

1 **Measurement report: Molecular-level investigation of**  
2 **atmospheric cluster ions at the tropical high-altitude research**  
3 **station Chacaltaya (5240 m a.s.l.) in the Bolivian Andes**

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## 29 Abstract

30 Air ions are the key components for a series of atmospheric physicochemical  
31 interactions, such as ion-catalyzed reactions, ion-molecule reactions, and ion-induced  
32 new particle formation. They also control atmospheric electrical properties with effects  
33 on global climate. We performed molecular-level measurements of cluster ions at the  
34 high-altitude research station Chacaltaya (CHC; 5240 m a.s.l.), located in the Bolivian  
35 Andes, from January to May 2018 using an atmospheric pressure interface time-of-  
36 flight mass spectrometer. The negative ions mainly consisted of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$ ,  
37  $(\text{HNO}_3)_{0-2}\cdot\text{NO}_3^-$ ,  $\text{SO}_5^-$ ,  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$ , malonic acid-derived, and  
38  $\text{CHO}/\text{CHON}\cdot(\text{HSO}_4^-/\text{NO}_3^-)$  cluster ions. Their temporal variability exhibited distinct  
39 diurnal and seasonal patterns due to the changes in the corresponding neutral species'  
40 molecular properties (such as electron affinity and proton affinity) and concentrations  
41 resulting from the air masses arriving at CHC from different source regions. The  
42 positive ions were mainly composed of protonated amines and organic cluster ions, but  
43 exhibited no clear diurnal variation.  $\text{H}_2\text{SO}_4\text{-NH}_3$  cluster ions likely contributed to the  
44 new particle formation process, particularly during wet-to-dry transition period and dry  
45 season when CHC was more impacted by air masses originating from source regions  
46 with elevated  $\text{SO}_2$  emissions. Our study provides new insights into the chemical  
47 composition of atmospheric cluster ions and their role in new particle formation in the  
48 high-altitude mountain environment of the Bolivian Andes.

## 49 1 Introduction

50 Air ions regulate the electrical properties of the atmosphere by serving as carriers of  
51 electrical charges (Williams, 2009). They also play an important role in atmospheric  
52 chemistry by participating/catalyzing ion-molecule reactions and ion-induced new  
53 particle formation (NPF, Hirsikko et al., 2011). The formation of tropospheric ions is  
54 initiated through simple-structured ions, such as  $\text{O}^+$ ,  $\text{N}_2^+$ ,  $\text{O}^-$ , and  $\text{O}_2^-$ , mainly from  
55 radioactive decay in the soil (e.g., radon and gamma radiation), thunderstorm activity  
56 (lightning), and galactic cosmic rays (GCR). These ions can transfer their charges to  
57 other compounds, leading to the subsequent production of an assortment of ions, such  
58 as the bisulfate ion ( $\text{HSO}_4^-$ ), nitrate ion ( $\text{NO}_3^-$ ), hydronium ion ( $\text{H}_3\text{O}^+$ ), and ammonium  
59 ion ( $\text{NH}_4^+$ ; Smith and Spang, 1995; Hirsikko et al., 2011). Depending on their sizes,  
60 air ions are usually classified into cluster ions (diameter  $\leq 1.6$  nm) that are charged  
61 molecules or molecular clusters, and charged particles (diameter  $> 1.6$  nm; Hirsikko et  
62 al., 2005, 2011; Komppula et al., 2007).

63 Cluster ions exist almost always in the troposphere and can undergo frequent ion-  
64 molecule reactions during their lifetime ( $\sim 100$  seconds; Manninen et al., 2010; Hirsikko  
65 et al., 2011). Their chemical composition, in addition to the initial ionization, also  
66 depends on the concentrations of the parent neutral species (Eisele, 1986). Bianchi et  
67 al. (2017) showed that the diurnal cycle of negative organic ions followed the variations  
68 of their neutral molecules' concentrations in a boreal forest, since the higher

69 concentrations of neutral molecules would result in a larger probability of them being  
70 charged. Moreover, molecular properties of the neutral species, such as electron affinity  
71 (EA) and proton affinity (PA), are also important for determining cluster ion  
72 composition. Cluster ions derived from molecules with higher EA (e.g., HSO<sub>4</sub> and NO<sub>3</sub>)  
73 or PA (e.g., trimethylamine (C<sub>3</sub>H<sub>9</sub>N) and pyridine (C<sub>5</sub>H<sub>5</sub>N)) tend to obtain the ambient  
74 negative or positive charge, respectively (Ferguson and Arnold, 1981; Hirsikko et al.,  
75 2011). Because of the strong EA or PA, it is almost unlikely that the ions derived from  
76 those molecules will further transfer their charges to other neutral compounds via ion-  
77 molecule reactions. Thus, these negative (HSO<sub>4</sub><sup>-</sup> and NO<sub>3</sub><sup>-</sup>) and positive (C<sub>3</sub>H<sub>10</sub>N<sup>+</sup> and  
78 C<sub>5</sub>H<sub>6</sub>N<sup>+</sup>) ions are usually more abundant than other ions in the atmosphere (Eisele, 1986;  
79 Ehn et al., 2010; Bianchi et al., 2017; Frege et al., 2017). In contrast, charge transfer  
80 occurs more easily for ions derived from neutral species of lower EA or PA.

81 Atmospheric cluster ions can contribute to new particle formation (NPF) via ion-  
82 induced nucleation (Yu, 2010). Since the discovery of this mechanism in the first cloud  
83 chamber study in the early 1900s (Wilson, 1911), ion-induced nucleation has been  
84 known as an important source of atmospheric aerosol particles. Recently, a series of  
85 chamber studies conducted at the CLOUD (Cosmics Leaving Outdoor Droplets) facility  
86 at CERN (the European Centre for Nuclear Research) have shown that aerosol  
87 nucleation rates are substantially enhanced in the presence of some specific cluster ions,  
88 such as sulfuric acid – ammonia (H<sub>2</sub>SO<sub>4</sub> – NH<sub>3</sub>) cluster ions (Kirkby et al., 2011;  
89 Schobesberger et al., 2015), pure H<sub>2</sub>SO<sub>4</sub> cluster ions (Kirkby et al., 2011), and organic  
90 cluster ions (Kirkby et al., 2016). Field measurements have also suggested the important  
91 role of atmospheric ions in ion-induced nucleation (Manninen et al., 2010; Hirsikko et  
92 al., 2011; Rose et al., 2018; Jokinen et al., 2018; Yan et al., 2018; Beck et al., 2021).  
93 Among them, the onsets of high-altitude NPF events, compared to those occurring in  
94 the lower troposphere, are often associated with more abundant cluster ions (Lee et al.,  
95 2003; Venzac et al., 2008; Boulon et al., 2010; Sellegri et al., 2019). Such increases are  
96 due to the higher GCR intensity and lower condensation sink (CS) in the high-altitude  
97 regions. As a result, potentially larger contributions of cluster ions to aerosol formation  
98 would be expected (Smith and Spang, 1995; Hirsikko et al., 2011).

99 However, the molecular-level understanding of ambient cluster ions and their influence  
100 on NPF in high-altitude environments (in the troposphere) is still very limited. Two  
101 mountaintop studies in the Alps show that, depending on the air mass origins, NPF  
102 could be triggered by sulfuric acid-ammonia clusters, or nitrate (or sulfuric acid)  
103 clustering with highly oxygenated organic molecules (Bianchi et al., 2016; Frege et al.,  
104 2017). Another study in the Himalayas found that NPF was mainly driven by organic  
105 vapors of biogenic origin (Bianchi et al., 2021). Recently, frequent and intensive NPF  
106 events were observed at the high-altitude research station Chacaltaya (CHC; 16.3505°  
107 S, 68.1314° W; 5240 m a.s.l.) located in the Bolivian Andes (Rose et al., 2015a), but  
108 the exact mechanism and the role of cluster ions in aerosol nucleation process remain  
109 unclear. Therefore, a detailed investigation of cluster ions at CHC, including their

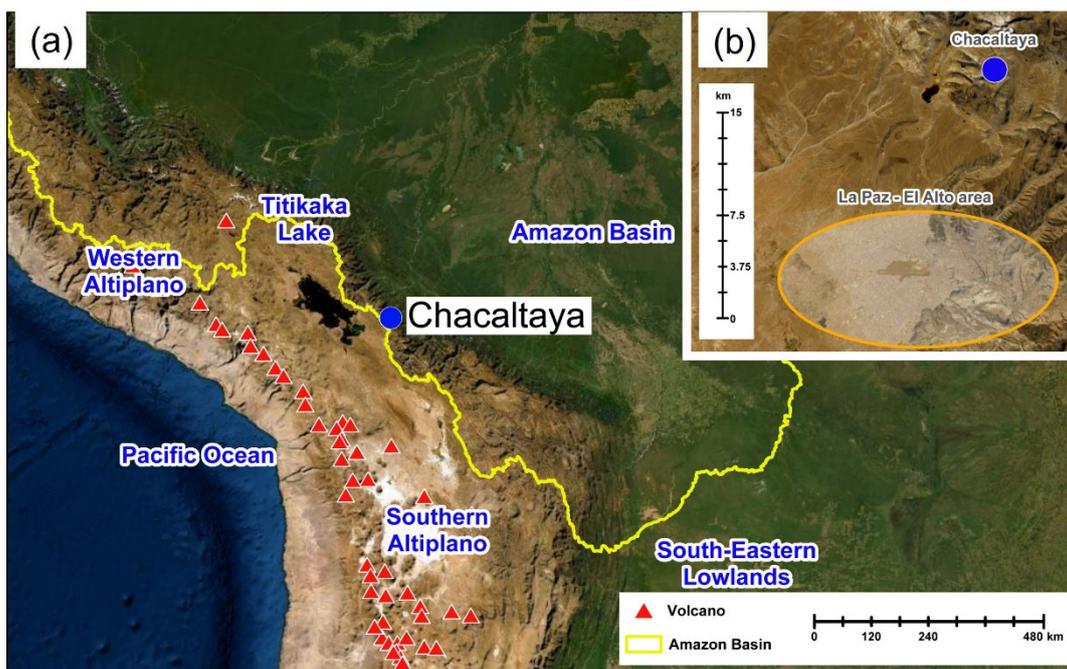
110 molecular composition, temporal variation (diurnal and seasonal), and source regions,  
111 is needed in order to understand their role in atmospheric processes such as NPF in the  
112 study regions.

113 Here we present measurements of atmospheric ions from January to May 2018 at CHC.  
114 The dataset is part of the Southern hemisphere high ALTitude Experiment on particle  
115 Nucleation And growth (SALTENA) field experiment campaign (Bianchi et al., 2022).  
116 During the study period, the sampled air masses originated from various source regions,  
117 such as the Amazon Basin to the east and the Altiplano and the Pacific Ocean to the  
118 west (Fig. 1a; Aliaga et al., 2021). Temporal evolution (diurnal and/or seasonal  
119 variations) of both negative and positive ion composition are investigated, and their  
120 potential connections with source regions and NPF are discussed. Our study thus adds  
121 important observational information on a better understanding of atmospheric ions and  
122 provides new insights into their role in high-altitude NPF in the troposphere of the  
123 Bolivian Andes.

