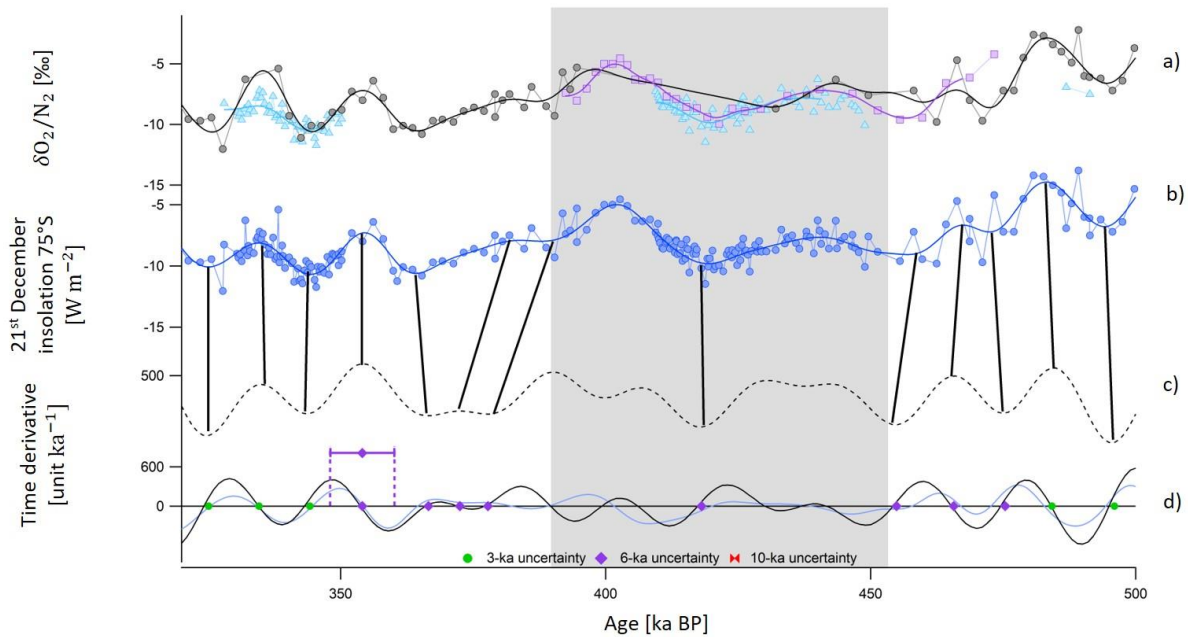
175 L349 – "superior" in English implies "better". I think "greater than" is better phrasing

176 **Author's response:** The change will be made.

177 L372 – the discarding of "tie points" worries me. Doesn't this imply that you don't understand the  
178 underlying mechanisms that link the measurements parameter on the target tuning parameter? If you are  
179 discarding tie points all together, should the uncertainty for the tie points you keep be increased to  
180 respect that the relationship the ties are based on are not stationary?

181 **Author's response:** I think the word "discarding" was poorly chosen. Over the period of MIS 11 (gray  
182 frame in the Figure), it is impossible to match  $\delta O_2/N_2$  and insolation variations as they do not resemble  
183 each other. For instance, two peaks in the insolation curve (dashed black line) only correspond to one  
184 peak in the  $\delta O_2/N_2$  data (blue circles). Hence, there is no tie point in the first place to be discarded. We  
185 suggest the following modifications at lines 370-372: *"In such cases, the uncertainty associated with*  
186 *each tie point is ranging from 6 to 10 kyr (precession half period) and some extrema in the target are*  
187 *not used to tune the record (5 extrema over MIS 11 out of 63 over the last 800 kyr)."*

188 We argue that the uncertainty of 3 kyr for the  $\delta O_2/N_2$  tie points is enough. It was evaluated by Bazin et  
189 al. (2016) who examined three  $\delta O_2/N_2$  records from Vostok, Dome Fuji and EDC ice cores over MIS  
190 5 and detected some site-specific  $\delta O_2/N_2$  variations. This observation, along with the presence of a 100  
191 kyr periodicity in the  $\delta O_2/N_2$  record and the difficulty of identifying  $\delta O_2/N_2$  mid-slopes and maxima,  
192 led them to recommend the use of a 3 kyr uncertainty.



