

1 Gravity Inversion Method Using L_0 -norm Constraint
2 with Auto-adaptive Regularization and Combined
3 Stopping Criteria.

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Abstract

We present a gravity inversion method that can produce compact and sharp images, to assist the modeling of non-smooth geologic features. The proposed iterative inversion approach makes use of L_0 -norm stabilizing functional, hard and physical parameter inequality constraints, and depth weighting function. The method incorporates an auto-adaptive regularization technique, which automatically determines a suitable regularization parameter and error weighting function that helps to improve both the stability and convergence of the method. The auto-adaptive regularization and error weighting matrix are not dependent on the known noise level. Because of that, the method yields reasonable results even if the noise level of the data is not known properly. The utilization of an effectively combined stopping rule to terminate the inversion process is another improvement that is introduced in this work. The capacity and the efficiency of the new inversion method were tested by inverting randomly chosen synthetic and measured data. The synthetic test models consist of multiple causative blocky bodies, with different geometries and density distributions that are vertically and horizontally distributed adjacent to each other. Inversion results of the synthetic data show that the developed method can recover models that adequately match the real geometry, location, and densities of the synthetic causative bodies. Furthermore, the testing of the improved approach using published real gravity data confirmed the potential and practicality of the method in producing compact and sharp inverse images of the subsurface.

Keywords— Gravity data, Iterative inversion, L_0 -norm constraint, Auto-adaptive regularization, Stopping criteria, Compact image.

1 Introduction

Gravity measurements have been used in a wide range of geophysical prospecting and investigations, such as in mineral explorations, engineering and environmental problems as well as archeological site investigations (Hinze et al., 2013, p. 20). In general, gravity inversion is a process that is used to determine the density, size, shape, and location of complex subsurface causative bodies from an observed gravity anomaly, by using different mathematical modeling techniques. Thus, inversion of gravity data constitutes an important step in the quantitative interpretation since the reconstruction of density contrast models markedly increases the amount of information that can be extracted from the gravity data.

However, a principal difficulty with the gravity data inversion is the inherent non-uniqueness and instability that also exists in any geophysical method (Al-Chalabi, 1971; Blakely, 1996, p. 216). In other words, for the given observed gravity data there are many equivalent density distributions that can reproduce the same field data. The standard approach used to select acceptable solutions, that are geologically reasonable, is to use additional information about the problem by making assumptions on the following aspects: (1) about the model parameters (existing information on the subsurface structure from geological or other geophysical hindsight) and (2) about the data parameters (statistical properties of the inexact data, e.g. Gaussian distribution of errors). Based on these assumptions there are two approaches in gravity inversion: The first approach fixes the density and vary the geometry. This approach is nonlinear in nature and has been studied by many authors, for instance, Lelievre et al. (2015); Camacho et al. (2002) and Camacho et al. (2011). The second approach, which also is the one used in this work, fixes the geometry and vary the density. This approach is linear in nature and has been investigated by many researchers (Li & Oldenburg, 1998; Boulanger & Chouteau, 2001).

In an effort to introduce more qualitative prior information, Last & Kubik (1983) in particular, developed a method called compact gravity inversion. Their strategy utilizes the compactness stabilizer to minimize the area (in 2D) or volume (in 3D) occupied by the causative body, which is equivalent to maximizing its compactness. Barbosa & Silva (1994) generalized the compact inversion method by making use of compactness along several axes using Tikhonov's regularization. In 2006 Silva and Barbosa further developed the Compact inversion method with the so-called 'interactive inversion' which estimates the location and geometry of several density anomalies. They simplified their old method (Barbosa & Silva, 1994) to improve computational performance. The generalized compact and interactive inversion strongly need a priori information to yield an accurate estimation.

The compactness stabilizer (Last & Kubik, 1983) also known as the minimum support stabilizer (Port-

niaguine & Zhdanov, 1999) has been borrowed and implemented by other researchers in various geophysical inversion methods (Ajo-Franklin et al., 2007; Stocco et al., 2009; Fei et al., 2018; Feng et al., 2020; Varfinezhad et al., 2020). As it was demonstrated by a number of researchers (Zhdanov & Tolstaya, 2004; Rezaie et al., 2017; Feng et al., 2020; Varfinezhad et al., 2022), this stabilizer is known to yield a compact or focused geophysical model with sharp boundaries. Apart from the inversion methods which produce focused images mentioned above, sparse geophysical inversion approaches derived from L_p -norm ($0 \leq p \leq 1$) stabilization have been developed by many researchers. For instance, sparse seismic reflectivity inversion method (Li et al., 2017), direct current resistivity data inversion algorithm (Singh et al., 2018), magnetic data sparse inversion method (Li et al., 2018; Fournier et al., 2020), sparse gravity data inversion technique (Vatankhah et al., 2017; Peng & Liu, 2021), to mention only a few. Some instability of the original compact gravity inversion algorithm of Last & Kubik (1983) was reported by Lewi (1997, p. 87) when the data is contaminated with noise. Then Lewi (1997, p. 89) has improved the original compact inversion by introducing a new approach to the 3D compact gravity inversion. The problem with Lewi (1997, p. 89) method arises when dealing with a multiple-source model, where the inversion algorithm tends to concentrate densities towards the surface regardless of the true depth of the causative bodies. In overcoming this drawback, Gebre & Lewi (2022) improved the compact gravity inversion method by incorporating a new depth weighting function. In this paper, we present a gravity inversion method that can produce compact and sharp images, to assist the modeling of non-smooth, blocky geologic features with sharp boundaries. The proposed approach is based on the authors' previous work (Gebre & Lewi, 2022), to which the reader is referred for further details, with the following two main differences and advancements. The first is proposing and incorporating an auto-adaptive regularization and error weighting function. This has improved the fast convergence of the method while keeping its stability. The second is the implementation of combined stopping criteria to terminate the iteration after an appropriate number of steps. The developed method uses an iteratively reweighted least squares (IRLS) minimization algorithm in combination with L_0 -norm stabilizer, depth weighting and physical parameter inequality constraint to estimate a compact and sharp density contrast model of the subsurface.

2 Methodology

2.1 The 2D model

Most fixed geometry gravity inversion algorithms, including the one presented here, employ rectangular prismatic elements, to discretize the subsurface, owing to their flexibility in constructing complex models (Silva & Barbosa, 2006; Commer, 2011; Grandis & Dahrin, 2014). A 2-D model is obtained by discretization of the subsurface under the survey area into a large number of infinitely long horizontal rectangular prisms, with the infinitely long dimension oriented in the invariant y -direction, with variations in densities only assumed for the X and Z directions. The 2-D model is illustrated in Fig. 1. The density contrasts are constant inside each cell only and can vary individually. Here we have used equal dimensions for the cells. However, the algorithm is flexible, to accommodate non-regular size cells. Gravity stations indicated by ∇ symbols are located at the centers of the upper faces of the rectangular blocks in the top layer. This discretization scheme of the subsurface allows us to calculate the gravitational attraction caused by each rectangular block separately.

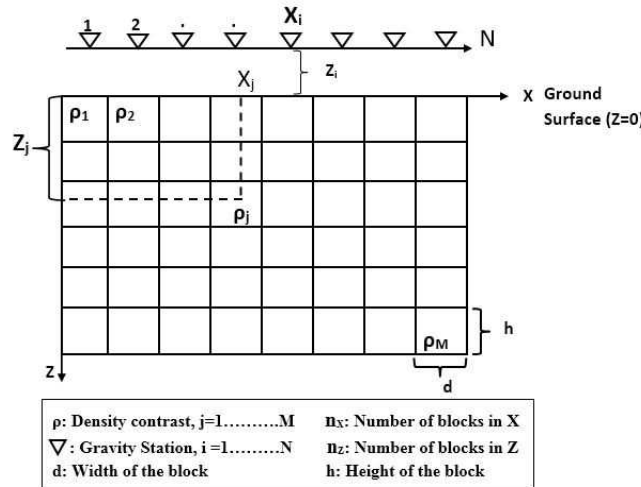


Figure 1: A 2-D model of the subsurface under a gravity profile. Gravity stations (X_i) are located at the centers of the blocks, indicated by the ∇ symbols.

2.2 Forward modelling

After discretization of the modeling space into a set of elementary rectangular blocks, the total vertical component gravity response calculated at the i^{th} observation point g_i is the sum of the gravity contributions generated by each of the individual rectangular element, on all points belonging to the observation

102 grid and it is given by:

$$g_i = \sum_{j=1}^M a_{ij} \rho_j + e_i \quad i = 1, 2, 3, \dots, N \quad (1)$$

103 where ρ_j is the density of the j^{th} prism; N denotes the numbers of observations; a_{ij} is the contribution
104 of j^{th} prism to the gravity value on i^{th} observation point and e_i is the noise associated with i^{th} data point.
105 The kernel a_{ij} is the forward operator that maps from the physical parameter space to the data space.
106 The exact mathematical expression of the kernel used here is presented by Last & Kubik (1983) which
107 is adopted from Nagy (1966) to which the reader is referred for more detail mathematical development.
108 In matrix notation Eq. (1) can be written as:

$$\mathbf{g} = \mathbf{A}\boldsymbol{\rho} + \mathbf{e} \quad (2)$$

109 where \mathbf{g} is an N-dimensional vector containing the gravity values, $\boldsymbol{\rho}$ is an M-dimensional model vector
110 of densities, \mathbf{A} is the N x M kernel matrix, and \mathbf{e} represents the noise vector at data points. Equation
111 (2) constitutes the gravity forward modeling, i.e. used to calculate the predicted gravity anomalies
112 (theoretical data) for a known subsurface density contrast (model $\boldsymbol{\rho}$).

