

1 RESPONSE FILE:
2 The comments from referees will be indicated by a C.
3 The author responses will be indicated by an A.
4
5
6
7 Referee #1
8
9 1) C: This study uses mantle convection simulations to address the influence of pressure-,
10 temperature- and composition-dependence of thermal conductivity on the fate of a dense
11 primitive layer sitting at the bottom of the mantle, mimicking the LL(S)VPs on the Earth.
12 While previous studies have investigated the effect of those various dependencies of
13 thermal conductivity have been investigated individually, the novelty of this work is to
14 consider consider the interplay between them in the context of the Earth's lower mantle
15 structure. While I think that the results are worth being published, their presentation in
16 the current manuscript is hard to follow and lacks a clear narrative thread, and more
17 importantly, lacks in interpretation. Therefore, I think the paper would benefit from a
18 thorough re-organization of the results section.
19
20 A: We have decided to expand the introduction and rearrange the results section so that
21 the manuscript is easier to follow.
22 CHANGE: Paragraphs were added discussing the dependencies of thermal conductivity.
23
24
25 General
26
27 2) C: The study adopts a rather complex baseline set up (mixed basal + internal heating, phase
28 transition, yield stress) to study the effect of pressure, temperature and composition on
29 thermal conductivity. While my preference goes to simpler models when trying to unravel
30 such systematic trends, one can argue that the current set-up is relevant for an Earth-like
31 case, and I think it is fair.
32
33 A: Yes, the mantle convection model set-up we consider is intended to be Earth-like.
34 However, we do not intend on reproducing the Earth mantle's history. Within this framework
35 we have been able to examine the effect of the thermal conductivity model on the evolution
36 of the primordial layer at the bottom of the mantle.
37 CHANGE: no change necessary
38
39 3) C: I find the introduction a bit too light. I would expect some more background on the
40 physics of the various dependencies of thermal conductivity, and in particular that the
41 trends be clearly announced: increase of pressure results in increase of thermal
42 conductivity, increase in temperature results in decrease in thermal conductivity, increase
43 in iron content results in decrease in thermal conductivity (it seems). Also, it seems that
44 considering a variable thermal conductivity is particularly relevant for compressible
45 convection. If this is indeed the case, I would expect more time to be spent discussing
46 why.
47
48 A: We agree with R1 that the introduction can be expanded. Regarding the thermal conductivity,
49 each of the component (dependency) of the conductivity model will be discussed. In addition, we
50 will now address why conductivity is relevant for compressible mantle convection.
51 CHANGE: changes outlined in response to point 1)
52
53 4) C: In my opinion, the main weakness of the paper in its current state is the lack of apparent
54 organization in its exploration of the parameter space, which makes it pretty confused and
55 fails at highlighting general tendencies in the effect of the different dependences of the
56 thermal conductivity. This is probably due in part to the simultaneous changes in
57 investigated parameters and diagnostic quantities from one section to the other (first time-
58 evolution of the heat flux varying KC and n, then instantaneous mantle structures varying
59 everything, and finally time-evolution of entrainment again varying KC and n). While I
60 seem to get the idea in the progression, I had a really hard time getting a clear picture
61 out of this section. Maybe starting by isolating the effect of each dependencies, and then
62 considering their correlations could make things simpler. Maybe also distinguishing more
63 systematically between the effect on mantle flow (pile structure) and on the time-
64 evolution (e.g. of the CMB and surface heat fluxes) could help. Another possibility would
65 be to add dependencies successively to a fiducial model, for which a preferred value of KD
66 would be selected before considering the effect of n, and in turn a preferred value of n
67 would be selected before adding a third layer with KC. That could spare some cases of the
68 parameter space if there are reason to think they are less relevant. These are suggestion,
69 hopefully that can help. Anyway, one important thing which is lacking in my opinion is also
70 some theoretical speculation on the expected effect of the parameters that should come
71 before presenting the results, and would help their interpretation (e.g. we expect the
72 effect of increasing KD to increase the thermal conductivity in the lower mantle and thus to
73 make it more stable to convection (it decreases Ra) by homogenizing its temperature,
74 potentially building heat... etc.). I think it is better for the reader to be incited to think in
75 advance of seeing the results.
76
77 A: We consider the suggestions that R1 outlines for reorganizing the results section. Introducing the
78 reference case $K_D = 2.5$ first is too much too quickly for the reader. We now separate all the
79 effects so that the progression of the results is easier to digest.
80 CHANGE: The subsections of the results are now:
81 3.1) Effect of a purely depth- dependent conductivity
82 3.2) Effect of a temperature- and depth- dependent conductivity
83 3.3) Including the effect of composition- dependent conductivity
84 3.4) Long-term stability of thermochemical reservoirs featuring mineral physics derived conductivities
85
86 Figure 3 in 3.1) now features field snapshots for purely depth- dependent cases.
87 Figure 4 in 3.2) and Figure 5 in 3.3) now highlight field snapshots for the
88 non-zero end-member linear depth- dependent cases.
89 Figure 6 in 3.4) now features field snapshots for different thermal conductivity models featuring
90 a mineral conductivity derived depth- dependence K_{DH} .
91 Figure 7 in 3.4) shows the corresponding timeseries for cases in Figure 6.
92 Old Figure 7 is moved to Figure 8 and the corresponding timeseries is included in Figure 9.
93
94 Specific
95
96 5) C: This article will probably mainly be read by people familiar with the equations of mantle
97 convections. Nevertheless, I think it would be beneficial to write down the conservation
98 equations (they are not even in the supplementary materials!), at least the heat
99 conservation where the thermal conductivity appears, as it would make clear where the
100 supplementary mechanism induced by varying thermal conductivity operate.
101
102 A: The heat conservation equation was added to the Methods section to help the reader
103 understand where the conductivity model influences mantle dynamics.
104 CHANGE: as above.
105
106 6) C: A plot of the thermal conductivity profile corresponding to the reference state for the
107 various KD, KC and n (not all of them but the few most relevant combinations, e.g. the
108 cases presented in Figure 2) would be insightful. It could be a supplementary panel in
109 Figure 1, or a stand-alone figure.
110
111 A: A new figure with a plot of the initial thermal conductivity profiles for cases featuring
112 mineral physics defined conductivities is now included so that the conductivity reductions
113 from different temperature- and composition- dependencies is clarified.
114 CHANGE: A new Figure 2 is inserted into the Methods section to highlight initial conductivity

