The 2021 volcanic eruption in La Palma Island and its impact on ionospheric scintillation as measured from GNSS reference stations, GNSS-R, and GNSS-RO

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Abstract. Ionospheric disturbances induced by seismic activity have been studied in the last years by many authors, showing an impact both before and after the occurrence of earthquakes. In this study, the ionospheric scintillation produced by the 2021 La Palma volcano eruption is analyzed. The "Cumbre Vieja" volcano was active from 19 September to 13 December 2021, and many magnitude 3-4 earthquakes were recorded, with some of them reaching magnitude 5. Three methods: GNSS reference monitoring, GNSS Reflectometry (GNSS-R) from NASA CYGNSS, and GNSS Radio Occultation (GNSS-RO) from COSMIC and Spire constellations are used to compare and evaluate their sensitivity as proxies of earthquakes associated with the volcanic eruption. To compare the seismic activity with ionospheric scintillation, earthquakes' generated energy, and percentile 95 of the intensity scintillation parameter (S_4), measurements have been computed every 6 h intervals for the whole duration of the volcanic eruption. GNSS-RO has shown the best correlation between earthquakes' energy and S_4, with values up to 0.09 when the perturbations occur around 18 h after the seismic activity. GNSS reference monitoring stations data also shows some correlation 18 h after and 7-8 days after. As expected, GNSS-R is the one that shows the smallest correlation, as the ionospheric signatures get masked by the signature of the surface where the reflection is taking place. Additionally, the three methods show a smaller correlation during the week before earthquakes, even though, given the small magnitude of the seismic activity, the correlation is barely detectable in this situation, and difficult to use in any application to find earthquake proxies.

1 Introduction

Ionospheric disturbances such as scintillation constitute a notable problem for satellite communications, global navigation satellite systems (GNSS), and Earth observation systems, notably at P and L-bands. They can disturb the signals making it difficult or even impossible to transmit the correct information through the ionosphere. Nevertheless, they can also be seen as an opportunity to detect, measure or infer other physical phenomena, not necessarily related to the ionosphere itself. For example, in the last decades, several studies have shown that ionospheric disturbances can occur during solar eclipses (Das

et al., 2022), or geomagnetic storms (Ding et al., 2007; Li et al., 2008), which are due to causes external to the Earth, coming from the Sun or near space weather.

In addition, other "internal" causes have recently shown that they can impact the ionosphere too. There are studies relating severe atmospheric phenomena such as cyclones or hurricanes to ionospheric anomalies (Ke et al., 2019; Xu et al., 2019; Kamogawa, 2006). Anomalous variations in the Total Electron Content (TEC) and peaks in the ionospheric scintillation have been detected during the passage of a large cyclone or hurricane, which creates gravity waves that couple to the lower ionosphere, yielding to ionospheric disturbances.

Another source of perturbations in the ionosphere is related to the lithosphere, and the seismic activity, as supported by many studies that are being published in the last years by (Liu et al., 2004; Pulinets, 2004; Kamogawa, 2006; Pulinets and Davidenko, 2014; Pulinets et al., 2021; Molina et al., 2021, 2022). The physical mechanisms behind this interaction are still not very clear, but there are several research paths open. Some of them explain this interaction by the generation of low-frequency electromagnetic signals from the underground rock under huge pressures during the earthquake preparation period. Other authors explain the interaction by changes in the surface electric potential due to the piezoelectric effect in the underlying rocks, which can induce changes in the ionosphere's TEC.

This study looks for ionospheric anomalies related to the seismic activity generated by a recent volcano eruption on the Spanish island of La Palma (Canary Islands). In the past, other studies have analyzed the impact of volcanic eruptions on the ionosphere (De Ragone et al., 2004; Shults et al., 2016; Astafyeva, 2019; Yong-Qiang et al., 2006). For example, the recent Tonga submarine eruption on 12 January 2022 created traveling ionospheric disturbances (TID) from the eruption site (Themens et al., 2022). In this case, the eruption was so strong that the gravity waves generated within the atmosphere traveled to the ionosphere, and then propagated concentrically all around the globe producing these perturbations. The La Palma eruption that happened last year, wasn't as explosive as the Tonga one, but it was much longer in time, and it had significant seismic activity associated with it.

1.1 The September 2021 "Cumbre Vieja" Volcanic eruption in La Palma (Spain)

La Palma is a volcanic island located in the North West of the Canary Islands archipelago (Spain) in the Atlantic Ocean, at ~500 km from the coast of Africa. The island has relatively low volcanic activity with only three eruptions in the last century, and seven in the last 500 years, even though it is one of the islands in the archipelago with the highest potential risks (Fernández et al., 2021).

The last eruption started at 13:43 UTC on 19 September 2021, near the former volcano "Cumbre Vieja", and it lasted for 85 days until 13 December 2021, when it was declared finished.

Preceding the eruption, for eight days since 11 September a series of earthquakes were registered in the region where the eruption took place.

Around 6000 earthquakes occurred during this time frame with magnitudes ranging from 1mbLg to 3.8 mbLg (mbLg is the magnitude unit used by the Instituto Geográfico Nacional (IGN) to characterize the earthquakes on the island). In this period, the epicenters migrated North, approaching to the eruptive cone location, at the same time as the hipocenter depths rose starting

55   from 15 km below the surface to near-surface depths. The evolution of the locations and depths of this precursory earthquake swarm is shown in Fig. 1.
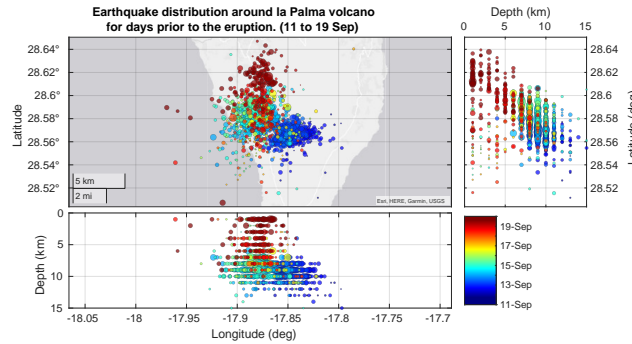
Figure 1. Series of earthquakes preceding the volcanic eruption from 11 September indicating their position, depths and date.

During the 10 days following the start of the eruption, while the lava, gasses, and ashes were being expelled from the volcano, the seismic activity was relatively low. At the begging of October, around 15 days after the start of the eruption, the seismic activity increased again and remain stable until the end of the eruption, which was declared to be on 13 December.

2   Data sources and methods

This study is focused on the simultaneous analysis of several sources of data measuring the ionospheric scintillation with GNSS signals and correlating them with the seismic activity related to the eruptive event in La Palma in 2021.