113 **2.3 Inverse Modeling**

114 Our objective in solving gravity inverse problems is given the observed gravity data (\mathbf{g}), we seek a
115 solution that gives a density distribution $\boldsymbol{\rho}$ which predicts the observed data with a certain noise level
116 and at the same time, satisfies certain constraints. For the model presented here, the density vector
117 $\boldsymbol{\rho}$ is related to the predicted gravimetric field \mathbf{g} by the linear expression given in Eq. (2). Like the
118 majority of practical inverse problems arising in geophysical modeling gravity inversion is an ill-posed
119 problem. Moreover, usually we have less number of the observed gravity data than the number of the
120 model parameters which makes the system an under-determined problem. A standard way to solve such
121 ill-posed and under-determined problems, according to regularization theory (Tikhonov et al., 2013), is
122 minimization of the following objective function (Φ) which is the combination of data fidelity or misfit
123 functional (Φ_d) and stabilizing functional (stabilizer) term ($S(\boldsymbol{\rho})$):

$$\Phi = \Phi_d + \ell^2 S(\boldsymbol{\rho}) \quad (3)$$

124 Here the misfit functional is $\Phi_d = \|\mathbf{W}_e(\mathbf{A}\boldsymbol{\rho} - \mathbf{g}^{obs})\|_2^2$ and \mathbf{W}_e is error weighting diagonal matrix.
125 In Eq. (3), ℓ is a regularization parameter that controls the trade-off between the data fidelity and the
126 stabilizing term. Choosing a small value improves the data fit but the recovered models have highly
127 oscillatory artificial structures (which is equivalent to under-regularization). On the other hand, a large
128 value of ℓ leads to a large misfit value between the observed and predicted data and a small norm of the
129 model (over-regularizing the solution). Thus, the choice of a suitable value for ℓ is very important.
130 The choice of the stabilizing functional, in Eq. (3), depends on the desired model features that are to
131 be recovered. There are several types of stabilizers that have been developed and implemented in the
132 inversion of potential field data, which can roughly be divided into two categories: (I) Smooth stabilizer
133 which uses L_2 -norm of the model parameters or gradient of the model parameters (Li & Oldenburg,
134 1998; Cella & Fedi, 2012; Paoletti et al., 2013). (II) Non-smooth stabilizer which uses L_1 -norm or
135 L_0 -norm directly on the model parameters or on the gradient of the model parameters (Bertete-Aguirre
136 et al., 2002; Sun & Li, 2014; Li et al., 2018; Utsugi, 2019). Inversion methods that utilize a smooth stabi-
137 lizer produce model typically characterized by smooth features, and hence have difficulties in recovering
138 blocky structures or non-smooth distributions that have sharp boundaries or abrupt changes in physical
139 properties (Farquharson, 2008). To overcome this problem, non-smooth stabilizers that help to produce
140 compact and sharp models have been applied successfully (Zhdanov, 2009; Meng et al., 2018). Since
141 we are interested in developing a gravity inversion method that can produce compact and sharp models,
142 we use a non-smooth stabilizer through the L_0 -norm on the model parameters and will be discussed in
143 the next subsection. In general, with all mentioned stabilizers Eq. (3) needs to be solved by using an
144 iterative minimization algorithm. In this work, we use the IRLS algorithm to estimate the solution and
145 it is described below.

146 Using the classical weighted L_2 -norm stabilizing functional $S(\boldsymbol{\rho}) = \|\mathbf{W}_c(\boldsymbol{\rho} - \boldsymbol{\rho}_F)\|_2^2$ in the objective
147 function Φ (Eq. (3)) and minimizing by applying the standard weighted-damped least-square optimiza-
148 tion, the estimated density distribution in matrix notation can be given by (Menke, 1989, p. 55):

$$149 \quad \boldsymbol{\rho}^{k+1} = \boldsymbol{\rho}_F^k + \left[[\mathbf{W}_c^k]^{-1} \mathbf{A}^T \left(\mathbf{A} [\mathbf{W}_c^k]^{-1} \mathbf{A}^T + \ell^2 [\mathbf{W}_e^k]^{-1} \right)^{-1} \mathbf{g}_r^k \right] \quad (4)$$

150 where the superscript k denotes that variable at k^{th} iteration and \mathbf{W}_c^k is a combined weighting matrix.
151 $\boldsymbol{\rho}_F^k$ is reference density vector, which is from prior information or calculated at each iteration. $\mathbf{g}_r^k =$
152 $\mathbf{g}^{obs} - \mathbf{A}\boldsymbol{\rho}_F^k$ represents residual data vector computed at each iteration. Computation of the regularization

153 parameter ℓ in Eq. (4) will be described in Sect. 2.3.3. In this work, the combined weighting matrix
 154 (\mathbf{W}_c^k) is defined as a product of three different diagonal matrices, L₀-norm constraint matrix ($\mathbf{W}_{L_0}^k$),
 155 depth weighting (\mathbf{W}_z) and hard constraint matrix (\mathbf{W}_h^k).

$$\mathbf{W}_c^k = \mathbf{W}_{L_0}^k \mathbf{W}_z \mathbf{W}_h^k \quad (5)$$

156 2.3.1 L₀-norm Constraint

157 The L₀-norm is commonly defined as the number of nonzero elements in a vector. Because there is no
 158 analytical formula that meets the mathematical requirement to be regarded as L₀-norm, the approximate
 159 expression is usually used to convert the L₀-norm into an equivalent norm for the suitability of computa-
 160 tion. In literature (Zhao et al., 2016; Li & Yao, 2020) that discusses the inversion of potential field data,
 161 different L₀-norm approximate stabilization functions have been developed and implemented to obtain
 162 focused images and sharp boundaries. Meng (2016) used a hyperbolic tangent function to approximate
 163 the L₀-norm and applied it to the 3D inversion of gravity gradient tensor data. Meng et al. (2018) pro-
 164 posed an exponential mathematical function to approximate the L₀-norm for 3D gravity sparse inversion.
 165 In this paper, the minimum support functional, which is also called compactness constraint originally
 166 proposed by Last & Kubik (1983) and then further extended by Portniaguine & Zhdanov (1999) to
 167 include a reference model is selected which can be expressed as follows:

$$L_0(\rho) = \sum_{j=1}^M \frac{(\rho_j - \rho_j^{apr})^2}{(\rho_j - \rho_j^{apr})^2 + \varepsilon} \quad (6)$$

168 In our case to avoid the requirement of a prior model, we set $\rho_j^{apr} = 0$ and hence Eq. (6) can be rewritten
 169 as follows (Sun & Li, 2014):

$$L_0(\rho) = \sum_{j=1}^M \frac{\rho_j^2}{\rho_j^2 + \varepsilon} \quad (7)$$

170 where ε is a focusing parameter. Application of $L_0(\rho)$ as stabilizer in minimization process of the
 171 objective function (Eq. (3)) leads to the following choice of an L₀-norm constraint matrix \mathbf{W}_{L_0} which is
 172 given by (Last & Kubik, 1983):

$$[\mathbf{W}_{L_0}]_j = ([\rho_j]^2 + \varepsilon)^{-1} \quad (8)$$

173 Based on Eq. (8) the k^{th} iteration diagonal elements of the L_0 -norm constraint matrix ($\mathbf{W}_{L_0}^k$) can be
 174 formulated as follows:

$$[\mathbf{W}_{L_0}^k]_{jj}^{-1} = [\rho_j^{k-1}]^2 + \varepsilon \quad (9)$$

175 The focusing parameter ε is a very important parameter. Its main purpose is to avoid singularities when
 176 $\rho_j \rightarrow 0$. The parameter ε is a small number and in general, we are interested in the case where $\varepsilon \rightarrow 0$
 177 because a small value leads to very compact models. However, this may introduce instability. On the
 178 other hand, if ε is chosen large the L_0 -norm compactness constraint has no influence on the compactness
 179 of the model which means it results in a smooth solution. Figure 2 shows the comparison of the minimum
 180 support stabilizing functional for different values of ε to demonstrate the impact of the choice of different
 181 values of ε further. From Fig. 2, one can see that as ε becomes large the minimum support stabilizing
 182 function loses its property and behaves more like the minimum length L_2 -norm stabilizer which results
 183 in undesirable smoothness in the model though it improves the stability. Therefore, it is essential to
 choose an optimal value of ε .

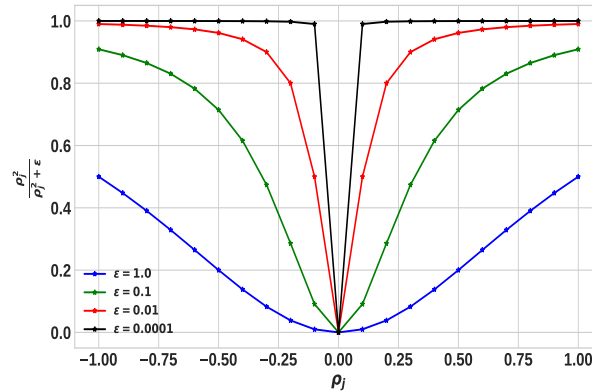


Figure 2: Comparison of the minimum support stabilizing function for different values of ε .

