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2	Effects of topographic and meteorological parameters
3	on the surface area loss of ice aprons in the Mont-Blanc massif (European Alps)
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15	Abstract
16	Ice aprons (IAs) are part of the critical components of the Alpine cryosphere. As a result of the
17	changing climate over the past few decades, deglaciation has resulted in a surface decrease of IAs,
18	which has not yet been documented, except for a few specific examples. In this study, we quantify
19	the effects of climate change on IAs since the mid-20 th century in the Mont-Blanc massif (western
20	European Alps). We then evaluate the role of meteorological parameters and the local topography
21	in the behaviour of IAs. We precisely mapped the surface areas of 200 IAs using high-resolution
22	aerial and satellite photographs from 1952, 2001, 2012 and 2019. From the latter inventory, the
23	surface area of the present individual IAs ranges from 0.001 to 0.04 km ² . IAs have lost their surface
24	area over the past 70 years, with an alarming increase since the early 2000s. The total area, from
25	7.93 km ² in 1952, was reduced to 5.91 km ² in 2001 (-25.5 %) before collapsing to 4.21 km ² in
26	2019 (-47 % since 1952). We performed a regression analysis using temperature and precipitation
27	proxies to understand better the effects of meteorological parameters on IA surface area variations.
28	We found a strong correlation between both proxies and the relative area loss of IAs, indicating
29	the significant influence of the changing climate on the evolution of IAs. We also evaluated the

role of the local topographic factors in the IAs area loss. At a regional scale, factors like direct
solar radiation and elevation influence the behaviour of IAs, while others like curvature, slope, and
size of the IAs seem to be rather important on a local scale.

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Key words: Ice aprons, surface area loss, topographic factors, meteorological parameters, Mont
Blanc massif.

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38 **1 Introduction**

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The predicted shift in climate dynamics over the next decades will undoubtedly have severe consequences on the high mountain environments, primarily on glacier extent (Rafiq and Mishra, 2016; Kraaijenbrink et al., 2017; IPCC, 2021), permafrost (Magnin et al., 2017) and ice and snow cover (Rastner et al., 2019; Guillet and Ravanel, 2020). The effects of climate change on glaciers constitute a remarkably well-discussed topic in the scientific community (Yalcin, 2019).

Meteorological parameters (mainly temperature and precipitation) are the main driving forces responsible for these changes (Scherler et al., 2011; Bolch et al., 2012; Davies et al., 2012). Shifting temperature and precipitation trends lead to the advance or retreat of glaciers both in volume and surface area (Liu et al., 2013; Yang et al., 2019). On a regional and global scale, many authors have studied the impacts of climate warming on glacier retreats and, consequently, on the hydrology of the mountain environments (*e.g.*, Baraer et al. (2012); Sorg et al. (2014); Frans et al. (2016); Coppola et al. (2018)).

However, as observed by Furbish and Andrews (1984), Oerlemans et al. (1998), Hoelzle et al.
(2003) and Salerno et al. (2017), glaciers present in the same climate regime can respond to climate
change in different ways. The local climate variations can partly explain these variable responses.
However, many of these variations result from different morphometric (size, shape, length) and
topographic (altitude, slope, aspect, curvature, terrain ruggedness) characteristics.

57 Several studies have been devoted to understanding the linkage between topographic factors and 58 the response of glacier/ice bodies (*e.g.*, Davies et al., 2012; De Angelis, 2014; Salerno et al., 2017).

60 World Glacier Monitoring Service (WGMS) monitors glacier changes in all the major mountain

61 regions of the world. However, most mapping and monitoring studies on a global scale focus on

62 massive glaciers since they are generally assessable and easier to monitor compared to other ice

63 features (Liu et al., 2013).

64 Studies are rare for small glaciers or ice bodies, which generally show a more pronounced response

to climate change (Oerlemans and Reichert, 2000; Triglav-Čekada and Gabrovec, 2013; Fischer

66 et al., 2015). This has led to a critical gap in our understanding of their behaviour and mass balance

67 estimates. As part of this trend, ice aprons (IAs), sometimes also referred to as 'rock faces partially

covered with ice' (Gruber and Haeberli, 2007; Hasler et al., 2011), have also received poorattention from the scientific community.

These small ice accumulations on steep rock slopes are commonly found in all significant 70 71 glacierized basins worldwide. However, a concrete and well-summarized definition for IAs is still missing from the literature. Previously, many authors like Benn and Evans (2010), Singh et al. 72 (2011) and Cogley et al. (2011) tried to define IAs, but the most precise definition for IAs up to 73 now can be found in Guillet and Ravanel (2020) for the Mont-Blanc massif (MBM; European 74 Alps). These authors defined IAs as "very small (typically smaller than 0.1 km² in extent) ice 75 bodies of irregular outline, lying on slopes $>40^{\circ}$, regardless of whether they are thick enough to 76 77 deform under their weight". The small spatial extent of the IAs makes them very difficult to map 78 and monitor. Also, they are typically present in extremely challenging topographies on isolated 79 steep slopes. Cogley et al. (2011) specified that IAs are "lying above the head of a glacial bergschrund which separates the flowing glacier ice from the stagnant ice, or a rock headwall". 80

Because of their presence on steep slopes, IAs are essential natural elements for the practice of 81 82 mountaineering, especially in famous destinations like MBM (Barker, 1982). IAs are passing 83 points for many classic mountaineering routes (Mourey et al., 2019). Hence, the loss of IAs is a 84 severe threat to the iconic practice of mountaineering, inscribed in 2019 by UNESCO on the Representative List of the Intangible Cultural Heritage of Humanity. IAs on steep rock walls also 85 carry the critical role of covering steep rock slopes and preventing them from direct exposure to 86 direct solar radiation, thus partly preventing the warming of the underlying permafrost. In addition, 87 88 a recent study by Guillet et al. (2021) showed that the ice present at the base of the Triangle du 89 Tacul IA could be older than 3 ka, making IAs a potential important glacial heritage.

Guillet and Ravanel (2020) showed that IAs in the MBM have lost mass since the Little Ice Age (LIA). Based on six different IAs, their study also showed an acceleration in the shrinkage since the 1990s. They linked the loss of IA area with meteorological parameters, mainly air temperature and precipitation. It was thus the first documented evidence that IAs have been losing ice volume due to the changing climate. However, since this study was local and based on only a few IAs, the authors could not consider other factors, such as the local topography critical for small glacier bodies (Hock, 2003; Laha et al., 2017).

97 Thus, to overcome these limitations, we propose a large-scale analysis to ascertain the relationship of the area loss of IAs with the meteorological parameters, mainly air temperature and 98 precipitation, using a more comprehensive database (c. 200 IAs) covering the whole MBM. The 99 large inventory of IAs has been surveyed thanks to high-resolution aerial and satellite images from 100 101 1952, 2001, 2012 and 2019. Further, based on our inventory, we also evaluate the impacts of the topographic/geometric controls on the area changes of IAs. For this, we consider the size of IA, 102 103 elevation/altitude, slope, curvature, Topographic Ruggedness Index (TRI), direct solar radiation and permafrost conditions (classified together as topographic factors) based on past studies on 104 105 similar themes (e.g., Oerlemans et al., 1998; Warren, 2008; DeBeer and Sharp, 2009; Jiskoot et 106 al., 2009; Davies et al., 2012; Salerno et al., 2017).

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109 2 Study area and the impacts of climate change in the region

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111 The Mont-Blanc massif (Fig. 1) is located in the north-western (external) Alps between France, 112 Switzerland, and Italy. It covers $c.550 \text{ km}^2$ and displays some of the highest peaks in the European 113 Alps; a dozen peaks have elevations greater than 4000 m a.s.l. MBM thus shows a significant 114 variation in the elevation range throughout the massif; the lowest point of the massif is at 1050 m 115 a.s.l. (Chamonix) and the highest, the top of Mont Blanc at 4808 m a.s.l.