The same physical mechanism is involved in the three techniques used to measure the ionospheric scintillation through GNSS signals, and it is based on the fact that in the ionosphere, when the electron density is perturbed, the propagation of
65   electromagnetic waves gets affected, in particular in the GNSS bands and at lower frequencies.

Ionospheric scintillation refers to the rapid fluctuations in phase and/or intensity of the electromagnetic signal received after crossing the ionosphere. From now on, the study will be centered only on the intensity, also called "amplitude scintillation". It is usually measured as the normalized standard deviation of the intensity of a radio electromagnetic signal after crossing the ionosphere, and it is computed with eqn. 1.

$$70 \quad S_4 = \sqrt{\frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}}, \tag{1}$$

where I is the signal's intensity, and $\langle \cdots \rangle$ represents the average of a certain period of time, usually in the order of tens of seconds.

In the current state-of-the-art there are three ways of measuring this ionospheric amplitude scintillation with GNSS signals: (1) by static ground-based GNSS reference monitoring, (2) by GNSS-R (GNSS Reflectometry), and (3) by GNSS-RO (GNSS

75   Radio Occultation). The novelty and interest of this work is the analysis of the effects that the same physical phenomenon, a volcanic eruption, produces in the ionosphere, as observed by three different techniques measuring the same magnitude at the same frequency.

2.1   Ground stations data

Since the appearance of GNSS, ground-based monitoring GNSS stations have been operating and continuously monitoring the
80   signals emitted by these satellites. GNSS monitoring ground station networks have been sensing the behavior of the ionosphere for many years, providing TEC and scintillation data required to make the ionospheric corrections needed for navigation, and to assess the quality of the service.

    One disadvantage of this technique is the sparse spatial coverage compared to other satellite-based techniques. GNSS stations are typically sparsely installed at fixed, single ground locations, only providing data for the local region. Fortunately for this
85   study, there are two ground stations close to the volcano, one station is located on the same island (La Palma), which from now on, will be identified as LPAL, and the second one in Maspalomas (MAS1), a town South of the Gran Canaria Island, which is around 250 km southeast of the volcano. Figure 2 shows a map displaying the two ground stations.
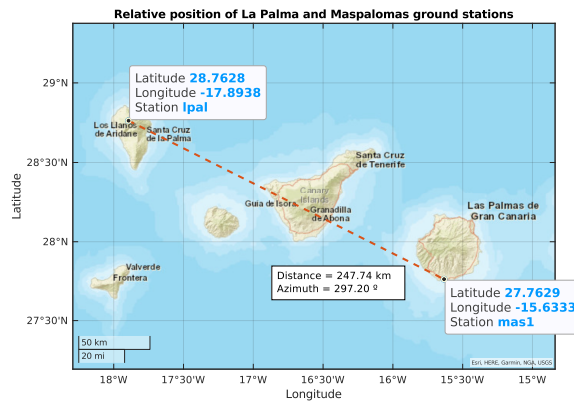
Figure 2. Relative position of LPAL and MAS1 ground stations with respect to the La Palma Island, which is the top left one. Map: Esri, HERE, Garmin, NGA, USGS.

    The GNSS monitoring ground station data contains the measurement for every minute of all the GNSS satellites in view for each station (Juan et al., 2018; Rovira-Garcia et al., 2020). The database has been pre-processed using the algorithms described
90   in (Juan et al., 2017) work, and it includes the variables in Table 1:

    The S_4 data here is obtained by using eqn. 1 averaging over a period of 1 minute, and using the direct intensity received in ground stations from the GNSS satellites in view.

    A clear trend between the elevation angle and the measured scintillation has been observed. The smaller the elevation angle, the larger the S_4, which can be attributed to the longer path within the ionosphere of the lower elevation rays, which may suffer

Table 1. GNSS monitoring ground stations database variables

Variable Name	Description
Year	Year of the measurement
DoY	Day of the Year
Seconds	Second of the day
Satellite	Satellite number
	1 to 33: GPS
	37 to 70: GLONASS
	71 to 120: GALILEO
	121+: discarded
Azimuth (deg)	Azimuth from the station (North at 0º)
Elevation (deg)	Elevation from the station (0º – 90º)
S4	Amplitude scintillation calculated over 60 s
Sigma_phi	σ_ϕ from the L1 signal calculated over 60 s

95 from stronger scintillation and multipath propagation. This behavior is displayed in Fig. 3. Therefore, data with an elevation lower than 30º has been discarded.
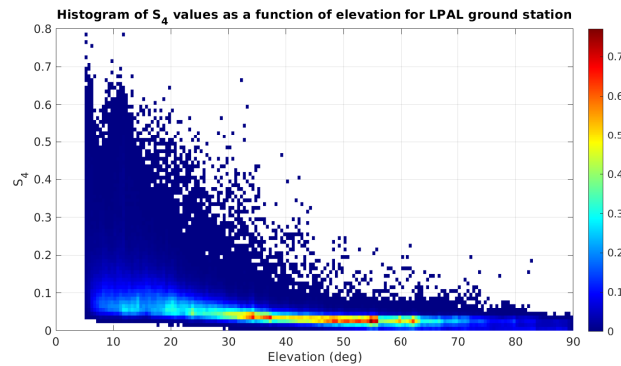
Figure 3. Histogram of S_4 values with respect to the elevation angle for La Palma ground station data.

Also, the dependence on azimuth has been studied. A certain dependence of the number of measurements vs. the azimuth has also been found. As shown in Fig. 4, the most probable azimuths for receiving GNSS signals at the LPAL station are +45º and -45º.
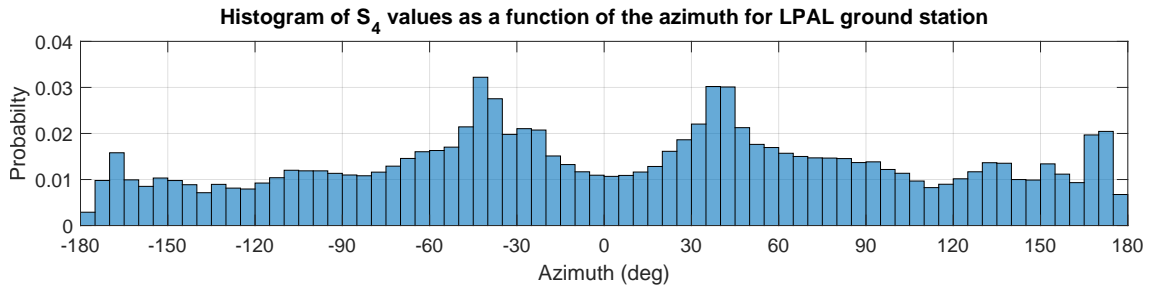
Figure 4. Histogram of the amount of S_4 measurements as a function of the azimuth.