184

185 In previous investigations e.g Last & Kubik (1983) and Guillen & Menichetti (1984) the parameter ε
 186 was assigned a value close to machine precision ($\approx 10^{-11}$ to 10^{-15}). Alternatively, Zhdanov &
 187 Tolstaya (2004) introduced a trade-off curve method, similar to the L-curve technique, to select ε by
 188 computing the model objective for the current model estimate over a range of values for ε . However,
 189 as pointed out by Ajo-Franklin et al. (2007) setting ε to values near machine precision results in severe
 190 instability as $\rho_j \rightarrow 0$ and the approach of Zhdanov & Tolstaya (2004) often yields trade-off curves with
 191 corners that are not well defined. Therefore it is better to fix ε at a reasonable value determined by
 192 experience, typically between 10^{-4} to 10^{-7} (Ajo-Franklin et al., 2007). Accordingly, in the present

193 work based on several numerical simulation tests, the value 10^{-6} is assigned just for the inversion
 194 examples presented in the manuscript. Note that the developed method is flexible to use different values
 195 of ε .

196 2.3.2 Error weighting

197 According to compact inversion method proposed by Last & Kubik (1983), the k^{th} iteration error weight-
 198 ing matrix \mathbf{W}_e^k is defined as:

$$[\mathbf{W}_e^k]^{-1} = \text{diag} \left(\mathbf{A} [\mathbf{W}_{L_0}^k]^{-1} \mathbf{A}^T \right) \quad (10)$$

199 Even though \mathbf{W}_e^k expressed by Eq. (10) is applied by many authors (Guillen & Menichetti, 1984;
 200 Barbosa & Silva, 1994; Ghalehnoee et al., 2017; Gebre & Lewi, 2022), some instability was reported
 201 by Lewi (1997, p. 87) in using \mathbf{W}_e^k in scenarios such as complicated geological geometry and when the
 202 data is contaminated with noise. To overcome this problem Lewi (1997, p. 90) proposed a weighting
 203 matrix that makes use of the following equation:

$$[\mathbf{W}_e^k]^{-1} = \left[\frac{[\sigma_\rho^2]^k}{1 + [\sigma_e^2]^k} \right] \mathbf{I} \quad (11)$$

204 where \mathbf{I} represents identity matrix, and σ_ρ^2 and σ_e^2 are model and error variances respectively that are
 205 given by:

$$[\sigma_e^2]^k = \frac{\sum_{i=1}^N \{ \mathbf{g}_i - \sum_{j=1}^M a_{ij} [\boldsymbol{\rho}_j^{k-1}] \}^2}{N - 1} \quad (12)$$

207

208

$$[\sigma_\rho^2]^k = \frac{\sum_{j=1}^M [\boldsymbol{\rho}_j^{k-1}]^2}{M - 1} \quad (13)$$

209 The term in square brackets in Eq. (11) can be considered as regularization parameter (Silva & Barbosa,
 210 2006; Lewi, 1997, p. 90). Based on several numerical experiments done in the present work it was
 211 observed that this term can sometimes ends up in a larger value which may result in over-regularization
 212 of the solution. For this reason, in the present study, a new error weighting matrix \mathbf{W}_{ne}^k is introduced
 213 and it is given as:

214

$$[\mathbf{W}_{ne}^k]^{-1} = \text{diag} \left(\mathbf{A} \left[\mathbf{W}_z \left(\frac{[\sigma_\rho^2]^k}{1 + [\sigma_e^2]^k} \right) \mathbf{W}_h^k \right] \mathbf{A}^T \right) \quad (14)$$

215 Let us represent the terms in square brackets by \mathbf{W}_n^k as follows:

$$216 \quad \mathbf{W}_n^k = \mathbf{W}_z \left(\frac{[\sigma_\rho^2]^k}{1 + [\sigma_e^2]^k} \right) \mathbf{W}_h^k \quad (15)$$

217 where \mathbf{W}_z and \mathbf{W}_h^k are diagonal depth and hard constraint matrices respectively and will be described in
 218 the next subsections. Then the error weighting matrix in Eq. (14), the one introduced and implemented
 219 here becomes:

$$[\mathbf{W}_{ne}^k]^{-1} = \text{diag} \left(\mathbf{A} \mathbf{W}_n^k \mathbf{A}^T \right) \quad (16)$$

220 **2.3.3 Auto-adaptive Regularization Parameter Estimation**

221 Choosing a suitable value for the regularization parameter is a crucial part of the inversion process. The
 222 precise value of the regularization parameter depends on the noise level associated with the observed
 223 data. Thus, the higher value of ℓ refers to the higher noise level of the data points. Several methods
 224 have been proposed to choose the appropriate value of regularization parameter, and are reviewed in
 225 the literature (Farquharson & Oldenburg, 2004; Vatankhah et al., 2014) and standard texts for example
 226 Vogel (2002, pp. 97-109) and Aster et al. (2018, p. 57). Particularly, depending on the noise level a
 227 constant value of ℓ , throughout the inversion, has been chosen by many authors (Silva & Barbosa, 2006;
 228 Ghalehnoee et al., 2017). In other works, for example Zhdanov (2009) and Rezaie et al. (2017) the
 229 parameter ℓ has been iteratively updated in each iteration.

230 As pointed out in previous works (Farquharson & Oldenburg, 2004; Gholami & Aghamiry, 2017) in-
 231 stead of using a constant value of ℓ , dynamic re-adjustment throughout the iterative scheme might be
 232 a superior approach. Taking this into account, in the present work ℓ is updated in each iterative step.
 233 In our implementation, to select an optimal regularization parameter at each iteration, we proposed an
 234 auto-adaptive regularization method. This method leads to an automatic update of the regularization
 235 parameter at each and every iteration. The basic principle including its procedure in relation to the
 236 formally known adaptive regularization approach which was proposed by Zhdanov (2002, p. 55) and
 237 implemented by many authors (Zhdanov, 2009; Rezaie et al., 2017) is as follows. In adaptive regular-
 238 ization approach the initial value of the regularization parameter ℓ^1 is updated at each iteration step by
 239 (Zhdanov, 2002, p. 55):

$$\ell^k = \ell^1 q^k \quad (17)$$

240 where q , as described by Zhdanov (2002, p. 55), is damping factor which decreases from iteration to
 241 iteration. Its initial value is empirically determined having a value between zero and one. It is obvious
 242 that the trial and error selection of the value for q requires computational work . The presented auto-
 243 adaptive regularization method overcomes this problem and the iterative values ℓ^k are determined by the
 244 following formula:

$$\ell^k = \ell^{k-1} \left[\frac{|g^{obs} - A\rho|_{max}^{k-1}}{|g^{obs} - A\rho|_{max}^k} \right] \quad (18)$$

245 where the term in the square bracket is an adjusting factor that is automatically determined at each
 246 iterative step and $|g^{obs} - A\rho|_{max}$ is the maximum absolute value of the residual data elements. In
 247 the auto-adaptive regularization method, choosing a suitable initial value of (ℓ_o) is essential. Based
 248 on a number of synthetic and real data simulations done in this work we recommend the following in
 249 choosing a reasonable value of ℓ_o : Firstly, the initial value of ℓ should be within the range $0 < \ell_o \leq 1$.
 250 Secondly, the precise value of ℓ_o depends on the noise level related to the observed data. When the
 251 probable or expected noise level of the data is higher, a larger value ℓ_o is a reasonable choice to avoid
 252 unwanted and false anomalies due to noise. In contrast, when the probable or expected noise level is less
 253 a small value of ℓ_o should be chosen. Once an appropriate initial value ℓ_o is given as an input, then for
 254 subsequent iterations Eq. (18) is used to determine ℓ^k . The advantage of the auto-adaptive regularization
 255 scheme is its capability to automatically determine a suitable regularization parameter, in the course of
 256 the optimization process, depending on the automatically determined adjusting factor.

257 **2.3.4 Physical Parameter Inequality Constraint**

258 To produce a physically meaningful model from a gravity inverse solution, the usage of lower and upper
 259 bound constraints on the recovered density contrast is beneficial (Silva et al., 2001; Grandis & Dahrin,
 260 2014). Lower and upper bounds can be obtained from a prior information such as geological investiga-
 261 tions in conjunction with published density values of rocks, well-logging, and/or laboratory tests. Many
 262 procedures such as gradient projection approach (Wang & Ma, 2007; Lelièvre et al., 2009), transform
 263 function approach (Pilkington, 2008) and logarithmic barrier approach (Li & Oldenburg, 2003) have
 264 been applied in different inversion schemes to implement this constraint. However, with regard to L₀-
 265 norm stabilizer based gravity inversion methods an effective method is the direct utilization of lower
 266 and upper density constraints (Meng et al., 2018). Hence, in this work the direct density bound inequal-
 267 ity constraint is used, that is at each iteration density contrast of each rectangular block is bounded by

268 minimum and maximum density constraint function given by:

$$269 \quad [\rho^k]_j = \begin{cases} [\rho_{max}]_j & \text{if } [\rho^k]_j > [\rho_{max}]_j \\ [\rho^k]_j & \text{if } [\rho_{min}]_j < [\rho^k]_j < [\rho_{max}]_j \\ [\rho_{min}]_j & \text{if } [\rho^k]_j < [\rho_{min}]_j \end{cases} \quad (19)$$

270 By using this function, if k^{th} iteration ρ_j of any block exceeds one of its bounds, then it will be fixed at
271 the violated bound.