## 124 **2 Methods**

### 125 **2.1 Measurement site description**

126 The high-altitude research station CHC is ~140 m below the summit of Mount  
127 Chacaltaya (5380 m a.s.l.) with an open view to the south and west (Andrade et al.,  
128 2015). The La Paz – El Alto metropolitan area (with 1.7 million inhabitants) is ~1 – 1.6  
129 km lower (in altitude) and ~15 km south of CHC (Fig. 1b). The seasonal meteorological  
130 conditions at CHC depend on the cycle between the wet (November to March; wet-to-  
131 dry transition period in April) and dry (May to September; dry-to-wet transition period  
132 in October) seasons driven by large-scale tropical circulation (Rose et al., 2015a;  
133 Bianchi et al., 2022). This pattern also affects the source regions of air masses arriving  
134 at CHC (Aliaga et al., 2021). Additionally, due to the strong diurnal cycle of the  
135 planetary boundary layer (PBL) height and the thermally-induced winds in the  
136 mountainous terrain, CHC is often affected by polluted PBL transported from the La  
137 Paz – El Alto metropolitan area during daytime (Wiedensohler et al., 2018) whereas at  
138 night CHC is located in the residual layer or tropical free troposphere (Rose et al., 2015b;  
139 Chauvigné et al., 2019).



140

141 Figure 1 (a) True-color satellite image showing the location of CHC (blue circle) and  
 142 its surrounding area. The yellow line presents the boundary of the Amazon Basin. Red  
 143 triangles denote the volcanoes in this area. (b) A zoomed-in true-color satellite image  
 144 showing the distance between CHC and the La Paz – El Alto metropolitan area (orange  
 145 circle). Image sources: Esri, DigitalGlobe, GeoEye, i-cubed, USDA FSA, USGS, AEX,  
 146 Getmapping, Aerogrid, IGN, IGP, swisstopo, and the GIS User Community.

## 147 2.2 Instrumentation

148 All the instruments involved in this study were installed in a temperature-controlled  
 149 measurement room (at  $\sim 25^\circ\text{C}$ ). All data are reported in local time (UTC-4).

### 150 2.2.1 Measurements of atmospheric cluster ions

151 The composition of cluster ions was measured by an atmospheric pressure interface  
 152 time-of-flight mass spectrometer (APi-TOF, Aerodyne Research Inc. & Tofwerk AG).  
 153 The APi-TOF consists of an atmospheric pressure interface (APi) module and a time-  
 154 of-flight (TOF) mass spectrometer. The APi module allows the instrument to sample  
 155 ions from ambient air directly. The positive or negative ions within the sampled airflow  
 156 are focused and guided by two quadrupoles and an ion lens, with a gradually decreasing  
 157 pressure (from atmospheric pressure to  $\sim 10^{-4}$  mbar), before entering the TOF mass  
 158 spectrometer ( $\sim 10^{-6}$  mbar). A more detailed description of this instrument is given in  
 159 Junninen et al. (2010). In this study, ambient air was sampled through a  $\sim 1.5$  m stainless  
 160 steel tube with a total sample flow of 14 standard liters per minute (SLPM) to ensure  
 161 laminar flow during sampling, and 0.8 SLPM of the total flow entered the APi-TOF.

162 During the wet season, the APi-TOF was first operated in negative mode to measure  
 163 negative cluster ions (January), and then switched to positive mode to measure positive

164 cluster ions (February to March). During the wet-to-dry transition period (April) and  
165 dry season (May), the instrument was changed back to negative mode to investigate the  
166 potential seasonality of negative ion composition. It is important to note that, similar to  
167 Frege et al. (2017), for better characterization of the connection between cluster ions  
168 and NPF, we only included the ion data observed under the cloud-free condition in this  
169 study (to avoid influence from, e.g., lightning activity). CHC was considered to be  
170 affected by clouds when relative humidity (RH) exceeded 95 %, as suggested by a  
171 previous study at the same location (Rose et al., 2015a). From January to May 2018,  
172 the proportions of cloud-free hours to the total measurement time were 72 %, 78 %,  
173 79 %, 98 %, and 98 %, respectively (Fig. S1).

## 174 **2.2.2 Measurements of H<sub>2</sub>SO<sub>4</sub> and oxidized organic molecules**

175 Concentrations of H<sub>2</sub>SO<sub>4</sub> and oxidized organic molecules (OOM) were measured using  
176 a nitrate ion (NO<sub>3</sub><sup>-</sup>) based chemical ionization atmospheric pressure interface time-of-  
177 flight mass spectrometer (CI-APi-TOF, Aerodyne Research Inc. & ToFwerk AG;  
178 Jokinen et al., 2012). The instrument is a combination of the APi-TOF and a chemical  
179 ionization (CI) unit, which has been widely used to measure H<sub>2</sub>SO<sub>4</sub> in the atmosphere  
180 (Jokinen et al., 2012; Bianchi et al., 2016; Zha et al., 2018). In this study, a soft X-ray  
181 source (L9490, Hamamatsu) was used to charge nitric acid (HNO<sub>3</sub>) in a sheath flow of  
182 20 SLPM to produce the reagent ion, NO<sub>3</sub><sup>-</sup>. H<sub>2</sub>SO<sub>4</sub> and OOM in the sample flow (10  
183 SLPM) were then charged by either proton transfer or the formation of an adduct with  
184 the reagent ion during the ~200 ms residence time in the CI unit. A calibration factor of  
185  $1.5 \times 10^{10} \text{ cm}^{-3}$  for H<sub>2</sub>SO<sub>4</sub> was determined (with sampling loss corrected) according to  
186 the approach by Kürten et al. (2012). The same calibration coefficient was adopted for  
187 determining OOM concentrations in this study, which could result in an  
188 underestimation of their concentrations due to a lower charging efficiency of OOM than  
189 H<sub>2</sub>SO<sub>4</sub> by NO<sub>3</sub><sup>-</sup> (Hyytinen et al., 2015).

## 190 **2.2.3 Auxiliary measurements**

191 The number concentration and size distribution of atmospheric ions and neutral  
192 particles were measured with a neutral cluster and air ion spectrometer (NAIS, Airel  
193 Ltd., Mirme and Mirme, 2013). The instrument can detect naturally charged air ions  
194 and total particles with mobility diameters from 1.4 – 50 nm and 3 – 50 nm, respectively.  
195 It is important to note that the size range of the NAIS was corrected based on a side-  
196 by-side comparison with an updated version of NAIS (designed for measurements  
197 under low pressure environments; Mirme et al., (2010)) at CHC. Thus, the size range  
198 of detection was slightly different from the traditionally reported ranges (i.e., 0.8 – 42  
199 nm and 2 – 42 nm, respectively; Manninen et al., 2010). The details of the instrument  
200 used can be found in Rose et al. (2017).

201 Particle number size distributions between 10 to 500 nm were measured by a Mobility  
202 Particle Size Spectrometer (MPSS; Wiedensohler et al., 2012), and the data was used

203 for calculating the CS, which represents the loss rate of condensing vapors and cluster  
204 ions on pre-existing particles (Kulmala et al., 2001).

205 Meteorological parameters, such as temperature, RH, and global radiation, were also  
206 measured simultaneously at CHC. Detailed descriptions can be found in Bianchi et al.  
207 (2022).

## 208 **2.3 Simulation of air mass origin and history**

209 To understand the source regions and transport pathways of the air masses arriving at  
210 CHC, we used the results of air mass history analysis obtained from FLEXPART-WRF  
211 simulations described in Aliaga et al. (2021). In brief, a Lagrangian transport and  
212 dispersion model (FLEXPART-WRF; version 3.3.2; Brioude et al., 2013) was used to  
213 calculate the air mass history during the campaign period. The backward simulation  
214 was driven by the high spatial and temporal resolution meteorological output from the  
215 Weather Research and Forecasting model (WRF; version 4.0.3; Skamarock et al., 2019).  
216 In the simulation, twenty thousand particles were continuously released every hour  
217 from a 10 m deep layer (0 – 10 m a.g.l.) at CHC and traced back in the atmosphere for  
218 96 hours. The output of the FLEXPART-WRF is the source-receptor relationship (SRR,  
219 in seconds), which is calculated for each geographical grid cell included in the  
220 simulation. The SRR value depends on the particle’s residence time and the number of  
221 particles in the output grid cells. Clustering analysis was conducted by applying a series  
222 of pretreatments (e.g., log-polar grid transformation and grid cell pre-processing) and a  
223 k-means clustering algorithm (Lloyd, 1982) to the calculated SSR dataset (see Aliaga  
224 et al. (2021) for more details).

### 225 **2.3.1 Major air mass pathways**

226 Six air mass pathways (PW) representing air masses arriving at CHC were determined  
227 from the clustering analysis. They are named based on their clock positions from CHC  
228 (e.g., 03\_PW indicates the pathway with its centroid located at the 3 o’clock direction  
229 (east, 90°) of CHC, Fig. 2a). Characteristics of these air mass pathways, such as source  
230 region, transport distance, and transport time, were distinct from each other (Table 1).  
231 A detailed description of the air mass pathways and their characteristics can be found  
232 in Aliaga et al. (2021).

233 The influence of each air mass pathway on CHC varied with time, and was estimated  
234 by its SRR percentage ( $SRR[\%]_{pathway}$ ) as in equation (1):

$$235 \quad SRR[\%]_{pathway} = \frac{SRR_{pathway}}{SRR_{total}} \times 100 \quad (1)$$

236 where  $SRR_{pathway}$  and  $SRR_{total}$  are the residence time of a specific air mass pathway and  
237 in total (96 hours = 345600 seconds) in the simulation, respectively.

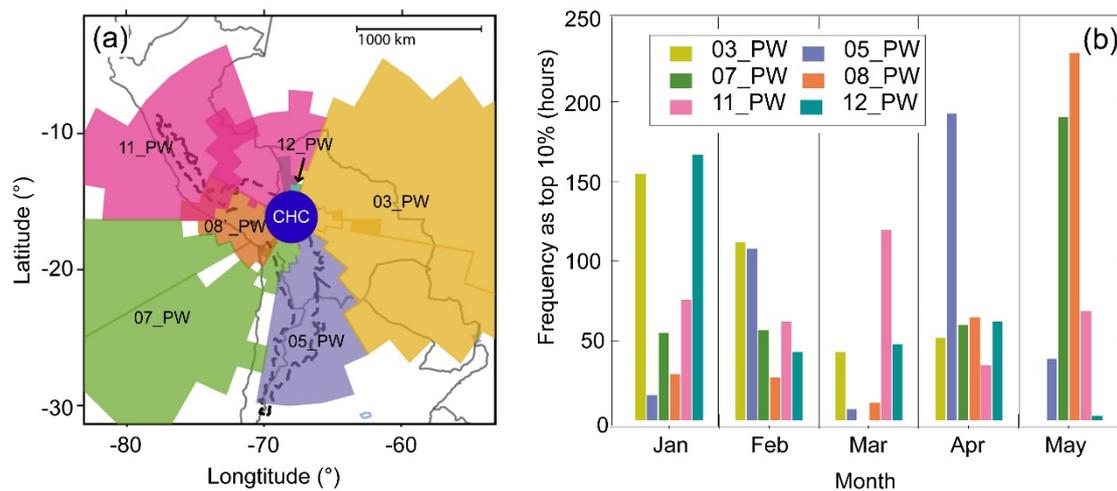
238 Table 1. Overview of the six air mass pathways extracted from Aliaga et al. (2021).

