

Characteristics of fine particle matters at the top of Shanghai Tower

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Abstract. To investigate the physical and chemical processes of fine particle matters at mid-upper planetary boundary layer (PBL), we conducted one-year continuous measurements of fine particle matters (PM), chemical composition of non-refractory submicron aerosol (NR-PM₁) and some gas species (including sulfur dioxide, nitrogen oxides and ozone) at an opening observatory (~600 m) at the top of Shanghai Tower (SHT), which is the Chinese 1st and World's 2nd highest building located in the typical financial central business district of Shanghai, China. This is the first report for the characteristics of fine particles based on continuous and sophisticated online measurements at the mid-upper level of urban PBL. The observed PM_{2.5} and PM₁ mass concentrations at SHT were 25.5±17.7 and 17.3±11.7 $\mu\text{g m}^{-3}$ respectively. Organics, nitrate (NO₃) and sulfate (SO₄) occupied the first three leading contributions to NR-PM₁ at SHT, accounting for 35.8 %, 28.6 % and 20.8 % respectively. The lower PM_{2.5} concentration was observed at SHT by 16.4 % compared with that near surface during the observation period. It was attributed to the decreased nighttime PM_{2.5} concentrations (29.4 % lower than surface) at SHT in all seasons due to the complete isolations from both emissions and gas precursors near surface. However, daytime PM_{2.5} concentrations at SHT were 12.4-35.1 % higher than those near surface from June to October, resulted from unexpected larger PM_{2.5} levels during early to middle afternoon at SHT than surface. We suppose the significant chemical production of secondary aerosols existed in mid-upper PBL because strong solar irradiance, adequate gas precursors (e.g., NO_x) and lower temperature were observed at SHT favorable for both photochemical production and gas-to-particle partitioning. This was further demonstrated by the significant increasing rate of oxygenated organic aerosols and NO₃ observed at SHT during 8:00-12:00 in spring (7.4 % h^{-1} and 12.9 % h^{-1}), autumn (9.3 % h^{-1} and 9.1 % h^{-1}) and summer (13.0 % h^{-1} and 11.4 % h^{-1}), which cannot be fully explained by vertical mixing. It was noting that extremely high NO₃ was observed at SHT both in daytime and nighttime in winter, accounting for 37.2 % in NR-PM₁, suggesting the efficient pathway from heterogeneous and gas oxidized formation. Therefore, we highlight the priority of NO_x reduction in Shanghai for the further improvement of air quality. This study reported greater daytime PM_{2.5} concentrations at the height of 600 m in urban PBL compared with surface measurement, providing insight into their potential effects on local air quality, radiation forcing, and cloud/fog formations. We propose that the efficient production of secondary aerosol in mid-upper PBL should be cognized and explored more comprehensively by synergetic observations in future.

1 Introduction

The fine particle matters (PM) can absorb and scatter solar radiation, and act as cloud condensation nuclei. Thus, they can impact on Earth's energy budget directly and indirectly (Yu et al., 2006). Therefore, PM observations are important and necessary. Although worldwide surface (SUR) PM observation networks are reinforcing (Zhang and Cao, 2015; Solomon et al., 2014), the characterization of PM vertical distribution remains uncertain. The PM profiles can be acquired through ground-based lidar (Pappalardo et al., 2014), or airborne measurements (Kulmala et al., 2004). For aircraft observation, the advantage is a platform where flexible instruments can be equipped. For example, the aircraft studies gained aerosol size and composition during ACE-Asia field campaign with an Aerodyne aerosol mass spectrometer (AMS) (Bahreini et al., 2003). For lidar observation, the continuous long-term and high spatial resolution data can be achieved at the same time (Liu et al., 2021; Voudouri et al., 2020). Nevertheless, PM concentration retrieval based on lidar extinction coefficient highly depends on aerosol size distribution, aerosol composition, and atmospheric relative humidity assumptions, which are highly uncertain (Tao et al., 2016).

AMS technique is frequently applied in field observation to analyze PM chemical composition (including organics, nitrate, sulfate, ammonium, and chloride) (Frohlich et al., 2015; Zhang et al., 2007). Based on organic aerosol (OA) data of AMS, the source apportionment of OAs can be performed through positive matrix factorization (PMF) (Zhang et al., 2011). The common outcomes of OA PMF source apportionment are hydrocarbon-like and oxygenated OA (HOA and OOA, respectively). As part of ACE-Asia field campaign, the AMS was deployed in Asia for the first time (Zhou et al., 2020). As mentioned before, vertical PM composition observation needs a platform, which aircraft, mountain, tower, or high-altitude building can serve as. The observations conducted at mountain, tower and high-altitude building make up the "blind zone" of aircraft and lidar observations.

For aircraft observation, the height can reach free troposphere. Previous aircraft studies discovered OA formation in nighttime planetary boundary layer (PBL) (Brown et al., 2013; Pratt et al., 2012) and shallow cumulus clouds (Wonaschuetz et al., 2012). Besides, the distinct vertical distributions of PM chemical species were revealed (Brooks et al., 2019; Aldhaif et al., 2018; Liu et al., 2019). Zhao et al. (2020b) studied the vertical dispersion of size-resolved carbonaceous aerosols by comparing data at near surface level and hilltop. Based on volatile organic compound (VOC) measurements at the heights of 118m and 488m at Canton Tower, Mo et al. (2020) estimated the emission flux of VOC and secondary OA (SOA) formation potential using a mixed layer gradient technique. The measurements on a 300m research tower in a suburban area near Denver showed that the sampling site was under the influence of aged air masses at heights between 40m and 120m, while the fresh emissions below 40m (Ozturk et al., 2013). Based on a 325m meteorological tower in Beijing, previous researchers studied the vertical distribution of chemical species through a series of field campaigns (Chen et al., 2015; Zhao et al., 2020c). Zhou et al. (2018b) found that the differences between PM chemical species at SUR and 260m originated from the

different impacts of regional transport and local emission on primary and secondary species. Xie et al. (2019) presented the
65 contribution of brown carbon to aerosol absorption.

Although previous studies made significant contribution to understanding PM vertical characteristics, long-term observations of PM and their chemical composition in the middle and upper boundary in high density residential area were in lack. Shanghai is one of the most densely populated megacities in the world. In this study, we present one-year continuous
70 observation of $PM_{2.5}$ and PM_1 mass concentrations at the top of 632 m high Shanghai Tower (SHT) in Shanghai, together with the observation of PM_1 chemical composition. In section 2, we describe the measurement sites, instruments and analysis methods. In section 3, we discuss the general characteristics, seasonal variations and diurnal cycles of both SHT and SUR PM. Then, a conclusion is presented in section 4.

2 Experimental

75 2.1 Measurement site

Shanghai seats in the east of Yangtze River Delta region of China and is under the influence of northern subtropical monsoon. As mentioned before, the measurement site is located on a platform (~600 m) at the top of SHT (121.501°E, 31.236°N) in Lujiazui Finance and Trade Zone, a typical central business district with local emissions mainly from dense transportation. As the world's second highest construction that has been finished in the world, SHT stands out in the skyline
80 (Figure 1). To compare the PM characteristics between SUR and SHT, SUR PM data and meteorological data were collected at Pudong Environmental Monitoring Center (PEMC) site and Pudong Meteorological Bureau (PMET) site respectively. PEMC site (121.534°E, 31.229°N, about 3.2 km east of SHT) deployed by Shanghai Environmental Monitoring Center belongs to the national air quality monitoring network, providing hourly concentrations of $PM_{2.5}$, PM_{10} , sulphur dioxide (SO₂), carbon monoxide (CO), ozone (O₃), and nitrogen dioxide (NO₂) for this study. Hourly meteorological measurements
85 including 2 m air temperature, relative humidity (RH), 10 m horizontal wind speed and direction were obtained at PMET site (121.548°E, 31.222°N, about 4.7 km east of SHT) (Pan et al., 2019), which is a standard meteorological observatory managed by Shanghai Meteorological Bureau. Both PEMC site and PMET site are referred to as SUR site in the following discussion. All data are presented in Beijing Standard Time (BJT), which is 8 h ahead of Universal Time Coordinated (UTC).

2.2 Instrumentation

90 The Aerodyne quadrupole-type Aerosol Chemical Speciation Monitor (Q-ACSM) was equipped at SHT to analyze non-refractory PM_1 (NR- PM_1) (Canagaratna et al., 2007) chemical components, including sulfate (SO₄), nitrate (NO₃), ammonium (NH₄), chloride (Chl) and organics (Org) (Ng et al., 2011b), with a time resolution of ~15 minutes. The Q-ACSM was deployed for one year from April 17, 2019 to April 16, 2020. The particles greater than 2.5 μ m were removed

through a PM_{2.5} cyclone (Model URG-2000-30ED) in front of the sampling line. The particles were then dried with a naftion 95 dryer (Perma Pure, Model MD-700-36S-1) before passing through the ACSM inlet. Moreover, PM₁ and PM_{2.5} mass concentrations were collected simultaneously at a time resolution of 5 minutes by using the Thermo Scientific Model 5030 SHARP monitor. The nitrogen oxide (NO-NO₂-NO_x) and SO₂ data were collected by the Thermo Scientific Model 42i and 43i, respectively. Both the gas and aerosol analyzers are deployed in a cabin equipped with air condition (Figure 1), providing continuous measurements with high reliability for this study.