Figure 5 shows the percentile 95% of S_4 values as a function of time and azimuth, and it can be observed that the largest values are obtained, for both stations, mostly when the GNSS satellites are located southwards, near 180º azimuth.
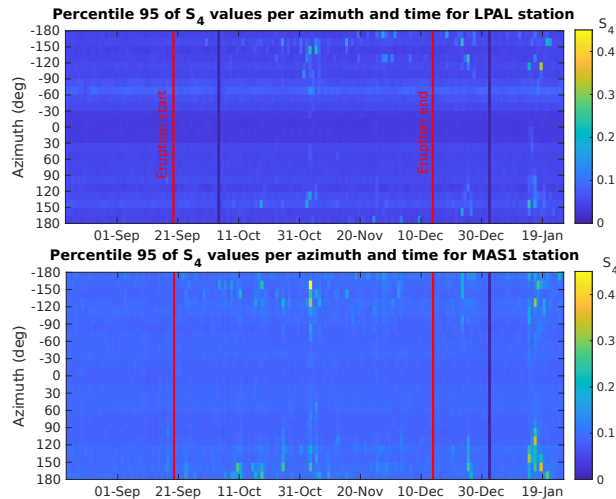
Figure 5. Percentile 95% of S_4 values as a function of the azimuth (in vertical axis) and time (in horizontal axis) for LPAL and MAS1 ground stations respectively, where it can be observed that most of the S_4 peaks are detected around the South direction.

2.2 GNSS-R data

GNSS-Reflectometry has also proven to be a good way to detect ionospheric scintillation over calm oceanic regions (Molina and Camps, 2020). The open-access NASA CYGNSS GNSS-R constellation, which started providing science data in March 2017, was utilized for this study. CYGNSS orbit's inclination is around 35º, so the coverage is from 40ºS to 40ºN, which includes the latitude of the islands (28.5ºN).
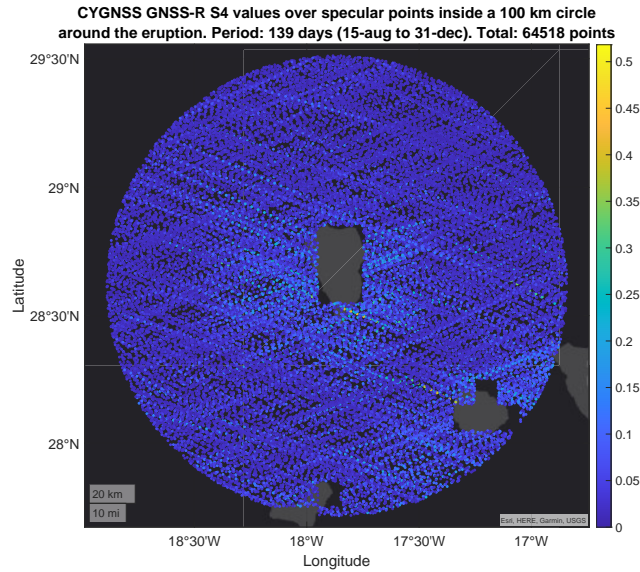
Figure 6. Position of CYGNSS GNSS-R specular reflection points inside a circle of 100 km around La Palma Island for the whole period studied, indicating the S_4 in colors. Map: Esri, HERE, Garmin, USGS.

The eight satellites comprising the CYGNSS constellation are continuously tracking up to 4 GPS satellites in view, taking measurements at a sample rate of 2 Hz, providing good availability of data in the region close to La Palma. During the full 139-day eruptive period, about 65 000 points were recorded within a radius of 50 km around the island.

110   The location of each point corresponds to the specular reflection point of each trajectory between the GPS satellite and the CYGNSS receiver, as shown in Fig. 6. The track of these reflected signals crosses the ionosphere twice because the height of the CYGNSS satellites is around 520 km, which is above the typical height of the ionosphere's maximum density (~350 km). As GNSS-R cannot distinguish if the scintillation was generated in the ascending or descending paths, the specular reflection point is used to locate the measurements.

115   GNSS-R data has been processed following the methodology detailed in (Molina and Camps, 2020), doing the moving average and standard deviation of the SNR for a 10 s window, and computing the S_4 index with eqn. 1, in which I has been computed from the SNR of the Delay Doppler Map (DDM) of the CYGNSS product.

As the performance of the GNSS-R technique to estimate the scintillation is affected by the sea surface roughness, another filter has been applied. A wavy water surface destroys the signal coherence, making it impossible to infer the scintillation 120   suffered along the path. Using data from maritime buoys around the islands, the model SIMAR detailed in Puertos del Estado (2020) documentation extrapolates the wave height in a grid of points along the ocean. Figure 7 shows the position of these points around La Palma Island and the ones selected to estimate the wave height during the eruptive period. The wave height extracted from them is compared to the detected scintillation from CYGNSS in Fig. 8. It can be observed that the high values of scintillation can appear only when the sea roughness is small, as indicated by the color-shaded areas. For example, the peaks

125   around 4 September appear during a period of waves lower than 1.5 m. Similar behavior is observed around 25 September, 16 October, or 2 November.

Figure 7. Set of SIMAR model points around the island, indicating the ones selected to extract the Significant Wave Height (SWH) around La Palma. Dataset downloaded from Puertos del Estado (2022) website. Map: Sentinel-2 cloudless 2021 by EOX IT Services GmbH.
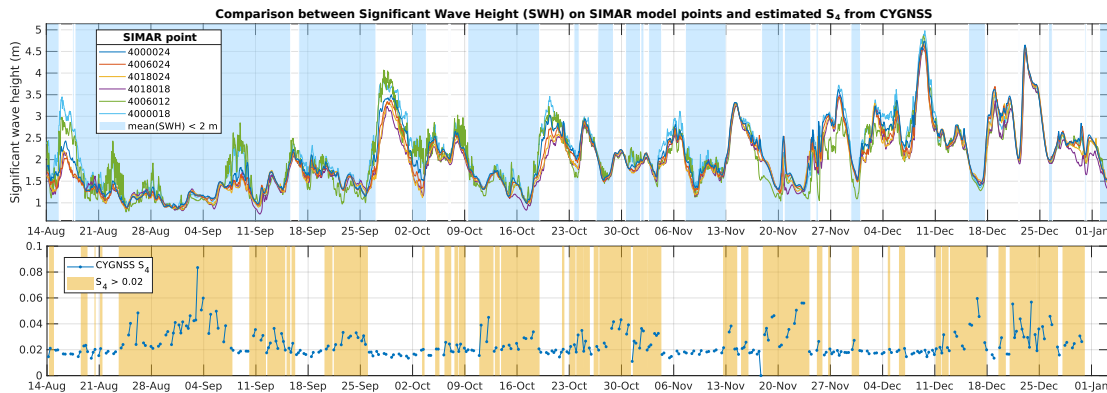
Figure 8. Significant Wave Height (SWH) from 6 SIMAR model points every hour around La Palma, compared to the estimated S_4 from CYGNSS GNSS-R data. Light blue shaded area marks the periods with an average SWH less than 2 m and the yellow area marks the intervals in which the S_4 is larger than 0.02, showing that most of the S_4 peaks only appear when the SWH is low.