Because of its high elevation, the MBM is also the most glacierized massif in the French Alps (Gardent et al., 2014). There are about 100 glaciers often bordered by steep rock walls, including l2 glaciers larger than 5 km². The steep and irregular terrain facilitates the development of many unique ice bodies like cold-based hanging glaciers or IAs. Figure 2 shows two examples of the locations of the IAs on the steep N faces from the study region.

As a result of an asymmetry of the massif, 6 of the largest glaciers are located on its NW French 121 side, where slopes are gentler than the Italian side and glaciers are well fed by the westerly winds 122 123 while melting is reduced by the protection of the shaded North faces. The SE Italian side is characterized by smaller glaciers and generally steeper slopes bounded by very high sub-vertical 124 rock walls. This asymmetry is also evidenced by the difference in the meteorological conditions 125 126 observed on the two sides of the massif. For example, the Mean Annual Air Temperature (MAAT) recorded in Chamonix (at 1044 m a.s.l.) is +7.2°C while that in Courmayeur (1223 m a.s.l.) is + 127 10.4°C (Deline et al., 2012). Comparing the annual MAAT values from 1934 to today shows that 128 MAAT increased by > 2.1°C in Chamonix (*MétéoFrance* data). Moreover, the increase in MAAT 129 from 1970 to 2009 was almost four times faster than from 1934 to 1970 (Mourey et al., 2019). Not 130 only at lower elevations, but the MAAT also increased by 1.4°C at elevations exceeding 4000 m 131 132 a.s.l. between 1990 and 2014 (Gilbert and Vincent, 2013). The MBM has experienced nine summers characterized by heatwaves (where maximum temperatures for at least three consecutive 133 134 days exceed a heatwave temperature threshold defined for the region) since 1990 (1994, 2003, 2006, 2009, 2015, 2017, 2018, 2019 and 2020), with the one as recent as 2018 being the second 135 136 (after 2003) hottest. The average annual precipitation recorded for Chamonix is 1,288 mm, and 854 mm for Courmayeur (Vincent, 2002). The precipitation rates in the MBM have remained 137 138 relatively constant since the end of the LIA, but there is a noticeable decrease in the number of 139 snowfall days relative to the total precipitation days below 2700 m a.s.l. (Serguet et al., 2011).

140 Global warming has led to a general retreat trend of the MBM glaciers since the end of the LIA despite small re-advances culminating in 1890, the 1920s and the 1980s (Bauder et al., 2007). The 141 recorded loss of glacier surface area was 24 % of the total area from the end of the LIA to 2008 142 (Gardent et al., 2014). The reported loss of ice thickness is also noteworthy. For example, the loss 143 144 of ice thickness at the front of the "Mer de Glace" glacier (1650 m a.s.l.) from 1986 to 2021 is 145 145 m; the Argentière glacier (1900 m a.s.l.) has lost 80 m in thickness from 1994 to 2013 (Bauder et al., 2007). At 3550 m a.s.l., the surface of the Géant glacier also lowered by 20 m between 1992 146 147 and 2012 (Ravanel et al., 2013). The glacier retreat and shrinkage concur with the Equilibrium Line Altitude (ELA) that rose by 170 m between 1984 and 2010 in the western Alps (Rabatel et 148 al., 2013). As a result of the loss of ice volume, the density of open crevasses has considerably 149 increased, along with an increase in bare ice areas. In some instances, ice volume loss leads to 150 instability of steep slopes, and serac falls from the front of warm and cold glaciers are more 151

152	frequent (Fischer et al., 2006). This latter process can be typical during the warmest periods of the
153	year (Deline et al., 2012). Warming trends also intensify moraine erosion, resulting in an increase
154	in rockfall and landslide events (Deline et al., 2015; Ravanel et al., 2018). Degradation/warming
155	is another critical concern for permafrost (e.g., Haeberli and Gruber, 2009).
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158	3 Data description
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160	This section describes all the datasets obtained from diverse sources used in this study (Table. 1).
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162	3.1 Digital Elevation Model
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164	Since one of the main aims of our study was to perform a joint analysis of the behaviour of small
165	ice bodies and the local topography, it was paramount to have a robust high resolution and accurate
166	Digital Elevation Model (DEM) for the study region. To avoid the uncertainties that most global
167	DEMs are plagued with and to overcome the problem of different DEM origins on the French and
168	Italian sides of the MBM, we built our own DEM. As part of the CNES Kalideos Alps project,
169	stereoscopic sub-meter resolution optical images from the Pleiades constellation were acquired.
170	Using the pair of stereo panchromatic images (25/08/2019), a 4 m resolution DEM was computed
171	using the Ames Stereo Pipeline (ASP), an open-source processing chain developed by Shean et al.
172	(2016). The parameters used for the processing were kept the same as those of Marti et al. (2016).
173	The second part of the processing involved accurately co-registering the newly built DEM with an
174	existing reference DEM of high precision and accuracy. For this purpose, we used the automatic
175	DEM co-registration methodology given by Nuth and Kääb (2011). As a 'reference', we used a 2
176	m LiDAR DEM for the area around the Argentière glacier (8 * 2.5 km spatial extent) (Fig. 3a)
177	built by the Institut des Géosciences de l'Environnement (IGE) to co-register the 'source' 4 m
178	Pleaides DEM (Fig. 3b) generated in the previous step. A precisely co-registered, high-resolution,
179	robust 4 m DEM was obtained at the end of the processing steps. More detailed information about
180	the processing parameters for DEM generation and co-registration can be found in Kaushik et al.
181	(2021). We used this DEM to compute topographic parameters like slope, aspect, curvature,
182	elevation, TRI, mean annual rock surface temperature (MARST) and direct solar radiation.

184 **3.2 Optical aerial and satellite images**

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This study relies on high-resolution aerial and satellite images (Table 1). Working with data from 186 different sources allowed us to tap into the wealth of data for comparison. Spanning over seven 187 188 decades and covering the entire MBM, ortho-images for 1952, 2001 (0.5 m resolution) and 2015 (0.2 m resolution) were downloaded from Géoportail IGN (French Institut national de 189 l'information géographique et forestière), while the panchromatic and XS images from SPOT 6 190 and Pleiades at 2.2 m and 0.5 m respectively were downloaded from the Kalideos Alps website. 191 Considering the small dimensions of the ice bodies, we could only work with high-resolution 192 optical images covering the entire MBM. We were thus limited by only one set of excellent quality 193 194 images for 1952 and 2001, as very high-resolution images for this study period were unavailable from any other source. Hence our mapping exercise relied only on the ortho-images for these two 195 196 time periods. For 2012 and 2019, we had data from multiple sources (Pleiades, SPOT and orthoimages) to deal with the problems associated with the lack of coverage, cloud cover, illumination, 197 198 shadow, and seasonal snow cover that made visual interpretation difficult. We used a combination of Pleiades and SPOT 6 XS images for mapping the IAs boundaries, with validation of results 199 200 conducted with the help of the ortho-images. To avoid overestimating the extent of IAs, we utilized 201 images acquired at the end of the summer period (late August or early September). Considering 202 that our optical images came from many sources, it was necessary to accurately co-register all images. We used the automatic 'image to image' co-registration tool in ENVI 5.6. The process 203 204 included locating and matching several feature points called tie points in a 'reference' image and a 205 'warped' image selected for co-registration. Here, we used the Pleiades panchromatic image of 206 2019 as a reference, and all the warped images were accordingly co-registered. Both coarse and 207 fine co-registration procedures were performed, and the co-registration process was stopped when the RMSE values achieved were less than half the pixel resolution of the warped image based on 208 the recommendations of Han and Oh (2018). A more detailed description of the co-registration 209 210 process was discussed in Kaushik et al. (2021).