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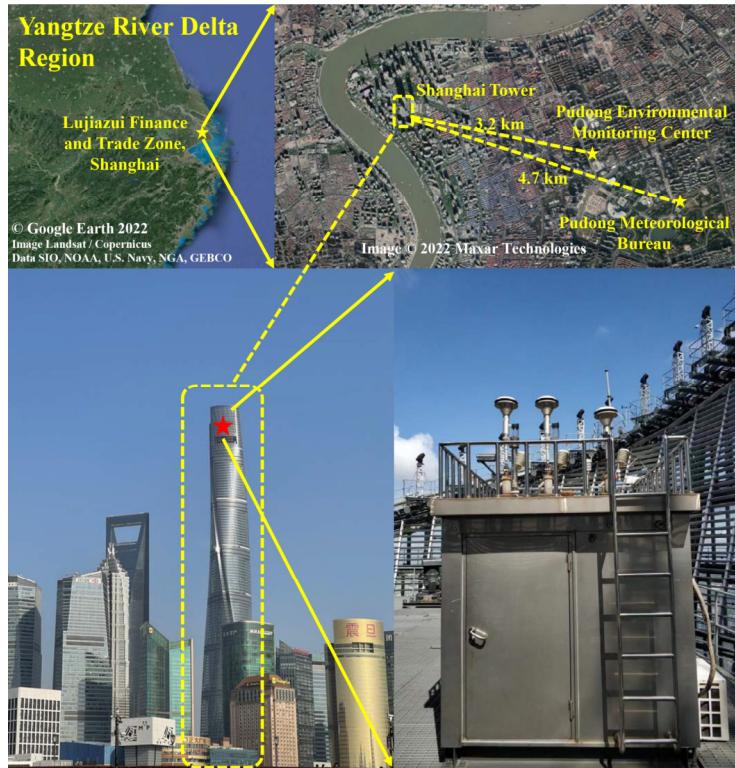


Figure 1: The deployment of SHT site. The red star denotes the platform at the top of SHT.

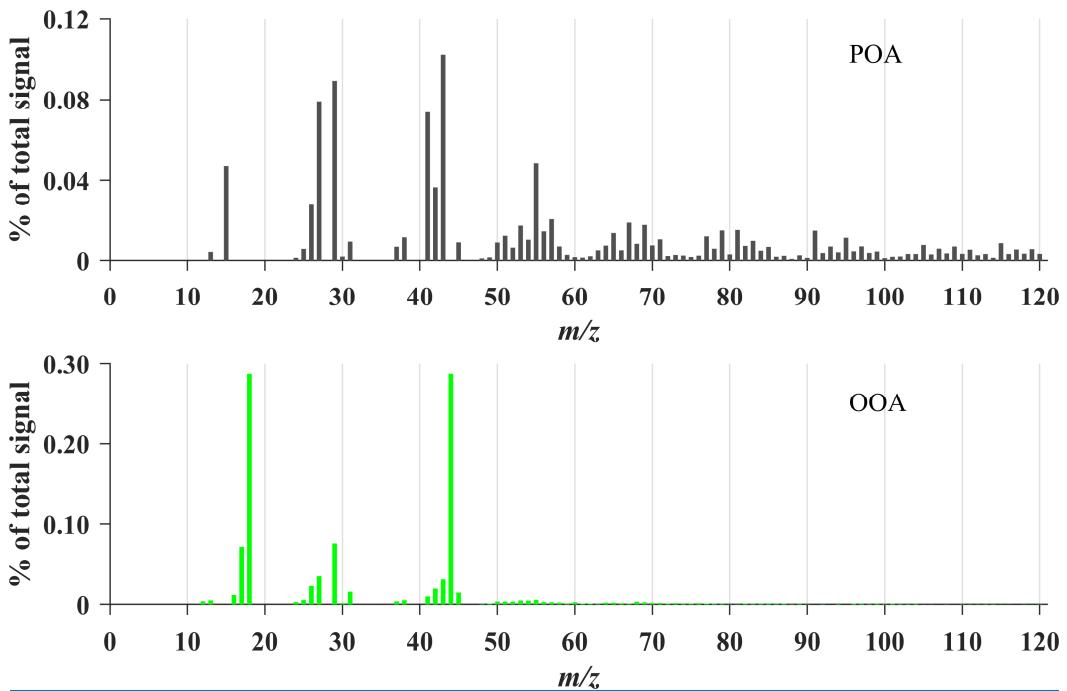
2.3 ACSM data analysis

The ACSM chemical species concentrations were determined from the ion signals measured by aerosol mass spectrometer, 105 using the ACSM Local software (version 1.6.1.0, released on October, 2017) within Igor Pro (Wave Metrics, Inc., USA). The relative ionization efficiency (RIE) values were determined as 7.99 and 0.81 for NH₄ and SO₄ through ionization efficiency (IE) calibrations following the procedures brought by Ng et al. (2011b). RIE values were set as defaults for NO₃ (1.1), Org (1.4), and Chl (1.3). The value of collection efficiency (CE) was taken as 0.5 based on previous field studies in China (Zhao et al., 2020a; Chen et al., 2015; Huang et al., 2012). The value of 0.5 is reasonable as aerosol particles were 110 dried, and the mass fraction of ammonium nitrate (29%) was below the threshold value (40 %) that affects CE (Middlebrook

et al., 2012). Besides, the ratio of measured NH₄ and predicted NH₄ was 0.78, indicating that the particle acidity was weak, and had little effects on CE. The composition dependent CE was investigated according to the algorithm brought up by Middlebrook et al. (2012) and resulted in no significant changes.

2.4 Source apportionment

115 The organics data were further examined by source apportionment using PMF (Canonaco et al., 2013) with an Igor-based source finder tool (SoFi version 6.G). Only *m/z* lower than 120 were included in the source apportionment analysis. The unconstrained two-factor situation (Figure 2S4) was chosen for following discussions. One factor was recognized as a mixture of primary OA (POA). The corresponding profile had hydrocarbon-like fragments (C_nH_{2n-1} and C_nH_{2n+1}; particularly *m/z* 27, 29, 41, 43, 55, 57, 67, and 71) as in HOA, higher ratio of *m/z* 55 than *m/z* 57 as in COA (cooking OA), and 120 distinctive polycyclic aromatic hydrocarbons (PAHs) fragments as in CCOA (coal combustion OA) (Duan et al., 2019). The mixture of POA factors was also reported in previous ACSM studies (Sun et al., 2018). The other factor had an obvious OOA signature with a profile of significantly higher contribution (28.8%) of *m/z* 44 (CO₂⁺) as compared with previous ambient AMS/ACSM datasets (Ng et al., 2011a), meaning that OOA reaching SHT experienced sufficient chemical aging processes. After increasing the number of unconstrained factors (Figure S2S1), no extra meaningful factors were interpreted. 125 We also tried performing PMF analysis separately for each season (Figure S3-62-5), the POA factors were mixed with OOA feature (prominent *m/z* 44 signal) in 2-factor solutions for all four seasons. We did ME-2 analysis with a priori POA profile from the unconstrained two-factor solution for the entire research period, two OOA factors could be identified as a MO-OOA (more oxidized OOA) and a LO-OOA (less oxidized OOA) (Figure S7-96-8). Then, we compared the mass concentrations of ME-2 OA factors with those of unconstrained factors. The Pearson correlation coefficients (R²) between 130 the two methods were 0.97 and 1.00 for POA and OOA, respectively. However, a portion of 22.3% of unconstrained POA mass further split into OOA in the ME-2 solution. As lack of simultaneous measurements of surface chemical components, the unconstrained 2-factor solution of PMF was adopted in consideration of focusing on PM differences between SHT and SUR.



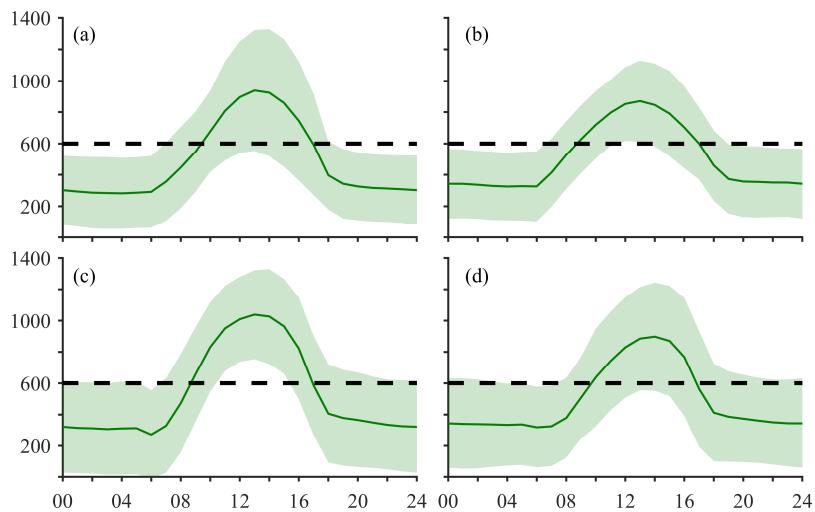
135 **Figure 2: Mass spectra profiles of OA factors for unconstrained two-factor situation.**

2.5 PBL height analysis

To diagnose the relative location of the SHT site to PBL, the seasonal and diurnal changes of PBL height (PBLH) were introduced. We obtained PBLH at SHT from the nearest ERA5 gridded reanalysis data (Hersbach et al., 2020) (<https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-single-levels?tab=form>, last access: 27 November 2022).

140 The ERA-PBLH is calculated utilizing a bulk Richardson method, which was widely used for both convective and stable boundary layers (Kim, 2022). According to Wang et al. (2018), the ERA data tend to overestimate PBLH at nighttime, but underestimate PBLH during daytime in Eastern China by comparing with PBLH calculated from radiosonde sounding data. Overall, the reanalysis data can capture the diurnal and seasonal cycle of PBL structure.