Figure 9 shows the comparison between the Significant Wave Height (SWH) in the horizontal axis vs. the detected value of S_4, confirming what was observed in the timeline in Fig. 8. In this correlation, we can define a noise floor at 0.02 to remove all the values that are prone to be affected by the sea surface roughness. In the study, we compare the results for both cases: when

130   using all data without filtering and when using only S_4 values above 0.02.
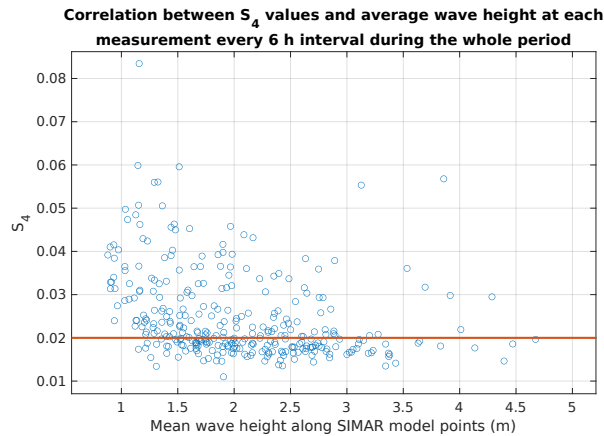
Figure 9. Correlation between the wave height and the detected scintillation index S_4 using GNSS-R from CYGNSS.

2.3 GNSS-RO data

The GNSS Radio-Occultation method is another way to retrieve information about the ionosphere using GNSS signals. In this case, the signal emitted by the GNSS satellites is received by the receiver onboard a LEO satellite when they are setting under or rising above the horizon. The use of this technique has the advantage of not being affected by ground reflection disturbances as in GNSS-R. For the same reason, land and oceanic regions can be studied indistinctly.

For this study, Spire and open-access data from COSMIC-2 (UCAR/NCAR, 2019) have been used. Spire Global (Jales et al., 2020; Irisov et al., 2018) operates a constellation of more than 80 3U CubeSats that can perform GNSS-RO, and more recently GNSS-R as well. From Spire, measurements from around 58 000 GNSS-RO occultations in the region around La Palma Island from 15 August to 31 October 2019 have been used.

Cosmic data includes all the occultations of GPS satellites as seen from the constellation of COSMIC-2 LEO satellites. COSMIC-2, also known as FORMOSAT-7, is a constellation of 6 LEO mini-satellites (300 kg) that were launched on 25 June 2019 into a 24º inclination orbit. The Level 1b podTc2 dataset contains the information used in this study, which is detailed in Table 2.
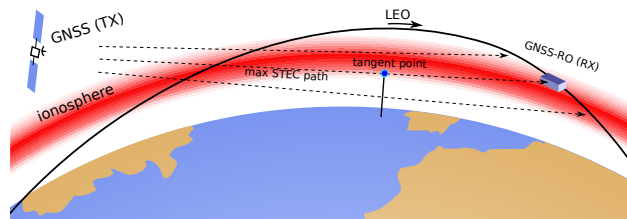
Figure 10. Schematics of the GNSS-RO technique to measure the ionosphere, indicating the location of the tangent point in the path with maximum STEC value.

Table 2. Level 1b "podTc2" database variables available from COSMIC-2 database.

Variable Name	Description
Time (GPS seconds)	GPS time of the measurements (seconds from 1 Jan 1980)
TEC (TECU)	Total Electron Content along LEO-GPS link (slant TEC)
caL1_SNR (V/V)	Signal to Noise ratio on the L1 channel, CA code
pL2_SNR (V/V)	Signal to Noise Ratio on the L2 channel, P code
x_LEO (km) y_LEO (km) z_LEO (km)	LEO position (ECF) at time of signal acquisition.
x_GPS (km) y_GPS (km) z_GPS (km)	GPS position (ECF) at time of signal acquisition.
S4	S_4 scintillation index

With these data, the tangent point, (i.e., the point in the trajectory which is closer to the Earth), can be computed from the
145   GPS and LEO satellite positions. For each of these points, the corresponding S_4 value has been computed, and their coordinates will be used later to filter by distance around the volcano. Given that the occultations have been recorded during long periods, including high elevation angles (almost 90º in some cases), it is proposed a method to filter the data points that correspond to the rays traversing the ionosphere.

Using the slant TEC (STEC) value during the occultation, we can assume that the maximum value coincides with the path
150   in which the ray crosses its longest path in the ionosphere, as shown in Fig. 10. Selecting the points that have a STEC larger than 80% of the maximum TEC in each occultation leaves the points that are inside the ionosphere, more likely to suffer from scintillation, as show in Fig. 11.