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214 **3.3 Meteorological data**

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216 Proxies to define accumulation and ablation phases were built to explore the correlated variations 217 in the surface area of IAs with the changing climate. A similar study for 6 IAs was performed by Guillet and Ravanel (2020); we aim to test the validity of their results with a more extensive 218 219 database (c.200 IAs) in the entire MBM. Since the IAs are spread across different elevation ranges, we tested the results using the SAFRAN reanalysis product (Vernay et al., 2019) that produces 220 gridded temperature, precipitation, wind speed, and other datasets of meteorological variables at 221 an hourly time step. This data was available as NetCDF files from 1958 for all the French massifs, 222 at every 300 m elevation belts, at 0, 20, 40° slopes, and for all eight aspects (N, NE, E, SE, S, SW, 223 W, NW). Our study is based on two different meteorological datasets described in the sub-sections 224 225 below to compare the differences.

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3.3.1 Meteorological datasets used by Guillet and Ravanel, 2020

The first part of our analysis follows a similar methodology followed by Guillet and Ravanel 228 229 (2020) in their analysis. Like their study, we used homogenized weather records from the Col du Grand Saint Bernard (GSB), located close to the MBM at 2469 m a.s.l. and provided by 230 "MeteoSwiss" and from the Aiguille du Midi (AdM) cable car station (3810 m a.s.l.). GSB 231 232 represents a similar climatological regime as the MBM, and the weather records were available 233 for an extended period starting from the 1860s. Such long-term weather records were unavailable from any weather station in the MBM. Since all IAs are located at elevations above the elevation 234 235 of the GSB weather station, it was necessary to transform the weather records to an elevation closer to the average elevation range of the IAs. For this reason, it was necessary to transform the data 236 237 from the GSB station using the weather records from the AdM weather station (data available since 238 2007). Guillet and Ravanel (2020) found a strong correlation between the monthly averaged AdM and GSB temperature records and were able to transform the GSB temperatures using a linear 239 model: 240

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$$T_{AdM} = \alpha T_{GSBi} + \beta + r_i, \tag{1}$$

where
$$\alpha = 0.87$$
 (slope), $\beta = -7.7^{\circ}$ C (intercept) and r (residuals) with zero mean.

No transformation for the precipitation values was performed as this relation is tough to establish and not always linear (Smith, 2008). Hence, the original GSB precipitation values were used for the analysis. Using these weather records, Guillet and Ravanel (2020) found a robust correlation between ablation and accumulation proxies and the surface area change of 6 IAs. We used the same datasets to test for similar potential relationships for *c*. 200 IAs, and the results are shown in Sect. 5.3.

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3.3.2 Meteorological datasets used in this study for comparison

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Since the study from Guillet and Ravanel (2020) involved a small number of IAs, the disparity 254 arising from elevation differences of IAs (in turn, the temperature and precipitation coming from 255 256 weather stations at a fixed elevation) could have been minimized or not well represented. We decided to use the SAFRAN reanalysis product and checked for similar potential relationships of 257 258 climate variables with the surface area change of IAs. The first problem we encountered was that the SAFRAN data starts from 1958, while our first images date from 1952. Therefore, for 259 260 comparison, it was essential to interpolate the missing data for the six years before 1958 (Fig. 4). Like the previous methodology, we looked for a linear relationship between the SAFRAN 261 262 temperature data (at 2400 m a.s.l. elevation belt) and the GSB temperature data. We again found 263 a strong correlation between the two datasets (Fig. 5) which helped us transform the data using:

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$$T_{SAFRAN2400} = \alpha T_{GSB i} + \beta + r_i, \qquad (2)$$

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where $\alpha = 1.01$ (slope), $\beta = -1.35^{\circ}$ C (intercept) and r (residuals) with zero mean.

For the SAFRAN data estimated (2400 m a.s.l.) from 1952, we extrapolated the data for all elevation bands. We used a standard gradient of -0.53° C/100 m increase of elevation based on the observations of Magnin et al. (2015) for the MBM.

As previously stated, a similar relationship for precipitation was tough to establish. Hence, for the analysis, we used the SAFRAN precipitation data from 1958 and extrapolated the precipitation values from the GSB weather station to all elevation bands of SAFRAN data before 1958 (six years up to 1952). However, taking a cue from the previous study of Guillet and Ravanel (2020), we expect this impact to be insignificant when considering the results over seven decades.

- 276 **4 Methods**
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4.1 Mapping the surface area of IAs from high-resolution satellite images

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IAs boundaries were manually delineated/digitized by the first author of this paper to maintain 280 281 data consistency in a GIS environment for 1952, 2001, 2012 2019. The problem of seasonal snow, which can lead to an overestimation of surface areas, was avoided using images at the end of the 282 ablation period. The differentiation of IAs from other snow/ice bodies relies on the slope angle 283 (we only consider ice bodies on slopes $> 40^{\circ}$ to be IAs) and whether they are thick enough to 284 deform under their own weight and show movement, like in the case of hanging glaciers. The slope 285 mask to remove areas with slopes $< 40^{\circ}$ was built in ArcGIS 10.6 using the Pleaides DEM. Figure 286 287 6 shows the variations in the surface areas of IAs over the study period. It also highlights the importance of high-resolution images because of the small dimensions of our studied ice bodies. 288 289 However, this data is not always available in the best quality for the past periods as we could only very accurately map 200 IAs (out of the total 423 IAs reported in Kaushik et al., 2021) for all the 290 291 periods. These 200 IAs were selected carefully after a detailed visual inspection and considering 292 issues related to shadow and illumination. Since a point-based correlation analysis (with 293 meteorological and topographic parameters) requires very high accuracy and precision of mapping, any significant uncertainty would have resulted in a major bias in our correlation 294 295 estimates. To avoid this, we used only 200 of the best mapped IAs for the correlation analysis. However, for the estimation of the total area of IAs in 1952, 2001, 2012 and 2019, as described in 296 297 Sect. 5.2, we use the complete database of 423 IAs with the assumption that overall, for the entire 298 database, the uncertainty in the mapping (\pm of the surface area) cancels out eventually and becomes 299 insignificant.

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301 4.2 Generation of topo-climatic parameters

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The relative area loss of IAs for three time periods, *i.e.* 1952 to 2001, 2001 to 2012 and 2012 to 2019, is analyzed with all topographic factors. The area loss is expressed as a relative percentage of the area lost between the first observation and the next. Authors like Salerno et al. (2017) have also used absolute values, but for our study, this would not give a fair estimation for the analysis as it generates a bias based on the size of IAs. The factors we considered for our analysis are
elevation, slope, aspect, curvature, TRI, direct solar radiation (all estimated in ArcGIS 10.6),
MARST, and size of the IAs. The topographic parameters are generated using the 4 m Pleaides
DEM described in Sect. 3.1.

311

312 Direct solar radiation:

Direct solar radiation (DSR) measures the potential total insolation across a landscape or at a 313 314 specific location. On a local scale, components such as topographic shading, slope, and aspect control the radiation distribution (Olson and Rupper, 2019). For estimating the DSR, the viewshed 315 algorithm was run based on a uniform sky and a fixed atmospheric transmissivity value of 1. Sabo 316 et al. (2016) showed the application of these algorithms in areas of rough topography. The total 317 318 DSR (DSR_{tot}) for a given location is calculated as the sum of the DSR (Dir $_{\theta,\alpha}$) from all the sun sectors (calculated for every sun position at 30 minutes intervals throughout the day and month for 319 320 a year):

321 322

$$DSR_{tot} = \sum DSR_{\theta,\alpha}$$
⁽³⁾

- 323 The direct solar radiation ($Dir_{\theta,\alpha}$) with a centroid at zenith angle (θ) and azimuth angle (α) is 324 calculated using the following equation:
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$$DSR_{\theta,\alpha} = S_{Const} * (\beta^{m(\theta)}) * SunDur_{\theta,\alpha} * SunGap_{\theta,\alpha} * cos(AngIn_{\theta,\alpha}),$$
(4)

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where S_{Const} is the solar constant with a value of 1367 W/m², β is the transmissivity of the atmosphere (averaged over all wavelengths) for the shortest path (in the direction of the zenith), m(θ) is the relative optical path length, measured as a proportion relative to the zenith path length, SunDur_{θ,α} is the time duration represented by the sky sector, SunGap_{θ,α} is the gap fraction for the sun map sector and AngIn_{θ,α} is the angle of incidence between the centroid of the sky sector and the axis normal to the surface.