145 As shown in Figure 32, the autumn found the highest PBLH for its prevailing synoptic of the continental high pressure (characterized as weak winds, strong solar radiation, and dry weather), favorable for the PBL development. PBLH in four seasons presented similar diurnal variations. The PBL started to develop at 06:00-08:00 before reaching a daily top at 13:00-14:00, and then decreased until stabilizing after sunset (18:00-19:00). However, the summertime PBL had the longest development period (06:00-19:00), while the wintertime PBL had the shortest (08:00-18:00). At nighttime, the observatory 150 at SHT generally stood on top of stable PBL despite the deviations. Whereas the time PBL top reaching SHT site varied during the day. Nevertheless, the PBL had contact with SHT top even for the lower bound of deviation, indicating inevitable mass exchanges between SHT and SUR during the daytime.



155 **Figure 32: Diurnal variations of the reanalysis PBLH in spring (a), summer (b), autumn (c), and winter (d) at the grid box where the Shanghai Tower (SHT) site is in. The solid line represents the mean value, and the shaded area stands for the standard deviation. The dash lines represent the altitude (~600 m) of the SHT site.**

3 Results and discussions

160 First, it should be noted that we omitted the PM originating from transport outside Shanghai throughout the discussion. However, the seasonal winds induced by Asia monsoon are quite different in upstream (ocean or land, mostly natural or 165 anthropogenic in background) and could impact much at SHT than on the surface. We analyzed the transport pathway at the height of 100 m and 600 m in each season, using 72 h back trajectory from HYbrid Single Particle Lagrangian Integrated Trajectory (HYSPLIT) model. Though the two heights had similar tracks (Figure S9), the small departures might lead to different source origins. This factor should be explored in future studies. Second, PBLH is crucial for the vertical structure 165 analysis, and direct observations of PBLH are in need to give precise view of the boundary layer processes.

3.1 Overview of PM levels and chemical compositions

3.1.1 PM_{2.5} and PM₁ concentrations

170 The averaged PM_{2.5} concentration (Table 1) at SHT during the observation period was $25.5 \pm 17.7 \mu\text{g m}^{-3}$, about 16.4% lower than that ($30.5 \pm 20.7 \mu\text{g m}^{-3}$) at SUR. In which, the PM_{2.5} measurements at SHT were generally consistent with those obtained from sensor-based instruments at the same platform reported by Hao et al. (2022) but only covered the period from June to November of 2019. The averaged PM₁ concentration at SHT was $17.3 \pm 11.7 \mu\text{g m}^{-3}$ (68% of PM_{2.5}), also lower than reported surface PM₁ measurements in Shanghai (e.g., Qiao et al., 2015; Zhou et al., 2018a). The PM₁/PM_{2.5} ratio at SHT

was comparable with those reported by Qiao et al. (2016) and close to 0.69 presented by Zhou et al. (2018a), suggesting the main contribution of PM_{1} to $PM_{2.5}$. R^2 between SHT and SUR was 0.61 for hourly $PM_{2.5}$ concentrations. By contrast, R^2 of 175 $PM_{2.5}$ between PEMC and other surface sites in Pudong district were all higher than 0.89, indicating more significant inhomogeneity of PM distribution in vertical than in horizontal. The relatively lower R^2 in vertical direction suggests distinct origins, transformations and fates of PM at upper PBL which need to be explored.

3.1.2 Chemical compositions

The averaged concentration of NR- PM_{1} measured by ACSM was $16.4 \pm 3.6 \mu g m^{-3}$, which was little lower than the PM_{1} concentration from SHARP 5030, indicating small black carbon existence. Among NR- PM_{1} , the averaged concentrations of 180 chemical species were $3.4 \pm 2.2 \mu g m^{-3}$ (20.9%) for SO_4 , $4.7 \pm 5.3 \mu g m^{-3}$ (28.6%) for NO_3 , $2.1 \pm 1.7 \mu g m^{-3}$ (12.9%) for NH_4 , $1.9 \pm 1.5 \mu g m^{-3}$ (11.4%) for POA, $4.0 \pm 2.8 \mu g m^{-3}$ (24.6%) for OOA, $0.3 \pm 0.2 \mu g m^{-3}$ (1.6%) for Chl. In general, OA, NO_3 and SO_4 were the first three leading contributors to NR- PM_{1} , consistent with the ACSM measurements at 260m Beijing 185 tower (Chen et al., 2015). Similar with previous surface observations in Shanghai (e.g., Zhu et al., 2021; Zhao et al., 2020a), the OA dominated PM_{1} with 35.8% contribution at SHT, in which, POA and OOA comprised 31.5% and 68.5% respectively, and the fractions were very close to the 260m observations (39% and 61%) in Beijing (Chen et al., 2015). The NO_3 at SHT (28.6%) had larger contribution than documented surface measurements (about 1-26%, summarized in Table S1) (Zhao et al., 2020a; Zhu et al., 2021). We found significantly higher NO_3 mass fractions in spring and winter at SHT than previous 190 surface studies despite sampling sites, instruments, and years. In summer, the fraction at a rural site (Zhao et al., 2020) was found lower but close to this study. According to Cui et al. (2022), the proportion of NO_3 was as high as 26.0% in late autumn of 2018. For a similar period (November) in 2019, the ratio at SHT was 27.2%. We also gathered surface water-soluble NO_3 from MARGA (Monitor for AeRosols and Gases in ambient Air) observations at PEMC site during the study period for the exact same period of this study, further supporting the higher portion of NO_3 at SHT. The R^2 between NR- PM_{1} and SHARP PM_{1} was 0.82, indicating consistency between measurements of ACSM and SHARP 5030. Both NR- PM_{1} and 195 SHARP PM_{1} showed a decreasing frequency in mass concentration (Figure S10), while a positive-skewed distribution was found for PM_{1} at SUR in previous study in Shanghai (Zhao et al., 2020a). These results can be attributed to that SHT is far away from emission sources, and influenced by lower PM background concentration than SUR.

3.1.3 Meteorological elements

The observatory at SHT is close to the top of PBL, observed airmass was less affected by direct exchange of heat and 200 moisture from surface. The meteorology at SHT presented lower temperature and less relative humidity (RH) than those at SUR. For example, the mean temperature at SHT was 3-4 °C lower than SUR in different seasons. In terms of RH at SHT, it was nearly consistent with that at SUR in summer and autumn, while about 5-10% lower in spring and winter. Both temperature and RH at SHT showed consistent seasonal variations with those at SUR.

205 The differences between maximum and minimum temperature at SHT were greatest (5.5 °C) in Spring, and smallest (4.1 °C) in Autumn (shown in Table 1). In comparison, the daily ranges of temperature at SUR were greater than SHT in all seasons, with largest range (8.8 °C) in Spring, and smallest (6.9 °C) in Summer. As the daily maximum (minimum) temperature always shows around noontime (midnight), greater temperature differences between SHT and SUR were presumed during daytime than those during nighttime. Accordingly, temperature-sensitive particle formations, for example gas-to-particle
210 partitions were expected to have great differences between two altitudes during the daytime. The daily maximum RH at SHT were lower than SUR in all seasons, yet the daily minimum RH at SHT were higher than SUR. Unlike the temperature, the daily minimum (maximum) RH can always be found during daytime (nighttime). Thus, the higher daytime and lower nighttime RH were expected at SHT than those at SUR, leading to higher daytime and lower nighttime chemical productions from potential heterogeneous reactions at SHT.

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Table 1: The seasonal and annual averaged concentrations of aerosol species (μg m⁻³) and meteorological parameters.

The “dmean”, “dmax” and “dmin” mean the daily average, maximum and minimum.

		Spring	Summer	Autumn	Winter	Annual
Aerosol Species (μg m⁻³)						
SHT	PM ₁	18.6±11.3	16.7±10.8	14.8±8.5	19.4±14.8	17.3±11.7
	PM _{2.5}	25.5±14.2	22.4±13.0	22.3±13.5	31.4±24.7	25.5±17.7
	SO ₄	3.0±1.9	4.2±2.2	3.1±1.8	3.3±2.4	3.4±2.2
	NO ₃	4.8±4.8	3.3±3.2	3.4±2.9	7.2±7.6	4.7±5.3
	NH ₄	2.0±1.5	1.9±1.3	1.9±1.1	2.6±2.3	2.1±1.7
	Chl	0.2±0.2	0.1±0.1	0.3±0.2	0.4±0.3	0.3±0.2
	OA	6.1±3.8	6.6±5.2	5.0±2.9	5.8±4.1	5.9±4.2
	POA	1.9±1.4	2.4±2.1	1.5±1.0	1.7±1.2	1.9±1.5
	OOA	4.1±2.6	4.2±3.2	3.5±2.1	4.2±3.0	4.0±2.8
SUR	PM _{2.5}	29.0±15.8	24.7±12.9	24.3±14.3	43.7±29.1	30.5±20.7
Meteorological parameters						
SHT	T-mean (°C)	13.3±5.3	22.8±3.1	15.9±4.7	5.9±3.7	14.5±7.4
	T-dmax (°C)	16.2±5.7	25.4±3.3	18.0±4.9	8.2±4.2	17.0±7.7
	T-dmin (°C)	10.7±5.3	20.7±3.2	13.9±4.7	3.6±3.3	12.2±7.5
	RH-mean (%)	61.1±21.5	79.6±9.0	74.9±11.6	72.1±15.4	71.9±16.6
	RH-dmax (%)	74.9±19.2	88.9±4.8	84.8±9.0	82.3±11.5	82.7±13.3
	RH-dmin (%)	46.3±23.6	67.4±13.6	61.9±14.4	59.6±19.6	58.8±19.8
SUR	T-mean (°C)	16.2±4.6	26.5±3.1	19.7±4.8	8.6±3.0	17.7±7.6
	T-dmax (°C)	20.7±5.5	30.3±3.6	23.7±5.0	12.3±3.9	21.8±7.9
	T-dmin (°C)	11.9±4.6	23.4±3.2	16.1±5.3	5.2±3.3	14.2±7.8
	RH-mean (%)	71.0±15.1	82.8±8.1	76.7±10.5	77.5±13.4	77.0±12.7
	RH-dmax (%)	94.4±7.3	97.3±4.0	95.1±7.8	94.8±8.9	95.4±7.3
	RH-dmin (%)	46.1±23.1	63.1±13.1	53.3±16.0	55.0±21.5	54.4±19.7