Figure 12 shows the location of the points with STEC larger than the 80% of the maximum one and within a circle of 1000 km radius around the eruption coordinates.
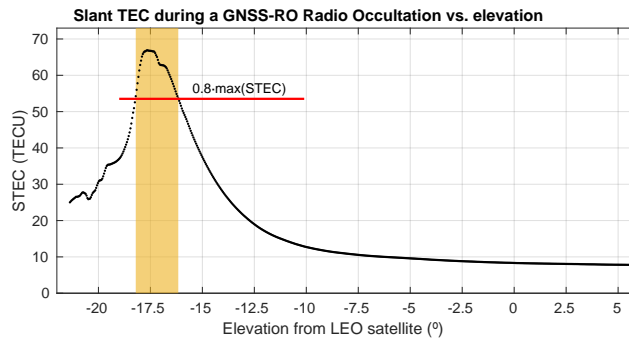
Figure 11. STEC variation during a GNSS occultation as a function of the elevation angles seen from a COSMIC-2 satellite on 19 September 2021.
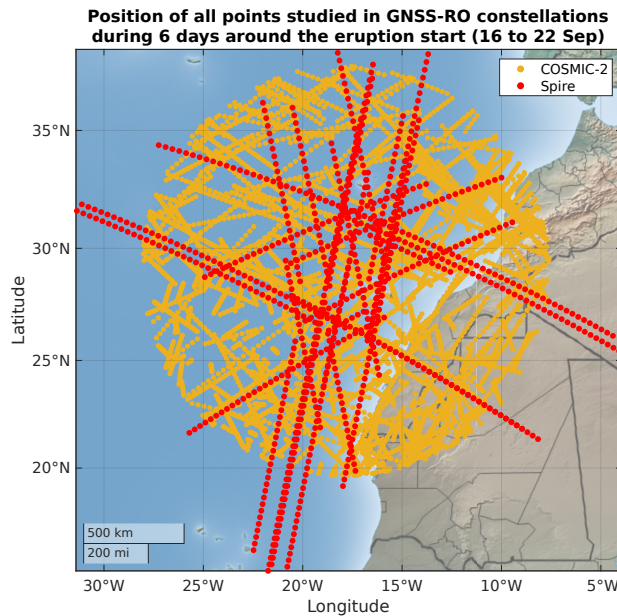
Figure 12. Location of Spire (in red) and COSMIC-2 (in yellow) measurements during 6 days (±3 days) around the start of the volcanic eruption on 19 September. Map: made with Natural Earth.

2.4   Seismic activity data

The database of earthquakes has been retrieved from the Spanish Instituto Geográfico Nacional (IGN) (2022). For the whole duration of the eruption, around 9200 earthquakes have been recorded in the database, which includes information about their time, magnitude, location, and depth.

Earthquakes are represented in Fig. 13, indicating their depth in the vertical axis during the whole eruption period (yellow shade). As observed in the figure, there is a precursor seismic activity close to the surface with decreasing depths, then a

relatively calm period of 8-days, preceding a roughly stable activity with earthquakes at two differentiated depths, but with homogeneous magnitudes in each group. When the eruption ends, the seismic activity lasted for approximately 15 more days with decreasing magnitude and frequency.
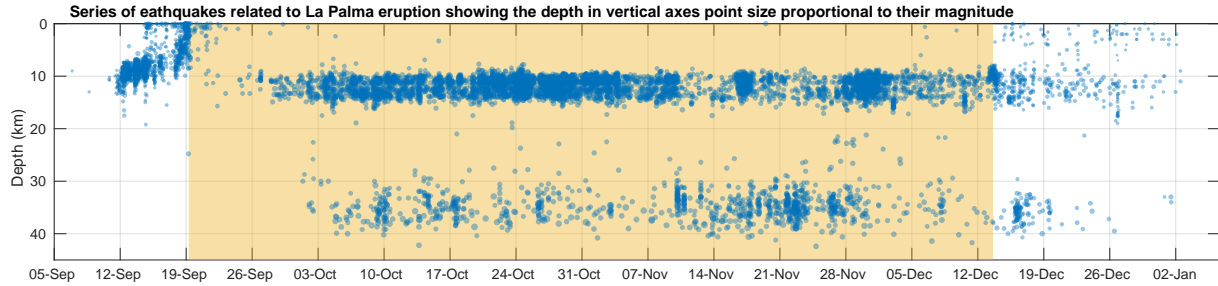
Figure 13. Series of earthquakes associated with the volcanic eruption in La Palma Island, indicating their depth in the vertical axis, and their magnitude proportional to the point size. The volcanic eruption is active during the period shaded in yellow.

To allow the comparison between the seismic activity and the corresponding ionospheric scintillation indicator, it is proposed to use the energy generated per earthquake and then integrate all the earthquakes' energy in a time interval. The rationale behind the selection of these metrics is that, whichever mechanism is involved in the perturbation of the ionosphere, the larger the energy dissipated into the environment, the larger the induced perturbations should be. The formula used to compute the energy per earthquake is taken from the work of (Gutenberg and Richter, 1955):

$$\log E = 5.8 + 2.4m, \tag{2}$$

where m is the magnitude of the earthquake, and E is the dissipated energy in Ergs.

Figure 14 shows a temporal histogram of the magnitude of earthquakes during the eruption period, indicating in vertical the magnitude, and in colors the number of earthquakes per bin. Red arrows mark the beginning and end of the eruption, respectively. The orange line represents the integrated energy every 6-hour intervals, computed for each earthquake using eqn. 2.

2.5 Geomagnetic and solar activity data

An important factor that impacts the ionosphere and can produce ionospheric scintillation is due to the geomagnetic perturbations and the solar weather. The geomagnetic data used in this study is the planetary index, K_p, which is an internationally recognized index usually used in the aerospace technologies and physical models of the geomagnetic environment. It is obtained from geomagnetic perturbations produced by the solar wind, and it is measured from the K indices of 13 observatories around the world located outside the auroral zone.

The data it is gathered by the GFZ German Research Centre for Geosciences (Matzka et al., 2021), from where it has been downloaded the period corresponding to the la Palma volcanic eruption. The K_p is presented in 3-hour intervals.
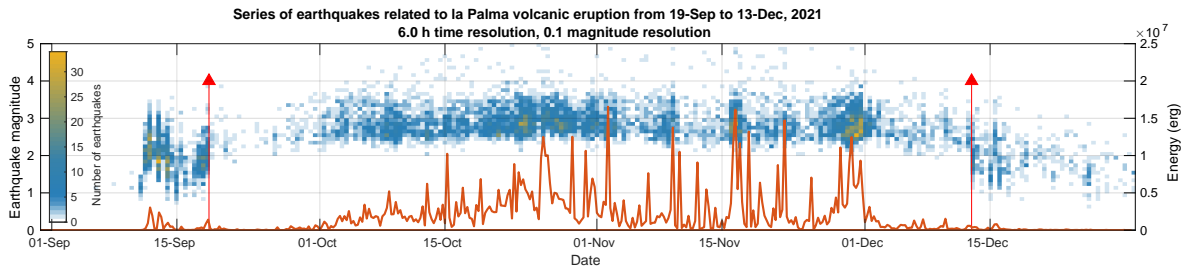
Figure 14. Seismic activity associated with the La Palma volcano eruption for the whole period, showing in colors the histogram of magnitudes and the integrated generated energy with an orange line. The red arrows mark the start and end of the volcanic eruption.

The solar activity data has been taken into account by studying the solar flux at the radiofrequency range 10.7 cm (F10.7). The solar flux is one of the main sources of geomagnetic and ionospheric perturbations. Its value is expressed in Solar Flux Units (SFU), and it is recorded with a periodicity of one day from 1947.