Pathway	Direction to CHC	Representative source region	Transport distance (km) <sup>1</sup>	Transport time (hour)
			Median (25 – 75 %)	Median (25 – 75 %)
03_PW	East	Amazon Basin and Eastern/South-Eastern Lowlands	518 (413–608)	51 (45-57)
05_PW	South and Southeast	South-Eastern Lowlands and Southern Altiplano	428 (303-567)	45 (36-52)
07_PW	Southwest	The Pacific Ocean, coastal area, Western Altiplano, and La Paz – El Alto	721 (577-896)	54 (45-61)
08_PW	West	Western Altiplano and Titicaca lake, coastal area	238 (198-279)	36 (29-43)
11_PW	North and Northwest	Amazon Basin, Western Altiplano, coastal area	465 (326-563)	53 (46-59)
12_PW	North	Amazon Basin	76 (49 -95)	27 (21-33)

239 <sup>1</sup>Distance between CHC and each pathway's center point (see Fig. 2a).

### 240 2.3.2 Identification of representative periods for each air mass 241 pathway

242 Air mass history analysis shows that the air sampled at CHC was typically a mixture of  
243 multiple pathways. Thus, the cluster ion composition observed during the study period  
244 was often influenced by multiple source regions concurrently. To characterize the  
245 influence of every single pathway on cluster ion composition, periods when an air mass  
246 pathway exerted its largest impact on CHC (the highest 10% of its  $SRR[\%]_{\text{pathway}}$  values;  
247 Fig. 2b) during the whole study period are identified as the representative periods of  
248 the specific pathway. For instance, the representative periods of 03\_PW (covering, e.g.,  
249 the Amazon Basin) are more frequently seen during wet season (highest in January),  
250 whereas 08\_PW (covering, e.g., Altiplano region) has most of its representative periods  
251 in dry season (highest in May). Note that  $SRR[\%]_{\text{pathway}}$  of any individual pathway  
252 rarely reached 40 % due to the relatively short study period (see Fig. S2), and thus the  
253 representative periods cannot be directly identified via  $SRR[\%]_{\text{pathway}}$  values. In contrast,  
254 such representative periods were determined by using a certain threshold (e.g., > 70 %)   
255 in a previous study at CHC, which was based on a more than 6-year dataset (Koenig et  
256 al., 2021).



257

258 Figure 2 Influence of the six air pathways on CHC from January to May 2018. (a)  
 259 Horizontal profile of the air mass pathways, adapted from Aliaga et al. (2021). (b)  
 260 Frequency of the representative periods for each pathway (the highest 10% of their  
 261 corresponding  $SRR[\%]_{\text{pathway}}$ ) in different months.

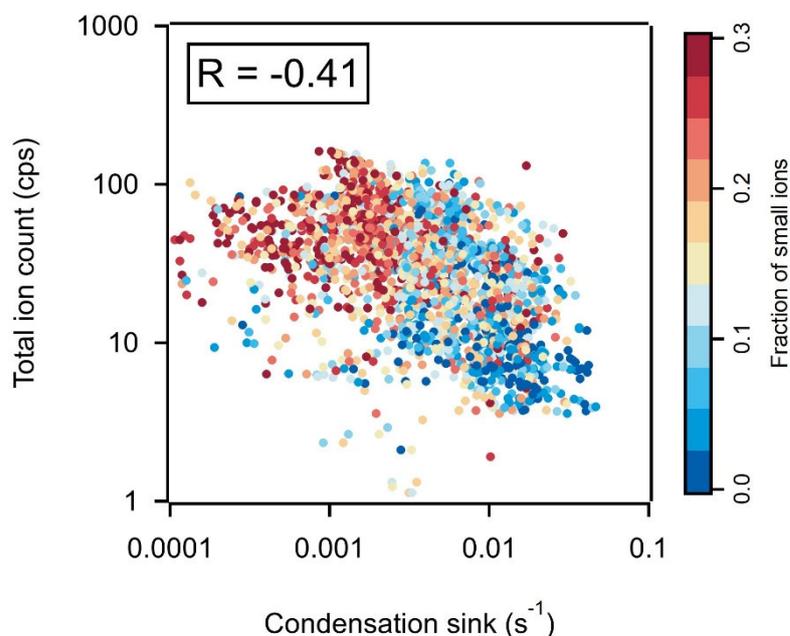
## 262 3 Results and discussion

### 263 3.1 Variation in total ion count

264 During the study period, the total ion count (TIC) observed by APi-TOF at CHC varied  
 265 between <10 and 100 counts per second (cps; Fig. 3 and Fig. S3). Similar variations in  
 266 the TIC were also observed in the long-term cluster ion measurements at the high-  
 267 altitude station Jungfraujoch in Switzerland (JFJ; 3454 m a.s.l.; Frege et al., 2017) and  
 268 were attributed to the seasonal changes of ion precursor and sink. Like the winter–  
 269 summer seasonality of the JFJ, CHC and its adjacent mountain areas are frequently  
 270 covered by snow in wet season and mostly free of snow in dry season (Bianchi et al.,  
 271 2022; Koenig et al., 2021). Thus, in contrast with the generally stable GCR flux  
 272 (primarily controlled by the decadal scale solar cycle; Shuman et al., 2015), a reduced  
 273 radioactive decay from the soil and a lower ion production rate could be expected at  
 274 CHC in wet season than in dry season.

275 However, the TIC measured by APi-TOF was significantly higher in wet season ( $41 \pm$   
 276  $23$  cps, mean  $\pm$  standard deviation) and the wet-to-dry transition period ( $56 \pm 32$  cps)  
 277 than in dry season ( $14 \pm 11$  cps; Fig. S3). Considering the slight negative correlation  
 278 (Pearson's correlation coefficient ( $R$ ):  $-0.41$ ; Fig. 3) between the TIC and CS  
 279 (representing the loss rate of condensing vapors and cluster ions on pre-existing  
 280 particles; Kulmala et al., 2001), the observed TIC fluctuation may be related (at least  
 281 partially) to the varying CS ( $\sim 1 \times 10^{-4} \text{ s}^{-1}$  in wet season to  $\sim 5 \times 10^{-2} \text{ s}^{-1}$  in the dry season).  
 282 Moreover, the cluster ions measured by APi-TOF usually account for only a small  
 283 fraction of the total atmospheric ions (Rose et al., 2018). Changes in the fraction of  
 284 small ions in total atmospheric ions can potentially lead to a fluctuation in TIC (Frege

285 et al., 2017). This is illustrated in Figure 3 that a smaller TIC determined from APi-TOF  
286 is associated with a lower fraction of smaller ions (< 2 nm) observed by NAIS (mostly  
287 cluster ions). However, for better characterization of the influences of different ion  
288 composition on CHC and their diurnal and seasonal relative changes, we normalized  
289 the observed ion signal to the TIC for APi-TOF measurements.



290

291 Figure 3 Correlation between the TIC measured by APi-TOF and condensation sink,  
292 colored by the fraction of small ions (defined as concentrations of ions with diameter  
293 <2 nm to the total ion concentrations) determined from NAIS data. Data are shown in  
294 the time resolution of 1 hour.

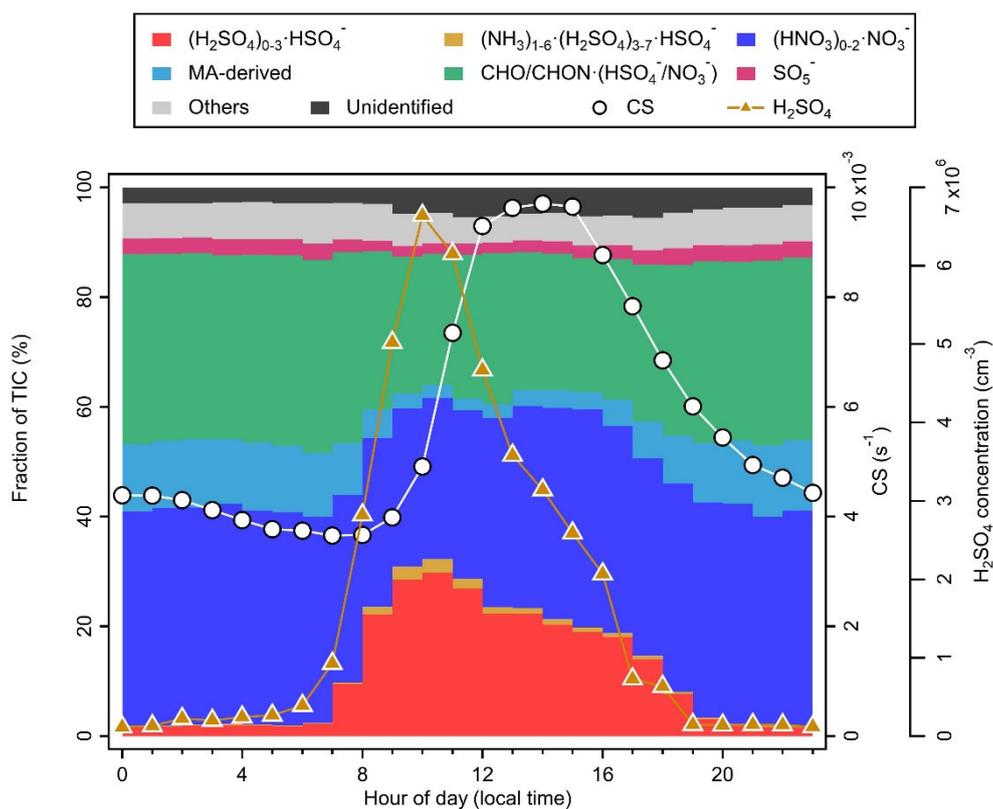
## 295 3.2 Negative ions

### 296 3.2.1 Main negative ions and their diurnal variation

297 A number of negative ions were consistently observed at CHC throughout the study  
298 period (i.e., in the wet, wet-to-dry transition, and dry seasons). Based on the chemical  
299 composition of the observed negative ions, we classified them into eight groups as  
300 follows: sulfuric acid ((H<sub>2</sub>SO<sub>4</sub>)<sub>0-3</sub>•HSO<sub>4</sub><sup>-</sup>), nitric acid ((HNO<sub>3</sub>)<sub>0-2</sub>•NO<sub>3</sub><sup>-</sup>), SO<sub>5</sub><sup>-</sup>, sulfuric  
301 acid-ammonia ((NH<sub>3</sub>)<sub>1-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>3-7</sub>•HSO<sub>4</sub><sup>-</sup>), malonic acid-derived (MA-derived;  
302 including C<sub>3</sub>H<sub>3</sub>O<sub>4</sub><sup>-</sup>, C<sub>3</sub>H<sub>4</sub>O<sub>4</sub>•NO<sub>3</sub><sup>-</sup>, and C<sub>3</sub>H<sub>4</sub>O<sub>4</sub>•HSO<sub>4</sub><sup>-</sup>), oxidized organic molecules  
303 (CHO/CHON•(HSO<sub>4</sub><sup>-</sup>/NO<sub>3</sub><sup>-</sup>)), others (other identified negative ions, such as IO<sub>3</sub><sup>-</sup>), and  
304 unidentified ions. The campaign-average diurnal variations of these eight negative ion  
305 groups are shown in Figure 4. Different diurnal patterns of each negative ion group  
306 were observed, mainly due to changes in concentrations of the parent neutral species  
307 and their unique physicochemical properties (e.g., the EA and PA of molecules;  
308 Ferguson and Arnold, 1981; Bianchi et al., 2017; Hirsikko et al., 2011).