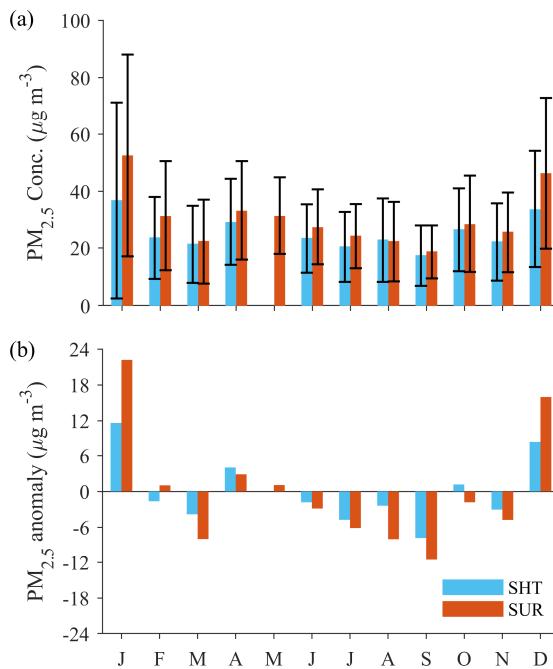
3.2 Seasonal changes

3.2.1 Monthly variations of PM_{2.5} at SHT and SUR

220 As shown in Figure 43a, the monthly variations of PM_{2.5} at SHT and SUR were generally consistent, higher in winter and lower in late summer to early autumn. The highest monthly PM_{2.5} concentrations at SHT (36.9 $\mu\text{g m}^{-3}$) and SUR (52.6 $\mu\text{g m}^{-3}$) both took place in January, while the minimums (17.5 $\mu\text{g m}^{-3}$ for SHT, and 18.8 $\mu\text{g m}^{-3}$ for SUR) appeared in September. The significant monthly change of PM_{2.5} is resulted from the distinct primary emissions (aerosol and its gas precursors), chemical transformation, transports and diffusions as well as wet removals in different seasons. In winter, more pronounced 225 transport and much shallower PBL are conducive to PM_{2.5} accumulation near surface, resulting in higher SUR PM_{2.5} loadings in Shanghai. It is noting that daytime PBL developments usually carry aerosols from surface to high altitude by turbulence, exerting opposite impacts on PM_{2.5} variations near surface and at upper altitudes. As a result, local emissions and regional transports were expected to be responsible for the similar patterns of PM_{2.5} monthly variation at SHT and SUR, rather than PBL changes.

230 **3.2.2 Monthly anomaly of PM_{2.5} at SHT and SUR**

The anomaly was defined as the monthly deviation from annual average. By calculating the anomaly, we intended to find monthly changes relative to the whole year. The comparison of SHT and SUR PM anomalies allows us to see the consistency of monthly features at two altitudes. For both SHT and SUR, negative PM_{2.5} anomalies were found in most months, because of significant positive anomalies in January and December (Figure 43b). The PM_{2.5} concentrations in 235 January and December were 11.6 $\mu\text{g m}^{-3}$ and 22.2 $\mu\text{g m}^{-3}$ higher than the annual averages at SHT and SUR, respectively. Therefore, more stringent attentions should be paid to PM mitigation during these two months. The monthly changes of PM_{2.5} anomalies presented very similar patterns at SHT and SUR, with largest positive anomaly in January, highest negative one in August and September. However, the SHT anomaly oscillated more flatly than SUR, for example, the relative 240 anomaly (PM anomaly divided by yearly-averaged value) during December to January (39.7%) and August to September (-20.2%) was much lower than those at SUR (62.8% and -32.2%). It could be partially explained by weaker influences from surface emissions and air pollutants at higher altitudes, due to nighttime isolations and shallow PBL in winter discussed in following section. It was noting that PM_{2.5} anomalies were opposite in February and October compared with those at SHT and SUR, indicating different leading roles on PM_{2.5} budgets between surface and mid-upper PBL.



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Figure 43: Monthly variations of (a) PM_{2.5} concentrations ($\mu\text{g m}^{-3}$) at SHT and SUR, (b) PM_{2.5} anomalies ($\mu\text{g m}^{-3}$) at SHT and SUR. The monthly-averaged PM_{2.5} in May at SHT is not presented because of low data collection efficiency (36%).

3.2.3 Relative changes between monthly-averaged PM at SHT and SUR

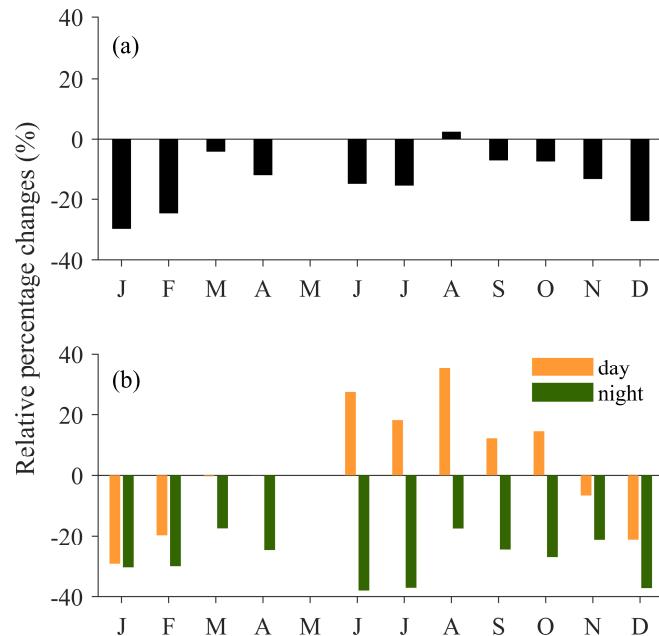
250 Since lower PM_{2.5} concentrations were observed at SHT compared with SUR, the relative percentage changes (RPC) ((PM_{2.5},
 SHT – PM_{2.5}, SUR)/PM_{2.5}, SUR *100%) in Figure 54 were calculated to quantify their discrepancy. The RPC (Figure 54a)

exhibited the lowest value (-27.2%) in winter (December, January and February), and generally consistent values (-8.0%, -9.2%, and -9.1% for spring, summer, and autumn) in the other seasons. ~~The lowest RPC in winter could be attributed to the most shallow PBL height (Pan et al., 2019), where the PM vertical diffusion was significantly inhibited.~~ It was noting that

255 PM_{2.5} concentration observed at SHT was slightly higher (2.4%) than SUR in August. Given that SHT was farther from the direct emission sources than SUR, the PM_{2.5} at SHT tended to have lower concentration than SUR as in the other months despite vertical mixing during the daytime. Thus, the higher PM_{2.5} at SHT in August indicated extra aerosol productions at mid-upper PBL.

260 The exchange of air pollutants between SHT and SUR only exists in daytime due to turbulent mixing. Such mixing process between SHT and SUR would terminate at night due to stable stratification. It could be presumed that nighttime PM_{2.5} observed at SHT (600 m) was independent from that at SUR. Therefore, it is necessary to compare the daytime and nighttime PM_{2.5} separately between SHT and SUR. In Figure 54b, nighttime PM_{2.5} concentrations observed at SHT were consistently lower than SUR, with RPC ranging from -20% to -38%. In addition, nighttime RPC presented very weak

265 monthly variations, suggesting that PM_{2.5} observed at the height of SHT was mostly isolated from both air pollutants and primary emissions near surface at night. In comparison, daytime RPC presented strong monthly variabilities. Different from the negative RPCs of daily PM_{2.5} presented in Figure 54a, daytime PM_{2.5} concentrations at SHT were 20-40% higher than those at SUR from June to October. As is known, turbulence induces the vertical mixing of PM_{2.5} to eliminate the PM_{2.5} gradient within the entire PBL. Since there are no direct primary sources at SHT, the higher PM_{2.5} measurements mean
 270 additional physical or chemical origins existed at this height. Thus, the daytime RPC modulated the seasonal signature of the total RPC, presenting the lowest RPC in winter and implying which suggests the necessity of looking into PM diurnal changes.



275 **Figure 54: Monthly variations of the relative percentage changes (%) between SHT and SUR for (a) all time, (b)**
daytime (08:00-19:00) and nighttime (20:00-07:00). The relative percentage changes are calculated as $(PM_{2.5, SHT} - PM_{2.5, SUR})/PM_{2.5, SUR} * 100\%$.