272 In each iteration step the procedure to compute the hard constraint matrix \mathbf{W}_h^k (Boulianger & Chouteau,
273 2001) and the reference density vector ρ_F^k is determined as follows: The diagonal elements of \mathbf{W}_h^k are
274 fixed at ε or 1.0. When a prior geological and geophysical information are able to provide the initial
275 value of density contrast of the j^{th} specific cells, then these values are assigned to the corresponding
276 $[\rho_F^k]_j$. Simultaneously, the corresponding diagonal elements of $[\mathbf{W}_h^k]_{jj}$ are set to be ε . During the in-
277 version process, if the j^{th} elements of estimated density values falls out of inequality constraint limits
278 defined by ρ_{min} and ρ_{max} , then $[\rho_F^k]_j$ will be fixed at the violated bound density itself and $[\mathbf{W}_h^k]_{jj}$
279 assigned to be ε . On the other hand, if the elements of the estimated density did not exceed its bounds
280 (i.e. lies between the limits), $[\mathbf{W}_h^k]_{jj}$ and $[\rho_F^k]_j$ are assigned to be 1.0 and 0.0 respectively.

281 Using \mathbf{W}_h^k any blocks whose density is known from a priori information or exceeds the density con-
282 straint limit, the algorithm will automatically freezes this block in the next iteration by assigning a very
283 small weight to it. Whereas, ρ_F^k is used to remove the gravity effects of those cells that have crossed
284 the inequality constraint limit from the observed gravity data. That is applied to compute the reduced
285 gravity data vector $\mathbf{g}_r^k = \mathbf{g}^{obs} - \mathbf{A}\rho_F^k$ in Eq. (4) of the inversion algorithm. In other word, at each
286 iterative step the inversion of subsequent iteration will be performed using reduced gravity data vector.

287 2.3.5 Depth weighting

288 It is well known that gravity data, like any potential field data, has no inherent depth resolution. The
289 reconstructed model structures by the inversion process tend to concentrate near the surface regardless of
290 the true depth of the causative bodies (Li & Oldenburg, 1996). This happens because the inverse solution
291 of model construction is a linear combination of kernel, whose amplitudes rapidly decay with depth. The
292 problem can be overcome by introducing a depth weighting matrix to counteract the natural decay of
293 kernel with depth (Li & Oldenburg, 1998). Depth weighting is designed to ensure that all cells have
294 equal likelihood to accommodate the sources, not just those at shallow levels that are most sensitive to

295 the observed data. Depth weighting is used and its effect is investigated by different authors (Pilkington,
 296 2008; Commer, 2011). Based on Gebre & Lewi (2022), the recently proposed depth weighting function
 297 is given as follows:

$$298 \quad w_{zj} = (aZ_j + c_o)^{-\tau} \quad (20)$$

299 where Z_j is the mean depth of the j^{th} cell and a , c_o and τ are adjustable parameters. The values of
 300 the three adjustable parameters are computed by optimizing $w_z(z)$ to match with the actual gravity
 301 kernel values utilizing nonlinear least-squares minimization (Virtanen et al., 2020). Accordingly, for all
 302 inversions in this work the depth weighting matrix similar to the one used by Gebre & Lewi (2022) is
 303 employed (Eq. (21)):

$$304 \quad [\mathbf{W}_z]_{jj} = \mathit{diag}(w_{zj}) \quad (21)$$

305 where \mathbf{W}_z is diagonal $M \times M$ depth weighting matrix.

306 2.3.6 Stopping Criteria

307 It is clear that if the iterations are stopped too early, then a reasonable solution of the inverse problem may
 308 not be obtained. On the other hand, too many iterations may waste computer time without increasing the
 309 overall solution qualities. Thus, an important aspect of any iterative inversion method is to decide when
 310 the iterations should be terminated. A number of stopping criteria have been proposed and employed to
 311 terminate iterative inversion algorithms (Borges et al., 2015; Levin & Meltzer, 2017). Commonly used
 312 stopping criteria are based on a norm of the residual vector (i.e. the norm of the difference between
 313 estimated and observed data). For instance, a noise level, i.e. $\chi^2 = \|\mathbf{W}_d(\mathbf{g}^{obs} - \mathbf{A}\boldsymbol{\rho})\|_2^2$, where a
 314 diagonal data weighting matrix \mathbf{W}_d , whose i^{th} element is the inverse of the standard deviation of the
 315 noise at each data point, is used by Boulanger & Chouteau (2001) and Vatankhah et al. (2017). Other
 316 criteria for stopping gravity inversion procedure are based on simple *misfit* or the Root Mean Square
 317 Error (*RMSE*) between the observed data and predicted data produced by the recovered model (see, for
 318 example Rezaie & Moazam (2017)). The expressions used to estimate these criteria are the following:

$$319 \quad \mathit{misfit} = \left(\frac{\sum_{i=1}^N (\mathbf{g}_i^{obs} - \mathbf{g}_i^{cal})^2}{\sum_{i=1}^N (\mathbf{g}_i^{obs})^2} \right)^{\frac{1}{2}} \quad (22)$$

$$320 \quad \mathit{RMSE} = \frac{(\sum_{i=1}^N (\mathbf{g}_i^{obs} - \mathbf{g}_i^{cal})^2)^{\frac{1}{2}}}{N} \quad (23)$$

322 Ekinci (2008) also introduced other possible criterion, namely the parameter variation function (*smv*)
 323 which is defined as:

$$324 \quad smv = \left(\sum_{j=1}^M (\rho_j^k - \rho_j^{k-1})^2 \right)^{\frac{1}{2}} \quad (24)$$

325

326 The most widely used approach is to quit the iterative process when one of the above criteria are below
 327 a given tolerance (the level of observational error). However, in practical applications a precise value
 328 for such tolerance is rarely known; rather, only some possibly vague idea of the desired quality of
 329 the numerical approximation is at hand. Moreover, it has been pointed out by Rao et al. (2018) that
 330 stopping iteration based solely on the norm of the residual is neither safe nor a robust solution. The
 331 non-uniqueness and instability of the gravity inverse problem further complicates the usage of only
 332 one of the aforementioned stopping criteria. To overcome these issues, a combination of the *misfit*
 333 and *smv* has been utilized in this paper. Therefore, the iterative procedure continues until one of the
 334 following stopping criteria is met: (I) the maximum number of iteration (k_{max}) given by the user is
 335 reached or (II) the difference between two consecutive iteration values of *smv* and *misfit* have reached
 336 the target values. That means for the second criterion both the conditions $|smv^{k-1} - smv^k| \leq \tau$ and
 337 $|misfit^{k-1} - misfit^k| \leq \mu$ must be satisfied at the same time. In all demonstrations considered in this
 338 work, after testing different values, the parameter τ is assigned to $\sqrt{2M}$; and μ to 0.005. Where M is
 339 again the total number of model parameters. The effectiveness of the proposed termination criteria will
 340 be illustrated by using synthetic tests.

341 2.4 Computational procedure

342 The solution of the linear system of equations in Eq. (2) will be carried iteratively using the information
 343 about the misfit and density from successive iteration. The input parameters for the inversion proce-
 344 dure are: (1) Kernel matrix (\mathbf{A}) and discretized subsurface model (mesh) and its initial approximation
 345 reference density model ρ_F if exists based on a priori information; (2) Observed gravity anomaly (\mathbf{g})
 346 at measurement points (\mathbf{x}); (3) Maximum number of iteration (k_{max}) and the constant β ; (4) Lower
 347 ρ_{min} and upper ρ_{max} density bounds and initial ℓ_o value. In summary, the steps taken to carry out the
 348 inversion process consists the followings.

- 349 1. For $k = 0$, if there is no a priori information, \mathbf{W}_{L_0} , \mathbf{W}_c , \mathbf{W}_n and \mathbf{W}_h are identity matrices,
 350 $\rho_F = 0$. \mathbf{W}_z and \mathbf{W}_{ne} are computed through Eq. (21) and (16) respectively, after this, the first

- 351 iteration model parameters solution is obtained by Eq. (4).
- 352 2. The elements of \mathbf{W}_h and ρ_F are updated as explained in preceding section, then \mathbf{W}_{L_0} is calcu-
353 lated using Eq. (9) and then \mathbf{W}_c using Eq. (5).
- 354 3. Compute the value of σ_ρ and σ_e using expressions (13) and (12) respectively. Then calculate \mathbf{W}_n
355 using Eq. (15).
- 356 4. To remove the effect of those blocks that have crossed the maximum target density, evaluate the
357 reduced data $\mathbf{g}_r^k = \mathbf{g}^{obs} - \mathbf{A}\rho_F^k$. Then compute the current ℓ with Eq. (18) and \mathbf{W}_{ne} with Eq.
358 (16).
- 359 5. Carrying out the inversion through Eq. (4).
- 360 6. Application of bounded constraints on density are carried out as discussed in the preceding sec-
361 tion.
- 362 7. Now a forward modelling procedure will be carried out using Eq. (2) to compute the gravity
363 anomaly \mathbf{g}^{cal} from the estimated model in the previous iteration.
- 364 8. Data *misfit* (Eq. (22)) and *smv* (Eq. (24)) are computed using \mathbf{g}^{cal} from step 7, and obtained
365 model parameters from the previous and current iteration.
- 366 9. Test if the stopping criteria are fulfilled. If the termination criteria are satisfied the iteration
367 terminates and obtained results are stored and plotted. Otherwise, using the current estimated
368 density model, move to the next iteration k by going to the second step and continue the iterative
369 procedure until the stopping criteria are fulfilled.