309  $(\text{HNO}_3)_{0-1}\cdot\text{NO}_3^-$  and  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  were among the highest in the signal of all  
310 negative ion groups in all seasons, making up 37 % (whole day) and 20 % (daytime;  
311 07:00 – 19:00, and hereafter) of negative ions at CHC during the study period,  
312 respectively. The EA of their neutral molecules ( $\text{HSO}_4$  and  $\text{NO}_3$ ) is higher than that of  
313 most of the neutral species in the atmosphere, and thus hinders the direct electron  
314 transfer from  $\text{HSO}_4^-$  and  $\text{NO}_3^-$  to other molecules through ion-molecule reactions  
315 (Ferguson and Arnold, 1981). As a result, these ion groups were found to dominate  
316 negative cluster ions at CHC and also other locations, such as a number of remote sites  
317 in the United States (Eisele, 1986), a boreal forest site in Finland (Ehn et al., 2010;  
318 Bianchi et al., 2017), and the JFJ in Switzerland (Frege et al., 2017).

319 Distinct diurnal patterns were observed for the  $(\text{HNO}_3)_{0-1}\cdot\text{NO}_3^-$  and  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$   
320 ion groups.  $(\text{HNO}_3)_{0-1}\cdot\text{NO}_3^-$  exhibited a relatively flat diurnal pattern (see Fig. 4) with  
321 similar fractions at daytime (35 %) and nighttime (19:00 – 07:00; 39 %). Such a diurnal  
322 pattern could result from the high EA of the  $\text{NO}_3$  molecule (4.01 eV and an additional  
323  $\sim 1$  eV per  $\text{HNO}_3$ ; Ferguson and Arnold, 1981), and its relatively abundant (usually  
324 several ppbv) parent neutral species (e.g.,  $\text{HNO}_3$  and  $\text{N}_2\text{O}_5$ ) with multiple sources in the  
325 atmosphere (e.g., anthropogenic emission and lightning; Martin et al., 2007). In contrast,  
326  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  exhibited a strong diurnal variation. While the fraction of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$   
327 remained low (2 %) during nighttime, it started to increase after sunrise (shortly  
328 after 07:00) and reached a maximum (30 %) at around 10:00. Despite an EA comparable  
329 to that of the  $\text{NO}_3$  molecule (4.75 eV for  $\text{HSO}_4$ ; Wang et al., 2000), the strong diurnal  
330 variation of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  is a result of the photochemical production of neutral  
331  $\text{H}_2\text{SO}_4$ . The influence of neutral  $\text{H}_2\text{SO}_4$  on  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  is indicated by their  
332 similar diurnal patterns ( $R$ : 0.52; see Fig. S4a). Similarly, a higher level of  $(\text{NH}_3)_{1-6}\cdot$   
333  $(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$  was only observed with the presence of abundant  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$   
334 during daytime. It is also important to note that the decreases of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  and  
335  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$  at around noontime (12:00; see Fig. 4) coincided with an  
336 enhanced CS, indicating the influence of a higher ion sink in addition to the decrease  
337 in neutral  $\text{H}_2\text{SO}_4$  concentration (Boulon et al., 2010; Frege et al., 2017).



338

339 Figure 4 Diurnal variation of the negative ions (fraction of TIC), CS, and neutral H<sub>2</sub>SO<sub>4</sub>  
 340 concentrations at CHC, averaged over the periods when negative ions were measured  
 341 (i.e., January, April, and May 2018).

342 The MA-derived ion group is distinguished from the CHO/CHON·(HSO<sub>4</sub><sup>-</sup>/NO<sub>3</sub><sup>-</sup>) ion  
 343 group due to its abundance in total organic ions. This group of ions, mainly composed  
 344 of C<sub>3</sub>H<sub>3</sub>O<sub>4</sub><sup>-</sup>, is formed from the deprotonation of malonic acid, with a higher EA (~4.60  
 345 eV) than that of the NO<sub>3</sub> molecule (4.01 eV; Ravi Kumar et al., 2005). Malonic acid,  
 346 similar to HNO<sub>3</sub>, has multiple origins in the atmosphere, such as primary and secondary  
 347 anthropogenic sources, biogenic sources, and the degradation of larger organic  
 348 compounds (Braban et al., 2003). It is also one of the main dicarboxylic acids which  
 349 make up a substantial fraction of total carbon in aerosol particles (Kawamura and  
 350 Bikkina, 2016). However, the concentration of gas-phase malonic acid is less well  
 351 documented and is estimated to be in the range of 10<sup>7</sup> to 10<sup>9</sup> cm<sup>-3</sup>, up to three orders of  
 352 magnitude higher than the typically reported ambient H<sub>2</sub>SO<sub>4</sub> concentration (Fang et al.,  
 353 2020). Thus, the fraction of the MA-derived ion group was high during nighttime (12 %;  
 354 see Fig. 4). In contrast, its fraction decreased significantly (to 5 %) during daytime due  
 355 to the increase of (H<sub>2</sub>SO<sub>4</sub>)<sub>0-3</sub>·HSO<sub>4</sub><sup>-</sup>, which has an even higher EA (4.75 eV for HSO<sub>4</sub><sup>-</sup>;  
 356 Wang et al., 2000).

357 The CHO/CHON·(HSO<sub>4</sub><sup>-</sup>/NO<sub>3</sub><sup>-</sup>) ion group, with an overall molecular formula of C<sub>2</sub>-  
 358 <sub>15</sub>H<sub>2-26</sub>O<sub>2-13</sub>N<sub>0-2</sub>·NO<sub>3</sub><sup>-</sup>/HSO<sub>4</sub><sup>-</sup>, constituted a significant fraction of negative ions (31 %)

359 at CHC. These organic ions are formed through the adduction between primary charge  
360 carriers, such as  $\text{HSO}_4^-$  and  $\text{NO}_3^-$ , and neutral OOM. These OOM are likely the  
361 oxidation products of the volatile organic compounds (VOC) from the Amazon, the  
362 Altiplano, and the adjacent La Paz – El Alto metropolitan area. Their chemical  
363 composition is potentially affected by the changing air pathways covering different  
364 VOC source regions (Aliaga et al., 2021), and by the different conditions during  
365 daytime/nighttime due to the evolution of different atmospheric layers (Beck et al.,  
366 2022). While the diurnal variation was relatively small (34 % for the nighttime and 27 %  
367 for the daytime), the ion composition of  $\text{CHO/CHON}\cdot(\text{HSO}_4^-/\text{NO}_3^-)$  could also be  
368 significantly different between daytime and nighttime due to the availability of the  
369 charging ions (see more discussions in Section 3.2.2). A previous study from a boreal  
370 forest shows that organic ions are mainly composed of  $\text{CHO/CHON}\cdot\text{NO}_3^-$  during  
371 nighttime, and that the fraction of  $\text{CHO/CHON}\cdot\text{HSO}_4^-$  increases with the  $\text{HSO}_4^-$  signal  
372 during daytime (Bianchi et al., 2017). This is also shown by the slightly positive  
373 correlation between the  $\text{CHO/CHON}\cdot\text{HSO}_4^-$  signal fraction and the total neutral OOM  
374 concentration during daytime ( $R$ : 0.25; see Fig. S4b), whereas no clear dependence was  
375 found between  $\text{CHO/CHON}\cdot(\text{HSO}_4^-/\text{NO}_3^-)$  and the total neutral OOM concentration.