The dataset used is the Penticton Solar Radio Flux at 10.7 cm (National Research Council Canada (NRC), 2023), which contains two variables: the observed solar flux at Earth, and the adjusted solar flux, which compensates the varying distance from the Sun to the Earth. In our case, the actual flux arriving to the Earth is which is impacting the ionosphere, therefore, the observed flux has been used.

Both geomagnetic disturbances and solar activity data are analyzed in this work in relationship with the scintillation index to complete the study and explain, or discard some signatures found in the ionospheric perturbations.

3 Results and discussion

The three GNSS techniques studied to measure the ionospheric scintillation are correlated to the seismic activity induced by the volcanic eruption. In each of the cases, instantaneous measurements of the S_4 index at every geographic coordinate have been recorded. To integrate this information into something comparable to the integrated energy dissipated by earthquakes, measurements were averaged and integrated into 6 h bins as the earthquake's energy ones. This 6 h period is long enough to include many measurements and reduce noise, but short enough to allow tracking possible variations within the day.

Figure 15 shows the comparison between the integrated energy generated by the earthquakes and the different methods to estimate the S_4 ionospheric scintillation index used around La Palma Island. Figure 15a shows the geomagnetic perturbations measured by the planetary index (K_p), and the solar activity represented by the solar flux F10.7. Figure 15b shows the integrated earthquake energy every 6 h interval over a yellow background indicating the time when the eruption was active. Figure 15c shows the ground station data obtained from LPAL and MAS1 stations, and then the percentile 95% computed every 6 h intervals. Figure 15d shows the GNSS-R data from CYGNSS, in the same 6 h intervals. And finally, Fig. 15e shows the GNSS-RO data. The percentile 95% of the S_4 values is shown in these plots after filtering them by distance to the eruption: 300 km in blue color and 1000 km in red color.
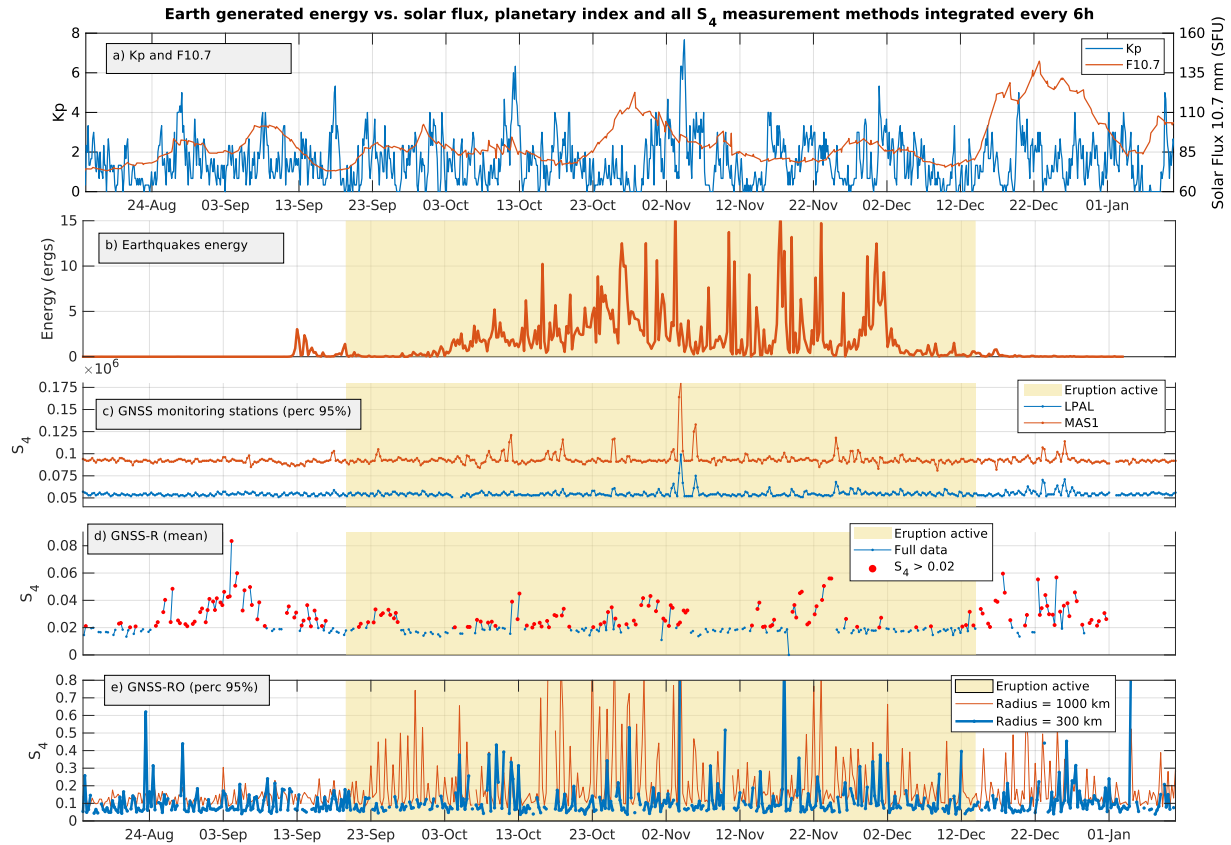
Figure 15. Comparison of a) geomagnetic planetary index and solar flux with b) integrated earthquake's energy every 6 h interval vs. estimated S_4 from the three sources: c) percentile 95 of S_4 obtained from ground stations in La Palma and Maspalomas, d) mean S_4 obtained from CYGNSS GNSS-R, and e) percentile 95% of S_4 obtained from GNSS-RO with two different filtering distances.

A first visual inspection of these data shows a certain correlation between the seismic energy and the estimated scintillation. For example, the largest peak in the seismic activity on 3 November at 9 h UTC almost matches with the peaks in the GNSS monitoring ground stations and GNSS-RO measurements, both at the 6-hours interval at 0 h UTC on 4 November. Similarly, the second largest peak in the seismic activity on 17 November at 15 h UTC, has a corresponding replica in the 300 km radius GNSS-RO measurements on 18 November 0 h UTC.

It can be observed that GNSS monitoring ground stations data presents an offset between LPAL and MAS1 stations, but they are highly correlated most of the time as both stations can sense the region of the ionosphere likely to be perturbed by the eruptive activity.

Figure 15d shows the CYGNSS GNSS-R data, which seems the most uncorrelated measurement. As mentioned previously, this can be explained as the GNSS-R reflectometry is affected by the sea surface conditions to correctly estimate the ionospheric

scintillation index. The red points in this Fig. 15d are the ones with an S_4 value larger than 0.02 to filter out the values that are more prone to be affected by sea roughness, as explained in Sect. 2.2.