3.2.4 Monthly variations of chemical species at SHT

As shown in Figure 65, the NR-PM₁ at SHT presented consistent variability and slight departures with SHARP PM₁,
 280 providing well insight to investigate the seasonal contributions of chemical composition. OA, NO₃ and SO₄ occupied the three leading contributions to NR-PM₁ at SHT, accounting for 36%, 28.6% and 20.9% respectively. Their proportions in NR-PM₁ presented distinct seasonal variations. In general, SO₄ and OA occupied higher fractions in summer, and lower proportions in winter. For example, SO₄ had largest portion of 26.1% in NR-PM₁ in summer, while the lowest of 17.1% in winter. Similarly, OA made up the most proportion of 39.5% in summer, and the least of 29.9% in winter. Both OOA and

285 POA showed consistent seasonal contributions to NR-PM₁ with OA. Larger fractions of OOA and SO₄ in NR-PM₁ exhibited
in summer were attributed to relative stronger oxidation capacity and higher moisture conductive to both gas and aqueous
transformations, which were also observed by other mass spectrometer studies near surface (Dai et al., 2019; Hu et al.,
2016). Different from OOA, HOA was mainly emitted by vehicles especially in cold months. Some documents found that
HOA dominated the organics in wintertime Beijing due to enhanced primary emissions from heating season (Duan et al.,
290 2020; Zhang et al., 2013; Zhou et al., 2019). Zhu et al. (2021) also reported a slight decrease of HOA fraction in summer
than other seasons in urban Shanghai. In this study, the POA observed at SHT comprised highest fraction of 14.6% in NR-
PM₁ and 35.8% in organics in summer. In comparison, the ratios were 8.7% and 28.3% in winter, indicating that POA
transport and mixing process were significantly inhabited in winter than in summer. It was expected that NO₃ presented
opposite seasonal pattern to SO₄ and OA. The fraction was found highest (37%) in winter and lowest (21%) in summer,
295 because of the temperature-dependent gas-particle partition. As for NH₄, its seasonal variation was found insignificant.

300 The daytime and nighttime mass fractions were also shown in Figure 65. As results of vertical mixing, the larger portions of
primary species (POA and Chl) during daytime were notable, especially for summer and autumn. The changes of OOA, NO₃,
and NH₄ were slight, with increase of OOA and NH₄, but decrease of NO₃ from nighttime to daytime. Accordingly, SO₄ saw
lower fraction in NR-PM₁ during the daytime. More diurnal features of NR-PM₁ can be found in section 3.3.4.

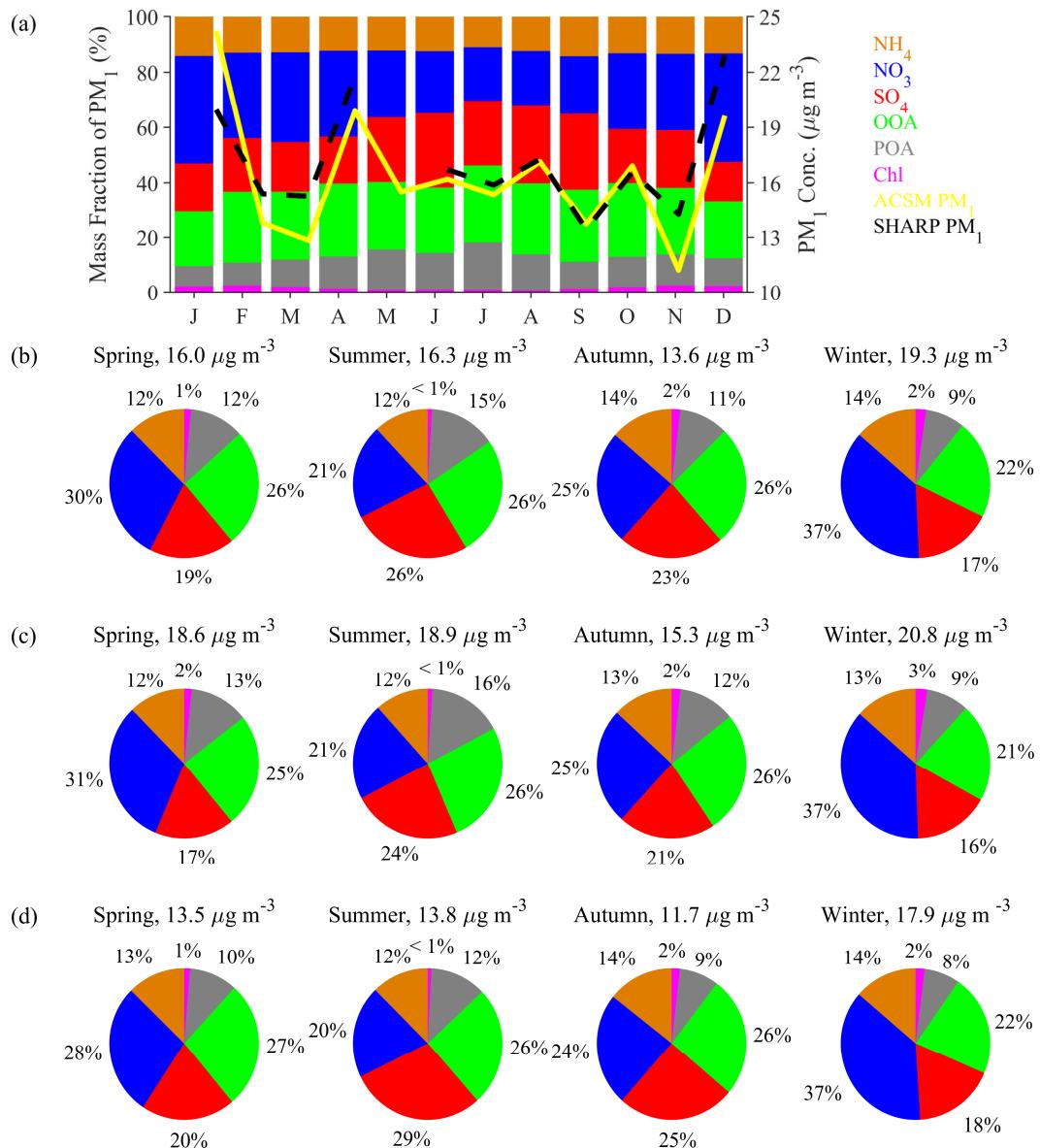


Figure 65: The monthly averaged (a) and seasonal averaged (b-d) mass fractions (%) of NR- PM_1 at SHT. The mass fractions (%) are calculated based on all (b), daytime (c), and nighttime (d) data. The monthly averaged mass concentrations ($\mu\text{g m}^{-3}$) of NR- PM_1 are also shown. The solid and dashed line represent SHARP PM_1 and NR- PM_1 , respectively.

305

3.3 Diurnal changes

3.3.1 Diurnal variations of $\text{PM}_{2.5}$ at SHT and SUR

The air pollutants observed at mid-upper PBL have different origins from those near surface. They are effectively connected

310 in daytime by turbulence, while absolutely isolated at night. As a result, the diurnal variations of air pollutants at bottom and

upper PBL need to be investigated by synergetic observations. In this section, the diurnal characteristics of PM_{2.5} and chemical compositions at SHT are displayed and compared with SUR to discuss the related dynamical and chemical processes at mid-upper PBL.

315 The mean diurnal variations of PM_{2.5} observed at SHT and SUR are displayed in Figure 76, exhibiting different patterns in all seasons. The PM_{2.5} diurnal cycle near surface have been fully documented by previous studies, driven by primary emission, PBL evolution and chemical transformation etc. In general, there existed a clear morning peak of PM_{2.5} concentration around 6:00-7:00 in all seasons, due to substantial air pollutants were released from rush transportations and accumulated in shallow PBL. After sunrise, PBL gradually develops by turbulence which transports PM_{2.5} from surface to
320 high latitude. As a result, PM_{2.5} concentration near surface decreased until early afternoon, then increased again after sunset due to depressed PBL and kept stable at night. However, the PM_{2.5} concentrations observed at SUR presented stable or slight enhancement during the PBL developing period (10:00-15:00) in summer and autumn, differing from the clear PM_{2.5} descent during the same period in spring and winter. It was also reported by Pan et al. (2019), they found the significant PM_{2.5} enhancement around noontime in summer in Shanghai downtown, explained by that rapid production of secondary aerosols
325 related to strong photochemistry in summer offsetting the aerosol loss by vertical mixing.

The PM_{2.5} diurnal variation observed at SHT presented similar unimodal pattern in all seasons. Peak PM_{2.5} concentration was observed around midnoon (12:00) as a result of the rapid increasing of PM_{2.5} concentration since early morning, and notable decline in the afternoon. The early morning PM_{2.5} enhancement at SHT was presumed to be resulted from the PBL
330 development, transporting higher concentration of aerosols from surface to upper layer. Accordingly, PM_{2.5} reduction at SUR was observed simultaneously in Figure 76 due to vertical mixing. It was interesting to note that higher PM_{2.5} at SHT was observed than SUR around noontime (10:00 to 15:00) in spring, summer and autumn, which could not be attributed to vertical mixing. According to the turbulent theory, materials including heat, moisture and air pollutants in PBL are supposed to be mixed homogeneously by vertical mixing. Therefore, when PM_{2.5} concentration at SHT exceed that at SUR before
335 noontime in Figure 76, turbulence would mix the higher loadings of aerosol from upper PBL to surface. It could be also demonstrated by more significant PM_{2.5} anti-correlations between SUR and SHT during morning to early noontime (6:00 to 10:00) when observed PM_{2.5} at SHT was lower than SUR. Therefore, there must exist other processes responsible for the higher PM_{2.5} concentrations appeared around noontime at SHT. This was also observed by Hao et al. (2022) that the daytime PM_{2.5} concentration at 585m was larger than that at 25m by 10-15 $\mu\text{g m}^{-3}$. We supposed significant chemical formation of
340 secondary aerosols as the dominant process for the occurrence of greater PM_{2.5} levels at SHT around noontime, which were discussed in the following section.