370 **3 Synthetic Model Test**

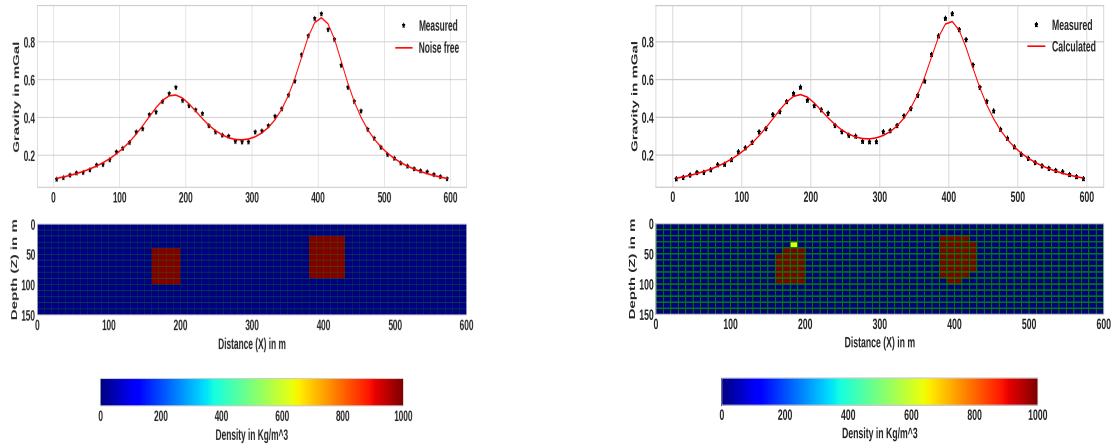
371 To evaluate the functionality and efficiency of the method, the developed procedure was tested on sev-
372 eral synthetic model examples. The examples presented here are randomly chosen to demonstrate: (I)
373 the applicability of the proposed auto-adaptive regularization technique (Eq. (18)) and error weighting
374 function (Eq. (16)); (II) the performance of the method in producing compact and sharp images of the
375 causative bodies; (III) the effectiveness of the combined stopping criterion. The forward and the inverse
376 problem were carried out using the procedure described in the preceding sections. In the inversion of the

377 synthetic examples, the same subsurface discretization as the one used in generating the synthetic data
378 (Forward modeling) is used. All the inversion tests are performed on a Desktop computer (11th Gen
379 Intel(R) Core(TM) i7-11700, 2.50GHz processor). For the first and second synthetic examples presented
380 in this work: (I) The model region was discretized into 60 x 15 rectangular cells and the dimensions of
381 each cell were taken as 10 x 10 m, in the X and Y directions respectively. (II) The synthetic gravity data
382 were computed at 60 data points that are centered in each cell at the top side of the model, to produce
383 data at 10 m sample interval. (III) The computed gravity data are contaminated with Gaussian noise
384 that has a standard deviation that amounts to 4 % of the magnitude at each data point with zero mean
385 (Farquharson, 2008; Rezaie et al., 2017).

386 The first synthetic data inversion has been done for the model presented in Fig. 3(a). For this synthetic
387 model the causative bodies are two rectangular structures elongated differently in the horizontal and
388 vertical directions and located at different depths. The causative bodies have the same density contrast
389 1000 kg/m^3 . The density of the causative bodies are given relative to the zero density of uniform back-
390 ground. Figure 3(a) upper panel shows noise free (solid line) and noise contaminated (star dots) gravity
391 data. Separate inversion runs, for three different ℓ_o values (0.2, 0.3 and 0.4), were performed with the
392 developed inversion method. Note that, for subsequent iterations the proposed auto-adaptive regular-
393 ization technique (Eq. (18)) is used to compute ℓ for each case. At the beginning of the inversion, the
394 iterations are initialized with $\rho_F = 0$ and $\mathbf{W}_h = \mathbf{W}_c = \mathbf{W}_n = \mathbf{W}_{L_0} = \mathbf{I}$. The lower limit density
395 contrasts of all cells is zero ($\rho_{min} = 0$) and the upper bound $\rho_{max} = 1000 \text{ kg/m}^3$.

396 The results of the inversion by using the developed method for three different ℓ_o values are shown in
397 Figs. 3(b) and 4. The corresponding data fit between the predicted (solid line) and the actual contam-
398 inated (stars) gravity data are also shown. Comparing the inversion results with the original synthetic
399 model in Fig. 3(a), the inversion has sufficiently recovered the true models. The depth, geometry, and
400 density distributions of the synthetic causative bodies were recovered adequately. This can confirm
401 the applicability of the proposed auto-adaptive regularization technique (Eq. (18)) and error weighting
402 function (Eq. (16)). Notice that the results also indicate the robustness and stability of the developed
403 inversion method for different ℓ_o values. The average computation time to finish the inversion is approx-
404 imately 16.3 seconds.

405 The second synthetic model is more complicated and consists of two causative bodies placed at various
406 depth. The bodies have different sizes, shapes, and density contrasts. The first causative body is a verti-
407 cal rectangular block, with density contrast 2000 kg/m^3 , placed at 40 m depth and the second body is

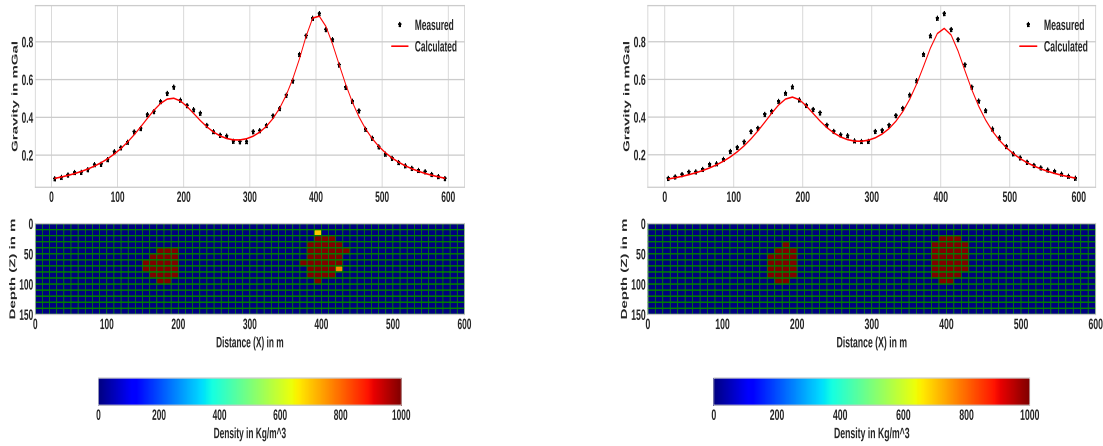


(a) The lower panel represents 2-D synthetic model, which constitutes two isolated rectangular bodies located at various depths and the top panel shows the gravity anomaly due to these two subsurface rectangular bodies.

(b) The lower panel represents the subsurface, as a result of the proposed inversion method using $\ell_o = 0.3$ and the top panel shows the synthetic data together data derived from the model.

Figure 3: The first synthetic model and the result of the inversion.

408 a dipping dike with density contrast 3000 kg/m^3 at 20 m depth. The synthetic model is shown in the
 409 lower part of Fig. 5(a) and the generated noise-corrupted and noise free gravity data are shown on the
 410 upper part. Using the generated synthetic data, the inversion was initiated by assigning an initial zero
 411 density to each cell. We set initial $\ell_o = 0.3$. The density contrast limits are bounded between lower
 412 bound $\rho_{min} = 0$ and the upper bound $\rho_{max} = 3000 \text{ kg/m}^3$. Even though a maximum iteration of 20
 413 was set, the *misfit* and *smv* between two consecutive iterations gradually fall below the threshold set
 414 after the 14th iteration. The total computation time is approximately 15.73 seconds. In Fig. 5(b), the
 415 resulting model from the inversion of the second synthetic model (Fig. 5(a)) using the proposed method
 416 is presented. As can be seen in Fig. 5(b) upper panel the modeled gravity data (solid line) fits adequately
 417 with the synthetic data. The result, presented in Fig. 5(b) lower panel, indicates an acceptable recon-
 418 struction of the synthetic multi-sources and multi-shape bodies that are located at different depths. The
 419 true shape, location and density of the causative bodies are recovered adequately. Like the first example
 420 the reproduced images of the localized multiple sources are compact and sharp (Fig. 5(b) lower panel).
 421 For the third and fourth synthetic examples: (I) The subsurface model was discretized into 100 x 20
 422 rectangular cells. Each cell has a size of 50 m in X and Z directions. (II) The synthetic gravity data were
 423 computed on 100 data points with a sample spacing of 50 m. The third synthetic model includes two
 424 dipping dikes in opposite directions. The causative 2-D bodies have different sizes and the same density

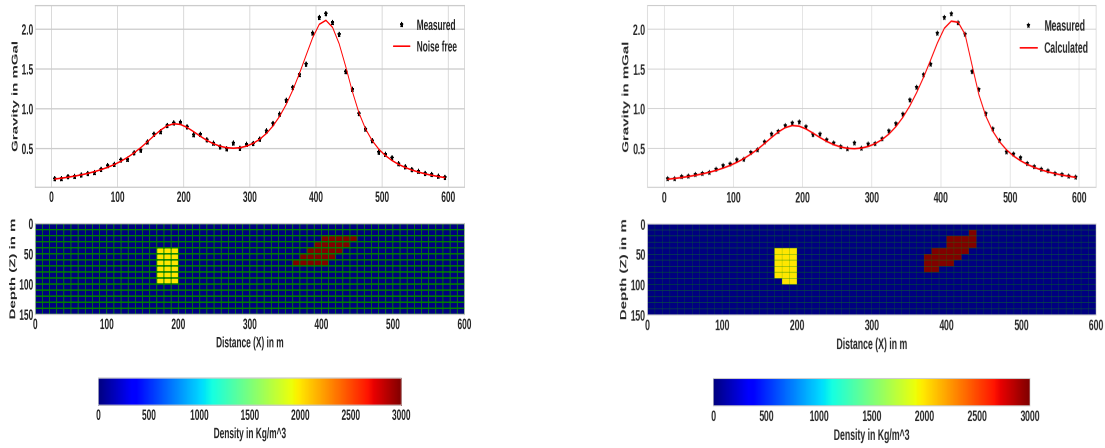


(a) Using $\ell_o = 0.2$.