Figure 15e shows the GNSS-RO data after being filtered by their distance to the eruption site: 1000 km in a red line, and 300 km in a blue line. The 1000 km one shows more and higher peaks during almost the whole period, which may be indicating that it is being affected by other sources of perturbations, different from the volcanic eruption.

To do a quantitative analysis between the different GNSS data sources, and allow a better comparison of them, a linear correlation between each pair of data (earthquakes energy vs. each of the GNSS measurement methods) has been performed. Before correlating each signal, they have been shifted by a certain amount of time from -10 days to +10 days, in steps of 6 h, equal to the sampling rate for all signals. This way we can also see if the impact of earthquakes in the ionosphere is a precursor or a consequence of it.

After the temporal shift, using the corresponding pair of points (S_4 vs. integrated earthquake energy), a least-squares linear correlation is computed, obtaining for each case its Pearson correlation coefficient, R. Then, for each shifted time, the values of R over time are plotted in Fig. 16. In all cases, the x-axis indicates the amount of time shifted, being negative when the scintillation is a precursor of the earthquakes, and the y-axis is the correlation coefficient R.
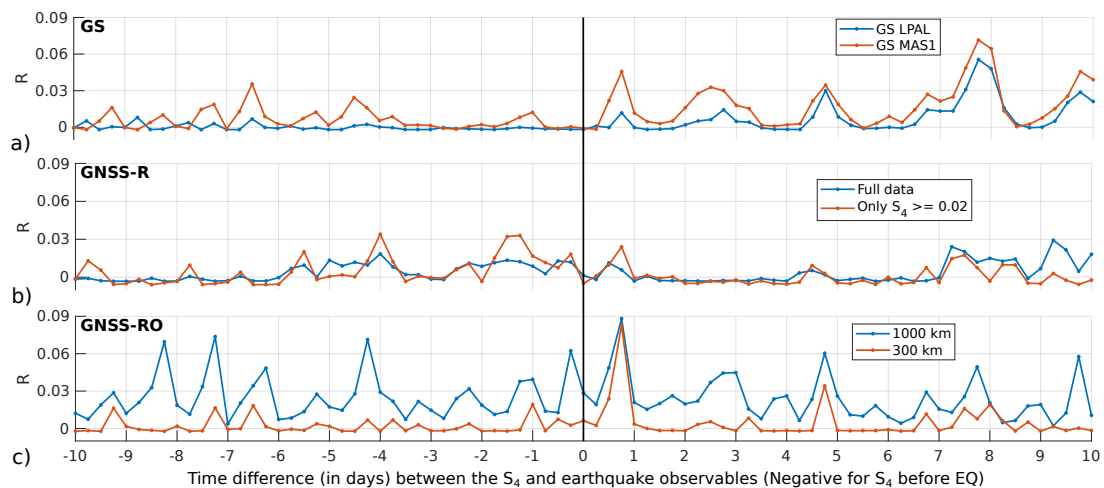
Figure 16. Results of the correlation analysis for each GNSS method and each shifted period in the x-axis: a) GNSS monitoring ground stations, b) GNSS-R data, and c) GNSS-RO data.

These results confirm some of the features commented qualitatively in the previous paragraphs. For example, Fig. 16b proves that the GNSS-R method presents the smallest correlation for all shifting times. Even though, it presents a weak correlation in several points, which tends to be larger when using only the S_4 values larger than 0.02. GNSS-R data present peaks of correlation from -7 to -4 days, also from -2 to +1, and a last set of peaks from +7 to +10.

For GNSS monitoring ground stations data in Fig. 16a the largest peaks of correlation occur when the scintillation is produced around 8 days later, and also 5 or 3 with smaller intensity. Additionally, there is a peak, which is bigger in Maspalomas than in the La Palma monitoring station, at +18 h from the eruption, which supports the fact observed previously in Fig. 15.

Figure 16a also shows that the Maspalomas (MAS1) station always has a higher correlation than the La Palma station, even if it is on the same island. One possible explanation for this is that the lower the elevation angle, the larger the detected scintillation is, increasing the possibility of detection of small ionospheric signatures. This is why larger correlations are found in the GNSS-RO measurements.

In the results for GNSS-RO in Fig. 16c it can be seen the largest correlation peak occurs at +18 h after the seismic activity, as for the GNSS monitoring ground stations data. It can also be seen that the data filtered by a 1000 km radius is noisier than the one filtered by a 300 km radius, because the last ones are more related to the eruptive activity than to other external causes. Even though, for the 1000 km radius curve, there are some correlation peaks when the seismic activity precedes the ionospheric perturbations in 8, 7, or 4 days approximately. Also, another peak is found at +4.75 days after the seismic activity, with a smaller replica in the 300 km radius, being the second larger peak for this case.

In Fig. 16, a cross-correlation between two or more techniques is observed sometimes. There are several points that present correlation peaks approximately at the same time interval, reinforcing the results presented and confirming that the different methods can actually detect the same signatures related to the seismic activity.

The same technique has been also applied to correlate the ionospheric scintillation to geomagnetic and space weather indicators: planetary index (K_p), and solar flux (F10.7), respectively. Figure. 17 shows the correlation analysis for the planetary index and the three methods to estimate the S_4: GS, GNSS-R, and GNSS-RO.

It can be observed a peak when the EQs occur from 1 to 2 days before the S_4 perturbations only in the ground station's method. These peaks reach a value of 0.09 for the MAS1 station, and 0.05 for the LPAL station. For the rest of the time windows, the correlation is almost null. The same happens for the GNSS-R and GNSS-RO results, where the correlation is always less than 0.03, and it can be considered negligible.

In Fig. 18 it is shown the result for the correlation between the solar weather variable, F10.7 and the S_4 from each method used. In this case, the solar flux is more correlated with the scintillation when using the GNSS-R method from CYGNSS data, with values up to 0.09. It is less correlated for the ground stations and the GNSS-RO method, being almost zero all the time for the last one.