The final map of DSR is the sum of values calculated at an hourly time step for every pixel, as per the resolution of the DEM used. The values of solar radiation are given in W/m^2 . Higher values for solar radiation indicate higher insolation, while lower values suggest low insolation. We prefer 337 DSR over the aspect for our analysis to avoid bias due to local shading on sun-exposed faces,338 considering the slope angle associated with the aspect.

339

340 Elevation:

Elevation strongly influences the meteorological conditions within the same region, significantly 341 342 altering the precipitation, temperature, and wind regime even at a local scale. Generally, higher elevations receive more precipitation and experience lower temperatures and higher wind speeds. 343 344 Hence regions at higher elevations, especially above the ELA, should favour more accumulation than ablation. However, wind-driven snow at higher elevations does not readily accumulate on 345 steep slopes. Some IAs may take advantage of the leeward conditions at lower elevations and 346 sustain for more extended periods. Similar results for large glaciers have previously been reported 347 348 by Bhambri et al. (2011) or Pandey and Venkataraman (2013).

349

350 Mean slope:

Slope angle strongly influences ice velocities of glaciers, mass flux, and the hydrology of the 351 352 mountain environments. Its influence on avalanche transport of snow over the glacier surface has been discussed previously (e.g., Oerlemans, 1989; Hoelzle et al., 2003; DeBeer and Sharp, 2009). 353 354 Numerous studies have also reported that slope is the single most crucial terrain parameter that controls glacier responses to climate change (Furbish and Andrews, 1984; Oerlemans et al., 1998; 355 356 Jiskoot et al., 2009; Scherler et al., 2011). In mountainous regions, the terrain slope strongly influences snow accumulation. On steep slopes, accumulation in the temperature range of $-5 - 0^{\circ}$ 357 358 C can accumulate on steep slopes. Slope likewise plays a key role when calculating other terrain 359 parameters and indices.

360

361 Mean annual rock surface temperature:

MARST estimates the average annual temperature of the rock surface governed mainly by the potential incoming solar radiation (PISR) and the mean annual air temperature (MAAT). The method for estimating MARST is described by Boeckli et al. (2012) and Magnin et al. (2019). The estimation is based on a multiple linear regression model with the form:

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(5)

367
$$Y = \alpha + \sum_{i=1}^{k} \theta i X^{i} + \varepsilon,$$

369 where Y is the value for MARST, α is the intercept term, $\theta i X^i$ represents the model's k variables 370 (PISR and MAAT) and their respective coefficients, and ε residual error term distributed equally 371 with the mean equal to 0 and the variance $\sigma^2 > 0$. For predicting the values of MARST in steep 372 slopes, we use the equation:

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$$MARST_{(pred)} = \alpha + PISR * b + MAAT * c, \qquad (6)$$

375

where α is the MARST_{pred} value when PISR and MAAT are equal to 0, and b and c are the respective coefficients of PISR and MAAT at measured RST positions. These coefficients were calibrated by Boeckli et al. (2012) (rock model 2) for the entire European Alps using a set of 53 MARST measurement points. The MAAT of the 1961-1990 period was used to calculate MARST, representing a steady state.

381 The values for MARST are calculated in $^{\circ}$ C and, for our study region, range from -12 to 10 $^{\circ}$ C.

382 MARST is also an important criterion to check for the very likely presence of permafrost below

- the IAs, which likely allows the formation and existence of IAs.
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385 **Topographic Ruggedness Index:**

The topographic Ruggedness Index (TRI) measures the ruggedness of the landscape. TRI was 386 calculated based on the methodology proposed by Sappington et al. (2007). It is calculated as a 3-387 dimensional dispersion of vectors (x, y, z components) normal to the grid cells considering the 388 slope and aspect of the cell. The magnitude of the resultant vector in a standardized form (vector 389 strength divided by the number of cells in the neighbourhood) measures the ruggedness of the 390 391 landscape. Higher values of TRI thus suggest a more rugged and sporadic terrain, which could 392 block the downward movement of the snow and subsequently lead to the formation of a weak layer, destabilizing the snowpack and leading to small avalanches resulting in mass wasting 393 394 (Schweizer, 2003). Since IA surfaces are smooth, the TRI values calculated at the surface of the IA are always low. Hence, we consider the TRI values by taking a buffer of 20 m around the IA 395

boundary delineated for the first observation (1952). The mean TRI value from this buffer isconsidered for our analysis.

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400 Curvature:

401 Curvature, estimated as a second derivative of the surface, defines the shape of the slope. Curvature is considered an essential factor because it can define snow accumulation or ablation rates for a 402 403 surface. Generally, two types of curvature profiles are known, plan and profile. For our analysis, we only used the profile curvature as it defines the shape of the slope in the steepest direction. 404 From a theoretical point of view, erosion processes prevail in convex (negative values) profile 405 curvature locations, while deposition is predominant in concave (positive values) profile curvature 406 407 locations. The curvature values define how strongly the slope is convex (lower negative values) or concave (higher positive values). That is why curvatures can be considered essential in the 408 409 accumulation and ablation rates of a glacier or ice body. Like TRI, the IAs tend to show flat curvature profiles if we consider their surface. Hence, we estimate the curvature values around the 410 411 same buffer as the TRI and use this for further analysis.

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414 **4.3 Proxies for ablation and accumulation**

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To eventually correlate changes in surface area of IAs with the changing climate, we use the 416 417 temperature and precipitation data from the transformed GSB weather records and SAFRAN reanalysis product (see Sect. 3.3) to build proxies for accumulation and ablation. The proxy for 418 419 ablation was built by estimating the annual sum of positive degree-days (PDD) computed from the 420 normal probability distribution centered around the mean monthly temperature. Estimation of the PDD is based on the empirical relation, which states that the melting rate is proportional to the 421 surface-air temperature excess above 0°C. Several methods for estimating PDD have been 422 423 proposed by Braithwaite and Olesen (1989), Braithwaite (1995), and Hock (2003). However, the 424 method proposed by Calov and Greve (2005) also accounts for stochastic variations in temperature 425 during the computation of PDD. The formula for the estimation of the PDD using this method is given by: 426

428

$$PDD = \int_0^A dt \left[\frac{T_{ac}^2}{\sqrt{2\pi}} \exp\left(-\frac{T_{ac}}{2\sigma^2} \right) + \frac{T_{ac}}{2} \operatorname{erfc} \left(-\frac{T_{ac}}{\sqrt{2\sigma}} \right) \right]$$
(7)

429

T_{ac} is the annual temperature cycle (monthly mean temperatures estimated in °C for the entire 430 year), σ is the standard deviation of the temperature from the annual cycle, A = 1 year, and *erfc* is 431 432 the conventional error function built-in in all programming languages.

After computing the PDD, we calculate the cumulative PDD (CPDD) by taking the sum of all the 433 annual PDD values for each observation period (*i.e.* 1952-2001, 2001-2012 and 2012-2019). This 434 value of CPDD is then used as a proxy for ablation (Braithwaite and Olesen, 1989; Vincent and 435 436 Vallon, 1997).