193 **Figure.** Alignment of  $\delta\text{O}_2/\text{N}_2$  and insolation between 500 and 300 ka BP. (a) EDC raw  $\delta\text{O}_2/\text{N}_2$  old data (black  
 194 circles for data of Extier et al. (2018) and purple squares for data of Landais et al. (2012)), outliers (grey crosses)  
 195 and filtered signal (black and purple lines). EDC raw  $\delta\text{O}_2/\text{N}_2$  new data (blue triangles, this study) and filtered  
 196 signals (blue line). The  $\delta\text{O}_2/\text{N}_2$  data are plotted on AICC2012 ice timescale. (b) Extrema in the compiled filtered  
 197  $\delta\text{O}_2/\text{N}_2$  dataset (blue plain line) are identified and matched to extrema in the (c) 21<sup>st</sup> December insolation at 75°  
 198 South plotted on a reversed y-axis and on the age scale given by Laskar et al. (2004) (dash line). The matching  
 199 peaks are linked by black vertical bars. (d) The 0 value in the time derivative of insolation (black line) and of the  
 200 filtered  $\delta\text{O}_2/\text{N}_2$  dataset (blue line) corresponds to extreme values in the signals. The determined tie points between  
 201  $\delta\text{O}_2/\text{N}_2$  and insolation are depicted by markers on the horizontal line. Green circles are attached to a 3 kyr  
 202 uncertainty and purple squares are associated with a 6 kyr uncertainty (purple horizontal error-bar represented at  
 203 354.1 ka BP). Between 390 and 475 ka BP, all extrema are not tuned to the target due to the poor resemblance  
 204 between the signal and insolation (see gray frame).

205 L396 – The authors should not use “continuous” to describe the discrete gas measurements. These  
 206 samples are still quite sparse. Instead, the authors should emphasize increase in sample resolution and  
 207 the reduction in the largest gaps.

208 **Author’s response:** We agree and several changes will be made to remove the adjectives “continuous”  
 209 or “discontinuous” when designating the gas records.

210 Figure 5 – I find the match points between d18O-O2 and speleothem d18O to be unconvincing. What  
 211 features are being matched and what features aren’t seems arbitrary. Maybe this would be improved by  
 212 showing the uncertainty

213 **Author’s response:** This point was also raised by the Referee 1 and we agree to modify the Fig. 5 so  
 214 that the uncertainty is shown.

215 L500 – I’m concerned the 6ka uncertainty is way to small. 6ka seems reasonable for the actual matches,  
 216 but shifting the tie points based by 5ka based on whether there is a Heinrich-like event is not well  
 217 founded. This really needs process modeling for support. Since that is outside the scope of the study, I  
 218 recommend increasing the uncertainty at least 10 ka (5ka since you don’t know what to tune to and 5ka  
 219 for the murky matches themselves).

220 **Author’s response:** Please don’t forget that 6 ka is a 1 sigma uncertainty, so the 2 sigmas uncertainty  
 221 is 12 kyr, which seems enough for us.

222 Jouzel et al. (2002) presented the drawbacks of assuming a constant phase between  $\delta^{18}\text{O}_{\text{atm}}$  and insolation  
223 which is a key assumption of the orbital tuning approach. To evaluate the uncertainty of the phasing  
224 between  $\delta^{18}\text{O}_{\text{atm}}$  and insolation, Parrenin et al. (2001) assumed that the number of precessional cycles  
225 can be counted in the  $\delta^{18}\text{O}_{\text{atm}}$  record. For them, this assumption “*is straightforward considering how*  
226 *clearly this cycle is imprinted in the  $\delta^{18}\text{O}_{\text{atm}}$  series*” and implies that “*ice and gas chronologies are*  
227 *assigned to pass through a succession of large doors with a width of 6 kyr (1/4 of a precession cycle)*”.  
228 The authors estimated this width by combining glaciological modeling and orbital tuning.  
229 We chose to stick with the recommendation of Jouzel et al (2002) and to use a 6-kyr uncertainty, which  
230 also allows to remain coherent with other orbital dating studies already conducted (Bazin et al., 2013).  
231 Yet, it is possible to run another dating experiment with the uncertainty increased to 10 kyr if needed.

232 L576 – As mentioned above, I don’t really understand what you are doing to get a coherent scenario.  
233 Are there other firm models which get better agreement? And what are the climate forcings?

234 **Author’s response:** We agree that the standard normalization is not necessary and are willing to let the  
235 rough modeled LID values.

236 L588 – Why are you not using the tie points to WAIS Divide directly? These ties are well established  
237 in Buizert et al. 2018. The WAIS Divide timescale is more accurate than GICC05 as demonstrated by  
238 Svensson et al. 2020 who had to shift the dates of GICC05 more than WDC14 for the bipolar matches.

239 **Author’s response:** Although we agree, we would prefer to remain coherent with the AICC2012 study,  
240 that is to say to update the timescale AICC between 60 and 800 ka BP while keeping GICC05 between  
241 60 and 0 ka BP. However, we understand fully this comment and will implement a correspondence  
242 between AICC2023 and WD2014 age models in the dataset submitted to PANGAEA. In the new version  
243 of the manuscript we will insist that for now, we focus mostly on the 60-800 ka BP age interval and  
244 stipulate that the WD2014 age model is more accurate over the last 60 kyr.

245 Ideally, one possible future development would indeed be to include the high-resolution information  
246 from WAIS Divide (and other cores). To do so, the WAIS Divide ice core should be added to the  
247 Paleochrono experiment along with the ties established by Buizert et al. 2018 and background  
248 glaciological scenarios that need to be determined. This development is beyond the scope of this study.