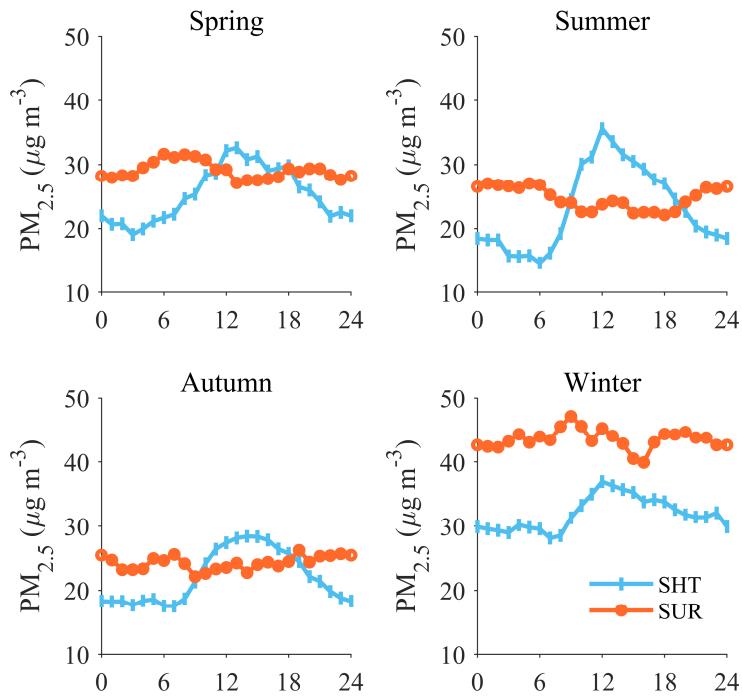
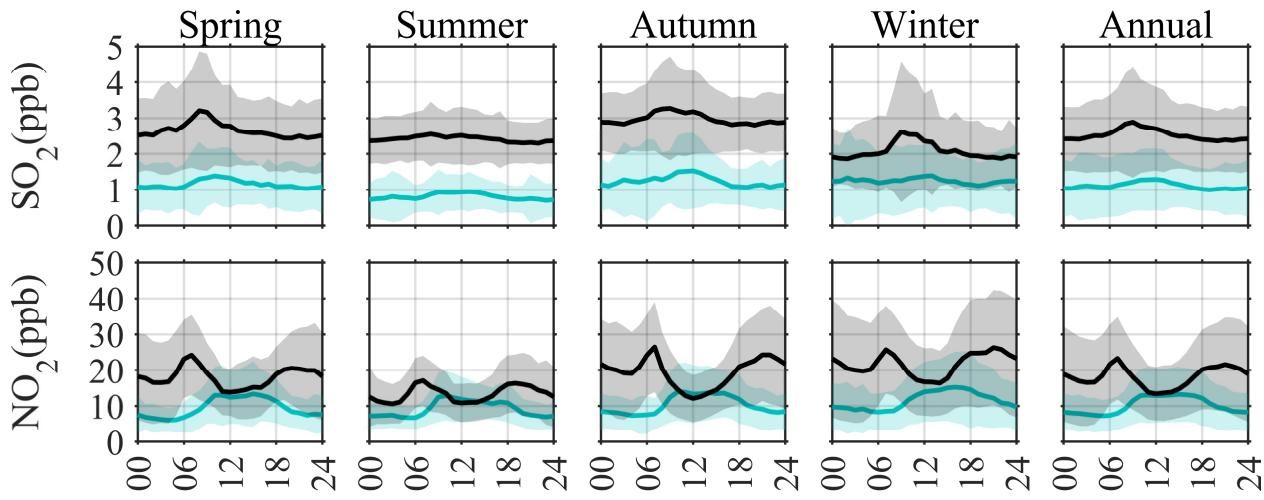


Figure 76: Diurnal variations of PM_{2.5} mass concentrations ($\mu\text{g m}^{-3}$) at SHT and SUR in four seasons.

345 **3.3.2 Discuss on the daytime aerosol production at SHT**

As showed in Figure 76, the peak PM_{2.5} concentration at SHT was highest (~40 $\mu\text{g m}^{-3}$) in summer, it could not origin from air pollutants near surface by vertical mixing because the PM_{2.5} at SUR was relatively lower in summer. In addition, the PM_{2.5} at SHT presented significant increasing rate even after its exceedance to SUR around 10:00, indicating that there must exist remarkable origins of aerosols at SHT, especially in summer. Therefore, chemical transformation from both gas and aqueous pathways were presumed to be the major process for promoting the aerosol productions at SHT, resulting in the significant peak PM_{2.5} concentration.

Chemical production of aerosol in daytime is mainly driven by gas phase and aqueous phase transformations from the gas precursors, including SO₂, NO₂ and VOCs. We further compared the gas species observed at SHT and SUR in Figure 87. It 355 was found that SO₂ concentrations were low (<4 ppb) at both SHT and SUR, and SO₂ concentration at SUR was about 2 times higher than SHT. A slight increase (2.1-19.2%) of SO₂ at SHT during 8:00-12:00 could be found in four seasons. Similar as SO₂, lower NO₂ were observed at SHT than SUR. However, NO₂ was found well mixed around noontime in all seasons. NO₂ at SHT in different seasons rose by 21.8-61.4% from 8:00 to 12:00, when NO₂ were reduced at ground level, indicating the effects of vertical mixing. Therefore, secondary formation of inorganic aerosols could be expected at SHT 360 with adequate gas precursors.



365 **Figure 87:** Diurnal variations of SO_2 (ppb) and NO_2 (ppb) observed at SHT (blue line) and SUR (black line) in four seasons and the entire observation period. The line stands for mean value. The shaded area represents the standard deviation.

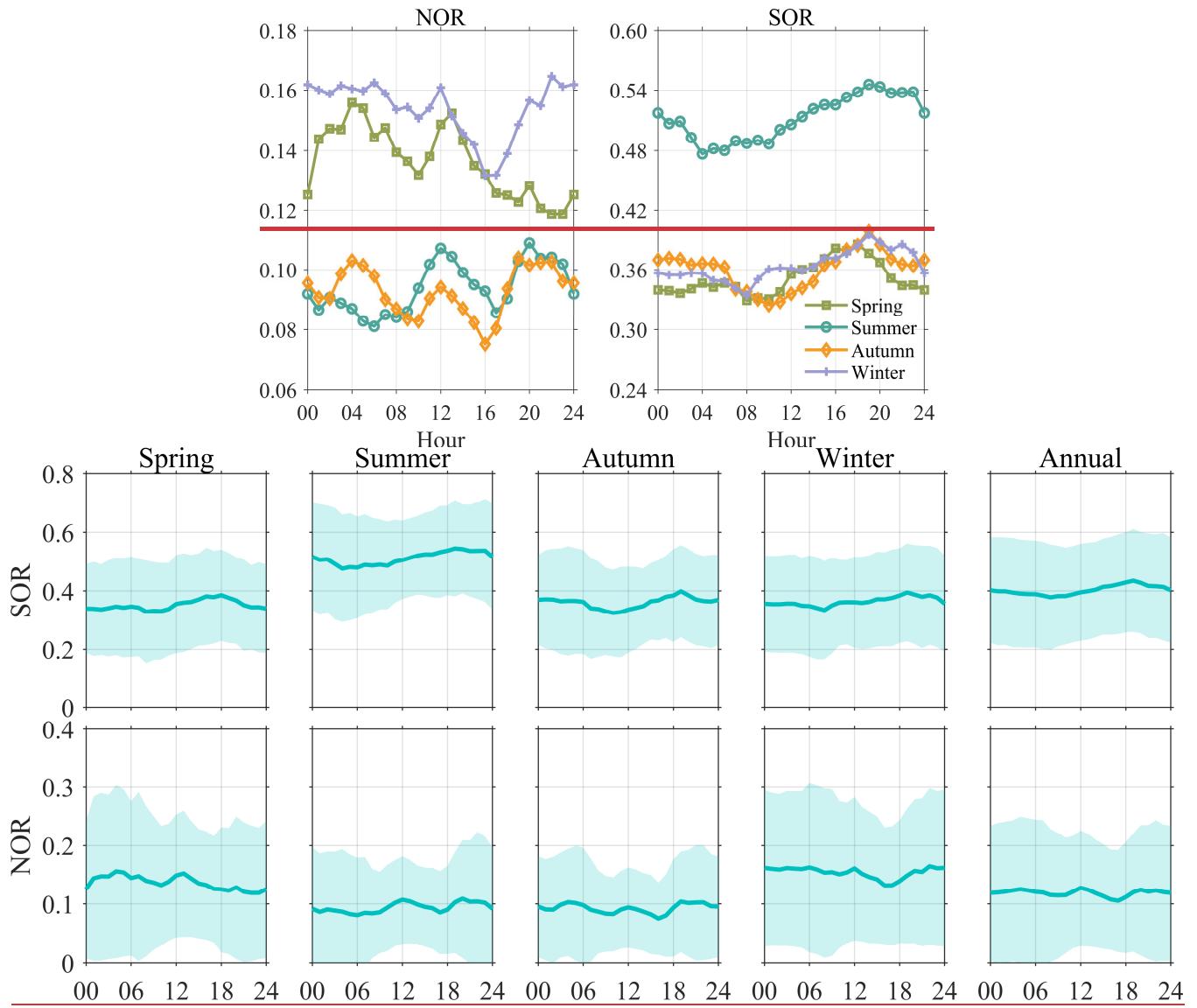
370 Both gas and aqueous formation of secondary aerosols in daytime are greatly determined by atmospheric oxidants, such as OH , H_2O_2 radicals and O_3 . In which OH is produced from the photodissociation of O_3 (clean atmosphere) or $\text{HONO}/\text{H}_2\text{O}_2$ (polluted atmosphere) in troposphere. So that solar radiation is vital for OH production. Apparently, stronger solar radiation

375 could be expected at SHT than that near surface due to shorter optical range of sunlight and hardly attenuation from buildings or ~~vegetables~~ vegetation, providing excellent photolysis capacity for OH production. We further examined the diurnal variations of sulfur oxidation ratio (SOR) and nitrogen oxidation ratio (NOR) at SHT in different seasons in Figure

380 98. SOR and NOR are indicators of the secondary formation of SO_4 and NO_3 (Zhang et al., 2020). The SOR and NOR are calculated as $[\text{SO}_4]/([\text{SO}_4]+[\text{SO}_2])$, and $[\text{NO}_3]/([\text{NO}_3]+[\text{NO}_2])$, respectively. $[\text{x}]$ stands for the molar concentration of x . The

385 SOR in summer (0.51) was significantly higher than those in the other seasons (0.36). The diurnal cycle of SOR was similar in each season, with the highest SOR appeared around 19:00, when RH reached the peak of the day (Figure 109). However, the diurnal variations of NOR were not reproducible between the seasons and not as straightforward to interpret as those of SOR.