115 profiles. A paragraph describing these variations is also included.

116
117 7) C: I don't really understand why simulations run until 11 Gyr, but some values are averaged
118 around 4.5 Gyr. Anyway, the averaged values reported in Table 1 do not seem to be used.

119
120 A: The simulation time is long (~11 Gyr) to allow the system to reach a statistically steady
121 state and to investigate how the heterogeneous conductivity will affect the long-term evolution
122 of the primordial layer. The simulation averages are taken towards the start of simulations'
123 statistically steady state (approximately 4.5 Gyr). Again, we do not intend to model the exact
124 evolution of the Earth's mantle. The conditions for simulating Earth's history
125 (i.e., initial conditions, decaying heat sources, and cooling bottom boundary) are not included
126 in our model setup. In the current version of the manuscript, the averaged values were inset
127 within the annulus snapshots and their trends were discussed within the results subsections.
128 Specific values were not used often so that the text would not be inundated with numbers.
129 We now incorporate more references to the Table 1 averages into the manuscript.

130 CHANGE: An updated Table 1 (including supplemental cases) is referred to in the updated results section.

131
132 8) C: Dimensional and non-dimensional quantities are often mixed, which lacks a bit of rigor.

133
134 A: The values presented in the manuscript will be presented as dimensional.
135 CHANGE: as above. Colour bar for conductivity cases is now modified so that one can also interpret
136 the field data as non-dimensional.

137
138 Technical
139 L. 48: "compositional- dependent" - "composition-dependent"
140 A, CHANGE: for clarity, "composition- dependent" will now be used throughout the manuscript.

141
142 L. 84: "a quadratic that smoothly" - "a quadratic curve that smoothly"
143 A, CHANGE: inserted the word 'curve'.

144
145 Figure 1: the markers for "<200 ppm water" are hard to see, please change for a more
146 contrasting colour.
147 A, CHANGE: markers for "< 200 ppm water" are changed to bright green to be more contrasting with the background.