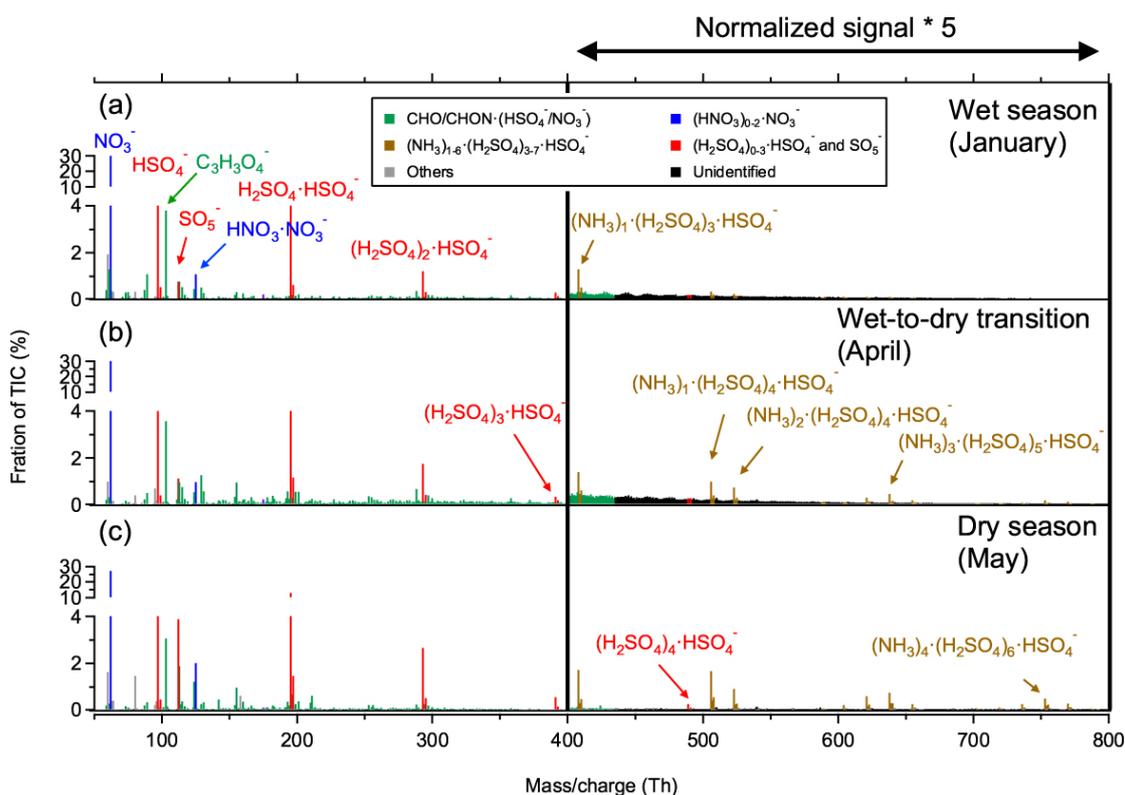
376 The  $\text{SO}_5^-$  ion group, consisting of  $\text{SO}_5^-$  ions and/or  $\text{O}_2\cdot\text{SO}_3^-$  cluster ions (Bork et al.,  
377 2013; Frege et al., 2017), exhibited a lower fraction (<5 %) than the aforementioned  
378 ion groups during the study period. Similar to that of the  $(\text{HNO}_3)_{0-1}\cdot\text{NO}_3^-$  ion group, no  
379 diurnal pattern was evident for the  $\text{SO}_5^-$  ion group. This may be the result of its different  
380 major formation pathways during daytime and nighttime (Bork et al., 2013; Frege et al.,  
381 2017). Daytime production of  $\text{SO}_5^-$  ions is likely associated with photo-oxidation of  
382  $\text{SO}_2$  (similar to the formation pathway of  $\text{H}_2\text{SO}_4$ ; Ehn et al., 2010; Schobesberger et al.,  
383 2015). This is shown in the positive correlation ( $R$ : 0.46 for daytime data in Fig. S4c)  
384 between the neutral  $\text{H}_2\text{SO}_4$  concentration and the signal fraction of the  $\text{SO}_5^-$  ion group.  
385 During nighttime, however, the  $\text{SO}_5^-$  ion group is mainly composed of  $\text{O}_2\cdot\text{SO}_3^-$  cluster  
386 ions, which are possibly formed via the oxidation of  $\text{SO}_2$  with  $\text{O}_3^-$  (producing  $\text{SO}_3^-$ ),  
387 and subsequent addition of  $\text{O}_2$  (Bork et al., 2013).

### 388 3.2.2 Seasonalities of negative ions

389 For better seasonality comparison at high-altitude CHC, we calculated the average mass  
390 spectra of the negative ion groups for each season (Fig. 5 for daytime and Fig. S5 for  
391 nighttime). Distinct seasonalities (wet season, wet-to-dry transition period, and dry  
392 season) were found for the majority of the negative ion groups at CHC, including  
393  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$ ,  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$ ,  $\text{SO}_5^-$ , and organic cluster ions, as shown  
394 in the averaged daytime mass spectra (Fig. 5; more detailed reason will be discussed  
395 below). However, the signals of some other negative ion groups, e.g., MA-derived ions  
396 and  $(\text{HNO}_3)_{0-2}\cdot\text{NO}_3^-$ , were generally stable (with differences  $\leq 20$  %) across the seasons.  
397 Such unclear seasonalities can be attributed to the high EA (Ferguson and Arnold, 1981;  
398 Ravi Kumar et al., 2005) and/or the stability of the parent neutral species (Martin et al.,

399 2007; Kerminen et al., 2000; Bikkina et al., 2021). Similar patterns can also be found  
 400 in the average nighttime mass spectra among the seasons (Fig. S5).

401  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  group exhibited much higher contribution in dry season (May) than  
 402 in wet season (January) and wet-to-dry transition period (April). The daytime fraction  
 403 of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  increased continuously from 16 % in wet, 20 % in wet-to-dry  
 404 transition period, to 30 % in dry season. The maximum number of  $\text{H}_2\text{SO}_4$  molecules  
 405 increased concurrently from 2 to 4 in the cluster ions (i.e., from  $(\text{H}_2\text{SO}_4)_2\cdot\text{HSO}_4^-$  to  
 406  $(\text{H}_2\text{SO}_4)_4\cdot\text{HSO}_4^-$ ). Similar trends were also found for other  $\text{H}_2\text{SO}_4$ -related ions, such as  
 407 the  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$  and  $\text{SO}_5^-$  during daytime (Fig. 5).



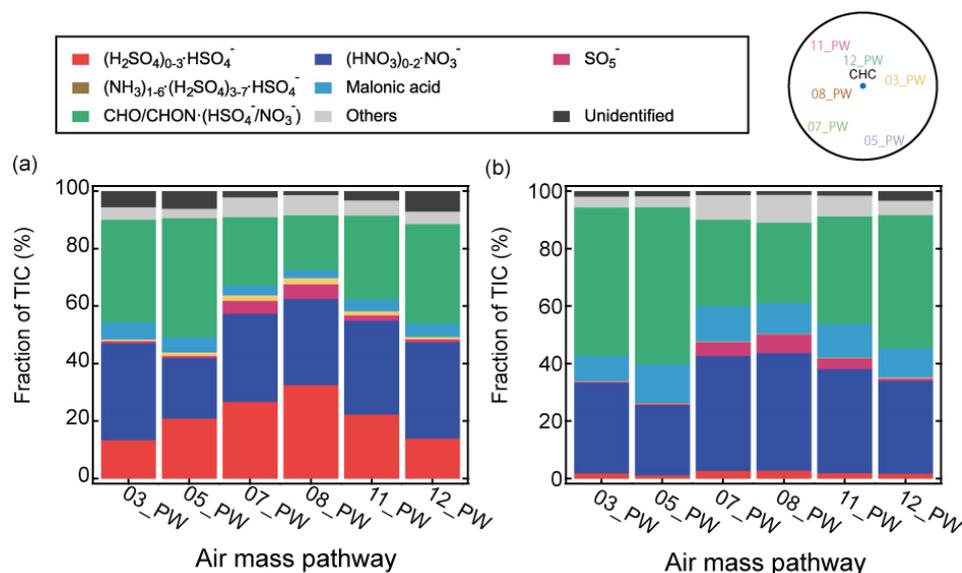
408  
 409 Figure 5 Mass spectra of negative ions at CHC averaged between 07:00 – 19:00 in (a)  
 410 wet season (January), (b) wet-to-dry transition period (April), and (c) dry season (May).  
 411 The normalized signal intensities from 400 Th to 800 Th are multiplied by a factor of 5  
 412 for better visualization.

413 The seasonal variations of the aforementioned  $\text{H}_2\text{SO}_4$ -related ion groups are likely due  
 414 to the changes in neutral  $\text{H}_2\text{SO}_4$  (Fig. S6) linked to the changing synoptic-scale wind  
 415 patterns carrying different air masses with varying  $\text{SO}_2$  (Bianchi et al., 2022). The air  
 416 mass pathways 07\_PW and 08\_PW, covering the Western and Northern Altiplano  
 417 plateau (see Table 1 and Fig. 2a), where active volcanic degassing of  $\text{SO}_2$  has been  
 418 reported (Moussallam et al., 2017; Carn et al., 2017), had their largest influence on  
 419 CHC in dry season (i.e., May; see Fig. 2). The corresponding daytime fractions of  
 420  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  from these two pathways (Fig. 6a) were also the highest (27 % and  
 421 32 %, respectively). In contrast, air mass pathways 03\_PW and 12\_PW, originating in

422 the Amazon Basin and Eastern/South-Eastern Lowlands, exerted their most significant  
423 impact on CHC in wet season (i.e., January) with lower daytime fractions of  $(\text{H}_2\text{SO}_4)_0\text{-}$   
424  $3\cdot\text{HSO}_4^-$  (13 % and 14 %, respectively). The low fractions of  $\text{H}_2\text{SO}_4$ -related cluster ions  
425 in wet season are also consistent with the lower  $\text{SO}_2$  level in the Amazon Basin  
426 compared to the Altiplano plateau (Andreae et al., 1990). As for the wet-to-dry  
427 transition period (i.e., April), 05\_PW covering both the South-Eastern Lowlands and  
428 Southern Altiplano plateau (where volcanic degassing is also significant; Carn et al.,  
429 2017) had an evident influence on CHC, resulting in a substantial level of  $\text{H}_2\text{SO}_4$ -  
430 related cluster ions (21 % for daytime). It is also noted that, because of the much lower  
431 nocturnal neutral  $\text{H}_2\text{SO}_4$  concentrations, the nighttime fractions of  $\text{H}_2\text{SO}_4$ -related  
432 cluster ions in all air mass pathways (Fig. 6b) were generally low (< 3 %) and no clear  
433 seasonality was found.

434 The organic cluster ion group exhibited a distinct seasonal variation than the  $(\text{H}_2\text{SO}_4)_0\text{-}$   
435  $3\cdot\text{HSO}_4^-$ . The signal fraction of organic cluster ions was higher in wet season (31 % for  
436 daytime and 32 % for nighttime) than in dry season (23 % for daytime and 27 % for  
437 nighttime; Fig. 5 and Fig. S5), but it was highest for the wet-to-dry transition period  
438 (46 % for daytime and 52 % for nighttime; see Fig. 6).

439 The seasonal changes of organic cluster ions could be due to the combined effect of  
440 different meteorological conditions and VOC from different air mass origins (see Fig.  
441 2). The air masses that originated from the Amazon Basin and Lowlands (03\_PW and  
442 12\_PW) showed their largest impact on CHC in wet season (i.e., January). They  
443 contained higher fractions of organic cluster ions, which were 35 % and 34 % for  
444 daytime, and 50 % and 45 % for nighttime, respectively (Fig. 6). In dry season (i.e.,  
445 May), however, the changes in air mass origin towards the Altiplano plateau and the  
446 Pacific Ocean led to a lower content of organic cluster ions. The organic cluster ion  
447 fractions for 07\_PW and 08\_PW (largest influence on CHC in dry season) in May were  
448 23 % and 19 % for daytime, and 29 % and 27 % for nighttime, respectively. As for the  
449 wet-to-dry transition period (i.e., April), due to the combined influences of biogenic  
450 and anthropogenic VOC sources from 05\_PW (evident impact on CHC in April),  
451 covering the South-Eastern Lowlands and the Southern Altiplano plateau, the  
452 corresponding organic cluster ion fractions from this air mass pathway were also the  
453 highest (41 % for daytime and 53 % for nighttime; see Fig. 6).