437 The calculation of the proxy for accumulation is more tricky because we only consider the yearly sum of precipitation occurring at a temperature between -5 and 0° C, as only snowfall within this 438 439 temperature range is believed to accumulate/adhere to steep slopes (Kuroiwa et al., 1967; Guillet 440 and Ravanel, 2020; Eidevåg et al., 2022). The temperature-dependent indicator function can be 441 written in the following form:

442

 $\chi_{i}(T, (t)) = \begin{cases} 1 \text{ if } -5^{\circ}C \leq T(t) \leq 0^{\circ}C \\ \\ 0 \text{ otherwise} \end{cases}$ 443 444

445 446

447 4.4 Surface area model

448

449 Using the proxy for ablation and accumulation, Guillet and Ravanel (2020) proposed a surface 450 area model to estimate the differences in the surface areas of IAs between different time steps due 451 to the time-integrated changes in meteorological parameters. The main goal is to look for a 452 potential linear relationship between climate variables and the changes in surface areas of IAs, 453 using a multivariate regression model. The equation for the model can be written as:

454

455
$$S_{m}(t) = S(t_{0}) - \int_{t_{0}}^{t} (\alpha 1 CPDD(t) - \chi i(T(t))\alpha 2 A(t))dt + \beta + \varepsilon(t)$$
(9)

456

(8)

where $S_m(t)$ corresponds to the modelled surface area at time t; similarly, CPDD(t) and A(t) 457 represent the proxies for ablation and accumulation; S(t=0) is the first available measurement; α 458 459 and $\alpha 2$ are the coefficients of linear regression, β is the intercept, and ϵ the residual. $\chi(T, t)$ accounts for precipitation occurring in the [-5°C, 0°C] temperature range and is given by the temperature-460 dependent indicator function given in equation (8). The area of IAs at each time step was calculated 461 462 using the surface area model (with the temperature and precipitation proxies), and we hereafter refer to this area as modelled area. The measured area is the surface area we delineated using high-463 464 resolution optical images.

465

466 **4.5 Uncertainty estimations**

467

468 Since this study uses data from different sources and periods, uncertainties of different origins might have been introduced to delineate the IA boundaries. A good estimation of these 469 470 uncertainties is thus crucial to have a fair estimation of the significance of the results (Racoviteanu et al., 2008; Shukla and Qadir, 2016; Garg et al., 2017). Some sources of uncertainty in this study 471 472 could arise from (1) errors inherent to the aerial images and satellite-derived datasets, (2) errors resulting from inaccurate co-registration of data from various sources, (3) errors produced while 473 474 generating the high-resolution DEM from stereo images, and (4) conceptual errors linked with 475 defining the boundaries of IAs in all images. Quantifying the errors inherent in processing all 476 datasets used is challenging, and this is out of the scope of this paper. A detailed accuracy assessment of the DEM generation and co-registration process is provided in Sect. 5.1 and 3.1, 477 478 respectively. Quantifying errors resulting from the manual delineation of IA boundary is also challenging, but we have previous guidelines from Paul et al. (2017) for the quality and consistency 479 480 assessment of manual delineations.

One way to assess the area uncertainty is to perform multiple digitizations of the same surface and calculate the mean area deviation (MAD), taking the first digitization as a reference (Meier et al., 2018). Considering this, the first author performed three digitizations for 50 IAs on images from 1952, 2001, 2012 and 2019, considering different challenges associated with aerial and satellite images like shadow and illumination. MAD provides a percentage estimate of how the final area calculated varies across multiple digitizations for each polygon. MAD values are affected by the size of the polygon manually digitized. Previously, authors like Paul et al. (2013), Fischer et al. (2014) and Pfeffer et al. (2014) have reported an increase in the uncertainty of manual digitizations
with a decrease in the size of the polygons. With this in mind, we also digitized IAs of different
sizes ranging from 0.001 km² to 0.01 km².

- **5 Results**

5.1 Accuracy of the DEM

Figure 7a shows the stable surfaces (after eliminating glacier boundaries, trees, and forests) we used for our co-registration process, and Figure 7b displays the difference in elevation between the reference DEM and the source DEM before co-registration. Figure 7c presents the results after the co-registration process considering all the surfaces (stable and non-stable), and Figure 7d shows the difference considering only the stable areas after masking out non-stable areas using the glacier boundaries provided by the Randolf Glacier Inventory (RGI v6.0) (Consortium, 2017). The source DEM was translated using the corresponding shift values x = -5.03 m, y = 6.00 m, and z = 3.22 m The distribution of errors can be visualized by a histogram of the sampled errors, where the number of errors (frequency) within certain predefined intervals is plotted (Höhle and Höhle, 2009). Figure 8 shows the histogram of the errors Δh (elevation difference between the reference and source DEM) in meters for the stable areas. The accuracy estimates before and after the co-registration are shown by the normalized median absolute deviation (nmad) and the median value calculated together. As can be seen, the nmad and median values before the co-registration process for stable areas were 5.16 and -5.06, respectively. After the co-registration process, the value dropped to 1.98 for the nmad and -0.14 for the median value. This suggests a good correlation between the high-resolution LiDAR DEM used as a reference and the Pleaides DEM we built.

519 5.2 Total area loss of ice aprons in the Mont-Blanc massif over seven decades

520

521 The total area of IAs mapped in 1952 was 7.932 km². It dropped to 5.915 km² in 2001. The surface area dropped to 4.919 km² in 2012 and then to 4.21 km² in 2019 (Fig. 9). This implies that from 522 1952 to 2019, IAs lost ~47 % of their original area. It corresponds to an average surface area loss 523 of 0.78 km² per decade. However, the percentage area loss from 1952 to 2001 was ~25 % compared 524 to ~29 % relative area loss from 2001 to 2019. This is an alarming rate: IAs have lost more relative 525 area during the 18 recent years (with an average area loss of 1.15 % per year) compared to the 50 526 527 years before 2001 (0.5 % per year average area loss). Figure 10 shows the MAD values for 50 IAs in 1952, 2001, 2012 and 2019. We did not observe 528 an increase in MAD values with decreasing size of the IAs, mainly because the number of samples 529

we used is comparatively less than that in the previous studies. Overall, the mean MAD observed for all years was \pm 6.4 %. The MAD for the IAs digitized on the orthophotos from 1952 was \pm 6.68 %, while for 2001, it was \pm 7.2 %. The MAD for 2012 and 2019 was \pm 6.32 % and \pm 5.50 %, respectively.

534

535 **5.3 Influence of changing climate on the area loss**

536

537 Figure 11 shows the trend of PDD increase over the years in the MBM. All elevations, except 4800 538 m a.s.l., show an increasing trend of PDD values from 1952 to 2019. Specifically, for the year 1952, since we have only one year for the longer-term analysis, it is interesting to look in detail at 539 540 the climatic conditions prevailing in the region around this period. Looking at the GSB records, the average temperature in the region during the past ten years before 1952 was -0.987 °C, with 541 542 average summer temperatures (July, August and September) being 6.783 °C. 1947 was particularly hot, with the annual average temperature recorded at -0.275 °C and average summer temperatures 543 at 8.266 °C. The next thirty years after 1952 were more favourable, with average annual 544 temperatures at -1.523 °C and average summer temperatures at 5.256 °C (GSB data provided as 545 supplementary material 1). Since 1952 was coming at the back of considerably warmer years, a 546 547 significant reduction in the surface area of IAs can be expected during this period. Looking at the weather records, the conditions after 1952 for the next thirty years were more favourable. 548