390 NOR in spring and winter (0.15) were about 1.6 times larger than that in summer and autumn (0.09). Despite the significant increase of NO_2 during the daytime, the diurnal cycle of NOR kept roughly stable In addition, NOR did not see notable decrease from 8:00 to 12:00, when NO_2 increased significantly, indicating that NO_3 formation at SHT was evident but not necessarily more efficient than SUR.



385 **Figure 8: Diurnal variations of SOR and NOR observed at SHT in all seasons.**

Figure 9: Same as Figure 8 but for SOR (upper panel) and NOR (lower panel) at SHT.

As was discussed in Figure 6S, OOA, NO_3 and SO_4 were the major species in NR- PM_{1} . Their production and variation are closely related to meteorology, especially temperature and RH. For example, lower temperature was favorable for 390 productions of NO_3 and some semi-volatile species in OA. Whereas higher RH played the important roles in the aqueous pathway of SO_4 formation. Recent studies also reported significant production of low-volatile OA by aqueous phase

reactions (Chen et al., 2021). Considering the importance of meteorology, the diurnal variation of temperature and RH at SHT were displayed in Figure 109 and compared with SUR. In general, the SHT RH was found higher than SUR from 09:00 to 19:00 in all seasons. Therefore, the SHT atmosphere provided better conditions for the aqueous phase production of secondary aerosols during daytime. Similar to RH, the diurnal changes of air temperature were more visible for SUR. Overall, temperature at SHT was 4.4 °C lower than SUR during daytime, compared with 2.1 °C during nighttime. Furthermore, PM differences between SHT and SUR were found keen to both temperature and RH differences (Figure S11). On one hand, the lower temperature at SHT might correspond to a stronger vertical temperature gradient promoting vertical mixing of PM. What's more, the lower temperature was favorable for partition of NO_3 and semi-volatile organic species to particles.

According to above discussions, we suggest that the gas precursors, atmospheric oxidants and meteorology observed at SHT were all appropriate for aerosol formations through gas and aqueous pathway. As a result, efficient production of secondary aerosols could be expected at SHT in daytime, leading to the higher $\text{PM}_{2.5}$ concentration than SUR.

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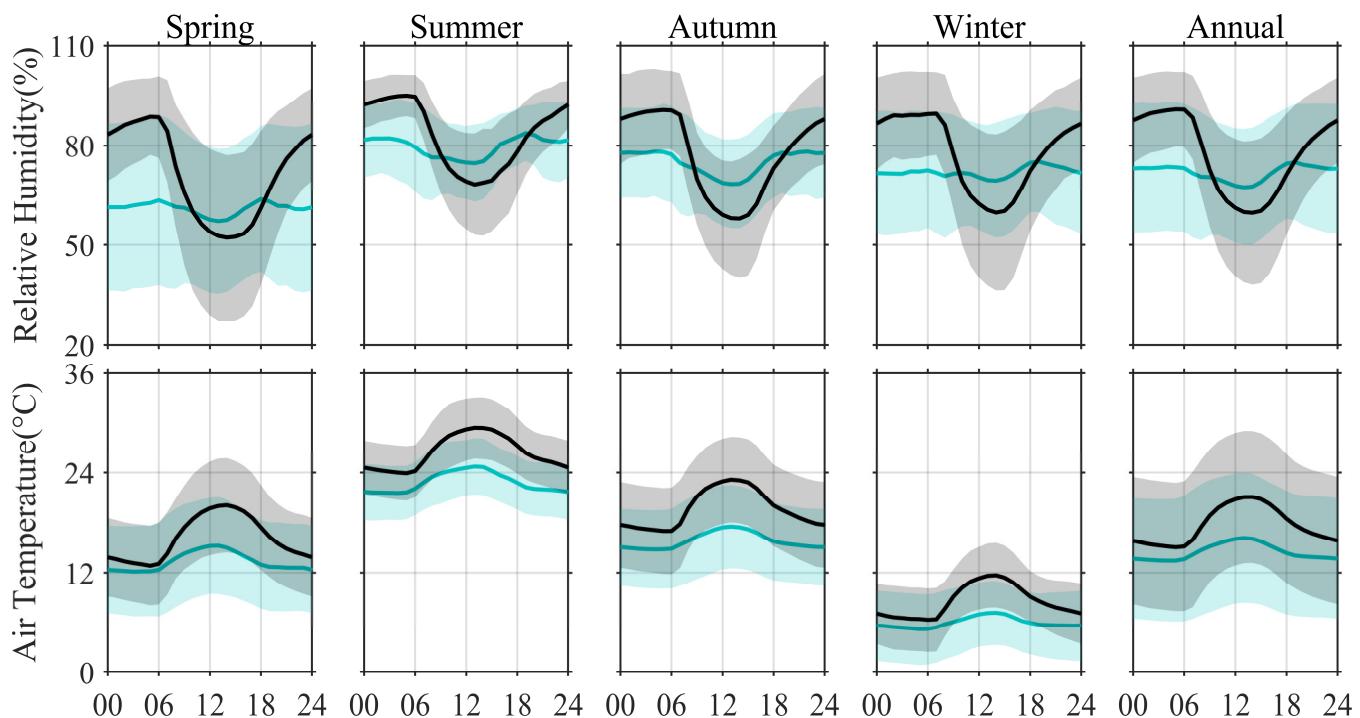


Figure 109: Same as Figure 87 but for relative humidity (%), upper panel) and air temperature (°C, lower panel).

3.3.3 Relative amplitudes of PM at SHT and SUR

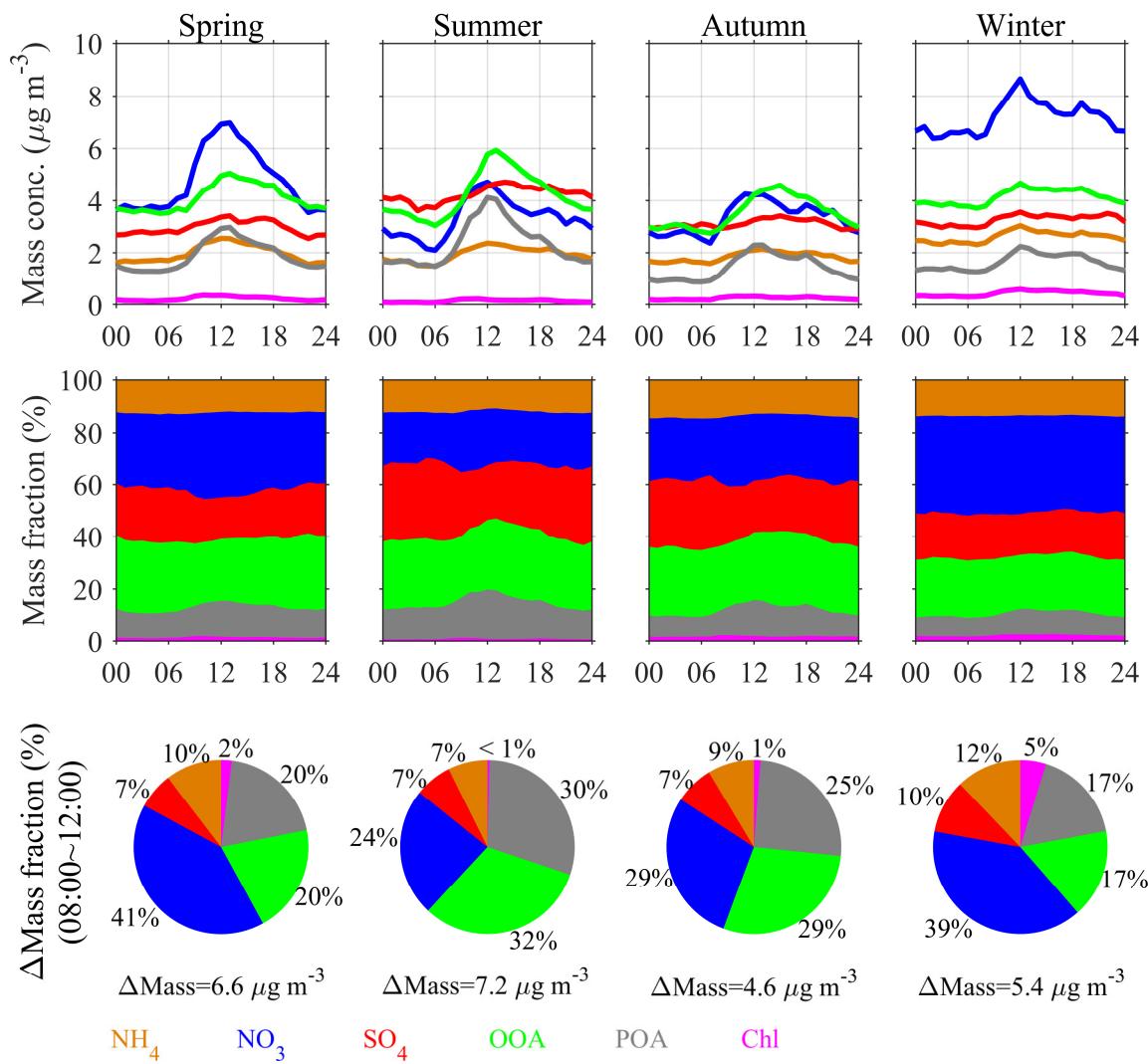
Besides diurnal tendency, diurnal amplitude of $\text{PM}_{2.5}$ variations also presented clear distinctions between SHT and SUR. We 410 introduced $(C_{\max} - C_{\min})/C_{\text{mean}} * 100\%$ to estimate the relative amplitudes (Ramps) of atmospheric species. C_{\max} , C_{\min} , and C_{mean} represent the maximum, minimum, and average of diurnal concentrations. The Ramps of $\text{PM}_{2.5}$ at SUR were 15.5%, 20.0%, 16.9% and 16.6% in spring, summer, autumn and winter, while those at SHT were 53.4%, 89.9%, 48.9%, and 27.7% respectively. Apparently, the diurnal $\text{PM}_{2.5}$ at SHT exhibited much larger Ramps than SUR, indicating more significant 415 amplitude. This was because nighttime $\text{PM}_{2.5}$ at SHT was much lower by isolating from surface. As a result, the nighttime $\text{PM}_{2.5}$ at SHT were 30-40% lower than that at SUR. While in daytime, the $\text{PM}_{2.5}$ departures between SHT and SUR were 420 greatly narrowed to 10-20% due to turbulent mixing. In addition, chemical production still played important roles in higher Ramps at SHT. It could be demonstrated by the highest Ramps happened in summer at SHT due to the significant production 425 of aerosols in daytime. In winter, the $\text{PM}_{2.5}$ enhancement during late morning to noontime was still observed at SHT, but with weaker amplitude than those in other seasons. Since the $\text{PM}_{2.5}$ concentration at SHT during this period was continuously lower than SUR, vertical mixing made positive contributions to $\text{PM}_{2.5}$ concentration at SHT. It was supposed 430 that the chemical production of aerosols at SHT was significantly inhibited due to weak solar radiation and few precursors from vertical mixing in winter except for NO_3 . While the mean $\text{PM}_{2.5}$ level was still highest in winter (Figure 76), resulted from the greater nighttime $\text{PM}_{2.5}$ level of $30 \mu\text{g m}^{-3}$, significantly higher than those in other seasons ($15\text{-}20 \mu\text{g m}^{-3}$).