(b) Using $\ell_o = 0.4$.

Figure 4: Inversion results, using different ℓ_o values, for the first synthetic model given in Fig. 3(a).

425 contrast that amounts to $1000 \text{ kg}/\text{m}^3$ in a homogeneous background zero density. The top part of the
 426 shallower dipping dike lies at a depth of 200 m and that of the deeper dike at a depth of 250 m. The
 427 computed gravity data were contaminated by uncorrelated Gaussian noise whose standard deviation was
 428 equal to 4% of the difference between the maximum and the minimum anomaly and zero mean. The
 429 synthetic model and the corresponding data are shown in Fig. 6 at lower and upper panels respectively.
 430 The inversion process was commenced by setting the densities of all cells to zero. The initial value of
 431 ℓ_o was set to 0.4. The bounding density ranges were set to a minimum value $\rho_{min} = 0$ and maximum
 432 value $\rho_{max} = 1000 \text{ kg}/\text{m}^3$. The maximum number of iterations was set to 20. Here, the inversion
 433 converged after the 13th iteration and the total computation time is approximately 66.49 seconds. The
 434 resulting model and the inverted data using the proposed method are shown in Fig. 7(b). For the sake
 435 of comparison keeping all inversion parameters the same, the synthetic data was also inverted with the
 436 classical L_2 -norm regularized inversion approach and the obtained result is shown in Fig. 7(a). As it
 437 can be seen from the lower panel of Fig. 7(b), unlike the model in Fig. 7(a), the developed method was
 438 able to produce a compact and sharp model successfully. The other concern, which can be seen from
 439 the result in Fig. 7(a), is that the target density contrast values are underestimated in the case of the
 440 conventional L_2 -norm inversion. In contrast, the geometrics, locations, and densities of both anomalous
 441 structures were adequately recovered with the presented inversion method (see Fig. 7(b)).
 442 The fourth synthetic model consists of two different rectangular anomalous bodies (Fig. 8(a) lower
 443 panel). The anomalous structures have different dimensions and are buried at different depths. The top



(a) Synthetic model consisting of a dipping dike and vertical rectangular block and the corresponding gravity data.

(b) The density model obtained by inverting the gravity data using the developed method. The predicted data as a result of inversion process are shown on the top panels (solid line).

Figure 5: The second example synthetic model and the corresponding inversion result.

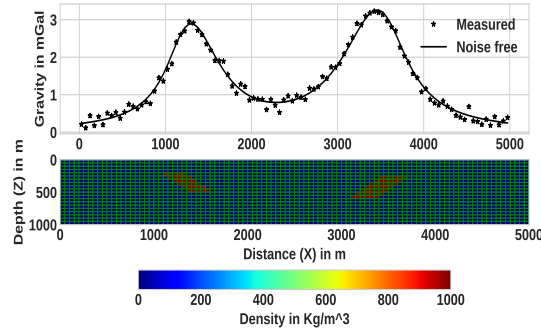
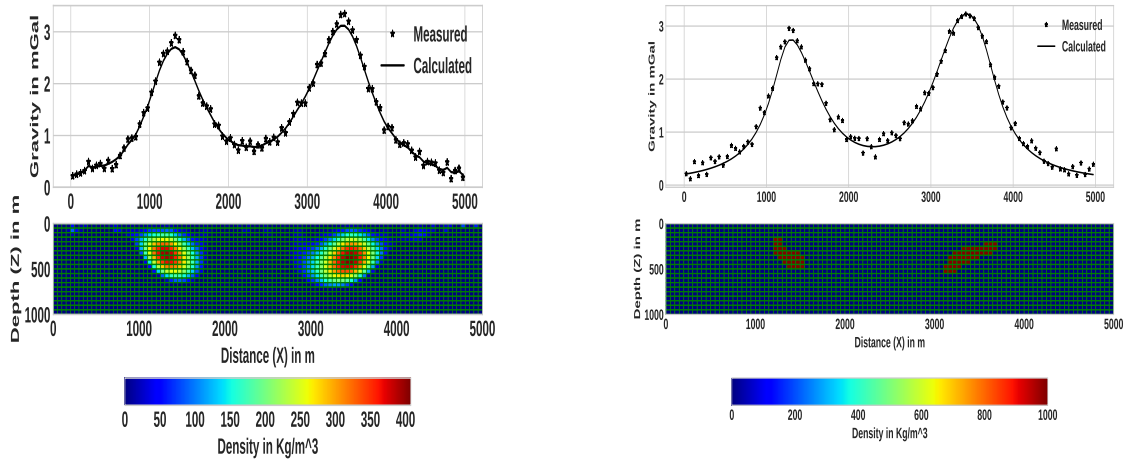


Figure 6: The third synthetic model that comprises two dikes at various depths with the density contrast that amounts to 1000 kg/m^3 and the corresponding gravity data.

444 of the first rectangular block is placed at a depth of 200 m and its density contrast is -1000 kg/m^3 while
 445 the top of the second block is placed at a depth of 250 m and has a density contrast of 1000 kg/m^3 .
 446 Different density contrast, size, and depth of adjacent structures have been considered to show the ability
 447 of the presented inversion method in reconstructing true parameters for these models. In this synthetic
 448 example, the computed data are contaminated by Gaussian noise with a standard deviation of 3% of the
 449 difference between the maximum and the minimum anomaly.

450 For the current example, the inversion process was initialized by setting the initial value of $\ell_o = 0.5$.
 451 The lower bound for the density constraint $\rho_{min} = -1000 \text{ kg/m}^3$ and the upper bound $\rho_{min} = 1000$
 452 kg/m^3 . Similar to the previous examples, though the maximum number of iterations was set to be 20,



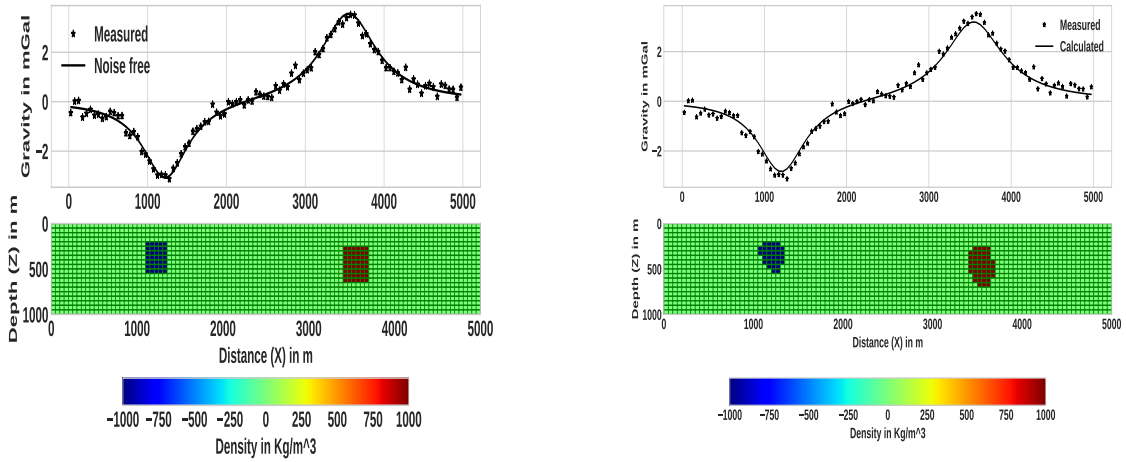
(a) Using the conventional minimum norm (L_2 -norm) smooth stabilizer and the corresponding data fit.

(b) Using the presented method

Figure 7: Inversion results of the third synthetic example in Fig. 6.

453 the iterative step terminated when the proposed combined criterion is satisfied after 11 iterations. The ap-
 454 proximate running time required to finish the inversion is 55.64 seconds. Figure 8(b) lower panel shows
 455 the recovered density contrast model. The corresponding fits between synthetic (stars) and predicted
 456 data (line) are shown in the upper panel of the same figure. We can see that the recovered rectangular
 457 bodies are compact and have sharp boundaries. The obtained results also indicate that the depth and
 458 density contrast of the anomalous rectangular bodies have been determined sufficiently.

459 Here, the effectiveness and the advantage of the proposed combined stopping criterion are illustrated
 460 by comparing it with another commonly used stopping condition. For this reason, the inversion process
 461 was performed again with the developed inversion method using only the misfit function ($|misfit^{k-1} -$
 462 $misfit^k|$) as a stopping condition. Note that, for comparison purposes, all the other inversion parame-
 463 ters are set the same except for the stopping criterion. The resulting recovered density contrast models
 464 and the data fit are presented in Fig. 9. The corresponding values of the $misfit$ and smv as a function
 465 of iteration number are also shown in Fig. 10(a). For the sake of comparison, the $misfit$ and smv
 466 when using the proposed combined stopping criterion for the same data set are also presented in Fig.
 467 10(b). The stopping condition $|misfit^{k-1} - misfit^k| \leq \mu$ was reached after 5 iterations, as shown
 468 in the curve of Fig. 10(a) before the true density distribution has been recovered fully. In other words,
 469 the estimated models are not satisfactory because densities lower than the target density are observed
 470 around the edges of the anomalous bodies (Fig. 9). This indicates that unlike the result presented in
 471 Fig. 8(b), where the proposed combined stopping condition is used, quitting the iterative process only



(a) Synthetic model consisting of two rectangular bodies at various depths with different density contrast and the corresponding noise free and contaminated gravity anomalies.