Similarly, Figure 12 shows the variations in the accumulation rates (average annual accumulation 549 550 per period) for all elevation bands. The results show only that part of the snowfall which is 551 expected to accumulate on the steep slopes. Except for the highest elevation band, i.e. 4800 m a.s.l., accumulation rates at all elevation bands show a general decreasing trend. For example, at 552 the 3900 m a.s.l. elevation band, accumulation rates fell from 32mm/year from 1952 -2001 to 28 553 554 mm/year from 2001 to 2012, and further to 18 mm/year between 2012 - 2019. This shows that temperatures in the MBM are increasing, while on the other hand, accumulation on steep slopes is 555 556 decreasing over time. Figure 13 shows the annual variation of the accumulation on steep slopes 557 at different elevations. The first observation from this trend shows that very little precipitation accumulates on steep slopes in the winter months, while accumulation occurs almost entirely in 558 559 the summer months. Further, the accumulation is more significant at higher elevations (4200 -560 4500 m a.s.l.) in the summer months than at lower elevations. At lower elevations, accumulation is predominantly observed in pre-summer (May) and post-summer (October) months. 561

Figure 14a presents the correlation between the ratio of the mean measured surface area at time t, S(t), to the initial area, $S(t_0)$, with the ratio of the mean modelled surface area using the GSB transformed data to the initial area for 2001 and 1952. We consider the ratio of $S(t)/S(t_0)$ as an indicator to estimate the area loss between the two time periods. A high ratio value (*i.e.* value close to 1) in the present context indicates that the relative surface area loss of IAs between the two periods is comparatively less than that of IAs whose ratio is closer to 0. A value larger than 1 indicates an increase in the surface area over time.

From the results, we do not see a strong correlation (r = 0.73) between the modelled area (from GSB transformed climate data) and the measured area for the 200 IAs spread across the MBM (Fig. 14 a). However, the correlation improves significantly (r = 0.86) when we use the SAFRAN data based on different elevations and remodel the surface area for each IA (Fig. 14 b). This can be seen from the values of R², Pearson's r, RMSE and the p-value estimates from the T-test achieved from both datasets (Table 2). The best-fitting line presents a slope of 1.0 and an intercept of 0.0.

Both figures show that IAs at lower elevations (blue to green colour and small tick size) generally
show lower ratios values than IAs at higher elevations (yellowish colours and bigger tick size).
This implies that the elevation of the IAs potentially plays a crucial role in their response to the
changing climate. Overall, the surface area of IAs decreased throughout the massif from 1952 to

2001 except for 4 IAs, which showed an increase in surface area. These 4 IAs are: 2 IAs on the N 580 and NW face of Rochers Rouges Inferieurs (~4350 m a.s.l. and 4050 m a.s.l.) near the Grand 581 582 Plateau, 1 IA on the NE face of Col de la Brenva (~4160 m a.s.l.) and 1 IA on the S face of Col du Bionnassay (~4050 m a.s.l.). As observed, all these IAs are located at elevations higher than 4000 583 m a.s.l. As an exception, it can be expected that a few IAs could show an increase in the surface 584 585 area. However, this increase is not dramatic (~10 % increase). The results, however, reaffirm the proficiency of the proposed surface area model in predicting new IA states from the accumulation 586 587 and ablation proxies. Similar results were observed for the other two time periods, *i.e.* 2001-2012 and 2012-2019, as seen in Table 2. 588

589 590

591 **5.4 Influence of the local topography and other factors on the area loss of IAs**

592

Each parameter, as described in Sect. 4.2 was individually regressed with the relative area loss of IAs for the three periods, and their influence was assessed by the coefficient of determination (R^2) and Pearson's r-value.

A joint analysis of the surface area lost by the IAs and the direct solar radiation reveals a strong 596 597 correlation between the values of DSR and the relative surface area loss of IAs for all the threetime periods (1952-2001, 2001-2012 and 2012-2019) (Fig. 15a; Table 3). However, this is the first 598 599 evidence of the potential negative impact of solar radiation on small ice bodies like IAs. Previous analysis of Guillet and Ravanel, 2020 with the climate variables indicated a potential relationship 600 between the elevation and the surface area loss of IAs. This is somewhat statistically significant 601 602 from the regression analysis, as we found a negative correlation between the surface area loss and 603 the mean elevation of the IAs (Fig. 15b; Table 3). A further comparison of the IAs (200 IAs) 604 distribution with elevation and aspect shows that most IAs (~ 77 % of the total number) exist at elevations above 3200 m a.s.l. Further, most IAs (~ 56.5 %) exist in the northern aspects (N, NW, 605 606 NE), while the E and W aspects are the least favoured (supplementary material 2). In addition, we 607 found a moderate positive correlation between the average MARST values and the surface area 608 loss of IAs. The correlation observed was not very significant compared to the previous two factors. It indicates that the effect of rock surface temperatures on the area loss of IAs is not strong 609 on a regional scale. (Fig. 15c; Table 3). However, this relationship needs to be examined in a more 610

611 site-specific and localized area to understand better its impact on the surface area loss of IAs. We
612 also observed that the correlation was higher for a more extensive observation period (1952-2001)
613 than for shorter periods. This could suggest that the influence of rock surface temperatures
614 potentially becomes more prominent with a more extensive observation period.

A similar analysis of IAs area loss with the TRI showed a weak positive correlation (Fig. 15d; 615 616 Table 3). An increase in TRI values (*i.e.* increase in terrain ruggedness) may result in more ice area loss on a site-specific scale, but this relationship is hard to observe globally. Like the results 617 618 from the analysis with MARST, the strongest correlation was again observed for the largest study period. Further, like the TRI, we also found a weak correlation between the terrain slope and 619 curvature with the surface area loss of IAs. We must note that our criteria for selecting IAs already 620 limit us to areas with slope angles steeper than 40° (Fig.15e; Table 3). Hence it was difficult to 621 622 observe any significant impact of terrain slope on the rate of area loss of IAs. Similarly, terrain curvature seems to have the most negligible impact (Fig. 15f; Table 3). As cited in Sect. 4.2. 623 624 previous studies may have shown that terrain curvatures could play an essential role in the erosion and accumulation dynamics on steep slopes, but this is not the case for IAs in the MBM. We 625 626 performed the last comparison between the relative surface area loss of IAs with their initial area. 627 Our results were similar to the one Lopez et al. (2010) reported, as we did not find any correlation 628 between the two quantities (Fig. 15g; Table 3).

629

630

631 **6. Discussion:**

632

633 6.1 Area loss of ice aprons and the role of the changing climate

634 As observed from the results in Sect. 5.2, IAs have been losing surface area at an alarming rate. 635 This rate of surface area loss is disconcerting because, compared to the glaciers in the MBM, the IAs are losing their area at a higher rate (~24 % for glaciers from the end of LIA till 2008, according 636 to Gardent et al., 2014, while IAs have lost ~47% surface area in the last 70 years). The small size 637 of IAs makes them more vulnerable to global warming than large glaciers. An increase in average 638 639 annual temperatures and a decrease in precipitation rates put the existing IAs at risk of losing their mass entirely before the end of this century. In addition, considering that the effects of local 640 topography are also more pronounced in the case of IAs than for large glaciers, continuous 641

monitoring of these critical ice bodies has become imperative. Results discussed in Sect. 5.3 642 indicated the strong influence of temperature and precipitation on the surface area changes of IAs. 643 644 The results raise further questions regarding the sensitivity of the IAs to extreme weather events, like the heatwaves experienced in the study region. Unfortunately, our sampling rate does not 645 allow us to quantify the effects of individual extreme weather events. Nevertheless, there is a 646 647 strong argument in favour that these events, especially in the past two decades, cause the IAs to lose mass more rapidly than in the previous decades. Further, the heatwaves occurring during 648 649 winter and midsummer, when the IA surfaces are free of snow, will have the worst adverse effect. 650 As suggested by Meehl and Tebaldi (2004), with an increase in the intensity and frequency of extreme events in the coming decades, understanding the effects of climate variables on the 651 sensitivity of IAs is even more critical. 652

Further, several authors have previously also accounted for the variations in solar radiation in mass-balance modelling studies (Huss et al., 2009; Thibert et al., 2018). Our results showed a strong correlation of DSR with area change, making this argument stronger. However, since the focus was to show the impact of climate variables separately, we preferred a temperature-index model as the first approach. However, we expect solar radiation to play a significant role in the sensitivity of ice aprons, and future studies on ice apron evolutions in the 21st century should address this question.