148
149 L. 118: "We first defined a purely depth- dependent reference case characterized by
150 depth- dependence, $K_D = 2.5$, with lower mantle conductivities comparable to current
151 estimates" in Figure 1, it seems that $K_D = 10$ is the closest match to lower mantle
152 estimates.
153 A, CHANGE: The case with a purely depth- dependent thermal conductivity and $K_D = 2.5$ is considered as an analogue
154 for models that already account for the combined depth- and temperature- dependences. In the latter (hypothetical)
155 model of thermal conductivity, the depth- dependent component would have conductivity values that are much greater
156 than those defined by $K_D = 2.5$. HOWEVER, this section has changed and the sentence is no longer relevant.
157 No change necessary.

158
159 L. 121: "by approximately 75%" it seems less than that in Figure 2.
160 A, CHANGE: "by approximately 75%", this refers to a comparison of heat flow values averaged about 4.5 Gyr for cases #4 and #8.
161 This comparison is between a purely depth- dependent conductivity case (#4) and heterogeneous conductivity case (#8).
162 HOWEVER, this section has changed and this sentence no longer refers to cases with $K_{(D)} = 2.5$. No change necessary.

163
164 L. 122: "in agreement with Li et al. (2022) findings" - "in agreement with Li et al.
165 (2022)'s findings"
166 A, CHANGE: fixed the citation to read "... Li et al., (2022)'s findings".

167
168 L. 150: "We observed that T_{prim} increased with greater temperature dependence (top-to-
169 bottom rows in Figure 3)." I don't see it.
170 A: The mean temperature of primordial material is indicated by T_{prim} (not to be confused with the global mean
171 temperature of the system $\langle T \rangle$). T_{prim} is inset within the annulus and is indicated for each case. T_{prim} can be seen
172 decreasing for cases with the same K_D value as n is increased. In addition, the temperature field is offset with respect to the
173 CMB temperature so that the differences in hotter temperatures within the piles is easier to see. The boundary between regular
174 mantle material and primordial material is indicated by a green contour. When n is increased, the red colour within these
175 contours becomes more saturated.
176 CHANGE: No change necessary. This trend is still visible in the new Figure 4.

177
178 For the snapshot figures, when a parameter is held constant, write it in the caption rather
179 than for each snapshot. That will lighten a bit the figures.
180 A, CHANGE: Snapshot figures have been revised so that superfluous conductivity model labels are omitted.

181
182 Figure 5: Please alternate colormaps between the different quantities plotted. In
183 particular, don't take a "divergent" colormap for the primordial mantle one which only has
184 two extremal values, which are both very dark and hard to distinguish in the current plot.
185 A, CHANGE: The colormaps for each different field are now alternated. Conductivity field
186 snapshots now have a colormap that can also be interpreted in a non-dimensional values
187 (i.e., the conductivity ratio).

188
189 L. 172: "and is equivalent to" - "and which is equivalent to" ?
190 A, CHANGE: inserted the word 'which'.

191
192 Figure 6: Same as figure 5: it is very hard to know which colormap corresponds to the snapshot
193 and which corresponds to the time-evolution plots. Also please change colours
194 for the onset of instability (the magenta is pretty hard to see).
195 A, CHANGE: The colormap is changed. The onset of instability is now indicated by a dashed cyan line.

196
197 L. 189: "may not compensate (or be exceeded) by" - "may not compensate (or be
198 exceeded by)"
199 A, CHANGE: The parantheses have been moved.

200
201 Supplementary Materials
202 L. 43: "and the depth variation of thermal expansivity imply a depth average" - "and the
203 depth variation of thermal expansivity implies a depth average"
204 A, CHANGE: "imply" -> "implies".

205
206 L. 176: "The density anomalies [...] is calculated" - "The density anomalies [...] are
207 calculated"
208 A, CHANGE: "is" -> "are".

209
210 Figure S1: the y-axis is non-dimensional height, not depth.
211 A, CHANGE: ALL vertical axes in radial profiles now correspond to depth values.

212
213 Figure S4: What do line styles correspond to?
214 A, CHANGE: The case numbers have been added to the legend so that the line styles are now clear.
215 Line style descriptions are also included in the captions.

216
217 Referee #2

218
219 C: In this paper the authors carried out numerical experiments of thermochemical mantle
220 convection by varying the spatial changes in thermal conductivity to investigate the
221 temporal changes in the distributions of dense "primordial" materials initially imposed
222 above the bottom boundary. My honest impression is, unfortunately, that the manuscript
223 is very hard to follow because of its poor organization and description from the reasons
224 summarized below. I therefore strongly suggest that the authors should thoroughly revise
225 the manuscript before the reconsideration.