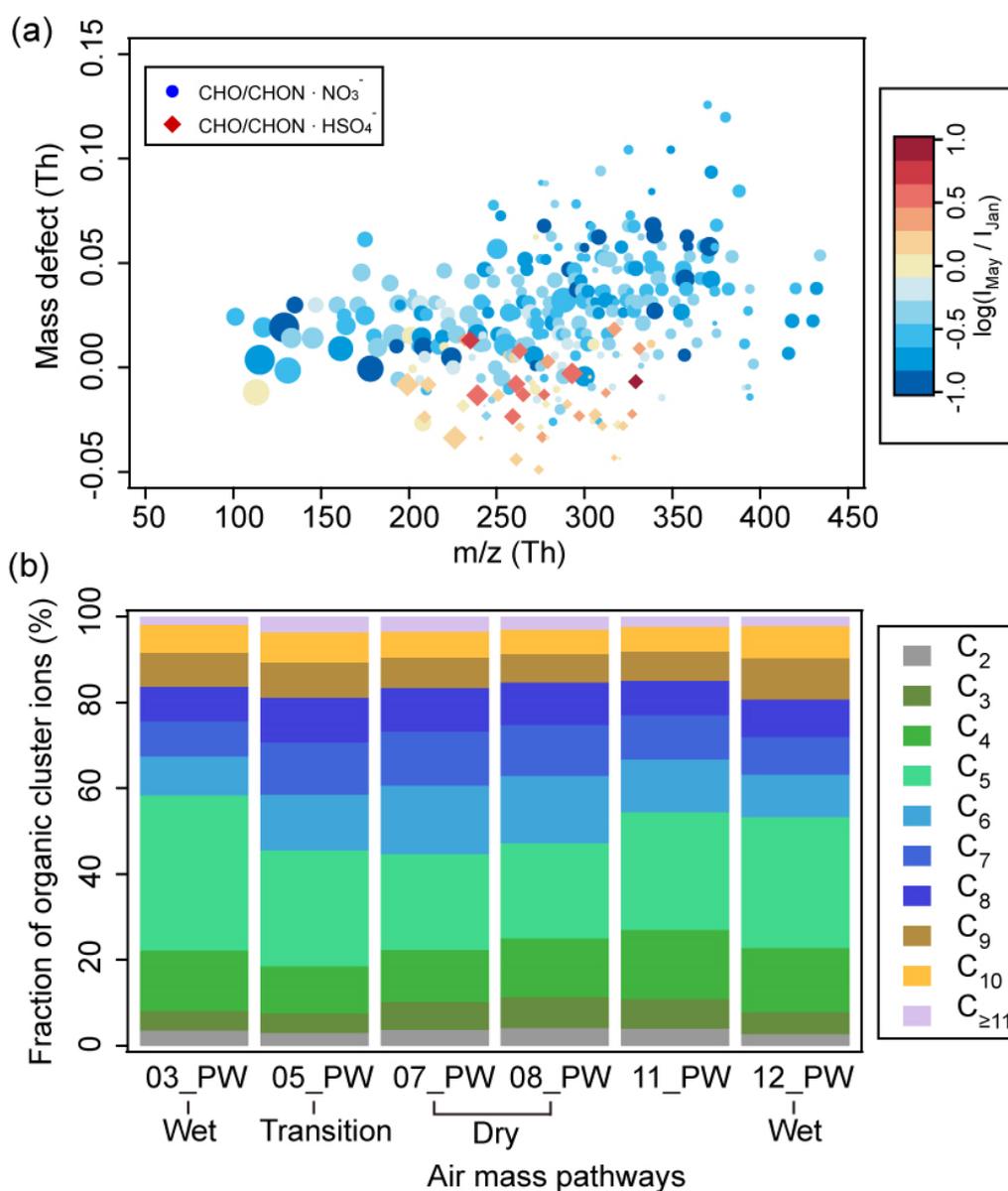
454

455 Figure 6 The fractions of the negative ion groups observed at CHC determined during  
 456 the representative periods of each air mass pathway (described in Section 2.3.2) for **(a)**  
 457 daytime (07:00 – 19:00) and **(b)** nighttime (19:00 – 07:00). A sketch of the horizontal  
 458 profile of the air mass pathways (Fig. 2a) is shown in the upper right corner for clarity.

459 A further investigation of the organic ion group shows that the seasonal trends of the  
 460 individual organic ions also varied (Fig. 7 for daytime and Fig. S7 for nighttime).  
 461 Whereas the majority of the organic cluster ions at CHC were more abundant during  
 462 wet season (Fig. 7a and Fig. S7a), fractions of CHO/CHON·HSO<sub>4</sub><sup>-</sup> increased during the  
 463 dry season. The observed increases of CHO/CHON·HSO<sub>4</sub><sup>-</sup> cluster ions could be  
 464 associated with the increased HSO<sub>4</sub><sup>-</sup>/NO<sub>3</sub><sup>-</sup> ratios in dry season (Fig. 5 and Fig. S5).  
 465 Similar increases of CHO/CHON·HSO<sub>4</sub><sup>-</sup> cluster ions were also found to relate to the  
 466 ratio of HSO<sub>4</sub><sup>-</sup>/NO<sub>3</sub><sup>-</sup> in a boreal forest environment (Bianchi et al., 2017). In addition,  
 467 changes in OOM composition between wet and dry seasons may also play a role (Fig.  
 468 7b and Fig. S7b), as NO<sub>3</sub><sup>-</sup> tends to cluster with OOM containing hydroxyl and  
 469 hydroperoxyl functional groups (Hyttinen et al., 2015) while some other observed  
 470 OOM may be more efficiently charged by HSO<sub>4</sub><sup>-</sup>.

471 The seasonal variations of the individual organic cluster ions are likely caused by  
 472 different air masses (Fig. 7b and Fig. S7b). The air masses influenced by tropical  
 473 rainforest vegetation from the Amazon Basin are dominated by isoprene (C<sub>5</sub>H<sub>8</sub>)  
 474 emissions and isoprene oxidation products (Bianchi et al., 2022). This region  
 475 corresponds to 03\_PW and 12\_PW (largest impact on CHC in wet season in January)  
 476 consisting of relatively higher fractions of organic cluster ions with OOM containing  
 477 4-5 carbon atoms (50 % and 46 % for nighttime, and 29 % and 32 % for daytime,  
 478 respectively). In contrast, when the air masses were more influenced by the Altiplano  
 479 plateau (i.e., 05\_PW, 07\_PW, and 08\_PW, with more anthropogenic emissions and less  
 480 vegetation) in wet-to-dry transition period and dry season, organic cluster ions with 6-  
 481 8 carbon atoms, potentially originating from anthropogenic sources (e.g., toluene

482 ( $C_7H_8$ ); Huang et al., 2019; Cai et al., 2022), were of higher contributions. The signal  
 483 fractions of these organic ions were thus the highest in these air mass pathways,  
 484 accounting for 36-39 % for nighttime and 37-39 % for daytime. For all the air mass  
 485 pathways, fractions of organic ions with more than 9 carbon atoms were relatively low  
 486 (<10 %). This might be due to their lower volatilities compared to OOM with smaller  
 487 carbon numbers (Donahue et al., 2012), resulting in a larger probability of them being  
 488 removed during their transport to CHC (e.g., condensing on pre-existing particles).

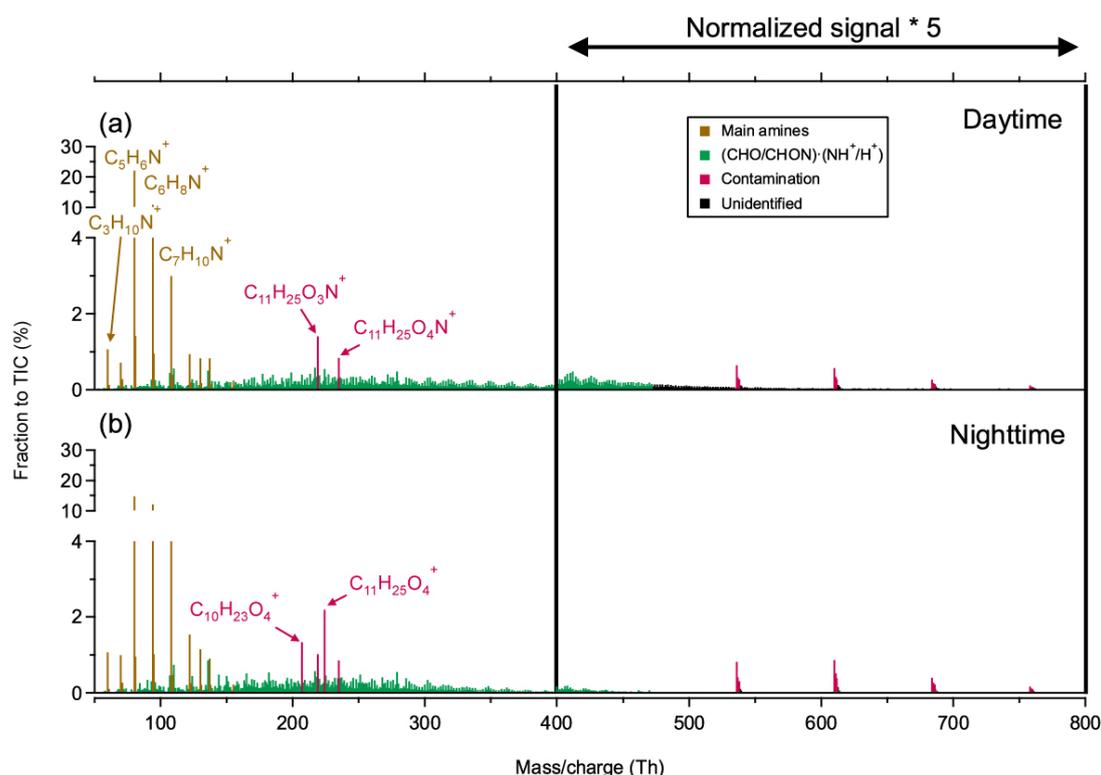


489  
 490 Figure 7 (a) Mass defect plot of organic cluster ions during nighttime (19:00-07:00).  
 491 The color code indicates ratios (in log scale) between median signals of each ion  
 492 detected in May (IMay) of dry season and January (IJan) of wet season. The marker  
 493 size is proportional to the log-transformed median signals of ions in May. (b) Fraction  
 494 of organic cluster ions from different air mass pathways as a function of carbon atom  
 495 numbers during nighttime (19:00-07:00). A similar figure based on daytime data  
 496 (07:00-19:00) is in the supplementary information (Figure S7). Note that MA-derived

497 ions were not included in this figure.