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661

662 6.2. Area loss of ice aprons and the role of topographic parameters

663

Since ice/glacier bodies within the same climate regime can also respond to climate change 664 665 differently, the last part of the analysis (Sect. 5.4) was dedicated to understanding the linkages of 666 local topographic factors to the surface area loss of IAs. As reported previously by Salerno et al. (2017), some local topographic factors influence the response of IAs to climate change more 667 668 significantly than others. A first analysis showed that IAs that receive more solar radiation from the sun throughout the year lose their surface area faster than those that receive less DSR. Similar 669 670 results for other regions in mountain environments have also been reported previously by 671 Oerlemans and Klok (2002); Mölg (2004); Johnson and Rupper (2020). Incoming solar radiation

is an essential component of all surface energy and mass balance models. But the significance of 672 DSR on the surface area loss of small ice bodies like IAs is reported for the first time in our study. 673 674 Further, the correlation between elevation and surface area loss of IAs was the second most significant of all topographic factors. IAs located at lower elevations are potentially subject to 675 more intense degradation and lose their surface area faster than those at higher elevations. On a 676 677 more local scale, other topographic factors could also play a critical role in the surface area variations of IAs. However, elevation seems to be a dominant causative factor on a regional scale. 678 679 Elevation strongly influences meteorological conditions (temperature, precipitation, and wind 680 speeds) and permafrost; this likely strongly influences the durability of IAs in the context of changing climate. Hantel et al. (2012) suggested that the median summer snowline for the Alps to 681 be at 3083 ± 121 m a.s.l. (1961 – 2010), while Rabatel et al. (2013) documented the regional ELA 682 683 at 3035 ± 120 m a.s.l. (1984 – 2010). Rabatel et al., 2013 further described the rising of the ELA to 3250 ± 135 m a.s.l. during the 2003 heatwave. Subsequent heatwaves of 2006, 2015 and 2019 684 would have likely resulted in similar scenarios (Hoy et al., 2017). Since ~77 % of the total IAs 685 reported in this study exist at elevations above 3200 m a.s.l., the rising of the ELA in future climate 686 687 scenarios risks more IAs towards faster degradation and disappearance. An example of this is the case of the IA on the north face of Aiguille des Grands Charmoz (3445 m a.s.l.), which completely 688 689 disappeared during the 2017 summer heatwave (Guillet and Ravanel, 2020).

690 In addition, on a local scale, some correlation between the rock surface temperatures and the area 691 loss of IAs was observed from the analysis. Guillet et al. (2021) suggested that IAs are cold ice bodies that exist predominantly on permafrost-affected rock walls. They further reported 692 693 temperatures <0°C at the base of the ice core taken from the IA on the north face of Triangle du Tacul (3970 m a.s.l.). Heating from rock surfaces is predominantly the cause of permafrost 694 695 degradation, which further affects mountain slope stability leading to an increased rock mass 696 wasting (Magnin et al., 2017). Cold surfaces demonstrate more ice cohesion with the underlying surfaces, while a rise in surface temperatures decreases basal cohesion, increasing the sliding 697 process and leading to more ice flow (Deline et al., 2015). Thus, it is likely that underlying 698 699 permafrost conditions aid the sustainability of IAs in the long term, and an increase in rock surface 700 temperatures around IAs could result in IAs losing mass more rapidly.

Kaushik et al. (2021) further showed that most IAs exist in extremely rugged terrains: 51 % of the

total IAs mapped exist in the TRI's high and very high ruggedness class, while only 8 % exist in

the low ruggedness. Thus, comparing the terrain ruggedness to the area loss of IAs makes sense since the topography around the snow/ice bodies can critically influence their stability (Deline et al., 2015). Increasing terrain ruggedness is associated with slope instability and further ice volume loss. However, for our study, this relation was not very pronounced, showing that the topography's ruggedness does not substantially affect the area loss of IAs.

Previous analysis by Kaushik et al., 2021 also showed that most IAs in the MBM (83 %) lie at 708 709 mean slopes between 40° and 65°. Increasing slope steepness limits accumulation, while 710 avalanches further scour away snow from the surface of the IA, thus exposing the ice directly to the sun and wind (Vionnet et al., 2012). However, the differences in slope angles of the IAs were 711 not a dominant factor affecting the rates of area loss. A plausible explanation for this could be that 712 since we limit the slope criteria to $>40^{\circ}$ and most IAs lie in the range of 40 to 65 ° slope, the effect 713 of terrain slope is not as well pronounced as it would be between low ($< 15^{\circ}$) and extreme slopes 714 (>65°). Similar results were observed by Li et al. (2011), as they observed very slight variations 715 in area loss for small glaciers with differences in slope. They suggested other local topographic 716 717 factors could mitigate the effects of slope in case of small ice/snow bodies.

Similarly, terrain curvature also has a negligible effect on the surface area loss of IAs. As suggested by Alkhasawneh et al., 2013; Yanuarsyah and Khairiah, 2017, convex profile curvature favours the erosion processes, while in locations with concave curvature, the deposition process can be predominant. Over time, the terrain curvature can be a dominant factor in the dynamics of glacier/ice bodies, but this relation was not established for our study.

At last, a comparison of the rate of surface area loss of IAs with the original size of the IAs was 723 724 performed, and we observed no correlation between the two factors. Although previous studies by 725 Paul et al. (2004), Jiskoot et al. (2009), and Garg et al. (2017) have shown the correlation between 726 the size of the ice/glacier bodies with the area loss, this is not evident in our case. Unlike previous studies, which considered different glaciers ranging in size from less than a km² to several hundred 727 km², IAs are small ice bodies (0.0005 km² to 0.1 km²). Hence, it is plausible that the effect of IA 728 size on area loss rate is not pronounced in our case. Similar results were shown by Lopez et al. 729 730 (2010), who analyzed 72 glaciers in South America, and reported no correlation between the 731 glacier length and the area loss of glaciers.

Another critical factor to consider, along with the impact of topography, is the role of avalanches
in the erosion and deposition processes on the IA surface. Analysis of the ice core from the N- face

of Triangle du Tacul showed that IAs are almost immobile cold ice bodies (Guillet et al., 2021).
Hence IAs do not directly participate in feeding the larger glacier systems below them. However,
the avalanches triggered above can bring fresh snow/debris and lead to erosion or deposition on
the IA surface. We expect this factor to also play a role in the area change dynamics of the IA,
which we have not considered as part of this study. Although this impact on the scale of an IA is
tough to determine, future studies should focus on ascertaining this impact at least on a local sitespecific scale.

741

742 **7. Conclusions**

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This study makes the first attempt to understand the dynamics of IAs concerning the changing 744 745 climate and topographic factors at a regional scale. IAs are very small ice features but highly critical components of the mountain cryosphere. Because of the difficulties associated with their 746 monitoring and relative unimportance to mountain hydrology, no studies solely based on their 747 evolution on a regional scale have been performed before. This paper presented an analysis of 200 748 749 IAs spread throughout the MBM and existing in different topographic settings to understand their dynamics in the context of climate warming. For this purpose, we accurately mapped the IAs on 750 751 very high-resolution aerial and satellite images available for 1952, 2001, 2012 and 2019. Using our extensive database of IAs, we compared the total area variation of IAs for three periods. 752 753 Further, we also attempted to establish a relationship between the surface area lost by IAs with 754 meteorological parameters (*i.e.* temperature and precipitation) and their associated topographic 755 parameters.