(b) The lower panel shows recovered density contrast model obtained by inverting the gravity data using the developed method, while the upper one shows the associated fits between the synthetic data that is taken from (a) and the predicted response.

Figure 8: The fourth synthetic model example and the corresponding inversion result.

472 with $|misfit^{k-1} - misfit^k| \leq \mu$ criterion produces a premature solution that is before the maximum
 473 compactness is achieved.

A number of other numerical experiments we carried out showed that there are situation where either

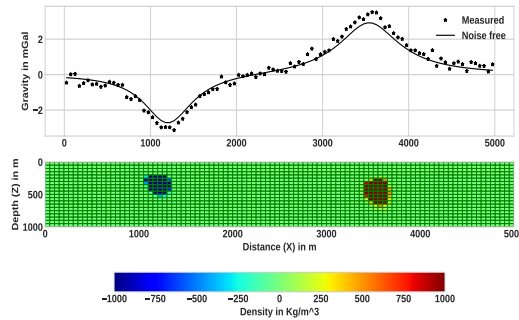
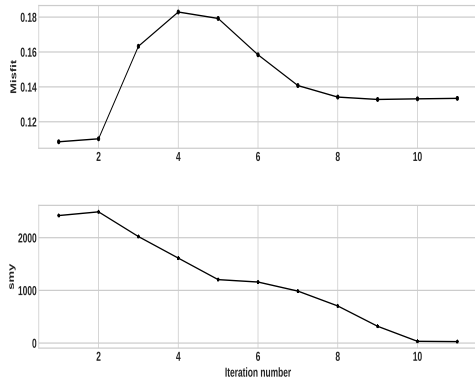


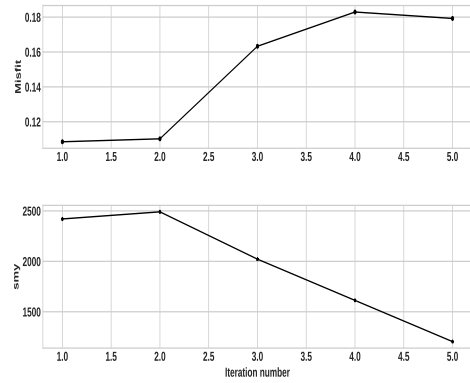
Figure 9: Inversion result obtained using only the commonly used criterion ($|misfit^{k-1} - misfit^k|$) and the corresponding data fit (upper panels) for the synthetic example in Fig. 8(a). The obtained density model shows that compact and sharp model is not approximately achieved due to the termination before the iterative procedure has reached convergence.

474

475 $|misfit^k|$ or $|misfit^{k-1} - misfit^k|$ fall below the given threshold values, at earlier iterations, before the
 476 true density is fully recovered. Thus, it is hard to take only one criterion as a termination condition. As
 477 stated in Sect. 2.3.6, it has been mentioned that the same has also be pointed out in number of previous
 478 works (Rao et al., 2018). Whereas, in the case of the proposed criterion that is when both the conditions



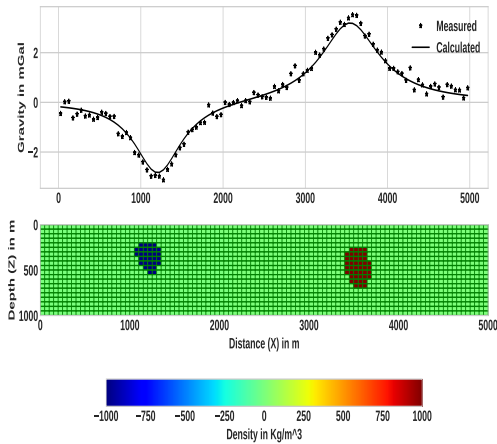
(a) Using the proposed combined stopping condition



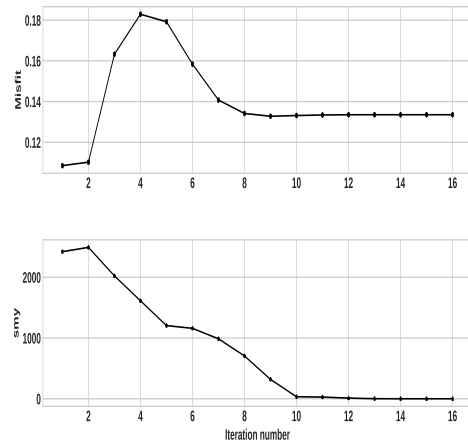
(b) Using only $|misfit^{k-1} - misfit^k| \leq \mu$

Figure 10: The progression of $misfit$ and sm_y in the course of the iteration during the inversion of the fourth example synthetic data.

479 $|sm_y^{k-1} - sm_y^k| \leq \tau$ and $|misfit^{k-1} - misfit^k| \leq \mu$ are satisfied at the same time the inversion
 480 process yields an acceptable model. This clearly illustrates the advantage of using the proposed stopping
 criterion and its effectiveness in quitting the iterative scheme after optimal number of iterations. To fur-



(a) The obtained recovered density model (lower panel) and the corresponding data fit (upper panel).



(b) Progression of $misfit$ (top panel) and sm_y (lower panel) in the course of the iterative procedure.

Figure 11: Late iteration termination (at 16th iteration) inversion result and the corresponding $misfit$ and sm_y variation with iteration number for the fourth example in Fig 8.

481

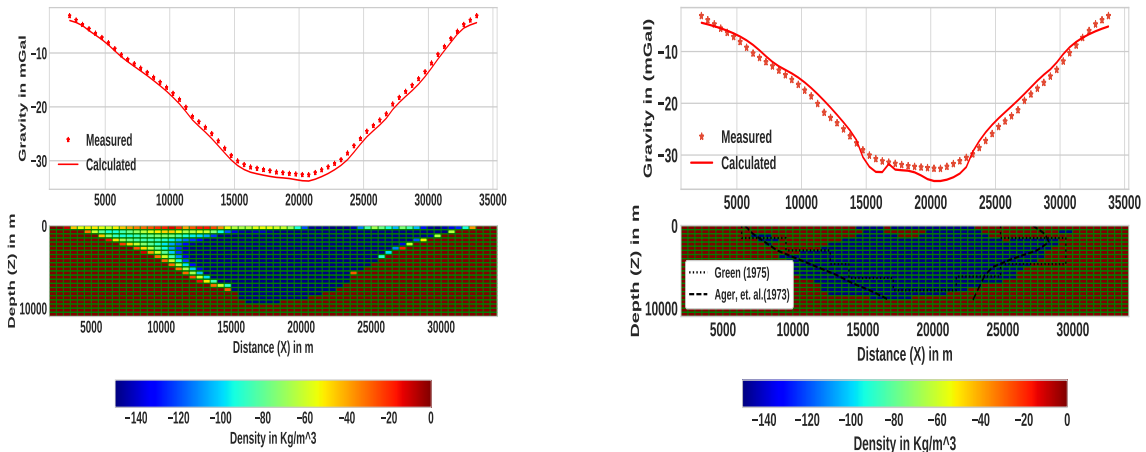
482 ther illustrate the effectiveness of the proposed combined criterion, the inversion process is allowed to
 483 continue to the 16th iteration and the model as a result of this is presented in Fig. 11(a). The progression
 484 of the $misfit$ and sm_y in the course of the iterative procedure are also given in Fig. 11(b). As it can be
 485 seen from the result (Fig. 11(b)) the solution obtained at subsequent iterations, after the 11th iteration

486 where the iteration is terminated with the proposed stopping condition, remains virtually unaltered. This
487 can also be observed from the *misfit* and *smv* variation curves shown in Fig. 11(b), in such that after
488 11th iteration the *misfit* and *smv* values remain literally unchanged. Moreover, the results also indicate
489 the appropriateness of the suggested threshold values μ and τ used in the proposed stopping criterion.
490 The other thing one can observe from the results in Fig. 11 is the stability of the developed inversion
491 method. This can also illustrate the effectiveness of the newly proposed auto-adaptive regularization
492 technique (Eq. (18)) and error weighting function (Eq. (16)).

493 In general, the presented method was tested with noise contaminated data that are generated from dif-
494 ferent geometries, locations, sizes, and densities contrasts of causative bodies and it has successfully
495 recovered all models. Moreover, all the reconstructed images of the presented synthetic models are
496 compact and sharp. Numerous synthetic data inversions were performed to analyze the impact of the
497 density contrast bounds. The obtained results, which are not presented here, suggest that the values of
498 density contrast bounds have a significant effect on the results, and hence to recover a feasible model a
499 good knowledge of the density bounds is vital. This also pointed out by number of authors, for exam-
500 ple Vatankhah et al. (2017); Li et al. (2018) and Utsugi (2019), in the case of inversion methods that
501 use non-smooth stabilizers (L_1 -norm or L_0 -norm). Provided that the lower and upper density contrast
502 bounds are chosen properly, this inversion technique produces acceptable solutions. Therefore, as it was
503 demonstrated using synthetic examples, the proposed method has effectively and efficiently recovered
504 the synthetic models. Generally, the tests performed on different geometry synthetic models showed that
505 the method gives acceptable results for localized multi-sources anomalies at different depths with sharp
506 features.

507 **4 Real Data Test**

508 To test the method in the real world, where the gravity data is contaminated with noise the improved
509 algorithm is implemented on gravity data acquired on different published geologic settings. The first one
510 is taken from Green (1975) by carefully digitizing the residual gravity data. As it was given in Green
511 (1975) the data was measured over the Guichon Creek batholith in south-central British Columbia. For
512 the details about the measurement and geology the reader is referred to Ager et al. (1973) and Ager
513 (1972). The residual gravity profile is digitized at a regular intervals of 0.5 km to produce a total of 64
514 data points as shown in Fig. 12 (star marks).