### 498 3.3 Positive ions

499 Several positive cluster ion groups were consistently observed in February and March  
500 (i.e., wet season) during the study period. Based on their chemical composition, the  
501 positive cluster ions measured at CHC are classified into four groups (Fig. 8): (1) a  
502 series of protonated amines, including trimethylamine ( $C_3H_9N\cdot H^+$ ), pyridine  
503 ( $C_5H_7N\cdot H^+$ ), aniline ( $C_6H_7N\cdot H^+$ ), and benzylamine ( $C_7H_9N\cdot H^+$ ); (2) organic cluster  
504 ions consisting of OOM (identified as  $C_{3-24}H_{6-39}O_{2-12}N_{0-2}$ ) clustered with positive  
505 charge carriers such as protons ( $H^+$ ), ammonium ( $NH_4^+$ ), and aminium ( $NH^+$ ) ions; (3)  
506 contamination ions; and (4) unidentified ions (likely organic ions in higher masses;  
507 Bianchi et al., 2021). Contamination in the positive cluster ions includes  
508 ethylhexylglycerin (e.g.,  $C_{11}H_{24}O_3\cdot NH^+$ ), which is widely used in cosmetics (Aerts et  
509 al., 2016), and polydimethylsiloxane (e.g.,  $(C_2H_6OSi)_7\cdot NH_4^+$ ) possibly from instrument  
510 tubing (Bianchi et al., 2014). In contrast to the negative cluster ions, the four positive  
511 cluster ion groups were generally stable with smaller diurnal variability over the study  
512 period (Fig. 9). This is similar to the diurnal patterns determined in previous studies in  
513 a boreal forest environment (Ehn et al., 2010) and at the JFJ (Frege et al., 2017).  
514 However, due to the unavailable measurements of the corresponding neutral species  
515 (e.g., amines), the exact reason for such weak diurnal variations observed in different  
516 locations remains unclear.

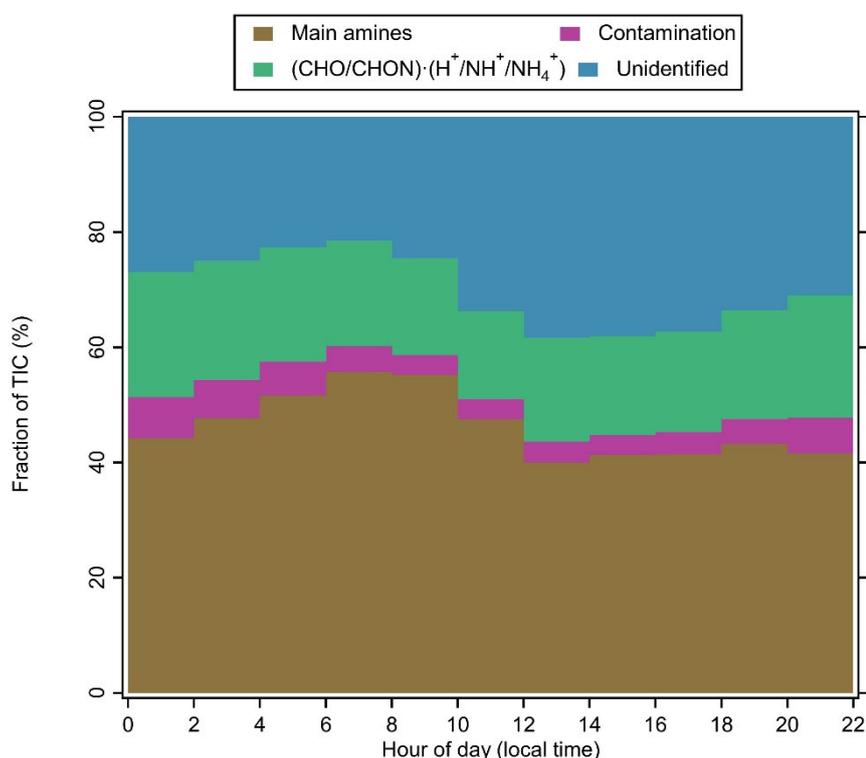


517

518 Figure 8 Averaged mass spectra of positive ions at CHC in February and March 2018  
519 when APi-TOF was operating in positive ion mode (see section 2.2.1), during (a)

520 daytime (07:00 – 19:00) and (b) nighttime (19:00 – 07:00). The normalized signal  
521 intensities from 400 Th to 800 Th are multiplied by a factor of 5 for better visualization.

522 The protonated amines were the most abundant positive ion group (46 %), with no  
523 significant diurnal variations. Nighttime contributions of this ion group (47 %) were  
524 similar to its daytime contributions (45 %; Fig. 9). They also dominated the positive  
525 ion spectra observed in different environments, such as a boreal forest (Ehn et al., 2010),  
526 the JFJ (Frege et al., 2017), and the free troposphere (Schulte and Arnold, 1990). Their  
527 sources have not been fully identified (Kosyakov et al., 2020), but they are widely used  
528 as solvents and dyes (Sims et al., 1989), which may be potential sources of these ions  
529 observed at CHC.

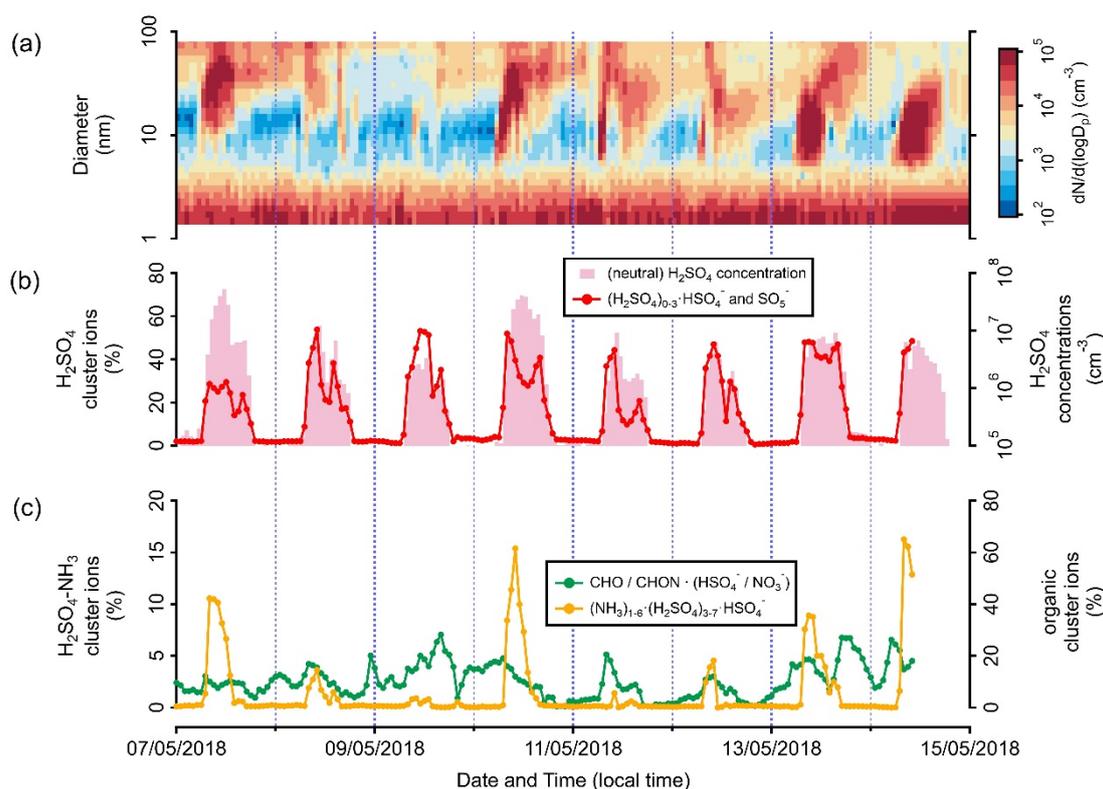


530  
531 Figure 9 Diurnal variation of positive ion groups at CHC, averaged over measurements  
532 in February and March 2018 (when APi-TOF was operating in positive ion mode, see  
533 section 2.2.1).

534 Positive organic cluster ions were also relatively abundant (19 %) at CHC during the  
535 wet season. Similar to the negative organic ions in wet season (Fig. 6), this reflects the  
536 influence of air masses originating from the Amazon Basin and Eastern/South-Eastern  
537 Lowlands (e.g., 03\_PW and 11\_PW). Differences in the positive organic ion signals  
538 between nighttime (21 %) and daytime (18 %) were small, which is similar to the  
539 negative organic cluster ions (see Fig. 3). A further investigation of the relationship  
540 between these positive ions and their neutral species is, unfortunately, not possible due  
541 to the unavailability of CI-APi-TOF data in February and March caused by instrumental  
542 issues.

543 **3.4 Potential connections between atmospheric ions and new particle**  
 544 **formation events**

545 During the SALTENA campaign from January to May 2018, NPF events were  
 546 frequently observed at CHC (Fig. S8). While most of them occurred from April (the  
 547 wet-to-dry transition period, 21 events) to May (dry season, 26 events), NPF events  
 548 seldom occurred during wet season from January to March (8 events in total). This is  
 549 consistent with a previous study performed at CHC (Rose et al., 2015a), which also  
 550 found that NPF events mainly occurred during dry season.

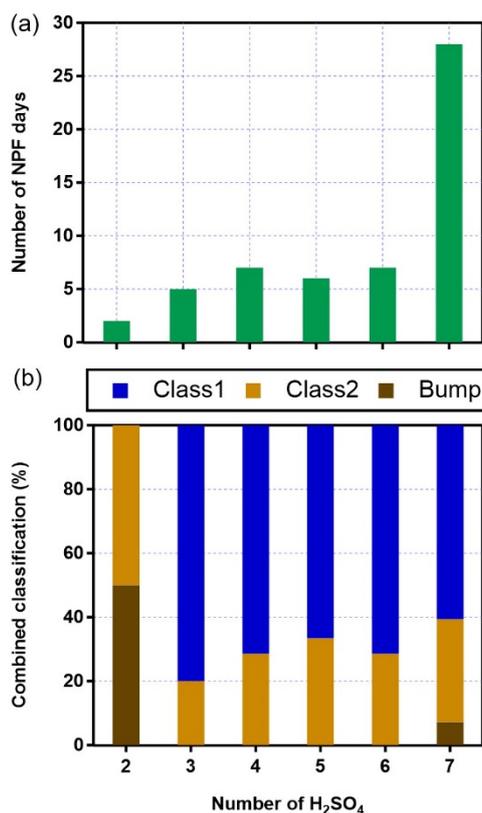


551

552 Figure 10 Time series of the (a) size distribution of aerosol particles (measured with  
 553 NAIS and MPSS), (b) signal fraction of the  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$  ion group and neutral  
 554  $\text{H}_2\text{SO}_4$  concentration, and (c) signal fractions of the  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$   
 555 and negative organic cluster ion groups, observed at CHC from 7 to 14 May 2018 when  
 556 NPF occurred frequently.