756

757 Some important outcomes from the study are:

Over the study period 1952-2019, IAs have lost their surface area at a very alarming rate.
 The total area of IAs in MBM was 7.93 km² in 1952, dropping to 5.91 km² in 2001, 4.91 km² in 2012, and 4.21 km² in 2019 (~ 47% drop in total surface area in less than three quarters of a century).

The observed rate of relative area loss in the last 18 years (~29 %) is more than the overall area loss during the 48 previous years (1952-2001; ~ 26 %).

- Results from the analysis of IA's surface area loss and meteorological parameters (i.e.
 temperature and precipitation) conclusively proved the strong impact of these parameters
 on the behaviour of small ice bodies like IAs.
- Further analysis of IAs surface area loss with different topographic parameters showed a strong correlation of IAs surface area loss with the DSR and elevation. Other factors like MARST, TRI, and mean terrain slope could also play an important role locally, but their effect is not significant regionally. Terrain curvature and the size of the IAs were not found to impact the IA's surface area loss significantly.
- 772

Looking at the melting rate of IAs and the future predictions of global climate change, it is evident that these small and critical ice bodies are most vulnerable to adverse impacts. It is hard to imagine any of the IAs surviving the next few decades with increasing temperatures at the present and future melting rates. The loss of IAs will thus be the loss of crucial glacial heritages and playgrounds for the iconic practice of mountaineering. Hopefully, this study forms a basis to encourage further studies on IAs.

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Author contributions: SK designed the study and drafted the paper, which all co-authors revised.
LR and FM helped in data interpretation and analysis. YY and ET proofread the manuscript and
provided valuable inputs for improving the overall quality of the paper. DC processed and provided
the DEM used for the study.

785

786 **Data availability:** The ice apron inventory will be made available on demand.

787

788 **Competing Interests:** The authors declare that they have no conflict of interest.

789

790 Acknowledgements:

791 This research is part of the USMB *Couv2Glas* and *GPClim* projects. Pleiades data were acquired

792 within the CNES *Kalideos-Alpes* project and successfully processed under the program "*Emerging*"

risks related to the 'dark side' Alpine cryosphere". We also thank C. Vincent of the Institut des

- 794 Géosciences de l'Environnement (IGE) for providing the LiDAR DEM of the Argentière glacier
- 795 area.
- 796
- 797

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1084 Tables:

Data type	Source	Resolution (m/time)	Acquisition time/pe- riod
	Orthoimages IGN	0.2	July 2015
	Pleiades 1A PAN	0.5	25/08/2019,
			19/08/2012
Optical	Sentinel 2	10	12/09/2019
	SPOT 6	2.2	14/09/2019
	Pleiades 1A XS	2	19/08/2012
	Orthoimages IGN	0.5	July 2001
	Orthoimages IGN	0.5	1952
Meteorological	Col du Grand-Saint Bernard weather station (2469 m a.s.l.)	daily	1952 - 2019
	Aiguille du Midi weather sta- tion (3840 m a.s.l.)	daily	2007-2018
	SAFRAN reanalysis	daily	1958 - 2019

Table 1: Datasets used for the study.

Time pe- riod	Slope	Intercept	R ²	Pearson's r	RMSE (km ²)	p value
1952 - 2001	0.79	0.12	0.53	0.73	0.010	< 0.001
2001 - 2012	0.70	0.26	0.56	0.75	0.102	< 0.001
2012 - 2019	0.89	0.04	0.63	0.80	0.097	< 0.001
			(a)	·		

Time pe- riod	Slope	Intercept	R ²	Pearson's r	RMSE (km ²)	p value
1952 - 2001	1.01	-0.04	0.73	0.86	0.075	< 0.001

2001 - 2012	0.74	0.22	0.67	0.82	0.086	< 0.001	
2012 - 2019	1.37	-0.39	0.83	0.91	0.071	< 0.001	
(b)							

1090	Table 2. I	inear regression	parameters and	correlation	metrics fo	r each i	period (a)	using	GSB
1030	1 auto 2. L	mear regression	parameters and	conclation	metrics 10	1 Cach	periou (a)	using	ODD

1091 transformed data and (b) using the SAFRAN reanalysis product.

Variable	Time-period	R ²	Pearson's r
Direct Solar	1952 - 2001	0.64	0.79
Kadiation	2001 - 2012	0.67	0.81
	2012 - 2019	0.51	0.72
	1952 – 2001	0.61	-0.78
	2001 - 2012	0.57	-0.75
Elevation	2012 - 2019	0.51	-0.71
	1952 - 2001	0.40	0.63
	2001 - 2012	0.34	0.58
MARST	2012 - 2019	0.27	0.52
	1952 - 2001	0.37	0.60
TRI	2001 - 2012	0.30	0.55
	2012 - 2019	0.32	0.57
	1952 - 2001	0.29	0.54
Slope	2001 - 2012	0.25	0.50
	2012 - 2019	0.21	0.46
	1952 - 2001	0.06	-0.26
Curvature	2001 - 2012	0.03	-0.18
	2012 - 2019	0.06	-0.24
	1952 - 2001	0.04	-0.22
Size of IA	2001 - 2012	0.06	-0.26
	2012 - 2019	0.04	-0.22

- 1093
- 1094 Table 3: Linear regression parameters and correlation metrics for each studied parameter.
- 1095
- 1096 Figures:



Figure 1: The Mont-Blanc massif (Western European Alps). 200 IAs (red stars) were digitized
accurately on high-resolution images. The glacier outlines (in blue) comes from Gardent et al.,

- 1100 2014. The green line shows the border between France, Italy and Switzerland.
- 1101



Figure 2: Ice aprons and their locations in the MBM, a. IAs on the N face of Grandes Jorasses(4208 m a.s.l.) and b. IAs on the headwall of the Argentière glacier separated by a bergschrund

1106 (3280 m a.s.l.).



- 1110 Figure 3: (a) The reference LiDAR DEM of the Argentière glacier used for co-registration, (b)
- 1111 the source Pleiades DEM used for further analysis.





1114 Figure 4: SAFRAN reanalysis product temperature time-series from 1952-2019 for different

elevations in the MBM. The figure shows the variation of the mean annual temperatures for the entire study period.

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1121 Grand Saint Bernard (GSB) and the SAFRAN reanalysis data at 2400 m a.s.l.



1124 Figure 6: IAs extent delineated on high-resolution images: (a) orthophotos 1952, (b) orthophotos

- 1125 2001, (c) Pleaides panchromatic 2012, (d) Pleaides panchromatic 2019. The different colour
- 1126 polygons represent the surface area for each date.





1129 Figure 7: Stepwise Pleiades DEM accuracy assessment (a) the surfaces used for coregistration

1130 (b) elevation difference before coregistration (c) elevation difference after coregistration

1131 considering all areas (d) elevation difference after coregistration considering only the stable1132 areas.



Figure 8: DEM Error (elevation difference between the reference and source DEM) distributionfor stable areas









Figure 10: The distribution of MAD values based on multiple digitizations of the IAs area for allperiods.



Figure 11: The variation of annual PDD values estimated based on monthly mean temperaturesat different elevations in the MBM from 1952 to 2019.



Figure 12: Variation of the average accumulation rates on steep slopes at different elevations foreach period of observation



1154 Figure 13: Accumulation (steep slopes) trends for the year 2019 at different elevations



Figure 14: Correlation between the mean normalized measured and modelled surface areas (a)
with GSB data transformed to AdM data and (b) with SAFRAN reanalysis data at time t. The
colour and size of the ticks represent the mean elevation of the IA.



Figure 15: Scatter plots showing relationships between topographic factors and the area loss of IAs from 1952 to 2019. a) Direct solar radiation, b) elevation, c) MARST, d) TRI, e) slope, f) curvature and g) size of the IAs.