These last two results indicates that there is a stronger correlation between the ionospheric scintillation and the earthquakes than with solar weather or geomagnetic perturbations. Only in the cases for the GNSS-R compared with the solar flux, and the ground stations with the planetary index, the influence of earthquakes sometimes appears to be smaller than from the space weather variables. This could mean that the influence of earthquakes, is, in general, locally a stronger factor to perturb the ionosphere. It is clear that the space weather during longer periods of time and wider regions, should be one of the main factors impacting the ionosphere. These results defend the idea that the local perturbations generated by an upcoming earthquake can be larger than the ones originated by the solar weather, and can be, therefore, detectable with instruments.
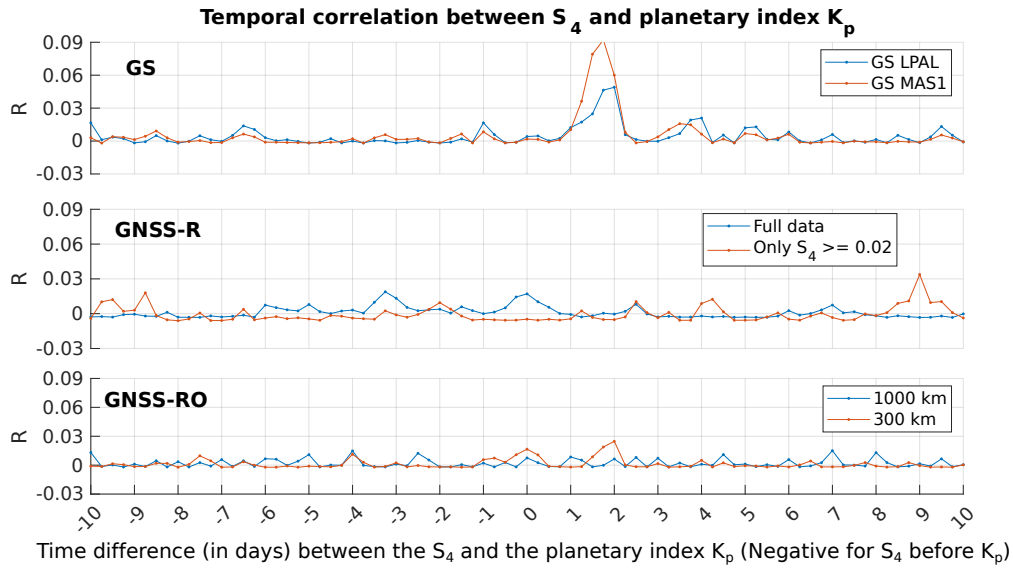
Figure 17. Correlation analysis between the planetary index (K_p) and S_4 from all GNSS methods: a) GNSS monitoring ground stations, b) GNSS-R data, and c) GNSS-RO data.
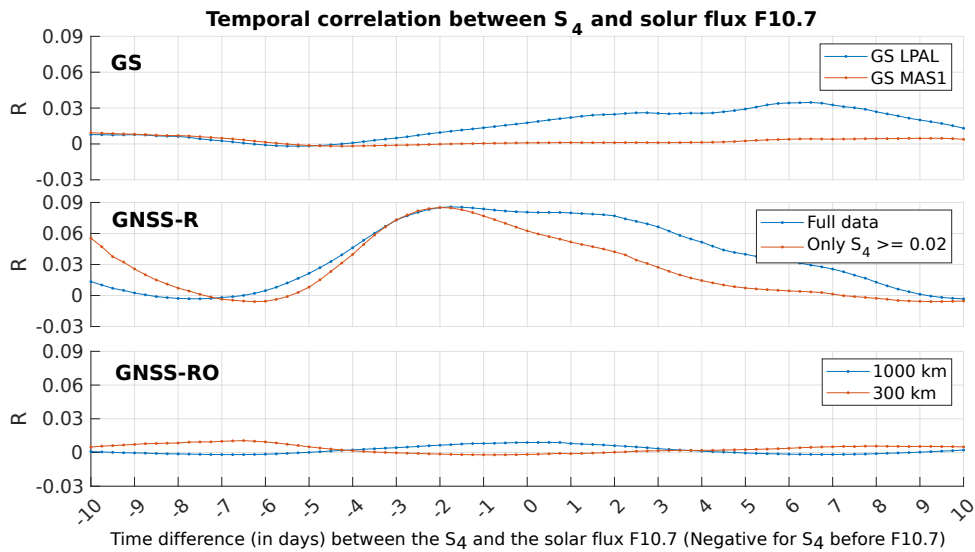
Figure 18. Correlation analysis between the solar flux (F10.7) and S_4 from all GNSS methods: a) GNSS monitoring ground stations, b) GNSS-R data, and c) GNSS-RO data.

4 Conclusions

An analysis of the impacts of the La Palma eruption on ionospheric scintillation has been made by correlating three different ways to measure it by using GNSS signals (GNSS monitoring ground stations, GNSS-R, and GNSS-RO) with the seismic activity related to the eruption.

This allows comparing the 3 methods' performance to detect tiny signals in the ionosphere produced by seismic activity. A detectable correlation for the GNSS monitoring ground stations, and the GNSS-RO methods is present. The GNSS-R technique resulted in the least favorable technique for this methodology, which may be explained by the rough sea state (SWH > 2 m) for the majority of the observed eruption period. The correlation peaks found in the other two methods are obtained after computing the data for the complete duration of the eruption, from 19 September to 13 December 2021. They show the largest peaks at +18 h after the seismic activity, with a correlation coefficient R of around 0.09 and 0.05 for GNSS-RO and GNSS-GS respectively. Even though they are not very high, they are detectable, and they can be related to the direct energy transfer from the earthquakes to the ionosphere by mechanical gravity waves (pressure waves coupling the atmosphere and the ionosphere), as other studies have reported in other eruptions.

Some correlation has also been found when the earthquakes occur some days before the ionospheric scintillation, mainly in the GNSS-R method, and in the GNSS-RO. In this last one, these peaks occurs in -4, -7 and -8 days.

In the case of the La Palma volcanic eruption, the pre-earthquake ionospheric perturbations may be produced through a piezoelectric effect caused by the severe seismic activity under La Palma during the whole duration of the eruption. As some studies affirm (Qian et al., 2001), the piezoelectric effect induced by the large underground rocks' pressure can induce electric charges in the surface generating perturbations in the ionosphere's electron density during some days before the earthquakes.

The general conclusion of this study is that the small correlation found between the earthquakes and the ionospheric scintillation using the Pearson coefficient makes very difficult to use these proxies, at present, in practical applications. This can be due to the small magnitude of the earthquakes associated with this volcanic eruption. Studying earthquakes with higher magnitudes the results may show clearer correlations. The importance of this study is that —to authors' knowledge— the ionospheric scintillation derived from three different GNSS-based methods, namely ground-based monitoring stations, GNSS-RO and GNSS-R, have been used to analyze ionospheric scintillation and their potential correlation with seismic activity, discarding other sources of geomagnetic activity or space weather events.

Code availability. Data analysis' code is not available. The analysis followed standard statistical routines and can be reproduced by the explanations given in the text.

Author contributions. All authors contributed conceptualization, data curation, formal analysis, visualization, and manuscript revision. Additionally, CM performed software implementation, and original draft writing, GGC contributed in software implementation, and AC was in charge of the project administration, funding acquisition, and resources provision.

Competing interests. The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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