226
227 A: In addition to the comments made by R1, we have decided to thoroughly revise the manuscript
228 so that it is easier to follow.

229 CHANGE: Relevant changes have been outlined in responses to Referee #1.
230
231 C: 1. One of my major criticism is that the main issues in this study were not well described
232 either in the abstract or introduction section. In my understanding the major intention of
233 the authors is to reduce the thermal conductivity in the deep mantle to much lower levels
234 than those simply expected from the dependence on temperature and pressure (or
235 depth), in order to induce an "instability" from the initial layer of dense materials within a
236 sufficiently short period of time. The authors should clearly indicate their ultimate
237 motivation in earlier parts of the manuscript.
238
239 A, CHANGE: In alignment with the comments of R1, the introduction section was expanded and
240 the intentions of our study were made clearer.
241
242 C: 2. I am not well convinced of the meaningfulness of the onset time of "instability" from
243 the initial layer of dense materials. It seems to me that the authors had assumed that the
244 deformation of the basal layer of dense materials occurs only in an "intrinsic" manner
245 owing to the thermal buoyancy. However, several earlier studies had demonstrated the
246 ultimate importance of "extrinsic" deformation induced by the cold subduction from the
247 top surface. I do not therefore think that the onset time of "instability" can be a good
248 measure to investigate the influence the thermal conductivity at depth on the thermal
249 buoyancy in the basal layer of dense materials.
250
251 A: We calculate the onset time for instability from the temporal variations in the average
252 height of dense material. It is true that these variations in average height do not discriminate
253 between 'intrinsic' (i.e., thermal buoyancy) or 'extrinsic' (i.e., downwellings) deformation.
254 From examining the average height timeseries and animations of the fluid flow we can see the
255 influence of downwellings. The first downwellings imprint on the initial dense layer but do
256 not result in a rapid uplift of material (sufficient to eject blobs of dense material).
257 Once the initial dense layer has organized into piles, downwellings tend to move dense material
258 laterally over the CMB but not rapidly increasing the pile height. Furthermore, we find that it is easier
259 for downwelling currents to push primordial material that has been made lighter due to their retained heat.
260 We agree that downwellings are important in deforming the dense primordial layer.
261 CHANGE: In section 3.4) this mechanism is discussed in addition to the thermal effect regarding instability.
262
263 C: 3. I felt quite odd with the authors' exaggerated claim on the overall profiles of thermal
264 conductivity including "We find that the temperature- and depth- variations combined
265 characterize the mean conductivity ratio from top-to-bottom" in the abstract. It is quite
266 obvious from the assumptions made by the authors themselves.
267 A: Right. This statement is a bit of a tautology. The intention for this statement is that for
268 Earth-like models that consider a variable thermal conductivity, it may be ill advised to isolate
269 for just one dependency. On one hand, to simulate a 2-pile configuration, a purely depth- dependent
270 conductivity may be employed, but its top-to-bottom ratio should implicitly emulate the temperature
271 and composition effects. On the other hand, a purely temperature- dependent conductivity will result
272 in the entrainment of a dense primordial layer. By assuming a parametrized conductivity model, it is
273 predictable what the mean conductivity ratio from top-to-bottom will be (having specified the temperature
274 contrast and the dependencies of the conductivity model).
275 CHANGE: This sentence was omitted and changed to read:
276 "The mean conductivity ratio from bottom-to-top indicates the relative competition between the
277 decreasing effect with increasing temperature and the increasing effect with increasing depth. We find that,
278 when depth- dependence is stronger than temperature- dependence, a mean conductivity ratio > 2 will result
279 in long-lived primordial reservoirs."
280
281 C: 4. In Section 2.1 the authors should state the reason why the phase change from
282 perovskite (pv) to post-perovskite (ppv) is ignored in their numerical model. If they
283 believe that the pv-ppv phase change is of little importance on the dynamic behaviors of
284 the basal layers of dense materials, the authors should make clear the reason why.
285
286 A, CHANGE: In the methods section, we now discuss the reasons why the phase change from perovskite (pv)