(a) Using the conventional minimum norm (L_2 -norm) smooth stabilizer.

(b) Using the presented method

Figure 12: The observed gravity anomaly over Guichon Creek batholith in south-central British Columbia (after Green (1975)) and its inversion results. Digitized data (star marks) with calculated data (solid line) are shown on the top panels of each subfigure. The corresponding recovered density contrast models are shown on the bottom. For comparison, the results obtained by Ager et al. (1973), which was obtained from drilling and Green (1975) are also presented.

515 For the inversion, the source volume beneath the anomaly was divided into 64×22 square lattice with
 516 dimensions of each cell being 0.5 km in both X and Z-directions. Based on the a prior information from
 517 Ager (1972) density values were constrained between the limits $\rho_{min} = -150 \text{ kg/m}^3$ and $\rho_{max} = 0.001$
 518 kg/m^3 . We start the inversion with a homogeneous initial model in which every block has the same
 519 zero density and an initial ℓ_o value of 0.48. The inversion was terminated after 9th iteration because the
 520 stopping criteria are fulfilled. The resulting model is presented in Fig. 12(b). For comparison, the results
 521 obtained by Ager et al. (1973) and Green (1975) are also included in Fig. 12(b). In addition, using the
 522 same inversion parameters we have performed L_2 -norm regularized inversion and the obtained result is
 523 shown in Fig. 12(a). The shape, real extent of the anomaly, and depth to bottom from the developed
 524 method are very close to the true geological feature (Ager et al., 1973) which was obtained from drilling.
 525 That means the implementation of the presented method resulted in a better solution compared to Green
 526 (1975) and the conventional L_2 -norm inversion. Note that, this reasonable result is obtained by using
 527 only the density contrast limits as a prior information.

528 The second test on measured gravity data is carried out using the published data by Last & Kubik
 529 (1983) over the Woodlawn massive sulfide ore body, New South Wales, Australia. The residual anomaly
 530 of the area consisting of 61 data measurements, sampled every 5 m, is digitized from Last & Kubik

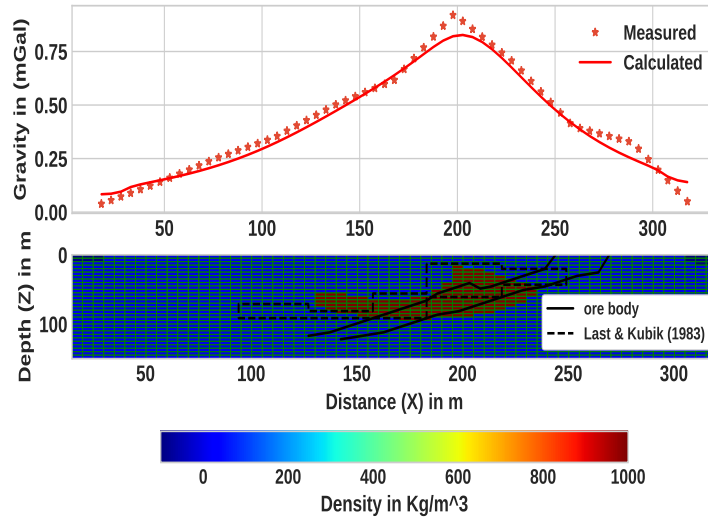


Figure 13: An observed gravity anomaly over the Woodlawn ore body, New South Wales (After Last & Kubik (1983)) and the its inversion result. The digitized data (star marks) is shown together with calculated data (solid line) on the top panel. The corresponding recovered density contrast model after 11th iteration shown on bottom panel and the ore body proved by drilling is shown with solid line. The recovered body density contrast is represented by the color scale bar.

531 (1983). The details about the data measurement and the geology of the area are discussed in Whiteley
 532 (1981). The model subsurface was divided into 61 by 30 blocks with a dimension of 5 m in both X- and
 533 Z-direction. Inverse modeling was performed with bounding constraints $\rho_{min} = -600$ and $\rho_{max} = 1000$
 534 kg/m^3 . The initial given value for ℓ_o is 0.6. The final solution was obtained after the 11th iteration. The
 535 reconstructed model including the final model of Last & Kubik (1983) are shown in Fig. 13. The cross-
 536 section of the ore body verified by drilling (Whiteley, 1981) is also shown in the figure. The recovered
 537 model is approximately coincident with the shape, depth of burial and density of the known ore body.
 538 Areas of misfits in the current and previous works are believed to be caused by the termination of the
 539 original data at both ends before it reaches the background level. Thus, this can be additional evidence
 540 that the presented method can be successfully applied to real data.

541 5 Conclusion

542 We have presented an alternative gravity inversion method that can produce compact and sharp images
 543 by using the L_0 -norm stabilizing functional that helps to model geological features with non-smooth,
 544 blocky geologic bodies. Physical parameter inequality constraints, and depth weighting are integrated

545 into the procedure. The method also incorporates an auto-adaptive regularization technique, which auto-
546 matically determines a suitable regularization parameter at every iteration, and an error weighting func-
547 tion that helps to improve both the stability and convergence of the method. One of the strongest sides of
548 the proposed auto-adaptive regularization and error weighting matrix is that they are not dependent on a
549 priori knowledge of the noise level. Because of that, the method can yield reasonable results even when
550 the noise level of the data is not known properly. We implemented a combined stopping criteria and
551 illustrated its effectiveness to terminate the iterative inversion process after an optimal number of steps.
552 To illustrate the efficiency and the capacity of the proposed procedure numerous synthetic tests were
553 done. From these, four synthetic examples were presented. According to the results from these syn-
554 thetic examples, the method can be applied for multi-source localized bodies located at different depths
555 and having different geometries with sharp features. Furthermore, the method proved to be efficient in
556 resolving causative bodies both vertically and laterally and produced compact and sharp images. The
557 obtained results also indicate that the method behaves well with different noise levels embedded in the
558 data and still retains its stability. This can confirm the robustness and stability of the developed inversion
559 method for different noise levels. The method was also tested on measured gravity data. We obtained
560 geologically acceptable models and the results showed that our approach is effective and reliable. From
561 a computational point of view, the method is efficient and can be easily run on a personal computer just
562 in a few seconds. In conclusion, the developed method is advantageous in such that it is stable, efficient,
563 and resolves sharp subsurface features with acceptable resolving capacity. In geophysical exploration
564 gravity data are more often used to image complex 3D structures of the subsurface, hence further de-
565 velopment of the method to 3D is crucial. Accordingly, future work will deal with the extension of the
566 presented method to 3D gravity inversion algorithm.

567 **Data Availability**

568 The authors confirm that the real data supporting the findings of this study are available within the
569 articles:

570 1. Green, W. R. (1975) Inversion of gravity profiles by use of a backus-gilbert approach. *Geophysics*,
571 40, 763–772 . and its supplementary material.

572 2. Last, B. and Kubik, K. (1983) Compact gravity inversion. *Geophysics*, 48, 713–721. and its
573 supplementary material

574 **Author contributions**

575 MGG developed the methodology; EL supervised the research work; MGG wrote the manuscript draft;
576 EL reviewed and edited the manuscript.

577 **Competing interests**

578 The authors declare that they have no known competing financial interests or personal relationships that
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723 **Figure Captions**

724 Figure 1. A 2-D model of the subsurface under a gravity profile. Gravity stations (X_i) are located
725 at the centers of the blocks, indicated by the ∇ symbols.

726 Figure 2. Comparison of the minimum support stabilizing function for different values of ε .

727 Figure 3. The first synthetic model and the result of the inversion.

728 Figure 4. Inversion results, using different ℓ_o values, for the first synthetic model given in Fig.
729 3(a).

730 Figure 5. The second example synthetic model and the corresponding inversion result.

731 Figure 6. The third synthetic model that comprises two dikes at various depths with the density
732 contrast that amounts to 1000 kg/m^3 and the corresponding gravity data

733 Figure 7. Inversion results of the third synthetic example in Fig. 6.

734 Figure 8. The fourth synthetic model example and the corresponding inversion result.

735 Figure 9. Inversion result obtained using only the commonly used criterion ($|misfit^{k-1} -$
736 $misfit^k|$) and the corresponding data fit (upper panels) for the synthetic example in Fig. 8(a).
737 The obtained density model shows that compact and sharp model is not approximately achieved
738 due to the termination before the iterative procedure has reached convergence.

739 Figure 10. The progression of $misfit$ and smv in the course of the iteration during the inversion
740 of the fourth example synthetic data.

741 Figure 11. Late iteration termination (at 16th iteration) inversion result and the corresponding
742 $misfit$ and smv variation with iteration number for the fourth example in Fig. 8(a).

743 Figure 12. The observed gravity anomaly over Guichon Creek batholith in south-central British
744 Columbia (after Green 1975) and its inversion result. Digitized data (star marks) with calculated
745 data (solid line) shown on the top panel. Corresponding recovered density contrast model after
746 9th iteration shown on the bottom panel. The recovered body density contrast is represented by
747 the color scale bar. For comparison, the results obtained by Ager (1973), which was obtained
748 from drilling and Green (1975) are also presented.

749 Figure 13. An observed gravity anomaly over the Woodlawn ore body, New South Wales (After
750 Last and Kubik (1983)) and the its inversion result. The digitized data (star marks) is shown
751 together with calculated data (solid line) on the top panel. The corresponding recovered density
752 contrast model after 11th iteration shown on bottom panel and the ore body proved by drilling is
753 shown with solid line. The recovered body density contrast is represented by the color scale bar