3.3.4 Diurnal variation of chemical compositions at SHT

425 The diurnal variations of chemical species at SHT were further examined in Figure 1140. The major species, such as NO_3 , SO_4 and OOA exhibited unimodal pattern similar as $\text{PM}_{2.5}$, with the peak concentration appearing around noontime. These were different from the 260m observations at Beijing reported by Chen et al. (2015), their peak concentrations of secondary species usually appeared at 20:00, indicating the distinct controlling process for $\text{PM}_{2.5}$ diurnal variation between lower and upper PBL. The pronounced increasing rate of NO_3 and OOA were observed during early morning ($\sim 8:00$) to midnoon 430 ($\sim 12:00$) in spring, summer and autumn. For example, the increasing rates were estimated as $0.39 \mu\text{g m}^{-3} \text{h}^{-1}$ ($9.3\% \text{ h}^{-1}$) and $0.29 \mu\text{g m}^{-3} \text{h}^{-1}$ ($8.5\% \text{ h}^{-1}$) for NO_3 and OOA respectively, further demonstrating the significant chemical productions at SHT. NO_3 presented significant larger levels in winter and spring compared with other species. Both daytime and nighttime NO_3 435 accounted for the largest fractions of 36% in winter and 30% in spring in NR- PM_1 . Zang et al. (2022) suggested that both heterogeneous hydrolysis of dinitrogen pentoxide (N_2O_5) and the gas-phase OH oxidation of NO_2 were the important pathways for nitrate formation in Shanghai. It was noting that nighttime NO_3 in winter was extremely high up to $6\text{-}7 \mu\text{g m}^{-3}$, even exceeded daytime peak concentration in other seasons, indicating the most efficient heterogeneous production at winter night. In addition, NO_3 increasing rate in winter morning was estimated as $0.42 \mu\text{g m}^{-3} \text{h}^{-1}$, significantly higher than other species. It was attributed to the synergistic impacts of lower temperature favorable for the gas-to-aerosol partition, and higher NO_2 concentration promoting gas-phase HNO_3 productions. The enhanced level of NO_3 with altitude in PBL was also

440 reported by other studies (Zhou et al., 2018b). OOA presented most significant diurnal variability in summer, suggesting that
441 OOA formation was more sensitive to atmospheric oxidation. SO_4 formation in daytime is mainly driven by aqueous
442 reactions. Therefore, higher SO_4 at SHT in summer was expected from plenty moisture and strong solar radiation
443 accelerating its aqueous transformation, despite low SO_2 level, favorable diffusion, and wet scavenging condition of
444 atmosphere in summer. However, SO_4 presented lower diurnal variabilities in all seasons compared with NO_3 and OOA. In
445 general, distinct mass fractions of increased PM_1 (during 8:00-12:00) were observed with great contributions from NO_3 and
446 organics in spring (80.9%), summer (85.4%) and autumn (83.0%). The notable differences of mass fractions between SHT
447 and previously documented surface measurements indicate that the increased PM_1 at SHT were combined results of strong
448 chemical production as well as vertical mixing.

450 The diurnal Ramps of chemical species in the NR- PM_1 were further estimated. It was clear that POA had the most significant
451 diurnal amplitude, with the Ramps of 89.1%, 112.1%, 99.0%, and 59.4% in spring, summer, autumn and winter respectively.
452 Similarly, another primary species, Chl also had comparable Ramps of 95.8%, 102.0%, 58.7%, and 66.4% in each season,
453 suggesting the dominant impact of turbulence on the diurnal variations of primary species. In comparison, the secondary
454 species exhibited much lower diurnal Ramps, with 49.7% for NO_3 , 41.9% for OOA, 34.2% for NH_4 , and 17.5% for SO_4 . The
455 larger Ramps of primary species at SHT could be well understood, their nighttime concentrations were very low due to the
456 isolation from surface emissions, while in daytime the concentrations increased due to vertical mixing. In comparison, the
457 Ramps reduction of secondary species mostly resulted from the chemical formations at night. The pronounced formation of
458 secondary aerosols through heterogeneous pathway at night has been widely observed near surface by many studies. The
459 recent study from Zang et al. (2022) reported that heterogeneous pathway accounted for 68% of NO_3 production in winter in
460 Shanghai. It could be found in Figure 1140 that the nighttime concentrations of secondary species were significantly higher
461 than primary ones nearly in all seasons, suggesting that there also existed active heterogeneous formation of aerosols at night
462 at upper PBL level except summer.



|465 **Figure 1110:** Diurnal variations of NR-PM₁ mass concentration ($\mu\text{g m}^{-3}$) (upper panel) and mass fraction (%) (middle panel) at SHT in four seasons. Mass fractions (%) of increased NR-PM₁ during 08:00-12:00 (lower panel) in four seasons.

4 Conclusions

This study presents one-year continuous observation of the fine PM mass concentrations and chemical compositions at the 470 top of 632 m high Shanghai Tower from April 17, 2019 to April 16, 2020. The results show that SHT PM_{2.5} concentration was 16.3% lower than SUR during the entire observation period. Through investigating the diurnal changes, we found uniformly lower nighttime PM_{2.5} at SHT as results of isolations from surface emission. However, the daytime PM_{2.5} presented significant monthly changes with unexpected higher concentrations than SUR from June to October. Other than

surface, the SHT PM exhibited a consistent unimodal diurnal change in four seasons, with rapid increase of PM mass concentration starting from late morning, and a peak in the early afternoon. Combining the adequate precursors and lower temperature at SHT during daytime, we suggest strong chemical production of secondary species from both photochemical reactions and gas-to-particle partitioning at mid-upper PBL. Moreover, the averaged mass fraction of increased PM (during 8:00-12:00) at SHT revealed notably high proportions of NO_3 and organics, further demonstrating the contributions from chemical production as well as vertical mixing. In addition, we found high NO_3 concentration at SHT for both daytime and nighttime winter, implying efficient gas-phase and heterogeneous formation.

~~It should be noted that we omitted the PM originating from transport outside Shanghai throughout the discussion. However, the seasonal winds induced by Asia monsoon are quite different in upstream (ocean or land, mostly natural or anthropogenic in background) and could impact much at SHT than on the surface. This factor should be explored in future studies. Besides, PBLH is crucial for the vertical structure analysis, and direct observations of PBLH are in need to give precise view of the boundary layer processes.~~

Data availability

Hourly data used in this study are deposited at National Earth Observational Data Center (<https://chinageoss.cn/datassharing/datasetDetails/630094ef42544e709be88207>, last access: 27 November 2022), which provides open access to its data. Raw data of ACSM are archived at Shanghai Key Laboratory of Meteorology and Health, and are available on request by contacting the corresponding author.

Author contributions

JMX designed the experiments and the research. CQY, LP, WG, YG, QF and FY provided experimental assistance and the analytical method. CQY and JMX analyzed the data and performed research. All authors commented on the manuscript.

495 Competing interests

The authors declare that they have no conflict of interest.

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