287 to post-perovskite (ppv) is ignored in our numerical model. The effects of the pv-ppv transition
288 properties on the stability and structure of primordial reservoirs has been investigated previously by
289 Li et al., (2015). There are many controlling parameters for the pv-ppv transition including the temperature
290 of the CMB, the viscosity contrast between pv and ppv, and the viscosity contrast between ppv and primordial
291 material that can affect the stability of piles. For instance, weak ppv (i.e., low viscosity contrast between
292 pv and ppv) and a low Tcmb (i.e., Tcmb ~ 3350 K) can result in entrainment of primordial reservoirs.
293 Because of the model setup we consider in our study, the inclusion of a pv-ppv transition will result in the
294 entrainment of a dense primordial reservoir. Thus, the pv-ppv phase transition will mask the effect of thermal
295 conductivity on the stability of primordial reservoirs that we are examining.
296
297 C: 5. I was quite disgusted to see that the profile given by equation (2) is denoted by
298 "KD=9.185". Such a denotation should be used only for the profiles given by equation (1)
299 !!
300 A, CHANGE: All usage of the label "K_D = 9.185" will be replaced with "K_{DH}" to indicate that this
301 depth profile was defined by parameterizations published in Deschamps and Hsieh, (2019).
302
303 C: 6. Near equation (3) the authors should indicate the magnitude of the reduction in
304 thermal conductivity due to the increase in temperature within the modeled domain by the
305 choice of n=0.5 and n=0.8.
306
307 A, CHANGE: The magnitude of the reduction in thermal conductivity depends on temperature (which changes with depth).
308 A simple calculation was added to the methods section to show how much the conductivity is reduced at CMB temperatures
309 for different n and K_{C} values. The reduction in conductivity will also be shown by the inclusion of a new figure
310 which shows the initial conductivity profiles for cases featuring K_{DH} with different temperature- and composition- dependence.
311
312 C: 7. Near equation (4) and later, "compositional correction" should be rephrased with
313 "compositional dependence". The word "CORRECTION" could imply that the numerical
314 experiments without the compositional dependence in thermal conductivity are
315 meaningless.
316
317 A, CHANGE: In alignment with the comments of R1, all usage of "compositional correction" was rephrased to "composition- dependence".
318
319 C: 8. In Section 3.1 "QCMB" is used without explicitly defined.
320
321 A, CHANGE: Q_{CMB} and Q_{SURF} will be clearly defined in Section 3.1.
322
323 C: 9. To show the 2-D distributions of thermal conductivity in Figures 4 and 5, the authors
324 should show the ratio of thermal conductivity to its surface value (K_S) rather than the
325 values of conductivity itself.
326
327 A, CHANGE: As suggested by R1, we will keep consistent with dimensional values. We currently present 2D distributions of thermal
328 conductivity in Figures 4 and 5. By plotting the conductivity fields relative to the surface value k_{S}, it would be converted
329 to the non-dimensional conductivity field. The colour map is now adjusted so that conductivity values can be interpreted
330 nondimensionally with colour. That is, values < 3 W/m/K (ratio < 1) tend to blue-ish hues and values > 3 W/m/K (ratio > 1) tend to brown hues.
331
332 C: 10. The paper by Marzotto et al. (2020) has not been cited anywhere in the main text.
333
334 A, CHANGE: Marzotto et al., (2020) will be removed from the references. Conductivity data points included in Figure 1 are from the
335 references cited within the caption.
336
337 C: 11. Near equation (8) of the Supplement, the spatial coordinates should not be in
338 Cartesian (x,y,z) but in 2-D polar (r,phi) in this study. Similarly in equation (12) of the
339 Supplement, the coordinate is not in 3-D polar (r,theta,phi).
340
341 A, CHANGE: Coordinates for 3D geometry (x,y,z) or (r,theta,phi) will be replaced by 2D spherical annulus coordinates (r,phi).
342

343

344 C: 12. In Section 3.1 of the Supplement, I think that the authors can use the potential
345 temperature instead of the "adiabatic correction" $a(z)$.

346

347 A, CHANGE: The potential temperature definition will also be stated in addition to the adiabatic correction.

348

349 C: 13. The paper by Xu et al. (2004) has not been cited anywhere in the Supplement.

350

351 A, CHANGE: Xu et al., (2004) is removed from the references. This reference was included in the methods section discussion on the
352 thermal conductivity model (Section 2.2) and had been moved from the supplement to the main text. The reference to Klemens, (1960)
353 in the supplement is also be removed for the same reason.

354