557 Previous field studies at high-altitude mountain sites have shown that NPF events can  
 558 be triggered by different compounds, such as low-volatile neutral OOM (Bianchi et al.,  
 559 2021), neutral  $\text{H}_2\text{SO}_4$  and OOM (Bianchi et al., 2016), and  $\text{H}_2\text{SO}_4\text{-NH}_3$  cluster ions  
 560 (Frege et al., 2017). While the signal fractions of the negative organic cluster ions did  
 561 not seem to have a strong correlation with the onset of the NPF events, the fractions of  
 562 the  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$  (associated with  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$ ) always increased  
 563 concurrently with the number concentration of small particles (see example NPF events  
 564 on, e.g., 7, 10, 13, and 14 May 2018; Fig. 10). The number of NPF days increased when

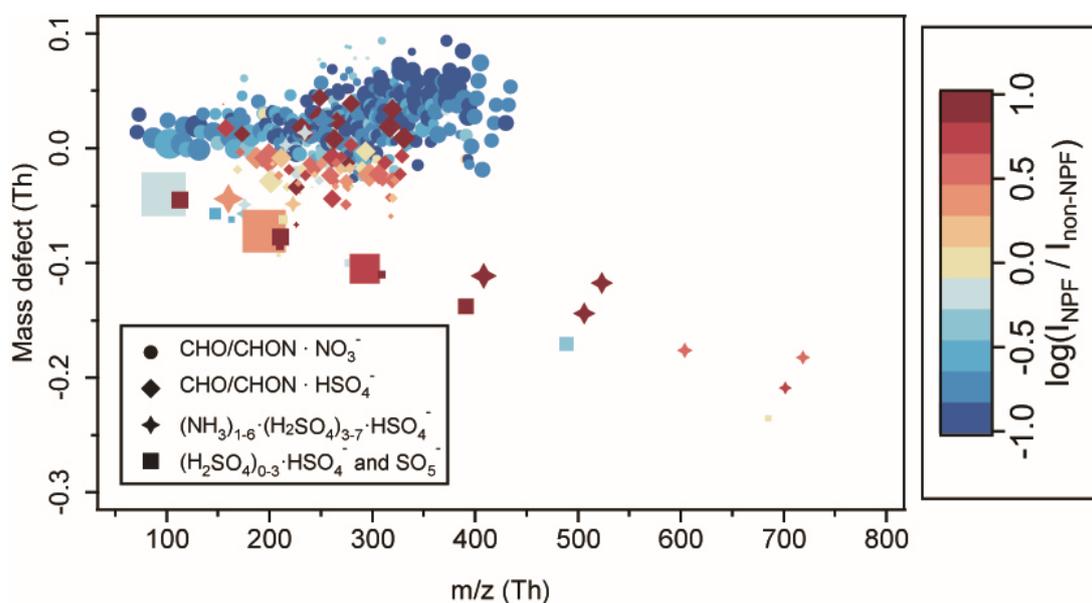
565 more H<sub>2</sub>SO<sub>4</sub> molecules were present in the (NH<sub>3</sub>)<sub>1-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>3-7</sub>•HSO<sub>4</sub><sup>-</sup> cluster ion (Fig.  
 566 11a). In particular, more than half of the NPF events (28 out of the total 55 events) were  
 567 observed in the presence of (NH<sub>3</sub>)<sub>4-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>7</sub>•HSO<sub>4</sub><sup>-</sup>. The majority (35 events) of all  
 568 the NPF events exhibited clear nucleation and growth processes (i.e., Class 1 events;  
 569 the classification is defined following the approach by Yli-Juuti et al. (2009); Fig. 11b).  
 570 In contrast, only Class 2 (similar to Class 1 but with less clarity) and bump events (early  
 571 growth of the newly formed particles is interrupted) were observed when only  
 572 (H<sub>2</sub>SO<sub>4</sub>)<sub>2</sub>•HSO<sub>4</sub><sup>-</sup> was observed.



573

574 Figure 11 Connection between the maximum number of H<sub>2</sub>SO<sub>4</sub> (in addition to HSO<sub>4</sub><sup>-</sup>)  
 575 observed in the (H<sub>2</sub>SO<sub>4</sub>)<sub>2</sub>•HSO<sub>4</sub><sup>-</sup> and (NH<sub>3</sub>)<sub>1-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>3-7</sub>•HSO<sub>4</sub><sup>-</sup> cluster ions during  
 576 NPF events and **(a)** the number of NPF days; **(b)** proportions of different NPF classes.

577 Moreover, higher levels (up to an order of magnitude) of (NH<sub>3</sub>)<sub>1-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>3-7</sub>•HSO<sub>4</sub><sup>-</sup>  
 578 and (H<sub>2</sub>SO<sub>4</sub>)<sub>0-3</sub>•HSO<sub>4</sub><sup>-</sup> ions as well as negative organic cluster ions charged by HSO<sub>4</sub><sup>-</sup>  
 579 were also observed during the NPF days (Fig. 12). In contrast, other negative ion groups  
 580 (e.g., the majority of the negative organic cluster ions charged by NO<sub>3</sub><sup>-</sup>) were more  
 581 abundant during the non-NPF days. Our observations indicate a potentially important  
 582 role of (NH<sub>3</sub>)<sub>1-6</sub>•(H<sub>2</sub>SO<sub>4</sub>)<sub>3-7</sub>•HSO<sub>4</sub><sup>-</sup> cluster ions in NPF events at CHC from January to  
 583 May 2018, particularly in wet-to-dry transition period and dry season.



584

585 Figure 12 Mass defect plot of differences in negative cluster ion composition between  
 586 NPF and non-NPF days. The negative ion composition of NPF events was averaged  
 587 over all NPF days from 08:00 to 12:00 in January, April, and May 2018 (when APi-  
 588 TOF was operating in negative ion mode, see section 2.2.1) at CHC. The ion  
 589 composition of non-NPF days was averaged over non-NPF days from 08:00 to 12:00  
 590 for the same period. The x-axis is the exact mass of cluster ions, and the y-axis is the  
 591 mass defect. The color code indicates ratios (in log scale) between median signals of  
 592 each ion determined in NPF events ( $I_{\text{NPF}}$ ) and non-NPF periods ( $I_{\text{non-NPF}}$ ). The marker  
 593 size is proportional to the log-transformed median signals of ions observed in the NPF  
 594 events that occurred in January, April and May. Note that  $(\text{HNO}_3)_{0-2} \cdot \text{NO}_3^-$  cluster ions  
 595 were not included here.

596 The majority of the observed NPF events occurred when CHC was more impacted by  
 597 air masses originating from source regions with elevated  $\text{SO}_2$  emissions (05\_PW,  
 598 07\_PW, and 08\_PW). This is similar to the observations from the high-altitude station  
 599 JFJ (Frege et al., 2017). Moreover, consistent with the previous findings from Rose et  
 600 al. (2015a), CHC was affected by the frequent arrival of air masses from the Pacific  
 601 Ocean during the dry season (see also Fig. 2). And the NPF during the dry season at  
 602 CHC seems to be triggered once the air masses arrive on the continent (Rose et al.,  
 603 2015). However, these air masses of marine origin may not directly contain the  
 604 nucleating species (also reflected in the poor connection between NPF events and the  
 605 levels of methanesulfonic acid ion ( $\text{CH}_3\text{SO}_3^-$ ) or  $\text{IO}_3^-$ ; data not shown). In addition,  
 606 inconsistent time series of  $\text{CH}_3\text{SO}_3^-$  and  $\text{IO}_3^-$  prohibit us from drawing further  
 607 conclusions on the role of marine-derived compounds on NPF at Chacaltaya during the  
 608 dry season. In addition, the fraction of large positive organic cluster ions (mass range  
 609 from 500 to 800 Th; Fig. S9) was found to increase during the NPF events in the wet  
 610 season. These organic cluster ions usually contained at least ten carbon atoms (Fig. S10).  
 611 Such large positive organic ions have been found to contribute to NPF in the Himalayas  
 612 (Bianchi et al., 2021), and thus, their contribution to NPF can not be completely ruled

613 out at CHC as well.

#### 614 **4. Conclusion**

615 In this study, both negative and positive atmospheric ions were measured at a high-  
616 altitude research station (CHC) in the Bolivian Andes for five months, from January to  
617 May 2018, using an APi-TOF mass spectrometer. Negative ions were mainly composed  
618 of  $(\text{H}_2\text{SO}_4)_{0-3}\cdot\text{HSO}_4^-$ ,  $(\text{HNO}_3)_{0-2}\cdot\text{NO}_3^-$ ,  $\text{SO}_5^-$ ,  $(\text{NH}_3)_{1-6}\cdot(\text{H}_2\text{SO}_4)_{3-7}\cdot\text{HSO}_4^-$ , MA-derived,  
619 and CHO/CHON $\cdot(\text{HSO}_4^-/\text{NO}_3^-)$  ion groups. Positive ions mainly consisted of a series  
620 of protonated amines ( $\text{C}_{3-7}\text{H}_{7-9}\text{N}\cdot\text{H}^+$ ) and organic cluster ions  
621 CHO/CHON $\cdot(\text{H}^+/\text{NH}_4^+/\text{NH}^+)$ . Distinct diurnal variation was observed for the negative  
622 ions, and attributed mainly to the changes in the corresponding neutral species'  
623 concentrations and/or their EA / PA. An example is  $\text{H}_2\text{SO}_4$ -related cluster ions, the diel  
624 temporal variation of which was mainly due to the photochemical production of neutral  
625  $\text{H}_2\text{SO}_4$  during daytime. Strong seasonality of negative ions was also found, such as for  
626  $\text{H}_2\text{SO}_4$ -related cluster ions owing to changes in  $\text{SO}_2$  and the resulting neutral  $\text{H}_2\text{SO}_4$   
627 concentrations. The seasonal variation was mainly because of the differences in source  
628 regions of air masses arriving at CHC from wet to dry seasons. In contrast, no  
629 significant diurnal variation was observed for the positive ions. The comparison  
630 between NPF and non-NPF days infers that  $\text{H}_2\text{SO}_4$ - $\text{NH}_3$  cluster ions contribute to the  
631 aerosol nucleation process at CHC, particularly in wet-to-dry transition period and dry  
632 season when CHC was more impacted by air masses originating from source regions  
633 with elevated  $\text{SO}_2$  emissions. The results further indicate that atmospheric ion  
634 composition at CHC is directly affected by air masses from different source regions.

635 Measurements of atmospheric ions in the field will improve understanding of  
636 atmospheric physical and chemical processes in the study regions, as the ions play  
637 important roles in atmospheric chemistry through participation in or catalysis of ion-  
638 molecule reactions and ion-induced new particle formation. Our study thus provides  
639 new insights into the chemical composition of atmospheric ions and their potential role  
640 in high-altitude NPF in the Bolivian Andes where both natural (e.g., biogenic and  
641 volcanic) and anthropogenic emissions are important.

642

643 **Data availability:** The data that are involved in the figures can be found in  
644 [doi.org/10.5281/zenodo.7271286](https://doi.org/10.5281/zenodo.7271286) (Zha et al., 2022).

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647 M.A., C.M., and F.B. collected the data and operated the instruments during the  
648 measurement campaign. Q.Z. and W.H. wrote the manuscript with contributions from  
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665

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